

Fidelity for kicked atoms with gravity near a quantum resonance

Rémy Dubertrand,^{1,2} Italo Guarneri,³ and Sandro Wimberger^{2,4}

¹*School of Mathematics, University of Bristol, University Walk, Clifton, Bristol BS8 1TW, United Kingdom*

²*Institut für Theoretische Physik, Ruprecht-Karls-Universität, Philosophenweg 19, D-69120 Heidelberg, Germany*

³*Center for Nonlinear and Complex Systems, Università dell'Insubria, Via Valleggio 11, I-22100 Como, Italy*

⁴*Heidelberg Center for Quantum Dynamics, Philosophenweg 12, D-69120 Heidelberg, Germany*

(Received 6 January 2012; published 12 March 2012)

Kicked atoms under a constant Stark or gravity field are investigated for experimental setups with cold and ultracold atoms. The parametric stability of the quantum dynamics is studied using the fidelity. In the case of a quantum resonance, it is shown that the behavior of the fidelity depends on arithmetic properties of the gravity parameter. Close to a quantum resonance, the long-time asymptotics of the fidelity is studied by means of a *pseudoclassical* approximation introduced by Fishman *et al.* [*J. Stat. Phys.* **110**, 911 (2003)]. The long-time decay of fidelity arises from the tunneling out of pseudoclassical stable islands, and a simple ansatz is proposed which satisfactorily reproduces the main features observed in numerical simulations.

DOI: [10.1103/PhysRevE.85.036205](https://doi.org/10.1103/PhysRevE.85.036205)

PACS number(s): 05.45.Mt, 03.65.Sq, 37.10.Jk

I. INTRODUCTION

The stability of quantum evolution against parametric changes of the quantum Hamiltonian is a subject of wide theoretical and experimental interest. A widely used concept is the fidelity introduced by Peres [1], and the closely related Loschmidt echo [2,3], which is built as the interference pattern between states that are obtained by propagating the same initial state under Hamiltonians, e.g., \hat{H}_0 and \hat{H} , which are slight perturbations of each other. A standard definition of the fidelity is

$$F(\tau) = |\langle \psi | e^{i\hat{H}_0\tau/\hbar} e^{-i\hat{H}\tau/\hbar} | \psi \rangle|^2. \quad (1)$$

The behavior of fidelity in time is known to display some universal properties that reflect the underlying classical dynamics [2,3]. Such properties have been mostly explored for the case of systems, which are chaotic in the classical limit. In this paper, we will study the fidelity in the mixed phase-space regime. For this, the system under consideration is the quantum kicked rotor [4–6].

The main motivation for our analysis here is twofold. First, there has been a growing interest over the last decade in the dynamics of the quantum kicked rotor (and its variants) at and close to the so-called quantum resonances [7], theoretically (see, e.g. [8–16]) as well as experimentally (see, e.g., [17–23]). Second, only recently concepts have been developed to actually access the fidelity in setups based on cold or ultracold quantum gases. The used techniques range from interferometric methods in either internal atomic states [24] or in the center-of-mass motion of the atoms [23] to the time reversal of the dynamics by exploiting the properties of the quantum resonant motion [21,25].

In this paper, we study the quantum kicked rotor under the additional influence of a Stark or gravity field [8,17]. In Sec. II, the Hamiltonian of one kicked atom is quickly reviewed, and the fidelity, which is the main quantity studied here, is precisely defined for our system. We then report on a subtle dependence of the fidelity on the arithmetic properties of the relevant parameters (Sec. III), which is a result that may be interesting for future precise measurements of fundamental constants (see the discussions in Refs. [21,23,26]). Section IV

is devoted to the dynamics close to quantum resonance. Based on the pseudo- or ϵ -classical formalism developed by Fishman *et al.* [8], we explain the overall behavior of the fidelity using ϵ -classical phase-space densities and quantum tunneling rates from the stable resonance island to the surrounding chaotic sea in phase space. Some technical details are found in the Appendices.

II. THE KICKED ROTOR WITH GRAVITY

We are interested in the quantum dynamics of a particle moving in a line, periodically kicked in time, and subject to a constant Stark or gravity field. It is described by the Hamiltonian in dimensionless variables (such that $\hbar = 1$) [8]:

$$\hat{H}(\hat{x}, \hat{p}, \tau) = \frac{\hat{p}^2}{2} - \frac{\eta}{T} \hat{x} + k \cos(\hat{x}) \sum_{t \in \mathbb{Z}} \delta(\tau - tT). \quad (2)$$

The kicking period is T , the kicking strength is k , and t is a discrete time variable that counts the number of kicks. The parameter η yields the change in momentum produced by the constant field in one kicking period. In the accelerated frame of reference [8], the potential experienced by the particle is periodic in space and so the quasimomentum β is conserved by the evolution. With the chosen units, β takes all values between 0 and 1. Using Bloch theory, the particle dynamics can then be identified with that of a family of quantum rotors, labeled by the values of β . For the β rotor (i.e., the rotor in the family to which a given value β of the quasimomentum is affixed), the evolution from immediately after the $(t-1)$ th kick to immediately after the t th kick is described by the unitary propagator [8],

$$\hat{U}_{\beta,k,\eta}(t) = e^{-ik \cos(\hat{\theta})} e^{-iT/2(\hat{N} + \beta + \eta t + \eta/2)^2}, \quad (3)$$

and the evolution operator over the first t kicks is

$$\hat{U}_{\beta,k,\eta}^t \equiv \hat{U}_{\beta,k,\eta}(t-1) \hat{U}_{\beta,k,\eta}(t-2) \cdots \hat{U}_{\beta,k,\eta}(1) \hat{U}_{\beta,k,\eta}(0), \quad (4)$$

where $\hat{\mathcal{N}}$ is the momentum operator,

$$\hat{\mathcal{N}} = -i \frac{d}{d\theta},$$

with periodic boundary conditions. The time-dependent Hamiltonian that generates the quantum evolution corresponding to (4) is then

$$\hat{\mathcal{H}}(\hat{\mathcal{N}}, \hat{\theta}, \beta, \tau) = \frac{1}{2} \left(\hat{\mathcal{N}} + \beta + \frac{\eta}{T} \tau \right)^2 + k \cos(\hat{\theta}) \sum_{t \in \mathbb{Z}} \delta(\tau - tT). \quad (5)$$

