# Theory of radiative and rare $B$ decays 

Gino Isidori

INFN, Laboratori Nazionali di Frascati, Via E. Fermi 40, I-00044 Frascati, Italy


#### Abstract

We present a concise theoretical overview of radiative and rare $B$ decays mediated by flavour-changing neutralcurrent transitions of the type $b \rightarrow s(d) \gamma$ and $b \rightarrow s(d) \bar{\ell} \ell$.


## INTRODUCTION

Thanks to the efforts of $B$ factories, the exploration of the mechanism of quark-flavour mixing is now entering a new interesting era. The precise measurements of mixing-induced CP violation and tree-level allowed semileptonic transition have provided an important consistency check of the SM, and a precise determination of the Cabibbo-Kobayashi-Maskawa (CKM) matrix. The next goal is to understand if there is still room for new physics (NP) or, more precisely, if there is still room for new sources of flavour symmetry breaking close to the electroweak scale. From this perspective, radiative and rare $B$ decays mediated by flavour-changing neutral current (FCNC) amplitudes represent a fundamental tool (see e.g. Ref. [1]).

Beside the experimental sensitivity, the conditions which allow to perform significant NP searches in rare decays can be summarized as follows: i) decay amplitude dominated by electroweak dynamics, and thus enhanced sensitivity to non-standard contributions; ii) small theoretical error within the SM, or good control of both perturbative and non-perturbative corrections. In the rest of this talk we shall analyze at which level these conditions are satisfied in various decay modes.

## INCLUSIVE FCNC $B$ DECAYS

Inclusive rare $B$ decays such as $B \rightarrow X_{s} \gamma, B \rightarrow X_{s} \ell^{+} \ell^{-}$ and $B \rightarrow X_{s} \nu \bar{v}$ are the natural framework for highprecision studies of FCNCs in the $\Delta B=1$ sector [2]. Perturbative QCD and heavy-quark expansion form a solid theoretical framework to describe these processes: inclusive hadronic rates are related to those of free $b$ quarks, calculable in perturbation theory, by means of a systematic expansion in inverse powers of the $b$-quark mass.

The starting point of the perturbative partonic calculation is the determination of a low-energy effective Hamil-
tonian, renormalized at a scale $\mu=\mathscr{O}\left(m_{b}\right)$, obtained by integrating out the heavy degrees of freedom of the theory. For $b \rightarrow s$ transitions -within the SM- this can be written as

$$
\begin{equation*}
\mathscr{H}_{\mathrm{eff}}=-\frac{G_{F}}{\sqrt{2}} V_{t s}^{*} V_{t b} \sum_{i=1}^{10, v} C_{i}(\mu) Q_{i}+\text { h.c. } \tag{1}
\end{equation*}
$$

where $Q_{1 \ldots 6}$ are four-quark operators, $Q_{8}$ is the chromomagnetic operator and

$$
\begin{align*}
Q_{7} & =\frac{e}{4 \pi^{2}} \bar{s}_{L} \sigma_{\mu v} m_{b} b_{R} F^{\mu v} \\
Q_{9} & =\frac{e^{2}}{4 \pi^{2}} \bar{s}_{L} \gamma^{\mu} b_{L} \bar{\ell} \gamma_{\mu} \ell \\
Q_{10} & =\frac{e^{2}}{4 \pi^{2}} \bar{s}_{L} \gamma^{\mu} b_{L} \bar{\ell} \gamma_{\mu} \gamma_{5} \ell \\
Q_{v} & =\frac{e^{2}}{4 \pi^{2} s_{w}^{2}} \bar{s}_{L} \gamma^{\mu} b_{L} \bar{v}_{L} \gamma_{\mu} v_{L} \tag{2}
\end{align*}
$$

are the leading FCNC electroweak operators. Within the SM, the coefficients of all the operators in Eq. (2) receive a large non-decoupling contribution from topquark loops at the electroweak scale. But the $m_{t}$ dependence is not the same for the four operators, reflecting a different $S U(2)_{L}$-breaking structure, which can be affected in a rather different way by new-physics contributions [3, 4].

The calculation of the rare decay rates then involves three distinct steps: i) the determination of the initial conditions of the Wilson coefficients at the electroweak scale; ii) the evolution by means of renormalizationgroup equations (RGEs) of the $C_{i}$ down to $\mu=\mathscr{O}\left(m_{b}\right)$; iii) the evaluation of the hadronic matrix elements of the effective operators at $\mu=\mathscr{O}\left(m_{b}\right)$, including both perturbative and non-perturbative QCD corrections. Each of the three steps must be taken to matching orders of accuracy in powers of the strong coupling constants $\alpha_{s}$ and of the large logs generated by the RGE running. The interesting short-distance (electroweak) dynamics that we


FIGURE 1. Representative diagrams for the mixing of fourquark operators into $Q_{7}$ (left) and $Q_{9}$ (right).
would like to test enters only in the first step; the following two steps are fundamental ingredients to reduce and control the theoretical error.

