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Monitoring the time-averaged discharge rates, volumes and emplacement style of large lava flows by using MIROVA system: the case of the 2014-2015 eruption at Holuhraun (Iceland)

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25 **Abstract**

26 The 2014-2015 eruption at Holuhraun has produced more than 1.5 km³ of basaltic magma and can be
27 considered one of the major effusive events of the last two centuries in the world. During this eruption
28 the MIROVA system (a volcanic hot-spot detection system based on MODIS middle infrared – MIR
29 - data) has been used to detect and locate the active portions of the lava flow(s), and to measure the
30 heat radiated by the growing lava field. According to these data the eruption was characterized by a
31 slow decay of the radiant power, accompanied by a change in the lava transport mechanism that
32 shifted from open channels, at the beginning of the eruption, to lava tubes, during the last months of
33 activity. Despite the evident evolution of lava transport mechanism, we found that the overall
34 decreasing trend of the thermal flux was mainly controlled by the exponential decline of lava
35 discharge rates, while the increasing insulation of the flow field had a strong impact in transporting
36 efficiently the lava at the distal flow front(s). Our results suggest the apparent time averaged lava
37 discharge rates (*TADR*), derived from satellite thermal data, may fluctuate around the real effusion
38 rate at the vent, especially in the case of large lava flows emplacing in nearly flat conditions. The
39 magnitude and frequency of these fluctuations are mainly controlled by the emplacement dynamic,
40 (i.e. occurrence of distinct major flow units), while the transition from channel- to tube-fed lava
41 transport mechanism play only a minor role ($\pm 30\%$) in the retrieval of *TADR* using MIR data . When
42 the *TADR* values are integrated to calculate erupted lava volumes, the effects of pulsating
43 emplacement dynamic become smoothed and the eruptive trend become more clear.

44 We suggest that during the Holuhraun's eruption, as well as during many other effusive eruptions,
45 the MIR-derived radiant flux essentially mimic the overall trend of lava discharge rates, with only a
46 minor influence due to the emplacement style and evolving eruptive conditions. From a monitoring
47 and operational perspective, MIROVA demonstrates to be a very valuable tool to follow and,
48 possibly, forecast the evolution of on-going effusive eruptions.

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1. Introduction

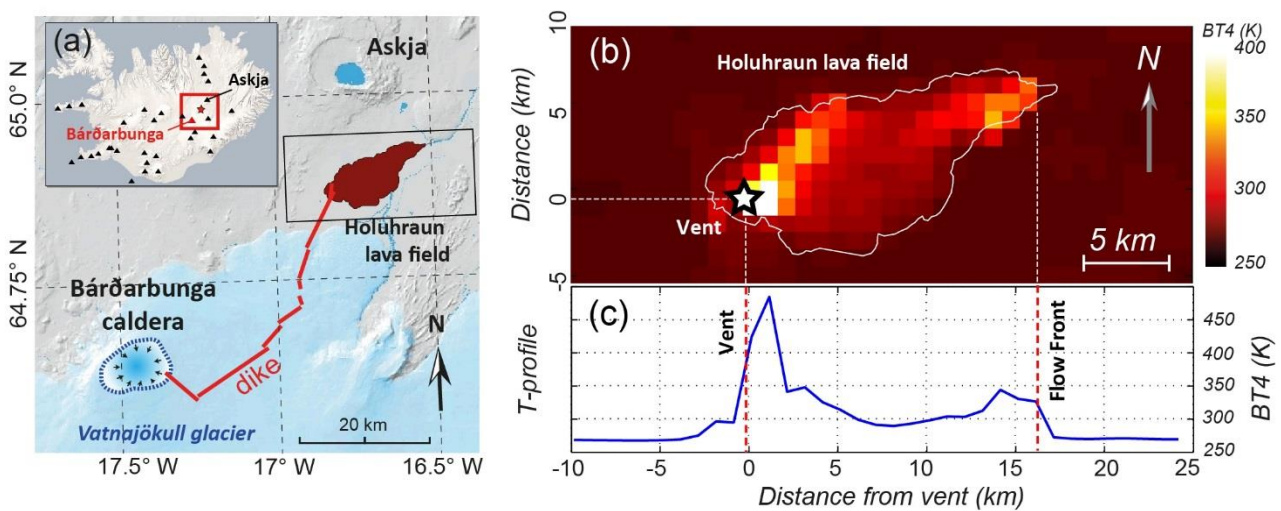
On August 29th 2014, one of the largest effusive eruption of the last 3 centuries took place along the Eastern Volcanic Zone (EVZ) of Iceland, about 45 km north-east of Bárðarbunga volcanic system (Gudmunsson, et al., 2014). The new eruption followed 15 days of sustained seismicity and accompanied the propagation of a 45 km long dike (Sigmundsson et al., 2015) that unlocked an historical eruptive path named Holuhraun (Sigurðsson and Sparks, 1978). The emission of lava persisted at very high rate for 180 days up to February 27th, 2015 when the activity was declared over (Gislason et al., 2015). The eruption was characterized by a slow decay of effusion rate (Coppola et al., 2017) that accompanied a coeval slow subsidence of the Bárðarbunga caldera (Gudmundsson et al., 2016). The clear link between the two processes provided evidences of an inelastic eruption whereby the lateral magma withdrawal was essentially linked to the gravity-driven collapse of the summit caldera (Coppola et al., 2017, Gudmundsson et al. 2016). About 84 km² of nearly flat land was covered by the new lava field that reached a maximum extend of about 18 x 5 km (Figure 1) and thickness up to 40 m (Gislason et al., 2015, Pedersen et al. 2017).

The prohibitive environmental conditions that might characterize these latitudes made field observations at the eruptive site very hard, especially on a continuous basis and for several months during the winter time. In these conditions, space-based thermal data have been extremely useful since they allowed a safe detection, location and quantification of the radiant flux produced by the effusive activity. In particular, infrared data acquired by MODIS (Moderate Resolution Imaging Spectroradiometer) and elaborated through the MIROVA volcanic hotspot detection system, (Coppola et al., 2016; www.mirovaweb.it), were delivered to the Icelandic Meteorological Office (IMO) as part of its daily operational monitoring activity of the Holuhraun eruption.

Since the beginning of the eruption the main tasks of satellite thermal data were: (i) to provide information about the location of active lava flow areas and front, (ii) to give an estimation of lava discharge rates and erupted volumes and finally (iii) to identify the ongoing effusive trend (steady,

75 waning or waxing). The latest two points became a priority after a couple of months when the winter
 76 season and the low-light conditions made the eruptive site hardly accessible and field observations
 77 became more and more rare and difficult. The calculation of effusion rates from satellite thermal data
 78 (the so called “thermal proxy”), become one of the main topic discussion among a thermal remote
 79 sensing group, established in that occasion by Dr. Thor Thordarson (Harris et al., 2016a). One of the
 80 main questions discussed within the group was “whether the subtle, but persistent decay of thermal
 81 flux indicated by satellite measurements, was caused by a real, slow decline of the effusion rate or
 82 reflected the increasing insulation of the growing lava field and the development of a lava tube system
 83 (Pedersen et al., 2017)

84



85

86 **Figure 1.** (a) Location of the Bárðarbunga-Holuhraun volcanic system showing the ice covered, 10-km-wide caldera of
 87 Bárðarbunga and the 2014-2015 lava field at Holuhraun. (b) Example of MODIS-derived thermal image (spatial
 88 resolution of 1 km) showing the Brightness Temperature at 4 μm (BT4) recorded on February 03rd, 2015 over the
 89 Holuhraun lava field (white line). The star indicates the approximate location of the vent. (c) Temperature profile along
 90 W-E direction, obtained by calculating the maximum BT4 along each vertical (N-S) line. The position of the flow front
 91 is marked by the easternmost thermally anomalous pixel (see the text for details).

92

93

94

95 In this paper we present time-averaged lava discharge rates and volumes, derived by using MIROVA
96 system, and we compare the results with field and independent measurements collected during and
97 after the Holuhraun eruption. The comparison reveals the role of changing emplacement style in the
98 lava discharge rates calculation and outlines the contribution of MIROVA in safely tracking this large
99 effusive eruptions from space.

100

101 **2. Time-averaged discharge rates from satellite thermal data: background and open** 102 **questions**

103 The relationship between effusion rates and thermal emissions of lava flows has received increasingly
104 attention during the past three decades (a full list of 46 papers published between 1990 and 2005 is
105 reported in [Harris 2013](#)). The methods, limits and applications of this approach are part of a book
106 expressly focussed on detecting, modelling and responding to effusive eruptions ([Harris et al., 2016b](#)),
107 whereas exhaustive overview of the physic behind mass and energy flow through a lava flow system
108 is provided by a series topical works ([Pieri and Baloga 1986](#), [Crisp and Baloga 1990](#), [Harris et al.,](#)
109 [1997](#), [Wright et al., 2001](#), [Harris et al., 2007](#), [Harris and Baloga 2009](#); [Dragoni and Tallarico 2009](#),
110 [Coppola et al, 2013](#); [Garel et al., 2012, 2014, 2015](#), [Tarquini \(2017\)](#), among others). Here, we outline
111 the basic principles of this technique and we summarize the open questions that have been addressed
112 in this work.

