

Contrasting crustal evolution processes in the Dharwar craton: Insights from detrital zircon U–Pb and Hf isotopes

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Abstract

New *in situ* U–Pb–Hf analyses of detrital zircons from across the Archaean Dharwar craton indicate significant juvenile crustal extraction events at ~3.3 and 2.7 Ga, and continuous extraction from 3.7–3.3 Ga. Reworking in the older western block at ~3.0 Ga marks the onset of cratonisation, most likely due to ‘modern’ plate tectonic processes, while reworking in both the western and younger eastern block at 2.55–2.50 Ga indicates accretion of the two terranes and final cratonisation much later than in most other Archaean terranes (~2.7 Ga). Different patterns of disturbance to the zircon U–Pb systematics reflect variations in both the U content of parent rocks and later metamorphic conditions. Tectonic links are observed between the Kaapvaal and western Dharwar cratons, and between the north China and eastern Dharwar cratons, though none of these links necessarily requires a consanguineous origin.

Keywords: Dharwar craton, zircon, U–Pb, Hf model ages, crustal evolution

1. Introduction

Archaean cratons provide a critical window into early Earth dynamics, preserving a record of crustal evolution processes that include the start of ‘modern’ plate tectonics and the development of supercontinents (e.g., Campbell and Allen, 2008; Næraa et al., 2012). However, these same processes can also destroy or rework substantial volumes of crust, such that only ~7% of the continental crust is older than 2.5 Ga, and the oldest extant Archaean terrane is only ~3.9 Ga (e.g., Hawkesworth et al., 2010; Mojzsis et al., 2014; Shirey and Richardson, 2011). Where Archaean rocks are exposed, bulk techniques such as Pb and Nd

1 isotopes can provide considerable information about continental formation, but these may
2 have been altered by later metamorphic events.
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5 Another method is to examine sedimentary units, which can preserve fragments of the
6 crust that are no longer exposed at the Earth's surface. Resistant detrital minerals such as
7 zircon have particular use in these studies, as they incorporate a range of isotopic and
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9 geochemical tracers and can survive multiple crystallisation and/or sedimentary events (e.g.,
10 Fedo et al., 2003). Modern *in situ* techniques combined with careful imaging allow direct
11 correlation of crystallisation ages with geochemical information from each growth zone of a
12 mineral, providing context for interpretation even without the parent rock(s). In this manner,
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14 a more complete record of a craton's evolution may be obtained, with the benefit of
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16 contributions from contrasting isotopic systems.
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27 The Dharwar craton of southern India is one such Archaean block, comprising >2.7
28 Ga trondjemite-tonalite-granodiorite (TTG) gneisses, volcano-sedimentary belts (>3.0 and
29 2.9–2.6 Ga) and 2.7–2.5 Ga calc-alkaline to potassic granitoids (e.g., Chadwick et al., 2007;
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31 Jayananda et al., 2000, 2006; Jayananda et al., 2013a; Manikyamba and Kerrich, 2012).
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33 These rocks preserve evidence for several cycles of supracrustal formation, deformation,
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35 metamorphism and granitic activity during the Precambrian (Table 1). Significant work has
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37 been undertaken on these units with bulk techniques, especially ϵNd , identifying considerable
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39 reworking of older crust into younger units (summary in Dey, 2013; Mohan et al., 2013a;
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41 Peucat et al., 2013), although very little zircon Hf isotopic work has been published to date
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43 (e.g., Mohan et al., 2013b; Santosh et al., 2014; Sarma et al., 2012). However, several
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45 important issues remain intensely debated, including the timing and nature of juvenile crust
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47 formation and crustal reworking events in the Dharwar craton and how they correlate with
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49 events recognised globally, the tectonic context of these events, and differences in the
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51 evolutionary histories between the two main blocks of the craton (e.g., Chardon et al., 2011;
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1 Dey, 2013; Jayananda et al., 2013b; Maibam et al., 2011; Manikyamba and Kerrich, 2012;
2 Peucat et al., 2013; Santosh et al., 2014). This study examines zircons from three supracrustal
3 units from different stratigraphic levels as a window into the evolution of the continental
4 crust from a new angle.
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10 11 12 **2. Geological background**

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14 The southern Indian peninsula consists of a northern Archaean domain (Dharwar
15 Craton) and a southern, chiefly Proterozoic, domain (e.g., Bhaskar Rao et al., 2003). The
16 Dharwar Craton is divided in turn into older western (WDC) and younger eastern (EDC)
17 blocks, although a separate central block has also been proposed (e.g., Jayananda et al., 2006;
18 Peucat et al., 2013; Swaminath and Ramakrishnan, 1981). The WDC is dominated by 3.35–
19 3.0 Ga TTG gneisses, most of which report ϵNd values equal to or less than CHUR,
20 indicating reworking of older crustal material (e.g., Beckinsale et al., 1980; Dey, 2013; Meen
21 et al., 1992). Highly deformed >3.0 Ga Sargur Group greenstone belts are interlayered with
22 the TTG units, and reached peak metamorphic conditions of ~8 kbar and 700–750°C at ~2430
23 Ma (Janardhan et al., 1982; Jayananda et al., 2013b; Raith et al., 1983). These belts are in
24 turn unconformably overlain by the younger (2.9–2.6 Ga) Dharwar-type greenstone belts,
25 consisting of metasedimentary and metavolcanic rocks which are divided into the lower
26 Bababudan and upper Chitradurga groups (e.g., Kumar et al., 1996; Swaminath and
27 Ramakrishnan, 1981). Minor ~2.60 Ga potassic granites record final cratonisation in the
28 WDC (Jayananda et al., 2006).
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51 The EDC is dominated by 2.7 Ga Kolar-type greenstone belts and 2.7–2.5 Ga calc-
52 alkaline felsic plutonic and volcanic rocks (e.g., Chardon et al., 2002; Dey et al., 2012;
53 Jayananda et al., 2000, 2013a). The Kolar-type belts are predominantly greenschist to
54 amphibolite facies metabasalts— both plume-related (basalts, komatiites, alkaline basalts)
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1 and arc-related (basalts, boninites, Nb-enriched basalt-adakites)—with subordinate felsic
2 volcanic rocks and metasediments (e.g., Balakrishnan et al., 1991; Jayananda et al., 2013a;
3 Manikyamba and Kerrich, 2012; Naqvi et al., 2006). Younger granitoids include gneissic
4 TTGs, syn-tectonic high-Mg diorites or sanukitoids, the atypical ‘Closepet-type’ granitoid
5 and K-rich leucogranites (e.g., Dey et al., 2003, 2009, 2012; Jayananda et al., 1995, 2000;
6 Moyen et al., 2001, 2003). Older (3.3–3.0 Ga) granitoids and metasediments are now only
7 present as remnants, though old and highly variable Nd model ages and xenocrystic zircon
8 U–Pb ages around the EDC suggest these units were once much more abundant (summary in
9 Dey, 2013; Jayananda et al., 2000; Maibam et al., 2011). Accretion of the EDC onto the
10 WDC at ~2.5 Ga was followed by the formation of widespread platformal basins (1.9–0.6
11 Ga), including the Cuddapah basin in the far east, which received detritus from the
12 surrounding Archaean basement (Saha and Mazumder, 2012; Saha and Tripathy, 2012).
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29 Structural comparisons between the two blocks provide important clues to late
30 Archaean crustal evolution processes. Chadwick et al. (2000, 2007) argued that the EDC was
31 formed during the Neoarchaean in a convergent setting where oceanic lithosphere was
32 subducted in a WNW direction below a Mesoarchaean foreland continental margin (i.e., the
33 WDC). The Neoarchaean schist belts of the EDC were therefore interpreted as intra-arc
34 basins, whereas those of the WDC (the Dharwar Supergroup) were considered marginal or
35 back-arc basins. Some workers have suggested that the WDC, already cratonised and
36 thickened to a large extent by 2.6 Ga, was deformed moderately by transcurrent shear
37 deformation and shortening at 2.55–2.50 Ga, creating the N–S to NNW trending structural
38 fabric of the Dharwar craton (e.g., Chardon et al., 2011). The EDC, however, evolved in a hot
39 Neoarchaean orogen characterised by magmatic accretion, consistent with the presence of
40 high temperature and low pressure metamorphic assemblages (Jayananda et al., 2012). This
41 block behaved differently due to lateral constrictional flow of viscous lower crust under
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1 terminal Neoproterozoic convergence (Chardon et al., 2011). These conclusions are supported
2 by geophysical work which has identified a much thicker keel underneath the WDC than the
3 EDC (e.g. Borah et al., 2014; Gupta et al., 2003).
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7 Samples were collected from three stratigraphic levels and both blocks of the Dharwar
8 craton (Fig. 1, Table 1). In the WDC, the presumably Mesoarchaean Sargur belt contains
9 quartzites, metapelites, banded magnetite quartzites, Mn-rich rocks, crystalline limestones,
10 amphibolites and metaultramafic rocks. While upper amphibolite to lower granulite facies
11 metamorphism and multiple episodes of deformation have obliterated the original
12 depositional structures of these supracrustal units, a shallow continental margin basin setting
13 can be envisaged from the rock association (Janardhan, 1994; Swaminath and Ramakrishnan,
14 1981).
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18 Sample SRG-2 was collected from a fuchsite-bearing quartzite band of the Sargur belt
19 in the type area of the Sargur supracrustals, about 4.5 km SE of Sargur town (76°26'20"E,
20 11°59'31"N). The rock is dominated by quartz grains with sutured contacts (Fig. 2a), in which
21 often large patchy quartz grains are surrounded by an irregular secondary granular matrix of
22 finer quartz grains. The subparallel arrangement of thin muscovite-rich layers imparts a
23 distinct foliation within the rock, and suggests it was originally a quartz-rich sedimentary
24 rock with a subordinate clay matrix which suffered considerable post-depositional
25 deformation, re-crystallisation and sub-grain formation. Deposition can be constrained to
26 between ~3.4 and 3.1 Ga based on whole-rock Sm–Nd isochron ages of 3125±120 Ma and
27 3352±110 Ma from related mafic-ultramafic units, and a U–Th–Pb chemical age of 3082±66
28 Ma for monazite from near the western margin of the Chitradurga belt, interpreted as the age
29 of metamorphism of the Sargur Group (Hokada et al., 2013; Jayananda et al., 2008;
30 Mukherjee et al., 2012).
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1 Sample CH3 was collected about 6 km north of Janakal village (76°19'09"E,
2 13°57'51"N) from a quartzite band at the western margin of the Chitradurga greenstone belt
3 (WDC). This area exposes northwest-trending mature clastic sediments, including quartz
4 arenites and quartz-pebble conglomerates, intercalated with metabasalt bands belonging to
5 the lower part of the Bababudan Group of the Dharwar Supergroup (Geological Survey of
6 India, 2006; Hokada et al., 2013). They are metamorphosed to the upper greenschist to lower
7 amphibolite facies. Sedimentological studies of the mature lower Bababudan conglomerates
8 and arenites implied deposition in a braided fluvial plains developed over a stable gneiss-
9 granite basement (Srinivasan and Ojakangas, 1986). The basement underlying the
10 Chitradurga greenstone belt consists of 3.35–3.0 Ga TTG gneisses and >3 Ga supracrustal
11 rocks of the Sargur belt, while overlying mafic volcanic rocks reported a whole-rock Sm–Nd
12 isochron age of 2.911 ± 0.049 Ga (e.g. Hokada et al., 2013; Jayananda et al., 2008; Kumar et
13 al., 1996; Peucat et al., 1993). Sample CH3 comprises mainly coarse to very coarse sand-
14 sized grains of mono- and poly-crystalline quartz, surrounded by smaller quartz grains (Fig.
15 2b). Subordinate thin muscovite flakes are arranged in thin subparallel layers defining a
16 foliation, which in places swerve around the larger quartz grains. It is evident that the rock
17 has suffered deformation and recrystallisation forming a tight interlocking mosaic of grains
18 with sutured contacts.

