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Background Radiation Measurements at High Power Research Reactors

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Abstract

1 Research reactors host a wide range of activities that make use of the intense neutron fluxes generated at these
2 facilities. Recent interest in performing measurements with relatively low event rates, e.g. reactor antineutrino detec-
3 tion, at these facilities necessitates a detailed understanding of background radiation fields. Both reactor-correlated
4 and naturally occurring background sources are potentially important, even at levels well below those of importance
5 for typical activities. Here we describe a comprehensive series of background assessments at three high-power re-
6 search reactors, including γ -ray, neutron, and muon measurements. For each facility we describe the characteristics
7 and identify the sources of the background fields encountered. The general understanding gained of background pro-
8 duction mechanisms and their relationship to facility features will prove valuable for the planning of any sensitive
9 measurement conducted therein.

Location	Reactor	Thermal power (MW)	Latitude	Longitude	Altitude (m)	Fast Neutron Flux
NIST	NBSR	20	39.13° N	77.22° W	105	1.0
ORNL	HFIR	85	35.93° N	84.31° W	259	1.1
INL	ATR	110	43.59° N	112.96° W	1435	3.1

Table 1: Facility parameters, including reactor thermal power, geographic location, and predicted fast neutron fluxes relative to NBSR.

1. Introduction

Research reactors have for decades been important facilities for an enormous variety of activities including, but by no means limited to, isotope production, transmutation, materials and reactor studies, teaching and training, and fundamental physics investigations [1]. More than 250 research reactor facilities are operational or planned in 57 countries [2]. The large neutron flux generated by a controlled fission chain reaction enables such activities. Typically experiments are conducted within or close to the reactor core, or using neutron beams generated through the thermal moderation and collimation of fission neutrons. With the addition of specialized moderators at cryogenic temperatures, neutron beams with flux rates on the order of $10^9 \text{ cm}^{-2}\text{s}^{-1}$ can be produced and efficiently guided tens of meters away from the reactor to both reduce reactor related backgrounds, i.e. high-energy γ -rays and fast neutrons, and provide more space for experiment deployment [3].

Recently, there has been renewed interest in using research reactors as a source for another product of the fission process: electron antineutrinos ($\bar{\nu}_e$). On average approximately six $\bar{\nu}_e$ result from each fission reaction in a reactor via the beta decay of neutron rich daughter nuclei. The ability to site a $\bar{\nu}_e$ detector close to a research reactor would enable a sensitive search for additional sterile neutrino states suggested as an explanation for anomalous results in several neutrino oscillation experiments [4, 5], the observation of coherent neutrino nuclear scattering (CNNS) [6, 7], and testing the hypothesis of neutrino-induced decay-rate modulation [8, 9]. A measurement of the reactor $\bar{\nu}_e$ energy spectrum performed at a research reactor fueled by ^{235}U would help constrain uncertainties in the prediction of reactor $\bar{\nu}_e$ emissions and may provide additional information on short-lived daughter states that contribute to decay heat uncertainties [10]. Using measurements of research reactor $\bar{\nu}_e$ emissions is also of interest for nuclear safeguards and non-proliferation applications, allowing verification of operator declarations [11, 12, 13].

Conducting a reactor $\bar{\nu}_e$ measurement in a research reactor facility, however, carries a significant challenge. In contrast to typical activities performed at such facilities, the expected signal event rates are low (100s to 1000s of events per day for ton-scale detectors). Therefore strong suppression and an excellent understanding of all background sources is required. To obtain the broadest sensitivity to the possible existence of additional neutrino states [5] and to maximize the event rate for a $\bar{\nu}_e$ energy spectrum measurement, such an experiment should be placed as close to the reactor core as practical. At this close proximity γ -rays and neutrons produced by reactor operation can not be neglected, and indeed may be the dominant background source. This in is contrast to most reactor $\bar{\nu}_e$

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37 experiments (e.g. [14, 15, 16]), which are sited 10s to 1000s of meters from the reactor core(s) of interest. Similar
38 considerations apply to searches for CNNS and nuclear decay-rate modulation experiments.

39 In preparation for PROSPECT [10, 17], we have therefore conducted a comprehensive background radiation sur-
40 vey at three reactor facilities in the US. Our goal is to characterize the background radiation fields generally encoun-
41 tered at research reactor facilities, to understand the sources of those backgrounds, and to develop background miti-
42 gation strategies appropriate for low-background experiments generally. While obviously essential for the planning
43 and execution of PROSPECT, we expect this study to provide valuable insight into background sources, intensities,
44 and mitigation strategies for other research reactor facility users.

45 The outline of this paper is as follows. In Sec. 2 we describe the research reactor facilities examined in this study,
46 highlighting features that influence the background fields encountered. In Sec. 3 we describe the instruments used to
47 perform background measurements at each of the selected facilities and the results obtained in Sec. 4. In Sec. 5 we
48 use the measurements to illustrate characteristics of reactor-correlated backgrounds at these facilities. Finally, using
49 the understanding of the background radiation fields gained, we describe the steps taken to mitigate reactor-correlated
50 backgrounds in a prototype detector deployment in Sec. 6.

51 **2. Reactor Facilities**

52 The National Institute of Standards and Technology (NIST) [18], the Oak Ridge National Laboratory (ORNL)
53 [19], and the Idaho National Laboratory (INL) [20] operate powerful, highly compact research reactors and have
54 identified potential sites for the deployment of compact $\bar{\nu}_e$ detectors at distances between 4–20 m from the reactor
55 cores. Important parameters for these facilities are summarized in Table 1. While designed for a variety of purposes,
56 all three are research reactors with active user programs. All use similar Highly Enriched Uranium (HEU) fuel and
57 operate at typical peak thermal powers of 20 MW, 85 MW, and 110 MW respectively. While having much lower
58 power than typical commercial reactors, the availability of sites at short baselines roughly compensates in terms of
59 available $\bar{\nu}_e$ flux. Importantly, these facilities operate single cores with refueling and maintenance outages of signif-
60 icant length, thus allowing precise characterization of natural background during reactor off periods. Nonetheless,
61 placing extremely sensitive detectors at such short baseline locations requires careful assessment of both natural and
62 reactor generated background radiation.

63 The potential deployment sites at these facilities include locations as close as practical to the reactor core (“near”,
64 5–10 m) and at slightly greater separation (“far”, 15–20 m). A wide variety of measurements have been performed
65 at each location, as will be described in later sections. In the following, we describe the general features of these
66 locations with a focus on those relevant to the background measurements performed. Broadly speaking, backgrounds
67 at the near locations exhibit significant reactor correlations since they are as close to the reactor as practical, while at
68 the far locations that have more intervening shielding and typically greater separation from plant systems there is little
69 or no reactor correlation observed.

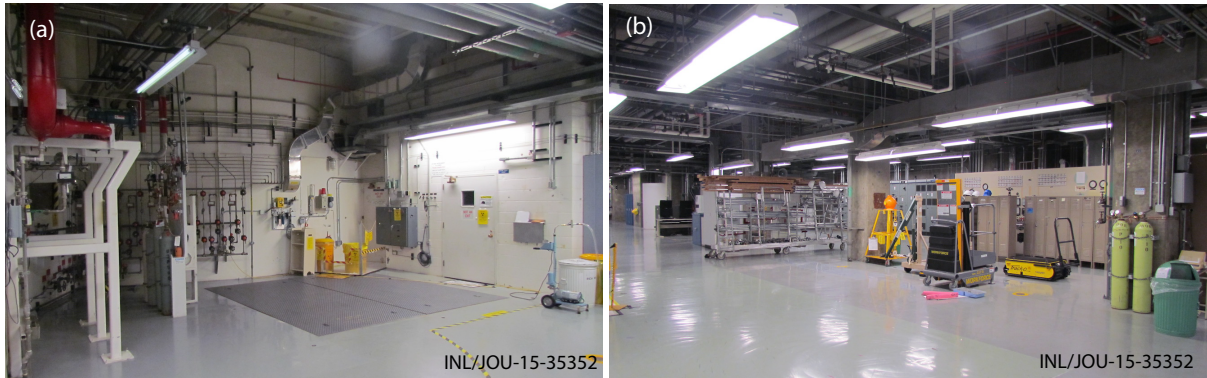


Figure 1: Photographs of the near (a) and far (b) locations studied at ATR.

70 At all facilities considered here the thermal neutron flux at the periphery of the reactor vessel varies by less than
71 10% over the course of a cycle with the reactor held at nominal power. While this variation may result in background
72 rate variations within the facility, other background generation mechanisms described later (e.g. scattering from beam
73 lines) will also result in time dependent variations.

74 In addition to variations in background due to site design, variation in cosmogenic background rates are expected
75 due to differences in facility location and elevation (Table 1). Tools developed for Single Event Upset (SEU) pre-
76 dictions can be used to estimate the relative fast neutron flux at each location, relative to a reference location [21].
77 These estimates predict a minor cosmogenic background difference between the National Bureau of Standards Reac-
78 tor (NBSR) at NIST and the High Flux Isotope Reactor (HFIR) at ORNL, while the higher altitude of the Advanced
79 Test Reactor (ATR) at INL leads to a significantly higher cosmogenic neutron flux, absent any shielding or enhance-
80 ment effects due to the local surroundings or potential overburden. The actual overburden available at these facilities
81 will depend upon the precise location and size of a deployed detector since it will depend in detail upon the facility
82 layout (floor and roof thicknesses, wall locations and thicknesses, etc). The muon measurements presented in Sec. 4.6
83 provide an indication of the relative cosmogenic flux at each location, accounting for construction details and altitude
84 effects.

85 2.1. ATR Locations

86 ATR is a light water moderated reactor that uses fuel made from a U_3O_8 -Al dispersion clad in and burnable ^{10}B
87 poison aluminum and has reactivity control elements composed of beryllium and hafnium. While the data presented
88 here were collected during a typical cycle with a thermal power of ≈ 110 MW, this parameter and the power distri-
89 bution within the core can be varied based upon the needs of in-core experiments. Some cycle-to-cycle variation in
90 background can therefore be expected, but given limited available measurement time this possibility is not explored
91 in this work. The ATR core is sited below-grade and adjacent basement levels contain the possible detector deploy-
92 ment locations. While this below-grade siting provides some cosmogenic attenuating overburden, this is offset by the
93 higher cosmic ray rate encountered at 1435 m elevation of the facility. Cosmogenic background rate measurements

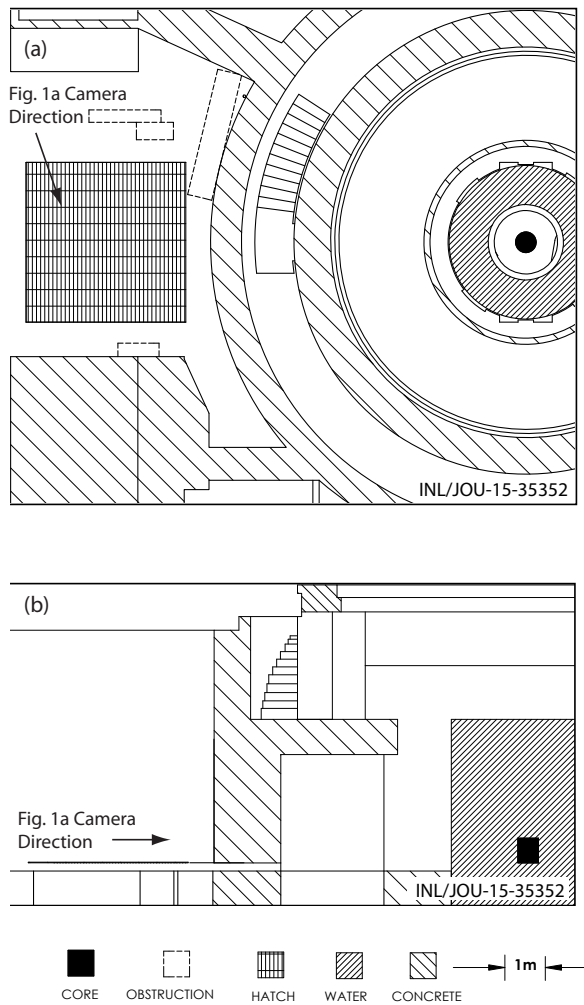


Figure 2: (a) Plane and (b) elevation views of the ATR near location.

94 were therefore of particular interest at this site.

95 Background measurements were performed in the following locations at ATR (Figs. 1 and 2):

96 • **Near: First sub-basement hatch area**

97 Located ≈ 6 m below-grade, this site is directly beneath a large service hatch that provides crane access to the
 98 lower levels of the facility. Therefore, despite being below-grade, there is relatively little overburden provided
 99 by the facility structure directly above. Several plant systems containing small amounts of reactor primary
 100 coolant are located in this area.

101 • **Far: Second sub-basement storage area**

102 This below-grade location is a possible far detector location. It is located ≈ 12 m below-grade, and is a relatively
 103 open location used for equipment storage and staging. There is one significant plant system that passes through

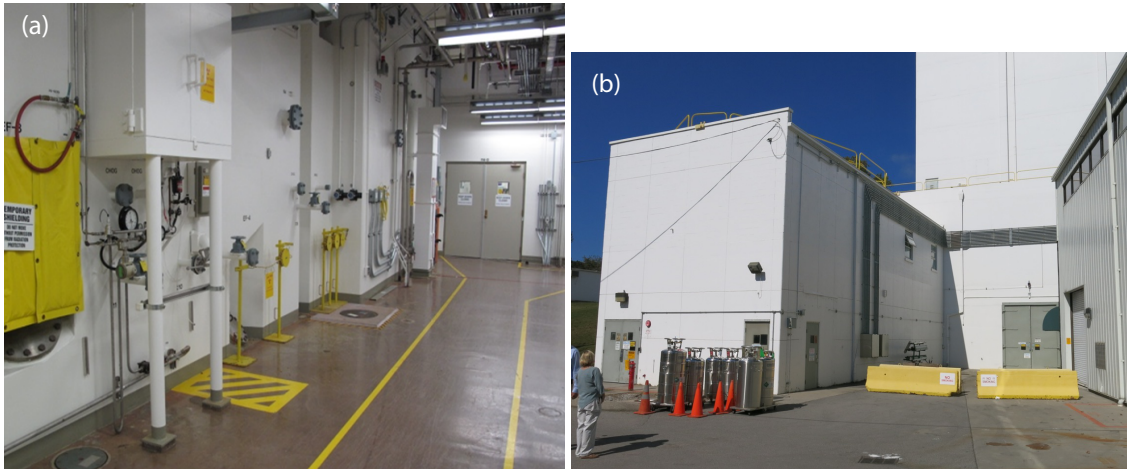


Figure 3: Photographs of the near (a) and far (b) locations studied at HFIR.

