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1 Progress on Perovskite Materials and Solar Cells with Mixed Cations and Halide Anions

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6 **ABSTRACT:** Organic–inorganic halide perovskite materials (e.g., MAPbI₃, FAPbI₃, etc.; where MA = CH₃NH₃⁺, FA =
7 CH(NH₂)₂⁺) have been studied intensively for photovoltaic applications. Major concerns for the commercialization of perovskite
8 photovoltaic technology to take off include lead toxicity, long-term stability, hysteresis, and optimal bandgap. Therefore, there is
9 still need for further exploration of alternative candidates. Elemental composition engineering of MAPbI₃ and FAPbI₃ has been
10 proposed to address the above concerns. Among the best six certified power conversion efficiencies reported by National
11 Renewable Energy Laboratory on perovskite-based solar cells, five are based on mixed perovskites (e.g., MAPbI_{1-x}Br_x,
12 FA_{0.85}MA_{0.15}PbI_{2.55}Br_{0.45}, Cs_{0.1}FA_{0.75}MA_{0.15}PbI_{2.49}Br_{0.51}). In this paper, we review the recent progress on the synthesis and
13 fundamental aspects of mixed cation and halide perovskites correlating with device performance, long-term stability, and
14 hysteresis. In the outlook, we outline the future research directions based on the reported results as well as related topics that
15 warrant further investigation.

16 **KEYWORDS:** perovskite, solar cell, mixed cations, mixed halides, stability, hysteresis

1. INTRODUCTION

17 Organic–inorganic halide perovskite materials (hereinafter
18 denoted as “perovskites”) continue to attract worldwide
19 attention, shown by a rapid increase in the number of
20 publications per year (Figure 1a).¹ CH₃NH₃PbI₃ and
21 CH₃NH₃PbBr₃ perovskites were first incorporated as light
22 harvesters in liquid dye-sensitized solar cells in 2009 by Kojima
23 et al.² In 2012, a significant breakthrough was made with the
24 introduction of all solid-state CH₃NH₃PbI₃ perovskite solar
25 cells by Kim et al.³ Since then, perovskite solar cells have been
26 under the spotlight with a cluster of fundamental scientific
27 discoveries^{4–39} as well as breakthroughs in solar-to-electricity
28 power conversion efficiencies (PCEs).⁴⁰ The highest certified
29 efficiency of 22.1% as of today is only a few percent points shy
30 of that of the best single-crystalline silicon solar cells.^{40–43}
31 Perovskite solar cells are considered as the most promising
32 candidate for the next generation high efficiency solar cell
33 technology that is compatible with low-cost, low-temperature
34 processing, flexible substrates, and large-area (module)
35 fabrication processes.^{4,10,23,25,44–58}

Highly efficient perovskite solar cells are composed of 36
perovskites that have an ABX₃ three-dimensional (3D) 37
structure and are commonly composed of an organic/inorganic 38
monovalent cation, A = (methylammonium (MA), CH₃NH₃⁺; 39
formamidinium (FA), CH(NH₂)₂⁺; Cs⁺; Rb⁺), a divalent metal 40
cation, B = (Pb²⁺; Sn²⁺), and halide anion motif X₃ = (I⁻; Br⁻; 41
Cl⁻). Pristine methylammonium lead iodide perovskite 42
(MAPbI₃) has been intensively used as light harvesting material 43
from earlier times of perovskite solar cell develop- 44
ments.^{2,3,42,43,59} Perovskite compositions with single ions 45
occupying each of A, B, and X sites (e.g., MAPbI₃, FAPbI₃, 46
etc.) are hereinafter denoted as “simple perovskites” and has 47
been intensively studied (Figure 1a). In earlier studies, 48
perovskites with mixed halides including minute quantities of 49
Cl⁻ in the X₃ motif of MAPbI₃ (i.e., MAPbI_{3-x}Cl_x) were 50
employed,^{60,61} and currently there is still a high predominance 51
of employing high contents of iodine-based perovskite 52

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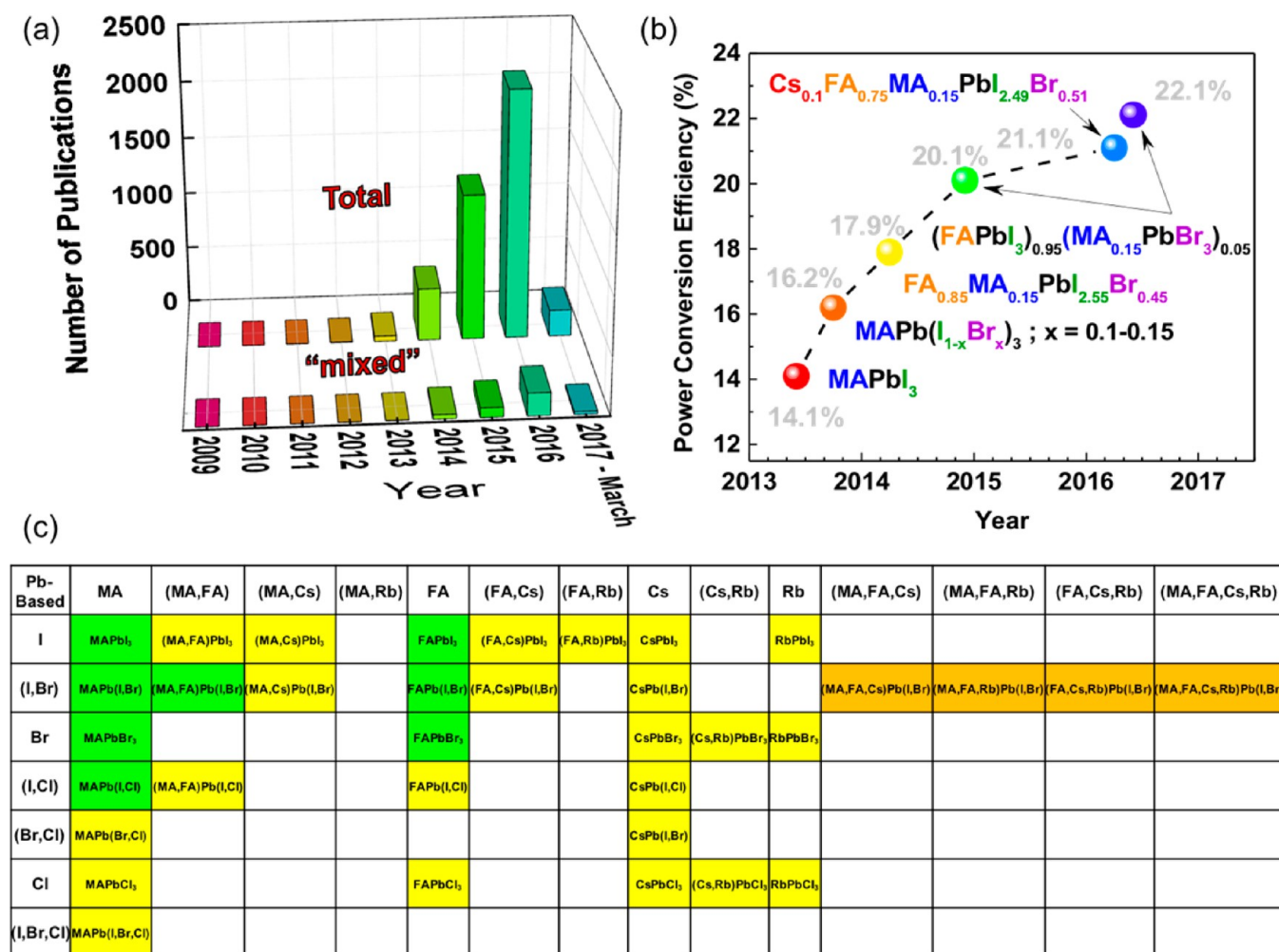


Figure 1. (a) Number of publications retrieved from Web of Science (Thomson Reuters) as a function of the year using the following search keywords: “perovskite” AND “solar” AND “cell” (AND “mixed”).¹ (b) Best research perovskite solar cells efficiencies certified by NREL.⁴⁰ For each of certified efficiencies, the chemical compositions of mixed perovskites are specified. (c) matrix of several simple and mixed cations and mixed halides Pb-based perovskites generated by the permutation of A = (MA,FA,Cs,Rb) and X = (I,Br,Cl). The perovskite compositions marked in yellow are documented in the literature. The green ones correspond to widely studied in the literature due to suitability for optoelectronic applications. The orange ones are the ones receiving great attention. Finally, the white boxes are perovskite compositions not still explored.

53 deposition recipes for photovoltaic applications.^{13,24–26} In fact,
 54 except for the first certified efficiency record reported for hybrid
 55 perovskite research cells using phase pure MAPbI₃,⁵⁹ the
 56 following subsequent five National Renewable Energy Labo-
 57 ratory (NREL) records with publicly disclosed information
 58 used A and/or X site mixed hybrid perovskites^{62–66} (herein-
 59 after denoted as “mixed perovskites”), see Figure 1b. However,
 60 based on Figure 1a, the fraction of the number of publications
 61 employing mixed perovskites with respect to the total number
 62 are still relatively low, ~9% in 2016 and ~10% as of April 15th,
 63 2017.¹ A question naturally arises whether the future trends for
 64 perovskite photovoltaic technology should be focused on mixed
 65 perovskites. In this review, we outline systematically the
 66 benefits to employ mixed cations and halides in perovskite
 67 materials. For example, we will highlight the following
 68 advantages: (i) higher performance, (ii) increased stability,
 69 (iii) enhanced carriers charge transport, (iv) enabling band gap
 70 tuning, (v) less pronounced hysteresis, (vi) a thorough
 71 understanding of fundamental aspects of Pb-based mixed
 72 cations and halides may lead to alternative Pb-free perovskites
 73 by rational designs. In this review, only the Pb-based mixed
 74 cations and halides perovskites are to be covered based on the

fact that substantial knowledge has been accumulated with 75
 several congruent conclusions for Pb-based mixed cations and 76
 halides perovskites. In Figure 1c, we marked the Pb-based pure 77
 and mixed cations and halides perovskites that have been 78
 reported in the literature. The color coding signifies: (yellow, 79
 green, orange) = reported in the literature; green = widely 80
 studied composition; orange = receiving attention recently; 81
 white = not reported composition. Following our survey 82
 (Figure 1c) and aiming a comprehensive description of mixed 83
 cations and mixed halides perovskites, we start with the 84
 description of the simpler systems of mixed A cations (FA/ 85
 MA)PbI₃, (MA/Cs)PbI₃, (FA/Cs)PbI₃, (FA/Rb)PbI₃ perovskites 86
 (section 2) and the mixed X halides MAPb(I/Cl), 87
 FAPb(I/Cl), MAPb(I/Br), FAPb(I/Br), MAPb(Br/Cl), 88
 MAPb(I/Br/Cl), CsPb(I/Br), CsPb(Br/Cl), CsPb(I/Cl) perovskites 89
 (section 3). Based on important concepts described in 90
 sections 2 and 3, the more complex systems of simultaneously 91
 mixed A and mixed X (MA/Cs)Pb(I/Br), (FA/MA)Pb(I/Br), 92
 (FA/MA)Pb(I/Cl), (FA/Cs)Pb(I/Br), (Cs/FA/MA)Pb(I/Br), 93
 (Rb/FA/MA)Pb(I/Br), (Rb/Cs/FA)Pb(I/Br), (Rb/Cs/FA/ 94
 MA)Pb(I/Br) perovskites are described in section 4. 95

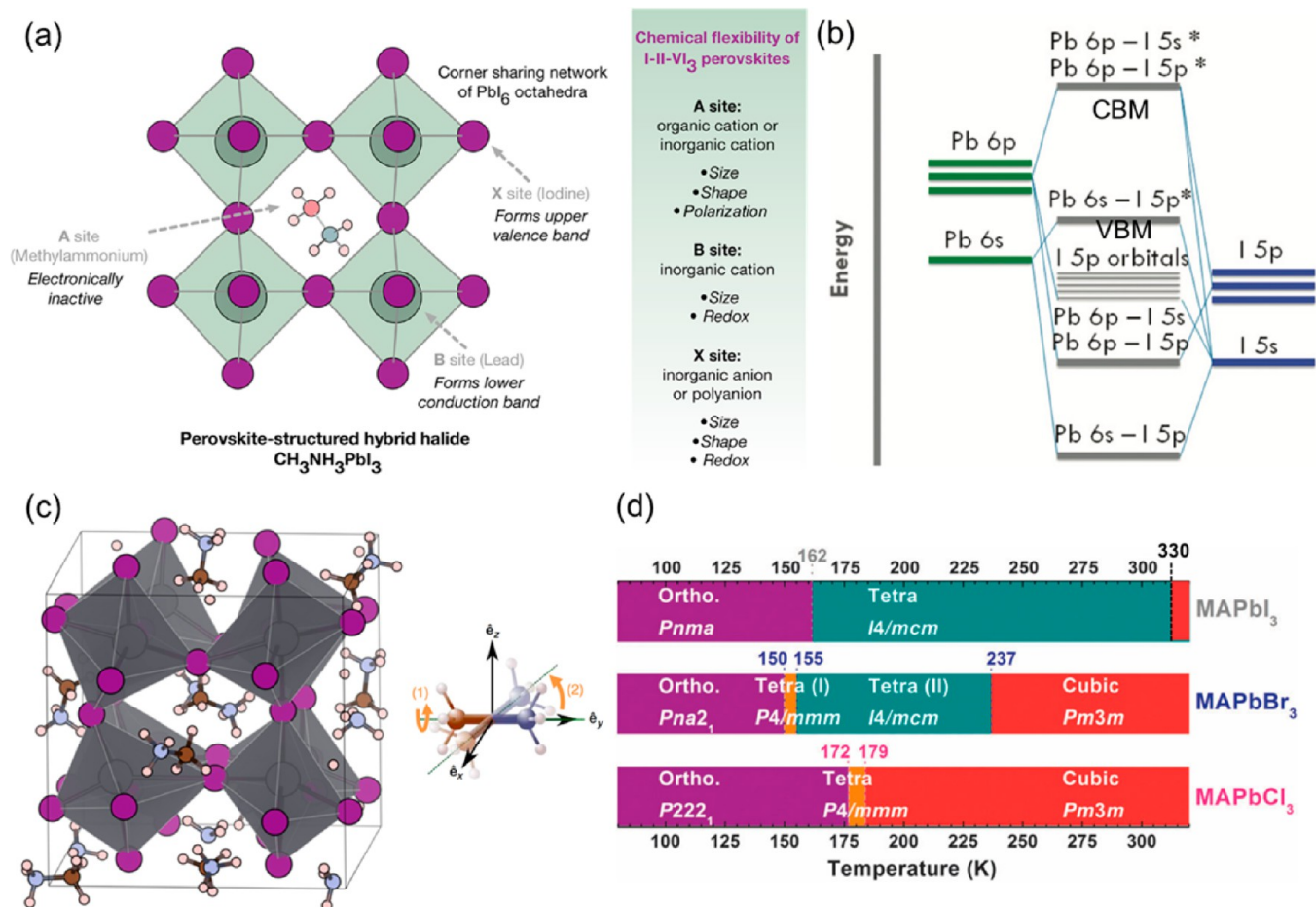


Figure 2. (a) Illustration of a $\text{CH}_3\text{NH}_3\text{PbI}_3$ perovskite structure with A, B, and X lattice sites. The redox chemistry of ions influence the valence and conduction band energies and orbital composition. Reprinted with permission from ref 67. Copyright 2015 American Chemical Society. (b) Illustration of energy bands in $\text{CH}_3\text{NH}_3\text{PbI}_3$ relative to isolated p and s atomic orbital energies of Pb and I. The CH_3NH_3^+ cation does not introduce states at the band edge. Reprinted with permission from ref 68. Copyright 2014 The Royal Society of Chemistry. (c) Dynamics of CH_3NH_3^+ cations generating distortions in the PbI_6 octahedra. Reprinted with permission from refs 69 and 70. Copyright 2015 American Chemical Society and Copyright 2015 Nature Publishing Group. (d) Summary of $\text{CH}_3\text{NH}_3\text{PbX}_3$ -based ($X = \text{I}, \text{Br}, \text{Cl}$) perovskites and space groups adopted as a function of temperature. Reprinted with permission from ref 71. Copyright 2016 The Royal Society of Chemistry.

96 Before starting to describe more complex mixed cation
 97 perovskite systems [e.g., (Rb/Cs/MA/FA)Pb(I/Br) or (MA/
 98 FA)Pb(I/Br)], we briefly outline our current understanding on
 99 structure–property relationship of simple perovskite systems
 100 (e.g., MAPbI_3 and FAPbI_3) and then gradually add one element
 101 (cation or halide) at a time and introduce the new properties of
 102 the material induced by the addition of this new element.
 103 MAPbI_3 is the most widely studied hybrid halide perovskite for
 104 photovoltaic applications^{13,24–26,72} and therefore its fundamen-
 105 tal properties were investigated most thoroughly.^{12,14,67,73–82} At
 106 room temperature, the Goldschmidt’s tolerance factor (t) of
 107 MAPbI_3 is 0.91^{83,84} with ionic radii of $\text{Pb}^{2+} = 0.132$ nm, $\text{I}^- =$
 108 0.206 nm, and $\text{MA}^+ = 0.18$ nm⁷ suggesting the tetragonal phase,
 109 which was confirmed experimentally by MAPbI_3 single crystal
 110 X-ray diffraction (XRD) data.^{7,9} It was reported that the PbI_6
 111 octahedra in MAPbI_3 are corner connected and the MA^+
 112 cations are filled in the octahedral interstices (Figure 2a).⁶⁷
 113 According to density functional theory (DFT) calculations, the
 114 ions within the inorganic PbI_6 octahedra have electronic
 115 configuration of Pb: $5d^{10}6s^26p^0$ and I: $5p^6$. The Pb $6s6p$ –I $5p$
 116 interactions are responsible for chemical bonding in PbI_6
 117 octahedra and generation of two bands: the valence band
 118 maximum (VBM) is formed of an antibonding (σ^*) Pb $6s$ –I $5p$

interactions, while the conduction band minimum (CBM) is
 119 formed of empty Pb $6p$ orbitals⁷³ and/or from Pb $6p$ –I $5p$
 120 interactions (Figure 2b).^{68,72,85–87} The MA^+ molecular units
 121 form bonding (σ) states deep (5 eV below Fermi level) in the
 122 VBM and they do not hybridize with the PbI_6 octahedra near
 123 VBM or CBM. Therefore, the MA^+ cations were proposed to
 124 do not contribute directly toward the band structure, but play a
 125 major role providing structural stability by charge compensa-
 126 tion within the PbI_6 octahedra based largely on electrostatic
 127 (van der Waals) interactions.^{15,26,74} Because of this lack of
 128 electronic interaction between MA^+ and PbI_6^{4-} , free carriers in
 129 MAPbI_3 are likely to diffuse along the corner shared PbI_6 chains
 130 in the crystal lattice.^{88,89} The VBM electrons that are
 131 photoexcited to CBM empty states determine the band gap
 132 (E_g) (Figure 2b).⁸⁵ A wide range of E_g values between 1.5 and
 133 1.61 eV have been reported for MAPbI_3 ^{14,15,80,90} with the E_g
 134 values generally higher for polycrystalline films possibly due to
 135 quantum confinement.^{7,90} Because the optimal E_g for a single-
 136 junction solar cell is between 1.1 and 1.4 eV on the basis of the
 137 Shockley–Queisser limit,⁹¹ efforts have been made to modify
 138 the MAPbI_3 structure.^{14,15}

In addition, material instabilities have been widely
 140 reported.^{28,29,92,93} It has been proposed that perovskites have 141

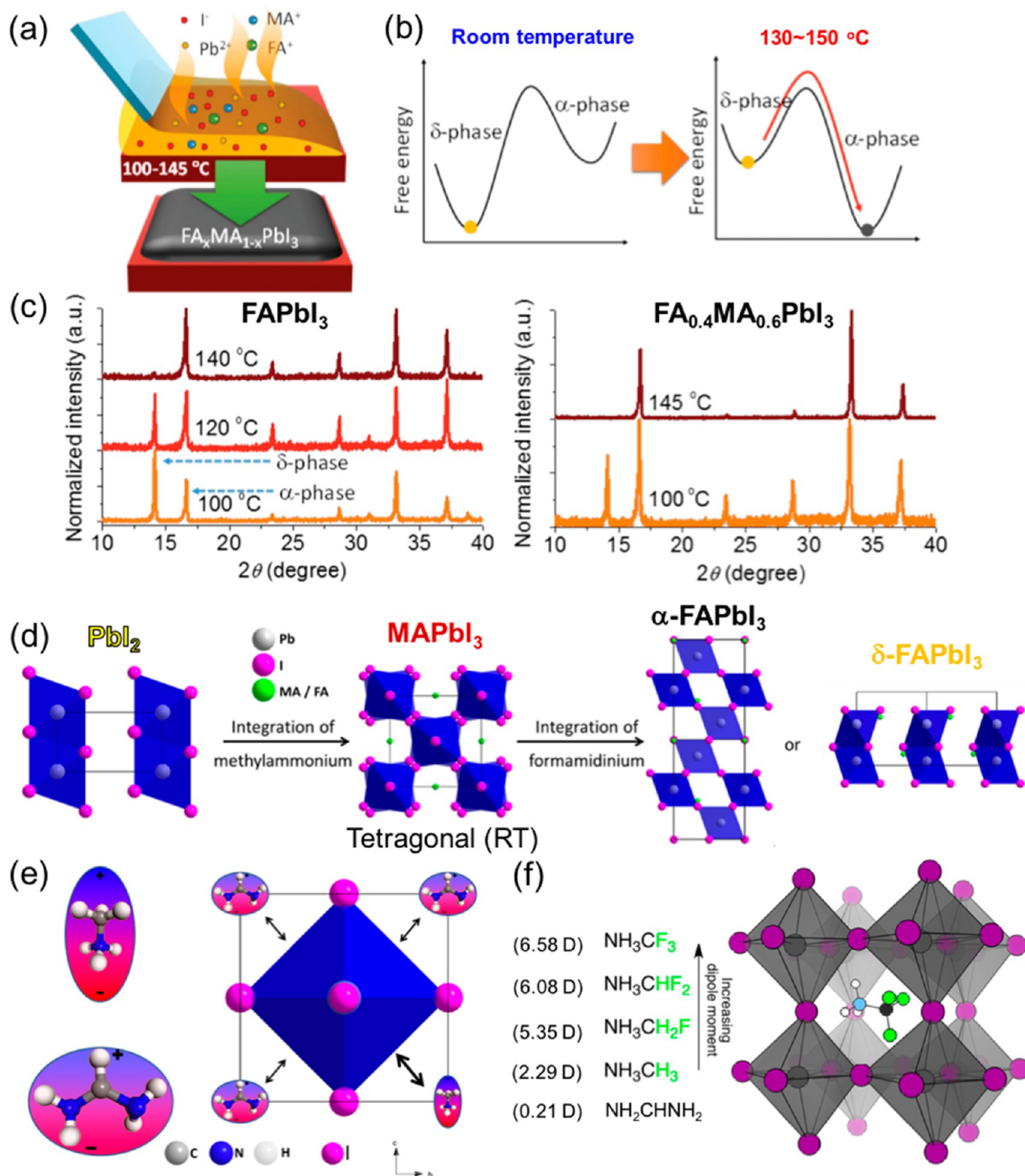


Figure 3. (a) Illustration of $\text{MA}_{0.6}\text{FA}_{0.4}\text{PbI}_3$ perovskite film formation by doctor-blade method. (b) Schematic diagrams of free energy versus α -phase and δ -phase of FAPbI_3 and $\text{MA}_x\text{FA}_{1-x}\text{PbI}_3$ perovskites. (c) XRD patterns for FAPbI_3 and $\text{MA}_x\text{FA}_{1-x}\text{PbI}_3$ perovskites as a function of annealing temperature. The reflection at $2\theta = 12^\circ$ and 14° are correlated to δ - and α -phase of FAPbI_3 , respectively. Reprinted with permission from ref 111. Copyright 2016 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim. (d) Crystal structures of PbI_2 where PbI_6 octahedra are connected via shared plane of three iodide ions. Incorporation of MA leads the layered octahedra to form a 3D MAPbI_3 network in tetragonal phase by reducing the number of shared iodide ions from three to one. Incorporation of $<20\%$ FA to MAPbI_3 structure, the tetragonal phase collapses and the trigonal phase of $\alpha\text{-FAPbI}_3$ is formed (3D network). However, at room temperature, the hexagonal phase of $\delta\text{-FAPbI}_3$ can also be formed. (e) Schematic illustration of the mechanism for the $\text{MA}_x\text{FA}_{1-x}\text{PbI}_3$ stabilization based on the dipole moments of cations. Reprinted with permission from ref 106. Copyright 2015 American Chemical Society. (f) Theoretically calculated dipole moments for selected cations. Reprinted with permission from ref 114. Copyright 2014 American Chemical Society.

Table 1. Summary of Perovskite Solar Cell Performances Employing Mixed Cations and/or Halide Perovskites^{a,b}

Solar cell architecture	Perovskite thickness (nm)	E_g (eV)	Electrode active area (cm ²)	J_{sc} (mA/cm ²)	V_{oc} (V)	FF (%)	PCE (%)	Hysteresis	Stability	ref.
FTO/c-TiO ₂ /mp-TiO ₂ /MA _{0.6} FA _{0.4} PbI ₃ /spiro/Au	~300 nm	1.53	0.285	18.15	1.027	71.5	13.4	—	—	102
ITO/PEDOT:PSS/MA _{0.6} FA _{0.4} PbI ₃ /PC61BM/Ca/Ag	~300 nm	— ^c	0.4	18.95	0.943	73.1	13.0	Hysteresis present.	—	109
FTO/c-TiO ₂ /(MAPbI ₃)(FAPbI ₃)/Au [No HTL]	—	1.53–1.54	0.09	9.58	0.77	54	4.0	—	—	103
FTO/c-TiO ₂ /MA _{0.13} FA _{0.87} PbI ₃ /spiro/Au	—	1.52 ± 0.02	0.0831	15.7	1	56	8.73	—	—	106
FTO/c-TiO ₂ /mp-TiO ₂ /MA _{0.6} FA _{0.4} PbI ₃ /spiro/Au	— ^c	— ^c	0.16	20.87	0.975	69.97	14.23	MA _{0.6} FA _{0.4} shows small hysteresis; Other compositions show hysteresis.	Storage in air and dark; RH < 60%; 120 days.	104
ITO/PtAA/MA _{0.6} FA _{0.4} PbI ₃ /ICBA/C ₆₀ /BCP/Cu [Doctor blade]	— ^c	1.55	—	23	1.03	77	18.3	Small hysteresis.	Storage in air; ~25 °C and RH 20%–60%; 30 days.	111
ITO/PEDOT:PSS/MA _{0.6} FA _{0.4} PbI ₃ /PC ₆₁ BM/Ag	—	—	0.1	20.96 ± 0.14	0.979 ± 0.016	64.7 ± 1.2	13.28 ± 0.37	Small hysteresis.	—	110
FTO/c-TiO ₂ /mp-TiO ₂ /MA _{0.6} FA _{0.4} PbI ₃ /WO ₃ /Ag [CVD]	—	1.57	0.2	20.85	1.04	73.15	15.86	Negligible hysteresis.	Storage in air; 30 days.	112
FTO/ZnO/MA _{0.6} FA _{0.4} PbI ₃ /spiro/Ag	~315	—	0.045	22.39	0.984	60.66	13.4	Hysteresis present.	Storage in N ₂ ; 21 days.	105
FTO/c-TiO ₂ /C60/MA _{1-x} FA _x PbI ₃ /spiro/Au [LP-VASP]	~340	1.55	0.09	22.51	1.00	73.56	16.48	Small hysteresis.	Storage in N ₂ ; 288 h.	113
FTO/c-TiO ₂ /mp-TiO ₂ /MA _{0.25} FA _{0.75} PbI ₃ /spiro/Ag	~300	~1.545	1	17.12	0.88	52.80	8.00	Hysteresis present.	—	107
FTO/c-TiO ₂ /MA _{0.7} FA _{0.3} PbI ₃ /spiro/Ag	—	—	—	22.03	1.03	75	17.02	Hysteresis present.	—	108
ITO/PEDOT:PSS/MA _{0.9} Cs _{0.1} PbI ₃ /PC60BM/Al	38 ± 9	1.54	0.13	10.10	1.05	73	7.68	—	Storage in air; ~20 °C and RH 11%; 42.5 h.	115
FTO/c-TiO ₂ /mp-TiO ₂ /MA _{0.91} Cs _{0.09} PbI ₃ /spiro/Au	~400	— ^c	0.09	22.57	1.06	76	18.1	Hysteresis present.	Storage in air; ~85 °C; 1 h.	116
FTO/c-TiO ₂ /Cs _{0.1} FA _{0.9} PbI ₃ /spiro/Ag or Au	— ^c	1.55	0.125	23.5	1.06	66.3	16.5 ^d	Hysteresis present.	Test 1: Exposure to white light; Sample temperature ~60 °C; RH < 50%. Test 2: Storage in dark; 25 °C; RH 85%. Storage in a desiccator and dark; RH 15%; 350 h. Storage in air; 1000 h.	100
FTO/c-TiO ₂ /Cs _{0.13} FA _{0.85} PbI ₃ /spiro/Ag	—	1.52	—	20.39	1.06	74	16.1 ^d	Hysteresis present.	—	101
FTO/c-TiO ₂ /mp-TiO ₂ /Cs _{0.2} FA _{0.8} PbI ₃ /spiro/Au	~435	1.56	0.16	21.5	1.017	70.1	15.69	—	—	117
FTO/c-SnO ₂ /C60-SAM/Cs _{0.2} FA _{0.8} PbI ₃ /spiro/Au	— ^c	— ^c	0.08	21.73 ± 0.51	1.03 ± 0.02	72.37 ± 1.18	16.18 ± 0.50 ^d	Hysteresis present.	Storage in a drybox under room light; 10 days.	118
FTO/c-SnO ₂ /C60-SAM/Cs _{0.2} FA _{0.8} PbI ₃ + 0.5 mol % Pb(SCN) ₂ /spiro/Au	— ^c	— ^c	0.09	21.94 ± 0.31	1.06 ± 0.01	77.77 ± 1.51	18.16 ± 0.54 ^d	Hysteresis present.	Storage in air; ~25 °C and RH ~ 40%; 4 h.	119
FTO/c-TiO ₂ /Cs _{0.8} FA _{1-x} PbI ₃ /spiro/Ag (x not determined)	— ^c	— ^c	0.12	20.4	1.09	74	16.4	Hysteresis present.	—	120
FTO/NiO/Cs _{0.15} FA _{0.85} PbI ₃ /PC61BM/PFN-Br/Ag	— ^c	— ^c	0.09	20.81	1.035	71.4	15.38	—	—	121
FTO/c-TiO ₂ /mp-TiO ₂ /Rb _{0.05} FA _{0.95} PbI ₃ /spiro/Au	~480 (capping layer)	1.53	—	23.93	1.07	67	17.16	Smaller hysteresis compared to FAPbI ₃ .	Storage in air; ~25 °C and RH ~ 55%; 1000 h.	121
FTO/c-TiO ₂ /mp-TiO ₂ /Rb _{0.05} FA _{0.95} PbI ₃ /spiro/Au	—	— ^c	0.65	23.8	1.03	65.9	16.2	—	Storage in air; ~25 °C and RH ~ 55%; 4 weeks.	122

