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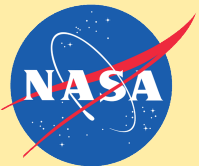
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Stable Isotope Chemistry in Titan Haze Aerosol

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Introduction and Motivation

Why study Titan?

- Titan is the **largest satellite of Saturn** and has atmosphere composed of nitrogen and a few percent methane (CH₄).
- The **atmosphere is believed to be similar to that of early Earth**.
- The **haze layer of Titan is rich in organic chemistry** and can give new insights into **prebiotic chemistry and planetary habitability**.

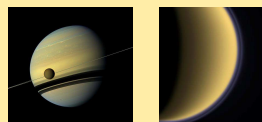


Figure 1. Images from the Cassini-Huygens mission. Titan is shown in orbit around Saturn (left) and the haze layer of Titan's atmosphere is observed as a hazy halo around the planet. Images courtesy of JPL-NASA

Why study isotopes on Titan?

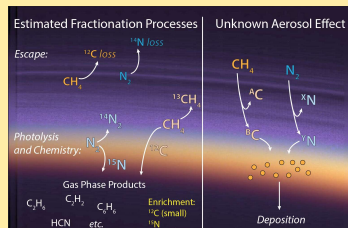
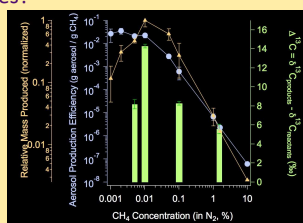


Figure 2. Known and unknown pathways to explain the isotopic fractionation occurring in Titan atmosphere.

- The measurements of stable isotope ratios give information on the **history and evolution of the atmosphere**.
- Measurements from Titan indicate that the ¹²C/¹³C ratio in CH₄ is similar to the protosolar ratio, which suggests that the **CH₄ is relatively young**.¹
- The ¹⁴N/¹⁵N is similar to the **protosolar value of NH₃** based on comet measurements.^{2,5}

Are aerosols a sink for stable isotopes?

- Current models of the observed isotope ratios on Titan do not incorporate isotopic fractionation resulting from organic aerosol formation and subsequent deposition onto the surface of Titan** (Figure 2).
- Initial studies have shown that **fractionation direction and magnitude are dependent on the initial bulk composition of the gas mixture**³ (Figure 3).



Why study aromatic compounds?

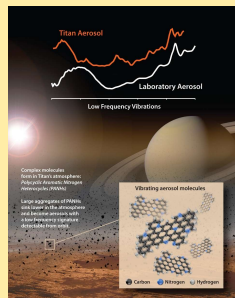


Figure 4. Far-infrared spectra from the Cassini spacecraft as compared to the spectra of laboratory aerosols produced from aromatic compounds. Image courtesy of JPL-NASA.

- Cassini-borne instruments **have detected benzene in Titan's atmosphere**.
- Aromatics, such as benzene (below, left), has been shown to be an **important pathway in aerosol formation**.
- Far-IR spectral feature of Titan's haze layer is similar to that of aerosols produced from aromatic compounds.³ (Figure 4)
- Though not observed in situ, pyridine (N-containing aromatic, below) is a likely product of Titan chemistry and produces laboratory aerosol with a strong Far-IR feature.

Materials and Methods

Aerosol Production

- Titan aerosol analogs are produced in the laboratory to study their fractionation.**
- Gas mixtures used are trace gases (CH₄, benzene (C₆H₆), and pyridine (C₅H₅N) with nitrogen (N₂)) in mixing chamber as shown in Figure 6.
- The **gas mixture is irradiated with far-UV light (115-400 nm)** that leads to aerosol production. a quartz filter (Figure 7).
- Aerosol samples are collected in an inert, ex situ environment (Ar, N₂ or vacuum) and processed for Isotope Ratio Mass Spectrometry (IRMS) Analysis

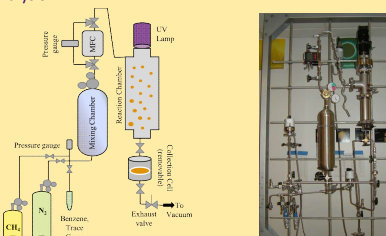


Figure 6. Left: Setup for generating Titan aerosol analogs. Right: Actual system for generating the aerosols. Note that the tanks of the reactant gases (N₂ and CH₄) are not shown in the picture.