104 this area: a ceiling mounted pipe that returns a small amount of primary coolant from a ^{16}N power monitoring
 105 system to the main coolant loop.

106 2.2. HFIR Locations

107 HFIR is a light water moderated reactor that uses fuel made from a $\text{U}_3\text{O}_8\text{-Al}$ dispersion and burnable ^{10}B poi-
 108 son clad in aluminum. The fuel elements are surrounded by a beryllium neutron reflector. Thin cylindrical control
 109 elements containing europium, tantalum, and aluminum are used to maintain a constant thermal power of 85 MW
 110 throughout each reactor cycle. The HFIR reactor core is located near grade at an elevation of 259 m. Background
 111 measurements were performed in two locations at HFIR (Figs. 3 and 4):

112 • Near: Experiment hall

113 The potential location is in a broad corridor on the building level above the core. Large concrete blocks on
 114 the level below provide substantial shielding, but the shielding wall at the deployment level is considerably
 115 thinner and contains a number of penetrations. In addition, there is a shielded enclosure at this location used
 116 intermittently for measurements of activated gas. Measurements were taken at several positions throughout this
 117 location.

118 • Far: Loading area

119 The potential far location is located outdoors on the level below the near location, adjacent to an area in which
 120 neutron beam experiments are conducted. Measurements were taken either in the outdoor location or, during
 121 inclement weather, in an adjacent steel clad building that supplied little overburden.

122 2.3. NBSR Locations

123 NBSR is a heavy water moderated reactor that uses fuel made from a $\text{U}_3\text{O}_8\text{-Al}$ dispersion clad in aluminum. The
 124 NBSR building is located slightly above grade at an elevation of 106 m. Background measurements were performed

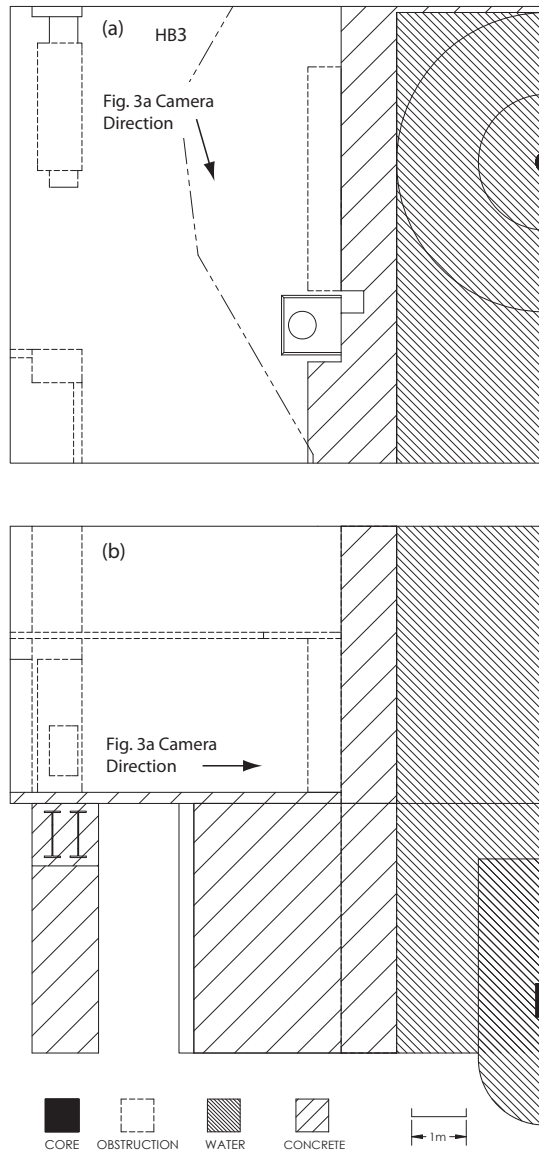


Figure 4: (a) Plan and (b) elevation views of the HFIR near location. The location of the HB3 beamline on the floor below is indicated in the top panel.

125 in two locations at NBSR:

126 • **Near: Thermal column**

127 This location is an area within the NBSR confinement building immediately adjacent to the biological shield
 128 surrounding the core. This location was designed to provide high-flux thermal neutron beams (thermal column),
 129 but is currently decommissioned. The moderator and shielding for these sources are still in place however, and
 130 result in very low neutron penetration from the core. Photographs and cross-sections of the thermal column
 131 shielding and possible detector location are shown in Figs. 5 and 6. Unique to the NBSR site, this location has

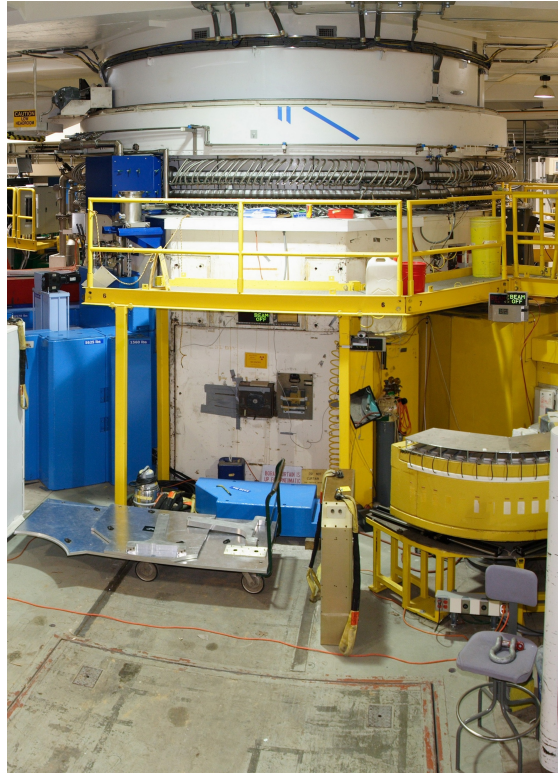


Figure 5: Photograph of the near location studied at NBSR.

132 neutron scattering instruments to either side. These are sources of both thermal neutrons and prompt γ -rays
133 arising from neutron capture. Above the thermal column area is a cooling water manifold that is part of the
134 biological shield cooling system. As described in the sections to follow, this is a source of γ -rays from ^{16}N that
135 illuminates roughly half of the potential detector area. Measurements have been taken under a variety of reactor
136 and adjacent instrument operating conditions.

137 • **Far: Loading area and high-bay**

138 There are two potential far locations at NBSR: outside the confinement building and in a high-bay area adjacent
139 to the confinement building. Since Health Physics surveys indicated no reactor-correlated background in these
140 locations, representative measurements of naturally occurring radiogenic and cosmogenic background were
141 carried out in a laboratory space nearby.

142 **3. Background Measurement Techniques**

143 All of the locations studied have low radiation fields from a Health Physics perspective (that are typically of
144 greatest concern to reactor operators and users), i.e. personnel dose rates are low. For context, an area is typically
145 designated a “radiation area” and will need controlled access given a $50 \mu\text{Sv/h}$ (5 mrem/h) γ -ray field. At 2.5 MeV this
146 corresponds to roughly $1 \times 10^3 \text{ cm}^{-2}\text{s}^{-1}$, which represents a significant flux relative to expected neutrino interaction

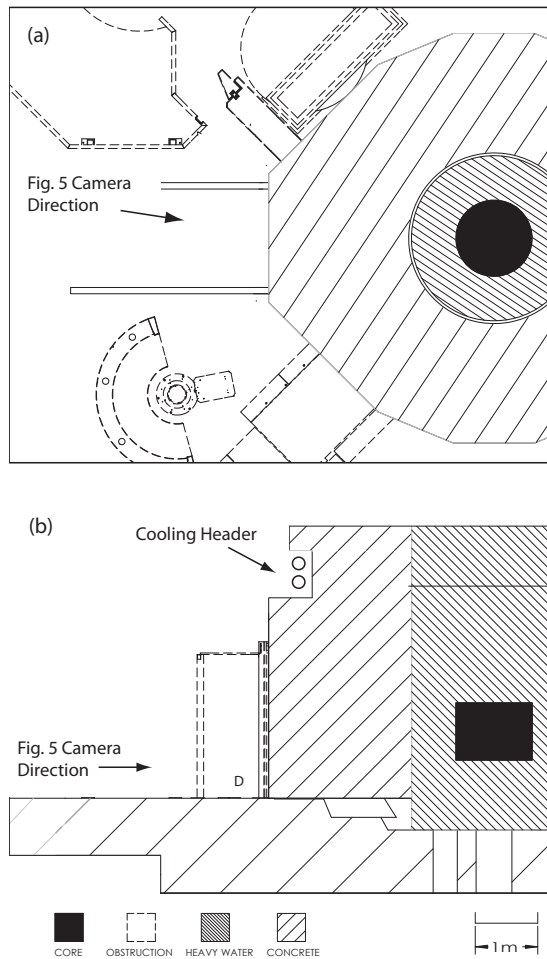


Figure 6: (a) Plane and (b) elevation views of the NBSR near location.

147 rates. In many cases, the background levels of interest in this study are beneath the sensitivity of common radiation
 148 survey instruments. Therefore, we have assembled a suite of measurement instruments with greater sensitivity, the
 149 ability to provide spectral information, and the ability to separately measure the most important backgrounds for
 150 low-background experiments.

151 Where possible, compact portable instruments were selected that could be easily transported between the three
 152 facilities, so as to provide a robust relative background comparison. These relative measurements are augmented by
 153 higher resolution or higher precision absolute flux measurements using devices available for use at only one or two
 154 locations. Such measurements proved valuable in determining the sources of particular backgrounds. The types of
 155 measurement performed during this survey were high and low resolution γ -ray spectroscopy, fast and thermal neutron
 156 flux measurements, muon flux measurements, and fast neutron spectroscopy. Where possible and appropriate, angular
 157 and spatial variations of the background fields have been measured. In particular, we have sought to localize γ -ray
 158 sources corresponding to particular site features like piping or shield wall penetrations, since in principle it is possible

159 to substantially reduce such sources using localized shielding.

160 3.1. γ -ray Measurements

161 Interaction of γ -rays is likely to dominate the singles rate in a $\bar{\nu}_e$ detector in the locations examined. Typically, low-
162 background experiments are most concerned with γ -rays with energies up to 2.614 MeV emitted by small amounts of
163 radioisotopes such as ^{60}Co , ^{40}K , and ^{208}Tl found in construction materials. These γ -ray emissions can be effectively
164 controlled through shielding and careful material selection and screening. For operation at a research reactor, there
165 are several important differences. First, the relatively compact spaces available and floor loading limitations at a
166 research reactor facility constrain the shielding that can be used. Second, short-lived radioisotopes with high-energy
167 γ -ray emissions can be present in reactor facilities due the higher neutron fluxes present. Of course, at all sites
168 under consideration there is significant shielding incorporated in the facility design that eliminates direct transport of
169 radiation from the reactor core. Nonetheless, there are several mechanisms that can produce elevated γ -ray background
170 rates. These include:

- 171 • **Local neutron interactions.**

172 Fast and thermal neutrons transported through shielding or scattered from beams can interact with material in
173 the local environment. In particular, neutron interactions with water and iron in structural steel can result in
174 high-energy γ -ray emissions.

- 175 • **Activation product transport in plant piping.**

176 Short-lived radioisotopes produced in water exposed to high neutron fluxes near the reactor core can be trans-
177 ported in plant piping to locations close to the measurement locations.

- 178 • **γ -ray transport through shielding.**

179 The shielding between a location of interest and a high intensity background source (e.g. pipe carrying activated
180 water) may not attenuate the emitted γ -ray flux to levels comparable with natural background. Seams or piping
181 penetrations in shielding walls may allow a scattering path for γ -ray (and neutrons) that results in a localized
182 “hot-spot”.

183 Because each of these sources would be expected to depend on specific features of a site, detailed characterization
184 of the γ -ray flux present in each location is necessary. Spectroscopic studies identifying particular radioisotopes
185 present and surveys indicating emission locations can aid in determining the production mechanism.

186 A variety of γ -ray spectroscopy instruments were used for these measurements. The same easily-transported
187 NaI(Tl) moderate-resolution crystal spectrometer was used at each site, with measurements providing a robust relative
188 comparison of the overall rate and the general shape of the energy spectrum at each location. The NaI(Tl) detector
189 used was a Bicron [22] model 2M2 (2” right cylindrical crystal size). A Bridgeport qMorpho [23] Data Acquisition

190 (DAQ) system was used in a Multi-Channel Analyzer (MCA) mode to collect measured spectra in the 0–12 MeV
191 electron equivalent energy range.

192 Higher resolution instruments supplied by the host institutions were used for more complete characterization
193 of the γ -ray fields at each location. While the differences in efficiency between these instruments precluded direct
194 comparisons between collected spectra, the higher energy resolution allows specific γ -ray lines to be identified.

195 At ATR, a 2" LaBr₃(Ce) scintillation detector (St. Gobain BrillanCe [24]) was used. This material has good
196 detection efficiency and a very good resolution for an inorganic crystal (\approx 3% FWHM at 662 keV). An Ortec Di-
197 giBASE [25] MCA was used for spectra collection.

198 At NBSR, a Canberra High-Purity Germanium detector (Model CPHA7.5-37200S) was used. The crystal is a
199 closed-end coaxial geometry of 55 mm length and 62.5 mm diameter. The health-physics group at the NIST Center
200 for Neutron Research (NCNR) carries out regular energy calibrations of the detector and these data are combined
201 with Monte Carlo calculations to determine absolute efficiency as a function of energy. The photo-peak efficiency at
202 6 MeV is roughly 0.025%.

203 At HFIR, a standard n-type High-Purity Germanium detector with 15% intrinsic efficiency was used to perform
204 high-resolution background measurements.

205 In addition, several centimeters of configurable lead shielding was used at both HFIR and NBSR to make di-
206 rectional measurements of γ -ray fields. As will be discussed in subsequent sections, collimated measurements were
207 useful in locating specific hot spots at these sites.

208 3.2. Neutron Measurements

209 Neutron backgrounds at each site were also measured with several instruments. Dose measurements were recorded
210 primarily as a simple method of characterizing thermal neutrons. Particular attention was focussed on fast neutrons
211 as these are an important background for a reactor $\bar{\nu}_e$ measurement that can mimic the signature of inverse-beta decay
212 (IBD) events. Fast neutron backgrounds at each site predominantly fall into two categories: cosmogenic fast neutron
213 and reactor-related fast neutrons.

214 Cosmogenic neutrons can be produced in the atmospheric column above the detector, in structures (e.g. buildings)
215 surrounding the detector, or within the detector itself, especially in high-Z passive shielding materials. These neutrons
216 range in energy from thermal to many GeV with a spectral shape that is reasonably well known. Figure 7 shows the
217 most recent JEDEC standard spectrum for fast neutrons (JESD89a) at sea-level [26], based upon measurements by
218 Gordon *et al.* using an array of Bonner spheres [27].

