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Phase transition in U(1) configuration space: Oscillons as remnants of vortex-antivortex annihilation

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We show that the annihilation of vortex-antivortex pairs can lead to very long-lived oscillon states in 2d Abelian Higgs models. The emergence of oscillons is controlled by the ratio of scalar and vector field masses, $\beta = (m_s/m_v)^2$ and can be described as a phase transition in field configuration space with critical value $\beta_c \approx 0.13(6) \pm 2$: only models with $\beta < \beta_c$ lead to oscillonlike remnants. The critical behavior of the system obeys a power law $O(\beta) \sim |\beta - \beta_c|^o$, where O is an order parameter indicating the presence of oscillons and $o = 0.2(2) \pm 2$ is the critical exponent.

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I. INTRODUCTION

In relativistic and nonrelativistic field theories, models that exhibit spontaneous symmetry breaking support a plethora of topologically stable field configurations or defects [1], static solutions of the nonlinear equations of motion with properties determined by the topology of the vacuum manifold. One of the simplest examples is the kink in 1d $\lambda\phi^4$ models, the result of a discrete symmetry breaking: the vacuum has two disconnected points and the kink interpolates between them. Once formed, kinks cannot be destroyed. Unless, of course, they meet an antikink. Low-momentum kink-antikink scattering leads to the formation of breathers, long-lived time-dependent bound states characterized by large-amplitude oscillations of the scalar field [2]. At larger velocities, breathers have been shown to form only at certain windows or resonances [3]. Although $\lambda\phi^4$ breathers have not been seen to decay in numerical studies [3], they may decay via nonperturbative effects [4], radiating their energy to spatial infinity.

In 2d, the simplest stable topological defect with localized energy is the Nielsen-Olesen vortex found in Abelian Higgs models (AHMs) [1]. These vortices are of great interest in many areas of physics, e.g., as prototypes for studying nonperturbative effects in the standard model and its extensions, or, in the nonrelativistic limit, as Abrikosov vortices in the Landau-Ginzburg theory of superconductivity [5]. Vortex-like solutions may also have a variety of cosmological roles, being formed during a phase transition at the grand unified theory (GUT) scale or thereafter [1,6] or during reheating after inflation [7].

Inspired by the existence of long-lived breathers in 1d $\lambda\phi^4$ models, in this work we investigate the annihilation of bound configurations of vortices and antivortices in AHMs. Vortices and antivortices carry equal and opposite topological charge. This allows for their annihilation. The only adjustable parameter in the model is $\beta \equiv \lambda/2g^2$, the ratio

of squared masses of the scalar and vector fields. We will show that for β smaller than a critical value β_c , vortex-antivortex (henceforth vav) annihilation does give rise to higher-dimensional breatherlike configurations known as oscillons [8], also characterized by long-lived oscillations of the Higgs field magnitude. They were found in 2d [9], 3d [10], and higher-dimensional real scalar field models [11]. Their properties have been studied in great detail [12]. Long-lived oscillatory solutions have also been found in the decay of sphalerons in the 1d AHM [13]. Recently, oscillons were found in 3d gauged SU(2) models [14] and gauged SU(2) \times U(1) models [15] when the Higgs mass is twice the W^\pm boson mass.

In the above references, the procedure was to solve the relevant equations of motion assuming or not spherical symmetry and using an initial profile that approximates the oscillon solution. So long as certain conditions are satisfied [11], the system relaxes into oscillons as it evolves in time. Here, we show that oscillons may also emerge dynamically, as remnants of vav annihilation. (Previous attempts to find oscillons in vortex-vortex scattering were unsuccessful [16].) Our finding complements recent results where oscillons were shown to form after a rapid quench in real scalar field models with symmetric [17] and asymmetric [18] potentials. The fact that oscillons may emerge dynamically raises their stakes considerably: they can play a crucial role during spontaneous symmetry breaking in a variety of physical systems and situations, from the early universe to high energy collisions, and possibly in superfluids and superconductors.

