Teukolsky-Starobinsky identities for arbitrary spin

E. G. Kalnins

Department of Mathematics and Statistics, University of Waikato, Hamilton, New Zealand

W Miller Jr

School of Mathematics, University of Minnesota, Minneapolis, Minnesota 55455

G. C. Williams

Department of Mathematics and Statistics, University of Waikato, Hamilton, New Zealand

(Received 18 October 1988; accepted for publication 21 June 1989)

The Teukolsky-Starobinsky identities are proven for arbitrary spin s. A pair of covariant equations are given that admit solutions in terms of Teukolsky functions for general s. The method of proof is shown to extend to the general class of space-times considered by Torres del Castillo [J. Math. Phys. 29, 2078 (1988)].

I. INTRODUCTION

Gravitational and electromagnetic perturbations in Kerr geometry are known to be intimately connected to Teukolsky functions. 1 This came about because of investigations by Teukolsky who showed that in the Newman-Penrose formalism² separable solutions were possible for certain Maxwell and Weyl scalars in Kerr geometry. The resulting separable solutions are known as Teukolsky functions. In addition to the problem of gravitational and electromagnetic perturbations these functions reappear when the neutrino³ and Rarita-Schwinger4 equations are solved in a background of Kerr geometry. These functions satisfy what are known as the Teukolsky-Starobinsky identities. In this work we prove these identities for any spin s. This result is established relatively easily. One of the difficulties with the Kerr metric is that for s > 2 these functions do not appear to come from any covariant equation. We rectify this situation by introducing covariant equations that admit Teukolsky functions for general s as their solutions. No claim is made that these equations have physical significance. Finally we note that the method of proof applies to the more general class of space-time studied by Torres del Castillo⁵ who proved these results for $s \le 2$.

II. THE TEUKOLSKY-STAROBINSKY IDENTITIES

We consistently use in this article the spinor notation of Penrose and Rindler⁶ and the null tetrad formalism of Chandrasekhar.⁷ Specifically we restrict ourselves to the Kinnersley null tetrad of vectors with components

$$l^{a} = (1/\sqrt{2}\Delta)(r^{2} + a^{2}, \Delta, 0, a),$$

$$n^{a} = (1/\sqrt{2}\tilde{\rho}\tilde{\rho}^{*})(r^{2} + a^{2}, -\Delta, 0, a),$$

$$m^{a} = (1/\sqrt{2}\tilde{\rho})(ia\sin\theta, 0, 1, i\csc\theta)$$

$$\overline{m}^{a} = (1/\sqrt{2}\tilde{\rho}^{*})(-ia\sin\theta, 0, 1, -i\csc\theta),$$
ere

where

$$\Delta = r^2 - 2Mr + a^2$$
, $\rho^2 = r^2 + a^2 \cos^2 \theta$

and

2925

$$ilde{
ho}=r+ia\cos heta.$$
 J. Math. Phys. **30** (12), December 1989

The Kerr solution has the line element

$$ds^{2} = \left(1 - \frac{2Mr}{\tilde{\rho}\tilde{\rho}^{*}}\right)dt^{2} - \frac{\tilde{\rho}\tilde{\rho}^{*}}{\Delta}dr^{2} - \tilde{\rho}\tilde{\rho}^{*}d\theta^{2}$$
$$-\left((r^{2} + a^{2}) + \frac{2a^{2}Mr\sin^{2}\theta}{\tilde{\rho}\tilde{\rho}^{*}}\right)\sin^{2}\theta d\phi^{2}$$
$$+ \frac{4aMr\sin^{2}\theta}{\tilde{\rho}\tilde{\rho}^{*}}dt d\phi. \tag{3}$$

The differential operators \mathcal{D}_n , \mathcal{D}_n^{\dagger} , \mathcal{L}_n , and \mathcal{L}_n^{\dagger} are defined as

$$\mathcal{D}_{n} = \partial_{r} + iK/\Delta + 2n[(r - M)/\Delta],$$

$$\mathcal{D}_{n}^{\dagger} = \partial_{r} - iK/\Delta + 2n[(r - M)/\Delta],$$

$$\mathcal{L}_{n} = \partial_{\theta} + Q + n \cot \theta,$$

$$\mathcal{L}_{n}^{\dagger} = \partial_{\theta} - Q + n \cot \theta,$$
(4)

where

$$K = (r^2 + a^2)\sigma + am$$
 and $Q = a\sigma \sin \theta + m \csc \theta$. (5)

Teukolsky functions P_{+} , and P_{-} , in the variable r satisfy

$$(\Delta \mathcal{D}_{1-s} \mathcal{D}_{0}^{\dagger} - 2(2|s|-1)i\sigma r)P_{+s} = \lambda P_{+s},$$

$$(\Delta \mathcal{D}_{1-s}^{\dagger} \mathcal{D}_{0} + 2(2|s|-1)i\sigma r)P_{-s} = \lambda P_{-s},$$
(6)

where $s = \frac{1}{2}, 1, \dots$. The first result proven is the following theorem.

