# Efficient Estimation of Tighter Bounds for Worst Case Execution Time of Programs 

Kelly Leahy

In this paper, we will present a framework for the statistical analysis of the execution time of program units. We will show alternative methods for computing the distribution of the execution times and provide justification for the use of each of the methods presented. We will estimate the worst-case execution time (WCET) of the program units using several methods and compare the results of these methods. We will also present a new method for estimating the WCET, based on the theory of extreme value distributions.

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# Efficient Estimation of Tighter Bounds for Worst Case Execution Time of Programs (Masters Thesis) 

Authors: Leahy, Kelly

December 15, 2004


#### Abstract

In this paper, we will present a framework for the statistical analysis of the execution time of program units. We will show alternative methods for computing the distribution of the execution times and provide justification for the use of each of the methods presented. We will estimate the worst-case execution time (WCET) of the program units using several methods and compare the results of these methods. We will also present a new method for estimating the WCET, based on the theory of extreme value distributions.


# SEVER INSTITUTE OF TECHNOLOGY MASTER OF SCIENCE DEGREE 

## THESIS ACCEPTANCE

(To be the first page of each copy of the thesis)

DATE: December 15, 2004

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This student's thesis, entitled Efficient Estimation of Tighter Bounds for Worst Case Execution Time of Programs has been examined by the undersigned committee of three faculty members and has received full approval for acceptance in partial fulfillment of the requirements for the degree Master of Science.
$\qquad$
$\qquad$

# WASHINGTON UNIVERSITY SEVER INSTITUTE OF TECHNOLOGY DEPARTMENT OF COMPUTER SCIENCE AND ENGINEERING 

EFFICIENT ESTIMATION OF TIGHTER BOUNDS FOR WORST CASE EXECUTION TIME OF PROGRAMS<br>by<br>Kelly P. Leahy<br>Prepared under the direction of Ron K. Cytron

A thesis presented to the Sever Institute of Washington University in partial fulfillment of the requirements for the degree of

Master of Science
May, 2005
Saint Louis, Missouri

# WASHINGTON UNIVERSITY SEVER INSTITUTE OF TECHNOLOGY DEPARTMENT OF COMPUTER SCIENCE AND ENGINEERING 

# ABSTRACT <br> EFFICIENT ESTIMATION OF TIGHTER BOUNDS FOR WORST CASE EXECUTION TIME OF PROGRAMS 

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| :---: |
| ADVISOR: Ron K. Cytron |
| May, 2005 |
| Saint Louis, Missouri |

In this paper, we will present a framework for the statistical analysis of the execution time of program units. We will show alternative methods for computing the distribution of the execution times and provide justification for the use of each of the methods presented. We will estimate the worst-case execution time (WCET) of the program units using several methods and compare the results of these methods. We will also present a new method for estimating the WCET, based on the theory of extreme value distributions.
to my loving wife Nadia, for all of her patience and support

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Kelly P. Leahy
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## Chapter 1

## Introduction

The estimation of worst case execution times (WCET) for a program requires a distributional assumption for the random variable representing the run time. This assumption may be explicit, or implicit, and may involve the use of one or more of the well known probability distributions. Much of the existing literature on this subject centers on the use of an extreme value distribution, such as the Gumbel or Weibull distributions (see [7]). Other existing literature takes a low level approach to estimation of the WCET by attempting to model the precise architecture on which the task is running and simulate the program itself (see [3], [19], [23], [13], [9]). Ernst and Ye [10] present a model using basic blocks as the components of the execution time, but do not perform a statistical analysis of these times. Instead, they compute their estimates of the WCET deterministically by simply summing the estimates. We propose a method in which the distributions of the individual elements are computed, and then the total distribution is computed. The total distribution is then used to estimate the WCET.

The disadvantage of low-level methods is that they require an intimate knowledge of the target architecture and the source program. Often it is simply not feasible
to build such a model for each target architecture. On the other hand, a disadvantage of the methods using extreme value distributions is that they often give extremely high WCET estimates. These high estimates can lead to wasted cycles, since the tasks will often complete in a much shorter amount of time than that prescribed by the WCET model.

We present a method for estimating a more aggressive bound on the WCET that is more reasonable for a single scheduling of a given task. This more aggressive bound may cause the deadlines to be missed from time to time, but allow the scheduler to schedule many more tasks than the bounds given by the extreme value distributions. We also describe some methods of using a combination of the extreme value WCET bound and our single-sample WCET bound. Using a linear combination of the bound provided by the extreme value distribution and that provided by our method can alleviate some of the concern that our bound is too small. Of course, a scheduler should recalculate the deadline from time to time, if it finds that the deadlines are being missed with a greater frequency than is desired.

We describe a model in which a task is composed of several components that are assumed to have independent execution times (from each other). These components may represent separate function calls, basic blocks, loop iterations, or any other highlevel components that satisfy the assumption of independence. Of course, it is possible that some components do not satisfy the independence assumption. We will assume, for the purposes of this thesis, that this assumption is satisfied.

### 1.1 The Model

Our model is one for describing the execution times of a given task. For such a task, we will assume that there are several components that are executed to perform this
task, and that that the execution time for the task is given by an equation similar to

$$
\begin{equation*}
T=\sum_{i=1}^{M} T_{i} \tag{1.1}
\end{equation*}
$$

where $T$ is the random variable representing the total execution time for a task, and the $T_{i}$ are the execution times for the individual components. We assume that $M$ is a constant, and that the individual $T_{i}$ are mutually independent random variables (they need not be identically distributed).

We make no specific assumptions about the individual distributions of the $T_{i}$; instead, we provide a method for estimating the distribution of $T$ from estimates of the distributions of the individual $T_{i}$. We describe methods for estimating the $T_{i}$ in systems that can have multiple environment states that greatly affect the execution time of some components. An example of one such system is that in which the main processor has a data and/or memory cache. In such a system, each $T_{i}$ may be dependent on the cache state when the component is to be executed (cold cache / hot cache). In this case, the assumption of independence of the $T_{i}$ may be questionable, but we still assume that they may be independent, since the data cache and instruction cache state of one component is often independent of the cache states of another.

### 1.2 Our Methods

We describe several methods for computing the distribution of $T$, as given in Equation 1.1. Because $T$ is a sum of a fixed number of random variables, its probability function (pf) is given by a convolution of the probability functions of the individual random variables on the right-hand side of Equation 1.1. Analytically, the convolution of two random variables $X$ and $Y$, given by probability functions $f_{X}$ and $f_{Y}$
respectively, is written as $f_{X} * f_{Y}$, and is defined as

$$
\begin{equation*}
f_{X+Y}(z)=\left(f_{X} * f_{Y}\right)(z)=\int_{-\infty}^{\infty} f_{X}(x) f_{Y}(z-x) \mathrm{d} x \tag{1.2}
\end{equation*}
$$

where $z$ is the value of the sum $X+Y$ at which the pf should be evaluated.
For some simple distributions (Gaussian, Gamma family), a closed-form expression for the convolution exists, permitting direct computation of the composite distribution (of $T$ ) when all components are of applicable type. For more complicated models, however, the composite distribution must be numerically estimated. It is these situations with which we are most concerned.

For distributions of the discrete type (we will use these as numerical approximations of those of the continuous type), the convolution of two random variables $X$ and $Y$, given by probability functions $f_{X}$ and $f_{Y}$ respectively, is given by

$$
\begin{equation*}
f_{X+Y}(z)=\left(f_{X} * f_{Y}\right)(z)=\sum_{z=x+y} \sum_{X}(x) f_{Y}(z-x) \tag{1.3}
\end{equation*}
$$

where the summation is over all possible values of $x$ and $y$ that sum to $z$.
As can be seen from this formula for the calculation of the convolution, its apparent complexity is $O(\bar{X} \bar{Y})$, where $\bar{X}$ and $\bar{Y}$ denote the number of possible values for $X$ and $Y$ respectively. We next describe a few methods that can allow us to reduce the complexity of this computation, using Fourier analysis.

### 1.2.1 Convolution using Discrete Distributions

Our first method of computing the composite distribution of $T$ involves an approximation of the distributions of the $T_{i}$ by use of a discrete distribution. We assume that this is, in fact, how the data was already gathered when timing was performed on the
different components. For this reason, this is computationally the simplest method we have at our disposal. For our purposes, we assume that the discrete distribution is given as a frequency distribution. This method is described in Chapter 3. The input to this method is the sample data (as frequency distributions) corresponding to the components represented by the $T_{i}$, and the output is the approximate distribution of the random variable $T$ (as a frequency distribution).

### 1.2.2 Convolution by Discrete Fourier Transform

Due to the computational complexity $\left(O\left(N^{M}\right)\right.$ if each $T_{i}$ has $N$ samples) of the discrete convolution method described in the previous subsection, we are interested in other methods to compute the distribution of $T$. One such method, also using the discrete distributions for the $T_{i}$, is the use of the Fast Fourier Transform (FFT) of the probability functions to perform the convolution. The Fourier transform (the FFT is a fast implementation of the discrete Fourier transform (DFT)) transforms the input function from the time domain to the frequency domain. The advantage of this is that the convolution operation in the time domain is transformed into multiplication in the frequency domain. This property is known as the Fourier convolution theorem. Stated in the typical Fourier notation, we have

$$
\begin{equation*}
f_{T}(t)=\left(f_{T_{1}} * f_{T_{2}} * \cdots * f_{T_{M}}\right)(t)=\mathcal{F}^{-1}\left[\prod_{i=1}^{M} \mathcal{F}\left[f_{T_{i}}\right]\right] \tag{1.4}
\end{equation*}
$$

where $\mathcal{F}[f]$ represents the Fourier transform of a function $f$, and $\mathcal{F}^{-1}[\varphi]$ represents the inverse Fourier transform of a function $\varphi$. In this method (with the discrete distributions), we use the DFT, rather than the continuous Fourier transform (CFT). However, the convolution theorem is also true for the DFT, so we can use it for
the computation of the convolution of the individual $T_{i}$ probability functions. This method is presented in Chapter 4.

### 1.2.3 Convolution by Characteristic Function

One of the disadvantages of convolution by the use of the DFT is that we must use discrete distributions as the input to the method. If, on the other hand, we wish to approximate the distributions of our components $\left(T_{i}\right)$ using continuous distributions (standard or otherwise), we need to be able to compute the CFT of each probability function (for each $T_{i}$ ) and multiply these CFTs together to compute the CFT of the convolution. It turns out that the Fourier transform of the probability function of a distribution is nearly identical to the characteristic function (often well known and with many convenient properties) of the same distribution. We use the definition of the characteristic function for the distribution of a random variable $T$ with distribution function $F(t)=P(T \leq t)$ from [27] (paragraph 4.1), written as

$$
\begin{equation*}
\varphi(\omega)=\int_{-\infty}^{\infty} e^{i \omega t} \mathrm{~d} F(t) \tag{1.5}
\end{equation*}
$$

whereas we use the definition of the Fourier transform from [17] (page 317), written as

$$
\begin{equation*}
\varphi(\omega)=\int_{-\infty}^{\infty} e^{i \omega t} f(t) \mathrm{d} t \tag{1.6}
\end{equation*}
$$

One may immediately see that these two functions are, in fact, identical, and so the characteristic function of a distribution may be used interchangebly with the continuous Fourier transform of the probability function of that distribution. It should be noted that most literature defines the CFT with -itx in the exponent of the exponential factor of the transform, rather than itx as we have here. This point will
be important in Chapter 5 when we must produce input for the inverse FFT algorithm. Once we have the CFT of the pf of the distribution, we then must use some method to compute the inverse Fourier transform of this convolution. We describe a method by which we compute the product of the CFTs of the individual distributions of the $T_{i}$, then compute the inverse DFT of a discretized version of the convolution's CFT. This has the advantage of being more accurate than the fully-discrete method, in that sampling and truncation is performed only at the final step of the computation, rather than prior to the use of the DFT. It also allows us to quickly $(O(N))$ compute the CFT of the convolution, then compute the inverse DFT of the convolution using the FFT for a total cost below that of the fully-discrete method described above. This method is described in Chapter 5.

### 1.3 Examples

We will use one running example throughout all chapters that will illustrate all of the methods described herein. The example will be presented in parts. The data for the example will be presented in Appendix B, and each of the method chapters will cover the application of one of the methods to this example. The example involves a task built from 4 components. The first two of these components are dependent on the system environment (at the start of their respective execution); the other two are not. We will model the first two components with mixture distributions (described in Chapter 2) and the other two components with standard distributions.

In order to test our methods, we will use simulation to generate execution times consistent with our model of the task. We will train the system with these samples (that is, we will fit our distributions to these samples), and then we will
generate separate samples (from the same model) to test the ability of our estimators to predict the WCET.

### 1.4 Mathematical Preliminaries

Much of the theory put forth here requires some knowledge from Probability and Statistics, Fourier Analysis, and general Mathematics. We will present references to some of this knowledge, and will repeat the most important items in the chapters in which they are first used. For a general review of probability and statistics, we refer the reader to one of [15], [12], or [27].

### 1.5 Our Contribution

While others have investigated the WCET by use of extreme value distributions, we present methods for estimating the run-time distribution of the program based on a breakdown of several independent components. We then show how to proceed from this estimate of the run-time distribution to a WCET estimate using the nonasymptotic extreme value distribution. We show that this method produces a tighter bound on the WCET than that provided by the asymptotic Gumbel distribution. We also present several methods that can allow the calculation of the convolutions necessary in our model to be performed in a manner that is less computationally intensive than the traditional methods for computing convolutions. The CFT method we present here is not found elsewhere in Computer Science literature, so far as the author can tell.

## Chapter 2

## Estimating the Component <br> Distributions

### 2.1 Introduction

The first step in the estimation of the total execution-time distribution of a task is the estimation of the distributions of the individual components that comprise the task. The time distributions for these components may be analytical distributions, estimated by the programmer based on hardware specifications, or they may be based on samples gathered during profiling of the task in question. For our purposes, we assume that the data come from the latter source. We do not concentrate on the methods of gathering this data, but rather refer the reader to [22] or [11] for ideas on how to gather timing samples for real-time analysis.

We use a simple program as the example for our task, given in Figure 2.1. This program has four main parts, which we call parts $A, B, C$, and $D$. We assume that timing data has been gathered from all of these parts. We present this timing data in Figure B.1, in Appendix B.

```
program example;
begin
    procA;
    procB;
    procC;
    procD;
end.
```

(part $A$ )
(part B)
(part $C$ )
(part $D$ )

Figure 2.1: An Example Program

The first two parts (represented by procA and procB) are assumed to be dependent upon the system environment. The others (procC and procD), are assumed to be independent. The dependent parts will be modeled with mixture distributions, while the independent parts will be modeled with standard distributions. We determine the standard distributions that best fit parts $C$ and $D$ using the method of maximum likelihood estimation (MLE) of the parameters of the distributions, using the distribution with the greatest value of the likelihood function as our choice of distribution (see [28], [26], [8]). For simplicity, we will restrict our analysis to a relatively small number of continuous probability distributions. We will use three different distributions: Gamma, Normal, Log-Normal. In practice, one might use a catalog of many distributions, but the process we describe for choosing the distribution is the same, regardless of the number of distributions in our catalog.

### 2.2 Fitting the Simple Distributions

In order to choose the distribution to use for each of the two parts $C$ and $D$, and to estimate the values of the parameters of these distributions, we must compute the MLEs of the parameters for each distribution based on the data samples. During the process of computing the MLEs, we also compute the value of the likelihood function
for these parameter values. This value is the maximum value of the likelihood function over all possible parameter choices, given the data as observed. First, we review the process of maximum likelihood estimation.

### 2.2.1 Maximum Likelihood Estimation

The process of computing the MLE estimate of the parameters of a distribution involves maximizing the likelihood function over all choices of parameter values. The likelihood function is defined as

$$
\begin{equation*}
L(\Psi ; \mathbf{x})=\prod_{i=1}^{N} f_{X}\left(x_{i} ; \Psi\right) \tag{2.1}
\end{equation*}
$$

where $\Psi$ is the vector of parameters to the pf of the distribution, $f_{X}$ is the pf of the distribution, and $\left\{x_{i}\right\}_{i=1}^{N}$ is the sample data from which $\Psi$ is estimated. The MLE estimate of $\Psi$ is the value of $\Psi$ that maximizes the value of $L$ over all possible values of $\Psi$. As it is often more convenient in optimization problems, we may instead maximize the log-likelihood function, defined as the logarithm of $L$ written as

$$
\begin{equation*}
\ell(\Psi ; \boldsymbol{x})=\log L(\Psi ; \mathbf{x})=\sum_{i=1}^{N} \log f_{X}\left(x_{i} ; \Psi\right) \tag{2.2}
\end{equation*}
$$

## The Normal Distribution

The MLE estimators for the Normal (Gaussian) distribution with pf

$$
\begin{equation*}
f_{X}(x ; \mu, \sigma)=\frac{1}{\sigma \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{x-\mu}{\sigma}\right)^{2}} \tag{2.3}
\end{equation*}
$$

are

$$
\begin{align*}
\hat{\mu} & =\frac{1}{N} \sum_{i=1}^{N} x_{i}  \tag{2.4}\\
\hat{\sigma}^{2} & =\frac{1}{N} \sum_{i=1}^{N}\left(x_{i}-\hat{\mu}\right)^{2} \tag{2.5}
\end{align*}
$$

Note that the estimator for $\sigma$ is not the same as the traditional (unbiased) estimator $s$ often used in statistical literature ([27], [28], [17]). We are working within the MLE framework, so we use the MLE estimator as our estimate of the variance.

## The Log-Normal Distribution

The Log-Normal distribution is a distribution for which the logarithm of the random variable is normally distributed. It is given by the pf

$$
\begin{equation*}
f_{X}(x ; \mu, \sigma)=\frac{1}{x \sigma \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{\log x-\mu}{\sigma}\right)^{2}} \tag{2.6}
\end{equation*}
$$

with parameter MLE estimates given by

$$
\begin{align*}
\hat{\mu} & =\frac{1}{N} \sum_{i=1}^{N} \log x_{i}  \tag{2.7}\\
\hat{\sigma}^{2} & =\frac{1}{N} \sum_{i=1}^{N}\left(\log x_{i}-\hat{\mu}\right)^{2} \tag{2.8}
\end{align*}
$$

See Appendix A for derivation of the MLE of the Log-Normal distribution.

## The Gamma Distribution

The Gamma distribution is defined by the pf

$$
\begin{equation*}
f_{X}(x ; \alpha, \beta)=\frac{1}{\Gamma(\alpha) \beta^{\alpha}} x^{\alpha-1} e^{-x / \beta} \tag{2.9}
\end{equation*}
$$

where $\Gamma(\alpha)$ is the well-known Gamma function (see [15], [31], or [17]) from advanced calculus.

Unfortunately, the Gamma distribution does not have a closed form expression for the MLE estimates of its parameters. We describe a method for estimating the parameters based on the Newton-Raphson method ([4], [29], [14], and [24]) in Appendix A.

### 2.2.2 MLE Estimates for the Training Data

Now that we have MLE estimators defined for the three different standard distributions we will be using, we may compute the MLE estimates of the parameters for each of the three distributions for each of our two data sets (for parts $C$ and $D$ ), and compute the loglikelihood function for each of the distributions at these choices of parameters. The first step in computing these results is finding the values of the statistics used in our analysis. These statistics are the sample variance, the sample mean, the log-sample variance, and the log-sample mean (these are the terms we use for the sample variance and mean of the logarithm of the observations). We denote
these respectively by $S_{X}^{2}, \bar{x}, S_{\log X}^{2}$, and $\overline{l(x)}$, given by

$$
\begin{aligned}
S_{X}^{2} & =\frac{1}{N} \sum_{i=1}^{N}\left(x_{i}-\bar{x}\right)^{2} & S_{\log X}^{2} & =\frac{1}{N} \sum_{i=1}^{N}\left(\log x_{i}-\overline{l(x)}\right)^{2} \\
\bar{x} & =\frac{1}{N} \sum_{i=1}^{N} x_{i} & \overline{l(x)} & =\frac{1}{N} \sum_{i=1}^{N} \log x_{i} .
\end{aligned}
$$

The values of these statistics are tabulated in Figure 2.2.

Table 2.2: Statistics for MLE Estimation

| Part | $S_{X}$ | $\bar{x}$ | $S_{\log X}$ | $\overline{l(x)}$ |
| :---: | :---: | :---: | :---: | :---: |
| $C$ | 18.77356302 | 103.1843 | 0.18612334 | 4.61960194 |
| $D$ | 14.62697266 | 148.7574 | 0.09687508 | 4.99758541 |

## Normal and Log-Normal Distributions

Using these statistics, we can immediately estimate the parameters for the Normal and Log-Normal distributions. The estimates are given in Figure 2.3.

Table 2.3: Normal and Log-Normal Results $(C, D)$

|  | Normal |  |  |  |  | Log-Normal |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Part | $\hat{\mu}$ | $\hat{\sigma}$ | $-\ell(\Psi)$ |  | $\hat{\mu}$ | $\hat{\sigma}$ | $-\ell(\Psi)$ |  |
| $C$ | 103.1843 | 18.7736 | 435.1388 |  | 4.6196 | 0.1861 | 435.7195 |  |
| $D$ | 148.7574 | 14.6270 | 410.1806 |  | 4.9976 | 0.0969 | 408.2191 |  |

## Gamma Distribution

In order to estimate the parameters for the Gamma distribution, we use the iterative process from Section A.3.1. The results are tabulated in Figure 2.4.

Table 2.4: Gamma estimation using Newton's method

| $k$ | $\alpha^{(k)}$ | $\beta^{(k)}$ | $T_{1}^{(k)}$ | $T_{2}^{(k)}$ | $T_{3}^{(k)}$ | $\delta^{(k)}$ | $\mathrm{NLL}^{(k)}$ |  |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | Results for part C |  |  |  |  |  |  |
| 0 | 30.20883 | 3.41570 | 0.00027 | 0.00000 | 0.01673 | 0.49417 | 434.82077 |  |  |
| 1 | 29.71779 | 3.47122 | -0.00027 | -0.02726 | 0.01701 | 0.00783 | 434.81422 |  |  |
| 2 | 29.72562 | 3.47122 | -0.00000 | 0.00000 | 0.01701 | 0.00006 | 434.81412 |  |  |
| 3 | 29.72568 | 3.47122 | 0.00000 | 0.00000 | 0.01701 | 0.00000 | 434.81412 |  |  |
| 4 | 29.72568 | 3.47122 |  |  |  |  | 434.81412 |  |  |
|  |  |  |  | Results for part $D$ |  |  |  |  |  |
| 0 | 103.43041 | 1.43824 | -0.00011 | 0.00000 | 0.00485 | 2.35849 | 408.72488 |  |  |
| 1 | 105.78868 | 1.40544 | -0.00052 | -0.07733 | 0.00474 | 0.05207 | 408.71308 |  |  |
| 2 | 105.84075 | 1.40548 | -0.00000 | 0.00000 | 0.00474 | 0.00305 | 408.71164 |  |  |
| 3 | 105.84379 | 1.40544 | 0.00000 | -0.00000 | 0.00474 | 0.00000 | 408.71164 |  |  |
| 4 | 105.84379 | 1.40544 |  |  |  |  | 408.71164 |  |  |

## Choosing the Distribution

Now that we have calculated the parameter estimates for each of our distributions in parts $C$ and $D$, we can choose the distribution that best fits our data under the MLE criterion. The distribution (from our choices) that best fits the sample data will be the one with the smallest negative log-likelihood (NLL) or equivalently, the
largest likelihood. The NLL for each of the distributions for each part is tabulated in Figure 2.5.

Table 2.5: NLL for Parts $C$ and $D$

| Part | Normal | Log-Normal | Gamma | Best Choice |
| :---: | ---: | ---: | :---: | :---: |
| $C$ | 435.13882 | 435.71948 | 434.81412 | Gamma |
| $D$ | 410.18058 | 408.21910 | 408.71164 | Log-Normal |

The best choice from the distributions in our catalog for parts $C$ and $D$ are Gamma and Log-Normal respectively. We take the parameters for these distributions from Figures 2.3 and 2.4. Our choices lead to the distribution assumptions given in Figure 2.6.

Table 2.6: Distributions for Parts $C$ and $D$

| Part | Distribution | Parameters |
| :---: | :---: | :---: |
| C | Gamma | $\hat{\alpha}=29.72568$ |
|  |  | $\hat{\beta}=3.47122$ |
| D | Log-Normal | $\hat{\mu}=4.99759$ |
|  |  | $\hat{\sigma}=0.09688$ |

### 2.3 Fitting the Complex Distributions

We now have estimates of the distributions for Parts $C$ and $D$ - the simple distributions. Our goal now is to find applicable distributions and their parameters for the
environment-dependent parts of our program (Parts $A$ and $B$ ). We will use a mixture model to approximate the distribution of Parts $A$ and $B$.

A mixture model is a model that describes random variables that can come from one of a group of more than one distribution, where the distribution is randomly chosen. For example, consider an experiment in which we have two urns, each containing a mix of colored balls. In the first, we have $30 \%$ red balls and $70 \%$ black balls. In the second, we have $50 \%$ red balls and $50 \%$ black balls. The mixture model describes the probability of choosing a red (or a black) ball given that we randomly choose the urn from which to draw the ball. In this experiment, there is a 50-50 chance of drawing from a given urn, and a $30 \%$ chance of drawing red from the first urn (given it was the one selected) and a $50 \%$ chance of drawing red from the second (again, given it was the one selected). Our probability distribution (for the choice of a red ball) is given by

$$
\begin{equation*}
\operatorname{Pr}(X=\mathrm{red})=0.5(0.3)+0.5(0.5) \tag{2.10}
\end{equation*}
$$

In general, a mixture model is characterized by a $\mathrm{pf} f_{Z}$ of the form

$$
\begin{equation*}
f_{Z}(z ; \Psi)=\sum_{i=1}^{N-1} \pi_{i} f_{X_{i}}(z ; \Psi)+\left(1-\sum_{i=1}^{N-1} \pi_{i}\right) f_{X_{N}}(z ; \Psi) \tag{2.11}
\end{equation*}
$$

or, if we define $\pi_{N}$ such that $\sum_{i=1}^{N} \pi_{i}=1$, then we can simply write

$$
\begin{equation*}
f_{Z}(z ; \Psi)=\sum_{i=1}^{N} \pi_{i} f_{X_{i}}(z ; \Psi) \tag{2.12}
\end{equation*}
$$

We will use Equation 2.12 to describe the pf of a mixture distribution for the remainder of this document, under the understanding that $\pi_{N}$ is a function of the $\pi_{i}$ for $i=1,2, \ldots N-1$, rather than a parameter itself. The $\Psi$ given in the equations is
the vector of parameters for all of the mixture components (each of the $X_{i}$ is called a mixture component, with $\pi_{i}$ known as its mixture proportion and $f_{X_{i}}$ its pf ).

