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1 **Ultramafic clasts from the South Chamorro serpentine mud volcano reveal a**
2 **polyphase serpentinization history of the Mariana forearc mantle**

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16 ABSTRACT

17 Serpentine seamounts located on the outer half of the pervasively fractured Mariana
18 forearc provide an excellent window into the forearc devolatilization processes, which
19 can strongly influence the cycling of volatiles and trace elements in subduction
20 zones. Serpentinized ultramafic clasts recovered from an active mud volcano in the
21 Mariana forearc reveal microstructures, mineral assemblages and compositions that
22 are indicative of a complex polyphase alteration history. Petrologic phase relations
23 and oxygen isotopes suggest that ultramafic clasts were serpentinized at
24 temperatures below 200 °C. Several successive serpentinization events represented
25 by different vein generations with distinct trace element contents can be recognized.
26 Measured Rb/Cs ratios are fairly uniform ranging between 1 and 10, which is
27 consistent with Cs mobilization from sediments at lower temperatures and lends
28 further credence to the low-temperature conditions proposed in models of the

29 thermal structure in forearc settings. Late veins show lower fluid mobile element
30 (FME) concentrations than early veins, suggesting a decreasing influence of fluid
31 discharge from sediments on the composition of the serpentinizing fluids. The
32 continuous microfabric and mineral chemical evolution observed in the ultramafic
33 clasts may have implications as to the origin and nature of the serpentinizing fluids.
34 We hypothesize that opal and smectite dehydration produce quartz-saturated fluids
35 with high FME contents and Rb/Cs between 1 and 4 that cause the early pervasive
36 serpentinization. The partially serpentinized material may then be eroded from the
37 basal plane of the suprasubduction mantle wedge. Serpentinization continued but the
38 interacting fluids did not carry the slab-flux signature, either because FME were no
39 longer released from the slab, or due to an en route loss of FMEs. Late chrysotile
40 veins that document the increased access of fluids in a now fluid-dominated regime
41 are characterized by reduced trace element contents with a slightly increased Rb/Cs
42 ratio near 10. This lack of geochemical slab signatures consistently displayed in all
43 late serpentinization stages may indicate that the slab-derived fluids have been
44 completely reset (i.e. the FME excesses were removed) by continued water-rock
45 reaction within the subduction channel. The final stage of diapiric rise of matrix and
46 clasts in the conduits is characterized by brucite-dominated alteration of the clasts
47 from the clast rim inward (independent of the intra-clast fabric relations), which
48 corresponds to re-equilibration with alkaline, low-silica activity fluids in the rising mud.
49

50 *Keywords:* Serpentinization, polyphase alteration, mud volcano, fluid mobile
51 elements recycling, hydrated mantle wedge, forearc peridotites, subduction zone.

52

53

54 **1. Introduction**

55

56 Subduction zones are sites of intense mass transfer between the subducting plate
57 and the Earth's mantle. The incoming slab of oceanic lithosphere (sediments, as well
58 as variably altered mafic and ultramafic rocks) releases a generally high but variable
59 fluid flux (Bebout 2007). As a consequence of rheological contrasts between the
60 juxtaposed rock-types, shearing and metasomatism evolve a complex *mélange* zone
61 atop the slab (i.e., the subduction channel) where rocks and fluids are extensively
62 modified (Cloos and Shreve 1988a, 1988b). The inventory of slab lithologies and the
63 thermal structure determine where fluids leave the slab and where they enter the
64 subduction channel (Schmidt and Poli, 1998; Reynard, 2013; Spandler and Pirard,
65 2013). In particular, the flux of fluids is affected by the local setting, e.g., the slab
66 dipping angle and the presence, or absence, of an accretionary sedimentary wedge.
67 During subduction, the downgoing slab progressively releases water (Hyndman and
68 Peacock, 2006): at shallow depths by porosity compaction and dehydration reactions
69 of sediments, at intermediate depths by dehydration of sediments and altered
70 oceanic crust, and, at greater depths, by continued dehydration of oceanic crust and
71 deserpentinization of the subducting lithospheric mantle (Schmidt and Poli, 1998;
72 Rüpke et al., 2002, 2004). The geochemical signature of the fluids' source lithologies
73 is recorded by fluids and rocks affected by subduction zone processes: fore arc
74 springs, metasomatized *mélange* material, serpentized mantle peridotites, and arc
75 volcanism inherit the distinctive trace element contents of the reacting fluids
76 (Parkinson and Pearce, 1998; Fryer et al., 1999; Mottl et al., 2003, 2004; Savov et
77 al., 2005a; Hulme et al., 2010).

78

79 Mud volcanoes are a characteristic feature of the forearc of the Mariana subduction
80 zone system (Fryer et al., 1985; 1999; 2000). Located along the outer half of the

81 deeply fractured Mariana forearc, diapiric mud volcanism has formed large edifices
82 from extrusion of blueschist and serpentine mud, up to 50 km in diameter and up to 2
83 km high (Fryer et al., 1990, 1999; Oakley et al., 2007). Systematic studies of pore
84 waters and vent fluid compositions suggest that the fluids fueling the serpentine mud
85 volcanoes are released from breakdown of hydrous phases in the subducting slab.
86 Most noticeably, these fluids have lower chloride, Mg, Ca, Sr, Li, and Si
87 concentrations, and higher pH, alkalinity, K, Na, Rb, Cs and Ba contents than
88 seawater, indicating slab dehydration and water-rock interactions in the conduits
89 feeding the mud volcanoes (Mottl, 1992; Benton et al., 2001; Mottl et al., 2003, 2004;
90 Wheat et al., 2008; Hulme et al., 2010). The occurrence of blueschist fragments in
91 the mud and rare metabasic clasts in the near-trench forearc region may indicate the
92 involvement of upward migrating fluids and melange material in the subduction
93 channel from deeper parts of the slab-mantle interface (e.g., Bebout, 2007).
94 However, geobarometry on metabasite schists from South Chamorro Seamount
95 (Fryer et al., 2006; 250-300 °C at a pressure of ~0.4-0.50 GPa, ~12-15 km) and
96 Conical Seamount (Maekawa et al., 1993; temperatures of ~175-250 °C and
97 pressures of 0.5-0.65 GPa, ~15-19.5 km) indicate their origin in depths slightly above
98 the base of the feeding conduits. Therefore, the erupted clast and mud materials
99 record changes in the composition of the slab-sourced fluids, progressively released
100 and accumulated in shallow and intermediate depths during subduction.

101

102 Trace element bulk rock compositions of serpentinized ultramafic clasts and
103 serpentine mud have been used to gain deeper insights into the composition of the
104 serpentinized forearc mantle (Savov et al., 2007; Deschamps et al., 2011), and to
105 deduce potential parent rocks of the serpentinizing fluids (e.g., Hulme et al., 2010).
106 At South Chamorro seamount, Ocean Drilling Program Leg 195 drilled Site 1200,
107 where ultramafic clasts occur within the mud matrix (Straub and Layne, 2003;
108 D'Antonio and Kristensen, 2004; Fryer and Salisbury 2006). These clasts record a

109 complex serpentinization history and provide an excellent window into the prevailing
110 fluid-rock interactions at, and above, the interface between the mantle wedge and the
111 subducting lithosphere. Moreover, the microfabrics of the highly serpentinized
112 harzburgite and dunite clasts exhibit evidence for multiple fracturing events in the
113 forearc mantle.

114

115 Here, we present a detailed investigation of trace element contents of the individual,
116 successive serpentine generations and the primary mantle minerals. As many of
117 these clasts have preserved domains representing early-stage serpentinization, we
118 can reconstruct much of the history of fluid-rock reaction in the Mariana forearc and
119 identify potential sources of the serpentinizing fluids.

120

121

122 **2. Setting and samples**

123

124 *South Chamorro Seamount*

125 Site 1200 is located on a knoll at the summit of south Chamorro Seamount in a water
126 depth of 2930 m (Salisbury et al., 2002). Hole A was drilled with a rotary core barrel
127 to a depth of 147.2 meters below seafloor (mbsf) and recovered mainly serpentinized
128 ultramafic clasts and minor amounts of the mud matrix. South Chamorro Seamount
129 (Fig. 1A) is one out of two dozen active mud volcanoes of the Mariana forearc,
130 located 85 km west the trench where the Mesozoic Pacific Plate is being subducted
131 west-northwestward beneath the West Philippine Plate (Fryer et al., 1990; Fig. 1B).
132 The serpentine mud volcano was drilled in 2000 at Site 1200 of the Ocean Drilling
133 Program (ODP) during Leg 195. Estimates of the slab depth beneath Site 1200
134 range between 26 km (Fryer, 1996; 1999) and 20 km (Oakley et al., 2005, 2007;
135 Fryer and Salisbury 2006). The seamount is composed of unconsolidated flows of
136 serpentine and blueschist mud with clasts of serpentinized mantle peridotite and

137 minor blueschist fragments (Fryer and Fryer, 1987; Fryer, 1996; Salisbury et al.,
138 2002; Fryer et al., 2006). The mudflow matrix consists of c. 90 % serpentine, as clay-
139 to silt-sized lizardite, chrysotile and minor antigorite. Accessory minerals of the
140 mudflow matrix include brucite, talc, mixed-layer smectite/illite clays, chlorite, relict
141 grains of primary silicates and Cr-spinel, amphibole and garnet. The sizes of clasts
142 recovered from Site 1200 vary from a few mm to >1 m in diameter. The clasts
143 represent lithologies from both the subducting plate and the suprasubduction-zone
144 mantle. In Hole 1200A, the peridotite clasts are predominantly harzburgites and
145 minor dunites. Serpentinization is extensive and ranges between ~40 % and 100 %
146 with an average alteration of 95 % for harzburgites and 75 % for dunites (D'Antonio
147 and Kristensen, 2004).

148 In this study, nine samples from the marginal parts of highly serpentinized clasts of
149 harzburgite and dunite were selected for detailed petrographic and microanalytical
150 geochemical investigations (Table 1). A second set of adjacent samples from the
151 clasts' centers was chosen for thin section microscopy to identify the mineralogical
152 and textural differences between clasts' rims and cores (Table 1).

153

154

155 **3. Analytical methods**

156

157 *Petrography*

158 Mineral phases were identified by polarized light microscopy. In addition, a
159 *Panalytical X'Pert PRO* X-ray diffractometer (XRD) at the University of Bremen (Cu
160 K α radiation, 45 kV, 40 mA) was used to identify serpentine modifications in the
161 samples selected for oxygen isotope analysis (see below). Data were collected in the
162 continuous mode, between 3 to 85 $^{\circ}2\theta$, in steps of 0.017 $^{\circ}2\theta$. Spectra were
163 processed with *High-Score Plus* software.

164

165 *Electron microprobe*

166 Major element analyses of mineral compositions were performed with a *JEOL JXA-*
167 *8900* electron microprobe at the University of Kiel (Germany). For most silicate
168 minerals an accelerating voltage of 15 kV was used, with exception of garnet, spinel
169 and iron oxides (20 kV). Measurement spot sizes were typically 1-5 μm in diameter.
170 Standards were either natural minerals or synthetic materials. The CITZAF method of
171 Armstrong (1995) was used for correction of the raw counts. Micro-scale element
172 mappings and backscattered electron (BSE) images of rock microfabrics were used
173 to complement microscopic observations.

174

175 *Laser ablation ICP-MS*

176 Analyses of major and trace element concentrations were carried out on thin sections
177 by laser ablation inductively coupled plasma mass spectrometry (LA-ICP-MS) using
178 a Thermo Element2 coupled to a NewWave UP193ss (wavelength 193 nm) at the
179 Department of Geosciences of Bremen University. Helium (0.8 l/min) was used as
180 carrier gas and argon (0.8-0.9 l/min) was added as make-up gas; plasma power was
181 1200 W. Samples and standards were ablated with an irradiance of $\sim 1 \text{ GW/cm}^2$, and
182 a pulse rate of 5 Hz. The laserbeam diameter was typically between 25 and 100 μm .
183 Prior to ablation the blank intensities were acquired during 25 s. All isotopes were
184 analyzed at low resolution with five samples in a 20% mass window and a total dwell
185 time of 25 ms per isotope. Depending on the mineral composition, either silicon or
186 magnesium was used as an internal standard element, which had previously been
187 determined by electron microprobe analysis. Calibration was done using glass
188 standards NIST610 and NIST612, with preferred averages from Jochum et al.
189 (2011). Analytical quality was controlled by repeated analyses of the international
190 glass standard BCR2-G (Jochum et al., 2006).

191

192 *Micro-Raman investigations*

193 Serpentine phases were identified using a Horiba LabRAM HR confocal Raman
194 spectrometer equipped with a 532 nm laser, an astigmatic flat field spectrograph with
195 a focal length of 800 mm, and a multichannel air-cooled (-70 °C) CCD detector.
196 Spectra were collected using a 100x objective with a numerical aperture of 0.9. Two
197 diffraction gratings with 600 and 1800 grooves / mm were used. A confocal hole
198 diameter of 100 μm was chosen for most analyses. Spectra were collected for 5
199 seconds with 3-5 accumulations between 100 cm^{-1} and 1280 cm^{-1} .

200

201 *Oxygen isotope analysis*

202 For oxygen isotope analysis, serpentine minerals of the different generations were
203 separated by crushing and hand picking, and purity and mineralogy were checked by
204 X-ray diffraction (see above) and visual inspection. The oxygen isotope analyses
205 were carried out at the University of Michigan following the procedure of Alt et al.
206 (2010). Samples were reacted with BrF_5 in externally heated nickel reaction vessels
207 (Clayton and Mayeda, 1963), and O_2 was converted to CO_2 gas by reaction with
208 heated carbon rods. Oxygen isotope ratios were measured using a Finnigan Delta-S
209 mass spectrometer. These data were normalized to a value of 9.6‰ for NBS 28
210 quartz standard ($\pm 0.2\text{‰}$, 1σ), compared to the accepted value of 9.58 ‰.

211

212

213 **4. Results**

214

215 The observations presented here are based on nine extensively serpentinized
216 harzburgite and dunite clasts sampled from different depth intervals of Hole 1200A
217 (see Table 1). In particular, the clast material selected for thin section preparation
218 and analytical investigation was collected from near-rim regions that show
219 considerable influence of late-stage alteration events.

220

221

222 4.1 Petrography

223

224 The serpentized ultramafic clasts exhibit a variety of microfabrics that range from
225 virtually undeformed to strongly deformed (see Figs. 2 and 3). The degree of
226 serpentization we observed in thin section varies between 75 % and almost 100 %,
227 but can be highly variable on the scale of centimeters. Even within the same thin
228 section, weakly serpentized regions can be observed adjacent to regions with
229 abundant veins of different generations.

230

231 Relicts of primary olivine with average grain diameters of around 50 μm (some grains
232 are up to 250 μm in diameter) show microgranulation and kink-banding in some
233 areas. The size of the original olivine grains in the precursor rock was up to 3 mm.
234 Relicts of orthopyroxene show signs of crystal deformation (undulatory extinction and
235 cleavage planes); former grains were up to 4 mm in size. Clinopyroxene is preserved
236 as solitary crystals (diameter up to 100 μm) either in proximity to orthopyroxene, or
237 intimately intergrown with chromian spinel. In addition, clinopyroxene is present as
238 exsolution lamellae in orthopyroxene. In a dunite sample, spinel exhibits rhomb-
239 shaped euhedral faces, whereas in the harzburgites studied it commonly shows
240 embayments indicative of resorption.

241

242 *Fabric relations*

243

244 Serpentinization of the primary minerals led to pseudomorphic mesh and hourglass
245 textures after olivine and bastite after orthopyroxene (and locally of clinopyroxene).
246 Serpentine is present in veins of several generations. Besides serpentine group
247 minerals, brucite is a common secondary phase. It is observed in intimate

248 intergrowth with serpentine in veins, and as single crystals in the centers of
249 recrystallized mesh textures.

250

251 Heazlewoodite and awaruite (Ni_3Fe) are present in trace amounts. The atomic
252 metal/sulfur ratio of heazlewoodite is near stoichiometric and varies between 1.59
253 and 1.68. Small amounts of Fe account for a representative composition of
254 $(\text{Ni}_{2.97}\text{Fe}_{0.03})\text{S}_2$. Heazlewoodite can be present in the centers of hour glass structures,
255 or as rare occurrence in veins, rimmed by awaruite. Moreover, awaruite is observed
256 on trails in the center of serpentine veins in mesh-textured olivine (Figs. 3A and B).
257 Awaruite and in a few instances wairauite (CoFe) are present in halos around type (i)
258 veins and in type (ii) veins (vein types as detailed below). Magnetite is developed in
259 early veins, which predate pervasive serpentinization as characteristic (double)
260 central vein filling (D'Antonio and Kristensen, 2004; Klein et al., 2009). Furthermore,
261 magnetite was found in late, extension-related veins as syntactically grown crystals.
262 Garnet of andraditic composition (with neglectable Al and Mg contents, and with a
263 hydroandradite (hydrogarnet) component, calculated from Si deficiency) is present in
264 two samples. It is either homogeneously distributed as small euhedral crystals (up to
265 $50\ \mu\text{m}$ in diameter) or clustered in anhedral crystals (individual crystals up to c. $70\ \mu\text{m}$,
266 clusters up to $300\ \mu\text{m}$). In either case, garnet growth was located in mesh-
267 textured domains, commonly at the boundary between mesh core and rim.
268 Apparently, those parts of the grains growing into the mesh core developed anhedral
269 morphology, contrasting with the euhedral habit of the faces that grew towards the
270 mesh rim. A representative garnet composition is $(\text{Ca}_{2.98}\text{Mg}_{0.04})(\text{Fe}_{1.98}\text{Al}_{0.01}\text{Cr}_{0.00})$
271 $(\text{Si}_{2.81}\text{H}_{40.19})\text{O}_{12}$, the hydroandradite component shows values between 0.14 and
272 0.27.