19 Sample GLCH-1 is a very coarse-grained sandstone collected about 4.5 km east of
20 Dorigallu village (78°04'51"E, 14°25'20"N), along the south-western margin of the
21 Proterozoic pericratonic Cuddapah basin (EDC). The sample belongs to the lower part of the
22 Gulcheru Quartzite Formation, which consists of unmetamorphosed, subhorizontal beds of
23 conglomerates, sandstones and shales deposited in a basin-margin fluvio-aeolian environment
24 (Basu et al., 2014). This formation, the lowermost unit of the Cuddapah Supergroup,
25 unconformably overlies 2.7–2.5 Ga basement granitoids and greenstone belts, which contain

1 minor Palaeoproterozoic granites, syenites and mafic dykes that limit the maximum age of
2 the Gulcheru Formation to around 2.1 Ga (e.g. Chardon et al., 2002; Chardon et al., 2011;
3 Dey et al., 2014b; French and Heaman, 2010; Jayananda et al., 2013a; Kumar et al., 2012;
4 Suresh et al., 2010). A mafic sill intruding the lower part of the Cuddapah Supergroup was
5 dated to between 1.8–1.9 Ga (Anand et al., 2003; Bhaskar Rao et al., 1995; French et al.,
6 2008), implying that initial sedimentation in the Cuddapah basin started at around 1.9–2.0
7 Ga. Framework grains, commonly surrounded by dark Fe-oxide rims, form a bimodal size
8 distribution within a fine-grained sericitic matrix (~10% of the rock). These framework
9 minerals consist of large 2500–500 μm rounded to subrounded grains of polycrystalline and
10 monocrystalline quartz with subordinate quartzite, chert and highly altered feldspar, and
11 interstitial fine sand-sized grains of subangular to subrounded quartz with minor chert and
12 altered feldspar (Fig. 2c).
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31 **3. Analytical methods**

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34 Zircons were separated from ~10 kg of samples at the Indian School of Mines by
35 crushing in a mortar and pestle and collecting the 375–75 μm fraction using disposable nylon
36 mesh. Further processing by Mozley Table, heavy liquids (bromoforn and methylene iodide)
37 and handpicking yielded a clean zircon cut. Grains were then cast into epoxy mounts,
38 polished to half height and photographed by cathodoluminescence (CL) imaging on a JEOL
39 JSM-6060 LV scanning electron microscope, all at the University of Portsmouth. All grains
40 were examined to identify growth zoning and contaminating features such as cracks and
41 inclusions. Typical CL images are presented in Fig. 3, and all data are presented in
42 Supplementary Table 1.
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56 U–Pb ages were measured by laser ablation quadrupole mass spectrometry (LA-Q-
57 ICP-MS) at the University of Portsmouth after Jeffries et al. (2003), using an Agilent 7500cs
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1 coupled to a New Wave Research UP-213 Nd:YAG laser. A 30 μm spot was rastered along a
2 45–60 μm line. Grains were analysed using the ribbon method to avoid an analytical bias
3 towards ‘nice’ grains (e.g. Mange and Maurer, 1992). The amount of ^{204}Pb in these analyses
4 was below the detection limit, and no common Pb correction was undertaken. Ratios were
5 calculated using an in-house spreadsheet based on LamTool (Kořler et al., 2008), measuring
6 GJ-1 as the primary standard and Plešovice as the secondary; all uncertainties were
7 propagated in quadrature. Average $^{206}\text{Pb}/^{238}\text{U}$ and $^{207}\text{Pb}/^{206}\text{Pb}$ ratios of GJ-1 were
8 0.09761 \pm 0.00178 and 0.06012 \pm 0.00206 (n=188, 2SD), respectively, and of Plešovice
9 0.05432 \pm 0.00142 (n=38, 95%) and 0.05321 \pm 0.00191 (n=38, 2SD), consistent with published
10 references (Jackson et al., 2004; Sláma et al., 2008). Previous work under the same
11 conditions, using 91500 as the secondary standard before it was polished away, yielded
12 $^{206}\text{Pb}/^{238}\text{U}$ and $^{207}\text{Pb}/^{206}\text{Pb}$ ratios of 0.17642 \pm 0.00395 and 0.07465 \pm 0.00243 (n=37, 2SD),
13 respectively. The resulting average $^{207}\text{Pb}/^{206}\text{Pb}$ age of 1059 \pm 97 Ma (n=37, 2SD) provides
14 confidence that using such a young primary standard yields correct $^{207}\text{Pb}/^{206}\text{Pb}$ ages on
15 Archaean samples, albeit with large uncertainties. Only ages from a single growth zone and
16 avoiding irregular features such as cracks and inclusions were used; concordant ages are
17 those with less than 10% difference between their $^{207}\text{Pb}/^{206}\text{Pb}$ and $^{206}\text{Pb}/^{238}\text{U}$ ages. Concordia
18 plots and discordant arrays were calculated using Isoplot 3.07 (Ludwig, 2003).

19 U, Pb and Th concentrations were calculated for each analysis using the average
20 values for GJ-1 given in Jackson et al. (2004). Concentrations in Plešovice from this study
21 averaged 361 \pm 86, 18 \pm 4 and 42 \pm 20 (n=37, 2SD), respectively, approximately half the mean
22 provided for pristine grains but consistent with the minimum observed values for pristine
23 material. Both GJ-1 and Plešovice have highly variable ‘accepted’ concentrations, but
24 previous work using 91500 as a secondary standard against GJ-1 yielded concentrations of
25 112 \pm 30, 20 \pm 6 and 40 \pm 14, slightly higher but within uncertainty of published values
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1 (Wiedenbeck et al., 2004). The range of concentrations observed in the Dharwar zircons
2 greatly exceeds these uncertainties, suggesting that while the absolute values should be
3 interpreted with caution, the patterns observed across each sample are sound.
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7 Hafnium isotope ratios were measured by LA-MC-ICP-MS at University College
8 Dublin, following Hawkesworth and Kemp (2006). Every separated grain with growth zoning
9 of sufficient size and a U–Pb age <10% discordant was analysed, using 40 µm spots directly
10 over the U–Pb tracks. JMC-475 averaged $^{176}\text{Hf}/^{177}\text{Hf}=0.282144\pm 0.00004$ (2SD), yielding a
11 correction of +0.000016, while regular analyses of Mud Tank zircon (Yb/Hf ~0.001)
12 produced a JMC-corrected $^{176}\text{Hf}/^{177}\text{Hf}$ ratio of 0.2825043 ± 0.000025 (n=115, 2SD). An
13 exponential $^{173}\text{Yb}/^{171}\text{Yb}$ mass bias correction was applied using an in-house spreadsheet.
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15 Analyses of the Temora-2 standard zircon (Yb/Hf ~0.07) produced a JMC-corrected
16 $^{176}\text{Hf}/^{177}\text{Hf}$ ratio of 0.282677 ± 0.000019 (n=12, 2SD), consistent with the published value of
17 0.282680±0.000024 (2SD; Woodhead et al., 2004). To further test the reliability of this
18 correction over the range of Yb/Hf observed in this study, a synthetic ‘zircon’ solution of 200
19 ppb JMC-475 and 0.5 ppb natural Lu was doped with a natural Yb standard. No effect was
20 observed up to a Yb/Hf of ~0.06 ($^{176}\text{Lu}/^{177}\text{Hf}=0.282144\pm 0.0009$, 2SD), while concentrations
21 up to a Yb/Hf of ~0.33 reported slightly lower corrected $^{176}\text{Lu}/^{177}\text{Hf}$ ratios
22 (0.282136±0.00014, 2SD), though still within uncertainty of undoped JMC-475.
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26 εHf values were calculated based on a two-stage model, using the bulk Earth
27 (chondrite uniform reservoir; CHUR) $^{176}\text{Hf}/^{177}\text{Hf}$ and $^{176}\text{Lu}/^{177}\text{Hf}$ from Bouvier et al. (2008),
28 depleted mantle (DM) $^{176}\text{Hf}/^{177}\text{Hf}$ and $^{176}\text{Lu}/^{177}\text{Hf}$ from Griffin et al. (2002), and the Lu decay
29 constant from Söderlund et al. (2004). The slight lowering observed for high Yb analyses
30 translates into a +0.2 εHf unit shift at 3 Ga, which is insignificant compared to a typical 2σ
31 precision of 1.2 εHf units. Model ages were calculated assuming an average crustal
32 $^{176}\text{Lu}/^{177}\text{Hf}$ value of 0.015 (Griffin et al., 2002). If the value of Chauvel et al. (2014) is used
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1 instead, individual model ages decrease by <100 Ma, on the same order as the 2σ precision
2 on individual U–Pb ages, but the overall conclusions do not change.
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6 7 **4. Results**

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9 The two samples from the WDC (SRG-2 and CH3) present quite different age ranges
10 and Concordia distributions. Th older SRG-2 sample yielded 93 valid analyses, nearly all of
11 which are concordant, but form a continuous smear along Concordia from *c.* 3600 to 2750
12 Ma (Fig. 4a). While it is possible each age records a unique event within this interval, the
13 typically homogeneous and unzoned appearance in CL (Fig. 3) and very high correlation
14 coefficients (ρ) are more consistent with disturbance during later high grade metamorphism
15 (e.g., Halpin et al., 2012; Whitehouse and Kemp, 2010). Concordia is comparatively flat in
16 the Archaean due to the longer half life of ^{238}U , so ancient Pb loss typically drags analyses
17 along the curve, making it much less apparent. In these cases, additional evidence for or
18 against resetting may come from comparing the U–Pb ages to a more resilient isotopic
19 system, such as Hf isotopes, which are discussed later.
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36 In the younger CH3 sample (n=33), nearly half (n=16) are concordant, with the
37 remaining ages plotting along an apparent discordant array or scattering to older discordant
38 ages (Fig. 4b). These older ages indicate at least two more source areas/rocks were
39 contributing to the detritus in this sample; however, constraining the older source(s) is
40 impossible without any clear pattern to these ages. Arranging the data younger than 3.3 Ga in
41 order of decreasing $^{207}\text{Pb}/^{235}\text{U}$, and alternately rejecting the highest and lowest analyses until
42 a minimum MSWD is reached, yields a line with intercepts at 527 ± 420 and 3284 ± 92 Ma
43 (n=13, MSWD=8.3). The lower intercept age may indicate resetting during the Pan-African
44 event (~550 Ma; Brandt et al., 2014), since half of the analyses which define it contain <100
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1 ppm U (Fig. 5b; e.g., Mezger and Krogstad, 1997), but could also be geologically
2 meaningless.
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5 In the EDC, sample GLCH-1 (n=65) yields slightly fewer concordant ages (n=25; Fig.
6 4c). Following the same method as for CH3, two discordant arrays between can be calculated
7 between 817 ± 280 and 2647 ± 40 Ma (n=18, MSWD=4.2) and 55 ± 460 and 2511 ± 61 Ma
8 (n=13, MSWD=3.5). The former consists of mostly low U zircons (Fig. 5c), and the lower
9 intercept may correspond to regional magmatism identified in the Southern Granulite Terrane
10 at ~800 Ma (Brandt et al., 2014). The latter contains only two analyses with <100 ppm U, but
11 coincides with terminal crustal reworking and potassic granitic magmatism at 2.52–2.51 Ga
12 (e.g., Dey et al., 2014b; Jayananda et al., 1995, 2013a).
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24 Another method for identifying Pb loss, especially the cryptic loss suspected in SRG-
25 2, is to plot $^{176}\text{Hf}/^{177}\text{Hf}$ ratios against the measured U–Pb crystallisation ages (Fig. 5). Hf is
26 bound into the zircon structure by direct substitution for Zr, so any change to the Pb
27 systematics will lead to lines of similar $^{176}\text{Hf}/^{177}\text{Hf}$ ratios over a range of U–Pb ages. These
28 sub-horizontal arrays mimic evolution lines of $^{176}\text{Lu}/^{177}\text{Hf}=0.005$, as is seen in sample
29 GLCH-1 which displayed clear evidence for Pb loss in Fig. 4c. Two identical lines can be
30 drawn for the data from SRG-2, one through those reporting $^{176}\text{Hf}/^{177}\text{Hf}$ ratios of ~0.2808 and
31 the other through those of ~0.2806. As such, the analyses falling along these lines are
32 interpreted to have experienced Pb loss similar to that of GLCH-1 and CH3, despite the lack
33 of discordant arrays on a standard Concordia plot.
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48 ϵHf values have been plotted at both their measured and ‘best estimate’ U–Pb
49 crystallisation ages (Fig. 7). Against their measured ages, ϵHf values in both GLCH-1 and
50 CH3 are nearly all positive (0–5), while the majority in SRG-2 are slightly positive to
51 negative (-10–2). These values correspond to extraction ages of 2800–2700 Ma for GLCH-1
52 and ~3300 Ma for CH3, while individual model ages in SRG-2 fall between 3700 and 3300
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1 Ma. Estimated ages are determined as either the upper intercept of the relevant discordant
2 array (GLCH-1 and CH3) or as the age at which the relevant $^{176}\text{Lu}/^{177}\text{Hf}$ evolution line
3 intersects the data array clustered along CHUR (SRG-2). These estimated U–Pb ages and the
4 resulting Hf model ages are also presented in the Supplementary data table.
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10 11 **5. Discussion**

