

# Geochemical and isotopic insights into the assembly, evolution and disruption of a magmatic plumbing system before and after a cataclysmic caldera-collapse eruption at Ischia volcano (Italy)

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**Abstract** New geochemical and isotopic data on volcanic rocks spanning the period ~75–50 ka BP on Ischia volcano, Italy, shed light on the evolution of the magmatic system before and after the catastrophic, caldera-forming Monte Epomeo Green Tuff (MEGT) eruption. Volcanic activity during this period was influenced by a large, composite and differentiating magmatic system, replenished several times with isotopically distinct magmas of deep provenance. Chemical and isotopic variations highlight that the pre-MEGT eruptions were fed by trachytic/phonolitic

magmas from an isotopically zoned reservoir that were poorly enriched in radiogenic Sr and became progressively less radiogenic with time. Just prior to the MEGT eruption, the magmatic system was recharged by an isotopically distinct magma, relatively more enriched in radiogenic Sr with respect to the previously erupted magmas. This second magma initially fed several SubPlinian explosive eruptions and later supplied the climactic, phonolitic-to-trachytic MEGT eruption(s). Isotopic data, together with erupted volume estimations obtained for MEGT eruption(s), indicate that >5–10 km<sup>3</sup> of this relatively enriched magma had accumulated in the Ischia plumbing system. Geochemical modelling indicates that it accumulated at shallow depths (4–6 km), over a period of ca. 20 ka. After the MEGT eruption, volcanic activity was fed by a new batch of less

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differentiated (trachyte-latite) magma that was slightly less enriched in radiogenic Sr. The geochemical and Sr–Nd-isotopic variations through time reflect the upward flux of isotopically distinct magma batches, variably contaminated by Hercynian crust at 8–12 km depth. The deep-sourced latitic to trachytic magmas stalled at shallow depths (4–6 km depth), differentiated to phonolite through crystal fractionation and assimilation of a feldspar-rich mush, or ascended directly to the surface and erupted.

**Keywords** Ischia volcano · Magmatic plumbing system · Radiogenic isotopes · Geothermometry · Feldspar assimilation · Caldera collapse

## Introduction

Caldera-forming eruptions are among the most destructive natural phenomena on the planet and can catastrophically disperse huge volumes of fragmented magma across large areas in hours to days. Understanding the processes that control the genesis and evolution of the large-volume silicic magmatic systems that feed caldera-forming eruptions, as well as their fate, is important in terms of hazard assessment (Druitt et al. 2012; Cashman and Sparks 2013). In order to decipher these processes, geological, geochemical and petrological investigations of volcanic rocks emplaced before, during and after large caldera-forming eruptions are needed (e.g. Bachmann and Bergantz 2008; Charlier et al. 2008; Charlier and Wilson 2010). These can help understand both the processes responsible for the temporal and spatial assembly of large magma bodies and the response of a magmatic plumbing system following its partial destruction. Numerous studies of this type have been carried out to establish the complexity of these magmatic systems and the important roles that magma composition, temperature, crystallinity and magma mixing play in their evolution (e.g. Foden 1986; Bacon and Druitt 1988; Bachmann and Bergantz 2008; Charlier and Wilson 2010; Pabst et al. 2008). Nevertheless, the origin of large volumes of silicic magma and the relative roles of fractional crystallization of mantle-derived mafic magmas, crustal contamination and/or remelting of crustal intrusive bodies in their petrogenesis remain a matter of debate. Recent work has revealed the short geological timescales over which critical processes within magma chambers can operate (Burgisser and Bergantz 2011; Druitt et al. 2012).

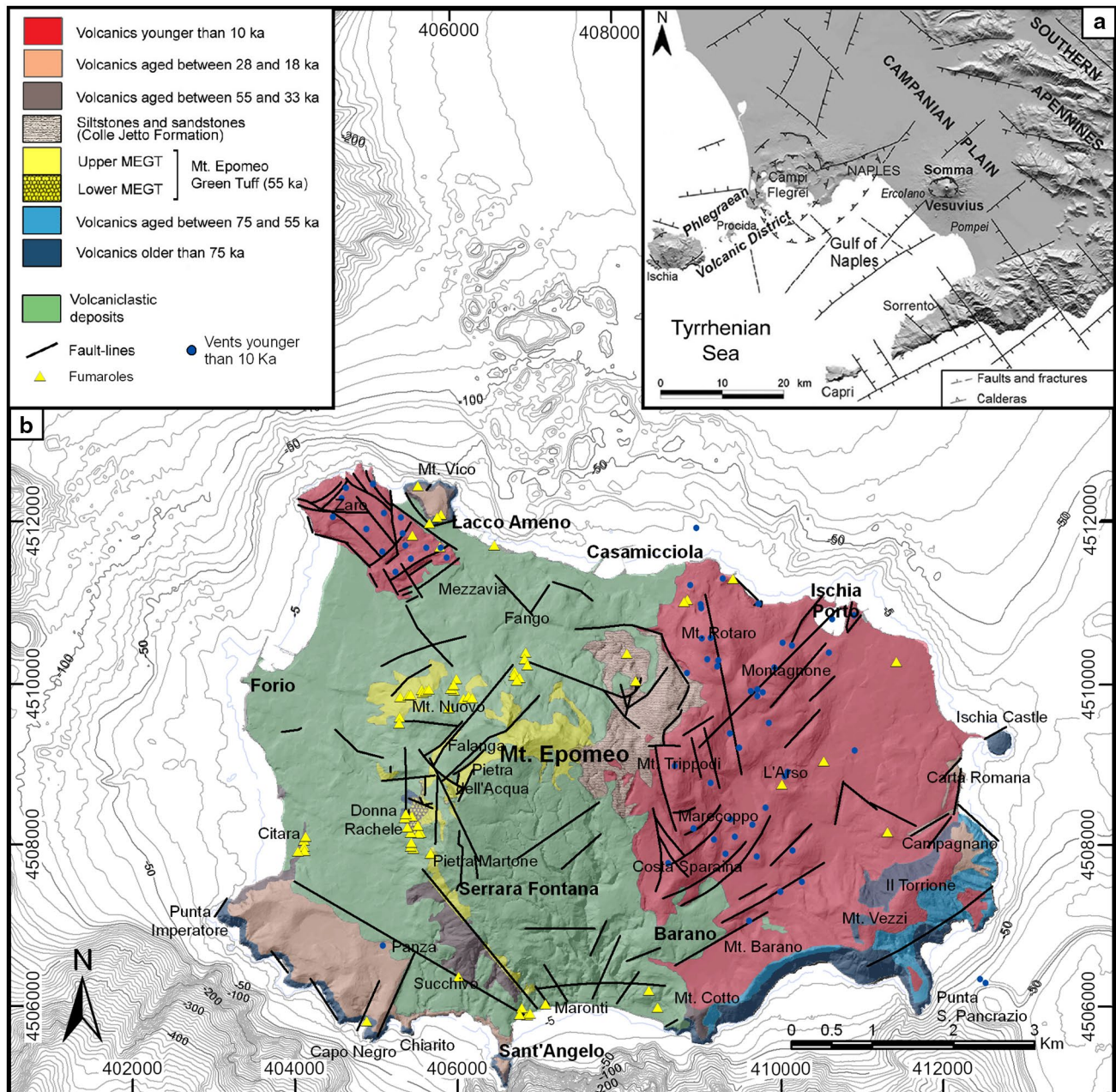
This paper presents a detailed study of the mineralogy and trace, major and isotope geochemistry of the magmas erupted at Ischia volcano, Italy, before, during and after the catastrophic, caldera-forming Monte Epomeo Green Tuff (MEGT) eruption. Ischia has received relatively scant scientific attention compared with neighbouring Campi

Flegrei and Vesuvius volcanoes. The aim was to establish the characteristics of the magmatic system prior to and after the caldera-forming eruption, and to investigate the processes that control the genesis and evolution of the voluminous silicic magmatic reservoir. The island of Ischia (Fig. 1) is an instructive place to investigate this because the stratigraphic sequence, spanning the most recent caldera-forming eruption (MEGT eruption; Vezzoli 1988), is well known (e.g. Brown et al. 2008), and both pre- and post-caldera volcanic deposits are exposed outside the caldera depression. We outline new geochemical and isotopic data on the volcanic succession of the 75–50 ka BP period of activity and use this to illustrate how the Ischia magmatic system prepared for the caldera-forming eruption and how it started to recover. Geochemical modelling constrains the depth and history of the pre- and post-MEGT magmas that were contaminated by Hercynian continental crust at 8–12 km depth and changed in composition from latite to phonolite through fractional crystallization and assimilation of feldspar from a crystal mush at 4–6 km depth.

## Geological and magmatic history of Ischia

Ischia, an island volcano located at the NW corner of the Gulf of Naples, Southern Italy, is a densely populated active resurgent caldera (Fig. 1). The oldest dated rocks on the island, 150 ka old, are poorly exposed along the southern coast and comprise part of an ancient island, about which little is known. Geophysical data indicate that Ischia is the remnant of an older volcanic complex that extended to the west (Orsi et al. 1999; Bruno et al. 2002; Paoletti et al. 2013). Petrological studies (Rittmann and Gottini 1981; Di Girolamo et al. 1979, 1995; Poli et al. 1987; Ghiara et al. 1979; Crisci et al. 1989; Civetta et al. 1991; D'Antonio et al. 2007, 2013) show that Ischia belongs to the low K-Series (Appleton 1972). Its volcanic products range in composition from shoshonite to latite, trachyte and phonolite (most abundant), according to the total alkali versus silica (TAS) classification diagram (Le Maitre et al. 1989). Phenocryst assemblages include alkali-feldspar, plagioclase, biotite, salite, diopside and olivine as major phases, and apatite, sphene, titanomagnetite and alkali-amphibole as the most common accessory phases (depending on whole rock composition).

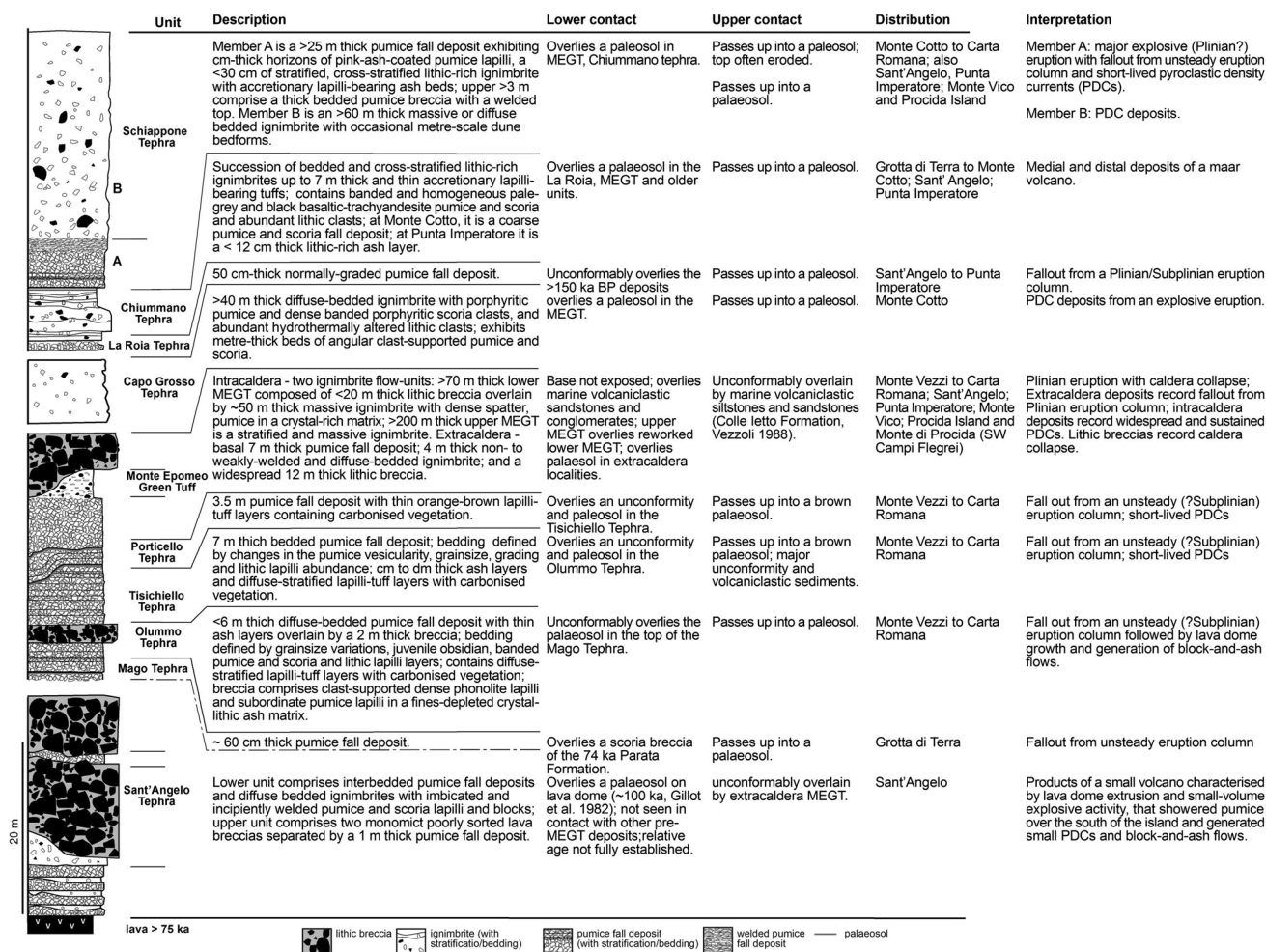
Since 150 ka BP, Ischia has experienced discontinuous volcanic activity (Gillot et al. 1982; Vezzoli 1988; Orsi et al. 1996; de Vita et al. 2006, 2010; Fig. 1). Between 150 and 75 ka BP, trachytic and phonolitic lavas, lava domes and pyroclastic rocks were erupted. Between 75 and 60 ka BP, explosive eruptions emplaced a succession of pumice fall deposits, block-and-ash flow deposits and ignimbrites (Forcella et al. 1982; Rosi et al. 1988; Vezzoli 1988; Brown et al.



