

1 **Travertine precipitation in the Paleoproterozoic Kuetsjärvi**
2 **Sedimentary Formation, Pechenga Greenstone Belt, NE**
3 **Fennoscandian Shield**

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14 **Abstract**

15

16 Precambrian travertines, tufas and speleothems either formed rarely or they have not been identified in previous
17 studies. In the absence of high pCO₂ soils in Paleoproterozoic, karst solution and speleothem formation occurred
18 by processes distinct from those commonly found in present-day low temperature karst environments. However,
19 the high pCO₂ atmosphere could itself have encouraged karst formation. The Paleoproterozoic Kuetsjärvi
20 Sedimentary Formation of the Pechenga Greenstone Belt, NW Russia, includes abundant terrestrial carbonate
21 precipitates. These precipitates were sampled from a drillcore representing a complete section of the ca. 120-m-
22 thick formation and were investigated for C and O isotopes, acid-soluble elemental contents and petrography.
23 The newly obtained results were used to constrain the origins of the precipitates and to illuminate different
24 terrestrial carbonate types. The investigated drillcore includes abundant small-scale cavities and veins, which are
25 commonly filled with dolomite and quartz. Dolomite crusts are found both in the cavities and on

26 bedding/erosional surfaces. Dolomite cements coat uneven surfaces and surficial rock fragments. The surficial
27 dolomite crusts form distinct and discrete layers, whereas the cements do not. The cavity and vein fills are likely
28 post-depositional in origin, whereas the surficial dolomite crusts and dolomite cements are likely syn-
29 depositional precipitates. The investigated precipitates often show $\delta^{13}\text{C}$ values lower than those reported from
30 their host rocks, suggesting the influence of an external carbon source. Petrographic features and geochemical
31 data suggest dissolution and precipitation of carbonate material originating from deep-sourced CO_2 -bearing
32 fluids, likely at high earth surface temperatures.

33

34 Keywords: Travertine, speleothem, carbonate, dissolution, degassing, Paleoproterozoic

35

36 **1. Introduction**

37

38 Reports of Precambrian travertines, tufas and speleothems are rare because they either rarely
39 formed or have not been identified (see Brasier, 2011; Brasier et al., 2013). Identifying these
40 deposits in drillcore can be problematic as examples may be mistaken for signs of post-
41 depositional alteration. Distinction between travertines, tufas and speleothems is also
42 challenging and perhaps unnecessary (e.g. Brasier, 2011; Rogerson et al., 2014).

43 Travertine and tufa have been classified in many ways. The word “tufa” was originally used
44 for both volcanic ash and soft, poorly consolidated freshwater carbonate, but later “calcareous
45 tufa” was used for freshwater carbonates (e.g. Pentecost, 1993; Pentecost and Viles, 1994).
46 Today, many sedimentologists use “calcareous tufa” or “tufa” for the softer varieties
47 (unsuitable for building), whereas the word “travertine” is used for harder freshwater
48 carbonates (e.g. Pentecost and Viles, 1994). However, travertine and tufa have also been

49 classified according to their fabrics, morphology, geochemistry and water temperature at time
50 of deposition (e.g. Ford and Pedley, 1996; Pentecost and Viles, 1994).

51 In some cases the word “travertine” is used for both thermal (above ambient) and ambient
52 water temperature chemical carbonate precipitates, and can be divided into (i) thermal or
53 thermogene travertines and (ii) meteogene travertines (e.g. Pentecost and Viles, 1994).
54 Thermal travertines are normally precipitated from hot waters and their carrier CO₂ primarily
55 originates from interaction of hot rock and CO₂-rich fluid (Pentecost and Viles, 1994). The
56 carried CO₂ for the precipitation of meteogene travertines is derived from soils and epigeal
57 atmosphere (Pentecost and Viles, 1994).

58 Alternatively, the word “travertine” can be used for thermal/hydrothermal precipitates and the
59 word “tufa” for cool/near-ambient water precipitates (e.g. Ford and Pedley, 1996; Pedley,
60 1990; Riding, 1991). In modern environments macrophytes can be used to distinguish
61 between travertine and tufa: travertines lack macrophyte remains, but tufas are commonly
62 characterized by macrophytes (e.g. Ford and Pedley, 1996). As there were no plants during
63 Precambrian, distinguishing between travertine and tufa is difficult or even impossible
64 (Brasier, 2011).

65 Modern speleothems are usually formed under high pCO₂ soils. These both provide carbonic
66 acid for carbonate bedrock dissolution and provide a contrast with the low pCO₂ of the cave
67 atmosphere that enables degassing of drip-waters causing carbonate mineral precipitation
68 (e.g. Fairchild et al., 2000; Frisia and Borsato, 2010). Paleoproterozoic karst and speleothem
69 probably formed in some other way, as there is no convincing evidence that high pCO₂ soils
70 existed at that time (see Brasier, 2011). A high pCO₂ atmosphere could have caused carbonic
71 acid formation and hence karst formation. Carbon dioxide degassing leading to precipitation
72 of carbonates in high pCO₂ atmospheres is possible in certain circumstances (see Brasier,

73 2011). One of the most appealing explanations for Paleoproterozoic karst and speleothem
74 formation is the common-ion effect, involving dissolution of Ca sulfate (e.g. Calaforra et al.,
75 2008; Wigley, 1973a) or incongruent dissolution of dolomite (e.g. Wigley, 1973b).
76 Explanations may alternatively involve deep-sourced CO₂ (e.g. Duliński et al., 1995;
77 Pentecost and Viles, 1994), possibly connected to hydrothermal activity.

78 In this paper, we use the term “travertine” in a collective way for the ancient deposits we
79 describe. The term “speleothem” is here used for cavity fills and the word “cement” for
80 dolomites binding or coating surface rock fragments and erosional surfaces.

81 Carbonate rocks extremely enriched in ¹³C were deposited world-wide during the
82 Paleoproterozoic (ca. 2200 – 2060 Ma), marking a global positive δ¹³C excursion recorded in
83 sedimentary carbonates (e.g. Baker and Fallick, 1989a, 1989b; Karhu and Holland, 1996).
84 The Kuetsjärvi Sedimentary Formation (KSF) of the Pechenga Greenstone Belt (NW Russia)
85 records this excursion (e.g. Karhu, 1993; Karhu and Melezhik, 1992; Melezhik et al., 2005,
86 Salminen et al., 2013a). The KSF exposures and cores are excellently preserved, and provide
87 an opportunity for studying Paleoproterozoic carbonate rock environments and precipitation
88 processes. In addition to stratified dolostone and limestone, hot-spring associated travertines
89 have previously been interpreted from the KSF (Melezhik and Fallick, 2001). Carbonates
90 associated with dissolution surfaces and small-scale cavities (epikarst) have also been found
91 in the KSF (Melezhik et al., 2004), along with calcrete (caliche) and dolocrete (e.g. Melezhik
92 et al. 2004; Melezhik and Fallick, 2003). These calcrete and dolocrete instances were
93 interpreted to have formed by capillary rise and evaporation (Melezhik et al., 2004; Melezhik
94 and Fallick, 2003). Moreover, Melezhik and Fallick (2005) reported probable CaSO₄
95 pseudomorphs in the KSF sabkha- or playa- carbonates.

96 In this study, discrete morphologies of carbonate precipitates were investigated from a
97 drillcore of the KSF. These included dolomite crusts, dolomite cements, and carbonate fills in
98 small-scale cavities and veins. A detailed petrographic picture of the precipitates was
99 constructed. Several samples were analyzed for the isotope composition of C and O and acid-
100 soluble abundances of selected elements. Results were compared to those of the more
101 common stratified dolostone and limestone rocks from the same drillcore (Salminen et al.,
102 2013a).

103 The goals of this research were to (1) identify and characterize different kinds of Precambrian
104 terrestrial carbonates, (2) decipher the origins of the investigated carbonate precipitates, and
105 (3) provide additional information on the depositional setting of the KSF.

106

107 **2. Geological background**

108

109 **2.1 Geological setting**

110

111 This investigation is based on the samples from Core 5A, which was drilled by ICDP
112 (International Scientific Continental Drilling Program) FAR-DEEP (Fennoscandia Arctic
113 Russia – Drilling Early Earth Project) from the KSF of the Pechenga Greenstone Belt, NW
114 Russia (**Fig. 1**). The geological setting of the core has previously been described by Salminen
115 et al. (2013a, 2013b) and is briefly summarized below.

116 (Approximate location of Figure 1.)

117 The Pechenga Greenstone Belt is a section of a larger (ca. 1000 km long) belt in the north-
118 eastern part of the Fennoscandian Shield (e.g. Melezhik and Sturt, 1994). This larger belt has

119 been interpreted as an intracontinental rift developed into an intercontinental rift with a
120 subsequent aborted oceanic phase and arc-continent collision (e.g. Melezhik and Sturt, 1994).
121 More extensive opening followed by oceanic floor subduction and arc-continent collision has
122 also been suggested (Berthelsen and Marker, 1986). The Pechenga Greenstone Belt has been
123 divided into the North and South Pechenga groups (e.g. Melezhik and Sturt, 1994). The KSF
124 belongs to the North Pechenga Group, which is composed of four paired sedimentary-
125 volcanic cycles (e.g. Melezhik and Sturt, 1994).

126 The Pechenga Greenstone Belt rocks underwent metamorphic alteration ranging from
127 prehnite-pumpellyite to amphibolite facies (Petrov and Voloshina, 1995). Core 5A was
128 drilled from biotite-actinolite phase of the greenschist facies. The KSF has also been exposed
129 to epigenetic alteration (Melezhik, 1992).

130 The thickness of the KSF varies from 20 to 120 m. Its thickness in Core 5A is ca. 117 m. The
131 KSF was deposited on a paleo-weathering crust developed on the basaltic andesites of the
132 Ahmalahti Formation and its depositional top is defined by the first basalts of the Kuetsjärvi
133 Volcanic Formation (Predovsky et al., 1974).

134 The minimum depositional age (U-Pb) of the KSF is 2058 ± 2 Ma (Melezhik et al., 2007),
135 inferred from detrital zircons in volcanoclastic conglomerate within the Kuetsjärvi Volcanic
136 Formation and the overlying Kolosjoki Sedimentary Formation. Martin et al. (2013) obtained
137 a depositional age (U-Pb) of 2056.6 ± 0.8 Ma for the Kolosjoki Sedimentary Formation. A
138 robust maximum depositional age (U-Pb) is 2505.1 ± 1.6 Ma, obtained from the Mount
139 Generalskaya gabbro-norite intrusion (Amelin et al., 1995). This intrusion is unconformably
140 overlain by basal conglomerate of the Neverskrukk Formation, which is the lowermost
141 formation of the North Pechenga Group.

142 Travertines of probable hot-spring origin were reported from the KSF by Melezhik and
143 Fallick (2001) and Melezhik et al. (2004). Melezhik and Fallick (2001) reported two types of
144 travertines from the KSF: (1) laminated crusts formed on a pure carbonate substrate and
145 capped by stratified, stromatolitic dolostone, and (2) small-scale mounds formed on a
146 carbonate substrate and buried under red siltstone and sandstone. In addition to travertines,
147 several other kind of subaerial exposure surfaces have been reported from the KSF (Melezhik
148 et al., 2004; Melezhik and Fallick, 2001, 2003). These include dissolution surfaces and
149 epikarst, erosional surfaces, calcrete/caliche, dolocrete, silica sinters. Dissolution surfaces and
150 epikarst included voids and small-scale cavities.

151

152 **2.2 Core 5A: lithostratigraphy and depositional setting**

153

154 A detailed lithostratigraphic description of Core 5A and an interpretation of its depositional
155 setting can be found in Salminen et al. (2013a, 2013b) with a brief summary provided below.
156 The lithostratigraphy of the core is presented in **Fig. 2**.

157 (Approximate location of Figure 2.)

158 Salminen et al. (2013b) divided the KSF in Core 5A into four informal members. From oldest
159 to youngest these are the Arkosic, Lower Dolostone, Quartzite and Upper Dolostone
160 members. A Fe-picrite dyke (ca. 5 m thick) separates the Arkosic and Lower Dolostone
161 members. Sedimentary carbonates of Drillcore 5A include micritic, sparry and stromatolitic
162 dolostone, dolarenite, micritic limestone and calcarenite. Siliciclastic rocks include arkosic to
163 quartzitic sandstone, graywacke, siltstone, shale and calcareous siliciclastic rocks (marl).

164 Both Lower and Upper Dolostone members show marks of erosional and dissolution
165 processes. Dolomite crusts, surficial cements and carbonate-filled veins and small dissolution
166 cavities are common (totaling at ca. 18 vol.%) in both of these members (see section 4; **Fig.**
167 **2**; Salminen et al., 2013a). Similar-looking dolomitic crusts and carbonate fills have
168 previously interpreted as hot-spring associated travertines (Melezhik et al., 2004; Melezhik
169 and Fallick, 2001).

170 The depositional setting of the Arkosic and Quartzite members has been assigned to delta-
171 plain, pro-delta, deltaic environments influenced by tides (Salminen et al., 2013b). The basal
172 part of the Quartzite member has been interpreted as a deposit from a flooding event
173 (Salminen et al., 2013b). The Lower Dolostone member was interpreted to have deposited in
174 a shallow-lacustrine setting and the Upper Dolostone member in a tidally influenced (hence
175 sporadically marine) shallow-lacustrine environment (Salminen et al., 2013b).

176

177

178

179 **3. Materials and methods**

180

181 Samples for this study were obtained from different depth-intervals of Core 5A. All analyses
182 were performed at the Department of Geosciences and Geography, University of Helsinki
183 using techniques described in Salminen et al. (2013a).

184 Subsamples were micro-drilled with 1.0 or 1.8 mm diamond grinding bits from carbonate
185 crusts, cements and fills. The number of subsamples from each carbonate type reflects its
186 abundance in the core. All carbonate subsamples were analyzed for the isotope composition

187 of C and O. The larger subsamples (those of ca. 10 – 20 mg; drilled with the 1.8 mm bit),
188 were also analyzed for acetic acid-soluble elements by ICP-MS (described below).

189 In addition, several subsamples were also obtained from the enclosing carbonate rocks,
190 immediately over- and underlying the cavity at ca. 62.1 m. These subsamples were only
191 analyzed for C and O isotopes, to compare/contrast with values obtained from the cavity fills.

192 Carbonate mineralogy (calcite or dolomite) of the larger subsamples was determined by
193 using Mg/Ca ratios, whereas in the 1–2 mg subsamples it was determined by XRD or through
194 a qualitative acid test (reaction with ca. 3 M HCl).

195 The acid-soluble Mg, Ca, Mn, Sr, Fe and Ba contents were determined using an Agilent
196 7500ce/cx ICP-MS. Powdered samples (9 –13 mg) were leached in 5 ml of ca. 0.5 M acetic-
197 acid under room temperature for 16 h. The final results were calculated in relation to the acid-
198 soluble fraction. Including the dissolution step, the average precision (1σ) was 14.6 % for the
199 samples from the depth 79.55 – 81.50 m (excluding the sample from 81.45 m) and 7.5 % for
200 all other samples.

201 The C and O isotope composition of the samples was determined using a Thermo Finnigan
202 Delta Plus Advantage isotope ratio mass spectrometer in continuous-flow mode. The isotope
203 compositions of ca. 125 – 170 μg powdered samples were measured on phosphoric acid
204 liberated CO_2 gas (in 90°C for $\geq 1\text{h}$). The C and O isotope compositions are expressed in
205 conventional δ -notations as a per-mil difference from the international VPDB standard. An
206 in-house dolomite quality standard indicated a long-term reproducibility (1σ) of 0.10‰ for C
207 and 0.22‰ for O (n=59). An in-house calcite standard indicated a long-term reproducibility
208 (1σ) of 0.08‰ for C and 0.17‰ for O (n=7).

209 Petrography and mineralogy of the carbonate precipitates were investigated in several thin
210 sections under optical microscope and by visual examination of drillcore samples. In
211 addition, selected thin sections were investigated semi-quantitatively (EDS, standardless)
212 with an electron microanalyzer (Jeol JXA-8600 superprobe, accelerating voltage 15 kV,
213 beam current 1 nA, focused beam).

214

215 **4. Petrography of the precipitates: description and interpretation**

216

217 Core 5A contains five distinct features. These are (1) cavities, (2) veins, (3) dolomite crusts
218 filling cavities, (4) dolomite crusts precipitated on bedding/erosional surfaces, and (5)
219 dolomite cements encrusting surface rock fragments and uneven surfaces. New petrographic
220 observations and interpretations of these allow us to add substantially to the previously made
221 brief descriptions of the Core 5A rocks (Salminen et al., 2013a, b; Brasier et al., 2013). The
222 general description of the five types of precipitates is provided in Table 1 and the investigated
223 samples are described in Inline Supplementary Table 1.

224 (Approximate location of Table 1.)

225 Insert Inline Supplementary Table 1 here.

226

227 **4.1 Small-scale dissolution cavities (type 1)**

228

229 **Description.** Dissolution cavities are found in all types of carbonate. The dolostone
230 immediately surrounding the cavity often appears recrystallized or powdery. The heights of

231 the cavities vary from a few centimeters to ca. ten centimeters. Cavities are walled with
232 rim(s) of fibrous (radiating), bladed and/or sparitic dolomite; some cavities also contain
233 calcite fills. Carbonate fills show both upwards and downwards growth. Small white
234 stalactites and stalagmites are found in places, e.g. at ca. 60.6 m. These stalactites/stalagmites
235 have concentric bedding-parallel shapes and they are composed of sparitic dolomite, walled
236 with a rim of bladed dolomite. The latest infill is usually quartz and/or dolospar (**Figs. 3-4**),
237 but in places it is quartz sandstone. Pyrite (**Fig 3b**) and dolomite ooids have also been
238 identified in some cavities.

