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## Metamorphic Evolution of the Sierra de Maz: Implications for the Timing of Terrane Accretion on the Western Margin of Gondwana

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**METAMORPHIC EVOLUTION OF THE SIERRA DE MAZ: IMPLICATIONS FOR THE  
TIMING OF TERRANE ACCRETION ON THE WESTERN MARGIN OF GONDWANA**

By

Andrew Tholt

Accepted in Partial Completion  
of the Requirements for the Degree  
Master of Degree Name

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Andrew Tholt

June 11, 2018

**METAMORPHIC EVOLUTION OF THE SIERRA DE MAZ: IMPLICATIONS FOR THE  
TIMING OF TERRANE ACCRETION ON THE WESTERN MARGIN OF GONDWANA**

A Thesis  
Presented to  
The Faculty of  
Western Washington University

In Partial Fulfillment  
Of the Requirements for the Degree  
Master of Science

by  
Andrew Tholt  
June 2018



## Abstract

The Mesoproterozoic MARA terrane of South America has been interpreted to collide and subduct beneath the Gondwana margin of South America in the Cambrian. To test the proposed tectonic model, I combine metamorphic petrology and geochronology along with quantitative thermobarometry to constrain metamorphic events and peak conditions within the Sierra de Maz of the MARA terrane and the adjacent Sierra de Ramaditas. Foliations within the Sierra de Maz are steeply east dipping, with a major sinistral reverse shear zone separating the two major units (Zaino and Maz Complexes) of the northern range. Garnet Lu-Hf records a single metamorphic event in the Zaino Complex at ~425 Ma. Combined garnet Lu-Hf and monazite U-Pb reveal a polymetamorphic history in the Maz Complex with two distinct metamorphic events at ~1.2 Ga and ~425 Ma. Both Complexes were deformed by sinistral transpression between 435 and 415 Ma. Preserved across these Complexes is a Silurian-Devonian amphibolite to granulite inverted metamorphic field gradient that is compatible with observed and modeled gradients from other transpressive orogens. The Ramaditas complex records lower-pressure granulite facies metamorphism at ~460 Ma, was deformed at ~426 Ma, and is correlated with the Famatina arc rather than its previously proposed affiliation to units in the Sierra de Maz. Units of the Sierra de Maz do not share metamorphic histories with nearby ranges within the proposed extent of the MARA terrane. Age dissimilarities require that portions of the MARA terrane were tectonically juxtaposed likely as a result of transpressive translation along the Gondwana margin.

## **Acknowledgements**

Special thanks to my advisor, Dr. Sean Mulcahy, for his dedication to my education and his willingness to share knowledge. I'll never be able to thank Sean enough for allowing me to be a part of this project and to work with everyone involved. Thanks to my committee member, Dr. Elizabeth Schermer, for insightful discussion both in and outside of the classroom and for inspiring me to be a better scientist at every point of my career. Thanks to my committee member, Dr. Bill McClelland, for his holistic mentoring in the field and during data collection in the laboratory. Thanks to Dr. Sarah Roeske for her insights in the field and her commitment to mentoring students. Thanks to my stellar field assistant Wes Johns, and to Dr. Matt Coble of Stanford University.

This work was supported in part by funding from the Western Washington University graduate school and Department of Geology. Funding was also supported by the National Science Foundation Grant No. 1550110. Any opinions, findings, and conclusions or recommendations expressed in this material are those of the authors and do not necessarily reflect the views of the National Science Foundation.

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## Introduction

Collisional orogens are among the most complex tectonic settings on Earth and exhumed metamorphic rocks record processes that take place during convergence, collision, and exhumation. Classically held models of orogeny invoke a simple cycle of crustal thickening followed by exhumation (e.g., Thompson and England, 1984) and successfully reproduce pressure-temperature paths preserved in many metamorphic terranes. Exhumed orogens, however, often record multiple episodes of thickening and exhumation, magmatic addition, and/or orogen parallel translation which give rise to complicated metamorphic histories (e.g. Brown and Talbot, 1989; Mulcahy et al., 2014). Testing proposed tectonic models in these orogens requires detailed metamorphic and structural studies in order to reconstruct the most accurate pressure, temperature, and deformation history.

The western margin of South America preserves a poorly understood and complex orogenic belt that is critical to Paleozoic plate reconstructions and the relative positions of Laurentia and Gondwana. Basement exposures within the Maz, Arequipa, and Rio Apa (MARA) ranges span almost the entire length of South America, are located immediately west of the Famatina continental arc, with the prevailing tectonic model based largely on observations in the Western Sierras Pampeanas of northwest Argentina (Figure 1). The MARA terrane is interpreted as Laurentian-derived crystalline basement that preserves two metamorphic episodes at ~1200 Ma and between 460-430 Ma (Lucassen and Becchio, 2003; Casquet et al., 2006). The current tectonic model assumes that MARA rifted from Laurentia around 570 Ma and then collided and was partially subducted beneath the Rio de la Plata craton around 530 Ma (Rapela et al.,

2016; Figure 2). Direct evidence for ~570 Ma rifting comes from a carbonatite intrusion found within the Sierra de Maz of the MARA terrane, but high temperature rift-related metamorphic ages have yet to be documented. Evidence for a ~530 Ma collision is indirectly inferred from cessation of arc volcanism within the Rio de la Plata craton but direct evidence for this event is absent within MARA (Rapela et al., 2016; Figure 2).

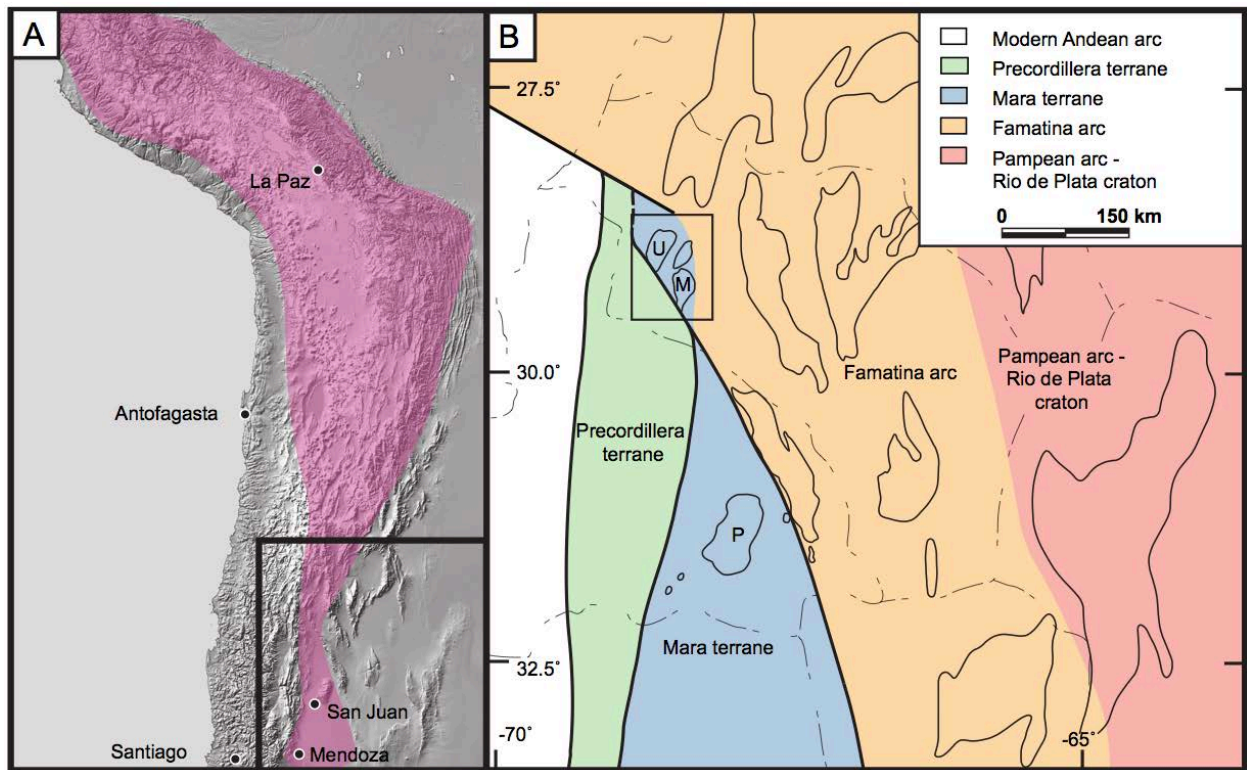


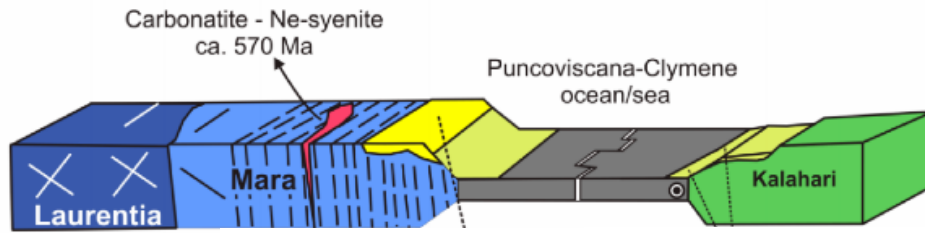
Figure 1: A) Proposed extent of the MARA terrane (magenta) and geographic location of map area in panel B. B) Geologic map of the Western Sierra Pampeanas, modified from Mulcahy et al. 2011. Abbreviations: M, Sierra de Maz; P, Sierra de Pie de Palo; and U, Sierra de Umango. Thick line weights indicate proposed terrane boundaries, thin weights outline important mountain ranges, dash-dot lines are major roadways. Boxed region is the extent of Figure 3.

Existing metamorphic ages from the MARA terrane are difficult to reconcile with the proposed tectonic model as this region records seemingly different metamorphic and deformational histories. Basement exposures in the Sierra de Pie de Palo (Figure 1) record lower-crustal Cambrian to Middle Ordovician shortening along low angle shear zones (Mulcahy, 2007; Mulcahy, 2011) between ~515-430 Ma. In contrast, the Sierra de Maz exposes mid-crustal rocks with steep fabrics interpreted to record oblique to strike-slip deformation of unknown age. Inferred metamorphic U-Pb ages within the Sierra de Maz of 460-430 Ma reported by Casquet et al. (2003) are not explained by the Rapela (2016) model. The ages imply that potentially, 1) the collision and juxtaposition of the MARA terrane with the Rio de Plata craton is younger than assumed, 2) metamorphic ages within the MARA terrane represent cooling ages or are misinterpreted, or 3) metapelitic rocks recording 530-520 Ma metamorphism and deformation have yet to be sampled.

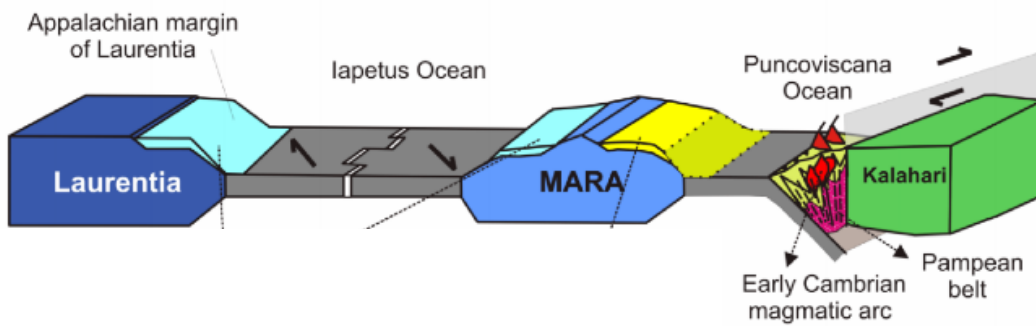
In order to test the Rapela et al. (2016) tectonic model for the MARA terrane, I combine quantitative thermobarometry with Lu-Hf and U-Pb geochronology to constrain the conditions and timing of metamorphism and deformation within the Sierra de Maz of the MARA terrane. The data suggest that rocks within the Sierra de Ramaditas are barometrically and chronologically distinct from the rest of the MARA terrane and likely represent a contact aureole more closely associated to the adjacent Famatina arc. The Sierra de Maz records Silurian to Devonian sinistral transpression and an amphibolite to granulite inverted metamorphic field gradient.



(a) ca. 630 Ma. **Opening of the transient Puncoviscana/Clymene ocean/sea**



(b) 540-530 Ma. **Opening of the Iapetus Ocean and coeval Pampean oblique subduction**



(c) 530-520 Ma. **Pampean oblique collision**

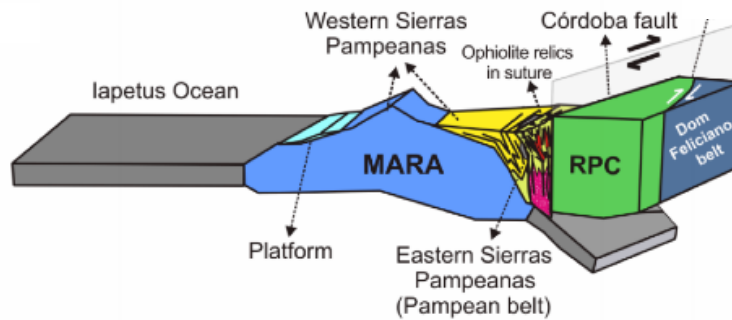


Figure 2: Proposed tectonic model for the MARA terrane modified from Rapela et al. 2016. RPC: Rio de la Plata Craton.

## **Geology of the Sierra de Maz**

The Sierra de Maz is situated between the allochthonous Devonian Precordillera terrane and the Ordovician Famatina continental magmatic arc (Figure 1B) and was first described by Kilmurray and Dalla Salda (1971) as three sub-parallel metamorphic domains trending NNW-SSE separated by first-order faults and shear zones (Figure 3). Foliation dips steeply to the east and metamorphic grade increases from west to east (Kilmurray and Dalla Salda, 1971; Casquet et al., 2006). From west to east, previous workers have divided the domain into distinct units: the Taco Complex, the Zaino Complex, the Maz Complex, and the Sierra de Ramaditas.

The El Taco Complex of the Sierra de Maz (Kilmurray and Dalla Salda, 1971), also referred to as the Western Domain by Casquet et al. (2008a), contains mostly marble, calc-silicate, and quartz-mica schists, with lesser amphibolite and orthogneiss. Titanite U-Pb ages of 443 and 431 Ma (Lucassen and Becchio, 2003) and a garnet Sm-Nd age of 301 Ma (Porcher et al., 2004) were reported from the Taco Complex with no quantitative estimates of metamorphic conditions.

The El Zaino Complex lies structurally above the El Taco Complex and was described by Casquet et al. (2008a) as a succession of Fe-rich garnet-chlorite schists of unknown age. Porcher et al. (2004) described the Zaino Complex as a sequence of metasediment of variable bulk compositions containing schist, calc-silicate, amphibolite, and marble, although their descriptions came solely from observations within the adjacent Sierra de Espinal to the north (Figure 1B). There are no published ages from this unit and thermobarometry is lacking.

The Maz Complex lies structurally above the Zaino Complex and is consists of mafic and pelitic gneiss, a Paleoproterozoic anorthosite massif, ~845 Ma and ~774 Ma A-type granitoids, and ~570 Ma carbonatites (Casquet et al., 2008a; Casquet et al., 2008b; Colombo et al., 2009; Rapela et al., 2010). Lucassen and Becchio (2003) reported a titanite U-Pb age of 431 Ma from a calc-silicate in the eastern portion of the Maz Complex. Casquet et al. (2006) recognized 1208 Ma metamorphic rims on zircon in orthogneiss and attributed them to Grenvillian metamorphism. Zircon cores from the same sample resulted in peak ages around 1700 and 1880 Ma and were interpreted to represent a late Paleoproterozoic provenance age for metasediments of the Maz Complex (Casquet et al., 2006). Casquet et al. (2006) also estimated granulite P-T conditions of 780 °C and 0.8 GPa from a metabasite adjacent to the eastern orthogneiss. Porcher et al. (2004) reported three whole rock Sm-Nd ages of 1040, 969, and 462 Ma from within the the Maz Complex. Three P-T estimates made by Porcher et al. (2004) from within the Maz Complex span a range of conditions from 630 - 771 °C and 5 - 7 kbar, but lack individual uncertainty estimates with at least one analysis plotting outside of the kyanite stability field as reported in the sample assemblage. Lucassen and Becchio (2010) also reported scattered P-T estimates from within the Maz Complex that range from 650 – 800 °C and 0.5 - 0.9 GPa. These previous studies have not systematically examined metamorphic conditions and timing within the Maz Complex as it relates to the history of deformation.

The Sierra de Ramaditas is a lesser range to the immediate southeast of Sierra de Maz (Figure 3). Porcher et al. (2004) described Ramaditas as a sequence of garnet-sillimanite migmatitic paragneiss, marble, and amphibolite and correlated the region

with the El Taco Complex. Casquet et al. (2008) obtained a weighted mean U-Pb zircon age of 442 Ma from within Sierra de Ramaditas and correlated it with units in the northern Sierra de Espinal.

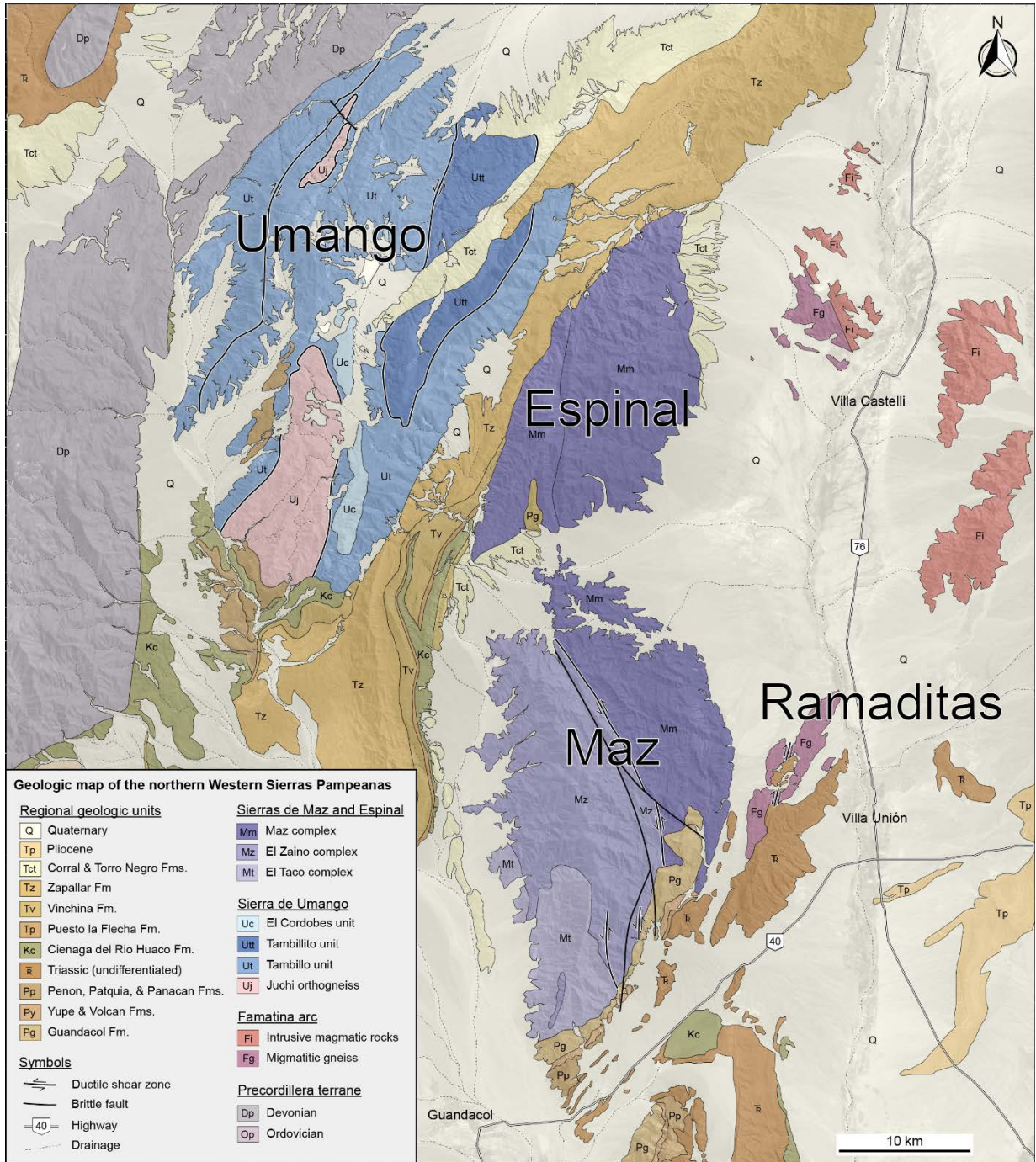


Figure 3: General geologic map of the Western Sierras Pampeanas and the Sierras de Maz, Espinal, Umango. The Sierra de Ramaditas is denoted by the map unit Fg (compiled and adapted after Meira et al., (2012) and references therein)

## Lithologic Units

This study follows the original naming convention established by Kilmurray and Dalla Salda (1971) with modifications based on observed differences in protolith, metamorphic assemblage, and metamorphic age (Figure 4). From west to east, we recognize four distinct units: the Taco Complex, the Zaino Complex, the Maz Complex, and the Ramaditas Complex of the Famatina arc (Figure 3).

The Taco Complex occurs only in the southwestern Sierra de Maz and consists of isoclinally folded schist, orthogneiss, paragneiss, phyllite, amphibolite, quartzite, and marble (Figure 3 and 5A). The boundary between the Taco Complex and the Zaino complex was not directly observed in this study. Pelitic assemblages of variable protolith contain  $\text{Grt-Bt-Qtz}\pm\text{Ms}\pm\text{Sil}\pm\text{Chl}$ . Amphibolite layers range in thickness from sub- to meter-scale and contain the assemblage  $\text{Amp-Qtz-Bt-Pl}\pm\text{Ms}\pm\text{Grt}$ . The dominant foliation dips from  $55^\circ\text{E}$  to sub-vertical with multiple zones of grain size reduction throughout the unit.

