Supporting information for the main manuscript:

## "Strain budget of the Ecuador-Colombia subduction zone: a stochastic view"

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This electronic supplement is a collection of additional figures referenced in the main article. These figures were added to ensure the precision of the description of our method and results.

## Supplementary text T1

Following Ye et al. (2014) and Nocquet et al. (2017), we compare waveforms of the 1942 earthquake recorded at the DBN station (De Bilt, Netherlands) with stochastic waveform predictions at the same station for the 2016 Pedernales slip distribution.

We compute displacement Green's functions for each subfault patch using the Kikuchi-Kanamori program (Kikuchi and Kanamori, 2003; Kikuchi, Masayuki and Kanamori, Hiroo, 1982). For comparison, we then convolve predicted stochastic waveforms with the instrumental response of the Galitzin seismometer that recorded the 1942 earthquake (pendulum and galvanometer periods  $T_p=T_g=25$  s and gain factor  $V_m=310$ ; Charlier and Van Gils, 1953).

In Fig. S10a, we first compare 1942 waveforms with predictions of the kinematic slip model (i.e., for the posterior distributions of slip, rise-times, rupture velocities) and hypocenter location obtained for the 2016 Pedernales earthquake. Model predictions show poor fit to the 1942 earthquake waveform. In Fig. S10b, we then compute predictions for the same kinematic slip distribution, but with a hypocenter location between the two slip asperities. With that hypocenter location, model predictions have a very good fit to the 1942 waveform. Finally, in Fig. S10c, we predict waveforms for a slip distribution on the megathrust interface, but updip of the actual 2016 rupture. Notice, that the dip is different due to the variation of the slab interface geometry with depth. We also correct the slip amplitude for the variation of shear modulus in our velocity model (cf., Fig. S2). Similarly to the previous case, the hypocenter is located between the two slip asperities. In this scenario, we are also able to explain the 1942 waveform. It illustrates that the teleseismic P-waveform is mostly sensitive to the relative location of the hypocenter and slip asperity rather than the absolute location of the earthquake.

Supplementary movie M1: Variability in the Ecuador-Colombia geodetic coupling solution The animation is made with 150 models randomly selected in the posterior population represented by the background colour. Grey lines are the 2 m contour intervals of 150 co-seismic models also randomly selected in the posterior population.

Supplementary movie M2: Temporal evolution of co-seismic slip of the 2016 Pedernales earthquake. (left) Posterior mean model of the cumulative slip. The bottom-right inset shows the stochastic source time function. (right) Incremental slip on the fault. The red star marks the inverted posterior mean hypocenter location.

Supplementary movie M3: Variability in the 2016 Pedernales earthquake co-seismic slip distribution solution. The animation is made with 200 models randomly selected in the posterior population.

Satellite N° of data Corr. length Orbit Acquisition dates Std. ALOS-2 07/02/16 - 01/05/16 130 2.88 km ascending 5.3 mmALOS-2 descending 01/04/16 - 29/04/16 483 9.2 mm 11.90 km12/04/16 - 24/04/16 Sentinel-1A descending 380 5.0 mm15.0 km

Table S1: InSAR observations used in this study.

Station	Type	Filter corner frequencies		
		East	North	Up
bahi	HRGPS	0.015Hz - 0.08Hz	0.015Hz - 0.08Hz	0.015Hz - 0.08Hz
$\operatorname{cabp}$	HRPGS	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
ecec	HRPGS	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
flfr	HRPGS	0.015 Hz - 0.08 Hz	$0.015 \mathrm{Hz}$ - $0.08 \mathrm{Hz}$	$0.015 \mathrm{Hz}$ - $0.08 \mathrm{Hz}$
mlec	HRPGS	0.015 Hz - 0.08 Hz	$0.015 \mathrm{Hz}$ - $0.08 \mathrm{Hz}$	$0.015 \mathrm{Hz}$ - $0.08 \mathrm{Hz}$
$\operatorname{momp}$	HRPGS	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
onec	HRPGS	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
pdns	HRPGS	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
ISPT	Strong motion	N/A	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
PDNS	Strong motion	0.037 Hz - 0.08 Hz	0.037 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
LGCB	Strong motion	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
AATC	Strong motion	0.028 Hz - 0.08 Hz	0.028 Hz - 0.08 Hz	0.032 Hz - 0.08 Hz
AES1	Strong motion	N/A	N/A	0.015 Hz - 0.08 Hz
AMNT	Strong motion	N/A	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
APED	Strong motion	0.035 Hz - 0.08 Hz	0.035 Hz - 0.08 Hz	0.035 Hz - 0.08 Hz
ATON	Strong motion	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
AV18	Strong motion	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz
AV21	Strong motion	$0.032\mathrm{Hz}$ - $0.08\mathrm{Hz}$	0.015 Hz - 0.08 Hz	0.015 Hz - 0.08 Hz

