

Swarm Precise Orbit Determination and ionospheric signatures

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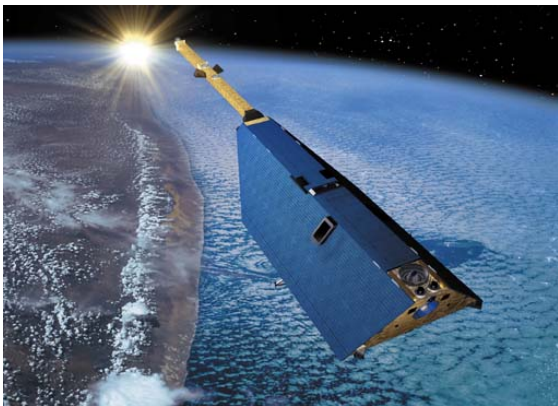
AIUB

Astronomical Institute
University of Bern



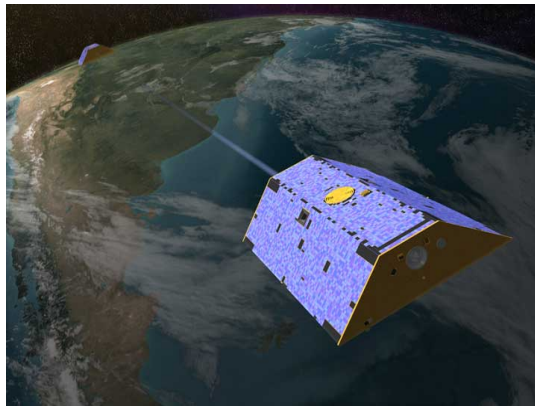
Low Earth Orbiters (LEOs)

CHAMP



CHallenging
Mini**sat**ellite **P**ayload

GRACE



Gravity **R**ecovery **A**nd
Climate **E**xperiment

GOCE



Gravity and
steady-state **O**cean
Circulation **E**xplorer

But there are many more missions equipped with GPS receivers ...

Jason



Jason-2



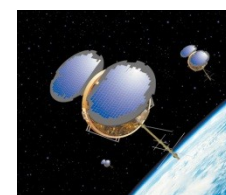
MetOp-A



Icesat



COSMIC



LEO Constellations

TanDEM-X



Swarm



Sentinel

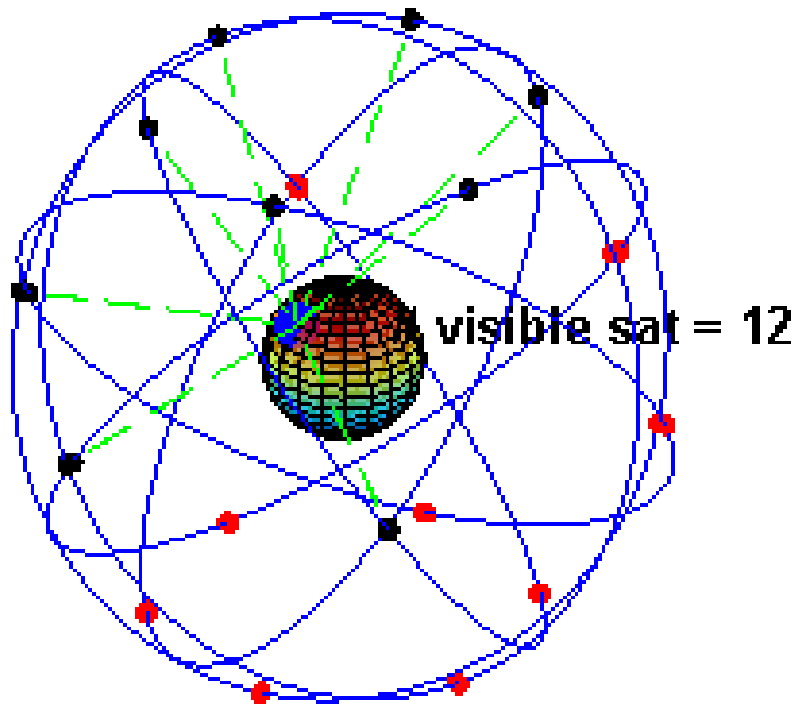


and of course, in the future

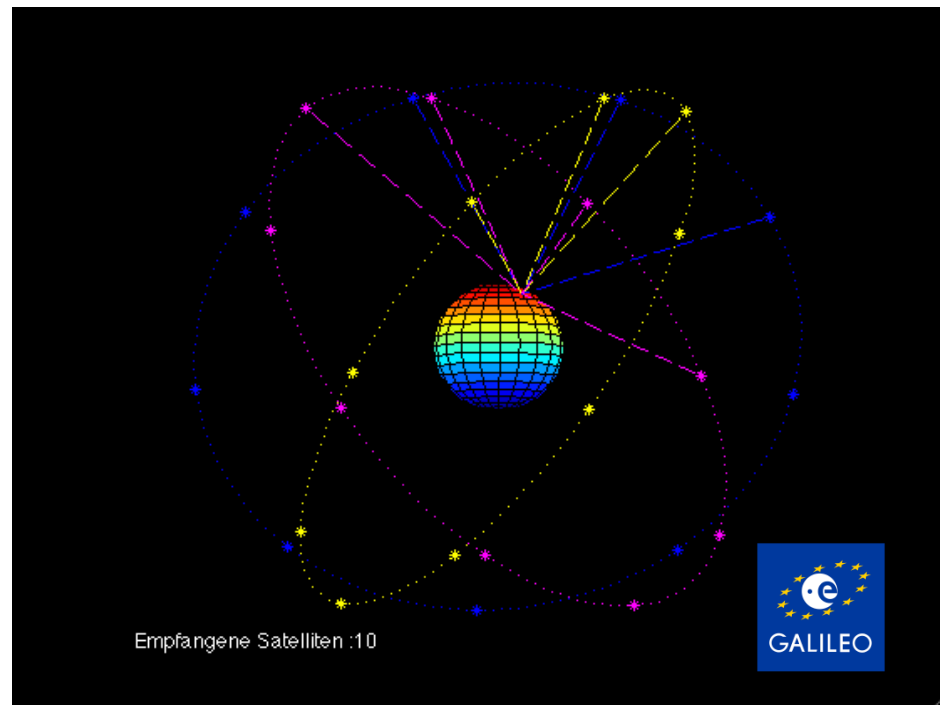


Global Navigation Satellite Systems (GNSS)

GPS

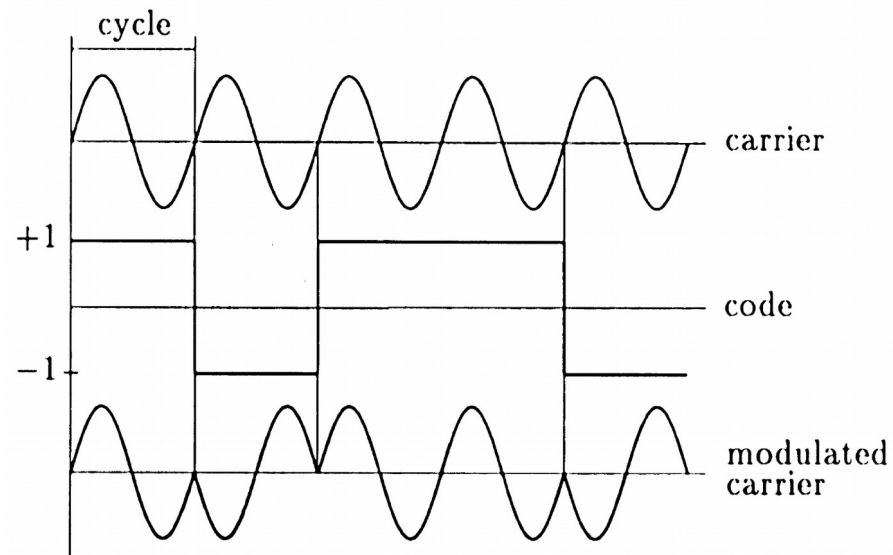
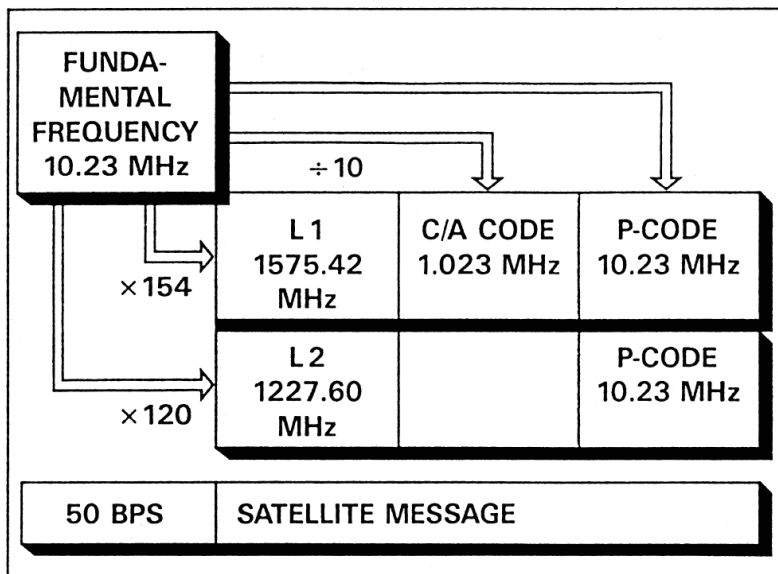