Theorem 1: If $s = \frac{1}{2}, 1, ...$ then

$$\Delta^{s} \mathcal{D}_{0}^{2s} \left[\Delta \mathcal{D}_{1-s}^{\dagger} \mathcal{D}_{0} + 2(2s-1)i\sigma r \right]$$

$$= \left[\Delta \mathcal{D}_{1-s} \mathcal{D}_{0}^{\dagger} - 2(2s-1)i\sigma r \right] \Delta^{s} \mathcal{D}_{0}^{2s}. \tag{7}$$

Proof: By induction on s. Noting that for $s = \frac{1}{2}$

$$\Delta^{1/2} \mathcal{D}_0(\Delta \mathcal{D}_{1/2}^{\dagger} \mathcal{D}_0) = (\Delta \mathcal{D}_{1/2} \mathcal{D}_0^{\dagger}) \Delta^{1/2} \mathcal{D}_0. \tag{8}$$

If we now assume the result is true for a given s then

$$\begin{split} &\Delta^{s+1/2} \mathcal{D}_{0}^{2(s+1/2)} \left[\Delta \mathcal{D}_{1-(s+1/2)}^{\dagger} \mathcal{D}_{0} + 2(2(s+\frac{1}{2})-1)i\sigma r \right] \\ &= \Delta^{s+1/2} \mathcal{D}_{0}^{2s+1} \left[\Delta (\mathcal{D}_{1-s}^{\dagger} - (r-M)/\Delta) \mathcal{D}_{0} + 4si\sigma r \right] \\ &= \Delta^{1/2} \mathcal{D}_{-s} \left[\Delta^{s} \mathcal{D}_{0}^{2s} (\Delta \mathcal{D}_{1-s}^{\dagger} \mathcal{D}_{0} + 2(2s-1)i\sigma r) + \Delta^{s} \mathcal{D}_{0}^{2s} (2i\sigma r - (r-M) \mathcal{D}_{0}) \right] \end{split}$$

(2) and

0022-2488/89/122925-05\$02.50 © 1989 American Institute of Physics

2925

$$\begin{split} & \left[\Delta \mathcal{D}_{1-(s+1/2)} \, \mathcal{D}_{0}^{\dagger} - 2(2(s+\frac{1}{2})-1)i\sigma r \right] \Delta^{s+1/2} \, \mathcal{D}_{0}^{2(s+1/2)} \\ & = \Delta^{1/2} (\Delta \mathcal{D}_{1-s} \, \mathcal{D}_{1/2}^{\dagger} - 4si\sigma r) \mathcal{D}_{-s} \Delta^{s} \, \mathcal{D}_{0}^{2s} \\ & = \Delta^{1/2} \mathcal{D}_{-s} (\Delta \, \mathcal{D}_{1/2}^{\dagger} \, \mathcal{D}_{-s} - 4si\sigma r) \Delta^{s} \, \mathcal{D}_{0}^{2s} + 4si\sigma \Delta^{s+1/2} \, \mathcal{D}_{0}^{2s} \\ & = \Delta^{1/2} \mathcal{D}_{-s} \left[\Delta \left(\mathcal{D}_{1-s} + (2s-1) \frac{r-M}{\Delta} - \frac{2iK}{\Delta} \right) \left(\mathcal{D}_{0}^{\dagger} - 2s \frac{r-M}{\Delta} + \frac{2iK}{\Delta} \right) - 4si\sigma r \right] \Delta^{s} \, \mathcal{D}_{0}^{2s} + 4si\sigma \Delta^{s+1/2} \, \mathcal{D}_{0}^{2s} \\ & = \Delta^{1/2} \mathcal{D}_{-s} \left[(\Delta \mathcal{D}_{1-s} \, \mathcal{D}_{0}^{\dagger} - 2(2s-1)i\sigma r) \Delta^{s} \, \mathcal{D}_{0}^{2s} + (2i\sigma r - 2s - (r-M) \mathcal{D}_{-s}) \Delta^{s} \, \mathcal{D}_{0}^{2s} + 4si\sigma \Delta^{s} \, \mathcal{D}_{0}^{2s-1} \right]. \end{split}$$

Note that

$$\Delta^{s} \mathcal{D}_{0}^{2s} (2i\sigma r - (r - M)\mathcal{D}_{0}) = (2i\sigma r - (r - M)\mathcal{D}_{-s})\Delta^{s} \mathcal{D}_{0}^{2s} + \Delta^{s} (4si\sigma \mathcal{D}_{0}^{2s-1} - 2s \mathcal{D}_{0}^{2s})$$

$$= (2i\sigma r - 2s - (r - M)\mathcal{D}_{-s})\Delta^{s} \mathcal{D}_{0}^{2s} + 4si\sigma \Delta^{s} \mathcal{D}_{0}^{2s-1}. \tag{11}$$

Thus subtracting (10) from (9) and making use of (11) we have

$$\Delta^{s+1/2} \mathcal{D}_{0}^{2(s+1/2)} \left[\Delta \mathcal{D}_{1-(s+1/2)}^{\dagger} \mathcal{D}_{0} + 2(2(s+\frac{1}{2})-1)i\sigma r \right]
- \left[\Delta \mathcal{D}_{1-(s+1/2)} \mathcal{D}_{0}^{\dagger} - 2(2(s+\frac{1}{2})-1)i\sigma r \right] \Delta^{s+1/2} \mathcal{D}_{0}^{2(s+1/2)}
= \Delta^{1/2} \mathcal{D}_{-s} \left[\Delta^{s} \mathcal{D}_{0}^{2s} \left(\Delta \mathcal{D}_{1-s}^{\dagger} \mathcal{D}_{0} + 2(2s-1)i\sigma r \right) - \left(\Delta \mathcal{D}_{1-s} \mathcal{D}_{0}^{\dagger} - 2(2s-1)i\sigma r \right) \Delta^{s} \mathcal{D}_{0}^{2s} \right].$$
(12)