II. THE MODEL

The Abelian Higgs Lagrangian density is

$$\mathcal{L} = \mathcal{D}_\mu \phi^\dagger \mathcal{D}^\mu \phi - \frac{1}{4} F^{\mu\nu} F_{\mu\nu} - \frac{\lambda}{4} (\phi^\dagger \phi - \eta^2)^2, \quad (1)$$

where $\mathcal{D}_\mu = \partial_\mu - igA_\mu$. The scalar and vector masses are, respectively, $m_s = \sqrt{\lambda}\eta$ and $m_v = \sqrt{2}g\eta$. Their ratio defines the parameter $\beta = (m_s/m_v)^2 = \lambda/2g^2$. In the nonrelativistic limit, $\beta = 1$ defines the boundary between

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Type I ($\beta < 1$) and Type II ($\beta > 1$) superconductors. With the scaling $A_\mu \rightarrow \eta^{-1} A_\mu$, $\phi \rightarrow \eta^{-1} \phi$ and $x \rightarrow \eta g x$, the energy is $E = \eta^2 \int_M d^2x \mathcal{H}$, where the Hamiltonian density in the temporal gauge $A_0 = 0$ is

$$\mathcal{H} = \frac{1}{2}(E^2 + B^2) + \partial_t \phi^\dagger \partial_t \phi + \mathcal{D}_i \phi^\dagger \cdot \mathcal{D}_i \phi + \frac{\beta}{2}(\phi^\dagger \phi - 1)^2. \quad (2)$$

From Eq. (1), one derives the equations of motion

$$\mathcal{D}_\mu \mathcal{D}^\mu \phi = -\frac{\lambda}{2} \phi (|\phi|^2 - \eta^2); \quad (3)$$

$$\partial_\mu F^{\mu\nu} = j^\nu = 2g \text{Im}\{\phi^\dagger \mathcal{D}^\nu \phi\}.$$

Nielsen-Olesen vortices are solutions satisfying $\phi(\mathbf{r}) = e^{in\theta} f(r)$ and $A(\mathbf{r})_\theta = -(n/gr)\alpha(r)$, with boundary conditions at spatial infinity [$r = (x^2 + y^2)^{1/2} \rightarrow \infty$] $f(r), \alpha(r) \rightarrow 1$. n is the winding number. At the vortex center, $f(0) = \alpha(0) = 0$. Vortices have a quantized magnetic flux $\Phi_B = 2\pi n/g$. For a vav pair, the net flux is zero. Quantities are measured in units of $\eta = 1$.

III. LATTICE IMPLEMENTATION

We work on the temporal gauge $A_0 = 0$ and follow the Hamiltonian implementation for lattice gauge theories described in Refs. [16,19]. The discrete Hamiltonian is

$$H = \sum_x \mathcal{H}(\pi_x^\alpha, \phi_x^\alpha) \delta x^2, \quad (4)$$

where the density is

$$\mathcal{H} = \pi_x^\dagger \pi_x + \vec{\mathcal{D}} \phi_x^\dagger \cdot \vec{\mathcal{D}} \phi_x + \frac{1}{2} \vec{E}_x \cdot \vec{E}_x + \frac{1}{2\delta x^2} \sum_{i \neq j} (A_{x+i}^i - A_x^i - A_{x+i}^j + A_x^j)^2 + V(\phi_x^\dagger \phi_x), \quad (5)$$

and the lattice covariant derivative is $\mathcal{D}_\mu \phi = (e^{-ig\delta x A_x^\mu} \phi_{x+\mu} - \phi_x) \delta x^{-1}$. The Hamiltonian density is invariant under that lattice gauge transformations

$$\begin{aligned} \phi_x &\rightarrow e^{i\chi_x} \phi_x; & \pi_x &\rightarrow e^{i\chi_x} \pi_x; \\ A_x^\mu &\rightarrow A_x^\mu + (\chi_{x+\mu} - \chi_x)/(g\delta x); & E_x^\mu &\rightarrow E_x^\mu. \end{aligned} \quad (6)$$

The equations of motion are then

$$\begin{aligned} \partial_t \pi_x^\alpha &= -\partial_{\phi_x} \sum_x \mathcal{H}(\pi_x^\alpha, \phi_x^\alpha); \\ \partial_t \phi_x^\alpha &= \partial_{\pi_x} \sum_x \mathcal{H}(\pi_x^\alpha, \phi_x^\alpha) \end{aligned} \quad (7)$$

for each of the fields $\pi^\alpha = \{\pi, E_j\}$ and $\phi^\alpha = \{\phi, A_j\}$. The temporal evolution follows a second-order leapfrog scheme with $\delta t < \delta x/\sqrt{2}$. Provided that the initial conditions satisfy the lattice Gauss's law constraint,

$$\frac{1}{\delta x} \sum_i (E_{x-i}^i - E_x^i) = i(\pi_x \phi_x^\dagger - \pi_x^\dagger \phi_x), \quad (8)$$

it will remain satisfied during the evolution. This can be implemented by, e.g., setting all temporal derivatives to zero as the simulation starts. Energy is conserved to $O(\delta t^2)$. We used $\delta t = 0.02$ and $\delta x = 0.2$ and checked that the results are consistent for $\delta x = \{0.1 \leftrightarrow 0.3\}$. Larger values compromise the accuracy of the results.

