From passive degassing to violent Strombolian eruption: the case of the 2008 eruption of Llaima volcano, Chile. Dawn C.S. Ruth<sup>1,2\*†</sup>, Elizabeth Cottrell<sup>2</sup>, Joaquín A. Cortés<sup>3</sup>, Katherine A. Kelley<sup>4</sup>, Eliza S. Calder<sup>3</sup> \* Corresponding author: sdawn@ntu.edu.sg † Now at Earth Observatory of Singapore, Nanyang Technological University, Singapore, 639798 1. Department of Geology, State University of New York at Buffalo, Buffalo, NY 14260, USA 2. Smithsonian Institution, National Museum of Natural History, Washington, DC, 20560, USA 3. School of Geosciences, University of Edinburgh, Scotland, EH9 3JW 4. Graduate School of Oceanography, University of Rhode Island, Narragansett, RI, 02882, USA 

#### **ABSTRACT**

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On 1 January 2008 Llaima volcano, a basaltic andesite stratocone in southern Chile, entered a phase of violent Strombolian eruption. Llaima, like many passively degassing systems, has experienced prolonged (decades-long) periods of persistent summit degassing from its open vent. The rapid transition from long-lived passive degassing to violent explosive eruption occurred with limited precursory monitoring signals. This paper is motivated by the desire to understand what occurs in these systems when that switch takes place. To this end, we study the products of the 2008 violent Strombolian eruption of Llaima volcano, Chile. We present new textural analyses of scoria and geochemical data from five whole rock samples, troctolite glomerocrysts with and without Cr-spinel, and 182 olivine-hosted melt inclusions from tephra samples. Two populations of scoria ("brown" and "black") are distinguished by variable crystallinity and vesicularity, but are geochemically indistinguishable. Black scoria contain abundant microlites with tabular to acicular morphologies and convolute vesicles up to 1.75 mm in effective diameter. The brown scoria tend to have fewer, acicular microlites, abundant matrix glass, and round vesicles with a narrower size distribution constrained to <0.4 mm in diameter. Overall, textures from the black and brown scoria evidence a textural maturation process where shallow system magma becomes more crystal rich and likely rheologically stiffer as a result of prolonged passive degassing. The Cr-spinel and Cr-spinel free troctolite glomerocrysts have plagioclase and olivine compositions of An<sub>65-92</sub> and Fo<sub>81</sub>, respectively. Cr# in the Cr-spinel range from 26-37; these are consistent with magma originating from the deeper plumbing system. Whole rock compositions for the tephra average 51 wt% SiO<sub>2</sub>, 18 wt% Al<sub>2</sub>O<sub>3</sub>, and ~6 wt% MgO. The major elements of olivine-hosted melt inclusions range from 49 to 56 wt% SiO<sub>2</sub>, and 3.72 to 7.76 wt% MgO; there was no distinct compositional difference between olivine-hosted melt

inclusions sourced from the different scoria. Melt inclusion volatile contents range from below detection up to 2.95 wt%  $H_2O$ , and 1973 ppm  $CO_2$  (though not in the same melt inclusion).  $H_2O$  and  $CO_2$  concentrations are consistent with open system degassing and, when compared with fractionation indices (e.g.  $K_2O$ ), indicate coupled degassing and crystallization throughout the system. The majority of melt inclusions denote a single liquid line of descent indicative of plagioclase and olivine fractionation. Entrapment pressures range from 8-342 MPa and fall into two groups: 8-100 MPa (300 m to ~4 km depth) and > 100 MPa (4-14 km), revealing that this eruption tapped a deep plumbing system.

We propose here that passive degassing at Llaima is maintained by periodic small-batch magma injections. Consequently, due to extensive degassing the upper plumbing system magma crystallized and increased in viscosity. Before the 2008 eruption, some volatiles sourced from the repeatedly-injected magmas, exsolved from the inferred crystal mush, and ascending from deeply-sourced degassing magmas gradually accumulated within the crystal mush and beneath the stiffened conduit magma. Our results support a model where eruption triggering occurred when magma injection remobilized the mush and, importantly, unlocked the accumulated gasses that ascended rapidly, and generated the observed violent Strombolian explosive activity. Our proposed model contrasts with those models for explosive mafic volcanism which require rapid magma ascent under closed system degassing conditions. Importantly, our proposed mechanism provides a means for systems with dominantly open system degassing behavior to transition from passive degassing to explosive eruptions.

- **Keywords:** Basaltic andesite; eruption trigger; magma recharge; melt inclusions; open vent
- volcanism; passive degassing; regulated gas transport

# INTRODUCTION

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On the 1st of January 2008, Llaima volcano in the Southern Andean Volcanic Zone, Chile, entered a new phase of intense explosive activity. The eruption, described as a violent Strombolian type eruption, was characterized by highly pulsatory fire fountaining and the development of a 10 km high ash plume (Ruth and Calder, 2014). Prior to the eruption the volcano exhibited persistent passive degassing at least since its previous eruption in 1994 (OVDAS, 2008). Passive degassing (sometimes called persistent degassing, or quiescent degassing) is characterized by volatile outgassing, usually from a summit vent, but is commonly associated with limited or negligible magma net flux out of the system. In many cases volcanoes exhibiting this behavior are typified by mafic magmas and open vents, and the persistent degassing can continue for months to years (Rosi et al., 2000; Rose et al., 2013). In contrast, violent Strombolian eruptions (often referred to as paroxysms, or cineritic eruptions) are characterized by moderately-sized, sustained ash plumes (> 10 km in altitude) and dispersed tephra deposits (Francis et al 1990; Valentine and Gregg, 2008; Houghton and Gonnermann, 2008). Violent Strombolian eruptions are known to occur periodically at volcanoes that commonly present passive degassing, including, Stromboli, Italy (Barberi et al., 1993; Rosi et al., 2000; Rosi et al., 2013); Llaima and Villarrica, in Chile (Petit-Breuihl, 2006; OVDAS, 2008), Masaya, Nicaragua (Williams, 1983); and Fuego, Guatemala (Rose et al., 2008). Where these violent explosive eruptions occur from open, already passively degassing systems, their onset can be sudden, sometimes with little to no advance warning or signals in monitoring data. The hazards associated with such events (ballistics, tephra fall, lava flows, and sometimes lahars) can be challenging to ameliorate due to the sudden onset nature, especially if the volcano is located in a region of high population (e.g. Rose et al., 2008; Rose et al., 2013; Rosi et al., 2013).

Understanding the mechanisms that initiate this rapid change in behavior, from passive degassing to violent explosive activity, is therefore a key challenge to improving our capacity to deal with explosive basaltic volcanism.

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Investigations on this style of explosive basaltic volcanism have been undertaken at Stromboli, Mt. Etna, Parícutin, and Cerro Negro (Allard, 1997; Roggensack et al., 1997; Métrich et al., 2001; Lautze and Houghton, 2007; Pioli et al., 2008; Johnson et al., 2010; Métrich et al., 2010). Violent Strombolian eruptions necessitate not only high mass fluxes of magma (e.g., mass eruption rates 10<sup>3</sup>-10<sup>5</sup> kg s<sup>-1</sup>), but also, critically, high instantaneous gas fluxes (Pioli et al., 2009; 2009). In several eruption cases, the triggering of these events have been attributed to bouts of recharge of mafic magma that ascend rapidly through the system, e.g. Stromboli (Métrich et al., 2001; Bertagnini et al., 2003; Métrich et al., 2010), Parícutin (Lurh, 2001; Pioli et al., 2008), and Cerro Negro (Roggensack et al., 1997). Deeply-sourced magma injection into a relatively shallow reservoir is inferred from the volatile contents of mineral-hosted melt inclusions. The inferred rapid closed system ascent and subsequent degassing of H<sub>2</sub>O in these magmas then drive the sporadic explosive eruptions, often producing texturally bimodal scoria populations with chemically distinct olivine hosted melt inclusions (e.g., Métrich et al., 2010). However this model does not work for systems where erupted products record dominantly open system degassing behavior (e.g. Anak Krakatau, Mandeville et al., 1998), or systems that are not well characterized by either closed or open system degassing (e.g. Johnson et al., 2010; Oppenheimer et al., 2011). Alternative mechanisms that have been proposed to explain the transition from passive degassing or low intensity eruptions to violent explosive eruptions include changes in magma system geometry (Palma et al., 2011; Genereau et al., 2010) and changes in the fluid dynamics during magma ascent (Jaupart, 1998; Houghton and Gonnermann, 2008).

Studies of the products of the 2008 eruption of Llaima have been undertaken by Bouvet de Maisonneuve et al. (2012a, 2012b) and Ruth and Calder (2014). Ruth and Calder (2014) report on morphological, textural and dispersal characteristics of the diverse tephra types, including distinctive plate tephra. Bouvet de Maisonneuve et al. (2012a) present olivine hosted melt inclusion data for the 2008 scoria, as well as for three historical eruptions in 1957, and ~1850 AD (separated into upper and lower tephras). Their melt inclusion dataset evidences a persistent shallow plumbing system (< ~130 MPa) through this time, but provides limited evidence for consistent degassing and fractionation trends. For the 2008 eruption, mineral zoning and chemistry data suggest repeated pre-eruption magma injection (Bouvet de Maisonneuve et al., 2012b). Based on these findings, they propose that the 2008 eruption was triggered by remobilization of a crystal mush (i.e. magmatic regions with >50% crystals) by magma recharge. The Bouvet de Maisonneuve et al. (2012a, 2012b) work provides an excellent framework for understanding reservoir processes associated with several of the recent major eruptive episodes at Llaima, but key aspects of the 2008 eruption remain unaccounted for.

In this work we focus specifically on the products of the 2008 eruption. We extend the work undertaken by both Bouvet de Maisonneuve et al., (2012a, 2012b) and Ruth and Calder (2014) and use a combined textural and geochemical analysis that both complements previous work, and provides new information about the shallow and deeper system. Specifically, we build on the previous work in the follow ways: 1) Ruth and Calder (2014) identified texturally heterogeneous tephra produced during the violent explosive phase of the 2008 eruption. Here we use textural data, including vesicle and crystal size distributions, from these different tephra to investigate crystallization and degassing processes in the shallow system and discuss these in the context of the explosive eruptive activity. This information is important because shallow

degassing and crystallization are considered critical processes that can lead to increases in the intensity of mafic eruptions (e.g. Sable et al., 2006; Houghton and Gonnermann, 2008). 2) From the different tephra types we provide major, trace, and volatile elements from 182 olivine-hosted melt inclusions, thereby extending the work of Bouvet de Maisonneuve et al. (2012a, 2012b). Our comprehensive geochemical data show evidence of a deeper, more CO<sub>2</sub>-rich plumbing system than previously reported. 3) We consider it important that any conceptual model for the Llaima plumbing system places the 2008 eruptive episode in the context of the persistently active passive degassing system, which typified the activity both before that eruption occurred and afterwards. Although Bouvet de Maisonneuve et al. (2012a, 2012b) focused dominantly on the nature of the magmatic plumbing and crystal mush processes, the role of gas accumulation was not investigated. Here, bearing in mind that high gas fluxes resulting in highly pulsatory lava fountaining is a key driver of violent Strombolian eruptions (e.g. Pioli et al., 2009), we discuss potential scenarios for gas accumulation within the Llaima plumbing system. Finally we propose a model for the transition in activity style from passive degassing to violent explosive eruptions.

# LLAIMA VOLCANO AND THE 1 JANUARY 2008 ERUPTION

Llaima volcano (38°41′45″S 71°43′54″W, 3179 m a.s.l) is a basaltic andesite stratocone located in the Southern Volcanic Zone (SVZ) of the Chilean Andes about ~600 km south of Santiago (Fig. 1). Llaima is one of the most historically active volcanoes in the SVZ with 54 eruptions since 1640 (Naranjo and Moreno, 2005; Petit-Breuihl, 2006). Statistical analysis of the historic record for eruptions of VEI ≥ 2 indicates an eruption recurrence interval of ~6 years (Dzierma and Wehrmann, 2010). Major explosive activity producing significant eruption plumes (>10 km) and associated tephra deposits occurred most recently in 1957 (Hantke, 1962; Naranjo

and Moreno, 2005), with a similar, but lower intensity eruption occurring in May 1994 (Moreno and Fuentealba, 1994). The 1994 eruption lasted 8 hours, produced lava fountains, a 7 km altitude plume and associated tephra deposit, and lahars down the western flanks; the eruption was characterized at VEI 2 (Moreno and Fuentealba, 1994).

On 1 January 2008, between 1730 and 1820 local time (GMT - 04:00), a summit eruption began at Llaima. The eruption was characterized by violent pulsating Strombolian and fire fountaining activity which produced a plume reaching ~11 km high. Lava flows were generated on the west flank, and lahars were generated in the Rio Calbuco and Rio Captrén valleys, respectively (Global Volcanism Program, 2008) and resulted in the evacuation of the Conguillío National Park. The main explosive phase lasted ~13 hours, but eruptive activity continued with reduced but intermittent Strombolian activity continuing through July 2008. A second explosive phase occurred on 3 April 2009; and the extended eruptive phase ended by July 2009. In all, the 2008-2009 eruptive phase produced lava flows extending 3 km from the vent onto the western and southern flanks, a tephra deposit with a minimum eruption volume of 1.3 x 10<sup>6</sup> kg m<sup>-3</sup> (Ruth and Calder, 2014), pyroclastic flow deposits over snow (i.e. mixed avalanches) associated with an observed vent on the (north) eastern flank, and lahars formed from the interaction of lava and summit glacier. Retrospectively, minor increases in seismic activity and minor ash explosions were recognized in May and June 2007 but otherwise there was little/no precursory activity (OVDAS, pers. comm.). There is also no evidence that significant edifice deformation occurred prior to January 2008 (Fournier et al., 2010; Remy et al., 2015; Delgado et al., submitted).

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## SAMPLE COLLECTION AND PREPARATION

The tephra samples studied here are from the initial phase of the 2008 eruption, collected in January 2009 and January 2011 from a total of 94 different sites in the fall deposit on the eastern flanks of the volcano (Ruth and Calder, 2014). The tephra samples were characterized in the field based on hand sample color, texture, and morphology as either black scoria, brown scoria, or plate tephra. The black scoria and plate tephra only differ in shape, and can otherwise be considered similar (Ruth and Calder, 2014). From these, twenty thin sections of representative tephra types were made for textural analysis using conventional petrography and back-scattered electron (BSE) imaging. BSE images were collected using the Hitachi SU-70 scanning electron microscope at the University at Buffalo South Campus Instrument Center (SCIC) and the FEI Nova NanoSEM 600 in the Department of Mineral Sciences at the National Museum of Natural History (NMNH), Smithsonian Institution.

Tephra were separated by type (black and brown scoria, and plate tephra) and were lightly crushed using a jaw crusher and shaker sieved with screen sizes of 25, 40, and 80 mesh (e.g., openings of 710, 425, and 180  $\mu$ m, respectively). Whole rock analyses were conducted on the > 25 mesh separate, as well as gravel-sized hand-picked samples where homogeneity could be ensured. Olivine phenocrysts were hand-picked from the 25 and 40 mesh separates of the crushed tephra.

Clear, naturally quenched, glassy melt inclusions hosted in olivine crystals were selected for analysis (see supplemental material for database of melt inclusions) (Fig. 2). Melt inclusions containing daughter crystals were excluded. Some melt inclusions showed visible fractures but the fractures did not extend to the crystal surface. Melt inclusions contained zero to one vapor bubble. We include in the supplemental data the approximate location in the host olivine of the

melt inclusions. Acceptable samples were mounted in epoxy in single-grain mounts to create wafered samples with the inclusions double-side exposed and polished. A total of 182 melt inclusions were analyzed.

