

THE SIGNIFICANCE OF ATYPICAL HIGH-SILICA IGNEOUS ROCKS

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A dissertation submitted to the faculty of the University of North Carolina at Chapel Hill in partial fulfillment of the requirements for the degree of Doctor of Philosophy in the Department of Geological Sciences.

Chapel Hill
2017

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ABSTRACT

RYAN E. FRAZER: The significance of atypical high-silica igneous rocks
(Under the direction of Drew S. Coleman)

The origins of high-silica igneous rocks are debated, as they may be products of high-degree fractional crystallization or low-degree partial melting. They may play a role in the generation of intermediate igneous rocks and are responsible for large, ash-rich volcanic eruptions. High-silica granites and rhyolites in the Sierra Nevada, California, and the Colorado Mineral Belt (CMB) are investigated using isotope geochemistry to better understand how they bear on these issues.

Zircon U-Pb geochronology identifies two intrusive suites comprising large volumes of high-silica granites emplaced in the mid-Cretaceous Sierra Nevada batholith: the 106-98 Ma Shaver Intrusive Suite (SIS) in the central part of the batholith, and the 103-100 Ma Kearsarge intrusive suite (KIS) on the Sierra Crest and Owens Valley. High-silica granites in both suites have relatively high concentrations of middle rare earth and high field strength elements. Data for these and other discrete high-silica plutons in the batholith suggest they were derived from titanite-free sources in the deep crust, unlike similarly felsic parts of zoned intrusive suites. Despite similar trace element signatures, SIS and KIS high-silica granites have divergent isotopic compositions. High-silica granites of the SIS have supracrustal O in zircon, crustal Sr and Nd whole rock isotopic compositions, and negative Ce anomalies suggesting the SIS granites may have been derived from oceanic sedimentary sources. In contrast, KIS granites have mantle-like

isotopic compositions. The location and geochemistry of the KIS suggests it may have resulted from backarc magmatism in the mid-Cretaceous Sierra.

Volcanic and plutonic rocks in the central CMB were emplaced during the Laramide orogeny and subsequent Oligocene-Eocene volcanic flare-up. Strontium and Nd data suggest the 63-39 Ma Twin Lakes pluton and igneous rocks as young as 24 Ma were derived from a persistent mafic lower crust or enriched lithospheric mantle source. In contrast, the ~35 Ma Grizzly Peak Tuff and resurgent plutons are isotopically dissimilar from each other and the CMB as a whole, suggesting derivation by partial melting of ancient felsic lower crust. This distinct source could account for the lack of Mo mineralization in the Grizzly Peak caldera relative to other high-silica parts of the CMB.

To Sarah,
who makes me a better person every day

ACKNOWLEDGMENTS

I am forever thankful to have had Drew Coleman as my advisor for seven years and two degrees. His knowledge and humor helped me through thick and thin, making me a better scientist, teacher, and person. I thank him and Ryan Mills, who mentored me as an officemate, committee member, and friend, for introducing me to coffee. I am indebted to Allen Glazner for recruiting me to Carolina and making me a more thoughtful geologist. Committee members Kevin Stewart and John Bartley provided valuable insights that improved this work considerably. I owe a debt of gratitude to Jade Star Lackey for introducing me to Sierra Nevada geology, and for support and patience as I continually expanded and reworked chapter 1 of this dissertation. Jennifer Wenner and Matt Zimmerer kindly provided samples that were essential to this work. My dissertation benefitted from conversations and help from numerous graduate students over the years, in particular Courtney Beck, Tom Chapman, Adam Curry, Maggie Ellis, Connor Lawrence, Roger Putnam, and Josh Rosera. I especially thank Sean Gaynor for his assistance with fieldwork in the Alabama Hills and Grizzly Peak, his lab and science expertise, and his sunny disposition.

This work was generously funded by the Geological Society of America, the National Science Foundation, Sigma Xi, the UNC Department of Geological Sciences Martin Fund, and a UNC Graduate School Dissertation Completion Fellowship.

Finally, I would like to thank my family. My parents, Brett and Tricia, who unwaveringly encouraged me and always seemed to know how to help. And my wife, Sarah, who supported me through all of this, and with whom I am so lucky to build a life and a family.

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CHAPTER 1: INTRODUCTION

The relationships between intrusive and extrusive igneous rocks have been discussed and disputed for centuries (e.g., Lyell, 1838; Daly, 1933). However, methodological improvements of the past two decades have breathed new life into the debate. For example, advances in ultraclean lab procedures and pre-treatment of zircons (e.g., Parrish, 1987; Mattinson, 2005) now permit accurate and precise dating of individual zircon grains and fragments to 0.1% (e.g., Samperton et al., 2015). These enhancements have led to the application of high-precision geochronology to multiple samples from single plutons, and dozens of samples dated from entire intrusive suites. Coleman et al. (2004) were among the first to date an intrusive suite at high sample density by U-Pb in zircon, finding that emplacement of the Tuolumne Intrusive Suite in the Sierra Nevada batholith spanned a ~10 Ma period, and a single pluton was intruded over at least 4 Ma. That study raised questions regarding the links between plutonic and volcanic rocks because it implied that large magma chambers may be ephemeral. Glazner et al. (2004) extended this line of thinking, suggesting that intrusion rates may govern the fate of magmas in the upper crust.

Numerous studies in the past decade have used Coleman et al. (2004) and Glazner et al. (2004) as touchstones for investigations of plutonic-volcanic connections. A review of those works is beyond the scope of this introduction. However, much of the work since 2004 has focused on intermediate-composition rocks (e.g., Bachmann et al., 2007; Lipman, 2007; Wotzlaw et al., 2013; Mills and Coleman, 2013; Lee and Bachmann, 2014). This is logical because granodiorite plutons are common in batholiths such as the Sierra Nevada (Bateman, 1992), and

dacites comprise the spectacular crystal-rich ash-flow tuffs known as monotonous intermediates (Hildreth, 1981). I contributed to this body of work with my Masters (Frazer et al., 2014), in which I investigated the postulated links between the 4500 km³ Mount Givens Granodiorite and compositionally and volumetrically similar monotonous intermediates, such as the 5000 km³ Fish Canyon Tuff. I concluded that zircon in both plutonic and volcanic intermediate rocks faithfully reflects their respective accumulation histories and therefore, despite their superficial similarities, the exposed Mount Givens Granodiorite accumulated too slowly to have supported a Fish Canyon-like super-eruption.

High-silica igneous rocks (i.e., greater than 70 wt% SiO₂; granites and rhyolites *sensu lato*) have not been overlooked in the volcano-pluton debate (e.g., Glazner et al., 2008; Bachmann and Bergantz, 2008; Coleman et al., 2012; Lundstrom and Glazner, 2016). High-silica igneous rocks are important because they may represent either very high-degree fractional crystallization (e.g., Tuttle and Bowen, 1958; Lee and Morton, 2015) or low-degree partial melting (e.g., Sisson et al., 2005; Stevens et al., 2007) of more mafic progenitors, and because they also bear on the elemental and isotopic compositions of intermediate-composition rocks that may result from magma mixing (Eichelberger, 1975; Wenner and Coleman, 2004; Gray et al., 2008). Debate continues as to whether chemical signatures in high-silica igneous rocks reflect fractional crystallization (Keller et al., 2015; Lee and Morton, 2015) or partial melting (Glazner et al., 2015; Clemens and Stevens, 2016). Another question in this debate is whether high-silica intrusive rocks represent new additions to the crust, or are the products of recycling pre-existing crust. Some models call for high-silica magmas to be the extensively recycled (Bindeman and Simakin, 2014), or to be the results of fractional crystallization after homogenization of primitive mantle and Proterozoic rocks in deep crustal hot zones (Lee and Bachmann, 2014). In contrast, others

suggest that high silica magmas may be generated through partial melting of juvenile mafic materials that were themselves recently derived from the mantle (Coleman et al., 1992; Wenner and Coleman, 2004), thus representing new crustal growth.

This dissertation investigates the occurrence and significance of high-silica granites and rhyolites from mid-Cretaceous parts of the Sierra Nevada batholith, California, and the Cenozoic central Colorado Mineral Belt using compositional and isotopic analyses of whole rocks and the mineral zircon. The first two chapters concern the mid-Cretaceous Sierra Nevada. Chapter one examines the compositionally diverse Shaver Intrusive Suite; whereas its largest pluton is granodiorite, like the Tuolumne and other Late Cretaceous Sierra suites, the suite also comprises nearly a dozen discrete high-silica granite plutons. Many of those plutons preserve chemical and isotopic evidence suggesting they represent significant crustal recycling. Chapter two investigates the origins of the Alabama Hills block east of the Sierra Nevada, which includes a high-silica granite that is now known to be ~102 Ma and therefore is similar in age to high silica granites on the Sierra crest. New age and isotopic data suggest those granites (part of the informal Kearsarge intrusive suite) may be mantle-derived and represent new crustal additions, in contrast to the Shaver granites. The Shaver and Kearsarge suites are part of a ~105-100 Ma magmatic event that appears to have favored granitic magma production relative to earlier and later magmatism in the batholith. High-silica granites such as these also appear to be compositionally appropriate end members for mixing with mafic magma to generate the voluminous granodioritic suites of the Sierra Crest magmatic event. Chapter three extends the analysis to extrusive rocks. The Grizzly Peak Tuff, an Eocene rhyolitic caldera-forming system in the central Colorado Mineral Belt, has a more enriched isotopic signature than older and younger igneous rocks adjacent to the caldera. Moreover, the tuff and its resurgent plutons are isotopically distinct from rocks throughout

the mineral belt. The tuff's isotopic composition suggests it was derived in from felsic Proterozoic lower crust, and thus represents a large degree of crustal recycling. In contrast, the relatively constant isotopic compositions of igneous rocks preceding and following eruption of the tuff imply a long-lived, voluminous source inferred to be either enriched lithospheric mantle or lower crust, or a combination of the two. With these studies, my dissertation seeks to demonstrate the utility high-silica rocks have for understanding deep magma sources and processes, including recycling of old crust and production of new crust.

CHAPTER 2: AGES AND SOURCES OF THE SHAVER INTRUSIVE SUITE: SNAPSHOTS OF CRUSTAL RECYCLING IN THE CENTRAL SIERRA NEVADA BATHOLITH, USA

INTRODUCTION

Resolving the hierarchy of magmatic processes leading to the assembly of plutons, intrusive suites, and batholiths is critical to deciphering the geodynamics of magmatic arcs. Studying these systems and subsystems in detail provides greater insight into the overarching tectonic controls of arc magmatism, the tempo and mode of magma transport, and the thermal and mass balance of arcs, which are loci of crustal growth and recycling. The Cordilleran batholiths of western North America record a protracted period of Mesozoic magmatism. Despite decades of study, the controls on magma production and emplacement, including the importance of tectonic controls (e.g. Glazner, 1991; Tikoff and Teyssier, 1992; Tobisch et al., 1995), particularly in the Cretaceous Sierra Nevada batholith of California (Bateman, 1992), remain poorly understood.

Recent studies have centered on the emplacement and hybridization of middle and upper crustal plutons. In particular, high-precision U-Pb zircon dating of intrusive suites belonging to the 98-86 Ma Sierra Crest magmatic event (Coleman and Glazner, 1997) has fueled debate about how mapped plutons may be built incrementally over ~10 Ma timescales, far longer than the thermal lifetime of magma bodies in the upper crust (Glazner et al., 2004; Annen, 2009). The Late Cretaceous suites comprise older mafic units, intermediate- to late-stage megacrystic granodiorites, and late stage granitic intrusions, with suites recording both mantle and crustal sources

(Gray et al., 2008) or simply an enriched mantle source (Coleman et al., 1992). In contrast, Early Cretaceous magmatism in the western parts of the batholith resulted in chemical and temporal variations that are not as well organized, with isotopic data suggesting an increase in crustal inputs over time (Lackey et al., 2012; Holland et al., 2013).

The centrally-located Shaver Intrusive Suite (SIS; Fig. 1) presents an opportunity to assess the transition between contrasting older and younger magmatic styles in the batholith because it superficially shares qualities of both Early and Late Cretaceous magmatism in the Sierra. The SIS is similar to the older suites in that it is not concentrically zoned, yet it is generally more felsic than the western suites (Bateman, 1992). Like many of the younger intrusive suites of the batholith, the SIS is predominantly composed of granodiorite. However, it contrasts with many of the Late Cretaceous suites because the granitic and dioritic plutons of the SIS are discrete intrusions lacking gradational contacts (Bateman and Wones, 1972a). It is also an intriguing location to study changing magmatic styles and sources because the SIS is at the nexus of several proposed geochemical provinces and boundaries, including the $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.706$ line, the Panthalassan/North American lithospheric break, and the strongly contaminated and reduced zone (Kistler and Peterman, 1973; Ague and Brimhall, 1987, 1988; Kistler, 1990). Thus, it may represent a critical transition in the Cretaceous Sierran arc.

GEOLOGIC BACKGROUND

The SIS (Fig. 1) is located in the central Sierra Nevada batholith of California, flanked on the west by the 124–105 Ma Fine Gold Intrusive Suite (Lackey et al., 2012) and on the east by the 98–91 Ma Mount Givens Granodiorite (Frazer et al., 2014).

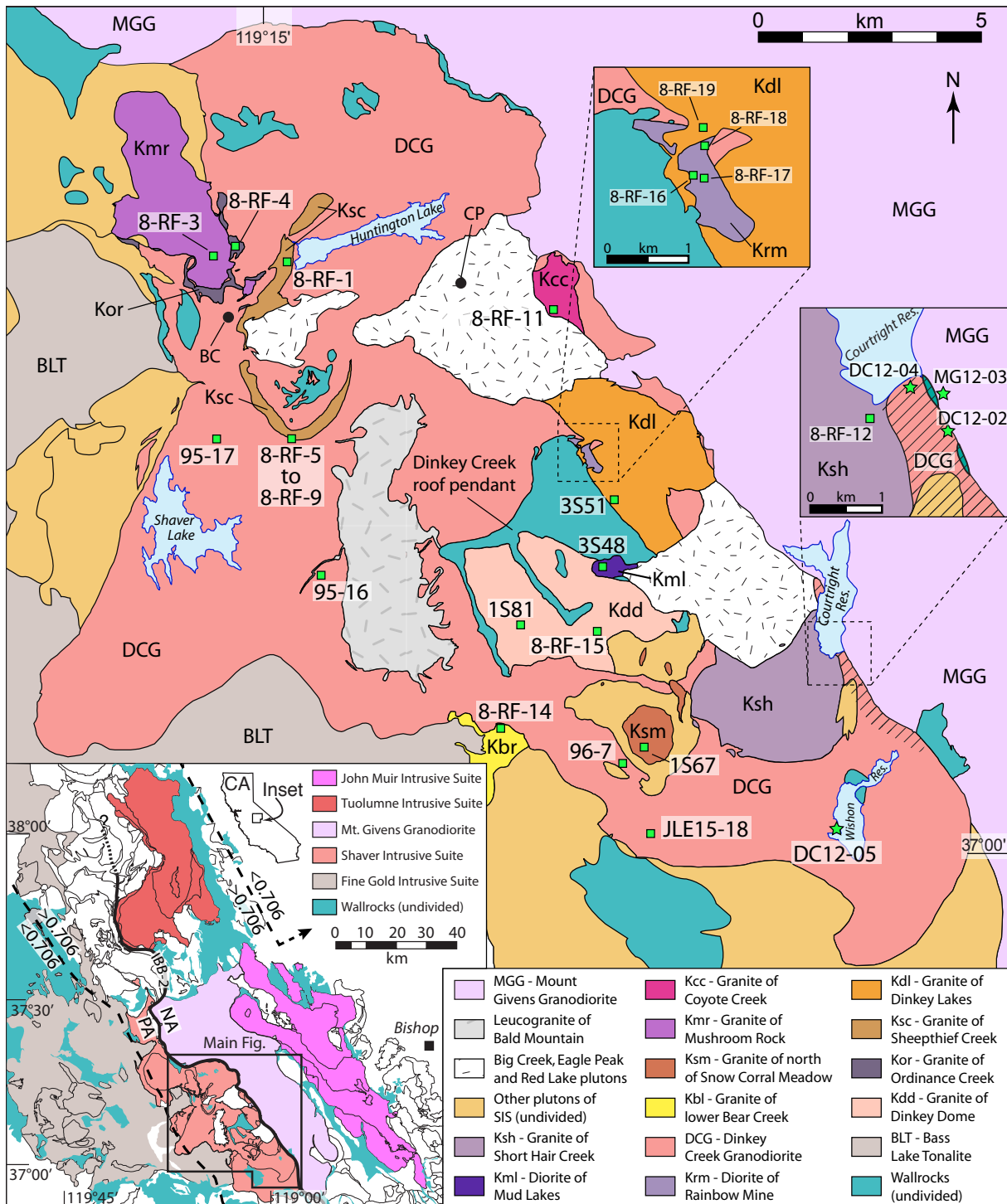


Figure 1. Generalized bedrock geologic map of the Shaver Intrusive Suite. Informal names of individual plutons are given in the legend and Table 1. Inset gives location of Shaver Intrusive Suite relative to other Cretaceous intrusive suites of the central Sierra Nevada as well as approximate locations of the $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.706$ line, intrabatholithic break 2, and the PA-NA (Panthalassan-North American) boundary after Kistler and Peterman (1973), Kistler (1990, 1993), Saleeby and Busby (1993) and Lackey et al. (2008, 2012). Green symbols: samples analyzed in this study; green stars: samples dated by ID-TIMS; black circles: locations of Cenozoic volcanic pipes (BC – Big Creek; CP – Chinese Peak). Maps after Bateman (1965), Bateman et al. (1971), Bateman and Wones (1972a), Lockwood and Bateman (1976), Bateman (1992) and Lackey et al. (2008).

Wallrocks

Metamorphic wallrocks in and around the SIS are dominated by quartzite, with subordinate pelitic hornfels, schist, calc-silicate hornfels, and marble (Huber, 1968; Bateman et al., 1971; Bateman and Wones, 1972a; Lockwood and Bateman, 1976). The Dinkey Creek roof pendant is the most prominent wallrock mass exposed in the SIS, and it has been proposed to belong to terranes of various ages from the late Proterozoic to early Cambrian Snow Lake block (Lahren and Schweickert, 1989), to the Triassic-Jurassic Kings sequence (Kistler, 1993; Saleeby and Busby, 1993).

Shaver Intrusive Suite plutons

The SIS comprises numerous intermediate and felsic plutons plus scattered diorite bodies (Bateman, 1992). It is dominated by the Dinkey Creek Granodiorite (DCG), which ranges in composition from tonalite to granite, but is typically granodiorite in both equigranular and megacrystic facies (Bateman, 1992). The DCG contains abundant mafic enclaves that are hypothesized to result from magma mingling and stirring of a stratified magma chamber (Barbarin et al., 1989; Dorais et al., 1990; Cruden et al., 1999). Mafic orbs and comb layering are present near Deer Creek in a tonalitic portion of the DCG (Moore and Lockwood, 1973; Dodge and Kistler, 1990). Two shear zones are preserved at the DCG's margins, including the Kaiser shear zone in the north and the Courtright shear zone in the southeast (Tobisch et al., 1995). The Courtright shear zone is particularly well studied and records a two-stage strain history of weak extension followed by contraction, around 90 Ma (Tobisch et al., 1993; Renne et al., 1993). The DCG has previously been dated by isotope dilution thermal ionization mass spectrometry (ID-TIMS) at 104 Ma in its southwestern lobe (Stern et al., 1981) and 102 ± 1 Ma in the Courtright shear zone

(Fig. 1; Tobisch et al., 1993). Beck (2013) reported chemically abraded (CA) ID-TIMS data for single-zircon fractions in two samples from the DCG ranging from 101.4–100.6 Ma. Using laser ablation multi collector inductively coupled plasma mass spectrometry (LA-MC-ICP-MS), Lackey et al. (2012) dated a mafic enclave and its granodiorite host from the southern portion of the DCG, yielding indistinguishable ages of 99.4 ± 1.3 Ma and 101.5 ± 1.2 Ma, respectively.

This study focuses only on plutons within and adjacent to the DCG; additional large plutons of the SIS of Bateman (1992), including the granodiorites of McKinley Grove, Whiskey Ridge, Stevenson Creek, and the granite of Shuteye Peak, extend to the north and south of the DCG and are not considered here in part because of their uncertain ages. For example, the granite of Shuteye Peak west of Shaver Lake was dated by Stern et al. (1981) at 102 Ma, whereas Lackey et al. (2012) dated a different exposure of the same pluton at ~114 Ma.

Previous field studies proposed that the DCG is intruded by the all of the areally subordinate plutons in the SIS, including the granites of Dinkey Dome, Short Hair Creek, Sheepthief Creek, lower Bear Creek, Mushroom Rock, north of Snow Corral Meadow, and Ordinance Creek (Hamilton, 1956; Bateman et al., 1971; Bateman and Wones, 1972a; Lockwood and Bateman, 1976; Guy, 1980; Bateman, 1992; Cruden et al., 1999). This group of high-silica plutons will hereafter be informally referred to as the Shaver granites. Bateman (1992) interpreted the DCG foliations that parallel the Shaver granites to suggest that the smaller felsic plutons postdate the DCG, deforming the DCG as they were emplaced. However, beyond field relationships, little work has been done on the Shaver granites.

The Shaver granites are discontinuous and feature at least six plutons as defined by Bateman and Wones (1972a; 1972b) using texture and mineralogy (Fig. 1). Several plutons have separate exposures or are previously unnamed. Herein each exposure is informally named after

a landmark within it or nearby. Some of the Shaver granites are associated with small bodies of diorite, of unknown age (Bateman, 1992). Two of the Shaver granites are nested: the Mushroom Rock pluton is mantled by the slightly more mafic Ordinance Creek pluton, and the granite north of Snow Corral Meadow is surrounded by the granodiorite of McKinley Grove.

Shaver Intrusive Suite petrology and geochemistry

Past geochemical analyses of the Shaver granites show they are more siliceous (73–78 wt% SiO₂) than the DCG (48–71 wt% SiO₂; average is 65 wt% SiO₂). Moreover, some of the granites have magmatic garnet and muscovite (Fig. 2), and more unusual accessory mineral assemblages locally that include andalusite, sillimanite, molybdenite, and/or urananite (Hamilton, 1956; Bateman and Wones 1972b; Guy, 1980; Ague and Brimhall, 1988; Hinke, 2002, Lackey et al. 2006). Cordierite, a key indicator of S-type granites, is not recognized in the SIS (Ague and Brimhall, 1987). Ague and Brimhall (1988) assigned the SIS to a belt of “strongly contaminated and reduced” plutons on the basis of mineralogy and mineral compositions, particularly the greater amounts of ilmenite over magnetite, the relative lack of titanite, and the presence of Fe-rich biotite.

Isotopically, the DCG has relatively high initial ⁸⁷Sr/⁸⁶Sr (⁸⁷Sr/⁸⁶Sr_i) values compared to nearby suites (Dodge et al., 1986; Kistler, 1990). The SIS is east of the ⁸⁷Sr/⁸⁶Sr_i = 0.706 line, inferred to be an expression of the premagmatic boundary between Proterozoic North American continental lithosphere to the east and accreted arc terranes to the west (Fig. 1; Kistler and Peterman, 1973). There are no published isotopic data for the Shaver granites.

Sub-arc xenolith suites

Cenozoic alkali basalt lavas of the San Joaquin volcanic field erupted through the SIS

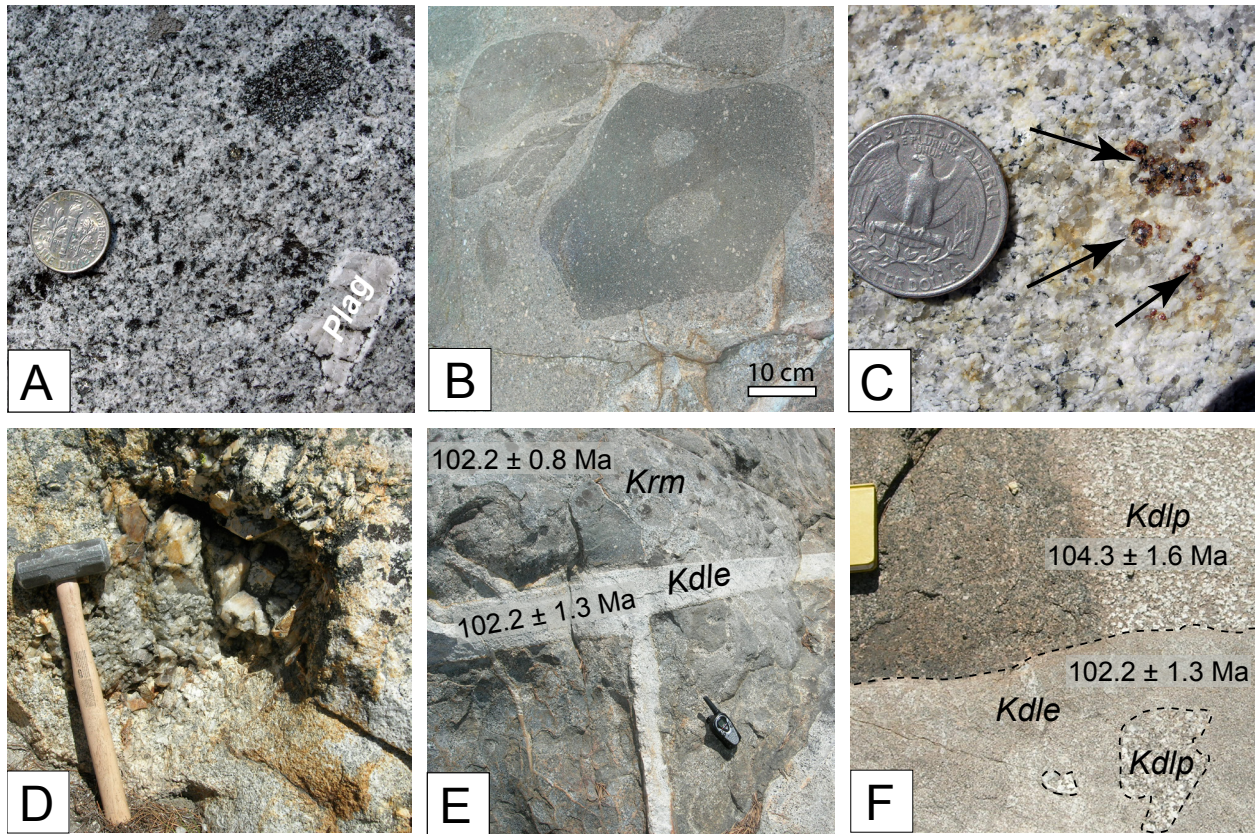


Figure 2. Field relations in the SIS. (A) Ordinance Creek pluton showing rapakivi rims on feldspar. (B) Mafic magmatic enclaves within the diorite of Mud Lakes. (C) Red garnet glomerocrysts in the Coyote Creek pluton. (D) Large miarolitic cavity in the Coyote Creek pluton with inward radiating euhedral K-feldspar and quartz crystals. Hammer handle length approximately 30 cm. (E) Dike of the equigranular facies of the Dinkey Lakes pluton intruding the diorite of Rainbow Mine. Laser ablation weighted mean ages shown were not acquired on this outcrop. Radio length approximately 15 cm. (F) Field relationships between porphyritic and equigranular facies of the Dinkey Lakes pluton. Laser ablation weighted mean ages shown were not acquired on this outcrop. Length of side of field notebook approximately 12 cm.

host nearly all known xenoliths of lower crustal and upper mantle origin from the Sierran arc (Fig. 1; Dodge et al. 1986, 1988), including Cretaceous-aged samples (Ducea and Saleeby, 1998). These xenolith suites provide understanding of the dynamics of magma formation and extraction in the Sierran arc (Mukapadahay and Manton, 1994, Ducea, 2001; 2002; Chin et al., 2014). Thus, the SIS is one of the few locations in the Cordillera where the connectivity of the petrogenetic record between coeval granitoids and xenoliths from the base of the arc can be evaluated.

METHODS

Sample collection and preparation

Samples from the SIS were collected for zircon separation and whole rock chemical analyses. In addition, samples of the Mount Givens Granodiorite corresponding to those dated by Frazer et al. (2014) were analyzed for their whole rock isotopic compositions.

Geochronology aliquots were crushed using standard techniques and zircon was isolated by gravimetric (water table and methylene iodide) and magnetic concentration methods. Analyzed zircons were hand-picked under a binocular microscope from the non-magnetic split on a Frantz magnetic separator. Grains were selected to be clean, inclusion-free, unbroken and typical of the population. Whole rock geochemical aliquots were pulverized to powder using standard crushing techniques outlined below.

U-Pb geochronology

LA-MC-ICP-MS

Most of the geochronologic data were collected at the University of Arizona LaserChron Center by LA-MC-ICP-MS. Zircon grains were mounted along with chips of Sri Lankan zircon standard SL-2 (Gehrels et al., 2008) in 1” epoxy rounds. Most mounts were initially characterized by scanned color cathodoluminescence (CL) at California State University, Bakersfield on a Hitachi S-3400N SEM at 10.0 kV.

Laser ablation analyses of all samples except those from the DCG and the Courtright shear zone were performed with a New Wave/Lambda Physik DUV193 Excimer laser operating at a wavelength of 193 nm using a spot diameter of 25-35 microns. The ablated material was carried with helium gas into the plasma source of a GV Instruments Isoprobe, which was equipped

with a flight tube of sufficient width that U, Th, and Pb isotopes were measured simultaneously. All measurements were made in static mode, using Faraday detectors for ^{238}U and ^{232}Th , an ion-counting channel for ^{204}Pb , and either faraday collectors or ion counting channels for ^{208}Pb - ^{206}Pb . Ion yields were ~ 1 mv per ppm. Each analysis consisted of one 20-second integration on peaks with the laser off (for backgrounds), 20 one-second integrations with the laser firing, and a 30 second delay to purge the previous sample and prepare for the next analysis. The ablation pit was ~ 15 microns in depth.

Zircon U-Pb isotopic data from samples 95-16, 95-17 and 96-7 from the DCG were collected using a New Wave 193 nm ArF laser ablation system coupled to a Nu Plasma HR inductively-coupled plasma mass spectrometer (ICP-MS) at the University of Arizona following the methods of Gehrels et al. (2008) and Lackey et al. (2012). See <http://sites.google.com/a/laserchron.org/laserchron/> for additional information.

For all laser ablation samples, inter-element fractionation of Pb/U is generally $\sim 20\%$, whereas fractionation of Pb isotopes is generally less than 2%. In-run analysis every fifth measurement of fragments of SL-2, a Sri Lankan zircon crystal with known age of 564 ± 4 Ma (2σ error), is used to correct for this fractionation (Gehrels et al., 2008). The uncertainty resulting from the calibration correction is generally $\sim 1\%$ (2σ) for both $^{206}\text{Pb}/^{207}\text{Pb}$ and $^{206}\text{Pb}/^{238}\text{U}$ ages. The analytical data are reported in Appendix 1. Uncertainties shown in these tables are at the 2σ level, and include only measurement errors. The systematic error, which includes contributions from the standard calibration, age of the calibration standard, composition of common Pb, and U decay constants, is generally 1–2% (2σ). Common Pb correction was accomplished using the measured ^{204}Pb and assuming an initial Pb composition using the model of Stacey and Kramers (1975) with uncertainties of 1.0% for $^{206}\text{Pb}/^{204}\text{Pb}$ and 0.3% for $^{207}\text{Pb}/^{204}\text{Pb}$.

ID-TIMS

Four samples (DC12-02, MG12-03, DC12-04, and DC12-05) were analyzed by CA-ID-TIMS on a VG Sector 54 in the Department of Geological Sciences at the University of North Carolina at Chapel Hill following the sample preparation and analytical methods of Frazer et al. (2014). Data processing and age calculations were completed in the applications Tripoli and ET_Redux (Bowring et al., 2011; McLean et al., 2011). Corrections for initial Th/U disequilibrium (Mattinson, 1973; Schmitz and Bowring, 2001) were made in ET_Redux assuming the measured whole rock Th/U value (Table 3) for each sample approximates the magmatic value; this adjustment increased the ages of individual analyses by up to 100 ka.

$\delta^{18}\text{O}$ analysis

Bulk aliquots of zircon (4–5 mg) were picked and cleaned with HNO_3 , HCl , and HF acid, with the latter removing radiation-damaged material from the zircon grains, which can alter $\delta^{18}\text{O}$ (Valley, 2003). Zircon crystals were powdered with a tungsten carbide mortar and pestle to facilitate fluorination (Lackey et al., 2008). Oxygen isotopes were analyzed by laser fluorination at the Stable Isotope Lab at the University of Oregon using a 35W New Wave CO_2 IR laser in a reagent atmosphere of bromine pentafluoride to liberate O_2 with the F^- anion. Liberated O_2 was passed through a hot Hg pump to remove excess F, then converted to CO_2 by reaction with a hot graphite rod. CO_2 gas was analyzed for $\delta^{18}\text{O}$ with a Finnigan MAT 253 mass spectrometer. Values are reported as $\delta^{18}\text{O}$ in ‰ relation to Vienna Standard Mean Ocean Water (VSMOW). Unknowns were corrected to Gore Mountain garnet (accepted value = 5.80‰; Valley et al., 1995) analyzed throughout the analytical session.

Whole rock elemental geochemistry

Whole rock powders of all samples except DC12-02, MG12-03, DC12-04, and DC12-05 were prepared in a Rocklabs® tungsten carbide head and mill. Powdered sample and flux were mixed at a 2:1 ratio, typically 3.5 g powder to 7.0 g dilithium tetraborate. The vortex-blended mixture was fused to a glass bead in a graphite crucible at 1000°C for 10 minutes, reground, fused a second time, polished on diamond laps, and analyzed. Major oxides and 17 trace elements (Ba, Ce, Cr, Cu, Ga, Nb, Ni, Pb, Rb, Sc, Sr, Th, U, V, Y, Zn, Zr) are analyzed on the same fused bead using a 3.0 kW Panalytical Axios WD-XRF at Pomona College equipped with PE, LiF 200, LiF 220, GE, and PX1 crystals. Concentrations were determined using reference calibration curves defined by fifty-five certified reference materials that span a range of natural igneous, metamorphic, and sedimentary rock compositions. Oxides (in wt%) are totaled pre-normalization based on iterating analysis after determination of initial loss on ignition in the first analysis. Total Fe is reported as Fe_2O_3 . Typical analytical uncertainties for major elements in absolute wt%, determined by fusion of duplicate powders, are $\pm 0.15\%$ (SiO_2); 0.06–0.07% (Fe_2O_3 , Al_2O_3); 0.02–0.05% (Na_2O , K_2O , MgO , CaO); $<0.004\%$ (TiO_2 , MnO , P_2O_5). Trace element uncertainties at $\pm 2\sigma$ vary, with most trace elements reproducible at better than 2 ppm: (Cu, Ga, Nb, Pb, Rb, Th, U, Y, Zr). Others are reproducible at 3 ppm (Cr, Sc, Sr, V), and a handful have higher uncertainties, in parentheses in ppm: Zn (6); Ni (7); Ba (40). Rare earth elements for samples 8-RF-1, 8-RF-3, 8-RF-19, 1S67, 1S81, 3S48, and 3S51 were measured by ICP-MS at Vrije University using the methods of Klaver et al. (2015).

Whole rock aliquots of samples DC12-02, MG12-03, DC12-04, and DC12-05, plus the samples of Frazer et al. (2014), were crushed to powder in a SPEX Shatterbox® alumina swing mill and analyzed at Actlabs (Ontario, Canada). Samples were dissolved by fusion in a lithium

metaborate/tetraborate mixture. Major elements and Ba, Sr, Y, Zr, Sc, Be and V were analyzed by inductively coupled plasma-optical emission spectroscopy (ICP-OES), with remaining trace elements and REE analyzed by ICP-MS. Relative uncertainties ($\pm 2\sigma$) for major elements were less than 2% for all oxides except MgO (3%), MnO (5%), and P₂O₅ (16%). Trace elements analyzed at Actlabs that are reproducible ($\pm 2\sigma$) at better than 2 ppm include Ag, Be, Cs, Ge, Hf, Sb, Sn, Ta, Th, Tl, U, and W. Other elements are reproducible within 3 ppm (Co, Mo, Nb, Sc), 4 ppm (Ga, Y) and others have higher uncertainties, in parentheses in ppm: Pb (9); Cu (14); Rb and Zr (15); V (19); Ni (23); Sr (25); Ba (31); Zn (42); Cr (49). Rare earth elements reproducible ($\pm 2\sigma$) at 0.2 ppm or better are Eu, Ho, Lu, Tb, and Tm; REE reproducible at 0.2–0.5 ppm or better are Dy, Er, Gd, Pr, Sm, and Yb; others are reproducible at higher uncertainties, in ppm: Nd (1.2); La (1.5); Ce (2.1).

Strontium and neodymium isotopic analyses

Whole rock powders of samples 8-RF-1, 8-RF-3, 8-RF-19, 1S67, 1S81, 3S48, and 3S51 were dissolved in Teflon® bombs and separated using column chromatography following standard procedures outlined in Heumann and Davies (2002). Both Sr and Nd isotope ratios were measured in static mode on a ThermoElectron Triton Plus TIMS at Vrije University. The Sr fractions were loaded onto Re filaments using a TaF₅ activator to enhance ionization. Strontium isotope ratios were measured using a static multi-collection routine. An analysis consisted of 20 blocks of 10 cycles with an integration time of 8 s per cycle. Ratios of ⁸⁷Sr/⁸⁶Sr and ⁸⁴Sr/⁸⁶Sr were corrected for mass fractionation using an exponential law and ⁸⁶Sr/⁸⁸Sr ratio of 0.1194. Replicate analyses (n = 58) of the Sr standard NBS 987 on load sizes of 100 ng yielded average ⁸⁷Sr/⁸⁶Sr and ⁸⁴Sr/⁸⁶Sr ratios of 0.710242 ± 0.000008 (2σ) and 0.056492 ± 0.000004 (2σ). Neodymium

isotopes were measured on double Re filaments and were corrected for mass fractionation using an exponential law and $^{146}\text{Nd}/^{144}\text{Nd}$ ratio of 0.7219. Analyses of the standard BHVO-2 ($n = 9$) yielded $^{143}\text{Nd}/^{144}\text{Nd} = 0.512982 \pm 0.000005$ (2σ) and $^{145}\text{Nd}/^{144}\text{Nd} = 0.348405 \pm 0.000003$ (2σ); analyses of the standard BCR-2 ($n = 4$) yielded $^{143}\text{Nd}/^{144}\text{Nd} = 0.512636 \pm 0.000004$ (2σ) and $^{145}\text{Nd}/^{144}\text{Nd} = 0.348406 \pm 0.000004$ (2σ).

Whole rock powders of samples 8-RF-4, 8-RF-6 through -9, 8-RF-11, 8-RF-12, 8-RF-14, 8-RF-16, 8-RF-17, 8-RF-19, DC12-02, MG12-03, DC12-04, and DC12-05, and those of Frazer et al. (2014) were dissolved in HF and HNO_3 in Teflon® (Parr) bombs at 180°C for 48 hours. Following dissolution, samples were dried down to release Si, then immediately fluxed in concentrated HCl for 4-8 hours. Samples were subsequently aliquoted for Sr and Nd chemistry. Strontium was purified using Sr-spec cation exchange resin after the methods of Lundblad (1994) and loaded on single Re filaments with TaF_5 . Neodymium was purified using a three-stage column chemistry procedure after Harvey and Baxter (2009) and loaded on single Re filaments in a Ta_2O_5 - H_3PO_4 slurry. Strontium isotopic analyses were accomplished on a VG Sector 54 TIMS and Nd isotopes were analyzed on an Isotopx Phoenix TIMS, both housed at the University of North Carolina at Chapel Hill. Strontium was analyzed as a metal in dynamic multicollector mode with $^{88}\text{Sr} = 3\text{V}$. Neodymium was analyzed as an oxide in dynamic multicollector mode with $^{142}\text{Nd}^{16}\text{O} = 1\text{V}$. Isotope ratios for both elements were corrected for mass fractionation using an exponential law. Strontium isotopic ratios were normalized to $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$; Nd isotopic ratios were normalized to $^{146}\text{Nd}/^{144}\text{Nd} = 0.7219$. Replicate analyses of the NBS 987 Sr standard yielded $^{87}\text{Sr}/^{86}\text{Sr} = 0.710265 \pm 0.000015$ (2σ ; $n = 26$). Replicate analyses of the Nd standard JNdi yielded $^{143}\text{Nd}/^{144}\text{Nd} = 0.512098 \pm 0.000012$ (2σ ; $n = 18$).

All isotopic data are corrected to initial values using concentration and geochronologic

data presented in this study. Strontium data from the two labs are related by correcting present-day $^{87}\text{Sr}/^{86}\text{Sr}$ ratios of unknowns measured in each lab by the deviation of each lab's NBS 987 measurement from $^{87}\text{Sr}/^{86}\text{Sr} = 0.710250$, then corrected for decay to initial values. Neodymium data for the two labs are related to the La Jolla Nd standard, then corrected for decay to initial values. UNC Nd unknowns are corrected using a value of 0.999967, based on the difference between JNdi measurements in this study and La Jolla-normalized measurements of JNdi by Tanaka et al. (2000). Vrije University Nd unknowns are corrected using a value of 0.999997, based on the average difference between BCR-2 and BHVO-2 measurements in this study and the La Jolla-normalized measurements of the same standards by Weis et al. (2006). For age corrections of samples dated by LA-ICP-MS, the weighted mean age is used; for samples dated by ID-TIMS, the midpoint of the stated age range is used.

RESULTS

Cathodoluminescence

Cathodoluminescence reveals that most zircons have typical magmatic zonation, varying in style from sample to sample (Fig. 3; cf. Corfu et al., 2003). Many samples have oscillatory zonation, typified by the Sheepthief Creek and north of Snow Corral Meadow plutons. Sector zonation is present in more mafic samples, such as the two diorite plutons. Because no systematic pattern of U-Pb age is discernible with CL, images are shown without age spot locations. Laser ablation spots were typically placed midway through grains in order to cover the most representative domain (e.g., sample 8-RF-5 shows laser spots on the grains).



Figure 3. Cathodoluminescence (CL) images of representative zircons from dated samples. Most zircons have oscillatory or sector zoning, indicative of typical magmatic growth. Some notable features, such as inclusions or obvious inherited cores, are labeled. Images for all samples except 8-RF-5 are color CL collected at California State University – Bakersfield. Sample 8-RF-5 analyzed with black and white CL at Pomona College. Images for all samples are at the same scale.

U-Pb analyses

LA-MC-ICP-MS

Laser ablation data are reported as weighted mean ages with uncertainty given at the 95% confidence level after Ludwig (2003). Analyses were excluded from age interpretations in cases where grains were evidently inherited or antecrystic (Miller et al., 2007), or showed signs of obvious Pb-loss. Age interpretations are conservative and omit few analyses so as to not discount geologic complexities in the data. For samples 8-RF-4 and 8-RF-12, two age interpretations are provided because several high-precision analyses have significant weight on calculated weighted

mean ages. However, the preferred ages are those that include more analyses, yielding higher mean squares of weighted deviates (MSWD), thus reflecting potential zircon spectra complexity.

Zircon U-Pb data from most LA-MC-ICP-MS samples are concordant. Two age groups are evident: the Dinkey Dome pluton at 119-117.5 Ma, and the bulk of the SIS from ca. 106.5–98 Ma (Fig. 4; Table 1; Appendix 1). Three SIS samples (8-RF-4, 8-RF-5, and 3-S-51) are older than the DCG outside of uncertainty; one sample (8-RF-12) was younger than the DCG. Sample 1S81 of the Dinkey Dome pluton yielded the largest spectrum of ages. Three xenocrystic analyses include an analysis with a $^{207}\text{Pb}/^{206}\text{Pb}$ age of 2784 ± 63 Ma, a normally discordant analysis with a $^{207}\text{Pb}/^{206}\text{Pb}$ age of 979 ± 88 Ma, and a concordant analysis with a $^{206}\text{Pb}/^{238}\text{U}$ age of 140 ± 4 Ma. Several other fractions are subtly normally discordant near 120 Ma and have Proterozoic $^{207}\text{Pb}/^{206}\text{Pb}$ ages (Appendix 1). In addition, four analyses are 66–60 Ma, and likely reflect Pb-loss. Sample 8-RF-5 of the Sheepthief Creek pluton yielded an inherited zircon with a $^{207}\text{Pb}/^{206}\text{Pb}$ age of 1427 ± 62 Ma.

The Dinkey Lakes and Coyote Creek plutons, originally mapped as parts of the granite of Dinkey Dome (Bateman and Wones, 1972a), are instead coeval with the rest of the SIS. Two samples of the Dinkey Lakes pluton yield ages of 104.3 ± 1.6 and 102.2 ± 1.3 Ma, with the younger age overlapping the DCG. The $^{206}\text{Pb}/^{238}\text{U}$ weighed mean age of 100.1 ± 1.9 Ma for the Coyote Creek pluton (8-RF-11) also overlaps the DCG; however, most individual analyses were slightly normally discordant so it is not considered as robust as the other U-Pb data.

ID-TIMS

Three samples from the DCG and one sample of the adjacent Mount Givens Granodiorite were dated by ID-TIMS (Figs. 1, 5). All ages are concordant within 2σ analytical uncertainty

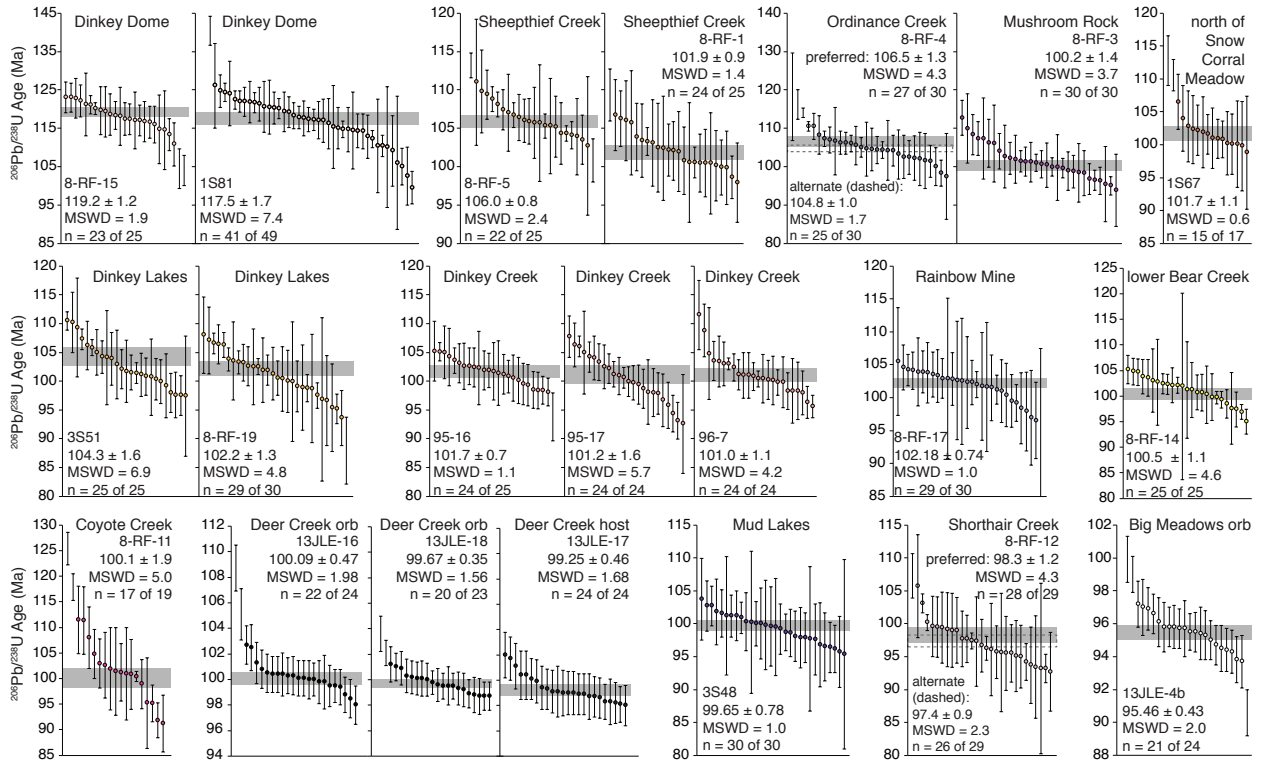


Figure 4. Results of laser ablation geochronology shown for each sample as $^{206}\text{Pb}/^{238}\text{U}$ age-sorted spectra. Analyses that are either obviously xenocrystic or affected by Pb-loss are omitted here; see Supplemental File for complete data. Note different ordinate age scales. Samples that are from same pluton (e.g., Dinkey Lakes) or similar locations (e.g., Ordinance Creek and Mushroom Rock) are shown at same scale. Individual analyses shown with corresponding 2σ analytical uncertainty; analyses not included in weighted mean calculations are shown as error bars only. Preferred weighted mean age interpretations and uncertainties at 95% confidence level (after Ludwig, 2003) shown by horizontal gray bars for each sample. Alternative interpretations of weighted mean ages and uncertainties for two samples given by dashed boxes; analyses omitted from alternative interpretations indicated by empty circles. Data for Deer Creek and Big Meadows (Eisenberg, 2014) included for comparison.

