

*Large Hadron Collider Project*

**LHC Project Report 256**

## **VUV Photoemission Studies of Candidate LHC Vacuum Chamber Materials**

R. Cimino<sup>1</sup>, I.R. Collins<sup>2</sup>, V. Baglin<sup>2</sup>

### Abstract

In the context of future accelerators and, in particular, the beam vacuum of the Large Hadron Collider (LHC), a 27 km circumference proton collider to be built at CERN, VUV synchrotron radiation (SR) has been used to study both qualitatively and quantitatively candidate vacuum chamber materials. Emphasis is given to show that angle and energy resolved photoemission is an extremely powerful tool to address important issues relevant to the LHC, such as the emission of electrons that contribute to the creation of an electron cloud which may cause serious beam instabilities. Here we present not only the measured photoelectron yields (PY) from the proposed materials, prepared on an industrial scale, but also the energy and, in some cases, the angular dependence of the emitted electrons when excited with either a white light (WL) spectrum, simulating that in the arcs of the LHC or monochromatic light in the photon energy range of interest. The effects on the materials examined of WL irradiation and/or ion sputtering, simulating the SR and ion bombardment expected in the LHC, were investigated. The studied samples exhibited significant modifications, in terms of electron emission, when exposed to the WL spectrum from the BESSY TGM7 beamline. Moreover, annealing and ion bombardment also induce substantial changes to the surface thereby indicating that such surfaces would not have a constant electron emission during machine operation. Such characteristics may be an important issue to define the surface properties of the LHC vacuum chamber material and are presented in detail for the various samples analysed.

1 INFN sezione di Trieste

2 LHC Vacuum Group, CERN

Administrative Secretariat  
LHC Division  
CERN  
CH - 1211 Geneva 23  
Switzerland

Geneva, 8 December 1998

## Table of contents

<b>1. INTRODUCTION</b>	<b>3</b>
<b>2. EXPERIMENTAL</b>	<b>4</b>
2.1 The photon source	4
2.2 The data collection	6
2.3 The samples	7
<b>3. RESULTS AND DISCUSSION</b>	<b>9</b>
3.1 Photon energy dependence and angular distributions of the EDC of Au	9
3.2 Understanding the effects of WL irradiation and surface conditioning	11
3.2.1 Photoemission with monochromatic light: Cu	12
3.2.2 Photoemission with monochromatic light: Cu-el.	13
3.2.3 Photoemission with monochromatic light: TiZr.	14
3.2.4 Photoemission with monochromatic light: St 707.	17
3.2.5 Photoemission with monochromatic light: Pd.	18
3.2.6 Photoemission with monochromatic light: TiN.	19
3.3 The photon energy dependence of the PY (Constant Final State spectroscopy).	20
3.4 White light photoelectron yields of the as-received samples	22
3.5 White light photoelectron yields of the surface conditioned samples	23
3.6 Energy Distributions of the WL EDC	24
3.7 Generation of ‘low dose’ WL spectra.	26
3.8 The energy distribution of the secondary electrons in the LHC.	29
<b>4. CONCLUSIONS</b>	<b>31</b>

## 1. Introduction

The LHC will provide two counter circulating proton beams with colliding energies of nominally 14 TeV in the center of mass, requiring superconducting bending magnets operating in superfluid helium at 1.9 K. In order to reduce the cryogenic power consumption at 1.9 K in the arcs, the heat load induced by the beam will be intercepted on a beam screen, which operates at 5–20K. In the arcs of the machine molecules will be pumped through the pumping slots, distributed along the length of the beam screen, and onto the surrounding cold bore held at 1.9 K.

The emitted synchrotron radiation (SR) from the circulating protons, with a critical energy of 44.1 eV, is a major consideration for the design of the vacuum system. Its radiated power induces a heat load of 0.2 W/m per beam and may: i) stimulate gas desorption of weakly and tightly bound gases from the walls of the vacuum system either directly by photons or mediated by electrons [1,2], ii) create photoelectrons which can be accelerated, to an average energy of 380 eV [3], towards the opposite wall by the positive space charge of the bunched beam leading to additional gas desorption and heat loads on the cryogenic system, iii) create secondary electrons which may contribute to electron multipacting [2]. This latter phenomenon is a resonant effect where a cloud of secondary electrons oscillate, in phase with the bunched beam, between opposite walls of the vacuum chamber. Simulations performed by considering the maximum of the secondary electron yield (SY),  $\delta$ , as the representative parameter of the SY suggest that if  $\delta$  exceeds a critical value [3], the cloud may increase exponentially, leading to an unmanageable heat load and ultimately to beam loss.

In addition to the SR, ion induced desorption is another major consideration in the design of the LHC. Positive ions, created by ionisation of the residual gas and repelled by the positive space charge of the beam, bombard the vacuum chamber wall and desorb tightly and weakly bound molecules which would lead, to a pressure instability if careful design is not considered [2]. Consequently, significant modifications of the industrially ex-situ prepared beam screen can be expected during machine operation. Such surface cleaning induced by SR ion and electron bombardment, may therefore result in changes in the electron yields as a function of machine operation.

Calculations based on simulations are in progress to estimate the cryogenic heat load due to photoelectrons and secondary electrons and, in doing so, provide information on the sensitivity of the heat loads and electron cloud to the input parameters [3,4]. It has been predicted that only the secondary electrons emitted with a kinetic energy in the range between 1 to 6 eV play a role in multipacting. Realistic input values for the photoelectron yield (PY), the energy and angular distribution of the emitted electrons, the reflectivity, *etc.* from technical surfaces are required for design optimisation.

SR photoemission is an ideal tool to provide this information. Preliminary experiments have been performed giving qualitative information [5]; here quantitative information on the PY are presented. In addition the detailed energy and angular distribution of the photoemitted electrons are presented. Data were acquired using both ‘white light’ (WL) SR and monochromatic light as the excitation probe. Studies were performed to investigate the important issues of the cleaning effects of SR illumination and/or surface conditioning, such as ion bombardment or annealing.

It should be pointed out that the data presented here were performed with samples held at room temperature whilst the LHC beam screen will be held at between 5 and 20 K. In order to simulate closely the LHC conditions, it will be necessary to measure samples at cryogenic temperatures in order to study issues such as the effect of physisorbed gases on the PY.

Experimental details are given in section 2 including information of each sample preparation, followed in section 3 by the experimental results and discussion. A summary of the results and an outlook on future work will be presented in sect. 4.

## 2. Experimental

The experimental section will be divided into three parts describing the photon source used, the samples studied with their detailed sample preparation and the details of the data acquisition.

### 2.1 The photon source

In the LHC, given the mass of the circulating particles and their energy, the SR critical energy is 44.1 eV, so that the maximum flux emitted by the machine is in the ultraviolet up to the soft x-ray region. To simulate as accurately as possible such a spectrum one has either to run an electron accelerator so that SR has the same critical energy [6] or to use existing SR sources, which have generally higher critical energy, *i.e.* higher photon flux in the hard X ray region, and filter their emission through a beamline. The latter can be achieved by attenuating the high energy part of the spectrum by multiple reflections from the different optical elements or by using monochromatic light in the range of interest for LHC.

There are no existing beamlines designed to deliver either monochromatic light and a WL spectrum identical to the one of the LHC. These studies employed a Toroidal Grating Monochromator beamline (TGM7) [7] installed on the BESSY synchrotron light source, Berlin, a 805 MeV electron storage ring. The optical layout of such monochromator is shown in Figure 1. The grating can either be used in diffraction mode or in reflection mode in order to produce monochromatic light or WL respectively. It is important to estimate the photon flux to quantitatively analyse the observed WL irradiation effects and total PY measurements obtained from the different samples. The WL spectrum was calculated and measured experimentally for the optical layout of Figure 1, when using the grating as a mirror (in its ‘zero order’ position). The BESSY bending magnet photon flux and the transmission of each of the optical elements as a function of its material and angle of incidence, has been calculated using the LBL codes available on the Web [8]. The total irradiated area was estimated to be of the order of  $1 \text{ mm}^2$ .

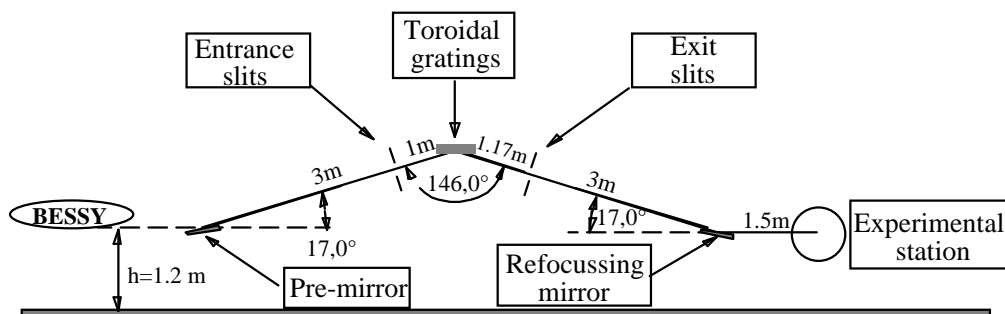


Figure 1: TGM 7 optical layout

The so-calculated flux has been corrected by the angular acceptance of the beamline (6 mrad vertical x 15 mrad horizontal) and the used slits sizes to give the spectrum shown in Figure 2. The approximations made for this calculation are such that the spectrum in Figure 2 gives an over estimated photon flux, *i.e.* losses due to imperfect alignment of the optical elements, *etc.* are not considered. However, the integrated flux of  $8.9 \cdot 10^{14}$  ph/s/mm<sup>2</sup> obtained from the calculated spectrum provides a consistent value with the actual photon flux measured experimentally from a clean gold reference sample (as shown in section 3.4).

Figure 2 shows that the spectrum so-obtained (having a cut-off mainly determined by the poor reflectivity of the optical elements at energies higher than 150 eV) is, in shape, a distorted replica of the LHC bending magnet spectrum. However, the number of photons/s/mm<sup>2</sup> impinging on the sample is roughly 1000 times larger, per unit area, than that in the LHC. This considerably more intense photon flux was exploited to study the effects of high dose irradiation on the samples.

On the other hand, as shown in Table 1, when using monochromatic light in the energy range between 18 and 110 eV, photon fluxes were comparable to that expected in the LHC for a mean value of BESSY beam current which decayed from 700 mA at injection to 200 mA at the end of a run.

The flux reduction between WL and monochromatic light is not surprising since, when using the grating in its diffraction mode an efficiency loss of around a factor 1000 is expected. It should be pointed out that the photon dose has to be considered carefully when discussing WL electron emission spectrum (and measured total currents) since significant surface cleaning may occur and therefore may only be representative of strongly ‘irradiated samples’.

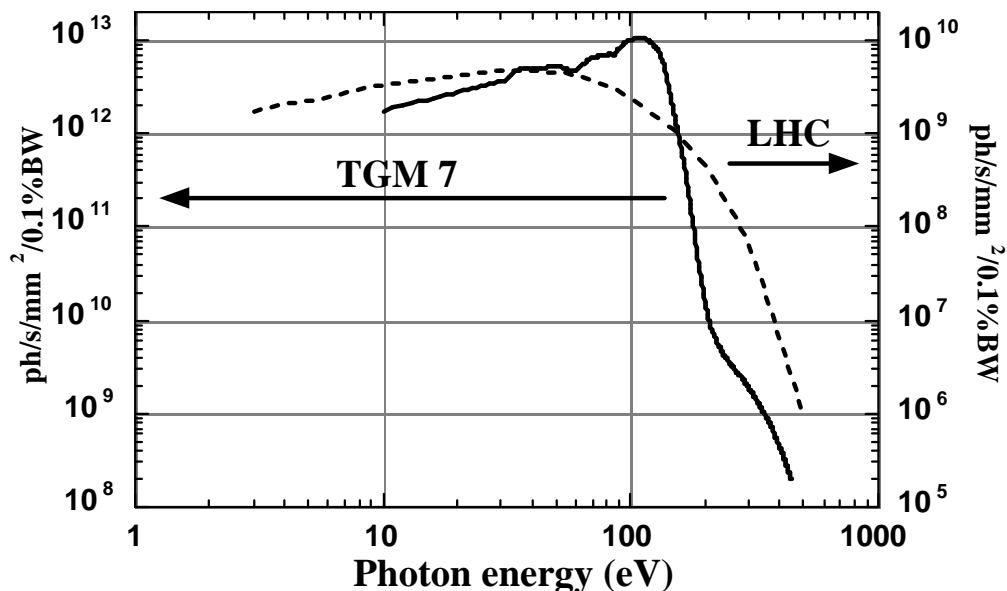


Figure 2: Comparison between the LHC bending magnet photon flux with that obtained with the WL spectrum of TGM7 beamline in BESSY. The integrated photon flux is  $5.2 \cdot 10^{11}$  and  $8.4 \cdot 10^{14}$  ph/s/mm<sup>2</sup> respectively.

## 2.2 The data collection

The samples (described in section 2.3) were installed *via* a fast load–lock into a UHV ( $10^{-10}$  Torr range)  $\mu$ -metal experimental chamber equipped with a VG ARUPS 500 angle resolving electron spectrometer. The surfaces were characterized *via* Constant Final State (CFS) spectroscopy and angle resolving photoemission using photon energies between 18 and 110 eV spanning the 44.1 eV critical energy spectrum of the LHC. In addition, the TGM 7 WL spectrum was used to perform angle resolved photoemission measurements. The beamline slits define a photon resolution from 200 meV to 400 meV for 18 eV to 110 eV photon energy respectively.

