

OPTICAL IDENTIFICATION OF 664 OHIO SOURCES USING ACCURATE RADIO AND OPTICAL POSITIONS MEASURED BY THE TEXAS INTERFEROMETERS

FRANK D. GHIGO

The University of Texas at Austin

Received 1977 April 4; accepted 1977 May 23

ABSTRACT

Results of optical identification work are reported for 664 radio sources selected from the Ohio 1415 MHz survey. Radio positions were measured at 365 MHz to an accuracy of about 2" with the Texas broad-band synthesis interferometer. Forty previously unpublished radio positions are given. Optical positions of 0".5 to 0".7 accuracy were measured on the Palomar Sky Survey (PSS) plates with a two-coordinate laser-interferometer measuring machine. These errors are consistent with those derived by comparing the positions with other accurate optical positions. Background objects were counted on the PSS plates and were used along with counts of objects within 20" of the radio sources to estimate the radial distribution of the radio-optical offsets of the true identifications. This analysis showed that about 30% of the identifications have radio-optical separations in excess of those expected from the position errors. This effect appears to be primarily due to resolution of large-angular-size sources by the interferometer system. The completeness and reliability of the identifications are discussed. A total of 319 identifications are suggested. Galaxies and probable galaxies with $m_v \leq 19.5$ number 102 (15%). Quasars, probable quasars, and BL Lacertae-type objects with $m_v \leq 19.5$ number 136 (20%). One hundred eighteen new or revised identifications are found. Thirteen identification candidates, found to have stellar spectra, are probably random coincidences. Two probable BL Lacertae-type objects (2207+020 and 2217+018) appear to have steep radio spectra.

Subject headings: galaxies: general — quasars — radio sources: general

I. INTRODUCTION

The Ohio radio source survey (Rinsland, Dixon, and Kraus 1975, and references therein) covers virtually the whole area between declinations -36° and $+63^\circ$, and contains almost 20,000 sources, making it by far the most extensive survey at 1415 MHz or any higher frequency. Clearly, a completely identified Ohio survey would be of great value. Unfortunately, its position uncertainties (1' to 10') are too great to permit reliable identifications, and its flux densities, for sources fainter than 1 Jy,¹ are too uncertain ($\geq 50\%$; Jauncey and Niell 1971; Bridle *et al.* 1972) for statistically useful samples to be defined. The present work is a step toward alleviating the first problem by measuring positions accurately. The sample presented here is a representative selection of Ohio sources in three areas of sky covering 0.842 steradians. These data will be suitable for statistical studies of quasars and radio galaxies when accurate 1415 MHz flux densities become available for the Ohio sources in these regions.

The selection effects in this sample are discussed in § II. The Texas radio observations are summarized in § III. In § IV the Texas and Ohio positions are compared.

The optical measurements are described in § V, in which the table of optical data is also found. Comments on individual objects are in § VI, the

Texas optical positions are compared with other accurate optical positions. The results of counting background objects on the PSS plates are given in § VII. The effect of resolved sources and the completeness and reliability of the sample are discussed in § VIII.

The identification content of the sample is summarized in § IX. A discussion of the neutral stellar object (NSO) identifications and the possibility of radio stars in the sample is given. Among the NSOs, two probable BL Lacertae types appear to have steep radio spectra, although these objects more typically have flat spectra.

II. SELECTION EFFECTS IN THE SAMPLE

The regions of sky observed are the three areas listed in Table 1. Objects in these regions were scheduled for observing in accordance with the requirements of the broad-band synthesis interferometer (BSI) of the University of Texas Radio Astronomy Observatory (UTRAO), which takes data on the meridian. The instrument is discussed by Douglas *et al.* (1973). A 3 minute observation of each source was made, so no more than about 200 sources could be observed in a single night. A complete observing pass is five nights of observation, to allow scans to be made on all eight interferometer baselines. Five passes were made through each of the three regions. Eighty percent of the Ohio sources could be scheduled in five

¹ 1 Jansky (Jy) = 10^{-26} W m⁻² Hz⁻¹.

GHIGO

 TABLE 1
 SUMMARY OF RADIO OBSERVATIONS

	REGION		
	1	2	3
Right Ascension.....	8 ^h to 15 ^h	12 ^h to 18 ^h 5	20 ^h 5 to 1 ^h 5
Declination.....	23° to 33°	10° to 20°	-2° to 10°
Dates of Observations.....	1971 March-April	1971 May	1971 Aug.-Sept.
Area (steradians).....	0.282	0.287	0.273

passes. The remaining 20% were missed because they were in areas where more than five sources occurred in a 3 minute interval of right ascension.

Incompleteness also occurs because the Ohio survey is confusion-limited. Stull (1973) has found that about 30% of Ohio sources are blends of two or more sources. About one-third of the blended sources which could be scheduled on the BSI were observed and good positions measured. The remainder were missed because no source could be seen within several arc minutes of the listed Ohio catalog position.

The loss of sources from the final sample because of the scheduling procedure and the effect of Ohio catalog confusion means that sources are preferentially lost from crowded regions. This should not seriously bias the final sample since there appears to be no evidence of radio source clustering (e.g., Webster 1976).

The present sample is also biased by the selection in spectral index (by observing a 1415 MHz survey at 365 MHz) and in angular size (since the BSI fringe size is about 1', whereas the Ohio beam is about 10' × 40'). The number of sources that would be missed from a complete sample at 1415 MHz was estimated from the spectral index distribution found for the 1400 MHz NRAO survey (Bridle *et al.* 1972), and from the distribution of angular sizes in the 3CR catalog (Macdonald, Kenderdine, and Neville 1968; Elsmore and Mackay 1969; Mackay 1969). These estimates show that about 5% of the Ohio catalog was lost due to the spectral index distribution, and about 13% were lost because they had angular sizes $\geq 1'$.

One thousand two hundred forty-seven Ohio sources occur in the selected regions with $S_{1415} \geq 0.3$ Jy. The estimates of combined losses due to scheduling, blending of sources, angular size being too large, and radio spectrum being too flat appear to be adequate to account for our having found only 532 sources having $\gamma S_{365} \geq 0.5$ Jy and $S_{1415} \geq 0.3$ Jy in the present sample [γ is the BSI fringe visibility. The BSI fringe separation is 52" sec (30:1 - δ) in declination, and 56"7 in right ascension].

A number of sources with flux densities below the aforementioned limits are included in the data to be presented, bringing the total number of sources listed in this paper to 664.

III. RADIO OBSERVATIONS

The BSI and its data reduction are described by Douglas *et al.* (1973). The radio positions of most

sources in this sample are given by Douglas *et al.* (1973), Ghigo and Owen (1973), and Sharp and Bash (1975). Twenty-one new positions, 19 corrections of published lobe-shifted positions, and 177 positions improved by averaging positions from two or more BSI observing programs are presented in Table 2. The median radio position errors for the sample of 664 sources are 2".2 in right ascension and 1".5 in declination. The true position errors will be somewhat in excess of the listed errors for resolved sources, which are indicated by having *R* or *RR* appended to the lobe-shift class, which appears in column (8) of Table 2. For a position (α , δ), the allowed lobe-shift positions (α^* , δ^*) are given by

$$(\alpha^*, \delta^*) = (\alpha + k_\alpha l_\alpha, \delta + k_\delta l_\delta), \quad (1)$$

where

$$l_\alpha = 3^s782 \text{ sec } \delta$$

$$l_\delta = 52^s22 \text{ sec } (30:1 - \delta)$$

and

$$(k_\alpha + k_\delta) = \text{an even integer.}$$

The key to Table 2 is as follows:

Column (1): Source name in Parkes notation.

Column (2): Notes regarding the position: \$ means the position is the mean of two or more positions from different BSI observing programs; *L* means that a lobe-shift position was previously published but is corrected here; *N* means that no BSI position has previously been published for this source.

Column (3): The 1950.0 right ascension (h, m, s) and its error (s).

Column (4): The 1950.0 declination (degrees, minutes, seconds) and its error (arcsec).

Column (5): *N* and *M* are the number of independent observations in right ascension and declinations, respectively. *N* = 3 and *M* = 5 for a complete, single-pass observation (all eight baselines).

Column (6): Galactic coordinates, l^{II} and b^{II} (degrees).

Column (7): Measured flux density (γS_{365} , where S_{365} is the source's total flux density) and its error, in Jy, at the observed centroid.

Column (8): Lobe-shift and resolution class (these are more fully described in Ghigo and Owen 1973): class 1—probability of lobe shift less than 1%; classes

TABLE 2
NEW AND REVISED RADIO POSITIONS FROM THE TEXAS INTERFEROMETER

1	2	3	4	5	6	7	8											
SOURCE	NOTE	RIGHT ASCENSION (1950.0)	DECLINATION (1950.0)	N	M	L2	B2	FLUX DENSITY	CLASS	OHIO	4C	OTHER NAMES						
0003-003	\$	00 03 48.81	.05 -00 21 08.1	.7	5	8	99 -61	11.68	.70	1	B-007	-00.01	3C2	P	G	N		
0010+005	\$	00 10 36.98	.05 00 35 0 .0	.7	11	19	103 -60	3.55	.25	1R	B+017	+00.01	3C5	P	G	N		
0016+093	\$	00 16 20.58	.07 09 23 27.9	.8	15	26	110 -52	1.56	.10	1	B+027							
0018+052	\$	00 18 22.07	.15 05 16 43.3	1.5	4	7	109 -56	1.67	.12	2BR	B+030	+05.04						
0019-000	\$	00 19 51.62	.06 -00 01 43.4	.8	5	9	107 -62	2.90	.20	1	B-032							
0025+006	\$	00 25 59.48	.14 00 38 35.1	1.7	5	8	111 -61	1.08	.12	2AR	B+044							
0029+013	\$	00 29 34.05	.10 01 20 22.4	1.0	4	8	113 -61	2.24	.17	2A	B+050			P				AO
0033+079	\$	00 33 41.21	.21 07 58 35.7	1.7	5	8	116 -54	.88	.12	2BR	B+056							
0034+077	\$	00 34 01.35	.18 07 43 35.3	2.5	4	7	116 -55	1.21	.17	2AR	B+056							
0034-014	\$	00 34 30.56	.05 -01 25 38.5	.7	6	10	115 -64	9.62	.59	1R	B-057	-01.03	3C15	P	G	N		
0038+097	L	00 38 14.50	.13 09 46 36.8	1.6	1	2	119 -53	5.29	.59	2AR	B+063.6	+09.02	3C18	P	G	N		
0043+000	\$	00 43 08.01	.11 00 04 45.1	1.2	7	11	120 -62	1.66	.14	2AR	B-072.4	-00.05		P				AO
0059+056	\$	00 59 36.48	.11 05 39 09.5	1.3	5	9	128 -57	1.76	.14	2A	B+099.5	+05.07		P				
0106+013	\$	01 06 04.54	.06 01 19 0 .6	.8	6	10	132 -61	3.30	.22	1	C+012	+01.02		P	G			
0109+025	\$	01 09 13.03	.16 02 35 11.5	1.6	5	9	133 -60	.91	.10	2ARR	C+015	+02.03		P				
0109+026	\$	01 09 42.46	.11 02 41 45.8	1.2	6	10	133 -59	1.44	.12	2A	C+015	+02.03		P				
0115-016	\$	01 15 42.22	.07 -01 36 16.8	.9	5	8	138 -63	4.39	.28	2A	C-025	-01.07		P				
0118-001	\$	01 18 21.24	.13 -00 10 48.6	1.4	6	9	139 -62	1.44	.13	2AR	C-031	-00.08		P				
0122-005	\$	01 22 43.62	.08 -00 34 07.8	.9	9	16	141 -62	1.79	.13	2AR	C-038	-00.10		P	G			
0801+303	\$	08 01 34.86	.06 30 21 10.7	.6	6	11	191 28	3.52	.21	2AR	J+302	+30.13	B2					
0802+243	\$	08 02 35.28	.08 24 18 32.2	.7	5	11	198 26	2.98	.20	2AR	J+204	+24.16	B2	3C192.0	P	G	N	
0814+294	\$	08 14 0 .92	.10 29 27 32.9	.9	6	12	193 30	1.28	.08	2R	J+223	+29.28	B2					
0815+238	\$	08 15 06.14	.08 23 49 35.3	.7	26	43	199 29	1.25	.10	1	J+227	+23.19		P				
0817+307	\$	08 17 26.03	.07 30 44 19.9	.6	10	21	192 31	1.52	.08	2A	J+329	+30.15	B2					
0824+294	\$	08 24 21.40	.05 29 28 41.4	.4	8	19	194 33	5.73	.27	1R	J+240	+29.29	B2	3C200.0	G	N		
0828+324	N	08 28 14.96	.29 32 29 26.9	2.2	4	7	191 34	.55	.08	2BR	J+348		B2					G
0828+325	N	08 28 28.32	.35 32 30 18.7	1.9	4	7	191 34	.51	.08	2BR	J+348	+32.25	B2					G
0831+241	\$	08 31 40.39	.15 24 08 52.8	1.1	13	23	200 33	.74	.08	2A	J+253		B2					AO
0835+256	\$	08 35 52.87	.06 25 37 50.5	.5	18	34	199 34	2.73	.14	2A	J+260	+25.22	B2					P
0838+325	\$	08 38 06.82	.12 32 35 43.7	1.2	5	8	191 36	1.23	.12	2B	J+363	+32.26	B2					
0840+299	\$	08 40 09.28	.12 29 55 17.0	.9	5	13	194 36	1.43	.10	2RR	J+267	+29.31	B2					
0845+298	\$	08 45 28.03	.16 29 49 28.4	1.9	13	20	195 37	.54	.06	2BR	J+275		B2					
0853+291	\$	08 53 0 .47	.05 29 10 06.5	.5	34	57	196 38	2.26	.13	2AR	J+288	+29.32	B2					N
0855+280	\$	08 55 10.83	.06 28 02 29.2	.5	7	16	198 39	5.91	.35	1R	J+292	+28.21	B2	3C210.0	G	N		
0858+292	\$	08 58 05.25	.05 29 13 34.5	.5	18	31	196 40	4.10	.20	1R	J+297	+29.33	B2	3C213.1	G	N		
0903+258	\$	09 03 18.74	.13 25 49 37.0	1.3	9	14	201 40	1.11	.10	2ARR	K+207	+25.23	B2					P
0905+231	L	09 05 06.67	.17 23 06 08.3	1.9	3	4	205 40	1.43	.17	2B	K+210							P
0912+297	N	09 12 53.47	.36 29 45 54.4	2.7	2	4	197 43	.61	.12	2B	K+222		B2					
0914+257	N	09 14 16.09	.24 25 44 54.6	2.1	3	5	202 42	.78	.12	2B	K+224		B2					
0932+241A	\$	09 32 31.49	.16 24 08 37.3	1.8	26	43	206 46	.40	.06	2BR	K+255		B2					
0932+254	\$	09 32 40.08	.09 25 24 11.7	.8	7	11	204 46	1.66	.12	2AR	K+254	+25.26	B2					P
0932+241B	\$	09 32 56.18	.19 24 10 21.5	2.2	26	43	206 46	.40	.05	2BR	K+255		B2					
0934+255	\$	09 34 17.30	.18 25 34 23.3	1.1	12	22	204 47	.70	.08	2AR	K+259	+25.27	B2					P
0939+266	N	09 39 39.00	.17 26 40 47.8	1.6	3	5	203 48	1.57	.16	2B	K+267	+26.29	B2					P
0945+247	\$	09 45 55.73	.08 24 46 42.0	.7	8	18	206 49	1.59	.10	2B	K+277	+24.21	B2					
0949+246	\$	09 49 10.22	.08 24 36 38.2	.7	6	12	206 50	3.58	.24	2R	K+285	+24.22	B2					P
0953+254	\$	09 53 59.37	.24 25 29 29.8	1.9	8	12	205 51	.62	.10	1	K+290		B2					G
0955+326	\$	09 55 25.43	.05 32 38 23.9	.4	11	21	194 52	3.47	.18	1	K+393	+32.33	B2	3C232	P	G	N	
0958+256	\$	09 58 35.82	.10 25 40 57.5	.9	5	9	205 52	1.70	.12	2A	K+299	+25.30	B2					P
0958+290	\$	09 58 54.83	.08 29 01 21.5	.6	5	12	200 53	10.09	.59	3R	L+200	+29.35	B2	3C234.0	G	N		
1003+290	\$	10 03 54.15	.09 29 03 56.9	.8	5	9	200 54	1.65	.12	2AR	L+207	+29.36	B2					
1015+277	\$	10 15 0 .69	.08 27 47 08.7	.7	4	7	203 56	3.45	.24	2A	L+227	+27.21	B2					G
1015+294	\$	10 15 36.53	.15 29 29 24.6	1.4	5	9	200 56	.85	.10	2B	L+228	+29.38	B2					
1016+329	\$	10 16 26.65	.10 32 56 53.4	.9	6	10	194 57	1.48	.12	2AR	L+327	+32.35	B2					
1026+256	N	10 26 58.46	.35 25 38 50.9	4.3	3	5	208 58	.33	.10	2BR	L+245		B2					
1028+281	\$	10 28 10.37	.07 28 11 23.5	.6	14	30	203 59	1.70	.10	2AR	L+247	+28.26	B2					
1032+265	\$	10 32 30.67	.10 26 30 51.3	1.0	6	10	207 60	1.26	.10	2AR	L+254		B2					
1036+323	\$	10 36 02.56	.09 32 21 43.7	.7	5	9	195 61	2.22	.14	2	L+360	+32.36	B2					
1044+297	\$	10 44 39.89	.09 29 44 13.8	.7	5	10	201 63	1.93	.13	2	L+274	+29.39	B2					
1055+314	\$	10 55 05.24	.10 31 29 58.2	.6	12	22	197 65	1.63	.05	2BR	L+391		B2					
1057+307	\$	10 57 22.69	.08 30 42 46.0	.9	6	10	198 65	2.16	.14	2BR	L+396	+30.20	B2					N
1118+237	\$	11 18 04.11	.05 23 44 21.8	.4	11	21	218 69	5.91	.32	1	M+230	+23.27	B2	3C256.0	P	G	N	
1132+303	\$	11 32 16.28	.06 30 22 01.7	.6	6	11	198 73	4.34	.26	1	M+356	+30.22	B2	3C261	P	G	N	
1134+265	\$	11 34 52.26	.13 26 35 49.2	1.4	12	23	211 73	.73	.08	2BR	M+258	+26.33	B2					P
1146+255	\$	11 46 55.16	.10 25 31 12.3	.9	13	19	217 76	1.33	.10	2B	M+278		B2					
1151+295	\$	11 51 38.21	.06 29 32 50.0	.5	4	10	200 77	4.81	.29	2AR	M+286	+29.44	B2					
1153+317	\$	11 53 44.13	.04 31 44 46.7	.3	19	39	190 77	7.17	.33	1R	M+389	+31.38	B2					G
1156+295	\$	11 56 57.82	.06 29 31 26.9	.6	7	13	199 78	2.80	.17	1R	M+295	+29.45	B2					

TABLE 2—Continued

1	2	3	4	5	6	7	8								
SOURCE	NOTE	RIGHT ASCENSION (1950.0)	DECLINATION (1950.0)	N	M	L2	B2	FLUX DENSITY	CLASS	OHIO	4C	OTHER NAMES			
1221+186	\$	12 21 14.46	.12	18 37 44.4	1.2	7 11	265 79	1.14	.10	2A	N+135	+18.34			
1221+164	\$	12 21 20.18	.09	16 24 45.4	.9	8 13	271 77	1.44	.10	2A	N+136	+16.32	3C271	P	N
1231+248	\$	12 31 56.13	.09	24 48 15.4	1.0	6 9	246 85	1.59	.12	2AR	N+253	+24.27		B2	
1234+252	\$	12 34 44.91	.12	25 14 32.0	1.4	5 7	246 86	1.39	.12	2A	N+257	+25.41		B2	
1235+196	L	12 35 07.03	.22	19 41 08.9	1.8	3 5	280 82	1.00	.14	2A	N+158				
1241+166	\$	12 41 27.52	.04	16 39 16.5	.5	15 22	293 79	9.42	.55	1R	N+169	+16.34	3C275.1	P	G
1241+246	N	12 41 49.11	.25	24 39 48.2	2.6	2 3	272 87	.93	.16	2BR	N+271			B2	
1250+291	\$	12 50 11.28	.13	29 08 05.1	1.0	5 8	114 88	1.24	.10	2A	N+283			B2	
1250+119	L	12 50 21.47	.23	11 55 38.6	1.7	3 5	304 74	1.58	.20	2AR	N+184				P
1251+159	\$	12 51 03.03	.05	15 58 40.8	.6	6 10	305 78	5.17	.32	2R	N+185	+15.40	3C277.2	P	N
1252+119	\$	12 52 07.75	.11	11 57 22.0	1.0	12 20	306 74	.97	.08	1	N+187				P
1256+281	L	12 56 57.32	.23	28 10 56.8	2.2	3 5	57 88	.68	.12	2BR	N+294			B2	
1259+150	\$	12 59 09.93	.11	15 02 20.3	2.0	6 10	314 77	.95	.10	2BRR	N+199	+15.42			
1303+192	\$	13 03 19.88	.10	19 17 42.6	.9	6 10	326 81	1.70	.12	2ARR	P+105	+19.43			
1308+120	\$	13 08 23.68	.08	12 05 48.5	.8	6 10	320 74	2.26	.16	2A	P+116	+12.47			
1317+179	\$	13 17 54.85	.06	17 58 53.8	.6	6 11	340 78	3.97	.25	2ARR	P+129.8	+17.56			P
1318+113	\$	13 18 49.72	.07	11 22 30.5	.8	6 10	328 72	6.09	.40	1	P+131	+11.45			P
1319+270	\$	13 19 52.94	.06	27 01 26.9	.6	5 9	32 83	3.29	.20	1	P+233	+27.25	B2		P
1319+230	L	13 19 57.14	.16	23 00 53.6	1.2	2 4	3 82	1.77	.20	2BR	P+235	+23.35			P
1321+181	\$	13 21 11.26	.09	18 10 04.1	1.1	9 14	343 78	1.45	.10	2AR	P+135	+18.37			
1323+321	\$	13 23 57.96	.04	32 09 43.2	.4	19 33	67 81	7.61	.41	1	P+340			B2	
1325+126	\$	13 25 26.52	.14	12 38 41.3	1.7	6 9	335 73	.97	.10	2BR	P+144	+12.48			
1326+150	\$	13 26 08.09	.08	15 03 23.1	.8	6 10	340 75	2.18	.16	2A	P+145	+15.43			
1328+254	\$	13 28 15.94	.03	25 24 37.7	.3	19 32	22 81	15.42	.62	1	P+247	+25.43	B2	3C287.0	P
1328+141	L	13 28 49.08	.12	14 08 36.7	1.5	6 10	340 74	1.03	.10	2BR	P+148	+14.48			
1328+307	\$	13 28 49.66	.04	30 45 59.0	.4	14 28	56 81	28.55	1.55	1	P+348	+30.26	B2	3C286.0	G
1331+189	\$	13 31 39.70	.13	18 58 56.6	1.2	7 11	354 77	1.08	.10	3	P+152				
1334+288	\$	13 34 13.45	.07	28 49 39.4	.7	9 14	44 80	1.88	.10	2A	P+257	+28.32	B2		
1335+170	L	13 35 33.56	.13	17 03 46.1	1.0	13 20	351 75	.96	.08	2A	P+160				
1338+288	\$	13 38 22.44	.11	28 52 48.0	1.1	6 10	44 79	1.11	.10	2A	P+263.9	+28.33	B2		
1338+107	\$	13 38 35.68	.07	10 47 16.1	.8	6 10	340 70	2.84	.18	2A	P+164	+10.36			
1339+122	\$	13 39 30.25	.17	12 13 46.7	1.2	6 11	343 71	.86	.10	2BR	P+167	+12.49			
1341+144	\$	13 41 57.48	.06	14 24 17.5	.6	6 10	349 72	4.68	.28	2	P+169	+14.49			P
1345+125	\$	13 45 06.22	.04	12 32 20.5	.4	12 20	347 70	8.35	.41	1	P+175	+12.50			P
1348+161	\$	13 48 33.76	.10	16 06 48.0	1.5	6 10	357 72	1.25	.10	2RR	P+182	+16.37			
1350+113	\$	13 50 28.71	.10	11 21 54.1	1.4	4 6	348 68	3.37	.25	2B	P+185	+11.46			
1352+164	\$	13 52 16.36	.07	16 29 27.5	1.2	6 7	359 72	3.30	.21	2RR	P+187	+16.38	3C293.1	P	N
1354+195	\$	13 54 41.95	.04	19 33 47.4	.5	17 27	9 73	3.72	.18	1R	P+191	+19.44			P
1357+173	L	13 57 02.34	.25	17 21 33.2	2.3	3 5	4 71	.82	.16	2AR	P+195.3				
1358+244	N	13 58 49.84	.14	24 29 07.7	1.6	3 4	26 74	1.36	.16	2B	P+298	+24.29	B2		P
1359+103	\$	13 59 45.33	.18	10 20 42.7	1.3	6 10	351 66	.91	.10	2BR	P+199	+10.37			
1400+162	\$	14 00 20.64	.09	16 14 23.3	.9	6 10	3 70	1.82	.13	2	Q+100	+16.39			
1401+123	\$	14 01 06.30	.10	12 20 29.9	1.2	7 11	355 67	1.34	.10	2R	Q+101				
1407+178	\$	14 07 53.46	.08	17 48 09.7	.9	5 9	9 69	3.29	.10	2BR	Q+112	+17.57			P
1408+141	\$	14 08 34.06	.09	14 10 56.7	.8	6 10	1 67	2.01	.13	2AR	Q+114	+14.51			
1409+133	\$	14 09 13.45	.14	13 18 25.6	1.5	6 10	0 66	.99	.10	2	Q+116	+13.51			
1417+144	\$	14 17 18.39	.13	14 26 54.3	1.1	5 8	5 66	1.43	.13	2A	Q+128				
1417+272	\$	14 17 44.68	.06	27 15 26.6	.6	6 10	38 70	3.35	.20	2A	Q+229	+27.28	B2		G
1420+198	\$	14 20 41.03	.05	19 48 54.8	.6	6 9	18 68	9.24	.54	1R	Q+134	+19.46	3C300.0	P	G
1421+122	\$	14 21 04.68	.06	12 13 26.4	.8	11 16	2 64	2.01	.12	1	Q+135				
1421+191	\$	14 21 49.46	.08	19 06 26.3	.8	6 10	17 67	2.72	.18	2	Q+136	+19.47			
1425+287	\$	14 25 28.52	.06	28 46 49.9	.6	8 13	42 69	3.39	.21	2B	Q+243	+28.35	B2		
1426+131	\$	14 26 20.05	.09	13 09 58.3	.9	6 10	6 63	1.90	.14	2R	Q+145	+13.52			
1431+163	N	14 31 06.56	.20	16 20 41.3	1.7	3 5	13 64	.97	.13	2B	Q+153	+16.40			
1433+177	\$	14 33 36.17	.07	17 42 34.9	.7	6 10	17 64	2.71	.18	1	Q+155	+17.59			G
1435+248	\$	14 35 34.39	.11	24 52 03.5	1.1	6 9	33 66	1.25	.10	2A	Q+261	+24.32	B2		
1436+115	N	14 36 51.63	.34	11 35 47.7	2.8	3 5	6 60	.56	.13	2B	Q+162				
1439+111	L	14 39 02.89	.12	11 08 12.9	1.4	6 10	6 59	1.34	.14	2AR	Q+165				
1439+137	\$	14 39 40.73	.13	13 42 02.6	1.5	5 10	11 61	1.02	.10	3	Q+166	+13.53			P
1443+232	N	14 43 29.74	.16	23 15 13.5	1.5	6 10	30 64	1.12	.12	2B	Q+272				
1444+180	\$	14 44 49.07	.06	18 00 23.7	.6	6 11	20 62	3.82	.24	2	Q+175	+17.61			P
1445+168	N	14 45 24.34	.19	16 48 40.5	2.0	3 5	18 61	.75	.13	2AR	Q+177	+16.42			P
1445+167	\$	14 45 48.07	.16	16 46 58.9	2.3	3 5	18 61	1.20	.12	2BR	Q+177	+16.42			P
1449+144	\$	14 49 16.45	.15	14 26 19.6	1.4	7 12	14 59	1.16	.08	2BRR	Q+181.9	+14.55			
1454+183	\$	14 54 21.44	.13	18 20 20.7	1.1	4 8	22 60	1.75	.16	2BR	Q+191	+18.39			
1501+127	L	15 01 28.34	.19	12 42 06.9	2.1	3 5	14 56	.99	.14	2BR	R+102				
1502+106	L	15 02 0 16	.10	10 41 18.8	1.2	7 11	11 54	1.38	.12	2A	R+103				G
1504+163	\$	15 04 49.75	.10	16 18 46.0	.9	6 10	21 57	1.80	.13	2	R+108	+16.44			
1505+123	\$	15 05 18.52	.13	12 19 32.3	1.2	4 8	14 55	1.68	.13	2B	R+110	+12.54	3C312		
1509+102	\$	15 09 04.62	.09	10 12 59.7	1.0	6 10	12 53	1.84	.13	2A	R+117	+10.40			
1509+158	\$	15 09 52.34	.06	15 51 38.0	.7	6 10	21 55	3.04	.20	1	R+118	+15.45			P
1512+104	N	15 12 59.54	.21	10 29 38.3	2.3	3 5	13 52	.85	.14	2A	R+122				
1524+101	\$	15 24 21.70	.12	10 09 37.2	1.4	5 7	15 50	1.54	.14	2A	R+140.5	+10.43			
1527+158	\$	15 27 27.28	.10	15 49 59.5	1.4	5 7	24 51	1.63	.13	2A	R+146	+15.48			
1530+137	\$	15 30 54.22	.07	13 42 27.4	.7	5 9	21 50	3.38	.21	1	R+152	+13.55			
1532+139	\$	15 32 02.11	.11	13 59 16.9	1.2	6 10	22 50	1.72	.14	1	R+1				

OPTICAL IDENTIFICATION OF OHIO SOURCES

TABLE 2—Continued

1	2	3		4		5		6		7		8			
SOURCE	NOTE	RIGHT ASCENSION (1950.0)		DECLINATION (1950.0)		N	M	L2	B2	FLUX	DENSITY	CLASS	OHIO	4C	OTHER NAMES
1544+118	L	15 44 48.57	.25	11 52 53.6	2.9	3	5	21	46	.80	.14	2BR	R+174		
1551+179	S	15 51 03.94	.13	17 58 34.2	1.5	8	12	30	47	1.12	.10	2BR	R+185.5	+17.64	
1554+165	N	15 54 34.00	.44	16 32 28.2	3.3	3	5	29	46	.52	.13	2BR	R+191		
1602+178	S	16 02 54.21	.14	17 51 47.9	2.0	6	8	31	44	1.18	.12	2AR	S+106	+17.66	
1616+138	S	16 16 42.70	.07	13 48 42.9	.8	6	10	28	40	2.55	.17	2A	S+130	+13.58	P N
1621+132	L	16 21 11.78	.15	13 13 02.8	1.4	3	5	28	39	1.56	.18	2A	S+135		G
1641+173	S	16 41 34.69	.05	17 21 21.1	.6	12	19	35	36	9.23	.60	1	S+170	+17.70	3C346.0 P G N
1645+174	S	16 45 28.03	.05	17 25 27.7	.5	24	40	36	35	3.77	.24	1	S+176	+17.71	P G N
1650+111	N	16 50 41.11	.52	11 10 46.7	3.0	2	4	29	31	.53	.16	2BR	S+185		
1653+163	L	16 53 29.36	.31	16 18 40.7	2.1	3	5	35	33	.78	.14	2BR	S+188		
1707+117	N	17 07 14.75	.16	11 44 13.6	2.2	3	5	32	28	1.12	.16	2AR	T+112.1		
1722+105	S	17 22 19.90	.25	10 33 33.2	1.4	5	9	33	24	1.22	.14	2A	T+136.8		
1725+167	S	17 25 21.67	.14	16 46 54.2	1.7	5	7	39	26	1.42	.16	2AR	T+143	+16.48	
1727+174	S	17 27 53.98	.09	17 26 37.3	1.1	6	10	40	25	1.98	.16	2BR	T+146	+17.76	
1732+160	S	17 32 28.08	.05	16 02 27.3	.7	6	10	39	24	5.54	.35	1	T+154	+16.49	P
1739+107	S	17 39 20.11	.12	10 45 40.5	1.6	9	14	35	20	1.15	.12	2BR	T+165.7	+10.53	
1740+162	L	17 40 36.46	.15	16 14 08.5	1.4	2	4	40	22	1.72	.16	2BR	T+168	+16.52	
1743+173	L	17 43 22.57	.24	17 21 10.3	2.3	3	5	42	22	.92	.16	2A	T+172		G
1744+114	N	17 44 36.61	.19	11 28 22.1	2.0	3	5	36	19	1.08	.16	2AR	T+175		
1751+153	L	17 51 59.45	.22	15 21 06.6	3.2	6	10	41	19	.49	.10	2BR	T+186		
1754+159	N	17 54 19.31	.19	15 54 06.5	1.9	3	5	41	19	1.02	.16	2A	T+190		G
1755+153	N	17 55 53.56	.23	15 18 12.0	1.9	5	9	41	18	.74	.12	2A	T+193		
1758+111	S	17 58 08.82	.18	11 06 12.8	1.2	8	13	37	16	1.53	.14	2A	T+197	+10.54	
1759+138	S	17 59 21.67	.06	13 51 22.8	.7	6	10	40	17	7.34	.45	1	T+198	+13.66	P
1810+170	S	18 10 45.43	.13	17 05 13.8	2.0	7	11	44	16	1.03	.12	2A	U+119	+17.78	
1822+127	S	18 22 42.73	.10	12 46 39.3	1.0	5	9	41	11	1.96	.16	2A	U+139	+12.64	
2042+045	S	20 42 15.70	.11	04 33 09.5	1.2	5	9	51	-22	1.65	.14	2A	W+071	+04.71	
2045+068	S	20 45 44.35	.04	06 50 10.8	.5	17	30	54	-22	7.42	.41	1	W+076	+06.68	3C424.0 P G N
2055+055	S	20 55 59.65	.09	05 31 12.1	.9	4	7	54	-25	2.56	.21	2AR	W+093	+05.78	
2108+039	S	21 08 40.32	.09	03 58 42.8	1.1	5	8	55	-28	2.68	.20	2A	X+014	+03.50	P
2115+056	S	21 15 11.06	.17	05 39 35.8	1.6	5	8	57	-29	1.14	.13	2AR	X+027	+05.79	
2120+044	S	21 20 50.79	.12	04 27 03.4	1.4	5	8	57	-30	1.62	.16	2A	X+035	+04.74	
2128+048	S	21 28 02.55	.05	04 49 04.1	.6	17	28	59	-32	3.43	.20	1	X+046		P G
2128+089	S	21 28 54.24	.06	08 59 19.4	.7	8	13	63	-29	3.23	.22	1	X+049	+08.62	P G
2134+004	S	21 34 05.06	.09	00 28 25.8	1.2	6	10	55	-35	1.95	.16	1R	X+057		P G
2142+077	S	21 42 39.42	.12	07 46 31.5	1.2	5	9	64	-33	1.46	.13	2A	X+071	+07.56	
2142+042	S	21 42 46.99	.07	04 17 40.7	.8	6	11	61	-35	3.40	.24	2A	X+072	+04.75	P G
2145+067	S	21 45 36.11	.04	06 43 40.6	.4	22	37	64	-34	3.95	.18	1	X+076.1	+06.69	P G
2150+053	S	21 50 54.34	.07	05 22 08.9	.8	5	9	63	-36	3.80	.25	2A	X+085	+05.81	P
2152+017	S	21 52 0.95	.31	01 44 06.7	2.4	4	7	60	-38	1.06	.12	2AR	X+089		
2203+056	S	22 03 50.25	.08	05 41 32.4	1.0	6	10	66	-38	2.51	.18	2A	Y+007	+05.82	P
2207+020	S	22 07 0.21	.06	02 03 57.4	.8	7	12	63	-41	3.08	.20	2A	Y+011	+02.54	P
2209+080	S	22 09 32.04	.06	08 04 25.6	.6	6	11	70	-37	4.71	.30	1	Y+015.8	+08.64	P G
2226+089	S	22 26 38.34	.05	08 59 0.2	.6	10	16	74	-40	4.04	.21	1	Y+045	+08.67	P
2241+013	S	22 41 16.83	.09	01 20 40.7	.9	7	13	71	-48	1.75	.13	2A	Y+069	+01.72	
2308+017	S	23 08 17.55	.12	01 45 07.2	1.2	5	8	79	-52	1.59	.14	2A	Z+014.2	+01.73	
2309+090	S	23 09 56.52	.05	09 03 08.9	.6	8	14	86	-46	8.31	.41	1	Z+016	+09.73	3C456.0 P G N
2314+038	S	23 14 02.31	.04	03 48 54.9	.6	11	18	83	-51	17.97	1.18	1	Z+024	+03.57	3C459.0 P G N
2320+079	S	23 20 03.95	.07	07 55 33.4	.8	17	28	88	-49	1.50	.10	1	Z+033		P G
2330+015	S	23 30 43.27	.11	01 33 52.2	1.2	5	8	87	-55	1.60	.14	2A	Z+051.2		
2344+092	S	23 44 03.90	.07	09 14 05.2	.7	8	13	97	-50	2.51	.16	1	Z+073.5	+09.74	P G
2345+048	S	23 45 58.45	.10	04 50 53.6	1.2	6	11	95	-54	1.47	.12	2AR	Z+077	+04.83	
2349+014	S	23 49 22.39	.05	-01 25 57.5	.7	8	14	92	-60	4.11	.26	1R	Z+082	-01.61	P G
2350+057	S	23 50 21.11	.06	05 43 51.1	.7	5	9	98	-54	3.56	.22	2A	Z+084	+05.86	P

2, 2A, 2B—lobe-shift probability $\sim 5\%$; class 3—lobe-shift probability about 20%. An “R” appended to these class numbers means that there is evidence of resolution. “RR” indicates severe resolution. The lobe-shift rate appears to be about 50% for classes 3R and 3RR.

