

# Characterization of Iron in Lake Towuti Sediment

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## Abstract

Sediments collected from Lake Towuti, an ultramafic-hosted lake in Indonesia, preserve a visible alternating pattern of red and green sediments due to variations in clay mineral and Fe-oxide composition and abundance consistent with changes in iron oxidation state through time. Spectral, mineralogical, and chemical analyses on soils, river, and sediment samples from across the lake and its catchment were carried out to better understand the starting composition of these sediments and the processes that affected them before and after deposition. Despite high Fe abundances in all samples and abundant Fe oxides in lateritic source regions, mineralogical analyses (X-ray diffraction (XRD) and Mössbauer spectroscopy) of the modern lake sediment show almost no well-crystalline iron oxides. In addition, sequential Fe extractions suggest an increasing proportion of easily extractable, poorly crystalline (X-ray amorphous)

14 material with burial depth. XRD, bulk chemistry, and visible-near infrared  
15 (VNIR) spectral reflectance measurements demonstrate that clay mineralogy  
16 and bulk chemistry can be inferred from VNIR data. These results provide  
17 evidence for variations in Fe mineralogy and crystallinity based on location  
18 in this source to sink system. Understanding how the mineralogy and chem-  
19 istry of sediments within a ferruginous lake basin are affected by transport,  
20 chemical alteration, physical alteration, and deposition from source to sink on  
21 Earth, and the degree to which these trends and underlying processes can be  
22 inferred from chemical and spectral properties, may provide useful direction  
23 in assessing paleoenvironmental conditions in other terrestrial lakes as well as  
24 ancient lacustrine environments preserved in the stratigraphic record of Mars.

25 **Keywords:** *Fe oxides; sedimentation; redox-stratified lake; Mars analogue*

## 26 1 Introduction

27 The oxidation state of redox-sensitive elements in lake sediment and cores can  
28 encode information on water column conditions and potentially the relative position  
29 of the oxycline and/or shoreline. Thus, tracking oxidation state changes and mineral  
30 hosts of iron during sediment weathering, transport, deposition, and diagenesis can  
31 elucidate varying redox conditions in modern and ancient depositional environments.

32 Iron is of particular interest as it is ubiquitous in mafic systems found on Earth  
33 and Mars. Here we examine a suite of samples from Lake Towuti, Indonesia, to  
34 characterize the nature of Fe in the sediment of a modern terrestrial redox-stratified  
35 lake. We analyze sediment sampled from soils, rivers, and the lakebed itself in an  
36 attempt to discriminate between processes affecting the mineral hosts of Fe during  
37 weathering in the source region, fluvial transport, and early diagenesis in the sedi-  
38 ment column. We also examine if key mineralogical and chemical trends, including  
39 variations in clay mineralogy, Fe mineralogy and Fe oxidation state, can be inferred  
40 from Mössbauer and near-infrared reflectance spectroscopy of the bulk sediment.  
41 Understanding spatial variations in the mineralogy and chemistry of the uppermost

42 surface sediment across the modern lakebed, and the causes of these variations, can  
43 provide a foundation for interpreting sediments in deeper lake cores. Characterizing  
44 the hosts of Fe across the modern transport and depositional system can thus aid in  
45 the reconstruction of paleoenvironmental conditions in the broader Lake Towuti re-  
46 gion. Information on the nature and evolution of Fe-bearing phases in Lake Towuti  
47 may also provide insight into the chemical and mineralogical evolution of an ancient  
48 redox-stratified lake purported to exist on Mars (Hurowitz et al. 2017). Assessing re-  
49 lationships between bulk chemical and spectral properties can also lend new insight  
50 into how rapid, non-destructive techniques such as visible-near infrared reflectance  
51 spectroscopy may be used to decipher past aqueous and atmospheric conditions in  
52 fine-grained sediments and mudstones.

## 53 **2 Background**

### 54 **2.1 Malili Lakes: Physical hydrological and geological setting**

55 Lake Towuti is the largest lake in the Malili Lakes system, a set of five tectonic  
56 lakes located near the equator on Sulawesi island, Indonesia (2.75°S, 121.5°E, 318  
57 m a.s.l.). It is situated in the East Sulawesi Ophiolite which is composed primar-  
58 ily of unserpentinized harzburgite, dunite, and serpentinized lherzolite that are a  
59 source of large quantities of iron and other metals delivered to the lake (Kadaru-  
60 man et al. 2004). Lake Towuti is fed by several rivers that drain a catchment of  
61 approximately 1280 km<sup>2</sup> (Hasberg et al. 2018). The main inlet, the Mahalona River,  
62 drains two smaller upstream lakes, Matano and Mahalona, which run almost exclu-  
63 sively through a mix of serpentinized and unserpentinized peridotite, as well as some  
64 presently unidentified ultramafic terrain (Costa et al. 2015). The Mahalona River is  
65 also fed by the Lampenisu catchment to the northeast, and sources a large deposit of  
66 coarse-grained sediment associated with a prograding delta extending approximately  
67 10 km into Lake Towuti (Vogel et al. 2015). Several other smaller rivers also feed  
68 into Lake Towuti, with catchments composed of serpentinized and unserpentinized

69 peridotite and, on the east side of the lake, areas of metasediment and limestone  
70 (Costa et al. 2015). The major outlet, the Larona River, drains Lake Towuti to the  
71 west where the bedrock is unidentified ultramafic and minor sandstone/siltstone  
72 (Fig. 1, Costa et al. 2015). Because of its significant depth, ranging up to  $\sim 200$   
73 meters, and tropical setting, Towuti is a thermally and redox-stratified lake (Costa  
74 et al. 2015; Vuillemin et al. 2016). The entire lake system is sulfate-poor (Crowe  
75 et al. 2008; Vuillemin et al. 2016).

## 76 **2.2 Lake Towuti: Studies of sediment and the water column**

77 Lake Towuti possesses a rare continuous record of tropical sedimentation span-  
78 ning several glacial cycles, and the system has been used to study sediment re-  
79 sponse to short- and long-term climate changes (Russell et al. 2014; Vogel et al.  
80 2015; Russell et al. 2016). Sediment in the Lake Towuti catchment is sourced from  
81 a mafic/ultramafic provenance (Kadarusman et al. 2004), thus has elevated concen-  
82 trations of redox sensitive elements. Consequently, the lake sediment may provide  
83 a record of redox conditions in the lake over time. Prior studies have examined  
84 changes during the past  $\sim 60$  kyrs and demonstrate that trends in Fe and other  
85 redox-sensitive trace elements record more oxidizing conditions during dry periods  
86 (Costa et al. 2015). Redox cycling in the system is suggested by the presence of a  
87 down-core pattern of alternating red and green silty clays (Costa et al. 2015; Vogel  
88 et al. 2015). Consistent with this, previous work has shown that that red sediment  
89 is enriched in iron, presumed to be ferric oxides, that form when cold, dry periods  
90 induce deep mixing and oxygenation of the lake, resulting in oxidation of the bot-  
91 tom waters (Costa et al. 2015). This is a useful heuristic model for changes in lake  
92 sediment geochemistry through time, but to date there has been little attempt to  
93 investigate the mineralogy of iron sources and sinks in the present-day lake, nor min-  
94 eralogical changes that accompany the variations in sedimentary Fe concentrations  
95 in the past.

96 Initial studies of modern sedimentation in the lake examined a limited number

107 of surface sediment samples from a transect across the northern part of the lake  
108 (Weber et al. 2015) and samples from the Mahalona delta and nearby rivers (Goudge  
109 et al. 2017). Both studies concluded that the major types of clay minerals in those  
110 samples could be identified from VNIR reflectance spectra and that the strength of  
111 absorption features within the spectra were correlated with chemical composition  
112 of the bulk sample. Recent studies used a larger suite of surface sediment samples  
113 to track sedimentation across the lake, concluding that the Mahalona River is the  
114 main source of serpentine and its influence relative to kaolinite-bearing rivers can be  
115 traced in sample X-ray diffraction (XRD) patterns, mid-infrared (MIR) spectra, and  
116 elemental abundances (Morlock et al. 2018; Hasberg et al. 2018). However, these  
117 studies focused on silicate mineral sources and sinks, not iron oxides. This study  
118 presents analyses for a set of samples used in previous studies (Goudge et al. 2017;  
119 Hasberg et al. 2018; Morlock et al. 2018) that span the Lake Towuti system from  
120 source (catchment soils) to sink (lake surface sediment samples distributed across  
121 the entire basin). We examine if previously reported chemical and clay mineral  
122 trends associated with VNIR spectral features are observed within sediments from  
123 Lake Towuti as a whole. In addition to characterizing clay mineralogy, we also  
124 evaluate differences in the mineral host(s) of Fe between these samples and determine  
125 the extent to which VNIR spectra are sensitive to changes in Fe-mineralogy, bulk  
126 mineralogy, and bulk chemistry.

## 117 **3 Methods**

### 118 **3.1 Sampling**

119 Samples were collected in 2013 (Costa et al. 2015; Goudge et al. 2017) and  
120 during the 2015 Towuti Drilling Project (Russell et al. 2016), and they consist of  
121 three different sample types: lateritic soils from the catchment, inflowing river bed  
122 sediment, and surficial sediment from the lake floor (Fig. 1, Table 1).

123 *Lateritic soils* were collected from six separate, exposed weathered soil profiles.

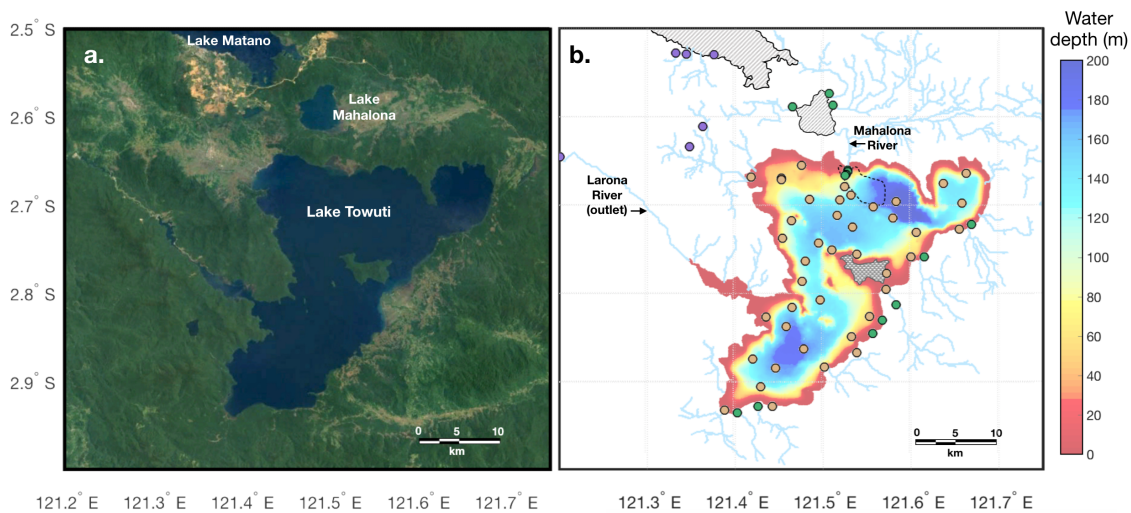


Figure 1: The Malili Lakes System and associated rivers. *a.* Landsat composite image of the Malili Lakes System, including the southeastern tip of Lake Matano, Lake Mahalona, and Towuti. *b.* Bathymetry of Lake Towuti (m). Rivers are in light blue and the delta of coarse-grained material carried in by the Mahalona River is outlined by a dashed line. Brown points designate locations of surface sediment sample locations, green are river bedload samples, and purple are laterite samples.

124 Because of limitations in accessibility, all sampling sites were west or northwest of the  
 125 lake. The 20 soil samples were collected at different depths within the soil profiles,  
 126 spanning the top of the soil bed to bedrock. Some samples were coarse-grained with  
 127 distinct pebbles and many had visible pieces of organic material (Morlock et al.  
 128 2018).

129 *River bedload* samples were collected from twelve separate rivers. The 16 total  
 130 samples were primarily collected near the river mouths, with four samples collected  
 131 along the largest inlet, the Mahalona River (Weber et al. 2015; Goudge et al. 2017).

132 *Lake surface sediment* samples were collected from the sediment/water interface  
 133 using a polnar grab sampler (UWITEC Corp., Austria). The 42 samples analyzed for  
 134 this study were mostly fine-grained (most dominated by silt and clay-size material)  
 135 and relatively homogeneous in appearance (Hasberg et al. 2018; Morlock et al. 2018).

136 *Lake sediment core* samples were collected in 2015 by the International Conti-  
 137 nental Scientific Drilling Program Towuti Drilling Project (ICDP-TDP) from Lake  
 138 Towuti (Russell et al. 2016). These cores were subsampled at the LacCore facility  
 139 at the University of Minnesota in 2016.

Table 1: Sample location information, grain size (from Hasberg et al. 2018), and analyses performed. (- indicates not measured or not applicable. VNIR sp. = VNIR reflectance spectroscopy, Ex. 1 and Ex. 2 are sequential extractions described in Table 2, Mossb. = Mössbauer spectroscopy, XRD = X-ray diffraction, ICP = ICP-AES for elemental abundances.)

