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Origin of the Enhanced Photoluminescence Quantum Yield in MAPbBr₃ Perovskite with Reduced Crystal Size

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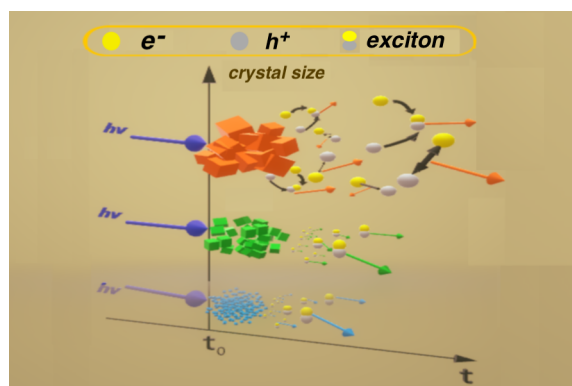
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Methylammonium lead bromide perovskite (MAPbBr₃) has been widely investigated for applications in visible perovskite light-emitting diodes (LEDs). Fine-tuning of the morphology and of the crystal size, from the microscale down to the quantum confinement regime, has been used to increase the photoluminescence (PL) quantum yield (QY). However, the physical processes underlying the PL emission of this perovskite remain unclear. Here, we elucidate the origin of the PL emission of polycrystalline MAPbBr₃ thin-films by different spectroscopic techniques. We estimate the exciton binding energy, the reduced exciton effective mass and the trap density. Moreover, we confirm the coexistence of free carriers and excitons, quantifying their relative population and mutual interaction over a broad range of excitation densities. Finally, the enhanced PLQY upon crystal size reduction to the micro- and nanometer scale in the presence of additives, is attributed to favored excitonic recombination together with reduced surface trapping thanks to efficient passivation by the additives.

TOC GRAPHICS



Hybrid organic-inorganic perovskites are emerging as highly promising new-generation energy materials. They are intensively studied for photovoltaic applications, thanks to their high absorption coefficients, low exciton binding energies and high carrier mobilities.^{1,2,3} In addition, aiming to reduce the global energetic consumption, perovskites with higher exciton binding energies, wider band gaps and narrow emission bandwidth are excellent candidates for energy-efficient photonic and optoelectronic applications, such as light-emitting diodes (LEDs).^{4,5,6,7,8} Methylammonium lead bromide (MAPbBr₃) is the most studied perovskite for LEDs due to its easy thin-film processing and its green emission. However, the low photoluminescence quantum yield (PLQY) of MAPbBr₃ limits its use in efficient light-emitting applications. The main strategy to foster the emission in this material has been the control over the crystal growth, in particular by limiting the grain size to the micro- and nanoscale.^{9,10,11} Despite growing knowledge about the perovskite morphology and its optoelectronic performance, the physical mechanisms governing the photoluminescence (PL) in MAPbBr₃ thin-films remain unclear. Both free carriers¹² and excitons¹³ have been suggested as the emissive species, but their interplay has been studied only at high carrier densities¹⁴ and recently at low temperature.¹⁵ Moreover, a systematic correlation between the reduction of the crystal size and the origin of the enhanced PLQY is missing. In this context, the dependence of the PL properties on the excitation density needs to be examined more exhaustively, since it brings insights about the origin of the PL.^{16,17}

Here, we elucidate the PL properties of solution-processed perovskite thin-films with different morphologies and crystal grain sizes by steady-state and time-resolved spectroscopy. Three different systems are compared: neat polycrystalline MAPbBr₃ (grain size of the order of ~1 μm), a polymer-MAPbBr₃ blend with crystal size in the tens to hundreds nanometer range (50 ± 25 nm, see SEM image Fig. S1),¹⁸ and a small molecule-MAPbBr₃ blend with crystal size in the weak quantum confinement regime (9.1 ± 1.6 nm, see TEM image Fig. S1).⁹ All the material preparation (compounds storage, solution preparation, thin film deposition) as well as the photophysical characterization is described in the SI and was carried out in inert atmosphere. For simplicity, the samples will be referred to as “polycrystalline perovskite”, “microcrystals” and “nanocrystals”. We show the co-existence of both free-carriers and

excitons in the polycrystalline perovskite and monitor their interaction over a broad range of excitation densities. We also find that the increased PLQY with crystal size reduction is due to a bright excitonic population even at low excitation density, together with reduced surface trapping thanks to passivation by the additives.

