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# A Surface Evolution Approach to Probabilistic Space Carving

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## Abstract

We present a 3D photography method that generates a texture-mapped three-dimensional model of a scene computed from multi-view calibrated two-dimensional photographs. Our approach first performs probabilistic space carving, which results in a 3D grid of voxel probabilities that describe the likelihood of a voxel existing in the model. We then employ a three-dimensional geodesic active surface to extract a most likely conformally-weighted minimal surface from the voxel probabilities. This surface is then polygonalized and texture-mapped, yielding in a 3D model easily rendered with standard graphics hardware.

## 1. Introduction

The objective of this work is to reconstruct a three-dimensional shape from a set of two-dimensional images in a formal statistical manner through the use of a geometric active surface model which allows us to easily incorporate geometric smoothness constraints.

We begin with an array of voxels where each voxel is assigned a probability of existing in the reconstructed shape. In addition, each voxel is assigned a spherical Gaussian color distribution. Thus when a voxel is projected onto an image, the color of each resulting pixel is a sample point drawn from that voxel's color distribution. This is a useful model because a voxel will often project to several pixels per image and this assumption does not require all of these image pixels to have identical RGB values.

This approach differs from Space Carving [7] because each voxel is assigned a probability for it existing in the model. This probability may be regarded as a continuous measure of photo-consistency. Since this measurement is non-negative, it is naturally used as a weighting factor in the framework of geometric active surface models. Specifically, our goal will be to seek a closed 3D surface of mini-

**Table 1. Notation used in this paper.**

<b>Indices</b>	
$i$	Image index $\in \{1 \dots n\}$
$x, y$	Pixel co-ordinate index
$k, l, m$	Voxel index ( $m$ is depth)
<b>Image data</b>	
$\mathcal{I}^i$	Image $\mathcal{I}^i$
$\mathcal{I}_{xy}^i$	Pixel $x, y$ of image $\mathcal{I}^i$
$\mathcal{R}^i$	An image region within image $i$
$\mathcal{R}_{klm}^i$	Region of $\mathcal{I}^i$ projected by $klm$
<b>Geometric variables</b>	
$\mathbf{P}^i$	Projection matrix for image $\mathcal{I}^i$
$\mathbf{H}_m^i$	Homography from plane $m$ to $\mathcal{I}^i$
<b>Statistical variables</b>	
$D$	All available data $\{\{\mathcal{I}^i\}, \{\mathbf{P}^i\}\}$
$\exists_{klm}$	$\exists_{klm} \in \{0 = \text{empty}, 1 = \text{exists}\}$
$\mathcal{V}_{klm}$	The model for a voxel at $klm$
$\mathcal{W}_{klm}^i$	The missing voxel model for image $i$
<b>Gaussian parameters</b>	
$\vec{\mu} = (\mu_R \ \mu_G \ \mu_B)$	Gaussian mean in RGB space
$\sigma_R^2, \sigma_G^2, \sigma_B^2$	Gaussian variance in RGB space

mal area but where the local area element on the surface is weighted in inverse proportion to the probabilities assigned to the nearby voxels. In this manner, the surface will be attracted to more photo-consistent locations within the 3D grid (since smaller weighted areas arise at such locations) while remaining smooth as well.

Note that only the voxel probabilities, which define the local area weighting factors, are ‘‘voxelized’’ in this framework. The surface itself is never represented as a set of discrete voxels, but is free to position itself continuously in between voxel locations. The reconstructed surface may be interpreted geometrically as a minimal surface with respect to a conformally Euclidean metric defined by this proba-

bilistic measure of photo-consistency. This surface can then be polygonalized and texture-mapped, producing a model well suited to rendering with standard graphics hardware.

## 2. Computing Voxel Probabilities

Consider pixel  $\mathcal{I}_{xy}^i$  which is located at co-ordinates  $(x, y)$  in the  $i$ -th image. The probability that  $\mathcal{I}_{xy}^i$  is drawn from the spherical Gaussian distribution  $\{\vec{\mu}, \sigma\}$  is denoted  $P(\mathcal{I}_{xy}^i | \vec{\mu}, \sigma)$  and can be estimated using (1).

$$P(\mathcal{I}_{xy}^i | \vec{\mu}, \sigma) = \frac{1}{(\sigma\sqrt{2\pi})^3} e^{-\frac{1}{2\sigma^2} \|\mathcal{I}_{xy}^i - \vec{\mu}\|^2} \quad (1)$$