(Appendix 2). Preferred ages for ID-TIMS data are given as ranges due to the increasing difficulty of interpreting high-precision zircon data without corroborating compositional information (Samperton et al., 2015). Subsequent interpretations do not depend on precise crystallization ages. Note the age ranges provided here may not capture the full crystallization history of the samples (Glazner and Sadler, 2016), assuming all analyses are free from inheritance and Pb-loss. For several samples, weighed mean ages and 2σ analytical uncertainties are provided for comparison.

Sample DC12-02, collected in the DCG in the most mylonitized portion of the Courtright shear zone ~45 meters from the contact with the Mount Givens Granodiorite, yields 6 zircons

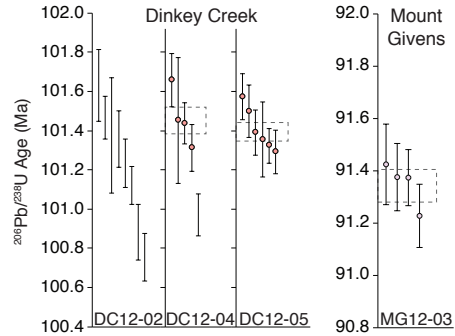


Figure 5. Results of ID-TIMS geochronology from the Courtright shear zone, shown as $^{206}\text{Pb}/^{238}\text{U}$ age-sorted spectra. Note different age scales on ordinates. All analyses are concordant and corrected for initial Th-disequilibrium using the whole rock Th/U measured for each sample. Preferred age interpretations are the age range of each sample including 2σ analytical uncertainties. Alternative age interpretations (dashed boxes) are weighted mean ages and corresponding 2σ analytical uncertainties calculated using analyses indicated by circles; see text for ages and explanation. Note that the Mount Givens Granodiorite is not part of the Shaver Intrusive Suite.

spread along concordia from 101.8–100.6 Ma (Fig. 5). Sample DC12-04 was collected in the further from the contact with the Mount Givens pluton and was less deformed. Four of five zircons yield an age range of 101.8–101.2 Ma, with a weighted mean age of 101.45 ± 0.07 Ma (MSWD = 4.9). One younger zircon is likely affected by Pb-loss. Sample DC12-05, collected several km west of the mapped shear zone, did not appear deformed in hand sample. Five zircons range from 101.7–101.2 Ma and yield a weighted mean age of 101.39 ± 0.05 Ma (MSWD = 3.5). A sample of the Mount Givens Granodiorite collected 20 m east of the shear zone and the DCG yields four zircons from 91.6–91.1 Ma and a weighted mean age of 91.34 ± 0.06 Ma (MSWD = 1.8). This is slightly older than other samples of the pluton from the northern end of Courtright Reservoir (Frazer et al., 2014). Samples DC12-02 and DC12-04 are within uncertainty of the 102 ± 1 Ma age of Tobisch et al. (1993).

Bulk geochemistry

Most samples from the SIS follow linear trends on Harker diagrams (Fig. 6; Table 2), except for the Deer Creek orbs and host, which appear to be relatively depleted in oxides such as TiO_2 and Al_2O_3 (Eisenberg, 2014). The Shaver granites follow the similar trends on Hark-

er diagrams, and most have greater than 70 wt% SiO₂. The A/CNK index (molar ratio Al₂O₃/[CaO+Na₂O+K₂O]), indicates the SIS is mostly metaluminous (A/CNK < 1), with some samples, including some of the Shaver granites, reaching mildly peraluminous values (A/CNK = 1.0 – 1.1). Magnesium number (MgO/[MgO+FeO_(tot)]100) decreases linearly with increasing SiO₂ in the SIS, except in the Shaver granites, which are depleted in MgO relative to FeO_(tot). Magnesium number in the Deer Creek locality follows a positive trend with SiO₂, counter to the rest of the SIS (Eisenberg, 2014).

Trace element data generally follow linear trends relative to SiO₂ with some scatter (Tables 3, 4; Fig. 7). Elements including Sr, Sc, V, and the middle REEs decrease with increas-

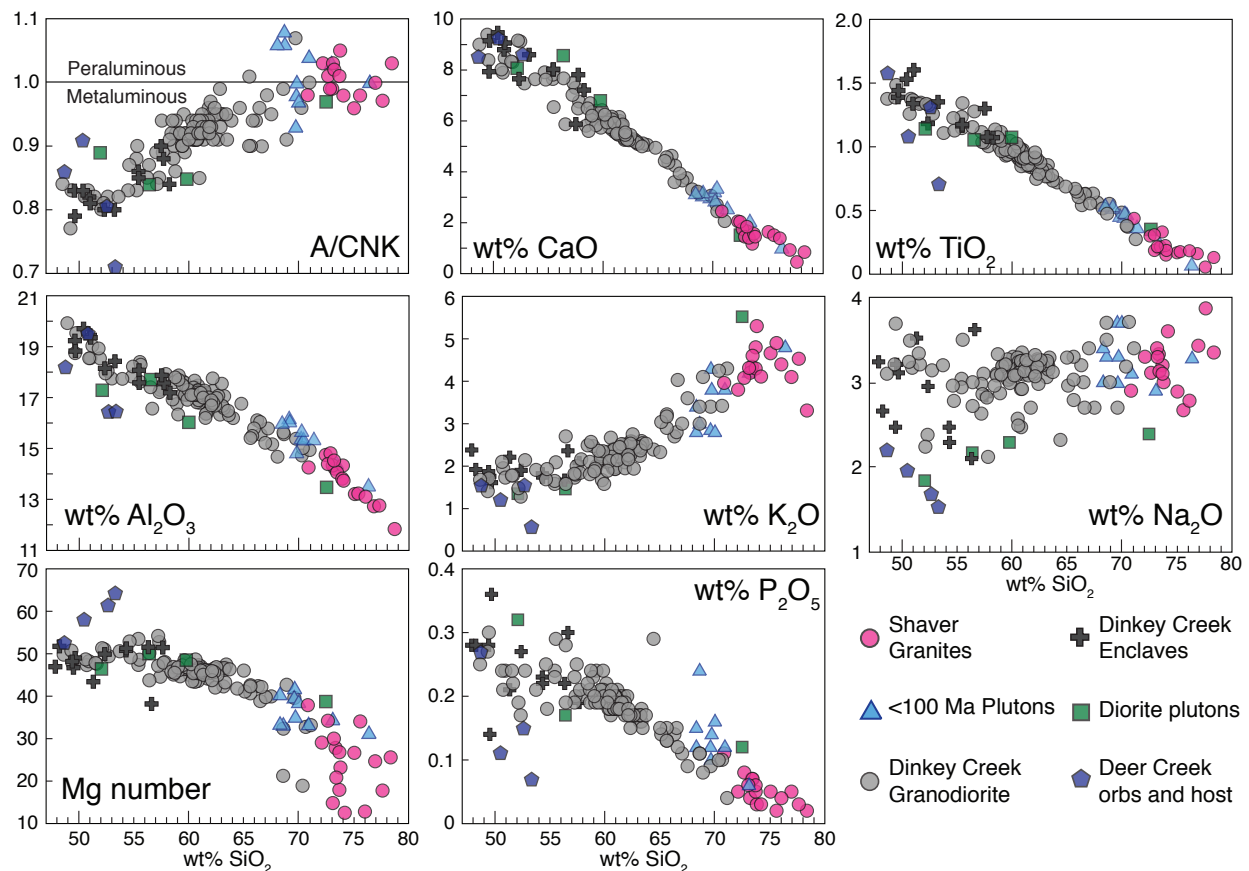


Figure 6. Representative Harker diagrams of whole rock analyses from the Shaver Intrusive Suite. Samples mostly follow linear trends, although there is general scatter in Na₂O, and there is significant scatter in P₂O₅ data in mafic compositions. Data are from this study, Barbarin et al. (1989), Lackey et al. (2008), and Hahm et al. (2014). The ratio A/CNK is calculated as the molar ratio Al₂O₃/[(CaO+Na₂O+K₂O)]; Mg number calculated as MgO/(MgO+Fe₂O₃·0.89981). Plutons designated as “<100 Ma” include Shorthair Creek (this study) and the leucogranite of Bald Mountain (Hahm et al., 2014).

ing SiO_2 , whereas Ba, Rb, and Pb generally increase with increasing SiO_2 . Niobium and Cs also appear to increase with increasing SiO_2 , but with more scatter. Zirconium, Th, and U increase until approximately 65–70 wt% SiO_2 , then decrease. The light and heavy REEs and Y show no apparent pattern with SiO_2 . The SIS exhibits a wide range of REE concentrations, with the Shaver granites being most variable (Fig. 8). Limited data for the DCG suggest it has similar REE characteristics to the adjacent Mount Givens Granodiorite (Fig. 8). Europium anomalies ($\text{Eu}/\text{Eu}^* = \text{Eu}_N/(\text{Sm}_N \cdot \text{Gd}_N)^{1/2}$, where subscript “N” indicates the element normalized to chondrite values of McDonough and Sun, 1995), mostly fall in the range 1.4–0.6 and become more negative as SiO_2 content increases. The granite north of Snow Corral Meadow is an exception, with positive Eu/Eu^* of 2.5 and 75.6 wt% SiO_2 . Negative Ce anomalies as low as $\text{Ce}/\text{Ce}^* = 0.4$, where $\text{Ce}/\text{Ce}^* = \text{Ce}_N/(\text{La}_N \cdot \text{Pr}_N)^{1/2}$, are present in several of the Shaver granites. Cerium anomalies in the Shaver granites do not correlate with SiO_2 , but do correlate with $^{87}\text{Sr}/^{86}\text{Sr}$, ϵNd_1 , and $\delta^{18}\text{O}(\text{Zrc})$ (Fig. 14).

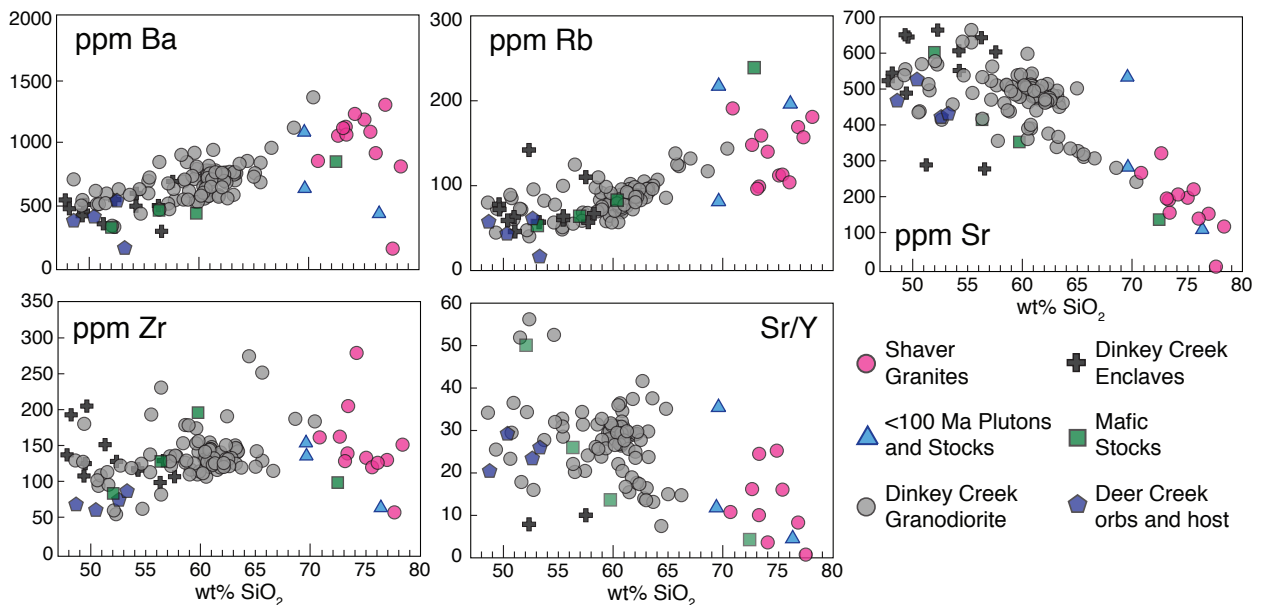


Figure 7. Harker diagrams for selected trace elements measured in this study. Trends are broadly linear across silica compositions. Note that the Shaver granites mostly have similar Zr concentrations to the DCG. Additional data sources as in Fig. 6.

Oxygen isotope ratios in zircon

Values of $\delta^{18}\text{O}(\text{Zrc})$ vary by 1.40‰ (Table 5), with the lowest $\delta^{18}\text{O}(\text{Zrc})$ (6.60‰) in a sample from the Dinkey Lakes pluton (8-RF-19) and the highest (8.00‰) from the Mushroom Rock pluton (8-RF-3). All of the $\delta^{18}\text{O}(\text{Zrc})$ values are above those for mantle-derived zircon ($5.3 \pm 0.3\text{‰}$; Valley et al., 1998). Additional $\delta^{18}\text{O}(\text{Zrc})$ data from the SIS are reported by Lackey et al. (2006, 2008) and Eisenberg (2014). Including those data show the Dinkey Lakes pluton has the greatest $\delta^{18}\text{O}(\text{Zrc})$ range in the suite. Its two samples, collected 3 km apart, yield $\delta^{18}\text{O}(\text{Zrc}) = 6.60\text{‰}$ and 8.03‰ . In contrast to the Dinkey Lakes pluton, many of the Shaver granites that are spatially associated have similar $\delta^{18}\text{O}(\text{Zrc})$ values, such as the nested Mushroom Rock (8.00‰) and Ordinance Creek (7.91‰) plutons, and two samples from the Sheepthief Creek pluton (7.15–7.18‰). In addition, diorite $\delta^{18}\text{O}(\text{Zrc})$ values are similar to the peraluminous plutons and the DCG (Fig. 9; Lackey et al., 2008).

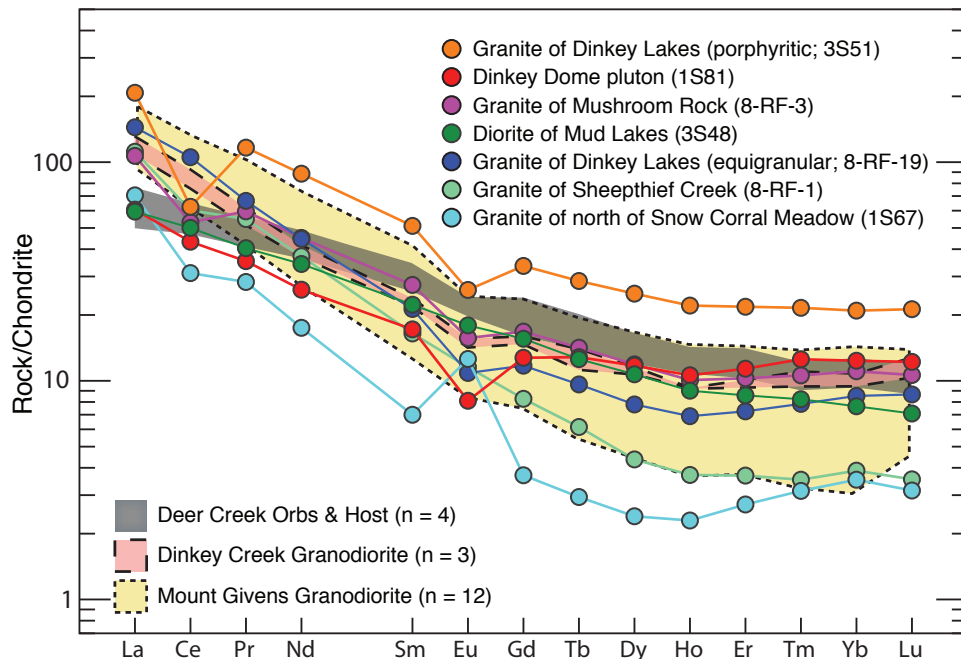


Figure 8. Rare earth element abundances for samples analyzed in this study normalized to chondritic composition of McDonough and Sun (1995). Individual analyses for each of the Shaver granites and the diorite of Mud Lakes are shown; granite REE concentrations span the entire range of concentrations measured in the Mount Givens Granodiorite. Note that field for the Dinkey Creek Granodiorite represents analyses of only three samples. Field for Mount Givens pluton includes sample MG12-03 of this study and plus data for samples from Frazer et al. (2014).

Radiogenic Isotopes

Most SIS samples have $^{87}\text{Sr}/^{86}\text{Sr}_i$ above 0.707; these are the highest $^{87}\text{Sr}/^{86}\text{Sr}_i$ values recognized in this part of the Sierra Nevada batholith (Figs. 9, 10; Table 6). The highest $^{87}\text{Sr}/^{86}\text{Sr}_i$ (0.7106) is from a leucocratic portion of the diorite of Rainbow Mine. The lowest value (0.6878) from the Coyote Creek pluton is reported for completeness, but will not be considered further because it is below that of basaltic achondrites (Papanastassiou and Wasserburg, 1968). Sample 8-RF-19 of the Dinkey Lakes pluton has the lowest likely robust $^{87}\text{Sr}/^{86}\text{Sr}_i$ (0.7054). In contrast, the other Dinkey Lakes pluton sample yields $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.7087$. There is no clear trend in $^{87}\text{Sr}/^{86}\text{Sr}_i$ with SiO_2 . The Sr isotopic data broadly agree with those of Dodge and Kistler (1990).

Neodymium isotopic data negatively correlate with $^{87}\text{Sr}/^{86}\text{Sr}_i$. Whereas the Sr isotopic data are the most radiogenic in the central part of the batholith, the Nd isotopic data are approximately equivalent to data from the Tuolumne and Whitney intrusive suites (Fig. 10). As with

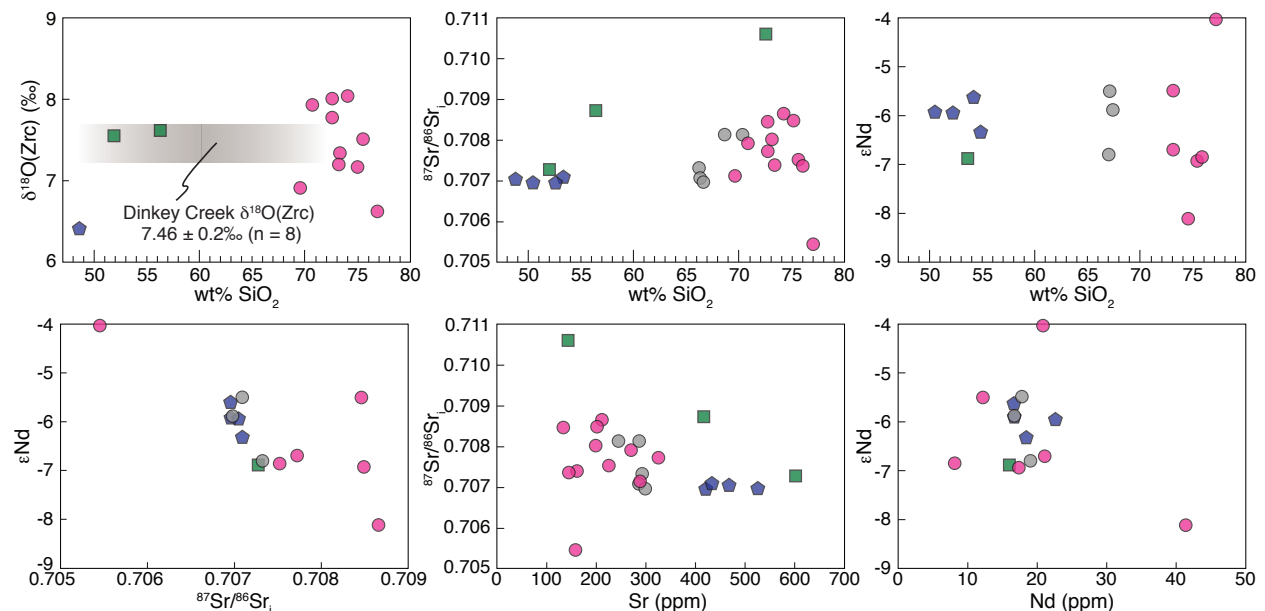


Figure 9. Results of isotopic analyses of the Shaver Intrusive Suite. Symbols as in Figures 6 and 7. Note the general scatter of the Shaver granites in all isotope systems relative to other SIS rocks. Radiogenic isotopic data are initial values corrected for decay using the corresponding age of the sample. Plot of $\delta^{18}\text{O}(\text{Zrc})$ also includes results from Lackey et al. (2006, 2008). Samples of Dinkey Creek Granodiorite analyzed for $\delta^{18}\text{O}(\text{Zrc})$ (Lackey et al., 2008) do not have corresponding SiO_2 data and thus the field for that pluton is shown as gradational in wt% SiO_2 .

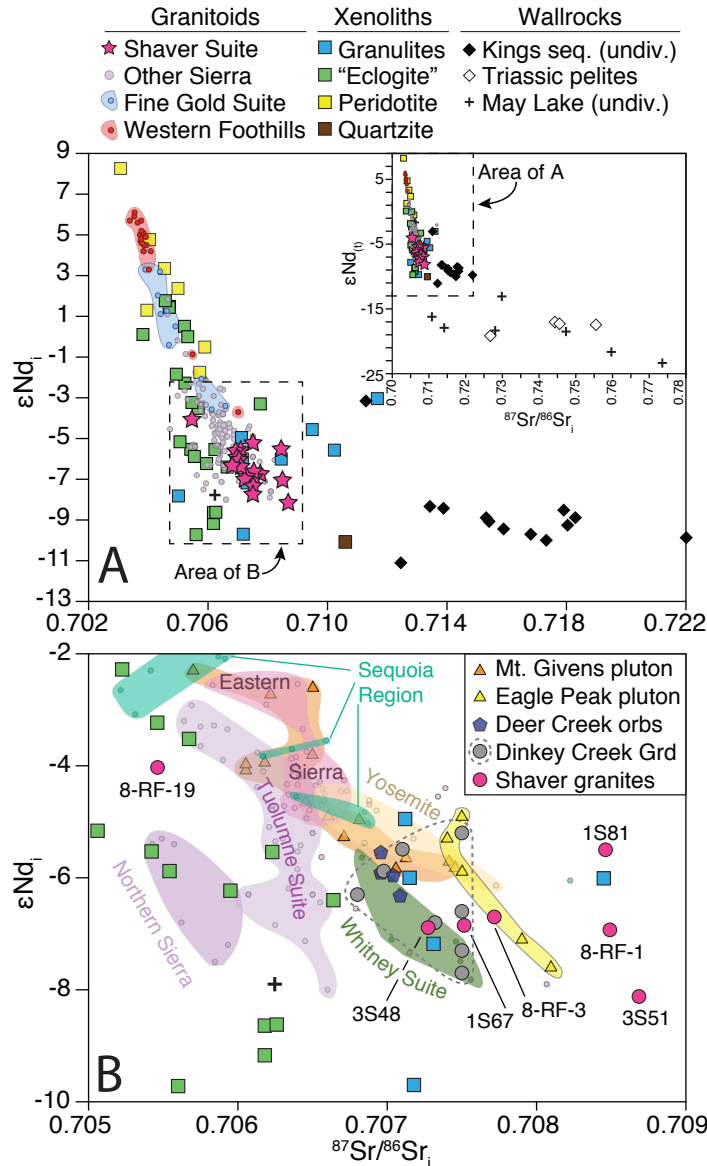


Figure 10. Results of Sr and Nd isotopic analyses for the Shaver Intrusive Suite. All data are shown as initial values corrected for decay using the corresponding age of the sample; xenolith and wallrock data are corrected to 100 Ma. (A) Data for the Shaver Intrusive Suite (pink stars) relative to published data for the Sierra Nevada batholith. Inset shows full extent of wallrock analyses. (B) Results for the Shaver Intrusive Suite relative to the Cretaceous batholith and some xenolith and wallrock analyses as in (A). Data for other suites and regions are from: Eagle Peak (Hill et al., 1988); Eastern Sierra (Coleman et al., 1992; Wenner and Coleman, 2004); Fine Gold (Truschel, 1996); Mount Givens (this study); Northern Sierra (Cecil et al., 2012); Sequoia (Holland et al., 2013); Shaver (Barbarin et al., 1989; Kistler and Fleck, 1994; this study); Tuolumne (Kistler et al., 1986; Gray et al., 2008); Western Foothills (Clemens-Knott, 1992); Whitney (Hirt and Glazner, 1995; Hirt, 2007); Yosemite (Ratjeski et al., 2001). Data for xenoliths are from: Domenick et al. (1983), Dodge et al. (1986), Mukhopadhyay and Manton (1994), and Ducea and Saleeby (1998), and Ducea (2001, 2002). Data for wallrocks are from: Kings sequence (Zeng et al., 2005); May Lake (Mills et al., 2009), and Triassic pelites (Zeng et al., 2005).

⁸⁷Sr/⁸⁶Sr_i, the Dinkey Lakes pluton has very heterogeneous Nd isotopic composition (εNd_i = -4.03 to -8.12). There is no correlation between εNd_i and SiO₂ (Fig. 9).

DICUSSION

Geochronology in the Shaver Intrusive Suite

Geochronologic data document emplacement of the SIS over ~8 Ma, from 106.5 to 98.3 Ma, although taking uncertainty into account permits emplacement duration from 7–11 Ma (Fig.

11). With an exposed area of $\sim 670 \text{ km}^2$ and relief of $\sim 2.2 \text{ km}$, the SIS has an approximate volume of 1500 km^3 . Therefore, the average magma flux over the lifespan of the suite ranges from $0.0001\text{--}0.0002 \text{ km}^3/\text{a}$. This magma flux is similar to those calculated for other plutons and intrusive suites in the Sierra Nevada and elsewhere (Mills and Coleman, 2013; Frazer et al., 2014). Although there is debate about how plutonic volumes are calculated, particularly in regards to their apparent thicknesses (e.g., Lipman and Bachmann, 2015), field evidence suggests the SIS is tabular (Cruden et al., 1999).

The new geochronology indicates the Shaver granites were emplaced throughout the history of the SIS (Fig. 11). These data preclude past field interpretations that called for the Shaver granites to be significantly younger than the DCG (Bateman, 1992). However, the uncertainties in the LA-ICP-MS data permit all but two of the Shaver granites to be contemporaneous with

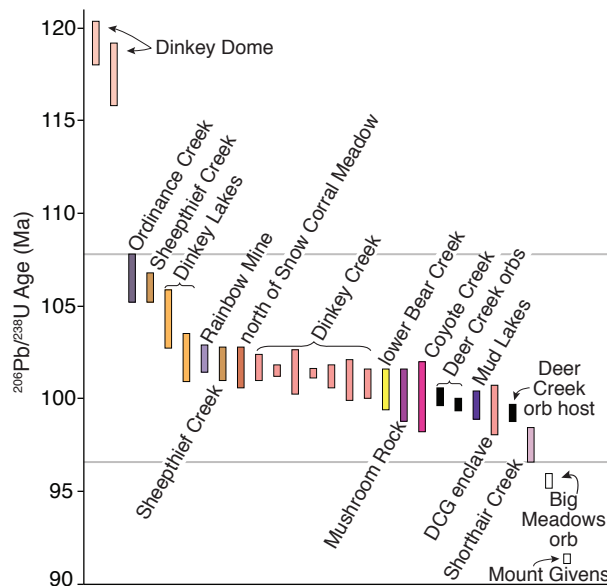


Figure 11. Compilation of zircon U-Pb geochronology in the Shaver Intrusive Suite ordered by age (laser ablation data: weighted means and 95% confidence interval; TIMS data: total age range with 2s analytical uncertainties). Interpreted age range of the Shaver Intrusive Suite indicated by horizontal gray lines. The uncertainties of the data permit only the Ordinance Creek and sample 8-RF-5 of Sheepthief Creek to be considered distinctly older than the DCG. Laser ablation ages for a mafic enclave and its host Dinkey Creek Granodiorite, as well as the Deer Creek orbicular locality, are also included (Lackey et al., 2012; Eisenberg, 2014). Ages for a Big Meadows orb (Eisenberg, 2014) and Mount Givens Granodiorite (sample MG12-03, this study) are included for comparison but are not part of the SIS.

or slightly younger than the DCG (Fig. 11). Samples 8-RF-4 from the Ordinance Creek pluton (106.5 ± 1.3 Ma) and 8-RF-5 from the Sheepthief Creek pluton (106.4 ± 1.2 Ma) are older, including uncertainty, than any dated sample of the DCG. The DCG adjacent to the Ordinance Creek pluton is undated, whereas the Sheepthief Creek sample is at minimum 2.5 Ma older than sample 95-17 of the DCG (101.2 ± 1.6 Ma), 2 km away (Fig. 1). Data for these two samples suggest, but do not require, that field relationships indicating the DCG has been cut by those plutons are inconclusive or incorrect.

The Dinkey Dome pluton (Kdd, Fig. 1) is a clear case where geochronologic data are at odds with previous field mapping, as it is some 15 Ma older than the DCG. The pluton's intimate association with the Dinkey Creek roof pendant is notable because the pendant is a part of the hypothesized Snow Lake block, which some have suggested was dextrally transported 400 km north in the Early Cretaceous (Lahren and Schweickert, 1989; Grasse et al., 2001). Recent LA-ICP-MS detrital zircon analyses of the Snow Lake pendant and others near the Tuolumne Intrusive Suite have cast doubt on the need for 400 km of displacement (Memeti et al., 2010b), but up to 200 km is still possible (Kistler, 1993; Saleeby and Busby, 1993; Lewis and Girty, 2001; Chapman et al., 2015). Memeti et al. (2010b) suggested movement on the hypothesized Mojave-Snow Lake fault must have occurred between 145 Ma and 102 Ma. Because of the chemical and isotopic similarities between the Dinkey Dome pluton and the 106-98 Ma Shaver granites, movement of the block may have been complete by the time of the pluton's emplacement ca. 119 Ma.

New CA-ID-TIMS U-Pb zircon data document emplacement of the DCG in the area of the Courtright shear zone ca. 102-101 Ma, with the Shorthair Creek pluton emplaced ca. 98 Ma. Both plutons are deformed, although deformation is most intense in the easternmost DCG where

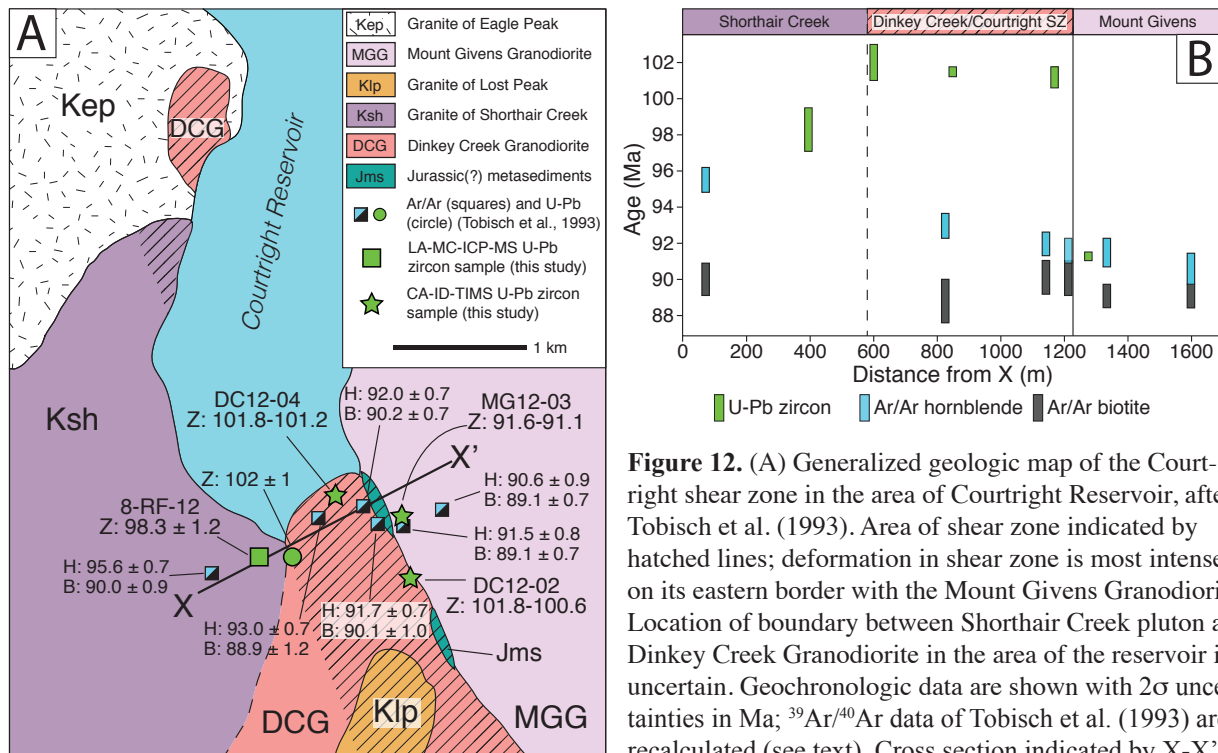
it is in contact with the Mount Givens Granodiorite, which was emplaced ca. 91.3 Ma and does not appear deformed. Recalculated $^{40}\text{Ar}/^{39}\text{Ar}$ data from Tobisch et al. (1993) show that hornblende $^{40}\text{Ar}/^{39}\text{Ar}$ ages in and adjacent to Mount Givens Granodiorite overlap within uncertainty of its new zircon U-Pb age (Fig. 12). Biotite $^{40}\text{Ar}/^{39}\text{Ar}$ ages for all samples are synchronous and within ~ 2 Ma of the 91.3 Ma zircon U-Pb age of the Mount Givens Granodiorite. Renne et al. (1993) argued that the $160^\circ\text{C}/\text{Ma}$ cooling rate indicated by the data may have been the result of rapid uplift and/or exhumation ca. 90 Ma, when the Mount Givens pluton was emplaced and the shear zone was inferred to be active. However, the data permit an alternative interpretation. Cooling rates in non-deformed plutonic rocks are now routinely calculated to be $>100^\circ\text{C}/\text{Ma}$ (Davis et al., 2012; Barboni et al., 2014; Samperton et al., 2015). Therefore, the $^{40}\text{Ar}/^{39}\text{Ar}$ data within and immediately adjacent to the Mount Givens Granodiorite may simply record the thermal effects of incremental emplacement and concomitant rapid cooling of small magma batches. This hypothesis raises the possibility that strain in the Courtright shear zone mostly accumulated prior to emplacement of the Mount Givens Granodiorite ca. 91.3 Ma.

Cerium anomalies and the source of the Shaver granites

Cerium anomalies occur when Ce^{3+} is oxidized to Ce^{4+} , thus decoupling the behavior of Ce relative to the trivalent REEs. Negative cerium anomalies are common in supracrustal rocks that have interacted with seawater, which has a large negative Ce anomaly (Elderfield and Greaves, 1982; Murray et al., 1990). Negative Ce anomalies as low as 0.85 are commonly measured in basalts erupted from arcs where the subducting sediment column also has a negative Ce anomaly (e.g., Hole et al., 1984; Plank and Langmuir, 1998). Such anomalies are rarely reported in plutonic or high-silica igneous rocks (Meen, 1990), and when present they are often attributed

to surficial weathering processes in part because they do not correlate with other geochemical or isotopic signatures (Sawka et al., 1986; Rooney et al., 2010; Lee et al., 2013). However, the correspondence of the most negative Ce anomalies with the highest $\delta^{18}\text{O}(\text{Zrc})$ values in the Shaver granites suggests a supracrustal source (Fig. 13). Furthermore, the high $^{87}\text{Sr}/^{86}\text{Sr}_i$ and low ϵNd_i values suggest the supracrustal source was likely not altered oceanic crust, as has been proposed for the Fine Gold Intrusive Suite (Lackey et al., 2012). Thus, a sedimentary source is plausible.

Metasedimentary wallrocks are common in the central Sierra Nevada, and could possibly have served as magma sources if similar rocks were present deep in the crust, where partial melting is more efficient (Annen and Sparks, 2002). For example, leucosomes and pelitic host rocks in a migmatite complex of the Kings sequence in the southern Sierra Nevada have Ce/Ce^*



line. Z – zircon; H – hornblende; B – biotite. (B) Samples are projected onto cross section X-X' with age on the ordinate. Hornblende Ar/Ar dates for DCG samples collected closest to the Mount Givens pluton are within uncertainty of the Mount Givens pluton's zircon U-Pb age, suggesting the Ar system was reset in those samples. In contrast, the hornblende Ar/Ar date for the Shorthair Creek pluton appears unaffected by the Mount Givens pluton's thermal aureole. All biotite Ar/Ar data overlap ca. 90 Ma, suggesting that deformation in the shear zone was complete by that time. Jurassic(?) metasedimentary wallrocks are omitted from cross section for clarity.

of ~ 0.7 (Zeng et al., 2005). However, wallrocks in the SIS are dominated by quartzite, with little pelite (Bateman, 1992). There is also disagreement over the affinity of SIS wallrocks (e.g., Kistler and Bateman, 1966; Lahren and Schweickert, 1989; Kistler, 1993; Saleeby and Busby, 1993) making a genetic link between the Kings sequence and the Shaver granites tenuous.

The presence of Cenozoic volcanic pipes that have erupted through and near the SIS (Fig. 1) permits evaluation of connections to the deep crust. Common garnet pyroxenite xenoliths in the basalts may be residues complementary to the voluminous Cretaceous batholith, and have Ce/Ce^* as low as 0.84 (Ducea, 2002). Rare metaquartzite xenoliths, which are interpreted as residues after melting of impure quartzite, have Ce/Ce^* as low as 0.64 (Chin et al., 2013).

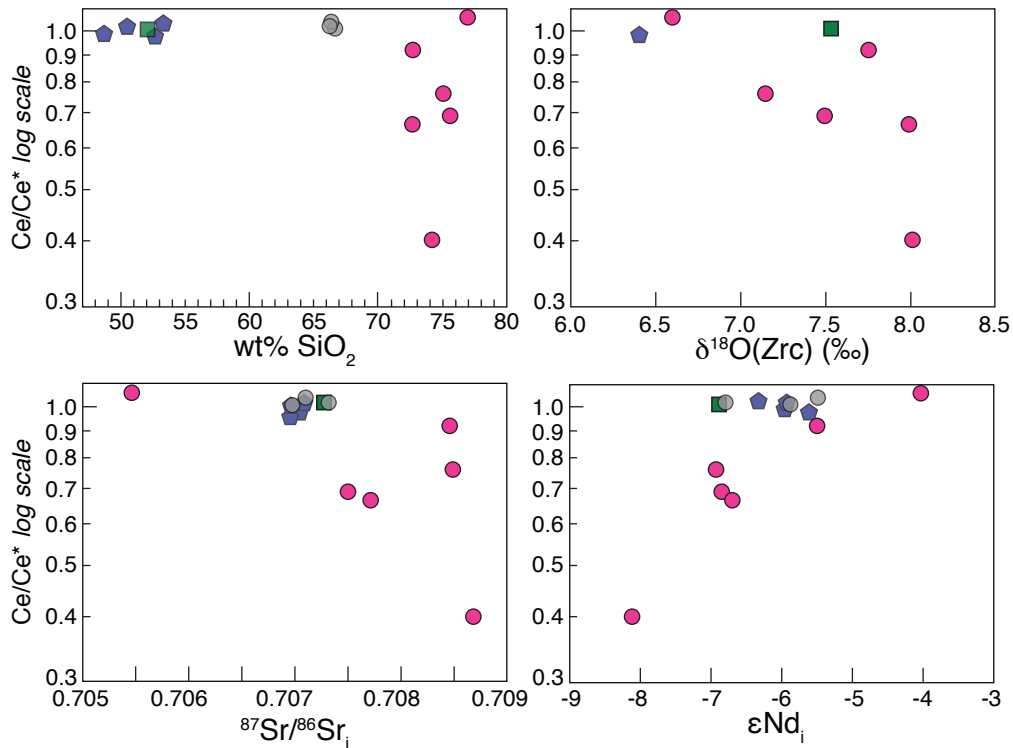


Figure 13. Cerium anomalies (Ce/Ce^*) relative to compositional and isotopic parameters measured in this study. Cerium anomalies are restricted to the Shaver granites, but not all granites have an anomaly. Sample 8-RF-19 is the most silicic sample yet has the most positive anomaly ($Ce/Ce^* = 1.08$). In contrast to silica, Ce anomalies correlate well with isotopic parameters. The correlation with $\delta^{18}O(Zrc)$ is notable because it indicates the whole rock anomaly accords with crystal-scale isotopic compositions. Likewise, the correlation with ϵNd_1 is notable because Ce and Nd are both light REEs and should therefore share a similar derivation. Symbols as in Fig. 9 and 10. See text for Ce anomaly calculation. Note that the ordinate is shown on a logarithmic scale.

Both xenolith types have mid-Cretaceous ages that overlap the SIS, and document the presence of negative Ce anomalies in the deep crust and/or upper mantle during assembly of the SIS. The metaquartzites in particular bear superficial similarity to exposed SIS wallrocks (Kistler and Bateman, 1966), and may also be isotopically appropriate matches for the Shaver granites, with one xenolith recording $^{87}\text{Sr}/^{86}\text{Sr}_{100\text{ Ma}} = 0.710637$ and $\epsilon\text{Nd}_{100\text{ Ma}} = -10.3$ (Ducea and Saleeby, 1998). However, Sierran quartzites are infertile (Mills et al., 2009), and modeled melts complementary to the metaquartzite xenoliths are much more aluminous than the Shaver granites (Chin et al., 2013).

Lower crustal Ce-anomalous rocks could have been part of a fertile pre-existing wall-rock terrane (Zeng et al., 2005), underthrust from the backarc (Ducea, 2001; Chin et al., 2013), or transported down the subduction zone with the Farallon slab (von Huene and Scholl, 1991). Lower Jurassic to mid-Cretaceous shales and cherts of the Franciscan assemblage on the Marin headlands have prominent negative Ce anomalies (Murray et al., 1990). If a similar package of mostly oceanic fertile sediments were thrust into the mantle wedge, it could have become a partial melting source. The variability of Ce anomaly magnitude in the Shaver granites, along with correlated isotopic compositions, may be attributed to entrainment of peritectic and accessory phases from the source (Villaros et al., 2009) and/or disequilibrium partial melting (Davies and Tommasini, 2000; Zeng et al., 2005). Variable Eu anomalies could perhaps be controlled by differing contributions of peritectic plagioclase (Clemens et al., 2011). A pelitic source for the Shaver granites is not ideal because the granites do not have cordierite or distinctly high A/CNK (e.g., Miller, 1985), but it is difficult to envision a scenario where non-sedimentary rocks form a significant part of the source because the Ce anomalies would be diminished by involvement of non-sedimentary rocks. Therefore, partial melting of a metasedimentary protolith and peritectic

assemblage entrainment (Clemens and Stevens, 2016) appear to be the simplest methods for generating the Shaver granites. This implies that the Shaver granites are mostly recycled supracrustal material.

Origins of diorites and the Dinkey Creek Granodiorite

In contrast to the Shaver granites, which have distinct mineralogy, Ce anomalies, and limited emplacement-level interaction with the rest of the SIS, the DCG and more mafic parts of the suite are superficially more similar to the large intrusive suites of the Sierra Crest. However, models for the origins of the mafic and intermediate parts of the SIS must account for their high $^{87}\text{Sr}/^{86}\text{Sr}_i$, low ϵNd_i , and high $^{18}\text{O}(\text{Zrc})$ values relative to those parts of the batholith.

Some hypotheses call for a large component of primitive mantle to generate chemical and isotopic diversity in intrusive suites (DePaolo, 1981; Nelson et al., 2013). However, they cannot account for the heterogeneities in the SIS. Data for the SIS do not follow curvilinear mixing trends on isotope-silica and isotope-element plots, as would be expected from simple mixing of primitive mantle and Proterozoic crust. Lee and Bachmann (2014) hypothesized that relatively “flat” isotopic trends with varying silica content, which are suggested by the $\delta^{18}\text{O}(\text{Zrc})$ and ϵNd_i data for intermediate and mafic SIS rocks (Fig. 9), could be explained by mixing primitive mantle and lower crust followed by fractional crystallization. However, elements such as P_2O_5 and Zr have mostly linear trends on Harker diagrams and thus do not require a fractional crystallization origin (Figs. 6, 7). In addition, it would be difficult to achieve the evolved isotopic compositions in mafic end members of the SIS by assimilation and fractional crystallization processes. Following Nelson et al. (2013), and using high-alumina olivine tholeiite (Bacon et al., 1997) and average continental crust (Rudnick and Gao, 2003) for end members (Table 7), bulk assimilation

of 40–45% continental crust is necessary to achieve the isotopic compositions of most of the SIS (i.e., $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.707$). Although such mixing would reproduce the suite's isotopic compositions, the new mafic end member would have ~53 wt% SiO_2 before subsequent fractional crystallization is taken into account, whereas the most mafic rocks present in the suite have 48 wt% SiO_2 (Fig. 9).

In contrast, others have suggested enriched mantle as an important component in the Sierra Nevada batholith, particularly in the well-studied Sierra Crest suites (Coleman et al., 1992; Wenner and Coleman, 2004). Enriched mantle is an appealing source for the mafic and intermediate parts of the SIS because it can account for its evolved radiogenic isotopic compositions. However, the SIS has elevated $\delta^{18}\text{O}(\text{Zrc})$ relative to the eastern Sierra (Lackey et al., 2008). This suggests that if enriched mantle played a role in the genesis of the SIS, it likely also interacted with supracrustal rocks in order to acquire the elevated $\delta^{18}\text{O}$ signature (Fig. 14). Both Kings sequence rocks and granulite xenoliths erupted through the Cenozoic Chinese Peak pipe are well-characterized and constitute plausible end members with high $\delta^{18}\text{O}$ values and appropriate $^{87}\text{Sr}/^{86}\text{Sr}$ (Fig. 1). Simple mixing arrays suggest incorporation of 0-40% crustal component into enriched mantle to account for the O isotope composition of mafic and intermediate parts of the SIS (Fig. 14). However, if Kings sequence rocks were involved in the production of the mafic and intermediate parts of the SIS, they must not have had Ce anomalies (Fig. 13; c.f. Zeng et al., 2005).

The Shaver Intrusive Suite and transitions in the Sierra Nevada batholith

The underlying causes for the structure, chemical and isotopic compositions of the SIS are enigmatic. Tobisch et al. (1995) showed that the SIS was emplaced during a period of dy-

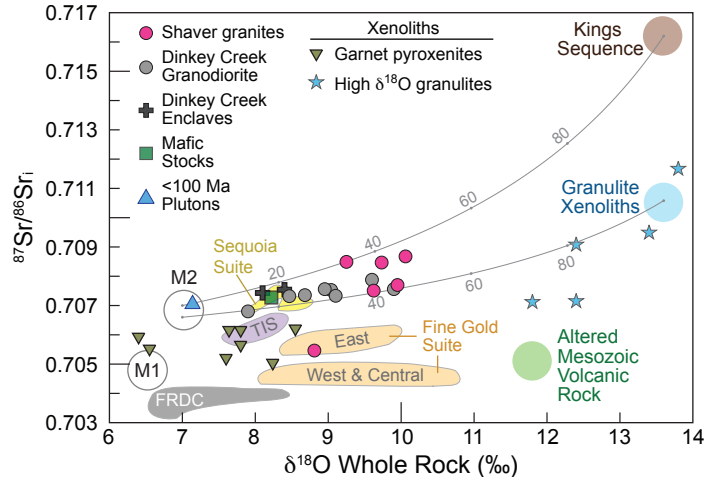


Figure 14. Strontium and oxygen isotope sources for the Shaver Intrusive Suite. M1 and M2 are postulated mantle reservoirs. M1 was suggested as a reservoir for the Dinkey Creek Granodiorite and Tuolumne Intrusive Suite by Lackey et al. (2012). However, the data presented here suggest the Shaver Intrusive Suite requires a more evolved source, designated M2. Simple mixing curves from M2 ($\delta^{18}\text{O} = 7\text{‰}$, $\text{Sr}_i = 0.7066\text{--}0.707$, 550 ppm Sr) to high- $\delta^{18}\text{O}$ end-members of Kings Sequence ($\delta^{18}\text{O} = 13.6\text{‰}$, $\text{Sr}_i = 0.7162$, 190 ppm Sr) and granulite xenoliths ($\delta^{18}\text{O} = 13.4\text{‰}$, $\text{Sr}_i = 0.71058$, 190 ppm Sr) bound the Shaver Intrusive Suite array. Although the Sr and O isotopic data for the entire suite are satisfied by this mixing model, Ce anomaly data suggest alternative mechanisms for the generation of the Shaver granites. The mafic and intermediate members of the SIS require 0–40% crustal component similar to Kings sequence rocks and/or granulite xenoliths to acquire their high $\delta^{18}\text{O}$ values. Kings sequence source is a composite of $\delta^{18}\text{O}$ data from Lackey et al. (2006) and Sr concentration and isotopic data from Zeng et al. (2005a). Granulite xenolith data are a composite from Dodge et al. (1986), corrected to 100 Ma. Data for the Sequoia Intrusive Suite are from this study, with additional $\delta^{18}\text{O}$ data from Lackey et al. (2008) and Sr data from Chen and Tilton (1991). Additional Sierran data are from Masi et al. (1981), Clemens-Knott (1992), Kistler and Fleck (1994), and Truschel (1996). FRDC – Foothills ring dike complexes; TIS – Tuolumne Intrusive Suite. Altered Mesozoic volcanic rock field from Lackey et al. (2012), after Staudigel et al. (1995).

namic tectonic activity in the Sierra Nevada batholith, as recorded by ductile shear zones preserved in and near the suite. In a review of global geodynamics that included data from Tobisch et al. (1995), Matthews et al. (2012) noted a variety of tectonic markers that may have recorded a worldwide tectonic reorganization at 105–100 Ma, perhaps triggered by the cessation of subduction to the east of Australia and New Zealand. However, it is unclear how or why such an event would influence the compositions and emplacement dynamics of magmas in the SIS. Whereas other suites of similar age also feature high-silica granites (Holland et al., 2013; Putnam et al., 2015; Frazer et al., in prep.), none preserve voluminous granodiorites like the DCG.