Figure 1(a) shows the steady-state PL and absorption spectra of the three MAPbBr₃ thin-films. Both the absorption onset and PL bands are blue-shifted as the size of the perovskite crystals decreases. The exciton Bohr radius of the bulk material has been reported to be 2 nm,¹⁹ while Malgras et al.²⁰ have observed a quantum confinement effect in MAPbBr₃ crystals with a diameter up to 7.1 nm. The average size of our nanocrystals, 9.1±1.6 nm, can thus justify the existence of weak quantum confinement and the subsequent blue-shift compared to the bigger crystals and grains. On the other hand, the blue-shift observed for the microcrystals cannot be attributed to quantum confinement.²¹ Two other possible explanations can be given. First, a change in the dielectric screening of the excitons induced by the presence of the additive (see Table S1 in the SI) could lead to the blue-shift.²² Second, according to Grancini et al.,²³ a stronger distortion in the Pb-I bond has been observed upon crystal size reduction in MAPbI₃, and D’Innocenzo et al. attributed the PL blue-shift in ~100 nm crystals to this strain effect.²⁴ The lack of a clear excitonic feature in the absorption spectra as well as the broad photoluminescence peaks, point to a rather broad distribution of particles size in the micro- and nanocrystalline materials (as confirmed in Fig. S1).

Figure 1(b) shows the time-resolved photoluminescence (TRPL) dynamics of the three samples under similar excitation density (as calculated in the SI). A difference in the PL dynamics obtained at different positions of the same sample has been reported in previous studies.^{25,26} In our study, the probed area was too large (8 mm laser spot, see Table S2 in SI) to distinguish any local variation of the excited state deactivation dynamics, and the results are an average over the microscale inhomogeneity. The PL decay for the micro- and nanostructured films is slower compared to the polycrystalline material, which, together with a higher oscillator strength for the emission, can contribute to the much higher PLQY (increase from 1.2% to 80%, Table 1). This trend contradicts previous reports of faster dynamics in nanocrystals,^{20,24} which we discuss below in terms of traps. The average PL lifetimes are reported in Table 1, together with the

radiative (k_{rad}) and non-radiative recombination ($k_{\text{n.rad}}$) rate constants. The increase of PLQY in the micro- and nanocrystals correlates with an increase in k_{rad} (higher oscillator strength) and a decrease in $k_{\text{n.rad}}$. Radiative processes include direct band-to-band recombination such as exciton recombination or free electron-hole recombination. Non-radiative recombination can occur if excitons or free carriers are trapped, via exciton-exciton annihilation or by three-body Auger processes involving free charges.

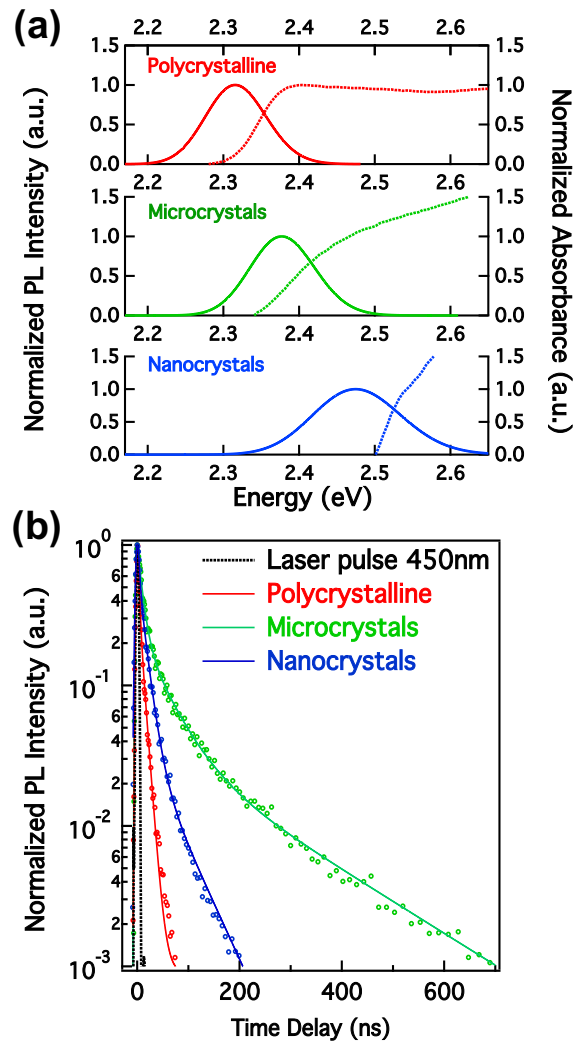


Figure 1. (a) Steady-state PL (left) and absorption spectra (right) and (b) TRPL dynamics of the three samples, recorded with excitation at 450 nm with 3 ns duration pulses and repetition rate of 10 Hz at a common carrier density of $\sim 10^{16}$ cm^{-3} . The detection is at the wavelength of the maximum PL intensity.

Table 1. Average PL lifetime, PLQY, radiative and non-radiative constants for the perovskite film series.