Now consider the image region  $\mathcal{R}^i$  which is the set of pixels  $\{\mathcal{I}_{xy}^i\}$  from image  $\mathcal{I}^i$ . The probability that these pixels are drawn from the distribution  $\{\vec{\mu}, \sigma\}$  can be calculated using (2). This probability is denoted  $P(\mathcal{R}^i | \vec{\mu}, \sigma)$  and assumes that the image pixels are independent.

$$P(\mathcal{R}^i | \vec{\mu}, \sigma) = \prod_{xy \in \mathcal{R}^i} P(\mathcal{I}_{xy}^i | \vec{\mu}, \sigma) = \prod_{xy \in \mathcal{R}^i} \frac{e^{-\frac{\|\mathcal{I}_{xy}^i - \vec{\mu}\|^2}{2\sigma^2}}}{(\sigma\sqrt{2\pi})^3} \quad (2)$$

The next problem that needs to be discussed is the method used to decide whether or not a voxel should be part of the reconstruction. Each voxel  $klm$  projects to a region  $\mathcal{R}_{klm}^i$  in image  $\mathcal{I}^i$ . The image data in this region is described by one of two models. Either the data observed at  $\mathcal{R}_{klm}^i$  is the result of the projection of a voxel, or it is described by some other feature in the reconstruction as the voxel is empty. The latter case is referred to as the case for the empty voxel. Bayes' theorem is then used to determine which of the two cases is more likely, and is discussed in Section 2.3.

### 2.1. The statistical model for voxel projection

The first possibility is that the voxel  $klm$  is present in the reconstruction. This case is denoted  $\exists_{klm} = 1$  and is shown diagrammatically in Figure 1. In this case, the image data in region  $\mathcal{R}_{klm}^i$  is described by the statistical color distribution  $\mathcal{V}_{klm}$  of voxel  $klm$ , provided that the voxel is not occluded in this viewpoint. Remember, each voxel is represented by a spherical Gaussian distribution in RGB space. This means that  $\mathcal{V}_{klm}$  has four degrees of freedom that need to be estimated, which are  $\{\mu_R, \mu_G, \mu_B, \sigma^2\}$  (see Table 1). The probability that the observed data in image  $\mathcal{I}^i$  is described by the voxel model  $\mathcal{V}_{klm}$  is obtained by considering the pixels in region  $\mathcal{R}_{klm}^i$  and is denoted  $P(\mathcal{I}^i | \exists_{klm}=1, \mathcal{V}_{klm})$ .

$$P(\mathcal{I}^i | \exists_{klm}=1, \mathcal{V}_{klm}) = P(\mathcal{R}_{klm}^i | \vec{\mu}, \sigma) = \prod_{xy \in \mathcal{R}_{klm}^i} P(\mathcal{I}_{xy}^i | \vec{\mu}, \sigma) \quad (3)$$

Equation (3) gives the probability for the data observed in one particular image, given that the voxel exists in the

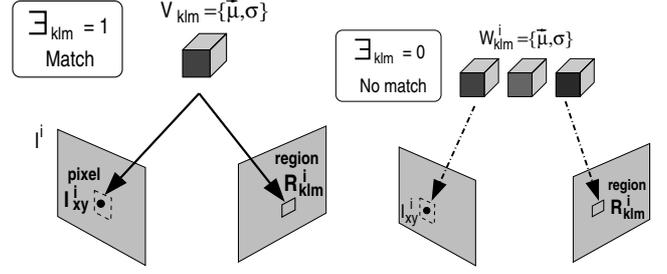


Figure 1. The models for the projection of a voxel (left) and for a missing voxel (right).

reconstruction. The combined probability for observing all of the images can be calculated by assuming that each of the images is independent.

$$P(D | \exists_{klm}=1, \mathcal{V}_{klm}) = \prod_i P(\mathcal{I}^i | \exists_{klm}=1, \mathcal{V}_{klm}) \quad (4)$$

Unfortunately, the parameters of the voxel model  $\mathcal{V}_{klm}$  are not known, so they have to be marginalized<sup>1</sup> by integrating over all possible values for the distribution.

$$P(D | \exists_{klm}=1) = \int_{\mathcal{V}_{klm}} P(D | \exists_{klm}=1, \mathcal{V}_{klm}) P(\mathcal{V}_{klm}) d\mathcal{V}_{klm} \quad (5)$$

The prior term  $P(\mathcal{V}_{klm})$  is a distribution that describes the model parameters of  $\mathcal{V}_{klm}$  (see [2, 1] for details).