Galileo



Other GNSS are already existing (GLONASS) or being built up (Galileo, Beidou), but there are no multi-GNSS spaceborne receivers (with open data policy) in LEO orbit yet.

GPS Signals



Bits encoded on carrier by phase modulation:

Signals driven by an **atomic clock**

Two **carrier signals** (sine waves):

- L_1 : $f = 1575.43$ MHz, $\lambda = 19$ cm
- L_2 : $f = 1227.60$ MHz, $\lambda = 24$ cm

- **C/A-code** (Clear Access / Coarse Acquisition)
- **P-code** (Protected / Precise)
- **Broadcast/Navigation Message**

Improved Observation Equation

$$L_i^k = \rho_i^k - c \cdot \Delta t^k + c \cdot \Delta t_i + \cancel{I_i^k} + \cancel{I_t^k} + \lambda \cdot N_i^k + \Delta_{rel} - c \cdot b^k + c \cdot b_i + m_i^k + \epsilon_i^k$$

ρ_i^k	Distance between satellite and receiver	←	Satellite positions and clocks
Δt^k	Satellite clock offset wrt GPS time	←	are known from ACs or IGS
Δt_i	Receiver clock offset wrt GPS time		
T_i^k	Tropospheric delay	←	Not existent for LEOs
I_i^k	Ionospheric delay	←	Cancels out (first order only) when forming the ionosphere-free linear combination:
N_i^k	Phase ambiguity		
Δ_{rel}	Relativistic corrections		
b^k	Delays in satellite (cables, electronics)		
b_i	Delays in receiver and antenna		
m_i^k	Multipath, scattering, bending effects		
ϵ_i^k	Measurement error		

$$L_c = \frac{f_1^2}{f_1^2 - f_2^2} L_1 - \frac{f_2^2}{f_1^2 - f_2^2} L_2$$

Geometric Distance

Geometric distance ρ_{leo}^k is given by:

$$\rho_{leo}^k = |\mathbf{r}_{leo}(t_{leo}) - \mathbf{r}^k(t_{leo} - \tau_{leo}^k)|$$

\mathbf{r}_{leo} Inertial position of LEO antenna phase center at reception time

\mathbf{r}^k Inertial position of GPS antenna phase center of satellite k at emission time

τ_{leo}^k Signal traveling time between the two phase center positions

Different ways to represent \mathbf{r}_{leo} :

- **Kinematic** orbit representation
- **Dynamic** or **reduced-dynamic** orbit representation

Kinematic Orbit Representation (1)

Satellite position $\mathbf{r}_{leo}(t_{leo})$ (in inertial frame) is given by:

$$\mathbf{r}_{leo}(t_{leo}) = \mathbf{R}(t_{leo}) \cdot (\mathbf{r}_{leo,e,0}(t_{leo}) + \delta\mathbf{r}_{leo,e,ant}(t_{leo}))$$

\mathbf{R} Transformation matrix from Earth-fixed to inertial frame

$\mathbf{r}_{leo,e,0}$ LEO center of mass position in Earth-fixed frame

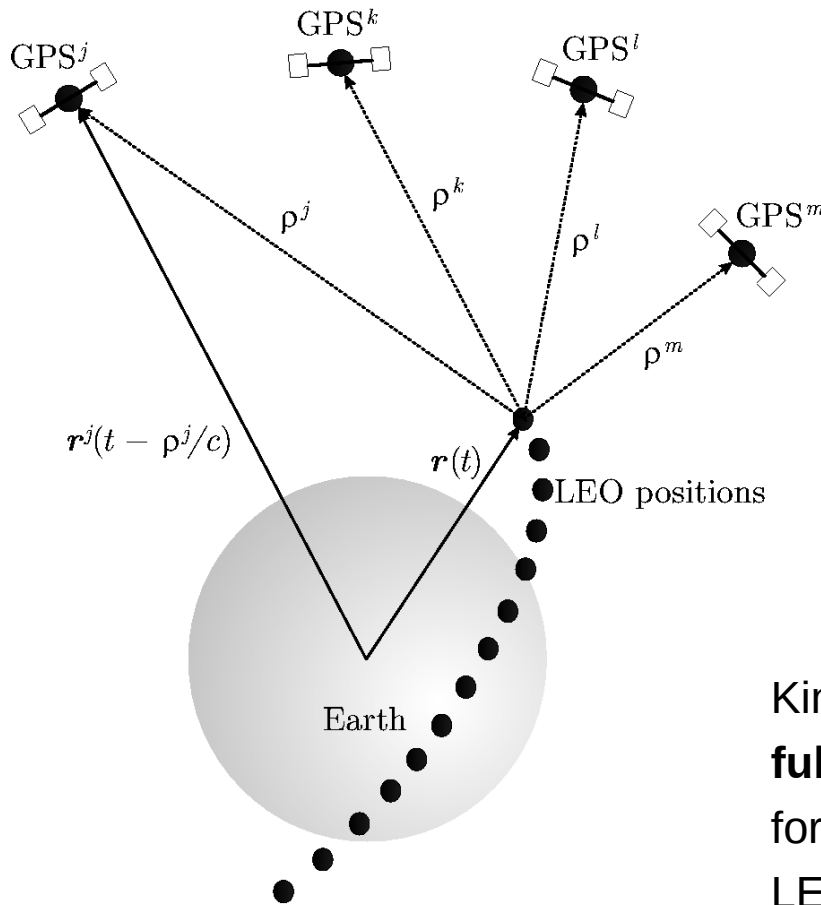
$\delta\mathbf{r}_{leo,e,ant}$ LEO antenna phase center offset in Earth-fixed frame

Kinematic positions $\mathbf{r}_{leo,e,0}$ are estimated for each **measurement epoch**:

- Measurement epochs **need not** to be identical with nominal epochs
- Positions are **independent** of models describing the LEO dynamics.