A direct consequence of this result is that $\Delta \mathcal{D}_0^{2s} P_{-s}$ is a solution of the Teukolsky equation for P_{+s} . Similarly it may be proven that

$$\Delta^{s} \mathcal{D}_{0}^{\dagger 2s} \left[\Delta \mathcal{D}_{1-s} \mathcal{D}_{0}^{\dagger} - 2(2s-1)i\sigma r \right] = \left[\Delta \mathcal{D}_{1-s}^{\dagger} \mathcal{D}_{0} + 2(2s-1)i\sigma r \right] \Delta^{s} \mathcal{D}_{0}^{\dagger 2s}, \tag{13}$$

i.e., $\Delta^s \mathcal{D}_0^{\dagger 2s} P_{+s}$ is a solution of the Teukolsky equation for P_{-s} . By suitable choice of the relative normalization of the functions we can write the following results:

$$\Delta^{s} \mathcal{D}_{0}^{2s} P_{-s} = D_{s} P_{+s}, \Delta^{s} \mathcal{D}_{0}^{2s} P_{+s} = D_{s}^{*} P_{-s},$$
 (14)

where D_s is some complex constant. These are the Teukolsky-Starobinsky identities known to be true for $s = \frac{1}{2}$, 1, $\frac{3}{2}$, 2. For the variable θ we can prove a similar result.

Theorem 2: If $s = \frac{1}{2}$, 1,... then

$$\mathcal{L}_{1-s}\mathcal{L}_{2-s}\cdots\mathcal{L}_{s-1}\mathcal{L}_{s}\left[\mathcal{L}_{1-s}^{\dagger}\mathcal{L}_{s}+2(2s-1)\sigma a\cos\theta\right]$$

$$=\left[\mathcal{L}_{1-s}\mathcal{L}_{s}^{\dagger}-2(2s-1)\sigma a\cos\theta\right]\mathcal{L}_{1-s}\mathcal{L}_{2-s}\cdots\mathcal{L}_{s-1}\mathcal{L}_{s}.$$
(15)

Proof: Again using induction we note that for $s = \frac{1}{2}$

$$\mathcal{L}_{1/2}(\mathcal{L}_{1/2}^{\dagger} \mathcal{L}_{1/2}) = (\mathcal{L}_{1/2} \mathcal{L}_{1/2}^{\dagger}) \mathcal{L}_{1/2}. \tag{16}$$

Then

$$\mathcal{L}_{1-(s+1/2)}\mathcal{L}_{2-(s+1/2)}\cdots\mathcal{L}_{(s+1/2)-1}\mathcal{L}_{s+1/2}\left[\mathcal{L}_{1-s}^{\dagger}\mathcal{L}_{s}+2(2(s+\frac{1}{2})-1)\sigma a\cos\theta\right]$$

$$=(1/\sqrt{\sin\theta})\mathcal{L}_{-s}\mathcal{L}_{1-s}\cdots\mathcal{L}_{s-1}\mathcal{L}_{s}(\mathcal{L}_{s}^{\dagger}\mathcal{L}_{s}+4s\sigma a\cos\theta)\sqrt{\sin\theta}$$

$$=(1/\sqrt{\sin\theta})\mathcal{L}_{-s}\mathcal{L}_{1-s}\cdots\mathcal{L}_{s-1}\mathcal{L}_{s}[\mathcal{L}_{1-s}^{\dagger}\mathcal{L}_{s}+2(2s-1)\sigma a\cos\theta-\cot\theta\mathcal{L}_{s}+2\sigma a\cos\theta]\sqrt{\sin\theta}$$
(17)

and

$$\left[\mathcal{L}_{1-(s+1/2)} \mathcal{L}_{s+1/2}^{\dagger} - 2(2(s+\frac{1}{2})-1)\sigma a \cos\theta \right] \mathcal{L}_{1-(s+1/2)} \mathcal{L}_{2-(s+1/2)} \cdots \mathcal{L}_{(s+1/2)-1} \mathcal{L}_{(s+1/2)}
= (1/\sqrt{\sin\theta}) (\mathcal{L}_{-s} \mathcal{L}_{s}^{\dagger} - 4s\sigma a \cos\theta) \mathcal{L}_{-s} \mathcal{L}_{1-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} \sqrt{\sin\theta}
= (1/\sqrt{\sin\theta}) \mathcal{L}_{-s} (\mathcal{L}_{s}^{\dagger} \mathcal{L}_{-s} - 4s\sigma a \cos\theta) \mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} \sqrt{\sin\theta}
- 4s\sigma a \sqrt{\sin\theta} \mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} \sqrt{\sin\theta}.$$
(18)