### 2.2. The statistical model for an empty voxel

The second possibility is that the voxel  $klm$  does not exist. This case is denoted  $\exists_{klm} = 0$  and is shown in Figure 1. In this case, the non-voxel is projected into each of the images, and the image samples that are obtained are projections of *different* voxels. Unfortunately, it is not known which voxels these are until the entire scene has been reconstructed. To get around this difficulty, the voxel projections  $\mathcal{R}_{klm}^i$  are assumed to be locally independent. This means that a missing voxel can be represented by a set of independent statistical models  $\{\mathcal{W}_{klm}^i\}$ , one for each image. Again, each of these models is represented by a spherical Gaussian  $\{\mu_R, \mu_G, \mu_B, \sigma^2\}$  distribution.

The probability that the observed data in image  $\mathcal{I}^i$  is described by the voxel model  $\mathcal{W}_{klm}^i$  is obtained by considering the pixels in region  $\mathcal{R}_{klm}^i$ .

$$P(\mathcal{I}^i | \exists_{klm}=0, \mathcal{W}_{klm}^i) = P(\mathcal{R}_{klm}^i | \mathcal{W}_{klm}^i) = \prod_{xy \in \mathcal{R}_{klm}^i} P(\mathcal{I}_{xy}^i | \vec{\mu}, \sigma) \quad (6)$$

<sup>1</sup>To marginalize a parameter using Bayes' rule, integrate over all possible values for the parameter, multiplied by the prior at that value, and divide by the total probability.

Since the specific value of  $\mathcal{W}_{klm}^i$  is not known, it is necessary to marginalize this parameter by integrating over all possible values of  $\mathcal{W}_{klm}^i$  multiplied by the prior term  $P(\mathcal{W}_{klm}^i)$ . This prior term describes the distribution of the model parameters of  $\mathcal{W}_{klm}^i$  (see [2, 1] for details).

$$P(\mathcal{I}^i | \exists_{klm}=0) = \int_{\mathcal{W}_{klm}^i} P(\mathcal{I}^i | \exists_{klm}=0, \mathcal{W}_{klm}^i) P(\mathcal{W}_{klm}^i) d\mathcal{W}_{klm}^i \quad (7)$$

Equation (7) calculates the probability of the data in a particular image given that it is the result of an independent voxel. This probability can be calculated for each of the images independently and the combined result is given by:

$$\begin{aligned} P(D | \exists_{klm} = 0) &= \prod_i P(\mathcal{I}^i | \exists_{klm} = 0) \\ &= \prod_i \int_{\mathcal{W}_{klm}^i} P(\mathcal{I}^i | \exists_{klm} = 0, \mathcal{W}_{klm}^i) P(\mathcal{W}_{klm}^i) d\mathcal{W}_{klm}^i \end{aligned} \quad (8)$$

Visibility can be addressed by analyzing the existence of voxels along a ray between voxel  $klm$  and  $\mathcal{I}^i$ , as described in [2, 1].

Two probabilities have now been calculated. The first is the probability of the image data given that it is described by a single voxel (5). The second is the probability of the image data given that it is described by the projection of independent voxels (8). The final step is to calculate which of these two cases is more likely.

### 2.3. How to make a decision about a voxel

The probability of a voxel existing  $P(\exists_{klm} = 1 | D)$  is determined using Bayes' Theorem from the probabilities of the two possible voxel models. These are the model for a voxel existing  $P(D | \exists_{klm} = 1)$ , and the model for a voxel being empty  $P(D | \exists_{klm} = 0)$ . The prior probabilities for a voxel existing, or not existing, are denoted:  $P(\exists = 1) + P(\exists = 0) = 1$ . These probabilities can be combined using Bayes' Theorem:

$$P(\exists_{klm} = 1 | D) = \frac{P(D | \exists_{klm} = 1) P(\exists = 1)}{P(D | \exists_{klm} = 1) P(\exists = 1) + P(D | \exists_{klm} = 0) P(\exists = 0)} \quad (9)$$

## 3. Computing the 3D Surface

The result from the previous section of this paper is a volume of probabilities. Now we want to use these probabilities to find a most likely reconstructed surface  $S$ . Following the same intuition as Faugeras and Keriven [5] using the 3D extension of geodesic active contours [3, 11], we will use our probabilistic measure of photo-consistency

to non-uniformly weight the standard Euclidean metric of the volume space by a conformal factor  $\Phi$ , and then seek a minimal surface (surface of least area) where its area is measured according to this conformally weighted metric. Faugeras and Keriven did not employ such a probabilistic measure but used normalized cross correlation measures as their weighting factor instead.

