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Relation	

LETTER

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Sequential analysis of carbonaceous materials in Hayabusa-returned samples for the determination of their origin

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Abstract

Preliminary results of the analyses of five carbonaceous materials (particle size of approximately 50 μm) from the Hayabusa spacecraft sample catcher, including their texture, chemistry, and chemical/isotopic compositions, are summarized. The carbonaceous particles underwent sequential analysis using a series of microanalytical instruments located at several research institutes and universities. Collected particles were initially classified into four categories: two categories containing extraterrestrial silicate particles, one category containing metal and quartz particles consistent with contamination from the sample catcher or sample manipulation tools, and a final category containing carbonaceous particles. Analysis of this final category was the main focus of this study. Through examination of the carbonaceous materials, the appropriate analytical processes for sample transportation and handling were optimized to minimize sample damage and terrestrial contamination. Particles were investigated by transmission electron microscopy/scanning transmission electron microscopy, and Ca-carbonate inclusions were found in one particle. In a different particle, a heterogeneous distribution of silicon in a uniform C, N, and O matrix was found. Though further analysis is required for a strict determination of particle origin, the differences in the microstructure and elemental distribution of the carbonaceous particles suggest multiple origins.

Keywords: Hayabusa; Carbonaceous material; Sequential analysis; Microstructural observation by TEM/STEM

Findings

Introduction

More than 450 tiny particles have thus far been collected by the Hayabusa spacecraft sample catcher (Yada et al. 2014). These particles were initially classified into four categories based on their chemical composition obtained from analysis by a field emission scanning electron microscope (FE-SEM) with energy dispersion spectrometer (EDS). Particles classified as categories 1 and 2 were mainly composed of silicate materials and were confirmed as Itokawa regolith particles by the Hayabusa sample preliminary examination team (HASPET) (Ebihara et al. 2011; Nagao et al. 2011; Nakamura et al. 2011; Noguchi et al.

2011; Tsuchiyama et al. 2011; Yurimoto et al. 2011). Category 4 particles were defined as particles containing synthetic material originating from the sample catcher, the micro-manipulator, and the clean chamber (e.g., aluminum flake, quartz glass, and stainless steel). However, the origins of category 3 particles, which were defined as particles composed mainly of carbon, were still unknown. The first analysis of category 3 particles (hereafter, first examination) was performed in parallel with the preliminary examination of silicate materials (categories 1 and 2). The difficulties in handling and/or processing of these small carbonaceous samples without introducing terrestrial contamination resulted in the collection of insufficient information from the first examination to explore the origin of category 3 particles. Kitajima et al. (2011) and Naraoka et al. (2012) investigated the existence of organic materials on the surface of category 1 and 2 particles during preliminary

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examinations of these silicate materials. They concluded that organic material, such as that found in the category 3 particles, was not detected on the surface of category 1 and 2 particles. Thus, there is currently no information on the origin of category 3 particles.

After preliminary examination of the silicate material, new analytical facilities and members were introduced to the HASPET for reexamination of the category 3 particles (hereafter, second examination). In order to examine the origin of category 3 particles by utilizing different chemical and physical information, the handling processes of the carbonaceous particles were improved and multiple microanalyses were applied to single particles. In this paper, the newly developed analytical scheme for analysis of the category 3 particles, including sample processing and analytical techniques developed during the first and second examinations, are reported. In addition, the results of the second examination of the category 3 particles are briefly summarized.

Description of category 3 particles

Fifty-eight out of 459 particles were classified as category 3 by FE-SEM with EDS analysis at the Extraterrestrial Sample Curation Center of JAXA (ESCuC/JAXA). Six particles were lost during sample transfer from the SEM holder to the storage slide glass using an electrostatically controlled manipulation system (for a complete description, see Yada et al. 2014). Thus, 52 category 3 particles were stored in the clean chamber at ESCuC/JAXA under purified nitrogen (N_2) gas. Category 3 samples were found to be mainly composed of carbon (C), nitrogen (N), and oxygen (O), with trace amounts of fluorine (F) and sulfur (S). The sizes of the category 3 particles were similar to that of category 1 and 2 particles, ranging from 15 to 208 μm in diameter.

Figure 1 shows backscattered electron (BSE) images of selected category 3 particles. From the BSE images, the category 3 particles were further divided into three sub-categories, types 1 to 3, based on their texture and chemical composition. Type 1 was defined as blocky particles, which were found to have irregular shapes and small white grains on their surfaces; the grains were found to be composed of silicate, iron sulfide, stainless steel, and aluminum (Al). These blocky type 1 particles were the largest subcategory within category 3 (38 out of 52 particles). The silicate grains were found to be olivine, pyroxene, plagioclase, iron sulfide, and potassium feldspar. These minerals were also components of category 1 and 2 particles. Therefore, it was assumed that these small silicate grains were also probably fragments of Itokawa particles. The small surface grains of stainless steel and Al were most likely contaminants from the sample catcher or the clean chamber. Type 1 particles also often included chloride (Cl) salts such as sodium

chloride (NaCl), potassium chloride (KCl), and calcium chloride ($CaCl_2$).

Type 2 particles were defined by their fibrous shape, and 6 out of 52 category 3 particles were defined as type 2. Type 3 particles were deemed faint type, having a very smooth surface marked by very few tiny surface particles. All type 3 particles had an N peak in their EDS spectrum, and 8 out of 52 particles were classified as this particle type.

The differences in the texture and chemical composition of each defined particle type within category 3 might indicate a different origin for each particle type. Alternatively, these differences may indicate that these particles were initially from the same source but underwent different irradiation processing or other alterations before their recovery from the sample catcher. Table 1 shows the characteristics of all category 3 particles. Pictures and other information on category 3 particles are available in the sample catalogue of Hayabusa-returned samples (<http://hayabusao.isas.jaxa.jp/catalog/cat3/>).

