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1 **Slowing Rates of Regional Exhumation in the Western Himalaya: Fission**
2 **Track Evidence from the Indus Fan**

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12

13 **Abstract**

14 We use apatite fission track (AFT) ages from sediments recovered by International Ocean
15 Discovery Program in the Laxmi Basin, Arabian Sea, to constrain exhumation rates in the
16 western Himalaya and Karakoram since 15.5 Ma. With the exception of a Triassic population in
17 the youngest 0.93 Ma samples AFT ages are overwhelmingly Cenozoic, largely <25 Ma,
18 consistent with both a Himalaya-Karakoram source and rapid erosion. Comparison of the
19 minimum cooling age of each sample with depositional age (lag time) indicates an acceleration
20 in exhumation between 7.8 and 7.0 Ma, with lag times shortening from ~6.0 m.y. between 8.5
21 and 7.8 Ma to being within error of zero between 7.0 and 5.7 Ma. Sediment supply at that time

22 was largely from the Karakoram and to a lesser extent the Himalaya based on U-Pb zircon ages
23 from the same samples. This time coincides with a period of drying in the Himalayan foreland
24 caused by weaker summer monsoons and Westerly winds. It also correlates with a shift of
25 erosion away from the Karakoram, Kohistan and the Tethyan Himalaya towards more erosion of
26 the Lesser, Greater Himalaya and Nanga Parbat, as shown by zircon U-Pb provenance data and
27 especially after 5.7 Ma based on Nd isotope data. Samples younger than 7.0 Ma have lag times
28 ~4.5 m.y., similar to Holocene Indus delta sediments.

29

30 Keywords: International Ocean Discovery Program, Fission track, erosion, Himalaya, Indus Fan,
31 monsoon

32

33 **Introduction**

34 If we are to understand how the evolving climate of Asia has impacted the tectonic
35 development of the Himalaya and Tibetan Plateau, or vice versa, we must use the sedimentary
36 records in basins adjacent to these mountain ranges in order to reconstruct the long-term history
37 of exhumation caused by erosion. Thermochronology measurements on bedrock currently
38 exposed at the surface only provide constraints on the most recent stages of the cooling history of
39 those particular units. By definition older bedrock has been removed so that the older erosional
40 history can only be reconstructed through study of the sedimentary record. However, interpreting
41 the sedimentary record can be complicated if burial of sediment resets sensitive low temperature
42 thermochronometers, eliminating the cooling history of the source bedrocks (Carter 1999).
43 Although higher temperature methods (e.g., muscovite Ar/Ar dating) (Szulc et al. 2006; White et

44 al. 2002) can be useful in examining past erosion and are resistant to resetting these have the
45 disadvantage of being less sensitive to changes in the rates of exhumation by erosion because
46 they require a greater amount of exhumation between isotopic closure and exposure at the
47 surface. Nonetheless, detrital apatite fission track (AFT) can also have resolution problems,
48 because single grain ages are often imprecise, especially for young grains with very low track
49 counts.

50 A number of studies have examined the history of erosion in the Himalaya using the
51 foreland basin sediment record, in particular sedimentary rocks belonging to the Miocene-
52 Pliocene Siwalik Group (Baral et al. 2015; Bernet et al. 2006; Cerveny et al. 1989; Chirouze et
53 al. 2015; Chirouze et al. 2013; Ghosh & Kumar 2000; Najman 2006; van der Beek et al. 2006).
54 Although this stratigraphic unit has provided useful information about past patterns and rates of
55 erosion the quality of information from AFT thermochronology has been limited due to resetting
56 caused by post-deposition burial, especially in the lower parts of the section (van der Beek et al.
57 2006). In addition, the foreland basin sequence at any one particular location will typically
58 reflect the rivers that are flowing from the Himalaya at that point, providing a localized record.
59 Although this may be very useful for examining single rivers, it is often hard to judge how
60 effective each sequence might be in reconstructing erosion at the regional scale. For example,
61 because the trunk Indus River lies on the western edge of the drainage, Siwalik Group rocks in
62 the eastern parts of the catchment provide no information about how its sediment load may have
63 evolved.

64 In this study we present AFT data from new scientific boreholes in the western Indian
65 Ocean in order to derive a regional image of changing erosion rates within the Western
66 Himalayas since ca. 15.5 Ma, and in particular after 9 Ma. Use of the International Ocean

67 Discovery Program (IODP) boreholes in the Laxmi Basin (Fig. 1)(Pandey et al. 2016b) has the
68 advantage that the sediment thickness is low (<1.1 km) and the geothermal gradient is 53°C/km
69 and 57°C/km at Sites U1456 and U1457 respectively (Pandey et al. 2016b). Although these are
70 high values this means that even the base of the section will fall below temperatures required to
71 cause significant annealing or resetting of fission tracks in apatite, i.e. ~60–110°C (Green 1989)
72 and therefore the original cooling history of the bedrock sources will be preserved. All but one of
73 the samples were recovered from depths shallower than 722 mbsf, implying no more than 38°C
74 burial temperature at the present maximum burial depth. The deepest sample (U1456E-19R-3,
75 10-20 cm) was recovered from a depth of 1103 mbsf but the fission track ages are older than the
76 depositional age, indicating that this too is not reset.,

77 Constraining rates of bedrock source cooling caused by erosion driven by rock uplift can
78 help identify locations of active tectonics and the rates and patterns of mountain growth.
79 However, climate change may also play a role in relation to variations in precipitation rate that
80 are linked to the intensity of the South Asian monsoon. This is known to have varied
81 significantly throughout the Cenozoic (Betzler et al. 2016; Gupta et al. 2015; Kroon et al. 1991;
82 Prell et al. 1992; Quade et al. 1989). Debate continues concerning the history of strengthening of
83 the South Asian monsoon, but increasingly there is a consensus that the climate began to dry
84 after 8 Ma (Behrensmeyer et al. 2007; Clift 2017; Singh et al. 2011), following a period of
85 maximum intensity in the middle Miocene (Clift et al. 2008). It has been suggested that it is the
86 strength of the summer monsoon rains during the middle Miocene that resulted in rapid
87 exhumation of the Greater Himalaya at that time driven by strong erosion (Clift et al. 2008). If
88 that is true one might predict that the rate of erosion since that time was also coupled with
89 monsoon intensity. However, work within the foreland sedimentary rocks of the Siwalik Group

90 in Nepal shows that the rate of exhumation in the central Nepalese Himalaya remained
91 essentially constant after 8 Ma (van der Beek et al. 2006). In contrast, the same study argued that
92 rates of erosion had increased between 8 and 3 Ma in Western Nepal, despite the fact that both
93 sections lie within the Ganges drainage system, which is wetter than the Indus basin considered
94 here (Bookhagen & Burbank 2006). In contrast, AFT data from Ocean Drilling Program (ODP)
95 Sites 717 and 718 on the Bengal fan showed that rapid rates of exhumation of the bedrock
96 sediment sources to the Ganges-Brahmaputra basin has been ongoing since the middle Miocene
97 (Corrigan & Crowley 1990). Reappraisal of this data by van der Beek et al. (2006) indicated
98 relatively constant lag times (i.e., the difference between the depositional age and the AFT
99 cooling) since 9 Ma, suggestive of uniform erosion rates.

100 There are few constraints over how erosion rates might have changed during the
101 Pleistocene. While some have argued that the onset of northern hemisphere glaciation (NHG) has
102 intensified rates of erosion during the last couple of million years (Clift 2006; Métiévier et al.
103 1999; Zhang et al. 2001), other workers, drawing on cosmogenic isotope data (Willenbring &
104 von Blanckenburg 2010), suggest that continental weathering rates have remained essentially
105 steady-state during the Neogene and especially the Plio-Pleistocene. Such an observation does
106 not require faster sediment delivery to the ocean, although this was proposed from a global data
107 compilation implying a steady state supply of sediment spanning tens of millions of years (Sadler
108 & Jerolmack 2014). Here we provide the first detailed AFT constraints on erosion rates in the
109 Western Himalaya, within the Indus basin, in order to see whether the temporal evolution in that
110 region mirrors that found in Nepal and in the Ganges-Brahmaputra drainage basin.

111 Over the period since 15.5 Ma considered by this study the Western Himalaya have
112 experienced significant tectonic changes. The Lesser Himalayas were brought to the surface

113 because of duplexing above the Main Boundary Thrust (MBT) (Huyghe et al. 2001; Mugnier et
114 al. 1994), coupled with focused erosion since the Late Miocene. There is continued debate about
115 when exposure of the Lesser Himalaya might have occurred. Early studies suggested that the
116 MBT initiated around 10–11 Ma (Meigs et al. 1995) allowing the Lesser Himalayan Duplex to
117 form and be uplifted and then eroded. Work from the Siwalik Group in Northwest India points to
118 an initial exposure of the Lesser Himalaya at ca. 9 Ma followed by more widespread exposure
119 after 6 Ma (Najman et al. 2009), although this may be only applicable to the Beas River area
120 (Fig. 1). Nd and zircon U-Pb data from IODP Sites U1456 and U1457 now suggest initial
121 exposure after 8.3 Ma and widespread unroofing after 1.9 Ma (Clift et al. 2019b). Other
122 potentially important sources of sediment to the submarine fan include the Nanga Parbat massif
123 that is located next to the Indus River in the Western Syntaxis (Fig. 1). Provenance studies from
124 the Indus River downstream of Nanga Parbat indicate that this massif has only limited sediment
125 generating potential at the present time (Clift et al. 2002b; Garzanti et al. 2005; Lee et al. 2003),
126 despite the start of uplift ca. 6 Ma (Chirouze et al. 2015). In contrast, its eastern equivalent
127 (Namche Barwe) is believed to be a major source of sediment to the Brahmaputra (Garzanti et al.
128 2004; Stewart et al. 2008). Bedrock thermochronology measurements testify to Nanga Parbat
129 being very rapidly exhumed in the recent geologic past (Zeitler et al. 1993), but this does not
130 seem to generate much of the sediment in the river downstream of that point (Alizai et al. 2011).
131 Zircon fission track (ZFT) and Nd isotope data in the western part of the Siwalik ranges in
132 Pakistan indicate that this massif and other Himalayan units in the western syntaxis may have
133 become more important as a sediment source after around 6 Ma (Chirouze et al. 2015). The
134 sedimentary record in the Indus Fan may also been affected by large-scale drainage capture.
135 Neodymium isotope measurements on samples from an industrial drill site on the Indus shelf, as

136 well as limited ODP samples from the upper fan, were used to argue that the eastern tributaries
137 of the Indus River were only been captured into the modern system after 5 Ma (Clift & Blusztajn
138 2005). However, this is contradicted by combined ZFT and Nd isotope data that supports relative
139 stability in drainage patterns but changing rates of erosion in the Himalaya and Karakoram since
140 the Miocene (Chirouze et al. 2015).

141

142 **Regional Setting**

143 IODP Expedition 355 sampled sediments from the Indus Fan deposited within the Laxmi
144 Basin offshore western India (Fig. 1). Although the Laxmi Basin is separated from the main
145 Arabian Basin by the Laxmi Ridge, the bathymetry of the basin and the orientation of active
146 channels (Mishra et al. 2016) indicates that the primary source of sediment to the coring
147 locations would be the Indus River, with lesser input from peninsular rivers such as the Tapti and
148 Narmada. Initial petrographic-based interpretations of the sediments made shipboard during the
149 expedition suggested that there were limited amounts of sediment delivery from Western India,
150 and tend to be found only in the youngest parts of the section (Pandey et al. 2016a).