Sample No.	Type	Latitude	Longitude	Water depth (m)	Depth in laterite bed (m)	Depth in core (m)	% sand	% silt	% clay	Analyses performed
133	River	-2.573333	121.509050	0.28	-	-	-	-	-	VNIR sp., XRD
134	River	-2.573333	121.509050	0.28	-	-	-	-	-	VNIR sp., XRD
135	River	-2.758320	121.617050	0.70	-	-	-	-	-	VNIR sp., XRD
136	River	-2.927380	121.428300	0.40	-	-	-	-	-	VNIR sp., XRD
137	River	-2.660290	121.529940	1.00	-	-	-	-	-	VNIR sp., XRD
138	River	-2.662050	121.530200	0.50	-	-	-	-	-	VNIR sp., XRD
139	River	-2.665210	121.528220	1.50	-	-	-	-	-	VNIR sp., XRD
140	River	-2.666140	121.527440	2.50	-	-	-	-	-	VNIR sp., XRD
141	River	-2.721880	121.670770	0.80	-	-	-	-	-	VNIR sp., XRD
142	River	-2.721880	121.670770	1.40	-	-	-	-	-	VNIR sp., XRD
143	River	-2.845450	121.558220	0.50	-	-	-	-	-	VNIR sp., XRD
144	River	-2.588817	121.467333	-	-	-	-	-	-	VNIR sp., XRD
145	River	-2.587000	121.513300	0.50	-	-	-	-	-	VNIR sp., XRD
146	River	-2.830231	121.569557	-	-	-	-	-	-	VNIR sp., XRD
147	River	-2.935080	121.405030	0.70	-	-	-	-	-	VNIR sp., XRD
148	River	-2.812530	121.585230	1.00	-	-	-	-	-	VNIR sp., XRD
226	Laterite (bed 1)	-2.528883	121.346800	0.00	0.40	-	-	-	-	VNIR sp., XRD, ICP
227	Laterite (bed 1)	-2.528883	121.346800	0.00	1.00	-	-	-	-	VNIR sp., XRD, ICP
228	Laterite (bed 1)	-2.528883	121.346800	0.00	1.50	-	-	-	-	VNIR sp., XRD, ICP
229	Laterite (bed 1)	-2.528883	121.346800	0.00	2.20	-	-	-	-	VNIR sp., XRD, ICP
230	Laterite (bed 2)	-2.528100	121.334867	0.00	0.10	-	-	-	-	VNIR sp., XRD, ICP
506	Laterite sample 230 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., XRD
518	Laterite sample 230 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD
231	Laterite (bed 2)	-2.528100	121.334867	0.00	2.00	-	-	-	-	VNIR sp., XRD, ICP
507	Laterite sample 231 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp.
519	Laterite sample 231 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp.
232	Laterite (bed 3)	-2.644950	121.203617	0.00	0.30	-	-	-	-	VNIR sp., XRD, ICP
233	Laterite (bed 3)	-2.644950	121.203617	0.00	1.00	-	-	-	-	VNIR sp., XRD, ICP
234	Laterite (bed 4)	-2.653533	121.350733	0.00	0.40	-	-	-	-	VNIR sp., XRD, ICP
235	Laterite (bed 4)	-2.643533	121.350733	0.00	1.20	-	-	-	-	VNIR sp., XRD, ICP
508	Laterite sample 235 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp.
520	Laterite sample 235 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp.
236	Laterite (bed 4)	-2.643533	121.350733	0.00	0.05	-	-	-	-	VNIR sp., XRD, ICP
237	Laterite (bed 5)	-2.611200	121.365600	0.00	0.20	-	-	-	-	VNIR sp., XRD, ICP
238	Laterite (bed 5)	-2.611200	121.365600	0.00	1.50	-	-	-	-	VNIR sp., XRD, ICP, Mossb.
509	Laterite sample 238 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., XRD
521	Laterite sample 238 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD
239	Laterite (bed 5)	-2.611200	121.365600	0.00	2.5	-	-	-	-	VNIR sp., XRD, ICP
240	Laterite (bed 6)	-2.529517	121.378400	0.00	0.10	-	-	-	-	VNIR sp., XRD, ICP
241	Laterite (bed 6)	-2.529517	121.378400	0.00	1.00	-	-	-	-	VNIR sp., XRD, ICP
242	Laterite (bed 6)	-2.529517	121.378400	0.00	1.50	-	-	-	-	VNIR sp., XRD, ICP
243	Laterite (bed 6)	-2.529517	121.378400	0.00	2.50	-	-	-	-	VNIR sp., XRD, ICP
244	Laterite (bed 6)	-2.529517	121.378400	0.00	3.50	-	-	-	-	VNIR sp., XRD, ICP
245	Laterite (bed 6)	-2.529517	121.378400	0.00	3.00	-	-	-	-	VNIR sp., XRD, ICP

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Sample No.	Type	Latitude	Longitude	Water depth (m)	Depth in laterite bed (m)	Depth in core (m)	% sand	% silt	% clay	Analyses performed
247	Surface	-2.668183	121.420633	1.99	-	-	0.034	0.795	0.170	VNIR sp., XRD, ICP
248	Surface	-2.669217	121.454833	72.54	-	-	0.001	0.715	0.284	VNIR sp., XRD, ICP
249	Surface	-2.654700	121.477517	41.80	-	-	0.060	0.690	0.250	VNIR sp., XRD, ICP
250	Surface	-2.671183	121.454833	80.37	-	-	0.008	0.732	0.261	VNIR sp., XRD, ICP, Mossb.
251	Surface	-2.678817	121.526600	129.01	-	-	0.050	0.787	0.163	VNIR sp., XRD, ICP
252	Surface	-2.694317	121.520983	175.81	-	-	0.016	0.752	0.232	VNIR sp., XRD, ICP
253	Surface	-2.688333	121.533500	179.13	-	-	0.014	0.779	0.207	VNIR sp., XRD, ICP
254	Surface	-2.695733	121.585100	271.82	-	-	0.007	0.689	0.304	VNIR sp., XRD, ICP
255	Surface	-2.693483	121.486783	173.87	-	-	0.000	0.692	0.308	VNIR sp., XRD, ICP
256	Surface	-2.701850	121.559167	236.27	-	-	0.003	0.729	0.268	VNIR sp., XRD, ICP
510	Surface sample 256 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp.
522	Surface sample 256 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp.
257	Surface	-2.675250	121.638417	184.54	-	-	0.000	0.585	0.415	VNIR sp., XRD, ICP
258	Surface	-2.697633	121.659883	182.22	-	-	0.014	0.662	0.324	VNIR sp., XRD, ICP
259	Surface	-2.727433	121.656883	1.01	-	-	0.545	0.361	0.094	VNIR sp., XRD, ICP
511	Surface sample 259 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., XRD
523	Surface sample 259 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD
260	Surface	-2.714633	121.581683	240.47	-	-	0.000	0.674	0.326	VNIR sp., XRD, ICP
261	Surface	-2.711333	121.518033	207.21	-	-	0.007	0.761	0.232	VNIR sp., XRD, ICP
262	Surface	-2.717317	121.466550	142.94	-	-	0.000	0.639	0.361	VNIR sp., XRD, ICP
263	Surface	-2.724583	121.535833	213.48	-	-	0.015	0.730	0.255	VNIR sp., XRD, ICP
264	Surface	-2.737617	121.456067	113.75	-	-	0.000	0.647	0.353	VNIR sp., XRD, ICP
265	Surface	-2.663617	121.664683	95.07	-	-	0.004	0.658	0.338	VNIR sp., XRD, ICP
266	Surface	-2.742833	121.497067	182.03	-	-	0.000	0.643	0.357	VNIR sp., XRD, ICP, Mossb.
267	Surface	-2.730683	121.607867	118.22	-	-	0.001	0.670	0.330	VNIR sp., XRD, ICP
268	Surface	-2.758267	121.601700	35.60	-	-	0.605	0.355	0.039	VNIR sp., XRD, ICP
269	Surface	-2.795283	121.573617	1.42	-	-	0.333	0.554	0.113	VNIR sp., XRD, ICP
270	Surface	-2.777133	121.573917	3.84	-	-	0.726	0.239	0.035	VNIR sp., XRD, ICP, Mossb.
512	Surface sample 270 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp.
524	Surface sample 270 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., Mossb.
271	Surface	-2.755533	121.540467	137.23	-	-	0.001	0.681	0.318	VNIR sp., XRD, ICP
272	Surface	-2.750450	121.512283	190.65	-	-	0.000	0.681	0.319	VNIR sp., XRD, ICP
273	Surface	-2.763400	121.481800	183.57	-	-	0.000	0.644	0.356	VNIR sp., XRD, ICP
274	Surface	-2.786183	121.478083	153.82	-	-	0.000	0.621	0.379	VNIR sp., XRD, ICP
275	Surface	-2.815450	121.466950	187.85	-	-	0.000	0.572	0.428	VNIR sp., XRD, ICP
276	Surface	-2.826200	121.437283	35.98	-	-	0.000	0.553	0.447	VNIR sp., XRD, ICP
277	Surface	-2.874033	121.422017	171.33	-	-	0.000	0.586	0.414	VNIR sp., XRD, ICP
513	Surface sample 277 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., XRD
525	Surface sample 277 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD
278	Surface	-2.932000	121.390400	0.54	-	-	0.421	0.462	0.116	VNIR sp., XRD, ICP
279	Surface	-2.905200	121.431267	161.28	-	-	0.029	0.621	0.351	VNIR sp., XRD, ICP
280	Surface	-2.884450	121.448050	235.65	-	-	0.000	0.601	0.399	VNIR sp., XRD, ICP
281	Surface	-2.862533	121.479833	241.34	-	-	0.000	0.642	0.358	VNIR sp., XRD, ICP
282	Surface	-2.807383	121.498967	184.13	-	-	0.000	0.657	0.343	VNIR sp., XRD, ICP
283	Surface	-2.883300	121.503900	15.78	-	-	0.103	0.706	0.191	VNIR sp., XRD, ICP
514	Surface sample 283 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp.
526	Surface sample 283 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp.
284	Surface	-2.848783	121.534367	67.92	-	-	0.011	0.757	0.232	VNIR sp., XRD, ICP
285	Surface	-2.825667	121.554883	63.61	-	-	0.073	0.827	0.100	VNIR sp., XRD, ICP
286	Surface	-2.837217	121.459983	234.13	-	-	0.000	0.580	0.420	VNIR sp., XRD, ICP, Mossb.
515	Surface sample 286 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., XRD

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Sample No.	Type	Latitude	Longitude	Water depth (m)	Depth in laterite bed (m)	Depth in core (m)	% sand	% silt	% clay	Analyses performed
527	Surface sample 286 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, Mossb.
287	Surface	-2.927383	121.444717	1.37	-	-	0.782	0.199	0.019	VNIR sp., XRD, ICP
516	Surface sample 287 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp.
528	Surface sample 287 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp.
288	Surface	-2.867050	121.540550	1.01	-	-	0.591	0.343	0.066	VNIR sp., XRD, ICP
97	Core	-2.717480	121.514980	-	-	2.36	-	-	-	VNIR sp., XRD, Mossb.
489	Core sample 97 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., ICP
497	Core sample 97 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, ICP
99	Core	-2.717480	121.514980	-	-	7.72	-	-	-	VNIR sp., XRD, Mossb.
490	Core sample 99 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., ICP
498	Core sample 99 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, ICP
104	Core	-2.717480	121.514980	-	-	14.78	-	-	-	VNIR sp., XRD
491	Core sample 104 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., ICP
499	Core sample 104 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, ICP
113	Core	-2.717480	121.514980	-	-	26.52	-	-	-	VNIR sp., XRD
492	Core sample 113 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., ICP
500	Core sample 113 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, ICP
114	Core	-2.717480	121.514980	-	-	27.12	-	-	-	VNIR sp., XRD
493	Core sample 114 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., ICP
501	Core sample 114 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, ICP
116	Core	-2.717480	121.514980	-	-	30.16	-	-	-	VNIR sp., XRD
494	Core sample 116 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., ICP
502	Core sample 116 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, ICP
118	Core	-2.717480	121.514980	-	-	32.63	-	-	-	VNIR sp., XRD
495	Core sample 118 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., ICP
503	Core sample 118 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, ICP
121	Core	-2.717480	121.514980	-	-	75.84	-	-	-	VNIR sp., XRD
96	Core sample 121 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., ICP
504	Core sample 121 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, ICP
122	Core	-2.717480	121.514980	-	-	76.59	-	-	-	VNIR sp., XRD
505	Core sample 122 after Extraction 1	-	-	-	-	-	-	-	-	Ex. 1, VNIR sp., ICP
517	Core sample 122 after Extraction 2	-	-	-	-	-	-	-	-	Ex. 1 & 2, VNIR sp., XRD, ICP

## 140 **3.2 Reflectance spectroscopy**

141 The mineralogy of each sample was determined by reflectance spectroscopy  
142 (VNIR-IR wavelengths) and powder X-ray diffraction. Samples were relatively fine-  
143 grained (spanning sand to clay; for full details on grain size see Goudge et al. 2017)  
144 and so were lightly powdered using a corundum mortar and pestle. Reflectance  
145 spectra were acquired for dried, powdered samples over a wavelength range of 0.35-  
146 4.75  $\mu\text{m}$  using a combination of spectrometers in the NASA Reflectance Experiment  
147 LABoratory (RELAB) facility housed at Brown University. An Analytical Spectral  
148 Devices (ASD) FieldSpec3 spectroradiometer was used for the visible-near infrared  
149 (VIS-NIR, 0.35-2.5  $\mu\text{m}$ ) range and a Thermo Nexus 870 FTIR spectrometer was  
150 used for the near-infrared to mid-infrared (NIR-MIR, 0.8-4.75  $\mu\text{m}$ ) range. ASD  
151 reflectance spectra were acquired relative to a white Spectralon standard, whereas  
152 FTIR reflectance data were acquired relative to diffuse gold (both standards from  
153 Labsphere). The two sets of spectral data were spliced together at 1.7  $\mu\text{m}$  to provide  
154 a continuous spectrum.

155 VNIR spectral parameters, including band depth, were calculated for absorptions  
156 characteristic of certain minerals common to the Malili Lakes system (Fig. 2). Band  
157 depth is a measurement of the strength of an absorption feature and was calculated  
158 based on the method of Clark and Roush (1984). A variety of factors can influence  
159 the strength of an absorption band in reflectance spectra, including particle size and  
160 strong spectral contrast between absorbing species (e.g., VNIR-transparent minerals  
161 such as quartz mixed with highly absorbing, opaque minerals such as ilmenite or  
162 magnetite). All samples studied here were ground to roughly equivalent particle  
163 size and lack significant abundances of opaque phases (weak or absent peaks for  
164 magnetite and ilmenite in XRD data), thus band depth values (absorption strength)  
165 may provide a reasonable proxy for relative mineral abundance. Major mineral  
166 phases considered in the analysis of spectral data are discussed below and were  
167 based on previous studies (Goudge et al. 2017; Weber et al. 2015), XRD patterns,  
168 and qualitative analysis of the spectra themselves.