12 *5.1 Crustal evolution from individual samples*

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16 Clear differences are observed both between the WDC and EDC, as well as between
17 the two samples from the WDC, and comparing these data provides greater insight into
18 crustal processes during a critical transitional period in Earth's history (Fig. 7). The two
19 samples from the WDC contain overlapping ranges of U–Pb ages, but mutually exclusive ϵHf
20 distributions. The older SRG-2 sample contains zircon cores up to 3555 Ma, associated with
21 both oscillatory (magmatic) and unzoned (metamorphic or metamict) patterns in CL images.
22 Similar ages are reported by Nutman et al. (1992), who obtained detrital zircon U–Pb ages
23 from Sargur-type schist belts and concluded that 3.58–3.13 Ga granitoids supplied the
24 detritus. The same authors noted a younger age component of 3.13–2.96 Ga, which they
25 interpreted as the result of high-grade metamorphism due to intrusion of surrounding
26 granitoids.
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43 More negative ϵHf values correlate with younger U–Pb ages and a transition to
44 overgrowths on existing cores, recording U–Pb ages between 2550–2450 Ma. However, ϵHf
45 values plot very close to CHUR until ~3000 Ma, before suddenly shifting to negative values
46 and overlapping older zircons in ~2550 Ma greywackes from the Gadag greenstone belt
47 (Sarma et al., 2012). Since older U–Pb ages typically report lower U concentrations (Fig. 5a),
48 younger zircons with higher U concentrations, and hence more radiation damage, were more
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1 affected by open system behaviour during later metamorphism (e.g., Kooijman et al., 2011;
2 Mezger and Krogstad, 1997).
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4 The smear of U–Pb ages along Concordia (Fig. 4a) also complicates the determination
5 of Hf model ages for the parent rocks of detritus in SRG-2. Model ages (T_{DM}) for grains
6 reporting U–Pb ages older than ~3000 Ma should be the most reliable, as neither the U–Pb
7 nor Hf systems should have been disturbed. T_{DM} calculated for these grains using the typical
8 upper continental crust $^{176}\text{Lu}/^{177}\text{Hf}$ of 0.015 range between 3700 and 3500 Ma, with a small
9 number recording model ages up to 4200 Ga. In addition, one zircon core (3466 Ma) plots
10 very close to the DM line, recording a model age of 3500 Ma. Examining the younger grains,
11 which most likely experienced Pb loss, provide an overall model age of ~3400 Ma. Together,
12 these data suggest the Sargur belt was receiving detritus from rocks extracted over ~300 Ma
13 and evolved to zircon-bearing compositions within 400 Ma of extraction, consistent with
14 global averages (e.g., Lancaster et al., 2011).
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31 By contrast, CH3 contains very few concordant U–Pb analyses, all of which report
32 positive ϵHf values. These zircons overlap the range of U–Pb ages reported elsewhere in the
33 Chitradurga belt and its northern extension, the Gadag belt (e.g., Hokada et al., 2013; Sarma
34 et al., 2012). Very few zircons preserve magmatic zoning in CL, and the youngest concordant
35 magmatic U–Pb age is 3200 Ma. Once again the oldest ages are associated with the lowest U
36 contents (Fig. 5b), while the discordance pattern observed in the U–Pb ages may be due to
37 nanoscale recrystallisation during greenschist facies metamorphism (e.g., Hay and Dempster,
38 2009b). Apart from a single core dated at 3438 Ma, which plots within the SRG-2 data field
39 and reports a T_{DM} of ~3700 Ma, nearly all the CH3 zircons plot within uncertainty of each
40 other (U–Pb) and the DM (ϵHf), and report model ages within 300 Ma of crystallisation at
41 ~3200 Ma. Because these data plot so close to the DM, the choice of $^{176}\text{Lu}/^{177}\text{Hf}$ in the parent
42 rock has little impact on the calculated model age. Together, these observations suggest the
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1 detritus in this unit was derived from a single point source, extracted from the mantle ~3300
2 Ma and reworked to a zircon-bearing composition at ~3200 Ma. Since these zircons record a
3 similar range of U–Pb ages and ϵHf values as the northern Gadag belt (Sarma et al., 2012),
4 rocks of this description were most likely common within the source area for these
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Sample GLCH-1 reflects the younger rocks found in the EDC. Much like sample
CH3, few U–Pb ages are concordant, in this case between 2400 and 2700 Ma, with the rest
falling along two very similar discordant arrays (Fig. 4c). Roughly half the analyses were on
growth layers with oscillatory zoning in CL, but these analyses were just as likely to be
discordant as those with other forms of zoning. The youngest concordant magmatic age,
based on oscillatory zoning in CL imaging, is ~2500 Ma, coincident with all concordant
metamorphic ages (homogeneous, usually dark, appearance in CL), but does not represent the
time of deposition as this basin developed during intracontinental rifting at ~1900 Ma (e.g.,
French et al., 2008; Saha and Tripathy, 2012). Evidence for exposed Palaeoarchaeon material
amongst the more abundant Neoarchaeon–Palaeoproterozoic basement comes from several
sources. Detrital zircons from the younger Owk Shale Fm. (Kurnool Group; Table 1),
overlying the Cuddapah Basin sediments and traditionally believed to be Neoproterozoic in
age, also record U–Pb ages between 2690 and 2340 Ma, with a single grain at 3300 Ma
(Bickford et al., 2013), while remnants of Mesoarchaeon granitic gneisses and
Palaeoarchaeon to Mesoarchaeon whole rock Nd depleted mantle model ages and inherited
zircons yielded by Neoarchaeon granitoids and felsic volcanic rocks of the EDC (Dey, 2013;
Dey et al., 2014b; Jayananda et al., 2013a; Mohan et al., 2013a).

Nearly all concordant grains have positive ϵHf values, which correspond to Hf model
ages between 2700 and 3000 Ma. Since nearly all data overlap within uncertainty, though,
these most likely derive from a single source with an extraction age between 2700 and 2800

1 Ma depending on the $^{176}\text{Lu}/^{177}\text{Hf}$. Whole rock ϵNd work on basement granitoids and
2 greenstones exposed in the immediate vicinity of GLCH-1 has identified juvenile crustal
3 extraction from 2700–2600 Ma followed by extensive reworking in the EDC from 2580–
4 2520 Ma (Dey et al., 2014b, a; Jayananda et al., 2013a), consistent with the results of GLCH-
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10 1. Rocks with similar U–Pb and ϵHf appears to be common both within and outside the
11 Dharwar craton, as similar zircons are reported in the older Aravalli sediments to the
12 northwest (Kaur et al., 2011, 2013) and felsic volcanic units from southern India (Praveen et
13 al., 2014). Basement anorthosites from the Sittampundi Complex, to the south of the Dharwar
14 craton, also record the same distribution of zircon U–Pb ages, a regional metamorphic event
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22 ~715 Ma and ϵHf values which directly overlap those in this sample (Mohan et al., 2013b).
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26 *5.2 Crustal evolution in the wider Dharwar craton*

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29 While each detrital sample only records a small part of the craton's evolutionary
30 history, together they provide a more complete picture. Although collected far from each
31 other, the zircons in both SRG-2 and CH3 overlap published data from around the WDC,
32 suggesting they are representative of broader crustal processes in the region. The oldest
33 zircons in SRG-2 form a continuous array around CHUR until ~3000 Ma, suggesting rocks
34 which evolved to zircon-bearing compositions were extracted continuously from the mantle
35 from ~3700 to at least 3400 Ma. In addition, both SRG-2 and CH3 record zircons >3000 Ma
36 with similar U–Pb ages but different ϵHf values. Together, these observations suggest the
37 WDC contains discrete fragments of crust extracted at different times, in keeping with the
38 model of pre-modern plate tectonics whereby it was similarly easy to make and destroy new
39 crust (e.g., Hawkesworth et al., 2009).
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56 TTG gneisses aged between 3400–3200 Ma are well exposed in the WDC (e.g.,
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1 Peucat et al., 1993). While reworking of older crust is associated with the extraction of
2 juvenile material (e.g., Meen et al., 1992), the preponderance of zircons of these ages
3 reporting $\epsilon_{\text{Hf}} > \text{CHUR}$ suggests little mixing between juvenile and pre-existing crust, a
4 conclusion also reflected by ϵ_{Nd} data from these units (e.g., Devaraju et al., 2007; Jayananda
5 et al., 2008; Peucat et al., 1993). Regional metamorphism is first noted at ~3100–3000 Ma
6 (Jayananda et al., 2013b), which coincides with both crystallisation in the juvenile source of
7 CH3 and the onset of Pb loss in SRG-2. These oldest zircons report the lowest U contents
8 (typically <100 ppm), but there is no clear cut-off in U concentration between the presence or
9 absence of open system behaviour. While U(-Th-Rb) depletion in whole rocks is linked to
10 sub-solidus metamorphic conditions (900–1000°C, ~10 kbar; Cohen et al., 1991), Meen et al.
11 (1992) report depletion on this scale did not occur until the ~3000 Ma event in the WDC.