Sample preparation and analysis method

For the preliminary examinations of category 1 and 2 particles (i.e., silicate materials), the samples were fixed on the top of a 5- μm fiber with epoxy resin or glycol phthalate (e.g., Nakamura et al. 2011) to facilitate sample handling and transfer. However, these adhesive materials, which included water and ethanol, could be a source of potential contamination and/or damage during the handling of category 3 (carbonaceous) particles. As such, category 3 samples should be handled in a completely dry environment, without adhesive, in order to avoid sample damage and contamination. Thus, for the first examination, category 3 particles were transferred using a diamond or quartz sample-holder under a purified N_2 atmosphere. However, the samples were difficult to handle using the manipulator without any adhesive material or liquid, resulting in the loss of a particle (RA-QD02-0008) during handling.

For the second examination, samples were fixed on 0.3-mm-thick metal plates (gold (Au) or indium (In)) on the SEM sample holder (Figure 2). The samples were pressed onto the Au or In plates by 5-mm-thick sapphire windows installed in a special pressure apparatus so that samples could be observed through the glass during pressing. Au or In were chosen as sample substrates to minimize potential contamination and charging artifacts during SEM, NanoSIMS (secondary ion mass spectroscopy), and time-of-flight (ToF)-SIMS analyses.

The subset of category 3 samples allocated for the first and second preliminary examinations were RA-QD02-0008, RA-QD02-0120, RA-QD02-0180, RB-QD04-0001, RB-QD04-0037-01, and RB-QD04-0047-02 (Figure 1).

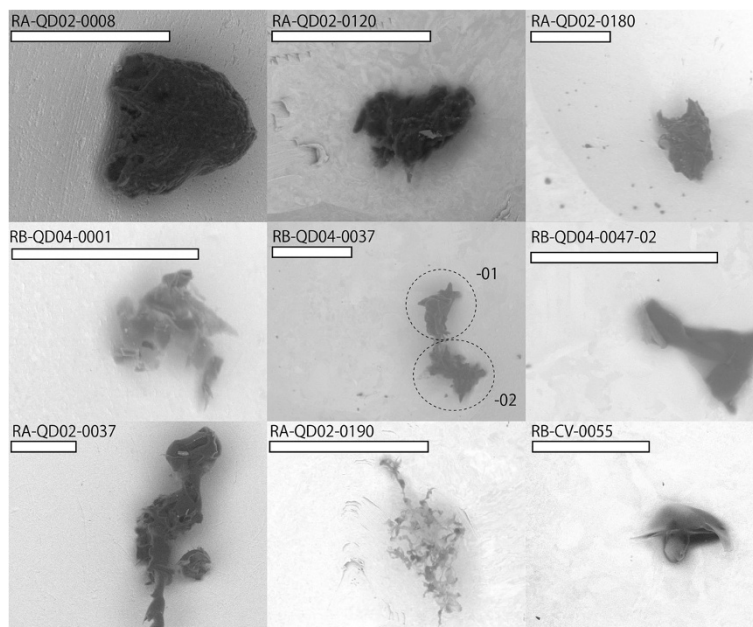


Figure 1 Backscattered electron images of category 3 particles. White bars show 50 μm . Particles allocated to the first and second examinations are shown in the top and middle rows of the figure. RA-QD02-0190 is classified as fibrous type (type 2), and RB-QD04-0001 and RB-CV-0055 are classified as faint type (type 3).

All samples were first analyzed by FE-SEM-EDS (Hitachi S-4300E/N, SU6600, ESCuC/JAXA, Hitachi, Tokyo, Japan) during the characterization process required for nominal curation, referred to as 'initial description' (Yada et al. 2014). This initial description separated the returned samples into the four categories described above.

Four particles, RA-QD02-0008, RA-QD02-0120, RB-QD04-0001, and RB-QD04-0047-02, were previously chosen to be analyzed in the first examination and, thus, had been exposed to the terrestrial atmosphere. Two of these particles, RA-QD02-0008 and RB-QD04-0001, were analyzed by Fourier transform infrared spectroscopy (FT-IR) and Raman spectroscopy in the first examination (Kitajima et al. 2014), while RB-QD04-0047-02 and RA-QD02-0120 were returned to ESCuC/JAXA without any analyses from the first examination. Thus, these samples were minimally affected by brief atmospheric exposure during the first examination. Samples allocated to the first examination were stored for 6 months in a purified N_2 desiccator after analysis and then allocated to the second examination.

Except for RA-QD02-0008, particles allocated to the first examination were pressed on an Au plate under terrestrial atmosphere at the beginning of the second examination. The two new particles allocated solely to the second examination, RA-QD02-0180 and RB-QD04-0037-01, were pressed onto In plates under

purified N_2 gas. Several types of microanalysis instruments located in different laboratories (NanoSIMS ion microprobe: CAMECA NanoSIMS 50 L at Kochi Institute for Core Sample Research/JAMSTEC, ToF-SIMS: ULVAC-PHI TRIFT III at Nagoya University, FT-IR: JASCO IRT-5000 at ESCuC/JAXA and Perkin-Elmer Spectrum One at Kyushu University, Raman spectroscopy: JASCO NRS-5100 at ESCuC/JAXA and JEOL JRS System 2000 at Kyushu University) were used on the pressed samples. If evidence of extraterrestrial material (i.e., isotopic anomalies of H, C, and N in organics) was found, the second examination was immediately closed and a more detailed analysis commenced.