Directly northeast from the Taco Complex lies the Zaino Complex, spanning almost the entire length of the Sierra de Maz (Figure 3). The eastern boundary of the Zaino Complex is defined by the NNW-SSE trending Maz shear zone. Casquet et al. (2008) described this unit as Fe-rich  $\text{Grt-Chl}$  schists (Figure 5B), however, we recognize additional metacarbonate, calc-silicate, amphibolite, and pelitic schist. Amphibolite is present on the sub-meter scale and contains the assemblage  $\text{Qtz-Plg-Amp}\pm\text{Ms}$ . Pelitic schists have variable assemblages that change systematically from west to east across the Zaino Complex (Fig 4?). In the most westerly regions of the Zaino Complex, schists contain the assemblage  $\text{Qtz-Plg-Ms-Chl-Bt}\pm\text{Grt}$ . Moving east, the dominant pelitic

assemblage becomes Qtz-Plg-Ms-Bt-Grt (Figure 5C) before crossing into the most eastern portions where metapelites contain the assemblage Qtz-Plg-Ms-Grt-Bt-St (Figure 5D). The Zaino Complex is isoclinally folded with post-kinematic garnets growing over a steeply inclined NW-SE striking axial planar cleavage (Figure 4; Figure 5C). Foliation primarily dips eastward with stretched micas defining shallow lineations (Figure 4).

The Maz Complex extends eastward from the Maz shear zone to the termination of the range and has been divided into four parallel groups based on protolith differences (Figure 4). Immediately east of the Maz shear zone is the first subunit termed Mm1 that contains amphibolite, orthogneiss, and less common pelitic layers. Metapelites found nearest to the Maz shear zone contain the assemblage Qtz-Plg-Ms-St-Grt with syn to post-kinematic garnet porphyroblasts. This unit is isoclinally folded. The next sub-parallel zone to the east, Mm2, is a sequence of paragneiss, amphibolite, and interleaved orthogneiss (Figure 5F). Paragneiss from Mm2 contains pre to syn-kinematic garnet with the assemblage Qtz-Plg-Ms-Bt-St-Grt-Chl±Ky (Figure 5E). Rare outcrops within Mm2 as well as Mm4 contain pre-kinematic garnet megacrysts several centimeters in diameter with chlorite rinds. Orthogneiss of Mm2 contains the assemblage Amp-Qtz-Bt-Plg-Ms-Gt with post kinematic garnet. Within Mm4 is massif-type anorthosite body interleaved within the dominant NNW-SSE foliation and labeled as Mm3 in Figure 4. Directly east of the anorthosite massif is the subunit Mm4, with the dominant lithologies being intermediate to mafic orthogneiss. Potassium feldspar bearing augen gneiss outcrops in multiple locations within Mm4 and can be seen both cutting and being deformed by the latest foliation event. Paragneiss assemblages in the

most eastern portions of Mm4 contain pre-kinematic garnets and the assemblage Qtz-Plg-Ms-Bt-St-Grt-Sil-Kfs (Figure 5G).

Rocks of the Ramaditas Complex within the Sierra de Ramaditas lie unconformably below Triassic red beds (Figure 3). The unit consists of metacarbonate, migmatite, intermediate to felsic orthogneiss, amphibolites and metapelite. The unit is isoclinally folded and the dominant foliation dips to the east following the general trend of units within the Sierra de Maz. Weakly deformed felsic Kfs-Bt pegmatites cross cut the dominant NNW-SSE foliation. Paragneiss from the Ramaditas Complex contains the assemblage Grt-Bt-Kfs-Qtz-Plg-Bt±Ctd±Sil±Ms.



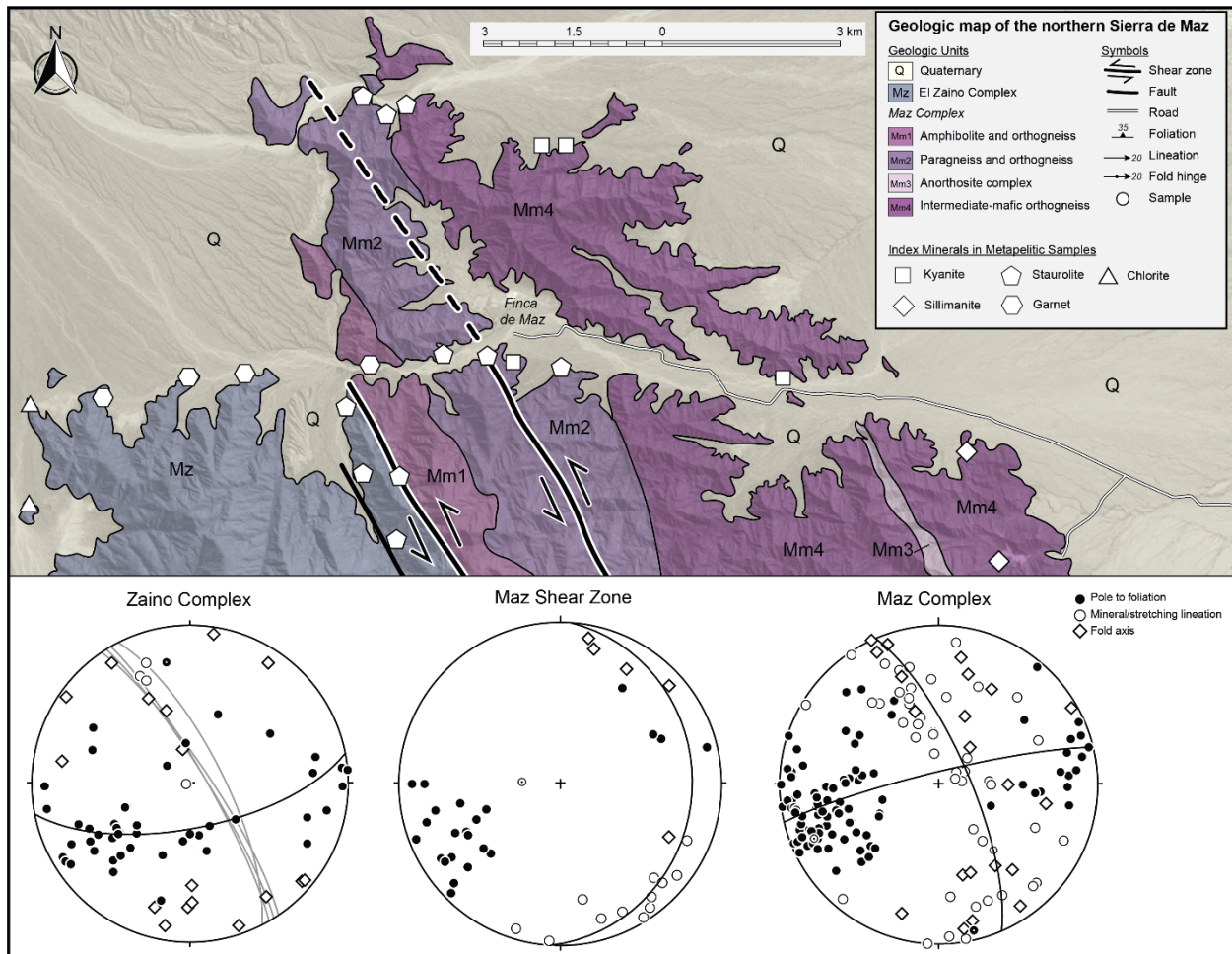


Figure 4 Geologic map of northern Sierra de Maz illustrating variation in metamorphic grade across the Zaino Complex, Maz shear zone, and the Maz Complex. Pelitic index minerals, denoted by various polygons, indicate an east to west inverted metamorphic gradient across the northern Sierra de Maz. Stereonets display poles to foliations, fold axes, and lineations from within the Zaino and Maz Complexes as well as the Maz shear zone. Grey lines in the Zaino Complex stereonet are orientations of axial planar cleavage, black lines are calculated great circle fits to field measurements.



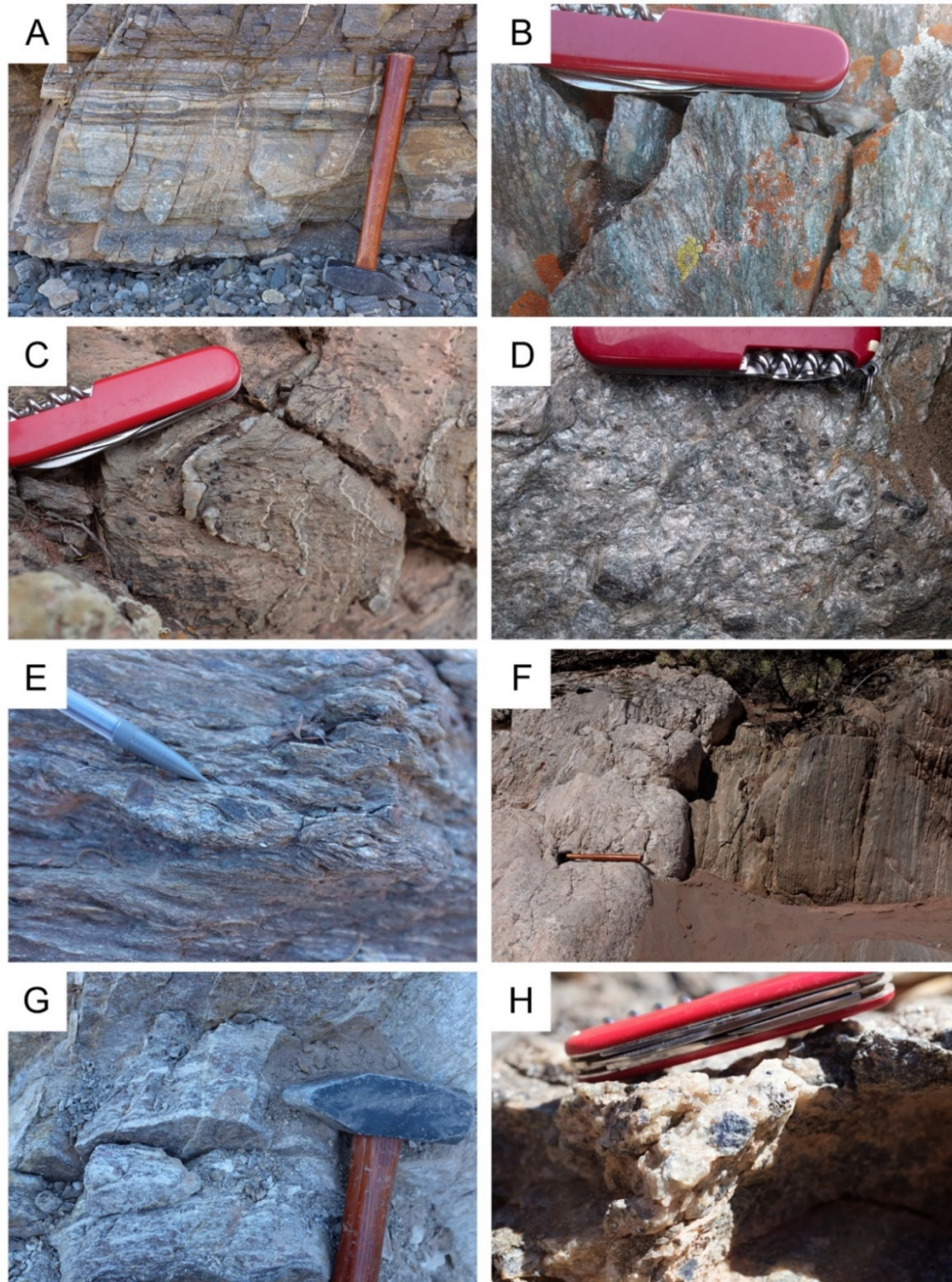


Figure 5: Outcrop photographs of units within the Sierra de Maz. Scales for photos are either a mechanical pencil, a ~50cm long rock hammer, or a ~10cm long pocket knife. A) Isoclinal folding and gneissic banding within units of the Taco Complex (sample AT-062). B) Chlorite schist of the Zaino Complex (sample AT-017). C) Post-kinematic euhedral garnets overgrowing axial planar cleavage within the Zaino Complex (sample AT-021). D) Qtz-Plg-Ms-Bt-Grt-St schist found within the most eastern exposures of the Zaino Complex. (sample AT-072). E) Pre- to syn-kinematic Qtz-Plg-Grt-Ms-Bt-St schist within subunit Mm2 of the Maz Complex. F) Paragneiss (right side of photo, darkly colored) and orthogneiss (left side of photo, lightly colored) within the Maz Complex subunit Mm2 (sample AT-068). G) granulite facies metapelitic outcrop within the Maz Complex subunit Mm4 (sample AT-028). H) Blue, pseudo-hexagonal cordierite grain found within the Sierra de Ramaditas.

## Ductile Shear Zones

Previous studies (Porcher et al., 2004; Casquet et al., 2006) have recognized the La Vibora brittle fault as the boundary between El Taco and El Zaino Complexes, however no previous work has related ductile shearing within Sierra de Maz to the metamorphic history of the region. This study recognizes the Maz shear zone as a major structural boundary separating the Zaino Complex in west from and the Maz Complex in the east. The Maz shear zone is a meter-scale ultramylonite zone of reverse-sinistral shear (Houlihan et al., 2017) that trends NNW-SSE and dips steeply east (Figure 6A; Figure 4). Lineations are shallow to sub-horizontal and are defined by stretched mica, quartz, and aligned amphibole. Even though the ultramylonite Maz and Finca shear zones are delineated on the map as a discrete line, those lines mark the regions of highest strain, but deformation extends over a zone of more than 7 km wide (Figure 6B).

While the Maz shear zone is a prominent mylonite-ultramylonite zone, smaller regions of mylonite and ultramylonite occur within the Maz Complex. One such region, located within the subunit Mm2 of the Maz Complex is referred in this study as the Finca shear zone. This sub-meter ultramylonite is concentrated within metapelite and orthogneiss around a region of amphibolite. Mylonitic fabrics in the Finca shear zone are sub-parallel to Maz shear zone with similar shallow lineations and a sinistral shear sense (Figure 6C). Pelitic units adjacent to this shear zone contain coarse, zoned garnet of varying size that show pre, syn, and post-kinematic relationships to the most recent shear fabric.

The Ramaditas shear zone is a meter-scale mylonite within the Sierra de Ramaditas that has not been described by previous studies of this range. The trend of the shear zone is roughly N-S with foliation steeply east dipping to sub vertical. Lineations are poorly preserved within the main mylonite zone, but are generally observed to be oblique. The shear sense is dominantly top-to-the-west thrust with a locally weakly developed right-lateral shear sense indicators. A coarse, felsic dike with centimeter sized grains of Kfs-Qtz-Ms can be seen cutting across the main mylonite zone and is also boudinaged and deformed within the mylonitic foliation (Figure 6D). This relationship suggests that shear is pre- to syn-tectonic with respect to the felsic dike. Leucocratic melts within mafic orthogneiss near the main mylonite zone contain the assemblage Qtz-Kfs-Grt while pelitic units near the shear zone contain the assemblage Qtz-Kfs-Bt-Sil.



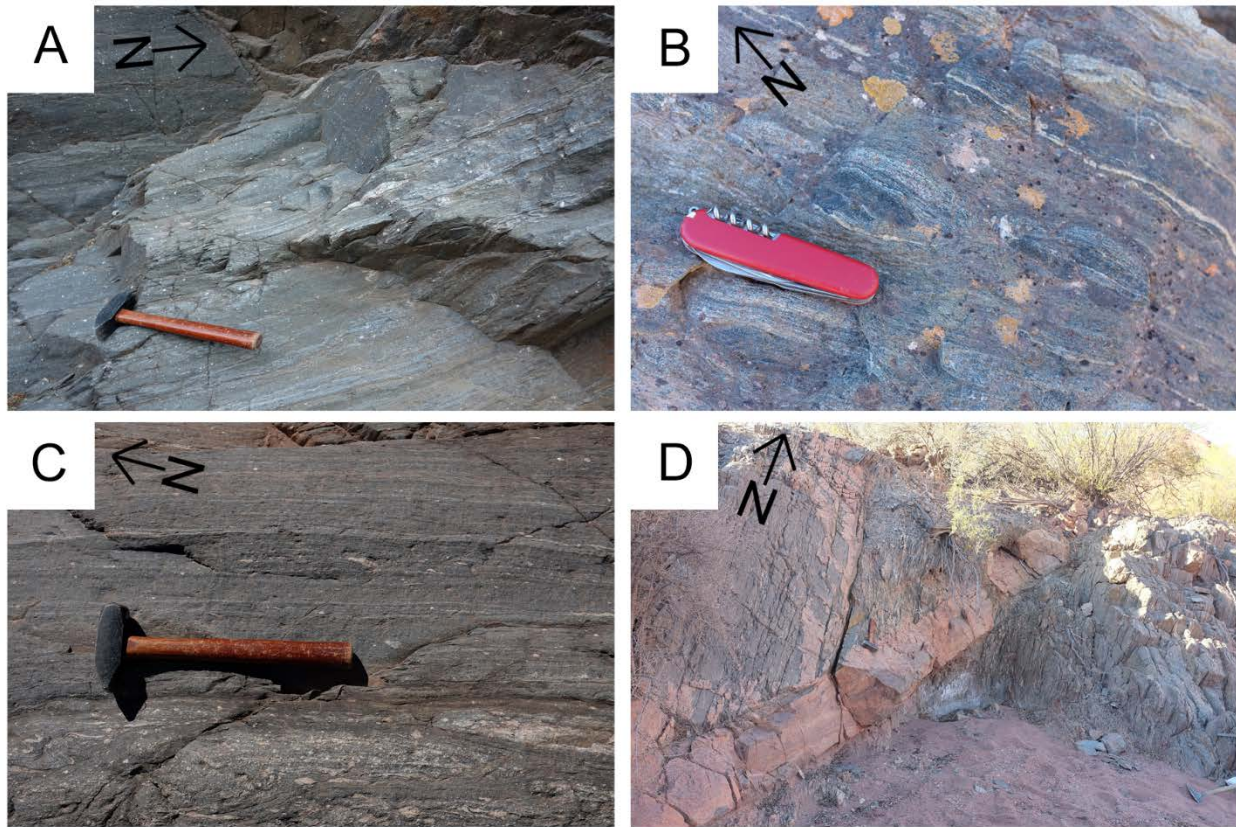


Figure 6: Outcrop photographs of sheared units within the Sierra de Maz. Scales for photos are either ~50cm long rock hammer or a ~10cm long pocket knife. A) Steeply dipping ultramylonite zone from the Maz shear zone (proximal to sample AT-076). B) A relict orthogneiss fabric from the Maz Complex subunit Mm2 that is boudinaged by foliation related to the Maz shear zone (sample AT-015). C) Half meter ultramylonite from the Finca shear zone within the Maz Complex subunit Mm2 (sample AT-068). D) Steeply dipping Ramaditas shear zone with coarse Kfs dike that both cross-cuts and is boudinaged by the shear zone.

## Summary

Pelitic index minerals suggest an amphibolite to granulite facies inverted sequence is preserved within the Maz Complex with pelitic index minerals grading from staurolite to kyanite and eventually sillimanite + potassium feldspar from west to east and structurally upward. The presence of the low-pressure phase cordierite (Figure 5H) within the Sierra de Ramaditas points to significantly different metamorphic conditions in this unit compared to high-pressure kyanite-bearing assemblages within the Sierra de Maz. Outcrop geology suggests that foliation development related to the Maz shear zone post-dates much of the metamorphic development within the Maz Complex and is a major event in the development of fabrics related to the Zaino Complex. This notion is most strongly supported in relict fabrics and assemblages observed within Maz Complex subunit Mm2 (Figure 6B) that are otherwise absent from the Zaino Complex. Sinistral kinematics and shallow lineations within steeply dipping foliations suggest that movement along the Maz shear zone accommodated transpressive deformation.

## **Petrology and Thermobarometry**

Samples were collected along a west-to-east transect across the Sierras de Maz (Figure 4; Table 1; Table 2) and Ramaditas in order to (1) identify the metamorphic conditions of the various lithological units, (2) compare differences in the metamorphic conditions across major structures, and (3) compare the metamorphic conditions between the Sierra de Maz and Sierra de Ramaditas. Mineral compositions were determined by electron probe microanalysis (EPMA) with a Cameca SX-100 at UC Davis. Operating conditions including spot sizes and accelerating voltages are reported in DR1 of the digital appendix. Temperature and pressure estimations were determined with multiequilibrium thermobarometry using average PT method in THERMOCALC (Powell and Holland, 1994) and are summarized in Figure 7 and Table 3.

## **Taco Complex**

High grade pelitic schist from the EL Taco Complex contains Grt-Bt-Ky-Sil-Plg-Kfs-Qtz (Figure 8A). Samples from the El Taco Complex were not analyzed with EPMA, thus only qualitative estimates of temperature and pressure can be made. The presence of sillimanite and K-feldspar along with the absence of muscovite implies that this rock is beyond the second sillimanite reaction. A lack of cordierite and orthopyroxene within this sample places pressure estimates between 4-12 kilobars and temperature estimates between 750-900 °C (Spear and Cheney, 1999).

