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Panel Session III - Astronomical Search for Origins Program

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Astronomical Search for Origins Program

Eric P. Smith

Origins Theme Scientist,
James Webb Space Telescope Program Scientist

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Origins Theme's Two Defining Questions

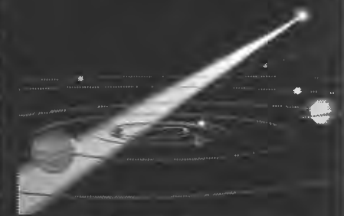
Where Did We Come From?



Tracing Our Cosmic Roots

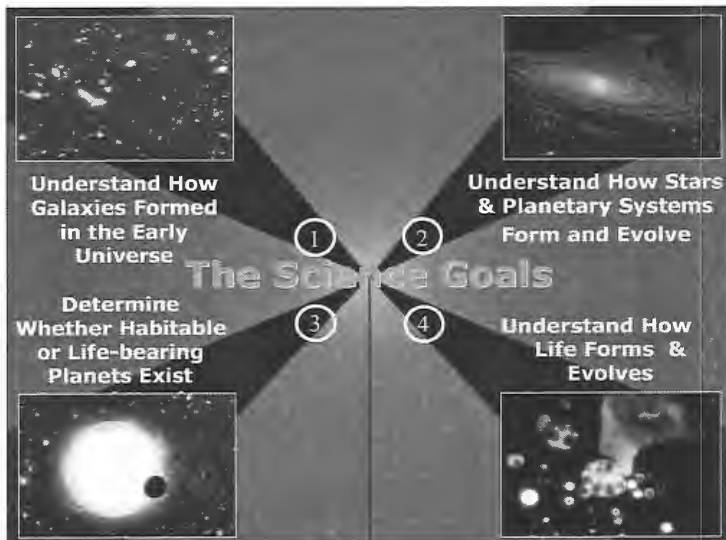
- Formation of galaxies, stars, heavy elements, planetary systems and ... life on the early Earth

Are We Alone?



Search for Life Outside the Solar system

- Remote detection of biological activities on planets beyond our solar system



Understand How the Galaxies Formed in the Early Universe

- Determine the role of gravity in the emergence of galaxies
- Establish how the birth and aging of a galaxy influence the chemical composition that is available to stars, planets, and living organisms



- HST, COBE and WMAP gave us our first hints
- Spitzer and JWST will complete our understanding

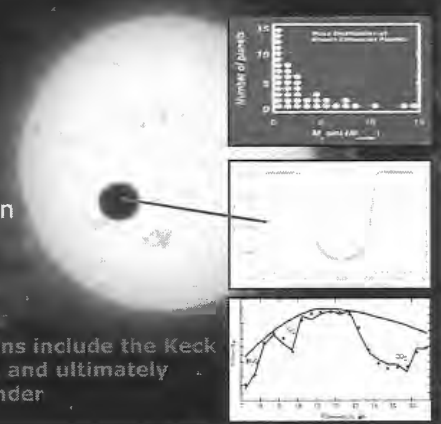
Understand How Stars and Planetary Systems Form and Evolve

- Discover planetary systems forming around young stars
- Characterize the planets and planetary systems around stars



Determine Whether Habitable or Life-bearing Planets Exist around Nearby Stars

- Determine how common habitable worlds are in the Universe
- Establish how to recognize the signatures of life on other worlds

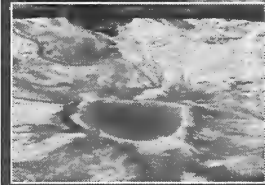


- Key Science missions include the Keck Interferometer, SIM, and ultimately Terrestrial Planet Finder

