A THEORETICAL MODEL FOR WIDE-BAND IR-ABSORPTION MOLECULAR SPECTRA AT ANY PRESSURE:

FICTION OR REALITY?

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Plan



- **2** Non-Markovian Energy-Corrected Sudden approach
- Output to high-density IR spectral
- 4 Applications to CO₂ IR spectra at low densities
- **5** Conclusions & perspectives

Outline of the needs







Planets: http://www.space.com/11187-earth-magnetic-field-solar-wind.html

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Outline of the needs



Outline of the needs

STANDARD APPROACH for allowed absorption

Niro et al., JQSRT 88, 483 (2004)





STANDARD APPROACH: PROBLEMS



- **6** One-sided sum rules \rightarrow asymptotic behavior σ^{-3} in the far wing
 - ightarrow negative intensities in low-frequency wing

General sum rules hold ONLY in the non-Markovian case

Bulanin et al., JQSRT 31, 521 (1984)

Symmetrized spectral density:

$$S^{(r)}(\sigma) = \frac{1}{\pi} \operatorname{Re} \sum_{k,k'} A_k^{(r)} \left[i(\sigma - \sigma_k) + n_2 \Gamma^{(r)}(\sigma) \right]_{k,k'}^{-1} A_{k'}^{(r)}$$

r - rank of the coupling tensor

$$\sigma_k$$
 - proper frequencies of the free active molecule
 n_2 - gas density
 $A_k^{(r) \ 2}$ - intensities of separate lines $A_k^{(r)} = (-1)^{J_i} \sqrt{(2J_i + 1)(2J_f + 1)} \begin{pmatrix} J_i & r & J_f \\ 0 & 0 & 0 \end{pmatrix} \sqrt{\frac{\rho_i + \rho_f}{2}}$

General form of the off-diagonal relaxation matrix elements:

$$\Gamma_{if,i'f'}^{(r)}(\sigma) = -\left[1 + \exp(-hc\sigma/kT)\right]/2n_{if}n_{i'f'}$$

$$\times \sum_{L} (2L+1)F_{if,i'f'}^{Lr} \left[\rho_i \Phi_{L}(\sigma - \sigma_{f'i}) + \rho_{i'} \Phi_{L}(\sigma - \sigma_{fi'})\right]$$

Properties of the non-Markovian relaxation matrix

Symmetry (detailed balance):

$$\Gamma_{if,i'f'}^{(r)}(\sigma) = \Gamma_{i'f',if}^{(r)}(\sigma)$$

Time-reversal symmetry:

 $\Gamma_{if,i'f'}^{(r)}(\sigma) = \Gamma_{fi,f'i'}^{(r)*}(-\sigma)$

Double-sided sum rules:

$$\sum_{i'f' \neq if} \Gamma_{if,i'f'}^{(r)}(\sigma) A_{i'f'}^{(r)} = \sum_{i'f' \neq if} A_{i'f'}^{(r)} \Gamma_{i'f',if}^{(r)}(\sigma) = -\Gamma_{if,if}^{(r)}(\sigma) A_{if}^{(r)}$$

Positive definition:

$$\sum_{if,i'f'} B_{if}^{(r)} \Gamma_{if,i'f'}^{(r)}(\sigma) B_{i'f'}^{(r)} \ge 0, \forall B^{(r)}$$

ECS modeling of the correlation function

Buldyreva & Bonamy, PRA 60, 370 (1999)

Re
$$\Phi_{L}(\sigma) = \frac{2}{1 + \exp(-hc \,\sigma/kT)} \Phi_{L}^{class}(\sigma)$$

 $\Phi_{L}^{class}(\sigma) = Q_{L}^{'} \Omega(\sigma)$
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 $for all conducts of the second second$

 $\Omega(\sigma$

Sp in

Modifications for IR absorption bands r = 1

Buldyreva & Daneshvar, JCP 139, 164107 (2013)

Linear molecule with a bending mode $(v_1v_2^lv_3)$ can be considered as a symmetric top with $K \equiv l$ IOSA: Green, JCP 70, 816 (1979); JCP 90, 3603 (1989)

(I) Isolated-line intensities

$$A_{k}^{(r)} = (-1)^{J_{i}+l_{f}} \sqrt{(2J_{i}+1)(2J_{f}+1)} \begin{pmatrix} J_{i} & r & J_{f} \\ l_{i} & l_{f}-l_{i} & -l_{f} \end{pmatrix} \sqrt{\frac{\rho_{i}+\rho_{f}}{2}}$$

(II) Relaxation matrix

$$F_{if,i'f'}^{Lr} = (-1)^{r+l_i+l_f} \begin{pmatrix} J'_i & L & J_i \\ -l_i & 0 & l_i \end{pmatrix} \begin{pmatrix} J_f & L & J'_f \\ l_f & 0 & -l_f \end{pmatrix} \begin{bmatrix} J_i & J_f & r \\ J'_f & J'_i & L \end{bmatrix}$$

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(III) Vibration-rotation interactions:

(A) Isolated-line intensities: Correction via Herman-Wallis factors

$$F(m) = (1 + a_1 m + a_2 m^2 + a_3 m^3)^2 \qquad P-\& R-$$

$$F(m) = (1 + b_2 m^2)^2 \qquad Q-$$

$$m = -J \quad (P); \quad m = J \quad (Q); \quad m = J+1 \quad (R)$$

$$A_k^{(1)} \to A_k^{(1) HW} = A_k^{(1)} \sqrt{F(m)}$$

Coriolis interactions between vibrational levels:

Hartmann et al., JCP 110, 7733 (1999)

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- If sum rules are restricted to one band, unperturbed dipole elements should be used
- Modifications of rel. matr. for Coriolis effects have no influence on the absorption shape

(B) Relaxation matrix modeling:



CO₂: Effective correction allowed for diagonal elements 12

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Buldyreva et al., JCP 138, 164117 (2013)

 $CO_2 - CO_2 \quad 2v_1 + v_3$



HW factors: Rothman et al., JQSRT 48, 537 (1992)

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 $CO_2 - CO_2 = 3v_3$



HW factors: Rothman et al., JQSRT 48, 537 (1992)

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Filippov et al., PCCP 15, 13826 (2013)



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Buldyreva & Daneshvar, JCP 139, 164107 (2013)

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SOIR spectra of the atmosphere of Venus (Venus Express)



Courtesy : A.C. Vandaele (IASB)







0.029 atm Q-branch zoom



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Daneshvar et al., JQSRT 2014 (under revision)

Conclusions & perspectives

Main results

- Immediate good agreement with expt OCS-He spectra up to 300 atm without any additional fitting parameter
- Satisfactory agreement for CO₂ up to nearly liquid phase (+ 1 fitting parameter)
- Excellent agreement with low-density expt CO₂ spectra down to 0.03 atm
- The model has a universal character (IR, Raman iso, aniso)

Further improvements/tests/applications

- Studies of extremely low pressures (narrowing effects)
- Modeling of I m Γ

5)

- Studies of far wings (non-Markovian effects)
- Modeling of temperature dependences
- Calculations of correlation functions from ab initio PES
- Studies of other bands and other molecular systems 16 June 2014 ISMS