## ANALYTICAL METHODS

## X-ray Fluorescence

Whole rock major elements and some trace elements (Rb, Sr, Y, Zr, V, Ni, Cr, Nb, Ga, Cu, Zn, Co, Ba, La, Ce, U, Th, Sc, Pb) for five tephra samples were analyzed with a PANanalytical 2404 X-ray fluorescence spectrometer (XRF) hosted at Franklin and Marshall College, Pennsylvania. Major element analyses were conducted on dried powder mixed with lithium tetraborate. The mixture was melted in a platinum crucible and then quenched into a flat, glass disk. Total iron as  $Fe_2O_3$  was reported from the XRF analysis and the speciation of iron was measured using titration on a separate aliquot from the crushed samples (after Reichen and Fahey, 1962). Loss on ignition was determined after samples were heated for one hour at 950 °C. The whole-rock powder was mixed with a high purity Copolywax powder and then pressed into a briquette. Trace elements analyses were conducted on this briquette. Calibration curves were constructed with standards listed in Abbey (1983) and Govindaraju (1994). One sample was run in triplicate to determine reproducibility. For  $SiO_2$ ,  $Al_2O_3$ , CaO,  $Na_2O$  the replicate analyses showed a < 1% difference;  $Fe_2O_3$ ,  $TiO_2$ , MgO, and  $K_2O$  had a 1-2% difference. Replicate  $P_2O_5$  and MnO analyses were identical.

## **Electron Microprobe**

Major element analyses of carbon-coated melt inclusions, matrix glass, plagioclase, and olivine were collected using the five-spectrometer JEOL 8900 electron microprobe (EMPA) located at The National Museum of Natural History, Smithsonian Institution. Select plagioclase and matrix glass analyses were obtained with the JEOL 8900 electron microprobe at Cornell University. Spinel inclusions in the glomerocrysts were analyzed using the five-spectrometer JEOL JXA-8430 field emission electron microprobe located at the Facilities for Analysis, Classification and Testing (FACTS) at Nanyang Technological University, Singapore. Glass analyses were collected on at least five spots per melt inclusion with a 10 µm beam diameter, 15 keV accelerating voltage, and 10 nA beam current (see supplemental materials for details). To minimize alkali loss, we analyze Na and K first, and for shorter on peak durations (10 sec). Our chosen wide beam diameter also minimizes alkali loss (e.g. Hayward, 2011). These techniques are reasonable since analysis of standards as unknowns (e.g. VG-2) returns Na<sub>2</sub>O totals within 3% of reported values. Sulfur and chlorine were run after the major element analysis on a minimum of three new spots with 15 keV, 80 nA beam current, and 10 µm beam diameter. Sulfur peaks were initially obtained via a peak search routine on individual melt inclusions. Since the peak position did not shift more than 2%, the peak position was determined using the search routine on scapolite during standardization. The closest (~5-10 µm from the melt inclusion) olivine probe point to the melt inclusion was considered the host olivine composition. Olivine analyses were collected with 15 keV, 40 nA beam current, and 1 µm beam diameter. Plagioclase (in thin section and the glomerocrysts) analyses were obtained with an accelerating voltage of 15 keV, 15 nA beam current, and 5 µm beam diameter. Beam conditions for the spinel analyses were: 15 keV, 20 nA, and 1 µm beam diameter. For all analyses, backgrounds on each side of the peak were counted each for half the peak count times. Beam drift was assessed by

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running standards (glass: VG-2; plagioclase: Bytownite; olivine: San Carlos olivine) as an unknown throughout the analytical run and calculating the percent deviation from reported values.

# **Fourier Transform Infrared Spectroscopy**

H<sub>2</sub>O and CO<sub>2</sub> concentrations were determined by transmission Fourier Transform Infrared spectroscopy (FTIR) on 49 of the 182 melt inclusions using the Bio-Rad Excalibur spectrometer or the Nicolet 6700 Analytical FTIR hosted at The National Museum of Natural History, Smithsonian Institution, following the methods of Luhr (2001). To minimize atmospheric interference on the spectra, the system (bench, microscope, and samples) was continually purged with dry, CO<sub>2</sub>-free air. The spectral region of interest, 1000-6000 cm<sup>-1</sup>, was obtained with a liquid nitrogen-cooled MCT detector, KBr beam splitter, and tungsten halogen source. Aperture size was selected to ideally obtain three independent spectra per melt inclusion; they ranged from 11x22 μm to 44x44 μm. For smaller melt inclusions, the spectra measurements overlapped. Thickness data were collected either with a piezometric micrometer, and/or indirectly with the wavelength of fringes from 2000-2700 cm<sup>-1</sup> (Nichols and Wysoczanski, 2007).

Water concentrations in the melt inclusions were calculated using the total  $H_2O$  peak at 3535 cm<sup>-1</sup> (Stolper, 1982a, 1982b). Dissolved  $CO_3^{2-}$  concentrations were quantified using the asymmetric stretch doublet between 1515 cm<sup>-1</sup> and 1435 cm<sup>-1</sup> (Dixon et al., 1995). The background-corrected peak intensities were obtained through an Excel fitting routine which builds an optimized, synthetic signal that accounts for the vertical offset, the 1630 cm<sup>-1</sup> total water peak, the  $CO_3^{2-}$  doublet, and a straight line (Newman et al., 2000).

When fringe amplitude prevented a straightforward assessment of the intensity of the carbonate doublet, we calculated the CO<sub>2</sub> content of the fringe-filtered signal. The amplitude and wavelength of the fringes were modeled for each spectrum between 1800 and 2000 cm<sup>-1</sup> using the curve fitting module in MATLAB (see supplemental material for the MATLAB script). The user imports the FTIR absorbance and wavenumber data via a Microsoft Excel file and selects a region where the script will determine the fringes signal. The fringes signal between 1800 and 2000 cm<sup>-1</sup> can be modelled simply as a linear function overprinted by a sine function. Note that this region may be overprinted by the Si-O signal from analyzed olivine (Nichols and Wysoczanski, 2007); if overprinting is present then this technique is not suitable to correct for fringes. The script fits a sine curve to the fringes, subtracts this from the original spectrum, and outputs the results into a comma-separated text file. CO<sub>2</sub> values were then re-calculated for the fringe-subtracted signal using the Excel fitting routine described above. The difference between the fringe-filtered and unfiltered calculation is less than 5% (Fig. 3), indicating that the presence of fringes does not significantly affect the quantification of carbonate in these samples. In some cases, the CO<sub>3</sub><sup>2-</sup> doublet was completely overwhelmed by the fringes and CO<sub>2</sub> values could not be reliably retrieved. The  $CO_2$  detection limit for these melt inclusions (up to 80 µm thick) was ~ 400 ppm. For these samples, we only report H<sub>2</sub>O values. In total, we performed this correction on spectra from 12 melt inclusions (see supplemental material for raw and fringe-corrected data). Water was quantified using the spline function built-in in the Thermo Fischer OMNIC spectral analysis software package. Peak heights were collected on the background corrected

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spectral analysis software package. Peak heights were collected on the background corrected spectra. Molar absorptivities at 1515 and 1435 cm<sup>-1</sup> for C and 1635, 4500, and 5200 cm<sup>-1</sup> for H species were corrected for major element composition according to Dixon et al. (1995). A

constant value for molar absorptivity of 63 mol<sup>-1</sup> cm<sup>-1</sup> was applied to all 3535 cm<sup>-1</sup> peaks as it is not sensitive to composition (Dixon et al., 1995; Kelley and Cottrell, 2012).

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#### **Secondary Ion Mass Spectrometry**

Secondary ion mass spectrometry (SIMS) was performed on 39 additional melt inclusions because the detection limit for CO<sub>2</sub> by SIMS is lower (~20 ppm) relative to FTIR (~50 ppm) (Hervig et al., 2003). Additional volatile analyses (H<sub>2</sub>O, CO<sub>2</sub>, S, Cl, F) were collected on a subset of olivine-hosted melt inclusions at the Northeast National Ion Microprobe Facility, Woods Hole Oceanographic Institution, using the Cameca IMS 1280 secondary ion mass spectrometer following the analytical procedure of Helo et al. (2011). These samples were polished to 0.5 micron grit (single-side only), mounted in indium and gold coated samples were then placed under vacuum (<5x10<sup>-9</sup> torr) and bombarded with a 15 µm, 1.2-1.5 nA, 10kV Cs<sup>+</sup> ion beam to produce the secondary ions (12C, 16O1H, 19F, 30Si, 32S, and 35Cl) with 30Si as the reference mass. The samples were pre-sputtered for 4 minutes and the primary beam rastered over an area of 30 x 30 µm. Analysis continued after pre-sputtering with an aperture reduced to 15 x 15 µm. Count times for <sup>16</sup>O<sup>1</sup>H and <sup>12</sup>C were 10 s for 10 cycles for a total of 100 s. <sup>19</sup>F, <sup>30</sup>Si, <sup>32</sup>S, and <sup>35</sup>Cl count times were 5 s for 10 cycles for a total of 50 s. Calibration curves for measured  ${}^{12}\text{C}/{}^{30}\text{Si}$ .  ${}^{16}\text{O}^{1}\text{H}/{}^{30}\text{Si}$ .  ${}^{19}\text{F}/{}^{30}\text{Si}$ .  ${}^{32}\text{S}/{}^{30}\text{Si}$ .  ${}^{35}\text{Cl}/{}^{30}\text{Si}$  versus the respective volatile concentration were built from the analysis of 11 standard basaltic to andesite glasses (49-57 wt% SiO<sub>2</sub>) (Helo et al., 2011) using a York regression (see supplement for calibration curves). R<sup>2</sup> values were 0.99 for all volatiles measured; the reported error includes the measurement error and the error in the linear fit (see supplemental material).

# **Laser Ablation Inductively Coupled Plasma Mass Spectroscopy**

129 melt inclusions were analyzed for thirty-nine trace elements using laser ablation inductively coupled plasma mass spectroscopy (LA-ICPMS). These analyses were done using the Thermo X-Series II quadrupole ICP-MS coupled with a New Wave UP 213 Nd-YAG laser ablation system located at the Graduate School of Oceanography, University of Rhode Island. Spot sizes ranged from 20 − 80 µm depending on melt inclusion size. Ablation time was maximized for the wafered samples with an energy output of 60% and 5 Hz repeat rate. Background and washout were collected for 30 seconds, and the maximum analytical time was 60 seconds. Our methodology was modified from Kelley et al. (2003) and Lytle et al. (2012). We used <sup>43</sup>Ca as an internal standard and calibrated on eight natural glasses from the USGS (BIR-1g, BHVO-2g, BCR-2g) and Max Planck Institute (GOR132-G, StHls-G, T1-G, ML3B-G, KL2-G); reference values are as reported in Kelley et al. (2003), Jochum et al. (2006), and Lytle et al. (2012). Acceptable linear calibration curves had R² > 0.995 (see supplemental material for more information). For Ag, Pt, and W, a lower R² value is reported due to limited studies on the standards used. Lab reproducibility is ≤ 5% for most trace elements.

# X-ray Absorption Near Edge Structure Spectroscopy - (micro-XANES)

Two melt inclusions (blk3-2011 and blk1-2010) were analyzed *in situ* for Fe<sup>3+</sup>/ $\sum$ Fe ratios using micro X-ray Absorption Near Edge Structure (micro-XANES) spectroscopy with precision of  $\pm 0.005$  (Cottrell et al., 2009). Following the methods of Cottrell et al. (2009), the two melt inclusions were analyzed on beamline X26a located at the National Synchrotron Light Source, Brookhaven National Lab. Spectra that included olivine interference were excluded (e.g., Kelley and Cottrell, 2009; Brounce et al., 2014).

# TEXTURAL ANALYSIS METHODS AND DEFINITIONS

# **Definitions**

Textural analysis of crystal assemblages provides information about the kinetics of magmatic processes, especially crystallization and degassing (Hammer, 2008). We define phenocrysts here as crystals that are > 100 μm and microphenocrysts as crystals that range from 30 to 100 μm. Microlites are crystals under 30 μm. Vesicles are described as rounded or convolute (i.e. irregular perimeters). Qualitative crystal shape terms used in sample characterization include acicular, tabular, swallow-tail and hopper morphologies (e.g. Lofgren, 1974; Couch et al., 2003; Crabtree and Lange, 2011) (Fig. 4). Acicular crystals are characterized by needle shape habits whereas tabular crystals are rectangular. Swallow-tail plagioclase crystals have forked terminations, often with acicular habits. Hopper (or skeletal) morphology plagioclase crystals are those with voids often in the center (Fig. 4).

#### **Vesicle Size Distributions**

We measured the vesicle size distributions (VSDs) to quantify the textural differences between the scoria (e.g. Polacci et al., 2006; Moitra et al., 2013). BSE images at x70 and x90 magnification were made binary for vesicles or crystals. Vesicle walls were manually drawn in the images in the case of breakage during sample preparation. First, we determined the phenocryst areas by manually tracing each crystal in an image; these values were subtracted from the total area measured. Then, where necessary, we reconstructed all bubble walls if portions of the septum were visible to quantify the maximum number of vesicles per image. More than 1000 vesicles were measured per thin section for two black and brown scoria samples, respectively. The vesicle data obtained include: area, major axis, minor axis, and perimeter. Equivalent radius

and diameter were calculated from the area values. Since we are interested in an internal comparison, here we report 2D values to avoid complications due to stereological conversion (Polacci et al., 2006).

# **Crystal Size Distributions**

We used BSE images to collect crystal size data of plagioclase crystals from the brown and black scoria. Long and short axes were measured using ImageJ freeware (http://rsb.info.nih.gov/ij/). We estimated crystal aspect ratios using the CSDSlice technique (Morgan and Jerram, 2006) and considered the textural fabric as massive; we did not correct for crystal roundness. Using petrography and built-in features of ImageJ, we determined the total slide area, volumetric phase abundance, and vesicularity. With these data we calculated the crystal size distributions (CSDs) of the plagioclase using the CSD Corrections 1.51 (http://www.uqac.ca/mhiggins/csdcorrections.html) after Higgins (2006). Higgins (2000) reports maximum errors of ±0.5 ln units in the population density.

## PETROGRAPHIC RESULTS

#### Scoria Textures

The black scoria usually occurs as sub-angular lapilli, although flatter examples (i.e. "plate tephra" - Ruth and Calder, 2014) are common in the proximal (< 9 km) deposit. The brown scoria only occurs as sub-angular, highly vesicular lapilli. General crystallinity (groundmass + phenocrysts) for the black scoria is between 50 to 60 vol%, whereas the brown scoria crystallinity ranges from 10 to 15 vol% (Ruth and Calder, 2014). Average vesicularity for both scoria types was derived from density measurements (after Houghton and Wilson, 1989).

Black scoria vesicles tend to be large, convoluted, with an average vesicularity of 79%. Conversely, vesicles in the brown scoria tend to be smaller and rounder, with an average vesicularity of 88% (Ruth and Calder, 2014).

Phenocryst phase assemblage is plagioclase (~1-2 vol% in brown scoria, 4-7 vol% in black scoria), olivine (<1 vol% in both scoria), rare titanomagnetite and orthopyroxene (<< 1 vol%). Plagioclase is the dominant mineral in the groundmass and is found in all scoria types. Titanomagnetite and orthopyroxene are present as groundmass phases in the black scoria, but are rarely observed in the brown scoria.

Plagioclase phenocrysts show significant signs of disequilibrium including sieve textures and compositional zoning (also observed by Bouvet de Maisonneuve et al., 2012b). Some plagioclase phenocrysts show additional features such as dendrite-like or swallow-tail terminations (Fig. 5a). Plagioclase microphenocrysts have an euhedral, tabular crystal habit and instances of hollow cores, hopper shapes, and minor swallow-tail terminations. They show little to no sieve textures and are sometimes normally zoned; these are found only in the black scoria.

A wide range of microlite morphologies are observed in the black scoria, where habits range from tabular to acicular, often in an interlocking-like configuration (Fig. 5b). Within the brown scoria, microlite habits are dominantly acicular, though tabular microlite are sometimes observed. Crystal free matrix glass is limited within the black scoria, whereas it is more abundant in the brown scoria (Fig. 5c). Additionally, microlite-rich textural domains, similar to the black scoria, are present in the brown scoria (Fig. 5d). Late-stage oxide microlites exhibiting cruciform habits are often found along the boundary separating the two textural domains (Ruth and Calder, 2014). Additional images are provided in the supplemental material. A summary of the textural features of the brown and black scoria is given in Table 1.