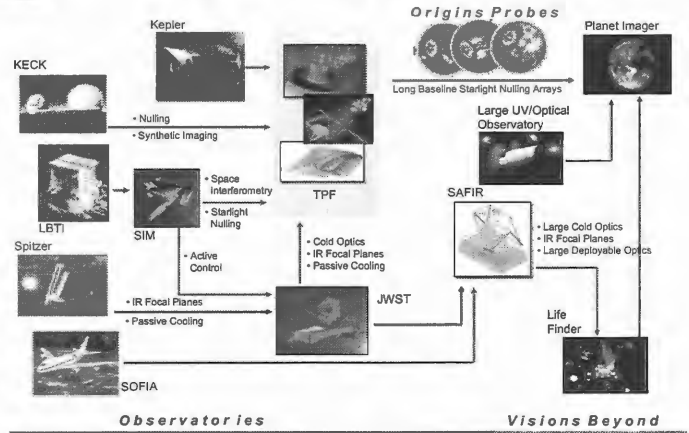
Understand How Life Forms & Evolves

- Determine the principles governing the organization of matter into living systems
- Determine limits of life in environments analogous to conditions on other worlds

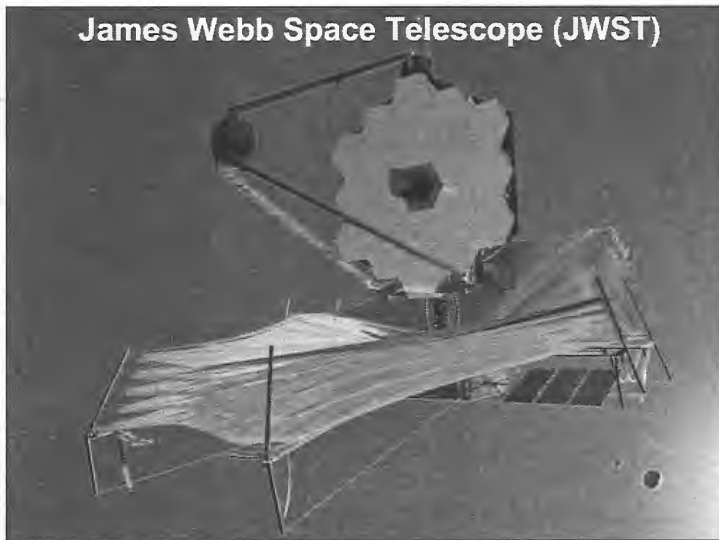
• Astrobiology research to identification of



Origins Program Missions



James Webb Space Telescope (JWST)



JWST Overview

- **Mission** (Part of NASA's *Astronomical Search for Origins* theme)
 - Detect light from the first objects in the universe
 - Study of the birth and evolution of galaxies
 - Understand star and planet formation
- **Description**
 - Sensitive (large and cold) infrared (0.6-28 μ m) space observatory
 - Orbit at Sun-Earth Lagrange Point 2 (L2 - about 1.5 million km from Earth)
 - Launch in 2011
 - 5 year lifetime (10 year goal)
 - ~\$3B total value:
 - ~\$2.5B U.S. (~\$1.9B development + ~\$0.6B ops) + ~\$0.4B equivalent value from ESA + ~\$0.1B equivalent value from CSA [nearest \$0.1B]
- **Organization**
 - Project Lead: NASA's Goddard Space Flight Center
 - International collaboration w/ European and Canadian Space Agencies (ESA & CSA)
 - Prime Contractor: Northrop Grumman Space Technology
 - Other mission hardware & software: University of Arizona, JPL, STScI
- **Status**
 - Preliminary Design (Formulation/Phase B)

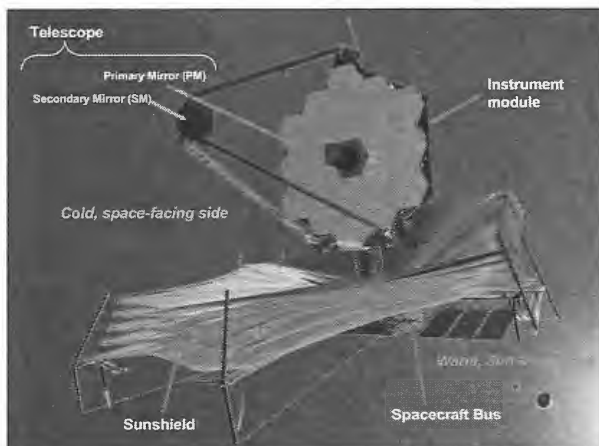
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JWST Observatory



