

Stratospheric control of planetary waves

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Key Points:

- Coupled evolution of the stratospheric mean state and planetary waves drives half of their amplification prior to sudden warmings
- Imposing sudden-warming-like anomalies creates a downward-migrating region of local wave mean-flow interaction confined to the stratosphere
- The dependence of the equatorward shift of the tropospheric jet on the height of the imposed anomalies is quantified

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Abstract

The effects of imposing at various altitudes in the stratosphere zonally symmetric circulation anomalies associated with a stratospheric sudden warming are investigated in a mechanistic circulation model. A shift of the tropospheric jet is found even when the anomalies are imposed only above 2 hPa. Their influence is communicated downwards through the planetary wave field via three distinct mechanisms. First, a significant fraction of the amplification of the upward fluxes of wave activity prior to the central date of the warming is due to the coupled evolution of the stratospheric zonal mean state and the wave field throughout the column. Second, a downward-propagating region of localized wave, mean-flow interaction is active around the central date, but does not penetrate the tropopause. Third, there is deep, vertically synchronous suppression of upward fluxes following the central date. The magnitude of this suppression correlates with that of the tropospheric jet shift.

1 Introduction

The influence of the stratosphere on surface weather and climate is of interest not only for possible associated gains in medium-range to seasonal forecasting [*Sigmond et al.*, 2013; *Scaife et al.*, 2015] and for its role in a changing circulation resulting from changing greenhouse gases [*Manzini et al.*, 2014], but also because there are pathways to surface impacts for a variety of specific middle atmosphere forcings, including solar forcing [*Kodera and Kuroda*, 2002; *Ineson et al.*, 2011], volcanic eruptions [*Muthers et al.*, 2014], and the quasibiennial oscillation [*Gray et al.*, 2004]. The mechanisms invoked often include a connection from the forcing to the occurrence of stratospheric sudden warmings (or other dynamical behaviour of the polar vortex) and from there to surface impacts.

There is very clear evidence that forcing in the lower stratosphere does influence the tropospheric circulation. The strongest case in observations arises from the Antarctic ozone hole, which is believed to have led to an observed poleward shift of the Southern Hemisphere surface westerlies [see *Previdi and Polvani*, 2014, for a recent review]. Modeling studies have shown that imposing similar anomalies in the Arctic polar vortex leads to a surface response as well [e.g. *Douville*, 2009; *Hitchcock and Simpson*, 2014]. In the Northern Hemisphere, one of the major motivations for these studies has arisen from considering the consequences of stratospheric sudden warmings [*Baldwin and Dunkerton*, 2001]. But whether a sudden warming can be considered an ‘externally imposed’ strato-

43 spheric anomaly is not clear since warmings are driven by planetary-scale Rossby waves
44 whose sources are predominantly tropospheric.

45 The strength of surface impacts from forcings higher in the stratosphere, such as
46 solar cycle effects, is less clear. A continuing cause of confusion in this subject is the
47 relation between observed time evolution of circulation anomalies and the existence of
48 mechanisms for downward influence. The fact that anomalies in the upper stratosphere
49 are often observed to precede those at lower levels does not in itself imply that the upper
50 level anomalies are the cause of the lower level anomalies. This possibility was clearly
51 demonstrated by *Plumb and Semeniuk* [2003] in an idealized model of polar stratospheric
52 variability [*Holton and Mass*, 1976], where constraining the mean state above a fixed level
53 led to little impact on the evolution of the flow below. Similarly, the descent of the strato-
54 spheric cold anomaly which follows a subset of major sudden warmings known as Polar-
55 night Jet Oscillation (PJO) events can largely be explained by the vertical gradient in ra-
56 diative timescales, again requiring no downward influence [*Hitchcock et al.*, 2013a]. Al-
57 though significant downward influence from high-altitude solar effects has been argued in
58 several cases [*Gray*, 2003; *Ineson et al.*, 2011], the strength of this influence and the rele-
59 vant mechanisms remain unclear not least due to the strongly chaotic evolution of both the
60 tropospheric jet and the stratospheric vortex.

61 Downward influence within the stratosphere is thought to arise through two types of
62 pathways [*Plumb and Semeniuk*, 2003; *Hardiman and Haynes*, 2008]. The first is through
63 the zonally symmetric circulations associated with the maintenance of a balanced state;
64 these depend only weakly on the zonal mean state itself, and the effects decay exponen-
65 tially with distance through which the downward influence extends [e.g. *Haynes et al.*,
66 1991].

67 The second broad class relies on interactions between planetary-scale Rossby waves
68 and the zonal mean stratospheric state, and depends strongly upon the latter. These can
69 be local in character; one commonly invoked mechanism [*Matsuno*, 1971; *Kodera and*
70 *Kuroda*, 2002; *Ineson et al.*, 2011] involves the presence of a critical line for quasi-
71 stationary waves, or more generally a layer where winds are weak, below which there is
72 strong dissipation of the waves. If this absorption is sufficiently local and coherent, this
73 leads to deceleration of the zonal mean winds below the existing anomaly and thus de-
74 scent of the region of absorption. However, it was argued by *Plumb and Semeniuk* [2003]

75 that a mechanism of this type does not necessarily lead to downward influence, in the
76 sense of downward propagation of information.

77 Non-local influence by the waves can arise in the presence of reflection [*Perlwitz*
78 *and Harnik, 2004; Shaw and Perlwitz, 2013*] or resonances [*Plumb, 1981; Matthewman*
79 *and Esler, 2011; Albers and Birner, 2014*], note however that the two mechanisms are not
80 exclusive of each other. These mechanisms imply a significant degree of stratospheric
81 control over the quasi-stationary waves throughout the depth of the atmosphere. While
82 such control can be clearly demonstrated in highly idealized contexts [*Coughlin and Tung,*
83 *2005*], the relevance of this type of downward influence in real stratospheric sudden warm-
84 ings remains an issue of current debate [*Albers and Birner, 2014*].

85 Since the troposphere-stratosphere system is highly chaotic, each of these mecha-
86 nisms may appear to be relevant for particular initial conditions. It is therefore essential
87 to quantify their ‘deterministic’ effects, which survive averaging over some non-trivial en-
88 semble. We present in this paper a series of numerical experiments which demonstrate a
89 deterministic response to stratospheric forcing and the dependence of this response on the
90 height above which the forcing is applied.

91 We briefly present the model setup in section 2, providing more complete details in
92 the appendix. Results are given in section 3, and conclusions are presented in section 4.