113 To describe volumetric flow rates of erupted lavas, we use the terminology given by [Harris et al.](#)
114 [\(2007\)](#). We use the generic term *effusion rate*, to describe the instantaneous rate at which lava is
115 erupted from the vent at any time. The term *mean output rate (MOR)* is used to describe the final
116 volume of erupted lava divided by the total duration of the eruption. Finally, we used the term *time-*
117 *averaged lava discharge rate (TADR)* to describe the volume of lava emplaced during a specific time
118 interval ([Harris et al., 2007](#)). This definition better applies to the use of satellite-based methods, for

119 measuring the changes in lava volume over a given period of time prior the image acquisition (Wright
120 et al., 2001).

121 Currently, two main methods exist to estimate *TADR*s from space-based thermal data; the method of
122 Harris et al. (1997), simplified later by Wright et al. (2001), and the method of Coppola et al., (2013).

123 Both the methods rely on the original heat balance approach (Pieri and Baloga, 1986), stating that the
124 *mean output rate MOR* ($\text{m}^3 \text{s}^{-1}$) of a cooling-limited lava flow, is related to its final plan area A
125 (attained when the flow achieves its final length, L), by:

126

$$127 \quad MOR = \frac{\varepsilon \sigma T_e^4}{\rho C_p (T_0 - T_{stop})} A(L) \quad (\text{eq. 1})$$

128

129 where ρ (kg m^{-3}) and C_p ($\text{J kg}^{-1} \text{K}^{-1}$) are the bulk density and heat specific capacity of lava,
130 respectively, T_e (K) is the effective radiation temperature of the flow, T_0 (K) is the lava eruption
131 temperature, and T_{stop} (K) is the temperature of the flow front at the time the flow has cooled to a halt.

132 In this framework, the effective radiation temperature is defined as “the temperature at which the flow
133 would radiate if it had a constant surface temperature throughout the emplacement of the flow” (Crisp
134 and Baloga, 1990).

135 Harris et al., (1997), applied this approach to satellite thermal data, developing what is actually called
136 the “thermal proxy”. Further works (Wright et al., 2001; Harris et al. 2007, 2009) refined and
137 simplified this method suggesting that the equation 1, when applied to satellite radiance data, provides
138 discharge rates that are not necessarily averaged over the total duration of an eruption, but rather over
139 “some time” prior to the satellite acquisition. Here, the effective radiation temperature, T_e , takes a
140 specific meaning, since it represents the average surface temperature over “some time” prior to the
141 satellite acquisition, weighted according to the radiative heat flux (Harris and Baloga, 2009).
142 Accordingly, Harris et al., (2007) proposed to use the term “*time-averaged discharge rate*” (*TADR*),
143 in order to describe the discrete measurements of lava flux retrieved from single satellite thermal

144 images. [Wright et al., \(2001\)](#) also argued that all values except *MOR* (or *TADR*) and *A* are assumed
145 a priori in equation 1 so that the thermal proxy reduces to a simple relationship whereby:

146

$$147 \quad TADR = Ax \quad (\text{eq. 2})$$

148

149 Here *x* is an empirical parameter (m s^{-1}), that describes the appropriate compositional flow parameters
150 (ρ , C_p) as well as thermal insulation (T_e) and cooling (T_0-T_{stop}) conditions (cf. [Harris and Baloga](#)
151 [2009](#)). In this formulation the area of the active flow area, *A* (m^2) is also dependent on the insulation
152 condition expressed by T_e , and is directly retrieved from the pixel integrated spectral radiance, R_λ (W
153 $\text{m}^{-2} \text{sr}^{-1} \mu\text{m}^{-1}$), according to:

154

$$155 \quad A = \sum_{i=1}^{npix} \frac{R_\lambda - L_\lambda(T_{bk})}{L_\lambda(T_e) - L_\lambda(T_{bk})} A_{pixel} \quad (\text{eq. 3})$$

156

157 where T_{bk} is the background temperature (K), L_λ is the Plank function for wavelength λ ($\text{W m}^{-2} \text{sr}^{-1}$
158 μm^{-1}), and A_{pixel} is pixel area (1 km^2 for MODIS). By assuming two end-member radiating
159 temperatures (T_e) for the hot and cold models (for example 500°C and 100°C for the channel- and
160 tube-fed flow type, respectively), the equation 3 allows the user to calculate a range of plausible active
161 flow areas, responsible for the observed radiance ([Harris et al., 2007](#)).

162 As stressed by [Harris and Baloga \(2009\)](#), the values for *x* have to be set on a case-by-case, thus leaving
163 a wide arbitrariness to the user that may adjust all the unknowns in equation (1) to achieve a best-fit
164 with available and independent field data. The TIR (thermal infrared) bands are generally used in this
165 approach under the assumption that the surface area of high temperature cracks is small and provides
166 a negligible contribution to the pixel integrated radiance at $10 \mu\text{m}$ to $12 \mu\text{m}$ ([Harris and Baloga 2009](#)).

167 This approach (equations 2 and 3) was the one used by [Bonny et al., \(2018\)](#) to estimate the volume
168 of lava erupted during the 2014-2015 eruption at Holuhraun by using MODIS TIR data.

169

170 The method of [Coppola et al., 2013](#), does not hold with the calculation of active flow areas, but is
171 based on the empirical relationship that directly links the *TADR* with the Volcanic Radiant Power
172 (*VRP*) calculated via the MIR method ([Wooster et al., 2003](#)). The MIR-derived *VRP* is a measurement
173 of the heat flux radiated almost exclusively by the portions of lava flows having effective temperature
174 (T_e) above 600 K, and is calculated as:

175

$$176 \quad VRP = 18.9 \cdot A_{\text{pixel}} \cdot \sum_{i=1}^{npix} (R_{MIR,alert} - R_{MIR,bk})_i \quad (\text{eq. 4})$$

177

178 where $R_{MIR,alert}$ is the pixel integrated MIR radiance of the i^{th} alerted pixel, $R_{MIR,bk}$ is the MIR radiance
179 of the background, A_{pixel} is the pixel size (1 km² for the resampled MODIS pixels), and 18.9 is a
180 constant of proportionality (see [Wooster et al., 2003](#)). The *TADR* is then calculated by using a single
181 coefficient called radiant density (c_{rad} in J m⁻³):

182

$$183 \quad TADR = \frac{VRP}{c_{\text{rad}}} \quad (\text{eq. 5})$$

184

185 that describes the relationship between volumetric and radiant flux appropriate for the observed
186 eruption ([Coppola et al., 2013](#)). By analyzing a large compositional spectrum of recent lava flows,
187 [Coppola et al., \(2013\)](#) proposed an empirical method to calculate the c_{rad} , based on the silica content
188 of the erupted lavas:

189

$$190 \quad c_{\text{rad}} = 6.45 \times 10^{25} \times (X_{\text{SiO}_2})^{-10.4} \quad (\text{eq. 6})$$

191 where X_{SiO_2} is the silica content of the erupted lavas (wt%).
192 This method allows to take into account the strong effects that the bulk rheology has on the spreading
193 and cooling processes of active lavas, and, according to the authors, permit to estimates $TADR$ with
194 an uncertainty of $\pm 50\%$ (Coppola et al., 2013). Most importantly, it allows to calculate the appropriate
195 radiant density, once the first chemical analysis of the erupted lavas become available (Coppola et
196 al., 2017). Since the MIROVA system provides automatically the VRP (Coppola et al., 2016), the
197 radiant density approach (equations 5 and 6) was the one used by Coppola et al., 2017 to estimate
198 $TADR$ and erupted lava volume during the Holuhraun eruption.

199
200 Although the two methods derive from the original heat balance approach, valid for cooling-limited
201 lava flows (eq. 1), it is important to remark that they do not necessarily yield exactly the same results
202 (cf. Bonny et al., 2018, Coppola et al., 2017). This is because the wavelengths used by the two
203 methods are different (R_{TIR} and R_{MIR} , respectively) but also because of the arbitrariness of the input
204 coefficients, as x for the Harris' method and c_{rad} for the Coppola's method. Nevertheless, both the
205 methods provide an upper and lower boundary estimates of $TADR$ that are *linearly* related to the
206 pixel-integrated spectral radiances (cf. eq. 1 to 5). It follows that whatever the method used, a constant
207 $TADR$ would be translated into a constant radiance detected by space.