All the measurements were performed with the sample biased negatively with respect to ground to allow electrons with a very low kinetic energy to be measured accurately by the analyser. The chosen bias of -9.5 V (obtained from a battery) was sufficient to eliminate sample space charge problems and to permit one to measure the low energy part of the photoemission spectrum with the analyser working in optimum conditions. Indeed, more than -5 V bias produced spurious effects on the shape of the low kinetic energy secondaries due to the simultaneous collection of secondaries from the sample and from those created in the analyser itself.

Slow electrons (with kinetic energies close to 0 eV) could be measured confidently in our set up due to the use of a  $\mu$ -metal UHV chamber that shielded efficiently stray magnetic fields acting near the sample. It is noted that even the Earth's magnetic field or spurious fields from ion pumps *etc.* may affect strongly the detection of electrons with very low kinetic energies.

The samples were mounted on a partially electrically isolated sample holder (resistance to earth was around 1 G $\Omega$ ). Such non-perfect isolation and the use of a negative 9.5 V bias, allows us to measure accurately a sample drain current only with WL illumination (measured current was in  $\mu$ A range). It was impossible to measure with the required accuracy a sample drain current in the range of nA typical range when exciting with monochromatic light.

The photon flux, proportional to the drain current on the last gold-plated refocusing mirror after the exit slit of the monochromator (Figure 1), was measured during data acquisition. The measured drain currents were converted into photon flux using the quantum efficiency of clean gold, as cited in the literature [9].

Table 1 shows such data in the following way: in the first column is reported some selected photon energies for which the flux will be given; in the second column the mirror current reading is reported per 100 mA ring current; in the third column the Au efficiency values are shown [9]; in the fourth column the so-obtained flux values are given in units of ph/s/mm<sup>2</sup>/0.1%bw; in the last column an estimated error of the quoted flux is given. Such errors are quite large and have been extracted taking into account the variation of the ring current in time, the difficulties to estimate correctly the spot size on the sample, *etc.* As stated in section 2.1, the monochromatic photon fluxes depicted in Table 1 agree, within its error, with the required flux to be obtained to simulate the LHC spectrum shown in Figure 2.

Table 1: Measured monochromatic flux on the last refocussing gold mirror of the TGM7 beamline.

Photon Energy (eV)	$I_{\text{mirror}}$ (nA)	Au eff. [9]	Photon Flux ph/s/mm <sup>2</sup> /0.1%bw	Error ph/s/mm <sup>2</sup> /0.1%bw
20	0.08	0.091	$2.5 \cdot 10^9$	$1.3 \cdot 10^9$
30	0.20	0.073	$7.7 \cdot 10^9$	$3.8 \cdot 10^9$
50	0.50	0.057	$2.5 \cdot 10^9$	$1.3 \cdot 10^9$
70	0.50	0.053	$2.6 \cdot 10^9$	$1.3 \cdot 10^9$
100	0.40	0.034	$3.3 \cdot 10^9$	$1.7 \cdot 10^9$

The WL photon flux estimation has been obtained by correcting the measured drain current from the last refocussing mirror by a weighted estimate of the energy dependent gold efficiency curve [9]. Such measured values are reported in Table 3 for different ring currents and agree well with the calculated WL integral flux (of  $8.9 \cdot 10^{14}$  ph/s/mm<sup>2</sup>) shown in Figure 2.

All measurements were performed with photons incident at 45° and electrons emitted perpendicular to the surface, unless otherwise stated.

### 2.3 The samples

The studied samples are listed in

Table 2 with a detailed description of their preparation to produce the analysed surface. Sample Cu has been already measured in a previous experiment on a different beam line [5].

Table 2: Sample and sample preparation details. The naming key will be used throughout the rest of this paper to refer to each sample studied.

Naming key	Sample	Sample preparation
Au	Evaporated gold Reference sample	A clean Au thin (>5 μm) film was prepared by evaporation of pure Au, by resistive heating of W crucible, onto a gold foil in UHV ( $p < 3 \cdot 10^{-9}$ mbar)

Cu	OFE copper co-laminated on high-Mn content stainless steel	OFE copper co-laminated onto high-Mn content (9% wt.) stainless steel (UNS21904) and annealed at 900°C for 15 minutes (from UGINE, France). It was chemically cleaned by degreasing in perchloroethylene vapor at 121°C, then deoxidizing by immersion in an alkaline detergent, pickling in HCl, passivation in a chromic acid/H <sub>2</sub> SO <sub>4</sub> bath, rinsing in demineralized water and finally drying with N <sub>2</sub> gas. The sample was held one year at atmospheric pressure.
Cu-sp.	Cu sputtered	A 'clean' copper sample was obtained by <i>in-situ</i> 2 keV Ar <sup>+</sup> sputtering (P = 1 10 <sup>-5</sup> mbar) for 30 minutes of the Cu sample.
Cu-ab.	Cu air baked	OFE copper co-laminated onto high-Mn content (9% wt.) stainless steel (UNS21904) and annealed at 920°C for 7.5 mins under H <sub>2</sub> (from UGINE, France). It was chemically cleaned by degreasing in perchloroethylene vapor at 121°C, then deoxidizing by immersion in an alkaline detergent, pickling in HCl, passivation in a chromic acid/H <sub>2</sub> SO <sub>4</sub> bath, rinsing in demineralized water and finally drying with N <sub>2</sub> gas. Baked at 350°C for 24hrs, vented to air whilst at 350°C for 15 mins., evacuated at 350°C and cooled to RT. The sample was held three months at atmospheric pressure.
TiN	TiN on extruded Al alloy	6063 Al Alloy (0.5% Mg, 0.5% Si) and 187 nm TiN (DC reactive sputtered, Ti cathode, 20%N/Ar, 230mTorr total pressure).
TiN-sp.	TiN sputtered	<i>In-situ</i> 2 keV Ar <sup>+</sup> sputtering (P=1.7 10 <sup>-5</sup> mbar) for 5 minutes of the TiN sample .
TiZr	~1µm TiZr onto OFE copper	~1 µm TiZr film deposited onto OFE copper after having first removed 120 µm of Cu by electropolishing then another 10 µm as a result of a chemical treatment in: Sulfamic acid (5g/l), Ammonium Citrate (1g/l), n-butanol (50ml/l) and 30% vol. H <sub>2</sub> O <sub>2</sub> (50ml/l).
TiZr-ac.	Activated TiZr with Procedure A	Annealing of TiZr in UHV (pressure during annealing was less than 10 <sup>-8</sup> mbar) at 330±30°C for 4 hours (Procedure A).
TiZr II	~1µm TiZr on OFE copper	A second identical TiZr sample.
TiZrII-ac.	Activated TiZr with Procedure B	Annealing of TiZr II in UHV (pressure during annealing was less than 5x10 <sup>-9</sup> mbar) at 300±30°C for 2.25 hours (procedure B).
Cu-el.	Cu electropolished	Cu electropolished (150 µm removed).



Cu-el.-an.	Annealed electropolished Cu with procedure B	Annealing of Cu-el. sample in UHV (pressure during annealing was less than $5 \cdot 10^{-9}$ mbar) at $330 \pm 30^\circ\text{C}$ for 2.25 hours (Procedure B).
Pd	0.3 $\mu\text{m}$ Pd on OFE Cu	0.3 $\mu\text{m}$ Pd film deposited onto OFE copper after having first removed 120 $\mu\text{m}$ of Cu by electropolishing then another 10 $\mu\text{m}$ as a result of a chemical treatment in: Sulfamic acid (5g/l), Ammonium Citrate (1g/l), n-butanol (50ml/l) and 30% vol $\text{H}_2\text{O}_2$ (50ml/l).
Pd-an.	Annealed Pd with Procedure C	Annealing of Pd sample in UHV (pressure during annealing was less than $5 \cdot 10^{-9}$ mbar) at $300 \pm 30^\circ\text{C}$ for 8 hours (Procedure C)
St 707	St 707 <sup>TM</sup> from SAES Getters	St 707 <sup>TM</sup> (Zr 70%, V 24.6%, Fe 5.4%) as described by the SAES Getters data-sheet.
St 707-ac.	Activated St 707 <sup>TM</sup> with Procedure C	Annealing of St 707 <sup>TM</sup> sample in UHV (pressure during annealing was less than $5 \cdot 10^{-9}$ mbar) at $300 \pm 30^\circ\text{C}$ for 8 hours (Procedure C)
Al	Al alloy	Aluminium alloy AA6060, degreased with perchloroethylene vapour at $121^\circ\text{C}$ , etched with alkaline detergent and rinsed.
Al-sp.	Al sputtered	<i>In-situ</i> 2 keV $\text{Ar}^+$ sputtering ( $P=1.7 \cdot 10^{-5}$ mbar) for 5 minutes of Al sample.

### 3. Results and Discussion

The results obtained are divided into seven sections, each one addressing a particular key point, namely:

- 3.1) The photon energy dependence and angular distributions of the Energy Distribution Curves (EDC) Au.
- 3.2) The use of monochromatic SR to understand the effect of WL irradiations and surface conditioning effects.
- 3.3) The photon energy dependence of the PY (CFS).
- 3.4) The measured WL exposed PY of the as-received samples obtained from the sample drain current.
- 3.5) The measured WL exposed PY of the conditioned samples obtained from the sample drain current.
- 3.6) The EDC of such WL exposed samples.
- 3.7) The generation of 'low dose' WL EDC.
- 3.8) The energy distribution of the secondary electrons in the LHC.

#### 3.1 Photon energy dependence and angular distributions of the EDC of Au

The clean Au sample was studied in detail, not only that it may be considered as a possible solution to be used for the LHC beam screen surface, but because it was necessary to study a well known and characterized sample in order to calibrate our experimental set-up. All the electron

energy analyser calibrations, and most of the angle resolved studies, have been performed from this surface. It provided a reference sample from which estimates of confidence levels with which the data from the other studied samples could be treated. In Figure 3 are shown a full series of EDC's from the clean Au as a function of photon energy. Valence band structures, Fermi edges, core levels and secondary electron tails are clearly identified. The energy calibrations and resolutions can be deduced from the high energy cut-off (Fermi level). All EDCs were taken with a 20 V pass energy on the analyser, which, together with the photon energy resolution, correspond to a total electron energy resolution (FWHM) ranging from 0.2 eV to 0.4 eV for 20 eV and 110 eV photon energies, respectively. It should be noted that the onset of the secondary tail is directly related with the material work function (WF). The known clean polycrystalline Au WF of 5.1 eV [14] can be then used to extract WFs for the different samples studied. The WF is an important parameter since it can be simply related to PY. A high(low) WF material is expected to have a low(high) PY, giving a simple rule-of-thumb to predict the relevant behaviour of a surface in terms of PY. We will discuss this aspect in more detail in a future publication.

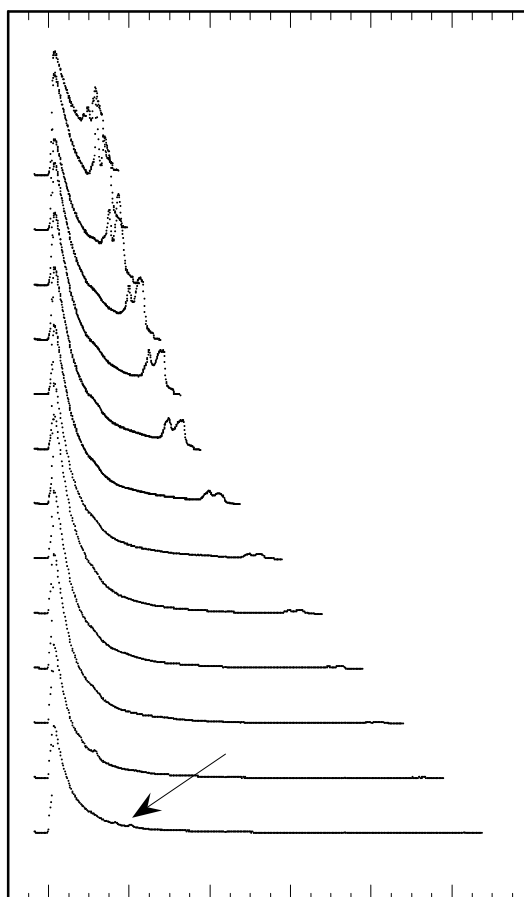


Figure 3: A complete series of EDC's taken from the Au sample as a function of photon energy.

Angle resolved EDC of the valence band electrons from clean Au are shown in Figure 4. Similar data were recorded from some of the other samples but are not shown here. In the inset of this figure the energy integrated intensity is plotted as a function of angle, exhibiting a rather flat maximum around  $20^\circ$  off normal emission and, then, decreases with angle. Very probably there exist some systematic errors such that the maximum of the distribution is not at normal emission, as one might expect. At this point these data serve to demonstrate that photoemission can provide both information on the intensity of the photoemitted electrons versus energy and emission angle. This information may be useful for the modeling of photoemitted electrons in the LHC and may be implemented in the simulations. In the LHC the electrons will be photoemitted mostly in the presence of a strong dipole field (8.4 T), so that their energy and angular distributions may play a role in the generated heat loads and beam perturbations due to the electron cloud.