JWST Animation



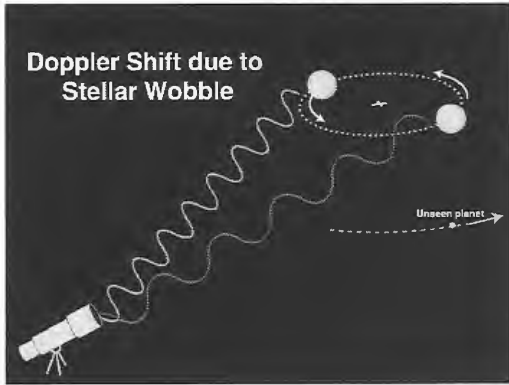
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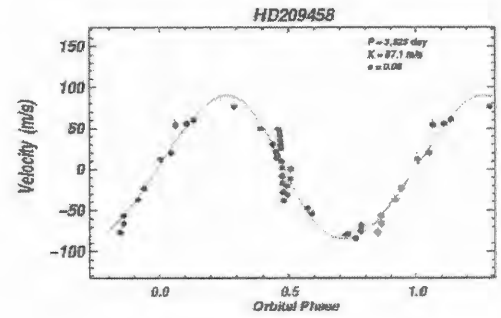
Radial Velocity Planet Detection



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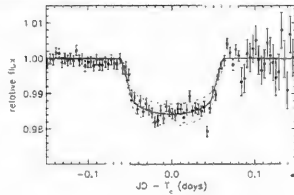
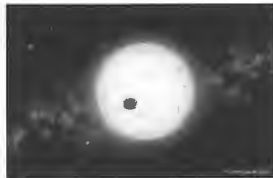
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Transits Reveal Planet

- Transits of a planet orbiting HD 209458 determine properties of another Solar System
 - Previously known from radial velocity
 - Inclination= 87.1°
 - Mass= 0.63 M_{Jup}
 - Radius = 1.27-1.6 R_{Jup}
 - Density= 0.27-0.38 g/cc < Saturn
- Spectroscopy of transit signal can probe planet's atmosphere



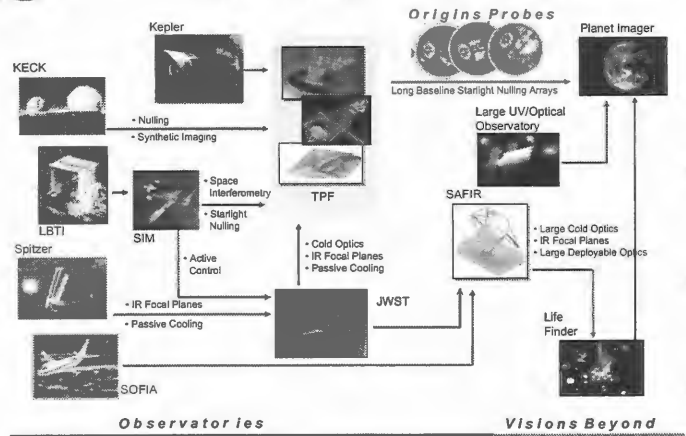
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Origins Program Missions



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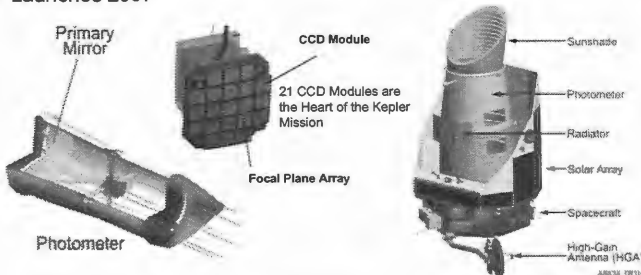
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The Kepler Mission

Use transit photometry to detect Earth-size planets

- Wide field-of-view 0.95-meter (3-foot) diameter telescope
- Monitor 100,000 stars (every 15 minutes) for 4 years
- Enough precision (20 ppm) to detect transits of Earth-size planets
- Launches 2007



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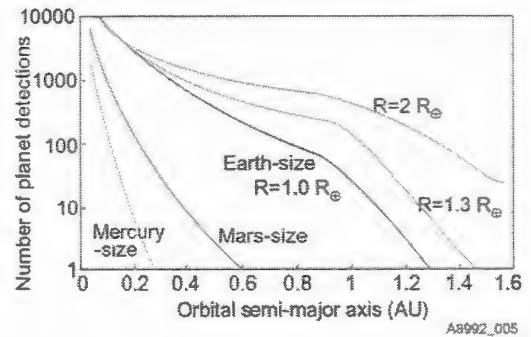
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The Kepler Mission