The first two steps (initial conditions and RGEs) are process independent and are common also to exclusive modes. Nonetheless, the organization of the leading-log (LL) series is not the same for the three underlying partonic processes, or the four operators in Eq. (2):
$b \rightarrow s \gamma$. Here only $Q_{7}$ has a non-vanishing matrix element at the tree level. The large logarithms generated by mixing of four-quark operators into $Q_{7}$ (see Fig. 1 play a very important role and enhance the partonic rate by a factor of almost three [5]. Since this mixing vanishes at the one-loop level, a full treatment of QCD corrections beyond lowest order is a rather non-trivial task. This has been achieved already a few years ago, thanks to the joint effort of many authors (see e.g. Ref. [2, 6] and references therein), and is nowadays a rather mature subject. All the ingredients of the partonic calculation have been cross-checked by more than one group. In particular, very recently an independent confirmation of the three-loop mixing of $Q_{7}$ and $Q_{1 \ldots 6}$ [7] - till few months ago the only piece of the calculation performed by one group only - has been presented [8].
$b \rightarrow s \ell^{+} \ell^{-}$. The three operators with non-vanishing tree-level matrix elements are $Q_{7}, Q_{9}$ and $Q_{10}$. Similarly to $Q_{7}, \mathrm{QCD}$ corrections are very important also for $Q_{9}$. Since $Q_{9}$ mixes with four-quark operators already at the one-loop level (see Fig. [1), the organization of the LL series for $b \rightarrow s \ell^{+} \ell^{-}$is different than in $b \rightarrow s \gamma$ : the NLL level is much simpler (no three-loop mixing involved), but less precise. An accuracy below the $10 \%$ level on the decay rate (a precision similar to the NLL level in $b \rightarrow s \gamma$ ) is reached here only with a NNLL calculation. All the missing ingredients to reach this goal has finally become available. In particular, NNLL initial conditions and anomalous dimension matrix can be found in [9] and [8], respectively.

It is worth to stress that the impact of QCD corrections is very limited in the axial-current operator $Q_{10}$. This operator does not mix with four-quark operators and is completely dominated by short-distance contri-
butions. Together with $Q_{V}, Q_{10}$ belongs to the theoretically clean $\mathscr{O}\left(G_{F}^{2}\right)$ hard-GIM-protected part of the effective Hamiltonian. Thus observables more sensitive to $Q_{10}$, such as the forward-backward (FB) lepton asymmetry in $b \rightarrow s \ell^{+} \ell^{-}$, have a reduced QCD uncertainty and a stronger sensitivity to possible non-standard phenomena.
$b \rightarrow s V \bar{v}$. In this case only $Q_{v}$ is involved. Similarly to $Q_{10}$, QCD corrections play a very minor role since there is no mixing with four-quark operators. As a result, the only non-trivial step of the perturbative calculation for $b \rightarrow s v \bar{v}$ decays is the determination of the initial condition of $C_{V}$ at the electroweak scale: this is known with a precision around $1 \%$ within the SM 10].

These two processes-independent steps of the calculation, namely the determination of the effective Hamiltonian renormalized at a low scale $\mu=\mathscr{O}\left(m_{b}\right)$, can easily be transferred from the $b \rightarrow s$ case to the $b \rightarrow d$ one. The only difference is the richer CKM structure of the $b \rightarrow d$ Hamiltonian, with two independent non-negligible terms $\left(V_{t d}^{*} V_{t b}\right.$ and $\left.V_{u d}^{*} V_{u b}\right)$.

The situation is very different for the last step of the calculation, namely the evaluation of the hadronic matrix elements. The latter strongly depend on the specific process and the specific observable we are interested in (e.g. fully inclusive rate or differential distribution). In the following we shall review the results of this step (and thus the final numerical predictions) for some of the most interesting $b \rightarrow s$ observables.

## $B \rightarrow X_{s} \gamma$ [the most effective " $N P$ killer"]

The inclusive $B \rightarrow X_{s} \gamma$ rate is the most precise and clean short-distance information that we have, at present, on $\Delta B=1$ FCNCs. Combining the precise measurements by ALEPH, BaBar, Belle and CLEO, the world average reads [11]

$$
\begin{equation*}
\mathscr{B}\left(B \rightarrow X_{s} \gamma\right)^{\exp }=(3.34 \pm 0.38) \times 10^{-4} \tag{3}
\end{equation*}
$$

On the theory side, the NLL partonic calculation of the matrix elements, performed first in Ref. [12] for the leading terms, has recently been cross-checked and completed in Ref. [13]. Perturbative corrections due to higher-order electroweak effects have also been analyzed (see Ref. [14] and references therein).

Non-perturbative $1 / m_{b}$ corrections are well under control in the total rate. In particular, $\mathscr{O}\left(1 / m_{b}\right)$ corrections vanish in the ratio $\Gamma\left(B \rightarrow X_{s} \gamma\right) / \Gamma\left(B \rightarrow X_{c} \ell v\right)$, and the $\mathscr{O}\left(1 / m_{b}^{2}\right)$ ones are known and amount to few per cent [15]. Also non-perturbative effects associated to charmquark loops have been estimated and found to be very small [16, 17, 18]. The most serious problem of nonperturbative origin is related to the (unavoidable) experimental cut in the photon energy spectrum that prevents
the measurement from being fully inclusive [17, 19]. With the present cut by CLEO $E_{\gamma}>2.0 \mathrm{GeV}$ [20], this uncertainty is smaller but non-negligible with respect to the error of the perturbative calculation. The latter is around $10 \%$ and its main source is the uncertainty in the ratio $m_{c} / m_{b}$ that enters through charm-quark loops [21].