We will study the fidelity that measures the stability of the evolution (4) with respect to changes of the parameter k . For a given β rotor, this fidelity is defined by

$$F_\beta(k_1, k_2, \eta, t) = \left| \langle \hat{\mathcal{U}}_{\beta, k_1, \eta}^t \psi | \hat{\mathcal{U}}_{\beta, k_2, \eta}^t \psi \rangle \right|^2. \quad (6)$$

Moreover, having in mind experimental situations with cold atoms [9,11,18,23,24], we will also consider the case when the initial state of the atomic cloud is an incoherent mixture of plane waves with a distribution $\rho(\beta)$ of the quasimomentum. In this case, the fidelity is given by [12]

$$F(k_1, k_2, \eta, t) = \left| \int_0^1 \rho(\beta) \langle \hat{\mathcal{U}}_{\beta, k_1, \eta}^t \psi | \hat{\mathcal{U}}_{\beta, k_2, \eta}^t \psi \rangle d\beta \right|^2. \quad (7)$$

III. FIDELITY AT A QUANTUM RESONANCE

In the gravity free case ($\eta = 0$), Eq. (3) describes the standard kicked-rotor (KR) dynamics, and so-called KR resonances [7,13,27] occur whenever T is commensurate to 2π . Then, for special values of quasimomentum β , the energy of the β rotor asymptotically increases quadratically as $t \rightarrow \infty$. In the presence of gravity, the asymptotic quadratic growth of energy at certain values of β is still possible [28]. Here we study the behavior of fidelity in the presence of gravity and for the case of a main KR resonance, i.e., $T = 2\pi l$ (with integer l). Denoting $\psi_\beta(t) \equiv \hat{\mathcal{U}}_{\beta, k, \eta}^t \psi_\beta(0)$, one may explicitly compute [8]

$$\begin{aligned} \langle \theta | \psi_\beta(t) \rangle &= e^{-i\alpha(\beta, \eta, t)} e^{-ikA(\theta, \beta, \eta, t)} \\ &\times \langle \theta - t(2\beta + 1)\pi l - \pi l \eta t^2 | \psi_\beta(0) \rangle, \end{aligned} \quad (8)$$

where $\alpha(\beta, \eta, t)$ is a global phase, and

$$A(\theta, \beta, t) = \sum_{r=0}^{t-1} \cos[\theta - (2\beta + 1)\pi l r - 2\pi l r \eta t + \pi l \eta r^2]. \quad (9)$$

From now on, we assume that the initial state is a plane wave, $|\psi_\beta(0)\rangle = |n_0\rangle$, i.e.,

$$\langle \theta | \psi_\beta(0) \rangle = \langle \theta | n_0 \rangle = \frac{e^{in_0\theta}}{\sqrt{2\pi}}. \quad (10)$$

The fidelity is directly obtained from (8) using the method described in Ref. [12]. First, write $A(\theta, \beta, t) = \text{Re}(e^{i\theta} W_t)$, with

$$W_t \equiv W_t(\eta, \beta) = \sum_{r=0}^{t-1} e^{-i\pi l(2\beta+1)r} e^{-i2\pi l r \eta t + i\pi l \eta r^2}. \quad (11)$$

Then, from (8), it follows that the wave function after the t th kick is given in momentum representation by

$$\begin{aligned} \langle n | \hat{\mathcal{U}}_{\beta, k_1, \eta}^t | \psi \rangle &= e^{ni \arg(W_t)} \frac{1}{2\pi} \int_0^{2\pi} e^{i(n_0-n)\theta} e^{-ik|W_t| \cos \theta} d\theta, \\ &= e^{ni \arg(W_t)} (-i)^{n_0-n} J_{n_0-n}(k|W_t|), \end{aligned} \quad (12)$$

where $J_{n_0-n}(\cdot)$ is the Bessel function of the order of $n_0 - n$. Using this along with the addition formula of Bessel functions (see, e.g., 7.15.(31) in Ref. [29]), one finds

$$\langle \hat{\mathcal{U}}_{\beta, k_1, \eta}^t \psi | \hat{\mathcal{U}}_{\beta, k_2, \eta}^t \psi \rangle = J_0(\Delta k |W_t|), \quad (13)$$

where we defined the perturbation parameter $\Delta k = k_2 - k_1$, and so the fidelity of a single β rotor is given by

$$F_\beta(k_1, k_2, \eta, t) = |J_0(\Delta k |W_t|)|^2. \quad (14)$$

A. Asymptotics of the fidelity for one single rotor

From Eqs. (12) and (14), it is clear that the long-time asymptotics of the wave-packet propagation, and of the fidelity as well, are determined by the behavior of $|W_t|$ as $t \rightarrow \infty$. One may write

$$W_t = e^{i\Phi(\eta, \beta, t)} \mathfrak{W}(\beta, \eta, t), \quad (15)$$

where $\Phi(\eta, \beta, t) = -\pi l(2\beta + 1 + \eta t)t$, and $\mathfrak{W}(\eta, \beta, t)$ is a quadratic Weyl sum:

$$\mathfrak{W}(\eta, \beta, t) = \sum_{r=1}^t e^{i\pi l(2\beta+1)r} e^{i\pi l \eta r^2}. \quad (16)$$

The asymptotic behavior of such sums as $t \rightarrow \infty$ is known to depend on the arithmetic nature of the number η , i.e., on whether it is rational or irrational, and, in the latter case, on its Diophantine properties [30]. First of all, as the behavior of Weyl sums may be quite erratic, we resort to the time-averaged fidelity defined by

$$\langle F_\beta(k_1, k_2, \eta, t) \rangle_N \equiv \frac{1}{N} \sum_{t=0}^{N-1} F_\beta(k_1, k_2, \eta, t), \quad N \gg 1, \quad (17)$$

which has a smoother dependence on time than the original fidelity.