Velocities cannot be provided

Kinematic Orbit Representation (2)



A kinematic orbit is an ephemeris at **discrete** measurement epochs

Kinematic positions are **fully independent** on the force models used for LEO orbit determination

(Švehla and Rothacher, 2004)

Kinematic Orbit Representation (3)

Measurement epochs
(in GPS time)

Positions (km)
(Earth-fixed)

* 2016 6 1 0 0 17.99999974			
PL49 -2023.517746 -3061.130332 5742.844473	0.262691		
* 2016 6 1 0 0 18.99999989			
PL49 -2026.734429 -3066.703833 5738.746569	0.111976		
* 2016 6 1 0 0 19.99999974			
PL49 -2029.949393 -3072.273033 5734.641440	0.261262		
* 2016 6 1 0 0 20.99999979			
PL49 -2033.162630 -3077.837924 5730.529099	0.210548		
* 2016 6 1 0 0 21.99999984			
PL49 -2036.374136 -3083.398504 5726.409546	0.159835		
* 2016 6 1 0 0 22.99999979			
PL49 -2039.583909 -3088.954755 5722.282787	0.209119		
* 2016 6 1 0 0 23.99999984			
PL49 -2042.791949 -3094.506686 5718.148843	0.158440		
* 2016 6 1 0 0 24.99999979			
PL49 -2045.998248 -3100.054291 5714.007709	0.207778		
* 2016 6 1 0 0 25.99999974			
PL49 -2049.202791 -3105.597545 5709.859370	0.257064		
* 2016 6 1 0 0 26.99999979			
PL49 -2052.405584 -3111.136450 5705.703844	0.206352		

Clock correction to nominal epoch (μ s), e.g., to epoch 00:00:20

Excerpt of kinematic Swarm-C positions at begin of 1 June, 2016

The kinematic orbits may be downloaded at ftp://ftp.unibe.ch/aiub/LEO_ORBITS/

Dynamic Orbit Representation (1)

Satellite position $\mathbf{r}_{leo}(t_{leo})$ (in inertial frame) is given by:

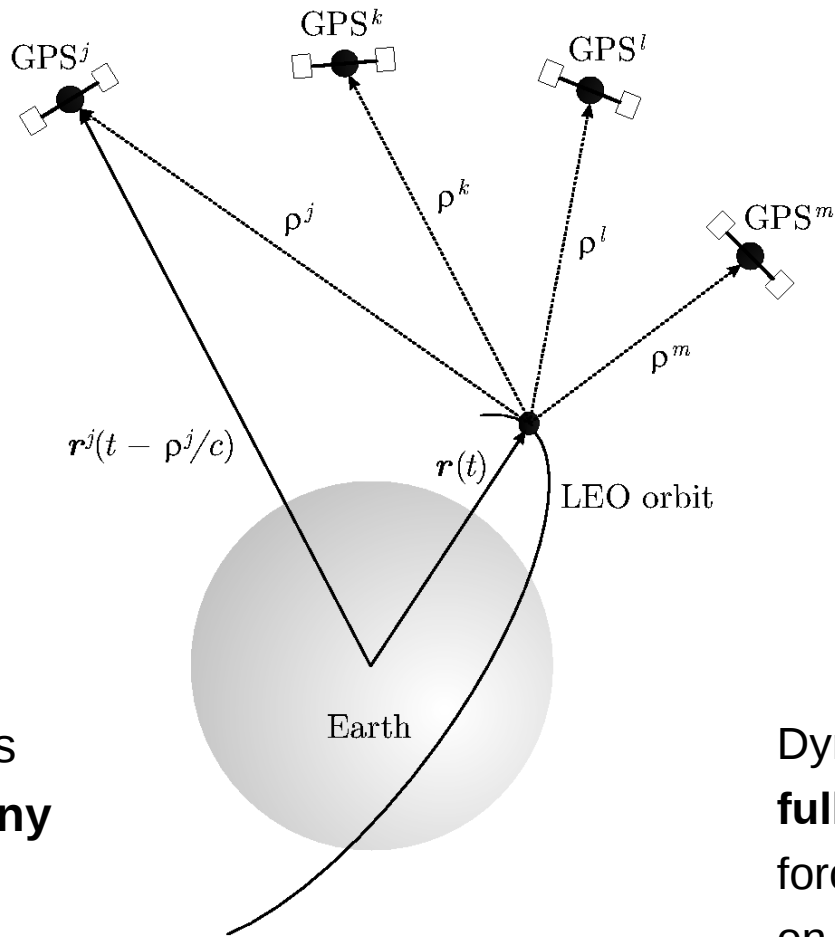
$$\mathbf{r}_{leo}(t_{leo}) = \mathbf{r}_{leo,0}(t_{leo}; a, e, i, \Omega, \omega, u_0; Q_1, \dots, Q_d) + \delta\mathbf{r}_{leo,ant}(t_{leo})$$

$\mathbf{r}_{leo,0}$	LEO center of mass position
$\delta\mathbf{r}_{leo,ant}$	LEO antenna phase center offset
$a, e, i, \Omega, \omega, u_0$	LEO initial osculating orbital elements
Q_1, \dots, Q_d	LEO dynamical parameters

Satellite trajectory $\mathbf{r}_{leo,0}$ is a particular solution of an **equation of motion**

- One set of **initial conditions** (orbital elements) is estimated per arc.
Dynamical parameters of the force model on request

Dynamic Orbit Representation (3)



Dynamic orbit positions may be computed at **any epoch** within the arc

Dynamic positions are **fully dependent** on the force models used, e.g., on the gravity field model

Reduced-Dynamic Orbit Representation (1)

Equation of motion (in inertial frame) is given by:

$$\ddot{\mathbf{r}} = -GM \frac{\mathbf{r}}{r^3} + \mathbf{f}_1(t, \mathbf{r}, \dot{\mathbf{r}}, Q_1, \dots, Q_d, P_1, \dots, P_s)$$

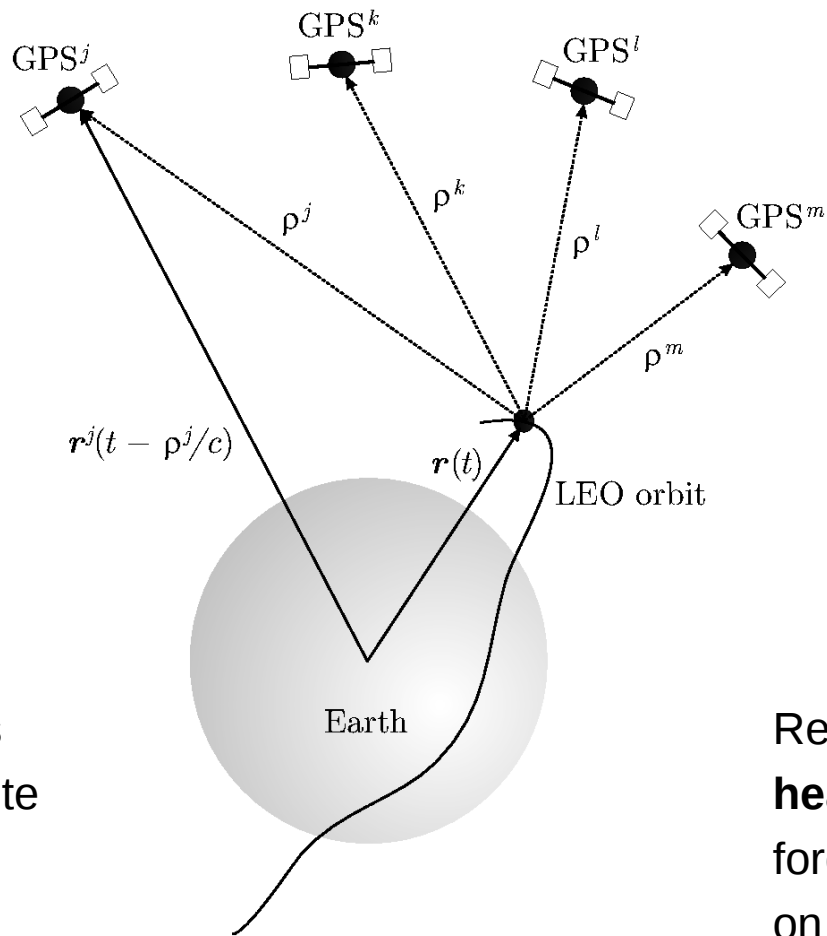
P_1, \dots, P_s

Pseudo-stochastic parameters

Pseudo-stochastic parameters are:

