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H₂O CONTENTS IN OLIVINE-HOSTED MELT INCLUSIONS FROM PRIMITIVE MAGMAS IN THE NORTHERN CASCADE ARC

By

Steven D. Shaw

Accepted in Partial Completion Of the Requirements for the Degree Master of Science

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MASTER'S THESIS

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Steven D. Shaw July 2011

H₂O CONTENTS IN OLIVINE-HOSTED MELT INCLUSIONS FROM PRIMITIVE MAGMAS IN THE NORTHERN CASCADE ARC

A Thesis Presented to The Faculty of Western Washington University

In Partial Fulfillment Of the Requirements for the Degree Master of Science

> by Steven D. Shaw July 2011

Abstract

The subducting Juan de Fuca plate is the hot endmember of slabs worldwide, and its unique thermal character prompts debate about the role of fluid-flux melting versus decompression melting in the Cascade arc. While slow subduction of this hot slab is expected to result in strong dehydration prior to reaching sub-arc depths, there is no consensus on whether the slab is entirely dehydrated at this point, or whether volcanism is the result of water-poor, decompression melting, or fluid-flux melting. I provide the first measurements of pre-eruptive volatile contents in olivine-hosted melt inclusions from primitive magmas in the northern region of the arc, at Mount Baker and Glacier Peak. These volatile contents and melt inclusion compositions are used to model mantle melting processes.

Low-K olivine tholeiite (LKT) and calc-alkaline basalt (CAB) melt inclusions at Glacier Peak have minimum H₂O concentrations of 2.0 and 2.2 wt. % and fO2 of Δ QFM +1.1 and +1.5, respectively. The evolved compositions of these melt inclusions in both lava types (host olivine: Fo₈₅₋₈₉) are corrected to mantle values by addition of \leq 15 wt. % olivine, and the results suggest that the minimum water contents in the parental magmas are 1.7 wt. % and 2.0 wt. %. Measured values themselves may be low due to degassing at crustal depths. The Mount Baker Schreiber's Meadow cinder cone (CAB) has minimum H₂O concentrations of 2.3 wt. %, though these contents cannot be adequately disentangled from potential crustal involvement and/or magma mixing/mingling to be corrected to mantle values.

Results of modeling indicate that both LKT and CAB magmas at Glacier Peak are the result of 13-15% fluid-fluxed melting of a compositionally heterogeneous mantle source, last equilibrated near the base of the crust. Source regions are interpreted to contain both an ocean island basalt (OIB)-like component and a mid-ocean ridge basalt (MORB)-like component. Minimum H₂O contents suggested in the source region are between 0.21 and 0.28 wt %. This is in contrast to southern regions of the Cascade arc, where LKT magmas are considered to be the result of dry decompression melting of a MORB-like source.

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Introduction

Fluid-flux melting is commonly accepted as the primary driver of arc magmatism worldwide (Anderson, 1974; Fyfe and McBirney, 1975; McBirney, 1969; Sisson and Grove, 1993); however, decompression melting does play a part in some arc magmatism (Cervantes and Wallace, 2003; Grove et al., 2002; Johnson et al., 2009; Sisson and Bronto, 1998; Sisson and Layne, 1993). The Cascade volcanic arc (Figure 1) has been described as a "hot" endmember arc worldwide in terms of subducting plate temperature (Green and Harry, 1999; Leeman et al., 2005a) and its unique thermal structure has prompted debate about the relative roles of fluid-flux (wet) melting versus decompression (dry) melting in magma generation. Resolving this question has centered on characterization of volatile contents and trace element geochemistry of primitive, mafic magmas, which frequently occur in monogenetic cinder cones throughout the arc.

The Cascade arc displays diversity in the geochemical composition of these most primitive magmas. Three major compositional types have been described: calc-alkaline basalts and basaltic-andesites (CAB), low potassium olivine tholeiites (LKT), and ocean island basalt (OIB)-like basalts (Bacon, 1997; Borg et al., 1997; Conrey et al., 1997; Green and Harry, 1999; Leeman et al., 2005a; Reiners, 2000). These compositional types have been attributed to generation by distinct sources and processes (e.g. Bacon, 1997; Leeman et al., 2005a; Leeman et al., 1990).

Thorough characterization and interpretation of the origin of Cascade arc LKTs and CABs has been an ongoing process (Bacon, 1997; Borg et al., 1997; Conrey et al., 1997; Green and Harry, 1999; Green and Sinha, 2005; Grove et al., 2002; Le Voyer et al., 2010; Leeman et al., 2005a; Leeman et al., 1990; Leeman et al., 2005b; Moore and DeBari, submitted; Reiners et al., 2000; Rowe et al., 2009; Ruscitto et al., 2010b; Schmidt et al., 2008; Sisson and Layne, 1993; Streck et al., 2007). LKTs have low potassium (≤ 0.6 wt. % K₂O), high aluminum (≥ 16 wt. % Al₂O₃) and are distinct; their geochemical signatures resemble those of mid-ocean ridge basalts (MORB) (Bacon, 1997; Leeman et al., 2005a). Like MORB, they have low abundances of incompatible elements; however they often have a small enrichment in large-ion lithophile elements (LILE) such as Ba, Sr and Pb (Bacon, 1997; Leeman et al., 2005a; Schmidt et al., 2008). The CAB and basaltic andesites have variable enrichments of LILEs, which is commonly known as an arc signature (Bacon, 1997; Leeman et al., 2005a). The LKTs have been proposed to be a result of decompression (dry) melting, while the origin of the CABs is less clear.

A model by Grove et al. (2002) explains generation of distinct CAB and LKT compositions at Mt. Shasta, in the southern Cascade arc, as melts of variably fluxed subarc mantle. To discern source characteristics, they use water content estimates, trace element and isotopic abundances, and experimental pressure and temperature constraints. They propose that while mantle convection results in shallow decompression melting (represented by LKTs); calc-alkaline basalts are caused by fluid-flux melting. Mantle material carried up with convection currents towards the wedge corner would melt due to decompression, depleting the mantle. This depleted source would descend along the slab/wedge interface, gaining fluids released from the subducting slab, which would cause fluid-flux melting (represented by the CABs) (Figure 2) (Grove et al., 2002). In this model, H_2O contents of calc-alkaline primitive magmas are expected to overlap with typical arc H_2O concentrations, between 1-6 wt. % (Wallace, 2005), whereas low potassium tholeiites are expected to have low H_2O , less than 1 wt. % (Sisson and Bronto, 1998; Sisson and Layne, 1993; Wallace, 2005).

In contrast to the Grove et al. model, Leeman et al. (2005a) propose that both LKT and CAB in the southern Washington Cascades are produced by melting at different depths and temperatures in a vertical profile beneath the arc. Decompression melting of the mantle wedge produces LKT lavas. CAB lavas are then formed when rising decompression melts heat mantle components that retain a trace element arc-signature (enrichment of LILE trace elements that fractionate in fluid phases) from previous episodes of arc magmatism (Leeman et al., 2005a). These "fossilized" slab components can explain the slight LILE enrichments seen in LKTs within the Cascade arc and LILE enrichments seen far east of the arc axis (Hart et al., 1984). Further, they predict that, because the subducting slab becomes younger in the northern portion of the Cascade arc (due to oblique subduction—Figure 1) any fluid-flux component would become negligible; as the slab becomes hotter, slab dehydration would occur closer to the plate margin, eliminating the fluid component (Leeman et al., 2005a). While Mt. Shasta CAB lavas have been shown to have a fluid-flux component (Le Voyer et al., 2010; Ruscitto et al., 2010a; Sisson and Layne, 1993), Leeman et al. (2005a) suggest that this may be an anomalous feature of that region that is unrepeated elsewhere.

The best test of the relative roles of fluid-flux and decompression melting is to directly measure pre-eruptive volatile contents in primitive magmas. In the southern

Cascades (at Mount Shasta and Medicine Lake), melt inclusion studies support the low $(<1 \text{ wt\% } \text{H}_2\text{O})$ nearly anhydrous nature of LKT magmas, and the wet (up to 5 wt % $\text{H}_2\text{O})$ nature of the CAB magmas (Anderson, 1974; Le Voyer et al., 2010; Sisson and Layne, 1993). In the central arc, recent work by Ruscitto et al. (2010b) has shown that CAB in the Oregon Cascades have initial water contents that overlap with arcs worldwide (1.7 to 3.6 wt.%). Their interpretation is that these water contents (and their correlation to proxies for water) indicate a small amount of slab-derived fluid that is added to the mantle wedge.

Volatile characterization has yet to be extended to the northernmost segment of the Cascade arc, the Garibaldi volcanic belt (GVB) (Figure 1). No direct measurements of volatile contents in primitive magmas in the GVB have been made. Relatively primitive mafic magmas are rare, but do exist in this segment, and share the compositional diversity of primitive magmas shown in southern regions (Moore and DeBari, submitted; DeBari et al., submitted).

Characterization of primitive melt volatile contents is commonly undertaken through analyses of melt inclusions in early-formed minerals such as olivine in basaltic tephra. As olivine crystals grow, they entrap small parcels of melt. These melt inclusions are preserved as glass within their host crystals, and volatiles are dissolved within that glass. The advantage of analyzing melt inclusions is that they represent the most primitive melt composition that can be sampled directly. Once trapped, they are not affected by further magmatic processes common in the crust, such as fractionation of

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multiple minerals, assimilation or mixing (Danyushevsky et al., 2002; Kent, 2008; Sobolev and Chaussidon, 1996).

In this paper, I present measured magmatic volatile concentrations from Mt. Baker and Glacier Peak in the GVB and add to the growing dataset of primitive magma compositions along the Cascades arc. These data are crucial to assessing melt generation models as they pertain to the northern arc.

Geologic Setting

The Cascade volcanic arc results from subduction of the Juan de Fuca, Explorer, and Gorda lithospheric plates beneath the western edge of North America (Figure 1). The arc is approximately 1200 km in length and was initially split into five segments based on volcanic center density and output (Guffanti and Weaver, 1988). More recently, Schmidt et al. (2008) redefined the segments based on isotopic variability, splitting the arc into four segments, and redefining the southern Washington Cascade segment based on its geochemically distinct basement rock, the Siletz terrane.

In the southern and central segments, where volcanic output is relatively high, volcanic vents are densely spaced and eruptive activity extends both into the forearc and backarc regions. This has been interpreted as the result of regional extension (Hughes and Taylor, 1986). In the northern segment, the Garibaldi Volcanic Belt (GVB), volcanic output is low, vents are sparsely separated, and there is no eruptive activity noted in the forearc and backarc regions. The GVB is ~330 km long, beginning just south of Glacier Peak and ending at Mt. Meager in British Columbia. The crust in this region is composed of metamorphosed oceanic terranes (Brown, 1987) between 40-45 km thick and the depth to the slab/wedge interface is generally considered to be between 80-100 km (McCrory et al., 2004; Mooney and Weaver, 1989; Parsons et al., 1998). The age of the slab beneath Glacier Peak and Mount Baker is estimated at < 22 Ma (Green and Harry, 1999; Wilson, 1988).

Glacier Peak, the southernmost GVB volcanic center, is a dacitic stratovolcano that has been active for the past 700 ka (Tabor and Crowder, 1969). Low-K olivine

tholeiite (potential "decompression type") and calc-alkaline basalt (potential "fluid-flux type") tephras have erupted near Glacier Peak at the Whitechuck and Indian Pass cinder cones, respectively (DeBari et al., submitted). Taylor (2001) described the petrography and geochemistry of four mafic lavas in the Glacier Peak area, including the Indian Pass (CAB) and Whitechuck (LKT) cinder cones used in this study. Whitechuck tephra are generally aphyric, with \leq 5% phenocrystic assemblage dominated by olivine with minor plagioclase. Clinopyroxene does not occur as phenocrysts, but is found in the groundmass. Indian Pass tephras have a slightly higher percentage of phenocrysts $(\leq 10\%)$, again dominated by olivine (~7%), with minor plagioclase and no clinopyroxene, other than in the groundmass (Taylor, 2001). Whereas Indian Pass lavas meet the major and trace element standards for a primitive magma established by the Basaltic Volcanism Study Project (1981) (MgO > 6 wt. % and Mg# >55; Ni>100 ppm and Cr >200 ppm), Whitechuck LKT lavas have MgO > 6 wt.%, low Ni (between 55-75) ppm) and low Cr (150-175 ppm) indicative of fractionation of olivine and probably chromite.

Based on whole rock analyses and mineral chemistry, Taylor (2001) establishes that these mafic cones are likely derived from two distinct mantle sources, a depleted mantle, similar to mid-ocean ridge basaltic sources (MORB), for the Whitechuck LKT, and a more enriched source for the CABs (including Indian Pass). All lavas have the characteristic high LILE/HFSE indicative of a subduction input (either fossilized or contemporaneous). Whitechuck LKTs are interpreted as the result of high degrees of anhydrous melting of a depleted source. High LILE/HFSE ratios are explained as the sign of a fossilized subduction component. The Indian Pass CABs are interpreted to result from lower degrees of fluid-fluxed melting of a less depleted source.

Mount Baker, the northern neighbor of Glacier Peak by ~90 km, is an andesitic stratovolcano. Calc-alkaline basalt (potential "fluid flux type") tephras and flows are present to the south of Mount Baker, at the monogenetic Schreiber's Meadow cinder cone and at the Sulphur Creek and Lake Shannon lava flows. An LKT-like composition (potential "decompression melt type") is found in the Park Butte flow (Moore and DeBari, submitted).

Moore and DeBari (submitted) distinguish three distinct primitive end-member magma types present at Mount Baker, an LKT-like basalt, a CAB endmember, and a high-Mg basaltic-andesite. One of the lava flow units, the CAB Sulphur Creek flow, is associated with the Schreiber's Meadow cinder cone used in this study. The Sulphur Creek lava flow is porphyritic, and contains $\leq 25\%$ modal phenocrysts, mostly plagioclase ($\leq 20\%$), with some olivine ($\leq 5\%$) and minor ($\leq 1\%$) clinopyroxene (Moore and DeBari, submitted). All mafic lavas studied have significant textural evidence for disequilibrium, as well as chemical disequilibrium between whole rock and mineral phases, suggesting an unidentified crustal modification of the magma or magma mixing. However, the lack of incompatible element variation and primitive isotopic signatures suggest minimal addition of crustal material. Thus Moore and DeBari (submitted) conclude that the incompatible trace element variations in these lavas are useful for characterizing mantle sources. All lavas include the high LILE/HFSE ratios typical of arcs, and LKT-like lavas have some of the highest LILE/HFSE ratios, which is unlike other Cascade LKTs. Through melt modeling, Moore and DeBari (submitted) establish that the three endmember types are the result of two distinct mantle sources and variable melt degree, degree of differentiation and/or enrichment in slab contribution. The LKT-like basalt is interpreted to derive from a depleted spinel lherzolite source and displays the strongest enrichment in slab-derived fluids. The CAB flows (including Sulphur Creek) are interpreted to derive from a less-depleted spinel lherzolite source with the smallest enrichment in slab-derived fluids. The high-Mg basaltic-andesite end-member is interpreted to originate from a depleted garnet lherzolite source that has an intermediate enrichment in slab-derived fluids.