Sample	τ_{av} (ns)	PLQY	$k_{rad} \times 10^7$ (s ⁻¹)	$k_{n.rad.} \times 10^7$ (s ⁻¹)
Polycrystalline perovskite	5.4	1.2%	0.22	18.2
Microcrystals	21.4	80%	3.73	0.93
Nanocrystals	12.3	80%	6.5	1.62

The time-integrated PL spectra of the three samples were collected at different excitation densities, under excitation with pulses of 3 ns duration at 450 nm. Interestingly, the integrated PL spectra of the micro- and nanocrystals were found to be independent of the excitation density, while a clear change in the spectral shape was observed for the polycrystalline film (Figures 2a and S2). A possible explanation for this behavior might be re-absorption due to an increased penetration depth,^{27,28} but we exclude this here, as the thickness of the perovskite film is only 60 nm. Furthermore, a change in the PL spectra due to photo-degradation of the sample with increasing excitation density was excluded, since we showed repeatability of the observations when measuring the data again on the same sample spot at zero time delay t_0 (see Fig. S3). We also found that the emission properties do not change under continuous WL illumination (Fig. S4). The observed spectral changes in the polycrystalline perovskite include a shift, a spectral narrowing and a reduction of the PL intensity on the high-energy side of the spectrum at high excitation intensities. Analysis of the integrated PL spectra with a single Gaussian function was not satisfactory. The sum of two Gaussian functions (eq. 1) was needed to reproduce the experimental PL spectra, suggesting the co-existence of two emission bands due to excitons and free-charges (Figures 2a and S5).

$$I_{PL} = A \cdot e^{-\left(\frac{E-E_1}{w_1}\right)^2} + B \cdot e^{-\left(\frac{E-E_2}{w_2}\right)^2} \quad (1)$$

In eq. 1, A and B are the amplitudes, E_1 and E_2 the energies of the maximum PL intensity and w_1 and w_2 the widths of the low- and high-energy Gaussian, respectively. E_1 , E_2 and the width of the low-energy Gaussian (w_1) were kept constant, while A and B and the width of the high-energy Gaussian (w_2) were left as free parameters. w_1 was fixed to the value obtained by analyzing the excitonic peak present in the

absorption spectrum of the polycrystalline sample with a single Gaussian (78 meV), as shown in the inset of Figure 2(a). The very good agreement between the experimental low-energy PL band and the fit, as well as the small Stoke's shift of 49 meV, confirm the excitonic origin of this emission band.

The width of the high-energy Gaussian w_2 , which we assign to free carrier emission, was found to increase only after a certain carrier density. This behavior is characteristic of the Burstein-Moss shift due to band filling with free carriers upon photo-excitation at increasing carrier densities.^{29, 30, 31, 32, 33} Figure 2(b) shows the evolution of w_2 at different $n_o^{3/2}$, where n_o is the excitation density (Burstein-Moss analysis, eq. 2):

$$\Delta E_g^{BM} = \frac{\hbar^2}{2\mu^*} (3\pi^2 n_o)^{2/3} \quad (2)$$

In eq. 2, ΔE_g^{BM} is the change in the optical band gap, μ^* is the exciton reduced effective mass and \hbar is the reduced Planck constant.³² From the slope of the high excitation density part of Figure 2b, a value of $0.051 \cdot m_o$ for the reduced effective mass was extracted (m_o being the electron mass). This value is smaller but on the same order of magnitude compared to previous reports.^{34, 35, 36, 37, 38, 39, 40} According to Manser et al.³² and Christians et al.³³, the abrupt increase of w_2 that is observed at an excitation density of $7 \cdot 10^{15} \text{ cm}^{-3}$ ($n_o^{2/3} \sim 3.6 \cdot 10^{10} \text{ cm}^{-3}$) can be attributed to the process of trap filling (which competes with band filling at low excitation densities). This threshold can thus be used as a measure of the trap density in the material. The value is in agreement with the trap densities of polycrystalline perovskite films.⁴¹ The plateau at higher excitation densities ($> 10^{17} \text{ cm}^{-3}$) is likely related to band-gap renormalization, becoming more prominent due to the increased electron-electron interactions,^{42, 43} but also due to the decreased population of free-carriers. Indeed, the emission spectra measured at high excitation densities become increasingly dominated by the excitonic band (Figure 2a).