93 **2 Model Setup**

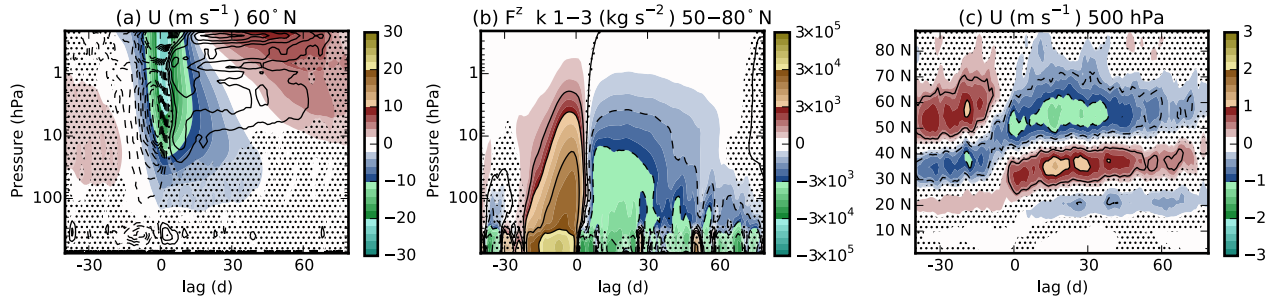
94 The numerical experiments are carried out with a version of the Reading Intermedi-
95 ate General Circulation Model (IGCM), a dry dynamical core. The model configuration is
96 based on a modified setup of *Polvani and Kushner [2002]*, appropriate for a perpetual mid-
97 winter configuration. Details of the numerics and relaxation temperature profile are given
98 in the appendix.

99 To produce a stationary wave field, surface topography is specified as a mountain,
100 Gaussian in latitude ϕ and longitude λ , centered in the Northern Hemisphere:

$$101 \quad \Phi_s = gh_0 \exp \left(- \left(\frac{\phi - \phi_0}{\Delta\phi} \right)^2 - \left(\frac{\lambda}{\Delta\lambda} \right)^2 \right). \quad (1)$$

102 The height h of the mountain is 3 km, centered at $\phi_0 = 45^\circ$ N with $\Delta\phi = \Delta\lambda = 15^\circ$. The
103 stationary wave field has a stronger component of zonal wave number one than two.

104 The base run is integrated for 100,000 days with the first 10,000 days discarded to
 105 remove the influence of an initial period of transient behaviour in the tropics. In the re-
 106 maining 90,000 days, the model produces 465 stratospheric sudden warmings defined in
 107 terms of the reversal of the zonal mean eastward winds at 60° N and 10 hPa, discarding
 108 those reversals not preceded by 20 days of eastward winds.



109 **Figure 1.** Composite anomalies from 465 sudden warmings in the base run. (a) Zonal mean zonal wind at
 110 60° N (colors) and acceleration due to the convergence of EP flux due to planetary-scale eddies averaged from
 111 40°-80° N (contours, interval $0.5 \text{ m s}^{-1} \text{ d}^{-1}$). (b) Vertical component of the EP flux due to planetary-scale
 112 eddies averaged from 50°-80° N. (c) Zonal mean zonal wind at 500 hPa. Stippling indicates regions where the
 113 composite mean differs from zero at the 95% confidence interval as estimated by a t-test.

114 A composite over these events is constructed relative to the date of this wind reversal,
 115 which we define as the central date. The evolution of this composite, as an anomaly
 116 from the time mean, is shown in Fig. 1. Anomalous westward winds arise first in the up-
 117 per stratosphere, about ten days prior to the central date. Just prior to the central date,
 118 the 2 m s^{-1} contour reaches 100 hPa. The lower stratospheric wind anomalies persist for
 119 about 45 days, with more rapid recovery to anomalous eastward winds in the upper strato-
 120 sphere (Fig. 1a). Consistent with the standard understanding of the dynamics of these
 121 events, the wind anomalies are forced by the angular momentum transported by planetary-
 122 scale Rossby waves, as measured by the Eliassen-Palm (EP) flux due to zonal wave num-
 123 bers 1 to 3 [computed following *Andrews et al.*, 1987, from daily instantaneous output].
 124 The upward wave fluxes amplify over two weeks prior to the wind reversal (Fig. 1b), and
 125 are subsequently suppressed until nearly 60 days after the wind reversal, consistent with
 126 composites of similar events in reanalyses and comprehensive models [*Hitchcock et al.*,
 127 2013b]. The anomalous divergence of these fluxes is shown in Fig. 1a, revealing a deep
 128 region of strong convergence prior to the wind reversal, followed by anomalous diver-

144

Table 1. Summary of model integrations

Run	Time (days)	p_b	p_t
base	100,000		
<i>c1</i>	37,000	2 hPa	0.8 hPa
<i>c8</i>	37,000	10 hPa	6 hPa
<i>c30</i>	37,000	40 hPa	20 hPa
<i>c70</i>	37,000	90 hPa	50 hPa
<i>s1</i>	740×160	2 hPa	0.8 hPa
<i>s8</i>	740×160	10 hPa	6 hPa
<i>s30</i>	740×160	40 hPa	20 hPa
<i>s70</i>	740×160	90 hPa	50 hPa
<i>m30</i>	465×80	90 hPa	50 hPa
<i>m20</i>	465×80	90 hPa	50 hPa

129 gence corresponding to the suppressed vertical fluxes. To a large degree these anomalies
 130 are explained by the vertical derivative of the vertical flux. During the recovery phase of
 131 the stratospheric event the tropospheric jet shifts equatorward (Fig. 1c). Significant wind
 132 anomalies are seen prior to the wind reversal in both the troposphere and stratosphere.

133 A series of further ensembles of integrations are then carried out following the method-
 134 ology of *Hitchcock and Simpson* [2014, henceforth HS14]. For each ensemble, a control
 135 integration (*c1*, *c8*, *c30*, *c70*) is first carried out. This is achieved by relaxing the zonally
 136 symmetric component of the circulation towards the time-averaged state of the base run
 137 (X_c , where X denotes the divergence, vorticity, or temperature). The rate of the relaxation
 138 varies linearly ($q = 1$) from $\tau_0 = 6$ h above p_t to zero below p_b :

$$139 \quad K(p) = \begin{cases} \tau_0^{-1} & \text{if } p < p_t, \\ \tau_0^{-1} \left(\frac{p-p_b}{p_t-p_b} \right)^q & \text{if } p_t < p < p_b, \\ 0 & \text{if } p > p_b. \end{cases} \quad (2)$$

140
 141 The height at which this relaxation is performed is varied by setting p_b and p_t according
 142 to Table 1. The region above p_b is referred to as the nudging layer. The effects of this
 143 nudging are discussed further in the supplementary material.