Since we use periodic boundary conditions, to prevent energy from the initial condition to come back to the configuration we use a lattice much larger than the region occupied by the vav pair. However, as oscillons are very long-lived, we also use an adiabatic absorbing wall [9] implemented by adding gauge invariant $\gamma(x, y)\pi^\alpha$ terms to the evolution equations, where $\gamma(x, y)$ has support only on a band near the lattice boundary and slowly increases from $0 \rightarrow 0.1$.

IV. VORTEX-ANTIVORTEX ANNIHILATION: RESULTS

We set the initial conditions by using an *ansatz* that approximates a vav pair separated by an initial distance δ_0 . This is then allowed to evolve in time under the equations of motion, Eq. (3), at high viscosity [we set $\gamma(x, y) = 1$] until it settles into the vav configuration. At large δ_0 and for $\beta \leq 1$ (the case of interest to us) we computed numerically the total energy of a vortex, finding a good fit for

$$E_v = 2\pi\beta^{1/5}, \quad (9)$$

which agrees near the region $\beta \sim 1$ with the result of [20]. Also, the contribution in gauge fields to the vortex energy, $E_g = (E^2 + B^2)/2$, can be fitted as

$$E_g = \frac{\pi}{\sqrt{6}} \beta^{1/4}. \quad (10)$$

After the initial *ansatz* settles into a vav pair, we let them approach until their centers reach the distance $\delta_i = 4m_s^{-1}$, large enough to avoid any tail overlap [20]. At this point, we set the viscosity to zero and let the pair attract and scatter. We then repeat the experiment for different values of β .

In Fig. 1 we show a sequence of snapshots of an annihilation process for $\beta = 0.08$, where an oscillon is formed. We plot both the amplitude $|\phi|^2$ and the magnetic field B_z . One can see the outgoing wave front of radiation (center) as the vortices coalesce into the oscillon and its oscillating dipole magnetic field (right).

In Fig. 2 we plot the associated energy (\mathcal{H}) and Chern-Simons (n_{CS}) densities, where $n_{CS} = (4\pi)^{-1} \epsilon_{\alpha\beta\gamma} A^\gamma F^{\alpha\beta}$. Note that oscillons have a distinctive signature (localized oscillations of n_{CS} , top right) and thus could have an interesting role in theories with current anomalies of the form $\partial_\mu J_B^\mu \propto F\tilde{F}$ leading to baryogenesis, including dur-

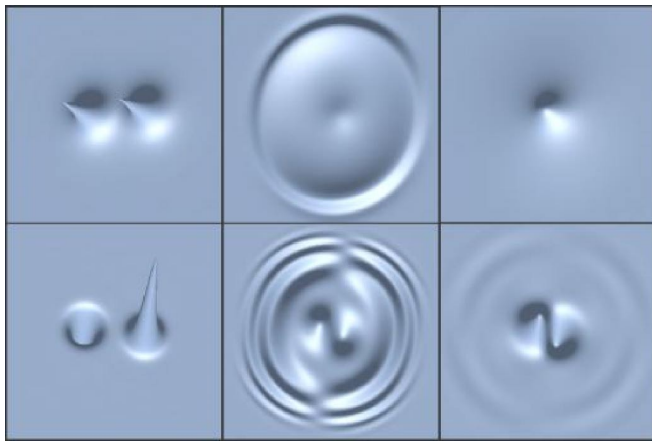


FIG. 1 (color online). Snapshots of $|\phi|^2$ (top) and B_z (bottom) for a configuration with $\beta = 0.08$. Time proceeds from the initial vav pair on the left to the oscillon on the right.

ing postinflationary reheating [21], a possibility we are exploring.

Vortex-antivortex annihilation can lead to two possible outcomes, depending on the value of β : either the vav pair quickly radiates all its energy to spatial infinity, $(+) + (-) \rightarrow 0$, or it settles into an intermediate, long-lived oscillatory state before radiating its energy to spatial infinity, $(+) + (-) \rightarrow (+-) \rightarrow 0$. In order to investigate the process quantitatively, we measured the energy in a disk (E_{disk}) surrounding the vav pair. The results are shown in Figs. 3 and 4. In Fig. 3 we show the energies in a disk of radius $r = 15$ for different values of β . In Fig. 4 we show the total energy and the associated energy in gauge fields for $\beta = 0.052$. The inset shows that there are two characteristic frequencies, approximately set by the scalar and gauge field masses, m_s and m_v . For $\beta = 0.052$, one obtains excellent agreement with $m_v \sim 4.4m_s$.