By dividing the amplitudes of the two Gaussians by their sum, the relative population of the species contributing to each of the two emission bands at different excitation densities was estimated (Figure 2c). Already at the lowest excitation density, almost 40% of the excited species are excitons. At higher excitation densities, the exciton population increases while the free carrier population decreases almost to zero. This is an indication of the favored conversion of free carriers to excitons at high excitation

densities, which results in the inversion of the majority population in the excited state. This conversion has been previously observed for MAPbBr₃,^{14,15} and other perovskite materials.⁴⁴ As a consequence, the energy difference between the maximum PL intensity of the two Gaussians can be used to estimate the exciton binding energy to 110 meV, in agreement with previously reported values for MAPbBr₃ (ranging between 35 and 320 meV, depending on the crystal size and the dimensionality).^{38,37,14,45,46} The value is further influenced by the chemical environment and dielectric constant of the surroundings.^{47, 48, 49}

Figure 2(d) shows the ratio of the squared free-carrier population over the population of excitons. The Saha equation, correlating this ratio of free carriers and excitons to the excitation density in hot electron-hole plasma, was used to model the obtained data:^{29, 5, 44, 50 51}

$$R = \frac{n_{e,h}^2}{n_x} = \frac{1}{n} \left(\frac{2\pi\mu^* k_B T}{h^2} \right)^{3/2} e^{-\frac{E_B}{k_B T}} \quad (3)$$

In eq. 3, $n_{e,h}$ is the population of free carriers and n_x is the population of excitons, both divided by their sum, which accounts for the total population. μ^* is the exciton reduced effective mass (estimated with the Burstein-Moss model), k_B the Boltzmann constant, h is the Planck constant, T the temperature in Kelvin and E_B the exciton binding energy of 110 meV. The calculation of the uncertainties of each parameter contributing to the error bars shown in Figure 2(d) is reported in the SI. The good agreement of the experimental data with the Saha model confirms the coexistence of free carriers and excitons in polycrystalline MAPbBr₃ thin films, even at low carrier densities. This is important, as this carrier concentration range corresponds to the working regime of optoelectronic devices, such as LEDs and solar cells.⁵²

In contrast to the polycrystalline perovskite, neither the microcrystals nor the nanocrystals show any variation in the shape of the integrated PL spectrum with increasing excitation density (Fig. S2). This indicates that only a single excited species, which we assign to excitons, governs the emission. The direct formation of excitons upon photo-excitation is expected to be enhanced for the nanocrystals compared to the polycrystalline perovskite, because of the weak quantum confinement. A larger exciton binding energy, that favors the formation of excitons upon photo-excitation, has also been observed for small

MAPbBr₃ crystals even outside the quantum confinement regime.⁵³ Changes in the dielectric environment due to the high volumetric fraction of the additive might enhance this effect (Table S1). This justifies an enhanced excitonic character for both the nano- and microcrystals. The higher electron-hole coupling for radiative excitonic recombination also explains the higher oscillator strength of the emission (higher k_{rad} , Table 1) compared to the polycrystalline perovskite, where both excitons and free charges contribute to the emission. In agreement, k_{rad} is highest in the nanocrystals (highest binding energy due to the weak confinement).

