

Syracuse University

**SURFACE**

---

Physics

College of Arts and Sciences

---

6-7-2003

## Radiative Decays Involving $a_0(980)$ and $f_0(980)$ in the Vector Meson Dominance Model

Joseph Schechter

*Department of Physics, Syracuse University, Syracuse, NY*

Deirdre Black

*Jefferson Lab*

Masayasu Harada

*Nagoya University*

Follow this and additional works at: <https://surface.syr.edu/phy>

 Part of the [Physics Commons](#)

---

### Recommended Citation

Schechter, Joseph; Black, Deirdre; and Harada, Masayasu, "Radiative Decays Involving  $a_0(980)$  and  $f_0(980)$  in the Vector Meson Dominance Model" (2003). *Physics*. 278.

<https://surface.syr.edu/phy/278>

This Article is brought to you for free and open access by the College of Arts and Sciences at SURFACE. It has been accepted for inclusion in Physics by an authorized administrator of SURFACE. For more information, please contact [surface@syr.edu](mailto:surface@syr.edu).

## Radiative Decays involving $a_0(980)$ and $f_0(980)$ in the Vector Meson Dominance Model <sup>\*)</sup>

Deirdre BLACK<sup>(a)</sup>, Masayasu HARADA<sup>(b)</sup> and Joseph SCHECHTER<sup>(c)</sup>

<sup>(a)</sup> *Theory Group, Jefferson Lab, 12000 Jefferson Ave., Newport News, VA 23606, USA*

<sup>(b)</sup> *Department of Physics, Nagoya University, Nagoya, 464-8602, Japan*

<sup>(c)</sup> *Department of Physics, Syracuse University Syracuse, NY 13244-1130, USA*

(Received February 1, 2008)

We summarize some features of the vector meson dominance model which was recently proposed for studying radiative decays involving the scalar mesons. Using the experimental values of  $\Gamma(a_0 \rightarrow \gamma\gamma)$ ,  $\Gamma(f_0 \rightarrow \gamma\gamma)$  and  $\Gamma(\phi \rightarrow a_0\gamma)$  as inputs, we show that the model predicts a large hierarchy between  $\Gamma(a_0 \rightarrow \omega\gamma)$  and  $\Gamma(a_0 \rightarrow \rho\gamma)$  as well as between  $\Gamma(f_0 \rightarrow \omega\gamma)$  and  $\Gamma(f_0 \rightarrow \rho\gamma)$ .

### §1. Introduction

According to the recent theoretical and experimental analysis, there is a possibility that nine light scalar mesons exist below 1 GeV, and they form a scalar nonet.<sup>1)</sup> In addition to the well established  $f_0(980)$  and  $a_0(980)$  evidence of both experimental and theoretical nature for a very broad  $\sigma$  ( $\simeq 560$ ) and a very broad  $\kappa$  ( $\simeq 900$ ) has been presented. However, the properties of the nonet members such as quark structure and interactions with other mesons are not well known. It is interesting to study the properties of these light scalar mesons, which would be of great importance for a full understanding of QCD in its nonperturbative low energy regime.

In particular, the reactions  $\phi \rightarrow f_0\gamma$  and  $\phi \rightarrow a_0\gamma$  have recently been observed<sup>2)</sup> with good accuracy and are considered as useful probes of scalar properties. The theoretical analysis was initiated by Achasov and Ivanchenko<sup>3)</sup> and followed up by many others.<sup>4)</sup> The models employed are essentially variants of the single  $K$  meson loop diagram to which a  $\phi$ -type vector meson, a photon and two pseudoscalars or a scalar are attached.

In this paper we summarize some features of the vector meson dominance model which was recently proposed in Ref. 5) for studying radiative decays involving the scalar mesons. Our model predicts that a large hierarchy between  $\Gamma(a_0 \rightarrow \omega\gamma)$  and  $\Gamma(a_0 \rightarrow \rho\gamma)$  as well as between  $\Gamma(f_0 \rightarrow \omega\gamma)$  and  $\Gamma(f_0 \rightarrow \rho\gamma)$ .