Sampling and Analytical Methods

Detailed sampling and analytical methods are presented in Appendix A. A summary of these methods is provided below.

Tephra samples for this study were collected from three monogenetic cinder cones—two from Glacier Peak (a primitive calc-alkaline basalt [CAB] and a primitive low-potassium olivine tholeiite [LKT]) and one from Mount Baker, a primitive CAB). LKT lavas are present at Mt. Baker, but none have associated tephra and were not sampled for this study. Olivine crystals from tephras in this study were hand-picked, temporarily mounted individually on round glass slides, and doubly polished to expose melt inclusions on two parallel sides of the host olivine. Melt inclusions were glassy, between 50 and 150 microns in diameter, generally round or oval-shaped and contained vapor bubbles (Figure 3). After polishing, host olivine crystals were removed from the glass slide for analysis.

Fourier Transform Infrared Spectroscopy

Analysis of H_2O and CO_2 in the melt inclusions was performed with a Thermo-Nicolet Nexus 670 FTIR spectrometer at the University of Oregon. Each unmounted olivine crystal was individually placed upon a NaCl wafer. Spectral analyses were repeated between three and four times, each analysis encompassing 256 individual scans. Aperture size ranged from 50—120 μ m depending on inclusion diameter. Inclusion thickness, initially determined with a digital micrometer measurement of the olivine wafer, was corroborated through the use of reflectance fringes. To accomplish this, a secondary scan using reflected light was performed with each crystal, and thickness calculated with the wavelength of reflectance fringes, as per Wysoczanski and Tani (2006). Thickness measurements typically agreed to within 5% and calculated measurements were used in almost all cases. In cases in which the fringes were indistinct the initial micrometer measurement was used. H₂O and CO₂ concentrations were calculated using the Beer-Lambert law (Dixon and Pan, 1995). Because the CO₂ absorption coefficient is dependent on the glass composition, final CO₂ concentrations were not calculated until after microprobe analysis. Determination of peak heights for CO₂ required subtraction of background noise (unpublished program, S. Newman). Initial H₂O and CO₂ concentrations, uncorrected for post-entrapment olivine crystallization (explained below), are reported in Appendix A. Due to uncertainties both in inclusion thickness and in absorbance peak heights, average uncertainty at 1 σ in H₂O and CO₂ concentrations is 11%. Accuracy of the FTIR technique is expected to be ±10% for H₂O and ±20% for CO₂ (Dixon, 1991).

Electron Microprobe

After FTIR analysis, olivine hosts were washed, mounted in epoxy and the associated melt inclusions were analyzed for major element composition with the Cameca SX100 electron probe at the University of Oregon. Beam conditions were as follows: 15 kV accelerating voltage, 20 nA beam current, and a beam diameter of 10 microns. Element count times were as follows: 10 s (Mn), 20 s (Ca, S, Cl, Ti), 30 s (Mg, Si), 40 s (K, Al, P), 60 s (Fe). Olivine host compositions were also analyzed, at a spot located a distance of ~100 microns from both the edge of the host crystal and the inclusion. Intensities were corrected by either linear off-peak background correction (K,

Mn, S, Cl, Ti and P) or the mean atomic number (MAN) background correction procedure (Na, Si, Al, Mg, Fe and Ca) of Donovan and Tingle (1996). Measured standard concentrations were within 0.07% of published values, except S, which was within 0.92 % of published values. Percent standard deviation, based on replicate analyses of each inclusion, was \leq 5% for all elements except Mn, which was \leq 7%.

The oxidation states of the magmas were determined through sulfur speciation (S^{6+}/S_{total}) , which was obtained from S K α peak shifts relative to the peak shift of anhydrite. Sulfur peak shifts were measured with the JEOL 8500F field emission electron microprobe at Washington State University (as per Carroll and Rutherford, 1988; Wallace and Carmichael, 1994). Beam conditions were as follows: 15kV accelerating voltage, 30 nA beam current, and a beam diameter of 5 microns. Count times were 5 s for standards and 30 s for unknowns. To avoid oxidation of sulfur during analysis, the beam position was moved 1 μ m each 60 s of analysis. Troilite and pyrrhotite standards were run before and after the glass inclusions to monitor accuracy. Pyrite and anhydrite standards were analyzed at the start and end of the run in order to calculate sulfur speciation (Rowe et al., 2007). Precision, based on replicate analyses, was better than 10% at 1 σ for all samples, except sample IP_003 (± 13%).

LA-ICP-MS

Trace element compositions of melt inclusions were measured with a laser ablation inductively coupled mass spectrometer (VG PQ ExCell quadrupole ICP-MS and NewWave DUV 193 nm ArF Excimer laser) located at Oregon State University. Beam conditions were 3 Hz, beam size was 50 µm with a 45 s total dwell time. Dwell time per element was 10 milliseconds. Glass standard GSE-1G was used for calibration and ⁴³Ca was used as an internal standard. Accuracy at 1σ was $\leq 5\%$ from accepted values. Precision was $\leq 5\%$ for trace elements Ti, V, Sr, Y, Zr, Ba and Ce, $\leq 10\%$ for trace elements Sc, Rb, Nb, La, Pr, Nd, Pb, $\leq 15\%$ for Sm, Eu, Hf, and Th, and $\leq 20\%$ for Gd, Dy, Er, Yb, Ta and U.

Melt Inclusions and data treatment

Inclusions can undergo a number of possible post-entrapment compositional changes. These include post-entrapment crystallization of olivine on the host mineral, diffusive loss of Fe to the host olivine and diffusive loss or gain of H_2O to the host magma (Danyushevsky et al., 2000; Gaetani and Watson, 2000; Portnyagin et al., 2008).

As inclusions cool, there is continued crystallization of olivine on the rim of the inclusion, referred to as post-entrapment crystallization (PEC) (Sobolev, 1996). The effect on major element composition was corrected by adding in olivine (in steps of 0.1 wt. %) until equilibrium between host and inclusion was achieved (Sobolev and Chaussidon, 1996). The effect of PEC on trace element concentrations was corrected by utilizing the batch equilibration crystallization equation, using the % PEC to determine F (the amount of melt remaining) and using the partition coefficients for olivine presented in Appendix C.

Loss of iron occurs when continued inclusion-rim olivine crystallization is slow enough to achieve equilibrium with the rest of the host crystal. If each new increment of olivine re-equilibrates (through iron diffusion), the Mg# of the inclusion stays relatively constant with the host. When olivine is added back into the original composition to correct for PEC, the inherited FeO^{T} will be low (Danyushevsky et al., 2000). Iron loss was corrected using the FeLoss computer program, developed by Danyushevsky et al. (2000), that requires an estimate of the correct FeO^{T} of the inclusion as a starting assumption. This estimate was arrived at by first tentatively correcting the inclusion compositions by adding in incremental olivine until equilibrium with the host was achieved. These compositions were compared to those of whole rock samples from the same units (Figure 4). If the FeO^{T} of the inclusions did not overlap with that of the whole rock (at a particular MgO wt. %), an estimate of FeO^{T} was made that would bring the inclusions in line with whole rock compositions (e.g. Vigouroux et al., 2008). Further olivine addition was then required to bring the inclusions into equilibrium with their hosts.

Portnyagin et al. (2008) demonstrate that diffusive exchange of water can occur over extremely short time-scales (less than 2 days). For this study, using quickly cooled inclusions (tephra) decreases the likelihood of post-entrapment diffusive loss of water. Diffusive loss of water before final ascent and eruption is still possible, however, and is addressed below.

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Results

Measured values, melt inclusions and olivine

The twenty-three measured melt inclusion compositions are primarily basalt, with some basaltic-andesite, and one andesite (47.6-57.1 wt. % SiO₂). These compositions were subsequently corrected for post-entrapment crystallization of olivine on the crystal-inclusion interface, and potential loss of iron through diffusive exchange (see procedure above). Uncorrected, measured compositions are presented in Appendix A, corrected compositions are presented in Table 1. Melt inclusions generally share LKT and CAB composition characteristics with their host lavas.

Olivine crystals that host melt inclusions range in composition from Fo_{80} - Fo_{89} and are presented in Table 2. The most primitive hosts are seen at Indian Pass (Fo_{87} - Fo_{89}). Whitechuck olivine hosts range from Fo_{85} - Fo_{86} , whereas Schreiber's Meadow olivine hosts show the most variability and range from Fo_{80} - Fo_{86} .

Correcting for post-entrapment crystallization requires knowing Fe^{2+}/Mg ratios of both melt and host olivine. Host olivine can be assumed to have $Fe^{2+}/Fe^{T}=1$, thus Fe^{2+}/Mg ratios were obtained from the microprobe analyses reported in Table 2. In the melt, Fe^{2+}/Fe^{T} is constrained by oxygen fugacity. Oxygen fugacity of inclusions was determined through sulfur speciation (see sulfur speciation procedure above) using the empirically derived equation of Wallace and Carmichael (1994), relating sulfur speciation and oxygen fugacity of magmas. Oxidation states (relative to the Quartz-Fayalite-Magnetite (QFM) buffer) for the Whitechuck LKT cinder cone are Δ QFM +1.0 to +1.5 (average Δ QFM +1.3). Oxidation states for Indian Pass CAB inclusions range from Δ QFM +1.6 to +1.8 (average Δ QFM +1.7). Oxidation states for Schreiber's Meadow CAB range from Δ QFM +1.0 to +1.3 (average Δ QFM+1.2). These fugacity values require an initial Fe²⁺/Fe^T of 0.84 for Whitechuck, 0.80 for Indian Pass, and 0.83 for Schreiber's Meadow, calculated as described in Borisov and Shapkin (1990).

DeBari et al. (submitted) do not report fO_2 for Glacier Peak magmas, though they note that it is probably between QFM and the nickel-nickel oxide (NNO) buffers, which is similar, though slightly lower than these results. These oxygen fugacities add to those determined at Mount Baker by Moore and DeBari (submitted). They calculated oxygen fugacity for two of Mount Baker's endmember primitive magma types based on Fe-Ti oxides: the Park Butte LKT-like flow (Δ QFM +2.3 to +2.4; average +2.3), and the Tarn Plateau high-Mg basaltic andesite (Δ QFM +1.4 to +3.2; average +2.1). Neither of Mount Baker's calc-alkaline lava flows (the primitive Lake Shannon or more differentiated Sulphur Creek) contain an ilmenite endmember suitable for these calculations. However, as they noted, temperatures calculated from the Park Butte and Tarn Lake mineral pairs were substantially low, indicating sub-solidus re-equilibration of Fe-Ti oxides.

The olivine Fo values from Table 2, along with appropriate distribution coefficient (K_D) values for Fe/Mg equilibria in olivine-liquid pairs, were used to calculate equilibrium melt inclusion Mg#s. Kd values were determined as per Toplis (2005). Average Kd values were: 0.32 ±0 .01 (Whitechuck), 0.31 ±0.02 (Indian Pass), and 0.32 ±0.02 (Schreiber's Meadow).

Major element compositions of melt inclusions were corrected for postentrapment crystallization of olivine. Trace element and volatile concentrations were also corrected for post-entrapment crystallization, though due to their generally incompatible nature in olivine, these corrections were minor. Olivine was incrementally added back into the melt inclusion compositions until the inclusion compositions were in equilibrium with their host olivine. Most Whitechuck inclusions required additional olivine \leq 4.7 wt. % (one inclusion required 8.2 wt. %). Indian Pass inclusions required additional olivine between 0.6 and 5.2 wt. % (average 3.5 wt. %). Schreiber's Meadow inclusions required additional olivine between 0.9 to 7.0 wt. % (average 3.6 wt. %).

Oxygen fugacity must also be corrected for post-entrapment crystallization; as Fe^{2+} is added back into the system, the Fe^{2+}/Fe^{T} ratio will change. Oxygen fugacity was corrected by calculating new fO2 as each new 0.1 wt. % increment of olivine was added back into the inclusion composition. Corrected oxidation states for the Whitechuck LKT cinder cone are $\Delta QFM + 0.8$ to +1.4 (average $\Delta QFM + 1.1$). Oxidation states for Indian Pass CAB inclusions range from $\Delta QFM + 1.4$ to +1.5 (average $\Delta QFM + 1.5$). Oxidation states for Schreiber's Meadow CAB range from $\Delta QFM + 0.7$ to +1.1 (average $\Delta QFM + 1.0$). Corrected oxidation states are presented in Table 1.

A total of eight inclusions were corrected for post-entrapment iron-loss (Figure 4). Three inclusions from Whitechuck required an average of 0.84 wt. % addition of FeO^{T} . Three inclusions from Indian Pass required an average of 0.44 wt. % addition of FeO^{T} . Two inclusions from Schreiber's Meadow required an average of 0.68 wt. % addition of FeO^{T} (Table 1).

Corrected major element compositions for melt inclusions in the calc-alkaline lavas (Schreiber's Meadow at Mt. Baker and Indian Pass at Glacier Peak) generally fall within the medium-K calc-alkaline basalt to basaltic-andesite field (Figure 5). However, one Indian pass melt inclusion is a high-K basaltic andesite. One Schreiber's Meadow melt inclusion is a high-K andesite and one is a low-K tholeiitic basaltic andesite. Melt inclusions for the Whitechuck LKT extend from the low-K tholeiite field to the boundary between LKT and medium-K basalt (Figure 5). These compositions are consistent with whole rock analyses of these sites and also overlap with calc-alkaline and tholeiitic compositions along the Cascade arc (Figure 5).

Major elements plotted on Harker diagrams (all elements are compared on an anhydrous basis) show good overlap with the bulk rock compositions at Whitechuck (Figure 6). Indian Pass major elements are slightly offset (i.e.: higher CaO, Na₂O, TiO₂, Al_2O_5 and lower MgO, FeO^T all at lower SiO₂) from whole rock compositions, yet cluster tightly, except for sample IP_005, which is more evolved and shows evidence of significant degassing (see below). Schreiber's Meadow inclusions show the most extreme scatter in comparison with bulk rock from the associated Sulphur Creek lava flow (Figure 7).

Trace element compositions are generally more variable than major elements seen in the melt inclusions. However, Whitechuck trace element compositions are relatively homogeneous, and overlap tightly with bulk rock compositions. Indian Pass compositions are slightly offset to higher (Ba, La, Nd, Sr) or lower (Rb, Yb, and Zr) values at lower SiO₂ (Figure 8). Schreiber's Meadow inclusions are the most variable, and do not group tightly near bulk rock compositions (Figure 9). Rare earth element patterns and trace element diagrams reflect the varying degrees of overlap or, alternatively, scatter seen in Harker diagrams. Whitechuck rareearth element (REE) patterns overlap those of bulk rock analyses. Both Indian Pass and Schreiber's Meadow REE patterns are more variable than their respective bulk rock compositions—however average values for Indian Pass compositions overlap whole rock compositions (Figs 10, 11). Even with the variability, trace element patterns from melt inclusions corroborate data recorded by whole rocks. Glacier Peak LKT have flatter REE patterns and lower abundances of high-field strength elements (HFSE) and middle REE than CAB at Mount Baker and at Glacier Peak. Glacier Peak CAB have steeper REE than CAB at Mount Baker. LILE elements are enriched in all three units, and some LILE/HFSE ratios are higher at Whitechuck (Ba/Nb ~75, Pb/Ce~0.18) in comparison to either Schreiber's Meadow (Ba/Nb ~ 51, Pb/Ce~0.12) or Indian Pass CAB (Ba/Nb ~51, Pb/Ce~0.12). Proxies for H₂O (Ba/Zr and Pb/Ce) span the range of LKT in the southern arc (Figure 12), though Pb/Ce extends to higher values than elsewhere.