After the first analyses of the pressed samples, ultrathin sections (UTSs) of RA-QD02-0120 and RB-QD04-0047-02 ($20 \mu\text{m} \times 1 \mu\text{m} \times 100 \text{nm}$) were extracted by a focused ion beam (FIB, Hitachi FB2200, ESCuC/JAXA). These UTSs were mounted on copper (Cu) grids by tungsten deposition, and the Cu grids were fixed on polyurethane elastomers during sample transfer. X-ray-absorption near edge structure (XANES) (ALS/Berkley and UVSOR/IMS) and (scanning) transmission electron microscopy ((S)TEM, JEOL JEM-2800, JEM-2100 F, JEOL Ltd., Tokyo, Japan) were examined to further analyze the UTSs. Characteristics of the samples and a summary of the applied analyses are reported in Table 2. Details of NanoSIMS (Ito et al. 2014), ToF-SIMS (Naraoka et al. 2014), FT-IR and Raman spectroscopy (Kitajima et al.

Table 1 List of category 3 particles

Name	Size ^a (μm)	Type	Elements ^b	Surface grains
RA-QD02-0008	50	Type1: blocky	C, N, O	Al
RA-QD02-0012	100	Type1: blocky	C, O	FeS
RA-QD02-0018	50	Type1: blocky	C, O, Cl	Al, ol, lpx
RA-QD02-0037	60	Type1: blocky	C, O	ol, lpx
RA-QD02-0040-01	30	Type1: blocky	C, O, Si	Lpx
RA-QD02-0078	100	Type1: blocky	C, O, Cl, F, Si	Lpx
RA-QD02-0091	30	Type1: blocky	C, O, Si, Cl	
RA-QD02-0120	26	Type1: blocky	C, O	SUS
RA-QD02-0134	21	Type1: blocky	K, Cl	
RA-QD02-0180	55	Type1: blocky	C, O, K, Na, Cl	
RA-QD02-0181	55	Type1: blocky	C, N, O, Na, K, Cl	
RA-QD02-0182	46	Type1: blocky	C, N, O, Na, K, Cl	
RA-QD02-0183	38	Type1: blocky	C, N, O, Na, K, Cl	
RA-QD02-0190	58	Type2: fibrous	C, N	
RA-QD02-0222	121	Type2: fibrous	C, O, Mg Si	
RB-CV-0005	37	Type1: blocky	C, Na, Cl	
RB-CV-0006	27.9	Type1: blocky	C, N, O	
RB-CV-0007	43.1	Type1: blocky	C, N, O	
RB-CV-0008	56.1	Type1: blocky	C, N, O	
RB-CV-0012	73.2	Type2: fibrous	C, N, O	
RB-CV-0017	51.7	Type1: blocky	C, N, O	
RB-CV-0019	32.6	Type3: faint	C, N, O	
RB-CV-0020	68.6	Type1: blocky	C, O	
RB-CV-0021	49.3	Type2: fibrous	C, O	
RB-CV-0027	59	Type1: blocky	C, N, O	
RB-CV-0029	86	Type1: blocky	C, O, Na, Cl	Pl
RB-CV-0031	61	Type2: fibrous	C, N, O	
RB-CV-0032	59	Type3: faint	C, N, O	
RB-CV-0035	25	Type1: blocky	C, N, O, Na, Cl	hpx, Al
RB-CV-0041	49	Type1: blocky	C, N, O, Na, Cl	
RB-CV-0047	48	Type1: blocky	C, O, Al, Ca, Si	
RB-CV-0049	37	Type3: faint	C, N, O	
RB-CV-0052	40	Type1: blocky	C, O, Na, Cl	
RB-CV-0055	45	Type3: faint	C, N, O, Al	
RB-CV-0065	33	Type1: blocky	C, N, O	
RB-CV-0066	29	Type1: blocky	C, N, O, Na, Cl	
RB-CV-0068	55	Type3: faint	C, N, O	Ol
RB-CV-0074	26	Type1: blocky	C, N, O, Na, Cl	Al
RB-CV-0077	100	Type3: faint	C, N, O	
RB-CV-0078	208	Type2: fibrous	C, O	
RB-CV-0079	68	Type1: blocky	C, N, O	K-feld
RB-CV-0080	83	Type1: blocky	C, N, O	K-feld
RB-QD04-0001	19	Type3: faint	C, N, O	ol, Al
RB-QD04-0031	64	Type1: blocky	C, O, Al	

Table 1 List of category 3 particles (Continued)

RB-QD04-0035	59	Type1: blocky	C, N, O, Na, K, Cl, Al	
RB-QD04-0037-01	48	Type1: blocky	C, N, O	Ol
RB-QD04-0037-03	15	Type3: faint	C, N, O	
RB-QD04-0047-01	66	Type1: blocky	C, O	
RB-QD04-0047-02	28	Type1: blocky	C, O	
RB-QD04-0048	56	Type1: blocky	C, O	Pl
RB-QD04-0052	96	Type1: blocky	C, F, O, Al, Ti	
RB-QD04-0078	19	Type1: blocky	C, O	Ol

^aLength of the major axis; ^bfound in FE-SEM-EDS analysis. ol, olivine; pl, plagioclase; K-fld, K feldspar; hpx, high-Ca pyroxene; lpx, low-Ca pyroxene; SUS, stainless steel; FeS, iron sulfide.

2014), and XANES (Yabuta et al. 2014) analyses are described in papers in this issue.