219 Work by Kowatari *et al.* has shown that the shape of the cosmogenic neutron spectrum varies little between dif-
220 ferent locations [28]. However, the total flux of these neutrons, particularly the thermal part of the spectrum, depends
221 on the local conditions present, including altitude, geomagnetic cutoff, solar activity, weather, and the presence of
222 high-Z material that may cause spallation from high-energy particles in cosmic ray showers including muons and fast
223 neutrons.

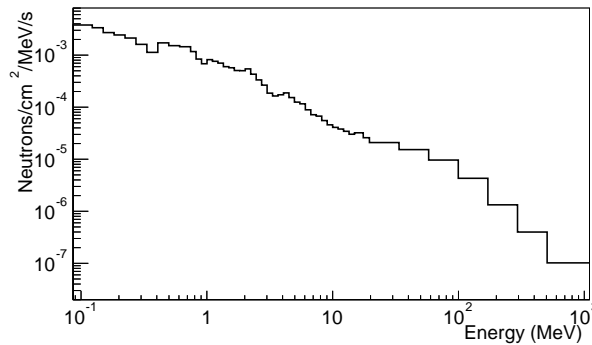


Figure 7: The JEDEC standard fast neutron spectrum recorded at sea level in New York [26].

224 Reactor-correlated fast neutron backgrounds can be expected to have different characteristics. Neutrons produced
 225 in the reactor will follow a fission spectrum, with very few neutrons expected at energies >10 MeV. Since the reactors
 226 are surrounded by moderating material and, of course, no line of sight exists between active fuel elements and the
 227 locations under consideration, any reactor-correlated neutron can be expected to have undergone multiple scattering
 228 interactions and therefore to have degraded energy. This fast neutron source is of less concern with regard to $\bar{\nu}_e$ -
 229 mimicking correlated background but could still be a significant source of singles background from neutron capture
 230 in an antineutrino detector. Similar to the preceding discussion of localized γ -ray sources, reactor-correlated neutron
 231 background can be expected to correspond to penetrations, beam instruments, or other shielding leakage paths.

232 3.2.1. Fast Neutron Measurements

233 As with the γ -ray background measurements, two classes of instruments were used to assess the fast neutron
 234 backgrounds at the potential reactor sites. First, a well characterized fast neutron spectrometer, FaNS-1, was used
 235 at NBSR and HFIR to validate the assumption that the higher-energy (>1 MeV) portion of the cosmogenic neutron
 236 energy spectrum varies little between locations and to measure the absolute flux at those sites. This device was
 237 difficult to transport, so it was not used at ATR. In addition, measurements were taken at the three facilities using a
 238 small portable fast neutron recoil counter. Due to its small size and operating method, this device could not readily
 239 provide absolute flux and spectral information but, analogous to the use of the NaI(Tl) γ -ray detector, it provides a
 240 robust basis for a relative site comparison.

241 The FaNS-1 spectrometer is a fast neutron detector consisting of segments of BC-400 plastic scintillator with ^3He
 242 proportional counters positioned between [29, 30]. Each of the six optically-decoupled plastic scintillator segments
 243 are $9\text{ cm} \times 18.5\text{ cm} \times 15\text{ cm}$, for a total active volume of 15 liters. Light from each segment is collected by pairs
 244 of Photo-Multiplier Tubes (PMTs) attached to cylindrical light-guides. The signal from each PMT passes through an
 245 asymmetric splitter circuit that produces two signals, a delayed full-amplitude signal and one attenuated by a factor of
 246 nine. Each pulse pair was waveform digitized. This approach allows for the construction of a linear response over a
 247 large dynamic range. The six 1" diameter ^3He proportional counters are filled with 4.0 bar ^3He and 1.1 bar of natural

248 krypton and have high thermal neutron capture efficiency. All six helium-counter signals were combined in one fan
249 in/out module and digitized.

250 FaNS-1 operates via the concept of capture-gated spectroscopy. A fast neutron enters the detector, where it
251 thermalizes through multiple (n,p) scatters. After thermalizing, it randomly walks until it is captured by a ^3He counter
252 or leaves the volume. Thus the signature of a fast neutron is a scintillator signal followed by a delayed neutron capture.
253 The neutron energy is determined by the quantity of light detected in the PMTs. Segmentation reduces the effect of
254 non-linear light yield and improves the energy resolution of the spectrometer. The energy response of the device was
255 determined via irradiation with well calibrated ^{252}Cf , 2.5 MeV and 14 MeV generator sources and detailed MCNP
256 models of detector response. By examining the time separation between a scatter-like event and a capture-like event,
257 it is possible to differentiate accidental coincidences, which are uniform in time, and correlated coincidences, which
258 have a distinct exponential distribution. The rate of accidental coincidences is driven by the product of the ^3He trigger
259 rate and the scintillator signal rate. If either of these sustains a substantial increase from non-fast neutron interactions,
260 e.g. γ -ray interactions in the scintillator or thermal neutron captures in the ^3He tubes, it will degrade the ability for
261 FaNS-1 to determine the fast neutron rate. Backgrounds such as these limited the measurements that could be carried
262 out at the NBSR near location.

263 Additionally, a small stilbene detector was taken to all three reactor sites. This device comprised a 2" trans-stilbene
264 crystal, a 2" PMT, and a Bridgeport eMorph DAQ system packaged in a small aluminum tube. Power and readout
265 were supplied via a USB connection to a laptop computer. Trans-stilbene is an organic crystal with good Pulse Shape
266 Discrimination (PSD) properties that allow for the identification of fast neutron recoil events. The crystal used for
267 these measurements was grown in a materials development laboratory at the Lawrence Livermore National Laboratory
268 (LLNL) [31].

269 3.2.2. Neutron Dose and Thermal Measurements

270 Neutron dose measurements were performed at NBSR and HFIR using neutron survey instruments. The detector
271 used at NBSR consisted of a 23 cm diameter cadmium-loaded polyethylene sphere surrounding a BF_3 tube. The
272 cadmium loading is designed to create a neutron response such that the instruments directly reads dose rates in units
273 of rem/hr ($\approx 2.7 \times 10^5 \text{ cm}^{-2}\text{s}^{-1}$ for thermal energies [32]). A thermal spectrum was assumed in the conversion to flux.
274 The HFIR detector was similar, except that it was calibrated to read in s^{-1} and an absolute efficiency was provided by
275 the instrument manufacturer. Since the two similar instruments were primarily used for assessing spatial variations,
276 no attempt was made to perform relative or absolute response calibrations. Calibrated bare BF_3 tubes were also used
277 to measure the approximate thermal neutron flux at several locations.

278 3.3. Muon Measurements

279 Cosmic ray measurements are important as they indicate the amount of overburden provided by reactor buildings
280 and other structures at these shallow sites. The cosmic muon rate and angular dependence was measured using a
281 telescope detector comprised of 4 scintillation paddles spaced at varying distances vertically as shown in Fig. 8.

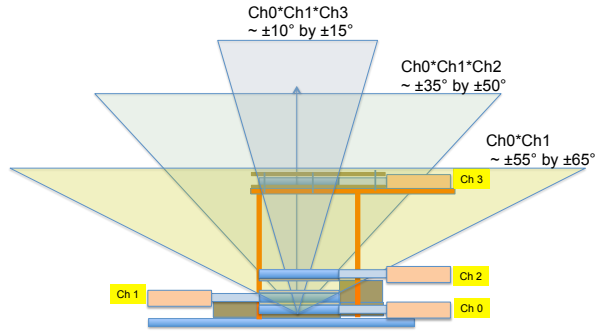


Figure 8: The angular acceptances for the muon telescope instrument used at all sites is determined by the coincidence requirement enforced between the 4 plastic scintillator paddles.

282 Requiring coincidences between the paddles is equivalent to restricting the muon angular acceptance of the telescope.
 283 Up to three angular ranges (approximately $\pm 10^\circ \times \pm 15^\circ$, $\pm 35^\circ \times \pm 50^\circ$, and $\pm 55^\circ \times \pm 65^\circ$) can be measured at the same
 284 time.

285 The $25 \text{ cm} \times 15 \text{ cm} \times 2.5 \text{ cm}$ scintillator paddles were constructed from Eljen EJ-200 plastic scintillator. Each
 286 paddle was connected to a 5 cm ADIT B51D01 PMT by a trapezoidal shaped acrylic light guide. A Bridgeport
 287 Instruments hvBase-P-B14D10 provided the high voltage and served as the voltage divider for the 10-dynode chain.
 288 Data acquisition was provided by a Bridgeport Instruments qMorpho-2010 ADC. The DAQ had the capability to
 289 record individual waveforms and could individually control the gain settings of each PMT channel. The qMorpho
 290 contains four 20 MHz multichannel analyzers with 10 bit resolution and was controlled through a USB interface.

291 This simple telescope cannot discriminate between particle type as it can only identify particles creating time
 292 coincident hits in the separated scintillator paddles. Cosmic rays near sea level are an admixture of muons, hadrons,
 293 electrons, photons and neutrons. According to the Review of Particle Properties (Particle Data Group) the integral
 294 rate of muons $\geq 1 \text{ GeV}/c$ is $\approx 60\text{--}70 \text{ m}^{-2}\text{sr}^{-1}\text{s}^{-1}$ and follows a roughly $\cos^2\theta$ angular distribution [33]. The number of
 295 electrons and positrons is very energy dependent with rates of $30 \text{ m}^{-2}\text{sr}^{-1}\text{s}^{-1}$ above 10 MeV, $6 \text{ m}^{-2}\text{sr}^{-1}\text{s}^{-1}$ above 100
 296 MeV, and $0.2 \text{ m}^{-2}\text{sr}^{-1}\text{s}^{-1}$ above 1 GeV. Protons and neutrons $\geq 1 \text{ GeV}/c$ add $\approx 0.9 \text{ m}^{-2}\text{sr}^{-1}\text{s}^{-1}$. Even the small amount
 297 of material in the roof over a typical laboratory space can reduce the observed rate by a few percent consistent with
 298 low energy electron fraction quoted above. Adding 3 mm of lead between paddles reduces the overall coincidence rate
 299 in the lab by roughly 5%. The roofs over the confinement buildings at both the NBSR and HFIR are $\approx 0.5 \text{ m}$ concrete
 300 and reduced the coincident rate by $\approx 17\text{--}19\%$. Our data show that the scintillator energy spectra are consistent with
 301 minimum-ionizing tracks and follow a roughly $\cos^2\theta$ angular distribution as expected from muons. Therefore we will
 302 assume that the rates measured with the telescope near the reactor are due to muons with a small $\leq 5\%$ contribution
 303 from other particles.

304 4. Background Measurement Results

305 In this section we describe the results obtained from measurement campaigns at each reactor site.

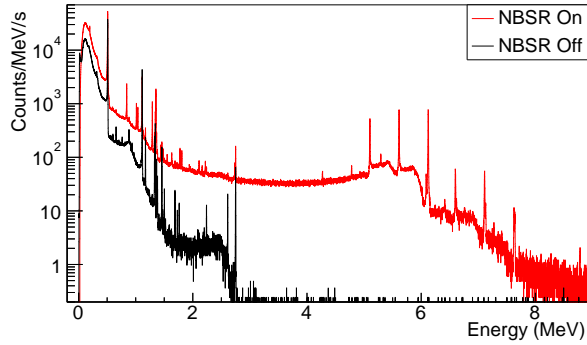


Figure 9: Example HPGe γ -ray spectra taken with the NBSR on and off. Prominent lines, and associated escape peaks and Compton continua, are evident. The line sources are identified in Table 2.

Isotope	Reaction	Source Material	Energy (keV)	$t_{1/2}$
^{187}W	$^{186}\text{W}(n,\gamma)^{187}\text{W}$	Unknown	479.5	23.9h
-	annh.		511.0	
^{82}Br		Fission Product	554.3	35.3h
^{208}Tl		Structural Material	583.2	
^{214}Bi		Radon	609.3	
^{82}Br		Fission Product	619.0	35.3h
^{137}Cs		Fission Product	661.6	
^{187}W	$^{186}\text{W}(n,\gamma)^{187}\text{W}$	Unknown	685.8	23.9h
^{82}Br		Fission Product	776.5	35.3h
^{27}Mg	$^{27}\text{Al}(n,p)^{27}\text{Mg}$	Fuel Cladding, Structural Material	843.8	9.5m
^{27}Mg	$^{27}\text{Al}(n,p)^{27}\text{Mg}$	Fuel Cladding, Structural Material	1014.5	9.5m
^{60}Co	$^{59}\text{Co}(n,\gamma)^{60}\text{Co}$	Stainless Steel	1173.2	
^{41}Ar	$^{40}\text{Ar}(n,\gamma)^{41}\text{Ar}$	Air	1293.6	1.8h
^{60}Co	$^{59}\text{Co}(n,\gamma)^{60}\text{Co}$	Stainless Steel	1333.2	
^{24}Na	$^{27}\text{Al}(n,\alpha)^{24}\text{Na}$	Fuel Cladding, Structural Material	1368.6	15.0h
^{40}K		Structural Material	1460.9	
^{214}Bi		Radon	1764.5	
^2H	$^1\text{H}(n,\gamma)^2\text{H}$	Water, HDPE	2223.2	
^{55}Fe	$^{54}\text{Fe}(n,\gamma)^{55}\text{Fe}$	Steel	2469.9	fs
^{208}Tl		Structural Material	2614.5	
^{24}Na	$^{27}\text{Al}(n,\alpha)^{24}\text{Na}$	Fuel Cladding, Structural Material	2754.0	15.0h
unkn.			5297.0	
^{55}Fe	$^{54}\text{Fe}(n,\gamma)^{55}\text{Fe}$	Steel	5507.5	fs
^{16}N	$^{16}\text{O}(n,p)^{16}\text{N}$	Water	6128.6	7.2s
^{57}Fe	$^{56}\text{Fe}(n,\gamma)^{57}\text{Fe}$	Steel	6318.8	fs
^{16}N	$^{16}\text{O}(n,p)^{16}\text{N}$	Water	7115.2	7.2s
^{57}Fe	$^{56}\text{Fe}(n,\gamma)^{57}\text{Fe}$	Steel	7631.1	fs
^{57}Fe	$^{56}\text{Fe}(n,\gamma)^{57}\text{Fe}$	Steel	7645.5	fs
^{28}Al	$^{27}\text{Al}(n,\gamma)^{28}\text{Al}$	Fuel Cladding, Structural Material	7724.0	2.2m
^{16}N	$^{16}\text{O}(n,p)^{16}\text{N}$	Water	8869.0	7.2s
^{55}Fe	$^{54}\text{Fe}(n,\gamma)^{55}\text{Fe}$	Steel	8886.4	fs
^{55}Fe	$^{54}\text{Fe}(n,\gamma)^{55}\text{Fe}$	Steel	9297.8	fs

Table 2: Radionuclides identified as contributing to the γ -ray background at NBSR and ATR. Listed are likely production reactions, source materials, γ -ray energy, and half-life (for short-lived reactor-correlated products). Many of these γ -rays should be expected at similar facilities, though relative line strengths could vary considerably.