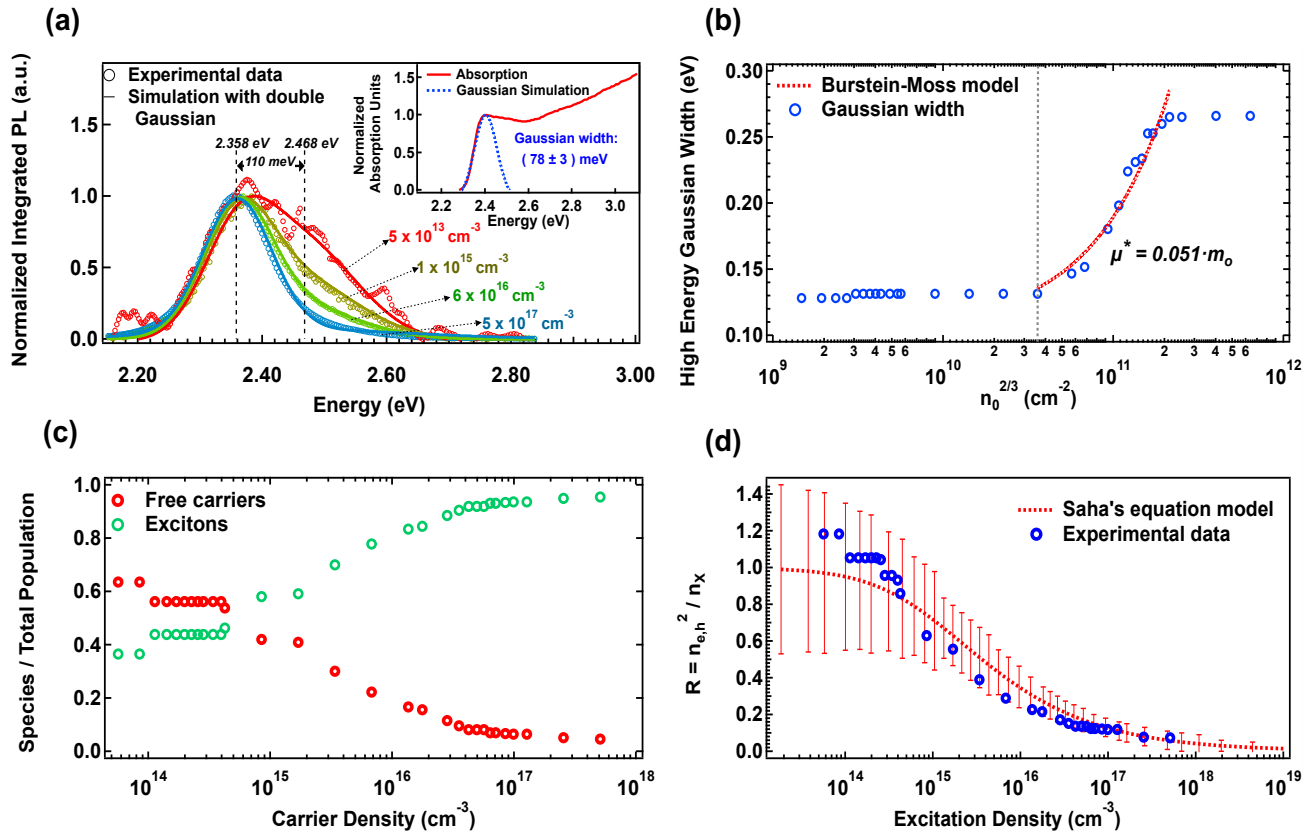


Figure 2. (a) Time-integrated PL spectra obtained under different excitation densities, and simulated by a double Gaussian function, for the polycrystalline perovskite film. The absorption of the same sample, with the excitonic peak analyzed by a single Gaussian function is shown in the inset. (b) Width of the high-energy Gaussian (w_2) versus $n_0^{2/3}$, analyzed by a linear equation. (c) Population of free-carriers and excitons over the total population of the excited state. (d) Ratio of the squared population of free carriers over that of the excitons at different carrier densities. A theoretical model, with our experimentally derived values implemented into the Saha equation, is shown on the same graph, together with error bars.

Further evidence that both excitons and free carriers are present in the polycrystalline perovskite, while excitons are predominant for the micro- and nanocrystal, is shown in Figure 3. This depicts the fluence-dependence of the excited-state population rate, which can be described by a polynomial function containing the generation rate G_0 and decay rates of different orders (k_1 , k_2 and k_3), see eq. 4.

$$-\frac{dn}{dt} = -G_0 + k_1n + k_2n^2 + k_3n^3 \quad (4)$$

Here, n is the density of either free carriers or of excitons, and it has been shown that the recombination terms in a free carrier semiconductor have different origins compared to excitonic materials.^{54,55,56} For free carriers, k_1 is the monomolecular decay indicative of free carrier trapping, k_2 is the bimolecular electron-hole emission, and k_3 is the three-body Auger recombination term.^{57, 58, 59} In the case of an excitonic semiconductor, k_1 represents the monomolecular excitonic and/or trap-assisted recombination, and k_2 the exciton-exciton annihilation.^{60,61,62,63} The third order term k_3 is absent since there are no free carriers formed upon photo-excitation,⁶³ although it might become relevant if excitons dissociate to free carriers.

By combining data from nanosecond-resolved TRPL with nano- and femtosecond transient absorption (TA) spectroscopy (450 nm excitation), dn/dt was obtained over a broad regime of excitation densities and time scales, as shown in Figure 3 (see SI and Fig. S6 for details). The number of polynomial terms necessary to describe the data for the different samples provides information about the recombination process. We found above that both excitons and free carriers co-exist in polycrystalline MAPbBr₃. In agreement, a complete description of the recombination rates can be obtained using three polynomial terms. For the micro- and nanocrystals containing only excitons, two polynomial terms are sufficient. The smaller $k_{\text{non-rad}}$ compared to k_{rad} (Table 1) suggests that the radiative recombination of the excitons is more efficient compared to monomolecular trapping, thus governing the monomolecular term (see discussion below), while exciton-exciton annihilation governs the bimolecular term. The rate constants that were derived by the aforementioned analysis are reported in Table 2.