Both pressed samples and UTs were transferred by parcel delivery service to each institute or university within a N₂-purged environment. Prior to the analysis of category 3 particles, analysis procedures, including sample transfer, were tested using carbonaceous particles collected at the spacecraft and fairing assembly building (SFA) of the Tanegashima Space Center/JAXA (hereafter SFA particle), insoluble organic matter (IOM) of the Antarctic meteorite A-881458 (CM2), and synthetic Mg-Fe-Ni olivine particles. All test particles were in the same size range as the Hayabusa-returned samples. NanoSIMS isotope analysis of H, C, and N measured terrestrial isotopic ratios for the SFA particles and extraterrestrial isotopic signatures for the IOM from A-881458 (Ito et al. 2014). From this data, it was confirmed that the sample handling system and transportation process were suitable for these organic samples.

Evaluation of possible sample damage and contamination during analysis

Evaluation of sample damage by SIMS analysis The purpose of this study was to apply several microanalytical techniques to a single particle to suppress any errors caused by the instruments used during the sequential analysis. In order to investigate the change of a sample's isotopic ratio as a function of sample analysis sequence order, RB-QD04-0047-02 was analyzed by NanoSIMS before and after ToF-SIMS analysis. The isotopic values of the samples from NanoSIMS analysis obtained prior to ToF-SIMS analysis were $\delta D = +103\text{‰}$ (vs SMOW), $\delta^{13}C = +3\text{‰}$ (vs PDB), and $\delta^{15}N = -4\text{‰}$ (vs air), and the values obtained after ToF-SIMS analysis were $\delta D = +177\text{‰}$, $\delta^{13}C = +14\text{‰}$, and $\delta^{15}N = +6\text{‰}$ (Ito et al. 2014). The isotopic ratios tended to be slightly enriched in the heavier isotope after ToF-SIMS analysis. However, the change in δD was smaller than the expected isotopic anomalies found in the organic material of ordinary chondrites (δD of approximately +4000‰, e.g., Alexander et al.

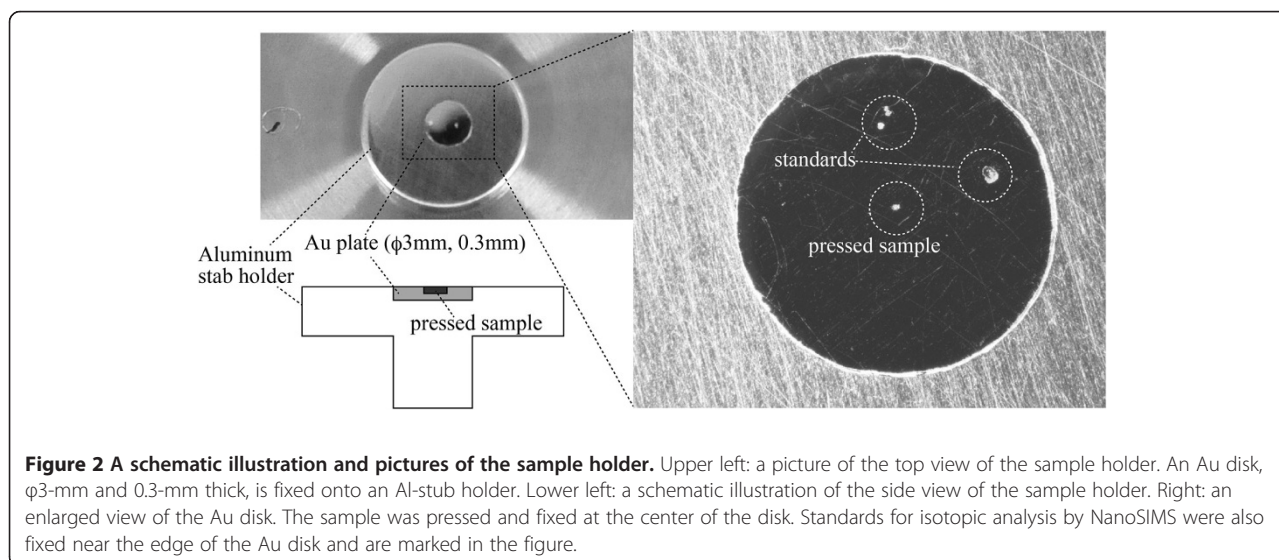


Figure 2 A schematic illustration and pictures of the sample holder. Upper left: a picture of the top view of the sample holder. An Au disk, $\phi 3$ -mm and 0.3-mm thick, is fixed onto an Al-stub holder. Lower left: a schematic illustration of the side view of the sample holder. Right: an enlarged view of the Au disk. The sample was pressed and fixed at the center of the disk. Standards for isotopic analysis by NanoSIMS were also fixed near the edge of the Au disk and are marked in the figure.

Table 2 Details of the samples and analysis

Sample name	Size ^a (μm)	Pressed on	Type	Elements ^b	Analysis flow
RA-QD02-0008	50	(Lost)	Blocky	C, N, O, Al	FT-IR, Raman (Lost)
RA-QD02-0120	26	Au	Blocky	C, O	NanoSIMS, FT-IR, Raman, XANES, TEM
RA-QD02-0180	55	In	Blocky	C, O, K, Na, Cl	ToF-SIMS, NanoSIMS, Raman
RB-QD04-0001	19	Au	Faint	C, N, O	NanoSIMS, FT-IR, Raman
RB-QD04-0037-01	48	In	Blocky	C, N, O	FT-IR, ToF-SIMS
RB-QD04-0047-02	28	Au	Blocky	C, O	NanoSIMS, FT-IR, Raman, ToF-SIMS, NanoSIMS, XANES, TEM
S01 (SFA particle)	30	Au	Blocky	C, O	NanoSIMS
IOM1 (A-881458, CM2)	50	Au	Blocky		NanoSIMS, FT-IR, Raman, XANES, TEM
NI1 (synthetic Mg-Fe-Ni olivine)	50	Au	-	-	XANES, TEM

^aLength of the major axis; ^bfound in FE-SEM-EDS analysis.