208 This point was challenged by recent laboratory experiments and physical modelling of cooling
209 viscous gravity currents (Garel et al., 2012, 2014, 2015) according to which for a given magma
210 discharge rate the heat radiated by the flow surface reaches a steady value only after a transient time.
211 According to the cooling-limited flow model (eq. 1), this transient time would correspond to the time
212 required for the lava flow front to reach its maximum length (L) and to cool to a halt ($T_{front} = T_{stop}$).
213 Hence, during rapid changes of effusion rates the thermal signal would be "buffered" in time, because
214 it takes time for the higher flow rate to propagate downstream and cause perceptible increases in flow
215 area (Harris and Baloga 2009). The transient time thus reflect the dynamic response of the thermal
216 structure toward the appropriate cooling-limited conditions expressed by the new $TADR$ vs. A

217 relationship (i.e. new steady-state condition). [Tarquini \(2017\)](#) also suggests that this transient time is
218 typical of non-equilibrium steady-state systems, and is related to a structural relaxation time that
219 operate particularly on long-lived lava flows. In this non-equilibrium framework, the active lava flow
220 units represent “dissipative structures” that promote significant fluctuations in the radiative power,
221 even in the case of a constant supply. Hence, the radiant flux would be modulated by a pulsating
222 emplacement dynamic, associated to the occurrence of flow diversions that resets the system further
223 from its maximum extension (cf. [Tarquini 2017](#)).

224 Other open questions remain challenging, especially for operational use of the thermal proxy during
225 effusive crisis. For example, based on its theoretical model, [Garel et al. \(2012\)](#) suggests that when
226 lava tubes form, the surface thermal signal will not reflect the flow dynamics, because the low crust
227 temperatures do not reflect a potentially high flow rate of hot lava underneath. According to the
228 authors, this would pose a serious problem, since many lava flows exhibit evolving emplacement
229 styles throughout the eruption, often characterized by increasing insulation conditions and gradual
230 formation of lava tubes. Similarly, it is still unclear how to determine whether an eruption has ceased
231 based on thermal data, or whether the lava stored underneath a cooled crust can potentially emerge
232 suddenly and spread rapidly ([Garel et al., 2015](#)). In these terms the 2014-2015 eruption at Holuhraun
233 represents a key-case since allowed us to address all these points over the largest effusive eruptions
234 occurred in the past three centuries.

235

236 **3. Methods: Analysis of MODIS-MIR data**

237 ***3.1 The MIROVA system***

238 MIROVA (Middle Infrared Observation of Volcanic Activity) is an automated global hot spot
239 detection system (www.mirovaweb.it) based on near-real time processing of Moderate Resolution
240 Imaging Spectroradiometer (MODIS) infrared data ([Coppola et al. 2016](#)). The system completes
241 automatic detection and location of thermal anomalies, and provides a quantification of the Volcanic

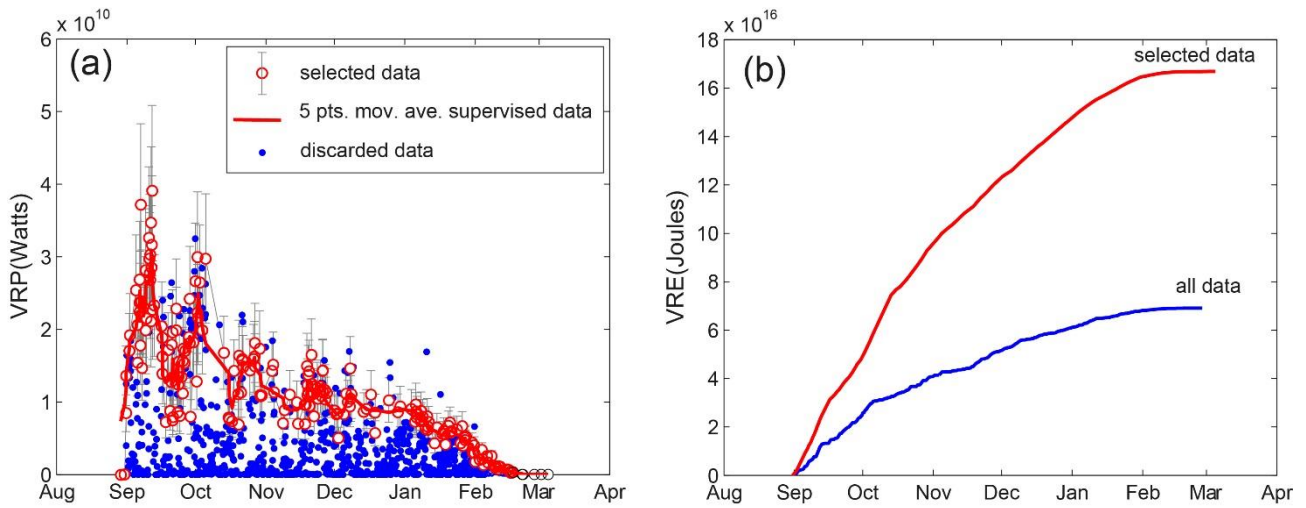
242 Radiative Power (*VRP*), within 1 to 4 hours of each satellite overpass. This is achieved through a
243 hybrid algorithm (fully described in Coppola et al., 2016) based on MIR radiance data ($\sim 4 \mu\text{m}$)
244 recorded at 1-km spatial resolution, as shown in Fig. 1b. Two MODIS instruments (carried on two
245 NASA's satellites: Terra and Aqua) deliver approximately 4 images per day for every target volcano
246 located at equatorial latitudes. However, due to the polar, sun-synchronous orbit, the two satellites
247 increase their sampling time at high latitudes, providing, at least 6 to 10 overpasses per day over
248 Iceland. An example of thermal images elaborated by the MIROVA system during the Holuhraun
249 eruption is given in Fig. 1b.

250

251 ***3.2 Volcanic radiative power and energy of the 2014-2015 eruption at Holuhraun***

252 Between August 29th, 2014 and March 4th, 2015, MIROVA detected hotspots in 1105 images over a
253 total of 1623 MODIS overpasses above Iceland ($\sim 68.1\%$). The volcanic radiative power (*VRP*) ranged
254 from ~ 39.1 GW, during the initial stage of the effusion, to less than 10 MW just before the end of the
255 eruption (Figure 2a). During the course of the eruption, we visually inspected all the acquired scenes
256 and we identified a large number of images acquired in cloudy conditions, and/or under poor
257 geometrical conditions that strongly deformed and affected the thermal anomaly at ground level.
258 These images were discarded from further analysis (blue circles in Fig. 2a) so that a supervised dataset
259 of only 206 high-quality images ($\sim 12.7\%$ of the total MODIS overpasses) was used to provide a
260 robust quantification of *VRP* produced by the eruption (red circles in Fig. 2a). As a whole we
261 estimated that the Holuhraun eruption radiated approximately 1.6×10^{17} J into the atmosphere (red
262 curve in Fig. 2b), with a time-averaged radiant flux of about 10.5 GW.

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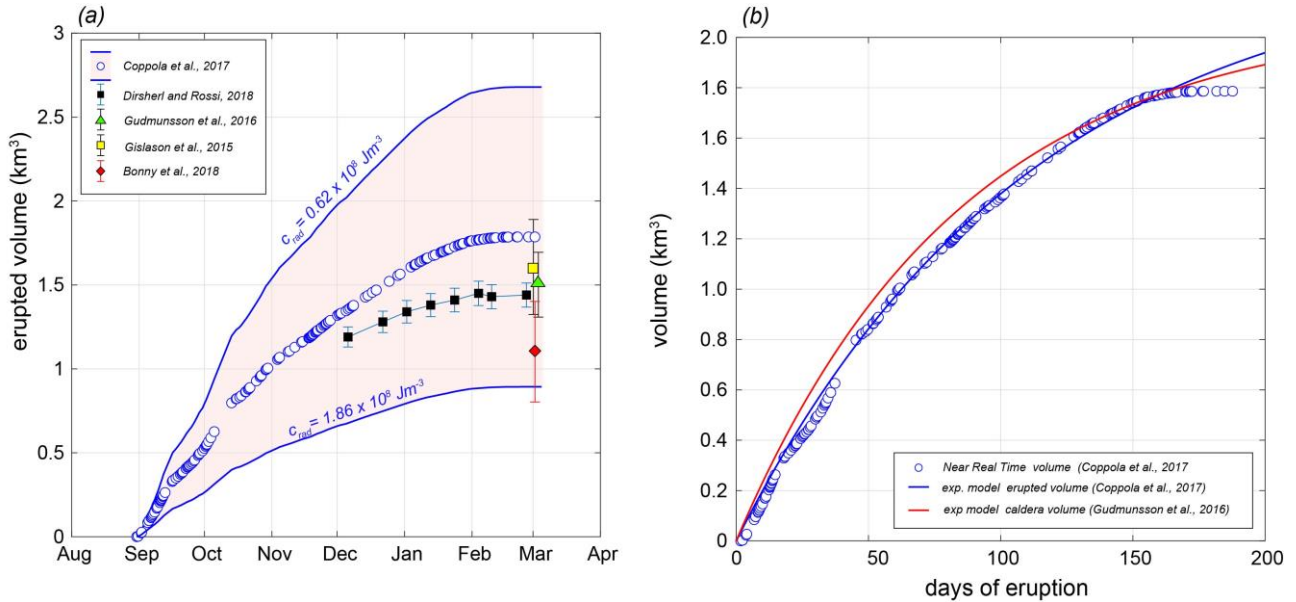
265 **Figure 2.** (a) Timeseries of *VRP* retrieved from MODIS-MIROVA system during the 2014-2015 eruption at Holuhraun.
 266 Blue circles represent data discarded because of cloudy conditions or poor viewing geometry. The data selected for *VPR*
 267 calculation are the white circles. The red line is a 5-points-moving average of *VRP*. A *VRP* below 10 MW coincides with
 268 the end of the eruption. (b) Volcanic Radiative Energy (*VRE*) calculated by trapezoidal integration of selected data only
 269 (red curve) or all the data (blue curve). Note the strong influence of data selection on the final total energy. Approximately
 270 1.6×10^{17} J were radiated into the atmosphere (red curve), with a mean radiant flux of about 10.5 GW.