- additional empirical parameters characterized by a priori known **statistical properties**, e.g., by expectation values and a priori variances
- useful to **compensate** for deficiencies in dynamic models, e.g., deficiencies in models describing non-gravitational accelerations
- often set up as **piecewise constant accelerations** to ensure that satellite trajectories are continuous and differentiable at any epoch

Reduced-Dynamic Orbit Representation (2)



Reduced-dynamic orbits are well suited to compute LEO orbits of **highest quality**

Reduced-dynamic orbits **heavily depend** on the force models used, e.g., on the gravity field model

(Jäggi et al., 2006; Jäggi, 2007)

Reduced-dynamic Orbit Representation (3)

Position epochs
(in GPS time)

Positions (km) &
Velocities (dm/s)
(Earth-fixed)

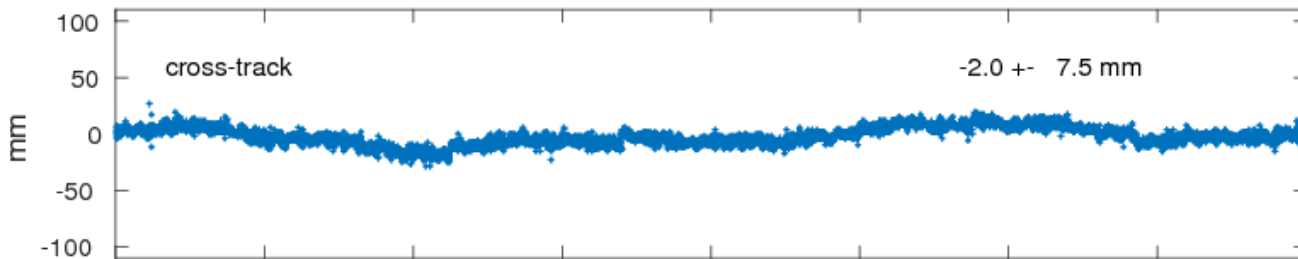
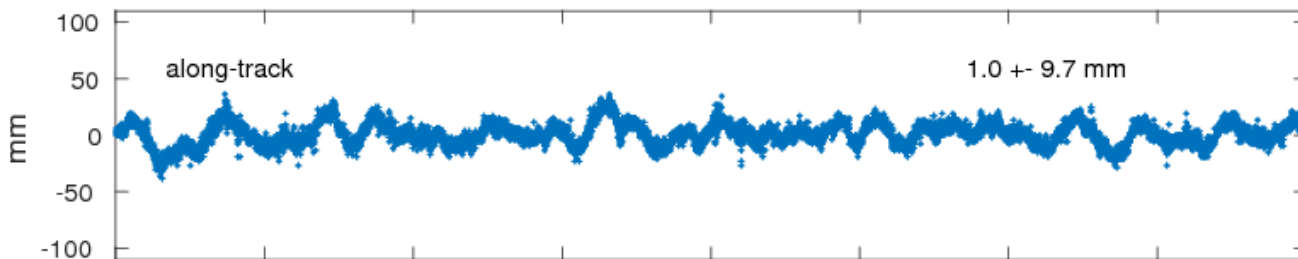
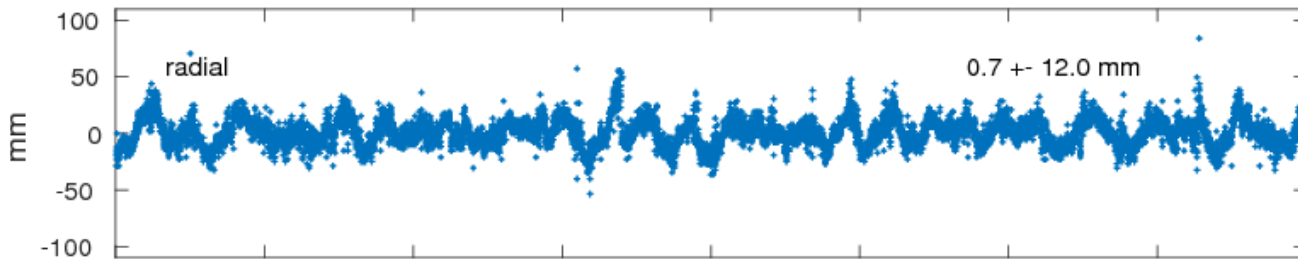
* 2016 6 1	0	0	0.00000000		
PL49	-1965.328762	-2960.079621	5815.366063	999999.999999	
VL49	-32476.530949	-56518.428574	-39633.949261	999999.999999	
* 2016 6 1	0	0	10.00000000		
PL49	-1997.722965	-3016.388318	5775.367094	999999.999999	
VL49	-32311.097194	-56097.834133	-40363.154274	999999.999999	
* 2016 6 1	0	0	20.00000000		
PL49	-2029.949403	-3072.273033	5734.641439	999999.999999	
VL49	-32141.000143	-55670.464832	-41087.301898	999999.999999	
* 2016 6 1	0	0	30.00000000		
PL49	-2062.003415	-3127.727011	5693.194205	999999.999999	
VL49	-31966.250891	-55236.380456	-41806.300697	999999.999999	
* 2016 6 1	0	0	40.00000000		
PL49	-2093.880357	-3182.743574	5651.030585	999999.999999	
VL49	-31786.861194	-54795.641569	-42520.059993	999999.999999	
* 2016 6 1	0	0	50.00000000		
PL49	-2125.575594	-3237.316095	5608.155863	999999.999999	
VL49	-31602.843520	-54348.309592	-43228.489711	999999.999999	
* 2016 6 1	0	1	0.00000000		
PL49	-2157.084506	-3291.438018	5564.575411	999999.999999	
VL49	-31414.211010	-53894.446726	-43931.500489	999999.999999	

Clock corrections
are not provided

Excerpt of reduced-dynamic Swarm-C positions at begin of 1 June, 2016

Orbit Differences KIN-RD (Swarm-C)