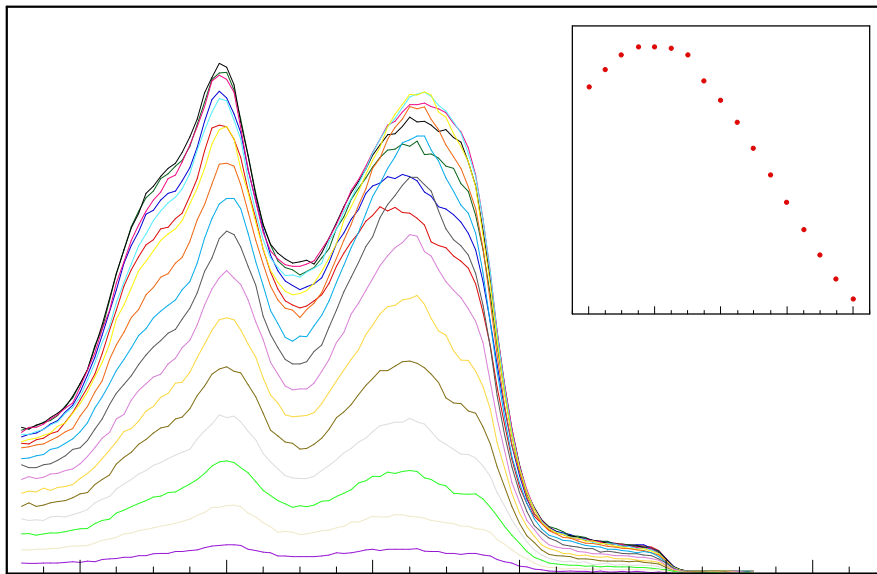


Figure 4: Angle resolved valence band photoemission spectra from clean Au taken with 50 eV photons, referenced to the Fermi level. The energy integrated intensity as a function of emission angle is shown in the inset.

### 3.2 Understanding the effects of WL irradiation and surface conditioning

As discussed in section 2.1 the number of photons/s/mm<sup>2</sup>/0.1% bandwidth impinging on the sample with monochromatic light is comparable to that on the beam screen from the SR of the LHC bending magnet, whereas the TGM 7 WL spectrum is some 1000 times more intense. This high photon flux thus corresponds to 5 to 50 hours of LHC irradiation ( $10^{16}$  -  $10^{17}$  photons/s/mm<sup>2</sup>).

One major advantage when using monochromatic light is therefore the substantially reduced photon flux with respect to WL flux thereby simulating the initial conditions in the LHC. The monochromated light can be used efficiently as a tool to monitor and follow changes due to cleaning effects, induced by intense WL irradiation, ion bombardment or heating, without significantly modifying the surface by the measurement itself.

The following is divided into sections for a selection of the samples studied.

### 3.2.1 Photoemission with monochromatic light: Cu

Cu is currently the base line design material for the surface of the LHC beam screen. The resistivity of the vacuum chamber wall material is an important design criteria for the LHC. The image currents of the beam flowing in the surface of the material (over a given skin depth) leads to a beam impedance and a heat load on the cryogenic system. In fact, it is proposed to coat the beam screen, fabricated from a high Mn content stainless steel, with 50  $\mu\text{m}$  of Cu.

Photoemission spectra, taken with a photon energy of 30 eV, from the Cu sample as a function of conditioning with WL and Ar ion sputtering (sample Cu-sp.), are shown in Figure 5.

It is important to notice that the shape and the electron energy distribution of the as-received spectrum is structureless apart for the existence of a very high emission of secondary electrons at low kinetic energies. The as-received EDC is similar with previous measurements performed on the same sample using a different experimental configuration (beam line and analyser) [5].

Dramatic changes of the emitted electrons, both in the low energy part (mainly secondary electrons) and in the region closer to the Fermi energy (which is a distorted replica of the valence band density of states), are evident as a function of conditioning. The WL irradiation ( $10^{16}$ - $10^{17}$  ph/mm<sup>2</sup>) partially cleans the dirty surface of the as-received sample, but does not clean it completely since the resulting spectrum is somewhat different to that exhibited by the essentially clean Cu surface, obtained by argon sputtering of the Cu sample. Ar ion sputtering (estimated dose  $8 \cdot 10^6$  Ar<sup>+</sup>.cm<sup>-2</sup>) of Cu produces a nearly atomically clean copper surface, as confirmed by the clear appearance, in the valence band part of the spectrum, of a Fermi edge and the Cu 3d bands at about 3 eV below it. The spectrum is similar to the one published in the literature for clean polycrystalline copper [10]. The WL exposure of this Ar sputtered surface does not change the measured spectrum (data not show for clarity) confirming that the WL has an effect only in the presence of a contaminant covered overlayer.

It is important to note, however, that the so-called WL cleaning effect is a very complex process which strongly depends on the detailed chemical structure of the contaminant layer and on the way it is bound to the underlying surface. In this, an important parameter can indeed be the time of irradiation *i.e.* the dose given. Such an aspect has been not studied in detail since beamtime was limited although initial investigations on the exposure dependence of the WL exposed sample indicated that irradiation for few minutes and more than an hour produced similar spectra. This is consistent with the heavy dose given even after few seconds of WL irradiation. Such doses have been shown to be sufficient to affect the desorption yield of the surface (hence its composition) [11, 12]. It can be speculated that the WL exposure is, rather than cleaning the surface, changing it by cracking chemisorbed molecules to smaller units which can then no longer be photodesorbed. This is clearly supported by the striking difference seen in Figure 5 between the spectrum for the WL irradiated sample and the one obtained for clean copper: after WL illumination no Fermi level is observed, and the Cu 3d band have more the shape of the one observed in oxidized copper than of clean copper [10].

The shifts in the cutoff region observed confirms that WF changes will result from surface conditioning. In the case of sputtered Cu it exhibits a higher WF than the as-received sample and, therefore, is expected to have a lower PY. This result is consistent with the PY data which will be discussed in sec. 3.4, 3.5.

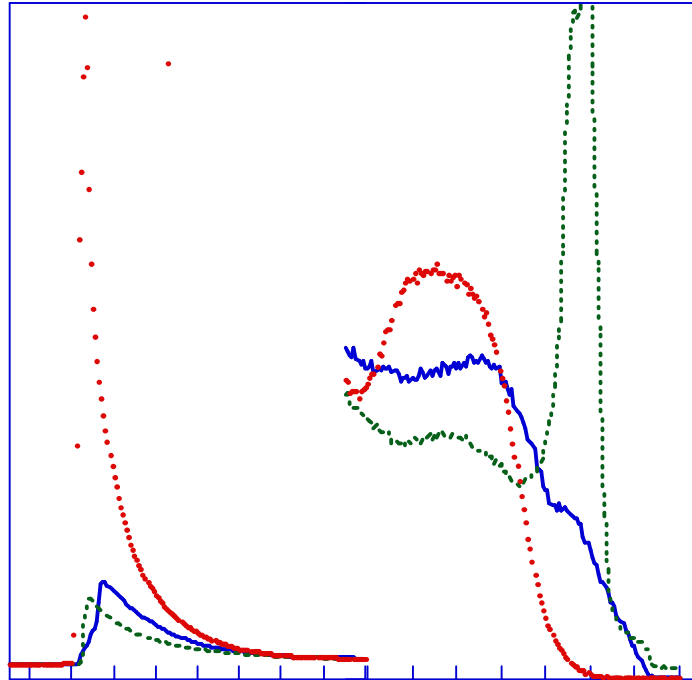


Figure 5: Photoemission spectra of Cu taken at 30 eV photon energy of: a) as-received; b) after WL exposure; c) after Ar<sup>+</sup> Sputtering (Cu-sp.).

### 3.2.2 Photoemission with monochromatic light: Cu-el.

It is interesting to understand if SR irradiation effects are comparable to the one obtained with annealing. To address this question an electropolished copper sample, Cu-el., was studied after different annealing treatments.

Figure 6 shows photoemission spectra taken at 50 eV photon energy of such a Cu sample in three conditions: a) as received, b) after SR illumination ( $10^{16}$ - $10^{17}$  ph/mm<sup>2</sup>); c) after UHV annealing (pressure during annealing was less than  $5 \cdot 10^{-9}$  mbar) at  $300 \pm 30$  °C for 2.25 hours (sample Cu-el-an.).

The as-received material gives a very high emission of very low energy electrons (as in the case of Cu), which is reduced in a similar way both after SR irradiation and after UHV annealing. However, the two surface treatments, affect differently the surface electronic properties, as demonstrated by the strong differences observed in the Cu 3d valence band structures. The annealing procedure seems to produce a cleaner surface than WL exposure, since the Cu 3d valence band is more pronounced. This is in line with the observed higher activation energy of the

molecules at a surface after exposure to SR than after annealing [12]. However, the annealing is less efficient than Ar sputtering in terms of cleaning the surface.

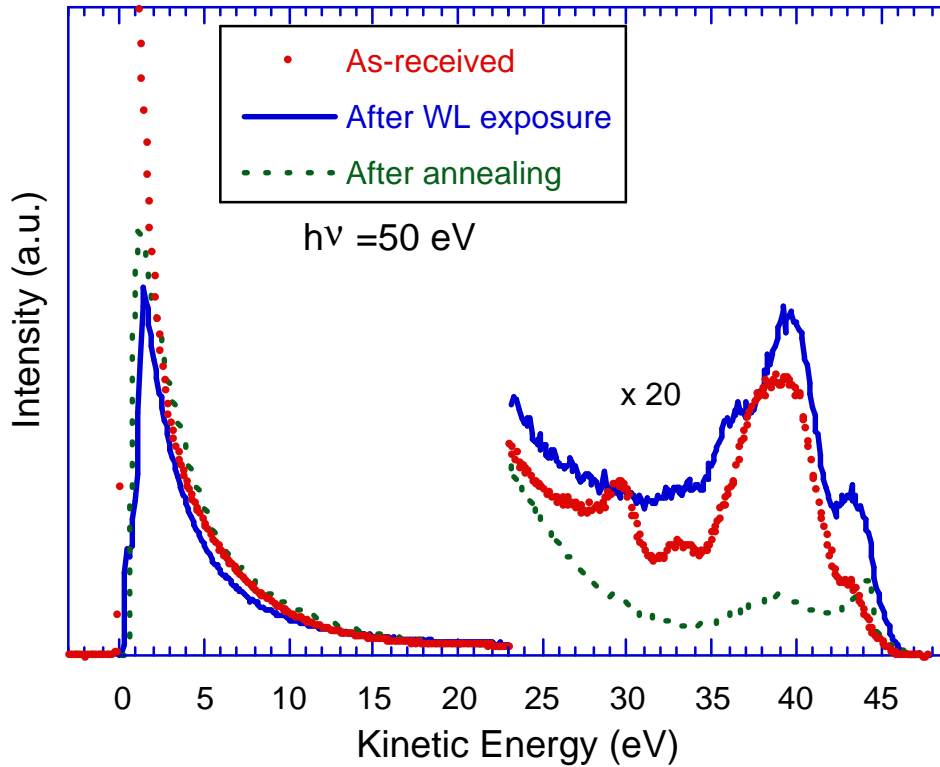


Figure 6: Photoemission spectra of Cu-el. taken at 50 eV photon energy of: a) as-received; b) after WL exposure; c) after annealing in UHV (pressure during annealing was less than  $5 \cdot 10^9$  mbar) at  $300 \pm 50$  °C for 2.25 hours (Cu-el.-an.).

### 3.2.3 Photoemission with monochromatic light: TiZr.

A family of candidate materials for the LHC vacuum chambers are Non-Evaporable Getters (NEG) since they provide an elegant solution for distributed pumping in conductance limited vacuum systems [13]. In particular these materials are being proposed in the warm sections of the machine and in the experimental beampipes. In order to pump, NEG's must be activated by heating. It is therefore of interest to study their behavior under SR illumination after activation.

A similar experiment, to that described previously in section 3.2.2, was performed on a  $\sim 1\mu\text{m}$  TiZr film deposited onto OFE Cu. The spectra are shown in Figure 7. The as-received surface shows a very high emission of low energy electrons.

Again illumination with WL SR ( $10^{16}$ - $10^{17}$  ph/mm<sup>2</sup>), induces strong spectral changes: the secondary emission is reduced and structure closer to the Fermi edge in the top of the valence band appears. This is again a signature of the SR assisted cleaning.

More interestingly, the annealing of such a surface causes dramatic changes in its nature: the intensity of secondary electrons reduces and a clear Fermi level is observed as a signature of a clean metallic surface. The effect of annealing causes the removal of the oxide passivation layer present on the as-received surface. This sample was annealed in a UHV pressure of less than

$1 \cdot 10^{-8}$  mbar and at  $330 \pm 30$  °C for 4 hours. From these data the atomic cleanliness of the surface cannot be evaluated since valence band photoemission, for instance, is not sensitive to hydrogen, but the metallicity of the surface after activation, as shown from the appearance of a sharp Fermi edge at 46.8 eV kinetic energy is clearly demonstrated.

As in the case of Cu-sp., the valence band structures are also unaffected by further WL illumination, perhaps suggesting that the surface is atomically clean. A closer examination of the secondary electron distribution, however, indicates that the WL modifies the distribution, both by affecting its intensity, and by shifting the zero kinetic energy cutoff. This shift is a signature of a related change in the work function of the material, possibly due to desorption of atomic species (like H) which might be present on the surface even after activation. This behavior is not surprising given the fact that it is known that a variation of a very small quantity, in the submonolayers coverage range, on a surface [14] can induce dramatic changes in the work function (up to some eV in some cases). This aspect is an essential issue for LHC since the PY, one of the most important parameters in the simulations of the heat loads and electron cloud issues, is obviously strongly dependent on the work function of the chosen material.