†

Table 1. continued

Solar cell architecture	Perovskite thickness (nm)	E_g (eV)	Electrode active area (cm^2)	J_{sc} (mA/cm^2)	V_{oc} (V)	FF (%)	PCE (%)	Hysteresis	Stability	ref.
FTO/c-TiO ₂ /mp-TiO ₂ / C ₈₀ 2FA _{0.8} Pb _{1.2} Br _{0.16} / Spiro/Au	—	1.58	0.16	21.9	1.073	74.2	17.35	—	Storage in air; 1000 h.	117
FTO/c-TiO ₂ /mp-TiO ₂ /(MAP- bI ₃) _{1-x} (CsPbBr ₃) _x / Spiro/Au	360	— ^c	0.09	20.9 ± 0.42	1.07 ± 0.01	71 ± 2	15.9 ± 0.52	Small hysteresis	(1) Degradation test of unencapsulated devices under UV light (365 nm, 364 mW/cm ²) (2) Stability of unencapsulated devices in ambient air (25 °C, 20–30% RH) 2000 h aging test.	123
FTO/Nb-TiO ₂ /FA _{0.84} MA _{0.16} Pb ₁ (I _{0.84} Br _{0.16})/ Spiro/Au	~500	— ^c	0.12	23.4	1.006	74	17.6	—	2000 h aging test.	124
FTO/c-TiO ₂ /mp-TiO ₂ /FAPbI ₃ -MABr/ Spiro/ Au	~450	—	0.096	22.17	1.07	67.38	15.98	Small hysteresis	Storage in air (~50% RH, 23 °C) without encapsulation for 1000 h.	125
ITO/PEDOT:PSS/(FAP- bI ₃) _{0.8} (MAPbBr ₃) _{0.2} /C60/BCP/Ag	~280 ^c	—	0.1	20.6	0.88	65.9	12	—	—	126
FTO/c-TiO ₂ /mp-TiO ₂ /(FAPbI ₃) _{0.83} Pb (MABr ₃) _{0.15} / Spiro/Au	mp-TiO ₂ with PVSK in- filtrated (~200 nm) + capping layer (~650 nm)	1.57– 1.58	0.126	22.5	1.13	69	17.6	Yes	—	127
FTO/c-TiO ₂ /mp-TiO ₂ /(FA/MA)Pb(I/Br) ₃ / Spiro/Au	mp-TiO ₂ with PVSK in- filtrated (~150 nm) + capping layer (~400 nm)	—	0.36	21	1.01	69	14.5	Small hysteresis	ISOS-O-2 protocol for 1000 h.	128
ITO/PEDOT:PSS/(FAP- bI ₃) _{0.8} (MAPbBr ₃) _{0.2} /C60/BCP/Ag	—	— ^c	0.1	20.1 ± 0.5	0.87 ± 0.01	66.4 ± 1.2	11.8 ± 0.20	Small hysteresis	—	129
FTO/c-TiO ₂ /mp-TiO ₂ /(FA/MA)Pb(I/Br) ₃ / Spiro/Au PbI ₂ :FAI = 1.05 (PbI ₂ excess)	~400 ^c	1.6	0.16	24.6	1.16	73	20.8	No hysteresis	Sealed cells using epoxy and stored in a desiccator in dark for 1 month.	130
FTO/c-TiO ₂ /mp-TiO ₂ /FA _{0.7} MA _{0.3} Pb (I _{1.6} Br _{1.6})/ Spiro/Au	~400 ^c	1.64	0.16	23.7	1.14	76	20.67	Hysteresis varies according to composition.	—	90
ITO/PEDOT:PSS/MA _{0.88} FA _{0.20} PbI _{3-y} Cl _y / PC61BM/C60/LiF/Ag	~280	1.58	0.11	21.55 ± 0.55	1.10 ± 0.01	75 ± 2	17.45	No hysteresis when MA _{0.88} FA _{0.20} PbI _{3-y} Cl _y is annealed below 110 °C.	—	131
FTO/SnO ₂ /PC60BM/FA _{0.83} C _{80.17} Pb (I _{0.8} Br _{0.4}) ₃ / Spiro/Ag	— ^c	1.74	0.0919	19.4	1.2	75.1	17.1	—	—	132
ITO/NiO/FA _{0.83} C _{80.17} Pb(I _{0.6} Br _{0.4}) ₃ /LiF/ PC60BM/SnO ₂ /ZTO/ITO/LiF/Ag	— ^c	1.63	1	18.7	0.98	78.8	14.5	—	single-junction perovskite device with no additional encapsulation during 1000 h of continuous maximum-power-point tracking	133
FTO/SnO ₂ /C60/C _{80.1} FA _{0.83} Pb(I _{0.4} Br _{0.4}) ₃ / Spiro/Au	—	—	0.0919	23	1.06	75	18.3	—	Aged under full AM1.5 spectrum at V_{oc} in air without UV filter.	134
FTO/c-TiO ₂ /mp-TiO ₂ / C ₈₀ 2FA _{0.8} Pb _{1.2} Br _{0.16} / Spiro/Au	—	1.49	0.16	21.9	1.073	74.2	17.35	—	Storage in air; 1000 h.	117
FTO/c-TiO ₂ /mp-TiO ₂ / C _{80.05} (FA _{0.83} MA _{0.17}) _{0.95} Pb(I _{0.83} Br _{0.17}) ₃ / Spiro/Au	~500 ^c	1.6 ^c	0.16	22.69 ± 0.75	1.132 ± 0.025	74.8 ± 1.8	19.2 ± 0.91	Yes	Under constant illumination and maximum power tracking for 250 h; Cells held at room temperature.	65
FTO/c-TiO ₂ /mp-TiO ₂ /C _{80.1} FA _{0.8} MA _{0.1} Pb (I _{0.83} Br _{0.17}) ₃ / Spiro/Au	~500	— ^c	0.16	21.5	1.155	73	18.1	Yes	300 h storage; dry air and in dark.	135
FTO/c-TiO ₂ /mp-TiO ₂ / Rb _{0.1} (FAPbI ₃) _{0.9} (MAPbBr ₃) _{0.1} /PTAA/Au 15% PbI ₂ excess 5% RbI doping	— ^c	— ^c	0.16	22.86 ± 0.37	1.123 ± 0.011	69.6 ± 1.9	17.68 ± 0.91	Exhibit some degree of hysteresis	(1) Under continuous light illumination; cells kept under ambient conditions (40–50% RH, 25 °C).	136

Table 1. continued

Solar cell architecture	Perovskite thickness (nm)	E_g (eV)	Electrode active area (cm^2)	J_{sc} (mA/cm^2)	V_{oc} (V)	FF (%)	PCE (%)	Hysteresis	Stability	ref.
FTO/c-TiO ₂ /mp-TiO ₂ / Rb _{0.05} FA _{0.98} MA _{0.15} PbI _{2.55} Br _{0.45} /spiro/Au	—	— ^c	0.5	22.5	1.168	74.7	19.6	—	(2) Steady-state efficiency tracking maximum power point under N ₂ environment.	122
FTO/c-TiO ₂ /mp-TiO ₂ /Rb _{0.05} Cs _{0.10} FA _{0.85} Pb (I _{0.88} Pb _{0.17}) ₃ /spiro/Au	—	—	0.16	22.3	1.157	75	19.3	Yes	Storage in air; ~25 °C and RH ~55%; 4 weeks.	137
FTO/c-TiO ₂ /mp-TiO ₂ / Rb _{0.05} Cs _{0.05} (FA _{0.88} MA _{0.17}) _{0.99} Pb(I _{0.88} Br _{0.17}) ₃ / PTAA/Au	~500	1.63	0.16	22.8	1.180	81	21.8	Yes	Device aged for 500 h at 85 °C under continuous full illumination and maximum power tracking in a nitrogen atmosphere.	—
FTO/c-TiO ₂ /mp-TiO ₂ /Rh(5%) in FA _{0.75} (MA _{0.6} Cs _{0.4}) _{0.25} PbI ₂ Br/PTAA/Au	mp-TiO ₂ with PVSK in- filtrated (~180 nm) + capping layer (~400 nm)	1.73	0.1764	18.3	>1.1 ^c	— ^c	17.4 (steady state power)	Negligible	(1) Light stability for 12 h under continuous 1 sun illumination and applied maximum power point voltage in N ₂ environment with (controlled 25 °C). (2) 12 h—light/dark cycles.	138

^aPerovskite thickness, bandgap (E_g), electrode active area, solar cell parameters of short-circuit current (J_{sc}), open-circuit voltage (V_{oc}), fill factor (FF), and power conversion efficiency (PCE) are indicated. Additional notes on hysteresis and stability are indicated. ^bAbbreviations: ITO = indium tin oxide; FTO = fluorine doped tin oxide; polyTPD = poly(*N,N'*-bis(4-butylphenyl)-*N,N'*-bis(phenyl)benzidine); c-TiO₂ = compact layered TiO₂; spiro-MeOTAD = 2,7'-7,7'-tetrakis(*N,N*-di-*p*-methoxyphenylamine)-9,9'-spirobifluorene; PTAA = poly[bis(4-phenyl)(2,4,6-trimethylphenyl)amine]; PEDOT:PSS = poly(3,4-ethylenedioxythiophene); poly(styrenesulfonate); poly(styrenesulfonate); ICBM = indene-C60 bisadduct; BCP = 2,9-dimethyl-1,10-phenanthroline; PC60BM = [6,6]-phenyl-C60 butyric acid methyl ester; PC61BM = [6,6]-phenyl C61 butyric acid methyl ester; PFN-Br = poly[(9,9-bis(3'-(*N,N*-dimethyl)-*N*-ethylammonium)propyl)-2,7-fluorene]-*alt*-2,7-(9,9-dioctylfluorene)]dibromide. C60-SAM = fullerene-self-assembled-monolayer. Pb(SCN)₂ = lead(II) thiocyanate. PFN-Br = poly[(9,9-bis(3'-(*N,N*-dimethyl)-*N*-ethylammonium)propyl)-2,7-fluorene]-*alt*-2,7-(9,9-dioctylfluorene)] dibromide. ZTO = zinc tin oxide. ^cThickness and absorption measured, but values not explicitly stated. ^dAveraged values based on forward and reverse scans reported.

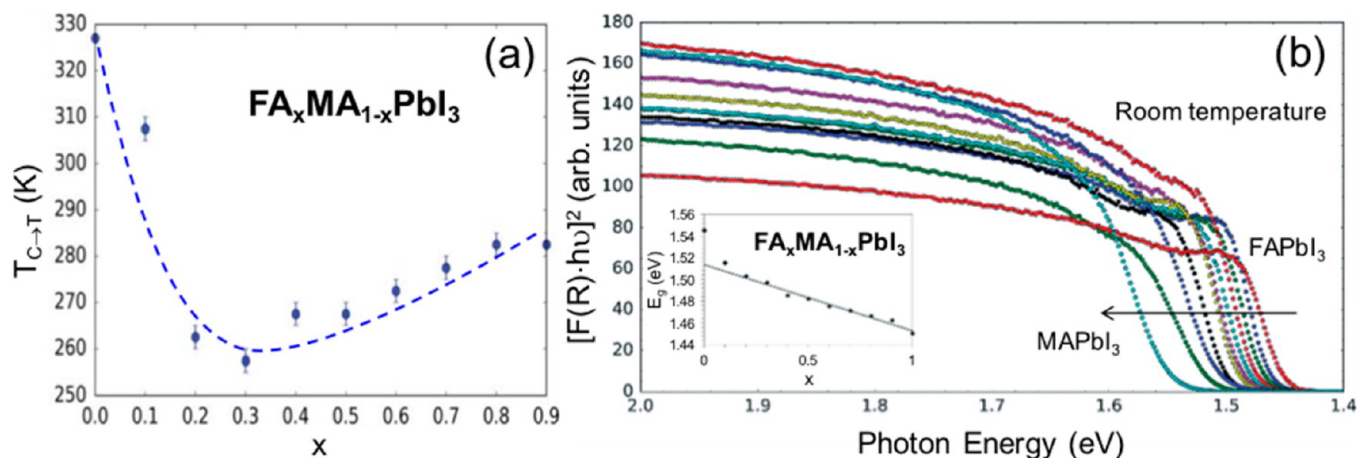


Figure 4. (a) Phase transition temperatures from cubic to tetragonal ($T_{C\rightarrow T}$) and (b) Tauc plots of UV-vis reflectance spectra measured at room temperature and (inset) extracted band gap (E_g) of $\text{FA}_x\text{MA}_{1-x}\text{PbI}_3$ ($0 \leq x \leq 1$) perovskites. Reprinted with permission from ref 83. Copyright 2016 The Royal Society of Chemistry.

142 an inherently “soft” structure (“plastic crystals”⁹⁴) and therefore
 143 several dynamical phenomena characterized by different time
 144 scales exist in perovskites (from femtosecond up to minutes)
 145 influencing strongly on the overall optoelectronic properties
 146 (e.g., energy levels, bandgap, balanced/imbalanced charge
 147 transport, etc.) of perovskites (Figure 2c).^{69,70,95–97} Further-
 148 more, MAPbI_3 undergoes a phase transition from the tetragonal
 149 to cubic phase at $\sim 54^\circ\text{C}$ (Figure 2d), a temperature that is of
 150 relevance during typical solar cell operation.^{9,49} Therefore, it is
 151 imperative to develop perovskites insensitive to composition
 152 instabilities under stress conditions for solar cell operation. For
 153 certification, solar modules must operate successfully in the
 154 temperature range between -40 and $+85^\circ\text{C}$.⁹⁸ In contrast to
 155 MAPbI_3 , FAPbI_3 was reported to be free from phase transition
 156 between a wide temperature range from 25 to 150°C .^{99,100}
 157 Chemical compositional engineering⁶³ (or alloying^{8,101}) was
 158 shown to be an effective strategy (e.g., substitution of MA^+ by
 159 other cations such as FA^+ , Cs^+ , Rb^+ , Figure 1c) to further
 160 enhance the optoelectronic properties of MAPbI_3 as evidenced
 161 by reported certified efficiencies (Figure 1b).^{5,7,90}

2. MIXED A CATIONS

162 **2.1. Binary (FA/MA)PbI₃ System.** Perturbing the A cation
 163 size can influence the optical properties by deforming the BX_6
 164 octahedron network. A larger (e.g., $\text{FA}^+ = 0.19\text{--}0.22\text{ nm}$) or
 165 smaller (e.g., $\text{Cs}^+ = 0.167\text{ nm}$, $\text{Rb}^+ = 0.152\text{ nm}$) cation causes
 166 the lattice to expand or contract leading to the change of B–X
 167 bond length, which has been shown to influence E_g .⁷ Pellet et
 168 al. reported the first mixed A cation describing the bandgap
 169 tunability of $\text{MA}_x\text{FA}_{1-x}\text{PbI}_3$ -based solar cells by varying the
 170 ratio of MA to FA.¹⁰² The MA/FA mixed perovskites were
 171 synthesized by dipping the predeposited PbI_2 with premixed
 172 $\text{MA}_x\text{FA}_{1-x}\text{I}$ solution in isopropanol. They showed that the
 173 $\text{MA}_{0.6}\text{FA}_{0.4}\text{PbI}_3$ composition resulted in the best PCE of 13.4%
 174 (Table 1) with the absorption edge extending up to $\sim 810\text{ nm}$
 175 ($E_g \sim 1.53\text{ eV}$), which was also similar to the pure FAPbI_3
 176 perovskite. Later several works have been published showing
 177 better $\text{MA}_x\text{FA}_{1-x}\text{PbI}_3$ film quality (higher film coverage, fewer
 178 pin-holes, higher crystal quality, larger grain sizes, smaller
 179 roughness within the grains, and alternative selective
 180 contacts),^{103–108} employing inverted structure,^{109,110} and
 181 using perovskite deposition techniques other than spin-coating
 182 (Table 1),^{111–113} such as doctor blading (Figure 3a),¹¹¹

chemical vapor deposition (CVD),¹¹² and low-pressure vapor-
 assisted solution process (LP-VASP)¹¹³ (Table 1). Based on
 Table 1, among optimized MA/FA ratios, there is a consensus
 that $\text{MA}_{0.6}\text{FA}_{0.4}\text{PbI}_3$ composition results in the best PCE as
 high as 18.3%.¹¹¹ Deng et al.¹¹¹ employed additionally Cu as
 the top electrode in their devices and observed high device
 performances. Chen et al.¹¹³ reported devices with larger active
 area of 1 cm^2 generating PCE of 8%, while their smaller cells
 (0.09 cm^2) generated PCE of 16.48%.
 191

The highest PCE for MAPbI_3 solar cells in a standard device
 architecture was reported to reach $\sim 20\%$,^{139,140} whereas in an
 inverted structure, the highest PCE of 18.1% was reported.¹⁴¹
 Up to now, employing complex strategies (e.g., solvent-
 engineering, HPbI_3 precursor, organic-cation displacement),
 the best PCE reported for FAPbI_3 perovskite solar cells reached
 13.5–18%, which is somehow lower than the MAPbI_3 -based
 ones.^{64,99,142–144} Because of the larger ionic radius of FA^+
 compared to MA^+ , FAPbI_3 perovskites are expected to have a
 smaller E_g (1.47 to 1.55 eV depending on the fabrication
 methods¹⁴⁵) compared to MAPbI_3 .^{99,143,146,147} As the con-
 sequence, higher PCE would be expected for FAPbI_3 solar cells
 as light harvesting range extends further into the near-infrared.
 Comparative works have shown that the short-circuit current
 (J_{sc}) of FAPbI_3 devices is higher than that of MAPbI_3 ,
 confirming the ability of extended light-absorption of FAPbI_3
 perovskites. The main loss in PCE comes from the poor fill
 factor (FF).^{143,148} Several studies proposed that there is a
 fundamental limitation associated with the phase instability of
 the pure FAPbI_3 in ambient conditions (either in the form of
 polycrystalline film^{8,63,106,149} or single crystal¹⁵⁰). As a matter of
 fact, FAPbI_3 possess two polymorphs (Figure 3b–
 213 d).^{8,82,106,119,149–151} FAPbI_3 crystallizes at room temperature
 as (i) photoinactive nonperovskite, hexagonal $\delta\text{-FAPbI}_3$
 215 perovskite (“delta or yellow phase”, space group $P6_3mc$ ⁸²)
 216 formed from face-sharing PbI_6 octahedra (Figure 3d), and (ii)
 217 photoactive trigonal $\alpha\text{-FAPbI}_3$ perovskite (“alpha or black
 218 phase”, space group $P3m1$ ⁸² or $Pm\bar{3}m$ ¹⁴⁹) formed at higher
 219 temperatures, $\sim 125\text{--}165^\circ\text{C}$ (Figure 3d).^{65,82,101,111,121} Similar
 220 polymorphism was reported for $\alpha,\delta\text{-CsPbI}_3$ system.^{65,121} After
 221 $\alpha\text{-FAPbI}_3$ is formed, a slow phase transformation to $\delta\text{-FAPbI}_3$ is
 222 reported when kept at room temperature, which leads to
 223 unstable device operation.^{8,111} It is interesting to note that the
 224 incorporation of MA^+ into FAPbI_3 structure results in a much
 225

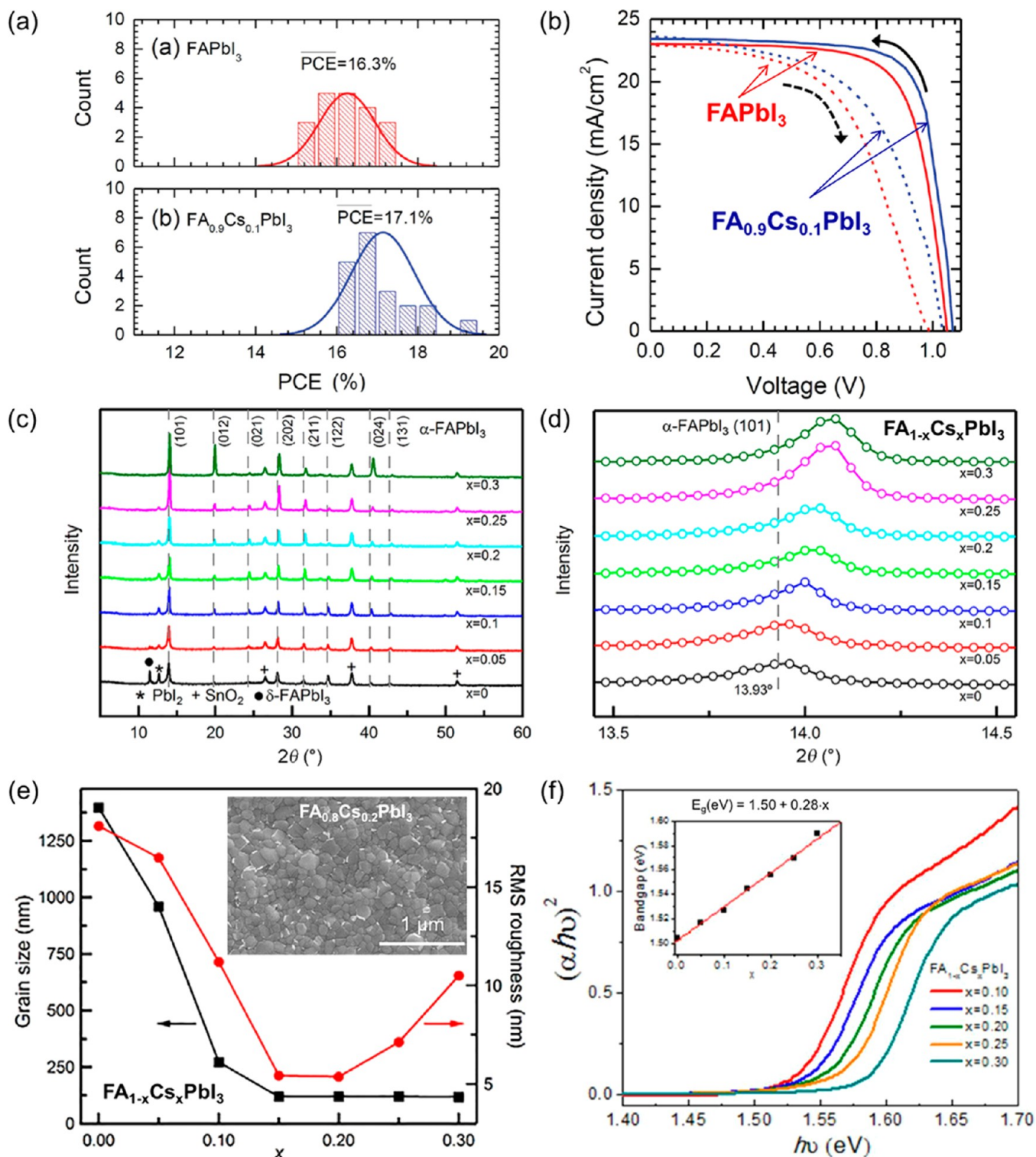


Figure 5. (a) Statistical distribution of perovskite solar cell PCE based on FAPbI_3 (red) and $\text{FA}_{1-x}\text{Cs}_x\text{PbI}_3$ (blue) and (b) $J-V$ curves of the best performing devices for FAPbI_3 (red) and $\text{FA}_{1-x}\text{Cs}_x\text{PbI}_3$ (blue) at reverse scans (solid lines) and forward scans (dotted lines). Reprinted with permission from ref 100. Copyright 2015 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim. (c,d) XRD patterns and (e) grain size and surface RMS roughness of $\text{FA}_{1-x}\text{Cs}_x\text{PbI}_3$ with $x = 0$ to 0.30. Reprinted with permission from ref 118. Copyright 2016 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim. (f) Tauc plots and extracted band gaps (E_g , inset) for $\text{FA}_{1-x}\text{Cs}_x\text{PbI}_3$ with $x = 0$ to 0.30. Reprinted with permission from ref 101. Copyright 2015 American Chemical Society.

226 more stable $\text{MA}_x\text{FA}_{1-x}\text{PbI}_3$ perovskite structure,^{83,152,153} its
 227 origins of which were further investigated by Binek et al.¹⁰⁶ It
 228 has been proposed that the incorporation of a smaller cations
 229 (MA^+) with a large dipole moment exhibits stronger interaction

with the PbI_6 octahedra, which stabilizes the 3D arrangement of
 230 $\alpha\text{-FAPbI}_3$ with little lattice shrinkage or changes in the optical
 231 properties (Figure 3e). MA^+ has a dipole moment ten times
 232 higher than that of FA^+ (Figure 3f).¹¹⁴ Similarly, based on DFT
 233

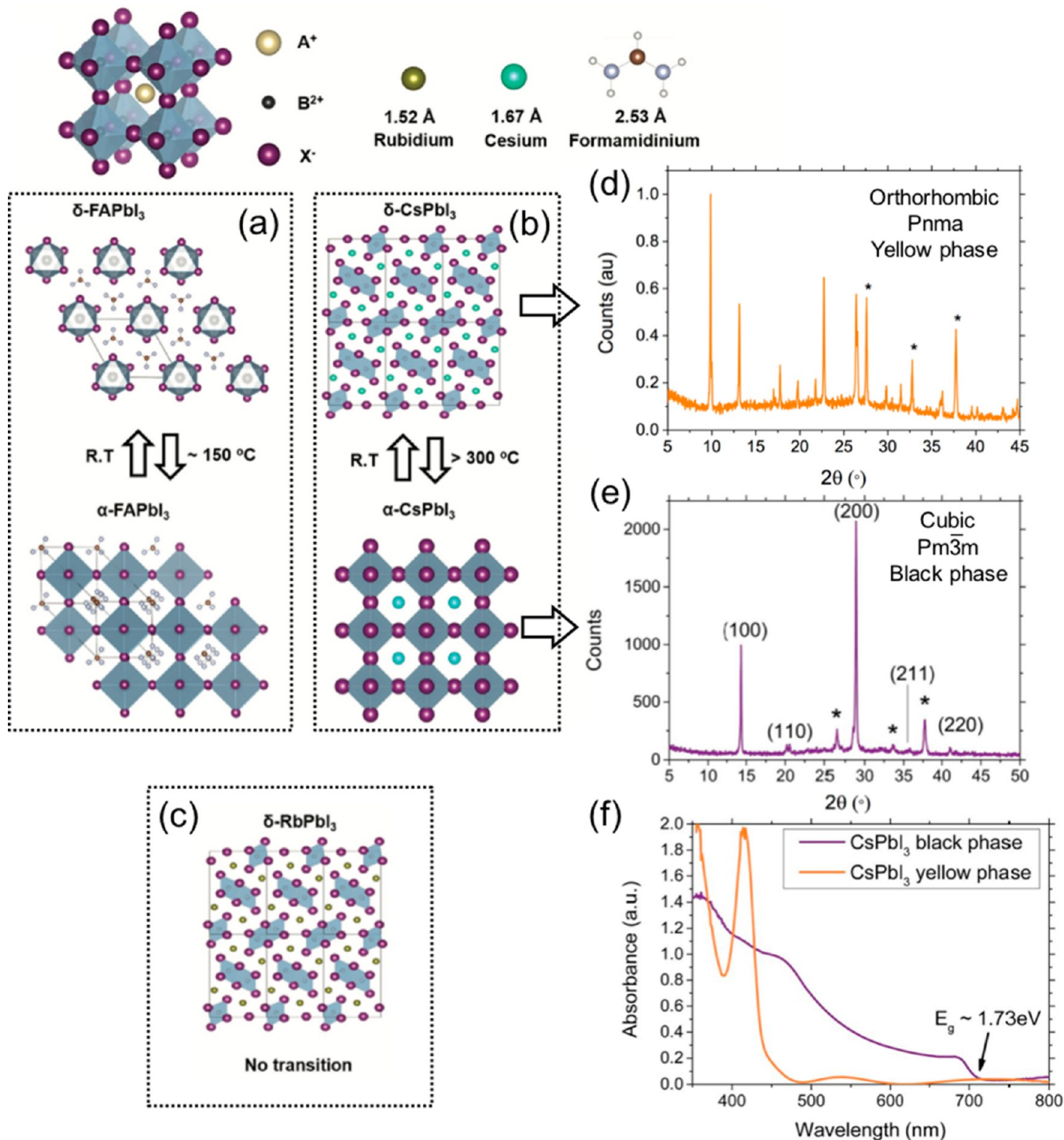


Figure 6. Illustration of ABX₃ perovskite crystal structure and possible candidates for the A-site cation (FA⁺, Cs⁺, Rb⁺) with corresponding ionic radius. (a) Schematic representations of the polymorphs in (b) α,δ -FAPbI₃ and (c) α,δ -CsPbI₃ perovskites. The $\alpha \rightarrow \delta$ phase transitions for CsPbI₃ is much higher than in FAPbI₃. At room temperature, the δ -phases are preferentially stabilized. Reprinted with permission from ref 121. Copyright 2017 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim. (d) XRD of CsPbI₃ thin film in yellow phase and (e) black phase. Peaks with * correspond to the FTO substrate. (f) Absorbance spectra of black and yellow phases of CsPbI₃ thin films. The measurements were done in air, with the black phase perovskite coated with poly methyl methacrylate (PMMA) to minimize transformation into the yellow phase. Reprinted with permission from ref 155. Copyright 2015 The Royal Society of Chemistry.

234 calculations El-Mellouhi et al.¹⁵⁴ proposed that CH₃PH₃⁺,
 235 CH₃SH₂⁺, and SH₃⁺ cations could also result in stronger
 236 cohesion between PbI₆ octahedra and A cation within the
 237 crystal and attain enhanced stability while maintaining a
 238 suitable bandgap for solar cell applications. As described later

in section 2.3, Cs⁺ that has no dipole moment also helps
 239 stabilize the α -FAPbI₃ phase. The transition from tetragonal
 240 MAPbI₃ to trigonal α -FAPbI₃ is observed when the FA⁺/MA⁺
 241 ratio in the structure is >80%.¹⁰⁶ The phase transition behavior
 242 as a function of FA_xMA_{1-x}PbI₃ (0 ≤ x ≤ 1) composition and
 243

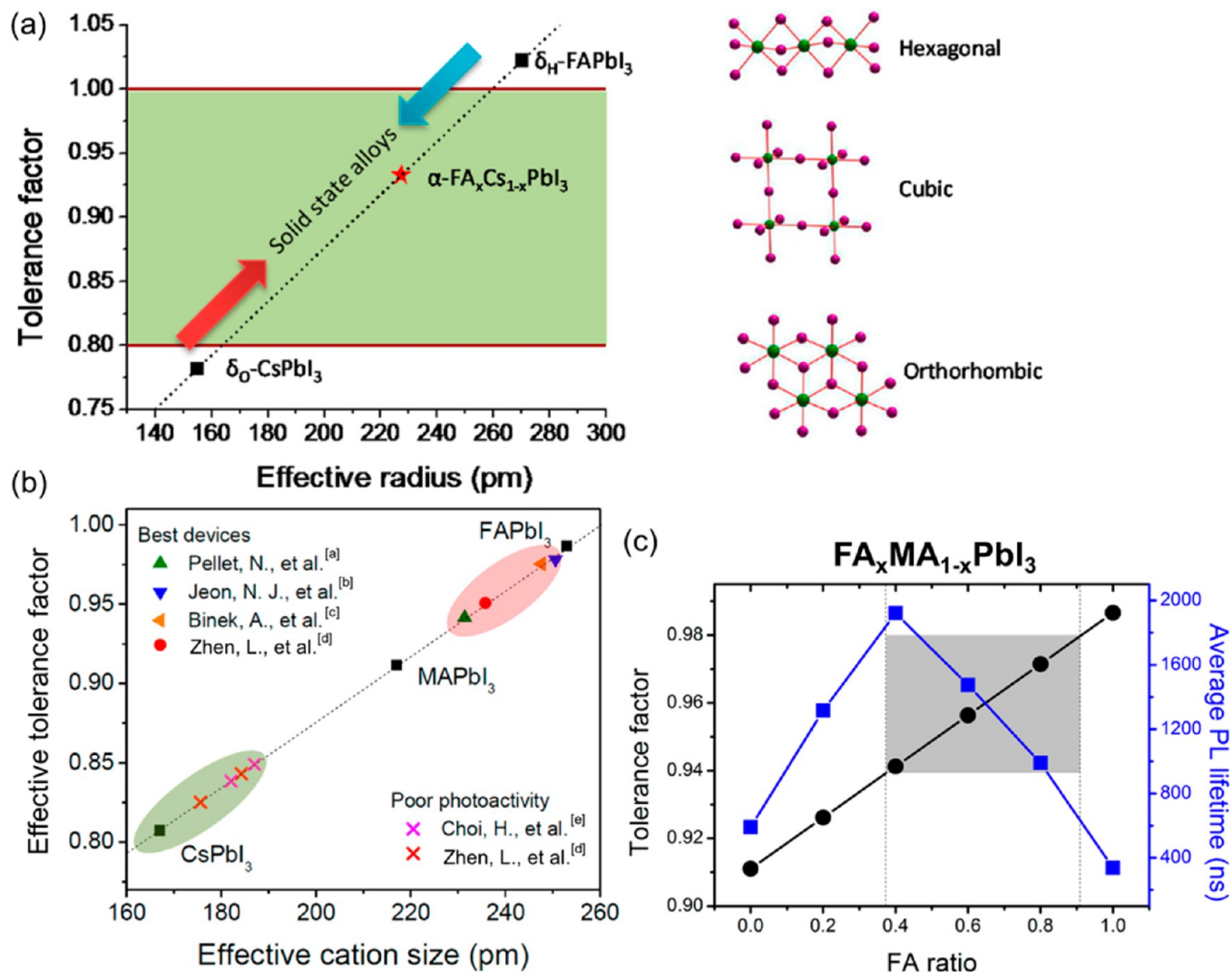


Figure 7. (a) Goldschmidt tolerance factor as empirical index for predicting stable perovskite crystal structure. Tolerance factor between 0.8 and 1.0 (green area) is favorable for cubic perovskite structure, and larger (>1) or smaller (<0.8) values generate nonperovskite structures. (b) Summary of the effective tolerance factors of different mixed perovskite alloys from the literature. [a] MA_{0.6}FA_{0.4}PbI₃,¹⁰² [b] (FAPbI₃)_{1-x}(MAPbBr₃)_x ($x = 0.15$),⁶³ [c] MA_{0.13}FA_{0.87}PbI₃,¹⁰⁶ [d] Cs_{0.15}FA_{0.85}PbI₃,¹⁰¹ [e] MA_{0.9}Cs_{0.1}PbI₃.¹¹⁵ Reprinted with permission from ref 101. Copyright 2016 American Chemical Society. (c) Correlation between Goldschmidt tolerance factor (black circles) and PL (blue squares) of the mixed cation FA_xMA_{1-x}PbI₃ perovskites. Reprinted with permission from ref 152. Copyright 2016 American Chemical Society.