The distinct composition of the SIS, particularly the Shaver granites, may have arisen from a fortuitous combination of prolonged magmatism and a preexisting metasedimentary

framework. All Cretaceous U-Pb zircon ages in the Sierra Nevada from 36.5°N to 38.15°N using data acquired on zircon pre-treated by air- or chemical-abrasion when dated by TIMS, or zircon dated by LA-ICP-MS were compiled (Fig. 15). Limiting the dataset to these modern techniques allows for better recognition of zircons affected by inheritance or Pb-loss. The compilation reveals that rather than a smoothly migrating magmatic arc from west to east as past work has indicated (Chen and Moore, 1982; Nadin and Saleeby, 2008), magmatism in the central Sierra Nevada progressed step-wise, with the temporal step coincident with the location of the Courtright shear zone and other proposed intrabatholithic faults (Lahren and Schweickert, 1989; Kistler, 1990, 1993; Saleeby and Busby, 1993). Magmatism began at 124-120 Ma west of the intrabatholithic divide, and began ca. 98-95 Ma east of the divide. The 103-100 Ma Kearsarge intrusive suite is an exception whose eastward location may have been controlled by the presence of pre-existing fractures (Frazer et al., in prep.) Whereas the magmatic front did not migrate smoothly as previously suggested, the cessation of magmatism did progress steadily eastward throughout the Cretaceous. A line fit by eye across the batholith suggests the trailing magmatic cessation migrated at a rate of 2.7 mm/a (Fig. 15), agreeing with the rate determined by Chen and Moore (1982).

The data suggest that locations in the farthest eastern reaches of each magmatic “step” (i.e., the SIS in western half of the batholith, and the Sierra Crest suites in the eastern half) may have been the loci of prolonged magmatic activity. For the SIS, this may have led to up to 15–20 Ma of thermal priming leading to the subsequent melting of metasedimentary rocks interpreted to have been metamorphosed under the SIS at 103 ± 10 Ma (Chin et al., 2013). This hypothesis predicts that similar thermal priming could have occurred on the eastern margin of the eastern “step” in the suites of the Sierra Crest magmatic event (Coleman and Glazner, 1997), though perhaps to

a lesser degree due to the shorter period of magmatism recorded there. It is also notable that the location of the eastern margin of each step coincides with the presence of well-documented shear zones (Tikoff and Teyssier, 1992; Tobisch et al., 1993, 1995; Tikoff and de Saint Blanquat, 1997; Nadin et al., 2016).

Despite the stepwise function of magmatism at current exposure levels, it is possible the magma generation zone in the asthenosphere progressed gradually inboard of the subduction zone due to slab flattening or subduction erosion (Coney and Reynolds, 1977; van Huene and Scholl, 1991). However, instead of ascending magmas migrating along with the generation zone, they continued to make use of the area of the intrabatholithic divide as a conduit to the mid- and upper crust in the west. By the end of magmatism in the SIS (~98 Ma), the conduit became untenable because the magma generation front was too far east; after 98 Ma, magmas made use of other pathways to begin building the Sierra Crest intrusive suites. This model is analogous to

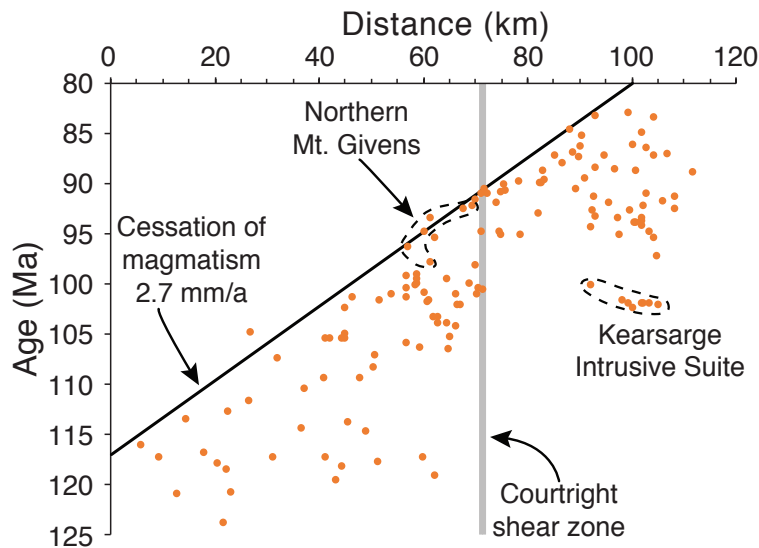


Figure 15. Data compilation of zircon U-Pb ages from the Cretaceous central Sierra Nevada batholith (36.5°N to 38.15°N) relative to their distance from a batholith-parallel line in the Central Valley oriented N35°W. Only samples dated by LA-ICP-MS, SIMS, SHRIMP, or ID-TIMS (using small fractions of pre-treated zircons) are included in this analysis. The results indicate that the 2.7 mm/a trend of Chen and Moore (1982) is present, but it tracks the eastward cessation of magmatism. The magmatic “front” appears to move across the batholith in a stepwise manner ca. 98 Ma, approximately demarcated by the location of the Courtright shear zone. Exceptions to this pattern include the northern part of the Mount Givens Granodiorite, which is younger than 98 Ma where the northern part of the pluton extends to the west. In addition, the location of the Kearsarge intrusive suite is enigmatic, and may have been controlled by pre-existing fractures (Frazer et al., in prep.).

hot spot volcanic chains, where the spacing between volcanoes may be controlled by lithospheric thickness and fractures that capture magma ascending from the source until the overriding plate has moved too far away (Vogt, 1974).

CONCLUSIONS

New U-Pb zircon geochronology shows that the Shaver Intrusive Suite was emplaced over an ~8 Ma period, from 106–98 Ma. Within the uncertainty of the laser ablation data, emplacement of granites, granodiorites, and diorites was essentially contemporaneous, although the only significant evidence for interactions between magmas of different compositions at emplacement level is the presence of mafic magmatic enclaves in the Dinkey Creek Granodiorite that dominates the suite. Prominent negative Ce anomalies in several of the peraluminous granites, coupled with the presence of Ce-anomalous metasedimentary xenoliths in Cenozoic volcanic rocks that erupted through the suite, suggest the granites are the products of partial melting of a metasedimentary source. It is hypothesized that the granites' variable REE concentrations and Ce anomaly magnitudes were governed by peritectic assemblage entrainment. Thus, the granites represent nearly 100% recycling of previously existing crustal material. In contrast, the bulk of the suite besides the granites was likely mostly derived from enriched mantle lithosphere with varying amounts of a supracrustal component such as Kings sequence-type rocks, or granulite. The contrasting mantle and crustal contributions to the SIS may be due to prolonged magmatism in the area of the SIS, which until ~98 Ma was the easternmost portion of the Cretaceous central Sierra Nevada batholith. If magmatism occurred in the area of the SIS since ~122 Ma, then the lower crust may have become thermally and chemically primed to partially melt metasedimentary rocks of varying fertility (impure quartzite to pelites) to generate the Shaver granites.

TABLE 1. SAMPLE SUMMARY

Sample	IGSN*	Name/Description	Rock type [†]	Analyses performed [§]	UTM-N [#]	UTM-E	Age ± 2σ**
8-RF-1	P4700000J	Sheepthief Creek	granite	Zrc, δ ¹⁸ O-Sr _i -εNd, WR	4122494	301547	101.9 ± 0.9
8-RF-3	P4700000K	Mushroom Rock	granite	Zrc, δ ¹⁸ O-Sr _i -εNd, WR	4122957	298625	100.2 ± 1.4
8-RF-4	P4700000L	Ordinance Creek	granite	Zrc, δ ¹⁸ O-Sr _i , WR	4123156	299198	106.5 ± 1.3
8-RF-5	P4700000M	Sheepthief Creek	granite	Zrc, δ ¹⁸ O, WR	4115181	301322	106.0 ± 0.8
8-RF-6	P47000018	Sheepthief Creek	granite	Sr _i , WR	4115056	301347	ND
8-RF-7	P47000019	Sheepthief Creek	granite	Sr _i , WR	4115089	301349	ND
8-RF-8	P4700001A	Dinkey Creek	granodiorite	Sr _i , WR	4114995	301316	ND
8-RF-9	P4700001B	Dinkey Creek	granodiorite	Sr _i , WR	4114966	301296	ND
8-RF-11	P4700000N	Coyote Creek ^{††}	granite	Zrc, Sr _i , WR	4120042	312716	100.1 ± 1.9
8-RF-12	P4700000O	Shorthair Creek	granite	Zrc, δ ¹⁸ O-Sr _i , WR	4105142	324281	98.3 ± 1.2
8-RF-14	P4700000P	Lower Bear Creek	granite	Zrc, δ ¹⁸ O-Sr _i , WR	4102479	309971	100.5 ± 1.1
8-RF-15	P4700000Q	Dinkey Dome	granite	Zrc, Sr _i , WR	4106172	314037	119.2 ± 1.2
8-RF-16	P4700000R	Rainbow Mine ^{††}	leucodiorite	Sr _i , WR	4113986	313947	ND
8-RF-17	P4700000S	Rainbow Mine ^{††}	diorite	Zrc, δ ¹⁸ O-Sr _i , WR	4113953	314114	102.18 ± 0.74
8-RF-18	P4700000T	Rainbow Mine ^{††}	diorite	WR	4114384	314160	ND
8-RF-19	P4700000U	Dinkey Lakes ^{††}	granite	Zrc, δ ¹⁸ O-Sr _i -εNd, WR	4114626	314158	102.2 ± 1.3
1S67 ^{§§}	P4700000V	North of Snow Corral Meadow	granite	Zrc, δ ¹⁸ O-Sr _i -εNd, WR	4102088	315437	101.7 ± 1.1
1S81 ^{§§}	P4700000W	Dinkey Dome	granite	Zrc, δ ¹⁸ O-Sr _i -εNd, WR	4106978	311033	117.5 ± 1.7
3S48 ^{§§}	P4700000X	Mud Lakes ^{††}	diorite	Zrc, δ ¹⁸ O-Sr _i -εNd, WR	4109522	315054	99.63 ± 0.76
3S51 ^{§§}	P4700000Y	Dinkey Lakes ^{††}	granite	Zrc, δ ¹⁸ O-Sr _i -εNd, WR	4111836	315339	104.3 ± 1.6
95-16 ^{##}	P4700000Z	Dinkey Creek	granodiorite	Zrc, δ ¹⁸ O	4107269	300768	101.7 ± 0.7
95-17 ^{##}	P47000010	Dinkey Creek	granodiorite	Zrc, δ ¹⁸ O	4114386	298468	101.2 ± 1.6
96-7 ^{##}	P47000011	Dinkey Creek	granodiorite	Zrc, δ ¹⁸ O	4100640	315668	101.0 ± 1.1
13JLE-4a	P47000012	Big Meadows host	gabbro	Sr _i -εNd, WR	4067046	340330	ND
13JLE-4b	P47000013	Big Meadows orb core	gabbro	Zrc, Sr _i -εNd, WR	4067046	340330	95.57 ± 0.40
13JLE-15	P47000014	Deer Creek orb core	gabbro	Sr _i -εNd, WR	4097406	315964	ND
13JLE-16	P47000015	Deer Creek orb core	gabbro	Zrc, Sr _i -εNd, WR	4097406	315964	100.09 ± 0.47
13JLE-17	P47000016	Dinkey Creek (Deer Creek orb host)	gabbro	Zrc, δ ¹⁸ O-Sr _i -εNd, WR	4097406	315964	99.25 ± 0.46
13JLE-18	P47000017	Deer Creek orb core	gabbro	Zrc, Sr _i -εNd, WR	4097406	315964	99.67 ± 0.35
DC12-02	P4700000D	Dinkey Creek	granodiorite	Zrc, Sr _i -εNd, WR	4104697	325337	101.8-100.6
MG12-03	P4700000E	Mount Givens	granodiorite	Zrc, Sr _i -εNd, WR	4105368	325121	91.6-91.1
DC12-04	P4700000F	Dinkey Creek	granodiorite	Zrc, Sr _i -εNd, WR	4105322	324680	101.8-101.2
DC12-05	P4700000G	Dinkey Creek	granodiorite	Zrc, Sr _i -εNd, WR	4097808	324206	101.7-101.2

*IGSN: International Geo Sample Number. Access sample metadata at www.geosamples.org.

[†]*Sensu lato*

[§]Analysis abbreviations: Zrc – zircon U-Pb geochronology; δ¹⁸O – δ¹⁸O in zircon; Sr_i – whole rock ⁸⁷Sr/⁸⁶Sr; εNd – whole rock ¹⁴³Nd/¹⁴⁴Nd; WR – whole rock elemental composition

[#]Sample locations given in NAD 83, UTM zone 11 north

**Preferred age and uncertainty given in Ma. ND: not dated. Note age ranges given for DC12-02 through DC12-05. See text for geochronologic methods used

^{††}Informal pluton name used in this paper

^{§§}δ¹⁸O and WR data from Lackey et al. (2006)

^{##}δ¹⁸O data from Lackey et al. (2008)

TABLE 2. MAJOR ELEMENT DATA

Sample	SiO ₂ wt%	TiO ₂ wt%	Al ₂ O ₃ wt%	Fe ₂ O ₃ wt%	MnO wt%	MgO wt%	CaO wt%	Na ₂ O wt%	K ₂ O wt%	P ₂ O ₅ wt%	Total
8-RF-1*	75.06	0.18	13.14	1.69	0.03	0.31	1.59	2.89	4.66	0.05	99.60
8-RF-3*	72.66	0.32	14.32	2.40	0.04	0.63	2.09	3.11	4.08	0.08	99.73
8-RF-4*	70.83	0.47	14.17	3.67	0.06	1.13	2.45	2.90	3.96	0.11	99.75
8-RF-5*	73.36	0.22	14.15	2.16	0.03	0.42	1.71	3.33	4.34	0.07	99.79
8-RF-6*	73.12	0.21	14.51	2.11	0.03	0.33	1.84	3.29	4.29	0.06	99.79
8-RF-7*	76.05	0.18	13.12	1.65	0.02	0.22	1.38	2.78	4.39	0.04	99.82
8-RF-8*	70.35	0.38	15.02	3.27	0.06	0.68	2.45	3.18	4.25	0.10	99.73
8-RF-9*	68.59	0.47	15.59	3.89	0.07	0.95	3.20	3.49	3.39	0.11	99.75
8-RF-11*	77.63	0.06	12.75	0.55	0.02	0.06	0.45	3.87	4.53	0.03	99.95
8-RF-12*	69.63	0.45	14.75	3.38	0.07	1.22	2.94	2.99	4.29	0.10	99.82
8-RF-14*	73.42	0.23	13.98	2.41	0.04	0.32	1.44	3.13	4.60	0.07	99.64
8-RF-16*	72.47	0.35	13.43	3.04	0.06	0.97	1.49	2.39	5.52	0.12	99.84
8-RF-17*	56.37	1.05	17.67	8.18	0.13	4.13	8.52	2.17	1.46	0.17	99.85
8-RF-18*	59.76	1.07	16.01	7.74	0.12	3.68	6.75	2.29	2.19	0.21	99.82
8-RF-19*	76.96	0.17	12.69	1.21	0.05	0.20	0.92	3.43	4.10	0.05	99.78
13JLE-4a*	56.93	0.84	19.19	6.55	0.10	2.37	6.58	3.06	1.60	3.06	100.28
13JLE-4b*	52.78	1.27	17.62	9.33	0.15	4.46	7.73	2.81	1.33	2.81	100.29
13JLE-15*	53.30	0.70	16.47	8.27	0.14	7.42	10.96	1.55	0.54	0.07	99.42
13JLE-16*	50.46	1.08	19.53	7.74	0.11	5.35	9.19	1.95	1.18	0.11	96.70
13JLE-17*	48.66	1.58	18.20	9.95	0.17	5.55	8.52	2.19	1.51	0.27	96.60
13JLE-18*	52.64	1.31	16.45	8.04	0.11	6.39	8.60	1.67	1.52	0.15	96.88
DC12-02 [§]	66.34	0.506	14.72	4.82	0.077	1.46	3.83	3.01	3.37	0.10	98.23
MG12-03 [§]	64.45	0.608	15.90	5.60	0.082	1.74	4.46	4.08	2.18	0.20	99.30
DC12-04 [§]	66.64	0.543	15.04	4.87	0.079	1.56	3.99	3.12	3.38	0.12	99.34
DC12-05 [§]	66.22	0.573	15.01	5.49	0.083	1.64	4.13	3.10	3.37	0.11	99.72

*Data acquired by X-ray fluorescence at Pomona College

[§]Data acquired by inductively coupled plasma optical emission spectroscopy at Actlabs (Ontario, Canada)

TABLE 3. TRACE ELEMENT DATA

Sample	Sc ppm	V ppm	Ba ppm	Sr ppm	Y ppm	Zr ppm	Cr ppm	Co ppm	Ni ppm	Cu ppm	Zn ppm	Ga ppm	As ppm	Rb ppm	Nb ppm	Mo ppm	Cs ppm	Hf ppm	Ta ppm	Pb ppm	Th ppm	U ppm
8-RF-1*	BDL [†]	9	1176	202	8	134	4	NA [§]	BDL	5	46	15	NA	112	10	NA	5	NA	NA	21	11	3
8-RF-3*	5	29	1049	324	20	163	5	NA	BDL	49	84	18	NA	148	14	NA	6	NA	NA	23	12	4
8-RF-4*	8	48	851	270	25	162	6	NA	BDL	5	83	19	NA	191	17	NA	12	NA	NA	21	19	8
8-RF-5*	5	15	1119	196	8	140	4	NA	BDL	4	58	16	NA	100	11	NA	7	NA	NA	21	13	4
8-RF-6*	8	48	1119	285	31	189	9	NA	4	BDL	69	20	NA	117	13	NA	BDL	8	NA	18	15	3
8-RF-7*	5	35	1351	245	26	185	8	NA	2	BDL	64	18	NA	144	14	NA	9	6	NA	18	16	3
8-RF-8*	BDL	11	918	145	5	128	9	NA	3	BDL	31	14	NA	104	8	NA	BDL	4	NA	16	6	BDL
8-RF-9*	4	20	1106	199	7	131	7	NA	5	BDL	42	16	NA	96	9	NA	BDL	5	NA	18	12	BDL
8-RF-11*	BDL	5	163	13	17	59	BDL	NA	BDL	4	49	14	NA	157	10	NA	5	NA	NA	31	13	4
8-RF-12*	6	59	638	287	24	137	11	NA	BDL	6	69	17	NA	218	16	NA	12	NA	NA	33	32	18
8-RF-14*	4	12	1058	161	16	205	6	NA	BDL	6	78	18	NA	159	18	NA	6	NA	NA	23	20	5
8-RF-16*	11	44	846	143	33	100	7	NA	BDL	5	61	17	NA	239	18	NA	14	NA	NA	28	30	15
8-RF-17*	25	226	463	416	16	129	24	NA	BDL	13	89	19	NA	64	10	NA	8	NA	NA	13	7	5
8-RF-18*	24	164	440	355	26	196	30	NA	5	14	87	18	NA	83	14	NA	7	NA	NA	17	16	9
8-RF-19*	BDL	5	1292	158	19	131	6	NA	BDL	4	53	14	NA	169	15	NA	8	NA	NA	27	23	7
13JLE-4a*	11.2	172.4	477.8	588.3	11.6	60.3	12.4	58.3	-0.30	16.7	99.7	20.1	14.7	72.1	11.4	2.5	4.6	2.6	9.2	1.7	12.4	2.6
13JLE-4b*	19.2	228.9	401.9	534.3	17.3	93.6	29.5	155.1	4.30	31.5	128.8	19.9	24.0	46.7	11.5	3.9	2.5	2.6	7.9	0.2	9.3	4.1
13JLE-15*	30.6	210.8	161.2	432.6	17.4	86.7	198.2	171.8	22.7	9.4	95.6	15.8	-5.0	15.4	7.0	2.7	5.3	2.4	NA	-4.7	4.6	0.9
13JLE-16*	24.9	285.6	408.8	525.3	17.6	62.0	62.8	137.3	21.6	17.2	91.4	19.5	4.8	43.7	6.4	1.8	2.5	1.4	NA	-1.7	2.6	1.5
13JLE-17*	32.9	251.4	372.8	466.3	22.9	68.8	32.3	154.9	14.0	34.7	131.1	20.1	9.1	58.2	12.9	3.4	4.6	2.8	NA	-0.6	6.0	0.2
13JLE-18*	31.5	312.0	530.0	420.6	17.8	75.2	87.2	115.6	34.8	16.9	99.2	18.7	4.7	61.1	6.7	2.1	9	1.8	NA	-0.7	3.4	2.2
DC12-02 [#]	8	74	633	285	13	115	260	9	BDL	20	60	16	BDL	138	7	4	5	3.3	1.7	19	22.7	6.3
MG12-03 [#]	7	91	418	579	8	137	40	13	30	40	90	21	BDL	95	8	BDL	4.1	3.7	0.9	16	10.0	5.6
DC12-04 [#]	9	79	611	298	13	132	230	10	BDL	20	70	17	BDL	153	8	4	6.9	3.8	1.4	19	26.0	9.1
DC12-05 [#]	9	84	590	292	15	127	40	11	BDL	10	60	17	BDL	154	9	3	3.5	3.5	1.3	18	23.1	5.6

*Data measured by X-ray fluorescence at Pomona College

[§]NA = element not analyzed[†]BDL = below detection limit[#]Data for Sc, V, Ba, Sr, Y, and Zr measured by inductively coupled plasma optical emission spectrometry at Actlabs (Ontario, Canada). All other elements measured by inductively coupled plasma mass spectrometry at Actlabs

TABLE 4. RARE EARTH ELEMENT DATA

Sample	La ppm	Ce ppm	Pr ppm	Nd ppm	Sm ppm	Eu ppm	Gd ppm	Tb ppm	Dy ppm	Ho ppm	Er ppm	Tm ppm	Yb ppm	Lu ppm	Ce/Ce*	Eu/Eu*
8-RF-1 [§]	26.45	36.82	5.22	17.34	2.52	0.67	1.70	0.23	1.11	0.21	0.61	0.09	0.66	0.09	0.76	0.99
8-RF-3 [§]	25.32	32.46	5.66	21.03	4.20	0.91	3.45	0.53	3.02	0.57	1.70	0.27	1.88	0.27	0.66	0.73
8-RF-19 [§]	34.22	64.49	6.32	20.82	3.26	0.63	2.41	0.36	1.98	0.39	1.20	0.20	1.45	0.22	1.06	0.68
1S67 [§]	16.71	19.03	2.69	8.16	1.07	0.73	0.76	0.11	0.61	0.13	0.45	0.08	0.60	0.08	0.69	2.48
1S81 [§]	14.37	26.48	3.35	12.19	2.63	0.47	2.62	0.48	2.98	0.60	1.88	0.32	2.10	0.31	0.92	0.55
3S48 [§]	14.08	30.63	3.84	15.98	3.41	1.04	3.20	0.47	2.72	0.51	1.42	0.21	1.30	0.18	1.01	0.96
3S51 [§]	49.21	38.24	11.07	41.36	7.82	1.51	6.89	1.07	6.36	1.25	3.61	0.55	3.56	0.54	0.40	0.63
13JLE-4a [†]	30.80	54.26	5.47	18.45	2.79	1.06	2.03	0.29	1.51	0.31	0.89	0.13	0.94	0.13	1.01	1.36
13JLE-4b [†]	18.59	37.94	4.81	19.40	4.19	1.16	3.56	0.53	3.07	0.60	1.63	0.22	1.63	0.24	0.97	0.92
13JLE-15 [†]	14.17	33.75	4.45	18.42	3.85	1.26	3.51	0.52	3.06	0.60	1.70	0.24	1.71	0.24	1.03	1.04
13JLE-16 [†]	11.76	28.63	3.91	16.71	4.03	1.22	3.31	0.52	3.00	0.61	1.69	0.23	1.63	0.22	1.02	1.02
13JLE-17 [†]	18.00	39.62	5.23	22.63	5.30	1.43	4.96	0.75	4.17	0.82	2.35	0.31	2.15	0.30	0.99	0.85
13JLE-18 [†]	12.58	28.37	3.87	16.64	3.95	1.12	3.57	0.53	3.21	0.64	1.72	0.24	1.57	0.23	0.98	0.91
DC12-02 [#]	30.2	54.4	5.32	17.8	3.4	0.79	2.9	0.4	2.6	0.5	1.5	0.23	1.7	0.26	1.04	0.77
MG12-03 [#]	26.2	51.6	5.73	20.2	3.5	0.93	2.7	0.4	1.9	0.3	0.9	0.14	0.9	0.13	1.02	0.92
DC12-04 [#]	25.3	46.0	4.78	16.7	3.2	0.86	2.9	0.4	2.7	0.5	1.5	0.24	1.5	0.25	1.01	0.86
DC12-05 [#]	30.9	56.4	5.75	19.0	3.5	0.86	3.1	0.5	2.8	0.5	1.6	0.27	1.7	0.31	1.02	0.80

*Ce and Eu anomalies calculated as $Ce_N/(La_N \times Pr_N)^{1/2}$ and $Eu_N/(Sm_N \times Gd_N)^{1/2}$, respectively, where subscript N denotes concentration normalized to chondritic values of McDonough and Sun (1995)

[§]Data measured by inductively coupled plasma mass spectrometry (ICP-MS) at Vrije University

[†]Data measured by ICP-MS at University of California – Santa Barbara

[#]Data measured by ICP-MS at Actlabs (Ontario, Canada)

TABLE 5. OXYGEN ISOTOPE VALUES FOR ZIRCON

Sample	Pluton	Rock type*	$\delta^{18}\text{O}(\text{Zrc})$ (‰)	$\pm 1\sigma$
8-RF-1	Sheepthief Creek	granite	7.15	0.03
8-RF-3	Mushroom Rock	granite	8.00	0.03
8-RF-4	Ordinance Creek	granite	7.91	0.03
8-RF-5	Sheepthief Creek	granite	7.19	0.03
8-RF-12	Shorthair Creek	granite	6.92	0.03
8-RF-14	Lower Bear Creek	granite	7.33	0.03
8-RF-17	Rainbow Mine	diorite	7.61	0.03
8-RF-19	Dinkey Lakes	granite	6.60	0.03
13JLE-4a	Big Meadows (orb core)	gabbro	6.97	0.10
13JLE-17	Dinkey Creek (Deer Creek orb host)	tonalite	6.40	0.10
1S67 [†]	North of Snow Corral Meadow	granite	7.50	0.04
1S81 [†]	Dinkey Dome	granite	7.76	0.04
3S48 [†]	Mud Lakes	diorite	7.54	0.04
3S51 [†]	Dinkey Lakes	granite	8.02	0.04
95-16 [§]	Dinkey Creek	granodiorite	7.35	0.01
95-17 [§]	Dinkey Creek	granodiorite	7.32	0.09
96-7 [§]	Dinkey Creek	granodiorite	7.58	0.04

**Sensu lato*[†]Oxygen isotope data from Lackey et al. (2006)[§]Oxygen isotope data from Lackey et al. (2008)

TABLE 6. SR AND ND ISOTOPIC DATA

Sample	Rb ppm	Sr ppm	$^{87}\text{Sr}/^{86}\text{Sr}$	$\pm 2\sigma$ abs*	$^{87}\text{Sr}/^{86}\text{Sr}_{\text{norm}}$	$^{87}\text{Rb}/^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}_i$	Sm ppm	Nd ppm	$^{143}\text{Nd}/^{144}\text{Nd}$	$\pm 2\sigma$ abs*	$^{147}\text{Sm}/^{144}\text{Nd}$	$^{143}\text{Nd}/^{144}\text{Nd}_{\text{norm}}$	$^{143}\text{Nd}/^{144}\text{Nd}_i$	ϵNd_i
8-RF-1 [†]	88	202	0.710318	0.000008	0.710326	1.261	0.708500	2.52	17.34	0.512210	0.000006	0.088	0.512211	0.512153	-6.91
8-RF-3 [†]	100	324	0.708991	0.000007	0.708999	0.893	0.707727	4.20	21.03	0.512245	0.000007	0.121	0.512246	0.512167	-6.67
8-RF-4 [†]	191	270	0.711013	0.000011	0.710998	2.047	0.707899								
8-RF-6 [#]	96	199	0.710119	0.000010	0.710104	1.396	0.708001								
8-RF-7 [#]	104	145	0.710504	0.000010	0.710489	2.076	0.707362								
8-RF-8 [#]	144	245	0.710605	0.000009	0.710590	1.701	0.708124								
8-RF-9 [#]	117	285	0.709856	0.000009	0.709841	1.188	0.708119								
8-RF-11 [#]	157	13	0.738072	0.000010	0.738056	35.047	0.688205								
8-RF-12 [#]	218	287	0.710192	0.000010	0.710177	2.198	0.707106								
8-RF-14 [#]	159	161	0.711475	0.000013	0.711460	2.858	0.707378								
8-RF-16 [#]	239	143	0.717620	0.000010	0.717605	4.840	0.710576								
8-RF-17 [#]	64	416	0.709372	0.000009	0.709357	0.445	0.708711								
8-RF-19 [†]	187	158	0.710426	0.000014	0.710434	3.425	0.705459	3.26	20.82	0.512363	0.000006	0.095	0.512364	0.512301	-4.01
1S67 [§]	105	225	0.709472	0.000003	0.709480	1.350	0.707528	1.07	8.16	0.512209	0.000009	0.079	0.512210	0.512158	-6.82
1S81 [§]	205	134	0.715860	0.000010	0.715868	4.430	0.708471	2.63	12.19	0.512305	0.000004	0.130	0.512306	0.512206	-5.47
3S48 [§]	77	601	0.707798	0.000007	0.707806	0.371	0.707281	3.41	15.98	0.512241	0.000005	0.129	0.512242	0.512158	-6.86
3S51 [§]	155	211	0.711807	0.000011	0.711815	2.123	0.708663	7.82	41.36	0.512166	0.000007	0.114	0.512167	0.512089	-8.08
DC12-02 ^{††}	138	285	0.709103	0.000009	0.709088	1.401	0.707053	3.4	17.8	0.512303	0.000005	0.116	0.512320	0.512244	-5.16
MG12-03 ^{††}	95	579	0.707223	0.000010	0.707208	0.475	0.706592	3.5	20.2	0.512332	0.000006	0.105	0.512349	0.512286	-4.57
DC12-04 ^{††}	153	298	0.709119	0.000010	0.709104	1.490	0.706961	3.2	16.7	0.512283	0.000004	0.116	0.512283	0.512223	-5.55
DC12-05 ^{††}	154	292	0.709519	0.000010	0.709504	1.526	0.707304	3.5	19.0	0.512233	0.000004	0.111	0.512250	0.512176	-6.47
MG10-01 ^{††}	165	370	0.708802	0.000011	0.708802	1.290	0.707045	4.4	23.1	0.512289	0.000004	0.115	0.512306	0.512234	-5.49
MG10-02 ^{††}	192	217	0.709531	0.000011	0.709516	2.560	0.706165	1.9	13.2	0.512370	0.000005	0.087	0.512387	0.512335	-3.61
MG10-03 ^{††}	198	103	0.713377	0.000010	0.713362	5.565	0.706030	0.6	4.5	0.512357	0.000005	0.081	0.512374	0.512325	-3.78
MG10-04 ^{††}	241	284	0.709278	0.000009	0.709263	2.456	0.706039	2.1	12.5	0.512377	0.000004	0.102	0.512394	0.512536	-3.64
MG10-05 ^{††}	121	388	0.708687	0.000009	0.708672	0.902	0.707448	3.2	19.1	0.512281	0.000005	0.101	0.512298	0.512235	-5.47
MG10-07 ^{††}	137	400	0.707989	0.000008	0.707974	0.991	0.706694	2.5	14.8	0.512313	0.000005	0.102	0.512330	0.512269	-4.91
MG10-08 ^{††}	194	352	0.708873	0.000010	0.708858	1.595	0.706795	3.8	25.0	0.512321	0.000004	0.092	0.512338	0.512283	-4.64
MG11-01 ^{††}	130	470	0.707556	0.000012	0.707541	0.800	0.706452	6.1	33.2	0.512347	0.000004	0.111	0.512364	0.512294	-4.30
MG11-02 ^{††}	139	388	0.708527	0.000010	0.708512	1.037	0.707108	3.6	21.0	0.512291	0.000004	0.104	0.512308	0.512243	-5.31
MG11-03 ^{††}	129	495	0.707401	0.000009	0.707386	0.754	0.706383	3.4	18.7	0.512376	0.000004	0.110	0.512393	0.512535	-3.75
MG11-04 ^{††}	137	288	0.708125	0.000010	0.708110	1.376	0.706195	3.7	18.3	0.512450	0.000005	0.122	0.512467	0.512389	-2.41
MG12-01 ^{††}	139	445	0.708638	0.000009	0.708623	0.904	0.707398	4.6	25.0	0.512291	0.000006	0.111	0.512308	0.512239	-5.40

Note: Subscript "norm" indicates ratio normalized to facilitate interlaboratory comparison. See text for rationale. Initial isotopic values calculated using normalized ratios and U-Pb zircon ages given in Table 1. For ID-TIMS age ranges, the midpoint was used. Ages used for undated samples are: 8-RF-6 and 8-RF-7 (106 Ma); 8-RF-8 and 8-RF-9 (102 Ma); 8-RF-16 (102.18 Ma)

*Analytical uncertainty only. See text for reproducibility of standards at Vrije University (VU) and University of North Carolina at Chapel Hill (UNC)

[†]Rb concentration by isotope dilution (ID) at VU; Sr concentration by X-ray fluorescence (XRF) at Pomona College (PC); Sm, Nd concentrations by ICP-MS at VU; isotopic ratios by TIMS at VU

[#]Rb, Sr concentrations by XRF at PC; isotopic ratios by TIMS at UNC

[§]Rb concentration by ID at VU; Sr concentration from Lackey et al. (2006); Sm, Nd concentrations by ICP-MS at VU; isotopic ratios by TIMS at VU

^{††}Rb, Sm, Nd concentrations by ICP-MS at Actlabs (Ontario, Canada); Sr by ICP-optical emission spectroscopy at Actlabs; isotopic ratios by TIMS at UNC

CHAPTER 3: THE ALABAMA HILLS GRANITE AND IMPLICATIONS FOR DEXTRAL OFFSETS IN OWENS VALLEY, MID-CRETACEOUS SIERRAN MAGMATISM, AND THE GENESIS OF GRANODIORITES

INTRODUCTION

Understanding the origins of the Alabama Hills, in Owens Valley, CA, and their relationship to the adjacent Sierra Nevada has been a goal of geologists since Josiah Whitney described them as “forming a sort of detached outlier of the Sierra, and composed of similar granitic and metamorphic rocks” (Whitney, 1872, p. 137). Their composition (Lee, 1906; Knopf, 1918) and well-documented historical seismic activity (Whitney, 1872; Gilbert, 1884; Hobbs, 1910) typify the geology of eastern California. Coupled with their ease of access between the Sierra Nevada and Inyo Range, they are an intriguing target for study.

The Quaternary history of the Alabama Hills is well-documented due to interest in modern activity on the Lone Pine fault, including the 1872 earthquake (M_w 7.8-7.9; Hough and Hutton, 2008) that devastated Lone Pine and produced a dramatic and well-preserved fault scarp near the eastern base of the hills. This has led to a wealth of study on the Lone Pine and Owens Valley faults (Lubetkin and Clark, 1988; Beanland and Clark, 1994; Lee et al., 2001) and their relationships to larger-scale tectonic systems (e.g., Wesnousky and Jones, 1994; Unruh et al., 2003; Le et al., 2007; Stevens et al., 2013). The Alabama Hills have also proven fruitful for the application of cosmogenic nuclides to understand bedrock erosion rates, earthquake recurrence, and debris flow deposition (Bierman et al., 1995a, 1995b; Nichols et al., 2006).

Pre-Quaternary study of the Alabama Hills has mostly been limited to the Jurassic metavolcanic and metasedimentary rocks that comprise the eastern half of the hills (Dunne and Walker, 1993; Dunne et al., 1998; Sorensen et al., 1998), as well as the Independence dike swarm that intrudes those units (e.g., Moore and Hopson, 1961; Chen and Moore, 1979; Bartley et al., 2007). Relatively little work has been done on the granite pluton that comprises the western half of the hills. A zircon U-Pb age of 85 Ma (Chen and Moore, 1982) is similar to the youngest parts of the Sierra Crest magmatic event, including the adjacent Whitney Intrusive Suite (Coleman and Glazner, 1997; Davis, 2010), and has led some to speculate the Alabama Hills are a down-dropped block from the Whitney suite (Ali et al., 2009).

High-precision U-Pb zircon geochronology, whole rock elemental and radiogenic isotopic data are used to test the links between the Alabama Hills Granite and the main Sierra Nevada batholith. These data demonstrate the Alabama Hills Granite is mid-Cretaceous in age and belongs to a suite of plutons of similar ages and isotopic compositions that lies to the north of the Whitney Intrusive Suite. The data permit evaluation of the relationship of the Alabama Hills block to the Sierra Nevada, and the extent of high-silica granite in the mid-Cretaceous batholith. Finally, the significance of high-Zr, high-silica granite for the generation of both voluminous granitoids and high-silica rhyolites is explored.

GEOLOGIC BACKGROUND

The Alabama Hills are a low range located in Owens Valley, west of the town of Lone Pine, CA and east of the Sierra Nevada range front (Figs. 1, 2). The range was first mapped in reconnaissance by Knopf (1918) and includes the Jurassic volcanic complex of the Alabama Hills (Dunne et al., 1998; Stone et al., 2000) and the Cretaceous Alabama Hills Granite (Stone et

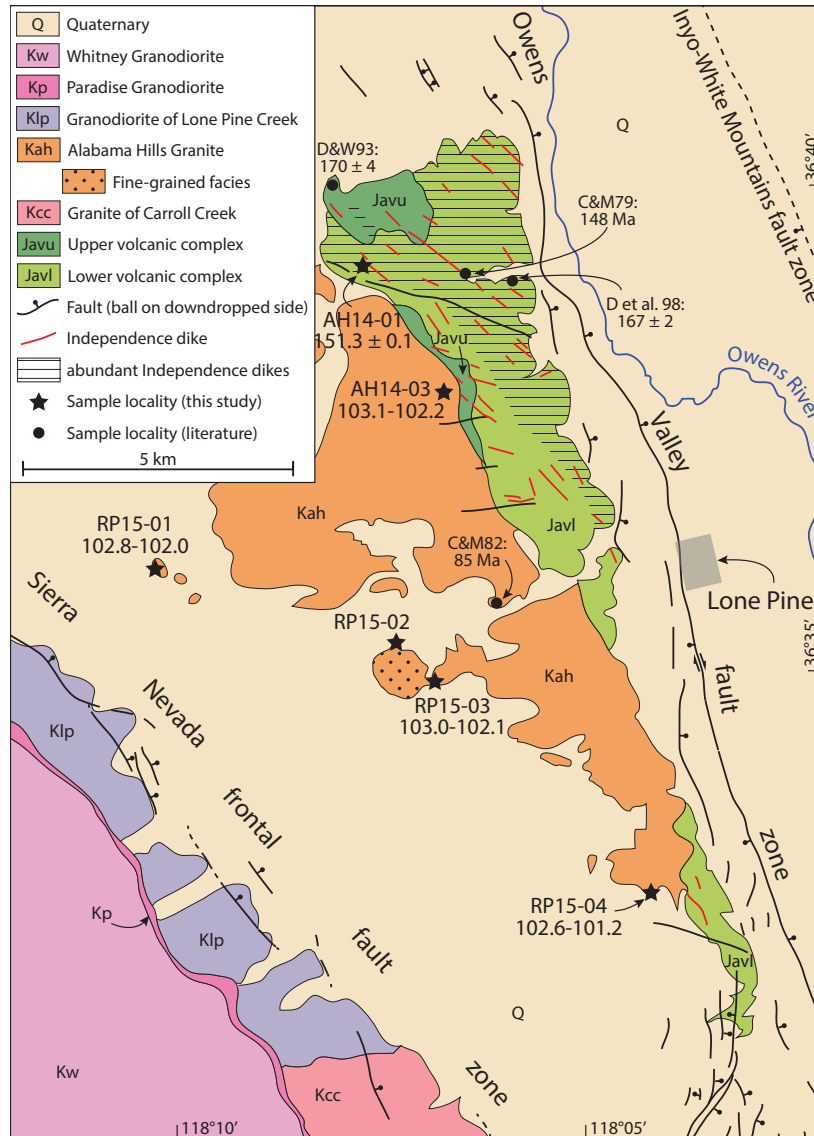


Figure 1. Generalized geologic map of the Alabama Hills, Owens Valley, CA, after Stone et al. (2000). Zircon U-Pb ages given in Ma as weighted means with 2σ uncertainty, or as age ranges (see text for discussion). Note that whereas all individual Independence dikes are mapped, two samples (AH14-01 and literature data of C&M79) were specifically collected from Independence dikes. Literature age prefixes: C&M79 – Chen and Moore (1979); C&M82 – Chen and Moore (1982); D&W93 – Dunne and Walker (1993); D et al. 98 – Dunne et al. (1998).

al., 2000). The volcanic complex of the Alabama Hills is also pervasively invaded by the Jurassic Independence dike swarm (Chen and Moore, 1979; Carl et al., 1998).

The western part of the Alabama Hills is dominated by the Alabama Hills Granite (Stone et al., 2000). The granite is exposed as distinctive inselbergs and ridges over an area of approximately 30 km². The pluton is mostly equigranular granite, although textures range from mildly porphyritic to fine-grained (Stone et al., 2000). Color index is below 10%, with accessory min-

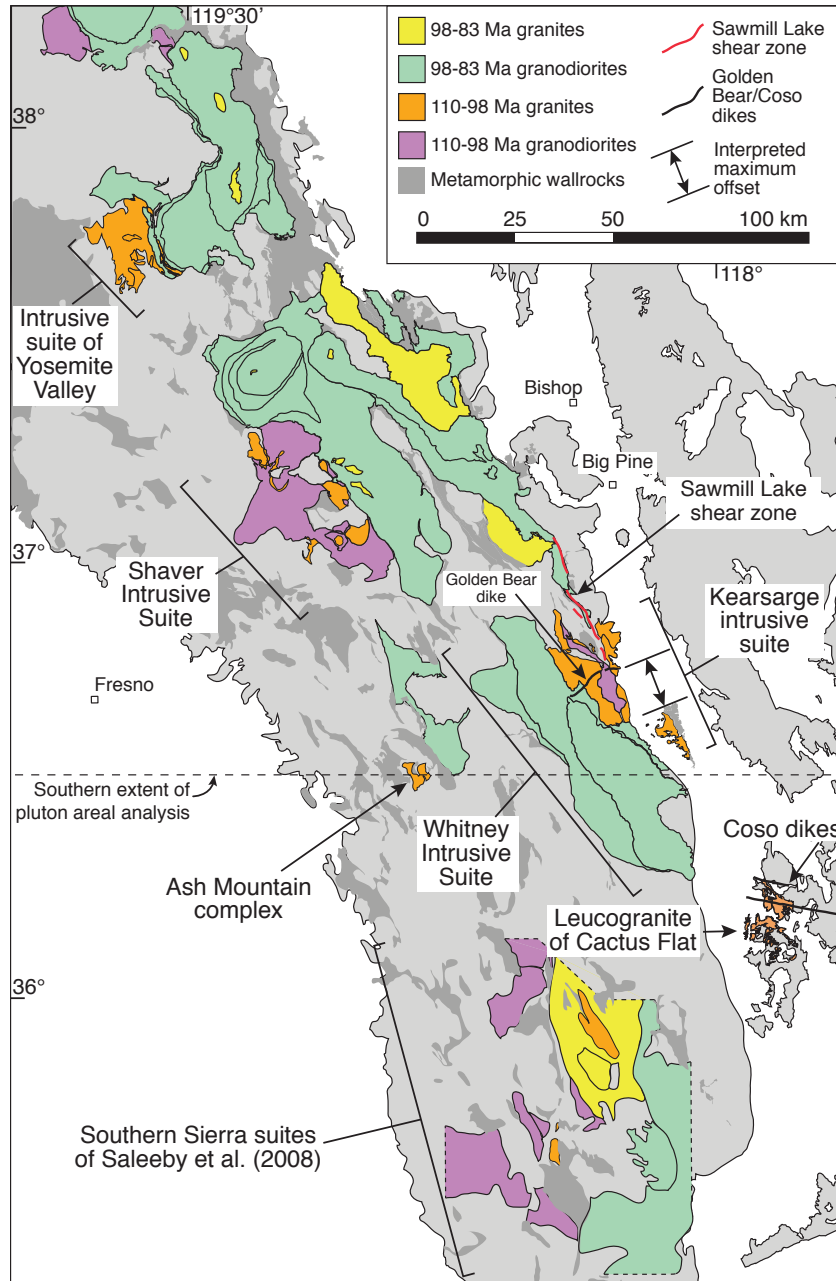


Figure 2. Generalized bedrock geologic map of the central Sierra Nevada batholith indicating localities mentioned in the text. The Alabama Hills Granite is grouped with the Kearsarge intrusive suite, which includes the Bullfrog, Independence, and Dragon plutons in the main batholith, and may also include the leucogranite of Cactus Flat in the Coso Range. Maximum interpreted offset of Alabama Hills relative to the main batholith is approximately 10 km based on the distance to the Golden Bear dike. Mid- and Late Cretaceous plutons shown in color are those considered “well-dated” in Figure 8. Horizontal dashed line indicates southern extent of plutons included in analysis of Figure 8. Only those parts of the intrusive suites that have been reliably dated are shown. For example, the intrusive suite of Yosemite Valley is limited to the map units that are dated by Ratajeski et al. (2001) and Putnam et al. (2015). The western part of the El Capitan Granite of Bateman (1992) is omitted due to a 114 Ma age (Putnam et al., 2015). Light gray part of batholith indicates plutonic rock undivided by age or composition. Bedrock geology after Moore (1963, 1978, 1981), Huber (1983), Moore and Sisson (1987), Bateman (1992), Moore and Nokleberg (1992), Sisson and Moore (1994), Stone et al. (2000), Wahrhaftig (2000), Lackey et al. (2008), and Saleeby et al. (2008). Evolution Basin Alaskite modified after Wenner and Coleman (2003) and Davis et al. (2012). Tuolumne Intrusive Suite modified after Memeti et al. (2010b). Sawmill Lake shear zone after Bartley et al. (2012).

erals including biotite, magnetite, ilmenite, and titanite (Richardson, 1975; Ague and Brimhall, 1988a; Stone et al., 2000). Ague and Brimhall (1988a) assigned the Alabama Hills Granite to their I-SC (strongly contaminated I-type) group of plutons. Lackey et al. (2008) measured $\delta^{18}\text{O}$ in two bulk zircon aliquots from a single sample, finding $\delta^{18}\text{O}$ in zircon = $6.41 \pm 0.08 \text{ ‰}$ (2σ). Chen and Moore (1982) dated a sample from the central part of the pluton using TIMS U-Pb geochronology on bulk zircon aliquots (~ 10 mg each), yielding a concordant age of 85 Ma.