239 Some cavities are found in possible surficial dolomite crusts. These include e.g. the cavity at
240 ca. 60.2 m. This cavity is walled with small white stalactites and stalagmites (**Fig. 5a**), similar
241 to those in the cavity at ca. 60.6 m. The remaining space of the cavity is filled with quartz and
242 patches of fibrous/sparitic dolomite. The cavity is underlain by red, beige and white, banded
243 dolomite crust and overlain by white dolomite crust and a possible silica sinter.

244 **Interpretation.** Most of the dolomite rims in the cavities are best interpreted as flowstones.
245 However, the small-scale stalactites and stalagmites probably indicate precipitation from
246 dripping water. Dolomite rims are relatively symmetrically walling the cavities. These
247 symmetrically rimmed cavities are interpreted to have been formed in dominantly phreatic
248 conditions.

249

250 (Approximate location of Figures 3-5.)

251 *4.1.1 Special small-scale dissolution cavity cases*

252

253 Cavities at ca. 69.6 m and 72.7 m appear somewhat different to other examples and are
254 therefore described here separately.

255 *The dissolution cavity at 69.6 m* has a complex structure, exhibiting three types of carbonate
256 fill (**Fig. 3b**). The outermost zone includes thin laminae of white dolomite and quartz. The
257 second zone includes yellowish, sparitic calcite, which is followed by thin laminae of white
258 dolomite and quartz. These laminae also form microstalactite-like structures. These laminae
259 are sealed with a thin rim of fibrous dolomite. The remaining space, in the center of the
260 cavity, is filled with quartz, dolospar and pyrite. These three fabrics represent the latest infills
261 of the cavity. The precipitates of this cavity, like the rim of fibrous dolomite, are best
262 interpreted to have precipitated in phreatic to vadose conditions.

263 *The cavity at 72.7 m* is filled with siliceous sparitic dolomite with quartz-filled fenestrae-like
264 structures. (**Fig. 4b**). The morphologies of some of these fenestrae resemble sulfate
265 pseudomorphs. The cavity ceiling is composed of white, thinly-laminated, sparitic dolomite,
266 which also forms stalactite-like structures. Similar laminae are not found at the base of the
267 cavity. The non-symmetrical structure of the cavity likely indicates carbonate mineral
268 precipitation in the vadose zone.

269

270 **4.2 Veins (type 2)**

271

272 **Description.** Some structures likely represent veins, rather than filled cavities. These
273 structures include thin (ca. 0.2–2 cm), bedding-parallel vugs (**Figs. 6a-b**). Some of these have
274 relatively straight borders, whereas others have more curved margins. Veins are walled with
275 rims of sparitic and bladed dolomite. the bladed dolomite shows inward growth. Quartz
276 and/or dolospar represent the latest infills; some veins are almost entirely filled with dolospar.

277 **Interpretation.** These veins have symmetrical shapes, indicating mineral precipitation within
278 a fluid-filled space (e.g. fracture). In some veins, the flooring is thicker than the ceiling. This
279 could indicate precipitation within a gas- and fluid-filled space.

280 (Approximate location of Figure 6.)

281

282 *4.2.1 Special vein case*

283

284 A vein at ca. 53.7 m shows color and mineralogical zoning (**Fig. 6c**). This vein is larger (with
285 a height of ca. 8-9 centimeters) than other veins in the core. Its outer rim is composed of
286 white dolomite, whereas the outer rim is capped by a grayish calcitic and dolomitic layer. The
287 center of the vein is filled with orange calcite. The color zoning is more pronounced here than
288 in the cavities. Pyrite is also found in this vein. This vein may have formed during burial
289 diagenesis.

290 **4.3 Large cavities with dolomite crust (type 3)**

291

292 **Description.** The cavities with dolomite crusts are up to ca. 20 cm thick. The flooring in
293 these cavities is thicker than the ceiling. The remaining space between the flooring and the
294 ceiling is filled with dolospar, but cavities often appear as collapsed or deformed. The
295 flooring (red, pink, beige, white, and/or gray) usually consists of banded, laminated or
296 massive dolomitic crust, which is capped with thinly-laminated gray-and-white crust. In
297 places the ceiling is composed of a dolomitic crust (white, gray or yellow) underlain by small
298 thinly-laminated gray-and-white stalactites (**Fig. 5b**), but elsewhere the crust is lacking and
299 there are only stalactites. The thinly-laminated crusts and stalactites are often deformed (**Figs.**

300 **5b).** They usually consist of alternating white and/or gray laminae of fibrous, radiating
301 dolomite (forming cones) and sparitic dolomite.

302 **Interpretation.** The cavities with dolomite crusts have probably been deposited in vadose
303 conditions. These cavities have non-symmetrical structure; the cavity flooring is thicker than
304 the ceiling coating. The basal part of the flooring is often characterized by red coloration,
305 which possibly implies deposition from an oxidizing fluid.

306

307 *4.3.1 Special large cavity with dolomite crust case*

308

309 In the cavity at ca. 31.4 m (**Fig. 5c**), the entire flooring is thinly laminated and also includes
310 dolospar and quartz filled fenestrae. The stalactites in the cavity consist of laminae of small
311 (micritic) grain size; the laminae are coloured white, gray, black or reddish (**Fig. 5c**). This
312 cavity also shows non-symmetrical structures so its fill probably precipitated in the vadose
313 zone.

314

315 **4.4 Surficial dolomite crusts (type 4)**

316

317 **General description and interpretation.** Surficial dolomite crusts were formed on bedding
318 and erosional surfaces. The crusts are composed of bedding parallel layers with uneven upper
319 surfaces and contain curved or mound-shaped laminae. Hence, the petrography and
320 morphology of these carbonate precipitates differs from normal stratified dolostone and
321 limestone of the KSF. We tentatively interpret such crusts as authigenic precipitates in

322 surface settings. The red coloration at the base of some crusts may imply deposition from an
323 oxidizing fluid. Below, different types of the surficial crusts are described in more detail.

324 **Crust at 30.3 m.** One inferred surficial dolomite crust in the Upper Dolostone member (at ca.
325 30.3 m, **Inline Supplementary Figure 1**) shows a reddish lamination at its base, followed by
326 beige laminated, banded or massive dolomite. This crust includes small vugs/veins, which are
327 filled with white, grey and black sparry carbonate. The upper part of the crust **has a chalky**
328 **micritic appearance. The crust is capped with a rim of fibrous dolomite and quartz.**

329 **(Insert Inline Supplementary Figure 1 here)**

330 **White dolomite crusts.** There are several inferred surficial dolomitic crusts composed
331 mainly of white micritic to sparitic dolomite. The crust at ca. 68.3 m shows thin lamination
332 (**Fig. 7**). These laminae form mound-like or oncoidal structures. These crusts are capped by a
333 thin (<1 cm), upwards-growing rim of fibrous dolomite which is sealed with a crystalline
334 silica sinter (**Fig. 7**). The crust at 68.3 m is overlain by a sandstone layer (**Fig. 7**).

335 **Crust at 27.4 m.** The layer at ca. 27.4 m has a complex internal structure (**Fig. 8**). Its basal
336 part is composed of beige, laminated, micritic dolomite and contains some quartz-filled
337 fenestrae. The basal part is overlain by a partially vuggy interval. The latter contains several
338 small vugs, which are surrounded by beige, laminated, micritic dolomite and grayish
339 (siliceous), structureless, sparitic dolomite. The vugs are walled by white (inner) and yellow
340 (outer) laminae of fibrous dolomite. The remaining space in the vugs is filled with dolospar
341 and quartz. The yellow and white laminae also form small (ca. 1 cm) mounds. The vuggy
342 interval is sealed with a rim of white sparitic dolomite. These vugs probably formed
343 simultaneously with the surrounding dolomites. The uneven surface of the entire layer is
344 capped with a thin siliceous film (interpreted as a sinter or a silcrete layer). The overall
345 structural pattern suggests a surficial dolomite crust.

346 (Approximate location of Figures 7-8.)

347

348 **4.5 Dolomite cements (type 5)**

349

350 Dolomite cements are found in the Lower Dolostone member (**Fig. 9**), chiefly at ca. 76-78 m
351 and at 83 m. These cements consist of white feather-shaped (Figs. 9a-c) , fibrous, radial,
352 bladed and/or sparitic dolomite. Dolomite cements do not form distinct bedding-parallel
353 layers (in contrast to the surficial dolomite crusts), but they are found on uneven surfaces and
354 capping or corrosively replacing sandstone fragments. Sometimes they form mound-like
355 accumulations (**Figs. 9a-b**), which often have exhumed upper surfaces. There are also thinly-
356 laminated cements (**Fig. 9b**). Dolomite cements usually show upward growth and appear as
357 syn-depositional precipitates. Sometimes they are also found filling space between the
358 sandstone fragments, showing downward growth (**Fig. 9d**); quartz or dolospar is the latest
359 infill.

360 (Approximate location of Figure 9.)

361 Although the dolomite cements and quartz also fill void spaces between sandstone fragments,
362 the cements are best interpreted as syn-depositional precipitates, as these precipitates are
363 usually cementing surface rock fragments and uneven erosional surfaces.

364

365 **5. Geochemical results**

366

367 **5.1 Electron microprobe data**

368

369 Carbonate cavity fills, crusts and cements were analyzed semi-quantitatively with an electron
370 microanalyzer to determine mineralogies. The reddish dolomite crust in the cavity flooring at
371 ca. 81.5 m includes mainly dolomite but also calcite, quartz, talc, phlogopite, muscovite, Fe-
372 poor chlorite, potassium feldspar and fluorapatite. Calcite was found filling the remaining
373 space between dolomite crystals. Dolomite and minor calcite and talc were identified in the
374 cavity at ca. 60.6 m. The quartz-filled fenestrae in the layer at ca. 72.7 m contain calcite and
375 dolomite inclusions. Apatite and potassium feldspar were also identified in that layer. No
376 traces of evaporite minerals were found in the investigated fenestrae. Two samples from the
377 dolomite cements are composed of mainly dolomite with minor calcite, quartz and talc.
378 Several minor minerals were also identified in the dolomite cements, including fluorapatite,
379 potassium feldspar and chlorite. Possible sulfate pseudomorphs (now dolomite and calcite)
380 were recognized in a secondary electron image (**Inline supplementary figure 2**) in the
381 dolomite cement at ca. 76.9 m.

382 (Insert Inline Supplementary Figure 2 here.)

383 Both the electron microprobe and the microscope investigations suggest that dolomite is a
384 primary mineral. There are no petrographic indications for dolomitization. Instead, the
385 presence of talc could indicate that the minor calcite has been formed via de-dolomitization.

386

387 **5.2 Acid-soluble elemental contents (ICP-MS)**

388

389 The acid-soluble elemental contents and the Mg/Ca and Mn/Sr ratios of the carbonate fills,
390 crusts and cements are presented in **Inline Supplementary Table 2**.

391 *(Insert Inline Supplementary Table 2 here.)*

392

393 The Mg/Ca ratios vary from 0.01 to 0.71 (n=65). Based on the Mg/Ca ratios, most of the
394 subsamples were dolomitic. Two subsamples were calcitic. Two subsamples were mixtures
395 of dolomite and calcite.

396 On the core scale, the Mn and Fe contents show subtle upwards decreasing trends. The Mn
397 contents vary from 42 to 3900 ppm (average 640 ppm, n = 65, **Fig. 10**). The abundances of
398 Fe range between 110 and 9600 ppm (average 1600 ppm, n = 65, **Fig.10**). The Fe
399 concentrations correlate strongly with the Mn concentrations ($r = 0.57$, n = 65, **Fig. 11**).

400 The Sr contents show an upwards increasing trend in the core, varying from 30 to 310 ppm
401 (average 96 ppm, n=65, **Fig. 10**). The Sr and Mn contents do not show co-variation in the
402 investigated carbonate precipitates (**Fig. 11**). The Ba contents do not show any clear
403 stratigraphic trend, varying from <0.48 to 51 ppm (average 6.2 ppm, n=65, **Fig. 10**).

404 *(Approximate location of Figures 10-11.)*

405

406 *5.2.1 Detailed investigations of distinct cavities and crusts*

407

408 Several subsamples from the cavities at ca. 79.6 m and ca. 81.4 m were analyzed for
409 elemental contents. In the cavity at 79.6 m (**Fig. 12**), the highest Mn and Fe contents were
410 obtained from a stalactite and stalagmite in the cavity center. The Fe contents show a subtle

411 upwards increase in the upper part of the cavity (in the ceiling). The Mn, Sr and Ba
412 concentrations do not exhibit any clear trends. In the cavity at 81.4 m (**Fig. 13**), the Mn and
413 Fe contents show generally upwards increasing (cyclical) trends. The Sr concentration is
414 variable, demonstrating no clear trend across the cavity, whereas the Ba content shows a
415 subtle upwards decrease.

416 (Approximate location of Figures 12-13.)

417 The vein at ca. 53.7 m shows mineralogical and color zoning. The orange infill in the center
418 is composed of calcite ($Mg/Ca = 0.01$), whereas the grayish rim surrounding the orange infill
419 has a mixed dolomite-calcite composition ($Mg/Ca = 0.20$). The white outermost rim consists
420 of dolomite ($Mg/Ca = 0.64$). The Mn and Ba concentrations decrease from the calcitic infill
421 towards the dolomitic rim. The highest Fe and Sr contents were obtained from the mixed
422 calcite-dolomite sample. All analyzed carbonate phases have a high Mn content (750–2700
423 ppm), whereas the Sr abundances (62–150 ppm) are higher or similar to those obtained from
424 other carbonate phases from similar depths.

425 The cavity at ca. 72.7 m includes silica filled fenestrae. As mentioned earlier, some of these
426 fenestrae morphologically resemble sulfate pseudomorphs, although no traces of sulfates
427 have been detected. High Mn (3900 ppm) and Fe (9600 ppm) contents were measured from
428 this interval. In contrast, the Sr (35 ppm) and Ba (0.82 ppm) contents are relatively low.

429

430

431 **5.3 Carbonate C and O stable isotopes**

432

433 The $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values of all analyzed samples are presented in **Inline Supplementary**
434 **Table 2.**

435 In cavity fills, crusts and cements (n = 139), the $\delta^{13}\text{C}$ values vary from 4.0 to 8.0‰ averaging
436 at 6.8‰ (**Fig. 2**). The $\delta^{18}\text{O}$ values range from -15.1 to -8.4‰, with an average value of -
437 12.6‰ (**Fig. 2**). In general, cavity/vein fills, surficial crusts and cements show similar
438 variations of $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values.

439 The $\delta^{13}\text{C}$ and the $\delta^{18}\text{O}$ values show moderate positive correlation ($r = 0.50$, $n = 139$, **Fig. 11**)
440 in the carbonate fills, crusts and cements. Lower $\delta^{13}\text{C}$ values are often linked to higher Mn
441 concentrations (**Fig. 11**). The Sr contents of the precipitates do not show correlation with the
442 $\delta^{13}\text{C}$ or $\delta^{18}\text{O}$ values (**Fig. 11**). No significant co-variation is observed between the $\delta^{18}\text{O}$
443 values and the Mn contents (**Fig. 11**).

444 *5.3.1 Detailed investigations of distinct cavities and crusts*

445

446 A series of subsamples was obtained from some cavity fills. In individual crusts and cavities,
447 the $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values show moderate or even strong positive correlation (**Fig. 14**). Clearly
448 decreasing $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ trends from the walls towards the centers were found in the cavities
449 at ca. 83.0 m (**Fig. 15**) and ca. 79.6 m (**Fig. 12**). In the cavity at ca. 81.4 m (**Fig. 13**), the
450 lowest $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values were obtained from the upper part, which included thinly-
451 laminated gray-and-white crust and stalactites. The $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values show subtle
452 upwards decreasing trends in the flooring of that cavity, followed by subtle upwards
453 increasing $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ trends.

454 *(Approximate location of Figures 14-15.)*

455 The cavity fills at ca. 62.1 m show similar or lower $\delta^{13}\text{C}$ values (6.1 – 7.5‰, average 7.0‰)
456 than those (7.3 – 7.8‰, average 7.6‰) in the carbonates enclosing the cavity (**Fig. 16**).

457 These rocks over- and underlying the cavity show a subtle, symmetrical decrease in their $\delta^{13}\text{C}$
458 values towards the cavity. The $\delta^{18}\text{O}$ values are relatively similar or slightly higher in the
459 cavity fills (from -12.5 to -9.2‰, average -11.5‰) than in the rocks enclosing it (from -12.3
460 to -11.3‰, average -11.9‰) (**Fig. 16**).

461 (Approximate location of Figure 16.)

462 In the surficial crust at ca. 30.3 m (**Fig. 17**), the $\delta^{13}\text{C}$ values first show a cyclic, upwards
463 decreasing trend (in the lowermost third of the crust), followed by a subtle tendency towards
464 higher values. The $\delta^{18}\text{O}$ values show an increasing trend in the lowermost part of the crust,
465 followed by a decreasing trend, and after that, variable $\delta^{18}\text{O}$ values. The upper half of the
466 surficial dolomite crust at 68.3 m shows a small (ca. 0.5‰) increase in $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$
467 upwards in the section (**Fig. 18**); the lower half of the crust was not investigated. These crusts
468 at 30.22 m and 68.26 m are different in their petrography.

469 (Approximate location of Figures 17-18.)