**Fig. 1** Geological map of Ischia. **a** Structural sketch map of the Neapolitan volcanic area, **b** geological map of Ischia showing the outcrop of rocks of different ages on the island (modified after Della Seta et al. 2012)

2008). The last 55 ka of activity on the island have been divided into three magmatic cycles based on stratigraphical, geochronological, geochemical, and Sr isotopic data (Poli et al. 1989; Civetta et al. 1991). The first cycle (55–33 ka BP) started with the caldera-forming MEGT eruption. Several phases of caldera collapse may have occurred during this period (Brown et al. 2008; Sbrana et al. 2009), and a  $\sim 10 \times 7$ -km-diameter caldera was formed (Vezzoli 1988; Tibaldi and Vezzoli 1998). The start of the second cycle (28–18 ka BP) was marked by the arrival of magmas that

were less enriched in radiogenic Sr and that progressively mixed with the residual, more radiogenic and more differentiated magma residue from the first cycle. This recharge event probably initiated block resurgence in the caldera (Civetta et al. 1991). The third cycle (10000 BP–1302 AD) was characterized by the eruption of latitic to phonolitic magmas with a wide range of isotope compositions. Petrographical, geochemical and isotopic characteristics of the latites (the Molarà, Vateliero, Cava Nocelle and Arso eruptions) erupted over the past 3 ka suggest mingling and mixing among



**Fig. 2** Summary volcanic stratigraphy for the period ~75–55 ka on Ischia (modified from Brown et al. 2008)

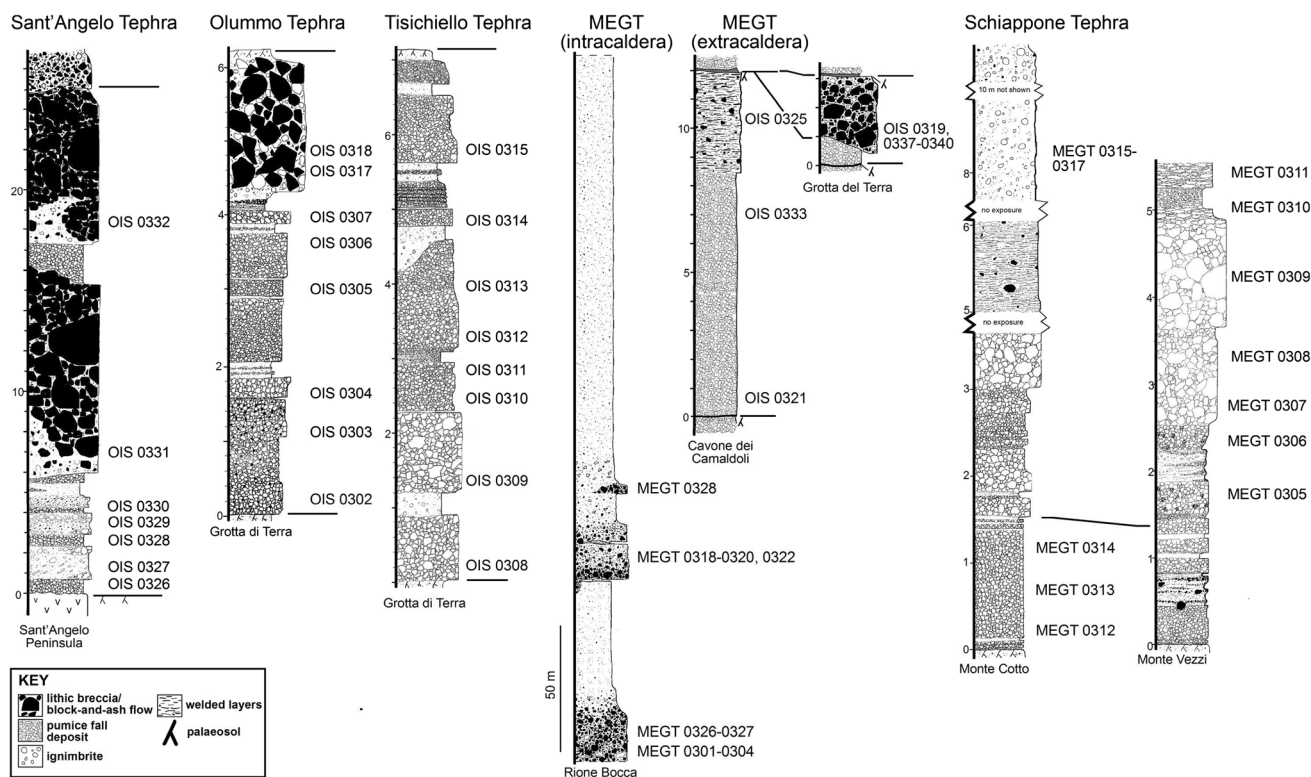
variably evolved, small magma batches and the incorporation of crystals inherited from previous eruptions (e.g. Civetta et al. 1991; Di Girolamo et al. 1995; Piochi et al. 1999; D'Antonio et al. 2013). Data from olivine-hosted melt inclusions (MIs) in latites require gas–melt equilibria between 3 and 18 km depth, suggesting the presence of both deep and shallow magma reservoirs (Moretti et al. 2013).

The genesis of Ischia magmas has been studied by Tonarini et al. (2004), D'Antonio et al. (2007, 2013) and Moretti et al. (2013). In particular, Moretti et al. (2013) suggest that magma genesis occurs in a mantle wedge modified by CO<sub>2</sub>-rich fluids produced by devolatilization of subducted terrigenous-pelagic metasediments of the Ionian oceanic lithosphere. These mantle-derived magmas, mostly propelled by CO<sub>2</sub>-rich fluids, rose to mid-upper crustal levels, where they stagnated and differentiated, and were subsequently periodically tapped as distinct magma batches. The arrival of these magma batches at shallow depths (Civetta et al. 1991; Piochi et al. 1999, and references therein) resulted in alternating periods of resurgence,

intense volcanic activity and quiescence (Orsi et al. 1991; 1996; de Vita et al. 2006). The Monte Epomeo resurgent block records a net uplift of >900 m over ~30 ka (see Vezzoli 1988; Orsi et al. 1991; Tibaldi and Vezzoli 1998). Repeated flank failures have occurred during uplift, resulting in numerous debris flows (Tibaldi and Vezzoli 2004; de Vita et al. 2006; Rapolla et al. 2010; Della Seta et al. 2012) that have built a large submarine debris apron. The present active volcanic state of the island is indicated by historical ground movements (Buchner et al. 1996), seismicity (e.g. the 1883 Casamicciola earthquake) and by fumaroles and thermal springs (see Caliro et al. 1999; Inguaggiato et al. 2000; Di Napoli et al. 2009, 2011, 2013).

### Volcanology and stratigraphy of the analysed volcanic units

Samples in this study come from volcanic products erupted on Ischia between ~75 and 50 ka BP. Stratigraphic



**Fig. 3** Locations of samples within the different volcanic rock units on Ischia. Stratigraphic logs modified from Brown et al. (2008). Thicknesses in metres (note variable scale)

and geologic data on these eruptive units, summarized in Figs. 2 and 3, are given in detail by Brown et al. (2008). This period leads up to, and includes part of, the first cycle of post-MEGT volcanic activity on Ischia (Civetta et al. 1991). We chose to study this time span because it brackets the 55 ka BP MEGT eruption—the largest eruption recorded on the island. The volcanic products of this period outcrop in coastal sections on the southern half of the island, as well as on Procida and western Campi Flegrei (Fig. 1). The geological record of the pre-MEGT activity is dominated by the products of explosive volcanic eruptions and includes numerous pumice fall deposits, ash layers, thin ignimbrites and block-and-ash flow deposits (Sant' Angelo to Porticello Tephra, Fig. 2). Products of the MEGT eruption outcrop in the resurgent caldera block in the centre of the island, along the south coast, on Procida and on the mainland. At least two thick ignimbrites (>70 and >200-m thick, respectively) ponded in the caldera. They are separated by a metre of volcanoclastic sediments and may record two phases of subsidence. Extra-caldera deposits of the MEGT include a coarse-grained (Plinian) pumice fall deposit, a widespread lithic breccia and a distal ignimbrite in south-west Campi Flegrei (see Brown et al. 2008). Post-MEGT volcanic activity included Plinian or Subplinian eruptions recorded

by ignimbrites and pumice fall deposits (Capo Grosso, La Roia and Schiappone Tephra, Figs. 2, 3) and phreatomagmatic eruptions (e.g. the Chiummano maar eruption). Several large tuff ring-forming eruptions at ~45 ka BP (e.g. Citara eruptions, Vezzoli 1988) support the notion that caldera collapse increased the access of seawater to erupting magma.

Estimating the volumes of the eruptions during this period is difficult. Most of the deposits are only intermittently exposed along the south, east and west coastlines of Ischia and only a few are found on neighbouring Procida and in Campi Flegrei. A large proportion of the erupted material was dispersed across the sea. Distal ash layers from the MEGT have been found more than 300 km away (Tomlinson et al. accepted, and references therein). The volume of the MEGT eruption can be estimated from the size of the caldera, which measures  $10 \times 7$  km, and has a minimum average subsidence of 200 m. This suggests a volume of ~9–15 km<sup>3</sup> of erupted magma.

The stratigraphic positions of samples analysed in this study are given in Fig. 3 for eruptive units where more than three samples were taken. One sample was taken from each of the Mago and La Roia Tephra pumice fall deposits, one from the Chiummano Tephra and two from the Porticello Tephra (one each from the bottom and top of the unit).

## Analytical procedure

All collected samples are either composite pumices from individual layers or obsidian clasts. Each analysed sample comprised ~500 g of pumice clasts, which were similar in colour, texture and crystal content. All samples were washed in distilled water multiple times, each time checking for the presence of halite. After dechlorinating, the exteriors of pumice clasts were removed with a hand saw, washed again multiple times in distilled water, crushed to lapilli-size particles, then ground and homogenized in an agate mortar. Major and trace elements were analysed at Centre de Recherches Pétrographiques et Géochimiques, Nancy (France). Powders were analysed for major oxides and Sc by inductively coupled plasma–atomic emission spectrometry and for the remainder of trace elements by inductively coupled plasma–mass spectrometry (ICP–MS). Precision is variable: 1–5 %, for major elements, with the exception of P (<10 %); <5 %, for trace element contents higher than 50 ppm; 5–10 %, for trace element contents in the range 10–50 ppm, and 5–15 %, for trace element contents in the range 1–10 ppm.