145 The effects of the zonally-symmetric anomalies associated with the composite sud-
 146 den warming are then determined in a further set of integrations. The ensembles (*s1*, *s8*,
 147 *s30*, *s70*) consist of a set of 740 ‘nudged’ integrations, initialized from the corresponding
 148 control integration at intervals of 50 days; the large ensemble size was found to be neces-
 149 sary to achieve a statistically robust signal, particularly in *s1*. Each integration is carried
 150 out for 160 days, and is nudged by relaxing the circulation according to (2) towards the
 151 time-evolving composite of the sudden warmings (Fig. 1). The composite values are de-
 152 noted $X_s(t)$, where t is defined relative to the central date of the sudden warming. The
 153 reference state X_r to which the circulation is relaxed is defined by interpolating smoothly
 154 over 10 days (t_0) from the climatological mean to the time-varying composite, starting 40
 155 days prior to the central date ($t_s = -40$ d):

$$156 \quad X_r = X_c + r(t, t_0)(X_s(t) - X_c), \quad (3)$$

$$157 \quad r(t, t_0) = \begin{cases} \sin^2(\frac{\pi}{2}(t - t_s)/t_0) & \text{if } t - t_s < t_0, \\ 1 & \text{otherwise.} \end{cases} \quad (4)$$

159 The composite anomalies starting about a month prior to the wind reversal at 10° N
 160 60 hPa are thus imposed in each ensemble member through the nudging defined by (2).
 161 Since each ensemble member is initialized with an essentially random initial condition
 162 drawn from the control run, any signal in the ensemble average relative to the control
 163 integration can be interpreted as a ‘deterministic’ response to the imposed stratospheric
 164 anomalies, independent of the initial conditions. The responses shown below are com-
 165 puted by differencing the nudged integrations (comprising, e.g. *s70*) from the correspond-
 166 ing time period in the control run (e.g. *c70*).

167 3 Results

172 Figure 2 shows the anomalies for each ensemble of the same quantities shown for
 173 the base run in Fig. 1. The nudging layer is indicated in the first and second column of
 174 panels by horizontal dashed (p_b) and solid (p_t) lines. In all cases the high-latitude wind
 175 response is reproduced to a good approximation within the nudging layer. When the nudg-
 176 ing is imposed higher in the stratosphere (*s1*, *s8*; Figs. 2a,b), wind anomalies are also
 177 produced one to two scale heights below the level of the nudging. Accompanying these
 178 anomalies is a region of EP-flux convergence which descends over time, following the
 179 wind anomalies, consistent with the local wave-mean flow interaction mechanism pro-

180 posed by *Matsuno* [1971]. A small region of convergence is also apparent in Fig. 2c, but
181 does not descend below 100 hPa, and no such feature is seen in Fig. 2d, suggesting this
182 mechanism is only active within the stratosphere. Note that there are artifacts in the flux
183 divergence near the lower level of the nudging layers when these are imposed at higher al-
184 titudes; these can be reduced by further smoothing the profile of nudging rates and do not
185 affect our conclusions (see Fig. S3 and discussion in supplementary material).

186 The suppression of vertical EP fluxes seen during the recovery phase of the events
187 in Fig. 1b and in the nudging experiments of HS14 is also reproduced here (Figs. 2e-h).
188 The reduction is strongest in *s70* (Fig. 2h), comparable to that seen in the base run, and
189 weakens with the altitude of the imposed anomalies. Significant suppression is still ob-
190 tained even in *s1* when the anomalies are imposed near the stratopause. For the week or
191 two around the central date of the imposed warming, particularly in the cases *s8* and *s30*,
192 the negative flux anomalies emerge at successively lower altitudes, corresponding to the
193 descending region of absorption. This indicates at least some of these anomalies arise due
194 to filtering by the mean flow. However, the anomalies in *s1* (Fig. 2e) arise synchronously
195 throughout the depth of the stratosphere at layers with no strong mean flow anomalies.
196 This suggests that simple ideas of filtering or reflection of propagating modes (for which
197 flux anomalies would be expected to propagate with a bounded vertical group speed) can-
198 not explain all of these anomalies.

199 In all cases anomalous fluxes arise within the troposphere below 300hPa. How-
200 ever the magnitude of the anomaly averaged over the timescale for recovery for the vor-
201 tex is small compared with the time variation that appears in the ensemble mean shown in
202 Figs. 2e-h, and while this time variation seems likely to be internal variability rather than
203 a systematic signal, simple estimates suggest that to verify this convincingly a much larger
204 ensemble would be required. Constraining the details of the time-dependence of the flux
205 anomalies with the troposphere therefore remains a significant challenge.

206 Figures 2g-h also show that significant enhancement of the upward fluxes are also
207 obtained during the 30 days prior to the central date in *s70* and *s30*, indicating a role for
208 the stratospheric mean state in the amplification of waves responsible for sudden warm-
209 ings. In contrast to the flux anomalies after the central date in *s70* and *s30* that are similar
210 in magnitude to the base run composite, the flux anomalies prior to the central date are
211 a factor of 10 weaker to those in the corresponding period in Fig. 1b. This enhancement

212 was not obtained by HS14, likely because anomalies were imposed only at the time of the
 213 stratospheric wind reversal.

214 On first sight this result appears to suggest that the stratospheric influence is only
 215 responsible for about 10% of the amplification of the waves prior to the central date. It
 216 does firmly establish that the troposphere must be in a favourable state in order for the full
 217 amplification to occur, since if the wave amplification was independent of the tropospheric
 218 state, it would have been fully recovered in the nudged ensemble. However, this does not
 219 quantify the role of the stratospheric state in the amplification of the waves when tropo-
 220 sphere is in such a favourable state.

221 To investigate this role further, an additional pair of ensembles, *m30* and *m20*, are
 222 also produced. These are 80-day integrations initialized from the base run 30 and 20 days
 223 (respectively) prior to the central date of each of the warmings composited in Fig. 1. They
 224 are relaxed towards the time average of the base run (X_c) with the same profile of relax-
 225 ation rates used by *s70*. The relaxation is switched on smoothly using $r(t, 5d)$. By pre-
 226 venting the stratospheric mean state from evolving with the amplifying waves, these en-
 227 sembles test for the role of the stratospheric mean state in this amplification.