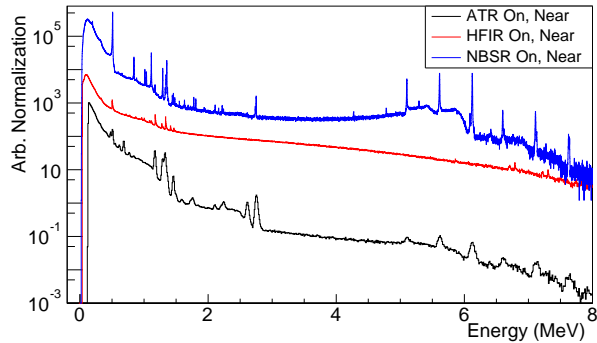


Figure 10: Comparison of the high resolution γ -ray spectra measured at the near locations using different HPGe detectors at HFIR and NBSR and a LaBr_3 detector at ATR. Spectra are collected with the reactors operating at nominal power. Note that spectra are offset in normalization for clarity and the detectors have different response functions, therefore this comparison only illustrates the general features of the γ -ray backgrounds in these locations.

4.1. High Resolution γ -ray Spectroscopy Results

High energy resolution γ -ray spectra were acquired at most of the locations of interest. At HFIR measurements were not possible in the outdoor far location. Data taken at the HFIR near location with the reactor off are thought to provide a reasonable representation of what would be encountered there. Similarly, no high energy resolution data was taken at either far site at NBSR. Again, reactor-off data taken inside the confinement building should provide a reasonable approximation of likely backgrounds.

Typical raw HPGe spectra acquired at the NBSR near location with both the reactor-on and the reactor-off are shown in Fig. 9. Clearly visible in the reactor-on spectrum are the prompt γ -ray lines associated with $^{16}\text{O}(n,p)^{16}\text{N}$ (6128.6 keV) and $^{56}\text{Fe}(n,\gamma)^{57}\text{Fe}$ (7631.1 keV and 7645.5 keV) as well as the associated single and double 511 keV escape peaks and Compton edges. Much of the apparent continuum is a function of detector response and thus a full understanding of the source spectrum would require a full deconvolution. However, it is evident that the high energy resolution of the HPGe instruments allows for identification of prominent lines contributing to the spectrum. It is also clear that there is a substantial increase in γ -ray background when the reactor is operating, particularly at higher energies. Many short-lived isotopes, some with high energies in the range 6–9 MeV are produced by a variety of mechanisms. We have identified the isotopes that make the largest contributions to γ -ray background at NBSR and ATR in Table 2 [34, 35].

Considering Table 2 more fully, the short half-life of many of the observed isotopes is notable. Those with half-lives measured in seconds or less are likely produced by neutron interactions in the immediate vicinity of the measurement location. For example, neutron leakage fields can interact with Fe in structural steel components giving rise to the observed ^{55}Fe and ^{57}Fe γ -ray lines. Similarly, neutron interactions with water or HDPE shielding give rise to ^{16}N and ^2H emissions. Isotopes with half-lives measured in minutes-to-hours can also be produced in this way, but can additionally be produced in shielded regions with high neutron flux, e.g. in primary or secondary cooling loops, and then transported in plant piping to the measurement locations. One prominent example is ^{24}Na which can be produced from trace amounts of dissolved ^{27}Al in cooling water when it is exposed to the large neutron flux at or near

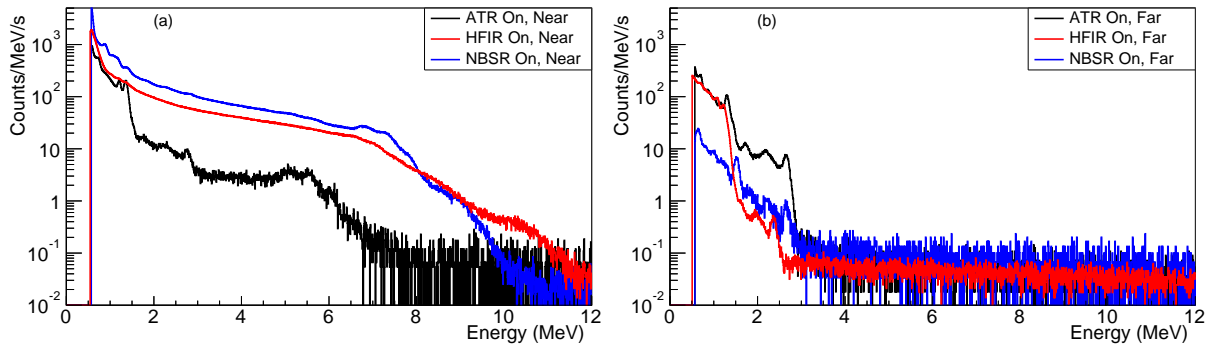


Figure 11: Comparison of NaI(Tl) spectra acquired at each reactor site, in both the (a) near and (b) far locations. Spectra are collected with the reactors operating at nominal power. The near location data are averaged over several measurement positions representing the extent of the available deployment footprint to account for the effects of positional variations.

330 the reactor core. Note that all of these isotopes will have decayed substantially within ≈ 1 day of reactor shutdown.

331 The high resolution γ -ray spectra acquired at the three near locations are compared in Fig. 10. Note that since
 332 each detector has different efficiency, it is not possible to draw conclusions about the relative intensity of the γ -ray
 333 flux at each location from this comparison. Instead, we can discern important differences relating to the background
 334 sources at the sites. The most obvious difference is the relatively featureless spectrum observed at HFIR. We interpret
 335 this to mean that the primary source of γ -ray background at the HFIR near location is γ -rays that are down-scattered
 336 as they propagate from intense radiation sources through shielding material. Given that the broad continuum observed
 337 extends to high energies, neutron interactions on steel and water are the likely source for the majority of these γ -rays.
 338 As will be discussed in Sec. 5, the emission of the down scattered γ -ray continuum is strongly correlated with the wall
 339 closest to the reactor, and there is evidence of a neutron capture γ -ray source within the HFIR near location as well.

340 The γ -ray background at both ATR and NBSR shows clear line structure implying that a significant fraction of
 341 that background is due to decays that occur locally with little intervening shielding. The high-energy features can be
 342 attributed to neutron interactions on steel and water, shedding light on local neutron backgrounds. As demonstrated by
 343 the spectral unfolding process to be described in Sec. 4.3, there is also a down-scattered continuum in these locations
 344 which can be attributed to locally produced γ -rays or the sort of incomplete shielding observed at HFIR.

345 4.2. Moderate Resolution γ -ray Spectroscopy Results

346 As described previously, a NaI(Tl) inorganic crystal spectrometer was used to obtain data for a relative comparison
 347 of the reactor sites. This device was calibrated using ^{60}Co sources. Spectra representative of the near and far detector
 348 locations at each site are compared in Fig. 11. Significant differences exist between the near and far locations, and
 349 among the sites themselves. The spectral shapes observed at the near locations are consistent with those observed in
 350 the high resolution measurements (Fig 10). Absolute γ -ray fluxes estimated using the unfolding procedure described
 351 in Sec. 4.3 to account for the detector response are given in Table 3.

352 These lower resolution spectra display features similar to that observed in Fig. 10. There is a considerable high-
 353 energy γ -ray background (> 3 MeV) at every near location. This can attributed to short-lived isotopes produced by

Location	Flux 1–3 MeV ($\text{cm}^{-2}\text{s}^{-1}$)	Flux 3–10 MeV ($\text{cm}^{-2}\text{s}^{-1}$)
ATR near	3.7	0.3
HFIR near	5.4	4.3
NBSR near	11.7	7.7
ATR far	1.7	-
HFIR far	1.7	-
NBSR far	0.1	-

Table 3: Approximate γ -ray fluxes measured with the 2" NaI(Tl) detector at the three reactor sites. While statistical errors on these values range between 0.1–1%, a conservative 10% relative systematic is assumed for the unfolding procedure used since an absolute efficiency calibration was not performed. No values for the far sites are reported in the upper energy range since there is no significant γ -ray contribution to the spectra in these cases. Note that the NBSR far site represents a typical laboratory background spectrum dominated by naturally occurring radioactivity.

354 neutron reactions either at or nearby the near locations. Contrasting the potential near locations we observe variation
 355 in the total γ -ray flux and its character. The counts recorded above 7 MeV at NBSR and HFIR imply a larger thermal
 356 neutron background, leading to neutron capture on structural steel. The feature at ≈ 2.7 MeV observed at the ATR
 357 far location is due to ^{24}Na produced near the reactor and transported in piping. At all sites, there is an indication
 358 of a continuum background due to scattered γ -rays leaking through shielding walls, or local γ -ray scattering from
 359 surrounding material. Based upon the relative lack of peak structure, it appears that HFIR has a more significant
 360 down-scattered component.

4.3. γ -ray Spectrum Unfolding

The measured γ -ray spectra are strongly dependent upon instrument response, as evidenced by the prominent escape peaks and Compton edge features visible in Fig. 9. To obtain an accurate representation of the absolute γ -ray flux for use in shielding studies, we must account for both the structure the response imprints upon the measured spectra and the energy-dependent detection efficiency of the γ -ray instruments. There exists a rich literature describing statistical unfolding or inversion techniques for problems such as this. Many difficulties can arise in applying unfolding algorithms, producing error estimates for unfolded quantities, and in selecting appropriate regularization parameters and/or convergence techniques. Furthermore, there is typically no guarantee that the solution obtained is unique. Nonetheless, this is still a useful exercise for our purpose: obtaining a reasonable estimate of the γ -ray source term for propagation through simulations of proposed detector shielding configurations. The “reasonableness” of an unfolded solution can be readily assessed by convolving it with the detector response function and making a qualitative comparison with the measured spectrum. Here we describe the method used to unfold the various γ -ray measurements, using the near location measurement at ATR with the $\text{LaBr}_3(\text{Ce})$ detector as an example. Since this detector was only used at ATR, and the measured γ -ray flux at ATR was the lowest of the three sites, this unfolding procedure was also performed for the $\text{NaI}(\text{Tl})$ measurements at all locations to estimate absolute fluxes for inter-comparison (Table 3). While pileup has been neglected in what follows due to the relatively small γ -ray detectors used, it will clearly be an important consideration for cubic meter scale $\bar{\nu}_e$ detectors.

The following data processing steps were taken prior to performing the spectral unfolding. For the $\text{LaBr}_3(\text{Ce})$ detector, background due to internal La and ^{40}K radioactivity was subtracted using a background run taken in a shielded enclosure. All measured spectra were calibrated using known line positions in the 479–7645.5 keV energy range. This calibration also provided a measurement of the detector resolution as a function of energy.

Detector response functions were generated using a dedicated GEANT4 simulation. Electron energy depositions in the crystal volume of the detector package (including an aluminum casing and readout PMT) were recorded. Detector resolution effects were accounted for by convolving the simulation result with the energy resolution function determined during calibration. For the $\text{LaBr}_3(\text{Ce})$ detector the simulated response was validated against measurements in an INL laboratory using ^{137}Cs and ^{60}Co sources. Furthermore, the simulation-predicted efficiency was in good agreement with tabulated values supplied by the detector manufacturer. The $\text{LaBr}_3(\text{Ce})$ response was then simulated over the energy range of interest (0–8 MeV). Simulated γ -rays were propagated towards the detector model uniformly from all directions. The response function was generated with a bin size of 20 keV for the incident γ -ray energy, and a non-linear binning matching the experimental data for the detected energy.

An example of the generated response function, convolved with the experimentally measured resolution, is shown in Fig. 12 for γ -ray lines due to ^{16}N . These and 14 other prominent line responses were also generated so that contributions from monoenergetic lines could be directly subtracted from the measured spectra. Doing so leaves only the relatively smooth down-scattered continuum to unfold which presents an easier task. The monoenergetic line response was estimated by fitting the sum of relevant line responses and a smooth background model to the data in

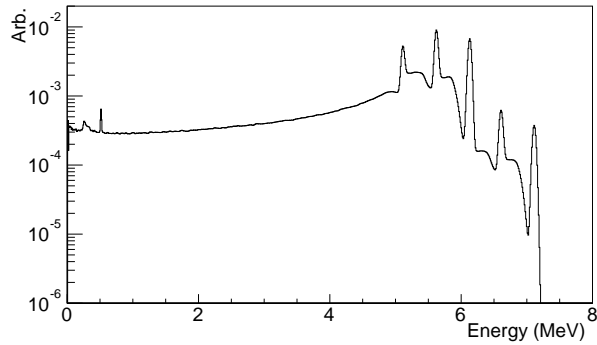


Figure 12: Simulated response of the $\text{LaBr}_3(\text{Ce})$ detector to the 6.13 and 7.12 MeV γ -rays emitted by ^{16}N . The prominent features in this response are full absorption peaks, single and double escape peaks, and the summed Compton continuum from both lines.

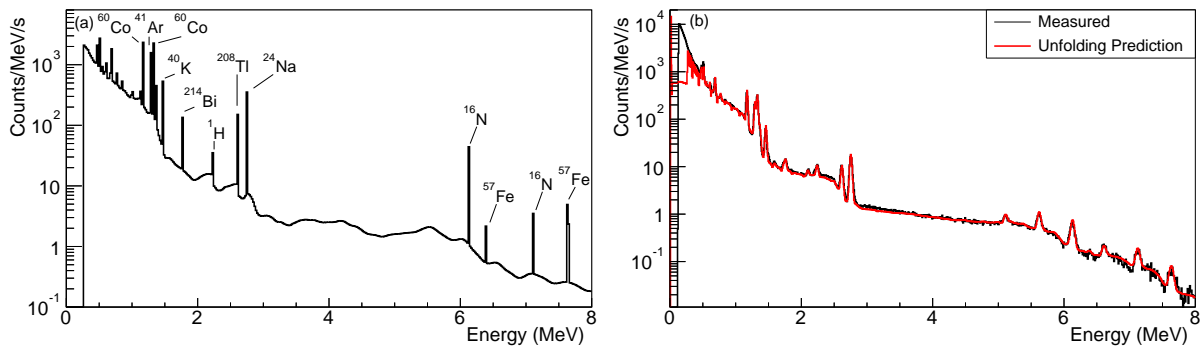


Figure 13: (a) The γ -ray spectrum incident upon the $\text{LaBr}_3(\text{Ce})$ detector at the ATR near location, as predicted by the unfolding of the measured spectrum. Prominent line sources are identified. (b) A comparison of the measured γ -ray spectrum with that predicted from the unfolded source term and the simulated detector response. Note that the unfolding procedure accounts for escape peaks and Compton scattering events in the measured spectrum. The residual continuum in (a) is due to γ -rays that have down-scattered in the surrounding environment interacting in the detector.