The concept that the distribution of the program part (Parts $A$ and $B$ ) depends on which state the system is in translates precisely to a mixture distribution. Each of the $\pi_{i}$ represents the probability that the system is in state $i$ with the $f_{X_{i}}$ representing the distribution when the system is in state $i$. A mixture model is, in fact, a direct translation of the law of total probability. The law of total probability (see [15] for instance) states that if a set of events ( $A_{i}$ for $i=1,2, \ldots N$ ) is mutually exclusive and exhaustive ${ }^{1}$ then

$$
\begin{equation*}
P(B)=\sum_{i=1}^{N} P\left(B \mid A_{i}\right) P\left(A_{i}\right) \tag{2.13}
\end{equation*}
$$

where the notation $P\left(B \mid A_{i}\right)$ represents the probability of the event $B$ occuring given that the event $A_{i}$ is certain. To see the relationship between this law of total probability and the mixture pf given by Equation 2.12, we write $P(B)=f_{Z}(z), \pi_{i}=P\left(A_{i}\right)$, and $P\left(B \mid A_{i}\right)=f_{X_{i}}(z)$. We have already assumed that the system may be in only one state at a time, and that the system must be in one of the states, so the $A_{i}$ in our case are, of course, mutually exclusive and exhaustive events.

We refer the interested reader to [21] for an excellent discourse on the applications and theory of mixture models.

[^0]For the purposes of fitting the data from Parts $A$ and $B$, we use a twocomponent mixture model, made up of components from each of our three distributions. We will not use "mixed" mixtures where the components are from different distributions, so we will have three possible models to fit. If the distributions of the components were likely to be different (as may very well be the case in a real system), we would need to fit each possible (or feasible) combination of component distributions. An example of such a mixture would be a distribution where $10 \%$ of the time, the distribution is a Normal distribution, and $90 \%$ of the time it is a Gamma distribution. We will not use such a mixture, though there is no theoretical reason why such a mixture could not exist. As in the previous section, the maximum likelihood criterion would be used to select which combination of mixture distributions provides the best fit to the data.

### 2.3.1 Maximum Likelihood Estimation

There are several ways in which the MLE estimates for the parameters of a mixture distribution may be obtained. All of these methods are numerical in nature, as there exist no closed form results for any of the distributions in which we are interested. We will concentrate on the use of the EM algorithm ([20], [21]) for the estimation of the mixture parameters.

In a mixture model, there are two types of explicit parameters: the component proportions, and the component parameters. The first of these refers to the $\pi_{i}$ in the mixture pf. If the mixture has $M$ components, there are $M-1$ of these parameters. The second type of parameter refers to the parameters of the component distributions themselves. If we assume that all component distributions are of the same type (as we will in this document), then the number of these parameters is $M P$ where $P$ is the
number of parameters for the component distribution. Therefore, we are estimating $3 M-1$ parameters for each of our types of component distributions (each has 2 parameters). In our case, $M=2$, so there are 5 parameters to estimate for each type of component distribution.

## The EM Algorithm

The EM algorithm is a two-stage algorithm for MLE in the face of missing data. While the reader may note that it doesn't appear that we are missing data (our samples aren't truncated or censored), our problem can be formulated as a missingdata problem by assuming that there are some data, denoted $z_{i j}$, associated with each of the data points $x_{i}$. These data $z_{i j}$ are indicators that indicate which state the system was in for the execution corresponding to $x_{i}$. They are $0 / 1$ variables, with $z_{i j}=0$ indicating "not in state $j$ during execution leading to time $x_{i}$ " and $z_{i j}=1$ indicating the alternative (in state $j$ ). Had we been able to gather this data, it would be quite simple to estimate the parameters of the components (since we would know the distribution to which each observation belongs) and the proportions (since we would know how many observations came from each system state).

For example, in the Normal distribution case, we would have

$$
\begin{align*}
\hat{\mu}_{j} & =\frac{\sum_{i=1}^{N} z_{i j} x_{i}}{\sum_{i=1}^{N} z_{i j}}  \tag{2.14}\\
\hat{\sigma}_{j}^{2} & =\frac{\sum_{i=1}^{N} z_{i j}\left(x_{i}-\hat{\mu}_{j}\right)^{2}}{\sum_{i=1}^{N} z_{i j}} \tag{2.15}
\end{align*}
$$

and

$$
\begin{equation*}
\hat{\pi}_{j}=\frac{1}{N} \sum_{i=1}^{N} z_{i j} \tag{2.16}
\end{equation*}
$$

The Process The process of applying the EM algorithm is as follows:

1. Compute an initial estimate of the parameters of the distributions based on an initial partitioning of the observations into groups corresponding to the mixture components.
2. Using the computed estimates for the component distribution parameters, compute the probabilities that each of the data points resulted from each of the mixture components. Use these probabilities to estimate the $z_{i j}$ using the equation

$$
\begin{equation*}
\hat{z}_{i j}=\mathrm{E}\left[z_{i j}\right]=\frac{\pi_{j} f_{X_{j}}\left(x_{i} ; \hat{\theta}_{j}\right)}{\sum_{k=1}^{M} \pi_{k} f_{X_{k}}\left(x_{i} ; \hat{\theta}_{k}\right)} \tag{2.17}
\end{equation*}
$$

where $\hat{\theta}_{j}$ is the current estimate of the parameters for mixture distribution $j$, and $M$ is the number of mixture components.
3. Using the estimates $\hat{z}_{i j}$ from step 2, estimate the parameters of the component distributions (see Equations A.5* and A.6* for the Normal distribution, Equations A.11* and A.12* for the Log-Normal distribution, or Equations A.13*, A.14* and A.15* for the Gamma distribution).
4. If the result has converged to within the desired tolerance, stop; otherwise return to step 2

### 2.3.2 MLE Estimates for the Training Data

For our purposes, we will initialize our estimates for the EM algorithm by simply dividing the training data in two halves, the fifty smallest and the fifty largest. We then use this division to estimate the parameters of our three distributions, for the starting point of the EM algorithm. The statistics computed on this data are given in Figure 2.7.

Table 2.7: Statistics for the EM Algorithm

| Part | Group | $S_{X}$ |  | $S_{\log X}$ | $\overline{l(x)}$ | $\hat{\pi}_{j}$ |
| :---: | :---: | ---: | ---: | :---: | :---: | :---: |
| A | Low | 1.35859522 | 38.3052 | 0.03627907 | 3.64493752 | 0.5 |
|  | High | 430.93016683 | 310.6080 | 1.42489530 | 4.62455229 | 0.5 |
| B | Low | 5.22430583 | 43.2008 | 0.13365915 | 3.75753905 | 0.5 |
|  | High | 414.87919156 | 276.5026 | 1.21920932 | 4.67256012 | 0.5 |

## The Normal Mixtures

First, we estimate the parameters for the Normal mixtures. The results for Part $A$ are tabulated in Figure 2.8. The results for Part $B$ are tabulated in Figure 2.9. The first row of each table (iteration one) is the initialization provided in Figure 2.7. The second and later iterations are the results of applying the EM algorithm to the initial estimates of the parameters given in row 1.

Table 2.8: Part $A$ EM for Normal Components

| Iter | $\pi_{1}$ | $\hat{\mu}_{1}$ | $\hat{\mu}_{2}$ | $\hat{\sigma}_{1}$ | $\hat{\sigma}_{2}$ | NLL |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0.50000 | 38.30520 | 310.60800 | 1.35860 | 430.93017 | 433.64794 |
| 2 | 0.77469 | 41.06717 | 633.10545 | 2.54341 | 412.26863 | 349.04243 |
| 3 | 0.85860 | 39.86482 | 991.68282 | 2.26691 | 75.04354 | 306.90075 |
| 4 | 0.86000 | 39.83500 | 1001.41786 | 2.26794 | 20.22373 | 294.90933 |
| 5 | 0.86000 | 39.83500 | 1001.41786 | 2.26794 | 20.22373 | 294.90933 |

Table 2.9: Part $B$ EM for Normal Components

| Iter | $\pi_{1}$ | $\hat{\mu}_{1}$ | $\hat{\mu}_{2}$ | $\hat{\sigma}_{1}$ | $\hat{\sigma}_{2}$ | NLL |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0.50000 | 43.20080 | 276.50260 | 5.22431 | 414.87919 | 518.24196 |
| 2 | 0.74908 | 49.99793 | 487.79806 | 7.38597 | 464.48401 | 443.34094 |
| 3 | 0.87641 | 49.40430 | 943.05710 | 8.30094 | 236.95495 | 422.04111 |
| 4 | 0.88999 | 48.94692 | 1057.08244 | 8.38641 | 31.95495 | 403.49570 |
| 5 | 0.89000 | 48.94685 | 1057.17273 | 8.38651 | 30.68959 | 403.47817 |
| 6 | 0.89000 | 48.94685 | 1057.17273 | 8.38651 | 30.68959 | 403.47817 |

## The Log-Normal Mixtures

Now that we have estimated the Normal-component mixtures, we can move on to the Log-Normal and Gamma mixtures. The Log-Normal estimates are given by Figure 2.10 and 2.11.

Table 2.10: Part A EM for Log-Normal Components

| Iter | $\pi_{1}$ | $\hat{\mu}_{1}$ | $\hat{\mu}_{2}$ | $\hat{\sigma}_{1}$ | $\hat{\sigma}_{2}$ | NLL |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0.50000 | 3.64494 | 4.62455 | 0.03628 | 1.42490 | 422.93776 |
| 2 | 0.69615 | 3.92630 | 4.61232 | 0.26613 | 1.64867 | 490.01643 |
| 3 | 0.78568 | 3.80686 | 5.33674 | 0.13785 | 1.49541 | 430.57377 |
| 4 | 0.83998 | 3.70945 | 6.36724 | 0.06261 | 0.95026 | 352.38005 |
| 5 | 0.85951 | 3.68344 | 6.89583 | 0.05689 | 0.17912 | 318.42644 |
| 6 | 0.86000 | 3.68313 | 6.90897 | 0.05692 | 0.02004 | 294.63614 |
| 7 | 0.86000 | 3.68313 | 6.90897 | 0.05692 | 0.02004 | 294.63614 |

Table 2.11: Part $B$ EM for Log-Normal Components

| Iter | $\pi_{1}$ | $\hat{\mu}_{1}$ | $\hat{\mu}_{2}$ | $\hat{\sigma}_{1}$ | $\hat{\sigma}_{2}$ | NLL |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0.50000 | 3.75754 | 4.67256 | 0.13366 | 1.21921 | 491.87104 |
| 2 | 0.64877 | 4.21344 | 4.21802 | 0.40855 | 1.55641 | 526.07859 |
| 3 | 0.72727 | 4.11043 | 4.49404 | 0.28738 | 1.64076 | 498.43958 |
| 4 | 0.79453 | 4.00911 | 5.01139 | 0.20661 | 1.62081 | 467.61674 |
| 5 | 0.83470 | 3.93301 | 5.63920 | 0.16739 | 1.51421 | 448.04447 |
| 6 | 0.85653 | 3.90063 | 6.09217 | 0.16192 | 1.38700 | 443.11100 |
| 7 | 0.86894 | 3.88811 | 6.38258 | 0.16395 | 1.23241 | 441.61787 |
| 8 | 0.87740 | 3.88290 | 6.59212 | 0.16795 | 1.03711 | 440.03889 |
| 9 | 0.88645 | 3.87777 | 6.84796 | 0.17535 | 0.57681 | 433.05540 |
| 10 | 0.89000 | 3.87542 | 6.96292 | 0.17781 | 0.02955 | 405.38911 |
| 11 | 0.89000 | 3.87542 | 6.96293 | 0.17781 | 0.02896 | 405.38472 |
| 12 | 0.89000 | 3.87542 | 6.96293 | 0.17781 | 0.02896 | 405.38472 |

The astute reader may notice that the NLL on the Log-Normal mixtures actually increases from the first to the second iteration. This is a result of the poor initial estimate given by the choice of the median as the "dividing line". One should notice that after the first iteration, the NLL decreases monotonically.

## The Gamma Mixtures

As before, the process of estimating the Gamma mixtures is a bit more complicated than that of the Log-Normal and Normal mixtures, due to the fact that the Gamma MLEs are not closed-form. However, the general idea is the same, only the MLE (the

M step) is replaced by the solve for the Gamma MLEs based on the initial estimates given by $\bar{x}$ and $S^{2}$. The results are tabulated in Figure 2.12 for Part $A$ and Figure 2.13 for Part $B$.

Table 2.12: Part $A$ EM for Gamma Components

| Iter | $\pi_{1}$ | $\hat{\alpha}_{1}$ | $\hat{\alpha}_{2}$ | $\hat{\beta}_{1}$ | $\hat{\beta}_{2}$ | NLL |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0.50000 | 771.60823 | 0.56034 | 0.04964 | 554.31932 | 423.01425 |
| $2^{\dagger}$ | 0.72282 | 482.01746 | 0.65914 | 0.08166 | 799.15633 | 423.00408 |
| 3 | 0.83418 | 353.32292 | 5.98702 | 0.11252 | 142.32138 | 358.87962 |
| 4 | 0.86000 | 308.97979 | 2405.30944 | 0.12892 | 0.41634 | 331.93477 |
| 5 | 0.86000 | 308.97919 | 2477.36193 | 0.12892 | 0.40423 | 294.67962 |
| 6 | 0.86000 | 308.97919 | 2477.36193 | 0.12892 | 0.40423 | 294.67660 |

Table 2.13: Part $B$ EM for Gamma Components

| Iter | $\pi_{1}$ | $\hat{\alpha}_{1}$ | $\hat{\alpha}_{2}$ | $\hat{\beta}_{1}$ | $\hat{\beta}_{2}$ | NLL |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0.50000 | 60.26261 | 0.64396 | 0.71688 | 429.38127 | 496.82920 |
| 2 | 0.68900 | 56.31266 | 0.62847 | 0.83627 | 651.84801 | 492.82888 |
| $3^{\dagger}$ | 0.81232 | 45.11140 | 1.66844 | 1.08018 | 384.08715 | 455.59212 |
| $4^{\dagger}$ | 0.87498 | 36.67997 | 7.92605 | 1.33867 | 117.95610 | 440.82549 |
| 5 | 0.89000 | 32.82081 | 1188.05476 | 1.49134 | 0.88797 | 427.51405 |
| 6 | 0.89000 | 32.82080 | 1190.55002 | 1.49134 | 0.88797 | 404.19959 |
| 7 | 0.89000 | 32.82080 | 1190.55002 | 1.49134 | 0.88797 | 404.19958 |
| 8 | 0.89000 | 32.82080 | 1190.55002 | 1.49134 | 0.88797 | 404.19958 |

[^1]
## Choosing the Distribution

As before, we choose the distribution family that provides the best MLE (lowest NLL) as the distribution to represent the data. The NLL values for the mixtures we have selected as candidates for our estimates are given in Figure 2.14.

Table 2.14: NLL for Parts $A$ and $B$

| Part | Normal | Log-Normal | Gamma | Best Choice |
| :---: | ---: | ---: | :---: | :---: |
| $A$ | 294.90933 | 294.63614 | 294.67660 | Log-Normal |
| $B$ | 403.47817 | 405.38472 | 404.19958 | Normal |

### 2.4 Summary of Results

In the previous two sections, we estimated the parameters for the four distributions of Parts $A, B, C$, and $D$ from our example program. Two of these distributions were estimated as mixtures, and two as simple distributions. Parts $A$ and $B$ were estimated as mixtures, with the parameters given by Figure 2.15. Parts $C$ and $D$ were estimated as simple distributions, with the parameters given in Figure 2.16.

Table 2.15: Distributions for Parts $A$ and $B$

| Part | Distribution |  | Parameter Estimates |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| A | Log-Normal | $\hat{\pi}_{1}=$ | 0.86000 | $\hat{\pi}_{2}=$ | 0.14000 |
|  |  | $\hat{\mu}_{1}=$ | 3.68313 | $\hat{\mu}_{2}=$ | 6.90897 |
|  |  | $\hat{\sigma}_{1}=$ | 0.05692 | $\hat{\sigma}_{2}=$ | 0.02004 |
| $B$ | Normal | $\hat{\pi}_{1}=$ | 0.89000 | $\hat{\pi}_{2}=$ | 0.11000 |
|  |  | $\hat{\mu}_{1}=$ | 48.94685 | $\hat{\mu}_{2}=$ | 1057.17273 |
|  |  | $\hat{\sigma}_{1}=$ | 8.38651 | $\hat{\sigma}_{2}=$ | 30.68959 |

Table 2.16: Distributions for Parts $C$ and $D$

| Part | Distribution | Parameters |  |
| :---: | :---: | :---: | :---: |
| $C$ | Gamma | $\hat{\alpha}=$ | 29.72568 |
|  |  | $\hat{\beta}=$ | 3.47122 |
|  | Log-Normal <br> $D$ | $\hat{\mu}=$ | 4.99759 |
|  |  | $\hat{\sigma}=$ | 0.09688 |

## Chapter 3

## Discrete Convolution

In this chapter, we present the method of discrete convolution. This method is the simple brute force method of computing convolutions. We describe this method in the context of two different data sources (both based on the example data given in Appendix B). The first data source we consider is the raw data from the example data set. The second data source is the empirical pf of the distribution, smoothed by kernel density smoothing [27].

### 3.1 The Convolution Process

Consider a set of data points with associated frequencies. Each point in the data set can be represented by the ordered pair $\left(f_{i}, x_{i}\right)$, where $f_{i}$ is the frequency associated with the point $x_{i}$. In the case of a regular data sample that is ungrouped, we may let $f_{i}=1$ for all $i$. A data set gathered in this way is known as a frequency distribution if

$$
\sum_{i} f_{i}=1
$$

and the $x_{i}$ are unique. The first condition can be satisfied by a process known as "normalization". When normalizing a frequency distribution, each $f_{i}$ is divided by $\sum_{i} f_{i}$ in order to cause all $f_{i}$ to sum to one. The second condition can be satisfied by grouping all $x_{i}$ with a given value together $\left(\left(f_{j}, y_{j}\right)=\left(\sum_{x_{i}=y_{j}} f_{i}, y_{j}\right)\right)$.

The convolution of two frequency distributions $F$ and $G$, given in the normalized and unique-point form described above, is then

$$
\begin{equation*}
Z=\left\{(u, z) \mid(\exists(v, z) \in F \oplus G) \wedge u \sum_{(a, z) \in F \oplus G} a\right\} \tag{3.1}
\end{equation*}
$$

where the operator $\oplus$ is defined such that

$$
X \oplus Y=\{(f \cdot g, x+y) \mid(f, x) \in X,(g, y) \in Y\} .
$$

An Example. Consider the two frequency distributions $F$ and $G$ defined as

$$
F=\{(0.1,1),(0.5,2),(0.3,3),(0.1,5)\}
$$

and

$$
G=\{(0.3,1),(0.2,3),(0.5,7)\}
$$

The convolution of these two frequency distributions is given by

$$
\begin{aligned}
Z=\{ & (0.03,2),(0.15,3),(0.11,4),(0.1,5),(0.09,6), \\
& (0.07,8),(0.25,9),(0.15,10),(0.05,12)\}
\end{aligned}
$$

A simple algorithm for computing the discrete convolution $Z$ of $F$ and $G$ is given by Algorithm 1. The result of this algorithm will be a set $Z$ that satisfies the two requirements given above for frequency distributions.

```
Algorithm 1 Compute the convolution \(Z=F * G\)
Require: \(F, G\) frequency distributions
Ensure: \(Z\) is frequency distribution of \(F * G\)
    \(Z \leftarrow \emptyset\)
    for all \((f, x) \in F\) do
        for all \((g, y) \in G\) do
            if \(\exists(u, x+y) \in Z\) then
            \(Z \leftarrow(Z-\{(u, x+y)\}) \cup\{(u+f g, x+y)\}\)
            else
                    \(Z \leftarrow Z \cup\{(f \cdot g, x+y)\}\)
            end if
        end for
    end for
```


### 3.2 Extension to $M$ Distributions

The algorithm given in the previous section for computing the convolution is only for two frequency distributions. In order to compute the convolution of more than two distributions, we recognize that the output of the convolution operation is itself a distribution, so we perform the convolutions in series. For instance, if we wish to compute the convolution

$$
Z=F_{1} * F_{2} * F_{3}
$$

we notice that we may simply compute

$$
Z_{2}=F_{1} * F_{2}
$$

and then compute

$$
Z=Z_{2} * F_{3}
$$

using Algorithm 1. It is easy to see that this can be extended to any finite number of distributions being convolved together.

In general, to convolve $M$ distributions, the convolution operation must be performed $(M-1)$ times. Since the convolution operation can be shown to be $O\left(N_{1} N_{2}\right)$ if there are $N_{1}$ points in the first distribution and $N_{2}$ points in the second, the convolution of $M$ distributions is $O\left(N^{M}\right)$. This point is not obvious, but is apparent after recognizing that each convolution potentially produces an output that is the size of the Cartesian product of the elements of the two input sets.

In later chapters, we present other methods of computing the convolution that require less computation time than this method. One advantage of this method, however, is that it is the only method that can handle arbitrary frequency distributions without any sampling (approximation) error. The other methods will require points to be equally spaced on the number line. Such a requirement is not necessary for the brute-force convolution method.

### 3.3 Using the Results of Convolution

One of the disadvantages of discrete convolution, as described here, is that it has the tendency to magnify the effects of sampling artifacts in the component distributions. Sampling artifacts are due to the fact that we are approximating the real distribution of a component run-time with a discrete set of samples. These samples are not, in themselves, necessarily very representative of the source distribution. Without some form of smoothing, either by grouping, approximation by analytical function, or kernel methods, the distribution does not assign probability at values for which no observation was encountered.

As an example of sampling artifacts, consider an experiment in which a fair coin is tossed 50 times, with the number of heads recorded. If this experiment is
performed 10 times, resulting in the sample values

$$
\begin{array}{llllllllll}
29 & 29 & 19 & 21 & 32 & 28 & 23 & 26 & 22 & 23
\end{array}
$$

we obtain the frequency distribution

$$
(19,0.1),(21,0.1),(22,0.1),(23,0.2),(26,0.1),(28,0.1),(29,0.2),(32,0.1)
$$

which doesn't even include the most likely value (the mode -25 ) of the true distribution. If we use this frequency distribution, without any smoothing or grouping, we would be led to believe that the distribution is bimodal with modes as 23 and 29, and no probability of the values 1 through $18,20,24,25,27,30,31$, or 33 through 50 .

If we convolve the distributions in their current form (as unsmoothed frequency distributions), the errors given by the original samples propagate to the entire distribution and produce even more holes than those found in the original frequency distributions. Consider, in reference to our example, that we would like to predict the number of "heads" given by two separate trials of the experiment described above. If we use the frequency distribution, we would have the convolution of this distribution with itself. This convolution is the frequency distribution given by Figure 3.1. In the table, $\hat{p}(x)$ is the empirical distribution, and $p(x)$ is the probability function of the true distribution of the convolution at these points. $\tilde{p}(x)$ is described later in this paragraph. Those points with values significantly close to zero or one are left out of the table. Here, $p$ and $\hat{p}$ are the probability functions of the distribution (true and empirical). We measure the error by taking the sum of the squared differences between the empirical and actual probability functions. Using this error measure, the error in the approximation of $p$ by $\hat{p}$ is 0.01135 . If instead, we estimate (using

MLE) the parameter $\pi$ of the binomial distribution (assuming the number of trials is known to be 50), we estimate from our original distribution of values the result $\hat{\pi}=0.504$. Using this estimate to compute $\tilde{p}$ (the distribution of the convolution of two $b(50,0.504)$ random variables), we have the $\tilde{p}$ given in Figure 3.1. The error given by the same measure applied to $\tilde{p}$ (versus $p$ ) is 0.00018 . The reason for this substantial decrease in measured error is that the distribution does not have so many holes, since it was smoothed by using an analytical distribution before convolution occurred. Of course, often we are unable to determine the underlying analytical distribution of a dataset (if any is appropriate at all). In these more common cases, we may still attempt to smooth the empirical data by kernel methods as discussed in the next section. The entries in the fifth column of Figure 3.1 are described in the next section.