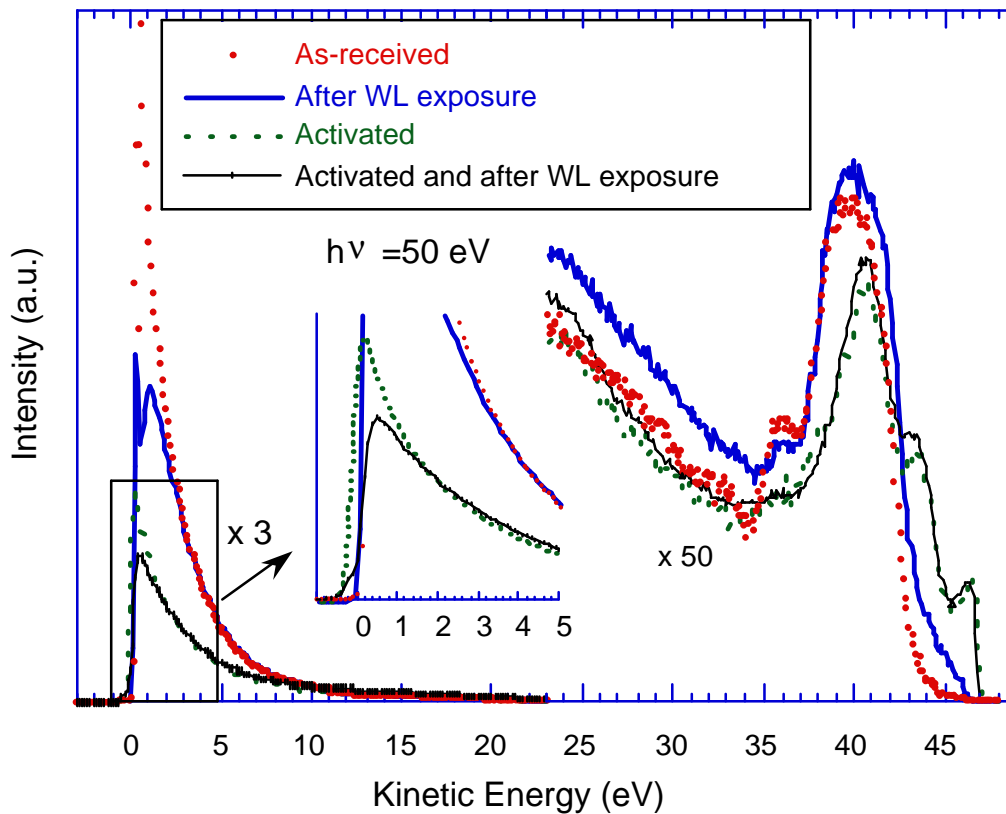


Figure 7 : Photoemission spectra of TiZr taken at 50 eV photon energy of: a) as-received; b) after WL exposure; c) after annealing in UHV (pressure during annealing was less than  $10^{-8}$  mbar) at  $330 \pm 30$  °C for 4 hours (TiZr-ac.), d) after WL exposure of the annealed surface.

To verify if the behaviour reported in Figure 7 is dependent on the chosen annealing conditions, the experiment was repeated on a second TiZr sample, prepared in an identical fashion to the first. The TiZr II sample was annealed following a slightly different procedure than the TiZr. It was annealed in a UHV pressure of less than  $5 \cdot 10^{-9}$  mbar and at  $300 \pm 30$  °C for 2.25 hours.

The spectra reported in Figure 8 show qualitative agreement with the spectra from the TiZr sample, discussed above. However, they clearly indicate that the detailed behavior of a surface is indeed strongly dependent on the fine details of its history and activation process.

The as-received spectrum from the TiZr II sample exhibited an even sharper secondary emission at low kinetic energies than that from the TiZr sample. This is due to the fact that the 50 eV spectrum reported here was the first spectrum measured from the surface, whereas the corresponding spectrum from the TiZr sample was taken after some monochromatic light exposure performed at 20, 30 and 50 eV photon energies (but before WL exposure).

The different annealing conditions result in observable changes in surface composition, and even if this second treatment exhibits a Fermi edge (indicative of essentially a metallic surface) the observed spectrum differs from the one reported in Figure 7. It should be noticed that the second annealing process produces a surface which also metallic but much more sensitive to SR irradiation: the secondary electron distribution exhibit larger changes in intensity and energy position (0.25 eV) reported in the inset of Figure 8 than the one reported in Figure 7.

These data clearly demonstrate the importance of defining a standard procedure to obtain reproducible surface conditions.

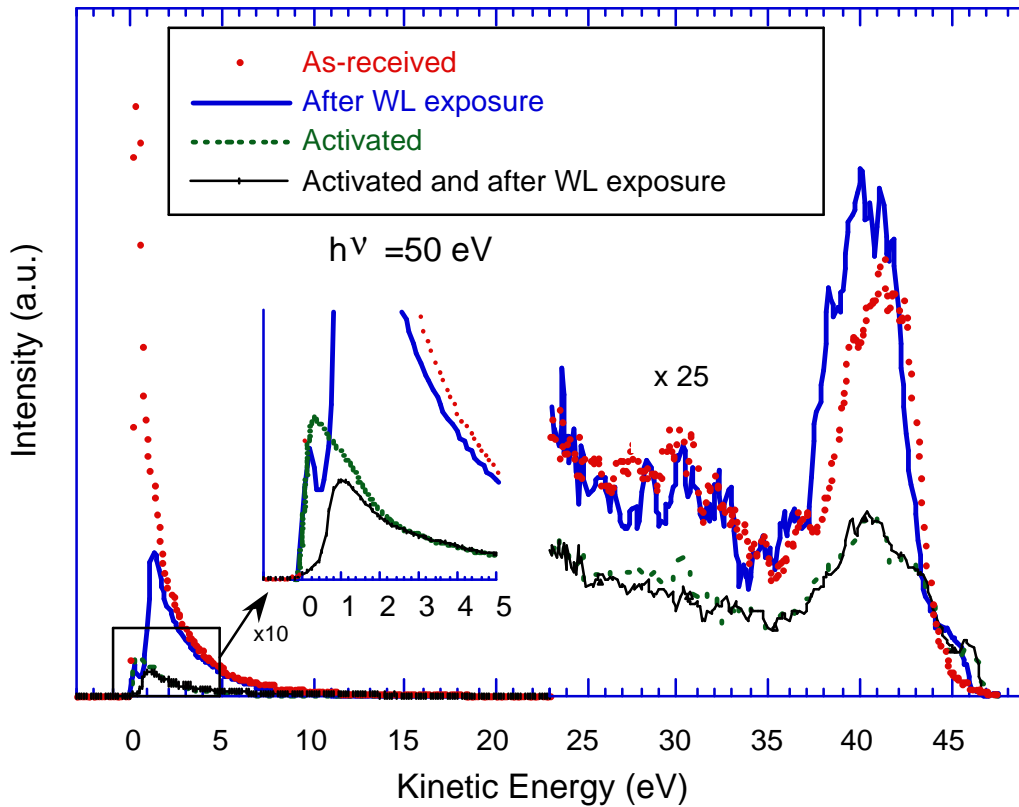


Figure 8 : Photoemission spectra of TiZrII taken at 50 eV photon energy of: a) as-received; b) after WL exposure; c) after annealing in UHV (pressure during annealing was less than  $5 \cdot 10^{-9}$  mbar) at  $300 \pm 30^\circ\text{C}$  for 2.25 hours (TiZrII-ac.), d) after WL exposure of the annealed surface.



### 3.2.4 Photoemission with monochromatic light: St 707.

A second type of non-evaporable getter material, namely St 707 getter alloy (Zr 70%; V 24.6%; Fe 5.4%), produced industrially by SAES Getters, has been studied. Figure 9 shows photoemission spectra taken at 50 eV photon energy from a St 707 sample in three conditions: a) as-received, b) after WL illumination ( $10^{16}$ - $10^{17}$  ph/mm<sup>2</sup>); c) after UHV annealing (pressure during annealing was less than  $5 \cdot 10^{-9}$  mbar) at  $300 \pm 30$  °C for 8 hours.

Once again, the as-received material exhibits a very high emission of very low energy electrons while the valence band signal is very weak and does not show any signal at the Fermi level.

Illumination with a WL dose of about  $10^{16}$ - $10^{17}$  ph/mm<sup>2</sup> induces quite dramatic changes to the surface. Not only, the secondary tail decreases significantly while the high energy electron emission increases by more than a factor 5 but no lineshape changes are observed.

As expected, annealing changes very significantly the surface from the appearance of a well defined Fermi level at 46.1 eV implying that activating the St 707 getter produces a metallic surface. This spectrum is very different from the one obtained just after WL irradiation, indicating that the WL exposure is not as efficient in cleaning the surface. Together with these changes in the high energy part of the spectrum a significant intensity reduction and a large energy shift in the low energy region is observed.

Subsequent irradiation of the activated getter with WL does not affect significantly the valence band while, as in the case of TiZr, the secondary tail shifts back to higher energy with a reduction in intensity.

In this case there is evidence that WL irradiation induces additional modifications of the activated surface, observed by the appearance of the double low energy cutoff after WL re-exposure. This is most probably due to the coexistence of differently cleaned regions during activation. This additional modifications induced by WL may be due to removal of 'light' contaminants from the surface which do not have consequences on the valence band line shape, but only in the cut off region.

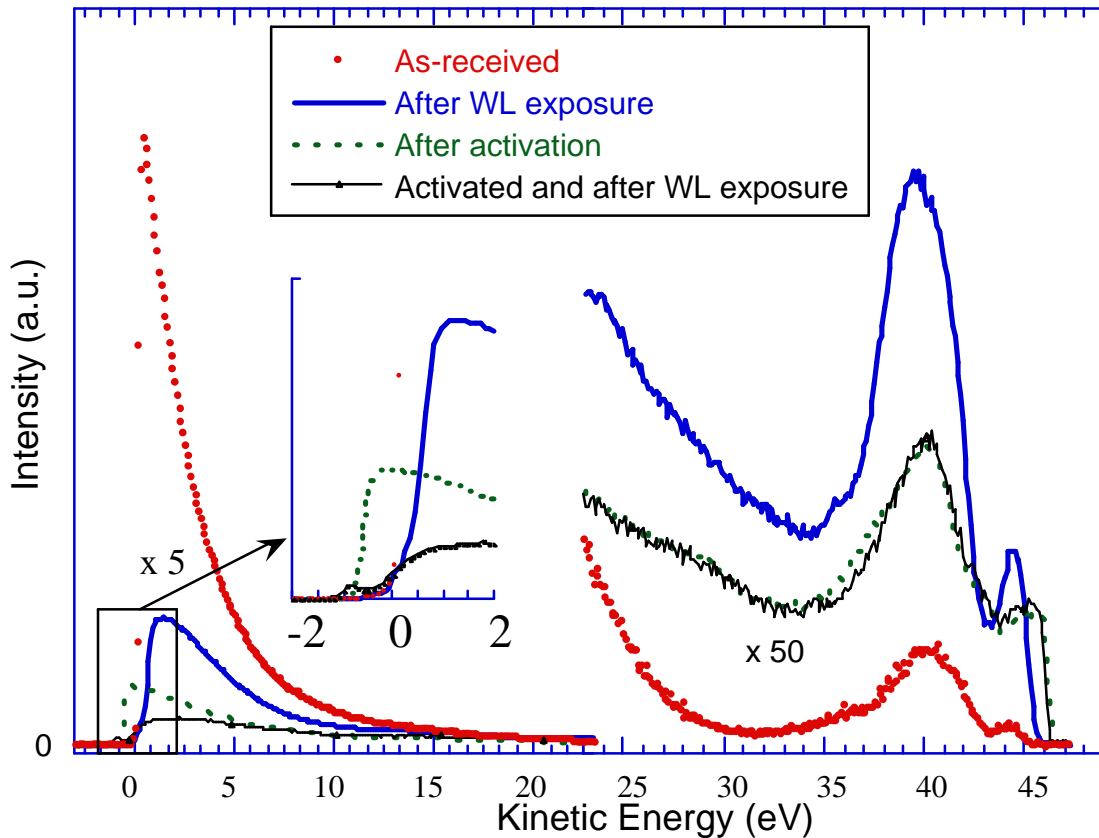


Figure 9 : Photoemission spectra of St 707 taken at 50 eV photon energy of: a) as-received; b) after WL exposure; c) after annealing in UHV (pressure during annealing was less than  $5 \cdot 10^9$  mbar) at  $300 \pm 30^\circ\text{C}$  for 8 hours (St 707-ac.), d) after WL exposition of the annealed surface.

### 3.2.5 Photoemission with monochromatic light: Pd.

Another interesting sample studied was Pd due to its affinity to  $\text{H}_2$ , potentially increasing the pumping speed of a NEG, once heated to about  $150^\circ\text{C}$  [15]. In Figure 10, the selected spectra obtained from such a sample according to different preparation procedures are shown.