Column (9): Other names. B2 = Bologna, P = Parkes, G = Green Bank, AO = Arecibo, N = NRAO.

IV. TEXAS-OHIO POSITION COMPARISON

The distribution of position differences between the Ohio positions and the BSI positions is shown in Figure 1 for three flux density ranges. This gives an indication only of the Ohio position errors, since the BSI position errors are much smaller. The extended tails of the distributions are caused primarily by the presence of those blend components which were seen

by the BSI as separate sources. If one ignores the tails of the distributions, the position differences are consistent with the Ohio position errors as given by Brundage *et al.* (1971), namely $\sigma_\alpha \approx 10''$ and $\sigma_\delta \approx 8'$ for $S_{1415} = 0.2$ Jy, $\sigma_\alpha \approx 7''$ and $\sigma_\delta \approx 5'$ for $S_{1415} = 0.4$ Jy, and $\sigma_\alpha \approx 4''$ and $\sigma_\delta \approx 3'$ for $S_{1415} = 1.0$ Jy.

V. OPTICAL DATA

Accurate positions were determined on the *National Geographic Society-Palomar Observatory Sky Survey* (PSS) plates for all objects within 25" of each radio source in the sample, and also for a few (typically three) stars within 4' of each radio source for use as secondary position standards.

a) Machine Measurements

Measurements were made with a two-coordinate laser-interferometer measuring machine, built for the

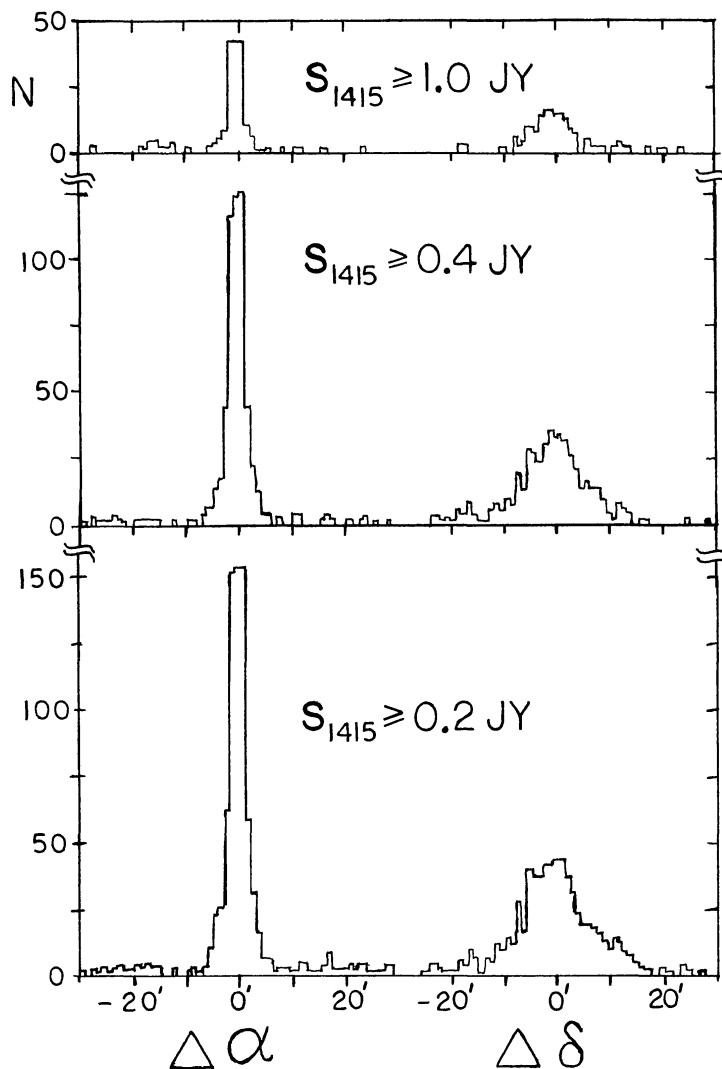


FIG. 1.—Ohio-Texas position differences

University of Texas Radio Astronomy Observatory by Opto Mechanisms, Inc. Position standards were from the Smithsonian Astrophysical Observatory (SAO) catalog. A local second-order least-squares fit was done in a 2° square centered on each radio source. About 10 to 12 reference stars determined each fit. The equations of condition were

$$\alpha - \alpha_0 = G_1 + G_2x + G_3y + G_4x^2 + G_5y^2 + G_6xy, \quad (2)$$

$$\delta - \delta_0 = D_1 + D_2x + D_3y + D_4x^2 + D_5y^2 + D_6xy, \quad (3)$$

where G_i and D_i ($i = 1, 2, \dots, 6$) are parameters to be determined, (α, δ) are the equatorial coordinates, (α_0, δ_0) denote the field center, and (x, y) are rec-

tangular coordinates on the plate relative to the plate center. If an insufficient number of stars were available in the 2° square area, a linear fit was done, involving only the first three terms in the above expressions.

For each field, measurements were made on both red and blue PSS plates. Positions were determined separately on the two plates. The mean of the two positions is quoted in Table 3 for objects which occurred on both plates.

The internal standard errors of the least-squares fit were in the range $0''.2$ to $0''.4$. Estimates of the measurers' personal equation effects were also in this range. As yet undiagnosed problems with the measuring apparatus appear to contribute an additional uncertainty of $\pm 0''.3$ to $\pm 0''.4$. Thus the overall standard errors in the resulting optical positions given in Table 3 are $0''.5$ to $0''.7$. In § VI we shall see that these estimates are consistent with comparisons with other accurate optical positions.

TABLE 3—Continued

1										2																																																																																																																																																																																																																																																																																																																																																	
POSITION(1950.0)					OFFSETS					MAG TYPE IO					POSITION(1950.0)					OFFSETS					MAG TYPE IO																																																																																																																																																																																																																																																																																																																																		
h	m	s	°	'	Δ _R	Δ _D	Δ _S	Δ _B	Δ _V	h	m	s	°	'	Δ _R	Δ _D	Δ _S	Δ _B	Δ _V	h	m	s	°	'	Δ _R	Δ _D	Δ _S	Δ _B	Δ _V	h	m	s	°	'	Δ _R	Δ _D	Δ _S	Δ _B	Δ _V																																																																																																																																																																																																																																																																																																																				
01	03	20	37	06	12	15	11	4	5	17.0	N	C	(C)	PQ	B	0*	01	28	42	54	06	08	05	2	-24.0	5.7	(20.0)	(S)	01	28	42	54	06	08	05	2	-24.0	5.7	(20.0)	(S)	(1)01	03	16	50	06	10	25	3	-52.8	-104.4	(1)01	28	42	54	06	06	55	3	-23.3	-64.3	(2)01	03	22	30	06	14	09	3	33.7	119.7	(2)01	28	42	54	06	07	15	1	86.2	-40.4	(3)01	03	18	70	06	09	21	2	-19.9	-168.5	(3)01	28	44	52	06	09	39	3	6.2	99.8																																																																																																																																																																																																																																																							
01	03	59	77	08	59	19	6	-3.1	-1.9	20.0	R			PG	A	0	01	28	58	96	00	17	56	2	.5	2.4	16.0	R	D	(D)	CG	A	31	(1)01	04	00	55	09	01	15	3	8.4	113.8	(1)01	28	59	73	00	18	10	6	12.0	16.9	17.0	R	S	(S)	(2)01	03	51	99	08	59	08	7	-118.5	-12.7	(1)01	29	04	06	00	18	28	5	77.0	34.7	(2)01	04	10	21	08	56	45	0	151.5	-156.4	(2)01	28	54	42	00	16	40	7	-67.6	-73.1	(3)01	04	10	21	08	56	45	0	151.5	-156.4	(3)01	29	00	64	00	20	28	5	25.6	154.7																																																																																																																																																																																																																																						
01	06	04	52	01	19	00	6	-.4	.1	19.0	B	S	(S)	Q	A	13	08	01	34	98	30	21	09	6	1.5	-1.1	18.5	B	S	(S)	Q	A	16	(1)01	06	01	99	01	18	54	1	-38.3	-6.4	08	01	34	96	30	21	26	3	1.2	15.6	20.0	R	(2)01	06	15	30	01	18	07	1	161.2	-53.4	(1)08	01	36	75	30	20	31	5	24.4	-39.2	(3)01	05	55	55	01	20	56	7	-134.9	116.1	(2)08	01	24	88	30	19	00	4	-129.3	-130.4	(1)01	09	12	41	02	35	47	8	-9.4	36.2	(3)08	01	51	15	30	21	23	3	210.9	12.5	(2)01	09	07	97	02	36	11	4	-75.9	59.9	08	02	35	52	24	18	26	6	3.3	-5.6	17.0	N	D	(D)	CG	B	46	(3)01	09	15	55	02	33	18	0	37.7	-113.5	08	02	36	47	24	18	20	3	16.2	-11.9	15.0	N	S	(S)	(1)01	09	42	25	02	41	43	5	-3.2	-2.2	19.5	R	D	(D)	PG	A	38	(1)08	02	33	68	24	16	59	0	-21.8	-93.2	(1)01	09	42	82	02	41	58	5	5.5	12.8	13.5	N	S	(S)	08	05	01	40	26	59	54	3	-3.9	10.4	20.0	N	D	(D)	C	EF	*	(1)01	09	40	71	02	40	17	0	-26.2	-88.7	08	05	01	37	26	59	59	5	-4.4	15.5	19.0	B	D	(D)	08	05	01	37	26	59	59	5	-4.4	15.5	19.5	R	D	(1)01	11	55	94	-00	15	15	6	-.4	-5.6	19.5	R	D	08	05	01	37	26	59	59	5	-4.4	15.5	19.5	R	D	(1)01	11	55	51	-00	15	34	3	-6.9	-24.3	(1)08	04	52	83	27	00	13	8	-118.5	29.9	(2)01	11	59	45	-00	14	36	3	52.3	33.7	(2)08	05	07	52	27	01	26	1	77.9	102.2	(3)01	11	50	66	-00	15	42	4	-79.6	-32.4	(3)08	05	13	91	26	59	00	8	163.3	-43.1
01	12	43	89	-01	42	55	2	-6.1	1.3	18.0	B	S	(S)	Q	B	8	08	07	10	33	27	40	36	5	-8.9	5.7	16.5	N	S	(S)	PQ	C	0*	(1)01	12	43	-01	42	57	20	NP	(1)08	07	01	70	27	39	56	5	-123.6	-34.2	(1)01	12	40	49	-01	43	20	7	-57.0	-24.2	(2)08	07	01	23	27	40	05	4	-129.8	-25.4	(2)01	12	52	80	-01	42	15	4	127.6	41.1	(3)08	07	00	69	27	41	32	5	-137.0	61.7	(3)01	12	34	33	-01	43	11	8	-149.4	-15.3	08	08	06		30	05	2		21	SF	20.0	R	D		EF	(1)01	13	52	25	01	11	34	3	-101.8	-69.4	(1)08	08	06	07	30	05	58	0	8.5	31.7	(2)01	14	02	54	01	15	02	1	52.5	138.4	(2)08	08	04	44	30	06	16	1	-12.7	49.8	(3)01	14	01	66	01	15	40	6	39.3	176.9	(3)08	08	11	26	30	05	02	0	75.8	-24.2																																																																																																																																																																											
01	14	50	45	07	26	00	1	2.1	.4	19.0	R	D		PG	A	22*	08	09	49	40	32	52	06	4	7.2	-6.9	20.0	R		>	G	17	(1)01	14	49	47	07	26	30	6	-12.6	30.9	(1)08	09	47	42	32	53	01	6	-17.8	48.2	(2)01	14	52	57	07	25	37	4	33.6	-22.3	(2)08	09	53	81	32	53	46	3	62.7	92.9	(3)01	14	47	98	07	25	36	1	-34.6	-23.6	(3)08	09	49	50	32	50	19	9	8.5	-113.5																																																																																																																																																																																																																																																															
(1)01	15	37	15	-01	38	07	3	-76.0	-110.5	EF	*	08	10	08	60	31	17	37	2	18.1	-16.8	17.0	R	S	(S)	EF	(2)01	15	51	33	-01	35	46	6	136.6	30.2	(1)08	10	04	88	31	17	46	9	-29.6	-7.1	(3)01	15	26	39	-01	35	27	4	-237.3	49.4	(2)08	10	10	77	31	18	27	8	45.9	33.8	(3)08	10	07	76	31	16	54	0	7.3	-60.0																																																																																																																																																																																																																																																																															
01	15	43	62	02	42	19	5	1.3	-1.6	17.5	B	S	(S)	Q	A	7	08	14	00	06	29	27	45	2	-11.2	12.3	14.5	N	S	(S)	EF	(1)01	15	42	70	02	41	54	7	-12.7	-26.4	(1)08	13	58	36	29	27	45	8	-33.4	12.9	(2)01	15	46	00	02	40	28	0	36.9	-113.1	(2)08	13	54	77	29	26	09	4	-80.2	-83.5	(3)01	15	35	96	02	43	32	3	-113.6	71.2	(3)08	14	14	09	29	26	53	7	172.1	-39.2																																																																																																																																																																																																																																																																
01	16	24	21	08	14	09	2	-.9	-.6	20.0	N			PQ	A	38	08	15	06	22	23	49	36	6	1.2	1.3	19.0	B	S	(S)	PQ	A	23	(1)01	16	23	50	08	14	15	1	-11.3	5.3	08	15	05	73	23	49	31	2	-5.6	-4.1	20.0	R	(1)01	16	26	17	08	14	20	2	28.3	10.4	08	15	06		23	49	4		15	SF	20.0	(2)01	16	26	72	08	13	49	0	36.4	-20.7	(1)08	15	01	93	23	50	41	7	-57.7	66.4	(3)01	16	30	43	08	12	47	8	91.5	-82.0	(2)08	15	15	32	23	49	46	2	125.9	10.9	(1)01	16	36	96	05	16	29	1	36.8	77.7	(3)08	15	03	58	23	47	18	5	-35.1	-136.7	(2)01	16	43	84	05	15	40	0	139.7	28.6	08	17	25	89	30	44	18	7	-1.9	-1.2	19.0	R	D		G	A	16*	(3)01	16	25	12	05	14	37	2	-140.0	-34.2	08	17	25	58	30	44	18	1	-5.9	-1.9	19.0	N	D	(D)	(1)01	18	21	64	-00	10	49	2	6.0	-.6	(1)08	17	26	04	30	44	45	7	.0	25.7	(1)01	18	19	95	-00	11	12	5	-19.3	-23.9	(2)08	17	24	87	30	45	22	2	-15.1	62.2	(2)01	18	17	38	-00	12	23	7	-58.0	-95.1	(3)08	17	19	08	30	44	25	2	-89.6	5.3	(3)01	18	25	67	-00	13	28	3	66.4	-159.7																																																																																										
01	18	26	12	03	28	30	3	10.5	-4.0	19.0	B	S	(S)	Q	C	13*	08	18	44	23	29	32	18	6	-.9	-.9	19.0	R	D		PG	A	0	(1)01	18	26	15	03	28	51	2	11.0	16.8	08	18	44		29	32	2		10	SF	20.0	(1)01	18	25	96	03	26	33	7	8.1	-120.6	(1)08	18	41	61	29	32	11	6	-35.1	-7.9	(2)01	18	32	70	03	30	00	2	109.0	85.9	(2)08	18	47	07	29	32	12	0	36.1	-7.6	(3)01	18	21	14	03	30	40	8	-64.1	126.5	(3)08	18	37	77	29	31	20	1	-85.2	-59.4																																																																																																																																																																																																																																									
(1)01	18	57	60	09	41	24	3	-124.5	-38.4	EF	08	20	36	61	29	38	11	2	3.9	.2	19.0	B	S	(S)	Q	A	38	(2)01	19	04	55	09	44	46	4	-21.9	163.7	(1)08	20	37		29	38	2		14	F	(20.0)	(1)01	19	30	68	-00	06	46	2	-26.1	-5.0	(2)08	20	33	88	29	38	12	9	-31.6	1.9	(2)01	19	34	72	-00	06	33	6	34.5	7.6	(2)08	20	32	63	29	37	40	8	-48.0	-30.2	(3)01	19	31	10	-00	07	12	4	-19.8	-31.2	(3)08	20	37	24	29	39	31	6	12.2	80.6																																																																																																																																																																																																																																															
01	19	32	15	-00	33	41	0	-6.7	2.2	20.0	R			PG	B	0	08	24	21	41	29	28	40	6	.2	-.8	20.0	R	D		PG	A	46	(1)01	19	38	06	-00	33	08	6	81.9	34.5	(1)08	24	14	94	29	28	38	6	-84.2	-2.8	(2)01	19	38	32	-00	32	53	5	85.8	49.7	(2)08	24	25	98	29	30	13	7	59.8	92.3	(3)01	19	24	86	-00	33	33	8	-116.0	9.4	(3)08	24	22	30	29	25	57	3	11.8	-164.2																																																																																																																																																																																																																																																														
01	22	43	58	-00	34	04	9	-.7	3.0	20.0	B	S	(S)	PQ	A	0	08	27	25		32																																																																																																																																																																																																																																																																																																																																						

TABLE 3—Continued

1							2																						
POSITION(1950.0)			OFFSETS		MAG	TYPE	ID	POSITION(1950.0)			OFFSETS		MAG	TYPE	ID														
h	m	s	$\Delta\alpha$	$\Delta\delta$				h	m	s	$\Delta\alpha$	$\Delta\delta$																	
09	20	08.00	31	12	18.4	2.6	-4.6	19.0	B	S(S)	PO	A	16	(2)09	39	47.90	26	39	53.6	119.3	-54.2								
09	20	07.10	31	12	29.7	-8.9	10.6	19.5	B					(3)09	39	46.68	26	38	20.8	102.9	-147.0								
(1)09	20	01.86	31	13	31.2	-76.2	72.1							09	41	50.22	26	08	32.0	-4	-1.3	18.5	B	S(S)	Q	A	42		
(2)09	20	06.33	31	10	31.8	-18.9	-107.2							(1)09	41	49.90	26	08	31.5	-4.7	-1.7	18.5	B	S(S)					
(3)09	20	16.84	31	11	54.2	116.1	-24.8							(2)09	41	49.15	26	09	25.1	-14.8	51.8								
09	20	48.46	31	20	48.7	-1	3.5	18.0	B	S(S)	Q	A	41	(2)09	41	42.22	26	09	21.3	-108.1	48.0								
(1)09	20	53.56	31	20	16.3	65.3	-29.0							(3)09	41	56.65	26	05	43.1	86.3	-170.2								
(2)09	20	43.14	31	21	18.1	-68.2	32.9							09	42	25.73	27	39	21.3	5.7	5.0	19.5	N		PG	B	0		
(3)09	20	39.21	31	21	04.1	-118.6	18.8							09	42	24.27	27	39	21.5	-13.8	5.2	(20.0)							
09	22	29.80	31	00	19.8	-11.9	-5.7	(19.0)						(1)09	42	29.64	27	38	42.8	57.6	-33.5								
09	22	29.44	31	00	39.0	-16.5	13.5	19.5						(2)09	42	18.00	27	38	11.3	-97.0	-65.0								
(1)09	22	32.50	31	00	45.5	22.7	20.0							(3)09	42	40.26	27	39	24.5	198.8	8.2								
(2)09	22	27.68	31	01	40.9	-39.2	75.5							(1)09	42	43.88	26	54	05.9	74.0	-38.5							>	17
(3)09	22	30.00	30	58	32.7	-9.4	-112.8							(2)09	42	39.68	26	56	19.9	17.7	95.5								
09	22	33.03	32	13	08.1	1.1	16.2	18.0	B	S(S)	EF			(3)09	42	38.78	26	52	51.4	5.7	-113.0								
09	22	32.06	32	12	33.8	-11.2	-18.1	17.0	N	S(S)				(1)09	43	30.48	25	46	11.6	-18.2	-58.6								
(1)09	22	29.44	32	12	20.0	-44.3	-31.9							(2)09	43	19.33	25	49	06.6	-168.8	116.3								
(2)09	22	28.78	32	12	20.0	-52.7	-31.9							(3)09	43	47.62	25	44	16.2	213.3	-174.1								
(3)09	22	43.78	32	12	56.5	137.6	4.6							09	45	55.91	24	46	42.7	2.4	-7	19.0	R		PG	A	38		
09	22	43.36	23	22	57.4	-23.7	.4	19.0	R	S	EF			09	45	56.21	24	46	39.7	6.5	-2.3	17.0	R	D(D)					
(1)09	22	47.92	23	25	06.4	39.1	129.4							(1)09	45	52.18	24	46	48.1	-48.4	6.0								
(2)09	22	46.44	23	20	04.4	18.7	-172.5							(2)09	45	58.49	24	45	41.4	37.5	-60.6								
(3)09	22	32.16	23	25	02.6	-177.9	125.6							(3)09	46	02.16	24	47	28.3	87.6	46.3								
(1)09	23	14.51	26	12	03.2	108.2	-48.4							09	47	20.14	24	47	43.2	2.6	-19.0								
(2)09	23	16.84	26	14	28.0	139.6	96.4							(1)09	47	08.17	24	49	52.2	-160.4	110.0								
09	23	16.82	25	47	35.1	-3.8	-2.8	8.5	N	S(S)	OR	41*		(2)09	47	34.74	24	45	37.1	201.4	-145.1								
(1)09	23	21.84	25	47	56.2	64.1	18.4							09	49	10.01	24	36	36.0	-2.9	-2.2	18.5	R	S	PG	A	7*		
(2)09	23	05.03	25	49	06.5	-163.1	88.6							09	49	09.48	24	36	29.8	-10.1	-8.4	19.5							
(1)09	25	22.02	29	16	17.6	49.7	-64.2							(1)09	49	05.54	24	35	53.0	-63.8	-45.2								
(2)09	25	12.92	29	18	53.9	-69.3	92.1							(2)09	49	02.87	24	37	58.2	-100.3	80.0								
(3)09	25	08.72	29	18	21.0	-124.2	59.2							(3)09	49	08.89	24	39	31.9	-18.2	173.7								
09	27	07.21	25	13	17.1	.8	-1.0	18.5	N	S(C)	PQ	A	0*	09	49	12.70	28	42	37.1	2.5	-3.1	(20.0)	(D)	PG	A	17*			
(1)09	27	03.33	25	14	43.2	-51.8	85.2							(1)09	49	02.17	28	44	41.5	-136.2	121.4								
(2)09	27	17.36	25	12	37.4	138.6	-40.6							(2)09	49	22.62	28	44	47.9	133.0	127.8								
(3)09	27	08.54	25	10	18.4	18.9	-179.6							(3)09	49	26.27	28	43	52.1	180.9	72.0								
(1)09	27	33.35	31	24	28.5	-27.3	-38.3							09	53	59.72	25	29	33.9	4.7	4.1	16.5	B	S(S)	Q	B	3		
(2)09	27	42.10	31	24	08.9	84.7	-58.0							(1)09	54	05.43	25	30	19.2	82.0	49.4								
(3)09	27	36.00	31	22	12.5	6.5	-174.4							(2)09	53	52.99	25	31	22.4	-86.5	112.6								
09	31	49.52	31	04	19.1	-3.7	-3.6	20.0	B	S(C)	PQ	B	16	(3)09	54	07.62	25	27	33.6	111.7	-116.2								
(1)09	31	40.23	31	05	07.9	-123.1	45.2							09	55	25.36	32	38	23.2	-9	-4.6	17.0	B	S(S)	Q	A	16		
(2)09	31	55.74	31	06	16.7	76.2	113.9							(1)09	55	27.65	32	37	28.6	28.0	-55.2								
(3)09	32	01.22	31	04	40.7	146.6	18.0							(2)09	55	28.68	32	39	34.7	41.0	70.8								
09	32	30.67	24	08	33.8	-11.3	-3.6	19.0	R	C	C	PG	C	0*	(3)09	55	18.88	32	38	06.4	-82.7	-17.5							
09	32	32.66	24	08	27.8	16.0	-9.5	19.0	R	C				(1)09	58	33.33	25	41	22.9	-33.6	25.4								
09	32	31.55	24	09	00.1	.8	22.8	15.0	R	S(S)				(2)09	58	39.08	25	41	01.6	44.1	4.1								
09	32	32.99	24	08	25.0	20.5	-12.3	18.0	R	C				(3)09	58	36.64	25	39	08.2	11.1	-109.2								
09	32	31.68	24	08	12.6	2.6	-24.7	11.0	N	S(S)				(1)09	58	57.45	29	01	37.3	34.3	15.9								
09	32	32	24	08	3	25	SP	17.0	R	C(C)				(2)09	58	49.79	29	03	46.2	-66.1	144.7								
(1)09	32	39.79	24	07	08.6	113.6	-88.8							(3)09	58	43.14	29	02	27.2	-153.3	65.8								
(2)09	32	29.74	24	07	39.5	-147.1	-57.8							(1)10	03	55.17	29	01	43.0	13.4	-133.9								
(3)09	32	42.97	24	08	58.6	157.1	21.2							(2)10	04	07.03	29	04	24.7	168.9	27.8								
09	32	40.04	25	24	09.4	-5	-2.3	18.0	R	C(D)	PG	A	32	(3)10	04	06.90	29	02	28.7	167.2	-88.2								
09	32	41.17	25	24	00.9	14.8	-10.9	18.0	R	S(S)				10	04	30.82	23	06	07.7	.9	-2.7	20.0	B	(S)	PQ	A	0		
09	32	41	25	24	2	19	F	20.0	R	D				(1)10	04	27.76	23	05	09.6	-41.4	-60.7								
(1)09	32	37.51	25	24	10.4	-34.8	-1.4							(2)10	04	29.94	23	07	31.7	-11.3	81.4								
(2)09	32																												

TABLE 3—Continued

1		2		3		4		5		6		7		8					
POSITION(1950.0)		OFFSETS		MAG		TYPE		ID		POSITION(1950.0)		OFFSETS		MAG		TYPE		ID	
		$\Delta\alpha$	$\Delta\delta$									$\Delta\alpha$	$\Delta\delta$						
10 ^h 12 ^m 02 ^s .19	23° 15' 57".8	19.2	-12.4	19.0	R	C				10 ^h 36 ^m 52 ^s .40	31° 25' 41".0	4.3	12.5	18.5	R	C(C)	EF		
(1)10 11 58.63	23 14 59.1	-29.5	-71.1							10 36 53.12	31 25 16.8	13.6	-11.7	20.0	R				
(2)10 12 02.71	23 17 37.0	26.7	86.8							(1)10 36 53.39	31 27 12.8	17.0	104.3						
										(2)10 36 49.24	31 23 23.9	-36.0	-124.6						
										(3)10 36 41.60	31 23 46.7	-133.9	-101.8						
10 13 13	25 28.8	24	P	18.5	R	C(C)	EF			10 37 42.76	30 13 38.6	1.1	-1	17.0	N	D(D)	G	A	16
(1)10 13 17.95	25 29 17.1	33.0	26.8							(1)10 37 40.58	30 14 18.7	-27.2	40.0						
(2)10 13 15.09	25 27 24.9	-5.8	-85.4							(2)10 37 50.62	30 14 15.5	103.0	36.8						
(3)10 13 06.74	25 31 29.5	-118.9	159.2							(3)10 37 51.77	30 13 47.5	117.9	8.8						
10 13 36.53	25 40 24.2	5.8	-9.5	(20.0)			EF			10 39 48.55	30 05 23.3	-13.0	-2.8	18.5	R	C	EF		
10 13 36.27	25 40 21.2	2.3	-12.5	(20.0)						(1)10 39 51.20	30 04 08.3	21.5	-77.8						
(1)10 13 40.90	25 40 30.8	64.9	-2.9							(2)10 40 00.23	30 06 15.5	138.6	49.3						
(2)10 13 29.16	25 39 56.5	-93.8	-37.1							(3)10 39 58.27	30 07 28.2	113.1	122.0						
(3)10 13 42.00	25 41 55.7	79.8	82.0							10 40 31.44	31 46 44.8	5.1	-1	20.0	R	D	G	B	16
10 14 24.67	27 14 03.7	11.1	-8.4	(19.0)B	(S)		EF			10 40 31.10	31 46 51.8	7	6.8	15.0	N	D(D)			
10 14 22	27 14.2	16	P	20.0	R					(1)10 40 31.06	31 47 32.8	-1	47.9						
10 14 24	27 14.0	17	SF	(20.0)	(S)					(2)10 40 28.60	31 44 54.0	-31.1	-110.9						
10 14 25.14	27 14 10.3	17.3	-1.8	(19.0)B	(S)					(3)10 40 47.15	31 46 09.4	205.4	-35.5						
(1)10 14 19.29	27 14 17.0	-60.7	4.9							10 44 40.18	29 44 11.9	3.8	-1.9	18.5	R	S(S)	PG	A	16
(2)10 14 17.86	27 14 41.9	-79.8	29.7							10 44 41.16	29 44 26.8	16.5	13.0	18.5	C				
10 15 00.34	27 47 06.0	-4.6	-2.7	19.0	B	C(S)	Q	B	44	10 44 41	29 43.9	21	SF	19.0	R	C			
(1)10 15 05.76	27 49 18.3	67.3	129.6							(1)10 44 37.10	29 42 34.6	-36.3	-99.2						
(2)10 14 50.93	27 48 51.1	-129.5	102.4							(2)10 44 50.45	29 43 28.9	137.6	-45.0						
(3)10 15 16.53	27 46 32.1	210.3	-36.6							(3)10 44 45.37	29 46 36.7	71.4	142.8						
10 15 36	29 29.4	6	F	20.0			PG	B	16	(1)10 47 15.74	28 04 51.8	27.3	-2.1						
10 15 37.32	29 29 14.9	10.3	-9.8	20.0	R					(2)10 47 18.26	28 05 59.5	60.6	65.6						
(1)10 15 34.18	29 30 17.1	-30.7	52.4							(3)10 47 21.31	28 07 53.3	100.9	179.4						
(2)10 15 40.98	29 28 50.5	58.0	-34.2							10 47 34.82	28 46 53.3	6	10.3	19.0	R	C	PG	C	32
(3)10 15 29.88	29 30 11.3	-86.8	46.6							10 47 34.41	28 46 31.8	-4.8	-11.2	20.0	B	(S)			
(1)10 16 26.01	32 55 22.7	-8.1	-90.7				EF			10 47 33.67	28 46 46.5	-14.5	3.6	19.5	R	S			
(2)10 16 31.67	32 55 35.8	63.1	-77.6							(1)10 47 33.36	28 47 26.6	-18.6	43.6						
(3)10 16 28.97	32 59 47.2	29.1	173.8							(2)10 47 43.75	28 46 31.4	118.1	-11.5						
10 17 47.85	31 53 11.8	6.1	-3	20.0	R	(C)	PQ	B	16	(3)10 47 45.49	28 45 49.0	140.8	-54.0						
(1)10 17 49.03	31 53 43.5	21.2	31.4							10 48 02.34	30 33 42.6	1.3	4.3	(20.5)					
(2)10 17 58.67	31 52 41.9	143.9	-30.2							10 48 03.57	30 33 34.0	17.1	-4.3	20.0	R	D(D)	O	A	0
(3)10 17 47.33	31 50 34.4	-6	-157.8							(1)10 48 18.20	30 31 34.7	206.1	-123.6						
10 19 39.87	30 56 15.1	-1.4	6.5	18.0	B	S(S)	Q	B	39	(1)10 52 33.16	28 55 28.1	117.9	108.7						
(1)10 19 31.48	30 56 49.6	-109.2	41.0							(2)10 52 37.12	28 53 15.1	169.9	-24.4						
(2)10 19 49.20	30 54 56.3	118.7	-72.3							(3)10 52 41.26	28 54 03.3	224.2	23.8						
(3)10 19 31.89	30 57 57.4	-104.0	108.8							(1)10 52 38.39	28 29 55.4	19.0	37.6						
(1)10 20 49.45	28 51 07.3	86.8	-126.7				EF			(2)10 52 31.81	28 28 03.1	-67.7	-74.8						
10 21 09.03	29 04 29.1	-5.5	-19.3	20.0	B	D(D)	EF			(3)10 52 45.17	28 28 31.6	108.4	-46.2						
(1)10 21 04.55	29 03 36.8	-64.2	-71.6							10 53 30.35	24 11 44.0	6.7	-8.8	14.5	N	S(S)	EF	C	
(2)10 21 05.67	29 06 13.0	-49.5	84.6							10 53 28.60	24 12 06.3	-17.3	13.5	(19.0)B	(C)				
(3)10 21 21.00	29 04 27.6	151.4	-20.8							(1)10 53 22.92	24 10 17.1	-95.0	-95.6						
10 26 58.52	25 38 51.3	7	4	19.5	N		C	PG	A	0	(2)10 53 28.54	24 14 21.1	-18.1	148.3					
10 26 58.82	25 38 33.7	4.8	-17.2	18.5	R	C				(3)10 53 30.74	24 09 20.1	11.9	-152.6						
(1)10 26 56.03	25 39 04.5	-32.9	13.7							(1)10 53 53.64	28 01 40.8	-10.0	84.6						
(2)10 27 01.08	25 38 30.2	35.4	-20.6							(2)10 53 52.62	28 02 51.8	-23.5	155.6						
(3)10 26 55.31	25 39 17.8	-42.7	26.9							(3)10 53 56.29	28 03 01.8	25.1	165.6						
(1)10 27 28.31	32 48 53.3	-27.4	143.9				EF			10 54 43.58	26 54 05.4	19.0	0	19.0	R	C	>	O	17
(2)10 27 22.40	32 44 09.5	-102.0	-139.9							(1)10 54 40.55	26 54 22.4	-21.6	16.9						
(3)10 27 45.36	32 47 27.6	187.6	58.2							(2)10 54 48.30	26 54 31.5	82.0	26.0						
(1)10 28 24.30	23 21 30.4	239.1	37.3				EF			(3)10 54 32.81	26 53 53.7	-125.2	-11.7						
10 28 10.03	28 11 27.4	-4.4	3.9	19.5	B	(S)	PQ	B	38	(1)10 55 02.34	31 29 14.6	-37.1	-43.7						
(1)10 28 11.31	28 08 53.9	12.4	-149.6							(2)10 55 00.87	31 28 59.4	-66.1	-58.8						
(2)10 28 13.79	28 08 53.7	45.2	-149.8							(3)10 54 58.26	31 29 19.5	-89.4	-38.8						
(3)10 27 58.00	28 13 16.2	-163.5	112.8							10 56 50	28 37.9	15	NF	20.0	R	EF			
10 28 34.48	30 02 25.2	1.3	2.7	18.0	R	S	PG	A	16	10 56 48.24	28 37 53.3	-21.3	7.6	17.0	R	S(S)			
10 28 33.20	30 02 04.4	-15.3	-18.2	19.0	R					(1)10 56 37.15	28 37 39.1	-167.4	-6.6						
(1)10 28 30.28	30 02 04.3	-53.2	-18.2							(2)10 56 33.22	28 36 36.9	-219.1	-68.8						
(2)10 28 40.34	30 04 22.5	77.5	120.0							(1)10 57 09.31	30 43 05.6	-172.6	19.5						
(3)10 28 26.60	30 00 01.5	-101.0	-141.0							11 03 37.78	28 13 32.0	6.9	7.4	20.0	R	EF			
10 29 59.68	25 18 11.5	1.3	12.3	(20.0)	(D)		EF			(1)11 03 23.15	28 12 22.9	-186.6	-61.7						
(1)10 30 10.39	25 17 45.9	146.4	-13.3							(2)11 03 54.27	28 14 09.2	224.8	44.7						
(2)10 30 16.47	25 17 53.8	228.8	-5.3							11 05 58.07	31 29 09.0	11.8	15.5	20.0	R	EF			
10 32 30.61	26 30 48.7																		