The easiest case is when η is rational: $\eta = p/q$, with p and q mutually prime integers. In that case, setting $r = 2jq + \nu$ in the sum (16) with j as a nonnegative integer and $0 \leq \nu \leq 2q - 1$, the sum may be rewritten in the form

$$\mathfrak{W}(\eta, \beta, t) = C(\eta, \beta, t) B(\eta, \beta, t),$$

where

$$\begin{aligned} C(\eta, \beta, t) &= \sum_{j=0}^{[t, 2q]} e^{-4i\pi l \beta j q}, \\ B(\eta, \beta, t) &= \sum_{\nu=0}^{[t, 2q]} e^{i\pi l(2\beta+1)\nu} e^{i\pi l \nu^2 p/q}, \end{aligned} \quad (18)$$

having denoted by $[t, 2q]$ the integer part of $t/(2q)$ and $\{t, 2q\} = t \bmod(2q)$. The factor $B(\eta, \beta, t)$ is a periodic function of t with period $2q$. The explicit calculation of the sum on the right-hand side of the first equation shows that $C(\eta, \beta, t)$ is

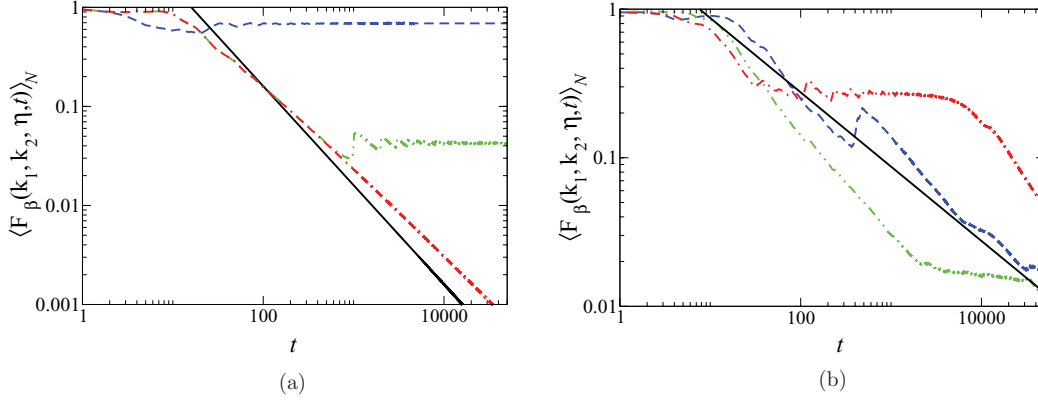


FIG. 1. (Color online) Fidelity for the kicked rotor with gravity at the quantum resonance $T = 2\pi$ for $k_1 = 0.8\pi$ and $k_2 = 0.7\pi$. The initial state is a plane wave with zero momentum, $n_0 = 0$. Several choices of η and β are shown: (a) $\eta = 0.1$. Solid black line: const/t . Blue dashed line: $\beta = 0.23$. Green double dot-dashed line: $\beta = 0.499$. Red dot-dashed line: $\beta = 0.5$. (b) Solid black line: const/\sqrt{t} . Blue dashed line: $\eta = (\sqrt{5} - 1)/2$ (the golden ratio), $\beta = 0.23$. Green double dot-dashed line: $\eta = \pi$, $\beta = 0.23$. Red dot-dashed line: $\eta = \pi$, $\beta = 0.5$.

- (i) quasiperiodic for β irrational,
- (ii) periodic for β rational and $2\beta q$ noninteger, and
- (iii) linear, i.e., $C(\eta, \beta, t) = t$, when $2\beta q$ is an integer.

Such facts have the following implications on wave-packet dynamics, on the one hand, and on the behavior of the fidelity, on the other hand. Two cases have to be distinguished, according to whether or not $2\beta q$ is an integer. In the former case, a quantum resonance occurs. Indeed, using that $|W_t|$ is equal to t times a periodic function of t , from Eq. (12) and from the well-known asymptotics of the Bessel functions at large argument and fixed order (see, e.g., 7.13.1 (3) in [29]),

$$J_n(x) \sim \sqrt{\frac{2}{\pi x}} \cos\left(x - \frac{n\pi}{2} - \frac{\pi}{4}\right), \quad x \rightarrow \infty, \quad (19)$$

we find that the probability in the n th momentum eigenstate decays in time like $1/t$. Hence, the wave packet spreads linearly in time in momentum space, and the energy quadratically increases. Instead, if $2\beta q$ is not an integer, then Eq. (12) shows that the amplitude of the evolving wave function in any momentum eigenstate oscillates quasiperiodically in time, so no unbounded spreading in momentum space occurs.

A similar reasoning based on Eqs. (14) and (17) shows that for all nonresonant values of β , the time-averaged fidelity saturates to a nonzero value in the limit $N \rightarrow \infty$ [Fig. 1(a)]. Instead, at resonance ($2\beta q = \text{integer}$) from Eqs. (14), (19) (with $n = 0$), and (17), the fidelity is seen to asymptotically decay to zero like $\ln(N)/N$ [case $\beta = 0.5$ in Fig. 1(a)].

The case of irrational η is much more difficult, as the behavior of Gauss sums crucially depends on the Diophantine properties of η [30]. Here we limit ourselves to a heuristic analysis. For strongly irrational η , one may naively picture W_t as a sort of random walk in the complex plane, suggesting that $|W_t|$ grows like \sqrt{t} in some average sense. Thanks to (14) and (19), an asymptotic decay of the average fidelity like $1/\sqrt{t}$ is then expected. This is roughly numerically confirmed for the case when η is equal to the golden ratio in Fig. 1(b). However, the actual decay displays strong fluctuations because it depends on the continued fraction expansion of η ; notably large partial quotients in the latter may cause fidelity to behave

as in cases of rational η over significant time scales, e.g., for $\eta = \pi = [3, 7, 15, 1, 292, 1, 1, \dots]$ in Fig. 1(b).

In order to mimic the experimental setups based on cold atoms [9,11,12,18,23,24], we consider the case when the initial state of the kicked atoms is an incoherent mixture of plane waves with a uniform density of β , i.e., $\rho(\beta) = 1$. The expression (7) is computed as an average over a large number of randomly chosen values of β . The result does not vary significantly when the number of values of β exceeds a few thousand. We observe a sharp difference in the asymptotic regime depending on whether or not η is rational; see Fig. 2. We again show the time-averaged quantity:

$$\langle F(k_1, k_2, \eta, t) \rangle_N \equiv \frac{1}{N} \sum_{t=0}^{N-1} F(k_1, k_2, \eta, t). \quad (20)$$

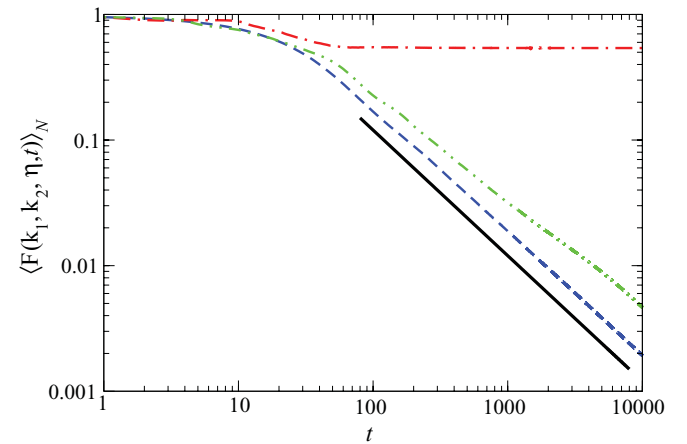


FIG. 2. (Color online) Fidelity for an ensemble of kicked rotors with gravity at $T = 2\pi$ for $k_1 = 0.8\pi$ and $k_2 = 0.7\pi$. The initial state is a plane wave with $n_0 = 0$. Several choices of η are shown. The integral over β in Eq. (7) is computed via a Riemannian sum with 5000 randomly chosen β 's. Solid black line: const/t . Blue dashed line: $\eta = \text{golden ratio}$. Green double dot-dashed line: $\eta = \pi$. Red dot-dashed line: $\eta = 0.1$.