Plagioclase and olivine (i.e. troctolite) glomerocryst clusters are also present in all tephra. They are subdivided into two groups based on textural features and mineralogy (Fig. 6). One suite of glomerocrysts shows anhedral, equigranular, plagioclase and olivine with limited amounts of matrix glass containing round vesicles (Fig. 6a). The plagioclase phenocrysts are minimally sieve-textured. Subhedral to euhedral spinel are observed as inclusions in pristine plagioclase and olivine crystals; these are hereafter referred to as Cr-spinel troctolite glomerocrysts. In the other suite of glomerocrysts, the plagioclase phenocrysts show significant sieve textures, the olivine phenocrysts are much smaller and less abundant, matrix glass is generally more abundant with convolute and round vesicles, and the spinel phase is rare (Fig. 6b). These are referred to as Cr-spinel-free troctolite glomerocrysts. Cr-spinel troctolite glomerocrysts are generally found in the brown scoria, whereas the Cr-spinel-free troctolite glomerocrysts are found in the brown and black scoria. The glomerocrysts were noted but not studied by Bouvet de Maisonneuve et al. (2012a), and were thought to originate from a crystal mush. This inference is reasonable given that the glomerocrysts mineral chemistry is similar to that of the scoria. Additionally, it is unlikely that they are assimilated country rock given that the underlying basement comprises evolved Tertiary intrusives (Lucassen et al., 2004).

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## **Crystal and Vesicle Size Distributions**

Results from the vesicle textural analysis are shown in Table 2. Vesicle number densities are similar for the brown scoria samples (e.g. 124 and 142), whereas the black scoria samples exhibit higher and lower values (e.g. 106 and 166). The larger vesicle number density value for 1Black is puzzling given that the crystallinity and vesicularity are distinctly different than the brown scoria (see below). One explanation is that the textures could reflect mingling of magma

at different textural stages (e.g. Lautze and Houghton, 2007). The black scoria have a larger average vesicle diameter (0.074 mm) and a wider range of vesicle sizes (<0.01 to 1.74 mm), whereas vesicles in the brown scoria are smaller (0.044-0.055 mm) and are characterized by narrower range of sizes (<0.01 to 0.4 mm) (Fig. 7). The black and brown scoria have broadly similar vesicle size distributions (VSDs) with linear segments (Fig. 7), indicative of steady state nucleation and growth (Mangan and Cashman, 1996). However, the VSD of the black scoria diverges from the brown scoria around 0.1 mm (see Fig. 7 inset), and is characterized by a lower slope indicative of longer growth rates. The brown scoria shows a small plateau in the smallest vesicle diameters, which may reflect the initial stages of vesicle ripening (Mangan and Cashman, 1996).

The statistical summary for the plagioclase crystal size data is presented in Table 3. Crystal number density (CND), the number of crystals per unit area, generally decreases from the black scoria (including plate tephra for comparison) to brown scoria. Sample BrownLL has an exceptionally large CND and the average crystal size for this sample (~4 µm) is almost half the average for the other brown scoria (7 and 9 µm). The black and brown scoria have generally concave-up crystal size distributions (CSDs), whereas the plate tephra have a kinked profile (Fig. 8). The plate tephra have a significant component of larger crystal sizes, which manifests as another kink in the CSD; it is unclear to what extent this might be a sampling artifact. However, the plate tephra formed from ruptured bubble films, probably from magma that resided near the conduit margins (Ruth and Calder, 2014). Magma near the conduit margins likely had the longest residence times, thus producing the highest crystallinities and largest crystals. One sample of brown tephra, brn1, exhibits textures similar to that of the black scoria, which could suggest it is a transitional texture. The concave-up distribution has been attributed to super exponential

nucleation events (i.e. growth rate is a function of crystal size; Marsh, 1998) while a kinked profile is commonly observed in igneous systems and interpreted as a signature of magma mixing (Marsh, 1998; Higgins, 2006; Vinet and Higgins, 2010). Overall, the slope decreases from the brown to black scoria and plate tephra. Similarly, plot intercepts decrease slightly from the brown to black scoria and plate tephra (see Fig. 8 insets). Changes in slope can be attributed to increased residence time and/or coarsening (Higgins, 2006). Plot intercepts are not expected to change with increased residence time, however they do decrease with coarsening (Vinet and Higgins, 2010). BrownLL has the highest intercept of all the tephra, which combined with the high crystal number density, suggests a rapid nucleation event. Since the plot intercepts decrease in these datasets, the change in slope might result from coarsening via Ostwald ripening. Both plate tephra and one black scoria (1Black) samples show a kink in their CSD at the smallest crystal sizes, which is not observed in the brown or black scoria. This extra kink could suggest that these samples may have experienced coarsening with a nucleation event right before eruption. Similar distributions and interpretations were presented theoretically (Marsh, 1998) and in samples from Soufrière Hills volcano (Higgins and Roberge, 2003).

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## **GEOCHEMICAL RESULTS**

## **Whole Rocks - Major Elements**

The whole rock composition of the tephra produced during the 2008 eruption is basaltic with ~51 wt% of SiO<sub>2</sub> (n=5 whole rock samples), consistent with previously reported whole rock chemistry from recent Llaima products (Naranjo and Moreno 2005; Reubi et al., 2011; Bouvet de Maisonneuve et al., 2012a). These and additional geochemical data are provided in the supplemental material. The black scoria samples have SiO<sub>2</sub> values of 51.2 to 51.59 wt%; brown

scoria samples have similar SiO<sub>2</sub> values, at 51.20 to 51.22 wt%. These data suggest, unlike the textural data, that there is no discernible geochemical difference in whole rock chemistry between the black and brown scoria. This finding is consistent with texturally bimodal tephra reported elsewhere (e.g. Métrich et al., 2010; Cimarelli et al., 2010).

# **Phenocrysts - Major Elements**

Mineral chemistry was determined on the phenocryst phases, the plagioclase microphenocrysts, and all phases in the glomerocrysts (Table 4). The composition of the isolated plagioclase phenocrysts and microphenocrysts ranges from  $An_{85-65}$  with an average of  $An_{69}$  (Fig. 9). Isolated olivine phenocryst compositions range between  $Fo_{83-66}$ . Plagioclase within the Cr-spinel-free troctolite glomerocrysts ranges from  $An_{92-65}$ , whereas the plagioclase in the Cr-spinel troctolite glomerocrysts is slightly more calcic at  $An_{92-75}$ . Olivine composition within both types of glomercrysts clusters around  $Fo_{81}$ . Zoning information is limited to a small subset (n=14) of olivine crystals; these are dominantly reversed zoned, which is consistent with previously reported observations (Bouvet de Maisonneuve et al., 2012a). The spinel phase is Cr and Al-rich with Cr# (molar Cr/(Cr+Al+Fe<sup>3+</sup>)) ranging from 26-37. Cr-spinel in conjunction with high An plagioclase is consistent with origins in the deeper part of the magmatic plumbing system (e.g. Streck et al., 2002). Conversely, the lack of spinel and slightly lower An plagioclase in the Cr-spinel-free troctolite glomerocrysts is consistent with origins in the shallower parts of the plumbing system.

# Post Entrapment Modification of Melt Inclusions and Fe Oxidation State

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Melt inclusions may experience compositional modification after trapping. Processes may include crystallization, melting, re-equilibration with the host mineral, and elemental diffusion through the host mineral (for a review, see Kent, 2008). In order to undertake meaningful analysis of the data, estimating the extent of these modifications is important.

Post entrapment modifications through Fe loss or crystallization were determined by comparing the Fe-Mg exchange coefficients (K<sub>D</sub>) of the melt inclusion and host olivine (Putirka, 2008) (Fig. 10). Melt inclusion compositions were corrected for post entrapment crystallization by adding host olivine in 0.01% increments until  $K_D(Fe^{2+}-Mg)^{ol-liq} = 0.3$  (Roeder and Emslie, 1970) using the Ford et al. (1983) olivine-melt model in the PETROLOG 3 software package (Danyushevsky and Plechov, 2011). Fe<sup>3+</sup>/\sumset Fe ratios were measured via XANES on two melt inclusions; blk3-2011 yielded  $0.194 \pm 0.002$  whereas blk1-2010 was 0.163. These values were used for post entrapment corrections for these samples. For samples that were not analyzed with XANES, our post entrapment crystallization (PEC) corrections assumed an oxygen fugacity equivalent to the nickel-nickel oxide (NNO) oxygen buffer typical of arc systems (Toplis and Carrol, 1995) and a Fe<sup>3+</sup>/\(\sumes\) Fe ratio of 0.2, based on literature values for other arc systems (e.g., Cervantes and Wallace, 2003; Lloyd et al., 2013). We tested this assumption using equation 7 from Kress and Carmichael (1991) to estimate the oxygen fugacity buffer with a range of Fe $^{3+}/\Sigma$ Fe values (Table 5). When 20% Fe is Fe $_2$ O $_3$ , the buffer is  $\Delta NNO$  +0.06 to +0.14 for the samples shown. At Fe<sup>3+</sup>/ $\Sigma$ Fe = 15%, the buffer is much closer to  $\Delta$ NNO (-0.72 to -0.63), or  $\Delta$ QFM. The buffer did not change considerably with changing pressures in the model. Under this assumption, most samples experienced < 10% post entrapment crystallization.

We assessed Fe loss to the host crystal (e.g. Danyushevksy et al., 2000) by comparing the FeO\* and Mg# of the melt inclusions (anhydrous, normalized data) to the FeO\* and Mg# trend of whole rock samples (Fig. 11). Melt inclusion compositions are consistent with the whole rock array, but extend to significantly more mafic compositions. Four samples showed evidence of post-entrapment Fe-loss. These inclusions were all less than 40 microns, considerably smaller than other melt inclusions in the suite, which ranged from <20 to 425 microns in diameter (average 68 µm). As such, they could re-equilibrate faster than larger melt inclusions and therefore, their measured compositions may not reflect actual magmatic compositions (Cottrell et al., 2002). Due to well documented post entrapment complications from Fe-loss (e.g. Danyushevsky et al., 2000), these melt inclusions were filtered from the suite.

Water concentrations in olivine-hosted melt inclusions may be affected by post entrapment diffusion of H, and possibly molecular water, through the olivine crystal lattice (Métrich and Wallace, 2008; Johnson et al., 2008; Portnyagin et al., 2008). Three lines of evidence argue against significant diffusive water loss from the majority of our inclusions. First, significant water loss may cause extensive (>10%) post entrapment olivine crystallization by increasing the inclusion liquidus temperature (Lloyd et al., 2013; Bucholz et al., 2013). We therefore do not include any inclusions with > 10 % post-entrapment crystallization in this study. Second, H loss may initiate precipitation of fine magnetite within the melt inclusion (Danyushevsky et al., 2002; Métrich and Wallace, 2008), though this has been contested in other studies that propose that proton diffusion into the inclusion is rapid and would mitigate any changes in oxidation state (Gaetani et al., 2012; Bucholz et al., 2013). We do not include any inclusions with daughter crystals in this study. Third, comparisons of olivine-hosted melt inclusions sourced from ash, lapilli, and bombs show that samples sourced from lapilli, such as

those we report on here, may experience limited amounts of water diffusion (Lloyd et al., 2013). We consider diffusive water loss minimal, though it cannot be entirely excluded. In summary, going forth we consider only glassy melt inclusions that are free of daughter crystals although vapor bubbles may be present, are without significant Fe loss, and have less than 10% post entrapment crystallization.

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Recent studies have shown that CO<sub>2</sub> concentrations in melt inclusions are also altered by post-entrapment processes (e.g. Neave et al., 2014; Wallace et al., 2015; Aster et al., in preparation). In particular, post-entrapment modifications of the melt inclusion via cooling and crystallization causes a decrease in pressure. CO<sub>2</sub> partitions into the resultant vapor bubble because of the pressure dependence of its solubility. This can be augmented if there is H<sup>+</sup> loss via diffusion into the host mineral. Ultimately, these processes lower the CO<sub>2</sub> content of the melt inclusion by as much as 40 to 90% (Wallace et al., 2015; Aster et al., 2016). Recovering the original CO<sub>2</sub> contents of the melt inclusion requires knowing the volume of the vapor bubble and the CO<sub>2</sub> density within the bubble; however, we do not have these data for our suite of inclusions. Alternatively, original melt inclusion CO<sub>2</sub> contents may be estimated assuming that a bubble forms during thermal contraction associated with crystallization and cooling (Aster et al., 2016); the bubble volume can be estimated with crystallization and thermal contraction models. However, this correction assumes that the olivine have higher entrapment temperatures than eruption temperatures. Of the 182 melt inclusions analyzed in this study, entrapment temperatures were lower than eruption temperatures for 171 melt inclusions (calculated using the technique outlined in Aster et al. (2016)). These melt inclusions are less likely to have experienced thermal contraction. This suggests that the Aster et al. (2016) protocol may not be appropriate for these samples. We report uncorrected CO<sub>2</sub> values and consider these as minima.

## **Major Elements in Melt Inclusions**

Our melt inclusion dataset show a wide compositional range, with the most mafic and most evolved compositions between 49 and 56 wt% SiO<sub>2</sub> respectively, expanding on the previously reported compositional range (Bouvet de Maisonneuve et al., 2012a) (Table 6). The matrix glasses show similar trends to the melt inclusions, but with a smaller range of values for all elements (e.g., 52 - 55 wt% of SiO<sub>2</sub>). The majority of the samples (152 of 182 melt inclusions) are consistent with a single liquid line of descent (LLD) fractionating of olivine and plagioclase (Fig. 12). CaO and Al<sub>2</sub>O<sub>3</sub> trends do not change slope with respect to SiO<sub>2</sub>, suggesting that clinopyroxene did not crystallize during the evolution of the melt. These observations are consistent with sample petrography, as the dominant phenocryst assemblage is plagioclase + olivine. MnO and FeO<sup>T</sup> (i.e. FeO + 0.8998Fe<sub>2</sub>O<sub>3</sub>) do not exhibit obvious trends with respect to SiO<sub>2</sub>. Na<sub>2</sub>O and K<sub>2</sub>O behave as incompatible elements for the range of SiO<sub>2</sub> presented, all increasing with increasing SiO<sub>2</sub> content.

We observe a tight LLD for  $TiO_2$  and  $P_2O_5$  vs  $SiO_2 < 52$  wt%. After 52 wt%  $SiO_2$ , we observe a factor of five range in  $TiO_2$  and  $P_2O_5$  values, falling into three groups (see Fig. 12). One group follows the main LLD. The second group shows decreasing  $TiO_2$  and  $P_2O_5$  (23 of 182 melt inclusions). For this subset,  $Al_2O_3$  and  $Na_2O$  show slightly elevated concentrations with respect to the main LLD. These findings are consistent with localized assimilation of a lithology more evolved than the main LLD, possibly Tertiary intrusives (e.g. Lucassen et al., 2004) or evolved ancestral magmas potentially with Ti-magnetite (Reubi et al., 2011). A third group shows higher values of  $TiO_2$  and  $P_2O_5$  (7 of 182 melt inclusions). These samples also have some of the lowest CaO,  $Al_2O_3$ , and  $Na_2O$  concentrations in the dataset. These data are consistent with

formation in a residual, interstitial melt in the crystal mush (e.g. Bouvet de Maisonneuve et al., 2012a).

The P<sub>2</sub>O<sub>5</sub> values reported here are significantly higher than those reported in Bouvet de Maisonneuve et al. (2012a) (see Fig. 12). Our melt inclusion data are consistent with observed matrix glass and whole rock trends. Additionally, the whole rock values for P<sub>2</sub>O<sub>5</sub> reported by Bouvet de Maisonneuve et al. (2012a) overlap with the whole rock P<sub>2</sub>O<sub>5</sub> values reported in our study. These differences could be attributed to differences in analytical methodology such as count times, or primary standards.

The chemical compositions of melt inclusions sourced from the black and brown scoria overlap (Fig. 13); therefore, in this study we treat the melt inclusions as a single set. This observation contrasts with those of Métrich et al. (2010), who propose, for the Stromboli system, that olivine sourced from Stromboli brown scoria (i.e., "golden pumice" in their paper) originate from a different magma than their black scoria.