The values obtained for the polycrystalline films are in agreement with values reported for MAPbBr₃ by Yang et al.,⁶⁴ but are higher than those observed by Richter et al.,¹² which we attribute to different preparation conditions. The higher monomolecular recombination rate we observe in MAPbBr₃ compared to free carrier perovskites such as MAPbI₃,^{52,57} can be explained considering that not only monomolecular trapping but also excitonic recombination takes place in this material (we have shown the co-existence of the two species). The value of k_1 is almost the same for all investigated samples. Either excitonic recombination or trapping is included in this monomolecular term. We will show that excitonic recombination becomes more important with the reduction of the crystal size, while trapping becomes less pronounced, resulting in the increased contribution of radiative recombination for the micro- and nanocrystals (Table 1). The bimolecular recombination constant, k_2 , is smaller for the microcrystals compared to the polycrystalline sample, as it reflects only exciton-exciton annihilation and not free carrier recombination. The larger k_2 for the nanocrystals compared to the microcrystals is due to the enhanced exciton-exciton annihilation because of their spatial confinement. Despite the increase in k_2 , it is still lower compared to other perovskite nanocrystals.⁶⁵ Importantly, the fact that the third order polynomial term, including the three-body Auger recombination, is not necessary for the micro- and nanocrystals, indicates that excitons created directly upon photo-excitation are stable and do not dissociate to free carriers even at high excitation fluencies. The generation of stable excitons was also shown by Sarritzu et al. in MAPbBr₃ crystals.¹⁵

Table 2. Polynomial coefficients obtained by the analysis of the data shown in Figure 3 by a second or third order polynomial function.

Sample	Polynomial Order	$k_1 \times 10^7 \text{ (s}^{-1}\text{)}$	$k_2 \times 10^{-10} \text{ (cm}^3\text{s}^{-1}\text{)}$	$k_3 \times 10^{-28} \text{ (cm}^6\text{s}^{-1}\text{)}$
Polycrystalline perovskite	3 rd	1.8 ± 0.8	1.0 ± 0.9	5.3 ± 0.2
Microcrystals	2 nd	1.6 ± 0.2	0.37 ± 0.02	--
Nanocrystals	2 nd	1.5 ± 0.2	1.3 ± 0.1	--

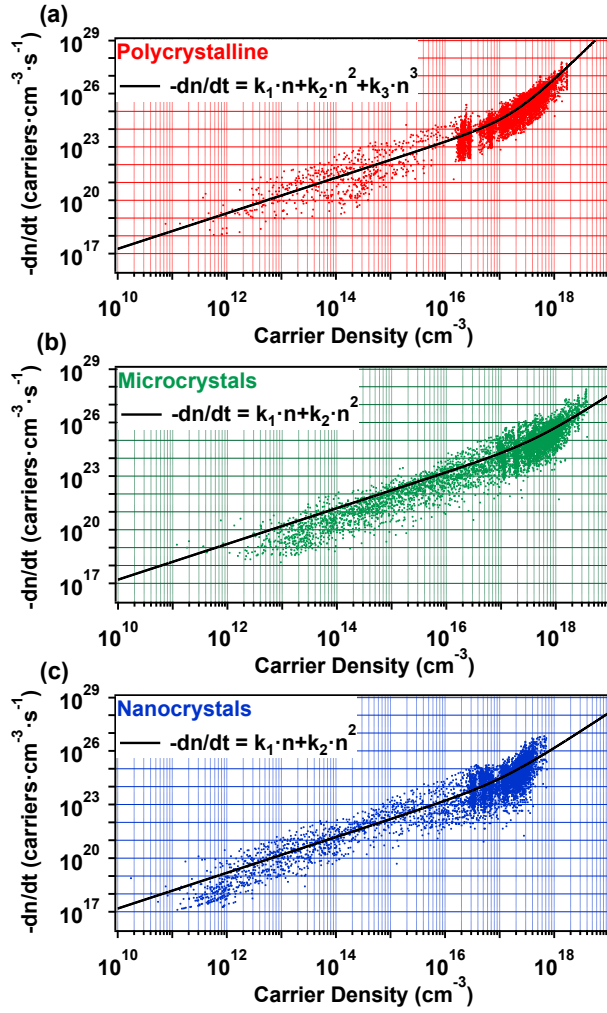


Figure 3. Population rate versus carrier density derived from ns-TRPL, ns-TA and fs-TA measurements for the three systems. The black line represents an analysis with a 3rd order polynomial for the case of the polycrystalline perovskite (a) and with a 2nd order polynomial in the case of the microcrystals (b) and the nanocrystals (c).