Now note that we can write

$$\mathcal{L}_{s}^{\dagger} \mathcal{L}_{-s} - 4s\sigma a \cos \theta = (\mathcal{L}_{1-s} + (2s-1)\cot \theta - 2Q)(\mathcal{L}_{s}^{\dagger} - 2s\cot \theta + 2Q) - 4s\sigma a \cos \theta$$

$$= \mathcal{L}_{1-s} \mathcal{L}_{s}^{\dagger} - 2(2s-1)\sigma a \cos \theta - \cot \theta \mathcal{L}_{s}^{\dagger} + 2s\csc^{2}\theta - 2m\cot \theta \csc \theta \tag{19}$$

observing the identities

$$\mathcal{L}_a \mathcal{L}_{a+1} \cdots \mathcal{L}_{b-1} \mathcal{L}_b \cos \theta = \cos \theta \mathcal{L}_a \mathcal{L}_{a+1} \cdots \mathcal{L}_{b-1} \mathcal{L}_b - (b-a+1) \sin \theta \mathcal{L}_{a+1} \mathcal{L}_{a+2} \cdots \mathcal{L}_{b-1} \mathcal{L}_b \quad (20)$$
 and

$$\mathcal{L}_{a}\mathcal{L}_{a+1}\cdots\mathcal{L}_{b-1}\mathcal{L}_{b}\cot\theta = \cot\theta\mathcal{L}_{a-1}\mathcal{L}_{a}\cdots\mathcal{L}_{b-2}\mathcal{L}_{b-1} - (b-a+1)\mathcal{L}_{a}\mathcal{L}_{a+1}\cdots\mathcal{L}_{b-2}\mathcal{L}_{b-1}$$
 (21) and noting that

2926 J. Math. Phys., Vol. 30, No. 12, December 1989

Kalnins, Miller, Jr., and Williams 2926

$$(1/\sqrt{\sin\theta}) \mathcal{L}_{-s} \left[\mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} (-\cot\theta \mathcal{L}_{s} + 2\sigma a \cos\theta) \right.$$

$$- \left(-\cot\theta \mathcal{L}_{s}^{\dagger} + 2s \csc^{2}\theta - 2m \cot\theta \csc\theta \right) \mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} \right] \sqrt{\sin\theta}$$

$$+ 4s\sigma a \sqrt{\sin\theta} \mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} \sqrt{\sin\theta}$$

$$= (1/\sqrt{\sin\theta}) \mathcal{L}_{-s} \left[-\cot\theta \mathcal{L}_{-s} \mathcal{L}_{1-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} + 2s \mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} \right.$$

$$+ 2\sigma a \cos\theta \mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} - 4s\sigma a \sin\theta \mathcal{L}_{2-s} \mathcal{L}_{3-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s}$$

$$- (-\cot\theta \mathcal{L}_{s} + 2Q \cot\theta + 2s \csc^{2}\theta - 2m \cot\theta \csc\theta) \mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} \right] \sqrt{\sin\theta}$$

$$+ 4s\sigma a \sqrt{\sin\theta} \mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} \sqrt{\sin\theta}$$

$$= (1/\sqrt{\sin\theta}) \mathcal{L}_{-s} \left[-\cot\theta \mathcal{L}_{-s} + 2s + 2\sigma a \cos\theta + \cot\theta \mathcal{L}_{s} - 2Q \cot\theta - 2s \csc^{2}\theta \right.$$

$$+ 2m \cot\theta \csc\theta \right] \mathcal{L}_{1-s} \mathcal{L}_{2-s} \cdots \mathcal{L}_{s-1} \mathcal{L}_{s} \sqrt{\sin\theta}$$

$$= 0, \tag{22}$$

we have established that

$$\mathcal{L}_{1-(s+1/2)}\mathcal{L}_{2-(s+1/2)}\cdots\mathcal{L}_{(s+1/2)-1}\mathcal{L}_{s+1/2}\left[\mathcal{L}_{1-s}^{\dagger}\mathcal{L}_{s}+2(2(s+\frac{1}{2})-1)\sigma a\cos\theta\right]$$

$$-\left[\mathcal{L}_{1-(s+1/2)}\mathcal{L}_{s+1/2}^{\dagger}-2(2(s+\frac{1}{2})-1)\sigma a\cos\theta\right]\mathcal{L}_{1-(s+1/2)}\mathcal{L}_{2-(s+1/2)}\cdots\mathcal{L}_{(s+1/2)-1}\mathcal{L}_{(s+1/2)}$$

$$=(1/\sqrt{\sin\theta})\mathcal{L}_{-s}\left[\mathcal{L}_{1-s}\mathcal{L}_{2-s}\cdots\mathcal{L}_{s-1}\mathcal{L}_{s}(\mathcal{L}_{1-s}^{\dagger}\mathcal{L}_{s}+2(2s-1)\sigma a\cos\theta\right]$$

$$-(\mathcal{L}_{1-s}\mathcal{L}_{s}^{\dagger}-2(2s-1)\sigma a\cos\theta)\mathcal{L}_{1-s}\mathcal{L}_{2-s}\cdots\mathcal{L}_{s-1}\mathcal{L}_{s}]\sqrt{\sin\theta}$$
(23)

and the result is proven. In this case the Teukolsky equations are defined as

$$(\mathcal{L}_{1-s}^{\dagger} \mathcal{L}_{s} + 2(2s-1)\sigma a \cos \theta)S_{+s} = -\lambda S_{+s},$$

$$(\mathcal{L}_{1-s} \mathcal{L}_{s}^{\dagger} - 2(2s-1)\sigma a \cos \theta)S_{-s} = -\lambda S_{-s}.$$
(24)

Consequently Theorem 2 tells us that we can upon suitable renormalization, find a constant C, such that

$$\mathcal{L}_{1-s}\mathcal{L}_{2-s}\cdots\mathcal{L}_{s-1}\mathcal{L}_{s}S_{+s}=C_{s}S_{-s}.$$
 (25)

