

1 Depth of diamond formation obtained from single periclase  
2 inclusions

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21 **ABSTRACT**

22 Super deep diamonds (SDDs) are those that form between ~300 and ~1000 km in  
23 the Earth's mantle. They comprise only 1% of the entire diamond population but play a  
24 pivotal role in geology, as they represent the deepest direct samples from the interior of  
25 our planet. Ferropericlasite, (Mg,Fe)O, is the most abundant mineral found in SDDs and,  
26 when associated with low-Ni enstatite, which is interpreted as retrogressed bridgmanite,  
27 is considered proof of a lower-mantle origin. As this mineral association in diamond is  
28 very rare, the depth of formation of most ferropericlasite inclusions remains uncertain.  
29 Here we report geobarometric estimates based on both elasticity and elasto-plasticity  
30 theories for two ferropericlasite inclusions, not associated with enstatite, from a single  
31 Brazilian diamond. We obtained a minimum depth of entrapment of 15.7( $\pm$ 2.5) GPa at  
32 1830( $\pm$ 45) K [ $\approx$ 450( $\pm$ 70) km depth], placing the origin of the diamond-inclusion pairs at  
33 least near the upper mantle-transition zone boundary and confirming their super-deep  
34 origin. Our analytical approach can be applied to any type of mineral inclusion in  
35 diamond and will expectedly allow better insights into the depth distribution and origin of  
36 SDDs.

## 37 INTRODUCTION

38 Diamonds, and the mineral inclusions they trap during their growth, are pristine  
39 samples from the Earth's mantle and provide information on processes operating in  
40 inaccessible regions of our planet. This information is particularly valuable if it can be  
41 combined with depth estimates. Based on the mineral inclusions, the majority of  
42 diamonds (99%) originate within the lithosphere (Stachel and Harris, 2008). The other  
43 1% are sub-lithospheric and formed at depths between 300 and ~1000 km, and hence are

44 called super-deep diamonds (hereafter SDDs) (Walter et al., 2011; Pearson et al., 2014;  
45 Smith et al., 2016; Nestola et al., 2018).

46       Based on experimental evidence, bridgmanite and ferropericlase (fper) are the  
47 most abundant minerals in the lower mantle, comprising approximately ~75 and ~17  
48 wt%, respectively (Stixrude and Lithgow-Bertelloni, 2012, and references therein). On  
49 decompression, bridgmanite inverts to Al-rich, low-Ni enstatite (Stachel et al., 2000),  
50 while fper can remain stable to room pressure. Early inclusion work (Harte et al., 1999;  
51 Stachel et al., 2000) concluded that fper was a lower mantle mineral, especially when  
52 found in the same diamond as low-Ni enstatite, but the findings of fper in association  
53 with olivine and jeffbenite in some SDDs (Hutchison et al., 2001) cast doubt on that  
54 conclusion. Indeed, ringwoodite is in equilibrium with fper at 24 GPa (Brey et al., 2004)  
55 and it could have later reverted to olivine, whereas jeffbenite is only stable up to 13 GPa  
56 (Armstrong and Walter, 2012), even if its origin is still controversial. In addition, the  
57 observation of droplets of Fe-Ni alloys in some Fe-enriched fpers induced Hayman et al.  
58 (2005) to outline a model which ascribes the Fe-rich character to equilibration with  
59 silicates in the deeper part of the lower mantle (1700-2900 km). On the other side,  
60 synthesis of fper and diamond by carbonate melt-peridotite reactions (Thomson et al.,  
61 2016) suggested that fper inclusions with variable Fe contents can form at lower upper-  
62 mantle to transition-zone depths. The presence of nanometric exsolutions of  
63 magnesioferrite in some fper inclusions (Harte et al., 1999; Wirth et al., 2014; Kaminsky  
64 et al., 2015) lead Palot et al. (2016) to propose an origin in the uppermost part of the  
65 lower mantle, but Uenver-Thiele et al. (2017a,b) showed that magnesioferrite cannot  
66 exsolve directly from fper in the lower mantle.

