

**LEPTON-FLAVOUR-VIOLATION IN SUSY MODELS
WITH AND WITHOUT R-PARITY ^a**

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We discuss Lepton-Flavour-Violating phenomena such as $\mu \rightarrow e\gamma$, $\mu \rightarrow eee$, and $\mu \rightarrow e$ conversion in nuclei in SUSY models with and without R-parity. We stress that experimental searches for all the LFV processes are important to distinguish between the different models.

Recently, the atmospheric neutrino experiment SuperKamiokande has announced evidence for non-zero neutrino mass, which indicates the existence of physics beyond the Standard Model (SM). Generally in models that accommodate neutrino oscillations, lepton-flavour-violating phenomena (LFV) can occur not only in the neutrino, but also in the charged-lepton sector. In this paper, we will discuss LFV in processes such as $\mu \rightarrow e\gamma$, $\mu \rightarrow eee$, and $\mu \rightarrow e$ conversion in nuclei in two models, which are extensions of the minimal SUSY SM (MSSM): one is the MSSM with right-handed neutrinos (without R-parity violation),¹ the other is the MSSM with R-parity violation.² Especially the latter model can also be tested by the HERA experiments;³ however, the low energy muon-decay experiments provide very stringent bounds on some of the couplings, which are not constrained by HERA. We will also discuss the different features of LFV in these two models.

1 LFV in the MSSM with right-handed neutrinos

In this section, we consider the MSSM with right-handed neutrinos. The superpotential in the lepton sector is given by

$$W = f_e^i H_1 E_i^c L_i + f_\nu^i V_{ij} H_2 N_i^c L_j + \frac{M_R}{2} N_i^c N_i^c, \quad (1)$$

where for simplicity we assumed that the right-handed neutrino mass matrix is proportional to the unit matrix. In this framework, the small neutrino masses arise through the seesaw mechanism; if the right-handed neutrino mass scale M_R is much larger than the electroweak scale, we can obtain very tiny neutrino masses: $m_{\nu i} = (m_{\nu i}^D)^2 / M_R$, $\nu_{\text{mass}}^i = V_{ij} \nu_{\text{flavor}}^j$, where m_ν^D is a Dirac neutrino mass ($m_{\nu i}^D = \frac{f_\nu^i v \sin \beta}{2}$), and the mass eigenstates (ν_{mass}) of neutrinos are related to the flavour eigenstates (ν_{flavor}) via the mixing matrix V . An important

^aTalk given at the 8th International Workshop on Deep Inelastic Scattering and QCD (DIS 2000), Liverpool, England, 25-30 April 2000

point is that, because of the LFV mixing V , LFV is induced in the slepton mass terms even if we assume universal scalar mass (m_0) at the gravitational scale $M = 2 \times 10^{18} \text{GeV}$. We can calculate the LFV in slepton masses by solving the renormalization group equations numerically. The approximate solution is given by $(\Delta m_L^2)_{ij} \simeq -\frac{|f_\nu^k|^2 V_{ki} V_{kj}}{16\pi^2} (6 + 2a_0^2) m_0^2 \log(M/M_R)$. Note that large neutrino Yukawa couplings and large mixing V induce large flavour-violating masses. The LFV masses $(\Delta m_L^2)_{ij}$ generate LFV processes, e.g. $(\Delta m_L^2)_{23}$ induces $\tau \rightarrow \mu\gamma$, and $(\Delta m_L^2)_{12}$ generates $\mu \rightarrow e$ flavour violation. The atmospheric neutrino experiments indicate large mixing V_{32} . This large mixing can induce large branching ratio for $\tau \rightarrow \mu\gamma$ ^{1,4}. Moreover, some of the solar neutrino solutions suggest large V_{21} , which can induce large $\mu \rightarrow e$ flavour violation.⁴ In Fig. 1, we present the branching ratios of $\mu \rightarrow e\gamma$ and $\mu \rightarrow e$ conversion in Ti assuming that

$$V = \begin{pmatrix} \frac{1}{\sqrt{2}} & \frac{1}{2} & \frac{1}{2} \\ \frac{-1}{\sqrt{2}} & \frac{1}{2} & \frac{1}{2} \\ 0 & \frac{-1}{\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix}, \quad m_{\nu i} = (0, 0.006, 0.055)_i \text{ eV}, \quad (2)$$

which provides a solution to the atmospheric neutrino problem, as well as a large-angle MSW solution to the solar neutrino problem.^b We also checked that $\text{Br}(\mu \rightarrow eee)/\text{Br}(\mu \rightarrow e\gamma) \simeq 6 \times 10^{-3}$. Therefore the present and future experiments with a sensitivity of 10^{-14} for the $\mu \rightarrow e\gamma$ rate⁵ and 10^{-16} for the $\mu \rightarrow e$ conversion rate in Al^{6c} can probe the LFV in a large region of the parameter space of this model.