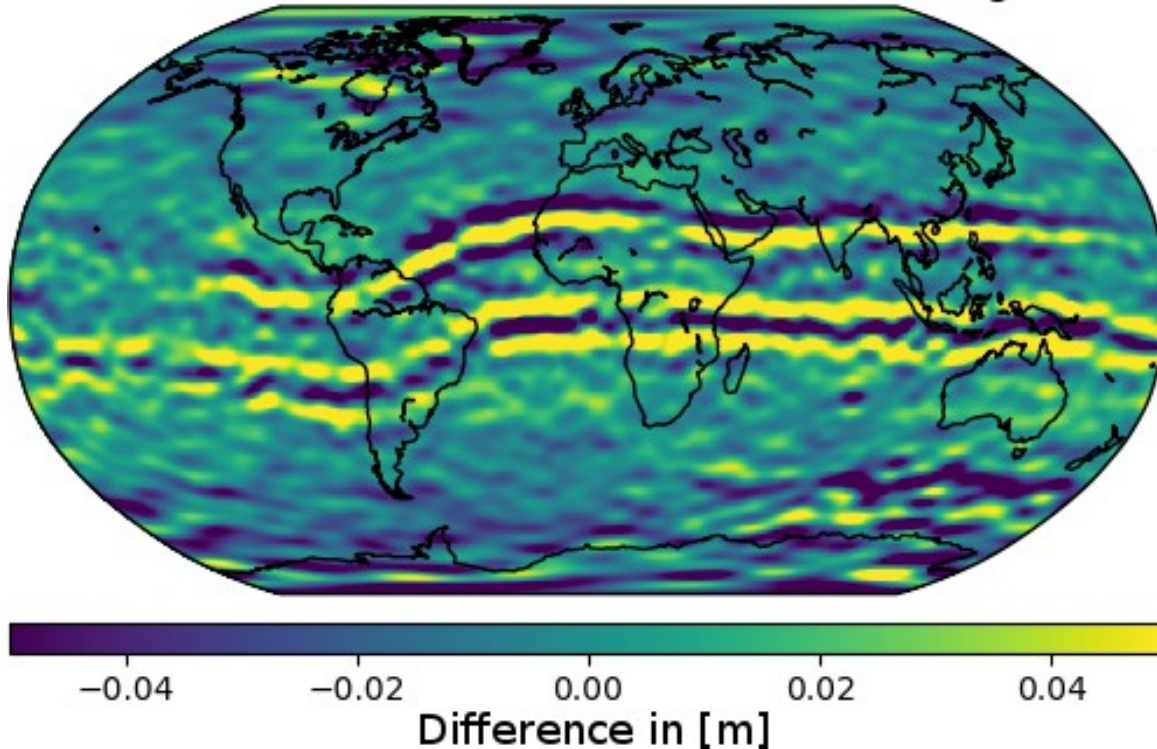
Differences at epochs of kin. positions



Hours of 1 June, 2016

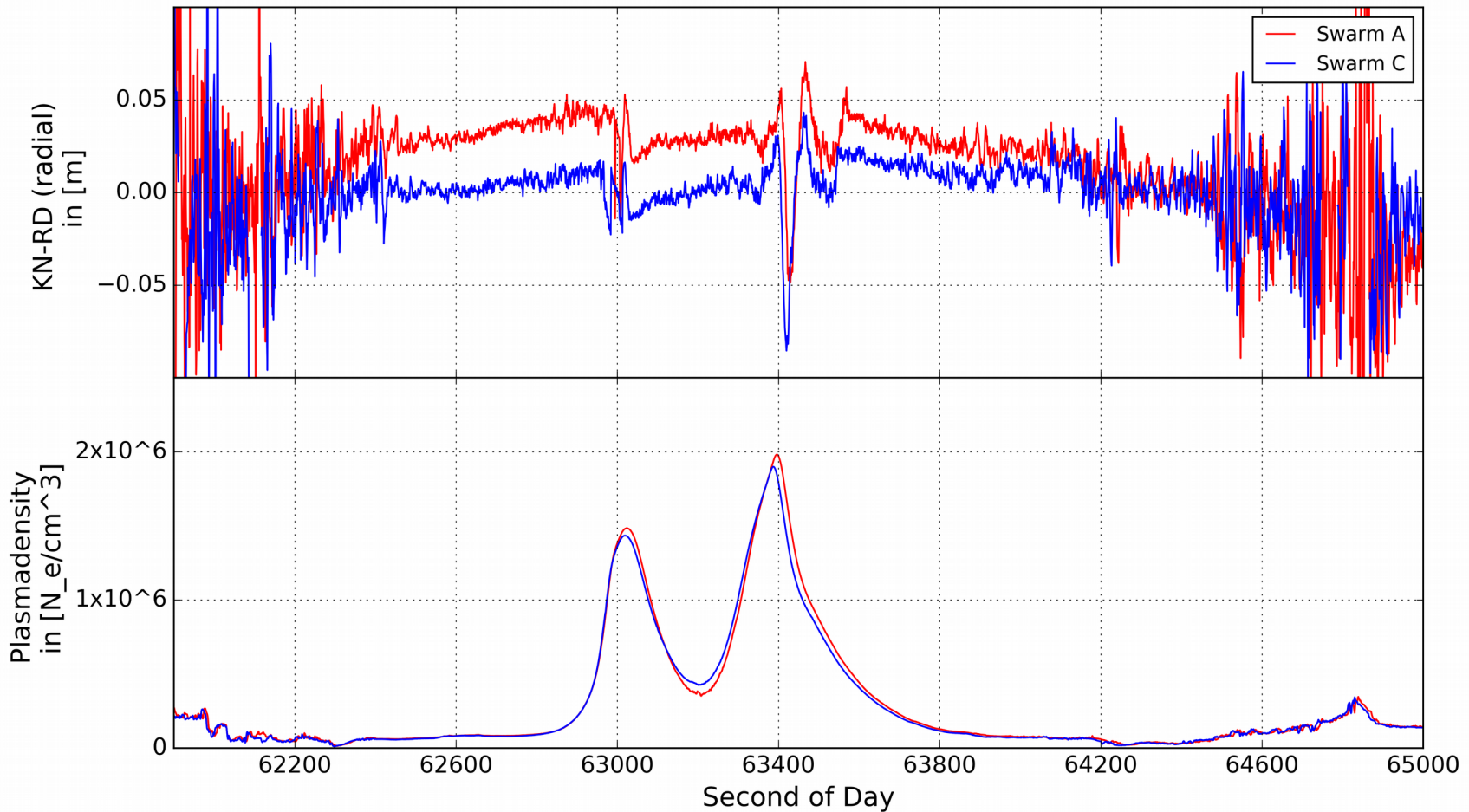
Swarm A Gravity Field

Swarm-A, Nov 2014, 400km Gauss, unweighted



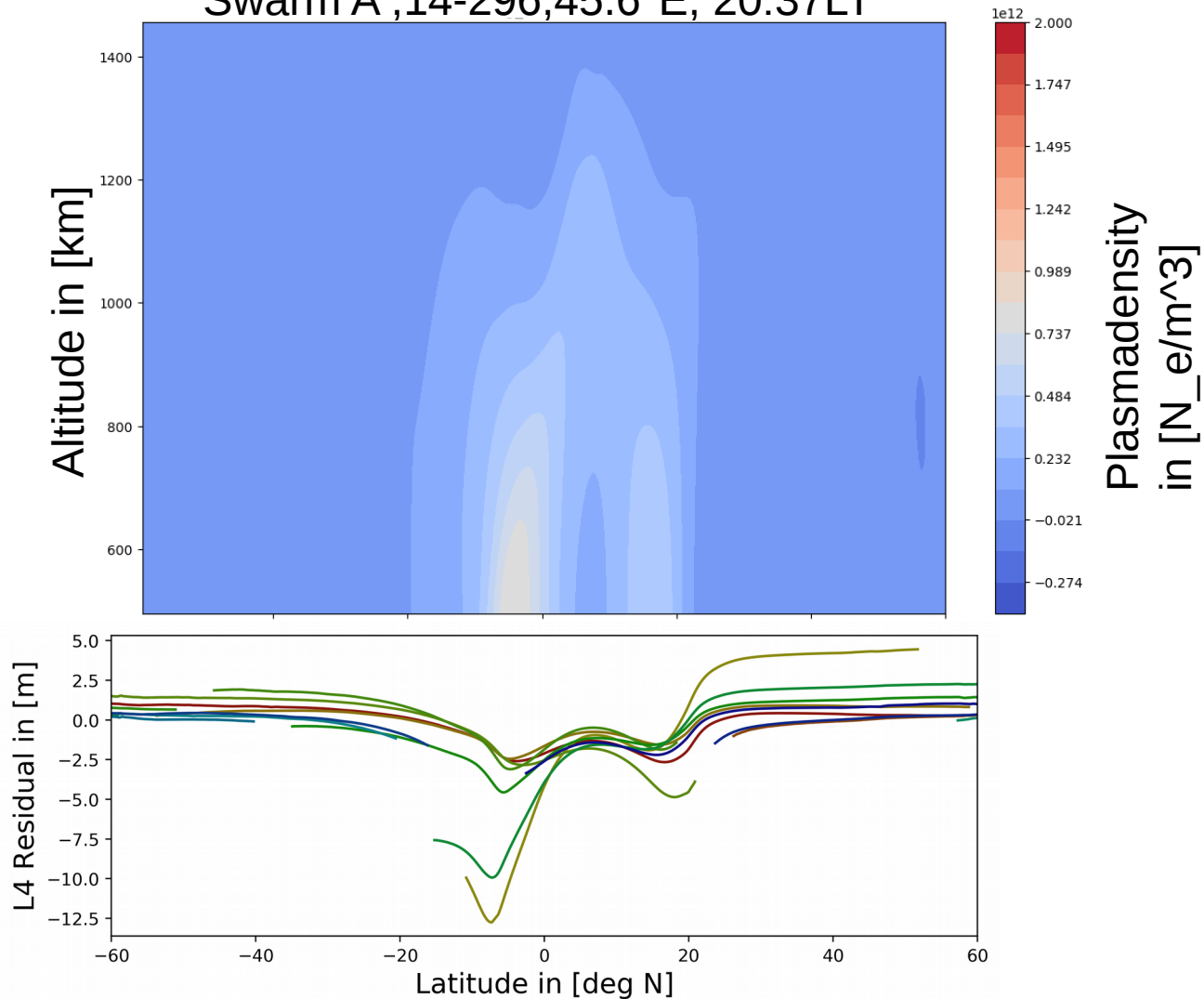
Geoid height differences, static GRACE gravity field AIUB-GRACE03S - Swarm A gravity field, November 2014