Similarly one may prove the identity

$$\mathcal{L}_{1-s}^{\dagger} \mathcal{L}_{2-s}^{\dagger} \cdots \mathcal{L}_{s-1}^{\dagger} \mathcal{L}_{s}^{\dagger} \times \left[\mathcal{L}_{1-s} \mathcal{L}_{s}^{\dagger} - 2(2s-1)\sigma a \cos \theta \right]$$

$$= \left[\mathcal{L}_{1-s}^{\dagger} \mathcal{L}_{s} + 2(2s-1)\sigma a \cos \theta \right]$$

$$\times \mathcal{L}_{1-s}^{\dagger} \mathcal{L}_{2-s}^{\dagger} \cdots \mathcal{L}_{s-1}^{\dagger} \mathcal{L}_{s}^{\dagger}$$
(26)

from which we can write

$$\mathcal{L}_{1-s}^{\dagger} \mathcal{L}_{2-s}^{\dagger} \cdots \mathcal{L}_{s-1}^{\dagger} \mathcal{L}_{s}^{\dagger} S_{-s} = C_{s} S_{+s}. \tag{27}$$

These are the Teukolsky-Starobinsky identities known to be true for $s = \frac{1}{3}, \frac{1}{3}, \frac{2}{3}$. The question we now ask is what if any significance do the Teukolsky functions have for general s. Before giving a covariant equation that works for general s let us recapitulate how things work in the case of the Rarita-Schwinger field. The Rarita-Schwinger equation written in spinor notation is

$$\nabla^{AA'}F_{AB}{}^{B'}=0, \tag{28}$$

where $F_{ABB'} = F_{(AB)B'}$. We can construct a coupled system of equations as follows. Let

$$h_{ABC} = \nabla_{(AA'} F_{BC)}^{A'}. \tag{29}$$

Then h_{ABC} satisfies a first-order equation as follows:

$$\nabla^{AA'} h_{ABC} = \nabla^{AA'} \nabla_{(AB'} F_{BC)}^{B'}$$

$$= \frac{1}{3} \nabla^{AA'} \nabla_{AB'} F_{BC}^{B'} + \frac{2}{3} \nabla^{AA'} \nabla_{(BB'} F_{C)A}^{B'}. \tag{30}$$

2927 J. Math. Phys., Vol. 30, No. 12, December 1989 Using the Rarita-Schwinger equation and the symmetry in the indices B and C we write

$$\frac{1}{3}\nabla^{AA'}\nabla_{AB'}F_{BC}^{B'} = \frac{1}{3}\nabla^{AA'}\nabla_{BB'}F_{AC}^{B'} + \frac{1}{3}\nabla_{B}^{A'}\nabla_{AB'}F_{C}^{A'}^{B'}
= \frac{1}{3}\nabla^{AA'}\nabla_{(BB'}F_{C)A}^{B'}.$$
(31)

Consequently (30) becomes

 $(\delta^* + 2\alpha + \beta^* - \tau^*)F_{11}^{0}$

$$\nabla^{AA'}h_{ABC} = \nabla^{AA'}\nabla_{(BB'}F_{C)A}^{B'}$$

$$= \nabla_{(BB'}\nabla^{AA'}F_{C)A}^{B'} + [\nabla^{AA'},\nabla_{(BB')}]F_{C)A}^{B'}$$

$$= [\nabla^{AA'},\nabla_{(BB')}]F_{C)A}^{B'}$$

$$= -\epsilon^{A'}_{B'}\Psi^{A}_{BC}^{M}F_{MA}^{B'}$$

$$= \Psi_{BC}^{AM}F_{AM}^{A'}. \tag{32}$$

The pair of equations (29), (32) when written in Newman-Penrose notation become

$$(D-\rho)h_{111} - (\delta^* + 3\pi + \alpha)h_{110} = -\Psi_2 F_{110}, \tag{33}$$

$$(D-2\rho)h_{110} - (\delta^* + 2\pi - \alpha)h_{100} = \Psi_2 F_{100}, \tag{34}$$

$$(D-3\rho)h_{100}-(\delta^*+\pi-3\alpha)h_{000}=-\Psi_2F_{000},\qquad(35)$$

$$(\delta + 3\beta - \tau)h_{111} - (\Delta + \gamma + 3\mu)h_{110} = -\Psi_2 F_{111}, \quad (36)$$

$$(\delta + \beta - 2\tau)h_{110} - (\Delta - \gamma + 2\mu)h_{100} = \Psi_2 F_{101}, \quad (37)$$

$$(\delta - \beta - 3\tau)h_{100} - (\Delta - 3\gamma + \mu)h_{000} = -\Psi_2 F_{001},$$
 (38) and

$$(D - \rho^*)F_{00}^{0'} + (\delta - \alpha^* - 2\beta + \pi^*)F_{00}^{1'} = h_{000}, \quad (39)$$

$$+ (\Delta + 2\gamma - \gamma^* + \mu^*) F_{11}^{1'} = h_{111}, \tag{40}$$

$$2[(D - \rho^* + \rho)F_{10}^{0'} + (\delta + \pi^* - \alpha^* + \tau)F_{10}^{1'}] + [(\delta^* + \beta^* - 2\alpha - \tau^* - 2\pi)F_{00}^{0'} + (\triangle + \mu^* - \gamma^* - 2\gamma - 2\mu)F_{00}^{1'}] = 3h_{100},$$
 (41)