273

274

275 *Pervasive serpentinization*

276 Pervasive serpentinization of olivine resulted in the formation of mesh (with fresh
277 olivine in the cores) or hourglass textures (see Figs 2, and 3A and B). Former olivine
278 grain boundaries in mesh textures are outlined by awaruite; in most samples the
279 mesh texture is free of magnetite. Pseudomorphic serpentinization of orthopyroxene
280 (and locally of clinopyroxene) led to the formation of bastites. Former interstitial
281 spinel is apparently unaffected by alteration and is (in places) surrounded by
282 serpentine coronae formed from breakdown of orthopyroxene and olivine (as
283 described by Seyler et al., 2001). Brucite was not observed in pervasive mesh
284 textures. However, it is present in interstices between clinopyroxene (sub)grains, in
285 regions of mesh textures that were affected by vein serpentinization, or recrystallized
286 hourglass textures (see Figs. 2A, 2G). Serpentine in mesh textures and bastites has
287 been identified as lizardite by Raman spectroscopy (for serpentine polymorphs
288 identified in the polyphase serpentinization fabrics, see Table 4), which is consistent
289 with investigations of similar textures in other serpentinites (e.g., Dilek et al., 1997;
290 Andreani et al., 2007; Boudier et al., 2010; Pabst et al., 2011; Schwartz et al., 2012).

291

292

293 *Veins*

294 The most striking fabric element present in the clasts is the abundant veining, which
295 comprises different vein generations.

296

297 In harzburgites, several stages of veining and associated serpentinization can be
298 distinguished:

299

300 Early magnetite-bearing veins (predating pervasive serpentinization)

301 Serpentine veins containing magnetite were found in one of the thin sections studied
302 (see Fig. 2A). In these transgranular veins magnetite occurs as central vein fillings or
303 as double trails along the vein rim. This fabric type has been reported in other

304 ultramafic clasts recovered from Site 1200 (e.g., D'Antonio and Kristensen, 2004), as
305 well as in serpentinized abyssal peridotites from the Mid-Atlantic Ridge (recovered
306 from Leg 209, Hole 1274A, see Klein et al., 2009). These veins are likely an early
307 feature, preceding and maybe initiating the pervasive serpentinization of pristine
308 ultramafic rock (Salisbury et al., 2002).

309

310 Type (i) veins

311 This vein type postdates pervasive serpentinization and features lizardite with brucite
312 and awaruite; veins are straight without marked conformance with the rock fabric
313 (see Fig. 2D1, and 2E). The walls of type (i) veins may show finger-like
314 serpentinization fronts. These "fingers" seem to emanate into interstices of mesh
315 textured domains formed during a preceding pervasive serpentinization event. In
316 places, these veins developed mm- to cm-wide halos. In places, veins of type (i)
317 have been reactivated by later fluid fluxing events of type (iv) veins.

318

319 Type (ii) veins

320 These chrysotile-brucite veins (Table 4) encompass regions of weakly serpentinized,
321 microgranular olivine and large orthopyroxene crystals (Fig. 2B-E). They are
322 brownish in color and up to 2 mm wide. Type (ii) veins do not show a preferred
323 orientation and appear to have formed to accommodate expansion which is not related
324 to expansion during pervasive serpentinization. However, the orientation of the fibers
325 with respect to the vein walls documents an anisotropic stress field. It is unclear,
326 whether this feature represents growth or recrystallization of the vein serpentine.

327

328 Type (iii) veins

329 Extensional chrysotile veins (known as "Frankenstein" type, see e.g., Salisbury et al.,
330 2002) have developed in halos along type (i) veins (where they may intersect type (ii)
331 veins; Fig. 2D, 2E, and Table 4). They are related to the accommodation of local

332 stresses during the overall isotropic expansion of the surrounding rock. A similar vein
333 geometry has been described in detail from serpentinitized orthopyroxenite dykes in a
334 dunite matrix from the Leka Ophiolite Complex, Norway (Iyer et al., 2008a, 2008b).

335

336 Type (iv) veins

337 Late-stage extensional chrysotile veins document multiple crack-seal events (as
338 described by Bons et al., 2012) either in undeformed regions with preserved granular
339 fabric, or in serpentinites hosting type (i) and (ii) veins where they crosscut and offset
340 the earlier vein generations (see Fig. 2F).

341

342 Extensional magnetite-bearing veins

343 One thin section contains magnetite, present in deformation-related extensional
344 veins, which developed as syntaxial veins from a single crack-seal event. Large
345 crystals extend from one wall to the other (as confirmed with high-resolution X-ray
346 microtomography).

347

348 Some dunites host syntaxial serpentine veins (ribbons). The ribbons separate
349 regions, in which recrystallization of serpentine and brucite can be observed (see
350 Fig. 2G). The extension accommodated by the ribbons is clearly anisotropic.

351

352

353 *Late Browning*

354 A "browning" of the clasts' margins is developed in few samples (e.g., 3R-1, 7R-2,
355 13R-2, 16R-1), which is related to an inward-migrating alteration front (Fig. 2H). This
356 browning is independent of the intra-clast fabric relations comprised by the different
357 vein generations. Browning led to high proportions of brucite in the rock that
358 proceeded either localized along former grain or sub-grain boundaries or in a

359 pervasive manner. The latter may develop by coalescence of grain boundary scale
360 replacement.

361

362

363 4.2 Mineral chemistry

364

365 *Primary minerals*

366 Olivine in both harzburgite and dunite is Mg-rich ($X_{Mg} = [Mg/(Mg+Fe)]$ of 0.92) (Figs.
367 4A-D, Table 2), which is typical for mantle wedge peridotite (Bonatti and Michael,
368 1989; Ishii et al., 1992). Orthopyroxene has X_{Mg} values between 0.91 and 0.92, Al_2O_3
369 contents from 1.3 to 2.5 wt.%, and contains 0.6 to 1.2 wt.% Cr_2O_3 (see Figs. 4A-D,
370 and 5A-C). X_{Mg} ratios of clinopyroxene range between 0.94 and 0.95, with 0.96 - 2.02
371 wt.% Al_2O_3 , and 0.24 - 1.09 wt.% Cr_2O_3 . Chromian spinel is variable in composition
372 with X_{Cr} [$Cr/(Cr+Al)$] values between 0.48 and 0.63, and X_{Mg} ranging from 0.51 to
373 0.59 in the pervasively serpentinized regions in harzburgites.

374 In the two most strongly serpentinized harzburgites and dunites (9R-1 and 11R-1),
375 spinel approaches ferritchromite composition ($X_{Cr} = 0.69 - 0.93$ and $X_{Mg} = 0.22 -$
376 0.43). Ferritchromite is commonly regarded as an indicator of Cr-spinel alteration
377 (e.g., Mellini et al., 2005; Merlini et al., 2009; Saumur and Hattori, 2013).

378

379 *Serpentine minerals*

380 The polyphase serpentinization history of the ultramafic clasts is apparent in distinct
381 mineral compositions of the successive serpentine generations (Figs. 4A-D and 5A-
382 C). Pure chrysotile/lizardite has a molar $(Mg+Fe^{2+})/Si$ ratio of 1.5. However, since the
383 majority of serpentine mineral EMPA analyses of the early vein generations of type
384 (i) and (ii) exhibits significantly lower SiO_2 and higher MgO contents, it appears that
385 minor brucite is intergrown with serpentine on a submicroscopic scale (Table 3).

386

387 Compositions of pseudomorphic serpentine minerals formed during the pervasive
388 serpentinization phase were found to be indicative of their precursor mineral in terms
389 of Al_2O_3 and CaO contents of the serpentine phase (Fig. 5A-C). Serpentine in mesh
390 and hourglass textures has X_{Mg} values of 0.90 - 0.94. Serpentine in mesh and
391 hourglass textures does contain as little Cr_2O_3 as the olivine precursor, but exhibits
392 higher Al_2O_3 and CaO as olivine. These serpentines in harzburgites ($\text{Al}_2\text{O}_3 > 0.1$ to <
393 1.0 wt.%, CaO: 0.1 - 0.6 wt.%) tend to be richer in both components than dunites
394 ($\text{Al}_2\text{O}_3 < 0.1$ wt.%, CaO: 0.08 - 0.2 wt.%). In recrystallized hourglass textures in
395 dunite, the X_{Mg} of serpentine coexisting with brucite is slightly higher (0.94-0.96) than
396 X_{Mg} of serpentine in unrecrystallized regions. No differences between mesh rim and
397 core compositions have been observed. Bastite textures consist of serpentine, with
398 X_{Mg} values between 0.89 - 0.94. Most bastites retain the high Al_2O_3 contents (> 1.0
399 wt.%), as well as the CaO (0.1 - >1.0 wt.%) and Cr_2O_3 (0.1 - 1.0 wt.%) values of the
400 pyroxene precursor (see Fig. 5A-C). Analyses of serpentine in halo regions of type (i)
401 veins and in type (ii) veins show marked SiO_2 deficits (Fig. 4B-C), indicating
402 intergrowth with submicroscopic brucite, and low X_{Mg} values (an average X_{Mg} of 0.80
403 in Brc-Srp intergrowth, and down to X_{Mg} of 0.40 in coexistence with Cpx).
404 Serpentinites in the halos of type (i) veins show identical contents of Al_2O_3 and CaO
405 as the pseudomorphs outside the halos. Serpentine in type (iii) and type (iv) veins
406 shows SiO_2 contents expected for pure serpentine and X_{Mg} values of 0.89 - 0.96 (see
407 Fig. 4D). Later veins of type (ii)-(iv) also exhibit the characteristic Al_2O_3 and CaO
408 contents of the pseudomorphs present in their host rocks (see Fig. 5A-C).

409

410 Using Micro-Raman investigations of the polyphase serpentinization fabrics, we have
411 identified two serpentine polymorphs: Lizardite and chrysotile (see section 4.1 and
412 Table 4). Lizardite has been found in the rims of mesh textures, surrounding olivine
413 cores, and in bastites. Moreover, it is present in mesh/hourglass and bastite
414 pseudomorphs in regions that were affected during the halo formation around type (i)

415 veins. Later vein generations (type ii-iv) are comprised of chrysotile. In the clasts
416 investigated by micro-Raman, no antigorite was identified.

417

418 *Brucite*

419 Brucite is present either as readily identifiable crystals (up to 100 μm diameter) in
420 interstices or recrystallized areas, or fine-grained as vein fill. The occurrence of
421 brucite submicroscopically intergrown with serpentine minerals has been recognized
422 from microprobe analyses based on SiO_2 deficits, MgO excesses, and low totals (see
423 above, and Table. 3). Brucite-serpentine intergrowths in harzburgites were also
424 observed in type (i) and type (ii) veins and point to brucite compositions of $X_{\text{Mg,Bruc}}$
425 0.79 - 0.85. Brucite is apparently lacking in the pervasive serpentinization stage as
426 well as in type (iii) and in type (iv) veins.

427

428 In harzburgites, interstitial brucite is present in type (i) veins or halos, typically with
429 MnO contents from 0.19 to 0.53 wt.%, and X_{Mg} values between 0.79 and 0.90. In
430 places, type (i) veins contain brucite with slightly elevated MnO contents between
431 0.94 and 2.20 wt.% and low X_{Mg} values of 0.59 - 0.63 (these MnO values correspond
432 with brucite analyses given in D'Antonio and Kristensen, 2004). A single occurrence
433 of brucite in contact with clinopyroxene features 1.65 to 2.19 wt.% MnO and X_{Mg}
434 values between 0.36 and 0.47. In dunites, single crystals of brucite can be observed
435 where cores of hourglass textures have recrystallized (Fig. 2G). The X_{Mg} values of
436 the recrystallized brucite range between 0.83 to 0.85.

437

438

439 4.3 Trace element mineral chemistry

440

441 Trace element concentrations of serpentine provide insights into potential precursor
442 phases and compositions of serpentinization fluids.

443

444 Ni and V contents of serpentine are different for olivine and pyroxene precursors.
445 Figure 6A shows that serpentine (or regions of serpentine-brucite intergrowth)
446 formed at the expense of olivine can contain up to 0.5 wt.% Ni (with an average
447 around 0.3 wt.%), corresponding to the Ni content of the primary olivine. Or rather, Ni
448 is most likely measured as fine-grained awaruite, present in the laser-ablation spots).
449 In contrast, bastite-serpentine (i.e. orthopyroxene precursor) has < 0.1 wt.% Ni,
450 corresponding to the average Ni concentration of ortho- and clinopyroxene (c. 700
451 ppm). Likewise, V concentrations of serpentine (Fig. 6B) replacing olivine are mostly
452 below 1 ppm, which reflects the V concentration of its precursor. In bastite textures,
453 V contents are up to 80 ppm, corresponding to 60-80 ppm V in orthopyroxene. On
454 rare occasions, bastitization of clinopyroxene has been observed in completely
455 serpentinized clast rims (identified by the shape of the former grain). Here, vanadium
456 contents are slightly higher as in orthopyroxene bastites (60-120 ppm). The Ni-V
457 variation diagram (Fig. 6C) shows the discriminatory significance of these trace
458 elements in identification of the serpentine parent mineral. The variation in the Cr-Ni
459 diagram (not shown) is identical apart from the higher absolute concentrations of
460 chromium (section 4.2), corresponding with data from Kodolanyi et al. (2012).

461

462 B, Ba, Sr, Rb and Cs data of regions with submicroscopic intergrowth of serpentine
463 and brucite show increasing trace element concentrations with increasing
464 Si/(Mg+Fe)-ratio, which indicates that these elements are incorporated in the
465 serpentine structure and are not present in brucite (Figs. 7A-E). Therefore, the
466 element concentrations of the mixed analyses of serpentine-brucite regions have
467 been recalculated to pure serpentine (indicated by the curved lines in Figs. 7A-E).

468

469 Regardless of the nature of the precursor phase and presumed formation stage, all
470 serpentinization products are enriched in B (Figs. 7A and 7F), and calculated

471 serpentine endmember compositions have up to 200 ppm boron (mostly between 80-
472 100 ppm). Ba, Sr, Rb, and Cs are also commonly enriched in the serpentinization
473 products and veins relative to the precursor phases (Fig. 7B-E). In contrast to B,
474 however, these elements show distinct differences in concentration between
475 serpentine formed during different stages of veining and fluid-rock interaction. Most
476 notably, early vein generations show high trace element concentrations compared
477 with later veins (Table 4).

478

479 On average, serpentine of the early serpentinization stages (i.e., in mesh rims of
480 pervasive stage, in type (i) veins, or in the halo around type (i) veins) exhibit a strong
481 enrichment in Rb (0.8 ppm), Cs (0.4 ppm), Ba (> 10 ppm), and Sr (> 50 ppm). For Ba
482 and Sr the enrichment is unrelated to the nature of the precursor mineral, while
483 bastites, and interstitial serpentine of regions with only weakly serpentinized
484 orthopyroxene and microgranular olivine crystals, have even higher concentrations of
485 Rb (up to 2.6 ppm) and Cs (up to 1.2 ppm; Figs. 7D and E). In the later vein
486 generations, remarkable differences in the enrichment of either (Rb and Cs) or (Ba
487 and Sr) occur. In type (ii) veins, average Ba (3.7 ppm, up to 11.3 ppm) and Sr (13.6
488 ppm, up to 27.7 ppm) concentrations are also high. In type (iii) and (iv) veins,
489 serpentines show average concentrations of Ba (0.4 ppm) and Sr (3 ppm), which
490 markedly lower than in the earlier vein generations (types i and ii), but still
491 significantly higher than those of the primary minerals (Figs. 7B and C). In contrast,
492 the average concentrations of Rb and Cs in serpentines of type (ii)-(iv) veins are low
493 (<0.1-0.3 ppm) and similar to those of the primary phases Figs. 7D and E).

494

495

496 4.4 Oxygen isotope compositions of serpentine minerals

497

498 The measured oxygen isotope compositions of serpentine minerals range between
499 6.1 and 10.5 $\delta^{18}\text{O}$, consistent with serpentinite clasts from Conical, Torishima, and
500 Ogasawara Seamounts (Alt and Shanks, 2006; Sakai et al., 1990). Table 5 provides
501 an overview of the oxygen isotope data for the mineral separates along with details
502 about mineralogical composition of the different samples (as determined by XRD
503 analyses). The oxygen isotope data do not show significant differences between
504 different vein generations or serpentine polymorphs. Note that two mineral separates
505 contained a noticeable fraction of brucite (16R-2).