TABLE 3—Continued

1		2		3		4		5		6		7		8						
POSITION(1950.0)		OFFSETS		MAG		TYPE		ID		POSITION(1950.0)		OFFSETS		MAG		TYPE		ID		
		$\frac{\Delta\alpha}{\cos\delta}$ $\frac{\Delta\delta}{\sin\delta}$										$\frac{\Delta\alpha}{\cos\delta}$ $\frac{\Delta\delta}{\sin\delta}$								
(1)11	18 02.82	23 44 18.6	-17.7	-3.2	18.5	N	C(C)	>G	0*	(1)12	01 09.46	30 53 16.9	8.1	-59.6						
(1)11	18 07.07	23 42 02.3	40.6	-139.5						(2)12	01 14.12	30 53 45.1	68.1	-31.4						
(2)11	17 51.47	23 44 02.9	-173.7	-18.9						(3)12	01 02.36	30 54 25.5	-83.3	9.0						
(3)11	17 56.32	23 42 03.3	-107.1	-138.5																
(1)11	21 41.84	31 29 35.9	-35.1	14.8					EF	(1)12	01 54.18	11 46 53.6	16.2	59.5						EF
(2)11	21 39.29	31 27 48.2	-67.6	-92.9						(2)12	01 58.36	11 46 22.6	77.5	28.5						
(3)11	21 40.67	31 31 13.7	-50.0	112.6						(3)12	01 45.13	11 45 36.6	-116.7	-17.5						
(1)11	25 26.90	26 01 35.3	4.7	-22.8	19.0	R	D(D)	C	EF *	(1)12	02 10.82	15 18 10.9	25.3	-15.8					C	EF
(1)11	25 23.94	26 01 55.0	-35.3	-3.2						(2)12	02 05.50	15 18 50.0	-51.7	23.3						
(2)11	25 26.75	26 02 37.7	2.6	39.5						(3)12	02 13.28	15 18 04.5	60.9	-22.2						
(3)11	25 26.60	25 59 20.9	.6	-157.2																
(1)11	26 38.03	29 23 28.6	-60.6	108.5					EF	(1)12	02 19.68	29 46 47.4	1.9	-43.0						EF *
(2)11	26 46.83	29 19 44.0	54.4	-116.1						(2)12	02 05.12	29 47 02.8	-187.7	-27.6						
(3)11	26 37.07	29 23 58.4	-73.1	138.3						(3)12	02 09.11	29 44 58.6	-135.7	-151.8						
(1)11	32 16.24	30 22 01.7	-.5	-0	18.5	B	S(S)	Q	A 16	(1)12	02 50.09	26 49 29.8	-3.1	-109.8					>G	17
(1)11	32 23.60	30 20 54.9	94.8	-66.8						(2)12	02 57.45	26 52 27.2	95.4	67.6						
(2)11	32 02.94	30 21 49.5	-172.6	-12.2						(3)12	02 50.85	26 53 19.5	6.9	120.0						
(3)11	32 28.83	30 20 07.9	162.4	-113.7																
(1)11	33 11.77	26 13 39.6	-3.5	6.5	17.0	N	S(S)	EF	*											
(1)11	33 07.75	26 13 15.1	-57.6	-18.1																
(2)11	33 03.99	26 14 38.3	-108.1	65.1						(1)12	02 51.64	32 09 08.4	-66.3	45.8						EF
(3)11	33 04.07	26 15 25.2	-107.1	112.0						(2)12	02 50.51	32 07 10.9	-80.6	-71.7						
										(3)12	02 59.57	32 10 12.1	34.5	109.5						
(1)11	34 59.67	26 34 59.3	99.4	-49.8					>Q	17	(1)12	03 07.15	26 42 17.8	105.0	-3.7					EF *
(2)11	34 59.15	26 38 30.2	92.4	161.1							(2)12	02 47.48	26 43 30.2	-158.6	68.7					
(3)11	35 08.27	26 34 08.2	214.7	-101.0																
(1)11	35 26.05	31 21 32.8	-148.4	17.6					>G	17										
(2)11	35 53.95	31 21 01.0	209.0	-14.2							12 03 22.63	10 59 35.6	1.4	-.4	17.5	B	S(S)	Q	A 40	
(3)11	35 53.47	31 19 41.9	202.9	-93.3							(1)12	03 25.10	11 00 22.0	37.8	46.0					
											(2)12	03 24.33	11 01 13.9	26.5	97.9					
											(3)12	03 14.75	10 59 55.7	-114.6	19.7					
(1)11	36 47.32	29 54 14.3	3.9	-9.5	20.0	R		PG	B 0	(1)12	03 21.68	26 56 51.4	-38.3	61.2						EF
(1)11	36 51.53	29 51 27.1	58.6	-176.7						(2)12	03 18.10	26 55 56.2	-86.2	6.1						
(2)11	36 34.19	29 52 35.2	-166.8	-108.6																
(3)11	37 04.96	29 54 58.0	233.3	34.2																
(1)11	36 56.04	30 24 36.6	65.2	58.5					EF		12 04 14.91	32 09 13.0	-17.2	-15.4	18.5	R	D(D)	C	EF	
(2)11	36 55.01	30 22 08.2	51.9	-90.0						(1)12	04 14.34	32 09 35.6	-24.6	7.2						
(3)11	37 03.30	30 24 31.2	159.2	53.1						(2)12	04 09.65	32 11 02.6	-84.0	94.2						
										(3)12	04 10.66	32 07 21.4	-71.2	-127.0						
(1)11	44 38.39	25 39 42.0	-8.9	-17.3	17.0	N	S(S)	EF			12 06 31.27	23 26 32.0	-17.8	-16.3	18.5	R	S		EF	
(1)11	44 41.93	25 38 29.4	39.0	-89.9							12 06 33.38	23 26 26.9	11.2	-21.5	18.0	R	S			
(2)11	44 48.75	25 41 28.9	131.2	89.6						(1)12	06 28.22	23 27 16.7	-59.8	28.3						
(3)11	44 52.41	25 38 53.9	180.7	-65.4						(2)12	06 41.34	23 28 23.8	120.7	95.4						
										(3)12	06 20.76	23 27 46.6	-162.5	58.2						
(1)11	46 32.87	26 51 40.3	-4.7	38.7					EF		12 07 34.19	11 50 34.5	2.8	-2.5	19.5	B	(S)	PQ	A 0	
										(1)12	07 28.56	11 51 31.1	-79.9	54.2						
										(2)12	07 25.23	11 51 54.8	-128.8	77.8						
(1)11	46 55.26	25 31 14.3	1.4	2.0	18.5	B	S(S)	PQ	A 42											
(1)11	46 55.85	25 31 25.0	9.3	12.6	20.0	R														
(2)11	47 00.96	25 29 07.1	78.6	-125.2																
(3)11	46 59.29	25 28 54.6	56.0	-137.7						(1)12	10 06.13	24 13 47.5	-67.7	-70.1						EF
(3)11	46 56.41	25 33 45.7	17.0	153.4						(2)12	10 11.63	24 16 56.7	7.5	119.1						
										(3)12	10 10.27	24 17 20.4	-11.0	142.8						
(1)11	47 10.97	28 54 41.0	-15.6	3.8	19.5	B	D(D)	EF												
(1)11	47 15.53	28 55 36.3	44.3	59.1							12 10 59.28	13 24 02.0	.7	.4	19.0	B	(S)	Q	A 40	
(2)11	47 16.39	28 53 02.8	55.5	-94.4						(1)12	11 06.14	13 24 30.9	100.8	29.3						
(3)11	47 14.74	28 52 27.4	33.9	-129.7						(2)12	10 52.47	13 26 15.4	-98.6	133.9						
										(3)12	11 01.22	13 21 10.9	29.0	-170.7						
(1)11	47 44.01	24 34 34.5	3.5	3.1	17.0	N	S(S)	Q	A 41											
(1)11	47 44.12	24 34 14.2	5.0	-17.2	20.0	B	(C)				12 11 53.67	14 19 39.3	-2.0	-.3	19.0	R	D(D)	PG	A 39	
(1)11	47 45.08	24 34 34.6	18.0	3.2	20.0	N					12 11 55.47	14 19 38.7	24.1	-.9	17.0	N	C(C)			
(1)11	47 48.74	24 33 27.0	68.0	-64.4						(1)12	11 53.29	14 18 22.4	-7.4	-77.2						
(2)11	47 45.41	24 36 30.2	22.6	118.8						(2)12	11 44.84	14 19 53.6	-130.2	14.0						
(3)11	47 53.11	24 34 19.9	127.5	-11.5						(3)12	11 44.77	14 22 35.7	-131.2	176.1						
(1)11	51 37.97	29 32 49.6	-3.2	-.4	19.0	R	C	C	G A 16	(1)12	12 34.26	17 46 55.4	-114.5	14.1						EF
(1)11	51 38.59	29 32 48.3	4.9	-1.7	19.0	R	D			(2)12	12 46.46	17 45 00.6	59.7	-100.7						
(1)11	51 38.37	29 32 41.4	2.1	-8.6	20.0	R				(3)12	12 46.70	17 45 00.8	63.1	-100.5						
(1)11	51 37.78	29 32 29.9	-5.7	-20.1	18.5	R	D													
(1)11	51 41.37	29 33 58.2	41.1	68.2							12 13 17.82	32 08 14.5	-2.2	-.8	(19.0)	B	(D)	PQ	A 15	
(2)11	51 37.91	29 30 47.0	-4.0	-123.0																

TABLE 3—Continued

1								2																	
POSITION(1950.0)				OFFSETS				MAG	TYPE	ID	POSITION(1950.0)				OFFSETS				MAG	TYPE	ID				
1	2	3	4	$\Delta\alpha$	$\Delta\delta$	$\Delta\alpha$	$\Delta\delta$				1	2	3	4	$\Delta\alpha$	$\Delta\delta$	$\Delta\alpha$	$\Delta\delta$							
(3)12	18	43	22	28	29	44	3	-234.0	-54.8		12	41	15	68	31	40	45	3	8.2	4.3	20.0 R	EF			
12	19	23		31	47	6					12	41	14	27	31	40	48	9	-9.7	3.9	(20.0)				
(1)12	19	21	73	31	48	20	1	-18.0	46.0		12	41	15	96	31	40	45	2	11.8	.2	(20.0)				
(2)12	19	24	08	31	48	54	6	11.9	80.6		12	41	14	10	31	40	34	8	-11.9	-10.2	(19.0)	(S)			
(3)12	19	08	27	31	45	41	7	-189.6	-112.4		(1)12	41	12	89	31	39	11	6	-27.4	-93.4					
12	21	14	60	18	37	43	4	2.0	-1.0	18.5 B S(S)	0	A	40	(2)12	41	15	51	31	43	00	5	6.0	135.5		
(1)12	21	15	91	18	37	23	2	20.7	-21.2		(3)12	41	31	69	31	40	35	0	212.6	-10.0					
(2)12	21	12	56	18	37	03	0	-26.9	-41.4		12	41	27	56	16	39	18	2	.6	1.7	20.0 B (S)	Q A 13			
(3)12	21	19	46	18	37	44	7	71.1	.3		(1)12	41	27	94	16	40	11	0	6.0	54.6					
(1)12	21	24	38	29	59	04	8	89.0	-104.0	C EF	(2)12	41	21	86	16	38	42	5	-81.3	-34.0					
(2)12	21	26	47	29	58	50	2	116.1	-118.5		(3)12	41	25	87	16	37	50	6	-23.6	-85.9					
(3)12	21	02	70	30	00	37	2	-192.6	-11.6		12	41	34	22	25	25	09	8	15.3	6.5	16.0 N S(S)	EF *			
12	21	20	05	16	24	49	2	-2.0	3.8	20.0 B	(1)12	41	33	91	25	23	10	1	11.2	-113.2					
12	21	20	77	16	24	38	1	8.4	-7.4	20.0 R D	(2)12	41	29	50	25	27	24	9	-48.5	141.6					
(1)12	21	25	17	16	25	2		24	N	(20.0)	(3)12	41	49	38	25	23	54	4	220.7	-68.9					
(2)12	21	18	22	16	26	36	1	-28.2	110.6		12	41	50	70	24	39	52	5	21.6	4.3	18.0 R D(D)	EF			
12	22	01	83	26	29	50	6	-15.2	-1.5	19.0 R D	(1)12	41	45	09	24	39	19	8	-54.9	-28.4					
(1)12	22	00	25	26	31	01	2	-36.4	69.2		(2)12	41	47	98	24	41	02	3	-15.4	74.1					
(2)12	21	56	07	26	31	13	4	-92.6	81.3		(3)12	41	47	33	24	42	29	5	-24.4	161.4					
(3)12	21	54	15	26	28	43	0	-118.3	-69.1		12	44	45	97	12	53	20	5	8.2	-2.9	18.0 B S(S)	PQ B 0			
(1)12	22	11	76	24	13	32	3	75.0	-153.5		(1)12	44	43	33	12	53	16	2	-30.3	-7.2					
(1)12	23	33	33	17	39	40	9	71.6	41.5	EF	(2)12	44	42	10	12	51	11	4	-48.3	-132.1					
(2)12	23	26	41	17	40	46	4	-27.3	106.9		(3)12	44	50	17	12	56	01	8	69.6	158.4					
(3)12	23	25	59	17	40	50	3	-39.0	110.9		(1)12	44	52	92	32	44	19	8	85.9	-26.0					
12	27	01	58	18	06	55	4	1.0	-.9	19.0 R D	(2)12	44	47	08	32	47	07	0	12.2	141.2					
(1)12	27	06	88	18	06	25	3	76.5	-31.0		(3)12	44	54	89	32	42	50	5	110.8	-115.3					
(2)12	27	09	57	18	06	32	1	114.9	-24.2		12	45	37	82	18	54	32	9	2.9	-.4	18.0 B S(S)	Q A 42			
(3)12	26	52	71	18	07	48	7	-125.5	52.3		(1)12	45	37	47	18	53	48	2	-2.0	-45.1					
(1)12	27	05	35	16	18	23	8	-6.1	-61.4	EF	(2)12	45	34	28	18	54	18	8	-47.3	-14.5					
(2)12	27	08	51	16	21	01	9	39.5	96.8		(3)12	45	38	70	18	52	19	6	15.5	-133.7					
(3)12	26	55	78	16	18	20	8	-143.8	-64.4		(1)12	45	45	02	17	02	24	4	65.6	16.8					
12	28	19	53	25	41	36	3	-6.8	-.9	(19.5) B (D)	(2)12	45	40	32	16	59	35	0	-1.8	-152.6					
(1)12	28	18	30	25	40	41	8	-23.5	-55.4		12	45	46	56	11	33	41	9	4.4	-2.8	13.0 R S(S)	EF B *			
(2)12	28	28	54	25	39	34	0	115.0	-123.2		(1)12	45	53	24	11	32	14	8	102.7	-90.0					
(3)12	28	24	39	25	44	24	0	58.8	166.7		(2)12	45	57	28	11	34	54	1	162.0	69.4					
(1)12	30	53	27	23	02	40	5	-44.9	-20.9	EF	(3)12	45	32	22	11	32	15	0	-206.3	-89.7					
(2)12	31	02	37	23	03	22	5	80.7	21.0		(1)12	48	40	77	32	26	12	0	-16.3	48.3					
(3)12	31	06	03	23	04	36	5	131.3	95.0		(2)12	48	46	03	32	24	11	9	50.2	-71.9					
12	31	56	34	24	48	15	0	2.8	-.4	18.5 B S(S)	(3)12	48	50	88	32	25	11	6	111.7	-12.2					
(1)12	31	53	03	24	49	37	7	-42.2	82.3		12	50	11	35	29	08	08	0	.9	2.9	17.5 R S(D)	C G A 2			
(2)12	31	49	25	24	48	43	1	-93.6	27.6		(1)12	50	10	62	29	08	07	2	-8.7	2.1	19.5 R C				
(3)12	32	22	23	29	34	01	7	-52.5	-12.9	EF	(2)12	50	03	31	29	07	00	5	-104.4	-64.6					
(2)12	32	31	84	29	34	46	4	72.8	31.7		(3)12	50	54	09	29	07	12	3	-225.2	-52.9					
(3)12	32	32	97	29	32	55	2	87.5	-79.4		(1)12	50	28	48	29	06	44	5	225.4	-80.6					
(1)12	34	07	37	30	32	28	0	-19.0	45.9	EF	(1)12	50	23	37	11	56	15	8	27.9	37.2					
(2)12	34	16	31	30	31	26	0	96.4	-16.1		(2)12	50	23	21	11	54	50	4	25.6	-48.2					
(3)12	34	14	27	30	29	58	1	70.1	-104.1		(3)12	50	26	26	11	54	15	8	70.3	-82.8					
12	34	45	24	25	14	34	6	4.6	2.6	19.5 B (S)	(1)12	51	07	59	15	59	24	2	65.7	43.4					
(1)12	34	44	32	25	13	46	4	-7.9	-45.6		(2)12	51	06	51	16	00	57	5	50.2	136.7					
(2)12	34	59	35	25	13	24	1	195.9	-67.8		(3)12	51	02	54	16	01	09	6	-7.0	148.9					
(3)12	35	00	29	25	13	20	2	208.8	-71.8		12	52	07	70	11	57	21	4	-.7	-.6	16.5 B S(S)	Q A 13			
12	35	06	46	19	41	10	3	-8.0	1.4	19.5 B (S)	(1)12	52	08	57	11	56	20	8	12.0	-61.1					
(1)12	35	05	67	19	40	23	7	-19.2	-45.2		(2)12	52	14	72	11	58	31	9	102.2	69.9					
(2)12	35	03	93	19	42	12	1	-43.8	63.2		(3)12	52	51	59	92	11	59	04	1	-114.8	102.1				
(3)12	35	00	82	19	40	53	5	-87.7	-15.4		12	52	13	25	25	9	24	NF	20.0 R	EF					
12	36	42	42	32	46	55	6	.7	-1.8	(20.5)	(1)12	52	12	32	25	27	15	6	5.4	94.2					
(1)12	36	42	20	32	47	18	7	-2.0	21.3	(20.0)	(2)12	52	07	11	25	24	01	3	-65.1	-100.1					
(2)12	36	43	60	32	46	35	7	15.7	-21.7		(3)12	52	06	67	25	27	56	9	-71.1	135.5					
(3)12	36	24	95	32	48	13	8	-219.5	76.4		12	53	20	57	31	02	30	0	-17.9	8.2	19.0 R D(D)	EF			
(1)12	36	59	29	32	44	57	1	213.5	-120.3		(1)12	53	24	55	31	02	34	4	33.2	12.7					
(1)12	37	31	10	23	07	38	4	-66.0	-43.6	EF *	(2)12	53	24	70	31	01	29	1	35.2	-52.6					
(2)12	37	34	59	23	10	47	0	-17.8	145.0		(3)12	53	18	38	31	00	55	5	-46.0	-86.2					
(3)12	37	25	99	23	07	23	6	-136.4	-58.4		12	56	57	78	28	11	05	2	6.0	8.4	15.5 R S(S)	C G C 43			
12	38	20	72	24	21	50	2	16.7	4.0	20.0 N D(D)	(1)12	56	58	51	28	10	51	7	15.7	-5.1	14.9 R D(D)				
(1)12	38	22	57	24	22	04	8	41.9	18.6		(1)12	57	04	37	28	12	22	9	93.2	86.1					
(2)12	38	23	16	24	21	42	3	49.9	-3.9		(2)12	56	51	08	28	13	21	5	-82.5	144.6					
12</																									

TABLE 3—Continued

1	2	3	4	5	6	7	8	1	2	3	4	5	6	7	8
POSITION(1950.0)		OFFSETS		MAG	TYPE	ID		POSITION(1950.0)		OFFSETS		MAG	TYPE	ID	
		$\Delta\alpha$	$\Delta\delta$							$\Delta\alpha$	$\Delta\delta$				
13 03 19.87	19 17 44.1	-4.1	1.5	20.0	R	PG A 0*		13 19 43.98	13 53 56.1	-16.4	1.9	19.5	R C	EF	
(1)13 03 17.43	19 17 20.1	-34.7	-22.4					(1)13 19 47.60	13 52 46.6	36.2	-67.6				
(2)13 03 15.93	19 18 06.5	-55.9	23.9					(2)13 19 39.97	13 51 44.0	-74.9	-130.2				
(3)13 03 25.12	19 18 54.6	74.2	72.0					(3)13 19 45.57	13 51 18.6	6.6	-155.6				
13 04 42.77	25 02 59.3	-6.8	-2.7	19.0	R S	PG R 0		(1)13 19 49.75	27 01 01.0	-42.6	-25.8			C EF	*
(1)13 04 47.08	25 02 01.3	51.7	-60.7					(2)13 19 43.29	27 02 17.1	-128.9	50.3				
(2)13 04 35.44	25 02 10.3	-106.4	-51.6					(3)13 19 44.23	27 03 15.9	-116.4	109.1				
(3)13 04 40.41	25 05 12.2	-38.9	130.3												
13 04 45.27	24 23 43.0	3.8	7.6	18.5	N S(S)	PG R 0		13 19 57.16	23 00 55.2	.3	1.6	20.0	R D	PG A 0	
(1)13 05 00.01	24 24 04.5	205.2	29.1					(1)13 19 56.26	23 01 54.0	-12.1	60.3				
(2)13 04 29.96	24 21 59.2	-205.2	-96.2					(2)13 20 02.97	22 59 32.2	80.5	-81.5				
(3)13 05 02.06	24 23 34.6	233.2	-8					(3)13 20 00.83	23 03 38.5	50.9	164.9				
13 06 29.25	12 20 21.3	11.9	14.6	17.0	N S(S)	EF		13 20 28.57	32 31 47.8	-7	-2.5	(20.0)	(D)	PQ A 17*	
(1)13 06 29.83	12 20 01.1	20.4	-5.6					(1)13 20 30.44	32 32 12.5	22.9	22.1				
(2)13 06 33.60	12 19 37.0	-70.9	-29.7					(2)13 20 27.13	32 31 04.3	-19.0	-46.0				
(3)13 06 33.99	12 20 45.6	81.5	38.9					(3)13 20 25.99	32 30 57.1	-33.4	-53.2				
(3)13 06 26.83	12 22 25.5	-23.6	138.8												
13 06 33.12	27 24 10.5	1.7	1.3	18.5	B S(S)	Q A 42		13 20 40.46	29 57 25.0	-16.4	11.2	19.0	N D(D)	C EF	
(1)13 06 21.96	27 24 00.2	-147.0	-9.0					13 20 42	29 57.5	20	NF	19.0	D		
(2)13 06 31.25	27 26 53.6	-23.2	164.4					(1)13 20 37.94	29 57 28.3	-49.1	14.5				
(3)13 06 43.41	27 27 03.7	138.7	174.4					(2)13 20 47.54	29 58 41.8	75.7	88.0				
								(3)13 20 34.37	29 55 11.4	-95.5	-122.4				
13 07 04.37	12 10 22.1	-8.4	-4.3	19.0	B(S)	Q B 1		(1)13 21 16.13	18 09 56.4	69.5	-7.7				EF
(1)13 07 05.34	12 10 27.5	20.5	1.1					(2)13 21 16.75	18 08 18.0	78.3	-106.1				
(2)13 07 03.30	12 09 00.3	-24.0	-86.0					(3)13 21 09.87	18 07 13.8	-19.9	-170.3				
(3)13 06 51.52	12 10 20.2	-196.8	-6.1												
(3)13 07 18.86	12 09 01.8	204.1	-84.5					13 23 57.94	32 09 43.0	-2	-2	20.0	R D(D)	PG A 16	
								13 23 58.62	32 09 40.3	8.4	-2.9	19.0	R C		
13 08 07.50	32 36 40.3	-4.3	-1.1	19.0	B S(S)	PQ A 16*		(1)13 24 05.57	32 10 48.4	96.6	65.2				
(1)13 07 56.88	32 36 04.6	-138.5	-36.8					(2)13 23 47.21	32 10 06.0	-136.5	22.9				
(2)13 08 14.42	32 33 47.7	83.1	-173.7					(3)13 24 09.07	32 07 51.4	141.1	-111.8				
(3)13 07 49.69	32 36 06.0	-229.4	-35.4												
13 08 23.66	12 05 48.5	-3	-1	(20.0)		PQ A 3R		13 24 31	23 00.8	18	S	18.0	R S	EF	*
(1)13 08 28.28	12 06 18.0	67.4	29.5					(1)13 24 16.73	23 03 21.1	-204.7	132.6				
(2)13 08 10.62	12 05 08.5	-191.6	-40.0												
13 08 29.50	18 15 33.5	-2.2	-2.5	18.0	B S(S)	Q A 39		(1)13 25 10.83	16 25 19.8	-26.9	6.4				EF
(1)13 08 19.70	18 15 31.9	-141.8	-4.1					(2)13 25 11.37	16 24 12.4	-19.2	-61.0				
(2)13 08 41.23	18 17 28.7	164.9	112.6					(3)13 25 18.13	16 25 10.6	78.0	-2.8				
(3)13 08 42.25	18 13 06.0	179.4	-150.1												
13 08 58.95	17 13 46.1	-14.3	.0	(20.0)	(C)	C EF		13 25 13.29	32 06 58.1	.4	-6	18.5	R C(D)	C G A 16	
(1)13 09 08.16	17 13 08.0	117.7	-38.1					(1)13 25 12.60	32 07 31.1	-8.3	32.4				
(2)13 08 58.79	17 16 07.0	-16.6	140.9					(2)13 25 10.83	32 07 46.3	-30.9	47.5				
(3)13 09 11.49	17 12 57.4	165.3	-68.0					(3)13 25 00.30	32 05 56.8	-164.7	-62.0				
(1)13 09 03.34	31 26 30.9	17.4	-93.8			EF		13 25 20.08	12 28 53.5	5.6	.7	20.0	R D	PG B 0	
(2)13 08 57.85	31 29 58.1	-52.7	113.5					(1)13 25 17.07	12 26 59.6	-38.5	-113.2				
(3)13 08 52.59	31 29 35.9	-120.0	91.3					(2)13 25 29.25	12 28 48.0	139.9	-4.8				
13 09 05.38	25 56 23.6	-5.2	-1.8	18.5	N(D)	PG B 3B		(3)13 25 07.52	12 26 15.0	-178.3	-157.8				
(1)13 09 04.10	25 56 28.0	-22.5	2.6					13 25 26.44	12 38 39.8	-1.2	-1.6	19.5	N D(D)	PG A 0	
(1)13 09 03.43	25 56 36.3	-31.5	10.9					13 25 28.18	12 38 43.0	24.2	1.6	19.0	N D(D)		
(2)13 09 10.43	25 54 30.0	62.9	-115.4					(1)13 25 23.08	12 36 47.2	-50.4	-114.1				
(3)13 09 03.90	25 59 18.6	-25.2	173.2					(2)13 25 23.41	12 41 24.5	-45.5	163.1				
								(3)13 25 15.10	12 40 37.7	-167.2	116.3				
13 09 28	32 44.1	20	N	20.0	R	>G 17		(1)13 25 54.78	15 38 04.5	47.5	-10.9				EF
(1)13 09 37.04	32 42 53.3	102.8	-50.2					(2)13 25 57.04	15 39 17.3	80.2	61.9				
(2)13 09 37.85	32 43 15.3	113.0	-28.2					(3)13 25 45.99	15 36 24.2	-79.5	-111.2				
(3)13 09 23.75	32 42 05.8	-65.0	-97.7					13 26 07.64	15 03 25.2	-6.5	2.1	20.0	R	PG R 0*	
(1)13 09 28.39	18 55 41.9	-34.0	18.7			EF		13 26 08.81	15 03 40.4	10.5	17.3	19.5	C(C)		
(2)13 09 30.52	18 56 58.9	-3.9	95.7					(1)13 26 10.01	15 02 34.1	27.9	-49.0				
(3)13 09 32.77	18 57 03.9	28.1	100.6					(2)13 26 00.20	15 03 13.9	-114.2	-9.2				
(1)13 09 58.86	25 33 12.6	-107.0	-36.4			EF		(3)13 26 16.08	15 03 06.9	115.8	-16.2				
(2)13 10 16.88	25 34 18.3	136.8	29.3					(1)13 27 42.93	16 46 47.4	9.5	23.8				EF
13 17 26.35	19 57 19.4	-2.7	-1	20.0	R	PG A 0		(2)13 27 37.00	16 46 53.9	-75.6	30.4				
(1)13 17 26.80	19 57 19.7	3.7	.2					13 28 15.89	25 24 38.2	-7	.5	18.0	B C	Q A 45	
(1)13 17 26.92	19 59 29.8	5.4	130.4					(1)13 28 16.84	25 24 59.7	12.2	22.0				
(2)13 17 37.29	19 57 27.9	151.6	8.5					(2)13 28 15.28	25 25 46.9	-9.0	69.2				
(3)13 17 11.12	19 56 28.5	-217.3	-51.0					(3)13 28 18.71	25 26 24.5	37.6	106.8				
13 17 37.31	25 48 27.2	12.5	13.1	16.5	R S(S)	C EF		(1)13 28 45.41	14 08 34.3	-53.3	-2.4				C EF
(1)13 17 35	25 47.9	23	SP	19.0	R S			(2)13 28 52.90	14 08 39.8	55.5	3.1				
(1)13 17 37.32	25 48 36.7	12.6	22.6					(3)13 28 54.76	14 06 08.4	82.7	-148.3				
(2)13 17 40.19	25 47 23.4	51.4	-50.7					13 28 49.65	30 45 58.9	-1	-1	17.0	B S(S)	Q A 16	
(3)13 17 43.94	25 48 00.6	102.0	-13.6					(1)13 28 51.39	30 45 03.0	22.4	-56.0				
13 17 54.39	17 58 58.9	-6.6	5.1	20.0	R	PG R 24		(2)13 28 42.59	30 45 24.8	-91.1	-34.2				
(1)13 17 54.26	17 58 48.8	-8.5	-5.0					(3)13 28 43.42	30 48 02.9	-80.4	123.9				
(1)13 17 54.25	17 59 47.4	-8.7	53.6					(1)13 30 12.38	23 24 06.7	-44.6	-22.2				EF
(2)13 17 57.27	17 57 41.7	33.7	-72.1					(2)13 30 18.73	23 26 37.4	42.8	128.5				
(3)13 18 05.51	17 59 33.1	152.0	39.3					(3)13 30 25.41	23 23 28.1	134.8	-60.8				
13 18 49.65	11 22 31.3	-1.1													