METHODS

Six samples were collected from the Alabama Hills for geochronological and geochemical analyses. Sample AH14-01 was collected from a diabase dike in the northern part of the volcanic complex of the Alabama Hills. Samples AH14-03, RP15-01, RP15-03, and RP15-04 were taken from the main body of the Alabama Hills Granite with the aim of maximizing spatial distribution. Samples were equigranular, although RP15-03 and RP15-04 featured some alkali feldspar crystals up to 1 cm in length with rapakivi texture. Sample RP15-02 was collected from a prominent hill mapped as a fine-grained facies of the Alabama Hills Granite (Stone et al., 2000). Only fresh, unweathered samples were selected for whole rock geochemical and isotopic analyses.

Whole rock elemental and isotopic geochemistry

Whole rock aliquots of samples RP15-01 through RP15-04 were crushed to powder in a SPEX Shatterbox® alumina swing mill and analyzed at Actlabs (Ontario, Canada). Samples were dissolved by fusion in a lithium metaborate/tetraborate mixture; major elements and Ba, Sr, Y, Zr, Sc, Be and V were analyzed by inductively coupled plasma-optical emission spectroscopy (ICP-OES), with remaining trace elements and REE analyzed by ICP-MS. Uncertainties ($\pm 2\sigma$

relative) for major elements were less than 2% for all oxides except MgO (3%), MnO (5%), and P₂O₅ (16%). Trace elements analyzed at Actlabs that are reproducible ($\pm 2\sigma$ absolute) at better than 2 ppm include Ag, Be, Cs, Ge, Hf, Sb, Sn, Ta, Th, Tl, U, and W. Other elements are reproducible within 3 ppm (Co, Mo, Nb, Sc), 4 ppm (Ga, Y) and others have higher uncertainties, in parentheses in ppm: Pb (9); Cu (14); Rb and Zr (15); V (19); Ni (23); Sr (25); Ba (31); Zn (42); Cr (49). Rare earth elements reproducible ($\pm 2\sigma$) at 0.2 ppm or better are Eu, Ho, Lu, Tb, and Tm; REE reproducible at 0.2-0.5 ppm or better are Dy, Er, Gd, Pr, Sm, and Yb; others are reproducible at higher uncertainties, in ppm: Nd (1.2); La (1.5); Ce (2.1).

For isotope geochemistry analyses, whole rock powders were dissolved in HF and HNO₃ in Teflon® (Parr) bombs at 180°C for 48-72 hours. Following dissolution, samples were dried down then immediately fluxed in 6M HCl for 16 hours. Samples were subsequently aliquoted for Sr and Nd purification by ion exchange column chromatography. Strontium was purified using Sr-spec cation exchange resin after the methods of Lundblad (1994) and loaded on single Re filaments with TaF₅. Neodymium was purified through a three-stage column chemistry procedure after Harvey and Baxter (2009) and loaded on single Re filaments in a Ta₂O₅-H₃PO₄ slurry. Strontium isotopic analyses were accomplished on a VG Sector 54 TIMS and Nd isotopes were analyzed on an Isotopx Phoenix TIMS at the University of North Carolina at Chapel Hill. Strontium was analyzed as a metal in dynamic multicollector mode with ⁸⁸Sr = 3V; Nd was analyzed as an oxide in dynamic multicollector mode with ¹⁴²Nd/¹⁶O = 1V. Strontium and Nd isotope ratios were corrected for mass fractionation using an exponential law; Sr isotopic ratios were normalized to ⁸⁶Sr/⁸⁸Sr = 0.1194; Nd isotopic ratios were normalized to ¹⁴⁶Nd/¹⁴⁴Nd = 0.7219. Replicate analyses of the NBS 987 Sr standard yielded ⁸⁷Sr/⁸⁶Sr = 0.710265 ± 0.000015 (2σ; n = 26). Replicate analyses of the Nd standard JNdi yielded ¹⁴³Nd/¹⁴⁴Nd = 0.512098 ± 0.000012 (2; n = 18). All

isotopic data are corrected to initial values using elemental concentration values from this study and a crystallization age of 102 Ma.

Geochronology

Geochronology was performed by CA-ID-TIMS on either a VG Sector 54 or Isotopx Phoenix-X62 in the Department of Geological Sciences at the University of North Carolina at Chapel Hill following the sample preparation and analytical methods of Frazer et al. (2014). Zircons for each sample were thermally annealed for 48 hours at 900°C and then chemically abraded in HF and HNO₃ acids for 16 hours at 220°C (Mundil et al., 2004; Mattinson, 2005). Individual grains were then selected under a binocular microscope for dissolution and isotope dilution using a mixed ²⁰⁵Pb-²³³U-²³⁶U tracer after Parrish and Krogh (1987). Dissolution and chemical purification methods for U and Pb were modified after Krogh (1973) and Parrish (1987). Data processing and age calculations were completed using the applications Tripoli and ET_Redux, developed as part of the EARTHTIME initiative (Bowring et al., 2011; McLean et al., 2011), and include 2σ analytical uncertainties only. Corrections for initial Th/U disequilibrium (Mattinson, 1973; Schmitz and Bowring, 2001) were made in ET_Redux. Corrections for samples RP15-01 through RP15-04 were made assuming the measured whole rock Th/U value for each respective sample approximates the magmatic value; this adjustment increased the ages of individual analyses by up to 100 ka. Samples without corresponding whole rock concentration data used the following Th/U values and rationales: AH14-01: 3.5 (average of Independence dike samples AH95-8B, AH95-8C, and ID95-11 of Glazner et al., 2008a); AH14-03: 4.4 (average of samples RP15-01, RP15-03, and RP15-04 from this study).

RESULTS

Whole rock chemical analyses

Samples of Alabama Hills Granite have 73-77 wt% SiO₂ and small ranges of major element concentrations except sample RP15-02 from the fine-grained facies, which has slightly lower Fe₂O₃, MnO, MgO, CaO, and TiO₂, and higher K₂O (Fig. 3). Trace element concentrations are variable (Fig. 4). Samples RP15-01 and RP15-03 have higher Zr and Cr concentrations than other samples (Table 2). Sample RP15-02 is depleted in Ba, Sr, Y, and Zr relative to the equigranular facies, whereas it has higher Rb and U concentrations. All samples contain at least 13 ppm Nb and Y/Sr ratios above 0.1 (Fig. 4).

Samples from the equigranular facies have La_N/Lu_N ratios near 10, and Gd_N/Lu_N near 1, indicating a “flat” chondrite-normalized heavy rare earth element (REE) pattern (subscript “N” indicates concentration normalized to chondritic values of McDonough and Sun, 1995). The same samples also have Eu/Eu* (calculated as Eu_N/(Sm_N·Gd_N)^{1/2}) near 0.4, whereas sample RP15-02 has no discernable Eu anomaly (Fig. 5). Sample RP15-02 has low REE concentrations and a distinct REE pattern with low La_N/Lu_N and Gd_N/Lu_N.

Radiogenic isotopic analyses

There is a wide range of ⁸⁷Sr/⁸⁶Sr_i ratios among the samples analyzed, from 0.7030 to 0.7058. The equigranular samples have more restricted ⁸⁷Sr/⁸⁶Sr_i values (0.7052-0.7058), whereas RP15-02, with Rb/Sr = 11.75, yielded ⁸⁷Sr/⁸⁶Sr_i = 0.7030. Neodymium isotopic compositions are restricted; calculated εNd at 102 Ma for the four Alabama Hills Granite samples ranges -2.4 to -3.1 (Table 2; Fig. 6).

Zircon geochronology

As modern high-precision zircon U-Pb data for intrusive rocks have become difficult to

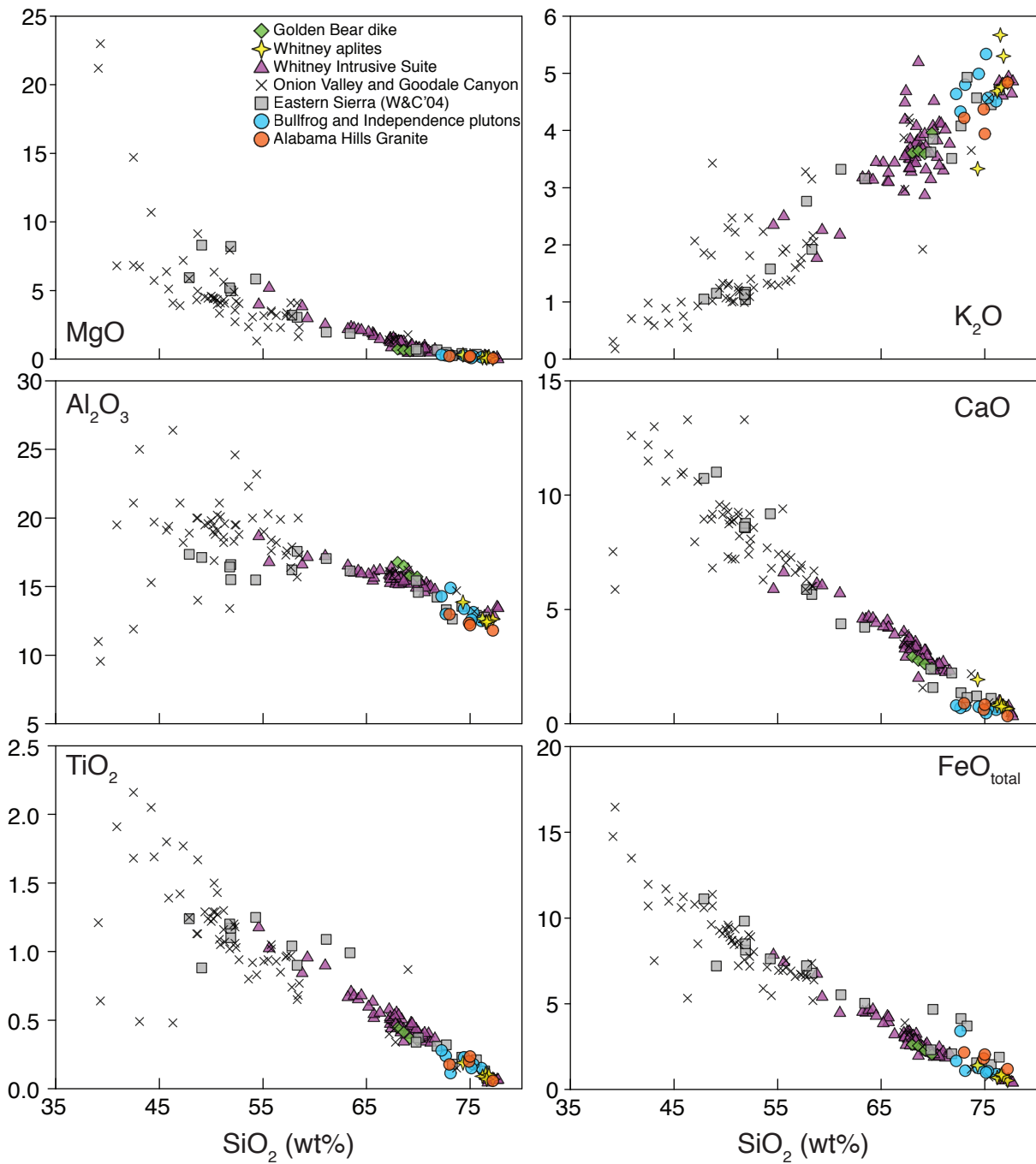


Figure 3. Major element data for the Alabama Hills Granite and other plutonic rocks of the eastern Sierra Nevada batholith. Ordinate units are wt% of the oxide indicated. The Alabama Hills Granite cannot be distinguished from other high-silica rocks, including aplites and leucogranites from the Whitney Intrusive Suite, and the Bullfrog and Independence plutons. Literature data from Bradford (1995), Sisson et al. (1996), Wenner and Coleman (2004), Kylander-Clark et al. (2005), Hirt (2007), Glazner et al. (2008b), and Coleman et al. (2012). W&C'04 – Wenner and Coleman (2004).

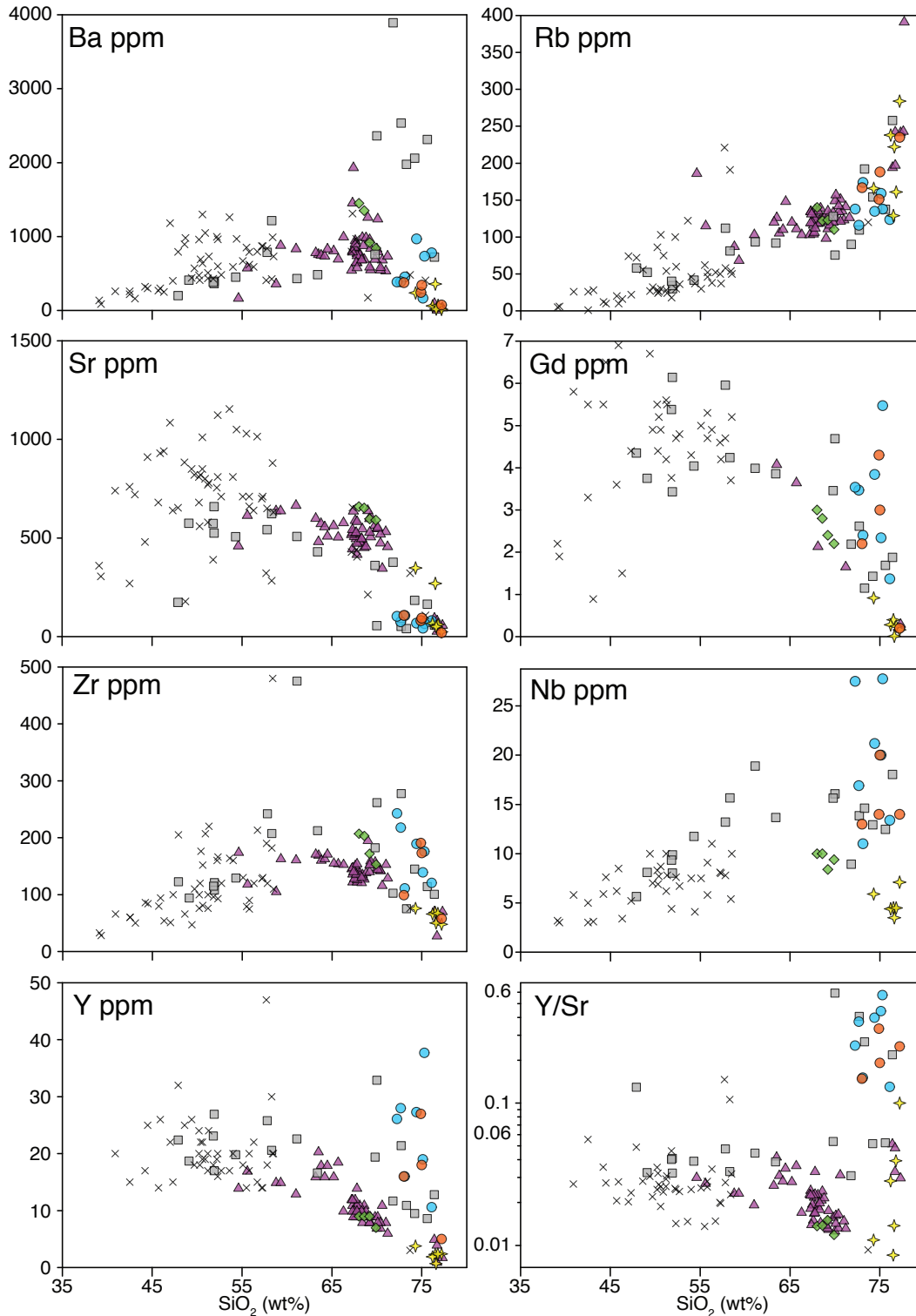


Figure 4. Trace element data for the Alabama Hills Granite and other plutonic rocks in the eastern Sierra Nevada. Data for the main facies of the Alabama Hills Granite are distinct from Whitney Intrusive Suite data with enrichments in HFSEs including Zr, Y, and Nb, as well as middle REEs including Gd. Sample RP15-02 from the fine-grained facies tracks high-silica rocks of the Whitney suite except for Y/Sr, where it plots with the Alabama Hills Granite and Bullfrog and Independence plutons. Symbols and data sources as in Fig. 3.

interpret geologically (Miller et al., 2007; Samperton et al., 2015; Schaltegger et al., 2015), and because the interpretations do not depend on precise relationships between samples, weighted mean ages are not assigned to the Alabama Hills Granite samples (Table 3; Fig. 7). Instead, the range of the 2σ analytical uncertainties for individual analyses from each sample is provided. A weighted mean age is calculated for dike sample AH14-01, which likely had a simpler thermal history.

Sample AH14-01 was collected from a diabase dike in the northern Alabama Hills metavolcanic complex. Four concordant zircon analyses yield Th-corrected $^{206}\text{Pb}/^{238}\text{U}$ ages spanning 1.3 Ma from 152.0-150.7 Ma. They yield a weighted mean age of 151.34 ± 0.13 Ma with a mean square of weighted deviates (MSWD) of 8.0.

Samples AH14-03, RP15-01, RP15-03, and RP15-04 are all concordant within uncertain-

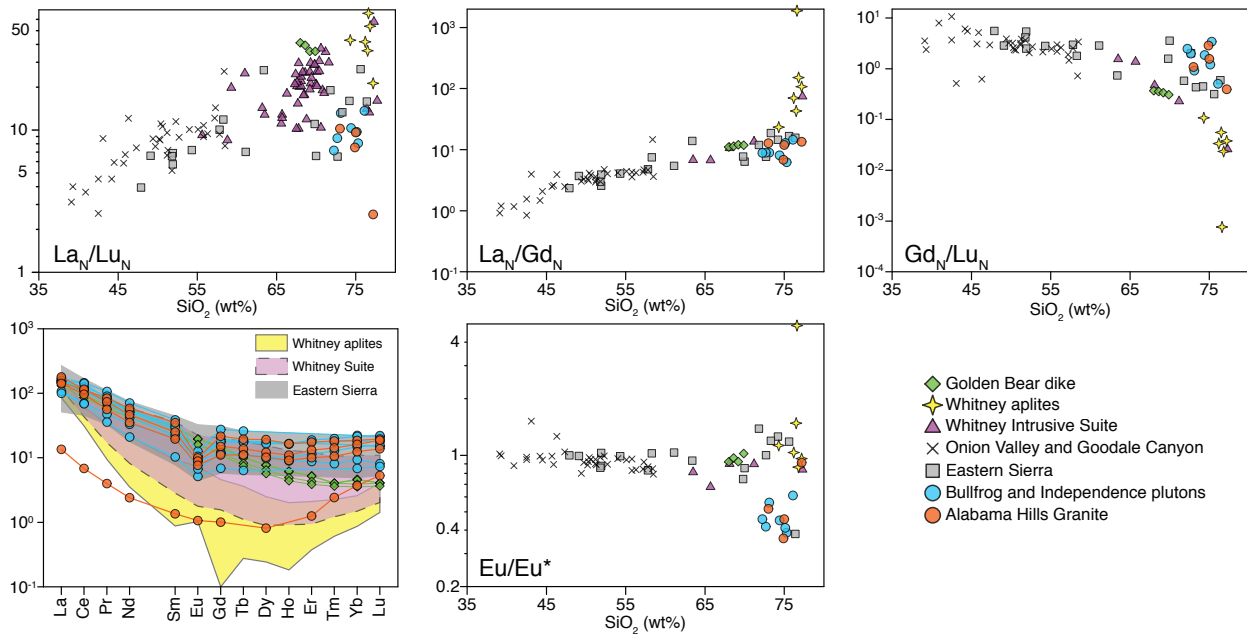


Figure 5. Rare earth element for the Alabama Hills Granite and other plutonic rocks in the eastern Sierra Nevada. Except for sample RP15-02, the Alabama Hills Granite has flatter REE patterns, with lower La_N/Lu_N and higher Gd_N/Lu_N , than Whitney Intrusive Suite rocks of similar silica concentrations. In addition, main facies Alabama Hills Granite samples are distinguished from Whitney suite rocks by having Eu/Eu^* near 0.4. In contrast to the Whitney Intrusive Suite, REE data for the Bullfrog and Independence plutons match the Alabama Hills Granite well. The low light REE concentrations in sample RP15-02 are likely a primary feature, as suggested by ϵNd_i values similar to the other Alabama Hills Granite samples. Data sources as in Fig. 3.

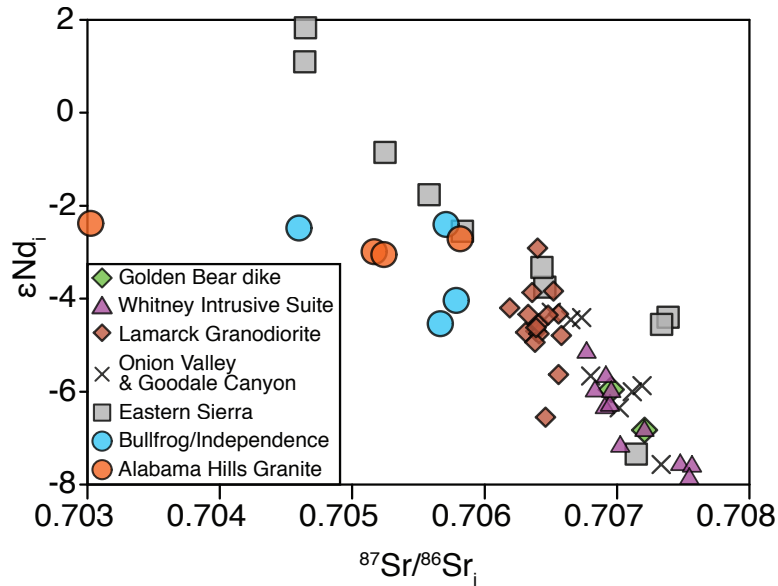


Figure 6. Strontium and Nd isotopic data corrected to initial values for the Alabama Hills Granite and other plutonic rocks of the eastern Sierra Nevada. Strontium is less radiogenic and Nd is more radiogenic in the Alabama Hills Granite compared to the Whitney Intrusive Suite. In addition, the Alabama Hills Granite data are matched well by Sr-Nd data for the Bullfrog and Independence plutons. Note that we interpret the lowest $^{87}\text{Sr}/^{86}\text{Sr}_i$ datum for the Alabama Hills Granite (sample RP15-02) to be compromised by post-crystallization open system processes, though the Sm-Nd system is considered robust (see text). Data sources as in Fig. 3, except: Lamarck Granodiorite (Coleman et al., 1992); Whitney Intrusive Suite (Hirt, 2007).

ty. Sample AH14-03 gave the oldest age spectrum, with seven individual zircons ranging 103.1-102.2 Ma. Seven zircons from sample RP15-03 yielded a nearly identical spectrum from 103.0-102.1 Ma. Four zircons each were dated in samples RP15-01 and RP15-04. Sample RP15-04 gave an age range of 102.8-102.0 Ma. Three zircons from RP15-01 overlap in age from 102.6-101.2 Ma, although the total range is encompassed by one zircon with high analytical uncertainty due to a relatively poor Pb isotopic analysis (Fig. 7). One younger grain has an age of 99.5 ± 0.3 Ma. Excluding the two youngest grains from sample RP15-01, 20 individual zircon analyses from the Alabama Hills Granite span 1.3 Ma from 103.1-101.8 Ma.

DISCUSSION

Interpretation of new geochronology

The 151.34 ± 0.13 Ma age for a mafic dike from the northern Alabama Hills metavolca-

nic complex suggests it is correlative with the Late Jurassic Independence dike swarm. To our knowledge, this is the most precise date yet for any dikes from the swarm. Although it is older than the nominal ~ 148 Ma age of the Late Jurassic portion of the swarm (Chen and Moore, 1979), the swarm may contain dikes as old as 152 ± 3 Ma (unpub. data by J.M. Mattinson, E.R. Schermer, and C.J. Busby cited in Schermer and Busby, 1994).

New data for the Alabama Hills Granite indicate it was assembled from 103-102 Ma; one grain with an age of 99.5 ± 0.3 Ma from sample RP15-01 is interpreted to reflect Pb-loss that was incompletely mitigated by chemical abrasion pre-treatment (Fig. 7). Although this study did not analyze the same outcrop dated by Chen and Moore (1982; precise location indicated in Stone et al., 2000), the samples dated here are likely representative of zircon crystallization ages throughout the pluton because of their spatial distribution. Note that the full zircon age range in the pluton may be 60% longer than what we have measured due to undersampling (Glazner and Sadler, 2016).

The 17-18 Ma difference between these new data and the 85 Ma age of Chen and Moore

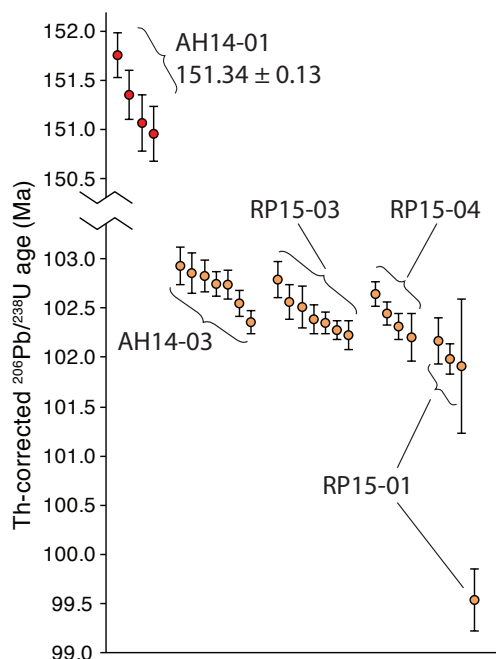


Figure 7. New CA-ID-TIMS age-sorted data for a mafic dike (interpreted as a member of the Independence dike swarm) from the Jurassic volcanic complex of the Alabama Hills (AH14-01) and the Alabama Hills Granite (four other samples). Each filled circle represents a single zircon analysis; vertical bars indicate 2σ analytical uncertainty. Note change in scale on ordinate. Weighted mean and 2σ analytical uncertainty provided for sample AH14-01 (see text for other sample age interpretations). Youngest zircon from sample RP15-01 interpreted to reflect Pb-loss.

(1982) is likely due to pervasive Pb-loss in the large fractions they dated (10 mg zircon each). These data suggest that geochronology performed on bulk zircon fractions, often without physical or chemical pre-treatment to remove Pb-loss zones, should be interpreted with caution because the data may not accurately reflect zircon crystallization ages. Large discrepancies (more than 5 Ma) between older and modern U-Pb dating techniques have also been documented nearby in the Independence pluton (Kylander-Clark et al., 2005), in the northern Mount Givens Granodiorite (Frazer et al., 2014), in the eastern Bass Lake Tonalite (Lackey et al., 2012), and in the Mineral King pendant (Klemetti et al., 2014).

Relationship of the Alabama Hills Granite to the Sierra Nevada batholith

Whitney Intrusive Suite

Chen and Moore's (1982) geochronology for the Alabama Hills pluton (85 Ma) and Whitney Intrusive Suite (87-83 Ma) led some workers to suggest they were originally contiguous and shared a genetic history (Ali et al., 2009), although Stone et al. (2000) considered the two bodies to be separate due to their dissimilar mineralogy. New zircon data for the Whitney suite (90.6-84.8 Ma; Davis, 2010) and geochronologic and geochemical data for the Alabama Hills Granite permit reevaluation of the links between the two. The age data demonstrate the Alabama Hills pluton is 12-17 Ma older than Whitney Intrusive Suite, precluding a genetic link between them. In addition, trace element and REE data are distinctive. The main body of the Alabama Hills Granite is enriched in HFSE as well as middle and heavy REEs relative to high-silica rocks of the Whitney Intrusive Suite (Figs. 4, 5). Europium anomalies are also more pronounced in the Alabama Hills samples. The adjacent intrusive complexes also appear to have different sources, as the Alabama Hills Granite is isotopically more primitive than the Whitney Intrusive Suite (Fig.

6), except in $\delta^{18}\text{O}$ in zircon (Lackey et al., 2008). These data, in sum, suggest different petrogenetic processes led to the high-silica granites in the Alabama Hills and the Whitney suite.

The fine-grained facies of the Alabama Hills Granite (sample RP15-02) appears to be a late-stage differentiate. The scoop-shaped REE pattern is similar to Sierra Nevada aplites interpreted to have been removed from titanite-bearing sources (Glazner et al., 2008), but it differs in having an enigmatic depletion in light REEs (Fig. 5). The REE pattern is probably a primary feature because the sample yielded ϵNd_i similar to other Alabama Hills samples (Table 2, Fig. 6). Although the Sm-Nd systematics are likely robust, the Rb-Sr system was apparently disturbed by open-system processes, yielding unrealistic $^{87}\text{Sr}/^{86}\text{Sr}_i$ near 0.703.

Mid-Cretaceous plutons near the Sierra crest: the Kearsarge intrusive suite

Several plutons north of the Alabama Hills along the Sierran crest and range front appear to be plausible temporal and geochemical matches for the Alabama Hills Granite (Fig. 6). These include the high-silica Bullfrog and Independence plutons, which are bisected by the more mafic Dragon pluton (Moore, 1963, 1981; Stone et al., 2000), and the smaller Sardine and Diamond plutons. Modern zircon geochronology indicates the Independence pluton is 102.5 ± 0.2 Ma, and the Bullfrog pluton has an age of ~ 100 Ma (Kylander-Clark et al., 2005; Davis, 2010). Bulk zircon geochronology mostly agrees with the modern techniques, and suggests the Dragon, Sardine, and Diamond plutons may have also been emplaced in the period 106-102 Ma (Chen and Moore, 1982; Saleeby et al., 1990), overlapping the Alabama Hills Granite. Compositional and isotopic analyses of the Sierran plutons also bear similarity to the Alabama Hills Granite, and strengthen the significance of those only dated by bulk zircon techniques. The Bullfrog and Independence plutons are enriched in HFSE, middle and heavy REEs, have high Y/Sr, and comparable Eu

anomalies to the main Alabama Hills pluton (Figs. 4, 5; Wenner and Coleman, 2004; Kylander-Clark et al., 2005; Coleman et al., 2012). Isotopic data for the plutons range from $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.7045$ to 0.7058 and $\epsilon\text{Nd}_i = -2.4$ to -4.5 (Fig. 6; Saleeby et al., 1990; Wenner and Coleman, 2004; Kylander-Clark et al., 2005). The plutons also have relatively homogeneous $\delta^{18}\text{O}(\text{Zrc})$, ranging 4.79 - 5.40% (Lackey et al., 2008). Collectively, the isotopic data suggest all of these plutons were derived from an enriched mantle source or juvenile mantle derivative with mantle-like O, $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.705$ - 0.706 , and $\epsilon\text{Nd}_i \approx -2.5$ to -4.5 . These enriched mantle characteristics are slightly less evolved than those proposed for the nearby Lamarck Granodiorite (Coleman et al., 1992).

On the basis of robust chemical, isotopic, and age similarities between the Bullfrog, Independence, and Alabama Hills plutons, we hereafter group them into the informal Kearsarge intrusive suite, after numerous geographic features in the suite that bear the name Kearsarge. The Dragon, Diamond, and Sardine plutons are also included based on their isotopic similarities to the Alabama Hills, Bullfrog, and Independence plutons, although additional geochronological analyses are required to confirm such an assignment. The McGann pluton, a quartz-poor quartz monzonite to the north of the Independence pluton (Moore, 1963), could also be correlative but has no age or chemical data to test the relationship.

Intrusive rocks of the Kearsarge intrusive suite are intimately associated with the Oak Creek metavolcanic pendant. The pendant exposes mostly Middle Jurassic rocks, but also includes subordinate mid-Cretaceous metavolcanic rocks dated to ~ 110 - 109 Ma by bulk zircon geochronology (Saleeby et al., 1990). The rhyolite sill has $^{87}\text{Sr}/^{86}\text{Sr}_i$ similar to the Kearsarge intrusive suite, of 0.7052 , whereas two analyses of the ash-flow tuff yielded $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.7070$ and

0.7062. These data suggest that mid-Cretaceous volcanic rocks in the pendant may have a similar source to the Kearsarge intrusive suite, and perhaps a close temporal history.

Tectonic origins of the Alabama Hills block

The Alabama Hills block is highly dilated by the Independence dike swarm; realignment of the Alabama Hills with similar high-dilation zones in the Sierra Nevada batholith could require 55 km of dextral offset between them (Bartley et al., 2007). Because isotopic data suggest the Alabama Hills Granite is part of the Kearsarge intrusive suite, such hypothesized dextral offset would have had to accumulate between the emplacement of highly-dilating Independence dikes in the Alabama Hills (~151.3 Ma) and emplacement of the Alabama Hills Granite ca. 103 Ma. Although a large amount of offset is possible, it is unnecessary. First, the swarm strikes across Owens Valley at a low angle, so correlations are subject to considerable uncertainty (Bartley et al., 2007). Second, there is no preserved evidence for Late Jurassic-Early Cretaceous dextral shear zones in the eastern Sierra, although such shear zones could have been obliterated by Cretaceous magmatism or covered by sediments in Owens Valley. Third, only the southern limit of the high dilation zone of the Independence dikes is exposed in the Alabama Hills. There, dilation values drop from 29% in the northern Alabama Hills to 4% just west of Lone Pine (Bartley et al., 2007). A similar southerly decrease in dilation values occurs in the Woods Lake basin (44% to 2%), approximately along strike of the dike swarm from the Alabama Hills (Bartley et al., 2007). Glazner et al. (2008a) also noted compositional similarities between dikes in the Alabama Hills and Woods Lake basin, suggesting these areas may have been in close proximity when the swarm was emplaced. Altogether, it appears large amounts of offset between the Ala-

bama Hills and the Sierra Nevada batholith are not necessary. Smaller amounts of offset may be possible.

The relationship between the Kearsarge intrusive suite exposed in the Sierra Nevada, proper, and the Alabama Hills Granite suggests the pluton could have originally been emplaced further north, adjacent to the range-front Independence pluton. Such a scenario must account for the distinctive ~15-20 m wide, ~83.4 Ma Golden Bear dike, which is not present in the Alabama Hills (Kylander-Clark et al., 2005). If the Alabama Hills Granite was emplaced adjacent to the rest of the Kearsarge intrusive suite, it must have been displaced dextrally by at least 10 km, and perhaps over 30 km, prior to emplacement of the Golden Bear dike 30 Ma later. However, another possibility is that the Alabama Hills Granite was always south of the latitude of the Golden Bear dike, permitting at most ~10 km of dextral offset of the block to its current location. This small amount of offset is preferred because it is simplest, and because there are no structures in the Alabama Hills Granite, the other Kearsarge intrusive suite plutons, or younger Whitney Intrusive Suite plutons that suggest range-front dextral shear. It is also possible the Alabama Hills Granite extends further north than its current exposure; numerous workers using have recognized a shallowly buried bedrock ridge extending north from the Alabama Hills approximately 12 km to the northeast, truncating near Independence (Kane and Pakiser, 1961; Pakiser et al., 1964; Stevens et al., 2013; Chaulk, 2016).

If the Alabama Hills have moved little relative to the main Sierra Nevada batholith since the Late Jurassic, it suggests the block is a suitable piercing point for areas highly-dilated by the Independence dike swarm, which is estimated to be offset by 75 km across Owens Valley (Bartley et al., 2007). Dextral offset of 65 km inferred by Kylander-Clark et al. (2005) is also strengthened by the recognition of the Kearsarge intrusive suite. Kylander-Clark et al. (2005) originally

correlated the leucogranite of Cactus Flat in the Coso Range with only the Bullfrog and Independence plutons. The addition of the Alabama Hills Granite and other small, isotopically similar granitoids from Sierran range front makes this mid-Cretaceous suite more prominent (Fig. 2). Thus, the leucogranite of Cactus Flat, which has a low negative Eu anomaly ($\text{Eu}/\text{Eu}^* = 0.44$) and similar $^{87}\text{Sr}/^{86}\text{Sr}_i$ (0.7060), may have also been part of the Kearsarge intrusive suite, which was subsequently dismembered by dextral (Kylander-Clark et al., 2005; Bartley et al., 2007) and normal (Moore, 1963, 1981; Stone et al., 2000; Ali et al., 2009) faulting.

The Enigmatic mid-Cretaceous central Sierra Nevada

For the purposes of the following discussion, the Cretaceous batholith is broken into three intervals of time: the Early Cretaceous, from 124-110 Ma; the mid-Cretaceous, from 110-98 Ma; and the Late Cretaceous, from 98-83 Ma. The latter period includes the Sierra Crest magmatic event of Coleman and Glazner (1997), extended to slightly younger ages due to new geochronology (Kylander-Clark et al., 2005). These temporal groups broadly correspond to the western, axial, and eastern batholithic belts of Nadin and Saleeby (2008), although recognition of the Kearsarge intrusive suite in the eastern Sierra suggests that east-west spatial patterns require additional consideration.

Controls on the location of the Kearsarge intrusive suite

Workers have long recognized that plutons of the Cretaceous Sierra Nevada batholith become progressively younger towards the east through intrusive relationships and isotopic dating (e.g., Calkins, 1930; Stern et al., 1981), with Chen and Moore (1982) calculating eastward magmatic front migration at 2.7 mm/a. These early hypotheses predict mid-Cretaceous plutons to be located in the axial part of the batholith, between Early Cretaceous and Late Cretaceous

rocks. Numerous suites dated by modern methods follow this prediction (Fig. 2), and additional geochronology for plutons between those suites may strengthen the relationship. The Kearsarge intrusive suite stands in contrast to this pattern, as it is 45-60 km inboard of the Ash Mountain complex when measured perpendicular to the ~N20W trend of the Cretaceous batholith (Fig. 2).

The migrating magmatic front has commonly been attributed to flattening of an eastward-subducting Farallon slab, although the mechanisms causing flattening are debated (Coney and Reynolds, 1977; Chen and Moore, 1982; Tobisch et al., 1995; Nadin et al., 2016). The recently-proposed westward-dipping subduction hypothesis is intriguing, where slab rollback causes the migrating pattern of ages (e.g., Hildebrand, 2009, 2013). However, westward-dipping subduction requires the Kearsarge suite to have been emplaced close to the trench, where water supply would be limited to low P-T sediment dewatering reactions (Kerrick and Connolly, 2001; Rüpke et al., 2004). Trenchward emplacement is also problematic as it calls for little interaction with the mantle wedge, which is difficult to reconcile with the mantle-like isotopic data for the suite (Fig. 6; Lackey et al., 2008).

Another explanation for the Kearsarge intrusive suite's location is that it is an expression of backarc magmatism over an eastward-dipping subduction zone. The Central Andes backarc may serve as an analog. There, magmatism has been documented as far as 300 km east of the active Western Cordillera in association with strike-slip systems or lineaments (Riller et al., 2001; Chernicoff et al., 2002; Gioncada et al., 2010). These examples are intriguing because the Sierra Nevada may share some commonalities with them. For example, the Kearsarge intrusive suite is located near the southernmost recognized portion of the Sawmill Lake shear zone, which deforms Jurassic plutonic and metavolcanic rocks of the Oak Creek pendant (Fig. 2; Bartley et al., 2012) and was active through ~95 Ma (Mahan et al., 2003; Stearns and Bartley, 2014). The shear

zone is apparently not present in the larger plutons of the Kearsarge intrusive suite. However, the small Sardine pluton appears to be deformed along with mid-Cretaceous metavolcanic wallrocks of the Oak Creek pendant (Saleeby et al., 1990). Thus, the Sawmill Lake shear zone, or some correlative deeper structure, could have served to focus magmas of the Kearsarge intrusive suite; a similar process has been invoked for the crack-seal assembly of the 95 Ma late synkinematic McDoogle pluton (Mahan et al., 2003; Bartley et al., 2008; Stearns and Bartley, 2014).

The Independence dike swarm suggests an alternative to the Sawmill Lake shear zone for focusing the Kearsarge intrusive suite. Coleman et al. (2000) documented 94-90 Ma dikes in the swarm that superficially resemble Jurassic dikes in composition and orientation. The similarities between Cretaceous dikes and Jurassic dikes despite significant changes in convergence direction (i.e., sinistral to dextral; Engebretson et al., 1985; Glazner, 1991) suggests that regional strain may not be a major factor controlling the orientation of intrusive rocks in the batholith. Coleman et al. (2000) hypothesized that the common orientation of Jurassic and Cretaceous dikes may be controlled by fractures in pre-existing (i.e., Middle Jurassic or older) rocks, or by the orientation of the North American margin as realized by a relative lack of coupling between North America and the subducting Farallon plate, causing extension perpendicular to the plate's margin (Tobisch et al., 1986). Thus, the Kearsarge intrusive suite may have taken advantage of similar pre-existing fractures that led it to intrude inboard of other mid-Cretaceous rocks. If Jurassic fractures or the orientation of the continental margin were controlling factors for the inboard extension of ~105-100 Ma magmatism, this hypothesis predicts that undiscovered mid-Cretaceous plutons may be present along a similar trend (i.e., N30°W to N40°W; Tikoff and Teyssier, 1992; Coleman et al., 2000) between the Kearsarge intrusive suite and the ~105-100 Ma arc in the central

part of the batholith. Therefore, plutons from the Marion Peak quadrangle (Moore, 1978) make intriguing targets to test this hypothesis.

Are mid-Cretaceous plutonic rocks different from the rest of the batholith?

The new geochronology and geochemistry presented here adds to a growing high-quality dataset documenting mid-Cretaceous magmatism in the Sierra Nevada batholith, including the 106.6-103.5 Ma intrusive suite of Yosemite Valley (Ratajeski et al., 2001; Nelson et al., 2013; Putnam et al., 2015), the 106.5-98 Ma Shaver Intrusive Suite (Frazer et al., in prep.) and the 105.5-102.5 Ma Ash Mountain complex in Sequoia National Park (Holland et al., 2013). These intrusive suites superficially appear to be more felsic than the preceding and following magmatism in the batholith (Clemens-Knott and Saleeby, 1999; Lackey et al., 2012; Coleman and Glazner, 1997). The Dinkey Creek Granodiorite of the Shaver Intrusive Suite is the only well-dated voluminous granodiorite in the mid-Cretaceous batholith (Frazer et al., in prep.). However, the suite also includes numerous texturally and chemically distinct granite plutons (Frazer et al., in prep.).

To test whether the mid-Cretaceous batholith is more felsic than the Early or Late Cretaceous, a simple analysis of the areal extent and composition of Cretaceous plutons in the central Sierra Nevada batholith was performed. Plutons were categorized by their USGS compositional designations (Bateman, 1992). A more optimal methodology for investigating spatio-temporal compositional trends would involve regular sampling on a grid pattern to avoid biasing, as has been carried out in the Peninsular Ranges batholith (Baird et al., 1979; Lee et al., 2007; Lee and Morton, 2015). Unfortunately, despite thousands of geochemical analyses in the central Sierra Nevada, a structured sampling regimen has not been undertaken. The method used here is obvi-

ously a crude measure, but the large time bins used and the stark results make the analysis illustrative of broad compositional trends.

Two sets of analyses were performed in the Cretaceous batholith from 36.5-38.25°N (Fig. 8). The first included only plutons considered to be “well-dated”. As demonstrated in the Alabama Hills and elsewhere, older geochronologic data may be compromised by unresolved Pb-loss or inheritance due to the use of large fractions or a lack of physical or chemical pre-treatment (e.g., Krogh, 1982; Mattinson, 2005). Therefore, in the “well-dated” analysis, only plutons dated by LA-ICP-MS, secondary ion mass spectrometry, sensitive high resolution ion microprobe, or ID-TIMS on pre-treated zircons were used. One exception was the inclusion of mafic ring dike complexes in the western Sierra (Clemens-Knott and Saleeby, 1999) because such mafic rocks are unlikely to contain inherited zircon, and if affected by Pb-loss, the rocks would be older and would not affect the younger age groups. The second analysis performed was “complete”, and tentatively assigns all mapped plutons from 36.5-38.25°N to age groups based on pluton location, field data from USGS mapping, and cautious interpretation of existing geochronologic data. The well-dated analysis is considered more robust, although the complete analysis is useful in determining whether plutons dated by modern geochronology are biased toward particular compositions.

The compositions of the Early and Late Cretaceous parts of the batholith accord with conventional wisdom (Fig. 8). The Early Cretaceous batholith is the most mafic, with more than 70 area% comprising tonalite in both the well-dated and the complete datasets, and the remaining areas taken up by mafic ring dike complexes and smaller proportions of more felsic rocks including granodiorite, trondhjemite, and granite. The Late Cretaceous batholith is dominated by the granodiorites of the Sierra Crest magmatic event (Coleman and Glazner, 1997), covering

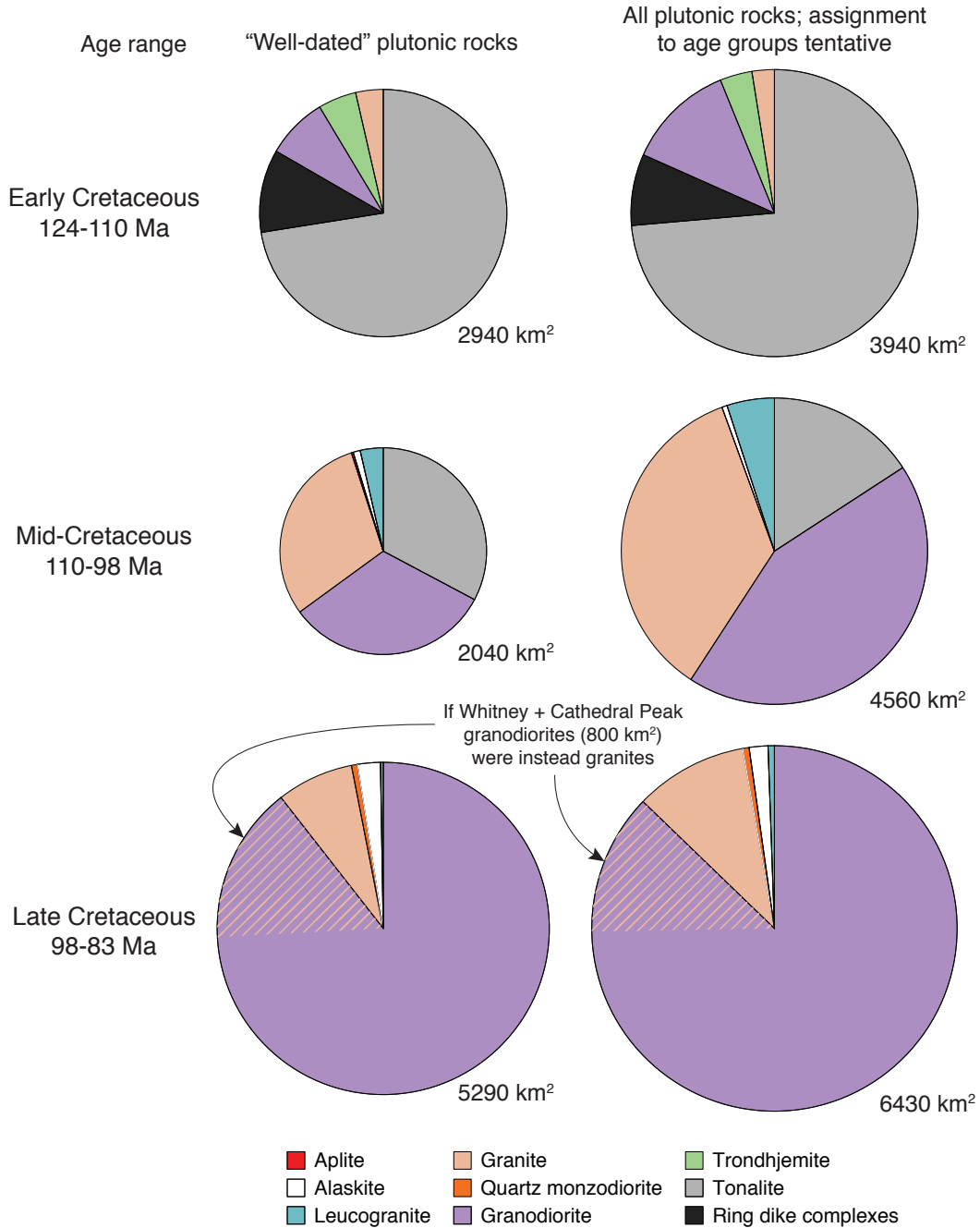


Figure 8. Areal analyses of Cretaceous intrusive rocks of the Sierra Nevada by age grouping. Rock compositional type taken from descriptions in USGS publications where available. Rationale for age ranges given in text. In left column, “well-dated” plutonic rocks are those dated by U-Pb zircon using LA-ICP-MS, SIMS, SHRIMP, or TIMS on single or several zircons pretreated by air- or chemical abrasion such that the effects of Pb-loss and inheritance are minimized compared to bulk fractions. For the right column, plutons were assigned to age groups using USGS-defined intrusive relationships and cautiously-interpreted legacy age data. The relatively similar compositional breakdowns between well-dated and complete areal analyses for each age group suggests there is little compositional bias in the rocks that have been dated, and that granites are more common in the mid-Cretaceous (110-98 Ma) compared to the Late Cretaceous (98-83 Ma). In the Late Cretaceous group, hatched lines show the effects of changing the designations of the Cathedral Peak and Whitney plutons from granodiorite to granite. In both cases, granodiorite dominates granite in the Late Cretaceous. Diorites and gabbros are not included in the analyses due to a general lack of age control.

more than 85 area% in both datasets. Granites are also somewhat common in the Late Cretaceous, comprising 7.5 area% and 10 area% in the well-dated and complete datasets, respectively. Therefore, granodiorite:granite ratios by area in the Late Cretaceous are ~11:1 to ~8.5:1. Because plutons commonly extend across compositional fields in QAP diagrams (e.g., many analyses of the Cathedral Peak Granodiorite fall in the granite field; Bateman, 1992), the effect of changing the designation of the Cathedral Peak and Whitney plutons to granite was also tested. Despite a combined area of 800 km² in the two plutons, the results show that Late Cretaceous granodiorites comprise nearly 75 area% in both datasets and still dwarf granites by a ratio of ~3:1 by area (Fig. 8).