506

507

508 **5. Discussion**

509

510 *5.1 Low-temperature serpentinization conditions in the Mariana mantle wedge*

511

512 The serpentinized ultramafic clasts exhibit a variety of microfabrics that indicate a
513 polyphase serpentinization history. A common feature during all stages of
514 serpentinization is that little magnetite was formed (if at all). In contrast to samples
515 from mid-ocean ridge settings, mesh textures lack magnetite tracing former olivine
516 (sub-)grain boundaries. Instead, the serpentinized ultramafic clasts are composed of
517 serpentine \pm brucite \pm iron alloys (mostly awaruite). This magnetite-free assemblage
518 has been found as late-stage formation in rocks from magma-poor rifted continental
519 margins (Beard et al., 2009), in the forearc of subduction zones (Lafay et al., 2012),
520 and in low-temperature serpentinization experiments (batch reactor: Seyfried et al.,
521 2007; flow-through experiment: Godard et al., 2013). In a recent study, Klein et al.
522 (2014) propose that decreasing abundance of magnetite is indicative of decreasing
523 alteration temperatures. Instead of magnetite, Fe-rich brucite forms at temperatures
524 below ca. 200°C (cf. Klein et al., 2009). These low formation temperatures would be
525 consistent with temperatures predicted for a depth of 25 km by models of the thermal

526 structure of a cold oceanic subduction zone (Izu-Bonin, see Hyndman and Peacock,
527 2003).

528

529 Metabasite clasts may serve as an additional constraint on p,T-conditions that prevail
530 at the slab-mantle interface recorded at variable depth levels reaching from
531 intermediate to near the base of the conduit of the mud diapir. In the case of South
532 Chamorro Seamount and Conical Seamount, metabasite schists indicate low
533 temperatures of 250-300 °C (at a pressure of •0.4-0.50 GPa; Fryer et al., 2006) and
534 ~175-250 °C (at a pressure of 0.5-0.65 GPa; Maekawa et al., 1993), in depths
535 slightly above the base of the feeding conduits.

536

537 The $\delta^{18}\text{O}$ values for serpentine in this study (6.1 to 10.8 ‰) are similar to those from
538 Conical seamount and to those dredged from Torishima and Ogasawars Seamounts
539 (Sakai et al., 1990; Alt and Shanks, 2006). Using oxygen isotope fractionation
540 between serpentine and magnetite, Alt and Shanks (2006) calculated higher
541 temperatures of 300-375 °C for serpentinite clasts from Conical Seamount. The
542 dominance of antigorite in some of the Conical Seamount samples suggests that
543 higher estimated temperatures may be real, as antigorite is stable at higher
544 temperatures (above 300 °C, Schwartz et al., 2012), but these temperature estimates
545 were based on mineral separates from bulk rocks (Alt and Shanks, 2006), which
546 integrate multiple stages of serpentinization and magnetite formation, leading to
547 larger uncertainties. In the rocks studied in this work, no pairs of coexisting
548 serpentine and magnetite were at hand for temperature estimation. Since the
549 presence of magnetite is restricted to scarce occurrence in vein types not related to
550 the main serpentinization phases (see above), it is not very likely, that a magnetite
551 separated from bulk rock powders would be coeval with chrysotile separated from
552 late veins. Rather, fluid-rock ratios, the fluids as such, and maybe also the
553 temperatures prevalent at early mesh formation (yielding magnetite, given

554 appropriate conditions) would differ from late-stage conditions governing the
555 formation of late-stage, extensional chrysotile veins. However, by constraining some
556 boundary conditions for the fluid-rock interaction we can estimate serpentinization
557 temperatures of the 1200A clasts.

558

559 Using the serpentine-water oxygen isotope fractionation from Saccocia et al. (2009),
560 temperatures of 175-240 °C are calculated assuming equilibrium between serpentine
561 in Table 5 and fluids sampled from South Chamorro seamount ($\delta^{18}\text{O} = +2.5\text{‰}$; Fryer
562 et al., 2006). Lower temperatures of 145-200 °C are calculated assuming equilibrium
563 with seawater (0‰). Stable isotope compositions of serpentinites from Conical
564 seamount indicate derivation of serpentinizing fluids from the subducting slab (Alt
565 and Shanks, 2006). If serpentinizing fluids contained a component of sediment pore
566 waters, which have negative $\delta^{18}\text{O}$ values (Lawrence and Drever, 1981), then
567 equilibrium temperatures for serpentine are even lower (125-170 °C for -2‰ fluid;
568 Saccocia et al., 2009). These temperature estimates are consistent with low
569 temperatures inferred from mineralogy of Chamorro samples and with thermal
570 models for cold subduction (Klein et al., 2009; Hyndman and Peacock, 2003).

571

572 The above estimates suggest relatively low temperatures (below 250 °C) for the
573 serpentinite samples from South Chamorro seamount (Site 1200, distance from
574 trench: 85 km), which seemingly is in contrast to the higher serpentinization
575 temperatures of 300 - 375 °C estimated for samples from Conical seamount (Site
576 780, distance from trench: 90 km; distance from Site 1200: almost 1000 km) by Alt
577 and Shanks (2006). While the high-temperature polymorph antigorite is the dominant
578 serpentine phase in several samples from Conical seamount, it is scarce in South
579 Chamorro and other Mariana forearc seamounts (see section 5.2). On the contrary,
580 the absolute concentrations of Cs, Rb, Ba, and Sr observed in serpentine in this

581 study indicate that the serpentinizing fluids of the South Chamorro samples derive
582 from sediment dehydration at temperatures of ~ 150 °C (section 5.3), similar to the
583 estimations by Alt and Shanks (2006) for Conical. We consider this an expression of
584 both a differing subduction zone geometry and the variation in the thermal structure
585 of the slab and wedge: the mud volcanoes can sample the entire thickness of the
586 wedge, and can include samples from any time within the 40 million year history of
587 the subduction zone. So they could include early higher-temperature serpentinites,
588 which could be one explanation for high-temperature samples at Conical seamount.
589 Certainly in a "mature" subduction zone, without mantle counterflow in the wedge,
590 the low temperatures as shown by the Chamorro samples are expected. Moreover,
591 the source region of the ultramafic material mobilized as clasts and erupted in mud
592 flows does not necessarily lie within the wedge. Besides serpentinization of conduit
593 wall rock, also material from the basal plane of the suprasubduction mantle wedge,
594 partly serpentinized and eroded in shallower parts of the decollement, may rise
595 through the conduit system and be erupted within the mudflows (section 5.4).

596

597

598 *5.2 Textures and serpentine polymorphs*

599

600 Different low-temperature serpentine polymorphs are common in serpentinized
601 abyssal peridotites. In the present work, micro-Raman investigations demonstrate
602 the occurrence of two serpentine polymorphs in distinct textural context. Lizardite is
603 present in the rims of mesh textures, surrounding olivine cores, in bastites, and in
604 early type (i) veins. Later vein generations are comprised of chrysotile. Antigorite,
605 which is commonly seen as indicative for higher-temperature serpentinization, could
606 not be identified. In previous studies of ultramafic clasts from ODP Leg 195, Site
607 1200 (D'Antonio and Kristensen, 2004; Pabst et al., 2011), lizardite and chrysotile,
608 but not antigorite, were identified. At Conical Seamount, Ishii et al. (1992), and

609 Saboda et al. (1992) observed antigorite in thin sections of serpentinized dunite
610 (Hole 780C) and harzburgite clasts (Hole 779A). Alt and Shanks (2006) observed
611 antigorite in serpentinized harzburgites and dunites from Site 779 and 780, being the
612 dominant serpentine phase in several samples. In contrast, antigorite was putatively
613 identified by Fryer and Mottl (1996) in muds at Holes 778A and 779 A (in 3 out of 16
614 samples) but appears absent in Site 780 muds. As stated in Savov et al., (2005a),
615 scarcity or absence of antigorite in the Mariana forearc serpentinites (apart from
616 Conical Seamount) are noted by many authors (e.g., Saboda et al., 1992; Zack et al.,
617 2004; Murata et al., 2009; Pabst et al., 2011), while chrysotile and lizardite are
618 common.

619

620 In numerous studies of oceanic serpentinites, the growth of distinct serpentine
621 polymorphs under different reaction conditions (e.g., varying water rock ratios,
622 tectonic regimes etc.) was noted. For instance, lizardite in rims around olivine mesh
623 centers, and in bastites, has been reported from the mid-ocean ridge Site 920, e.g.,
624 by Dilek et al., (1997, by XRD), Andreani et al., (2007, by TEM; 2013, by Raman),
625 from Sites 1272 and 1274 by Kodolanyi et al., (2012, by Raman), and from
626 ophiolites, e.g., by Boudier et al., (2010, TEM), Hirauchi et al., (2010, Raman), and
627 Schwartz et al., (2012, Raman). The clasts from Mariana forearc mantle may exhibit
628 a great diversity with respect to (i) the differing degree of serpentinization, (ii) the
629 differences in the source regions within the mantle wedge, and (iii) the relative timing
630 of serpentinization stages in the individual clasts. Pabst (2009) found either lizardite
631 or chrysotile both in mesh rims and mesh centers, and also the occurrence of
632 polygonal serpentine is reported. Pabst (2009) found bastites composed of lizardite,
633 in other samples of polygonal serpentine. Kodolanyi et al. (2012) reported the
634 occurrence of lizardite in mesh rims and bastites, and of chrysotile in mesh centers
635 and bastites. In the present study, lizardite has been found in early serpentinization
636 textures, and chrysotile in intermediate and late vein generations.

637

638 In most of the aforementioned studies, chrysotile has been found in extensional
639 veins, crosscutting pervasive serpentinization and in earlier vein generations.
640 Furthermore, Andreani et al., (2007, 2013), mention the involvement of polygonal
641 serpentine in the vein mineralization. In general, the preferential occurrence of
642 chrysotile in expansional regimes with higher water-rock ratios has also been noted
643 for the ultramafic clasts from the Mariana forearc mud volcanoes (e.g., D'Antonio and
644 Kristensen, 2004; Pabst et al., 2011; and described in more detailed by Pabst, 2009).

645

646 In the present study, the pseudomorphic serpentinization textures suggest that early
647 pervasive serpentinization took place under static conditions (e.g., Wicks and
648 Whittaker, 1977; O'Hanley 1996), yielding isotropic expansion. In the course of the
649 successional veining events, the expansional regime gradually became more
650 anisotropic. The formation of type (i) veins points to clearly anisotropic stress,
651 however related to merely minor expansion. Type (ii) veins do not show a preferred
652 orientation and appear to have accommodated considerable overall isotropic
653 expansion. The extensional chrysotile veins of the "Frankenstein" type (iii) veins
654 accommodate local stresses in the halos along type (i) veins, associated with the
655 overall isotropic expansion of the surrounding rock. Late-stage extensional type (iv)
656 chrysotile veins formed in a clearly anisotropic stress regime and led to anisotropic
657 expansion, documented as multiple crack-seal events (as described by Bons et al.,
658 2012).

659

660

661 *5.3 In-situ mineral analyses reveal distinct FME patterns for successive*
662 *serpentinization events and potential fluid source rocks*

663

664 Since the proportions of the different serpentine generations in the bulk sample and
665 the fraction of brucite in Srp-Brc intergrowth regions are highly variable, it is difficult
666 to unravel the formation history of the clasts using trace element contents of bulk
667 rocks. For instance, Savov et al., (2007) observed no correlation between the degree
668 of serpentinization (expressed as LOI wt.%) and the abundance of FME and Th. The
669 detailed mineral analyses of serpentine, in contrast, facilitate the recognition of
670 different trace element contents in distinct veins and microfabrics, which can be
671 linked to distinct episodes in the polyphase serpentinization history of mantle
672 material.

673

674 The observed trace element contents in the clasts from South Chamorro Seamount
675 decrease with progressive serpentinization. These elements are incorporated in the
676 serpentine structure and are not present in brucite (Figs. 7A-E; section 4.3). While
677 later serpentinization stages recorded in the clasts - and the trace element
678 concentration measured in serpentine of the mud matrix - are characterized by
679 decreased FME concentrations, both bastites and interstitial serpentine in
680 microgranular olivine regions formed during preceding serpentinization events share
681 higher FME contents.

682

683 In order to explain the observed variation of serpentine trace element concentrations,
684 selected mineral analyses are compared with literature data on related bulk rock and
685 fluid compositions (Fig. 8). Concentrations of Cs and Rb (Fig. 8A) from the late
686 serpentine type (iii) and type (iv) are similar to the Cs/Rb values reported by Savov et
687 al. (2007, and 2005) for the mud matrix. In contrast, products of all earlier veining
688 and serpentinization events yield higher Cs and Rb values, which are similar to the
689 variation of bulk compositions reported by Savov and coworkers for the serpentinized
690 clasts. The same systematic difference can be seen in the concentrations of Sr and
691 Ba (Fig. 8B). Trace element contents of sediments (from Plank and Langmuir, 1998)

692 and variously altered oceanic crust (AOC; Kelley et al., 2006) as potential source
693 rocks of the serpentinizing fluids are also plotted in Fig. 8, together with analyses of
694 "fresh peridotite" from Meffré et al., (2012).

695

696 Rb, Cs, Sr, and Ba concentrations in sediments are higher than in both mud matrix
697 and serpentine veins in the clasts. Altered oceanic crust (AOC) may have higher Sr
698 and Ba contents than either mud matrix or clast serpentine, and may show similar Cs
699 contents as the mud matrix and serpentine in clasts, depending on the alteration
700 stage. However, AOC contains slightly less Cs than the early clast serpentine
701 generations. In contrast, Rb concentrations of strongly altered oceanic crust may be
702 higher than those of mud and late clast serpentine, and even higher than those of
703 early clast serpentine generations. Thus, both sediments and altered ocean crust
704 could serve as source for higher Sr and Ba values in the serpentinizing fluid, but only
705 sediments and very highly altered oceanic crust can be considered to discharge high
706 Cs,Rb fluids. The main sources of fluid to be released in a low-T environment such
707 as the forearc region of a cold subduction zone are thought to be related to sediment
708 compaction and from sediment dehydration reactions involving opal (at 30 - 80 °C)
709 and smectite (120 - 150 °C; see Hyndman and Peacock , 2003; Fryer and Salisbury,
710 2006), with little or no participation of AOC as fluid source.

711

712 The mobilization of volatiles (B and NH₄) and incompatible elements (As, Be, Cs, Li,
713 Pb, Rb) in hydrothermal fluids at relatively low temperatures (~ 300 °C) was
714 demonstrated by You et al. (1996) in laboratory hydrothermal autoclave experiments
715 with sediments of the decollement zone at Site 808 in Nankai Trough, Japan. The
716 absolute concentrations of Cs, Rb, Ba, and Sr observed in serpentine in this study,
717 indicate that fluids derived from sediment dehydration may account for high trace
718 element contents in serpentine formed during the pervasive stage and the early vein
719 generations. Regardless of the absolute Cs and Rb contents of lizardite and

720 chrysotile, the observed Rb/Cs ratios are fairly uniform within the different serpentine
721 generations and average between 1 and 10 (average Rb/Cs ratios in early
722 serpentines (mesh, bastites, vein types I and II) are between 2-4, associated with
723 high absolute elemental concentrations; late serpentine exhibits averaged Rb/Cs
724 ratios of 7 (type III) and 8 (type IV), associated with low absolute elemental
725 concentrations; Fig. 8A). It is possible that subducted sediments containing a high
726 fraction of pelagic clays release fluids high in Rb and Cs, and account for the high
727 trace element contents of serpentine formed in the pervasive stage and the early
728 vein generations. The uniformly low Rb/Cs ratios ≤ 10 observed in the serpentine
729 minerals from Site 1200A are consistent with a sedimentary origin of the
730 serpentinizing fluids, which were released at low to moderate temperatures (50 - 300
731 °C), indicating a higher mobility of Cs relative to Rb.

732

733 The systematic variation of trace element contents of pore fluid compositions of
734 several seamounts with distance from the Mariana trench, revealed in geochemical
735 studies (Fryer et al., 1999; Mottl et al., 2003, 2004; Hulme et al., 2010), support the
736 hypothesis of sediments as a main source of slab fluids in the forearc region. Figures
737 8A and 8B shows the trace element contents of the pore fluids from South Chamorro
738 and other Mariana forearc seamounts as summarized in Mottl et al. (2004). Water
739 compositions from the following seamounts, in the order of increasing distance from
740 the trench, are also plotted: Bottom seawater (distance: 0 km), Pacman Baseball Mitt
741 (52 km), Blue Moon (55 km), Pacman summit (64 km), Big Blue (70 km), S.
742 Chamorro S351 spring (85 km), Seamount Site 1200 (85 km), and Conical Site 780
743 (90 km). The observed concentrations are consistent with increased mobilization of
744 trace elements from sediments with increasing temperature and depth of slab, as
745 predicted for cold subduction zones (Hyndman and Peacock, 2003; Plank and
746 Langmuir, 1998; Fryer, 2012, and references therein). The accuracy of these

747 predictions is also supported by the similarity of the fluids in terms of Rb and Cs (Fig.
748 8A) observed in the experimental charges of You et al. (1996). The Ba concentration
749 might be influenced by complex interactions with sulphate (Fig. 8B). Admittedly,
750 although the fluids discharged in experimental runs document the mobilization of
751 trace elements from solids into the fluid phase, the fractionation of trace elements
752 into the secondary minerals during serpentinization is unknown. However, as
753 indicated by the similarity in trace element contents of late-stage vein and mud
754 serpentine, the fluids constituting the mud might be quite similar to those of the late
755 stages of serpentinization. A caveat of this interpretation is that the different fluid
756 advection velocities in the different seamounts and different extents of seawater
757 entrainment and mixing/precipitation (ie. of aragonite, affecting Sr concentrations)
758 cannot be considered.