271

272 **3.3 Time-averaged discharge rates and volume calculation**

273 During the course of the eruption, we calculated *TADRs* and erupted lava volumes by using the radiant
 274 density approach described above (eq. 5 and 6). The preliminary chemical analysis of Holuhraun
 275 lavas ($\text{SiO}_2 = 50.5$ wt%), provided in early September 2014 by the University of Iceland
 276 (http://earthice.hi.is/bardarbunga_2014), were used to calculate (eq. 6) a radiant density comprised
 277 between 0.62×10^8 and 1.86×10^8 J m^{-3} . These two values allowed us to provide an upper and lower
 278 boundary limits for lava volume calculation, with the mean value being the most likely (Fig. 3a).
 279 Given a final *VRE* equal to 1.64×10^{17} J (Fig. 2b), this method provided a first, timely assessment of
 280 the erupted lava volume equal to 1.75 ± 0.88 km^3 (Fig. 3a). This was in good agreement with the flow
 281 volume estimated a-posteriori ($1.4 - 1.6$ km^3) by using independent datasets (Gislason et al. 2015,
 282 Gudmundsson et al., 2016, Dirshel and Rossi, 2018; Bonny et al., 2018). The inferred satellite-

283 derived *TADR* and erupted volumes were updated continuously and delivered via emails to the teams
 284 of the IMO and University of Iceland in charge of the monitoring.
 285



286
 287 **Figure 3.** (a) Lava volumes measured during the 2014-2015 eruption at Holuhraun. The colored field envelop the upper
 288 and lower boundary limits calculated using the radiant density approach (eqs. 5 and 6). Best-estimate volume calculation
 289 (blue circles), and relative uncertainty ($\pm 50\%$) were provided in near real time to IMO and University of Iceland.
 290 Independent measurements of final lava flow volumes given by several authors are also shown. (b) Comparison between
 291 lava field volume measured by MODIS (blue circles) and exponential models of erupted lava volume (Coppola et al.,
 292 2017) and caldera volume (Gudmunsson et al., 2016). The good correlation indicates a strong link between to the gravity-
 293 driven collapse of Bárðarbunga caldera and the gradual decrease of magma-static pressure driving the effusive eruption
 294 at the vent (cf. Gudmunsson et al., 2016, Coppola et al., 2017, Dirsheri and Rossi., 2018).

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297 **4. Eruptive trend derived from satellite thermal data**

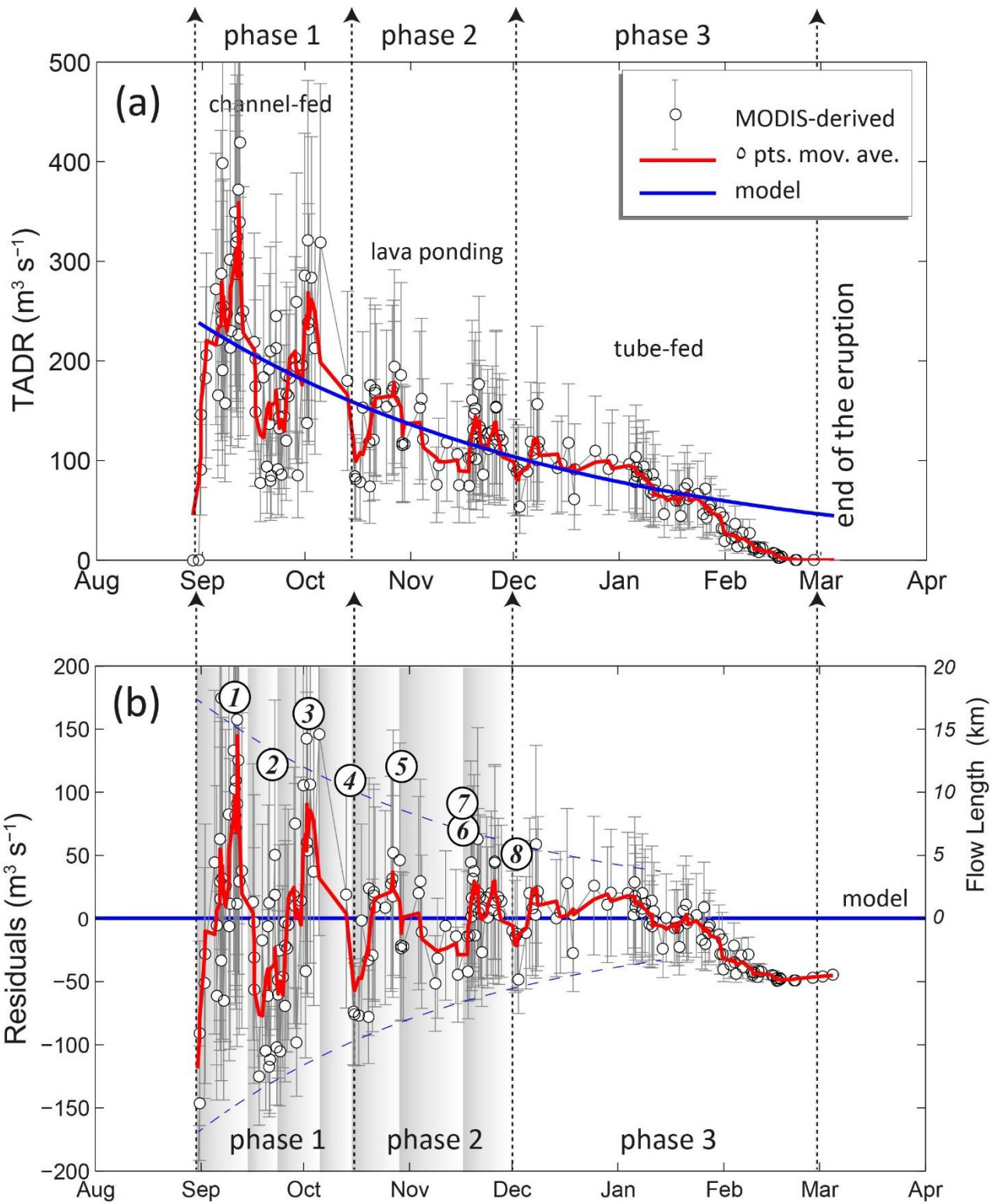
298 One of the most significant feature of this eruption was the overall exponential trend of the satellite-
 299 derived cumulative volume (Fig. 3b). As outlined by recent studies, this trend mirrors almost perfectly
 300 the slow subsidence of the Bárðarbunga caldera floor (Fig. 3b) and suggests a strong connection
 301 between source caldera/reservoir dynamic, and the eruptive dynamics 45 km distant ([Gudmundsson](#)

302 [et al., 2016](#); [Coppola et al., 2017](#)). According to these works, the exponential decay of effusion rates
303 was driven by a decrease of magma-static pressure, associated to the shrinking of the magma chamber
304 and the gravity-driven collapse of Bárðarbunga caldera. In this scenario, the instantaneous effusion
305 rate at the vent follows a perfectly exponential curve (blue line in Fig 4a) that has been used as a
306 benchmark for our MODIS-derived *TADRs*' (red line in Fig 4a). Hence, the residuals between the
307 two rates (observed - modeled; Fig. 4b) enhance temporary variations of the source process, or
308 instabilities of the thermal proxy due to lava flow emplacement dynamic. In the following section,
309 we discuss the effusive trend in the light of the evolving emplacement style observed in the Holuhraun
310 lava field.

311 According to [Pedersen et al. \(2017\)](#) the Holuhraun eruption can be subdivided in three main phases,
312 characterized by evolving lava transport processes.