To further investigate the recombination mechanism in the MAPbBr₃ thin films, the time-integrated PL intensity (at the spectral maximum) was plotted as a function of the excitation density (Figure 4, only data below the threshold of exciton-exciton annihilation and Auger recombination is shown). The result was analyzed with a power law function of the form $PL \sim F^k$, where F is the excitation fluence or density, and k is a real number exponent, which provides information about the order of the recombination process.^{13,66,67}

In the polycrystalline perovskite, the value of 1.5 agrees with the co-existence of free carriers (bimolecular emission) and excitons (monomolecular emission), and with other measurements on perovskite thin-films.¹³ For the microcrystals and nanocrystals, the dependence of the integrated PL intensity on the excitation density becomes linear ($k \approx 1$), indicating predominantly monomolecular recombination. A similar behavior was observed when plotting PL_0 , the maximum of the PL decay trace at t_0 within our 3 ns time resolution, as a function of fluence (Fig. S3), confirming the above-mentioned results and the photostability of the samples (the PL_0 data was obtained with the same samples right after the time-integrated PL measurements). The monomolecular recombination in the micro- and nanocrystals could originate from either excitonic recombination (radiative) or trap-assisted recombination (non-radiative).⁶⁸ To distinguish between the two, the TRPL dynamics were recorded under different excitation densities.

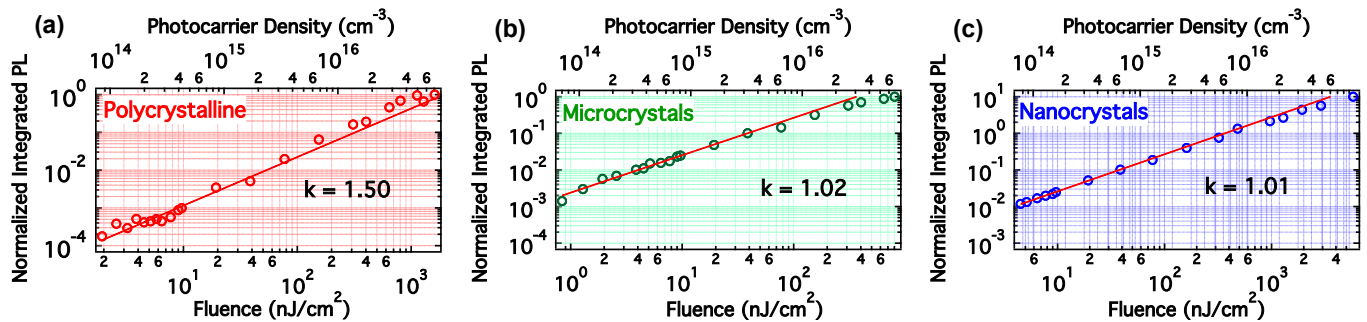


Figure 4. Graphs of the time-integrated PL intensity (at the maximum) versus the excitation density (or fluence) for (a) the polycrystalline perovskite, (b) the microcrystals and (c) the nanocrystals.

For the polycrystalline sample (Figure 5a), the dynamics become slower as the excitation density increases until a value of $7 \cdot 10^{15} \text{ cm}^{-3}$. According to the analysis with the Burstein-Moss equation (Figure 2b) this excitation density coincides with the one for which traps are filled. By further increasing the excitation density, the dynamics become faster again (inset of Figure 5a), due to higher order interactions such as Auger recombination. A similar trend has been observed in previous studies.^{69,59,70} We note that a non-reversible behavior of the TRPL dynamics when increasing and decreasing the excitation density has been reported due to photoinduced changes of the sample (e.g. photoinduced trap formation or ion

migration).^{71,72} We could exclude a significant contribution of such effects here, since the emission properties were reproducible and did not change in the presence of a continuous white light illumination (see Fig. S3, S4 and discussion in the SI). To explain the slower dynamics upon trap filling, we consider a distribution of trap states as suggested by Yamada et al.⁶⁹ and Wright et al.⁷³ Filling of the trap states that lie deep in the bandgap (and which cause non-radiative recombination) leaves only shallow traps unfilled. Chirvony et al. showed that carriers temporarily trapped in the shallow states can be thermally de-trapped, leading to delayed PL,⁷⁴ and explaining the long-lived (hundreds of ns) tail in the PL dynamics at the intermediate fluences. We have used the carrier density for which the inversion of the dynamics takes place to estimate a trap density of $4 \cdot 10^{14} \text{ cm}^{-3}$ for the microcrystals and of $5.9 \cdot 10^{13} \text{ cm}^{-3}$ for the nanocrystals (Figures 5b, c). These values are much smaller compared to polycrystalline MAPbBr₃, suggesting that these materials have little traps and are therefore suitable candidates for applications in LEDs.⁶⁸

The reduction in trap density between the polycrystalline sample and the films with reduced crystal size is consistent with the strong reduction of $k_{\text{nr,rad}}$ (Table 1). It shows that the monomolecular recombination in the latter samples is predominantly excitonic. It also explains the slower PL decay in the micro- and nanocrystals in Figure 2b. In the polycrystalline perovskite, the deep traps are filled at excitation densities where higher-order effects are already significantly present and reduce the overall excited-state lifetime. In contrast, delayed PL is seen more effectively at lower fluences in the systems with reduced crystal size. It is worth to mention that the reduction of the crystal size of MAPbBr₃ should in fact lead to an increase of surface trap density due to the higher surface to volume ratio of smaller structures.⁷⁵ However the additives present in the micro- and nanocrystals apparently passivate the surface traps, leading to the reduced trap density. This then enhances the PLQY by reducing $k_{\text{nr,rad}}$ and by favoring delayed PL. Finally, it is worth mentioning that the improved film morphology with the reduction of the crystal size could also result in better out-coupling of the emitted light, as shown by Richter et al.¹² Although this factor cannot be excluded, it would not affect the nature of the emissive state and the PL lifetime, as we report here.