151 The Laxmi Basin itself dates from the latest Cretaceous when India began to separate from
152 the Seychelles (Bhattacharya et al. 1994; Pandey et al. 1995). Following the onset of India-Asia
153 collision, ca. 50–60 Ma (DeCelles et al. 2014; Najman et al. 2010), the uplift and erosion of the
154 Himalaya has resulted in a huge flux of sediment into the Arabian Sea. Although the Indus Fan is
155 much smaller than the Bengal Fan it is nonetheless the second largest sediment body on Earth
156 and is believed to have accumulated sediment eroded from the mountains at least since 45 Ma
157 (Clift et al. 2001).

158 Drilling during Expedition 355 recovered a section that penetrated to basement at Site
159 U1457 (Fig. 2), but because of large-scale mass wasting (Dailey et al. 2019) the most complete
160 erosional record only spans the last 10.8 m.y., with much of the older sediment either missing,
161 due to erosion or non-deposition, or not sampled. Coring was undertaken at two sites, Site U1456
162 in the central part of the Laxmi Basin, as well as at Site U1457 located on the flanks of the
163 Laxmi Ridge (Fig. 1). In general, the sediment at Site U1456 tended to be coarser grained (Fig.
164 2). The entire sedimentary cover is also more complete at Site U1456 than at Site U1457. The
165 coarse-grained, sandy sediment that forms the focus of this study was taken from both sites and
166 is the product of turbidity current flows. Nonetheless, significant parts of the section are fine-
167 grained muddy facies together with carbonate-rich intervals and these are interbedded with sandy
168 turbidite material caused by sedimentation on depositional lobes within the middle fan (Fig. 2).
169 There are also interbeds of calcareous-rich pelagic material that reflect times when the main
170 Indus-sourced depocentre was located to the west of the Laxmi Ridge, so that the primary clastic
171 flux from the Indus River was not reaching the drilling area. Because the drilling sites are located
172 above the carbonate compensation depth (CCD) it was possible to date the age of sedimentation
173 using a combination of nannofossil and foraminifera biostratigraphy coupled with
174 magnetostratigraphy that provides a relatively robust age model (Pandey et al. 2016b). Drilling
175 was able to penetrate a thick mass transport deposit (MTD) deposited just before 10.8 Ma
176 (Calvès et al. 2015), but at Site U1456 coring was able to recover a short interval below the
177 MTD, providing a single sample that is substantially older than any of the other sediments
178 recovered and which has been approximately dated at 15.5 Ma (Pandey et al. 2016a). At Site
179 U1457 all fan sediment predating the mass wasting event had been removed so that our studies
180 are restricted to the section younger than 10.8 Ma at that location.

181 We apply the AFT thermochronology dating method to this sediment in order to understand
182 how the source rocks that provided material to the Arabian Sea evolved in their cooling and
183 exhumation history since the middle Miocene. Fission track studies are a well-established
184 method for looking at bedrock unroofing and potentially also sediment provenance if the source
185 regions themselves are sufficiently well defined and if cooling ages are relatively constant in a
186 source area (Carter 1999; Green et al. 1989; Laslett et al. 1987). In a complex area like the
187 western Himalaya cooling ages vary across tectonic blocks and through time so that the
188 interpretation of the AFT ages is contingent on supporting provenance data and cannot be used to
189 constrain provenance by themselves. In this study we draw on zircon U-Pb age data from these
190 same boreholes (Clift et al. 2019b). Simple comparison of modern bedrock AFT ages and detrital
191 AFT ages in sediments more than around a million years old is not justifiable because the
192 cooling rates of the bedrock will change on such timescales.

193

194 **Methodology**

195 Low-temperature AFT central ages reflect cooling through 60–125°C over time scales of 1–
196 10 m.y. (Green et al. 1989). Fission tracks form continuously through time at an abundance
197 determined by the concentration of ^{238}U in the host apatite grain (Haack 1977). The method has
198 been a widely used and is effective for studying exhumation history and provenance of shallow-
199 buried sediment (Carter 2007; Gallagher et al. 1995). Samples were taken where suitable sandy
200 material was available at both IODP sites, as shown in Figure 2 and Table 1. Some of the apatites
201 were extracted from the same samples analysed for detrital zircon U-Pb dating by Clift et al.
202 (2019b).

203 Following mineral separation AFT analysis was performed at the London Geochronology
204 Centre based at University College, London, UK. Polished grain mounts of apatite were etched
205 with 5N HNO₃ at 20°C for 20 seconds to reveal the spontaneous fission-tracks. Subsequently the
206 uranium content of each crystal was determined by irradiation, which induced fission in a
207 proportion of the ²³⁵U. The induced tracks were registered in mica external detectors. The
208 samples for this study were irradiated in the irradiated in the FRM 11 thermal neutron facility at
209 the University of Munich, Germany. The neutron flux was monitored by including Corning glass
210 dosimeter CN-5, with a known uranium content of 11 ppm, at either end of the sample stack.
211 After irradiation, sample and dosimeter mica detectors were etched in 40% HF at 20°C for 25
212 minutes. Only crystals with sections parallel to the c-crystallographic axis were counted, as these
213 crystals have the lowest bulk etch rate. To avoid biased results through preferred selection of
214 apatite crystals the samples were systematically scanned, and each crystal encountered with the
215 correct orientation was analysed, irrespective of track density. The results of the fission track
216 analysis are presented in Table 2 and online Supplementary Table 1. Because the chi test, used to
217 detect extra Poisson variation, does not show how much over dispersion to be present in the
218 dataset we include the central age and its percentage relative error because this provides a
219 measure of the extent of age dispersion. It is also useful when there are low track counts (young
220 ages) as the chi test is unreliable under these conditions.

221

222 **Results**

223 Because all samples showed evidence of over-dispersion we examined the range of single
224 grain AFT ages in each sample using a combination of kernel density estimates (KDE) plots
225 following the method of Vermeesch (2012) and the radial diagrams of Galbraith (1990)(Fig. 3).

226 Plots that combine both types of data presentation are known as abanico plots (Dietze et al.
227 2016). In the radial plots the single grain ages are plotted away from a central point on the left
228 side of each diagram, with higher accuracy measurements plotted closer to the right-hand curved
229 y-axis against which the ages are measured. This approach allows populations of grains with
230 similar ages but varying degrees of uncertainty to be identified as arrays. In this particular study
231 we focus on the identification of a minimum age population extracted using the algorithm of
232 Galbraith (2005) that clusters in an array and trends towards the y-axis on the right-hand side of
233 each diagram. This avoids problems associated with a general purpose, multi-component mixture
234 model that can give a biased estimate of the minimum age towards younger values with
235 increasing sample size. The radial plots show if there is a single source (single array) or multiple
236 sources, if there are more than one array. Figure 3 and Table 2 show samples that have a second
237 age component (P2) as defined by ten or more grains. In all cases the majority of analysed grains
238 defines the minimum age and represents the time at which the dominant bedrock sources cooled
239 through the AFT partial annealing zone (PAZ).

240 In each case we also show the calculated depositional age derived from the shipboard
241 biostratigraphy and magnetic stratigraphy (Fig. 3). The minimum ages are older than or
242 concordant with the depositional age, as might be expected in a relatively shallow borehole in
243 which the temperatures are not elevated above those known to reset fission tracks in apatite
244 crystals. All samples have minimum ages less than 20 Ma, and P2 AFT ages are all less than 40
245 Ma (apart from the youngest sample) post-dating the initial collision of India and Asia. There are
246 particularly noteworthy concentrations of grain ages between 3 and 20 Ma. 50% of samples have
247 a minimum age younger than 10 Ma. The minimum age gets younger with decreasing
248 depositional age but not in a systematic way. The age difference between the minimum age and

249 deposition age is <5 m.y. for most samples, i.e., short lag times, but increases for samples
250 deposited between 7.84 and 8.2 Ma, as well as 7.07–7.28 Ma. The youngest sample (U1456A-
251 11H-6, 60-69 cm) is unlike many of the others in showing significantly older AFT ages (Fig. 3).

252 The youngest deposited sample is anomalous in having a minimum age population of
253 20.7 Ma, despite only having been deposited around 930 ka (Fig. 3A). This may be due to the
254 sample containing fewer apatites, with only 24 grains being countable, which is the smallest
255 number out of all samples analysed. This is in strong contrast with the much younger minimum
256 ages of the directly underlying samples. It is only the very oldest sample (~15 Ma, U1456E-19R-
257 3, 10–20 cm) which also has a minimum age of that value, but that sample has a short lag time
258 (Fig. 3W). We can assess the possible impact of low grain numbers on the critical minimum age
259 result in Figure 4. This plot shows that there is no correlation between the number of grains and
260 the minimum age, only reinforcing the fact that samples with low numbers of grains have more
261 uncertainty in the result, but not causing short lag times.

262 The core is not altered or veined and the modern maximum burial temperature of the
263 samples with lag times close to zero is far too cool to have affected the AFT ages. The ages are
264 within error of the depositional age, not resolvably younger. Although sample U1456D-12R-1
265 30-36 cm has a minimum age population slightly younger (6.6 ± 1.5 Ma) than the calculated
266 depositional age (7.0 Ma) but within error of that value and need not be reset. Moreover, the
267 young ages are also accompanied by older age populations that are also consistent with the
268 sediment not being thermally reset, as well as with the modern borehole temperatures being well
269 below the apatite partial annealing zone (556 mbsf (29.4°C) at Site U1456, 572–590 mbsf (32.6–
270 33.6°C) at Site U1457).

271

272 **Discussion**

273 The fact that all of our AFT ages are relatively young and mostly postdate widely
274 accepted times of India-Asia collision is a clear indication that they are derived from
275 Himalayan/Karakoram sources supplied by the Indus River and, with the exception of the
276 youngest sample, not from peninsular India. Ancient rocks of the Indian peninsula have not been
277 substantially deformed and uplifted during the Cenozoic and basement apatite fission track ages
278 are mostly Jurassic-Cretaceous. Although they range as young as 54 Ma (Gunnell et al. 2003;
279 Kalaswad et al. 1993), 95% of the ages measured are older than 100 Ma, averaging 228 Ma (Fig.
280 5H). This is somewhat older than most of the grain ages in Sample U1456A-11H-6, 60-69 cm.,
281 but does match the P2 older population in that sample (Table 2). Nonetheless, the minimum age
282 population of 20.7 ± 3.8 Ma requires a Himalaya-Karakoram provenance for 14 of the 24 grains
283 measured. U-Pb zircon ages from this same sample (Clift et al. 2019b) show that 8.25% of the
284 grains date to <200 Ma that require derivation from the Indus River because such zircon ages can
285 only be generated by erosion from Kohistan or Karakoram sources. Zircon grains older than 300
286 Ma could be from the peninsula or the Himalaya. This youngest sample seems likely to be of
287 mixed provenance, with material from both the Indus and the peninsula. For the other samples
288 the AFT data argue strongly for the sand at these drilling sites being entirely derived from the
289 Indus River because they are generally much younger than AFT ages from the western margin of
290 peninsular India and broadly consistent with the AFT ages derived from sands that are definitely
291 of Indus derivation (Clift et al. 2004; Clift et al. 2010).

292 Some information can also be derived about where the sediments may be coming from
293 within the possible source ranges if we refer to the bedrock data that has been measured onshore,
294 as summarized in Figure 5. Comparison of these sources and detrital data is only valid for the
295 youngest sediments because young bedrock AFT ages do not inform us about the cooling of
296 these sources in the older geologic past, only the cooling of the rocks now exposed. We note that
297 the different ranges within the Indus basin have a number of distinctive peaks and that some of
298 these are distinct in terms of their AFT age spectra. We note that the Greater and Lesser
299 Himalaya have relatively similar fission track ages, clustering around 3–4 Ma, but with some
300 ranging to ca. 1 Ma, at least in the Sutlej Valley (Thiede et al. 2004), and that these also overlap
301 with ages known from the Karakoram, especially the eastern Karakoram (Wallis et al. 2016) and
302 the Yasil Dome lying in the Karakoram immediately north of the Nanga Parbat Massif (Poupeau
303 et al. 1991). The Karakoram however, also have bedrock AFT ages that range to older values,
304 suggestive of earlier exhumation in at least parts of that block, most notably in the west and their
305 continuation into the Hindu Kush (Zhuang et al. 2018). The very youngest grains are measured
306 around the Nanga Parbat Massif (Zeitler 1985), while the oldest are found in the Transhimalayan
307 Ladakh Batholith (Kirstein et al. 2009) and Deosai Plateau (van der Beek et al. 2009). The
308 Tethyan Himalaya have also yielded older AFT ages in the central Himalaya (Li et al. 2015), but
309 have not been dated within the Indus catchment. Uplift and erosion in the mountains around the
310 Indus Suture and located to the north of the Greater Himalaya are widely accepted to have
311 initiated earlier and then mostly slowed as the exhumation shifted into the Greater and Lesser
312 Himalayan ranges (Searle, 1996).