TABLE 3—Continued

1								2																					
POSITION(1950.0)				OFFSETS				MAG TYPE ID				POSITION(1950.0)				OFFSETS				MAG TYPE ID									
1	2	3	4	Δα	Δδ	5	6	7	8	9	10	1	2	3	4	Δα	Δδ	5	6	7	8	9	10						
(1)13	32	18.17	31	53	11.8	-32.2	93.3					(1)13	50	03.14	31	41	33.4	4.1	0.0	15.6	N	D(D)	G	A	14				
(2)13	32	18.83	31	49	20.8	-23.8	-137.6					(2)13	50	00.58	31	41	17.9	-32.7	-15.5										
(3)13	32	08.08	31	51	50.3	-160.7	11.8					(3)13	50	03.67	31	40	45.0	6.9	-48.4										
(1)13	32	37.19	32	04	07.3	.3	39.0			EF		(1)13	50	07.14	31	42	49.5	51.1	76.0										
(2)13	32	37.95	32	02	03.1	9.9	-85.2					(2)13	50	26.43	11	21	26.2	-33.5	-27.9			EF							
(3)13	32	20.23	32	00	59.1	-215.3	-149.2					(3)13	50	30.64	11	20	55.3	28.4	-58.8										
(1)13	33	14.61	27	35	27.1	17.0	172.2			EF	*	(1)13	50	27.21	11	20	51.7	-22.1	-62.4										
(2)13	33	25.88	27	33	47.9	166.9	73.0					(1)13	51	31.46	32	06	36.7	2.2	2.0	20.0	R		PG	A	17				
(3)13	33	00.24	27	33	30.9	-174.1	56.0					(2)13	51	32.55	32	06	40.9	16.1	6.1	20.0	R								
(1)13	33	57.62	23	43	17.3	-15.4	-17.8	(20.0)		EF		(1)13	51	36.40	32	06	22.2	65.0	-12.5										
(2)13	33	59.29	23	42	36.2	7.6	-59.0					(2)13	51	25.87	32	08	12.1	-68.7	97.3										
(3)13	33	49.52	23	44	16.9	-126.6	41.7					(3)13	51	36.28	32	08	17.7	63.5	102.9										
(3)13	33	56.70	23	41	07.3	-28.0	-147.9					(1)13	52	13.47	17	30	34.6	5.0	-23.0	19.0	R	S(S)	EF						
(1)13	33	53.84	26	26	36.1	-113.3	66.8			C	EF	(1)13	52	15.15	17	30	59.0	29.0	1.4										
(2)13	34	02.96	26	23	06.0	9.2	-143.3					(2)13	52	14.74	17	31	27.8	23.2	30.2										
(3)13	34	02.02	26	27	56.2	-3.4	146.9					(3)13	52	18.53	17	33	13.9	77.4	136.3										
(1)13	34	12.28	28	50	00.4	-15.4	21.0			C	>G	17	(1)13	52	16.12	16	29	34.3	-3.4	6.8	19.0	R	D	C	G	B	46		
(2)13	34	13.44	28	50	32.1	-.1	52.8					(2)13	52	24.02	16	28	30.1	110.2	-57.4										
(3)13	34	04.14	28	49	25.3	-122.3	-14.1					(3)13	52	07.08	16	27	26.7	-133.4	-120.8										
(1)13	35	36.06	17	06	03.4	35.8	137.3			EF		(1)13	53	35.33	18	38	19.3	-6.7	27.8			EF	*						
(2)13	35	21.75	17	03	07.5	-169.5	-38.6					(2)13	53	38.75	18	38	06.5	42.0	15.0										
(3)13	35	20.87	17	03	31.1	-182.0	-15.0					(3)13	53	37.83	18	38	54.9	28.8	63.4										
(1)13	35	51.97	32	19	28.8	62.7	108.1			EF		(1)13	54	42.06	19	33	44.2	1.5	-3.2	16.5	B	S(S)	Q	A	4				
(2)13	35	41.46	32	20	23.7	-70.6	163.0					(2)13	54	45.94	19	33	43.2	56.4	-4.2										
(1)13	38	23	28	52.8	12	F	20.0	D(D)	EF	*	(2)13	54	40.00	19	32	56.6	-27.6	-50.9											
(2)13	38	23.49	28	53	02.5	13.8	14.5	19.5	R			(1)13	54	48.35	25	52	01.1	-.5	-4.5	18.5	B	S(S)	Q	A	41				
(3)13	38	22.57	28	52	02.0	1.7	-46.0					(2)13	54	48.48	25	52	16.6	1.2	11.1	19.0	R	S							
(1)13	38	27.85	28	52	54.0	71.1	6.0					(1)13	54	46.44	25	49	55.2	-26.3	-130.4										
(3)13	38	18.41	28	51	21.1	-53.0	-86.8					(2)13	54	37.03	25	51	18.3	-153.3	-47.3										
(1)13	38	26.15	29	34	26.7	-11.1	-12.1	19.0	R	C	>G	17	(3)13	54	44.04	25	49	14.5	-58.7	-171.1									
(2)13	38	37.28	29	34	56.8	134.1	17.9					(1)13	57	01.71	17	21	32.6	-9.0	-.6	20.0	D		PG	B	0*				
(3)13	38	23.58	29	37	21.3	-44.6	162.5					(2)13	57	01.95	17	21	52.3	-5.5	19.1	17.0	N	S(S)							
(1)13	38	20.56	29	32	12.6	-84.0	-146.3					(3)13	57	01.48	17	21	16.9	-12.2	-16.3	18.5	N	S(S)							
(1)13	38	36.40	10	47	28.2	10.6	12.1	20.0	B	D(S)	EF	*	(1)13	56	59.64	17	22	31.6	-38.6	58.4									
(2)13	38	36.25	10	46	58.4	8.5	-17.7	(19.5)	B	(S)			(2)13	57	04.51	17	23	16.9	31.1	103.8									
(3)13	38	34.13	10	47	30.1	-22.8	14.0					(3)13	57	10.05	17	19	35.4	110.4	-117.8										
(1)13	38	35.60	10	46	44.8	-1.2	-31.3					(1)13	57	51.57	27	33	46.8	3.4	2.9	19.0	R	D	C	PG	A	44			
(3)13	38	31.66	10	46	21.8	-59.2	-54.3					(2)13	57	50.14	27	33	51.1	-15.6	7.3	20.0	N	D(D)							
(1)13	39	27.82	12	14	30.4	-35.6	43.7			EF		(1)13	57	57.32	27	31	33.7	79.8	-130.2										
(2)13	39	26.25	12	14	01.0	-58.7	14.3					(2)13	58	02.15	27	30	48.0	144.1	-175.8										
(3)13	39	25.60	12	14	35.4	-68.1	48.7					(3)13	57	38.49	27	36	19.9	-170.5	156.1										
(1)13	39	31.06	26	37	34.5	2.9	3.4	19.0	R	C	C	G	A	0*	(1)13	58	41.72	24	28	00.8	-110.7	-67.0			EF				
(2)13	39	31.48	26	37	32.8	8.5	1.8	(19.0)	R	(D)			(2)13	58	58.93	24	30	24.1	124.1	76.4									
(3)13	39	30.84	26	37	19.7	-.1	-11.4	13.0	R	D(D)			(3)13	58	58.59	24	30	55.6	119.5	107.9									
(1)13	39	31.72	26	37	37.9	11.8	6.9	18.5	R	D			(1)13	59	46.95	10	20	38.5	23.9	-4.2	20.0	R		EF					
(2)13	39	29.49	26	37	31.6	-18.1	-.6	13.0	R	D(D)			(2)13	59	48.03	10	21	11.2	39.7	28.4									
(3)13	39	29.04	26	37	37.1	-24.2	6.1	18.0				(3)13	59	37.64	10	21	23.2	-113.6	40.4										
(1)13	39	30.91	26	36	16.8	.9	-74.2					(3)13	59	53.37	10	21	13.8	118.5	31.0										
(1)13	40	46.79	31	58	48.9	-4.3	-2.6	20.0	B	(C)	PG	B	16	(1)14	00	20.60	16	14	20.7	-.7	-2.5	16.5	B	S(S)	Q	A	40		
(2)13	40	58.85	31	57	49.7	149.1	-61.7					(2)14	00	20.42	16	14	15.7	-3.2	-7.6	19.0	R	D							
(3)13	40	58.25	31	57	22.2	141.4	-89.2					(3)14	00	21.38	16	14	36.8	10.7	13.5	19.5	R								
(1)13	41	54.86	14	23	45.1	-38.0	-32.4			EF		(1)14	00	21.87	16	14	24.5	17.7	1.3	19.0	B	S(S)							
(2)13	42	00.01	14	25	21.4	36.8	63.9					(2)14	00	20.91	16	15	40.6	3.8	77.3										
(3)13	41	52.51	14	24	43.4	-72.1	25.9					(3)14	00	12.28	16	12	29.5	-120.4	-113.8										
(1)13	45	06.28	12	32	19.9	.9	-.6	16.5	R	D(D)	G	A	13	(3)14	00	29.30	16	12	25.6	124.6	-117.7								
(2)13	45	01.32	12	32	53.0	-71.8	32.5					(1)14	01	06.08	12	20	37.3	-3.2	7.5	16.0	R	S(S)	EF	*					
(3)13	45	11.79	12	32	17.0	81.5	-3.5					(2)14	00	53.00	12	21	24.6	-194.9	54.7										
(1)13	45	56.77	12	30	28.7	-138.4	-111.9					(3)14	00	52.49	12	19	58.0	-202.4	-31.9										
(1)13	45	37.87	26	13	17.7	8.1	8.2	20.0	R	D	PG	C	0	(3)14	01	20.36	12	20	49.7	206.1	19.9								
(2)13	45	39.36	26	13	57.4	28.2	47.8					(1)14	04	06.24	18	46	49.4	-9.1	-4.4	(20.5)		EF							
(3)13	45	43.55	26	12	51.2	84.6	-18.4					(2)14	04	11.82	18	47	24.0	70.1	30.2										
(1)13	45	28.93	26	15	30.2	-112.2	140.6					(3)14	04	19.60	18	47	33.5	180.6	39.7										
(1)13	45	53.88	24	29	54.																								

TABLE 3—Continued

1	2	3	4	5	6	7	8	1	2	3	4	5	6	7	8
POSITION(1950.0)		OFFSETS		MAG	TYPE	ID		POSITION(1950.0)		OFFSETS		MAG	TYPE	ID	
		$\Delta\alpha$	$\Delta\delta$							$\Delta\alpha$	$\Delta\delta$				
14 ^h 07 ^m 52 ^s .15	31° 38' 17".7	4.6	15.0	19.0	R S	EF	*	(1)14 ^h 21 ^m 49 ^s .62	19° 04' 02".0	2.1	-144.3			EF	
(1)14 07 58.65	31 36 30.4	87.6	-92.2					(2)14 21 36.02	19 07 35.4	-190.6	69.1				
(2)14 07 41.84	31 38 14.1	-127.1	11.4					(3)14 21 34.21	19 07 28.2	-216.2	61.9				
(3)14 07 53.87	31 35 36.9	26.5	-145.8												
(1)14 07 54.28	17 47 47.2	11.7	-22.5			EF		14 22 53.68	12 52 57.7	-1.3	4.4	19.0	B S(S)	PQ A 0	
(2)14 07 50.04	17 47 24.3	-48.8	-45.5					14 22 55.13	12 53 01.2	20.0	8.0	18.0	N S(S)		
(3)14 07 47.10	17 47 46.4	-90.7	-23.3					(1)14 22 52.91	12 53 42.9	-12.4	49.7				
								(2)14 22 56.15	12 52 06.5	34.9	-46.8				
								(3)14 22 48.85	12 52 25.9	-71.9	-27.4				
14 07 58.41	14 52 07.7	-20.6	12.2	19.0	R C	EF		(1)14 24 45.27	10 37 00.2	-87.1	-3.6			EF	
(1)14 08 01.94	14 52 37.9	30.5	42.4					(2)14 24 45.28	10 36 36.4	-87.0	-27.3				
(2)14 08 05.55	14 51 11.3	82.7	-44.2					(3)14 24 51.17	10 35 13.8	-2	-110.0				
(3)14 07 50.83	14 50 35.6	-130.5	-79.9												
14 08 34	14 10.9	9 F		20.0	R D	PG B 0		(1)14 25 22.95	26 44 34.2	20.4	-17.4			Q C 15 [#]	
14 08 33.68	14 11 20.6	-5.4	23.9	15.0	N S(S)			(2)14 25 21.80	26 45 38.9	5.1	47.2				
(1)14 08 28.28	14 10 56.3	-84.0	-4					(3)14 25 16.17	26 42 41.3	-70.3	-130.4				
(2)14 08 28.01	14 10 43.9	-88.0	-12.7												
(3)14 08 41.24	14 11 04.1	104.4	7.5					14 25 28.32	28 46 44.4	-2.6	-5.5	16.5	N S(S)	PQ B 0	
14 09 12.62	13 18 10.3	-12.1	-15.3	20.0	R D	EF	*	(1)14 25 27.20	28 49 03.4	-17.4	133.5				
(1)14 09 12.03	13 18 00.5	-20.7	-25.0					(2)14 25 34.99	28 48 42.2	85.0	112.2				
(2)14 09 11.97	13 16 59.3	-21.6	-86.3					(3)14 25 39.33	28 44 45.7	142.1	-124.2				
(3)14 09 14.60	13 20 41.0	16.9	135.5					14 26 20.18	13 09 57.0	1.9	-1.3	18.5	R D	G A 40	
(1)14 09 51.02	29 23 39.4	26.0	-131.2			EF		14 26 19.65	13 09 47.0	-5.7	-11.3	19.5	R C		
(2)14 09 49.39	29 23 05.0	4.7	-165.5					(1)14 26 21.20	13 09 38.7	16.9	-19.6				
(3)14 09 58.57	29 23 30.4	124.8	-140.2					(2)14 26 27.03	13 08 55.5	101.9	-62.8				
								(3)14 26 09.76	13 08 57.2	-150.2	-61.1				
(1)14 10 06.16	11 52 51.6	106.9	-8.6			EF		14 26 32.59	29 32 26.5	.0	-2.4	18.5	B S(S)	Q A 16	
(2)14 10 00.60	11 54 48.6	25.4	108.4					14 26 33.50	29 32 37.0	11.8	8.1				
(3)14 09 51.43	11 51 29.4	-109.3	-90.8					(1)14 26 28.83	29 32 18.9	-49.0	-10.1				
								(2)14 26 39.79	29 32 54.4	93.9	25.5				
								(3)14 26 17.06	29 33 03.9	-202.6	34.9				
14 10 13.56	29 35 28.1	9.4	-6.4	20.0	R	EF		14 27 28.40	30 34 17.5	-19.7	-15.1	17.0	N S(S)	EF	
14 10 12.69	29 35 15.1	-2.0	-19.3	20.0	R			(1)14 27 25.63	30 33 47.8	-55.5	-44.9				
(1)14 10 02.28	29 35 55.2	-137.9	20.7					(2)14 27 33.16	30 33 28.6	41.7	-64.1				
(2)14 10 15.10	29 38 27.4	29.5	172.9					(3)14 27 39.93	30 37 28.6	129.1	175.9				
(3)14 09 54.88	29 32 57.5	-234.3	-157.0					(1)14 28 17.47	25 26 02.9	50.7	-63.7			EF	
(1)14 10 17.18	15 28 45.7	-36.9	-27.0			C EF		(2)14 28 10.29	25 28 42.0	-46.5	95.4				
(2)14 10 23.71	15 29 59.8	57.5	47.1					(3)14 28 20.24	25 28 13.6	88.2	67.0				
(3)14 10 23.17	15 27 01.2	49.6	-131.5					14 28 48.18	26 03 16.6	-1	-3.9	20.0	R	PG A 0	
14 13 38	25 23.9	17 NF	20.0			EF		(1)14 28 49.84	26 03 50.0	22.3	29.5				
14 13 38	25 23.4	24 SF	20.0	(D)				(2)14 28 53.62	26 02 31.4	73.2	-49.1				
(1)14 13 34.10	25 24 04.8	-44.9	22.8					(3)14 28 48.77	26 00 58.6	8.0	-141.9				
(2)14 13 37.47	25 22 10.3	.8	-91.7					14 30 26.82	25 08 28.4	-3.1	-6.7	17.0	N D(D)	G B 32	
(3)14 13 26.74	25 22 07.0	-144.6	-95.0					(1)14 30 19.29	25 09 28.9	-105.3	53.9				
14 13 44.48	24 30 39.8	20.6	-8.5	17.0	R D(D)	PG C 0		(2)14 30 36.43	25 10 00.7	127.4	85.7				
(1)14 13 33.28	24 32 45.7	-132.3	117.3					(3)14 30 38.34	25 07 53.1	153.3	-41.9				
(2)14 13 48.14	24 33 33.6	70.6	165.2					14 30 36.10	17 15 57.9	-70.8	-17.3			EF	
(3)14 13 39.71	24 27 50.6	-44.5	-177.7					(2)14 30 35.68	17 15 24.7	-76.9	-50.4				
14 15 41.59	17 17 13.0	1.4	-5	18.5	B S(S)	Q A 1		(3)14 30 44.38	17 14 40.1	47.9	-95.1				
(1)14 15 45.36	17 14 35.3	55.4	-158.3					(1)14 31 13.65	16 20 32.4	102.1	-8.9			EF	
(1)14 17 15.56	30 20 50.2	26.7	-92.3			EF		(2)14 31 07.37	16 22 53.8	11.7	132.5				
(2)14 17 18.36	30 24 52.5	62.9	150.0					(3)14 30 51.44	16 18 09.6	-217.7	-151.9				
(3)14 17 28.04	30 21 05.5	188.2	-77.0					14 33 06	31 53.8	6 F	20.0			PQ B 17	
14 17 18.83	14 27 10.3	6.3	16.0	12.0	N S(S)	EF		14 33 08	31 53.8	22 F	20.0				
(1)14 17 15.44	14 27 47.1	-42.8	52.8					(1)14 32 56.83	31 55 00.7	-123.3	75.2				
(2)14 17 12.79	14 25 24.6	-81.4	-89.7					(2)14 33 21.20	31 53 50.7	187.0	5.2				
(1)14 17 48.93	27 15 33.4	56.7	6.8			EF		(3)14 33 15.86	31 51 07.4	119.0	-158.1				
(2)14 17 34.15	27 16 07.7	-140.3	41.1					(1)14 33 20.18	23 58 26.0	-29.0	39.3			EF	
(1)14 17 57.96	30 16 57.0	25.5	-167.1					(2)14 33 21.46	23 58 36.6	-11.4	49.9				
(2)14 17 52.80	30 22 39.9	-61.4	175.7					(3)14 33 25.39	23 59 12.2	42.4	85.5				
(3)14 18 12.52	30 18 57.7	213.9	-46.4					14 33 36.12	17 42 35.6	-6	.6	18.5	B S(S)	Q A 1	
14 19 19.31	31 32 43.1	.4	2.3	19.5	B C	PQ A 16		(1)14 33 39.80	17 40 08.8	51.9	-146.1				
14 19 20.65	31 32 40.7	17.6	-1	19.0	R C			(2)14 33 25.77	17 45 01.3	-148.6	146.3				
(1)14 19 24.16	31 31 01.4	62.4	-99.3					14 35 29.97	30 26 00.8	-2.4	8.7	20.0	R	C PG C 0	
(2)14 19 23.16	31 34 46.0	49.6	125.3					14 35 30.56	30 25 44.6	5.2	-7.5	20.0	R		
(1)14 19 33.40	18 02 21.4	26.5	-12.0			EF		14 35 29.94	30 25 35.6	-2.8	-16.4	19.0	R D		
(2)14 19 34.50	18 02 32.0	42.1	-1.3					14 35 31.22	30 26 06.5	13.8	14.4	18.5	R D		
(3)14 19 33.52	18 01 59.9	28.2	-33.5					(1)14 35 34.02	30 26 44.7	50.1	52.6				
14 20 21.21	32 36 48.1	6.9	8.0	18.0	B S(S)	Q C 41		(2)14 35 28.62	30 23 58.8	-19.9	-113.2				
(1)14 20 15.71	32 37 48.8	-62.6	68.7					(3)14 35 42.14	30 26 25.6	155.1	33.5				
(2)14 20 16.53	32 33 44.4	-52.2	-175.8					14 35 34.34	24 52 03.0	-6	-5	19.0	B (S)	Q A 32	
(3)14 20 31.32	32 39 01.9	134.6	141.8					14 35 35.07	24 52 19.8	9.3	16.3	16.0	N S(S)		
14 20 40.03	19 49 13.2	-14.2	18.4	19.5	R D	G C 46*		14 35 34.04	24 51 44.5	-4.7	-19.0	19.0	N C(C)		
(1)14 20 37.98	19 49 11.1	-43.0	16.3					(1)14 35 33.81	24 52 40.1	-7.8	36.7				
(2)14 20 44.24	19 49 05.7	45.3	10.9					(2)14 35 30.59	24 51 12.6	-51.7	-50.8				
(3)14 20 39.73	19 49 41.0	-18.3	46.1					(3)14 35 24.88	24 52 13.2	-129.3	9.7				
14 21 04.66	12 13 26.3	-4	-1	18.0	B S(S)	Q A 41		(1)14 35 45.82	28 33 39.0	-174.2	-64.2			EF	
14 21 03.62	12 13 19.3	-15.6	-7.1	19.5	R S			(2)14 36 10.03	28 36 58.8	144.7					

TABLE 3—Continued

1								2											
POSITION(1950.0)				OFFSETS		MAG	TYPE	ID	POSITION(1950.0)				OFFSETS		MAG	TYPE	ID		
1	2	3	4	$\Delta\alpha$	$\Delta\delta$				1	2	3	4	$\Delta\alpha$	$\Delta\delta$					
(2)14	36	04	16.0	12	59	18.2			(2)14	45	31.94	16	44	26.9	-231.6	-152.0			
(3)14	36	13.48	12	59	16.4	135.2	134.2		(1)14	49	13.61	13	55	39.1	66.1	-4.6	EF		
	14	36	51.87	11	35	43.2	3.5	-4.6	18.0	B	S(S)								
	14	36	51.83	11	35	56.8	3.0	9.1	19.0	N	D(D)								
	14	36	50.44	11	35	47.1	-17.5	-4.6	17.0	R	S(S)								
(1)14	36	48.83	11	35	43.5	-41.2	-4.3		14	49	16.52	14	26	21.5	1.1	1.9	20.0	R	
(2)14	36	55.69	11	36	11.9	59.7	24.1		14	49	16.07	14	26	18.9	-5.4	-7	18.5	N	
(3)14	36	47.30	11	34	49.8	-63.7	-57.9		14	49	17.28	14	26	11.0	12.0	-8.6	20.0	B	
	14	39	03.42	11	08	06.4	7.8	-6.6	(20.0)										
	14	39	04.19	11	07	59.7	19.1	-13.2	20.0	R									
(1)14	38	56.59	11	09	26.2	-92.7	73.3		(1)14	49	15.47	14	26	45.7	-14.2	26.1			
(2)14	39	06.77	11	06	01.3	57.1	-131.6		(2)14	49	19.21	14	26	34.8	40.1	15.2			
(3)14	39	09.82	11	09	54.0	102.0	101.1		(3)14	49	22.09	14	26	32.1	81.9	12.5			
	14	39	40.73	13	42	01.1	-0	-1.5	18.0	B	S(S)								
	14	39	41.11	13	41	57.4	5.5	-5.2	19.5	B	(D)								
	14	39	40.00	13	42	01.3	-10.6	-1.3	20.0	R	S								
	14	39	41.67	13	41	58.0	13.7	-4.6	17.0	R	S(S)								
(1)14	39	38.83	13	41	49.6	-27.7	-13.0		(1)14	51	15.06	19	05	49.2	-4.8	-24.7			
(2)14	39	42.98	13	41	39.1	32.8	-23.5		(2)14	51	10.76	19	06	35.9	-65.8	21.9			
(3)14	39	44.33	13	42	11.6	52.5	9.0		(3)14	51	23.93	19	05	47.2	120.9	-26.8			
	14	39	54.09	32	47	05.8	2.0	-2	19.5	B	(S)								
	14	39	54.34	32	47	26.5	5.1	20.5	18.5	R	D(D)								
(1)14	39	54.91	32	47	32.8	12.3	26.8		14	51	43.03	27	00	44.1	1.6	3.4	19.5	B	
(2)14	39	56.85	32	44	29.3	36.7	-156.7		14	51	43.91	27	00	55.2	13.3	14.5	20.0	R	
	14	40	33.20	30	54	47.8	-3.5	-24.1	19.5	R	S								
(1)14	40	24.82	30	54	21.0	-111.3	-50.9		(1)14	52	08.14	25	51	48.9	-40.6	-12.5			
(2)14	40	20.08	30	54	57.3	-172.2	-14.6		(2)14	52	08.52	25	50	56.1	-35.5	-65.4			
(3)14	40	43.11	30	52	44.0	124.2	-147.8		(3)14	52	11.36	25	53	58.5	2.8	117.1			
(1)14	42	01.92	14	11	40.4	-6.6	-49.2		14	52	25.06	30	08	06.7	-2.1	-2	19.0	B	
(2)14	42	00.34	14	11	34.6	-29.6	-55.0		(1)14	52	23.50	30	08	25.4	-22.4	18.5			
(3)14	41	48.97	14	11	33.6	-194.9	-55.9		(2)14	52	24.94	30	09	00.0	-3.8	53.1			
	14	42	16.17	19	34	05.4	-2.2	-4.3	18.5	N									
	14	42	15.68	19	34	21.7	-9.1	11.9	16.0	N	S(D)								
	14	42	15.37	19	33	59.1	-13.5	-10.6	17.5	R	D								
	14	42	17.10	19	33	55.6	10.9	-14.1	19.0	N	D(D)								
(1)14	42	27.60	19	33	35.8	159.3	-33.9		(1)14	54	21.12	30	23	52.7	8.4	-2.1	16.5	B	
(2)14	42	06.80	19	32	10.2	-134.6	-119.5		(1)14	54	23.84	30	24	14.8	43.6	20.0			
(3)14	42	05.74	19	36	20.8	-149.6	131.1		(2)14	54	24.19	30	23	41.1	48.1	-13.7			
	14	42	44.13	17	02	08.6	6.7	7.6	19.0	R	C								
	14	42	42.35	17	02	05.7	-18.9	4.8	18.5	R	C(D)								
(1)14	42	45.25	17	02	13.1	22.6	12.2		(2)14	54	21.34	30	22	56.8	11.3	-58.1			
(2)14	42	42.52	17	01	10.8	-16.5	-50.2		14	54	21.37	18	20	19.6	-9	-1.1	19.5	B	
(3)14	42	49.45	17	01	50.0	82.9	-10.9		14	54	21.93	18	20	30.7	7.0	10.1	20.0	B	
	14	42	50.49	10	11	12.4	-3	.1	18.5	B	S(S)								
	14	42	51.34	10	11	17.8	12.2	5.4	19.0	R	S								
	14	42	51.74	10	11	06.4	18.2	-5.9	17.5	B	S(S)								
(1)14	42	49.13	10	11	48.4	-20.4	36.1		(1)14	54	46.80	13	59	13.5	66.5	-46.7			
(2)14	42	47.19	10	10	24.1	-49.1	-48.2		(2)14	54	39.51	13	58	29.4	-39.6	-90.8			
(3)14	42	51.69	10	09	31.0	17.4	-101.3		(3)14	54	37.61	14	01	56.6	-67.3	116.4			
	14	43	20.31	26	36	25.7	-3.8	2.3	18.5	B	C(S)								
(1)14	43	23.80	26	36	41.5	43.1	18.1		14	55	31.43	24	46	54.5	6.4	-10.9	19.5	R	
(2)14	43	24.49	26	37	13.8	52.3	50.4		14	55	30.69	24	46	47.8	-3.6	-17.6	19.5	R	
(3)14	43	09.83	26	34	42.6	-144.2	-100.8		14	55	29.89	24	46	48.2	-14.5	-17.1	19.5	R	
	14	43	29.81	23	15	13.5	1.0	.0	18.5	B	S(S)								
	14	43	29	23	15.1	9	5	20.0	S	(S)									
	14	43	30.52	23	14	59.1	10.7	-14.4	20.0	R									
	14	43	28.42	23	15	09.9	-18.3	-3.6	19.0	R									
(1)14	43	31.41	23	15	03.3	23.0	-10.2		(1)14	55	22.34	11	56	05.3	-162.3	-17.3			
(2)14	43	25.28	23	13	28.1	-61.5	-105.4		14	55	44.03	14	46	23.5	3.4	16.6	20.0	R	
(3)14	43	29.58	23	12	33.7	-2.2	-159.8		(1)14	55	41.75	14	45	19.9	-29.7	-47.0			
	14	43	38.09	17	51	02.2	.3	-3.0	15.5	R	D(D)								
	14	43	37.61	17	50	53.5	-6.5	-11.7	(19.0)	B	(S)								
(1)14	43	37.45	17	50	36.0	-8.9	-29.2		(2)14	55	45.46	28	44	17.4	10.9	31.8			
(2)14	43	38.09	17	49	55.0	.3	-70.2		(2)14	55	48.76	28	43	48.3	54.2	2.6			
(3)14	43	33.38	17	50	26.9	-67.0	-38.3		(3)14	55	40.77	28	43	03.0	-50.9	-42.7			
	14	44	37.16	23	03	29.9	-8	-2.7	20.0	R									
	14	44	37.09	23	03	35.1	-1.8	2.6	19.0	N	D(S)								
	14	44	37.09	23	03	09.8	-1.8	-22.7	17.5	N	S(S)								
(1)14	44	35.71	23	03	03.4	-20.8	-29.2		(1)14	56	17.71	18	42	24.5	25.5	-28.9			
(2)14	44	38.87	23	05	28.5	22.7	115.9		(2)14	56	16.60	18	41	46.1	9.8	-67.4			
(3)14	44	28.09	23	02	40.7	-126.0	-51.9		(3)14	56	16.99	18	44	46.5	15.3	113.0			
	14	44	48.66	18	00	36.3	-5.9	12.7	19.0	R	D								
	14	44	48.02	18	00	08.9	-15.0	-14.8	(18.5)	B	(S)								
(1)14	44	48.39	18	01	03.8	-9.8	40.1		14	58	00.06	14	45	16.8	2.3	10.0	19.0	R	
(2)14	44	46.94	18	01	21.3	-30.4	57.6		(1)14	57	58.21	14	45	00.0	-24.6	-6.9			
(3)14	44	53.39	18	00	52.0	61.7	28.3		(2)14	58	00.71	14	45	39.2	11.6	32.3			
	14	45	23.15	16	47	32.5	-17.1	-68.0		(3)14	58	01.52	14	45	35.1	23.5	28.2		
(2)14	45	22.44	16	49	49.5	-27.3	69.0		(1)14	58	46.61	31	24	57.8	26.2	-3.6			
(3)14	45	15.30	16	47	47.7	-129.9	-52.8		(2)14	58	48.15	31	25	00.9	45.9	-5			
	14	45	48.48	16	47	09.3	5.9	10.4	17.5	N	(S)								
(1)14	45	49.94	16	45	24.0	26.9	-94.9		(3)14	58	34.26	31	25	16.6	-13				

TABLE 3—Continued

1		2		3		4		5		6		7		8					
POSITION(1950.0)		OFFSETS		MAG		TYPE		ID		POSITION(1950.0)		OFFSETS		MAG		TYPE		ID	
α δ		$\Delta\alpha$ $\Delta\delta$		$\Delta\alpha$ $\Delta\delta$						α δ		$\Delta\alpha$ $\Delta\delta$							
(3)15 01 39.94	12 41 59.2	169.7	-7.7							(2)15 27 27.56	15 48 20.4	4.0	-99.1						
15 02 00.16	10 41 16.3	0	-2.5	19.5	B	S(S)	0	A	21	(3)15 27 18.85	15 50 16.5	-121.7	17.1						
(1)15 02 01.78	10 40 36.9	23.9	-41.9							15 27 36.30	19 14 49.3	-2.6	-3	20.0	B		PQ	A	38
(2)15 02 05.20	10 42 39.9	74.2	81.0							15 27 36.14	19 14 40.8	-4.8	-8.7	(20.0)					
(3)15 01 53.65	10 39 36.8	-96.0	-102.0							15 27 37.07	19 15 07.7	8.4	18.1	19.5	R				
15 04 50.62	16 18 38.3	12.6	-7.7	17.5	B	S(S)	EF			(1)15 27 42.35	19 14 43.5	83.2	-6.0						
15 04 50.30	16 19 10.4	8.0	24.3	18.5	B	S(S)				(2)15 27 28.97	19 14 46.2	-106.4	-3.3						
(1)15 04 53.90	16 17 35.9	59.7	-70.1							(3)15 27 35.61	19 16 42.2	-12.4	112.6						
(2)15 04 52.77	16 20 59.8	43.5	133.7							15 30 54.24	13 42 28.0	.4	.6	19.0	B	(S)	Q	A	39
(3)15 04 41.47	16 17 20.4	-119.1	-85.6							15 30 54.78	13 42 07.0	8.2	-20.4	(20.0)	(D)				
15 05 18.85	12 19 16.2	4.9	-16.1	18.5	R	S	EF			(1)15 30 57.26	13 41 53.3	44.3	-34.1						
(1)15 05 12.79	12 18 52.3	-84.0	-60.0							(2)15 30 51.10	13 43 03.9	-45.5	36.5						
(2)15 05 15.51	12 21 19.5	-44.2	107.2							(3)15 30 56.17	13 41 27.6	28.4	-59.8						
(3)15 05 10.98	12 20 11.6	-110.6	39.2							15 30 55.27	15 27 00.3	-2.2	16.0	18.5	N	S(S)	EF		
15 06 53.70	14 17 35.1	-11.6	-1.5	18.5	B	S(S)	PQ	C	19	(1)15 30 55.36	15 27 28.7	-1.0	44.4						
(1)15 06 55.88	14 17 51.7	20.1	15.0							(2)15 30 58.87	15 26 33.5	49.8	-10.7						
(2)15 06 52.29	14 19 05.6	-32.1	88.9							(3)15 31 00.38	15 24 58.3	71.7	-105.9						
(3)15 06 47.86	14 16 53.1	-96.5	-43.6							15 32 02.13	13 59 15.7	.3	-1.2	18.5	R	D(D)	PG	A	21
15 08 25.25	10 53 53.7	-8.2	22.8	19.0	B	S(S)	EF			15 32 03.07	13 59 21.9	14.1	5.0	18.0	N	S(S)			
(1)15 08 22.21	10 53 15.3	-53.0	-15.6							15 32 03	13 59.6	25	NF	19.5	R				
(2)15 08 21.45	10 53 51.3	-64.2	20.5							(1)15 32 06.51	13 59 31.7	64.1	14.8						
(3)15 08 27.78	10 54 50.1	29.0	79.3							(2)15 32 04.89	14 00 16.0	40.5	59.1						
15 09 05.51	10 12 44.8	13.1	-14.9	17.0	R	S(D)	EF			(3)15 32 04.89	14 00 38.1	40.5	81.2						
15 09 06.21	10 12 55.6	23.5	-4.2	19.0	B	(D)				(1)15 32 31.23	19 27 01.7	-1.8	-64.9						
(1)15 09 07.03	10 13 57.9	35.5	58.1							(2)15 32 32.45	19 30 04.8	15.5	118.1						
(2)15 09 10.00	10 12 48.0	79.4	-11.7							(3)15 32 41.89	19 28 35.4	149.0	28.7						
(3)15 08 58.56	10 13 33.6	-89.6	33.8							(1)15 34 18.47	17 26 27.6	18.0	30.7						
15 09 52.46	15 51 39.6	1.8	1.6	18.5	B	S(S)	Q	A	1	(2)15 34 12.52	17 26 17.5	-67.1	20.6						
(1)15 09 45.35	15 52 36.5	-100.9	58.5							(3)15 34 25.45	17 26 04.2	117.9	7.3						
(2)15 09 52.12	15 49 35.7	-3.2	-122.3							(1)15 35 10.30	13 54 25.2	-24.3	-10.9						
(3)15 09 54.55	15 53 41.9	31.9	123.9							(2)15 35 07.45	13 55 39.3	-65.7	63.2						
(1)15 11 49.07	15 54 22.5	-98.5	75.7							(3)15 35 13.06	13 52 39.3	15.9	-116.8						
(2)15 11 50.51	15 51 14.0	-77.8	-112.8							15 37 08.30	16 13 53.3	11.8	1.0	18.5	R	S(S)	EF		
(3)15 11 46.26	15 54 29.8	-139.1	83.0							15 37 07.11	16 13 41.2	-5.3	-11.2	18.5	N	S(S)			
15 13 00.74	10 29 25.5	17.6	-12.8	20.0	B	(D)	EF			15 37 08.53	16 14 04.0	15.2	11.7	19.0					
(1)15 13 02.60	10 29 58.1	45.0	19.9							(1)15 37 06.58	16 14 56.9	-13.0	64.5						
(2)15 13 02.49	10 31 51.0	-104.0	132.8							(2)15 37 06.95	16 12 46.6	-7.6	-65.8						
(3)15 13 09.92	10 28 19.6	153.0	-78.7							(3)15 37 03.27	16 14 24.9	-60.6	32.5						
15 14 39.80	18 41 20.1	-2.0	-5	20.0	B	(D)	PG	A	39	15 37 45.72	14 29 43.2	6.7	-16.2	19.0	R	D(D)	EF		
(1)15 14 37.43	18 40 46.0	-35.7	-34.6							(1)15 37 44.58	14 30 39.4	-9.9	40.0						
(2)15 14 36.40	18 42 03.1	-50.3	42.5							(2)15 37 44.30	14 27 18.2	-13.9	-161.1						
(3)15 14 45.85	18 41 07.3	83.9	-13.3							(3)15 37 56.21	14 29 23.2	159.0	-36.1						
15 18 10.74	15 37 13.2	4.2	13.4	17.5	N	(S)	PQ	C	0*	15 38 30.25	14 57 21.8	-6	-1.0	16.5	B	S(S)	Q	A	41
15 18 11.10	15 36 47.1	9.3	-12.7	19.5	N					15 38 31	14 57.6	15	NF	19.5	R	C			
(1)15 18 06.15	15 35 10.6	-62.1	-109.2							(1)15 38 31.20	14 56 29.7	13.2	-53.1						
(2)15 18 13.68	15 34 35.8	46.6	-144.0							(2)15 38 34.86	14 56 41.2	66.2	-41.7						
(3)15 17 59.92	15 38 20.1	-152.2	80.3							(3)15 38 31.99	14 59 21.1	24.6	118.2						
15 19 48.24	10 52 10.4	-8	4	18.0	R	D	PG	A	38	15 38 53	15 53.4	13	SF	20.0	R				
(1)15 19 46.08	10 50 44.3	-32.5	-85.7							15 38 53.03	15 53 55.7	3.7	21.6	19.5	R	S	EF		
(2)15 19 48.51	10 54 08.6	3.3	118.6							(1)15 38 50.95	15 54 51.6	-26.4	77.5						
(3)15 19 41.55	10 53 35.9	-99.3	85.9							(2)15 38 59.47	15 53 48.3	96.7	14.3						
(1)15 21 04.29	13 44 20.6	54.0	73.4							(3)15 38 59.15	15 52 43.7	92.0	-50.4						
(2)15 20 57.08	13 44 29.4	-51.1	82.2							15 40 03.63	18 05 38.4	1.4	-1.3	18.5	B	S(S)	Q	A	38
(3)15 21 00.49	13 41 11.2	-1.4	-116.0							15 40 04.69	18 05 53.4	16.4	13.8	13.0	N	S(S)			
15 21 33.30	11 06 17.5	-1.1	-1.0	19.0	R	D	C	PG	A	40	(1)15 40 08.76	18 04 54.4	74.5	-45.2					
(1)15 21 34.90	11 07 19.0	22.5	60.5							(2)15 40 04.43	18 07 19.0	12.8	99.4						
(2)15 21 36.03	11 05 24.6	39.1	-53.9							(3)15 40 02.18	18 03 38.2	-19.3	-121.5						
(3)15 21 27.67	11 06 42.1	-84.0	23.6							15 41 28.90	18 56 44.9	-4.0	1.0	19.0	B	(S)	PQ	A	38
15 22 39.47	11 18 15.5	6.2	-1.8	18.5	B	S(S)	PQ	B	38	(1)15 41 30.25	18 57 33.0	15.2	49.2						
15 22 40.18	11 18 19.0	16.6	1.7	19.5	B	(S)				(2)15 41 24.76	18 55 39.1	-62.7	-64.7						
(1)15 22 37.33	11 17 00.7	-25.3	-76.6							(3)15 41 32.51	18 54 15.2	47.3	-148.7						
(2)15 22 46.47	11 17 56.0	109.1	-21.3							15 43 40	19 16.2	25	N	19.0	R	D	EF		
(3)15 22 41.70	11 16 17.9	39.0	-119.4							(1)15 43 42.99	19 16 22.7	31.6	36.7						
15 23 13.07	16 06 30.3	-13.6	-8.9	19.0	B	S(S)	EF			(2)15 43 40.09	19 16 35.3	-9.5	49.2						
(1)15 23 15.57	16 06 38.5	22.4	-6	19.5	R	(D)				(3)15 43 37.56	19 15 16.5	-45.4	-29.6						
(2)15 23 10.13	16 05 31.4	-56.0	-67.8							15 44 50	11 52.9	22	F	2					