234 Figures 3a,b show the difference in the same fluxes shown in Figs.2e-h between the
 235 integrations in *m30* and *m20* and the corresponding periods in the base run, respectively.
 236 When the stratospheric mean state is prevented from evolving, this amplification is re-
 237 duced. When the nudging is switched on 30 days prior to the central date (*m30*), the am-
 238 plification is weakened by 50% relative to that seen in the base run (Fig. 1b). The strato-
 239 spheric constraint must be imposed sufficiently early in the evolution, however; in *m20*
 240 where the nudging is switched on 20 days prior to the central date relatively weaker sup-
 241 pression of the amplification is seen only in the upper troposphere. We have confirmed
 242 that the full amplification seen in Fig. 1b is not a result of chaotic error growth but is in
 243 fact due to the stratospheric constraint (see Fig. S4).

244 The stratospheric mean state does therefore play a significant role in the ampli-
 245 fication of the waves responsible for the breakdown of the vortex in these simulations,
 246 as suggested by the theories of *Plumb* [1981] and *Matthewman and Esler* [2011] and
 247 demonstrated by the model experiments of *Scott and Polvani* [2004]. On the other hand
 248 it is clear from Figs 2g,h that the troposphere must also be in a favourable state for the

249 amplification of the waves to occur. Characterizing the nature of these tropospheric states
250 in detail is beyond the scope of this work.

251 Returning to Fig. 2, panels (i-l) show the tropospheric wind anomalies in each en-
252 semble. An equatorward shift of the jet persisting until nearly 60 days following the cen-
253 tral date is obtained in all cases. The magnitude of the anomalies decrease with the height
254 of the imposed anomalies, with *s70* producing anomalies nearly as strong as those in the
255 base run.

256 Recent work [HS14, *Smith and Scott, 2016; Hitchcock and Simpson, 2016*] has high-
257 lighted the importance of the planetary-scale wave field for communicating the influence
258 of the stratospheric anomalies to the tropospheric jet during the recovery phase of strato-
259 spheric sudden warmings. Figure 3c shows the time-averaged upper tropospheric zonal
260 wind anomalies (from 30° to 40° N) and vertical EP flux anomalies in each ensemble.
261 These quantities are proportional in the forced response. The present experimental design
262 cannot directly attribute the suppressed vertical wave fluxes to the imposed stratospheric
263 anomalies (as opposed to being determined by the evolution of the tropospheric flow);
264 however, as was found by HS14 in a comprehensive model, the vertical EP fluxes are not
265 correlated with the tropospheric jet variability in the base and control runs suggesting the
266 suppressed fluxes determine the jet response, not the reverse.

267 4 Conclusions

268 We have demonstrated that anomalies associated with stratospheric sudden warm-
269 ings, even when imposed in the upper stratosphere (*s1*) within a layer representing only
270 0.2% of the mass of the atmosphere, can impart a significant, robust impact on the waves
271 and mean flow below, in both the stratosphere and troposphere. This has been achieved
272 with a mechanistic circulation model through a set of numerical experiments that identifies
273 a deterministic impact by averaging across a large ensemble of integrations, each with a
274 different tropospheric initial condition. The experiments clearly reveal how the response to
275 imposed zonally-symmetric stratospheric anomalies vary with the height (p_b) above which
276 they are imposed.

277 In all cases we find that the influence of the anomalies extends well below p_b . In
278 cases where p_b lies well above the tropopause, there is a clear, localized region of wave,
279 mean-flow interaction which emerges below the region of nudging when the imposed

280 westward anomalies are strongest, producing a descent of the westward anomalies (Fig. 2a-
281 d). This signal weakens in cases where p_b lies closer to the tropopause, and in no case
282 does it penetrate the tropopause, suggesting that the mechanisms for downward influence
283 within the stratosphere are distinct from those responsible for the suppression of upward
284 wave flux within the troposphere after the warming. One reason this mechanism may be
285 restricted to the stratosphere is the presence of the strong wave guide at the edge of the
286 polar vortex; another reason may be that the coherence of the signal is lost in the presence
287 of strong tropospheric variability.

288 There is some similarity to the mechanism described by *Matsuno* [1971] for the evo-
289 lution of sudden warmings including the downward migration of wind anomalies but it is
290 important to note that our experiments have established that there is genuine downward
291 propagation of information. This contrasts with the *Plumb and Semeniuk* [2003] charac-
292 terisation of the *Matsuno* [1971] mechanism as similar to the *Plumb* [1977] model of the
293 equatorial quasibiennial oscillation, in which there is no downward propagation of infor-
294 mation. Note also that the downward propagation of the zonal flow (and flux divergence)
295 anomalies seen at the onset of the event in Fig. 2a,b is much slower than the downward
296 migration seen in the base run composite (Fig. 1a). Thus, rather as is the case for the role
297 of the stratospheric flow in enhancing upward wave fluxes prior to the warming (see dis-
298 cussion below), the mechanism responsible for the downward propagation may be an im-
299 portant part of the evolution of sudden warmings, but it must be accompanied by other
300 physical effects.

301 The imposed anomalies suppress vertical fluxes of wave activity throughout the
302 depth of the atmosphere during the recovery phase of the imposed warmings (Fig. 2e-
303 h), in agreement with the results of HS14. When the anomalies are imposed in the lower
304 stratosphere, the flux anomalies are as large as those found in the free running integration
305 (Fig. 1b). The flux anomalies weaken as p_b is reduced. When the circulation anomalies
306 are imposed in the middle or upper stratosphere, the flux anomalies arise nearly simul-
307 taneously throughout the depth of the stratosphere, suggesting the possible relevance of
308 barotropic modes for this coupling.

309 In all cases an equatorward shift of the tropospheric jet is obtained over much of the
310 recovery period of the imposed warming. The structure of the wind anomalies are only

311 weakly dependent on p_b , but their magnitude reduces as p_b reduces, and correlates with
312 the tropospheric upward wave flux anomalies.

313 When the anomalies are imposed lower in the stratosphere, significant enhancement
314 of the vertical wave fluxes are found prior to the central date of the imposed warming.
315 The amplification in the nudged ensembles is only of the order of 10% of that in the base
316 run composite. However, when the stratospheric zonal mean is constrained to its time
317 mean state sufficiently early during the onset of the warming, the amplification of the
318 waves is found to be reduced by about 50%. This provides strong and novel evidence in
319 a full primitive-equations model for the coupled evolution of waves and the mean state
320 during the onset of a warming, expected, for example, from the ideas of resonant amplifi-
321 cation [Plumb, 1981; Matthewman and Esler, 2011].