67 In aiming to identify a method for determining the depth of origin of fper  
68 inclusions completely independent of mineral paragenesis, Hutchison (1997) combined  
69 sophisticated thermoelastic modelling with measurements of periclase cell parameters  
70 before and after release from diamonds from the São Luiz River, Juina, Brazil and  
71 Guinea. Hutchison and Harris (1998) reported an absolute minimum depth of formation  
72 of 320 km (equivalent to an entrapment pressure,  $P_{\text{trap}}$ , of 11 GPa) uncorrected for the  
73 brittle deformation evident in the diamond host. This study provided strong evidence for  
74 super deep origins for the samples analyzed, however, limitations were imposed by  
75 uncertainties in the Gandolfi camera measurement technique available at the time and full  
76 quantification of plastic and brittle diamond deformation. In this study, we have been  
77 able to extend the original work with improved certainty and propose an updated method  
78 for determining minimum  $P_{\text{trap}}$  applied to two fper inclusions in a further diamond from  
79 São Luiz (sample AZ1, Fig. 1). The reverse calculation of  $P_{\text{trap}}$  was performed by  
80 applying the elastic geobarometry approach (Angel et al., 2014; 2015a,b; 2017),  
81 including the full geometry of the inclusions based on a realistic 3D reconstruction  
82 (Mazzucchelli et al., 2018), coupled with a new elasto-plastic model to account for  
83 plasticity of the diamond host at high temperature.

## 84 **METHODS**

### 85 **Sample**

86 The diamond investigated in this study (Fig. 1) was recovered in the mid to late  
87 1980s from alluvial deposits of the São Luiz river in the Juina area of Mato Grosso State,  
88 Brazil. The sample contains two main black tabular inclusions, identified as fper  
89 [(Mg<sub>0.60</sub>Fe<sub>0.40</sub>)O; see below] by SCXRD (see Supplemental Information). The smaller

90 one, whose longest dimension is  $\sim 160 \mu\text{m}$ , is named AZ1\_1; the bigger one, whose  
91 longest dimension is  $\sim 340 \mu\text{m}$ , is named AZ1\_2.

## 92 **Synchrotron X-ray Tomographic Microscopy**

93 This non-destructive, high-resolution technique creates three-dimensional maps of  
94 the variations of the X-ray attenuation coefficient within a sample. X-ray micro-  
95 tomography experiments were carried out at the Swiss Light Source (SLS) at TOMCAT,  
96 a beamline for TOMographic Microscopy and Coherent rAdiology experimenTs  
97 (Stampanoni et al., 2006). Measurements were performed at 13.5 keV in order to  
98 maximize contrast. A total of 1501 X-ray radiographs were acquired from different  
99 angular positions around a vertical rotation axis for each sample. The used imaging setup  
100 consisted of a  $20 \mu\text{m}$  thick LuAG:Ce scintillator screen, a  $20\times$  objective and a sCMOS  
101 (PCO.edge) camera. The tomographic reconstruction was performed using optimized  
102 routines based on the Fourier Transform Method (Marone and Stampanoni, 2012). The  
103 resulting volume consisted of 2160 axial slices of  $2560 \times 2560$  pixels, with a pixel size of  
104  $0.33 \mu\text{m}$ .

## 105 **Single-Crystal X-ray Diffraction (SCXRD)**

106 SCXRD measurements were performed on the fper inclusions both before and  
107 after release from their diamond host at the Department of Geosciences (University of  
108 Padova). X-ray data were collected using a Rigaku Oxford Diffraction *SuperNova* single-  
109 crystal diffractometer, equipped with a Dectris *Pilatus* 200 K area detector and with a  
110 *Mova* X-ray microsource. A monochromatized MoK $\alpha$  radiation ( $\lambda = 0.71073 \text{ \AA}$ ), working  
111 at 50 kV and 0.8 mA, was used. The sample-to-detector distance was 68 mm. Data  
112 reduction was performed using the CrysAlisPro software (Rigaku Oxford Diffraction).

113 **Field Emission Gun—Scanning Electron Microscopy (FEG-SEM)**

114 The two fper inclusions were first extracted by mechanical crushing of the host,  
115 then polished in a three-step process and finally carbon coated. FEG-SEM measurements  
116 were carried out at the Department of Physics and Astronomy (University of Padova),  
117 using a Zeiss SIGMA HD FEG-SEM microscope operating at 20 kV, with a spot size of  
118 ~1 nm. Imaging was performed using an InLens secondary electron detector.

119 Compositional analysis was performed using an energy dispersive X-ray spectrometer  
120 (EDX by Oxford Instruments). The spatial resolution in microanalysis was of ~1  $\mu\text{m}$ .