For rational values of η , the fidelity is observed to saturate toward a finite value. This is not surprising because this is precisely the expected behavior for all values of β in a set of full measure. On the contrary, the fidelity decays like $1/N$ for irrational η ; see Fig. 2. This is roughly explained noting that besides the decaying prefactor $\sim |W_t|^{-1/2}$ in the asymptotics (19) (with $n = 0$), one more mechanism of decay is introduced by ensemble averaging, which affects the rapidly oscillating part of the Bessel function. The observation of the fine details of the number of theoretical properties of η is certainly of high interest for precision measurements [21,23,25,26], but, at the same time, presents a big challenge to experimental resolution in current setups.

IV. FIDELITY CLOSE TO A QUANTUM RESONANCE

A. Reminder of the ϵ -semiclassics and fixed points

When the kicking period is close to a quantum resonance,

$$T = 2\pi l + \epsilon, \quad l = \text{integer}, \quad |\epsilon| \ll 1, \quad (21)$$

we will implement a technique of quasiclassical approximation originally described in Ref. [8] and therein termed “ ϵ -classical approximation.” Introducing a rescaled kick strength $\tilde{k} = |\epsilon|k$, and a new momentum operator

$$\hat{I} = |\epsilon|\hat{\mathcal{N}} = -i|\epsilon|\frac{d}{d\theta}, \quad (22)$$

the propagator (3) may be rewritten as

$$\hat{U}_{\beta,k,\eta}(t) = \exp\left[-\frac{i}{|\epsilon|}\tilde{k}\cos(\hat{\theta})\right] \exp\left[-\frac{i}{|\epsilon|}\hat{\mathcal{H}}_{\beta}(\hat{I},t)\right], \quad (23)$$

where

$$\hat{\mathcal{H}}_{\beta}(\hat{I},t) = \frac{1}{2}\text{sgn}(\epsilon)\hat{I}^2 + \hat{I}\left[\pi l + T\left(\beta + \eta t + \frac{\eta}{2}\right)\right]. \quad (24)$$

The small parameter ϵ plays the formal role of a Planck constant in Eq. (23), so, close to a quantum resonance, the quantum dynamics mirrors an ϵ -classical dynamics, which is immediately inferred from (23) and (24). After changing the ϵ -classical momentum variable from I to J given by

$$J = I + \text{sgn}(\epsilon)[\pi l + T\beta + T\eta t + T\eta/2], \quad (25)$$

the ϵ -classical dynamics is described by the following map that relates the variables J and θ from immediately after the t th kick to immediately after the $(t+1)$ th one:

$$\begin{aligned} J_{t+1} &= J_t + \tilde{k}\sin(\theta_{t+1}) + \text{sgn}(\epsilon)T\eta, \\ \theta_{t+1} &= \theta_t + \text{sgn}(\epsilon)J_t \bmod 2\pi. \end{aligned} \quad (26)$$

If considered on the 2-torus, this map (26) has a fixed point at $J = 0$ and $\theta = \theta_0$ if

$$\sin\theta_0 = -\text{sgn}(\epsilon)\frac{T\eta}{\tilde{k}}, \quad (27)$$

and this fixed point is stable if and only if

$$0 \leq \tilde{k}|\cos\theta_0| \leq 4 \quad \text{and} \quad \cos\theta_0 = -\text{sgn}(\epsilon)|\cos\theta_0|. \quad (28)$$

Such stable fixed points give rise to stable islands immersed in a chaotic sea. From (25), it is immediately seen that in physical momentum space, such islands travel at constant velocity,

$$v = -\frac{T\eta}{\epsilon}. \quad (29)$$

As islands trap some of the particle’s wave packet, they give rise to experimentally observable quantum accelerator modes [17,20]. Because of such modes, a quadratic growth of energy is observed over significant time scales, both in the falling frame and in the laboratory frame [8]. However, such quadratic growth eventually comes to an end as the modes decay due to tunneling out of the stable islands [16]. Smaller accelerating islands may also exist, associated with higher-order fixed points of map (26) [14,15]; however, we will restrict ourselves to the ones described above.

B. Long-time asymptotics of the fidelity

Typical numerical results illustrating the time dependence of fidelity (6) are shown in Figs. 3–6. In all of those simulations, the parameters were chosen in ranges where significant, experimentally detectable accelerator modes exist (see Appendix A for details). In general, a very short initial transient is observed (typically up to one or a few hundred kicks, depending on parameters), marked by a very quick drop. A clear, relatively long exponential decay follows. This is sometimes followed by yet another stage of exponential decay, at a slower rate and with stronger fluctuations. This general behavior is qualitatively understood as follows. The initial sharp decay is due to the part of the initial wave packet that lies in the chaotic component of either of the two dynamics (defined by the two different kick strengths), and is rapidly carried away. The fidelity is thereafter dominated by the parts of the wave packet which are trapped inside the islands. For