244 temperature ($150 \text{ K} \leq T \leq 300 \text{ K}$) was investigated
 245 experimentally by Weber et al.⁸³ At 300 K, the FA_xMA_{1-x}PbI₃
 246 with $0.2 \leq x \leq 1$ composition revealed a cubic structure with
 247 unit cell dimensions of $\sim 6.3 \text{ \AA}$. However, the tetragonal phase
 248 was observed for MAPbI₃ ($x = 0$) and FA_{0.1}MA_{0.9}PbI₃ ($x = 0.1$)
 249 perovskites. Their summary data (Figure 4a) extracted from
 250 variable temperature single crystal XRD measurements show a
 251 nonmonotonic behavior of the phase transition temperatures
 252 from cubic to tetragonal ($T_{C \rightarrow T}$). Starting from $x = 0$ (MAPbI₃)
 253 showing $T_{C \rightarrow T} \sim 327 \text{ K}$, $T_{C \rightarrow T}$ decreased sharply reaching ~ 257
 254 K at $x = 0.3$. Subsequent increase in the FA ($0.3 \leq x \leq 0.9$)
 255 content leads to a slow steady increase in $T_{C \rightarrow T} \sim 298 \text{ K}$ at $x =$
 256 0.9 (Figure 4a). UV-vis measurements on FA_xMA_{1-x}PbI₃ ($0 \leq$
 257 $x \leq 1$) revealed a linear decrease in the optical band gap with
 258 increase in the FA content from $0.2 \leq x \leq 1$ (Figure 4b),
 259 corresponding to the range of compositions adopting the cubic
 260 unit cell at room temperature. MAPbI₃ ($x = 0$) and
 261 FA_{0.1}MA_{0.9}PbI₃ ($x = 0.1$) showed deviation from the linear

trend due to the tetragonal structure stabilized at room 262
 temperature. 263

2.2. Binary (MA/Cs)PbI₃ System. To the best of our 264
 knowledge, there are only two reports about the (MA/Cs)PbI₃ 265
 system. The first study was conducted by Choi et al.,¹¹⁵ in 266
 which they optimized Cs_xMA_{1-x}PbI₃ devices with $x = 0.1$ 267
 (inverted structure: ITO/PEDOT:PSS/Cs_{0.1}MA_{0.9}PbI₃/ 268
 PCBM/Al) achieved a PCE of 7.68%. The second study was 269
 performed by Niu et al.¹¹⁶ reporting a higher efficiency (reverse 270
 scan) of 18.1% in their optimized Cs_xMA_{1-x}PbI₃ devices with x 271
 = 0.09 (regular structure: FTO/c-TiO₂/mp-TiO₂/ 272
 Cs_{0.09}MA_{0.91}PbI₃/Spiro-MeOTAD/Au) (Table 1). When the 273
 PCEs of Cs_xMA_{1-x}PbI₃ devices are compared to their 274
 individual reference cells, i.e., MAPbI₃, both studies observed 275
 increase of $\sim 20\%$ ¹¹⁵ and $\sim 15\%$,¹¹⁶ respectively. Although, the 276
 phase transitions for (MA/Cs)PbI₃ system were not studied, 277
 Niu et al.¹¹⁶ showed better thermal stability of Cs_xMA_{1-x}PbI₃ 278
 with $x = 0.09$ and $x = 0.20$ than that of MAPbI₃. Contrary to 279
 MA⁺ (section 2.1), which has a large dipole moment, Cs⁺ 280

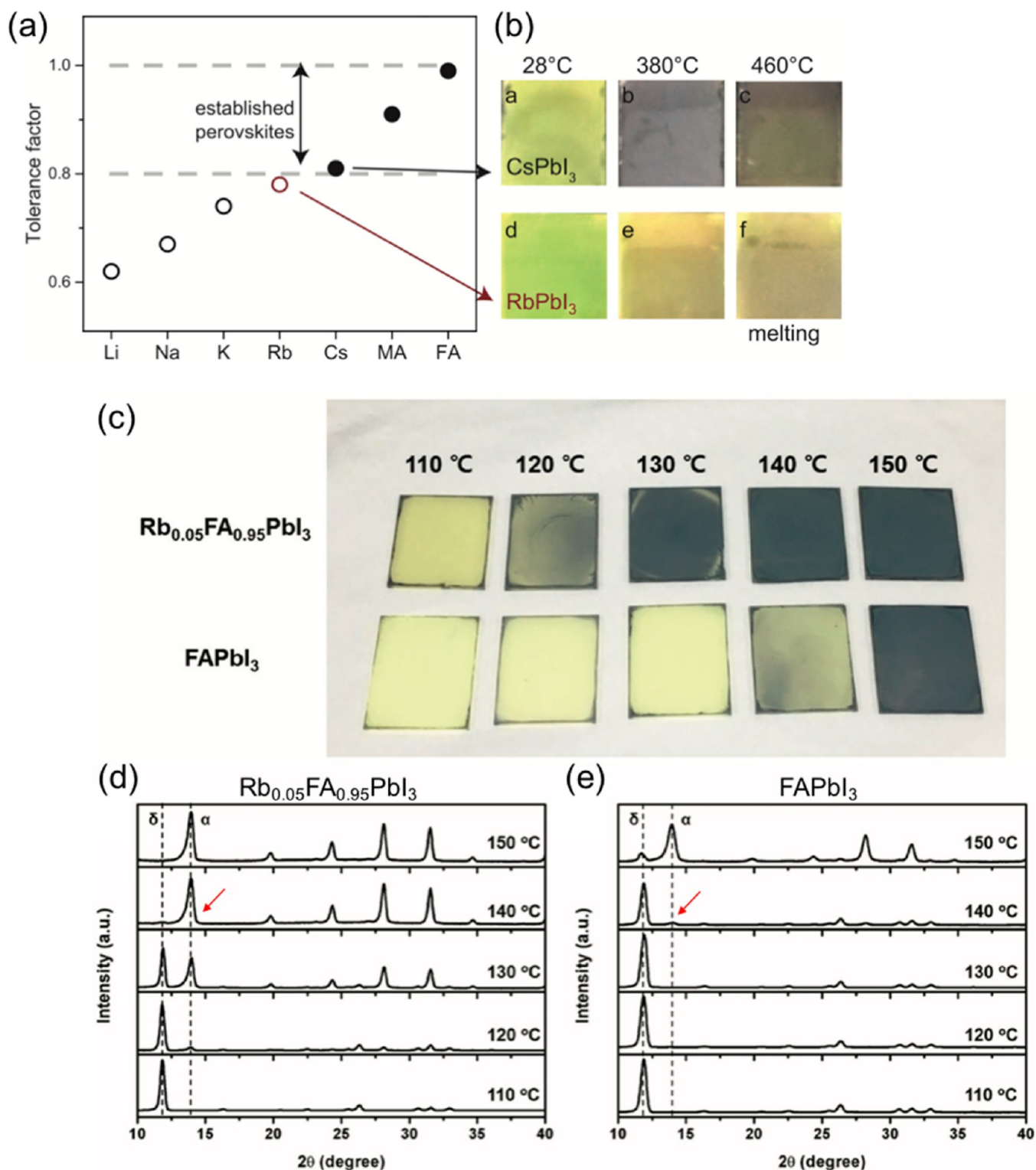


Figure 8. (a) Goldschmidt tolerance factor of APbI₃ (A = Li, Na, K, Rb, Cs, MA, FA) perovskites. Empirically, perovskites with a tolerance factor between 0.8 and 1.0 (dashed lines) show a photoactive black phase (solid circles) as opposed to nonphotoactive phases (open circles). Rb (red open circle) is very close to this limit. (b) CsPbI₃ and RbPbI₃ at 28, 380, and 460 °C. Irreversible melting occurs at 460 °C. RbPbI₃ never shows a black phase. Reprinted with permission from ref 137. Copyright 2016 American Association for the Advancement of Science (AAAS). (c) Photographs and XRD patterns of (d) Rb_{0.05}FA_{0.95}PbI₃ and (e) FAPbI₃ films after heating at specified temperatures for 5 min. Diffraction peaks corresponding to α - and δ -FAPbI₃ are indicated. Reprinted with permission from ref 121. Copyright 2016 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim.

281 cation (ionic radius = 0.167–0.181 nm) has no dipole moment,
 282 therefore it would be interesting to study the fundamental
 283 aspects of stabilization of MA/Cs and FA/Cs combinatorial
 284 systems.

2.3. Binary (FA/Cs)PbI₃ System. The number of reports
 on (FA/Cs)PbI₃ system is also scarce (Table 1).^{100,101,118–120}
 In all these studies, similar to the Cs_xMA_{1-x}PbI₃ case, it is
 noticed that Cs_xFA_{1-x}PbI₃ with Cs quantities in the range of x 285
 286
 287
 288

289 = 0.1 to 0.2 shows some enhancement in PCEs when compared
 290 to pure FAPbI₃ (Table 1, Figure 5a). These enhancements in
 291 PCEs can go as high as ~5%,¹⁰⁰ ~13%,¹⁰¹ ~16%.¹¹⁸ Hysteresis
 292 phenomena (Figure 5b) are present in all these studies
 293 employing devices with a standard structure of (i) FTO/c-
 294 TiO₂/Cs_xFA_{1-x}PbI₃/spiro-MeOTAD/Au (or Ag)^{100,101,119} and
 295 (ii) FTO/c-SnO₂/C60-SAM/Cs_{0.2}FA_{0.8}PbI₃/spiro-MeOTAD/
 296 Au.¹¹⁸ Encapsulated Cs_xMA_{1-x}PbI₃ devices were reported to
 297 show some promises regarding long-term stability tests under
 298 continuous white light illumination (250 h)¹⁰⁰ or unencapsu-
 299 lated devices under storage at low relative humidity (~15%) for
 300 350 h.¹⁰¹ Lee et al.¹⁰⁰ proposed that partial substitution of FA⁺
 301 by Cs⁺ leads to the contraction of the cuboctahedral volume
 302 (Figure 5c,d), and consequently enhances the interactions
 303 between FA⁺ and iodine. Calculated lattice parameters of
 304 Cs_xMA_{1-x}PbI₃ vary from ~6.363 Å (*x* = 0) down to ~6.310 Å
 305 (*x* = 0.25).¹⁰¹ Full width at half-maximum (fwhm) was reported
 306 to broaden with increasing Cs content (Figure 5d), which was
 307 correlated with a decrease in grain size (Figure 5e).^{100,118} An
 308 exception is observed in the work by Li et al.¹⁰¹ who observed
 309 enhancement in grain size as a function of Cs addition.
 310 Nevertheless, proposed additional strategies such as the use of
 311 Pb(SCN)₂ additive¹¹⁸ and microstructure-mediated δ→α phase
 312 transformation methods¹¹⁹ in perovskite film formation help
 313 further enlarge grain sizes. The lattice contraction of
 314 Cs_xMA_{1-x}PbI₃ with increasing Cs concentration (Figure 5d)
 315 leads to increase in E_g (Figure 5f) from ~1.50 eV (*x* = 0) up to
 316 ~1.59 eV (*x* = 0.3) resulting in lower J_{sc} (and consequently
 317 diminishing PCEs).^{100,101,118,119}

318 Similar to the α,δ-FAPbI₃ perovskite (Figure 3b), CsPbI₃ also
 319 shows the polymorphism; however, the α→δ phase transition
 320 temperature is much higher (>300 °C) than in the FAPbI₃ case
 321 (Figure 6a). Although the black α-CsPbI₃ has E_g of ~1.67 to
 322 ~1.73 eV (Figure 6f),^{82,155} solar cells based on CsPbI₃ have
 323 exhibited low PCEs (<2.9%),^{115,155,156} which was attributed to
 324 the structural instability, i.e., favorable α→δ phase trans-
 325 formation at room temperature (Figure 6b,d,e).^{155,157} Typical
 326 E_g for the yellow δ-CsPbI₃ is ~2.82 eV (Figure 6f).¹⁵⁵ Despite
 327 these complex behaviors of FAPbI₃, CsPbI₃ and predominance
 328 of δ-phases at room temperature, it is interesting to note that
 329 mixing small quantities of Cs with FA substantially enhances
 330 the stability of Cs_xFA_{1-x}PbI₃.^{100,101,118,119} For example, the δ→
 331 α phase transformation in FAPbI₃ that occurs at ~125–165 °C
 332 (Figure 6a)^{65,82,101,111,121} can be lowered down to room
 333 temperature when Cs/FA ratio of 45 at. % is incorporated in
 334 the Cs_xFA_{1-x}PbI₃.¹⁰¹ This decrease of phase transition
 335 temperature suggests that the stability of black phase α-
 336 FAPbI₃ can be stabilized even at room temperature, which is
 337 important for solar cell applications.¹⁰¹ An explanation for the
 338 enhanced stability of Cs_xMA_{1-x}PbI₃ perovskites was proposed
 339 by Li et al.¹⁰¹ based on the empirical Goldschmidt tolerance
 340 factor (*t*).^{158–160} For a composition of A_xB_{1-x}PbI₃, the effective
 341 tolerance factor (*t*_{effective}) is given by (eq 1). The atomic-ratio
 342 weighted average of two different cations is used as the
 343 estimated effective cation size (*r*_{effective}) as shown in (eq 2).

$$t_{\text{effective}} = \frac{r_{\text{effective}} + r_{\text{I}^-}}{\sqrt{2}(r_{\text{Pb}^{2+}} + r_{\text{I}^-})} \quad (1)$$

$$r_{\text{effective}} = xr_{\text{A}^+} + (1-x)r_{\text{B}^+} \quad (2)$$

346 (Cs/FA)PbI₃ perovskite materials tend to form δ-phase
 347 orthorhombic perovskite structure (e.g., δ₀-CsPbI₃) for *t* < 0.8,
 348 a cubic structure for 0.8 < *t* < 1 (e.g., Cs_xFA_{1-x}PbI₃), and a

hexagonal nonperovskite structure for *t* > 1 (e.g., δ_H-FAPbI₃)
 (Figure 7a).¹⁰¹ Alloying FAPbI₃ with a high *t* value and CsPbI₃
 with a low *t* value, *t*_{effective} can be tuned to be between 0.8 and
 1.0 in Cs_xMA_{1-x}PbI₃ perovskites, which favors a stable
 perovskite structure. This concept was extended to other
 alloy systems to verify its validity (Figure 7b).¹⁰¹ For the
 different alloy compositions reported (see Figure 7b), [a]
 MA_{0.6}FA_{0.4}PbI₃,¹⁰² [b] (FAPbI₃)_{1-x}(MAPbBr₃)_x (*x* = 0.15),⁶³
 [c] MA_{0.13}FA_{0.87}PbI₃,¹⁰⁶ [d] Cs_{0.15}FA_{0.85}PbI₃,¹⁰¹ [e]
 MA_{0.9}Cs_{0.1}PbI₃,¹¹⁵ the best performing devices are for perov-
 skites with *t*_{effective} around 0.94–0.98 (red circle).¹⁰¹ On the
 other hand, poor-photoactive mixed perovskites assemble at the
 low tolerance factor region with *t* < 0.85 (green area).¹⁰¹ In
 summary, Li et al.¹⁰¹ proposed that *t*_{effective} (eqs 1 and 2) can be
 a simpler figure of merit to roughly predict stable structures of
 mixed perovskite alloys. Furthermore, they extracted that *t*_{effective}
 of approximately 0.95 but not exceeding 1 is good for
 maintaining a cubic perovskite structure. However, an *t*_{effective}
 lower than 0.85 would cause too much distortion in the lattice
 leading easily to nonperovskite structures.¹⁰¹ Returning to the
 (MA/FA)PbI₃ system, Dai et al.¹⁵² found that all mixed
 FA_xMA_{1-x}PbI₃ nanorods presented longer photoluminescence
 (PL) lifetimes than the pure FAPbI₃ and MAPbI₃ nanorods,
 (blue squares, Figure 7c). The longer PL lifetimes at the
 composition around FA_{0.4}MA_{0.6}PbI₃ (Figure 7c) was correlated
 with a lower density of defects and consequently resulting in
 high performance.^{102,104,105,107–113} Based on PL measurements,
 Dai et al.¹⁵² identified that the FA/MA molar ratio region in the
 range of 37.5–91% would lead to high performing solar cells
 (gray area, Figure 7c).

2.4. Binary (FA/Rb)PbI₃ System. Recently, Rb⁺ cation,
 which has an even smaller ionic radius of 0.152 nm than Cs⁺
 (Figure 6), has also received attention due to the viability of
 enhancing further both the efficiency and stability of Rb-mixed
 perovskite solar cells.^{121,122,137} Due to the intrinsic non-
 perovskite structure of δ-RbPbI₃ (Figure 6c and 8a,b), it has
 been rarely investigated as a light harvesting material for
 perovskite solar cell.^{121,122,137,157} In fact, RbPbI₃ is known to
 have only the δ-phase (E_g ~ 2.7–3.1 eV)^{121,161–163} and the α-
 phase has not been reported.^{121,122,137,157,164} As shown by
 heating experiments of RbPbI₃ films at different temperatures
 (Figure 8b), RbPbI₃ at 28 °C is yellow; upon heating to 380 °C
 still the yellow color persists in RbPbI₃ while CsPbI₃ (as
 comparison) has transformed to its black α-CsPbI₃. Further
 heating up to 460 °C causes eventually both RbPbI₃ and
 CsPbI₃ to melt, and RbPbI₃ does not show a black phase
 (Figure 8b).^{121,137}

The (FA/Rb)PbI₃ system was studied systematically and
 independently by Park et al.¹²¹ and Zhang et al.¹²² Both studies
 report that only a small Rb quantity (*x* ≤ 0.05) can be
 incorporated into the Rb_xFA_{1-x}PbI₃ perovskite lattice, other-
 wise phase segregation will occur. As shown in Figure 8c, the
 formation of the black α-phase Rb_{0.05}FA_{0.95}PbI₃ started at 120
 °C, whereas for FAPbI₃ started at a higher temperature of 140
 °C and completely darkened at 150 °C. XRD patterns of
 Rb_{0.05}FA_{0.95}PbI₃ (Figure 8d) and FAPbI₃ (Figure 8e) confirms a
 slightly lower transition temperature for the Rb_{0.05}FA_{0.95}PbI₃
 compared to FAPbI₃.^{121,122} Further differential scanning
 calorimetry (DSC) results also confirmed a temperature
 difference of ~10 °C when probing the δ→α phase transitions
 of Rb_{0.05}FA_{0.95}PbI₃ and FAPbI₃.¹²¹ Park et al.¹²¹ described that
 the area under the transition peak in DSC curve is indicative of
 enthalpy of transition (Δ*H*). The area of the peak

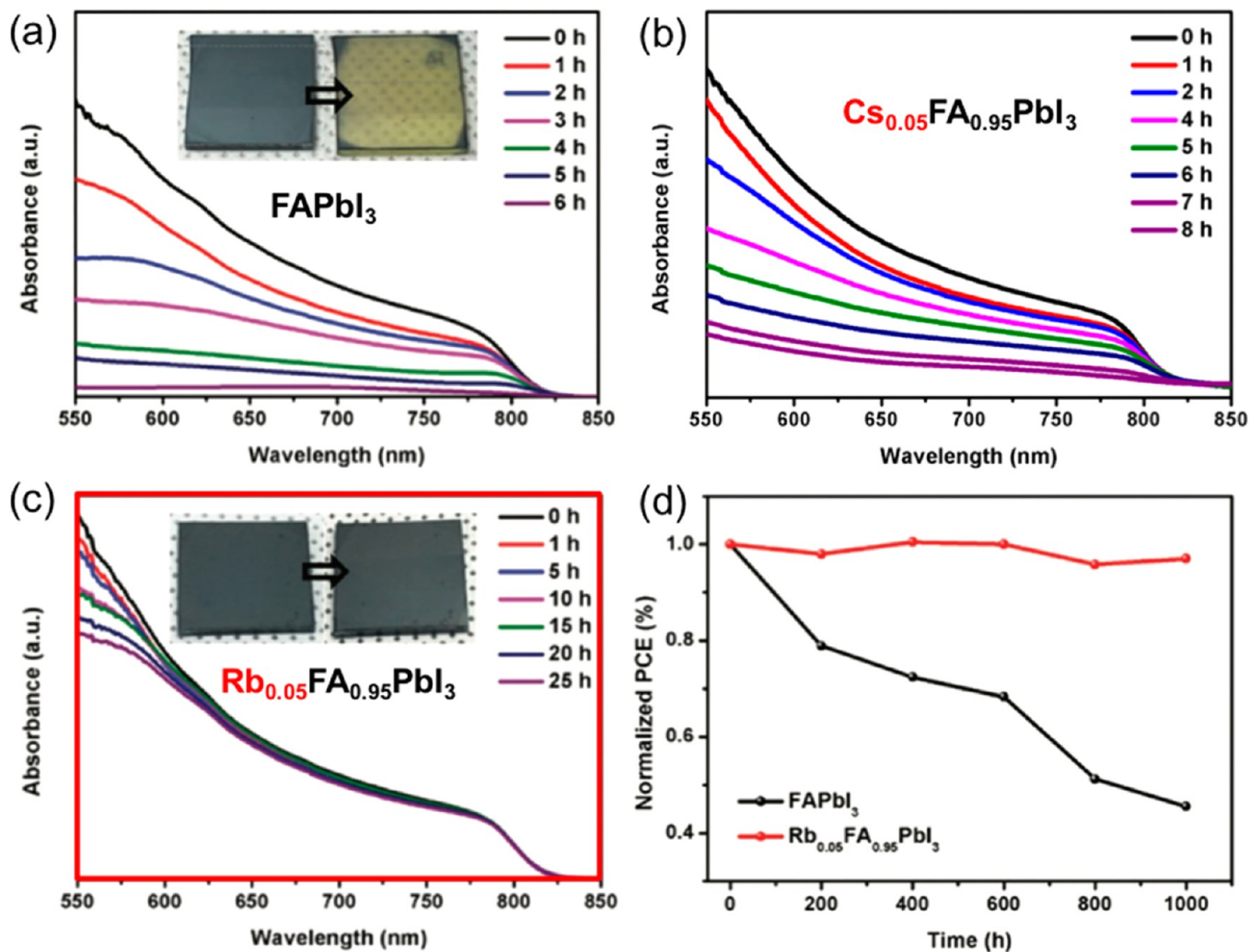


Figure 9. Stability tests (85% RH, 25 °C, and dark) of (a) FAPbI₃, (b) Cs_{0.05}FA_{0.95}PbI₃, and (c) Rb_{0.05}FA_{0.95}PbI₃ perovskite films monitored by the changes in absorbance as a function of time. The enhanced stability of Rb_{0.05}FA_{0.95}PbI₃ films is highlighted from this comparison experiments. (d) Comparison of the stability of FAPbI₃ and Rb_{0.05}FA_{0.95}PbI₃-based unencapsulated devices kept under ambient conditions (average 55% RH) and in the dark. Reprinted with permission from ref 121. Copyright 2017 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim.

412 corresponding to the perovskite $\delta \rightarrow \alpha$ phase transition was
 413 observed to be smaller for Rb_{0.05}FA_{0.95}PbI₃ than FAPbI₃,
 414 indicating that perovskite transition to α -phase is more
 415 favorable when Rb is present in FAPbI₃. It was also reported
 416 that the $\delta \rightarrow \alpha$ phase transformation kinetics of Rb_{0.05}FA_{0.95}PbI₃
 417 is faster (~ 60 s) than FAPbI₃ (~ 4 min).¹²¹

418 Park et al.¹²¹ and Zhang et al.¹²² showed that perovskite
 419 devices based on Rb_{0.05}FA_{0.95}PbI₃ when compared to FAPbI₃
 420 devices, outperformed in PCE from 11.1% \rightarrow 16.15% (on
 421 average) and 14.9% \rightarrow 16.2%, respectively. More importantly,
 422 the superior enhanced stability against moisture of Rb-mixed
 423 perovskites (Figure 9) was highlighted in these two
 424 works.^{121,122} Figure 9a,b,c shows the absorbance change with
 425 time for FAPbI₃, Cs_{0.05}FAPbI₃, and Rb_{0.05}FA_{0.95}PbI₃ films kept
 426 in a sealed box at 85% RH and 25 °C.¹²¹ In the case of FAPbI₃,
 427 the film degraded after 6 h indicated by the formation of δ -
 428 phase and/or PbI₂ (Figure 9a). Slightly better stability can be
 429 inferred when Cs is incorporated forming Cs_{0.05}FAPbI₃ (Figure
 430 9b). However, no obvious changes in the absorbance onset
 431 (~ 800 nm) can be observed for Rb_{0.05}FA_{0.95}PbI₃ films even
 432 after 25 h air exposure (Figure 9c). Cs⁺ and Rb⁺ cations have
 433 chemically and electrostatically similar properties if not equal

because they belong to the same alkali metal group. The
 434 stability properties in Cs_{0.05}FAPbI₃, and Rb_{0.05}FA_{0.95}PbI₃
 435 perovskites are mainly dictated by the small differences in the
 436 ionic radius for hexacoordinated Cs⁺ (0.167 nm) and Rb⁺
 437 (0.152 nm). This ionic radius difference is not significant, and it
 438 is remarkable the enhanced stability of Rb_{0.05}FA_{0.95}PbI₃
 439 compared to that of Cs_{0.05}FA_{0.95}PbI₃ (Figure 9b,c). Park et al.
 440 et al.¹²¹ and Zhang et al.¹²² also compared the long-term stabilities
 441 (~ 1 month) of complete devices (FTO/c-TiO₂/mp-TiO₂/
 442 Rb_{0.05}FA_{0.95}PbI₃/Spiro-MeOTAD/Au) evaluating under ambi-
 443 ent air storage conditions without any encapsulation. 444
 445 Rb_{0.05}FA_{0.95}PbI₃-based solar cells maintained a PCE $>97\%$ ¹²¹
 446 and $>90\%$ ¹²² of their initial efficiencies corresponding to the
 447 fresh devices.

448 Yi et al.¹¹⁷ proposed that the stability of Cs_xFA_{1-x}PbI₃
 449 perovskites could be rationalized based on the structural and
 450 thermodynamics arguments of internal energy variation (ΔE)
 451 and entropic gains (ΔS , mixing entropy concept) (Figure 10a).
 452 They calculated the free energies, $\Delta F = \Delta E - T \cdot \Delta S$ (black
 453 squares, Figure 10a), of Cs_xFA_{1-x}PbI₃ with δ -, β -, α -phases
 454 considering different Cs contents ($0 \leq x \leq 1$). A range in Cs
 455 composition was identified to lead to $\Delta F < 0$ (black squares, 455

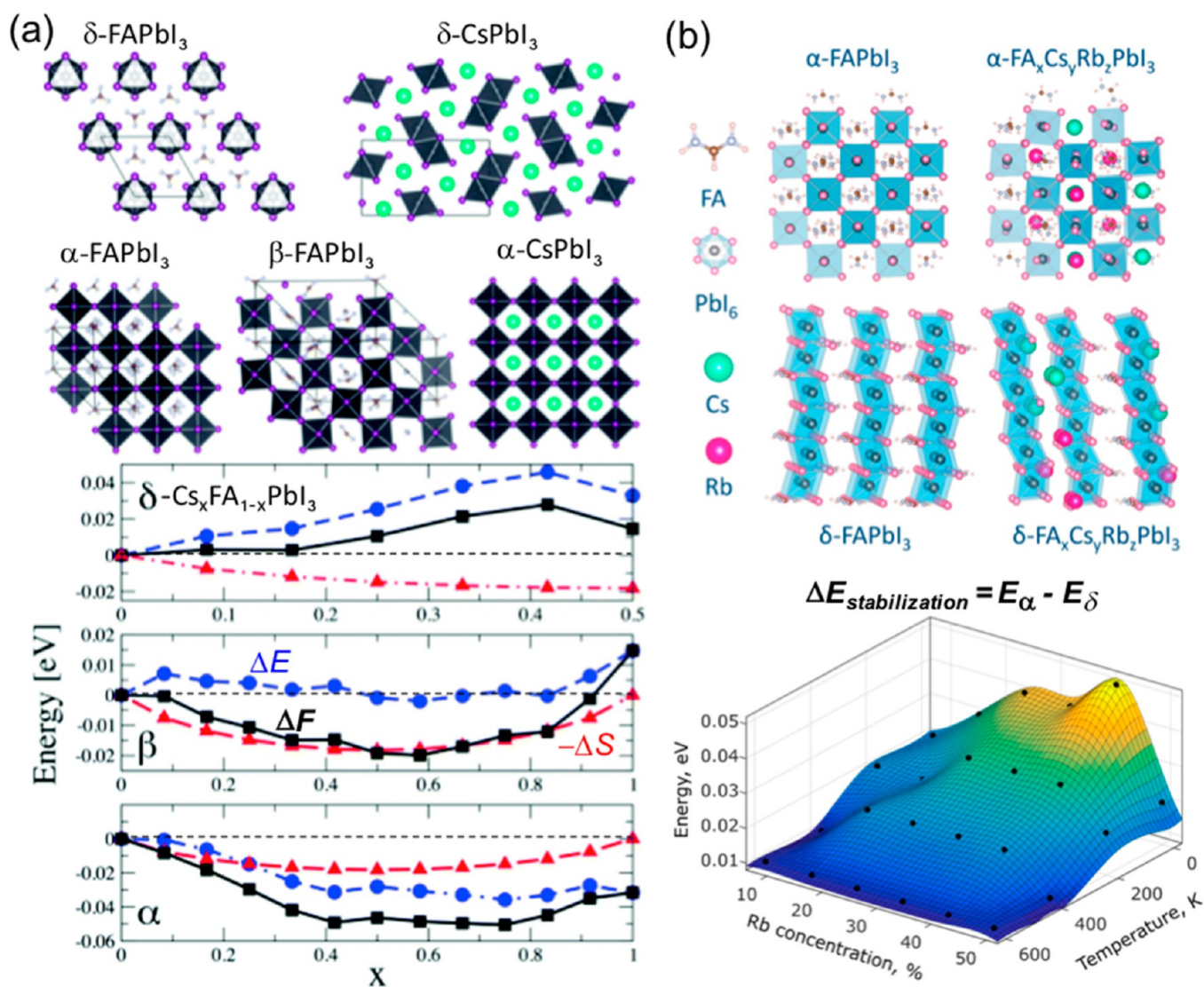


Figure 10. (a) Crystal structure of various polymorphs of FAPbI₃ and CsPbI₃. The stability of Cs_xFA_{1-x}PbI₃ was rationalized based on structural and thermodynamic arguments. Variation of internal energy ΔE (blue circles), mixing entropy contribution $-\Delta S$ (red triangles), and free energy $\Delta F = \Delta E - T \cdot \Delta S$ (black squares). $\Delta F < 0$ (below the dashed lines) signifies stability is observed. Reprinted with permission from ref 117. Copyright 2016 The Royal Society of Chemistry. (b) Structures of pure FAPbI₃ and mixed FA_xCs_γRb₂PbI₃ in α - and δ -phases. Reprinted with permission from ref 157. Copyright 2017 American Chemical Society.

456 Figure 10a), which signifies stable configuration. The
 457 stabilization energy was calculated to be in the order of 0.05
 458 eV ($\sim 2k_B T$ at room temperature) and 0.02 eV ($\sim 0.8k_B T$) per
 459 stoichiometric unit for α - and β -phases, respectively. Based on
 460 this model, the $\delta \rightarrow \alpha$ or β transition temperature is reduced by
 461 ~ 200 – 300 K when going from pure FAPbI₃ to the mixed
 462 Cs_xFA_{1-x}PbI₃ perovskites. In a follow-up work, Syzgantseva et
 463 al.¹⁵⁷ have recently performed additional computational study
 464 comparing the influences of MA⁺, Cs⁺, Rb⁺ cation substitutions
 465 on the stabilization energies of FAPbI₃ perovskite (Figure 10b)
 466 based on the same arguments described by Yi et al.¹¹⁷ The
 467 relative stabilization energies ($\Delta E_{\text{stabilization}} = E_{\alpha} - E_{\delta}$, Figure
 468 10b) of α - (E_{α}) with respect to δ -phase (E_{δ}) were determined
 469 using simulations of a series of α and δ structures, in which FA⁺
 470 cations were successively substituted by MA⁺, Cs⁺, and Rb⁺. Cs⁺
 471 and Rb⁺ were shown to be more efficient in stabilizing the
 472 perovskite than MA⁺ based on the balance of ΔE and ΔS .^{117,157}
 473 The introduction of Cs⁺ and/or Rb⁺ thermodynamically favors

the formation of new perovskite phases, bringing the system
 into a new equilibrium state.¹⁵⁷

To achieve a systematic understanding of the influences of
 mixing halides in the ABX₃ structure, in the next section we will
 first describe simpler systems constraining the A cation to
 single elements (MA⁺, FA⁺, or Cs⁺) and allowing to vary halide
 part (I/Br, I/Cl, or I/Br/Cl) in the X site.

3. MIXED X HALIDE ANIONS

A handful number of reviews described and summarized the
 new optoelectronic properties of mixed perovskites when
 mixed halide ions are incorporated in perovskites.^{5,7,8,14,15,24,26}
 Below we present the current views of the structure–property
 relationship based on a survey of recently published papers
 summarized for MAPb(I/Cl), FAPb(I/Cl), MAPb(I/Br),
 FAPb(I/Br), MAPb(Br/Cl), CsPb(I/Br), CsPb(Br/Cl), CsPb-
 (I/Cl), MAPb(I/Br/Cl).