Supplementary table reports major and trace elements contents of the majority of the collected samples. Due to the intense unavoidable alteration of some of MEGT samples, only trace element data are used for plots and geochemical modelling. Sr- and Nd-isotopic compositions were determined by Thermal Ionization Mass Spectrometry at the Istituto Nazionale di Geofisica e Vulcanologia, Sezione di Napoli-Osservatorio Vesuviano, using a ThermoFinnigan Triton TI<sup>®</sup> multicollector mass spectrometer. 0.1 g of whole rock powders were leached with cold 6 N HCl for 10 min and with warm 6 N HCl for 10 min, then rinsed several times in pure MilliQ<sup>®</sup> water. Sr and Nd were separated by conventional ion-exchange chromatographic techniques. Due to the low Sr content of a large number of samples and to the clear evidence of the effects of seawater chemical alteration, for the majority of the samples, we have analysed ca. 0.1 g of feldspars. Mineral fractions were ultrasonically cleaned in diluted hydrofluoric acid (7 %) and then rinsed with MilliQ<sup>®</sup> water, before dissolution. In few cases, isotope analysis on different aliquots of mineral fractions was performed. Measured <sup>87</sup>Sr/<sup>86</sup>Sr- and <sup>143</sup>Nd/<sup>144</sup>Nd-isotope ratios were normalized for within-run isotopic fractionation to <sup>86</sup>Sr/<sup>88</sup>Sr = 0.1194 and <sup>146</sup>Nd/<sup>144</sup>Nd = 0.7219, respectively. Sr blanks were on the order of less than 0.5 ng during the period of chemistry processing. The mean measured value of <sup>87</sup>Sr/<sup>86</sup>Sr for NIST-SRM 987 and of <sup>143</sup>Nd/<sup>144</sup>Nd for La Jolla, during the period of measurements, was 0.710204 ± 0.000015 (2σ, N = 72) and 0.511834 ± 0.000009 (2σ, N = 32), respectively. The external reproducibility 2σ is calculated according to Goldstein et al. (2003). Sr- and Nd-isotope

ratios have been normalized to the recommended values of NIST-SRM 987 (<sup>87</sup>Sr/<sup>86</sup>Sr = 0.71025) and La Jolla (<sup>143</sup>Nd/<sup>144</sup>Nd = 0.51185) standards, respectively. Sr- and Nd-isotope data for the 75–45 ka volcanic rocks and separated feldspars are reported in Supplementary table. Given the high Rb/Sr, the <sup>87</sup>Sr/<sup>86</sup>Sr ratios measured on whole rocks have been corrected for <sup>87</sup>Rb decay since eruption. Major elements content of minerals used in modelling were determined by Electron Microprobe analyses. Details on the analytical procedures are in Di Napoli et al. (2013).

## Results

### Petrography

The investigated pre-MEGT and MEGT rocks are variably vesicular and porphyritic, with a total phenocryst content of pumice clasts in the range 5–25 vol%, whereas scoria, obsidian and lava clasts may reach 40 vol% (Table 1). Alkali-feldspar is the dominant phenocryst; some crystals reach 2–3 mm in diameter. Additional phenocrysts phases include clinopyroxene, plagioclase and black mica (in decreasing order of abundance): opaque oxide minerals and apatite are common accessory phases. Alkali-feldspar varies from sanidine-Na-sanidine-anorthoclase; clinopyroxene is ferroan diopside with variable Fe content; plagioclase is andesine; black mica is ferrian phlogopite; and opaque oxide is Ti-magnetite. Reverse zoning is often observed in plagioclase and alkali-feldspar crystals, many of which are rounded and show textural signs of resorption and/or recrystallization.

The mineral assemblages vary according to the chemical composition of the rock. Trachytes are characterized by dominant alkali-feldspar, followed by clinopyroxene, plagioclase, ferrian phlogopite, opaque oxide and apatite; the groundmass is mostly glassy with rare microlites of alkali-feldspar and sometimes ferrian phlogopite and clinopyroxene, when optically resolvable. Phonolites are characterized by a rarity or absence of plagioclase, by dominant alkali-feldspar, very minor amounts of clinopyroxene, black mica and opaque oxide and by additional phases such as alkali-amphibole, nepheline and sphene in a few obsidian and lava samples. Of the latter mineral phases, sphene and alkali-amphibole occur as microphenocrysts or microlites, whereas nepheline is confined to the groundmass, along with abundant opaque oxides (Table 1). The occurrence of optically distinct alkali-feldspar overgrowths around alkali-feldspar phenocrysts may suggest mixing among magmas of variable composition (e.g. Ginibre and Wörner 2007).

Petrographic evidence for alteration occurs in some samples from these older units, particularly in the Sant'Angelo, Olummo and MEGT deposits, carbonate as either patches

**Table 1** Summary of petrographic features of 75-50 ka Ischia volcanic rocks

Stratigraphic unit	TAS classification	Phenocrysts		Groundmass	
		Type and relative abundance of minerals	P.I. vol%	Texture	Additional minerals
Schiappone	LT/TR	alk-fd > pl, cpx > bt, mt ± ol	5–15	Vitric, microcrystalline	
Chiummano	LT	alk-fd > pl, cpx > bt, ol, mt	25	Hypocrystalline	
Capo Grosso	TR	alk-fd > pl > cpx > bt > mt	15	Hypocrystalline	
La Roja	TR	alk-fd > pl ≫ cpx, bt, mt	12	Hypocrystalline	
Upper MEGT	TR	alk-fd > pl ± cpx > bt > mt	15–36	Felty, perlitic	ap
Lower MEGT	TR/PH	alk-fd ≫ cpx, bt, pl > mt ± sph	13–41	Vitric, felty, fluidal	ap
Porticello	TR/PH	alk-fd ≫ cpx ± mt	2–4	Vitric	
Tisichiello	TR/PH	alk-fd ≫ cpx, bt > mt	5–7	Vitric	
Olummo	TR/PH	alk-fd ≫ bt > cpx ± pl > mt	5–15	Vitric, microcrystalline	
Mago	PH	alk-fd	2–3	Vitric	
Sant' Angelo	PH	alk-fd ≫ cpx, bt ± pl, mt, sph, ne	5–20	Vitric, microcrystalline	amph
Barano	TR	alk-fd ≫ cpx, bt > pl > mt	15	Felty, microcrystalline	

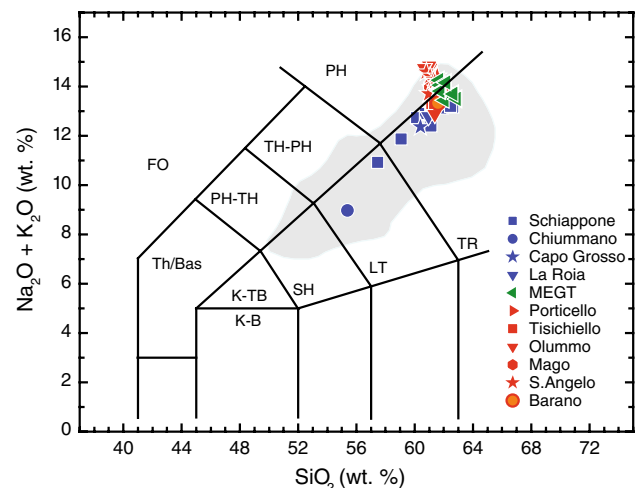
Key for mineral names: *alk-fd* alkali-feldspar, *pl* plagioclase, *bt* biotite, *cpx* clinopyroxene, *mt* titanomagnetite, *ol* olivine, *amph* amphibole, *sph* sphene, *ne* nepheline, *ap* apatite, *P.I.* porphyricity index

in the groundmass, or as veins cross-cutting phenocrysts; partial or complete transformation of clinopyroxene into Fe-oxy-hydroxides and carbonate, and of ferrian phlogopite into chlorite. The groundmass glass of these altered samples is variably transformed in secondary clay minerals and/or analcite and carbonate. More information on the secondary mineralogy of MEGT rocks and their implications for the hydrothermal circulation at Ischia can be found in Di Napoli et al. (2013).

The petrographic features of post-MEGT rocks are similar to those of both pre-MEGT and MEGT rocks. However, post-MEGT rocks include few latites and trachytes (from Chiummano and Schiappone units). The latites are characterized by more abundant plagioclase and clinopyroxene and rare alkali-feldspar and olivine phenocrysts with respect to the trachytes (Table 1).

#### Geochemistry and isotope geochemistry

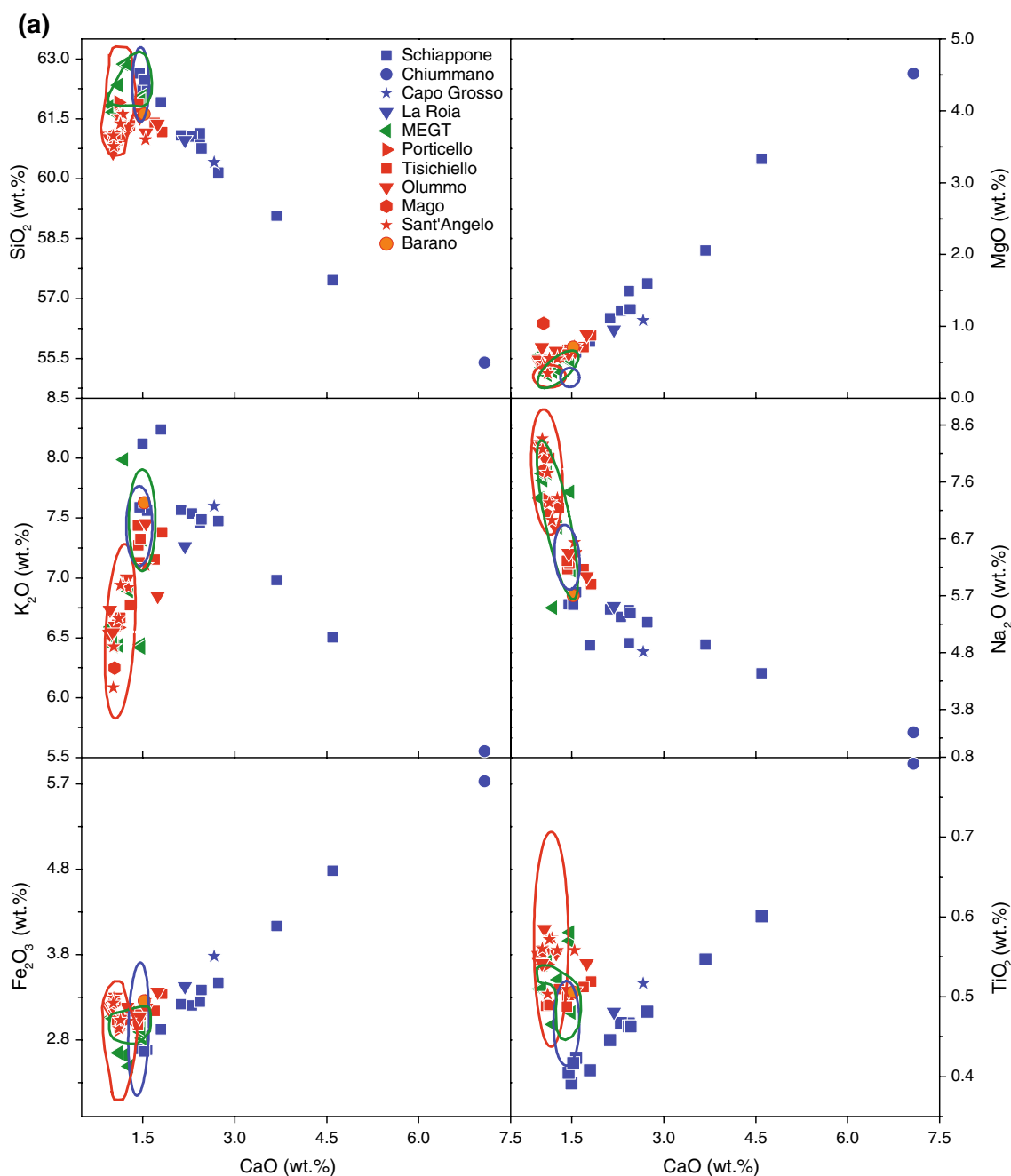
We present new geochemical and isotopic data for volcanic rocks erupted on Ischia between ~75 and 50 ka BP. Figure 4 compares the new data with previously published geochemical data on Ischia volcanic rocks in a TAS plot. The analysed volcanic products range in composition from latite to phonolite. In Fig. 5, whole rock data of the analysed samples are plotted together with fields encompassing major and trace element data acquired on glasses from the same samples (see Tomlinson et al. accepted). In these plots, CaO content has been used as differentiation index. Analyses of pumice from an older ignimbrite (Barano ignimbrite, ~150 ka BP, Gillot et al. 1982) on the southeast coast of Ischia are included as a reference to the earlier magmatic history.