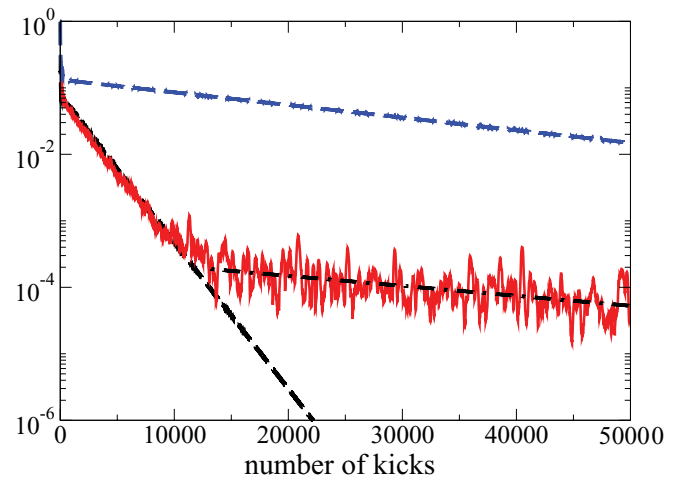


FIG. 3. (Color online) Comparison between the survival probability (black and blue dashed lines) and the fidelity (red solid line). $T = 5.86$, $\epsilon = T - 2\pi$, $\eta = 0.01579T$, $\beta = 0.48984326$, $k_2 = 0.8\pi$, and $k_1 = 0.7\pi$. The upper blue dashed line corresponds to k_1 and gives the tunneling rates $\Gamma_1 = 5.1 \times 10^{-4}$, whereas the lower black dashed line corresponds to k_2 and gives $\Gamma_2 = 4.4 \times 10^{-5}$. The dot-dashed line corresponds to an exponential fit with a decay rate $\Gamma = 3.5 \times 10^{-5}$.

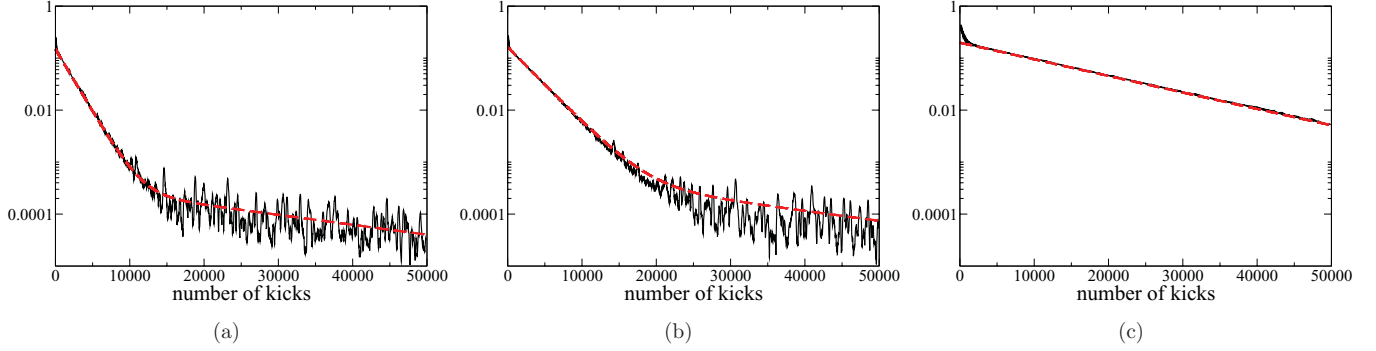


FIG. 4. (Color online) Comparison between the ansatz Eq. (30) (red dashed line) and the smoothed fidelity (black solid line) at $T = 5.86$, $\epsilon = T - 2\pi$, $\eta = 0.01579T$, $\beta = 0.48984326$, $k_2 = 0.8\pi$, and (a) $k_1 = 0.7\pi$, (b) $k_1 = 0.72\pi$, (c) $k_1 = 0.83\pi$.

this reason, in order to bring this stage of fidelity decay into full light, we choose our initial state in the form of a Gaussian state mostly located inside one island. The islands which correspond to the k_1 and to the k_2 dynamics are slightly displaced with respect to each other; however, they travel in momentum space with the same velocity (29). Therefore, the mismatch between the k_1 and the k_2 dynamics is, in ϵ -classical terms, mostly produced by (i) different structures inside the islands, and (ii) the decay from the islands into the chaotic sea due to dynamical tunneling. Concerning (i), the different rotational frequencies in the two islands are expected to produce quasiperiodic oscillations of the fidelity [10,31], so (ii) should be the main mechanism responsible for the mean fidelity decay. This leads to the following crude description. The main contribution to fidelity comes from the part of each factor in the scalar product in Eq. (6) which is trapped in the respective traveling island. Hence, the decay of fidelity is determined by the decay of each part, which is in turn determined by its respective tunneling rate into the chaotic sea, which we will denote by Γ_i for $i = 1, 2$. Then, a simple, self-explanatory ansatz for the asymptotic decay of fidelity is

$$F(t) \propto \mu(\mathcal{A}_1 \setminus \mathcal{A}_2) e^{-\Gamma_1 t} + \mu(\mathcal{A}_2 \setminus \mathcal{A}_1) e^{-\Gamma_2 t} + \mu(\mathcal{A}_2 \cap \mathcal{A}_1) e^{-(\Gamma_1 + \Gamma_2) t}, \quad (30)$$

where μ is the classical invariant measure of phase-space sets, and \mathcal{A}_i is the island around the fixed point associated with k_i .

This ansatz was found to satisfactorily reproduce the actual decay of fidelity in our numerical checks. In our simulations, the phase-space areas appearing in Eq. (30) were estimated, as described in Appendix B. Quantum decay rates were found as follows: for both the k_1 and the k_2 dynamics, we numerically computed the probability in a momentum range centered on the accelerator mode. This measures the amount of the initial probability, which travels within the accelerator mode. This quantity is called the survival probability and is shown in Fig. 3. Fitting the long-time decay of this probability with an exponential function gives us an estimate of the tunneling rate; see an example in Fig. 3. In our numerical simulations, we take the following momentum range:

$$[n(t) - 15, n(t) + 15], \quad n(t) = n_0 + vt, \quad (31)$$

where n_0 and v are given, respectively, by (A3) and (29). We checked that the peak traveling ballistically has a width less than 30 (in two photonic recoil units for the experiment; see, e.g., [11]) with our choice of parameters.

We note that the right-hand side in Eq. (30) is defined up to a proportionality factor. Moreover, it crucially depends on k_1 and k_2 , because island sizes and tunneling rates vary when the kicking strength is changed. In our simulations, the proportionality factor was chosen such as to fit the earlier regime of exponential decay (approximately between 100 and 10^4 kicks in Fig. 3, for instance). In our numerical computations, the initial state is given by (A2) with a width $\sigma^2 = 0.25$.