3.1. Binary MAPb(I/Cl) and FAPb(I/Cl) Systems. The
 MAPb(I/Cl) material system is by far one of the most studied

Table 2. Summary of Experimental Conditions for MAPb_{1-3-x}Cl_x Synthesis Methods and Measurement Conditions and Measurement Conditions Extracted from References 60, 165–199, 209^a

ref. (Year)	Sample structure	MAPb _{1-3-x} Cl _x fabrication method	MAPb _{1-3-x} Cl _x fabrication environmental conditions	Experimental techniques	Measurement environmental conditions	Cl detected?	Main conclusions
165 (2017)	ITO/PEDOT:PSS/PVSK	(1) 40 wt % dissolving MAI and PbCl ₂ in 3:1 molar ratio in DMF. (2) Postannealing at 100 °C in a N ₂ glovebox from 10 to 100 min. (3) ~250 thickness.	N ₂ glovebox	TOF-SIMS XPS GIXRD	Vacuum Vacuum N ₂ flux	Yes	(1) MAPbCl ₃ forms as intermediate phase during crystallization. (2) After completion of PVSK crystallization some Cl detected by TOF-SIMS. (3) Cl is majorly located at the interface and within the underlying PEDOT:PSS. Atomic ratio of Cl:I ~ 3:97 determined by XPS.
166 (2016)	FTO/c-TiO ₂ /PVSK	Redissolution and crystal grain growth via spray coating: (1) MAI and PbCl ₂ in 3:1 molar ratio dissolved in IPA. (2) Centrifuge. (3) MAPb _{1-3-x} Cl _x redissolved in DMF:GBL.	Ambient	XPS	Vacuum	Yes	(1) EDX analysis showed compositions of MAPbCl _{0.27} MAPbCl _{0.35} MAPbCl _{0.45} MAPbCl _{0.75} MAPbCl _{0.25} . (2) 3D crystal structure for MAPb _{1-3-x} Cl _x proposed. (1) EDX showed Cl:I ~ 0.61. (2) XRD showed MAPbCl ₃ phase. (3) Assuming 4% Cl incorporation in MAPb _{1-3-x} Cl _x lattice, the ratio of MAPb _{1-3-x} Cl _x and MAPbI ₃ phases will be ~65:35. This is close to expected reaction: 3MAI + 3PbI ₂ → 2MAPbI ₃ + MAPbCl ₃ .
167 (2016)	Glass slide	Single-crystalline nanofibers; Stock solutions of (MAI+PbCl ₂ /DMF, 0.03 M) and (MAI+PbI ₂ /DMF, 0.03 M) at different volume ratios.	—	EDX, XRD, UV-vis, PL	—	Yes	(1) EDX analysis showed compositions of MAPbCl _{0.27} MAPbCl _{0.35} MAPbCl _{0.45} MAPbCl _{0.75} MAPbCl _{0.25} . (2) 3D crystal structure for MAPb _{1-3-x} Cl _x proposed. (1) EDX showed Cl:I ~ 0.61. (2) XRD showed MAPbCl ₃ phase. (3) Assuming 4% Cl incorporation in MAPb _{1-3-x} Cl _x lattice, the ratio of MAPb _{1-3-x} Cl _x and MAPbI ₃ phases will be ~65:35. This is close to expected reaction: 3MAI + 3PbI ₂ → 2MAPbI ₃ + MAPbCl ₃ .
168 (2016)	FTO/c-TiO ₂ /mp-TiO ₂ /PVSK	Vapor-assisted solution deposition: 1 M PbI ₂ in DMF spin coated and substrate heated at 110 °C (15 min); PbI ₂ film placed 2 mm above MAI source in an oven; Postannealing at 70 °C (30 min).	—	EDX, XRD	—	Yes	(1) EDX analysis showed compositions of MAPbCl _{0.27} MAPbCl _{0.35} MAPbCl _{0.45} MAPbCl _{0.75} MAPbCl _{0.25} . (2) 3D crystal structure for MAPb _{1-3-x} Cl _x proposed. (1) EDX showed Cl:I ~ 0.61. (2) XRD showed MAPbCl ₃ phase. (3) Assuming 4% Cl incorporation in MAPb _{1-3-x} Cl _x lattice, the ratio of MAPb _{1-3-x} Cl _x and MAPbI ₃ phases will be ~65:35. This is close to expected reaction: 3MAI + 3PbI ₂ → 2MAPbI ₃ + MAPbCl ₃ .
169 (2015)	FTO/c-TiO ₂ /PVSK (and) SIN/PVSK	(1) MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF. (2) Postannealing at 95 °C for 0, 15, 60, 90, 120 min in dry air (3 ppm of H ₂ O).	—	XANES	He atmosphere	Yes	(1) $x = 0.05 \pm 0.03$ Cl atoms per formula unit (i.e., Pb atom). (2) Cl may be incorporated in the film directly (e.g., substitution or interstitial) or indirectly (e.g., grain boundary or poor crystalline region).
170 (2015)	ZnS optical prism/PEDOT:PSS/PVSK	Multicycle spin-coating process: PbI ₂ in DMF spin-coated on substrate; MAI/MAI (4:1 molar ratio) in IPA spin-coated on PbI ₂ film; Postannealing at different temperatures.	—	PTIR	Under N ₂ flow	Yes	(1) Mild annealing (60 °C, 1 h) leads to a MAPb _{1-3-x} Cl _x film with Cl-rich ($x < 0.3$) and Cl-poor phases. (2) Further annealing (110 °C) leads to a homogeneous Cl-poorer ($x < 0.06$) phase.
171 (2015)	Quartz	Vacuum coevaporation of MAI and PbCl ₂ ; Postannealing at 70, 80, 90, 100, 110, 120 °C (1 h) in N ₂ glovebox; ~530 nm thickness.	Vacuum	XRD, EDX, UV-vis, PL	—	Yes	(1) As-grown MAPb _{1-3-x} Cl _x showed 0.13 at. % Cl. (2) Further annealing (110 °C) leads to a homogeneous Cl-poorer ($x < 0.06$) phase.
172 (2015)	c-TiO ₂ /PVSK	(1) MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF. (2) Postannealing at 120 °C for 45 min.	—	Ion chromatography	Films dissolved in DI water and H ₂ SO ₄ (0.15 M)	Yes	(1) MAPb _{1-3-x} Cl _{0.06} composition determined.
173 (2015)	ITO/PEDOT:PSS/PVSK	Two-step method: (1) 1.5 M PbCl ₂ /PbI ₂ dissolved in DMF and spin coated. (2) substrate dipped in MAI dissolved in IPA at 70 °C.	Ambient air with 20 °C and 20% RH.	Potentiometric titration method, XRD, UV-vis	Ultrapure water	Yes	(1) Measured $n_{Cl}/(n_{Cl} + n_I)$ mole fraction content in PVSK according to (PbCl ₂ /PbI ₂): 0.012 ± 0.008 (0.10); 0.073 ± 0.012 (0.20); 0.185 ± 0.015 (0.30); 0.220 ± 0.015 (0.40). (2) Band gap varies from 1.54 to 1.59 eV with increasing Cl. (3) XRD shows shifts in (110) and (220) peaks.

Table 2. continued

ref. (Year)	Sample structure	MAPb _{1-x} Cl _x fabrication method	MAPb _{1-x} Cl _x fabrication environmental conditions	Experimental techniques	Measurement environmental conditions	Cl detected?	Main conclusions
174 (2015)	FTO/c-TiO ₂ /PVSK	One-step method. (1) MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF. (2) Postannealing at 90 °C (2.5 h) and then 120 °C (15 min).	N ₂ glovebox	HAXPES, FY-PES	Vacuum	Yes	$n_{Cl}/(n_{Cl} + n_I)$ mole fraction of 0.056 ± 0.015 determined. (1) Cl is depleted from top surface. (2) Upper limit on the amount of Cl in MAPb _{1-x} Cl _x : $x < 0.07$ and $x < 0.40$ to depths of ~10 and ~26 nm, respectively. (3) High Cl concentration ($x > 0.40$) deep in the film and near PVSK/TiO ₂ interface.
175 (2015)	FTO/c-TiO ₂ /PVSK	Two-step method: (1) PbI ₂ in DMF spin coated. MAI/MACl (10:1 in weight) dissolved in IPA spin coated on top of dry PbI ₂ film. (2) PbI ₂ in DMF spin coated. MACl dissolved in IPA spin coated on top of dry PbI ₂ film. Both films annealed in air at 135 °C.	Dry air (Dew Point of -70 °C)	XRD, XPS	–	Yes	(1) Cl incorporation improve carrier transport across heterojunction interfaces, rather than within perovskite crystal. (2) XRD showed ~0.5% difference of the unit cell volume. (3) Very weak Cl 2p signal in XPS; Cl has negligible impact on the original crystal structure.
176 (2014)	Glass/PVSK (for optical and structural determination) (and) Si/TiO ₂ (for GIWAXS)	(1) MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF. (2) Postannealing at 95 and 100 °C for various times.	PVSK preparation in glovebox. Post annealing in air and N ₂ for comparison.	GIWAXS, in situ XRD, XPS, XRF, EDX, UV-vis	In situ XRD in ambient; XPS and EDX in vacuum; UV-vis in glovebox	Yes	(1) MAPb _{1-x} Cl _x with maximum of $x = 0.15$ determined based on lattice volume. (2) Cl content of samples annealed at 95 °C for 120 min, varies from 0 to 0.3 Cl per formula unit determined by XRF. (3) XPS and EDX did not show Cl. (4) Small Cl amount incorporated in the crystal lattice; Most of Cl resides at grain boundary or interface with substrate. Larger band-bending at grain boundaries with Cl incorporation.
177 (2015)	FTO/c-TiO ₂ /PVSK	Two-step method: (1) 1 M PbI ₂ in DMF. (2) Substrate dipping into MACl + MAI dissolved in IPA. Layer-by-layer growth: PbCl ₂ thermally evaporated in vacuum; Dipping in MAI solution (IPA). Similarly a second PbCl ₂ thermally evaporated on MAPb _{1-x} Cl _x and dipped into MAI solution. These cycles are repeated until desired film thickness is attained.	–	KPFM	–	Yes	(1) Cl atomic ratio was 8.3%. The ratio above is much higher than that of solution-cast counterpart (1.4%). Chloride located at TiO ₂ /PVSK
178 (2015)	ITO/PEDOT:PSS/PVSK	(1) 1.5 wt % dissolving MAI and PbCl ₂ in 3:1 molar ratio in DMF. (2) Postannealing at 100 °C for 1 h. (3) ~15 nm thickness.	N ₂ atmosphere	XPS	Vacuum	Yes	(1) There exists a miscibility gap for MAPbI ₃ and MAPbCl ₃ in the 0.5 < x < 0.95 regime for MAPbI ₃ Cl _{1-x} . (2) Cl incorporation stabilizes the cubic phase of MAPbI ₃ at room temperature. (1) Cl content in the film is miniscule, but detected by XPS (~1 at %) and TGA (~1.3 at %). (2) EDX did not detect Cl. (3) MAPbI _{3-x} Cl _x shows tetragonal <i>I4/mcm</i> structure.
179 (2014)	FTO/c-TiO ₂ /PVSK	Vacuum coevaporation of MAI and PbCl ₂ .	Vacuum	In situ XRD, EDX	Vacuum	Yes	
180 (2014)	Soda lime glass/PVSK	(1) 0.88 M PbI ₂ + 2.64 M MAI. (2) 0.88 M PbI ₂ + 0.88 M MAI + 0.88 M MACl. (3) Films were dried at 70 °C (10 min) and annealed at 100 °C (45 min).	Glove box	TGA, XPS, EDX	TGA in N ₂ flow (40 mL/min) and ~10 mg of PVSK; XPS and EDX in vacuum	Yes	
181 (2015)	FTO/c-TiO ₂ /PVSK						

Table 2. continued

ref. (Year)	Sample structure	MAPbI _{3-x} Cl _x fabrication method	MAPbI _{3-x} Cl _x fabrication environmental conditions	Experimental techniques	Measurement environmental conditions	Cl detected?	Main conclusions
182 (2014)	FTO/PVSK	MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF; Postannealing at 100 °C (50 min).	N ₂ glovebox	XPS depth profile (Ar sputtering), EDX	Vacuum	Yes	(1) Very weak Cl signal in XPS of ~1% Cl detected for the top 20 nm. (2) EDX probed Cl content of 0.7 at. %. There is a note that detection limit of EDX setup is ~1 at. %.
183 (2014)	ITO/PEDOT:PSS/PVSK	MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF; Postannealing at 90 °C (2 h).	–	XPS	Vacuum	Yes	(1) Cl/(Cl + I) ~ 2.2% detected by XPS. (2) XRD shows similar diffraction pattern to MAPbI ₃ ; therefore, only small fraction of Cl can be incorporated.
184 (2014)	ITO/TiO ₂ /PVSK	22.5–60 wt % MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMSO; Postannealing at 100 °C (50 min).	N ₂ glovebox (<1 ppm of O ₂ and H ₂ O)	XPS	Vacuum	Yes	XPS on the 60 wt % precursor concentration, showed atomic composition of C:N:Pb:I:Cl = 1.04:1.1:0.5:2.02:0.99.
185 (2013)	Glass slide (and) FTO/c-TiO ₂ /mp-TiO ₂ /PVSK	(1) 20 wt % MAI and PbCl ₂ (3:1 molar ratio) dissolved in DMF; Postannealing at 100 °C (45 min). (2) 20 wt % MAI and PbCl ₂ (1:1 molar ratio) dissolved in DMF; Postannealing at 100 °C (10 min).	Spin coating in N ₂ flux.	EDX	Vacuum	Yes	(1) Reliable values of the Cl content not achieved as the amplitude of Cl-related signal was observed to decrease during the measurement. (2) Based on DFT, Cl incorporation is allowed at low concentrations (below 3–4%) in MAPbI ₃ .
60 (2012)	FTO/c-TiO ₂ /np-TiO ₂ (or Al ₂ O ₃)/PVSK	MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF; Postannealing at 100 °C for 45 min.	Ambient air	EDX	Vacuum	Yes	MAPbI ₃ Cl had I/Cl ratio ~2:1.
186 (2017)	FTO/c-TiO ₂ /PC61BM/PVSK	1 M PbCl ₂ :PbI ₂ (molar ratios of 0:1, 1:3, 11: 3:1)	N ₂ glovebox (RH < 10%)	XRD, XPS	–	No	(1) Cl can be incorporated into precursor film in the form of PbCl or PbCl ₂ and it was found in the final PVSK film in MAPbCl ₃ , PbI ₂ or PbCl ₂ . (2) No evidence for MAPbI _{3-x} Cl _x phase observed. (3) MAPbI _{3-x} Cl _x phase is relatively unstable or possess higher formation energy.
187 (2017)	FTO/c-TiO ₂ /PVSK	MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF; Postannealing at 100 °C (110 min).	Glove box	XRD, UV-vis, XPS	–	No	XPS shows negligible amount of Cl in fully annealed films.
188 (2016)	Glass	PbI ₂ and PbCl ₂ dissolved in DMSO spin coated on glass and loaded in XRD stage. MAI powder dispersed around the film and sample stage controllably heated.	He atmosphere	in situ XRD, TOF-SIMS	In situ XRD in He	No	(1) In the presence of reactive I ⁻ ions, neither Br ⁻ nor Cl ⁻ can be incorporated into the perovskite crystal lattice. (2) TOF-SIMS mapping revealed that Cl ⁻ resides in the grain boundaries, possibly in the form of amorphous Cl ⁻ -based compounds.
189 (2016)	FTO/c-TiO ₂ /PVSK	MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF; Postannealing at 90 °C (2 h), 100 °C (10 min), 130 °C (5 min).	–	in situ XRD	N ₂ atmosphere	No	(1) Initially, the film crystallizes in MAPbCl ₃ and is fully converted to MAPbI ₃ after a certain time under heating. (1) Below detection limits of XPS.
209 (2016)	MAPbI(Cl) single crystal	Supporting info in ref209.	–	XRD, XPS, TOF-SIMS	–	No	(2) Trace amounts of Cl detected by TOF-SIMS.
190 (2015)	ITO	(1) Sequential vacuum evaporation of MAI (15 nm) and then PbCl ₂ . (2) coevaporation.	Vacuum	In situ XPS	Vacuum	No	(1) No Cl detected at the initial PbCl ₂ deposition (0.2–0.5 nm); Below detection limit of XPS (~0.1 at. %) (2) As grown and annealed (100 °C, 1 h) MAI:PbI ₂ (3:1) films 15 nm did not show Cl in XPS.
191 (2015)	FTO/c-TiO ₂ /PVSK	(1) PbI ₂ dissolved in DMF and spin coated. (2) CYD: First reacted with MAI (120 °C, 5 min) and then with MAI at 145 °C, 100 min.	–	XRD, XPS, EDX	XPS and EDX in vacuum.	No	Cl content not detectable in XPS and EDX.

Table 2. continued

ref. (Year)	Sample structure	MAPb _{1-x} Cl _x fabrication method	MAPb _{1-x} Cl _x fabrication environmental conditions	Experimental techniques	Measurement environmental conditions	Cl detected?	Main conclusions
192 (2015)	ITO/PEDOT:PSS/PVSK	MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF; Postannealing at 60 °C (2 h).	Postannealing in vacuum and glovebox	XPS	Vacuum	No	Cl content of 1.36% at. % assigned as MACl residue.
193 (2014)	FTO/c-TiO ₂ /PVSK	(1) Two-step: PbI ₂ ; MAI/MACl (95:5 wt %) in IPA. (2) Two-step: PbCl ₂ ; MAI in IPA. (3) One-step: PbCl ₂ :MAI = 1:3 Several (Table S2 in ref 194) dissolved in DMF; Postannealing at 90 °C for 2–3 h.	–	STEM-EDS	Vacuum	No	No Cl feature detected within the detection limit of EDS (~1000 ppm).
194 (2014)	ITO/PEDOT:PSS/PVSK	Several (Table S2 in ref 194) dissolved in DMF; Postannealing at 90 °C for 2–3 h.	N ₂ glovebox	TEM, SAED, EDX, XPS	Vacuum	No (below detection limit)	(1) EDX showed I:Pb ~3 at. % ratio. (2) Cl was below detection limit of EDX (~0.1 at. %). (3) Presence of Cl impacts on the microstructure and orientation of PVSK films. (4) SAED patterns indexed to tetragonal MAPbI ₃ .
195 (2014)	FTO/c-TiO ₂ /PVSK	MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF; Postannealing at 120 °C (30–60 min).	Ambient atmosphere	EDX, XPS	Vacuum	No	(1) No signs of Cl in XPS. (2) EDX showed no Cl (or is below detection limit, <1% mol). Material reported MAPb _{1-x} Cl _x is a combination of MAPbI ₃ and MAPbCl ₃ phases.
196 (2014)	Glass/mp-TiO ₂ /PVSK	MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF and DMSO mixed in 7:3 by volume; Postannealing at 140 °C (30 min) in dry air box (RH < 10%); Samples dried in vacuum (5 × 10 ⁻³ Torr, 1 h)	Dry air box (RH < 10%)	XRD	Room temperature in air; RH 30–50%; 4 h measurement.	No	Longer annealing (45 min) leads to complete loss of Cl (within EDX detection limit of 1%).
197 (2014)	FTO/c-TiO ₂ /mp-TiO ₂ /PVSK	PbI ₂ , MAI, MACl dissolved in DMF with molar ratio of 1:1:1 (x varies from 0 to 2; max dissolvable x is ~2.8); Postannealing at 100 °C (5–45 min).	Ambient air	EDX	Vacuum	No	Cl was below detection limit of EELS, EDX, XPS.
198 (2014)	FTO/c-TiO ₂ /PVSK	Two-step: PbI ₂ in DMF; Spin coating (~200 nm film thickness); MAI and MACl dissolved in IPA. Substrate immersed.	–	EELS, EDX, XPS	Vacuum	No	Cl sublimes as MACl leaving only MAPbI ₃ as observed in XRD.
199 (2014)	FTO/mp-TiO ₂ /PVSK	MAI and PbCl ₂ in 3:1 molar ratio dissolved in DMF; Postannealing at different temperatures (60–200 °C) and times (20 to 0.17 h)	Ambient air	XRD	Ambient	No	

“Abbreviations: See Table 1. PVSK = “MAPb_{1-x}Cl_x”; DMF = N,N-dimethylformamide; GBL = γ -butyrolactone; IPA = 2-propanol; XPS = X-ray photoelectron spectroscopy; AR-XPS = angle-resolved XPS; hard XPS (HAXPES); fluorescence yield X-ray absorption spectroscopy (FY-XAS); TOF-SIMS = time-of-flight secondary ion mass spectrometry; GIXRD = grazing-incidence X-ray diffraction; XANES = X-ray absorption near edge structure; EDS/EDX = energy-dispersive X-ray spectroscopy; XRF = X-ray fluorescence spectroscopy; GIWAXS = grazing-incidence wide-angle X-ray scattering; TEM = transmission electron microscopy, SAED = select area electron diffraction, KPFM = Kelvin probe force microscopy, PTIR = photothermal-induced resonance; TGA = thermogravimetric analysis, PL = photoluminescence, STEM-EDS = scanning transmission electron microscopy EDS. EELS = electron energy loss spectroscopy.

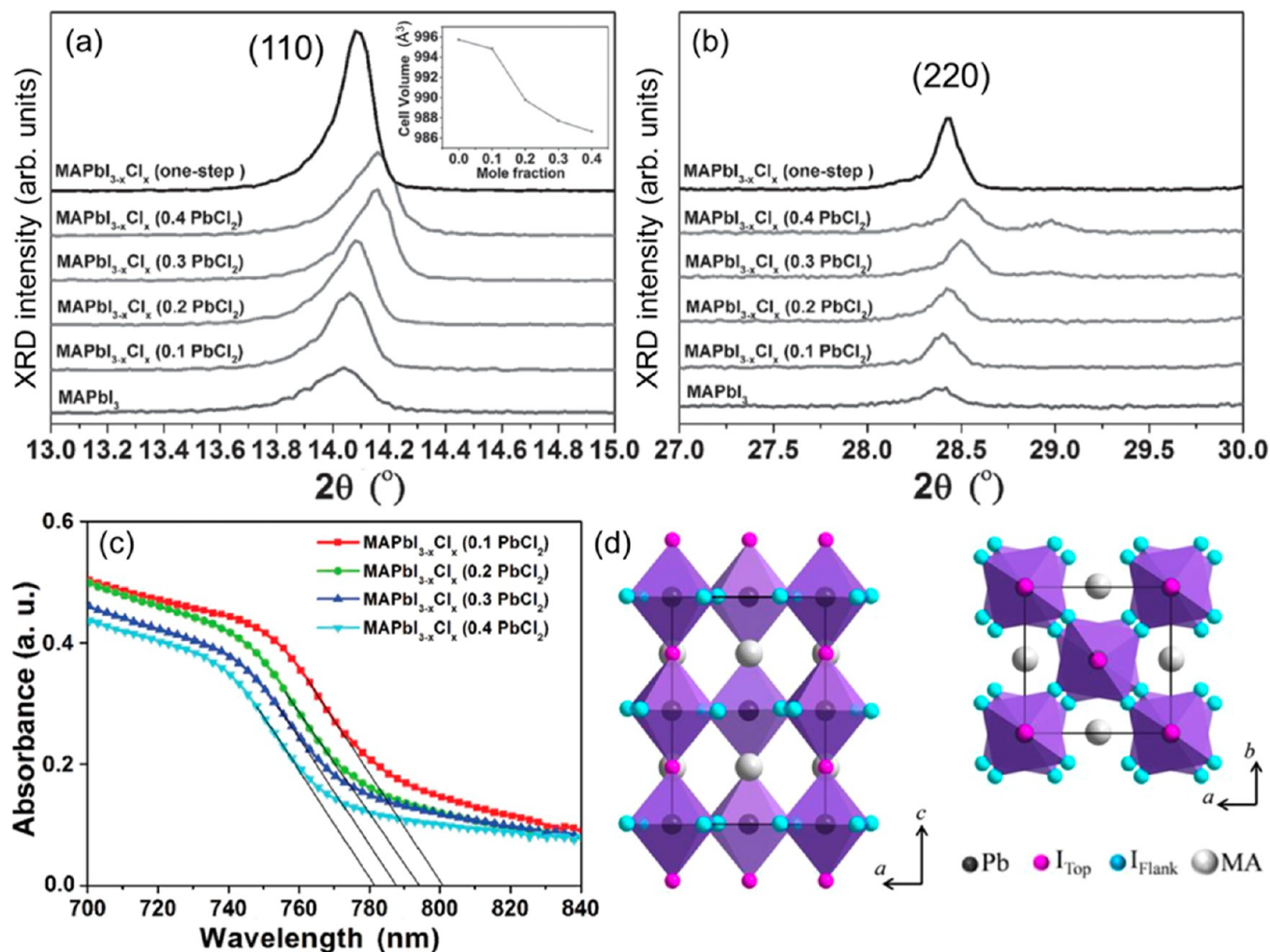


Figure 11. XRD pattern of (a) (110) and (b) (220) lattice planes of $\text{MAPbI}_{3-x}\text{Cl}_x$ perovskites with different PbCl_2 mole fractions by a two-step method. $\text{MAPbI}_{3-x}\text{Cl}_x$ perovskites formed by one-step method and pure-phase MAPbI_3 are shown as comparison. (c) UV-vis of $\text{MAPbI}_{3-x}\text{Cl}_x$ perovskites with different PbCl_2 mole fractions by a two-step method. Reprinted with permission from ref 173. Copyright 2017 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim. (d) Proposed schematic representation of tetragonal crystal structure of $\text{MAPbI}_{3-x}\text{Cl}_x$ where Cl atoms partly substituted I at I_{flank} (blue) positions forming different contents of x . Reprinted with permission from ref 167. Copyright 2017 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim.

491 binary mixed perovskite.^{60,61,165–200} From earlier re-
 492 ports,^{2,3,21,47,201,202} PbI_2 and MAI precursors are employed
 493 for the synthesis of MAPbI_3 . Furthermore, if one-step method
 494 is employed, generally equimolar ratio of PbI_2 :MAI (i.e., 1:1) is
 495 used. Lee et al.⁶⁰ reported the first deposition approach
 496 employing PbCl_2 and 3-fold excess of MAI (i.e., 1:3). The
 497 dissolution of PbCl_2 in N,N -dimethylformamide (DMF) is
 498 difficult when the MAI: PbCl_2 molar ratio is lower than
 499 3:1.^{194,195} Although Cl incorporation has been reported to
 500 improve optoelectronic properties (e.g., diffusion lengths for
 501 electron and holes of ~ 130 and ~ 90 – 105 nm in MAPbI_3
 502 increases by 10 times in $\text{MAPbI}_{3-x}\text{Cl}_x$, ~ 1069 and ~ 1213 nm,
 503 respectively),^{200,203} the controversial question whether Cl can
 504 in fact be incorporated into the crystal lattice of MAPbI_3 is still
 505 not completely settled. Reviews on this topic can be found in
 506 refs^{8,14,15,169,204}. The reported studies show contradicting results
 507 regarding detection or absence of Cl in $\text{MAPbI}_{3-x}\text{Cl}_x$ perovskite
 508 (Table 2). From a total of thirty seven studies surveyed by us
 509 (Table 2), we find that twenty two confirmed the presence of
 510 Cl in $\text{MAPbI}_{3-x}\text{Cl}_x$ based on an extensive number of analytical
 511 tools such as X-ray photoelectron spectroscopy

(XPS),^{166,175,178,181,183,184} angle-resolved XPS (AR-XPS),¹⁷⁹
 XPS depth-profile,¹⁸² hard XPS (HAXPES),¹⁷⁴ fluorescence
 yield X-ray absorption spectroscopy (FY-XAS),¹⁷⁴ time-of-flight
 secondary ion mass spectrometry (TOF-SIMS), in situ XRD,¹⁸⁰
 glazing-incidence XRD (GIXRD),¹⁶⁵ X-ray absorption near
 edge structure (XANES),¹⁶⁹ energy-dispersive X-ray spectroscopy
 (EDX or EDS),^{167,168,171,182,185} X-ray fluorescence
 spectroscopy (XRF),¹⁷⁶ grazing-incidence wide-angle X-ray
 scattering (GIWAXS), select area electron diffractogram
 (SAED), photothermal-induced resonance (PTIR),¹⁷⁰ Kelvin
 probe force microscopy (KPFM),¹⁷⁷ thermogravimetric anal-
 ysis (TGA),¹⁸¹ ion chromatography,¹⁷² potentiometric titra-
 tion.¹⁷³ However, there are also at least 15 studies^{186–199}
 arriving at the conclusion that Cl is absent in the final
 perovskite films or is below the detection limit of the
 instruments (e.g., EDX detection limit is 0.1–1 at.
 %^{194,195,197} or ~ 1000 ppm¹⁹³ and XPS detection limit is 0.1
 at. %¹⁹⁰). It has been argued that (1) as a result of the larger
 difference in ionic radii, Cl incorporation yields low miscibility
 with iodine;^{15,176,185} (2) the $\text{MAPbI}_{3-x}\text{Cl}_x$ phase is metastable
 or possess higher formation energies;¹⁸⁶ (3) PbCl_2 is the least

soluble Pb-halide in DMF that may lead to PbCl₂ particles;^{194,195} (4) experimental evidence that Cl were found at the perovskite/substrate interface (i.e., the affinity of Cl to TiO₂^{169,176,179} and PEDOT:PSS¹⁶⁵ is high) and/or grain boundaries;^{169,170,176,177} (5) the postannealing step (100 °C, 45 min) leads to sublimation of Cl in the form of MAcl;^{55,181,182,192,199} (6) MAPbI_{3-x}Cl_x and MAPbI₃ perovskites show very similar UV-vis and XRD patterns indicating no or a very small amount of Cl can be incorporated;^{15,60,186,191,197-199} (7) crystallization (crystallographic texture, crystalline orientation) and morphological (polycrystalline grain structure) improvements instead of Cl-incorporation.^{167,182,194,195,205}

In Table 2, we compare the relevant parameters such as sample conditions (sample preparation method and environment, sample structure, perovskite thicknesses) as well as technical aspects of measurement conditions (vacuum versus ambient pressure and/or under inert gas environment). We find that the different results from these studies may stem from (1) the various sample preparation conditions, for instance, vapor-based techniques (e.g., vacuum codeposition, vapor-assisted solution deposition at ambient, chemical vapor deposition) versus solution methods (one-step and two-step methods; use of MAcl or PbCl₂ as source for Cl) and (2) the influence of excitation probes and environment during measurements. Zhao and Zhu¹⁹⁷ showed that Cl can be incorporated in MAPbI₃ employing MAcl precursor by controlling the annealing time (~20 min). On the contrary, Chen et al.¹⁷⁵ reported that it is difficult to incorporate Cl using the MAcl precursor. PbCl₂ was proposed to be a better way to introduce Cl. Li et al.¹⁷³ employed a two-step dipping method for the fabrication of MAPbI_{3-x}Cl_x films, which were deposited on PEDOT:PSS. The PbCl₂:PbI₂ with various mole fractions were dissolved at 85 °C in DMF and spin coated on PEDOT:PSS films (after the solutions were cooled to room temperature). The substrates were then dipped into a preheated MAI solution (70 °C, 60 s) dissolved in 2-propanol. Finally, the substrates were postannealed at a temperature of 75 °C for 20 min. It is noticed that in several works large variations in the postannealing temperature is employed for the perovskite crystallization, for example in the one-step method, 90–120 °C for 45–120 min.^{60,165,169,172,176,179,181,182,184,185,189,195,199} The fine-tuning of annealing temperature and time will play a major role for the delicate incorporation of Cl into the perovskite crystal lattice. The XRD data on MAPbI_{3-x}Cl_x films using the two-step method by Li et al.¹⁷³ show that the (110) and (220) diffraction peaks shift as a PbCl₂ concentration increases (Figure 11a,b). Although the shift is small, it is above the instrumental sensitivity and not negligible. They performed further XRD data fitting extracting the lattice parameters and unit cell volumes (inset in Figure 11a). In addition, the XRD pattern shows that the Cl-content in perovskites prepared by the two-step dipping method is higher than the ones fabricated by the one-step method. Furthermore, no (110) diffraction peak (15.68°) related to MAPbCl₃ was found indicating no phase segregation occurring (Figure 11a,b).^{173,186,196} The UV-vis of perovskite films show the shift of absorption edge toward a shorter wavelength with increasing PbCl₂:PbI₂ mole fraction (Figure 11c) corroborating with XRD on solid-solution formation. Zhang et al.¹⁶⁷ synthesized single-crystalline nanofibers (NFs) of MAPbI_{3-x}Cl_x perovskites in the tetragonal phase at room temperature with the Cl-content between 0 ≤ x

≤ 0.75. Furthermore, based on SAED patterns it was suggested that I ions are not substituted by Cl ions along the [001] direction of single-crystalline NFs, but the Cl incorporation takes place along the [100] or [010] directions, i.e., within the crystal ab-plane (Figure 11d). In other words, Cl-inclusion occurred along the [001] direction (apical positions) because of high intensity of (110) diffraction planes observed for MAPbI_{3-x}Cl_x perovskites.^{15,206-208} Pistor et al.¹⁸⁰ and Luo et al.¹⁷¹ synthesized MAPbI_{3-x}Cl_x perovskites under vacuum conditions by coevaporation of MAI and PbCl₂ sources. Interestingly, upon subsequent postannealing treatments, they observed that a phase transition from tetragonal to a cubic phase (space group *Pm* $\bar{3}$ *m* with lattice constant of 6.276 Å) takes place in MAPbI_{3-x}Cl_x perovskites, which was also reported to be stabilized after the cooling at room temperature. As comparison, MAPbCl₃ adopts a cubic structure at room temperature (Figure 2d). More recently, a recipe for the synthesis of single crystal MAPb(I/Cl) was reported by Lian et al.²⁰⁹ and their TOF-SIMS measurements confirmed the trace amount of Cl in the MAPb(I/Cl) bulk film.