Orbit Differences KIN-RD and Plasmadensity



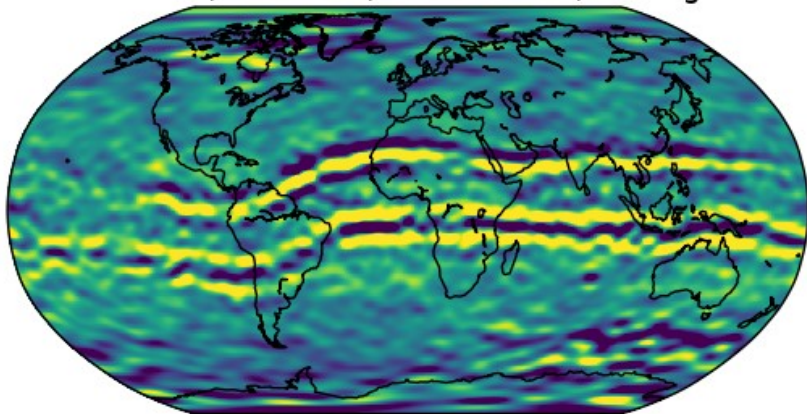
Orbit Differences KIN-RD and Plasmadensity

Swarm A ,14-296,45.6°E, 20:37LT

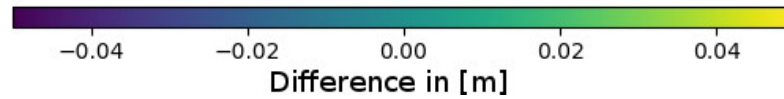
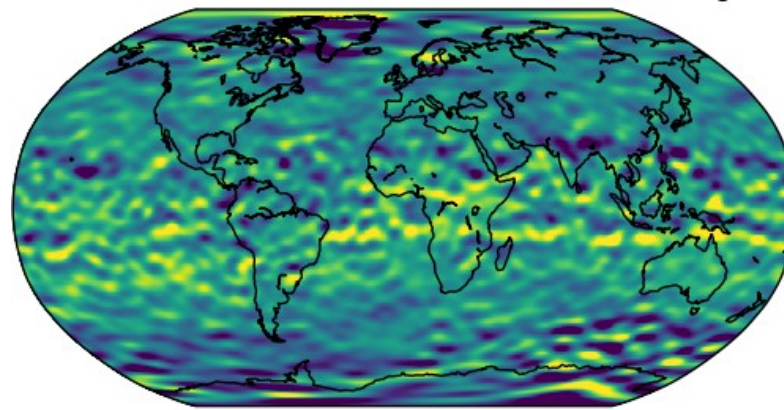


Previous Studies: AIUB-Cutoff

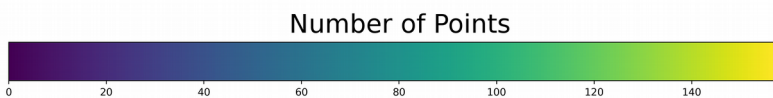
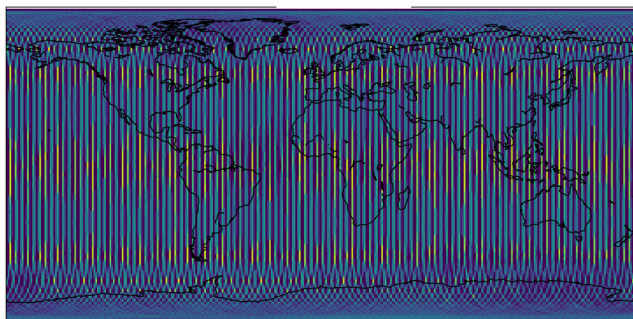
Swarm-A, Nov 2014, 400km Gauss, unweighted



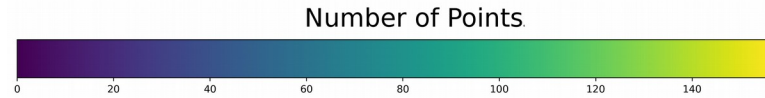
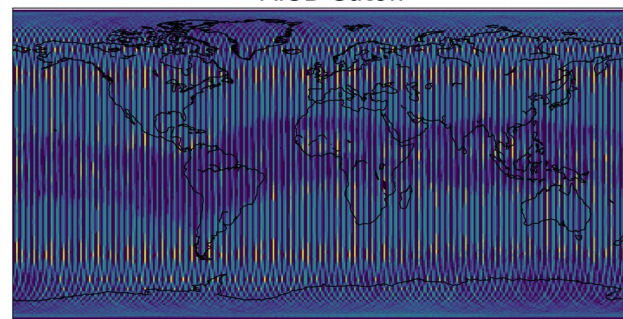
Swarm A, Nov 2014, AIUB RINEX screening



Original Datapoints



AIUB Cutoff

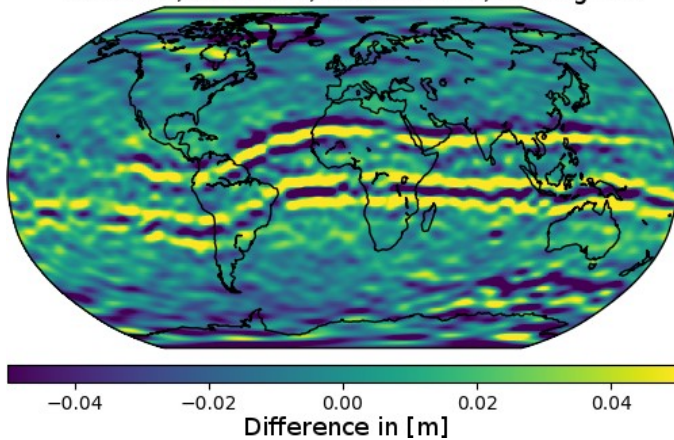


Previous Studies: Graz-ROTI

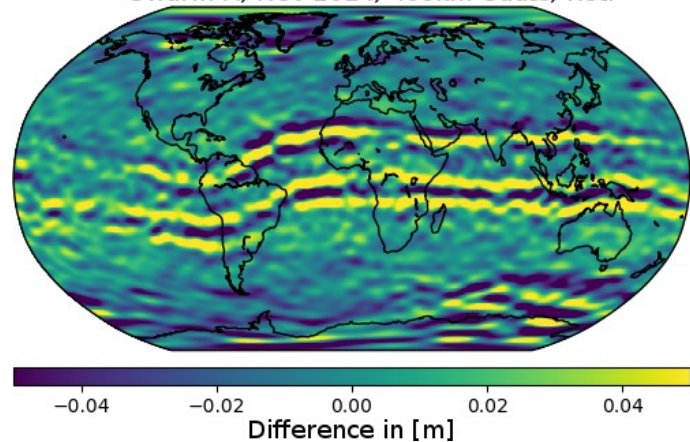
$$ROTI = \sqrt{\frac{\langle \Delta TEC^2 \rangle - \langle \Delta TEC \rangle^2}{dt^2}}$$

- Applied to ΔTEC , moving window, 31s
- $\sigma = 10 \cdot ROTI \cdot \sigma_0$ if $ROTI > 0.1$
- Developed for GOCE