According to the detailed analysis of theoretical errors presented in Ref. [21], the SM expectation is

$$
\begin{equation*}
\mathscr{B}\left(B \rightarrow X_{s} \gamma\right)^{\mathrm{SM}}=(3.73 \pm 0.30) \times 10^{-4} \tag{4}
\end{equation*}
$$

in good agreement with Eq. (3). It must be stressed that the overall scale dependence is very small: for $\mu \in$ $\left[m_{b} / 2,2 m_{b}\right]$ the central value moves by about $1 \%$. The error in Eq. (4) is an educated guess about the size of possible NNLL terms. In particular, the largest source of uncertainty is obtained by the variation of $\bar{m}_{c}(\mu) / m_{b}^{\text {pole }}$ for $\mu \in\left[m_{c}, m_{b}\right]$. A critical discussion about the error in Eq. (4), with alternative more conservative estimates, can be found in Ref. [22].

The comparison between theory and experiments in $\mathscr{B}\left(B \rightarrow X_{s} \gamma\right)$ is a great success of the SM and has led us to derive many significant bounds on possible newphysics scenarios. For instance, the $\mathscr{B}\left(B \rightarrow X_{s} \gamma\right)$ constraint is one of the main obstacles to build consistent models that predict a sizable difference between $\mathscr{A}_{\mathrm{CP}}\left(B \rightarrow \phi K_{S}\right)$ and $\mathscr{A}_{\mathrm{CP}}\left(B \rightarrow \psi K_{S}\right)$. This constraint can be avoided (see e.g. Ref. [23] and references there in), but the resulting models require a considerable amount of fine tuning. By far more natural are the so-called MFV models [4], where $\mathscr{A}_{\mathrm{CP}}\left(B \rightarrow \phi K_{S}\right) \approx \mathscr{A}_{\mathrm{CP}}\left(B \rightarrow \psi K_{S}\right)$ and deviations from the SM in $\mathscr{B}\left(B \rightarrow X_{s} \gamma\right)$ do not exceed the $10 \%-30 \%$ level [4]. Improved measurements of $\mathscr{B}\left(B \rightarrow X_{s} \gamma\right)$ are certainly useful to further constrain this possibility. However, since the experimental error has reached the level of the theoretical one, it will be very difficult to clearly identify possible deviations from the SM, if any, in this observable.

Hopes to detect new-physics signals are still open through the CP -violating asymmetry

$$
\begin{equation*}
\Delta \Gamma_{\mathrm{CP}}\left(B \rightarrow X_{s} \gamma\right)=\frac{\Gamma\left(B \rightarrow X_{s} \gamma\right)-\Gamma\left(\bar{B} \rightarrow X_{s} \gamma\right)}{\Gamma\left(B \rightarrow X_{s} \gamma\right)+\Gamma\left(\bar{B} \rightarrow X_{s} \gamma\right)} \tag{5}
\end{equation*}
$$

This is expected to be below $1 \%$ within the $\operatorname{SM}$ [22, 24], but could easily reach $\mathscr{O}(10 \%)$ values beyond the SM, even in the absence of large effects in the total $B \rightarrow X_{s} \gamma$ rate. This is indeed one of the main expectations in models with sizable NP effects in $\mathscr{A}_{\mathrm{CP}}\left(B \rightarrow \phi K_{S}\right)$. The present measurement of $\Delta \Gamma_{\mathrm{CP}}\left(B \rightarrow X_{s} \gamma\right)$ is consistent with zero [11], but the sensitivity is still one order of magnitude above the SM level.

## $B \rightarrow X_{s} \ell^{+} \ell^{-}$[the present frontier]

Both Belle [25] and BaBar [26] have recently announced a clear evidence $(\approx 5 \sigma)$ of the $B \rightarrow X_{s} \ell^{+} \ell^{-}$decay. The two results are compatible and are both based on a semi-inclusive analysis (the hadronic system is reconstructed from a kaon plus 0 to 4 pions, with at most one $\pi^{0}$ ). Their combination [11]

$$
\begin{equation*}
\mathscr{B}\left(B \rightarrow X_{s} \ell^{+} \ell^{-}\right)^{\exp }=\left(6.2 \pm 1.1_{-1.3}^{+1.6}\right) \times 10^{-6} . \tag{6}
\end{equation*}
$$

represents a very useful new piece of information about $\Delta B=1$ FCNCs, with considerable margin of improvement in the near future.

In principle, these decays offer a phenomenology reacher than $B \rightarrow X_{s} \gamma$, with more than one interesting observable. The joint effort of several groups has recently allowed to evaluate at the NNLL level all the matrix element necessary for the two main kinematical distributions: the dilepton spectrum [27, 28, 29] and the lepton FB asymmetry [30, 31].