The second possible scenario for the discrepancies in the reports could lie on the technical aspects of the measurement conditions. Most of chemical analytical tools (XPS, EDX, TOF-SIMS, etc.) require high vacuum under which volatile species contained in perovskite films can desorb.¹⁷⁴ Furthermore, the situation can be more drastic when perovskite films are exposed to a probing beam (e-beam, UV, X-ray, etc.) that may lead to beam-induced artifacts during measurement. Colella et al.¹⁸⁵ mentioned that reliable values of Cl concentration could not be achieved as the amplitude of the Cl-related signal was observed to decrease during the measurement in EDX. Starr et al.¹⁷⁴ emphasized the importance of employing nondestructive techniques (e.g., HAXPS and FY-XAS) to guarantee that the measurement itself will not modify the original chemical composition of the perovskite samples. Based on the extensive analytical measurements (Table 2), our current understanding is that the incorporation of Cl ions into the perovskite crystal lattice is viable only in a small quantity (<1 at. %), but sufficient to induce new material properties.¹⁷³ Alternatively, Cl-based additives may be another way to facilitate Cl-incorporation into the perovskite lattice. HCl,²¹⁰⁻²¹² NH₄Cl,²¹³ CaCl₂,²¹⁴ tetraphenylphosphonium chloride (TPPCL),²¹⁵ and butylphosphonic acid 4-ammonium chloride (4-ABPACL)²¹⁶ additives were reported to play an important role in the formation of high quality films as well as enhancement of device stability; however, the question whether chloride ions enter the perovskite crystal lattice is still under debate.²¹⁷

To our knowledge, there are only two reports focusing on the FAPb(I/Cl) system.^{218,219} Lv et al.²¹⁸ employed the one-step method by mixing PbCl₂ and FAI in a molar ratio of 1:3 in DMF followed by stirring at 60 °C for 30 min. The solution was spin coated on FTO/c-TiO₂/mp-TiO₂, and the FAPbI_{3-x}Cl_x films were formed by postannealing in an oven at a temperature ranging from 120 to 170 °C for 30 min. Based on XRD data, annealing at 140 °C generated FAPbI_{3-x}Cl_x perovskite in the tetragonal phase (space group *P3m1*, *a* = *b* = 8.977(7) Å, *c* = 10.890(2) Å). A slight *c* lattice parameter contraction of ~1.1% compared to the FAPbI₃ perovskite (*a* = *b* = 9.000(8) Å, *c* = 11.012(2) Å) was attributed to the partial substitution of Cl⁻ into the perovskite structure. EDX was also employed to estimate the chlorine content, but it was barely detectable.²¹⁸ Wang et al.²¹⁹ employed NH₄Cl, MAcl, and FACL additives to suppress the formation of yellow δ -phase

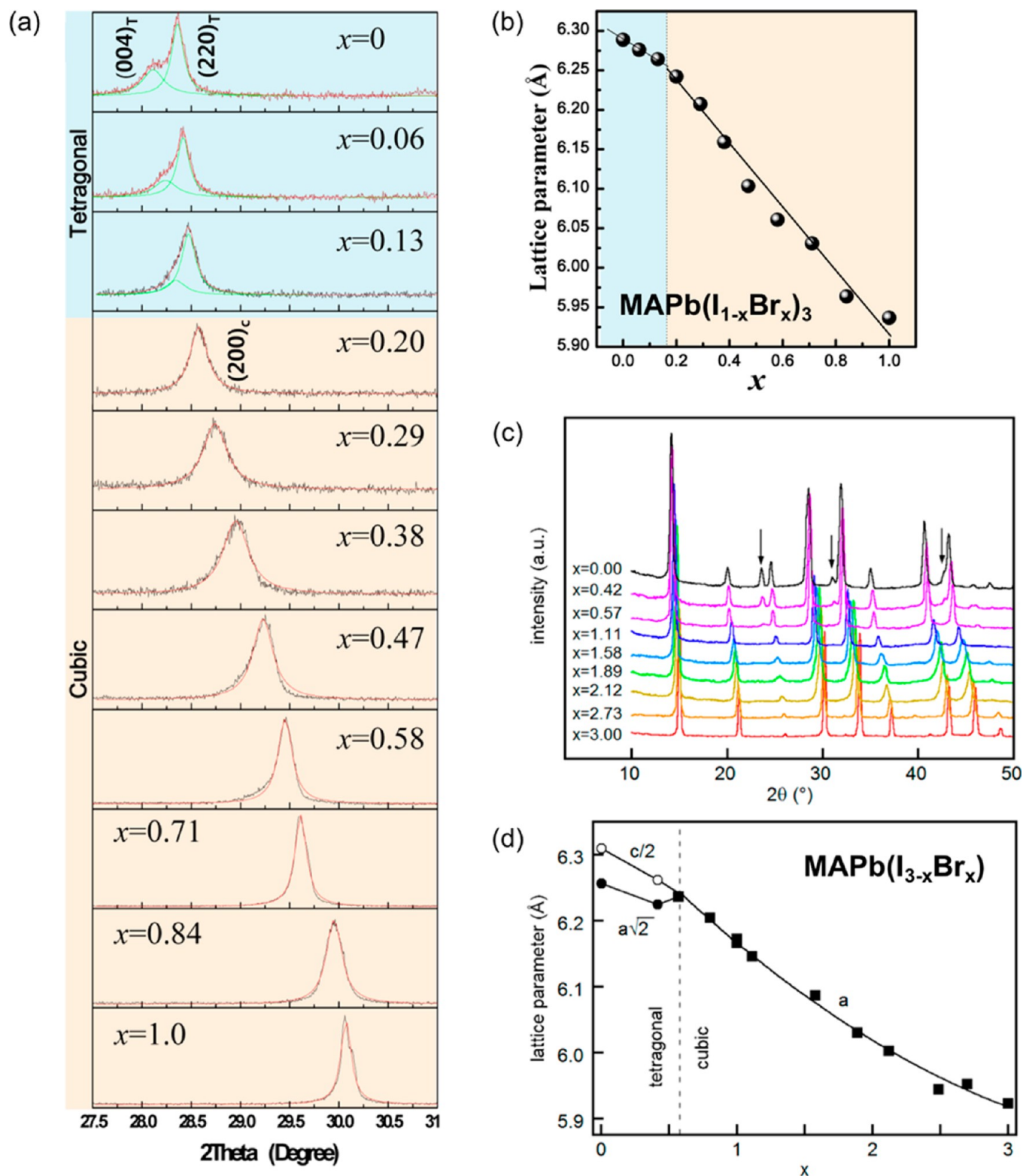


Figure 12. (a) Narrow-range XRD pattern of MAPb(I/Br) perovskites with varying Br:I molar ratio concentrations magnified in the region of the tetragonal $(004)_T$ and $(220)_T$ and cubic $(200)_C$ peaks. (b) Lattice constants of pseudocubic or cubic as a function of Br:I molar ratio concentrations. (c) Wide-range XRD pattern of MAPb(I/Br) perovskites and (d) extracted lattice parameters. (a,b) Reprinted with permission from ref 220. Copyright 2013 American Chemical Society. (c,d) Reprinted with permission from ref 226. Copyright 2015 American Chemical Society.

659 FAPbI₃. It has been proposed that suppressing the formation of
 660 δ -FAPbI₃ phase at all stages of the formation of FAPbI₃ on mp-
 661 TiO₂ is essential for achieving high-purity black α -FAPbI₃
 662 perovskite phase.

3.2. Binary MAPb(I/Br) and FAPb(I/Br) Systems. 663
 664 Contrary to the MAPb(I/Cl) material system described in
 section 3.1, the substitution of I⁻ by Br⁻ ions has been widely 665
 demonstrated and to effectively tune the band gap of 666

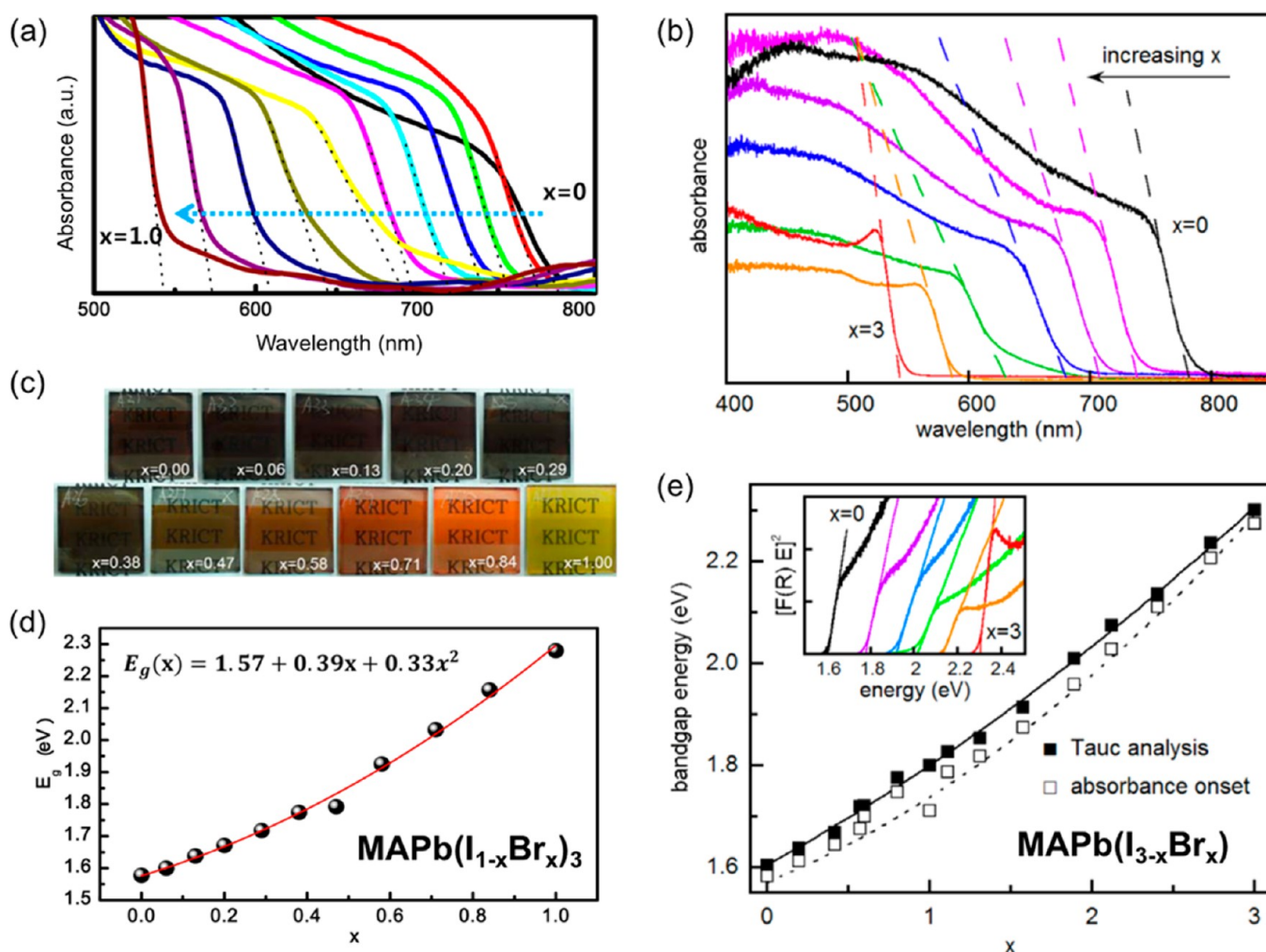


Figure 13. (a,b) UV-vis absorbance spectra of MAPb(I/Br) perovskites with varying Br:I molar ratio concentrations. (c) Photographs of MAPb(I/Br) perovskite films. (d,e) Band gap energies as a function of Br:I concentrations extracted from the absorbance onsets and Tauc analysis. Inset in (e) shows representative Tauc plots used to determine E_g . (a,c,d) Reprinted with permission from ref 220. Copyright 2013 American Chemical Society. (b,e) Reprinted with permission from ref 226. Copyright 2015 American Chemical Society.

perovskites.²²⁰ The photovoltaic properties of mixed MAPbI_{3-x}Br_x were first demonstrated by Noh et al.²²⁰ reporting an efficiency of 12.3%. Jeon et al.⁶² achieved a certified PCE of 16.2% (Figure 1b) by the solvent-engineering method, which enables the deposition of uniform and dense perovskite films of MAPbI_{3-x}Br_x ($x = 0.10-0.15$). Later alternative deposition methods such as vacuum deposition,²²¹ low-pressure vapor-assisted solution (LP-VASP),²²² and printing²²³ have been demonstrated. In the work by Noh et al.,²²⁰ the Br content (<10%) gave the best initial efficiency due to a lower band gap, but higher Br contents (>20%) provided a better high-humidity shelf life stability (RH 55%). This was correlated with a tetragonal to pseudocubic structural transition (at $x = 0.13$) arising from a higher t factor due to the smaller ionic radius of Br (Figure 12). Mixed perovskites composed of two different perovskite crystals with similar lattice constants follow the Vegard's law. According to this law, a linear dependence of the lattice parameter with composition is expected, in the absence of strong electronic effects.^{220,224-228} Fedeli et al.²²⁶ claimed that in the cubic regime ($x \geq 0.57$, MAPbI_{3-x}Br_x), a small deviation from the Vegard's law was observed indicating additional interactions in the mixed-halogen (Figure 12d). Furthermore, a blue-shift of the absorption edge, i.e., increase in

band gap following a quadratic function of the Br content was extracted from absorbance onsets, $E_g(x) = 1.57 + 0.39x + 0.33x^2$ (eV) by Noh et al.²²⁰ (Figure 13a,c,d). Fedeli et al.²²⁶ emphasized that slight variations (a few tens of meV) in E_g values are obtained when derived from the Tauc analysis and compared to E_g values extracted from the absorbance onset (Figure 13b). The $E_g(x)$ expression that is independent of the film properties (thickness, surface roughness related scattering) was derived based on Tauc plots (Figure 13e):

$$E_g(x) = E_{I_3} + (E_{Br_3} - E_{I_3} - b)\frac{x}{3} + b\left(\frac{x}{3}\right)^2 \quad (3)$$

where E_{I_3} and E_{Br_3} are the bandgaps of MAPbI₃ (1.604 eV) and MAPbBr₃ (2.307 eV), respectively, and b (0.175 eV) is the so-called bowing parameter that accounts for the effects of composition disorder on the conduction and valence band edges. Based on the low b value extracted, the authors concluded that compositional disorder is low in MAPb(I/Br) perovskites (Figure 14a). Based on DFT calculations, Mosconi et al.²⁰⁶ proposed that Br can occupy both apical and equatorial positions in the PbX₆ octahedra. Based on combined first-principles total energy calculations with statistical mechanical treatments (evoking the energy and entropy of mixing), Brivio

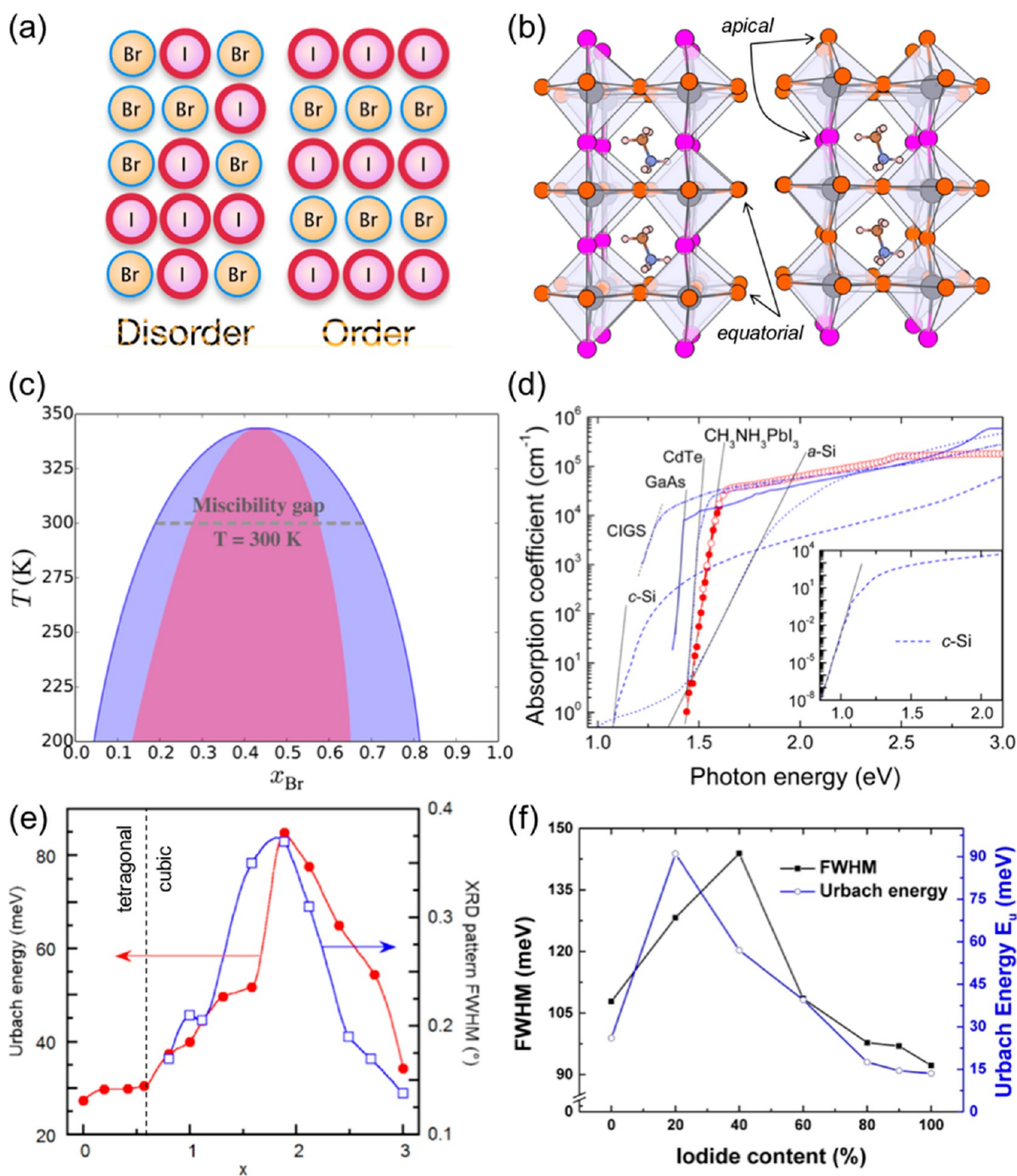


Figure 14. (a) Schematic illustration of disordered and ordered systems in MAPb(I/Br). (b) Identified stable ordered structures for MAPbI₂Br₂ and MAPbI_{1/2}Br_{3/2} (layered with iodine at the apical positions) that minimize internal strain arising from the size mismatch between I and Br. (c) Calculated phase diagram of MAPb(I_{1-x}Br_x)₃ alloy. Purple and pink lines are binodal and spinodal lines, respectively. The dashed horizontal line corresponds to miscibility gap at room temperature. Thermodynamically stable alloy can be formed only in the white region. Reprinted with permission from ref 229. Copyright 2016 American Chemical Society. (d) Effective absorption coefficient of MAPbI₃ perovskite compared with other technologically relevant photovoltaic materials, amorphous-silicon (a-Si), GaAs, CIGS, CdTe, and crystalline silicon (c-Si), all corresponding to room temperature measurements. For each material, the slope of the Urbach tail is shown. The inset shows the data for c-Si down to low absorption values. Reprinted with permission from ref 230. Copyright 2014 American Chemical Society. (e) Urbach energy and fwhm of the XRD pattern in the cubic phase composition range as a function of Br:I molar ratios in MAPb(I/Br) perovskites. Reprinted with permission from ref 226. Copyright 2014 American Chemical Society. (f) Urbach energy and fwhm of the PL peak as a function of iodide concentration in MAPb(I/Br) perovskites. Reprinted with permission from ref 231. Copyright 2014 American Chemical Society.

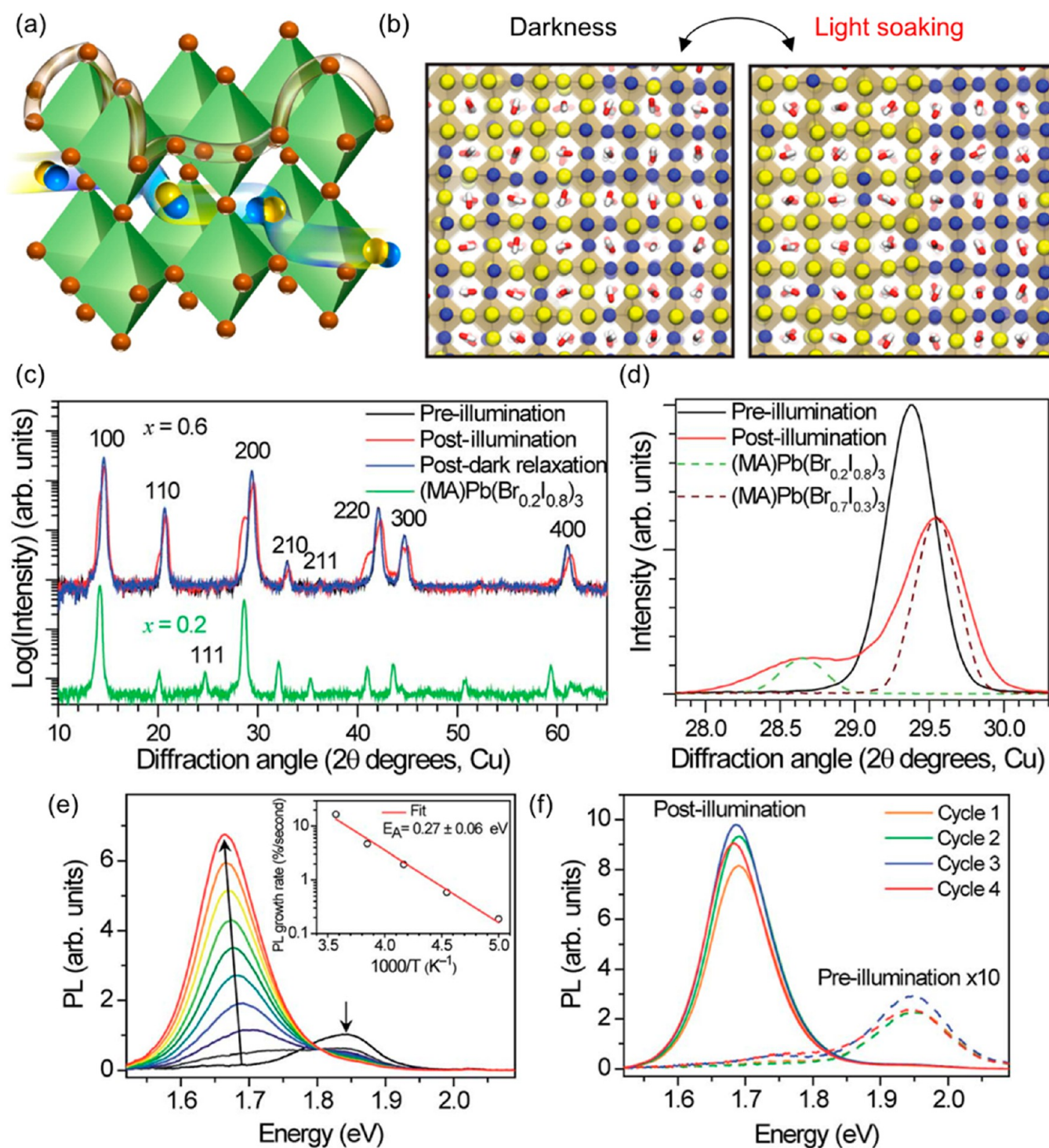


Figure 15. (a) Illustration of proposed migration path of I⁻ ions along the I⁻-I⁻ edge of the PbI₆ octahedra in MAPbI₃. MA⁺ migration was also corroborated experimentally. Reprinted with permission from ref 232. Copyright 2016 American Chemical Society. (b) Schematic illustration of phase separation (Light soaking) and reversibility (darkness) in MAPb(I/Br) system, where yellow and blue spheres represent I⁻ and Br⁻, respectively; the red and white pill shapes represent MA. Reprinted with permission from ref 233. Copyright 2017 American Chemical Society. (c) XRD pattern of MAPbBr_{0.6}I_{0.4} film before (black) and after (red) white-light soaking (~50 mW/cm²; ~0.5 sun) for 5 min, and subsequently after 2 h in the dark (blue). XRD of MAPbBr_{0.2}I_{0.8} (green) is shown for comparison. (d) XRD peak of MAPbBr_{0.6}I_{0.4} film before (black) and after (red) white-light soaking (~50 mW/cm²) for 5 min. XRD of MAPbBr_{0.2}I_{0.8} (dashed green) and MAPbBr_{0.7}I_{0.3} (dashed brown) are shown for comparison. (e) PL spectra of MAPbBr_{0.4}I_{0.6} film over 45 s in 5 s increments under 475 nm, 15 mW/cm² at 300 K. Inset shows temperature dependence of initial PL growth rate. (f) PL spectra of MAPbBr_{0.6}I_{0.4} film after sequential cycles of illumination (475 nm, 15 mW/cm²) for 2 min followed by 5 min in the dark. Reprinted with permission from ref 234. Copyright 2015 The Royal Society of Chemistry.

711 et al.²²⁹ identified the two most stable configurations 712 corresponding to ordered structures of MAPbI₂Br₂ and

MAPbI_{1/2}Br_{5/2} (Figure 14b). Both structures were formed 713 with iodine ions located at the apical locations. Furthermore, 714

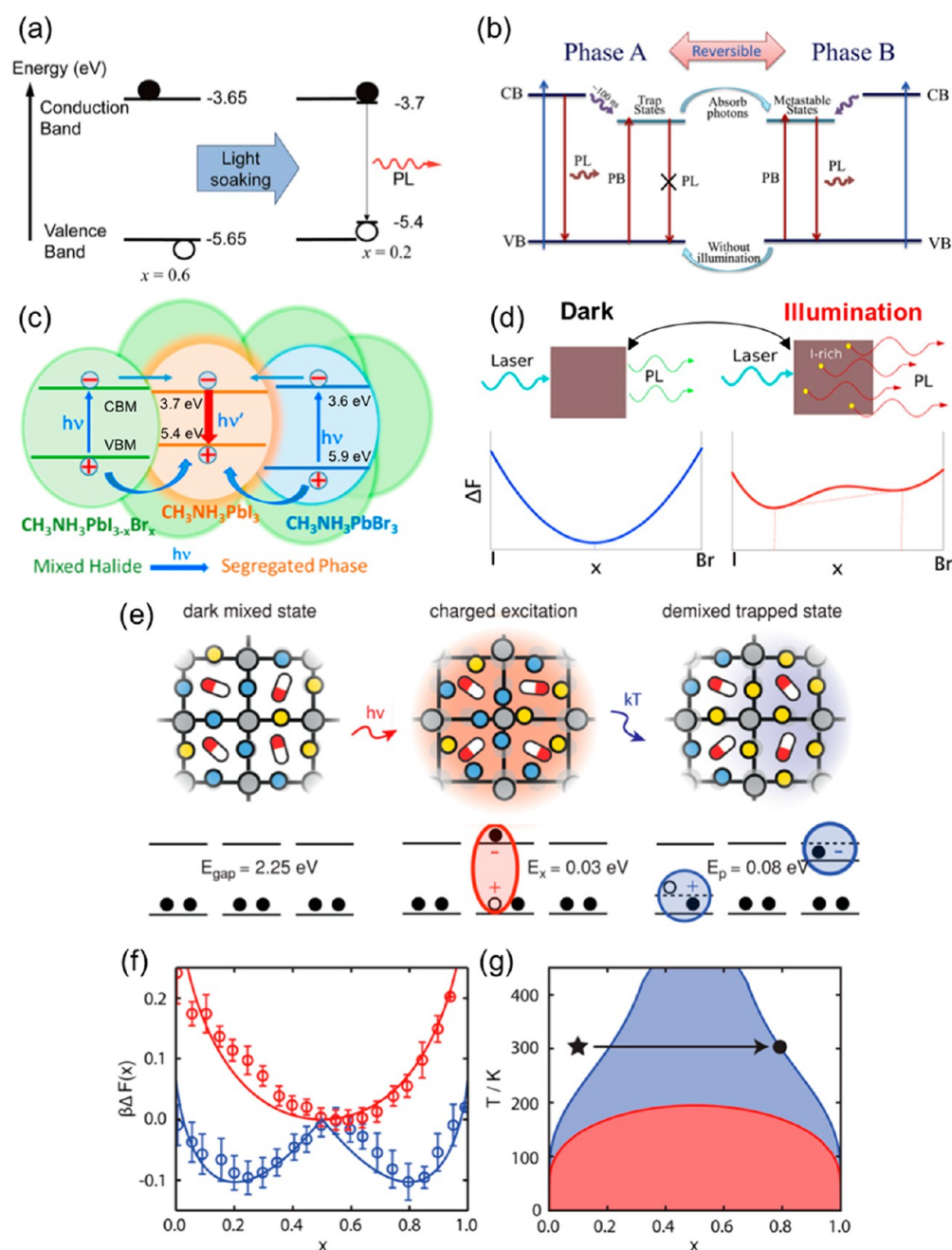


Figure 16. (a) Illustration of proposed mechanism for photoinduced trap states formation within the band gap in $\text{MAPbBr}_{0.4}\text{I}_{0.6}$. Photogenerated holes or excitons are hypothesized to stabilize the formation of iodine-rich domains (e.g., $\text{MAPb}(\text{Br}_{0.2}\text{I}_{0.8})_3$) that dominate PL. Reprinted with permission from ref 234. Copyright 2015 The Royal Society of Chemistry. (b) Illustration for the proposed two-step mechanism for light induced segregation: initial photon absorption will generate trap states in the band gap after 100 ns; these traps have lifetime of 1 μs (Phase A). If subsequent coming photons are absorbed by these traps, it transforms to a metastable state (Phase B), which behave as new energy band and can lead to PL emission. Reprinted with permission from ref 235. Copyright 2016 Elsevier Ltd. (c) Illustration of $\text{MAPb}(\text{I}/\text{Br})_3$ with each of the domains (iodide- and bromide-rich) having characteristic conduction band minimum (CBM) and valence band maximum (VBM) values will lead to iodide-rich domains to serve as sinks for the photogenerated charge carriers. Reprinted with permission from ref 236. Copyright 2016 American Chemical Society. (d) Schematic Helmholtz free-energy (ΔF) curves in light and in dark predicting phase segregation. The red vertical dashed lines show that the lowest attainable free energy in light occurs when the material segregates into I-rich and Br-rich phases, whereas the blue dashed line shows that the lowest energy corresponds when the material remains in a single phase in dark. Reprinted with permission from ref 237. Copyright 2016 American Chemical Society. (e) Photoinduced polaron trapping and associated energy scales. Yellow and blue spheres represent I^- and Br^- , respectively; the red and white pill shapes represent MA. Lead atoms represented by gray circles. (f) Free energies per unit cell for $\text{MAPb}(\text{I}_x\text{Br}_{1-x})_3$ with varying composition in the ground (red) and photoexcited (blue) states, computed from MD simulations (circles) and mean field theory (solid lines). (g) Temperature–composition phase diagram in the ground (red) and photoexcited state (blue). Areas beneath the red and blue coexisting curves indicate demixed states. The arrow through the phase diagram from initial state (star) to demixed state (circle) correspond to experimental observation for $\text{MAPb}(\text{I}_{0.1}\text{Br}_{0.9})_3$. Reprinted with permission from ref 233. Copyright 2017 American Chemical Society.