**Fig. 4** Total alkali-silica plot for the volcanic rocks erupted in the study period, on Ischia. Grey shaded area represents all analyses from Ischia. Data from Vezzoli (1988), Crisci et al. (1989) and Civetta et al. (1991)

#### Pre-MEGT eruptions

The oldest eruption in the studied succession is recorded by Sant' Angelo Tephra: a >25-m-thick sequence of pumice fall deposits and thin ignimbrites capped by two block-and-ash flow deposits from a dome-forming eruption (Brown et al. 2008; Fig. 2). The eruption is thought to be of small volume, with the vent located ca. 1 km from the outcrop on the Sant' Angelo peninsula (Fig. 1). The composition of the Sant' Angelo magma (Figs. 5, 6) shows a slight variation through time from more- to less differentiated phonolite, back up to more differentiated phonolite



**Fig. 5** Harker diagrams. **a** CaO versus major elements contents; **b** CaO versus selected trace elements contents. *Red symbols* refer to products older than MEGT (pre-MEGT), *green symbols* to the MEGT

and *blue symbols* to products younger than the MEGT eruption (post-MEGT). Fields are built one data relative to glasses from the same samples (Tomlinson et al., accepted)

(CaO from ~1.1 up to 1.5 wt%, then to ~1.0 wt%; Zr from ~800 to 600 ppm, then up to ~1,100 ppm). Most samples have a slight peralkalinity (Agpaitic Index, A.I. = molar  $(\text{Na}_2\text{O} + \text{K}_2\text{O})/\text{Al}_2\text{O}_3 = 0.98\text{--}1.08$ , Supplementary table). The  $^{87}\text{Sr}/^{86}\text{Sr}$  of Sant'Angelo feldspar phenocrysts is almost constant through the sequence ( $0.706557 \pm 6$  to  $0.706642 \pm 6$ ; hereafter, the quoted error refers to the last digit; Fig. 6).  $^{143}\text{Nd}/^{144}\text{Nd}$  ranges from  $0.512547 \pm 6$  to

$0.512537 \pm 6$ , from the base of the stratigraphic section upward (Supplementary table).

The following eruptive unit is Mago Tephra. It is a poorly preserved <60-cm-thick pumice fall deposit comprised of fine- to medium-grained pumice lapilli. It is considered to have been a small eruption vented from somewhere in the centre of the island. Only one sample was analysed. Its CaO and Zr contents are ~1.0 wt% and



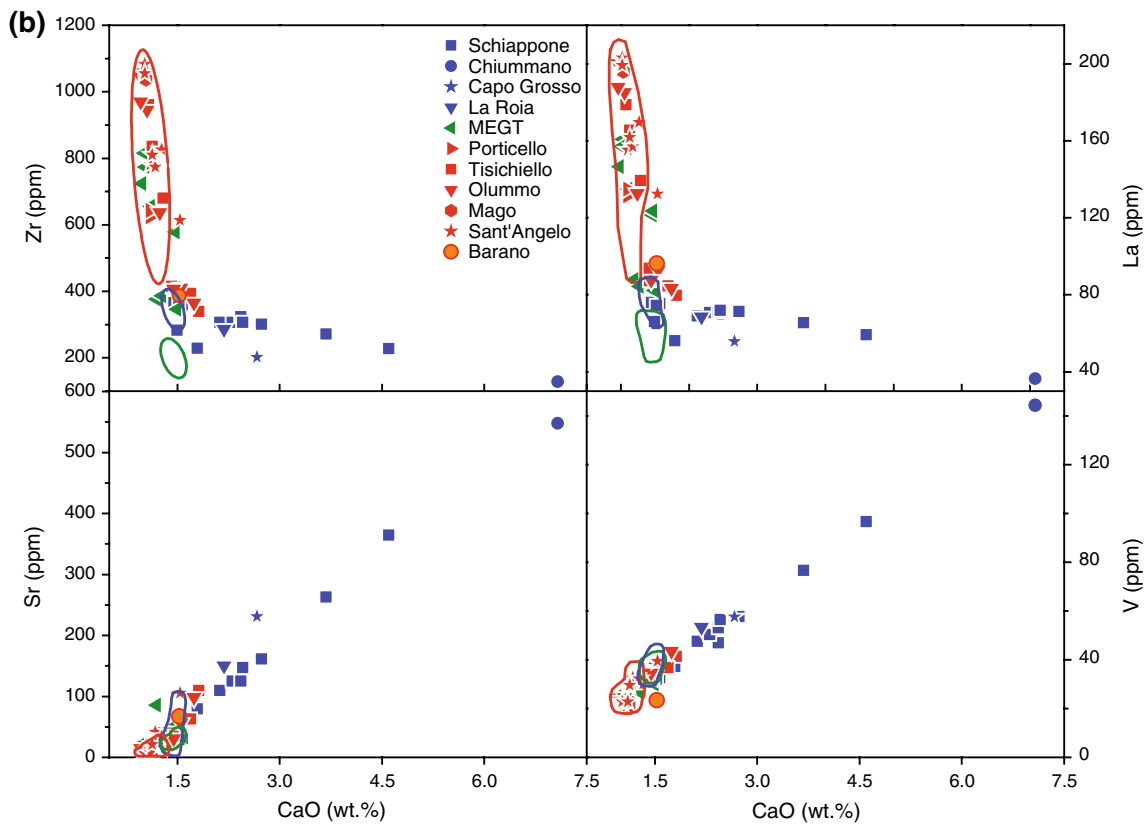
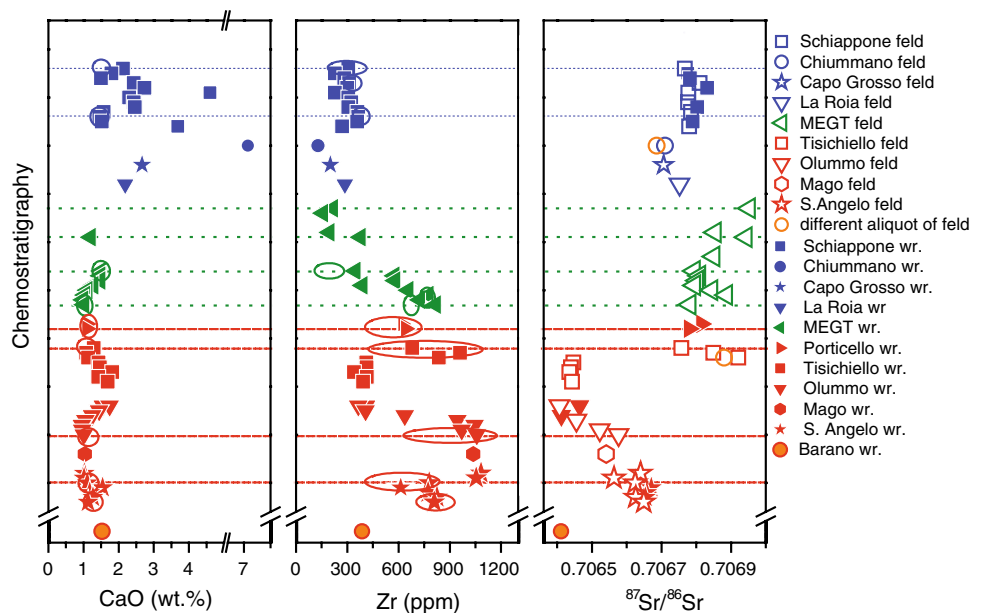


Fig. 5 continued

**Fig. 6** Chemostratigraphy. Red symbols refer to products older than MEGT (pre-MEGT), green symbols to the MEGT and blue symbols to products younger than the MEGT eruption (post-MEGT). Fields are built on data relative to glasses from the same samples (Tomlinson et al., accepted)



~1,000 ppm, respectively (Figs. 5, 6), indicating that the Mago magma (A.I. = ~1.1) was compositionally similar to the last-erupted Sant'Angelo magma, but was slightly less enriched in radiogenic Sr ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.706537 \pm 6$ ). Its  $^{143}\text{Nd}/^{144}\text{Nd}$  is similar to that measured in the last-erupted

Sant'Angelo Tephra products ( $0.512536 \pm 5$ ; Supplementary table).

The Olummo Tephra is a 6-m-thick sequence of pumice fall deposits capped by a block-and-ash flow deposit. The Olummo magma shows a wide range of

composition (Figs. 5, 6). Initially, a more differentiated phonolitic magma (CaO = ~1.0 wt%; Zr = ~1,100 ppm; A.I. = 1.1) was erupted, similar to the Mago and to the last-erupted Sant'Angelo magmas. During the course of the Olummo eruption, magma composition changed to trachyte (CaO = ~1.6 wt%; Zr = ~400 ppm; A.I. = 0.9).  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope ratios of whole rock and separated feldspar phenocrysts gradually changed, during the course of the eruption, from  $0.706569 \pm 6$  to  $0.706416 \pm 6$ , while  $^{143}\text{Nd}/^{144}\text{Nd}$  remained unchanged for all the analysed samples (~0.51254; Fig. 6; Supplementary table).

The >7-m-thick Tisichiello Tephra ( $76 \pm 3$  ka, based on correlation with the varved Lago Grande di Monticchio core, Tomlinson et al., accepted) reaches 7 m in thickness and comprises a complex succession of pumice fall deposits and thin ignimbrites (Brown et al. 2008). It was potentially a larger-volume eruption than the preceding eruptions. The first-erupted Tisichiello magma is trachytic in composition (CaO = ~1.7 wt%; Zr = ~400 ppm; A.I. = 0.9) and compositionally and isotopically similar ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.706449 \pm 6$ ;  $^{143}\text{Nd}/^{144}\text{Nd} = 0.512550 \pm 5$ ) to the last-erupted Olummo magma (Fig. 6; Supplementary table). The late-erupted Tisichiello magma is phonolitic in composition (CaO = ~1.1 wt%; Zr = ~950 ppm; A.I. = 1.1), more radiogenic ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.70687 \pm 6$  to  $0.706732 \pm 6$ ) and is separated from the first-erupted trachytic magma by a relative wide Sr isotopic gap. Its  $^{143}\text{Nd}/^{144}\text{Nd}$  is lower than that of the trachytic Tisichiello magma ( $0.512536 \pm 6$ ; Supplementary table).

The last-recorded eruption before MEGT is the >3.5-m-thick Porticello tephra (Brown et al. 2008;  $59 \pm 2$  ka, based on correlation with the varved Lago Grande di Monticchio core, Tomlinson et al., accepted), which is composed of <4 m of pumice fall deposits and thin ash layers. Products from this eruption are trachytic/phonolitic in composition (CaO = ~1.1 wt%; Zr = ~600 ppm; A.I. = 1.0) and have  $^{87}\text{Sr}/^{86}\text{Sr}$  ( $0.706752 \pm 6$  to  $0.706784 \pm 6$ ) in the range of those of the late products of the Tisichiello eruption.  $^{143}\text{Nd}/^{144}\text{Nd}$  ( $0.512539 \pm 5$ ) is similar, within the analytical errors, to the value characterizing the last-erupted Tisichiello magma.

#### MEGT eruption

The 55 ka MEGT eruption was characterized by the emission of a variety of volcanic products ranging in composition from trachyte to phonolite. Early phases of the eruption are recorded outside the caldera by a 7-m-thick phonolitic-to-trachytic pumice fall deposit (CaO = ~1.0 wt%; Zr = 800–700 ppm; A.I. ~ 1) overlain by a phonolitic welded ignimbrite (CaO = ~1.1 wt%; Zr = 770 ppm; A.I. ~ 1; Figs. 3, 5, 6). This is overlain by lithic breccias considered to record a phase of caldera collapse. Pumice

and dense glassy clasts in these breccias are of less differentiated trachytic/phonolitic composition (CaO = ~1.5–1.1 wt%; Zr = ~350–650 ppm; A.I. = ~1). Intra-caldera deposits comprise two thick ignimbrites. The lower ignimbrite has basal lithic breccias (see Fig. 3) and is compositionally similar (CaO = ~1.2 wt%; Zr = ~370 ppm; A.I. = ~0.9) to the lithic breccias exposed at the coast (Brown et al. 2008). The magma that fed the upper ignimbrite was less differentiated (Zr = 150–210 ppm).