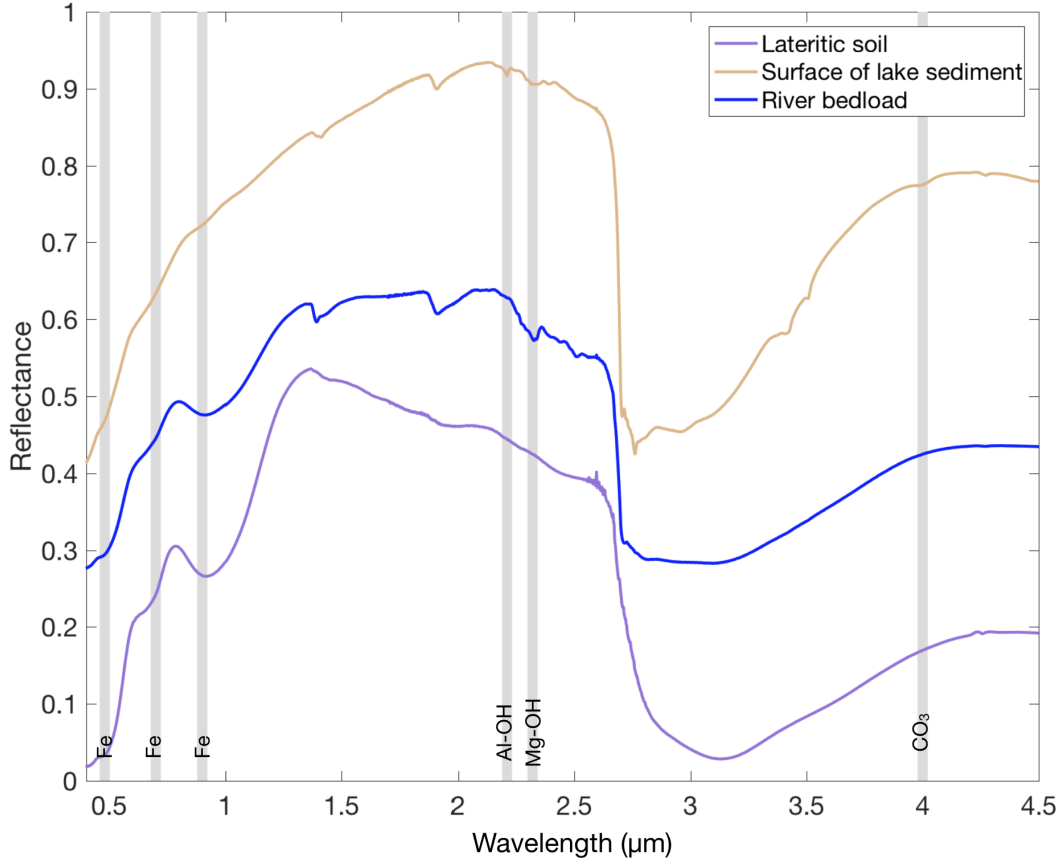


Figure 2: A representative spectrum from each sample type (soil, river, and surface sediment), offset for clarity. Important absorptions are highlighted at 0.48, 0.7, and 0.9  $\mu\text{m}$  (Fe) 2.21  $\mu\text{m}$  (Al-OH), 2.32  $\mu\text{m}$  (Mg-OH), and 4.0  $\mu\text{m}$  (carbonate).

169 *Mg serpentine*: Trioctahedral Mg phyllosilicates in the serpentine group (e.g.  
 170 antigorite, chrysotile, lizardite) were identified based on a prominent asymmetric ab-  
 171 sorption centered at  $\sim 2.32 \mu\text{m}$  caused by vibrations (combination bend and stretch)  
 172 of the Mg-OH bond (Weber et al. 2015).

173 *Al clays*: Absorptions near  $\sim 2.2 \mu\text{m}$  are caused by combination stretch and bend  
 174 of OH and metal-OH (Weber et al. 2015). The precise location of the band at 2.21  
 175  $\mu\text{m}$  and the asymmetric shoulder near 2.16  $\mu\text{m}$  seen in many sample spectra from  
 176 previous studies of Lake Towuti sediment are most consistent with Al-OH vibrations  
 177 in kaolinite, whereas other Al-bearing phyllosilicates (e.g. montmorillonite, illite,  
 178 muscovite) exhibit a broader, more symmetric absorption in this region (Clark et al.  
 179 1990; Bishop et al. 2008; Weber et al. 2015; Goudge et al. 2017).

180 *Fe smectite*: Spectra of the dioctahedral ferric smectite nontronite exhibit a  
 181 characteristic absorption at 2.28  $\mu\text{m}$  caused by vibration of the Fe-OH bond (Clark

182 et al. 1990; Bishop et al. 2008; Weber et al. 2015). Nontronite was not specifically  
183 analyzed in previous studies of more limited sample sets, but all spectra in this  
184 study were examined for the possible presence of this mineral.

185 *Other Fe absorptions:* The presence of  $\text{Fe}^{2+/3+}$  in octahedral and tetrahedral  
186 coordination can give rise to a wide range of complex and sometimes overlapping  
187 absorptions in the 0.35-2.5  $\mu\text{m}$  wavelength range (Burns, 1993). These features are  
188 due to electronic absorptions (e.g. intervalence charge transfer, oxygen-metal charge  
189 transfer, and crystal field splitting). Features associated with  $\text{Fe}^{2+/3+}$  in oxides and  
190 poorly crystalline materials of interest to this study include absorptions at 0.48, 0.7,  
191 and 0.9 $\mu\text{m}$  (see Fig. 2). Spectra for the subset of samples processed for sequential  
192 iron extraction (described below) were also acquired for the dried samples after  
193 each extraction step. This allows for an assessment of changes in absorptions due  
194 to relative changes in  $\text{Fe}^{2+/3+}$  during the extraction, which helps to link observed  
195 spectral absorptions to the presence of X-ray amorphous versus crystalline Fe phases.

196 *Carbonate:* A broad  $\text{CO}_3$  absorption ( $\nu_1 + \nu_3$  vibration mode) is located at 3.8-  
197 4.0  $\mu\text{m}$  (Gaffey et al. 1993; Sutter et al. 2007). This band is stronger than overtone  
198 absorptions observed at shorter wavelengths ( $<2.5 \mu\text{m}$ ), does not overlap with OH  
199 absorptions associated with clay minerals, and is not as affected by  $\text{H}_2\text{O}$  or  $\text{CO}_2$   
200 absorptions (Wagner and Schade 1996).

### 201 **3.3 X-ray Diffraction**

202 Powder X-ray diffraction (XRD) measurements were made for all samples to  
203 provide qualitative information on phase identification, including measurements of  
204 d-spacing for clay minerals. Dried, powdered samples were analyzed on a Bruker  
205 D2 Phaser XRD with a  $\text{Cu K}\alpha$  source. Sample cups were filled without packing to  
206 reduce preferential orientation. XRD patterns were analyzed for presence/absence  
207 of relevant peaks, in part using the patterns of synthesized nanophase iron (oxy-  
208 hydr)oxides discussed in Sklute et al. (2018) and reference patterns from data in the  
209 American Mineralogist Crystal Structure Database (AMCSD) using the CrystalD-

210 diffract software by CrystalMaker database (Fig. 3). Measurements were taken from  
211 5 to 90 °2θ at 0.02029° step size.

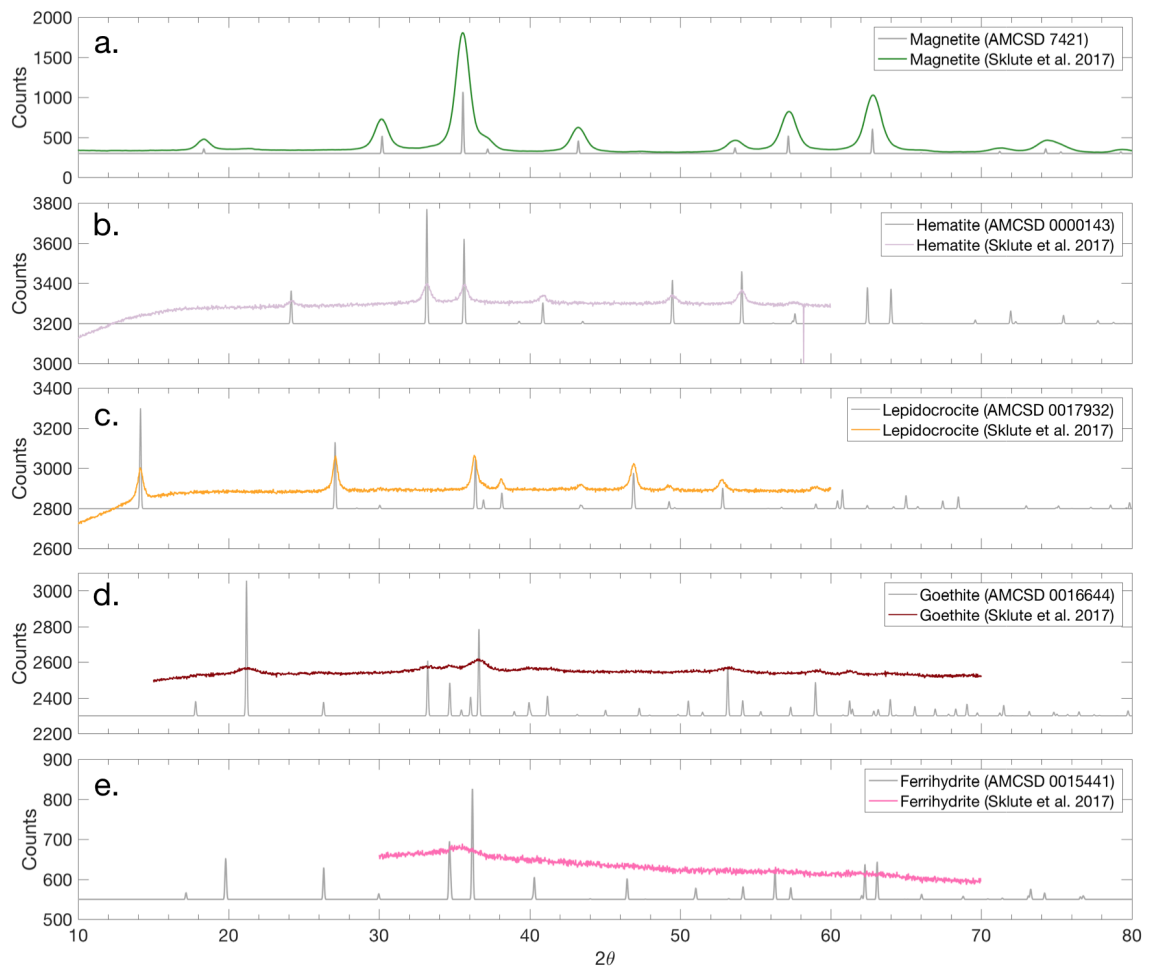


Figure 3: XRD patterns of iron oxides synthesized in Sklute et al. (2018) (colored lines), with AMCSD reference patterns in gray: *a.* magnetite, *b.* hematite, *c.* lepidocrocite, *d.* goethite, and *e.* ferrihydrite

### 212 3.4 ICP-AES

213 Major and minor elemental abundances were determined by inductively coupled  
214 plasma-atomic emission spectrometry (ICP-AES) at Brown University. Major (Al,  
215 Ca, Cr, Fe, K, Mg, Mn, Na, Ni, P, Si, Ti) and trace (Co, Cr, Cu, Mo, Ni, Sc, Sr, V,  
216 Zn, Zr) elements were analyzed in the 42 surface samples after they were dissolved  
217 using flux fusion (Murray et al. 2000). To do so, each sample ( $0.04 \pm 0.005g$ ) was  
218 gently mixed with lithium metaborate flux ( $0.16 \pm 0.005g$ ) in a graphite crucible  
219 and fused at 1050°C for 10 minutes. Each fused bead was then dropped into 20

220 mL of 10% nitric acid ( $\text{HNO}_3$ ), shaken for 30-60 minutes, filtered with a 0.45  $\mu\text{m}$   
 221 Gelman filter, and 5 mL of the final fused solution diluted in 35 mL of 10%  $\text{HNO}_3$ .  
 222 The same process was done for standard reference materials (RGM-1, NIST1646a,  
 223 BIR-1, BHVO-2, NIST2711, BCR-2, and DTS-2B).

224 Once diluted, the samples were analyzed on a Jobin Yvon JY2000 ICP-AES.  
 225 Concentration data are calibrated to the fluxed standard reference materials run  
 226 with each batch. A drift solution comprising a small amount of each diluted sample  
 227 was analyzed every 12 measurements to monitor and correct for any change in signal.  
 228 The detector did not saturate during these measurements.

### 229 3.5 Sequential Iron Extraction

230 To determine the relative amounts of highly reactive iron (e.g. amorphous Fe  
 231 phases, ferrihydrite, lepidocrocite) and crystalline iron oxy/hydroxides (e.g. hematite,  
 232 goethite) in the samples, a two-stage sequential iron extraction was carried out on  
 233 several laterite and surface sediment samples using a protocol modified from Poulton  
 234 and Canfield (2005), described in Table 2. Fe in silicates was not removed by these  
 235 extractions.

236

Table 2: Sequential Fe Extraction Protocol

Extraction	Target Phases	Duration
Hydroxylamine HCl	Amorphous iron oxides (e.g. ferrihydrite)	48 hours
Sodium dithionite	Crystalline iron oxides (goethite and hematite)	2 hours

237 The hydroxylamine HCl extraction was created by mixing 69.49 g Hy-HCl into  
 238 1 L of a 25% acetic acid solution. The extraction was performed in 50 mL plastic  
 239 falcon tubes, where 100 mg of sediment and 10 mL hydroxylamine HCl were mixed  
 240 and placed on a shaker table and reacted for 48 hours at room temperature.

241 The sodium dithionite extraction was prepared by mixing 50 g of sodium dithion-  
 242 ite and 51.6 g sodium citrate in 1 L DI water. Acetic acid was then added until

243 the pH reached 4.8. Samples residues from the hydroxylamine HCl extraction were  
244 reacted with 10 mL of the sodium dithionite in falcon tubes on a shaker table for 2  
245 hours at room temperature.

246 After each extraction, the falcon tubes were centrifuged and supernatant poured  
247 into a vial. The residual sediment was then washed with deionized water three  
248 times, centrifuged each time, and poured into the same vial. The residual sediment  
249 was freeze dried and spectra and XRD patterns were collected to assess whether the  
250 extractions were truly removing the relevant Fe oxides. The supernatant was dried  
251 at 60°C, rehydrated in 10 mL 2% nitric acid, diluted 1:100 in 2% nitric acid, and  
252 measured for total iron concentration on the ICP-AES.

### 253 **3.6 Mössbauer spectroscopy**

254 Mössbauer spectroscopy is sensitive to small energetic changes around Fe atoms  
255 and can be used to characterize coordination environments of Fe oxides in complex,  
256 heterogeneous samples (Sklute et al. in review). Mössbauer measurements were ac-  
257 quired at Mt. Holyoke College with the sample at 4 K, 130 K, and 295 K on a Web  
258 Research Co. (now See Co.) W100 spectrometer using a  $\sim 75\text{-}65$  mCi  $^{57}\text{Co}$  source  
259 in rhodium. Low temperature spectra were obtained using a Janis Research Co.  
260 Model 850 4 K closed-cycle helium compression system. The spectra were fit using  
261 the Mex\_disd program and the fits were then used to derive several Mössbauer pa-  
262 rameters (Dyar et al. 2006; Sklute et al. in review). These include center shift (CS),  
263 which reflects the s-electronic charge density, and is affected by bond characteristics,  
264 valence state, and coordination environment. This shift is presented in Mössbauer  
265 spectra as a velocity (mm/s) shift relative to  $\alpha$ -Fe foil. Another relevant parameter  
266 is quadrupole splitting (QS) of nuclear energy levels, which creates the distinctive  
267 doublets seen in Mössbauer spectra. Changes in Fe valence state or changes to the  
268 crystal lattice that affect the coordination or bonding environment will affect this  
269 splitting and the observed doublets and sextets. Finally, the hyperfine field (hff),  
270 which results from the interaction of the nuclear magnetic moment with the net

271 magnetic field at the nucleus, creates the distinct sextet pattern seen in Mössbauer  
272 spectrum. While the hff is sensitive to changes in the bonding environment and the  
273 crystal lattice, it is also sensitive to crystallinity and grain size. For example, grain  
274 size and composition affect the temperature at which the transition from doublet to  
275 sextet occurs, which allows ferric and ferrous Fe in different coordination environ-  
276 ments to be distinguished by collecting spectra at multiple temperatures (4 K, 130  
277 K, and 295 K here).