396 several energy ranges. The background model was generated using the Sensitive Nonlinear Iterative Peak clipping
 397 algorithm (SNIP) [36] implemented in the TSSpectrum class of the ROOT analysis package [37].

398 Finally, an unfolding algorithm [38] is applied to the residual continuum. This is done within several energy
 399 ranges where the count rate is similar, to aid convergence. The predicted source term for the measured continuum is
 400 assembled piece-wise and the monoenergetic line contributions added. The results of this procedure for the $\text{LaBr}_3(\text{Ce})$
 401 measurement taken at the ATR near location are shown in Fig. 13. The detector response predicted from the unfolded
 402 source term is in good qualitative agreement with that measured and can be readily used for detector shielding simu-
 403 lation studies.

404 4.4. Neutron Dose and Thermal Measurement Results

405 Shown in Fig. 14 are the results of neutron dose and thermal neutron flux (*italics*) measurements taken in the HFIR
 406 near location with the reactor at the nominal operating thermal power of 85 MW. The neutron dose data represent two
 407 data sets taken roughly six months apart. These measurements generally agreed at the 10% level, except where large
 408 gradients were observed. In these cases the disagreement is likely due to inaccuracies in locating the instruments.

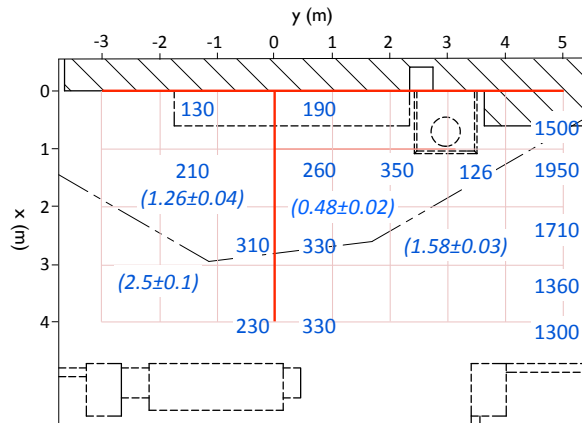


Figure 14: A pictorial representation of neutron dose rates (measured in nSv/h) and thermal neutron rates in italics ($\text{cm}^{-2}\text{s}^{-1}$) at the HFIR near location roughly 15 cm ($z = 0.15$) above the floor. Measurements are plotted on a one meter square grid referenced to the reactor wall ($x = 0$) and the smallest baseline ($y = 0$). The reactor core is centered at $(x, y, z) = (-4.06, 0, -3.85)$.

409 As is evident in the figure, considerable spatial variation was observed, in particular, a strong increase in neutron
 410 rate to both the left and right of the proposed detector location. The likely cause of this spatial variation is a large
 411 shielding structure on the level below that terminates in approximately this area (the dotted polygon shaped region in
 412 the Fig. 4). This structure is probably shielding the central area from scattered neutrons originating from the neutron
 413 beamlines on the lower level, while in other regions of this space they can propagate through the floor. The effect is
 414 particularly pronounced on the right side above the cold neutron source and guides. Consistent with this hypothesis,
 415 dose measurements taken above the cold-neutron beamline shielding, but below the experimental level floor, were
 416 $2.35 \mu\text{Sv/h}$. Similarly, the dose rate at $(x = 1 \text{ m}, y = 1 \text{ m})$ dropped by a factor of two when the HB3 beamline shutter
 417 was closed (lower floor as indicated in Fig. 4). In this scenario, localized shielding would be difficult, but may still be
 418 possible since relatively thin layers of borated materials can be very effective for thermal neutron suppression.

419 Measurements were taken at the NBSR near location and in a lab space far from the confinement building as
 420 a reference point. The near location measurements were taken multiple times, approximately 2 m from the face
 421 of the reactor biological shield indicated in Fig. 5, with adjacent instruments on. Dose rates (as described above)
 422 were $1.44 \mu\text{Sv/h}$ which, assuming a spectrum centered around energies close to the maximum detector efficiency,
 423 corresponds to an approximate flux of $2\text{--}3 \text{ cm}^{-2} \text{ s}^{-2}$. This is four times the rate observed with the adjacent instruments
 424 off. For context, the rate in the far lab space was 22 nSv/h , consistent with natural backgrounds. These rates were
 425 fairly constant between measurements. A bare BF_3 tube was used to measure the thermal flux in the same location.
 426 The flux of $2 \text{ cm}^{-2}\text{s}^{-1}$ indicates that the spectrum is likely peaked at lower near-thermal energies. Such measurements
 427 were not taken at ATR, but the relatively low flux of neutron-capture γ -rays observed at that site implies the thermal
 428 neutron flux is also low.

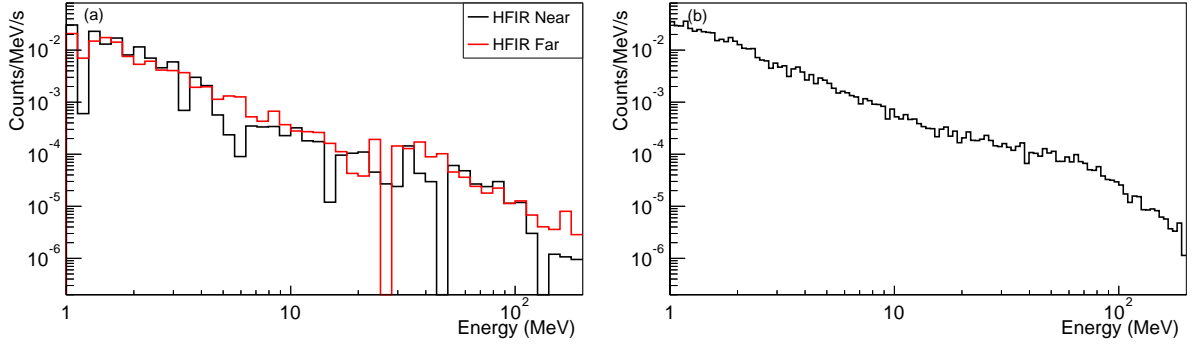


Figure 15: The cosmogenic neutron induced energy spectrum recorded at the (a) HFIR near and far locations and (b) NBSR far location.

Location	Exposure (h)	Flux ($E_n > 1$ MeV) ($\text{cm}^{-2}\text{s}^{-1}$)
HFIR near	12	$(4.1 \pm 0.3) \times 10^{-3}$
HFIR far	8	$(4.4 \pm 0.3) \times 10^{-3}$
NBSR far	156	$(5.6 \pm 0.1) \times 10^{-3}$

Table 4: Cosmogenic neutron background measurements conducted with FaNS-1 at the HFIR near location and the NBSR and HFIR far locations. Quoted uncertainties are statistical only.

4.5. Fast Neutron Measurement Results

4.5.1. Fast Neutron Spectrum Measurements with FaNS-1

Measurements of the cosmogenic neutron spectra (Fig. 15) and fluxes (Table 4) at NBSR and HFIR were performed using FaNS-1. Reactor off measurements were taken at HFIR at the near and far locations and the NBSR far location. The sensitivity of FaNS-1 to cosmogenic neutrons has been simulated using MCNPX. An isotropic distribution of neutrons following the JEDEC standard spectrum was launched at the detector and the sensitivity, in neutrons detected per incident neutron fluence, for neutron energies above 1 MeV was determined to be 10.3 ± 2.5 ($\text{n}/(\text{n}/\text{cm}^2)$) [30]. This is akin to the efficiency weighted by the cosmogenic spectrum times the cross-sectional area. This sensitivity is then used to convert a measured count rate in s^{-1} into the incident flux in $\text{cm}^{-2}\text{s}^{-1}$.

Comparing the NBSR measurement with the HFIR far location we see a slight deficit in the HFIR flux which can possibly be explained by the presence of a large (10 m–12 m) concrete wall that shadows the location. The HFIR far and near measurements are comparable. Note the similarity of the spectra shape between each site. This similarity reinforces the previously discussed notion that the spectrum of cosmogenic neutrons does not vary significantly between sites. It is important to note that these fluxes have not been corrected for fluctuations in the barometric pressure and solar cycle. These environmental parameters are known to influence the total flux by 10–20% (-0.73% per millibar change in pressure) [30, 39]. However, from a qualitative point of view, we find that the difference between NBSR and HFIR is minimal.

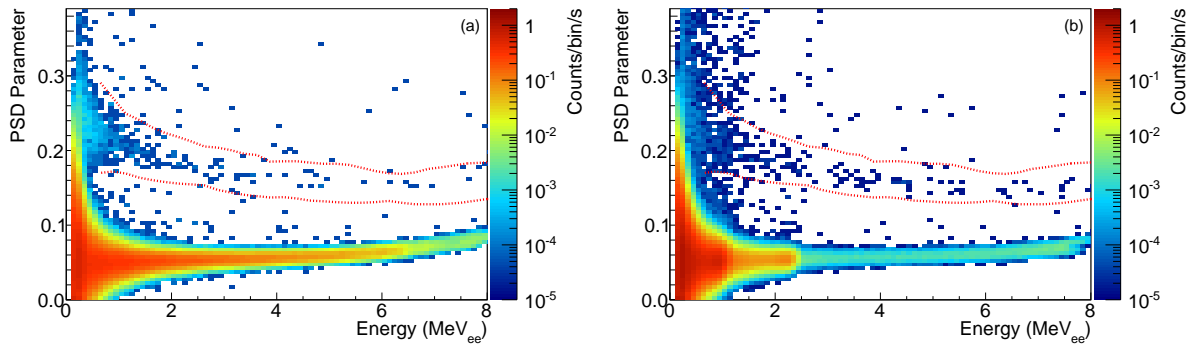


Figure 16: Stilbene PSD measurements performed at the (a) HFIR and (b) ATR far locations. Spectra are collected with the reactors operating at nominal power. The two horizontal bands correspond to γ -rays (lower) and fast neutron recoil interactions (upper). The region used for fast neutron rate measurements is indicated by dashed lines. The higher γ -ray rate encountered at ATR due to ^{24}Na causes particle misidentification at energies $< 1.5 \text{ MeV}_{ee}$.

446 4.5.2. Fast Neutron Relative Rate Measurements with a Portable Stilbene Detector

447 The stilbene detector system records list mode data for each event. A digital filtering algorithm that mimics the
 448 function of an analog constant fraction discriminator is applied to a stream of waveform samples to derive a trigger.
 449 Two integrals of waveform samples are acquired relative to the trigger time: a “full” integral summing the total PMT
 450 charge resulting from an interaction in the crystal and a “tail” integral summing charge produced primarily by the
 451 slow component of the scintillator response. Since more heavily ionizing particles, like recoil protons, preferentially
 452 excite long-lived states in the scintillator, the ratio of the “tail” to “full” integrals can be used to distinguish particle
 453 type. With this DAQ setup it was not possible to record the full waveforms corresponding to each event, only these
 454 integrals. Subsequently there is a potential for misidentification if the triggering algorithm is disrupted by pulse pileup
 455 or baseline disturbances caused by relatively high interactions rates.

456 Example fast neutron measurements using the stilbene detector are shown in Fig. 16 for the ATR and HFIR far
 457 locations. As is conventional for fast neutron PSD measurements, the electron equivalent energy (MeV_{ee}) of an event
 458 is plotted against the “tail” to “full” ratio PSD parameter. This allows the energy dependence of the PSD parameter
 459 to be readily observed. The energy scale is determined using calibration sources and background γ -rays observed at each
 460 site. Two clear horizontal bands are observed in the data, the lower corresponding to electron depositions (primarily
 461 from γ -rays) and the upper corresponding to neutron induced recoil protons. As the total event energy decreases the
 462 width of these bands increases due to worsening photo-statistics causing direct spread in the ratio and larger jitter in
 463 the trigger time determination. The PSD parameter range used for neutron counting was determined by performing a
 464 Gaussian fit to the neutron band as a function of total energy using a high statistics background dataset taken at LLNL.
 465 The selection band is set $\pm 3\sigma$ about the mean PSD parameter value found for a particular energy range.

466 Comparing the HFIR and ATR measurements we see that the larger γ -ray background at the ATR far location
 467 causes misidentification at energies $< 1.5 \text{ MeV}_{ee}$. Subsequently, this is the lower threshold implemented for all site-to-
 468 site comparisons. At some locations, notably the NBSR near location and the HFIR near location close to the reactor

Location	Rate ($\times 10^{-3} \text{s}^{-1}$)	
	4–14.5 MeV _{nr}	10–14.5 MeV _{nr}
ATR near	4.7 ± 0.3	1.0 ± 0.1
HFIR near	2.2 ± 0.2	0.3 ± 0.1
ATR far	1.8 ± 0.2	0.4 ± 0.1
HFIR far	3.5 ± 0.2	0.6 ± 0.1
NBSR far	2.8 ± 0.2	0.8 ± 0.1

Table 5: Relative fast neutron rates and associated statistical uncertainties measured at the three reactor sites. See text for additional comments.

469 wall, this misidentification was significant across almost all of the stilbene detector energy range. These measurements
470 were therefore not included in the comparison. Values reported for HFIR are the average of measurements taken at
471 the middle and rear of the potential detector location, while values for NBSR are based on data taken in a nearby
472 laboratory.

473 The conversion from electron equivalent deposition energy to recoil proton energy (denoted MeV_{nr}) for this ma-
474 terial is obtained from [40]. The 1.5 MeV_{ee} lower threshold corresponds to a proton energy of 4 MeV_{nr}, while the
475 dynamic range of the eMorpho DAQ implies an upper limit of 14.5 MeV_{nr}. As well as giving the rate in the full
476 range of comparable sensitivity, we also calculate the rate in the range 10–14.5 MeV_{nr} where there will be effectively
477 no contribution from reactor-correlated fission spectrum neutrons, allowing us to make a comparison based only on
478 cosmogenic fast neutron interactions.

479 The measured fast neutron rates for the near and far locations at all reactor sites are given in Table 5. All measure-
480 ments were performed with the reactors operating. The values for NIST were recorded in a nearby laboratory, since
481 the γ -ray background encountered in the near location was too high across much of the energy range of interest. The
482 rates given in Table 5 for NBSR are therefore a lower bound in the 4–14.5 MeV_{nr} range since any possible fission
483 neutron contribution is not included, and an upper bound in the 10–14.5 MeV_{nr} range since the attenuating affect of
484 the reactor confinement building is not included.