Table 3.1: Discrete Convolution (Binom. Example)

| X | $\hat{p}(x)$ | $p(x)$ | $\tilde{p}(x)$ | $\tilde{\tilde{p}}(x)$ |
| ---: | :---: | :---: | :---: | :---: |
| $\leq 28$ | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 29 | 0.00000 | 0.00001 | 0.00001 | 0.00000 |
| 30 | 0.00000 | 0.00002 | 0.00002 | 0.00000 |
| 31 | 0.00000 | 0.00005 | 0.00004 | 0.00000 |
| 32 | 0.00000 | 0.00011 | 0.00008 | 0.00000 |
| 33 | 0.00000 | 0.00023 | 0.00018 | 0.00000 |
| 34 | 0.00000 | 0.00046 | 0.00035 | 0.00000 |
| 35 | 0.00000 | 0.00086 | 0.00068 | 0.00000 |
| 36 | 0.00000 | 0.00156 | 0.00124 | 0.00150 |
| 37 | 0.00000 | 0.00270 | 0.00218 | 0.00200 |

Table 3.1: Discrete Convolution (Binom. Example)

| X | $\hat{p}(x)$ | $p(x)$ | $\tilde{p}(x)$ | $\tilde{\tilde{p}}(x)$ |
| :---: | :---: | :---: | :---: | :---: |
| 38 | 0.01000 | 0.00447 | 0.00368 | 0.00600 |
| 39 | 0.00000 | 0.00711 | 0.00594 | 0.00900 |
| 40 | 0.02000 | 0.01084 | 0.00921 | 0.01900 |
| 41 | 0.02000 | 0.01587 | 0.01370 | 0.02300 |
| 42 | 0.05000 | 0.02229 | 0.01955 | 0.03350 |
| 43 | 0.02000 | 0.03007 | 0.02680 | 0.03800 |
| 44 | 0.05000 | 0.03895 | 0.03527 | 0.04450 |
| 45 | 0.06000 | 0.04847 | 0.04460 | 0.04500 |
| 46 | 0.04000 | 0.05796 | 0.05419 | 0.04850 |
| 47 | 0.04000 | 0.06659 | 0.06327 | 0.05000 |
| 48 | 0.06000 | 0.07353 | 0.07098 | 0.05300 |
| 49 | 0.06000 | 0.07803 | 0.07654 | 0.06300 |
| 50 | 0.06000 | 0.07959 | 0.07933 | 0.07250 |
| 51 | 0.10000 | 0.07803 | 0.07903 | 0.07200 |
| 52 | 0.09000 | 0.07353 | 0.07568 | 0.06600 |
| 53 | 0.02000 | 0.06659 | 0.06964 | 0.05900 |
| 54 | 0.04000 | 0.05796 | 0.06159 | 0.04700 |
| 55 | 0.08000 | 0.04847 | 0.05234 | 0.04300 |
| 56 | 0.01000 | 0.03895 | 0.04274 | 0.04200 |
| 57 | 0.04000 | 0.03007 | 0.03352 | 0.03800 |
| 58 | 0.06000 | 0.02229 | 0.02526 | 0.03050 |
| 59 | 0.00000 | 0.01587 | 0.01827 | 0.02800 |

Table 3.1: Discrete Convolution (Binom. Example)

| X | $\hat{p}(x)$ | $p(x)$ | $\tilde{p}(x)$ | $\tilde{p}(x)$ |
| :---: | :---: | :---: | :---: | :---: |
| 60 | 0.02000 | 0.01084 | 0.01268 | 0.02300 |
| 61 | 0.04000 | 0.00711 | 0.00845 | 0.01600 |
| 62 | 0.00000 | 0.00447 | 0.00540 | 0.01250 |
| 63 | 0.00000 | 0.00270 | 0.00331 | 0.00800 |
| 64 | 0.01000 | 0.00156 | 0.00195 | 0.00300 |
| 65 | 0.00000 | 0.00086 | 0.00109 | 0.00200 |
| 66 | 0.00000 | 0.00046 | 0.00059 | 0.00150 |
| 67 | 0.00000 | 0.00023 | 0.00030 | 0.00000 |
| 68 | 0.00000 | 0.00011 | 0.00015 | 0.00000 |
| 69 | 0.00000 | 0.00005 | 0.00007 | 0.00000 |
| 70 | 0.00000 | 0.00002 | 0.00003 | 0.00000 |
| 71 | 0.00000 | 0.00001 | 0.00001 | 0.00000 |
| 72 | 0.00000 | 0.00000 | 0.00001 | 0.00000 |
| $\geq 73$ | 0.00000 | 0.00000 | 0.00000 | 0.00000 |

Since $\hat{p}$ was never smoothed, it is not really fair to compare the error in $\hat{p}$ to that of $\tilde{p}$. If instead, we compute a smoothed version of $\hat{p}$ by assuming that the resulting distribution is binomial ${ }^{1}$, our results should be much better. This is relatively easy in this situation because a convolution of two instances of a binomial distribution with the same success rate $(\pi)$ is simply a binomial distribution with the trials equal to the sum of the trials for the convolved distributions. Therefore, if we convolve

[^2]$b(50, \pi)$ with $b(50, \pi)$ we get $b(100, \pi)$. We can then estimate $\pi$ from our empirical distribution by using the MLE estimator of $\pi$. This gives us an estimate of $\pi$ that is 0.504 . It should come as no surprise that this value is the same as the estimate we obtained from the individual distributions, since the expected value of the sum of two random variables is the sum of the expected values, and the MLE estimate of $\pi$ is based on the sample mean, divided by the number of trials. Before, we estimated $p$ by $\mathrm{E}_{X}[X] / N$, now we estimate it by $2 \mathrm{E}_{X}[X] / 2 N$.

### 3.3.1 Smoothing Using a Kernel

We were able to smooth the empirical distribution by estimating with an analytical distribution. However, often there is no obvious analytical distribution that is applicable to the data. In this case, we often want to smooth the empirical distribution. We can turn to a kernel function to help us smooth the distribution. One use of a kernel function is to smooth the distibution so that there are no "holes" in the empirical distribution. Holes are points of zero probability between points of non-zero probability, where we would expect the true distribution to have non-zero probability. For instance, in the empirical distribution of Figure 3.1, there is zero probability of the value 39 , while there is non-zero probability for 38 and 40 . In the real distribution, we can see that there is probability of 0.00711 for the value 39. Even without knowing the true distribution, we might suspect that the distribution should be somewhat bell shaped. Such an assumption leads to the belief that there should be no holes in intermediate values of the probability function.

There are many classes of kernel functions [30] [25]. The main requirements of a kernel function are that the sum of the kernel function over all values is equal to
one. For illustrative purposes, we use a simple discrete kernel function defined as

$$
\begin{equation*}
K_{j}(u)=\{(-2,0.15),(-1,0.2),(0,0.3),(1,0.2),(2,0.15)\} \tag{3.2}
\end{equation*}
$$

where $u$ is the signed distance between $x$ and the point $x_{j}\left(x-x_{j}\right)$. The kernel function is used to weight nearby frequencies. This particular kernel function eliminates holes that are within 2 units of a non-zero probability point. For instance, to compute the estimate of $F$ at $x=39$, we use the points $x=37 \ldots 41$. The estimate is defined by the formula

$$
\tilde{\tilde{p}}(39)=\sum_{i=-2}^{2} K_{(39)}(i) \hat{p}(39+i)
$$

where $K_{(39)}$ is the kernel function defined above.
The estimate at 39 is 0.009 . Using this kernel function for an estimate of $p$, denoted $\tilde{\tilde{p}}$, and the error measure of the previous sections, we have a total error of 0.00229. This error is, of course, the smaller of the two errors of the non-analytical estimates of $p$. Mostly, this is the result of the "filling" of the holes in the distribution. Due to the relatively small number of sample values in our empirical distribution, we see a marked increase in the accuracy of the estimate using smoothing. The larger the number of samples in the empirical distribution, the less the need for smoothing.

### 3.4 An Example

We now provide an example of the topics of this chapter using the data provided in Appendix B. Our first example will use the raw data from the Appendix, while our second will show a method for smoothing the frequency distribution prior to convolution. We will also contrast the effects of smoothing before the convolution to those of smoothing the convolution of unsmoothed empirical distributions. In both
the smoothed and unsmoothed versions of the distribution, we will use a discrete frequency distribution with values at the non-negative integers.

In order to confine our distribution to the non-negative integers, we will round each observation to the nearest whole number. Alternatively, we could scale the distribution, but since we are talking about the number of heads in a series of coin tosses, it makes sense to confine our results to whole integers). We then will count the number of observations at each whole number. We will then divide the count at each whole number by the total number of observations (100 for each distribution). The result will be the frequency corresponding to the point in the frequency distribution. After computing this frequency distribution, the result is shown in Tables B.2, B.3, B. 4 and B. 5 in Appendix B.

After preparing the unsmoothed frequency distributions of the four parts, we can compute the convolution of these frequency distributions. This convolution is computed in the traditional manner as described in this chapter. The result of this convolution is tabulated in Figure B. 6 in Appendix B. The convolution p.f. is shown in comparison to the empirical probability function in Figure 3.2.

We now wish to compute the convolution of the smoothed versions of the individual probability functions. We will use the smoothing function defined by Equation 3.2 above. A comparison of the lower mode of part A's empirical probability function before and after smoothing can be found in Figure 3.3.

Once we have computed the smoothed versions of each of the probability functions (listed in Appendix B as Tables B.7, B.8, B. 9 and B.10), we can compute the convolution of these smoothed distribution (shown in Figure B.11).

It can seen by the results of both the unsmoothed convolution and the convolution of the smoothed components that the distribution is much smoother and much more precise than the original empirical distribution of the total run times (the last


Figure 3.2: Unsmoothed Convolution Distribution


Figure 3.3: Smoothing Comparison for Part A
column in Figure B.1). In fact, from looking at the results of the unsmoothed convolution, one might suspect that smoothing is not even necessary in order to provide a relatively accurate representation of the total run-time distribution.

## Chapter 4

## Convolution by Discrete Fourier

## Transform

In the previous chapter, we described a method for computing the discrete convolution of two distributions in the traditional manner. In this chapter, we describe a method for computing the convolution that requires far less computation than the traditional method. Since this method is applicable to discrete distributions, it can be used on frequency distributions or discretized continuous distributions. This method is particularly applicable to frequency distributions built from large samples and smoothed frequency distributions. For frequency distributions built from smaller samples, smoothing should typically be performed before applying the methods described in this chapter. If such smoothing is not applied, the result of the convolution (by Discrete Fourier Transform) will suffer similar problems to those seen in the previous chapter.

### 4.1 The Discrete Fourier Transform

The Discrete Fourier Transform (DFT) of a function $f$ is defined by

$$
\begin{equation*}
\hat{\phi}_{k}=\hat{\mathcal{F}}[f](k)=\sum_{j=0}^{N-1} f_{j} \exp \left(\frac{2 \pi i j k}{N}\right) \tag{4.1}
\end{equation*}
$$

where $f_{j}=f(A+j \Delta x), \Delta x=\frac{B-A}{N}$, and $f$ is a periodic function defined over the interval $[A, B)$. While the definition of the DFT requires a periodic function, the theory can be applied to nonperiodic functions by a process of mirroring the function in other ranges. For instance, if the function $f$ is defined over $[A, B)$, then we can make the function $f$ a periodic function satisfying the requirements of the DFT by defining the function $\tilde{f}$ recursively as

$$
\begin{array}{ll}
\tilde{f}(x)=f(x) & \text { for } x \in[A, B) \\
\tilde{f}(x)=\tilde{f}(x-(B-A)) & \text { for } x \geq B  \tag{4.2}\\
\tilde{f}(x)=\tilde{f}(x+(B-A)) & \text { for } x<A .
\end{array}
$$

With these definitions, $\tilde{f}$ is a periodic function defined at the points $A \pm j \Delta x$ for all natural numbers $j$. The function $\tilde{f}$ corresponds to $f$ within the interval $[A, B)$.

The DFT has several useful properties. The Discrete Fourier Transform operator is a linear operator. This means that

$$
\begin{equation*}
\hat{\mathcal{F}}[a f+b g]=a \hat{\mathcal{F}}[f]+b \hat{\mathcal{F}}[g] \tag{4.3}
\end{equation*}
$$

for any scalars $a$ and $b$ and functions $f$ and $g$ suitable for DFT application. These functions must, of course, have the same period. If they do not have the same period, one of the functions may be zero padded to the same range as the other and then
converted to periodic functions so that they have the same period. Another property of the transform is that the transform of a convolution is the product of the transforms of the convolved functions. This is known as the Fourier Convolution Theorem and has application to both the DFT and the Continuous Fourier Transform (CFT). The discrete form of this theorem can be stated mathematically as

$$
\begin{equation*}
\hat{\mathcal{F}}[f * g]=\hat{\mathcal{F}}[f] \hat{\mathcal{F}}[g] \tag{4.4}
\end{equation*}
$$

where again, $f$ and $g$ are functions of the same period. The convolution operator $*$ is defined for two functions $f$ and $g$ as

$$
\begin{equation*}
(f * g)(x)=\sum_{t_{1}} \sum_{t_{2}} f\left(t_{1}\right) g\left(t_{2}\right) \delta\left(x-t_{1}-t_{2}\right) \tag{4.5}
\end{equation*}
$$

with $\delta(u)$ defined as

$$
\delta(u)= \begin{cases}1, & u=0  \tag{4.6}\\ 0, & u \neq 0\end{cases}
$$

### 4.2 The Fast Fourier Transform

The Fast Fourier Transform (FFT) is a fast algorithm for the implementation of the DFT. It requires an input array of a size that is a power of 2 . It is capable of computing the DFT for an array of size $N$ on the order $O(N \log N)$. With the use of the Fourier Convolution Theorem, this means we can compute the convolution of $M$ distributions on the order of $O(M N \log N)$ which is, of course, the same as $O(N \log N)$ as $M$ is assumed to be fixed for a given problem.


Figure 4.1: The DFT Method

### 4.3 The DFT Method

The method of convolution by DFT involves steps for building the input array to the FFT algorithm (for each distribution to be convolved), running the FFT algorithm, multiplying the results of the FFT together, and running the inverse FFT algorithm (nearly identical to the FFT algorithm) to compute the results of the convolution. The results are illustrated in Figure 4.1.

### 4.3.1 Step 1 - Prepare the FFT Input

As mentioned previously, the FFT algorithm requires an array of a length that is a power of two. Also, the points at which the FFT is evaluated must be equidistant. Since the results of the FFT will be used to evaluate the convolution of the distributions, the range of the arrays must be sufficient to represent all non-zero probability of the convolution distribution.

Consider, for instance, the distribution of the previous chapter. The initial distributions were approximately binomial and could have non-zero probability only in the range $[0,50]$. The convolution (since we were only doing one convolution) could
have non-zero probability on the range $[0,100]$ (from the sum of the minimum values to the sum of the maximum values). If we wish to represent only integral values (as we do), the smallest array we can use to represent the input items is the array with indices 0 through 127 (128 entries).

In order to prepare the input array for the binomial distribution, we build an array with 128 elements, zeroing all elements except those with probability. We would fill the array using the entries in Figure 3.1. After filling the array as appropriate, we would proceed to the other steps of the DFT method.

This step should be done for each input distribution. If convolving a distribution with itself, this step must be performed only once.

### 4.3.2 Step 2 - Compute the FFT

There are several commercially available and public domain implementations of the FFT algorithm. For a description of the algorithm, see [6], [5] or [24]. For the DFT method described in this chapter, the implementation used must only be compatible with the inverse implementation used. However, in the CFT method, attention must be paid to the FFT method used as there are some differences in the format of the output array between various implementations. We will discuss this further in Chapter 5.

This step should be performed for each of the outputs of Step 1. If the same distribution is being convolved multiple times, only one FFT computation is required for that distribution.

### 4.3.3 Step 3 - Multiply the FFT Output

Once the FFT is computed for each input distribution, there should be two arrays for each distribution (the real part of the complex value and the imaginary part). They should be the same length as the input arrays produced in Step 1. In this step, we must multiply the complex outputs of the FFT with each other, once for each distribution. For instance, if we have three distributions, the FFT results from the first must be multiplied (one entry at a time) with the FFT results from the second. This result (of the multiplication) must then be mupltiplied by the FFT results from the third distribution. When performing the multiplication, it should be performed as a functional multiplication, not a dot product. The results of the functional multplication will be a real / imaginary pair of arrays of the same length as the original input arrays.

This step will be performed one time for each distribution other than the first. Therefore, if there are $M$ distributions to be convolved, it will be performed $M-1$ times.

### 4.3.4 Step 4 - Compute the iFFT

Given the output of Step 3, a single pair of real / imaginary arrays, we compute the inverse FFT of these results to obtain the probability function for the resulting distribution. This probability function will be the discrete probability function corresponding to the convolution of the individual probability functions of the input distributions. As mentioned in Step 2, the implementation of the iFFT should come from the same source as the implementation of the FFT, but otherwise, the choice of implementation is most likely not significant. The output of the iFFT will be a
pair of real/imaginary arrays, but if everything was done with sufficient precision the resulting imaginary array should be zero in all elements.

### 4.4 An Example

Returning to our example dataset, we can compute smoothed discrete distributions from the sample points by grouping our data into integral values and computing frequency distributions over these values. We will use the Gaussian kernel function

$$
\begin{equation*}
K_{h}(\delta)=\exp \left(-\frac{\delta^{2}}{2 h^{2}}\right) \tag{4.7}
\end{equation*}
$$

with $h=10$ for each of the four samples (for parts A through D). Because we are not covering the entire space of possible values for the kernel function (the kernel has a range of $(-\infty, \infty)$ ), we will need to normalize the distributions after applying the kernel. This has the effect of redistributing all probability from the tails of the distribution to the center of the distribution in proportion to the probability function at each point. For instance, if the tails of the distribution account for $20 \%$ of the total probability, then each point of the distribution will be "grossed up" by $25 \%\left(\frac{1-0.8}{0.8}\right)$. Figure 4.2 contains several example points from the smoothed distributions.

Since it will be necessary to have arrays of a length that is a power of 2 in order to perform the FFT step, and these arrays must be of sufficient size to avoid aliasing, we choose an array length of 4096 entries. The approximate maximum of $A+B+C+D$ is 4020 or less.

Table 4.2: Example Points from the Smoothed Distributions

| X | $\hat{p}_{A}(x)$ | $\hat{p}_{B}(x)$ | $\hat{p}_{C}(x)$ | $\hat{p}_{D}(x)$ |
| ---: | :---: | :---: | :---: | :---: |
| 37 | 0.032216 | 0.018057 | 0.000065 | 0.000000 |
| 100 | 0.000000 | 0.000011 | 0.020315 | 0.000257 |
| 150 | 0.000000 | 0.000000 | 0.001029 | 0.023462 |
| 990 | 0.002727 | 0.000190 | 0.000000 | 0.000000 |
| 1045 | 0.000760 | 0.001757 | 0.000000 | 0.000000 |

### 4.4.1 Computing the DFT of the Probability Functions

Now that we have computed the smoothed versions of the probability functions for the sample dataset, we can use the FFT algorithm to compute the DFT of each of the input arrays. We use the Fourier Analysis feature of Microsoft Excel ${ }^{\circledR}$ to compute the FFT. Several example points from this calculation are given in Figure 4.3. The Excel implementation, as do most, produces the array with positive angles first, followed by negative angles.

Table 4.3: Example Points from the DFTs

| X | $\hat{\varphi}_{A}(x)$ | $\hat{\varphi}_{B}(x)$ | $\hat{\varphi}_{C}(x)$ | $\hat{\varphi}_{D}(x)$ |
| ---: | :---: | :---: | :---: | :---: |
| 0 | $1.0000+0.0000 i$ | $1.0000+0.0000 i$ | $1.0000+0.0000 i$ | $1.0000+0.0000 i$ |
| 1 | $0.8631-0.1924 i$ | $0.8817-0.1765 i$ | $0.9870-0.1575 i$ | $0.9737-0.2261 i$ |
| 2 | $0.7138-0.1145 i$ | $0.7704-0.1219 i$ | $0.9483-0.3106 i$ | $0.8963-0.4400 i$ |
| 3 | $0.8302-0.0181 i$ | $0.8825-0.0909 i$ | $0.8851-0.4550 i$ | $0.7721-0.6302 i$ |

Table 4.3: Example Points from the DFTs

| X | $\hat{\varphi}_{A}(x)$ | $\hat{\varphi}_{B}(x)$ | $\hat{\varphi}_{C}(x)$ | $\hat{\varphi}_{D}(x)$ |
| :---: | :---: | :---: | :---: | :---: |
| 4093 | $0.8302+0.0181 i$ | $0.8825+0.0909 i$ | $0.8851+0.4450 i$ | $0.7721+0.6302 i$ |
| 4094 | $0.7138+0.1145 i$ | $0.7704+0.1219 i$ | $0.9483+0.3106 i$ | $0.8963+0.4400 i$ |
| 4095 | $0.8631+0.1924 i$ | $0.8817+0.1765 i$ | $0.9870+0.1575 i$ | $0.9737+0.2261 i$ |

### 4.4.2 Computing the DFT of the Convolution

Once we have calculated the DFT of each of the four distributions, we can convolve the distributions' probability functions by multiplying their DFTs. The multiplication performed is a complex product. Some example points from this product are given in Figure 4.4.

Table 4.4: Example Points from the Convolution DFT

| X | $\hat{\varphi}_{A+B+C+D}(x)$ | X | $\hat{\varphi}_{A+B+C+D}(x)$ |
| ---: | :--- | :--- | :--- |
| 0 | $1.0000+0.0000 i$ | 2048 | $0.0000+0.0000 i$ |
| 1 | $0.5516-0.5717 i$ | 4093 | $0.2068+0.7008 i$ |
| 2 | $0.2604-0.4979 i$ | 4094 | $0.2604+0.4979 i$ |
| 3 | $0.2068-0.7008 i$ | 4095 | $0.5516+0.5707 i$ |
| 2047 | $0.0000+0.0000 i$ |  |  |

### 4.4.3 Computing the Probability Function

In order to complete our process of computing the estimate of the probability function of the sum distribution, we must invert the DFT of the convolution computed in the previous section. Fortunately, this is quite simple as all FFT implementations can be used also to invert the results of the FFT to compute the inverse DFT. In the previous section, we multiplied all of the four DFTs of the individual components together to compute the DFT of the convolution. We now compute the inverse DFT of that product in order to obtain the convolution probability function. We again use the FFT implementation built into MS Excel to compute this inverse transform, and give some results from the array in Figure 4.5.

Table 4.5: Example Points from the Probability Function

| X | $\hat{p}_{A+B+C+D}(x)$ | X | $\hat{p}_{A+B+C+D}(x)$ |
| :---: | :---: | :---: | :---: |
| 250 | 0.00015262 | 1380 | 0.00079850 |
| 300 | 0.00428857 | 1460 | 0.00003856 |
| 350 | 0.00913601 | 2200 | 0.00007509 |
| 1300 | 0.00181381 | 2300 | 0.00012591 |

### 4.5 Computational Cost of the Method

Referring to Figure 4.1, we see that there are $M$ computations of the FFT algorithm in order to compute the DFTs of the input distributions. Since the cost of the FFT algorithm is $O(N \log N)$, and $M$ is a constant for each given problem (not related to the number of points at which the distribution is evaluated), the computation of the

FFT of all $M$ distributions is also $O(N \log N)$. This concludes step 2 of the method. We purposely ignore step one of the method as smoothing is optional and the cost of smoothing is very dependent on the choice of smoothing method. Additionally, smoothing would presumably be necessary for any method to be used if the sample size warrants smoothing. The other parts of step one (organizing the data into equally spaced buckets and appending zeroes as necessary to extend the array length to a power of two) are of negligible cost and so are also ignored.

The cost of computing step 3's product is on the order of $O(N)$ for each multiplication, and by the same argument as that for step 2 is in total $O(N)$. Therefore, the overall cost of the combination of steps 2 and 3 is also $O(N \log N)$ as $N \log N$ dominates $N$. The action taken in step 4 is simply the execution of the FFT algorithm (with the inverse option) which has the same cost as the forward FFT. Therefore, it is also $O(N \log N)$ keeping our total cost at $O(N \log N)$. By comparison, the computation required for the discrete convolution method is $O\left(N^{2}\right)$ due to the nature of the convolution operation.

### 4.5.1 The Cost of Smoothing

It should be noted, with all this discussion of computation cost of the method, that the choice of kernel function is very important if smoothing is to be performed. For example, the kernel we chose (Equation 4.7) for this example is very costly in terms of computation time. Ignoring the cost of computing the exponentials in the function, the evaluation of $p(x)$ using this kernel function requires evaluation at every sampled point for each frequency bucket. This feature alone makes the smoothing operation on the order of $O\left(N^{2}\right)$. On the other hand, the kernel used in the binomial example of Chapter 3 (see Equation 3.2) puts the smoothing operation on the order of $O(N)$.

From the results of the discrete unsmoothed convolution in Chapter 3, we can see that convolution is not even necessary if a less smooth resulting is satisfactory for the purposes of the analysis. If smoothing is desirable, but the performance burden is too great, a discrete kernel such as the one in Equation 3.2 can be used on the output of the convolution. The results of such a late smoothing will not be as profound as those of an early smoothing, but will still help to smooth the distribution a bit. Of course, if the goal of computing the output distribution is to compute a particular percentile of the distribution, such smoothing is not likely to be very useful.

## Chapter 5

## Convolution by Characteristic

## Function

In the previous chapter, we introduced a method of computing the convolution of several component distributions by means of the Discrete Fourier Transform (DFT). In this chapter, we present a method by which the convolution can be computed by means of the Continuous Fourier Transform (CFT) of the probability function of a continuous distribution. The CFT of the probability density function is known in probability theory as the Characteristic Function (CF). Accordingly, we will use the terms CFT and CF interchangeably throughout this chapter.

In order to use the method presented in this chapter, we must approximate the input distributions with a continuous distribution for which the CF is known. Unfortunately, the CF is not known in closed form for all distributions. A well known example of the lack of closed form CF is that of the Lognormal Distribution. The Lognormal is known (see [16] and [18]) to have a CF that is difficult to compute, due to the inability to expand the CF as a Taylor series based on the moments. The Taylor series can be shown to diverge and as such makes the CF difficult to compute. At
this time, the current body of research seems to imply a tendency towards believing a closed-form representation of the CF that is useful for computational purposes may not exist. At the very least, it has not yet been found.

However, for most distributions, the CF is known and is relatively easy to compute. For those distributions that have an easy to calculate CF, we will use the analytical CF. For those that have a CF that is difficult to compute, we will use the DFT of a discretized version of the PDF in place of the CF.

### 5.1 CFs of Common Distributions

In this paper, we have used three common continuous distributions: Normal, Gamma, and Lognormal. Of these three distributions, the first two have closed-form analytical CFs that can be easily computed. The third - Lognormal - does not. For the Normal distribution, the CF is given by

$$
\begin{equation*}
\varphi(t)=e^{i \mu t-\sigma^{2} t^{2} / 2} \tag{5.1}
\end{equation*}
$$

where $\mu$ and $\sigma$ are the well-known parameters of the distribution. For the Gamma distribution, given by probability density function

$$
\begin{equation*}
f(x)=\frac{1}{\Gamma(\alpha) \beta^{\alpha}} x^{\alpha-1} e^{-x / \beta} \tag{5.2}
\end{equation*}
$$

the CF is given by

$$
\begin{equation*}
\varphi(t)=\frac{1}{(1-i \beta t)^{\alpha}} . \tag{5.3}
\end{equation*}
$$

In Chapter 2, we discussed mixture distributions and their properties. Given a mixture distribution with the probability density function

$$
\begin{equation*}
f_{Z}(z)=\sum_{j=0}^{M-1} \pi_{j} f_{X_{j}}(z) \tag{5.4}
\end{equation*}
$$

which represents a mixture with $M$ components having the density functions $f_{X_{j}}$ and mixture proportion $\pi_{j}$, a little algebra will show that the CF of the mixture is given by

$$
\begin{equation*}
\varphi_{Z}(t)=\sum_{j=0}^{M-1} \pi_{j} \varphi_{X_{j}}(t) \tag{5.5}
\end{equation*}
$$

where the $\varphi_{X_{j}}$ are the CFs corresponding to the $f_{X_{j}}$.