In particular, the surface  $S$  will be chosen to minimize

$$E(S) = \int_S \Phi dA \quad \text{where} \quad \Phi = \frac{1}{1 + P(\exists_{klm} = 1 | D)} \quad (10)$$

by starting with an initial guess for  $S$  and then deforming it via the following gradient flow

$$\frac{\partial S}{\partial t} = \Phi H \vec{N} - (\nabla \Phi \cdot \vec{N}) \vec{N} \quad (11)$$

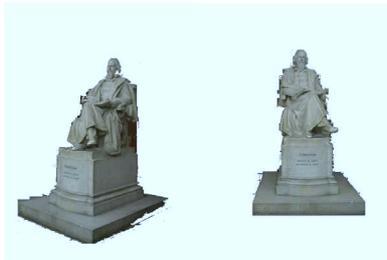
where  $H$  denotes mean curvature and  $\vec{N}$  the inward unit normal to  $S$ . Intuitively, regions that locally are low probability will have  $\Phi \approx 1$  and  $\nabla \Phi \approx 0$ . In such a region, the first term in equation 11 dominates, and the surface  $S$  smooths via evolution by its mean curvature. For regions that locally are high probability, the second term in equation 11 dominates, and the surface  $S$  flows towards and locks onto local maxima of the computed probability  $P(\exists_{klm} = 1 | D)$ . For other values of probability, there is a balance between these two terms.

We use level set methods of Osher and Sethian [8] to implement this flow since they allow easy handling of topological changes, since the surface  $S$  might need to break apart into multiple nearby pieces during evolution. When the flow is complete, the surface  $S$  can be extracted from the level set function using the Marching Cubes [6] algorithm, resulting in a polygonal representation that is easily texture-mapped using the original photographs and rendered in graphics hardware.

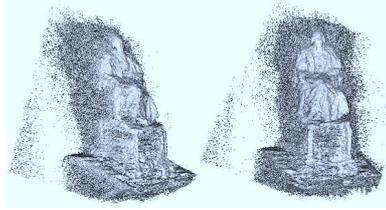
## 4. Experiments

The steps of this algorithm are illustrated in Figure 2. Two images of the Lord Tennyson statue are shown in Figure 2a (after segmenting away the background). The computed voxel probabilities are then examined with some simple heuristics to derive an initial estimate of the surface. Two views of the initial surface are shown in Figure 2b. This gives a rather poor reconstruction since no notion of geometric smoothness has been incorporated into the Bayesian probability estimates. Next, in Figure 2c we see the final reconstructed surface obtained by evolving the initial surface in Figure 2b via the flow of equation 11. Finally, in the bottom row of the figure we see six different views of the final reconstruction after texture mapping the original image data onto the result in Figure 2c.

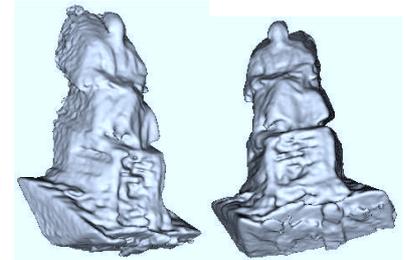
In practice, we found that the gradient field  $\nabla \Phi$  had a limited domain of influence. Consequently, if the initial



(a) Two of the input images



(b) Two views of the initial surface



(c) Two views of the final surface



**Figure 2. Reconstruction of Lord Tennyson statue.** The top row shows two of the input images followed by views of the initial and final surface. The bottom row shows many more views of the final reconstruction with texture mapping of the original image data.

guess for  $S$  was not near a maxima, the gradient term would have little influence on the surface evolution, and the surface  $S$  would not move towards the maxima. To address this issue, we implemented a diffusion of the gradient field using the technique of Xu and Prince [10]. We compute this gradient field diffusion by solving a set of decoupled partial differential equations. This is performed before the surface evolution of equation 11.

## 5 Conclusion

We have presented a method for computing a 3D model of a scene using multiple 2D photographs. Our technique employs a probabilistic framework for space carving and a geodesic active surface model for extracting a smooth, most likely surface.

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