## Zaino Complex

Pelitic schist assemblages from the Zaino Complex vary systematically from west to east across the unit, consistent with an increase in metamorphic grade. The most western outcrops contain abundant chlorite (Figure 8B) and transition into garnet-bearing assemblages over a structural distance less than 2 kilometers. Average PT calculations were determined for three pelitic assemblages using garnet rim and matrix mineral compositions with Qtz and H<sub>2</sub>O in excess. Sample AT-021a from the Zaino Complex contains the assemblage Grt-Bt-Ms-Plg-Qtz (Figure 8C). The primary foliation is an axial planar cleavage defined by alternating mica-rich and quartzo-feldspathic layers and cut by post-kinematic garnet. Garnet is compositionally zoned with a bell-shaped XSps profile and XMg concentrations that increase from core to rim suggestive of a single phase of prograde garnet growth (Figure 9A; Figure 10A). Garnet rim compositions paired with matrix Ms, Bt, and Plg record metamorphic conditions of  $8.2 \pm 1.7$  kbar and  $564 \pm 91$  °C. Sample AT-021b contains the assemblage Grt-Bt-Plg-Qtz-Amp with post-kinematic garnets that similarly cut across an axial planar cleavage. Garnet rim compositions and matrix Bt, Ms, and Plg record metamorphic conditions of  $7.3 \pm 1.6$  kbar and  $489 \pm 71$  °C. Sample AT-017 was the easternmost sample analyzed from the Zaino Complex and contains the assemblage Grt-Bt-Ms-Plg-Chl-Qtz. Garnet rim compositions and matrix Bt, Ms, Plg, and Chl record metamorphic conditions of  $8.2 \pm 1.4$  kbar and  $575 \pm 14$  °C. The most eastern metapelitic outcrops of the Zaino Complex contain the assemblage Grt-St-Bt-Ms-Plg-Qtz, although these samples were not analyzed for PT conditions, pelitic index assemblages qualitatively suggest that temperatures for this region were 550-650 °C (Spear and Cheney, 1989). Taken

together, the three samples analyzed using EPMA are interpreted to represent peak metamorphic conditions reached within the Zaino Complex after deformation.

### **Maz Shear Zone**

Two metapelites and one amphibolite were collected from within the Maz shear zone and analyzed for pressure and temperature conditions. Syn to post-kinematic garnets from the Maz shear zone are Fe-rich ( $X_{Alm} > 0.70$ ) with discrete zones apparent in major element maps (Figure 9B). Chemical zoning analysis reveal diffusively zoned garnets that are lacking in rim-ward increases in spessartine content. Fig 10b? Sample AT-076 from within the Maz shear zone contains the assemblage Grt-St-Bt-Ms-Pl and resulted in average PT calculations of  $7.7 \pm 1.1$  kbar and  $653 \pm 27$  °C (Figure 8E). Sample AT-073 from 0.5 km northwest along strike contains the assemblage Grt-St-Bt-Ms-Plg-Qtz and resulted in average PT conditions of  $8.3 \pm 1.4$  kbar and  $602 \pm 90$  °C. Sample AT-015c is a mafic assemblage of Grt-Bt-Plg-Amp-Qtz with post-kinematic garnets and aligned amphibole defining a mylonitic foliation that deforms and older fabric (Figure 5B; Figure 8F). The sample records average PT conditions of  $8.6 \pm 2$  kbar and  $652 \pm 153$  °C. These estimates are interpreted to reflect peak metamorphic conditions during deformation on the Maz shear zone.



## Maz Complex

Schist and gneiss from the Maz Complex exhibit a range of assemblages with index minerals indicating an increase in metamorphic grade from west to east across the unit (Figure 4). Twelve samples along an east-west transect of the Maz Complex were analyzed for pressure and temperature conditions (Table 2). The most westerly sample nearest to the Maz shear zone ultramylonite, sample AT-066, contains pre-kinematic garnets in the assemblage Grt-St-Bt-Ms-Plg-Qtz. Garnet rim compositions combined with matrix Bt, Ms, and Plg record metamorphic conditions of  $9.0 \pm 1.5$  kbar and  $706 \pm 34$  °C. Roughly 3 kilometers east of the Maz shear zone, sample AT-068 from within the Finca shear zone contains the assemblage Grt-St-Bt-Ms-Plg-Qtz. Garnets are pre-kinematic (Figure 8G) and possess irregular major element zoning, including elevated Mn and Ca along garnet fractures possibly from post-garnet fluid infiltration (Figure 9C). Rim to rim garnet compositional zoning generally reflects diffusive zoning with symmetrical increases in spessartine content suggesting multiple garnet growth events are preserved in this rock (Figure 10C). Garnet rim compositions paired with matrix St, Bt, Ms, and Plg compositions result in average PT conditions of  $7.2 \pm 1.1$  kbar and  $694 \pm 27$  °C. Less than a kilometer east of the Finca shear zone, sample AT-070 contains the assemblage Grt-Ky-St-Bt-Ms-Plg, with retrograde chlorite and resulted in average PT conditions of  $6.6 \pm 0.9$  kbar and  $624 \pm 17$  °C (Figure 8H). In the most eastern portions of the Maz Complex, sample AT-028 contains a metapelitic granulite facies assemblage of Grt-Sil-Bt-Kfs-Plg-Qtz (Figure 8I). Garnets are pre-kinematic and lack major element zoning in major element maps (Figure 9E) with rim to rim chemical analysis showing homogenized XSps compositions and weakly zoned

XMg (Figure 10D). Maximum XMg garnet rim compositions paired with matrix Bt, Kfs, and Plg resulted in average PT conditions of  $10.3 \pm 1.6$  kbar and  $876 \pm 56$  °C. Mafic orthogneiss from the eastern portions of the Maz Complex also contains the assemblage Grt-Cpx-Scp (Figure 8J) consistent with granulite facies metamorphic conditions.

### **Ramaditas Complex**

Several samples were collected from the Sierra de Ramaditas, however only a single metapelite contained enough relevant phases for PT calculations. Garnets from Ramaditas are pre-kinematic, Fe-rich ( $X_{Alm} > 0.70$ ), and major element maps lacking significant compositional zoning (Figure 8K; Figure 9E). Rim to rim compositional zoning however, shows rimward increases in XSp content and decreases in XMg that indicate diffusive cooling from high temperatures (Figure 10E). Sample AT-047 contains the assemblage Cd-Grt-Bt-Plg-Qtz and resulted in average PT conditions of  $5.5 \pm 1.5$  kbar and  $856 \pm 160$  °C. Additional samples collected from within the Sierra de Ramaditas have the assemblage Grt-Sil-Bt-Plg-Qtz (Figure 8L) but were not analyzed for average PT conditions. The presence of stable cordierite within metapelite outcrops indicates that high temperatures occurred at anomalously low pressure when compared with metapelite assemblages from the Sierra de Maz (Figure 7).

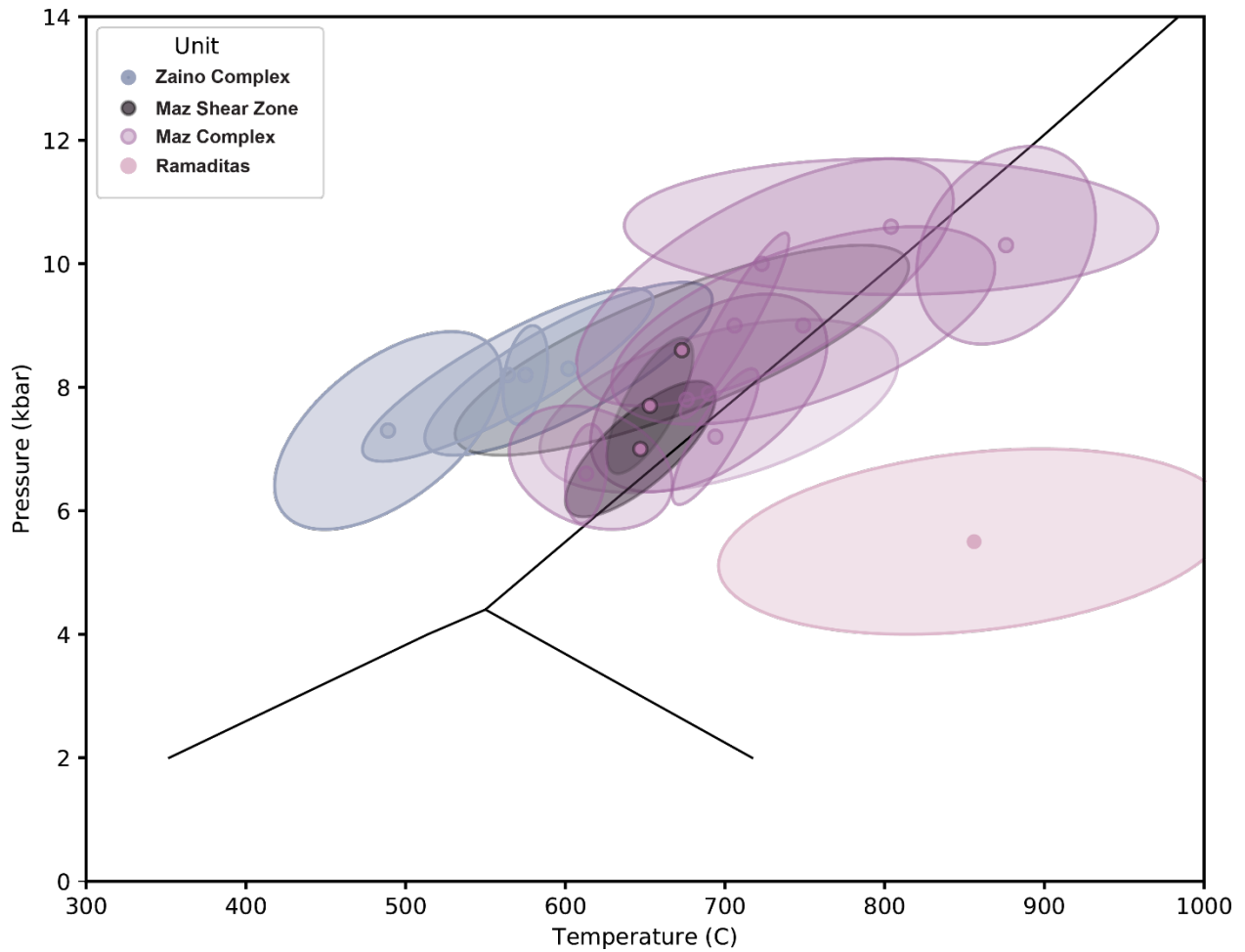


Figure 7: Pressure-temperature estimates and associated uncertainties for all samples analyzed from the Sierra de Maz and Sierra de Ramaditas. Conditions across the Sierra de Maz most closely reflect a Barrovian progression from west to east. The single analysis at anomalously higher temperatures and lower pressures than the rest of the data comes cordierite-bearing assemblages of the Sierra de Ramaditas.



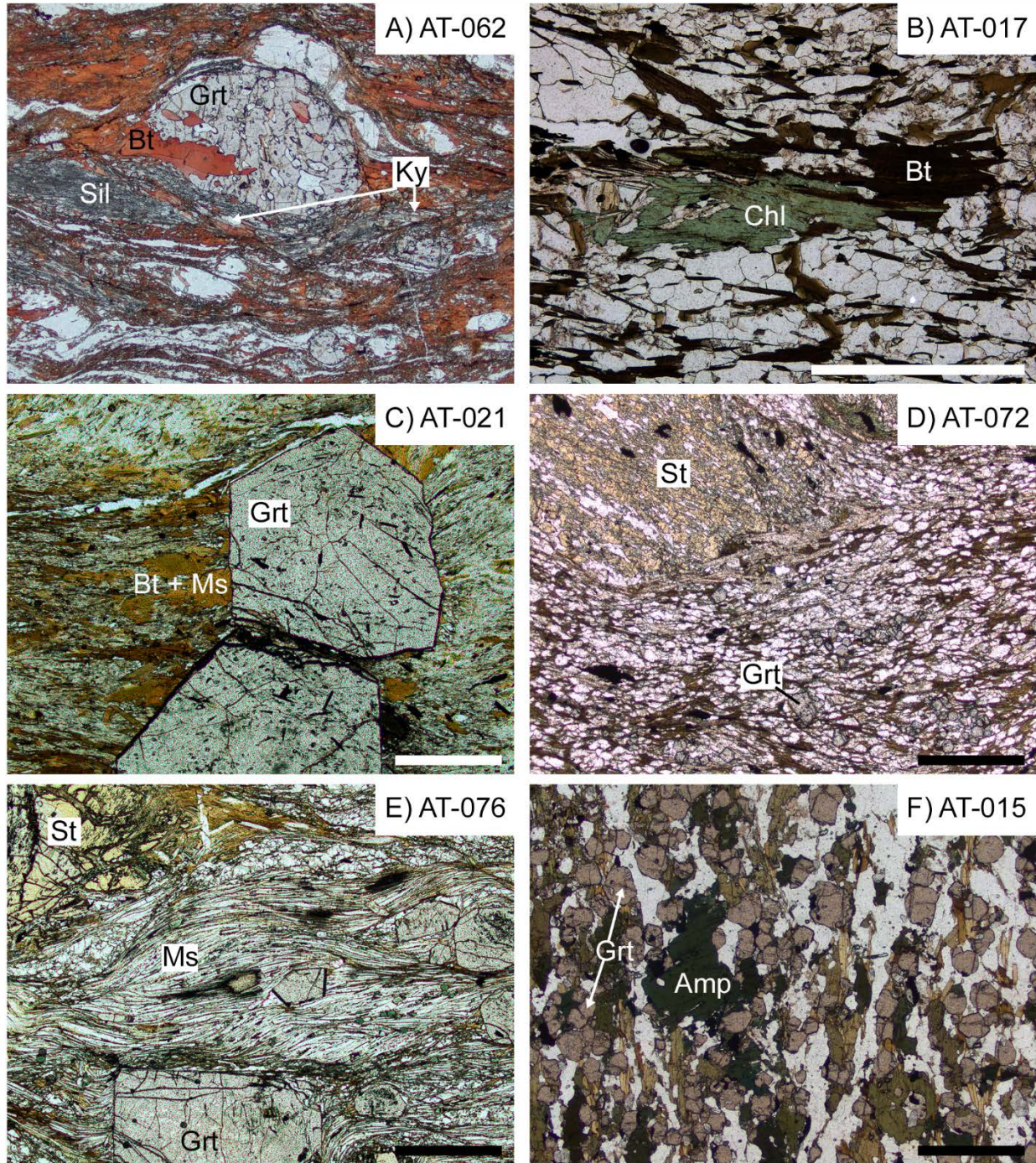


Figure 8: Photomicrographs of mineral assemblages from the Sierras de Maz and Ramaditas. Black and white scale bars in the bottom right corners of each photo are 1mm in length. A) Sil-Grt from the El Taco Complex. B) Bt-Chl schist from the Zaino Complex. C) Post-kinematic garnets from the Zaino Complex and the assemblage Grt-Bt-Ms-Plg. D) Grt-St-Bt schist from eastern units within the Zaino Complex. E) Grt-St-Ms-Plg-Bt assemblage from a metapelite near the Maz shear zone exhibiting syn to post-kinematic garnets. F) Mafic orthogneiss with the assemblage Grt-Bt-Ms-Amp just east of the Maz shear zone.



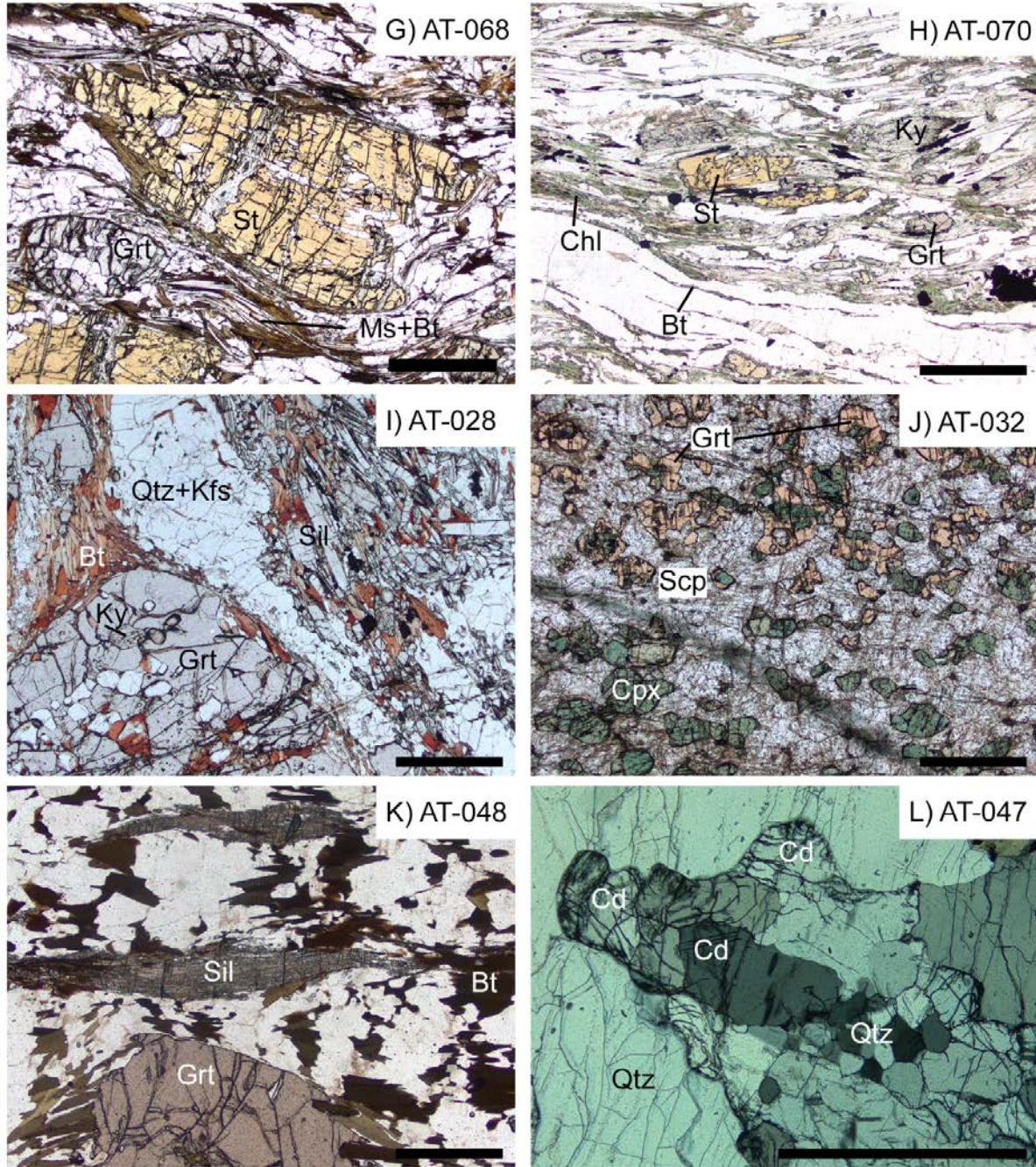


Figure 8 (continued): Photomicrographs of mineral assemblages from the Sierras de Maz and Ramaditas. Black and white scale bars in the bottom right corners of each photo are 1mm in length. G) Grt-St-Bt-Plg-Ms schist near the Finca shear zone of the Maz Complex. H) Grt-Ky-Bt-Plg-Ms schist with retrograde chlorite less than a kilometer east of the Finca shear zone. I) Eastern Maz Grt-Bt-Ms-Sil-Plg-Kfs gneiss with an inclusion of kyanite preserved within a relict garnet. J) Orthogneiss from the eastern units of the Maz Complex containing the assemblage Grt-Scp-Amp. K) Grt-Sil-Bt gneiss from the Ramaditas Complex. L) Cordierite from the Sierra Ramaditas with characteristic fractures that radiate from grain boundaries and terminate within interior portions of each grain. Fracture patterns and subtle twinning distinguish cordierite from quartz and feldspar.



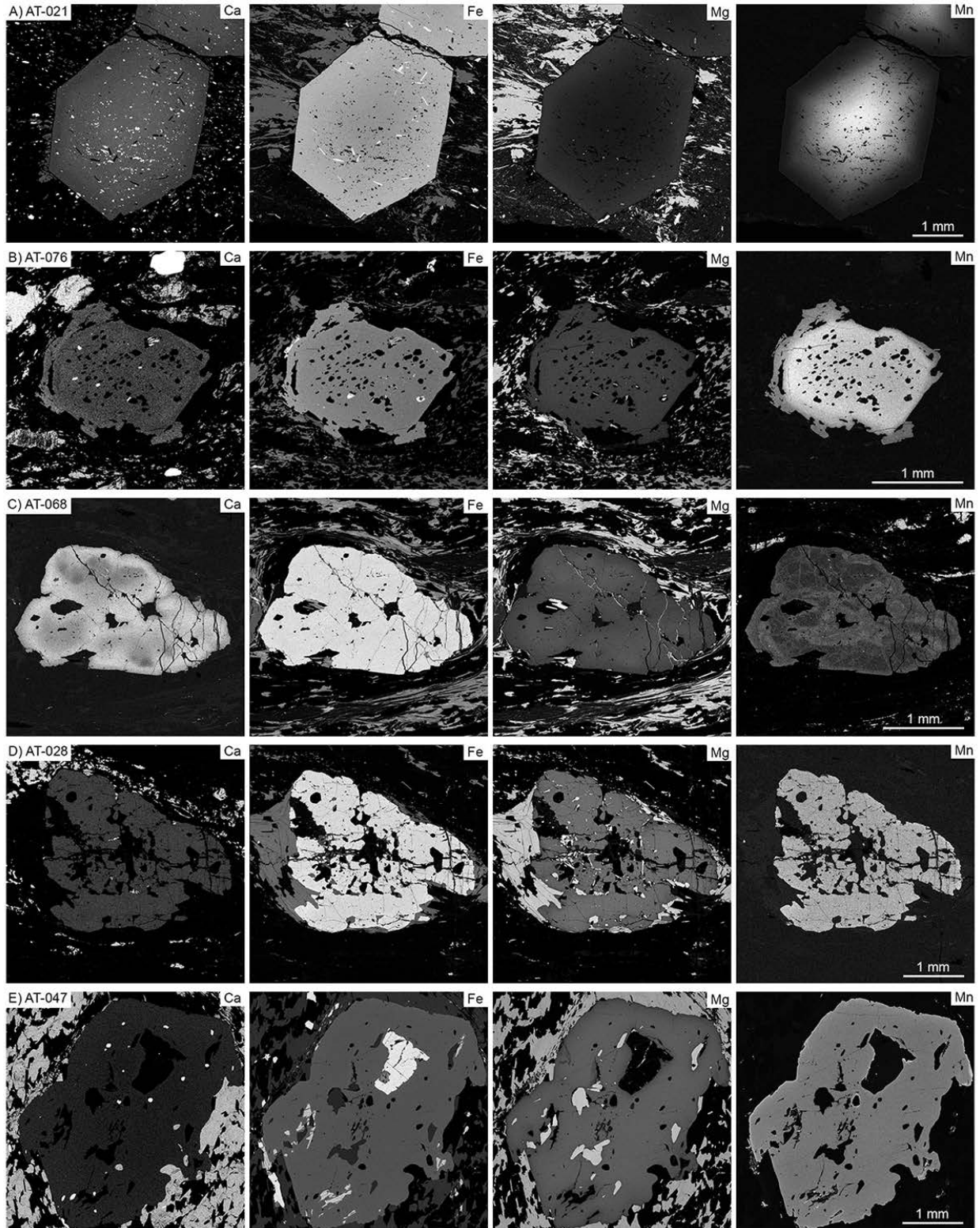


Figure 9: Zoning patterns of garnets chosen for combined thermobarometry and Lu-Hf geochronology. Light colors correspond to regions of high concentration.