313 The phase 1 (*31 August – 12 October 2014*) was characterized by lava transport confined to *open*
314 *channels* that led the formation of four consecutive lava flows emplaced side by side (no. 1–4 in Fig
315 4b). During this phase the discharge rates measured on the field were comprised between 350 and
316 $100 \text{ m}^3 \text{ s}^{-1}$ ([Pedersen et al., 2017](#)), in good agreement with satellite-derived values. The four flow
317 units reached distances of 16.9, 11.8, 16.3 and 11.0 km, respectively (Fig. 4a), and covered a total
318 area of 58.31 km^2 ([Pedersen et al., 2017](#)). Given a flow field volume of 0.78 km^3 estimated on October
319 12 (Fig. 3b), the mean flow thickness characterizing the Phase 1 was 13.4 m, consistent with the
320 thickness of each single flow units observed on the field ([Pedersen et al., 2017](#)). During this period
321 our *TADRs* measurements show the greatest oscillations around the modelled value (up to $\pm 60\%$),
322 resulting in sharp variations exactly in correspondence of the activations of the distinct flow units
323 (Fig 4b).

324



325

326 **Figure 4.** (a) *TADR* derived from MIROVA data (white circles) by using the radiant density approach (eqs. 5 and 6). The
 327 blue line represents the best-fit exponential model (Coppola et al., 2017) hereby considered the real effusion rate at the
 328 vent. (b) Residuals flow rates (left axis) obtained by subtracting the model from the observations (i.e. MODIS-derived
 329 *TADR* minus modelled effusion rates). Note how the residuals oscillate around the model showing a decreasing amplitude
 330 through time. The timing and amplitude of the residual pulses are coherent with the occurrence and maximum length of

331 the main channel-fed flow units numbered from 1 to 8 (data from Pedersen et al., 2017). The three phase of the eruption
332 described by Pedersen et al., 2017 are also shown.

333

334 The phase 2 (13 October – 30 November 2014) had lower effusion rates ($150 - 100 \text{ m}^3 \text{ s}^{-1}$) and was
335 characterized by the formation of a $< 1 \text{ km}^2$ lava pond, that acted as a distributor of subsequent lava
336 flows (i.e. flows no. 5–8; Pedersen et al., 2017). These flow units reached gradually shorter distances
337 (11.7, 8.4, 7.1, 5.4 km, respectively) and were accompanied by satellite-derived TADR's showing
338 decreasing oscillation amplitudes (Fig. 4b). Towards the end of this phase the occurrence of *inflation*
339 *plateaus* provided the first evidences of a growing lava tube system, with satellite-derived TADR's
340 becoming more stable (Fig. 5b). About 0.52 km^3 , of lava was erupted during this phase (Fig. 3a)
341 producing a net increment of the flow field area of 18.7 km^2 (Pedersen et al., 2017). By the end of
342 November the flow field reached an average thickness of 16.9 m.

343 During the third phase (1 December 2014 – 27 February 2015) the lava transport was mainly confined
344 to *lava tubes* which fed several breakouts and inflation plateaus along the initial flow units no. 1 and
345 2 (Pedersen et al., 2017). More than 19 km^2 of the flow field was resurfaced, with discharge rates
346 much more steady ($\pm 20\%$) and typically lower than $100 \text{ m}^3 \text{ s}^{-1}$. A sharp reduction of TADR was
347 recorded since 27 January 2015 and precluded the end of the eruption occurred one month later, on the
348 27 February 2015 (Fig.4). This reduction was interpreted by Coppola et al., (2017) as due to the
349 gradual closure of the magma path once the overpressure inside the dike had dropped below a critical
350 value. The lava tube system distributed about 0.4 km^3 during this phase, with the lava field reaching
351 a final extension of 85.4 km^2 . The inflation and resurfacing processes that operated through the tube
352 system led the average flow thickness to increase up to 20.2 m.

353

354 The emplacement of the major lava flows (phases 1 and 2) is outlined by the occurrence of apparent
355 TADR's "pulses" (Fig. 4a) oscillating around the modelled value for effusion rate (Fig. 4b). Pulsating
356 effusive activity has been recognized at several basaltic volcanoes and may results from different

357 processes such as pulsed magma supply, repeated accumulation and collapse of a foam layer at the
358 reservoir roof, or by processes of magma mixing occurring within the magma chamber (e.g., [Harris](#)
359 [and Neri 2002](#), [Lautze et al., 2004](#)). However, in the case of Holuhraun eruption, the link between the
360 effusive trend and the collapse of the caldera (Fig. 3b) claims for a simple, gravity-driven dynamic
361 ([Gudmundsson et al., 2016](#); [Coppola et al., 2017](#)) that would exclude, or minimize, the occurrence of
362 above mentioned processes. Hence, the *TADR*'s pattern overprinted to the exponential trend is likely
363 related to the emplacement dynamic occurred during the course of the eruption, and especially during
364 the phases 1 and 2. Laboratory experiments ([Garel et al., 2012, 2014](#)) and theoretical treatments
365 ([Tarquini 2017](#)), demonstrated that a constant effusion rate can actually produce pulses of the
366 radiative power (Fig. 4b) that reflect the non-equilibrium steady-state growth of compound lava flow
367 fields. According to [Tarquini \(2017\)](#) each flow unit will advance until the thermal equilibrium will
368 be reached and the flow front stop. Hence, the system is resettled to give origin to a new flow unit.
369 This seems to be exactly the dynamic occurred during the emplacement of the Holuhraun lava field,
370 in which each apparent *TADR*'s pulse correlates with a phase of lengthening of a distinct flow unit.
371 Notably, the amplitude of the pulses decreased with time as the overall effusion rate declined, the
372 maximum length attained by the flow units reduced, and the emplacement style evolved from
373 channel- to tube-fed (Fig. 4b).

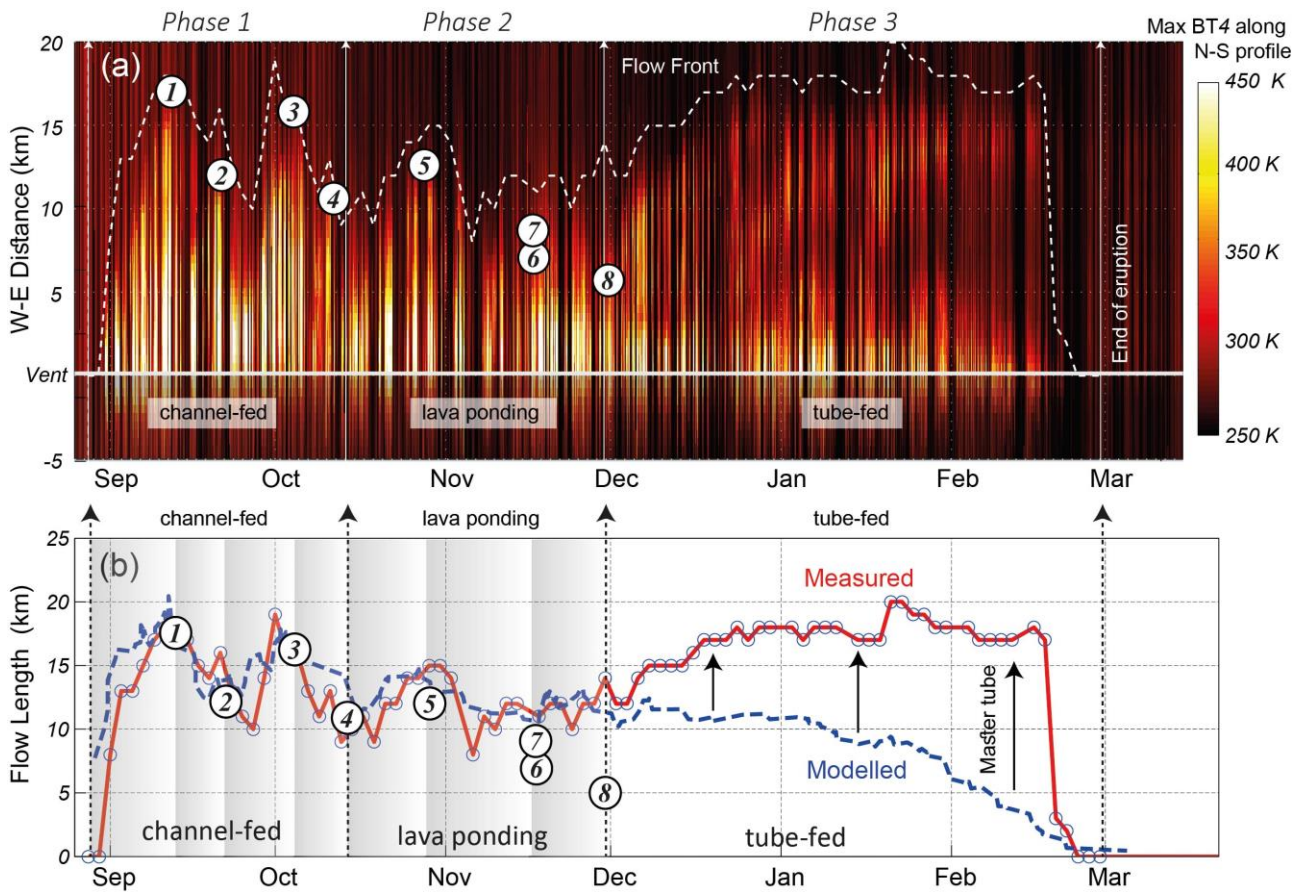
374

375 **5. Evolving emplacement style, evolving thermal structure**

376 In this section, we illustrate how the evolving emplacement style and the channel- to tube-fed
377 transition affected the surface thermal structure of the Holuhraun's lava field. In particular, we used
378 the temperature profiles (as the one shown in Fig. 1c) obtained from the MODIS image, in order to
379 track the position of the active flow fronts and all the hot lava surfaces emerging along the principal
380 flow path (from West to East). By stacking together all the W-E profiles the presence of active lavas
381 (i.e. the vent(s), open-channel(s), breakout(s), flow front(s), etc.) appears as high-temperature features