The first-erupted MEGT magma ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.706753 \pm 8$ ,  $^{143}\text{Nd}/^{144}\text{Nd} = 0.512534 \pm 5$ ) is isotopically similar to the Porticello magma and to the last-erupted Tisichiello magma. During the course of the MEGT eruption, Sr isotopic composition varies from  $0.706753 \pm 8$  to  $0.706907 \pm 6$ , while  $^{143}\text{Nd}/^{144}\text{Nd}$  remains constant ( $^{143}\text{Nd}/^{144}\text{Nd} \sim 0.512540$ ; Supplementary table).

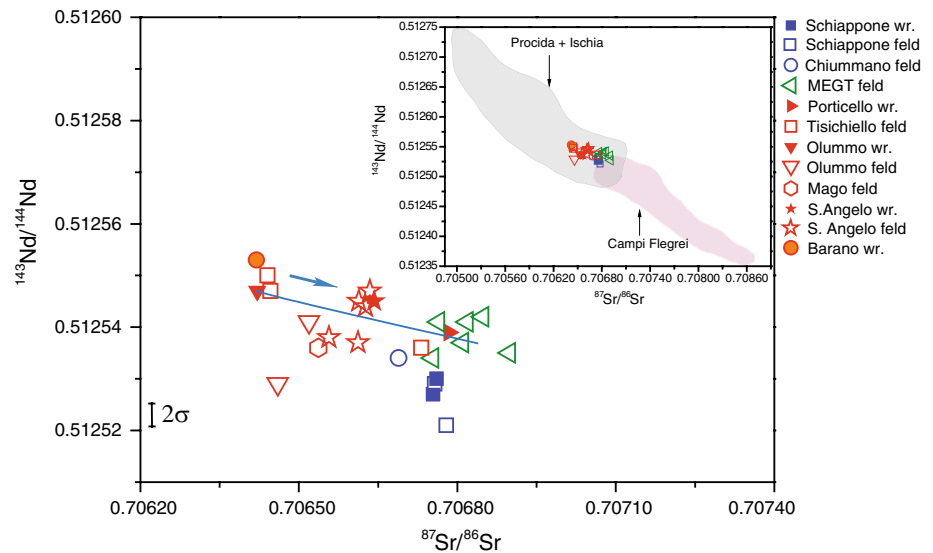
#### Post-MEGT eruptions

The >50-cm-thick La Roia Tephra is the first recorded eruption after the MEGT eruption (Brown et al. 2008). It comprises a 50-cm-thick trachyte pumice fall deposit with Zr contents slightly higher than the last-erupted MEGT magmas (CaO = ~2.2 wt%; Zr = 290 ppm; A.I. = ~0.9), but less enriched in radiogenic Sr ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.706732 \pm 6$ ; Figs. 5, 6) than MEGT.

The Capo Grosso Tephra is a 40-m-thick ignimbrite. The eruption tapped a trachytic magma (CaO = ~2.7 wt%; Zr = 200 ppm; A.I. = ~0.9) slightly less differentiated than the La Roia Tephra and characterized by  $^{87}\text{Sr}/^{86}\text{Sr}$  of  $0.706686 \pm 6$ , less enriched in radiogenic Sr than the last-erupted MEGT and the La Roia magmas (Fig. 6; Supplementary table). The Chiummano Tephra, interpreted as the deposits of a maar eruption in the south of the island, records the extrusion of the least evolved magma of the studied period, having a latitic composition (CaO = ~7.1 wt%; Zr = 130 ppm; A.I. = ~0.7). Two different aliquots of separated feldspar have  $^{87}\text{Sr}/^{86}\text{Sr}$  of  $0.706668 \pm 7$  and  $0.706691 \pm 6$ , while the whole rock has  $^{143}\text{Nd}/^{144}\text{Nd}$  of  $0.512534 \pm 6$  (Supplementary table). The Sr isotopes are similar to those measured for Capo Grosso rocks (Figs. 5, 6).

The youngest eruption in the studied period is the Schiappone Tephra ( $50.6 \pm 2$  ka, based on correlation with the varved Lago Grande di Monticchio core; Tomlinson et al., accepted) whose products outcrop extensively along the south coast of Ischia (Brown et al. 2008). It comprises a complex basal pumice fall deposit up to 6-m-thick overlain by a >60-m-thick ignimbrite. Pumice clasts within the middle part of the pumice fall deposit exhibit compositional banding, and the bulk composition varies from trachyte to latite during the course of the eruption (CaO from 1.4 to 5 wt%; Zr from 370 to 220 ppm;  $0.8 < \text{A.I.} < 0.9$ ). All the

**Fig. 7**  $^{87}\text{Sr}/^{86}\text{Sr}$  versus  $^{143}\text{Nd}/^{144}\text{Nd}$ -isotope compositions relative to the investigated products. In the main plot isotope ratios are plotted together with results from EC-AFC modelling (*blue line*). In the insert the 75–50 ka erupted products are compared with the fields built on the isotope compositions of Procida, Ischia Campi Flegrei erupted products, from literature



erupted products are characterized by homogeneous Sr- and Nd-isotopic compositions ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.70674\text{--}79$ ;  $^{143}\text{Nd}/^{144}\text{Nd} = 0.512527\text{--}30$ ) higher than La Roja, Chiummano and Capo Grosso products.

#### Chemical and isotopic trends

Magma feeding the pre-MEGT and MEGT eruptions were more differentiated (trachyte to phonolite) than those feeding the post-MEGT eruptions (latite to trachyte). With increasing degree of differentiation (decreasing CaO content; Figs. 5, 6), the data show that for the pre-MEGT magmas,  $\text{SiO}_2$ ,  $\text{Al}_2\text{O}_3$ ,  $\text{K}_2\text{O}$  first increase and then decrease; MnO, Zn, Y, Nb, Rb, Cs, rare earth elements (REE) except Eu and Th first remain constant and then increase;  $\text{Na}_2\text{O}$  increases, whereas MgO and  $\text{P}_2\text{O}_5$  (and V, Sr, Ba, Eu) decrease;  $\text{Fe}_2\text{O}_3$  and  $\text{TiO}_2$  first decrease and then increase. The pre-MEGT rocks match well with the highly evolved MEGT magma for major and trace elements, but are less enriched in radiogenic Sr, with the exception of the last-erupted Tisichiello and the Porticello magmas, both characterized by Sr isotope ratios similar to the first-erupted MEGT magma (Fig. 6). The highly evolved pre-MEGT and MEGT rocks are characterized by significantly higher  $\text{TiO}_2$ ,  $\text{Fe}_2\text{O}_3$ ,  $\text{P}_2\text{O}_5$  contents, compared with the post-MEGT rocks, for similar CaO content (Fig. 5). The post-MEGT magmas (latite to trachyte) display a wide range of composition (CaO from 7.0 to 1.5 wt%, Zr from 128 to 365 ppm; Fig. 5). The content of many trace elements (e.g. Zr, REE, Th) well discriminates post-MEGT with respect to pre-MEGT and the most evolved MEGT products (Fig. 6).

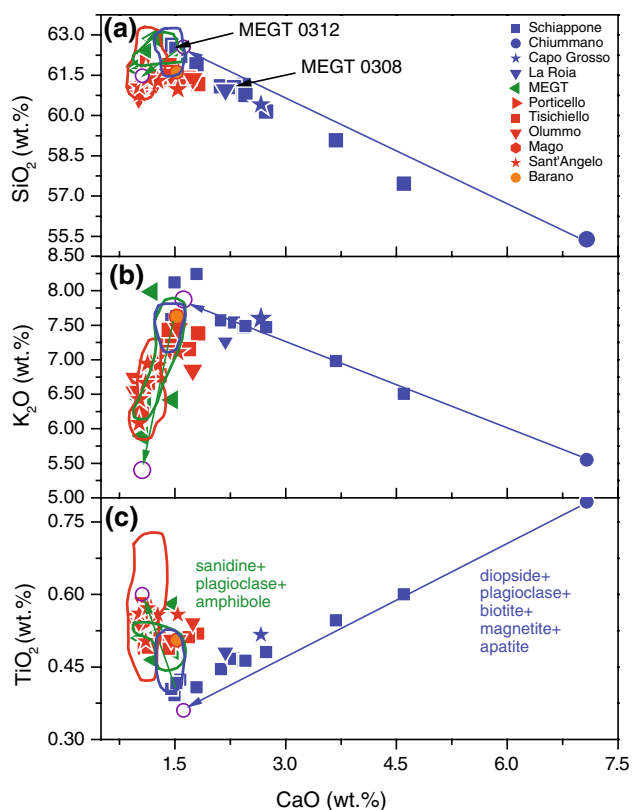
In conclusion, the 75–50 ka BP rocks show a wide range of  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.70642–0.70691). In particular, magmas extruded during the pre-MEGT eruptions show the widest

$^{87}\text{Sr}/^{86}\text{Sr}$  range (0.70642–0.70688), whereas those erupted at the end of Tisichiello eruption and during the Porticello, MEGT and post-MEGT volcanic activity are characterized by a more restricted  $^{87}\text{Sr}/^{86}\text{Sr}$  range (0.70667–0.70691). Nd-isotope ratios show a more restricted range throughout the studied period (0.51253–0.51255), although significant with respect to the analytical error. They show a slightly negative correlation with Sr isotope ratios (Fig. 7).

## Discussion

### Geochemical modelling

The overall compositional variation displayed by the investigated rocks (e.g. Fig. 8a–c) can be attributed, on first approximation, to fractional crystallization processes from a less- to a more evolved magma. This possibility has been explored by means of least-squared mass balance calculations (Stormer and Nicholls 1978). Since Harker diagrams exhibit abrupt changes at CaO ca. 1.5 wt%, two fractionation steps have been considered. For the first step, the composition of the Chiummano latite sample (OIS 0324) has been taken as parent magma. Modelling reveals that ca. 44 wt% removal of a solid assemblage, made up of Mg-rich clinopyroxene, labradoritic plagioclase, biotite, Ti-magnetite and apatite (in decreasing order of abundance), leads to a trachytic daughter magma similar to the first-erupted Schiappone magma (sample MEGT 0312 in Supplementary table, Table 3) ( $\sum r^2 = 0.25$ ; Fig. 8; Table 2). From the latter magma, a second step has been modelled, to reproduce the phonolitic composition of the  $\text{TiO}_2$  rich-glass sample OIS 0329 (Table 3 and Tomlinson et al. accepted). By subtracting ca. 63 wt% of a different mineralogical assemblage,



**Fig. 8** Results of least-squared mass balance calculations (Stormer and Nicholls 1978) have been plotted as purple circles. The first step (blue line) as been modelled from the composition of the poorly evolved (latite) Chiummano sample (OIS 0324) in order to get a trachytic daughter magma similar to the first-erupted Schiappone magma (MEGT 0312) (Table 2). From the latter magma, a second step of fractional crystallization (green line) allow us to partially explain the composition of the residual glasses. Conversely, if using the MELTS code, feldspar assimilation and fractional crystallization are required for reproducing, starting from a magma compositionally similar to the MEGT 0308, the chemical composition of the evolved trachyte magma represented by sample MEGT 0312 (step 2 in Table 3)

constituted by dominant alkali-feldspar, and subordinate andesinic plagioclase and alkali-amphibole (Fig. 8; Table 2), the calculation accounts only in part for the increase in  $\text{TiO}_2$  and  $\text{F}_2\text{O}_{3\text{tot}}$  contents, and the decrease in  $\text{SiO}_2$  shown by the most evolved, phonolitic pre-MEGT magma (Fig. 8).