TABLE 3—Continued

1	2	3	4	5	6	7	8	1	2	3	4	5	6	7	8
POSITION(1950.0)		OFFSETS		MAG	TYPE	ID		POSITION(1950.0)		OFFSETS		MAG	TYPE	ID	
α	δ	$\Delta\alpha$	$\Delta\delta$					α	δ	$\Delta\alpha$	$\Delta\delta$				
(3)16 ^h 39 ^m 58 ^s .09	19°08'53".3	-41".8	-3".0					(3)17 ^h 07 ^m 21 ^s .15	11°43'26".2	94".0	-47".4				
16 40 23.51	15 39 16.2	12.5	-13.7	20.0	R D	EF	17 08 43.93	19 34 05.5	18.2	-7.9	(19.5)	N		EF	
16 40 24.31	15 39 26.9	24.0	-3.0	19.0	R D		17 08 41.65	19 34 31.7	-13.9	18.3	16.5	N S(S)			
16 40 24.12	15 39 18.1	21.2	-11.8	19.0	B D(S)		(1)17 08 42.28	19 35 35.0	-5.1	81.6					
(1)16 40 26.33	15 40 56.5	53.1	86.6				(2)17 08 37.32	19 33 21.5	-75.2	-52.0					
(2)16 40 19.41	15 37 45.5	-46.7	-104.4				(3)17 08 45.90	19 32 40.5	46.2	-93.0					
(3)16 40 30.88	15 38 13.9	118.9	-76.0				17 10 11.55	15 39 40.2	.9	.3	18.0	N D(D)		G A 21	
16 41 34.53	17 21 20.5	-2.2	-.5	17.0	R D(D)	G A 46	(1)17 10 06.14	15 39 15.8	-77.2	-24.0					
16 41 34.40	17 21 11.3	-4.1	-9.7	20.0	R D		(2)17 10 17.92	15 39 22.6	93.0	-17.3					
16 41 34.14	17 21 35.8	-7.9	14.7	18.0	N S		(3)17 10 18.41	15 40 36.8	99.9	56.9					
(1)16 41 29.38	17 22 58.2	-76.0	97.1				17 11 59.43	19 52 36.6	-11.1	5.7	18.0	R S		EF	
(2)16 41 29.97	17 19 24.6	-67.6	-116.5				(1)17 11 56.71	19 51 42.8	-49.5	-48.0					
(3)16 41 23.70	17 22 05.5	-157.3	44.4				(2)17 12 03.31	19 53 52.2	43.6	81.4					
16 45 27.92	17 25 26.6	-1.5	-1.1	18.0	R S	G A 34	(3)17 12 04.17	19 50 45.8	55.8	-105.0					
16 45 28.88	17 25 30.3	12.2	2.6	18.0	B S		17 14 48.74	19 46 49.8	3.9	-23.8	19.0	N S(S)		EF	
(1)16 45 29.63	17 25 39.0	23.0	11.3				17 14 47.54	19 47 34.1	-13.0	20.6	14.0	N S(S)			
(2)16 45 28.40	17 25 57.6	5.4	29.9				(1)17 14 54.02	19 49 02.0	78.4	108.4					
(3)16 45 31.64	17 25 18.7	51.7	-9.0				(2)17 15 05.36	19 45 37.5	238.5	-96.0					
16 48 42.57	19 51 34.1	-1.1	-8.9	19.5	N D(D)	PG C 0	17 17 00.34	17 48 08.1	10.1	-4.2	18.5	B S(S)		PQ B 20	
16 48 42.73	19 51 29.3	1.2	-13.7	20.0	N D(D)		(1)17 17 02.80	17 47 21.1	45.2	-51.3					
16 48 44.13	19 51 45.0	20.9	2.1	19.0	R S		(2)17 17 07.40	17 48 27.2	111.0	14.8					
(1)16 48 36.97	19 51 41.7	-80.1	-1.3				(3)17 16 56.80	17 50 03.0	-87.0	110.7					
(2)16 48 39.48	19 50 16.0	-44.7	-7.0				17 17 48.51	13 00 41.8	11.2	7.1	19.5	R		PG C 0	
(3)16 48 42.66	19 54 23.5	.1	160.6				17 17 48.63	13 00 27.6	12.9	-7.0	18.5	N D(C)			
16 49 20.10	16 57 38.4	-12.9	-.8	20.0	R	EF	17 17 46	13 00 6	24	P	(19.5)	N (C)			
16 49 19.76	16 57 33.9	-17.7	-5.4	17.5	B S(S)		(1)17 17 51.22	13 00 49.8	50.8	15.2					
(1)16 49 23.38	16 58 04.2	34.2	25.0				(2)17 17 57.22	12 58 25.7	138.4	-128.9					
(2)16 49 28.97	16 57 25.0	114.4	-14.3				(3)17 18 01.06	12 59 30.3	194.5	-64.4					
(3)16 49 31.81	16 59 23.2	155.1	103.9				17 18 53.36	19 15 04.4	-3.1	17.1	19.5	R S		EF	
16 50 42.49	11 10 46.1	20.3	-.7	19.0	R	EF	(1)17 19 01.23	19 15 26.1	108.3	38.8					
(1)16 50 38.32	11 10 11.8	-41.1	-34.9				(2)17 19 01.53	19 12 54.2	112.7	-113.0					
(2)16 50 39.96	11 11 52.7	-16.9	65.9				(3)17 18 44.42	19 16 46.7	-129.7	119.4					
(3)16 50 46.62	11 10 45.8	81.1	-.9				17 20 59.02	17 38 17.5	-3.4	-3.9	19.5	R S		PG B 0	
16 52 19	13 51.8	16 S	18.5	R D	EF		17 21 00.38	17 38 04.3	16.1	-17.0	16.0	N S(S)			
16 52 18	13 52.0	17 P	19.5	B S(S)			(1)17 21 09.64	17 40 16.0	148.5	114.7					
16 52 20	13 52.0	19 F	(20.0)				(2)17 20 45.69	17 38 00.5	-193.9	-20.8					
(1)16 52 22.00	13 53 05.3	38.9	63.0				17 22 20.65	10 33 51.7	11.1	18.5	20.0	R		EF	
(2)16 52 13.84	13 52 17.4	-79.9	15.0				17 22 20.01	10 33 55.0	1.6	21.7	20.0	R			
(3)16 52 19.87	13 54 09.6	8.0	127.3				17 22 18	10 33 6	23 P	18.5	R S				
16 53 28.69	16 18 35.3	-9.7	-5.4	18.5	B S(S)	PQ C 0	(1)17 22 22.41	10 36 06.3	37.1	153.1					
16 53 28.51	16 18 46.6	-12.3	5.9	20.0	N		(2)17 22 08.25	10 33 54.3	-171.7	21.1					
16 53 29.35	16 18 30.7	-14.6	-10.0	18.0	R S(S)		17 25 21.53	16 47 04.5	-2.0	10.3	19.0	R D		EF *	
(1)16 53 33.99	16 19 15.0	66.6	34.3				17 25 22.41	16 46 52.5	10.7	-1.7	16.5	N S(S)			
(2)16 53 26.88	16 19 47.9	-35.8	67.2				(1)17 25 14.73	16 47 27.8	-99.6	33.6					
(3)16 53 33.66	16 17 35.9	61.9	-64.8				(2)17 25 34.53	16 47 24.2	184.7	30.1					
(1)16 53 51.89	18 18 26.6	-16.8	58.6				(3)17 25 34.23	16 48 56.2	180.4	122.0					
(2)16 53 47.26	18 18 47.4	-82.8	79.3				17 27 54.45	17 26 33.5	6.8	-3.8	20.0	N		PG B 0	
(3)16 53 59.87	18 18 42.7	96.9	74.7				17 27 55.08	17 26 42.2	15.7	4.9	18.5	R D			
16 58 22.37	14 52 59.9	-4.0	7.4	(20.0)		PG C 24 *	17 27 55.24	17 26 51.9	18.1	14.6	18.5	B (C)			
16 58 22.17	14 53 07.3	-6.9	14.9	18.5	D		(1)17 27 54.37	17 27 08.0	5.6	30.7					
16 58 23	14 52.6	20 SF	17.0	R S			(2)17 27 58.16	17 26 33.5	59.9	-3.8					
16 58 24	14 52.9	20 F	19.0	R S			(3)17 27 54.58	17 25 31.9	8.6	-65.4					
(1)16 58 17.81	14 52 21.6	-70.1	-30.9				17 29 43.18	15 05 00.5	-1.5	-3.0	18.5	R S(S)		PQ A 0	
(2)16 58 30.30	14 51 47.1	111.0	-65.3				17 29 43.59	15 04 57.7	4.4	-5.9	19.0	N S(S)			
(3)16 58 16.74	14 54 34.4	-85.5	101.9				17 29 42.85	15 04 59.2	-6.3	-4.4	18.0	N S(S)			
17 00 30.53	11 58 53.7	12.8	-5.6	18.5	R S	EF	17 29 42.96	15 04 49.6	-4.7	-14.0	18.5	R S(S)			
(1)17 00 30.41	11 58 02.3	11.1	-57.0				17 29 43.38	15 04 43.7	1.4	-19.9	17.0	N S(S)			
(2)17 00 29.26	12 00 15.5	-5.8	76.3				(1)17 29 37.64	15 05 05.1	-81.7	1.5					
(3)17 00 23.45	11 59 43.5	-91.1	44.2				(2)17 29 36.56	15 02 57.0	-97.3	-126.6					
17 00 41.10	18 02 55.0	-.7	.1	18.5	B S(S)	Q A 38	(3)17 29 44.00	15 02 14.6	10.5	-169.0					
(1)17 00 39.72	18 02 40.3	-20.4	-14.7				17 32 27.87	16 02 27.4	-3.0	.1	19.0	B S(S)		Q A 42	
(2)17 00 42.00	18 03 23.7	12.1	28.8				17 32 27.53	16 02 39.2	-7.9	11.9	18.5	R S			
(3)17 00 38.90	18 03 54.7	-32.2	59.7				17 32 28.92	16 02 12.1	12.1	-15.1	18.0	R S			
17 04 03.74	17 02 36.0	-1.2	-4.5	19.0	R D	PG A 38	17 32 26.83	16 02 37.2	-18.0	10.0	18.5	B S			
17 04 04.97	17 02 44.0	16.4	3.5	18.0	N C(S)		17 32 27.83	16 02 05.0	-3.7	-22.2	14.0	N S			
(1)17 04 09.28	17 02 25.7	78.2	-14.8				(1)17 32 30.58	16 00 25.7	36.0	-121.6					
(2)17 03 55.38	17 03 49.7	-121.2	69.1				(2)17 32 20.62	16 03 54.7	-107.5	87.4					
(3)17 03 53.40	17 03 21.0	-149.6	40.5				(3)17 32 28.97	16 05 26.3	12.9	179.0					
(1)17 05 51.62	18 21 12.8	10.3	44.0				17 36 42.02	16 08 16.9	-3.1	-.1	19.0	R D		PG A 0	
(2)17 05 54.25	18 20 05.1	47.8	-23.8				17 36 41.79	16 08 03.3	-6.4	-13.7	18.0	N S(S)			
(3)17 05 51.98	18 18 36.9	15.4	-112.0				17 36 41	16 08 3	17 P	20.0					
17 06 57.50	18 53 26.8	10.7	8.6	18.5	B S(S)	PQ C 0	(1)17 36 37.34	16 07 12.5	-70.5	-64.5					
(1)17 06 55.06	18 53 36.9	-24.0	18.7				(2)17 36 37.76	16 06 32.7	-64.5	-104.3					
(2)17 06 51.52	18 53 21.4	-74.2	3.2				(3)17 36 49.87	16 09 28.5	110.1	71.5					
(3)17 07 04.92	18 53 56.0	116.0	37.8				17 37 57.69	19 43 21.7	-18.8	16.4	15.0	B S(S)		EF	
17 07 14.73	11 44 09.9	-.2	-3.7	19.5	R D	PG A 0	(1)17 37 55.98	19 43 21.9	-42.8	16.6					

TABLE 3—Continued

1							2								
POSITION(1950.0)							POSITION(1950.0)								
OFFSETS							OFFSETS								
MAG TYPE ID							MAG TYPE ID								
(1)17 39 18.02	10 46 03.1	-30.7	22.6				17 58 09.5	11 05.9	20 SF	18.5 R S(C)					
(2)17 39 21.72	10 46 53.5	23.7	73.0				17 58 07	11 05.9	20 SF	18.0 N S					
(3)17 39 13.35	10 44 54.2	-99.6	-46.3				(1)17 58 10.05	11 05 42.3	18.0	-30.6					
17 39 24.10	17 21 42.6	-2.6	4.6	18.0 N D(D)	C G B 0*		(2)17 58 04.58	11 07 34.5	-62.6	81.7					
17 39 25.27	17 21 42.1	14.1	4.1	15.4 N D(D)			(3)17 58 13.13	11 08 01.6	63.4	108.8					
(1)17 39 29.84	17 22 01.6	79.4	23.6				17 59 21.38	13 51 19.4	-4.2	-3.4	(20.5)			PG A 0	
(2)17 39 23.95	17 23 13.5	-4.9	95.5				17 59 21.88	13 51 07.6	3.1	-15.3	18.5 B S(S)				
(3)17 39 25.30	17 23 18.4	14.4	100.4				17 59 21	13 51.1	20 S	18.0 R S					
17 40 36.36	16 14 09.5	-1.5	1.0	15.1 R D(D)	C G A 38		17 59 21.31	13 51 41.9	-5.2	19.1	18.0 N S(S)				
17 40 35.42	16 14 15.6	-15.0	7.2	18.5 R S(S)			17 59 22	13 51.7	24 NF	19.0 N D(D)					
17 40 35.17	16 13 56.4	-18.5	-12.0	19.5 R			(1)17 59 23.35	13 50 48.1	24.4	-34.8					
(1)17 40 37.28	16 13 45.8	11.8	-22.7				(2)17 59 17.43	13 51 22.9	-61.7	.1					
(2)17 40 32.96	16 14 56.5	-50.4	48.1				(3)17 59 20.98	13 52 31.9	-10.0	69.1					
(3)17 40 36.68	16 12 56.9	3.2	-71.6				17 59 25.20	12 27 59.0	-4	10.5	18.5 B S(S)			EF	
17 42 48.20	16 35 33.2	-13.8	-6.3	19.0 R S(S)	EF		17 59 25.53	12 27 38.2	4.5	-10.3	20.0 N				
17 42 47.89	16 35 33.9	-18.1	-5.5	19.0 B S(S)			17 59 26.33	12 27 52.3	16.2	3.8	18.5 R S(D)				
17 42 48.78	16 36 01.3	-5.4	21.8	17.0 N S(S)			17 59 24.05	12 27 47.9	-17.3	-6	18.5 B S(S)				
17 42 50	16 35.4	24 SF	18.5 B S				17 59 25.33	12 28 08.9	1.5	20.4	18.0 R S(S)				
(1)17 42 45.12	16 36 14.3	-58.0	34.9				17 59 26	12 27.5	22 SF	18.0 B S(S)					
(2)17 42 46.83	16 34 38.9	-33.4	-60.5				(1)17 59 22.87	12 27 51.2	-34.5	2.7					
(3)17 42 53.65	16 33 17.7	64.5	-141.7				(2)17 59 22.63	12 28 00.2	-38.0	11.7					
17 43 22.24	17 21 09.0	-4.8	-1.4	19.5 B (S)	PQ B 21		(3)17 59 29.88	12 28 29.0	68.1	40.5					
17 43 22.92	17 21 05.8	4.9	-9.5	19.5 B (S)			(1)17 59 44.62	18 10 40.0	-22.0	-41.3				EF	
17 43 23.05	17 21 22.4	6.8	12.0	16.5 N S(S)			(2)17 59 49.47	18 11 19.9	47.1	-1.4					
17 43 23.61	17 21 00.9	14.8	-9.4	19.0 N (S)			(3)17 59 43.71	18 12 25.8	-35.0	64.5					
17 43 21.69	17 20 51.4	-12.6	-18.9	17.0 N S(S)			18 04 42.58	15 42 58.7	4.5	-1.2	20.0 B			PG A 0	
17 43 23.57	17 20 51.5	14.3	-18.8	18.0 B S(S)			18 04 43.06	15 43 04.4	11.4	4.5	19.0 R D				
(1)17 43 17.61	17 20 53.0	-71.1	-17.3				18 04 41.59	15 42 51.5	-9.9	-8.4	(19.0)B (S)				
(2)17 43 22.39	17 22 36.6	-2.6	86.3				18 04 41.12	15 42 47.7	-16.6	-12.1	19.0 R S				
(3)17 43 17.98	17 20 01.7	-65.7	-68.6				18 04 42	15 43.4	22 N	19.5 N					
17 43 36.24	13 07 00.7	-1.6	2.2	19.5 R D	PG A 0		(1)18 04 40.84	15 42 42.9	-20.6	-16.9					
17 43 36.42	13 07 05.2	1.0	6.7	20.0 R			(2)18 04 41.21	15 42 31.5	-15.3	-28.3					
(1)17 43 34.59	13 07 02.1	-25.7	3.6				(3)18 04 43.11	15 43 35.6	12.2	35.7					
(2)17 43 36.81	13 06 32.0	6.7	-26.5				18 10 09.98	15 33 52.6	11.0	-1.5	19.0 N S(S)			EF	
(3)17 43 35.45	13 06 25.4	-13.2	-33.1				18 10 09.55	15 34 13.1	4.8	19.0	19.0 N D(S)				
17 44 36	11 28.4	1 SP	20.0 D(D)	PG A 0			18 10 10.12	15 33 38.7	13.1	-15.4	20.0 N				
17 44 35.94	11 28 36.2	-9.9	14.1	17.0 N S(S)			(1)18 10 08.57	15 32 57.2	-9.2	-56.9					
17 44 35.59	11 28 11.2	-15.0	-10.9	15.0 N S(S)			(2)18 10 11.91	15 32 47.9	39.0	-66.3					
(1)17 44 39.08	11 28 49.2	36.3	27.1				(3)18 10 19.92	15 33 50.0	154.7	-4.1					
(2)17 44 37.38	11 27 30.3	11.3	-51.8				18 10 45.36	17 05 14.7	-1.0	1.0	20.0 N			PG A 0	
(3)17 44 31.78	11 27 52.2	-71.0	-29.9				18 10 45.20	17 05 19.1	-3.3	5.3	18.0 R S(S)				
17 46 14.92	16 44 02.3	-9.0	-7.5	19.5 B (S)	PQ C 0*		18 10 45.67	17 05 05.8	3.5	-8.0	(19.5)B				
17 46 14.49	16 44 03.2	-15.3	-6.6	17.5 B S(S)			18 10 44.81	17 05 15.3	-8.9	1.5	16.0 N S(S)				
(1)17 46 11.30	16 43 51.9	-61.0	-17.9				(1)18 10 47.52	17 04 48.5	29.9	-25.2					
(2)17 46 16.57	16 46 14.1	14.7	124.3				(2)18 10 48.48	17 05 40.8	43.7	27.0					
(3)17 46 18.84	16 42 13.4	47.2	-116.4				(3)18 10 41.50	17 04 39.2	-56.4	-34.6					
17 51 59.26	15 21 03.6	-2.8	-3.0	20.0 N	PQ B 0*		18 16 34.80	18 39 08.2	-7	-9.1	19.5 R			EF	
17 51 59.76	15 21 10.6	4.5	4.0	19.0 B S(S)			18 16 33.77	18 39 10.5	-15.3	-6.8	18.5 R S(S)				
(1)17 52 03.44	15 21 01.5	57.8	-5.2				18 16 34.35	18 38 55.3	-7.1	-22.0	18.0 R S				
(2)17 51 53.64	15 20 11.1	-84.0	-45.5				(1)18 16 31.75	18 38 52.4	-44.1	-24.9					
(3)17 52 06.23	15 20 50.2	98.1	-16.4				(2)18 16 28.63	18 37 59.6	-88.4	-77.7					
17 54 18.98	15 54 03.9	-4.7	-2.5	19.0 B S(S)	PQ B 21		(3)18 16 42.90	18 38 49.3	114.4	-28.0					
17 54 18.16	15 54 17.4	-16.5	10.9	18.0 N S(S)			18 17 23.59	15 42 25.0	1.2	-6	20.0 N			PG A 0	
17 54 20.84	15 54 11.5	22.1	5.1	19.0 N (D)			18 17 23.24	15 42 24.5	-3.8	-1.0	20.0 B (D)				
17 54 17.63	15 54 08.2	-24.2	1.7	19.5 B S(S)			18 17 23.34	15 42 17.5	-2.3	-8.1	19.5 N				
(1)17 54 13.12	15 53 40.6	-89.3	-25.8				18 17 22.69	15 42 19.8	-11.7	-5.7	18.5 B S(S)				
(2)17 54 17.07	15 51 42.7	-32.2	-143.8				18 17 23.31	15 42 40.5	-2.8	15.0	18.5 N S(S)				
(3)17 54 08.77	15 53 07.5	-152.1	-59.0				18 17 22.32	15 42 14.8	-17.1	-10.7	19.0 N				
17 55 28.99	16 35 12.1	-4	-2.8	20.0 R	PG A 0		(1)18 17 21.78	15 42 11.2	-24.8	-14.3					
17 55 29.00	16 34 59.4	-2	-15.4	19.0 B S(S)			(2)18 17 19.27	15 42 05.3	-61.1	-20.2					
17 55 28.71	16 34 58.5	-4.4	-16.3	19.5 R S(D)			(3)18 17 21.76	15 41 20.5	-25.2	-65.1					
17 55 28.73	16 35 36.3	-4.1	21.4	18.5 B S(S)			18 20 08.39	17 58 38.6	-6.4	3.1	18.5 N S(S)			PQ B 0	
(1)17 55 25.58	16 35 02.6	-49.4	-12.3				18 20 08.52	17 58 29.9	-4.6	-5.6	18.5 N S(S)				
(2)17 55 27.94	16 34 18.9	-15.5	-56.0				18 20 09.21	17 58 43.1	5.2	7.7	18.5 N S(S)				
(3)17 55 34.31	16 35 14.0	76.1	-9				(1)18 20 09.92	17 59 24.7	15.5	49.2					
17 55 53.68	15 18 10.4	1.7	-1.6	18.5 N S(S)	PQ A 0		(2)18 20 02.39	17 59 18.8	-92.0	43.4					
17 55 53.84	15 18 05.8	4.0	-5.2	18.5 N S(S)			(3)18 19 58.06	17 57 45.0	-153.8	-50.5					
17 55 54.24	15 18 11.4	9.8	1.4	15.0 N S(S)			18 20 19.05	16 29 46.5	-6.8	9.8	17.0 N S(S)			EF	
17 55 53.69	15 18 27.1	1.8	15.1	19.5 N S			18 20 20.23	16 29 21.0	10.2	-15.6	16.0 N S(S)				
(1)17 55 56.15	15 20 05.2	37.4	113.2				18 20 20.73	16 29 29.0	17.4	-7.7	19.0 B S(S)				
(2)17 55 59.08	15 16 38.1	79.8	-93.9				18 20 18.17	16 29 32.2	-19.4	-4.5	17.0 N S(S)				
(3)17 55 49.71	15 20 18.0	-55.8	126.0				18 20 20.13	16 29 13.6	8.8	-23.0	18.5 N S(S)				
17 56 13.48	13 28 43.2	2.8	-3.2	19.5 R S	PG A 24		(1)18 20 23.89	16 29 55.5	62.9	18.9					
17 56 12.55	13 28 45.6	-10.8	-1.8	20.0 R			(2)18 20 16.57	16 28 20.9	-42.5	-75.7					
17 56 12.28	13 28 40.5	-14.8	-5.9	19.5 R			(3)18 20 16.98	16 31 25.9	-36.6	109.2					
17															

TABLE 3—Continued

1								2							
POSITION(1950.0)				OFFSETS		MAG TYPE		POSITION(1950.0)				OFFSETS		MAG TYPE	
1	2	3	4	$\Delta\alpha$	$\Delta\delta$	5	6	1	2	3	4	$\Delta\alpha$	$\Delta\delta$	5	6
18 22 41.72	12 46 36.6	-14.8	-2.7	18.0	R S(S)			21 05 49.32	03 21 08.5	-2.5	15.9	18.5	B S(S)	EF	
18 22 41.60	12 46 46.8	-16.5	7.5	18.0	B S(S)			(1)21 05 46.53	03 21 21.5	-39.3	28.9				
18 22 44.03	12 46 53.0	19.0	13.7	17.0	R S(S)			(2)21 05 51.82	03 21 54.1	39.9	61.6				
(1)18 22 43.51	12 47 25.7	11.4	46.4					(3)21 05 47.12	03 23 09.6	-30.4	137.0				
(2)18 22 38.10	12 45 25.6	-67.7	-73.7					(1)21 06 51.86	03 16 07.8	18.2	87.8			EF	
(3)18 22 33.64	12 47 34.4	-133.0	55.1					(2)21 06 46.12	03 15 57.7	-67.7	77.7				
								(3)21 06 54.05	03 12 51.0	51.0	-109.0				
18 23 06	16 15.5	3 NF	20.0			PG A 0		21 08 39.52	03 58 52.2	-11.9	9.4	19.5	R	EF	
18 23 05.67	16 15 32.4	-3.4	5.8	19.0	B S(S)			21 08 41.24	03 58 57.9	13.8	15.1	19.0	H S		
18 23 05.89	16 15 19.4	-3	-7.3	20.0				(1)21 08 44.03	03 58 11.4	55.6	-31.4				
18 23 05.06	16 15 26.3	-12.3	-4	18.0	N S(S)			(2)21 08 40.80	03 59 54.8	7.3	71.9				
18 23 06	16 15.6	13 NF	20.0					(3)21 08 34.93	03 57 11.3	-80.7	-91.6				
18 23 05.07	16 15 38.2	-12.1	11.5	19.0	R S			21 10 12.06	-01 46 28.6	-.7	5.0	19.0	B S(S)	Q B B	
18 23 07.24	16 15 34.2	19.2	7.6	19.5	N (S)			(1)21 10 14.37	-01 46 47.9	33.9	-14.2				
18 23 05.84	16 15 48.7	-1.0	22.0	19.5	B			(2)21 10 09.84	-01 47 57.4	-34.0	-83.8				
(1)18 23 02.60	16 15 22.3	-47.6	-4.4					(3)21 10 14.30	-01 44 37.4	32.9	116.2				
(2)18 23 10.09	16 16 10.3	60.3	43.6					21 11 04.68	04 33 50.4	11.7	-1.1	18.5	R S	EF	
(3)18 23 14.16	16 17 10.8	118.8	104.1					21 11 04.95	04 33 57.4	15.8	6.0	16.5	N S(S)		
18 28 55.99	17 41 39.5	-5.3	-7.3	20.0	B (S)	EF		(1)21 11 04.22	04 32 02.2	4.8	-109.2				
18 28 56.59	17 41 56.4	3.2	9.6	19.0	N S(S)			(2)21 11 13.62	04 33 12.4	145.4	-39.1				
18 28 55.31	17 41 52.4	-15.1	5.6	18.0	R S(S)			(3)21 11 01.39	04 36 37.8	-37.4	166.4				
18 28 55.19	17 41 56.2	-16.7	9.4	19.0	R S(S)			21 15 11.50	05 39 24.5	6.6	-11.4	18.5	N S	EF	
18 28 55.07	17 41 40.7	-18.6	-6.1	18.5	N S(S)			21 15 11	05 39.9	21 NF		19.0	R S		
(1)18 28 55.05	17 41 09.7	-18.8	-37.1					(1)21 15 08.72	05 40 26.5	-34.9	50.7				
(2)18 28 55.71	17 42 35.3	-9.3	48.5					(2)21 15 08.13	05 37 30.3	-43.7	-125.5				
(3)18 28 54.45	17 42 59.4	-27.4	72.6					(3)21 15 20.87	05 40 07.6	146.5	31.8				
20 34 03.98	02 46 50.5	6.1	19.4	16.0	N S(S)	EF		21 20 47.06	09 55 01.9	-2.7	2.4	19.0	B S(S)	Q A 41	
(1)20 34 04.68	02 46 50.2	18.1	19.1					21 20 47.64	09 55 00.8	5.8	1.4	18.5	R S(S)		
(2)20 34 01.29	02 46 32.6	-32.7	1.5					21 20 48	09 55.3	24 NF		18.5	N S		
(3)20 34 01.98	02 46 59.0	-22.3	27.9					(1)21 20 48.66	09 55 17.3	20.9	17.9				
20 34 24.54	03 59 29.1	16.2	.0	19.0	R	EF		(2)21 20 48.99	09 54 43.7	25.9	-15.7				
20 34 23.49	03 59 10.4	.6	-18.6	19.0	R			(3)21 20 44.77	09 55 24.6	-36.5	25.2				
(1)20 34 23.60	03 59 59.7	2.3	30.7					21 20 50	04 27.3	17 N	(19.0)B (S)	EF			
(2)20 34 15.68	03 59 30.6	-116.3	1.6					(1)21 20 50.89	04 27 26.9	1.5	23.5				
(3)20 34 21.60	03 57 35.8	-27.6	-113.3					21 22 55.63	05 15 42.9	-24.1	-6.2				
20 38 38.56	-01 22 26.0	2.0	-9.2	19.5	R	EF		(2)21 20 54.56	04 27 10.4	56.4	7.0				
20 38 38.40	-01 22 04.9	-4	11.8	(19.5)B				(3)21 20 49.90	04 25 55.0	-13.3	-68.4				
20 38 38.53	-01 22 35.5	1.6	-18.8	19.5	B			21 21 11.59	02 51 52.2	4.1	2.2	19.0	R S	PG B 0	
20 38 39.72	-01 22 16.5	19.4	.3	18.0	N S(S)			(1)21 21 09.59	02 51 23.5	-25.8	-26.5				
(1)20 38 44.30	-01 22 28.1	88.1	-11.3					(2)21 21 13.86	02 50 45.3	38.2	-64.6				
(2)20 38 32.03	-01 22 34.0	-95.9	-17.3					(3)21 21 16.17	02 52 33.7	72.8	43.8				
(3)20 38 43.31	-01 19 54.9	73.3	141.8					21 22 57	05 15.8	4 SP		20.0	R D	PG A 0	
20 42 15.85	04 33 06.4	2.3	-3.1	17.0	B S(S)	OB 41*		21 22 58.82	05 15 53.4	23.7	4.3	19.0	N S(S)		
20 42 14.77	04 33 05.4	-13.9	-4.1	19.5	B (S)			21 22 55.63	05 15 42.9	-24.1	-6.2	19.5	R		
20 42 16.64	04 32 56.8	14.1	-12.7	18.5	B S(S)			(1)21 22 56.41	05 14 47.1	-12.3	-62.0				
20 42 15.71	04 33 29.5	.3	20.0	18.0	B S(S)			(2)21 23 04.46	05 15 12.9	107.8	-36.2				
20 42 14	04 32.9	24 SP		19.5	N S(S)			(3)21 22 52.53	05 17 24.2	-70.4	95.1				
(1)20 42 13.68	04 33 44.6	-30.1	35.1					21 28 01.05	04 49 09.8	-22.4	5.8	18.5	R S(S)	EF	
(2)20 42 16.75	04 32 25.3	15.7	-44.2					(1)21 28 02.02	04 48 02.4	-8.0	-61.7				
(3)20 42 21.17	04 32 49.2	81.8	-20.3					(2)21 28 08.57	04 49 20.5	89.9	16.5				
20 42 51	03 13.7	9 NF	20.0	R		C PG C 0		(3)21 27 52.14	04 48 34.4	-155.6	-29.7				
20 42 50	03 13.7	12 NP	20.0	R				21 28 54.40	08 59 19.3	2.3	-.0	18.5	B S(S)	Q A 40	
20 42 50.90	03 13 18.4	-6.0	-14.8	19.0	R			21 28 54.89	08 59 20.8	9.6	1.4	19.0	N (S)		
20 42 52.46	03 13 38.0	17.4	4.9	20.0	B			(1)21 28 53.82	09 00 10.2	-6.3	50.8				
(1)20 42 49.38	03 13 26.2	-28.8	-7.0					(2)21 28 55.73	09 00 11.2	22.0	51.8				
(2)20 42 48.75	03 12 32.4	-38.1	-60.7					(3)21 28 54.23	08 57 34.7	-.1	-104.6				
(3)20 42 46.82	03 12 48.1	-67.1	-45.1					21 33 19.57	01 04 48.3	-1.2	2.9	19.0	B (S)	PQ A 10	
20 45 44.33	06 50 09.7	-.3	-1.1	18.0	R C(D)	G A 46		21 33 19.71	01 04 30.8	1.0	-14.6				
20 45 45.50	06 50 02.9	17.1	-7.9	17.5	R S(D)			(1)21 33 15.46	01 04 40.2	-62.9	-5.2				
(1)20 45 41.94	06 50 09.7	-35.9	-1.2					(2)21 33 10.14	01 05 18.5	-142.6	33.1				
(2)20 45 43.46	06 49 15.6	-13.2	-55.2					(3)21 33 28.90	01 03 20.7	138.8	-84.7				
(3)20 45 50.99	06 49 26.4	98.9	-44.4					21 34 05.21	00 28 25.0	2.2	-.9	17.0	B S(S)	Q A 36	
20 48 55.28	07 07 54.6	3.1	16.2	18.0	N S(S)	EF		(1)21 34 01.82	00 27 05.7	-48.7	-80.1				
20 48 55.42	07 08 02.2	5.1	23.8	19.0	R S			(2)21 34 11.44	00 27 15.6	95.7	-70.3				
(1)20 48 56.43	07 08 25.5	20.3	47.1					(3)21 33 57.62	00 29 12.5	-111.6	46.6				
(2)20 48 58.18	07 06 57.2	46.3	-41.2					21 35 30.72	03 45 18.0	-.2	-9.6	18.0	B S(S)	PQ B 38	
(3)20 48 57.21	07 08 40.5	31.8	62.1					(1)21 35 29.66	03 45 36.4	-16.1	8.7	19.0	R D		
20 50 48.99	05 26 12.7	12.0	9.5	20.0	R	EF		(2)21 35 33.85	03 46 25.8	46.6	58.2				
(1)20 50 45.96	05 26 24.2	-33.2	21.0					21 39 39.88	02 48 31.7	2.3	-13.2	19.0	R S	EF	
(2)20 50 41.39	05 25 15.0	-101.5	-48.2					(1)21 39 38.85	02 48 23.4	-13.2	-21.4				
(3)20 50 55.39	05 25 09.3	107.5	-53.8					(2)21 39 32.91	02 49 31.7	-102.2	46.8				
20 55 25	08 49.4	3 NP	20.0	R		PG A 0*		(3)21 39 37.34	02 50 37.9	-35.8	113.0				
20 55 25.98	08 49 18.5	6.0	-2.0	15.0	N S(S)			21 42 40.30	07 46 30.5	13.2	-1.0	20.0	R	EF	*
(1)20 55 28.85	08 50 07.5	48.7	47.0					21 42 40.36	07 46 33.8	14.0	2.3	20.0	R		
(2)20 55 22.10	08 48 09.9	-51.4	-70.6					21 42 39.12	07 46 51.9	-4.4	20.4	19.0	R		
(3)20 55 16.27	08 50 34.6	-137.8	74.1					21 42 40.88	07 46 22.7	21.8	-8.8	19.5	R		
20 56 01.13	05 31 03.8	22.0	-8.2	19.0	(S)	EF		(1)21 42 37.08	07 45 04.0	-34.8	-87.4				
(1)20 55 54.78	05 31 57.8	-72.6	45.7					(2)21 42 30.52	07 48 28.3	-132.2	116.8				
(2)20 56 05.29	05 30 48.7	84.3	-23.3					(3)21 42 51.42	07 44 49.4	178.4	-102.0				
(3)20 56 05.40	05 33 33.3	85.9	141.3					(1)21 42 48.77	04 17 37.9	26.7	-2.8			EF	*
20 58 38.58	01 54 37.3	.3	5.6	20.0	R	PQ B 0*									
20 58 38.61	01 54 43.5	.7	11.8	20.0	R										
20 58 39.36	01 54 52.7	12.0	20.9	17.0	B S(S)										
(1)20															