322 This constitutes an important asymmetry in the response, in the sense that impos-
323 ing the stratospheric anomalies prior to the central date only recovers a fraction of the
324 enhanced upward fluxes of wave activity, while the imposed stratospheric anomalies dur-
325 ing the recovery phase are sufficient to produce the full suppression. While the onset of
326 the warmings seem therefore to require appropriate configurations of both the stratosphere
327 and troposphere, the post-warming evolution seems only to require the configuration of the
328 stratospheric state.

329 These experiments reveal a substantial influence on the tropospheric circulation by
330 the full depth of the stratosphere, indicating clear potential for stratospheric forcings to
331 impact on the surface through the polar vortex. They reveal a variety of distinct mecha-
332 nisms by which the zonal mean flow and the planetary waves interact to communicate this
333 influence, highlighting in particular the potential for the stratospheric state to affect the
334 evolution of the waves over a deep region of the atmosphere.

335 **A: Temperature Relaxation Profile**

336 Numerical integrations are performed using a modified version of the Reading In-
337 termediate General Circulation Model (IGCM), version 1. The code integrates the dry
338 hydrostatic primitive equations on the sphere Hoskins and Simmons [1975] and has been
339 modified to use the angular-momentum conserving vertical discretization of Simmons and
340 Burridge [1981] on hybridized pressure levels. The model climate is determined by a lin-
341 ear relaxation towards an equilibrium temperature profile that is convectively stable but

342 baroclinically unstable [*Held and Suarez, 1994*]. All reference quantities below are defined
 343 on model levels (p is determined by setting $p_s = p_0$ with $p_0 = 1000$ hPa).

344 The radiative equilibrium temperature profile follows *Polvani and Kushner [2002]*,
 345 with several modifications; all notation below follows their definitions. The stratospheric
 346 profile is specified by

$$347 \quad T_{eq}^{strat} = T'_{US}(p) + W(\phi)T'_{PV}(p) \quad (\text{A.1})$$

348 in which the meridional weighting function W is the same as that used by *Polvani and*
 349 *Kushner [2002]* but with the vortex in the Northern Hemisphere. The polar vortex profile
 350 is specified by

$$351 \quad T'_{PV}(p) = T_T \left(\left(\frac{p}{p_T} \right)^{-R\gamma/g} - 1 \right) \quad (\text{A.2})$$

352 and is lowered by setting p_T to 300 hPa. The US Standard Atmosphere used outside the
 353 polar region is modified by reducing the temperature everywhere ($T'_{US} = T_{US} - 16.65$ K)
 354 so that $T'_{US}(p_T)$ is equal to 200 K and the tropopause in the equilibrium profile occurs at
 355 pressure levels closer to the Earth's tropopause. The hemispheric asymmetry parameter ϵ
 356 used by *Polvani and Kushner [2002]* is set to 0 K.

357 The profile of radiative damping timescales (above the boundary layer) is set to

$$358 \quad \alpha = \alpha_T + \frac{1}{2} \left(\tanh \left(\frac{z - z_s}{\sigma_z} \right) + 1 \right) (\alpha_S - \alpha_T) \quad (\text{A.3})$$

359 with $\alpha_T^{-1} = 40$ d and $\alpha_S^{-1} = 5$ d. The log-pressure height z is set to $H \log(p/p_0)$ where H
 360 = 7 km, and finally $z_s = 35$ km and $\sigma_z = 7$ km.

361 **Acknowledgments**

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363 The data from all model integrations are available from the authors upon request.

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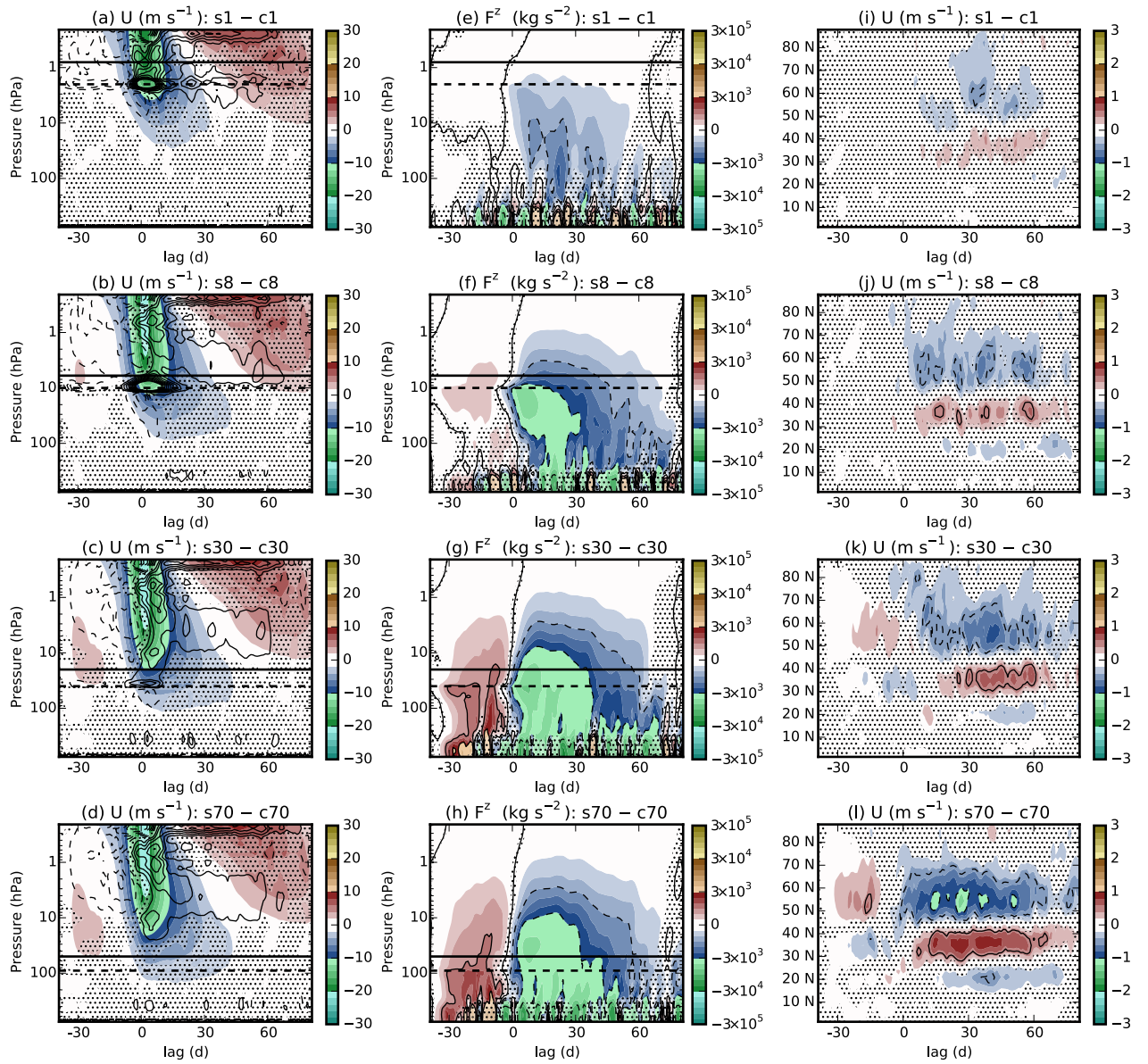
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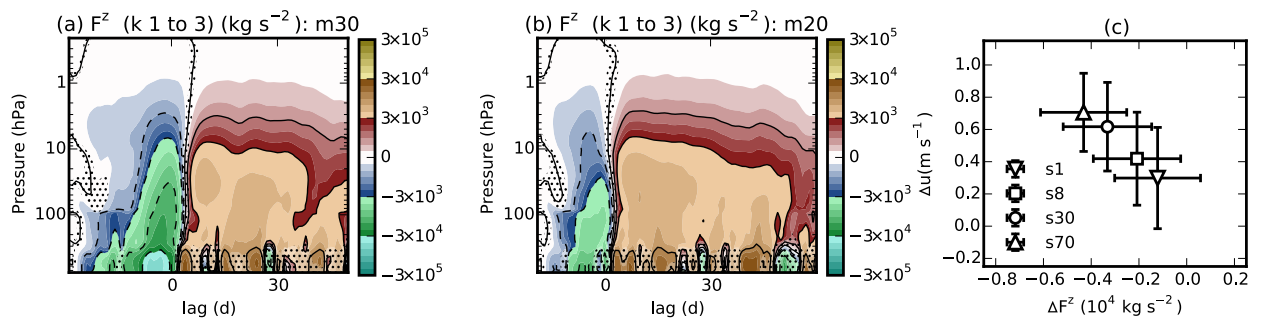
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168 **Figure 2.** Ensemble mean anomalies of $s1$, $s8$, $s30$, $s70$ relative to their respective controls for the same
 169 quantities shown in Fig. 1. The horizontal lines in panels (a-h) indicate the level at which the nudging is
 170 zero (dashed) and full strength (solid). Statistical significance is indicated as in Fig. 1, estimated using a
 171 paired-sample t-test.