2007). Therefore, measurement of a hydrogen isotopic anomaly should be possible after ToF-SIMS analysis. Contrastingly, the results of ToF-SIMS analysis were seriously altered by interference from Cesium (Cs), the primary ion beam of NanoSIMS, and by Au deposits from sputtering during the NanoSIMS analysis.

The effect of the ToF-SIMS analysis on FT-IR analysis was also investigated. Figure 3a shows the difference in the FT-IR spectrum of RB-QD04-0037-01 before and after ToF-SIMS analysis. The measured spectrum was greatly altered by ToF-SIMS analysis: a major peak between 2,500 and 2,000 cm⁻¹ disappeared, and the overall spectrum was significantly smoothed.

Evaluation of contamination through the FIB UTS extraction process Contamination transferred by the FIB from previous sample residue was evaluated by pressing a 50-μm fragment of the IOM of A-881458 (CM2) onto an Au plate and extracting a UTS from this pressed sample with the FIB. Before and after fabricating a UTS from the IOM, UTSs of synthetic Mg-Fe-Ni olivine were also extracted. The synthetic Mg-Fe-Ni olivine UTSs were analyzed by O-XANES. No changes were observed in the synthetic Mg-Fe-Ni olivine peak signals measured by O-XANES before and after the IOM fabrication (Figure 3b). Thus, the transfer of the carbonaceous material via the FIB was likely negligible and should not affect subsequent analyses of the carbonaceous material.

Evaluation of sample damage by electron beam of TEM/STEM on XANES spectra Yabuta et al. (2014) reported a change in the C-XANES spectra of RA-QD02-0120 before and after TEM analysis attributed to sample damage by the electron beam. This effect was reported in a previous study (e.g. De Gregorio et al. 2010) and should be considered in the sequential analysis.

Reconstruction of analytical flow for carbonaceous materials Through the discussed analyses, the flow of the sequential analysis was constructed using FT-IR, ToF-SIMS, NanoSIMS, Raman spectroscopy, XANES and TEM/STEM to minimize sample contamination during analysis (Table 3). The reconstructed flow was applied to the analysis of the previously unanalyzed RB-QD04-0037-01 and RA-QD02-0180 samples allocated to the second examination.

Analyses applied to each particle in the first and second examinations

Sample names, categories, sizes, observed EDS peaks, and techniques used for this study are summarized in Table 2. RA-QD02-0008 was analyzed only in the first examination. FE-SEM-EDS analysis identified C, N, O, and Al peaks in the sample. After the FT-IR and Raman spectroscopy measurements, the sample was lost during preparation for ToF-SIMS analysis.

RA-QD02-0120 was analyzed by NanoSIMS, FT-IR, Raman spectroscopy, XANES, and TEM in the second examination. This particle was analyzed using FT-IR and Raman spectroscopy after NanoSIMS analysis; thus, the spectra were probably altered (smoothed) by the effect described previously.

RA-QD02-0180 was only analyzed during the second examination and was found to have NaCl and KCl peaks as well as C and O peaks in the FE-SEM-EDS spectrum. The particle was carefully pressed onto an In plate in a purified N₂ environment and analyzed by ToF-SIMS. In order to avoid damage to the chlorides in the sample from atmospheric exposure, FT-IR analysis was not conducted on this sample. Currently, atmospheric exposure is unavoidable during FT-IR analysis at ESCuC/JAXA, though improvements to this system are being planned to address this issue.

RB-QD04-0001 was analyzed during the first examination by FT-IR and Raman spectroscopy. In the second examination, the particle was analyzed by NanoSIMS, FT-