Volatile concentrations

Volatile concentrations were corrected for post-entrapment crystallization (above) and are presented in Table 1. Corrected for post-entrapment crystallization, Whitechuck melt inclusion H₂O concentrations are 0.5-2.0 wt % (average 1.3 wt %). CO₂ concentrations range from 155 ppm to as high as 821 ppm (average 485 ppm), S concentrations are 1200-1690 ppm, and Cl is 910-1010 ppm.

Corrected Indian Pass H₂O contents are 1.1-2.2 wt. % (average 1.6 wt. %), CO₂ concentrations are between <50 ppm to 1720 ppm (average of measurable CO₂, 1140 ppm), S concentrations are 2550-3320 ppm, and Cl concentrations are 418-711 ppm.

Schreiber's Meadow inclusions have corrected H_2O contents of 0.8-2.3 wt. % (average 1.6 wt. %), CO₂ concentrations of <50 ppm-870 ppm (average of measurable 451 ppm), S concentrations between 170 (in the andesitic inclusion) and 1526 ppm, and Cl concentrations of 205-477 ppm (Table 1).

H₂O and CO₂ concentrations have been used to indicate entrapment pressures based on known solubilities in basaltic magmas. Vapor saturation pressures were calculated for each inclusion using the VolatileCalc program (Newman and Lowenstern, 2002) (Table 1, Figure 13). Entrapment pressures range from 6 to 374 MPa, though these calculated pressures may be affected by H₂O or CO₂ variability due to degassing, loss of CO₂ to a shrinkage bubble post-entrapment, or post-entrapment loss of H₂O due to diffusion. Therefore these pressures should be considered minimum entrapment pressures. The potential effects of degassing on measured volatile concentrations (and therefore calculated pressures) will be further developed below. Whitechuck inclusions record entrapment conditions between 44 and 202 MPa, Indian Pass between 82 and 374 MPa and Schreiber's Meadow between 6 and 255 MPa. These conditions correspond to crustal entrapment depths ranging between 0.2 and 12 km (assuming average upper crustal density of 2.8 g/cm³). Recorded maximum entrapment depths are: ~7 km (Whitechuck), ~12 km (Indian Pass) and ~9 km (Schreiber's Meadow).

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As vapor exsolves from a rising magma, the liquidus temperature of the magma increases, resulting in crystallization. Therefore, crystallization may be expected to correlate with loss of water (e.g. Sisson and Layne, 1993). At Schreiber's Meadow, with the exception of two inclusions, H₂O loosely correlates negatively with K₂O and Na₂O and positively with MgO—expected correlations for a magma that is exsolving H₂O and crystallizing as it rises (Figure 14).

Indian Pass H_2O concentrations show a loose negative correlation with K_2O , but H_2O does not show a correlation with decreasing MgO (Figure 14). Thus, it is difficult to make any interpretation of simultaneous crystallization accompanying degassing at Indian Pass.

At Whitechuck, no trends are seen with changes in H_2O content (Figure 14). An explanation for this may be rapid syneruptive degassing, which would not result in significant crystallization, as seen at Mount St. Helens (Blundy and Cashman, 2005).

Discussion

Volatiles

The use of olivine-hosted melt inclusions in primitive magmas is intended to discern volatile concentrations in the most primitive magma in a system and relate these concentrations to mantle source compositions and processes. However, interpretations are predicated on the assumption that these volatile contents are not the result of either crustal interaction or post-entrapment changes. These assumptions may be invalidated by pre-entrapment assimilation of crustal material or mixing with a second, volatile-rich magma. Complicating matters further, volatile content is constantly being adjusted due to degassing (volatile depletion) and pre-entrapment fractionation (incompatible volatile enrichment). There is also experimental evidence for open-system behavior of some volatiles such as H₂O (changed via hydrogen diffusion) (Portnyagin et al., 2008). I evaluate some of these complications below, in regards to measured H₂O contents at Glacier Peak and Mount Baker.

Volatile Variability at Glacier Peak

Pre-entrapment magma mixing is unlikely at Whitechuck or at Indian Pass as no chemical or petrographic evidence for this process is observed (this study; DeBari et al., submitted). Isotopic data presented by DeBari et al. (submitted) suggest that both Whitechuck and Indian Pass are internally isotopically homogeneous, in contrast to Glacier Peak's Lightning Creek basaltic andesite, which does show both chemical and petrographic evidence for extensive mixing with crustally-derived dacite. Thus, H₂O contents of Whitechuck and Indian Pass magmas are not expected to have been elevated by mixing with other, more H₂O-rich magmas.

Only minimal crystal fractionation is probable at Indian Pass given its nearly aphyric nature and high Cr and Ni contents in whole rock analyses, which are above established minimum values for a primitive composition (Ni> 100 ppm, Cr >200 ppm) (DeBari et al, submitted). Also, as noted earlier, H₂O contents show a loose anticorrelation with increasing K₂O wt. % (a proxy for fractionation) at Indian Pass. At Whitechuck, tephras are aphyric, but low Cr and Ni contents in whole rock analyses (DeBari et al., submitted) suggest higher amounts of pre-entrapment fractionation of olivine. Corrections for this additional crystal fractionation are made below.

Due to the primitive nature of Indian Pass CAB magmas, and the low Rb and Sr of Whitechuck LKT, similar to other arc LKTs, crustal contamination is also unlikely (Taylor, 2001). Thus, measured H₂O and other volatile concentrations are assumed to be indicative of mantle melt sources, enriched through variable amounts of fractionation, and potentially affected by degassing or diffusion.

Whitechuck

H₂O concentrations (corrected for post-entrapment crystallization) for the Whitechuck LKT magma (up to 2.0 wt. %) are within typical arc values, between 1 and 6 wt. % (Wallace, 2005). H₂O and other volatile contents show wide-ranging variability, despite the homogeneity of the magma (Figure 15). While low, the measured values are still higher than expected, as the source for the Whitechuck LKT magmas has been previously interpreted as relatively anhydrous (Taylor, 2001), similar to other Cascade LKT magmas. Measured H₂O concentrations and the slightly stronger arc trace element signature seen in this northern LKT are permissive of a more prominent role for fluidflux melting. Based on the relative primitiveness of the Whitechuck LKT and the lack of evidence for magma mixing or crustal contamination, I conclude that H₂O variability is likely due either to pre-entrapment degassing or post-entrapment diffusion through the host olivine or some combination of both. Thus, maximum measured H₂O contents at Whitechuck are taken as minimum values and are used when correcting for the effect of fractionation from mantle sources. These H_2O contents are also in good agreement with estimated H_2O contents (max ~2 wt. %) determined utilizing measured major element compositions of plagioclase phenocryst rims (collected by Taylor, 2001), the Ca/Na ratio of the melt (represented by the whole rock compositions), and the Ca/Na plagioclase hygrometer of Lange et al. (2009).

Indian Pass

 H_2O concentrations (corrected for post-entrapment crystallization) at Indian Pass (up to 2.2 wt. %) also overlap the low end of arcs (Table 1). While no evidence for crustal interaction processes is evident (as discussed above; DeBari et al., submitted), degassing and diffusion may have affected H_2O concentrations. However, measured H_2O contents are in good agreement with those determined by the hygrometer (max ~2 wt. %) of Lange et al. (2009). Again, I will assume that maximum measured values are representative of minimum values in the primitive magma.

Volatile Variability at Mount Baker

At the Sulphur Creek lava flow, associated with the Schreiber's Meadow cinder cone, there is significant petrographic and chemical evidence for disequilibrium, indicative of magma mixing (Moore and DeBari, submitted). It is therefore likely that H₂O content at Schreiber's Meadow have been affected by crustal interaction. There are much more primitive calc-alkaline basalts present at Mount Baker (i.e. Lake Shannon), yet these have no associated tephra (Hildreth et al., 2003) and are therefore not utilized in this study. H_2O contents at Schreiber's Meadow also overlap at the low end of arcs (up to 2.3 wt. %), and agree with the plagioclase hygrometer (max ~2 wt. %) of Lange et al. (2009). Decreasing H_2O with increasing alkalis, as well as a positive correlation with MgO (Figures 14,15), are consistent with degassing associated with crystallization and thus the maximum concentrations may be taken as minimum values within the crust. However, due to the significant evidence for disequilibrium (petrographic textures, variability of both major and trace elements), I do not refer to Schreiber's Meadow in the following discussions about mantle source compositions or associated H_2O contents in the southern Garibaldi volcanic belt.

Parental Magmas

While olivine-hosted melt inclusions record the most primitive magma accessible for analysis, they typically do not record compositions in equilibrium with their initial mantle source peridotite (olivine Fo≥90). Therefore, further corrections must be made in order to estimate the compositions of the parental melts of these primitive magmas. The simplest assumption is that the primitive magmas (as represented by the melt inclusions) have fractionated primarily olivine since last equilibration with the mantle. While the whole rock phenocryst assemblage contains minor (<1 %) plagioclase, this method of olivine addition is supported by isobaric, equilibrium crystallization modeling of the measured primitive melts by entering calculated parental melts into both the PETROLOG program (Danyushevsky and Plechov, 2008) and into the MELTS algorithm via the Adiabat interface (Smith and Asimow, 2005). This modeling indicates that olivine is the major crystallizing phase until 11% crystallization (both PETROLOG and MELTS) for Indian Pass compositions and until either 16% crystallization (PETROLOG) or 14% crystallization (MELTS) for Whitechuck compositions. After these amounts of crystallization are achieved, plagioclase becomes a fractionating phase, explaining its minor presence in the final phenocryst assemblage. Assuming mantle wedge olivine to be Fo₉₀, corrections for crystal fractionation (using the same method and partition coefficients as for post-entrapment crystallization) were made to equilibrate inclusion compositions to mantle compositions. Addition of 13.6-15% olivine was necessary to bring Whitechuck melt inclusion compositions to equilibrium with mantle olivine, and 5.5-7.7 wt. % olivine addition was required for Indian Pass. These calculated equilibrium mantle melt compositions (trace elements, major elements and volatiles) are presented in Table 3.

Temperature and Pressure

The ultimate aim of this study is to constrain melt generation processes in the northern Cascade arc. To do this, calculated equilibrium mantle melt compositions (and associated minimum H₂O contents) from Glacier Peak are used to constrain temperature and pressure conditions of the primary melt as it segregated from the mantle and to model source composition[s], potential slab inputs and their compositions.

Using the melt geothermometer and barometer of Lee et al. (2009) with calculated equilibrium mantle melt compositions, conditions of last equilibration with the mantle would have been 1278-1317 (\pm 40)°C and 1.2-1.6 (\pm 0.2) GPa for Whitechuck and
between 1237-1250 (\pm 40)°C and 0.9-1.4 (\pm 0.2) GPa for Indian Pass (Figure 16). The lowest calculated equilibration pressure at Indian Pass is obtained from sample IP_005, which has an anomalous composition (Figs. 6, 8), and is unlikely to be representative—other pressures range between 1.2 and 1.4 GPa.

Calculated pressures place equilibration at depths of $41-54 (\pm 7)$ km (mafic crustal density of 3.0 g/cm³) for Whitechuck and between 45-48 (\pm 7) km for Indian Pass. Crustal depth is considered to be between 40 and 45 km (Mooney and Weaver, 1989), therefore, the minimum depth of equilibration should be considered to be at or near the base of the crust (Figure 16). This would place depths within the spinel peridotite field. This is consistent with REE patterns, which are not highly depleted in HREE and enriched in LREE, as are those expected with magmas equilibrated within the garnet stability field. Surprisingly, within error, Whitechuck and Indian Pass magmas equilibrated at the same, or a quite similar depth. Calculated temperatures of Indian Pass CAB are 8-45°C cooler than the mantle dry solidus temperature of Hirschmann (2000) (33-45°C cooler when IP 005 is excluded) and Whitechuck equilibration temperatures are between 4°C cooler to 6°C hotter than the dry solidus. An adjustment of -35°C to the dry solidus calculation of Hirschmann (2000) is recommended by Katz et al. (2003), placing Indian Pass compositions at or near the solidus and Whitechuck compositions ~40°C hotter than the solidus. In either case, within error of the geothermometer and geobarometer (± 0.2 GPA and $\pm 40^{\circ}$ C), melts could have equilibrated at temperatures below the solidus (indicative of fluid-fluxed melting) or above the solidus (permissive of dry decompression melting).

Melting Processes

Measured H₂O values and calculated mantle equilibration temperatures all are permissive of fluid-fluxed melting. I evaluate the following potential mantle sources and melting scenarios to constrain whether they can produce the observed H₂O contents in Glacier Peak LKT and CAB magmas. These include 1) dry decompression melting of a source with no hydrous phases (limited hydration capacity in anhydrous minerals), 2) dehydration melting of a metasomatized mantle source with hydrous phases, or 3) fluidfluxed melting of a mantle source. In each case, the maximum measured H₂O contents (assumed to be the minimum content in the primitive magma), further corrected to mantle equilibration (Fo₉₀), are used [1.7 wt % for LKT, 2.0 wt. % for CAB (Table 3)]. To be permissible, melt processes must be compatible with measured H₂O concentrations, calculated primary melt compositions, and likely mantle source characteristics.

Mantle source characteristics can be identified by focusing on trace elements that are not considered to be enriched from a slab component (such as Nb, Zr, Y, Yb and Ti), known collectively as conservative elements. Plots of these elements, such as Nb and Y (Figure 17), indicate two distinct mantle sources for the LKT and CAB magmas; one more depleted (Whitechuck LKT) and one more enriched (Indian Pass CAB). For the Whitechuck LKT, the possible depleted mantle source is likely compositionally similar to a MORB source (Figure 17). This is further supported by spinel Cr#/olivine Fo systematics reported by Taylor (2001). Nb/Y systematics suggest that the enriched mantle source for the Indian Pass CAB may be similar to that of an enriched mid-ocean ridge basalt (EMORB) (Figure 17). Corrected Indian Pass trace elements were normalized to EMORB and plotted against an EMORB composition (which plots as a straight line at 1). Conservative elements from Indian Pass plot at or near 1 (excepting Zr), lending further support for an EMORB-like source (Figure 18). These source compositions (a depleted mantle source for Whitechuck LKT and an enriched mantle source for Indian Pass CAB) place an additional constraint on magma generation processes.