382 (pixels), whose intensity, contiguity and extension can be easily visualized as shown in Fig. 5a. We
383 assumed that the position of the flow front is represented by the easternmost pixel having a brightness
384 temperature (BT4) of 50°C higher than the background (white dashed line in Fig. 5a). Hence, the
385 distance between this pixel and the one located over the active vent (Fig. 1b), provides a measurement
386 of the active flow length at each satellite overpass (red line in Fig. 5b). Lava flow lengths measured
387 on the field during the phases 1 and 2 (Flows no. 1 to 8) are also plotted for comparison (Fig. 5b).

388 This analysis indicates that during the phases 1 and 2, the active flow fronts were thermally connected
389 to the vents, through a high-radiating feature representing the open channels. On the other hand, since
390 December 2014 the reduction of lava discharge rate caused a marked shortening of the open channels,
391 although the lava field was still characterized by the presence of active lavas at the distal front(s).
392 Notably, between the two hot, active zones (i.e. open channels and flow front) there low-temperature
393 zone corresponding to crusted and cooling lava surfaces (Fig. 5a). This is a clear thermal evidence
394 that since the beginning of the phase 3 the transport of lava occurred along a master tube system and
395 that the emplacement style started to be *tube-fed* dominated.



396

397 **Figure 5.** (a) Evolution of the along-path thermal structure of the Holuhraun lava flow. This plot is obtained by stacking
 398 together all the thermal profiles (max BT4 along each N-S line) as those illustrated in Fig. 1c. The white dashed line
 399 indicate the position of the active flow front (see the text for details); (b) Comparison between measured (red dashed line)
 400 and modelled (blue line) flow length. The Calvari & Pinkerton's model (eq. 7) assumes emplacement style dominated by
 401 *channel-fed*, cooling limited conditions. The discrepancy between measured and modelled flow distance after December
 402 2014 outline the increasing efficiency of *tube-fed* emplacement style in transporting the lava to the flow front.

403

404 As a pure theoretical exercise, the flow extension measured by using the temperature profiles has
 405 been compared (Fig. 5b) with the maximum distance (L_{max}) predicted by one of the simplest empirical
 406 model developed for single, cooling-limited, channel-fed lava flow (cf. Walker 1973; Harris and
 407 Rowland, 2009). In particular, we used the Calvari and Pinkerton's equation (1998) that takes into
 408 account exclusively the *TADR*s measurements for the estimation of flow length:

409

410

$$L_{max} = 10^{3.11} \times (TADR)^{0.47} \quad (\text{eq.7})$$

411 The coefficients of equation 7 were empirically determined by regression analysis of single channel-
412 fed Etnean lava flows lasting 1 to 12 days, a duration very similar to the single flow units observed
413 at Holuhraun. This model does not require any knowledge of several others parameters (i.e.
414 underlying slope, lava temperature, viscosity, velocity, channel width, channel depth, etc...) that are
415 necessary to predict flow length using other approaches (see [Harris and Rowland 2009](#) for a review),
416 but that are difficult, if not impossible to be constrained for each flow unit. However, despite its
417 simplicity, this empirical model of Calvari and Pinkerton 1998 allows to calculate the maximum
418 length of the Holuhraun's lava flow, and to see the effects of an emplacement style persistently
419 dominated by channel-fed, cooling-limited conditions (blue line in Fig. 5b).

420 During the phases 1 and 2, there is an excellent correlation between the measured and the modelled
421 flow lengths, both in terms of absolute values and pattern (Fig. 5b). This corroborate the fact that
422 during this period the emplacement style was effectively those for which eq.7 was proposed to
423 describe, that is single, channel-fed flows. Differently, since December 2014, there is a growing
424 discrepancy between the measured and modelled flow length (Fig. 5b), exactly in correspondence of
425 the formation of a master lava tube, as observed in the field ([Pedersen et al., 2017](#)).

426 From this picture, it is clear that the lower discharge rates characterizing the Phase 3, did not affect
427 the maximum extent of the lava flow, still reaching a distance of more than 15 km, just before the end
428 of the eruption (Fig. 5b). By paraphrasing [Harris and Rowland, \(2009\)](#) “the transition from poorly
429 insulated (*channel-fed*) to well insulated (*tube-fed*) regimes increases the length that a lava flow can
430 extend for a given *TADR*”.

431 Notably, we may calculate the radiant density ($c_{rad} = VRE/Vol$) typical of the tube-fed flow field, by
432 dividing the thermal energy measured during the Phase 3 ($VRE = \sim 0.4 \times 10^{17}$ J; Fig. 2b) by the
433 volume lava erupted in the same time window, derived from Tandem-X measurements ($Vol = 2.6 \times$
434 10^8 m³; Fig. 3a; [Dirscherl and Rossi 2018](#)). The resulting radiant density ($c_{rad} = \sim 1.5 \times 10^8$ J m⁻³) is
435 only slightly higher than the radiant density calculated in the same way for the bulk channel-fed flow
436 field emplaced during phases 1 and 2 ($c_{rad} = \sim 1.1 \times 10^8$ J m⁻³). This suggests that the variation of the

437 emplacement style produces an uncertainty in the *TADR*s measurement that does not exceed 30%,
438 when provided in near real time. In other words, our results point out that, the formation of a lava
439 tube does not affect substantially the amount of “active” lava radiating to the atmosphere, but simply
440 “made this lava to be transported further from the vent”. The lava tube thus act as a simple extension
441 of the volcanic conduit, by displacing the vent downstream, likely along the main axis of the
442 previously formed flow units. The recognition of this behavior in near real time allow to update the
443 vent location and has been proved to be a necessary improvement to correct estimation of the lava
444 flow path and the effective runout distance (Harris et al., 2018, this issue).

445

446 ***6. Operational use of MIROVA system during the effusive crisis***

447 By the end of November 2014, MIROVA started to be part of the monitoring system operative at the
448 Icelandic Meteorological Office (IMO) with the coverage of nine active volcanic systems. IMO is the
449 main institution in Iceland in charge of monitoring natural hazards and issuing forecast and warnings
450 in case of impending or on-going eruptions. It operates a multi-sensors network that includes
451 seismometers, GPS and gas sensors for the geophysical monitoring. During the Holuhraun eruption,
452 gas monitoring activity was improved and new instrumentations were deployed and maintained at the
453 eruption site for the entire duration of the event (Pfeffer et al. 2017). The MIROVA system showed
454 its strength and reliability in a period of the year when, in Iceland, the day light time is very short.
455 Several of the instrumentations and products that could normally be used to monitor an eruption,
456 started to fail or to give incomplete data sets due to harsh weather conditions and limited sunlight.
457 MIROVA provided a robust and continuous data set on a daily basis which was used from November
458 to March to follow the health and strength of the eruption, as well as the post-eruption phase.
459 On the operative side this system is revealed to be an important monitoring tool for following the
460 temporal evolution of the eruption and providing estimates of the erupted lava volume in near real
461 time (Coppola et al., 2017). Extrapolation of the effusive trend during the ongoing eruption may in

462 fact be used to forecast total erupted lava volume, which in turn is fundamental to drive some lava
463 flow modelling (i.e. [Tarquini et al., 2018](#)). Also, the recognition of emplacement style through the
464 analysis of flow' surface thermal structure (i.e. channel- or tube-fed lava flows) can be used to update
465 the position of active flow fronts, and to choose the appropriate modeling strategy, by tuning some
466 appropriate parameters that govern lava flow simulation's codes (i.e. [Tarquini et al., 2018, this issue](#)).
467 Most importantly the thermal data have proved to be very important for constraining the time of the
468 eruption end. A drastic change of *TADR* was observed since the end of January 2015 becoming clear
469 the second week of February 2015 (Fig. 4a). This change was interpreted by [Coppola et al. \(2017\)](#) as
470 due to the gradual closure of the magma path (dike) and indicated the potential timing of the end of
471 the eruption in quite a good advance, about one month before the declaration of ceased eruption.
472 These data were discussed within the Scientific Board Management meetings that were held from the
473 beginning of the eruption in collaboration with the Icelandic Civil Protection and the University of
474 Iceland. The very low *TADR* detected on 26 February 2017 ($0.2 \text{ m}^3 \text{ s}^{-1}$; Fig. 4a), supported the
475 decision for a field crew to flight over the eruptive site on February 27 and declare that the eruption
476 was over.