715 based on the Helmholtz free energy variation, a phase diagram
716 of MAPb(I/Br) was constructed predicting that (i) there is a
717 critical temperature of ~ 343 K, above which the solid solution
718 is stable for any Br/I composition; (ii) at 300 K, the alloy is not
719 stable against a phase separation in MAPbI₃ and MAPbBr₃ in
720 the range of compositions between $0.19 < x < 0.68$ (miscibility
721 gap). The lattice constant mismatch ($\sim 6\%$) between MAPbI₃
722 and MAPbBr₃ was associated with the instability of isovalent
723 solid solutions.²²⁹

724 Further investigations regarding disorder dependence on Br:I
725 ratio were conducted by analyzing the exponential decay of the
726 sub-band gap absorbance, commonly described by the empirical
727 Urbach rule, $A \propto \exp(E/E_0)$, where A is absorbance, E photon
728 energy, and E_0 the characteristic Urbach energy representing
729 the width of the exponential Urbach tail (Figure 14d).^{226,230,231}
730 Despite the fact that there is no theoretical derivation for the
731 Urbach rule, a general consensus exists that the Urbach tail in
732 crystalline semiconductors is related to the static (structural
733 disorder) and/or dynamic (phonon) disorder, that arises from
734 lattice point defects, dislocations, strain, deviation from ideal
735 stoichiometry, and grains.^{226,230,231} Pure MAPbI₃ and MAPbBr₃
736 perovskites were determined to have sharp absorption edges
737 with low Urbach energy of ~ 15 – 27 meV and ~ 23 – 34 meV,
738 respectively.^{226,230,231} Compared with other semiconductors
739 (Figure 14d), the Urbach energy of MAPbI₃ is close to the
740 value of GaAs (monocrystalline direct band gap semi-
741 conductor). Crystalline Si (c-Si) shows also a similar slope
742 below its bandgap, but because of its indirect band gap, it
743 occurs at much a lower value and shows signatures related to
744 phonon-assisted absorption (inset in Figure 14d). The disorder
745 evaluated on MAPbI_{3-x}Br_x perovskites through the Urbach
746 energy showed larger values reaching a maxima of ~ 85 meV
747 (MAPbI_{1.12}Br_{1.88}, Figure 14e)²²⁶ and ~ 90 meV (MAPbI_{1.2}Br_{1.8},
748 Figure 14f).²³¹ Interestingly, the dependence of x with these
749 Urbach energies was closely related to grain-size domains
750 (Figure 14e), and it was suggested that the sub-band gap
751 absorption is due to defect states localized at the grain
752 boundaries.²²⁶ It was also shown that the fwhm of the
753 photoluminescence (PL) peaks are correlated with the Urbach
754 energies (Figure 14f). This observation indicates that
755 absorption and emission arise from similar states.²³¹ Also, the
756 larger PL fwhm and Urbach energy with iodine concentrations
757 of 20–40% (Figure 14f) was observed to correlate with shorter
758 lifetimes observed in transient PL.²³¹

759 Although under storage conditions the MAPb(I/Br) system
760 exhibits enhanced stability, it has been reported that dynamical
761 processes (e.g., material degradation and phase segregation)
762 takes place when exposed to light. Misra et al.²²⁷ performed
763 photochemical stability tests of encapsulated films of
764 MAPbI_{3-x}Br_x ($x = 0.11, 0.16, 0.22, 1$) under accelerated
765 stressing conditions using concentrated sunlight (100 suns).
766 They observed that MAPbBr₃ was the most stable composition
767 exhibiting no degradation, whereas increasing iodine incorpo-
768 ration leads to accelerated photochemical degradation yielding
769 PbI₂ as a final remaining product. No degradation was recorded
770 for MAPbI_{3-x}Br_x under shelf-storage conditions, and degrada-
771 tion was observed to occur only under illumination conditions.
772 Currently, it is widely accepted that ion migration is one of the
773 causes for the anomalous photocurrent hysteresis phenomena
774 (Figure 15a).^{232,238,239} In addition, there are also concerns that
775 ion migration is closely correlated with inherent instability
776 issues in perovskite solar cells. For example, Hoke et al.²³⁴
777 reported first on the occurrence of a serious phase segregation

in MAPb(I/Br) films under illumination (coined as Hoke 778
effect^{237,240}). The initially homogeneous MAPb(Br_xI_{1-x})₃ (0.2 779
 $< x < 0.9$) perovskite films were observed to undergo 780
photoexcited phase-separation into two phases, one iodine- 781
rich and the other bromide-rich domain in the same film 782
(Figure 15b), as corroborated by the peak splitting probed in 783
XRD (Figure 15c,d) and PL (Figure 15e,f) measurements. As 784
shown in Figure 15d, when the MAPb(Br_{0.6}I_{0.4})₃ film was 785
illuminated with white light, the original diffraction pattern 786
splits into two peaks, suggesting formation of phases with larger 787
and smaller lattice constants. More interestingly, when allowed 788
to relax in dark, the XRD pattern returned to its original, single 789
phase state. Additional, PL (Figure 15f) and absorption spectra 790
acquired before and after light soaking confirmed this 791
reversibility (Figure 15b). Temperature dependence (200– 792
300 K) was performed to verify that phase segregation occurs 793
solely due to illumination and not from temperature increase 794
during illumination (Figure 15e). Although phase segregation 795
(disappearance of the original peak and rise of new lower- 796
energy peak) was observed at the lowest temperature of 200 K, 797
the time scale for the changes varied widely. The changes 798
occurred in ~ 1 min at 300 K, whereas at 200 K the phase 799
segregation took ~ 1 h to complete. The Arrhenius plot 800
behavior (inset in Figure 15e) was observed for this phase 801
segregation, where an activation energy of ~ 0.3 eV was 802
extracted. Interestingly, this value is similar to halide ion 803
migration activation energies reported previously, supporting 804
further that ionic transport plays a major role in perovskite solar 805
cells (Figure 15a,b).^{232,234,237} It is emphasized that even 806
without light illumination, the PL peak position was observed 807
to shift in MAPbBr_{1.2}I_{1.8} films from 1.68 to 1.94 eV after 2 808
weeks storage under dark and inert environment.²³¹ This was 809
attributed to a slow, but spontaneous room temperature phase 810
segregation within the miscibility gap (Figure 14c).²²⁹ 811

A number of reports proposed the microscopic origins for 812
the phase-segregation phenomena.^{74,233–237,240} Hoke et al. 813
proposed that the existence of initial iodide-rich domains 814
(nucleation points) in as prepared MAPb(I/Br) samples 815
stabilizes hole accumulation upon illumination (Figure 16a). 816
These accumulation of unbalanced charges (holes) provide 817
further driving enthalpy for halide segregation growing further 818
the iodide-rich domain sizes (Figure 15d,e). Yang et al.²³⁵ 819
employed PL and photomodulation (PM) spectroscopy to 820
study the MAPbBr_{1.7}I_{1.3} films. Using continuous wave (CW) or 821
10 ns pulsed laser with high repetition rates, they observed that 822
the initial PL ~ 640 nm (~ 1.94 eV) peak position shifted to a 823
longer wavelength of ~ 750 nm (~ 1.65 eV) after illumina- 824
tion.²³⁵ Interestingly, with the same 10 ns pulsed laser, but 825
using lower repetition rates (< 500 Hz), no matter what the 826
intensity of excitation light is, PL did not show shift in peak 827
position. Based on CW pump probe spectroscopy and transient 828
dynamics, Yang et al. proposed a two-step process: initial 829
photon absorption will generate trap states within the band gap, 830
but these traps will not lead to PL emission; these traps have 831
lifetime of 1 μ s (Phase A). If subsequent incident photons are 832
absorbed by these traps within the lifetime, it transforms to a 833
long-lived (\sim ms) “metastable” state (Phase B), which behaves 834
as new energy band and can lead to PL emission (Figure 16b). 835
These polar states (later identified as polaronic states²³³) were 836
hypothesized to locally drive ion migration. Yoon et al.²³⁶ 837
proposed that MAPb(I/Br) with each of the domains (iodide- 838
and bromide-rich) having characteristic conduction band 839
minimum (CBM) and valence band maximum (VBM) values 840

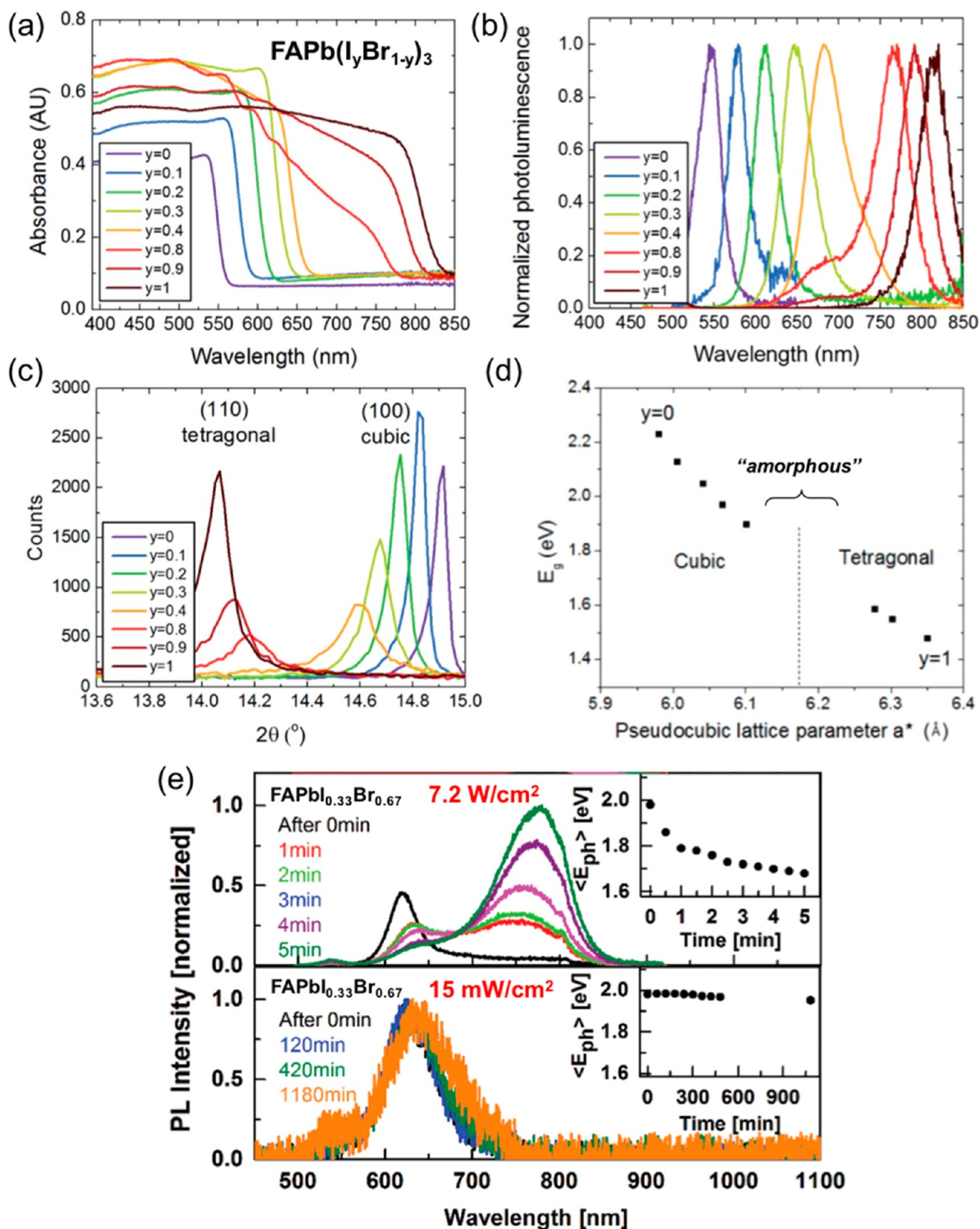


Figure 17. (a) UV-vis absorbance, (b) steady-state PL spectra, (c) XRD of FAPbI_xBr_{3-x} perovskites with varying x composition. (d) Variation of band gap as a function of pseudocubic lattice parameter extracted from XRD data. Reprinted with permission from ref 147. Copyright 2014 The Royal Society of Chemistry. (e) PL spectra of FAPb(Br_{0.67}I_{0.33})₃ film over 5 min of continuous illumination following excitation at 400 nm with intensity of 7.2 W/cm² and 15 mW/cm². Inset: Change of the average photon energy (E_{ph}) as a function of time. Reprinted with permission from ref 242. Copyright 2015 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim.

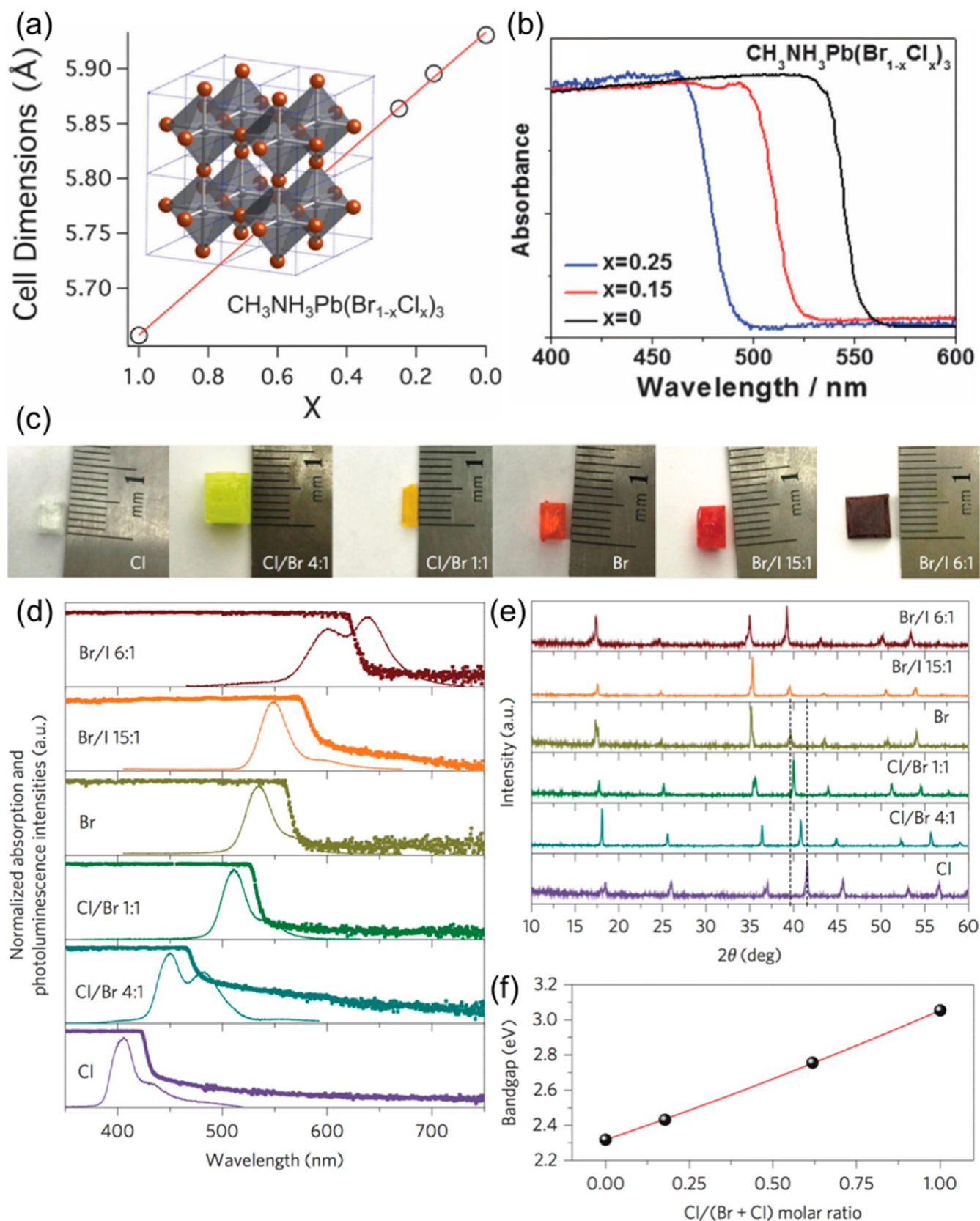


Figure 18. (a) Single crystal XRD cell dimensions and (b) UV-vis absorption spectra for $\text{MAPb}(\text{Br}_{1-x}\text{Cl}_x)_3$ ($x = 0, 0.15, 0.25$). Reprinted with permission from ref 247. Copyright 2015 The Royal Society of Chemistry. (c) Photographs of $\text{MAPb}(\text{Br}/\text{Cl})$ and $\text{MAPb}(\text{B}/\text{I})$ mixed halide perovskite single crystals with different halide compositions. (d) Absorption and PL spectra. (e) Powder XRD. (f) Crystal band gap extracted from PL as a function of $\text{Cl}/(\text{Cl} + \text{Br})$ molar ratio derived from XRD. Reprinted with permission from ref 246. Copyright 2015 Nature Publishing Group.

841 will lead to iodide-rich domains to serve as charge
842 recombination sites and thus contribute to the observed PL
843 emission (Figure 16c). Slotcavage et al.²³⁷ discussed the
844 thermodynamic origins of observed light-induced phase
845 segregation (Figure 16d). Bischak et al.²³³ performed combined
846 cathodoluminescence (CL) mapping to visualize the spatial
847 segregation into iodide- and bromide-rich domains. After
848 prolonged illumination, small clusters enriched in iodide were
849 observed to localize near grain boundaries. Their molecular
850 dynamics (MD) simulations were able to demonstrate that
851 photoinduced phase separation arises when charged excitations
852 generate enough lattice strain to destabilize the solid solution,
853 favoring phase-segregation. Their findings show that upon light
854 absorption weakly bound electron–hole pair (binding energy
855 ~ 0.03 eV) quickly dissociates into free carriers (Figure 16e).
856 Due to the ionic nature of the perovskite, these charges deform
857 the surrounding lattice through electron–phonon coupling.
858 The charge and the lattice deformation field that surrounds
859 together form a polaron that was predicted to have an average
860 size of 8 nm and binding energy of $E_p = 0.08$ eV (Figure
861 16e).²³³ This polaron-induced lattice distortion changes the
862 magnitude of the free energy as a function of bromide content
863 from one with one minimum (in dark, red curve in Figure 16f)
864 to another with two minima (under illumination, blue curve in
865 Figure 16f). They also determined a full temperature–
866 composition phase diagram for both ground- and photoexcited
867 states (Figure 16g). This graph shows that for MAPb(I_{0.1}Br_{0.9})₃
868 film (star symbol in Figure 16g), temperature variation over a
869 range of 50 K was insufficient to induce demixing, but does
870 increase the demixing rate under light illumination (circle in
871 Figure 16g). Their model suggests that naturally occurring
872 small variations in perovskite composition that exists before
873 illumination will yield iodide-rich domains that has a reduced
874 band gap. Therefore, perovskite compositional homogeneity
875 seems to be one of the determining factors that will favor
876 stability.^{233,237,240} It has been shown that photoinduced phase
877 segregation is minimized in highly crystalline perovskites.²⁴¹
878 Similarly, mixed halide FAPbI_{3-x}Br_x perovskites were also
879 synthesized with x varying from 0 to 1.^{147,242} Interestingly, both
880 reports described that they were unable to form crystalline
881 phases of FAPb(I_{1-x}Br_x)₃ perovskites with bromide contents in
882 the range $0.3 < x < 0.5$, as it was apparent from the absence of
883 significant UV–vis, PL, and XRD data (Figure 17a–d). The
884 term “amorphous”²⁴² was used implying that the crystalline
885 order is too short in the length scale detectable by XRD (Figure
886 17d).¹⁴⁷ Fundamental origins for this amorphous regime have
887 not been fully clarified. Phase segregation phenomena in
888 iodide- and bromide-rich domains were also reported for the
889 FAPb(I/Br) system.²³⁴ Rehman et al.²⁴² reported that the
890 intensity of the laser excitation spot played a major role in
891 producing phase segregation (Figure 17e). The PL emission of
892 FAPb(Br_{0.67}I_{0.33})₃ films recorded for laser excitation with an
893 intensity of 7.2 W/cm² following 5 min of continuous
894 illumination showed the gradual decrease and shift of the 620
895 nm (~ 2 eV) to a new dominant low-energy PL feature at ~ 785
896 nm (1.58 eV). On the other hand, when an identical
897 FAPb(Br_{0.67}I_{0.33})₃ film was illuminated at the same wavelength
898 (400 nm) with lower laser intensity of 15 mW/cm², no changes
899 in PL spectra were observed even after 1180 min (~ 20 h)
900 exposure (Figure 17e). Compared to the MAPb(I/Br) system,
901 it was reported that light intensity as low as 1.2 mW/cm² with
902 continuous illumination (10 min) can induce phase segrega-
903 tion.²³⁵ Because photovoltaic devices are typically tested under

AM1.5 conditions (100 mW/cm²), strategies to make both
MAPb(I/Br) and FAPb(I/Br) systems resilient to phase
segregation phenomena is needed.

3.3. Binary MAPb(Br/Cl) System. Contrary to the
MAPb(I/Cl) system described in section 3.1, Cl was confirmed
to coexist with I in the MAPb(Br/Cl) system.^{243–249}
Theoretical calculations determining the phase transitions of
temperature-halide compositions showed a relatively low
demixing critical temperature for MAPb(Br/Cl) system
(~ 140 K), which was similar to that of MAPb(Br/I) system
(~ 190 K).²³³ In comparison, for the MAPb(I/Cl) system, the
determined critical solution temperature of ~ 1800 K was well
above room temperature, explaining why this mixture is
unstable.²³³ The larger difference in ionic radii, Cl⁻ (1.67 Å)
< Br⁻ (1.84 Å) < I⁻ (2.07 Å)²⁵⁰ and the higher degree of ionic
character is explained as origin for the “easy” miscibility
between (Cl/Br) and (Br/I) system, but not for the (Cl/I)
system.^{15,233,250} This is also reinforced by the fact that growth
of halide mixed MAPbBr_{3-x}Cl_x and MAPbBr_{3-x}I_x single crystal
perovskites is possible in the full range of Br:Cl and Br:I molar
ratios confirming halide miscibilities (Figure 18).^{17,217,246,247}
The single crystals were synthesized from stoichiometric PbBr₂
and [y MACl + (1 - y)MABr] precursor solutions in DMF by
Zhang et al.²⁴⁷ employing the solvothermal method, which is
based on inverse solubility (Figure 18a,b). Fang et al.²⁴⁶
employed the cooling-induced crystallization method to growth
mixed halide perovskite single crystals (Figure 18c–f). The
authors constructed photodetector devices based on these
single crystals with a thickness of ~ 1 mm. The response
spectrum could be tuned from blue to red by varying the halide
composition of these single crystals. The main advantage lies on
the very narrowband fwhm (less than 20 nm) photo-
detection.²⁴⁶ Among the family of mixed halide perovskites,
MAPbBr_{3-x}Cl_x has a larger band gap (Figure 18b,d,f) than that
of iodine-based perovskites (MAPbI_{3-x}Cl₃ and MAP-
bI_{3-x}Br_x).²¹⁷ Therefore, most of works focused on perovskite
light-emitting devices where the emission color could be tuned
from the red to blue color by modulating the halide
composition in the perovskite.^{245,246,251–253}

3.4. Ternary MAPb(I/Br/Cl) System. A few reports
describe the triple halide-mixed MAPb(I/Br/Cl) system.^{254–256}
Similar to what was discussed in section 3.1, the amount of Cl
incorporated seems negligible. Nevertheless, Cl was shown to
impact device performance by improving coverage and
reducing crystal growth rate.^{254–256} On the other hand, Br
was observed to influence strongly the band gap of MAPb(I/
Br/Cl) system and consequently increase open-circuit voltage
(V_{oc}). In addition, Br incorporation was reported to stabilize
crystal lattice improving lifetime of device. These influences of
Br incorporation are similar to what was discussed in section
3.2. We observe that the ternary MAPb(I/Br/Cl) system
resembles a linear combination of MAPb(I/Cl) + MAPb(I/Br)
systems. Chiang et al.²⁵⁶ synthesized MAPb(I/Br/Cl) perov-
skites by a combined hot solution spin-coating and solvent
annealing film casting engineering obtaining high quality solid
solution films. Their inverted solar cell devices (ITO/
PEDOT:PSS/MAPbI_{3-x-y}Br_xCl_y/PC₆₁BM/Ca/Al) with active
areas of 0.1 cm² generated $J_{sc} = 19.25$ mA/cm², $V_{oc} = 1.10$ V,
FF = 78%, PCE = 16.52% (the average efficiency from 34
devices yields a small standard deviation: PCE = $15.61 \pm$
0.84%). Furthermore, the same authors up-scaled the process
and their designed inverted perovskite module composed of 9
cells in series with an active area of each cell equal to 4 mm × 966

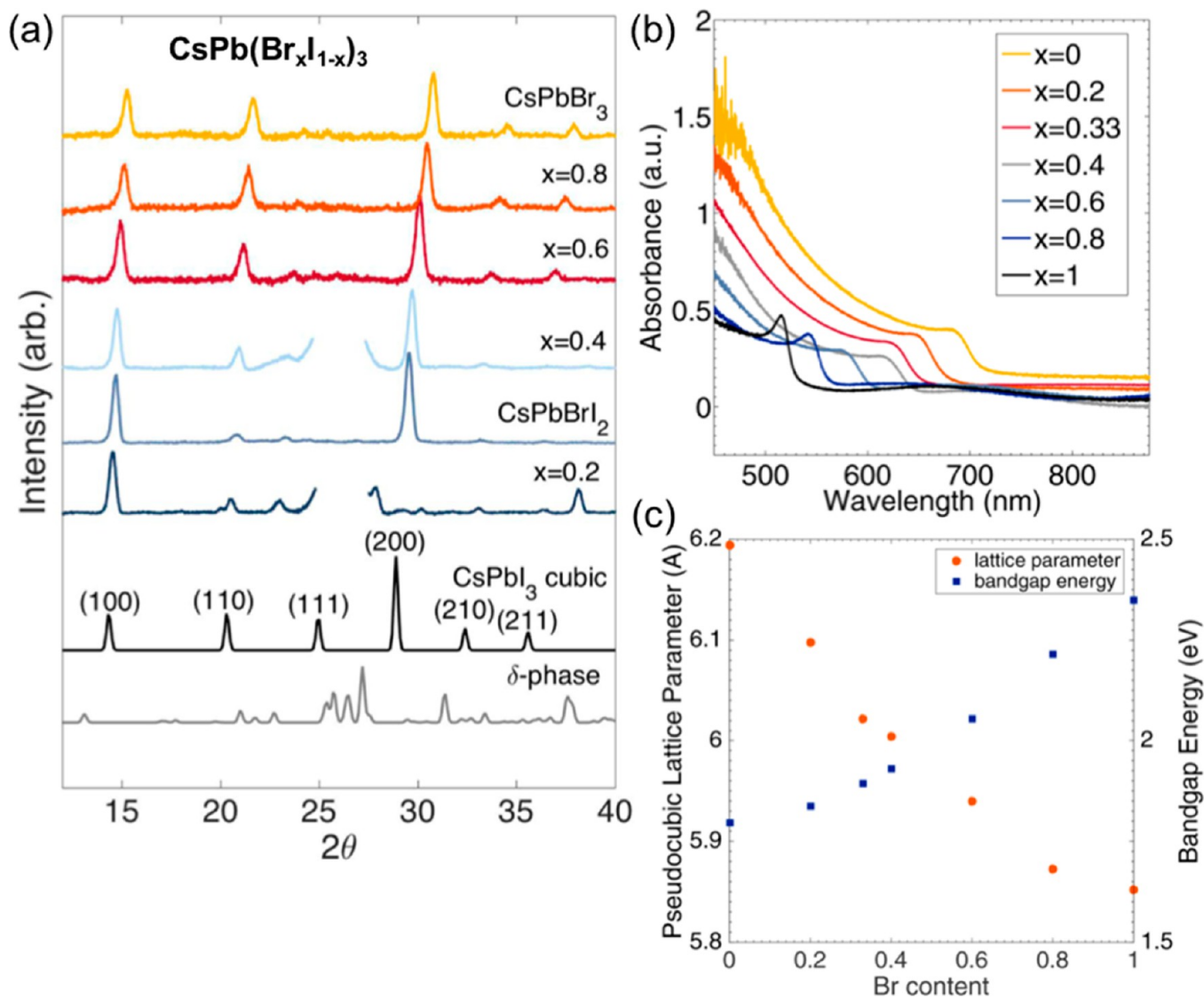


Figure 19. (a) XRD and (b) UV–vis absorption spectra for $\text{CsPb}(\text{Br}_x\text{I}_{1-x})_3$ ($0 \leq x \leq 1$) system. (c) Summary of lattice parameter changes based on the shift of the (100) peak and corresponding changes in the band gap energy extracted from the absorption onset. Reprinted with permission from ref 257. Copyright 2016 American Chemical Society.

967 70 mm (total active area = 25.2 cm²) generated $J_{sc} = 53.5$ mA,
 968 $V_{oc} = 9.05$ V (average per each cell is 1.06 V), FF = 74.4%,
 969 overall PCE = 14.3%.²⁵⁶

970 3.5. Binary $\text{CsPb}(\text{I}/\text{Br})$, $\text{CsPb}(\text{Br}/\text{Cl})$, and $\text{CsPb}(\text{I}/\text{Cl})$

971 **Systems.** Recent reports bring additional relevant insights
 972 into the $\text{CsPb}(\text{I}/\text{Br})$ system (Figure 19).^{98,250,257,258} For
 973 example, (i) contrary to the $\text{MAPb}(\text{I}/\text{Br})$ and $\text{FAPb}(\text{I}/\text{Br})$
 974 systems (section 3.2), halide phase-segregation into iodide- and
 975 bromide-rich domains (or Hoke effect) was observed to be
 976 minimized for $\text{CsPb}(\text{Br}_x\text{I}_{1-x})_3$ films with low Br concentrations
 977 ($x < 0.4$).²⁵⁷ However, for $x > 0.4$, phase separation was
 978 reported to occur under illumination (~ 1 sun). (ii) The
 979 CsPbI_2Br in particular, was studied in more details due to its
 980 suitable band gap energy for photovoltaic applications.
 981 CsPbI_2Br showed enhanced thermal stability when examined
 982 at 200 °C in an inert atmosphere²⁵⁷ and under solar cell
 983 operation conditions (85 °C in air).⁹⁸ (iii) The addition of
 984 bromide reduces the phase transition temperature. Good
 985 thermal stability of CsPbI_2Br alloy was also demonstrated at
 986 200 °C in inert atmosphere.²⁵⁸ The perovskite crystal forms at

987 ~ 350 °C for CsPbI_3 and ~ 230 °C for CsPbI_2Br , meaning a
 988 reduction of over 100 °C in the transition temperature.⁹⁸
 989 Device configuration of FTO/c-TiO₂/mp-TiO₂/ CsPbI_2Br -
 990 (~ 150 nm)/spiro-MeOTAD/Ag was demonstrated to generate
 991 the best efficiency of 9.84%, with $J_{sc} = 11.89$ mA/cm², $V_{oc} =$
 992 1.11 V, and FF = 75%.⁹⁸ Devices with an inverted structure,
 993 ITO/PEDOT:PSS/ CsPbI_2Br (~ 150 nm)/PC60BM/BCP/Al,
 994 were shown to generate a PCE of 6.80%, $J_{sc} = 10.9$ mA/cm²,
 995 $V_{oc} = 1.06$ V, and FF = $\text{PCE}/(J_{sc} \times V_{oc}) \sim 59\%$.²⁵⁷
 996 Yin et al.²⁵⁰ performed DFT calculations together with
 997 Monte Carlo simulations to systematically study the structural,
 998 electronic and energetic properties of mixed halide $\text{CsPb}(\text{I}/\text{Br})$,
 999 $\text{CsPb}(\text{Br}/\text{Cl})$, and $\text{CsPb}(\text{I}/\text{Cl})$ systems. They found that these
 1000 perovskites exhibit anomalous alloy properties that differ from
 1001 the conventional semiconductors (e.g., $\text{GaAs}_{1-x}\text{Sb}_x$). Generally,
 1002 in conventional isovalent semiconductors, the formation
 1003 energies are positive mostly attributed to the strain energy
 1004 (e.g., original ideal A and B bulk materials are stretched or
 1005 compressed to a new lattice constant of alloy A_{1-x}B_x). Halide
 1006 perovskites have strong ionic character, so Coulomb

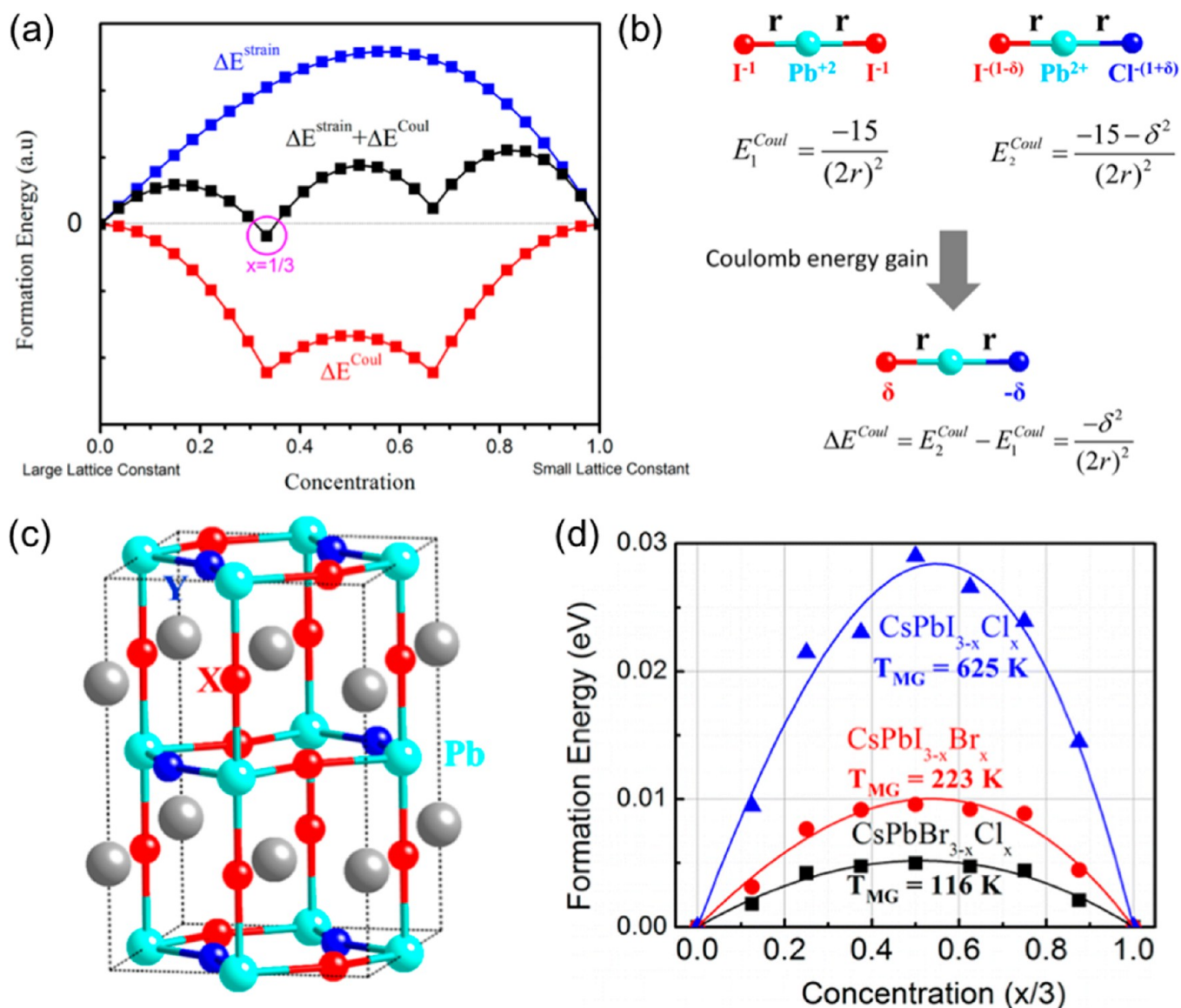


Figure 20. (a) Joint effects of strain (favoring demixing) and Coulomb (favoring mixing) energies for describing the formation energies of CsPb(X_{1-x}Y_x)₃ (X, Y = I, Br, Cl). This purely physical models lead the lowest energy at $x = 1/3$. (b) Model system (I–Pb–I trimer) shows the Coulomb energy gain process by charge transfer when halides are mixed. (c) Particular stable mixed-halide configuration for CsPbX_{3-x}Y_x (X, Y = I, Br, Cl; atomic size of X > Y) at $x = 1$. (d) Formation energies (per halogen atom) of CsPb(X_{1-x}Y_x)₃ (X, Y = I, Br, Cl) alloys calculated based on special-quasirandom-structures (scattered points). The solid lines are fitting based on the nine different points per halogen atom. The miscibility gap temperatures (T_{MG}) are indicated. Reprinted with permission from ref 250. Copyright 2014 American Chemical Society.

interactions play a major role. The Coulomb energy gain (Figure 20a,b) was identified to lower the overall formation energy (i.e., lower the strain energy) favoring certain stable ordered structures (Figure 20c). The authors gave a comprehensive example of a rigid ionic I–Pb–I trimer model (Figure 20b), where Pb and I have nominal 2+ and 1– oxidation states, respectively. In this case, Pb has Coulomb attraction with two I ions and two I ions have Coulomb repulsion between them, which equilibrates the entire system. Now, when one ion is replaced by a Cl ion, there will be electron transfer between I and Cl because the two ions have different electronegativities. Cl will be $(1 + \delta)^-$ charged and I will be $(1 - \delta)^-$ charged. As consequence, the total Coulomb energy (before and after) will be lowered by $\delta^2/(2r)^2$ (Figure 20b). The schematic shapes of Coulomb energy gain are shown in Figure 20a. The alloy system has the largest Coulomb energy

gain at $x = 1/3$. As described at the beginning of section 3.5, experimental reports by Sutton et al.⁹⁸ and Beal et al.²⁵⁷ seem to confirm this prediction, where CsPbI₂Br was demonstrated to show enhanced thermal and air stability. Yin et al.²⁵⁰ calculated the miscibility gap temperature (T_{MG}), which is defined as critical temperature that alloys can be fully mixed based on special-quasirandom-structures (Figure 20d). TMGs of 625, 223, and 116 K were extracted for CsPbI_{3-x}Cl_x, CsPbI_{3-x}Br_x, and CsPbBr_{3-x}Cl_x, respectively. These indicate that mixed-(Br/I) and (Br/Cl) are easily formed, whereas mixed-(Cl/I) perovskites are difficult to be formed.