If dry decompression melting is responsible for magma genesis, then the H₂O contents of the Glacier Peak LKT and CAB could be the result of varied amounts of partial melting of a nominally anhydrous mantle wedge. The storage capacity of H₂O in the mantle is between 300 and 1000 ppm (Dixon, 2001; Dixon et al., 2002; Jamtveit et al., 2001). Assuming the maximum H₂O storage capacity of the mantle, solving for the simple batch melting equation shows that the maximum partial melt of a mantle source containing ~ 1000 ppm H₂O, would be $\sim 3.8\%$ melt (Whitechuck) or ~ 3.5 wt. % melt (Indian Pass), assuming a mantle/melt partition coefficient for H₂O of 0.012 (Stolper and Newman, 1994; Workman and Hart, 2005). Then, solving the batch melting equation for each individual trace element utilizing these melt fractions (partition coefficients for incompatible trace elements in peridotite are also from Workman and Hart, 2005), the trace element signature of the mantle source peridotite can be determined for a dry melting scenario (with no hydrous phases). While this modeling does suggest two separate source regions (Figure 19), both are more depleted in conservative elements (Nb, Zr, Y, Yb, Ti) than the mantle source constraints determined above. Thus, dry decompression melting of a source with no hydrous phases is not a permissible melt generation scenario.

Alternatively, H₂O could potentially be stored in hydrous mantle phases.

Decompression-driven partial melting of these hydrous phases could account for high amounts of H₂O, including (and in excess of) measured concentrations, independent of fluid-fluxed melting. Such phases might include antigorite, amphibole, talc, chlorite or phlogopite. However, of these phases, only phlogopite is stable at pressures and temperatures expected within the mantle wedge beneath the arc (Fumagalli and Poli, 2005; Grove et al., 2006; Poli and Schmidt, 2002; Yoder Jr and Kushiro, 1969). Amphibole, for example, is only stable to ~ 1100 °C, limiting its presence to just above the subducting slab (Holloway and Burnham, 1972). Stoichiometrically, phlogopite contains K_2O and H_2O in a ratio of 2:1. Relationships between source region phases and abundances of H₂O, K₂O and Na₂O were explored on a ternary diagram for four arc volcanoes (Sisson and Layne, 1993), who determined that melting of phlogopite would shift compositions heavily towards K₂O (Figure 20). One would expect, then, that the K_2O/H_2O ratio at either Indian Pass (K_2O/H_2O : average 0.53, excluding degassed inclusion 0.65), or at Whitechuck (K₂O/H₂O average 0.28, or, excluding the three most degassed inclusions, average 0.73) would be much higher if derived from a source containing phlogopite. Thus, dry melting of a mantle source containing hydrous phases is incompatible with calculated primary melt compositions.

While both decompression melting scenarios can account for measured H_2O concentrations, they are either incompatible with mantle source compositions or with calculated primary melt compositions. Therefore, fluid-fluxed melting may account for generation of both LKT and CAB magmas. This is inconsistent with the complete

dehydration of the downgoing slab predicted in hot arcs (e.g. Leeman et al., 2005b) and supported by some thermal models of arcs (Hacker, 2008). However, newer thermal models suggest that while the sediment and crustal components of the subducting slab are entirely dehydrated, the slab can contain H₂O in serpentinized slab mantle, if only to depths of 115 km (van Keken et al., 2011).

To evaluate the possibility that fluid-fluxed melting can account for both LKT and CAB magmas at Glacier Peak; I model fluid-fluxed melting of a heterogeneous mantle wedge, starting with the constraints on mantle sources (depleted source for Whitechuck LKT and an enriched source for Indian Pass CAB) and following the method of Portnyagin et al. (2007). As mantle wedge source material has likely undergone some previous melt extraction (PME), Portnyagin et al. (2007) developed a method to calculate how much (if any) PME is required of a typical source composition (e.g., MORB or EMORB source) to most closely match the actual source that would be melted to create a calculated parental melt. Using this source composition, the degree of melting (F) of a mantle source (corrected for PME) can then be determined (e.g. Kelley et al., 2006; Kelley et al., 2010; Portnyagin et al., 2007). In addition to the amount of melting and PME, the procedure of Portnyagin et al. (2007) creates an estimate of the composition of the fluxed slab contribution.

The initial mantle source compositions are similar to MORB (Whitechuck LKT) and EMORB (Indian Pass CAB). Two depleted MORB sources are reported in the literature (Salters and Stracke, 2004; Workman and Hart, 2005), and both were evaluated as potential (pre-PME) Whitechuck LKT mantle sources. A commonly used EMORB source (for Indian Pass CAB) is not reported in the literature. However, a similarly heterogeneous mantle is noted in the Trans-Mexican Volcanic Belt, where enriched EMORB-like source regions are interpreted to result from previous melt extraction of an enriched OIB-like source followed by addition of a slab component (Vigouroux et al., 2008). Other workers have also considered CAB arc magmas to be the result of variably depleted OIB-like mantle sources contaminated with a subduction component (Johnson et al., 2009; Ruscitto et al., 2010b). To constrain CAB petrogenesis in the Central Oregon Cascade arc, Ruscitto et al. (2010b) assumed that three different central Cascade OIB compositions (Rowe et al., 2009) were each the result of 5% melting of variably enriched mantle sources. Using these three OIB compositions, they reconstructed initial mantle source compositions. Though there are no known eruptions of OIB-like magmas at Mount Baker or at Glacier Peak, an OIB-like component in the mantle wedge is interpreted throughout the Cascade arc (Conrey et al., 1997; Leeman et al., 1990; Rowe et al., 2009; Schmidt et al., 2008). Since an OIB-like component is a common Cascade arc feature, three Cascade arc OIB-sources from Ruscitto et al. (2010b) were evaluated as potential (pre-PME) Indian Pass CAB mantle sources.

The best fit for the average calculated Whitechuck parental melt indicates a mantle source similar to either the DMM of Salters and Stracke (2004), with previous melt extraction of 0.001 mass fraction, or a mantle source that is similar to the E-DMM of Workman and Hart (2005), with PME of 0.002 mass fraction. The best fit for the average Indian Pass parental melt indicates a source similar to that of the Central Oregon Cascade OIB sample WFO21 (Ruscitto et al., 2010b), with no previous melt extraction.

As mantle modal mineralogy is unconstrained beneath Glacier Peak, bulk partition coefficients (D) for melting within the spinel peridotite field were used (Kelley et al., 2006; Workman and Hart, 2005) (Appendix C). REE patterns and equilibration pressures of parental magmas support this assumption.

The degree of melting (F) required to produce Whitechuck LKT from a PMEadjusted mantle sources is either 15% melting (F: 0.151 ± 0.05) of a DMM source or 13.5% melting (F: 0.135 ± 0.067).of an E-DMM source. Indian Pass CAB can be generated by 13.7% melting (F: 0.137 ± 0.030 , bulk D from Workman and Hart (2005), or F: 0.137 ± 0.026 , bulk D from Kelley et al. (2006)) of a mantle source similar to that of Oregon Cascades sample WF021 (Ruscitto et al., 2010b).

Degrees of melting (above), input into the batch melting equation, were used to solve for the abundances of non-conservative trace elements and volatiles in the mantle source of parental magmas. The difference between these calculated abundances and those expected in the PME-corrected mantle source isolates the incompatible elements and volatiles attributable to the slab contribution. This provides a first-order constraint on the minimum H₂O present in the slab component and the amount of subduction component added to the mantle wedge. This modeling indicates that the Whitechuck LKT mantle source has been fluxed either with 0.34 wt. % subduction component (with the DMM initial source) or 0.30 wt. % subduction component (E-DMM initial source). The H₂O predicted to be present in the slab is either 0.28 wt. % (DMM) or 0.25 wt. % (E-DMM). Modeling indicates that the Indian Pass CAB mantle source has been fluxed wither with 0.28 or 0.29 wt. % subduction component containing either 0.21 or 0.22 wt.

% H₂O, depending on whether bulk partition coefficients are from Salters and Stracke (2004) or Workman and Hart (2005).

Conclusions

H₂O contents measured at primitive magmas in the southern Garibaldi Belt, at Mount Baker and Glacier Peak, and corrected for post-entrapment crystallization, range between 2.0 and 2.3 wt. %, overlapping the global range for arc volcanoes. Due to uncertain extents of crustal interaction, H₂O contents at Mount Baker cannot be related to parental values. Glacier Peak primitive LKT (Whitechuck cinder cone) and CAB (Indian Pass cinder cone) lavas, contain minimum H₂O contents of 2.0 and 2.2 wt. %, respectively. Parental magmas (in equilibrium with mantle olivine [Fo90]), are expected to have 1.7 and 2.0 wt. % H₂O. Trace element, isotopic data, and modeling suggest that these H₂O concentrations are unlikely to be the result of crustal interaction, decompression melting of the mantle wedge or of decompression melting of hydrous phases within the mantle wedge.

I conclude that both LKT and CAB at Glacier Peak are the result of between 13 and 15% fluid-fluxed melting of a compositionally heterogeneous source, last equilibrated near the base of the crust, with maximum temperatures at or near the dry peridotite solidus. Source heterogeneities likely include a MORB-like component and an OIB-like component. Minimum H₂O contents suggested in the slab component are between 0.21 and 0.28 wt %. This is in contrast to southern regions of the Cascade arc, where LKT magmas are considered as the result of dry decompression melting of a MORB-like source, though it is consistent with the more pronounced arc signature seen in this northern LKT.