In addition to the non-perturbative corrections due to the finite $b$ quark mass, $B \rightarrow X_{s} \ell^{+} \ell^{-}$transitions suffer from specific non-perturbative effects due to long-lived $c \bar{c}$ intermediate states ( $B \rightarrow X_{s} c \bar{c} \rightarrow X_{s} \ell^{+} \ell^{-}$). The heavyquark expansion, which allow to evaluate the $\Lambda_{\mathrm{QCD}} / m_{b}$ terms, is rapidly convergent and leads to small corrections for sufficiently inclusive observables [32, 33]. A consistent treatment of the second type of effects requires kinematical cuts in order to avoid the large nonperturbative background of the narrow $c \bar{c}$ resonances (see Fig. 2). These two requirements are somehow in conflict [28, 33]; nonetheless, we can identify two perturbative windows, defined by:

$$
\begin{array}{ll}
q^{2} \equiv M_{\ell^{+} \ell^{-}}^{2} \in\left[1 \mathrm{GeV}^{2}, 6 \mathrm{GeV}^{2}\right] & \text { (low) } \\
q^{2}>14.4 \mathrm{GeV}^{2} & \text { (high) }
\end{array}
$$

where reliable predictions can be performed [28]. It is worth to emphasize that the two regions have complementary virtues and disadvantages:

- Virtues of the low- $q^{2}$ region: reliable $q^{2}$ spectrum; small $1 / m_{b}$ corrections; sensitivity to the interference of $C_{7}$ and $C_{9}$; high rate.
- Disadvantages of the low- $q^{2}$ region: difficult to perform a fully inclusive measurement (severe cuts on the dilepton energy and/or the hadronic invariant mass); long-distance effects due to processes of the type $B \rightarrow \Psi X_{s} \rightarrow X_{s}+X^{\prime} \ell^{+} \ell^{-}$not fully under control; non-negligible scale and $m_{c}$ dependence.
- Virtues of the high-q ${ }^{2}$ region: negligible scale and $m_{c}$ dependence due to the strong sensitivity to $\left|C_{10}\right|^{2}$; easier to perform a fully inclusive measurement (small hadronic invariant mass); negligible long-distance effects of the type $B \rightarrow \Psi X_{s} \rightarrow$ $X_{s}+X^{\prime} \ell^{+} \ell^{-}$.


FIGURE 2. Dilepton spectrum of the inclusive $B \rightarrow X_{s} e^{+} e^{-}$ decay within the SM. Vertical axis: $d \mathscr{B}\left(B \rightarrow X_{s} e^{+} e^{-}\right) / d q^{2}$ in units of $10^{-7} \times \mathrm{GeV}^{-2}$; horizontal axis: $q^{2}$ in $\mathrm{GeV}^{2}$. The dotted line denotes the NNLL pure perturbative result, the full line includes an estimates of the non-perturbative $c \bar{c}$ effects [28].

- Disadvantages of the high- $q^{2}$ region: $q^{2}$ spectrum not reliable (only the integrated rate can be predicted); sizeable $1 / m_{b}$ corrections (effective expansion in $1 /\left(m_{b}-\sqrt{q_{\text {min }}}\right)$ [28] $)$; low rate.
Given this situation, we believe that future experiments should try to measure the branching ratios in both regions and report separately the two results. These two measurements are indeed affected by different systematic uncertainties (of theoretical nature) and provide a different short-distance information. The NNLL SM predictions for the two clean windows [28],

$$
\begin{gather*}
\mathscr{B}\left(B \rightarrow X_{S} \ell^{+} \ell^{-}\right)_{\text {low }}^{\mathrm{SM}}=(1.63 \pm 0.20) \times 10^{-6} \\
\mathscr{B}\left(B \rightarrow X_{s} \ell^{+} \ell^{-}\right)_{\text {high }}^{\mathrm{SM}}=(4.04 \pm 0.78) \times 10^{-7}, \tag{7}
\end{gather*}
$$

are still affected by a considerable error; however, in both cases the uncertainty is mainly of parametric nature and could be substantially improved in the future. In particular, the large error in the high $-q^{2}$ region is mainly due to the uncertainty in the relation between the physical $q^{2}$ interval and the corresponding interval for the partonic calculation (i.e. the uncertainty in the relation between $m_{b}$ and the physical hadron mass), which can be improved with better data on charged-current semileptonic modes. According to the recent analysis of Ref. [29], both results in (7) should be decreased by $\approx 4 \%$ to take into account the leading electroweak corrections.

The two results in (7) cannot be directly confronted with (6), which includes an extrapolation to the full $q^{2}$ spectrum. The updated SM expectation for this extrapolated branching ratio is $\left.(4.6 \pm 0.8) \times 10^{-6} \mid 28\right]$ (in agreement with the previous estimate of Ref. [34]), and it is consistent with the present experimental world average.

We stress that this prediction is already saturated by irreducible theoretical errors and, contrary to the results in (7), is very difficult to improve it further.

As anticipated, some of the most interesting shortdistance tests in $\mathscr{B}\left(B \rightarrow X_{s} \ell^{+} \ell^{-}\right)$decays can be performed by means of the FB asymmetry of the dilepton distribution:

$$
\begin{align*}
\mathscr{A}_{\mathrm{FB}}\left(q^{2}\right)= & \frac{1}{d \mathscr{B}\left(B \rightarrow X_{s} \ell^{+} \ell^{-}\right) / d q^{2}} \int_{-1}^{1} d \cos \theta_{\ell} \\
& \frac{d^{2} \mathscr{B}\left(B \rightarrow X_{s} \ell^{+} \ell^{-}\right)}{d q^{2} d \cos \theta_{\ell}} \operatorname{sgn}\left(\cos \theta_{\ell}\right) \tag{8}
\end{align*}
$$

where $\theta_{\ell}$ is the angle between $\ell^{+}$and $B$ momenta in the dilepton centre-of-mass frame. Here the SM predict a zero for $s_{0}=q_{0}^{2} / m_{b}^{2}=0.162 \pm 0.008[30,31]$ : a very precise prediction which could easily be modified beyond the SM, even in absence of significant non-standard effects on the total rate.