2 LFV in SUSY models with R-parity violation

Subsequently, we consider the MSSM with R-parity violation in which the small R-parity violation can potentially explain small neutrino masses. The superpotential is

$$W = f_e^i H_1 E_i^c L_i + \frac{f_{ijk}}{2} L_i L_j E_k^c + f'_{ijk} L_i Q_j D_k^c. \quad (3)$$

Here we assume that the baryon-number-violating terms $U^c D^c D^c$ are forbidden by baryon parity in order to avoid rapid proton decay.^d The Yukawa couplings f and f' violate lepton flavour number as well as lepton number. The LFV

^aIn the case with other solar neutrino solutions, see Ref.⁴

^bWe numerically checked that the $R(\mu \rightarrow e \text{ in Al}) \simeq 0.6 R(\mu \rightarrow e \text{ in Ti})$.

^cWe neglect R-parity violation in soft SUSY-breaking terms. This could be important for neutrino masses. Here we do not consider any particular models for neutrino masses.

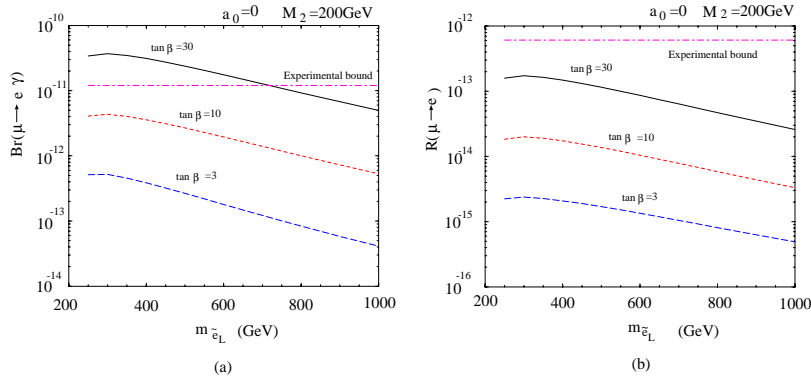


Figure 1: Event rates for (a) $\mu \rightarrow e\gamma$ (b) $\mu \rightarrow e$ conversion in Ti as a function of the left-handed selectron mass. We fixed the right-handed neutrino mass to $M_R = 10^{13}$ GeV.

experiments put the most stringent bounds on certain combinations of the Yukawa couplings f and f' .^{7,8} We list the bounds in Table 1. To understand an interesting feature of the LFV in this model, let us consider some simple examples. If only f_{131} and f_{231} are non-zero, the ratios of the event rates between the LFV processes are given by

$$\frac{\text{Br}(\mu \rightarrow e\gamma)}{\text{Br}(\mu \rightarrow eee)} = 1 \times 10^{-4}, \quad \frac{R(\mu \rightarrow e \text{ in Ti (Al)})}{\text{Br}(\mu \rightarrow eee)} = 2 \times 10^{-3} (1 \times 10^{-3}), \quad (4)$$

where we took all $m_{\tilde{l}}$ to be 100 GeV. The rate of $\mu \rightarrow eee$ is much larger than those of the other LFV processes. This is because $\mu \rightarrow eee$ is generated at tree level, while the rest of the LFV processes are induced at the one-loop level. If only f_{132} and f_{232} are non-zero, the event rates are given by

$$\frac{\text{Br}(\mu \rightarrow e\gamma)}{\text{Br}(\mu \rightarrow eee)} = 1.2, \quad \frac{R(\mu \rightarrow e \text{ in Ti (Al)})}{\text{Br}(\mu \rightarrow eee)} = 18 \quad (11), \quad (5)$$

where we took $m_{\tilde{l}} = 100$ GeV. Even though all processes are induced at the one-loop level in this case, the branching ratio of $\mu \rightarrow eee$ is still comparable to that of $\mu \rightarrow e\gamma$, and the rate of $\mu \rightarrow e$ conversion is even much larger because of a log-enhancement in the off-shell photon penguin contributions.⁸ Therefore $\mu \rightarrow eee$ and $\mu \rightarrow e$ conversion are important in this model. The interesting point is that the relations in Eqs.(4), (5) are quite different from those of the MSSM with right-handed neutrinos discussed in the previous section. Therefore, in order to distinguish between the different models, the study of all the LFV processes can be important. Note that at present the future proposal for