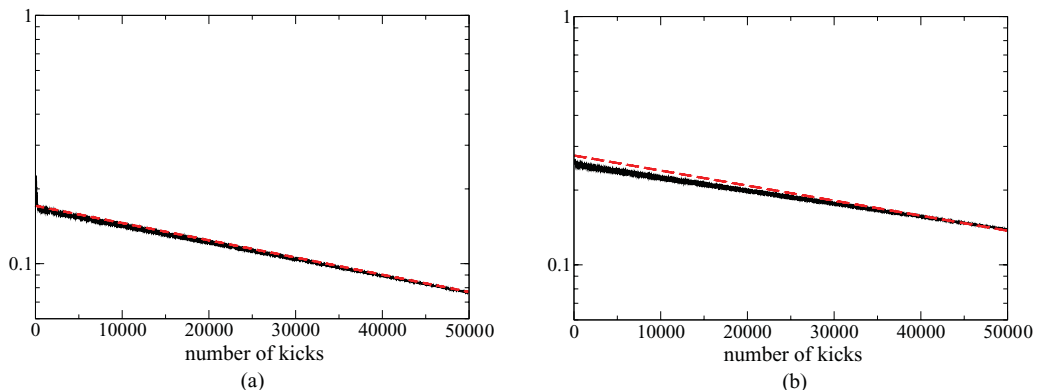


FIG. 5. (Color online) Same as in Fig. 4 at $T = 6.6$, $\epsilon = T - 2\pi$, $\eta = (\sqrt{5} - 1)/20$, $\beta = 0.123456789$, $k_2 = 2.5 + T\eta$, and (a) $k_1 = 2 + T\eta$, (b) $k_1 = 3.5 + T\eta$.

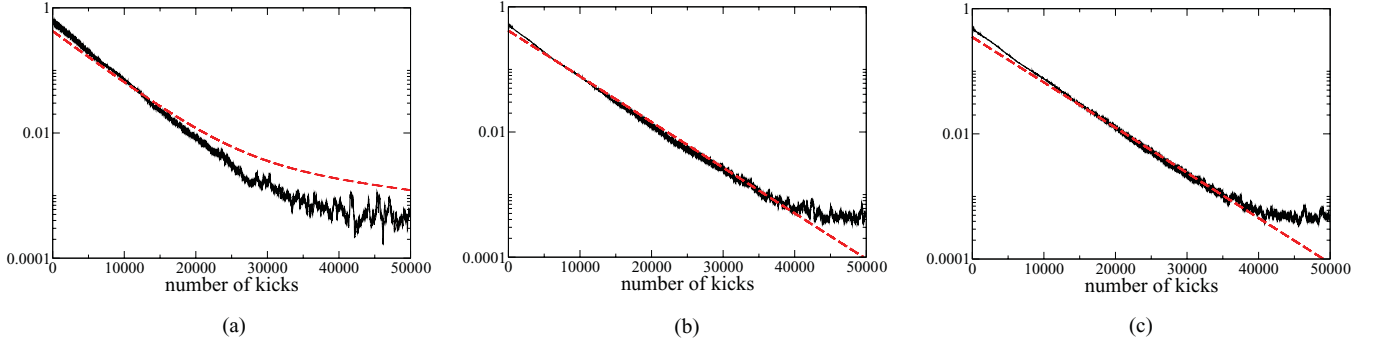


FIG. 6. (Color online) Same as Fig. 4 at $\epsilon = -0.5$, $T = 4\pi + \epsilon$, $\eta = 0.001$, $\beta = 0.123456789$, $k_2 = (1.35 + T\eta)/|\epsilon|$, and (a) $k_1 = (1 + T\eta)/|\epsilon|$, (b) $k_1 = (2 + T\eta)/|\epsilon|$, (c) $k_1 = (2.2 + T\eta)/|\epsilon|$.

As the fidelity is a wildly oscillating function, it is averaged over 200 kicks in order to clearly expose the mean long-time behavior. The results are shown in Figs. 4–6 for different sets of parameters, and will be discussed in the following.

The mean behavior of the fidelity for long times is well reproduced by the ansatz (30) when Γ_1 and Γ_2 are quite different: then the fidelity shows successively two different decay regimes, which are well reproduced by the pseudoclassical ansatz (30); see Figs. 4(a) ($\Gamma_1 \simeq 10\Gamma_2$) and 4(b) ($\Gamma_1 \simeq 5\Gamma_2$). On the contrary, when Γ_1 and Γ_2 are close to each other, one can see only one decay, which is still well reproduced by (30); see Fig. 4 ($\Gamma_1 \simeq \Gamma_2$) and Fig. 5 ($\Gamma_1 \simeq 3\Gamma_2$). For the sake of comparison, we are showing the same plots for $l = 2$ in Fig. 6. The quantum resonance is then $T = 4\pi$. It can be seen that the agreement is not as good in the latter case. One reason for this may be that higher-order ϵ -classical phase-space structures have a larger area, and hence may play a more important role, making estimates of the various areas in Eq. (30) more problematic. Both ϵ -classical phase spaces corresponding to k_1 and k_2 are displayed in Fig. 7, corresponding to Fig. 4(a) and Fig. 6(c). It is clear that the shape and the overlap between the two islands are qualitatively different in these two situations.

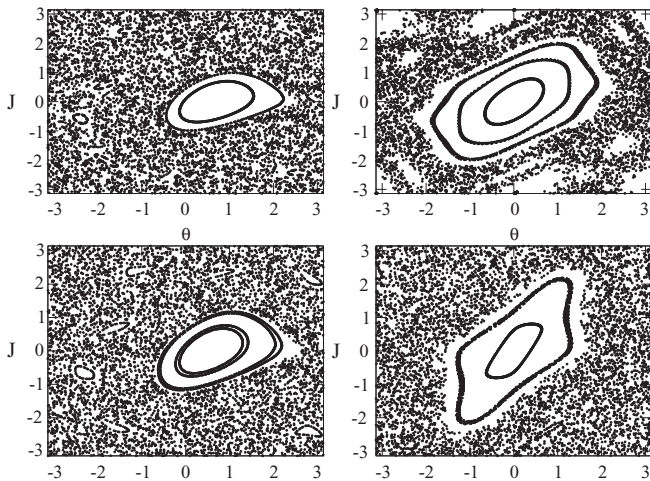


FIG. 7. Phase spaces corresponding to two sets of parameters. Left column: $T = 5.86$, $\epsilon = T - 2\pi$, $\eta = 0.01579T$. Top left: $k = 0.7\pi$. Bottom left: $k = 0.8\pi$. Right column: $\epsilon = -0.5$, $T = 4\pi + \epsilon$, $\eta = 0.001$. Top right: $k = (1.35 + T\eta)/|\epsilon|$. Bottom right: $k = (2.2 + T\eta)/|\epsilon|$.