313 Although many of the measured fission tracks at Nanga Parbat have ages of less than 1
314 Ma (Zeitler et al. 1989), clearly this could not have been the case before 1 Ma, when the fastest

315 cooled grains must have had ages within error of or older than 1 Ma. Lag times could however
316 have been short prior to 1 Ma. Consequently, direct comparison of the modern bedrock with the
317 detrital ages in old sediments is not appropriate for most of our samples. Because the cooling
318 rates of bedrock sources change on timescale of $>10^6$ yr, not only would the AFT ages have been
319 older in the past but we cannot assume that these sources still had the same lag times in the
320 geologic past. Different, higher temperature thermochronometers can constrain exhumation rates
321 during those earlier times and provide clues about lag times. We can however deduce that
322 because many of the grains AFT ages are relatively young (<15 Ma) and their lag times are short
323 that they were probably derived from fast exhuming sources in the Himalaya, Nanga Parbat or
324 Karakoram (Zeitler et al. 1993; Zhuang et al. 2018), rather than in Kohistan, the Transhimalaya
325 or Tethyan Himalaya where uplift and exhumation was mostly older. The cooling histories of
326 these latter sources imply that their AFT lag times would be mostly long during the Late
327 Miocene-Present (Fig. 5) (Kirstein et al. 2009; Krol et al. 1996; Searle 1996). Although some
328 young AFT ages <6.3 Ma have been recorded in the Ladakh Transhimalayan Batholith along the
329 Shyok Suture (Kirstein et al. 2009) these represent quite a small part of that tectonic block.
330 Zircon U-Pb ages from the same IODP sites imply that the Transhimalaya has not been a
331 dominant source during the period targeted by this study (maximum of 28% at 15.5 Ma and this
332 is likely a large overestimate because the Karakoram and Transhimalaya overlap in zircon U-Pb
333 ages) (Clift et al. 2019b).

334 The prevalence of short AFT lag times implies rapid exhumation in the dominant
335 sediment-producing sources close to the time of sedimentation. The AFT data require that little
336 sediment was stored for significant periods of geologic time between erosion in the mountain
337 sources and sedimentation on the Indus submarine fan because the difference/lag between

338 minimum ages and deposition is typically <4 m.y. (75% of samples), representing an upper limit
339 to the storage time. The lag time of a grain largely represents the time between cooling and
340 erosion. While the lag time also includes time spent during sediment transport, study of the
341 Quaternary Indus system indicates transport times of no more than $\sim 10^5$ y for the bulk of the
342 sediment delivered to the deep basin (Clift & Giosan 2014). Some of the sediment may be
343 recycled from foreland basin sedimentary rocks of the Siwalik Group and this would introduce
344 an additional lag into the sediment transport history. Secondary AFT age populations between 15
345 and 38 Ma (Table 2) would fit with this type of recycling. We can discount that these older ages
346 are coming from direct erosion of the slower cooled Ladakh Batholith or Tethyan Himalaya
347 because heavy mineral studies (Garzanti et al. 2005), trace element characteristics of detrital
348 amphiboles (Lee et al. 2003) and zircon U-Pb ages (Alizai et al. 2011) from the trunk Indus
349 River close to the Himalayan front show dominance by the Karakoram (especially the Southern
350 Karakoram Metamorphic Belt) over other sources in the modern upstream basin. That the
351 Siwalik Group sedimentary rocks themselves have not been entirely reset in AFT during burial is
352 known from studies in central Nepal (van der Beek et al. 2006) and these ranges could thus be a
353 source of the older AFT ages measured. Quantifying the amount of recycling out of the Siwalik
354 Ranges is impossible for our data because older grains could come from slow cooling sources or
355 from the Siwalik Group. However, the high abundance of short lag time grains suggests that the
356 degree of this recycling cannot be too large. Rates of incision in modern gorges cutting the
357 Siwalik Group in Nepal have been used to estimate that they account for no more than 15% of
358 the total flux (Lavé & Avouac 2001), while an isotope-based mass balance for the Ganges basin
359 indicates <10% of the mass flux in that drainage is from the Siwalik Group (Wasson 2003). A
360 contribution on that order to the Indus Basin would be consistent with the AFT data presented

361 here. The AFT data by themselves cannot resolve erosion from the Siwaliks, as they share older
362 AFT ages with sources in the Tethyan Himalaya, Kohistan and Transhimalaya.

363 On shorter timescales if sediment was being buffered on the floodplains, in the delta or
364 on the continental shelf then this is expected to have occurred only for a short amount of time,
365 essentially tens of thousands of years (Li et al. 2019). Storage and recycling on million-year
366 timescales would have resulted in longer lag times. When the lag times of our samples are 3–4
367 m.y. or some of this time must have been spent during transport. With the exception of storage
368 and recycling via Siwalik Group foreland sequences discussed above the assumption is that most
369 of this time would be spent prior to exposure and erosion because estimates of transport time in
370 the Quaternary Indus are just 10^5 y for the bulk of the sediment delivered to the deep basin (Clift
371 & Giosan 2014). Modern bedrock AFT data from the Greater and Lesser Himalaya and
372 Karakoram indicate this order of lag time at the present day (Fig. 5), without factoring in much
373 additional transport time. Our data are broadly consistent with the idea of rapidly uplifting
374 mountains being strongly eroded and so supplying most of the sediment into the Indus River
375 during the period of study since 15.5 Ma.

376 Combined Nd isotope and detrital zircon U-Pb age data from bulk sediment samples from
377 Sites U1456 and U1457 show that there was a change in provenance starting around 5.7 Ma
378 (Clift et al. 2019b). This analysis indicates more material coming from the Greater and Lesser
379 Himalaya and relatively less from the Karakoram after this time. The range of lag times in
380 sediments younger than 7.0 Ma is similar to those found at the Indus delta during the phase of
381 strong summer monsoon in the early Holocene, i.e. 2–5 m.y. (Fig. 6), when the provenance
382 constraints indicates that these were preferentially derived from Greater and Lesser Himalayan
383 sources (Clift et al. 2019b). In contrast, sediments older than 7.0 Ma have longer lag times (3.5–

384 8.8 m.y., average 6.0 m.y.) and are inferred to be more derived from the Karakoram, based on
385 their zircon U-Pb age spectra (Fig. 6) (Clift et al. 2019b). The fact that lag times of pre-7.0 Ma
386 samples are longer, like Indus Delta LGM sediments that have an AFT central age of 9 ± 1 Ma
387 (Clift et al. 2010) is consistent with a dominant Karakoram source.

388 That the Nd isotope provenance data change at around the same time as the AFT lag
389 times (after 5.7 Ma; Fig. 6) supports the idea that a change in provenance may account for at
390 least part of the changing AFT lag times at that time. The absence of the very short lag time
391 samples does mean that after 5.7 Ma there are no longer any significant fast eroding ranges in the
392 catchment. As noted above, the Crystalline Inner Lesser Himalaya are known to be experiencing
393 unroofing after ~ 6 Ma, at least in the vicinity of the Beas River catchment (Najman et al. 2009)
394 and the shift in the general character of the AFT age populations after 5.7 Ma may in large part
395 simply reflect more sediment delivery from the Greater and Lesser Himalayas, potentially related
396 to tectonic imbrication and rock uplift (Bollinger et al. 2004; Huyghe et al. 2001; Webb 2013).
397 Such a shift is consistent with the evolving provenance data in Laxmi Basin (Clift et al. 2019b).
398 The structural reconstructions of Webb (2013) for the western Himalaya propose that both the
399 Greater and Lesser Himalaya remained buried under the Tethyan Himalaya until after 5.4 Ma.
400 This would imply that the source of rapidly cooled grains before that time would be from the
401 Karakoram and Tethyan Himalaya.

402 The AFT ages can be used to constrain changing rates of exhumation in the bedrock
403 sources. Comparing depositional age against the AFT minimum age populations allows us to
404 assess the lag time between cooling of bedrock sources as they passed through the 60–110°C
405 partial annealing zone and their final deposition in the deep water of the Indian Ocean (Fig. 6). In
406 our analysis we further compare our results with those similar aged fluvial sedimentary rocks

407 from the Siwalik Group in Western and Central Nepal (van der Beek et al. 2006), as well as from
408 the Bengal Fan collected by ODP Leg 116 (Corrigan & Crowley 1990). It is clear that many of
409 these minimum age groups have relatively short lag times, which indicates fast cooling and
410 exhumation of bedrock sources. We note that both the oldest (15.5 Ma) sample from the Laxmi
411 Basin and a slightly younger sample from the Bengal Fan show lag times close to 4 m.y. in the
412 middle Miocene. This would imply exhumation rates of 1.1–1.4 km/m.y. assuming 25–35°C/km
413 geothermal gradients.

414 Unfortunately, we have little information between that time and ~8.5 Ma when the next
415 youngest dateable sandy sediment was deposited and preserved at the drilling sites. Although one
416 of the minimum age groups still lags by ~4.2 m.y., we note that this there is some scatter to
417 longer lag times of up to 8.8 m.y. between 8.5 and 7.0 Ma and with large uncertainties.
418 Combined zircon U-Pb (40–70 and 70–120 Ma grains) and bulk sediment Nd isotope (ϵ_{Nd} values
419 > -10) provenance data indicate that much of the sediment at that time was derived from the
420 Karakoram (Clift et al. 2019b). The zircon U-Pb budget over-represents the net flux from the
421 Himalaya because these bedrocks are >2.2 time more fertile with regard to zircon than the
422 Karakoram and Transhimalaya.

423 After 7.0 Ma lag times shortened significantly. Three samples from the Laxmi Basin
424 drilling sites are within error of the depositional age between 7.0 and 5.7 Ma, requiring
425 exhumation rates that were so rapid that we are unable to constrain the duration between cooling
426 through the PAZ (60–110°C) and sedimentation, i.e., lag times close to zero. This implies a
427 maximum rate of cooling in the sources at that time. All three of the fast cooling samples have
428 accompanying zircon U-Pb ages that show that they continue a trend towards more Himalayan
429 erosion but that there is not a sharp contrast with the sediment deposited before 7.0 Ma. After 5.7

430 Ma the change in Nd isotopes is especially marked implies that a change in provenance may be
431 responsible for the slowing of exhumation rates. Nonetheless, one sample, U1457C-43R-1 55-63
432 cm, deposited at 5.78 Ma, has a minimum age lag time 3.13 m.y., longer than the others. This
433 implies that not all sources were supplying large volumes of sediment at all times and that not all
434 bedrock sources were exhuming so quickly.