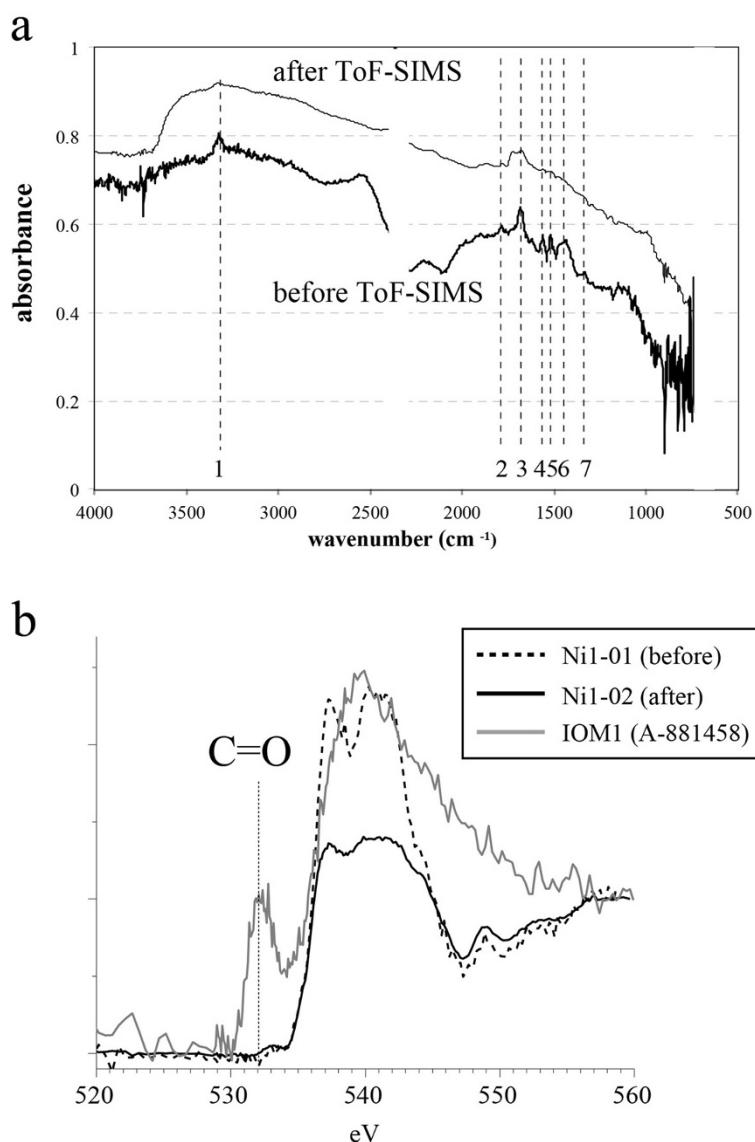


Figure 3 Results of the tests for sample damage by ToF-SIMS and for FIB cross-contamination. (a) Change of the FT-IR spectrum before and after ToF-SIMS analysis. Dashed lines show where the peak positions appear in the spectrum before ToF-SIMS analysis: (1) $3,300\text{ cm}^{-1}$ - N-H stretching; (2) $1,800$ and (3) $1,680\text{ cm}^{-1}$ - C = O stretching; (4) $1,560$ and (5) $1,510\text{ cm}^{-1}$ - N-H deformation; (6) $1,450\text{ cm}^{-1}$ - N-H bending; (7) $1,330\text{ cm}^{-1}$ C-N stretching. **(b)** O-XANES spectrum of UTSs of a synthetic Mg-Fe-Ni olivine sample before and after FIB fabrication of the IOM (A-881458, CM2) UTS. Vertical axis shows absorbance normalized by direct beam. A dashed line at 532 eV shows C-O bond absorption from IOM. Ni01-01 and Ni01-02 are the spectra from the UTSs before and after the FIB fabrication of IOM.

IR, and Raman spectroscopy. This was the only the particle classified into the category 3 type 3 (faint type) particles among the particles previously analyzed. During the NanoSIMS analysis, the pressed sample on the Au plate was almost completely sputtered away due to its small size. Thus, the sample could not be analyzed further.

RB-QD04-0037-01 was also only analyzed during the second examination. FE-SEM-EDS analysis showed C, N, and O, and olivine inclusions on the surface of the

particle. The sample was pressed onto an In plate and analyzed by FT-IR and ToF-SIMS.

The FE-SEM-EDS spectrum of RB-QD04-0047-02 was found to contain only C and O peaks. This particle was also analyzed by NanoSIMS, FT-IR, Raman spectroscopy, ToF-SIMS, XANES, and TEM/STEM. This particle was very hard compared to other particles, as its shape did not change during handling or after pressing onto the Au metal substrate. In comparison, other particles were deformed or crushed during this pressing process.

Table 3 Reconstructed analytical flow and effects on other measurements

Order	Method	Data	Sample damage
1	FT-IR	Molecular structure	-
2	ToF-SIMS	Elemental mapping and analysis of molecular species	Smoothing of FT-IR spectra, moderate change of isotope ratio of NanoSIMS analysis (Ito et al. 2014)
3	NanoSIMS	Isotope ratio and mapping	Large disturbance of ToF-SIMS result by Cs + implantation (Naraoka et al. 2014)
4	Raman spectroscopy	Molecular structure	Thermal damage by laser (not evaluated in this paper)
5	XANES	Molecular structure	Sample destruction by FIB sectioning
6	TEM/STEM	Nanoscale texture and elemental mapping	Sample destruction by FIB sectioning, electron beam damage on molecular structure (Yabuta et al. 2014)

Results of TEM/STEM analysis

Figure 4 shows the TEM/STEM images of RA-QD02-0120 and RB-QD04-0047-02. Examination of the UTS of RA-QD02-0120 revealed Ca-carbonate inclusions up to 300 nm in diameter. The matrix of the carbonaceous material showed uniform C, N, and O distributions. In the UTS of RB-QD04-0047-02, no distinct inclusions were apparent. In addition, only a heterogeneous distribution of Si was observed in the carbonaceous material from the EDS map.

Discussion and future work

Sequential analysis of category 3 particles allowed us to measure a number of chemical and structural characteristics, which could be related to the origin of these particles. No isotopic anomalies were found, suggesting these particles were not of extraterrestrial origin (Ito et al. 2014). However, as this finding does not immediately discount extraterrestrial origin, further analysis of this category for terrestrial contaminants was deemed necessary. FT-IR and Raman spectroscopy spectra of all samples showed differences between these category 3 samples and the meteoritic IOM (Kitajima et al. 2014), also suggesting that the category 3 particles are of terrestrial origin.

The UTS of the IOM of A-881458 was investigated using TEM/STEM (Figure 5). A nano-globule was found in the pressed IOM similar to previous findings from extraterrestrial matter (e.g., Nakamura-Messenger et al. 2002; Garvie and Buseck 2004; Nakamura-Messenger et al. 2006; Matsumoto et al. 2013). Thus, particles of extraterrestrial origin may be identified by the presence of these globules in the UTSs. These globules were not detected in any of the UTSs of category 3 particles. Together, the lack of an extraterrestrial isotopic signature, the difference between the FT-IR and Raman spectroscopy spectra of category 3 particles and meteoritic IOM, and the lack of nano-globules indicate a low possibility that category 3 particles are of extraterrestrial origin.