---

<sup>\*)</sup> Talk given by M. Harada at International Symposium on Hadron Spectroscopy, Chiral Symmetry and Relativistic Description of Bound Systems, February 24-26, 2003, Nihon University, Tokyo, Japan.

## §2. Vector Meson Dominance Model

Framework of the model proposed in Ref. 5) is that of a standard non-linear chiral Lagrangian containing, in addition to the pseudoscalar nonet matrix field  $\phi$ , the vector meson nonet matrix  $\rho_\mu$  and a scalar nonet matrix field denoted  $N$ . Under chiral unitary transformations of the three light quarks;  $q_{L,R} \rightarrow U_{L,R} \cdot q_{L,R}$ , the chiral matrix  $U = \exp(2i\phi/F_\pi)$ , where  $F_\pi \simeq 0.131$  GeV, transforms as  $U \rightarrow U_L \cdot U \cdot U_R^\dagger$ . The convenient matrix  $K(U_L, U_R, \phi)^{6)}$  is defined by the following transformation property of  $\xi$  ( $U = \xi^2$ ):  $\xi \rightarrow U_L \cdot \xi \cdot K^\dagger = K \cdot \xi \cdot U_R^\dagger$ , and specifies the transformations of “constituent-type” objects. The fields we need transform as

$$\begin{aligned} N &\rightarrow K \cdot N \cdot K^\dagger, \\ \rho_\mu &\rightarrow K \cdot \rho_\mu \cdot K^\dagger + \frac{i}{\tilde{g}} K \cdot \partial_\mu K^\dagger, \\ F_{\mu\nu}(\rho) &= \partial_\mu \rho_\nu - \partial_\nu \rho_\mu - i\tilde{g} [\rho_\mu, \rho_\nu] \rightarrow K \cdot F_{\mu\nu} \cdot K^\dagger, \end{aligned} \quad (2.1)$$

where the coupling constant  $\tilde{g}$  is about 4.04. One may refer to Ref. 7) for our treatment of the pseudoscalar-vector Lagrangian and to Ref. 8) for the scalar addition. The entire Lagrangian is chiral invariant (modulo the quark mass term induced symmetry breaking pieces) and, when electromagnetism is added, gauge invariant.

In Ref. 5), the strong trilinear scalar-vector-vector terms were included into the effective Lagrangian as

$$\begin{aligned} \mathcal{L}_{SVV} &= \beta_A \epsilon_{abc} \epsilon^{a'b'c'} [F_{\mu\nu}(\rho)]_{a'}^a [F_{\mu\nu}(\rho)]_{b'}^b N_c^c \\ &+ \beta_B \text{Tr}[N] \text{Tr}[F_{\mu\nu}(\rho)F_{\mu\nu}(\rho)] \\ &+ \beta_C \text{Tr}[NF_{\mu\nu}(\rho)] \text{Tr}[F_{\mu\nu}(\rho)] \\ &+ \beta_D \text{Tr}[N] \text{Tr}[F_{\mu\nu}(\rho)] \text{Tr}[F_{\mu\nu}(\rho)] . \end{aligned} \quad (2.2)$$

Chiral invariance is evident from (2.1) and the four flavor-invariants are needed for generality. (A term  $\sim \text{Tr}(FFN)$  is linearly dependent on the four shown). Actually the  $\beta_D$  term will not contribute in our model so there are only three relevant parameters  $\beta_A$ ,  $\beta_B$  and  $\beta_C$ . Equation (2.2) is analogous to the  $PVV$  interaction <sup>\*)</sup> which was originally introduced as a  $\pi\rho\omega$  coupling a long time ago.<sup>10)</sup> One can now compute the amplitudes for  $S \rightarrow \gamma\gamma$  and  $V \rightarrow S\gamma$  according to the diagrams of Fig. 1.