470 Carbonates in the vein at 53.7 m vary in their $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values. The calcitic phase (from
471 the central orange zone of the vein) shows a $\delta^{13}\text{C}$ value of 5.2‰ and a $\delta^{18}\text{O}$ value of -12.7‰.
472 The subsample representing a mixture of calcite and dolomite (from the grayish zone) shows
473 a $\delta^{13}\text{C}$ value of 5.9‰ and a $\delta^{18}\text{O}$ value of -13.5‰. The dolomitic subsample (from the white
474 zone) shows a value of 7.6‰ and a $\delta^{18}\text{O}$ value of -12.0‰. Thus, the $\delta^{13}\text{C}$ values increase
475 from the central calcitic zone outwards, towards the dolomitic zone. In contrast, the lowest
476 $\delta^{18}\text{O}$ value was obtained from the mixture of dolomite and calcite.

477

478 6. Discussion

479

6.1 Petrographic aspects

481

482 Many of the precipitates studied in Core 5A are post-depositional phases, including cavity
483 and vein fills (types 1-3). The term speleothem could be appropriate for the cavity fills.
484 Whether the waters from which they precipitated were thermal (e.g. Corbella et al., 2004;
485 Djidi et al., 2008; Dublyansky, 1995) or at ambient temperature remains to be substantiated.

486 The thinly-laminated crusts and stalactites in the type 3 cavities (**Fig. 5a, c**) are similar to
487 carbonate crusts coating modern travertines (e.g. Jones and Renaut, 2008). Moreover, some
488 of the studied carbonates (e.g. type 4) resemble the ~1.8-Ga-old Baker Lake Group laminar
489 crust travertines of Canada briefly described by Rainbird et al. (2006); those travertines are
490 associated with volcanic rocks.

491 All surficial dolomite crusts (type 3) are capped with a thin siliceous sinter or silcrete (e.g.
492 Fig. 7). This silica layer distinguishes such dolomite crust from “normal” stratified dolostones
493 of the KSF. Silica sinters in modern travertines are usually precipitated through rapid cooling
494 and/or evaporation (e.g. Jones et al., 1996; White et al., 1956; Rimstidt and Cole, 1983), and
495 do not form if surfaces are continually bathed in hot-spring water (Jones et al., 1996).
496 Melezhik and Fallick (2001) noted that the silica sinters in the travertines of the KSF may
497 have not been formed in a similar manner to the sinters in modern travertines. Rather, they
498 suggested that the silica was derived from quartz-rich carbonate rocks underlying the
499 travertines, and (hot?) silica-saturated fluids were transported by capillary rise, allowing
500 sinter to be subaerially precipitated by evaporation and cooling.

501 The investigated dolomite cements (type 5; **Fig. 9**) are often composed of feather-shaped
502 dolomite crystals. Crystals with feather-like shapes are well known in modern travertines
503 (e.g. Fouke et al., 2000; Guo and Riding, 1992; Rainey and Jones, 2009).

504 Corbella et al. (2004) reported hydrothermal calcite and sulfide stalactites, which outwardly
505 resemble some of the type 3 thinly-laminated crusts and the thicker crusts underlying them.
506 Corbella et al. (2004) described a model of karst and stalactite formation by (i) dissolution of
507 carbonates and (ii) precipitation of sulfides and later carbonates. Their model is based on the
508 mixing of two calcite-supersaturated hydrothermal fluids with different salinities, Ca and
509 bicarbonate ion concentrations and pH levels. They concluded that dissolution and
510 precipitation occur in separate but adjacent zones; dissolution occurs in the zone dominated
511 by brine-rich mixtures and precipitation takes place where low ionic strength groundwaters
512 dominate. A similar model might explain several of the KSF carbonates, whereby brines
513 generated through dissolution of evaporites mixed with low ionic strength groundwaters,
514 causing both cavity formation and carbonate precipitation.

515

516 **6.2 A comparison to the host carbonates**

517

518 The background sedimentary dolostones and limestones sampled in Core 5A were studied by
519 Salminen et al. (2013a). Below, their data are compared to the newly obtained isotopic and
520 geochemical results from the cavity fills, crusts and cements.

521 At the core scale, cavity fills, cements and crusts show in overall similar $\delta^{13}\text{C}$ variation (from
522 4 to 8‰) to that (from 5 to 8 ‰) reported by Salminen et al. (2013a) from the host carbonates
523 (**Fig. 2**). However, the carbonate fills, crusts and cements often show lower $\delta^{13}\text{C}$ values
524 compared to the host rocks from comparable depths (**Fig. 2**).

525 The $\delta^{18}\text{O}$ values obtained from cavity fills, crusts and cements show a somewhat different
526 range (from -15 to -8‰) compared to that reported from the host rocks (from -18 to -10‰;
527 Salminen et al., 2013a). This difference could be explained by different depositional

528 conditions or different fluids. For example, evaporation or a down-flow trend of rising water
529 temperature could explain increasingly higher $\delta^{18}\text{O}$ values in the precipitates.

530 At hand-specimen scale, the cavity fills, crusts and cements show more variation in the $\delta^{13}\text{C}$
531 and $\delta^{18}\text{O}$ values than that reported from their host carbonates (Salminen et al., 2013). For
532 example, within a distance of 10 cm, the carbonate fills, crusts and cements show up to 2‰
533 variation in their $\delta^{13}\text{C}$ values and up to 4‰ variation in their $\delta^{18}\text{O}$ values.

534 The investigated carbonate precipitates show coupled $\delta^{13}\text{C}$ and the $\delta^{18}\text{O}$ trends (**Figs. 11, 14**)
535 In host rocks, only carbonates near contacts with magmatic rocks show statistically
536 significant correlations between the $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values, whereas the bulk carbonates do
537 not, hence suggesting overall good preservation of the C isotope composition (Salminen et
538 al., 2013a).

539 Coupled $\delta^{13}\text{C}$ and the $\delta^{18}\text{O}$ trends in carbonate precipitates could indicate metamorphic
540 alteration (Valley, 1986). As the host rocks have generally retained their primary C isotope
541 compositions, it is unlikely that the C isotope composition of the carbonate fills, crusts or
542 cements was significantly affected by metamorphism. A linear correlation between the $\delta^{13}\text{C}$
543 and $\delta^{18}\text{O}$ values could instead indicate mixing between two chemically different fluids
544 (Banner et al, 1988; Banner and Hanson, 1990; Lohmann, 1988) or a thermal origin (e.g.
545 Chafetz and Guidry, 2003).

546 In some cavities, the $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values decrease from the cavity walls towards the cavity
547 center. This could result from the progressively decreasing degree of fluid/bedrock-
548 interaction. Thus higher $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values could indicate more interaction with the host
549 rocks. Alternatively, in some modern speleothem precipitation, kinetic isotope fractionation
550 of C and O results in correlations between the $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values (e.g. Mickler et al.,

551 2006). However, in the modern speleothems of Mickler et al. (2006), the $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$
552 values increase from the rim towards the center – and this is not the case in the KSF.

553 The surficial dolomite crust at ca. 30.3 m and the possible surficial crust at ca. 27.4 m both
554 show lower $\delta^{13}\text{C}$ values and generally lower $\delta^{18}\text{O}$ values compared to their host carbonates
555 (**Fig. 2**). In contrast, the similar crust at ca. 68.3 m shows slightly higher $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values
556 than its host carbonates (**Fig. 2**). This latter crust also shows a subtle trend of upwards
557 increasing $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values in its upper half (**Fig. 18**). This pattern might conceivably be
558 related to evaporation (e.g. Stiller et al., 1985). Alternatively, it might be related to
559 temporally (even seasonally?) increasing carbonate precipitation (causing progressive $^{12}\text{CO}_2$
560 loss) upstream of the site of crust deposition, from progressively cooler waters. This is seen
561 in some modern tufa and speleothem systems where mineral precipitation is driven by CO_2
562 degassing, since subterranean cave ventilation, CO_2 degassing and calcite precipitation, is
563 greater in the winter than the summer (e.g. Matsuoka et al., 2001). However the high pCO_2 of
564 the Paleoproterozoic atmosphere probably argues against such a degassing precipitation
565 mechanism for the KSF precipitates.

566

567 .

568 The investigated cavity fills, crusts and cements often show higher Mn contents than those in
569 their host carbonate rocks, whereas their Sr contents are similar (**Fig. 10**). High Mn contents
570 could imply a thermal origin for the waters, as hot-spring travertines are often characterized
571 by high Mn and Fe contents (e.g. Tanaka et al., 1994; Pentecost, 2005, pp.118). The
572 investigated precipitates show upwards decreasing Mn and Fe trends at the core scale (**Fig.**
573 **10**), This could indicate decreasing concentrations of soluble iron and manganese in the

574 carbonate mineral precipitating waters through time, consistent with a temporal change to a
575 more oxidizing environment (e.g. Pentecost, 2005; pp.116; Kah et al., 2012).

576 In the Upper Dolostone member, the precipitates are slightly less Fe rich than their host rocks
577 (**Fig. 10**). In contrast, the Mn contents are similar to or higher than those in the host rocks
578 (**Fig. 10**). This is perhaps explained by divalent Mn (which can be incorporated into the
579 carbonate mineral lattice) persisting longer in oxidizing travertine-depositing conditions than
580 Fe^{2+} (Pentecost, 2005, pp. 116).

581 The upwards increasing Sr contents might imply a transition to a more saline environment
582 and opening of a connection to the ocean. This idea is supported by published interpretations
583 of the depositional setting as Melezhik and Fallick (2005) and Salminen et al. (2013b)
584 suggested sea-water invasion during the deposition of the upper part of the KSF. Higher Sr
585 contents could also imply dissolution of the calcitic host rocks in the uppermost part of the
586 succession. Calcite in general contains less Mn and more Sr than dolomite (e.g. Gaucher et
587 al., 2007).

588 Samples distribute into four groups according to the co-variation plots (**Fig. 11**). (I) Majority
589 of the samples is plotted in the group, which is characterized by relatively high $\delta^{13}\text{C}$ values
590 (ca. 6–8 ‰), moderate to low $\delta^{18}\text{O}$ values (ca. -10 to -15‰), relatively low Mn contents (up
591 to 1000 ppm) and Sr contents (<200 ppm). (II) Another group includes some cavity and vein
592 fills (chiefly dolospar representing the **latest infill**) with **high $\delta^{18}\text{O}$ values** (ca. -9 to -8 ‰),
593 high $\delta^{13}\text{C}$ values (ca. 7-8 ‰), and low Mn (<500 ppm) and Sr contents (≤ 120 ppm). (III)
594 Second minor group includes different types of samples with **low $\delta^{13}\text{C}$ values** (<6 ‰), low
595 $\delta^{18}\text{O}$ values (≤ 11 ‰), high Mn contents (> 1000 ppm) and low to moderate Sr contents. (IV)
596 One more minor group consist of a few samples (cavity and vein precipitates and one

597 subsample from a surficial crust) from the Upper Dolostone member with high $\delta^{13}\text{C}$ values,
598 moderate $\delta^{18}\text{O}$ values, **high Sr contents** and low Mn contents.

599 Samples in the group I likely show also geochemical characteristics of the dissolved
600 dolomitic host rocks, whereas the group IV probably indicates dissolution of calcitic (and
601 dolomitic) host rocks. Group III samples perhaps represent “pure” deep-sourced fluid, as
602 implied by the low $\delta^{13}\text{C}$ values and high Mn contents. Group II represents the last dolomite
603 generation/phase.

604

605

606

607 **6.3 Comparison to previous investigations of the KSF travertines**

608

609 Some carbonates investigated in Core 5A (e.g. the cavity at ca. 81.4 m) resemble the
610 travertines studied by Melezhik and Fallick (2001) and Melezhik et al., (2004) from other
611 locations in the KSF. Melezhik and Fallick (2001) reported $\delta^{13}\text{C}$ values from 0.8 to 7.2‰ and
612 $\delta^{18}\text{O}$ values from -12.4 to -9.8‰ in a travertine crust. They also reported a much greater
613 range of $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ ratios (from -6.1 to 7.7‰ and from -18.3 to -9.1‰, respectively) in a
614 travertine mound. Even the lowest $\delta^{13}\text{C}$ values found in the current study are several per-mils
615 higher than the lowest $\delta^{13}\text{C}$ values reported by Melezhik and Fallick (2001). They explained
616 the observed $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ variations and trends by invoking initial CO_2 out-gassing at the
617 hot-spring orifice, down-flow trends of thermal waters (following modern examples e.g.
618 Friedman, 1970; Chafetz et al., 1991; Chafetz and Lawrence, 1994), mixing with ambient

619 waters, and possible evaporation. The travertine crusts were suggested to be precipitated in
620 shallow distal pools, periodically fed by the thermal springs (Melezhik and Fallick, 2001).

621 Upwards (within cavities/crusts) increasing or decreasing $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ trends and cyclical
622 variations of $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values were found in the investigated precipitates from Core 5A.
623 Such fluctuations could be explained by the model of Melezhik and Fallick (2001). The
624 upward $\delta^{13}\text{C}$ increase would then be explained by a down-flow trend, possibly connected to
625 evaporation and some mixing with ambient water in the depositional basin. The upwards
626 decreasing $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ trends could be explained by a resurgence in supply of hot-spring
627 water to the point of travertine precipitation (such as a distal pool). The cyclical $\delta^{13}\text{C}$ pattern
628 in the crust at ca. 30.3 m could indicate pulsed supply of the thermal waters, connected to
629 intervals of intense evaporation.

630 Melezhik and Fallick (2001) and Melezhik et al. (2004) found that travertines and the
631 stratified dolostone of the KSF plot along two entirely different lines in a $\delta^{13}\text{C}$ - $\delta^{18}\text{O}$ cross-
632 plot. Melezhik et al. (2004) found that travertines show positive correlation between the $\delta^{13}\text{C}$
633 and $\delta^{18}\text{O}$ values, whereas no significant correlation between them is found in the stratified
634 dolostone. Like the travertines investigated by Melezhik et al. (2004), the cavity fills, crusts
635 and cements in Core 5A show positive co-variation between the $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values (**Fig.**
636 **14**), while only the contact-altered host carbonate show co-variation between $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$
637 values (Salminen et al., 2013a). Thus, the investigated carbonate precipitates of Core 5A also
638 plot differently on a $\delta^{13}\text{C}$ - $\delta^{18}\text{O}$ cross-plot to their host rocks.

639 Melezhik and Fallick (2003) reported variable $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values from
640 dolospar/dolomicrospar cements and fills in fenestrae and voids. They did not find significant
641 differences between the $\delta^{13}\text{C}$ values of such cements and fills and their host dolomicrite.
642 They did find that the $\delta^{18}\text{O}$ values in the dolospar/dolomicrospar were somewhat lower to

643 those in the host dolomicrite, hence were explained by syn-depositional/early diagenetic
644 alteration processes. In Core 5A, the $\delta^{13}\text{C}$ values are commonly lower in the cavity fills,
645 crusts and cements than in their host carbonate rocks, suggesting that an external lower $\delta^{13}\text{C}$
646 carbon source is required. It is plausible that all the investigated carbonate precipitates from
647 Core 5A could have been precipitated from ^{13}C -depleted fluids.

648

649 **6.4 The origin of the cavity fills, crusts and cements in the KSF**

650

651 An external carbon source would explain the lower $\delta^{13}\text{C}$ values in the cavity fills, crusts and
652 cements compared to the surrounding host rock. That source could be (i) CO_2 -bearing
653 meteoric water, (ii) CO_2 derived from oxidation of organic carbon, or (iii) deep-sourced CO_2 .

654 As the sedimentary rocks in the KSF are devoid of organic carbon (e.g. Melezhik and Fallick,
655 2005), it seems unlikely that the external carbon comes directly from an organic source.

656 Ameteoric water with low $\delta^{13}\text{C}$ dissolved inorganic carbon is a possibility, though this would
657 have implications for the record of life in Paleoproterozoic terrestrial environments. The
658 external carbon could also be derived from deep-sourced CO_2 , as magmatic carbon
659 commonly has $\delta^{13}\text{C}$ values around -5 to -7‰ (e.g. Anderson and Arthur, 1983; Deines, 2002).

660 Several possible mechanisms could be invoked to explain how these carbonate rocks formed
661 under a high pCO_2 atmosphere and without high pCO_2 soils (Brasier, 2011).

662 Paleoproterozoic speleothem may have been formed **(1) under an ice cover due to warming**
663 **of waters** (Dreybrodt, 1982), and alternatively **(2) freezing of water** could have caused
664 carbonate precipitation in caves (e.g. Aharon, 1988). However, no petrographic or structural
665 evidence for glaciation is present in the KSF and these two models cannot explain the
666 observed $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ trends.

667 Precipitation may have been aided by **(3) cyanobacteria and microbes** (e.g. Merz, 1992;
668 Spiro and Pentecost, 1991). As stromatolites are found in the KSF, it is possible that
669 cyanobacteria influenced precipitation of crusts. However the formation of the cavity fills and
670 veins cannot be explained by photosynthetic cyanobacteria.

671 Carbonates are today usually dissolved by fluids rich in carbonic acid (H_2CO_3). However,
672 fluids rich in other acids may also cause carbonate dissolution. For example, **(4) sulfuric acid**
673 (H_2SO_4) has been reported to cause dissolution (e.g. Atkinson, 1983). Local H_2SO_4
674 production in the Paleoproterozoic was likely increased due to the oxygenation of the
675 atmosphere and the oxidation of sulfides (Bekker and Holland, 2012; Brasier, 2011). But
676 sulfuric acid dissolution alone could not itself explain the precipitation of the cavity fills.