#### **Olivine-Hosted Melt Inclusion Trace Elements**

The multi-element spider plot (Fig. 14) shows the range and average of key elements including select rare earth elements (REE) and large-ion lithophile elements, which are used to infer information about petrogenetic processes. All data collected with LA-ICPMS are plotted and overlain by the calculated average value; we also included results from previously reported whole rock trace element data from Llaima Holocene tephras for comparison (dark gray region; Jacques et al., 2014). Our trace element data show a signature generally consistent with subduction zone magmas as evidenced by enrichments of fluid-mobile elements (e.g., Rb, Ba, K, U, Pb), a negative Nb-Ta anomaly, and moderately enriched light REE with respect to heavy

REE, as previously reported elsewhere (e.g. Hickey-Vargas et al., 1986; Reubi et al., 2011; Jacques et al., 2014). However, several elemental signatures do exhibit some notable variation from those data in the literature. First, we observe positive Eu anomalies in some of our melt inclusions, and negative Eu anomalies in others. Second, Ba/La in our dataset ranges from 17-46, whereas literature values for Llaima range from 23-27 (e.g., Reubi et al., 2011; Jacques et al., 2014). Similarly, the range of observed Nb-Ta anomalies is wider than previously reported (see Fig. 14). Conversely, Ce/Pb values in our samples (2-4) overlap well with literature values (2-3) (Reubi et al., 2011; Jacques et al., 2014).

#### **Volatiles in Melt Inclusions and Matrix Glass**

Volatile contents within the melt inclusions were measured with the methodologies discussed earlier; the results are shown in Table 6.  $H_2O$  and  $CO_2$  were not measured in the matrix glass because it was not possible to prepare FTIR wafers without crystals that interfered with the path of the IR beam. S and Cl in the matrix glass were measured with EMPA and range from 6-600 ppm and 330-500 ppm, respectively.

FTIR results for H<sub>2</sub>O in the melt inclusions range from 0.36 to 2.97 wt%. Similarly, the SIMS analytical results have H<sub>2</sub>O contents ranging from 0.57 to 2.37 wt% H<sub>2</sub>O. H<sub>2</sub>O values correspond with those reported in previous work (Bouvet de Maisonneuve et al., 2012a) (Fig. 15). CO<sub>2</sub> values from the FTIR analysis range from 397 to 1973 ppm, contrasting with the lower CO<sub>2</sub> from SIMS analysis (22 to 274 ppm). We would expect the FTIR analysis to record the lower CO<sub>2</sub> contents. However, many of our samples exhibited fringes (e.g. Nichols and Wysoczanski, 2007), the presence of which could have obscured the small CO<sub>3</sub><sup>2-</sup> doublet that

would result from low CO<sub>2</sub> in the melt inclusions. The CO<sub>2</sub> values presented here are the highest measured in recent Llaima samples to date.

Sulfur contents from the EPMA analyses range from 23 to 1534 ppm. SIMS analyses report sulfur abundances from 92 to 1277 ppm. Similarly, chlorine results from EPMA show concentrations from 47 to 1342 ppm, whereas SIMS results have abundances from 236 to 641 ppm.

We were unable to measure volatiles using all techniques on each melt inclusion, thus direct comparisons between the techniques were not possible with this dataset. However, the measured values from the SIMS for H<sub>2</sub>O, S, and Cl overlap well with those reported from EPMA and FTIR (see Table 6). For S and Cl, the EPMA results show a wider range of values compared with the SIMS results. Similarly, the range of FTIR H<sub>2</sub>O concentrations is wider than those obtained via SIMS. CO<sub>2</sub> concentrations measured using SIMS were lower, on average, than those measured using FTIR; however, we were only able to retrieve CO<sub>2</sub> values using FTIR from 12 samples.

#### **Entrapment and Magma Temperature Calculations**

Entrapment temperature for each melt inclusion-olivine pair was estimated using olivine-liquid equilibria, which is based dominantly on the Mg partitioning behavior between olivine and the melt and has been calibrated for hydrous melts (Putirka et al., 2007; Putirka et al., 2008). For this study, we used equation 22 from Putirka (2008), which includes the effect of H<sub>2</sub>O on the temperature calculation. H<sub>2</sub>O contents measured with FTIR or SIMS were included for the respective melt inclusions. We assumed an anhydrous melt for those melt inclusions without H<sub>2</sub>O content data. Using Equation 22 from Putirka (2008), the melt inclusion entrapment

temperatures ranged from 1094 °C to 1233 °C, with an average of 1175 °C. This range of entrapment temperatures strongly suggests that olivine grew in a wide variety of magmatic thermal conditions, and that cooling-induced crystallization was likely an important process.

To further constrain magmatic temperatures in the shallow plumbing system, plagioclase-liquid equilibria, based on the alkali/alkali earth partitioning between the plagioclase and the melt (Putirka, 2008) was also considered. Assuming equilibrium and using the composition of the microphenocryst plagioclase and the anhydrous matrix glass, we estimated the average temperature just prior to eruption at 1146 °C (low 1109 °C, high 1167 °C). We compared these magmatic temperatures with those calculated from the plagioclase hygrometer proposed by Lange et al., (2009). Using the matrix glass H<sub>2</sub>O values (0.88 wt%, rounded to 0.9 wt% - Bouvet de Maisonneuve et al, 2012a), the average matrix glass composition (SiO<sub>2</sub>: 53.36 wt%, TiO<sub>2</sub>: 1.34 wt%, Al<sub>2</sub>O<sub>3</sub>: 15.56 wt%, FeO<sup>T</sup>:10.69 wt%, MgO: 5.69 wt%, CaO: 8.91 wt%, Na<sub>2</sub>O: 3.18 wt%, K<sub>2</sub>O: 0.76 wt%), the average plagioclase composition (An<sub>69</sub>), and pressure at 50 MPa, we calculated the minimum shallow magmatic temperatures as 1115 °C. This is lower than the results from the Purtirka model, but within the range of expected uncertainty (±25-30 °C) for geothermometers in general (Putirka, 2008). Based on this, we assume the calculated 1146 °C temperature as a maximum magmatic temperature.

#### **Melt Inclusion Entrapment Pressures**

The pressure-dependent solubility behavior of CO<sub>2</sub> and H<sub>2</sub>O can be used to estimate the entrapment pressures of the individual melt inclusions (Dixon et al., 1995; Métrich and Wallace, 2008). That H<sub>2</sub>O and CO<sub>2</sub> vary widely in our melt inclusions indicates melt inclusion formation (and associated olivine crystallization) occurred over a wide range of pressures. Two solubility

models have been commonly used to estimate entrapment pressures: VolatileCalc (Newman and Lowenstern, 2002) and the Papale et al. (2006) model. Moore (2008) established that the Papale model was the best choice for calc-alkaline magmas because VolatileCalc neglects the effect of Ca on CO<sub>2</sub> solubility. Note that we did not correct for the possible CO<sub>2</sub> migration into vapor bubbles in the melt inclusions and, therefore, consider these values as minimum entrapment pressures. The entrapment pressures for our suite of melt inclusions range from 8 to 342 MPa, based on the Papale model (Fig. 15). Assuming a crustal density of 2500 kg m<sup>-3</sup>, depths range between ~340 m and 13.9 km. We observe a strong peak between 50-100 MPa (1-4 km), suggesting a crystal rich region (i.e. the crystal mush). These data increase the depth of the crystallizing portion of the plumbing system as reported by Bouvet de Maisonneuve et al. (2012a) by a factor of three.

## **DISCUSSION**

# **System Wide Processes**

# Magmatic Evolution

The majority of our dataset shows evidence for a connected magmatic system that fractionates olivine + plagioclase. The most primitive melt inclusions follow the main liquid line of descent (LLD) (see Fig. 12). More compositional variability is observed in melt inclusions with > 52 wt% SiO<sub>2</sub>. Low TiO<sub>2</sub> melt inclusion compositions are most consistent with assimilation of more evolved lithologies, probably with SiO<sub>2</sub> values between 56-58 wt% and very low TiO<sub>2</sub> and P<sub>2</sub>O<sub>5</sub>. These could represent the basement Teritary intrusives (e.g. Lucassen et al., 2004) or evolved ancestral magmas (Reubi et al., 2011). Positive Eu anomaly values (1.10  $\pm$ 0.24) are indicative of plagioclase dissolution which supports assimilation. This assimilation

occurred within the inferred crystal mush zone as evidenced by entrapment pressures ranging from 50 to 150 MPa. Our interpretation is consistent with previous work that suggests limited assimilation occurs potentially as a late-stage shallow process (e.g. Reubi et al., 2011).

The high  $TiO_2$  melt inclusions have elevated  $P_2O_5$  and  $K_2O$ , and low values of  $Al_2O_3$  and CaO. Eu anomaly values average  $0.77 \pm 0.05$ , indicative of plagioclase fractionation. These data are consistent with melt inclusion formation in melts that fractionated a higher modal abundance of plagioclase than the main LLD inclusions. We propose that these melt inclusions formed from interstitial melts in crystal rich regions in the plumbing system (e.g. Bouvet de Maisonneuve et al., 2012a). Only two melt inclusions from the high  $TiO_2$  group have pressure estimates and these are 73 and 309 MPa, respectively. The latter entrapment pressure possibly suggests that crystal rich zones with evolved interstitial melts may be present in the deep system, as well as the crystal mush. Although we cannot quantitatively link the number of melt inclusions to magma volumes, the large number of melt inclusions analyzed (n=182) suggests that melt inclusions from the high  $TiO_2$  (n=7, 3.8% of the sample suite) and low  $TiO_2$  (n=23, 13% of the sample suite) groups both represent volumetrically minor amounts of magma compared to the main LLD. Our dataset contrasts with previously reported interpretations that suggest melt inclusions dominantly formed in isolated melts with different evolution histories (Bouvet de Maisonneuve et al., 2012a).

## Degassing and Crystallization

The melt inclusions display a wide range of CO<sub>2</sub> contents, from below detection to nearly 2000 ppm, but tightly cluster near 2 wt% water down to CO<sub>2</sub> contents of ~100 ppm. The population of inclusions with CO<sub>2</sub> below 100 ppm and with 2 wt% water argues for open system degassing (Fig. 15). In this work, we understand open system degassing as the process in which

volatiles exsolve, decouple from the parent magma and then escape through bubble ascent (i.e. fractional degassing) (Dixon and Stolper, 1995; Cashman, 2004). Our findings contrast with previous studies that suggest that explosive eruptions in mafic systems result from rapid magmatic ascent under closed system degassing conditions (e.g. Coltelli et al., 2005; Métrich and Wallace, 2008). For example, paroxysmal eruptions (i.e. violent Strombolian) at Stromboli show distinctly different degassing trends (Fig. 15a), which have been interpreted to record closed system degassing and ascent (Métrich et al., 2010). In closed system degassing, volatiles exsolve to form a vapor phase, and then remain with the parent melt (i.e. equilibrium degassing). When ascent is rapid, the vapor phase increases in volume fraction, and as it is coupled to the magma, it leads to explosive fragmentation. Although, passive degassing processes at the summit have been linked elsewhere to both open and/or closed system degassing processes in deeper plumbing systems (e.g. Burton et al., 2007), the melt inclusion evidence for open system degassing at Llaima indicates that volatiles, even at depth, are decoupled from the magma, and allowed to ascend. Therefore, to produce the 2008 violent Strombolian eruption, a mechanism must be invoked which can explain the observed transition from passive degassing to explosive activity, but is independent of the established deeply-rooted passive degassing system.

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Water loss through degassing promotes crystallization (e.g. Crabtree and Lange, 2011; Applegarth et al., 2013). Although crystallization does occur during syn-eruptive cooling and degassing, melt inclusion entrapment in olivine is not likely to occur syn-eruption. We explored the relationship between degassing and crystallization by comparing the respective proxy ratios (e.g. Blake et al., 2010; Johnson et al., 2010). Figure 16 displays the volatile/K<sub>2</sub>O ratios for H<sub>2</sub>O, S, Cl, and CO<sub>2</sub> versus MgO/K<sub>2</sub>O. If crystallization occurred without degassing, the data would follow a horizontal trend. If degassing was dominant, the data would follow vertical trend. Our

results show strong correlations between MgO/K<sub>2</sub>O and H<sub>2</sub>O/K<sub>2</sub>O, Cl/K<sub>2</sub>O, and S/K<sub>2</sub>O, indicating that degassing and crystallization are coupled processes. CO<sub>2</sub>/K<sub>2</sub>O correlated with MgO/K<sub>2</sub>O at lower CO<sub>2</sub> values (i.e. lower pressures), but do not correlate at higher CO<sub>2</sub> values. Indeed, the majority of the sample suite falls on the coupled crystallization and degassing trend, including samples with high volatile contents (e.g. high volatile/K<sub>2</sub>O ratios). Similarly, scoria matrix glass values fall on the trend with some of the lowest S/K<sub>2</sub>O and Cl/K<sub>2</sub>O values in the dataset. This indicates that the matrix glass was almost completely degassed due to crystallization. Some melt inclusions have similarly low S/K<sub>2</sub>O values, suggesting that the melt inclusions trapped largely degassed magma. These inferences suggest that crystallization and, importantly, degassing were coupled in the deep plumbing system (at least to 14 km depth), as well as the shallow system. These data combined with the wide range of olivine entrapment temperatures suggest that cooling-induced and degassing-induced crystallization processes may be equally important.

## **Shallow System Processes – Passive Degassing**

Textural Maturation and the Formation of Black and Brown Scoria

The brown scoria exhibit textural data indicative of rapid nucleation at relatively high undercooling, i.e. higher intercept on the CSD plot and high vesicle number density. The inferred initial stages of vesicle ripening in the brown scoria likely occurred during ascent in the conduit just before quenching. Similar interpretations and short timescales were presented for vesicle ripening in reticulite from Hawaii (e.g. Mangan and Cashman, 1996). Conversely, textural data from the black scoria suggest a rapid nucleation event but also more prolonged crystal growth, while vesicle textures suggesting steady-state nucleation with possibly some coalescence. These

data are indicative of longer residence times before eruption. Together, the observed textures are indicative of a continuum of textural maturation from the brown to black scoria that we propose results from background passive degassing. The presence of normally-zoned microphenocryst plagioclase predominantly in the black scoria supports this, as it suggests growth-dominated conditions associated with lower, possibly steady-state, rates of degassing (Cashman, 2004; Castro et al., 2005). In summary, the textural features of the black scoria reflect passive degassing conditions in the conduit and likely represent magma slowly circulating within the shallow plumbing system.

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Texturally bimodal scoria were first observed and described at Stromboli (Francalanci et al., 2004) and have since been observed in other systems (e.g. Croscat complex cone system – Cimarelli et al., 2010). The formation of texturally bimodal scoria in explosive systems has been attributed to two processes, depending on whether they are geochemically distinct, or not. Where scoria originated from geochemically-similar magma sourced from similar locations within the shallow plumbing system, textures are interpreted to record end members on a textural maturity continuum (Lautze and Houghton, 2007; Erlund et al., 2010; Cimarelli et al., 2010). In these cases, differences among scoria types are a function of the physical processes experienced. Alternatively, where distinctions can be made in the geochemistry, such as melt inclusion chemistry, the source magmas for the different scoria are chemically distinct and originate from different locations in the system (Métrich et al., 2001, 2010). At Llaima, the geochemical similarity of the scoria whole rock compositions (Fig. 12) and the fact that melt inclusions from the different scoria overlap (Fig. 13) indicate that differences are textural alone, and likely formed as a result of differing degassing histories, rather than from geochemically different magmas (at least with respect to major element chemistry).

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# **Plumbing System Conceptual Model**

Summarizing the scoria textures, olivine hosted melt inclusions, as well as findings from previous work (e.g. Bouvet de Maisonneuve et al., 2012a, 2012b, Ruth and Calder 2014) we present here a conceptual model of the plumbing system at Llaima from the vent to ~14 km depth (Fig. 17). We infer two magma accumulation regions based on the clustering of melt inclusion entrapment pressures (histogram in Fig. 15). The main magmatic region is the shallowest, extending from the conduit (300 m) to ~4 km. Magma that produced the black 'resident' scoria is likely sourced from this region. Most melt inclusions formed (or reequilibrated) at pressures from between 50-100 MPa. This strongly supports the crystal mush zone proposed in previous work (e.g. Bouvet de Maisonneuve et al., 2012a, 2012b). Crystallization of olivine and plagioclase occurred throughout the plumbing system as evidenced by the main LLD. Minor amounts of localized assimilation of evolved lithologies occurred in the crystal mush zone (e.g. low TiO<sub>2</sub> melt inclusions, 50-150 MPa entrapment pressures). High TiO<sub>2</sub> and P<sub>2</sub>O<sub>5</sub> melt inclusions with similar crystal mush pressures also evidence entrapment of interstitial melts. Reverse zoning in olivine crystals suggests this as the main region of magma mixing (Bouvet de Maisonneuve et al., 2012b; Ruth, 2014). This magma mixing probably produced the multiply zoned plagioclase and the sieve textured plagioclase phenocrysts, textures which form under dynamic conditions, especially with respect to magmatic H<sub>2</sub>O and temperature (Landi et al., 2004; Streck, 2008; Crabtree and Lange, 2011; Bouvet de Maisonneuve et al., 2012b). Since plagioclase crystals within the Cr-spinel-free troctolite glomerocrysts have similar sieve textures, we infer that they are also sourced from the crystal mush zone.