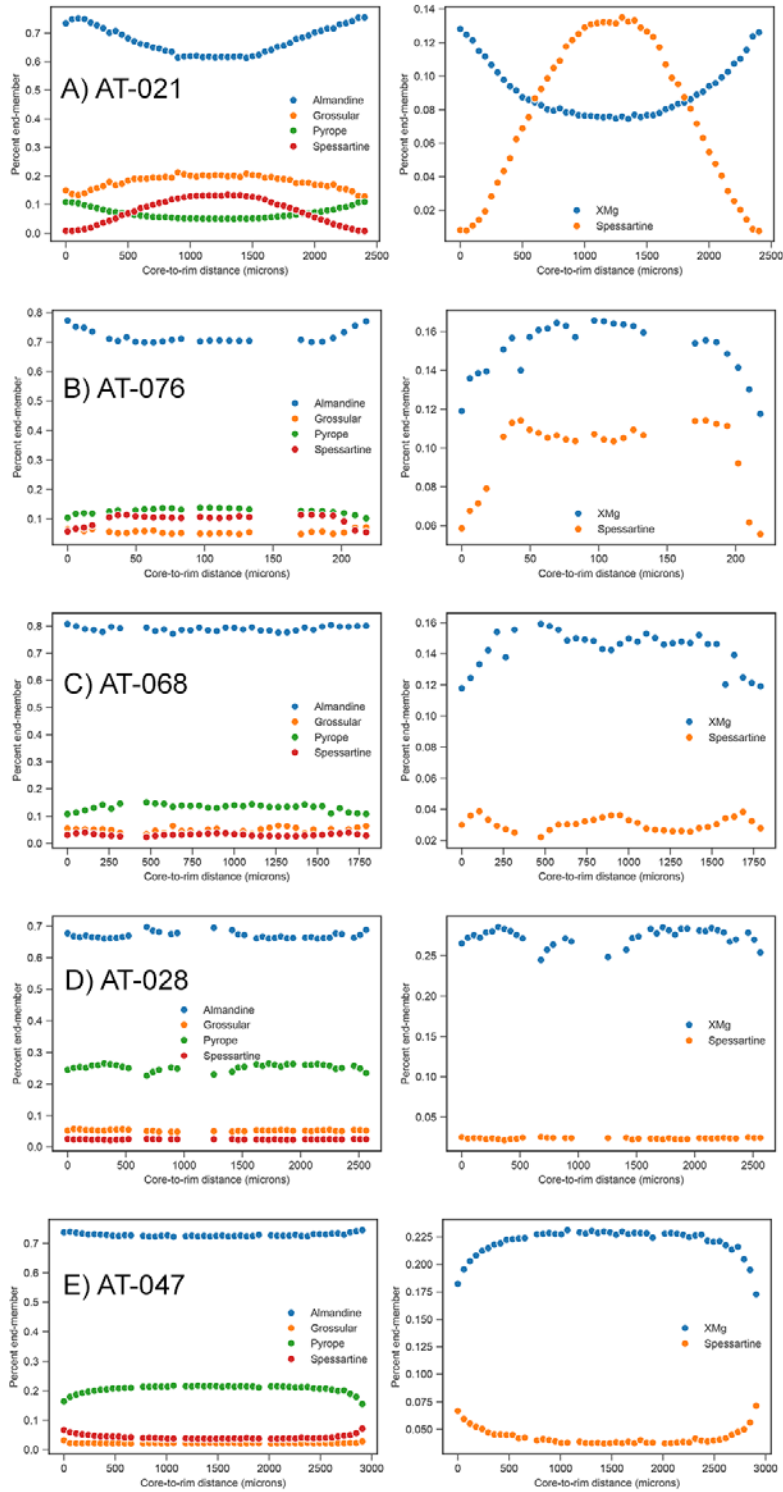


Figure 10: Figures within the left column depict garnet end-member compositions at given points along a transect. Figures in the right column depict elemental ratios at any given point. (XMg = Mg/Fe + Mg, Spessartine = Mn/ Fe + Mg + Mn + Ca).

## Garnet Lu-Hf Geochronology

Garnet-whole rock Lu-Hf ages for all samples were determined at Washington State University following the procedure outlined by Mulcahy et al. (2014). Garnet porphyroblasts were separated by hand crushing, magnetic separation, and hand-picking of individual grains. The concentrations of Lu and Hf were determined by isotope dilution with uncertainties estimated to be better than 0.5%. Uncertainties in  $^{176}\text{Lu}/^{177}\text{Hf}$  for the purposes of regression and age calculations are estimated at 0.5%. Reported errors on  $^{176}\text{Hf}/^{177}\text{Hf}$  represent within-run uncertainty expressed as  $2\sigma$  standard error. Estimated total uncertainty on individual  $^{176}\text{Hf}/^{177}\text{Hf}$  measurements for regressions and age calculations is 0.01%. Ages were calculated with the  $^{176}\text{Lu}$  decay constant values of Scherer et al. (2001) and Söderlund et al. (2004) and are reported with the associated 95% confidence interval. Concentrations and isotopic data for garnet Lu-Hf isochrons are listed in Table 4.

Rim to rim lutetium concentrations were determined by laser ablation inductively coupled plasma mass spectrometry (LA-ICPMS) in the AMSEC facilities at Western Washington University using an Agilent 7500ce ICPMS with a New Wave UP-213AS laser ablation accessory utilizing a 213nm UV Nd:YAG class IV laser. Time-resolved spectra with a 1 ms integration were collected using a 55  $\mu\text{m}$  diameter spot size. Working acquisition time was 120 seconds, including 30 seconds of background, and 90 seconds of sample ablation. All analyses used a 1500 W forward power, a 10 Hz repetition rate, a 13.7  $\text{J}/\text{cm}^2$  fluence, and a carrier gas flow rate of 400mL/min. Samples were calibrated to NIST610 values determined by Jochum et al. (2011), using an internal standard of aluminum determined by SEM analysis. Data were reduced using



the software GLITTER (van Achterburgh et al., 2000) to avoid ablated inclusions within garnet spots.

### **Zaino Complex**

A garnet-whole rock Lu-Hf isochron obtained from a Grt-Bt-Ms-Qtz-Plg metapelite (AT-021a) of the Zaino Complex resulted in an age of  $423 \pm 10$  Ma (MSWD = 4.6) (Figure 11A). Garnet from this sample are euhedral, post-kinematic to axial planar cleavage, and preserve prograde growth zoning (Figure 7C, 8A, 9A). Lutetium is concentrated within the cores of garnet porphyroblasts and systematically decreases towards the rim consistent with a single generation of garnet growth (Figure 12A). Garnets from this sample are bimodal in size with grains either larger than >2mm or between 1-2mm. The smaller grain size fraction had higher  $^{176}\text{Lu}/^{177}\text{Hf}$ . Despite the apparent size dependency of parent daughter ratios, the high MSWD and bell-shaped Mn and Lu profiles suggest a single period of protracted garnet growth (Kohn, 2009). The age from this sample is interpreted to date garnet growth in the Zaino Complex as result of burial along the Maz shear zone.

### **Maz Shear Zone**

Two samples were collected and analyzed from within the Maz shear zone. Sample AT-076b, a Grt-St-Ms-Plg-Bt schist within the Maz shear zone produced a garnet-whole rock Lu-Hf isochron of  $429 \pm 5$  Ma (MSWD=2.6) (Figure 11B). Garnets

from this sample are pre- to syn-kinematic with respect to the Maz shear zone (Figure 7E, 8B, and 9B). Lutetium distribution in this sample does not have a central peak and is concentrated on the rims of garnets, implying the age is biased towards the youngest period of garnet growth. Garnets from a mafic Grt-Bt-Plg-Amp orthogneiss (AT-015c) are post-kinematic with respect to the Maz shear zone and resulted in a Lu-Hf isochron of  $428 \pm 9$  Ma (MSWD=5.6) (Figure 11C). These two ages are interpreted to represent the age of garnet growth during the deformation along the Maz shear zone.

### **Maz Complex**

Two metapelites were sampled from the Maz Complex along a transect east from the Maz Shear zone. Resultant isochrons from samples AT-068 and AT-028 contained significant scatter and did not result in clear age interpretations. The oldest age calculated from sample AT-068 comes from a three-point isochron at  $759 \pm 54$  Ma while the youngest calculated age is  $591 \pm 59$  Ma. Sample AT-028 produced a less scattered isochron with an oldest age of  $1074 \pm 6.4$  Ma and a youngest age of  $1030.9 \pm 6.2$  Ma. The scattered Lu-Hf ratios and major and trace element zoning in the two samples suggest they did not remain a closed system and/or experienced multiple metamorphic events.

## Ramaditas

One Grt-Sil-Bt-Plg-Qtz gneiss, sample AT-047b, was chosen from within the Sierra de Ramaditas and produced a scattered isochron with an oldest age of  $461 \pm 1.7$  Ma and a youngest age of  $449 \pm 1.6$  Ma (Figure 11F). Lutetium zoning from this sample reflects a central peak, and asymmetric increases of lutetium towards each rim (Figure 12E). Despite the relative scatter in the produced isochron, this sample is clearly distinct from the  $\sim 420$ - $430$  Ma ages in the Zaino Complex and Maz shear zone and  $>591$  Ma ages found within the Maz Complex. The garnet age of  $461$ - $449$  Ma is interpreted to bracket peak metamorphism within the Sierras de Ramaditas.

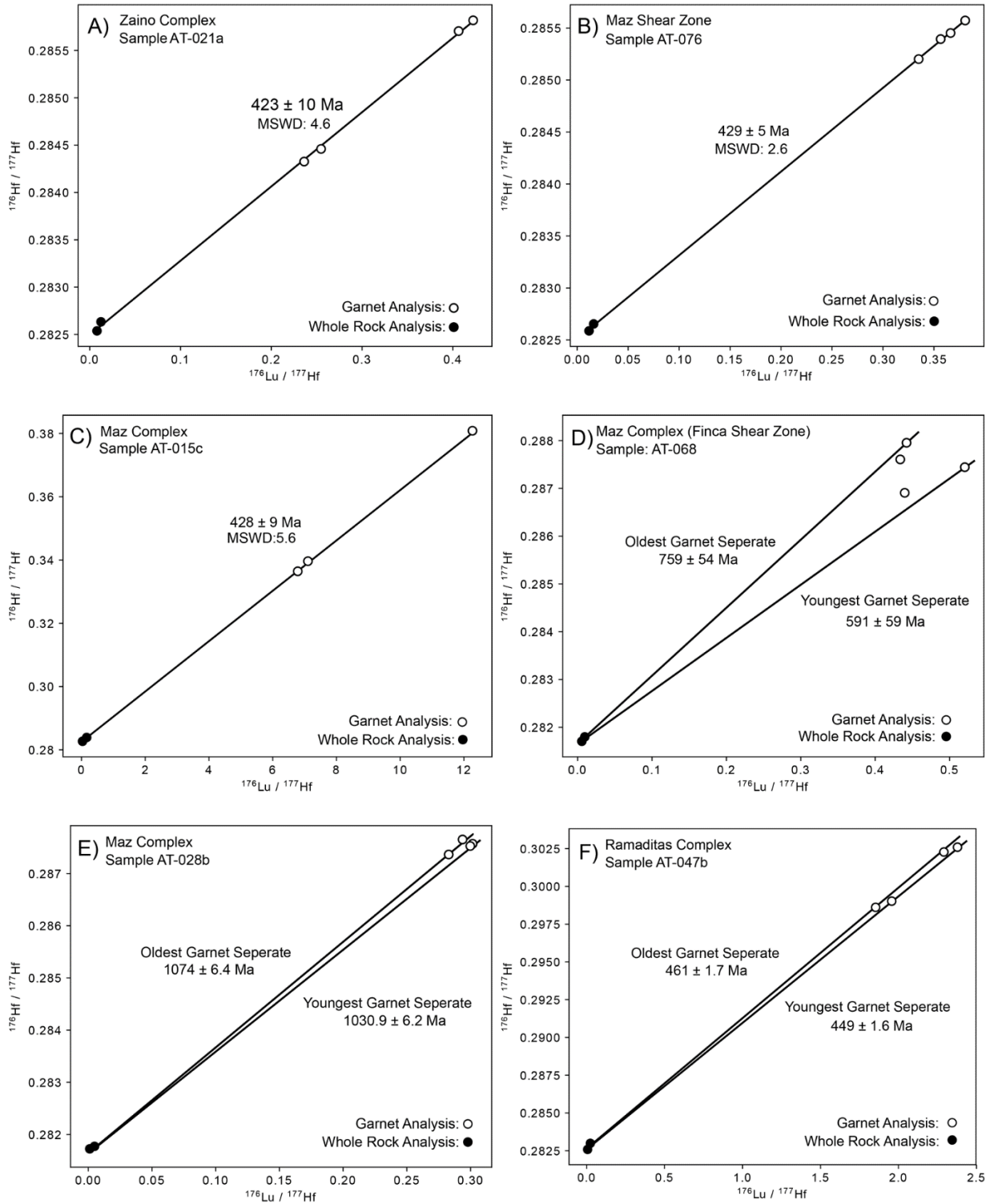


Figure 11: Garnet Lu-Hf ages. A) Lu-Hf isochron for sample AT-021. B) Lu-Hf isochron for sample AT-076. C) Lu-Hf isochron for sample AT-015. D) Scattered Lu-Hf ratios of sample AT-068. E) Scattered Lu-Hf ratios of sample AT-028. F) Scattered Lu-Hf ratios of sample AT-047.

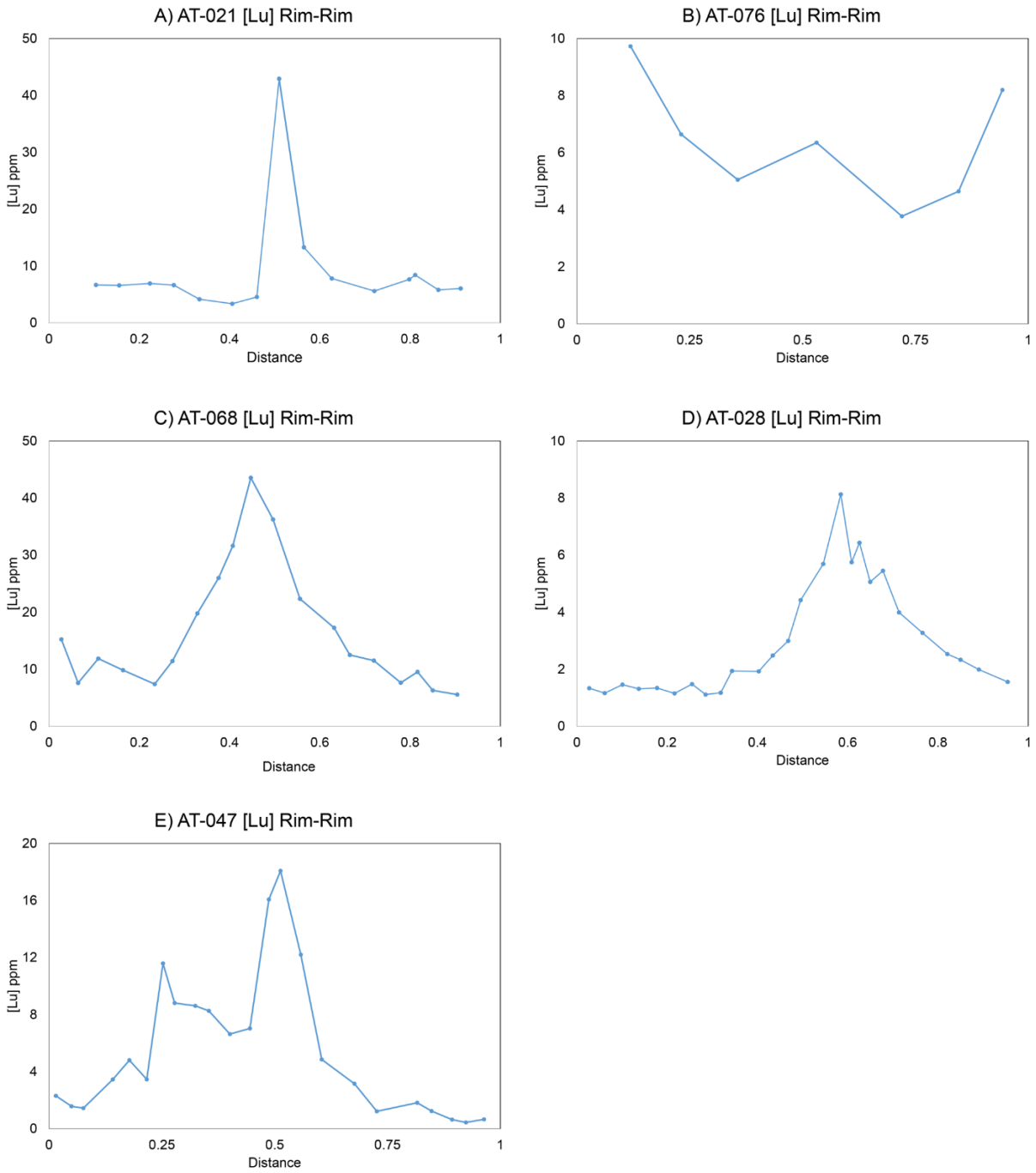


Figure 12: Garnet Lu zoning profiles. A) Lutetium concentration as a function of distance in sample AT-021. B) Lutetium concentration as a function of distance in sample AT-076 C) Lutetium concentration as a function of distance in sample AT-068. D) Lutetium concentration as a function of distance in sample AT-028. E) Lutetium concentration as a function of distance in sample AT-047.

## Monazite U-Pb Geochronology

In order to date the timing of deformation related to the Maz shear zone and interpret the cause of scatter in garnet Lu-Hf isochrons, monazite U-Th-Pb analysis was conducted on the sensitive high-resolution ion microprobe reverse-geometry (SHRIMP-RG) ion microprobe co-operated by U.S. Geological Survey and Stanford University in the SUMAC facility at Stanford University. Analyses of monazite were performed in-situ on cut pieces of polished thin sections, which were trimmed with a diamond wire saw and embedded in epoxy. Secondary ions were sputtered from the target spot using an  $O_2^-$  primary ion beam, which was accelerated at 10 kV and had an intensity varying from 2.2 to 3.1 nA. The analytical sputter pit was  $\sim 18 \times 20 \mu\text{m}$  in diameter and a depth of  $\sim 2\text{--}3 \mu\text{m}$ . Monazite trace element (Y, REE, U, Th) concentrations were calculated relative to standard 44069, which was calibrated based on the trace element concentration standard NAM monazite (Omarura, Namibia; Aleinikoff et al., 2012). Calculated model ages for monazite were standardized relative to standard 44069 ( $^{206}\text{Pb}/^{238}\text{U}$  age = 424 Ma; Aleinikoff et al., 2006), which were analyzed repeatedly throughout the duration of the analytical session. Data reduction for geochronology follows the methods described by Ireland and Williams (2003) using the MS Excel add-in programs Squid 2.51 and Isoplot3.76 of Ken Ludwig (2009; 2012). The measured  $^{206}\text{Pb}/^{238}\text{U}$  was corrected for common Pb using  $^{207}\text{Pb}$ , whereas  $^{207}\text{Pb}/^{206}\text{Pb}$  was corrected using  $^{204}\text{Pb}$  and a model Pb composition from Stacey and Kramers (1975). No additional error was propagated for the uncertainty in the common Pb composition. All reported  $^{206}\text{Pb}/^{238}\text{U}$  and  $^{207}\text{Pb}/^{206}\text{Pb}$  model ages and uncertainties ( $2\sigma$ ) include error summed in quadrature from the external reproducibility ( $1\sigma$  SD) of the standard 44069 during an individual analytical

session (24 hours). Concentrations and isotopic data for monazite U – Pb analyses are listed in Table 5

### **Maz Shear Zone**

Monazite in sample AT-076 from within the Maz shear zone occurs as elongate grains within the mylonitic foliation and no monazite was observed as inclusions within garnet. Back-scatter electron (BSE) imaging of monazite (Figure 13) shows patchy and somewhat irregular zoning of BSE-light and BSE-dark regions. Despite zoning irregularities, populations of monazite were indistinguishable based on rare earth element concentrations (Figure 14A). Fourteen analyses produced  $^{207}\text{Pb}$  corrected  $^{206}\text{Pb}/^{238}\text{U}$  ages ranging from ~402 to ~428 Ma with no significant age or trace element differences among regions of differing back-scatter intensity. The concordia age resulted in  $421 \pm 4.2$  Ma (MSWD = 1.03) while the inverse concordia age gave  $435 \pm 31$  Ma (MSWD = 4.0) with an overall weighted mean  $^{206}\text{Pb}/^{238}\text{U}$  age of  $418.7 \pm 3.3$  Ma (MSWD = 1.14) (Figure 14). The weighted mean age is interpreted to date growth of monazite within the foliation of the Maz shear after initial garnet growth and during deformation.