Expected Number of Planetary Discoveries if most Stars have Planets



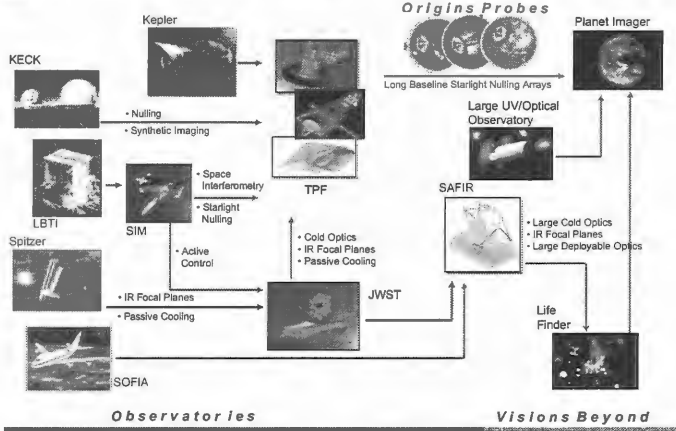
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Origins Program Missions



Space Interferometers: SIM

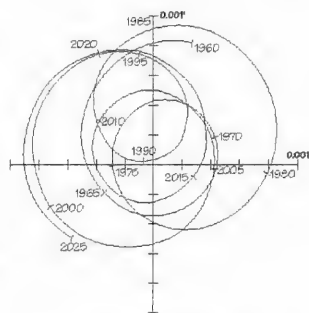
Space Interferometer Mission (SIM):

- Use astrometry to find planets orbiting nearby stars
 - From Jupiter's to a few Earth masses
- Carry out general astrometric program
- Demonstrate scientific operation of space interferometry
- Demonstrate TPF technologies and interferometric techniques
- Launch 2009
- 5+ year operation



Astrometric Search for Planets

- Measure star's positional wobble
 - Complements radial velocity
 - Most sensitive to massive planets far from their stars
- **Space Interferometry Mission** will be able to detect tiny, tiny stellar wobble due to Earth-mass planets around nearby stars
 - Measures positions of stars with precision equivalent to the size of an *amoeba in California as seen from Washington, DC!!*



Sun's wobble seen from 33 light-years away

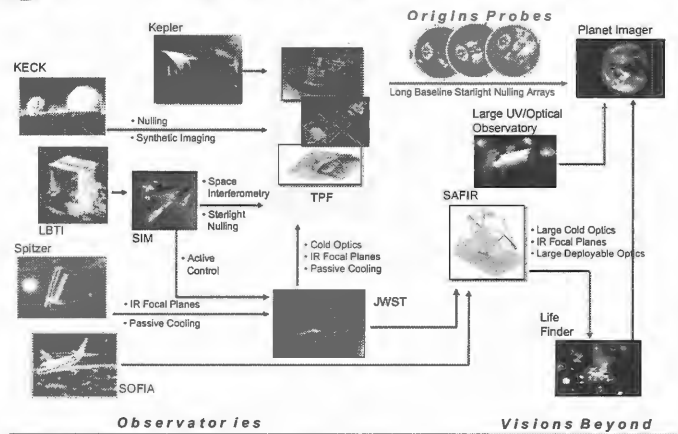
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Origins Program Missions



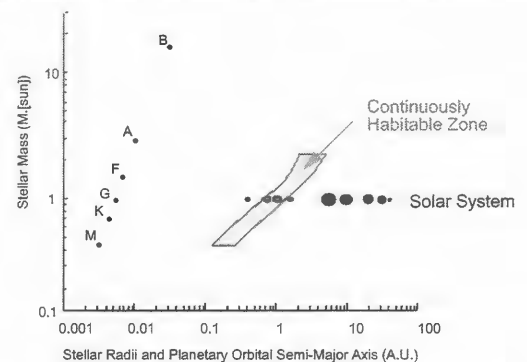
Terrestrial Planet Finder (TPF)