477 The thermal data were useful for providing a semi-quantitative indication of the strength of the gas
478 source affecting the ground level concentration ([Simmons et al., 2017](#)). Due to its extension the lava
479 field itself was a considerable source of volcanic gases. In addition, the gas released close to the
480 surface was more easily trapped within the boundary layer and largely affected the concentration at
481 ground. This fact suggested that the MIROVA detection and the *VRP* temporal variation could be
482 correlated with the amount of gases released at the source. In this way the *VRP* data have been used
483 by the forecasters at IMO to identify, in a qualitative way, when to expect an increase in the SO_2
484 values in areas located downwind the eruption site.

485

486

487 7. Conclusions

488 During the 2014-2015 Bárðarbunga-Holuhraun eruptive crisis, one of the main challenge for the
489 volcanological community was to monitor the effusive process in a safely, timely and routinely
490 manner. For the whole duration of the eruption, MIROVA provided a robust and continuous set of
491 thermal data that were used to depict the effusive trend and to estimates the erupted lava volume
492 although the changing emplacement conditions (from channel- to tube-fed flow) introduced
493 uncertainties in the interpretation of the data.

494 The retrospective analysis of this eruption provides an exceptional opportunity to study the
495 relationship between heat flux and *TADR* during the emplacement of a large compound lava field.
496 We show that the overall trend of thermal emission was related to the combined effect of two over
497 imposed patterns: (i) a main exponential decay of the effusion rate trend, governed by the source
498 processes and (ii) a secondary pulsating pattern related to the emplacement dynamic of the flow field.
499 We found that the magnitude and timing of these pulses were strictly related to the timing and length
500 scales of discrete flow units characterizing open-channel lava transport. Conversely, the formation of
501 lava tubes produces smaller instabilities and promotes lava to flow at greater distance in steady-state
502 thermal conditions. This process is clearly visible from the evolution of the thermal structure depicted
503 from the MODIS-MIROVA images and allow to track changes in the *lava transport mechanisms*
504 operating during the eruption (Fig. 5a). The results presented here clearly indicate that the calculation
505 of erupted lava volumes from the integration of satellite-derived *TADR* allow to smooth the short-
506 term perturbation associated to the emplacement dynamic and provide a robust way to depict source
507 eruptive trends. We thus regard the application of this methodology as a key-factor in volcano
508 monitoring and satellite-data-driven response to an effusive crisis (Harris et al., 2018).

509

510

511

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516

517 **References**

518 Bonny, E., T. Thordarson, R. Wright, A. Höskuldsson, I. Jónsdóttir (2018) The volume of lava
519 erupted during the 2014 to 2015 eruption at Holuhraun, Iceland: A comparison between satellite
520 and ground-based measurements, *J Geophys Res Solid Earth*, 123.
521 <https://doi.org/10.1029/2017JB015008>.

522 Calvari, S. and H. Pinkerton (1998) Formation of lava tubes and extensive flow field during the 1991–
523 1993 eruption of Mount Etna, *J. Geophys Res*, B103, 27291–27301.

524 Coppola, D., M. Laiolo, D. Piscopo and C. Cigolini (2013) Rheological control on the radiant density
525 of active lava flows and domes, *J. Volcanol. Geotherm. Res.*, 249, 39–48.
526 doi:10.1016/j.jvolgeores.2012.09.005.C.

527 Coppola, D., M. Laiolo, C. Cigolini, D. Delle Donne and M. Ripepe (2016) Enhanced volcanic hot-
528 spot detection using MODIS IR data: Results from the MIROVA system, in Harris A. J. L. , T.
529 De Groot, F. Garel and S. A. Carn (eds.) *Detecting, Modelling, and Responding to Effusive*
530 *Eruptions*, Geological Society, London, Sp. Pub. 426, p. 181–205, doi:10.1144/SP426.5.

531 Coppola, D., M. Ripepe, M. Laiolo and C. Cigolini (2017) Modelling Satellite-derived magma
532 discharge to explain caldera collapse, *Geology* 45(6), 523-526. doi:10.1130/G38866.1.

533 Crisp, J. and S. Baloga (1990) A method for estimating eruption rates of planetary lava flows, *Icarus*,
534 85(2), 512-515.

535 Crisp, J. and S. Baloga (1990) A model for lava flows with two thermal components, *J. Geophys.*
536 *Res.*, 95 (B2), 1255-1270.

537 Dirsherl, M. and C. Rossi (2018) Geomorphometric analysis of the 2014–2015 Bárðarbunga volcanic

538 eruption, Iceland, *Remote Sens. Environ.*, 204, 244-259.

539 Dragoni, M. and A. Tallarico (2009), Assumptions in the evaluation of lava effusion rates from heat
540 radiation, *Geophys. Res. Lett.*, 36, L08302. doi:10.1029/2009GL037411.

541 Garel, F., E. Kaminski, S. Tait and A. Limare (2012) An experimental study of the surface thermal
542 signature of hot subaerial isoviscous gravity currents: implications for thermal monitoring of
543 lava flows and domes, *J. Geophys. Res.* 117:B02205. doi: 10.1029/2011JB008698.

544 Garel, F., E. Kaminski, S. Tait and A. Limare (2014) An analogue study of the influence of
545 solidification on the advance and surface thermal signature of lava flows, *Earth Planet. Sci.*
546 *Lett.*, 396, 46–55.

547 Garel, F., E. Kaminski, S. Tait and A. Limare (2015) A fluid dynamics perspective on the
548 interpretation of the surface thermal signal of lava flows, , in Harris A. J. L. , T. De Groeve, F.
549 Garel and S. A. Carn (eds.) *Detecting, Modelling, and Responding to Effusive Eruptions*,
550 Geological Society, London, Sp. Pub. 426, 243-256. doi: 10.1144/SP426.6

551 Gíslason, S. R., G. Stefánsdóttir, M. A. Pfeffer, S. Barsotti, Th. Jóhannsson, I. Galeczka, E. Bali, O.
552 Sigmarsson, A. Stefánsson, N. S. Keller, A. Sigurdsson, B. Bergsson, B. Galle, V. C. Jacobo,
553 S. Arellano, A. Aiuppa, E. B. Jónasdóttir, E. S. Eiríksdóttir, S. Jakobsson, G. H. Guðfinnsson,
554 S. A. Halldórsson, H. Gunnarsson, B. Haddadi, I. Jónsdóttir, Th. Thordarson, M. Riishuus, Th.
555 Högnadóttir, T. Dürig, G. B. M. Pedersen, A. Höskuldsson and M. T. Gudmundsson (2015)
556 Environmental pressure from the 2014–15 eruption of Bárðarbunga volcano, Iceland,
557 *Geochem. Persp. Lett.*, 1, 84-93.

558 Gudmundsson, A., N. Lecoeur, N. Mohajeri and T. Thordarson (2014) Dike emplacement at
559 Bárðarbunga, Iceland, induces unusual stress changes, caldera deformation, and earthquakes,
560 *Bull. Volcanol.*, 76, 869–875. doi:10.1007/s00445-014-0869-8.

561 Gudmundsson, M. T., K. Jónsdóttir, A. Hooper, E. P. Holohan, S. A. Halldórsson, B. G. Ófeigsson,
562 S. Cesca, K. S. Vogfjörð, F. Sigmundsson, T. Högnadóttir, P. Einarsson, O. Sigmarsson, A. H.
563 Jarosch, K. Jónasson, E. Magnússon, S. Hreinsdóttir, M. Bagnardi, M. M. Parks, V.

564 Hjörleifsdóttir, F. Pálsson, T. R. Walter, M. P. Schöpfer, S. Heimann, H. Reynolds, S. Dumont,
565 E. Bali, G. H. Gudfinnsson, T. Dahm, M. J. Roberts, M. Hensch, J. M. Belart, K. Spaans, S.
566 Jakobsson, G. B. Gudmundsson, H. M. Fridriksdóttir, V. Drouin, T. Dürig, G. Aðalgeirsdóttir,
567 M. S. Riishuus, G. B. Pedersen, T. van Boeckel, B. Oddsson, M. A. Pfeffer, S. Barsotti, B.
568 Bergsson, A. Donovan, M. R. Burton, A. Aiuppa (2016) Gradual caldera collapse at
569 Bárðarbunga volcano, Iceland, regulated by lateral magma outflow, *Science*, 353.
570 doi:10.1126/science.aaf8988.

571 Harris, A. J. L. and M. Neri (2002) Volumetric observations during paroxysmal eruptions at Mount
572 Etna: Pressurized drainage of a shallow chamber or pulsed supply?, *J. Volcanol. Geotherm.*
573 *Res.*, 116, 79–95, doi:10.1016/S0377-0273(02)00212-3.