## EXCLUSIVE MODES

On general grounds, theoretical predictions for exclusive FCNC decays are more difficult. The simplest cases are processes with at most one hadron in the final state. Here there has been a substantial progress in the last few years, both by means of analytic approaches [35] and by means of Lattice QCD [36], but still the overall theoretical uncertainty is around $20 \%$ at the amplitude level. The largest source of uncertainty is typically the normalization of the hadronic form factors, an error that can be substantially reduced in appropriate ratios or differential distributions. These type of observables become particularly interesting in channels where, because of irreducible experimental problems, the short-distance amplitude cannot be extracted from corresponding inclusive modes. Two of such examples are the ratio

$$
\begin{equation*}
R_{\gamma}\left(\rho / K^{*}\right)=\frac{\mathscr{B}(B \rightarrow \rho \gamma)}{\mathscr{B}\left(B \rightarrow K^{*} \gamma\right)} \tag{9}
\end{equation*}
$$

and the normalized FB asymmetry in $B \rightarrow K^{*} \ell^{+} \ell^{-}$.
The ratio $R_{\gamma}\left(\rho / K^{*}\right)$ is one of the most promising tool to extract short-distance properties about the $b \rightarrow s \gamma$ amplitude. On the experimental side, the combination of the bounds on charged and neutral channels, in the isospin limit, leads to $R_{\gamma}\left(\rho / K^{*}\right)<0.047$ at $90 \%$ C.L. [37]. On the theory side, the $B \rightarrow V \gamma$ amplitudes which determine this ratio have been analyzed beyond naive factorization by several authors [38, 39, 40]. Within the SM one can write

$$
\begin{equation*}
R_{\gamma}\left(\rho / K^{*}\right)=\left|\frac{V_{t d}}{V_{t s}}\right|^{2} \frac{\left(m_{B}^{2}-m_{r h o}^{2}\right)^{3}}{\left(m_{B}^{2}-m_{K^{*}}^{2}\right)^{3}} \zeta^{2}(1-\Delta R), \tag{10}
\end{equation*}
$$



FIGURE 3. Bounds on the $\rho-\eta$ plane from the BaBar bound on $R_{\gamma}\left(\rho / K^{*}\right)$ : the two curves correspond to different values of $\zeta$ [39].
where $\zeta$ denotes the ratio of the form factors at $q^{2}=0$ in the $m_{b} \rightarrow \infty$ limit, and $\Delta R$ the additional $S U(3)$ (and isospin) breaking due to $1 / m_{b}$ and $\mathscr{O}\left(\alpha_{s}\right)$ corrections. The largest source of uncertainty is $\zeta$ : according to the light-cone sum rule estimate $\zeta=0.76 \pm 0.10$ [39], the present experimental bound on $R_{\gamma}\left(\rho / K^{*}\right)$ is about twice the SM expectation. However, preliminary Lattice results indicate a larger value, $\zeta=0.91 \pm 0.08$ [41], which would imply more stringent bounds on $\left|V_{t d} / V_{t s}\right|$. The two curves in Fig. 3 summarizes the present status.

While the inclusive $B \rightarrow X_{S} \ell^{+} \ell^{-}$rate is already accessible at the $B$ factories, a differential study of the inclusive FB asymmetry is beyond their present and near-future reach. More accessible from the experimental point of view is the FB asymmetry in $B \rightarrow K^{*} \ell^{+} \ell^{-}$ (defined as in Eq. 8) with $X_{s} \rightarrow K^{*}$ ). Assuming that the leptonic current has only a vector $(V)$ or axial-vector $(A)$ structure (as in the SM), the FB asymmetry provides a direct measure of the $A-V$ interference. Indeed, at the lowest-order one can write

$$
\mathscr{A}_{\mathrm{FB}}\left(q^{2}\right) \propto \operatorname{Re}\left\{C_{10}^{*}\left[\frac{q^{2}}{m_{b}^{2}} C_{9}^{\mathrm{eff}}+r\left(q^{2}\right) \frac{m_{b} C_{7}}{m_{B}}\right]\right\}
$$

where $r\left(q^{2}\right)$ is an appropriate ratio of $B \rightarrow K^{*}$ vector and tensor form factors 42]. There are three main features of this observable that provide a clear and independent short-distance information: 1) The position of the zero of $\mathscr{A}_{\mathrm{FB}}\left(q^{2}\right)$ in the low- $q^{2}$ region (see Fig. 4] [42]. As shown by means of a full NLO calculation [40], the experimental measurement of $q_{0}^{2}$ could allow a determination of $C_{7} / C_{9}$ at the $10 \%$ level. 2) The sign of $\mathscr{A}_{\mathrm{FB}}\left(q^{2}\right)$ around the zero. This is fixed unambiguously in terms of the relative sign of $C_{10}$ and $C_{9}$ [43]: within the SM one expects $\mathscr{A}_{\mathrm{FB}}\left(q^{2}>q_{0}^{2}\right)>0$ for $|\bar{B}\rangle \equiv|b \bar{d}\rangle$ mesons. 3) The relation $\mathscr{A}[\bar{B}]_{\mathrm{FB}}\left(q^{2}\right)=-\mathscr{A}[B]_{\mathrm{FB}}\left(q^{2}\right)$. This follows from the CP-odd structure of $\mathscr{A}_{\mathrm{FB}}$ and holds at the $10^{-3}$ level within the SM [43], where $C_{10}$ has a negligible CP-violating phase.