## 278 4 Results

### 279 4.1 Correlation of spectral characteristics with chemical data

280 Surface sediment samples show strong spatial variations in Al and Mg (Fig. 4*a,b*)  
281 with associated changes in silicate mineralogy, consistent with the results of previous  
282 studies. Values for Al are relatively low in the north near the Mahalona River delta  
283 and increase away from the river mouth and delta with movement toward more  
284 distal parts of the lake. Values for Mg exhibit the opposite trend: Mg abundance  
285 is elevated in the Mahalona River delta and drops quickly with distance from the  
286 delta.

287 Band depth values of the 2.21  $\mu\text{m}$  Al-OH (kaolinite) absorption in spectra of  
288 surface samples are correlated with  $\text{Al}_2\text{O}_3$  abundance (positive coefficient significant  
289 at  $p < 0.01$ ); for laterites the correlation is not significant (Fig. 4*c*). Spectra for  
290 laterite and surface samples exhibit positive correlations between band depths of the  
291 2.32  $\mu\text{m}$  Mg-OH (serpentine) absorption and MgO abundance (positive coefficient  
292 significant at  $p < 0.01$ ; see Table 4 for additional information on regression statistics)  
293 (Fig. 4*d*).



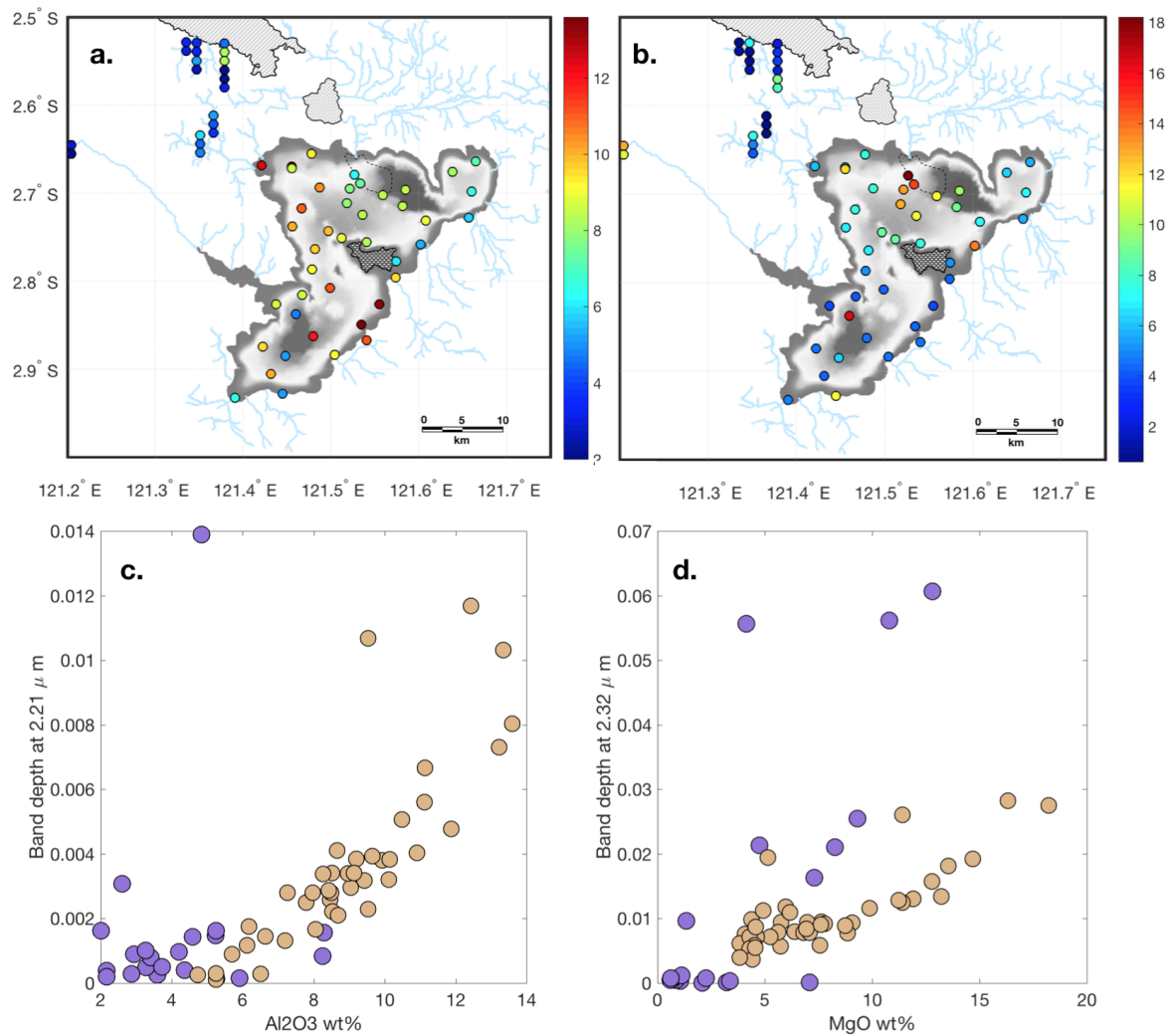


Figure 4: *a.*  $\text{Al}_2\text{O}_3$  (wt%) in surface and soil samples. *b.*  $\text{MgO}$  (wt%) in surface and soil samples. *c.* Band depth at  $2.21\ \mu\text{m}$  versus  $\text{Al}_2\text{O}_3$  abundance in surface (brown) and soil (purple) samples. *d.* Band depth at  $2.32\ \mu\text{m}$  versus  $\text{MgO}$  abundance in surface (brown) and soil (purple) samples.

Table 3: Chemistry results (wt%) from ICP-AES. (- indicates not measured.)

Sample No.	Type	Al <sub>2</sub> O <sub>3</sub>	FeO	MgO	SiO <sub>2</sub>	TiO <sub>2</sub>
133	River	-	-	-	-	-
134	River	-	-	-	-	-
135	River	-	-	-	-	-
136	River	-	-	-	-	-
137	River	-	-	-	-	-
138	River	-	-	-	-	-
139	River	-	-	-	-	-
140	River	-	-	-	-	-
141	River	-	-	-	-	-
142	River	-	-	-	-	-
143	River	-	-	-	-	-
144	River	-	-	-	-	-
145	River	-	-	-	-	-
146	River	-	-	-	-	-
147	River	-	-	-	-	-
148	River	-	-	-	-	-
226	Laterite (bed 1)	3.600	38.400	7.090	-	0.061
227	Laterite (bed 1)	3.280	38.100	1.050	-	0.047
228	Laterite (bed 1)	5.240	46.400	0.860	-	0.095
229	Laterite (bed 1)	2.600	24.300	1.340	-	0.029
230	Laterite (bed 2)	3.410	47.300	0.740	-	0.137
506	Laterite sample 230 after Extraction 1	-	-	-	-	-
518	Laterite sample 230 after Extraction 2	-	-	-	-	-
231	Laterite (bed 2)	2.930	48.600	1.130	-	0.108
507	Laterite sample 231 after Extraction 1	-	-	-	-	-
519	Laterite sample 231 after Extraction 2	-	-	-	-	-
232	Laterite (bed 3)	2.870	12.300	12.800	-	0.106
233	Laterite (bed 3)	2.170	14.900	10.800	-	0.047
234	Laterite (bed 4)	4.360	33.900	4.750	-	0.148
235	Laterite (bed 4)	4.840	18.300	4.130	-	0.170
508	Laterite sample 235 after Extraction 1	-	-	-	-	-
520	Laterite sample 235 after Extraction 2	-	-	-	-	-
236	Laterite (bed 4)	5.910	14.900	7.300	-	0.220
237	Laterite (bed 5)	5.250	49.100	0.610	-	0.111
238	Laterite (bed 5)	4.200	49.400	0.610	-	0.062
509	Laterite sample 238 after Extraction 1	-	-	-	-	-
521	Laterite sample 238 after Extraction 2	-	-	-	-	-
239	Laterite (bed 5)	3.730	49.100	0.620	-	0.059
240	Laterite (bed 6)	4.580	34.900	2.080	-	0.067
241	Laterite (bed 6)	8.250	33.800	3.220	-	0.068
242	Laterite (bed 6)	8.290	32.900	3.380	-	0.077
243	Laterite (bed 6)	2.170	40.900	2.250	-	0.043
244	Laterite (bed 6)	3.260	21.800	8.260	-	0.082
245	Laterite (bed 6)	2.000	14.300	9.310	-	0.050
247	Surface	12.429	12.458	5.976	49.106	0.882
248	Surface	13.229	15.138	5.731	42.950	0.922
249	Surface	9.423	20.636	7.641	39.266	0.480
250	Surface	8.666	13.083	11.918	38.961	0.431
251	Surface	6.180	11.687	18.219	41.885	0.255
252	Surface	7.782	14.102	13.216	39.235	0.334
253	Surface	7.257	12.184	14.688	39.143	0.307
254	Surface	8.470	17.191	9.882	37.126	0.323
255	Surface	10.483	15.998	7.774	38.222	0.515
256	Surface	8.483	15.360	11.400	37.006	0.333
510	Surface sample 256 after Extraction 1	-	-	-	-	-
522	Surface sample 256 after Extraction 2	-	-	-	-	-
257	Surface	8.058	23.128	6.378	30.811	0.259
258	Surface	6.635	23.003	6.967	29.796	0.207
259	Surface	5.698	25.302	5.749	28.869	0.126
511	Surface sample 259 after Extraction 1	-	-	-	-	-
523	Surface sample 259 after Extraction 2	-	-	-	-	-
260	Surface	8.510	16.311	9.066	35.024	0.353
261	Surface	7.969	13.427	12.775	38.703	0.324
262	Surface	10.906	16.371	7.558	39.555	0.494
263	Surface	8.419	15.271	11.242	38.404	0.330
264	Surface	10.121	17.291	6.794	37.649	0.439
265	Surface	7.198	23.427	5.618	31.635	0.226
266	Surface	9.931	15.659	8.863	39.460	0.427
267	Surface	9.044	19.058	7.076	35.082	0.373
268	Surface	5.280	21.700	13.554	33.676	0.168
269	Surface	9.541	12.580	4.374	52.465	0.576
270	Surface	6.118	27.767	5.134	35.524	0.168
512	Surface sample 270 after Extraction 1	-	-	-	-	-
524	Surface sample 270 after Extraction 2	-	-	-	-	-
271	Surface	8.262	16.582	7.618	33.699	0.335
272	Surface	8.969	15.428	8.748	36.008	0.365
273	Surface	9.656	16.290	6.933	35.348	0.387
274	Surface	9.209	15.486	5.249	32.634	0.354
275	Surface	8.523	22.386	4.074	32.930	0.294
276	Surface	8.688	23.757	3.835	30.356	0.277
277	Surface	9.532	23.344	4.625	31.792	0.293
513	Surface sample 277 after Extraction 1	-	-	-	-	-
525	Surface sample 277 after Extraction 2	-	-	-	-	-
278	Surface	6.501	29.400	4.922	23.804	0.093
279	Surface	10.145	18.993	4.293	34.179	0.390
280	Surface	5.246	30.272	6.168	27.696	0.074
281	Surface	11.879	17.809	4.537	37.579	0.456
282	Surface	11.129	13.502	4.300	40.793	0.498
283	Surface	9.133	17.855	4.566	33.168	0.364
514	Surface sample 283 after Extraction 1	-	-	-	-	-
526	Surface sample 283 after Extraction 2	-	-	-	-	-

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Sample No.	Type	Al <sub>2</sub> O <sub>3</sub>	FeO	MgO	SiO <sub>2</sub>	TiO <sub>2</sub>
284	Surface	13.594	14.020	4.422	42.879	0.589
285	Surface	13.339	12.371	3.821	46.219	0.669
286	Surface	4.726	20.371	16.331	33.981	0.038
515	Surface sample 286 after Extraction 1	-	-	-	-	-
527	Surface sample 286 after Extraction 2	-	-	-	-	-
287	Surface	5.250	27.322	11.414	33.082	0.052
516	Surface sample 287 after Extraction 1	-	-	-	-	-
528	Surface sample 287 after Extraction 2	-	-	-	-	-
288	Surface	11.125	20.991	4.545	31.899	0.321
97	Core	-	-	-	-	-
489	Core sample 97 after Extraction 1	-	4.989	-	-	-
497	Core sample 97 after Extraction 2	-	2.177	-	-	-
99	Core	-	-	-	-	-
490	Core sample 99 after Extraction 1	-	6.930	-	-	-
498	Core sample 99 after Extraction 2	-	4.353	-	-	-
104	Core	-	-	-	-	-
491	Core sample 104 after Extraction 1	-	5.383	-	-	-
499	Core sample 104 after Extraction 2	-	4.789	-	-	-
113	Core	-	-	-	-	-
492	Core sample 113 after Extraction 1	-	6.550	-	-	-
500	Core sample 113 after Extraction 2	-	2.754	-	-	-
114	Core	-	-	-	-	-
493	Core sample 114 after Extraction 1	-	5.655	-	-	-
501	Core sample 114 after Extraction 2	-	2.962	-	-	-
116	Core	-	-	-	-	-
494	Core sample 116 after Extraction 1	-	6.461	-	-	-
502	Core sample 116 after Extraction 2	-	4.212	-	-	-
118	Core	-	-	-	-	-
495	Core sample 118 after Extraction 1	-	3.947	-	-	-
503	Core sample 118 after Extraction 2	-	2.778	-	-	-
121	Core	-	-	-	-	-
496	Core sample 121 after Extraction 1	-	8.043	-	-	-
504	Core sample 121 after Extraction 2	-	7.093	-	-	-
122	Core	-	-	-	-	-
505	Core sample 122 after Extraction 1	-	13.852	-	-	-
517	Core sample 122 after Extraction 2	-	2.672	-	-	-

Table 4: Al and Mg Regression Results

	<i>Dependent variable:</i>			
	BD 2.21 $\mu\text{m}$ , surface (1)	BD 2.21 $\mu\text{m}$ , laterite (2)	BD 2.32 $\mu\text{m}$ , surface (3)	BD 2.32 $\mu\text{m}$ , laterite (4)
Al <sub>2</sub> O <sub>3</sub> , surface	0.001*** (0.0001)			
Al <sub>2</sub> O <sub>3</sub> , laterite		0.0002 (0.0004)		
MgO, surface			0.001*** (0.0002)	
MgO, laterite				0.004*** (0.001)
Constant	-0.005*** (0.001)	0.001 (0.002)	0.001 (0.001)	-0.003 (0.005)
Observations	42	20	42	20
R <sup>2</sup>	0.700	0.012	0.644	0.580
Residual Std. Error	0.002 (df = 40)	0.003 (df = 18)	0.004 (df = 40)	0.014 (df = 18)
F Statistic	93.113*** (df = 1; 40)	0.210 (df = 1; 18)	72.365*** (df = 1; 40)	24.901*** (df = 1; 18)

*Note:*

\*p&lt;0.1; \*\*p&lt;0.05; \*\*\*p&lt;0.01

294 Fe abundances are reported as %FeO (total Fe, not indicating redox state) and  
295 are very high in the lateritic catchment soils (Fig. 5*a.*), with concentrations of  
296 up to 49.4 wt%, and Fe abundances are more spatially variable between the lake  
297 sediment samples (Fig. 5*b.*). VNIR spectra indicate the presence of Fe-oxides and  
298 Fe-bearing silicates in both soils and lake sediment, where the latter may include  
299 nontronite in two lake samples and several laterite samples. Because of the complex  
300 and overlapping nature of spectral absorptions of Fe at VNIR wavelengths, (Fig. 6),  
301 two different spectral parameters are compared with ICP-determined Fe abundances  
302 (Fig. 5). The inverse slope of the 0.45  $\mu\text{m}$  to 0.48  $\mu\text{m}$  absorption band, which is  
303 observed in some Fe-oxides such as hematite and goethite and is similar to a measure  
304 of band area, is correlated with Fe abundance for both laterites and surface samples  
305 (positive coefficient significant at  $p < 0.01$ ) (Fig. 5*c.*). The band depth of the  $\sim 0.9$ -  
306 1.0  $\mu\text{m}$  absorption that is commonly observed in Fe-oxides and Fe-bearing silicates  
307 (e.g. nontronite) is also correlated with Fe abundance (positive coefficient significant  
308 at  $p < 0.05$  for surface samples, positive coefficient significant at  $p < 0.01$  for laterites)  
309 (Fig. 5*d.*). Although both spectral parameters exhibit a correlation with ICP-  
310 derived Fe content, the correlation is strongest for the spectral slope parameter in  
311 the lake surface samples (brown points in Fig. 5*c.*). The correlation between this Fe  
312 spectral slope parameter is much stronger than previously reported correlations of  
313 Fe abundance with spectral properties (Weber et al. 2015). Additional information  
314 on regression statistics is presented in Table 5.