TABLE 3—Continued

1								2																															
POSITION (1950.0)				OFFSETS		MAG	TYPE	ID	POSITION (1950.0)				OFFSETS		MAG	TYPE	ID																						
h	m	s	°	Δ	δ				h	m	s	°	Δ	δ																									
22	52	22.58	02	09	21.6	21.9	-4.3	17.5	N	S(S)	EF	(1)23	32	38.67	-01	47	44.7	-115.9	2.1	(1)23	32	40.73	-01	49	07.9	-85.0	-81.1	(2)23	32	47.30	-01	45	08.9	13.5	157.9				
(1)22	52	21.08	02	08	21.4	-1.6	-60.4	(1)23	34	08.67	08	33	19.7	2.5	19.9	19.5	R	PG C 19	(2)23	34	07.65	08	33	18.9	-12.7	19.1	18.0	R	D										
(2)22	52	29.08	02	07	24.8	119.3	-117.1	(1)23	34	09.68	08	31	44.7	17.5	-75.1	(2)23	34	13.22	08	34	40.2	69.9	100.4	(3)23	33	58.78	08	35	17.2	-144.3	137.4								
(3)22	52	31.66	02	08	08.4	158.0	-73.6	23	35	34.23	03	10	11.6	-1.6	-1.1	18.0	B	S(S)	Q A 38	(1)23	35	35.67	03	09	11.3	19.9	-60.3	(2)23	35	29.22	03	09	58.5	-76.7	-13.2				
22	53	15	-00	36	3.3	23	F	19.5	R	D	EF	(3)23	35	39.73	03	09	58.1	80.7	-13.5	23	38	24.69	04	14	37.6	1.1	-1.0	19.5	B	(S)	Q A 7								
(1)22	53	12.07	-00	36	55.7	-22.2	-35.9	(1)23	38	26.44	04	14	46.9	27.2	8.3	(2)23	38	23.03	04	15	01.2	-23.8	22.6	(3)23	38	16.45	04	12	42.7	-122.1	-115.9								
(2)22	53	16.64	-00	36	51.8	46.3	-32.0	23	38	56.82	03	00	47.1	.6	.6	18.0	N	D(D)	G A 31	(1)23	38	53.97	03	00	18.5	-42.2	-28.0	(2)23	39	00.74	03	00	23.8	59.2	-22.7				
(3)22	53	21.69	-00	36	19.9	122.1	-1	(3)23	38	57.66	03	02	29.1	13.2	102.7	23	44	03.73	09	14	05.8	-2.5	.6	17.0	B	S(S)	Q A 4												
(1)22	55	08.93	09	16	59.5	-36.9	7.1	(1)23	44	04.57	09	16	38.2	9.9	153.0	(2)23	43	49.87	09	12	55.7	-207.7	-69.5	(3)23	44	19.17	09	13	16.4	226.1	-48.7								
(2)22	55	16.07	09	16	20.2	68.8	-32.2	(2)23	45	05.10	03	53	49.7	-55.2	-65.1	(3)23	45	11.36	03	57	07.2	38.4	132.4	23	45	58.31	04	50	42.1	-2.1	-11.4	16.0	B	S(S)	PQ C 0				
(3)22	55	06.64	09	19	24.3	-70.9	151.9	(3)23	45	01.33	04	48	27.7	43.1	-145.9	(2)23	46	06.95	04	53	29.6	127.1	156.0	23	45	58.39	06	08	19.1	-2.1	-1.7	18.0	B	S(S)	Q A 41				
22	57	44.80	07	54	30.6	.6	.2	17.0	R	D	G A 22	(1)23	46	00.43	06	08	25.9	28.4	5.2	(2)23	45	54.17	06	08	44.0	-65.0	23.3	(3)23	45	54.99	06	07	18.3	-52.8	-62.5				
(1)22	57	44.98	07	54	22.4	3.3	-8.1	19.0	N	S(S)	(1)23	47	36.09	06	42	15.9	72.9	-93.2	(2)23	47	30.82	06	41	48.8	-5.7	-120.3	(3)23	47	34.96	06	46	09.4	56.1	140.3					
(2)22	57	38.89	07	52	15.8	-87.1	-134.7	23	49	22.26	-01	25	54.9	-1.9	2.6	16.0	B	S(S)	PQ A 6*	23	49	23.24	-01	26	01.9	12.7	-4.4	18.5	R	C(D)									
(3)22	57	32.64	07	54	39.6	-180.0	9.1	(1)23	49	31.67	-01	24	22.6	139.1	94.9	(2)23	49	10.07	-01	25	17.8	-184.8	39.7	(1)23	50	21.10	05	45	19.5	-1	88.4	(2)23	50	15.17	05	43	55.1	-88.6	4.0
22	59	03.25	08	21	48.1	-2.2	-1.6	20.0	R	PG A 0	(3)23	50	36.05	05	43	14.0	223.1	-37.1	23	53	02.29	-00	19	28.2	5.4	-0	19.0	R	C	PG B 38									
(1)22	59	03.96	08	22	05.0	-8.3	15.3	20.0	R	(1)23	53	04.74	-00	20	54.6	42.3	-86.4	(2)23	52	59.93	-00	16	54.4	-29.8	153.7	(3)23	52	53.32	-00	21	09.4	-129.1	-101.2						
(2)22	59	07.02	08	19	34.3	53.8	-135.4	(2)23	54	00.44	08	46	24.6	-85.5	51.1	(3)23	54	11.39	08	46	42.3	76.9	68.8	(1)23	54	14.18	08	43	20.9	118.2	-132.7								
(3)22	59	13.26	08	22	56.8	146.3	67.1	(3)23	54	14.18	08	43	20.9	118.2	-132.7	23	55	50.01	-01	01	23.8	-20.3	11.8	(1)23	55	56.03	-01	00	56.0	70.0	39.7								
(3)22	58	48.30	08	22	09.1	-224.1	19.4	(2)23	55	50.18	-00	59	57.0	-17.8	98.6	(3)23	55	46.48	-01	03	30.0	-73.3	-114.3	(1)23	56	08.42	03	19	54.5	-10.6	-22.9								
23	03	11.73	-00	52	22.2	-1.4	-3.0	18.5	R	G A 37	(2)23	56	07.13	03	21	15.1	-29.9	57.7	(3)23	56	04.45	03	19	50.1	-70.0	-27.4													
(1)23	03	11.68	-00	52	02.0	-2.2	17.2	20.0	R	23	55	08.42	03	19	54.5	-10.6	-22.9	23	56	07.13	03	21	15.1	-29.9	57.7	23	56	04.45	03	19	50.1	-70.0	-27.4						
(2)23	03	11.73	-00	51	39.7	-1.5	39.5	(1)23	45	04.97	03	55	01.0	-57.2	6.2	(2)23	45	05.10	03	53	49.7	-55.2	-65.1	(3)23	45	11.36	03	57	07.2	38.4	132.4								
(3)23	03	07.42	-00	53	59.3	-66.1	-100.1	(2)23	45	05.10	03	53	49.7	-55.2	-65.1	23	45	58.31	04	50	42.1	-2.1	-11.4	(1)23	46	01.33	04	48	27.7	43.1	-145.9								
(3)23	03	14.20	-00	49	34.0	35.6	165.2	(3)23	45	01.36	03	57	07.2	38.4	132.4	(2)23	46	06.95	04	53	29.6	127.1	156.0	23	45	58.39	06	08	19.1	-2.1	-1.7								
(1)23	04	13.10	00	40	47.5	62.8	11.9	(3)23	45	01.36	03	57	07.2	38.4	132.4	(1)23	46	00.43	06	08	25.9	28.4	5.2	(2)23	45	54.17	06	08	44.0	-65.0	23.3								
(2)23	04	14.75	00	40	54.6	87.5	18.9	23	45	58.31	04	50	42.1	-2.1	-11.4	(3)23	45	54.99	06	07	18.3	-52.8	-62.5	(1)23	47	36.09	06	42	15.9	72.9	-93.2								
23	04	53.66	03	50	11.6	2.5	.7	18.0	N	S(S)	G A 0*	(2)23	47	30.82	06	41	48.8	-5.7	-120.3	(3)23	47	34.96	06	46	09.4	56.1	140.3												
23	04	54.34	03	50	17.2	12.8	6.3	20.0	R	23	49	22.26	-01	25	54.9	-1.9	2.6	23	49	23.24	-01	26	01.9	12.7	-4.4	(1)23	49	31.67	-01	24	22.6	139.1	94.9						
23	04	52.09	03	50	16.1	-20.9	5.2	19.0	N	(1)23	49	31.67	-01	24	22.6	139.1	94.9	(2)23	49	10.07	-01	25	17.8	-184.8	39.7	(2)23	49	10.07	-01	25	17.8	-184.8	39.7						
(1)23	04	57.06	03	49	31.5	53.5	-39.4	(1)23	49	10.07	-01	25	17.8	-184.8	39.7	(1)23	50	21.10	05	45	19.5	-1	88.4	(2)23	50	15.17	05	43	55.1	-88.6	4.0								
(2)23	04	48.49	03	50	19.4	-74.8	8.5	(2)23	49	10.07	-01	25	17.8	-184.8	39.7	(3)23	50	36.05	05	43	14.0	223.1	-37.1	23	53	02.29	-00	19	28.2	5.4	-0								
(3)23	05	00.85	03	49	58.6	110.2	-12.4	(3)23	49	10.07	-01	25	17.8	-184.8	39.7	(1)23	53	04.74	-00	20	54.6	42.3	-86.4	(2)23	52	59.93	-00	16	54.4	-29.8	153.7								
23	05	43.01	02	12	48.8	-7.6	12.9	20.0	N	EF	(3)23	52	53.32	-00	21	09.4	-129.1	-101.2	(1)23	54	00.44	08	46	24.6	-85.5	51.1	(3)23	54	11.39	08	46	42.3	76.9	68.8					
(1)23	05	46.91	02	13	07.5	51.0	31.6	(1)23	54	00.44	08	46	24.6	-85.5	51.1	(3)23	54	14.18	08	43	20.9	118.2	-132.7	23	55	50.01	-01	01	23.8	-20.3	11.8								
(2)23	05	39.98	02	13	53.2	-52.9	77.4	(2)23	54	01.33	04	48	27.7	43.1	-145.9	(2)23	55	50.18	-00	59	57.0	-17.8	98.6	(3)23	55	46.48	-01	03	30.0	-73.3	-114.3								
(3)23	05	43.86	02	10	08.3	5.3	-147.5	(3)23	54	14.18	08	43	20.9	118.2	-132.7	(1)23	56	08.42	03	19	54.5	-10.6	-22.9	(2)23	56	07.13	03	21	15.1	-29.9	57.7								
(1)23	05	58.26	03	21	37.9	91.6	48.8	23	49	22.26	-01	25	54.9	-1.9	2.6	(3)23	56	04.45	03	19	50.1	-70.0	-27.4	23	55	08.42	03	19	54.5	-10.6	-22.9								
(2)23	05	49.44	03	18	52.6	-40.4	-116.5	(1)23	49	31.67	-01	24	22.6	139.1	94.9	(1)23	56	07.13	03	21	15.1	-29.9	57.7	(2)23	56	04.45	03	19	50.1	-70.0	-27.4								
(3)23	05	41.66	03	19	50.5	-157.0	-58.7	(2)23	49	10.07	-01	25	17.8	-184.8	39.7	(1)23	56	08.42	03	19	54.5	-10.6	-22.9	(2)23	56	07.13	03	21	15.1	-29.9	57.7								
(1)23	08	27.61	01	45	01.0	150.9	-6.2	(1)23	49	10.07	-01	25	17.8	-184.8	39.7	(3)23	56	04.45	03	19	50.1	-70.0	-27.4	(1)23	56	08.42	03	19	54.5	-10.6	-22.9								
(2)23	08	25.84	01	46	46.9	124.2	99.7	(2)23	49	10.07	-01	25	17.8	-184.8	39.7	(2)23	56	07.13	03	21	15.1	-29.9	57.7	(3)23	56	04.45	03	19	50.1	-70.0	-27.4								
(3)23	08	08.38	01	43	42.7	-137.5	-84.5	(3)23	49	10.07	-01	25	17.8	-184.8	39.7	(1)23	56	08.42	03	19	54.5	-10.6	-22.9	(2)23	56	0													

REFERENCES FOR TABLE 3

- 1 BALDWIN, J.A., BURBRIDGE, E.M., HAZARD, C., MURDOCH, H.S., ROBINSON, L.B. AND WAMPLER, E.J. 1973, *ASTROPHYS. J.* 185, 739.
- 2 BARBIERI, C. AND BERTOLA, F. 1972, *MONTHLY NOTICES ROY. ASTRON. SOC.* 156, 399.
- 3 BLAKE, G.M. 1970, *ASTROPHYS. LETTERS* 6, 201.
- 4 BOLTON, J.G. AND EKERS, J. 1966, *AUSTRALIAN J. PHYS.* 19, 471.
- 5 BOLTON, J.G. AND EKERS, J. 1966, *AUSTRALIAN J. PHYS.* 19, 559.
- 6 BOLTON, J.G. AND EKERS, J. 1966, *AUSTRALIAN J. PHYS.* 19, 713.
- 7 BOLTON, J.G., SHIMMINS, A.J. AND MERKELIJS, J. 1968, *AUSTRALIAN J. PHYS.* 21, 81.
- 8 BOLTON, J.G. AND WALL, J.V. 1970, *AUSTRALIAN J. PHYS.* 23, 789.
- 9 BROWNE, I.W.A. 1971, *NATURE* 231, 515.
- 10 BROWNE, I.W.A., BENTLEY, M., HAVES, P., MCEWAN, N.J., SPENCER, R.E. AND STANNARD, D. 1974, *NATURE* 252, 209.
- 11 BROWNE, I.W.A., CROWTHER, J.H. AND ADGIE, R.L. 1973, *NATURE* 244, 146.
- 12 BROWNE, I.W.A. AND MCEWAN, N.J. 1973, *MONTHLY NOTICES ROY. ASTRON. SOC.* 162, 21P.
- 13 CLARKE, M.E., BOLTON, J.G. AND SHIMMINS, A.J. 1966, *AUSTRALIAN J. PHYS.* 19, 375.
- 14 COLLA, G., FANTI, C., FANTI, R., GIOIA, I., LARI, C., LEQUEUX, J., LUCAS, R. AND ULRICH, M.H. 1975, *ASTRON. ASTROPHYS. SUPPL.* 20, 1.
- 15 FANTI, C., FANTI, R., FICARRA, A., FORMIGGINI, L., GIOVANNI, G., LARI, C. AND PADRIELLI, L. 1975, *NATURE ASTROPHYS. SUPPL.* 19, 143.
- 16 GRUEFF, G. AND VIGOTTI, M. 1972, *ASTRON. ASTROPHYS. SUPPL.* 6, 1.
- 17 GRUEFF, G. AND VIGOTTI, M. 1975, *ASTRON. ASTROPHYS. SUPPL.* 20, 57.
- 18 HAZARD, C., GULKIS, S. AND SUTTON, J. 1968, *ASTROPHYS. J.* 154, 413.
- 19 HAZARD, C., JAUNCEY, D.L. AND BACKER, D.C. 1970, *ASTRON. J.* 75, 1039.
- 20 HOSKINS, D.G., MURDOCH, H.S., ADGIE, P.L., CROWTHER, J.H. AND GENT, H. 1974, *MONTHLY NOTICES ROY. ASTRON. SOC.* 166, 235.
- 21 HOSKINS, D.G., MURDOCH, H.S., HAZARD, C. AND JAUNCEY, D.L. 1972, *AUSTRALIAN J. PHYS.* 25, 559.
- 22 HUNSTEAD, R.W. AND JAUNCEY, D.L. 1970, *MONTHLY NOTICES ROY. ASTRON. SOC.* 149, 91.
- 23 JAUNCEY, D.L. AND HAZARD, C. 1970, *ASTROPHYS. LETTERS* 7, 1.
- 24 JOHNSON, K.H. 1974, *ASTRON. J.* 79, 1006.
- 25 KRISTIAN, J. AND SANDAGE, A. 1970, *ASTROPHYS. J.* 162, 391.
- 26 KRISTIAN, J., SANDAGE, A. AND KATEM, B. 1974, *ASTROPHYS. J.* 191, 43.
- 27 LANG, K.R., SUTTON, J., HAZARD, C. AND GULKIS, S. 1970, *ASTROPHYS. J.* 160, 17.
- 28 MCEWAN, N.J., BROWNE, I.W.A. AND CROWTHER, J.H. 1975, *MEM. ROY. ASTRON. SOC.* 80, 1.
- 29 MERKELIJS, J.K. 1968, *AUSTRALIAN J. PHYS.* 21, 903.
- 30 MERKELIJS, J.K., SHIMMINS, A.J. AND BOLTON, J.G. 1968, *AUSTRALIAN J. PHYS.* 21, 523.
- 31 MERKELIJS, J.K. AND WALL, J.V. 1970, *AUSTRALIAN J. PHYS.* 23, 575.
- 32 OLSEN, E.T. 1970, *ASTRON. J.* 75, 764.
- 33 PENSTON, M.J., PENSTON, M.V. AND SANDAGE, A. 1971, *PUBL. ASTRON. SOC. PACIFIC*, 83, 783.
- 34 RADIVICH, M.M. AND KRAUS, J.D. 1971, *ASTRON. J.* 76, 683.
- 35 SHIMMINS, A.J., BOLTON, J.G. AND WALL, J.V. 1975, *AUSTRALIAN J. PHYS. ASTROPHYS. SUPPL.* 34, 63.
- 36 SHIMMINS, A.J., SEARLE, L., ANDREW, B.H. AND BRANDIE, G.W. 1968, *ASTROPHYS. LETTERS* 1, 167.
- 37 WALL, J.V. 1971, *AUSTRALIAN J. PHYS. ASTROPHYS. SUPPL.* 20, 1.
- 38 WILLS, B.J. 1976, *ASTRON. J.* 81, 1031.
- 39 WILLS, B.J., WILLS, D. AND DOUGLAS, J.N. 1973, *ASTRON. J.* 78, 521.
- 40 WILLS, D. AND BOLTON, J.G. 1969, *AUSTRALIAN J. PHYS.* 22, 775.
- 41 WILLS, D. AND WILLS, B.J. 1974, *ASTROPHYS. J.* 190, 271.
- 42 WILLS, D. AND WILLS, B.J. 1976, *ASTROPHYS. J. SUPPL. SER.*, 31, 143.
- 43 WILLSON, M.A.G. 1970, *MONTHLY NOTICES ROY. ASTRON. SOC.* 151, 1.
- 44 WILLSON, M.A.G. 1972, *MONTHLY NOTICES ROY. ASTRON. SOC.* 156, 7.
- 45 WYNDHAM, J.D. 1965, *ASTRON. J.* 70, 384.
- 46 WYNDHAM, J.D. 1966, *ASTROPHYS. J.* 144, 459.

b) Hand-measured Positions

Some faint objects within the 25" radius circles were missed during the plate measuring with the two-coordinate machine. To provide finding charts, all fields were photographed with a camera mounted on the measuring machine at two enlarged scales, 0".62 mm⁻¹ and 4".8 mm⁻¹. The photographs were examined to verify the objects measured on the machine and to see if any objects had been missed. Likely identifications which had been missed were measured on the machine. For other objects which had been missed in the first measuring session, crude positions ($\pm 1''$ to $\pm 2''$) were measured on the 0".62 mm⁻¹ prints using a plastic millimeter scale. Only the distance from the radio source and the octant were recorded, because only the radial distribution of objects was to be considered in the analysis in § VIII.

c) Magnitudes, Colors, and Morphological Types

Magnitudes (approximately m_v), colors, and morphological types were estimated by eye for all objects from the 0".62 mm⁻¹ and 4".8 mm⁻¹ photographs of both color PSS plates. For stellar objects visible on the red plates the visual magnitudes are accurate to about ± 0.5 to ± 1.0 mag. Magnitudes estimated from the blue plates are less accurate, as are magnitudes of galaxies. For Zwicky (1961–1968) galaxies, Zwicky's magnitude is listed in Table 3.

A color class of red (R), neutral (N), or blue (B) was assigned by comparing the relative sizes of the red-

plate and blue-plate images of an object, taking into account zero-point shifts between plates by comparisons of several nearby stars on the two plates.

Morphological classes of stellar (S), compact (C), or diffuse (D) were assigned, using the photographs. Obvious galaxies were called diffuse. In some cases the red-plate morphology was different from that on the blue-plate, so both are given in Table 3. Although colors and morphologies are given for virtually all objects in the table, these classes are probably meaningless for objects with $m_v \geq 19.5$.

d) Identification Method

An identification was suggested whenever an object occurred within 10" of a radio source. If the nearest object was further than 10", other criteria were used. Stellar objects further than 10" were usually not suggested unless bright or blue. Bright galaxies were usually taken to be the identification. Faint galaxies further than 10" were considered if the radio angular extent appeared large. If no nearby object appeared to be a likely identification, the source is listed as an empty field (EF). However, in many cases listed as EF, there are objects within 25". Future work on the brightness distribution of these sources may show that some of the EFs have visible identifications.

Quasars and BL Lacertae types with confirming spectra are labeled Q. Definite stellar objects without known spectra are "PQ" (probable quasars). Faint objects of uncertain morphology are called "PQ" if

they appear bluish. Galaxies clearly discernible as such are labeled "G." Faint identifications which do not appear very stellar or bluish are called "PG" (probable galaxy). A few very faint objects are listed simply as "O" (object).

e) Explanation of the Optical Data Table

Optical positions, magnitudes, colors, and morphological types are given in Table 3 for all optical objects found within 25" of 664 radio sources. If more than one object occurs within 25" they are listed in order of increasing radio-optical displacement. The identification is generally the closest object. If not, the situation is explained in the notes (§ Vg).

Identification charts, reproduced from the PSS plates, are given in Figure 2 (Plates 29–36) for all suggested identifications having no previously published chart. In a few cases of disagreement with a previously published identification, a chart showing the new identification is given. No charts are given for empty fields.

Secondary position standards are listed for each field. Their positions are preceded by a number in parentheses which identifies them on the chart.

References to previously published charts are given in the References to Table 3. The notes to the table (§ Vg) mention particularly interesting objects, disagreements with previous identifications, and additional references. More references may be found in the compilation by Véron and Véron (1974) and its updates (privately circulated).

The key to Table 3 is as follows:

Columns (1) and (2): Optical right ascension and declination (1950.0). A number in parentheses preceding a position is the index number of a secondary position standard which identifies it on the chart, if one appears in Figure 2.

Column 3: The offsets $\Delta\alpha$ and $\Delta\delta$ (both in arc seconds) from the radio positions (given in Table 2; Douglas *et al.* [1973], Ghigo and Owen [1973], or Sharp and Bash [1975]), in the sense optical minus radio. If the optical position of an object was not determined by the least-squares fit, but from hand measurements on the 0".62 mm⁻¹ prints, the approximate position is listed in columns (1) and (2); column (3) is the radial distance to the radio source in arc seconds (rms accuracy $\pm 1''$ to $\pm 2''$) and the octant in which the optical object lies with respect to the radio source (N = north, S = south, P = preceding, F = following).

Column (4): m_v , estimates from the red PSS plate, accuracy about ± 0.5 to ± 1.0 mag (see § Vc). If enclosed in parentheses, estimated from the blue plate.

Column (5): A color estimate, R = red, N = neutral, B = blue, and morphological types S = stellar, C = compact, and D = diffuse. The red plate morphological type is given first, followed by the blue plate type in parentheses.

Column (6): A code for the most probable identification (see § Vd). The codes are:

G, galaxy;

Q, quasar or BL Lacertae-type, confirmed by optical spectra;

PG, probable galaxy;

PQ, probable quasar or BL Lacertae-type object;

O, faint object;

> G, probable galaxy not visible on the PSS plates, but found by someone else using deeper plates;

> Q, as above, but a possible quasar or BL Lacertae-type object;

EF, empty field;

OB, bright star obscures radio position;

C, cluster or galaxy group in the vicinity (three or more galaxies within $\sim 2''$) (usually precedes one of the previous codes, e.g., C PG).

Column (7): The reliability class of the identification, based on position agreement. The object's apparent magnitude is not taken into account here, only the radio-optical position difference (r). See § VIII and Table 7.

A, $r \leq 5''$ (reliability is $\geq 90\%$);

B, $5'' < r \leq 10''$ (reliability is 70% to 90%);

C, $r > 10''$ (reliability is $< 70\%$).

Column (8): The reference to a finding chart. The key to the reference number appears below Table 3. If zero appears in this column, a chart is given in Figure 2. If an asterisk appears in this column there is a note about this object in § Vg.

f) Finding Charts

Charts are given in Figure 2 for the 118 objects listed with a zero in column (8) of Table 3. In the chart labels, "B" means the chart was copied from the blue, or "O," PSS plate, and "R" means it was copied from the red, or "E," plate. The reticle lines mark the radio source position. They do not necessarily mark a visible object. Index numbers of secondary position standards are indicated. Lines are drawn to indicate secondary position standards when there might be some doubt as to which object is meant.

g) Comments on Individual Identifications

The literature search on which these notes are based is reasonably complete to mid-1976.

0038+097 3C 18.—The 18.5 mag galaxy 24" N was suggested by Wyndham (1966). Since Fomalont and Moffet (1971) give an angular size of 17" EW by 41" NS, none of the three galaxies within 25" of the radio source can be ruled out.

0040+064.—Galaxy suggested by Merkelijn (1969) is 50" N. There is a 20th mag probable galaxy 3" from the radio source position, also found by Wills (1976).

0043+000.—Galaxy given by Merkelijn and Wall (1970) is 30" P and 15" S of the UTRAO radio position.

0055+063.—A 16th mag star 4" from radio source position has a stellar spectrum (Wills and Wills 1976), and may be obscuring the true identification.

0059+017.—A very faint image occurs on the red plate within 3" of the radio source. This may well be the identification.

0103+062.—The 17th mag stellar object has a diffuse jet extending NF, at similar position angle to radio structure double model fitted by Cotton, Owen, and Ghigo (1975).

0114+074.—This source is not the object suggested by Clarke, Bolton, and Shimmins (1966) and found to have $z = 0.861$ by Bolton (Agnew and Arp 1973).

0115-016.—The UTRAO radio position is 35" SF the galaxy suggested by Bolton and Ekers (1967). McEwan, Browne, and Crowther (1975) also find EF.

0118-001.—A 13.5 mag object found to have stellar spectrum (Wills and Wills 1974; Strittmatter *et al.* 1974).

0118+034.—The quasar has $z = 0.765$ (Bolton and Wall 1970). McEwan, Browne, and Crowther (1975) find an offset of 2"8 P and 5"2 S, in disagreement with our radio-optical offsets. Macdonald and Miley (1971) give complex structure, 35" diameter.

0128+040.—Fomalont (1971) gives 359" double, P.A. = 31°. A 19th mag galaxy 11" P, 12" N may be the identification.

0805+269.—Recent observations by the author with the NRAO² interferometer show that the Ghigo and Owen (1973) position is lobe-shifted. Thus the information listed for this source in Table 3 is irrelevant. Two sources were found in this vicinity. One is a 20" double at P.A. = 160° and having $S_{2695} = 122(12)$ mJy. Its centroid position (epoch 1950.0) is 08^h05^m01^s81(0^s04), +27°01'28"8(0"8). The identification is a 19th mag probable galaxy within 1" ± 1" of the radio centroid, in a faint cluster. The other source appears to be single, with $S_{2695} = 168(11)$ mJy. Its position (epoch 1950.0) is 08^h05^m34^s28(0^s02), +26°55'24"3(0"2). The identification seems to be a 20th mag object, but I have not measured its optical position.

0807+276.—The BSO ~ 35" NF the position was suggested by Grueff and Vigotti (1973) and Fanti *et al.* (1975).

0817+307.—"Dumbbell" galaxy. The radio position coincides with the NF component. Position angle of galaxy and radio structure model position angle (Cotton, Owen, and Ghigo 1975) coincide (both position angles are about 60°).

0827+235.—A 20th mag object suggested by Hazard, Gulkis, and Bray (1967) is about 35" NP the UTRAO radio position. There may be a very faint (> 20 mag) object within 3" of our position, visible only on the red plate.

0828+324 and 0828+325.—These seem to have large radio angular structure. They are probably two components of the 16th mag E galaxy about 1' F 0828+324. Gearhart *et al.* (1972) suggested a BSO 14" S of the E galaxy.

0829+280.—Wills and Wills (1976) find an M-type stellar spectrum for this object. It may be obscuring the true identification.

0838+325.—Cotton, Owen, and Ghigo (1975) give structure as 15" double at P.A. = 170° ± 50°. Major axis of galaxy is at P.A. ≈ 140°.

² The National Radio Astronomy Observatory is operated by Associated Universities, Inc., under contract with the National Science Foundation.

0840+299.—Olsen (1970) gives 18.5 mag object 50" S, shown by Schmidt (1974) to have a stellar spectrum.

0844+319.—Nearby object suggested by Olsen (1970) shown to be a quasar by Strittmatter *et al.* (1974) and Schmidt (1974). Westerbork radio structure (Colla *et al.* 1975) shows 15.5 mag E galaxy is the identification, not the quasar.

0907+245.—Four faint galaxies in the area. Hazard and Jauncey (1972) suggest the galaxy 23" SF.

0922+322.—An 18th mag BSO 16" N, suggested by Olsen (1970), shown to have stellar spectrum by Schmidt (1974).

0923+257.—An 8.5 mag star blots field. Wills and Wills (1974) found this object to have a stellar absorption spectrum.

0927+252.—Spectrum by Wills and Wills (1976) is inconclusive, but probably is that of a quasar.

0932+241A.—Source is 0.6 cluster radii from the center of an ED Zwicky cluster.

0932+241B.—Wills (1976) suggests the RSO 9" NF, but there is a 20-21 mag slightly blue object 7" SP which may be the identification.

0943+257.—A Zwicky galaxy, 15.6 mag, is 36" F, which may be the correct identification, since Cotton, Owen, and Ghigo (1975) find a 60" double model at P.A. = 50°.

0949+246.—Very faint object 6" NF the NF component of double object noted by Bolton, Shimmins, and Merkelijn (1968) is about 2" NF the radio position. Since it lies on a diffraction ring of a nearby bright star, its reality is in doubt. This object is not listed in Table 3, but the two objects noted by Bolton, Shimmins, and Merkelijn (1968) are.

0949+287.—The declination given by Grueff and Vigotti (1975) appears to contain a typographical error: They give 43 for the minutes of declination in their Table 2, where I found 42.

0958+256.—Low-frequency variable: Cotton (1976) found that the 365 MHz flux density decreased from 1.8 Jy to 1.2 Jy during the latter half of 1973.

0958+290, 3C 234.—Riley and Pooley (1975) find three components of this source. The UTRAO radio position is 15" F the extreme SP component. A 17.5 mag galaxy in a group coincides with the central component.

1047+280.—There is a 16th mag stellar object and an 18-19 mag galaxy near the NF lobe-shift position.

1055+314.—The radio position may be affected by resolution. The right ascension is one lobe P the B2 position, while the declinations agree (not an allowed lobe-shift from the B2 position).

1057+307.—The galaxy suggested by Grueff and Vigotti (1975) is 40" NP the UTRAO radio position. The B2 position coincides with the galaxy. The UTRAO position is probably affected by resolution.

1106+252.—Called an empty field by Wyndham (1966) and Kristian, Sandage, and Katem (1974). The galaxy 16" N is a possible identification since Cotton, Owen, and Ghigo (1975) find this to be a 30"-40" double source, P.A. = 45° ± 5°.

1118+237.—The wrong object was marked on the

finding chart given by Kristian and Sandage (1970). The correct object is barely visible on the PSS plates, and is at the radio position (see the erratum in *Ap. J.*, **205**, 308 [1976]).

1125 + 260.—Group of galaxies in this region. NGC 3689 is about 8' S, noted previously by Caswell and Wills (1967) and Aizu (1966).

1133 + 262.—Wills and Wills (1974) find this object to have a stellar absorption spectrum.

1202 + 297.—The UTRAO radio position is $\sim 1'$ SP the galaxy suggested by Grueff and Vigotti (1972).

1202 + 267.—There is an 18th mag galaxy near the SP lobe-shift position.

1222 + 264.—The radio position is 15" F the 18.5–19.0 mag galaxy suggested by Olsen (1970), but in view of the 35" angular size given by Cotton, Owen, and Ghigo (1975), it may be the identification.

1235 + 196.—Shimmins, Bolton, and Wall (1975) say empty field, but there is a 19.5 mag BSO 8" P.

1236 + 327.—Olsen (1970) suggests 18.5 mag galaxy $\sim 20''$ NF. Grueff and Vigotti (1972) say empty field. There is a 20.5–21 mag object at the radio position seen only on the blue plate.

1237 + 231.—Object suggested as a possible quasar by Hazard, Jauncey, and Backer (1970) is $\sim 40''$ SP.

1238 + 243.—There is a faint object within 5" of the radio source visible on the red plate only.

1241 + 254.—The 16th mag NSO 15" F has a stellar spectrum (Wills and Wills 1974).

1245 + 115.—The 13th mag RSO 5" SF has a stellar spectrum (Wills and Wills 1974).

1300 + 320.—The 18.5 mag galaxy 10" P is not the 16th mag galaxy suggested by Grueff and Vigotti (1972).

1303 + 192.—The wrong object was marked on the finding chart of Wills and Bolton (1969) and then was observed and found to be a star by Lynds and Wills (1972). The correct object, which Wills and Bolton meant to mark, is marked on the chart in Figure 2.