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27 However, the annealing temperature of zircon is only 600–650°C (Mezger and
28 Krogstad, 1997), in the upper amphibolite facies at typical crustal thicknesses. Below this
29 temperature, U-rich domains will preferentially accumulate radiation damage to the zircon
30 structure, rendering them open to Pb leakage in later metamorphic events. The zircons in
31 SRG-2 and CH3 all responded differently during the ~3000 Ma event, suggesting a change in
32 tectonic processes around this time which left zircons younger than this age preferentially
33 susceptible to later alteration. The regional scale and nature of this change therefore most
34 likely indicates a metamorphic event with significant variation in crustal depths and P–T
35 conditions, such as initial cratonisation. Pb loss most likely occurred during the next major
36 regional metamorphic episode at ~2550 Ma (e.g., Jayananda et al., 2013b), corresponding to
37 the youngest ages in SRG-2.

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If the ~3000 Ma event does mark the start of terrane accumulation and a significant change in tectonic setting in the WDC, clear evidence should be found in the chemical and physical structure of the block. Seismological work by Borah et al. (2014) identified that the

1 lithospheric keel is considerably thicker under the WDC than the EDC, as is the overall
2 thickness of the crust, suggesting cratonisation was of a longer duration in this block, with
3 strong, eclogitic-like material at the base of the crust to provide stability. Hoffmann et al.
4 (2011) considered bulk and trace element distributions in ~3000 Ma TTGs in the WDC which
5 indicate derivation from thickened mafic crust, with mixing between rutile-bearing and rutile-
6 free eclogitic residue. The event which depleted many WDC basement rocks in heat-
7 producing elements at ~3000 Ma left them resistant to both reworking and intrusion by
8 younger granites during later regional metamorphism at ~2550 Ma (Meen et al., 1992).
9 Together, it is clear the ~3000 Ma event marks a critical change in the evolutionary history of
10 the WDC, possibly the onset of 'plate tectonic' processes as suggested by large global studies
11 (e.g., Dhuime et al., 2012; Shirey and Richardson, 2011), and its effects have determined the
12 outcome of later intracrustal reworking.
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29 By contrast, the EDC is dominated by 2700–2500 granitoids and greenstone belts
30 which provided the detrital zircons in GLCH-1, with remnants of Meso- to Palaeoarchean
31 components (Dey et al., 2014b; Jayananda et al., 2000; Maibam et al., 2011). Typical
32 differences between extraction and crystallisation ages in this sample are 400–500 Ma,
33 consistent with the global average (e.g., Lancaster et al., 2011), suggesting 'modern' crustal
34 processes were established by this time. For instance, the ~2700 Ma Gadwal greenstone belt
35 preserves whole rock compositions consistent with minimal interaction with mature
36 continental crust, but considerable variation in ϵ_{Hf} and ϵ_{Nd} values (Khanna et al., 2014).
37 These authors concluded that the Gadwal belt records intra-arc variation, as is observed in
38 modern arcs such as the Aleutians (Yogodzinski et al., 1995; Yogodzinski et al., 2010), rather
39 than mixing. The lack of scatter in ϵ_{Hf} in the GLCH-1 zircons therefore reflects derivation
40 from a single, restricted source within one of these belts.
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1 Units of a similar age are also known in the WDC, particularly as felsic volcanic
2 layers (Jayananda et al., 2013a; Kumar et al., 1996; Mohan et al., 2014; Nutman et al., 1996).
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4 As such, the abundance of units preserving new and reworked crust in both blocks between
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6 2700–2400 Ma is the result of repeated accretion events concluding with regional
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8 metamorphism in both terranes up to the granulite facies at ~2520 Ma during final assembly
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10 of the Superia supercontinent (Jayananda et al., 2006, 2013a; Khanna et al., 2014; Peucat et
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12 al., 2013). It is this event which remobilised Pb in zircons from all three samples in this
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14 study, with the severity of that loss depending on the overall U content of each grain, the
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16 nature of any high-U domains within those grains and the metamorphic history each
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18 experienced (e.g., Hay and Dempster, 2009b, a; Malusà et al., 2013).
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26 *5.3 Correlations with other Archaean cratons*

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28 Similarities between Archaean terranes around the globe have led to many suggested
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30 correlations and palaeoreconstructions, with the inherent uncertainties increasing as we travel
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32 further back into the past. Rogers (1996) suggested the WDC formed the supercontinent Ur,
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34 together with the Pilbara and Kaapvaal cratons and East Antarctica, though this view is not
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36 entirely accepted (e.g., de Kock et al., 2009). Others have considered potential correlations
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38 between sequences in the WDC and the Kaapvaal or Slave cratons (e.g., Bleeker, 2003;
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40 Hokada et al., 2013) and the EDC with the Napier complex in East Antarctica and the N
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42 Australia craton (e.g., Hollis et al., 2014; Mohanty, 2011). Some authors have also suggested
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44 that the N China craton has a magmatic, sedimentological and metamorphic history similar to
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46 that of peninsular India, and they were possibly part of the same continent from the
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48 Mesoarchaeon to the Palaeoproterozoic (Hou et al., 2008; Zhao et al., 2003). Comparing the
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50 U–Pb and Hf zircon data from this study with published data from these belts provides
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52 another angle on this ongoing debate (Fig. 8a).
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1 The sub-horizontal spread of ϵHf values around CHUR >3000 Ma is only observed in
2 the three terranes within the Kaapvaal craton (Murchison-Northern Kaapvaal, Barberton
3 North and Barberton South) and the two WDC samples, while the kink in ϵHf values at
4 ~3000 Ma in SRG-2 is also observed in the three Kaapvaal terranes. Terranes of a similar age
5 elsewhere either appear to record short-duration ('episodic') events with vertical arrays, or
6 form long arrays parallel to $^{176}\text{Lu}/^{177}\text{Hf}$ evolution lines. While these observations may be an
7 artefact of preservation, it is also possible that the processes controlling crustal growth in the
8 Barberton and Dharwar cratons were different from those elsewhere. The most juvenile ϵHf
9 values within the sub-horizontal arrays form a line slightly above CHUR that mimics the
10 evolution of the depleted mantle. Over this time frame (~3500–3000 Ma), the difference
11 between extraction age (Hf) and crystallisation age (U–Pb) in these zircons only varies from
12 ~280–320 Ma, suggesting a continuous process of extraction and rapid evolution to zircon-
13 bearing compositions was interrupted only by the onset of cratonisation at ~3000 Ma. In this
14 case, the three main patterns observed in the global zircon data—vertical, horizontal and
15 diagonal—reflect coexisting methods of continental crustal production, and models of early
16 Earth evolution must incorporate them all to be truly representative.

17 Differences between the Barberton and WDC samples provide more clues to craton
18 formation in the Archaean. Unlike older greenstone belts in the Kaapvaal and Pilbara, the
19 Chitradurga belt (CH3) does not record reworking of pre-existing crust. Juvenile material
20 resembling that in CH3 is only rarely observed in zircons from any other craton, again
21 arguing against a single crustal evolution model (i.e., episodic plumes vs. constant accretion
22 via subduction) in the Palaeo- to Mesoarchaeon. However, the Neoarchaeon Chitradurga belt
23 is much younger than these other two belts, and the older sample (SRG-2) does provide clear
24 evidence of reworked crust, suggesting accretion via marginal basins was common at this
25 time (e.g., Kröner et al., 2013; Tessalina et al., 2010).

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In addition, most Archaean terranes record extensive crustal reworking at ~2700 Ma, thought to record the collisional phase of the Superia supercontinent, followed by relative quiescence (e.g., Campbell and Allen, 2008; Dey, 2013; Nance et al., 2014). While units of this age are abundant in both the Dharwar and Kaapvaal cratons, only the Dharwar craton records significant extraction and metamorphism after ~2700 Ma, reflecting cratonisation of the Kaapvaal at this time (e.g., Poujol et al., 2003; Rajesh et al., 2014). By contrast, the North China craton records significant juvenile extraction at ~2700 Ma, and both crustal reworking and granulite-facies metamorphism due to cratonisation at ~2500 Ma (e.g., Wan et al., 2014; Zhai and Santosh, 2011). These zircons fall on the same extraction lines as GLCH-1 and the southern Indian anorthosites, suggesting extraction of juvenile material was common at this time, but only preserved in those terranes which had yet to cratonise. Therefore, the Dharwar Craton preserves a more complete picture of Archaean crustal processes during this critical period than other terranes, and continental collisions were a prominent feature during the construction of Superia.

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Even if two terranes record similar physical processes, they do not necessarily share a consanguineous origin. For instance, the Limpopo belt, immediately to the north of the Barberton terranes, reports more negative zircon ϵ_{Hf} values and was derived from a high $^{238}\text{U}/^{204}\text{Pb}$ (μ) source (Barton, 1996), indicating terranes that are now adjacent need not have been extracted in close proximity to one another and vice versa. In addition, each terrane comprises a different suite of rocks. The Barberton greenstone belt is remarkably undeformed and consists of mafic–ultramafic volcanic rocks, greywackes, shales, cherts, felsic volcanoclastics, sandstones and conglomerates (e.g., Grosch et al., 2011; Kröner et al., 2013; Zeh et al., 2013). By contrast the Sargur belts (SRG-2) are dominated by highly deformed and metamorphosed quartzites, metapelites, metacarbonates, basalts, komatiites and barites (summaries in Jayananda et al., 2008; Meert et al., 2010); conglomerates and greywackes are

1 not reported. These observations suggest that while tectonically linked, they have been
2 geographically separate for most, if not all, of their histories.
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9 **6. Conclusions**

10 The Archaean Dharwar craton is a complex terrane comprising many discrete units accreted
11 onto each other through regular collisional events. Significant juvenile crustal extraction
12 events are recorded at ~3.3 and 2.7 Ga, with continuous extraction observed from 3.7–3.3 Ga.
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14 Cratonisation in the Dharwar craton occurred in two stages, initiating in the older WDC at
15 ~3000 Ma and the younger EDC at ~2500 Ma, and was completed considerably later than
16 most other Archaean cratons (~2500 Ma vs. ~2700 Ma). Repeated moderate-grade
17 metamorphic events created nanoscale zones of higher U concentrations within the zircon
18 structure, which preferentially leaked Pb during the cratonisation process. Further work is
19 needed to expand the zircon ϵ_{Hf} database to permit more complete comparisons between it
20 and more extensive U–Pb and ϵ_{Nd} records. Finally, tectonic links are observed between the
21 Barberton terranes of the Kaapvaal craton and the WDC, as well as the North China craton
22 and the EDC, though it does not necessarily require genetic links.
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Figure captions:

Figure 1: Geological map of the Dharwar craton (modified after the Vasundhara Project, Geological Survey of India, 1994) showing the location of samples. CSZ - Chitradurga shear zone, CGL - Closepet Granite. Sargur-type schist belts: B - Banasandra, G - Ghattihosahalli, Hn - Holenarasipur, J - J.C. Pura, K - Kalyadi, N - Nuggihalli, Sg - Sargur. Dharwar-type schist belts: Bb - Bababudan, C - Chitradurga, Ga - Gadag, Ku - Kudremukh, S - Sandur, Sh - Shimoga. Kolar-type schist belts: Gd - Gadwal, H - Hutti, Ka - Kadiri, HK - Hungund-Kushtagi, Ko - Kolar, R - Ramagiri, Rc - Raichur, V - Veligallu.

Figure 2. Photomicrographs of samples, all in XN. (a) SRG-2 from the Sargur belt. Deformed quartzite showing parallel alignment of muscovite layers imparting a distinct foliation. (b) CH3 from the Bababudan Group, Chitradurga belt. Deformed quartzite consisting dominantly of quartz with subordinate thin muscovite flakes, the latter arranged in thin subparallel layers defining a foliation. (c) GLCH-1 from the Gulcheru Formation, Cuddapah basin. A quartz arenite consisting of very coarse sand to granule-sized rounded to subrounded grains of quartz (both poly- and mono-crystalline) and chert (at the centre). Interstitial places are filled with fine sands of quartz.