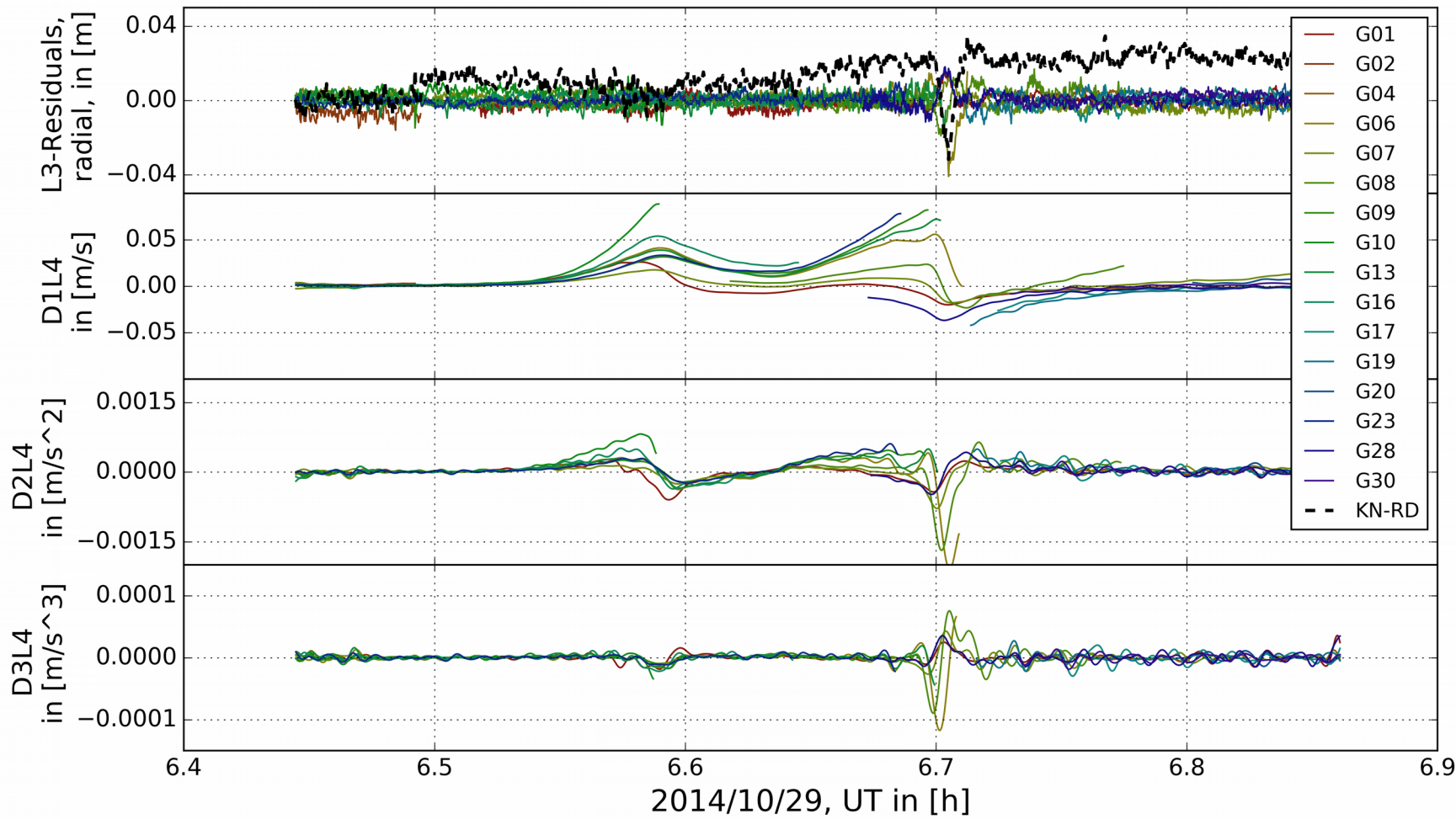
Swarm-A, Nov 2014, 400km Gauss, unweighted