This is the only case where WL exposure, with a dose of around  $10^{16}$ - $10^{17}$   $\text{ph}/\text{mm}^2$  does not change significantly the spectral shape both at low energy and in the valence band region. Only annealing affects the electron energy distributions, once again decreasing the secondary electron tail intensity and increasing the emission in the valence band region. The valence band region of the annealed sample, both before and after WL exposure, show a clear Fermi edge cut off at 46.5 eV, signature of an ‘almost clean’ metallic surface.

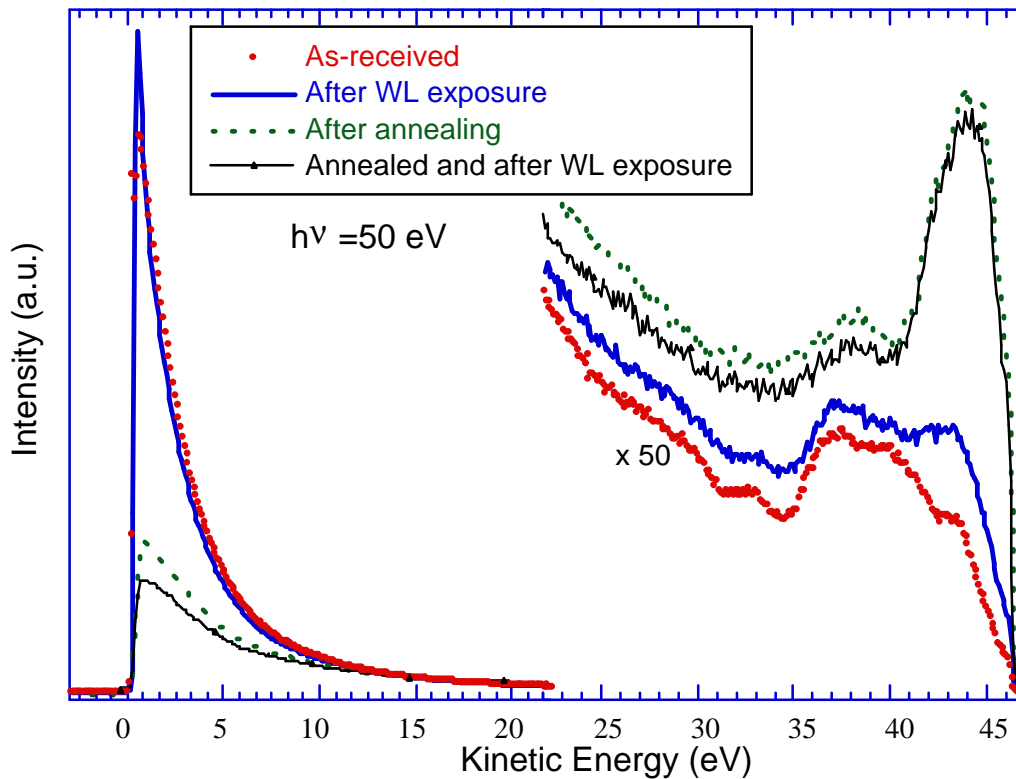


Figure 10 : Photoemission spectra of Pd taken at 50 eV photon energy of: a) as-received; b) after WL exposure; c) after annealing in UHV (pressure during annealing was less than  $5 \cdot 10^9$  mbar) at  $300 \pm 30^\circ\text{C}$  for 8 hours (Pd-an.), d) after WL exposure of the annealed surface.

### 3.2.6 Photoemission with monochromatic light: TiN.

The PEP-II low energy (positron) ring will operate at nominal energy of 3.1 GeV and current of 2.1 A [17]. A TiN coating has been chosen for the entire low energy ring in order to eliminate the predicted beam instabilities otherwise foreseen in that machine. For this reason experiments were performed on a TiN coated Al sample [16] which appeared black, carbon-like, and not Au-like characteristic of TiN. The sample was studied as-received, after  $\text{Ar}^+$  sputtering at 2 keV for 5 minutes in 1.7 Torr of Ar (estimated dose  $2 \cdot 10^6 \text{ Ar}^+ \cdot \text{cm}^{-2}$ ) and after dosing it with WL ( $10^{16}$ - $10^{17}$  ph/mm<sup>2</sup>).

The WL effect on the as-received sample was not studied and therefore we can only speculate that this sample will also behave in a similar way as the others studied.

The sputtering was used to obtain a 'cleaner surface'. The measured spectra, reported in Figure 11, show that the sputtering process changes the weight of the emitted electrons from the valence band to the low energy part of the spectrum. This effect is contrary to any of cleaning procedures used on the other samples analysed, perhaps requiring further investigations. However, it should be noted that the sample contains 43% oxygen and 22% carbon in the bulk [17] and therefore removal of the outermost layers by sputtering will not necessarily result in the exposure of a 'cleaner surface'. Not only the sputtering process affects the intensity of the electron

distribution but also the cut off energy, showing that the cleaning procedure reduces the work function of the surface by more than 2 eV.

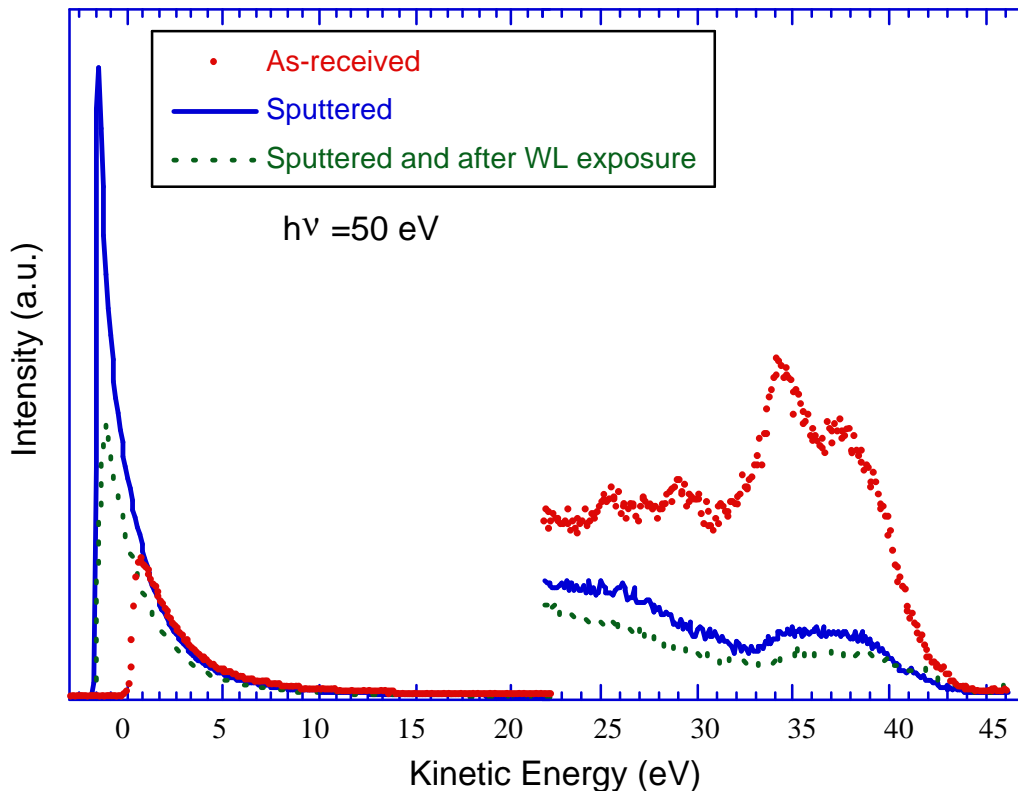


Figure 11 : Photoemission spectra of TiN taken at 50 eV photon energy of: a) as-received; b) after  $\text{Ar}^+$  sputtering (TiN-sp.), c) after WL exposure of the sputtered surface.

### 3.3 The photon energy dependence of the PY (Constant Final State spectroscopy).

In the arcs of the LHC the 44.1 eV critical energy SR, with a beam divergence of 0.55 mrad for 5 eV photons, will illuminate the beam screen at a mean incidence angle of 4.5mrad. Specular photon reflection from perfectly smooth surfaces, at such energies and incidence angles, might be expected to be close to 100 % [8]. Such calculations, at a given incidence angle, predict that the low energy photons (VUV) are reflected more readily than the higher energy photons (soft X-ray). Real surfaces with a finite roughness are far from being perfect reflectors in the VUV and a certain percentage of the incident photons will be adsorbed and/or diffused. Photons in the UV energy range can be reflected significantly from polished surfaces even when impinging at near normal incidence [18]. One may speculate therefore that a significant intensity of diffusely scattered UV light exists from a real surface. Experiments, using monochromatic light will be required to disentangle this question of photon reflection to produce a realistic estimate of the intensity and spectral distribution of the reflected and/or diffusely scattered photons. Nevertheless, it is expected that the spectrum of reflected and/or diffused WL will not resemble the one shown in Figure 2, but will be more weighted toward the low energy. This therefore leads to the importance of analysing the total yield of the different materials as a function of photon energy. In the case of a material which exhibits a very low cross section for photons with energies less than say 30 eV, then multiply

reflected/diffused light may be neglected. On the other hand multiply reflected/diffused photons will be an important consideration for a material exhibiting a high cross section for such low energy photons.

The availability of a tunable monochromatic light provides the possibility to study not only the energy distribution curves as a function of photon energy, as it was presented in section 3.1, but also the PY as a function of photon energy. The measurement was made by recording the number of secondary electrons emitted at  $1 \pm 0.2$  eV kinetic energy while scanning the photon energy from 18 eV to 110 eV. This so-called Constant Final State (CFS) spectroscopy is well known to mimic the shape of the total yield (*i.e.* drain current) emitted per monochromatic incident photon [10]. The total PY could not unfortunately be measured directly for technical reasons.

In Figure 12 CFS for four materials, namely: Au, Cu-el., TiZr-an. and Al are shown. It is evident that the four curves are different reflecting the different cross-sections of the atomic elements. While Au and Al are such that the maximum electron emission is caused by photons in the energy range between 30 and 100 eV, electrons from Cu el. and TiZr-an. are mainly caused by photons of 20 to 40 eV and 30 to 50 eV respectively. It is not possible here to quantify the implication of the available data for the LHC, nevertheless this preliminary study points to the need of a better understanding of the photon energy dependence of the emitted photoelectrons and of the specularly reflected and or diffusely reflected light.

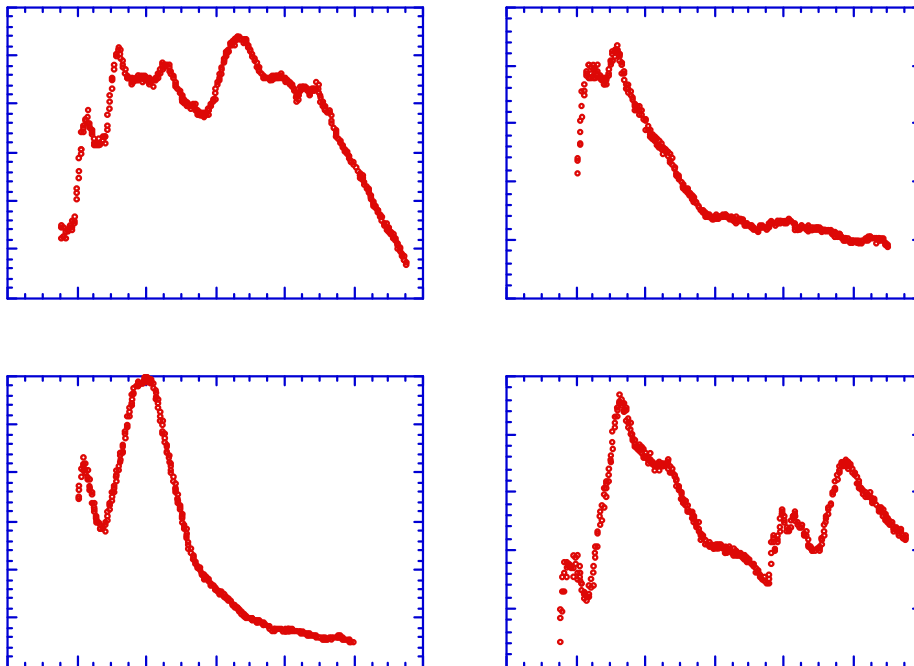


Figure 12: Constant Final State (CFS) spectra from Au, Cu-el., TiZr-an. and Al in the photon energy range from 18 to 110 eV. The collected electrons are emitted at  $1 \text{ eV} \pm 0.2 \text{ eV}$  kinetic energy.

### 3.4 White light photoelectron yields of the as-received samples

One of the important parameters relevant to the existence of a potentially serious electron cloud in the LHC is the PY per adsorbed photon. The measured drain currents and deduced PY of the as-received samples are indicated in Table 3. Also shown is the Au reference whose PY is in excellent agreement with literature [9]. It should be noted that the reported PY is that for the number of emitted electron per incident photon and not, as reported in other works, per adsorbed photon [19]. The error in the PY, given in Table 3, is estimated here only from the error in the photon flux (see section 2.3 for details). In Figure 13 these data are plotted as a function of decreasing PY. The plotted error bars are not the errors quoted in Table 3 (errors in determining the absolute flux) but are the reproducibility errors found from measuring two different TiZr samples (TiZr and TiZr II) and two different Cu-el. samples (the results from the second sample are not shown). Their PY were in agreement to within  $\pm 0.002$ . In other words, the absolute PY are accurate to within  $\pm 0.01$ , whereas the relative PY are accurate to within  $\pm 0.002$ .