In contrast to the Early and Late Cretaceous, the mid-Cretaceous has subequal amounts of granite and granodiorite by area (Fig. 8). If leucogranites are included with granites, then the more silicic rocks are areally superior. This result is at odds with the common notion that the youngest parts of the batholith are the most felsic (e.g., Bateman et al., 1963; Bateman, 1992; Nadin and Saleeby, 2008). The relatively small areal extent of the “well-dated” mid-Cretaceous batholith does not appear to bias the analysis toward granitic compositions; the “complete” mid-Cretaceous dataset contains subequal amounts of granodiorite and granites + leucogranites. Note that the large amount of “well-dated” tonalite in the mid-Cretaceous is due to the inclusion of the Bass Lake Tonalite’s “eastern domain” from Lackey et al. (2012).

We suggest that the large differences in granodiorite:granite areal ratios between the mid-Cretaceous (~1:1) and Late Cretaceous (~11:1) overcome the uncertainties inherent in the use of general pluton compositions. Even if undated granodiorites and granites were apportioned to the mid- or Late Cretaceous “complete” datasets in error, unrealistic scenarios are required to make the areal ratios comparable between the time periods. For example, if all granites of

unknown age are added to “well-dated” Late Cretaceous, and all granodiorites of unknown age are added to the “well-dated” mid-Cretaceous, granodiorite:granite areal ratios would remain larger in the Late Cretaceous, at $\sim 3.7:1$, than the mid-Cretaceous, $\sim 2.9:1$. Therefore, the granites exposed in the mid-Cretaceous batholith are greater in area than granites of the Late Cretaceous Sierra Crest magmatic event.

These results are broadly similar to the more robust measurements made in the Peninsular Ranges batholith. There, high-silica granites comprise a larger area in the western batholith (37%) than the eastern batholith (16%), perhaps because the western batholith is a relatively shallower exposure (Ague and Brimhall, 1988b; Lee and Morton, 2015). However, high-silica granites of the Sierran mid-Cretaceous appear to be abundant at both shallow and mid-crustal levels. For example, the Kearsarge intrusive suite was emplaced at low pressures (20-200 MPa; Richardson, 1975; Ague and Brimhall, 1988b; Sorensen et al., 1998), whereas the axial mid-Cretaceous suites were emplaced at higher pressures (230-400 MPa; Ague and Brimhall, 1988b; Tobisch et al., 1993; Putnam et al., 2015). Furthermore, if depth is a controlling factor in the expression of granites in the Sierra Nevada, more granites would be expected in the areas of the shallowly-emplaced Late Cretaceous Tuolumne (~ 150 MPa; Gray et al., 2008; Challener and Glazner, 2017) and Whitney intrusive suites (20-350 MPa; Ague and Brimhall, 1988b). Therefore, some other factor likely controls the genesis and expression of high-silica granites in the Sierra Nevada.

There are leucogranites, and there are leucogranites

Recent debate on the origins of high-silica granite has centered around whether high-silica magmas are distilled from batholith-scale crystal mushes in the mid- to upper crust (e.g.,

Bachmann and Bergantz, 2004; Hildreth, 2004; Gualda and Ghiorso, 2013; Lee and Morton; 2015) or are partial melts of deeper sources, including the crust and/or mantle (e.g., Wenner and Coleman, 2004; Sisson et al., 2005; Stevens et al., 2007; Clemens et al., 2010; Villaros et al., 2009; Tappa et al., 2011; Coleman et al., 2012). Here we comment on how new data for high-silica granites from the Sierra Nevada may bear on this lively and ongoing conversation.

There is little doubt that one class of high-silica granites—aplites—is the product of shallow-level crystal-liquid separation from a crystalline mush (Eichelberger et al., 2006; Glazner et al., 2008b). Low pressure haplogranitic compositions and trace element enrichments or depletions strongly support such an origin (Glazner et al., 2008b; Putnam et al., 2015). Aplites separated from titanite-bearing mushes have “scoop-shaped” REE patterns and Y depletions (Glazner et al., 2008b), whereas those derived from titanite-free sources do not (Bachmann and Bergantz, 2008; Putnam et al., 2015).

Evidence also suggests high-silica melt segregation occurs cyclically on scales of a km or less in large, incrementally-emplaced granodiorites, with trends toward similar trace element enrichments and depletions as aplites at high SiO₂ contents (Economos et al., 2010; Coleman et al., 2012). These high-silica magmas (“cycles”) may have segregated from a granodiorite mush, leading to gradational field and chemical relationships, but became trapped at the structural discontinuity of previously solidified granodiorite (Coleman et al., 2012).

The high-silica granite plutons studied here stand in sharp contrast to aplites and cycles, in some cases quite literally. In the field, many of the granite plutons of the Sierra Nevada are noted for their sharp intrusive contacts (e.g., Bateman, 1992; Holland et al, 2013). In places where contacts are gradational over scales of less than a meter, they are interpreted to reflect coeval emplacement of discrete magma batches rather than a genetic relationship where one

magma was derived from the other (Ratajeski et al., 2001; Putnam et al., 2015). The common absence of gradational contacts between high-silica granite plutons and granodiorites in the Sierra makes it difficult to envision a process through which the former is derived from the latter at emplacement level (e.g., Bachmann and Bergantz, 2004, 2008; Lee and Morton, 2015).

Whereas field evidence appears to contradict hypotheses that call for the generation of granites from granodiorites, perhaps the current exposures of the batholith are structurally biased (Lee and Morton, 2015). However, the chemical evidence against derivation of the high-silica granites from granodiorites in the Sierra Nevada is also compelling. For example, the Shaver Intrusive Suite (Frazer et al., in prep.) is the only well-dated mid-Cretaceous suite that comprises a large granodiorite pluton and discrete, largely coeval granites. Despite a temporal link, the granite plutons have more variable O, Sr, and Nd isotopic compositions than the granodiorite and other mafic rocks, precluding a scenario in which the granites were derived from the granodiorite.

High-silica granite plutons throughout the Sierra Nevada, regardless of age, are chemically distinct from aplites and cycles. Sierran high-silica granite plutons are enriched in HFSEs, Y, and middle REEs relative to aplites and cycles at similar SiO_2 concentrations, suggesting the plutons' magmas evolved in the absence of titanite, in contrast to cycles and many aplites (Coleman et al., 2012). Coleman et al. (2012) suggested three possible mechanisms for the titanite-absent signature in high-silica granite plutons. The first two occur at emplacement level. First, titanite could be absent from the crystallization sequence because of hot and reducing magmatic conditions (Bachmann and Bergantz, 2008). Although reducing conditions have been inferred in parts of the central Sierra (e.g., Shaver Intrusive Suite), many high-silica suites were emplaced under oxidizing conditions that typify most of the batholith (Ague and Brimhall, 1988a). Despite the

differing conditions, high-silica granites throughout the Sierra are enriched in middle REE like Gd (Fig. 9), suggesting the varying redox states did not affect titanite. Another possibility is that titanite crystallization was suppressed until after extraction of the high-silica granites (Colombini et al., 2011). However, titanite is present in the Fish Canyon Tuff (Bachmann et al., 2002), and is not uncommon in other Cordilleran-type dacites and rhyolites (see compilation in Glazner et al., 2015). This suggests titanite can appear before magmas are 50% crystalline. In addition, if high-silica plutons owe their high Gd contents to extraction prior to extremely late titanite crystallization (e.g., if titanite saturation occurs at 76 wt% SiO₂; Colombini et al., 2011), such a scenario predicts there should be slightly less silicic complementary cumulates that are depleted in Gd. However, a data compilation for the Sierra Nevada does not reveal the putative cumulates (Fig. 9).

If shallow fractionation does not satisfy the requirements for the generation of these high-silica granites, perhaps a deeper genesis is possible (Coleman et al., 2012). Experimental and isotopic data have been used to suggest Sierran granites were derived from partial melting in the lower crust, particularly juvenile, amphibole-bearing mafic rocks that were themselves derived from variably enriched mantle (e.g., Coleman et al., 1992; Sisson et al., 2005; Hollande et al., 2013). Alternatively, it has been suggested that the chemical variation in granitic plutons is controlled by the composition of the source, and heterogeneities are thus linked to source heterogeneity (e.g., Clemens et al., 2010; Clemens and Stevens, 2016). Specifically, Clemens et al. (2011) suggested that the sources of I-type granites are likeliest to be intermediate-to-mafic arc volcanic rocks.

It is difficult to envision a unifying model for the generation of the high-silica granite plutons in the Sierra Nevada in light of their varied isotopic compositions. The Kearsarge intrusive

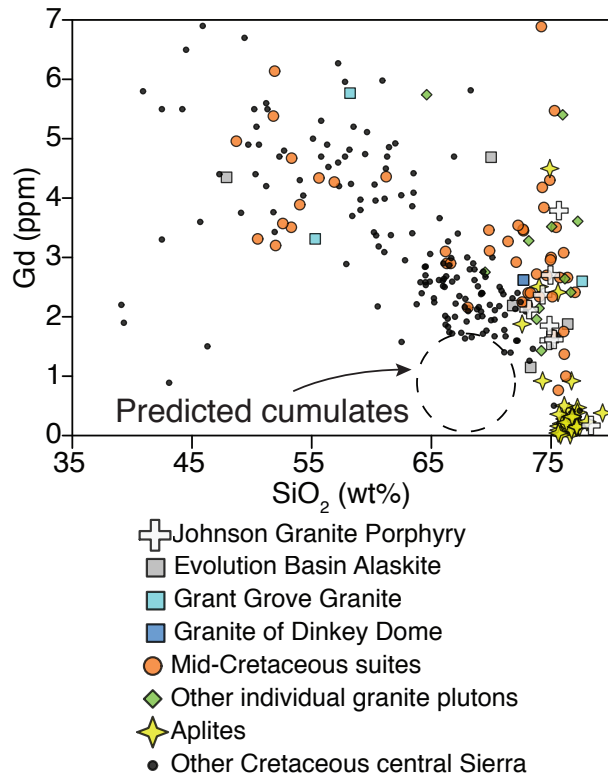


Figure 9. Plot of Gd concentration against silica concentration for Cretaceous plutonic rocks in the central Sierra Nevada. Data for notable granite plutons (Johnson Granite, Evolution Basin, Grant Grove, Dinkey Dome) plotted with separate symbols. Dashed circle indicates predicted composition of cumulates complementary to high-SiO₂ plutons if the high-Gd, high-SiO₂ plutons were separated from a mush prior to saturation of titanite (e.g., Colombini et al., 2011). The lack of the predicted cumulates suggests late saturation of titanite was uncommon in the Sierra Nevada. Data from Sisson et al. (1996), Truschel (1996), Ratajeski et al. (2001), Wenner and Coleman (2004), Kylander-Clark et al. (2005), Hirt (2007), Glazner et al. (2008b), Gray et al. (2008), Coleman et al. (2012), Holland et al. (2013), Putnam et al. (2015), Gaynor and Glazner (2016), Frazer et al. (in prep.), Frazer (unpub. data), and this study.

suite has mostly mantle-like isotopic compositions, including $\delta^{18}\text{O}(\text{Zrc})$ (Lackey et al., 2008), precluding an arc volcanic source as envisioned by Clemens et al. (2011). In contrast, the granites of the Shaver Intrusive Suite (Frazer et al., in prep.) have variable isotopic compositions and whole rock Ce anomalies suggesting they may have been derived from a variable metasedimentary source (e.g., Villaros et al., 2012). Yet both of these suites have high-silica granites that are enriched in HFSEs, Y, and middle REEs, with generally negative Eu anomalies. Thus, more work is necessary to identify the sources and mechanisms that led to the generation of these plutons and their distinct geochemical signatures.

High-silica granites and the genesis of granodiorites

Some workers have recently cast doubt on the possibility of a magma mixing origin for granodiorites based on the presence of curvilinear trends on Harker diagrams (Lee and Bachmann, 2014; Keller et al., 2015). For example, Lee and Bachmann (2014) suggested that mixing

processes are not significant in the generation of intermediate magmas because data compilations fail to show linear mixing trends between felsic and mafic rocks in P_2O_5 and Zr. However, data from large Sierran suites are equivocal (Fig. 10). Data for P_2O_5 in the Tuolumne and Whitney suites are nearly linear, particularly if the mafic end member is similar in composition to mafic enclaves. The Zr data for the Sierra Crest suites do follow a subtle curvilinear trend, with a peak between 60-65 wt% SiO_2 and a noted decrease in Zr at high silica concentrations (Fig. 10). If the putative magma mixing end members considered were only those compositions present in the Sierra Crest suites, then the Zr data appear to preclude a mixing origin.

However, the Sierran high-silica granites have similar or sometimes higher Zr concentrations to the Sierra Crest suites (100-200 or greater ppm Zr). Since the granodioritic intrusive

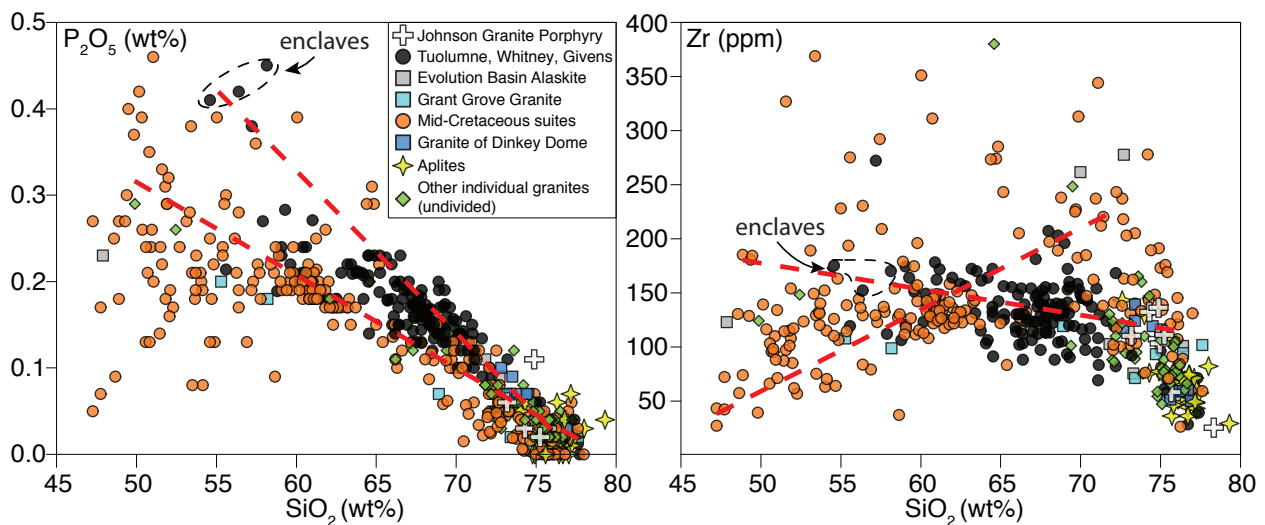


Figure 10. Plots of P_2O_5 and Zr against SiO_2 for three representative Sierra Crest intrusive suites (black circles; Tuolumne, Whitney, Mount Givens), as well as data for aplites from those intrusive suites (yellow stars), and other individual granite plutons as in Fig. 9. Phosphorus data for the Sierra Crest suites are approximately linear and do not appear to require an origin by fractional crystallization. Linear mixing trends between low- and high- SiO_2 parts of the mid-Cretaceous suites would produce granodiorite with lower P_2O_5 than the Sierra Crest suites, although if the mafic endmember is similar to enclaves, the pattern is reproduced. Note that the low- P_2O_5 , low SiO_2 rocks are subsets of the data of Holland et al. (2013) and Nelson et al. (2013). Zirconium data for the Sierra Crest suites are subtly curvilinear, with a noted decrease in Zr at high SiO_2 shown in the Sierra Crests intrusive suites and their aplites (black circles, yellow stars). Data for the mid-Cretaceous suites (orange circles) and the individual granite plutons (green diamonds) are highly scattered, although two hypothetical mixing trends (red dashed lines) between low- and high- SiO_2 components of those suites can reproduce the Zr concentrations of most Sierra Crest granodiorites. Note in particular that the high- SiO_2 granites have similar Zr concentrations (100-200 ppm) to granodioritic compositions. In contrast, the aplites and other low-Zr, high- SiO_2 intrusive rocks of the Sierra Crest suites are interpreted to be the products of shallow crystal-liquid separation (e.g., Glazner et al., 2008b; Coleman et al., 2012).

suites have largely flat Zr patterns with SiO₂ until the most felsic compositions, the high-Zr, high-silica granite plutons appear to be plausible end members for magma mixing scenarios. The low-Zr, high-silica parts of the intrusive suites are relatively small volume, generated at emplacement level by late crystal-liquid separation (e.g., Glazner et al., 2008b; Coleman et al., 2012). If magmas similar to the high-silica granite plutons were indeed end members in mixing to produce the voluminous Late Cretaceous suites, it may also explain the paucity of Late Cretaceous granites relative to the mid-Cretaceous (Fig. 8). Production of the Late Cretaceous suites would require large volumes of high-silica granitic magmas to mix with mafic magmas in order to generate the voluminous granodiorites. Perhaps high-silica magmas were produced at similar rates and volumes in the Late Cretaceous and the mid-Cretaceous, but in the Late Cretaceous Sierra Crest magmatic event, those high-silica magmas were largely consumed as mixing end members.

High-silica granites and volcano-pluton connections

Miller et al. (2003) noted a dichotomy in granitoids, grouping them into “hot” and “cold” classes on the basis of zircon inheritance and Zr saturation temperatures. Their hot granites lacked inherited zircon and averaged Zr saturation temperatures of 837°C Zr, whereas the cold granites had more than 10% “premagmatic” zircons and averaged 766°C (calculated using the calibration of Watson and Harrison, 1983). Such a dichotomy may also apply to the different types of high-silica rocks in the Sierra Nevada due to their contrasting Zr concentrations (Fig. 10). None of these rocks have more than 10% “premagmatic” zircon, although some of the Sierra Crest suites have antecrystic and/or xenocrystic zircon (e.g., Coleman et al., 2004; Miller et al., 2007; Frazer et al., 2014).

Zircon saturation temperatures were calculated for samples over 70 wt% SiO₂ from

individual high-silica granite plutons, Sierra Crest suites, and aplites using the updated parameterization of Boehnke et al. (2013). The results show that high-silica granite plutons have median zircon saturation temperatures of 721°C, with lower median saturation temperatures in high-silica parts of Sierra Crest suites (693°C) and aplites (661°C) (Fig. 11). The significance lies in the contrasts between the rock types, rather than their precise temperatures, which may not accurately reflect when most zircon in these rocks crystallized (Harrison et al., 2007). Broadly, it is logical that aplites would yield the lowest calculated temperatures, as they are interpreted as late-stage differentiates at shallow crustal levels (Eichelberger et al., 2006; Glazner et al., 2008b; Bachmann and Bergantz, 2008). More importantly, the data suggest that discrete high-silica granite plutons may have been slightly hotter as magmas than high-silica parts of Sierra Crest suites. After Miller et al. (2003), it follows that the individual granite plutons would have also been more eruptible than the high-silica members of the Sierra Crest or aplites. This supports the interpretations of Glazner et al. (2008b) and Coleman et al. (2012) that aplites and leucocratic cycles of the Sierra Crest suites resulted from emplacement level crystal-liquid separation. Those rock types are also poor matches for erupted high-silica rocks such as rhyolites (Glazner et al., 2015). In contrast, if the discrete high-silica granites were once relatively hot, crystal-poor (Miller et al., 2003) and eruptible, they may represent suitable intrusive equivalents to high-silica erupted rocks. It is also intriguing that the youngest volcanic rocks preserved in the Sierra Nevada batholith, such as the those at the Boyden Cave, Cinko Lake, Oak Creek, and Strawberry Mine pendants (Saleeby et al. 1990; Memeti et al., 2010b), and ash-flow tuffs and caldera collapse deposits in the Ritter Range (Fiske and Tobisch, 1994) are approximately mid-Cretaceous in age, when many high-silica granite plutons were emplaced (Fig. 8). Perhaps the distinct composition

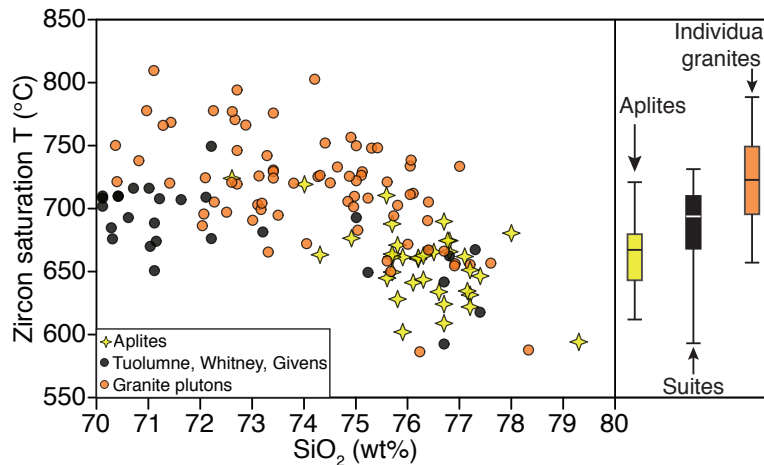


Figure 11. Compilation of calculated zircon saturation temperatures (Boehnke et al., 2013) for high-silica rocks of Sierra Crest intrusive suites (Tuolumne, Whitney, Mount Givens), individual granite plutons (Kearsarge, Ash Mountain, Shaver, Yosemite Valley, Evolution Basin, Dinkey Dome, Johnson Granite), and aplites. The pattern shows that individual granite plutons have higher zircon saturation temperatures than large intrusive suites or aplites at similar silica concentrations, suggesting that the individual granites may have been hotter as magmas than aplites or the high-silica portions of intrusive suites. Box-and-whisker plots in right panel elucidate the differences between the rock types; 1st and 3rd quartiles are contained in boxes; horizontal line indicates the median value. Whiskers extend to 95th and 5th percentiles. Data from Ratajeski et al. (2001), Wenner and Coleman (2004), Kylander-Clark et al. (2005), Hirt (2007), Lackey et al. (2006), Glazner et al. (2008b), Gray et al. (2008), Coleman et al. (2012), Holland et al. (2013), Putnam et al. (2015), Gaynor and Glazner (2016), Frazer et al. (in prep.), Frazer (unpub. data), and this study.

of mid-Cretaceous magmatism resulted in a more voluminous volcanic expression than the Late Cretaceous Sierra Crest magmatic event.

CONCLUSIONS

New geochronologic data indicate the Alabama Hills Granite is 103-102 Ma, about 17 Ma older than previously recognized. Although its individual outcrop area is relatively small, the recognition of this granite as mid-Cretaceous has implications for the Sierra Nevada batholith from Jurassic through Recent time. The data suggest the Alabama Hills Granite belongs to the heretofore unrecognized Kearsarge intrusive suite, which is characterized by its ~103-100 Ma age, silicic plutons, and mantle-like isotopic compositions ($^{87}\text{Sr}/^{86}\text{Sr}_i \approx 0.7045\text{-}0.7060$, $\epsilon\text{Nd}_i \approx -2$ to -5) relative to nearby Late Cretaceous rocks of the Sierra Crest. Without a link to the Whitney Intrusive Suite, the Alabama Hills block may have moved dextrally relative to the main Sierra Nevada batholith. However, it is likely it moved no more than 10 km dextrally since the

mid-Cretaceous, and perhaps since the intrusion of the ~151-148 Ma Independence dike swarm. As such, the Alabama Hills serve as a suitable piercing point for reconstructing dextral offset across Owens Valley, which previous workers have estimated at 65-75 km. The location of the Kearsarge intrusive suite some 60 km east of other mid-Cretaceous Sierran plutons suggests it may represent backarc magmatism that was focused along a pre-existing shear zone or fractures, similar to that observed in the Andes. The Kearsarge intrusive suite and other mid-Cretaceous intrusive complexes in the axial part of the Sierra Nevada batholith comprise approximately sub-equal amounts of granite and granodiorite, in contrast to the large suites of the Late Cretaceous Sierra Crest magmatic event, which are dominated by granodiorite. Unlike aplites and high-silica segregations in granodiorites, individual Sierran high-silica granite plutons appear to lack a titanite-removal signature, and it is suggested this is due to their origins deep in the crust. In addition, high-silica granite plutons are enriched in Zr relative to high-silica granites of the Late Cretaceous suites, suggesting the individual plutons: 1) may represent suitable end members for mixing with mafic melts to produce the voluminous granodiorites of the batholith; 2) may have once been hot, crystal-poor magmas that would have been capable of eruption, in contrast to the high-silica granites segregated from the Late Cretaceous suites at emplacement level.

TABLE 1. ALABAMA HILLS SAMPLE SUMMARY

Sample	Description	Age (Ma)	UTM-E*	UTM-N*	IGSN†
AH14-01	Independence dike (mafic)	151.34 ± 0.13	399189	4057055	P4700001C
AH14-03	Alabama Hills Granite	103.1-102.2	400282	4054245	P4700001E
RP15-01	Alabama Hills Granite	102.6-101.2	395073	4051373	P4700001F
RP15-02	Alabama Hills Granite (fine-grained facies)	Not dated	399492	4049916	P4700001G
RP15-03	Alabama Hills Granite	103.0-102.1	400230	4049235	P4700001H
RP15-04	Alabama Hills Granite	102.8-102.0	404115	4045307	P4700001I

Note: See text for zircon U-Pb age interpretations. Sample RP15-02 was not dated

*UTM coordinates given in NAD 83 datum, zone 11S

†IGSN – International Geo Sample Number. Additional sample metadata available at www.geosamples.org

TABLE 2. MAJOR AND TRACE ELEMENT DATA

Sample	RP15-01	RP15-02	RP15-03	RP15-04
Major element analyses* by ICP-OES				
SiO ₂	74.91	77.22	74.97	72.97
Al ₂ O ₃	12.34	11.81	12.19	12.98
Fe ₂ O ₃	2.00	1.32	2.27	2.40
MnO	0.061	0.017	0.076	0.056
MgO	0.15	0.06	0.22	0.21
CaO	0.62	0.34	0.83	0.89
Na ₂ O	3.96	3.31	3.79	4.05
K ₂ O	4.37	4.84	3.94	4.22
TiO ₂	0.199	0.059	0.236	0.178
P ₂ O ₅	0.07	0.06	0.05	0.05
LOI	0.72	0.61	0.55	0.58
Total	99.4	99.64	99.13	98.6
Trace element analyses† by ICP-OES				
Sc	2	2	3	2
Be	3	6	3	4
V	13	10	17	14
Ba	244	73	343	378
Sr	81	20	94	108
Y	27	5	18	16
Zr	191	58	173	99
Trace element analyses† by ICP-MS				
Cr	250	BDL	310	20
Co	2	BDL	2	2
Ni	BDL	BDL	BDL	BDL
Cu	10	BDL	20	BDL
Zn	40	BDL	40	BDL
Ga	16	17	16	17
Ge	1	3	2	2
Rb	151	235	188	167
Nb	14	14	20	13
Mo	4	BDL	5	BDL
Ag	0.8	BDL	0.6	BDL
Sn	2	BDL	2	2
Sb	0.6	BDL	0.8	0.6
Cs	1.2	2.3	2.6	1.6
La	34.9	3.2	42.6	33.5
Ce	68.6	4.2	70.5	59.3
Pr	7.86	0.37	6.85	5.24
Nd	26.6	1.1	21.2	16.1
Sm	5.2	0.2	3.7	2.9
Eu	0.56	0.06	0.5	0.43
Gd	4.3	0.2	3.0	2.2
Tb	0.7	BDL	0.5	0.4
Dy	4.7	0.2	3.0	2.5
Ho	0.9	BDL	0.6	0.5
Er	2.8	0.2	2.1	1.6
Tm	0.45	0.06	0.35	0.26
Yb	3.0	0.6	2.6	2.0
Lu	0.48	0.13	0.46	0.34
Hf	5.1	2.5	5.4	3.2
Ta	1.7	1.7	2.3	1.7
W	3	5	2	2
Tl	0.6	0.9	0.7	0.6
Pb	13	12	17	13
Th	11.5	14.8	30.1	19.3
U	3.2	14.9	9.2	3.0

Note: LOI = loss on ignition; BDL = below detection limit

*Major element data in wt%

†Trace element data in ppm

TABLE 3. SR AND ND ISOTOPIC DATA FOR THE ALABAMA HILLS GRANITE

Sample	Rb ppm	Sr ppm	$^{87}\text{Rb}/^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}$	$\pm 2\sigma$ abs*	$^{87}\text{Sr}/^{86}\text{Sr}_i$	Sm ppm	Nd ppm	$^{147}\text{Sm}/^{144}\text{Nd}$	$^{143}\text{Nd}/^{144}\text{Nd}$	$\pm 2\sigma$ abs*	$^{143}\text{Nd}/^{144}\text{Nd}_i$	ϵNd_i
RP15-01	151	81	5.397	0.713639	0.000010	0.705816	5.2	26.6	0.118	0.512447	0.000004	0.512368	-2.71
RP15-02	235	20	34.146	0.752518	0.000012	0.703025	0.2	1.1	0.110	0.512385	0.000006	0.512385	-2.38
RP15-03	188	94	5.790	0.713559	0.000010	0.705167	3.7	21.2	0.106	0.512354	0.000005	0.512354	-2.99
RP15-04	167	108	4.476	0.711725	0.000010	0.705238	2.9	16.1	0.109	0.512423	0.000006	0.512350	-3.05

Note: All isotopic data corrected for decay since 102 Ma.

*Analytical uncertainty only. See text for reproducibility of standards

TABLE 4. U-PB ZIRCON ISOTOPIC DATA FOR ROCKS FROM THE ALABAMA HILLS

Fraction	Composition			Isotopic Ratios						Ages (Ma)						Corr. coef.		
	U (ppm)	Pb (pg) ^a	Th/U ^b	Total common Pb (pg) ^c	²⁰⁶ Pb/ ²⁰⁴ Pb ^d	²⁰⁶ Pb/ ²³⁸ U ^e	Error (2σ%)	²⁰⁷ Pb/ ²³⁵ U ^e	Error (2σ%)	²⁰⁷ Pb/ ²⁰⁶ Pb ^e	Error (2σ%)	²⁰⁶ Pb/ ²³⁸ U ^g	Error (2σ abs.)	²⁰⁷ Pb/ ²³⁵ U ^g	Error (2σ abs.)		²⁰⁷ Pb/ ²⁰⁶ Pb ^g	Error (2σ abs.)
AH14-01, mafic Independence dike																		
F-1	123	24	0.99	1.3	1044	0.0238075	0.15	0.161912	1.24	0.049325	1.14	151.75	0.23	152.38	1.75	163.3	26.6	0.690
F-5	172	18	0.84	2.7	382	0.0236732	0.18	0.154235	3.11	0.047253	2.97	150.91	0.27	145.64	4.22	62.0	70.7	0.806
F-6	170	14	0.97	1.8	451	0.0236811	0.19	0.161554	2.45	0.049478	2.32	150.96	0.28	152.06	3.46	170.6	54.1	0.734
F-7	152	13	1.09	1.0	678	0.0236991	0.19	0.160700	1.80	0.049179	1.69	151.07	0.29	151.32	2.53	156.4	39.5	0.597
F-8	370	20	1.23	1.8	582	0.0237450	0.17	0.157259	2.85	0.048033	2.77	151.35	0.25	148.30	3.93	100.9	65.6	0.459
AH14-03, Alabama Hills Granite																		
F-1	193	36	1.34	1.1	1687	0.0160233	0.12	0.106515	0.89	0.048212	0.84	102.55	0.13	102.77	0.87	109.7	19.7	0.518
F-2	414	26	0.74	1.0	1479	0.0159911	0.12	0.105264	0.86	0.047742	0.80	102.36	0.12	101.62	0.84	86.5	18.9	0.612
F-3	440	29	0.98	1.1	1525	0.0160523	0.14	0.106896	0.88	0.048297	0.78	102.74	0.14	103.12	0.87	113.8	18.3	0.799
F-4	409	20	0.96	0.9	1231	0.0160816	0.19	0.106887	1.36	0.048205	1.17	102.93	0.20	103.11	1.33	109.4	27.6	0.972
F-5	174	35	1.08	1.9	985	0.0160533	0.12	0.106837	1.18	0.048268	1.11	102.75	0.12	103.07	1.16	112.4	26.1	0.626
F-7	313	23	0.96	2.5	507	0.0160652	0.16	0.105207	2.28	0.047496	2.17	102.82	0.16	101.57	2.20	74.2	51.7	0.686
F-8	247	18	0.84	2.9	366	0.0160708	0.20	0.106798	3.11	0.048198	2.97	102.86	0.20	103.03	3.05	109.0	70.2	0.703
RP15-01, Alabama Hills Granite																		
F-1	355	43	0.77	2.1	1171	0.0159218	0.67	0.103772	2.14	0.047291	1.66	101.91	0.68	100.25	2.04	62.9	39.5	0.796
F-2	337	15	0.84	1.4	589	0.0159624	0.23	0.103743	2.56	0.047158	2.37	102.17	0.23	100.23	2.44	56.2	56.6	0.819
F-3	383	44	0.72	2.8	926	0.0155474	0.32	0.102192	2.05	0.047692	1.79	99.54	0.32	98.80	1.93	83.0	42.5	0.827
F-5	412	36	0.79	2.6	802	0.0159322	0.15	0.104697	1.55	0.047681	1.44	101.98	0.16	101.10	1.49	82.4	34.1	0.764
RP15-03, Alabama Hills Granite																		
F-1	561	21	0.67	2.1	588	0.0159714	0.14	0.105472	1.89	0.047895	1.80	102.23	0.14	101.82	1.83	94.1	42.7	0.659
F-5	464	24	0.76	0.8	1730	0.0159911	0.11	0.106029	0.79	0.048089	0.73	102.35	0.11	102.33	0.77	103.7	17.1	0.620
F-6	363	49	0.65	1.2	2399	0.0159794	0.10	0.105842	0.60	0.048039	0.54	102.28	0.10	102.15	0.58	101.2	12.8	0.625
F-7	344	22	0.87	1.0	1306	0.0159978	0.14	0.106391	1.00	0.048233	0.93	102.39	0.15	102.66	0.98	110.7	21.9	0.575
F-8	244	10	0.79	1.1	551	0.0160252	0.18	0.105673	2.11	0.047826	2.01	102.57	0.18	102.00	2.05	90.7	47.7	0.615
F-9	378	16	0.78	1.4	666	0.0160600	0.18	0.106256	1.98	0.047985	1.85	102.79	0.19	102.53	1.93	98.5	43.8	0.742
F-10	197	18	0.89	1.3	781	0.0160168	0.21	0.105731	1.86	0.047877	1.70	102.51	0.21	102.05	1.81	93.2	40.2	0.809
RP15-04, Alabama Hills Granite																		
F-1	200	33	0.79	1.2	1533	0.0160048	0.11	0.106477	0.82	0.048251	0.75	102.45	0.12	102.74	0.80	111.6	17.8	0.653
F-2	331	26	0.85	1.0	1404	0.0159834	0.13	0.106109	0.87	0.048148	0.81	102.31	0.14	102.40	0.85	106.6	19.2	0.495
F-3	445	26	0.72	1.8	858	0.0160344	0.12	0.106412	1.39	0.048132	1.30	102.64	0.12	102.68	1.35	105.8	30.8	0.686
F-4	357	33	0.65	2.1	942	0.0159657	0.23	0.105836	1.69	0.048078	1.59	102.20	0.24	102.15	1.65	103.1	37.6	0.506

Note: All analyses performed on chemically abraded single zircon grains

^aTotal mass of radiogenic Pb.

^bTh contents calculated from radiogenic ²⁰⁶Pb and ²³⁰Th-corrected ²⁰⁶Pb/²³⁸U date of the sample, assuming concordance between U-Pb and Th-Pb systems.

^cTotal mass of common Pb.

^dMeasured ratio corrected for fractionation and spike contribution only.

^eMeasured ratios corrected for fractionation, tracer and blank. Pb blank ratios: ²⁰⁶Pb/²⁰⁴Pb = 18.687 ± 0.25; ²⁰⁷Pb/²⁰⁴Pb = 15.658 ± 0.25; and ²⁰⁸Pb/²⁰⁴Pb = 38.258 ± 0.5 (1σ abs.)

^fCorrected for initial Th/U disequilibrium using radiogenic ²⁰⁸Pb and Th/U_[magma], which is assumed to be approximated by the measured whole rock Th/U ratio.

^gIsotopic dates calculated using λ₂₃₈ = 1.55125E⁻¹⁰ (Jaffey et al. 1971) and λ₂₃₅ = 9.8485E⁻¹⁰ (Jaffey et al. 1971).

CHAPTER 4: TRACKING THE CENTRAL COLORADO MINERAL BELT THROUGH TIME: NEW ISOTOPIC DATA FOR THE GRIZZLY PEAK CALDERA AND TWIN LAKES PLUTON WITH IMPLICATIONS FOR PLUTON-VOLCANO CONNECTIONS AND MAGMA SOURCES

INTRODUCTION

The Colorado Mineral Belt is an important Cenozoic magmatic feature that extends from the Colorado Front Range to the Four Corners region (Fig. 1; Tweto and Sims, 1963; Chapin, 2012). Many of the region's largest economic metal deposits are located in the mineral belt (Wilson and Sims, 2003). Persistent magmatism in the belt coupled with its economic prominence has led to much study regarding the nature and style of magmatism there, particularly to understand the sources and mechanisms that concentrated deposits of Mo, Pb, Zn, Cu, Au, and Ag (e.g., Simmons and Hedge, 1978; Farmer and DePaolo, 1984; Stein and Hannah, 1985; Mutschler et al., 1987; Bookstrom et al., 1988; Stein and Crock, 1990; Shannon et al., 2004).

Past studies have focused on elucidating the histories of individual magmatic centers (e.g., Shannon, 1988; Mills and Coleman, 2013) or tracking regional patterns that vary in space and time (e.g., Stein, 1985; Stein and Crock, 1990). Understanding the spatiotemporal patterns of magmatism is important because the mineral belt developed through different tectonic regimes, including compression during the Laramide orogeny and extension during Rio Grande rifting (Chapin, 2012). Moreover, significant F-rich, high-grade porphyry Mo deposits ("Climax-type") at Urad-Henderson, Climax, and Mount Emmons, were emplaced in the northern, central, and

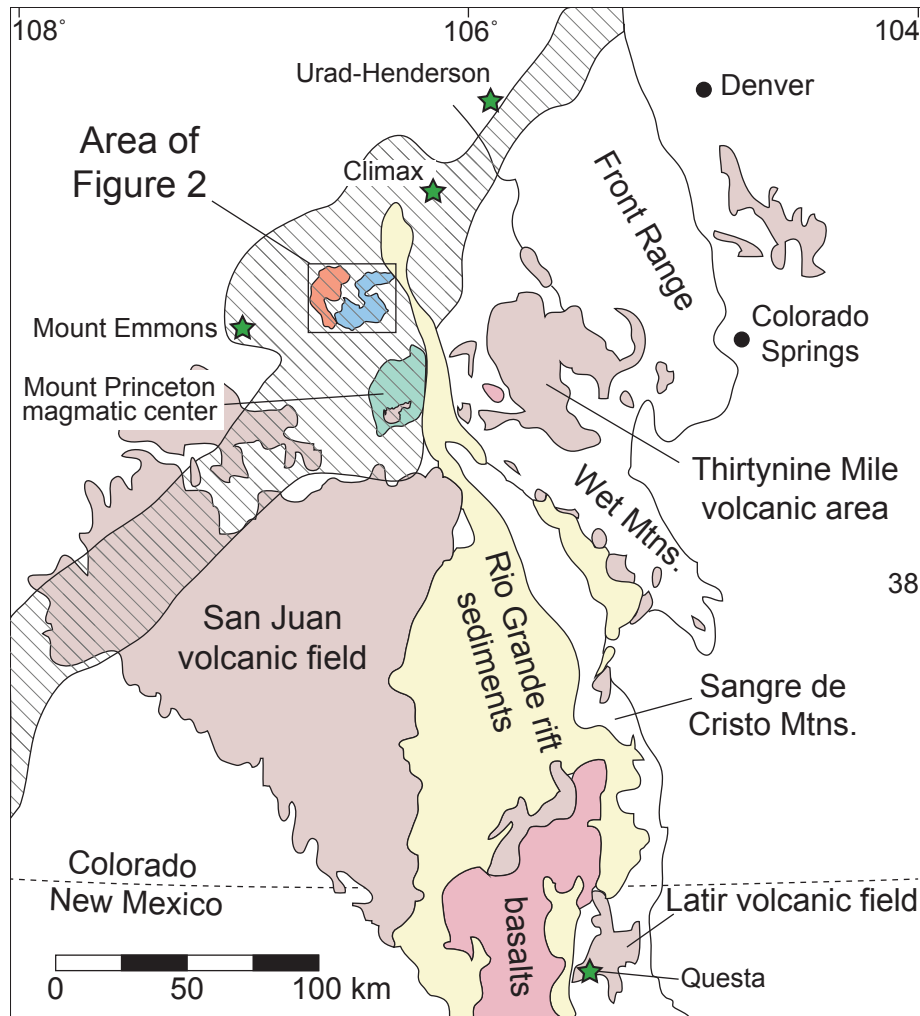


Figure 1. Generalized geologic map of central Colorado and northern New Mexico showing geologic features and locations discussed in the text. Grizzly Peak caldera (orange) and Twin Lakes pluton (blue) are noted by box. Locations of important fluorine-rich porphyry Mo deposits indicated by green stars; note the Grizzly Peak caldera is nearly equidistant between the Climax and Mount Emmons deposits. Location of Colorado Mineral Belt indicated by hatched lines, after Tweto and Sims (1963). Map after Lipman (2000) and McIntosh and Chapin (2004).

southern parts of the mineral belt, with differing isotopic compositions suggesting that lower crustal sources vary laterally (Stein, 1985; Stein and Crock, 1990).

This chapter aims to contribute to the understanding of temporal patterns in the Colorado Mineral Belt through study of the central Colorado Mineral Belt in the Sawatch Range (Fig. 1). This part of the mineral belt is well-suited to tracking secular changes in magma sources and styles because magmatism there occurred over at least 30 Ma, and includes both plutonic and volcanic episodes. This includes the Twin Lakes pluton (Fig. 2), a large granodiorite whose

63-39 Ma emplacement mirrors the Laramide history of the mineral belt (Feldman et al., 2012). Adjacent to the Twin Lakes pluton is the Grizzly Peak magmatic center, which comprises the well-characterized 300-600 km³ Grizzly Peak Tuff, plus resurgent plutons that intruded the intracaldera tuff soon after its eruption ca. 34.3 Ma (Fig. 2; Fridrich et al., 1991; McIntosh and Chapin, 2004). Through new geochemical, isotopic, and geochronological analyses, this chapter investigates the postulated links between the Grizzly Peak Tuff and its resurgent plutons, as well as the Grizzly Peak magmatic center's relationship to the adjacent Twin Lakes pluton and the Colorado Mineral Belt as a whole. The data also permit speculation about the lack of Mo deposits in the Grizzly Peak caldera and the ultimate source of Mo in the lower crust and/or mantle.

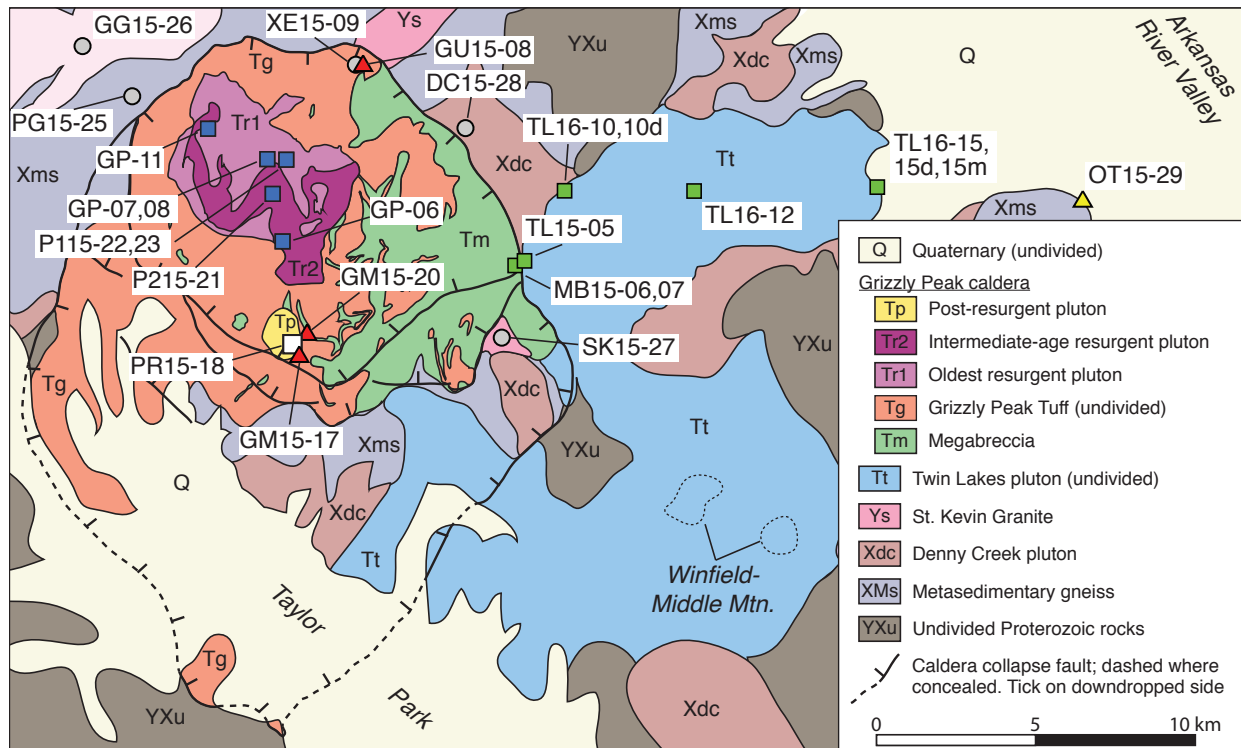


Figure 2. Generalized bedrock geologic map of the Grizzly Peak caldera and Twin Lakes pluton, northern Sawatch Range, Colorado. Locations of samples analyzed in this study indicated by symbols used in later geochemical figures. Outcrop area of tuff in Arkansas River Valley (sample OT15-29) is smaller than the symbol used. Locations of Mo-rich deposits at Winfield-Middle Mountain (Ranta, 1974; Stein and Crock, 1990) noted by dotted lines. Map after Fridrich et al. (1998).

GEOLOGICAL BACKGROUND

Proterozoic wallrocks

Proterozoic units dominate the wallrocks to the Twin Lakes pluton and Grizzly Peak caldera. Major units include metasedimentary gneisses that may be as young as 1701 Ma (Jones and Thrane, 2012), the deformed 1672-1656 Ma Denny Creek Granodiorite (Bickford and Boardman, 1984; Baker and Jones, 2008), and Mesoproterozoic plutons including the Grottos pluton and the ~1.4 Ga St. Kevin Granite (Nyman et al., 1994; Fridrich et al., 1998). Isotopic data for these rocks are limited. Feldspar Pb isotopic compositions for the Denny Creek and St. Kevin plutons are similar to each other (Aleinikoff et al., 1993). Present-day whole rock $^{87}\text{Sr}/^{86}\text{Sr}$ ratios for the St. Kevin Granite 40 km northeast of Grizzly Peak caldera range from 0.754 to 0.869 (Pearson et al., 1966).