Figure 3: Typical CL images from zircons in this study, indicating analytical locations and resulting data. Solid ellipses - U-Pb. Dashed circles - Hf. Data indicated as grain name (G), $^{207}\text{Pb}/^{206}\text{Pb}$ age $\pm 2\sigma$, and $\epsilon\text{Hf} \pm 2\sigma$, calculated at the indicated U-Pb age; % disc. indicates discordance when $>10\%$.

Figure 4: Concordia diagrammes for zircons from the three samples in this study, plotted at 2σ . Histogram and kernel density estimate curve of ages $<10\%$ in (a) plotted using the DensityPlotter tool of Vermeesch (2012). Core - innermost growth zone. Mantle - intermediate growth zone(s). Rim - outermost growth zone. Whole - no obvious internal zoning.

Figure 5: Plots of $^{207}\text{Pb}/^{206}\text{Pb}$ ages vs. U, Pb and Th concentrations in zircons from this study. Top row - U. Middle row - Pb. Bottom row - Th.

Figure 6: Plot of $^{176}\text{Hf}/^{177}\text{Hf}$ vs. U-Pb age for all data in this study, coded for growth zones as in Fig. 4, except mantles and rims are grouped as overgrowths (overs). Uncertainties plotted at 2σ .

Figure 7: Plot of ϵHf vs. U-Pb age for all data with U-Pb ages $<10\%$ discordant from this study, coded for growth zone as in Fig. 6. Uncertainties plotted at 2σ . Top: ϵHf values calculated using the measured $^{207}\text{Pb}/^{206}\text{Pb}$ ages. Bottom: ϵHf values calculated at the best estimate of the crystallisation age, based on the evidence for Pb loss in Figs. 4 and 6 (see text for discussion). CHUR - chondritic reservoir (bulk Earth). DM - depleted mantle; Griffin et al. (2002). NC - new crust; Dhuime et al. (2011). S08 - mantle model of Shirey et al. (2008).

Figure 8: Plot of ϵHf vs. U-Pb age for data $<10\%$ discordant in this study, plotted against published data from other Archaean terranes (references in the Supplementary Material). Data from this study are plotted at the best estimate U-Pb crystallisation ages, as discussed in

1 the text. Dark dashed lines indicate evolution of parental rocks with a $^{176}\text{Lu}/^{177}\text{Hf}$ of 0.015
2 extracted from the depleted mantle at the T_{DM} given in Ga; the line representing a $^{176}\text{Lu}/^{177}\text{Hf}$
3 of 0.005 used to adjust for Pb loss in SRG-2 is indicated in light grey. CHUR - chondritic
4 reservoir (bulk Earth). DM - depleted mantle; (Griffin et al.). Det - detrital. volcs - volcanics.
5 (a) Indian subcontinent. (b) Kalahari craton. MNK - Murchison-Northern Kaapvaal. BN -
6 Barberton north. GAB - Gaborone granite suite. NK-SMZ - Northern Kaapvaal, southern
7 marginal zone. LIM - Limpopo belt. BS - Barberton south. FRAN - Franciscan arc. (c) China,
8 Australia and Antarctica. JH - Jack Hills. (d) North Atlantic region and Baltica. Green -
9 Greenland.
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Figure 1 - colour for online only
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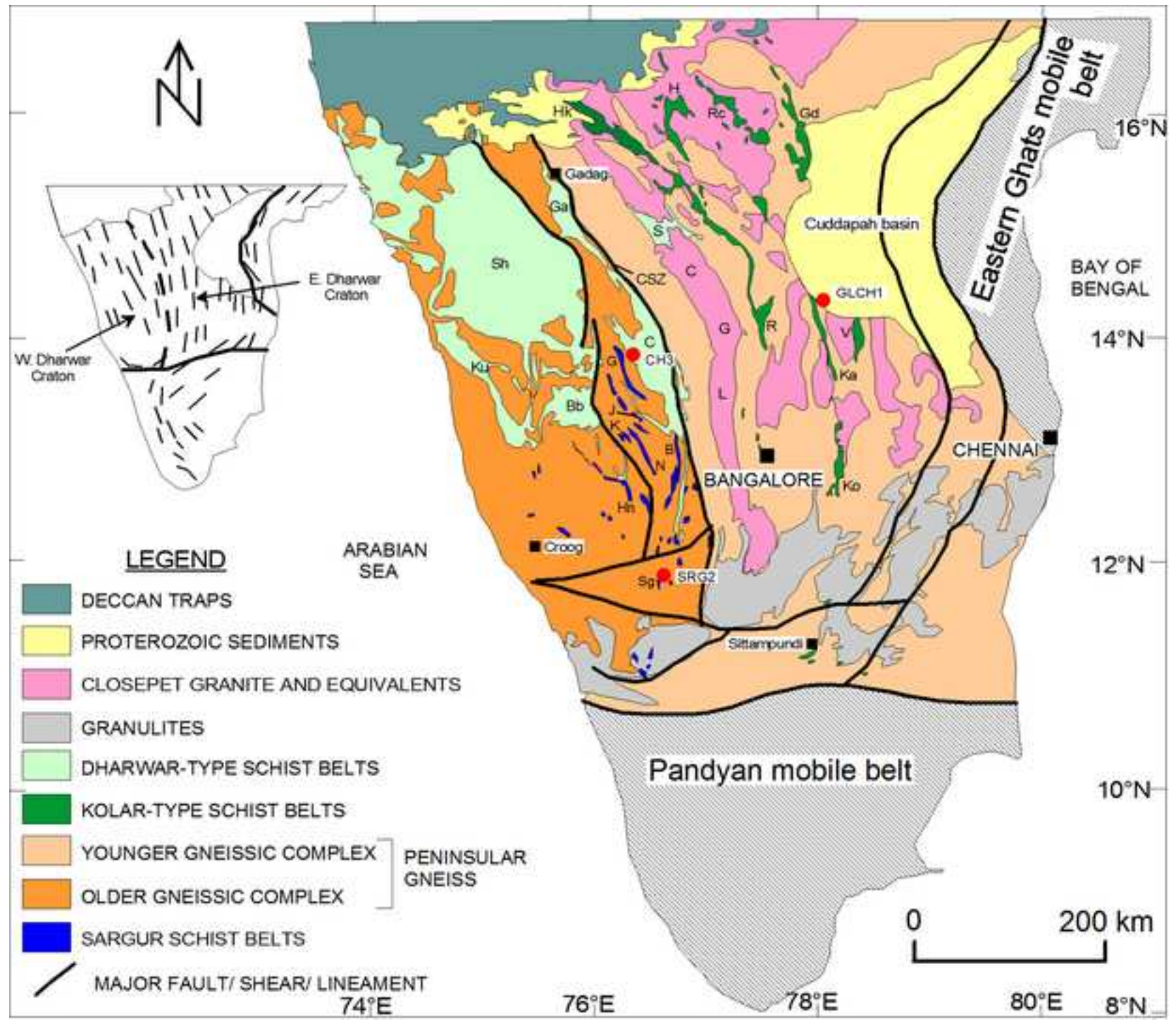


Figure 1 - B&W for print only

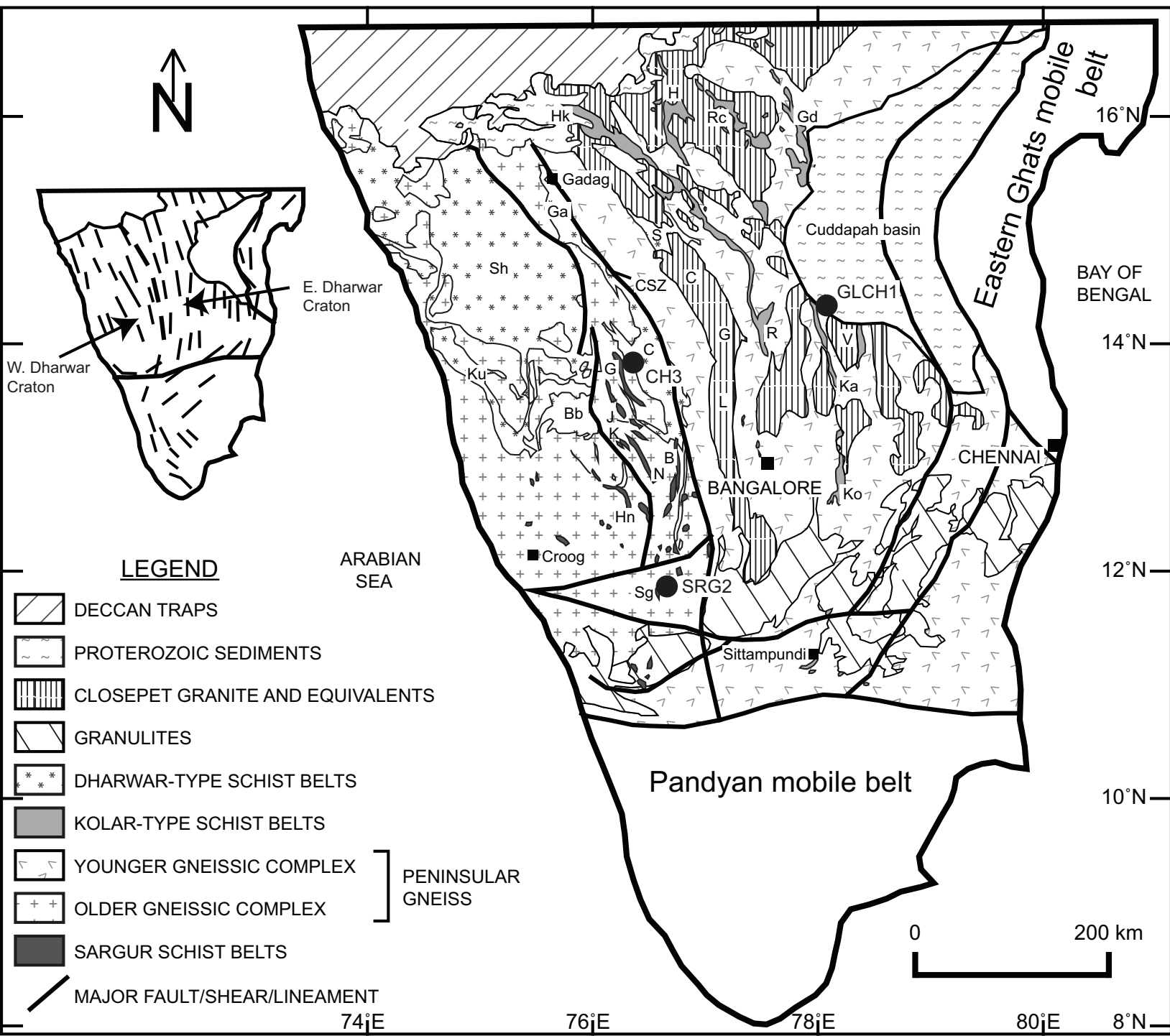
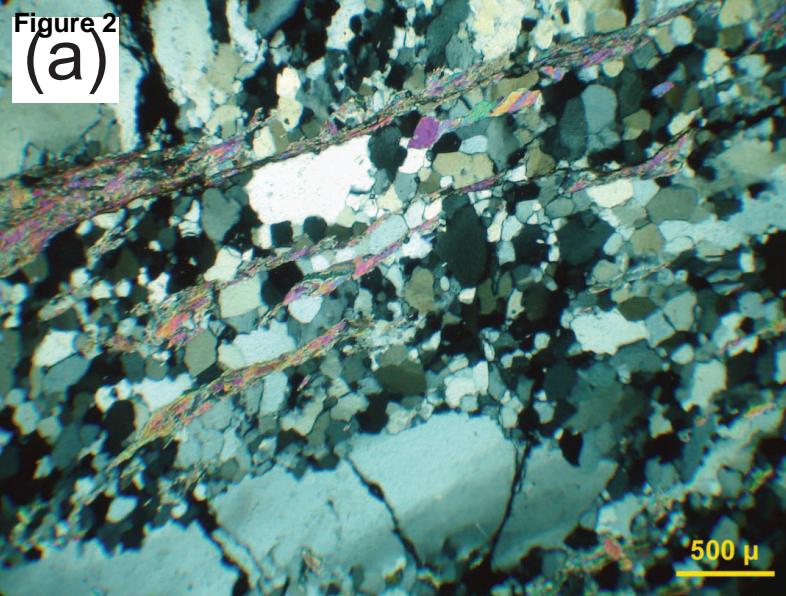
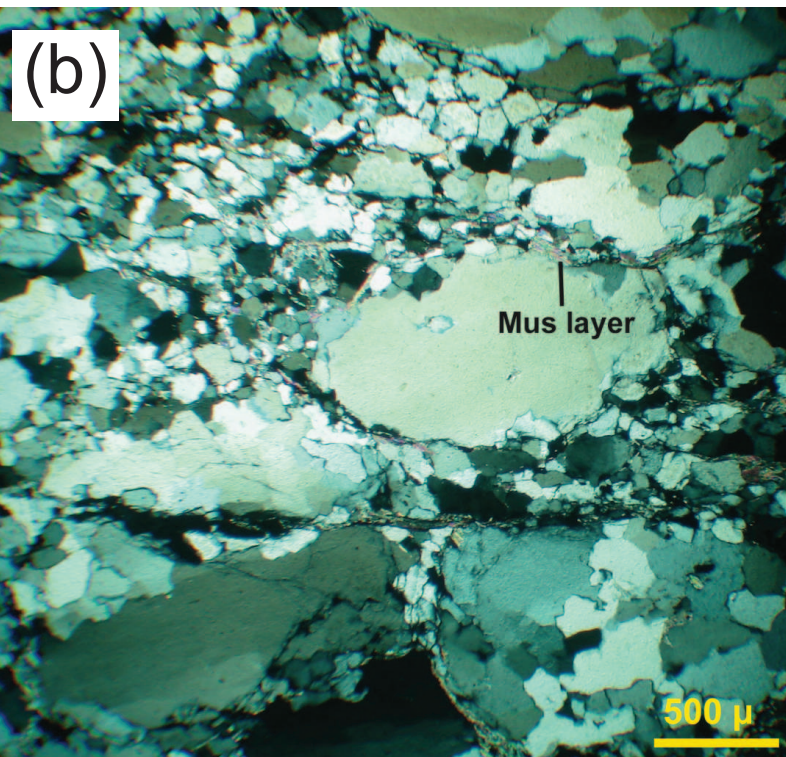