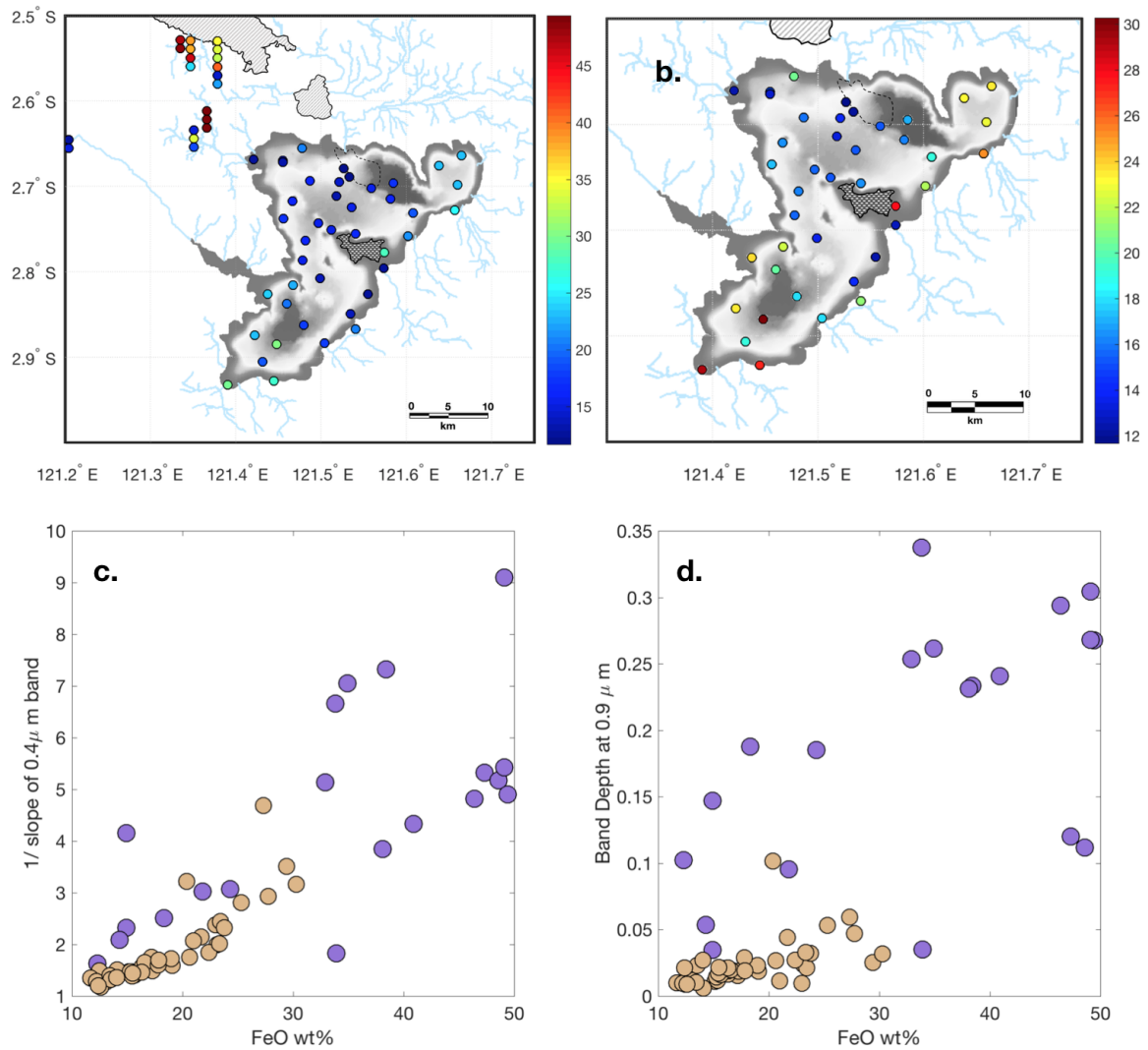


Figure 5: *a.* ICP-determined elemental abundance of FeO (wt%) in surface and soil samples. *b.* ICP-determined elemental abundance of FeO (wt%) in just surface samples. *c.* Flatness of the 0.4 μm band versus with FeO abundance in surface sediment (brown) and soil (purple) samples. *d.* Band depth at 0.9 μm band versus with FeO abundance in surface (brown) and soil (purple) samples.

Table 5: Fe Regression Results

	<i>Dependent variable:</i>			
	1/slope 0.4 $\mu\text{m}$ band, surface	1/slope 0.4 $\mu\text{m}$ band, laterite	BD 0.9 $\mu\text{m}$ , surface	BD 0.9 $\mu\text{m}$ , laterite
	(1)	(2)	(3)	(4)
FeO, surface	0.126*** (0.012)		0.005** (0.002)	
FeO, laterite		0.102*** (0.026)		0.008*** (0.002)
Constant	-0.437* (0.229)	1.107 (0.940)	0.118*** (0.034)	0.349*** (0.066)
Observations	42	20	42	20
R <sup>2</sup>	0.733	0.453	0.142	0.501
Residual Std. Error	0.384 (df = 40)	1.536 (df = 18)	0.058 (df = 40)	0.108 (df = 18)
F Statistic	109.810*** (df = 1; 40)	14.889*** (df = 1; 18)	6.626** (df = 1; 40)	18.067*** (df = 1; 18)

*Note:*

\*p&lt;0.1; \*\*p&lt;0.05; \*\*\*p&lt;0.01

315 The lateritic soil samples are characterized by high FeO abundances (25-50 wt%)  
 316 that are generally highest at the top of the soil bed and decrease with depth (Morlock  
 317 et al. 2018). One laterite, located west of the others near the Larona River, the  
 318 main outlet of Lake Towuti, has significantly lower FeO ( $\sim 14$  wt%; Fig. 8) and  
 319 elevated MgO. All soil samples examined here have moderate-to-low  $\text{Al}_2\text{O}_3$  values  
 320 (2-8 wt%) that, like FeO, tend to be highest in the samples closest to the surface.  
 321 VNIR spectra from these high-Fe samples are consistent with abundant hematite  
 322 and goethite, with much lower concentrations of other Fe-bearing phases such as  
 323 Fe-clay minerals or siderite.

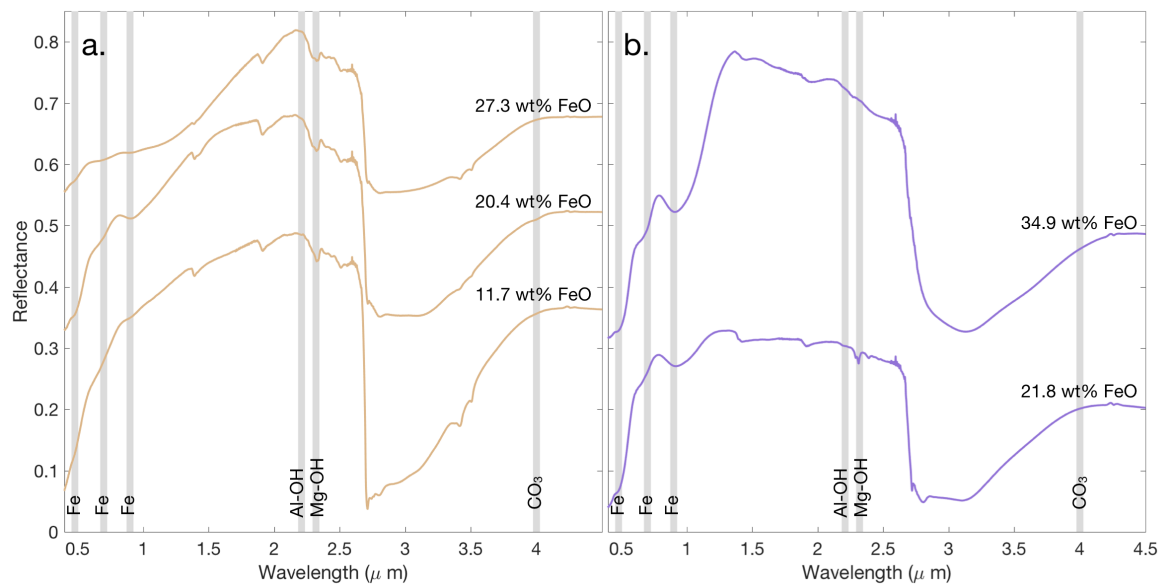


Figure 6: Example spectra, with relevant band locations highlighted, labeled with their ICP-determined FeO abundances (wt%). *a.* Three example surface sediment spectra, offset for clarity. The top two are from the southern end of the lake and the bottom spectrum is from the coarse-grained, serpentine-rich Mahalona river delta. *b.* Two example laterite spectra, offset for clarity. Top spectrum is from the top of one laterite bed (10 cm from the surface) and bottom spectrum is from the bottom of the same bed (3.5 m from surface). Note the increase in Fe absorption band strengths (0.4-1.5  $\mu\text{m}$ ) and decrease in Mg-serpentine band strength ( $\sim 2.32$   $\mu\text{m}$ ) in the spectrum for the upper soil sample compared with that of the lower soil sample, consistent with increased abundance of Fe oxides at the expense of serpentine during progressive weathering of the serpentinized bedrock.

324 Nontronite was identified in two shallow, coastal lake surface samples (Samples  
 325 259, 270) and potentially six laterite samples based on a diagnostic (Fe-OH) absorp-  
 326 tion feature at 2.28  $\mu\text{m}$  in their reflectance spectra (Fig. 7). In addition, spectra



327 for a number of the laterite samples exhibit two distinct absorptions near  $\sim 2.28$   
328  $\mu\text{m}$  and  $\sim 2.31 \mu\text{m}$  (purple spectrum, Fig. 7*a.*). The strength of these two features  
329 varies independently within the sample group as a whole, suggesting this apparent  
330 ‘doublet’ absorption is the result of two distinct mineral phases that vary in their  
331 relative abundance between samples (i.e., the doublet is not the result of a single  
332 mineral phase). The position of the  $2.28 \mu\text{m}$  feature is similar to that of Fe-OH in  
333 nontronite, and the position of the  $2.31 \mu\text{m}$  feature is similar to that of Mg-OH in  
334 saponite; neither is consistent with absorptions typical of serpentine. However, the  
335 absorptions in these laterite samples are narrower than what is commonly observed  
336 in reflectance spectra of pure nontronite or saponite, thus the origin of these spectral  
337 features is somewhat ambiguous. It is possible that the observed doublet is a result  
338 of a mixture of nontronite and saponite, perhaps poorly crystalline forms of these  
339 clay minerals.

## 340 **4.2 Iron mineralogy**

### 341 *XRD patterns*

342 Consistent with the ICP data (Fig. 5) and VNIR spectra (Figs. 6, 7), XRD  
343 patterns indicate that crystalline Fe-oxides are more abundant in laterite samples  
344 than in lake surface sediment samples (Fig. 8). These analyses are qualitative and  
345 based on relative peak intensity. Detailed examination of the laterite XRD patterns  
346 (Fig. 9*a-c.*) show sharp, relatively high goethite peaks in almost every laterite  
347 sample, sharp magnetite peaks in several samples, and hematite in few samples.  
348 In the surface sediment samples (Fig. 9*d-f.*), the peaks are broadened and shorter.  
349 There is still evidence of Fe oxides (goethite, magnetite, hematite) in the lake surface  
350 samples, but, similar to ICP results, they are less abundant compared to other iron  
351 phases such as serpentine, siderite, Cr spinel, olivine, amphibole, and pyroxene.  
352 There is also a notable paucity of well-crystalline Fe-oxides in lake sediment samples  
353 from the coarse-grained Mahalona delta deposits as well as samples located close  
354 to other river inlets (Fig. 9*g.*). Note that truly amorphous Fe will not appear as

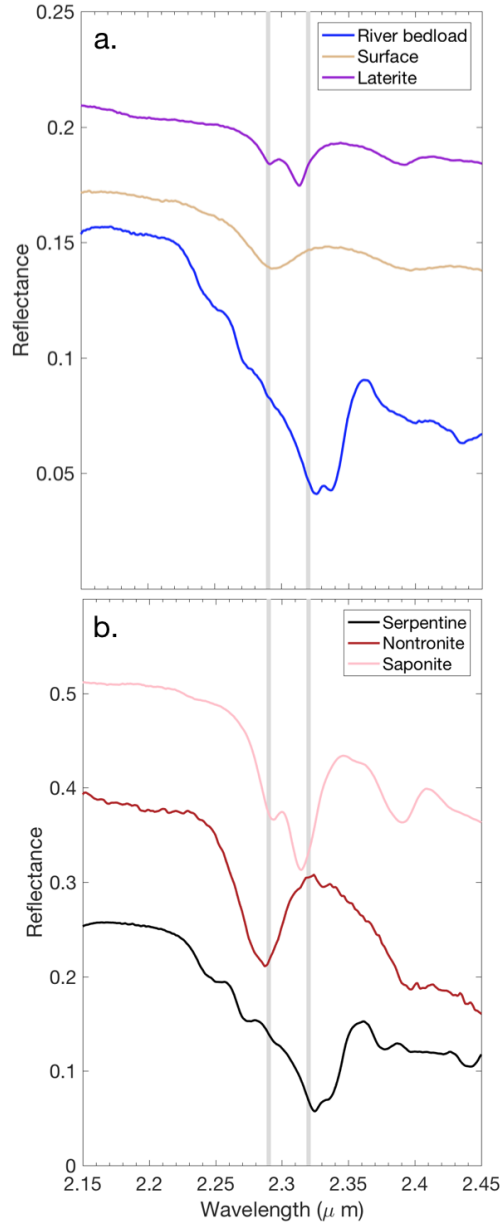


Figure 7: *a.* Four example reflectance spectra (offset for clarity): a laterite sample, surface sample with 2.28  $\mu\text{m}$  absorptions attributed to nontronite, and Mahalona River bedload sample with no nontronite but a strong 2.32  $\mu\text{m}$  feature attributed to serpentine. *b.* Laboratory spectra (offset for clarity) of serpentine (lizardite, NMNHR 4687), nontronite (NG-1), and saponite (SapCa-1).