Kalnins, Miller, Jr., and Williams

2927

$$2[(\delta^* + \beta^* - \tau^* - \pi)F_{10}^{0'} + (\Delta - \gamma^* + \mu^* - \mu)F_{10}^{1'}] + [(D - \rho^* + 2\rho)F_{11}^{0'} + (\delta - \alpha^* + 2\beta + \pi^* + 2\tau)F_{11}^{1'}] = 3h_{110}.$$
(42)

Considering Eqs. (35), (38), and (39) and putting

$$h_{100} = (1/\tilde{\rho}^*)H_1, \quad h_{000} = H_0,$$

$$F_{000} = (1/\sqrt{2}\tilde{\rho}^*)G_{000'}, \quad F_{001'} = (1/\sqrt{2}\rho^2)G_{001'},$$
(43)

$$\left(\mathcal{L}_{-1/2}^{\dagger} + \frac{2ia\sin\theta}{\tilde{\rho}^{*}}\right)H_{1} + \Delta\left(\mathcal{D}_{3/2}^{\dagger} - \frac{2}{\tilde{\rho}^{*}}\right)H_{0}
= -\Psi_{2}G_{001'},
\left(\mathcal{D}_{0} + \frac{2}{\tilde{\rho}^{*}}\right)H_{1} - \left(\mathcal{L}_{3/2} - \frac{2ia\sin\theta}{\tilde{\rho}^{*}}\right)H_{0} = -\Psi_{2}G_{000'},
\left(\mathcal{D}_{0} - \frac{1}{\tilde{\rho}^{*}}\right)G_{001'} - \left(\mathcal{L}_{-1/2}^{\dagger} - \frac{ia\sin\theta}{\tilde{\rho}^{*}}\right)G_{000'} = 2\rho^{2}H_{0}.$$
(44)

These equations imply that H_0 satisfies the separable equa-

$$(\Delta \mathcal{D}_1 \, \mathcal{D}_{3/2}^{\dagger} + \mathcal{L}_{-1/2}^{\dagger} \, \mathcal{L}_{3/2} - 4i\sigma\tilde{\rho}) H_0 = 0$$
 (45) admitting solutions $H_0 = \Delta^{-3/2} P_{+3/2} S_{+3/2}$. Similarly if Eqs. (33), (36), and (40) are considered then putting

$$h_{110} = (1/\tilde{\rho}^{*2})H_2, \quad h_{111} = (1/\tilde{\rho}^{*3})H_3,$$

$$F_{110'} = (1/\sqrt{2}\tilde{\rho}^{*3})G_{110'}, \quad F_{111'} = (1/\sqrt{2}\rho^2\tilde{\rho}^{*2})G_{111'},$$
(46)

we obtain

$$\left(\mathcal{Q}_{0} - \frac{2}{\tilde{\rho}^{*}}\right) H_{3} - \left(\mathcal{L}_{-1/2} + \frac{2ia\sin\theta}{\tilde{\rho}^{*}}\right) H_{2} = -\Psi_{2} G_{110'},
\left(\mathcal{L}_{3/2}^{\dagger} - \frac{2ia\sin\theta}{\tilde{\rho}^{*}}\right) H_{3} + \Delta \left(\mathcal{Q}_{-1/2}^{\dagger} + \frac{2}{\tilde{\rho}^{*}}\right) H_{2} = -\Psi_{2} G_{111'},
\left(\mathcal{L}_{-1/2} - \frac{ia\sin\theta}{\tilde{\rho}^{*}}\right) G_{111'} + \Delta \left(\mathcal{Q}_{-1/2}^{\dagger} - \frac{1}{\tilde{\rho}^{*}}\right) G_{110'} = 2\rho^{2} H_{3}.$$
(47)

The functions H_3 satisfies the separable equation

$$(\Delta \mathcal{D}_{-1/2}^{\dagger} \mathcal{D}_{0} + \mathcal{L}_{-1/2} \mathcal{L}_{3/2}^{\dagger} + 4i\sigma\tilde{\rho}) H_{3} = 0 \quad (48)$$

admitting solutions $H_3 = P_{-3/2}S_{-3/2}$. Two solutions to Eqs. (44) and (47) can be found

(1)
$$h_{000} = \Delta^{-3/2} P_{+3/2} S_{+3/2},$$

$$F_{000'} = \frac{1}{\sqrt{2} \tilde{\rho}^* \Psi_2} \left(\mathcal{L}_{3/2} - \frac{2ia \sin \theta}{\tilde{\rho}^*} \right) \Delta^{-3/2} P_{+3/2} S_{+3/2},$$

$$F_{001'} = \frac{-1}{\sqrt{2} \rho^2 \Psi_2} \Delta \left(\mathcal{D}_{3/2}^{\dagger} - \frac{2}{\tilde{\rho}^*} \right) \Delta^{-3/2} P_{+3/2} S_{+3/2},$$

and

2928

(2)
$$h_{111} = \frac{1}{\tilde{\rho}^{*3}} P_{-3/2} S_{-3/2},$$

$$F_{110'} = \frac{-1}{\sqrt{2} \tilde{\rho}^{*3} \Psi_2} \left(\mathcal{Q}_0 - \frac{2}{\tilde{\rho}^*} \right) P_{-3/2} S_{-3/2},$$

$$F_{111'} = \frac{-\Delta}{\sqrt{2} \rho^2 \tilde{\rho}^{*2} \Psi_2} \left(\mathcal{L}_{3/2}^{\dagger} - \frac{2ia \sin \theta}{\tilde{\rho}^*} \right) P_{-3/2} S_{-3/2}.$$