759

760 South Chamorro's pore water samples document the accumulated trace elements in
761 the fluid present in the subduction channel beneath the mud volcano, mobilized at all
762 temperatures and depths prevailing between trench and conduit. The multiphase
763 serpentinization events observed in the erupted clasts possibly also document the
764 release of serpentinizing fluids at different levels of the decollement: both microfabric
765 observations on the clasts and trace element analyses of the solids erupted by the
766 forearc mud volcanoes indicate a change in the fluid regime. On the one hand,
767 serpentines of the pervasive serpentinization stage and early vein generations
768 formed under low water/rock ratios and are characterized by high trace element
769 contents and Rb/Cs between 1 and 4. In contrast, serpentines in late extensional
770 veins and in the mud formed under higher water/rock ratios, and exhibit still high, but
771 slightly reduced trace element contents with a slightly increased Rb/Cs ratio near 10.
772 Thus, the increasing Rb/Cs ratio observed in successive serpentinization events
773 track the transformation of the compositional pattern of the slab-sourced fluids, being
774 progressively released and accumulated during continued subduction (e.g., Bebout

775 et al., 2007). In this context, late vein serpentine and mud would represent
776 serpentinization in greater depth nearer to the conduit, while serpentinization of the
777 pervasive stage and the early vein generations might have taken place in shallower
778 parts nearer to the trench. This is consistent with several studies of the Mariana
779 mudflow systems, that have mentioned the possibility that serpentinites formed at
780 shallower levels of the decollement might been carried in deeper parts of the
781 subduction zone, where they would contribute to magma genesis at the arc and
782 backarc regions (Kerrick and Connolly 2001; Savov et al., 2005b, 2007; Snyder et al.,
783 2005; Wei et al., 2005). Thus, clasts and matrix erupted at the mud volcanoes
784 represent a snapshot of these process.

785

786

787 *5.4 Generic model for the polyphase serpentinization history of Mariana forearc* 788 *mantle clasts*

789

790 The mantle wedge of a cold oceanic subduction zone imposes distinct boundary
791 conditions for fluid migration and mineral-fluid interaction at shallow levels. While the
792 fresh and cold mantle of the tapering wedge would not likely allow mantle convection,
793 the rheology of the serpentinized ultramafic rocks may confine the passage of fluids
794 to fracture bound focused flow, localized around faults or conduits (Fryer, 2012)
795 feeding the serpentine mud volcanoes. Concerning the interaction between mantle
796 and subducting plate, the location of the decoupling-coupling transition (Furukawa,
797 1993) is of fundamental importance for the advance of serpentinizing reactions in the
798 Mariana forearc mantle. Hyndman and Peacock (2003) propose a thermal structure
799 of a cold oceanic subduction zone (Izu-Bonin), with temperatures in the supra-
800 subduction mantle well below 300 °C. These authors suggest that a decreased
801 coupling between the subducting plate and mantle wedge is due to serpentinization
802 at the base of the mantle wedge seaward of the arc. Hyndman and Peacock (2003)

803 furthermore propose that the weak rheology and positive buoyancy of serpentinite
804 will act to isolate hydrated forearc mantle from the overall mantle wedge flow system.
805 By focusing shear deformation in the weak, serpentinitized mantle adjacent to the
806 subducting plate, suprasubduction mantle and downgoing slab are mechanically
807 decoupled at the plate interface. This idea is corroborated by findings of simple-shear
808 experiments at a temperature of 200 °C and a pressure of 1.0 GPa (Hirauchi et al.,
809 2010), indicating that the dominant flow mechanism of mesh-textured serpentine
810 (lizardite and chrysotile) evolves with increasing shear strain from inhomogeneous
811 semi-brittle flow to homogeneous ductile flow, rather than faulting, brecciation, or
812 fracturing. Moreover, the confirmation of plastic flow by intracrystalline deformation
813 within individual serpentine grains by Hirauchi et al. (2010) is consistent with textural
814 features observed in highly deformed natural serpentinites (i.e., ribbon texture; Viti
815 and Mellini, 1998; Hirauchi and Yamaguchi, 2007). More recently, deformation
816 experiments on lizardite (in the range of 150-400 °C, 1-8 GPa at varying strain rates)
817 and combined first-principles calculations of Amiguet et al. (2012) revealed strain-
818 rate independent plasticity of lizardite at low stress. Even at the potentially high strain
819 rates prevalent at the shallow mantle wedge (10-20 km depth) of a cold intra-oceanic
820 subduction zone, lizardite is capable of a strain localization into cm-scale shear
821 bands, which provides a plausible mechanism for mobilization of rock fragments in
822 shallower parts of the decollement (Hirauchi et al., 2010).

823

824 Two source regions of ultramafic material, mobilized as clasts and erupted in mud
825 flows, seem possible: (i) serpentinitization of conduit wall rock and mobilization as
826 clasts in the mud flow, and (ii) clasts of material from the basal plane of the
827 suprasubduction mantle wedge, partly serpentinitized and eroded in shallower parts of
828 the decollement. In scenario (i), fluids serpentinitizing the conduit wall rocks would
829 derive from the rising mud. The FME-enrichments characteristic of the early phase of
830 serpentinitization would imply a temporal change of the FME-contents of these mud

831 fluids that passing up the conduit, because none of the later generations of
832 serpentinite show this enrichment. Such an oscillation of the degree of FME-
833 enrichment in the mud rising through the conduit systems must have produced
834 likewise clasts with an FME-enrichment pattern reversed to what is actually
835 observed: low-FME concentrations in an early pervasive phase, and high-FME
836 concentrations in late vein stages, and large amounts of high-FME mud serpentine.

837

838 However, bulk analyses from clast and mud material (Savov et al., 2005a and 2007)
839 and the analyses of individual serpentine generations in this work uniformly show
840 higher FME in the clast serpentines than in the mud. It is not likely that the mantle
841 rocks forming the conduit walls beneath South Chamorro Seamount first were
842 serpentinitized pervasively by fluids emanating from the slab-mantle interface with
843 high-FME concentrations, which then shift to lower FME-concentrations in the later
844 stage of mobilization and massive mud rise. This shift in the trace element signature
845 of the serpentinitizing fluid is more plausibly explained by the second scenario outlined
846 above. Mantle material of the basal plane of the suprasubduction mantle wedge is
847 partly serpentinitized by fluids derived from the slab sediments in shallower parts of
848 the decollement. These fluids have uniformly high-FME contents which is inherited
849 by the rocks altered during these early serpentinitization stages. During descent to the
850 depth of the conduit beneath South Chamorro Seamount, the continued release of
851 fluids from the slab might alter the level of FME-enrichment of the fluid present in the
852 subduction channel that likewise govern the serpentinitization of the later clast vein
853 generations and the mud matrix serpentine.

854

855

856 *5.5 Summarized microfabric evolution of the polyphase serpentinitization of Mariana*
857 *forearc mantle clasts*

858

859 Figure 9 presents a schematic view of the relative chronology of the microfabric
860 evolution, summarizing the observed fabric relations of the polyphase
861 serpentinization history of Mariana forearc mantle. The first serpentinization event is
862 believed to have occurred along a discrete set of tiny fractures (mostly < 10 mm)
863 activated by regional brittle deformation (Salisbury et al., 2002), along which
864 formation of magnetite trails can be observed (see Fig. 9A, early magnetite veins).
865 This event was followed by a stage of pervasive serpentinization with low water-rock
866 ratios, which proceeded under a static, isotropic pressure regime, which led to the
867 development of mesh structures and bastitization of orthopyroxene (Fig. 9B). The
868 serpentine formed during this stage (lizardite) has likely been exposed to a relatively
869 unaltered sediment-derived fluid derived from the sedimentary covering the
870 subducting slab, which contained high concentrations of Cs, Rb, Sr and Ba.

871

872 The pervasive phase was followed by a phase of the establishment of fluid pathways,
873 generating focused fluid flow along reactivated early magnetite veins, or newly
874 developed type (i) veins, involving occasional isovolumetric metasomatism (SiO_2 -loss
875 and halo formation; see Fig. 9C). Metasomatism did not alter the trace element
876 contents present in the serpentine component in Srp-Brc regions. Later
877 serpentinization events document the loss of confining rock pressure and the onset
878 of volumetric expansion (see Fig. 9D) and the shift to higher water-rock ratios,
879 precluding the change from a rock-dominated to a fluid-dominated system:
880 extensional chrysotile veins of type (ii) contain a significant Brc fraction (see Fig. 4C)
881 with intermediate to low Cs and Rb contents, and high Ba and Sr content which are
882 comparable to serpentine of earlier phases. Most probably, the mobilization of wall
883 rock fragments into the mudflow (either in subduction channel or within conduit)
884 involves a transition from a closed system to an open system by development of fluid
885 pathways, similar to the sequence of serpentinization during unroofing of mantle in
886 slow-spreading mid-ocean ridge environments (Andreani et al., 2007).

887

888 Further expansion of the rock in the course of continued serpentinization could not
889 be followed by the already completely serpentinized halo at the rim, and this lead to
890 the formation of type (iii) "Frankenstein" veins (see Fig. 9E; Salisbury et al., 2002;
891 see also Iyer et al., 2008b). The serpentine (chrysotile) formed in this phase, and the
892 multiple crack-seal chrysotile veins of type (iv), which document marked anisotropic
893 expansion (see Fig. 9F), exhibit low trace Cs, Rb, Ba, and Sr contents which are
894 comparable to those of bulk mud analyses of Savov et al., (2005), and to spring or
895 pore water compositions of Site 1200. The type (iv) veins we could observe in our
896 samples are located in the rim regions of clast samples 3R-1_3-7, 16R-1_52-54, and
897 16R-2_7-9. The serpentine formed in both of the later vein generations documents a
898 further change in the serpentinizing fluids, suggesting a diminished influence of
899 solely sediment-derived fluids.

900

901 Microfabric information, phase relations (occurrence of the serpentine polymorphs
902 lizardite and chrysotile; presence or absence of brucite) and the correlation of FME-
903 concentrations of the successive serpentinization stages of the Mariana forearc
904 mantle clasts imply some constraints on the nature of the serpentinizing fluids. The
905 absence of brucite in the early serpentinization phase may be explained by the
906 quartz-saturated nature of unaltered sediment-derived fluids. Opal and smectite
907 dehydration produce high-FME (with Rb/Cs < 10; You et al., 1996), quartz-saturated
908 fluids that may initiate early pervasive serpentinization processes at shallow levels.
909 During the continued serpentinization in a still rock-dominated environment, brucite
910 becomes more common, because the high-FME-fluids are either consumed or
911 depleted by early pervasive serpentinization. Eventually, late chrysotile veins (types
912 iii and iv, in which brucite is absent) document the increased access of slab fluids in
913 a now fluid-dominated serpentinization regime. These fluids are still quartz saturated,
914 but geochemical slab signatures in trace elements are lacking, because the

915 composition of slab-derived fluids has been completely reset by serpentization
916 reactions and continued diagenetic to metamorphic processes within the subduction
917 channel (Moore and Saffer, 2001). The final stage of diapiric rise of matrix and clasts
918 in the conduits feeding the mud volcanoes is characterized by a very late alteration
919 feature (see Fig. 2H, or also Fig. F78 in Salisbury et al., 2002). Apparently, the
920 "browning" of the clasts from the rim inward (independent of the intra-clast fabric
921 relations) corresponds to re-equilibration with the rising mud fluids that lack buffering
922 by quartz-rich lithologies of the descending slab.

923

924

925 **6. Conclusions**

926

927 Observations on microfabric, petrography and trace element contents of
928 serpentized ultramafic clasts from South Chamorro Seamount provide novel
929 insights into the prevailing conditions in the shallow mantle beneath the Mariana
930 forearc. Detailed analyses of phase relations (low-temperature serpentine
931 polymorphs lizardite and chrysotile, absence of antigorite and magnetite) and mineral
932 chemistry of the successive serpentine generations present in the clasts, as well as
933 oxygen isotope data, reveal low-temperature conditions (≤ 250 °C), as proposed by
934 thermal models of cold oceanic subduction zones. Correlation of trace element
935 concentrations and ratios in serpentine from clasts and mud matrix with either mud-
936 volcano pore water analyses or experimental investigations of trace element
937 mobilization reveal fluids derived from sediments at low temperatures as potential
938 source of the serpentizing fluids.

939

940 The rheological properties of suprasubduction mantle material in a cold oceanic
941 subduction zone pose distinct boundary conditions on the serpentization

942 processes: (i) convection is absent in the tapering mantle wedge above the
943 decoupling-coupling transition; (ii) within the wedge, and likely on the interface
944 between subducting slab and overlying mantle, lizardite is capable of a strain
945 localization into sample-scale shear bands. Therefore, within the wedge, the passage
946 of fluids may be confined to fractures, which is the prerequisite for the transfer of
947 serpentized bouyant material to rise from the slab-mantle interface via conduits to
948 the mounds of the mud volcanoes. Moreover, due to the characteristic mechanical
949 properties of the low-temperature serpentine phases, partly serpentized mantle
950 material can be eroded at shallow levels of the decollement from the basal interface
951 downward in the subduction channel, and still preserve fabric records and chemical
952 signatures of the early stages of serpentization.

953

954 High concentrations of fluid mobile elements in the early stages of the polyphase
955 serpentization history of the clasts, in contrast to FME-poor serpentine present in
956 the late vein generations in these clasts and also in the mud matrix serpentine point
957 to a change in the nature of the serpentizing fluids throughout the history of
958 serpentization and eruption. We propose, that the continued microfabric and
959 mineral chemical evolution observed in the ultramafic clasts is consistent with (1) the
960 partial serpentization and erosion of mantle material off the basal plane of the
961 suprasubduction mantle wedge in shallower parts of the decollement, (2) the
962 subsequent descent in the subduction channel to the depth beneath South Chamorro
963 Seamount, and (3) the rise as clasts through the conduit towards the mound of the
964 mud volcano.

965

966 Sediments seem to play a vital role not only in the explanation of the observed trace
967 element ratios in the actual samples, but potentially also for the generation of fluids
968 under low-T conditions, and thus promoting the formation of mud volcanos in forearc
969 regions of nonaccretionary cold subduction zones. With this respect, additional

970 experimental data on the mobilization of FME from sediments are essential,
971 especially on the partition coefficients of these with the different serpentine
972 polymorphs. Moreover, research on the occurrence of certain serpentine polymorphs
973 in different serpentinization microfabrics may allow for more detailed mass (and
974 volume) balanced estimations. This will further the understanding of how subduction-
975 zone processes are connected to the hydration of ultramafic rocks.

976

977

978

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987

988

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1372

1373 **Figure Captions**

1374 Fig. 1

1375 (A) Bathymetric map of the Mariana arc-basin system showing the location of South
1376 Chamorro Seamount, 85 km from the trench (Leg 195, Site 1200A). Regional map
1377 prepared using GMT (Wessel and Smith, 1991). Elevation data from the
1378 SRTM30_PLUS dataset (Becker et al., 2009). (B) Schematic diagram (no vertical
1379 exaggeration) of the emplacement of serpentinite mud volcanoes on the outer
1380 Mariana Forearc (modified from Oakley et al., 2007). The asterisk marks the relative
1381 position of South Chamorro Seamount.

1382

1383 Fig. 2

1384 Thin section photomicrographs of serpentinitized ultramafic clasts documenting the
1385 polyphase serpentinitization history of the Mariana forearc mantle. (A) Sample 12R-1,
1386 8-11 cm, showing early magnetite-bearing veins predating the pervasive stage,
1387 developed linearly without marked conformance with the rock fabric. Back-scattered
1388 electron image. (B) Expansional chrysotile type (ii) veins (brown) in sample 16R-2,
1389 11-12 cm. Scanned thin section, plane-polarized light, image width 16 mm. (C)
1390 Close-up of type (ii) veins in sample 16R-2, 7-9 cm, crossing pervasively
1391 serpentinitized harzburgite with relict olivine subgrains. Photomicrograph, cross-
1392 polarized light, image width 1.7 mm. (D1 and D2) Type (iii) veins ("Frankenstein
1393 veins"), developed in the halo of type (i) veins and crosscutting type (ii) veins in
1394 sample 7R-2, 40-43 cm. Scanned thin section, plane- and cross-polarized,
1395 respectively Image width D1 is 11 mm. (E) Type (iii) veins in the halo of type (i) veins
1396 in sample 13R-2, 5-7 cm. Scanned thin section, cross-polarized light. Image height
1397 10 mm. (F) Late-stage type (iv) veins of the multiple crack-seal type, indicating
1398 clearly anisotropic expansion in sample 3R-1, 3-7 cm. Scanned thin section, plane-
1399 polarized light, image width 10.5 mm. (G) Dunite sample 9R-1, 21-23 cm. Ribbons
1400 contain submicroscopic intergrowth of serpentine and brucite. On the contrary, large,

1401 recrystallized brucite is observed in regions separated by the ribbons. Image height
1402 2.2 mm. (H) Samples 16R-1, 52-54 cm, and 16R-1, 62-64 cm, both from piece 11A,
1403 show browning, occurring from the clast's rim inward, which involves brucite formation
1404 in interstices. Scanned thin section, plane-polarized light. Image width 8 mm.