To further constrain the role of fractional crystallization from less- to more evolved Ischia magmas, and to investigate the possible occurrence of concomitant open-system evolution processes, we used the MELTS code (Ghiorso and Sack 1995; Asimow and Ghiorso 1998; Smith and Asimow 2005). In order to run the program properly, independent geothermobarometric constraints have been obtained by melt-crystal phase equilibria. In particular, by using the  $\text{TiO}_2$  exchange between melt and biotite (Righter and Carmichael 1996), we found that Ischia trachytes equilibrated

at a temperature of 930 °C. This value was obtained for both a  $\text{TiO}_2$  (and  $\text{Fe}^{2+}$ )-poor biotite, likely crystallized from a poorly evolved trachyte magma, and a  $\text{TiO}_2$  (and  $\text{Fe}^{2+}$ )-rich biotite crystallized from a slightly more evolved trachyte, represented by samples MEGT 0308 and MEGT 0312, respectively (Table 3). In the evolving magma, at the obtained temperature, biotite records an increase in oxidation from  $\log f\text{O}_2 = \text{NNO} + 0.75$  to  $\log f\text{O}_2 = \text{NNO} + 1.5$ , possibly accompanied by an increase of the population of  $\text{OH}^-$  groups in the X-site with respect to  $\text{F}^-$  and  $\text{Cl}^-$ . Remarkably, by using the mica-sanidine-magnetite geothermobarometer, proposed by Wones and Eugster (1965) and recently re-calibrated and applied to Campi Flegrei rocks (Fabbrizio et al. 2009), we again obtain a temperature of 930 °C (Table 3). The estimated temperature (as well as the obtained  $\log f\text{O}_2$ ) can be used by adopting the solubility model of Papale et al. (2006) to retrieve information about the water content under which Ischia magmas evolved. Dissolved  $\text{H}_2\text{O}$  contents calculated for  $\log f\text{O}_2 = \text{NNO} + 0.75$  and  $\log f\text{O}_2 = \text{NNO} + 1.5$  are 2.92 wt% (sample MEGT 0308) and 2.40 wt% (sample MEGT 0312), respectively.

If fugacities are approximated to partial pressures (disregarding the presence of  $\text{CO}_2$  in the system), and a crustal density of 2,500  $\text{kg}/\text{m}^3$  is assumed (P. Capuano personal communication) then the calculated  $\text{H}_2\text{O}$  contents correspond to depths of 2–4 km (52–100 MPa). If we consider the occurrence of dissolved  $\text{CO}_2$  in magmas, besides  $\text{H}_2\text{O}$  contents, and assume as  $\text{CO}_2$  contents the background degassing conditions of the Ischia magmas (Moretti et al. 2013), then the storage depths increase to 4–6 km (100–160 MPa). This depth range is in line with both estimates of major storage regions within the Ischia magmatic plumbing system based on geophysical modelling (Orsi et al. 1999; Paoletti et al. 2009, 2013), and recent MI data (Sbrana et al. 2009; Moretti et al. 2013). It also corresponds to the depths of magmas feeding the adjacent Campi Flegrei caldera (e.g. Mangiacapra et al. 2008; Arienzo et al. 2010, and references therein). Henceforth, for calculations relative to trachytic magmas we will use an average pressure value of 130 MPa corresponding to depths of ~5 km. Barometric estimates based on MIs data (Moretti et al. 2013) and geophysical data (Orsi et al. 1999) further suggest the existence of a deeper, less differentiated magmatic reservoir at 8–12 km depth, partially corresponding to the 8–10 km deep magmatic sill revealed by seismic tomography at Campi Flegrei and Vesuvio (Auger et al. 2001; Zollo et al. 2008) and constrained by MIs studies (e.g. Mangiacapra et al. 2008; Arienzo et al. 2010).

For MELTS calculations (see “Appendix”), we used the Chiummano latite (sample OIS 0324; Table 3) as a starting composition. For such a poorly differentiated magma we assumed as initial P and T 200 MPa and 1,200 °C, respectively, in agreement with current knowledge of magma

**Table 2** Summary of data and results of fractional crystallization modelling

From Chiummano latite (OIS 0324) to the TiO <sub>2</sub> -poor Schiappone trachyte (MEGT 0312)													
OXIDE (wt%)	Parent magma		Subtracted phases					Daughter magma	$\Sigma r^2$	Amount (wt%) of subtracted phases	Total amount (wt%) of subtracted phases <sup>a</sup>		
	Observed	Observed	Cpx	Pl	Bt	Mt	Ap	Calculated					
SiO <sub>2</sub>	55.71	62.65	51.09	52.65	41.24	0.06	0.53	62.72	0.25	Cpx	36.94	43.75	
TiO <sub>2</sub>	0.80	0.42	0.64	0.07	3.83	7.57	0.37	0.37		Pl	35.39		
Al <sub>2</sub> O <sub>3</sub>	17.14	18.92	3.88	29.66	11.99	2.59	0.01	18.99		Bt	22.05		
FeO <sub>tot</sub>	5.18	2.46	5.04	0.64	13.87	86.52	0.17	2.53		Mt	3.89		
MnO	0.14	0.14	0.08		0.90	1.21	0.10	0.08		Ap	1.72		
MgO	4.54	0.57	15.75	0.05	17.38	2.04	0.27	0.52					
CaO	7.12	1.53	23.39	12.46		0.02	59.61	1.63					
Na <sub>2</sub> O	3.44	5.56	0.13	3.95	0.84		0.18	5.16					
K <sub>2</sub> O	5.58	7.65		0.53	9.94		0.02	7.89					
P <sub>2</sub> O <sub>5</sub>	0.36	0.10			0.01		39.10	0.11					
TOT	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00					
From Schiappone latite (OIS 0324) to the TiO <sub>2</sub> -rich phonolite (glass of sample OIS 0320 <sup>b</sup> )													
	Parent magma		Subtracted phases			Daughter magma	$\Sigma r^2$	Amount (wt%) of subtracted phases	Total amount (wt%) of subtracted phases <sup>a</sup>				
	Observed	Observed	A-fd	Pl	Amp	Calculated							
SiO <sub>2</sub>	62.65	61.04	65.47	59.17	44.94	61.35	0.39	A-fd	76.78	62.58			
TiO <sub>2</sub>	0.42	0.61	0.16	0.00	2.61	0.62		Pl	16.94				
Al <sub>2</sub> O <sub>3</sub>	18.92	18.89	18.44	24.89	7.58	18.95		Amp	6.28				
FeO <sub>tot</sub>	2.46	3.04	0.58	0.55	18.13	3.30							
MnO	0.14	0.28	0.00	0.00	2.06	0.24							
MgO	0.57	0.32	0.02	0.04	9.66	0.38							
CaO	1.53	1.10	0.23	7.41	10.00	0.93							
Na <sub>2</sub> O	5.56	8.74	3.56	6.24	3.53	8.52							
K <sub>2</sub> O	7.65	5.93	11.54	1.70	1.48	5.58							
P <sub>2</sub> O <sub>5</sub>	0.10	0.06				0.14							
TOT	100.00	100.00	100.00	100.00	100.00	100.00							

Key for the names of minerals: *Cpx* clinopyroxene, *Pl* plagioclase, *Bt* biotite, *Mt* magnetite, *Ap* apatite, *A-fd* alkali-feldspar, *Amp* amphibole

<sup>a</sup> Relative to initial magma

<sup>b</sup> From Tomlinson et al., accepted.  $\Sigma r^2$  = sum of the squares of the residuals

storage at 8–12 km depths (Orsi et al. 1999; Moretti et al. 2013). The first modelled step of magma differentiation ends at 130 MPa (ca. 5 km depth). At such a pressure the residual liquid (72 %) reaches a composition close to that of the poorly evolved trachyte (sample MEGT 0308; Fig. 8; Table 3). From this point on, by using the MELTS code, it is practically impossible to reach the composition of the more evolved Schiappone trachyte (sample MEGT 0312) by fractional crystallization (either by changing fO<sub>2</sub> and/or water content), i.e. it cannot generate a melt with the high SiO<sub>2</sub> and Al<sub>2</sub>O<sub>3</sub> values detected in the natural sample (~62.6 and ~18.9 wt%, respectively; Table 3). On the

contrary, the system evolves towards more under-saturated compositions. MELTS code is calibrated for mafic systems and performs well for MORBS and alkalic mafic magmas. Despite that, it is possible to partially fill the compositional gap between the less evolved (sample MEGT 0308) and the more evolved (sample MEGT 0312) Schiappone trachyte (Fig. 8) by modelling isothermal–isobaric assimilation of felsic rocks (as suggested by geochemical modelling results, reported in the previous section) and still keeping active the fractional crystallization function (“Appendix”). The felsic rocks used for this purpose are feldspars crystallized from earlier magmas. In fact, only feldspars have

**Table 3** Constraints and results of the geochemical modelling (see text)

	Chiummano OIS 0324	Schiappone MEGT 0308.	Schiappone MEGT 0312	S. Angelo OIS 0329	MELTS Step 1 (from OIS 0324 to MEGT 0308)	MELTS Step 2 (from MEGT 0308 to MEGT 0312)	MELTS Step 3 (late phonolite)	MELTS Step 4 (residual liquid) <sup>a</sup>
SiO <sub>2</sub>	55.39	61.25	62.64	62.00	60.61	62.09	61.14	56.92
TiO <sub>2</sub>	0.79	0.47	0.42	0.70	0.57	0.53	1.05	1.73
Al <sub>2</sub> O <sub>3</sub>	17.04	18.64	18.91	18.80	20.22	19.77	18.77	17.37
FeO <sub>tot</sub>	5.73	2.92	2.46	2.90	1.45	1.01	1.13	0.88
MnO	0.14	0.13	0.14	0.30	0.19	0.18	0.43	2.17
MgO	4.51	1.22	0.57	0.30	0.82	0.76	1.21	0.62
CaO	7.08	2.31	1.53	1.26	3.22	2.50	2.67	1.94
Na <sub>2</sub> O	3.42	5.36	5.56	7.40	4.48	5.46	6.46	12.81
K <sub>2</sub> O	5.55	7.56	7.65	6.40	7.92	7.20	6.15	4.58
P <sub>2</sub> O <sub>5</sub>	0.36	0.13	0.10	n.a.	0.52	0.50	0.98	0.97
TOT	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00
<i>Summary of mineral geothermobarometry</i>								
Biotite								
Fe <sup>2+</sup> /Mg		0.27	0.18					
Fe <sup>3+</sup> /Fe <sup>2+</sup>		0.47	1.28					
X <sub>Ti</sub>		0.088	0.118					
X <sub>OH</sub>		0.71	0.93					
X <sub>F</sub>		0.288	0.0072					
X <sub>Cl</sub>		0.0045	0.0005					
Sanidine								
a <sub>San</sub>		0.45	0.68					
Magnetite								
a <sub>Mgt</sub>		0.71	0.71					
T (°C)		930	930					
fH <sub>2</sub> O (bar)		668	498					
fHF (bar)		0.4	0.07					
fHCl (bar)		335	18					
logfO <sub>2</sub>		NNO + 0.75	NNO + 1.44					
H <sub>2</sub> O <sup>b</sup>		2.90	2.40					
Prot(MPa) <sup>c</sup>		100	52					
Prot (MPa) <sup>d</sup>		~160	~100					
Prot (av. estimate)		130	75					

Table 3 continued

	Chiummano OIS 0324	Schiappone MEGT 0308.	Schiappone MEGT 0312	S. Angelo OIS 0329	MELTS Step 1 (from OIS 0324 to MEGT 0308)	MELTS Step 2 (from MEGT 0308 to MEGT 0312)	MELTS Step 3 (late phonolite)	MELTS Step 4 (residual liquid) <sup>a</sup>
<i>Summary of MELTS runs</i>								
Temperature (°C)								
Initial					1,200	930	930	930
Final					930	930	930	910
Pressure (MPa)					200	100	100	60
Initial					130	100	60 <sup>e</sup>	10
Final					3.46 <sup>f</sup>	3.34	2.910	1.47
H <sub>2</sub> O (wt%)		2.5						
Redox state					+1	+2	+2	+2
ΔFMQ					+0.45	+1.43	+1.43	+1.43
ΔNNO					72.02	103.41	40.68	3.24
Liquid (g)								
Fractionated phases (g)								
Olivine (Fo <sub>83</sub> )					2.46			0.47
Olivine (Fo <sub>67</sub> Teph <sub>25</sub> Fa <sub>8</sub> )								
Cpx Di-rich					18.54			0.61
Cpx Di-poor								
Plagioclase (An <sub>59</sub> )					4.77			46.85
Plagioclase (Ab <sub>57</sub> Or <sub>31</sub> An <sub>12</sub> )						62.56		
Feldspar (Or <sub>67</sub> )						13.40		
Feldspar (Or <sub>69</sub> )							0.47	0.47
Biotite							13.59	
Feldspar (Ab <sub>60</sub> An <sub>30</sub> Or <sub>10</sub> )								42.07
Feldspar (Ab <sub>37</sub> An <sub>5</sub> Or <sub>58</sub> )								
Feldspar (Or <sub>60</sub> )						0.62	42.07	1.72
Spinel							0.83	1.05
Apatite							0.22	
Water								
Solids (g)								
Biotite					0.17		>0.0	
Plagioclase (An <sub>53</sub> )					1.61			