"Like finding a firefly next to a searchlight on a foggy night"

- Detecting light from planets beyond solar system is hard:
 - Parent star is 10^6 to 10^9 times brighter
 - Planet within 1 AU of its parent star
- The science goal of TPF: identify terrestrial planets in the habitable zones of ~100-200 nearby stars and to examine the brightest detected planets for signs of life itself



The Habitable Zone* for Various Types of Stars



*The Habitable Zone is the range of distances from a star at which water can remain liquid on the surface of an orbiting planet

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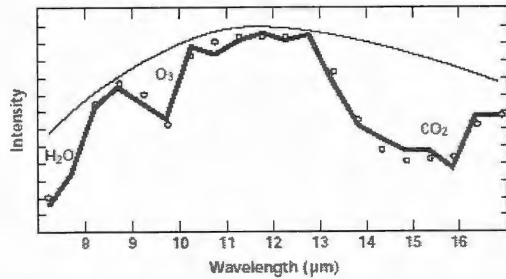
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Simulated TPF Spectrum of an Earth-like Planet



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Atmospheric Composition of an Early Earth-like World Harboring Life?

Main Constituents

CO₂, N₂, H₂O

Secondary Constituents

CO, H₂, H₂S, CH₄

Trace Constituents

Organo-sulfur compounds, hydrocarbons, other reduced gases

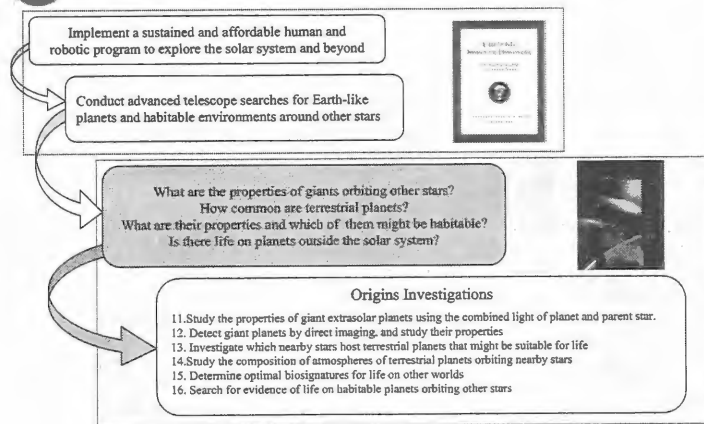
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Origins Science in the New Vision



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How do Origins Missions Align with the Vision?

Mission	Status	Investigations
HST	Operational	11,23
Keck Interf.	ΦC/D	11,12
LBTI	ΦC/D	11,12
Kepler	ΦB, 2007 launch	11,13
SIM	ΦB, 2009 launch	11,13,14
JWST	ΦB, 2011 launch	11,12
TPF	~ 2014 launch	11,12,13,14,15,16

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Changes to Missions in response to the Vision

- **JWST:** project considering white paper study by astrobiologists outlining observatory capabilities for optimal astrobiology science return
- **TPF:** Split TPF program into two components, coronagraphic and interferometric observatories and accelerate development on coronagraphic option.

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Why Fly a TPF Coronagraph and an Interferometer?

- Respond to presidents vision - "...conduct advanced telescope searches for Earth-like planets and habitable environments around other stars."
- The technology appears to be maturing for both
- Fly sooner -- A small to moderate sized coronagraph in ~2014
 - Starlight suppression technology must be fully demonstrated first (by 2006)
 - Enter phase A in 2007
- More capable, later --- Formation flying interferometer done jointly with ESA later towards end of decade
 - Continue technology development during coronagraph formulation and implementation
 - Collaboration with ESA on SMART3 (~2011)
 - Enter Phase A in ~2011
- Most compelling science --- Data from both wavelength regions provide robust assessment of habitability and biomarkers
- Doing both within ORIGINS line requires no near term augmentation

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ORIGINS Program Status

- Well Aligned with New NASA Vision
 - ORIGINS program already investigating planetary systems today
 - Existing programs adapting to vision
- 2003 was an important year for ORIGINS
 - Both SIM and JWST passed into Phase B and are meeting technology and schedule milestones
- Long Term issues
 - Develop technology program for the large telescopes of a decade from now