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Sample	WC		WC		WC		WĆ		WC		WC		WC		WC	
Sample	002_1	ls.e.	002_2	ls.e.	003	ls.e.	007	ls.e.	008	ls.e.	013	ls.e.	014	ls.e.	017	ls.e.
SiO ₂	50.86	0.32	50.71	0.15	50.05	0.31	51.44	0.24	50.75	0.11	50.28	0.21	50.98	0.66	50.65	0.25
TiO ₂	1.16	0.04	1.12	0.05	1.13	0.05	1.29	0.05	1.10	0.08	1.14	0.05	1.13	0.06	1.15	0.07
Al_2O_3	18.59	0.17	18.72	0.11	19.21	0.21	18.75	0.16	18.22	0.11	18.83	0.11	18.32	0.48	18.52	0.07
FeO	7.93	0.08	7.96	0.13	7.92	0.09	7.51	0.06	8.02	0.03	8.04	0.12	7.99	0.61	8.01	0.08
MnO	0.15	0.01	0.15	0.01	0.13	0.01	0.14	0.01	0.13	0.01	0.13	0.01	0.14	0.02	0.14	0.01
MgO	6.82	0.26	6.83	0.13	6.71	0.06	6.04	0.18	6.84	0.12	6.89	0.05	6.80	0.79	6.92	0.05
CaO	10.32	0.07	10.32	0.04	10.54	0.03	10.49	0.05	10.69	0.03	10.48	0.04	10.13	1.12	10.39	0.02
	3.43	0.15	3.44	0.13	3.50	0.09	3.52	0.06	3.46	0.05	3.43	0.09	3.72	0.40	3.47	0.14
K_2O	0.39	0.01	0.40	0.01	0.44	0.01	0.47	0.01	0.44	0.02	0.43	0.01	0.44	0.02	0.41	0.01
P_2O_5	0.22	0.01	0.21	0.01	0.22	0.01	0.23	0.00	0.22	0.02	0.21	0.01	0.21	0.00	0.20	0.01
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H_2U	1.5	0.2	1.0	0.1 50	0.5	0.1	2.0	0.1	1.9	0.1	1.1	20	0.0	0.1	1.9	0.1
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+ΔFMQ	1.4		1.4		1.0		1.0		1.4		1.1		1.1		1.3	
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Whitecht Sample	<i>uck LK</i> WC	T	WC 020_2	150	WC 023	150	Indian IP 001	Pass	<i>CAB(la</i> IP 003_1	ocatio	on:47°5 IP 005	9'58' Ise	<u>'',-121°(</u> IP 016	06'54 Ise	") IP 020	15.0
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Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃	uck LK WC 020_1 50.76 1.17 18.47	T <u>1s.e.</u> 0.24 0.06 0.53	WC 020_2 50.31 1.13 18.97	<i>1s.e.</i> 0.06 0.05 0.21	WC 023 50.05 1.13 17.82	<i>1s.e.</i> 0.26 0.06 0.15	<i>Indian</i> IP 001 50.38 1.38 17.55	Pass 1s.e. 0.27 0.11 0.08	CAB(ld IP 003_1 49.35 1.55 17.70	<i>ls.e.</i> 0.15 0.03 0.08	on: 47°5 IP 005 53.72 2.09 15.74	9'58' <u>Is.e.</u> 0.55 0.02 0.08	^{''} ,- <i>121</i> °(IP 016 50.19 1.49 17.88	06'54 <u>Is.e.</u> 0.23 0.05 0.19	") IP 020 50.10 1.42 17.52	<i>1s.e.</i> 0.34 0.04 0.29
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T	uck LK WC 020_1 50.76 1.17 18.47 8.03	T <u>1s.e.</u> 0.24 0.06 0.53 0.30	WC 020_2 50.31 1.13 18.97 8.01	1s.e. 0.06 0.05 0.21 0.19	WC 023 50.05 1.13 17.82 8.62	<i>1s.e.</i> 0.26 0.06 0.15 0.04	<i>Indian</i> IP 001 50.38 1.38 17.55 7.23	Pass 1s.e. 0.27 0.11 0.08 0.08	CAB(ld IP 003_1 49.35 1.55 17.70 7.21	<u>ls.e.</u> 0.15 0.03 0.08 0.07	<i>pn:47°5</i> IP 005 53.72 2.09 15.74 7.01	9,58, <u>1s.e.</u> 0.55 0.02 0.08 0.10	<u>',-121°(</u> IP 016 50.19 1.49 17.88 7.01	<u>1s.e.</u> 0.23 0.05 0.19 0.07	") IP 020 50.10 1.42 17.52 7.22	<i>1s.e.</i> 0.34 0.04 0.29 0.24
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13	T 1s.e. 0.24 0.06 0.53 0.30 0.01	WC 020_2 50.31 1.13 18.97 8.01 0.12	1s.e. 0.06 0.05 0.21 0.19 0.01	WC 023 50.05 1.13 17.82 8.62 0.16	1s.e. 0.26 0.06 0.15 0.04 0.00	<i>Indian</i> IP 001 50.38 1.38 17.55 7.23 0.11	Pass 1s.e. 0.27 0.11 0.08 0.01	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10	<i>1s.e.</i> 0.15 0.03 0.08 0.07 0.01	<i>m:47°5</i> IP 005 53.72 2.09 15.74 7.01 0.11	9 '58 ' <u>1s.e.</u> 0.55 0.02 0.08 0.10 0.02	<u>',-121°(</u> IP 016 50.19 1.49 17.88 7.01 0.12	<u>1s.e.</u> 0.23 0.05 0.19 0.07 0.00	") IP 020 50.10 1.42 17.52 7.22 0.10	<i>1s.e.</i> 0.34 0.04 0.29 0.24 0.01
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90	T 0.24 0.06 0.53 0.30 0.01 0.51	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18	WC 023 50.05 1.13 17.82 8.62 0.16 7.36	<i>1s.e.</i> 0.26 0.06 0.15 0.04 0.00 0.22	<i>Indian</i> IP 001 50.38 1.38 17.55 7.23 0.11 6.97	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24	<i>Is.e.</i> 0.15 0.03 0.08 0.07 0.01 0.23	<i>i</i> :47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59	9'58' <u>1s.e.</u> 0.55 0.02 0.08 0.10 0.02 0.06	IP 016 50.19 1.49 17.88 7.01 0.12 6.37	<u>1s.e.</u> 0.23 0.05 0.19 0.07 0.00 0.06	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16	<i>Is.e.</i> 0.34 0.04 0.29 0.24 0.01 0.57
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41	T 1s.e. 0.24 0.06 0.53 0.30 0.01 0.51 0.02	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14	1s.e. 0.26 0.06 0.15 0.04 0.00 0.22 0.05	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52	Is.e. 0.15 0.03 0.07 0.01 0.23 0.06	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07	',-121° IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80	1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04
Whitech Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36	Is.e. 0.24 0.06 0.53 0.30 0.01 0.51 0.02 0.24	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02 0.04	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90	1s.e. 0.26 0.06 0.15 0.04 0.00 0.22 0.05 0.12	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60	Is.e. 0.15 0.03 0.07 0.01 0.23 0.06 0.14	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09	',-121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52	1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41	Is.e. 0.24 0.06 0.53 0.30 0.01 0.51 0.02 0.24	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45	1s.e. 0.06 0.21 0.19 0.01 0.18 0.02 0.04 0.01	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47	1s.e. 0.26 0.06 0.15 0.04 0.00 0.22 0.05 0.12 0.01	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.03	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98	Is.e. 0.15 0.03 0.07 0.01 0.23 0.06 0.14 0.04	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73	9'58' <u>1s.e.</u> 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02	',-121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91	1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79	Is.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 0.21	Is.e. 0.24 0.06 0.53 0.01 0.51 0.02 0.24	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02 0.04 0.01 0.02	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22	1s.e. 0.26 0.06 0.15 0.04 0.02 0.05 0.12 0.01	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53	Is.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.03 0.02	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61	Is.e. 0.15 0.03 0.07 0.01 0.23 0.06 0.14 0.01	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78	9'58' <u>1s.e.</u> 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01	P 121°C IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56	1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02 0.00	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53	Is.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.20
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ S ^a	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 0.21 1650	Is.e. 0.24 0.06 0.53 0.30 0.01 0.51 0.02 0.24 0.03 0.01 70	WC 020 2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23 1740	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02 0.01 0.02 10	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22 1200	1s.e. 0.26 0.06 0.15 0.04 0.00 0.22 0.05 0.12 0.01 0.01	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53 2960	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.02 20	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61 2860	Is.e. 0.15 0.03 0.07 0.01 0.23 0.06 0.14 0.04 0.01	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78 89	9'58' 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01 10	P 121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56 2550 2550	1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.12 0.00 30	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53 3320	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.29
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ S ^a Cl	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 0.21 1650 921	Is.e. 0.24 0.06 0.53 0.30 0.01 0.51 0.02 0.24 0.03 0.01 30	WC 020 2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23 1740 929	1s.e. 0.06 0.21 0.19 0.01 0.18 0.02 0.01 0.02 0.03	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22 1200 917	1s.e. 0.26 0.06 0.15 0.04 0.00 0.22 0.05 0.12 0.01 40 40	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53 2960 446	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.02 20 30	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61 2860 520	Is.e. 0.15 0.03 0.07 0.01 0.23 0.06 0.14 0.04 0.01 10	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78 89 711	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01 10 10	',-121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56 2550 447	D6'54 1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02 30 30	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53 3320 418	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.01 90 40
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ S ^a Cl H ₂ O ^b	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 0.21 1650 921 1.3	Is.e. 0.24 0.06 0.53 0.30 0.01 0.51 0.02 0.24 0.03 0.01 70 30 0.1	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23 1740 929 0.8	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02 0.04 0.02 10 30 0.1	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22 1200 917 1.1	1s.e. 0.26 0.06 0.15 0.04 0.00 0.22 0.05 0.12 0.01 40 0.2	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53 2960 446 2.2	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.03 30 0.22 30 0.3	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61 2860 520 1.4	Is.e. 0.15 0.03 0.08 0.07 0.01 0.23 0.06 0.14 0.01 80 10 0.9	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78 89 711 1.1	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01 10 10 0.1	',-121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56 2550 447 1.6	D6'54 1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02 0.00 30 30 0.1	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53 3320 418 1.7	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.20 0.4 0.57 0.04 0.22 0.04 0.01 90 40 0.2
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ S ^a Cl H ₂ O ^b CO ₂ ^b	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 1650 921 1.3 430	Is.e. 0.24 0.06 0.53 0.30 0.01 0.51 0.02 0.24 0.03 0.01 30 0.01 30 0.1 34	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23 1740 929 0.8 155	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02 0.04 0.01 0.02 10 30 0.1 16	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22 1200 917 1.1 547	1s.e. 0.26 0.06 0.15 0.04 0.00 0.22 0.05 0.12 0.01 40 0.2 98	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53 2960 446 2.2 1724	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.03 20 30 0.3 259	CAB(ld IP 003 1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61 2860 520 1.4 387	Is.e. 0.15 0.03 0.08 0.07 0.01 0.23 0.06 0.14 0.04 0.01 9 48	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78 89 711 1.1 b.d.	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01 10 10 0.1 b.d.	',-121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56 2550 447 1.6 1233	D6'54 1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02 0.00 30 0.1 88	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53 3320 418 1.7 1200	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.22 0.04 0.21 156
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ S ^a Cl H ₂ O ^b CO ₂ ^b T (°C) ^c	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 0.21 1650 921 1.3 430 1138	Is.e. 0.24 0.06 0.53 0.01 0.51 0.02 0.24 0.03 0.01 30 0.1 34	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23 1740 929 0.8 155 1149	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02 0.04 0.01 0.02 10 30 0.1 16	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22 1200 917 1.1 547 1154	1s.e. 0.26 0.06 0.15 0.04 0.00 0.22 0.05 0.12 0.01 40 0.2 98	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53 2960 446 2.2 1724 1124	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.03 20 30 0.3 259	CAB(ld IP 003 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61 2860 520 1.4 387 1140	Is.e. 0.15 0.03 0.08 0.07 0.01 0.23 0.06 0.14 0.04 0.01 9 48	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78 89 711 1.1 b.d. 1137	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01 10 10 0.1 b.d.	',-121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56 2550 447 1.6 1233 1123	D6'54 1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02 0.00 30 30 0.1 88	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53 3320 418 1.7 1200 1137	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.21 0.02 0.02 0.03 0.04 0.21 156
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ S ^a Cl H ₂ O ^b CO ₂ ^b T (°C) ^c P(MPa) ^d	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 0.21 1650 921 1.3 430 1138 114	Is.e. 0.24 0.06 0.53 0.01 0.51 0.02 0.24 0.03 0.01 30 0.1 34	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23 1740 929 0.8 155 1149 44	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02 0.04 0.01 0.02 10 30 0.1 16	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22 1200 917 1.1 547 1154 133	1s.e. 0.26 0.06 0.15 0.04 0.05 0.12 0.01 40 40 98	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53 2960 446 2.2 1724 1124 374	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.03 0.02 30 0.3 259	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61 2860 520 1.4 387 1140 89	Is.e. 0.15 0.03 0.07 0.01 0.23 0.06 0.14 0.04 0.01 9.04 0.05	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78 89 711 1.1 b.d. 1137 82	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01 10 10 0.1 b.d.	',-121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56 2550 447 1.6 1233 1123 223	D6'54 1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02 0.00 30 30 0.1 88	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53 3320 418 1.7 1200 1137 238	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.21 0.02 0.04 0.22 0.04 0.01 90 40 0.2 156
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ S ^a Cl H ₂ O ^b CO ₂ ^b T (°C) ^c P(MPa) ^d	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 0.21 1650 921 1.3 430 1138 114 8.6	Is.e. 0.24 0.06 0.53 0.01 0.51 0.02 0.24 0.03 0.01 30 0.1 34	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23 1740 929 0.8 155 1149 44 86	1s.e. 0.06 0.21 0.19 0.01 0.18 0.02 0.04 0.01 0.02 10 30 0.1 16	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22 1200 917 1.1 547 1154 133 86	1s.e. 0.26 0.06 0.15 0.04 0.05 0.12 0.01 40 0.2 98	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53 2960 446 2.2 1724 1124 374 88 2.2	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.02 30 0.32 259	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61 2860 520 1.4 387 1140 89 89	Is.e. 0.15 0.03 0.08 0.07 0.01 0.23 0.06 0.14 0.04 0.01 80 10 0.9 48	m:47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78 89 711 1.1 b.d. 1137 82 88	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01 10 10 0.1 b.d.	P 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56 2550 447 1.6 1233 1123 223 87 0.12	D6'54 1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02 0.00 30 30 0.1 88	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53 3320 418 1.7 12000 1137 238 85	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.21 0.02 0.04 0.21 0.02 0.03 0.01 90 40 0.2 156
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ S ^a Cl H ₂ O ^b CO ₂ ^b T (°C) ^c P(MPa) ^d HostFo#	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 0.21 1650 921 1.3 430 1138 114 86 4.0	Is.e. 0.24 0.06 0.53 0.30 0.01 0.51 0.02 0.24 0.03 0.01 30 0.1 34	WC 020 2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23 1740 929 0.8 155 1149 44 86 8.2	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02 0.04 0.02 10 30 0.1 16	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22 1200 917 1.1 547 1154 133 86 3.5	1s.e. 0.26 0.06 0.15 0.04 0.02 0.05 0.12 0.01 40 0.2 98	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53 2960 446 2.2 1724 1124 374 88 3.0	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.03 20 30 0.32 259	CAB(ld IP 003_1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61 2860 520 1.4 387 1140 89 3.7	Is.e. 0.15 0.03 0.08 0.07 0.01 0.23 0.06 0.14 0.04 0.01 80 10 0.9 48	<i>m:47°5</i> IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78 89 711 1.1 b.d. 1137 82 88 5.2	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01 10 10 0.1 b.d.	',-121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56 2550 447 1.6 1233 1123 223 87 0.6	D6'54 1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02 0.00 30 30 0.1 88	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53 3320 418 1.7 1200 1137 238 88 5.1	1s.e. 0.34 0.029 0.24 0.01 0.57 0.04 0.22 0.04 0.22 10.4 0.57 0.04 0.22 0.04 0.22 156
Whitechi Sample SiO ₂ TiO ₂ Al ₂ O ₃ FeO ^T MnO MgO CaO Na ₂ O K ₂ O P ₂ O ₅ S ^a Cl H ₂ O ^b CO ₂ ^b T (°C) ^c P(MPa) ^d HostFo# %Ol fO2	uck LK WC 020_1 50.76 1.17 18.47 8.03 0.13 6.90 10.41 3.36 0.41 3.36 0.41 1650 921 1.3 430 1138 114 86 4.0 1.0	Is.e. 0.24 0.06 0.53 0.30 0.01 0.51 0.02 0.24 0.03 0.01 30 0.1 34	WC 020_2 50.31 1.13 18.97 8.01 0.12 6.77 10.22 3.65 0.45 0.23 1740 929 0.8 155 1149 44 86 8.2 0.8	1s.e. 0.06 0.05 0.21 0.19 0.01 0.18 0.02 0.04 0.01 0.02 10 30 0.1 16	WC 023 50.05 1.13 17.82 8.62 0.16 7.36 10.14 3.90 0.47 0.22 1200 917 1.1 547 1154 133 86 3.5 0.8	1s.e. 0.26 0.06 0.15 0.04 0.00 0.22 0.05 0.12 0.01 40 0.2 98	Indian IP 001 50.38 1.38 17.55 7.23 0.11 6.97 10.59 4.28 0.83 0.53 2960 446 2.2 1724 1124 374 88 3.0 1.4	Pass 1s.e. 0.27 0.11 0.08 0.01 0.34 0.05 0.14 0.03 0.02 20 30 0.32 259	CAB(ld IP 003 1 49.35 1.55 17.70 7.21 0.10 7.24 10.52 4.60 0.98 0.61 2860 520 1.4 387 1140 89 3.7 1.4	Is.e. 0.15 0.03 0.08 0.07 0.01 0.23 0.06 0.14 0.01 80 10 0.9 48	m: 47°5 IP 005 53.72 2.09 15.74 7.01 0.11 6.59 8.07 4.00 1.73 0.78 89 711 1.1 b.d. 1137 82 88 5.2 	9'58' 1s.e. 0.55 0.02 0.08 0.10 0.02 0.06 0.07 1.09 0.02 0.01 10 10 0.1 b.d.	',-121°(IP 016 50.19 1.49 17.88 7.01 0.12 6.37 10.80 4.52 0.91 0.56 2550 447 1.6 1233 1123 223 87 0.6 1.5	D6'54 1s.e. 0.23 0.05 0.19 0.07 0.00 0.06 0.04 0.16 0.02 0.00 30 0.1 88	") IP 020 50.10 1.42 17.52 7.22 0.10 7.16 10.91 4.11 0.79 0.53 3320 418 1.7 1200 1137 238 88 5.1 1.5	1s.e. 0.34 0.04 0.29 0.24 0.01 0.57 0.04 0.22 0.04 0.21 90 40 0.2 156

 Table 1: Major and Trace Element Data Corrected for Post-Entrapment Crystallization

Schreibe	r's Me	adow	CAB (l	ocati	on:48°0	59'62	",-121°	79'37	7")						
Sampla	SM		SM		SM		SM		SM		SM		SM		
Sample	001	ls.e.	005	ls.e.	025	1s.e.	027	ls.e.	029	ls.e.	030	1s.e.	032	ls.e.	
SiO ₂	52.74	0.14	51.85	0.28	49.28	0.59	51.74	0.48	52.65	0.83	57.12	0.42	52.99	0.43	
TiO ₂	1.26	0.02	1.29	0.05	0.67	0.04	1.33	0.08	1.82	0.04	1.93	0.08	1.22	0.04	
Al ₂ O ₃	17.56	0.20	18.94	0.09	18.75	0.64	18.73	0.41	16.53	0.42	17.14	0.06	18.79	0.36	
FeO ^T	8.73	0.16	6.51	0.14	10.37	0.14	7.47	0.08	9.06	0.27	6.48	0.15	6.49	0.06	
MnO	0.15	0.01	0.10	0.02	0.18	0.01	0.17	0.01	0.17	0.01	0.12	0.01	0.13	0.01	
MgO	5.11	0.02	5.70	0.13	6.25	0.39	4.63	0.07	5.06	0.47	3.35	0.08	4.96	0.03	
CaO	9.32	0.05	10.13	0.32	11.27	0.03	10.17	0.02	7.70	0.08	5.78	0.06	10.34	0.04	
Na ₂ O	4.41	0.10	4.17	0.16	2.47	0.09	4.51	0.21	4.99	0.11	5.17	0.26	3.75	0.47	
K ₂ O	0.35	0.02	0.78	0.17	0.40	0.01	0.79	0.02	1.34	0.05	2.10	0.02	0.83	0.01	
P_2O_5	0.20	0.01	0.41	0.01	0.16	0.01	0.33	0.00	0.52	0.01	0.69	0.01	0.38	0.01	
S ^a	983	10	1530	30	665	30	1480	20	1200	40	168	80	1480	40	
Cl	205	20	456	40	218	10	477	20	424	20	456	30	453	20	
H ₂ O ^b	0.97	0.10	2.27	0.15	2.05	0.20	1.82	0.22	1.34	0.09	0.76	0.08	1.70	0.10	
CO_2^{b}	167	18	870	57	276	27	624	76	358	23	b.d.		410	25	
T (°C) °	1105		1089		1100		1074		1094		1070		1082		
P(MPa) ^d	50		255		112		172		97		6		121		
HostFo#	£ 80		86		80		81		80		80		84		
%Ol	7.0		60		5.0		26		14		0.0		15		
add.	7.0		0.9		5.0		2.0		1.4		0.9		1.5		
fO2					1.0		1 1		0.0		07		1.0		
+ΔFMQ					1.0		1.1		0.9		0.7		1.0		
Mg#	56		65		56		57		54		52		62		