435 Although provenance data indicate mostly Karakoram sources, these rapidly cooled
436 grains could also be derived from the Himalaya. Zircon U-Pb ages allow us to discriminate
437 between erosion of Karakoram (40–120 Ma and Himalayan (>300 Ma) sources but we only
438 know that these are the largest sources at that time. However, the zircon ages only apply to these
439 minerals and the provenance cannot be transferred to the apatites so that we only know that there
440 were rapidly cooling areas between 7.0 and 5.7 Ma but not which range they are located in.
441 However, because there are large numbers of grains in the minimum age group it might
442 reasonably be expected that these are derived from bedrocks sources that also supply large
443 volumes of other mineral types. Between 7.0 and 5.7 Ma the longest lag time was 3.13 m.y. in
444 the sediment deposited at 5.87 Ma. This indicates an average cooling rate of at least
445 $35.1 \pm 9.7^\circ\text{C/m.y.}$, faster than the cooling rates of 12.5 to 26.1°C/m.y. between 8.2 and 7.0 Ma.
446 These are faster rates than those recorded in the Siwalik Group from Nepal (van der Beek et al.
447 2006), as well as sparse data from the Bengal Fan (Corrigan & Crowley 1990), although they are
448 within the uncertainties of the peak rates in Nepal at that time. However, in Nepal the sources
449 must have been Himalayan, not Karakoram. In the youngest part of the section (<4 Ma), which is
450 more dominated by Himalaya erosion (Clift et al. 2019b) these very short lag times are not
451 visible and are always more than 1.93 m.y., equivalent to approximate exhumation rates of ~2.3–
452 1.6 km/m.y. The moderate exhumation rates after 4 Ma compare with data from both the Bengal

453 Fan and from the Nepalese part of the Himalayan foreland. Both these sediment sequences are
454 dominated by Himalayan erosion (Bouquillon et al. 1990). Slowing of exhumation in the Indus
455 basin after 5.7 Ma is consistent with data from western Nepal (Karnali), but the slowing from
456 peak rates at 7.0 to 5.7 Ma is in contrast to conclusions of work from central Nepal (Surai and
457 Tinau Khola) that argued for relatively steady state cooling in that part of the mountain range
458 (van der Beek et al. 2006). The very youngest sample deposited at 930 ka stands out as having by
459 far the largest lag time and is inferred have a unique source, likely a mixture of sediment from
460 the Indus River and Peninsular India.

461 We can compare this pattern of accelerating exhumation before 7.0 Ma and then slowing
462 after 5.7 Ma with the climatic history (Fig. 6), while recognizing the shift in provenance that is
463 occurring at the same time. One of the most popular long-term proxies for monsoon intensity in
464 the Arabian Sea is the relative abundance of *G. Bulloides* offshore the margin of Arabia. The
465 abundance of *G. Bulloides* is largely a function of the availability of nutrients derived from
466 upwelling caused by the summer monsoon rains (Curry et al. 1992). There is little evidence for
467 such strong upwelling prior to around 13 Ma (Betzler et al. 2016). A general intensification of
468 upwelling is noted after 5.3 and 3.0 Ma (Gupta et al. 2015; Huang et al. 2007) (Fig. 6). However,
469 upwelling is not a direct proxy of rainfall and this apparent intensification does not reflect the
470 delivery of summer rains to the mountain front, because this proxy does not correlate with other
471 climatically sensitive indicators (Clift 2017).

472 Stable oxygen isotope data from the foreland basin instead agree with chemical
473 weathering data from the South China and Arabian Seas in arguing for relatively wet conditions
474 in the middle Miocene between 10 and 12 Ma (Dettman et al. 2001), followed by a decrease in
475 humidity particularly after around 6–8 Ma (Clift 2017; Singh et al. 2011). Moisture delivery to

476 this area from the winter Westerlies is also reconstructed to reduce around 7 Ma (Vögeli et al.
477 2017). The increasing lag time seen in the minimum age populations after 5.7 Ma would be
478 consistent with slower erosion and could be linked to weaker monsoon rainfall. Weaker monsoon
479 and Westerly rains would also reduce discharge and potentially slow the transport of sediment
480 across the flood plains. Increased aridity is consistent with decreasing strength of chemical
481 weathering seen in Indus Marine A-1 located on the Indus shelf (Clift et al. 2008), as well as Site
482 U1456 (Clift et al. 2019a), but largely postdates the carbon isotope transition from 8 to 6 Ma in
483 the foreland basin (Quade et al. 1989).

484 The acceleration in exhumation rates from 7.8 to 7.0 Ma generally coincides with the
485 climatic drying, which may seem counterintuitive. However, this also assumes that stronger
486 rains, sometimes modulated through glaciation, always increase erosion. There is evidence that
487 drier conditions, especially when this involves heightened seasonality, can increase erosion
488 provided the drying is not too extreme, but sufficient to reduce vegetation cover that reduces soil
489 erosion (Giosan et al. 2017). There is no evidence that the period of fast erosion at 5.7–7.0 Ma
490 was caused by faster India and Asia convergence. Indeed, convergence rates appear to have
491 slowed gradually during the Cenozoic (Clark 2012).

492

493 **Conclusions**

494 Apatite fission track ages derived from turbidite sediments from IODP Sites U1456 and
495 U1457 in the Laxmi Basin, eastern Arabian Sea, provide an opportunity to reconstruct changing
496 exhumation rates in the western Himalaya and Karakoram since 15.5 Ma, and especially since 9
497 Ma. AFT ages are mostly <50 Ma and demonstrate that the sediment is derived from the Indus

498 River, not peninsular India, except in the case of the youngest sample, deposited at 0.93 Ma.
499 Moreover, most samples show minimum age populations that are only slightly older than the
500 depositional age, implying fast rates of exhumation in the sources through this time. Lag times of
501 ~4 m.y. in the Middle Miocene imply exhumation rates of 1.1–1.4 km/m.y. After a period of
502 longer lag times (~6 m.y.) between 8.5 and 7.8 Ma these reach a minimum from 7.0 to 5.7 Ma,
503 when lag times were within error of zero. Provenance U-Pb zircon and Nd isotope data indicate
504 erosion dominantly in the Karakoram, but the AFT ages could have also come from Himalayan
505 sources, which were also important contributors at this time. The AFT data alone do not allow us
506 to discriminate which of the two ranges contained the fast exhuming sources. After 5.7 Ma lag
507 times lengthened to ~4.5 Ma, and exhumation rates slowed to 2.3–1.6 km/m.y. at the same time
508 that sediment supply came progressively more from the Himalaya and relatively less from the
509 Karakoram.

510 The time of peak exhumation correlates with the transition to a drier climate in the
511 foreland basin and of a weakening Westerly Jet. Erosion rates since 5.7 Ma are comparable or
512 slightly faster than those seen in the Nepalese parts of the Himalaya and the Bengal Fan. Slowing
513 exhumation rates after 5.7 Ma correlate with a drying climate and weaker summer monsoon rains
514 in the Late Miocene. There is a general shift in the AFT age populations from longer lag times,
515 more similar to the glacial era Indus River and associated with dominant erosion in the
516 Karakoram prior to 7 Ma, to shorter lag times and more erosion of the Himalaya, similar to the
517 Holocene Indus River after 5.7 Ma. The acceleration of exhumation as the climate dried between
518 7.8 and 7.0 Ma seems to imply a dominant tectonic control of erosion. The AFT data support
519 models that imply a non-linear relationship between summer monsoon rain strength and the
520 erosion of the western Himalaya.

521

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527 and an anonymous reviewer.

528

529

530 **Figure Captions**

531

532 Figure 1. Shaded bathymetric and topographic map of the Arabian Sea area showing the location
533 of the drilling sites within the Laxmi Basin. Map also shows the primary source ranges and the
534 major tributary systems of the Indus River, as well as smaller peninsular Indian rivers that may
535 have provided material to the drill sites. Magnetic anomalies are from Miles et al. (1993). KK =
536 Karakoram; NP = Nanga Parbat; K = Karnali; S = Surai Khola; T = Tinau Khola.

537

538 Figure 2. Simplified lithologic logs of the two drill sites considered in this study. Black arrows
539 show the location of the samples analysed. MTD = Mass Transport Deposit.

540

541 Figure 3. Radial plots and associated KDE spectra (abanico plots) showing the range of apatite
542 fission track ages for each of the samples considered within the study (Galbraith 1990). N_s —
543 number of spontaneous fission tracks; N_i —number of induced tracks. Single ages are plotted
544 with standard errors according to their precision ($1/\sigma$ on the 'x' axis). The error attached to each
545 plotted point is standardized on the y scale. The value of the age and the 2σ uncertainty can be
546 read off the radial axis by extrapolating lines from point 0,0 through the plotted age.

547

548 Figure 4. Cross plot of numbers of grains compared to minimum ages with 2σ uncertainties
549 displayed. There is no correspondence between the numbers of grains and the minimum age that
550 might bias the result of the lag time analysis.

551

552 Figure 5. KDE plots for the apatite fission track central ages of potential bedrock sources within
553 the headwaters of the Indus basin. Nanga Parbat data are from Warner et al. (1993), and Zeitler
554 (1985). Greater Himalaya data are from Kumar et al. (1995), Jain et al. (2000) and Thiede et al.
555 (2004). Lesser Himalaya data are from Thiede et al. (2004) and Vannay et al. (2004). Karakoram
556 data are from Foster et al. (1994), Zeitler (1985), Wallis et al. (2016) and Poupeau et al. (1991).
557 Kohistan data are from Zeitler (1985) and Zeilinger et al. (2001). Transhimalaya data are from
558 Kirstein et al. (2009; 2006), and Clift et al. (2002a). Tethyan Himalaya data are from Li et al.
559 (2015) and unpublished from Andrew Carter (UCL, 2017). Indian Peninsula data are from
560 Gunnell et al. (2003) and Kalaswad et al. (1993).

561

562 Figure 6. Lag time plot of detrital apatite fission track minimum ages showing the lag time
563 between the cooling and depositional ages. Note the minimum lag time achieved between 9 and
564 6 Ma. Siwalik data from Nepal is from van der Beek et al. (2006), Bengal Fan data is from
565 Corrigan and Crowley (1990). Monsoon records of *G. Bulloides* from Huang et al. (2007),
566 foreland basin $\delta^{14}\text{C}$ record from Quade et al. (1989). Sediment budget for Indus Fan from Clift
567 (2006). Evolution in the age spectra of zircon U-Pb ages and ϵ_{Nd} values is from Clift et al.
568 (2019b). Stippled area shows the time of the climatic transition to drier conditions in the foreland
569 basin.

570

571 Table 1. List of the samples with their depths and calculated depositional ages. Those samples
572 also analysed for detrital U-Pb zircon dating by Clift et al. (2019b) are highlighted.

573

574 Table 2. Summary of apatite fission track analytical data. Track densities are ($\times 10^6$ tr cm^{-2})
575 numbers of tracks counted (N) shown in brackets. Analyses by external detector method using
576 0.5 for the $4\pi/2\pi$ geometry correction factor. Ages calculated using dosimeter glass CN-5;
577 (apatite) $\zeta_{\text{CN5}} = 338 \pm 5$; calibrated by multiple analyses of IUGS apatite and zircon age standards
578 (Hurford 1990). $P\chi^2$ is probability for obtaining χ^2 value for ν degrees of freedom, where $\nu = \text{no.}$
579 crystals – 1. Central age is a modal age, weighted for different precisions of individual crystals
580 (see Galbraith (1990)). Minimum age model after Galbraith (2005). P2 used peak fitting
581 algorithm of Galbraith and Green (1990) where there are > 10 grains.