228 **Figure 3.** Ensemble mean anomalies from (a) *m30* and (b) *m20* of the vertical component of the EP flux
 229 due to planetary-scale waves, as an anomaly from the corresponding periods in the base run. Statistical signif-
 230 icance is computed and indicated as in Fig. 2. (c) Time averaged (days 30-60) zonal mean wind averaged over
 231 30° - 40° N and 500 hPa to 200 hPa plotted against vertical EP flux due to planetary-scale eddies averaged over
 232 50° - 80° N and 500 hPa to 200 hPa from *s1*, *s8*, *s30*, *s70*. Confidence intervals at the 95% level are indicated
 233 for each quantity.

Supporting Information for “Stratospheric control of planetary waves”

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Contents

1. Text S1 to S5
2. Figures S1 to S5

Introduction

Text S1.

Supplementary Figure 1 shows the climatology (time mean) of the zonal mean, zonal winds in each of the control runs *c1*, *c8*, *c30*, and *c70*. The differences between these climatologies and the base run climatology are shown by the contour lines with an interval of 0.5 m s^{-1} . This demonstrates that the effect of nudging the zonally symmetric component of the stratosphere to the climatological state of the base run has a minimal effect on the zonal mean basic state. The impact of the stratospheric nudging on the Northern Hemisphere troposphere amounts to less than a 0.25 m s^{-1} change in the tropospheric winds in all cases but *c70*, where a dipolar anomaly of approximately 0.5 m s^{-1} corresponding roughly to a poleward shift of the jet. The changes are small relative to the internal variability of the winds; it is therefore unlikely that these changes will have a significant effect on the response of the system to the imposed anomalies. The fact that the tropospheric response seen in Fig. 2l closely resembles the composite response shown in Fig. 1c also confirms this claim.

Text S2.

Supplementary Figure 2a shows the vertical profile of the mid-latitude average of the standard deviation of the zonal mean zonal wind in the base run and in each of the con-

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control runs, plus one additional control run, *c100*, with $p_b = 200$ hPa and $p_t = 80$ hPa. The boundary of the nudging layer for each case is indicated by the colored lines. Within the nudging layer where the relaxation is at its full strength, the internal variability is strongly suppressed; the variability is reduced to a lesser degree even below p_b . To some extent this need not indicate anything artificial - if variability in the stratosphere is driving some component of variability in the tropospheric flow as our experiments have demonstrated, eliminating the stratospheric variability should remove this component from the tropospheric variability as well. Nonetheless this figure demonstrates that the nudging layer cannot be moved much below 90 hPa without substantially constraining the tropospheric flow. Figure S3b shows the same quantity but computed as the ensemble spread over the nudged runs relative to their respective control runs, averaged over days 15 to 60 following the central date. This spread agrees closely with the internal variability.

Text S3.

Supplementary Figure 3 shows plots equivalent to those in Fig. 2 but for two additional ensembles, *s8w6h* and *s8w1d*, nudged with alternative profiles of the relaxational timescale, specified using different values for the parameters in (2). In both cases $q = 4$, $p_b = 10$ hPa and $p_t = 3$ hPa. The first, *s8w6h*, is an ensemble of 600 integrations with $\tau_1 = 6$ h (Fig. S2a,c,e), while second, *s8w1d*, is an ensemble of 400 integrations with $\tau_1 = 1$ d (Fig. S2b,d,f). Differences in both cases are taken from a control run relaxed to X_c with the corresponding nudging profile, but in the latter case the differences are taken from the time mean the control run, so they do not vanish at the onset of the integrations. These ensembles may be seen as corresponding to a profile intermediate between *s1* and *s8*, though we discuss these plots relative to the latter.

In both cases the EP flux convergence near p_b prior to the central date seen in Fig. 2b is no longer present. The strong anomalous divergence around the central date in Fig. 2b is reduced in these ensembles, and is weaker in *s8w1d*, consistent with the weaker vertical shear induced by the nudging. The descending region of anomalous convergence is still present, though again is somewhat weaker than in Fig. 2b. Nonetheless, the weakened vertical fluxes throughout the depth of the stratosphere, and the shift of the tropospheric jet seen in Fig. 2 are also present in these alternative ensembles; the mean jet shift in both cases is consistent with the uncertainty shown in Fig. 3c.