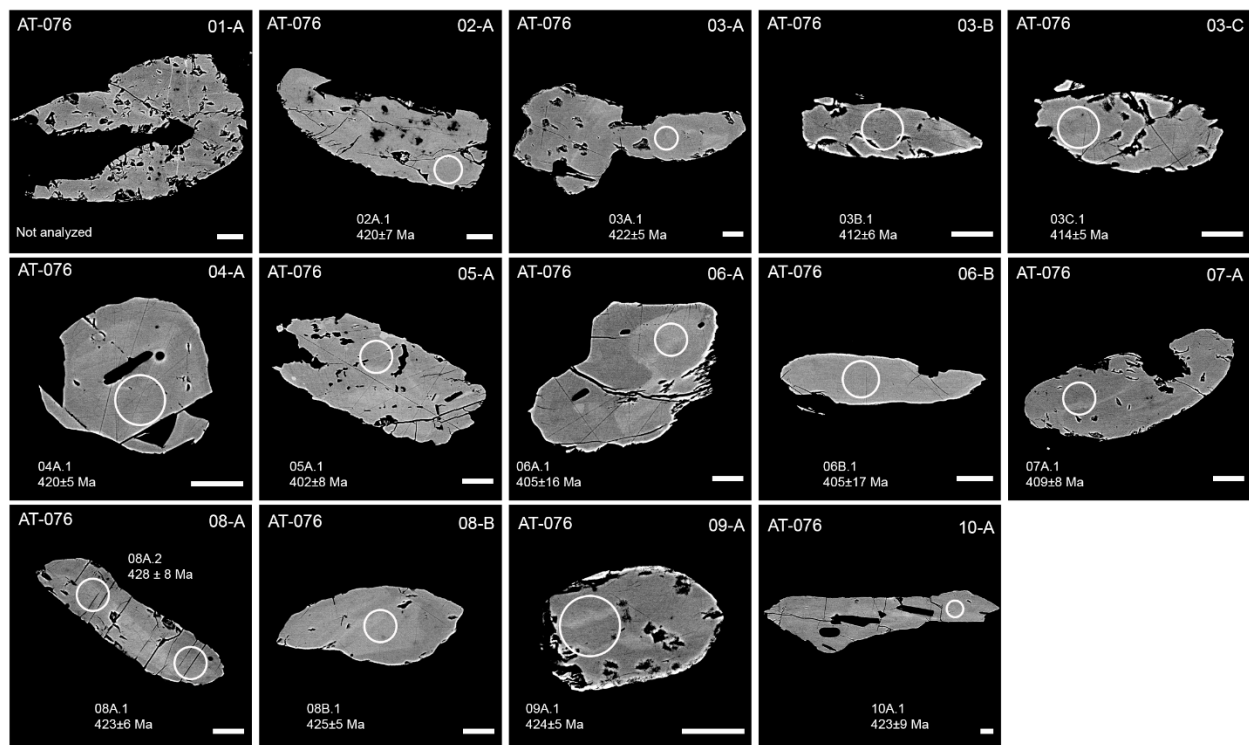


Figure 13: Back scatter electron images of analyzed monazite from sample AT-076 from the Maz shear zone. Circles outline ~20  $\mu\text{m}$  analytical spots. White bar at lower right corner is 10  $\mu\text{m}$ .



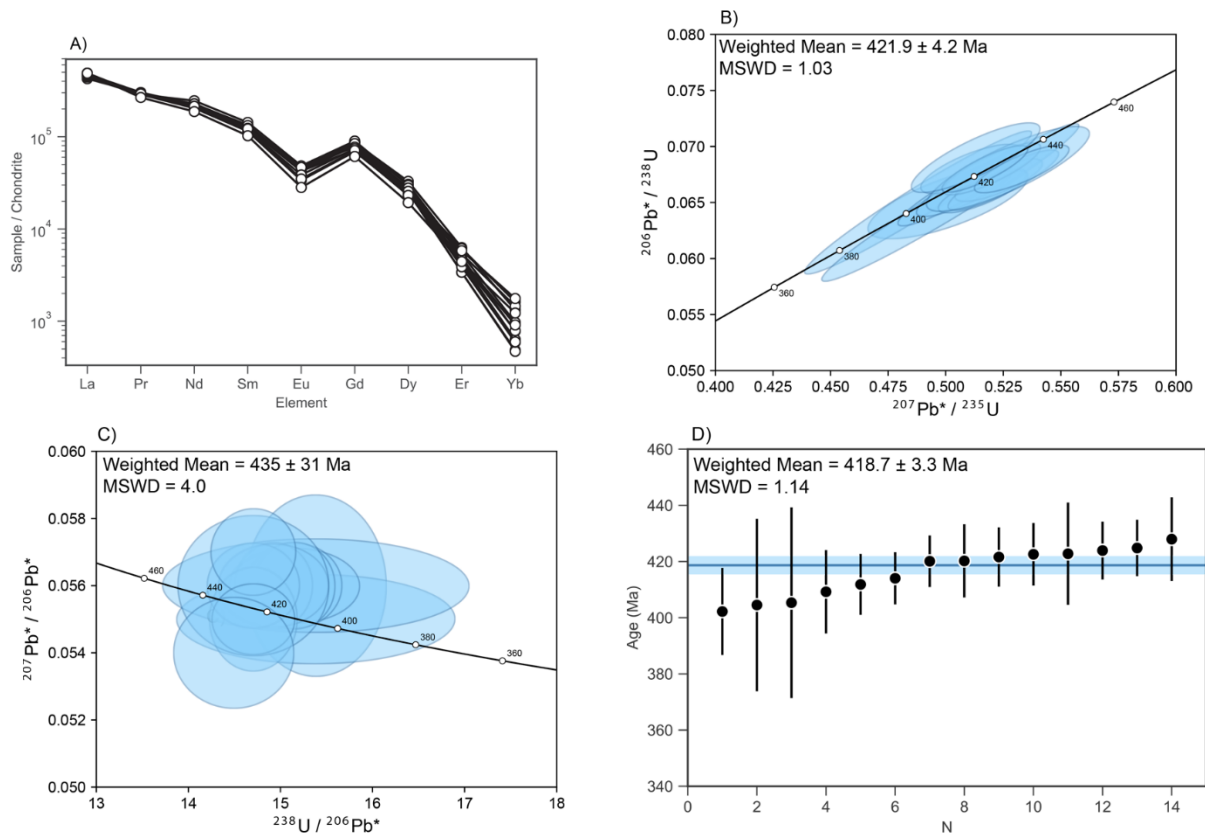


Figure 14: Monazite geochronology of sample AT-076 from the Maz shear zone. A) REE distributions of analyzed spots normalized to chondritic compositions. B) Concordia diagram with calculated age of  $421 \pm 4.2$  Ma ( $\text{MSWD} = 1.03$ ). C) Inverse concordia diagram with fitted age of  $435 \pm 31$  Ma ( $\text{MSWD} = 4.0$ ). D) Analyzed ages showing  $2\sigma$  error bars and a weighted mean age of  $418.7 \pm 3.3$  Ma ( $\text{MSWD} = 1.14$ ). Open circles represent ages excluded from the weighted mean calculation while closed circles indicate ages included in the weighted mean calculation.

## Maz Complex

Monazite from sample AT-068 of the Finca shear zone typically occurs as elongate grains within the mylonitic foliation and as inclusions in both garnet and staurolite. Back scatter electron imaging of monazites (Figure 15) reveal sector-zoned, and patchy regions of BSE-light and BSE-dark regions. Monazite are indistinguishable in rare-earth concentrations barring a single analysis from a monazite inclusion within an unfractured region in garnet. Twenty-two analyses of matrix monazite and monazite inclusions along fractures in garnet and staurolite produced ages ranging between ~381 and ~428 Ma, with weighted mean  $^{206}\text{Pb}/^{238}\text{U}$  age of  $410.6 \pm 2.4$  Ma (MSWD = 1.57) (Figure 16). A single grain of monazite in garnet not associated with fractures gave a  $^{204}\text{Pb}$  corrected  $^{207}\text{Pb}/^{206}\text{Pb}$  age of  $1149 \pm 12$  Ma. A regression line plotted through matrix grains, inclusionary grains along fractures, and the discordant grain has an upper intercept of  $1219 \pm 88$  Ma and a lower intercept of  $404 \pm 20$  Ma (MSWD = 2.2). The upper intercept is interpreted to represent the date monazite growth during initial metamorphism. The weighted mean age of matrix monazite is interpreted to date of monazite growth during deformation in the Finca shear zone.

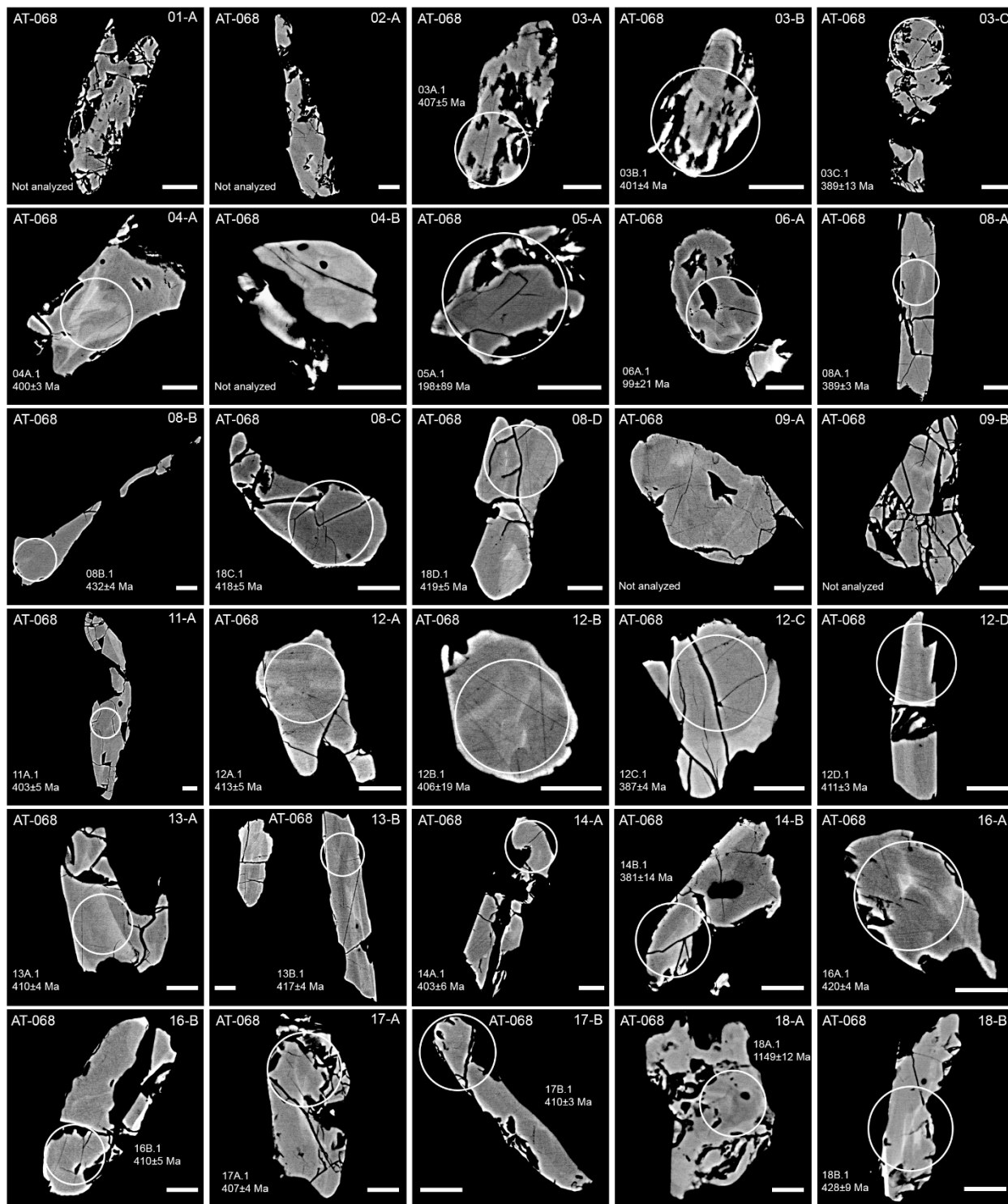


Figure 15: Back scatter electron images of analyzed monazite from sample AT-068 from the Finca shear zone. Circles outline ~20 μm analytical spots. White bar at lower right corner is 10 μm.

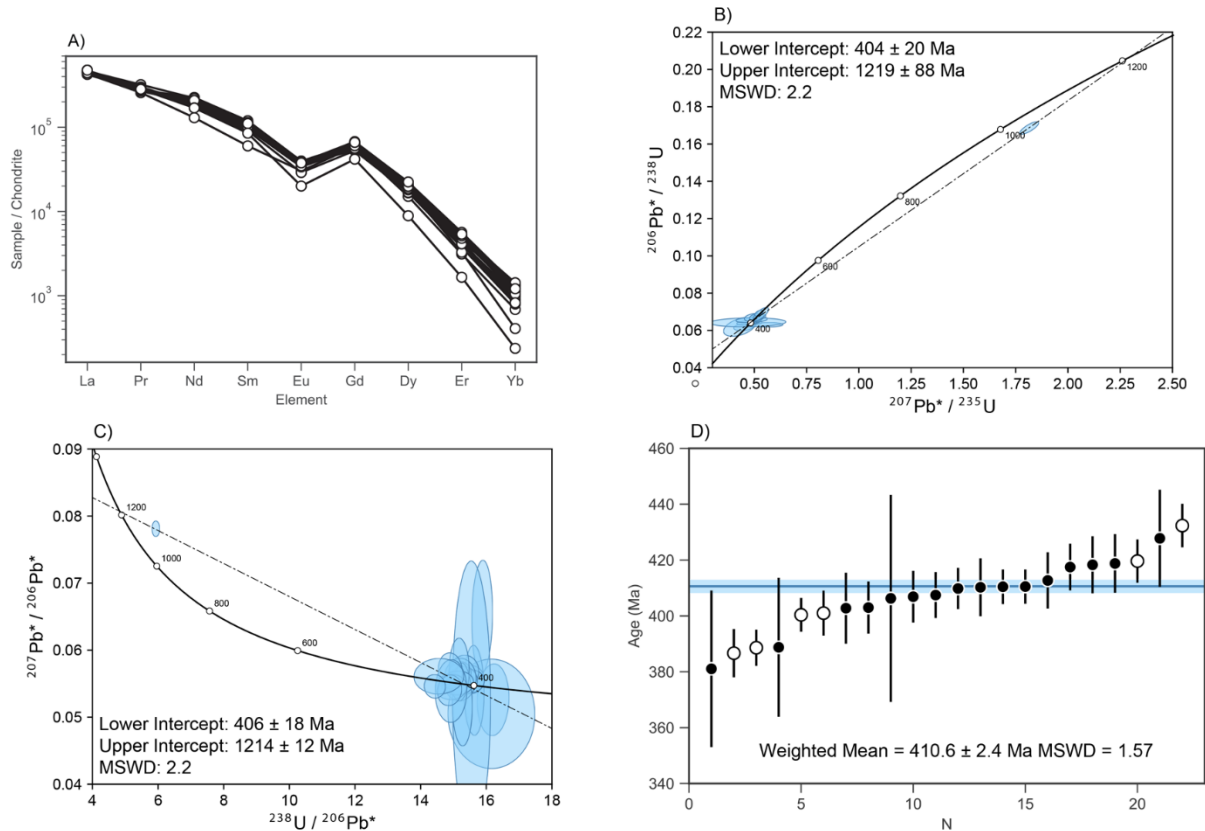


Figure 16: Monazite geochronology of sample AT-068 from the Finca shear zone. A) REE distributions of analyzed spots from sample AT-068 normalized to chondritic compositions. B) Concordia diagram with calculated upper intercept age of  $1219 \pm 88$  Ma and lower intercept age of  $404 \pm 20$  Ma (MSWD = 2.2). C) Inverse concordia diagram calculated upper intercept age of  $1214 \pm 12$  Ma and lower intercept age of  $406 \pm 18$  Ma (MSWD = 2.2) D) Analyzed ages showing  $2\sigma$  error bars and a weighted mean age of  $410.6 \pm 2.4$  Ma (MSWD = 1.57). Open circles represent ages excluded from the weighted mean calculation while closed circles indicate ages included in the weighted mean calculation.

Monazite from sample AT-028 occurs as elongate grains within the matrix of the foliation and as coarse inclusions in garnet up to 200 microns in length. Back-scatter electron imaging of monazite (Figure 17) show variable zoning patterns that can be broadly grouped into three main categories: 1) bright, oscillatory-zoned cores, 2) patchy overgrowths commonly forming around cores, 3) dark, unzoned asymmetrical rims on matrix grains. Two general populations can be identified through rare-earth element

concentrations of distinctly different enrichments and a continuous range that lies between these two (Figure 18A). Twenty-five analyses from nineteen grains with zoned cores produced concordant  $^{204}\text{Pb}$  corrected  $^{207}\text{Pb}/^{206}\text{Pb}$  ages that range from ~1191 to ~1262 Ma. The concordia plot resulted in an upper intercept age of  $1236 \pm 18$  Ma (MSWD = 1.8), while the inverse concordia plot gave an upper intercept of  $1230 \pm 21$  Ma (NSWD = 1.8). The weighted mean age for this population is  $1227.1 \pm 6.8$  Ma (MSWD = 1.55) (Figure 18). Dark, homogenous rims on matrix grains resulted in concordant  $^{207}\text{Pb}$  corrected  $^{206}\text{Pb}/^{238}\text{U}$  ages ranging from ~379 to ~424 Ma. The concordia plot has a lower intercept age of  $415 \pm 33$  Ma (MSWD = 1.8) while the inverse concordia plot has a lower intercept of  $404 \pm 16$  Ma (MSWD = 1.8). The weighted mean age of this population is  $413.2 \pm 5.0$  Ma (MSWD = 7.68) (Figure 18) with rim analyses exhibiting higher REE concentrations compared to core populations. Patchy recrystallized/overgrowth textures produced discordant ages that as young as ~900 Ma. The resultant ages and corresponding textures suggest that at least two distinct populations of monazite exist within sample AT-028 (Figure 18). Oscillatory zoning is commonly observed in igneous monazite (Engi, 2017), and thus the weighted mean age of that texture could represent an igneous formation age. The weighted mean for younger, concordant rim ages are interpreted to represent growth of metamorphic monazite in the matrix during shear within the Sierra de Maz.

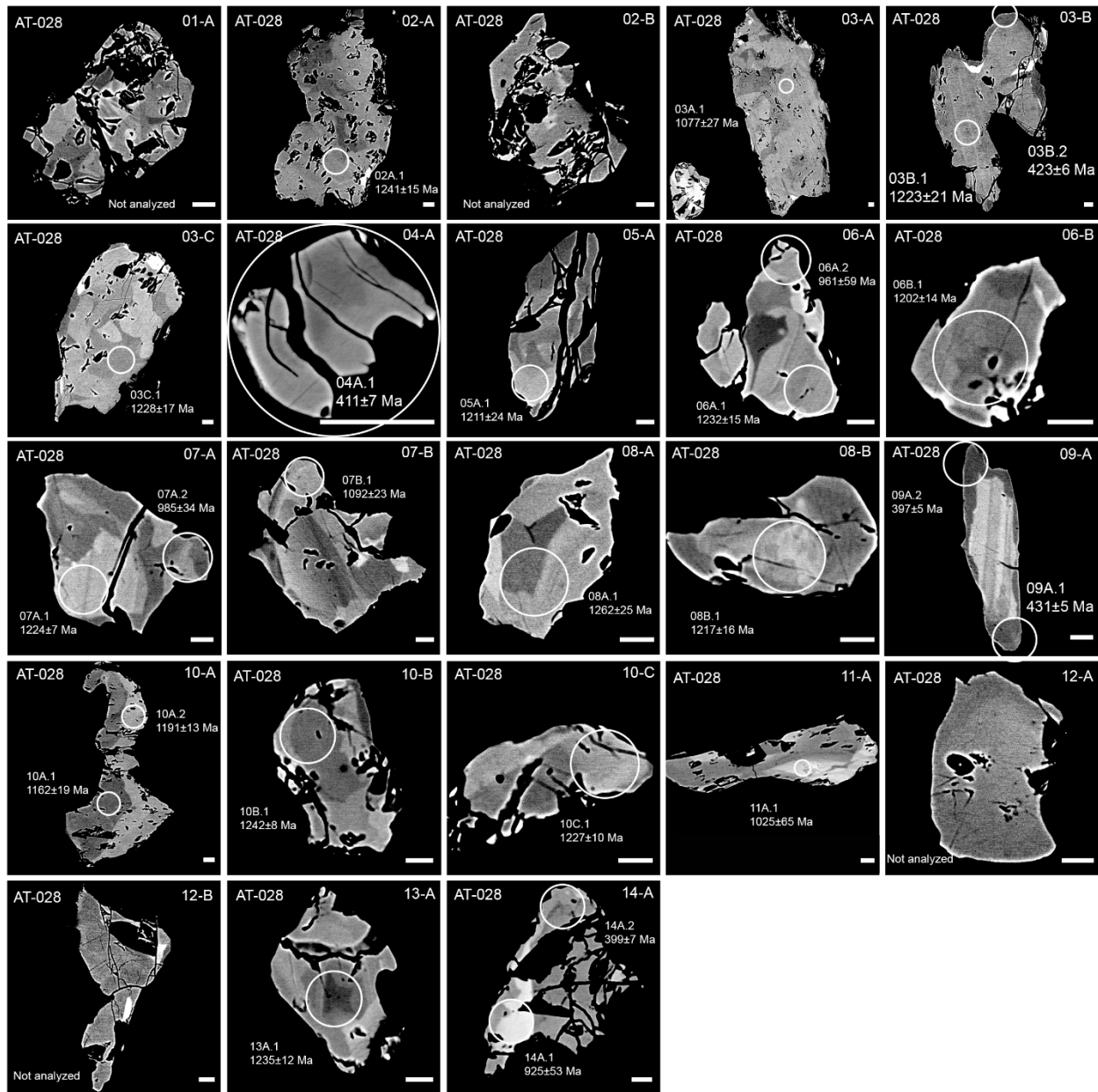


Figure 17: Back scatter electron images of analyzed monazite from sample AT-028 within the Maz Complex. Circles outline ~20  $\mu\text{m}$  analytical spots. White bar at lower right corner is 10  $\mu\text{m}$ .