Olivine hosted melt inclusions show that the deeper plumbing system extends from 4 to 14 km depth. Although we do not have many samples from the deep system, we infer information about the processes occurring within this region from melt inclusion chemistry and the Cr-spinel troctolite glomerocrysts. Degassing occurs at up to 14 km depth, and is coupled with crystallization. Extensive crystallization is suggested by the one melt inclusion with high TiO<sub>2</sub> and P<sub>2</sub>O<sub>5</sub> with high entrapment pressure, which may have formed in the interstitial melts within this region. Moreover, the presence of Cr-spinel in some troctolite glomerocrysts is consistent with deeper origins that have experienced limited fractionation (e.g. Streck et al., 2002); these are most likely to be found at depths lower than the main crystal mush zone. This potentially suggests a crystal rich zone in the deeper system (e.g. Annen et al., 2006). Similar crystal rich zones have been proposed for the deeper system beneath Mt. Etna (e.g. Laiolo and Cigolini, 2006).

## Changes in Magmatic Physical Properties Due to Passive and Syn-eruption Degassing

Magma viscosity is affected by crystal abundance and morphology, which can be altered due to degassing processes. Analog experiments show that increasing microlite abundance with respect to phenocryst abundance can potentially increase effective viscosity by up to three orders of magnitude (Cimarelli et al., 2011). The effective viscosity may increase further as a result of syn-eruption (or passive) degassing and crystallization. Heating experiments on volatile-poor porphyritic basalts reveal that rapid, significant degassing and crystallization (up to 35%) can occur in degassed conduit melts, resulting in additional dendritic to spherulitic crystal nucleation on existing crystals (i.e. overgrowths) (Applegarth et al., 2013, see their Figure 12 and 13). We estimated the viscosity indirectly using the MAGMA software package (version 2.50.0133)

http://www.lanl.gov/orgs/ees/geodynamics/Wohletz/KWare/Index.htm), the average whole rock composition, and the average total crystallinity of the brown and black scoria (Table 8). The majority of the crystallinity in both cases is microlites since the phenocryst content for both scoria types is less than 10%. The brown scoria magma has a vesicle-free effective viscosity ranging from 5.4×10<sup>1</sup> to 1.33×10<sup>2</sup> Pa s. The black scoria magma has a vesicle-free effective viscosity that is generally two to three orders of magnitude higher (2.29×10<sup>4</sup> to 5.64×10<sup>4</sup> Pa s). Our first order approximation of viscosity differences is consistent with findings reported in Cimarelli et al. (2011). Additionally, in the black and brown scoria we observe dendritic overgrowth textures similar to those found in Applegarth et al. (2013), especially around boundaries between textural domains (Fig. 5d), and near vesicles and fractures (see supplemental material Fig. S3). The proximity to vesicles, in particular, suggests formation resulting from late stage degassing. Based on the experimental results from the literature and our textural data, we propose that significant viscosity changes resulted from both passive degassing prior to, and syneruptive degassing during, the 2008 eruption at Llaima.

#### **Transitions in Eruptive Style at Llaima**

Maintenance of an open conduit commonly associated with passive degassing requires that heat and magma be continually replenished (Kazahaya et al., 1994; Stevenson and Blake, 1998). Heat and magma supply to the upper plumbing system can occur through density-driven convective mixing between the degassed, and cooler shallow magmas and relatively volatile-rich and hotter, deeper magmas (Kazahaya et al., 1994; Allard, 1997; Stevenson and Blake, 1998; Beckett et al., 2014). An additional mechanism for inducing axial mixing which is likely to be

important, is that the ascending gas flux in a conduit will itself drive a return flow of magma (Pansino, 2012).

We propose that the background activity at Llaima is maintained through periodic small-batch mafic injections, as evidenced by crystal zoning, into the base of a crystal mush (1-4 km depth) (Fig. 17). Resultant hybrid magmas with initially moderate H<sub>2</sub>O content (~ 2 wt%) are mobilized and entrained into the axial mixing processes. The ascent and mixing results in degassing and crystallization, and we speculate that this magma would ultimately result in the black scoria type products with the observed mature crystal and vesicle textures.

For the 2008 eruptive episode, magma injection and mush remobilization, evidenced by reversely zoned olivine (e.g. Bouvet de Maisonneuve et al., 2012a,b; Ruth, 2014), have been proposed as a triggering mechanism (Bouvet de Maisonneuve et al., 2012a). However, the lack of significant ground deformation prior to the eruption suggests that any magma injection into the system must have been volumetrically minor (Remy et al., 2015; Delgado et al., in preparation). Further, violent Strombolian eruptions are highly pulsatory in nature and require high mass and gas fluxes to drive them (Pioli et al., 2009). Therefore, any robust discussion of the triggers of such an eruption must account for substantial gas accumulation and its rapid ascent. Three mechanisms of potential gas accumulation are discussed below:

(i) Crystal mush zones may facilitate gas accumulation. Analog experiments suggest that small vesicles can become trapped within particle-rich suspensions (i.e. gas holdup), which may regulate gas flux through the plumbing system (Belien et al., 2010; Tran et al., 2015). Populations of small vesicles can be added to the crystal mush through volatile saturation and subsequent nucleation. A crystal mush zone at Llaima is inferred at depths between 1 and 4 km, based on the melt inclusion entrapment pressures. With the average whole rock composition, and

2 wt% H<sub>2</sub>O, saturation pressures (calculated with Papale et al. (2006)) are between 70 and 80 MPa (~ 3 km depth), thus it would be reasonable to expect vesicle nucleation within this region. Cr-spinel-free troctolite glomerocrysts have matrix glass with vesicles, lending further support to this hypothesis. These vesicles may have formed during syn-eruption ascent, but their convoluted morphologies suggest some maturity. Further, ascending CO<sub>2</sub>-rich vesicles from depth may also be partially trapped in the crystal mush (Thomas et al., 1993; Costa et al., 2006). The passive degassing behavior at the summit and the open system degassing signature of the melt inclusions indicate that gas bubbles within the system are mobile. Some mobility may occur due to crystal settling within the mush (e.g. Boudreau, 2016), with some bubble ascent feeding the passive degassing. However, significant bubble ascent might be inhibited by the surrounding crystals. We propose that gas accumulates within the mush until a critical gas volume is reached. Once the bubbles are released from the crystal mush via remobilization (i.e. unlocked) the accumulated gas would ascend rapidly to generate explosive activity (e.g. Vergniolle and Jaupart, 1990; Parfitt and Wilson, 1995).

(ii) Extensive passive degassing could result in increased viscosity for magma residing within the upper conduit. Over time, this magma would stagnate and potentially form a rheologically stiffer (but still permeable) cap, above which magma circulation does not take place. Experimental work suggests that rheological barriers lead to elevated gas overpressures, ultimately increasing the intensity of Strombolian eruptions (Del Bello et al., 2015). Although we do not have direct evidence for such rheological barrier, the black scoria and plate tephra from the 2008 eruption have textures consistent with prolonged residence and slightly higher viscosities (relative to crystal-free magmas of the same composition), as implied by their higher crystallinities (Table 8). Moreover, with increased viscosity, vesicle coalescence and growth

could be retarded (Cashman, 2004; Tran et al., 2015) further augmenting gas accumulation beneath the barrier.

(iii) Complex plumbing system geometry involving sills could provide additional traps for gas accumulation. Experimental results indicate that sills in passively degassing systems may serve as effective traps by segregating gas via buoyant exchange flow between the sill and conduit (Menand and Phillips, 2007). The plumbing system geometry of Llaima is not well constrained. Regional tectonics and the alignment of vents may indicate that the plumbing system follows a NE-SW orientation (Cembrano and Lara, 2008; Schonwalder-Angel, 2015). However, there is limited discussion of sills and/or other sub-horizontal bodies. Recent studies using InSAR inferred the presence and inflation (6-15 cm) of a sill to the west of the summit at ~ 5 km depth prior to a subsequent explosive phase on 3 April 2009 (Delgado et al, in preparation); there is assumed post-eruptive subsidence after this event. Although inflation was not observed before the 2008 eruption, the findings in Delgado et al. (in preparation) do suggest that a potential short-lived trap for magma and gas accumulation could have occurred.

We propose the following scenario to explain the transition to the violent Strombolian eruption at Llaima on 1 January 2008 (Fig. 17). The eruption occurred in a system that was already passively degassing which was likely sustained by repeated, minor, magma injection events. Upon injection these magma batches degassed volatiles that partially filtered through the crystal mush. Simultaneously, some of these volatiles, as well as volatiles ascending from depth, accumulated in the mush and under the stiffened conduit magma until a critical gas volume was reached. Disruption and remobilization of the mush by injecting magmas unlocked the accumulated gas phase, which ascended rapidly to generate the explosive activity at the vent. Clearing the upper portions of the shallow magma reservoir allowed rapid, and likely chaotic,

ascent of the deeper, relatively volatile-rich, magma from depths up to 14 km. Partial mixing between the influx, crystal mush and the resident magma produced the glomerocrysts and the crystal-rich domains represented by the clots of black scoria in the brown scoria (Fig 5b). Thus magma injection and gas accumulation processes equally contributed to eruption triggering.

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## **Explosive Eruptions in Mafic Systems**

Explosive mafic eruptions are commonly associated with closed system degassing and ascent of magmas from depth (e.g. Roggensack et al., 1997; Métrich et al., 2010). Bouvet de Maisonneuve et al. (2012a, 2012b) suggest that the dominant cause of the 2008 explosive eruption of Llaima was injection-related, heat-induced rheological changes in a crystal mush. We agree with Bouvet de Maisonneuve et al. (2012a, 2012b) that magma injection is an important contribution in triggering the eruption. However, the injection and remobilization model alone does not fully explain transitions from passive summit degassing that continuously leaks gas to an explosive eruption characterized by pulsatory behavior that must be related to gas accumulation (Pioli et al., 2009). Previous studies of mafic volcanoes elsewhere highlight the importance of accumulated gas in producing explosive and pulsatory behavior (e.g. Andronico et al., 2009; Pioli et al., 2008; Belien, 2011; Preece et al., 2016). Gas accumulation within crystal mushes has been invoked to explain eruptive behavior in intermediate to silicic systems (Sisson and Bacon, 1999; Pistone et al., 2015; Edmonds et al., 2016). We would have more insight into this process if comparable studies of gas accumulation within mafic systems were available. Additionally, magma reservoirs with accumulated gas are more compressible, and thus able to accommodate more influx with limited ground deformation (Voight et al., 2010); this could explain the lack of observed ground deformation prior to the 2008 activity. Therefore, we posit that gas accumulation within the crystal mush, as well as other traps within the upper plumbing system, and its subsequent sudden disruption and ascent is an equally important contribution to eruption triggering. The proposed mechanism provides an alternative means for mafic systems to generate explosive eruptions that does not rely on rapid ascent of magmas under closed system degassing conditions. This is especially pertinent for understanding major explosive eruptions in mafic systems like Llaima that exhibit a dominantly, open system degassing signature.

### **CONCLUSIONS**

We conducted an extensive geochemical and textural study of the products from the 2008 eruption of Llaima. Our aim was to establish how the processes associated with the explosive eruption initiate and occur, specifically in the context of the established background passive degassing of the system. Here we present the three main conclusions from this work.

- 1. Texturally bimodal scoria were observed and analyzed. Crystal and vesicle size distributions reveal quantitative texturally differences between the "brown" and "black" scoria from the 2008 eruption of Llaima Volcano. The black scoria samples generally have higher crystal number densities and larger vesicles, whereas the brown scoria have lower crystal number densities and smaller vesicles. Overall, this suggests increasing textural maturation from the brown to the black scoria. Black scoria likely formed from magma that experienced prolonged passive degassing and crystallization in the shallow conduit whereas the brown scoria is characterized by features resulting from rapid syneruptive ascent, degassing and crystallization.
- 2. Our dataset provides the first evidence for an extended deep, mafic plumbing system with olivine-hosted melt inclusions that sampled magmas that range from 49 to 56 wt% SiO<sub>2</sub>

at pressures from 8-342 MPa (300 m to  $\sim$  14 km). The majority of samples formed at pressures between 50 and 100 MPa, consistent with a shallow crystal mush. Major elements record a main LLD consistent with plagioclase and olivine fractionation in a well-connected system. Two smaller groups of melt inclusions show higher and lower  $TiO_2$  and  $P_2O_5$  with respect to the main LLD. Low  $TiO_2$  and  $P_2O_5$  samples suggest localized assimilation of more evolved lithologies. The high  $TiO_2$  samples may have formed in isolated pockets of evolved melt; they form at high and low pressures suggestive of extensive crystallization throughout the system.  $H_2O$  and  $CO_2$  in the majority of the melt inclusions show evidence for open system degassing, which is coupled with crystallization throughout the system.

3. Based on the combined textural and geochemical data, we propose that the main magmatic locus is a crystal mush, located between 1 and 4 km depth. Small volumes of magma from the deeper system are periodically injected, quasi steady state, into the shallow system. This contributes to maintaining passive degassing at the summit. We propose that concurrent with the passive degassing, gradual gas accumulation in the crystal mush and below the rheologically stiffer upper conduit magma occurred over time such that a critical gas volume in the system was reached, and the system overpressure was surpassed by magma injection prior to the 2008 eruption. Magma remobilization unlocked the accumulated gas, which was able to ascend rapidly and generate the observed explosive and pulsatory activity at the summit. Our proposed scenario provides an alternative model to those that posit that explosive mafic volcanism results from rapid ascent of magma due under closed system degassing conditions. This mechanism

provides a means for systems with dominantly open system degassing behavior to transition from passive degassing to violent explosive eruptions.

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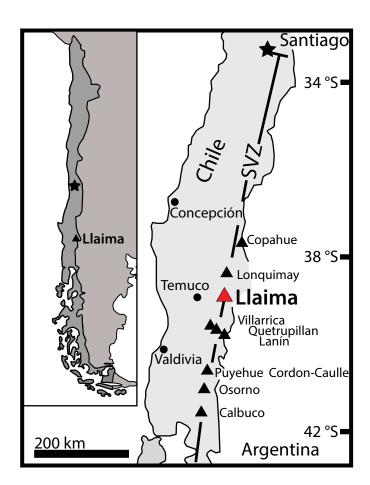


Figure 1. Location map of Llaima within the Southern Volcanic Zone of the Chilean Andes (33-46 °S). Shown for reference are the locations of select nearby volcanic systems, from north to south: Copahue, Lonquimay, Villarrica, Quetrupillan, Lanín, Puyehue Cordon-Caulle, Osorno, and Calbuco.

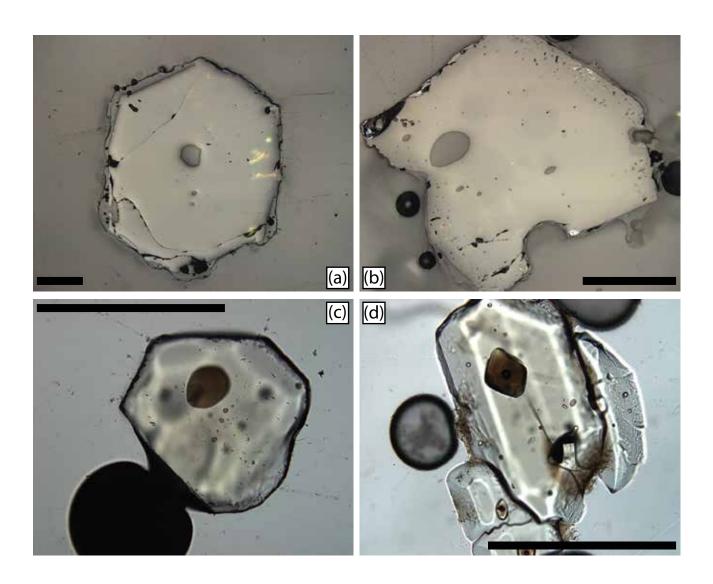


Figure 2. A range of different olivine and melt inclusion morphologies observed in the tephra samples. (A) and (b) are reflected light images whereas (c) and (d) are transmitted light images. All scale bars are  $500~\mu m$ .