Table 3 : PY per incident photon for the different as-received samples studied. In the second and third columns the sample drain currents and the last refocusing mirror drain currents are reported respectively; in the fourth column the estimated photon flux is given (see section 2.1 for details); in the last column the PY is reported with its absolute error.

Sample	I sample ( $\mu\text{A}$ )	I mirror ( $\mu\text{A}$ )	Photon flux (ph/s/mm <sup>2</sup> )	PY (e/ph)
Au	3.72	3.59	$5.6 \cdot 10^{14}$	$0.041 \pm 0.010$
Cu	3.8	2.4	$3.8 \cdot 10^{14}$	$0.063 \pm 0.010$
Cu-ab.	1.94	0.83	$1.5 \cdot 10^{14}$	$0.093 \pm 0.010$
TiN	8.3	4.2	$6.5 \cdot 10^{14}$	$0.080 \pm 0.010$
TiZr	2.4	1.1	$1.7 \cdot 10^{14}$	$0.088 \pm 0.010$
TiZr II	1.6	0.76	$1.2 \cdot 10^{14}$	$0.084 \pm 0.010$
Cu-el.	1.5	0.88	$1.4 \cdot 10^{14}$	$0.070 \pm 0.010$
Pd	0.87	0.48	$7.5 \cdot 10^{13}$	$0.072 \pm 0.010$
St 707	4.9	3.6	$5.7 \cdot 10^{14}$	$0.053 \pm 0.010$
Al	9.0	3.45	$5.4 \cdot 10^{14}$	$0.106 \pm 0.010$

Table 3 and Figure 13 show that the WL PY from the as-received surfaces exhibit values in the range  $(0.053 \text{ to } 0.106) \pm 0.010$ . The Al gives the highest PY twice that of the St 707. The fact that most of the PY appear not to differ substantially implies that the surface properties of these materials after the WL irradiation are similar. An explanation for this result may be that photo-cracking of similar molecules, found on all the as-received samples, produces similar fragments left behind on the surface after the intense WL SR. This surface layer will therefore smoothen out any expected differences in the PY of the different materials.

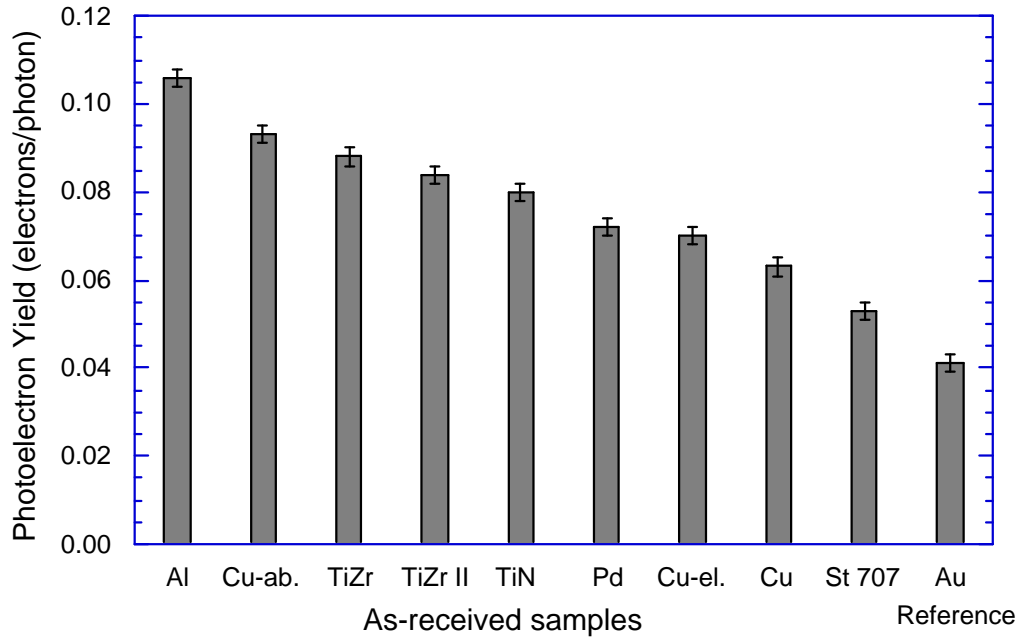


Figure 13 : Plot showing the PY during WL illumination, sorted as a function of decreasing values, for the different as-received sample.

It should be pointed out that the photon flux used to obtain these data was approximately 1000 times more intense than that expected in the LHC, therefore these reported PY are not representative of the initial PY to be expected in the LHC, rather they are those to be expected after a considerable time of machine operation (estimated to be 5 to 50 hours of LHC nominal operation). This aspect will be addressed in section 3.7. where reconstruction of ‘low dose’ white light spectra is presented.

### 3.5 White light photoelectron yields of the surface conditioned samples

For comparison, Figure 14 shows the measured PY of the as-received and conditioned samples. The most dramatic change on surface conditioning is seen for the Al sample where the PY increases by almost a factor of 8 after sputtering. Sputtering of the TiN sample has a similar, but less marked enhancement (50%), on its PY. This could be well explained by the decrease of work function observed in Figure 11. On the other hand, sputtering of the Cu sample causes a marginal reduction (16%) on its PY. As shown in section 3.2, sputtering can clean surfaces very efficiently by removing the outermost layers of the sample, resulting in the underlying material to become more exposed. The large variations between the PY of the WL exposed surfaces and the conditioned surfaces leads to the notion that the surfaces exposed to WL are effectively ‘passivated’ by their oxide layer.

Activation of TiZr, TiZr II or St 707 results in a reduced PY by between 38% and 32%. However, annealing of either the Pd or Cu-el. samples changes little the PY. The effect of activation of such NEG’s to 300°C partially removes the outermost surface oxide layers and therefore exposes the substrate material [20]. The applied annealing of the Pd and Cu-el. does not activate such surfaces, being insufficient to modify the outermost surface oxide layers and hence their PY.

The lowest PY of 0.035 was measured for the activated St 707 surface, being slightly lower than that of the evaporated Au reference surface.

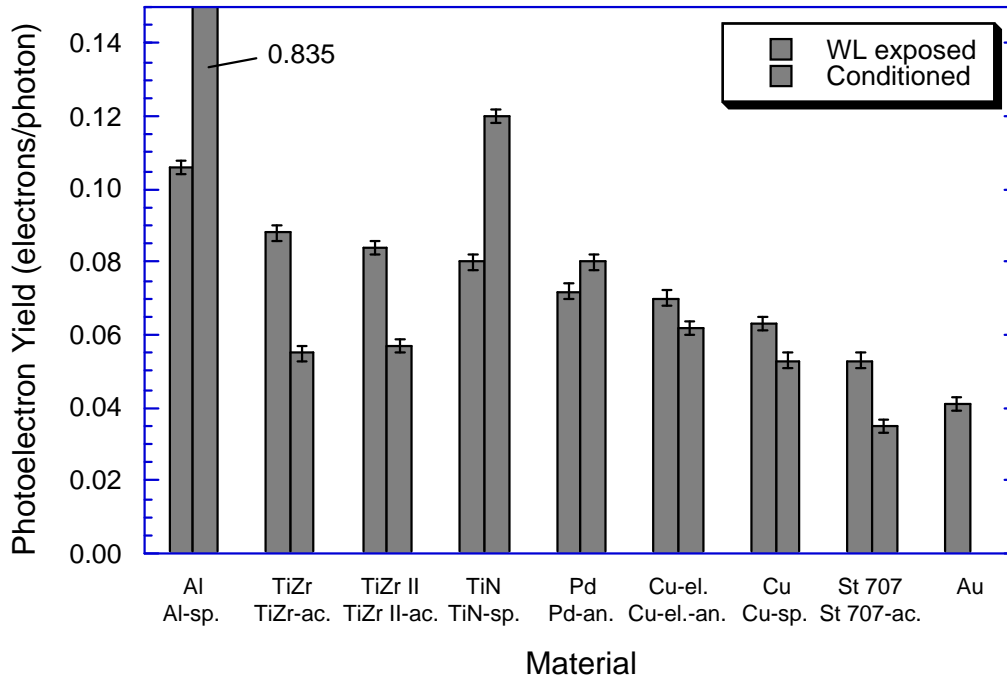


Figure 14 : WL PY for the samples in their as-received and conditioned states.

### 3.6 Energy Distributions of the WL EDC

In addition to the PY presented in section 3.4 & 3.5, the energy distribution of the PY is relevant to the LHC electron cloud phenomena (multipacting, activated gas desorption, heat loads *etc.*). Since the estimated heat loads and electron cloud growth are expected to be sensitive to the energy distribution of the emitted electrons [3, 4], realistic lineshapes should be included in the simulations. As an example the WL EDC for the Cu sample is shown in Figure 15. The most appropriate description of the energy distribution in this case, was found to be Lorentzian in nature centered at 0.64 eV and width,  $\sigma$ , of 3.7 eV. It is shown in the following that it is not possible to describe the photoelectron distributions of different materials with similar Lorentzians scaled to their PY intensities. In each separate case experimental input is required to obtain the realistic electron distribution to be used in the simulations.



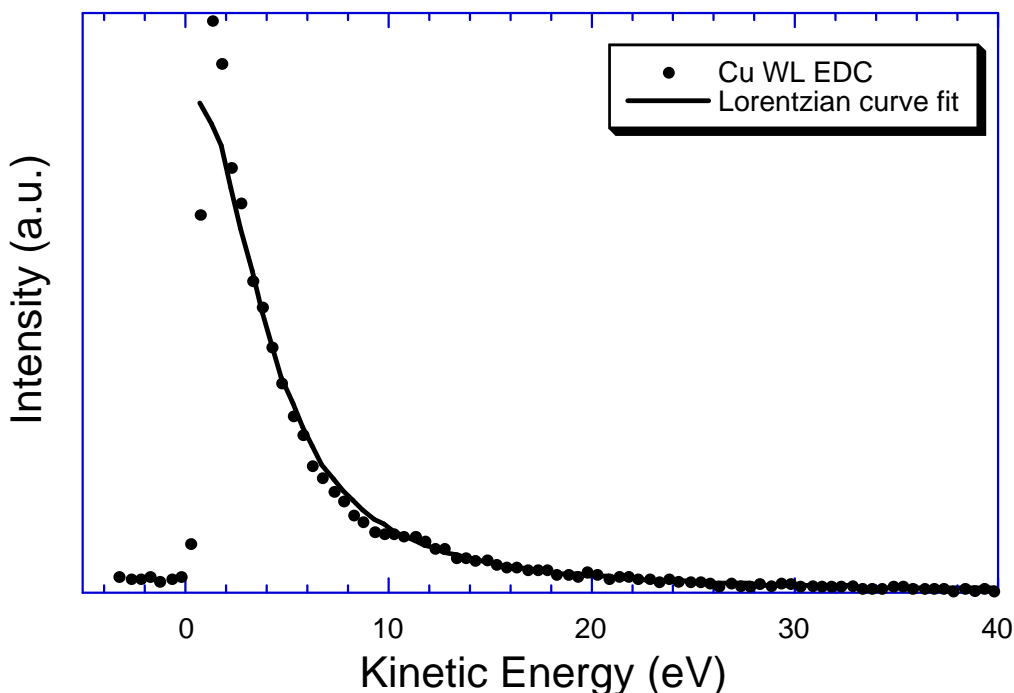


Figure 15: Lorentzian fit to Cu WL EDC.

In Figure 16 are shown the different WL excited energy distribution spectra for the different as-received samples studied (full line). There it is clear that the electron energy distributions are significantly different for the different samples, in contrast with the similar PY reported in section 3.4 and 3.5 and recalled in the insets of Figure 16. These WL exposed spectra clearly demonstrate that it is not only sufficient to measure the PY, but it is also important to measure the energy distribution of the emitted electrons. As an example, the Pd spectrum and the Cu-el. spectrum give nearly identical PY (0.072 and 0.070 respectively), but their energy distributions are considerably different; the former exhibits a sharp distribution of photoelectrons with mainly very low kinetic energies, the latter, together with a sharp emission of very low kinetic energy electrons, a broader peak is observed.

The WL EDC for the conditioned samples (dashed curves in Figure 16) demonstrate that significant changes in the lineshapes occur due to the conditioning. As an example, the WL exposed Cu-el. sample exhibits a sharp narrow peak at low energy and a broader peak, centered at 2 eV kinetic energy, with a characteristic tail to higher kinetic energies. After annealing the peak at 2 eV is shifted by 3 eV to higher energy and is considerably broadened.