677 Karst and speleothem-like precipitates can also be formed by local **(5) calcite**
678 **supersaturation due to a common-ion effect**: extra calcium is supplied e.g. from dissolution
679 of evaporites (e.g. Calaforra et al., 2008; Wigley, 1973a) or incongruent dissolution of
680 dolomite (e.g. Wigley, 1973b). This leads to calcite precipitation when the solution becomes
681 supersaturated with respect to calcite. The common-ion effect is found in polymineralic
682 systems (calcite and gypsum, or calcite and dolomite) that are initially calcite saturated (*see*
683 *Brasier, 2011*). Such a common-ion effect causes precipitation in some modern hot springs
684 (Pentecost, 1995) and in cold ($\sim 6.5^\circ\text{C}$) springs (Omelson et al., 2006) in the presence of
685 gypsum. It can also explain the precipitation of lake tufas (Dunn, 1953). Thus, the common-
686 ion effect could also explain the formation of the surficial crusts and cements in the KSF.

687 The **common-ion effect caused by dissolution of evaporites (5a)** may explain the formation
688 of the small-scale cavities and the precipitation of the carbonate fills and the surficial
689 precipitates in the KSF. Probable sulfate pseudomorphs have previously been reported from
690 the KSF (e.g. Melezhik and Fallick, 2005), and they were also found in this study. The

691 common-ion effect due to dissolution of Ca-evaporites (e.g. gypsum) results in precipitation
692 of calcite (e.g. Wigley, 1973a). Calcite, however, is only a minor constituent in the cavities of
693 the KSF, and there is no clear evidence of calcite recrystallizing to dolomite. Dolomites could
694 have been precipitated due to a common-ion effect caused by dissolution of Mg-rich
695 evaporites. However, pseudomorphs after Mg-rich evaporites have not been found in the
696 KSF.

697 The **common-ion effect caused by dissolution of dolomite (5b)** is a possible explanation, as
698 the host carbonates in the formation are mainly dolostones. However, Wigley (1973b) noted
699 that the incongruent dissolution of dolomite leads to the precipitation of calcite prior to
700 dolomite. Incongruent dissolution of dolomite is an unlikely explanation for the formation of
701 the precipitates. Thus, the common-ion effect could explain the karst solution in the KSF, but
702 does not obviously explain the precipitation of chiefly dolomitic precipitates.

703 Degassing can also happen due to **(6) boiling** (in addition to high temperature) in springs
704 (e.g. Arnósson, 1989). Temperatures well below the boiling point cause travertine
705 precipitation today (e.g. Fouke et al., 2000), but meteoric waters can also be heated when
706 percolating through the bedrock (e.g. Dublyansky, 1995; Djidi et al., 2008). Boiling meteoric
707 waters could cause carbonate precipitation, but might not explain all the observed $\delta^{13}\text{C}$ and
708 $\delta^{18}\text{O}$ trends. Moreover, heating of meteoric waters could be connected to supply of deep-
709 sourced CO_2 (Djidi et al., 2008).

710 **Deep-sourced CO_2 (7)** could explain the formation of the cavity fills, surficial dolomite
711 crusts and dolomite cements in the KSF (e.g. Duliński et al., 1995; Pentecost and Viles, 1994;
712 Yoshimura et al., 2004). Deep-sourced CO_2 could be either volcanic or produced by
713 metamorphic reactions, and is usually associated with hydrothermal activity (e.g. Pentecost
714 and Viles, 1994; Pentecost, 1995). Deep-sourced CO_2 can increase the dissolution rates as

715 well as carbonate precipitation even in ambient water temperatures (Yoshimura et al., 2004),
716 but some of the investigated precipitates have physical and geochemical characteristics (e.g.
717 veins that could have been fluid conduits; mound morphologies; feather-shaped crystals, high
718 Mn contents, co-variation and trends of $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$) that are at least consistent with deep-
719 sourced fluids and hot-spring travertine.

720 The isotope results from Core 5A samples do not show unambiguous evidence of a deep-
721 sourced origin or deposition in close vicinity to hot-spring vents. However, the carbonate
722 precipitates in the KSF could have precipitated from thermal waters distally from the vents,
723 as suggested by Melezhik and Fallick (2001).

724 The most likely explanation for the formation of the carbonate precipitates in the KSF is
725 apparently deep-sourced CO_2 , connected to hydrothermal activity. Boiling may have
726 contributed to the carbonate precipitation and deep-sourced fluids may have mixed with
727 meteoric waters. Thus, the cavity and vein fills in the KSF could represent hydrothermal karst
728 (“cave travertine”). Dolomite cements and some of the surficial crusts are likely travertine,
729 although some dolomite crusts could represent evaporite layers. The Mn/Sr ratios, elemental
730 contents and their patterns in the vein at 53.7 m in Core 5A could suggest formation during
731 burial.

732 If pseudomorphs or relicts of Mg-rich evaporites were later found in the KSF, the common-
733 ion effect connected to dissolution of evaporites could also explain the formation of the
734 cavities and the cavity fills, crusts and cements.

735

736 **6.5 The depositional setting of the KSF**

737

738 Salminen et al. (2013b) interpreted the depositional setting of Core 5A as deltaic and
739 lacustrine, but tidally influenced. Melezhik and Fallick (2005) concluded that the depositional
740 setting of the KSF was deltaic and shallow lacustrine. This new investigation supports the
741 previous interpretations of the depositional setting of the KSF.

742 Melezhik and Fallick (2001) pointed out that the modern travertines are only found in
743 terrestrial settings. Most of the investigated cavity fills, crusts and cements are likely
744 travertine or “cave travertine”. We infer they precipitated in terrestrial settings.

745 Melezhik et al. (2004) investigated various subaerial exposure surfaces in the KSF. Findings
746 of this study also support occasional subaerial exposure during the deposition of the KSF.
747 The red color in the dolomite crusts likely indicates oxic conditions, possibly in a subaerial
748 environment.. Likely evaporitic sulfate pseudomorphs were also found.

749 The question of how the deep-sourced CO₂ was supplied to the carbonate sediments in the
750 KSF is open. One possibility is through the veins (now cemented up); another is through syn-
751 depositional faulting, as the KSF was likely deposited in an intracontinental rift setting
752 (Melezhik and Sturt, 1994). Melezhik and Fallick (2005) provided a reconstruction of the
753 Kuetsjärvi rift-related lake, which includes hydrothermal springs and transverse faults.
754 Eruption of later volcanic rocks of the Kuetsjärvi Volcanic Formation may also have
755 provided a source for the formation of cavity and vein fills, which are apparently secondary
756 precipitates. However, this cannot explain the deposition of the surficial precipitates.

757

758 **7. Conclusions**

759

760 Core 5A was obtained from the Paleoproterozoic Kuetsjärvi Sedimentary Formation. Drilled
761 stratified carbonate rocks contain dolomite crusts, dolomite cements and abundant small-
762 scale cavities and veins. These cavities and veins are mainly filled with dolomite and quartz.
763 Dolomite crusts occur in the cavities and on bedding/erosional surfaces. Dolomite cements
764 are found on uneven surfaces encrusting sandstone fragments. Dolomite crusts, cements and
765 carbonate fills were analyzed for $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values and selected acid-soluble elements.
766 The newly obtained results were compared to the published results from the host carbonates.

767 The investigated carbonate precipitates show lower or similar $\delta^{13}\text{C}$ values to their host
768 carbonates. The lower $\delta^{13}\text{C}$ values suggest the presence of an external carbon source, which
769 was likely deep-sourced/hydrothermal CO_2 . Dolomite cements and the dolomite crusts on
770 bedding/erosional surfaces are likely thermal travertines. This is supported by the
771 petrographic and geochemical observations. In the absence of organic carbon rich soils, deep-
772 sourced/hydrothermal CO_2 also provides a very plausible explanation for the formation of the
773 cavities, veins and carbonate fills. Cavity fills may be interpreted as “cave travertine”. Some
774 veins may represent burial veins, or could have been conduits through which the travertine-
775 precipitating fluids passed.

776 The $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values show positive correlation. Both upwards increasing and decreasing
777 $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ trends were found in the dolomite crusts. These trends could be explained by
778 down-flowing thermal water possibly affected by evaporation and/or modified due to mixing
779 with the ambient waters. Dolomite crusts showing only upwards increasing $\delta^{13}\text{C}$ trends could
780 represent dolomitized evaporite layers. The $\delta^{13}\text{C}$ and $\delta^{18}\text{O}$ values commonly show increasing
781 trends from the middle of the cavity towards the cavity walls, probably indicating different
782 degrees of mixing with the host rocks and thermal waters.

783 Few reports have been published on Precambrian travertine or speleothem. The above new
784 detailed geochemical data and petrographic analyses of what may be Earth's earliest
785 travertines provide a glimpse of complexity in textures and facies likely quite similar to that
786 found in modern hot-spring settings. Consideration of several possible depositional
787 mechanisms leads to the conclusion that these are two billion year old hot-spring travertines
788 likely formed by similar processes to those of the modern world.

789

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791

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795

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1047 **Figure captions**

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1049 **Figure 1.** (a) The location of the KSF and the Pechenga Greenstone Belt. (b) The location of
1050 Core 5A in the Pechenga Greenstone belt. Modified from Melezhik and Fallick (2005)

1051 **Figure 2.** Lithostratigraphic profile of Core 5A (modified from Salminen et al., 2013a) and
1052 the C and O isotope composition of the surficial dolomite crusts and cements and carbonate
1053 fills in cavities and veins. The C and O isotope compositions of the host carbonates (micro-
1054 drilled samples; Salminen et al., 2013a) are shown for comparison. In addition, the estimated
1055 proportion of the carbonate crusts, cements and fills (Salminen et al., 2013a) is shown in the
1056 figure.

1057 **Figure 3.** Cavities. (a) Dissolution cavity at the depth of 62.1 m. The cavity is mainly filled
1058 with dolospar. Dolostone immediately surrounding the cavity is recrystallized and has a
1059 powdery appearance. (b) Dissolution cavity at the depth of 69.6 m with complex internal
1060 structure. The outer zone of the cavity fill includes (1) thin laminae of dolomite and quartz
1061 (also forming “mounds”) and (2) yellowish sparitic calcite. These are sealed with (3) a thin
1062 rim of fibrous, inward-growing dolomite. Quartz with some dolomite and (4) pyrite represent
1063 the latest fills in the cavity.

1064 **Figure 4.** Cavities. (a) Scanned image of a cavity at the depth of 83.0 m. The cavity is lined
1065 with several rims of sparitic dolomite, capped with a rim of fibrous dolomite. The remaining
1066 space is filled with dolospar and minor quartz. (b) Cavity (at the depth of 72.7 m) includes

1067 quartz-filled fenestrae, some of them appearing as sulfate pseudomorphs (e.g. in the circled
1068 area).

1069 **Figure 5.** Cavities with stalactites. (a) Quartz-filled and carbonate-walled cavity at the depth
1070 of 60.2 m. Note small white stalactites at the roof of the cavity. (b) Scanned slab showing
1071 gray-and-white, thinly-laminated crust (at the bottom) and stalactites in a cavity at a depth of
1072 81.4 m. (c) Gray, white and yellow thinly-laminated dolomitic crust (at the bottom) and
1073 stalactites (in the middle), overlain by pinkish dolomite, at a depth of 31.4 m.

1074 **Figure 6.** Veins. (a) A thin section showing a vein above a stromatolitic layer at the depth of
1075 35.1 m. (b) A thin section showing a vein in sandy micritic dolostone at the depth of 62.7 m.
1076 (c) A probable burial vein in dolarenite overlying a sandstone-filled cavity (at the bottom).
1077 The vein shows coloration and mineralogical zoning. The innermost, reddish part is calcite. It
1078 is lined with a mixture of gray calcite and dolomite. The outermost white rim is composed of
1079 dolomite. The depth is 53.7 m.

1080

1081 **Figure 7.** A surficial dolomite crust at the depth of 68.3 m. (a) Thin-section of the crust. (b)
1082 Thin-section image of the crust (stereo microscope). The micritic/sparitic crust at the base
1083 (grayish color in the photo) shows upward growth and thin lamination. This part of the crust
1084 is overlain by a rim of fibrous dolomite (black colored in the photo). The entire crust is sealed
1085 with a silica sinter (bright colored in the photo) and overlain by a sandstone layer (at the top
1086 of the photo).

1087 **Figure 8.** A probable surficial dolomite crust at 27.4 m. The crust is over- and underlain by
1088 sandy dolarenite. The basal part of the crust is composed of beige dolomitic bands overlain

1089 by partially vuggy interval. The crust is capped with white rim of sparitic dolomite and a thin
1090 film of siliceous precipitate (a sinter or a silcrete). (a) A photo of the crust. (b) Scanned slab.

1091 **Figure 9.** Dolomite cements. a) Scanned slab showing dolomite cement with feathery
1092 crystals, forming a mound-like structure at the depth of 76.4 m. b) Dolomite cements in a thin
1093 section. Fibrous dolomite cement forms thin crusts/mounds at the depth of 76.0 m. c) A thin
1094 section image of feathery dolomite cements (stereo microscope) at the depth of 76.0 m. d) A
1095 thin section image showing surficial cement (stereo microscope) at the depth of 76.9 m.

1096

1097 **Figure 10.** Acid-soluble elemental contents (ICP-MS) of the carbonate fills, cements and
1098 crusts. The elemental contents of the host carbonates (micro-drilled samples) have been
1099 adopted from Salminen et al. (2013a); the Ba contents for the host rocks have not been
1100 published.

1101 **Figure 11.** Cross-plots of selected variables in carbonate fills, crusts and cements.

1102 **Figure 12.** The C and O isotope composition and acid-soluble elemental compositions of
1103 micro-drilled subsamples from the cavity at 79.6 m. The subsamples were obtained from the
1104 circular spots shown in the slabs photo aside. The width of the slabs is ca. 2.5 cm.

1105 **Figure 13.** The C and O isotope composition and acid-soluble elemental compositions of
1106 micro-drilled subsamples from the cavity at 81.4 m. The subsamples were obtained from the
1107 circular spots shown in the slab photos aside. The width of the slabs is ca. 2.5 cm.

1108 **Figure 14.** A $\delta^{13}\text{C}$ - $\delta^{18}\text{O}$ cross-plot of the subsamples from the cavities at 79.6 and 81.4 m and
1109 the surficial dolomite crusts at 30.3 and 68.3 m.

1110 **Figure 15.** The C and O isotope composition of the micro-drilled subsamples from a cavity at
1111 83.0 m. The subsamples were micro-drilled from the circular spots shown in the slab photo
1112 aside. The width of the slab is ca. 2.5 cm.

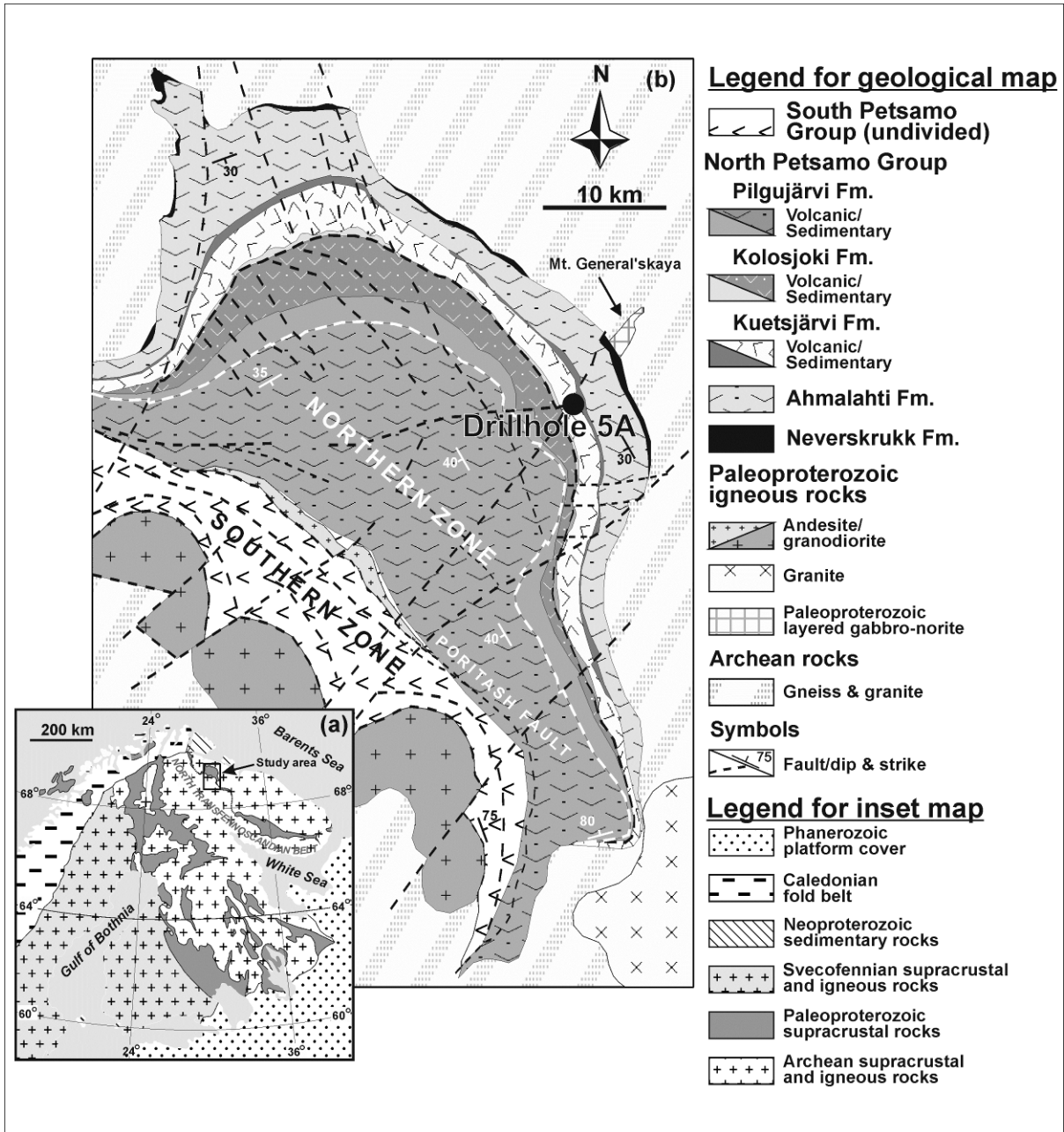
1113 **Figure 16.** The C and O isotope composition of the dolospar fill from the cavity at 62.1 m.
1114 The C and O isotope composition of the dolostone immediately over- and underlying the
1115 cavity is also shown in the figure. The subsamples were micro-drilled from the circular spots
1116 shown in the slab photo aside. The width of the slab is ca. 2.5 cm.