FIGURE 4. Zero of the forward-backward asymmetry in $B^{-} \rightarrow K^{*-} \ell^{+} \ell^{-}$at LO and NLO. The band reflects all theoretical uncertainties from parameters and scale dependence combined [40].

$$
B_{s, d} \rightarrow \ell^{+} \ell^{-} \quad[\text { the future frontier }]
$$

The purely leptonic decays constitute a special case among exclusive transitions. Within the SM only the axial-current operator, $Q_{10}$, induces a non-vanishing contribution to these decays. As a result, the short-distance contribution is not diluted by the mixing with four-quark operators. Moreover, the hadronic matrix element involved is the simplest we can consider, namely the $B$ meson decay constant

$$
\begin{equation*}
\langle 0| \bar{q} \gamma_{\mu} \gamma_{5} b\left|\bar{B}_{q}(p)\right\rangle=i p_{\mu} f_{B_{q}} \tag{11}
\end{equation*}
$$

Reliable estimates of $f_{B_{d}}$ and $f_{B_{s}}$ are obtained at present from lattice calculations and in the future it will be possible to cross-check these results by means of the $B^{+} \rightarrow$ $\ell^{+} v$ rate. Modulo the determination of $f_{B_{q}}$, the theoretical cleanliness of $B_{s, d} \rightarrow \ell^{+} \ell^{-}$decays is comparable to that of the golden modes $K_{L} \rightarrow \pi^{0} v \bar{v}$ and $B \rightarrow X_{s, d} v \bar{v}$.

The price to pay for this theoretically-clean amplitude is a strong helicity suppression for $\ell=\mu$ (and $\ell=e$ ), or the channels with the best experimental signature. Employing the full NLO expression of $C_{10}$ [10], we can write

$$
\begin{gathered}
\mathscr{B}\left(B_{s} \rightarrow \mu^{+} \mu^{-}\right)^{\mathrm{SM}}=3.1 \times 10^{-9}\left(\frac{\left|V_{t s}\right|}{0.04}\right)^{2} \\
\times\left(\frac{f_{B_{s}}}{0.21 \mathrm{GeV}}\right)^{2}\left(\frac{\tau_{B_{s}}}{1.6 \mathrm{ps}}\right)\left(\frac{m_{t}\left(m_{t}\right)}{166 \mathrm{GeV}}\right)^{3.12} \\
\frac{\mathscr{B}\left(B_{s} \rightarrow \tau^{+} \tau^{-}\right)^{\mathrm{SM}}}{\mathscr{B}\left(B_{s} \rightarrow \mu^{+} \mu^{-}\right)^{\mathrm{SM}}}=215
\end{gathered}
$$

The corresponding $B_{d}$ modes are both suppressed by an additional factor $\left|V_{t d} / V_{t s}\right|^{2}=(4.0 \pm 0.8) \times 10^{-2}$. The present experimental bound closest to SM expectations is the one obtained by CDF on $B_{s} \rightarrow \mu^{+} \mu^{-}$[11, 44]:

$$
\mathscr{B}\left(B_{s} \rightarrow \mu^{+} \mu^{-}\right)<9.5 \times 10^{-7} \quad(95 \% \mathrm{CL}),
$$

which is still very far from the SM level. The latter will certainly not be reached before the LHC era.

As emphasized in the recent literature [45, 46, 47], the purely leptonic decays of $B_{s}$ and $B_{d}$ mesons are excellent probes of several new-physics models and, particularly, of scalar FCNCs. Scalar FCNC operators, such as $\bar{b}_{R} s_{L} \bar{\mu}_{R} \mu_{L}$, are present within the SM but are absolutely negligible because of the smallness of downtype Yukawa couplings. On the other hand, these amplitudes could be non-negligible in models with an extended Higgs sector. In particular, within the MSSM, where two Higgs doublets are coupled separately to upand down-type quarks, a strong enhancement of scalar FCNCs can occur at large $\tan \beta=v_{u} / v_{d}$ [45]. This effect is very small in non-helicity-suppressed $B$ decays and in $K$ decays (because of the small Yukawa couplings), but could enhance $B \rightarrow \ell^{+} \ell^{-}$rates by orders of magnitude. As pointed out in Ref. [48], $\mathscr{O}(100)$ enhancements in $\mathscr{B}\left(B \rightarrow \ell^{+} \ell^{-}\right)$correspond to $\mathscr{O}(10 \%)$ breaking of universality in $\mathscr{B}\left(B \rightarrow K \mu^{+} \mu^{-}\right)$vs. $\mathscr{B}\left(B \rightarrow K e^{+} e^{-}\right)$. Therefore, the present search for $B \rightarrow \ell^{+} \ell^{-}$at CDF is already quite interesting, even if the sensitivity is well above the SM level. In a long-term perspective, the discovery of such processes is definitely one of the most interesting items in the $B$-physics program of hadron colliders.