### 5.2 The CFT Method

The characteristic function method described in this chapter is similar to the DFT method described in the previous chapter. The major difference between the methods is that the CF method requires fitting a continuous distribution to the data of the component distributions, and then uses these resulting distributions to compute the CF of the convolution. A pictoral description of the method is given in Figure 5.1.

Both the DFT method and the CF method eventually use the inverse FFT to compute the approximate probability function. However, the CF method assumes that a continuous distribution more accurately represents the underlying population from which the sampled times come. Once these continuous distributions are estimated, the CF method samples the CF of the convolution by computing the CFs of each of the components at the same set of points at which the iFFT is to be evaluated. Essentially, we are sampling in the frequency domain, rather than in the time domain, as is done in the DFT method.


Figure 5.1: The CF Method

The basic steps of the method are

1. Select appropriate continuous distributions for each of the components and determine the MLE estimates of their associated parameters,
2. Sample the CF of each of these component distributions at the appropriate points for iFFT application,
3. Compute the product of the CFs at the calculated points,
4. Use the results of step 3 as input to the iFFT algorithm. Normalize the output. The resulting array is an approximation to the probability function of the convolution.

### 5.2.1 Step 1 - Estimate the Distributions

In Chapter 2, we described how to estimate the underlying component population distributions with continuous distributions. We tabulated our results in Section 2.4. Using these results, we can find that for the distributions of Parts $A$ and $B$, we need the CF of the Lognormal and Normal distributions respectively. Since, as we mentioned in the beginning of this chapter, the Lognormal CF is not easy to calculate,
we will leave Part A out of our current consideration. The next chapter will show how to integrate the DFT and CF methods so that distributions for which the CF is unknown can also be used. For parts $C$ and $D$, we need the Gamma and Lognormal CFs respectively. Again, we will defer the computation of part $D$ to the next chapter. Since we are deferring parts $A$ and $D$ to the next chapter, we will compute the convolution of $B$ and $C$ in this chapter.

For part $B$, the CF is the CF of a mixture of two Normal (Gaussian) components, with mixture proportions and parameters given in Figure 2.15. The resulting CF is

$$
\begin{align*}
\varphi_{B}(t) & =0.89 \exp \left(48.94685 t i-(8.38651 t)^{2} / 2\right) \\
& +0.11 \exp \left(1057.17273 t i-(30.68959 t)^{2} / 2\right) \tag{5.6}
\end{align*}
$$

where $i=\sqrt{-1}$. With a little algebra, it can be shown that the real and imaginary parts of $\varphi_{B}$ are

$$
\begin{align*}
& \Re \varphi_{B}(t)=\pi_{1} \cos \left(\mu_{1} t\right) e^{-\frac{\sigma_{1}^{2} t^{2}}{2}}+\pi_{2} \cos \left(\mu_{2} t\right) e^{-\frac{\sigma_{2}^{2} t^{2}}{2}}  \tag{5.7}\\
& \Im \varphi_{B}(t)=\pi_{1} \sin \left(\mu_{1} t\right) e^{-\frac{\sigma_{t}^{2} t^{2}}{2}}+\pi_{2} \sin \left(\mu_{2} t\right) e^{-\frac{\sigma_{2}^{2} t^{2}}{2}} \tag{5.8}
\end{align*}
$$

where the $\pi_{j}, \mu_{j}$, and $\sigma_{j}$ are those from Figure 2.15. Using the equations above, we can calculate the entries in the real and imaginary arrays that are needed for input into the iFFT algorithm.

Similarly, we can compute the CF of part $C$ 's distribution (quite easily) as

$$
\begin{equation*}
\varphi_{C}(t)=\left(\frac{1}{1-3.47122 i t}\right)^{29.72568} \tag{5.9}
\end{equation*}
$$

which has real and imaginary parts given by

$$
\begin{align*}
& \Re \varphi_{C}(t)=\left(1+\beta^{2} t^{2}\right)^{-\alpha / 2} \cos (\alpha \arctan \beta t)  \tag{5.10}\\
& \Im \varphi_{C}(t)=\left(1+\beta^{2} t^{2}\right)^{-\alpha / 2} \sin (\alpha \arctan \beta t) \tag{5.11}
\end{align*}
$$

### 5.2.2 Step 2 - Sample the CF

In this step, we need to sample the CFs determined in step 1 for each of the component distributions (in our current example, just parts $B$ and $C$ ) at the appropriate points for evaluation of the iFFT. At this point, we need to discuss what such appropriate points might be. The FFT algorithm that we are using outputs two arrays, a real array and an imaginary array. These arrays are the real and imaginary parts of the DFT, respectively. Since we are skipping the application of the FFT and going straight to the CF, but wish to still use the iFFT, we must put our CF samples in the same form expected by the iFFT algorithm (which, of course, is the same form output by the FFT algorithm).

The astute reader may notice that we started with samples, so it might seem strange that we are again sampling. The reason we are sampling again is that the original sample data was used to estimate the parameters for the analytical distribution we use to represent the component, and the sampling we are doing here is to transition from the analytical CF to an array of values that can be used in the FFT algorithm.

The order of the elements in the arrays is, of course, significant to the iFFT algorithm. The FFT implementation we use for this paper puts the positive part of the DFT domain $(0,1, \ldots, N / 2-1)$ in the first $N / 2$ entries of the array, and the negative part of the domain in the second $N / 2$ entries. The negative part of the
domain is output in normal order, so the element at index $N / 2$ of the output array (0-based) corresponds to the DFT at $-N / 2$, and index $N / 2+1$ to $-(N / 2-1$ ), and so on. It should also be noted that the true domain of the Fourier Transform is from $-\pi$ to $+\pi$ in the frequency domain, and that this means that an argument of $k$ (in the DFT domain) corresponds to $k \frac{2 \pi}{N}$ in the Fourier domain.

As mentioned in Chapter 1, the definition of the Fourier Transform also comes into play. Our definition of the Fourier Transform (Equation 1.6) does not correspond to the one used by our FFT algorithm (MS Excel). To be precise, the CFT as we defined it is evaluated at the opposite of the argument at which Excel evaluates. Therefore, we must sample at the points $t^{*}=-t$ when sampling our CF.

### 5.2.3 Step 3 - Multiply the Results

We are now armed with the facts necessary to build the CF sample that corresponds to the output of the FFT algorithm (and of course to the input of the iFFT). Since we wish to compute the convolution of part $B$ 's distribution with that of part $C$, we simply multiply the CFs of $B$ and $C$ at each point at which we evaluate them. Accordingly, we need to evaluate the CF

$$
\begin{equation*}
\varphi(t)=\varphi_{B}(t) \varphi_{C}(t) \tag{5.12}
\end{equation*}
$$

at the points

$$
\begin{equation*}
t^{*}=-t=-\{[(j+N / 2) \quad \bmod N]-N / 2\} \frac{2 \pi}{N} \tag{5.13}
\end{equation*}
$$

for $j=0,1, \ldots, N-1$.
In order to do this calculation, we need to multiply $\varphi_{B}(t)$ with $\varphi_{C}(t)$. Let us compute one of the entries from our set of arrays as an example of the computation.

For example, for $j=3$, we have

$$
\begin{aligned}
t & =\{[(3+4096 / 2) \quad \bmod 4096]-4096 / 2\} \frac{2 \pi}{4096} \\
& =0.00460194
\end{aligned}
$$

so that

$$
\begin{aligned}
& \Re \varphi_{B}(-t)=0.88343247 \\
& \Im \varphi_{B}(-t)=-0.09099224 \\
& \Re \varphi_{C}(-t)=0.88601410 \\
& \Im \varphi_{C}(-t)=-0.45543744
\end{aligned}
$$

and

$$
\begin{aligned}
\Re \varphi(-t) & =\Re \varphi_{B}(-t) \Re \varphi_{C}(-t)-\Im \varphi_{B}(-t) \Im \varphi_{C}(-t) \\
& =0.74129235 \\
\Im \varphi(-t) & =\Re \varphi_{B}(-t) \Im \varphi_{C}(-t)+\Im \varphi_{B}(-t) \Re \varphi_{C}(-t) \\
& =-0.48296863 .
\end{aligned}
$$

### 5.2.4 Step 4 - Calculate the iFFT

The values of $t^{*}$ given by Equation 5.13 range from $-\pi$ to $\pi$ in equal steps. Having computed all of the entries in the array, we can compute the iFFT of this array to invert the CF of the convolution distribution.

Given the output of Step 3 (an array of numbers in FFT-output order), we can compute the iFFT of this array and obtain the estimate of the probability function


Figure 5.3: The Convolution PDF of $\mathrm{B}+\mathrm{C}$
for $B+C$. Some selected entries from the resulting probability function are given in
Figure 5.2, and a plot of the results are given in Figure 5.3.
Table 5.2: Selected Entries from $\hat{f}_{B+C}$.

| $z$ | $\hat{f}_{B+C}(z)$ | $z$ | $\hat{f}_{B+C}(z)$ | $z$ | $\hat{f}_{B+C}(z)$ |
| ---: | :---: | ---: | :---: | :---: | :---: |
| 100 | 0.00041065 | 101 | 0.00048661 | 102 | 0.00057386 |
| 148 | 0.01731210 | 149 | 0.01734352 | 150 | 0.01733304 |
| 151 | 0.01728132 | 210 | 0.00054769 | 1148 | 0.00115884 |
| 1153 | 0.00119979 | 1160 | 0.00121896 | 1165 | 0.00120454 |
| 1172 | 0.00114709 | 1178 | 0.00106790 | 1182 | 0.00100294 |

### 5.3 Computational Cost of the Method

The computational cost of computing the CFT is $O(N)$ for each component distribution, once the component distribution is approximated with a continuous distribution. Even in the case of the Lognormal distribution, the order cannot be $O\left(N^{2}\right)$ to compute the characteristic function because $N$ is not related in any way to $\varphi$ once the fit is performed (the only parameters to $\varphi$ are $t, \sigma$, and $\mu$ ). Since again we assume there is a fixed number of convolutions to be performed, the total convolution process is $O(N)$. However, since we need to compute the iFFT to return to the "time" domain (the $x$-domain), we need one iFFT computation at $\operatorname{cost} O(N \log N)$. This brings our total cost of computation to $O(N \log N)$. While this total cost is the same as the DFT method, we believe that when the DFT method is used on continuous fitted distributions (as this method is), the actual costs will be less for this method, assuming the CFT is not very difficult to compute. The difference in the cost of computation between the two methods is based on the difference between the cost of computing the CFT for the values of $t$ needed versus the cost of sampling and computing the DFT at the sampled points. In the case of using the DFT method with raw sampled data (without MLE fitting the component distributions and sampling from the probability functions) or smoothed data with a very simple smoothing kernel, the DFT method will probably be faster, though it may be less accurate (depending on the sample size).

## Chapter 6

## Combining the DFT and CFT

## Methods

In the previous two chapters, we described methods for computing an approximation to the convolved distribution by use of the DFT and CFT respectively. In Chapter 5, we could not apply the CFT method to Part $A$ or Part $D$ due to the fact that the Lognormal distribution does not have an easy to compute closed form CF ( $\varphi$ ). In this chapter, we show how to use the DFT method to compute the approximation to the CF of these two parts, while using the CFT method for the other two parts. The total cost of this method will also be $O(N \log N)$ since each of the two parts are of this order.

### 6.1 The Combined Method

Essentially, this method uses the DFT method from Chapter 4 to compute the convolution of Parts $A$ and $D$, while using the CFT method from Chapter 5 to compute the
convolution of Parts $B$ and $C$. The application of the DFT method to this problem is slightly different in this method, however.

### 6.1.1 Preparing the Distributions

In order to apply the DFT method in this problem, we would like to prepare our distributions of Parts $A$ and $D$ to match, as closely as possible, the results that we would get from applying the CFT method to these distributions. Instead of applying smoothing to our samples or using the DFT method directly on the raw sampled data, we wish to prepare input arrays with the probability functions of interest. In this case, we wish to prepare input arrays for Part $A$ that closely match the probability function of the Lognormal mixture that we fit to Part $A$ 's data. In the case of Part $D$, we wish to similarly prepare input arrays that closely match the probability function of the Lognormal distribution that we fit to the data.

In order to prepare the arrays, we sample the probability function of the applicable distributions at the appropriate points in the array. Since both of these distributions have the positive real line as their domain, we must cut them off at a certain point to avoid aliasing problems in the FFT. We will discuss the procedure for aliasing avoidance later in this chapter.

### 6.1.2 Computing the Convolution Distribution

After sampling from the probability function, we will apply the DFT method as in Chapter 4, obtaining the results of the convolutions of the distributions that are difficult for the CFT method. We will not apply the final step of the DFT method, however, so we will leave the results in the Fourier domain (the products of the FFTs).

For the distributions to which we can apply the CFT method, we apply the method. Again, as in the application of the DFT method to the other distributions, we will not apply the final step. Thus, after applying the CFT method, minus the final step, we will have the product of the approximated CFTs of the probability functions in the Fourier domain. We may now complete our computations by computing the product of the results from the DFT method and those from the CFT method. This product will be the approximated CFT of the total convolution, to which we may now apply the iFFT once (the final step of both methods). The result of the iFFT of this product will be the approximation of the probability function of the convolution.

### 6.2 An Example

We will use the CFT results obtained in Chapter 5 without adjustment, as they are precisely the same results that would be obtained by the combined method for Parts $B$ and $C$. We take the results just prior to applying the final iFFT in the CFT method, as prescribed in Section 6.1. For the DFT results, we must compute the sampled probability functions and then perform the DFT method on them, stopping prior to the final step.

The probability function for Part $A$ is the mixture probability function

$$
\begin{equation*}
f_{Z}(z)=\frac{\pi_{1}}{z \sigma_{1} \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{\log z-\mu_{1}}{\sigma_{1}}\right)^{2}}+\frac{\pi_{2}}{z \sigma_{2} \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{\log z-\mu_{2}}{\sigma_{2}}\right)^{2}} \tag{6.1}
\end{equation*}
$$

where the $\pi_{i}, \mu_{i}$, and $\sigma_{i}$ are those given in Figure 2.15. Whereas the probability function for Part $B$ is the lognormal probability function given by

$$
\begin{equation*}
f_{X}(x)=\frac{1}{x \sigma \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{\log x-\mu}{\sigma}\right)^{2}} \tag{6.2}
\end{equation*}
$$

where $\mu$ and $\sigma$ are those given in Figure 2.16.

We need to compute these functions at the same input values as our other distributions, but we need to clamp the distributions values to a range that will not allow overflow. For instance, since we chose an output array size of 4096 values (range of $[0,4095]$ ) we could restrict each of our distributions (rather arbitrarily) to a range of $[0,1023]$. However, this would not be the best choice, since two of our distributions have significant probability outside of this range and the others don't. A more appropriate choice might be to limit the two mixture distributions (which have significant probability at values above 1000) to the range $[0,1499]$ and the other two distributions to the range [0,499]. This way, the largest values from each of the distributions cannot produce a value with positive probability above 4095. Sometimes, depending on the tail weight of the distributions in question, such clamping is not necessary. This happens to be the case in our situation, since the probability of the total resulting distribution exceeding 4095 is negligible.

Once we have evaluated the probability function at the points in our array, we must normalize the array (so it sums to unity) by dividing each entry by the sum of all entries in the array. This, in effect, creates a discrete frequency distribution that is similar to the original continuous distribution that can be used as input to the FFT algorithm. We then apply the FFT algorithm to each of the two newly discretized distributions to obtain approximations to the CFTs. According to the DFT method and the adjustments prescribed in this chapter, we should also take the complex product of the two sets of output arrays prior to using the results in the combined method, but this is not important. We could just as easily treat these output arrays as sampled versions of the CFTs of the two distributions, and use these CFTs in the the multiplication step of the CFT method directly. In either case, once we take the product of these results with the product of the CFTs given by the CFT method, we can then apply the iFFT algorithm to obtain the probability function of


Figure 6.1: The Convolution PDF of all parts
the convolution. The results of applying this method for the example program are shown in Figure 6.1.

## Chapter 7

## Testing the Models

We have described how to compute estimates of the distributions of the total execution time using several methods. Now, we wish to test the resulting distributions by using them to predict the WCET of the program for a set of simulated values (from the same distributions used to generate our testing data). We will compare the average excess time for our prediction and the number of times our deadline is missed to the excess time for the Gumbel estimate. We will also use our estimated distribution to compute a revised Gumbel estimate that makes use of the more detailed information provided by our methods.

### 7.1 The Test Distributions

The distributions used to generate the training data provided in Appendix B are listed in Figure 7.1. From these test distributions, we simulate a test data set that is used to evaluate the WCET bounds determined several different ways. We simulate 100000 sample values from the total run-time distribution by simulating 100000 ordered 4 -tuples with an element from each of the four component distributions. For
the two mixture distributions, the element is determined by simulating a uniform random variate (on the unit interval) and one deviate from each of the two mixture component distributions. We choose the deviate from the first mixture component if the uniform deviate is less than $\pi_{1}$, the second otherwise. Such a method accurately models the mixture distribution as a random (nonuniform) sample from two separate distributions.

Table 7.1: Underlying Test Distributions

| Part | Distribution | Parameters |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| A | Mixture (Normal) |  | $=0.8$ | $\pi_{2}=$ | 0.2 |
|  |  |  | $=40$ | $\mu_{2}=$ | 1000 |
|  |  | $\sigma_{1}$ | $=2$ | $\sigma_{2}=$ | 20 |
| $B$ | Mixture (Gamma) | $\pi_{1}$ | $=0.85$ | $\pi_{2}=$ | 0.15 |
|  |  |  | $=25$ | $\alpha_{2}=$ | 350 |
|  |  |  | $=2$ | $\beta_{2}=$ | 3 |
| C | Gamma |  | $\alpha=25$ | $\beta=4$ |  |
| D | Lognormal |  | $\mu=5$ | $\sigma=0.1$ |  |

Due to space constraints, the simulated values are not tabulated in this paper, but are available on request from the author.

### 7.2 WCET Estimates

We use several WCET estimates and compare their performance in terms of the number of times the WCET bound is exceeded by our samples, the average value of
observations that exceed the bound, and the average excess of the bound over the total run-time. We contest that the WCET bound given by [7] is very high and leads to wasted cycles when used for scheduling. However, the percentile method we wished to use for our estimate of the WCET leads to bounds too low for practical use, as a substantial number of observations exceed this bound. We instead recommend a smaller-sample extreme value (EV) estimate similar in character to the Gumbel estimate, except using the exact (approximated) EV distribution of a finite sample size. The results of these various estimates of the WCET are compared in Figure 7.2. The estimation methods for the WCET tabulated in Figure 7.2 are

Gumbel (sample). The Gumbel distribution WCET estimate (as described in [7] with corrections ${ }^{1}$ ) based on the original sample of total run-times,

Gumbel (distribution). The same as Gumbel (sample), except using estimates of $\mu$ and $\sigma$ from the estimated total run-time distribution given by our method,

Percentiles. The percentiles of the total run-time distribution as estimated by our method,

Linear Combination. The WCET estimate given by a linear combination of the the Percentile estimate and the Gumbel (distribution) estimate using the coefficients 0.75 and 0.25 respectively,

[^3]Gumbel (max locmax). The Gumbel WCET estimated found by using a $\mu$ and $\sigma$ estimated from only the portion of the distribution where the largest (in $x)$ local maximum is found in the density,
$\mathbf{F}^{\mathbf{k}}(\mathbf{x})$. The WCET estimate given by the specified percentile of the excess value distribution (see Section 7.3) for sample size $k$.

### 7.3 The Exact EV Distribution

The Gumbel distribution used in [7] is the asymptotic distribution of the maximum value from a distribution. It represents the limit as the sample size tends to infinity of the maximal outcome in a random sample from any exponential family distribution (nearly all distributions of interest for our purposes are from this family). In simple terms, it is the probability distribution for the largest value that one would ever expect to see, given a few parameters about an existing sample. Because this distribution is asymptotic, it is very conservative in the bounds it provides. As can be seen from the data in Figure 7.2 for the Gumbel (sample) versus the $F^{k}(x)$ estimates, the average amount by which the WCET from the Gumbel method exceeds the actual run-time is much greater than that from the exact method.

The exact method is based on the same idea as the Gumbel distribution, but makes use of the actual distribution of the individual observations, rather than just the sample and a few sample statistics. Because of this, the real EV distribution based on our estimate of the distribution function $F(x)$ of the total run-time distribution provides a much tighter bound on the WCET. In general, the distribution function (probability that the variate does not exceed a given value) $F_{x_{(k)}}(x)$ of the maximum
value $x_{(k)}$ of a sample of size $k$ is given by

$$
\begin{equation*}
F_{x_{(k)}}(x)=\left(F_{X}(x)\right)^{k} \tag{7.1}
\end{equation*}
$$

where $F_{X}(x)$ is the distribution function underlying the sample observations themselves.

The EV method used here raises the estimated total run-time distribution function $F(x)$ (of the convolution) to various powers of $k$ and locates the $p^{\text {th }}$ percentile of this distribution. The percentile is then used as the WCET estimate, providing a much smaller bound than the Gumbel distribution. The method allows adjustment both by changing $k$ and by changing the desired percentile. Also, the underlying accuracy of $F(x)$ can be controlled by changing the number of points at which the CFs are sampled (or at which the FFT is applied in the discrete case).

Table 7.2: WCET Estimates and Results

| Estimate Type | Confidence | WCET <br> Bound | Count Over | Average Over | Average Excess |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Gumbel (sample) | 90\% | 3193 | - | - | 2607.37 |
|  | 95\% | 3448 | - | - | 2862.37 |
|  | 97.5\% | 3703 | - | - | 3117.37 |
|  | 99\% | 4039 | - | - | 3453.37 |
| Gumbel <br> (distribution) | 90\% | 3172 | - | - | 2586.37 |
|  | 95\% | 3420 | - | - | 2834.37 |
|  | 97.5\% | 3668 | - | - | 3082.37 |
|  | 99\% | 3994 | - | - | 3408.37 |
| Percentiles | 90\% | 1333 | 9769 | 1511.84 | 764.84 |
|  | 95\% | 1366 | 4971 | 1669.77 | 795.47 |
|  | 97.5\% | 1400 | 2423 | 1973.97 | 828.28 |
|  | 99\% | 2294 | 974 | 2334.02 | 1708.76 |
| Linear Combination | 90\% | 1793 | 1511 | 2309.68 | 1215.18 |
|  | 95\% | 1880 | 1511 | 2309.68 | 1300.87 |
|  | 97.5\% | 1967 | 1511 | 2309.68 | 1386.55 |
|  | 99\% | 2719 | , | 230.68 | 2133.37 |
| Gumbel <br> (max locmax) | 90\% | 2427 | 9 | 2435.47 | 1841.37 |
|  | 95\% | 2451 | , | 2453.64 | 1865.37 |
|  | 97.5\% | 2475 | - | - | 1889.37 |
|  | 99\% | 2506 | - | - | 1920.37 |
| $\begin{gathered} F^{k}(x) \\ k=1,000 \end{gathered}$ | 90\% | 2421 | 11 | 2433.65 | 1835.37 |
|  | 95\% | 2432 | 7 | 2437.48 | 1846.37 |
|  | 97.5\% | 2443 | 1 | 2453.64 | 1857.37 |
|  | 99\% | 2455 | - | - | 1869.37 |
| $\begin{gathered} F^{k}(x) \\ k=10,000 \end{gathered}$ | 90\% | 2455 | - | - | 1869.37 |
|  | 95\% | 2464 | - | - | 1878.37 |
|  | 97.5\% | 2473 | - | - | 1887.37 |
|  | 99\% | 2484 | - | - | 1898.37 |
| $\begin{gathered} F^{k}(x) \\ k=100,000 \end{gathered}$ | 90\% | 2483 | - | - | 1897.37 |
|  | 95\% | 2492 | - | - | 1906.37 |
|  | 97.5\% | 2499 | - | - | 1913.37 |
|  | 99\% | 2509 | - | - | 1923.37 |
| $\begin{aligned} & F^{k}(x) \\ k= & 1,000,000 \end{aligned}$ | 90\% | 2509 | - | - | 1923.37 |
|  | 95\% | 2517 | - | - | 1931.37 |
|  | 97.5\% | 2524 | - | - | 1938.37 |
|  | 99\% | 2535 | - | - | 1949.37 |

## Chapter 8

## Conclusion

We have shown that the use of our DFT and CFT methods are an efficient alternative to the traditional methods of computing the convolutions that arise in the estimation of run-time distributions built of several components. Additionally, we have shown that a tighter bound on the WCET can be provided by assuming a less-than-infinite sample size and using distributional estimates of the run-time distribution, along with the theoretical fixed-sample distribution of the maximum observed value (the $F^{k}(t)$ distribution function).

### 8.1 Future Work

In the development of WCET estimates, it would be desirable to have estimates of, or bounds on, the error of this estimate. Of course, an exact error estimate or bound requires an assumption for the number of times the task is to be run. Alternatively, if an asymptotic estimate is made, the number of times the task is to be run is assumed to be infinite. This assumption of the number of runs should be consistent with the underlying assumption of the WCET estimate.

In the estimate of the WCET error, it would also be desirable to understand the relationship between the input parameters relating to sampling, bucketization, and distribution choice and the underlying error of the estimate. Also, since there is obviously a relationship of some sort between these items and the error of the estimate, it would be nice to have a method of optimizing the choice of these parameters in order to minimize the resulting error.

In our work, we have assumed that a breakdown of the task into independently distributed components has been chosen appropriately. However, we have made no effort to develop methods for choosing this breakdown. It would be beneficial to have some theoretical support and some well defined algorithms or heuristics for the choice of this breakdown and the evaluation of the independence of the elements.

### 8.2 Some Observations

During the course of our research, we have made some observations for which we have no theoretical support, but nonetheless would like to present here. First of all, we have noticed that the separation of the task into several independent components has a profound effect on the sampling burden needed for confidence in the resulting distribution. We call this effect the independent sampling effect (ISE). The ISE can be understood through a simple example. Consider a task that can be broken up into three independent components. If we sample each of these components 10 times, this is equivalent (under the assumption of independence) to sampling the total distribution 1000 times.

The obvious advantage of the ISE is the number of samples required for confidence when using independent components is much smaller than the number of samples required for the same level of confidence in an estimate of only the total
run-time distribution. This greatly reduces the sampling burden and allows us to make relatively accurate estimates of the total run-time distribution with relatively few sampling runs of the program. Using the example of the last paragraph, we could run the program ten times, sampling it as three separate components (a total of 30 samples), and have the same level of confidence in the resulting estimate as we would have in an estimate using only the total run-time and a sample size (number of program executions) of 1000 runs. Of course, if it is expensive to run the program, the ISE is of substantial benefit to us.