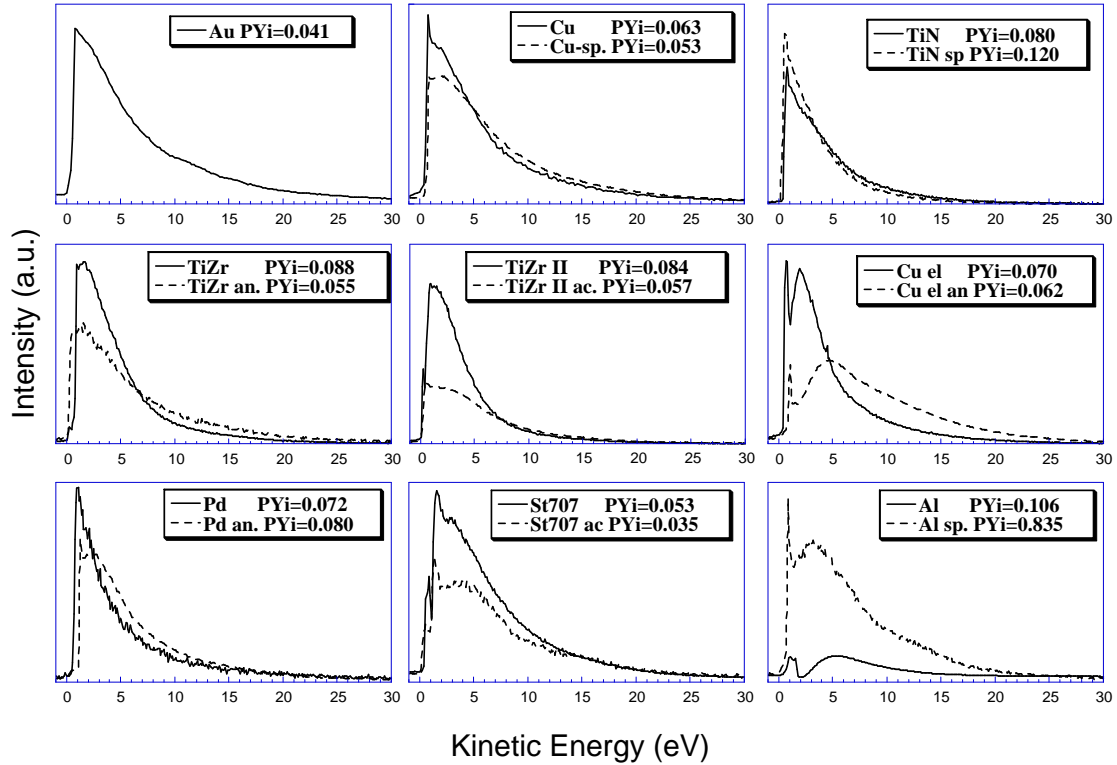


Figure 16: WL excited EDCs. The spectra from as-received (conditioned) samples are depicted with full (dashed) lines.

### 3.7 Generation of ‘low dose’ WL spectra.

At this point it is important to recall that the monochromatic spectra presented in section 3.2 suggest strongly that the WL spectra of the as-received and conditioned samples presented in the previous sections were subject to modifications due to the significant irradiation dose of WL during their acquisition ( $10^{16}$  to  $10^{17}$  ph/mm<sup>2</sup> corresponding to 5 to 50 hours of LHC nominal operation). In order to provide information on the initial PY expected in the LHC, *i.e.* before any significant SR radiation cleaning has occurred, a WL EDC can be reconstructed from a summation of the acquired monochromatic spectra, hereafter called a ‘low photon dose’ WL spectrum; such a photon dose corresponds to less than 1 minute of LHC nominal operation.

Such a spectrum, however, may be inaccurate due to the limited number of monochromated spectra from which it is generated. The analysis of the evaporated Au sample serves to test the validity procedure since clean Au should not exhibit any difference if the electron distribution has been measured before or after high photon dose irradiation. The resulting differences in the two spectra reported in the first graph therefore derive from the inaccuracy in the monochromated spectra summation. The observed agreement is quite good (the integral of the two curves differ by less than 13%) giving confidence to the procedure. These data allow us to determine an the absolute error bar on the value of the PY to be around 0.02 or less. In addition, lending further credance to the procedure, the reconstructed Cu spectrum from after a high photon dose, *i.e.* after WL illumination, is very similar to its corresponding WL EDC as can be seen in Figure 17 (second plot from the top, left panel) .

In Figure 17 a selection of a ‘low photon dose’ WL spectra are shown in comparison with their corresponding WL EDC from Figure 16. Here the ‘low photon dose’ WL spectrum has been normalized to the high energy tail of the WL EDC. It can be seen that significant differences

between the two spectra exist. In the case of the as-received Cu and TiZr the ‘low photon dose’ WL spectrum is dominated by a narrow and intense peak ( $< 1$  eV FWHM) at low kinetic energies which is almost totally absent in the latter (a small remanant structure is observed). Even once activated the TiZr ‘low photon dose’ WL spectrum has significantly more intensity below about 6 eV kinetic energy than the WL spectrum. These intensity and lineshape modifications can be attributed directly to sample cleaning by SR.

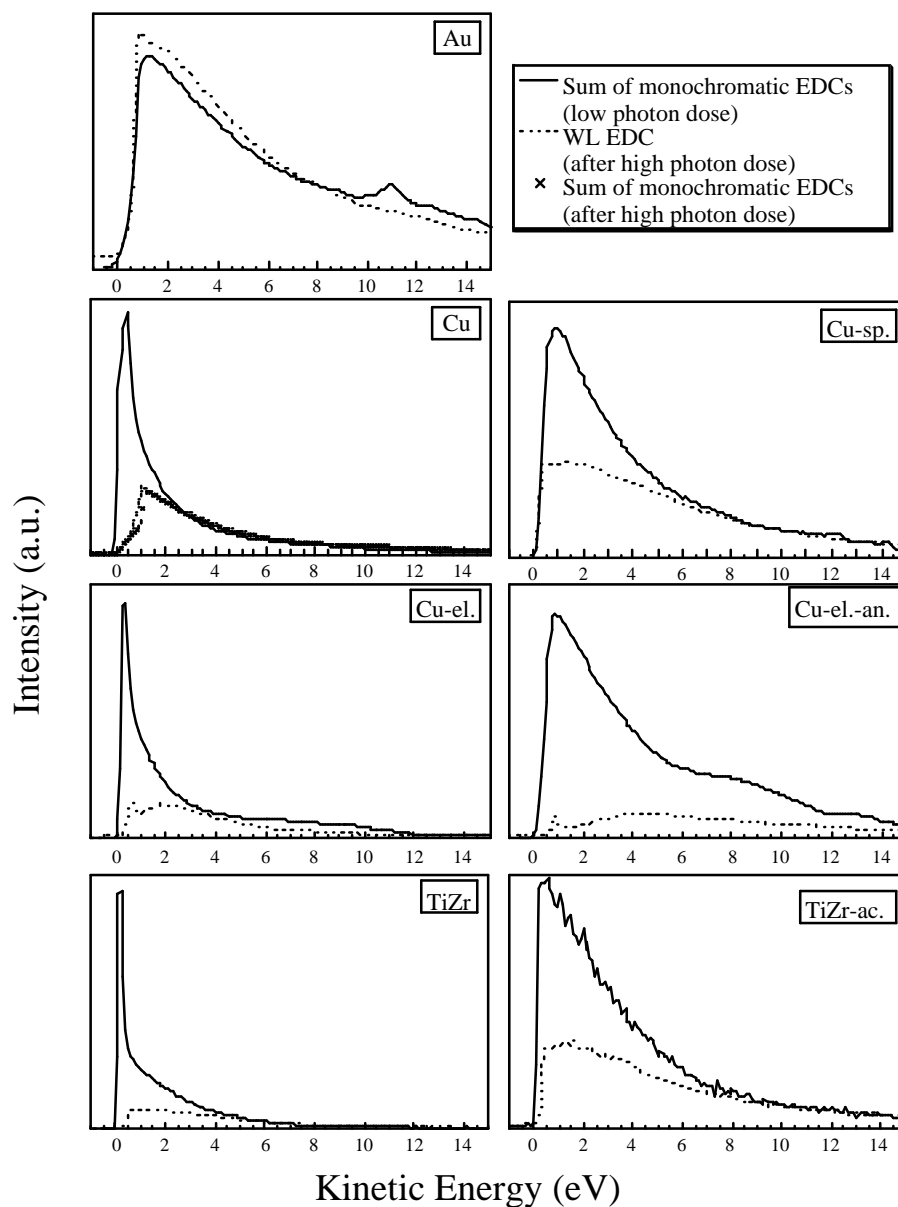


Figure 17 : Selected ‘low photon dose’ spectra generated from a summation of the monochromated EDCs (full line) compared with their corresponding WL EDC (dashed line). The sum of monochromatic EDCs (after high photon dose) is only shown for Cu (xxx line).

Figure 18 and Figure 19 most materials exhibit a more intense ‘low photon dose’ WL PY than that measured with WL. The most striking differences are seen for the as-received samples

where an intense narrow peak exists at low kinetic energy which is suppressed once irradiated with high dose WL. The air-baked Cu sample exhibits the strongest reduction in PY, by a factor of 5, whereas St 707 shows no significant modification, however, its EDCs are significantly different (not shown). The data shown in Figure 19, for the conditioned samples, indicate that WL induces additional modifications to the initial PY, reducing significantly the TiN-sp. and Cu-el-an surfaces. Indeed, in section 3.2.6 and 3.2.2 these surfaces were found not to be clean after conditioning.

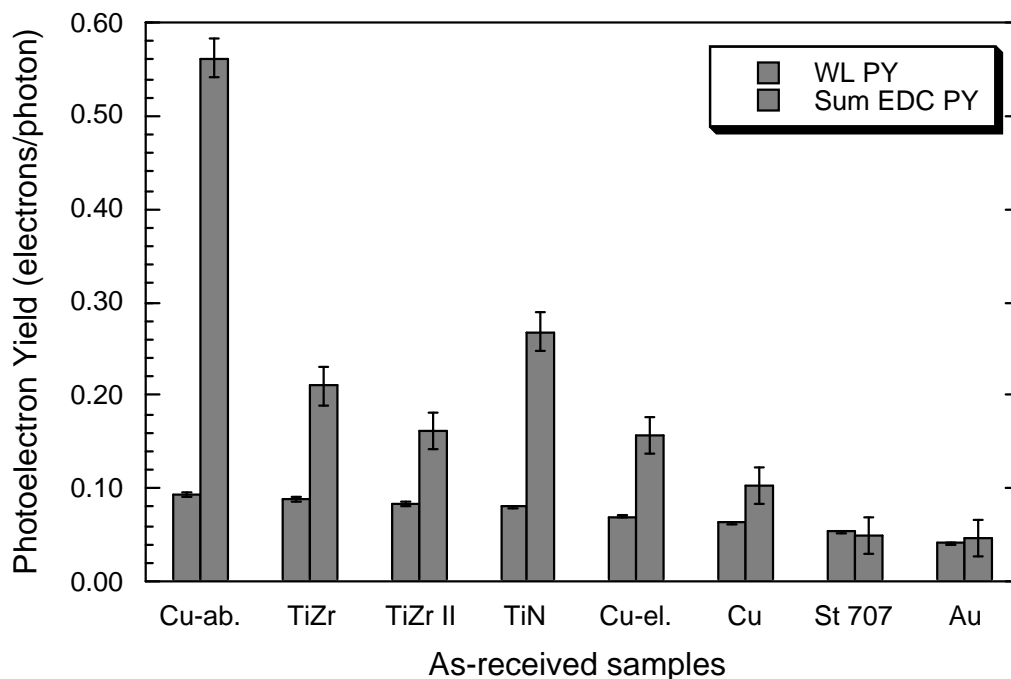


Figure 18 : Comparison between WL PY and its corresponding ‘low photon dose’ WL PY generated from a summation of the acquired EDCs for the as-received samples. Error bars are those generated from the Au reconstruction.

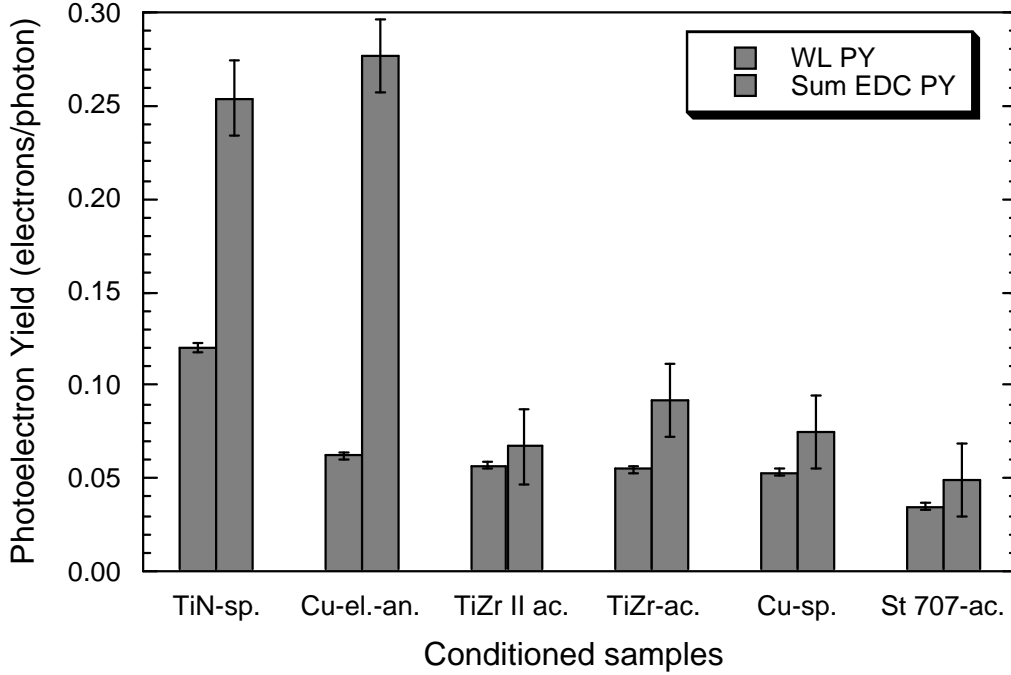


Figure 19 : Comparison between WL PY and its corresponding ‘low photon dose’ WL PY generated from a summation of the acquired EDCs for the conditioned samples. Error bars are those generated from the Au reconstruction.