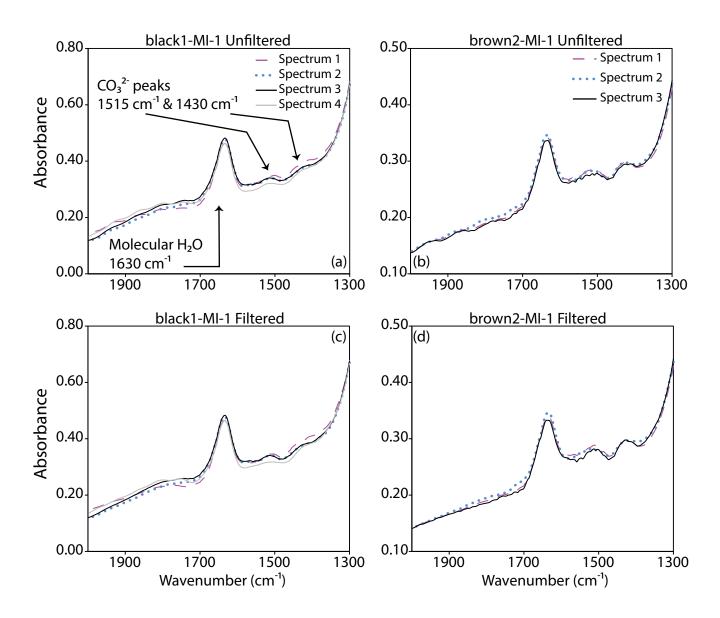


Figure 3. FTIR spectra showing the presence of fringes and the CO32- peaks. Panels (a) and (b) are unfiltered data whereas panels (c) and (d) show the fringes-filtered spectra. There was less than 5% difference between the raw and filtered spectra, indicating that the fringes did not affect the CO32- signal.

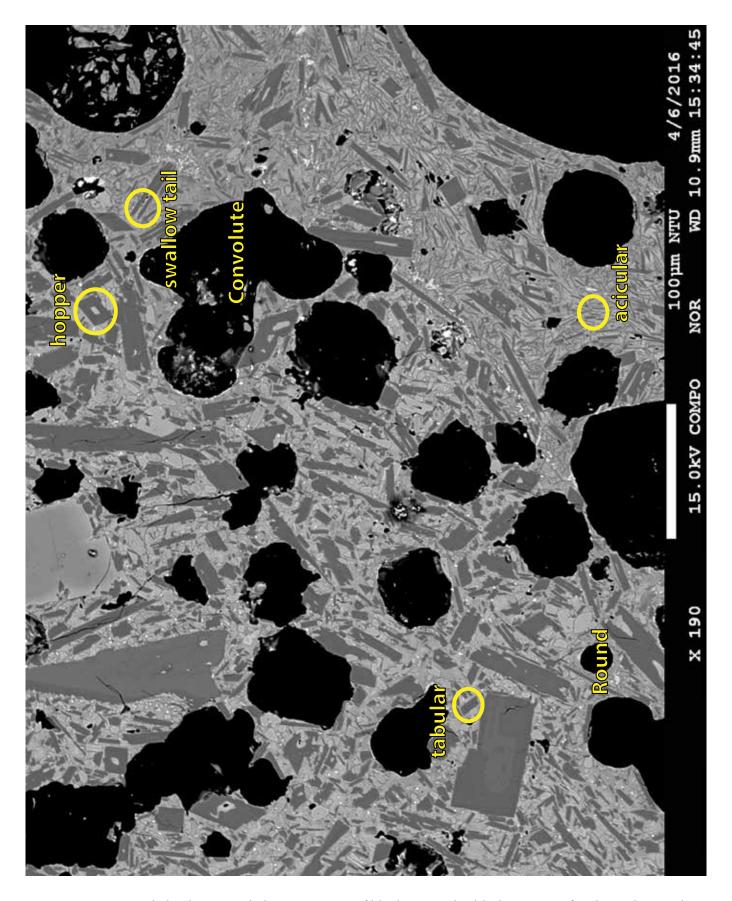


Figure 4. An example backscattered electron image of black scoria, highlighting specific plagioclase and vesicle textural features discussed in this study. With respect to plagioclase, we show hopper, swallow tail, tabular, and acicular morphologies. We highlight round and convolute vesicles.

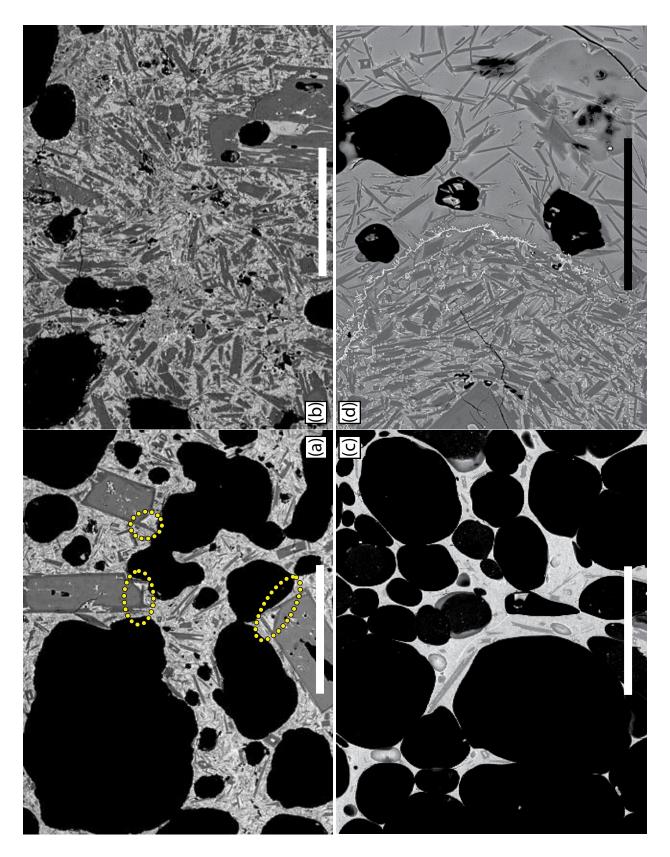


Figure 5. Backscattered electron images of the textural features of black and brown scoria. a) and b) The different microlite textures found in black scoria. In (a) there are abundant acicular plagioclase microlites (dark grey colour) alongside microphenocrysts and phenocrysts. Also note the circled plagioclase phenocrysts in showing the skeletal growths. In (b) the plagioclase microlites are dominantly tabular in morphology. c) The textural features of brown scoria, which show less abundant acicular microlites and abundant vesicles. d) The distinct boundary between two textural domains found in brown scoria. The boundary is highlighted by the spinel with cruciform morphology, suggesting rapid growth. The scale bar is  $100~\mu m$ .

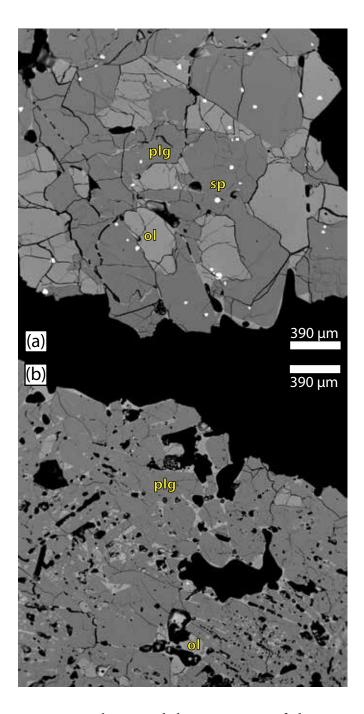


Figure 6. Backscattered electron images of plagioclase-olivine glomerocrysts. The bright phase is spinel, the medium gray phase is olivine, and the dark phase is plagioclase. a) Cr-spinel troctolite glomerocryst in equilibrium with olivine and spinel. Note there are a few vesicles in the surrounding matrix glass. b) Cr-spinel-free troctolite glomerocryst with significantly less olivine and no spinel present.

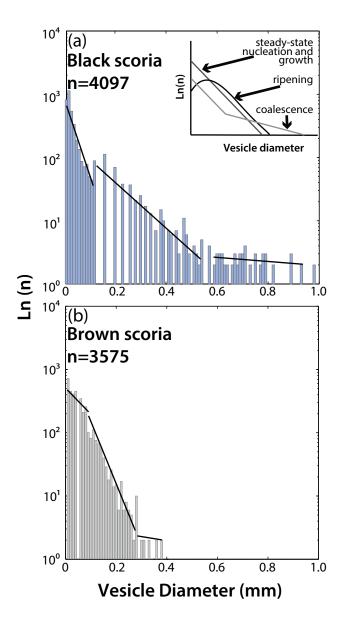


Figure 7. Vesicle size distribution histograms for a) black versus b) brown scoria. Inset figure shows interpretative models of different vesicle size histograms (after Mangan and Cashman, 1996). The lines on the histograms denote the approximate slope of each section. Note small plateau in the small vesicle sizes in the brown scoria, which are indicative of vesicle ripening (Mangan and Cashman, 1996). The black scoria exhibit two linear segments which suggests coalescence and residence.

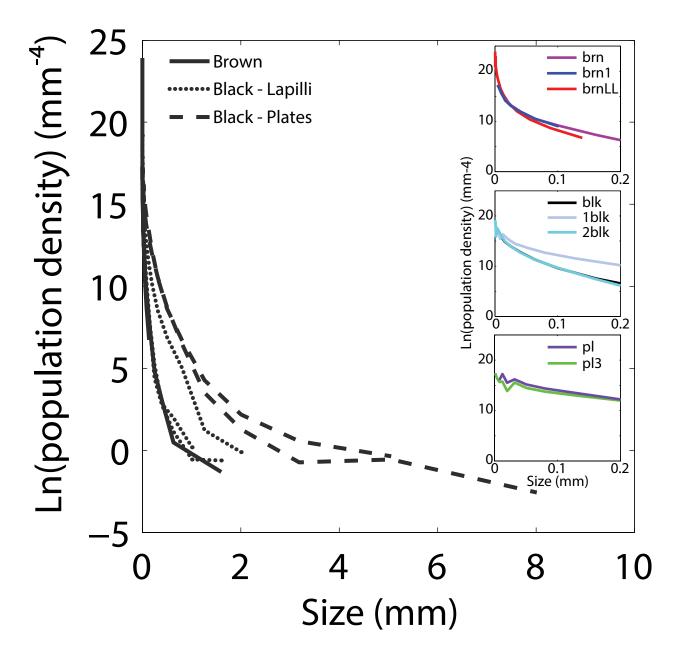


Figure 8. Crystal size distributions for the tephra erupted during the explosive phase. The insets show the breakdown by tephra type and sample. Note the decrease in slope from brown to black scoria to plates. Also note the decrease in y-intercept from brown to black scoria to plates.

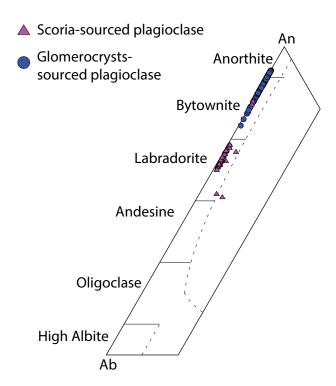


Figure 9. Modified plagioclase ternary diagram following disordered feldspar classification scheme of Deer et al. (1992). The triangles denote scoria-sourced microphenocryst plagioclase whereas the circles are plagioclases found in the glomerocrysts.

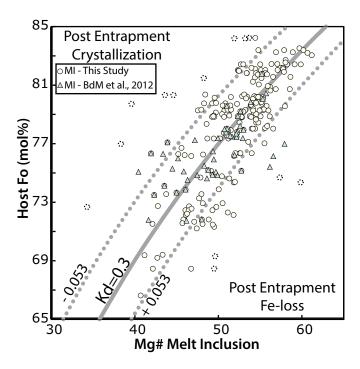


Figure 10. Rhodes diagram of Fo (mol%) of the host olivine and Mg# of the melt inclusion after Dungan et al., 1978; Rhodes et al., 1979). Samples are plotted assuming NNO, and Fe2O3 is 20% of the total Fe (Cervantes and Wallace, 2003). Those data in the top left experience post entrapment crystallization. Those data in the lower right experienced Fe-loss. The bold curve shows equilibrium which equates to Kd=0.3. The dotted lines are acceptable deviations from equilibrium (Putirka, 2008). The circles are melt inclusion analyses from this study whereas the triangles are analyses from the literature (Bouvet de Maisonneuve et al., 2012a). The dotted circles represent samples that were subsequently eliminated for further study due to likely post-entrapment modifications.

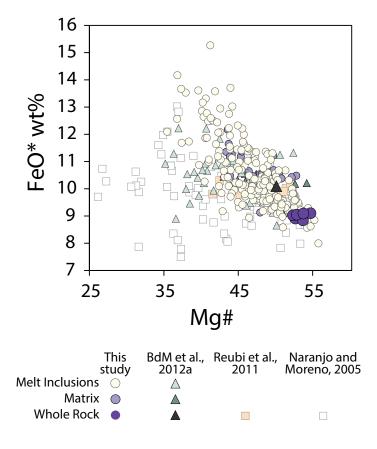


Figure 11. FeO\* (total from EMPA) versus the Mg# of the melt inclusion. Also plotted are the matrix glass, and whole rock analyses from this study, Bouvet de Maisonneuve et al. (2012a) (abbreviated as BdM et al. 2012a), Reubi et al. (2011), and Naranjo and Moreno (2005). If the melt inclusions plot along the same trend for the whole rock values then Fe-loss can be considered minimal.

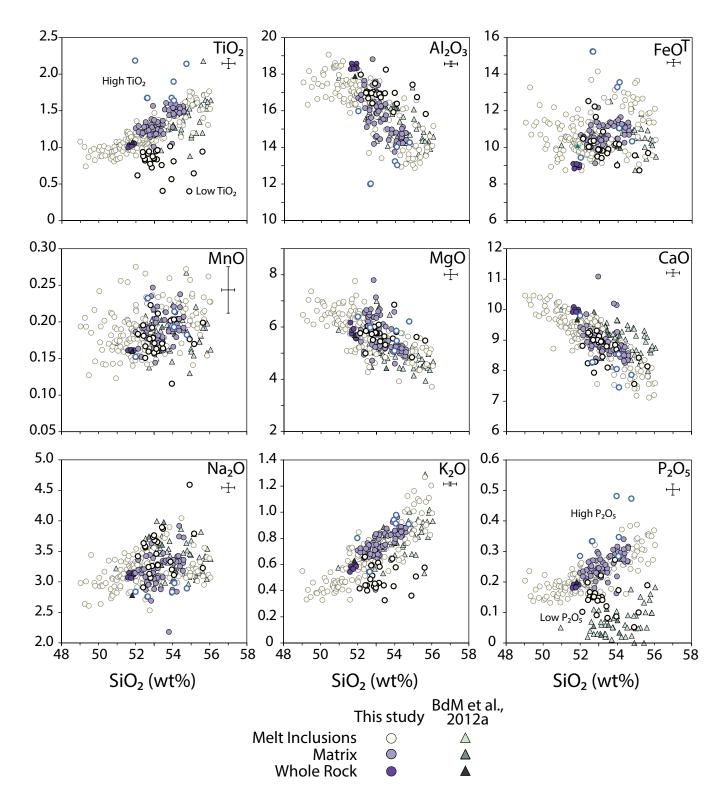


Figure 12. SiO2 diagrams showing the chemical range of the melt inclusions, matrix glass, and whole rock analyses. Samples from Bouvet de Maisonneuve et al. (2012a) are shown for comparison (triangles – same color scheme for whole rock, matrix glass, and melt inclusions). The crosses in the upper right corner represent the average standard deviation of the analyses. CaO, MgO, Al2O3 show negative correlations with SiO2, which is consistent with plagioclase and olivine crystallization. K2O, Na2O, P2O5, TiO2 show incompatible behavior. FeOT and MnO do not show any trends. All values are in wt%. Symbols with dark bold outlines highlight the low TiO2 and P2O5 samples whereas symbols with the lighter bold outlines denote samples with high TiO2 and P2O5.