In addition to the obvious advantage of the ISE, another advantage is the exhaustion effect of the ISE. For instance, in the case of the three components, 10 samples each as given in previous paragraphs, consider that two of the three components had distributions that are mixtures with a high and a low distribution. Consider also that the high elements of the mixtures are much less likely to occur than the low elements, for instance an $80 \%$ likelihood of the low element and a $20 \%$ likelihood of the high. Consider, as would be expected, that two of the observations from the 10 samples of each of these two components arose from the high distribution, but that none of the total run-time samples (runs) exhibited high elements from both of the two mixture distributions. In this case, the Gumbel method would use a maximum as a parameter of its estimate that is actually not anywhere close to the maximum of the theoretical distribution, since none of the runs exhibited the "high-high" behavior. However, when the components are convolved and the ISE occurs, an "artificial" maximum arises from the pairs of high samples that were encountered. Even though they occurred on different runs, the assumption of independence and the ISE allowed them to be used in the analysis together. In this small-sample case where there are gaping holes in the sample distribution, the ISE provides a great benefit and leads to a much greater accuracy in our estimate of WCET.

One final conclusion related to the ISE is that increasing the number of components into which the total distribution is broken gives much greater benefit than increasing the number of runs of the program used for sampling. This is a very nice result, assuming good methods are available for choosing appropriate breakdowns into independent components. One caveat, however, is that there is always a risk of reducing the sample size (number of sampling runs) to a point that features of the component distributions are missed. For instance, in the example of this chapter, if the sample size is reduced by too much, the likelihood of seeing any "high" elements in the components becomes relatively small. Also, if the "high" elements are much less likely (for instance $1 \%$ of the time, rather than $20 \%$ as in our example), the sampling burden is much higher on the component level and in turn the number of runs required will be much higher. That being said, the number of runs for the same level of confidence at the total run-time level only would be much higher, again due to the ISE.

## Appendix A

## Maximum Likelihood Estimation

In this appendix, we derive formulas and methods for the MLEs of the three distributions we use in this thesis. In all sections of this appendix, we assume the following:

1. a set of observations $\boldsymbol{x}=\left\{x_{i}\right\}_{i=1}^{N}$,
2. a random variable $X$ representing the distribution from which the observations are sampled, and
3. a vector of parameters $\boldsymbol{\Psi}$ for MLE fitting.

## A. 1 Normal Distribution

The MLE vector for the Normal distribution is $\boldsymbol{\Psi}=(\mu, \sigma)$. We wish to compute the estimate of this vector, $\hat{\boldsymbol{\Psi}}=(\hat{\mu}, \hat{\sigma})$. The likelihood function $L(\boldsymbol{\Psi})$ is defined by

$$
\begin{equation*}
L(\boldsymbol{\Psi} ; \boldsymbol{x})=\prod_{i=1}^{N} \frac{1}{\sigma \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{x-\mu}{\sigma}\right)^{2}} \tag{A.1}
\end{equation*}
$$

The $\log$-likelihood function $\ell(\boldsymbol{\Psi} ; \boldsymbol{x})$ is given by

$$
\begin{equation*}
-\ell(\boldsymbol{\Psi} ; \boldsymbol{x})=N \log \sigma \sqrt{2 \pi}+\frac{1}{2 \sigma^{2}} \sum_{i=1}^{N}\left(x_{i}-\mu\right)^{2} \tag{A.2}
\end{equation*}
$$

with partial derivatives (with respect to the parameters) given by

$$
\begin{equation*}
\frac{\partial \ell(\boldsymbol{\Psi})}{\partial \mu}=\frac{1}{\sigma^{2}} \sum_{i=1}^{N}\left(x_{i}-\mu\right) \tag{A.3}
\end{equation*}
$$

and

$$
\begin{equation*}
-\frac{\partial \ell(\boldsymbol{\Psi})}{\partial \sigma}=\frac{N}{\sigma}-\frac{1}{\sigma^{3}} \sum_{i=1}^{N}\left(x_{i}-\mu\right)^{2}, \tag{A.4}
\end{equation*}
$$

leading to the MLE estimates

$$
\begin{equation*}
\hat{\mu}=\bar{x}=\frac{1}{N} \sum_{i=1}^{N} x_{i} \tag{A.5}
\end{equation*}
$$

and

$$
\begin{equation*}
\hat{\sigma}^{2}=S^{2}=\frac{1}{N} \sum_{i=1}^{N}\left(x_{i}-\hat{\mu}\right)^{2} \tag{A.6}
\end{equation*}
$$

## A. 2 Log-Normal Distribution

The MLE vector for the Log-Normal distribution is $\boldsymbol{\Psi}=(\mu, \sigma)$. We wish to compute the estimate of this vector, $\hat{\boldsymbol{\Psi}}=(\hat{\mu}, \hat{\sigma})$. The likelihood function $L(\boldsymbol{\Psi})$ is defined by

$$
\begin{equation*}
L(\boldsymbol{\Psi} ; \boldsymbol{x})=\prod_{i=1}^{N} \frac{1}{x \sigma \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{\log x-\mu}{\sigma}\right)^{2}} . \tag{A.7}
\end{equation*}
$$

The $\log$-likelihood function $\ell(\boldsymbol{\Psi} ; \boldsymbol{x})$ is given by

$$
\begin{equation*}
-\ell(\boldsymbol{\Psi} ; \boldsymbol{x})=N \log (\sigma \sqrt{2 \pi})+\sum_{i=1}^{N} \log x_{i}+\frac{1}{2 \sigma^{2}} \sum_{i=1}^{N}\left(\log x_{i}-\mu\right)^{2}, \tag{A.8}
\end{equation*}
$$

with partial derivatives (with respect to the parameters) given by

$$
\begin{equation*}
\frac{\partial \ell(\boldsymbol{\Psi})}{\partial \mu}=\frac{1}{\sigma^{2}} \sum_{i=1}^{N}\left(\log x_{i}-\mu\right) \tag{A.9}
\end{equation*}
$$

and

$$
\begin{equation*}
-\frac{\partial \ell(\boldsymbol{\Psi})}{\partial \sigma}=\frac{N}{\sigma}-\frac{1}{\sigma^{3}} \sum_{i=1}^{N}\left(\log x_{i}-\mu\right)^{2}, \tag{A.10}
\end{equation*}
$$

leading to the MLE estimates

$$
\begin{equation*}
\hat{\mu}=\bar{z}=\frac{1}{N} \sum_{i=1}^{N} z_{i} \tag{A.11}
\end{equation*}
$$

and

$$
\begin{equation*}
\hat{\sigma}^{2}=S_{\log x}^{2}=\frac{1}{N} \sum_{i=1}^{N}\left(z_{i}-\hat{\mu}\right)^{2}, \tag{A.12}
\end{equation*}
$$

where

$$
z_{i}=\log x_{i} .
$$

## A. 3 Gamma Distribution

In this section, we provide a method for computing the maximum likelihood estimators of $\alpha$ and $\beta$, the parameters of the two-parameter Gamma distribution specified by the probability density function

$$
f(t ; \alpha, \beta)=\frac{t^{\alpha-1} e^{-t / \beta}}{\Gamma(\alpha) \beta^{\alpha}}
$$

We will compute on the observation vector $\boldsymbol{x}$ three statistics $\bar{x}, s^{2}$, and $\overline{l(x)}$. We will define these statistics as

$$
\begin{align*}
\bar{x} & =\frac{1}{N} \sum_{i} x_{i}  \tag{A.13}\\
s^{2} & =\frac{1}{N-1} \sum_{i}\left(x_{i}-\bar{x}\right)^{2}  \tag{A.14}\\
\overline{l(x)} & =\frac{1}{N} \sum_{i} \log x_{i} . \tag{A.15}
\end{align*}
$$

Our goal is to maximize the likelihood function $L(\Psi)$ for the parameters $\Psi=$ $\{\alpha, \beta\}$. The likelihood function in this case is

$$
L(\mathbf{\Psi})=\prod_{i} \frac{x_{i}^{\alpha-1} e^{-x_{i} / \beta}}{\Gamma(\alpha) \beta^{\alpha}}=\left(\Gamma(\alpha) \beta^{\alpha}\right)^{-N} \cdot\left(\prod_{i} x_{i}\right)^{\alpha-1} \cdot \exp \left(-\frac{1}{\beta} \sum_{i} x_{i}\right)
$$

As it is generally simpler to maximize the log-likelihood function, we calculate the log-likelihood function

$$
\begin{aligned}
\ell(\boldsymbol{\Psi})=\log L(\boldsymbol{\Psi}) & =-N[\log \Gamma(\alpha)+\alpha \log \beta]+(\alpha-1) \sum_{i} \log x_{i}-\frac{1}{\beta} \sum_{i} x_{i} \\
& =-N[\log \Gamma(\alpha)+\alpha \log \beta]+N(\alpha-1) \overline{l(x)}-\frac{N \bar{x}}{\beta}
\end{aligned}
$$

In order to make the optimization simpler, we can instead minimize the negative log-likelihood (NLL) divided by $N$ (since $N$ is a constant). We are then interested in minimizing the function

$$
\begin{equation*}
Z=\log \Gamma(\alpha)+\alpha \log \beta-(\alpha-1) \overline{l(x)}+\frac{\bar{x}}{\beta} . \tag{A.16}
\end{equation*}
$$

The vector $\Psi$ will be optimal if

$$
\begin{equation*}
F(\Psi)=\binom{Z_{\alpha}(\Psi)}{Z_{\beta}(\boldsymbol{\Psi})}=\mathbf{0} \tag{A.17}
\end{equation*}
$$

where the notation $Z_{x}(\Psi)$ denotes the partial derivative of $Z$ with respect to the variable $x$ evaluated at $\boldsymbol{\Psi}$.

The partial derivatives of $Z$ are given by

$$
\begin{equation*}
Z_{\alpha}(\Psi)=\psi(\alpha)+\log \beta-\overline{l(x)} \tag{A.18}
\end{equation*}
$$

and

$$
\begin{equation*}
Z_{\beta}(\boldsymbol{\Psi})=\frac{1}{\beta^{2}}(\alpha \beta-\bar{x}) \tag{A.19}
\end{equation*}
$$

where $\psi(\alpha)$ is the derivative of $\log \Gamma(\alpha)$.
However, since we are solving the system of equations $F=\mathbf{0}$, we can multiply $Z_{\beta}$ by $\beta^{2}$ to eliminate the distributed fraction ( $\beta$ is defined to be nonzero by the Gamma distribution).

Using these changes, we are interested in solving the system

$$
\begin{equation*}
\tilde{F}(\boldsymbol{\Psi})=\binom{\psi(\alpha)+\log \beta-\overline{l(x)}}{\alpha \beta-\bar{x}}=\mathbf{0} \tag{A.20}
\end{equation*}
$$

Newton's method for systems states ([4]) that the system can be solved by starting with an initial guess $\boldsymbol{\Psi}^{(0)}$, and iterating using the equation

$$
\begin{equation*}
\boldsymbol{\Psi}^{(k+1)}=\boldsymbol{\Psi}^{(k)}+\boldsymbol{\Delta}^{(k)} \tag{A.21}
\end{equation*}
$$

where

$$
\begin{equation*}
\boldsymbol{\Delta}^{(k)}=-\boldsymbol{J}^{-1}\left(\boldsymbol{\Psi}^{(k)}\right) \tilde{F}\left(\boldsymbol{\Psi}^{(k)}\right) \tag{A.22}
\end{equation*}
$$

with $\boldsymbol{J}$ being the Jacobian of the system $(\tilde{F})$.
Let us write $\tilde{F}_{1}$ and $\tilde{F}_{2}$ for the first and second elements of $\tilde{F}$ as defined by Equation A.20. Given this notation, and using the partial derivative notation as earlier, we can define the Jacobian of the system as

$$
\boldsymbol{J}(\boldsymbol{\Psi})=\left(\begin{array}{cc}
\left(\tilde{F}_{1}\right)_{\alpha} & \left(\tilde{F}_{1}\right)_{\beta}  \tag{A.23}\\
\left(\tilde{F}_{2}\right)_{\alpha} & \left(\tilde{F}_{2}\right)_{\beta}
\end{array}\right)=\left(\begin{array}{cc}
\psi^{(1)}(\alpha) & \frac{1}{\beta} \\
\beta & \alpha
\end{array}\right)
$$

with $\psi^{(1)}(\alpha)$ the first derivative of $\psi(\alpha)$. The Jacobian has the inverse $\boldsymbol{J}^{-1}(\boldsymbol{\Psi})$ given by

$$
\boldsymbol{J}^{-1}(\boldsymbol{\Psi})=\frac{1}{\alpha \psi^{(1)}(\alpha)-1}\left(\begin{array}{cc}
\alpha & -\frac{1}{\beta}  \tag{A.24}\\
-\beta & \psi^{(1)}(\alpha)
\end{array}\right) .
$$

Writing $T_{1}, T_{2}$, and $T_{3}$ as

$$
\begin{align*}
& T_{1}=\psi(\alpha)+\log \beta-\overline{l(x)}  \tag{A.25}\\
& T_{2}=\alpha \beta-\bar{x}  \tag{A.26}\\
& T_{3}=\alpha \psi^{(1)}(\alpha)-1, \tag{A.27}
\end{align*}
$$

we can write

$$
\begin{equation*}
\boldsymbol{\Delta}=-\boldsymbol{J}^{-1} \tilde{F}=\frac{1}{T_{3}}\binom{-\alpha T_{1}+T_{2} / \beta}{\beta T_{1}-\psi^{(1)}(\alpha) T_{2}} \tag{A.28}
\end{equation*}
$$

so that our update equations for $\alpha^{(k+1)}$ and $\beta^{(k+1)}$ are given by

$$
\begin{align*}
& \alpha^{(k+1)}=\alpha^{(k)}+\frac{1}{T_{3}^{(k)}}\left(-\alpha^{(k)} T_{1}^{(k)}+T_{2}^{(k)} / \beta^{(k)}\right)  \tag{A.29}\\
& \beta^{(k+1)}=\beta^{(k)}+\frac{1}{T_{3}^{(k)}}\left(\beta^{(k)} T_{1}^{(k)}-\psi^{(1)}\left(\alpha^{(k)}\right) T_{2}^{(k)}\right) \tag{A.30}
\end{align*}
$$

where the $T_{j}$ are given by

$$
\begin{align*}
& T_{1}^{(k)}=\psi\left(\alpha^{(k)}\right)+\log \beta^{(k)}-\overline{l(x)}  \tag{A.31}\\
& T_{2}^{(k)}=\alpha^{(k)} \beta^{(k)}-\bar{x}  \tag{A.32}\\
& T_{3}^{(k)}=\alpha^{(k)} \psi^{(1)}\left(\alpha^{(k)}\right)-1 . \tag{A.33}
\end{align*}
$$

## A.3.1 The Process

Armed with the results of this section, we can numerically compute the MLE of the parameters ( $\alpha$ and $\beta$ ) by the following procedure.

1. Let $\beta^{(0)}=\frac{s^{2}}{\bar{x}}$
2. Let $\alpha^{(0)}=\frac{\beta^{(0)}}{\bar{x}}$
3. Let $k=0$
4. Compute $T_{1}^{(k)}, T_{2}^{(k)}$, and $T_{3}^{(k)}$ using Equations A.31, A. 32 and A. 33
5. Compute $\alpha^{(k+1)}$ and $\beta^{(k+1)}$ using Equations A. 29 and A. 30 .
6. Compute $\delta^{(k+1)}=\left(\left(\alpha^{(k+1)}-\alpha^{(k)}\right)^{2}+\left(\beta^{(k+1)}-\beta^{(k)}\right)^{2}\right)^{1 / 2}$
7. Increment $k$
8. If $\delta^{(k)}>\epsilon$ then goto step 4 , otherwise - output $\alpha^{(k)}$ and $\beta^{(k)}$ as answer.

## A. 4 Adjustments for the EM Algorithm

In order to compute the MLE estimates of the parameters for a mixture distribution, we use the EM algorithm. The EM algorithm puts the computation of the MLE estimates for the distribution into a "missing data" context. The missing data are indicators identifying which of the mixture components gave rise to each data point. During each iteration of the EM algorithm, we are compute an estimate of the $z_{i j}$ (the indicator variables). These estimates, denoted $\hat{z}_{i j}$ are then used along with the observations (the $x_{i}$ ) to estimate both the component proportions and the parameters of the component distributions. In this section, we present the adjustments that must be made in light of this new information, to allow the computation of the MLE estimates for the parameters of the mixture.

## A.4.1 The Mixture Proportions ( $\pi_{j}$ )

The mixture proportions can be directly estimated by computing the ratio of the number of observations that were expected to have come from component $j$ to the total number of observations. This estimate is given by

$$
\begin{equation*}
\hat{\pi}_{j}=\frac{1}{N} \sum_{i=1}^{N} \hat{z}_{i j} \tag{A.34}
\end{equation*}
$$

## A.4.2 The Component Parameters

The component parameters must be estimated differently for each component distribution. In order to estimate the component parameters, we must adjust the statistics used by the MLE estimators given by Equations A.5, A.6, A.11, A.12, A.13, A. 14 and A.15. These equations must be replaced by the similarly labeled equations

$$
\begin{equation*}
\bar{x}=\frac{\sum_{i=1}^{N} \hat{z}_{i j} x_{i}}{\sum_{i=1}^{N} \hat{z}_{i j}} \tag{*}
\end{equation*}
$$

and

$$
\begin{equation*}
S^{2}=\frac{\sum_{i=1}^{N} \hat{z}_{i j}\left(x_{i}-\bar{x}\right)^{2}}{\sum_{i=1}^{N} \hat{z}_{i j}} \tag{*}
\end{equation*}
$$

for the Normal and Gamma distributions, and

$$
\begin{equation*}
\bar{z}=\overline{l(x)}=\frac{\sum_{i=1}^{N} \hat{z}_{i j} \log x_{i}}{\sum_{i=1}^{N} \hat{z}_{i j}} \tag{*}
\end{equation*}
$$

for the Log-Normal and Gamma distributions, and

$$
\begin{equation*}
S_{\log X}^{2}=\frac{\sum_{i=1}^{N} \hat{z}_{i j}\left(\log x_{i}-\bar{z}\right)^{2}}{\sum_{i=1}^{N} \hat{z}_{i j}} \tag{*}
\end{equation*}
$$

for the Log-Normal distribution.

## Appendix B

## Training Data and Tabular Results

Table B.1: Example timing data for the example program

| Obs. | A | $\mathbf{B}$ | $\mathbf{C}$ | $\mathbf{D}$ | Total |
| ---: | ---: | ---: | ---: | ---: | ---: |
| 1 | 981.87 | 44.57 | 123.48 | 144.99 | 1294.91 |
| 2 | 42.22 | 55.27 | 81.29 | 154.91 | 333.69 |
| 3 | 40.58 | 49.82 | 96.25 | 155.65 | 342.3 |
| 4 | 37.63 | 1031.44 | 109.84 | 160.73 | 1339.64 |
| 5 | 45.34 | 44.69 | 79.14 | 147.14 | 316.31 |
| 6 | 38.38 | 48.78 | 77.55 | 155.62 | 320.33 |
| 7 | 1040.32 | 49.86 | 102.44 | 160.86 | 1353.48 |
| 8 | 999.95 | 38.42 | 84.23 | 136.52 | 1259.12 |
| 9 | 38.88 | 42.87 | 98.95 | 144.89 | 325.59 |
| 10 | 36.72 | 50.18 | 106.65 | 141.68 | 335.23 |
| 11 | 41.14 | 1078.53 | 105.73 | 145.86 | 1371.26 |
| 12 | 39.51 | 48.66 | 106.82 | 135 | 329.99 |
| 13 | 38.46 | 49.6 | 74.72 | 143.86 | 306.64 |
| 14 | 41.86 | 46.43 | 124.39 | 162.19 | 374.87 |
| 15 | 39.6 | 64.17 | 123.74 | 153.74 | 381.25 |
| 16 | 43.56 | 38.98 | 83.27 | 128.37 | 294.18 |
| 17 | 34.62 | 59.57 | 111.3 | 145.36 | 350.85 |
| 18 | 37.58 | 56.12 | 93.11 | 155.16 | 341.97 |
| 19 | 40.52 | 43.22 | 68.95 | 146.22 | 298.91 |
| 20 | 971.61 | 61.83 | 95.89 | 138.51 | 1267.84 |
| 21 | 40.41 | 52.37 | 117.12 | 141.33 | 351.23 |
| 22 | 40.95 | 65.61 | 81.78 | 134.89 | 323.23 |
| 23 | 994.83 | 34.28 | 118.03 | 136.3 | 1283.44 |
| 24 | 40.83 | 46.81 | 120.58 | 156.48 | 364.7 |
| 25 | 38.87 | 66.6 | 113.64 | 144.59 | 363.7 |
| 26 | 38.32 | 38.85 | 111.28 | 163.42 | 351.87 |

Table B.1: Example timing data for the example program

| Obs. | $\mathbf{A}$ | $\mathbf{B}$ | $\mathbf{C}$ | $\mathbf{D}$ | Total |
| ---: | ---: | ---: | ---: | ---: | ---: |
| 27 | 37.67 | 51.44 | 131.75 | 171.11 | 391.97 |
| 28 | 40.11 | 46.52 | 117.14 | 151.18 | 354.95 |
| 29 | 41.81 | 49.66 | 119.37 | 143.77 | 354.61 |
| 30 | 43.54 | 48.26 | 94.96 | 169.67 | 356.43 |
| 31 | 36.92 | 1049.41 | 97.22 | 131.35 | 1314.9 |
| 32 | 39.01 | 39.15 | 90.15 | 132.28 | 300.59 |
| 33 | 1007.41 | 41.17 | 120.35 | 165.2 | 1334.13 |
| 34 | 40.11 | 47.89 | 98.6 | 165.85 | 352.45 |
| 35 | 981.36 | 1074.21 | 138.83 | 150.78 | 2345.18 |
| 36 | 42.65 | 39.19 | 126.49 | 123.22 | 331.55 |
| 37 | 36.81 | 41.92 | 89.25 | 171.79 | 339.77 |
| 38 | 987.98 | 58.37 | 94.22 | 118.63 | 1259.2 |
| 39 | 39.45 | 39.37 | 100.23 | 191.85 | 370.9 |
| 40 | 39.37 | 49.42 | 99.82 | 157.02 | 345.63 |
| 41 | 39.92 | 45.98 | 95.07 | 127.77 | 308.74 |
| 42 | 1010.35 | 55.95 | 101.87 | 183.1 | 1351.27 |
| 43 | 42.58 | 57.11 | 118.75 | 148.33 | 366.77 |
| 44 | 39.05 | 58.32 | 101.4 | 147.24 | 346.01 |
| 45 | 36.18 | 41.16 | 102.13 | 158.75 | 338.22 |
| 46 | 34.21 | 49.13 | 177.42 | 174.79 | 435.55 |
| 47 | 39.29 | 55.52 | 94.46 | 143.45 | 332.72 |
| 48 | 39.07 | 51.26 | 63.79 | 132.53 | 286.65 |
| 49 | 42.24 | 42.94 | 82.96 | 156.57 | 324.71 |
| 50 | 40.41 | 69.36 | 67.53 | 150.92 | 328.22 |
| 51 | 39.39 | 46 | 115.22 | 150.76 | 351.37 |
| 52 | 989.21 | 62.72 | 96.97 | 153.47 | 1302.37 |
| 53 | 42.21 | 45.33 | 108.4 | 162.4 | 358.34 |
| 54 | 37.92 | 38.69 | 91.42 | 141.98 | 310.01 |
| 55 | 37.88 | 54.11 | 95.53 | 170.43 | 357.95 |
| 56 | 39.04 | 1102.68 | 83.56 | 148.11 | 1373.39 |
| 57 | 36.56 | 46.51 | 101.04 | 136.64 | 320.75 |
| 58 | 38.56 | 51.96 | 99.64 | 152.55 | 342.71 |
| 59 | 40.84 | 46.06 | 94.14 | 153.98 | 335.02 |
| 60 | 42.09 | 50.7 | 113.38 | 151.12 | 357.29 |
| 61 | 993.7 | 50.12 | 97.88 | 165.8 | 1307.5 |
| 62 | 38.83 | 42.6 | 108.14 | 185.25 | 374.82 |
| 63 | 37.63 | 56.96 | 116.86 | 138.88 | 350.33 |
| 64 | 1002.91 | 1002.17 | 99.88 | 144.39 | 2249.35 |
| 65 | 39.87 | 43.36 | 130.45 | 174.64 | 388.32 |
| 66 | 42.35 | 42.42 | 109.53 | 153.53 | 347.83 |
| 67 | 42.45 | 47.68 | 63.34 | 131.53 | 285 |
| 68 | 38.84 | 1041.85 | 99.9 | 150.28 | 1330.87 |
| 69 | 38.67 | 1054.31 | 88.41 | 152.6 | 1333.99 |
| 70 | 43.27 | 68.4 | 108.88 | 144.14 | 364.69 |
|  |  |  |  |  |  |
|  |  |  |  |  |  |