The decay matrix element for  $S \rightarrow \gamma\gamma$  is written as  $(e^2/\tilde{g}^2)X_S \times (k_1 \cdot k_2 \epsilon_1 \cdot \epsilon_2 - k_1 \cdot \epsilon_2 k_2 \cdot \epsilon_1)$  where  $\epsilon_\mu$  stands for the photon polarization vector. It is related to the width by

$$\Gamma(S \rightarrow \gamma\gamma) = \alpha^2 \frac{\pi}{4} m_S^3 \left| \frac{X_S}{\tilde{g}^2} \right|^2, \quad (2.3)$$

---

<sup>\*)</sup> It was shown<sup>9)</sup> that the complete vector meson dominance (VMD) is violated in the  $\omega \rightarrow \pi^0\pi^+\pi^-$  decay which is expressed by  $PVV$  interactions. However, since the VMD is satisfied in other processes such  $\pi^0 \rightarrow \gamma\gamma^*$  as well as in the electromagnetic form factor of pion, we here assume that it holds in the processes related to  $SVV$  interactions.

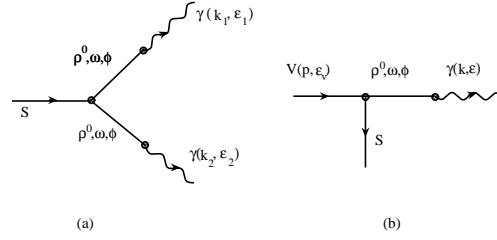


Fig. 1. Feynman diagrams for (a)  $S \rightarrow \gamma\gamma$  and (b)  $V \rightarrow S\gamma$ .

and  $X_S$  takes on the specific forms:

$$\begin{aligned}
 X_\sigma &= \frac{4}{9}\beta_A (\sqrt{2}s - 4c) + \frac{8}{3}\beta_B (c - \sqrt{2}s) , \\
 X_{f_0} &= -\frac{4}{9}\beta_A (\sqrt{2}c + 4s) + \frac{8}{3}\beta_B (\sqrt{2}c + s) , \\
 X_{a_0} &= \frac{4\sqrt{2}}{3}\beta_A .
 \end{aligned} \tag{2.4}$$

Here  $\alpha = e^2/(4\pi)$ ,  $s = \sin \theta_S$  and  $c = \cos \theta_S$  where the scalar mixing angle,  $\theta_S$  is defined from<sup>5)</sup>

$$\begin{pmatrix} \sigma \\ f_0 \end{pmatrix} = \begin{pmatrix} c & -s \\ s & c \end{pmatrix} \begin{pmatrix} N_3^3 \\ (N_1^1 + N_2^2)/\sqrt{2} \end{pmatrix} . \tag{2.5}$$

Furthermore ideal mixing for the vectors, with  $\rho^0 = (\rho_1^1 - \rho_2^2)/\sqrt{2}$ ,  $\omega = (\rho_1^1 + \rho_2^2)/\sqrt{2}$ ,  $\phi = \rho_3^3$ , was assumed for simplicity.

Similarly, the decay matrix element for  $V \rightarrow S\gamma$  is written as  $(e/\tilde{g})C_V^S \times [p \cdot k \epsilon_V \cdot \epsilon - p \cdot \epsilon k \cdot \epsilon_V]$ . It is related to the width by

$$\Gamma(V \rightarrow S\gamma) = \frac{\alpha}{3} |k_V^S|^3 \left| \frac{C_V^S}{\tilde{g}} \right|^2 , \tag{2.6}$$

where  $k_V^S = (m_V^2 - m_S^2)/(2m_V)$  is the photon momentum in the  $V$  rest frame. For the energetically allowed  $V \rightarrow S\gamma$  processes we have