## CONCLUSIONS

Rare FCNC decays of $B$ mesons provide a unique opportunity to perform high-precision studies of quarkflavour mixing. The $B \rightarrow X_{s} \gamma$ rate, where both experimental and theoretical errors have reached a comparable level around $10 \%$, represents the highest peak in our present knowledge of FCNCs. The lack of deviations from SM expectations in $\Gamma\left(B \rightarrow X_{s} \gamma\right)$ should not discourage the measurement of other clean and independent FCNC observables, such as the forward-backward asymmetry in $B \rightarrow X_{s} \ell^{+} \ell^{-}$or the $B \rightarrow \ell^{+} \ell^{-}$rates. Even if new physics will first be discovered elsewhere, the experimental study of these theoretically-clean observables would still be very useful to investigate the flavour structure of any new-physics scenario.

## Acknowledgements

I am grateful to the organizers of Beauty 2003 for the invitation and the financial support that allowed me to attend this interesting conference. This work is partially supported by the EC-Contract HPRN-CT-2002-00311 (EURIDICE).

## REFERENCES

1. A. J. Buras, hep-ph/0307203 Y. Nir, Nucl. Phys. Proc. Suppl., 117, 111 (2003) |hep-ph/0208080|; G. Isidori, Int. J. Mod. Phys. A, 17, 3078 (2002) |hep-ph/0110255|.
2. For a recent review see T. Hurth, Rev. Mod. Phys., 75, 1159 (2003) |hep-ph/0212304|.
3. S. Bertolini, F. Borzumati and A. Masiero, Phys. Lett. B, 192, 437 (1987); A. Ali, G. F. Giudice and T. Mannel, Z. Phys. C, 67, 417 (1995) |hep-ph/9408213|.
4. G. D'Ambrosio et al., Nucl. Phys. B, 645, 155 (2002) |hep-ph/0207036|; A. J. Buras et al., Phys. Lett. B, 500, 161 (2001) |hep-ph/0007085|; A. J. Buras, Acta Phys. Polon. B, 34, 5615 (2003) |hep-ph/0310208|.
5. S. Bertolini, F. Borzumati, A. Masiero, Phys. Rev. Lett., 59, 180 (1987); N. G. Deshpande et al., Phys. Rev. Lett., 59, 183 (1987).
6. A. J. Buras and M. Misiak, Acta Phys. Polon. B, 33, 2597 (2002) hep-ph/0207131.
7. K. Chetyrkin, M. Misiak and M. Munz, Phys. Lett. B, 400, 206 (1997); 425, 414 (1997) (E) |hep-ph/9612313|.
8. P. Gambino, M. Gorbahn and U. Haisch, Nucl. Phys. B, 673, 238 (2003) |hep-ph/0306079|.
9. C. Bobeth, M. Misiak and J. Urban, Nucl. Phys. B, 574, 291 (2000) |hep-ph/9910220|.
10. G. Buchalla and A. J. Buras, Nucl. Phys. B, 548, 309 (1999) |hep-ph/9901288|; M. Misiak and J. Urban, Phys. Lett. B, 451, 161 (1999) hep-ph/9901278.
11. M. Nakao, hep-ex/0312041
12. C. Greub, T. Hurth and D. Wyler, Phys. Rev. D, 54, 3350 (1996) hep-ph/9603404.
13. A. J. Buras, A. Czarnecki, M. Misiak and J. Urban, Nucl. Phys. B, 631, 219 (2002) hep-ph/0203135; 611, 488 (2001) |hep-ph/0105160|.
14. P. Gambino and U. Haisch, JHEP, 0110, 020 (2001) |hep-ph/0109058|.
15. A.F. Falk, M. Luke and M.J. Savage, Phys. Rev. D, 49, 3367 (1994) hep-ph/9308288.
16. M. B. Voloshin, Phys. Lett. B, 397, 275 (1997) |hep-ph/9612483|; A. Grant, A. Morgan, S. Nussinov and R. Peccei, Phys. Rev. D, 56, 3151 (1997) |hep-ph/9702380|.
17. Z. Ligeti, L. J. Randall and M. B. Wise, Phys. Lett. B, 402, 178 (1997) |hep-ph/9702322|.
18. G. Buchalla, G. Isidori and S. J. Rey, Nucl. Phys. B, 511, 594 (1998) |hep-ph/9705253|.
19. A. L. Kagan and M. Neubert, Eur. Phys. J. C, 7, 5 (1999) [hep-ph/9805303].
20. D. Cassel, these proceedings; S. Chen et al. [CLEO Collab.], hep-ex/0108032
21. P. Gambino and M. Misiak, Nucl. Phys. B, 611, 338 (2001) |hep-ph/0104034|.
22. T. Hurth, E. Lunghi and W. Porod, hep-ph/0312260
23. M. Ciuchini et al., eConf C0304052 (2003) WG307 |hep-ph/0308013|.
24. A. L. Kagan and M. Neubert, Phys. Rev. D, 58, 094012 (1998) |hep-ph/9803368];
25. J. Kaneko et al. [Belle Collaboration], Phys. Rev. Lett., 90, 021801 (2003) |hep-ex/0208029|.
26. B. Aubert et al. [BaBar Collaboration], hep-ex/0308016
27. H. H. Asatrian, H. M. Asatrian, C. Greub and M. Walker, Phys. Lett. B, 507, 162 (2001) |hep-ph/0103087|; Phys. Rev. D, 65, 074004 (2002) hep-ph/0109140|; Phys.