Table B.1: Example timing data for the example program

| Obs. | A | B | C | D | Total |
| ---: | ---: | ---: | ---: | ---: | ---: |
| 71 | 1044.11 | 41.98 | 93.36 | 130.72 | 1310.17 |
| 72 | 36.06 | 32.64 | 124.94 | 138.72 | 332.36 |
| 73 | 39.4 | 49.32 | 73.88 | 132.73 | 295.33 |
| 74 | 39.45 | 1111.46 | 106.74 | 142.23 | 1399.88 |
| 75 | 44.11 | 42.7 | 101.41 | 145.57 | 333.79 |
| 76 | 40.72 | 48.82 | 93.31 | 141.51 | 324.36 |
| 77 | 39.6 | 62.89 | 100.39 | 136.06 | 338.94 |
| 78 | 42.22 | 25.16 | 128.75 | 130.99 | 327.12 |
| 79 | 39.52 | 46.79 | 105.94 | 149.2 | 341.45 |
| 80 | 41.36 | 52.37 | 111.23 | 146.78 | 351.74 |
| 81 | 1014.24 | 30.97 | 102.51 | 151.26 | 1298.98 |
| 82 | 37.6 | 42.52 | 73.78 | 176.97 | 330.87 |
| 83 | 46.49 | 44.55 | 106.09 | 118.6 | 315.73 |
| 84 | 39.71 | 35.96 | 130.43 | 122.39 | 328.49 |
| 85 | 36.85 | 57.39 | 91.42 | 139.39 | 325.05 |
| 86 | 38.83 | 54.51 | 132.32 | 149.36 | 375.02 |
| 87 | 39.45 | 1034.04 | 114.3 | 139.44 | 1327.23 |
| 88 | 40.96 | 38.17 | 74.48 | 155.89 | 309.5 |
| 89 | 41.43 | 56.9 | 5.25 | 12.63 | 283.21 |
| 90 | 38.67 | 49.13 | 121.23 | 144.23 | 353.26 |
| 91 | 41.11 | 56.67 | 128.39 | 145.78 | 371.95 |
| 92 | 39.88 | 49.23 | 104.18 | 134.14 | 327.43 |
| 93 | 43.12 | 55.61 | 94.23 | 140.03 | 332.99 |
| 94 | 41.47 | 57.94 | 111.33 | 152.39 | 363.13 |
| 95 | 42.04 | 53.53 | 95.52 | 137.64 | 328.73 |
| 96 | 37.92 | 52.05 | 114.84 | 186.46 | 391.27 |
| 97 | 41.45 | 1048.8 | 143.85 | 149.79 | 1383.89 |
| 98 | 38.88 | 52.19 | 104.43 | 144.52 | 340.02 |
| 99 | 39.29 | 45.19 | 112.2 | 159.4 | 356.08 |
| 100 | 35.44 | 50.82 | 107.2 | 144.02 | 337.48 |
|  |  |  |  |  |  |

Table B.2: Unsmoothed Distribution for Part A

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{A}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{A}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{A}}(\mathbf{x})$ |
| ---: | ---: | ---: | ---: | ---: | ---: |
| 34 | 0.01 | 35 | 0.02 | 36 | 0.02 |
| 37 | 0.05 | 38 | 0.11 | 39 | 0.21 |
| 40 | 0.12 | 41 | 0.13 | 42 | 0.10 |
| 43 | 0.04 | 44 | 0.03 | 45 | 0.01 |
| 46 | 0.01 | 972 | 0.01 | 981 | 0.01 |
| 982 | 0.01 | 988 | 0.01 | 989 | 0.01 |
| 994 | 0.01 | 995 | 0.01 | 1000 | 0.01 |
| 1003 | 0.01 | 1007 | 0.01 | 1010 | 0.01 |
| 1014 | 0.01 | 1040 | 0.01 | 1044 | 0.01 |

Table B.3: Unsmoothed Distribution for Part B

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{B}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{B}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{B}}(\mathbf{x})$ |
| ---: | ---: | ---: | ---: | ---: | ---: |
| 25 | 0.01 | 31 | 0.01 | 33 | 0.01 |
| 34 | 0.01 | 36 | 0.01 | 38 | 0.02 |
| 39 | 0.06 | 41 | 0.02 | 42 | 0.03 |
| 43 | 0.07 | 45 | 0.05 | 46 | 0.04 |
| 47 | 0.04 | 48 | 0.03 | 49 | 0.08 |
| 50 | 0.06 | 51 | 0.04 | 52 | 0.05 |
| 54 | 0.02 | 55 | 0.02 | 56 | 0.04 |
| 57 | 0.05 | 58 | 0.03 | 60 | 0.01 |
| 62 | 0.01 | 63 | 0.02 | 64 | 0.01 |
| 66 | 0.01 | 67 | 0.01 | 68 | 0.01 |
| 69 | 0.01 | 1002 | 0.01 | 1031 | 0.01 |
| 1034 | 0.01 | 1042 | 0.01 | 1049 | 0.02 |
| 1054 | 0.01 | 1074 | 0.01 | 1079 | 0.01 |
| 1103 | 0.01 | 1111 | 0.01 |  |  |

Table B.4: Unsmoothed Distribution for Part C

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{C}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{C}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{C}}(\mathbf{x})$ |
| ---: | ---: | ---: | ---: | ---: | ---: |
| 59 | 0.01 | 63 | 0.01 | 64 | 0.01 |
| 68 | 0.01 | 69 | 0.01 | 74 | 0.03 |
| 75 | 0.01 | 78 | 0.01 | 79 | 0.01 |
| 81 | 0.01 | 82 | 0.01 | 83 | 0.02 |
| 84 | 0.02 | 88 | 0.01 | 89 | 0.01 |
| 90 | 0.01 | 91 | 0.02 | 93 | 0.03 |
| 94 | 0.04 | 95 | 0.02 | 96 | 0.04 |
| 97 | 0.02 | 98 | 0.01 | 99 | 0.02 |
| 100 | 0.06 | 101 | 0.03 | 102 | 0.03 |
| 103 | 0.01 | 104 | 0.02 | 106 | 0.03 |
| 107 | 0.04 | 108 | 0.02 | 109 | 0.01 |
| 110 | 0.02 | 111 | 0.04 | 112 | 0.01 |
| 113 | 0.01 | 114 | 0.02 | 115 | 0.02 |
| 117 | 0.03 | 118 | 0.01 | 119 | 0.02 |
| 120 | 0.01 | 121 | 0.02 | 123 | 0.01 |
| 124 | 0.02 | 125 | 0.01 | 126 | 0.01 |
| 128 | 0.01 | 129 | 0.01 | 130 | 0.02 |
| 132 | 0.02 | 139 | 0.01 | 144 | 0.01 |
| 177 | 0.01 |  |  |  |  |

Table B.5: Unsmoothed Distribution for Part D

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{D}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{D}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{D}}(\mathbf{x})$ |
| :---: | ---: | :---: | ---: | :---: | ---: |
| 119 | 0.02 | 122 | 0.01 | 123 | 0.01 |
| 126 | 0.01 | 128 | 0.02 | 131 | 0.03 |
| 132 | 0.02 | 133 | 0.02 | 134 | 0.01 |
| 135 | 0.02 | 136 | 0.02 | 137 | 0.02 |
| 138 | 0.01 | 139 | 0.05 | 140 | 0.01 |
| 141 | 0.01 | 142 | 0.04 | 143 | 0.01 |
| 144 | 0.06 | 145 | 0.05 | 146 | 0.04 |
| 147 | 0.03 | 148 | 0.02 | 149 | 0.02 |
| 150 | 0.02 | 151 | 0.06 | 152 | 0.01 |
| 153 | 0.03 | 154 | 0.03 | 155 | 0.02 |
| 156 | 0.04 | 157 | 0.02 | 159 | 0.02 |
| 161 | 0.02 | 162 | 0.02 | 163 | 0.01 |
| 165 | 0.01 | 166 | 0.02 | 170 | 0.02 |
| 171 | 0.01 | 172 | 0.01 | 175 | 0.02 |
| 177 | 0.01 | 183 | 0.01 | 185 | 0.01 |
| 186 | 0.01 | 192 | 0.01 |  |  |

Table B.6: Unsmoothed Distribution for Convolution

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | ---: | ---: | ---: | ---: | ---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 237 | 0.0002 | 238 | 0.0004 | 239 | 0.0004 |
| 240 | 0.0011 | 241 | 0.0027 | 242 | 0.0052 |
| 243 | 0.0041 | 244 | 0.0062 | 245 | 0.0096 |
| 246 | 0.0135 | 247 | 0.0158 | 248 | 0.0191 |
| 249 | 0.0243 | 250 | 0.0329 | 251 | 0.0439 |
| 252 | 0.0497 | 253 | 0.0633 | 254 | 0.0841 |
| 255 | 0.1141 | 256 | 0.1417 | 257 | 0.1641 |
| 258 | 0.2114 | 259 | 0.2718 | 260 | 0.3390 |
| 261 | 0.3747 | 262 | 0.4633 | 263 | 0.5646 |
| 264 | 0.6717 | 265 | 0.7960 | 266 | 0.9294 |
| 267 | 1.0989 | 268 | 1.2714 | 269 | 1.4821 |
| 270 | 1.6902 | 271 | 1.9703 | 272 | 2.2495 |
| 273 | 2.5639 | 274 | 2.9319 | 275 | 3.3193 |
| 276 | 3.7113 | 277 | 4.1960 | 278 | 4.7719 |
| 279 | 5.2993 | 280 | 5.9352 | 281 | 6.5611 |
| 282 | 7.3236 | 283 | 8.1553 | 284 | 8.9724 |
| 285 | 9.8698 | 286 | 10.7848 | 287 | 11.8535 |
| 288 | 12.8981 | 289 | 14.0605 | 290 | 15.4063 |
| 291 | 16.8229 | 292 | 18.1428 | 293 | 19.5799 |
| 294 | 21.1170 | 295 | 22.6542 | 296 | 24.4520 |
| 297 | 26.3358 | 298 | 28.1292 | 299 | 30.0043 |
| 300 | 32.1011 | 301 | 34.1585 | 302 | 36.3618 |
| 303 | 38.8238 | 304 | 40.9726 | 305 | 43.2978 |
| 306 | 45.8496 | 307 | 48.4962 | 308 | 51.1038 |
| 309 | 53.9138 | 310 | 56.8506 | 311 | 59.4863 |
| 312 | 62.3378 | 313 | 65.3592 | 314 | 68.4863 |
| 315 | 71.8616 | 316 | 75.2697 | 317 | 78.5651 |
| 318 | 81.5531 | 319 | 84.7625 | 320 | 88.0450 |
| 321 | 91.4246 | 322 | 94.7984 | 323 | 98.2855 |
| 324 | 101.4492 | 325 | 104.3412 | 326 | 107.5390 |
| 327 | 110.4036 | 328 | 113.2060 | 329 | 115.7463 |
| 330 | 117.9942 | 331 | 120.0053 | 332 | 121.8344 |
| 333 | 123.8301 | 334 | 125.5206 | 335 | 126.7851 |
| 336 | 127.3886 | 337 | 127.7312 | 338 | 127.8387 |
| 339 | 127.8949 | 340 | 128.0835 | 341 | 127.7878 |
| 342 | 126.8418 | 343 | 125.9688 | 344 | 124.5087 |
| 345 | 123.0813 | 346 | 121.6520 | 347 | 119.9535 |
| 348 | 117.8609 | 349 | 115.7310 | 350 | 113.5304 |
| 351 | 111.0966 | 352 | 108.8883 | 353 | 106.1318 |
| 354 | 103.5156 | 355 | 100.4236 | 356 | 97.6227 |
| 357 | 94.6687 | 358 | 91.8898 | 359 | 88.8445 |
| 360 | 85.6816 | 361 | 82.7584 | 362 | 79.4105 |
| 363 | 76.5277 | 364 | 73.4477 | 365 | 70.4693 |
|  |  |  |  |  |  |

Table B.6: Unsmoothed Distribution for Convolution

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | ---: | ---: | ---: | ---: | ---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 366 | 67.2713 | 367 | 64.4909 | 368 | 61.2848 |
| 369 | 58.3587 | 370 | 55.6777 | 371 | 52.9517 |
| 372 | 50.3328 | 373 | 47.7679 | 374 | 45.1724 |
| 375 | 42.4426 | 376 | 40.1944 | 377 | 37.7433 |
| 378 | 35.5680 | 379 | 33.4140 | 380 | 31.4748 |
| 381 | 29.4864 | 382 | 27.7476 | 383 | 25.8973 |
| 384 | 24.2039 | 385 | 22.5704 | 386 | 20.9388 |
| 387 | 19.6468 | 388 | 18.2725 | 389 | 17.0943 |
| 390 | 15.9051 | 391 | 14.9498 | 392 | 13.7596 |
| 393 | 12.8619 | 394 | 11.9711 | 395 | 11.0870 |
| 396 | 10.3833 | 397 | 9.6908 | 398 | 9.0543 |
| 399 | 8.4506 | 400 | 7.9508 | 401 | 7.4091 |
| 402 | 6.9829 | 403 | 6.5763 | 404 | 6.1956 |
| 405 | 5.8541 | 406 | 5.5203 | 407 | 5.1287 |
| 408 | 4.8430 | 409 | 4.6002 | 410 | 4.3841 |
| 411 | 4.1758 | 412 | 3.9745 | 413 | 3.6762 |
| 414 | 3.4186 | 415 | 3.2159 | 416 | 3.0191 |
| 417 | 2.8832 | 418 | 2.7374 | 419 | 2.5464 |
| 420 | 2.3935 | 421 | 2.2150 | 422 | 2.0191 |
| 423 | 1.8882 | 424 | 1.7743 | 425 | 1.6174 |
| 426 | 1.5065 | 427 | 1.4219 | 428 | 1.2836 |
| 429 | 1.2202 | 430 | 1.1090 | 431 | 1.0075 |
| 432 | 0.9365 | 433 | 0.8560 | 434 | 0.7985 |
| 435 | 0.7447 | 436 | 0.7017 | 437 | 0.6449 |
| 438 | 0.6138 | 439 | 0.5543 | 440 | 0.5220 |
| 441 | 0.4774 | 442 | 0.4431 | 443 | 0.4159 |
| 444 | 0.3866 | 445 | 0.3538 | 446 | 0.3215 |
| 447 | 0.3098 | 448 | 0.2913 | 449 | 0.2805 |
| 450 | 0.2704 | 451 | 0.2545 | 452 | 0.2166 |
| 453 | 0.2078 | 454 | 0.1842 | 455 | 0.1664 |
| 456 | 0.1604 | 457 | 0.1539 | 458 | 0.1431 |
| 459 | 0.1296 | 460 | 0.1104 | 461 | 0.0864 |
| 462 | 0.0738 | 463 | 0.0597 | 464 | 0.0571 |
| 465 | 0.0543 | 466 | 0.0459 | 467 | 0.0404 |
| 468 | 0.0357 | 469 | 0.0284 | 470 | 0.0258 |
| 471 | 0.0220 | 472 | 0.0162 | 473 | 0.0128 |
| 474 | 0.0106 | 475 | 0.0086 | 476 | 0.0075 |
| 477 | 0.0065 | 478 | 0.0043 | 479 | 0.0031 |
| 480 | 0.0018 | 481 | 0.0009 | 482 | 0.0005 |
| 483 | 0.0002 | 484 | 0.0001 | 1175 | 0.0002 |
| 1178 | 0.0001 | 1179 | 0.0003 | 1180 | 0.0002 |
| 1181 | 0.0002 | 1182 | 0.0002 | 1183 | 0.0004 |
| 1184 | 0.0010 | 1185 | 0.0007 | 1186 | 0.0006 |
|  |  |  |  |  |  |

Table B.6: Unsmoothed Distribution for Convolution

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 1187 | 0.0010 | 1188 | 0.0021 | 1189 | 0.0027 |
| 1190 | 0.0021 | 1191 | 0.0028 | 1192 | 0.0042 |
| 1193 | 0.0073 | 1194 | 0.0051 | 1195 | 0.0065 |
| 1196 | 0.0092 | 1197 | 0.0125 | 1198 | 0.0149 |
| 1199 | 0.0157 | 1200 | 0.0191 | 1201 | 0.0230 |
| 1202 | 0.0310 | 1203 | 0.0317 | 1204 | 0.0368 |
| 1205 | 0.0455 | 1206 | 0.0540 | 1207 | 0.0615 |
| 1208 | 0.0716 | 1209 | 0.0813 | 1210 | 0.0931 |
| 1211 | 0.1115 | 1212 | 0.1225 | 1213 | 0.1415 |
| 1214 | 0.1591 | 1215 | 0.1847 | 1216 | 0.2102 |
| 1217 | 0.2351 | 1218 | 0.2649 | 1219 | 0.2907 |
| 1220 | 0.3352 | 1221 | 0.3785 | 1222 | 0.4146 |
| 1223 | 0.4623 | 1224 | 0.5279 | 1225 | 0.5746 |
| 1226 | 0.6290 | 1227 | 0.7043 | 1228 | 0.7702 |
| 1229 | 0.8390 | 1230 | 0.9369 | 1231 | 1.0176 |
| 1232 | 1.0992 | 1233 | 1.2359 | 1234 | 1.3501 |
| 1235 | 1.4425 | 1236 | 1.5763 | 1237 | 1.7000 |
| 1238 | 1.8290 | 1239 | 2.0006 | 1240 | 2.1810 |
| 1241 | 2.3212 | 1242 | 2.4791 | 1243 | 2.6983 |
| 1244 | 2.8724 | 1245 | 3.0593 | 1246 | 3.3337 |
| 1247 | 3.5378 | 1248 | 3.7464 | 1249 | 4.0060 |
| 1250 | 4.2824 | 1251 | 4.5256 | 1252 | 4.8227 |
| 1253 | 5.1377 | 1254 | 5.3911 | 1255 | 5.6947 |
| 1256 | 6.0514 | 1257 | 6.3810 | 1258 | 6.7250 |
| 1259 | 7.1313 | 1260 | 7.4398 | 1261 | 7.7499 |
| 1262 | 8.1438 | 1263 | 8.5305 | 1264 | 8.9335 |
| 1265 | 9.3658 | 1266 | 9.8288 | 1267 | 10.1653 |
| 1268 | 10.5753 | 1269 | 10.9944 | 1270 | 11.4251 |
| 1271 | 11.9306 | 1272 | 12.3039 | 1273 | 12.7136 |
| 1274 | 13.1116 | 1275 | 13.5149 | 1276 | 14.0145 |
| 1277 | 14.4199 | 1278 | 14.8983 | 1279 | 15.2419 |
| 1280 | 15.5909 | 1281 | 15.9533 | 1282 | 16.3755 |
| 1283 | 16.8412 | 1284 | 17.1761 | 1285 | 17.5463 |
| 1286 | 17.7569 | 1287 | 18.0332 | 1288 | 18.3675 |
| 1289 | 18.6977 | 1290 | 18.9861 | 1291 | 19.2023 |
| 1292 | 19.3515 | 1293 | 19.4641 | 1294 | 19.7032 |
| 1295 | 19.9265 | 1296 | 20.1320 | 1297 | 20.1877 |
| 1298 | 20.3045 | 1299 | 20.3505 | 1300 | 20.4329 |
| 1301 | 20.5719 | 1302 | 20.6657 | 1303 | 20.6820 |
| 1304 | 20.6257 | 1305 | 20.6402 | 1306 | 20.6920 |
| 1307 | 20.7140 | 1308 | 20.7917 | 1309 | 20.7521 |
| 1310 | 20.6333 | 1311 | 20.6214 | 1312 | 20.5181 |
| 1313 | 20.5288 | 1314 | 20.4849 | 1315 | 20.4818 |
|  |  |  |  |  |  |

Table B.6: Unsmoothed Distribution for Convolution

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in $0.0001 ’ \mathbf{s})$ |  |  |  |  |  |
| 1316 | 20.4148 | 1317 | 20.2674 | 1318 | 20.2491 |
| 1319 | 20.0621 | 1320 | 20.0895 | 1321 | 19.9523 |
| 1322 | 19.8909 | 1323 | 19.7624 | 1324 | 19.5186 |
| 1325 | 19.4881 | 1326 | 19.4924 | 1327 | 19.3809 |
| 1328 | 19.2281 | 1329 | 19.0301 | 1330 | 18.7255 |
| 1331 | 18.6713 | 1332 | 18.5893 | 1333 | 18.4763 |
| 1334 | 18.2507 | 1335 | 18.0415 | 1336 | 17.7795 |
| 1337 | 17.5316 | 1338 | 17.4149 | 1339 | 17.1286 |
| 1340 | 16.9036 | 1341 | 16.5974 | 1342 | 16.3333 |
| 1343 | 16.0724 | 1344 | 15.8184 | 1345 | 15.5254 |
| 1346 | 15.3252 | 1347 | 14.9458 | 1348 | 14.6700 |
| 1349 | 14.3735 | 1350 | 14.0653 | 1351 | 13.8909 |
| 1352 | 13.6097 | 1353 | 13.3971 | 1354 | 13.0274 |
| 1355 | 12.8640 | 1356 | 12.6049 | 1357 | 12.3944 |
| 1358 | 12.1839 | 1359 | 11.9232 | 1360 | 11.5976 |
| 1361 | 11.2897 | 1362 | 11.1081 | 1363 | 10.8708 |
| 1364 | 10.6985 | 1365 | 10.4018 | 1366 | 10.1438 |
| 1367 | 9.8625 | 1368 | 9.6437 | 1369 | 9.4317 |
| 1370 | 9.1920 | 1371 | 8.9489 | 1372 | 8.7079 |
| 1373 | 8.4752 | 1374 | 8.2889 | 1375 | 8.0813 |
| 1376 | 7.8925 | 1377 | 7.6981 | 1378 | 7.5026 |
| 1379 | 7.3834 | 1380 | 7.1991 | 1381 | 7.0844 |
| 1382 | 6.9077 | 1383 | 6.7985 | 1384 | 6.6173 |
| 1385 | 6.4824 | 1386 | 6.3747 | 1387 | 6.2773 |
| 1388 | 6.2220 | 1389 | 6.1004 | 1390 | 5.9512 |
| 1391 | 5.7910 | 1392 | 5.6909 | 1393 | 5.5491 |
| 1394 | 5.4813 | 1395 | 5.3770 | 1396 | 5.2440 |
| 1397 | 5.1077 | 1398 | 4.9662 | 1399 | 4.8213 |
| 1400 | 4.6776 | 1401 | 4.5614 | 1402 | 4.4236 |
| 1403 | 4.3312 | 1404 | 4.1933 | 1405 | 4.0592 |
| 1406 | 3.9127 | 1407 | 3.7838 | 1408 | 3.6318 |
| 1409 | 3.5057 | 1410 | 3.3659 | 1411 | 3.2577 |
| 1412 | 3.1579 | 1413 | 3.0327 | 1414 | 2.8972 |
| 1415 | 2.7807 | 1416 | 2.6744 | 1417 | 2.5381 |
| 1418 | 2.4672 | 1419 | 2.3303 | 1420 | 2.2215 |
| 1421 | 2.1189 | 1422 | 2.0019 | 1423 | 1.8999 |
| 1424 | 1.7984 | 1425 | 1.7202 | 1426 | 1.6297 |
| 1427 | 1.5646 | 1428 | 1.4599 | 1429 | 1.3684 |
| 1430 | 1.2758 | 1431 | 1.2015 | 1432 | 1.1230 |
| 1433 | 1.0750 | 1434 | 1.0130 | 1435 | 0.9700 |
| 1436 | 0.9172 | 1437 | 0.8397 | 1438 | 0.7782 |
| 1439 | 0.7182 | 1440 | 0.6741 | 1441 | 0.6258 |
| 1442 | 0.5973 | 1443 | 0.5463 | 1444 | 0.5161 |
|  |  |  |  |  |  |
|  |  |  |  |  |  |

Table B.6: Unsmoothed Distribution for Convolution

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 1445 | 0.4833 | 1446 | 0.4397 | 1447 | 0.3996 |
| 1448 | 0.3727 | 1449 | 0.3470 | 1450 | 0.3222 |
| 1451 | 0.3061 | 1452 | 0.2785 | 1453 | 0.2548 |
| 1454 | 0.2300 | 1455 | 0.2181 | 1456 | 0.2049 |
| 1457 | 0.1930 | 1458 | 0.1788 | 1459 | 0.1690 |
| 1460 | 0.1560 | 1461 | 0.1424 | 1462 | 0.1304 |
| 1463 | 0.1300 | 1464 | 0.1278 | 1465 | 0.1261 |
| 1466 | 0.1235 | 1467 | 0.1067 | 1468 | 0.0972 |
| 1469 | 0.0917 | 1470 | 0.0865 | 1471 | 0.0862 |
| 1472 | 0.0871 | 1473 | 0.0815 | 1474 | 0.0770 |
| 1475 | 0.0701 | 1476 | 0.0608 | 1477 | 0.0556 |
| 1478 | 0.0557 | 1479 | 0.0461 | 1480 | 0.0506 |
| 1481 | 0.0478 | 1482 | 0.0412 | 1483 | 0.0392 |
| 1484 | 0.0327 | 1485 | 0.0283 | 1486 | 0.0246 |
| 1487 | 0.0204 | 1488 | 0.0208 | 1489 | 0.0211 |
| 1490 | 0.0175 | 1491 | 0.0161 | 1492 | 0.0151 |
| 1493 | 0.0142 | 1494 | 0.0124 | 1495 | 0.0101 |
| 1496 | 0.0109 | 1497 | 0.0107 | 1498 | 0.0091 |
| 1499 | 0.0086 | 1500 | 0.0073 | 1501 | 0.0075 |
| 1502 | 0.0096 | 1503 | 0.0073 | 1504 | 0.0098 |
| 1505 | 0.0078 | 1506 | 0.0054 | 1507 | 0.0047 |
| 1508 | 0.0031 | 1509 | 0.0033 | 1510 | 0.0044 |
| 1511 | 0.0052 | 1512 | 0.0058 | 1513 | 0.0056 |
| 1514 | 0.0040 | 1515 | 0.0032 | 1516 | 0.0020 |
| 1517 | 0.0014 | 1518 | 0.0016 | 1519 | 0.0023 |
| 1520 | 0.0013 | 1521 | 0.0013 | 1522 | 0.0010 |
| 1523 | 0.0004 | 1524 | 0.0003 | 1525 | 0.0001 |
| 1526 | 0.0001 | 2152 | 0.0002 | 2155 | 0.0001 |
| 2156 | 0.0003 | 2157 | 0.0002 | 2159 | 0.0002 |
| 2160 | 0.0002 | 2161 | 0.0007 | 2162 | 0.0004 |
| 2163 | 0.0001 | 2164 | 0.0006 | 2165 | 0.0010 |
| 2166 | 0.0010 | 2167 | 0.0009 | 2168 | 0.0012 |
| 2169 | 0.0014 | 2170 | 0.0018 | 2171 | 0.0020 |
| 2172 | 0.0018 | 2173 | 0.0019 | 2174 | 0.0031 |
| 2175 | 0.0031 | 2176 | 0.0037 | 2177 | 0.0045 |
| 2178 | 0.0041 | 2179 | 0.0053 | 2180 | 0.0059 |
| 2181 | 0.0068 | 2182 | 0.0061 | 2183 | 0.0075 |
| 2184 | 0.0100 | 2185 | 0.0090 | 2186 | 0.0108 |
| 2187 | 0.0131 | 2188 | 0.0131 | 2189 | 0.0143 |
| 2190 | 0.0161 | 2191 | 0.0162 | 2192 | 0.0187 |
| 2193 | 0.0233 | 2194 | 0.0224 | 2195 | 0.0215 |
| 2196 | 0.0275 | 2197 | 0.0302 | 2198 | 0.0309 |
| 2199 | 0.0365 | 2200 | 0.0351 | 2201 | 0.0370 |
|  |  |  |  |  |  |