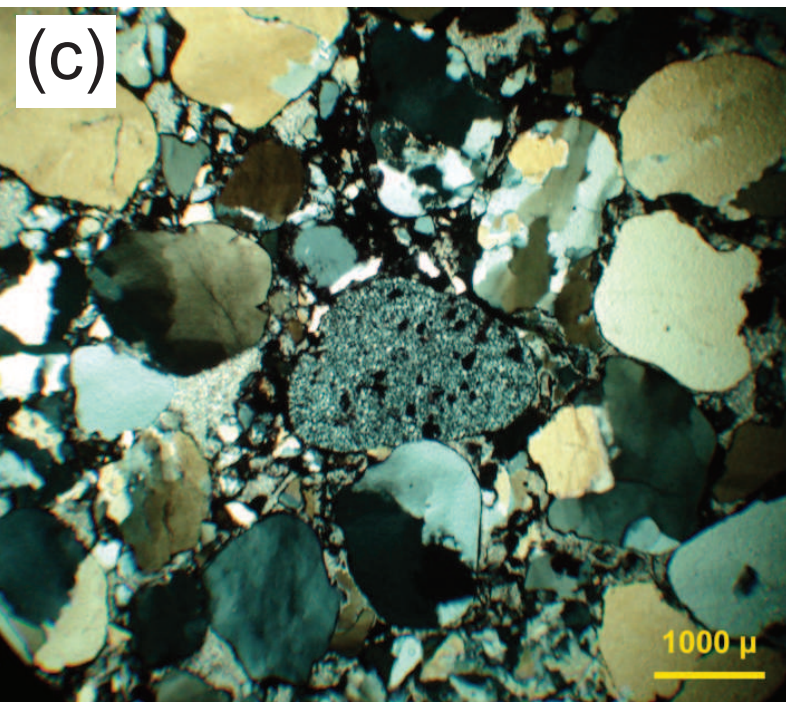
Figure 2
(a)

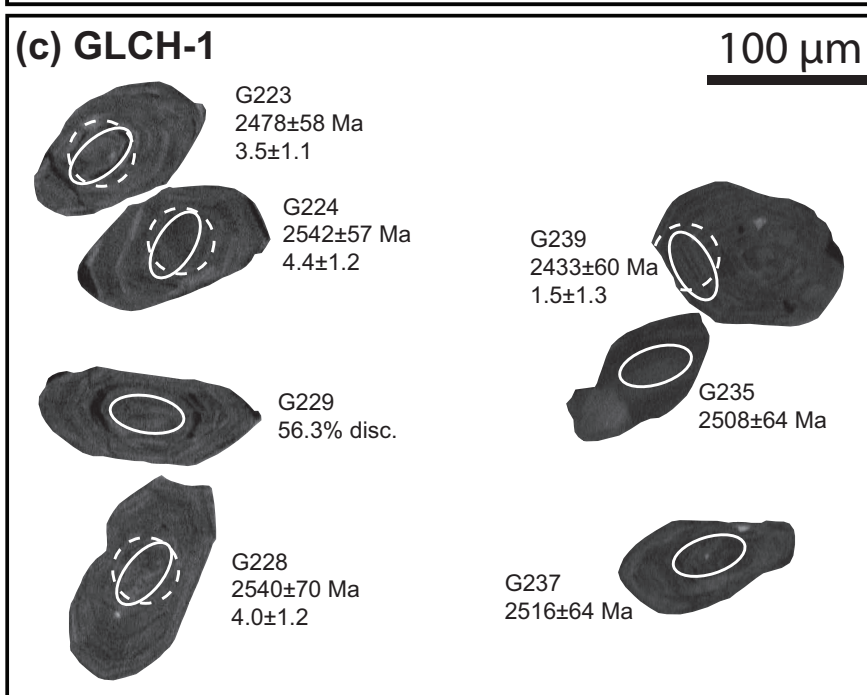
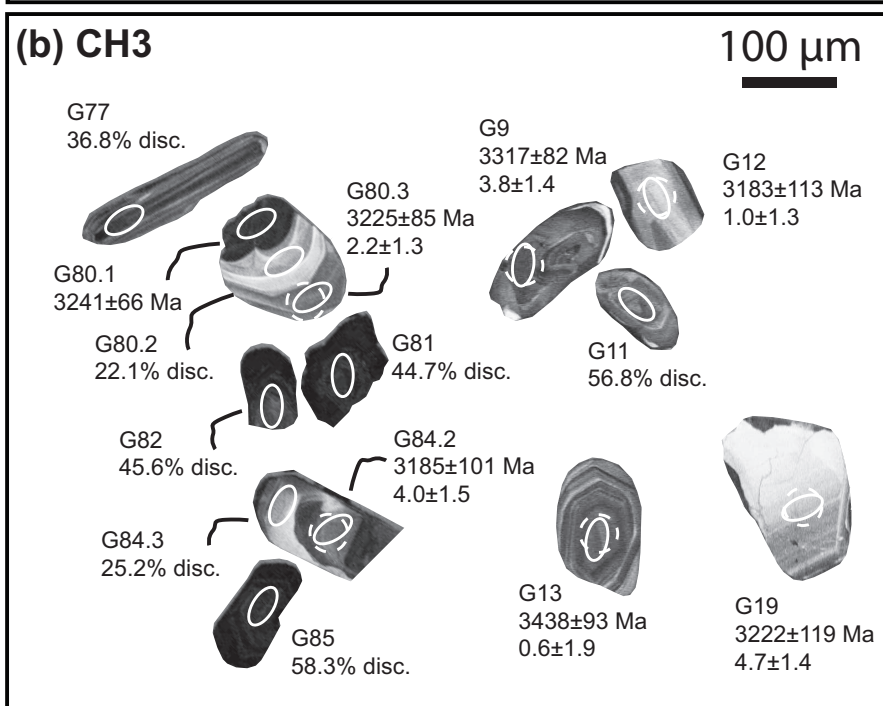
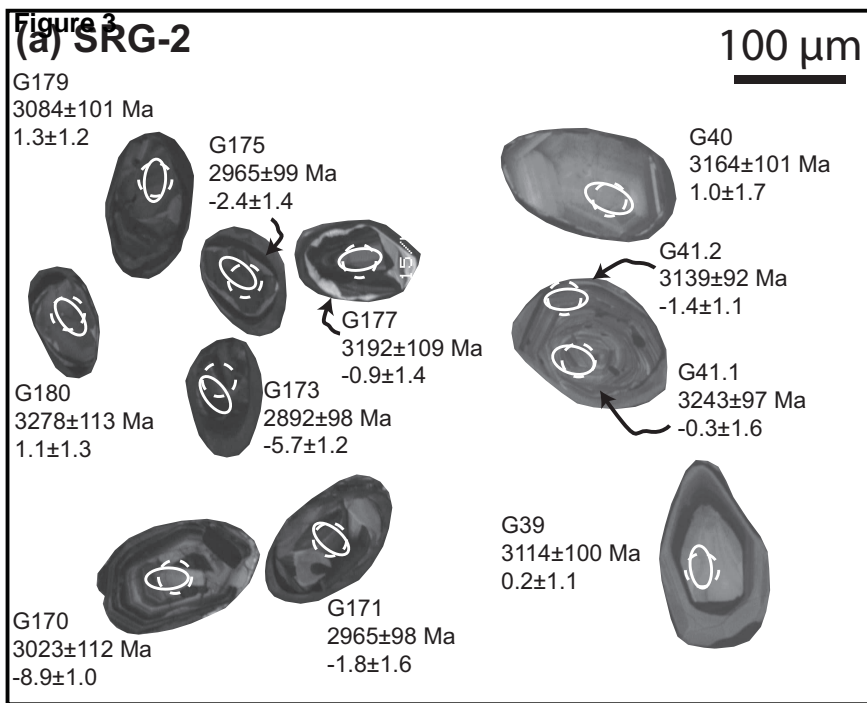


(b)



(c)