485 The ATR near location experiences the highest fast neutron rate, presumably due to the higher elevation of that site
486 which is not entirely offset by the overburden provided by the building structure. Conversely, the relatively deep ATR
487 far location has the lowest fast neutron rate. Comparing the HFIR near and far locations, we see that the near location
488 has a lower rate which is presumably due to the greater overburden provided by the reactor confinement building at
489 the near location relative to the (effectively) outdoor far location. The NBSR result is consistent with that at the HFIR
490 far location, which has similar elevation and overburden.

491 4.6. Muon Measurement Results

492 For this study cumulative histograms were recorded for two different trigger conditions: a twofold coincidence
493 between two of the lower paddles ($\approx \pm 55^\circ \times \pm 65^\circ$ angular range) or a threefold coincidence where the topmost
494 paddle was also required ($\approx \pm 10^\circ \times \pm 15^\circ$ angular range).

495 With a trigger threshold of ≈ 1 MeV and a coincidence window of 250 ns, accidental coincidences were found
496 to be negligible. Given a 2 MeV/cm energy deposition for minimum ionizing particles, a typical signal deposition in

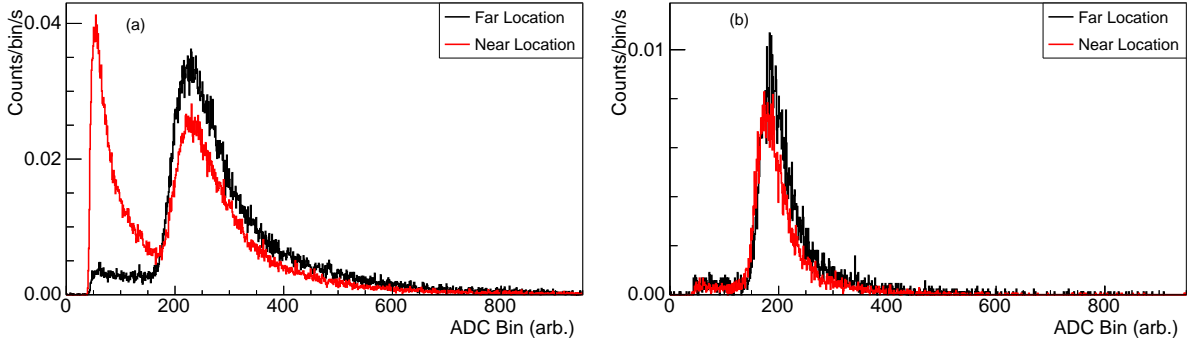


Figure 17: (a) Channel 1 spectrum requiring a twofold coincidence ($\text{Ch0}*\text{Ch1}$) at the NBSR near (red) and far (black) locations . A large correlated background from the high background γ -ray rate in the near location is apparent. (b) The same comparison for Channel 3 requiring a threefold coincidence ($\text{Ch0}*\text{Ch1}*\text{Ch3}$). No additional background is observed.

497 the paddles is ≈ 5 MeV, a value higher than most background γ -rays. In the high γ -ray fluxes encountered at some
 498 sites paddle singles rates were $< 250 \text{ s}^{-1}$, yielding a twofold accidental rate of $< 0.05 \text{ s}^{-1}$. This should be compared
 499 to the measured muon signal rate of $4\text{--}7 \text{ s}^{-1}$. The energy spectra for the twofold coincidence requirement are shown
 500 in Fig. 17a at NBSR. The minimum ionizing peak is evident in far location data. Muons clipping the edge of the
 501 scintillator paddle produce the flat shape in the bins lower than the Landau peak in simulation studies. Inside the
 502 reactor confinement building at the near location a large background was observed in the twofold coincidence spectra
 503 at low energy. This background is probably due to multiple scatter γ -ray interactions, since the accidental background
 504 previously calculated is too small to account for this feature. Requiring a threefold coincidence strongly suppresses
 505 this source of background (Fig. 17b). The coincidence spectra shape is consistent in the data taken inside and outside
 506 the NBSR confinement building, despite a factor of 200 increase in the singles rate inside the building.

507 Since the scintillator spectra requiring a threefold coincidence were consistent with clean muon signals, these
 508 measurements are used for the site comparison. Due to equipment damage during transport between the sites, the
 509 paddle separations were not identical for all measurements which had a small effect on the telescope acceptance
 510 efficiency. A geometry dependent correction factor was estimated via a simple Monte Carlo simulation. The resulting
 511 threefold coincidence rates are given in Table 6. Of the near locations, NBSR and HFIR have similar rates, while
 512 the higher rate observed at ATR is presumably due to the greater elevation at that site and the modest overburden
 513 provided by the crane access hatch at that location. The measured far location rates are very similar. At HFIR and
 514 NBSR these measurements were taken outside of the reactor confinement structures with reduced overburden relative
 515 to their respective near locations. Conversely, at ATR the far location is in a deeper basement level providing more
 516 overburden relative to the near location. For comparison, the rate measured at grade level at ATR was 0.85 s^{-1} .

517 Measurements at azimuthal angles of 0° , 45° , and 90° in different orientations relative to the reactor core were
 518 made at each of the sites. The measured rate at 90° was essentially zero. The data are consistent within errors with
 519 the expected $\cos^2\theta$ dependence. At 45° , the measured rates were lower by 10–40% when the telescope was oriented
 520 towards the more massive shielding structures surrounding the reactor cores.

Reactor	Rate at near location (s ⁻¹)	Rate at far location (s ⁻¹)
ATR	0.78 ± 0.03	0.68 ± 0.02
HFIR	0.59 ± 0.02	0.71 ± 0.03
NBSR	0.56 ± 0.01	0.69 ± 0.01

Table 6: Muon rates measured at the 3 possible near and far locations for the three-fold telescope. The far location at NBSR was a lab space whose rate should approximate the NBSR far location.

521 Translating the rates in Table 6 to an absolute muon flux requires a correction for the trigger efficiency and for
522 the solid angle acceptance. The trigger efficiency was measured with data using all four paddles stacked on top of
523 each other. An efficiency of 98.7% is assumed for all paddles. The solid angle acceptance of the coincidences was
524 calculated from a simple simulation. The average threefold acceptance with the extended paddle is 0.189 sr. The
525 twofold acceptance is 3.16 sr. Thus the fluxes at the near location obtained from the threefold measurements are
526 $79.5 \text{ m}^{-2}\text{sr}^{-1}\text{s}^{-1}$ at NBSR, $84.9 \text{ m}^{-2}\text{sr}^{-1}\text{s}^{-1}$ at HFIR, and $111.4 \text{ m}^{-2}\text{sr}^{-1}\text{s}^{-1}$ at ATR.

527 5. Characteristics of Reactor-Correlated Background

528 As noted above, there are three important sources of background encountered in research reactor facilities: natu-
529 rally occurring radioactivity in facility structures, cosmogenic background, and emissions correlated with reactor oper-
530 ations. Extensive discussion of naturally occurring and cosmogenic background can be found elsewhere (e.g. [27, 28]).
531 In this section we use the measurements described in Sec. 4 to examine the production mechanisms and other pertinent
532 characteristics of reactor-correlated backgrounds. In particular, in reference to the γ -ray lines identified in Table 2, it
533 is apparent that reactor-produced neutrons play a crucial role in elevated reactor-correlated γ -ray fluxes at the loca-
534 tions examined. Furthermore, elevated neutron rates at the locations are themselves a source of background for many
535 experiments. In this section we describe how the physical characteristics of a reactor facility influence the observed
536 background and make a qualitative comparison of the three facilities examined here.

537 The observation of significant spatial and/or temporal variations in reactor-correlated background rates at each
538 near location further illustrates the mechanisms at work. Here we give several indicative examples of the spatial
539 variation encountered at each site. These spatial variation studies were somewhat ad-hoc, being dependent upon the
540 particular configurations of detectors and shielding materials available at each site at the time of the measurements.
541 We qualitatively associate these variations with the following characteristics of reactor facilities:

- 542 • **Local concentrations of water, polyethylene, or iron.** In locations with thermal neutron leakage from the core
543 or beam lines, neutron interaction with these materials will produce prompt high-energy γ -rays;
- 544 • **Plant piping carrying water that has been exposed to high neutron fluxes.** Activated ^{16}O or trace impurities
545 in water can be transported outside of shielding walls. We do not believe significant ^{16}N was observed via this
546 pathway due to the relatively low flow in visible pipes and the short 7.3 s half-life involved. This mechanism

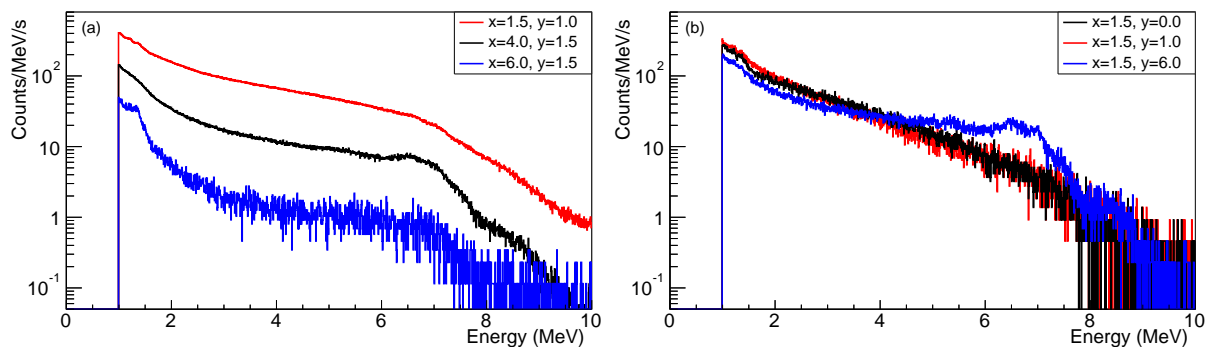


Figure 18: The γ -ray background at the HFIR near location has significant spatial variations when the reactor is operating at nominal power. These include variation with (a) distance from the wall closest to the reactor and (b) position along that wall.

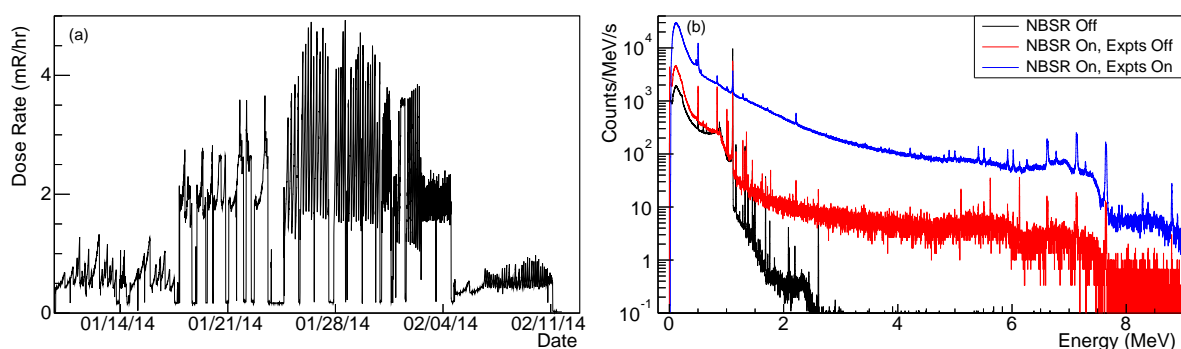


Figure 19: (a) The radiation dose rate due to background γ -rays as a function of time at the NBSR near location. (b) γ -ray spectra measured with the NIST reactor-off, with the reactor-on, and with the reactor-on and an adjacent neutron scattering experiment operating.

547 can cause activity to be transported a considerable distance from the reactor, depending upon the details of the
 548 site configuration;

- 549 • **Shielding walls or penetrations between the reactor and the measurement location.** The shielding between
 550 a location of interest and a high intensity background source (e.g. pipe carrying a large amount of primary
 551 coolant, or indeed the reactor core itself) may not attenuate the emitted γ -ray flux to levels comparable with
 552 natural background. Seams or piping penetrations in shielding walls may allow a scattering path for γ -rays that
 553 results in a localized “hot-spot”;
- 554 • **Experiments or other devices attached to neutron beamlines.** In facilities that support neutron scattering
 555 experiments, beamlines or experiments themselves can be significant sources of scattered neutrons and/neutron
 556 capture γ -rays. Large time variation can be expected from such sources during reactor on periods as experiments
 557 are reconfigured.

558 While the measurements here support the mechanism described above, we note that a more detailed survey at the site
 559 selected to host an experiment would be required to fully characterize the γ -ray background fields in order to optimize
 560 a shielding configuration.

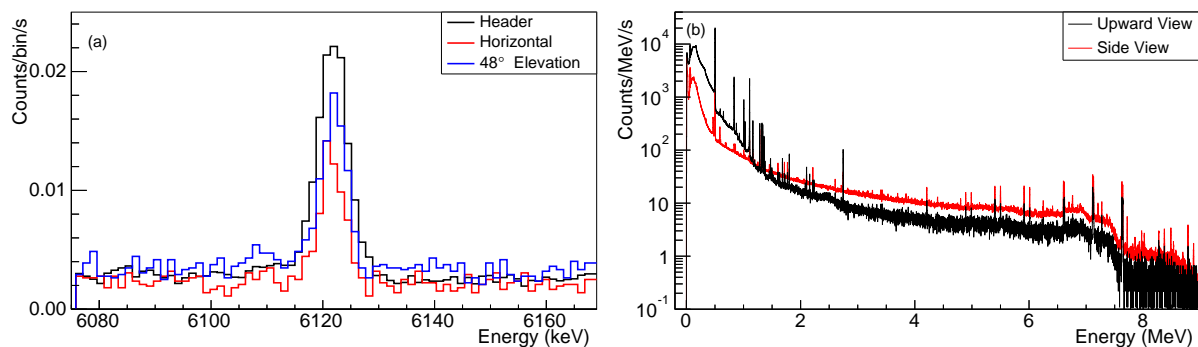


Figure 20: Examples of spatial variations observed at the NBSR. (a) The intensity of the observed ^{16}N line is correlated with header pipes carrying cooling water in the reactor thermal shield. (b) The intensity of the ^{56}Fe lines depends on the field of view with higher rates observed in the horizontal plane.

561 At HFIR considerable variation was found in the γ -ray flux with respect to proximity to the wall nearest the reactor
 562 (Fig. 4 and Fig. 18). This wall contains several penetrations which might be the source of the observed increase, or may
 563 simply not be sufficiently thick to completely attenuate emissions from activated water in the reactor pool. Collimated
 564 measurements indicate a higher flux in the direction of the wall, but not other directions. Removal of lead shielding
 565 about the NaI(Tl) detector in the vertical direction has little effect on the observed rates while removal of shielding
 566 in the direction of the wall closest to the reactor increased the rate by a factor of 4.5. Similarly, the detection rate
 567 measured with the unshielded NaI(Tl) detector exhibits a steep fall off as the distance from this wall was increased
 568 (Top panel, Fig. 18).