4. SIMULTANEOUS MIXED A-CATIONS AND MIXED X-HALIDE ANIONS

In the following sections, we continue to describe the more complex mixed-perovskite systems where the double (MA/

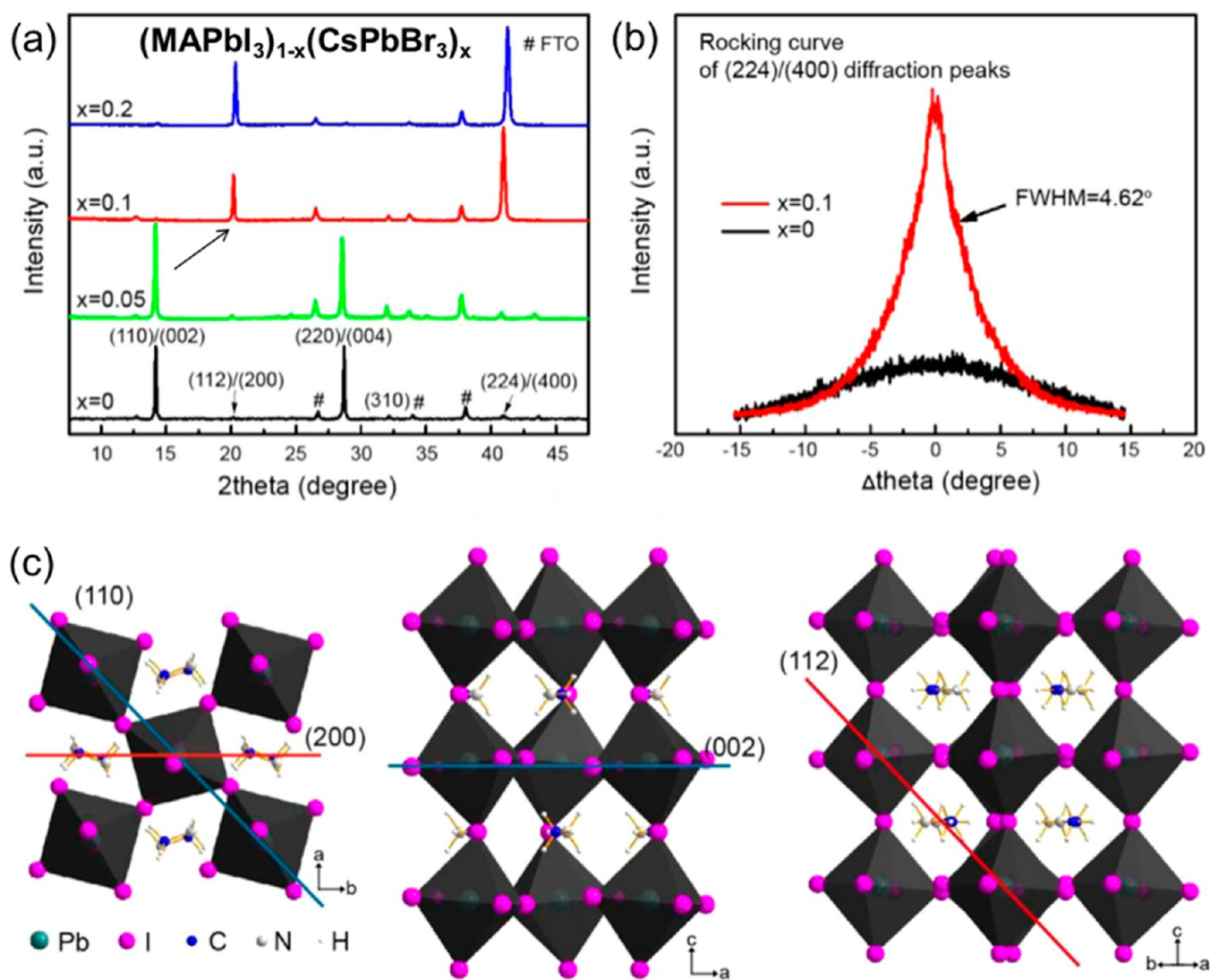


Figure 21. (a) XRD patterns of $(\text{MAPbI}_3)_{1-x}(\text{CsPbBr}_3)_x$ with x ranging from 0 to 0.2. (b) Rocking curve measurement of (224)/(400) diffraction peaks for $x = 0$ and 0.1. (c) Schematic illustrations of (110), (002), (112), and (200) perovskite crystal planes. Reprinted with permission from ref 123. Copyright 2016 Elsevier Ltd.

1037 FA),^{63,64,90,125–129,136,259} (Cs/MA) ,¹²³ $(\text{Cs}/$
 1038 $\text{FA})$,^{117,132–134,260,261} triple $(\text{Cs}/\text{FA}/\text{MA})$,^{65,135,262} and even
 1039 quadrupole $(\text{Rb}/\text{Cs}/\text{FA}/\text{MA})$ mixed cations^{7,137} are generally
 1040 followed by the mixed (I/Br), e.g., $(\text{Rb}/\text{Cs}/\text{FA}/\text{MA})\text{Pb}(\text{I}/\text{Br})$.
 1041 These newly developed material systems were reported to show
 1042 even higher stability¹³⁷ and efficiencies as shown in the plot of
 1043 certified efficiencies (Figure 1b).

1044 **4.1. (MA/Cs)Pb(I/Br) System.** Niu et al.¹²³ reported for the
 1045 first time on the synthesis of $(\text{MAPbI}_3)_{1-x}(\text{CsPbBr}_3)_x$ by a one-
 1046 step method dissolving MAI, PbI_2 , CsBr, PbBr_2 precursors in a
 1047 mixed solvent of γ -butyrolactone:DMSO = 7:3 vol %. The
 1048 fabricated solar cells exhibited an optimal performance at $x =$
 1049 0.1 with an average PCE of $15.9 \pm 0.52\%$ (Table 1). The best
 1050 cell showed $J_{sc} = 22.8 \text{ mA}/\text{cm}^2$, $V_{oc} = 1.05 \text{ V}$, FF = 73%, and
 1051 PCE = 17.6%. The stabilized power output (power output
 1052 measurement over time until a steady value is reached) of the
 1053 champion cell exhibited a PCE of 15.7% and photocurrent of
 1054 $19.6 \text{ mA}/\text{cm}^2$, which means hysteresis is present. Accelerated
 1055 stability of $(\text{MAPbI}_3)_{1-x}(\text{CsPbBr}_3)_x$ -based solar cells was
 1056 evaluated under UV irradiation (365 nm, $364 \text{ mW}/\text{cm}^2$)
 1057 without encapsulation. It was mentioned that as x increased, the

stability was significantly increased. For $(\text{MAPbI}_3)_{0.9}(\text{CsPbBr}_3)_{0.1}$, 75% of the initial performance was maintained after a total of 150 min irradiation. The same authors performed shelf life stability tests of unencapsulated $(\text{MAPbI}_3)_{0.9}(\text{CsPbBr}_3)_{0.1}$ devices in ambient air (dark, 25 °C, 20–30% RH) and observed that nearly 80% of the initial performance was maintained after 500 h.

The same work by Niu et al.¹²³ touches upon an interesting topic whether the photovoltaic efficiency is dependent on the facet-orientation in single grains of perovskites. This topic was discussed previously by Leblebici et al.²⁶³ Generally, MAPbI_3 and $\text{MAPbI}_3-x\text{Cl}_x$ systems exhibit preferred orientation along $\langle 110 \rangle$ and/or $\langle 002 \rangle$ directions (see XRD pattern corresponding to $x = 0$ in Figure 21a). Interestingly, the (112) and (200) facets increased in intensity substantially with $(\text{CsPbBr}_3)_x$ ($x = 0.1$ and 0.2) concentration (Figure 21a). The orientation of the film can be further verified by rocking curve measurements (Figure 21b), which provides information on the existence of preferential growth axis and its spread on the azimuthal angle. For $x = 0.1$, (224)/(400) planes were parallel to the substrate with an angular spread of $\pm 4.6^\circ$; meanwhile, for $x = 0$, no peaks

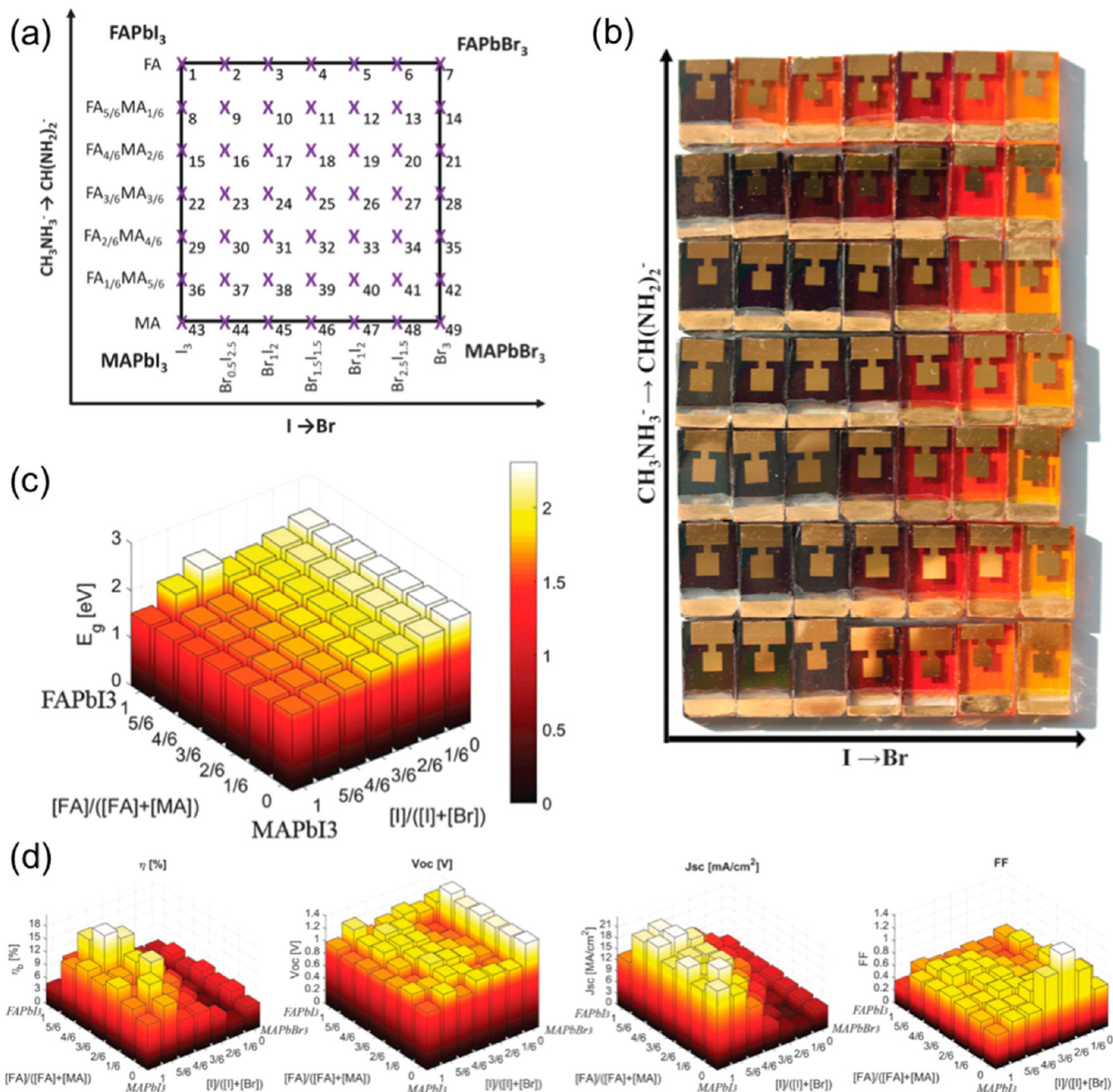


Figure 22. (a) Schematic illustration of the compositional matrix and (b) photo of fabricated cells of (FA/MA)Pb(I/Br) system. A total of 49 compositions were mapped. (c) Bar plot of the band gaps in the compositional space explored. (d) Solar cell device parameters for (FA/MA)Pb(I/Br)-based perovskite with different compositions. Reprinted with permission from ref 90. Copyright 2016 The Royal Society of Chemistry.

1079 were observed. When x is increased to 0.3 and above, new
 1080 peaks assigned to CsPb_2Br_5 were observed in XRD, which
 1081 indicates $x = 0.2$ is a threshold limit for the full-miscibility in
 1082 $(\text{MAPbI}_3)_{1-x}(\text{CsPbBr}_3)_x$. The origin for this preferential
 1083 orientational growth was further examined employing XPS
 1084 depth profile and DFT. Results from these measurements
 1085 showed a heavily Cs-incorporated perovskite preferentially
 1086 forms at the bottom layer. Based on DFT calculations of the
 1087 CsPbBr_3 structure, it was shown that (100) ($\gamma = 0.299 \text{ J}/\text{m}^2$)
 1088 and (112) ($\gamma = 0.314 \text{ J}/\text{m}^2$) planes showed a lower surface
 1089 energy than the (110) ($\gamma = 0.437 \text{ J}/\text{m}^2$) and (001) ($\gamma = 0.526$
 1090 J/m^2) planes. Therefore, it was concluded that the heavily Cs-
 1091 incorporated perovskite at the bottom of the film could initiate

the film growth along the $\langle 112 \rangle / \langle 200 \rangle$ directions under
 thermodynamic control (Figure 21c).

4.2. (FA/MA)Pb(I/Br) System. After the initial works by
 Jeon et al.⁶³ and Yang et al.,⁶⁴ the simultaneous FA/MA-cation
 mixed and I/Br-halide mixed perovskites are one of the most
 studied ternary mixed perovskite system for photovoltaic
 applications (Table 1).^{7,90,124–130,259,262,264–266} Out of six,
 three certified efficiencies reported by NREL is based on this
 perovskite system (Figure 1b). Jacobsson et al.⁹⁰ provided the
 entire compositional space of MA-FA–Br-I experimentally
 (Figure 22). In this study, a total of 49 precursor solutions with
 varying concentrations of PbI_2 , PbBr_2 , MAI, MABr, FAI, and
 FABr were used for perovskite synthesis (Figure 22a). The

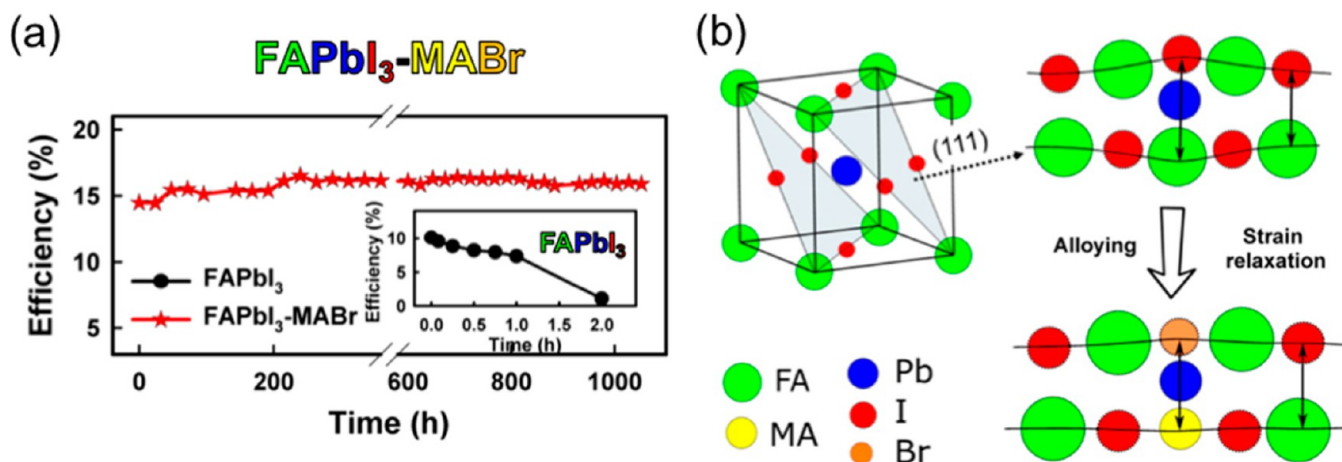


Figure 23. (a) Comparative PCEs of FAPbI₃-MABr- and FAPbI₃-based perovskite solar cells tested under storage conditions in air (~50% RH, 23 °C) without encapsulation. (b) Schematic illustration of strain relaxation when MABr is alloyed with FAPbI₃. Reprinted with permission from ref 125. Copyright 2016 American Chemical Society.

1105 films were systematically characterized by UV-vis, PL, XRD
1106 techniques. Differences in cation and halide compositions lead
1107 to large variation in the optical appearance of deposited films
1108 (Figure 22b,c). The distribution of extracted bandgaps in the
1109 compositional matrix (Figure 22c) was fitted with an empirical
1110 relation to provide a numerical description of the dependence
1111 of band gap with respect to MA/FA and Br/I compositions (eq
1112 4).

$$E_g(x, y) = 1.58 + 0.436x - 0.058y + 0.294x^2 + 0.0199x \cdot y \quad (4)$$

$$x = \frac{[\text{Br}]}{[\text{Br}] + [\text{I}]} \text{ and } y = \frac{[\text{FA}]}{[\text{FA}] + [\text{MA}]}$$

1114 Similar to the description in section 3.2, PL measurements
1115 revealed also the appearance of more than one peak for films
1116 with 50% bromide or more, excluding the pure bromide
1117 perovskites. The authors proposed two explanations, (i)
1118 perovskites with a high bromide concentration increase the
1119 density of trap states within the band gap that can act as
1120 recombination centers and (ii) phase separation with the
1121 formation of an iodine-rich domains leads to the peak at lower
1122 photon energies in PL (section 3.2). XRD measurements
1123 provide valuable experimental observations: (i) iodine-rich pure
1124 FA-samples (numbers 2 and 3 in Figure 22a,b) deviate from
1125 others and the dominant phase is not the photoactive
1126 perovskites; (ii) Disregarding these two compositions and the
1127 FAPbI₃ composition (due to yellow phase; number 1 in Figure
1128 22a,b), all other intended perovskites were formed and were
1129 the dominant phases; (iii) for compositions 1 to 38 (FA-rich
1130 devices), the experimental XRD data are in line with a cubic
1131 structure; (iv) it has been proposed that the transition border
1132 between room temperature tetragonal and cubic structures is
1133 found around FA_{1/6}MA_{5/6}PbBr₁I₂ (number 38). Full solar cell
1134 devices made with all 49 different compositions showed a
1135 widespread with efficiencies varying from 2.3% to 20.67%
1136 (Figure 22d). FA_{4/6}MA_{2/6}Pb(Br_{1/6}I_{5/6})₃ (number 16) yielded
1137 the best devices with a top efficiency of 20.67% ($J_{sc} = 23.7$ mA/
1138 cm², $V_{oc} = 1.14$ V, FF = 76%, Table 1). This is in agreement
1139 with other reports reporting high efficiencies around this
1140 composition.^{63,130} Hysteresis was also observed to vary as

mapping the compositional space. Interestingly, a trend can be
observed that hysteresis is small for devices with a high fraction
of both iodide and FA, and is higher for bromide- and MA-rich
devices.³⁰ Ion migration on time scales from 10⁻¹ to 10² s was
previously reported to explain hysteresis.^{7,259} Several reports
showed that regardless of solar cell architectures and
composition in perovskites, halide defects migrate and can
reversibly accumulate at the interfaces of selective contacts.
Furthermore, among several types of defects, iodide or bromide
vacancy generation has been calculated to show the lowest
formation energies (e.g., defects formed during sample
preparation, under bias, etc.), with further bromide vacancies
being favored over iodide.^{7,259} Tress et al.²⁶⁶ reported that
inverted hysteresis (forward bias scan generates higher PCE
than backward scan) is particularly pronounced in (FA/
MA)Pb(I/Br) systems. Zheng et al.¹²⁵ fabricated MAPbBr₃-
alloyed FAPbI₃ perovskites and solar cells based on these
materials showed enhanced shelf life stability of more than
1000 h under ambient air (~50% RH, 23 °C) without
encapsulation (Figure 23a). Based on XRD, the authors
proposed a model that α -FAPbI₃ has an anisotropic strained
lattice; higher strain in the (111) plane. In contrast, δ -FAPbI₃
is almost strain free explaining the favored $\alpha \rightarrow \delta$ phase transition
at room temperature (section 2.1). They proposed that the
strain in the (111) plane of α -phase is the driving force for the
phase transition to the δ -phase (Figure 23b). When MABr is
incorporated to alloy with FAPbI₃, the lattice size is reduced
and strain within the grain is relaxed. In this way, the
pseudocubic α -phase is stabilized at room temperature and
even under humid air. In a recent report, Domanski et al.²⁵⁹
investigated the impacts of cation defects formation and
migration in the (FAPbI₃)_{1-x}(MAPbBr₃)_x perovskite on its
corresponding solar cell performance and long-term stability
under operation conditions. Maximum power output was
monitored for three identically prepared devices (A, B, C in
Figure 24a). Devices A and B were continuously tracked over
100 h. These two devices exhibit very different profiles from
that of device A showing high instability than that of device B.
Because long-term stability is reflected from a convolution of
several mechanisms (physicochemical and electrical) taking
place within the perovskite solar cell, it is not surprising that
identically prepared devices age differently. However, the
authors identified that the initial decay in performance

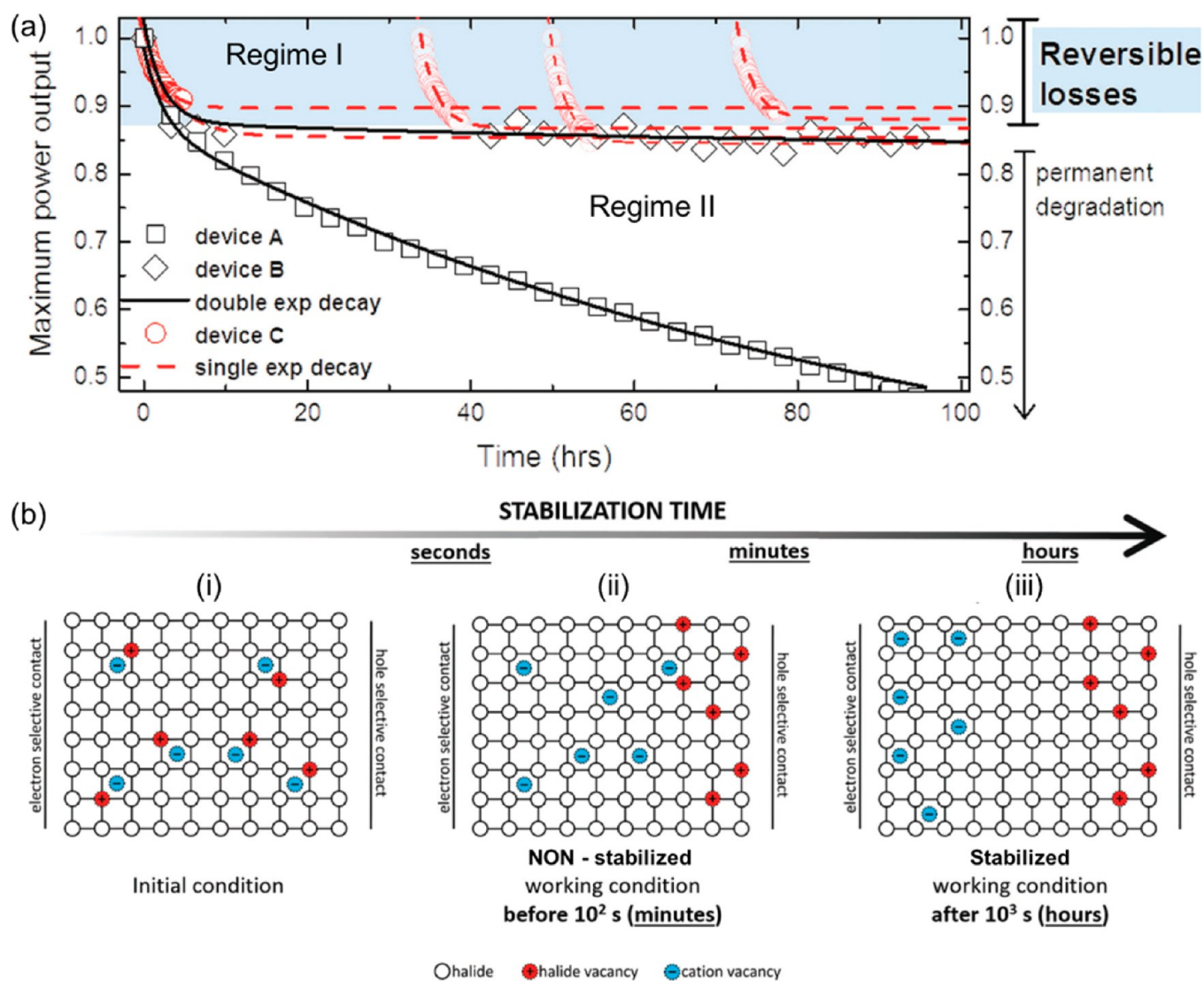


Figure 24. (a) Normalized maximum power output ($V \sim 0.85$ V) tested in N_2 environment for 3 identically prepared perovskite solar cells (devices A, B, C) measured under 1 sun. Devices A and B were continuously tracked for over 100 h. Device C was cyclically tracked 4 times for 5 h and it was left in dark at open circuit between consecutive measurements. (b) Schematic illustration showing the evolution of the ion distribution within the perovskite layer sandwiched between electron and hole selective contacts under working conditions on time scales from seconds to hours (i–iii). Reprinted with permission from ref 259. Copyright 2017 The Royal Society of Chemistry.

designated as regime I (Figure 24a) was unchanged. To isolate regime I from the subsequent degradation (regime II), the maximum power point for device C was stopped after only 5 h and repeated cyclically after resting the devices in dark for a different number of hours. Interestingly, the initial power output at each cycle was similar to the previous cycle, demonstrating that the initial performance losses are fully reversible (regime I) and can be separated from the subsequent permanent degradation (regime II). Degradation related to regime II (HTL, perovskite, top Au electrode related issues) was previously reported.^{65,267,268} Followed by a subsequent series of experiments, the authors concluded that migration of cation vacancies form an additional Debye layer at the interface with the electron selective contact, which inhibits charge extraction (Figure 24b). However, when the device is given several hours of rest in dark, the ionic distribution equilibrates to the initial state, which leads to recovery of the initial performance (Regime I, Figure 24b). Migration of cation vacancies was shown to exhibit significantly longer time scales

(>10³ s) than the halide vacancy migration (between 10⁻¹ and 10² s, Figure 24b).

4.3. (FA/MA)Pb(I/Cl) System. Mixed MA/FA cations and I/Cl halides perovskites with a chemical composition of $MA_{1-x}FA_xPb(I_{1-y}Cl_y)_3$ were first synthesized by Isikgor et al.¹³¹ These perovskite films were prepared by dissolving PbI_2 + $PbCl_2$ and MAI + FAI precursors in a cosolvent of GBL/DMSO (3:7 vol. ratio) and low postannealing temperature of 80–110 °C to avoid chlorine species sublimate and/or decompose in the form of MAI (described in section 3.1). A planar heterojunction perovskite solar cell based on an inverted structure with ITO/PEDOT:PSS/MA_{0.80}FA_{0.20}PbI_{3-y}Cl_y/PC61BM/C60/LiF/Ag generated average $J_{sc} = 21.55 \pm 0.55$ mA/cm², $V_{oc} = 1.10 \pm 0.01$ V, FF = 75 ± 2%, and PCE = 17.45% (Table 1). This performance employing MA_{0.80}FA_{0.20}PbI_{3-y}Cl_y ($E_g = 1.58$ eV) outperformed the control perovskite solar cells (MAPbI₃, MAPbI_{3-y}Cl_y, MAPbI_{3-y}Br_y, and MA_{1-x}FA_xPbI₃). The high photovoltaic performance was attributed mainly to (i) the long charge diffusion length

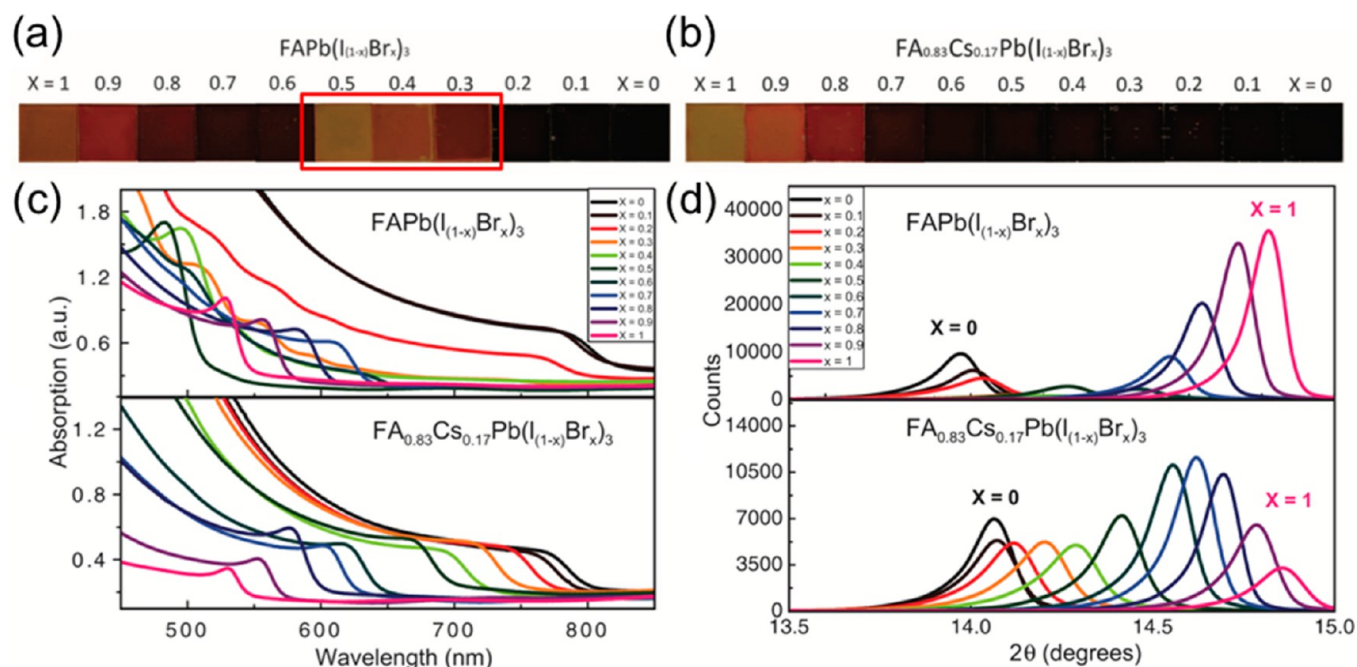


Figure 25. Photographs of (a) $\text{FAPb}(\text{I}_{1-x}\text{Br}_x)_3$ and (b) $\text{FA}_{0.83}\text{Cs}_{0.17}\text{Pb}(\text{I}_{1-x}\text{Br}_x)_3$ perovskite films with Br composition increasing from $x = 0$ to 1. (c) UV-vis absorbance spectra and (d) XRD patterns of $\text{FAPb}(\text{I}_{1-x}\text{Br}_x)_3$ and $\text{FA}_{0.83}\text{Cs}_{0.17}\text{Pb}(\text{I}_{1-x}\text{Br}_x)_3$ perovskites. Reprinted with permission from ref 132. Copyright 2016 American Association for the Advancement of Science (AAAS).

induced by mixed organic cations and mixed halides and (ii) suppression of the formation of undesirable yellow δ -phase of FAPbI_3 ($E_g = 2.8$ eV).