121 **Finite Element (FE) analysis**

122 The FE analysis was performed on the real 3D model built from the segmentation  
123 of the X-ray microtomographic data (Fig. 2). The surface of the model was smoothed to  
124 improve the quality of the final FE mesh and the final 3D model was then assembled  
125 placing the two inclusions in the diamond host. An elastically isotropic analysis was run  
126 with Simulia Abaqus, a commercial engineering package for FE analysis (for more  
127 details see Mazzucchelli et al., 2018). For the fper inclusions we used the isothermal bulk  
128 modulus  $K_{0\text{TR}} = 162(14)$  GPa from the Equation of State (EoS) as reported in Angel et al.  
129 (2017). This EoS was obtained fitting the original  $P$ - $V$ - $T$  data of Mao et al. (2011) up to  
130 2000 K and 50 GPa using a 3<sup>rd</sup>-order Birch-Murnaghan EoS combined with a Berman-  
131 type thermal expansion. The Reuss shear modulus  $G_{0\text{R}} = 87(2)$  GPa was obtained from  
132 the elastic constants reported by Jacobsen et al. (2002) for a fper with composition  
133  $(\text{Mg}_{0.63}\text{Fe}_{0.37})\text{O}$  that is close to the composition of our inclusions,  $(\text{Mg}_{0.60}\text{Fe}_{0.40})\text{O}$ . For  
134 diamond we used the  $K_{0\text{TR}} = 444(2)$  GPa from the  $P$ - $V$ - $T$  EoS reported by Angel et al.  
135 (2015a) and the  $G_{0\text{TR}} = 535$  GPa reported by Angel et al. (2015b).

## 136 **Elasto-plastic Model**

137       The calculation is split into two steps dividing the calculation into an isothermal,  
138 quasi-static decompression from  $P_{\text{trap}}, T_{\text{trap}}$  to  $P_{\text{room}}, T_{\text{trap}}$ , followed by an isobaric cooling  
139 to room temperature. The model is solved by inversion. The host-inclusion system is  
140 initially at  $P_{\text{room}}, T_{\text{room}}$  with the inclusion at the experimentally measured  $P_{\text{inc}}^{\text{exp}}$ . First, an  
141 entrapment temperature ( $T_{\text{trap}}$ ) is chosen and the over-pressure  $P_{\text{inc}}^{P_{\text{room}}, T_{\text{trap}}}$  developed in  
142 the inclusion during isobaric heating to  $P_{\text{room}}, T_{\text{trap}}$  is calculated adjusting the elastic  
143 properties of the host and the inclusion according to their EoS. A  $P_{\text{trap}}$  is guessed at the  
144 chosen  $T_{\text{trap}}$ , and the elasto-plastic deformation of the host and inclusion pressure are  
145 calculated during the quasi-static decompression of the host from  $P_{\text{trap}}, T_{\text{trap}}$  to  $P_{\text{room}}, T_{\text{trap}}$   
146 according to Campione (2018). The guessed  $P_{\text{trap}}$  is adjusted until the pressure calculated  
147 in the inclusion at  $P_{\text{room}}, T_{\text{trap}}$  matches the previously found  $P_{\text{inc}}^{P_{\text{room}}, T_{\text{trap}}}$ . The elastic  
148 properties for diamond are from Angel et al. (2015a) and from Zouboulis et al. (1998).  
149 The variation of  $\sigma_Y$  with  $T$  (between 1273 and 1823 K) was obtained from Weidner et al.  
150 (1994). The EoS of the inclusion is from Angel et al. (2017) as discussed in the main text.

## 151 **RESULTS AND DISCUSSION**

### 152 **Sample Analysis**

153       The 3D reconstruction (Fig. 2) revealed the absence of significant fractures at  
154 inclusion terminations. However, graphitization in haloes around the inclusions (Fig. 1)  
155 suggests that some pressure release by brittle deformation of the host diamond may have  
156 occurred. Both inclusions after release and polishing exhibited pervasively and  
157 homogeneously distributed exsolutions of magnesioferrite of about ~200 nm size, which  
158 often coalesced into chains of 2–3  $\mu\text{m}$  length and constituted ~6% of the total surface

159 area (calculated using the ImageJ software, Abràmoff et al., 2004). EDX analyses gave a  
160 composition of  $(\text{Mg}_{0.61}\text{Fe}_{0.39})\text{O}$  for AZ1\_1 and  $(\text{Mg}_{0.59}\text{Fe}_{0.41})\text{O}$  for AZ1\_2; therefore we  
161 consider them to have a similar approximate composition of  $(\text{Mg}_{0.60}\text{Fe}_{0.40})\text{O}$  (Figure  
162 DR1).