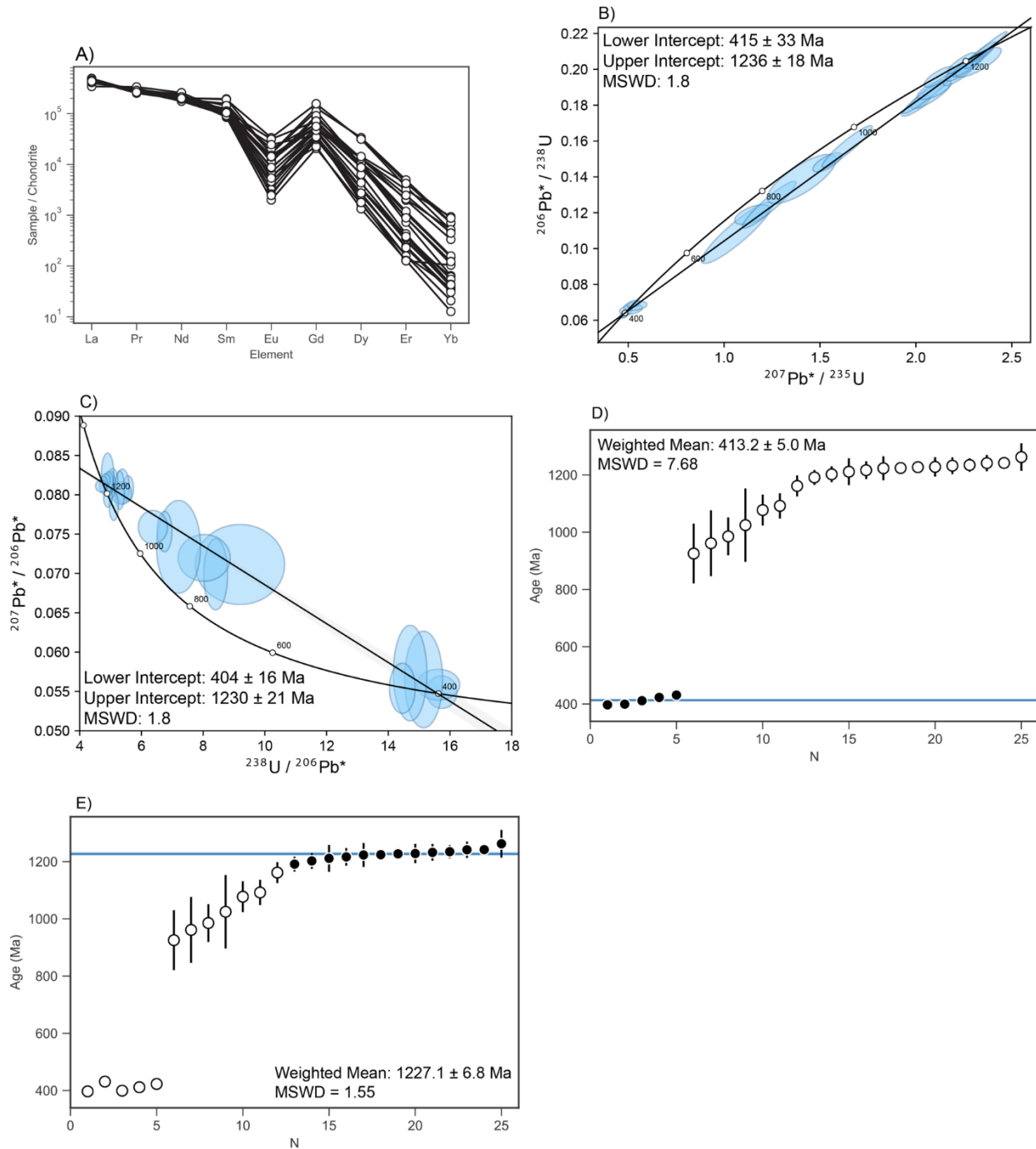


Figure 18: Monazite geochronology of sample AT-028 from the Maz Complex. A) REE distributions of analyzed spots from sample AT-028 normalized to chondritic compositions. B) Concordia diagram with calculated upper intercept age of  $1236 \pm 18$  Ma and lower intercept age of  $415 \pm 33$  Ma (MSWD = 1.8). C) Inverse concordia diagram calculated upper intercept age of  $1230 \pm 21$  Ma and lower intercept age of  $404 \pm 16$  Ma (MSWD = 1.8) D) Analyzed ages showing  $2\sigma$  error bars and a distinct monazite population with a weighted mean age of  $413.6 \pm 5.0$  Ma (MSWD = 7.68). E) Analyzed ages showing  $2\sigma$  error bars and a distinct monazite population with a weighted mean age of  $1227.1 \pm 6.8$  Ma (MSWD = 1.55). Open circles represent ages excluded from the weighted mean calculation while closed circles indicate ages included in the weighted mean calculation.

## Ramaditas Complex

Monazite from sample AT-047 of the Ramaditas Complex occurs as rounded to elongate grains within the matrix foliation and as anhedral inclusions in garnet. Monazite shows patchy and irregular zoning of BSE-light and BSE-dark regions (Figure 19). A general spread in the rare-earth element concentrations does not correlate with age. Fifteen grains produced concordant ages that range from ~418 and ~472 Ma. The concordia age for this sample is  $453.1 \pm 6.5$  MA (MSWD = 1.0) while the inverse concordia age is  $453.2 \pm 6.4$  Ma (MSWD = 1.0) (Figure 20). Texturally distinct monazite cores gave a  $^{207}\text{Pb}$  corrected  $^{206}\text{Pb}/^{238}\text{U}$  weighted mean of  $461.6 \pm 3.1$  Ma (MSWD = 1.6) and the monazite rims gave a  $^{207}\text{Pb}$  corrected  $^{206}\text{Pb}/^{238}\text{U}$  weighted mean of  $426.1 \pm 7.0$  Ma (MSWD = 1.02) (Figure 20). The weighted mean age of the older population is interpreted to represent growth of metamorphic monazite during peak granulite facies. The weighted mean age of the younger population could be interpreted as metamorphism and deformation associated with the Ramaditas shear zone.

## Summary

Monazite ages from three samples within the Sierra de Maz suggests that the deformation related to the Maz shear zone resulted in monazite growth across the range at ~415 Ma. The majority of monazite ages from within the Sierra de Ramaditas do not overlap within uncertainty of monazite ages found within the Sierra de Maz. Proterozoic monazite was analyzed in two samples from within the Sierra de Maz and are absent from rocks of the Sierra de Ramaditas. A summary table of all ages presented in this study can be found in Table 6.



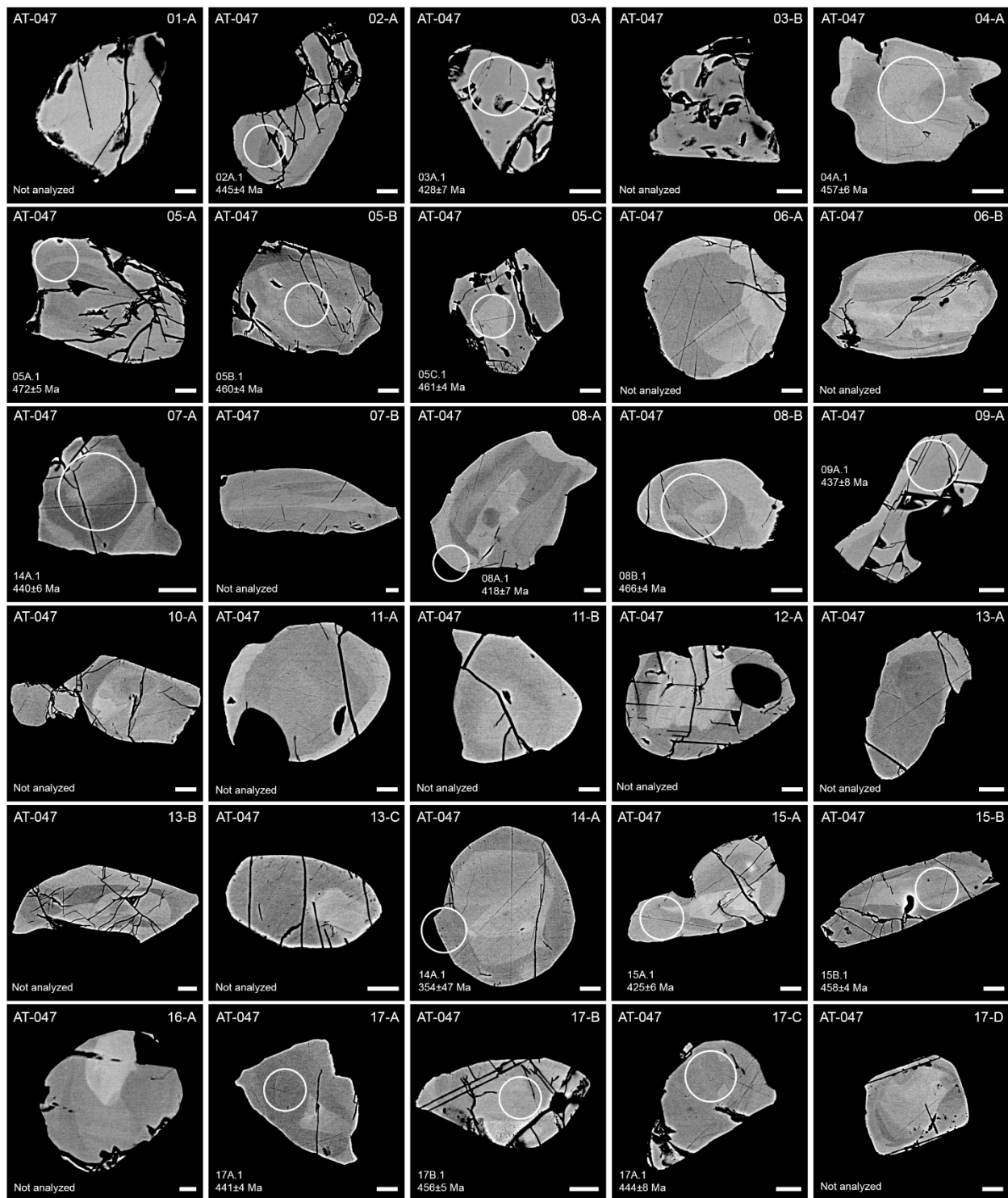


Figure 19: Back scatter electron images of analyzed monazite from sample AT-047 with white circles outlining ~20 μm diameter analysis spot. White bar at lower right corner is 10 μm.

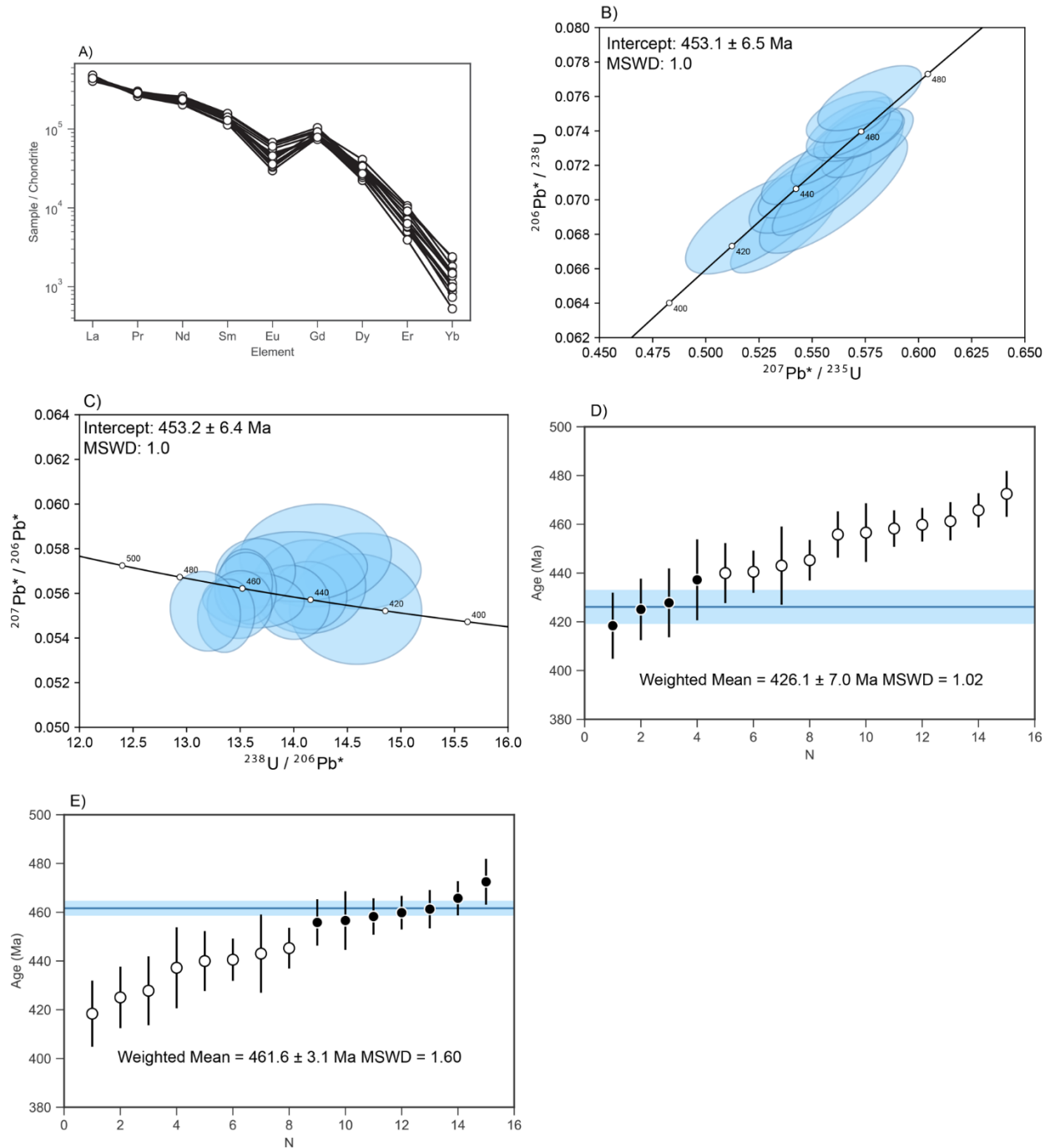


Figure 20: A) REE distributions of analyzed spots from sample AT-047 normalized to chondritic compositions. B) Concordia diagram with calculated age of  $453.1 \pm 6.5$  Ma (MSWD = 1.0). C) Inverse concordia diagram with fitted age of  $453 \pm 6.4$  Ma (MSWD = 1.0). D) Analyzed ages showing  $2\sigma$  error bars and a monazite population with a weighted mean age of  $426.1 \pm 7.0$  Ma (MSWD = 1.02). E) Analyzed ages showing  $2\sigma$  error bars and a monazite population with a weighted mean age of  $461.1 \pm 3.1$  Ma (MSWD = 1.60). Open circles represent ages excluded from the weighted mean calculation while closed circles indicate ages included in the weighted mean calculation.

## Discussion

Data from this study stand in contrast to the geotectonic model for the MARA terrane proposed by Rapela et al. (2016). Direct evidence of a Cambrian orthogonal collision and ~525 Ma partial subduction is still lacking within units of northern Sierra de Maz. Metamorphic ages presented in this study add to the growing dataset of ages that are not explained by the Rapela (2016) model. Silurian titanite ages within the Sierra de Maz were reported by Luccassen and Becchio (2003) and by Casquet et al. (2006), but were not related to a metamorphic event or deformation fabric. This study dates Silurian-Devonian deformation and metamorphism within the Sierra de Maz and allows for direct comparisons to be made between the various ranges of the southern MARA terrane.

Peak metamorphic conditions of Zaino, Maz, and Ramaditas metapelites are distinctly different according to observed assemblages, P-T approximations, and garnet zoning profiles. The Zaino Complex contains the lowest metamorphic grades in the Sierra de Maz with the dominant assemblage of Grt-Bt-Ms-Plg-Qtz (Figure 7C) and rare staurolite-bearing metapelites in the most eastern outcrops (Figure 7D). Pressure and temperature estimates from this unit range within uncertainty of the amphibolite to upper amphibolite facies and reflect the west to east increase in metamorphic grade suggested by pelitic index minerals (Figure 4). Post kinematic, prograde growth-zoned garnet (Figure 8A and Figure 9A) within Zaino peak assemblages suggest a single phase of metamorphism. Conversely, metapelites from the Maz Complex show evidence for multiple metamorphic events at higher temperatures and pressures than the Zaino Complex. A west to east increase in metamorphic grade within the Maz

Complex is apparent in pelitic index minerals, with temperatures and pressure estimates starting in the upper amphibolite facies with the assemblage Grt-St-Ms-Plg-Bt  $\pm$  Ky and ending in the granulite facies with the assemblage Grt-Bt-Plg-Kfs-Sil. The presence of pre-kinematic garnets (Figure 7G) and complicated zoning profiles from samples AT-068 (Figure 8C) and AT-028 (Figure 9D) indicate the unit experienced more than one metamorphic event. Kyanite inclusions in garnet from an eastern granulite sample AT-028 (Figure 7I) that contains sillimanite in the matrix implies that granulite conditions were attained during the most recent foliation-forming event. Metapelites from the Ramaditas Complex bear evidence for the lowest metamorphic pressures of the three studied units. The assemblage Cd-Grt-Bt-Plg-Qtz qualitatively suggests that peak metamorphic conditions were at least 2 kbars lower than the adjacent eastern units of the Maz Complex. Pressure and temperature calculations from Ramaditas sample AT-047 confirm that metamorphic pressures do not overlap within uncertainty with any other analysis within the Sierra de Maz (Figure 7). Furthermore, garnet from the Sierra de Ramaditas are diffusively zoned (Figure 9E) indicating a single event of slow cooling from high-T conditions, unlike garnet from the Sierra de Maz.

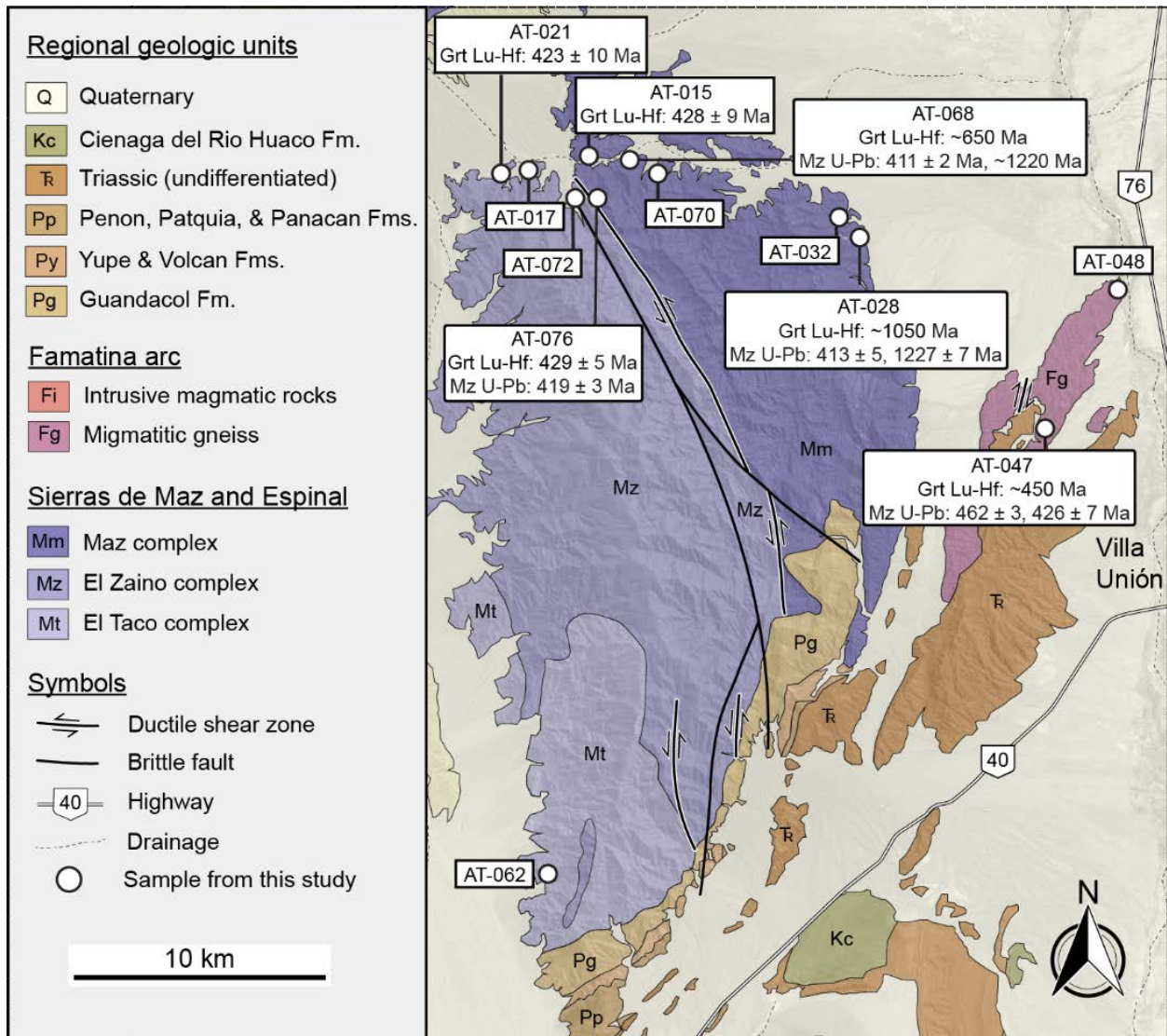


Figure 21: Geologic map of the Sierra de Maz with the locations and ages of samples dated in this study.