Table S2: Seismological data and filtering used in this study. We use a  $4^{th}$  order Butterworth bandpass filter.



Figure S1: **Parametrization of the megathrust interface used for the coupling inversion.** Coupling value is inverted at each nodes



Figure S2: Different models variability of the P-wave, S-wave, and density as a function of depth in central Ecuador. A layered model used in this study for Green's function [GF] calculations is plotted as a solid black line. The blue line represents the CRUTST2.0 model in the area (http://igppweb.ucsd.edu/~gabi/rem.html). The other models are from (Vallee et al., 2013; Bethoux et al., 2011; Nocquet et al., 2017). Grey histograms are the probability density function representing our confidence level on the elastic properties, as used to build the model prediction error.



Figure S3: **Posterior Median coupling model.** Thin black lines represent the fault parametrization. Coupling values are inverted at each nodes. Interseismic GPS displacement and predictions for the median model are plotted as black and blue arrows, respectively.



Figure S4: **Posterior Mean coupling model for a coarse parametrisation.** Same as Figure 2a. in the main text but obtained with a coarser fault parametrisation.



Figure S5: **Decimated InSAR observations, predictions, and residuals**. (a, d, g) Decimated InSAR observations inverted in this study. (b, e, h) Predictions for the posterior mean model. (c, f, i) Residuals of the Sentinel (top row), descending ALOS-2 (middle row), and ascending ALOS-2 (bottom row) interferograms.



Figure S6: Empirical covariance functions for the InSAR observations 1D empirical covariance functions and the associated best-fit exponential function for each tracks. For each image, we compute the empirical covariance as a function of the distance between pixels and then fit an exponential function to these covariances (Jolivet et al., 2012). This exponential function is then used to build the data covariance matrix used in the inversion.



Figure S7: **Parametrization of the megathrust interface used for the co-seismic inversion** The coloured plane represent the slab1.0 model (Hayes et al., 2012). Each subfault patch is a 15 km x 15 km square



Figure S8: Strong-motion observations and model predictions not presented in Figure 6 in the main text. The North (left) and vertical (right) components of each station are plotted around the map. For each waveform, the bold number indicates it's maximum amplitude. The station azimuth  $\Phi$  and distance d to the epicenter are also given. The black line is the recorded waveform. The gray lines are the stochastic predictions for our posterior model. The red line is the mean of the stochastic predictions.



Figure S9: Comparaison of co-seismic moment and moment deficit in the 1958 earthquake region. a) The background colour represents the coupling posterior mean model. The black dashed lines delimit four different areas where the co-seismic moment of the 1958 event and moment deficit for the 1906 - 1958 period are computed. b-e) Probability densities of the co-seismic moment released by the 1958 earthquake and the moment deficit accumulated between 1906 and 1958 in the different dashed area shown in a).



Figure S10: Comparison of model predictions and 1942 earthquake waveform recorded in the DBN station, Netherlands. (top) Slip model and hypocenter location (red star) used to compute the predictions shown in the bottom row. The model presented in a) results from the kinematic slip inversion of the 2016 earthquake. The models in b) and c) use a different hypocenter located between the two main slip asperities. The slip model in c) is the same as in a) and b), but located updip along the megathrust interface. Black lines in c) are slip contours of the original slip model. (bottom) East component waveform recorded at DBN for the 1942 earthquake (in black) and stochastic predictions (in grey) for the model shown on top. The red line is the posterior mean prediction. Predictions were convolved with the instrumental response of the Galitzin that recorded the event.

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