582

583 Supplementary Table 1. Single grain apatite fission track data.

584

585

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Lab No.	IODP Sample Name	Depositional Age (Ma)	Depth (mbsf)	AFT Minimum Age (Ma)	2 σ (Ma)	Number of grains	Zircon U-Pb ages
134-1	U1456A-11H-6 60-69 cm	0.93	97.60	20.70	3.80	24	Yes
134-2	U1456A-26F-3 50-58 cm	1.32	185.91	3.60	0.85	62	
134-3	U1456A-51F-3 100-110 cm	1.56	302.09	3.90	1.40	44	Yes
134-4	U1456A-61F-3 40-50 cm	1.92	345.32	6.50	1.10	45	Yes
177-1	U1456A-70F-2 10-16 cm	3.02	386.73	5.70	1.50	75	Yes
177-12	U1457C-31R-1 94-100 cm	3.17	474.25	5.10	1.80	52	
177-13	U1457C-33R-3 10-17 cm	3.43	499.10	6.40	1.20	49	Yes
177-2	U1456C-45X-3 45-51 cm	3.57	459.09	8.48	0.75	65	
134-6	U1456D-5R-1 12-20 cm	5.72	487.98	9.30	2.20	50	Yes
177-14	U1457C-41R-2 20-26 cm	5.78	572.16	5.91	0.83	46	
177-15	U1457C-42R-1 80-88 cm	5.82	580.40	6.40	1.10	55	
177-16	U1457C-43R-1 55-63 cm	5.87	590.53	9.00	1.20	57	Yes
177-3	U1456D-12R-1 30-36 cm	7.00	556.45	6.60	1.50	52	
177-4	U1456D-13R-1 30-38 cm	7.07	566.35	13.20	7.30	30	Yes
177-5	U1456D-15R-1 55-61 cm	7.28	586.00	15.80	1.90	50	
177-6	U1456D-19R-2 20-26 cm	7.66	625.73	11.90	1.80	40	
177-17	U1457C-51R-4 80-88 cm	7.78	675.16	12.00	3.20	51	
134-7	U1456D-22R-1 73-83 cm	7.84	653.50	15.48	0.97	69	Yes
134-10	U1457C-61R-1 8-18 cm	7.99	769.36	14.00	3.10	42	
177-8	U1456D-26R-2 37-43 cm	8.09	693.78	14.90	1.60	55	
177-9	U1456D-27R-2 100-106 cm	8.15	704.43	16.97	0.98	69	
177-10	U1456D-28R-1 40-46 cm	8.20	711.98	14.20	1.80	72	
134-8	U1456D-29R-2 24-34 cm	8.27	722.60	11.80	5.30	64	Yes
134-9	U1456E-19R-3 10-20 cm	15.58	1102.95	20.20	1.40	75	Yes

918

919 **Table 1**

920

	Lab No	Sample	Dep. Age	No. of	Dosimeter								Central Age	Minimum Age	P2 Age
			(Ma)	grains	ρ_d	Nd	ρ_s	Ns	ρ_i	Ni	$P\chi^2$	RE%	(Ma)	(Ma)	(Ma)
A	134-1	U1456A-11H-6 60-69 cm	0.93	24	1.583	4388	0.798	218	3.858	1440	0	111	61.2±14.9	20.7±3.8	223±28
B	134-2	U1456A-26F-3 50-58 cm	1.32	62	1.583	4388	0.108	308	3.555	11836	0	79	7.3±0.9	3.6±0.9	13.4±1.3
C	134-3	U1456A-51F-3 100-110 cm	1.56	44	1.583	4388	0.191	298	6.856	12192	0	70	6.8±0.9	3.9±1.4	7.2±0.9
D	134-4	U1456A-61F-3 40-50 cm	1.92	45	1.583	4388	0.178	349	5.498	11649	0	35.2	8.1±35.2	6.5±1.1	
E	177-1	U1456A-70F-2 10-16 cm	3.02	75	1.215	3367	0.206	446	4.539	11389	0	54.2	8.2±0.7	5.7±1.5	15.5±2.3
F	177-12	U1457C-31R-1 94-100 cm	3.17	75	1.215	3367	0.171	326	4.710	10359	0	51.5	6.8±0.6	5.1±1.8	12.7±2.1
G	177-13	U1457C-33R-3 10-17 cm	3.43	49	1.215	3367	0.211	313	4.528	8601	0	50.8	7.7±0.8	6.4±1.2	
H	177-2	U1456C-45X-3 45-51 cm	3.57	65	1.215	3367	0.349	474	4.737	9089	0	160	12.9±2.7	8.5±0.8	
I	134-6	U1456D-5R-1 12-20 cm	5.72	50	1.583	4388	0.272	314	6.211	7830	0	42.4	11.2±1.0	9.3±2.2	
J	177-14	U1457C-41R-2 20-26 cm	5.78	46	1.215	3367	0.186	236	3.801	6317	0	180	11.4±3.1	5.9±0.8	
K	177-15	U1457C-42R-1 80-88 cm	5.82	55	1.215	3367	0.179	361	4.073	9719	0	160	7.8±0.8	6.4±1.1	15.9±2.6
L	177-16	U1457C-43R-1 55-63 cm	5.87	80	1.215	3367	0.389	528	5.048	8747	0	12.4	13.7±1.6	9.0±1.2	29.4±1.2
M	177-3	U1456D-12R-1 30-36 cm	7.00	52	1.215	3367	0.241	347	4.004	6997	0	53.8	10.7±1.0	6.6±1.7	17.7±1.7
N	177-4	U1456D-13R-1 30-38 cm	7.07	30	1.215	3367	0.297	124	5.000	2061	2.1	44.7	11.4±1.5	11.4±1.5	
O	177-5	U1456D-15R-1 55-61 cm	7.28	50	1.215	3367	0.362	372	3.718	4683	0	39.2	16.5±1.3	15.8±1.9	
P	177-6	U1456D-19R-2 20-26 cm	7.66	40	1.215	3367	0.546	457	4.714	4931	0	73.4	19.9±2.6	11.9±1.8	28.0±4.7
Q	177-17	U1457C-51R-4 80-88 cm	7.78	51	1.215	3367	0.326	430	4.140	5605	0	40	14.7±1.2	12.0±3.2	19.9±1.6
R	134-7	U1456D-22R-1 73-83 cm	7.84	80	1.583	4388	0.424	799	6.226	12387	0	44.6	18.6±1.2	15.5±0.9	
S	134-10	U1457C-61R-1 8-18 cm	7.99	42	1.583	4388	0.353	468	5.490	7570	0	14.3	16.1±1.0	14.0±3.1	
T	177-8	U1456D-26R-2 37-43 cm	8.09	55	1.215	3367	0.337	403	3.651	5056	0	48.9	18.4±1.7	14.9±1.6	
U	177-9	U1456D-27R-2 100-106 cm	8.15	92	1.215	3367	0.309	605	3.710	7958	0	41.8	16.0±1.0	16.9±0.9	
V	177-10	U1456D-28R-1 40-46 cm	8.20	72	1.215	3367	0.499	639	5.203	7453	0	73.3	18.4±1.8	14.2±1.8	21.1±1.9
W	134-8	U1456D-29R-2 24-34 cm	8.27	72	1.583	4388	0.424	639	5.508	9347	0	48.6	19.3±1.4	11.8±5.3	38.8±3.8
X	134-9	U1456E-19R-3 10-20 cm	15.58	75	1.583	4388	0.462	873	4.957	9653	0	55.9	25.9±2.0	20.2±1.4	

Table 2. Summary of apatite fission track analytical data. Track densities are ($\times 10^6$ tr cm^{-2}) numbers of tracks counted (N) shown in brackets. Analyses by external detector method using 0.5 for the 4p/2p geometry correction factor. Ages calculated using dosimeter glass CN-5; (apatite) $z\text{CN}5 = 338 \pm 5$; calibrated by multiple analyses of IUGS apatite and zircon age standards (Hurford, 1990). $P\chi^2$ is probability for obtaining χ^2 value for ν degrees of freedom, where $\nu = \text{no. crystals} - 1$. Central age is a modal age, weighted for different precisions of individual crystals (see Galbraith (1990)). Minimum age model after Galbraith (2005). P2 used peak fitting algorithm of Galbraith and Green, (1990) where there are >10 grains.

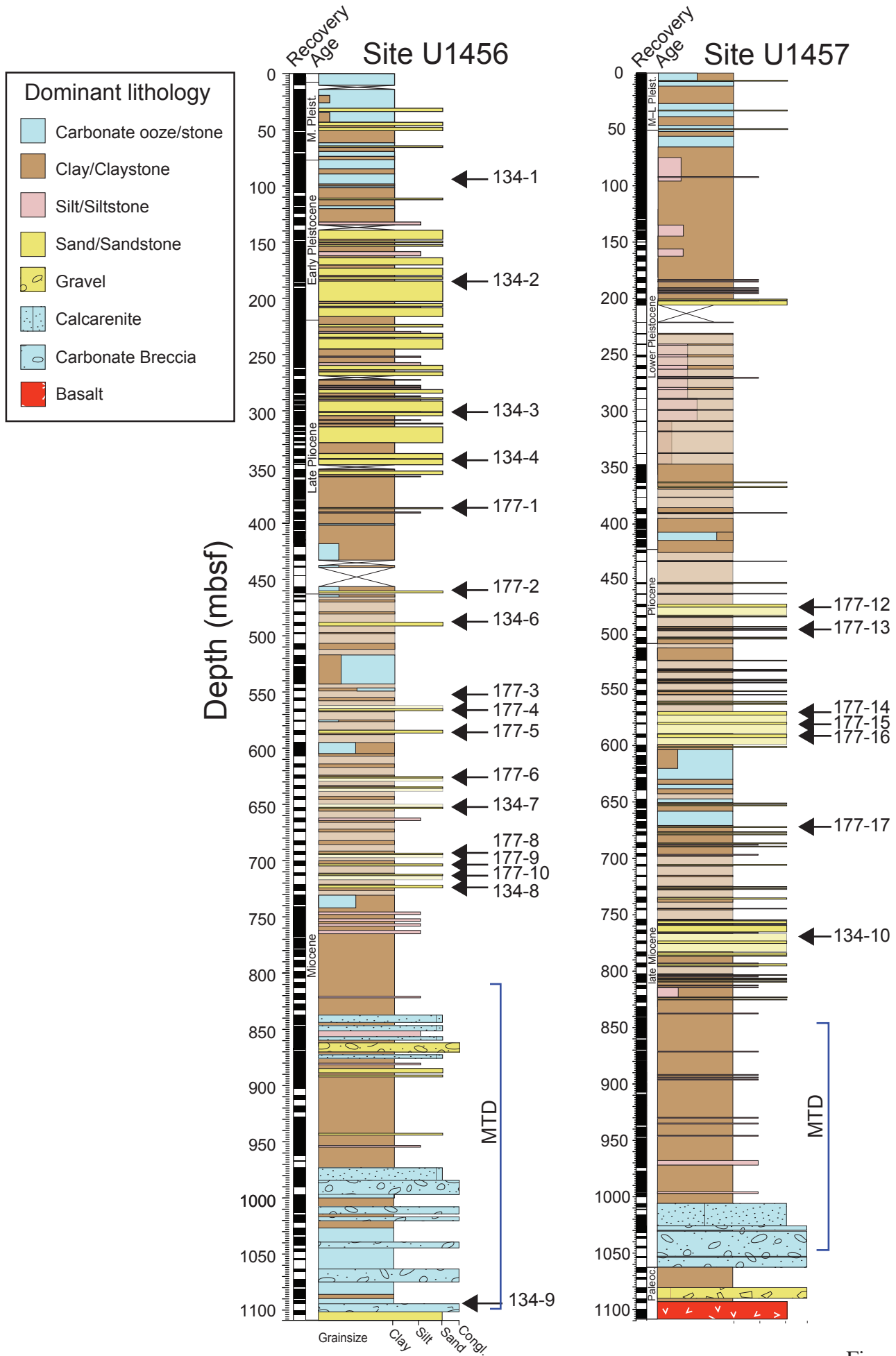


Figure 2
Zhou et al.