| | $\mu \rightarrow e\gamma$ | $\mu \rightarrow 3e$ | $\mu \rightarrow e$ in nuclei |
|----------------------|---------------------------------------|----------------------|---------------------------------------|
| $ f'_{131}f'_{231} $ | $2 \times 10^{-4} (7 \times 10^{-6})$ | 7×10^{-7} | $1 \times 10^{-5} (2 \times 10^{-7})$ |
| $ f'_{132}f'_{232} $ | $2 \times 10^{-4} (7 \times 10^{-6})$ | 7×10^{-5} | $1 \times 10^{-5} (2 \times 10^{-7})$ |
| $ f'_{133}f'_{233} $ | $2 \times 10^{-4} (7 \times 10^{-6})$ | 1×10^{-4} | $2 \times 10^{-5} (4 \times 10^{-7})$ |
| $ f'_{121}f'_{122} $ | $8 \times 10^{-5} (2 \times 10^{-6})$ | 7×10^{-7} | $6 \times 10^{-6} (1 \times 10^{-7})$ |
| $ f'_{131}f'_{132} $ | $8 \times 10^{-5} (2 \times 10^{-6})$ | 7×10^{-7} | $7 \times 10^{-6} (1 \times 10^{-7})$ |
| $ f'_{231}f'_{232} $ | $8 \times 10^{-4} (2 \times 10^{-6})$ | 4×10^{-5} | $8 \times 10^{-6} (1 \times 10^{-7})$ |
| $ f'_{111}f'_{211} $ | $7 \times 10^{-4} (2 \times 10^{-5})$ | 1×10^{-4} | $5 \times 10^{-6} (2 \times 10^{-7})$ |
| $ f'_{112}f'_{212} $ | $7 \times 10^{-4} (2 \times 10^{-5})$ | 1×10^{-4} | $4 \times 10^{-7} (7 \times 10^{-9})$ |
| $ f'_{113}f'_{213} $ | $7 \times 10^{-4} (2 \times 10^{-5})$ | 2×10^{-4} | $4 \times 10^{-7} (7 \times 10^{-9})$ |
| $ f'_{121}f'_{221} $ | $7 \times 10^{-4} (2 \times 10^{-5})$ | 2×10^{-4} | $4 \times 10^{-7} (6 \times 10^{-9})$ |
| $ f'_{122}f'_{222} $ | $7 \times 10^{-4} (2 \times 10^{-5})$ | 2×10^{-4} | $4 \times 10^{-5} (7 \times 10^{-7})$ |
| $ f'_{123}f'_{223} $ | $7 \times 10^{-4} (2 \times 10^{-5})$ | 3×10^{-4} | $5 \times 10^{-5} (9 \times 10^{-7})$ |
| $ f'_{131}f'_{231} $ | $2 \times 10^{-3} (6 \times 10^{-5})$ | 4×10^{-4} | $4 \times 10^{-7} (6 \times 10^{-9})$ |
| $ f'_{132}f'_{232} $ | $2 \times 10^{-3} (6 \times 10^{-5})$ | 5×10^{-4} | $9 \times 10^{-5} (2 \times 10^{-6})$ |
| $ f'_{133}f'_{233} $ | $2 \times 10^{-3} (6 \times 10^{-5})$ | 9×10^{-4} | $2 \times 10^{-4} (3 \times 10^{-6})$ |

Table 1: Present (Future) constraints on R-parity-violating couplings from LFV processes. The present limits (future expectations) on the event rates are given by $\text{Br}(\mu \rightarrow e\gamma) < 1.2 \times 10^{-11}$ (10^{-14}), $\text{Br}(\mu \rightarrow 3e) < 1.0 \times 10^{-12}$, and $\text{R}(\mu \rightarrow e \text{ in Ti (Al)}) < 6.1 \times 10^{-13}$ (10^{-16}). We took $m_{\tilde{\nu}} = m_{\tilde{t}_R} = 100$ GeV and $m_{\tilde{q}} = 300$ GeV.

$\mu \rightarrow eee$ is missing. Furthermore, we should stress here that not only low-energy muon probes but also other experiments³ should be encouraged in order to investigate the R-parity-violating SUSY since this theory is well-motivated from the recent neutrino data.

Acknowledgements

The author would like to thank A. de Gouvêa, G.F. Giudice, and S. Lola for useful discussions.

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