1308 + 326.—This had brightened to 17th mag and was found to have a BL Lacertae-type spectrum in 1976 April by Wills and Wills (Marsden 1976b). H. R. Miller reports it at $B \approx 14$ in 1974 April. On archival Harvard plates, it is usually > 17.5 mag, but in the 1940s reached 14.4 mag (Marsden 1976a).

1319 + 270.—Object suggested by Hazard, Jauncey, and Backer (1970) is about 40" NP the radio position.

1320 + 325.—Object suggested by Olsen (1970) is 38" SP the radio position, and was shown to be a star by Schmidt (1974). There is a faint object within 3" of the radio position, not visible on the red plate. Grueff and Vigotti (1975) call it neutral.

1324 + 230.—A 17.5 mag galaxy suggested by Hazard and Jauncey (1972) is 18" S of the radio position.

1326 + 150.—Hazard, Jauncey, and Backer (1970) give 15.5 mag galaxy 28" F, 49" S of the radio position.

1333 + 275.—I disagree with both the 17.5 mag galaxy 50" S suggested by Olsen (1970) and the 19.5 mag BSO 25" S given by Willson (1972).

1338 + 288.—The faint galaxy mentioned by Hazard,

Jauncey, and Backer (1970) may be one of the objects listed in Table 3.

1338 + 107.—Object noted by Wills and Bolton (1969) 60" SF the radio position was found to have a stellar spectrum (Lynds and Wills 1972).

1339 + 266.—VV 5.32.63. Compact group of galaxies.

1345 + 245.—Olsen (1970) suggested a 17.7 mag galaxy 40" SP.

1353 + 186.—The radio position listed in Ghigo and Owen (1973) is lobe-shifted. This was called to my attention by D. Wills (private communication) too late for correction of Table 3. Recent observations by the author on the four-element interferometer of the National Radio Astronomy Observatory show that the correct position (epoch 1950.0) is $13^{\text{h}}53^{\text{m}}39^{\text{s}}.87$ (0^s01), $+18^{\circ}36'57".3$ (0^s2), and that its angular diameter is about 0^s3. It coincides with the Seyfert nucleus in the galaxy Markarian 463. See Adams (1977) for optical data.

1357 + 173.—An 18.5 mag BL Lacertae object 20" SP found by Baldwin *et al.* (1973). The closest object to the radio source, a 20th mag probable galaxy 10" P, is probably the identification. There is evidence of resolution, so the correct identification is not clear.

1401 + 123.—Wills and Wills (1974) find the 16th mag RSO 8" NP to have a stellar absorption spectrum.

1407 + 177.—NGC 5490 (13.5 mag) is 30" SF. Very likely to be the correct identification. Cotton, Owen, and Ghigo (1975) give $20'' \pm 10''$ double, P.A. = $150^{\circ} \pm 30^{\circ}$, close to the radio-optical position angle.

1407 + 316.—Position affected by resolution. Grueff and Vigotti (1975) have noted a faint galaxy 70" NP.

1409 + 133.—There may be a very faint, possibly diffuse object within 5" of the radio position.

1420 + 198.—3C 300. Optical-radio P.A. = 142° for the 19.5 mag galaxy 23" NP. Cotton, Owen, and Ghigo (1975) find radio P.A. = 126° , Macdonald, Kenderdine, and Neville (1968) find P.A. = 135° . The UTRAO radio position centroid agrees with Macdonald *et al.* (1968).

1425 + 267.—Ton 202 ($z = 0.366$, 16th mag) is 50" N. Fanti *et al.* (1975) find three radio components, one coincident with Ton 202. The position is affected by resolution and agrees with none of the Fanti *et al.* positions. Reference star (2), given for this field in Table 3, is Ton 202.

1436 + 129.—Both objects within 10" have stellar spectra (Wills and Wills 1976).

1454 + 303.—Object 9" F has stellar spectrum (Wills and Wills 1974).

1455 + 287.—Olsen (1970) suggests two nearby galaxies, both within 1' of the radio source. Cotton, Owen, and Ghigo (1975) give structure as 45" double, so they cannot definitely be ruled out.

1504 + 163.—Object mentioned by Wills and Bolton (1969) is 25" N and has a stellar spectrum (Wills, private communication). A 17.5 mag BSO is closer: 12^s6 F, 7^s7 S.

1518 + 156.—The 17.5 mag NSO 13" N is called a BL Lacertae-type by Baldwin *et al.* (1973), but D. Wills (private communication) found that it has a stellar spectrum.

1521+137.—Caswell and Wills (1967) suggested the 15.5 mag galaxy 100" NF.

1523+161.—D. Wills (private communication) finds the 18.5 mag BSO 15" SP to be a star.

1556+123.—Cotton, Owen, and Ghigo (1975) find the source to be a 25" double with P.A. = 40°, which coincides with the direction from the radio position to the 19th mag galaxy 22" NF.

1601+116.—Wills and Bolton (1969) suggested a BSO 45" F, found to be a star (D. Wills, private communication).

1621+140.—There may be a very faint object within 5" of the radio position visible only on the blue plate in a group of galaxies. There is also an 18th mag peculiar galaxy about 15" from the SP lobe-shift position.

1636+106.—The object mentioned by Radivich and Kraus (1971) is 80" P.

1638+124.—Object 23" S, given by Wills and Bolton (1969), is a star (D. Wills, private communication).

1639+155.—Spectrum of this object by Lynds shows stellar features (D. Wills, private communication).

1658+148.—The galaxy 16" NP has a diffuse extension in the direction of the radio source. There is also a faint object 8" NP the radio position.

1725+167.—A 16.5 mag NSO 11" F has stellar spectrum (Wills and Wills 1976).

1739+173.—Clarke, Bolton, and Shimmins (1966) suggest the 16th mag galaxy 15" NF, but the 18th mag galaxy 5" NP is more likely.

1746+167.—Wills and Bolton (1969) suggest the 16.5 mag galaxy 40" N.

1751+153.—Possible double object. Radio source is on the line joining a 19th mag BSO and a 20th mag object, 10" apart.

2042+045.—A 17th mag BSO found to have a stellar absorption spectrum by Wills and Wills (1974).

2055+088.—A 15th mag NSO ~6" F has stellar spectrum (Wills and Wills 1976). The identification probably is the very faint galaxy within 3" of the radio position.

2058+019.—Jauncey and Hazard (1970) suggest the 19th mag BSO 24" S.

2142+077.—Jauncey and Hazard (1970) suggest a 17.5 mag BSO 29" P, 96" N of the radio position.

2142+042.—A 17th mag BSO 25" F noted by Wills and Bolton (1969), found to be a star (D. Wills, private communication).

2154-016.—The radio position may be affected by resolution. The RRE declination (McEwan, Browne, and Crowther 1975) is 1' N.

2158+068.—I disagree with both Gearhart *et al.* (1972), who suggest a 20th mag galaxy 71" SP, and with Hazard and Jauncey (1972), who suggest an 18.5 mag galaxy 27" N.

2248+067.—Clarke, Bolton, and Shimmins (1966) suggest the 18.5 mag galaxy ~30" S.

2304+038.—B. and D. Wills (private communication) found the spectrum of this object to be that of an emission-line galaxy, with $z = 0.154$.

2310+050.—The radio position is probably affected

by resolution of this extended radio source, found to be a triple source extending 161" at P.A. = 72° (Fomalont 1971). Kristian, Sandage, and Katem (1974) suggest a faint galaxy near the SF lobe-shift position, at the central component of the radio source.

2313+011.—The radio position is affected by resolution. Merkelijn (1969) and McEwan, Browne, and Crowther (1975) agree on a radio position 1' N of the UTRAO position and on an identification with an 18–20 mag peculiar galaxy. McEwan, Browne, and Crowther (1975) give angular size as 130".

2318+026.—I disagree with the Parkes 2700 MHz position which is 20" P, 22" N of the UTRAO radio position. Asymmetric structure (diam. ~50") was noted by McEwan, Browne, and Crowther (1975), who suggested an 18.6 mag quasar 6" F, 20" N of the UTRAO position, having $z = 1.968$ (D. Wills, private communication).

2349-014.—An N galaxy with some quasar-like features, $z = 0.174$. See Wills and Wills (1974) for discussion of this object and references.

VI. COMPARISON WITH OTHER ACCURATE OPTICAL POSITIONS

Table 4 shows the results of comparisons with other accurate optical positions. The position differences are in the sense "this paper minus other." The column headed "Approx. Errors" gives the position errors claimed by the authors of the papers being compared. In all cases, the population standard deviations σ_α and σ_δ are no larger than expected from the combination of the errors quoted in each case with my error estimate of 0".5 to 0".7.

Only one comparison shows a significant systematic difference. The large difference of $\langle \Delta\alpha \rangle = -1".36 \pm 0".16$ with respect to Véron's (1972) positions is not caused by one or two discrepant cases. All five objects have $\Delta\alpha$ in the range $-1"$ to $-2"$. A systematic error in the Véron (1972) list seems likely. No such problem is evident in the later list by Véron and Véron (1975), however.

Systematic right ascension differences between different position determinations of small groups of quasars have been noted (Hunstead 1971; Argue and Taylor 1974) and attributed to a possible systematic difference between the SAO and AGK3 systems. Such an effect is not supported by the results in Table 4. Furthermore, several direct comparisons of SAO and AGK3 positions show no systematic differences (Adgie 1974; Véron and Véron 1975; P. Hemenway, private communication).

VII. BACKGROUND COUNTS

Objects were counted in a 25" radius area centered 371" west of each radio source in the sample. This distance (6.5 BSI lobe shifts) was chosen to allow virtually no lobe-shifted sources to fall in the area, but to have the area close enough to the radio sources that the background counts would be, on the average, the same as in the radio source fields. Since the areas have

TABLE 4
OPTICAL POSITION COMPARISONS (this paper minus other)

	Approx. Errors	Ref. System	N	$\langle \Delta \alpha \rangle$	σ_α	$\langle \Delta \delta \rangle$	σ_δ
Wills, Wills, and Douglas 1973;							
Wills 1976.....	0 $^{\circ}$ 4–0 $^{\circ}$ 5	SAO	140	–0 $^{\circ}$ 07 \pm 0 $^{\circ}$ 05	0 $^{\circ}$ 55	+0 $^{\circ}$ 06 \pm 0 $^{\circ}$ 05	0 $^{\circ}$ 57
Johnson 1974.....	0 $^{\circ}$ 7	SAO	19	–0 $^{\circ}$ 31 \pm 0 $^{\circ}$ 23	1 $^{\circ}$ 00	+0 $^{\circ}$ 10 \pm 0 $^{\circ}$ 24	1 $^{\circ}$ 05
Hunstead 1971.....	0 $^{\circ}$ 4–0 $^{\circ}$ 5	SAO	7	+0 $^{\circ}$ 03 \pm 0 $^{\circ}$ 08	0 $^{\circ}$ 21	+0 $^{\circ}$ 14 \pm 0 $^{\circ}$ 12	0 $^{\circ}$ 33
Kristian and Sandage 1970.....	0 $^{\circ}$ 2–0 $^{\circ}$ 6	AGK3	6	–0 $^{\circ}$ 17 \pm 0 $^{\circ}$ 22	0 $^{\circ}$ 54	–0 $^{\circ}$ 18 \pm 0 $^{\circ}$ 27	0 $^{\circ}$ 66
Bergamini <i>et al.</i> 1973.....	1 $^{\circ}$ 0	AGK2	8	–0 $^{\circ}$ 37 \pm 0 $^{\circ}$ 28	0 $^{\circ}$ 79	–0 $^{\circ}$ 39 \pm 0 $^{\circ}$ 15	0 $^{\circ}$ 44
Véron 1972.....	0 $^{\circ}$ 5	AGK3	5	–1 $^{\circ}$ 36 \pm 0 $^{\circ}$ 16	0 $^{\circ}$ 35	+0 $^{\circ}$ 41 \pm 0 $^{\circ}$ 42	0 $^{\circ}$ 93
Véron and Véron 1975.....	0 $^{\circ}$ 5–0 $^{\circ}$ 6	AGK3	15	–0 $^{\circ}$ 23 \pm 0 $^{\circ}$ 17	0 $^{\circ}$ 65	–0 $^{\circ}$ 19 \pm 0 $^{\circ}$ 18	0 $^{\circ}$ 68

virtually the same sky distribution as the radio sources in the sample, these counts are optimum for comparing with counts in the radio source fields.

The counting was done on the 0 $^{\circ}$.62 mm $^{-1}$ and 4 $^{\circ}$.8 mm $^{-1}$ photographs mentioned in § Vb. The objects were classified as to color, magnitude, and morphological class as described in § Vc. Counts from both red and blue plates were combined. Objects visible on only one color plate were included in the total counts, so these counts are larger than counts confined to a single color would be.

The total area covered by the background count fields is 362 square arc minutes.

The counts of all objects, stars and galaxies included, to the PSS plate limit are given in Figure 3 as a function of galactic latitude. The error bars represent \sqrt{N} , where N is the number of objects counted in each latitude bin. Some of the points have been slightly shifted horizontally for clarity.

My counts at the north galactic pole are compared in Table 5 with counts by others, and are seen to be in reasonable agreement.

Counts of all objects and of various subclasses are given as a function of galactic longitude and latitude in Table 6. Regions I to VIII of the l^{II} - b^{II} plane, referred to in Table 6, are defined by Figure 4. The errors given in Table 6 are ρ/\sqrt{N} , where ρ is the surface density and N is the number of objects counted.

VIII. COMPLETENESS AND RELIABILITY

The completeness of an identified sample may be defined as the fraction of the number of visible identifications possessed by the sample that have actually been found. The reliability is the fraction of claimed identifications which are indeed correct. These quantities may be calculated from knowledge of the radio and optical position errors, the size of the search radius, the density of background objects, and the fraction of identifications which are visible. More complete discussions of completeness and reliability and derivations of formulae for calculating these quantities can be found in papers by Condon, Balonek, and Jauncey (1975) and by de Ruiter, Willis, and Arp (1977).

The radius of the area to be searched for identifications is generally taken to be 2 to 3 times the size of

the position errors for point sources. For sources of appreciable angular size, one usually increases the search radius in proportion to the angular size. To estimate completeness and reliability, the nominal position errors which go into the calculations are also increased in proportion to the angular sizes of the sources.

Due to the limited spatial frequency coverage of the BSI, accurate angular sizes could not be determined

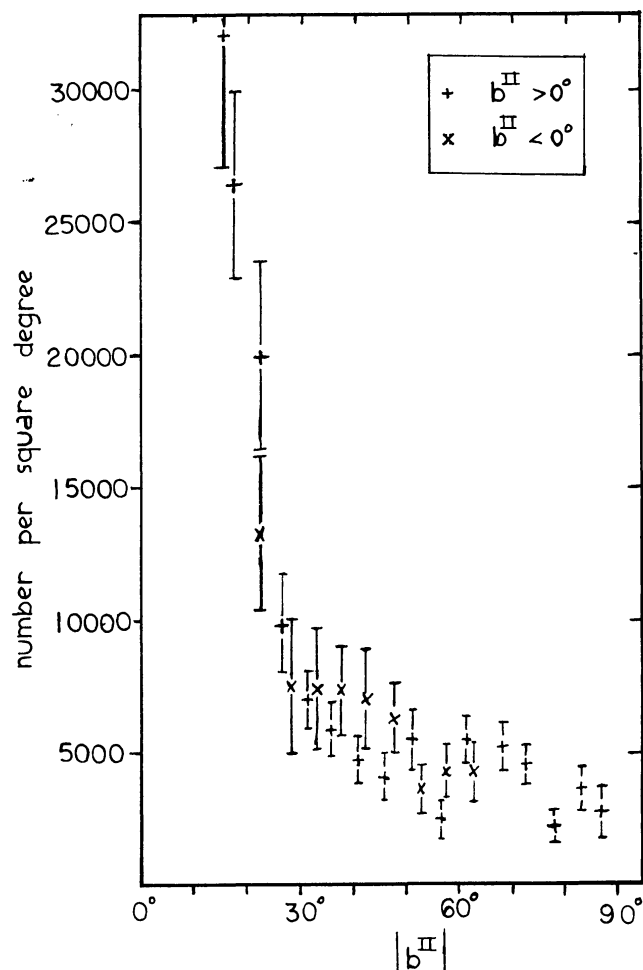


FIG. 3.—Counts of background objects plotted versus galactic latitude.

TABLE 5
VARIOUS COUNTS IN THE NORTH GALACTIC POLAR CAP*

Counts by	Objects Counted	Number per Sq. Degree
Brown 1973.....	Stars + galaxies to $B = 21$	2560 ± 130
Bergamini <i>et al.</i> 1973.....	All objects to PSS print limit	3000 ± 300
Braccisi <i>et al.</i> 1970.....	All objects to PSS print limit	2800
This paper.....	All objects to PSS plate limit $85^\circ \lesssim b^{\text{II}} \leq 90^\circ$	2970 ± 400

* Combining both star and galaxy counts.

for this sample. Nevertheless, the appropriate position error distribution for use in calculating completeness and reliability was determined by the following method: All objects were counted in $1''$ wide rings centered on the radio sources, out to a radius of $20''$. The background count level was subtracted from the resulting distribution, to produce an estimate of the distribution of radio-optical separations for the true identifications. This method works because the background counts are a good representation of the background at the radio sources, as explained in § VII. This procedure is believed to be freer from bias than using the distribution of radio-optical separations of the identifications found as described in § Vd, since the latter group includes a number of random coincidences.

The number of true identifications estimated by the above method for this sample is 307 ± 28 . This may be too high since a number of the radio sources probably lie in faint clusters of galaxies. The size of this effect is difficult to assess. Various estimates for this sample suggest that the number of identifications predicted is too high by no more than about 20, i.e., about 7%. The number of suggested identifications listed in Table 3 is 319, of which 20 to 30 are probably random coincidences. Thus the number of identifica-

tions found by considering individual cases is in approximate agreement with the statistical estimate.

The distribution of radio-optical separations, found as described above, can be fitted reasonably well with a two-component distribution of the following form:

$$n(r) = 0.7 \frac{r}{\sigma_1^2} \exp(-\frac{1}{2}r^2/\sigma_1^2) + 0.3 \frac{r}{\sigma_2^2} \exp(-\frac{1}{2}r^2/\sigma_2^2), \quad (4)$$

where n is proportional to the number of identifications between $r - 1$ and r arc seconds from a radio source, r is the radio-optical separation in arc seconds, $\sigma_1 = 2''.2$, and $\sigma_2 = 9''.5$. The quantities σ_1 and σ_2 represent combined radio and optical position errors; $2''.2$ is about the expected combined error for point sources.

The interpretation of equation (4) is that 70% of the identifications (represented by the first term of eq. 4) have no detectable offset of the optical position from the radio source centroid. It is assumed that the appreciable offsets of the remaining 30% (the second term of eq. 4) are due to these sources having large angular sizes. Such an effect is expected since, in the

TABLE 6
COUNTS OF OBJECTS IN VARIOUS REGIONS OF THE $l^{\text{II}}-b^{\text{II}}$ PLANE

Region.....	I	II	III	IV	V	VI	VII	VIII	All
Center of region l^{II}	195°	202°	...	25°	40°	60°	65°	115°	...
b^{II}	$+32^\circ$	$+57^\circ$	$+90^\circ$	$+55^\circ$	$+20^\circ$	-25°	-42°	-58°	...
Number of fields.....	38	93	164	163	44	18	58	86	664
All objects, to PSS limit*.....	5560	3260	3380	5790	23250	10630	6600	3990	5950
	± 980	± 480	± 370	± 480	± 1870	± 1970	± 870	± 550	± 240
All objects, $m_v \leq 19^*$	2200	1890	2140	3320	14950	8430	4400	2350	3710
	± 630	± 370	± 290	± 370	± 1500	± 1760	± 710	± 430	± 190
All objects, $m_v \leq 18^*$	1283	810	930	1460	7270	3300	2080	1210	1720
	± 460	± 240	± 190	± 240	± 1040	± 1100	± 480	± 310	± 130
All objects, $m_v \leq 17^*$	1100	670	500	580	3100	1470	810	830	860
	± 430	± 210	± 140	± 150	± 690	± 730	± 300	± 250	± 90
Stellar objects, $m_v \leq 19^*$	2080	1350	1570	2830	15300	8430	3530	1840	3180
	± 600	± 310	± 250	± 340	± 1520	± 1760	± 630	± 380	± 180
Diffuse objects, $m_v \leq 19^*$	350	500	600	730	1500	†	570	770	670
	± 250	± 190	± 160	± 170	± 470		± 250	± 240	± 80
BSOs, $m_v \leq 19^*$	†	280	200	530	3750	2570	230	80	570
		± 140	± 90	± 150	± 750	± 970	± 160	± 80	± 80

* Densities are given in number per square degree.

† No counts here.

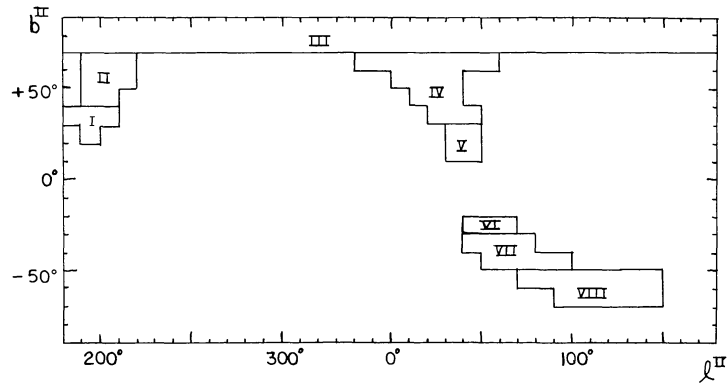


FIG. 4.—Observed regions in galactic coordinates

first place, a measurable difference between the optical and radio centroid may well occur in a source of large angular size, and, second, the BSI, due to its limited spatial frequency coverage, can measure an incorrect radio centroid for an asymmetric brightness distribution having an angular size greater than $10''$.

A test was made to see if some of the large radio-optical separations might occur for reasons other than the two outlined above. The 116 radio sources in this sample with no detectable angular extent ($\theta \lesssim 15''$) were analyzed as to distribution of radio-optical separation. In this subsample, no separations in excess of those expected for point sources were found. Thus no evidence was found for radio-optical offsets resulting from any other than angular size-related causes.

Assuming equation (4) to be the appropriate position error distribution for this sample, and using the derivation of formulae for completeness and reliability described by Condon, Balonek, and Jauncey (1975), the completeness and reliability of these identifications for various search radii was computed, with the results displayed in Table 7. On the basis of these calculations, reliability classes of "A," "B," and "C" were assigned to all suggested identifications as described in the key to Table 3.

Error distributions similar to equation (4) were also determined for the sources identified with galaxies and quasars. In general, galaxies have larger radio angular sizes than quasars. Consistent with this, $\sim 40\%$ of the galaxies and probable galaxies in the sample were found to have appreciable radio-optical separations ($r \gtrsim 5''$), while only 10% to 20% of the quasars and probable quasars exceed separations expected for point sources.

TABLE 7
COMPLETENESS AND RELIABILITY

Search Radius	Completeness	Reliability
$5''$	0.66	0.92
$10''$	0.80	0.81
$15''$	0.87	0.71
$20''$	0.91	0.63

IX. IDENTIFICATION RESULTS

A tabulation of the numbers of different kinds of identifications is given in Table 8. The identification classes have the same meaning as in Table 3. From Table 8, the total number of identifications visible on the PSS plates is 319. The number of type A identifications ($r \leq 5''$; reliability $\geq 90\%$) is given in Table 9. There are 205 of these, of which about 23 are expected to be spurious, based on an average background rate of 4.758×10^{-4} per square arc second.

It is impossible to distinguish between galaxy and stellar images near the plate limit, so statistics of occurrence of these objects are meaningful only above an apparent brightness limit somewhat above the plate limit. For $m_v \leq 19.5$, 102 (15%) of the sources are *G* or *PG*, and 136 (20%) are *Q* or *PQ*, having allowed for chance coincidences.

The counts of *G*, *PG*, *Q*, and *PQ* objects by magnitude are shown in Figure 5. The numbers of galaxies increase steadily to the plate limit, but the quasars seem to begin to cut off at a somewhat brighter level, the maximum occurring at 18.5 mag. This cutoff in quasar brightness has been noted in many other samples—for example, see Schmidt (1972).

a) Neutral Stellar Objects (NSOs)

Table 10 gives information for the 15 objects classified as neutral and stellar which lie within $5''$ of a

TABLE 8
IDENTIFICATION CONTENT*

Class	Number	Class	Number
EF.....	306	G.....	25
>G.....	8	PG.....	104
>Q.....	5	CG.....	16
C > G.....	2	CPG.....	14
CEF.....	18	Q.....	79
OB.....	6	PQ.....	76
O.....	3	CQ.....	2

* See key to Table 3 for explanation of classes.

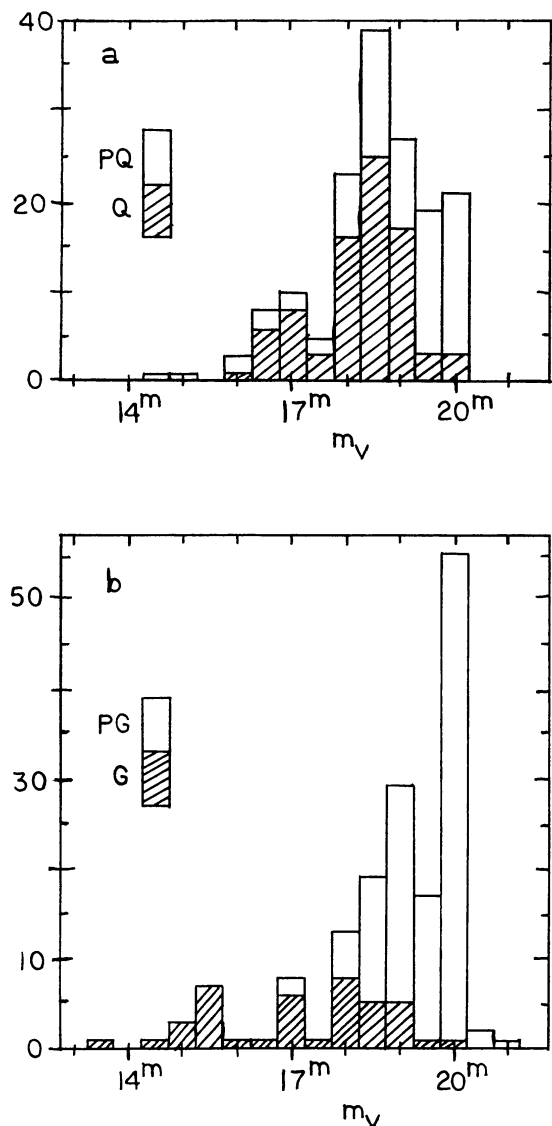


FIG. 5.—Magnitude histograms for the identified objects. (a) Probable and definite quasars or BL Lacertae-type objects. Includes classes Q, PQ, and C Q. (b) Probable and definite galaxies. Includes classes G, PG, C G, and C PG.

radio source position. From background counts, 4.7 of them are expected to be chance coincidences.

The spectral indices given in Table 10 were found from this expression:

$$\alpha \equiv \log (S_{1415}/S_{365})/\log (1415/365). \quad (5)$$

The 1415 MHz flux densities from the Ohio catalog were used and were multiplied by the factor 1.18 to put them on the absolute scale of B. J. Wills (1973). The 365 MHz flux densities are mostly total flux densities estimated by fitting brightness distribution models to

TABLE 9
CONTENT OF IDENTIFICATIONS WITHIN
5 ARCSEC

Classes	Number
G, PG, C G, and C PG.....	95
Q, PQ, and C Q.....	107
O.....	3
Total.....	205

the Texas data (Bash, Cotton, and Douglas 1974; Cotton, Owen, and Ghigo 1975). Measured total 365 MHz flux densities from the Texas short-baseline survey (Cotton 1974) were used when available. When only the interferometer flux density was available, an upper limit to α is given.

Optical spectra are available for most of these NSOs. They include BL Lacertae types, quasars, and N galaxies. The prevalence of BL Lacertae types among flat-spectrum NSOs has been noted, e.g., by Browne, Crowther, and Adgie (1973). Among the confirmed and likely BL Lacertae types in Table 10, the majority have $\alpha > -0.5$, but at least one (2207+020) has a steep spectrum ($\alpha < -0.5$). Although no confirming spectrum exists for 2217+018, its lack of UV excess may indicate that it is a BL Lacertae type. It, too, has a steep spectrum.

b) Radio Stars?

In this sample, 13 objects within 10" of radio sources were found to have stellar spectra (Wills and Wills 1974; Wills and Wills 1976; C. R. Lynds, private communication). The question may therefore be raised as to whether some of these stars are indeed radio sources. If so, they would be a new class of radio star, since the known types are too faint or too short-lived to be present in this sample.

To answer this question, comparisons with star counts were made. It was found that the expected number of random coincidences, even when confining the attention to radio sources at high galactic latitude, exceeded the number of stars found in this sample. Thus the 13 stars are almost certainly all random coincidences.

I am indebted to Dr. J. N. Douglas for considerable advice and encouragement; to Dr. D. Wills and Dr. B. J. Wills for many enlightening discussions and for allowing me to use some spectroscopic results prior to publication; to Mr. Chip Wolfe for programming; to Dr. P. D. Hemenway and Ms. E. Bozyan for assistance in the plate-measuring project; and to the crew of measuring machine operators. This work was supported by grants from the National Science Foundation and the National Geographic Society, and by the organized research program of the University of Texas.

TABLE 10
 NSOS WITHIN 5 ARCSEC OF A RADIO SOURCE

Source	Optical to Radio Distance	m_v	Spectral Index	Remarks	Ref.
0003-003 (3C 2).....	1"7	19.5	-0.79	QSO, $z = 1.037$	2
0013-005.....	4"2	19.0	-0.06	QSO, no z measured	6
0025+006.....	2"3	16.5	< -0.75	N-gal, $z = 0.105$	7
0912+297.....	1"2	16.5	< -0.06	Cont. spec., UV excess	3, 8
0923+257.....	4"7	8.5	< -0.57	Stellar spec.	7
0927+252.....	1"3	18.5	< -0.57	Inconcl. spec., prob. QSO	6
1147+245.....	4"7	17.0	-0.14	Cont. spec.	8
1219+285 (ON + 231).....	3"2	15.0	< +0.24	Cont. spec.	1
1604+159.....	2"1	18.0	< -0.41	Inconcl. spec., BSO	8
1755+153.....	2"3	18.5	< -0.80	Crowded area	
1758+111.....	4"8	19.0	< -0.61	Crowded area, object may be slightly blue	
2207+020.....	2"1	19.5	< -1.3	Cont. spec., BSO, no UV excess	5
2217+018.....	1"5	19.0	-0.82	No UV excess, BSO	4
2222+051.....	0"7	18.0	-0.93	QSO, $z = 2.324$	6
2304+038.....	2"6	18.0	-0.90	Emission line gal, $z = 0.154$, slightly nonstellar PSS	6

REFERENCES.—(1) Burbidge and Strittmatter 1972; Strittmatter *et al.* 1972; (2) Lynds 1967; (3) Fanti *et al.* 1975; (4) McEwan, Browne, and Crowther 1975; (5) Strittmatter *et al.* 1974; (6) D. Wills, private communications; (7) Wills and Wills 1974; (8) Wills and Wills 1976.

REFERENCES

- Adams, T. F. 1977, *Ap. J. Suppl.*, **33**, 19.
 Adgie, R. L. 1974, *Observatory*, **94**, 300.
 Agnew, D., and Arp, H. C. 1973, *Pub. A.S.P.*, **85**, 162.
 Aizu, K. 1966, *Pub. Astr. Soc. Japan*, **8**, 219.
 Argue, A. N., and Taylor, C. M. 1974, *Observatory*, **94**, 295.
 Baldwin, J. A., Burbidge, E. M., Hazard, C., Murdoch, H. S.,
 Robinson, L. B., and Wampler, E. J. 1973, *Ap. J.*, **185**, 739.
 Bash, F. N., Cotton, W. D., and Douglas, J. N. 1974, *A.J.*, **79**,
 1341.
 Bergamini, R., Braccisi, A., Colla, G., Fanti, C., Fanti, R.,
 Ficarra, A., Formigini, L., Gandolfi, E., Gioia, I., Lari, C.,
 Marano, B., Padrielli, L., Tomasi, P., and Vigotti, M. 1973,
Astr. Ap., **23**, 195.
 Bolton, J. G., and Ekers, J. 1967, *Australian J. Phys.*, **20**, 109.
 Bolton, J. G., Shimmins, A. J., and Merkelijn, J. 1968,
Australian J. Phys., **21**, 81.
 Bolton, J. G., and Wall, J. V. 1970, *Australian J. Phys.*, **23**, 789.
 Braccisi, A., Ficarra, A., Formigini, L., Gandolfi, E., Lari,
 C., Padrielli, L., and Tomasi, P. 1970, *Astr. Ap.*, **6**, 268.
 Bridle, A. H., Davis, M. M., Fomalont, E. B., and Lequeux,
 J. 1972, *A.J.*, **77**, 405.
 Brown, G. S. 1973, Ph.D. thesis, University of Texas at
 Austin.
 Browne, I. W. A., Crowther, J. H., and Adgie, R. L. 1973,
Nature, **244**, 146.
 Brundage, R. K., Dixon, R. S., Ehman, J. R., and Kraus, J. D.
 1971, *A.J.*, **76**, 777.
 Burbidge, E. M., and Strittmatter, P. A. 1972, *Ap. J. (Letters)*,
174, L57.
 Caswell, J. L., and Wills, D. 1967, *M.N.R.A.S.*, **135**, 231.
 Clarke, M. E., Bolton, J. G., and Shimmins, A. J. 1966,
Australian J. Phys., **19**, 375.
 Colla, G., Fanti, C., Fanti, R., Gioia, I., Lari, C., Lequeux, J.,
 Lucas, R., and Ulrich, M. H. 1975, *Astr. Ap. Suppl.*, **20**, 1.
 Condon, J. J., Balonek, T. J., and Jauncey, D. L. 1975, *A.J.*,
80, 887.
 Cotton, W. D. 1974, *UTRAO Internal Report* 44, November
 15.
 ———. 1976, *Ap. J. Suppl.*, **32**, 467.
 Cotton, W. D., Owen, F. N., and Ghigo, F. D. 1975, *A.J.*, **80**,
 353.
 de Ruiter, H. R., Willis, A. G., and Arp, H. C. 1977, *Astr. Ap.*
Suppl., **28**, 486.
 Douglas, J. N., Bash, F. N., Ghigo, F. D., Moseley, G. F., and
 Torrence, G. W. 1973, *A.J.*, **78**, 1.
 Elsmore, B., and Mackay, C. D. 1969, *M.N.R.A.S.*, **146**, 361.
 Fanti, C., Fanti, R., Ficarra, A., Formigini, L., Giovanni, G.,
 Lari, C., and Padrielli, L. 1975, *Astr. Ap. Suppl.*, **19**, 143.
 Fomalont, E. B. 1971, *A.J.*, **76**, 513.
 Fomalont, E. B., and Moffet, A. T. 1971, *A.J.*, **76**, 5.
 Gearhart, M. R., Lund, J. M., Frantz, D. J., and Kraus, J. D.
 1972, *A.J.*, **77**, 557.
 Ghigo, F. D., and Owen, F. N. 1973, *A.J.*, **78**, 848.
 Grueff, G., and Vigotti, M. 1972, *Astr. Ap. Suppl.*, **6**, 1.
 ———. 1973, *Astr. Ap. Suppl.*, **11**, 41.
 ———. 1975, *Astr. Ap.*, **20**, 57.
 Hazard, C., Gulkis, S., and Bray, A. D. 1967, *Ap. J.*, **148**, 669.
 Hazard, C., and Jauncey, D. L. 1972, *A.J.*, **77**, 621.
 Hazard, C., Jauncey, D. L., and Backer, D. C. 1970, *A.J.*, **75**,
 1039.
 Hunstead, R. W. 1971, *M.N.R.A.S.*, **152**, 277.
 Jauncey, D. L., and Hazard, C. 1970, *Ap. Letters*, **7**, 1.
 Jauncey, D. L., and Niell, A. E. 1971, *Nature Phys. Sci.*, **229**,
 223.
 Johnson, K. H. 1974, *A.J.*, **79**, 1006.
 Kristian, J., and Sandage, A. 1970, *Ap. J.*, **162**, 391.
 Kristian, J., Sandage, A., and Katem, B. 1974, *Ap. J.*, **191**, 43.
 Lynds, R. 1967, *Ap. J.*, **147**, 837.
 Lynds, R., and Wills, D. 1972, *Ap. J.*, **172**, 531.
 Macdonald, G. H., Kenderdine, S., and Neville, A. C. 1968,
M.N.R.A.S., **138**, 259.
 Macdonald, G. H., and Miley, G. K. 1971, *Ap. J.*, **164**, 237.
 Mackay, C. D. 1969, *M.N.R.A.S.*, **145**, 31.
 Marsden, B. G. 1976a, *IAU Circ.*, No. 2939.
 ———. 1976b, *IAU Circ.*, No. 2954.
 McEwan, J. J., Browne, I. W. A., and Crowther, J. H. 1975,
Mem. R.A.S., **80**, 1.
 Merkelijn, J. K. 1969, *Australian J. Phys.*, **22**, 237.
 Merkelijn, J. K., and Wall, J. V. 1970, *Australian J. Phys.*, **23**,
 575.
 Olsen, E. T. 1970, *A.J.*, **75**, 764.
 Radovich, M. M., and Kraus, J. D. 1971, *A.J.*, **76**, 683.
 Riley, J. M., and Pooley, G. G. 1975, *Mem. R.A.S.*, **80**, 105.
 Rinsland, C. P., Dixon, R. S., and Kraus, J. D. 1975, *A.J.*, **80**,
 759.
 Schmidt, M. 1972, *Ap. J.*, **176**, 273.