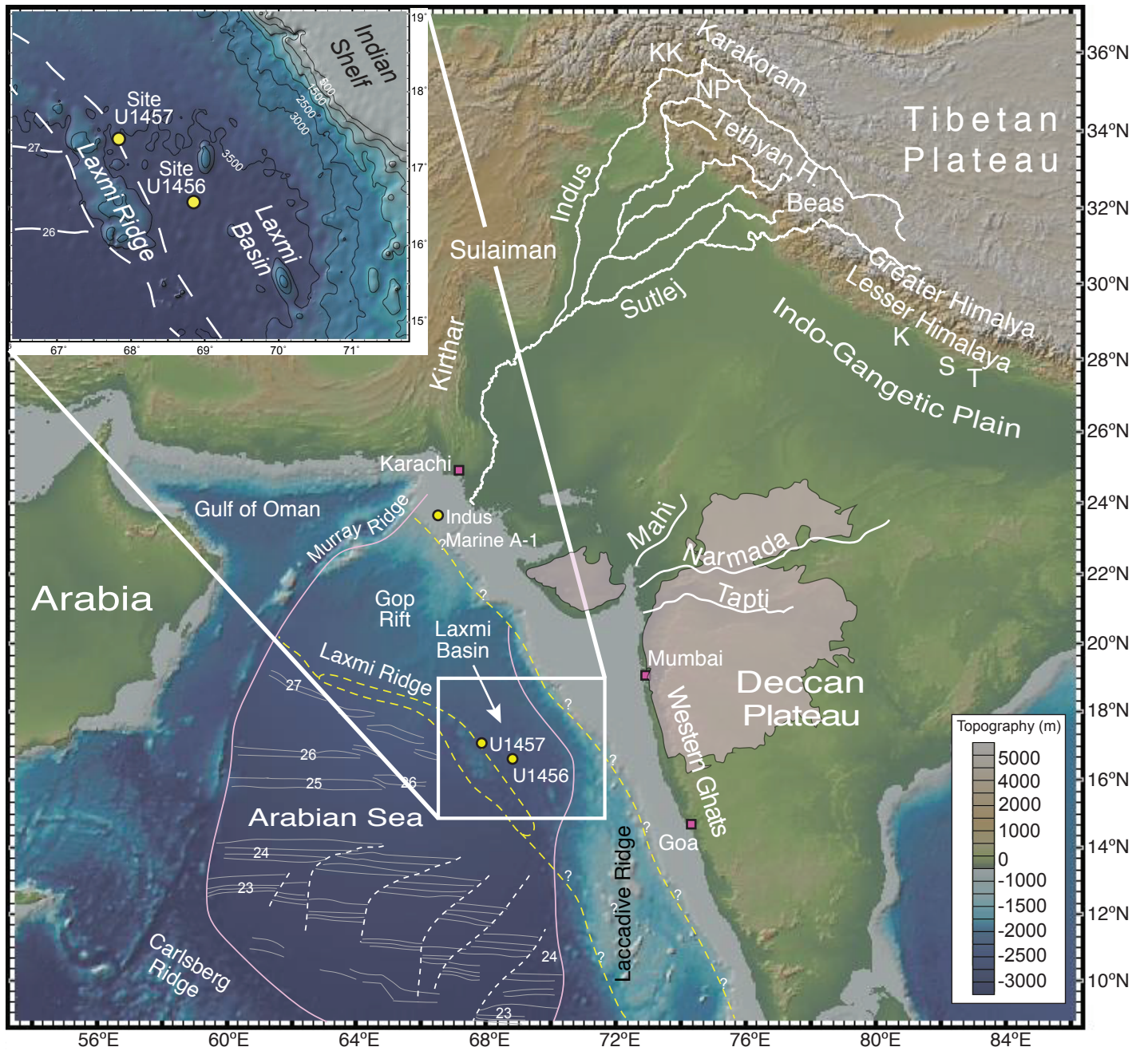


Figure 1
Zhou et al.

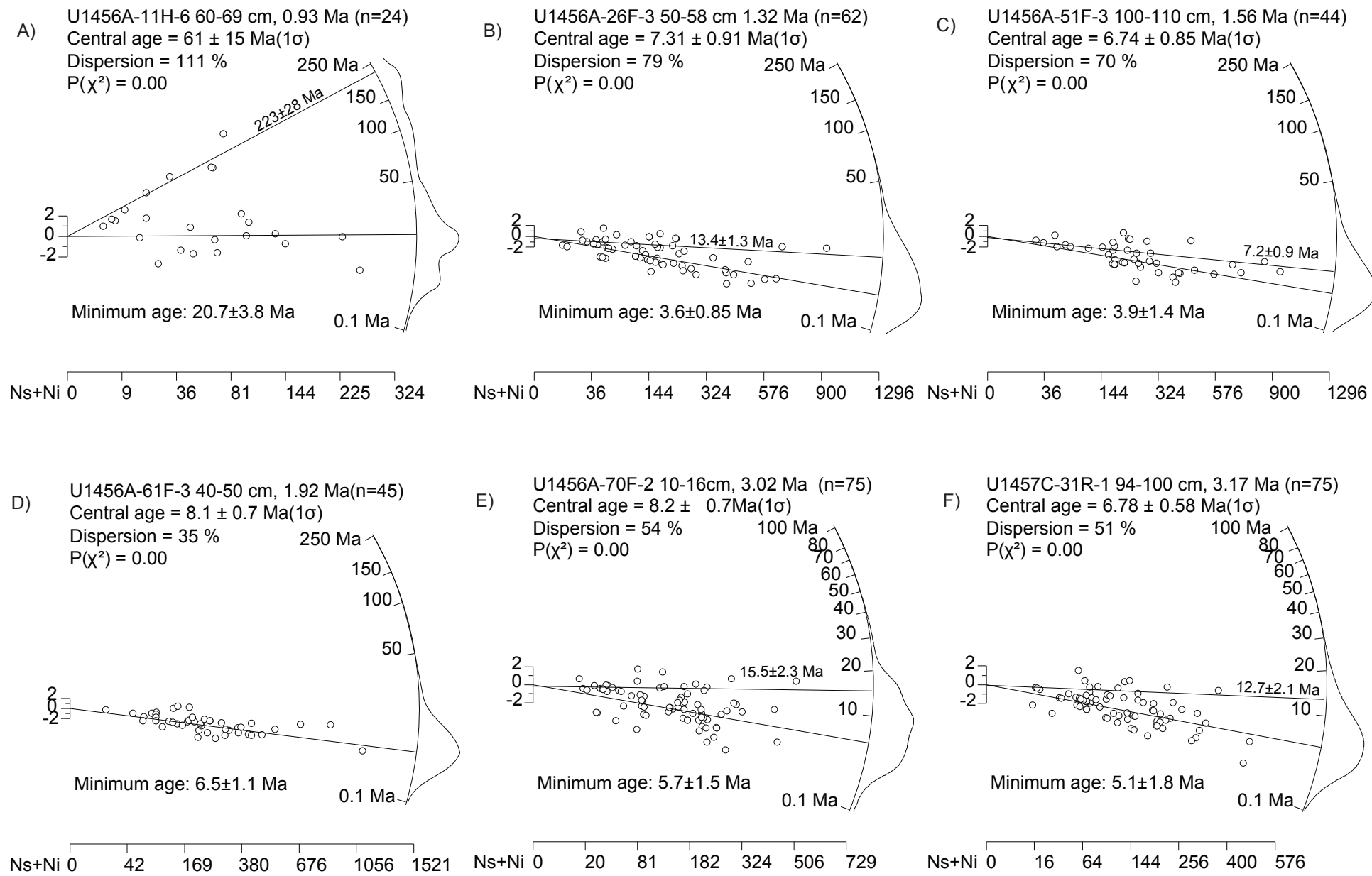
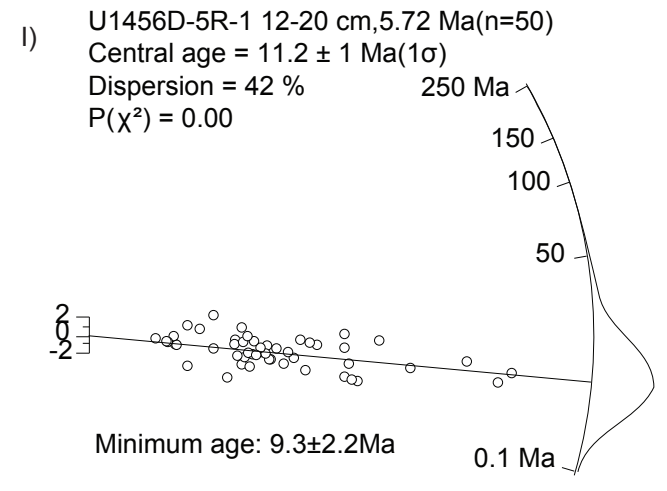
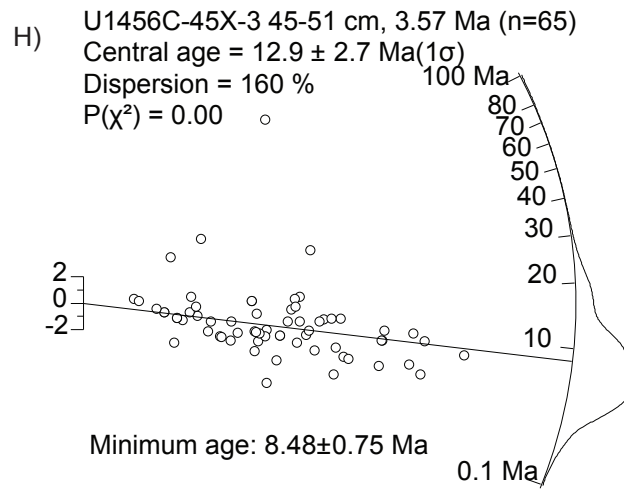
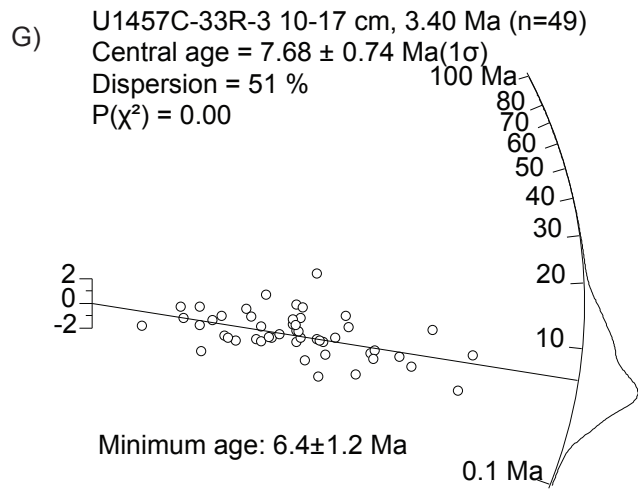


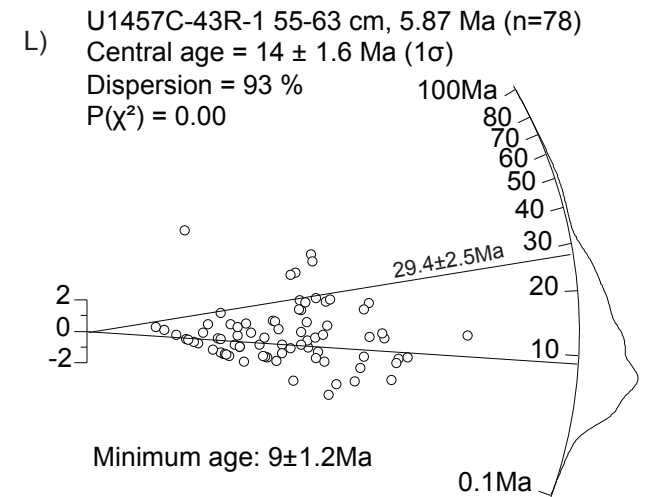
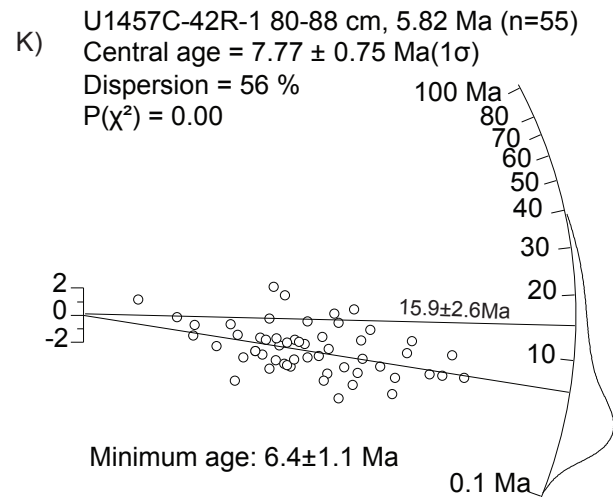
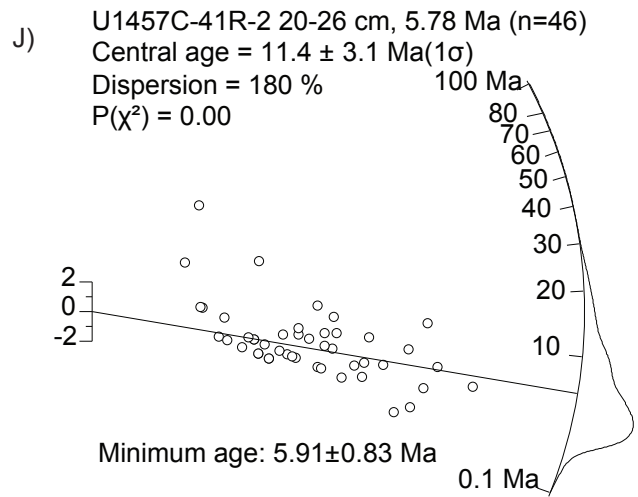
Figure 3
 Zhou et al.