$$\begin{aligned}
 C_\phi^{f_0} &= \frac{2\sqrt{2}}{3}\beta_A c - \frac{4}{3}\beta_B (\sqrt{2}c + s) + \frac{\sqrt{2}}{3}\beta_C (c - \sqrt{2}s) , \\
 C_\phi^\sigma &= -\frac{2\sqrt{2}}{3}\beta_A s - \frac{4}{3}\beta_B (c - \sqrt{2}s) - \frac{2}{3}\beta_C \left( c + \frac{1}{\sqrt{2}}s \right) , \\
 C_\phi^{a_0} &= \sqrt{2}(\beta_C - 2\beta_A) , \\
 C_\omega^\sigma &= \frac{2\sqrt{2}}{3}\beta_A (c + \sqrt{2}s) + \frac{2\sqrt{2}}{3}\beta_B (c - \sqrt{2}s) - \frac{2}{3}\beta_C (\sqrt{2}c + s) , \\
 C_{\rho^0}^\sigma &= -2\sqrt{2}\beta_A c + 2\sqrt{2}\beta_B (c - \sqrt{2}s) .
 \end{aligned} \tag{2.7}$$

In addition, the same model predicts amplitudes for the energetically allowed  $S \rightarrow V\gamma$  processes:  $f_0 \rightarrow \omega\gamma$ ,  $f_0 \rightarrow \rho^0\gamma$ ,  $a_0^0 \rightarrow \omega\gamma$ ,  $a_0^0 \rightarrow \rho^0\gamma$  and, if  $\kappa^0$  is sufficiently

heavy  $\kappa^0 \rightarrow K^{*0}\gamma$ . The corresponding width is

$$\Gamma(S \rightarrow V\gamma) = \alpha |k_S^V|^3 \left| \frac{D_S^V}{\tilde{g}} \right|^2, \quad (2.8)$$

where  $k_S^V = (m_S^2 - m_V^2)/(2m_S)$  and

$$\begin{aligned} D_{f_0}^\omega &= \frac{2}{3}\beta_A (-2c + \sqrt{2}s) + \frac{2}{3}\beta_B (2c + \sqrt{2}s) + \frac{2}{3}\beta_C (c - \sqrt{2}s), \\ D_{f_0}^{\rho^0} &= -2\sqrt{2}\beta_A s + 2\beta_B (2c + \sqrt{2}s), \\ D_{a_0}^\omega &= 2\beta_C, \\ D_{a_0}^{\rho^0} &= \frac{4}{3}\beta_A, \\ D_{\kappa^0}^{K^{*0}} &= -\frac{8}{3}\beta_A. \end{aligned} \quad (2.9)$$

### §3. Results

Let us summarize main points of the results obtained in Ref. 5) together with new results from our recent analysis.<sup>11)</sup>

We should stress again that all the different decay amplitudes are described by the parameters  $\beta_A$ ,  $\beta_B$  and  $\beta_C$ . Below we shall first illustrate the procedure for the model of a single putative scalar nonet.<sup>8)</sup>

We determine the value of  $\beta_A$  from the  $a_0 \rightarrow \gamma\gamma$  process. Substituting  $\Gamma_{\text{exp}}(a_0 \rightarrow \gamma\gamma) = (0.28 \pm 0.09) \text{ keV}$  (obtained using<sup>12)</sup>  $B(a_0 \rightarrow K\bar{K})/B(a_0 \rightarrow \eta\pi) = 0.177 \pm 0.024$ ) into Eqs. (2.3) and (2.4) yields

$$\beta_A = (0.72 \pm 0.12) \text{ GeV}, \quad (3.1)$$

where we assumed positive in sign. By using this value, the value of  $\beta_C$  is determined from  $\Gamma_{\text{exp}}(\phi \rightarrow a_0\gamma) = (0.47 \pm 0.07) \text{ keV}$  (obtained by assuming  $\phi \rightarrow \eta\pi^0\gamma$  is dominated by  $\phi \rightarrow a_0\gamma$ ) and Eq. (2.7) as