355 an amorphous “hump” as amorphous silica does, but rather raise the baseline of  
 356 the XRD pattern, making it difficult to understand qualitatively or quantitatively  
 357 through XRD alone.

358 *Sequential Fe extraction*

359 Sequential Fe extractions affecting the mineral hosts of iron were performed on  
 360 laterites, surface samples, and samples from sediment cores (Fig. 10). The laterites,

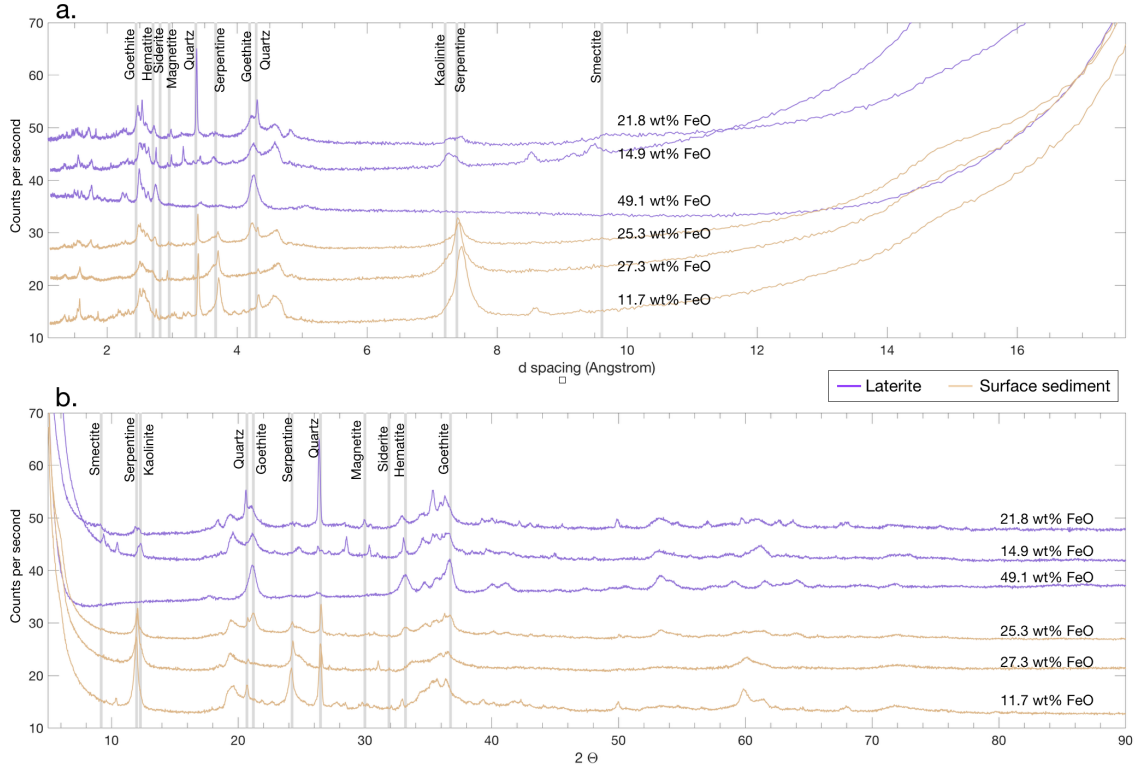


Figure 8: XRD patterns (offset for clarity) of three laterites (purple) and three surface samples (tan) presented in terms of *a.* d-spacing and *b.*  $2\theta$ . Patterns are offset for clarity and labeled with the samples' ICP-determined FeO abundance (wt%). Vertical gray bars highlight the peak location of relevant phases.

361 which are a dominant sediment source in this system, all exhibit significantly higher  
 362 total Fe concentrations and are dominated by crystalline Fe-oxides (e.g., goethite,  
 363 hematite) relative to lake sediment. Total Fe abundances are similar in the lake  
 364 surface and core samples, but the manner in which this Fe is partitioned between  
 365 mineral phases is quite different between the two sample groups. The proportion of  
 366 Fe in amorphous or highly reactive phases (i.e. the extraction 1 supernatant) versus  
 367 crystalline phases (i.e. extraction 2 supernatant) is roughly equivalent in surface  
 368 sediment samples. In contrast, samples from deeper within the sediment column  
 369 (core samples) contain more iron in what appear to be X-ray amorphous phases.

370 We note that not all of the iron is necessarily extracted by these dissolution  
 371 steps, and the summed supernatant Fe abundance of some samples is up to 9 wt%  
 372 Fe below that of the original surface sediment. This “excess iron” can be partially  
 373 attributed to the presence of Fe in silicates and other mineral phases that our se-

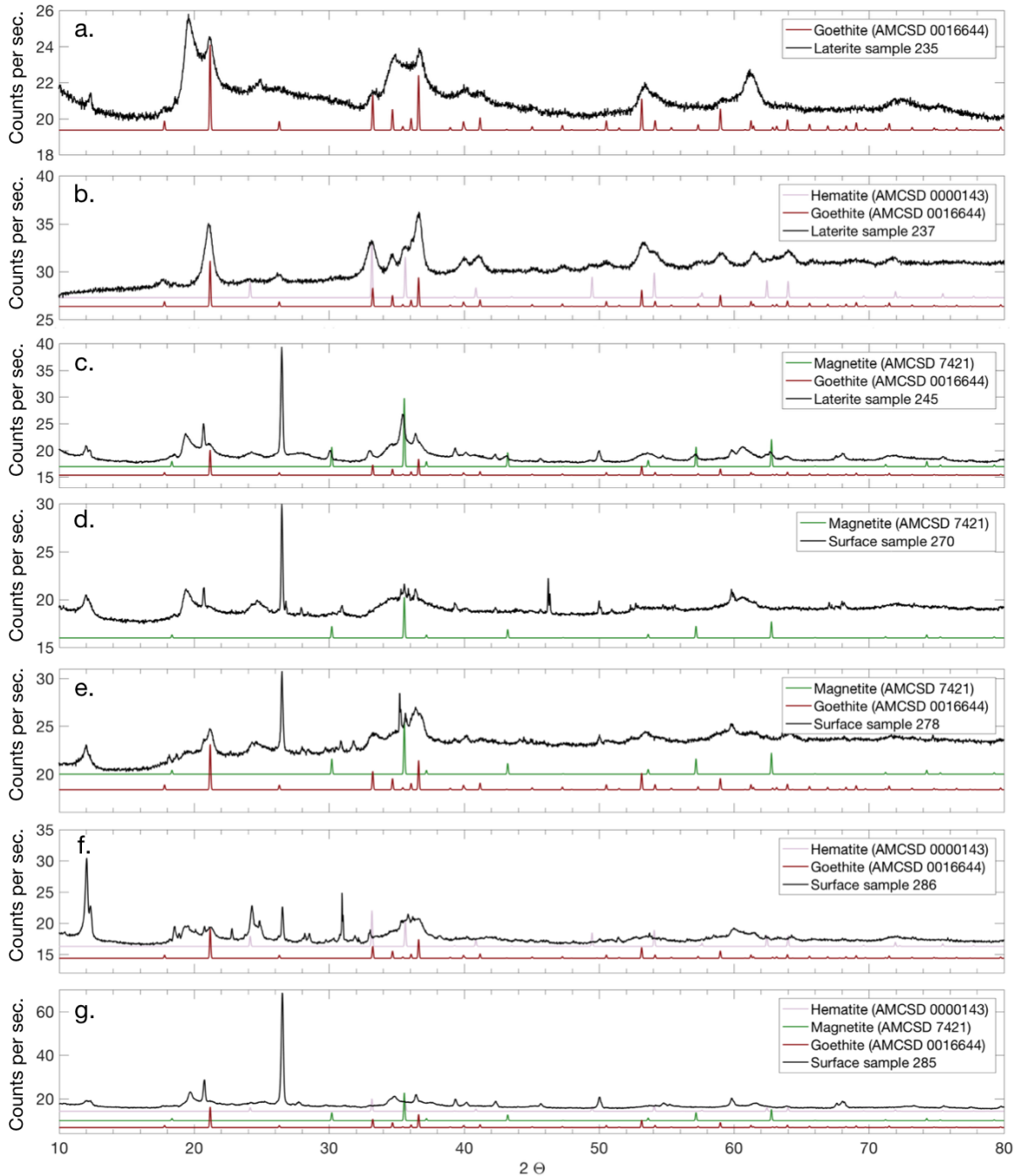


Figure 9: XRD patterns (offset for clarity) from (a-c.) laterites and (d-g.) surface sediment. a. A laterite with abundant goethite. b. A laterite with abundant goethite and hematite. c. A laterite with magnetite and a small amount of goethite. d. A surface sample with magnetite. e. A surface sample with magnetite and goethite. f. A surface sample with goethite and hematite. g. A surface sample with no discernible Fe oxides present. Note that the common sharp peak at 26.4  $2\theta$  is quartz.

374 sequential reaction protocol does not, and is not intended to, digest. However, XRD  
 375 analyses of sediments prior to and after the extractions suggest that some reactive  
 376 iron mineral phases survive the extraction process. For instance, the intensity of

377 the goethite peak is reduced but a portion appears to be resistant to the extractions  
 378 (Fig. 11). The intensity of the goethite peak is reduced but some goethite remains.  
 379 Processing the sample with extraction 2 a second time removed additional goethite,  
 380 resulting in near-complete removal of goethite peaks in the XRD patterns.  
 381

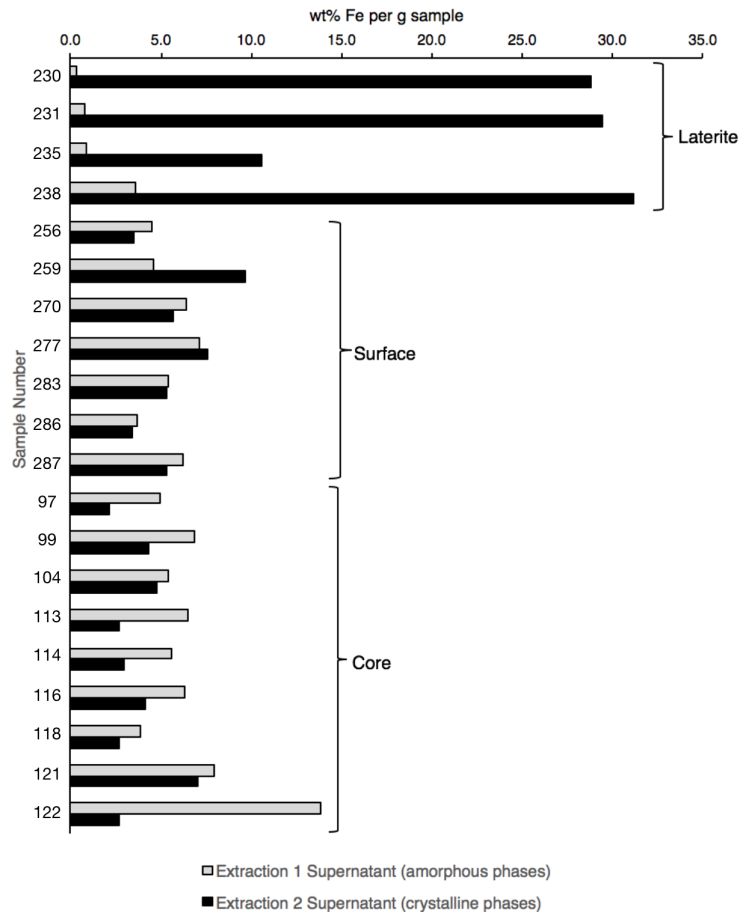


Figure 10: Comparison of the wt% Fe per g sample in each supernatant produced via extraction 1 (gray, amorphous phases) and extraction 2 (black, crystalline phases). Laterite samples are from depths of 10 cm to 2 m from the surface; core samples are from 2-77 m sediment depth, with one tephra (sample 20); surface samples represent sediment depths of <5 cm and are from varying locations, including one from the Mahalona delta (sample 5). Differences in the level of noise between patterns is due to differences in integration times. Values of  $2\theta$  were adjusted slightly to account for minor sample displacement, normalizing such that the primary quartz peak is centered at 26.6 degrees  $2\theta$ . Sample numbers are that of the original sediment; see Table 1 for more information.