We conclude from these further integrations that the EP flux artifacts near p_b are not playing a significant role in the evolution of the flow below the nudging layer.

Text S4.

Supplementary Figure 4 demonstrates that the amplification of the waves prior to stratospheric sudden warmings is fully recovered when the base run is restarted 30 (*m30c*) or 20 (*m20c*) days prior to the events. The panels show the evolution of the vertical EP flux in ensembles equivalent to *m30* and *m20* but with no stratospheric constraint, an as an anomaly from the evolution of the base run over the same period. Because the restarts are based on instantaneous output at a single timestep, the full information required by the leap-frog timestep used by the model to reproduce bit-for-bit evolution of the runs is not available, and this leads ultimately to diverging trajectories. However, this error growth only becomes significant well after the onset of the stratospheric event. This confirms that the effect demonstrated in Fig. 3a,b is in fact due to constrained stratospheric winds, not due to chaotic error growth.

Text S5.

Supplementary Figure 5 shows the effects of modifying the profile of the linear relaxation on the suppression of the wave amplification shown in Figs. 3ab. In each case an ensemble similar to *m30* or *m20* has been carried out. Panels a and b correspond directly to Figs. 3 a and b but for the nudging profile in (2) modified by setting $\tau_0 = 1$ d and $q = 4$; p_b is set to 90 hPa and p_t to 30 hPa. The weaker nudging strength has the expected effect of allowing for more amplification of the wave fluxes. Panel c shows an ensemble equivalent to *m30* but with $\tau_0 = 1$ d and $q = 4$; p_b set to 200 hPa and p_t to 80 hPa. Constraining the flow lower in the atmosphere has the effect of reducing the amplification of the wave fluxes.

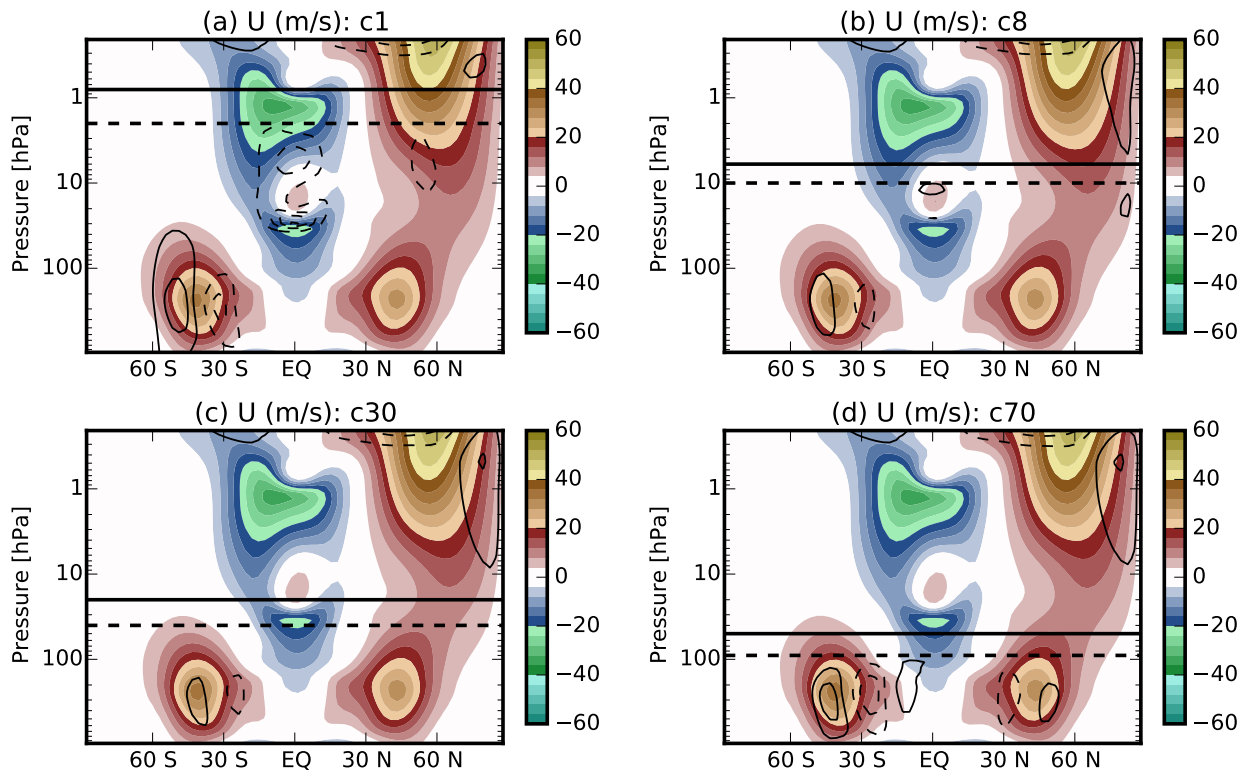


Figure 1. The colored contours show climatological (time mean) zonal mean zonal winds from each of the four control runs. Differences between these climatologies and that of the base run are indicated by the contour lines, shown at intervals of 0.5 m s^{-1} . The zero contour is omitted. The horizontal lines indicate the nudging layer in each control run as in Fig. 2.