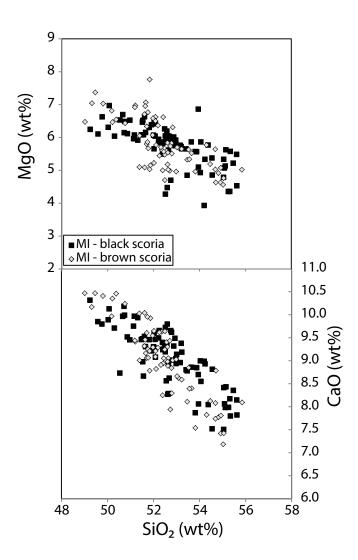


Figure 13. A comparison of melt inclusions sourced from the black and brown scoria to determine if there were distinct populations. The overlap of these data sets indicates that their respective melt inclusions are indistinguishable in terms of major elements (MgO and CaO).

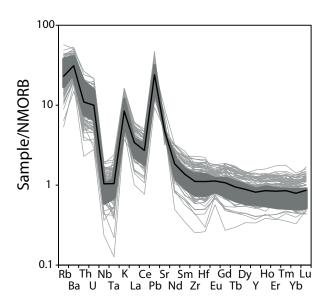


Figure 14. Multielement plot of trace elements in all analyzed melt inclusions with the average value shown in black. The dark gray field represents whole rock trace element data from Llaima Holocene tephra from Jacques et al. (2014). Note the subduction zone signature with the enriched large-ion lithophile elements, and light rare earth elements with respect to the heavy rare earth elements. Also note the range of Nb-Ta anomalies. Normalized values are from Sun and McDonough, 1989.

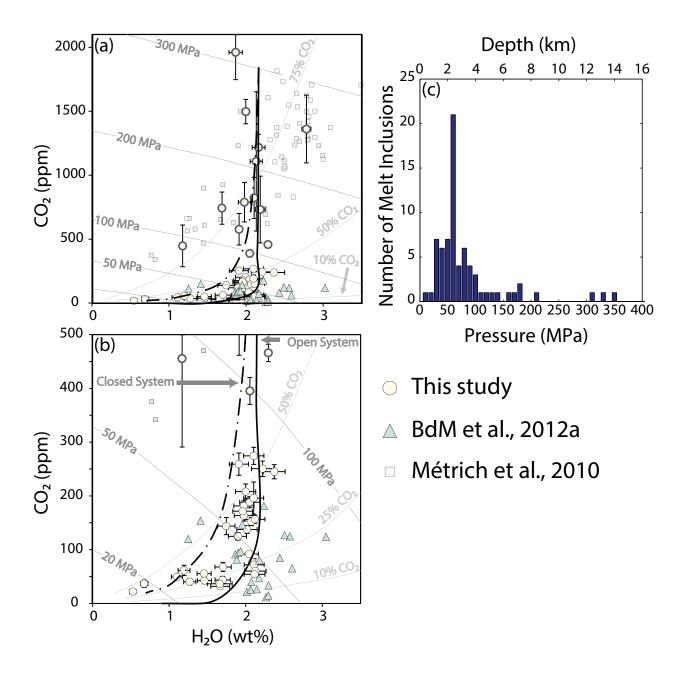


Figure 15. We used the model of Papale et al. (2006) to assess degassing style recorded by the melt inclusions. a) Plots of CO2 vs H2O from melt inclusions to show degassing behavior. Bold circles were analyzed with FTIR whereas the solid circles were analyzed with SIMS. The error bars are 2 sigma. Open (solid, black line) and closed (dash, black line) system degassing are modelled from 2000 ppm CO2, 2.15 wt% H2O. For the closed system, there is also an initial vapor phase of 1%. The dark gray lines are isobars and the light grey lines are isopleths showing the % CO2 in the vapor phase. All lines produced with Papale et al. (2006). b) Is a zoomed in version of a), note the vertical scale change. All symbols as in Figure 12. For comparison, melt inclusion volatile contents from paroxysmal eruptions at Stromboli reported in Métrich et al. (2010) are also plotted in a) and b). c) is a histogram showing the distribution of entrapment pressures for the plumbing system. Note the peak at 80 MPa.

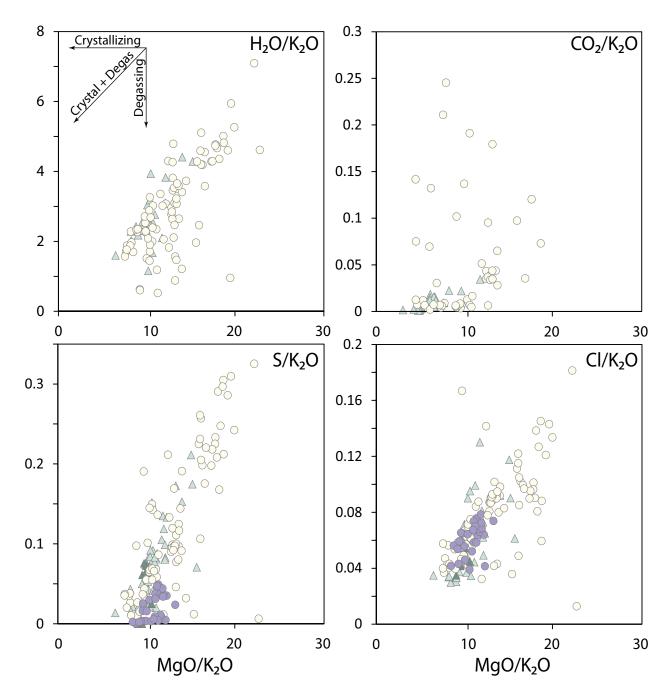


Figure 16. Volatiles ratios to K2O to show the relationship between degassing and crystallization. H2O, S, and Cl show good correlations with K2O, indicating that degassing and crystallization are coupled. CO2 shows more scatter, but at lower values (i.e. inferred shallower levels), degassing and crystallization are coupled. Matrix glass values for S and Cl provide information about degassing during the eruption. Both Cl and S in the matrix glass decrease, indicating syn-eruption degassing. Very low S values in the matrix glass suggest almost complete degassing whereas Cl did not completely degas.

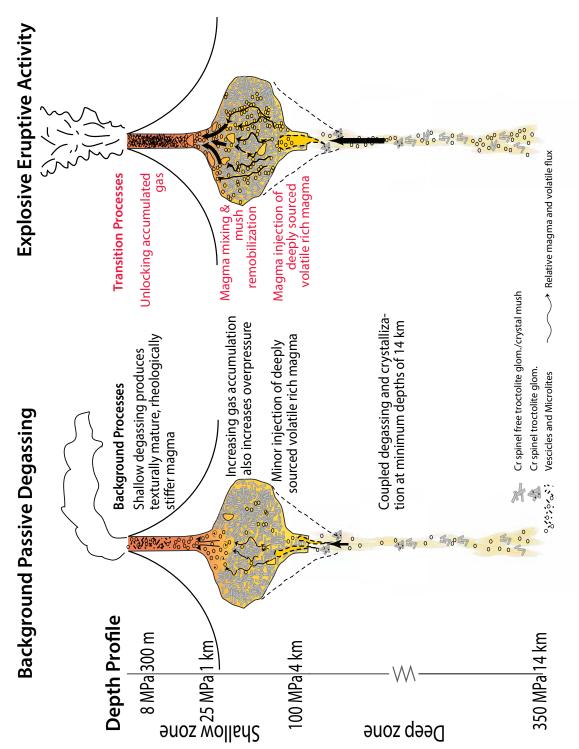


Figure 17. Schematic cartoon illustrating the plumbing system of Llaima and the key magmatic processes for the background passive degassing and transitions to the explosive eruption, respectively. Background passive degassing is characterized by continuous degassing from the summit, which results in the crystallization and rheological stiffening of the conduit magma. Minor volumes of magma are periodically injected into the crystal mush, which degasses and maintains the summit degassing. The deep plumbing system also experiences crystallization and degassing. Gas transport through the system is regulated mostly by the presence of the crystal mush, where vesicles get trapped within the interstitial spaces between crystals. Additional gas holdup may occur beneath the rheologically stiffer conduit magma. Gas may develop into larger bubbles, or more likely, foamy regions in the upper part of the crystal mush (the larger vesicles in the figure). The transition to the 2008 eruption at Llaima occurred after injection event that remobilized the crystal mush, and importantly, unlocked the accumulated gas that ascended rapidly to produce the observed explosive behavior at the summit.

Table 1. Summary table of crystalline and vesicle textures observed in the brown scoria and black lapilli scoria. For textural information on the black plate tephra, see Ruth and Calder (2014).

Crystallinity Micro-Vesicles (phenocrysts Phenocrysts Microlites Matrix glass Other phenocrysts +groundmass) plg: plg: tabular 79% 4-7 vol% and acicular, Black normally 50-60 vol% Limited interlocking zoned, convolute scoria plg, ol pristine configuration Domains plg: acicular, Brown 88%, 1-2 vol% Rarely with black 10-15 vol% sometimes Abundant round plg, ol observed scoria scoria tabular textures

Table 2. Textural input and output for vesicle size distribution measurements. Average diameter =  $2(area fraction/(\pi^*vesicle number density))^{1/2}$ .

, (,							
Sample	Total Area Measured (mm²)	Phenocryst area (mm²)	Total Area – Phenocryst Area	Number of vesicles measured	VND	Average diameter (mm)	SD (mm)
Brown	18.8	0.2	18.6	2315	124	0.055066	0.058757
brown	9.1	0.2	8.9	1260	142	0.044447	0.052181
1Black	13.7	0.6	13.0	2162	166	0.074175	0.155396
2Black	19.6	1.4	18.2	1935	106	0.074638	0.1865

Table 3. Textural inputs for crystal size distribution measurements. Crystal number density is the number of crystals per unit area.

arm area:								
Sample	Total Area (mm²)	Total Bubbles Area (mm²)	% Bubbles	Major axis	Intermediate axis	Minor axis	Number of crystals measured	CND
Brown	14.1374	10.41910152	73.7	2	1	1	5487	1476
Brown1	16.1369	11.8370804	73.4	2	1	1	5893	1371
BrownLL	1.5206	0.77426498	50.9	1.15	1.05	1	7485	10029
Black	6.9795	2.74364145	39.3	2	1	1	9298	2195
1Black	3.4565	1.72197592	49.8	10	1	1	8636	4979
2Black	7.511	3.7982712	50.6	2	1	1	9279	2499
Plate	2.2893	0.41159049	18	10	1	1	7917	4216
Plate3	3.7005	0.2743954	7.4	10	1	1	7332	2140

Table 4. Representative mineral analyses (olivine, plagioclase, spinel) obtained with the electron microprobe.

	BRN3-1- 2010	bk2-5-1- 2012	br3-1-1- 2012	br4-2-2- 2012	3 blk_ol_2- 52	15 brn_ol_3- 199	7 blk_ol_6- 128	6 blk_ol_5- 95
Type	Host ol	Host ol	Host ol	Host ol	Glom. ol	Glom. ol	Glom. ol	Glom. ol
SiO <sub>2</sub>	36.41	38.74	39.67	40.21	37.67	40.59	39.24	37.67
$TiO_2$	nd*	0.05	0.01	0.01	nd	nd	nd	nd
$AI_2O_3$	nd	0.02	0.01	0.02	nd	nd	nd	nd
$Cr_2O_3$	0.00	0.01	0.02	0.02	0.01	0.02	0.02	0.01
FeO	30.14	21.24	18.02	14.99	19.09	17.24	16.97	16.99
MnO	0.49	0.33	0.27	0.22	0.31	0.27	0.27	0.27
MgO	33.74	39.23	42.07	44.88	40.90	40.76	42.73	46.10
NiO	0.07	0.11	0.13	0.20	0.11	0.15	0.16	0.12
CaO	0.27	0.22	0.19	0.18	0.22	0.20	0.20	0.25
Total	101.11	99.94	100.38	100.72	98.32	99.22	99.60	101.40
Si	0.98	1.00	1.01	1.00	0.99	1.03	1.00	0.95
Ti	nd	0.00	0.00	0.00	nd	nd	nd	nd
Al	nd	0.00	0.00	0.00	nd	nd	nd	nd
Cr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Fe <sup>2+</sup>	0.68	0.46	0.38	0.31	0.42	0.37	0.36	0.36
Mn	0.01	0.01	0.01	0.00	0.01	0.01	0.01	0.01
Mg	1.35	1.51	1.59	1.67	1.60	1.55	1.62	1.73
Ni	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Ca	0.01	0.01	0.01	0.00	0.01	0.01	0.01	0.01
∑Cation	3.02	3.00	2.99	3.00	3.01	2.97	3.00	3.05
%Fo	66.25	76.42	80.39	84.03	78.98	80.58	81.54	82.64

Table 4. Continued.

-	13	15						
	brn_pl	brn_pl	Black-	WAQ-	WAQ-	Black-	Black-	ZRN -
	g_3-	g_5-	655	738	737	657	656	732
	236	269						
Туре	Glom.	Glom.	Scoria	Scoria	Scoria	Scoria	Scoria	Scoria
	plag							
$SiO_2$	44.46	44.07	53.75	52.52	52.57	52.18	49.38	47.70
$TiO_2$	0.03	0.00	0.036	0.175	0.068	0.036	0.007	0.018
$AI_2O_3$	33.95	34.39	28.75	29.11	29.89	30.00	32.56	32.81
FeO	0.57	0.54	1.22	1.08	0.92	0.84	0.95	0.58
MgO	0.07	0.07	0.31	0.33	0.26	0.26	0.30	0.12
CaO	18.08	18.39	11.80	12.24	13.21	13.54	13.08	16.85
$Na_2O$	1.09	0.87	4.24	4.14	3.87	3.74	3.32	2.01
$K_2O$	0.02	0.01	0.17	0.16	0.12	0.07	0.11	0.02
Total	98.27	98.34	100.28	99.75	100.90	100.67	99.70	100.11
Si	2.09	2.08	2.43	2.40	2.37	2.36	2.26	2.19
Ti	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.00
Al	1.88	1.91	1.53	1.56	1.59	1.60	1.76	1.78
Fe <sup>2+</sup>	0.02	0.02	0.05	0.04	0.03	0.03	0.04	0.02
Mg	0.00	0.01	0.02	0.02	0.02	0.02	0.02	0.01
Ca	0.91	0.93	0.57	0.60	0.64	0.66	0.64	0.83
Na	0.10	0.08	0.37	0.37	0.34	0.33	0.29	0.18
K	0.00	0.00	0.01	0.01	0.01	0.00	0.01	0.00
∑Cation	5.02	5.02	4.99	5.00	5.00	5.00	5.01	5.01
%An	90.07	92.05	59.97	61.41	64.90	66.42	68.05	82.18

Table 4. Continued

	17 Llaima5- 287	18 Llaima6- 294	18 Llaima7- 295	19 Llaima7- 296	20 Llaima8- 301	21 Llaima9- 304	24 Llaima12- 311
Type	oxide						
SiO <sub>2</sub>	0.08	0.10	0.11	0.20	0.16	0.05	0.11
$TiO_2$	0.61	1.25	1.25	0.63	1.40	1.03	1.46
$AI_2O_3$	31.19	27.79	27.29	32.94	20.80	27.52	21.87
$Cr_2O_3$	23.06	22.50	22.84	22.99	27.28	23.41	25.57
FeO	29.80	36.23	36.89	29.83	39.07	35.92	39.70
MnO	0.19	0.23	0.19	0.16	0.23	0.19	0.24
MgO	14.51	12.74	12.68	14.59	10.18	12.94	10.99
CaO	0.24	0.21	0.16	0.20	0.37	0.12	0.26
Total	99.69	101.05	101.40	101.53	99.49	101.18	100.20
Si	0.00	0.00	0.00	0.00	0.00	0.01	0.00
Ti	0.03	0.03	0.03	0.03	0.07	0.01	0.03
Al	0.94	0.86	0.79	0.74	0.53	1.11	0.77
Cr	0.55	0.60	0.62	0.73	0.66	0.52	0.68
Fe <sup>3+**</sup>	0.46	0.48	0.52	0.48	0.66	0.34	0.51
Fe <sup>2+**</sup>	0.45	0.48	0.50	0.50	0.60	0.37	0.52
Mn	0.01	0.01	0.01	0.01	0.01	0.00	0.01
Mg	0.55	0.53	0.51	0.50	0.46	0.62	0.48
Ni	0.00	0.00	0.00	0.00	0.00	0.00	0.00
∑Cation	2.99	3.00	2.99	2.99	3.00	2.99	2.99
Cr#	28.31	30.87	32.10	37.27	35.86	26.38	34.68

<sup>\*</sup>n.d. indicates not determined. \*\*Fe<sup>2+</sup> and Fe<sup>3+</sup> determined through charge balancing.