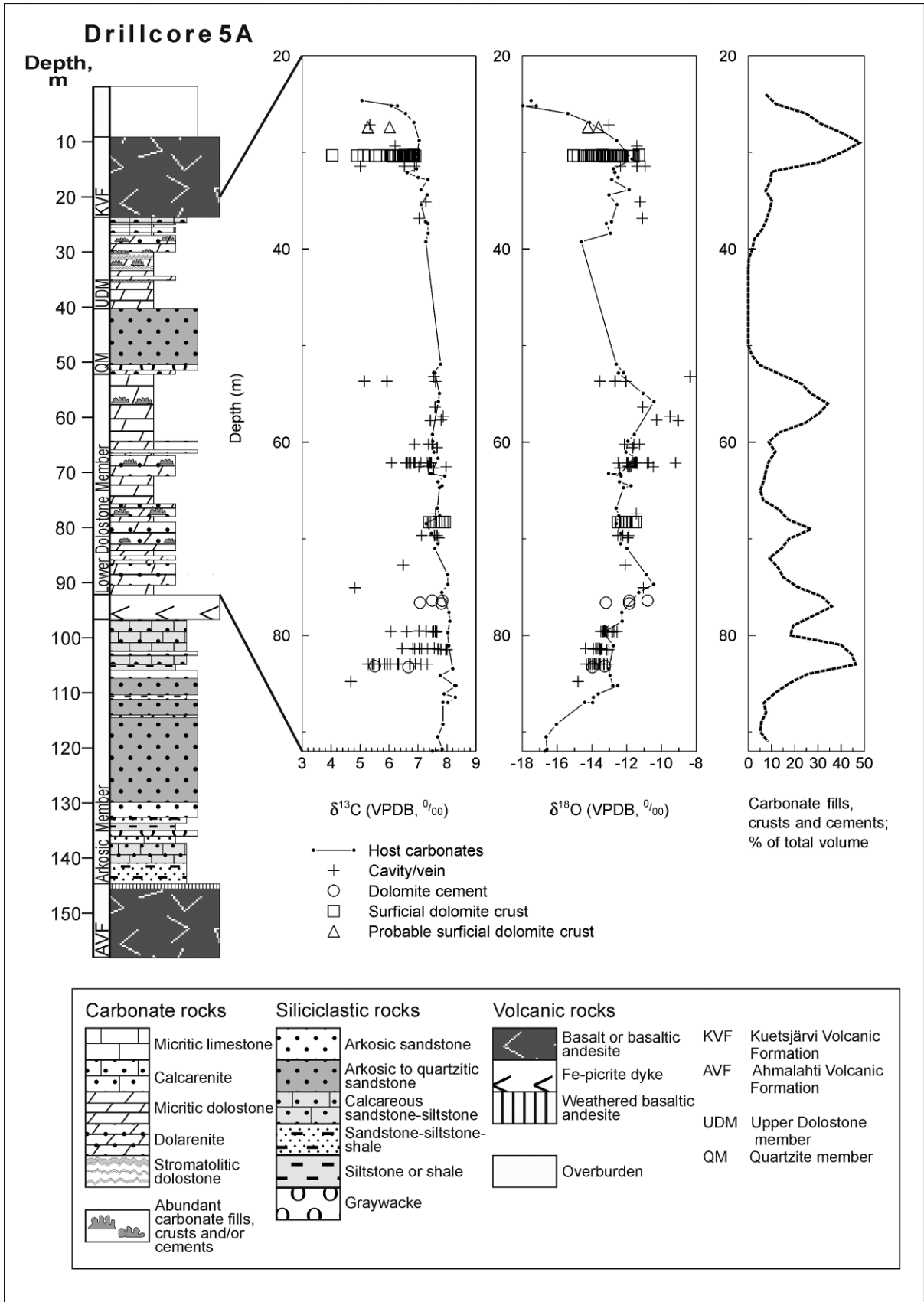
1117 **Figure 17.** The C and O isotope composition of the subsamples from the surficial dolomite
1118 crust at 30.3 m. The subsamples were obtained from the circular spots shown in the slab
1119 photo aside. The subsamples A-C were also analyzed for the acid-soluble elemental
1120 composition (results not shown in this figure). The width of the slab is ca. 2.5 cm.

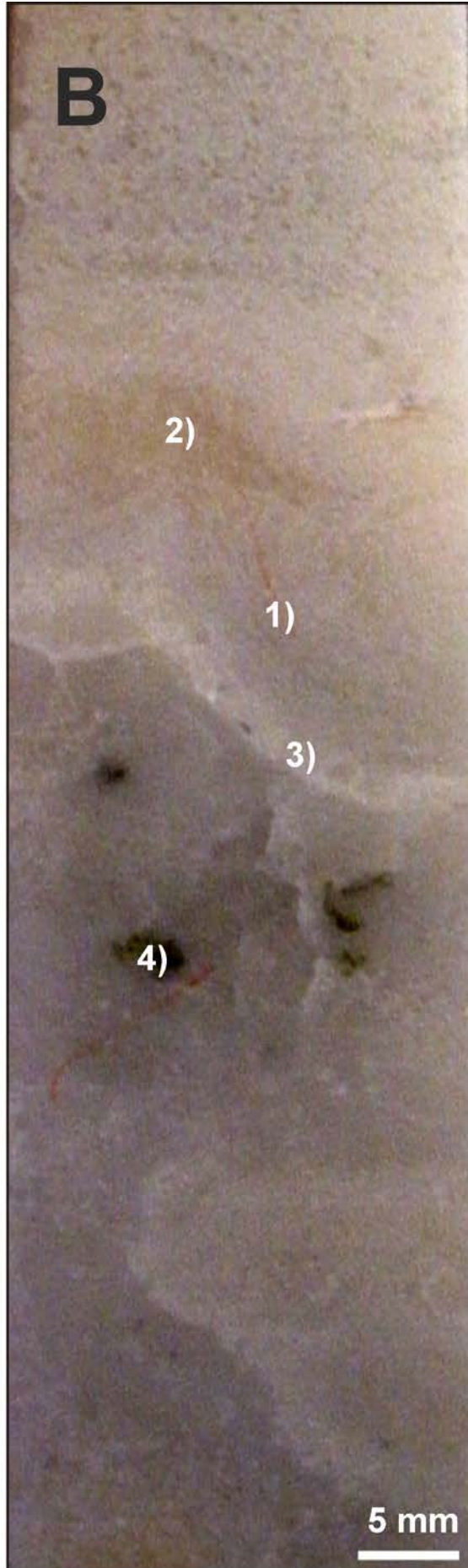
1121 **Figure 18.** The C and O isotope composition of the subsamples from the surficial dolomite
1122 crust at 68.3 m. The subsamples were micro-drilled from the circular spots shown in the slab
1123 photo aside. The width of the slab is ca. 2.5 cm.

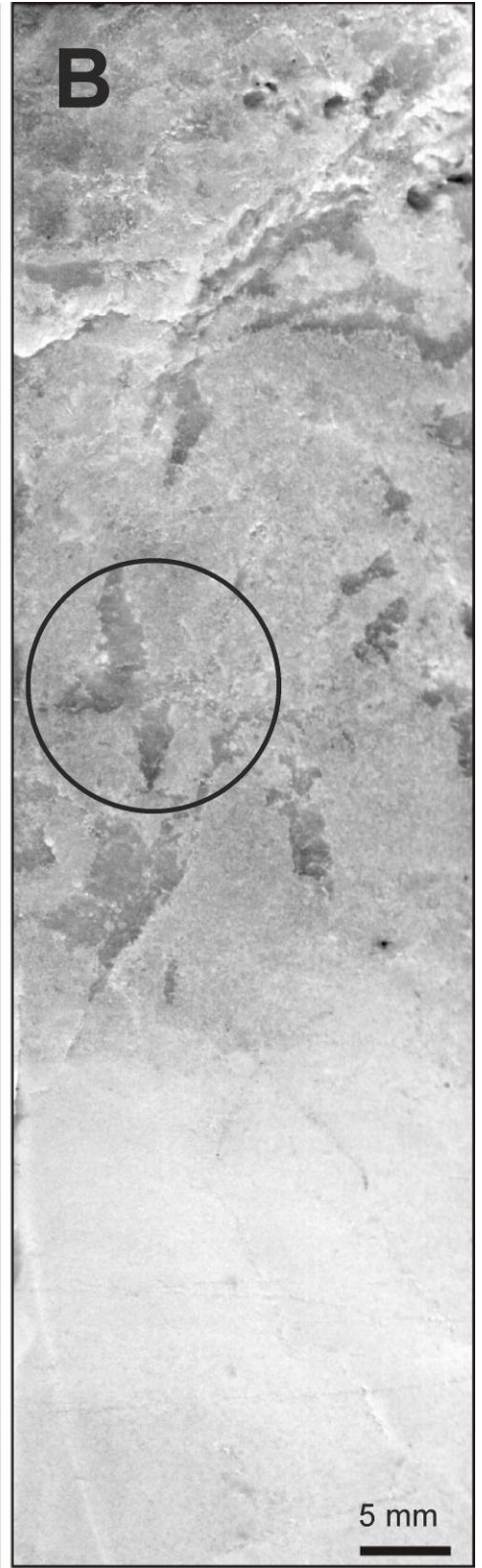
1124 **Inline Supplementary Figure 1.** Dolomite crust at the depth of 30.3 m. The crust starts
1125 with reddish dolomitic bands overlain by beige dolomite bands (the lower half), then
1126 followed by a crust with clotted microfabrics (the upper half). The dolomite crust contains
1127 small vugs filled with white, gray or black sparitic dolomite. There are also reddish fills in the
1128 vugs. (a) A photo of the crust. (b) Scanned slab of the crust showing also position of the
1129 micro-drilled subsamples.

1130 **Inline Supplementary Figure 2.** A secondary electron image showing possible sulfate
1131 pseudomorphs (not visible in a back-scatter image) in dolomite cement at the depth of 76.9
1132 m.