By inspecting Fig. 3, we identify an oscillon whenever the energy develops a plateau. To read off a value of the

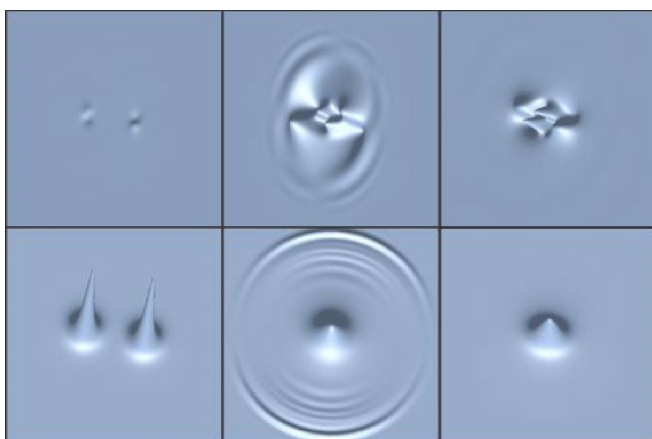


FIG. 2 (color online). Chern-Simons density (top) and energy density (bottom) for the same snapshots of Fig. 1.

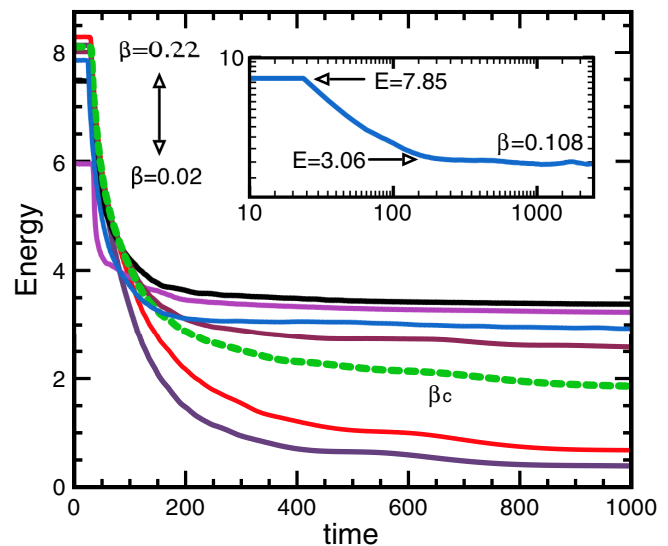


FIG. 3 (color online). Energy within a disk of radius $r = 15$ surrounding the initial vav pair for various values of β on a 100^2 lattice. The range of β is indicated on the top left. There are two regimes: for $\beta \geq \beta_c \approx 0.136$, the pair quickly annihilates, radiating its energy to spatial infinity. For $\beta < \beta_c$, the pair sheds some of its initial energy and settles into a long-lived intermediate state, where almost no energy is radiated. This state is the U(1) Abelian Higgs oscillon. Inset: change in energy (E) radiated as a function of $\ln t$ for $\beta = 0.108$. Note the “hockey stick” shape. The arrow pointing right denotes the oscillon’s energy, E_{osc} .

plateau energy we examine the energy vs $\ln t$ plot. When oscillons are present, the curve assumes a “hockey stick” shape. We read off the oscillon’s energy (E_{osc}) as the value of E_{disk} at the location of the elbow. [Arrow in inset in

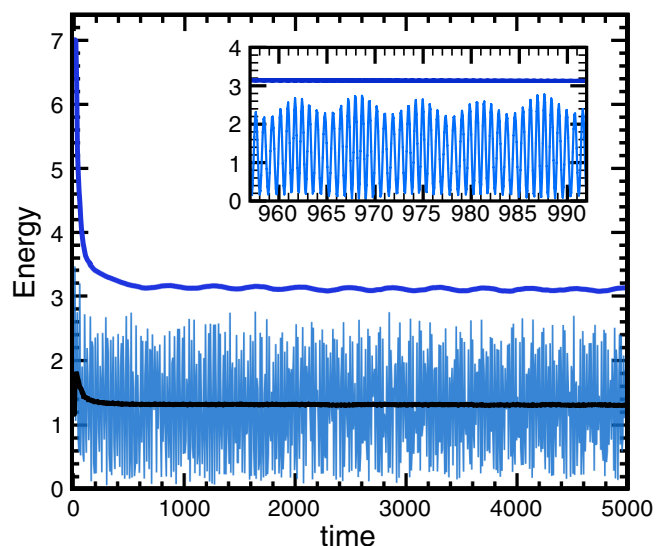


FIG. 4 (color online). Top curve: total energy for $\beta = 0.052$ in a disk of radius $r = 15$ surrounding the initial vav pair. Lower curve: energy in the gauge fields. Black line is the running average of the gauge energy. The inset shows the total energy and the energy in gauge fields for a small time interval.