1405

1406 Fig. 3

1407 Detailed view of pervasive serpentinization fabrics. (A) and (B) Sample 7R-2, 40-43
1408 cm. Mesh textures developed in the forearc setting lack the formation of magnetite
1409 along mesh rims. Instead, awaruite and submicroscopic intergrowth of brucite and
1410 serpentine is observed. (C) and (D) Qualitative analyses of the same sample,
1411 indicating the absence of brucite rims around olivine cores.

1412

1413 Fig. 4

1414 Compositions of primary and secondary minerals of the different serpentinization
1415 events, plotted as molar proportions of Si-Mg-Fe in ternary diagrams. Primary
1416 minerals are given for orientation, they are not part of the assemblage. (A1) and (A2)
1417 Serpentine analyses of the pervasive stage in harzburgites and dunites. In dunites,
1418 the analyses low in Si are from recrystallized hourglass textures. (B) Analyses of
1419 serpentine-brucite phases from type (i) veins and the adjacent halo region. (C)
1420 Composition of serpentine-brucite, submicroscopically intergrown in expansional type
1421 (ii) veins in harzburgite. (D) Serpentine analyses of type (iii) and type (iv) veins,
1422 indicating anisotropic expansion in the course of the serpentinization process.

1423

1424 Fig. 5

1425 Variations of SiO_2 , Al_2O_3 , CaO and Cr_2O_3 content in primary and secondary minerals.
1426 (A) SiO_2 wt.% vs. Al_2O_3 wt.%. Serpentine grown after pyroxene shows Al_2O_3 as high
1427 as the precursor mineral, whereas serpentine grown in later events or after olivine
1428 shows low to intermediate Al contents. (B) In terms of CaO , bastites show similar

1429 contents as orthopyroxene. On the contrary, serpentine grown from olivine or during
1430 later events is enriched in Ca relative to olivine. (C) Cr_2O_3 content of bastites overlap
1431 with pyroxene Cr contents, whereas serpentine formed from olivine or during later
1432 events contains as low Cr as the primary olivine.

1433

1434 Fig. 6

1435 Trace element (Ni, V) - $\text{Si}/(\text{Mg}+\text{Fe})$ relationships in primary and secondary minerals.
1436 (A) Ni concentrations detected in serpentine-brucite intergrowth regions positively
1437 correlate with the serpentine fraction in the mixed analyses. (B) V is neither found in
1438 serpentine-brucite intergrowth regions, nor in type (ii) veins. (C) The Ni-V variation
1439 diagram shows the discriminatory significance for identifying the serpentine precursor
1440 mineral.

1441

1442

1443 Fig. 7

1444 Concentrations of B, Ba, Sr, Rb, and Cs in serpentine and serpentine-brucite mixed
1445 analyses. (A) to (E) Boron contents in general increase with increasing Srp-fraction in
1446 the mixed Srp-Brc analyses, showing medium to high B values both for early and late
1447 vein generations. Barium and Sr show high contents also in the type (ii) veins, but
1448 not in type (iii) and type (iv) late-stage veins. Cs and Rb exhibit marked enrichment in
1449 early serpentinization stages, and low contents in all types of late extensional veins.
1450 (F) Boron-Cs variation diagram, showing the overall high B contents of the
1451 serpentinization fluids of the various stages, in comparison to Cs in Srp, which
1452 exhibits markedly different concentrations throughout the serpentinization history.

1453

1454 Fig. 8

1455 Diagrams of (A) Cs vs. Rb, and (B) Sr vs. Ba in primary minerals and serpentine of
1456 the various serpentine generations. In both diagrams, fields of bulk rock analyses of

1457 1200A clasts and mud from Savov et al., (2007, and 2005a) are given. For
1458 comparison, Cs and Rb contents of forearc peridotites (Meffré et al., 2012), altered
1459 oceanic crust (Kelley et al., 2006), and of sediments (Plank and Langmuir, 1998) are
1460 shown. Furthermore, compositions of natural (Mottl et al., 2004) and experimental
1461 fluids and solids (Y-You et al., 1996) are plotted. See text for further discussion.

1462

1463 Fig. 9

1464 Relative chronology of the microfabric evolution, summarizing the observed fabric
1465 relations and generalized trace element concentrations of the polyphase
1466 serpentinization history of Mariana forearc mantle. (A) First serpentinization event
1467 along a discrete set of tiny fractures. (B) Phase of pervasive serpentinization under a
1468 static, isotropic pressure regime. (C) Phase of establishment of fluid pathways,
1469 generating focused fluid flow along type (i) veins. (D) Onset of anisotropic expansion
1470 features and occurrence of extensional type (ii) veins. (E) Formation of type (iii) veins
1471 in outer rock/clast (?) halo regions of type (i) veins to compensate continued
1472 expansion of the inner parts. (F) Multiple crack-seal veins of type (iv), located in the
1473 rim regions of the clasts.

1474

1475

1476 **Tables**

1477

1478 Table 1.

1479 Samples (clasts) from South Chamorro Smt., Leg 195, Site 1200A.

1480

1481 Table 2.

1482 Representative electron microprobe analyses and in situ LA-ICP-MS data of relict

1483 primary minerals from Leg 195, Site 1200A, South Chamorro Smt.

1484

1485 Table 3.

1486 Representative electron microprobe analyses and in situ LA-ICP-MS data of

1487 serpentinization products from Leg 195, Site 1200A, South Chamorro Smt.

1488

1489 Table 4.

1490 Overview of selected properties of serpentine from different serpentinization events.

1491

1492 Table 5.

1493 Oxygen isotope data of serpentinization products from different serpentinization

1494 phases in ultramafic clasts from South Chamorro Smt., Leg 195, Site 1200A.

Figure1
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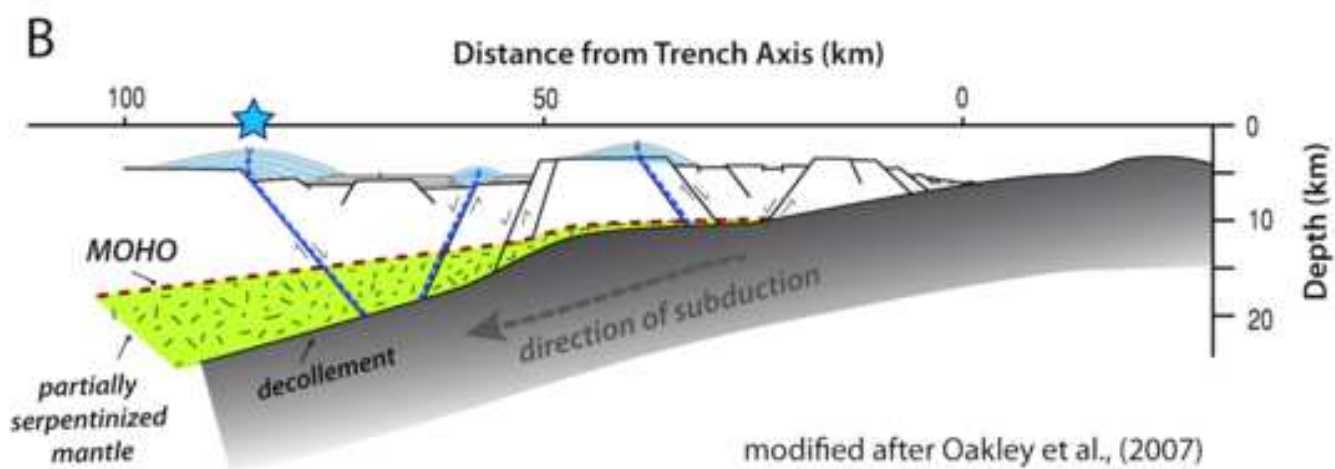
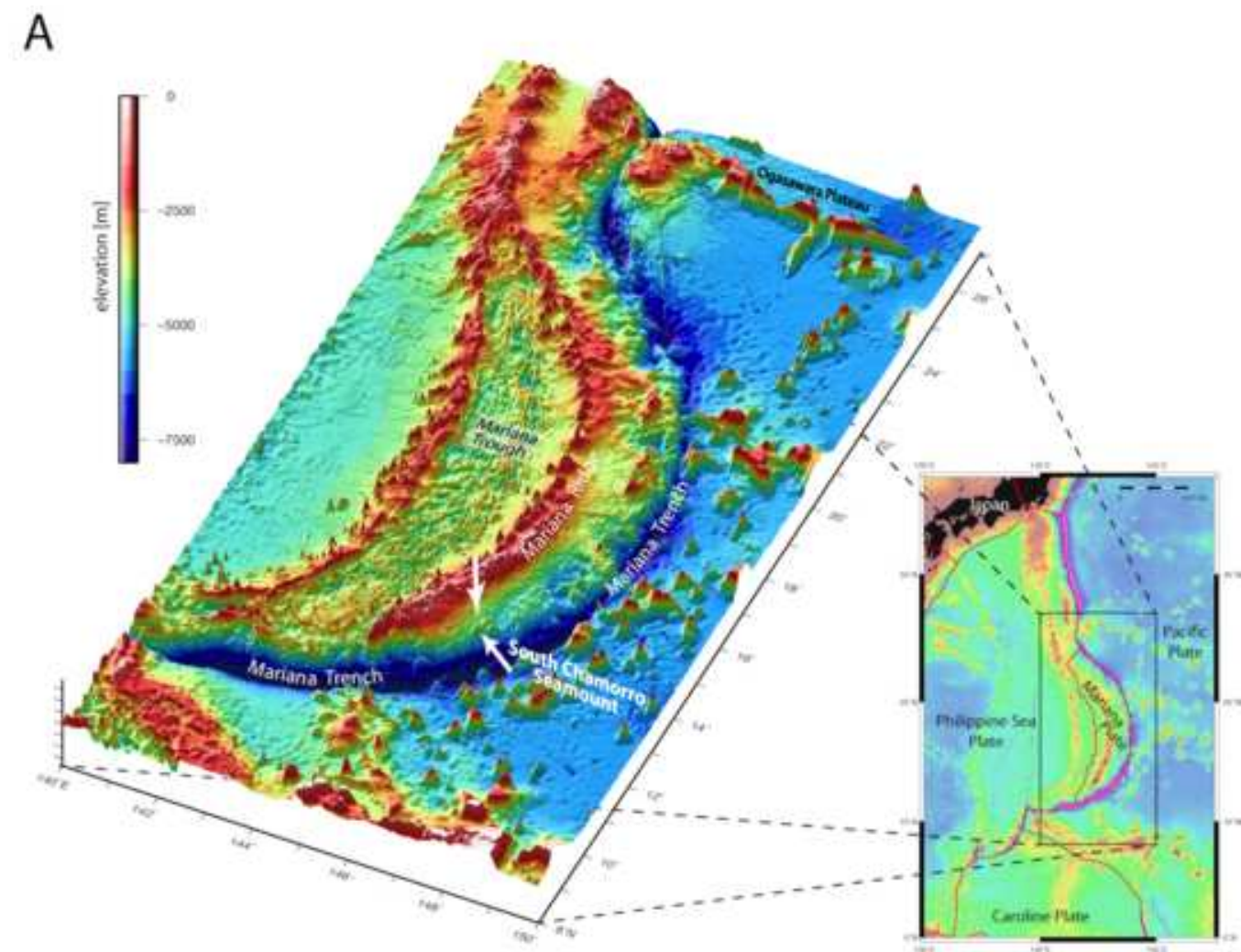


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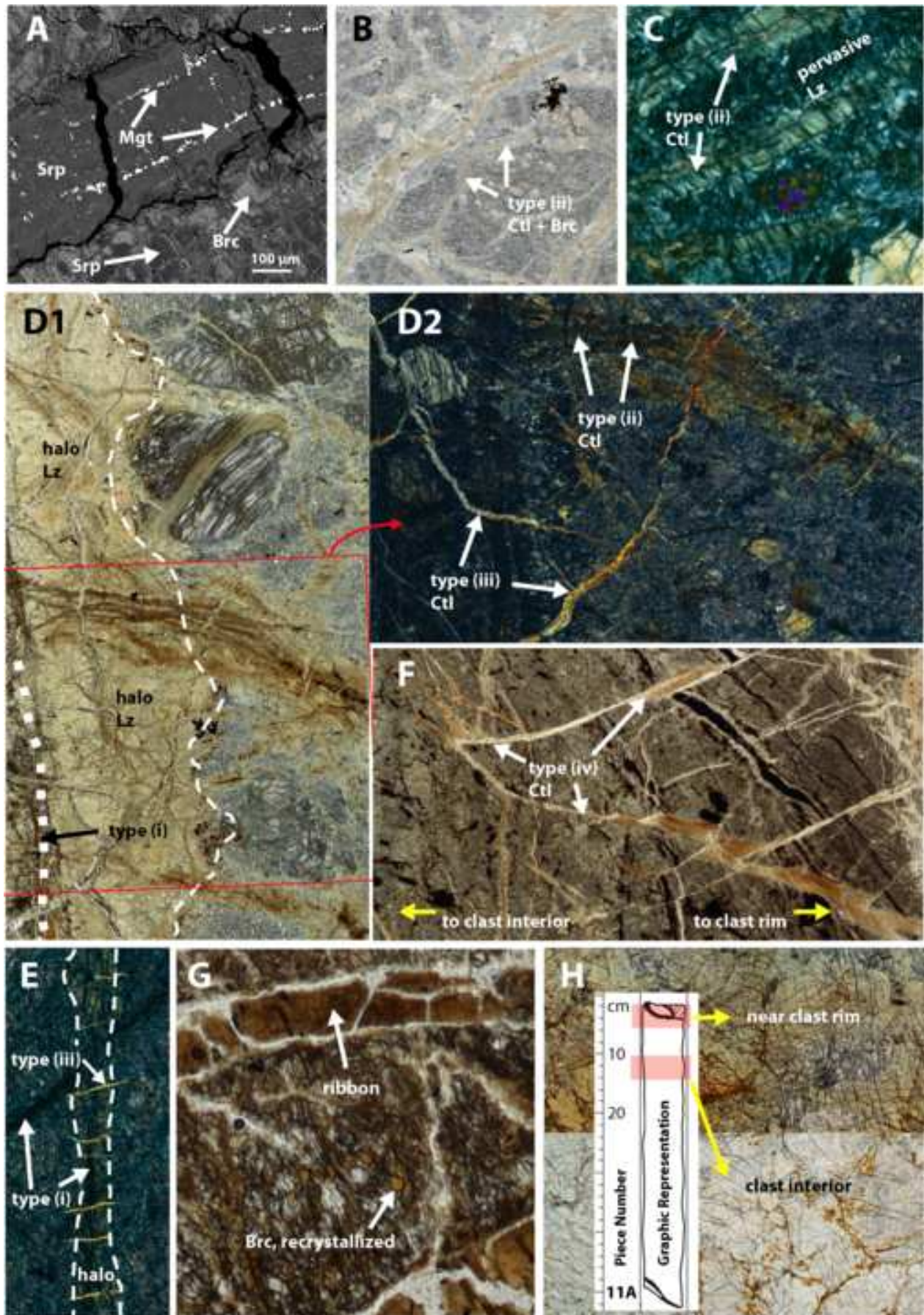