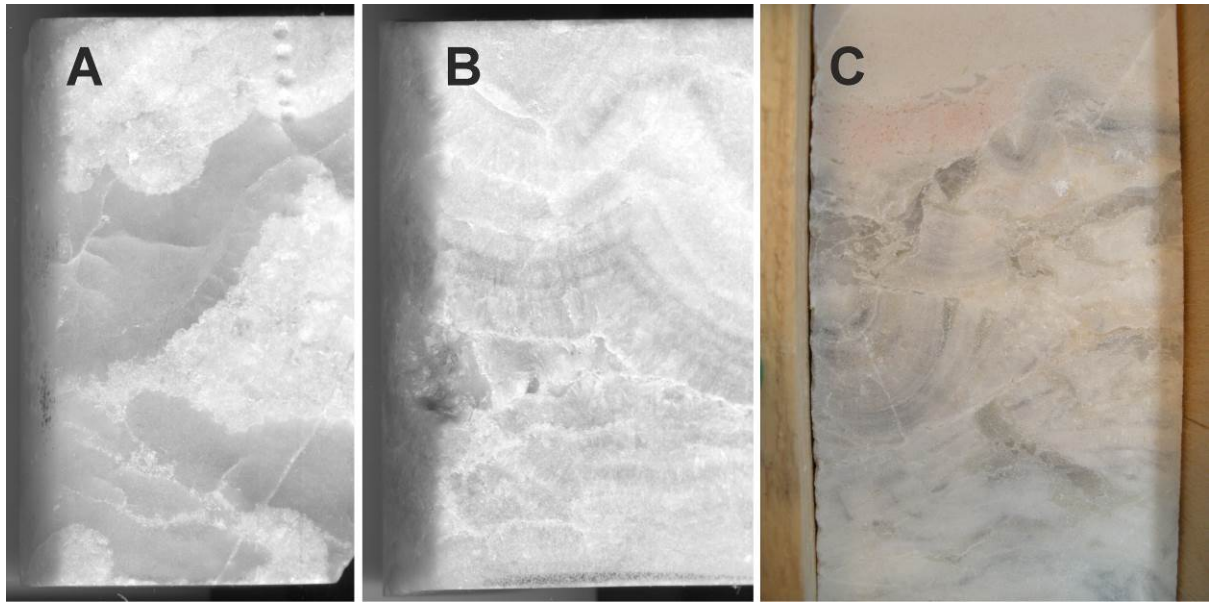




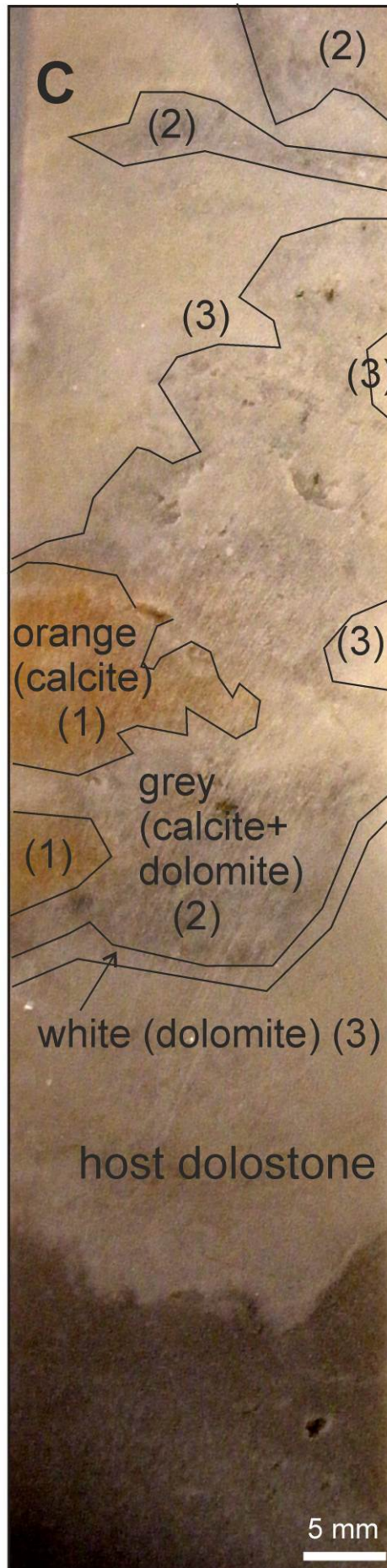


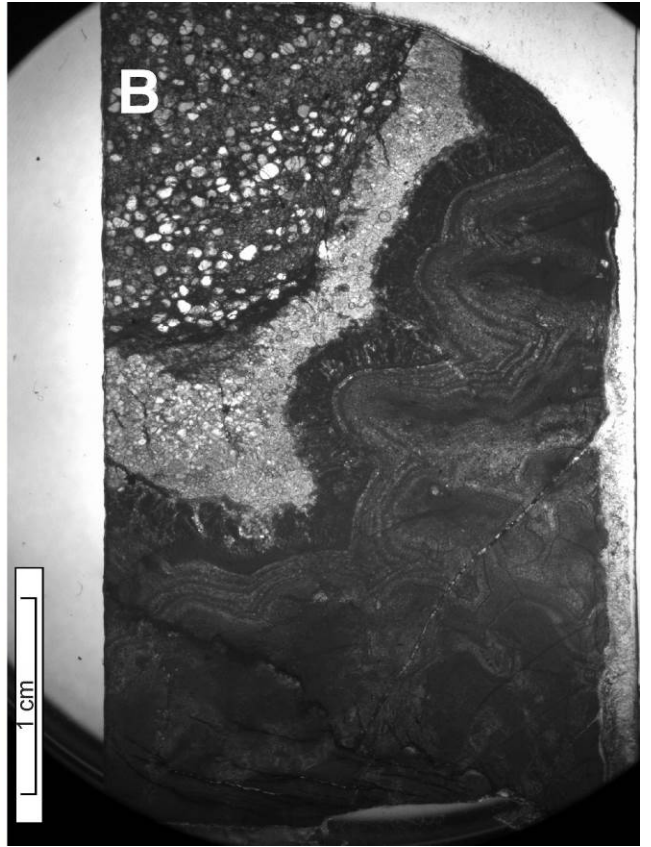


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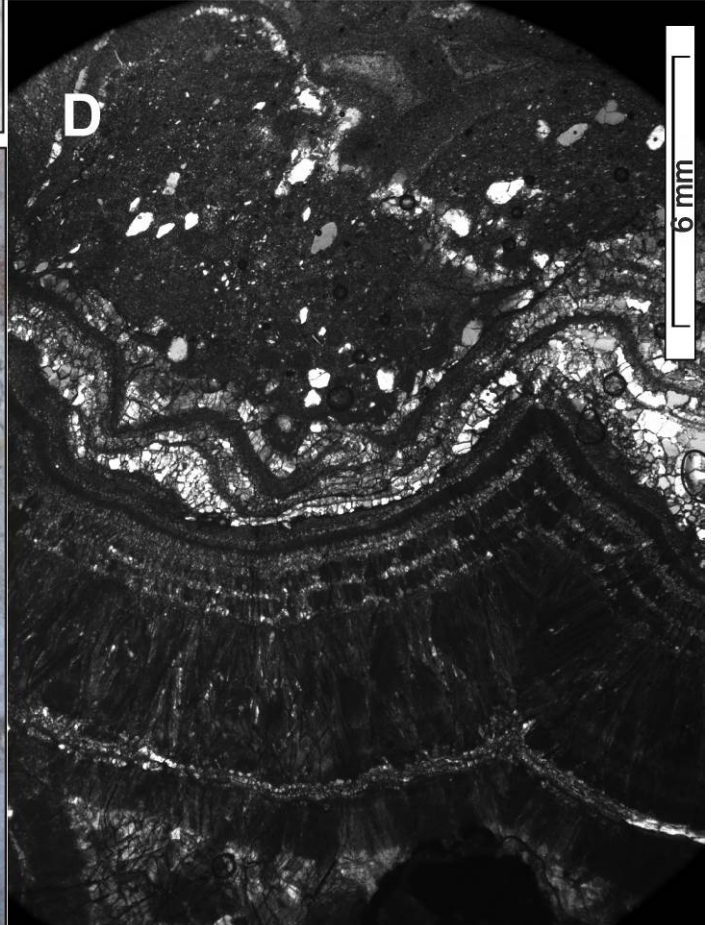
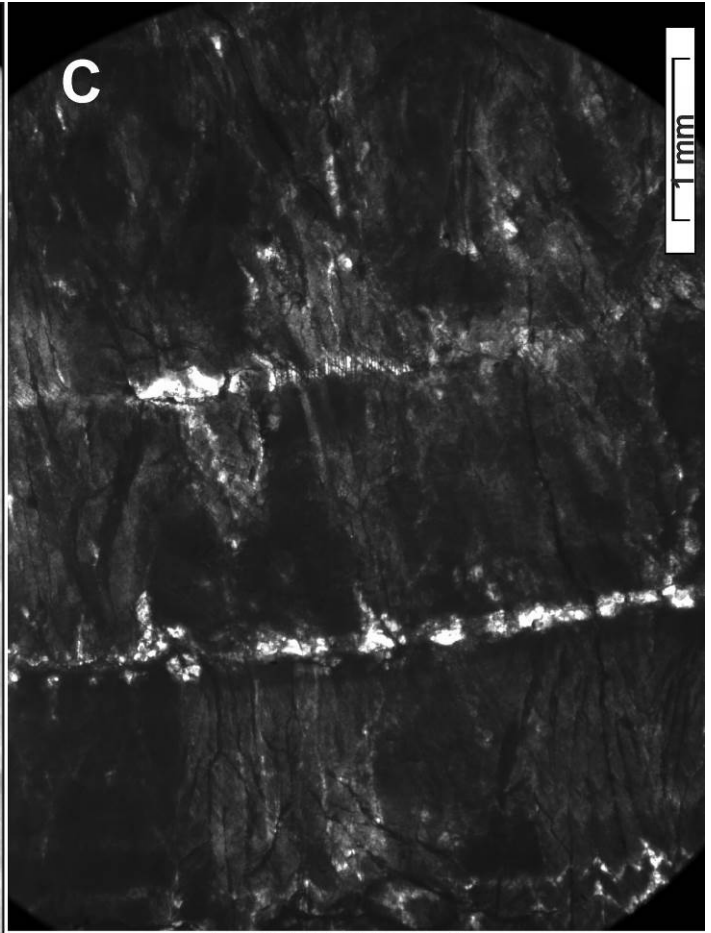


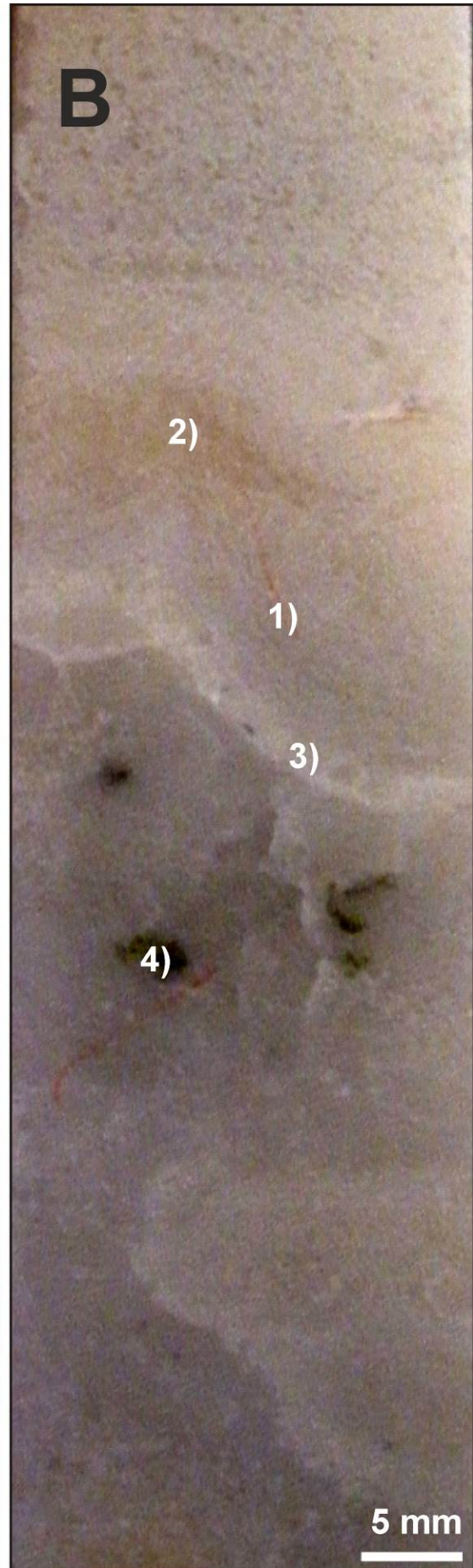


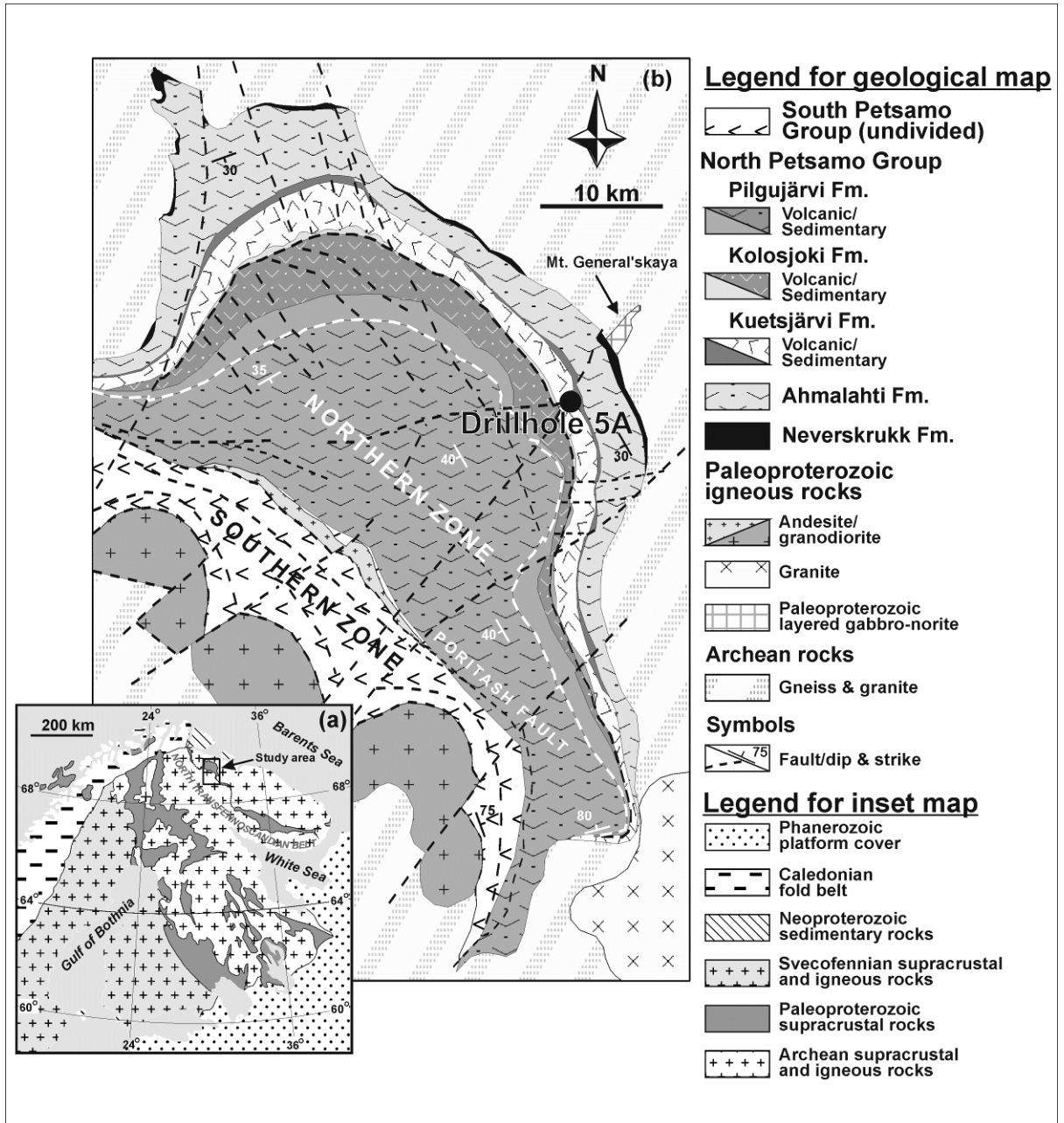
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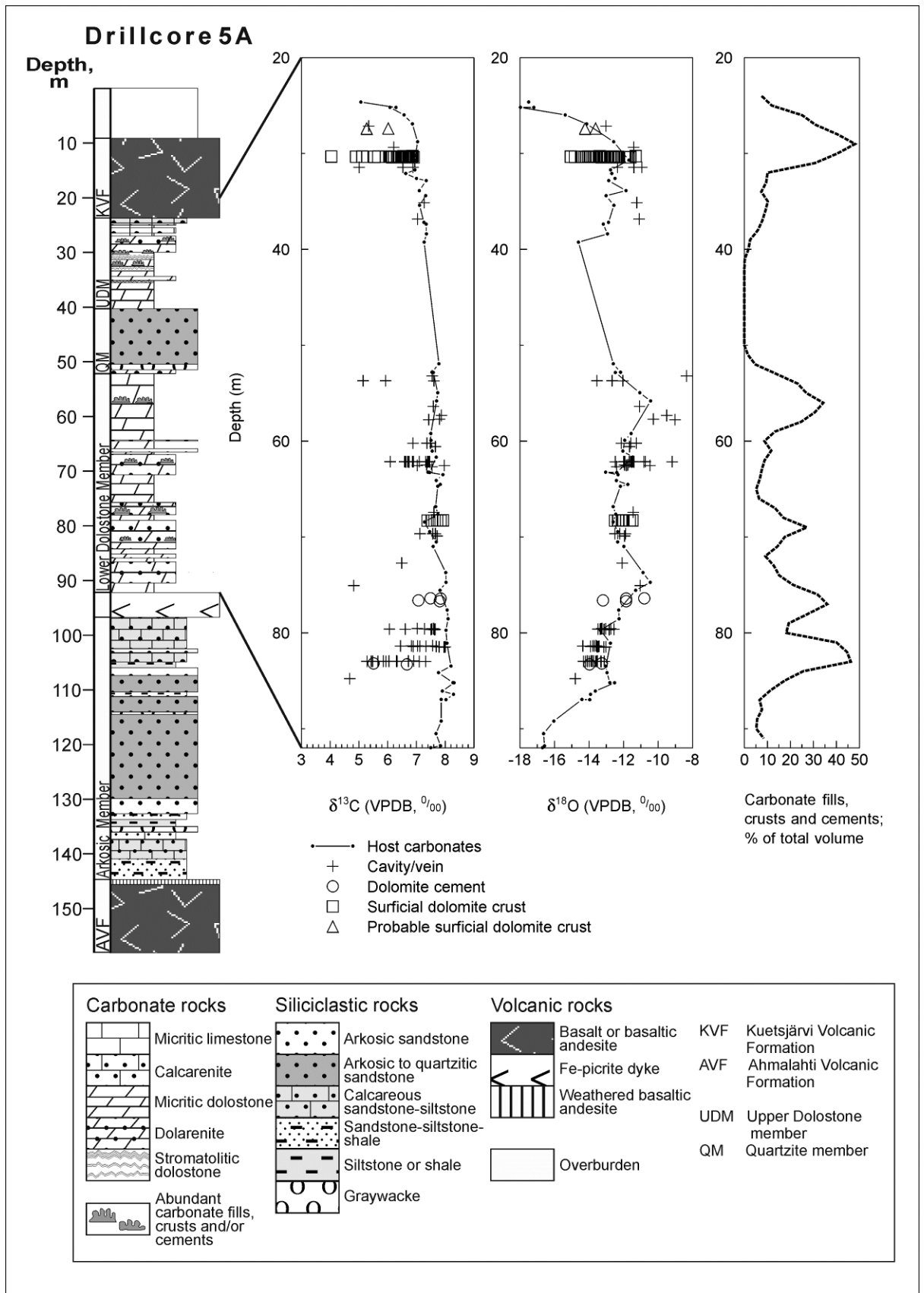


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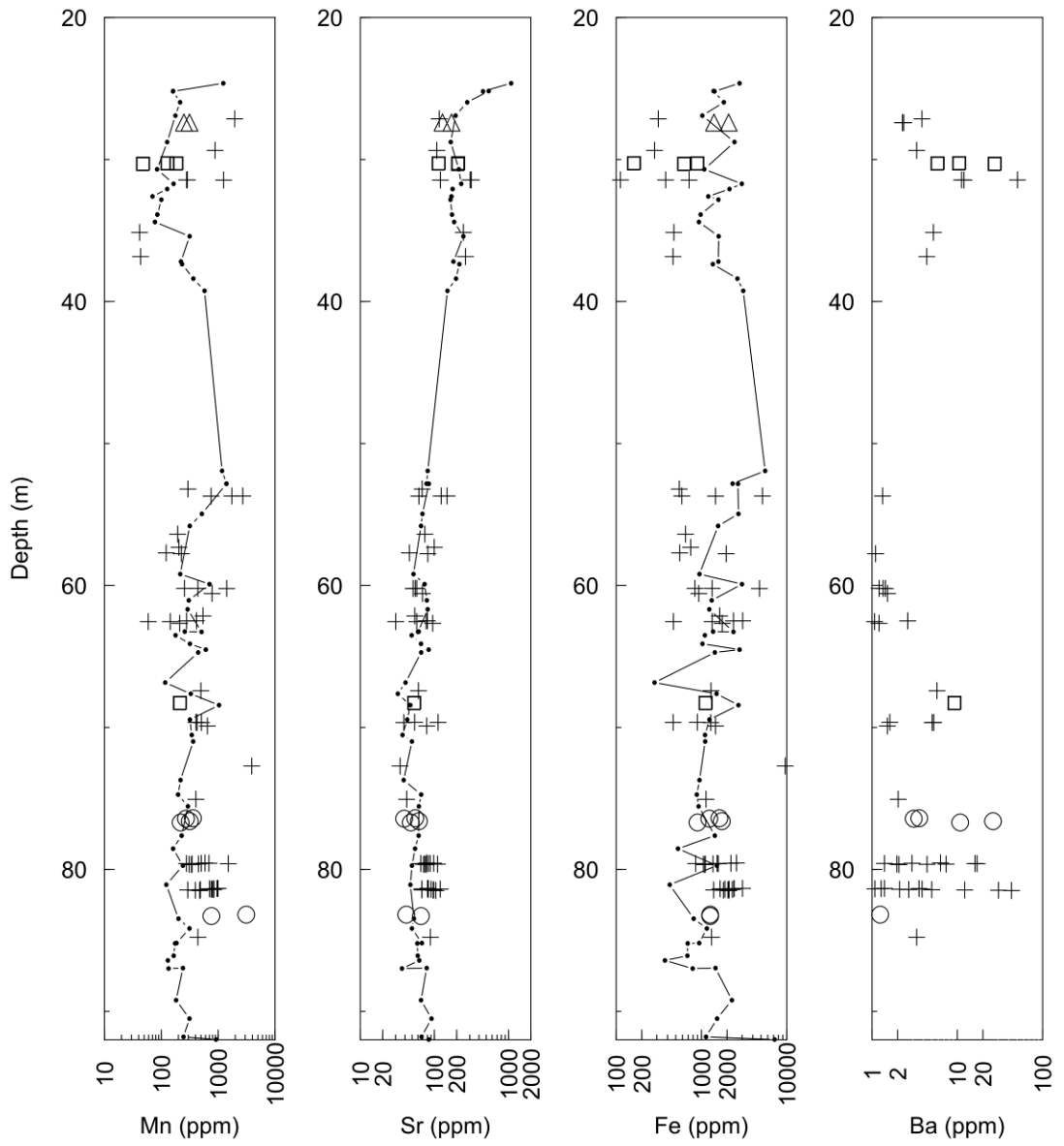