The distribution of metamorphic ages is variable among the Zaino, Maz, and Ramaditas Complexes. Of the three, the Zaino Complex has the simplest history with a single age of peak metamorphism inferred from the garnet Lu-Hf age of  $423 \pm 10$  Ma (MSWD = 4.6) from sample AT-021 (Figure 21). This garnet crystallization event overlaps within uncertainty of two additional garnet Lu-Hf ages from within the Maz shear zone of  $429 \pm 5$  Ma (MSWD = 2.6) and  $428 \pm 9$  Ma (MSWD = 5.6). When taken together, three similar ages of garnet growth are interpreted to record burial and metamorphism of the Zaino complex during oblique thrusting along the Maz shear zone.

The Maz Complex contains an older history of metamorphism not found within the Zaino Complex. Proterozoic garnet growth is inferred from the scattered isochron plot of sample AT-028 (Figure 11E). This older history is also evident in a U-Pb weighted mean age of  $1227.1 \pm 6.8$  Ma from monazite inclusions within garnets of the Maz Complex and overlaps with U-Pb ages from metamorphic overgrowths on detrital zircons first identified by Casquet et al. (2006). The event recorded by these two phases places rough constraints on metamorphism of rocks comprising the Maz Complex. The Maz Complex subsequently experienced thermal events from  $\sim 1.0$  Ga anorthosite intrusions, at  $\sim 845$  Ma and  $\sim 774$  Ma by granitoid intrusions, and at  $\sim 570$  Ma from carbonatite intrusions, all of which are currently absent from the Zaino and Ramaditas Complexes (Casquet et al., 2008a; Casquet et al., 2008b; Colombo et al., 2009; Rapela et al., 2010).

Ordovician metamorphic ages from the Ramaditas Complex are unique within the study area. Garnet Lu-Hf isochrons from sample AT-047 resulted in an older three-point isochron age of  $461 \pm 1.7$  Ma and a younger three-point isochron of  $449 \pm 1.6$  Ma.

While the validity of three-point isochrons may be suspect, when all analyses are taken together, the age of garnet growth for this region appears to be at least 30 my older than shear related metamorphism in the Sierra de Maz. Ordovician signatures are also evident in the U-Pb ages of matrix monazite from sample AT-047 with concordia and Terra-Wasserburg calculated ages of  $453.1 \pm 6.5$  Ma (MSWD = 1.0) (Figure 20B) and  $453.2 \pm 6.4$  Ma (MSWD = 1.0) (Figure 20C) respectively. Despite the multitude of thermal events experienced by the Maz Complex, Ramaditas is unique in its dominantly Ordovician signature despite the presence of Silurian monazite. Ordovician metamorphic ages, diffusively zoned garnets, and low pressure pelitic assemblages imply the Ramaditas Complex most likely recorded contact metamorphism related to Famatina arc magmatism.

Silurian - Devonian metamorphism observed within all of the units in the Sierra de Maz was a result of transpressive deformation along the Maz shear zone. Similar garnet and monazite ages in the Zaino Complex, Maz shear zone, Finca shear zone, and Maz Complex imply the inverted metamorphic gradient formed during deformation associated with Maz shear zone. These age interpretations are supported by field-based evidence of similar foliation orientations and sinistral-reverse kinematics. Thus, it can be assumed that the west to east inverted metamorphic sequence is direct result of transpression. Thermal modeling conducted by Thompson et al. (1997) demonstrated that low angles ( $\leq 30^\circ$ ) of oblique convergence (low ratios of pure/simple shear) can cause Barrovian to granulite conditions in transpressive orogens with long-lived periods of high temperatures. Shallow lineations and steeply dipping foliations across the Sierra de Maz indicate that the angle of obliquity was, on average, less than  $30^\circ$  within the

Maz shear zone (Figure 4) and within the geometric bounds proposed by Thompson et al., (1997). An unpublished Ar-Ar white mica cooling age of  $\sim 386.5 \pm 1.0$  Ma (E. Houlihan, written communication, 2018) from directly within the Maz shear zone implies temperatures stayed above mica closure for at least 50 my following garnet growth, consistent with the long-lived transpressive thermal regimes predicted by Thompson et al. (1997). The extrusional dynamics presented by Thompson et al. (1997) could explain a nominally younger age of  $410.6 \pm 2.4$  Ma from sample AT-068 suggesting that this area could have acted as a weak zone accommodating later shortening.

Similar metamorphic PT trends as a function of structural distance (Figure 22) have been observed in other transpressive orogens such as Kaoko Belt in Namibia and the Kalinjala Shear Zone in South Australia (Goscombe and Gray, 2009), but are slightly different than what is seen in the Sierra de Maz. In the Namibian orogen, granulite conditions are seen to exist within the orogen-scale median shear zone and diminish away from the shear zone in both directions. Within the northern Sierra de Maz, metamorphic conditions between major blocks (Maz shear zone) are calculated to have reached upper amphibolite facies with granulite conditions several kilometers east. This difference could be a preservation problem, with the true orogen median shear zone being just east of the Sierra de Maz, lost to erosion or possibly buried under sediments.



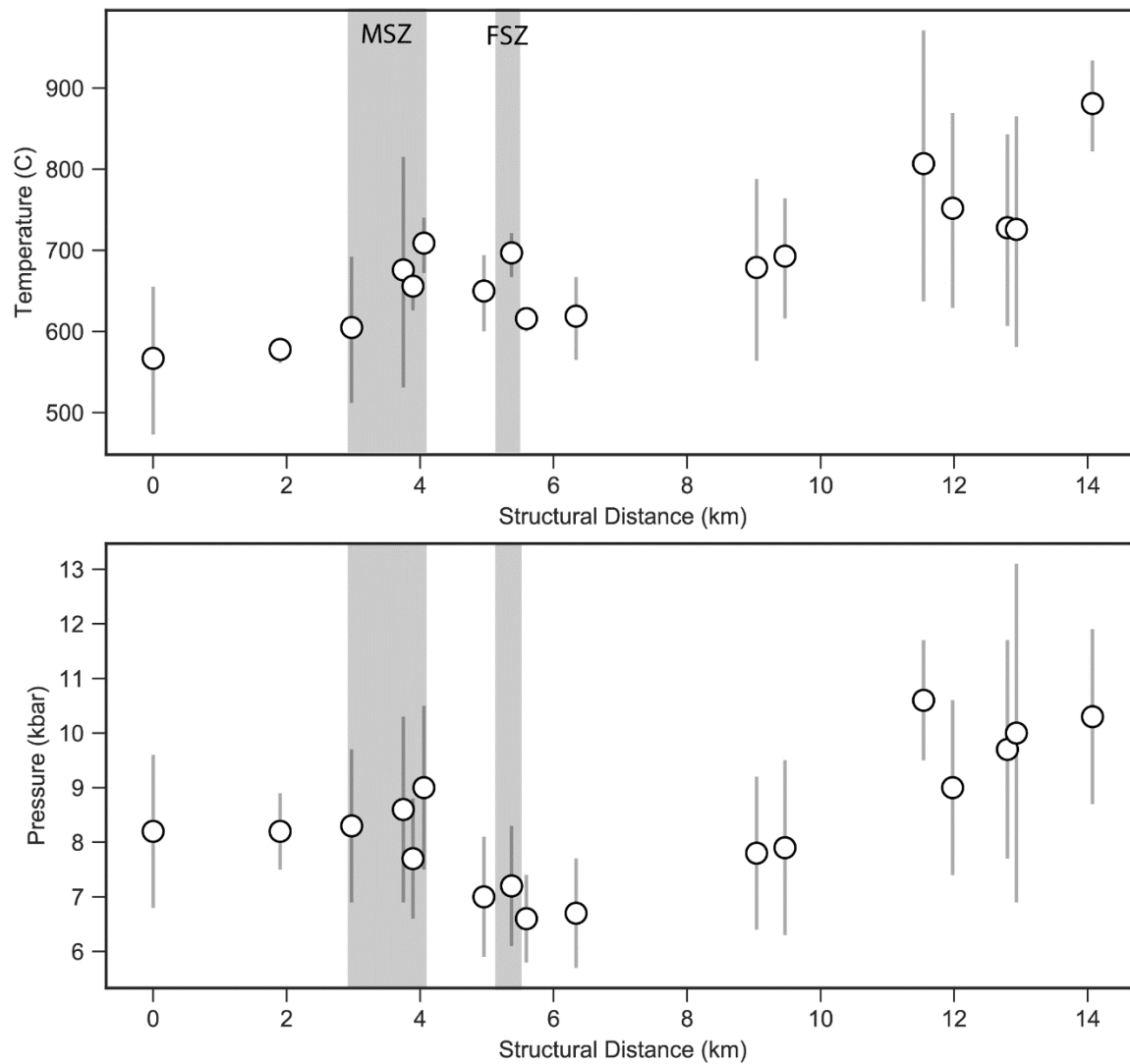


Figure 22: Temperature and pressure across the Sierra de Maz (MSZ = Maz shear zone; FSZ = Finsa shear zone).

Metamorphic ages from the Zaino and Maz Complexes do not support correlations with other proposed MARA units in the region (Figure 1). The Taco Complex in the southwest Sierra de Maz (Figure 3) contains unpublished metamorphic ages spanning 470 - 440 Ma (P. Webber, written communication, 2018). To the northwest within the Sierra de Umango, ~452 Ma monazite, ~446 Ma metamorphic zircon overgrowths (Varela et al., 2010) and ~440 Ma titanite (Luccasen & Becchio, 2003) are different from ages found within the Zaino and Maz Complexes. To the south, exposures in the Sierra de Pie de Palo (Figure 1) record oblique thrusting along low-angle shear zones between ~515 Ma and ~440 Ma (Mulcahy et al., 2011). The age data require that the Zaino and Maz Complexes of the Sierra de Maz do not share the same metamorphic histories preserved in the Sierra de Pie de Palo, the Sierra de Umango, or the Taco Complex. Continued work uncovering metamorphic cooling ages within the Taco Complex, the Sierra de Umango, and the Pie de Palo could provide the means to strengthen correlations between these ranges.

Considering the variable histories within the southern ranges of the MARA terrane that are unaccounted for in the Rapela et al. (2016) model, an additional model that accounts for Devonian-Silurian transpression must be proposed. Given the current distribution of ages, it must be assumed that the Sierra de Maz was at one time spatially isolated from its present-day neighboring ranges. The southern ranges of the MARA terrane could represent composite terranes, with site-specific accretion and translation histories that span from the Cambrian through the Late Devonian. On the basis of shared metamorphic ages, this study speculates that the Taco Complex, the Sierra de Umango, and the Sierra de Pie de Palo were more or less amalgamated

during the Ordovician. The absence of Ordovician Famatina ages within the Sierra de Maz suggests that these rocks were far enough away or at an isolated structural level to avoid significant arc interactions. Margin-scale transpression during the Silurian juxtaposed the Zaino Complex against the Maz Complex and deformed the contact aureole (Ramaditas Complex) of the Famatina arc. Transpression continued and translated the Sierra de Maz into its current position between ranges that experienced an Ordovician Famatinian history (Figure 23). Although the contact between the Zaino and Taco Complexes was not directly observed in this study, this model predicts a low angle thrust fault that is younger than the Silurian - Late Devonian deformation in the Zaino Complex. Additionally, this model predicts right-lateral faulting within the Sierra de Umango which is, at first order, supported by Meira et al. (2012) who documented right-lateral reverse-transpression with the Sierra de Umagno in the Ordovician and Silurian.

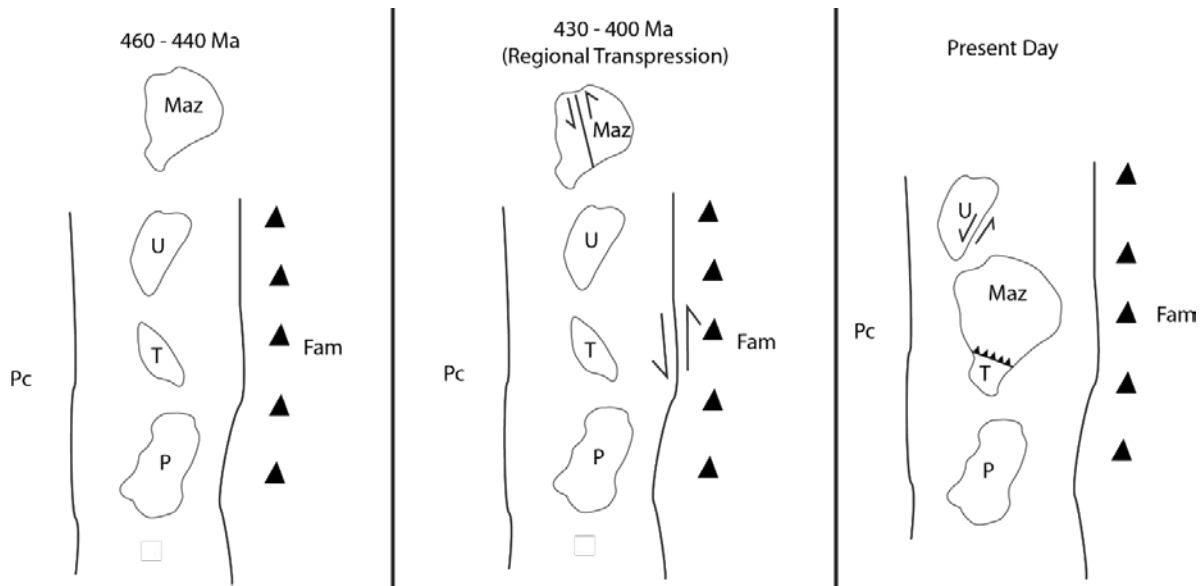


Figure 23: Map-view graphic of southern ranges of the MARA terrane depicting a simplified model for transpressional deformation and displacement of the Sierra de Maz (Pc = Precorderilla, Fam = Famatina arc, P = Pie de Palo, T = Taco Complex, U = Sierra de Umango).

## Conclusions

Metapelitic rocks from the Sierra de Maz contain evidence for regional transpression throughout the Silurian and Early Devonian. Metamorphic crystallization related to this event bears no Ordovician Famatinian age signature, is apparent in fabrics across the entire range, and overprints older events that are preserved in relict metapelitic assemblages of the Maz Complex. Garnet and a majority of monazite growth in the Sierra de Ramaditas are coeval with Famatinian arc magmatism and exhibit peak metamorphic conditions at distinctly lower pressures when compared to the Sierra de Maz. Data from this study imply that Silurian-Devonian sinistral transpression was a major accretion process that shaped the MARA terrane and that continued deformation juxtaposed the Sierra de Maz adjacent to rocks recording Famatinian histories.

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**Table 1: Sample Locations**

Sample	Latitude	Longitude	Sample	Latitude	Longitude
AT-004	-29.056	-68.782	AT-068	-29.186	-68.457
AT-005	-29.04	-68.783	AT-070	-29.187	-68.455
AT-006	-29.041	-68.782	AT-071	-29.193	-68.482
AT-007	-29.109	-68.749	AT-072	-29.192	-68.482
AT-009	-29.188	-68.466	AT-073	-29.195	-68.482
AT-011	-29.188	-68.47	AT-076	-29.201	-68.473
AT-013	-29.189	-68.473	AT-079	-29.188	-68.461
AT-015	-29.189	-68.474	AT-081	-29.187	-68.459
AT-017	-29.191	-68.494	AT-082	-29.186	-68.457
AT-019	-29.189	-68.5	AT-083	-29.188	-68.446
AT-020	-29.19	-68.508	AT-085	-29.19	-68.444
AT-021	-29.193	-68.514	AT-090	-29.19	-68.397
AT-026	-29.221	-68.366	AT-092	-29.19	-68.413
AT-028	-29.222	-68.367	AT-098	-29.194	-68.417
AT-032	-29.214	-68.374	AT-107	-29.197	-68.472
AT-041	-29.212	-68.388	AT-109	-29.201	-68.473
AT-042	-29.214	-68.379	AT-111	-29.193	-68.48
AT-044	-29.214	-68.378	AT-113	-29.199	-68.482
AT-046	-29.285	-68.282	AT-115	-29.21	-68.48
AT-047	-29.287	-68.283	AT-116	-29.195	-68.535
AT-048	-29.234	-68.26	AT-117	-29.195	-68.533
AT-049	-29.2	-68.38	AT-119	-29.194	-68.526
AT-056	-29.205	-68.391	AT-121	-29.157	-68.491
AT-057	-29.205	-68.393	AT-134	-29.158	-68.448
AT-059	-29.467	-68.493	AT-136	-29.156	-68.443
AT-060	-29.466	-68.492	AT-138	-29.158	-68.443
AT-062	-29.466	-68.489	AT-140	-29.16	-68.443
AT-064	-29.189	-68.48	AT-176	-29.403	-68.415
AT-065	-29.188	-68.477	AT-180	-29.403	-68.423
AT-066	-29.188	-68.471	AT-182	-29.401	-68.429

**Table 2: Sample Mineralogy**

Sample	Chl	Bt	Grt	St	Ky	Sil	Crd	Ms	Kfs	Plg	Qtz	Cpx	Amp	Ttn	Tur	Ap	Mnz	Rt	Ilm	Scp	Cal
AT-004b		X	X					X		X	X										
AT-004c		X								X	X		X								
AT-005		X								X	X										
AT-009	X	X	X					X			X										
AT-011		X	X							X	X		X	X	X						
AT-013		X	X					X	X	X	X			X							
AT-015		X	X							X	X		X		X			X			
AT-017	X	X	X					X		X	X								X		
AT-019		X	X					X		X	X							X	X		
AT-020		X	X					X		X	X				X			X			
AT-021a		X	X					X		X	X								X		
AT-021b		X	X					X		X	X							X	X		
AT-021c		X	X					X		X	X							X			
AT-026a		X	X					X		X	X										
AT-026b		X	X					X		X	X		X						X		
AT-028a		X	X			X			X	X	X			X	X		X		X		
AT-030a		X	X							X	X			X					X		
AT-030b		X	X							X	X			X					X		
AT-032a		X	X							X	X		X						X		
AT-032b		X	X							X	X		X	X					X	X	
AT-032c			X										X						X	X	
AT-041	X	X	X							X	X		X						X		
AT-042		X	X							X	X		X						X		X
AT-044		X	X							X	X			X					X		
AR-046		X	X							X	X				X						
AT-047a		X					X			X	X								X		
AT-048a		X	X										X						X		X
AT-048b		X	X			X				X	X	X	X								X
AT-049		X				X					X				X						

**Table 2 (cont.): Sample Mineralogy**

Sample	Chl	Bt	Grt	St	Ky	Sil	Crd	Ms	Kfs	Plg	Qtz	Cpx	Amp	Ttn	Tur	Ap	Mnz	Rt	Ilm	Scp	Cal
AT-056		X	X					X		X	X				X						X
AT-057		X	X							X	X		X		X				X		
AT-058		X	X							X	X		X	X							
AT-059		X	X						X	X	X								X		
AT-060	X	X	X	X				X		X	X			X	X				X		
AT-062								X		X	X								X		
AT-064		X	X						X	X	X				X				X		X
AT-065	X	X	X							X	X			X					X		
AT-066		X	X	X				X		X	X								X		
AT-070	X	X	X	X				X		X	X								X		
AT-071	X		X					X		X	X				X				X		
AT-072			X					X	X	X	X										
AT-073	X	X	X					X		X	X			X				X	X		
AT-076a		X	X	X				X		X	X				X						
AT-079	X	X	X	X				X		X	X								X		
AT-081		X	X					X		X	X										
AT-082	X	X	X							X	X				X				X		
AT-083		X	X		X			X		X	X			X					X		
AT-085		X	X							X	X								X		X
AT-090a		X	X							X	X								X		
AT-090b		X	X							X	X			X					X		
AT-092		X	X						X	X	X		X					X	X		
AT-098		X	X							X	X		X						X		
AT-107		X	X						X	X	X			X	X	X			X		
AT-109a		X	X					X		X	X								X		
AT-109b		X	X					X		X	X								X		
AT109c	X	X	X							X	X										
AT109d		X	X					X		X	X				X	X			X		
AT-109e		X	X					X		X	X								X		