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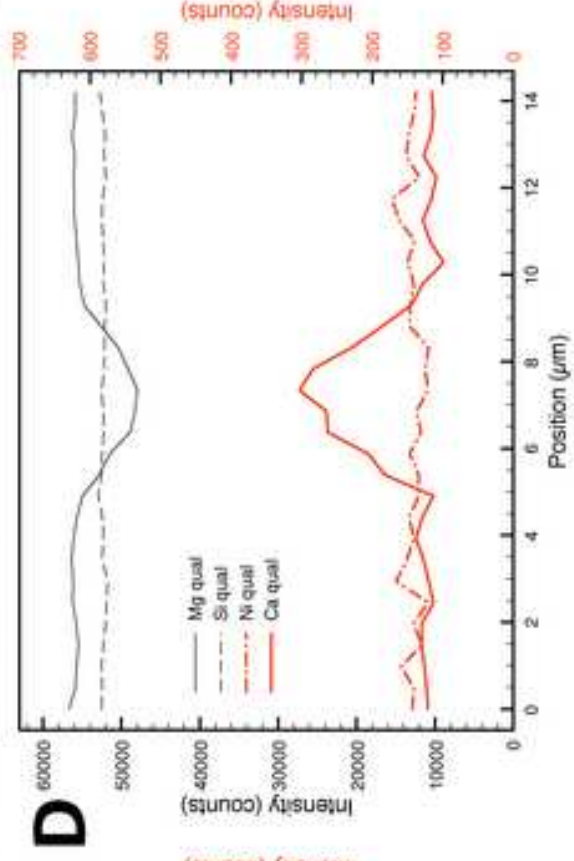
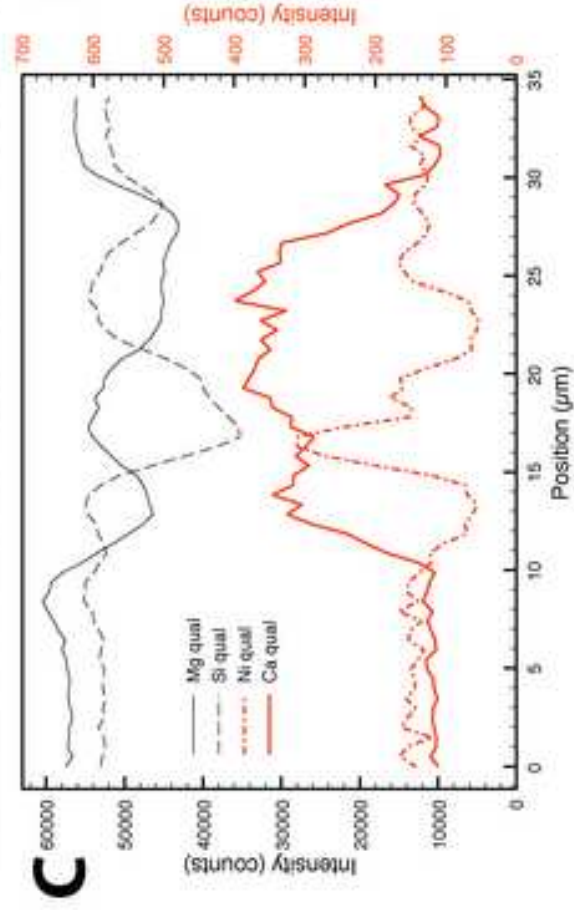
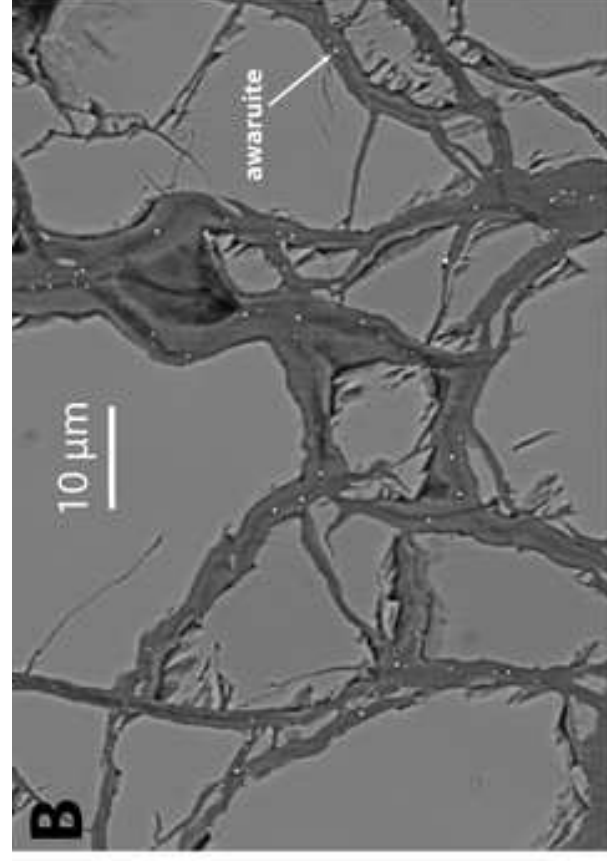
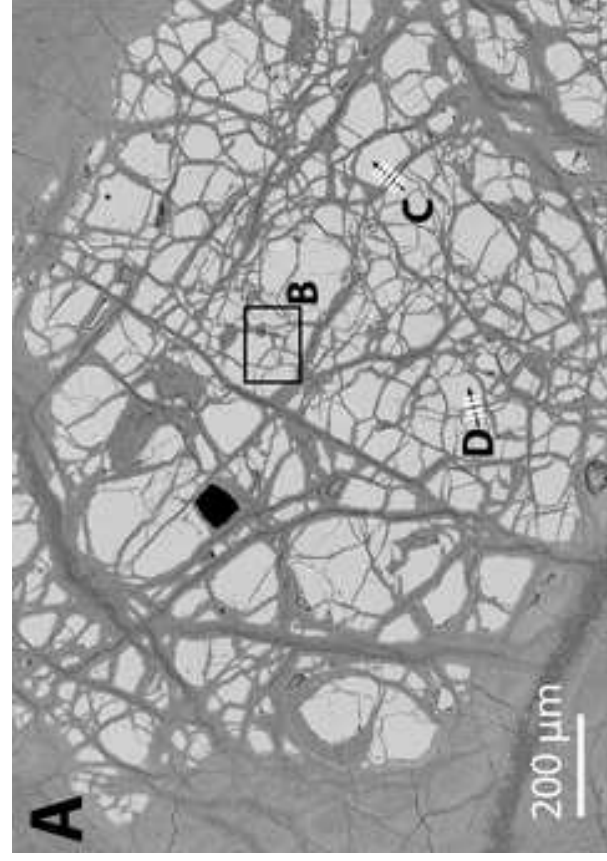


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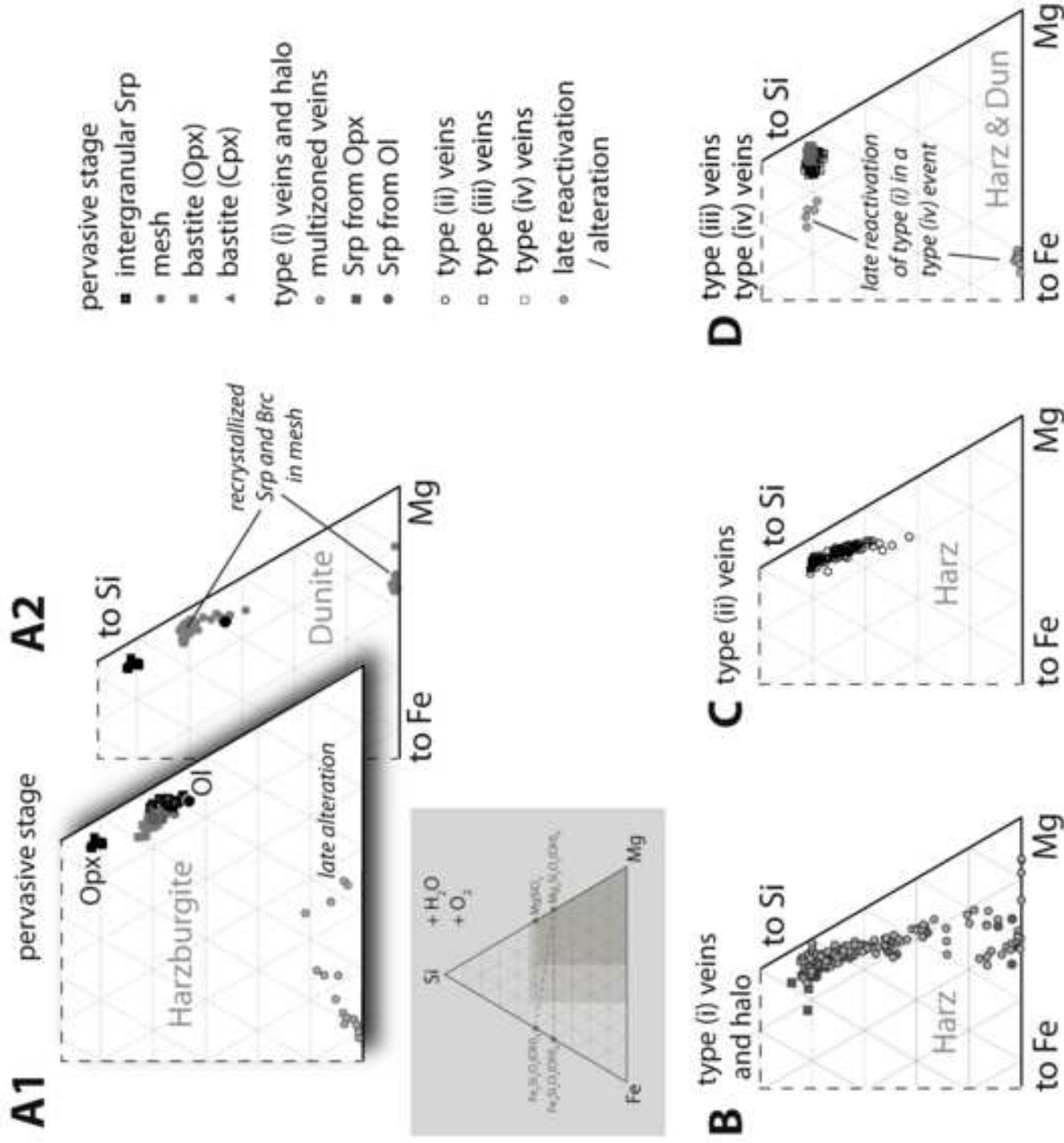


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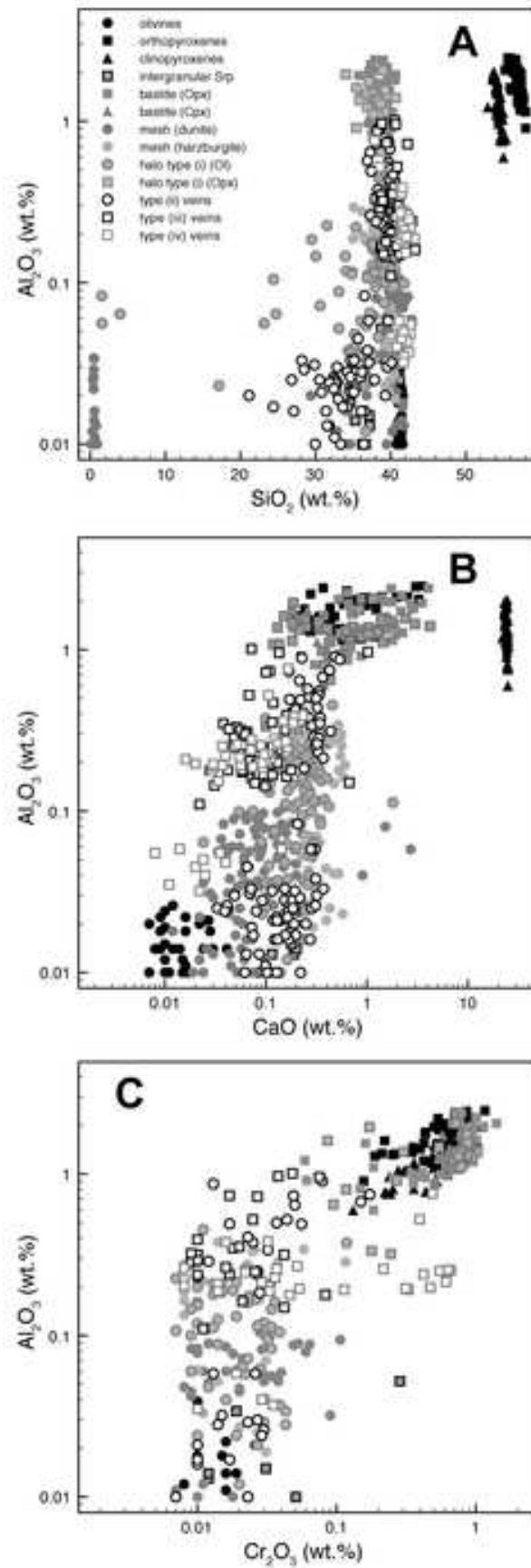


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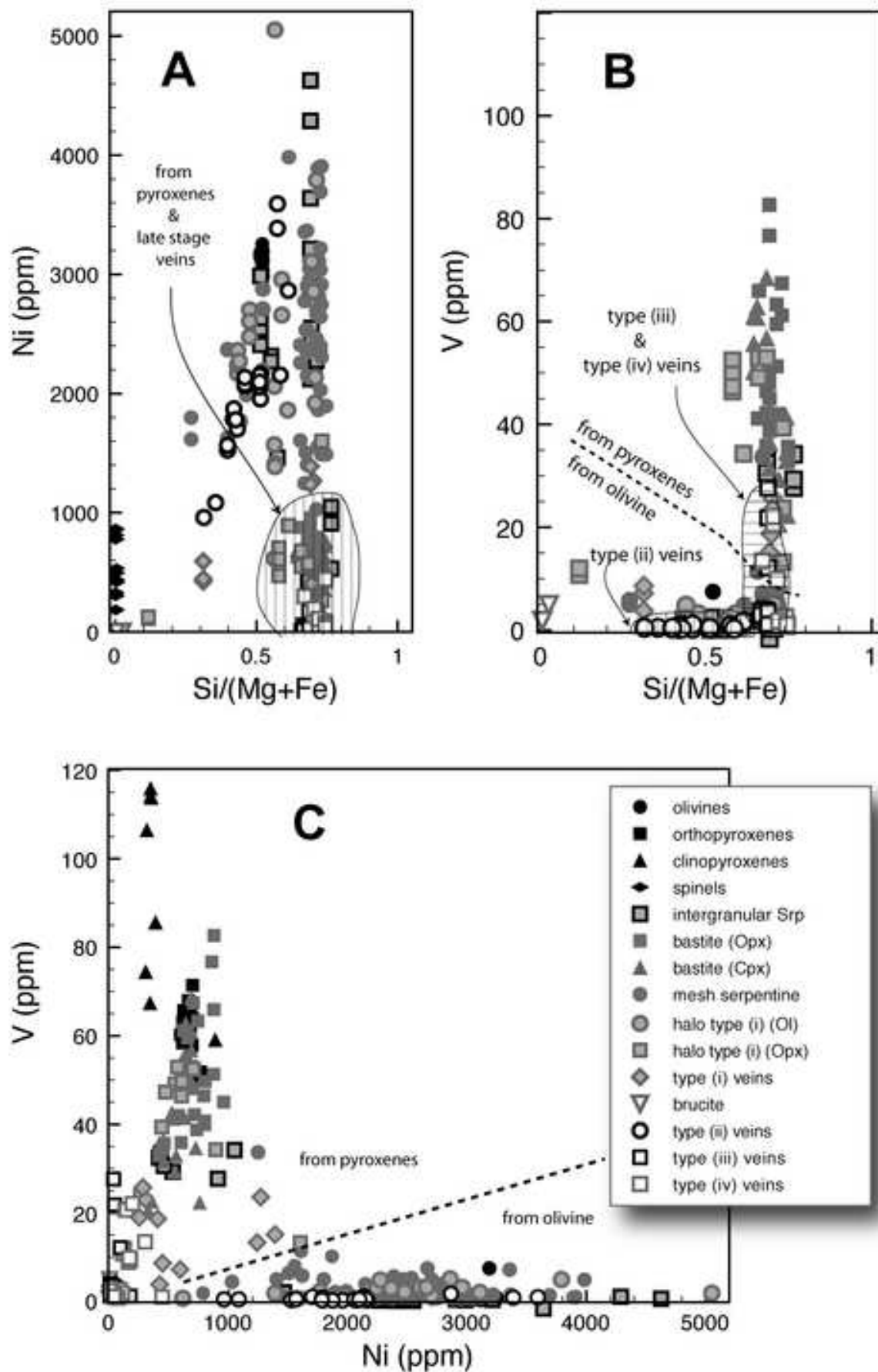