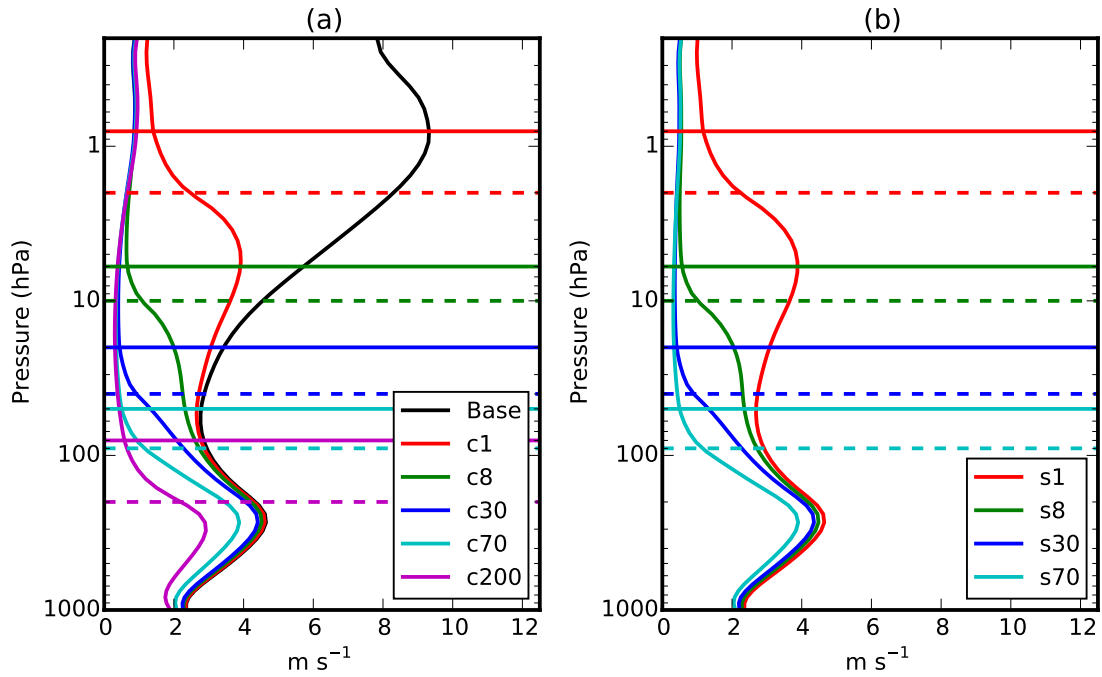


Figure 2. Vertical profile of the mid-latitude (30-60° N) average standard deviation of zonal mean zonal wind for (a) the internal variability of the control runs and (b) the ensemble spread of the nudged runs, relative to their respective controls. The internal variability of the base run is also shown in (a). For the nudged runs, the corresponding colored horizontal lines indicate the nudging layer in each control run as in Fig. 2.

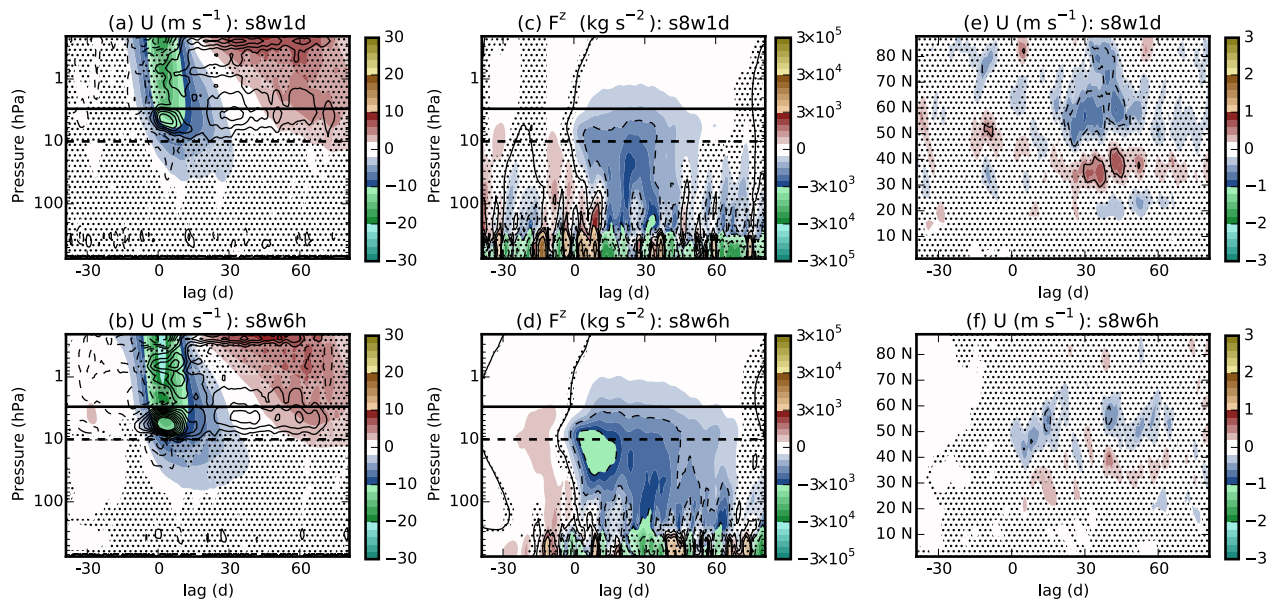


Figure 3. Equivalent to Fig. 2b,f,i but using two alternate profiles of relaxation timescales. See supplementary text S3 for full description.

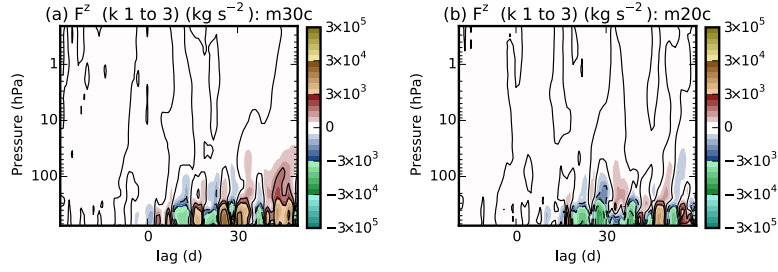


Figure 4. Evolution of the vertical EP flux in a control ensemble initialized from the base run (a) 30 days prior and (b) 20 days prior to the first four hundred sudden warming events. The fluxes are shown as an anomaly relative to the base run over the same periods.

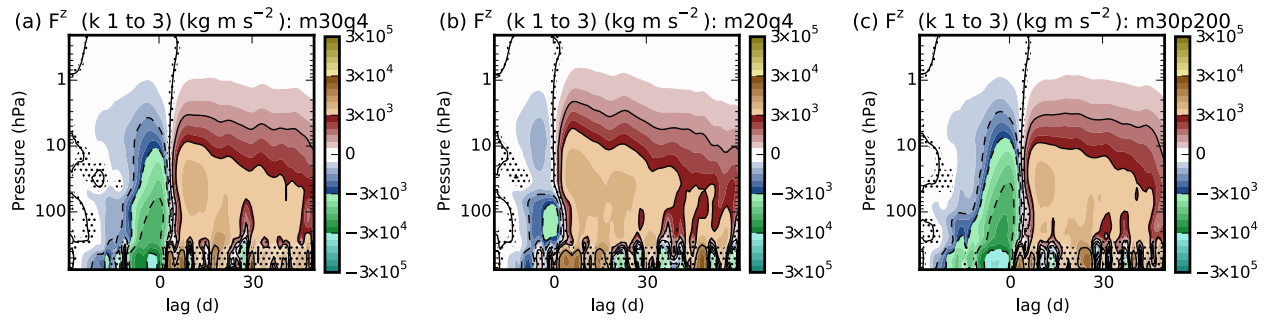


Figure 5. Equivalent to Fig. 3a,b, but for alternative nudging configurations. See text S5 for full description.