**Table 2 (cont.): Sample Mineralogy**

Sample	Chl	Bt	Grt	St	Ky	Sil	Crd	Ms	Kfs	Plg	Qtz	Cpx	Amp	Ttn	Tur	Ap	Mnz	Rt	Ilm	Scp	Cal
AT-111a	X	X	X	X				X		X	X								X		
AT-111b		X	X					X	X	X	X			X				X	X		
AT-111c		X	X						X	X	X			X	X				X		
AT-113	X	X	X	X				X		X	X								X		
AT-115	X	X	X					X		X	X								X		
AT-116a	X		X					X		X	X							X	X		
AT-116b	X	X	X					X		X	X								X		
AT-116c	X	X	X		X			X		X	X										
AT-117		X	X					X		X	X							X	X		
AT-119		X	X					X		X	X							X	X		
AT-121a		X	X					X		X	X							X	X		
AT-121b			X					X		X	X				X			X	X		
AT-134		X	X	X	X					X	X								X		
AT-136a			X		X					X	X					X		X	X		
AT-136b		X	X	X	X			X		X	X								X		
AT-138		X	X		X			X		X	X								X		
AT-140		X	X		X			X		X	X								X		
AT-176		X	X					X		X	X				X				X		
AT-178		X	X					X		X	X				X				X		
AT-180		X	X					X		X	X								X		
AT-182		X	X					X		X	X								X		

**Table 3: Pressure – Temperature summary**

Sample	Latitude	Longitude	Assemblage	°C	±2s	kbar	±2s	cor	σfit	fit
<i>El Zaino Complex</i>										
AT-021a	-29.19322	-68.51361	<i>Grt-Bt-Ms-Plg-Kfs</i>	564	91	8.2	1.4	0.861	0.82	1.61
AT-021b	-29.19322	-68.51361	<i>Grt-Bt-Amp-Plg</i>	489	71	7.3	1.6	0.559	1.45	1.54
AT-017	-29.19053	-68.49359	<i>Grt-Bt-Ms-Plg-Chl</i>	575	14	8.2	0.7	0.333	0.73	1.54
<i>Maz Shear Zone</i>										
AT-073	-29.19457	-68.48200	<i>Grt-Bt-Ms-Plg</i>	602	90	8.3	1.4	0.800	0.86	1.61
AT-015	-29.18856	-68.47404	<i>Grt-Bt-Plg-Amp</i>	673	142	8.6	1.7	0.795	1.24	1.73
AT-076	-29.20108	-68.47318	<i>Grt-Bt-Ms-Plg-St</i>	653	27	7.7	1.1	0.742	0.30	1.61
<i>Maz Complex</i>										
AT-066	-29.18795	-68.47084	<i>Grt-St-Bt-Ms</i>	706	34	9.0	1.5	0.928	0.53	1.61
AT-079	-29.18817	-68.46121	<i>Grt-St-Bt-Ms-Plg-Chl</i>	647	47	7.0	1.1	0.756	0.77	1.54
AT-068	-29.18615	-68.45705	<i>Grt-St-Bt-Ms-Plg</i>	694	27	7.2	1.1	0.840	0.72	1.54
AT-070	-29.18690	-68.45453	<i>Grt-St-Bt-Ms-Plg-Chl</i>	613	13	6.6	0.8	0.224	0.53	1.49
AT-083	-29.18825	-68.44635	<i>Grt-St-Bt-Ms</i>	616	51	6.7	1.0	-0.263	1.11	1.54
AT-098	-29.19443	-68.41742	<i>Grt-Bt-Plg-Amp</i>	676	112	7.8	1.4	0.486	0.57	1.96
AT-092	-29.18971	-68.41295	<i>Grt-Bt-Plg-Amp</i>	690	74	7.9	1.6	0.491	1.18	1.61
AT-056	-29.20469	-68.39146	<i>Grt-Bt-Plg-Kfs-Ms</i>	804	167	10.6	1.1	-0.016	0.97	1.96
AT-041	-29.21208	-68.38789	<i>Grt-Plg-Amp-Bt-Kfs</i>	749	120	9.0	1.6	0.581	0.25	1.73
AT-042	-29.21378	-68.37942	<i>Grt-Plg-Amp</i>	725	118	9.7	2.0	0.663	1.50	1.73
AT-044	-29.21397	-68.37800	<i>Grt-Bt-Ms-Kfs</i>	723	142	10.0	3.1	0.507		
AT-028	-29.22158	-68.36735	<i>Grt-Bt-Plg-Kfs</i>	878	56	10.3	1.6	0.272	0.61	1.73
<i>Ramaditas Complex</i>										
AT-047	-29.28666	-68.28350	<i>Grt-Bt-Plg-Cd</i>	856	160	5.5	1.5	0.261	1.10	1.96



**Table 4: Lu - Hf Summary**

Sample	Mineral	Lu (ppm)	Hf (ppm)	$^{176}\text{Lu}/^{177}\text{Hf}$	$^{176}\text{Hf}/^{177}\text{Hf}$	$\pm 2\sigma$
<i>AT-021</i>						
<i>423 ± 10 Ma MSWD = 4.6</i>						
	Garnet	7.63	2.66	0.4064	0.285705	15
	Garnet	5.73	1.93	0.4224	0.285818	15
	Garnet	4.71	2.63	0.2549	0.284460	15
	Garnet	3.19	1.91	0.2362	0.284326	15
	Whole Rock	0.495	8.82	0.0080	0.282538	15
	Whole Rock	0.432	4.96	0.0123	0.282634	15
<i>AT-076</i>						
<i>429 ± 5 Ma MSWD = 2.6</i>						
	Garnet	6.73	0.560	0.3350	0.285201	15
	Garnet	7.16	0.470	0.3566	0.285394	16
	Garnet	6.35	0.470	0.3663	0.285452	15
	Garnet	9.09	0.500	0.3806	0.285573	15
	Whole Rock	3.51	0.530	0.0117	0.282588	15
	Whole Rock	3.50	0.520	0.0162	0.282655	15
<i>AT-015</i>						
<i>428 ± 9 Ma MSWD = 5.6</i>						
	Garnet	7.15	0.144	7.105	0.339633	21
	Garnet	6.72	0.142	6.786	0.336439	20
	Garnet	7.20	0.085	12.27	0.380860	26
	Garnet	7.03	0.107	9.456	0.356599	23
	Whole Rock	1.88	8.74	0.0305	0.282703	15
	Whole Rock	1.97	1.75	0.1595	0.283902	15
<i>AT-068</i>						
<i>591 ± 59 Ma, 759 ± 54 Ma</i>						
	Garnet	6.51	2.09	0.4421	0.287954	15
	Garnet	6.84	2.24	0.4336	0.287606	15
	Garnet	6.69	2.16	0.4396	0.286907	15
	Garnet	6.58	1.80	0.5203	0.287443	15
	Whole Rock	0.423	10.6	0.0057	0.281706	15
	Whole Rock	0.236	3.47	0.0097	0.281799	15
<i>AT-028</i>						
<i>1031 ± 6.2 Ma, 1074 ± 6.4 Ma</i>						
	Garnet	1.46	0.730	0.2828	0.28732	15
	Garnet	1.81	0.850	0.3016	0.28753	15
	Garnet	2.00	0.950	0.2998	0.28748	15
	Garnet	1.81	0.880	0.2937	0.28761	15
	Whole Rock	0.050	6.710	0.0012	0.28168	15
	Whole Rock	0.040	1.120	0.0049	0.28173	15
<i>AT-047</i>						
<i>449 ± 1.6 Ma, 461 ± 1.7 Ma</i>						
	Garnet	10.55	0.810	1.854	0.298613	16
	Garnet	11.49	0.840	1.957	0.299025	16
	Garnet	11.84	0.710	2.378	0.302597	16
	Garnet	11.36	0.710	2.291	0.302277	16
	Whole Rock	0.270	6.21	0.0063	0.282584	15
	Whole Rock	0.180	1.04	0.0238	0.282992	15

Table 5: U - Pb Summary

Spot	Concentrations			Atomic Ratios						Age (Ma)			Weighted Average age	
	U	Th	Th/U	206Pb	f206	238U/206Pb	% err	207Pb/206Pb	% err	206Pb/238U	% err	206Pb/238U		err abs
AT-068	(ppm)	(ppm)		(ppm)	(%)		(1s)		(1s)		(1s)		(1s)	(2s)
68-36A.1	8568	9961	1.2	114	4.8	61.49	19.8	0.0577	114.3	0.0157	19.9	99	21	<b>410.6 ± 1.4 Ma</b> <b>MSWD: 1.57</b>
68-05A.1	574.3	2607	4.5	15	-0.3	32.23	45.4	0.0207	77.6	0.0300	45.4	198	89	
68-14B.1	3815	59150	15.5	200	8.9	14.96	2.4	0.0709	47.0	0.0622	3.8	381	14	
68-12C.1	2973	31970	10.8	158	1.1	16.01	1.1	0.0526	4.1	0.0617	1.2	387	4	
68-08A.1	4435	36268	8.2	237	2.9	15.63	0.8	0.0650	5.2	0.0630	0.8	389	3	
68-03C.1	2061	12485	6.1	110	-0.2	16.11	3.3	0.0505	6.5	0.0619	3.3	389	13	
68-04A.1	2411	24146	10.0	133	1.8	15.33	0.8	0.0540	5.0	0.0640	0.8	400	3	
68-03B.1	3676	33052	9.0	203	0.6	15.49	1.0	0.0556	1.5	0.0643	1.0	401	4	
68-14A.1	2728	33828	12.4	151	7.0	14.43	1.4	0.0529	15.5	0.0643	1.5	403	6	
68-11A.1	2780	23694	8.5	154	0.4	15.44	1.2	0.0553	1.3	0.0646	1.2	403	5	
68-12B.1	4261	58848	13.8	238	7.9	14.15	3.2	0.0702	46.9	0.0663	3.8	406	19	
68-03A.1	3534	25614	7.2	198	0.9	15.22	1.2	0.0567	1.8	0.0653	1.2	407	5	
68-17A.1	2553	24877	9.7	143	0.5	15.25	1.0	0.0555	1.5	0.0653	1.1	407	4	
68-13A.1	3181	30587	9.6	179	0.4	15.17	0.9	0.0566	1.2	0.0658	0.9	410	4	
68-16B.1	2904	25593	8.8	164	1.2	15.03	1.3	0.0556	2.3	0.0658	1.3	410	5	
68-17B.1	2635	23572	8.9	149	3.0	14.76	0.8	0.0519	4.8	0.0655	0.8	410	3	
68-12D.1	3271	27533	8.4	185	2.7	14.79	0.8	0.0571	3.3	0.0659	0.8	411	3	
68-12A.1	2622	26280	10.0	149	0.3	15.07	1.3	0.0551	1.3	0.0661	1.3	413	5	
68-13B.1	2955	25129	8.5	170	1.0	14.80	1.0	0.0569	1.9	0.0671	1.0	417	4	
68-08C.1	2767	20944	7.6	159	0.3	14.86	1.3	0.0545	1.4	0.0670	1.3	418	5	
68-08D.1	4073	26565	6.5	235	2.2	14.56	1.3	0.0539	3.5	0.0670	1.3	419	5	
68-16A.1	2059	21512	10.4	119	0.4	14.81	1.0	0.0540	1.7	0.0672	1.0	420	4	
68-18B.1	2752	28763	10.5	162	0.5	14.50	2.1	0.0561	1.9	0.0687	2.1	428	9	
68-08B.1	2847	21496	7.6	170	0.3	14.38	0.9	0.0546	1.3	0.0693	0.9	432	4	
68-18A.1	2405	15758	6.6	346	0.9	5.92	0.8	0.0781	0.6	0.1686	0.8	998	7	

Table 5 (cont.): U – Pb Summary

Concentrations			Atomic Ratios							Age (Ma)			Weighted Average age	
Spot	U	Th	Th/U	206Pb	f206	238U/206Pb	% err	207Pb/206Pb	% err	206Pb/238U	% err	206Pb/238U		err abs
AT-047	(ppm)	(ppm)		(ppm)	(%)		(1s)		(1s)		(1s)		(1s)	(2s)
47-14A.1	2328	18112	7.8	113	1.3	17.50	13.4	0.0552	11.0	0.0565	13.4	354	47	<b>461.6 ± 3.1 Ma</b>
47-08A.1	2595	23252	9.0	149	1.7	14.66	1.6	0.0547	7.6	0.0670	1.7	418	7	<b>MSWD: 1.6</b>
47-15A.1	1747	14417	8.3	102	0.2	14.64	1.5	0.0570	1.2	0.0683	1.5	425	6	<b>426.1 ± 7 Ma</b>
47-03A.1	2173	17327	8.0	128	0.5	14.50	1.7	0.0553	1.8	0.0686	1.7	428	7	<b>MSWD: 1.02</b>
47-09A.1	1497	11905	8.0	90	0.5	14.18	2.0	0.0577	1.6	0.0704	2.0	437	8	
47-07A.1	2645	18281	6.9	161	0.1	14.14	1.5	0.0563	1.5	0.0707	1.5	440	6	
47-17A.1	2193	15701	7.2	133	0.1	14.12	1.0	0.0559	1.1	0.0707	1.0	441	4	
47-17C.1	2003	18108	9.0	122	0.2	14.03	1.9	0.0572	1.1	0.0713	1.9	443	8	
47-02A.1	2048	16763	8.2	126	0.1	13.97	1.0	0.0556	1.2	0.0715	1.0	445	4	
47-17B.1	1867	9908	5.3	118	0.2	13.63	1.1	0.0569	1.1	0.0733	1.1	456	5	
47-04A.1	2230	15869	7.1	141	0.0	13.62	1.4	0.0557	0.9	0.0734	1.4	457	6	
47-15B.1	1663	9658	5.8	105	0.1	13.56	0.8	0.0563	1.2	0.0737	0.8	458	4	
47-05B.1	1890	15570	8.2	120	0.1	13.51	0.8	0.0563	1.1	0.0739	0.8	460	4	
47-05C.1	1862	12864	6.9	119	0.0	13.48	0.9	0.0555	1.1	0.0741	0.9	461	4	
47-08B.1	1904	16310	8.6	123	-0.1	13.35	0.8	0.0550	1.2	0.0748	0.8	466	4	
47-05A.1	1601	15183	9.5	105	0.0	13.15	1.0	0.0552	1.3	0.0759	1.0	472	5	

Concentrations			Atomic Ratios							Age (Ma)			Weighted Average age	
Spot	U	Th	Th/U	206Pb	f206	238U/206Pb	% err	207Pb/206Pb	% err	206Pb/238U	% err	206Pb/238U		err abs
AT-076	(ppm)	(ppm)		(ppm)	(%)		(1s)		(1s)		(1s)		(1s)	(2s)
76-05A.1	3147	13798	4.4	174	0.3	15.49	2.0	0.0563	1.9	0.0645	2.0	402	8	<b>418.7 ± 3.3 Ma</b>
76-06A.1	4317	30842	7.1	240	0.1	15.43	3.9	0.0547	1.0	0.0648	3.9	405	16	<b>MSWD: 1.14</b>
76-06B.1	3415	12904	3.8	190	0.2	15.37	4.4	0.0560	1.0	0.0650	4.4	405	17	
76-07A.1	3583	11590	3.2	202	0.1	15.24	1.9	0.0558	0.8	0.0656	1.9	409	8	
76-03B.1	3212	12469	3.9	182	0.2	15.12	1.4	0.0558	0.9	0.0660	1.4	412	6	
76-03C.1	2838	14470	5.1	162	0.2	15.04	1.2	0.0564	0.9	0.0664	1.2	414	5	
76-04a.1	3055	11045	3.6	177	0.2	14.82	1.1	0.0563	0.9	0.0674	1.1	420	5	
76-02A.1	2491	9839	4.0	144	0.1	14.83	1.6	0.0564	0.9	0.0675	1.6	420	7	
76-03A.1	2736	13320	4.9	159	0.1	14.78	1.3	0.0549	1.1	0.0676	1.3	422	5	
76-08A.1	4143	22688	5.5	241	0.2	14.73	1.4	0.0556	0.9	0.0678	1.4	423	6	
76-10A.1	3672	17579	4.8	214	0.2	14.72	2.2	0.0563	1.5	0.0679	2.2	423	9	
76-09A.1	2462	10491	4.3	144	0.3	14.67	1.3	0.0568	0.9	0.0681	1.3	424	5	
76-08B.1	4359	24239	5.6	255	0.0	14.67	1.2	0.0551	0.8	0.0681	1.2	425	5	
76-08A.2	2872	24016	8.4	169	0.1	14.56	1.8	0.0543	1.2	0.0685	1.8	428	8	

**Table 5 (cont.): U – Pb Summary**

Concentrations			Atomic Ratios						Age (Ma)					
Spot	U	Th	Th/U	206Pb	f206	238U/206Pb	% err	207Pb/206Pb	% err	206Pb/238U	% err	206Pb/238U	err abs	Weighted Average age
	(ppm)	(ppm)		(ppm)	(%)		(1s)		(1s)		(1s)		(1s)	(2s)
<b>AT-028</b>														
<b>28-09A.2</b>	2121	23486	11.1	116	0.3	15.70	1.3	0.0549	1.5	0.0635	1.3	397	5	<b>413.2 ± 5 Ma</b>
<b>28-14A.2</b>	2063	32781	15.9	113	0.3	15.59	1.8	0.0559	1.4	0.0640	1.8	399	7	<b>MSWD: 7.68</b>
<b>28-04A.1</b>	1316	17576	13.4	75	2.0	14.87	1.6	0.0565	4.4	0.0660	1.6	411	7	<b>1227.1 ± 6.8 Ma</b>
<b>28-03B.2</b>	940	32067	34.1	55	0.5	14.67	1.5	0.0573	4.3	0.0680	1.5	423	6	<b>MSWD: 1.55</b>
<b>28-09a.1</b>	1815	20322	11.2	108	0.5	14.39	1.2	0.0554	2.3	0.0692	1.2	431	5	
<b>28-06A.2</b>	905	17435	19.3	84	1.3	9.18	6.3	0.0711	2.9	0.1088	6.3	658	41	
<b>28-14A.1</b>	560	39184	69.9	57	1.0	8.39	1.8	0.0699	2.6	0.1190	1.8	719	13	
<b>28-07A.2</b>	1225	23442	19.1	130	1.1	8.02	4.2	0.0720	1.7	0.1245	4.2	750	31	
<b>28-11A.1</b>	538	31063	57.7	64	1.2	7.18	4.0	0.0734	3.2	0.1388	4.0	832	32	
<b>28-03A.1</b>	606	22367	36.9	76	1.0	6.74	1.4	0.0753	1.4	0.1481	1.4	884	12	
<b>28-07B.1</b>	604	16719	27.7	81	1.0	6.35	3.0	0.0759	1.1	0.1569	3.0	934	27	
<b>28-08B.1</b>	2465	22110	9.0	377	1.0	5.56	1.3	0.0808	0.8	0.1795	1.3	1057	13	
<b>28-03B.1</b>	696	31217	44.8	110	0.7	5.37	1.7	0.0811	1.1	0.1859	1.7	1093	18	
<b>28-06B.1</b>	983	18955	19.3	157	0.6	5.35	2.4	0.0802	0.7	0.1867	2.4	1098	26	
<b>28-05A.1</b>	743	24220	32.6	120	0.9	5.27	1.2	0.0806	1.2	0.1887	1.2	1110	12	
<b>28-10A.1</b>	560	21900	39.1	94	0.1	5.10	1.2	0.0786	0.9	0.1958	1.2	1153	13	
<b>28-02a.1</b>	888	23287	26.2	150	0.4	5.07	1.3	0.0818	0.7	0.1974	1.3	1157	15	
<b>28-03C.1</b>	692	20810	30.1	119	0.4	4.99	1.4	0.0813	0.9	0.2002	1.4	1174	15	
<b>28-06A.1</b>	750	19435	25.9	130	0.3	4.93	1.2	0.0814	0.8	0.2027	1.2	1187	14	
<b>28-13A.1</b>	1488	20797	14.0	259	0.3	4.92	1.4	0.0815	0.6	0.2029	1.4	1188	16	
<b>28-08A.1</b>	1097	16868	15.4	191	0.4	4.90	1.6	0.0827	1.3	0.2039	1.6	1193	19	
<b>28-10A.2</b>	1072	33311	31.1	187	0.0	4.91	1.2	0.0798	0.7	0.2035	1.2	1194	13	
<b>28-10B.1</b>	2855	24926	8.7	506	0.2	4.84	1.2	0.0818	0.4	0.2066	1.2	1209	14	
<b>28-10C.1</b>	1584	24342	15.4	285	0.1	4.77	1.1	0.0812	0.5	0.2093	1.1	1225	13	
<b>28-07a-1.1</b>	3436	19535	5.7	629	0.0	4.69	1.5	0.0811	0.4	0.2129	1.5	1246	18	

**Table 6: Lu – Hf and U – Pb Summary**

Sample	Unit	Latitude	Longitude	Age (Ma)	±	MSWD	Comments
<i>Lu-Hf Grt</i>							
AT-021	Zaino Complex	-29.19322	-68.51361	423	10	4.6	
AT-015	Maz Shear Zone	-29.18856	-68.47404	428	9	5.6	
AT-076	Maz Shear Zone	-29.20108	-68.47318	429	5	2.6	
AT-068	Maz Complex	-29.18615	-68.45705	759	54	-	oldest grt
AT-068				591	59	-	youngest grt
AT-028	Maz Complex	-29.22158	-68.36735	1074	6.4	-	oldest grt
AT-028				1031	6.2	-	youngest grt
AT-047	Ramaditas	-29.28666	-68.28350	461	1.7	-	oldest grt
AT-047				449	1.6	-	youngest grt
<i>U-Pb Mnz</i>							
AT-076	Maz Shear Zone	-29.20108	-68.47318	418.7	3.3	1.14	weighted mean
AT-068	Maz Complex	-29.18615	-68.45705	410.6	2.4	1.57	weighted mean
AT-068				1214	12	2.2	upper intercept
AT-028	Maz Complex	-29.22158	-68.36735	413.2	5	7.68	weighted mean
AT-028				1227.1	6.8	1.55	weighted mean
AT-047	Ramaditas	-29.28666	-68.28350	426.1	7	1.02	weighted mean
AT-047				461.6	3.1	1.6	weighted mean