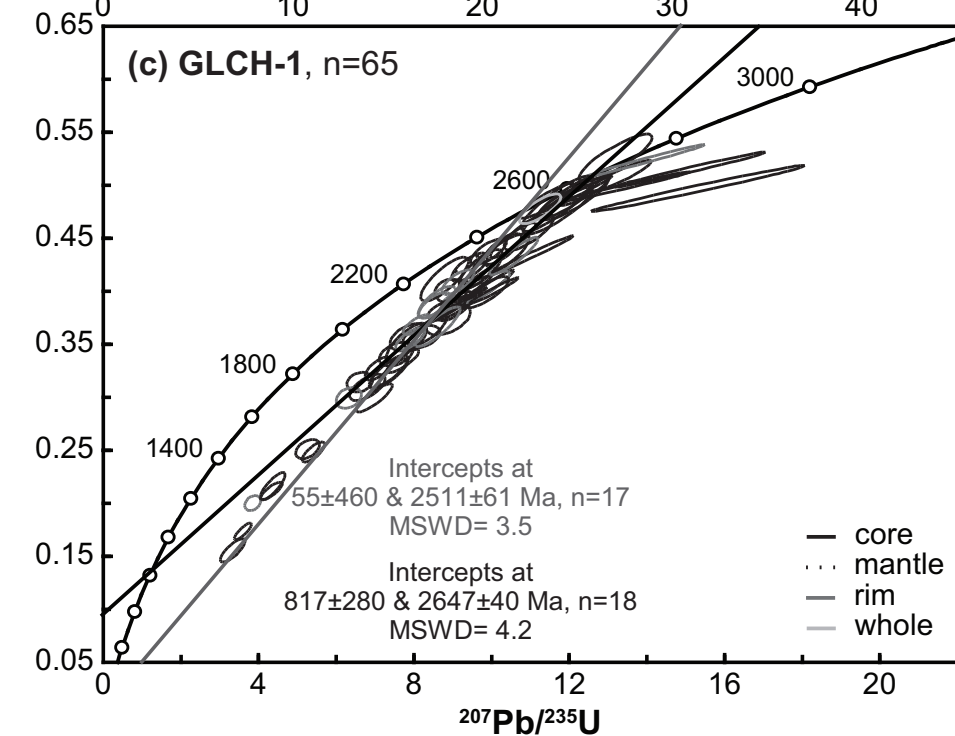
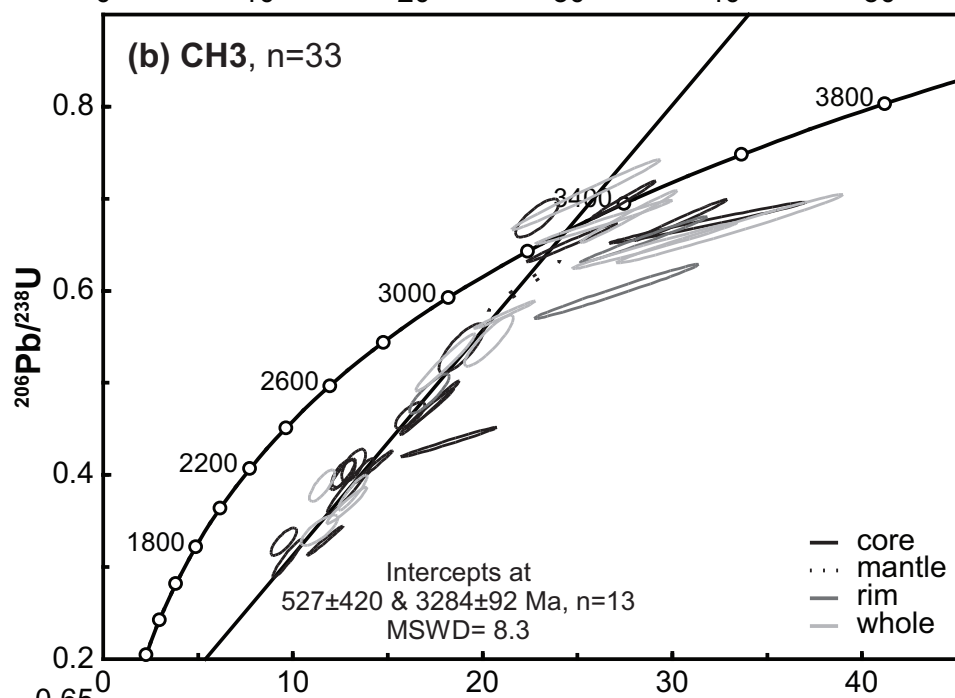
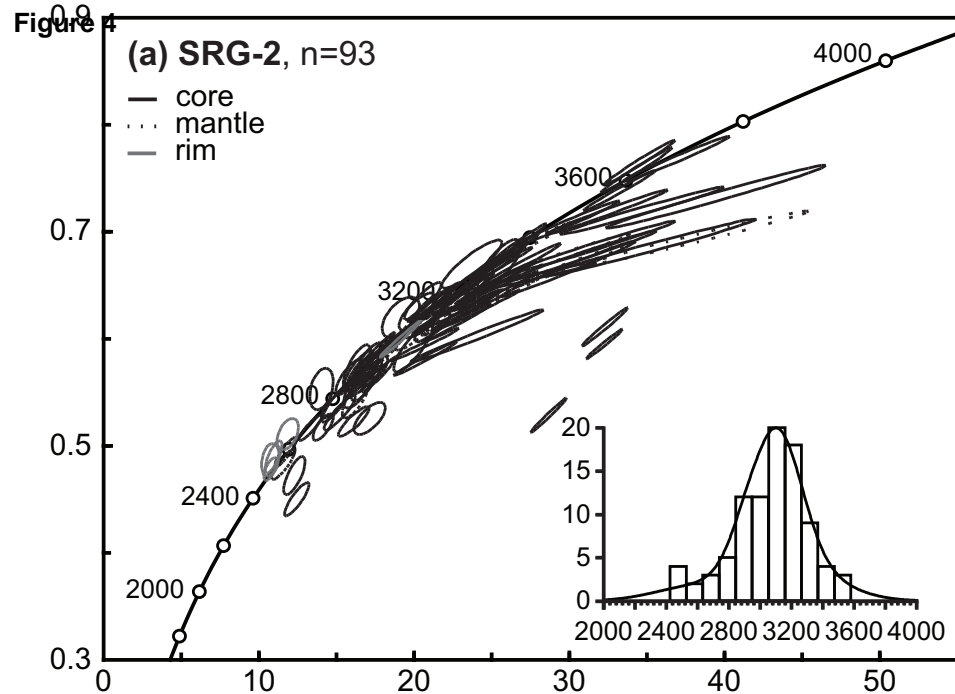