V. CONCLUSION

We have presented a theoretical analysis of the temporal dependence of fidelity for the quantum kicked rotor subject to an additional gravity field. The two major results concern the dynamics of this system at principal quantum resonances, i.e., at kicking periods $T = 2\pi l$ (l integer), and close to these resonances. In the former case, we arrive at analytical estimates for the decay of fidelity, which is highly sensitive to the arithmetic properties of the gravity parameter. Close to a resonance, we have used the ϵ -semiclassical method in order to describe the long-time asymptotics of the fidelity. The ansatz (30) based on semiclassical densities gives a good description of the long-time behavior of the quantum fidelity for both similar and different tunneling rates of the two compared nondispersive wave packets [32] centered at the accelerator mode islands in phase space [16]. This result highlights once more the utility of ϵ classics in describing the quantum evolution of the kicked rotor and its variants.

ACKNOWLEDGMENTS

It is a pleasure to thank D. Ullmo, J. Marklof, P. Schlagheck, and G. S. Summy for helpful discussions. R.D. and S.W. acknowledge financial support from the DFG through Program No. FOR760, the Helmholtz Alliance Program of the Helmholtz Association (Contract No. HA-216 Extremes of Density and Temperature: Cosmic Matter in the Laboratory), and within the framework of the Excellence Initiative through the Heidelberg Graduate School of Fundamental Physics (Grant No. GSC 129/1), the Frontier Innovation Fund, and the Global Networks Mobility Measures during R.D.'s stay at the University of Heidelberg when this work was done.

APPENDIX A: RELEVANT RANGE FOR THE PARAMETERS

It was observed numerically, when plotting the ϵ -classical phase portrait, that higher-order nonlinear resonances play a bigger role when l is increased. For this reason, we restrict in this paper mainly to $l = 1$. In the experiments (see, e.g., [24]), k typically runs from $0.3\pi \sim 0.94$ to $1.5\pi \sim 4.71$. In order to see an accelerator mode, one needs a stable fixed point of the classical map. Following (27), for a given k , one has a

fixed point when $\tilde{k} \geq T\eta$. Taking into account experimental constraints leads us to choose ϵ such that

$$0.4 < |\epsilon| < 1. \quad (\text{A1})$$

For the quantum fidelity, the initial state is chosen to be a Gaussian state,

$$\langle n | \psi_\beta(0) \rangle = \frac{e^{-(n-n_0)^2/4\sigma^2 + in\theta_0}}{(2\pi\sigma^2)^{1/4}}, \quad (\text{A2})$$

with

$$\sin \theta_0 = -\frac{T\eta}{\epsilon k}, \quad n_0 = \frac{2\pi m}{|\epsilon|} - \frac{1}{\epsilon} \left[\pi l + T \left(\beta + \frac{\eta}{2} \right) \right], \quad (\text{A3})$$

where m is an arbitrary integer. This initial state is centered on an accelerator mode for which $J = 0$. This mode is traveling (in momentum space) at the speed given by (29); see, e.g., [8]. The fidelity is computed by applying successively the operators (3) for two different values of the kicking strength: k_1 and k_2 . For each set of parameters, we keep k_2 fixed and vary k_1 . Our initial state is always chosen such as to follow the accelerator mode attached to k_1 : once T (and ϵ), η , and β are fixed, the initial state is centered in momentum space around n_0 defined by (A3).

APPENDIX B: ESTIMATING THE SIZE OF THE ISLANDS IN THE ϵ -CLASSICAL PHASE SPACE

First, we choose a set of parameters T and η for which we can have accelerator modes. Then, we vary k within the range of existence of these modes. One way to see this range is to compute numerically the area of the stable island in the ϵ -classical phase space; see Fig. 8. The area follows a bell shape as a function of k . In Fig. 8, some jumps are also visible (see, e.g., the blue dash-dotted line, between $k \simeq 4.424$ and $k \simeq 4.624$). We believe that this is due to the lack of precision when determining the island boundary and/or the breaking of the outermost tori and their remnants (cantori). The size of the stable island in the ϵ -classical phase space is computed by starting a fairly small number of trajectories outside the island. These are typically iterated

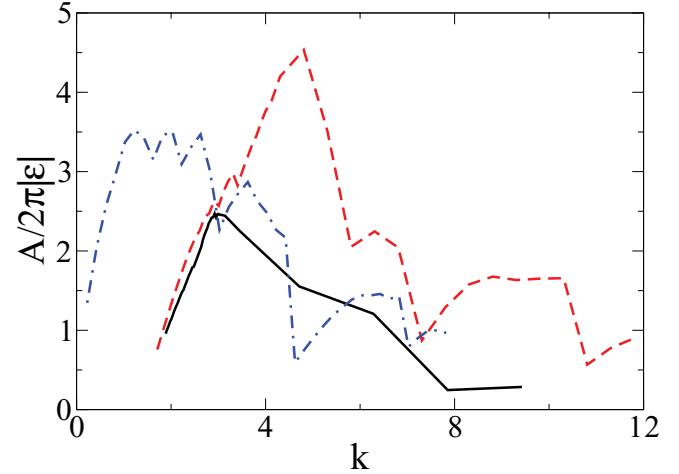


FIG. 8. (Color online) Area of the stable island in the ϵ -classical phase space as a function of the kicking strength for three sets of parameters. Black solid line: $T = 5.86$, $\epsilon = T - 2\pi \sim -0.42$, $\eta = 0.01579T$. Red dashed line: $T = 6.6$, $\epsilon = T - 2\pi \sim 0.32$, $\eta = \phi/10$. Blue dash-dotted line: $\epsilon = -0.5$, $T = 4\pi + \epsilon$, $\eta = 0.001$.

for a long time (10^8 kicks). Then, we move to polar coordinates (φ, I) centered at the fixed point under interest [33]. The boundary of the island is then determined by the curve $I(\varphi)$ defined in the following way. Using a grid of thickness $\delta\varphi$ along the φ axis, the boundary $I_i = I(\varphi_i)$ is given by

$$I_i = \min\{I_j, (\varphi_j, I_j) \text{ iterated points and } \varphi_i \leq \varphi_j \leq \varphi_i + \delta\varphi\}. \quad (\text{B1})$$

The size of the island is given by the area under this curve. Numerically, it is computed via a Riemannian sum.