Table 3 continued

	Chiummano OIS 0324	Schiappone MEGT 0308.	Schiappone MEGT 0312	S. Angelo OIS 0329	MELTS Step 1 (from OIS 0324 to MEGT 0308)	MELTS Step 2 (from MEGT 0308 to MEGT 0312)	MELTS Step 3 (late phonolite)	MELTS Step 4 (residual liquid) <sup>a</sup>
Feldspar (Ab <sub>62</sub> An <sub>26</sub> Or <sub>12</sub> )						>0.0		
Spinel				0.15		>0.0		
Feldspar (Or <sub>67</sub> )					8.12			
Nepheline (Na- Neph <sub>60</sub> )								0.04
Water							2.13	
Assimilant								
Phase added					50 wt% Na-feldspar (Ab <sub>70</sub> Or <sub>30</sub> )			
Phase added					50 wt% K-feldspar (Ab <sub>10</sub> Or <sub>90</sub> )			
Temperature					600 °C			
Pressure					100 Mpa			
Amount added (g)					80			
Total mass of the system				100	180	100	100	100

<sup>a</sup> This step is the prosecution of the previous one. It is put here to show the consequences of MELTS calculation on extreme fractionations; note the unreasonably high TiO<sub>2</sub>, MnO and P<sub>2</sub>O<sub>5</sub> values

<sup>b</sup> Water (in wt%) from application of the saturation model of Papale et al. (2006)

<sup>c</sup> CO<sub>2</sub> was not considered

<sup>d</sup> The possible presence of CO<sub>2</sub> was considered

<sup>e</sup> Nepheline precipitated in the subsequent step, by decompression to 10 MPa and cooling down to 910 °C

<sup>f</sup> Note that in this step, MELTS was run under water-undersaturated conditions



$\text{SiO}_2/\text{Al}_2\text{O}_3$  ratios that are comparable with that of the evolved trachyte. Results are reported in Table 3 (MELTS step 2). The occurrence of such a process can be inferred by petrographic investigations (rounded or recrystallized feldspar crystals and reversely zoned plagioclase and feldspar compositions). A similar process has been suggested recently for the Teide-Pico Vejo complex (Tenerife, Canary islands; Wiesmaier et al. 2012) for which the evolved phonolite magmas are interpreted as the result of AFC involving nepheline-syenites from previous eruptions.

Fractional crystallization driven by decompression allows magma that is compositionally similar to the Schiappone sample MEGT 0312 to evolve further and reach a composition similar to that of the pre-MEGT phonolitic magmas. During this decompression stage,  $\text{Fe}_2\text{O}_3$ ,  $\text{TiO}_2$ ,  $\text{Na}_2\text{O}$  and  $\text{MnO}$  increase,  $\text{SiO}_2$  decreases, while  $\text{CaO}$  remains approximately constant or slightly decreases. Results, summarized in Table 3 (Melts Step 3 and 4), corroborate the hypothesis that feldspar assimilation may represent an essential step to give rise to a compositional lineage extending from latite to phonolite in volcanic systems fed by evolved alkaline magmas. In such systems, the occurrence at depth of large amounts of feldspar-rich cumulates left by earlier eruptions, as inferred on volcanological and petrological grounds for the Campi Flegrei caldera (D'Antonio 2011), makes the process of extensive assimilation of such material very likely to occur.

### Isotopic modelling

We modelled the effects of crustal contamination on latitic magma to reproduce the overall compositional and isotopic range displayed by the 75–50 ka BP volcanic rocks, in line with that recently proposed for Campi Flegrei magmas (e.g. D'Antonio et al. 2007; Di Renzo et al. 2011; Arienzo et al. 2011). We used the energy constrained assimilation and fractional crystallization (EC-AFC) model of Spera and Bohron (2001) and assumed that the latitic magma ascended to a depth of 8–12 km, where it was contaminated by continental crust. Results of the modelling are plotted as a blue curve in Fig. 7, and the parameters used are listed in Table 4. We assigned to the parental magma the composition of the Chiummano latite, and the Sr- and Nd-isotope compositions of the least radiogenic Tisichiello magma (Supplementary table). The average composition of the Hercynian continental crust of the Calabrian region was used as an assimilant (Rottura et al. 1991). The contamination: (1) increases the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio; (2) decreases the  $^{143}\text{Nd}/^{144}\text{Nd}$  ratio, and (3) decreases magma temperature from 1,250 to 1,096 °C. Both temperature values are in line with temperature ranges for latitic melts (e.g. Cannatelli et al. 2007) and with MELTS simulations (see above). At such T values, the modelled daughter magma

has geochemical and Sr isotopic compositions close to that of the trachytic Schiappone magma.

### Magmatic evolution through time

The slight compositional variation exhibited by the Sant'Angelo magmas through time (from more differentiated to less differentiated phonolite, back up to more differentiated phonolite) coupled with a quasi-constant  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio, suggests closed system differentiation of the magma batch (Figs. 5, 6). However, the Sant'Angelo eruption did not follow the typical pattern of eruptions fed by zoned magma chambers, characterized by withdrawal of progressively less evolved, deeper magma with time (e.g. Smith and Bailey 1966; Blake and Ivey 1986a, b). Instead the eruption was reversely chemically zoned and the more differentiated melt exited the chamber at the end of the eruption. Such zoning has been documented in a number of other volcanic deposits (e.g. Brown et al. 1998; Goff et al. 2014; Takahashi and Nakagawa 2013) and several reasons have been proposed for it. These include: complex magma reservoir structure involving several isolated reservoirs of variable composition that are ruptured and tapped during an eruption (e.g. Brown et al. 1998; Takahashi and Nakagawa 2013); non-standard eruption dynamics (e.g. Blake and Ivey 1986a, b); or eruptions supplied from different depths or regions of a magma chamber (Spera and Crisp 1981). We cannot yet constrain the mechanism to account for the Sant'Angelo zonation. The similarity in isotopic and chemical composition among the last-erupted Sant'Angelo magma (phonolite,  $^{87}\text{Sr}/^{86}\text{Sr}$  from 0.706557 to 0.706626;  $\text{CaO} = 1.02$  wt%; Zr from 1,054 to 1,080 ppm), the Mago magma (phonolite,  $^{87}\text{Sr}/^{86}\text{Sr} = 0.706537 \pm 6$ ;  $\text{CaO} = 0.99$  wt%; Zr = 1,037 ppm) and the first-erupted Olummo magma (phonolite,  $^{87}\text{Sr}/^{86}\text{Sr}$  from 0.706569 to 0.706519;  $\text{CaO}$  from 0.97 to 0.92 wt%; Zr from 1,058 to 970 ppm; Figs. 5, 6), suggests that these eruptions were fed by a compositionally homogeneous reservoir (M1 magma in Fig. 9). However, evidence of geochemical zoning is given by the chemistry of the Olummo samples (Figs. 5, 6), which vary from an early extruded differentiated phonolite (Zr = ~1,100 ppm;  $\text{CaO} = \sim 1.0$  wt%) to a late trachyte (Zr = ~400 ppm;  $\text{CaO} = \sim 1.6$  wt%), indicating a progressive withdrawal of less differentiated magmas with time. Olummo magmas also show a slightly wider range of  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (0.70657–0.70642), suggesting replenishment of the pre-Olummo magma chamber by a less radiogenic, less differentiated trachytic magma (M2 magma, Fig. 9) that mixed with the trachy-phonolitic resident magma prior to eruption (Fig. 6).

The magma withdrawn during the Tisichiello eruption also exhibits a range of both chemical and isotopic compositions, with a compositional and isotopic gap (Figs. 5,

**Table 4** Parameters the EC-AFC model

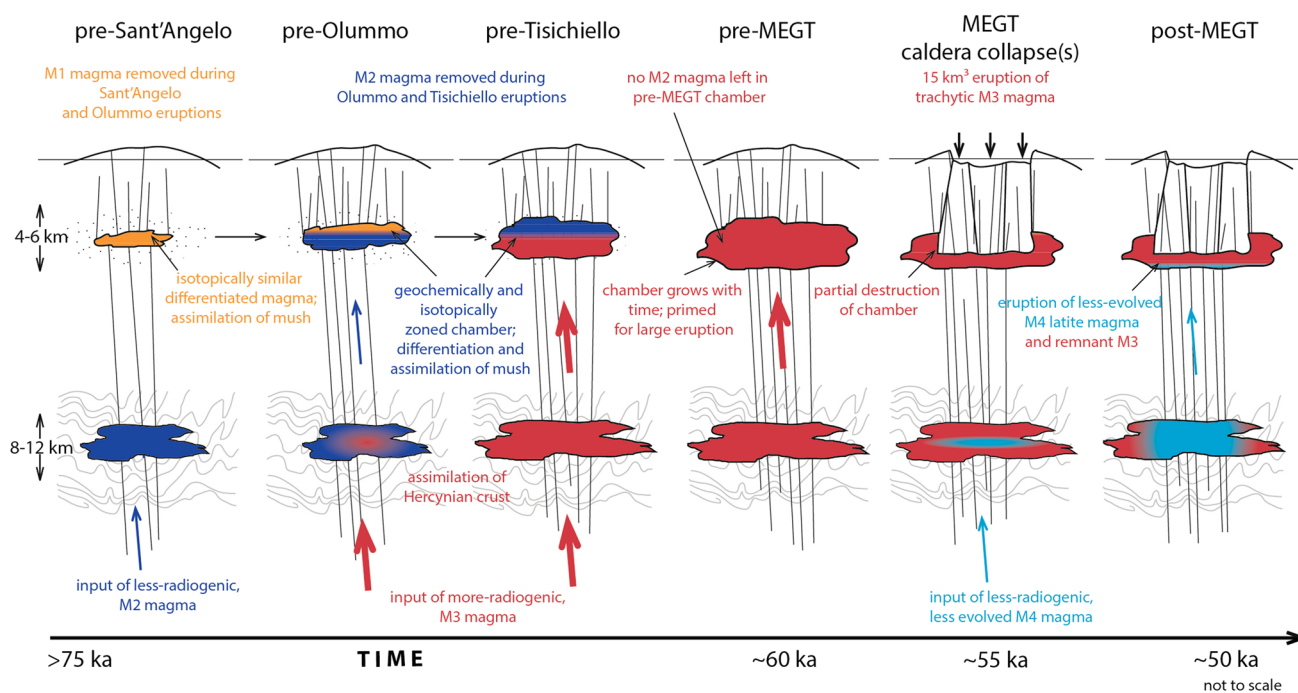
(a)			
Model parameters			
$T_{lm}$	1,250 °C		
$T_{m0}$	1,250 °C		
$T_{la}$	950 °C		
$T_{a0}$	400 °C		
$T_s$	700 °C		
$C_{pm}$	1,484 J/kg K		
$C_{pa}$	1,388 J/kg K		
$h_{cry}$	396,000 J/kg		
$h_{fus}$	270,000 J/kg		
$T_{eq}$	950 °C		
(b) End-member compositions			
Element	Sr (ppm)	Nd (ppm)	Th (ppm)
Concentration in magma	548	37	8
Bulk $D_0$	3.5	0.05	0.0001
Concentration in assimilant	298	36	5
Bulk $D_0$	0.02	0.04	0.0001
Isotope	$^{87}\text{Sr}/^{86}\text{Sr}$	$^{143}\text{Nd}/^{144}\text{Nd}$	Th (ppm)
Isotope ratio: magma	0.706420	0.512547	
Isotope ratio: assimilant	0.7140	0.5122	
Sr, Nd, Th of the contaminated magma (ppm)	356	46	14
Isotope ratio of the contaminated magma	0.706838	0.512537	
$T_{eq}$ equilibration temperature	$C_{p,m}$ magma isobaric specific heat capacity		
$T_{lm}$ magma liquidus temperature	$C_{p,a}$ assimilant isobaric specific heat capacity		
$T_{m0}$ initial magma temperature	$T_{la}$ wall rock liquidus temperature		
$h_{cry}$ crystallization temperature	$h_{fus}$ fusion enthalpy		
$T_{a0}$ initial wall rock temperature	$T_s$ wall rock solidus temperature		
$D_0$ distribution coefficient			