$$\beta_C = (7.7 \pm 0.5, -4.8 \pm 0.5) \text{ GeV}^{-1}. \quad (3.2)$$

We stress that the values of  $\beta_A$  and  $\beta_C$  obtained above are independent of the mixing angle  $\theta_S$ . It should be noticed that  $|\beta_A|$  is almost an order of magnitude smaller than  $|\beta_C|$ . As we can see from Eq. (2.9), the amplitude  $D_{a_0}^\omega$  is given by  $\beta_C$  while  $D_{a_0}^{\rho^0}$  is given by only  $\beta_A$ . Then, the large hierarchy between  $\beta_C$  and  $\beta_A$  implies that there is a large hierarchy between  $\Gamma(a_0 \rightarrow \omega\gamma)$  and  $\Gamma(a_0 \rightarrow \rho\gamma)$ . Actually, by using the values of  $\beta_A$  and  $\beta_C$  given in Eqs. (3.1) and (3.2), they are estimated as

$$\begin{aligned} \Gamma(a_0 \rightarrow \omega\gamma) &= (641 \pm 87, 251 \pm 54) \text{ keV}, \\ \Gamma(a_0 \rightarrow \rho\gamma) &= 3.0 \pm 1.0 \text{ keV}. \end{aligned} \quad (3.3)$$

This implies that there is a large hierarchy between  $\Gamma(a_0 \rightarrow \omega\gamma)$  and  $\Gamma(a_0 \rightarrow \rho\gamma)$  which is caused by an order of magnitude difference between  $|\beta_C|$  and  $|\beta_A|$ .

We next determine the value of  $\beta_B$  from the  $f_0 \rightarrow \gamma\gamma$  process.  $X_{f_0}$  in Eq. (2.4) depends on  $\beta_B$  as well as on  $\beta_A$  and the mixing angle  $\theta_S$ . Here we take the mixing angle as

$$\theta_S \simeq -20^\circ, \quad (3.4)$$

which is characteristic of  $qq\bar{q}\bar{q}$  type scalars.<sup>8)</sup> By using this and the value of  $\beta_A$  in Eq. (3.1),  $\Gamma_{\text{exp}}(f_0 \rightarrow \gamma\gamma) = 0.39 \pm 0.13 \text{ keV}$  yields

$$\beta_B = (0.61 \pm 0.10, -0.62 \pm 0.10) \text{ GeV}^{-1}. \quad (3.5)$$

This implies that  $|\beta_B|$  is on the order of  $|\beta_A|$ , and almost an order of magnitude smaller than  $|\beta_C|$ . Equation (2.9) shows that  $D_{f_0}^\omega$  includes  $\beta_C$  while  $D_{f_0}^\rho$  does not. Thus, we have a large hierarchy between decay widths of  $f_0 \rightarrow \omega\gamma$  and  $f_0 \rightarrow \rho\gamma$ : The typical predictions are given by

$$\begin{aligned} \Gamma(f_0 \rightarrow \omega\gamma) &= (88 \pm 17) \text{ keV}, \\ \Gamma(f_0 \rightarrow \rho\gamma) &= (3.3 \pm 2.0) \text{ keV}. \end{aligned} \quad (3.6)$$

This implies that there is a large hierarchy between  $\Gamma(f_0 \rightarrow \omega\gamma)$  and  $\Gamma(f_0 \rightarrow \rho\gamma)$  which is caused by the fact that  $|\beta_C|$  is an order of magnitude larger than  $|\beta_A|$  and  $|\beta_B|$ .

Let us check the dependence of the above results on the choice of the scalar mixing angle  $\theta_S$ . In Ref. 8), the value of  $\theta_S \simeq -90^\circ$  was obtained as another solution to reproduce the masses of the lightest scalar nonet, although the predicted value of  $f_0$ - $\pi$ - $\pi$  coupling is much larger than the value obtained in Ref. 13) by fitting to the  $\pi\pi$  scattering amplitude