Rev. D, 66, 034009 (2002) hep-ph/0204341.
28. A. Ghinculov, T. Hurth, G. Isidori and Y. P. Yao, hep-ph/0312128, hep-ph/0310187 Nucl. Phys. Proc. Suppl., 116, 284 (2003) |hep-ph/0211197|.
29. C. Bobeth, P. Gambino, M. Gorbahn and U. Haisch, hep-ph/0312090
30. A. Ghinculov, T. Hurth, G. Isidori and Y. P. Yao, Nucl. Phys. B, 648, 254 (2003) |hep-ph/0208088|.
31. H. M. Asatrian, K. Bieri, C. Greub and A. Hovhannisyan, Phys. Rev. D, 66, 094013 (2002) |hep-ph/0209006|; H. M. Asatrian, H. H. Asatryan, A. Hovhannisyan and V. Poghosyan, hep-ph/0311187
32. A. Ali, G. Hiller, L. T. Handoko and T. Morozumi, Phys. Rev. D, 55, 4105 (1997) hep-ph/9609449|.
33. G. Buchalla and G. Isidori, Nucl. Phys. B, 525, 333 (1998) hep-ph/9801456|.
34. A. Ali, E. Lunghi, C. Greub and G. Hiller, Phys. Rev. D, 66, 034002 (2002) |hep-ph/0112300|.
35. I. Stewart, these proceedings.
36. A. El-Khadra, these proceedings.
37. B. Aubert et al. [BaBar Collaboration], hep-ex/0207073
38. B. Grinstein and D. Pirjol, Phys. Rev. D, 62, 093002 (2000) hep-ph/0002216|; A. Ali and A. Y. Parkhomenko, Eur. Phys. J. C, 23, 89 (2002) |hep-ph/0105302|. S. W. Bosch and G. Buchalla, Nucl. Phys. B, 621, 459 (2002) [hep-ph/0106081.
39. A. Ali and E. Lunghi, Eur. Phys. J. C, 26, 195 (2002) |hep-ph/0206242|; T. Hurth and E. Lunghi, eConf C0304052 (2003) WG206 hep-ph/0307142|.
40. M. Beneke, T. Feldmann and D. Seidel, Nucl. Phys. B, 612, 25 (2001) hep-ph/0106067|.
41. D. Becirevic and J. Reyes, hep-lat/0309131
42. G. Burdman, Phys. Rev. D, 57, 4254 (1998) |hep-ph/9710550|.
43. G. Buchalla, G. Hiller and G. Isidori, Phys. Rev. D, 63, 014015 (2001) hep-ph/0006136|.
44. C.-J. Lin, these proceedings.
45. C. Hamzaoui, M. Pospelov and M. Toharia; Phys. Rev. D, 59, 095005 (1999) hep-ph/9807350|; K. S. Babu and C. Kolda, Phys. Rev. Lett., 84, 228 (2000) |hep-ph/9909476; G. Isidori and A. Retico, JHEP, 0111, 001 (2001) hep-ph/0110121; A. J. Buras et al., Nucl. Phys. B, 659, 3 (2003) hep-ph/0210145|; A. Dedes and A. Pilaftsis, Phys. Rev. D, 67, 015012 (2003) |hep-ph/0209306|; A. Dedes, Mod. Phys. Lett. A, 18, 2627 (2003) |hep-ph/0309233|;
46. C. S. Huang et al. Phys. Rev. D, 59, 011701 (1999) |hep-ph/9803460|; Phys. Rev. D, 63, 114021 (2001) |hep-ph/0006250|; S. R. Choudhury and N. Gaur, Phys. Lett. B, 451, 86 (1999) |hep-ph/9810307|; P.H. Chankowski and L. Slawianowska, Phys. Rev. D, 63, 054012 (2001) |hep-ph/0008046|; C. Bobeth et al., Phys. Rev. D, 64, 074014 (2001) |hep-ph/0104284]; Phys. Rev. D, 66, 074021 (2002) |hep-ph/0204225|; Z. Xiong and J. M. Yang, Nucl. Phys. B, 628, 193 (2002) |hep-ph/0105260|; A. Dedes et al., Phys. Rev. Lett., 87, 251804 (2001) |hep-ph/0108037|; S. w. Baek, P. Ko and W. Y. Song, Phys. Rev. Lett., 89, 271801 (2002) |hep-ph/0205259|; JHEP, 0303, 054 (2003) |hep-ph/0208112|; J. K. Mizukoshi, X. Tata and Y. Wang, Phys. Rev. D, 66, 115003 (2002)
|hep-ph/0208078|; C. S. Huang and X. H. Wu, Nucl. Phys. B, 657, 304 (2003) hep-ph/0212220); G. L. Kane,
C. Kolda and J. E. Lennon, hep-ph/0310042
47. G. Isidori and A. Retico, JHEP, 0209, 063 (2002)
|hep-ph/0208159]; R. Fleischer, G. Isidori and J. Matias, JHEP, 0305, 053 (2003) |hep-ph/0302229]; A. J. Buras,
Phys. Lett. B, 566, 115 (2003) |hep-ph/0303060|.
48. G. Hiller and F. Kruger, hep-ph/0310219