The distribution coefficients of Sr, Nd and Th for the evolving magma were assumed equal to 3.5, 0.05, and 0.0001, respectively, in agreement with the compatible behaviour of Sr and incompatible behaviour of Nd and Th during magma evolution. We assumed an initial temperature of 1,250 °C ( $T_m 0$ ) for the poorly differentiated magma, in line with the experimental estimates of Cannatelli et al. (2007). An initial temperature ( $T_{a0}$ ) of 400 °C was assigned to the crust. The liquidus temperature ( $T_{la}$ ) of the crust was assumed equal to 950 °C, its solidus temperature ( $T_s$ ) equal to 700 °C (e.g. Wyllie 1977). The equilibration temperature ( $T_{eq}$ ) was assumed equal to 950 °C. At that temperature, the modelled daughter magma has Sr = 356 ppm, Nd = 46 and Th = 14 ppm,  $^{87}\text{Sr}/^{86}\text{Sr} = 0.706838$  and  $^{143}\text{Nd}/^{144}\text{Nd} = 0.512537$ . The modelled daughter magma has geochemical and Sr isotopic composition close to that of the trachytic Schiappone magmas

6, 7). This strengthens arguments for the presence of a chemically and isotopically zoned magma chamber(s). Furthermore, the variation of the Tisichiello magma from trachyte to phonolite during the eruption suggests a complex withdrawal mechanism or magma reservoir arrangement similar to that inferred for the Sant'Angelo eruption. The early erupted Tisichiello magma is isotopically similar to the last-erupted Olummo magma, whereas the later, more differentiated Tisichiello magma is more enriched in radiogenic Sr and is separated from the earlier trachyte magma by a Sr isotopic gap (Fig. 6; Supplementary table). We explain such features as the result of an influx of a more

evolved and more radiogenic magma batch raised from the deep reservoir (8–12 km depth) that was more affected by crustal contamination with respect to the other pre-MEGT magmas.

The similarity of  $^{87}\text{Sr}/^{86}\text{Sr}$  among the last-erupted Tisichiello, Porticello and MEGT eruptions indicates that they were also fed by the magma of deep provenance more enriched in radiogenic Sr (M3 magma, Fig. 9). As the MEGT eruption extruded a large volume of magma (estimates by Brown et al. 2008 and Tomlinson et al. accepted, are in excess of 15 km<sup>3</sup> DRE) and isotopically distinct magma entered the system during the Tisichiello eruption (i.e. several thousand



**Fig. 9** Cartoon illustrating the sequence of arrival of different magma batches (M1–M4) into the deep and shallow magmatic plumbing system beneath Ischia for the studied period. Time and depth are not to scale

years before), the rate of magma production must have been high. Furthermore, the limited field data suggests that the size of the pre-MEGT eruptions increased through time from the Sant'Angelo to the Tisichiello eruption

The magma that fed the post-caldera-collapse La Roia eruption was trachytic, as was the last-erupted MEGT melt ( $\text{CaO} = \sim 2.2 \text{ wt\%}$ ;  $\text{Zr} = 290 \text{ ppm}$ ), although it was less enriched in radiogenic Sr. The magmas of the Capo Grosso and Chiummano eruptions were less differentiated although isotopically fairly similar to La Roia magma (Fig. 6). It follows that the La Roia, Capo Grosso and Chiummano eruptions tapped a new and isotopically distinct magma (M4 magma, Fig. 9) that rose from the deep reservoir after the MEGT eruption. The later change in composition from trachyte ( $\text{CaO} = 1.5 \text{ wt\%}$ ;  $\text{Zr} = 360 \text{ ppm}$ ) to latite ( $\text{CaO} = 4.6 \text{ wt\%}$ ;  $\text{Zr} = 230 \text{ ppm}$ ) during the Schiappone eruption, coupled with the slightly higher  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio (Figs. 5, 6) with respect to La Roia, Capo Grosso and Chiummano magmas, suggests that new, isotopically distinct magma arrived into the shallow system shortly before the Schiappone eruption.

## Conclusions

Volcanism on Ischia during the time span 75–50 ka BP was affected by variable magma differentiation processes at deep and shallow depths, such as replenishment, fractional

crystallization, mixing and assimilation. The Sr- and Nd-isotopic and geochemical variations through time suggest that the pre-MEGT volcanic activity was characterized by extrusion of magma batches poorly enriched in radiogenic Sr that mostly differentiated at shallow depth, where they attained a  $\text{TiO}_2\text{--Fe}_2\text{O}_3$ -rich phonolitic composition through assimilation of a feldspar-rich mush. In particular, during the oldest eruption studied (Sant'Angelo), a quasi-homogeneous magma was extruded (M1 magma, Fig. 9). This magma evolved at shallow depth mainly by fractional crystallization and assimilation. During the Olummo eruption, the magmatic system was recharged by a new, less differentiated and slightly less enriched in radiogenic Sr magma (M2 magma; Fig. 9) that also fed the first phases of the Tisichiello eruption. However, prior to or during the Tisichiello eruption, the magmatic system was recharged by a more enriched in radiogenic Sr magma (M3 magma; Fig. 9), potentially contaminated by both Hercynian crust at 8–12 km depth and by assimilation of feldspar mush at ca. 5 km depth. This magma later fed the climactic MEGT eruption. According to this scenario, the Tisichiello and Porticello events were precursors of the larger MEGT eruption. Volume estimations suggest that at least  $15 \text{ km}^3$  of magma accumulated in the shallow feeding system and assimilated residual feldspar cumulates prior to the MEGT eruption. After the MEGT eruption, influx of poorly evolved magma (M4 magma, Fig. 9) resulted in the chemical characteristics of the magmas that fed the later eruptions up to 45 ka BP.

The magmatic history and plumbing system structure explain the variation in magma composition from trachyte to phonolite as a rejuvenation of a sanidine-rich mush by arrival of new trachytic magma. In such a scenario, different P and T conditions are required for the stagnation and differentiation of the pre-MEGT and MEGT magmas, with respect to the post-MEGT magma. This process could have operated repetitively during the magmatic history of Ischia and could explain the eruption of latitic to phonolitic magmas with a wide range of Sr- and Nd-isotope compositions during the last 10 ka as the result of interaction between magmas and feldspar cumulates.

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## Appendix

### Fluid-mineral equilibria

Following Fabrizio et al. (2009), activity values for sanidine were taken by graphically reading the activity-composition plot in Waldbaum and Thompson (1969), those for magnetite by using the activity-composition expression of Woodland and Wood (1994) along the join magnetite-ulvospinel, and those for annite by the ionic model of Czamanske and Wones (1973). Calculations at 930 °C return  $f_{\text{H}_2\text{O}}$  values of 668 bars and 498 bars for the least (NNO + 0.75) and most oxidized (NNO + 1.44) trachyte (or trachy/phonolites), respectively. Furthermore, we computed the  $f_{\text{HF}}$  and  $f_{\text{HCl}}$  by using the expressions for chemical exchanges between fluoroannite/phlogopite–annite/fluorophlogopite and chloroannite/phlogopite–annite/chlorophlogopite pairs from Munoz (1984), based on the compositional deconvolution of Gunow et al. (1980) for the siderophyllite and annite components in biotites. We obtained  $f_{\text{HF}} = 0.4$  bars and  $f_{\text{HCl}} = 335$  bars at NNO + 0.75, and  $f_{\text{HF}} < 0.1$  bars and  $f_{\text{HCl}} = 18$  bars at NNO + 1.44.

### MELTS calculations

For the parental latite ascending from depth larger than 6 km, we lack a direct temperature estimate as well as information on volatile contents. By analogy with younger

than 3 ka Ischia latites (Moretti et al. 2013), temperatures around 1,200 °C and maximum water contents of 3–3.5 wt% can be suggested for the Chiummano latite. Of note is that the application of the CaO-in-glass thermometer of Cioni et al. (1998) returns a temperature of 1,060 °C.

MELTS calculations were thus initialized with a starting composition represented by the Chiummano latite (Supplementary table, Table 3) having 2.5 wt% of water,  $f_{\text{O}_2}$  conditions of QFM + 1 (i.e. NNO + ~0.6), the latter more reduced than biotites in poorly evolved trachytes, and initial P and T at 200 MPa and 1,200 °C, respectively. The crystallizing solid phases are those consistent with petrographic description (see also Table 3). A  $\nabla P/\nabla T$  gradient of 2.6 bar/°C was applied in such a way that 930 °C were reached at 130 MPa (5.5 km). In this P–T conditions, the poorly evolved trachyte magma forms. After crystallization of clinopyroxene, Ti-magnetite, plagioclase, minor olivine and biotite, the residual liquid (72 %) attains a composition close to MEGT 0308 trachyte sample (step 1 in Table 3). Interestingly, biotite starts crystallizing between 940 and 930 °C for the selected water content. An initial water content larger than 2.5 wt% would prevent crystallization of biotite.

The following step requires assimilation of felsic rocks to approach the composition of the MEGT 0312 sample; otherwise, fractional crystallization would only drive the system towards more silica-undersaturated compositions. The assimilant was composed by two feldspars (Na-sanidine and K-sanidine), whereas the assimilating magma was represented by the trachyte, derived by fractional crystallization from the parental latite at 930 °C. Table 3 in text reports the MELTS outcomes for the assimilation process, in which feldspars, initially at 600 °C, are isobarically and isothermally assimilated by the trachyte at 930 °C. Incremental assimilation of  $\text{P}_2\text{O}_5$ - and  $\text{TiO}_2$ -free feldspars gives results that are in line with the Schiappone trachy-phonolitic MEGT 0312 composition (step 2 in Table 3; Fig. 8), allowing a slight decrease of  $\text{P}_2\text{O}_5$  and  $\text{TiO}_2$  content (this latter favoured by concomitant precipitation of Ti-magnetite).

Alternatively, feldspar assimilation can involve directly the latite (1,200 °C), ascending between 130 and 100 MPa and infiltrating feldspar cumulitic residuals. If this is the case, during feldspar digestion, water is lost and dissolved  $\text{H}_2\text{O}$  falls down to 1 wt% in the final melt. In the absence of a detailed melt inclusion study, we will retain here the first scenario, involving feldspar assimilation from trachyte at 930 °C. It is also worth noting that during assimilation, fractionated solids, which amount are nearly equal in mass to that of assimilated feldspars, are dominated by feldspar  $\text{Or}_{60-67}$ . This implies that, after each assimilation event, feldspar cumulates are potentially left in place for future remobilization.

After feldspar assimilation, fractional crystallization drives again the residual liquid towards undersaturation, with SiO<sub>2</sub> reaching 55–56 wt% and Al<sub>2</sub>O<sub>3</sub> dropping inevitably, accompanied by a significant decrease of FeO<sub>tot</sub> and TiO<sub>2</sub>. Only a depressurization taking place right after assimilation allows to keep high silica and alumina contents, nearly constant FeO<sub>tot</sub> and to increase TiO<sub>2</sub>, consistent with observations (step 3 in Table 3).

After assimilation and during this final stage of isothermal decompression, biotite, two feldspars, spinel and apatite crystallize. This association is observed only by keeping QFM + 2. Lower oxygen fugacities promote olivine, whereas higher values do not allow Na<sub>2</sub>O to overcome K<sub>2</sub>O. Nepheline crystallizes at the very end ( $T = 910$  °C and  $P = 100$  bar; see MELTS step 4 in Table 3), when the residual liquid amounts to 3.5 wt%. However, liquid composition at that stage would be even more evolved than that in step 3 (Table 3).

Finally, we remark that MELTS cannot be used to investigate accessory phases such as sphene, apatite and Na-amphiboles, at least for the studied compositions, for which the role of fluids other than water (S-species, halogens) may have dramatic effects on the mixing properties of such solid solutions.

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