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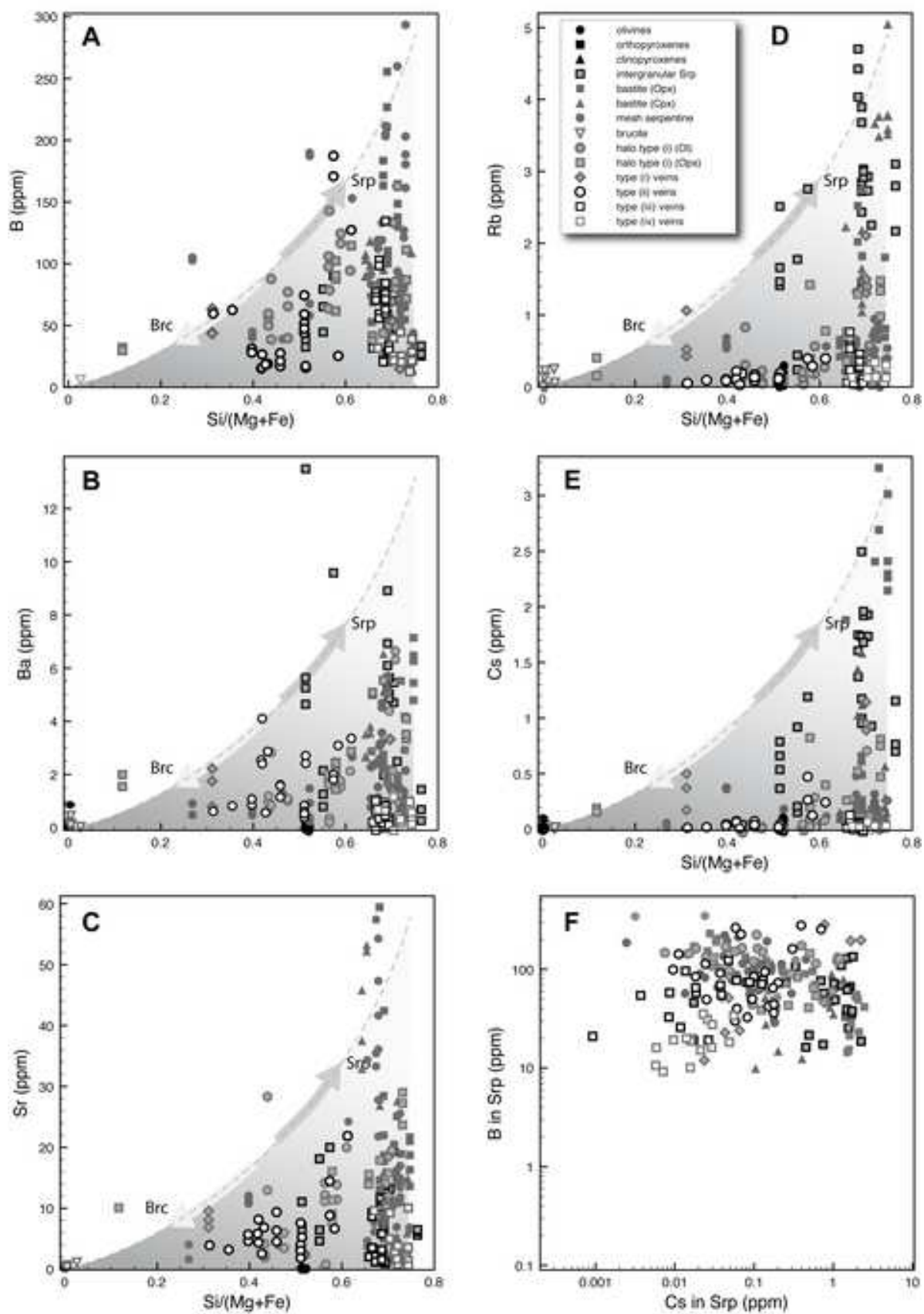
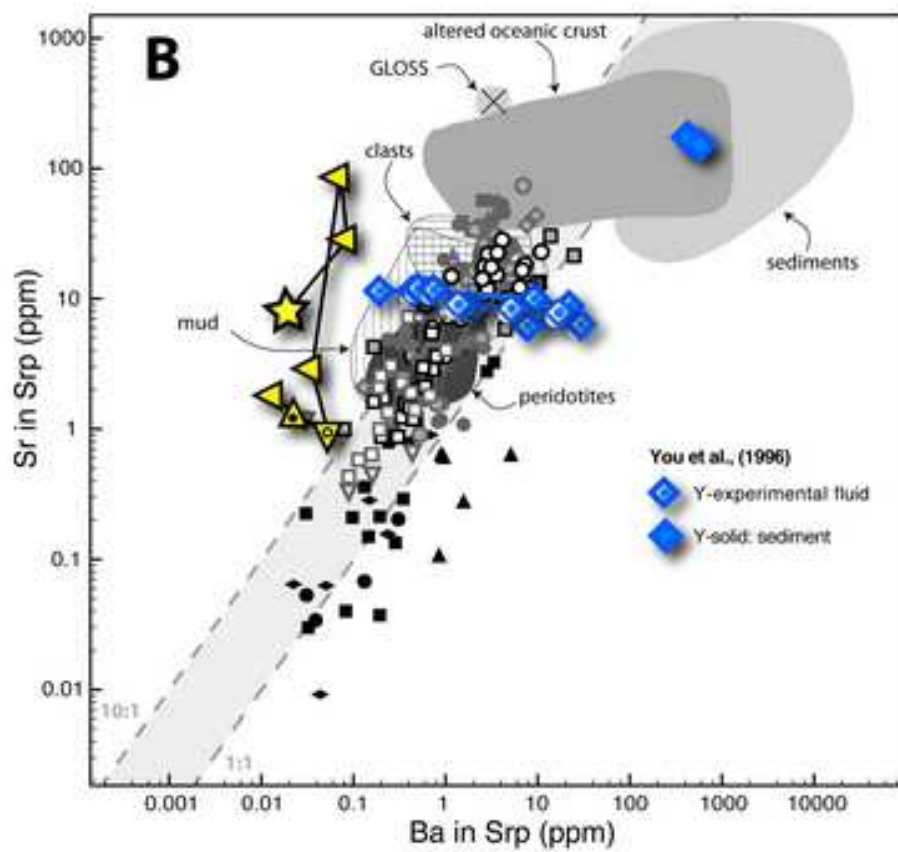
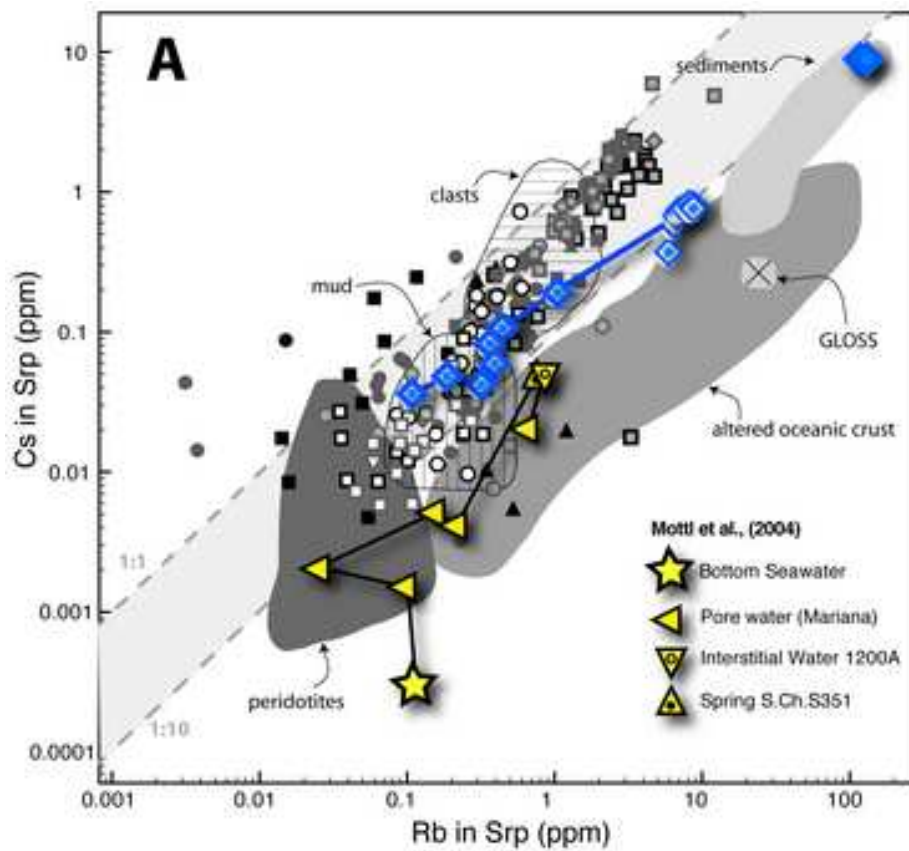
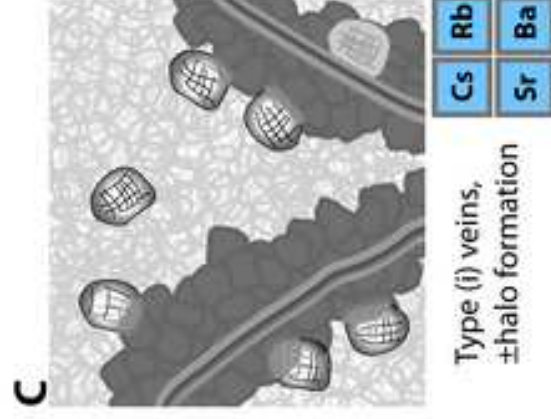
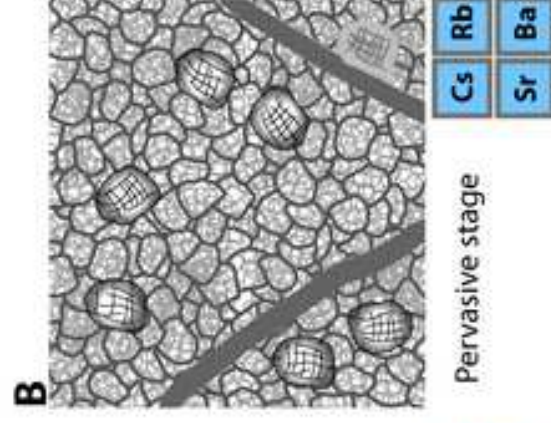
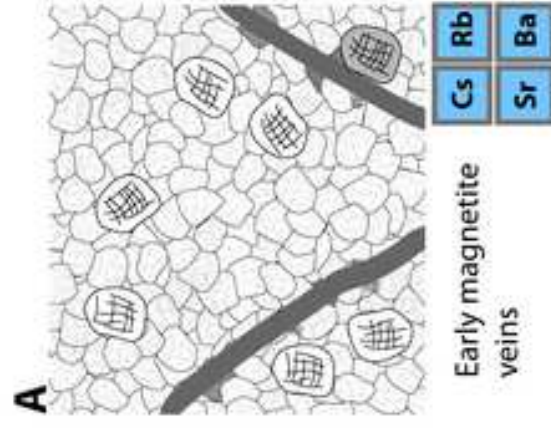


Figure8
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Rock-dominated system - isotropic pressure regime



Change to fluid-dominated system - anisotropic expansion

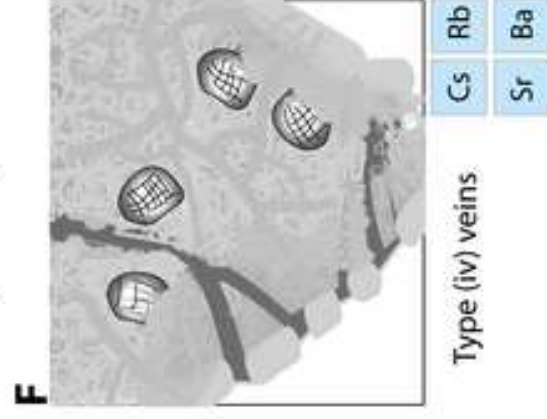
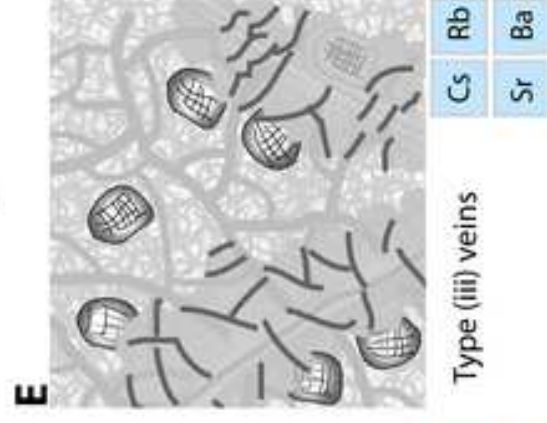
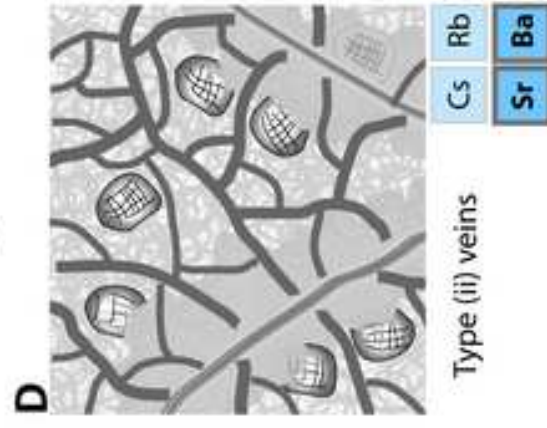


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Table 1. Samples (clasts) from South Chamorro Smt., Leg 195, Site 1200A.

Sample	Short hand	Subseafloor depth (mbsf)	Rock	% alteration
<i>near rim</i>				
1200A-3R-1_3-7	3R-1r	18.3	Dun	99
1200A-7R-2_40-43	7R-2r	52.1	Hz	65
1200A-9R-1_21-23	9R-1r	70.6	Dun	99
1200A-10R-1_25-27	10R-1r	80.2	Hz	95
1200A-11R-1_47-49	11R-1r	89.9	Hz	100
1200A-12R-1_8-10	12R-1r	99.1	Hz	100
1200A-13R-2_5-7	13R-2r	110.0	Hz	80
1200A-16R-1_52-54	16R-1r	138.1	Hz	98
1200A-16R-2_7-9	16R-2r	139.0	Hz	70
<i>far from rim</i>				
1200A-3R-1_15-17	3R-1c	18.35	Hz	99
1200A-7R-2_49-51	7R-2c	52.18	Hz	70
1200A-10R-1_25-27	10R-1c	80.15	hz	97
1200A-11R-1_49-50	11R-1c	89.89	Hz	70
1200A-12R-1_13-15	12R-1c	99.13	Hz	99
1200A-13R-2_15-17	13R-2c	110.1	Hz	85
1200A-16R-1_62-64	16R-1c	138.22	Hz	95
1200A-16R-2_11-12	16R-2c	139.06	Hz	75

Table 2. Representative electron microprobe analyses and in situ LA-ICP-MS data of relict primary minerals from Leg. 195, Site 1200A, South Chamorro Smt.

Sample	7R-2	16R-2	16R-1	7R-2	13R-2	16R-1	7R-2	16R-1	7R-2	16R-1	9R-1	16R-2
Analys. no.	27-25	22-103	94-61	6-10	28-104	20-53	17-35	7-43	156-144	5-100		
Mineral	Ol	Ol	Ol	Opx	Opx	Opx	Opx	Cpx	Cpx	Spl	Spl	
rock type	Hz	Hz	Hz	Hz	Hz	Hz	Hz	Hz	Hz	Dn	Hz	
wt. %												
SiO2	41.12	41.27	41.38	55.63	57.07	57.27	54.65	54.12	0.02	0.02	0.02	0.02
TiO2	0.00	0.00	0.00	0.00	0.02	0.02	0.00	0.03	0.03	0.03	0.01	0.01
Al2O3	0.01	0.00	0.02	2.47	1.50	1.31	1.18	1.18	5.97	5.97	24.79	24.79
Cr2O3	0.00	0.00	0.01	1.16	0.57	0.60	0.36	0.65	55.89	55.89	43.89	43.89
FeO	8.26	7.81	8.00	5.35	5.44	5.34	2.11	1.51	18.81	18.81	19.00	19.00
MnO	0.14	0.13	0.12	0.14	0.17	0.16	0.08	0.08	0.51	0.51	0.23	0.23
MgO	50.47	51.09	50.95	32.05	35.48	34.97	19.97	17.81	7.31	7.31	12.67	12.67
CaO	0.00	0.00	0.03	3.51	0.01	0.02	0.00	21.84	0.31	0.31	0.00	0.00
Na2O	0.00	0.00	0.00	0.00	0.01	0.02	0.00	0.05	0.05	0.00	0.00	0.00
K2O	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.01	0.00	0.00	0.00
Total	100.01	100.30	100.50	100.31	100.57	100.16	100.19	99.70	88.54	88.54	100.61	100.61
Oxygens	4	4	4	6	6	6	6	6	4	4	4	4
Si	1.00	1.00	1.00	1.93	1.95	1.97	1.97	1.97	0.00	0.00	0.00	0.00
Al	0.00	0.00	0.00	0.10	0.06	0.05	0.05	0.05	0.27	0.27	0.88	0.88
Ti	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Cr	0.00	0.00	0.00	0.03	0.02	0.02	0.01	0.02	1.69	1.69	1.05	1.05
Fe3+*	n. c.	n. c.	n. c.	n. c.	n. c.	n. c.	n. c.	n. c.	0.03	0.03	0.06	0.06
Mg	1.83	1.84	1.83	1.66	1.81	1.79	1.07	0.97	0.42	0.42	0.57	0.57
Fe2+	0.17	0.16	0.16	0.16	0.16	0.15	0.06	0.05	0.57	0.57	0.42	0.42
Mn	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.02	0.01	0.01
Ca	0.00	0.00	0.00	0.13	0.01	0.02	0.84	0.94	0.00	0.00	0.00	0.00
Na	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
K	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Sum	3.00	3.00	3.00	4.01	4.01	4.00	4.00	4.00	3.00	3.00	2.99	2.99
Mg/(Mg+Fe2+)	0.92	0.92	0.92	0.91	0.92	0.92	0.94	0.95	0.86	0.86	0.54	0.54
Cr/(Cr+Al)												
B [µg/g]	2.5	0.2	0.1	1.3	1.4	1.1	1.0	3.9	0.5	0.5	1.6	1.6
V [µg/g]	0.5	0.4	0.4	50.9	67.9	63.2	59.1	106.5	1002	1002	763	763
Ni [µg/g]	3169	3030	2983	745	663	616	386	314	188	188	866	866
Rb [µg/g]	0.0	-0.1	0.3	0.0	0.1	0.1	0.5	0.3	n. d.	n. d.	n. d.	n. d.
Sr [µg/g]	0.2	0.0	0.2	0.2	0.2	0.2	5.1	1.6	0.0	0.0	0.2	0.2
Cs [µg/g]	0.0	0.0	0.0	0.0	0.0	0.1	0.2	0.2	0.0	0.0	0.0	0.0
Ba [µg/g]	0.0	0.1	-0.7	0.2	0.0	0.2	0.6	0.3	0.0	0.0	0.2	0.2

Note: Hz: Harzburgite; Dn: Dunitite; * calculated assuming stoichiometry; n. d.: not determined; n. c.: not calculated; b. d. l.: below detection limit.