4.4. (FA/Cs)Pb(I/Br) System. With the aim to fabricate perovskite/silicon tandem solar cells, a number of reports synthesized mixed-cations and mixed-halides (i) to achieve high stability and (ii) to tune the band gap to an optimal value of ~ 1.75 eV in order to current-match the top perovskite cell and bottom Si cell ($E_g = 1.1$ eV).^{132,133,269} Based on the observations reported previously for binary mixed perovskites, (i) $\text{MAPb}(\text{I}/\text{Br})$ (section 3.2) is thermally unstable and suffers from halide segregation; (ii) $\text{FAPb}(\text{I}/\text{Br})$ shows better stability regarding halide segregation (section 3.2) and Br concentration helps tune the band gap to a desired value (~ 1.75 eV); (iii) $\text{FAPb}(\text{I}_{1-x}\text{Br}_x)_3$ with $0.3 < x < 0.6$, an amorphous unstable phase is formed caused by a transition from a trigonal ($x < 0.3$) and cubic ($x > 0.5$) (Figure 25a); therefore, this composition range should be avoided (section 3.2); (iv) mixing small quantities of Cs with FA substantially enhances the stability of $\text{Cs}_x\text{FA}_{1-x}\text{PbI}_3$ suppressing halide segregation (section 2.3); (iv) the $\delta \rightarrow \alpha$ phase transformation in FAPbI_3 can be lowered down to room temperature when Cs/FA ratio of 45 at. % is incorporated in the $\text{Cs}_x\text{FA}_{1-x}\text{PbI}_3$ (section 2.3), McMeekin et al.¹³² proposed the first $\text{FA}_y\text{Cs}_{1-y}\text{Pb}(\text{I}_{1-x}\text{Br}_x)_3$ system based on FAI , CsI , PbBr_2 , and PbI_2 precursors dissolved in DMF followed by addition of HI and HBr. Initially, they hypothesized that if FA is partially substituted by Cs, the structural instability in Br-to-I phase space could be shifted to higher energies, and thus achieve a structurally stable mixed halide perovskite with a band gap of 1.75 eV. Their trial with $\text{FA}_{0.83}\text{Cs}_{0.17}\text{Pb}(\text{I}_{1-x}\text{Br}_x)_3$ showed unexpected results that the region of structural instability was not observed (Figure 25b). Instead, a continuous series of dark films throughout the entire composition range ($0 \leq x \leq 1$) was observed and corroborated by UV-vis (Figure 25c) and XRD (Figure 25d) measurements. Over the entire of Br-to-I range, $\text{FA}_{0.83}\text{Cs}_{0.17}\text{Pb}(\text{I}_{0.6}\text{Br}_{0.4})_3$

composition was chosen based on the Vegard's law (section 3.2) leading to the band gap of 1.74 eV. These perovskites showed further improved structural stability and resistance to halide segregation when compared to $\text{MAPb}(\text{I}_{0.6}\text{Br}_{0.4})_3$ (section 3.2). Furthermore, under thermal stress conditions at 130°C , $\text{FA}_{0.83}\text{Cs}_{0.17}\text{Pb}(\text{I}_{1-x}\text{Br}_x)_3$ showed superior stability compared to $\text{MAPb}(\text{I}_{0.6}\text{Br}_{0.4})_3$.¹³² Planar heterojunction perovskite solar cell based on an inverted structure with $\text{FTO}/\text{SnO}_2/\text{PC60BM}/\text{FA}_{0.83}\text{Cs}_{0.17}\text{Pb}(\text{I}_{0.6}\text{Br}_{0.4})_3/\text{spiro-MeOTAD}/\text{Ag}$ generated $J_{sc} = 19.4$ mA/cm^2 , $V_{oc} = 1.2$ V, $\text{FF} = 75.1\%$, and $\text{PCE} = 17.1\%$ (Table 1). When combined with a 19% PCE c-Si, the feasibility of achieving $>25\%$ PC four-terminal tandem cells were demonstrated.¹³² Employing the same $\text{FA}_{0.83}\text{Cs}_{0.17}\text{Pb}(\text{I}_{0.6}\text{Br}_{0.4})_3$ perovskite, Busch et al.¹³³ demonstrated a two-terminal 1 cm^2 active area perovskite/tandem solar cell with $J_{sc} = 18.1$ mA/cm^2 , $V_{oc} = 1.65$ V, $\text{FF} = 79\%$, and $\text{PCE} = 23.6\%$. More recently, Zhang et al.²⁶⁹ performed DFT calculations determining three sets of x and y for $\text{FA}_y\text{Cs}_{1-y}\text{Pb}(\text{I}_{1-x}\text{Br}_x)_3$ systems, which have band gap around 1.75 eV. In addition, their refractive indices and extinction coefficients were also calculated. They found that $\text{FA}_{0.89}\text{Cs}_{0.11}\text{Pb}(\text{I}_{0.56}\text{Br}_{0.44})_3/\text{c-Si}$ tandem solar cells achieved the highest PCE among the three sets. The concept of perovskite-perovskite tandem cell employing $\text{FA}_{0.75}\text{Cs}_{0.25}\text{Sn}_{0.5}\text{Pb}_{0.5}\text{I}_3/\text{FA}_{0.83}\text{Cs}_{0.17}\text{Pb}(\text{I}_{0.6}\text{Br}_{0.4})_3$ with band gaps of 1.2 and 1.74 eV, respectively, was demonstrated to generate PCE as high as 20.3% in a four-terminal configuration.

Perovskite single-junction solar cells with regular structure (TiO_2 electron transport layer (ETL) and spiro-MeOTAD HTL) based on similar $\text{FA}_{0.8}\text{Cs}_{0.2}\text{PbI}_{2.84}\text{Br}_{0.16}$ composition, generated $J_{sc} = 23.3$ mA/cm^2 , $V_{oc} = 1.072$ V, $\text{FF} = 72.3\%$, and $\text{PCE} = 18.02\%$. A wider parameter space of the influences of mixed-cation and mixed-halide perovskites in $\text{Cs}_y\text{FA}_{1-y}\text{Pb}(\text{I}_{1-x}\text{Br}_x)_3$ was studied by Rehman et al.²⁶¹ They showed that a region for Cs concentration between $0.10 < y < 0.30$ leads to high crystalline quality, long charge-carrier lifetimes, and high charge-carrier mobilities. Within the $\text{Cs}_y\text{FA}_{1-y}\text{Pb}(\text{I}_{0.4}\text{Br}_{0.6})_3$

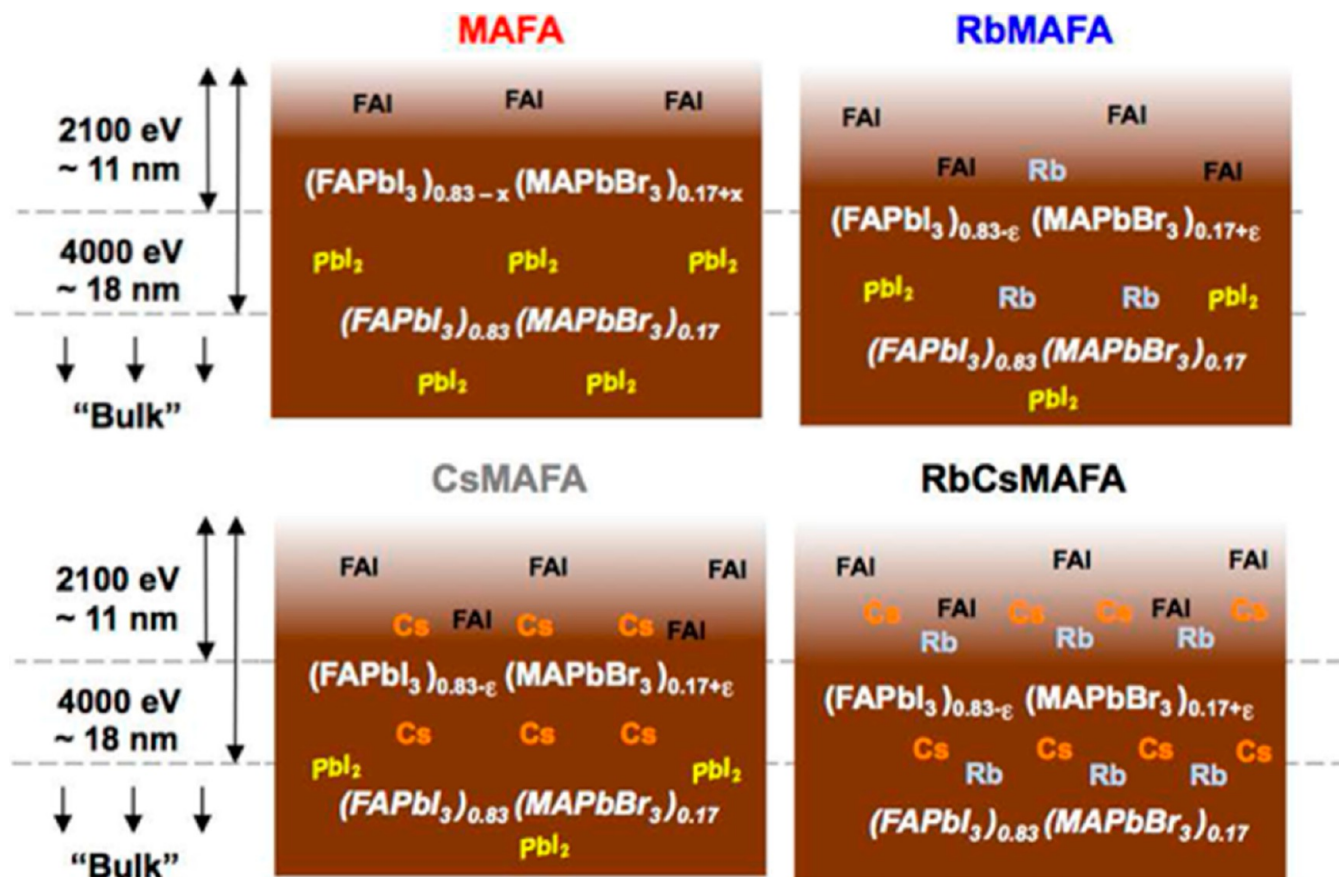


Figure 26. Schematic illustration of a depth-dependent chemical composition in MAFA = $(\text{MA}_{0.17}\text{FA}_{0.83})\text{Pb}(\text{I}_{0.83}\text{Br}_{0.17})_3, 0.1[0.83\text{PbI}_2, 0.17\text{PbBr}_2]_{0.95}$; RbMAFA = $0.05\text{RbI}[(\text{MA}_{0.17}\text{FA}_{0.83})\text{Pb}(\text{I}_{0.83}\text{Br}_{0.17})_3, 0.1[0.83\text{PbI}_2, 0.17\text{PbBr}_2]_{0.95}]_{0.95}$; CsMAFA = $0.05\text{CsI}[(\text{MA}_{0.17}\text{FA}_{0.83})\text{Pb}(\text{I}_{0.83}\text{Br}_{0.17})_3, 0.1[0.83\text{PbI}_2, 0.17\text{PbBr}_2]_{0.95}]_{0.95}$; RbCsMAFA = $0.05\text{RbI}[0.05\text{CsI}[(\text{MA}_{0.17}\text{FA}_{0.83})\text{Pb}(\text{I}_{0.83}\text{Br}_{0.17})_3, 0.1[0.83\text{PbI}_2, 0.17\text{PbBr}_2]_{0.95}]_{0.95}]_{0.95}$. The chemical formula are based on the precursor solution composition. A nonstoichiometric solution was used in excess of 10 mol % lead precursors leading to the nominal composition of MAFA stated above. Reprinted with permission from ref 270. Copyright 2017 American Chemical Society.

series, $y = 0.2$ showed the concentration leading the perovskite film exhibiting high charge-carrier mobilities of $\sim 18 \text{ cm}^2 \text{V}^{-1} \text{s}^{-1}$ and charge-carrier lifetimes of $\sim 80 \text{ ns}$. They demonstrated a correlation between high crystallinity and suppressed photo-induced halide segregation, i.e., short-range crystalline order or the presence of grain boundaries enables halide segregation by releasing lattice strain energy leading to iodide-rich and bromide-rich domains (section 3.2). Next, the halide-parameter space was investigated for $\text{Cs}_{0.17}\text{FA}_{0.83}\text{Pb}(\text{I}_{1-x}\text{Br}_x)_3$. They showed that once the perovskite contains enough fraction of Cs for the stability, high charge-carrier mobilities and diffusion lengths are obtained across the full I:Br range. Furthermore, the band gap varies linearly with the Vegard's law (section 3.2) across the full I:Br range.²⁶¹ Nonencapsulated $\text{Cs}_{0.17}\text{FA}_{0.83}\text{Pb}(\text{I}_{0.6}\text{Br}_{0.4})$ -based perovskite solar cells employing *n*-doped C60 as electron extraction layer in a planar heterojunction architecture, showed a 30-fold enhanced air stability compared to $\text{MAPbI}_{3-x}\text{Cl}_x$ -based devices under full spectrum solar illumination without encapsulation ($t_{80} \sim 600 \text{ h}$); and $t_{80} > 3420 \text{ h}$ when sealed.¹³⁴

4.5. (Cs/FA/MA)Pb(I/Br) System. Based on previous successful reports of Cs incorporation in $(\text{MA}/\text{Cs})\text{PbI}_3$ and $(\text{FA}/\text{Cs})\text{PbI}_3$ systems (sections 2.2 and 2.3), Saliba et al.⁶⁵ (a follow up work from the same group was reported by Matsui et al.¹³⁵) reported on the first triple cation mixed (Cs/FA/MA)Pb(I/Br) system generating stabilized power output of 21.1% that also holds a position in the certified NREL chart

(Figure 1b). The stability of these devices was tested in a nitrogen atmosphere held at room temperature under constant illumination and maximum power tracking for 250 h. The device efficiencies of triple cation perovskites with FTO/c-TiO₂/mp-TiO₂/Cs_{0.05}(FA_{0.83}MA_{0.17})_{0.95}Pb(I_{0.83}Br_{0.17})₃/spiro-MeOTAD/Au structure decayed from $\sim 20\%$ to $\sim 18\%$ within a few hours (Regime I,²⁵⁹ section 4.2) and then stayed stable for at least 250 h.

4.6. (Rb/FA/MA)Pb(I/Br) System. Duong et al.¹³⁶ and Zhang et al.¹²² studied the incorporation of Rb in high performing $(\text{FAPbI}_3)_{1-x}(\text{MAPbBr}_3)_x$ mixed perovskites. The motivation for studying this system was that inclusion of a smaller cation of Cs in $(\text{FAPbI}_3)_{1-x}(\text{MAPbBr}_3)_x$ lead to enhanced stability (section 4.5). Therefore, the incorporation of an even smaller cation would be interesting to be tested. The addition of 5% RbI in combination with excess PbI₂ was observed to eliminate the formation of yellow nonperovskite phase and enhance the crystallinity of the films. However, inclusion of more than 10% RbI resulted in the formation of a Rb-rich phase, which was detrimental for the cell performance.¹³⁶ A side-by-side comparison of solar cell devices based on $\text{Rb}_{0.05}\text{FA}_{0.80}\text{MA}_{0.15}\text{PbI}_{2.55}\text{Br}_{0.45}$ and $\text{Cs}_{0.05}\text{FA}_{0.80}\text{MA}_{0.15}\text{PbI}_{2.55}\text{Br}_{0.45}$ showed very similar performances (PCE $\sim 19.5\%$, Table 1).

4.7. (Rb/Cs/FA)Pb(I/Br) System. Syzgantseva et al.¹⁵⁷ reported theoretical calculations on the stabilization of FAPbI₃ by Cs/Rb incorporations. Enhanced stabilization of mixed

perovskites was demonstrated when Cs⁺ and Rb⁺ are employed instead of MA⁺. In addition, the increased phase stability of Cs⁺ and Rb⁺ incorporated systems comes with only a slight increase in the band gap. Saliba et al.¹³⁷ fabricated systematically Rb_{0.05}Cs_{0.05}FA_{0.90}Pb(I_{0.83}Br_{0.17})₃⁻, Rb_{0.05}Cs_{0.10}FA_{0.85}Pb(I_{0.83}Br_{0.17})₃⁻, and Rb_{0.10}Cs_{0.05}FA_{0.85}Pb(I_{0.83}Br_{0.17})₃-based solar cells showing efficiencies as high as 18.3%, 19.3%, and 18.7%, respectively (Table 1).

4.8. (Rb/Cs/FA/MA)Pb(I/Br) System. Based on their past experience with binary and triply mixed cations (RbFA, RbCsFA, RbMAFA), Saliba et al.¹³⁷ were able to synthesize the first perovskite containing four simultaneous cations (RbCsFAMA). The concentration of Rb was limited to 5 at % as they were aware that a higher concentration of Rb will lead to the formation of a Rb-rich phase, which is detrimental for the performance (section 2.4). Devices with the architecture of FTO/c-TiO₂/mp-TiO₂/Rb_{0.05}Cs_{0.05}(FA_{0.83}MA_{0.17})_{0.90}Pb(I_{0.83}Br_{0.17})₃/PTAA/Au generated stabilized efficiencies of up to 21.6% (averaged PCE was 20.2%, Table 1). Furthermore, PTAA was employed as hole transport layer (HTL) and these cells maintained ~95% of their initial performance when tested in N₂ atmosphere, but at elevated temperature of 85 °C for 500 h and under operation conditions (full illumination and tracking the maximum power point). Similar high stability was also confirmed by Duong et al.¹³⁸ fabricating FTO/c-TiO₂/mp-TiO₂/Rb(5%) doped FA_{0.75}(MA_{0.6}Cs_{0.4})_{0.25}PbI₂Br/PTAA/Au, which generated stabilized power output of ~17.4% (Table 1). The long-term high stability under illumination indicates that phase segregation in iodide- and bromide-rich domains phenomena (section 3.2) is significantly suppressed upon this cation engineering strategy.^{137,138} The charge transport within the RbCsMAFA perovskite layer is substantially faster than in CsMAFA, which is already much faster than in MAFA leading to the conclusion that Rb incorporation leads to defect-free perovskites.¹³²

Although these studies seem to infer a uniform solid solution, there is no clear evidence supporting that these cations are incorporated into the perovskite crystal structure forming a uniform solid solution. Recently, Philippe et al.²⁷⁰ reported a depth-dependent chemical composition in (Rb/Cs/FA/MA)-Pb(I/Br) perovskite films employing hard X-ray photoelectron spectroscopy (HAXPES, synchrotron radiation source). These results were compared systematically with perovskite materials with two (MA/FA) and three cations (CsMAFA and RbMAFA) to investigate the role of Cs and Rb cations. Two photon energies were employed, with the photon energy of 2100 eV to probe the surface of the sample (~11 nm) and the photon energy of 4000 eV to probe deeper (~18 nm; referred as bulk chemical composition). Quantifications were conducted for both photon energies and the I/Pb, Cs/Pb, and Rb/Pb ratios and its interpretation were summarized in Figure 26. For example, the Br/Pb and I/Br ratios correspond to the proportion of FA and MA. In the case of MAFA sample that is generally prepared with excess of PbI₂, HAXPS (4000 eV) confirms that the excessive PbI₂ can be found in the bulk of the material. Upon addition of a third cation (Cs⁺ or Rb⁺), less PbI₂ remains in the material. When the probe depth decreases, an increase of Br over Pb was observed indicating a slight bromide enrichment within the surface as compared to bulk as illustrated with (MAPbBr₃)_{0.17+ε}(FAPbI₃)_{0.83-ε}. This slight MAPbBr₃ enrichment in the surface was assigned to the presence of unreacted FAI at the surface and PbI₂ in the bulk leading to a deficiency in FAPbI₃. In the RbMAFA sample, only 2% and 1%

were detected at 4000 and 2100 eV, respectively, suggesting that Rb was mainly located in the bulk, but observed to easily migrate to the surface due to its small ionic radius. For the CsMAFA sample, ~6% detected Cs was found to distribute uniformly at both depths. Interestingly, the RbCsMAFA sample did not correspond to a simple superposition of RbMAFA and CsMAFA samples. Both Cs and Rb distributions were homogeneous over the ~18 nm indicating that Rb⁺ and Cs⁺ alkali metals act jointly in a way that Cs helps the incorporation of Rb into the perovskite compound.²⁷⁰ This is also in good agreement with recent theoretical calculations by Syzgantseva et al.¹⁵⁷ reporting that doping by a mixture of Cs and Rb has a synergetic effect on perovskite stabilization.

5. SUMMARY AND OUTLOOK

The exceptional performance of perovskites as light harvester materials for solar cells are ascribed to excellent material properties such as direct band gap, outstanding high absorption coefficient, abrupt optical band edge, large charge carrier diffusion length and low exciton binding energy. The earlier works focused on MAPbI₃ (E_g ~ 1.5–1.61 eV, section 1), α-FAPbI₃ (E_g ~ 1.47–1.55 eV, section 2.1), and α-CsPbI₃ (E_g ~ 1.67–1.73 eV, section 2.3) with preference for FAPbI₃ because of its smaller band gap. The optimal E_g for a single-junction solar cell is between 1.1 and 1.4 eV according to the Shockley–Queisser limit⁹¹ and therefore there is still plenty of room for improvement to achieve even higher efficiencies. So far, the highest PCE of pure MAPbI₃-based solar cell was reported to reach ~18–20%.^{139,140} As a comparison, the best PCE reported for pure FAPbI₃-based perovskite solar cell reached 13.5–18%,^{64,99,142–144} which is somewhat lower than MAPbI₃ (see section 2.1 for more details). CsPbI₃-based perovskite solar cell exhibited limited PCEs of <2.9% (section 2.3).^{115,155,156} Despite high efficiencies of MAPbI₃- and FAPbI₃-based solar cells, these simple hybrid perovskite materials fail to meet the required long-term stability under working conditions (e.g., when tracking the maximum power point (MPP) voltage under continuous light illumination), which is the main obstacle for this technology to reach commercialization.

The fundamental origins for the major PCE loss during solar cell operation includes (i) thermal, moisture, oxygen, bias, and light induced permanent degradation leading to PbI₂ and I-containing volatile species^{28,29,37,92,93,271,272} and (ii) polymorphism of α,δ-FAPbI₃ in ambient conditions (section 2.1 and Figure 3b–d). At room temperature, FAPbI₃ preferentially crystallizes to the δ-FAPbI₃, which possess a high E_g (~2.43 eV¹⁵¹) and is unsuitable for photovoltaic applications. The photoactive α-FAPbI₃ is stabilized only at higher temperatures of ~125–165 °C (Figure 3d). Note that phase transition (e.g., tetragonal to cubic in MAPbI₃) or polymorphism (α- to δ-phase in FAPbI₃) is not a permanent degradation phenomenon as they can be recovered reversibly.

The most promising approaches to improve degradation and phase/polymorphism stability in perovskite solar cells are the introduction of the mixing of A (binary, tertiary, or even quaternary mixed position) site cations and X site halide anions (I, Br, and Cl anions) (Figure 1b and Table 1), i.e., chemical compositional engineering⁸³ or alloying.^{8,101} In fact, A cation mixed perovskite systems of (FA/MA)PbI₃, (MA/Cs)PbI₃, (FA/Cs)PbI₃, (FA/Rb)PbI₃ (section 2) showed enhanced stability (Figure 8c and Figure 9) as compared to reference cells comprising simple perovskites (MAPbI₃ or FAPbI₃). At the current stage, it is difficult to pinpoint the binary composition

(mixed A cation) that leads to the most enhanced stability. This is because of the various synthesis methods of perovskite films adopted by the different groups that lead to different grain sizes, uneven coverage, pinhole formation, capping layer thickness, variations in stoichiometry (e.g., excessive lead halide or cation), nonuniform spatial distribution of mixed perovskites' constituents, and etc. Furthermore, variations in devices induced by processing imperfections make it difficult to properly identify intrinsic material and device properties when comparing among the individual and independent reports from the various groups (Table 1). In this sense, a universal deposition protocol²⁶² is desirable that allows side-by-side comparisons from a myriad of different chemical compositions to pinpoint the most promising chemical composition for photovoltaic applications. Fundamental microscopic origins of enhanced stability for (FA/MA)PbI₃¹⁰⁶ (section 2.1) as well as $t_{\text{effective}}$ (geometric Goldschmidt effective tolerance factor) for (FA/Cs)PbI₃¹⁰¹ (section 2.3) and thermodynamic considerations for (FA/Cs)PbI₃¹¹⁷ and (FA/Cs/Rb)PbI₃¹⁵⁷ (section 2.3) were proposed. For the (FA/MA)PbI₃, the incorporation of MA⁺ cations that has a high dipole moment leads to stronger interactions with PbI₆ octahedra and stabilizes the α -FAPbI₃ without significant lattice shrinkage or changes in the optical properties (Figure 3e).¹⁰⁶ It has been generalized that perovskites with $t < 0.8$ and $t > 1$ tend to form the δ -phase orthorhombic structure (e.g., δ_{O} -CsPbI₃) and the δ -phase hexagonal structure (e.g., δ_{H} -FAPbI₃), respectively.¹⁰¹ The cubic structure is only preferred for $0.8 < t < 1$ (Figure 7a). Alloying a high t FAPbI₃ and low t CsPbI₃, $t_{\text{effective}}$ can be tuned to be between 0.8 and 1.0 in (FA/Cs)PbI₃ perovskites, which favors a stable perovskite structure.¹⁰¹ Based on theoretical calculations, thermodynamic arguments were used to explain the stability of (FA/Cs)PbI₃¹¹⁷ and (FA/Cs/Rb)PbI₃¹⁵⁷ perovskite systems. Cs⁺ and Rb⁺ were predicted to be more efficient in stabilizing the FAPbI₃ perovskite than MA⁺ based on the balance of internal energy variation (ΔE), mixing entropy contributions ($-\Delta S$), and free energy $\Delta F = \Delta E - \Delta S$ (Figure 10). In short, the incorporation of Cs⁺ and/or Rb⁺ thermodynamically favors the formation of new perovskite phases and brings the system into a new equilibrium state. Systems with mixed X anion halides with A site constrained to a single monovalent cation, MAPb(I/Cl), FAPb(I/Cl), MAPb(I/Br), FAPb(I/Br), MAPb(Br/Cl), MAPb(I/Br/Cl), CsPb(I/Br), CsPb(Br/Cl), CsPb(I/Cl) were reported to show peculiar phenomena (section 3). MAPb(I/Cl) is one of the most studied binary mixed perovskite because of the strong interest from the field trying to answer the question whether Cl can in fact be incorporated into the lattice of MAPbI₃ (section 3.1). It has been proposed that Cl incorporation in FAPbI₃ structure helps stabilize the α -FAPbI₃ phase; however, further in-depth studies are needed as the reported works are scarce (section 3.1). Synthesis of MAPb(I/Br) system in the full range of Br:I composition ratio was reported by various groups where E_{g} can be tuned effectively (section 3.2 and Figure 13). These E_{g} values are closely related to the lattice parameter of MAPb(I/Br), which was demonstrated to closely follow the Vegard's law (section 3.2). Only a small deviation from the Vegard's law was observed in the cubic regime (Figure 12) indicating presence of additional interactions in the mixed-halogens.^{220,226} Under storage conditions in dark or low light intensity, MAPb(I/Br) with the Br content of ~20% provided enhanced stability even under relative humidity of 55%. This was closely correlated to the tetragonal to pseudocubic phase trans-

formation (Figure 12b,d) arising from the incorporation of smaller ionic radius of Br and consequently leading to a more compact tightly bound structure (higher t factor). However, when the initially homogeneous MAPb(I/Br) is exposed to light, a phase-separation into two phases, one I-rich domain and the other Br-rich domain form in the same film (Figure 15).^{234,237,240} A number of reports proposed the microscopic as well as macroscopic origins based on thermodynamics, energy levels (CBM and VBM), and light-induced generation of lattice strain leading phase-segregation phenomena (see section 3.2 and Figure 16). In comparison, the FAPb(I/Br) system was reported to be more resilient against phase segregation than MAPb(I/Br) system (section 3.2). The XRD analysis of FAPb(I/Br) system in the full range of Br:I composition ratio shows an interesting phenomenon that has an amorphous phase, which a fundamental understanding is still lacking (Figure 17). More interestingly, incorporation of small amounts of Cs was observed to suppress this amorphous phase (section 4.4). Based on the fact that halide perovskites have strong ionic character, Coulomb interactions were also proposed to play a major role (section 3.5 and Figure 20). The interplay of strain (favoring demixing) and Coulomb (favoring mixing) energies describes the formation energies of CsPb(X/Y) with X,Y = I, Br, Cl that as consequence dictates its stability (Figure 20).²⁵⁰ Based on these fundamental concepts described in sections 1 and 2, the more complex systems of simultaneously mixed A and mixed X (MA/Cs)Pb(I/Br), (FA/MA)Pb(I/Br), (FA/MA)Pb(I/Cl), (FA/Cs)Pb(I/Br), (Cs/FA/MA)Pb(I/Br), (Rb/FA/MA)Pb(I/Br), (Rb/Cs/FA)Pb(I/Br), (Rb/Cs/FA/MA)Pb(I/Br) perovskites were described in section 4. Out of six, five certified efficiencies reported by NREL were based on the (FA/MA)Pb(I/Br) perovskite system (Figure 1b). Enhanced stability under storage conditions was reported for (FA/MA)Pb(I/Br) system (Figure 23a). It was proposed that when MABr is incorporated to alloy with FAPbI₃, the lattice size is reduced and the strain forces (demixing) within the grain are relaxed. In this way, the pseudocubic α -phase is stabilized at room temperature and even under humid air (Figure 23b). However, (FA/MA)Pb(I/Br)-based solar cells under operation conditions (tracking the maximum power point), migration of halide vacancies as well as cation vacancies were proposed to take place (Figure 24).²⁵⁹ Cs incorporation in (FA/Cs)Pb(I/Br) system (section 4.4) was reported to suppress halide segregation (section 2.3) and formation of unstable amorphous phase (Figure 25a). Optimized compositions lead generally to higher E_{g} and therefore the (FA/Cs)Pb(I/Br) system was employed as the top cell in the tandem solar cell structure showing high promises (section 4.4).^{132,133,269} Reports on ternary (Cs/FA/MA)^{65,135,262} and quaternary (Rb/Cs/FA/MA) mixed cations^{7,137} have just bought new hopes that a "magical" composition may lead to a stable perovskite with even higher PCE.⁸⁴ Thermodynamics must underlie this "magical" composition. The concept of effective tolerance factor (section 2.3) may help as a starting point for identifying thermodynamically stable hybrid mixed cations and mixed halides perovskites. Multicomponent perovskite design to achieve a stable single pure phase can enable a creation of stable structures with optimal transport and optical properties. However, a strict control over the phase composition of mixed perovskites should be ensured at the preparation stage, i.e., mixed perovskite materials should be clean from admixtures of nonperovskite phases (section 3.2).^{120,157} Strategies to fabricate perovskite single-crystal-based solar cells were reported in the

literature^{17,273–275} and naturally it is of high interest to study the optoelectronic properties (e.g., surface and bulk trap densities) of mixed halide perovskite single crystals, which has significantly different morphological (e.g., grain boundary), structural (e.g., perovskite crystal orientation), as well as spatial chemical composition distribution compared to polycrystalline thin-films of perovskites.^{209,246,247}

Energy levels measurements (work function, CBM, and VBM) of mixed perovskites and its alignments with adjacent selective contact layers (ETL and HTL) are largely missing in the literature. To date, only a handful of publications provided insights into the energy diagrams, band bending, and interfacial dipole concepts studied on simpler MAPbI₃, MAPbI_{3–x}Cl_x, MAPbBr₃, MAPbBr_{3–x}Cl_x, CsPbBr₃ perovskite systems.^{276–288} Within perovskite solar cell structures, each underneath substrate may have influence on the energy levels of the top layer under examination by ultraviolet spectroscopy (UPS).^{282,289,290} Therefore, precise energy level alignments among stacked layers can only be obtained if UPS is performed on the individual top layers that are stacked on the actual prior layers following device structure. A few studies provided the energy-level alignments across all stacked functional layers within the perovskite solar cell, e.g., FTO/c-TiO₂ (+mp-TiO₂)/MAPbI₃ (or MAPbI_{3–x}Cl_x)/spiro-MeOTAD/Ag or Au, employing KPFM.^{291–293} Further studies on the energy level alignments of mixed perovskites across all stacked layers are expected to bring insights into the correlation with perovskite solar cell parameters.^{276–288}

Pb-free perovskites are still being investigated intensively.^{294–306} Computational approaches based on the tolerance factor (*t*) are generally employed to predict the geometrical stability of three-dimensional (3D) ABX₃ perovskite structures.^{15,160,307} Based on revised ionic radii, which considers greater covalency in metal-halide bonds, Travis et al.¹⁶⁰ found that only a handful of cations may be successfully placed on the B site of iodide-based perovskite: Sn, Yb, Dy, Tm, Sm, Ca, Sr. Filip and Giustino³⁰⁸ and Körbel et al.³⁰⁹ performed a systematic combinatorial search based on DFT over the entire periodic table. Starting from over 32 000 possible 3D ABX₃ compounds, Körbel et al. found 199 thermodynamically stable perovskites in the cubic structure.³⁰⁹ Considering the *E_g* values suitable for photovoltaic applications among these 199 perovskites, all ABX₃ structures with single cation and single anion were based on Sn and Ge halide perovskites. The overall conclusion from these studies is that Pb plays a key role in the optoelectronic properties of 3D ABX₃ perovskites and is unique among all single divalent metals in the periodic table. Great efforts³⁰⁶ have been made toward searching for low-dimensional perovskites with multivalent elements yielding a “3-2-9” (or A₃B₂X₉) 2D,^{81,310–312} quasi-2D,^{313–316} “3-1-5” single-chain, “4-1-6” single-octahedron structures, and “2-1-1-6” 3D double perovskite structures with compensated charges.^{81,296,299,306,317} Quasi-2D perovskite solar cells were reported to show enhanced moisture stability; however, it is often associated with poor photovoltaic properties. In a recent work, Xiao et al.³⁰⁶ compiled reported Pb-free halide perovskites and introduced the concept of electronic dimensionality. It was proposed that the higher photovoltaic performance is intimately associated with the isotropic 3D transport properties of photogenerated charges.^{313,315,316} In this sense, it explains the lower PCE reported for lower-dimensional perovskites as absorbers.³⁰⁶ The chemical composition engineering or alloying strategy has been scarcely explored for Pb-free perovskite

systems, and further exploration for a “magical” composition is expected to bring subsequent years of key findings for the photovoltaic community stimulating further interests of both academia (fundamental research) and industry (new technology).

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Author Contributions

Y.B.Q. conceived the idea, initiated, and supervised the work. All authors contributed to writing the paper.

Notes

The authors declare no competing financial interest.

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