382 *Mössbauer spectra*

383 Overall, Mössbauer data are in agreement with the observations of reflectance  
 384 spectra, XRD patterns, and sequential extractions and further indicate that the Fe

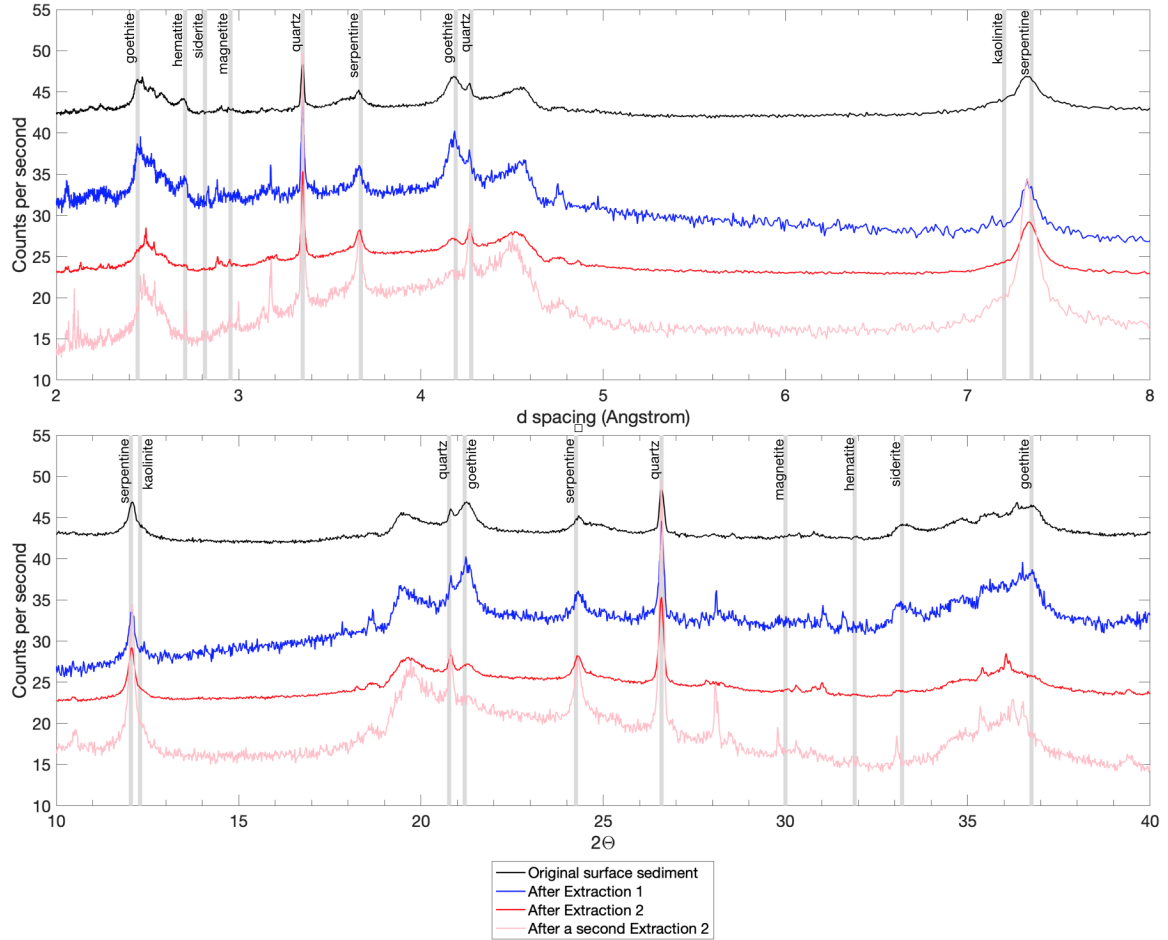


Figure 11: Zoomed-in XRD patterns (offset for clarity) of one surface sample (sample 259) throughout the sequential extraction process in terms of *a.* d-spacing and *b.*  $2\theta$ . Patterns are offset for clarity and diagnostic peaks of relevant phases are highlighted in gray.

385 mineralogy is extremely complex within the surface sediment. The samples contain  
 386 a large nanocrystalline component, both ferrous and ferric phases, and exhibit in-  
 387 terparticle effects causing broadening of the doublets and sextets during magnetic  
 388 ordering (Fig. 12). Fits to the Mössbauer patterns suggest a significant nanocrys-  
 389 talline component with both  $\text{Fe}^{2+}$  and  $\text{Fe}^{3+}$ . Spectra acquired at the lowest temper-  
 390 ature (4 K, Fig. 12) in particular suggest highly disordered, fine-grained material.  
 391 Only the finest-grained nanophase iron (oxyhydr)oxides would still be unsplit at  
 392 4 K, so at least 40% spectral area seems to be attributable to this form of iron  
 393 (oxyhydr)oxide.

394 Surface sediment sample 270 (Fig. 12*a-c.*) is from the northeastern portion of  
 395 the lake. At 295 K, the Mössbauer fits suggest two unremarkable pairs of ferrous

396 and ferric doublets. One ferrous doublet has a very low IS value and moderate QS  
397 value, likely attributable to a clay. There is no evidence for ilmenite, which should  
398 be readily apparent at room temperature. At 130 K, the ferrous phase QS increases  
399 only slightly, consistent with the presence of an Fe-bearing clay mineral. A broad  
400 sextet appears at this temperature (Fig. 12*b.*, orange curve) that is consistent with  
401 magnetite. At 4 K, the ferrous Fe undergoes strong magnetic orientation because of  
402 the higher symmetry of the electrons around ferrous Fe. The spectral fits are very  
403 complex at this temperature and suggest highly disordered, nanocrystalline phases  
404 with both ferrous and ferric components.

405 Surface sediment sample 286 (Fig. 12*d-f.*) is from the southern portion of the  
406 lake. At 295 K, the parameters are closest to those of ferrous clays or possibly  
407 pyroxene (and are inconsistent with olivine, feldspar, sulfate, or phosphate). There  
408 is evidence for two ferrous doublets and two ferric doublets at this temperature. At  
409 130 K, the ferric phases move to broad, disordered sextets, consistent with nanophase  
410 hematite, goethite, or possibly akaganeite. At 4 K, these sextets are even more  
411 prominent and the fits are indicative of significant disorder in the samples. Trace  
412 siderite could explain the broad, unsplit blue curve (Fig. 12*f.*) that disappears at  
413 higher temperatures. However, siderite is not apparent in these samples based on  
414 VNIR spectra or XRD patterns. If it is present, it is likely  $\leq 5\%$  in abundance, below  
415 the approximate detection limit for those methods but possibly still detectable in  
416 the Mössbauer data.

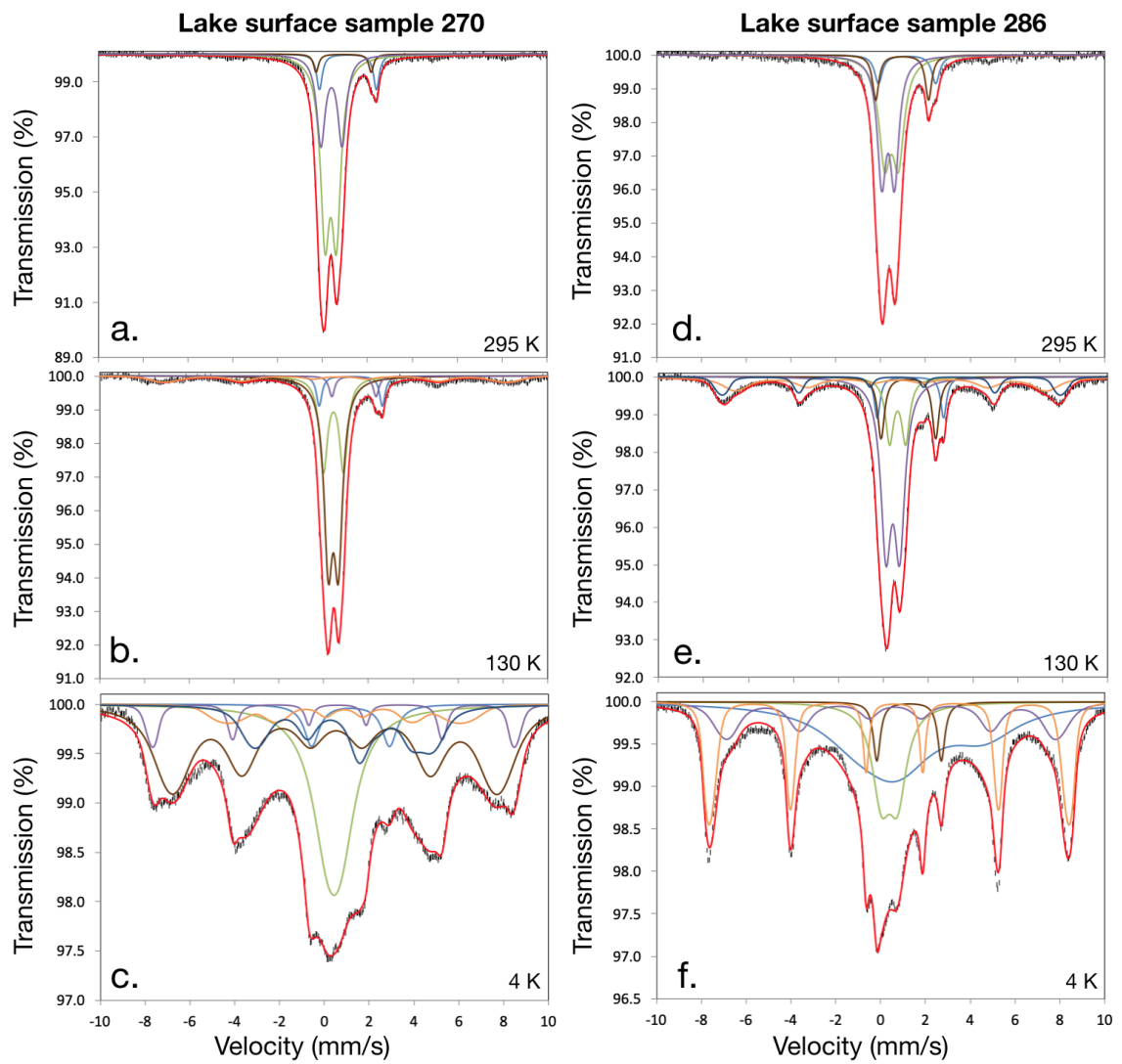


Figure 12: Mössbauer spectra (black line) and fits (colored lines) for surface sample 270 at *a.* 295 K, *b.* 130 K, and *c.* 4 K. Mössbauer spectra (black line) and fits (colored lines) for surface sample 286 at *d.* 295 K, *e.* 130 K, and *e.* 4 K.



## 417 5 Discussion

418 The dominant chemical variations observed in Lake Towuti laterites and surface  
419 sediments are in Mg, Al, and Fe content. In the lake, Mg and Al abundances are  
420 strongly linked to detrital mineralogy, particularly the relative abundance of detrital  
421 Mg-serpentine versus kaolinite from the catchment (Fig. 4, Table 4). The coarser-  
422 grained surface samples acquired closer to the Mahalona River are mainly composed  
423 of serpentine, which is inherently coarser-grained in this system because the river  
424 is transporting fragments of serpentinized peridotite bedrock that have undergone  
425 limited chemical alteration from the sediment source region. The abundance of Mg  
426 is thus a proxy for serpentine content which is in turn a proxy for grain size because  
427 serpentine is the dominant phase in the coarser-grained sediment. This can be seen  
428 in the spatial distribution of high-Mg lake surface samples that conforms to the  
429 boundaries of the Mahalona delta (Fig. 4*b*). In contrast, kaolinite (and thus Al)  
430 abundance increases in the finer-grained detrital component because it is forming  
431 in the soil weathering horizons. To the extent that these relationships hold true  
432 through geologic time, lateral and vertical variations in the relative abundance of  
433 serpentine and kaolinite may thus record important information about sediment flux,  
434 water depth/distance to shoreline, and surface weathering conditions through time.  
435 The ability of VNIR reflectance spectroscopy to accurately identify and distinguish  
436 kaolinite from serpentine, both of which are 7Å clay minerals, highlights the useful-  
437 ness of this technique for rapid, non-destructive analysis of sediments and sediment  
438 cores to help infer paleoenvironmental conditions in the Malili Lakes system.

439 Variations in the Fe mineralogy from source to sink reflect complex and dynamic  
440 cycling of Fe between a range of mineral hosts that vary in crystallinity and grain  
441 size. Elemental Fe abundances are very high in the catchment soils, which represent  
442 a major sediment source region, and are consistent with intense weathering in trop-  
443 ical laterites. XRD and spectral analyses of the catchment soils show that they are  
444 characterized by a range of crystalline Fe oxides, with goethite being the dominant  
445 phase accompanied by lesser amounts of hematite, likely derived from weathering of

446 primary Fe-bearing minerals (e.g., olivine and pyroxene). Magnetite is also present  
447 in several samples and appears to be more abundant at greater soil depths, consis-  
448 tent with at least some of the magnetite originating from unweathered bedrock, as  
449 suggested in Morlock et al. (2018). Magnetite is also removed and/or segregated  
450 during weathering, consistent with its almost complete absence in river samples and  
451 deep-water surface samples. Indeed, XRD analyses show that hematite and mag-  
452 netite are only present in a few of the shallow coastal samples and these phases are  
453 largely absent from deep water samples and from the Mahalona River delta. This  
454 is consistent with several processes, including early hydrodynamic settling of mag-  
455 netite in shallow water due to its higher density (Tamuntuan et al. 2015), reductive  
456 dissolution of hematite and magnetite in anoxic water, and variations in sediment  
457 source regions, particularly lithologies low in iron in the Mahalona catchment.

458 The broadening and decrease in peak height seen in XRD patterns of the lake  
459 surface samples (Figs. 8, 9) can be explained by a combination of reduction of Fe  
460 oxide abundance, decreased grain size, and disorder in the crystal lattice (Sklute  
461 et al. 2018). The majority of VNIR spectra for the lake surface samples do exhibit  
462 absorptions consistent with the presence of  $\text{Fe}^{2+}/3+$ , in agreement with the measured  
463 high  $\text{Fe}_{total}$  concentrations, but the shapes and positions of these features are not  
464 consistent with the presence of only well-crystalline goethite and/or hematite. The  
465 dominance of poorly crystalline Fe oxides in the lake surface sediments is also con-  
466 firmed by Mössbauer measurements, which are consistent with a significant fraction  
467 of the Fe in the lake sediments being in the form of nanocrystalline Fe oxides.

468 Within the lake itself, dissolved oxygen declines below 100 m depth and reaches  
469 anoxic conditions around 130 m depth, consistent with evidence that the water-  
470 sediment interface is oxic at shallow sites and anoxic and enriched in ferrous Fe at  
471 intermediate and deep sites (Vuillemin et al. 2016). Despite the decline (and in  
472 some cases absence) of XRD peaks associated with crystalline Fe-oxides and the  
473 dominance of such phases in the laterite source regions, ICP results show that many  
474 lake surface samples contain high concentrations of  $\text{Fe}_{total}$ . The total Fe content

475 in surface sediment and down-core samples are similar, but surface samples are  
476 roughly equal in their distribution of crystalline and amorphous Fe (based on Fe in  
477 the supernatant after each extraction step), whereas all core samples have more Fe in  
478 amorphous phases and a lesser amount in well-crystalline Fe-oxides such as goethite.  
479 Interestingly, goethite is present across the lake at the sediment/water interface,  
480 although greatly reduced in abundance compared with the catchment soils. The  
481 loss of goethite is likely due to its reduction in the anoxic water column and under  
482 anoxic conditions at shallow depths beneath the sediment-water interface. However,  
483 some goethite persists in the lake samples from the sediment/water interface, even  
484 those obtained from deep parts of the lake below anoxic waters. This, together with  
485 the persistence of goethite after sequential Fe extractions, suggest that some portion  
486 of the goethite may be more resistant to alteration/diagenesis. The cause of this is  
487 unknown but may result from goethite acting as a cementing agent, thus forming  
488 slightly larger particles that may be more resistant to dissolution. Although some  
489 goethite remains in the surface sediment of Lake Towuti, it is likely to be removed  
490 over time during burial and early diagenesis. Regardless of the persistence of some  
491 goethite, a key observation is that a significant fraction of Fe in the lake surface  
492 sediment cannot be accounted for by the crystalline Fe-oxides that are observed in  
493 XRD patterns, despite the fact that the majority of Fe likely enters the lake in this  
494 form. This suggests that much of the Fe within the lake may be in the form of  
495 poorly crystalline, X-ray amorphous phases that originate from the transformation  
496 of soil-derived Fe oxides by iron cycling and reduction within the lake.