569 As described in Sec. 3.2.2, an increased thermal neutron flux was observed to one side of the HFIR near location.
 570 One effect of this can be observed in measurements made with the unshielded NaI(Tl) detector. At the locations with
 571 elevated thermal neutron flux, high-energy capture γ -rays are more prominent (bottom panel, Fig. 18). Examination
 572 of the continuum portion of these spectra also sheds light on the source of this γ -ray background. The similarity in
 573 the intensity and shape of this continuum at locations along the wall closest to the reactor suggest that the source is
 574 not localized to a single penetration or narrow leakage path. Instead, it appears likely that the entire length of the wall
 575 is emitting downscattered γ -rays from the reactor pool.

576 At NBSR a number of spatial and temporal variations were observed. The variation in γ -ray background due to the
 577 operation of an adjacent neutron scattering instrument (MACS) is displayed in Fig. 19a. The top panel shows results
 578 from a logging dosimeter [41] that was used to record the γ -ray dose rate in the near location over a 1 month period.
 579 During the period spanning 1/18/14–1/31/14 the instrument was operated with cadmium thermal neutron shielding
 580 which increased the background rate (use of boron thermal neutron shielding should result in a substantial decrease
 581 in γ -ray dose). The detailed time structure seen in Fig. 19a is the result of different configurations of the neutron
 582 scattering instrument. The average dose during the period shown was $13 \mu\text{Sv/h}$. The effect of the operation of the
 583 adjacent instruments on the γ -ray energy spectrum is also displayed in Fig. 19b. Large increases in downscattered
 584 continuum background and ^{57}Fe emissions due to thermal neutron capture on structural steel are observed.

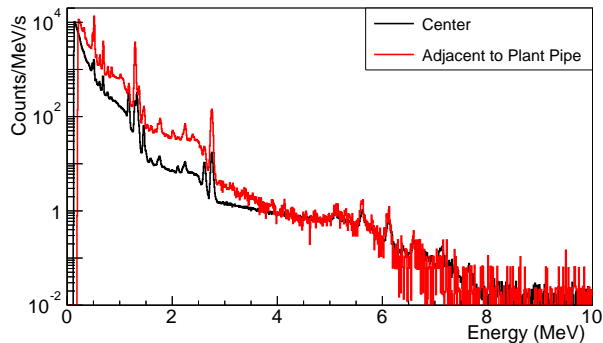


Figure 21: The γ -ray flux due to ^{24}Na (≈ 1.3 and ≈ 2.7 MeV) increases substantially when the $\text{LaBr}_3(\text{Ce})$ detector is moved from the center of the ATR near location to be adjacent to plant piping in the area and the reactor is operating at nominal power.

585 Significant spatial variation of the γ -ray background was also observed at NBSR. A primary source of background
 586 was identified as coming from the thermal shield cooling-water lines located above the proposed near location as
 587 indicated in Fig. 6a. The dominant 6.128 MeV line is clearly seen in Fig. 9. Measurements taken with 5 cm thick lead
 588 apertures that restricted the detector field of view to approximately 30° demonstrated qualitatively that these lines are
 589 originating in the header assembly. This is shown in Fig. 20a. As can be seen in Fig. 5, this source of background
 590 is partially shielded by the biological shield and illuminates roughly half of the potential near location. Fig. 20b
 591 compares data taken with two different apertures at a position roughly 50 cm from the face of the reactor biological
 592 shielding: a 2π upward view and an arrangement that views primarily the horizontal plane. The spatial dependence
 593 and lack of significant downscattering in these data suggest that the dominant source of higher-energy γ -rays are
 594 thermal neutron capture on the steel shielding surrounding the adjacent beamlines, consistent with the interpretation
 595 of Fig. 19b. They also indicate that the low energy part of the spectrum is dominated by overhead sources. The fact
 596 that the γ -ray backgrounds at NIST are highly directional, and in some cases, well localized, suggests that targeted
 597 shielding may be particularly effective.

598 At ATR, both the near and far locations have line-of-sight to piping carrying small amounts of water that has been
 599 in close proximity to the core. The length of these pipes and the relatively low flow rates they carry result in there
 600 being little ^{16}N activity observed from them. However, as demonstrated in Fig. 21, proximity to these pipes results
 601 in a substantial increase in the observed ^{24}Na γ -ray flux, most notably the line at 2.754 MeV. At the near location a
 602 piping manifold, used for monitoring water flows near control devices in the core, on the wall closest to the reactor
 603 is therefore the likely source of the observed ^{24}Na activity. At the far location, a small ceiling mounted pipe carrying
 604 primary coolant diverted to a power monitoring system is the ^{24}Na source.

605 6. Case Study: the PROSPECT Experiment at the High Flux Isotope Reactor

606 After an assessment process that considered the background characteristics described here in addition to logistical
 607 and engineering considerations, the PROSPECT collaboration decided to pursue PROSPECT Phase I at HFIR [10].
 608 Therefore the background characteristics of the HFIR near location were examined in greater detail. Described here

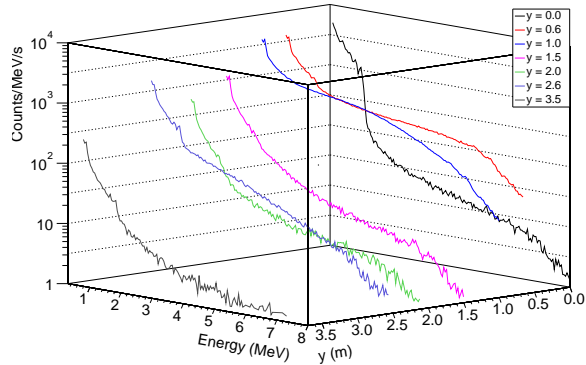


Figure 22: Measured energy spectra for an unshielded NaI(Tl) detector at different y locations along the wall surrounding the reactor pool, with the reactor operating at nominal power. All data are taken 0.1 m from the wall and 1 m above the floor ($x = 0.1$, $z = 1.0$). Several locations exhibit significantly higher rate, associated with penetrations in the wall described in the text.

609 are γ -ray surveys to more fully determine spatial variations in the background γ -ray flux, studies to develop detector
 610 shielding appropriate for the background encountered at this location, and results from a prototype detector deploy-
 611 ment to demonstrate background reduction.

612 6.1. Detailed Spatial γ -ray Surveys

613 The previously described NaI(Tl) measurements used for comparing reactor sites were unshielded measurements
 614 made within the expected near detector footprint. To better identify γ -ray background sources at HFIR, it was neces-
 615 sary to explore a wider range of positions with unshielded and shielded detectors. Over 200 NaI(Tl) measurements
 616 were made during three background measurement campaigns at HFIR with the reactor was operating at a thermal
 617 power of 85 MW. The grid shown in Fig. 14 provides a convenient reference for comparison of different positions.
 618 The y -axis measured distance along the wall surrounding the reactor water pool, the x -axis measured the distance
 619 from the wall and the z -axis measured the height above the floor, with $y = 0.0$ being in line with the reactor core.
 620 The complex spatial variations observed in Sec. 5 indicate that multiple sources contribute to the background at any
 621 given location. Lead shielding was used to restrict the angular acceptance of the NaI(Tl) during some measurements
 622 to indicate the spatial distribution of these background γ -ray sources.

623 The differing rate and energy spectra of the background sources along the reactor pool wall are illustrated in
 624 Fig. 22, where measured γ -ray spectra taken at different x positions are plotted. Two prominent hot spots are evident.
 625 A pipe directly through the concrete wall to the reactor water pool near $x = -0.04$ m is an intense source of lower-
 626 energy γ -rays (≤ 1.5 MeV). An unused beam tube between $y = 0.66$ – 1.0 m, pointing almost directly back to the
 627 reactor core, is the dominant source of higher-energy γ -rays (≥ 2 MeV) despite being filled with a concrete plug. Less
 628 prominent hot spots interrupt the general reduction in rate with increasing y at $y = 2.56$ m (a notch in the wall) and
 629 $y = 3.0$ m (above another unused beam tube in the floor).

630 General trends in the spatial variation of γ -ray backgrounds can be seen in Fig. 23, which displays integrated
 631 γ -ray counting rates between 1–10 MeV as a function of position. Contour plots at two different heights above the

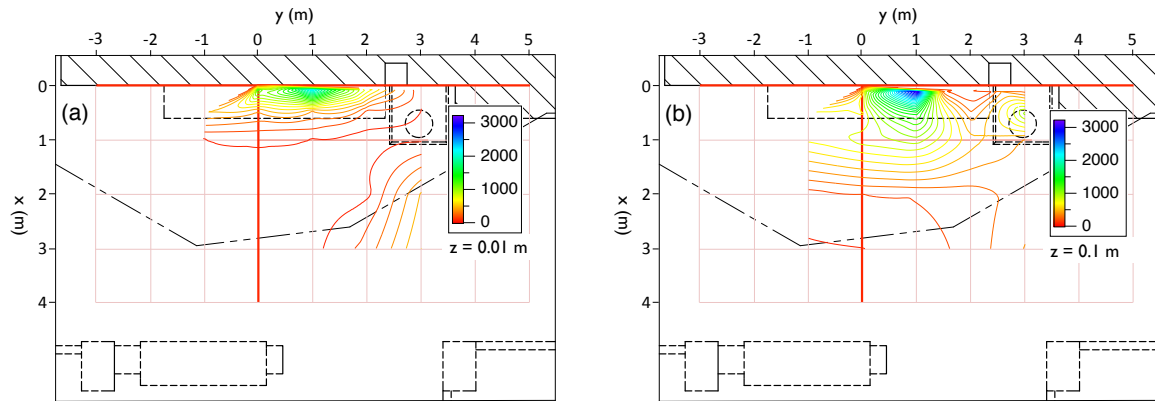


Figure 23: Measured count rates (s^{-1}) over the energy range of 1–10 MeV, for an unshielded NaI(Tl) detector either 10 cm above the floor (a) or 100 cm above the floor (b). The reactor was operating at nominal power. The reactor core is centered at $(x, y, z) = (-4.06, 0, -3.85)$.

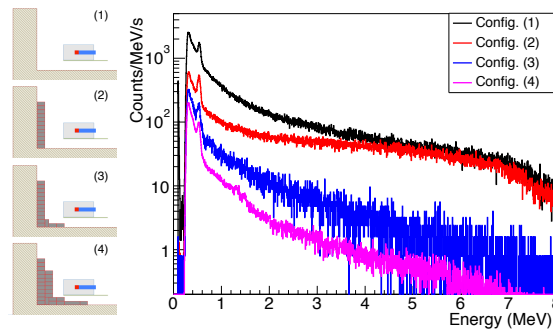


Figure 24: Measured energy spectra for a NaI(Tl) detector inside a horizontal lead collimator placed at $x = 0.5$ m, $z = 0.2$ m for different configurations of a 102 cm wide (y -axis in Fig. 23) lead wall, with the reactor operating at nominal power. This location is in front of a localized γ -ray background source.

632 floor are shown: (top) $z = 0.1$ m and (bottom) $z = 1.0$ m. Variation along the y -axis close to the wall ($x = 0.1$ m)
 633 follows the trends seen in Fig. 22. Integrated rates decrease along the y -axis as the distance from the reactor increases,
 634 consistent with the spectra shown in top of Fig. 18. The variation is most pronounced close to the floor as can be seen
 635 comparing the top and bottom of Fig. 23. This large reduction in background rate is attributed to the large concrete
 636 support monolith under this level whose outline can be seen as a dashed line in Fig. 23 or in the elevation view of
 637 Fig. 4. Backgrounds from the water pool much below the level of the floor are strongly suppressed.

638 Close to the reactor wall both the average γ -ray energy and rate are significantly lower 2 meters above the floor
 639 than at 1 meter. Rates below 1.5 MeV are a factor of 10 lower while rates ≈ 3 -6 MeV are nearly 100 times lower.
 640 However, further from the wall ($x \geq 0.7$ meter), the spectra at $z = 1$ and 2 meters are similar while rates just above the
 641 floor (monolith) are very low. These distributions imply that higher-energy γ -rays from the wall are emitted roughly
 642 at 45° to the vertical i.e. along the unused beam tube.

643 Measurements were taken with the NaI(Tl) detector inside a 10 cm thick rectangular lead well, intended to atten-
 644 uate all γ -rays not coming from directly beneath the detector. An intense local hot spot is observed near $y = -0.2$ m

Configuration (dimensions in cm)	Rate (Hz)	
	1-3 MeV	3-10 MeV
No Wall (1)	512.4	246.3
Add wall: 10 × 102 × 51 (2)	169.8	168.4
Add floor: 25 × 102 × 5 (3)	52.5	15.2
Add to floor: 25 × 102 × 10	32.2	10.5
Add to wall: 20 × 102 × 51	28.5	12.7
Extend floor: 30 × 20 × 5 (4)	15.5	3.0

Table 7: Integrated background rates for energy ranges 1–3 MeV and 3–10 MeV for sequential augmentation of a lead shielding wall. The shielding spans the range $y = 0.6$ – 1.0 m against the wall indicated in at $x = 0$ in Fig. 23. Wall dimensions are given as $x \times y \times z$ values, with (x, y, z) directions also as indicated in Fig. 23. Several configurations are pictorially represented in Fig. 24.

645 near the wall closest to the reactor. Away from the wall rates were uniformly low over the shielding monolith. Back-
646 ground rates increased with the detector over the relatively thin 15 cm concrete floor outside of the monolith footprint
647 ($x \gtrsim 2$ m). The level beneath the location being examined contains multiple neutron beam lines. Scattered beam
648 neutrons interacting with structural materials in that level or the floor itself are thought to be cause of the increased
649 γ -ray background rates observed past the shielding monolith.

650 A study of shielding effectiveness was conducted by varying the configuration of a lead wall in front of the beam
651 tube at $y = 0.6$ – 1.0 m and measuring background γ -ray rates (Fig. 24). The NaI(Tl) detector was placed between two
652 10 cm thick lead walls oriented perpendicular to the lead wall, thus limiting the detector acceptance in the horizontal
653 plane in directions other than the wall. Table 7 gives the background rates summed over the energy ranges 1–3 MeV
654 and 3–10 MeV for each wall configuration. Fig. 24 shows the background energy spectra at selected configurations.