Ns+Ni 0 20 81 182 324 506 729

Ns+Ni 0 20 81 182 324 506 729

Ns+Ni 0 25 100 225 400 625 900



Ns+Ni 0 16 64 144 256 400 576

Ns+Ni 0 16 64 144 256 400 576

Ns+Ni 0 12 49 110 196 306 441

Figure 3
 Zhou et al.

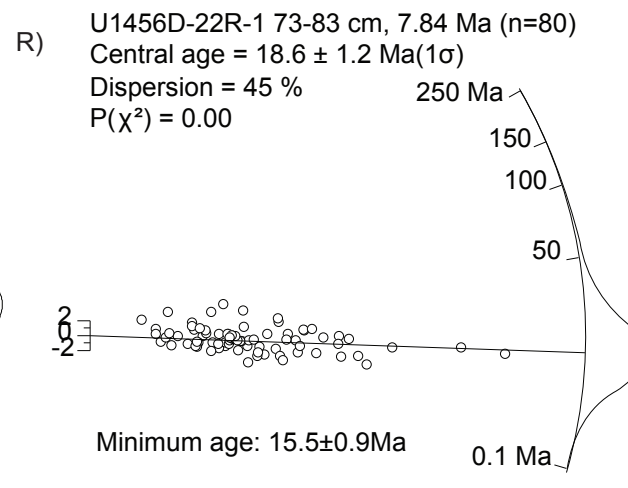
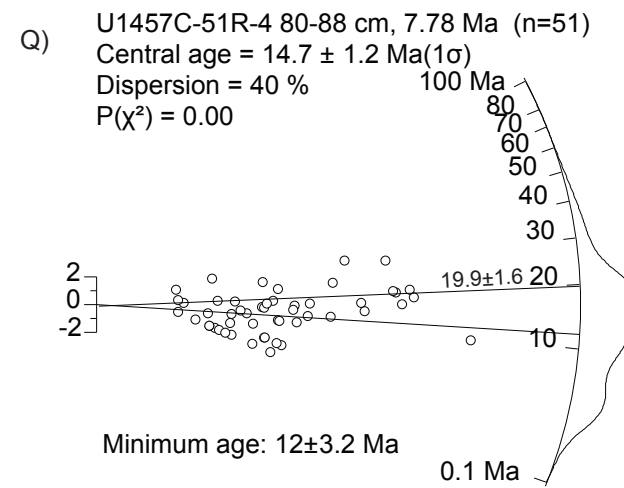
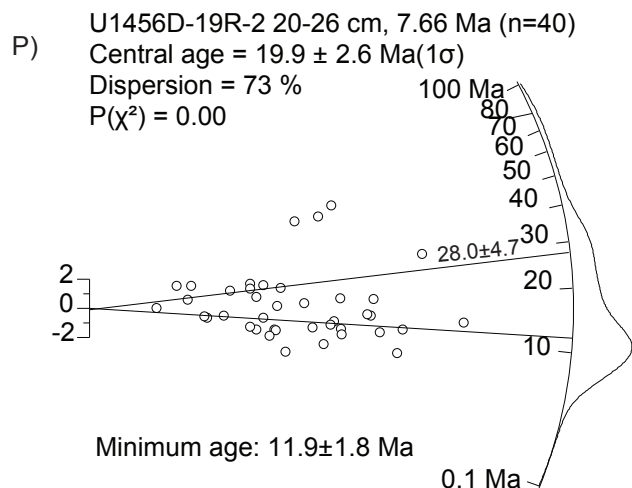
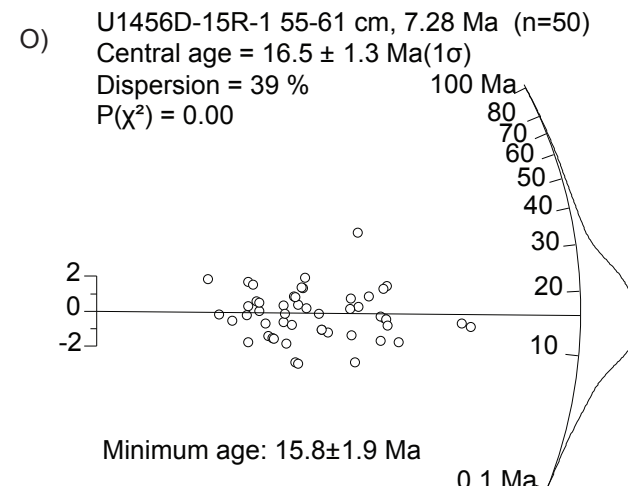
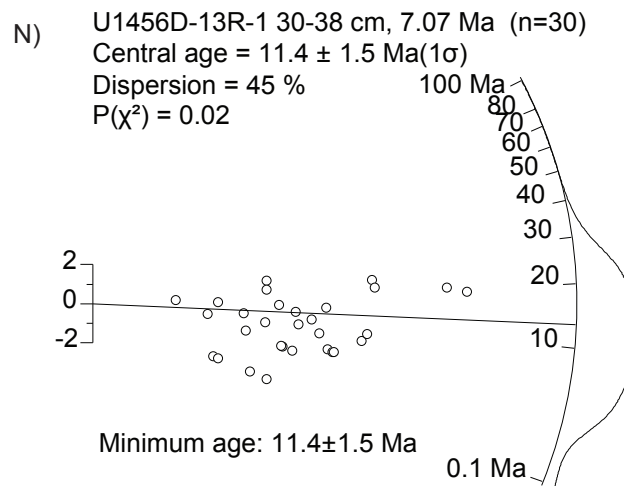
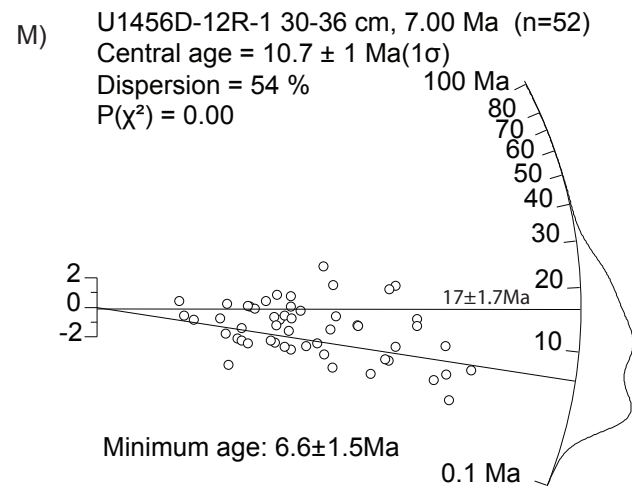


Figure 3
 Zhou et al.

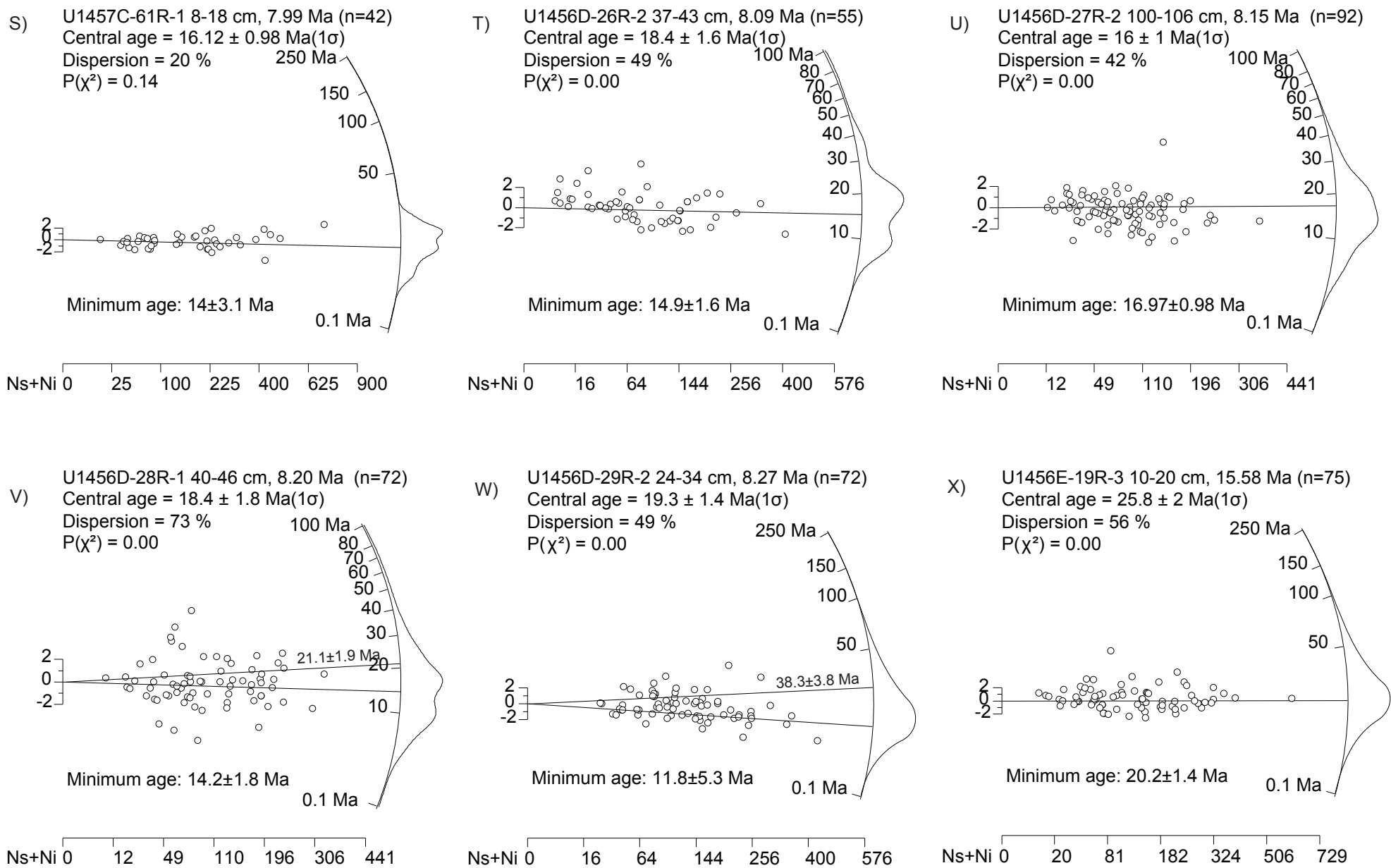


Figure 3
 Zhou et al.