Twin Lakes Granodiorite

The Twin Lakes pluton is an equigranular to porphyritic granodiorite with alkali feldspar megacrysts up to 15 cm long. The main body is gradational into leucocratic and mafic facies that dominate the southwestern portion of the pluton, including parts within the mapped extent of the Grizzly Peak caldera (Fridrich et al., 1998). The pluton is cut by numerous aplite dikes and contains modally layered zones interpreted to have resulted from shear segregation processes (Wilshire, 1969; Fridrich et al., 1998). The southeastern part, near Middle Mountain, comprises ~38 Ma rhyolites and granites that have been prospected for molybdenum as part of the Winfield mining district (Fridrich et al., 1998). A zircon U-Pb laser ablation-inductively coupled plasma-mass spectrometry (LA-ICP-MS) survey across the pluton indicates it was constructed from 63-39 Ma, with a gap from 57-43 Ma (Feldman, 2010). The gap mirrors the timing of a lull

in Laramide plutonism throughout the Colorado Mineral Belt (Feldman et al., 2012). Thermo-chronology by $^{40}\text{Ar}/^{39}\text{Ar}$ in K-feldspar suggests possible younger unexposed magmatism from 37-32 Ma (Feldman, 2010). Biotite $^{40}\text{Ar}/^{39}\text{Ar}$ ages are largely concordant with U-Pb zircon ages acquired from the same samples, except for the area of Middle Mountain where biotite ages are several Ma younger than zircon ages (Feldman, 2010). Initial Sr and Nd isotopic compositions include $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.7062\text{-}0.7072$ and $\epsilon\text{Nd} = -5.4$ to -10.1 (Farmer and DePaolo, 1984; Stein and Crock, 1990). A sample of mineralized granite at Winfield yielded $\epsilon\text{Nd}_{38\text{Ma}} = -8.0$ and $^{87}\text{Sr}/^{86}\text{Sr}_{38\text{Ma}} = 0.7073$ (Stein and Crock, 1990).

Grizzly Peak caldera

The Grizzly Peak Tuff is largely preserved as intracaldera fill in the 600 km³ Grizzly Peak caldera. Fridrich et al. (1991) estimated the pre-erosion volume of the intracaldera tuff to be 300 km³. Outflow tuff may have also had an original volume of 300 km³, but only two small remnants interpreted as outflow tuff are preserved (Fridrich et al., 1991). One remnant is located 20 km east in the Arkansas River Valley (Fig. 2; Fridrich et al., 1998). Pumice fiamme define a compositional range from 72-77 wt% SiO₂, grading from high-silica rhyolite at its stratigraphic base to low-silica rhyolite at the highest exposures (Fridrich and Mahood, 1987). Two small horizons in the Grizzly Peak Tuff contain more diverse pumice fiamme of seven distinct compositional groups (Fridrich and Mahood, 1987). Pumice groups 1-3 are equivalent to the rhyolitic pumice present throughout the intracaldera tuff. More mafic pumice compositions in the horizons range down to 57 wt% SiO₂ in group 7. The non-rhyolitic pumices comprise 5-10% of these horizons (Fridrich and Mahood, 1987).

Several plutons are interpreted as resurgent into the caldera. The Lincoln Gulch stock, in

the northern part of the caldera, includes three of the largest resurgent plutons, which are generally zoned from felsic granodiorite at their margins to mafic granodiorite at their cores (Fridrich and Mahood, 1984; Fridrich et al., 1998). Dike-like intrusions similar to the Lincoln Gulch stock are also preserved along the southeastern caldera ring fault (Fridrich et al., 1998). Following the main phase of resurgent intrusions, a series of small “post-resurgent” intrusions were emplaced in the central part of the caldera (Fridrich et al., 1991). Some of the dikes entrained large (up to 5 m diameter) boulders of pink, medium-grained granite interpreted to be a crystallized portion of the Grizzly Peak magma chamber carried from depth during resurgence (Fridrich et al., 1991).

Published isotopic analyses are limited to the Grizzly Peak Tuff. Johnson and Fridrich (1990) measured O, Sr, Nd, and Pb isotopic compositions for the compositionally diverse pumice fiamme groups. Oxygen, Sr, and Pb isotopic data are complicated and do not define simple trends from one fiamme group to another. Overall, $^{87}\text{Sr}/^{86}\text{Sr}_i$ in whole rock samples ranges from 0.7099 to 0.7111. Lead isotopic data are not corrected for decay but still fall in the field of Pb isotopic compositions defined by other central Colorado Mineral Belt rocks (Stein, 1985). Neodymium isotopic data change gradationally from group to group, but do not vary systematically with composition. The most mafic fiamme have $\epsilon\text{Nd}_i = -13.0$, whereas the most radiogenic fiamme are in group 3 (low-silica rhyolite), at $\epsilon\text{Nd}_i = -11.3$. The most felsic fiamme, in group 1, have $\epsilon\text{Nd}_i = -12.1$. The chemical and isotopic variability in the tuff has been attributed to primitive basalt that underwent fractional crystallization and assimilation of 20-40% Proterozoic crust (Johnson and Fridrich, 1990). Recent isotopic work in the K/Ca system shows that the Grizzly Peak Tuff possesses $^{40}\text{Ca}/^{44}\text{Ca}$ ratios higher than mantle values (Mills et al., 2013). Excess ^{40}Ca occurs in

reservoirs with high K/Ca ratios over time, such as Proterozoic continental crust. Therefore, Mills et al. (2013) suggested the Grizzly Peak Tuff has 40-70% ancient crustal component.

Sanidine $^{40}\text{Ar}/^{39}\text{Ar}$ geochronology indicates the Grizzly Peak Tuff erupted at 34.31 ± 0.12 Ma (2σ ; McIntosh and Chapin, 2004). Biotite K-Ar ages on the resurgent and post-resurgent plutons ranges from 36.8 ± 2.8 Ma to 32.9 ± 2.2 Ma (Kennecott, unpub. data recalculated by Fridrich et al., 1991; Fridrich et al., 1991). Biotite K-Ar data for a deposit interpreted as outflow Grizzly Peak Tuff in Arkansas River Valley yielded an age of 33.3 ± 2.0 Ma (2σ ; Fridrich et al., 1991).

Colorado Mineral Belt

The Colorado Mineral Belt is an enigmatic NE-SW feature comprising plutonic and volcanic rocks emplaced throughout its 500-km length from Laramide time (75-43 Ma) through the Miocene (Tweto and Sims, 1963; Chapin, 2012). Although there are different definitions for the extent of the mineral belt, the Twin Lakes pluton and Grizzly Peak caldera are centrally located in all of them (Tweto and Sims, 1963; Mutschler et al., 1987; Stein and Crock, 1990). The central part of the mineral belt also includes the 36-30 Ma Mount Princeton magmatic center to the southeast (Zimmerer and McIntosh, 2012; Mills and Coleman, 2013) and Mount Emmons and the Elk Mountains to the west (Fig. 1).

There are numerous hypotheses for the origins of the Colorado Mineral Belt that attempt to account for its economic importance, its orientation cutting across the “geologic grain” of Colorado (Tweto and Sims, 1963), and its unique spatiotemporal characteristics. The belt represents the only magmatism within an otherwise 1200-km-wide gap in Cordilleran magmatism during the Laramide (75-43 Ma; Chapin, 2012). Some workers have posited that mineral belt magma-

tism was primarily controlled by the presence of Precambrian shear zones and other lithospheric discontinuities that are subparallel to the belt (Tweto and Sims, 1963; Karlstrom and Humphreys, 1998; Wilson and Sims, 2003). Others have suggested the mineral belt is the result of a “leaky segment boundary” that may have been exploited as the exceptionally thick Wyoming craton and North American continental interior created an impediment to a flatly-subducting Farallon plate (Barker and Stein, 1990; Chapin, 2012). Isotopic work on the Colorado Mineral Belt has shown that magmatism related to Climax-type molybdenum mineralization is characterized by less radiogenic Nd ($\epsilon\text{Nd}_i = -8$ to -14) than other mineral belt rocks ($\epsilon\text{Nd}_i = -1$ to -10.6). Mineralizing rocks also have $^{87}\text{Sr}/^{86}\text{Sr}_i$ from 0.7095 to 0.740, whereas non-Climax type mineral belt rocks range from 0.705 to 0.710 (Stein, 1985; Stein and Crock, 1990). In contrast, Pb isotopic characteristics for mineralizing rocks are not distinct. Lead isotopic compositions vary regionally in all belt rocks, regardless of mineralization. Stein (1985) first demonstrated the northern mineral belt has the highest $^{208}\text{Pb}/^{204}\text{Pb}$ ratios, which progressively decrease in the central and southern mineral belt. Subsequent Pb isotopic work has largely corroborated the spatial pattern (Johnson and Fridrich, 1990; Mills, 2012).

METHODS

Samples selected for geochronological and geochemical analyses (Table 1) were chosen to limit altered zones as much as possible. Outer rinds affected by surficial weathering were removed on the outcrop using a hammer and chisel. A tile saw was used to remove additional weathered zones, if present. All samples were crushed using a jaw crusher. Crushed aliquots of bulk Grizzly Peak Tuff were picked by hand before further processing to avoid incorporation of weathered zones or lithic fragments. Aliquots of crushed samples selected for U-Pb zircon

geochronology were further disaggregated in a steel disc mill. Zircon was isolated using standard density (water table, methylene iodide) and magnetic (Frantz) techniques. Samples analyzed for elemental and isotope geochemistry were pulverized to powder in a SPEX 8530 Shatterbox alumina swing mill.

Whole rock powders of all samples were analyzed at Actlabs (Ontario, Canada). Samples were dissolved by fusion in a lithium metaborate/tetraborate mixture; major elements and Ba, Sr, Y, Zr, Sc, Be and V were analyzed by inductively coupled plasma-optical emission spectroscopy (ICP-OES), with remaining trace elements and REE analyzed by ICP-MS. Uncertainties are dependent on the abundances of the oxides and elements analyzed and for these samples are generally as follows: major elements uncertainties are less than 2% ($\pm 2\sigma$ relative) for all oxides except MgO (3%), MnO (5%), and P₂O₅ (16%). Trace elements analyzed at Actlabs that are reproducible ($\pm 2\sigma$ absolute) at better than 2 ppm include Ag, Be, Cs, Ge, Hf, Sb, Sn, Ta, Th, Tl, U, and W. Other elements are reproducible within 3 ppm (Co, Mo, Nb, Sc), 4 ppm (Ga, Y) and others have higher uncertainties, in parentheses in ppm: Pb (9); Cu (14); Rb and Zr (15); V (19); Ni (23); Sr (25); Ba (31); Zn (42); Cr (49). Rare earth elements reproducible ($\pm 2\sigma$ absolute) at 0.2 ppm or better are Eu, Ho, Lu, Tb, and Tm; REE reproducible at 0.2-0.5 ppm or better are Dy, Er, Gd, Pr, Sm, and Yb; others are reproducible at higher uncertainties, in ppm: Nd (1.2); La (1.5); Ce (2.1).

Whole rock isotopic analyses were prepared by weighing a mass of powder sufficient to provide aliquots of 1 μ g Sr, 200 ng Nd, and 100 ng Pb for each sample. Whole rock powders were digested in HF and HNO₃ in Teflon® (Parr) bombs at 180°C for 48 hours. Following digestion, samples were dried down then immediately fluxed in concentrated HCl on a 180°C hot plate for 16 hours. Samples were subsequently aliquoted for Sr, Nd, and Pb purification by ion exchange column chromatography. Strontium was purified using Sr-spec cation exchange resin

after the methods of Lundblad (1994) and loaded on single Re filaments with TaF₅. Neodymium was purified through a three-stage column chemistry procedure after Harvey and Baxter (2009) and loaded on single Re filaments in a Ta₂O₅-H₃PO₄ slurry. Lead was purified in a two-stage hydrobromic-hydrochloric column procedure modified after Krogh (1973). Strontium isotopic analyses were accomplished on a VG Sector 54 TIMS at the University of North Carolina at Chapel Hill. Strontium was analyzed as a metal in dynamic multicollector mode with ⁸⁸Sr = 3V. Neodymium isotopes were analyzed on an Isotopx Phoenix TIMS at the University of North Carolina at Chapel Hill, and was analyzed as an oxide in dynamic multicollector mode with ¹⁴²Nd¹⁶O = 1V. Strontium and Nd isotope ratios were corrected for mass fractionation using an exponential law, with Sr isotopic ratios normalized to ⁸⁶Sr/⁸⁸Sr = 0.1194 and Nd isotopic ratios normalized to ¹⁴⁶Nd/¹⁴⁴Nd = 0.7219. Lead isotopic analyses were completed on the VG Sector 54. Lead was analyzed as a metal in static multicollector mode with ²⁰⁸Pb = 1V. Lead isotope ratios for most samples were corrected for mass fractionation using a value of 0.165%/amu following replicate analyses of the NBS 981 Pb standard. A batch of samples analyzed at a later date on the same instrument was corrected for mass fractionation using a value of 0.150%/amu (see Table 4 for samples). Replicate analyses of the NBS 987 Sr standard yielded ⁸⁷Sr/⁸⁶Sr = 0.710265 ± 0.000015 (2σ; n = 26). Replicate analyses of the Nd standard JNdi yielded ¹⁴³Nd/¹⁴⁴Nd = 0.512098 ± 0.000012 (2σ; n = 18). All isotopic data for Cenozoic samples are corrected to initial values (i.e., values at the time of crystallization or eruption) using ages indicated in Table 1. Ages used are from this study or are approximated based on the LA-ICP-MS data of Feldman (2010). Isotopic data for Proterozoic samples are corrected for radiogenic ingrowth since 35 Ma to approximate their isotopic compositions at the time of the eruption of the Grizzly Peak Tuff.

Geochronology was performed by chemical abrasion-isotope dilution-thermal ionization

mass spectrometry (CA-ID-TIMS) on either a VG Sector 54 or Isotopx Phoenix-X62 in the Department of Geological Sciences at the University of North Carolina at Chapel Hill following the sample preparation and analytical methods of Frazer et al. (2014). Zircons for each sample were thermally annealed for 48 hours at 900°C, then chemically abraded in HF and HNO₃ acids for 16 hours at 220°C (Mundil et al., 2004; Mattinson, 2005). Individual grains free of cracks, obvious cores, or mineral inclusions were selected under a binocular microscope for dissolution. Dissolution and chemical purification methods for U and Pb were modified after Krogh (1973) and Parrish (1987), used the mixed ²⁰⁵Pb-²³³U-²³⁵U EARTHTIME tracer (Condon et al., 2015). Data processing and age calculations were completed using the applications Tripoli and ET_Redux, developed as part of the EARTHTIME initiative (Bowring et al., 2011; McLean et al., 2011), and include 2σ analytical uncertainties only. Corrections for initial Th/U disequilibrium (Mattinson, 1973; Schmitz and Bowring, 2001) were made in ET_Redux assuming the measured whole rock Th/U value for each sample approximates the magmatic value; this adjustment increased the ages of individual analyses by up to 100 ka. All subsequent zircon U-Pb data discussed in this chapter will be Th-corrected.

RESULTS

Grizzly Peak Tuff and resurgent plutons

Samples of bulk Grizzly Peak Tuff yielded SiO₂ concentrations from 69-73 wt%. Sample GU15-08, from the upper subunit of the Grizzly Peak Tuff near Independence Pass (Fridrich et al., 1998), was the most silicic, whereas bulk tuff samples from the structurally lower middle subunit in the central part of the caldera were slightly less silicic. Resurgent pluton samples ranged from 61-69 wt% SiO₂. One mafic enclave (sample GP-08) collected from the oldest

resurgent pluton has 56 wt% SiO₂. Samples of bulk tuff and resurgent plutons follow broadly linear trends on major element Harker diagrams (Fig. 3), with two exceptions. Sample P115-22 was collected from the oldest resurgent pluton is strongly foliated, unlike other resurgent pluton samples. It is distinguished from other resurgent pluton and tuff samples by having higher K₂O, and lower CaO, TiO₂, and P₂O₅. Sample GP-08, a mafic enclave from the resurgent pluton, falls below the linear trend on plots of Al₂O₃ and CaO (Fig. 3).

Trace element data for the tuff and resurgent plutons also follow broadly linear trends for most elements (Fig. 4). However, resurgent pluton sample P115-22 has higher Rb and lower Zr and Y concentrations than other tuff and resurgent pluton samples, and mafic enclave sample GP-08 has relatively low Sr compared to other resurgent pluton samples. Zirconium and Hf trends are also non-linear, with the Zr and Hf concentrations peaking at 61.5 wt% SiO₂. Samples of Grizzly Peak Tuff and the resurgent plutons have approximately chondritic Zr/Hf ratios (~37; McDonough and Sun, 1995) across the SiO₂ spectrum (Fig. 5). Ratios of Th/U are generally greater than 6 (Fig. 5). Chondrite-normalized REE patterns indicate that all samples are enriched in light REEs, with La_N/Lu_N ratios from 16-36 (subscript “N” denotes concentration normalized to chondritic values of McDonough and Sun, 1995), although Grizzly Peak Tuff samples have slightly lower ratios than similarly siliceous resurgent plutons (Fig. 5). Sample P115-22, which has distinctly lower REE concentrations than other tuff and resurgent pluton samples (Fig. 6). Bulk Grizzly Peak Tuff has negative Eu anomalies (where $Eu/Eu^* = Eu_N / (Sm_N \cdot Gd_N)^{1/2}$) of 0.72 to 0.89, whereas resurgent pluton samples (except P115-22) have Eu/Eu^* of 0.92 to 1.04 (Fig. 5).

Most tuff and resurgent pluton samples have $^{87}Sr/^{86}Sr_i$ between 0.7098 and 0.7108, except sample P115-22 with $^{87}Sr/^{86}Sr_i = 0.732$ (Fig. 7). Initial $^{87}Sr/^{86}Sr$ in the tuff is positively correlated with SiO₂ content, whereas the resurgent plutons have a negative correlation with SiO₂

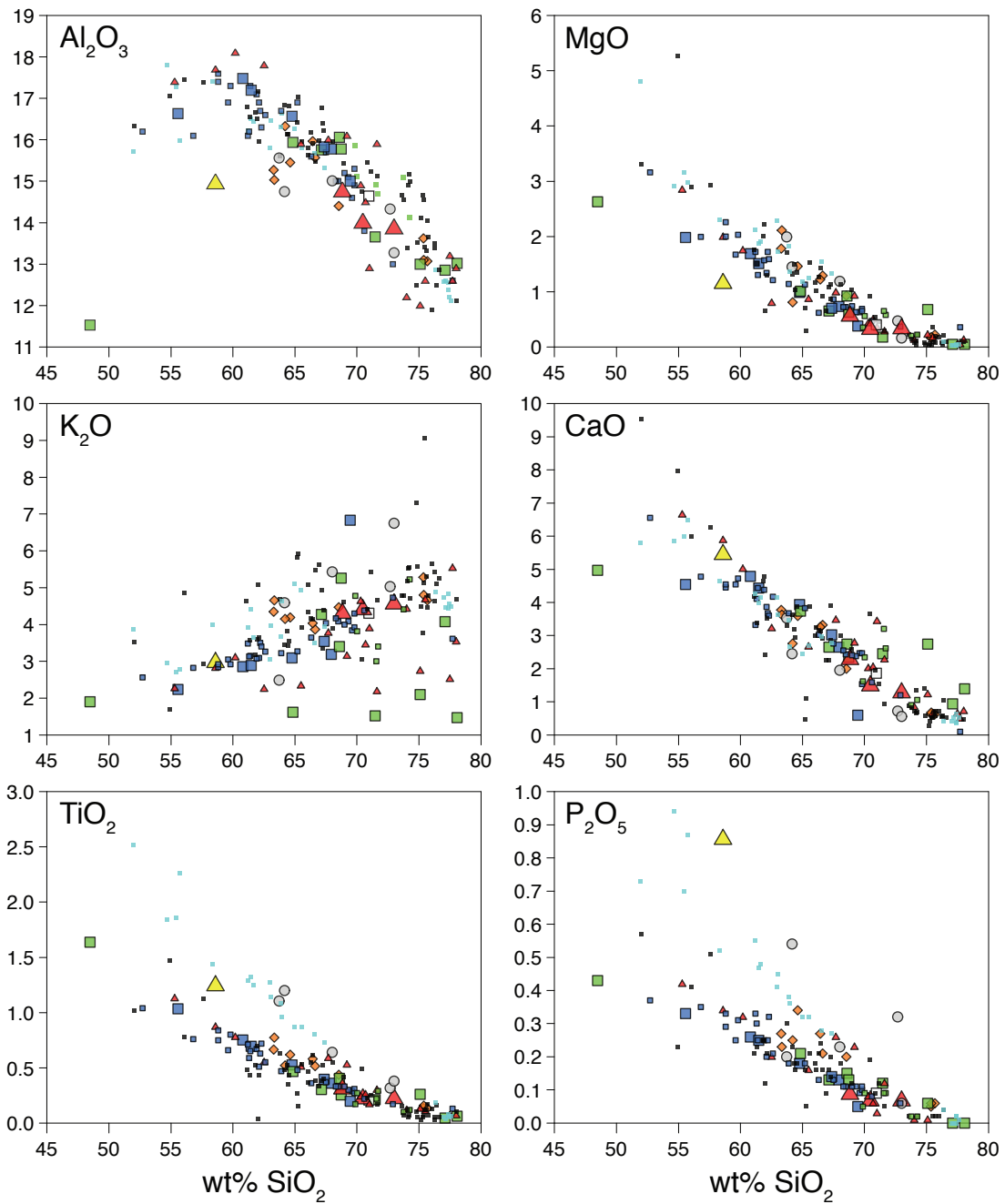


Figure 3. Major element Harker diagrams for samples analyzed in this study (large symbols) and the literature. Concentrations given in wt%. Most data for Cenozoic rocks from this study form linear trends and are similar to literature values for the same units. Exceptions include low- K_2O samples of the Twin Lakes pluton, a modally layered sample from the Twin Lakes pluton with low Al_2O_3 , CaO , and MgO , and the tuff in the Arkansas River Valley with low Al_2O_3 , MgO , and high P_2O_5 . Literature data sources are: Grizzly Peak Tuff and resurgent plutons – Fridrich (1987); Twin Lakes pluton and other Colorado Mineral Belt – Stein (1985); Neversummers – Jacob et al. (2015); Princeton area – Mills and Coleman (2013).

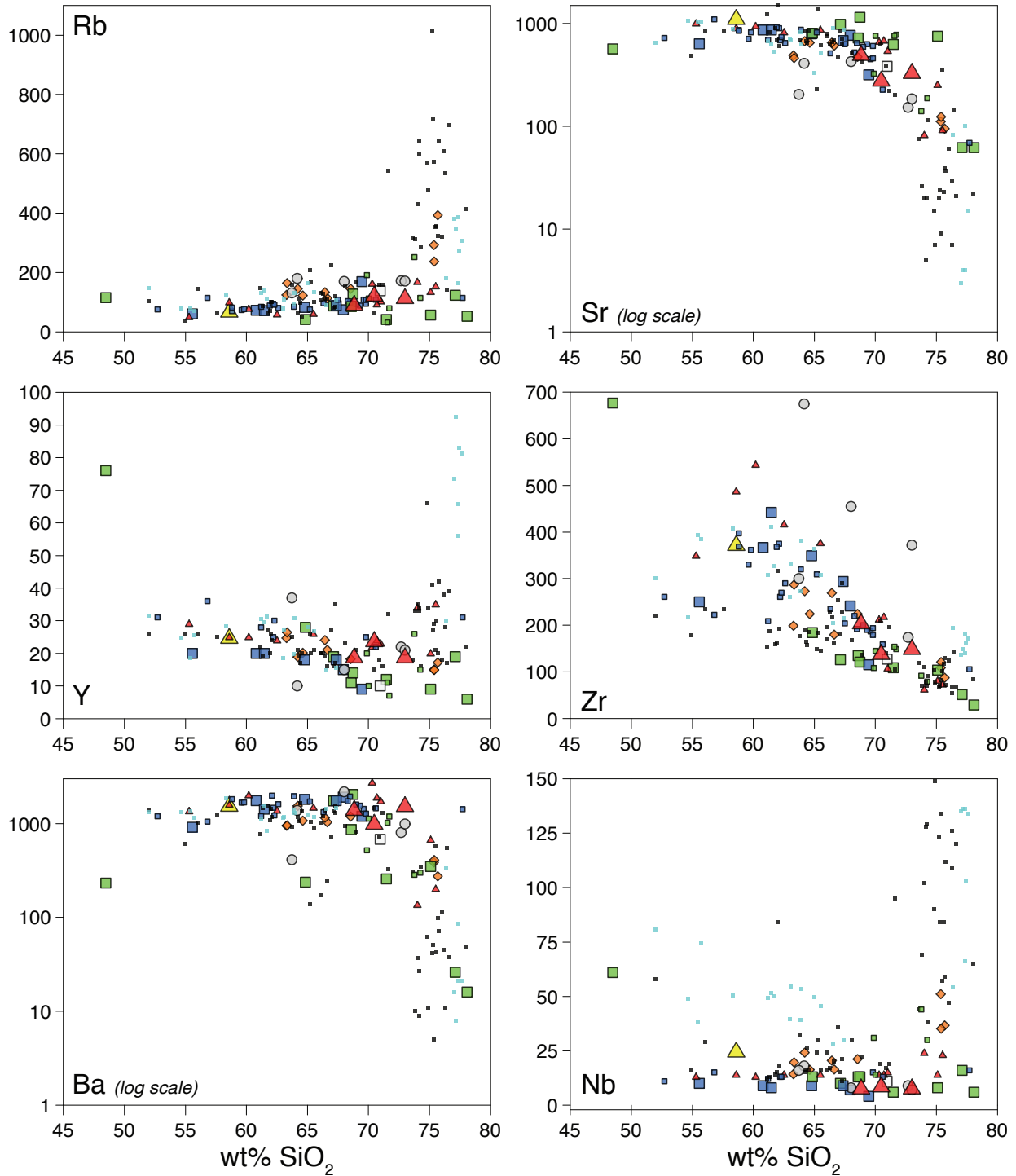


Figure 4. Trace element variation diagrams for samples from this study (large symbols) and other Cenozoic igneous rocks in Colorado. Trace element concentrations in ppm. Note log scale is used for ordinates for Ba and Sr. Symbols as in Figure 3. High-silica samples of Grizzly Peak Tuff (Fridrich, 1987; this study) do not have the extreme trace element enrichments or depletions as similarly-siliceous samples of the Colorado Mineral Belt (Stein, 1985) or Never Summer Mountains (Jacob et al., 2015). Twin Lakes pluton is distinct from Grizzly Peak Tuff with generally lower Rb, Zr, Y, and Ba, and has higher Sr at similar SiO_2 concentrations (except for aplite dikes). Note also the distinct off-trend trace element concentrations of the modally-layered sample of the Twin Lakes pluton, including very high concentrations of Y and Nb.

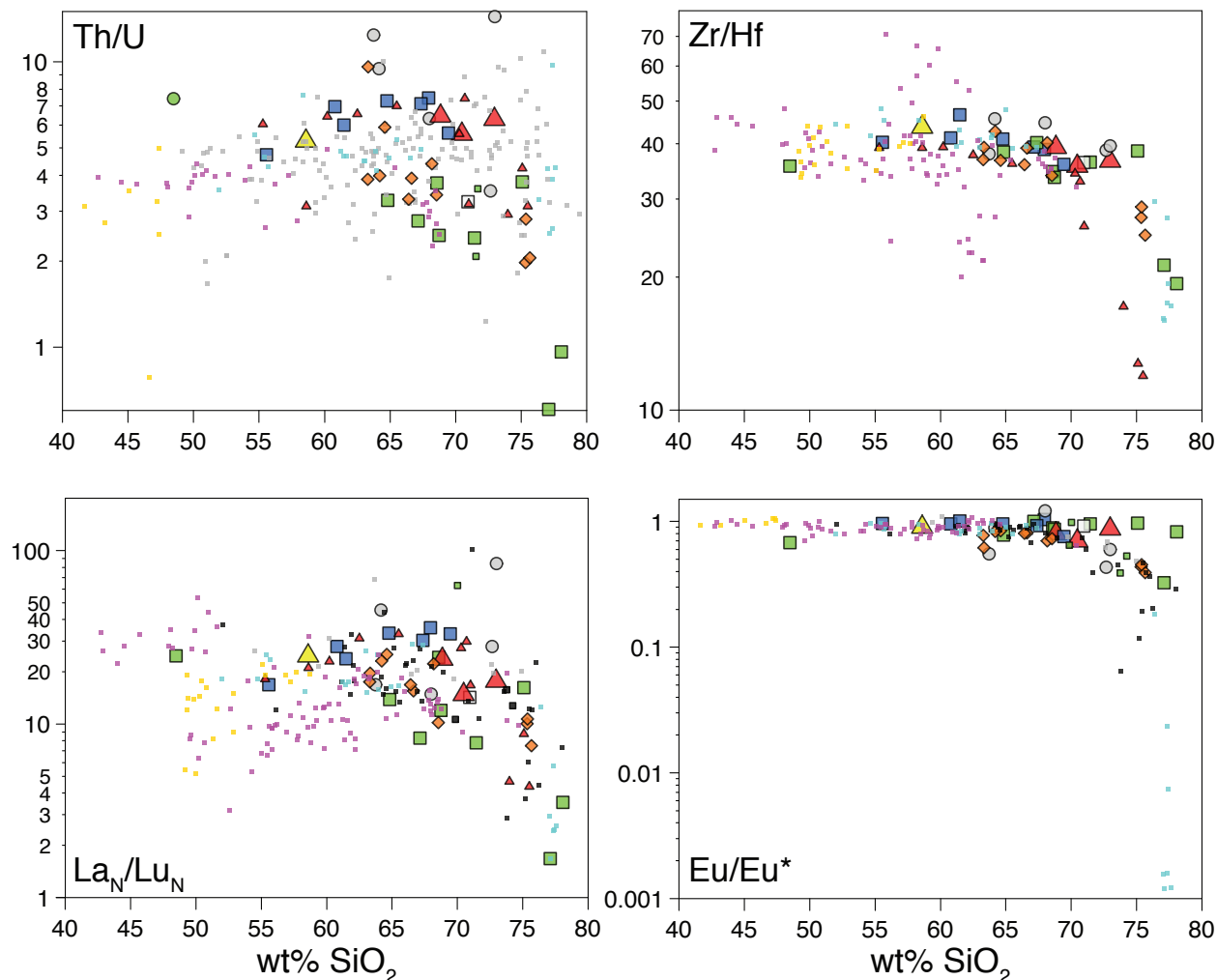


Figure 5. Trace element and REE ratios versus silica concentrations for samples from this study (large symbols) and other Cenozoic igneous rocks in Colorado. All ordinates are log scales. Samples of Grizzly Peak Tuff and resurgent plutons have high Th/U relative rocks from this study and most literature analyses in Colorado. Granitic xenolith (white square) plots with Twin Lakes data rather than Grizzly Peak data in Th/U. High-SiO₂ pumice fiamme analyses of Fridrich (1987) have sub-chondritic Zr/Hf ratios (chondrite Zr/Hf = 37; McDonough and Sun, 1995), whereas bulk tuff at same silica contents is approximately chondritic. Twin Lakes aplite dike samples have low Th/U, Zr/Hf, and La_N/Lu_N. See text for definition of normalized ratios and Eu/Eu* calculation. Symbols as in Figure 3. Additional symbols are purple squares: San Juan volcanic field (Bachmann et al., 2002; Parat et al., 2005; Parker et al., 2005; Lake and Farmer, 2015; Sliwinski et al., 2017); yellow squares: Rio Grande rift (Thompson et al., 1991; Cosca et al., 2014); gray squares: Thirtynine Mile volcanic field (Campbell, 1994).

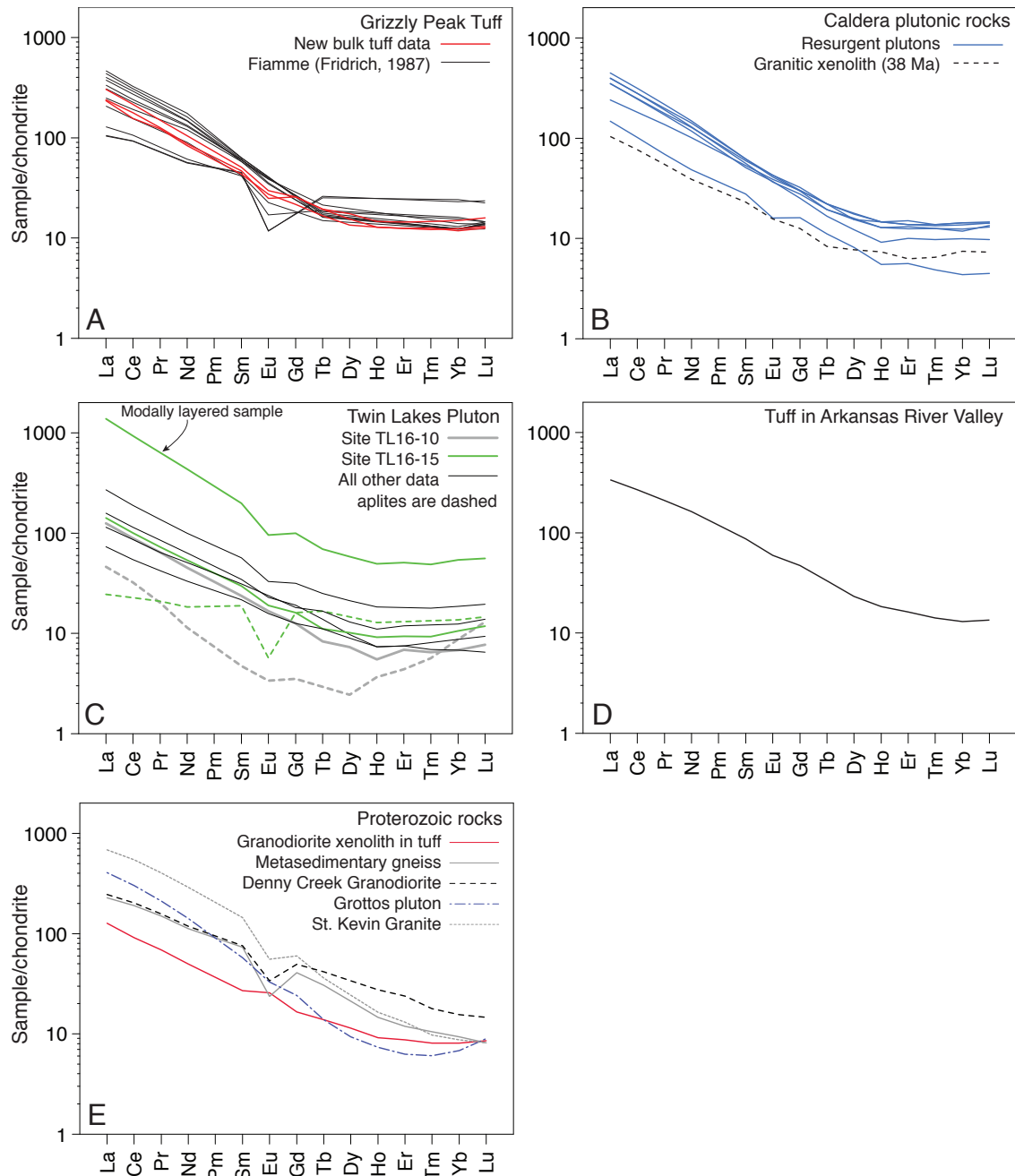


Figure 6. Rare earth element plots with concentrations normalized to chondritic values of McDonough and Sun (1995). (A) Grizzly Peak Tuff data include pumice fiamme of Fridrich (1987). High-silica fiamme have deeper negative Eu anomalies and higher heavy REE concentrations than bulk tuff. (B) Resurgent pluton samples have consistent REE concentrations and patterns. Only the REE-depleted sample P115-22 has an Eu anomaly. The 38 Ma granitic xenolith also has lower REE concentrations and is similar to Twin Lakes samples. (C) Twin Lakes pluton REE data vary widely within the main map unit, and local features (aplites and modal layering) extend the range of REE patterns. We interpret the aplite at site TL16-10 (scoop-shaped pattern) to have been locally derived, whereas the distinct “seagull” type pattern and isotopic compositions of aplite from site TL16-15 suggest that it may have been derived from a source different than its host granodiorite. (D) REE data for the tuff in the Arkansas River Valley are similar to the Grizzly Peak Tuff. However, isotopic and geochronologic data indicate it is a separate unit. (E) Proterozoic REE data vary widely. St. Kevin Granite has high light REE concentrations and a steep pattern, similar to other 1.4 Ga plutons in central Colorado (Stein and Crock, 1990). The granodiorite xenolith in the tuff has the only positive Eu anomaly ($\text{Eu}/\text{Eu}^* = 1.22$) in this study.

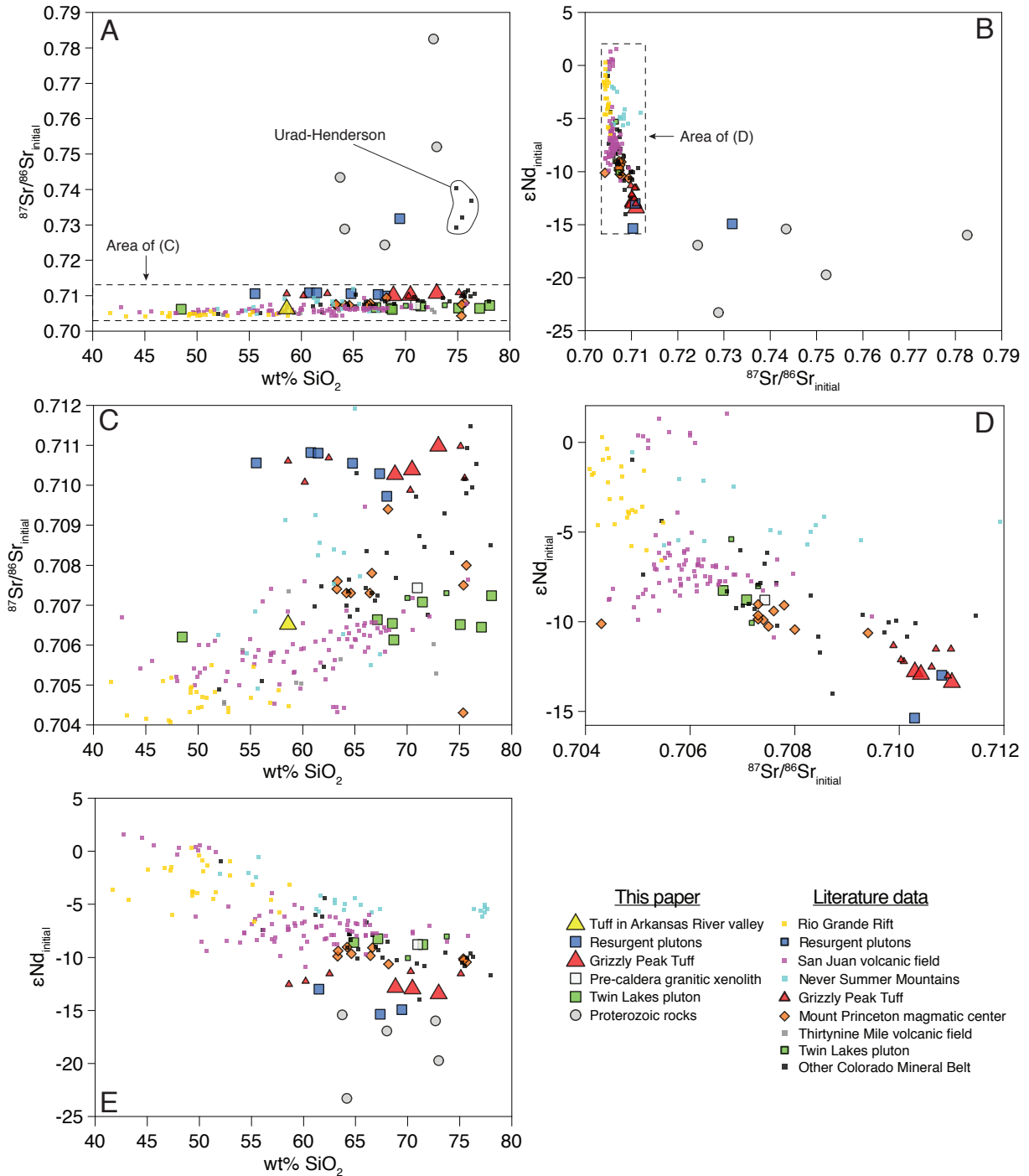


Figure 7. Strontium and Nd isotope data for samples from this study and other Cenozoic igneous rocks in Colorado. Panels (A), (C), and (E) are isotope-SiO₂ diagrams, with (C) showing broad positive correlation of ⁸⁷Sr/⁸⁶Sr_i with SiO₂ across all rocks, whereas (E) demonstrates that εNd_i and silica are negatively correlated. Although such patterns may suggest a mixing origin between primitive mantle and Proterozoic crust (Johnson, 1991) isotopic data for individual magmatic centers (Grizzly Peak, Twin Lakes, Mount Princeton, Never Summer Mountains) do not show correlations (panels C and E). (B) and (D) show that the Grizzly Peak Tuff and resurgent plutons have amongst the most radiogenic Sr and unradiogenic Nd in Cenozoic Colorado rocks, although Sr data for the Henderson deposit are a notable exception (panel A). Data sources as in Figure 3 and Figure 5, with additional data for Grizzly Peak (Johnson and Fridrich, 1990), Rio Grande rift (Johnson and Thompson, 1991), San Juan volcanic field (Ricuputi et al., 1995), and Thirtynine Mile volcanic field (Campbell, 1994).

content. Grizzly Peak Tuff samples have ϵNd_i ranging from -12.7 to -13.3, and the resurgent plutons range -13.0 to -15.5 (Fig. 7). All samples have slightly less radiogenic ϵNd_i at higher SiO_2 concentrations. Lead isotopic compositions for all samples except P115-22 are restricted to $^{206}\text{Pb}/^{204}\text{Pb}_i = 17.93\text{-}18.10$, $^{207}\text{Pb}/^{204}\text{Pb}_i = 15.55\text{-}15.57$, and $^{208}\text{Pb}/^{204}\text{Pb}_i = 38.41\text{-}38.68$ (Fig. 8). Sample P115-22 has more radiogenic $^{206}\text{Pb}/^{204}\text{Pb}_i$ and $^{207}\text{Pb}/^{204}\text{Pb}_i$, and less radiogenic $^{208}\text{Pb}/^{204}\text{Pb}_i$.

One zircon from the middle subunit of the Grizzly Peak Tuff (sample GM15-20) yielded a concordant $^{206}\text{Pb}/^{238}\text{U}$ age of 35.05 ± 0.15 Ma (Fig. 9). In contrast, six of seven zircons dated from three resurgent pluton samples (P215-21, P115-22, P115-23) are reversely discordant. Two zircons from resurgent pluton 1 (sample P115-23) are closest to concordia. A concordant grain yields an age of 34.85 ± 0.10 Ma, and a second grain is discordant and slightly younger. All three zircons from sample P115-21 of resurgent pluton 2 are reversely discordant. The range of their $^{206}\text{Pb}/^{238}\text{U}$ ages and 2 uncertainties are encompassed by one fraction measured at 34.43 ± 0.39 Ma. Sample P115-22 from resurgent pluton 1 yielded the most reversely discordant data in two zircons, which range in age from 34.47-33.90 Ma including 2σ uncertainties (Fig. 9).

Twin Lakes Granodiorite

Samples of typical Twin Lakes pluton range from 65-75 wt% SiO_2 . Two aplites are 77-78 wt% SiO_2 , and a modally layered sample has 48 wt% SiO_2 (Fig. 3). Excluding the modally layered sample, samples plot on linear trends on Harker diagrams for all major elements except K_2O and Na_2O , which are highly scattered. Trace element data are also typically linear, except for scatter in Ba and Rb (Fig. 4). The aplites are depleted in Sr and enriched in U and Pb relative to the granodiorite. The modally layered sample (TL16-15m) has high concentrations of V, Y, Zr, and Th, and low concentrations of Sr and Ba relative to its host granodiorite (TL16-15). Ratios

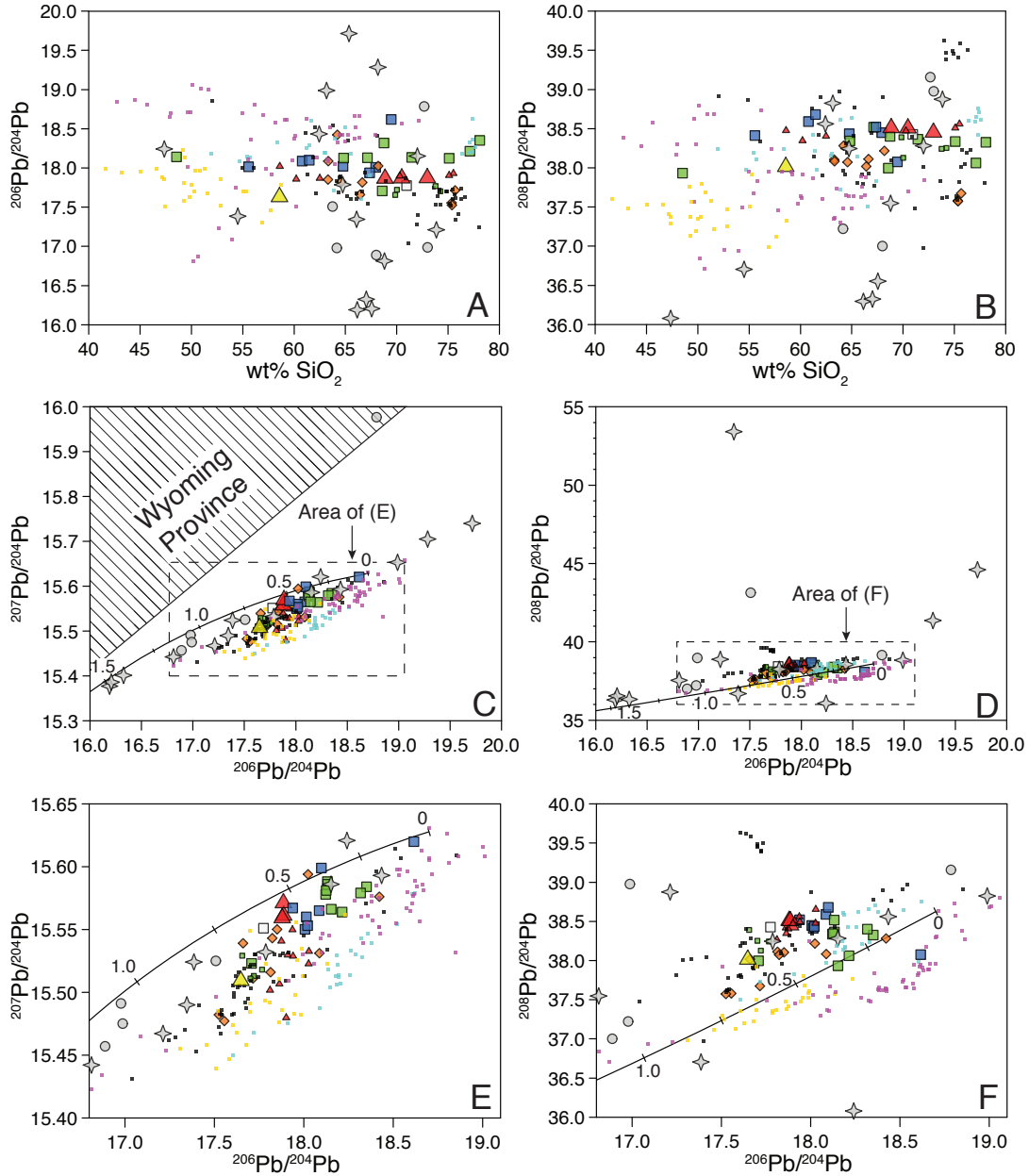


Figure 8. Lead isotope data for samples from this study and other Cenozoic igneous rocks from Colorado. (A) and (B) show there is no correlation of Pb isotope composition and silica. Panel (C) shows all $^{207}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$ data, and panel (D) shows all $^{208}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$ data. Isotopic data for Proterozoic rocks (Stein, 1985 – gray crosses; this study – gray circles) vary widely. High $^{207}\text{Pb}/^{204}\text{Pb}$ sample in panel (C) is metasedimentary gneiss sample PG15-25. Its uraniumogenic Pb composition is similar to Archean rocks of the Wyoming Province (hatched area; Wooden and Mueller, 1988). Panels (E) and (F) focus on Pb isotope data of Cenozoic samples. (E) Twin Lakes samples (green squares) with low $^{207}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$ ratios are from Eocene parts of the pluton, whereas high $^{207}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$ ratios are from older samples. (F) Samples from this study largely plot with $^{208}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$ data of central Colorado Mineral Belt samples of Stein (1985). Solid black lines on Pb-Pb diagrams are Pb isotope growth curves of Stacey and Kramers (1975); curve ages are given in Ga; tick marks are every 0.25 Ga. Data from this study, Mills (2012), Cosca et al. (2014), and Jacob et al. (2015) are whole rock data corrected to initial values. Colorado Mineral Belt data of Stein (1985) are feldspar compositions that are not corrected for decay but are interpreted to represent initial values. Cenozoic whole rock data of Johnson and Fridrich (1990), Johnson and Thompson (1991), Riciputi et al. (1995), Parat et al. (2005), and Lake and Farmer (2015), and Proterozoic whole rock data of Stein (1985), are not corrected for decay.