655 With the detector 0.5 m from the wall, a 10 cm thick lead wall reduced the γ -ray detection rate at energies below
656 3 MeV by a factor of 3. Extending the wall onto the floor by 25 cm significantly reduced the rate of higher-energy
657 γ -rays by as much as a factor of ten. Doubling the thickness of the floor layer further reduced rates, while doubling
658 the thickness of the vertical wall had little effect. Extending the floor bricks another 20 cm lowered the high-energy
659 γ -ray rate by an additional factor of four.

660 Both background sources and shadows were observed during these studies. The solid concrete monolith effectively
661 blocks any background sources directly beneath the location under consideration. Penetrations or relatively thin
662 sections in concrete structures were associated with higher backgrounds. In particular the beam tube near $y = 1.0$ m
663 was the dominant source of high-energy background. Less intense sources of higher-energy γ -rays were likely to be
664 associated with higher neutron fluxes at large y ($y \gtrsim 2$ m) or off the monolith ($x \gtrsim 3$ m). Accordingly, PROSPECT
665 aims to build a localized lead shielding structure against the wall and floor closest to the reactor and then remeasure
666 these background distributions before designing detector shielding.

667 6.2. Deployment of the PROSPECT2 Prototype at the HFIR Near Location

668 To test the efficiency of shielding and provide data for simulation validation, a prototype detector was deployed
669 at the HFIR near location. The detector is a right cylindrical acrylic vessel with an internal diameter of 12.7 cm

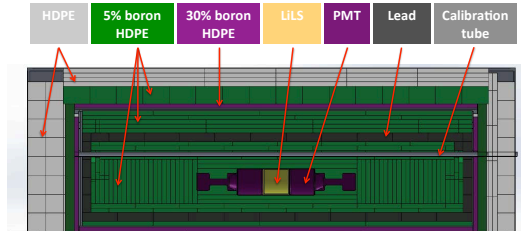


Figure 25: PROSPECT2 as installed at HFIR. The 5" cylindrical LS detector (yellow), PMTs and HV bases (purple) are surrounded by 5% borated polyethylene sheets (green), lead (dark grey), more 5% borated polyethylene sheet, an Al containment box (grey), 30% borated polyethylene sheet (purple), more 5% borated polyethylene sheet, and polyethylene sheet (light grey).

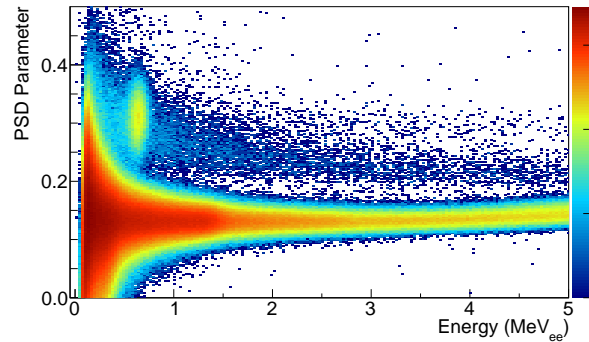


Figure 26: PSD parameter vs. electron equivalent energy for the shielded PROSPECT2 detector operated at HFIR with the reactor off.

670 containing 1.7 liters of organic liquid scintillator doped with 0.1% by weight ^6Li (LiLS). Since the active volume
 671 is almost 2 liters, the device is denoted as PROSPECT2; later prototypes of larger size follow a similar naming
 672 convention. Optical readout was via two 5 inch PMTs (ET9823KB [42]) coupled directly to each face of the vessel
 673 with EJ550 optical grease [43]. All other sides are covered with a diffuse reflective TiO_2 paint. Each PMT is readout
 674 using a CAEN V1720 waveform digitizer [44] sampling at 250 MHz with 12 bits per sample.

675 The PROSPECT2 detector was deployed within a multilayer shield enclosure designed to reduce both γ -ray and
 676 neutron fluxes. A diagram of the shielding configuration is shown in Fig. 25. The ≥ 50 cm thick shield consists of
 677 (from the outside in) 10-20 cm of high-density polyethylene, ≈ 20 cm of 5% borated polyethylene, 2.5 cm of 30%
 678 borated polyethylene, 5-10 cm of lead, and finally 10 cm more of borated polyethylene. In addition, a 10-cm-thick
 679 lead shield was placed over the beam port describe in Sec. 6.1 to locally shield that intense background source.

680 The guiding concepts behind this design are:

- 681 • Thermalize and capture low-energy neutrons in an outer layer of borated polyethylene to reduce high-energy
 682 capture γ -rays;

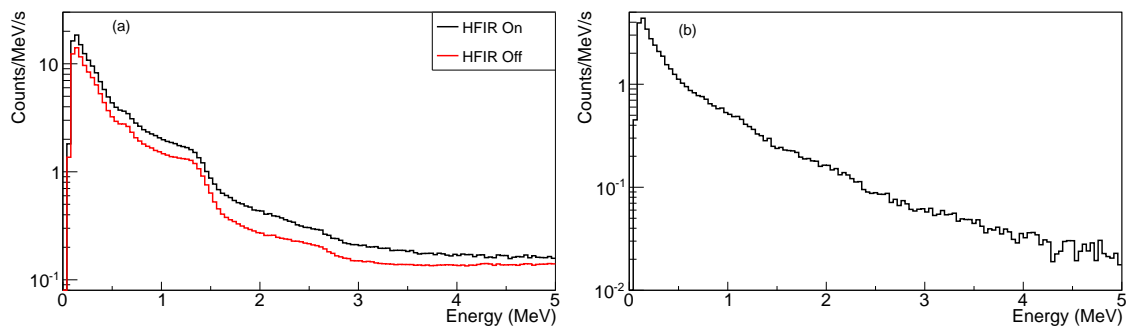


Figure 27: (a) The energy spectra of depositions in the PROSPECT2 detector are compared with the reactor on and off. (b) The residual after subtraction of the reactor-off spectrum from the reactor-on.

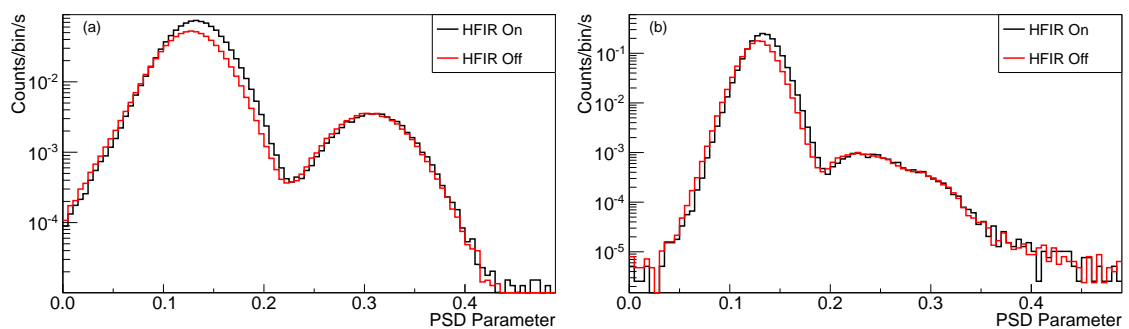


Figure 28: PSD parameter distributions are compared for reactor-on and reactor-off data for two energy ranges in the PROSPECT2 detector: (a) 0.5–0.8 MeV, corresponding to neutron capture on ${}^6\text{Li}$, and (b) 1–5 MeV. The similarity of the distributions at high PSD parameter values indicates the detection no reactor-correlated thermal or fast neutrons.

- 683 • Use a layer of high-Z material to stop external γ -rays as well as those produced from neutron capture in the
- 684 outer borated polyethylene later;
- 685 • Thermalize and capture any neutrons produced from cosmic rays interactions in the high-Z material in a second
- 686 layer of borated polyethylene.

687 The combination of ${}^6\text{Li}$ doping and PSD in the PROSPECT2 detector allows the same device to simultaneously

688 measure γ -ray, fast neutron recoil, and neutron capture rates. As with the stilbene detector described in Sec. 4.5.2

689 a PSD parameter is determined by taking the “tail” to “full” ratio of the PMT pulse. As demonstrated in Fig. 26,

690 interactions of each of these particle types fall in a different region of an PSD parameter vs. energy plot. We use

691 this capability to assess the effectiveness of the shield enclosure at reducing reactor generated backgrounds from each

692 of these particle types. Data sets totaling 109 hours with the reactor operational at a thermal power of 85 MW and

693 348 hrs with the reactor off were collected.

694 The electron-equivalent energy spectrum of all depositions in the PROSPECT2 detector is shown in Fig. 27 for

695 reactor-on and reactor-off data. Clearly, there is an increase in the detector interaction rate that can be attributed to

696 reactor generated particles. However, a considerable reduction in background is achieved compared to what would be

697 expected with no shielding. Using the reactor-correlated γ -ray fluxes given in Sec. 4.1, the unshielded PROSPECT2

Parameter space region	Background rate ($\times 10^3 \text{s}^{-1}$)	
	Reactor-on	Reactor-off
0.5–0.8 MeV, γ -like	960.1 \pm 1.5	700.6 \pm 0.7
0.5–0.8 MeV, n-like	58.0 \pm 0.4	58.6 \pm 0.2
1.0–5.0 MeV, γ -like	1719.5 \pm 2.1	1261.2 \pm 1.0
1.0–5.0 MeV, n-like	15.2 \pm 0.2	15.5 \pm 0.1

Table 8: Integrated rates for γ -like and neutron-like events in the PROSPECT2 detector for reactor-on and reactor-off conditions. The 1–5 MeV energy range approximately corresponds to inverse beta decay positrons, while the 0.5–0.8 MeV energy range corresponds to neutron capture on ${}^6\text{Li}$. Quoted uncertainties are statistical only.

698 detector would be estimated to have reactor related excess count rates of $\approx 4 \times 10^3 \text{s}^{-1}$ and $\approx 3 \times 10^3 \text{s}^{-1}$ in the
699 1–3 MeV and 3–10 MeV energy ranges, respectively. Instead rates of 1.38s^{-1} and 1.44s^{-1} are observed in these
700 respective ranges using the shielded detector.

701 The increased background occurs at energies $\lesssim 7$ MeV, with the greatest enhancement occurring at low energies.
702 We can use the PSD ability of the detector to infer the relative contribution of this increase background from different
703 particle types. PSD spectra are displayed in Fig. 28 for two energy ranges: 0.5–0.8 MeV, corresponding to the ${}^6\text{Li}$
704 neutron capture feature, and 1.0–5.0 MeV, corresponding to the region where positrons from the inverse beta-decay $\bar{\nu}_e$
705 interaction would be observed. In these projections, the features at lower values of the PSD parameter correspond to
706 electromagnetic interactions (predominately Compton scattering of γ -rays), while those at higher values correspond
707 to thermal neutron capture (Fig. 28a) or fast neutron recoils (Fig. 28b).

708 It is apparent from these figures and the integrals of the two PSD regions given in Table 8 that the increased
709 background is due to γ -ray interactions in the PROSPECT2 detector. That is, the shielding surrounding the detector
710 effectively eliminates any reactor generated thermal or fast neutrons. The reactor-correlated background observed
711 can be attributed to γ -rays produced outside and transported through the shield, or to γ -rays produced by neutron
712 capture interactions within the shield. Since care was taken not to include materials that produce high-energy γ -rays
713 within the shield, the high-energy excess observed in the PROSPECT2 detector is attributed to external production
714 and transport. While the γ -ray interaction rate changes with the reactor status, the rate of fast neutrons and neutron
715 captures in the PROSPECT2 detector are unchanged. This is an indication that any reactor-correlated neutron flux is
716 highly suppressed by the shielding package.

717 7. Conclusion

718 The background characteristics of three research reactor facilities have been measured. Both significant similari-
719 ties as well as important differences between the sites were encountered, and thus it is expected that these measure-
720 ments will inform work at research reactor sites generally. Features common to all sites, include:

- 721 • significant spatial variations in γ -ray and neutron backgrounds due to irregular shielding, localized leakage paths
722 through shielding, or the presence of piping carrying activated materials. Detailed site-specific characterization

723 of background is therefore essential to optimize a shielding design. In some cases, localized shielding applied
724 to compact background sources could be a cost and weight efficient approach to reducing detector backgrounds;

725 • higher reactor-correlated background rates are encountered at potential near detector locations, when compared
726 to far detector locations. This is not surprising, considering the near locations are closer to the reactors and
727 therefore have less shielding from that intense source, and/or are more likely to be proximate to plant systems
728 or other experiments that can transport radiation from the reactor to a detector location. A far detector may
729 therefore require less shielding than a near detector;

730 • neutron leakage and/or scattering is a significant background source, via neutron interactions on water, steel,
731 or other structural materials. The resulting high-energy γ -rays are relatively difficult to shield. Application of
732 relatively light neutron absorbing shielding to localized neutron sources could therefore be a cost and weight
733 efficient approach to reducing γ -ray backgrounds.

734 Features particular to ATR include:

735 • the lowest near site γ -ray background, due to relatively low thermal neutron leakage and good shielding from
736 the reactor, and few nearby plant systems. This is offset by the highest cosmogenic background flux (muon and
737 fast neutron), due to the high site elevation;

738 • the lowest cosmogenic background flux (muon and fast neutron) of any location at the potential far detector
739 location. This is due to the location being ≈ 12 m below grade in a basement. The far location γ -ray background
740 is the highest of any far site, but still significantly lower than the near locations;

741 • no expected or observed time variation of reactor-correlated γ -ray or neutron backgrounds.

742 Features particular to HFIR include:

743 • a large down-scattered γ -ray background coming from the entire length of the wall closet to the reactor at
744 the near location. This implies that an intense radiation source (likely the reactor pool) is being only partially
745 shielded. However, the flux falls rapidly as the distance to this wall increases, suggesting that localized shielding
746 applied along the length of the wall may be able to attenuate this flux in a cost and weight effective manner.

747 Features particular to NBSR include:

748 • both large spatial and temporal variations of γ -ray and thermal neutron backgrounds at the near location. This
749 is due to both the facility design and the operation of nearby experiments. Localized shielding may therefore
750 be able to attenuate these sources in a cost and weight effective manner. The γ -ray background encountered at
751 NBSR is similar to that at HFIR.

752 While the background surveys reported here should be useful in the preliminary design of an experiment, given
753 the considerable variation in background sources and intensity observed, a primary conclusion of this paper is that
754 any sensitive experiment intending to operate in such facilities must perform detailed assessment of the background in
755 the particular location of interest. Such detailed measurements conducted by the PROSPECT collaboration at HFIR
756 have illustrated the complex nature of the background fields in that facility as well as the ability to strongly suppress
757 backgrounds with well placed shielding. Deployment of the PROSPECT2 detector in a shielding enclosure verified
758 this conclusion, and importantly indicated that reactor-correlated neutron backgrounds can be essentially completely
759 suppressed.

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