Table 3. Representative electron microprobe analyses and in situ LA-ICP-MS data of serpentinization products from Leg 195, Site 1200A, South Chamorro Smt.

Sample	13R-2		13R-2		16R-2		16R-2		7R-2		7R-2		16R-2		16R-1		7R-2		16R-2		13R-2		3R-1		3R-1		9R-1			
	Analys. no.	72-126	77-128	245-132	239-63	261-50	280-49	74-83	153-37	66-40	210-125	44-114	#9-205	39-189	#8-174															
Mineral	Srp	Srp-Brc	Srp	Brc	Srp-Brc	Srp-Brc	Srp	Srp-Brc	Srp-Brc	Srp-Brc	Srp-Brc	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp		
fabric type	type (i)	type (i)	type (i)-late	type (i)-late	halo (mesh)	halo (mesh)	halo (mesh)	halo (mesh)	halo (mesh)	halo (Opx)	halo (Opx)	halo (Opx)	halo (Opx)	halo (Opx)																
wt.%																														
SiO2	40.86	21.24	41.45	0.51	36.42	34.87	40.04	27.47	34.04	38.08	40.87	40.16	40.73	41.00																
TiO2	0.00	0.00	0.01	0.06	0.01	0.00	0.01	0.00	0.05	0.00	0.00	0.02	0.05	0.00	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02	0.02
Al2O3	0.40	1.36	0.27	0.02	0.15	0.12	0.31	0.16	1.95	1.38	0.32	0.07	0.09	0.02	0.07	0.07	0.09	0.09	0.09	0.09	0.09	0.09	0.09	0.09	0.09	0.09	0.09	0.09	0.09	0.09
Cr2O3	0.13	0.04	0.00	0.02	0.02	0.00	0.00	0.03	0.17	0.81	0.00	0.02	0.00	0.00	0.02	0.02	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
FeO	5.52	8.82	5.08	42.00	5.83	6.65	4.29	6.99	6.36	6.36	4.07	3.94	4.46	2.85																
MnO	0.11	0.19	0.09	0.94	0.16	0.11	0.09	0.15	0.09	0.11	0.08	0.10	0.12	0.07	0.10	0.12	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07
MgO	37.97	41.31	38.13	32.58	38.55	39.31	38.23	39.33	37.28	34.95	37.98	37.67	38.05	39.80																
CaO	0.22	0.30	0.09	0.02	0.16	0.24	0.36	0.41	0.19	0.72	0.37	0.07	0.08	0.03	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07	0.07
Na2O	0.05	0.07	0.05	0.19	0.11	0.09	0.05	0.09	0.05	0.07	0.03	0.10	0.07	0.06																
K2O	0.02	0.02	0.01	0.00	0.00	0.02	0.01	0.01	0.02	0.03	0.02	0.01	0.01	0.00																
Total	85.34	73.91	85.19	76.37	81.56	81.61	83.49	74.81	81.15	82.65	83.89	82.28	83.78	83.88																
Oxygens	7	7	7	2	7	7	7	7	7	7	7	7	7	7																
Si	1.98	1.30	2.01	0.01	1.88	1.81	1.98	1.60	1.79	1.93	2.00	2.00	2.00	2.00																
Al	0.02	0.10	0.02	0.00	0.01	0.01	0.02	0.01	0.12	0.08	0.02	0.00	0.00	0.00																
Ti	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00																
Cr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.03	0.00	0.00	0.00	0.00																
Fe3+*	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00																
Mg	2.75	3.77	2.75	2.00	2.96	3.05	2.81	3.41	2.92	2.64	2.77	2.80	2.79	2.89																
Fe2+	0.22	0.45	0.21	0.82	0.25	0.29	0.18	0.34	0.28	0.27	0.17	0.16	0.18	0.12																
Mn	0.00	0.01	0.00	0.02	0.01	0.00	0.00	0.01	0.00	0.00	0.00	0.00	0.01	0.01																
Ca	0.01	0.02	0.00	0.00	0.01	0.01	0.02	0.03	0.01	0.04	0.02	0.00	0.00	0.00																
Na	0.00	0.01	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01	0.01																
K	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00																
Sum	5.01	5.66	4.99	1.99	5.12	5.19	5.02	5.40	5.15	5.01	5.15	5.01	5.00	5.01																
Mg/(Mg+Fe2)	0.92	0.89	0.93	0.58	0.92	0.91	0.94	0.91	0.91	0.91	0.91	0.91	0.94	0.96																
B [µg/g]	69.51	44.15	14.98	n. d.	94.50	106.01	85.59	38.62	61.71	43.65	86.87	84.54	61.13	161.41																
V [µg/g]	23.55	3.79	2.15	1.82	0.70	0.62	1.17	2.26	52.47	52.92	13.25	1.41	10.17	0.97																
Ni [µg/g]	1269	422	1	2	1862	1571	1766	1766	704	573	1601	3038	1867	3908																
Rb [µg/g]	2.11	1.06	0.16	0.14	0.27	0.21	0.58	0.31	0.35	1.29	1.48	0.92	0.73	0.00																
Sr [µg/g]	18.41	8.06	3.97	0.64	19.96	14.17	19.50	12.97	7.77	17.90	29.02	13.44	4.22	2.04																
Cs [µg/g]	1.14	0.50	0.03	n. d.	0.08	0.05	0.11	0.02	0.20	0.52	0.75	0.16	0.22	b. d. l.																
Ba [µg/g]	1.40	1.75	1.23	0.45	2.13	1.61	6.64	0.84	1.44	5.54	3.51	1.09	0.65	0.61																

Note: Hz: Harzburgite; Dn: Dunitite; * calculated assuming stoichiometry; n. d.: not determined; n. c.: not calculated; b. d. l.: below detection limit.

Table 3. Repr

Sample	9R-1	9R-1	16R-2	16R-2	16R-1	7R-2	13R-2	16R-2	16R-2	16R-1	7R-2	51-14	7R-2	57-7	16R-1	16R-2	
Analys. no.	120-56	147-162	203-121	#24-111	137-23	#3-18	41-117	14-12	51-49	51-14	57-7	86-55	184-112				
Mineral	Brc	Srp-Brc	Srp	Srp	Srp	gran. Ol	Srp-Brc	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp-Brc	
fabric type	Jn, recr. mesDn, mesh rim	Hz, mesh	Hz, mesh	Hz, mesh	gran. Ol	gran. Ol	gran. Ol	Bastite	Bastite	Bastite	Bastite	Bastite (Cpx)	Srp-Brc type (ii)				
wt. %																	
SiO2	0.47	34.11	40.26	39.94	39.21	39.96	35.89	39.98	39.29	40.59	38.38	38.42	36.17				
TiO2	0.00	0.00	0.05	0.02	0.03	0.02	0.10	0.01	0.01	0.00	0.00	0.06	0.00				
Al2O3	0.03	0.01	0.10	0.13	0.21	0.64	0.00	0.02	1.23	1.79	1.72	0.65	0.03				
Cr2O3			0.00	0.02	0.01	0.11	0.03	0.00	0.56	0.90	0.94	0.30	0.01				
FeO	16.12	6.74	4.90	6.00	6.61	5.35	5.61	3.38	6.44	7.69	5.92	5.38	5.85				
MnO	0.33	0.09	0.10	0.12	0.12	0.11	0.09	0.04	0.14	0.22	0.17	0.11	0.07				
MgO	54.40	41.85	38.50	36.08	36.43	36.48	41.16	40.46	33.36	31.93	33.51	36.60	40.40				
CaO	0.03	0.09	0.25	0.36	0.23	0.35	0.16	0.22	1.93	2.90	2.17	0.43	0.18				
Na2O	0.05	0.04	0.07	0.04	0.05	0.06	0.01	0.03	0.05	0.12	0.13	0.03	0.07				
K2O	0.02	0.00	0.03	0.04	0.00	0.03	0.00	0.04	0.02	0.05	0.02	0.03	0.01				
Total	71.95	82.99	84.31	82.90	83.03	83.35	83.47	84.27	83.19	86.31	83.16	82.16	82.85				
Oxygens	2	7	7.00	7	7	7	7	7	7	7	7	7	7				
Si	0.01	1.75	1.98	2.00	1.97	1.99	1.81	1.95	1.98	1.99	1.94	1.95	1.84				
Al	0.00	0.00	0.01	0.01	0.01	0.04	0.00	0.00	0.07	0.10	0.10	0.04	0.00				
Ti	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00				
Cr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.03	0.04	0.01	0.00				
Fe3+	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00				
Mg	1.69	3.20	2.82	2.70	2.73	2.71	3.10	2.94	2.51	2.33	2.53	2.77	3.06				
Fe2+	0.28	0.29	0.20	0.25	0.28	0.22	0.24	0.14	0.27	0.32	0.25	0.23	0.25				
Mn	0.01	0.00	0.00	0.01	0.01	0.00	0.00	0.00	0.01	0.01	0.01	0.00	0.00				
Ca	0.00	0.01	0.01	0.02	0.01	0.02	0.01	0.01	0.10	0.15	0.12	0.02	0.01				
Na	0.00	0.00	0.01	0.00	0.00	0.01	0.00	0.00	0.00	0.01	0.01	0.00	0.01				
K	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00				
Sum	1.99	5.25	5.02	4.99	5.02	4.99	5.18	5.05	4.97	4.95	5.00	5.03	5.17				
Mg/(Mg+Fe2)	0.86	0.92	0.93	0.91	0.91	0.92	0.93	0.96	0.90	0.88	0.91	0.92	0.92				
B [µg/g]	n. d.	187.33	42.80	41.93	84.22	70.14	90.26	28.23	54.77	20.69	67.23	84.17	25.35				
V [µg/g]	1.07	0.49	0.44	0.28	5.26	3.36	2.01	0.38	61.19	34.66	42.01	56.74	0.28				
Ni [µg/g]	2431	2710	2458	2288	2534	2277	1462	2906	623	443	579	688	2154				
Rb [µg/g]	0.23	0.13	3.00	2.93	0.54	2.25	2.76	2.85	3.76	3.51	3.66	2.52	0.29				
Sr [µg/g]	0.32	2.34	13.77	10.81	35.43	19.32	19.99	7.68	14.36	13.62	25.09	26.85	6.66				
Cs [µg/g]	n. d.	0.00	1.96	1.93	0.10	0.93	1.19	1.73	3.25	2.15	2.41	1.44	0.12				
Ba [µg/g]	0.09	0.39	4.85	5.45	3.08	2.49	9.58	6.93	2.93	4.79	5.52	6.53	3.09				

Note: Hz: Haz

Table 3. Repr







Sample	16R-2	9R-1	7R-2	7R-2	13R-2	11R-1	16R-1	16R-1	16R-1	16R-2	3R-1	3R-1
Analys. no.	185-122	#5-141	110-31	166-42	68-119	145-23	132-17	139-35	57-76	33-187	34-183	
Mineral	Srp-Brc	Srp-Brc	Srp-Brc	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp	Srp
fabric type	type (i)	type (ii)	type (ii)	type (iii)	type (iii)	type (iii)	type (iv)	type (iv)	type (iv)	type (iv)	type (iv)	type (iv)
wt. %												
SiO ₂	28.47	33.46	30.72	40.14	40.48	41.28	42.15	41.29	42.46	40.82	41.12	
TiO ₂	0.00	0.02	0.00	0.02	0.03	0.02	0.00	0.00	0.05	0.05	0.01	
Al ₂ O ₃	0.03	0.02	0.00	0.27	0.39	0.14	0.20	0.19	0.25	0.06	0.06	
Cr ₂ O ₃	0.02	0.00	0.00	0.00	0.01	0.01	0.05	0.00	0.55	0.00	0.00	
FeO	7.26	6.19	6.86	5.37	8.70	5.92	5.32	5.64	4.51	3.73	4.75	
MnO	0.14	0.13	0.12	0.13	0.15	0.10	0.09	0.10	0.09	0.10	0.11	
MgO	43.10	42.66	42.95	38.25	35.75	38.36	38.77	38.09	38.61	38.97	38.08	
CaO	0.22	0.06	0.19	0.07	0.16	0.04	0.03	0.09	0.08	0.04	0.14	
Na ₂ O	0.07	0.09	0.01	0.06	0.09	0.02	0.01	0.04	0.03	0.04	0.07	
K ₂ O	0.00	0.04	0.01	0.01	0.02	0.01	0.00	0.00	0.00	0.02	0.00	
Total	79.37	82.66	81.01	84.39	85.86	85.92	86.65	85.48	86.65	83.89	84.48	
Oxygens	7	7	7	7	7	7	7	7	7	7	7	
Si	1.56	1.72	1.64	1.97	1.99	1.99	2.01	2.00	2.01	2.00	2.01	
Al	0.00	0.00	0.00	0.02	0.02	0.01	0.01	0.01	0.01	0.00	0.00	
Ti	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
Cr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.00	0.00	
Fe ₃ ⁺	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
Mg	3.52	3.27	3.41	2.80	2.61	2.76	2.75	2.75	2.73	2.84	2.77	
Fe ₂ ⁺	0.33	0.27	0.31	0.22	0.36	0.24	0.21	0.23	0.18	0.15	0.19	
Mn	0.01	0.01	0.01	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.00	
Ca	0.01	0.00	0.01	0.00	0.01	0.00	0.00	0.00	0.00	0.00	0.01	
Na	0.01	0.01	0.00	0.01	0.01	0.00	0.00	0.00	0.00	0.00	0.01	
K	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	
Sum	5.44	5.28	5.37	5.02	5.01	5.01	4.99	5.00	4.97	5.00	4.99	
Mg/(Mg+Fe ₂)	0.91	0.92	0.92	0.93	0.88	0.92	0.93	0.92	0.94	0.95	0.93	
B [µg/g]	16.06	43.96	27.53	134.62	73.16	29.20	23.84	39.95	14.69	25.36	44.66	
V [µg/g]	0.21	0.47	0.27	21.83	3.03	1.17	9.30	22.08	2.62	0.95	1.79	
Ni [µg/g]	1872	2151	2078	43	57	7	175	193	89	71	109	
Rb [µg/g]	0.11	0.08	0.17	0.33	0.53	0.11	0.11	0.24	0.09	0.16	0.34	
Sr [µg/g]	4.35	3.73	9.39	1.62	9.06	0.97	2.97	9.26	2.58	0.80	1.67	
Cs [µg/g]	0.06	0.01	0.07	0.05	0.08	0.01	0.01	0.07	0.01	0.02	0.03	
Ba [µg/g]	2.40	0.57	1.61	0.46	1.01	0.35	0.44	1.16	0.60	0.21	0.31	

Note: Hz: Hat

Table4

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Table 4. Overview of selected properties of serpentine from different serpentinization events.

Serpentinization phase	Schematic fabric illustration	Secondary mineralogy	Serpentine geochemistry
Early magnetite veins		Srp, Mgt	n. a.
Pervasive stage		mesh: Lz* bastite: Lz*	mesh: < 0.5 wt.% Al ₂ O ₃ in Srp; bastite: up to 2.4 wt.% Al ₂ O ₃ ; Mesh: low V, high Ni; bastite: high V, low Ni; Cs, Rb, Sr, Ba
Type (i) veins		Liz*, Brc	up to 2.0 wt.% Al ₂ O ₃ ; Cs, Rb, Sr, Ba
Type (ii) veins		Ctl*, Brc	up to 0.9 wt.% Al ₂ O ₃ ; low V, high Ni; Cs, Rb, Sr, Ba
Type (iii) veins "Frankenstein"		Ctl*	up to 1.0 wt.% Al ₂ O ₃ ; V, low Ni; Cs, Rb, Sr, Ba
Type (iv) veins <i>multiple crack-seal</i>		Ctl*	up to 0.8 wt.% Al ₂ O ₃ ; V, low Ni; Cs, Rb, Sr, Ba

Note: n. a. not analyzed; *confirmed by micro-Raman; trace element contents: **bold**-relatively high values, plain-relatively low values.

Table 5. Oxygen isotope data of serpentinization products from different serpentinization phases in ultramafic clasts from South Chamorro Smt., Leg 195, Site 1200A.

Sample	% serp	fabric type	minerals #	$\delta_{18}O^*$
3R-1, 3-7 cm	100	type (iv) vein	Ctl	6.9
7R-2, 40-43 cm	97	halo region around type (i) vein	Liz, Ctl	7.6
10R-1_54-62 cm	100	vein, type (iv)	Ctl	6.4; 6.1
11R-1, 47-49 cm	100	vein in clast brown rim zone	Ctl	6.4
12R-1, 8-11 cm	100	bastite in brown clast rim zone	Ctl, Palyg, Sep, Sap (15 vol.%)	8.1
13R-2, 5-7 cm	100	chrysotile vein, type (iv)	Ctl	10.5
13R-2, 5-7 cm	100	serpentine, type (ii)		7.3
16R-2, 7-9 cm	97	type (ii) veins	Ctl, Liz, Brc (13 vol.%)	7.3
16R-2, 7-9 cm	97	type (i)/ reactivated by type (iv) vein	Liz, Ctl, Brc (4 vol.%)	6.6

Note: #determined by XRD; * VSMOW.