As we stressed above, the values of  $\beta_A$  and  $\beta_C$  are independent of the scalar mixing angle  $\theta_S$ . The value of  $\beta_B$  determined from  $\Gamma(f_0 \rightarrow \gamma\gamma)$  becomes

$$\beta_B = (1.1 \pm 0.1, 0.12 \pm 0.13) \text{ GeV}^{-1}. \quad (3.7)$$

Then the typical predictions for  $\Gamma(f_0 \rightarrow \omega\gamma)$  and  $\Gamma(f_0 \rightarrow \rho\gamma)$  are given by

$$\begin{aligned} \Gamma(f_0 \rightarrow \omega\gamma) &= (86 \pm 16) \text{ keV}, \\ \Gamma(f_0 \rightarrow \rho\gamma) &= (3.4 \pm 3.2) \text{ keV}. \end{aligned} \quad (3.8)$$

These predictions are very close to the ones in Eq. (3.6). This can be understood by the following consideration: From the expression of  $D_{f_0}^\omega$  in Eq. (2.9), we can see that it is dominated by the term including  $\beta_C$  which is proportional to  $(\cos\theta_S - \sqrt{2}\sin\theta_S)$ . Then, the approximate relation

$$\cos(-20^\circ) - \sqrt{2}\sin(-20^\circ) \simeq \cos(-90^\circ) - \sqrt{2}\sin(-90^\circ) \simeq 1.4 \quad (3.9)$$

implies that the value of  $D_{f_0}^\omega$  for  $\theta_S = -90^\circ$  is close to that for  $\theta_S = -20^\circ$ , and thus  $\Gamma(f_0 \rightarrow \omega\gamma)$  for  $\theta_S = -90^\circ$  to that for  $\theta_S = -20^\circ$ . As for  $\Gamma(f_0 \rightarrow \rho\gamma)$  we note that the following relation is satisfied for  $X_{f_0}$  in Eq. (2.4) and  $D_{f_0}^{\rho^0}$  in Eq. (2.9):

$$3X_{f_0} - 2\sqrt{2}D_{f_0}^{\rho^0} = -\frac{4}{3}\sqrt{2}\beta_A(c - \sqrt{2}s). \quad (3.10)$$

Since we use the experimental value of  $\Gamma(f_0 \rightarrow \gamma\gamma)$ , i.e.,  $X_{f_0}$  as an input, this relation implies that the predicted value of  $\Gamma(f_0 \rightarrow \rho\gamma)$  for  $\theta_S = -90^\circ$  is roughly equal to that for  $\theta_S = -20^\circ$ . From the above consideration we conclude that there is a large hierarchy between  $\Gamma(f_0 \rightarrow \omega\gamma)$  and  $\Gamma(f_0 \rightarrow \rho\gamma)$  for  $\theta_S = -20^\circ$  and  $\theta_S = -90^\circ$ :

$$\Gamma(f_0 \rightarrow \omega\gamma) \gg \Gamma(f_0 \rightarrow \rho\gamma) \quad \text{for } \theta_S = -20^\circ \text{ and } \theta_S = -90^\circ. \quad (3.11)$$

#### §4. Discussion

In this paper we showed the predictions of our model only for  $\Gamma(a_0 \rightarrow \omega\gamma)$ ,  $\Gamma(a_0 \rightarrow \rho\gamma)$ ,  $\Gamma(f_0 \rightarrow \omega\gamma)$  and  $\Gamma(f_0 \rightarrow \rho\gamma)$ . Predictions on several other processes such as  $\Gamma(\sigma \rightarrow \gamma\gamma)$  and  $\Gamma(\phi \rightarrow \sigma\gamma)$  can be seen in Ref. 5).

The value of  $\Gamma(\phi \rightarrow f_0\gamma)$  predicted in Ref. 5) is considerably smaller than the experimental value of  $\Gamma(\phi \rightarrow f_0\gamma)$ . We may need to include the effect of  $K$ -loop which would give a large enhancement as shown in Ref. 3). We leave the analysis with  $K$ -loop corrections to future publications.<sup>11)</sup>

#### Acknowledgements

We are happy to thank N. N. Achasov for a suggested correction to the first version of Ref. 5) and with A. Abdel-Rahiem and A. H. Fariborz for very helpful discussions. D.B. wishes to acknowledge support from the Thomas Jefferson National Accelerator Facility operated by the Southeastern Universities Research Association (SURA) under DOE contract number DE-AC05-84ER40150. The work of J.S. is supported in part by DOE contract DE-FG-02-85ER40231.