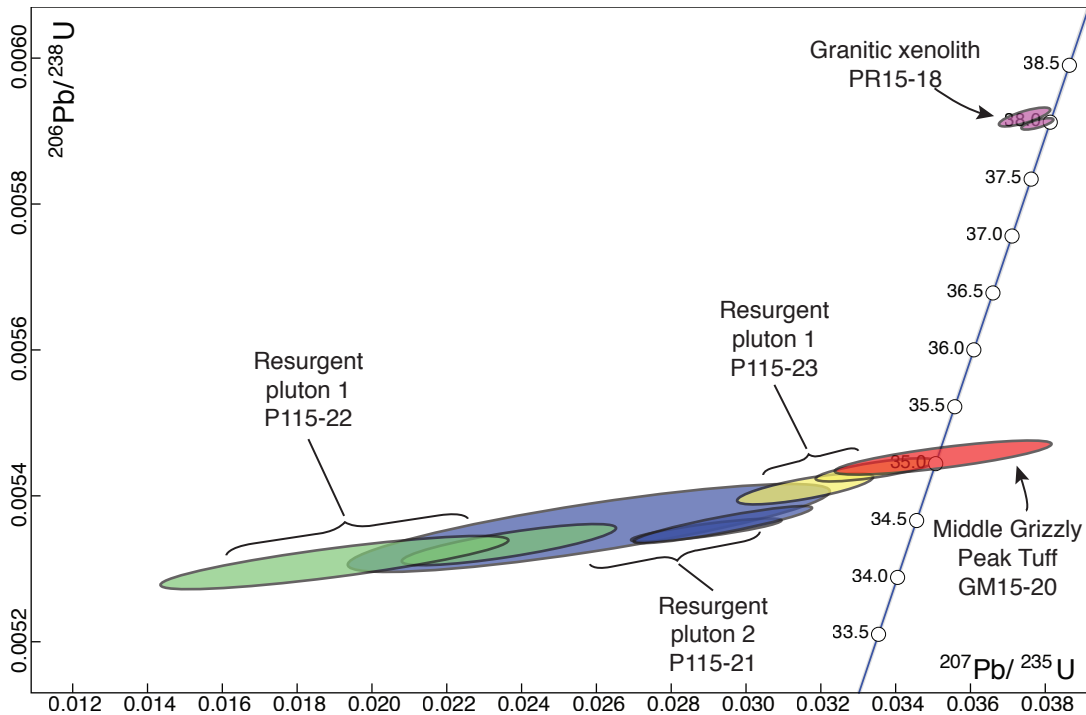


Figure 9. Zircon U-Pb concordia diagram of selected samples dated by CA-ID-TIMS in this study. The oldest analysis of resurgent pluton sample P115-23 is concordant and is indistinguishable within uncertainty of a concordant analysis of the Grizzly Peak Tuff, and suggests the pluton was emplaced very soon after eruption. We interpret the reversely discordant data of the other resurgent pluton samples to suggest they were likely also emplaced shortly after eruption. Data for the mafic tuff in the Arkansas River Valley (sample OT15-29), which includes one concordant and one reversely discordant analysis both near 24.5 Ma, not shown for clarity. One normally discordant zircon analysis from sample PR15-18 with a $^{207}\text{Pb}/^{206}\text{Pb}$ age of 1353 ± 10 Ma is not shown; see Table 5 for data. Ages along concordia are given in Ma.

of Th/U decrease with increasing SiO_2 , with markedly decreased ratios in the aplite dike samples (Fig. 5). Aplites also have sub-chondritic Zr/Hf ratios near 20. Typical Twin Lakes granodiorite samples have La_N/Lu_N between 8-25 and minor to absent negative Eu anomalies (Figs. 5, 6). The aplites follow very different patterns to the main pluton and to each other. Sample TL16-10d has a scoop-shaped pattern, whereas TL16-15d has a flat “seagull”-type pattern characterized by very low La_N/Lu_N and a negative Eu anomaly of 0.33 (Figs. 5, 6). The modally layered sample has elevated REE concentrations relative to its host granodiorite (Fig. 6), and a more pronounced negative Eu anomaly.

Initial $^{87}\text{Sr}/^{86}\text{Sr}$ in the Twin Lakes pluton ranges from 0.7061 to 0.7072. There does not appear to be any correlation with SiO_2 concentrations, although the aplite dikes have slightly

higher $^{87}\text{Sr}/^{86}\text{Sr}_i$ than their host rocks (Fig. 7). Neodymium isotopic data from three samples collected in close proximity to each other near the caldera margin (Fig. 2) show little variation around $\epsilon\text{Nd}_i = -8.6$. Lead isotopic data for all samples except TL15-12 fall in the range $^{206}\text{Pb}/^{204}\text{Pb}_i = 18.13-18.35$, $^{207}\text{Pb}/^{204}\text{Pb}_i = 15.56-15.59$, and $^{208}\text{Pb}/^{204}\text{Pb}_i = 38.06-38.52$. Uranogenic Pb data are distinct from the Grizzly Peak Tuff and resurgent plutons (Fig. 8). Sample TL15-12, collected from a ~42 Ma part of the Twin Lakes pluton (Feldman, 2010), has less radiogenic Pb isotopic compositions than all other Twin Lakes samples (Fig. 8).

Other Cenozoic rocks

Sample PR15-18 was collected from a large granite outcrop in the post-resurgent pluton map unit of Fridrich et al. (1998). It is a porphyritic biotite granite with pink alkali feldspar megacrysts up to 4 cm long. The granite has similar major and trace element abundances as felsic samples of the Twin Lakes pluton and resurgent plutons (Figs. 3, 4). Rare earth element abundances are similar to the Twin Lakes pluton, but lower than most of the resurgent pluton samples. The sample has $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.7074$ and $\epsilon\text{Nd}_i = -8.8$, and initial Pb isotopic compositions that are intermediate between sample TL15-12 and all other Twin Lakes samples (Fig. 8). One zircon yielded a concordant $^{206}\text{Pb}/^{238}\text{U}$ age of 37.985 ± 0.083 Ma, with a second zircon slightly reversely discordant at 38.046 ± 0.050 Ma. A third zircon is normally discordant and has a $^{207}\text{Pb}/^{206}\text{Pb}$ age of 1353 ± 10 Ma. Regression of a chord through the data yields a lower intercept age of 38.32 ± 0.33 Ma and an upper intercept age of 1410 ± 11 Ma.

Sample OT15-29 was collected near Twin Lakes Reservoir in the Arkansas River Valley from an outcrop mapped as an outflow deposit of the upper subunit of the Grizzly Peak Tuff (Fridrich et al., 1998). Pumice extracted from the red-weathered tuff yielded 59 wt% SiO_2 and

major element concentrations distinct from the all other analyzed Cenozoic rocks in this study (Fig. 3). Trace element concentrations are similar to mafic samples from the Grizzly caldera resurgent plutons, although it has slightly elevated Nb (Fig. 4). This tuff has $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.7065$, with Pb isotopic compositions similar to sample TL15-12 (Figs. 7, 8). Geochronologic data include one concordant fraction with a $^{206}\text{Pb}/^{238}\text{U}$ age of 24.46 ± 0.37 Ma and a reversely discordant zircon at 24.34 ± 0.22 Ma (Table 5).

Proterozoic rocks

Samples of four different wallrock units adjacent to the Grizzly Peak caldera, as well as a large (1 m diameter) granodioritic xenolith in the upper subunit of the tuff, were analyzed (Fig. 1). Silica contents range from 64-73 wt%; major elements scatter about the linear trends defined by Cenozoic samples in this study. In general, Proterozoic rocks have lower CaO, and higher TiO_2 and P_2O_5 , than Cenozoic samples at the same silica concentration. Trace elements are also relatively scattered, with high Zr concentrations in several samples (Fig. 4). Proterozoic rocks collected around the caldera generally have high La_N/Lu_N ratios and negative Eu anomalies (Figs. 5, 6). In contrast, the granodiorite xenolith has relatively low REE concentrations and a small positive Eu anomaly ($\text{Eu}/\text{Eu}^* = 1.22$). Isotopic compositions (corrected to 35 Ma) vary widely (Figs. 7, 8). Initial $^{87}\text{Sr}/^{86}\text{Sr}$ ranges 0.724-0.783 and ϵNd_i from -15.4 to -23.3. In all samples except the metasedimentary gneiss, $^{206}\text{Pb}/^{204}\text{Pb}_i$ ranges from 16.89-17.51 and $^{207}\text{Pb}/^{204}\text{Pb}_i$ from 15.46-15.53. In contrast, the metasedimentary gneiss (sample PG15-25) has higher $^{206}\text{Pb}/^{204}\text{Pb}_i$ and $^{207}\text{Pb}/^{204}\text{Pb}_i$ ratios (18.79 and 15.98). Thorogenic Pb is variable, from $^{208}\text{Pb}/^{204}\text{Pb}_i = 37.00$ -43.14.

DISCUSSION

Interpretation of new geochronology

Many of the geochronologic data are reversely discordant, and were acquired on both instruments used in this study. One possible cause for the discordant data is that zircon populations may have lost U-rich “parent” zones by hydrothermal or metamorphic processes after crystallization (Mattinson et al., 1996). Indeed, many of the zircons in this study have relatively high Th/U (Table 5). However, the reversely discordant data in this study are likely a result of preparation and/or analytical issues because: 1) some zircons with high Th/U are concordant; 2) some zircons with low Th/U are reversely discordant; 3) the whole rocks themselves also have high Th/U, suggesting the high Th/U of the zircons is primary (Table 2). Regardless of the origin of reverse discordance, the geochronologic data are interpreted with caution.

Rocks in Grizzly Peak caldera

The 35.05 ± 0.15 Ma age for a zircon from the Grizzly Peak Tuff is some 750 ka older than its 34.3 ± 0.12 Ma sanidine $^{40}\text{Ar}/^{39}\text{Ar}$ age (McIntosh and Chapin, 2004). Discrepancies between the U-Pb and Ar/Ar systems often occur in volcanic rocks, although they typically differ by less than 500 ka (Costa, 2008). For example, the weighted mean zircon U-Pb age of the Badger Creek Tuff from the Mount Princeton magmatic center (Fig. 1), is ~210 ka older than the 34.26 ± 0.06 Ma sanidine $^{40}\text{Ar}/^{39}\text{Ar}$ age (Zimmerer and McIntosh, 2012; Mills and Coleman, 2013). Recalculation of the Grizzly Peak Tuff Ar/Ar data using updated K decay constant and Fish Canyon sanidine age (Renne et al., 2011) yields an updated sanidine age of 34.87 ± 0.12 Ma, which is within uncertainty of the new zircon U-Pb age before taking into account decay

constant uncertainties. More zircon geochronology on the Grizzly Peak Tuff is required to confirm that the single zircon age is robust.

Geochronology for the resurgent plutons are difficult to interpret due to reverse discordance. However, a zircon from resurgent pluton 1 is concordant within uncertainty and overlaps with the zircon U-Pb age for the tuff. This suggests the pluton was emplaced shortly after eruption of the tuff. The rest of the discordant data for the resurgent plutons also suggest crystallization ages shortly following eruption. Short timespans between zircon U-Pb ages for the tuff and resurgent plutons are supported by field relationships (Fridrich et al., 1991) and are also observed in other Cenozoic calderas in the Rockies (Tappa et al., 2011; Mills and Coleman, 2013; Rosera et al., 2013).

Fridrich et al. (1991) speculated that granitic xenoliths similar to sample PR15-18 may represent completely solidified parts of the Grizzly Peak magma chamber that were subsequently entrained during resurgent magmatism. The concentration data largely corroborate the similarities between the tuff and granitic xenoliths, although sample PR15-18 has lower Ba and Th/U than the tuff (Figs. 4, 5). However, U-Pb zircon data for the granitic xenolith suggest a crystallization age approximately 3 Ma older than the Grizzly Peak Tuff (Fig. 9). One explanation for this discrepancy is that the zircon data are compromised by inheritance. If Grizzly Peak magma assimilated or otherwise interacted with slightly older granitoids, the Cenozoic data for sample PR15-18 could represent recycling of slightly older zircons, or mixing between inherited and autocrystic growth (Miller et al., 2007).

Whole rock isotopic data more clearly elucidate the relationship between the xenolith and the Grizzly Peak system. Sample PR15-18 has distinctly less radiogenic Sr and Pb, and more radiogenic Nd, than the tuff (Figs. 7, 8). Thus, the xenolith cannot represent a crystallized portion

of the Grizzly Peak magma chamber, and therefore the zircon U-Pb data likely reflect magmatism in the area before the Grizzly Peak system was established. The granitic xenolith appears to share numerous similarities to the Twin Lakes pluton. The interpreted crystallization age (38 Ma) is similar to the youngest parts of the Twin Lakes pluton (39 Ma; Feldman, 2010) and the Winfield-Middle Mountain area (38 Ma; Fridrich et al., 1998). In addition, inherited zircon characteristics are similar. The Twin Lakes pluton commonly has a 1400 Ma inherited component (Feldman, 2010), and a chord regressed through the granitic xenolith's concordant and normally discordant data yields an upper intercept age of 1410 ± 11 Ma. Combined with similar whole rock isotopic compositions (Figs. 7, 8; Stein, 1985; Stein and Crock, 1990), we suggest that this xenolith and others like it in the post-resurgent intrusions (Fridrich et al., 1998) may represent a western extension of the Twin Lakes pluton that was disrupted by the eruption of the Grizzly Peak Tuff and subsequent resurgent magmatism.

Outflow tuff in Arkansas River Valley

Two zircons from the tuff in the Arkansas River Valley suggest a crystallization age ca. 24.5 Ma, whereas K/Ar in biotite yielded an age of 33.1 ± 1.0 Ma (Fridrich, 1987). The discrepancy could be due to inherited Ar in the biotite, Pb-loss in the zircons, or a combination of the two. However, the U-Pb data are likely more robust because the zircons were chemically abraded (Mundil et al., 2004; Mattinson, 2005), and it is unlikely that two zircons of different sizes would record the same apparent age after experiencing Pb-loss. Therefore, the age of this tuff outcrop is best approximated by the zircon U-Pb data.

Significance of chemical and isotopic data for Proterozoic basement rocks

Lead isotopic data for the metasedimentary gneiss (sample PG15-25) are intriguing

because of its high uranogenic Pb isotope ratios (Fig. 8). Although the $^{206}\text{Pb}/^{204}\text{Pb}$ ratio is similar to the ~1800 Ma Idaho Springs Formation (Stein, 1985), sample PG15-25 also has a $^{207}\text{Pb}/^{204}\text{Pb}$ ratio near 16, which is much higher than data for Proterozoic whole rocks from northern and southern Colorado (Fig. 8; Aleinikoff et al., 1993; McLennan et al., 1995). High $^{207}\text{Pb}/^{204}\text{Pb}$ ratios relative to $^{206}\text{Pb}/^{204}\text{Pb}$ are reported in the Archean Wyoming Province, where 2.79 Ga rocks in the eastern Beartooth Mountains have similar present-day Pb isotopic compositions to the metasedimentary gneiss (Wooden and Mueller, 1988; Frost et al., 1998; Mueller and Frost, 2006; Mueller et al., 2010). Many of the Wyoming Archean rocks also have radiogenic Sr isotopic compositions similar to sample PG15-25 (Wooden and Mueller, 1988).

Lead isotopic and U-Pb zircon evidence suggests there is only minor influence of Archean material in Paleoproterozoic rocks in the northern Colorado, closest to the Wyoming province (Aleinikoff et al., 1993). However, more recent zircon and whole rock isotopic work has revealed Archean-aged zircons and the involvement of older crust in the genesis of Paleoproterozoic rocks in central Colorado (Hill and Bickford, 2001; Bickford and Hill, 2007; Bickford et al., 2008; Jones and Thrane, 2012). One Archean-aged zircon was also identified in LA-ICP-MS analyses of the Twin Lakes pluton (Feldman, 2010). Therefore, it is plausible that the metasedimentary gneiss, which is at least 1701 Ma (Jones and Thrane, 2012), acquired its Pb (and possibly Sr) from Archean rocks during the Paleoproterozoic.

A newly recognized tuff in the Arkansas River Valley

The data presented here indicate that the small tuff outcrop in the Arkansas River Valley cannot be a remnant of extracaldera Grizzly Peak Tuff. Significantly, the isotopic data do not permit a link because the tuff in the Arkansas River Valley has lower $^{87}\text{Sr}/^{86}\text{Sr}$, $^{206}\text{Pb}/^{204}\text{Pb}$,

$^{207}\text{Pb}/^{204}\text{Pb}$, and $^{208}\text{Pb}/^{204}\text{Pb}$ ratios (Figs. 7, 8). Moreover, U-Pb zircon geochronology yields consistent ages ca. 24.5 Ma, some ~10 Ma younger than the Grizzly Peak Tuff.

It is difficult to determine the significance of this unit. Outcrops are poorly exposed over only 0.03 km² and are deeply weathered (Fridrich et al., 1998). It erupted after initiation of Rio Grande rifting in this area, which began ~30-28 Ma and is characterized by bimodal magmatic compositions (Shannon, 1988; Chapin and Cather, 1994; Mills and Coleman, 2013; Ricketts et al., 2016). Its location on rift flank makes it difficult to determine if the small outcrop area is due to small-volume eruption, or to weathering and poor preservation as a consequence of rifting. The lack of volcanic rocks with similar ages in this part of Colorado (McIntosh and Chapin, 2004) suggests this eruptive unit may have been a local phenomenon. Around the same time, the youngest leucogranites in the Climax porphyry Mo system were emplaced ~35 km to the northeast (24 Ma; Bookstrom et al., 1988), and the small Raspberry Gulch rhyolite dome erupted in the Mount Princeton area (22 Ma; Campbell, 1994; Zimmerer and McIntosh, 2012).

Sources and magmatic processes in the Twin Lakes pluton

Magma sources

The restricted Sr isotopic range for the Twin Lakes pluton, from $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.7061$ - 0.7071 in rocks from 63-42 Ma (Feldman, 2010), suggests either a persistent and homogenous magma source(s), or that efficient and fortuitous mixing of different sources occurred throughout the pluton's diachronous history. The Pb and Nd data suggest the latter scenario is more likely, as they are variable in time and space. Three samples from the northwestern portion of the pluton have $\epsilon\text{Nd}_i \approx -8.6$, whereas Farmer and DePaolo (1984) measured a more radiogenic value of -5.4 in the northeastern part of the pluton. A sample of ~41 Ma Twin Lakes pluton from near

Winfield/Middle Mountain has $\epsilon\text{Nd}_i = -10.1$ (Stein and Crock, 1990; Feldman, 2010). The Pb isotopic data are more straightforward. All Paleocene samples have $^{206}\text{Pb}/^{204}\text{Pb}_i$ greater than 18.1, whereas Eocene samples have $^{206}\text{Pb}/^{204}\text{Pb}_i \approx 17.7$, suggesting a switch to a source with lower U/Pb ratios (Fig. 8). Thorogenic Pb data are more difficult to distinguish, but $^{208}\text{Pb}/^{204}\text{Pb}_i$ ratios appear slightly lower in younger samples.

The Pb isotopic data are intriguing because of recent hypotheses for the genesis of Mo-rich deposits. Pettke et al. (2010) suggested that Mo deposits are characterized by low uraniumogenic Pb ratios relative to otherwise similar Mo-poor intrusions, due to sourcing from ancient metasomatized subcontinental lithospheric mantle. In the case of the Twin Lakes pluton, the young, unradiogenic samples of Stein (1985) come from the Winfield/Middle Mountain area, which may be an eroded Climax-type Mo deposit (Ranta, 1974; Stein and Crock, 1990). Metasomatized subcontinental lithospheric mantle is also hypothesized to be enriched in light REE and large-ion lithophile elements, leading to low Sm/Nd and moderate Rb/Sr ratios (Kessel et al., 2005; Pettke et al., 2010). If the Eocene parts of the pluton were derived from such a source, it suggests that 1) the Paleocene Twin Lakes pluton was derived from a different source than the Eocene parts; 2) the similar Sr isotopic compositions between the Eocene and Paleocene parts of the pluton are coincidental.

Emplacement level processes

Isotopic data demonstrate that the modally layered sample (TL16-15m) has the same source as its host rock (TL16-15; Table 3; Figs. 7, 8). The modal layering's REE pattern is also parallel to the host rock's, but with a more pronounced negative Eu anomaly. It has previously been suggested that modal layering in the Twin Lakes resulted from flow sorting and deformation

processes that effectively segregated mafic minerals from the main granodiorite (Wilshire, 1969). However, the high viscosities of silicic magmas preclude sedimentary-type mechanisms for generating such features (Glazner, 2014). An alternative explanation that would satisfy the chemical data, including REE enrichment, and Al_2O_3 , K_2O , CaO , Na_2O , Ba, Sr, and Eu depletion relative to its host rock (Figs. 3, 4), is that it was once relatively “normal” granodiorite from which felsic liquid was removed (e.g., Coleman et al., 2011).

Two aplites collected from opposite ends of the northern part of the Twin Lakes pluton have starkly different REE patterns (Fig. 6). Sample TL16-10d, collected at the northwestern margin of the pluton (Fig. 2), has a scoop-shaped REE pattern and negligible Eu anomaly characteristic of aplites interpreted to have been separated from a titanite-bearing granodioritic mush (Glazner et al., 2008). In contrast, aplite TL16-15d was collected from a similarly-aged outcrop on the far eastern margin of the pluton, and has a flat REE pattern and a large negative Eu anomaly. Compared to the Oligocene-Miocene granites of Stein and Crock (1990), which also have prominent negative Eu anomalies, this aplite has low REE concentrations. Such a flat pattern in a pluton that contains titanite throughout (Fridrich et al., 1998; Feldman, 2010) is enigmatic. One possibility is that aplitic liquid was separated from a granodiorite mush prior to titanite crystallization (Colombini et al., 2011). However, it is plausible that this aplite dike was not derived from the granodiorite it intrudes. The $^{208}\text{Pb}/^{204}\text{Pb}$ data in particular permit the aplite to have been derived from a different source than the host rock (Fig. 8). Thus, the aplite on the eastern part of the Twin Lakes pluton could have been derived from a separate, titanite-free source from depth.

Source and history of the Grizzly Peak magmatic system

Hypotheses for the chemical and isotopic diversity of the Grizzly Peak Tuff have focused

on the heterogeneous tuff horizons that contain compositionally diverse pumice fiamme (Fridrich and Mahood, 1987; Johnson and Fridrich, 1990). For example, Zr-TiO₂ plots of pumice samples follow a liquid line of descent, allowing the possibility that the tuff resulted from fractional crystallization of a large magma chamber prior to eruption (Fig. 10; Fridrich and Mahood, 1987). Isotopic diversity may have been generated by assimilation of diverse Proterozoic rocks at emplacement level (Johnson and Fridrich, 1990). In addition, the resurgent plutons are interpreted to have been derived from the same magma as the tuff (Fridrich and Mahood, 1984).

Analyses of resurgent pluton samples fill the compositional space between all tuff compositions on the Zr-TiO₂ plot, suggesting that after eruption of the tuff, some portions of compositionally diverse melts could have remained in the chamber and mixed in varying proportions to produce the resurgent pluton compositions (Fig. 10). However, the isotopic data preclude such a scenario. At comparable SiO₂ concentrations, the resurgent plutons have less radiogenic Nd and more radiogenic Pb than the tuff (Figs. 7, 8). Alternatively, perhaps the leftover tuff magmas were modified by assimilation of Proterozoic wallrocks to generate the resurgent pluton isotopic compositions (e.g., Johnson and Fridrich, 1990). However, this is unlikely because assimilation is thermally limited in the upper crust (Glazner, 2007). Furthermore, if ancient rocks were the assimilant, the Sr isotopic compositions of the resurgent plutons should also diverge from that of the tuff.

In light of new compositional and isotopic data, we propose that the Grizzly Peak Tuff and the resurgent plutons were derived in large part from partial melting of felsic Proterozoic hydrous lower crust (e.g., Stein and Crock, 1990). The volumetrically minor mafic pumices in the tuff (57 wt% SiO₂) and mafic enclaves in the plutons could have been derived from small-scale compositional heterogeneities in the source, such as metasomatized amphibolite (e.g., Rapp and

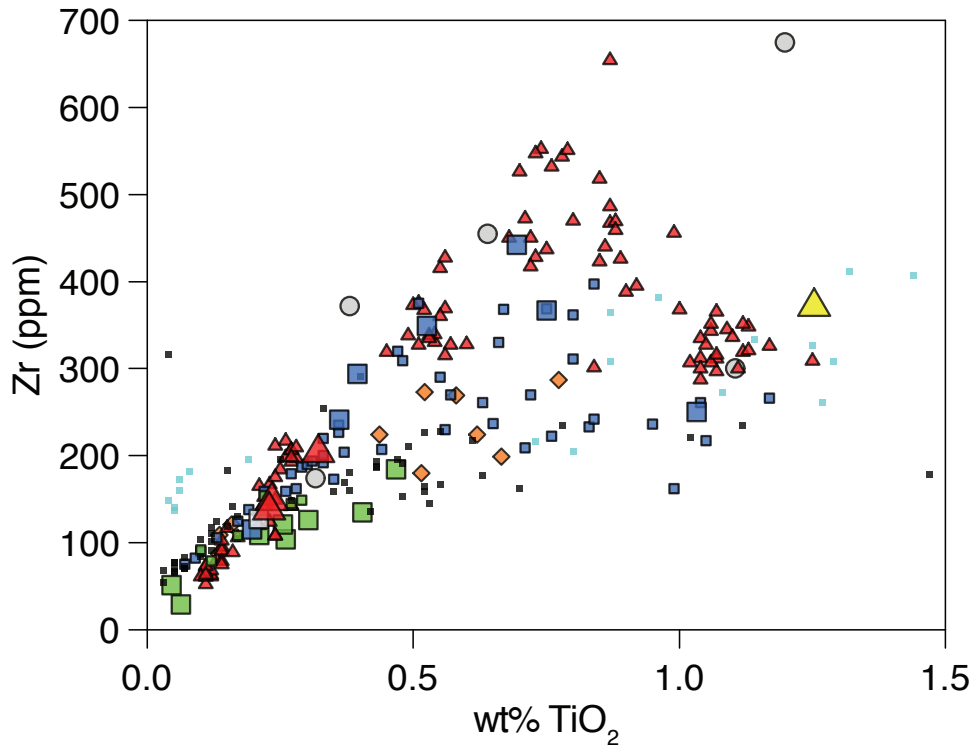


Figure 10. Plot of Zr-TiO₂ data for samples from this study (large symbols) and Cenozoic igneous rocks of Colorado. Data for the resurgent plutons fill in compositional gaps between pumice fiamme groups, suggesting the possibility that the plutons represent mixtures of unerupted Grizzly Peak magmas. However, Nd isotopic data for the plutons are less radiogenic than the tuff, making a mixing hypothesis untenable. Symbols and data sources as in Figure 3. Large amount of additional literature data for Grizzly Peak Tuff (small red triangles) due to inclusion of pressed powder X-ray fluorescence analyses of Fridrich (1987). Data presentation after Fridrich (1987) and Fridrich and Mahood (1987).

Wilson, 1995). A lower crustal source is supported by the high Th/U contents of in tuff and resurgent pluton whole rocks and zircons dated here (Fig. 5; Table 5). In addition, the source must have been ancient and felsic in order to generate high ⁴⁰Ca/⁴⁴Ca ratios (Mills et al., 2013). This model benefits from being simpler than calling upon coupled assimilation-fractional crystallization of primitive basalts, followed by a second stage of assimilation and fractional crystallization in the upper crust to generate the observed compositions (Johnson and Fridrich, 1990). In addition, partial melting can produce variable isotopic compositions such as those observed within the tuff and resurgent plutons (Figs. 7, 8; Stevens et al., 2007; Clemens et al., 2011; Villaros et al., 2012).

Relationships to the Colorado Mineral Belt

The Grizzly Peak Tuff represents an excursion in Sr and Nd isotopic compositions relative to the Colorado Mineral Belt and other rocks from the central Colorado volcanic field, including the Thirtynine Mile and San Juan volcanic fields, and Rio Grande rift magmatism (Fig. 11). Neodymium isotopic data for the tuff are distinctly unradiogenic compared to the dataset except for two much younger samples from the southern mineral belt (Stein and Crock, 1990). Strontium isotopic data for the tuff are relatively evolved, but are exceeded by a few mineral belt plutons (Simmons and Hedge, 1978; Stein and Crock, 1990), and the Never Summer Mountains outside the mineral belt (Jacob et al., 2015). In contrast, there is no distinction between the Grizzly Peak Tuff and other Cenozoic igneous rocks from Colorado in Pb isotopic compositions, including thorogenic Pb, which are similar to other central mineral belt rocks (Stein, 1985; Fig. 11).

Although temporal and spatial isotopic trends have been recognized in the Colorado Mineral Belt (Stein, 1985; Stein and Crock, 1990; Chapin, 2012), the long magmatic history of the central part of the mineral belt appears to be characterized by relatively stable isotopic compositions if the Grizzly Peak Tuff is considered an anomalous crustal melting event. For example, all Sr isotopic data from the 63-39 Ma Twin Lakes pluton, the ~38 Ma granitic xenolith in Grizzly Peak caldera, and the ~24.5 Ma mafic tuff in the Arkansas River Valley fall in the range $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.7061$ to 0.7072 . Limited Nd isotopic data for the same rocks fall in the range $\epsilon\text{Nd}_i = -5$ to -10 ; although the radiogenic end of the range in Sr data is anchored by just one sample (Farmer and DePaolo, 1984). Cenozoic rocks in the nearby Mount Princeton magmatic center have similar isotopic compositions (Figs. 7, 8; Mills, 2012). The data suggest a common source(s) was tapped continually over Cenozoic time in the central Colorado Mineral Belt, occasionally punc-

tuated by eruption of voluminous, isotopically anomalous tuffs (Grizzly Peak and Wall Mountain) that are rich in ancient crustal components (Johnson and Fridrich, 1990; Mills, 2012; Mills et al., 2013; this study).

In contrast to the relatively persistent Sr (and Nd?) isotopic compositions over time in the central mineral belt, Pb isotopic data in the central mineral belt show temporal trends, particularly in $^{206}\text{Pb}/^{204}\text{Pb}$. Stein (1985) attributed the broad decrease in uraniumogenic Pb ratios with time to

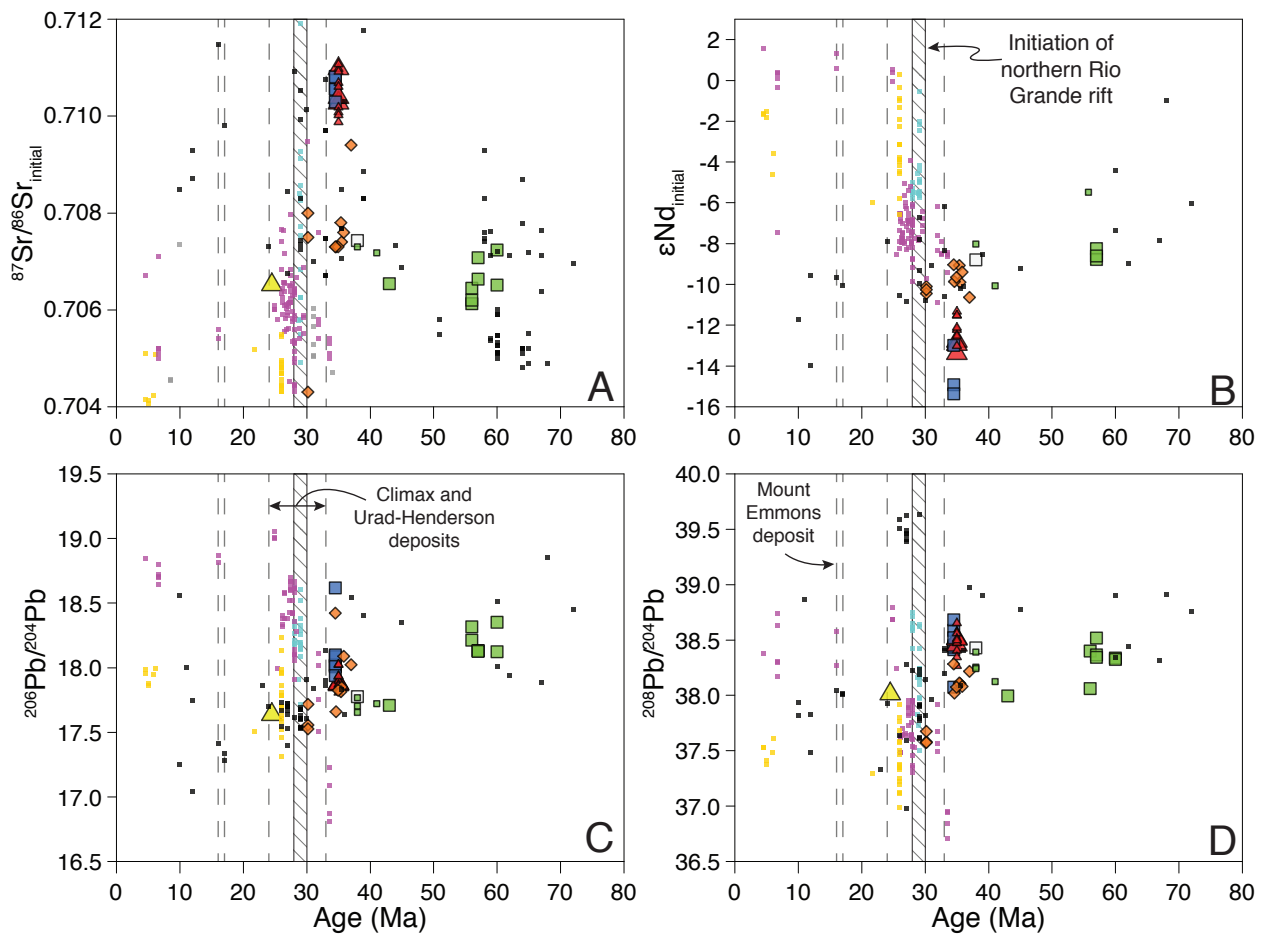


Figure 11. Diagrams showing variations through time in Colorado igneous rocks for (A) $^{87}\text{Sr}/^{86}\text{Sr}_i$; (B) ϵNd_i ; (C) $^{206}\text{Pb}/^{204}\text{Pb}$; (D) $^{208}\text{Pb}/^{204}\text{Pb}$. Grizzly Peak Tuff and resurgent plutons are distinct in Sr and Nd isotopic compositions (panels A and B). Secular trends in Colorado Mineral Belt data are apparent in all plots except (D) $^{208}\text{Pb}/^{204}\text{Pb}$. Mineral belt rocks have more radiogenic Sr, and less radiogenic Nd and $^{206}\text{Pb}/^{204}\text{Pb}$ through time. In the central Colorado Mineral Belt, there are no apparent trends in the Sr or Nd data (excluding the Grizzly Peak magmatic center), whereas uraniumogenic Pb mirrors trends in the mineral belt as a whole towards less radiogenic compositions, including Grizzly Peak data. Data symbols and sources as in Figures 3, 7, and 8. Colorado Mineral Belt dataset (small black squares) also includes data of Simmons and Hedge (1978). Ages of Mo deposits from Bookstrom et al. (1988), Stein and Crock (1990), and Shannon et al. (2004). Timing of northern Rio Grande rift initiation from Zimmerer and McIntosh (2004), Mills and Coleman (2013), and Ricketts et al. (2016).

the spatial distribution of mineral belt magmatism, as the youngest rocks are in the southwest and have the lowest $^{206}\text{Pb}/^{204}\text{Pb}_i$ ratios. Stein (1985) suggested the regional trend may be due to spatial compositional variability in a lower crustal source, or due to different melting depths. In contrast, the data presented here show that $^{206}\text{Pb}/^{204}\text{Pb}_i$ ratios may also vary secularly in a relatively restricted area (Fig. 11), as in the Twin Lakes pluton. These data suggest that early magmatism in the central Colorado Mineral Belt preferentially tapped sources with high U/Pb ratios, whereas lower U/Pb source(s) predominated later, although it is not clear why such a change occurred.

The relatively consistent Sr data coupled with the systematic changes in uraniumogenic Pb ratios through time in the central Colorado Mineral Belt (Grizzly Peak and Wall Mountain tuffs excepted) may be explained by several scenarios: 1) primitive mantle is the primary control on isotopic composition of the rocks, with assimilation-fractional crystallization producing the consistent Sr isotopic compositions. Lead isotopic compositions are controlled by the lower crust due to its much higher Pb contents relative to mantle melts (Johnson and Fridrich, 1990; Johnson, 1991). 2) Rocks are derived from metasomatized, relatively mafic lower crust of varying composition, with perhaps some metasomatized mantle playing a role (Stein and Crock, 1990; Jacob et al., 2015). 3) Isotopic compositions are largely controlled by enriched subcontinental lithospheric mantle (Coleman et al., 1992). The first scenario is probably least likely in the central part of the mineral belt, because it requires similar compositions and amounts of assimilated to be incorporated into primitive basalts, which must then differentiate to a similar degree repeatedly over a nearly 40 Ma interval. Production of isotopically homogeneous granitoids after assimilating isotopically heterogeneous Proterozoic rocks would be fortuitous.

The data cannot distinguish between the latter two models. Although the Grizzly Peak Tuff may have been derived from metasomatized felsic lower crust, that does not preclude other

rocks in the area from being derived from mafic lower crust as well (Stein, 1985; Stein and Crock, 1990). Indeed, the lower crust in Colorado is inferred to be isotopically heterogeneous (Lake and Farmer, 2015). If such heterogeneity existed vertically and laterally, derivation of most central Colorado Mineral Belt rocks by anatexis of lower crust is plausible, provided there was a sufficient volume of source rocks with the appropriate compositions. The possibility of an enriched lithospheric mantle source for the central part of the mineral belt is also intriguing because the lithospheric mantle is inferred to also be laterally heterogeneous. Enriched mantle under some parts of the Rocky Mountains reaches $^{87}\text{Sr}/^{86}\text{Sr}_i = 0.705$, $\epsilon\text{Nd}_i = -5$, and $^{206}\text{Pb}/^{204}\text{Pb}_i = 20.0$ (Roden et al., 1990; Bailey, 2010; Gonzales et al., 2010). The simplest scenario in the central mineral belt would involve a parcel of enriched subcontinental lithospheric mantle with isotopic characteristics similar to those analyzed here ($^{87}\text{Sr}/^{86}\text{Sr}_i = 0.706$, $\epsilon\text{Nd}_i = -8$, and $^{206}\text{Pb}/^{204}\text{Pb}_i = 17.6$). It is also possible that some combination of enriched lithospheric mantle and metasomatized mafic Proterozoic lower crust served as the source(s) for most of the Twin Lakes Granodiorite and other igneous rocks in this area (e.g., Stein and Crock, 1990).

Why is there not a large Mo deposit at Grizzly Peak?

The location and composition of the Grizzly Peak magmatic system are enticing because of their apparent similarities to nearby Climax-type deposits, including Climax, Urad-Henderson, and Mount Emmons (Fig. 1; Ludington and Plumlee, 1999). Numerous mining claims and workings for Mo, Au, Pb, Zn, and other metals in and around the caldera attest to the appeal of the Grizzly Peak magmatic system's economic potential (Wilson and Sims, 2003; Mineral Resources Data System, USGS, 2005). The existing geochemical dataset for the Grizzly Peak caldera (Fig. 4; Table 2; Fridrich, 1987) also appears similar to Climax-type deposits in most respects except,

critically, it lacks concentrated Mo. The 38 Ma granitic xenolith and some resurgent pluton samples analyzed here have Mo concentrations up to 9 ppm, or ~10 times the continental crust average (0.8 ppm; Rudnick and Gao, 2003), but samples of bulk Grizzly Peak Tuff have Mo below the detection limit (<2 ppm), and Fridrich (1987) did not analyze pumice fiamme for Mo.

One explanation for the lack of concentrated Mo assumes the magmatic system had a similar genesis to other Climax-type Mo systems, particularly the Questa deposit that was preceded by a 500 km³ rhyolitic tuff (Tappa et al., 2011; Zimmerer and McIntosh, 2012). If that was the case, it is possible that the current level of exposure in the Grizzly Peak system does not permit observation of deeper high-grade ore zones, which are typically restricted to 100-200 m of mineralizing intrusions (Ludington and Plumlee, 2009). Such a lack of exposure could be due to the preservation of intracaldera Grizzly Peak Tuff, or because the caldera has not been disrupted and tilted like the Questa magmatic system (Johnson et al., 1989; Rosera et al., 2013). Another possibility is that the Grizzly Peak system was not permitted to concentrate Mo, whether through low degree partial melting (Rosera et al., 2013) or extensive fractional crystallization (Johnson et al., 1989; von Quadt et al., 2011), and therefore Mo is “diluted” in the tuff and resurgent plutons. However, because Mo behaves incompatibly in silicate systems (Candela and Holland, 1986), both partial melting and fractional crystallization would be expected to concentrate Mo above crustal averages in the silicic rocks of the Grizzly Peak magmatic system, which is not observed.

Another explanation for the lack of exposed, concentrated Mo in the Grizzly Peak magmatic system is that despite its size and location, its timing did not allow it to produce significant Mo-rich fluids. Climax-type deposits are associated with extensional environments that may permit thermal and/or chemical interaction with the underlying mantle (Stein and Crock, 1990; Ludington and Plumlee, 2009; Chapin, 2012), and the Grizzly Peak system predates Rio Grande

rift extension (Ricketts et al., 2016). However, the nearby Climax deposit began developing during “tectonic relaxation” ca. 33 Ma, and thus a similar tectonic regime should also apply to the Grizzly Peak system (Bookstrom et al., 1988).

The distinct isotopic compositions of the Grizzly Peak Tuff and the resurgent plutons suggest source characteristics may be an important factor (Figs. 7, 11). Past isotopic surveys of the Colorado Mineral Belt sought out rocks associated with ores and important mining or prospecting activity (Stein, 1985; Stein and Crock, 1990). Therefore, the lack of Colorado Mineral Belt rocks with similar Nd and Sr isotopic compositions to the Grizzly Peak system is significant and suggests that source rocks with those isotopic compositions are not economically important. Since the isotopic compositions of the Grizzly Peak Tuff and its resurgent plutons are interpreted to have been inherited from a felsic Proterozoic lower crustal source, it suggests that felsic lower crust in this part of the mineral belt was not an important source of Mo. This is in contrast to proposals for derivation of Mo-rich deposits by anatexis of felsic lower crust elsewhere in the mineral belt (Stein and Crock, 1990). Some recent models that call for hybridized lower crust as the source Mo may offer an explanation (Richards, 2009, 2011; Rosera et al., 2013). Rosera et al. (2013) suggested that high-flux mantle inputs to the lower crust in the Questa magmatic system in northern New Mexico (Fig. 1) generated the 500 km³ Amalia Tuff and also constructed a hybridized lower crust rich in mantle-derived volatiles and metals such as Mo. Subsequent partial melting of the volatile-rich hybridized crust scavenged those metals and concentrated them in small volume, high-silica magmas that yielded a Climax-type porphyry system upon crystallization (Rosera et al., 2013). Indeed, the Sr and Nd isotopic compositions of mineralizing intrusions at Questa are more juvenile than the Amalia Tuff or earlier magmatism (Gaynor et al., 2015). However, there is no exposed evidence for a juvenile isotopic component after eruption of the

Grizzly Peak Tuff (Fig. 11). This suggests the Grizzly Peak magmatic system was either derived by mechanisms that did not generate a volatile-hybridized lower crust, or volatile-hybridized lower crust was generated but was not incorporated into the exposed Grizzly Peak magmatic system.

CONCLUSIONS

New isotopic and geochemical data in the central Colorado Mineral Belt demonstrate:

- 1) a Proterozoic metasedimentary gneiss adjacent to the Grizzly Peak caldera has Pb (and possibly other isotopic components) similar to the those found in the Archean Wyoming Province;
- 2) Cenozoic rocks adjacent to the Grizzly Peak caldera record broadly consistent Sr and Nd isotopic compositions over a 40 Ma period that spans the eruption of the Grizzly Peak Tuff;
- 3) the Grizzly Peak Tuff and its resurgent plutons are chemically and isotopically distinct relative to nearby Cenozoic igneous rocks including the Twin Lakes pluton, a granitic xenolith ~3 Ma older than the tuff, and an Oligocene mafic tuff exposed 20 km away in the Arkansas River Valley. The radiogenic Sr and highly unradiogenic Nd isotopic compositions of the Grizzly Peak Tuff and resurgent plutons, coupled with trace element indicators such as high Th/U, suggest they may have been generated by anatexis of Proterozoic lower crust. The resurgent plutons have slightly more enriched Sr, Nd, and Pb isotopic compositions than the tuff, precluding a complementary relationship between the two; the data imply compositional heterogeneities were inherited from the source. Preliminary zircon U-Pb geochronology on the tuff and reinterpretation of high-quality $^{40}\text{Ar}/^{39}\text{Ar}$ data suggest the Grizzly Peak Tuff erupted ca. 35 Ma; additional U-Pb zircon data suggest the resurgent plutons were emplaced shortly after eruption. We infer the broadly similar isotopic compositions present in the Cenozoic rocks of the central Colorado Mineral Belt to be

derived from a different lower crustal source(s) than the Grizzly Peak Tuff, or from enriched subcontinental lithospheric mantle. The lack of Mo enrichment in the Grizzly Peak Tuff and its resurgent plutons, despite their compositions and location in the mineral belt, suggests either that Mo-rich deposits are not exposed, or that the isotopically-enriched lower crustal source was not an important source of Mo.