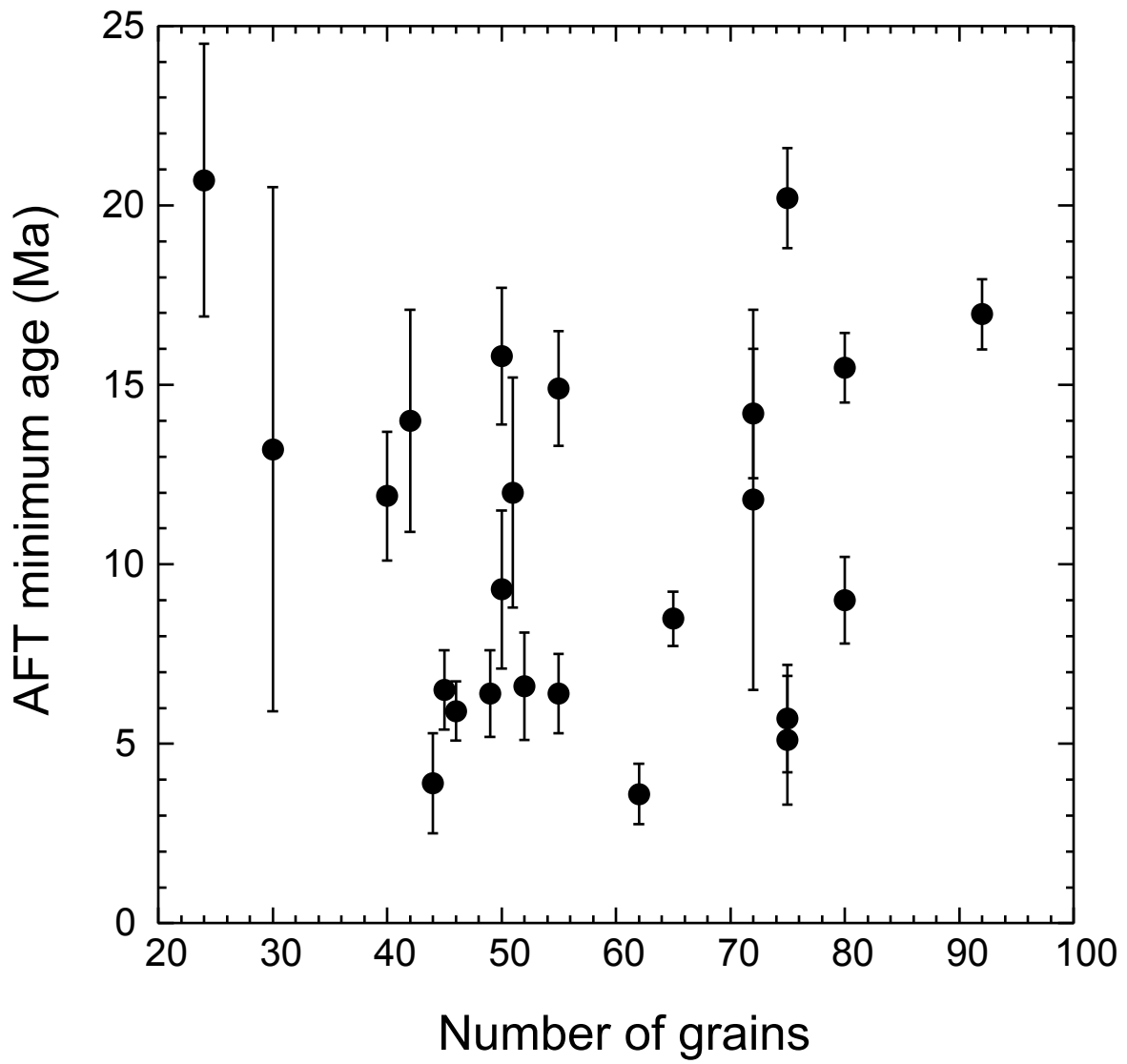


Figure 4
Zhou et al.

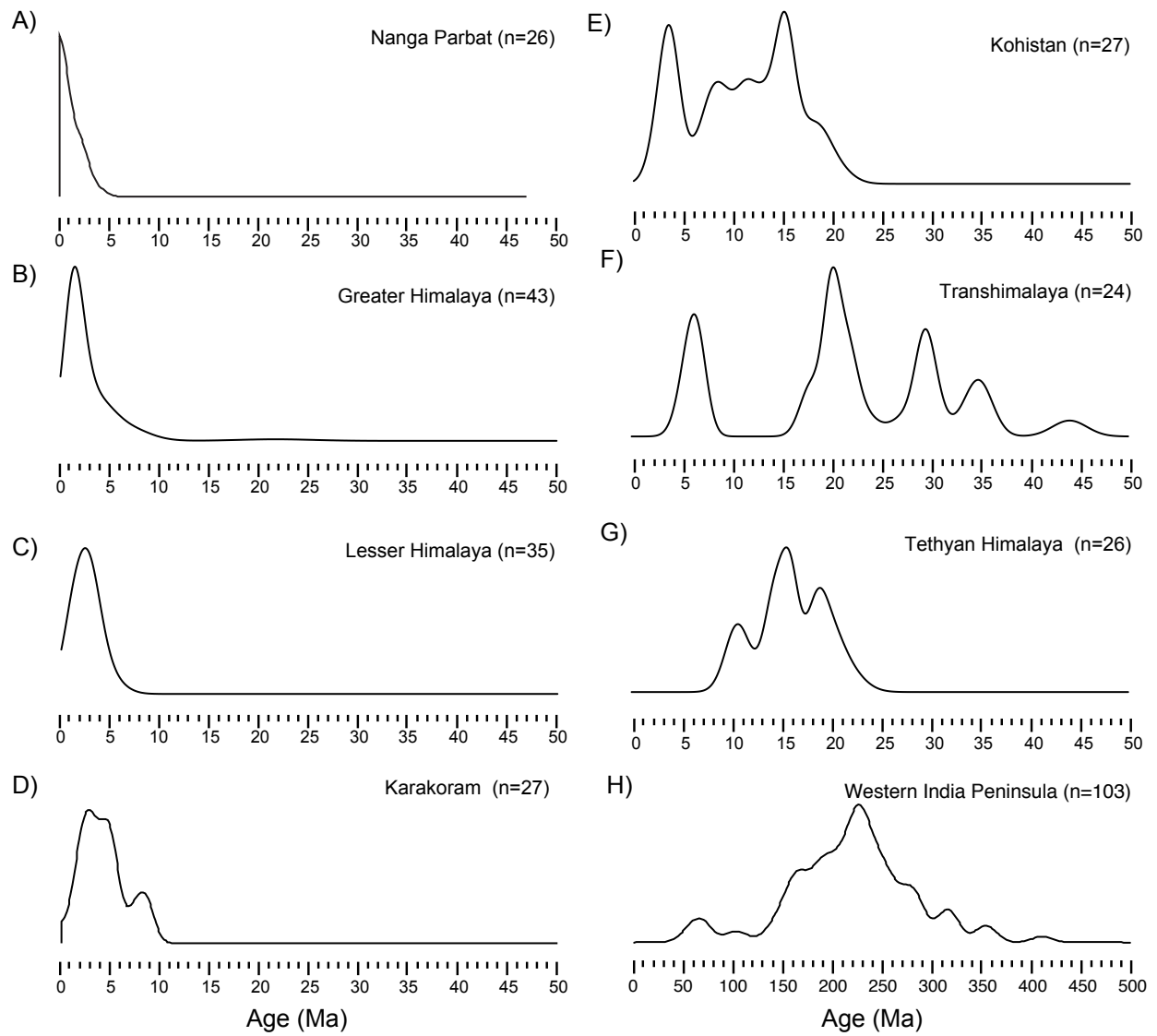


Figure 5
Zhou et al.

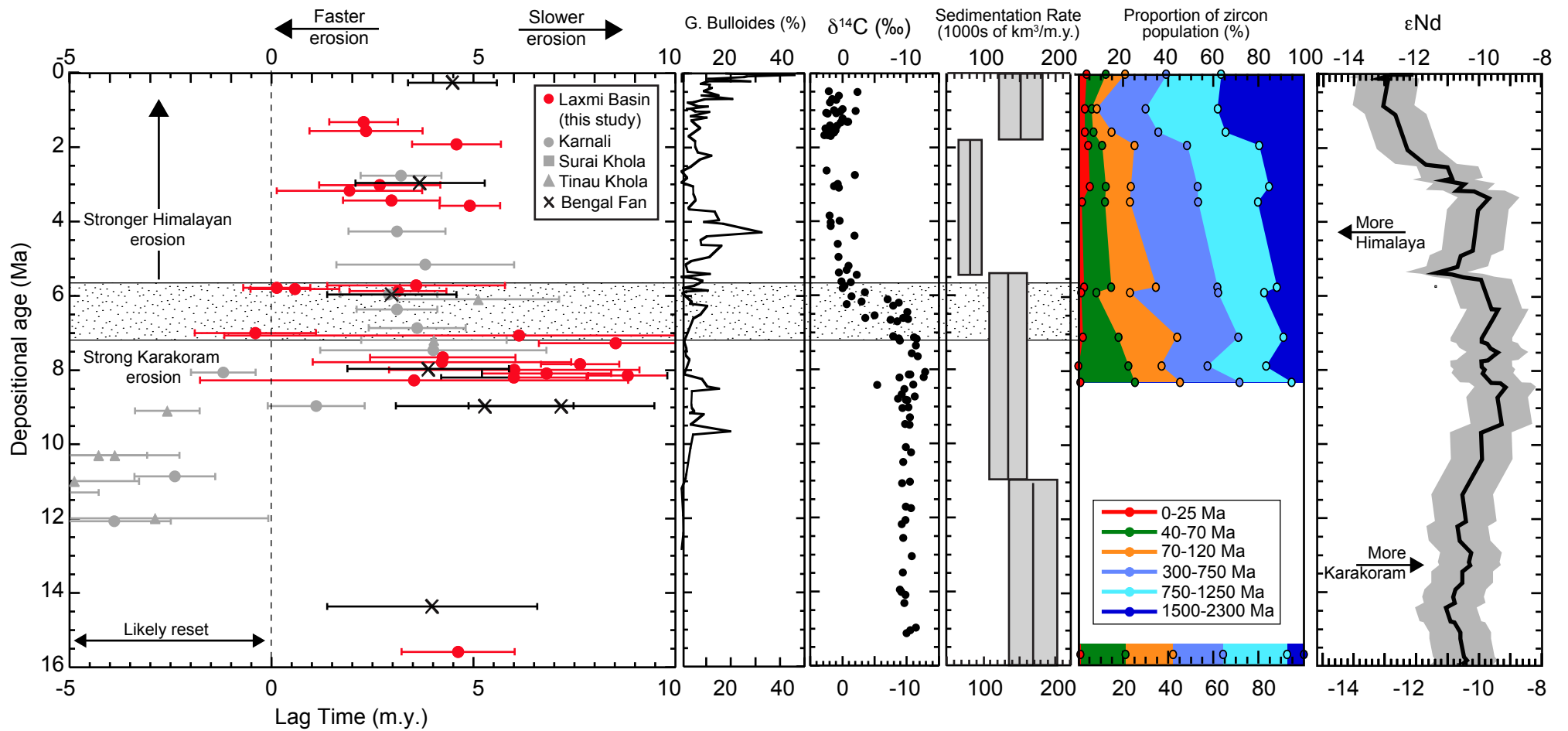


Figure 6
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