The measures of the different sets in Eq. (30) are computed by propagating a cloud of 10^4 classical points. The initial points are distributed following normal distributions with mean and width $(\theta_0, \sigma_\theta)$ in the θ direction and (J_0, σ_J) in the J direction. The areas in Eq. (30) reached stationary values after approximately 500 kicks. The measures needed in Eq. (30) are simply given by the number of points sitting on one or both of the stable islands associated to k_1 and k_2 , respectively.

- [1] A. Peres, *Phys. Rev. A* **30**, 1610 (1984).
 [2] R. A. Jalabert and H. M. Pastawski, *Phys. Rev. Lett.* **86**, 2490 (2001).
 [3] T. Gorin, T. Prosen, T. H. Seligman, and M. Znidaric, *Phys. Rep.* **435**, 33 (2006); P. Jacquod and C. Petitjean, *Adv. Phys.* **58**, 67 (2009).
 [4] G. Casati *et al.*, in *Stochastic Behavior in Classical and Quantum Hamiltonian Systems*, edited by G. Casati and J. Ford (Springer, Berlin, 1979), p. 334.
 [5] B. V. Chirikov, *Phys. Rep.* **52**, 263 (1979).
 [6] F. L. Moore, J. C. Robinson, C. F. Bharucha, B. Sundaram, and M. G. Raizen, *Phys. Rev. Lett.* **75**, 4598 (1995); C. F. Bharucha,

- J. C. Robinson, F. L. Moore, B. Sundaram, Q. Niu, and M. G. Raizen, *Phys. Rev. E* **60**, 3881 (1999).
 [7] F. M. Izrailev and D. L. Shepelyanskii, *Sov. Phys. Dokl.* **24**, 996 (1979) [*Theor. Math. Phys.* **43**, 553 (1980)].
 [8] S. Fishman, I. Guarneri, and L. Rebuffini, *Phys. Rev. Lett.* **89**, 084101 (2002); *J. Stat. Phys.* **110**, 911 (2003).
 [9] S. Wimberger, I. Guarneri, and S. Fishman, *Nonlinearity* **16**, 1381 (2003); *Phys. Rev. Lett.* **92**, 084102 (2004).
 [10] M. Abb, I. Guarneri, and S. Wimberger, *Phys. Rev. E* **80**, 035206(R) (2009).
 [11] M. Sadgrove and S. Wimberger, *Adv. At. Mol. Opt. Phys.* **60**, 315 (2011).

- [12] S. Wimberger and A. Buchleitner, *J. Phys. B* **39**, L145 (2006).
- [13] I. Dana and D. L. Dorofeev, *Phys. Rev. E* **73**, 026206 (2006); I. Guarneri, *Ann. Henri Poincaré* **10**, 1097 (2009).
- [14] I. Guarneri, L. Rebuzzini, and S. Fishman, *Nonlinearity* **19**, 1141 (2006).
- [15] I. Guarneri and L. Rebuzzini, *Phys. Rev. Lett.* **100**, 234103 (2008).
- [16] M. Sheinman, S. Fishman, I. Guarneri, and L. Rebuzzini, *Phys. Rev. A* **73**, 052110 (2006).
- [17] M. K. Oberthaler, R. M. Godun, M. B. d’Arcy, G. S. Summy, and K. Burnett, *Phys. Rev. Lett.* **83**, 4447 (1999); M. B. d’Arcy, R. M. Godun, M. K. Oberthaler, G. S. Summy, K. Burnett, and S. A. Gardiner, *Phys. Rev. E* **64**, 056233 (2001).
- [18] M. Sadgrove, S. Wimberger, S. Parkins, and R. Leonhardt, *Phys. Rev. Lett.* **94**, 174103 (2005); S. Wimberger, M. Sadgrove, S. Parkins, and R. Leonhardt, *Phys. Rev. A* **71**, 053404 (2005); M. Sadgrove, S. Wimberger, S. Parkins, and R. Leonhardt, *Phys. Rev. E* **78**, 025206(R) (2008).
- [19] M. Sadgrove, M. Horikoshi, T. Sekimura, and K. Nakagawa, *Phys. Rev. Lett.* **99**, 043002 (2007); M. Sadgrove and S. Wimberger, *New J. Phys.* **11**, 083027 (2009).
- [20] G. Behinaein, V. Ramareddy, P. Ahmadi, and G. S. Summy, *Phys. Rev. Lett.* **97**, 244101 (2006).
- [21] I. Talukdar, R. Shrestha, and G. S. Summy, *Phys. Rev. Lett.* **105**, 054103 (2010).
- [22] C. Ryu, M. F. Andersen, A. Vaziri, M. B. d’Arcy, J. M. Grossman, K. Helmerson, and W. D. Phillips, *Phys. Rev. Lett.* **96**, 160403 (2006).
- [23] S. Wu, A. Tonyushkin, and M. G. Prentiss, *Phys. Rev. Lett.* **103**, 034101 (2009); A. Tonyushkin, S. Wu, and M. G. Prentiss, *Phys. Rev. A* **79**, 051402(R) (2009).
- [24] S. Schlunk, M. B. d’Arcy, S. A. Gardiner, D. Cassettari, R. M. Godun, and G. S. Summy, *Phys. Rev. Lett.* **90**, 054101 (2003).
- [25] A. Ullah and M. D. Hoogerland, *Phys. Rev. E* **83**, 046218 (2011).
- [26] R. A. Horne, R. H. Leonard, and C. A. Sackett, *Phys. Rev. A* **83**, 063613 (2011).
- [27] S.-J. Chang and K.-J. Shi, *Phys. Rev. A* **34**, 7 (1986).
- [28] I. Dana and V. Roitberg, *Phys. Rev. E* **76**, 015201(R) (2007).
- [29] A. Erdelyi *et al.*, *Higher Transcendental Functions*, Vol. 2 (McGraw-Hill, New York, 1955).
- [30] G. H. Hardy and J. E. Littlewood, *Acta Math.* **37**, 193 (1914); J. H. Hannay and M. V. Berry, *Physica D* **1**, 267 (1980); M. V. Berry and J. Goldberg, *Nonlinearity* **1**, 1 (1988); J. Marklof, *Duke Math. J.* **97**, 127 (1999).
- [31] Y. Krivolapov, S. Fishman, E. Ott, and T. M. Antonsen, *Phys. Rev. E* **83**, 016204 (2011).
- [32] A. Buchleitner, D. Delande, and J. Zakrzewski, *Phys. Rep.* **368**, 409 (2002); S. Wimberger, P. Schlagheck, Ch. Eltschka, and A. Buchleitner, *Phys. Rev. Lett.* **97**, 043001 (2006).
- [33] P. Schlagheck (private communication).