Figure 7. Image of laboratory produced aerosol before IRMS processing

Isotope Ratio Mass Spectrometry (IRMS)

- IRMS is used to measure the relative abundance of isotopes in a given sample.**
 - The sample is combusted and is converted into CO₂ and N₂.
 - Carbon and nitrogen stable isotope values are reported in standard δ notation in per mil (‰) as defined by:
- $$\delta(\text{‰}) = \left[\left(\frac{R_{\text{sample}}}{R_{\text{standard}}} \right) - 1 \right] \cdot 1000$$
- where R_{sample} is the ratio of the heavy to light isotope (¹³C/¹²C or ¹⁵N/¹⁴N) and R_{standard} is the isotopic ratio of the standard.
- To determine the **isotope fractionation induced by the aerosol production**, the isotope ratios of the products and reactants are compared using the equation:

$$\Delta^{13}\text{C} = \delta^{13}\text{C}_{\text{products}} - \delta^{13}\text{C}_{\text{reactants}} \quad \text{or} \quad \Delta^{15}\text{N} = \delta^{15}\text{N}_{\text{products}} - \delta^{15}\text{N}_{\text{reactants}}$$

Results

Mixture	Pressure (Torr)	Collection Time (Hours)	$\Delta^{13}\text{C}$	$\Delta^{15}\text{N}$
50 ppm Benzene	100	~60	-4.3	9.6
50 ppm Benzene; 0.1% CH ₄	100	~120	13	-9.1
50 ppm Pyridine	100	~60	20	2.1
50 ppm Pyridine; 0.1% CH ₄	100	~120	30	2.1

- Significant differences in the isotopic ratios between the aerosols generated by the two different aromatic compounds (benzene and pyridine).**
- The addition of methane to the mixture significantly increases the aerosol production time
- The uncertainties of the data are still being assessed.

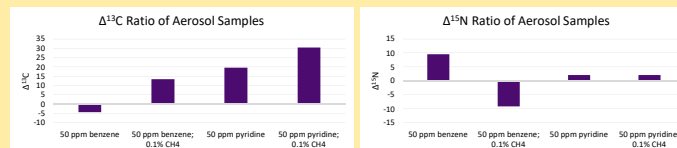


Figure 8. Comparison of the $\Delta^{13}\text{C}$ and $\Delta^{15}\text{N}$ values for the four different aerosol samples generated by aromatic gas mixtures in N₂

- The aerosols generated with the **pyridine/N₂ gas mixture are enriched in the ¹³C isotope**, while the aerosols generated by **benzene/N₂ gas mixture are depleted in the ¹³C isotope**.
- Methane significantly increases the $\Delta^{13}\text{C}$ ratio for the aerosols generated by both aromatic compounds.** This consistent previous research where methane generated aerosols, which were enriched in the ¹³C isotope.⁴
- The $\Delta^{15}\text{N}$ ratio for the benzene/N₂ generated aerosol seems to be significantly higher than the pyridine/N₂ generated aerosol but low amount of nitrogen in the sample could induce large uncertainties.
- The addition of methane to the benzene/N₂ gas mixture caused an enrichment in the ¹⁴N isotope over the ¹⁵N isotope in the generated aerosols.**
- Methane addition to the pyridine/N₂ gas mixture did not significantly change the $\Delta^{15}\text{N}$ ratio for the generated aerosols. This could be because the primary contribution to the N in the aerosol is from pyridine

Conclusions

- The aerosols produced in the laboratory setting demonstrate a change in isotopic ratio for ¹³C and ¹⁵N from the starting products.
- The addition of methane to a gas mixture appears to partially inhibit aerosol formation and increase collection time.
- Further work will need to be done to assess the effects of temperature and pressure on aerosol formation and isotopic ratio for ¹³C/¹²C, ¹⁵N/¹⁴N and D/H.

References

- ¹Nixon, C.A., et al., 2012. Isotopic ratios in Titan's methane: Measurements and modeling. *Astrophys. J.* **749** (2), 159.
- ²Rousselot, P. et al., 2014. Toward a unique nitrogen isotopic ratio in cometary ices. *Astrophys. J.* **780** (2), L17
- ³Sebree, J.A., et al., 2014. Titan aerosol analog absorption features produced from aromatics in the far infrared. *Icarus*. **236**, 146-152
- ⁴Sebree, J.A., et al., 2016. ¹³C and ¹⁵N fractionation of CH₄/N₂ mixtures during photochemical aerosol formation: Relevance to Titan. *Icarus*. **270**, 421-428
- ⁵Shinnaka, Y., et al., 2014. ¹⁴NH₂/¹⁵NH₂ ratio in Comet C/2012 S1 (ISON) observed during its outburst in 2013 November. *Astrophys. J.* **782** (2), L16.
- ⁶Trainer, M.G. et al., 2013. The Influence of benzene as a trace reactant in Titan aerosol Analogs. *Astrophys. J.* **766** (4)
- ⁷Trainer, M.G. et al., 2015. Titan aerosol formation as a sink for stable carbon and nitrogen isotopes. *EPSC*. **10**.
- ⁸Waite Jr., J.H., et al., 2007. The process of tholin formation in Titan's upper atmosphere. *Science*. **316** (5826) 870-875

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