- Schmidt, M. 1974, *Ap. J.*, **193**, 505.
 Sharp, J. R., and Bash, F. N. 1975, *A.J.*, **80**, 335.
 Shimmins, A. J., Bolton, J. G., and Wall, J. V. 1975, *Australian J. Phys. Ap. Suppl.*, **34**, 63.
 Strittmatter, P. A., Carswell, R. F., Gilbert, G., and Burbidge, E. M. 1974, *Ap. J.*, **190**, 509.
 Strittmatter, P. A., Serkowski, K., Carswell, R., Stein, W. A., Merrill, K. M., and Burbidge, E. M. 1972, *Ap. J. (Letters)*, **175**, L7.
 Stull, M. A. 1973, *A.J.*, **78**, 285.
 Véron, M. P. 1972, *Astr. Ap.*, **20**, 471.
 Véron, M. P., and Véron, P. 1974, *Astr. Ap. Suppl.*, **18**, 309.
 ———. 1975, *Astr. Ap.*, **42**, 1.
 Webster, A. 1976, *M.N.R.A.S.*, **175**, 71.
 Wills, B. J. 1973, *Ap. J.*, **180**, 335.
 ———. 1976, *A.J.*, **81**, 1031.
 Wills, B. J., Wills, D., and Douglas, J. N. 1973, *A.J.*, **78**, 521.
 Wills, D., and Bolton, J. G. 1969, *Australian J. Phys.*, **22**, 775.
 Wills, D., and Wills, B. J. 1974, *Ap. J.*, **190**, 271.
 ———. 1976, *Ap. J. Suppl.*, **31**, 143.
 Willson, M. A. G. 1972, *M.N.R.A.S.*, **156**, 7.
 Wyndham, J. D. 1966, *Ap. J.*, **144**, 459.
 Zwicky, F., Herzog, E., Wild, P., Karpowicz, M., and Kowal, C. T. 1961–1968, *Catalog of Galaxies and Clusters of Galaxies* (in 6 vols.; Pasadena: California Institute of Technology Press).

FRANK D. GHIGO: Physics Department, Brandeis University, Waltham, MA 02154

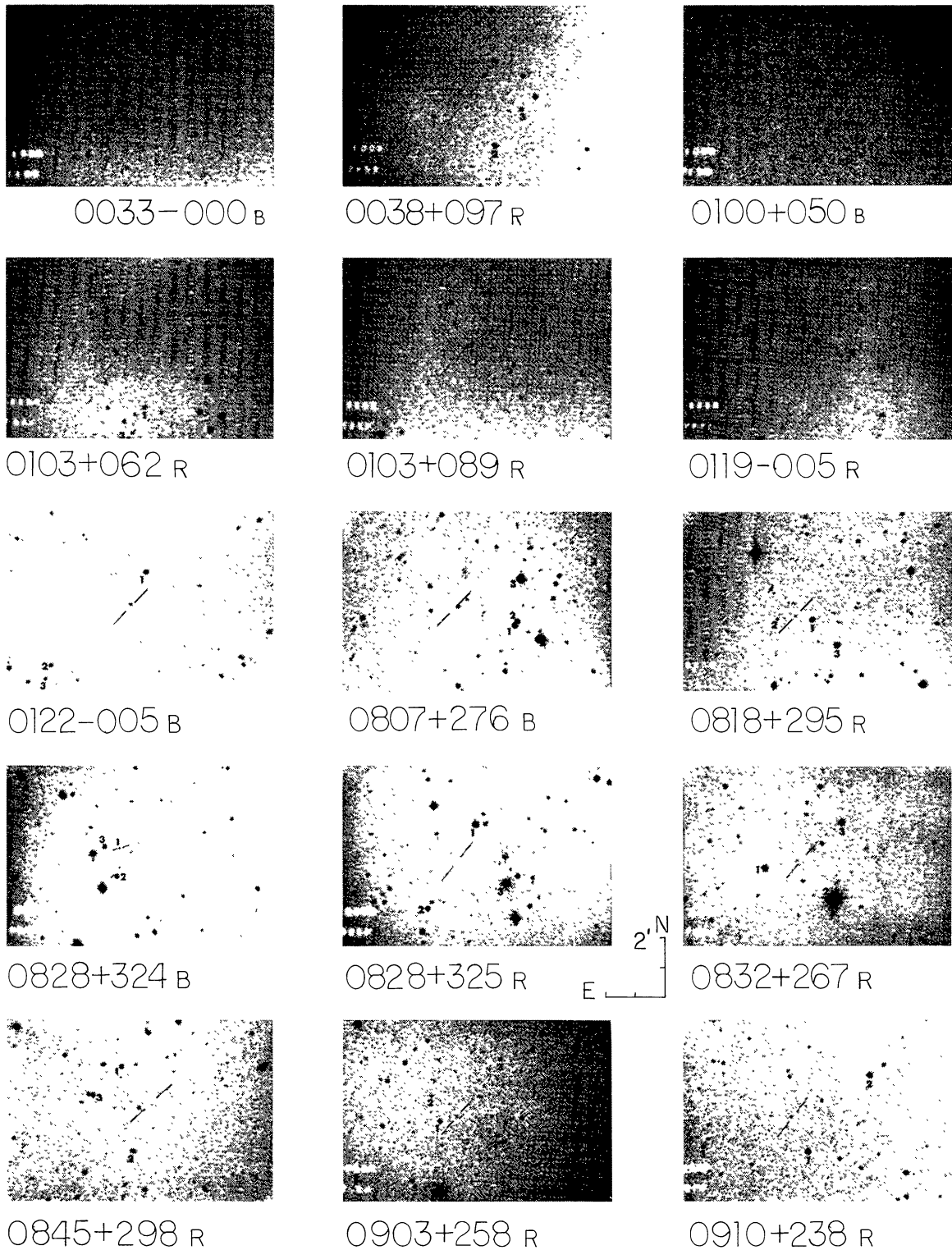
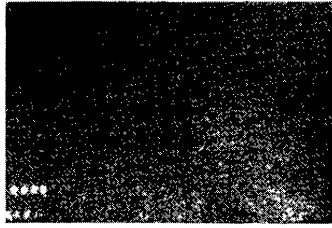


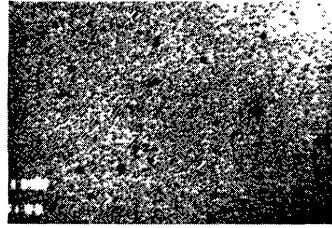
FIG. 2.—Identification charts for 118 new and revised identifications, reproduced, with permission, from the *National Geographic Society–Palomar Observatory Sky Survey* plates. The reticle lines indicate the radio position, not necessarily the identification.

GHIGO (see page 384)

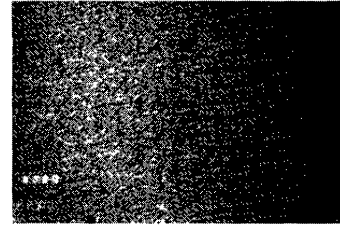
PLATE 30



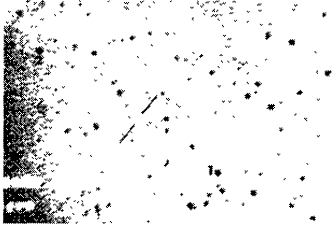
0912+253 R



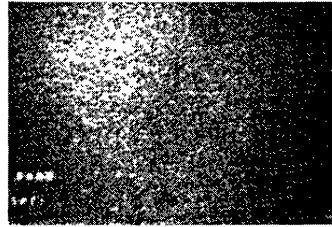
0914+257 R



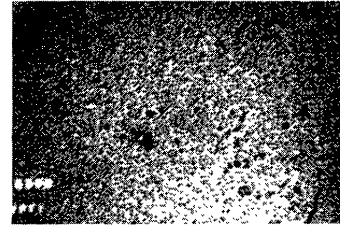
0916+263 B



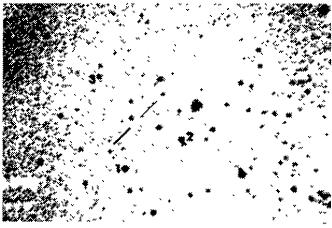
0919+313 B



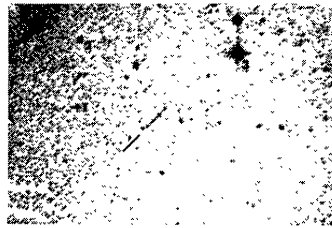
0927+252 B



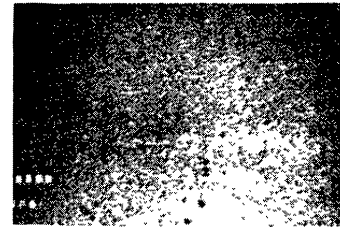
0932+241A R



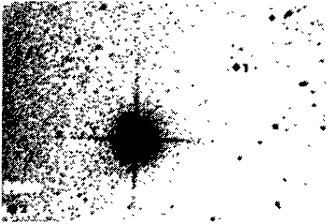
0932+241B R



0936+247 R



0942+276 R



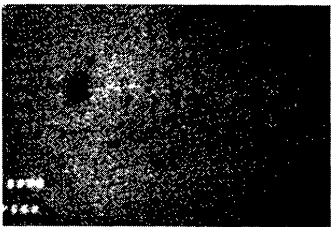
0947+248 R



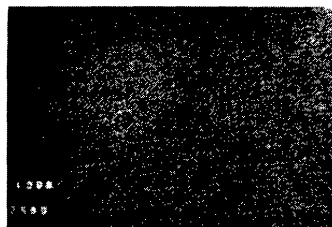
1004+231 B



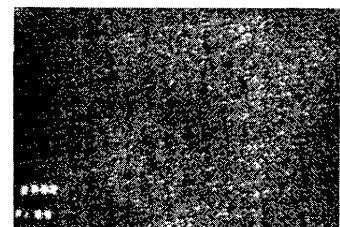
1026+256 R



1032+265 R



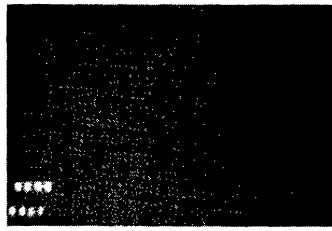
1032+260 B



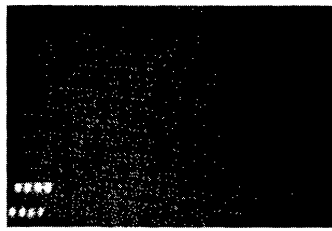
1048+305 R

FIG. 2.—Continued

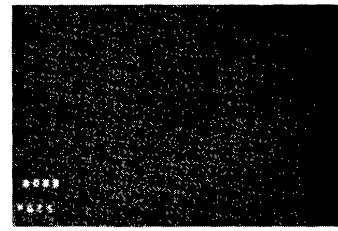
GHIGO (see page 384)



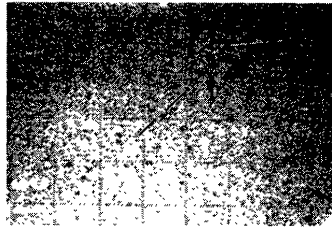
1106+252 R



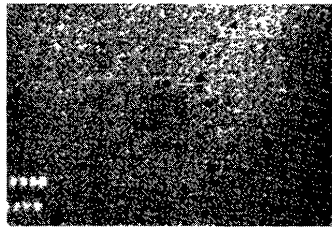
1118+237 R



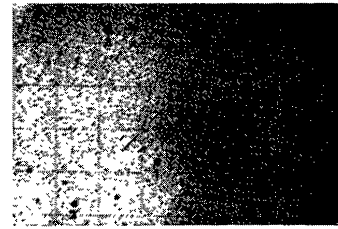
1136+299 R



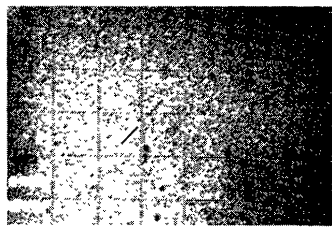
1207+118 B



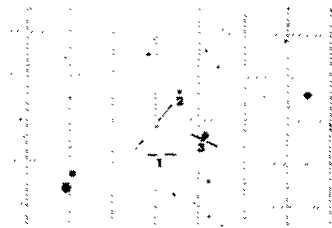
1222+264 R



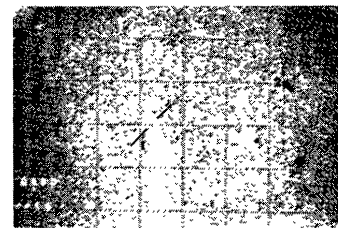
1228+256 B



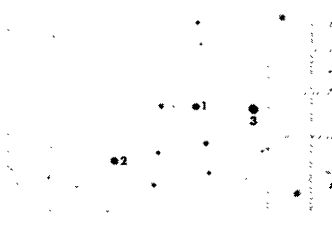
1234+252 R



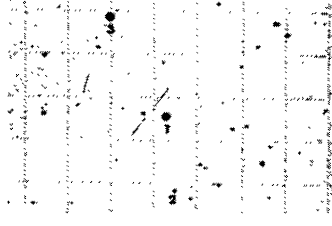
1235+196 B



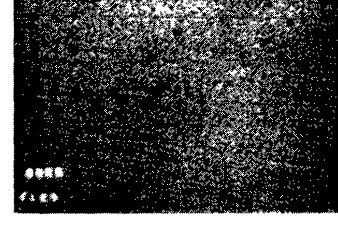
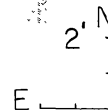
1236+327 B



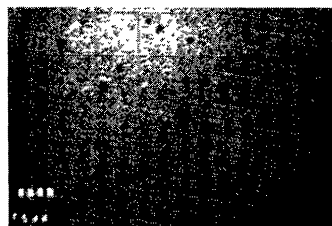
1239+168 R



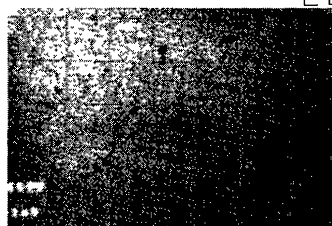
1244+128 R



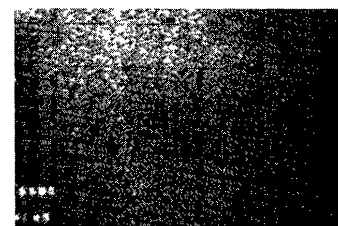
1300+320 R



1303+192 R



1304+250 R

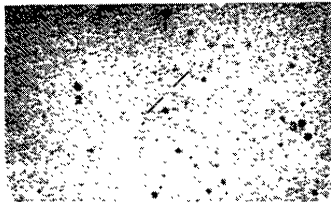


1304+243 R

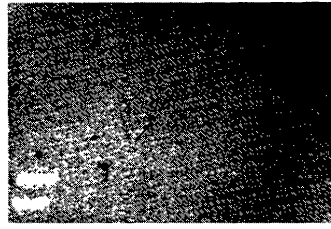
FIG. 2.—Continued

GHIGO (see page 384)

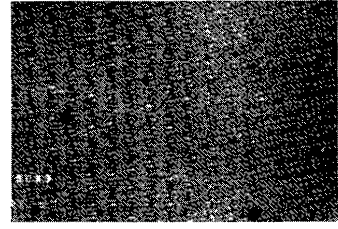
PLATE 32



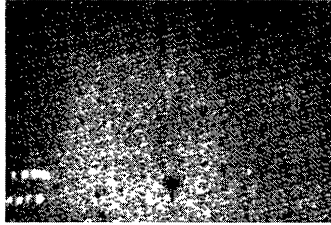
1317+199 R



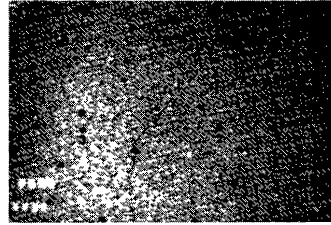
1319+230 R



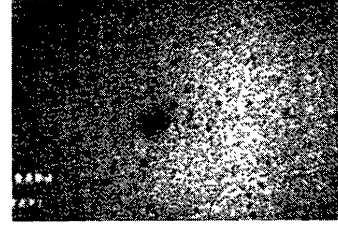
1325+124 R



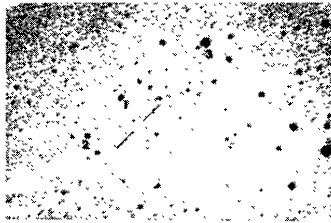
1325+126 R



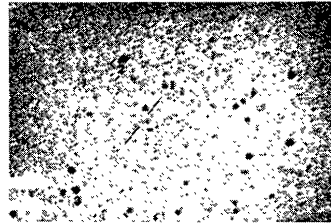
1326+150 R



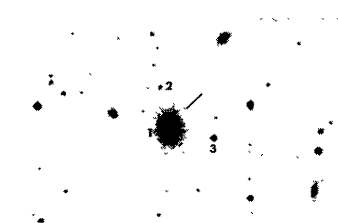
1339+266 B



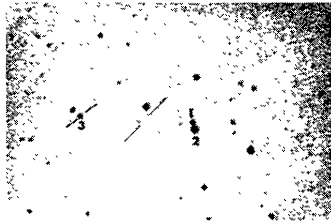
1345+262 R



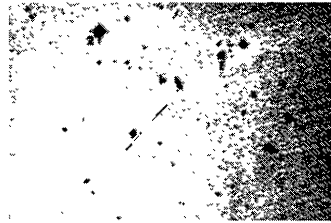
1357+173 R



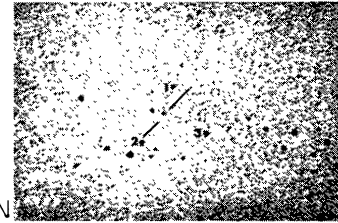
1407+177 R



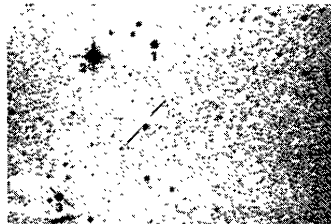
1408+141 R



1413+245 R



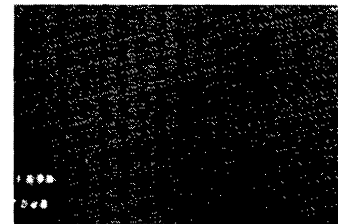
1422+128 B



1425+287 R



1428+260 R



1435+304 R

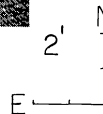


FIG. 2.—Continued

GHIGO (see page 384)

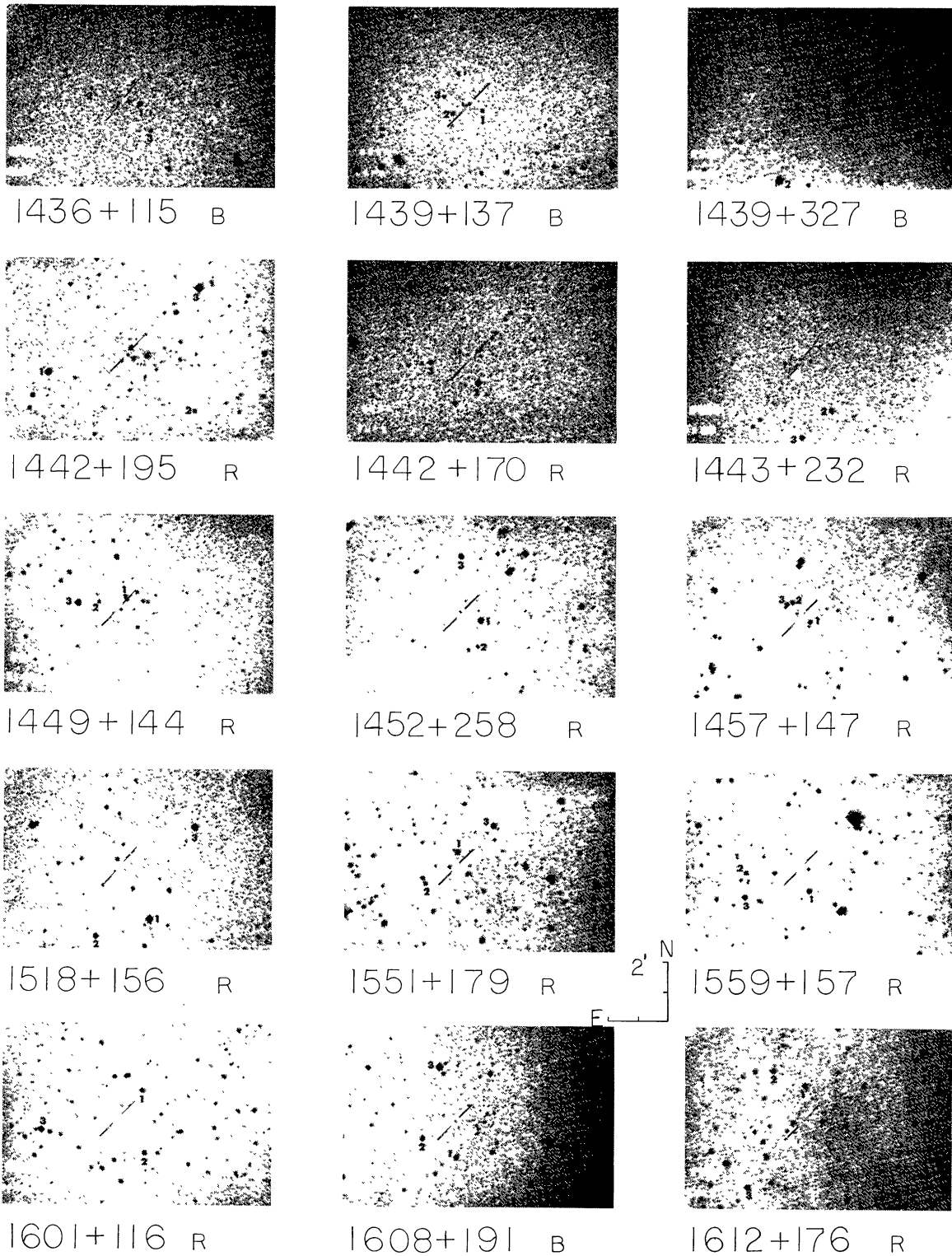
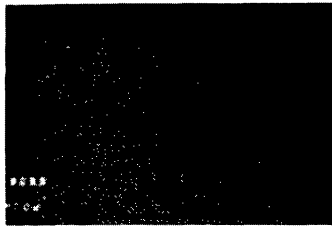


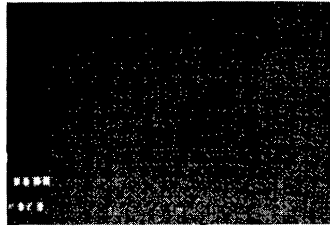
FIG. 2.—Continued

GHIGO (see page 384)

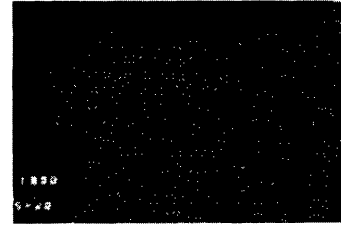
PLATE 34



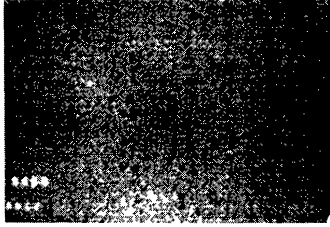
1618+108 B



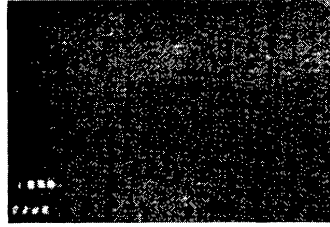
1620+127 R



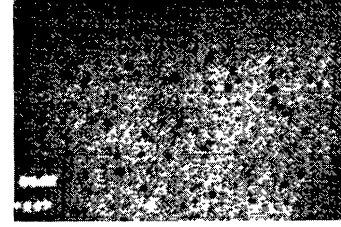
1621+140 R



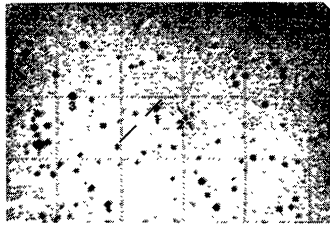
1627+138 R



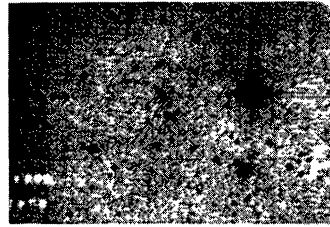
1648+198 R



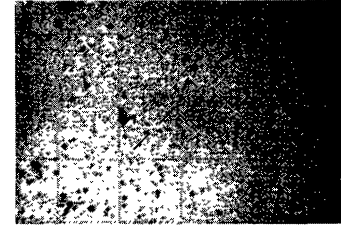
1653+163 R



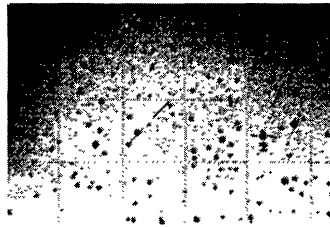
1706+188 R



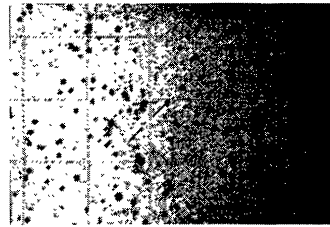
1707+117 R



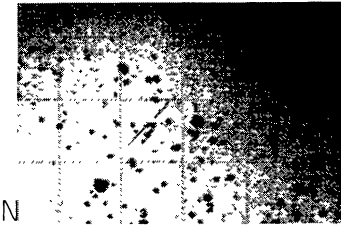
1717+130 R



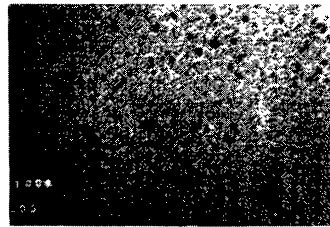
1720+176 R



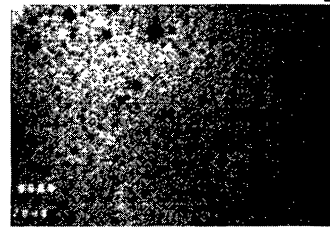
1727+174 R



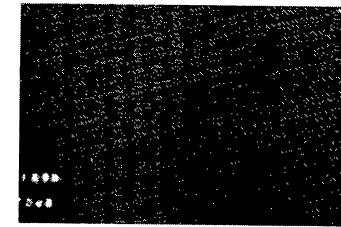
1729+150 R



1736+161 R



1739+107 R



1739+173 R

2' N
E

FIG. 2.—Continued

GHIGO (see page 384)

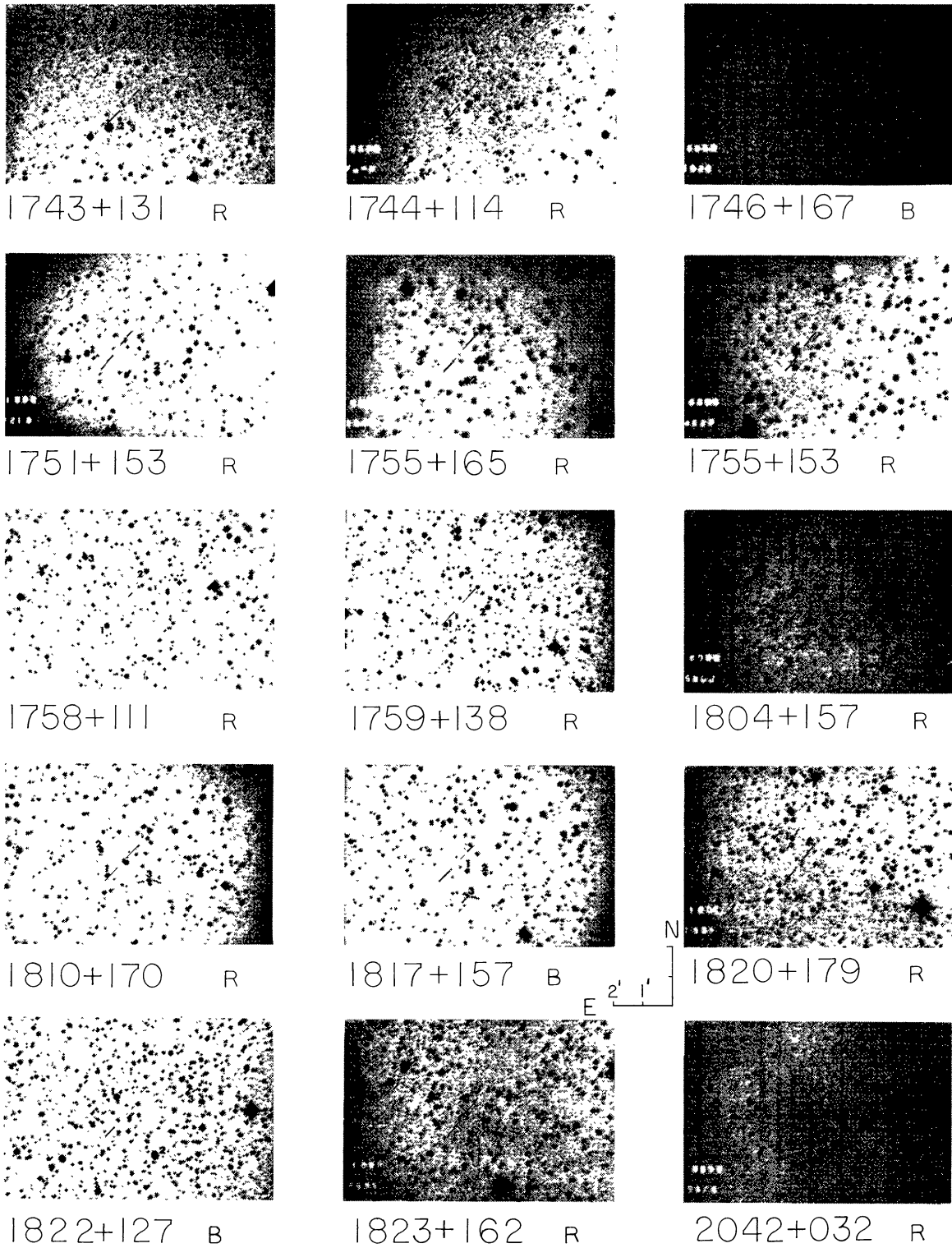
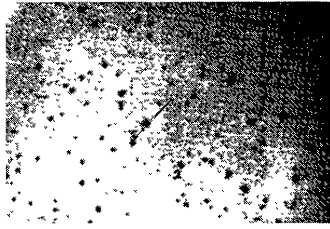


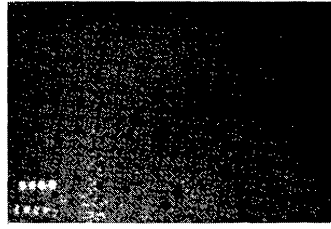
FIG. 2.—Continued

GHIGO (see page 384)

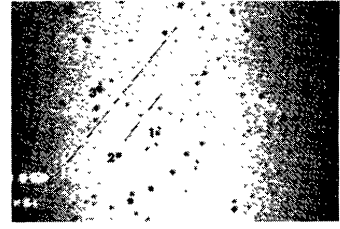
PLATE 36



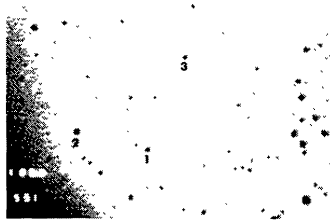
2055+088 R



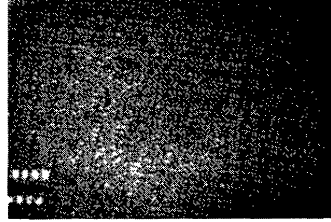
2058+019 B



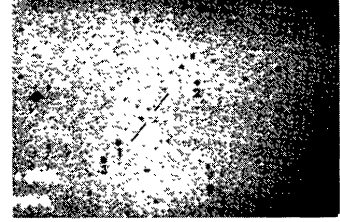
2121+028 B



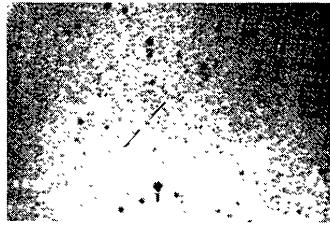
2122+052 R



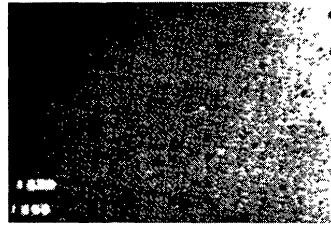
2206+019 R



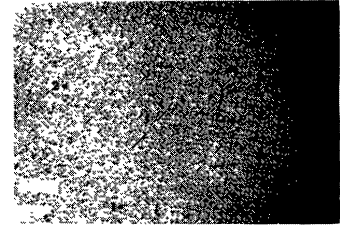
2215-000 R



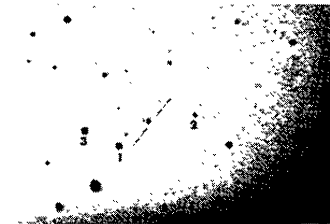
2225-019 R



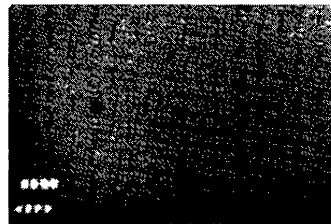
2238+033 R



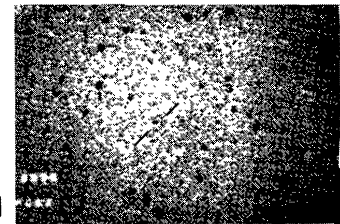
2259+083 R



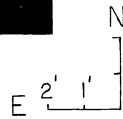
2304+038 B



2318+081 R



2330+015 R



2345+048 R

FIG. 2.—Continued

GHIGO (see page 384)