Table 1 (continued): Major and Trace Element Data Corrected for Post-Entrapment Crystallization

Major elements in wt. %, CO₂, S, Cl in ppm. Unless otherwise noted, analyses performed at University of Oregon CAMCOR lab. Fe reported as total FeO^{T a}Analysis performed at Washington State University ^bPerformed with FTIR at University of Oregon. ^c Calculated as per (Sugawara, 2000), corrected as per (Médard and Grove, 2008) ^d Entrapment pressure calculated with VolatileCalc (Newman and Lowenstern, 2002)

Whit	echuc	K																		
Sam ple	WC 002_ 1	1s.e.	WC 002_ 2	ls.e.	WC 003	ls.e.	WC 007	1s.e.	WC 008	1s.e.	WC 013	ls.e.	WC 017	ls.e.	WC 020_ 1	1s.e.	WC 020_ 2	ls.e.	WC 023	ls.e.
Ba	77	1	79	2	90	2	103	3	79	4	95	6	79	4	100	5	98	3	106	3
Ce	15.4	0.4	16.5	0.6	17.5	0.8	18.5	1.0	12.5	1.0	16.8	0.7	147	0.6	167	0.9	18.0	2.1	17.6	0.6
Dv	2.8	0.4	2.8	0.8	3.5	0.3	3.7	0.6	_	-	2.6	0.3	2.8	0.1	3.9	0.6	3.0	0.6	-	-
Er	2.0	0.2	1.9	0.2	1.1	0.2	1.4	0.0	3.0	0.8	1.5	0.3	1.4	0.2	1.3	0.3	2.2	0.3	2.1	0.3
Eu	1.0	0.1	1.0	0.3	0.9	0.1	1.4	0.1	_	-	1.0	0.1	0.9	0.1	0.7	0.2	1.1	0.1	1.1	0.1
Gd	2.9	0.7	3.4	0.6	3.0	0.3	4.1	0.8	-	-	3.0	0.3	2.3	0.2	3.4	0.4	2.5	0.6	2.7	0.5
Hf	1.9	0.2	1.7	0.2	1.8	0.2	1.4	0.1	-	-	1.4	0.2	1.7	0.2	3.4	0.9	3.8	0.5	2.0	0.2
La	5.1	0.3	5.1	0.4	4.9	0.2	6.0	0.5	-	-	5.9	0.3	4.6	0.1	5.4	0.6	6.0	0.5	5.9	0.2
Nb	1.1	0.1	1.0	0.0	1.1	0.1	1.0	0.1	-	-	1.1	0.1	1.1	0.1	1.1	0.2	1.8	0.1	1.4	0.1
Nd	10.9	0.4	10.7	0.3	11.4	0.8	16.0	1.3	5.4	1.0	11.0	0.5	10.8	1.0	10.7	0.7	10.7	1.6	11.0	0.8
Pb	2.9	0.3	2.0	0.3	2.4	0.2	3.4	0.1	3.7	0.6	1.8	0.3	2.2	0.3	3.6	0.6	3.7	0.5	3.1	0.2
Pr	2.8	0.2	2.5	0.1	2.5	0.2	2.7	0.1	1.9	0.6	2.2	0.1	2.2	0.1	2.5	0.2	2.4	0.2	2.8	0.2
Rb	3.6	0.2	2.2	0.2	4.2	0.2	3.7	0.2	-	-	3.7	0.2	2.6	0.4	4.9	0.5	5.0	0.7	4.7	0.6
Sc	31.3	1.9	30.3	0.6	32.5	1.8	32.0	1.4	84.0	2.1	29.0	1.0	23.2	1.7	40.1	4.5	41.5	4.3	29.6	1.5
Sm	2.6	0.3	2.5	0.2	3.2	0.3	3.8	0.5	-	-	-	-	-	-	-	0.9	3.4	0.7	2.7	0.3
Sr	540	9	534	17	530	13	533	18	457	35	553	16	535	14	539	16	549	11	514	9
Ta	-	-	0.1	0.0	-	-	-	-	-	-	-	-	-	-	-	-	0.2	0.0	-	-
Th	0.7	0.1	0.6	0.2	0.7	0.1	0.7	0.1			0.6	$\theta.\theta$	-	-	1.0	0.3	0.6	0.1	0.6	0.1
Ti ^a	6550	110	6480	110	6830	160	7380	250	5220	490	6390	160	6130	260	6470	160	7550	200	6710	130
U	0.1	0.0	-	-	0.2	0.1	0.3	0.1	-	-	0.1	0.0	0.2	0.0	0.3	0.0	0.2	0.1	0.2	0.0
V	209	7	217	8	249	9	232	11	266	22	221	9	211	15	246	12	252	9	211	10
Y	16.4	0.4	18.1	0.2	15.9	0.5	18.2	0.5	14.3	1.6	14.4	1.4	15.3	0.4	21.0	1.0	18.3	0.4	17.1	0.3
Yb	1.7	0.2	2.4	0.6	1.4	0.2	1.6	0.3	-	-	1.6	0.2	1.9	0.2	3.1	0.7	2.4	0.8	1.9	0.2
Zr	75	1.5	77	4.3	74	2.1	80	1.4	75	5.1	68	2.1	74	2.4	83	2.9	87	3.0	75	1.6

 Table 1 (continued): Major and Trace Element Data Corrected for Post-Entrapment Crystallization

 Whitechuck

All trace elements in ppm. Unless specified, all TE analyses performed with LA-ICP-MS and Oregon State University. ^a Calculated from TiO₂ analyzed with EPMA at University of Oregon.

India	in Pas	s CA	1 <i>B</i>		Schre	eiber	's Mee	adov	v CAB						
Sam ple	IP 1003_1	1s.e.	IP 016	1s.e.	SM 025	1s.e.	SM 027	1s.e.	SM 029	1s.e.	SM 030	1s.e.	SM 032	ls.e.	
Ba	416	7	344	9	201	12	258	15	382	19	549	23	247	10	
Ce	68.3	2.1	48.8	2.2	12.1	0.8	21.0	1.1	44.1	2.5	58.4	3.0	26.9	1.6	
Dy	3.0	0.3	3.1	0.4	2.1	0.3	1.8	0.2	7.5	0.6	5.6	1.0	1.5	0.5	
Er	1.2	0.5	2.7	0.5	0.7	0.2	-	-	4.0	0.3	4.3	0.3	2.0	0.7	
Eu	1.9	0.3	1.9	0.1	-	-	1.1	0.1	2.3	0.2	1.6	0.1	0.9	0.2	
Gd	4.0	0.9	4.0	0.4	-	-	4.6	1.2	5.9	0.6	5.2	0.7	1.7	0.2	
Hf	2.6	0.1	2.5	0.3	1.1	0.2	-	-	4.6	0.3	5.0	0.3	2.8	0.2	
La	23.3	1.4	20.8	0.7	4.0	0.3	8.4	0.6	19.8	0.8	23.4	1.1	10.8	0.6	
Nb	7.7	0.5	6.2	0.4	1.4	0.2	3.3	0.5	9.1	0.6	12.6	0.6	5.4	0.3	
Nd	30.2	1.6	24.3	1.2	8.2	0.6	12.8	1.1	24.9	1.2	30.9	1.2	16.1	1.5	
Pb	5.0	0.2	4.3	0.4	2.4	0.1	1.3	0.2	5.0	0.2	7.8	0.7	2.8	0.1	
Pr	7.7	0.3	6.2	0.3	1.0	0.1	1.6	0.2	6.6	0.5	7.2	0.4	2.6	0.2	
Rb	9.2	0.3	7.3	0.7	6.5	0.2	5.1	0.3	16.0	0.5	40.8	1.6	5.4	0.5	
Sc	24.8	2.2	24.9	1.0	28.7	1.6	31.9	1.5	32.1	2.4	20.2	2.1	26.5	1.0	
Sm	5.1	0.8	5.3	0.5	1.6	0.4	2.1	0.5	7.1	0.5	5.5	0.8	2.8	0.7	
Sr	1291	32	1191	37	753	33	509	24	470	19	388	17	526	20	
Та	0.4	0.1	0.3	0.0	-	-	2.0	0.7	0.5	0.1	1.1	0.1	-	-	
Th	2.9	0.4	2.0	0.1	-	-	-	-	2.3	0.1	5.0	0.7	-	-	
Ti ^a	8890	130	7810	210	3920	270	6980	470	1051 0	770	1200 0	800	6940	440	
U	0.9	0.1	0.6	0.1	0.3	0.1	-	-	0.9	0.1	1.4	0.2	-	-	
V	229	10	195	7	114	5	252	10	198	4	221	9	222	4	
Y	14.4	1.0	16.4	0.4	8.4	0.4	19.5	1.0	35.1	2.0	34.7	1.7	17.2	0.9	
Yb	0.9	0.2	1.2	0.1	1.2	0.4	2.2	0.5	1.8	0.6	2.0	0.4	-	-	
Zr	127	8.5	117	7.6	36	1.7	100	3.9	252	9.1	270	20.5	107	4.9	

Table 1 (continued): Major and Trace Element Data Corrected for Post-Entrapment Crystallization

All trace elements in ppm. Unless specified, all TE analyses performed with LA-ICP-MS and Oregon State University. ^a Calculated from TiO_2 analyzed with EPMA at University of Oregon. 1 s.e. (standard error)

Table 2: Host Olivine Compositions

Whitech	uck LK	T																
Sample	WC 002	1 s.e.	WC 003	1 s.e.	WC 007	1 s.e.	WC 008	1 s.e.	WC 013	1 s.e.	WC 014	1 s.e.	WC 017	1 s.e.	WC 020	1 s.e.	WC 023	1 s.e.
SiO ₂	40.22	0.20	39.11	0.12	40.04	0.10	39.64	0.21	40.39	0.13	40.34	0.05	40.19	0.12	40.30	0.21	40.36	0.20
MgO	45.40	0.11	45.47	0.15	44.73	0.05	45.08	0.10	45.09	0.08	45.06	0.12	44.98	0.22	45.49	0.32	45.60	0.09
FeO	13.63	0.18	13.69	0.21	14.40	0.20	13.61	0.10	13.50	0.35	13.54	0.20	13.37	0.19	13.69	0.18	13.45	0.29
CaO	0.17	0.00	0.17	0.00	0.17	0.00	0.17	0.00	0.18	0.00	0.17	0.00	0.17	0.00	0.17	0.00	0.17	0.00
MnO	0.20	0.00	0.21	0.00	0.22	0.01	0.21	0.01	0.20	0.01	0.20	0.00	0.20	0.00	0.20	0.01	0.20	0.01
Total	99.63		98.65		99.56		98.71		99.35		99.32		98.90		99.84		99.78	
Fo#	86		86		86		85		86		86		86		86		86	

Indian Pass CAB

Sample	IP 001	1 s.e.	IP 003	1 s.e.	IP 005	1 s.e.	IP 016	1 s.e.	IP 020	1 s.e.
SiO ₂	40.51	0.11	40.64	0.04	40.23	0.23	40.54	0.21	40.60	0.28
MgO	47.07	0.06	47.39	0.08	46.03	0.02	46.57	0.04	46.83	0.13
FeO	11.50	0.13	10.76	0.13	12.24	0.06	11.90	0.15	11.14	0.05
CaO	0.14	0.00	0.14	0.00	0.17	0.01	0.15	0.00	0.15	0.00
MnO	0.16	0.01	0.15	0.01	0.18	0.01	0.17	0.00	0.16	0.01
Total	99.68		99.47		99.10		99.62		99.18	
Fo#	88		89		88		87		88	

Schreiber's Meadow CAB

Sample	SM 001	1 s.e.	SM 005	1 s.e.	SM 025	1 s.e.	SM 027	1 s.e.	SM 029	1 s.e.	SM 030	1 s.e.	SM 032	1 s.e.	
SiO ₂	39.82	0.28	40.39	0.30	38.60	0.15	38.98	0.14	38.22	0.30	39.07	0.15	39.77	0.25	
MgO	41.67	0.25	45.69	0.07	40.66	0.56	42.16	0.03	41.21	0.49	40.88	0.27	44.00	0.50	
FeO	18.51	0.14	12.95	0.19	18.40	0.17	17.12	0.48	18.25	0.70	18.45	0.23	14.89	0.41	
CaO	0.18	0.00	0.17	0.00	0.23	0.00	0.19	0.01	0.18	0.01	0.20	0.01	0.17	0.00	
MnO	0.31	0.01	0.22	0.01	0.30	0.00	0.28	0.01	0.31	0.02	0.31	0.01	0.23	0.02	
Total	100.59		99.60		98.20		98.86		98.29		98.99		99.25		
Fo#	80		86		80		81		80		80		84		

All values reported in wt %. Olivine compositions analyzed with EPMA at University of Oregon