In the case of each solution we have given only the nonzero components. It is interesting to note that the spinor F_{ABA} satisfies the equation

$$\nabla^{A}_{C'}\nabla_{(AA'}F_{BC)}^{A'} = \Psi^{AM}_{BC}F_{AMC'}.$$
 (50)

However, the above choices do not satisfy the Rarita-Schwinger equation (28). It is possible to extend these equations to a set which has solutions in terms of Teukolsky functions for general s. If we consider the equations

$$\phi \nabla_{(A_1 A_1} F^{A_1'}_{A_2 \cdots A_{2s})} - \frac{1}{6} (2s - 3) (\nabla_{(A_1 A_1'} \phi) F^{A_1'}_{A_2 \cdots A_{2s}})
= \phi h_{A_1 \cdots A_{2s}},$$
(51)

$$\nabla^{AA'}h_{AA,\cdots A_{2s}} = (2s-1)(s-1)\Psi_{(A_2A_3)}{}^{BC}F^{A'}{}_{A_4\cdots A_{2s})BC},$$

where $\phi = 2I = \Psi_{ABCD} \Psi^{ABCD}$ (Ref. 8) then these equations admit analogous solutions, viz.

(1)
$$h_{0\cdots 0} = \Delta^{-s} P_{+s} S_{+s},$$

$$F_{0\cdots 01'} = \frac{-1}{\sqrt{2} \rho^2 (2s-1)(s-1) \Psi_2} \times \Delta \left(\mathscr{D}_s^{\dagger} - \frac{(2s-1)}{\tilde{\rho}^*} \right) \Delta^{-s} P_{+s} S_{+s},$$

$$F_{0\cdots 00'} = \frac{1}{\sqrt{2} \tilde{\rho}^* (2s-1)(s-1) \Psi_2} \times \left(\mathscr{L}_s - \frac{(2s-1)ia \sin \theta}{\tilde{\rho}^*} \right) \Delta^{-s} P_{+s} S_{+s},$$

(2)
$$h_{1\cdots 1} = \frac{1}{(\tilde{\rho}^*)^{2s}} P_{-s} S_{-s},$$

$$F_{1\cdots 10'} = \frac{-1}{\sqrt{2}(\tilde{\rho}^*)^{2s}(2s-1)(s-1)\Psi_2} \times \left(\mathcal{D}_0 - \frac{(2s-1)}{\tilde{\rho}^*}\right) P_{-s} S_{-s},$$

$$F_{1\cdots 11'} = \frac{-\Delta}{\sqrt{2}\tilde{\rho}(\tilde{\rho}^*)^{2s}(2s-1)(s-1)\Psi_2} \times \left(\mathcal{L}_s^{\dagger} - \frac{(2s-1)ia\sin\theta}{\tilde{\rho}^*}\right) P_{-s} S_{-s}, (52)$$

where $P_{+s}S_{+s}$ are separable solutions of

$$(\Delta \mathcal{D}_1 \mathcal{D}_s^{\dagger} + \mathcal{L}_{1-s}^{\dagger} \mathcal{L}_s - 2(2s-1)i\sigma\tilde{\rho})\Delta^{-s}P_{+s}S_{+s} = 0$$

and $P_{-s}S_{-s}$ separable solutions of

$$(\Delta \mathcal{D}_{1}^{\dagger} \mathcal{D}_{s} + \mathcal{L}_{1-s} \mathcal{L}_{s}^{\dagger} + 2(2s-1)i\tilde{qp})\Delta^{-s}P_{-s}S_{-s} = 0.$$

Equation (51) is a generalization of the Rarita-Schwinger equation although it does not in itself have obvious physical significance for $s > \frac{3}{2}$. We also note that the method of proof for the Teukolsky-Starobinsky identities can be successfully used in the general context of Torres del Castillo.⁵ Indeed we have the following result.

Theorem 3: If the operators \mathcal{D}_n and \mathcal{D}_n^{\dagger} are defined by

$$\mathcal{D}_{n} = \frac{\partial}{\partial r} + i \frac{q}{Q} + n \frac{Q^{(1)}}{Q} = Q^{-n} \mathcal{D}_{0} Q^{n}$$

$$\mathscr{D}_{n}^{\dagger} = \frac{\partial}{\partial r} - i \frac{q}{Q} + n \frac{Q^{(1)}}{Q} = Q^{-n} \mathscr{D}_{0}^{\dagger} Q^{n}, \qquad (53)$$

with the functions q and Q polynomials such that $q^{(3)} = 0$ and $Q^{(5)} = 0$, then for all integer s

Kalnins, Miller, Jr., and Williams

2928

J. Math. Phys., Vol. 30, No. 12, December 1989

$$Q^{s} \mathcal{D}_{0}^{\dagger 2s} \left[Q \mathcal{D}_{1-s} \mathcal{D}_{0}^{\dagger} - (2s-1)iq^{(1)} + \frac{1}{6}(s-1)(2s-1)Q^{(2)} \right]$$

$$= \left[Q \mathcal{D}_{1-s}^{\dagger} \mathcal{D}_{0} + (2s-1)iq^{(1)} + \frac{1}{6}(s-1)(2s-1)Q^{(2)} \right] Q^{s} \mathcal{D}_{0}^{\dagger 2s}, \tag{54}$$

with a similar complex conjugate identity also holding. This

theorem applies to all non-null orbit, type-D vacuum metrics given as for example by Torres del Castillo.⁵

ACKNOWLEDGEMENT

The work of W. M. was supported in part by the National Science Foundation under grant DMS 86-00372.