A Si distribution was detected in RB-QD04-0047-02 by TEM observation indicating that these particles are probably not of biological origin. ToF-SIMS analysis of RB-QD04-0037-01 also showed Si distribution over a

wide area on the surface of the pressed sample (Naraoka et al. 2014). Thus, silicon rubber was considered as a possible origin of the sample. However, FE-SEM-EDS showed the presence of N in RB-D04-0037-01, which was not apparent in the FE-SEM-EDS spectrum of a silicon-rubber standard sample. A uniform distribution of N, C, and O was also found utilizing a STEM-EDS elemental map of RA-QD02-0120 and RB-QD04-0047-02. The nitrogen signature in these samples was also detected by XANES analyses (Yabuta et al. 2014). These results suggest that the origin of these particles may not be silicon rubber or that the properties of these particles were altered from their original material by heat or cosmic ray irradiation during space flight.

Ca-carbonate inclusions were detected on the UTS of RA-QD02-0120 by TEM/STEM. However, Si was not found in the STEM-EDS map of the particle, indicating that this particle is different in origin from both RB-QD04-0037-01 and RB-QD04-0047-02. Though their exact origins are still uncertain, this data suggests that there may be several different category 3 particle sources. Due to its small mass, there is insufficient information on RB-QD04-0001 for the determination of its source.

In order to determine the origin of category 3 particles, possible contaminants, such as fluoro-rubber, silicon rubber, and Vectran, as well as extraterrestrial organic matter, must be characterized using the same analytical techniques used to analyze the Hayabusa-returned samples. Vectran, a polyarylate fiber, was used in the outer cover of the Hayabusa spacecraft sampler horn. Fluoro-rubber is used for the gloves of the clean chambers where the Hayabusa-returned samples are stored. Silicon rubber is generally used for the seal material of clean rooms. Particles collected from contamination coupons, which were exposed in the clean room during examinations of the Hayabusa 2 spacecraft (unpublished data), will also be analyzed. Examination of the contamination coupons will allow examination of potential contaminants prior to launch of the Hayabusa 2 spacecraft. By comparing the analytical results of potential contaminants and category 3 particles, the precise origin of category 3 particles may be determined. The effect of the long-term exposure of

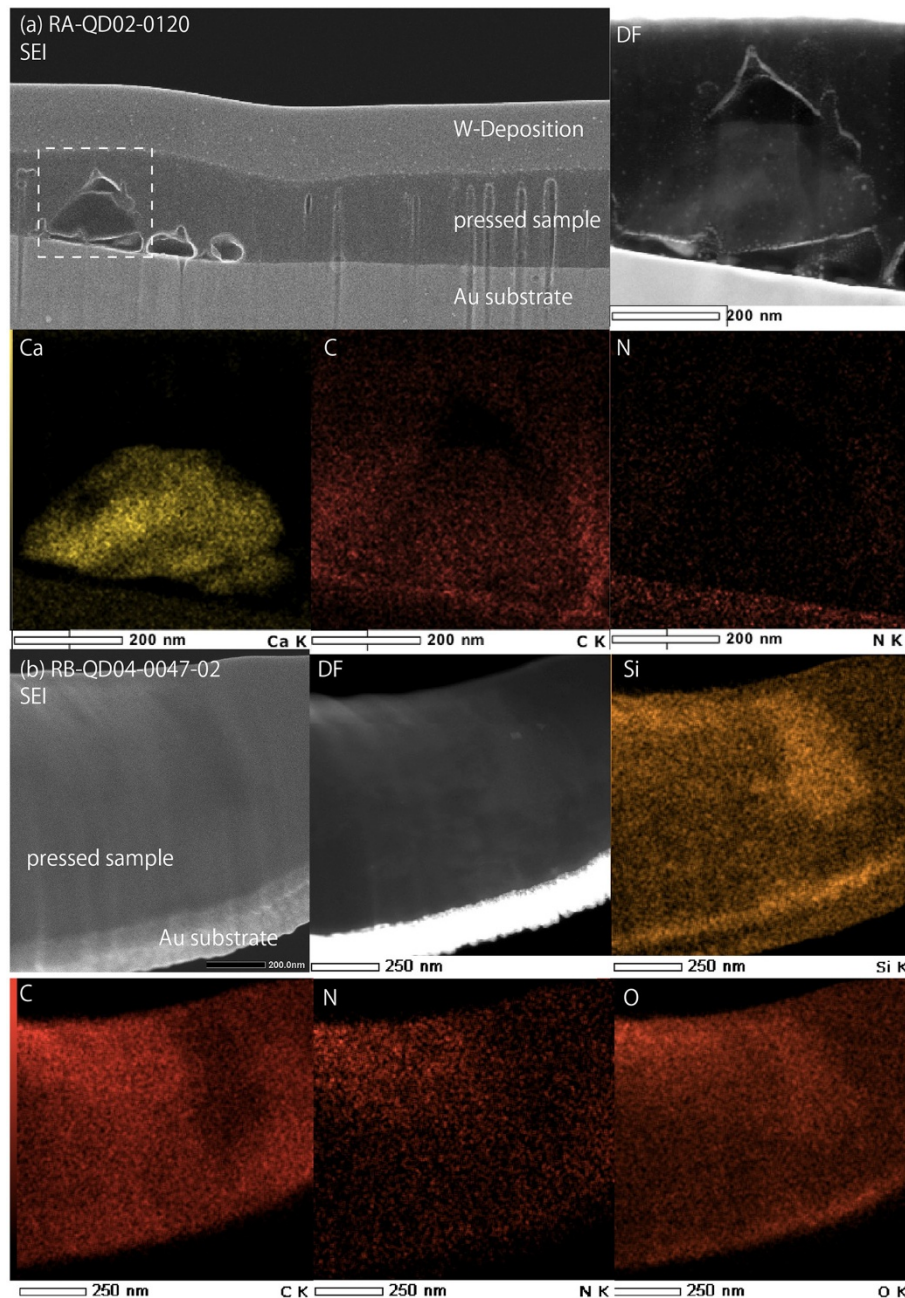


Figure 4 Images obtained from TEM/STEM-EDS (a, b). A dashed box in the secondary electron image (SEI) of RA-QD02-0120 shows the area of dark field (DF) image; and the EDS images show color elemental maps of Si, N, O, C, and Ca. More intense colors correspond to higher concentrations of these elements. Vertical artifacts in the SEI image are from FIB fabrication of the UTS.

category 3 particles to cosmic rays during operation of the spacecraft, which might cause a change in their structure and chemical composition, should also be considered. Collected samples and potential contaminants should also be investigated both before and after irradiation by ultraviolet rays or electron beam for further confirmation of sample origin.