Table B.6: Unsmoothed Distribution for Convolution

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 2202 | 0.0444 | 2203 | 0.0500 | 2204 | 0.0512 |
| 2205 | 0.0508 | 2206 | 0.0571 | 2207 | 0.0594 |
| 2208 | 0.0648 | 2209 | 0.0763 | 2210 | 0.0767 |
| 2211 | 0.0802 | 2212 | 0.0898 | 2213 | 0.0946 |
| 2214 | 0.0949 | 2215 | 0.1045 | 2216 | 0.1169 |
| 2217 | 0.1181 | 2218 | 0.1301 | 2219 | 0.1391 |
| 2220 | 0.1385 | 2221 | 0.1533 | 2222 | 0.1640 |
| 2223 | 0.1706 | 2224 | 0.1822 | 2225 | 0.1924 |
| 2226 | 0.1989 | 2227 | 0.2076 | 2228 | 0.2259 |
| 2229 | 0.2378 | 2230 | 0.2488 | 2231 | 0.2638 |
| 2232 | 0.2677 | 2233 | 0.2781 | 2234 | 0.2986 |
| 2235 | 0.3190 | 2236 | 0.3313 | 2237 | 0.3399 |
| 2238 | 0.3510 | 2239 | 0.3705 | 2240 | 0.3868 |
| 2241 | 0.4115 | 2242 | 0.4248 | 2243 | 0.4351 |
| 2244 | 0.4611 | 2245 | 0.4699 | 2246 | 0.4914 |
| 2247 | 0.5083 | 2248 | 0.5277 | 2249 | 0.5560 |
| 2250 | 0.5761 | 2251 | 0.6014 | 2252 | 0.6059 |
| 2253 | 0.6296 | 2254 | 0.6522 | 2255 | 0.6731 |
| 2256 | 0.7114 | 2257 | 0.7179 | 2258 | 0.7456 |
| 2259 | 0.7765 | 2260 | 0.7843 | 2261 | 0.8166 |
| 2262 | 0.8351 | 2263 | 0.8563 | 2264 | 0.8861 |
| 2265 | 0.9092 | 2266 | 0.9289 | 2267 | 0.9453 |
| 2268 | 0.9760 | 2269 | 1.0033 | 2270 | 1.0299 |
| 2271 | 1.0516 | 2272 | 1.0521 | 2273 | 1.0720 |
| 2274 | 1.1058 | 2275 | 1.1373 | 2276 | 1.1698 |
| 2277 | 1.1794 | 2278 | 1.1734 | 2279 | 1.1929 |
| 2280 | 1.2065 | 2281 | 1.2583 | 2282 | 1.2821 |
| 2283 | 1.2738 | 2284 | 1.2928 | 2285 | 1.2732 |
| 2286 | 1.3083 | 2287 | 1.3394 | 2288 | 1.3536 |
| 2289 | 1.3602 | 2290 | 1.3442 | 2291 | 1.3627 |
| 2292 | 1.3573 | 2293 | 1.3828 | 2294 | 1.3907 |
| 2295 | 1.3826 | 2296 | 1.3937 | 2297 | 1.3794 |
| 2298 | 1.3884 | 2299 | 1.3944 | 2300 | 1.3892 |
| 2301 | 1.3973 | 2302 | 1.3942 | 2303 | 1.3765 |
| 2304 | 1.3666 | 2305 | 1.3677 | 2306 | 1.3717 |
| 2307 | 1.3874 | 2308 | 1.3713 | 2309 | 1.3415 |
| 2310 | 1.3213 | 2311 | 1.3322 | 2312 | 1.3457 |
| 2313 | 1.3326 | 2314 | 1.3317 | 2315 | 1.2898 |
| 2316 | 1.2750 | 2317 | 1.2839 | 2318 | 1.2793 |
| 2319 | 1.2880 | 2320 | 1.2585 | 2321 | 1.2489 |
| 2322 | 1.2308 | 2323 | 1.2222 | 2324 | 1.2277 |
| 2325 | 1.2050 | 2326 | 1.2063 | 2327 | 1.1825 |
| 2328 | 1.1846 | 2329 | 1.1607 | 2330 | 1.1534 |
|  |  |  |  |  |  |
|  |  |  |  |  |  |

Table B.6: Unsmoothed Distribution for Convolution

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 2331 | 1.1564 | 2332 | 1.1382 | 2333 | 1.1312 |
| 2334 | 1.0970 | 2335 | 1.0923 | 2336 | 1.0788 |
| 2337 | 1.0842 | 2338 | 1.0736 | 2339 | 1.0560 |
| 2340 | 1.0281 | 2341 | 0.9998 | 2342 | 0.9981 |
| 2343 | 0.9887 | 2344 | 0.9936 | 2345 | 0.9701 |
| 2346 | 0.9374 | 2347 | 0.9202 | 2348 | 0.9089 |
| 2349 | 0.8994 | 2350 | 0.8854 | 2351 | 0.8722 |
| 2352 | 0.8543 | 2353 | 0.8277 | 2354 | 0.8210 |
| 2355 | 0.7984 | 2356 | 0.7939 | 2357 | 0.7717 |
| 2358 | 0.7615 | 2359 | 0.7461 | 2360 | 0.7226 |
| 2361 | 0.7091 | 2362 | 0.6873 | 2363 | 0.6844 |
| 2364 | 0.6588 | 2365 | 0.6466 | 2366 | 0.6179 |
| 2367 | 0.6054 | 2368 | 0.5952 | 2369 | 0.5836 |
| 2370 | 0.5637 | 2371 | 0.5359 | 2372 | 0.5198 |
| 2373 | 0.5024 | 2374 | 0.4966 | 2375 | 0.4758 |
| 2376 | 0.4643 | 2377 | 0.4451 | 2378 | 0.4276 |
| 2379 | 0.4121 | 2380 | 0.4035 | 2381 | 0.3840 |
| 2382 | 0.3769 | 2383 | 0.3589 | 2384 | 0.3536 |
| 2385 | 0.3342 | 2386 | 0.3238 | 2387 | 0.3116 |
| 2388 | 0.3030 | 2389 | 0.2975 | 2390 | 0.2791 |
| 2391 | 0.2745 | 2392 | 0.2586 | 2393 | 0.2561 |
| 2394 | 0.2421 | 2395 | 0.2366 | 2396 | 0.2247 |
| 2397 | 0.2170 | 2398 | 0.2073 | 2399 | 0.2029 |
| 2400 | 0.1957 | 2401 | 0.1833 | 2402 | 0.1781 |
| 2403 | 0.1667 | 2404 | 0.1651 | 2405 | 0.1534 |
| 2406 | 0.1542 | 2407 | 0.1408 | 2408 | 0.1407 |
| 2409 | 0.1283 | 2410 | 0.1254 | 2411 | 0.1167 |
| 2412 | 0.1129 | 2413 | 0.1103 | 2414 | 0.1052 |
| 2415 | 0.0978 | 2416 | 0.0915 | 2417 | 0.0922 |
| 2418 | 0.0816 | 2419 | 0.0824 | 2420 | 0.0756 |
| 2421 | 0.0762 | 2422 | 0.0668 | 2423 | 0.0668 |
| 2424 | 0.0603 | 2425 | 0.0580 | 2426 | 0.0555 |
| 2427 | 0.0529 | 2428 | 0.0514 | 2429 | 0.0452 |
| 2430 | 0.0444 | 2431 | 0.0395 | 2432 | 0.0415 |
| 2433 | 0.0354 | 2434 | 0.0357 | 2435 | 0.0306 |
| 2436 | 0.0324 | 2437 | 0.0286 | 2438 | 0.0275 |
| 2439 | 0.0245 | 2440 | 0.0234 | 2441 | 0.0239 |
| 2442 | 0.0203 | 2443 | 0.0190 | 2444 | 0.0167 |
| 2445 | 0.0173 | 2446 | 0.0159 | 2447 | 0.0159 |
| 2448 | 0.0116 | 2449 | 0.0127 | 2450 | 0.0113 |
| 2451 | 0.0108 | 2452 | 0.0096 | 2453 | 0.0094 |
| 2454 | 0.0083 | 2455 | 0.0072 | 2456 | 0.0074 |
| 2457 | 0.0067 | 2458 | 0.0065 | 2459 | 0.0046 |
|  |  |  |  |  |  |
|  |  |  |  |  |  |

Table B.6: Unsmoothed Distribution for Convolution

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$ |  |  |  |  | represented in $0.0001 ' s)$ |
| 2460 | 0.0056 | 2461 | 0.0051 | 2462 | 0.0049 |
| 2463 | 0.0034 | 2464 | 0.0040 | 2465 | 0.0041 |
| 2466 | 0.0040 | 2467 | 0.0032 | 2468 | 0.0033 |
| 2469 | 0.0028 | 2470 | 0.0029 | 2471 | 0.0030 |
| 2472 | 0.0024 | 2473 | 0.0025 | 2474 | 0.0022 |
| 2475 | 0.0022 | 2476 | 0.0020 | 2477 | 0.0022 |
| 2478 | 0.0013 | 2479 | 0.0021 | 2480 | 0.0013 |
| 2481 | 0.0014 | 2482 | 0.0011 | 2483 | 0.0016 |
| 2484 | 0.0007 | 2485 | 0.0011 | 2486 | 0.0010 |
| 2487 | 0.0009 | 2488 | 0.0006 | 2489 | 0.0005 |
| 2490 | 0.0007 | 2491 | 0.0005 | 2492 | 0.0002 |
| 2493 | 0.0003 | 2494 | 0.0007 | 2495 | 0.0004 |
| 2496 | 0.0001 | 2497 | 0.0002 | 2498 | 0.0004 |
| 2499 | 0.0003 | 2500 | 0.0001 | 2501 | 0.0001 |
| 2502 | 0.0002 | 2503 | 0.0004 | 2504 | 0.0001 |
| 2505 | 0.0002 | 2506 | 0.0001 | 2507 | 0.0003 |
| 2509 | 0.0002 | 2510 | 0.0001 | 2511 | 0.0001 |
| 2512 | 0.0001 | 2513 | 0.0001 | 2514 | 0.0001 |
| 2515 | 0.0001 | 2516 | 0.0001 | 2517 | 0.0001 |
| 2518 | 0.0001 | 2520 | 0.0001 | 2524 | 0.0001 |

Table B.7: Smoothed Distribution for Part A

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{A}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{A}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{A}}(\mathbf{x})$ |
| ---: | :--- | ---: | ---: | ---: | ---: |
| 32 | 0.0015 | 33 | 0.0050 | 34 | 0.0100 |
| 35 | 0.0195 | 36 | 0.0380 | 37 | 0.0755 |
| 38 | 0.1060 | 39 | 0.1360 | 40 | 0.1355 |
| 41 | 0.1205 | 42 | 0.0865 | 43 | 0.0590 |
| 44 | 0.0355 | 45 | 0.0170 | 46 | 0.0095 |
| 47 | 0.0035 | 48 | 0.0015 | 970 | 0.0015 |
| 971 | 0.0020 | 972 | 0.0030 | 973 | 0.0020 |
| 974 | 0.0015 | 979 | 0.0015 | 980 | 0.0035 |
| 981 | 0.0050 | 982 | 0.0050 | 983 | 0.0035 |
| 984 | 0.0015 | 986 | 0.0015 | 987 | 0.0035 |
| 988 | 0.0050 | 989 | 0.0050 | 990 | 0.0035 |
| 991 | 0.0015 | 992 | 0.0015 | 993 | 0.0035 |
| 994 | 0.0050 | 995 | 0.0050 | 996 | 0.0035 |
| 997 | 0.0015 | 998 | 0.0015 | 999 | 0.0020 |
| 1000 | 0.0030 | 1001 | 0.0035 | 1002 | 0.0035 |
| 1003 | 0.0030 | 1004 | 0.0020 | 1005 | 0.0030 |
| 1006 | 0.0020 | 1007 | 0.0030 | 1008 | 0.0035 |

Table B.7: Smoothed Distribution for Part A

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{A}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{A}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{A}}(\mathbf{x})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1009 | 0.0035 | 1010 | 0.0030 | 1011 | 0.0020 |
| 1012 | 0.0030 | 1013 | 0.0020 | 1014 | 0.0030 |
| 1015 | 0.0020 | 1016 | 0.0015 | 1038 | 0.0015 |
| 1039 | 0.0020 | 1040 | 0.0030 | 1041 | 0.0020 |
| 1042 | 0.0030 | 1043 | 0.0020 | 1044 | 0.0030 |
| 1045 | 0.0020 | 1046 | 0.0015 |  |  |

Table B.8: Smoothed Distribution for Part B

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{B}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{B}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{B}}(\mathbf{x})$ |
| ---: | ---: | ---: | ---: | ---: | ---: |
| 23 | 0.0015 | 24 | 0.0020 | 25 | 0.0030 |
| 26 | 0.0020 | 27 | 0.0015 | 29 | 0.0015 |
| 30 | 0.0020 | 31 | 0.0045 | 32 | 0.0055 |
| 33 | 0.0065 | 34 | 0.0065 | 35 | 0.0055 |
| 36 | 0.0075 | 37 | 0.0150 | 38 | 0.0195 |
| 39 | 0.0250 | 40 | 0.0235 | 41 | 0.0315 |
| 42 | 0.0270 | 43 | 0.0375 | 44 | 0.0345 |
| 45 | 0.0395 | 46 | 0.0345 | 47 | 0.0455 |
| 48 | 0.0480 | 49 | 0.0540 | 50 | 0.0540 |
| 51 | 0.0460 | 52 | 0.0350 | 53 | 0.0230 |
| 54 | 0.0235 | 55 | 0.0255 | 56 | 0.0335 |
| 57 | 0.0320 | 58 | 0.0265 | 59 | 0.0155 |
| 60 | 0.0090 | 61 | 0.0070 | 62 | 0.0100 |
| 63 | 0.0100 | 64 | 0.0100 | 65 | 0.0085 |
| 66 | 0.0080 | 67 | 0.0085 | 68 | 0.0085 |
| 69 | 0.0065 | 70 | 0.0035 | 71 | 0.0015 |
| 1000 | 0.0015 | 1001 | 0.0020 | 1002 | 0.0030 |
| 1003 | 0.0020 | 1004 | 0.0015 | 1029 | 0.0015 |
| 1030 | 0.0020 | 1031 | 0.0030 | 1032 | 0.0035 |
| 1033 | 0.0035 | 1034 | 0.0030 | 1035 | 0.0020 |
| 1036 | 0.0015 | 1040 | 0.0015 | 1041 | 0.0020 |
| 1042 | 0.0030 | 1043 | 0.0020 | 1044 | 0.0015 |
| 1047 | 0.0030 | 1048 | 0.0040 | 1049 | 0.0060 |
| 1050 | 0.0040 | 1051 | 0.0030 | 1052 | 0.0015 |
| 1053 | 0.0020 | 1054 | 0.0030 | 1055 | 0.0020 |
| 1056 | 0.0015 | 1072 | 0.0015 | 1073 | 0.0020 |
| 1074 | 0.0030 | 1075 | 0.0020 | 1076 | 0.0015 |
| 1077 | 0.0015 | 1078 | 0.0020 | 1079 | 0.0030 |
| 1080 | 0.0020 | 1081 | 0.0015 | 1101 | 0.0015 |
| 1102 | 0.0020 | 1103 | 0.0030 | 1104 | 0.0020 |
| 1105 | 0.0015 | 1109 | 0.0015 | 1110 | 0.0020 |
| 1111 | 0.0030 | 1112 | 0.0020 | 1113 | 0.0015 |
|  |  |  |  |  |  |

Table B.9: Smoothed Distribution for Part C

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{C}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{C}}(\mathbf{x})$ | x | $\hat{\mathbf{p}}_{\mathbf{C}}(\mathbf{x})$ |
| ---: | ---: | ---: | ---: | ---: | ---: |
| 57 | 0.0015 | 58 | 0.0020 | 59 | 0.0030 |
| 60 | 0.0020 | 61 | 0.0030 | 62 | 0.0035 |
| 63 | 0.0050 | 64 | 0.0050 | 65 | 0.0035 |
| 66 | 0.0030 | 67 | 0.0035 | 68 | 0.0050 |
| 69 | 0.0050 | 70 | 0.0035 | 71 | 0.0015 |
| 72 | 0.0045 | 73 | 0.0075 | 74 | 0.0110 |
| 75 | 0.0090 | 76 | 0.0080 | 77 | 0.0050 |
| 78 | 0.0050 | 79 | 0.0065 | 80 | 0.0070 |
| 81 | 0.0095 | 82 | 0.0120 | 83 | 0.0135 |
| 84 | 0.0115 | 85 | 0.0070 | 86 | 0.0045 |
| 87 | 0.0035 | 88 | 0.0065 | 89 | 0.0100 |
| 90 | 0.0105 | 91 | 0.0140 | 92 | 0.0175 |
| 93 | 0.0230 | 94 | 0.0280 | 95 | 0.0295 |
| 96 | 0.0275 | 97 | 0.0220 | 98 | 0.0260 |
| 99 | 0.0275 | 100 | 0.0340 | 101 | 0.0315 |
| 102 | 0.0290 | 103 | 0.0175 | 104 | 0.0170 |
| 105 | 0.0175 | 106 | 0.0230 | 107 | 0.0235 |
| 108 | 0.0235 | 109 | 0.0230 | 110 | 0.0205 |
| 111 | 0.0210 | 112 | 0.0190 | 113 | 0.0180 |
| 114 | 0.0135 | 115 | 0.0160 | 116 | 0.0145 |
| 117 | 0.0170 | 118 | 0.0145 | 119 | 0.0175 |
| 120 | 0.0125 | 121 | 0.0125 | 122 | 0.0105 |
| 123 | 0.0115 | 124 | 0.0115 | 125 | 0.0105 |
| 126 | 0.0095 | 127 | 0.0070 | 128 | 0.0095 |
| 129 | 0.0090 | 130 | 0.0125 | 131 | 0.0095 |
| 132 | 0.0090 | 133 | 0.0040 | 134 | 0.0030 |
| 137 | 0.0015 | 138 | 0.0020 | 139 | 0.0030 |
| 140 | 0.0020 | 141 | 0.0015 | 142 | 0.0015 |
| 143 | 0.0020 | 144 | 0.0030 | 145 | 0.0020 |
| 146 | 0.0015 | 175 | 0.0015 | 176 | 0.0020 |
| 177 | 0.0030 | 178 | 0.0020 | 179 | 0.0015 |
|  |  |  |  |  |  |

Table B.10: Smoothed Distribution for Part D

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{D}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{D}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{D}}(\mathbf{x})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 117 | 0.0030 | 118 | 0.0040 | 119 | 0.0060 |
| 120 | 0.0055 | 121 | 0.0065 | 122 | 0.0050 |
| 123 | 0.0050 | 124 | 0.0050 | 125 | 0.0035 |
| 126 | 0.0060 | 127 | 0.0060 | 128 | 0.0075 |
| 129 | 0.0085 | 130 | 0.0120 | 131 | 0.0160 |
| 132 | 0.0175 | 133 | 0.0195 | 134 | 0.0170 |
| 135 | 0.0180 | 136 | 0.0170 | 137 | 0.0225 |

Table B.10: Smoothed Distribution for Part D

| $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{D}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{D}}(\mathbf{x})$ | $\mathbf{x}$ | $\hat{\mathbf{p}}_{\mathbf{D}}(\mathbf{x})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 138 | 0.0215 | 139 | 0.0235 | 140 | 0.0225 |
| 141 | 0.0220 | 142 | 0.0265 | 143 | 0.0320 |
| 144 | 0.0420 | 145 | 0.0410 | 146 | 0.0400 |
| 147 | 0.0315 | 148 | 0.0250 | 149 | 0.0275 |
| 150 | 0.0265 | 151 | 0.0315 | 152 | 0.0285 |
| 153 | 0.0290 | 154 | 0.0265 | 155 | 0.0275 |
| 156 | 0.0245 | 157 | 0.0200 | 158 | 0.0140 |
| 159 | 0.0120 | 160 | 0.0110 | 161 | 0.0145 |
| 162 | 0.0120 | 163 | 0.0115 | 164 | 0.0100 |
| 165 | 0.0085 | 166 | 0.0080 | 167 | 0.0055 |
| 168 | 0.0060 | 169 | 0.0055 | 170 | 0.0095 |
| 171 | 0.0090 | 172 | 0.0080 | 173 | 0.0065 |
| 174 | 0.0055 | 175 | 0.0075 | 176 | 0.0060 |
| 177 | 0.0060 | 178 | 0.0020 | 179 | 0.0015 |
| 181 | 0.0015 | 182 | 0.0020 | 183 | 0.0045 |
| 184 | 0.0055 | 185 | 0.0065 | 186 | 0.0050 |
| 187 | 0.0035 | 188 | 0.0015 | 190 | 0.0015 |
| 191 | 0.0020 | 192 | 0.0030 | 193 | 0.0020 |
| 194 | 0.0015 |  |  |  |  |

Table B.11: Convolution After Smoothing

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| ---: | ---: | ---: | ---: | ---: | ---: |
| $(Z=A+B+C+D$, represented in 0.0001 s $)$ |  |  |  |  |  |
| 234 | 0.0001 | 235 | 0.0001 | 236 | 0.0003 |
| 237 | 0.0005 | 238 | 0.0009 | 239 | 0.0015 |
| 240 | 0.0023 | 241 | 0.0034 | 242 | 0.0047 |
| 243 | 0.0064 | 244 | 0.0086 | 245 | 0.0112 |
| 246 | 0.0145 | 247 | 0.0187 | 248 | 0.0238 |
| 249 | 0.0304 | 250 | 0.0387 | 251 | 0.0493 |
| 252 | 0.0627 | 253 | 0.0795 | 254 | 0.1006 |
| 255 | 0.1264 | 256 | 0.1579 | 257 | 0.1956 |
| 258 | 0.2404 | 259 | 0.2932 | 260 | 0.3550 |
| 261 | 0.4273 | 262 | 0.5114 | 263 | 0.6087 |
| 264 | 0.7209 | 265 | 0.8493 | 266 | 0.9954 |
| 267 | 1.1609 | 268 | 1.3478 | 269 | 1.5582 |
| 270 | 1.7942 | 271 | 2.0581 | 272 | 2.3521 |
| 273 | 2.6784 | 274 | 3.0396 | 275 | 3.4386 |
| 276 | 3.8781 | 277 | 4.3609 | 278 | 4.8897 |
| 279 | 5.4670 | 280 | 6.0951 | 281 | 6.7757 |
| 282 | 7.5106 | 283 | 8.3009 | 284 | 9.1484 |
| 285 | 10.0560 | 286 | 11.0274 | 287 | 12.0672 |
| 288 | 13.1790 | 289 | 14.3644 | 290 | 15.6229 |
| 291 | 16.9533 | 292 | 18.3550 | 293 | 19.8291 |
| 294 | 21.3783 | 295 | 23.0056 | 296 | 24.7129 |
| 297 | 26.5008 | 298 | 28.3693 | 299 | 30.3178 |
| 300 | 32.3457 | 301 | 34.4522 | 302 | 36.6361 |
| 303 | 38.8969 | 304 | 41.2348 | 305 | 43.6515 |
| 306 | 46.1479 | 307 | 48.7230 | 308 | 51.3740 |
| 309 | 54.0972 | 310 | 56.8913 | 311 | 59.7585 |
| 312 | 62.7028 | 313 | 65.7256 | 314 | 68.8219 |
| 315 | 71.9789 | 316 | 75.1793 | 317 | 78.4076 |
| 318 | 81.6537 | 319 | 84.9130 | 320 | 88.1817 |
| 321 | 91.4514 | 322 | 94.7061 | 323 | 97.9233 |
| 324 | 101.0769 | 325 | 104.1409 | 326 | 107.0909 |
| 327 | 109.9045 | 328 | 112.5624 | 329 | 115.0489 |
| 330 | 117.3515 | 331 | 119.4581 | 332 | 121.3529 |
| 333 | 123.0150 | 334 | 124.4212 | 335 | 125.5531 |
| 336 | 126.4035 | 337 | 126.9781 | 338 | 127.2892 |
| 339 | 127.3487 | 340 | 127.1615 | 341 | 126.7276 |
| 342 | 126.0460 | 343 | 125.1220 | 344 | 123.9677 |
| 345 | 122.6008 | 346 | 121.0391 | 347 | 119.2991 |
| 348 | 117.3947 | 349 | 115.3377 | 350 | 113.1388 |
| 351 | 110.8066 | 352 | 108.3498 | 353 | 105.7775 |
| 354 | 103.1020 | 355 | 100.3375 | 356 | 97.5004 |
| 357 | 94.6040 | 358 | 91.6598 | 359 | 88.6763 |
| 360 | 85.6617 | 361 | 82.6255 | 362 | 79.5775 |
|  |  |  |  |  |  |
|  |  |  |  |  |  |