#### References

- 1) See, e.g., the confereproceedings, S. Ishida et al, "Possible existence of the sigma meson and its implication to hadron physics", KEK Proceedings 2004, Soryushiron Kenkyu 102, No. 5, 2001; D. Amelin and A.M. Zaitsev, "Hadron spectroscopy", Ninth international conference on hadron spectroscopy, Protvino, Russia (2001).
- 2) M. N. Achasov *et al.* [SND Collaboration], Phys. Lett. B **479**, 53 (2000); R. R. Akhmetshin *et al.* [CMD-2 Collaboration], Phys. Lett. B **462**, 380 (1999); A. Aloisio *et al.* [KLOE Collaboration], arXiv:hep-ex/0107024.
- 3) N. N. Achasov and V. N. Ivanchenko, Nucl. Phys. B **315**, 465 (1989).
- 4) F. E. Close, N. Isgur and S. Kumano, Nucl. Phys. B **389**, 513 (1993); N. N. Achasov and V. V. Gubin, Phys. Rev. D **56**, 4084 (1997); Phys. Rev. D **57**, 1987 (1998); N. N. Achasov, V. V. Gubin and V. I. Shevchenko, Phys. Rev. D **56**, 203 (1997); J. L. Lucio Martinez and M. Napsuciale, Phys. Lett. B **454**, 365 (1999); arXiv:hep-ph/0001136; A. Bramon, R. Escribano, J. L. Lucio M., M. Napsuciale and G. Pancheri, Phys. Lett. B **494**, 221 (2000); F. E. Close and A. Kirk, Phys. Lett. B **515**, 13 (2001); N. N. Achasov, S. A. Devyanin and G. N. Shestakov, Phys. Lett. B **88**, 367 (1979); N. N. Achasov and A. V. Kiselev, Phys. Lett. B **534**, 83 (2002).
- 5) D. Black, M. Harada and J. Schechter, Phys. Rev. Lett. **88**, 181603 (2002); D. Black, M. Harada and J. Schechter, arXiv:hep-ph/0303223.
- 6) C. G. Callan, S. Coleman, J. Wess and B. Zumino, Phys. Rev. **177**, 2247 (1969).
- 7) M. Harada and J. Schechter, Phys. Rev. D **54**, 3394 (1996).
- 8) D. Black, A. H. Fariborz, F. Sannino and J. Schechter, Phys. Rev. D **59**, 074026 (1999).
- 9) See, e.g., M. Bando, T. Kugo and K. Yamawaki, Phys. Rept. **164**, 217 (1988); M. Harada and K. Yamawaki, to appear in Phys. Rept., arXiv:hep-ph/0302103.
- 10) M. Gell-Mann, D. Sharp and W. G. Wagner, Phys. Rev. Lett. **8**, 261 (1962). A general

formulation of the  $PVV$  interaction in a gauge invariant manner was given in: T. Fujiwara, T. Kugo, H. Terao, S. Uehara and K. Yamawaki, *Prog. Theor. Phys.* **73**, 926 (1985); O. Kaymakcalan, S. Rajeev and J. Schechter, *Phys. Rev. D* **30**, 594 (1984); O. Kaymakcalan and J. Schechter, *Phys. Rev. D* **31**, 1109 (1985); P. Jain, R. Johnson, U. G. Meissner, N. W. Park and J. Schechter, *Phys. Rev. D* **37**, 3252 (1988).

- 11) D. Black, M. Harada and J. Schechter, in preparation.
- 12) D. E. Groom *et al.* [Particle Data Group Collaboration], *Eur. Phys. J. C* **15**, 1 (2000) and 2001 partial update for edition 2002 (URL: <http://pdg.lbl.gov>).
- 13) M. Harada, F. Sannino and J. Schechter, *Phys. Rev. D* **54**, 1991 (1996).