Fig. 3.] We have observed that oscillon lifetimes are all in excess of $\tau_{\text{osc}} \geq 10^{(4-5)}$. For example, for $\beta = 0.098$, $\tau_{\text{osc}} \geq 5.0 \times 10^5$. Note that with the rescaling defined after Eq. (1), time is in units of $(\eta g)^{-1} = m_v/\sqrt{2}$. So, it is set by the couplings of the model and the symmetry-breaking scale. If this were the standard model, $(\eta g)^{-1}$ would be the inverse mass of the W boson and time would be in units of $(80 \text{ GeV})^{-1} \sim 8.2 \times 10^{-27} \text{ sec}$. We leave a detailed study of their extreme longevity for future work.

We identify $\beta_c = 0.13(6) \pm 2$ as the critical value separating the two regimes, as it exhibits the critical slowing down typical of continuous phase transitions. For $\beta > 0.2$, well outside the critical region, no oscillons form: vav pairs radiate their initial energy to spatial infinity so that by $t \sim 500$, $E_{\text{disk}} \leq 0.4 \ll E_{\text{osc}}$. [We neglect here some small remnant energy $E_{\text{disk}} < 0.1 - 0.2$, that persists long beyond the oscillon's demise.]

The quantity E_{osc}/E_v naturally lends itself as an effective order parameter describing the transition from vav pairs to oscillons in field configuration space. In Fig. 5 we plot E_{osc}/E_v vs β . The curve is well fitted for

$$E_{\text{osc}}/E_v \sim (|\beta - \beta_c|)^{0.2(2) \pm 2}. \quad (11)$$

Near criticality the exponent 0.2 will have small corrections [22]. Given the long integration times needed to extract a more precise value for the energy of the oscillons around β_c (akin to critical slowing down, as can be seen in Fig. 3), we did not attempt to find these corrections here. From the theory of continuous phase transitions, the critical point is related to infrared divergences. In fact, the effective size of the oscillon increases near β_c : disks of larger radii are needed to capture the full energy of the configuration. The associated decrease in the effective scalar field mass near the core [$\sim (A^2 - \beta)^{1/2}$] is the result of screening by the gauge field. Above criticality, the mass at the core becomes positive, eventually matching the asymptotic vacuum as the vav annihilates.

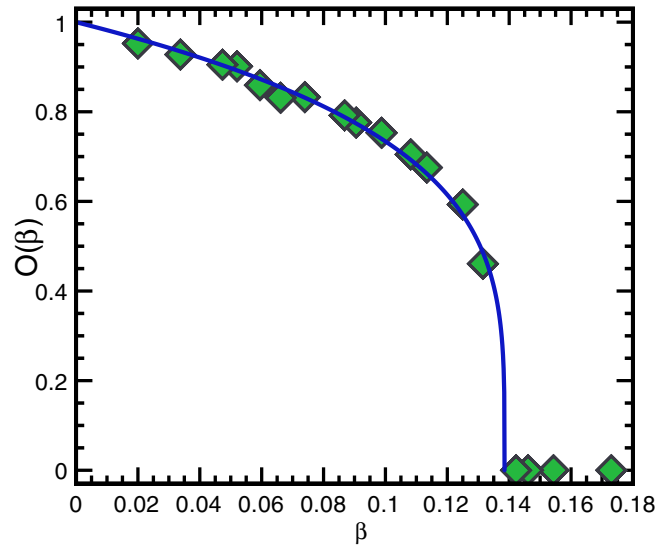


FIG. 5 (color online). Phase diagram for the vav \rightarrow oscillon transition. Plotted is the order parameter $O(\beta) \equiv E_{\text{osc}}/E_v$ as a function of β . The diamonds are numerical results, while the continuous line is the fit of Eq. (11).

Our results show that the annihilation of a vav pair in the 2d AHM can be described as a phase transition in configuration space, with complete annihilation corresponding to the “symmetric” phase $E_{\text{osc}}/E_v \rightarrow 0$, and oscillons corresponding to long-lived “broken-symmetric” states, with the associated order parameter attaining a nonzero value.

It will be interesting to investigate if oscillons can exist in superconductors, how their lifetimes vary with β , and if they can be created in non-Abelian models via, for example, low-momentum monopole-antimonopole scattering.

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