Table B.11: Convolution After Smoothing

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | ---: | ---: | ---: | ---: |
| $(Z=A+B+C+D$, represented in 0.0001 s $)$ |  |  |  |  |  |
| 363 | 76.5281 | 364 | 73.4856 | 365 | 70.4585 |
| 366 | 67.4553 | 367 | 64.4869 | 368 | 61.5644 |
| 369 | 58.6984 | 370 | 55.8952 | 371 | 53.1564 |
| 372 | 50.4806 | 373 | 47.8661 | 374 | 45.3149 |
| 375 | 42.8327 | 376 | 40.4288 | 377 | 38.1121 |
| 378 | 35.8894 | 379 | 33.7623 | 380 | 31.7294 |
| 381 | 29.7859 | 382 | 27.9279 | 383 | 26.1527 |
| 384 | 24.4609 | 385 | 22.8553 | 386 | 21.3390 |
| 387 | 19.9138 | 388 | 18.5781 | 389 | 17.3277 |
| 390 | 16.1562 | 391 | 15.0582 | 392 | 14.0291 |
| 393 | 13.0669 | 394 | 12.1706 | 395 | 11.3395 |
| 396 | 10.5719 | 397 | 9.8650 | 398 | 9.2150 |
| 399 | 8.6178 | 400 | 8.0693 | 401 | 7.5652 |
| 402 | 7.1013 | 403 | 6.6729 | 404 | 6.2759 |
| 405 | 5.9068 | 406 | 5.5634 | 407 | 5.2446 |
| 408 | 4.9488 | 409 | 4.6738 | 410 | 4.4158 |
| 411 | 4.1705 | 412 | 3.9348 | 413 | 3.7073 |
| 414 | 3.4890 | 415 | 3.2814 | 416 | 3.0855 |
| 417 | 2.9005 | 418 | 2.7242 | 419 | 2.5542 |
| 420 | 2.3889 | 421 | 2.2282 | 422 | 2.0736 |
| 423 | 1.9264 | 424 | 1.7879 | 425 | 1.6586 |
| 426 | 1.5380 | 427 | 1.4252 | 428 | 1.3193 |
| 429 | 1.2198 | 430 | 1.1266 | 431 | 1.0398 |
| 432 | 0.9599 | 433 | 0.8872 | 434 | 0.8215 |
| 435 | 0.7621 | 436 | 0.7081 | 437 | 0.6583 |
| 438 | 0.6119 | 439 | 0.5685 | 440 | 0.5277 |
| 441 | 0.4895 | 442 | 0.4539 | 443 | 0.4211 |
| 444 | 0.3911 | 445 | 0.3639 | 446 | 0.3396 |
| 447 | 0.3179 | 448 | 0.2982 | 449 | 0.2797 |
| 450 | 0.2619 | 451 | 0.2441 | 452 | 0.2264 |
| 453 | 0.2092 | 454 | 0.1928 | 455 | 0.1776 |
| 456 | 0.1635 | 457 | 0.1499 | 458 | 0.1364 |
| 459 | 0.1226 | 460 | 0.1087 | 461 | 0.0951 |
| 462 | 0.0825 | 463 | 0.0714 | 464 | 0.0620 |
| 465 | 0.0541 | 466 | 0.0474 | 467 | 0.0415 |
| 468 | 0.0361 | 469 | 0.0311 | 470 | 0.0265 |
| 471 | 0.0223 | 472 | 0.0185 | 473 | 0.0152 |
| 474 | 0.0124 | 475 | 0.0100 | 476 | 0.0081 |
| 477 | 0.0064 | 478 | 0.0049 | 479 | 0.0037 |
| 480 | 0.0027 | 481 | 0.0018 | 482 | 0.0012 |
| 483 | 0.0007 | 484 | 0.0004 | 485 | 0.0002 |
| 486 | 0.0001 | 1175 | 0.0001 | 1176 | 0.0001 |
| 1177 | 0.0001 | 1178 | 0.0001 | 1179 | 0.0002 |
|  |  |  |  |  |  |
|  |  |  |  |  |  |

Table B.11: Convolution After Smoothing

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 1180 | 0.0002 | 1181 | 0.0003 | 1182 | 0.0004 |
| 1183 | 0.0005 | 1184 | 0.0007 | 1185 | 0.0009 |
| 1186 | 0.0012 | 1187 | 0.0015 | 1188 | 0.0019 |
| 1189 | 0.0025 | 1190 | 0.0031 | 1191 | 0.0038 |
| 1192 | 0.0047 | 1193 | 0.0058 | 1194 | 0.0071 |
| 1195 | 0.0086 | 1196 | 0.0104 | 1197 | 0.0126 |
| 1198 | 0.0152 | 1199 | 0.0181 | 1200 | 0.0215 |
| 1201 | 0.0255 | 1202 | 0.0301 | 1203 | 0.0353 |
| 1204 | 0.0414 | 1205 | 0.0483 | 1206 | 0.0562 |
| 1207 | 0.0651 | 1208 | 0.0753 | 1209 | 0.0867 |
| 1210 | 0.0995 | 1211 | 0.1139 | 1212 | 0.1302 |
| 1213 | 0.1484 | 1214 | 0.1686 | 1215 | 0.1912 |
| 1216 | 0.2162 | 1217 | 0.2438 | 1218 | 0.2744 |
| 1219 | 0.3081 | 1220 | 0.3453 | 1221 | 0.3862 |
| 1222 | 0.4309 | 1223 | 0.4795 | 1224 | 0.5322 |
| 1225 | 0.5890 | 1226 | 0.6505 | 1227 | 0.7169 |
| 1228 | 0.7887 | 1229 | 0.8666 | 1230 | 0.9509 |
| 1231 | 1.0419 | 1232 | 1.1398 | 1233 | 1.2443 |
| 1234 | 1.3555 | 1235 | 1.4736 | 1236 | 1.5990 |
| 1237 | 1.7325 | 1238 | 1.8743 | 1239 | 2.0247 |
| 1240 | 2.1838 | 1241 | 2.3517 | 1242 | 2.5286 |
| 1243 | 2.7151 | 1244 | 2.9114 | 1245 | 3.1179 |
| 1246 | 3.3345 | 1247 | 3.5614 | 1248 | 3.7986 |
| 1249 | 4.0462 | 1250 | 4.3044 | 1251 | 4.5731 |
| 1252 | 4.8525 | 1253 | 5.1426 | 1254 | 5.4437 |
| 1255 | 5.7559 | 1256 | 6.0788 | 1257 | 6.4112 |
| 1258 | 6.7519 | 1259 | 7.0999 | 1260 | 7.4553 |
| 1261 | 7.8193 | 1262 | 8.1931 | 1263 | 8.5773 |
| 1264 | 8.9712 | 1265 | 9.3729 | 1266 | 9.7805 |
| 1267 | 10.1925 | 1268 | 10.6082 | 1269 | 11.0268 |
| 1270 | 11.4478 | 1271 | 11.8700 | 1272 | 12.2926 |
| 1273 | 12.7153 | 1274 | 13.1380 | 1275 | 13.5601 |
| 1276 | 13.9802 | 1277 | 14.3960 | 1278 | 14.8057 |
| 1279 | 15.2080 | 1280 | 15.6030 | 1281 | 15.9907 |
| 1282 | 16.3696 | 1283 | 16.7372 | 1284 | 17.0905 |
| 1285 | 17.4276 | 1286 | 17.7479 | 1287 | 18.0519 |
| 1288 | 18.3397 | 1289 | 18.6103 | 1290 | 18.8622 |
| 1291 | 19.0946 | 1292 | 19.3083 | 1293 | 19.5050 |
| 1294 | 19.6861 | 1295 | 19.8517 | 1296 | 20.0012 |
| 1297 | 20.1336 | 1298 | 20.2493 | 1299 | 20.3492 |
| 1300 | 20.4344 | 1301 | 20.5056 | 1302 | 20.5631 |
| 1303 | 20.6080 | 1304 | 20.6416 | 1305 | 20.6658 |
| 1306 | 20.6813 | 1307 | 20.6876 | 1308 | 20.6832 |
|  |  |  |  |  |  |

Table B.11: Convolution After Smoothing

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 1309 | 20.6674 | 1310 | 20.6407 | 1311 | 20.6056 |
| 1312 | 20.5647 | 1313 | 20.5197 | 1314 | 20.4702 |
| 1315 | 20.4148 | 1316 | 20.3523 | 1317 | 20.2819 |
| 1318 | 20.2042 | 1319 | 20.1199 | 1320 | 20.0299 |
| 1321 | 19.9346 | 1322 | 19.8351 | 1323 | 19.7324 |
| 1324 | 19.6269 | 1325 | 19.5176 | 1326 | 19.4022 |
| 1327 | 19.2779 | 1328 | 19.1437 | 1329 | 19.0004 |
| 1330 | 18.8504 | 1331 | 18.6951 | 1332 | 18.5338 |
| 1333 | 18.3641 | 1334 | 18.1828 | 1335 | 17.9885 |
| 1336 | 17.7813 | 1337 | 17.5625 | 1338 | 17.3337 |
| 1339 | 17.0958 | 1340 | 16.8495 | 1341 | 16.5956 |
| 1342 | 16.3353 | 1343 | 16.0696 | 1344 | 15.7994 |
| 1345 | 15.5248 | 1346 | 15.2467 | 1347 | 14.9666 |
| 1348 | 14.6870 | 1349 | 14.4106 | 1350 | 14.1395 |
| 1351 | 13.8744 | 1352 | 13.6153 | 1353 | 13.3614 |
| 1354 | 13.1117 | 1355 | 12.8652 | 1356 | 12.6199 |
| 1357 | 12.3741 | 1358 | 12.1259 | 1359 | 11.8754 |
| 1360 | 11.6238 | 1361 | 11.3732 | 1362 | 11.1250 |
| 1363 | 10.8794 | 1364 | 10.6353 | 1365 | 10.3916 |
| 1366 | 10.1480 | 1367 | 9.9050 | 1368 | 9.6639 |
| 1369 | 9.4257 | 1370 | 9.1913 | 1371 | 8.9615 |
| 1372 | 8.7371 | 1373 | 8.5193 | 1374 | 8.3091 |
| 1375 | 8.1073 | 1376 | 7.9146 | 1377 | 7.7310 |
| 1378 | 7.5565 | 1379 | 7.3903 | 1380 | 7.2314 |
| 1381 | 7.0790 | 1382 | 6.9323 | 1383 | 6.7917 |
| 1384 | 6.6575 | 1385 | 6.5302 | 1386 | 6.4088 |
| 1387 | 6.2916 | 1388 | 6.1758 | 1389 | 6.0593 |
| 1390 | 5.9411 | 1391 | 5.8216 | 1392 | 5.7020 |
| 1393 | 5.5828 | 1394 | 5.4636 | 1395 | 5.3428 |
| 1396 | 5.2191 | 1397 | 5.0919 | 1398 | 4.9617 |
| 1399 | 4.8300 | 1400 | 4.6984 | 1401 | 4.5679 |
| 1402 | 4.4384 | 1403 | 4.3092 | 1404 | 4.1790 |
| 1405 | 4.0474 | 1406 | 3.9144 | 1407 | 3.7812 |
| 1408 | 3.6489 | 1409 | 3.5190 | 1410 | 3.3920 |
| 1411 | 3.2680 | 1412 | 3.1463 | 1413 | 3.0265 |
| 1414 | 2.9082 | 1415 | 2.7911 | 1416 | 2.6756 |
| 1417 | 2.5616 | 1418 | 2.4491 | 1419 | 2.3382 |
| 1420 | 2.2291 | 1421 | 2.1222 | 1422 | 2.0182 |
| 1423 | 1.9178 | 1424 | 1.8211 | 1425 | 1.7279 |
| 1426 | 1.6375 | 1427 | 1.5494 | 1428 | 1.4631 |
| 1429 | 1.3794 | 1430 | 1.2990 | 1431 | 1.2230 |
| 1432 | 1.1520 | 1433 | 1.0856 | 1434 | 1.0229 |
| 1435 | 0.9627 | 1436 | 0.9040 | 1437 | 0.8467 |
|  |  |  |  |  |  |

Table B.11: Convolution After Smoothing

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in $0.0001 ' s)$ |  |  |  |  |  |
| 1438 | 0.7910 | 1439 | 0.7378 | 1440 | 0.6877 |
| 1441 | 0.6411 | 1442 | 0.5976 | 1443 | 0.5569 |
| 1444 | 0.5183 | 1445 | 0.4815 | 1446 | 0.4466 |
| 1447 | 0.4137 | 1448 | 0.3831 | 1449 | 0.3548 |
| 1450 | 0.3287 | 1451 | 0.3044 | 1452 | 0.2819 |
| 1453 | 0.2611 | 1454 | 0.2419 | 1455 | 0.2246 |
| 1456 | 0.2088 | 1457 | 0.1946 | 1458 | 0.1815 |
| 1459 | 0.1696 | 1460 | 0.1588 | 1461 | 0.1492 |
| 1462 | 0.1410 | 1463 | 0.1338 | 1464 | 0.1273 |
| 1465 | 0.1212 | 1466 | 0.1150 | 1467 | 0.1087 |
| 1468 | 0.1026 | 1469 | 0.0969 | 1470 | 0.0918 |
| 1471 | 0.0872 | 1472 | 0.0829 | 1473 | 0.0786 |
| 1474 | 0.0740 | 1475 | 0.0691 | 1476 | 0.0643 |
| 1477 | 0.0597 | 1478 | 0.0555 | 1479 | 0.0517 |
| 1480 | 0.0481 | 1481 | 0.0446 | 1482 | 0.0411 |
| 1483 | 0.0374 | 1484 | 0.0337 | 1485 | 0.0301 |
| 1486 | 0.0269 | 1487 | 0.0241 | 1488 | 0.0218 |
| 1489 | 0.0199 | 1490 | 0.0182 | 1491 | 0.0167 |
| 1492 | 0.0153 | 1493 | 0.0140 | 1494 | 0.0129 |
| 1495 | 0.0118 | 1496 | 0.0109 | 1497 | 0.0102 |
| 1498 | 0.0096 | 1499 | 0.0091 | 1500 | 0.0087 |
| 1501 | 0.0084 | 1502 | 0.0081 | 1503 | 0.0078 |
| 1504 | 0.0073 | 1505 | 0.0068 | 1506 | 0.0061 |
| 1507 | 0.0055 | 1508 | 0.0050 | 1509 | 0.0047 |
| 1510 | 0.0046 | 1511 | 0.0045 | 1512 | 0.0044 |
| 1513 | 0.0042 | 1514 | 0.0038 | 1515 | 0.0033 |
| 1516 | 0.0028 | 1517 | 0.0024 | 1518 | 0.0020 |
| 1519 | 0.0017 | 1520 | 0.0014 | 1521 | 0.0012 |
| 1522 | 0.0009 | 1523 | 0.0007 | 1524 | 0.0005 |
| 1525 | 0.0003 | 1526 | 0.0002 | 1527 | 0.0001 |
| 1528 | 0.0001 | 2152 | 0.0001 | 2153 | 0.0001 |
| 2154 | 0.0001 | 2155 | 0.0001 | 2156 | 0.0001 |
| 2157 | 0.0002 | 2158 | 0.0002 | 2159 | 0.0003 |
| 2160 | 0.0003 | 2161 | 0.0004 | 2162 | 0.0005 |
| 2163 | 0.0006 | 2164 | 0.0007 | 2165 | 0.0008 |
| 2166 | 0.0009 | 2167 | 0.0011 | 2168 | 0.0013 |
| 2169 | 0.0015 | 2170 | 0.0017 | 2171 | 0.0019 |
| 2172 | 0.0022 | 2173 | 0.0025 | 2174 | 0.0029 |
| 2175 | 0.0033 | 2176 | 0.0037 | 2177 | 0.0042 |
| 2178 | 0.0047 | 2179 | 0.0053 | 2180 | 0.0059 |
| 2181 | 0.0066 | 2182 | 0.0073 | 2183 | 0.0082 |
| 2184 | 0.0091 | 2185 | 0.0101 | 2186 | 0.0112 |
| 2187 | 0.0123 | 2188 | 0.0135 | 2189 | 0.0148 |
|  |  |  |  |  |  |

Table B.11: Convolution After Smoothing

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 2190 | 0.0162 | 2191 | 0.0177 | 2192 | 0.0194 |
| 2193 | 0.0211 | 2194 | 0.0230 | 2195 | 0.0250 |
| 2196 | 0.0272 | 2197 | 0.0295 | 2198 | 0.0320 |
| 2199 | 0.0347 | 2200 | 0.0376 | 2201 | 0.0406 |
| 2202 | 0.0438 | 2203 | 0.0471 | 2204 | 0.0506 |
| 2205 | 0.0543 | 2206 | 0.0583 | 2207 | 0.0626 |
| 2208 | 0.0673 | 2209 | 0.0723 | 2210 | 0.0775 |
| 2211 | 0.0829 | 2212 | 0.0884 | 2213 | 0.0943 |
| 2214 | 0.1005 | 2215 | 0.1071 | 2216 | 0.1141 |
| 2217 | 0.1215 | 2218 | 0.1291 | 2219 | 0.1371 |
| 2220 | 0.1454 | 2221 | 0.1540 | 2222 | 0.1630 |
| 2223 | 0.1724 | 2224 | 0.1820 | 2225 | 0.1921 |
| 2226 | 0.2026 | 2227 | 0.2135 | 2228 | 0.2249 |
| 2229 | 0.2365 | 2230 | 0.2485 | 2231 | 0.2607 |
| 2232 | 0.2733 | 2233 | 0.2864 | 2234 | 0.2999 |
| 2235 | 0.3139 | 2236 | 0.3282 | 2237 | 0.3429 |
| 2238 | 0.3580 | 2239 | 0.3737 | 2240 | 0.3898 |
| 2241 | 0.4064 | 2242 | 0.4232 | 2243 | 0.4403 |
| 2244 | 0.4576 | 2245 | 0.4754 | 2246 | 0.4938 |
| 2247 | 0.5130 | 2248 | 0.5327 | 2249 | 0.5528 |
| 2250 | 0.5731 | 2251 | 0.5934 | 2252 | 0.6139 |
| 2253 | 0.6348 | 2254 | 0.6563 | 2255 | 0.6783 |
| 2256 | 0.7008 | 2257 | 0.7235 | 2258 | 0.7462 |
| 2259 | 0.7689 | 2260 | 0.7915 | 2261 | 0.8142 |
| 2262 | 0.8370 | 2263 | 0.8599 | 2264 | 0.8830 |
| 2265 | 0.9062 | 2266 | 0.9294 | 2267 | 0.9527 |
| 2268 | 0.9757 | 2269 | 0.9983 | 2270 | 1.0205 |
| 2271 | 1.0422 | 2272 | 1.0639 | 2273 | 1.0856 |
| 2274 | 1.1072 | 2275 | 1.1282 | 2276 | 1.1484 |
| 2277 | 1.1677 | 2278 | 1.1863 | 2279 | 1.2045 |
| 2280 | 1.2224 | 2281 | 1.2399 | 2282 | 1.2564 |
| 2283 | 1.2718 | 2284 | 1.2860 | 2285 | 1.2996 |
| 2286 | 1.3125 | 2287 | 1.3248 | 2288 | 1.3359 |
| 2289 | 1.3457 | 2290 | 1.3541 | 2291 | 1.3613 |
| 2292 | 1.3677 | 2293 | 1.3734 | 2294 | 1.3782 |
| 2295 | 1.3821 | 2296 | 1.3850 | 2297 | 1.3871 |
| 2298 | 1.3882 | 2299 | 1.3886 | 2300 | 1.3880 |
| 2301 | 1.3864 | 2302 | 1.3839 | 2303 | 1.3808 |
| 2304 | 1.3772 | 2305 | 1.3733 | 2306 | 1.3689 |
| 2307 | 1.3636 | 2308 | 1.3576 | 2309 | 1.3509 |
| 2310 | 1.3439 | 2311 | 1.3368 | 2312 | 1.3295 |
| 2313 | 1.3215 | 2314 | 1.3128 | 2315 | 1.3037 |
| 2316 | 1.2946 | 2317 | 1.2856 | 2318 | 1.2768 |
|  |  |  |  |  |  |

Table B.11: Convolution After Smoothing

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 2319 | 1.2680 | 2320 | 1.2586 | 2321 | 1.2489 |
| 2322 | 1.2388 | 2323 | 1.2286 | 2324 | 1.2186 |
| 2325 | 1.2086 | 2326 | 1.1986 | 2327 | 1.1884 |
| 2328 | 1.1781 | 2329 | 1.1677 | 2330 | 1.1569 |
| 2331 | 1.1457 | 2332 | 1.1340 | 2333 | 1.1220 |
| 2334 | 1.1099 | 2335 | 1.0978 | 2336 | 1.0857 |
| 2337 | 1.0732 | 2338 | 1.0601 | 2339 | 1.0462 |
| 2340 | 1.0319 | 2341 | 1.0174 | 2342 | 1.0032 |
| 2343 | 0.9890 | 2344 | 0.9746 | 2345 | 0.9597 |
| 2346 | 0.9443 | 2347 | 0.9287 | 2348 | 0.9131 |
| 2349 | 0.8976 | 2350 | 0.8823 | 2351 | 0.8668 |
| 2352 | 0.8512 | 2353 | 0.8355 | 2354 | 0.8198 |
| 2355 | 0.8042 | 2356 | 0.7888 | 2357 | 0.7733 |
| 2358 | 0.7577 | 2359 | 0.7418 | 2360 | 0.7256 |
| 2361 | 0.7092 | 2362 | 0.6927 | 2363 | 0.6761 |
| 2364 | 0.6594 | 2365 | 0.6427 | 2366 | 0.6260 |
| 2367 | 0.6093 | 2368 | 0.5926 | 2369 | 0.5758 |
| 2370 | 0.5588 | 2371 | 0.5418 | 2372 | 0.5249 |
| 2373 | 0.5085 | 2374 | 0.4924 | 2375 | 0.4767 |
| 2376 | 0.4613 | 2377 | 0.4460 | 2378 | 0.4310 |
| 2379 | 0.4163 | 2380 | 0.4022 | 2381 | 0.3885 |
| 2382 | 0.3753 | 2383 | 0.3624 | 2384 | 0.3500 |
| 2385 | 0.3379 | 2386 | 0.3262 | 2387 | 0.3149 |
| 2388 | 0.3040 | 2389 | 0.2934 | 2390 | 0.2831 |
| 2391 | 0.2731 | 2392 | 0.2633 | 2393 | 0.2538 |
| 2394 | 0.2445 | 2395 | 0.2354 | 2396 | 0.2266 |
| 2397 | 0.2181 | 2398 | 0.2097 | 2399 | 0.2016 |
| 2400 | 0.1936 | 2401 | 0.1858 | 2402 | 0.1782 |
| 2403 | 0.1708 | 2404 | 0.1638 | 2405 | 0.1570 |
| 2406 | 0.1504 | 2407 | 0.1439 | 2408 | 0.1376 |
| 2409 | 0.1314 | 2410 | 0.1254 | 2411 | 0.1197 |
| 2412 | 0.1143 | 2413 | 0.1090 | 2414 | 0.1040 |
| 2415 | 0.0991 | 2416 | 0.0943 | 2417 | 0.0898 |
| 2418 | 0.0854 | 2419 | 0.0813 | 2420 | 0.0772 |
| 2421 | 0.0733 | 2422 | 0.0695 | 2423 | 0.0659 |
| 2424 | 0.0623 | 2425 | 0.0590 | 2426 | 0.0558 |
| 2427 | 0.0528 | 2428 | 0.0499 | 2429 | 0.0471 |
| 2430 | 0.0444 | 2431 | 0.0419 | 2432 | 0.0394 |
| 2433 | 0.0371 | 2434 | 0.0350 | 2435 | 0.0329 |
| 2436 | 0.0310 | 2437 | 0.0291 | 2438 | 0.0274 |
| 2439 | 0.0256 | 2440 | 0.0240 | 2441 | 0.0224 |
| 2442 | 0.0209 | 2443 | 0.0195 | 2444 | 0.0182 |
| 2445 | 0.0169 | 2446 | 0.0158 | 2447 | 0.0146 |
|  |  |  |  |  |  |
|  |  |  |  |  |  |

Table B.11: Convolution After Smoothing

| $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ | $\mathbf{z}$ | $\hat{\mathbf{p}}_{\mathbf{Z}}(\mathbf{z})$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(Z=A+B+C+D$, represented in 0.0001 's $)$ |  |  |  |  |  |
| 2448 | 0.0136 | 2449 | 0.0126 | 2450 | 0.0116 |
| 2451 | 0.0108 | 2452 | 0.0100 | 2453 | 0.0092 |
| 2454 | 0.0085 | 2455 | 0.0079 | 2456 | 0.0073 |
| 2457 | 0.0067 | 2458 | 0.0062 | 2459 | 0.0058 |
| 2460 | 0.0053 | 2461 | 0.0050 | 2462 | 0.0046 |
| 2463 | 0.0044 | 2464 | 0.0041 | 2465 | 0.0039 |
| 2466 | 0.0037 | 2467 | 0.0035 | 2468 | 0.0033 |
| 2469 | 0.0031 | 2470 | 0.0029 | 2471 | 0.0028 |
| 2472 | 0.0026 | 2473 | 0.0025 | 2474 | 0.0023 |
| 2475 | 0.0022 | 2476 | 0.0020 | 2477 | 0.0019 |
| 2478 | 0.0018 | 2479 | 0.0017 | 2480 | 0.0015 |
| 2481 | 0.0014 | 2482 | 0.0013 | 2483 | 0.0012 |
| 2484 | 0.0011 | 2485 | 0.0010 | 2486 | 0.0009 |
| 2487 | 0.0008 | 2488 | 0.0007 | 2489 | 0.0006 |
| 2490 | 0.0006 | 2491 | 0.0005 | 2492 | 0.0005 |
| 2493 | 0.0004 | 2494 | 0.0004 | 2495 | 0.0003 |
| 2496 | 0.0003 | 2497 | 0.0003 | 2498 | 0.0003 |
| 2499 | 0.0002 | 2500 | 0.0002 | 2501 | 0.0002 |
| 2502 | 0.0002 | 2503 | 0.0002 | 2504 | 0.0002 |
| 2505 | 0.0002 | 2506 | 0.0002 | 2507 | 0.0002 |
| 2508 | 0.0001 | 2509 | 0.0001 | 2510 | 0.0001 |
| 2511 | 0.0001 | 2512 | 0.0001 | 2513 | 0.0001 |
| 2514 | 0.0001 | 2515 | 0.0001 | 2516 | 0.0001 |
| 2517 | 0.0001 | 2518 | 0.0001 | 2519 | 0.0001 |

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[^0]:    ${ }^{1}$ This is the same as saying that their collective union is the sample space (all possible outcomes) and their intersection is the empty set. Written in mathematical notation this is

    $$
    \begin{aligned}
    & \bigcup_{i=1}^{N} A_{i}=\Omega \\
    & \bigcap_{i=1}^{N} A_{i}=\emptyset
    \end{aligned}
    $$

[^1]:    ${ }^{\dagger}$ The Gamma MLE solve for this iteration diverged for distribution 2 , so $\hat{\alpha}_{2}$ and $\hat{\beta}_{2}$ are estimates using the Method of Moments.

[^2]:    ${ }^{1}$ The number of heads resulting from a fixed number $(M)$ of tosses of a fair coin is represented by the Binomial distribution with parameters $n=M$ and $\pi=0.5$.

[^3]:    ${ }^{1}$ Some of the formulas given in [7] have errors in them. The most important of which is Equation 3 , which should be

    $$
    \theta_{i}(x)=\frac{G(x)-G\left(\max _{i}\right)}{1-G\left(\max _{i}\right)}
    $$

    A more precise description of the problems of [7] would argue that their Theorem 1 is not valid. In fact, the WCET estimate should be $\max _{i}+\omega_{i}$ if $\omega_{i}$ is the value of their $\theta_{i}$ that gives $\theta_{i}\left(\omega_{i}\right)=1-\epsilon_{i}$.