### 163 **Inclusion Residual Pressures**

164 X-ray analyses (Figure DR2) provided the lattice parameters and the relative unit-  
165 cell volumes reported in Table DR1. By comparing the unit-cell volumes before ( $V$ ) and  
166 after ( $V_0$ ) release from the diamond host and using the  $P$ - $V$ - $T$  equation of state (EoS) for  
167 fper reported in Angel et al. (2017), we obtained a residual pressure,  $P_{\text{inc}}$ , of  $1.84(\pm 0.65)$   
168 GPa for inclusion AZ1\_1 and of  $1.48(\pm 0.67)$  GPa for inclusion AZ1\_2. The high  
169 uncertainties in  $P_{\text{inc}}$  are due to the high uncertainty in the bulk modulus value of fper  
170 (Mao et al., 2011). Values of  $P_{\text{inc}}$  are consistent with  $1.29 (\pm 0.38)$  GPa for the Guinean  
171 diamond of Hutchison and Harris (1998) where in this case the uncertainty is confined to  
172 that of measurement of cell parameters.

### 173 **Depth of Formation of the Ferropericlase—Diamond Pair by Elasto-plastic**

#### 174 **Geobarometry**

175 Given the absence of significant fracture systems around the inclusions, the  
176 calculated  $P_{\text{inc}}$  can be linked to the depth of formation by elastic geobarometry. Standard  
177 elastic methods rely on simplified models which assume the inclusion is spherical and  
178 sitting isolated in an infinitely large host (e.g., Zhang, 1998). Mazzucchelli et al. (2018)  
179 extended the model to non-spherical inclusions and showed that platy inclusions develop  
180 a lower  $P_{\text{inc}}$  compared to more rounded inclusions. This is consistent with our  
181 measurements, which show a lower  $P_{\text{inc}}$  for the platy AZ1\_2 than for the more rounded



182 AZ1\_1. The method of Mazzucchelli et al. (2018) enabled us to calculate the appropriate  
183 geometrical correction factor ( $\Gamma$ ) for the two inclusions through an integration over their  
184 entire volumes. The  $\Gamma$  factors obtained in this way are  $-0.016(5)$  and  $-0.080(10)$  for  
185 inclusions AZ1\_1 and AZ1\_2, respectively. Applying the correction factor to our  
186 experimental determined residual pressures we obtained the corrected  $P_{\text{inc}}$  of  $1.87(\pm 0.66)$   
187 GPa and  $1.61(\pm 0.73)$  GPa for the inclusions, respectively.

188 We then calculated the entrapment isomeke for the two fper–diamond pairs using  
189 the corrected values for  $P_{\text{inc}}$  and the software EosFit-Pinc (Angel et al., 2017). Since both  
190 the host and the inclusion have cubic crystallographic symmetry, the effect of anisotropic  
191 elasticity is limited (see Anzolini et al., 2018), allowing the use of current isotropic  
192 elastic geobarometry models. To maintain consistency with the calculation of the  
193 geometrical factors, we used the  $P$ - $V$ - $T$  EoS of fper and diamond and the shear modulus  
194  $G_{\text{0TR}} = 535$  GPa of diamond all respectively reported previously (Angel et al., 2015a,b;  
195 2017). The intersection of the isomeke with the mantle adiabat, accounting for the  
196 isomeke and the adiabatic uncertainties, gave an entrapment pressure for AZ1\_1 of  $P_{\text{trap}} =$   
197  $13.5(\pm 1.8)$  GPa at a temperature  $T_{\text{trap}} = 1802(60)$  K and for AZ1\_2 of  $P_{\text{trap}} = 12.8(\pm 1.8)$   
198 GPa at a temperature  $T_{\text{trap}} = 1794(60)$  K (see Table DR2 and Fig. 3).