497 The changes in Fe mineralogy between sediment source (laterite) and sink (sed-  
498 imentation in the lake) must account for the increase in amorphous Fe material.  
499 If laterites accurately represent a primary sediment input into the lake (aside from  
500 the Mahalona delta region), some process or suite of processes acts to liberate iron  
501 from the ferric crystalline phases in the water column or shortly after deposition  
502 and promotes the formation of new nanocrystalline Fe-oxide phases. Interestingly,  
503 and despite Lake Towuti being redox-stratified, there is no clear correlation between

504 the relative abundance of crystalline Fe oxides in the surface sediments and water  
505 depth: most surface sediments that were sampled at water depths above the modern  
506 oxycline also exhibit less goethite and hematite than observed in soils even though  
507 total Fe content remains high. This, together with the presence of goethite in sedi-  
508 ments below the oxycline indicates there is less of a gradient in Fe mineralogy across  
509 the modern oxycline than might be expected if water column reduction reactions  
510 control the Fe mineralogy in the lake. Possible explanations for the apparent rel-  
511 ative “homogeneity” in Fe oxide mineralogy at different water depths include: (i)  
512 conditions, perhaps biologically mediated, in shallow water settings allow for some  
513 reductive dissolution of Fe, presumably along redox gradients at shallow depth in  
514 the sediment, or (ii) reduction of Fe oxides in deep water followed by frequent water  
515 column mixing to precipitate more oxidized species across the lake.

516 Fe cycling has been studied in the water column of Lake Matano, another of the  
517 Malili lakes that is non-sulfidic, permanently redox-stratified, and more than twice as  
518 deep as Lake Towuti (Crowe et al. 2011). Studies of its water column chemistry have  
519 shown that the surface mixed layer of Lake Matano is in a pH/pE range in which  
520 iron (hydr)oxides are stable, whereas the hypolimnion favors  $\text{Fe}(\text{OH})^{2+}$  and  $\text{Fe}^{2+}$   
521 (Crowe et al. 2008). It is possible that similar cycling promotes the transformation  
522 of the crystalline Fe oxide form to more nanophase components. Ferric oxides may  
523 settle below the oxycline where they are then rapidly reduced and/or altered to  
524 amorphous phases. Though we cannot exclude authigenic formation of goethite,  
525 based on the results presented here it seems more likely that some of the incoming  
526 goethite is relatively resistant and is slow to undergo reductive dissolution. The  
527 lake is very S-poor, thus appreciable quantities of pyrite do not form even below  
528 the oxycline. Fe that is reduced and dissolved at depth, either in the water column  
529 or in the sediment, can migrate upward to form new amorphous Fe phases in the  
530 water column or in very shallow sediment. Occasional lake mixing events may form  
531 substantial quantities of new  $\text{Fe}^{3+}$ -bearing minerals that are spread throughout the  
532 lake. This would be consistent with the conclusions of Tamuntuan et al. (2015)

533 that the main factor driving diagenesis within short sediment cores is iron oxide  
534 dissolution.

## 535 **5.1 Lake Towuti as an analogue for martian paleolakes?**

536 The mafic/ultramafic East Sulawesi Ophiolite shares some compositional sim-  
537 ilarities with the mafic martian crust (Bibring et al. 2005; McSween et al. 2009;  
538 Kadarusman et al. 2004). Examining chemical and mineralogical relationships in  
539 Mg/Fe-rich sediment in the Lake Towuti system may have implications for interpret-  
540 ing geochemical data of ancient lacustrine mudstones on Mars (Weber et al. 2015;  
541 Goudge et al. 2017).

542 The NASA Mars Curiosity rover is currently characterizing the chemistry and  
543 mineralogy of a  $\sim 5$  km thick sequence of strata informally known as Mt. Sharp  
544 that is hosted within Gale Crater (Grotzinger et al. 2014). The overall abundance  
545 of FeO has remained relatively constant for a large part of the stratigraphic section;  
546 however, the mineral hosts of Fe have changed tremendously (e.g. Hurowitz et al.  
547 (2017); Rampe et al. (2017)), leading to various theories on the structure of the lake,  
548 as discussed more below. In Lake Towuti, there is a similar overall steadiness in  
549 elemental  $Fe_{total}$  (Figs. 5, 10) with great diversity in Fe mineralogy (Fig. 8). While  
550 redox transformations of iron do support many types of microbial life, especially  
551 in warm tropical environments, this will catalyze reactions that may otherwise be  
552 thermodynamically feasible without them, increasing the rate but not necessarily  
553 fundamentally altering the way iron may be cycled in the environment (Davison  
554 1993). Though the tropical setting of the Towuti system is not a perfect analogue  
555 for Mars, and though biology and the presence of organic matter undoubtedly play  
556 an important role in redox reactions, it nevertheless provides a modern setting for  
557 understanding how to infer sediment-water interactions in a redox stratified lake  
558 via chemical and mineralogical information. In the case of Lake Towuti, VNIR  
559 reflectance spectra are linked to chemical and mineralogical trends in bulk sediment  
560 samples. Similarly, certain elements (Al, Mg) provide useful proxies for mineralogy.

561 Combined with bulk chemical measurements, VNIR spectra thus provide a rapid and  
562 non-destructive means to gain a first order understanding of sediment deposition and  
563 diagenesis across modern Lake Towuti and presumably through time, and similar  
564 approaches may be applied to lacustrine sequences on Mars.

565 The mudstones of Mt. Sharp in particular contain a range of hydrous and Fe-  
566 bearing minerals, including primary igneous phases (olivine, pyroxene), sulfides/sulfates,  
567 Fe-bearing clay minerals, hematite, and magnetite (Grotzinger et al. 2014; Vaniman  
568 et al. 2014; Rampe et al. 2017). Trends in the chemistry and mineralogy of these  
569 rocks are still being examined, but variations in Fe mineralogy and oxidation state  
570 have been observed as a function of stratigraphic position by the rover (Hurowitz  
571 et al. 2017) as well as from orbit (Milliken et al. 2010). Variations in iron mineral-  
572 ogy, and correlations between elements such as Zn, Ni, and Si, have been interpreted  
573 as evidence for changing martian environmental conditions during the evolution of  
574 the Gale crater lake. Hurowitz et al. (2017) recently suggested that some observa-  
575 tions are consistent with the presence of a redox-stratified lake in which oxidizable  
576 cations become enriched through photo-oxidation of reduced (ferrous) groundwaters  
577 that seep into the lake. In this model, mineral assemblages vary as a function of  
578 lake water depth, wherein magnetite and silica phases precipitate in the reducing  
579 deep water below the oxycline and hematite-phyllsilicate assemblages reflect more  
580 oxidizing, shallow water influenced by the surface environment. The transition from  
581 magnetite and clay-bearing facies to strata dominated by hematite and sulfate, in  
582 addition to recent identification of potential desiccation cracks in the mudstones,  
583 may be evidence of fluctuations in lake level. However, others have suggested that  
584 trends in redox-sensitive elements in Gale Crater indicate acid sulfate alteration of  
585 mafic materials (Yen et al. 2017; Rampe et al. 2017)

586 Characterizing the chemistry and mineralogy of sediments in a modern terrestrial  
587 redox stratified lake in a mafic/ultramafic catchment can also provide useful infor-  
588 mation on sediment transport and alteration processes, as well as evidence of what  
589 environmental factors may ultimately be recorded in the sedimentary rock record.

590 In the Malili Lakes system, VNIR spectral characteristics can be used to infer both  
591 chemistry and mineralogy (Weber et al. 2015; Goudge et al. 2017), and this has rel-  
592 evant applications for the study of martian paleolake basins. Strata in Gale Crater,  
593 for example, has orbital spectral evidence of clay minerals (Milliken et al. 2010),  
594 *in situ* XRD measurements that confirm the presence of clays in this crater, and  
595 evidence for significant fractions of X-ray amorphous components (Grotzinger et al.  
596 2014; Vaniman et al. 2014; Rampe et al. 2017). Rover payloads are inherently lim-  
597 ited in terms of analytical ability, and understanding source regions and hydrologic  
598 history through clay mineralogy in martian paleolakes requires distinguishing chem-  
599 ical and mineralogical trends using remote techniques such as VNIR spectroscopy  
600 (Ehlmann et al. 2008; Milliken et al. 2010). Lake Towuti is a useful analogue in  
601 that it provides proof of the link between chemistry, mineralogy, and spectroscopy  
602 in lacustrine sediments.

603 One intriguing observation in Lake Towuti sediments that may be useful for  
604 understanding martian counterparts is the relationship between the abundance of Si  
605 and Ti. A positive correlation between these elements has been observed at several  
606 sites on Mars, including lacustrine mudstones in Gale Crater and non-lacustrine  
607 rocks and soils in Gusev Crater. This trend has been cited as evidence for acid  
608 sulfate leaching in both locations (Squyres et al. 2008; Morris et al. 2008; Yen et al.  
609 2017; Rampe et al. 2017). A similar correlation is seen in Lake Towuti sediment  
610 samples (positive coefficient significant at  $p < 0.01$ ) (Fig. 13, Table 6), where it  
611 has been interpreted to be driven by trends in physical weathering and transport  
612 under different climate conditions (i.e. the integration of rainfall, erosion, and fluvial  
613 discharge in the area). An increase in Ti concentration in this system has been used  
614 as a proxy for increased physical weathering and runoff in the source region because  
615 this leads to an increase in the detrital influx of rutile and/or ilmenite (Russell  
616 et al. 2014). The similarity of the strong positive correlation between Si and Ti in  
617 the circum-neutral (Costa et al. 2015) Lake Towuti system suggests that lacustrine  
618 sediments in Gale Crater that exhibit this elemental relationship need not have been

619 formed in or affected by low-pH fluids, particularly if Ti-bearing phases are available  
620 and variable in the sediment source region.

621 As with any terrestrial analogue, the influence of organic matter and biology on  
622 the rate of redox processes in Towuti must be considered, though Lake Towuti is  
623 notable for being one of the least productive tropical lakes on the planet (Haffner  
624 et al. 2001). Nontronite is also a useful mineral for addressing the water oxidation  
625 state on Earth and Mars. There is evidence that this ferric smectite may form  
626 from ferrous precursors precipitated under initially reducing conditions (Muller and  
627 Forstner 1973; Harder 1976; Pedro et al. 1978; Decarreau and Bonnin 1986), and  
628 thus its presence could indicate a change in oxidation state. Nontronite has been  
629 identified in several places on Mars, including Gale Crater (Poulet et al. 2005; Bishop  
630 et al. 2008; Milliken et al. 2010; Ehlmann et al. 2009), and it seems to be present in  
631 some of the Towuti laterite samples and thus may represent a detrital component in  
632 the lake surface samples. However, if some of the nontronite in the lake sediments  
633 is authigenic then it could provide valuable information on changes in the redox  
634 interface in Lake Towuti at those locations, and future study of this phase in these  
635 samples is warranted. Similarly, *in situ* identification of nontronite by Curiosity  
636 could be used to constrain redox conditions and evolution in the Gale crater lake,  
637 particularly if it over or underlies beds with greater proportions of ferric oxides.

## 638 **6 Conclusions**

639 Sediment throughout the Malili Lakes system is highly variable in terms of min-  
640 eralogy and bulk chemistry, particularly with respect to the type and abundance of  
641 secondary Fe-bearing phases. VNIR spectral parameters are well correlated with the  
642 relevant ICP elemental abundances, demonstrating that reflectance spectra, which  
643 require little to no sample preparation, can provide a useful, rapid way to infer bulk  
644 chemistry of these sediments. The role and distribution of Fe in Lake Towuti is  
645 complex: abundant, easily characterizable crystalline Fe phases in lateritic soils are  
646 rapidly altered after deposition on the lake surface to complex forms of Fe, as seen



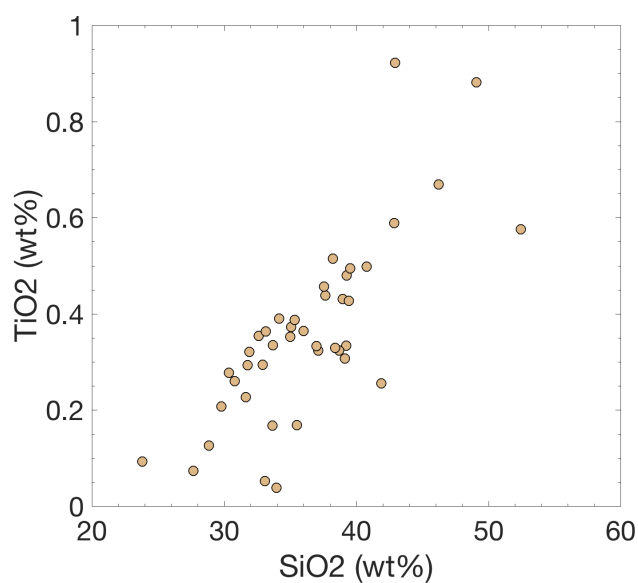


Figure 13: ICP-determined Si and Ti abundances for Lake Towuti surface samples.

Table 6: Si Ti Regression Results

	<i>Dependent variable:</i>
	TiO <sub>2</sub> , surface
SiO <sub>2</sub> , surface	0.026*** (0.003)
Constant	-0.594*** (0.125)
Observations	42
R <sup>2</sup>	0.597
Residual Std. Error	0.121 (df = 40)
F Statistic	59.187*** (df = 1; 40)
<i>Note:</i>	*p<0.1; **p<0.05; ***p<0.01

647 in XRD patterns, VNIR reflectance spectra, and Mössbauer spectra. The processes  
648 that drive observed changes in Fe mineralogy are not known with certainty, but  
649 Towuti is a redox-stratified lake with low primary productivity, thus the observed  
650 changes could be facilitated by Fe cycling across the oxycline. Consistent with this,  
651 sequential Fe extractions show a decrease in the ratio of crystalline to amorphous  
652 Fe phases through the transport system from laterites to surface sediment to cores.

653 The spatial variability in Fe hosts across the modern lake sediment surface is likely  
654 affected by variations in source composition as well as Fe cycling in the water column  
655 and early diagenetic processes. Understanding the effects of transport, deposition,  
656 and diagenesis on sediment composition, particularly changes in Fe mineralogy, in  
657 a terrestrial mafic/ultramafic system such as Lake Towuti will continue to aid in  
658 the interpretation and climatic reconstruction of this region, and these results may  
659 also have utility for interpreting Fe-rich mineral assemblages in lacustrine sediments  
660 observed on Mars.

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