APPENDIX

Here we list the value of the Teukolsky-Starobinsky constant $|D_s|^2$ for a number of values of s, where $\tilde{\alpha}^2 = \alpha^2 + ma/\sigma$: $|D_{1/2}|^2 = \lambda$,

$$|D_1|^2 = \lambda^2 - 4\sigma^2 \tilde{\alpha}^2,$$

$$|D_{3/2}|^2 = \lambda^2(\lambda + 1) - 16\sigma^2(\lambda \tilde{\alpha}^2 - a^2),$$

$$|D_2|^2 = \lambda^2 (\lambda + 2)^2 - 8\sigma^2 \tilde{\alpha}^2 \lambda (5\lambda + 6) + 96\sigma^2 a^2 \lambda + 144\sigma^4 \tilde{\alpha}^4 + 144\sigma^2 M^2,$$

$$|D_{5/2}|^2 = \lambda^2 (\lambda + 3)^2 (\lambda + 4) - 16\sigma^2 \tilde{\alpha}^2 \lambda (\lambda + 3)(5\lambda + 8) + 48\sigma^2 a^2 \lambda (7\lambda + 12) + 1024\sigma^4 \tilde{\alpha}^4 (\lambda + 1) - 3072\sigma^4 \tilde{\alpha}^2 a^2 + 1152\sigma^2 M^2 (\lambda + 2).$$

$$|D_3|^2 = \lambda^2 (\lambda + 4)^2 (\lambda + 6)^2 - 4\sigma^2 \tilde{\alpha}^2 \lambda (\lambda + 4) (35\lambda^2 + 252\lambda + 360) + 128\sigma^2 a^2 \lambda (\lambda + 4) (7\lambda + 15)$$

$$+ 16\sigma^4 \tilde{\alpha}^4 (259\lambda^2 + 1140\lambda + 900) - 2560\sigma^4 \tilde{\alpha}^2 a^2 (11\lambda + 15) + 25600\sigma^4 a^4 - 14400\sigma^6 \tilde{\alpha}^6$$

$$+ 576\sigma^2 M^2 ((3\lambda + 10)^2 - 100\sigma^2 \tilde{\alpha}^2).$$

$$\begin{split} |D_{7/2}|^2 &= \lambda^2 (\lambda + 5)^2 (\lambda + 8)^2 (\lambda + 9) - 32\sigma^2 \tilde{\alpha}^2 \lambda (\lambda + 5) (\lambda + 8) (7\lambda^2 + 63\lambda + 108) \\ &+ 288\sigma^2 a^2 \lambda (\lambda + 5) (7\lambda^2 + 65\lambda + 120) + 256\sigma^4 \tilde{\alpha}^4 (49\lambda^3 + 549\lambda^2 + 1728\lambda + 1296) \\ &- 4608\sigma^4 \tilde{\alpha}^2 a^2 (31\lambda^2 + 175\lambda + 180) + 57600\sigma^4 a^4 (5\lambda + 9) - 147456\sigma^6 \tilde{\alpha}^6 (\lambda + 3) \\ &+ 884736\sigma^6 \tilde{\alpha}^4 a^2 - 92160\sigma^4 M^2 (7\tilde{\alpha}^2 \lambda + 30\tilde{\alpha}^2 - 15a^2) + 5760\sigma^2 M^2 (3\lambda^3 + 45\lambda^2 + 220\lambda + 360), \end{split}$$

$$|D_4|^2 = \lambda^2 (\lambda + 6)^2 (\lambda + 10)^2 (\lambda + 12)^2 - 48\sigma^2 \tilde{\alpha}^2 \lambda (\lambda + 6) (\lambda + 10) (7\lambda^3 + 154\lambda^2 + 996\lambda + 1680)$$

$$+576\sigma^2a^2\lambda(\lambda+6)(\lambda+10)(7\lambda^2+78\lambda+168)+96\sigma^4\tilde{a}^4(329\lambda^4+7372\lambda^3+55484\lambda^2+156240\lambda+117600)$$

$$+28224000\sigma^{8}\tilde{\alpha}^{8}+28224000\sigma^{6}M^{2}\tilde{\alpha}^{4}+25401600\sigma^{4}M^{4}$$

$$+630\sigma^2M^2(75\lambda^4+2112\lambda^3+21568\lambda^2+96000\lambda+161280).$$

¹S. A.Teukolsky, Phys. Rev. Lett. 29, 1114 (1972).

²E. T. Newman and R. Penrose, J. Math. Phys. 3, 566 (1962).

³R. Güven, Proc. R. Soc. London Ser. A 356, 465 (1977).

⁴R. Güven, Phys. Rev. D 10, 2327 (1980).

⁵G. F. Torres del Castillo, J. Math. Phys. 29, 2078 (1988).

⁶R. Penrose and W. Rindler, Spinors and Spacetime Vol. 1 Two Spinor Cal-

culus and Relativistic Fields (Cambridge U.P., Cambridge, 1984).

⁷S. Chandrasekhar, *The Mathematical Theory of Black Holes* (Oxford U.P., London, 1984).

⁸D. Kramer, H. Stephani, E. Herlt, and M. MacCallum, Exact Solutions of Einstein's Field Equations (Cambridge U.P., Cambridge, 1980).