TABLE 1. SUMMARY OF SAMPLES FROM CENTRAL COLORADO MINERAL BELT

Sample	Map unit	Description	Age		UTM-E*	UTM-N*	IGSN [†]
			Age	source			
TL15-05	Tt	Twin Lakes pluton	57	1	368022	4323340	P4700001J
MB15-06	Tm	Twin Lakes megabreccia	57	1	367990	4323299	P4700001K
MB15-07	Tm	Twin Lakes megabreccia	57	1	367994	4323253	P4700001L
GU15-08	Tg	Upper Grizzly Peak Tuff	35	2	363360	4329185	P4700001M
XE15-09	Tg	Granodiorite xenolith in tuff	pЄ?	2	363360	4329185	P4700001N
GM15-17	Tg	Middle Grizzly Peak Tuff	35	2	360802	4320610	P4700001V
PR15-18	Tp	Granite xenolith in resurgent pluton	38	2	360898	4320589	P4700001W
GM15-20	Tgm	Middle Grizzly Peak Tuff	35	2	361190	4320835	P4700001Y
P215-21	Tr2	Intermediate age resurgent pluton	34.5	2	360216	4325534	P4700001Z
P115-22	Tr1	Oldest resurgent pluton	34.5	2	360634	4326478	P47000020
P115-23	Tr1	Oldest resurgent pluton	34.5	2	360629	4326480	P47000021
PG15-25	Xms	Metasedimentary gneiss	pЄ	3	356140	4328593	P47000023
GG15-26	Yg	Grottos pluton	pЄ	3	354416	4330144	P47000024
SK15-27	Ys	St. Kevin Granite	pЄ	3	367129	4321084	P47000025
DC15-28	Xdc	Leucocratic Denny Creek pluton	pЄ	3	366809	4326996	P47000026
OT15-29	Tav	Mafic tuff in Arkansas River valley	24.5	2	385541	4324606	P47000027
TL16-10	Tt	Twin Lakes pluton	60	1	369277	4325526	P4700002G
TL16-10d	Tt	Aplite in Twin Lakes pluton	60	1	369277	4325526	P4700002H
TL16-12	Tt	Twin Lakes pluton	43	1	373395	4325775	P4700002K
TL16-15	Tt	Twin Lakes pluton	56	1	378796	4324886	P4700002N
TL16-15d	Tt	Aplite in Twin Lakes pluton	56	1	378796	4324886	P4700002O
TL16-15m	Tt	Modally layered Twin Lakes pluton	56	1	378796	4324886	P4700002P
GP-06	Tr2	Intermediate age resurgent pluton	34.5	2	360334	4324228	
GP-07	Tr1	Oldest resurgent pluton	34.5	2	360125	4326646	
GP-08	Tr1	Oldest resurgent pluton	34.5	2	360125	4326646	
GP-11	Tr2	Intermediate age resurgent pluton	34.5	2	358220	4327733	

Note: Ages for Cenozoic samples given in Ma. Age sources are: 1 – Feldman (2010); 2 – this study; 3 – Fridrich et al. (1998). Ages for Twin Lakes samples are estimates based on proximity to dated samples of Feldman (2010). Ages for Grizzly Peak Tuff and resurgent pluton samples dated in this study are applied to non-dated samples of those same units

*UTM coordinates for samples in NAD 83 datum, zone 13S

[†]IGSN – International Geo Sample Number. Additional sample metadata available at www.geosamples.org. Samples GP-06 through GP-11 are courtesy of M. Zimmerer

TABLE 2. MAJOR AND TRACE ELEMENT ANALYSES

Sample	TL15-05	MB15-06	MB15-07	GU15-08	XE15-09	GM15-17	PR15-18	GM15-20	P215-21
Description	Twin Lakes pluton	(Twin Lakes) Mbreccia	(Twin Lakes) Mbreccia	Upper GPT	Grano- diorite xenolith	Middle GPT	Xeno. in resurgent pluton	Middle GPT	Middle resurgent pluton
Major element analyses* by ICP-OES									
SiO ₂	67.14	71.46	64.83	72.98	68.00	68.84	70.96	70.47	61.47
Al ₂ O ₃	15.76	13.66	15.94	13.88	15.01	14.77	14.64	14.02	17.2
Fe ₂ O ₃	3.11	2.66	5.24	1.91	3.52	3.02	2.53	2.2	5.53
MnO	0.095	0.055	0.132	0.026	0.038	0.050	0.053	0.058	0.073
MgO	0.65	0.18	1.01	0.35	1.19	0.58	0.40	0.34	1.51
CaO	2.65	2.45	3.74	1.31	1.95	2.32	1.86	1.52	4.43
Na ₂ O	3.77	4.45	4.44	2.95	2.67	3.35	3.87	3.38	3.43
K ₂ O	4.27	1.52	1.62	4.59	5.43	4.34	4.31	4.44	2.89
TiO ₂	0.303	0.21	0.467	0.234	0.64	0.321	0.209	0.229	0.695
P ₂ O ₅	0.13	0.12	0.21	0.07	0.23	0.09	0.09	0.07	0.25
LOI	1.33	1.98	1.74	1.55	1.21	1.23	0.51	2.05	2.85
Total	99.21	98.74	99.38	99.85	99.91	98.90	99.43	98.78	100.30
Trace element analyses [†] by ICP-OES									
Sc	4	3	6	5	12	6	3	5	9
Be	1	2	2	1	3	1	2	2	1
V	31	23	54	14	56	23	19	15	69
Ba	1748	258	238	1561	2178	1425	679	1012	1423
Sr	983	628	805	337	427	504	385	283	866
Y	19	12	28	19	15	19	10	24	20
Zr	126	109	184	151	455	206	127	140	442
Trace element analyses [†] by ICP-MS									
Cr	190	30	170	20	190	20	210	20	90
Co	4	3	6	2	6	3	3	2	8
Ni	BDL	BDL	BDL	BDL	20	BDL	BDL	BDL	BDL
Cu	10	BDL	50	BDL	20	BDL	10	BDL	20
Zn	70	50	120	50	60	80	50	100	80
Ga	21	18	25	16	19	18	20	17	21
Ge	2	1	2	1	1	1	1	1	1
Rb	87	42	42	116	170	93	138	124	71
Nb	10	6	13	8	8	8	11	9	8
Mo	3	BDL	3	BDL	3	BDL	4	BDL	9
Ag	BDL	BDL	BDL	BDL	1.3	BDL	BDL	BDL	BDL
Sn	2	1	2	1	2	1	2	1	2
Sb	0.7	0.6	0.6	0.7	0.7	0.6	0.6	0.7	BDL
Cs	1.2	1.8	1.3	3.6	5.7	2.8	1.3	2.7	2.1
La	27.2	17.3	63.9	55.5	30.1	72.1	24.7	56.6	82.6
Ce	52.9	33.3	116.0	95.7	55.9	132.0	46.9	109.0	153.0
Pr	5.96	3.92	12.7	11.4	6.37	13.9	5.07	11.7	16.4
Nd	23.2	15.1	45.7	38.1	22.8	47.8	17.7	39.6	58.2
Sm	4.6	3.2	8.4	6.4	4.0	7.5	3.4	6.9	8.8
Eu	1.34	0.89	1.86	1.54	1.45	1.67	0.88	1.40	2.40
Gd	3.6	2.5	6.3	4.3	3.3	5.1	2.5	5.1	6.0
Tb	0.6	0.4	0.9	0.6	0.5	0.7	0.3	0.7	0.8
Dy	3.2	2.2	5.2	3.3	2.8	3.8	1.9	4.3	3.8
Ho	0.6	0.4	1.0	0.7	0.5	0.7	0.4	0.8	0.8
Er	1.9	1.2	2.9	2.0	1.4	2.0	1.0	2.3	2.2
Tm	0.30	0.20	0.44	0.30	0.20	0.31	0.16	0.36	0.34
Yb	2.0	1.4	3.0	2.0	1.3	1.9	1.2	2.4	2.3
Lu	0.34	0.23	0.48	0.32	0.21	0.31	0.18	0.39	0.36
Hf	3.2	3.0	4.8	4.1	10.2	5.2	3.5	3.9	9.5
Ta	1.1	0.8	1.5	0.9	0.7	0.8	1.3	1.1	0.4
W	2	2	2	2	3	2	2	3	4
Tl	0.4	0.2	0.2	0.5	0.7	0.5	0.6	0.8	0.4
Pb	34	17	24	23	27	49	41	57	20
Th	7.2	5.8	15.7	17.2	9.5	18.3	16.8	19.2	13.8
U	2.6	2.4	4.8	2.7	1.5	2.8	5.2	3.4	2.3

TABLE 2 (CONTINUED). MAJOR AND TRACE ELEMENT ANALYSES

Sample	P115-22	P115-23	PG15-25	GG15-26	SK15-27	DC15-28	OT15-29	TL16-10	TL16-10d
Description	Oldest resurgent pluton	Oldest resurgent pluton	Metased gneiss	Grottos pluton	St. Kevin Granite	Denny Creek pluton	Tuff in Ark. Valley	Twin Lakes pluton	Twin Lakes aplite
Major element analyses* by ICP-OES									
SiO ₂	69.45	67.36	72.67	64.16	72.99	63.73	58.59	75.09	78.08
Al ₂ O ₃	15.00	15.83	14.33	14.75	13.27	15.56	14.97	13.00	13.02
Fe ₂ O ₃	2.23	3.13	3.45	6.47	2.89	8.03	7.88	2.49	0.60
MnO	0.022	0.073	0.034	0.073	0.034	0.085	0.086	0.069	0.017
MgO	0.38	0.70	0.47	1.45	0.16	1.99	1.17	0.68	0.05
CaO	0.60	3.02	0.72	2.45	0.56	3.53	5.48	2.74	1.39
Na ₂ O	3.48	3.56	1.54	3.06	1.95	3.18	3.36	3.28	5.19
K ₂ O	6.83	3.54	5.03	4.59	6.75	2.49	3.00	2.10	1.47
TiO ₂	0.197	0.395	0.316	1.199	0.380	1.105	1.253	0.260	0.063
P ₂ O ₅	0.05	0.14	0.32	0.54	0.06	0.20	0.86	0.06	BDL
LOI	0.63	2.26	1.29	0.87	1.23	0.85	2.14	0.67	0.39
Total	98.87	100	100.2	99.6	100.3	100.8	98.79	100.4	100.3
Trace element analyses [†] by ICP-OES									
Sc	2	6	5	8	5	22	13	3	2
Be	1	1	BDL	3	2	2	2	1	3
V	16	32	20	80	16	112	94	27	BDL
Ba	1199	1753	805	1385	992	413	1577	350	16
Sr	317	684	153	411	186	205	1127	756	62
Y	9	18	22	10	21	37	25	9	6
Zr	115	294	174	675	372	300	374	104	29
Trace element analyses [†] by ICP-MS									
Cr	30	100	30	190	30	230	80	BDL	BDL
Co	3	4	5	12	3	20	15	3	BDL
Ni	BDL	BDL	BDL	30	BDL	40	30	BDL	BDL
Cu	BDL	BDL	BDL	40	10	30	20	BDL	BDL
Zn	BDL	70	50	140	50	110	80	60	BDL
Ga	18	19	16	23	16	23	21	18	21
Ge	BDL	1	1	1	1	2	1	1	3
Rb	169	88	172	180	171	131	70	57	53
Nb	4	9	9	18	7	16	25	8	6
Mo	BDL	2	BDL	3	BDL	3	BDL	BDL	BDL
Ag	BDL	0.6	BDL	1.7	0.7	BDL	1.3	BDL	BDL
Sn	2	2	2	4	3	5	2	1	BDL
Sb	0.6	0.6	0.5	0.6	0.7	0.6	BDL	BDL	BDL
Cs	0.9	0.9	3.2	2.0	3.0	1.9	0.7	1.7	1.1
La	35.0	93.5	53.9	96.0	162.0	58.4	79.9	29.7	10.9
Ce	62.2	172.0	117.0	185.0	335.0	124.0	165.0	54.2	19.8
Pr	6.43	17.80	13.90	19.60	37.40	14.60	19.50	5.95	1.87
Nd	22.1	59.3	51.1	64.9	132.0	54.3	74.4	20.6	5.2
Sm	4.1	8.3	10.8	8.5	21.4	11.2	12.9	3.5	0.7
Eu	0.90	2.05	1.33	1.86	3.13	1.90	3.34	0.94	0.19
Gd	3.2	5.5	8.1	4.8	11.9	9.8	9.4	2.5	0.7
Tb	0.4	0.7	1.1	0.5	1.3	1.5	1.2	0.3	BDL
Dy	2.0	3.8	5.2	2.3	6.0	8.3	5.7	1.8	0.6
Ho	0.3	0.7	0.8	0.4	0.9	1.5	1.0	0.3	0.2
Er	0.9	2.0	1.9	1.0	2.1	3.8	2.6	1.1	0.7
Tm	0.12	0.31	0.26	0.15	0.24	0.44	0.35	0.16	0.14
Yb	0.7	2.0	1.5	1.1	1.4	2.5	2.1	1.1	1.4
Lu	0.11	0.32	0.20	0.22	0.20	0.36	0.33	0.19	0.32
Hf	3.2	7.3	4.5	14.8	9.4	7.9	8.5	2.7	1.5
Ta	0.5	0.7	0.9	1.1	0.7	1.0	1.3	0.5	0.4
W	3	2	2	3	2	2	BDL	1	BDL
Tl	0.7	0.4	0.7	0.9	0.8	0.5	BDL	0.2	0.1
Pb	38	21	41	23	31	16	11	20	44
Th	16.3	17.8	19.4	10.4	24.5	17.4	11.2	7.6	43.9
U	2.9	2.5	5.5	1.1	1.7	1.4	2.1	2.0	45.6

TABLE 2 (CONTINUED). MAJOR AND TRACE ELEMENT ANALYSES

Sample	TL16-12	TL16-15	TL16-15d	TL16-15m	GP-06	GP-07	GP-08	GP-11
Description	Twin Lakes pluton	Twin Lakes pluton	Twin Lakes aplite	Modally layered Twin Lakes	Middle resurgent pluton	Oldest resurgent pluton	Mafic enclave in GP-07	Middle resurgent pluton
Major element analyses* by ICP-OES								
SiO ₂	68.58	68.73	77.11	48.47	60.8	67.94	55.56	64.76
Al ₂ O ₃	16.06	15.78	12.85	11.53	17.48	15.78	16.63	16.57
Fe ₂ O ₃	3.41	2.7	0.84	23.09	5.27	3.07	7.74	3.94
MnO	0.050	0.090	0.029	0.465	0.099	0.062	0.146	0.086
MgO	0.92	0.63	0.05	2.63	1.69	0.73	1.98	0.99
CaO	2.74	2.34	0.94	4.97	4.79	2.65	4.54	3.94
Na ₂ O	3.92	3.42	3.76	2.74	3.2	3.66	3.75	3.55
K ₂ O	3.4	5.26	4.08	1.90	2.86	3.19	2.24	3.09
TiO ₂	0.404	0.255	0.045	1.638	0.751	0.361	1.033	0.525
P ₂ O ₅	0.15	0.13	BDL	0.43	0.26	0.13	0.33	0.18
LOI	0.37	0.3	0.2	2.49	2.41	1.06	4.74	2.67
Total	100.0	99.63	99.90	100.40	99.62	98.64	98.70	100.30
Trace element analyses [†] by ICP-OES								
Sc	4	4	2	23	10	6	18	8
Be	2	1	2	2	1	1	1	1
V	47	37	7	400	74	30	158	46
Ba	866	2048	26	232	1755	1923	918	1809
Sr	723	1155	62	566	867	772	634	794
Y	11	14	19	76	20	15	20	18
Zr	135	121	51	677	367	241	250	349
Trace element analyses [†] by ICP-MS								
Cr	BDL	BDL	BDL	30	BDL	BDL	30	BDL
Co	6	3	BDL	22	9	4	18	5
Ni	BDL	BDL	BDL	BDL	BDL	BDL	BDL	BDL
Cu	BDL	BDL	BDL	BDL	20	BDL	30	BDL
Zn	50	50	BDL	390	80	50	120	70
Ga	21	20	20	41	21	18	24	20
Ge	1	2	2	3	1	1	1	1
Rb	85	127	123	115	72	74	61	82
Nb	13	13	16	61	9	7	10	9
Mo	BDL	BDL	BDL	BDL	BDL	BDL	BDL	BDL
Ag	BDL	BDL	BDL	2.5	1.2	0.9	0.9	1.2
Sn	2	2	1	7	1	2	12	1
Sb	BDL	BDL	BDL	BDL	BDL	BDL	BDL	BDL
Cs	1.2	3.2	2.2	11.4	0.8	1.1	2.5	1.3
La	37.4	33.5	5.8	328.0	94.3	83.0	56.9	106.0
Ce	69.9	61.6	13.9	572	172	151	111	193
Pr	7.92	6.73	1.94	59.0	18.4	15.8	12.7	20.1
Nd	28.9	24.5	8.4	198.0	64.8	53.9	46.2	67.9
Sm	5.1	4.4	2.8	29.4	9.2	7.5	7.9	9.2
Eu	1.29	1.07	0.32	5.39	2.41	2.08	2.18	2.31
Gd	3.8	3.2	3.2	19.9	6.4	5.0	6.0	5.9
Tb	0.5	0.4	0.6	2.5	0.8	0.6	0.8	0.7
Dy	2.4	2.5	3.6	14.3	4.3	3.0	4.4	3.9
Ho	0.4	0.5	0.7	2.7	0.8	0.5	0.8	0.7
Er	1.2	1.5	2.1	8.1	2.2	1.6	2.4	2.1
Tm	0.17	0.23	0.33	1.20	0.33	0.24	0.34	0.31
Yb	1.1	1.7	2.2	8.7	2.2	1.6	2.3	1.9
Lu	0.16	0.29	0.36	1.38	0.35	0.24	0.35	0.33
Hf	3.9	3.6	2.4	19.0	8.9	6.2	6.2	8.5
Ta	1.0	0.9	1.2	3.9	0.5	0.5	0.5	0.5
W	BDL	1	BDL	3	BDL	1	BDL	BDL
Tl	0.2	0.5	0.5	0.6	0.3	0.3	0.2	0.2
Pb	17	33	52	23	14	35	12	18
Th	12.4	10.1	11.8	83.8	15.3	14.2	10.4	17.5
U	3.3	4.1	19.5	11.3	2.2	1.9	2.2	2.4

Note: LOI – loss on ignition; BDL – below detection limit; Mbrecchia – megabrecchia; GPT – Grizzly Peak Tuff; Xeno. – xenolith; Metased – metasedimentary; Ark. – Arkansas

*Major element data in wt%

[†]Trace element data in ppm

TABLE 3. SR AND ND ISOTOPIC ANALYSES

Sample	Age (Ma)	Rb ppm	Sr ppm	$^{87}\text{Rb}/^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}$	$\pm 2\sigma$ abs*	$^{87}\text{Sr}/^{86}\text{Sr}_t$	Sm ppm	Nd ppm	$^{147}\text{Sm}/^{144}\text{Nd}$	$^{143}\text{Nd}/^{144}\text{Nd}$	$\pm 2\sigma$ abs*	$^{143}\text{Nd}/^{144}\text{Nd}_i$	ϵNd_t
TL15-05	57	87	983	0.256	0.706840	0.000011	0.706633	4.6	23.2	0.1200	0.512186	0.000004	0.512142	-8.25
MB15-06	57	42	628	0.193	0.707237	0.000010	0.707080	3.2	15.1	0.1281	0.512162	0.000003	0.512115	-8.78
MB15-07	57							8.4	45.7	0.1111	0.512165	0.000004	0.512124	-8.60
GU15-08	35	116	337	0.996	0.711501	0.000010	0.711006	6.4	38.1	0.1016	0.511933	0.000004	0.511910	-13.32
XE15-09 [†]	35	170	427	1.154	0.724901	0.000009	0.724327	4.0	22.8	0.1061	0.511748	0.000003	0.511724	-16.95
GM15-17	35	93	504	0.534	0.710564	0.000010	0.710299	7.5	47.8	0.0949	0.511963	0.000004	0.511941	-12.72
PR15-18	38	138	385	1.037	0.707989	0.000011	0.707429	3.4	17.7	0.1161	0.512168	0.000003	0.512139	-8.79
GM15-20	35	124	283	1.268	0.711046	0.000010	0.710416	6.9	39.6	0.1054	0.511958	0.000004	0.511935	-12.87
P215-21	34.5	71	866	0.237	0.710926	0.000010	0.710810	8.8	58.2	0.0914	0.511948	0.000003	0.511928	-12.99
P115-22	34.5	169	317	1.546	0.732528	0.000010	0.731770	4.1	22.1	0.1122	0.511854	0.000004	0.511829	-14.92
P115-23	34.5	88	684	0.372	0.710473	0.000010	0.710291	8.3	59.3	0.0846	0.511825	0.000004	0.511806	-15.37
PG15-25 [†]	35	172	153	3.277	0.784155	0.000009	0.782526	10.8	51.1	0.1278	0.511802	0.000004	0.511773	-16.00
GG15-26 [†]	35	180	411	1.270	0.729430	0.000012	0.728799	8.5	64.9	0.0792	0.511417	0.000003	0.511417	-23.30
SK15-27 [†]	35	171	186	2.672	0.753408	0.000011	0.752080	21.4	132.0	0.0980	0.511604	0.000003	0.511582	-19.73
DC15-28 [†]	35	131	205	1.856	0.744355	0.000009	0.743433	11.2	54.3	0.1247	0.511831	0.000004	0.511802	-15.43
OT15-29	24.5	70	1127	0.180	0.706610	0.000011	0.706547							
TL16-10	60	57	756	0.218	0.706695	0.000008	0.706509							
TL16-10d	60	53	62	2.474	0.709345	0.000011	0.707236							
TL16-12	43	85	723	0.340	0.706747	0.000008	0.706539							
TL16-15	56	127	1155	0.318	0.706382	0.000008	0.706129							
TL16-15d	56	123	62	5.742	0.711015	0.000010	0.706447							
TL16-15m	56	115	566	0.588	0.706666	0.000010	0.706198							
GP-06	34.5	72	867	0.240	0.710937	0.000010	0.710819							
GP-07	34.5	74	772	0.277	0.709837	0.000011	0.709701							
GP-08	34.5	61	634	0.278	0.710693	0.000010	0.710557							
GP-11	34.5	82	794	0.299	0.710696	0.000009	0.710550							

Note: Subscript "t" indicates $^{87}\text{Sr}/^{86}\text{Sr}$ and ϵNd values corrected for radiogenic ingrowth since the listed age. All ages are interpreted as approximate crystallization or eruption ages except where noted

*Analytical uncertainty only. See text for reproducibility of standards

[†]Sample interpreted as Proterozoic; isotopic values corrected to Grizzly Peak Tuff eruption at 35 Ma

TABLE 4. PB ISOTOPIC ANALYSES

Sample	Age (Ma)	U ppm	Th ppm	Pb ppm	$^{206}\text{Pb}/$ ^{204}Pb	2σ (abs.)*	$^{207}\text{Pb}/$ ^{204}Pb	2σ (abs.)*	$^{208}\text{Pb}/$ ^{204}Pb	2σ (abs.)*	$^{206}\text{Pb}/$ $^{204}\text{Pb}_i$	$^{207}\text{Pb}/$ $^{204}\text{Pb}_i$	$^{208}\text{Pb}/$ $^{204}\text{Pb}_i$
TL15-05	57	2.6	7.2	34	18.175	0.010	15.590	0.008	38.558	0.022	18.132	15.588	38.519
MB15-06	57	2.4	5.8	17	18.212	0.010	15.589	0.008	38.430	0.019	18.133	15.586	38.367
MB15-07	57	4.8	15.7	24	18.237	0.009	15.586	0.008	38.466	0.020	18.125	15.581	38.345
GU15-08	35	2.7	17.2	23	17.928	0.012	15.563	0.012	38.550	0.037	17.887	15.562	38.466
XE15-09 [†]	35	1.5	9.5	27	16.907	0.011	15.458	0.012	37.041	0.035	16.889	15.457	37.003
GM15-17	35	2.8	18.3	49	17.900	0.010	15.560	0.009	38.568	0.025	17.880	15.560	38.526
PR15-18	38	5.2	16.8	41	17.822	0.014	15.553	0.012	38.478	0.032	17.775	15.551	38.428
GM15-20	35	3.4	19.2	57	17.907	0.016	15.573	0.016	38.560	0.043	17.886	15.572	38.522
P215-21	34.5	2.3	13.8	20	18.137	0.018	15.601	0.015	38.789	0.034	18.098	15.599	38.681
P115-22	34.5	2.9	16.3	38	18.643	0.007	15.621	0.006	38.123	0.014	18.617	15.620	38.075
P115-23	34.5	2.5	17.8	21	17.978	0.015	15.570	0.014	38.616	0.034	17.938	15.567	38.522
PG15-25 [†]	35	5.5	19.4	41	18.833	0.095	15.980	0.080	39.212	0.208	18.786	15.977	39.157
GG15-26 [†]	35	1.1	10.4	23	16.994	0.011	15.492	0.011	37.272	0.027	16.978	15.491	37.222
SK15-27 [†]	35	1.7	24.5	31	17.006	0.015	15.476	0.013	39.065	0.031	16.987	15.475	38.976
DC15-28 [†]	35	1.4	17.4	16	17.540	0.014	15.526	0.012	43.266	0.032	17.508	15.525	43.136
OT15-29	24.5	2.1	11.2	11	17.693	0.008	15.512	0.007	38.108	0.018	17.647	15.510	38.028
TL16-10 [§]	60	2.0	7.6	20	18.183	0.008	15.581	0.007	38.410	0.018	18.123	15.578	38.337
TL16-10d [§]	60	45.6	43.9	44	18.972	0.008	15.614	0.008	38.521	0.018	18.351	15.584	38.325
TL16-12 [§]	43	3.3	12.4	17	17.790	0.008	15.527	0.008	38.097	0.022	17.708	15.523	37.996
TL16-15 [§]	56	4.1	10.1	33	18.386	0.009	15.582	0.009	38.458	0.023	18.317	15.579	38.402
TL16-15d [§]	56	19.5	11.8	52	18.421	0.006	15.574	0.005	38.102	0.011	18.214	15.564	38.061
TL16-15m [§]	56	11.3	83.8	23	18.424	0.003	15.578	0.003	38.599	0.011	18.152	15.566	37.935
GP-06 [§]	34.5	2.2	15.3	14	18.140	0.008	15.568	0.007	38.710	0.018	18.086	15.565	38.588
GP-07 [§]	34.5	1.9	14.2	35	18.024	0.009	15.551	0.009	38.492	0.022	18.006	15.550	38.447
GP-08 [§]	34.5	2.2	10.4	12	18.077	0.010	15.563	0.009	38.508	0.023	18.015	15.560	38.412
GP-11 [§]	34.5	2.4	17.5	18	18.066	0.012	15.556	0.012	38.544	0.037	18.020	15.553	38.436

Note: Measured isotopic ratios corrected for fractionation using a value of 0.165%/amu except where noted. Subscript "i" indicates Pb ratio corrected for radiogenic ingrowth since the listed age. All ages are interpreted as crystallization or eruption ages except where noted

*Analytical uncertainty only

[†]Sample interpreted as Proterozoic; isotopic values corrected to Grizzly Peak Tuff eruption at 35 Ma

[§]Measured isotopic ratios corrected for fractionation using a value of 0.15%/amu. See text for explanation

TABLE 5. U-PB ZIRCON ANALYSES FOR SELECTED SAMPLES OF GRIZZLY PEAK TUFF, RESURGENT PLUTONS, AND TUFF IN ARKANSAS RIVER VALLEY

Fraction	Composition			Isotopic Ratios						Ages (Ma)						Corr. coef.	
	U (ppm)	Pb (pg) ^a	Th/U ^b	Total common Pb (pg) ^c	²⁰⁶ Pb/ ²⁰⁴ Pb ^d	²⁰⁶ Pb/ ²³⁸ U ^e	Error (2σ%)	²⁰⁷ Pb/ ²³⁵ U ^e	Error (2σ%)	²⁰⁷ Pb/ ²⁰⁶ Pb ^g	Error (2σ%)	²⁰⁶ Pb/ ²³⁸ U ^g	Error (2σ abs.)	²⁰⁷ Pb/ ²³⁵ U ^g	Error (2σ abs.)		²⁰⁷ Pb/ ²⁰⁶ Pb ^g
PR15-18, Granite xenolith in post-resurgent pluton. Sector 54																	
F-1	36	0.58	1.1	1961	0.005896	0.13	0.03776	1.1	0.04647	1.0	37.985	0.050	37.63	0.42	21	25	0.746
F-3	40	0.33	3.0	884	0.005904	0.22	0.03742	1.8	0.04598	1.7	38.046	0.083	37.30	0.67	-4	40	0.767
F-5	62	0.27	2.3	1748	0.069936	0.13	0.8355	0.57	0.08669	0.53	435.87	0.54	616.7	2.7	1353	10	0.445
GM15-20, Middle subunit Grizzly Peak Tuff. Phoenix X62																	
F-3	5	1.2	1.5	171	0.005439	0.43	0.0353	8.3	0.0470	7.9	35.05	0.15	35.2	2.8	50	190	0.800
P215-21, Intermediate age resurgent pluton. Sector 54																	
F-1	8	1.47	1.9	218	0.005340	0.30	0.0289	7.0	0.0393	6.7	34.42	0.10	28.9	2.0	-400	170	0.847
F-3	7	1.13	1.8	214	0.005347	0.48	0.0294	8.1	0.0398	7.7	34.46	0.16	29.4	2.4	-363	200	0.901
F-4	4	1.46	2.3	98	0.005343	1.1	0.0258	25	0.0350	24	34.43	0.39	25.8	6.4	-711	670	0.836
P115-22, Oldest resurgent pluton. Phoenix X62																	
F-1	6	0.71	1.9	211	0.005318	0.53	0.0236	12	0.0322	12	34.29	0.18	23.7	2.8	-944	340	0.804
F-2	5	1.18	3.0	111	0.005295	0.69	0.0190	25	0.0260	24	34.13	0.23	19.1	4.6	-1625	820	0.839
P115-23, Oldest resurgent pluton. Phoenix X62																	
F-1	10	1.70	1.9	275	0.005381	0.41	0.0313	6.0	0.0423	5.7	34.68	0.14	31.3	1.8	-211	140	0.775
F-2	6	1.42	1.2	277	0.005407	0.30	0.0333	4.9	0.0446	4.7	34.85	0.10	33.2	1.6	-75	110	0.759
OT15-29, Tuff in Arkansas River Valley. Sector 54																	
F-1	4	1.51	5.7	53	0.003789	1.5	0.0206	38	0.039	36	24.46	0.37	20.7	7.7	-390	950	0.897
F-2	2	1.66	1.2	112	0.003771	0.89	0.0197	19	0.0380	19	24.34	0.22	19.8	3.8	-488	500	0.878

Note: All analyses performed on chemically abraded single zircon grains. Instrument used indicated next to sample name.

^aTotal mass of radiogenic Pb.

^bTh contents calculated from radiogenic ²⁰⁸Pb and ²³⁰Th-corrected ²⁰⁶Pb/²³⁸U date of the sample, assuming concordance between U-Pb and Th-Pb systems.

^cTotal mass of common Pb.

^dMeasured ratio corrected for fractionation and spike contribution only.

^eMeasured ratios corrected for fractionation, tracer and blank. Pb blank ratios: ²⁰⁶Pb/²⁰⁴Pb = 18.687 ± 0.25; ²⁰⁷Pb/²⁰⁴Pb = 15.658 ± 0.25; and ²⁰⁸Pb/²⁰⁴Pb = 38.258 ± 0.5 (1σ abs.)

^fCorrected for initial Th/U disequilibrium using radiogenic ²⁰⁶Pb and Th/U_[magma], which is assumed to be approximated by the measured whole rock Th/U ratio.

^gIsotopic dates calculated using λ₂₃₈ = 1.55125E⁻¹⁰ (Jaffey et al. 1971) and λ₂₃₅ = 9.8485E⁻¹⁰ (Jaffey et al. 1971).

4-15	172	68	0.39	0.0527	3.6	0.1090	3.4	0.0152	1.7	0.35	97.0	1.7	105.0	3.4	302	66
4-16	414	180	0.43	0.0498	2.8	0.1023	2.1	0.0148	1.4	0.51	94.5	1.4	98.9	2.0	174	44
4-17*	715	526	0.74	0.0494	2.3	0.0964	1.8	0.0142	1.6	0.77	90.6	1.4	93.5	1.6	165	26
4-18	381	147	0.39	0.0493	2.7	0.1012	2.5	0.0150	1.7	0.64	95.7	1.7	97.9	2.3	158	40
4-19	1453	733	0.50	0.0485	2.2	0.0982	1.6	0.0148	1.4	0.87	94.4	1.3	95.1	1.5	121	21
4-20	711	340	0.48	0.0483	2.6	0.0967	2.5	0.0147	1.7	0.76	93.7	1.6	94.0	2.2	111	35
4-21	1116	558	0.50	0.0479	2.4	0.0967	1.9	0.0147	1.4	0.65	94.3	1.3	94.0	1.6	96	30
4-22	550	346	0.63	0.0478	2.6	0.0986	2.5	0.0149	1.6	0.74	95.3	1.5	95.4	2.3	85	37
4-23	265	139	0.53	0.0487	3.5	0.1015	3.2	0.0150	1.5	0.47	96.1	1.5	98.1	2.9	134	60
4-24	895	465	0.52	0.0489	2.5	0.1012	2.5	0.0149	1.7	0.75	95.4	1.6	97.8	2.3	142	34
13JLE-16, Deer Creek Gabbro Orb Core (in Dinkey Creek Granodiorite), IGSN: P47000015																
16-1	310	119	0.38	0.0502	3.0	0.1058	2.6	0.0156	1.6	0.52	99.5	1.6	102.1	2.5	191	48
16-2	659	244	0.37	0.0498	2.3	0.1056	1.8	0.0156	1.5	0.64	100.0	1.5	101.9	1.7	183	28
16-3	640	254	0.40	0.0540	2.5	0.1177	2.0	0.0157	1.4	0.65	100.3	1.4	113.0	2.1	368	34
16-4	475	166	0.35	0.0489	2.6	0.1062	2.2	0.0156	1.7	0.63	99.8	1.7	102.4	2.1	141	36
16-5	129	47	0.36	0.0519	3.7	0.1139	3.7	0.0156	2.1	0.48	99.8	2.1	109.4	3.8	282	72
16-6	286	102	0.36	0.0490	3.5	0.1092	3.4	0.0160	1.7	0.54	102.5	1.8	105.2	3.4	158	61
16-7	618	258	0.42	0.0494	2.7	0.1054	2.0	0.0155	1.4	0.51	98.8	1.4	102.0	2.0	168	40
16-8	273	111	0.41	0.0496	3.1	0.1073	2.6	0.0157	1.5	0.40	100.4	1.4	103.4	2.5	183	54
16-9	489	177	0.36	0.0500	2.8	0.1066	2.6	0.0156	1.3	0.56	100.0	1.3	102.8	2.6	197	45
16-10	273	226	0.83	0.0493	3.0	0.1045	3.1	0.0153	1.6	0.47	98.0	1.5	100.8	3.0	160	50
16-11	1613	1214	0.75	0.0485	2.2	0.1052	1.8	0.0157	1.5	0.89	100.5	1.5	101.5	1.7	118	19
16-12	256	115	0.45	0.0500	3.0	0.1092	2.3	0.0157	1.4	0.34	100.4	1.4	105.6	2.4	190	49
16-13*	147	47	0.32	0.0508	4.1	0.1140	4.6	0.0164	1.9	0.53	105.1	2.0	109.5	4.8	222	78
16-14	1321	600	0.45	0.0481	2.2	0.1037	1.6	0.0156	1.3	0.80	99.5	1.3	100.2	1.6	101	22
16-15	377	129	0.34	0.0497	2.8	0.1050	2.5	0.0154	1.6	0.61	98.5	1.5	101.4	2.4	178	43
16-16	204	127	0.62	0.0506	3.0	0.1099	2.4	0.0156	1.4	0.32	100.1	1.4	105.8	2.4	218	50
16-17*	390	229	0.59	0.0487	2.9	0.1148	2.9	0.0170	1.6	0.68	108.7	1.8	110.3	3.0	130	47
16-18	265	123	0.46	0.0503	3.3	0.1072	3.1	0.0157	1.4	0.42	100.1	1.4	103.3	3.1	203	59
16-19	810	286	0.35	0.0489	2.5	0.1057	2.1	0.0157	1.5	0.67	100.3	1.5	102.0	2.0	140	36
16-20	699	595	0.85	0.0486	2.5	0.1062	2.0	0.0157	1.7	0.72	100.4	1.7	102.5	1.9	126	35
16-21	946	583	0.62	0.0484	2.3	0.1069	2.1	0.0161	1.4	0.76	102.7	1.5	103.1	2.0	116	27
16-22	595	256	0.43	0.0491	2.5	0.1067	1.8	0.0158	1.6	0.59	101.3	1.6	102.9	1.7	143	37
16-23	703	292	0.41	0.0488	2.4	0.1068	1.7	0.0158	1.5	0.68	100.8	1.5	103.0	1.6	137	30
16-24	719	278	0.39	0.0485	2.6	0.1033	2.1	0.0155	1.5	0.66	99.4	1.4	99.8	2.0	124	36
13JLE-17, Dinkey Creek Granodiorite (Deer Creek orb host), IGSN: P47000016																
17-1	242	78	0.32	0.0495	3.5	0.1093	3.2	0.0160	1.8	0.53	102.0	1.8	105.3	3.2	169	60
17-2	446	175	0.39	0.0484	2.8	0.1032	2.9	0.0156	1.7	0.66	99.9	1.7	99.7	2.8	111	44
17-3	231	94	0.41	0.0495	3.3	0.1052	3.3	0.0155	1.7	0.65	99.1	1.7	101.9	3.2	169	56
17-4	681	403	0.59	0.0486	2.6	0.1018	2.4	0.0153	1.7	0.71	98.0	1.6	98.4	2.2	121	38
17-5	183	84	0.46	0.0496	4.0	0.1045	3.0	0.0154	1.8	0.15	98.7	1.8	100.9	2.8	170	73
17-6	165	63	0.38	0.0515	4.2	0.1113	4.0	0.0158	1.8	0.35	101.1	1.8	107.0	4.0	249	79
17-7	180	83	0.46	0.0530	3.2	0.1113	3.1	0.0154	1.6	0.58	98.6	1.6	107.1	3.2	328	56
17-8	227	91	0.40	0.0505	3.9	0.1101	3.5	0.0159	1.6	0.34	101.7	1.7	105.9	3.5	218	73
17-9	289	111	0.38	0.0491	2.9	0.1031	2.6	0.0154	1.6	0.60	98.7	1.6	99.6	2.5	151	45
17-10	214	91	0.43	0.0508	3.2	0.1071	3.1	0.0153	1.4	0.60	98.1	1.4	103.3	3.0	226	58
17-11	190	73	0.38	0.0493	4.2	0.1072	3.7	0.0157	1.7	0.32	100.1	1.6	103.3	3.7	156	78
17-12	177	83	0.47	0.0512	3.9	0.1083	4.0	0.0154	1.6	0.50	98.2	1.6	104.9	3.8	241	71
17-13	160	65	0.40	0.0517	4.2	0.1089	2.9	0.0155	1.6	0.20	98.9	1.6	104.9	3.0	260	78
17-14	228	69	0.30	0.0488	3.3	0.1040	2.9	0.0155	1.7	0.44	99.3	1.6	100.4	2.8	135	58
17-15	202	93	0.46	0.0515	3.4	0.1096	3.1	0.0154	1.4	0.45	98.3	1.4	105.6	3.1	253	62
17-16	137	57	0.41	0.0510	3.7	0.1086	3.8	0.0154	1.9	0.58	98.3	1.8	104.6	3.7	233	68
17-17	142	55	0.39	0.0504	4.4	0.1081	4.0	0.0155	1.8	0.27	99.0	1.8	104.1	3.9	214	87
17-18	151	60	0.39	0.0513	3.7	0.1086	3.7	0.0155	1.7	0.30	99.4	1.7	104.6	3.6	244	69
17-19	193	84	0.43	0.0495	3.3	0.1078	3.3	0.0157	1.9	0.52	100.4	1.9	103.8	3.3	171	59
17-20	190	77	0.41	0.0502	3.6	0.1088	2.6	0.0155	1.7	0.27	99.0	1.7	104.8	2.6	198	65
17-21	155	59	0.38	0.0506	3.4	0.1105	3.6	0.0155	2.0	0.60	99.1	2.0	106.3	3.7	216	62
17-22	133	58	0.43	0.0522	4.0	0.1127	4.0	0.0155	1.7	0.46	99.0	1.7	108.3	4.1	278	75
17-23	218	98	0.45	0.0498	3.1	0.1078	3.0	0.0157	1.9	0.60	100.4	1.9	103.9	2.9	181	52
17-24	129	59	0.46	0.0518	4.2	0.1101	4.0	0.0155	1.6	0.26	98.9	1.6	106.0	4.0	260	80
13JLE-18, Deer Creek Gabbro Orb Core (in Dinkey Creek Granodiorite), IGSN: P47000017																
18-1	87	42	0.48	0.0494	3.0	0.1061	2.0	0.0154	1.4	0.30	98.7	1.4	102.4	1.9	163	48
18-2	110	64	0.58	0.0493	2.8	0.1058	2.5	0.0155	1.4	0.53	98.8	1.4	102.0	2.4	157	47
18-3	308	184	0.60	0.0478	2.3	0.1026	1.3	0.0156	1.0	0.52	99.8	0.9	99.2	1.2	85	25
18-4	192	85	0.44	0.0514	5.6	0.1116	5.9	0.0157	1.3	0.73	100.1	1.3	107.1	5.8	153	35
18-5	192	76	0.40	0.0522	7.2	0.1130	8.8	0.0155	1.4	0.94	99.2	1.3	108.2	8.5	173	62
18-6*	110	72	0.65	0.0506	3.1	0.1638	2.8	0.0235	1.2	0.85	149.5	1.8	153.9	4.0	210	50
18-7*	308	143	0.46	0.0489	2.4	0.1076	1.6	0.0160	1.0	0.56	102.1	1.0	103.8	1.6	138	30
18-8	173	81	0.47	0.0505	2.4	0.1076	2.0	0.0155	1.3	0.60	99.3	1.3	103.7	2.0	209	33
18-9	167	58	0.35	0.0491	2.8	0.1074	2.6	0.0158	1.3	0.48	101.2	1.4	103.5	2.6	148	41
18-10	216	84	0.39	0.0486	3.2	0.1052	3.0	0.0157	1.3	0.84	100.2	1.3	101.5	2.8	105	32

18-11*	136	56	0.41	0.0486	2.6	0.1131	8.1	0.0162	1.3	0.92	103.6	1.4	108.2	7.4	126	38
18-12	663	551	0.83	0.0480	2.1	0.1057	4.2	0.0156	1.3	0.71	100.0	1.3	101.8	3.8	99	16
18-13	197	111	0.57	0.0489	2.5	0.1052	1.7	0.0156	1.1	0.28	99.5	1.0	101.5	1.6	141	33
18-14	254	94	0.37	0.0482	2.4	0.1030	1.7	0.0157	1.3	0.67	100.3	1.3	99.6	1.7	108	31
18-15	91	29	0.32	0.0493	3.0	0.1071	2.6	0.0158	1.3	0.46	100.9	1.3	103.3	2.5	158	49
18-16	188	114	0.61	0.0482	2.4	0.1030	1.9	0.0154	1.2	0.69	98.7	1.1	99.5	1.8	106	29
18-17	406	125	0.31	0.0486	2.2	0.1037	1.4	0.0156	1.3	0.69	99.6	1.3	100.1	1.3	123	19
18-18	158	84	0.53	0.0488	2.5	0.1048	2.0	0.0156	1.5	0.71	99.5	1.5	101.2	2.0	129	34
18-19	393	297	0.76	0.0487	2.1	0.1041	1.2	0.0155	1.2	0.74	98.9	1.1	100.5	1.2	136	19
18-20	104	39	0.38	0.0496	2.6	0.1081	2.2	0.0157	1.4	0.62	100.1	1.4	104.2	2.2	172	38
18-21	136	47	0.34	0.0493	2.7	0.1074	2.0	0.0158	1.1	0.58	101.0	1.1	103.5	2.0	160	38
18-22	182	78	0.43	0.0489	2.5	0.1040	1.8	0.0154	1.2	0.57	98.7	1.1	100.4	1.8	142	35
18-23	627	276	0.44	0.0481	2.2	0.1028	1.4	0.0156	1.0	0.74	99.5	1.0	99.4	1.3	103	20

APPENDIX 2: ID-TIMS ZIRCON U-PB DATA FOR THE SHAVER INTRUSIVE SUITE

Fraction	Composition				Isotopic Ratios						Ages (Ma)						Correlation coefficient	
	U (ppm)	Pb (pg) ^a	Th/U ^b	Total common Pb (pg) ^c	²⁰⁶ Pb/ ²⁰⁴ Pb ^d	²⁰⁶ Pb/ ²³⁸ U ^e (2σ%)	Error	²⁰⁷ Pb/ ²³⁵ U ^e (2σ%)	Error	²⁰⁷ Pb/ ²⁰⁶ Pb ^e (2σ%)	Error	²⁰⁶ Pb/ ²³⁸ U ^{f,g} (2σ abs.)	Error	²⁰⁷ Pb/ ²³⁵ U ^g (2σ abs.)	Error	²⁰⁷ Pb/ ²⁰⁶ Pb ^g (2σ abs.)		Error
DC12-02, Dinkey Creek Granodiorite, IGSN: P4700000D																		
F-1	934	102	0.39	6.7	976	0.0157572	0.14	0.103670	1.18	0.047717	1.10	100.88	0.14	100.16	1.12	85.3	26.1	0.574
F-7	1270	48	0.41	1.2	2588	0.0157948	0.10	0.103284	0.49	0.047426	0.45	101.12	0.10	99.80	0.47	70.7	10.6	0.554
F-8	1450	68	0.39	2.5	1714	0.0158352	0.29	0.099666	2.07	0.045648	1.66	101.38	0.29	96.47	1.90	-20.9	40.3	1.306
F-9	155	12	0.42	0.8	921	0.0157376	0.12	0.104017	1.26	0.047958	1.19	100.76	0.12	100.48	1.21	96.1	28.1	0.628
F-10	420	25	0.42	1.6	974	0.0158754	0.18	0.106039	1.48	0.048465	1.38	101.63	0.18	102.34	1.44	121.0	32.5	0.594
F-11	1300	103	0.41	3.0	2176	0.0158494	0.11	0.104361	0.53	0.047777	0.48	101.47	0.11	100.79	0.51	87.2	11.5	0.497
F-12	601	66	0.43	2.8	1493	0.0158322	0.14	0.105184	0.82	0.048206	0.76	101.36	0.14	101.55	0.79	108.3	17.9	0.503
F-14	1160	107	0.42	5.0	1356	0.0158128	0.12	0.103762	0.91	0.047613	0.84	101.23	0.12	100.24	0.87	79.0	20.0	0.598
MG12-03, Mount Givens Granodiorite, IGSN: P4700000E																		
F-1	477	92	0.56	2.3	2395	0.0142404	0.13	0.093833	0.65	0.047790	0.60	91.23	0.12	91.07	0.57	88.9	14.2	0.490
F-2	301	42	0.47	1.7	1558	0.0142628	0.12	0.093665	0.83	0.047629	0.75	91.37	0.11	90.91	0.72	80.9	17.7	0.728
F-3	642	114	0.52	1.4	4978	0.0142711	0.17	0.093787	0.55	0.047663	0.45	91.42	0.15	91.03	0.48	82.6	10.8	0.691
F-5	499	29	0.55	2.7	669	0.0142638	0.14	0.094558	1.71	0.048101	1.62	91.38	0.13	91.74	1.50	103.2	38.3	0.696
DC12-04, Dinkey Creek Granodiorite, IGSN: P4700000F																		
F-1	1010	54	0.41	4.6	756	0.0158257	0.12	0.105108	1.47	0.048169	1.40	101.31	0.12	101.48	1.42	107.6	33.0	0.603
F-2	1660	24	0.44	1.3	1200	0.0158455	0.10	0.103999	0.95	0.047602	0.90	101.44	0.11	100.46	0.91	79.5	21.3	0.587
F-5	448	31	0.34	0.8	2440	0.0158473	0.32	0.105780	1.91	0.048411	1.61	101.45	0.32	102.10	1.85	119.4	37.9	0.957
F-6	914	33	0.35	1.3	1628	0.0157714	0.11	0.104497	0.75	0.048054	0.69	100.97	0.11	100.92	0.72	101.9	16.3	0.566
F-7	197	7	0.40	0.8	579	0.0158801	0.14	0.104530	2.09	0.047740	2.02	101.66	0.14	100.95	2.01	86.4	48.0	0.533
DC12-05, Dinkey Creek Granodiorite, IGSN: P4700000G																		
F-1	771	58	0.39	1.4	2557	0.0158218	0.11	0.104538	0.56	0.047941	0.49	101.29	0.11	100.96	0.54	95.3	11.6	0.701
F-3	768	66	0.39	1.8	2385	0.0158371	0.12	0.105183	0.53	0.048190	0.47	101.39	0.12	101.55	0.51	107.6	11.1	0.556
F-4	990	36	0.40	3.3	709	0.0158316	0.19	0.104193	1.66	0.047754	1.59	101.35	0.19	100.64	1.59	86.0	37.6	0.429
F-5	633	58	0.40	1.6	2287	0.0158266	0.09	0.104866	0.52	0.048056	0.48	101.32	0.09	101.26	0.50	102.0	11.4	0.509
F-6	776	45	0.40	1.6	1801	0.0158659	0.12	0.104640	0.80	0.047833	0.73	101.57	0.12	101.05	0.77	91.0	17.3	0.596
F-7	625	38	0.40	1.9	1318	0.0158543	0.13	0.103623	1.10	0.047403	1.01	101.50	0.14	100.11	1.05	69.6	24.0	0.719

^aTotal mass of radiogenic Pb.

^bTh contents calculated from radiogenic ²⁰⁸Pb and ²³⁰Th-corrected ²⁰⁶Pb/²³⁸U date of the sample, assuming concordance between U-Pb and Th-Pb systems.

^cTotal mass of common Pb.

^dMeasured ratio corrected for fractionation and spike contribution only.

^eMeasured ratios corrected for fractionation, tracer and blank. Pb blank ratios: ²⁰⁶Pb/²⁰⁴Pb = 18.687 ± 0.25; ²⁰⁷Pb/²⁰⁴Pb = 15.658 ± 0.25; and ²⁰⁸Pb/²⁰⁴Pb = 38.258 ± 0.5 (1σ abs.)

^fCorrected for initial Th/U disequilibrium using radiogenic ²⁰⁸Pb and Th/U_[magma], which is assumed to be approximated by the measured whole rock Th/U ratio. Th/U ratios used: 3.60 (DC12-02), 1.79 (MG12-03), 2.86 (DC12-04), 4.13 (DC12-05).

^gIsotopic dates calculated using λ₂₃₈ = 1.55125E⁻¹⁰ (Jaffey et al. 1971) and λ₂₃₅ = 9.8485E⁻¹⁰ (Jaffey et al. 1971).

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