Figure 5

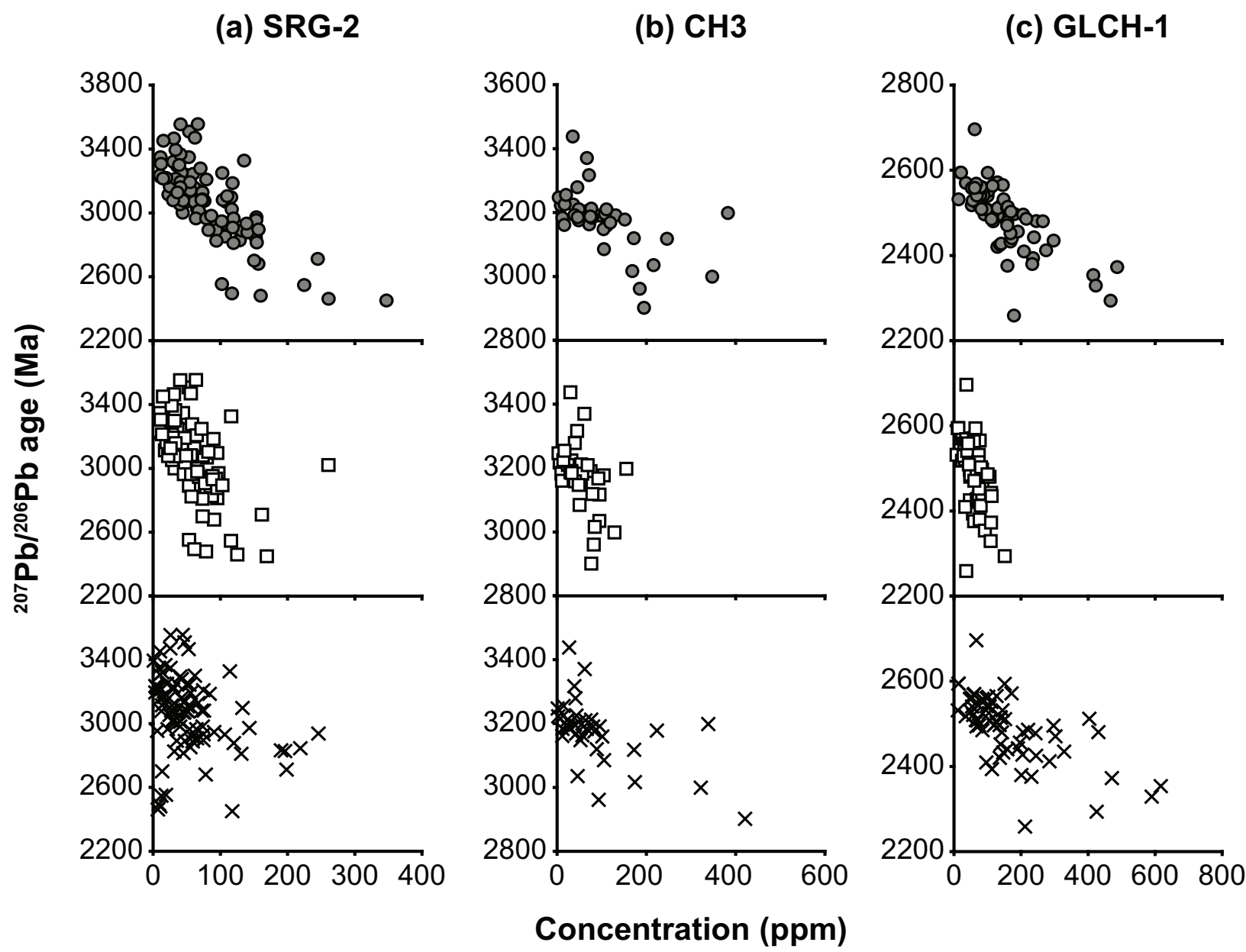


Figure 6

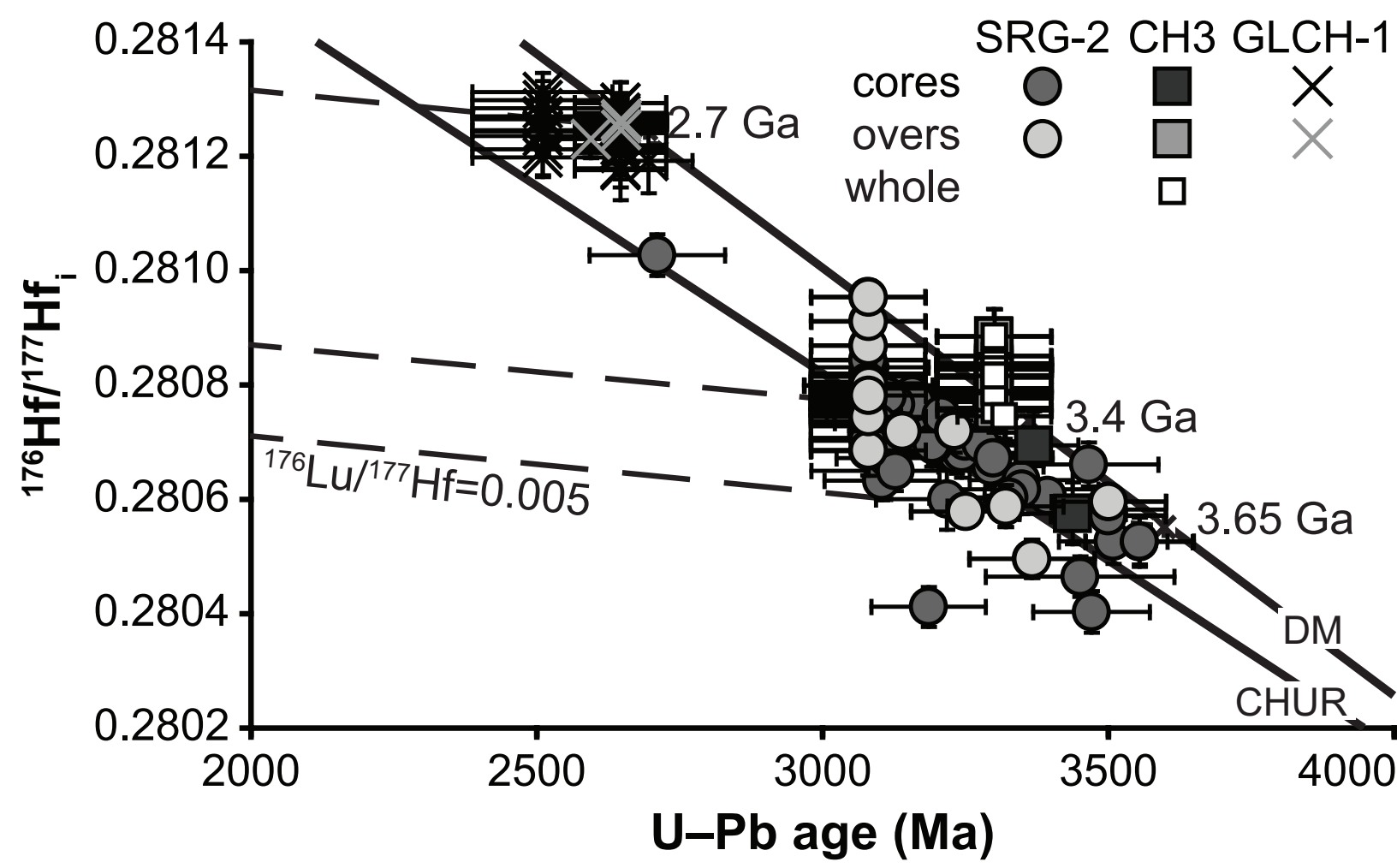


Figure 7

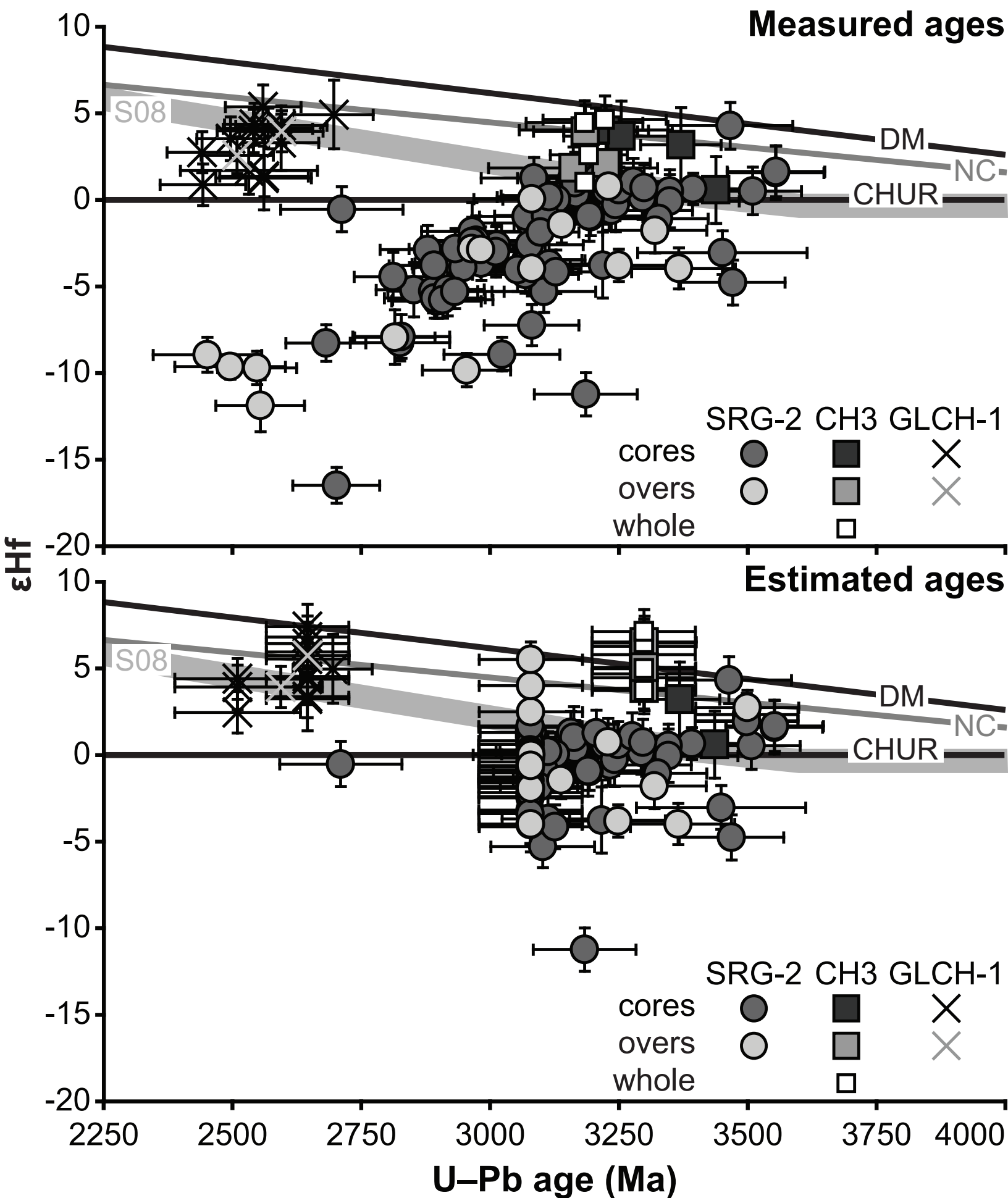


Figure 8

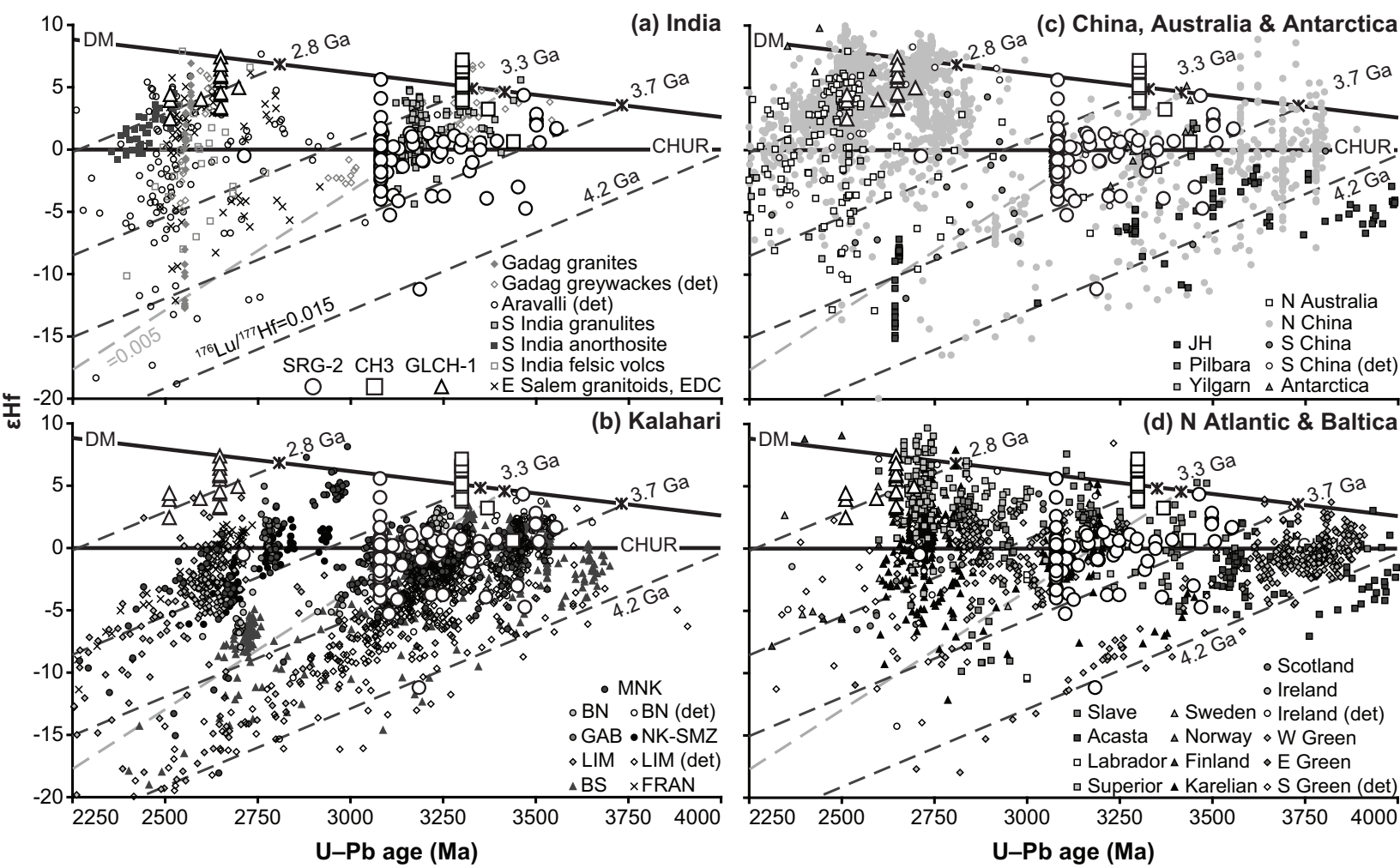


Table 1

Simplified Precambrian stratigraphy of the Dharwar craton. The sampled horizons are marked in red.

Western Dharwar Craton (WDC)		Eastern Dharwar Craton (EDC)
Badami Group (0.6 Ga?) ----- <i>Unconformity</i> -----	Kaladgi Supergroup (~Cuddapah Supergroup)	Kurnool Group, Bhima Group (0.6 Ga?) ----- <i>Unconformity</i> -----
Bagalkot Group ----- <i>Unconformity</i> -----		Cuddapah Supergroup (1.9-1.4 Ga?) ----- <i>Unconformity</i> -----
Younger granitoids (~2.6 Ga)		
Chitradurga Group (polymictic conglomerate, quartzite, greywacke, pelite, BIF, basalt)	Dharwar Supergroup (2.9-2.6 Ga)	TTG gneisses, granites (2.7-2.5 Ga)
Bababudan Group (oligomictic conglomerate, quartzite, basalt, BIF)		Kolar Group (~2.7 Ga) (basalt, komatiite, chert, BIF, pelite, felsic volcanics)
----- <i>Unconformity</i> -----		
Sargur Group (3.3-3.0 Ga?) (quartzite, pelite, calc- silicate, basalt, koma- tiite, BIF, layered mafic- ultramafics)	Granitic Gneisses (3.35-3.0 Ga)	----- Vestiges of older Gneisses (~ 3.3-3.1 Ga) and supracrustals (~ 3.3 Ga?)

*Highlights (for review)

- Terrane accumulation in the Dharwar craton began ~3 Ga in older Western block
- Final cratonisation at 2.5 Ga is much later than other Archaean cratons
- Tectonic, not genetic, links with the Kaapvaal and north China cratons