### 3.8 The energy distribution of the secondary electrons in the LHC.

The results discussed in the previous paragraphs have consequences if applied to the analysis of the secondary electron induced yield (SY). This process, preceding the photoelectron emission, contributes to the electron-cloud instability build-up and has been carefully addressed in recent theoretical studies [3,4]. An electron created by photoemission will be accelerated, up to a mean energy of 380 eV [3] towards the opposite wall of the beam screen, due to the contemporary presence of the proton beam. This electron will then create a cascade of secondaries that may be further accelerated against the opposite walls. The study of the energetics and the time structure of the multipacting problem, suggest that only secondary electrons with kinetic energy between 1 and 6 eV may play a role in the build-up of the electron cloud instability [3,4].

Secondary electrons are emitted from a surface by impinging photons or electrons. Their origin is due to the multiple scattering of a photoexcited electron or of the incident electron within the solid. This multiple scattering causes a cascade of electrons towards lower energy, so-called secondary electrons. These analogous processes of secondary electrons creation suggest that they will result in secondary electrons with similar energy distributions. However, their intensities will be quite different due to the longer mean free path of the photon to that of the electron. Despite this assumption, requiring careful experimental verification, the present data can be used to extract information for the LHC electron cloud phenomenon.

Taking an integral curve of the EDCs, scaled to the PY, one can obtain a relative PY ( $\propto$ SY) in a defined energy window. Table 4 compiles the calculated yields in the energy window 1 to 6 eV

together with the total PY. From this table, always assuming  $PY \propto SY$ , it is evident that, if we should judge the annealing process as a system to change the heat load only from the total yield measured we should conclude, for example, that annealing of the Cu-el. sample will decrease it only by about 10% (assuming that the heat load scales linearly with the number of emitted electrons). The possibility of computing the number of electrons emitted in the energy window shows that annealing the Cu-el. reduces the heat load by about 50% which is a significantly different conclusion than that drawn from the total yields. However, this conclusion cannot be extrapolated to other materials. Indeed, if in the case of TiZr we measure about 25 % reduction in total yield corresponding to about 50% reduction in the emission in the 1 to 6 eV energy window. In the case of Pd, annealing causes a 10% increase both in the yield and in the number of electrons emitted in the 1-6 eV energy window.

Table 4 : Low photon dose WL total yield, reconstructed from a summation of the EDCs, high photon dose WL total yield measured with WL, and corresponding WL yield in the energy window, 1 to 6 eV, for the different samples studied.

Sample	'low dose' WL YIELD (e/ph)	Energy window 'low dose' WL YIELD (e/ph)	WL YIELD (e/ph)	Energy window WL YIELD (e/ph)
Au	$0.047 \pm 0.020$	$0.022 \pm 0.020$	$0.041 \pm 0.002$	$0.023 \pm 0.002$
Cu	$0.103 \pm 0.020$	$0.051 \pm 0.020$	$0.063 \pm 0.002$	$0.042 \pm 0.002$
Cu-sp.	$0.075 \pm 0.020$	$0.046 \pm 0.020$	$0.053 \pm 0.002$	$0.026 \pm 0.002$
Cu-ab.	$0.562 \pm 0.020$	$0.290 \pm 0.020$	$0.093 \pm 0.002$	$0.058 \pm 0.002$
TiN	$0.268 \pm 0.020$	$0.125 \pm 0.020$	$0.080 \pm 0.002$	$0.044 \pm 0.002$
TiN-sp.	$0.254 \pm 0.020$	$0.132 \pm 0.020$	$0.120 \pm 0.002$	$0.075 \pm 0.002$
TiZr	$0.210 \pm 0.020$	$0.102 \pm 0.020$	$0.088 \pm 0.002$	$0.055 \pm 0.002$
TiZr-ac.	$0.092 \pm 0.020$	$0.055 \pm 0.020$	$0.055 \pm 0.002$	$0.027 \pm 0.002$
TiZr II	$0.162 \pm 0.020$	$0.073 \pm 0.020$	$0.084 \pm 0.002$	$0.053 \pm 0.002$
TiZr II ac.	$0.067 \pm 0.020$	$0.037 \pm 0.020$	$0.057 \pm 0.002$	$0.027 \pm 0.002$
Cu-el.	$0.157 \pm 0.020$	$0.069 \pm 0.020$	$0.070 \pm 0.002$	$0.042 \pm 0.002$
Cu-el.-an.	$0.277 \pm 0.020$	$0.146 \pm 0.020$	$0.062 \pm 0.002$	$0.021 \pm 0.002$
Pd	-	-	$0.072 \pm 0.002$	$0.043 \pm 0.002$
Pd-an.	-	-	$0.080 \pm 0.002$	$0.049 \pm 0.002$
St 707	$0.049 \pm 0.020$	$0.026 \pm 0.020$	$0.053 \pm 0.002$	$0.029 \pm 0.002$
St 707-ac.	$0.049 \pm 0.020$	$0.017 \pm 0.020$	$0.035 \pm 0.002$	$0.020 \pm 0.002$
Al	-	-	$0.106 \pm 0.002$	$0.038 \pm 0.002$
Al-sp.	-	-	$0.835 \pm 0.002$	$0.362 \pm 0.002$

It is clear that more studies are required both in case of WL illumination EDC and electron bombardment induced EDC. Here we demonstrate the importance of studying the energy distribution of the photoemitted electrons and speculate that a similar experimental approach should be necessary to understand the acceptable performance of LHC candidate materials in terms of their electron induced production of secondaries.

## 4. Conclusions

SR photoemission has been shown to be a powerful tool to study surface properties of relevance for the LHC vacuum chamber optimisation. It has been shown how the energy and angular distribution of the emitted photoelectrons can be measured. Irradiation of samples with monochromatic light have shown that the total PY is mainly due to secondary electrons. Monochromatic studies, after intense WL irradiation, of the as-received samples have shown that the cleaning mechanism induced by photons produce a much less cleaner surface than annealing or sputtering. Indeed, most of the sputtered samples and all of the activated getter materials exhibit a Fermi edge or valence band electrons, signature of the metalicity of the surface. CFS spectroscopy has shown that the photon energy dependence of photoemitted electrons is not constant and greatly differs from sample to sample. Intense WL irradiation of as-received samples results in a similar PY for all the studied samples, Al gives the highest PY (0.106) whilst Au the lowest (0.041). An explanation that the WL exposed PY do not differ greatly between most of the samples may be the result of photocracking of the surface molecules leaving behind similar fragments on the surface. However, pre-conditioning of the samples with sputtering, annealing or activation induces substantial differences in the PY. Sputtering resulted in the most dramatic changes. Annealing of TiZr and St 707 reduced their PY; the latter to the lowest value of all the measured samples. Annealing of either Pd or Cu-el samples has no significant effect on their PY. EDCs of such as-received and conditioned samples were measured and demonstrated that there are no “universal curve” which may describe their shape. The secondary electron tail of the as-received WL irradiated Cu sample was found to be Lorentzian in nature. WL EDCs of samples subjected to a low photon doses were reconstructed from monochromatic spectra. As-received samples exhibit an intense and narrow secondary electron tail whereas surface conditioned samples exhibit a broader secondary tail. All the samples (except clean gold) exhibit a higher PY than WL irradiated one. Assuming similar EDCs for electron induced secondary emission and for the measured photo-induced ones, we show that it is important to consider the detailed energy distribution of the emitted electrons. Such energy distribution is a key parameter in selecting the final vacuum chamber material and its detailed surface preparation. Also we note that an initially performant candidate material may strongly change its emission properties during machine operation due to WL irradiation and/or electron/ion bombardment. This observation, reported for the first time in this context, implies that a careful study of such changes will be necessary in order to guarantee a long term and reliable performance of the machine.

Further experiments should be carried out with the sample at cryogenic temperatures (i.e. including beam screen working conditions) where it is expected that cryosorbed molecules will induce strong modifications to the electron emission. Since photon reflection (specular and diffused) is an important issue for the electron cloud in the LHC photon energy dependant studies should be made. In addition, EDC's for electron stimulated secondary electron emission should be studied to provide better input to the understanding of the electron cloud problem. A measurement protocol, given scientific objectives as defined by the electron cloud instability as well as cryogenic cooling capacities and beam lifetime dominated by the gas density, has to be built in order to improve the powerfulness of this experimental method. Also, a devoted experimental set up [21] should be made available and be used also to study materials (like Be, Al, or NEG) different from the one of the beam screen which may be used in other parts of the machine.

## **Acknowledgments**

We gratefully acknowledge the financial support from the INFN–LHCLDS project, the CERN-INFN agreement n° K434/LHC and EC-LSI (HMC12) and BESSY light source for access to the TGM7 beamline. We would like to thank C. Benvenuti, P. Chiggiato and V. Ruzinov for supplying us with TiZr, Cu-el. and Pd samples; and R.E Kirby for supplying us with the TiN coated Al sample.



## References

- [1] “Comparaison of photodesorption yields using synchrotron radiation of low critical energies for stainless steel, copper, and electrodeposited copper surfaces”. J.Gómez-Goñi, O.Gröbner, and A.G.Mathewson, *J.Vac.Sci.Technol. A***12**(4), 1714(1994).
- [2] “The LHC Vacuum System”. O.Gröbner, *Vacuum* **47**, 6-8, 591 (1996) and PAC proceedings, Vancouver (1997).
- [3] “Simulations for the Beam-Induced Electron Cloud in the LHC Beam Screen with Magnetic Field and Image Charges”. O.S.Brüning, LHC-Project-Report-158, November 1997.
- [4] “The Electron Cloud Effect in the Arcs of the LHC”. M.A.Furman, LBNL-41482/CBP Note 247 or LHC-Project-Report-180, May 1998.
- [5] “VUV Synchrotron Radiation Photoemission Investigation of Proposed Materials for the Vacuum Chambers of the Large Hadron Collider”. I.R.Collins, A.G.Mathewson and R.Cimino, CERN Vacuum Technical Note 97-24, September 1997 and ECASIA’97 proceedings, Göteborg, Sweden (1997).
- [6] “The LEP pre-injector as a multipurpose facility”. J. P. Potier, L. Rinolfi. CERN-PS-98-016-LP and presented at EPAC’98, Stockholm, Sweden, (1998).
- [7] “Research at BESSY, A User’s Handbook”. second edition December 1995. Berliner Elektronenspeicherring-Gesellschaft für Synchrotronstrahlung mbH Lentzeallee 100, 14195 Berlin.
- [8] [http://www-cxro.lbl.gov/optical\\_constants/](http://www-cxro.lbl.gov/optical_constants/) and references therein.
- [9] “Schottky type photodiodes as detectors in the VUV and soft x-ray range”. M. Krumrey, E. Tegeler, J. Barth, M. Krisch, F. Schäfers, and R. Wolf. *Appl. Opt.* Vol. 27, No. 20,4336 (1988).
- [10] “Photoemission in solids”. Edited by L. Ley and M. Cardona, Springer-Verlag 1979
- [11] “Photon-stimulated desorption yields from stainless steel and copper-plated beam tubes with various pretreatments”. C. L. Foerster, H. Halama, and C. Lanni, *J. Vac. Sci. Technol. A* **8**(3) 2856 (1990).
- [12] “Gas desorption from an oxygen free high conductivity copper vacuum chamber by synchrotron radiation photons”. O. Gröbner, A. G. Mathewson, and P.C. Marin, *J. Vac. Sci. Technol. A***12**(3) 846 (1994).
- [13] “Non-evaporable getter films for ultrahigh vacuum applications”. C. Benvenuti, O. Chiggiato, F. Cicoira, and Y. L’Aminot, *J. Vac. Sci. Technol. A* **16**(1) 148 (1998).
- [14] “Work Function of Metals”. J. Hölzl and F.K. Schulte in *Solid Surface Physics* Springer-Verlag 1979.
- [15] C. Benvenuti, private communication.
- [16] “TiN coating of the PEP-II low-energy ring aluminium arc vacuum chambers.” K. Kennedy, B. Harteneck, G. Millos, M. Benapfl, F. King and R. Kirby. SLAC-PUB-7535 or LBNL 39862R1, May 1997 and PAC proceedings, Vancouver 1997.
- [17] R.E.Kirby, private communication.

- [18] “Neutral gas desorption and photoelectric emission from aluminum alloy vacuum chambers exposed to synchrotron radiation”. O.Gröbner, A.G.Mathewson, P.Strubin, E.Alge and R.Souchet. *J.Vac.Sci.Technol.* **A7**(2), 223(1989).
- [19] “Photoelectron yield and photon reflectivity from candidate LHC vacuum chamber materials with implications to the vacuum chamber design”. V. Baglin, I.R. Collins and O. Gröbner. CERN-LHC-Project-Report-206 and presented at EPAC’98, Stockholm, Sweden (1998).
- [20] “An investigation of Non Evaporable Getter films for Ultra High Vacuum applications”. D. Lawton, CERN Technical Report 1998.
- [21] “Proposal for the construction of an INFN-CERN-Sinicrotrone Trieste Beamline At ELETTRA”. R. Cimino *et al.* To be published.