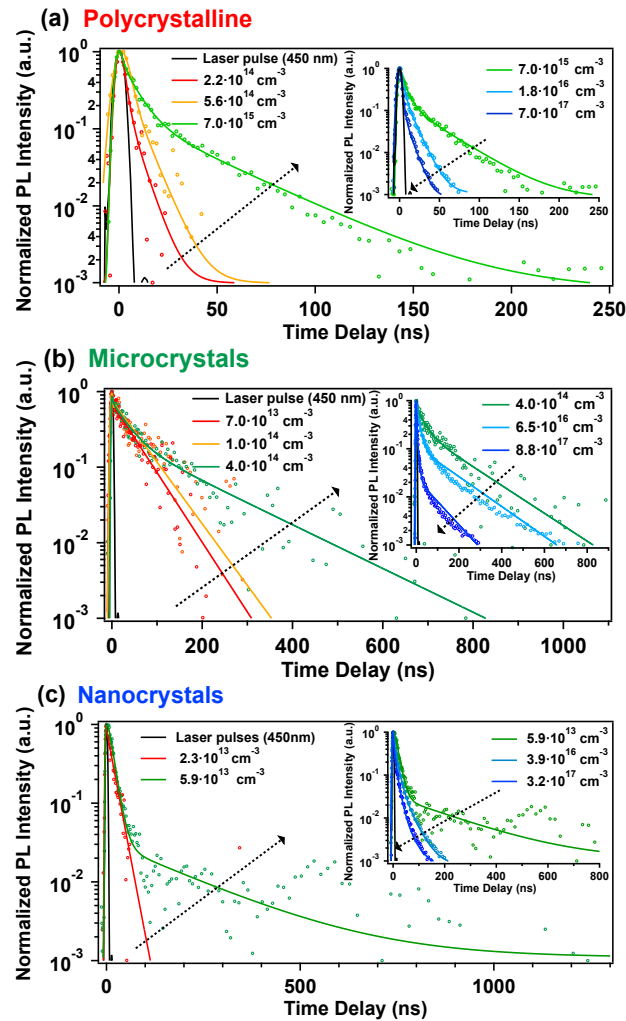


Figure 5. TRPL dynamics of (a) the polycrystalline perovskite, (b) the microcrystals and (c) the nanocrystals obtained at different excitation densities with excitation at 450 nm with pulses of 3 ns duration and a repetition rate of 10 Hz. The insets show the PL dynamics recorded at high excitation densities where the dependence of the lifetime on the excitation density is opposite than the one observed at low ones. The arrows indicate an increase in the excitation density.

In summary, the origin of the PL in polycrystalline MAPbBr₃, as well as in micrometer- and nanometer-sized perovskite crystals has been studied. By combining both time-integrated PL, time-resolved PL and transient absorption techniques, a broad range of excitation densities could be accessed. The results show that both free carriers and excitons co-exist in the polycrystalline perovskite, and that free carriers are converted into excitons at high excitation densities. On the other hand, in the micro- and nanocrystals,

directly photo-generated excitons are predominant. The contributions of non-radiative trapping and radiative exciton recombination upon crystal size reduction was studied. We found that the excitonic character is enhanced with the size reduction, while trapping is suppressed. Thus, the large increase of the PLQY for micro- and nanocrystals is attributed to a combination of increased oscillator strength of the radiative transition, due to the purely excitonic character of these materials, and of reduction of the surface traps by passivation induced by the additives.

ASSOCIATED CONTENT

Supporting Information.

Experimental details (sample preparation, steady-state absorption and PL techniques, ns-TRPL, ns- and fs-TA experiments), SEM and TEM pictures, calculation of the excitation densities, table with the content of additive in each sample and its ratio with the perovskite, time-integrated PL measurements, integrated PL spectra of the polycrystalline perovskite analyzed with a double Gaussian function, calculation of errors accompanying the Saha model and tables with the uncertainties of each parameter, graphs of the PL at different fluencies, graphs with the raw data from combined ns-TRPL/TA and fs-TA measurements and TRPL measurements under different conditions of illumination . (PDF)

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Notes

The authors declare no competing financial interests.

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