Sampla	WC											
Sample	002_1	002_2	003	007	008	013	014	017	020_1	020_2	023	Avg
SiO ₂	48.74	48.60	48.01	49.14	48.61	48.21	48.81	48.55	48.64	48.22	47.94	48.50
TiO ₂	1.00	0.97	0.97	1.11	0.95	0.98	0.97	1.00	1.01	0.97	0.96	0.99
Al_2O_3	16.07	16.17	16.55	16.05	15.70	16.24	15.79	16.01	15.94	16.32	15.22	16.01
FeO	7.14	7.16	7.16	6.91	7.22	7.23	7.20	7.19	7.22	7.22	7.69	7.21
Fe ₂ O ₃	1.22	1.22	1.21	1.14	1.23	1.23	1.22	1.23	1.23	1.22	1.31	1.23
MnO	0.13	0.13	0.12	0.12	0.11	0.11	0.12	0.12	0.11	0.10	0.14	0.12
MgO	11.52	11.56	11.54	11.16	11.65	11.66	11.62	11.61	11.65	11.65	12.40	11.64
CaO	8.92	8.92	9.08	8.98	9.21	9.04	8.73	8.98	8.99	8.79	8.66	8.94
Na ₂ O	2.96	2.97	3.01	3.01	2.98	2.96	3.21	3.00	2.90	3.14	3.33	3.04
K ₂ O	0.34	0.35	0.38	0.40	0.37	0.37	0.37	0.35	0.36	0.39	0.40	0.37
P_2O_5	0.19	0.18	0.19	0.20	0.19	0.18	0.18	0.17	0.18	0.20	0.18	0.18
H_2O	1.70	1.70	1.70	1.70	1.70	1.70	1.70	1.70	1.70	1.70	1.70	1.70
%Ol add	13.6	13.7	14.0	14.7	14.0	13.9	14.0	13.6	13.8	14.2	15.0	14.0
Ba	67	68	78	88	68	82		68	86	85	90	78
Ce	13.3	14.3	15.1	15.8	10.7	14.4		12.7	14.5	15.4	15.0	14.1
Cl (ppm)	798	799	817	830	810	780	870	792	794	797	780	806
Dy	2.4	2.4	3.0	3.2		2.3		2.4	3.3	2.6	1.8	2.6
Er	1.7	1.7	1.0	1.2	2.6	1.3		1.2	1.1	1.9	1.8	1.5
Eu	0.9	0.9	0.8	1.2	0.5	0.9		0.8	0.6	0.9	1.0	0.8
Gd	2.5	3.0	2.6	3.5	3.4	2.6		2.0	2.9	2.1	2.3	2.7
Hf	1.6	1.5	1.6	1.2		1.2		1.5	3.0	3.3	1.7	1.8
La	4.4	4.4	4.3	5.2	4.2	5.1		4.0	4.7	5.2	5.0	4.6
Nb	0.9	0.9	1.0	0.8	1.2	1.0		1.0	1.0	1.5	1.2	1.0
Nd	9.4	9.3	9.8	13.7	4.7	9.5		9.3	9.2	9.2	9.4	9.3
Pb	2.5	1.7	2.0	2.9	3.2	1.6		1.9	3.1	3.2	2.6	2.5
Pr	2.5	2.2	2.1	2.3	1.6	1.9		1.9	2.2	2.1	2.4	2.1
Rb	3.1	1.9	3.6	3.2	2.4	3.2		2.3	4.2	4.3	4.0	3.2
S (ppm)	1270	1280	1380	1380	1370	1470	1460	1140	1420	1500	1040	1340
Sc	27.8	26.8	28.7	28.1	74.3	25.7		20.6	35.5	36.6	25.9	33.0
Sm	2.2	2.1	2.7	3.3		1.5		1.4	0.4	3.0	2.3	2.1
Sr	467	462	457	456	393	477		463	466	472	438	455
Та	0.0	0.1		0.2	0.3	0.0				0.2		0.1
Th	0.6	0.6	0.6	0.6		0.5		0.3	0.9	0.5	0.5	0.6
U	0.1		0.2	0.2		0.1		0.2	0.2	0.2	0.2	0.2
V	182	189	216	200	231	192		184	214	219	181	201
Y	14.2	15.6	13.7	15.5	12.3	12.4		13.2	18.1	15.7	14.6	14.5
Yb	1.5	2.1	1.2	1.4	3.9	1.4		1.7	2.7	2.1	1.6	1.9
Zr	65	66	64	68	65	58		64	71	75	64	67

Table 3: Compositions corrected to mantle equilibrium (Fo₉₀) (Inclusion Mg#: ~73) and normalized to 100

Sample	IP 001	IP 003 1	IP 005	IP 016	IP 020	IP-Avg
SiO ₂	48.51	47.67	51.44	48.24	48.31	48.83
TiO ₂	1.23	1.39	1.86	1.31	1.27	1.41
Al ₂ O ₃	15.59	15.89	13.95	15.72	15.67	15.36
FeO	6.09	6.02	5.95	6.00	6.05	6.02
Fe ₂ O ₃	1.42	1.43	1.37	1.36	1.43	1.40
MnO	0.10	0.09	0.10	0.10	0.09	0.10
MgO	10.67	10.54	10.43	10.52	10.61	10.56
CaO	9.41	9.44	7.15	9.49	9.76	9.05
Na ₂ O	3.80	4.13	3.55	3.98	3.67	3.82
K ₂ O	0.74	0.88	1.54	0.80	0.71	0.93
P_2O_5	0.47	0.54	0.69	0.50	0.47	0.53
H ₂ O	1.91	1.91	1.91	1.91	1.91	1.91
%Ol add	6.5	5.5	6.9	7.7	5.9	6.5
Ba		393		318		356
Ce		64.6		45.1		54.8
Cl (ppm)	417	492	662	413	393	475
Dy		2.9		2.8		2.9
Er		1.2		2.5		1.8
Eu		1.8		1.8		1.8
Gd		3.8		3.7		3.7
Hf		2.4		2.3		2.4
La		22.0		19.2		20.6
Nb		7.3		5.7		6.5
Nd		28.6		22.4		25.5
Pb		4.8		3.9		4.4
Pr		7.3		5.7		6.5
Rb		8.7		6.8		7.8
S (ppm)	2560	2470	80	2200	2870	2030
Sc		23.6		23.3		23.5
Sm		4.9		4.9		4.9
Sr		1220		1100		1160
Ta		0.4		0.3		0.3
Th		2.7		1.8		2.3
U		0.9		0.6		0.7
V		217		181		199
Y		13.6		15.2		14.4
Yb		0.8		1.1		1.0
Zr		120		108		114

Table 3 (continued): Compositions corrected to equilibrium with mantle (Fo₉₀) and normalized to 100

$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	WC 023 48.98	VC WC W 13 014 01	WC 013	WC 000	WC	wa	ſ	uck LKT	Whitechu
WC WC	WC 023 48.98	VC WC W 13 014 01	WC 013	WG 000	WC	ILIC			
Comple 002_1 002_2 003 007 WC_008 013 014 017 020_1 020_2 0 SiO2 49.02 48.68 48.81 50.01 48.42 48.52 49.79 48.55 49.30 48.69 48.69	48.98	13 014 01	013		WC	WC	WC	WC	Samula
SiO ₂ 49.02 48.68 48.81 50.01 48.42 48.52 49.79 48.55 49.30 48.69 48	48.98		015	WC_008	007	003	002_2	002_1	Sample
	1 1 2	.52 49.79 48	48.52	48.42	50.01	48.81	48.68	49.02	SiO ₂
FiO ₂ 1.12 1.08 1.11 1.30 1.07 1.13 1.14 1.11 1.17 1.16 1	1.13	13 1.14 1.	1.13	1.07	1.30	1.11	1.08	1.12	TiO ₂
$\mathbf{M_2O_3}$ 17.97 17.98 18.84 18.88 17.65 18.70 18.54 17.80 18.50 19.53 1	17.93	.70 18.54 17	18.70	17.65	18.88	18.84	17.98	17.97	Al ₂ O ₃
$\mathbf{FeO}^{\mathrm{T}}$ 7.76 7.82 7.75 7.30 7.71 7.59 7.46 7.57 7.36 6.56 8	8.64	59 7.46 7.	7.59	7.71	7.30	7.75	7.82	7.76	FeO ^T
Ano 0.14 0.14 0.13 0.14 0.13 0.13 0.14 0.14 0.13 0.13 0	0.16	13 0.14 0.	0.13	0.13	0.14	0.13	0.14	0.14	MnO
MgO 6.34 6.43 6.22 3.83 5.67 5.18 4.83 6.55 5.22 3.74 5	5.59	18 4.83 6.	5.18	5.67	3.83	6.22	6.43	6.34	MgO
CaO 9.98 9.91 10.34 10.56 10.35 10.41 10.25 9.99 10.43 10.52 10	10.20	.41 10.25 9.	10.41	10.35	10.56	10.34	9.91	9.98	CaO
Na ₂ O 3.31 3.30 3.43 3.54 3.36 3.41 3.77 3.34 3.36 3.76 3	3.92	41 3.77 3.	3.41	3.36	3.54	3.43	3.30	3.31	Na ₂ O
K_2O 0.38 0.39 0.43 0.47 0.42 0.43 0.44 0.39 0.42 0.46 0	0.47	43 0.44 0.1	0.43	0.42	0.47	0.43	0.39	0.38	K ₂ O
P_2O_5 0.21 0.20 0.21 0.23 0.21 0.21 0.21 0.19 0.21 0.24 0	0.22	21 0.21 0.	0.21	0.21	0.23	0.21	0.20	0.21	P_2O_5
S 0.15 0.15 0.16 0.16 0.16 0.17 0.17 0.13 0.17 0.18 0	0.12	17 0.17 0.	0.17	0.16	0.16	0.16	0.15	0.15	S
CI 0.09 0.09 0.10 0.10 0.10 0.09 0.11 0.09 0.10 0.10	0.10	09 0.11 0.	0.09	0.10	0.10	0.10	0.09	0.09	Cl
$\mathbf{H}_2\mathbf{O}^c$ 1.5 1.8 0.5 2.1 2.0 1.2 0.6 1.9 1.3 0.9	1.2	.2 0.6 1	1.2	2.0	2.1	0.5	1.8	1.5	H ₂ O ^c
CO_2^{c} 620 820 330 680 770 250 290 510 450 170 5	570	50 290 5	250	770	680	330	820	620	CO ₂ ^c
Total 96.46 96.20 97.57 96.55 95.25 96.04 96.91 95.90 96.37 95.11 9	97.54	.04 96.91 95	96.04	95.25	96.55	97.57	96.20	96.46	Total

Appendix A: Uncorrected Major and Trace Element Data

Indian P	ass CA	В				Schreib	er's Me	adow C.	AB				
Sample	IP	IP	IP	IP	IP	SM	SM	SM	SM	SM	SM	SM	SM
Sample	001	003_1	005	016	020	001	005	007	025	027	029	030	032
SiO ₂	47.94	47.66	52.77	47.59	48.10	52.60	51.03	51.54	47.59	48.58	51.14	57.11	50.65
TiO ₂	1.35	1.54	2.14	1.41	1.42	1.32	1.34	1.27	0.68	1.22	1.78	1.94	1.16
Al ₂ O ₃	17.09	17.61	16.06	16.87	17.51	18.42	19.63	18.48	18.83	17.23	16.22	17.24	17.76
FeO ^T	6.40	6.13	6.69	6.92	6.26	7.55	5.31	6.07	9.69	8.29	8.70	6.18	6.33
MnO	0.11	0.10	0.12	0.11	0.10	0.16	0.10	0.12	0.18	0.16	0.17	0.12	0.12
MgO	5.61	5.86	4.26	6.16	5.07	2.63	3.10	3.65	4.14	4.79	4.36	3.07	5.34
CaO	10.32	10.47	8.23	10.19	10.90	9.77	10.49	10.34	11.32	9.35	7.55	5.82	9.77
Na ₂ O	4.17	4.58	4.08	4.27	4.10	4.63	4.33	4.27	2.48	4.15	4.90	5.20	3.54
K ₂ O	0.81	0.98	1.77	0.85	0.79	0.37	0.81	0.78	0.40	0.73	1.32	2.12	0.78
P_2O_5	0.51	0.60	0.80	0.53	0.52	0.21	0.43	0.39	0.16	0.30	0.51	0.70	0.36
S^{a}	0.30	0.29	0.01	0.26	0.33	0.10	0.16	0.16	0.07	0.15	0.12	0.02	0.15
Cl	0.05	0.05	0.08	0.05	0.04	0.02	0.05	0.05	0.02	0.05	0.04	0.05	0.05
H_2O^b	2.2	1.5	2.9	1.6	1.8	1.0	2.4	1.4	2.2	1.9	1.4	0.8	1.7
CO ₂ ^b	1380	310	b.d.	970	1000	180	940	360	290	640	360	b.d.	420
Total	94.74	100.00	97.15	95.32	100.00	100.00	96.85	97.22	95.62	95.03	96.88	99.65	96.03

Major elements in wt. %, CO₂ in ppm. Unless otherwise noted, analyses performed at University of Oregon CAMCOR lab. Fe reported as total FeO^T ^aAnalysis performed at Washington State University ^bPerformed with FTIR at University of Oregon.

Appendix A (continued): Uncorrected Major and Trace Element Data Whitechuck LKT

millena										
Sample	WC_002	WC_002_	<u>2 WC_003</u>	WC_007	' WC_008	WC_013	WC_017	WC_020_	<u>1 WC_020</u>	_2WC_023
Ba	77	79	91	108	81	98	79	104	107	110
Ce	15	17	18	19	13	17	15	17	20	18
Dv	2.8	2.8	3.5	3.9	-	2.7	2.8	4.0	3.2	-
Er	2.0	19	11	14	31	1.5	14	13	2.4	2.2
En	1.0	1.0	0.9	1.5	-	1.1	0.9	0.7	1.2	1.2
Cd	2.9	3.4	3.0	43	_	3 1	23	3.5	27	2.8
UU Uf	2.9	17	1.8	1.5		1.4	2.5	3.5	4.2	2.0
III La	5 1	5 1	1.0	6.2	-	1. 4 6.1	1.7	5.0	4.2	2.0
La	3.1 1.1	J.1 1.0	3.0	0.5	-	0.1	4.0	5.0 1.2	0.0	0.1
ND	1.1	1.0	1.1	1.0	-	1.2	1.1	1.2	1.9	1.5
Nd	11	11	11	1/	6	11	11	11	12	11.4
Pb	2.9	2.0	2.4	3.5	3.8	1.9	2.2	3.8	4.1	3.2
Pr	2.9	2.5	2.5	2.8	1.9	2.3	2.2	2.6	2.6	2.9
Rb	3.6	2.2	4.2	3.9	-	3.9	2.6	5.1	5.4	4.9
Sc	31	30	33	33	85	30	23	42	44	30
Sm	2.6	2.5	3.2	4.0	-	-	-	-	3.7	2.8
Sr	542	535	534	559	465	574	536	561	596	532
Та	-	0.1	-	-	-	-	-	-	0.2	-
Th	0.7	0.6	0.7	0.8	-	0.6	-	1.0	0.7	0.7
U	0.1	-	0.2	0.3	-	0.2	0.2	0.3	0.2	0.2
v	209	217	251	243	270	229	212	255	273	218
v	16	18	16	19	15	15	15	22	20	18
Vh	17	24	14	1.6	-	16	19	3 2	26	19
10 7r	75	2. 4 76	75	1.0 84	77	70	74	9.2 86	2.0	78
LA Indian De	73	70	75 Sahwaihar	04 Va Maada	CAP	70	/4	80	95	78
		ID 002 1	Schreider	S Meduc	W CAD	GM 020	GM 022			
Sample	IP_010	<u>IP_003_1</u>	<u>SM_025</u>	SIVI_027	SM_029	SM_030	SM_032			
Ba	346	432	211	265	387	554	251			
Ce	49	71	13	22	45	59	27			
Dy	3.1	3.1	2.2	1.8	7.6	5.7	1.5			
Er	2.7	1.3	0.8	-	4.0	4.3	2.0			
Eu	1.9	2.0	-	1.1	2.3	1.6	0.9			
Gd	4.0	4.2	-	4.8	6.0	5.3	1.7			
Hf	2.5	2.7	1.2	-	4.7	5.0	2.8			
La	21	24	4	9	20	24	11			
Nb	6.2	8.0	1.5	3.4	9.2	12.8	5.5			
Nd	24	31	9	13	25	31	16			
Pb	4.3	5.2	2.5	1.3	5.1	7.8	2.9			
Pr	6.2	8.0	1.0	1.6	6.7	7.3	2.7			
Rb	7	10	7	5	16	41	5			
Sc	25	26	30	33	32	20	27			
Sm	53	53	17	21	7 2	56	2.8			
Sn	1200	1340	790	520	480	390	530			
Te	0.3	0.5	170	2.0	400 0.5	11	550			
та Ть	20	2.0	-	2.0	2.3	1.1 5 1	-			
I II T	2.0	3.0	-	-	2.4 1.0	J.I 1 4	-			
U	0.0	0.9	0.5	-	1.0	1.4	-			
V	19/	257	120	239	201	225	225			
Y	17	15	9	20	36	35	17			
Yb	1.2	0.9	1.2	2.3	1.8	2.0	-			
Zr	117	132	38	102	256	274	108			

All trace elements in ppm. Analyses performed with LA-ICP-MS at Oregon State University

Appendix B: Extended Methods

Sample Preparation

Tephra samples were collected from as far down within the cinder cone deposits as possible, though not quite at the basal layer. Tephra size was limited to ≤ 1 cm, though most sample sets were smaller (3mm-0.1mm) for ease of separation. Sample sets were dried and sieved, to isolate tephra between 0.25 and 1.25 mm. Olivine and other dense minerals were isolated from matrix using heavy-liquid density separation techniques (lithium tetrabromoethane).