199 This estimate does not take into account plastic deformation in the diamond,  
200 which may accommodate part of the inclusion expansion during uplift to surface  
201 (Anzolini et al., 2016). Plastic deformation is well documented in diamond and,  
202 particularly, in SDDs (e.g., Cayzer et al., 2008), consistent with its low yield strength  
203 ( $\sigma_Y$ ) at high temperatures (Weidner et al., 1994). Therefore, the  $P_{\text{trap}}$  calculated from a  
204 purely elastic model is likely to be underestimated. To account for plastic deformation,

205 the elasto-plastic (EP) model for barometry proposed by Campione (2018) (see Methods)  
206 was applied to these data. The reverse calculation of  $P_{\text{trap EP}}$  as a function of  $T$  was solved  
207 by adjusting the  $\sigma_Y$  of diamond according to the experimental measurements of Weidner  
208 et al. (1994) and the elastic parameters for diamond and fper, previously noted. Since the  
209 EP model assumes that the inclusion is spherical, we applied this method only to the most  
210 rounded of the two inclusions, i.e., AZ1\_1. The best agreement between the calculated  
211  $P_{\text{trap EP}}(T)$  and the adiabat, with its uncertainty, is at 15.7( $\pm 2.5$ ) GPa and 1830( $\pm 45$ ) K  
212 [ $\approx 450(\pm 70)$  km depth]. Considering the uncertainties, this result is compatible with an  
213 origin in the lowermost upper mantle or, more probably, in the upper transition zone (Fig.  
214 3). Unfortunately, the depth obtained is constrained by a lack of experimental values of  
215  $\sigma_Y$  when temperatures are higher than  $\approx 1850$  K (Weidner et al. (1994) and the fact that  
216 the EP model only considers the deformation caused by over-pressurization of the  
217 inclusion with respect to the external lithostatic pressure (Campione, 2018). If external  
218 tectonic stresses act on diamonds during uplift through the sub-lithospheric mantle, they  
219 may promote additional plastic deformation, which may contribute to the release of part  
220 of the  $P_{\text{inc}}$  being built on the inclusion. Therefore, the  $P_{\text{trap EP}}$  value of 15.7( $\pm 2.5$ ) GPa for  
221 AZ1\_1, which corresponds to a depth of  $\approx 450(\pm 70)$  km, should be regarded as a  
222 minimum estimate.

223 In addition, models used in this work do not take into account the effect that the  
224 magnesioferrite exsolutions (see Supplementary Information) may have on  $P_{\text{inc}}$  and, in  
225 turn, on the calculated  $P_{\text{trap}}$ . However, given the small contrast in elastic properties  
226 between fper and magnesioferrite (Reichmann and Jacobsen, 2004) and the small volume

227 ratio (~6%) between these two minerals, the effect is probably limited and well within the  
228 uncertainties already accounted for in the calculations.

## 229 **CONCLUSIONS**

230 Our newly devised method for inclusion barometry, which incorporates an elasto-  
231 plastic treatment of the inclusion–host system and a correction for geometrical effects  
232 based on a real 3D model, can be extended to several other types of inclusions trapped in  
233 SDDs. Our analyses demonstrate that for the fper samples studied an origin much  
234 shallower than the transition zone (Thomson et al., 2016) can certainly be excluded.  
235 Although the present results may only provide minimum pressure estimates, they yield  
236 valuable constraints independent of mineral phase relations on the depth of origin of  
237 SDDs and thereby increase our knowledge of those inaccessible regions which play a key  
238 role in the Earth’s dynamics and deep carbon cycle.

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396

397 **FIGURE CAPTIONS**

398

399 Figure 1. The inclusion-bearing diamond studied in this work.

400

401 Figure 2. 3D model of the ferropericlasite inclusions built from the segmentation of the X-  
402 ray microtomographic dataset. It preserves the morphology of the two inclusions and  
403 their mutual distances and orientations and reveals the absence of significant fractures  
404 around the inclusions. The pressure calculated by FE analysis is not homogeneous within  
405 the inclusions. The final residual pressures ( $P_{inc}$ ) reported in the text are obtained for each  
406 inclusion as the average of the pressure over their entire volume and include also the  
407 uncertainty in the calculation.

408

409 Figure 3. Minimum entrapment pressures of the ferropericase inclusions determined by  
410 elastic and elasto-plastic models. The geotherm is calculated for a typical cratonic surface  
411 heat flow of 40 mW/m<sup>2</sup> (Hasterok and Chapman, 2011) and a mantle adiabat (Katsura et  
412 al., 2010; Trubitsyn and Trubitsyna, 2015). Entrapment pressures ( $P_{\text{trap}}$ ) calculated for  
413 inclusions AZ1\_1 and AZ1\_2 at various  $T$  with the purely elastic model are represented  
414 by blue and green diamonds, respectively. The  $P_{\text{trap EP}}$  calculated with the elasto-plastic  
415 model for inclusion AZ1\_1 at  $T$  consistent with the adiabat, and its uncertainty, is  
416 represented by the orange box.

417

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