574 Harris, A. J. L. and S. Baloga (2009) Lava discharge rates from satellite-measured heat flux, *Geophys.*
575 *Res. Lett.*, 36, L19302. doi:10.1029/2009GL039717.

576 Harris, A. J. L. and S. K. Rowland (2009) Effusion rate controls on lava flow length and the role of
577 heat loss; a review, in Thordarson, T., S. Self, G. Larsen, S. K. Rowland and A. Hoskuldsson
578 (eds) *Studies in Volcanology: The Legacy of George Walker*. Geological Society, London Sp.
579 *Pub. of IAVCEI*, 2, 33– 51.

580 Harris, A. J. L., S. Blake, D. A. Rothery, and N. F. Stevens (1997) A chronology of the 1991 to 1993
581 Etna eruption using AVHRR data: Implications for real time thermal volcano monitoring, *J.*
582 *Geophys. Res.*, 102(B4), 7985–8003, doi:10.1029/96JB03388.

583 Harris, A. J. L., J. Dehn, and S. Calvari (2007) Lava effusion rate definition and measurement: A
584 review, *Bull. Volcanol.*, 70, 1–22. doi:10.1007/s00445-007-0120-y.

585 Harris, A. J. L., S. Carn, J. Dehn, C. Del Negro, G. Guðmundsson, B. Cordonnier, T. Barnie, S.
586 Calvari, T. Catry, T. de Groeve, D. Coppola, A. Davies, M. Favalli, E. Fujita, G. Ganci, F.
587 Garel, J. Kauahikaua, K. Kelfoun, V. Lombardo, G. Macedonio, J. Pacheco, M. Patrick, N.
588 Pergola, M. Ramsey, R. Rongo, K. Smith, S. Tarquini, T. Thordarson, N. Villeneuve, P.
589 Webley, R. Wright and K. Zakzek (2016a) Conclusion: Recommendations and findings of the

590 RED SEED working group, in Harris A. J. L., T. De Groeve, F. Garel and S. A. Carn (eds.)
591 Detecting, Modelling, and Responding to Effusive Eruptions, Geological Society, London, Sp.
592 Pub. 426, p. 567-648. doi:10.1144/SP426.11.

593 Harris, A. J. L., T. De Groeve, S. Carn, F. Garel (2016b) Risk evaluation, detection and simulation
594 during effusive eruption disasters, in Harris A. J. L., T. De Groeve, F. Garel and S. A. Carn
595 (eds.) Detecting, Modelling, and Responding to Effusive Eruptions, Geological Society,
596 London, Sp. Pub. 426, p. 1-22. doi:10.1144/SP426.29.

597 Harris, A. J. L., O.M. Chevrel, D. Coppola, M.S. Ramsey, A. Hrysiewicz, S. Thivet, N. Villeneuve,
598 M. Favalli, A. Peltier, P. Kowalski, A Di Muro, J-L Froger, L. Gurioli (2018) Validation of
599 an integrated satellite-data-driven response to an effusive crisis: 2 the April–May 2018 eruption
600 of Piton de la Fournaise, Special Issue: MeMoVolc, Annals of Geophysics, (in press, this issue)

601 Lautze, N. C., A. J. L. Harris, J. E. Bailey, M. Ripepe, S. Calvari, J. Dehn, S. Rowland, and K. Evans-
602 Jones (2004) Pulsed lava effusion at Mount Etna during 2001, J. Volcanol. Geotherm. Res.,
603 137, 231–246. doi:10.1016/j.jvolgeores.2004.05.018

604 Pedersen, G. B. M., A. Höskuldsson, T. Dürig, T. Thordarson, I. Jónsdóttir, M. S. Riishuus, B. V.
605 Óskarsson, S. Dumont, E. Magnusson, M. T. Gudmundsson, F. Sigmundsson, V. J. P. B.
606 Drouin, C. Gallagher, R. Askew, J. Gudnason, W. M. Moreland, P. Nikkola, H. I. Reynolds, J.
607 Schmith and the IES eruption team (2017). Lava field evolution and emplacement dynamics of
608 the 2014–2015 basaltic fissure eruption at Holuhraun, Iceland, J. Volcanol. Geotherm. Res.,
609 10.1016/j.jvolgeores.2017.02.027.

610 Pfeffer, M. A., B. Bergsson; S. Barsotti; G. Stefánsdóttir; B. Galle, S. Arellano, V. Conde, A.
611 Donovan, E. Ilyinskaya, M. Burton, A. Aiuppa, R. C. W. Whitty, I. C. Simmons, Þ. Arason, E.
612 B. Jónasdóttir, N. S. Keller, R. F. Yeo, H. Arngrímsson, Þ. Jóhannsson, M. K. Butwin, R. A.
613 Askew, S. Dumont, S. von Löwis, Þ. Ingvarsson, A. La Spina, H. Thomas, F. Prata, F. Grassa,
614 G. Giudice, A. Stefánsson, F. Marzano, M. Montopoli and L. Mereu (2017) Ground-Based
615 Measurements of the 2014–2015 Holuhraun Volcanic Cloud (Iceland), Geosciences,8, 29.

- 616 Pieri, D. and S. M. Baloga (1986) Eruption rate, area, and length relationships for some Hawaiian
617 lava flows, *J. Volcanol. Geotherm. Res.*, 30 (1), 29-45.
- 618 Ramsey, M. and A. J. L. Harris (2013) Volcanology 2020: How will thermal remote sensing of
619 volcanic surface activity evolve over the next decade? *J. Volcanol. Geotherm. Res.*, 249, 217-
620 233.
- 621 Sigmundsson, F., A. Hooper, S. Hreinsdóttir, K. S. Vogfjörð, B. G. Ófeigsson, E. R. Heimisson, S.
622 Dumont, M. Parks, K. Spaans, G. B. Gudmundsson, V. Drouin, T. Árnadóttir, K. Jónsdóttir, M.
623 T. Gudmundsson, T. Högnadóttir, H. M. Fridriksdóttir, M. Hensch, P. Einarsson, E.
624 Magnússon, S. Samsonov, B. Brandsdóttir, R. S. White, T. Ágústsdóttir, T. Greenfield, R. G.
625 Green, A. R. Hjartardóttir, R. Pedersen, R. A. Bennett, H. Geirsson, P. C. La Femina, H.
626 Björnsson, F. Pálsson, E. Sturkell, C. J. Bean, M. Möllhoff, A. K. Braiden and E. P. S. Eibl
627 (2015) Segmented lateral dyke growth in a rifting event at Bárðarbunga volcanic system,
628 Iceland, *Nature*, 517, 191–195.
- 629 Sigurðsson, H. and R. S. J. Sparks (1978) Rifting episode in North Iceland in 1874–1875 and the
630 eruptions of Askja and Sveinagja, *Bull. Volcanol.*, 41, 149–167.
- 631 Simmons, I, M. A. Pfeffer, E. Calder, B. Galle, S. Arellano, D. Coppola and S. Barsotti (2017)
632 Extended SO₂ outgassing from the 2014-2015 Holuhraun lava flow field, Iceland, *Bull.*
633 *Volcanol.*, X. DOI: 10.1007/s00445-017-1160-6.
- 634 Tarquini, S. (2017) A review of mass and energy flow through a lava flow system: insights provided
635 from a non-equilibrium perspective, *Bull. Volcanol.*, 79, 64. [https://doi.org/10.1007/s00445-](https://doi.org/10.1007/s00445-017-1145-5)
636 [017-1145-5](https://doi.org/10.1007/s00445-017-1145-5).
- 637 Tarquini, S., M. de' Michieli Vitturi, E. Jensen, G. Pedersen, S. Barsotti, D. Coppola, M. A. Pfeffer
638 (2018) Modeling lava flow propagation over a flat landscape by using MrLavaLoba: the case
639 of the 2014–2015 eruption at Holuhraun, Iceland, Special Issue: MeMoVolc, *Annals of*
640 *Geophysics*, 61. <https://doi.org/10.4401/ag-7812>.

641 Walker, G. P. L. (1973) Lengths of lava flows. *Philosophical Transactions of the Royal Society*,
642 London, 274, 107–118.

643 Wooster, M. J., B. Zhukov and D. Oertel (2003) Fire radiative energy for quantitative study of
644 biomass burning: derivation from the BIRD experimental satellite and comparison to MODIS
645 fire products, *Remote Sens. Environ.*, 86, 83–107. [http://dx.doi.org/ 10.1016/S0034-](http://dx.doi.org/10.1016/S0034-4257(03)00070-1)
646 [4257\(03\)00070-1](http://dx.doi.org/10.1016/S0034-4257(03)00070-1).

647 Wright, R., S. Blake, A. J. L. Harris, and D. Rothery (2001) A simple explanation for the space-
648 based calculation of lava eruption rates, *Earth Planet. Sci. Lett.*, 192, 223 – 233.
649 [doi:10.1016/S0012-821X\(01\)00443-5](https://doi.org/10.1016/S0012-821X(01)00443-5).

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