Summary

The possible origin of the carbonaceous particles in Hayabusa-returned samples was investigated by reconstructing the preliminary examination team for category 3 particles and the flow. Category 3 particles were divided into different types based on their chemical composition and determined that there could have been multiple

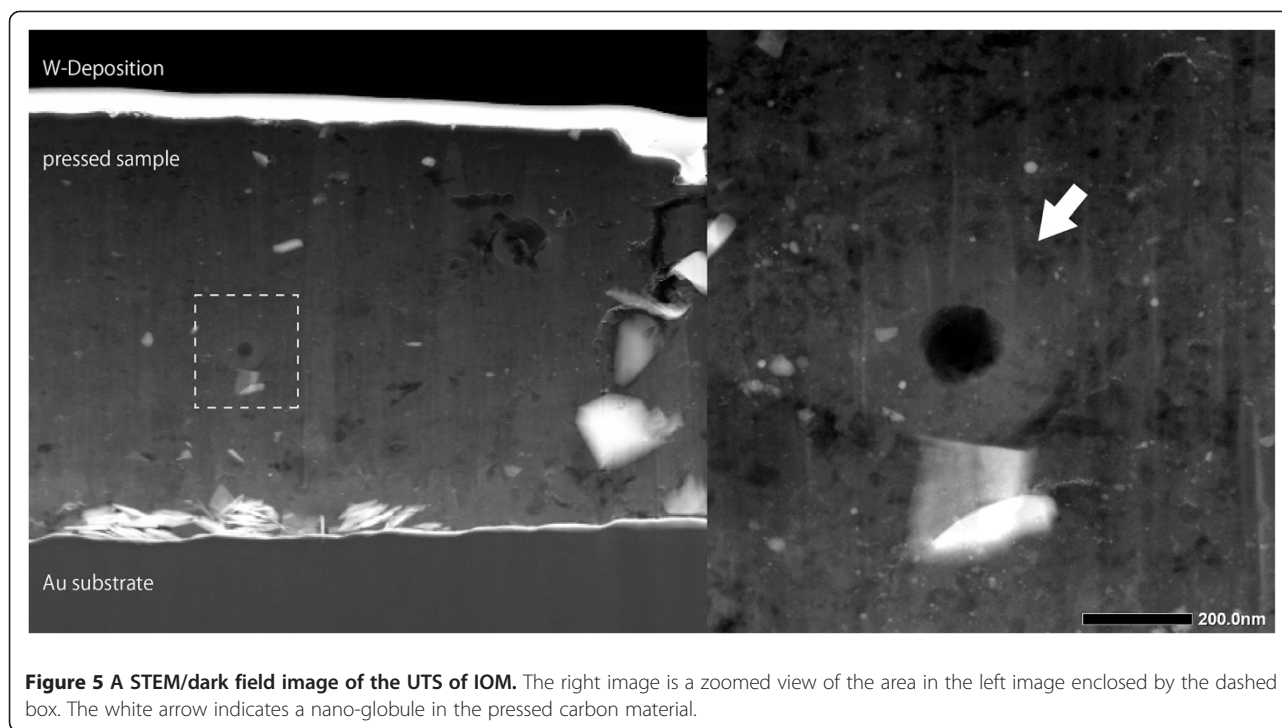


Figure 5 A STEM/dark field image of the UTS of IOM. The right image is a zoomed view of the area in the left image enclosed by the dashed box. The white arrow indicates a nano-globule in the pressed carbon material.

sources of the category 3 particles, although some category 3 particles may have the same origin. In future work, potential contaminants will be investigated using the same analysis flow developed in this study to minimize sample contamination during the optimized second examination.

The new optimized sample-handling system developed through these examinations, including transportation between institutes and the sequence of the analytical flow, will be important not only for the Hayabusa project, but also for sample handling, preparation, and analysis during future sample return missions such as Hayabusa 2 and Osiris-REX.

Abbreviations

ESCuC: Extraterrestrial Sample Curation Center; BSE: backscattered electron; EDS: energy dispersion spectrometry; FE-SEM: field emission scanning electron microscopy; FIB: focused ion beam; FT-IR: Fourier-transform infrared spectroscopy; HASPET: Hayabusa sample preliminary examination team; IOM: insoluble organic matter; NanoSIMS: nano-secondary ion mass spectrometry; (S)TEM: (scanning) transmission electron microscopy; SFA: spacecraft and fairing assembly building; ToF-SIMS: time-of-flight secondary ion mass spectrometry; UTS: ultra-thin section; XANES: X-ray-absorption near-edge structure.

Competing interests

The authors declare that they have no competing interests.

Authors' contributions

MU contributed to sample preparation, TEM analysis, image data interpretation, and manuscript preparation. IO operated TEM analysis. MU, YK, TY, and YI carried out the sample handling, electron beam analyses, and interpretation. MI, HN, FK, HY, YT, HM, YK, TO, and MA participated in the design of the research and interpretation. All authors read and approved the final manuscript.

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