Table 5. Oxygen fugacity calculations using equation 7 from Kress and Carmichael (1991). All calculations at the given pressure and temperature.

Sample	BLK1-1-2010	blk3-MI-1-2011	blk8-MI-1-2011	brn8-MI-1-2011
$SiO_2$	52.04	53.09	51.68	50.93
$TiO_2$	1.01	1.55	0.92	1.06
$AI_2O_3$	15.74	12.89	16.95	16.96
FeO*	9.38	11.01	9.81	9.11
MnO	0.20	0.19	0.16	0.15
MgO	4.16	5.35	4.81	5.44
CaO	8.76	7.48	9.02	9.02
$Na_2O$	3.29	3.04	3.24	3.33
K <sub>2</sub> O	0.64	1.04	0.49	0.60
$P_2O_5$	0.27	0.29	0.15	0.19
Sum	95.48	95.94	97.23	96.79
T (°C)	1147	1188	1131	1152
P (MPa)	50	50	50	50
Fe³+ <b>/ΣFe</b>	0.15	0.15	0.15	0.15
ΔNNO	-0.72	-0.63	-0.66	-0.71
Fe³+ <b>/∑Fe</b>	0.20	0.20	0.20	0.20
ΔNNO	+0.06	+0.14	+0.12	+0.07
Fe³+ <b>/∑Fe</b>	0.25	0.25	0.25	0.25
ΔΝΝΟ	+0.693	+0.779	+0.75	+0.71

Table 6. Representative major element and volatile concentrations in melt inclusions. These measurements are after post entrapment correction. Uncorrected data can be found in the supplemental material. Melt inclusions labeled int were located in between the core and the rim.

rim.								
Sample	br1-3- 3-2012	blk11- MI-5- 2011	br1-3- 2-2012	BRN6- 3 - 2010*	bk3-2- 1-2012	bk1-2- 1-2012	blk8- MI-4- 2011	blk7- MI-2- 2011
MI Location	int	rim	core	core	int	int	int	int
MI Size (um)	90	41	104	18	113	101	40	82
Bubble	no	no	yes	no	no	no	yes	no
% PEC	1	4	1	4	0	4	5	3
$SiO_2$	49.02	49.76	50.21	50.72	50.72	50.74	50.97	51.52
$TiO_2$	1.08	0.99	0.97	0.95	0.94	0.89	1.08	0.92
$Al_2O_3$	17.52	16.62	16.71	18.35	18.58	16.96	16.77	17.50
FeOt	11.34	12.29	11.34	9.19	9.45	11.30	11.17	10.05
$Fe_2O_3$	1.80	1.98	1.80	1.23	1.33	1.69	1.73	1.46
FeO	9.71	10.51	9.73	8.09	8.25	9.78	9.61	8.73
MnO	0.19	0.22	0.14	0.21	0.16	0.17	0.18	0.17
MgO	6.47	6.62	6.46	6.84	6.14	6.48	6.53	6.46
CaO	10.47	9.80	10.35	10.21	10.18	9.99	9.45	9.47
$Na_2O$	2.97	2.98	2.93	2.78	3.06	2.77	3.02	3.20
$K_2O$	0.55	0.38	0.50	0.42	0.48	0.38	0.48	0.40
$P_2O_5$	0.20	0.15	0.19	0.20	0.15	0.15	0.19	0.16
Total	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00
mg#	0.50	0.49	0.50	0.57	0.54	0.51	0.51	0.53
Host Fo	80	79	80	83	82	80	80	81
S (ppm)	n.d.	n.d.	n.d.	n.d.	n.d.	927	1244	1153
CI (ppm)	n.d.	n.d.	n.d.	n.d.	n.d.	511	549	488
F (ppm)	n.d.	n.d.	n.d.	n.d.	n.d.	305	0	0
$H_2O$ (wt%)	n.d.	0.37	n.d.	n.d.	n.d.	2.01	n.d.	2.05
H <sub>2</sub> O error	n.d.	0.14	n.d.	n.d.	n.d.	0.13	n.d.	0.02
CO <sub>2</sub> (ppm)	n.d.	n.d.	n.d.	n.d.	n.d.	138	n.d.	394
CO <sub>2</sub> error	n.d.	n.d.	n.d.	n.d.	n.d.	7	n.d.	25
Temp (°C)	1218	1180	1212	1179	1206	1173	1158	1185
Pressure (MPa)	n.d.	n.d.	n.d.	n.d.	n.d.	49	n.d.	81
Depth (km)	n.d.	n.d.	n.d.	n.d.	n.d.	2.0	n.d.	3.3
n d indicatos not	dotormina	24						

n.d. indicates not determined

Table 6. Continued.

Sample	blk9- MI-1- 2011	BUL4- 3- 2010*	BUL4- 1- 2010*	br2-1- 1-2012	brn5- MI-1- 2011	bk2-8- 1-2012	bk4-2- 2-2012	bk4-1- 1-2012
MI Location	rim	int	int	int	core	core	rim	core
MI Size (um)	34	16	20	30	50	158	65	184
Bubble	no	yes	yes	no	yes	no	no	no
% PEC	4	1	0	0	0	1	2	0
$SiO_2$	51.54	51.55	52.01	52.02	52.20	52.22	52.25	52.65
TiO <sub>2</sub>	1.22	1.02	1.11	1.18	1.10	1.10	1.04	1.68
$Al_2O_3$	15.86	17.54	17.01	17.55	17.59	17.09	16.96	11.99
FeOt	11.53	10.34	10.66	9.92	10.14	9.71	9.94	15.22
$Fe_2O_3$	1.90	1.57	1.65	2.21	2.26	1.50	1.52	3.14
FeO	9.82	8.93	9.18	7.94	8.11	8.36	8.57	12.39
MnO	0.20	0.25	0.22	0.20	0.19	0.16	0.17	0.23
MgO	6.52	5.74	6.08	5.22	5.25	5.94	5.84	5.99
CaO	8.98	9.59	8.82	9.65	9.51	9.49	9.65	8.29
$Na_2O$	3.04	3.04	3.01	3.54	3.24	3.34	3.21	2.76
$K_2O$	0.73	0.55	0.66	0.56	0.59	0.61	0.61	0.55
$P_2O_5$	0.19	0.21	0.26	0.17	0.19	0.19	0.19	0.33
Total	100.00	100.00	100.00	100.22	100.23	100.00	100.00	100.00
mg#	0.50	0.50	0.50	0.48	0.48	0.52	0.51	0.41
Host Fo	80	79	79	76	77	81	80	68
S (ppm)	400	n.d.	n.d.	512	454	553	577	175
CI (ppm)	545	n.d.	n.d.	515	597	495	532	454
F (ppm)	0	n.d.	n.d.	300	n.d.	311	324	289
H <sub>2</sub> O (wt%)	n.d.	n.d.	2.00	0.68	n.d.	1.70	1.61	2.04
H <sub>2</sub> O error	n.d.	n.d.	0.04	0.05	n.d.	0.10	0.10	0.12
CO <sub>2</sub> (ppm)	n.d.	n.d.	n.d.	37	n.d.	44	25	92
CO <sub>2</sub> error	n.d.	n.d.	n.d.	7	n.d.	7	2	5
Temp (°C)	1171	1193	1219	1197	1192	1196	1176	1190
Pressure (MPa)	n.d.	n.d.	n.d.	13	n.d.	31	26	73
Depth (km)	n.d.	n.d.	n.d.	0.5	n.d.	1.3	1.1	3.0

Table 6. Continued.

Sample	bk4-4- 1-2012	brn6- MI-1- 2011	bk3-4- 3-2012	Plate1- 2010*	blk9- Ml-2- 2011	brn4- MI-1- 2011	blk5- Ml-1- 2011	bk2-6- 1-2012
MI Location	int	core	int	core	int	int	core	int
MI Size (um)	90	93	105	33	38	106	157	102
Bubble	no	no	yes	no	yes	no	no	yes
% PEC	0	0	2	1	3	0	5	0
$SiO_2$	52.88	52.90	52.93	54.06	54.06	54.08	54.21	54.24
$TiO_2$	0.73	1.23	0.88	0.82	1.90	1.65	1.43	1.52
$AI_2O_3$	17.47	14.95	16.87	16.36	14.00	13.01	15.71	13.53
FeOt	9.48	14.13	9.86	10.15	11.04	13.54	10.92	12.23
$Fe_2O_3$	1.36	2.82	1.54	1.53	1.99	2.71	2.04	2.72
FeO	8.26	11.60	8.48	8.78	9.25	11.11	9.08	9.78
MnO	0.19	0.24	0.18	0.20	0.19	0.21	0.22	0.23
MgO	6.01	4.62	5.79	5.81	5.86	5.48	3.93	5.33
CaO	9.32	7.57	9.01	8.58	8.55	7.46	8.99	8.93
$Na_2O$	3.25	3.08	3.73	3.11	2.99	2.97	3.36	2.98
$K_2O$	0.44	0.73	0.46	0.58	0.92	0.98	0.72	0.64
$P_{2}O_{5}$	0.09	0.27	0.14	0.17	0.28	0.35	0.30	0.37
Total	100.00	100.00	100.00	100.00	100.00	100.00	99.99	100.28
mg#	0.53	0.37	0.51	0.51	0.49	0.42	0.39	0.44
Host Fo	81	68	80	80	79	71	72	72
S (ppm)	731	482	995	n.d.	895	104	618	n.d.
CI (ppm)	352	412	435	n.d.	491	423	512	n.d.
F (ppm)	218	n.d.	257	n.d.	n.d.	n.d.	n.d.	n.d.
$H_2O$ (wt%)	1.90	n.d.	1.96	1.98	n.d.	1.93	n.d.	n.d.
H <sub>2</sub> O error	n.d.	n.d.	0.13	0.08	n.d.	0.10	n.d.	n.d.
CO <sub>2</sub> (ppm)	125	n.d.	171	796	n.d.	n.d.	n.d.	n.d.
CO <sub>2</sub> error	7	n.d.	11	155	n.d.	n.d.	n.d.	n.d.
Temp (°C)	1215	1186	1178	1197	1176	1209	1224	1213
Pressure (MPa)	45	n.d.	55	151	n.d.	n.d.	n.d.	n.d.
Depth (km)	1.8	n.d.	2.2	6.1	n.d.	n.d.	n.d.	n.d.

Table 6. Continued.

Sample	br1-2- 1-2012	BLK8- 1- 2010*	br4-3- 3-2012	br1-5- 1-2012	brn14- MI-1- 2011	brn12- MI-1- 2011*	blk1- Ml-1- 2011	bk3-2- 2-2012	BUL1- 2010*
MI Location	core	int	rim	core	core	int	core	int	rim
MI Size (um)	131	60	111	42	92	123	236	57	60
Bubble	no	no	no	no	no	no	no	no	no
% PEC	0	5	1	0	0	0	0	2	0
$SiO_2$	54.31	54.38	54.47	54.92	54.93	55.03	55.05	55.15	56.04
$TiO_2$	1.55	1.35	1.15	1.77	1.70	1.76	1.75	0.65	1.53
$AI_2O_3$	12.91	16.05	16.65	13.02	13.36	12.85	13.34	16.72	15.13
FeOt	12.98	9.76	9.27	13.49	12.41	13.63	12.61	8.73	10.40
$Fe_2O_3$	2.50	1.54	1.53	2.85	2.42	2.93	2.56	1.30	1.81
FeO	10.73	8.38	7.90	10.93	10.23	10.99	10.31	7.55	8.78
MnO	0.22	0.16	0.17	0.20	0.26	0.24	0.24	0.17	0.23
MgO	5.77	5.77	5.19	4.63	5.07	4.56	4.79	5.62	4.51
CaO	7.82	8.05	8.12	7.42	7.81	7.18	7.50	8.43	7.60
$Na_2O$	3.08	3.12	4.04	3.14	3.09	3.11	3.09	3.78	3.06
$K_2O$	0.84	0.93	0.62	0.84	0.84	1.01	1.06	0.51	0.99
$P_2O_5$	0.26	0.29	0.16	0.28	0.29	0.34	0.32	0.10	0.32
Total	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00
mg#	0.44	0.51	0.50	0.38	0.42	0.37	0.40	0.53	0.44
Host Fo	73	80	80	69	73	69	72	82	75
S (ppm)	299	n.d.	n.d.	330	258	268	215	n.d.	n.d.
CI (ppm)	462	n.d.	n.d.	515	412	406	500	n.d.	n.d.
F (ppm)	478	n.d.	n.d.	606	n.d.	n.d.	n.d.	n.d.	n.d.
$H_2O$ (wt%)	1.91	2.17	n.d.	2.12	1.91	2.30	2.00	n.d.	1.69
H <sub>2</sub> O error	0.12	0.08	n.d.	0.14	0.04	0.04	0.04	n.d.	0.04
CO <sub>2</sub> (ppm)	259	1226	n.d.	61	584	n.d.	1508	n.d.	750
CO <sub>2</sub> error	21	103	n.d.	9	124	n.d.	96	n.d.	126
Temp (°C)	1225	1128	1182	1202	1201	1200	1196	1182	1175
Pressure (MPa)	108	205	n.d.	57	174	n.d.	342	n.d.	165
Depth (km)	4.4	8.4	n.d.	2.3	7.1	n.d.	13.9	n.d.	6.7

Table 7. Analytical results (range, average, and standard deviation) for melt inclusion volatiles measured with FTIR, SIMS, and EPMA.

$H_2O$ (wt%)	CO <sub>2</sub> (ppm)	S (ppm)	CI (ppm)	F (ppm)
0.36-2.94	394-1973			
$(1.8\pm0.61)$	$(939\pm473)$			
0.53-2.37	22-274	92-1277	236-641	155-606
$(1.76\pm0.44)$	$(110\pm77)$	$(639\pm290)$	(478±86)	$(322\pm89)$
		23-1534	47-1342	
		$(699 \pm 415)$	$(490\pm183)$	
	0.36-2.94 (1.8±0.61) 0.53-2.37	0.36-2.94 394-1973 (1.8±0.61) (939±473) 0.53-2.37 22-274	0.36-2.94 394-1973 (1.8±0.61) (939±473) 0.53-2.37 22-274 92-1277 (1.76±0.44) (110±77) (639±290) 23-1534	0.36-2.94 394-1973 (1.8±0.61) (939±473) 0.53-2.37 22-274 92-1277 236-641 (1.76±0.44) (110±77) (639±290) (478±86) 23-1534 47-1342

Standard deviations are 2 $\sigma$ . N equals number of analyses.

Table 8. Inputs and viscosity modeling results based on different crystallinity in the scoria. Model inputs include the average whole rock chemistry with all major elements in wt%, assumed water content of 1 wt%, and the temperature from the plagioclase melt thermometer. The reported values are for both the Shaw model and the Bottinga and Weill model.

	<u> </u>													
	Model Inputs													
SiO <sub>2</sub>	TiO <sub>2</sub>	$AI_2O_3$	Fe <sub>2</sub> O <sub>3</sub>	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	$K_2O$	$P_2O_5$	$H_2O$	T °C		
51.35							0.57	0.19	1.0	1150				
					(	8 MPa 15 MP					1Pa 30 MPa			
					Shaw	В8	W	Shaw	B&W	Sha	aw	B&W		
Black So	coria – 50	)% crysta	llinity		5.64x10 <sup>4</sup> 2.80x		x10 <sup>4</sup>	2.29x10 <sup>4</sup>	2.35x10 <sup>4</sup>	2.29	x10 <sup>4</sup>	2.35x10 <sup>4</sup>		
Brown Scoria – 10% crystallinity						$x10^{2}$	4.45x10 <sup>1</sup>	5.40x10 <sup>1</sup>	4.45	x10 <sup>1</sup>	$5.40x10^{1}$			

Model inputs include the average whole rock chemistry with all major elements in wt%, assumed water content of 1 wt%, and the temperature from the plagioclase melt thermometer. The reported values are for both the Shaw (1972) model and the Bottinga and Weill (1972) model.