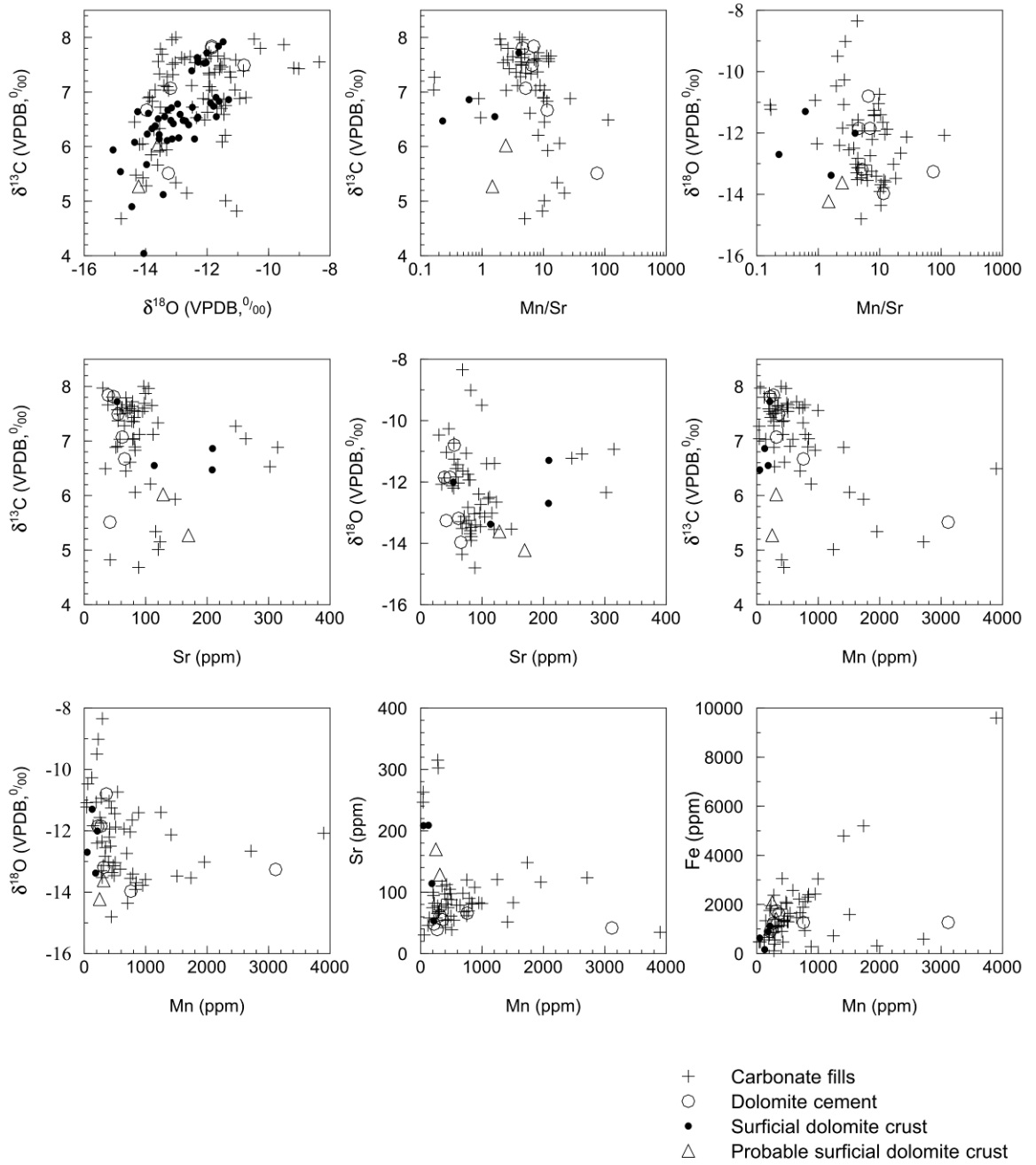




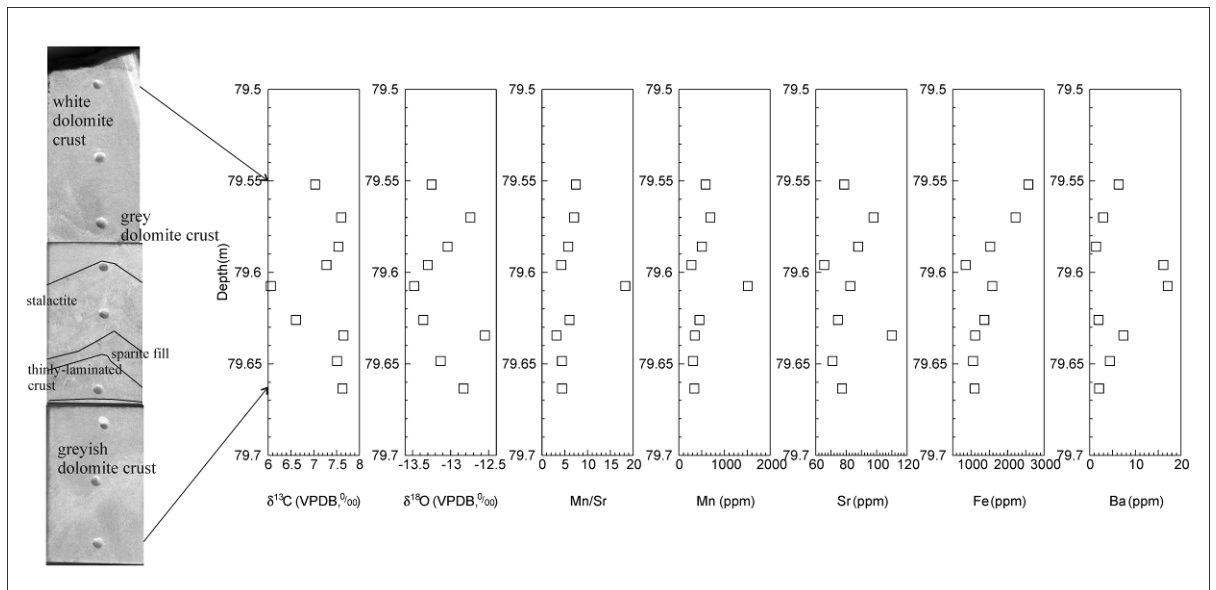


- Host carbonates (Salminen et al., 2013)
- + Cavity/vein
- Dolomite cements
- Surficial dolomite crusts
- △ Probable surficial dolomite crust