Once separated, olivine crystals were inspected under binocular light microscope and refractive index oil (refractive index of 1.65) to determine those most likely to contain appropriate melt inclusions. Inclusions appropriate to this study were primary inclusions (round or oval-shaped and not associated with olivine cracks), glassy (no evidence of secondary mineral growth), and larger than 50 microns. These inclusions were typically very light brown, completely transparent, and often contained a clearly visible shrinkage bubble.

Olivine crystals were individually mounted on round, 1 in diameter glass petrographic slides, using Crystalbond adhesive (a common thermoplastic mounting matrix) melted onto each slide on a hotplate. The smallest possible amount of Crystalbond was used that left enough surface area outside of the crystal to maintain a flat polishing surface. Too little Crystalbond would leave the crystal sitting above the thermoplastic where it would become rounded. Once the crystals were mounted, each one was lightly polished using 600 grit wet/dry sandpaper (used in automobile detailing). After this initial polish (3-4 passes), the crystal was inspected under light microscope (with a small drop of refractive index oil, if necessary) to establish the location and approximate depth of a good inclusion. Once identified, the olivine hosts were slowly polished by hand with 600 grit sandpaper until each inclusion was intersected. Intersected surfaces showed up as small, inner clear areas within inclusions as the light was turned down to its lowest level. Many inclusions were lost while gaining the experience to see this intersection.

Mounted crystals were then washed carefully in cold water (warm water may soften the thermoplastic) using dishwashing soap, to remove any embedded grit. This was important to avoid cross-contaminating the lap wheel with different grit sizes. Each crystal was polished for 3-5 minutes on a lap wheel with 6-micron diamond paste. Small amounts of olive oil (3-6 drops) were used to lubricate the wheel and to avoid removing the crystal from the thermoplastic accidently. While polishing a large set of inclusions, frequent observation under reflected light was done to make sure enough polishing compound remained on the cloth and that no cross-contamination of grit sizes from the previous step had occurred (noted by grooves within the crystal or the thermoplastic). Once polished, the mounted crystals were then washed again to remove any traces of the 6-micron paste and dried. Mounted crystals were then polished with 1-micron diamond paste for 1-3 minutes, inspecting them frequently under reflected light.

Mounted and polished crystals were then heated on a hotplate situated beneath a light microscope until the thermoplastic melted slightly. Using tweezers, each crystal was

flipped onto the polished side and pressed down gently to ensure that there was little thermoplastic beneath the crystal. This step was vital; calculation of the inclusion composition is dependent upon the wafer thickness, measured initially on the mounted host. After flipping, each inclusion was intersected from the other side of the crystal, using 600 grit sandpaper, then 6-micron and 1-micron paste. Olivine host crystals were easiest to deal with (and not break) in further steps if they were polished no thinner than 60 microns, and preferably closer to 100 microns. Very small, or thin, inclusions were not totally intersected on this side, and the extra olivine was subtracted out mathematically post-analysis. This required, however, that when these particular host-crystals were mounted for microprobe analysis that the intersected side was positioned face up to remain exposed. Further polishing (and washing between steps) was carried out as before.

Doubly polished crystals were then photographed (as much of the total crystal as possible), with the photographs immediately printed out and stored in a binder. This allowed for easy identification during later analysis, especially for crystals with multiple inclusions.

Before removing olivine-hosts from their mounts, each crystal thickness was measured with a digital micrometer. Slide thickness was measured in three different spots, averaged, and subtracted from the thickness of the combined crystal and slide.

Preparation for FTIR, EPMA and LA-ICPMS analysis

Each olivine-host was removed from the Crystalbond thermoplastic by setting it upside down in a watchglass filled with acetone. After approximately 30 seconds the crystals dropped from the slide into the watch-glass. Any remaining thermoplastic was dissolved from the crystal by refreshing with new acetone in at least three cycles of pipetting out and refilling. Once washed, the crystal was transferred to a slide holding a 13x2mm, round NaCl disc for FTIR analysis.

After analysis by FTIR, the crystals were mounted for microprobe analysis. Using double-sided tape mounted on a sheet of aluminum, locations were marked using a round 2.5 cm diameter, 1 cm thick plastic mount with 7 holes drilled into it by the machine shop at the University of Oregon. The mount was pressed down upon the tape, leaving an impression where each hole was located. Within each impression, between 1 and 5 crystals were placed (intersected side down, if not doubly intersected) on the tape. As each crystal was placed, its location was recorded on a map of the mount. Once all of the crystals were placed, the mount was set back over them.

Each hole in the mount was gently filled with a Cl-free epoxy and set aside to harden for 12 hours. After this initial hardening, the mount was placed in a 60°C oven for 24 hours to finish curing. After curing, the mount was removed and polished, using 15 micron, 6 micron and 1 micron diamond paste, with careful washing between each step. This mount was used for both EPMA and LA-ICPMS analysis.

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	Olivine	Lherzolite bulk partition	Lherzolite bulk
	partition	coefficient (Workman &	partition coefficient
	coefficients	Hart, 2005)	(Kelley et al. 2006)
Ba	0.0099	0.00012	0.00012
Ce	0.006	0.022	0.011
Cl	0.0013	0.003	0.003
CO ₂	0		
Dy	0.013	0.079	0.06
Er	0.0256	0.097	0.08
Eu	0.0074	0.05	0.035
Gd	0.01	0.056	0.04
H_20	0.001	0.012	0.012
K ₂ O		0.003	0.003
Hf	0.013		
La	0.0067	0.01	0.0075
Nb	0.01	0.0034	0.003
Nd	0.0059	0.031	0.02
Pb	0.0001		
Pr	0.006		
Rb	0.0098		
S	0		
Sc	0.17		
Sm	0.0066	0.045	0.03
Sr	0.014	0.025	0.011
Ta	0.03		
Th	0.0001	0.001	0.0015
Ti	0.02	0.058	0.04
U	0.002	0.0011	0.0011
V	0.06		
Y	0.01	0.088	0.095
Yb	0.0491	0.115	0.095
Zr	0.012	0.033	0.03

Appendix C: Partition Coefficients





MM, Meager Mountain; MC, Mount Cayley; MG, Mount Garibaldi; MB, Mount Baker; GP, Glacier Peak; MR, Mount Rainier; MSH, Mount St. Helens; MA, Mount Adams; SVF, Simcoe Volcanic Field; MH, Mount Hood; MJ, Mount Jefferson; TS, Three Sisters; NV, Newberry Volcano; CLV, Crater Lake Volcano; MMc, Mount McLoughlin; MLV, Medicine Lake Volcano; MS, Mount Shasta; LVC, Lassen Volcanic Center



Figure 2: Grove et al. (2002), proposed this model of magma generation at Mount Shasta—accounting for low potassium tholeiite (LKT) magmas (called here HAOT—high alumina olivine tholeiite) through decompression melting in the mantle wedge and calc-alkaline basalt and basaltic-andesite (CAB) magmas (called here BA) and another sub-type (primitive magnesian andesite-PMA) through fluid-flux melting of a depleted mantle that is descending along the downgoing slab. Figure from (Grove et al., 2002).



Figure 3: Typical melt inclusions used in this study. Inclusions are glassy, round or oval, and most contain a vapor bubble. Scale bars are 50 microns.



Figure 4 : Uncorrected inclusion data (grey diamonds) corrected, if necessary, for iron-loss (red, green and blue symbols). Circled inclusions (grey) are those that required this correction. Two inclusions from Schreiber's Meadow did not fall within the bounds of bulk rock compositions despite corrections (shown). Bulk rock data for Glacier Peak from DeBari et al., submitted. Bulk rock data for Mt. Baker from Moore and DeBari, submitted.



Figure 5 : K_2O vs. SiO₂ for corrected inclusion compositions (filled symbols) compared to Cascade arc primitive magmas. Arc LKTs and CABs are compiled from the GeoRoc database (Bacon, 1997; Conrey et al., 1997; Green and Sinha, 2005; Grove et al., 2002; Hildreth et al., 2003; Le Voyer et al., 2010; Leeman et al., 2005a; Rowe et al., 2009; Ruscitto et al., 2010; Sisson and Layne, 1993). Fields are from Gill (1981).



Figure 6: SiO_2 vs. major element for corrected Glacier Peak melt inclusions (red and blue symbols). Whole rock and dacite analyses of DeBari and et al. (submitted) shown for comparison. Inclusion IP_005 does not cluster with other compositions. Error bars are shown or are smaller than symbols. Oxides in wt. %.


Figure 7: SiO_2 vs. major element for corrected Mt. Baker melt inclusions (green symbols). Associated Sulphur Creek lava flow (Green, 1988; Hildreth et al., 2003, Moore and DeBari, submitted) and dacite analyses (Baggerman and DeBari, in press) shown for comparison. Error bars are shown or are smaller than symbols. Oxides are in wt. %.



Figure 8: Trace element comparisons of Whitechuck (LKT) and Indian Pass (CAB) inclusion compositions with whole rock and dacite analyses of DeBari and Taylor (submitted) shown for comparison. Error bars are shown or are smaller than symbols.



Figure 9: Trace element comparisons of Schreiber's Meadow (CAB) inclusion compositions with whole rock analyses from the associated Sulphur Creek lava flow (Green, 1988; Hildreth et al., 2003, Moore and DeBari, submitted) and dacite analyses (Baggerman and DeBari, in press) shown for comparison. Error bars are shown or are smaller than symbols.



Figure 10: Chondrite-normalized rare earth element patterns for melt inclusions . (a) Schreiber's Meadow melt inclusions and field showing range of associated lava flow whole rock compositions. (b) Indian Pass CAB melt inclusions and whole rock field. (c) Whitechuck melt inclusions and whole rock field. (d) Comparison of average melt inclusion compositions with NMORB and EMORB. Whole rock data from DeBari et al. (submitted) [Glacier Peak] and Moore and DeBari (submitted) [Mount Baker]. EMORB and NMORB compositions from Sun and McDonough (1989).



Figure 11: NMORB-normalized trace element patterns for corrected melt inclusions. (a) Schreiber's Meadow melt inclusions and field showing range of associated lava flow whole rock compositions. (b) Indian Pass CAB melt inclusions and whole rock field. (c) Whitechuck melt inclusions and whole rock field. (d) Comparison of average melt inclusion compositions with NMORB and EMORB. Whole rock data from DeBari et al. (submitted) [Glacier Peak] and Moore and DeBari (submitted) [Mount Baker]. EMORB and NMORB compositions from Sun and McDonough (1989).



Figure 12: Common H_2O proxies (Ba/Zr and Pb/Ce) for corrected inclusions (filled symbols) and whole rock (open symbols). Fields for Cascade LKT shown for comparison. Data sources: Whitechuck whole rock (DeBari et al., submitted) Medicine Lake, Lassen and Crater Lake (Bacon, 1997), Central Oregon (Rowe et al., 2009).



Figure 13: CO_2 vs. H_2O . Isobars showing minimum pressures at which inclusions were entrapped in crystallizing olivine hosts. Pressures calculated with VolatileCalc (Newman and Lowenstern, 2002). Error bars are shown or are smaller than symbols. Sample IP_005 (Glacier Peak CAB) is not shown, but has H_2O concentration of 1.05 wt. % and CO_2 concentrations below detection.



Figure 14: H₂O vs. alkalis and MgO. Error bars are shown or are smaller than symbols.



Figure 15: Volatiles vs. K₂O (using K₂O as a proxy for increasing fractionation) with corrected inclusion compositions. Volatiles for Whitechuck (red circles) have wide-ranging variability at similar degrees of fractionation, whereas Indian Pass and Schreiber's Meadow inclusions show decreasing volatile concentrations with increasing fractionation. Error bars are shown or are smaller than symbols.



Figure 16: Pressure and temperature of last equilibration of melts with the mantle calculated with the geobarometer and geothermometer of Lee et al. (2009), using inclusion compositions corrected to mantle equilibrium. Entrapment pressures calculated with VolatileCalc (Newman and Lowenstern, 2002), temperature calculated with Sugawara, 2000), both using inclusion compositions prior to correction to mantle values. For comparison, the High-Mg Basaltic Andesite (HMBA) Tarn Plateau primitive magma from Mount Baker (Moore and DeBari, submitted) is shown in purple. Slab depth is between 80-100 km (McCrory et al, 2004). Moho depth is between 40-45 km (Mooney and Weaver, 1989).



Figure 17: Nb vs.Y batch melting curves for four potential sources (diamonds and green and black stars), plotted in increments of 0.01 melt fraction (F), from 0.01 to 0.3. Trace elements from melt inclusions, corrected to mantle values (red circles and blue squares). Whole rock fields (dashed lines) show for comparison. NMORB and EMORB from Sun and McDonough (1989). E-DMM from Workman and Hart (2005). DMM from Salters and Stracke (2004). OIB sources from Ruscitto et. al (2010b). Bulk rock data from DeBari et al. (submitted).



Figure 18: EMORB-normalized trace elements from average corrected Indian Pass melt inclusions (blue squares) and from whole rock concentrations (white squares) of DeBari et al. (submitted). EMORB (white stars) from Sun and McDonough (1989).



Figure 19: Modeled trace element and REE patterns of potential mantle sources, for Whitechuck LKT (red circles) and Indian Pass CAB (blue squares) normalized to chondrite (Sun and McDonough, 1989) assuming maximum dry decompression of mantle with no hydrous phases (3.8 % partial melting for LKT, 3.5% partial melting for CAB). Conserved elements (Ti, Nb, Y, Yb) are more depleted than DMM (white crosses) (Salters and Stracke, 2004) or E-DMM (black stars) (Workman and Hart, 2005).



Figure 20: Comparison of alkalis and H_2O for corrected inclusions (filled symbols). Both Whitechuck and Indian Pass lavas have low amounts of K_2O , yet higher amounts of H_2O (and Na₂O). Anomalous inclusion (IP_005) is more evolved (Figs. 5, 6) and shows evidence for significant degassing (Figure 13). Plotted fields for Black Crater (decompression (dry) source example) and Fuego (fluid-flux (wet) source example) from Sisson and Layne (1993).