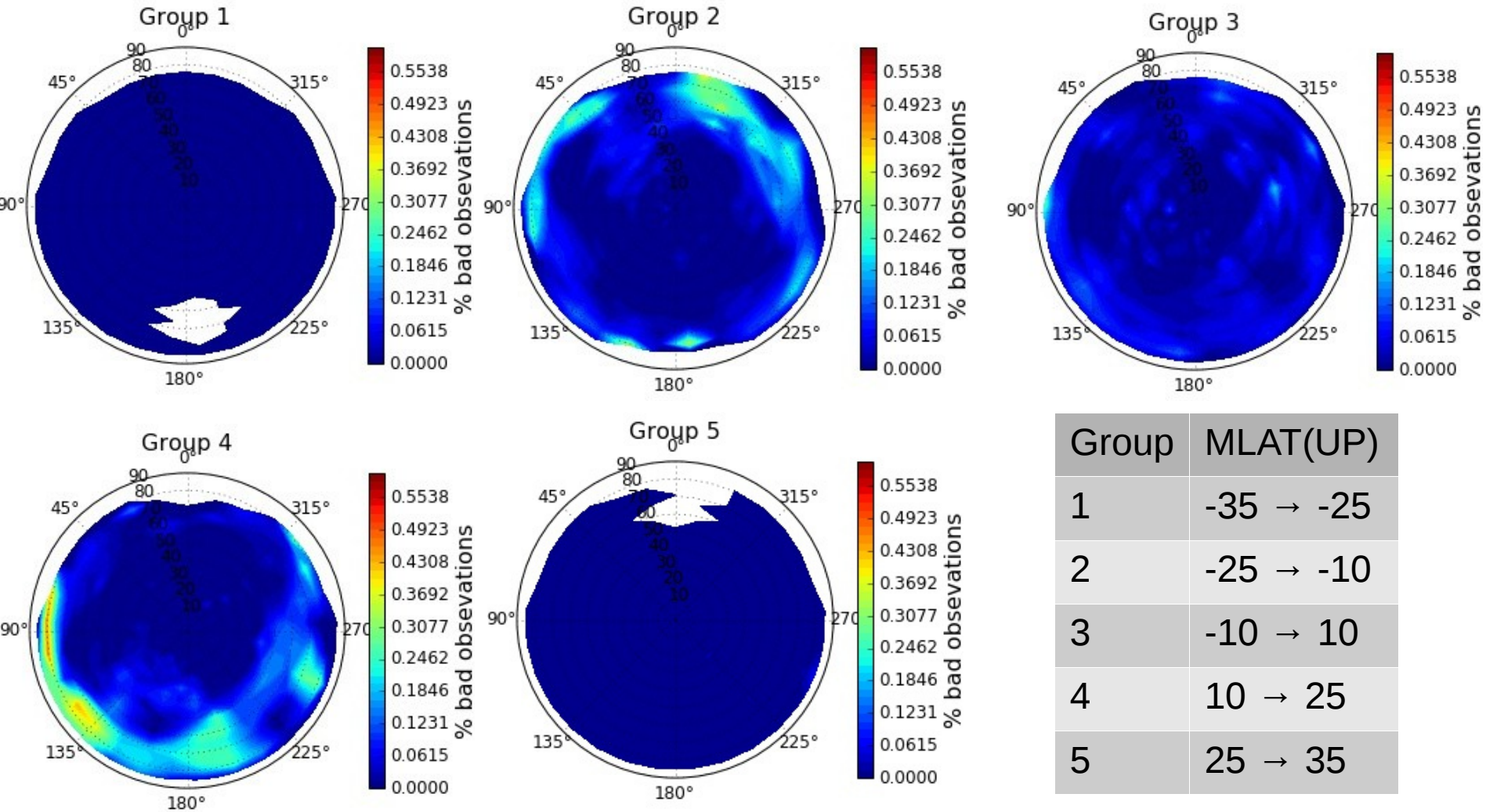
Swarm-A, Nov 2014, 400km Gauss, Roti



Phase Residuals and Time derivatives $L4$

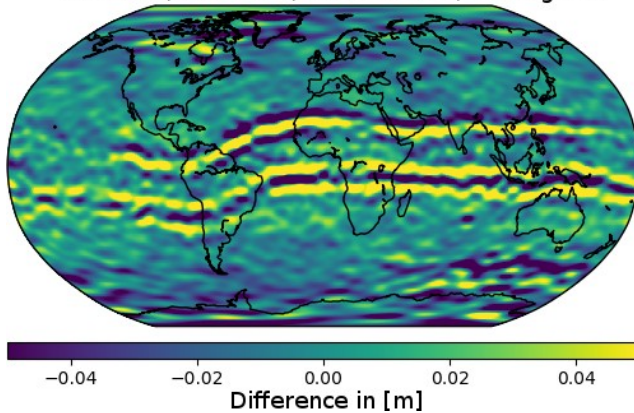


Elevations and azimuth of affected GPS satellites

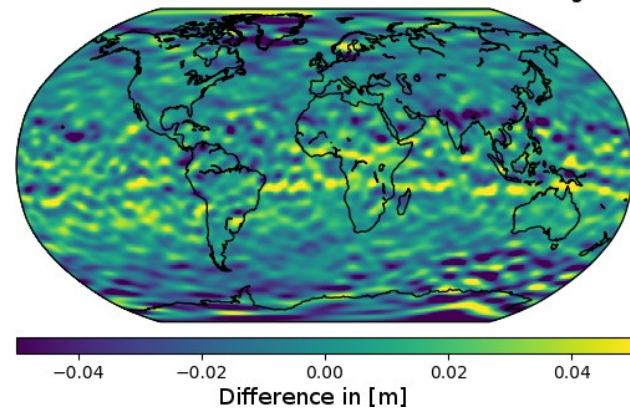


Elevations and azimuth of affected GPS satellites

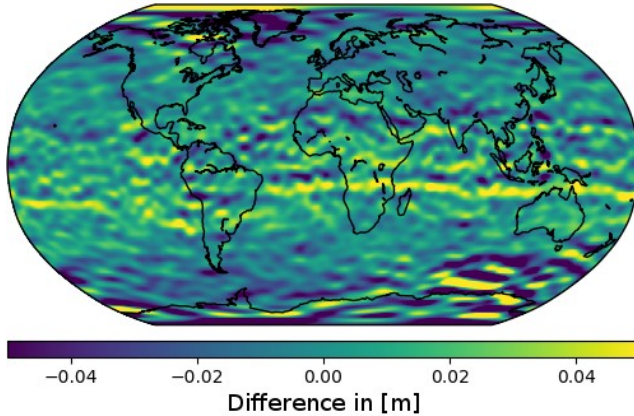
Swarm-A, Nov 2014, 400km Gauss, unweighted



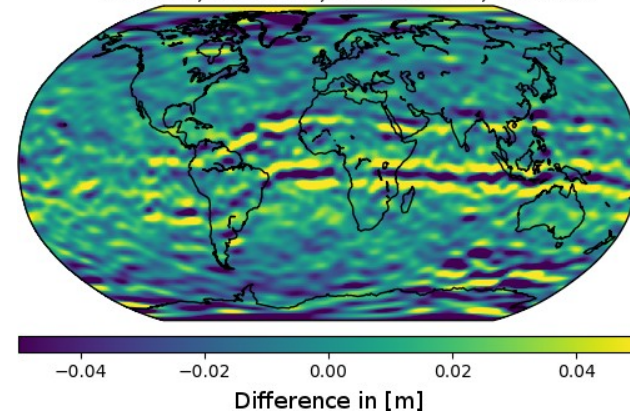
Swarm A, Nov 2014, AIUB RINEX screening



Swarm-A, Nov 2014, 400km Gauss, d^2L4/dt^2

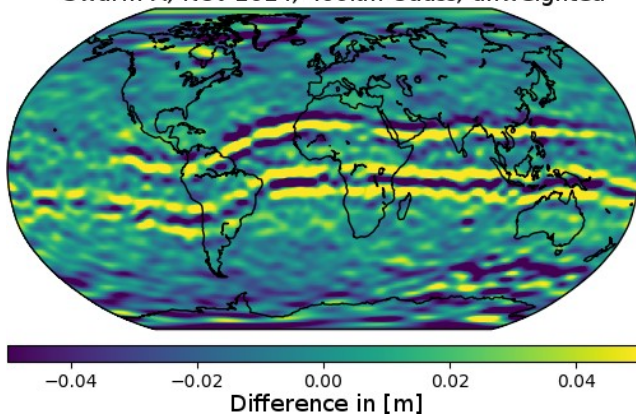


Swarm-A, Nov 2014, 400km Gauss, Model LS

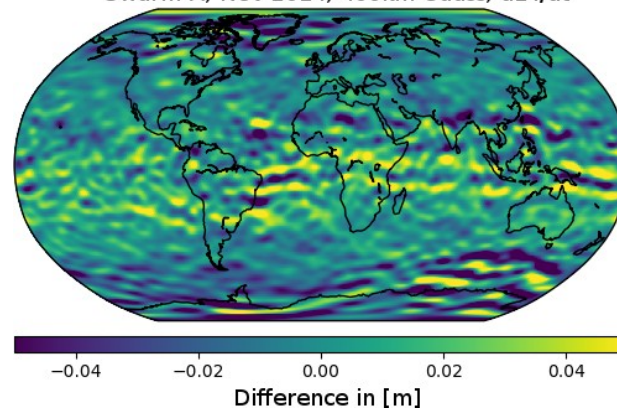


Elevations and azimuth of affected GPS satellites

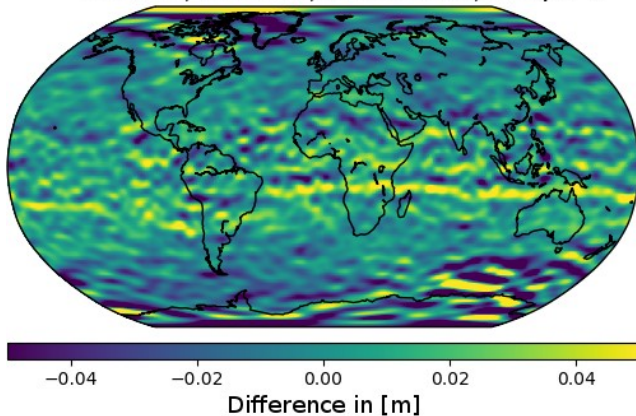
Swarm-A, Nov 2014, 400km Gauss, unweighted



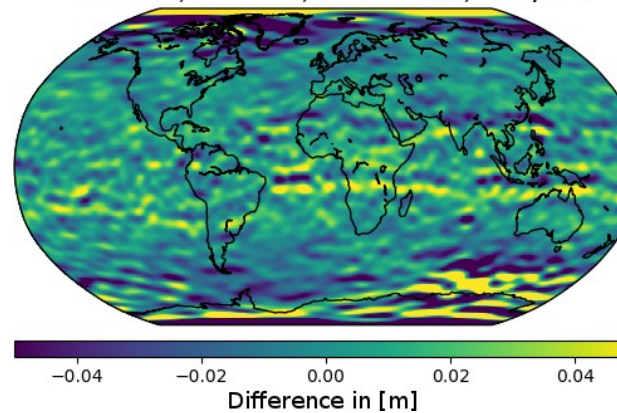
Swarm-A, Nov 2014, 400km Gauss, $dL4/dt$



Swarm-A, Nov 2014, 400km Gauss, d^2L4/dt^2



Swarm-A, Nov 2014, 400km Gauss, d^3L4/dt^3



Evaluation of mehtods

SLR-validation of reduced dynamic orbits:

Method	Mean[mm]	Std[mm]	RMS[mm]
Original	2.6	16.6	16.8
ROTI	2.7	17.3	17.5
AIUB Cutoff	7.4	32.1	32.9
Model	3.4	14.0	14.4
D1L4	0.3	22.1	22.1
D2L4	2.8	16.0	16.1
D3L4	4.1	13.4	14.4

Thank you for your attention