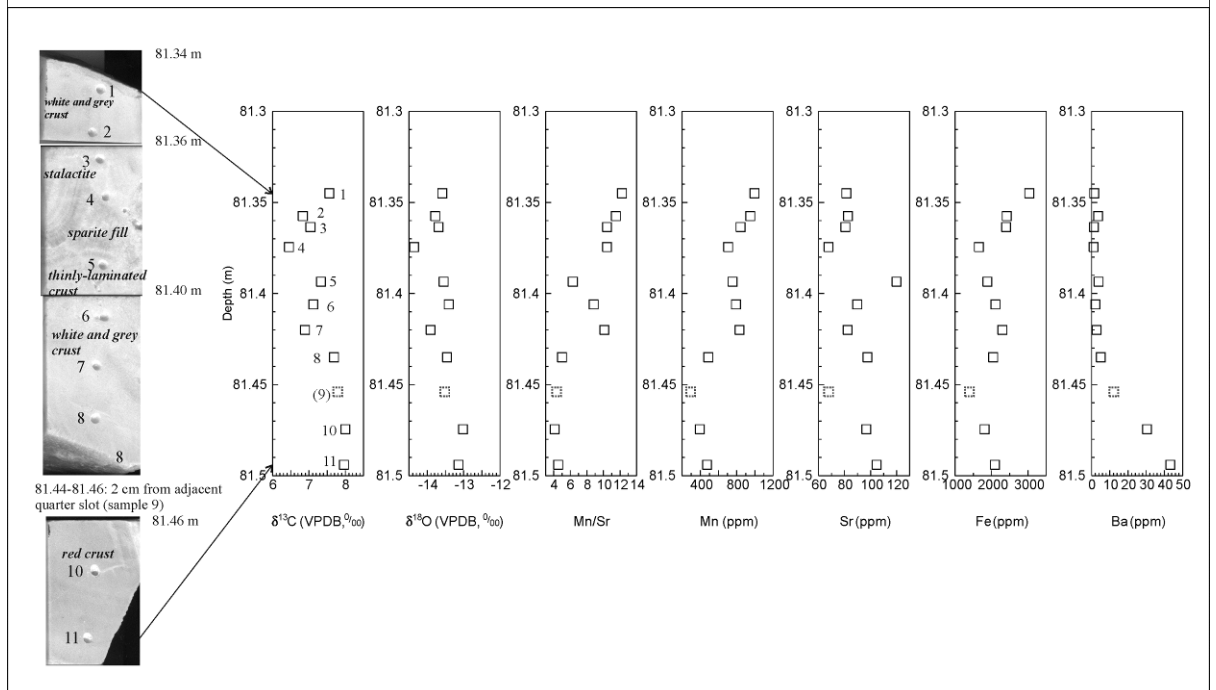




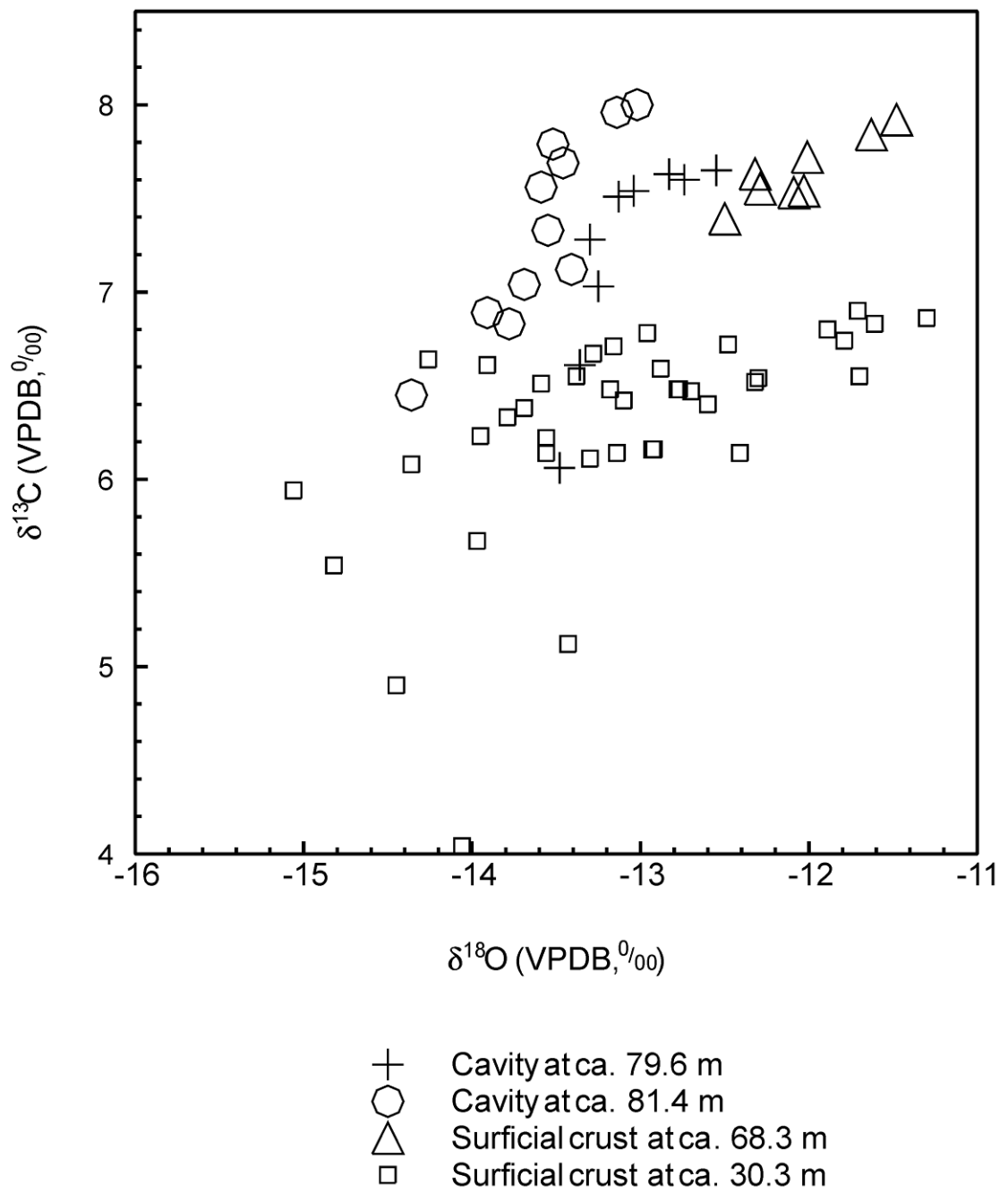
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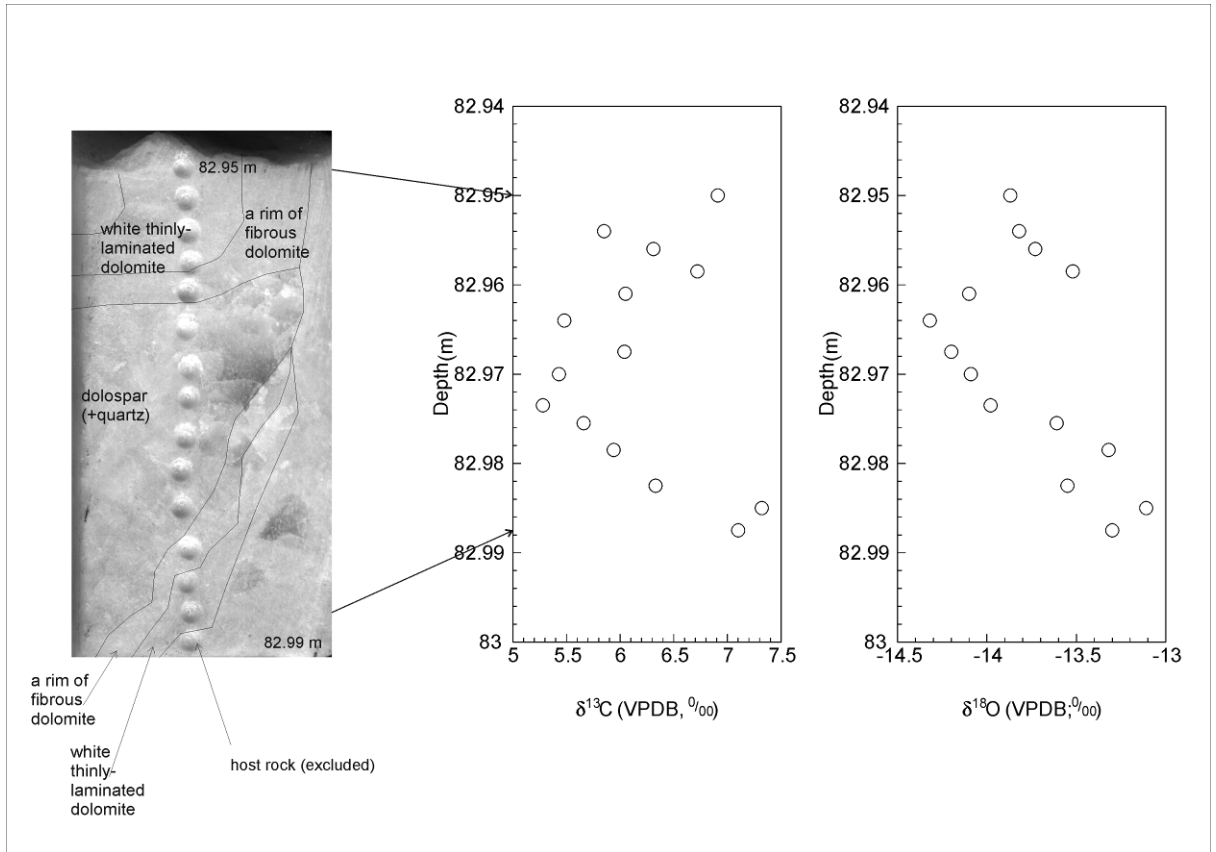
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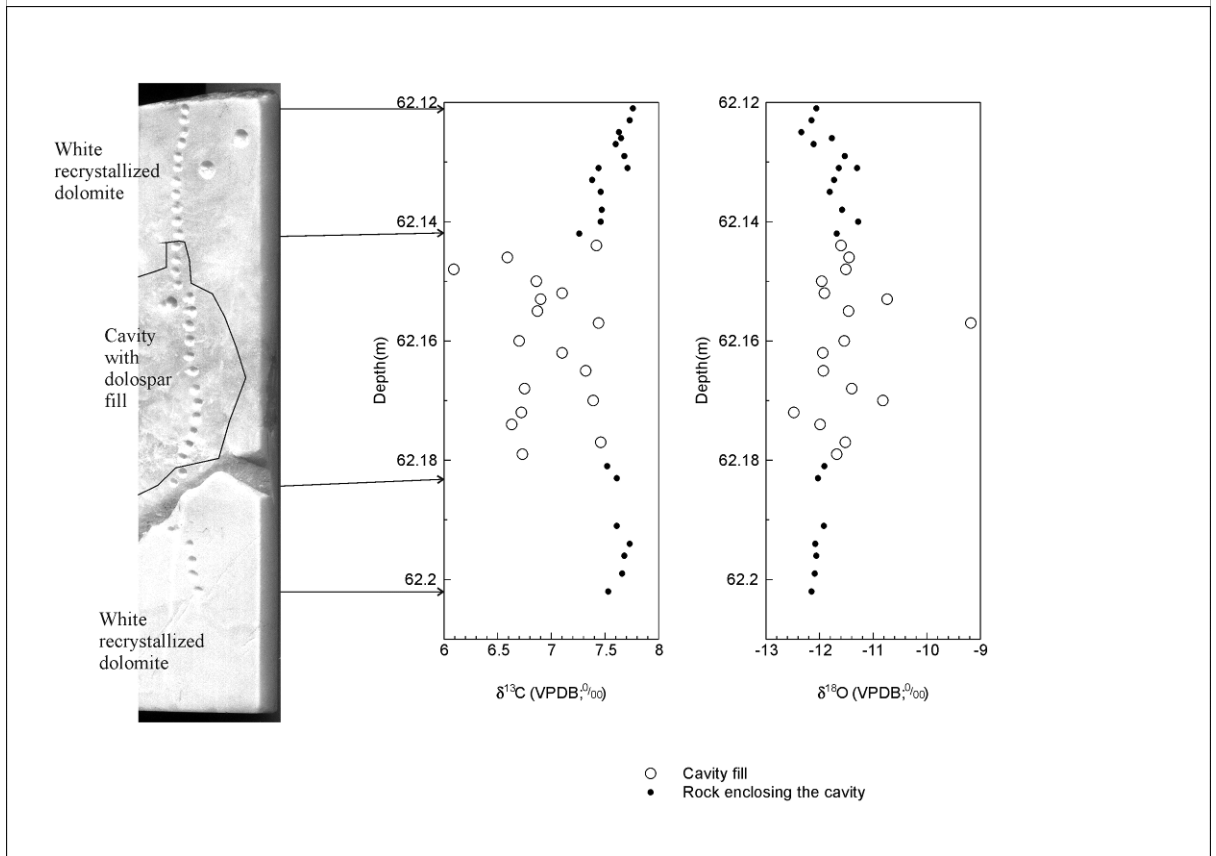
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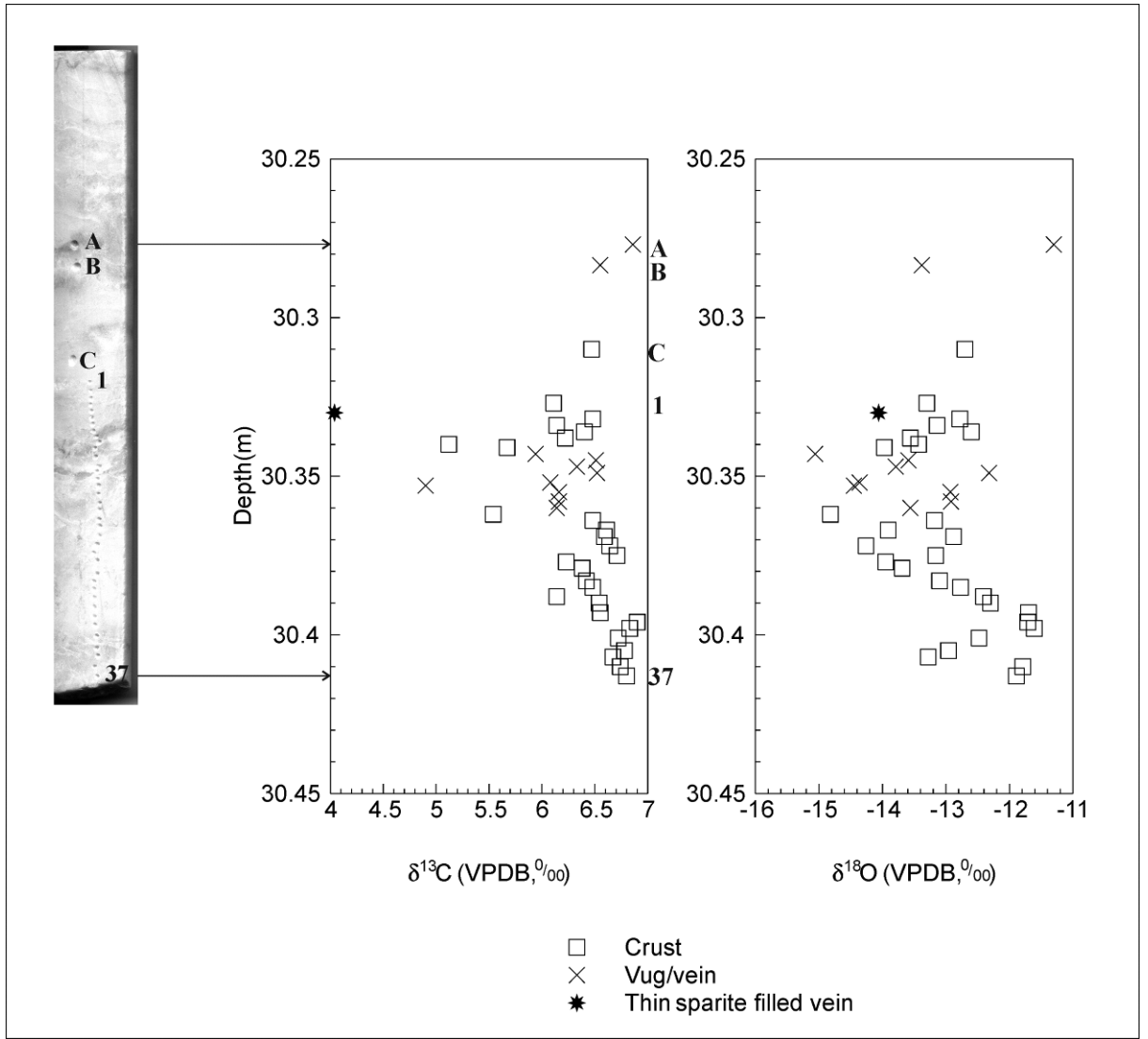
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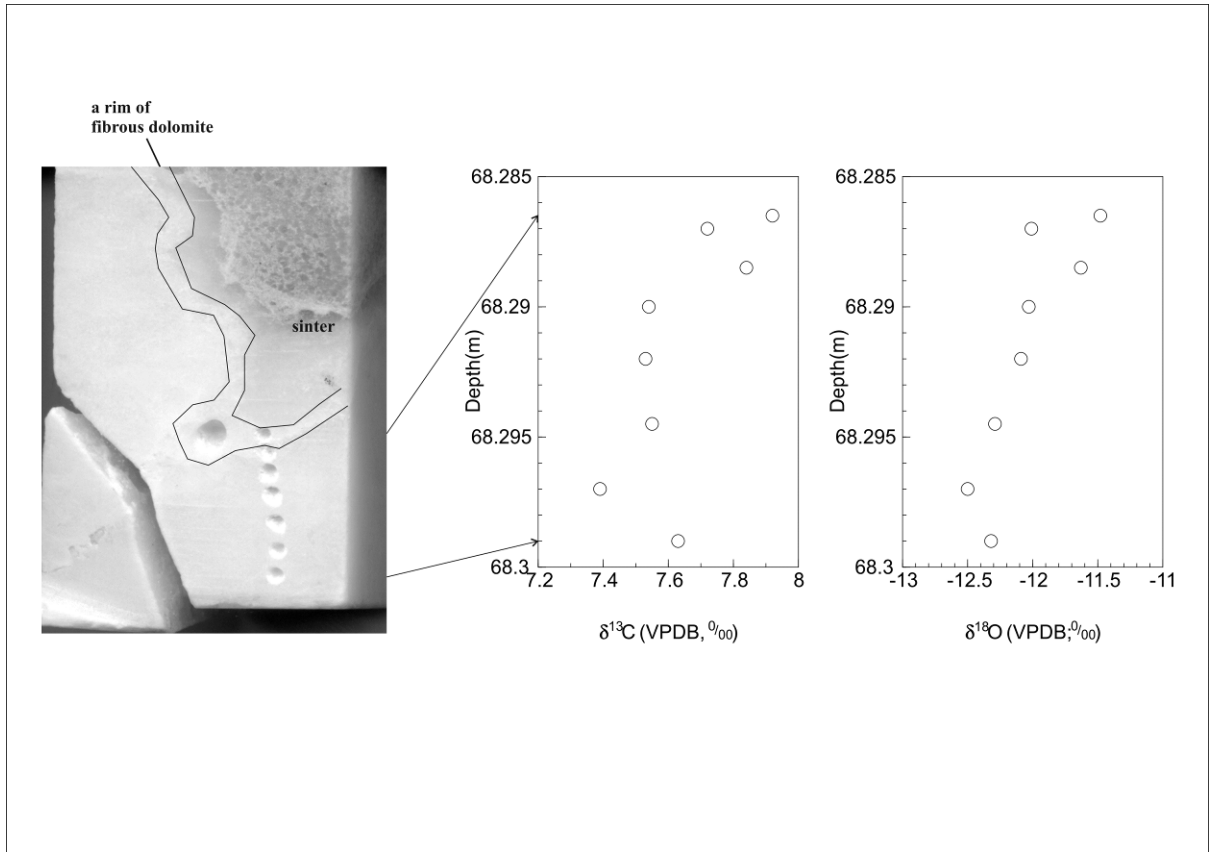
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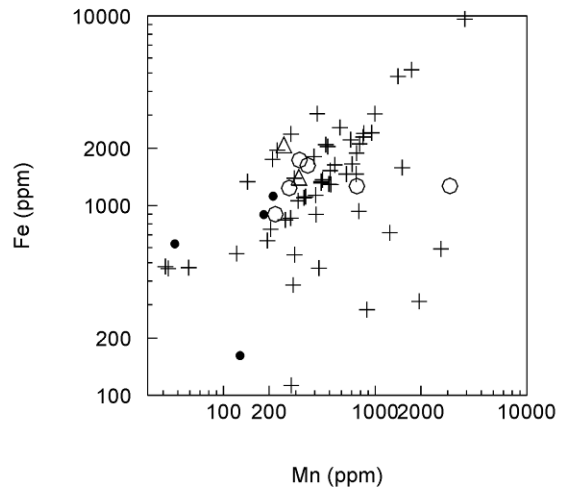
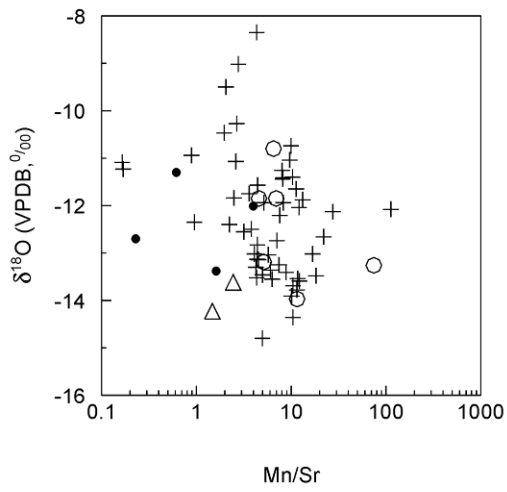
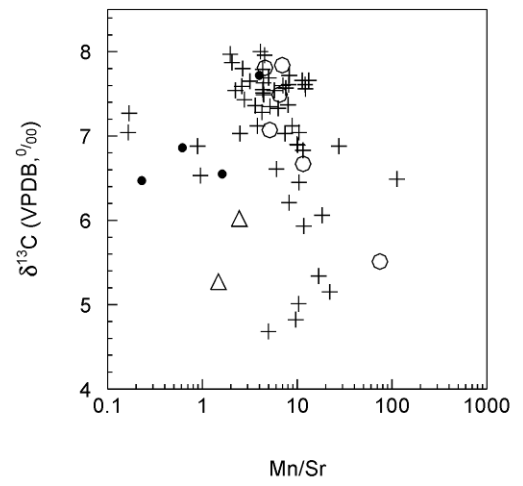
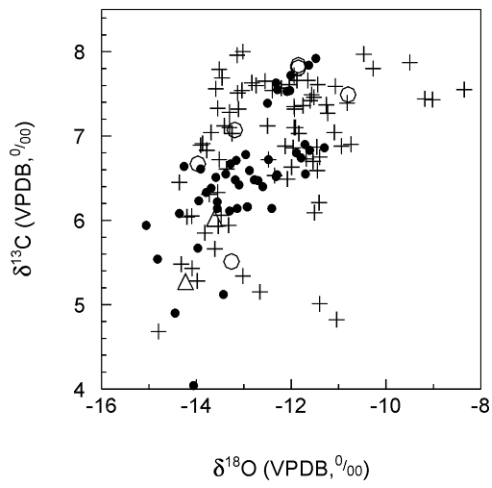
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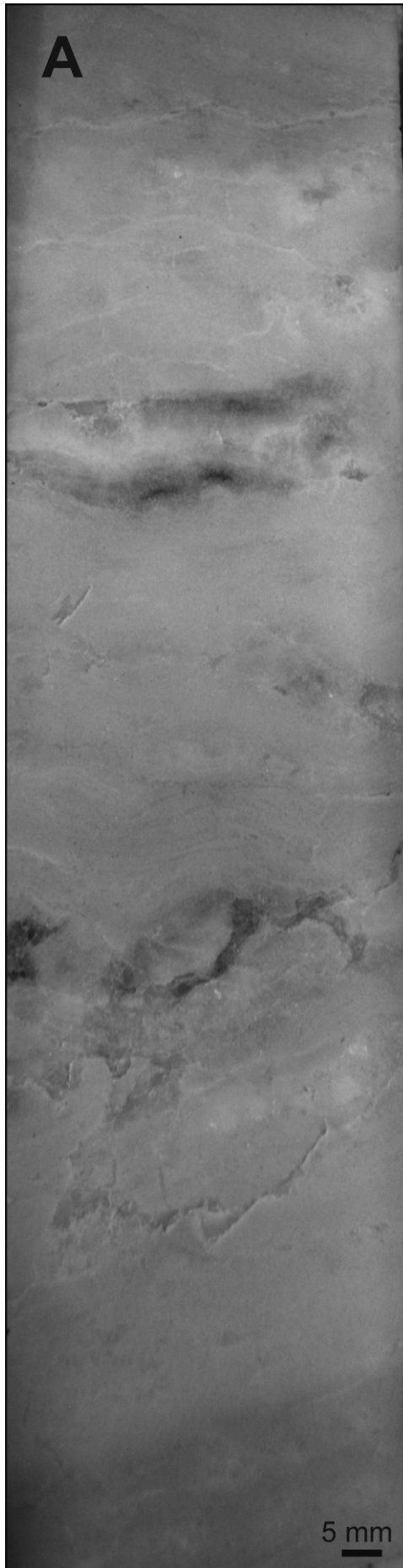
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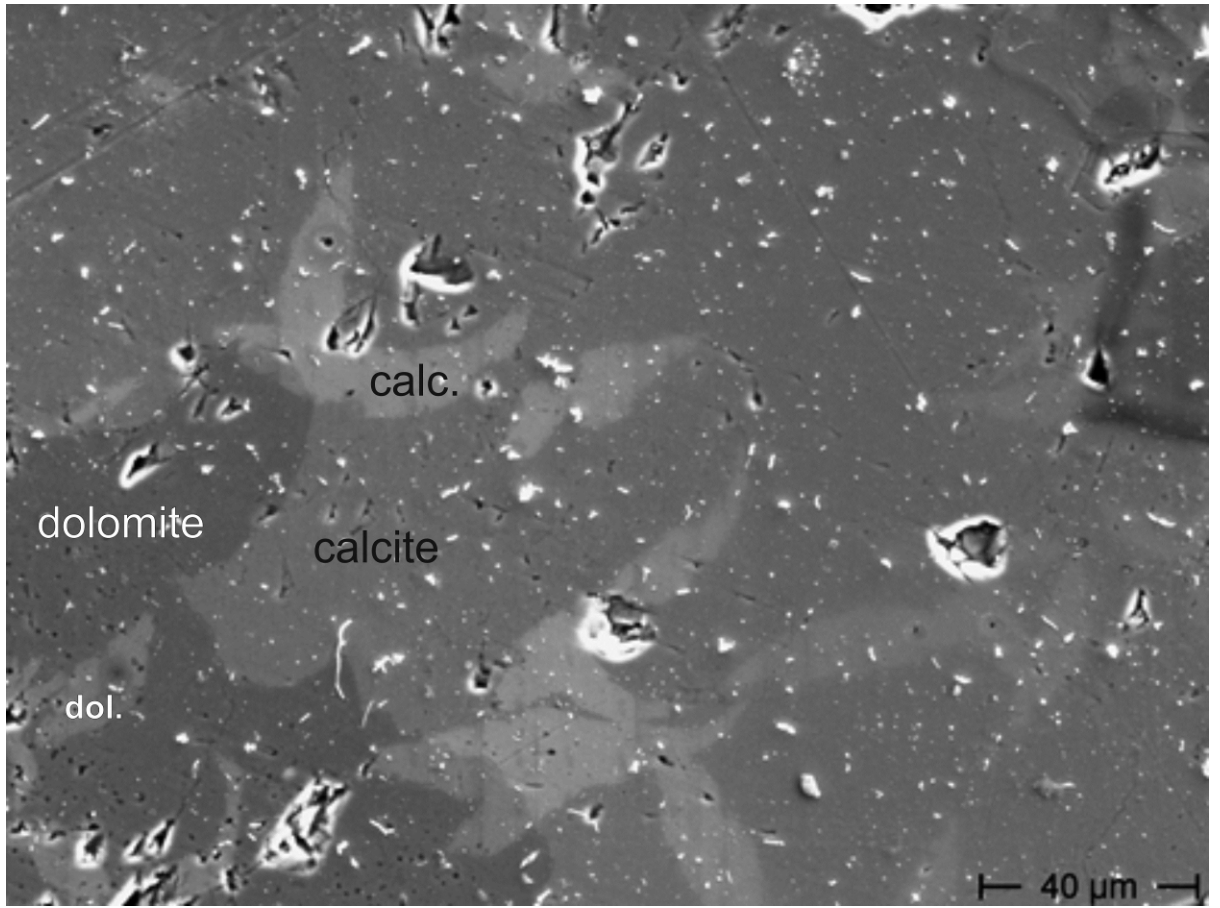


- + Carbonate fills
- Dolomite cement
- Surficial dolomite crust
- △ Probable surficial dolomite crust



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