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21 September 2011 Proc. R. Soc. A 2012 **468**, doi: 10.1098/rspa.2011.0137 first published online

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An operative method to obtain sharp nonlinear stability for systems with spatially dependent coefficients

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This paper explores an operative technique for deriving nonlinear stability by studying double-diffusive porous convection with a concentration-based internal heat source. Previous stability analyses on this problem have yielded regions of potential subcritical instabilities where the linear instability and nonlinear stability thresholds do not coincide. It is shown in this paper that the operative technique yields sharp conditional nonlinear stability in regions where the instability is found to be monotonic. This is the first instance, in the present literature, where this technique has been shown to generate sharp thresholds for a system with spatially dependent coefficients, which strongly advocates its wider use.

Keywords: energy method; double-diffusive convection; porous media

1. Introduction

Two of the key techniques employed in a stability analysis are the nonlinear energy method and the more widely used method of linearized instability. A linear instability analysis, by definition, only provides boundaries for instability due to the presence of nonlinear terms (cf. [Straughan 2004\)](#page-14-0). The use of the energy method to construct stability thresholds is therefore crucial to assess whether the linear theory accurately encapsulates the physics of the onset and behaviour of an instability. Another key advantage of performing a nonlinear energy analysis is that it provides rigorous conditions for stability without exploiting linearization or weakly nonlinear approximations.

Nonlinear stability analyses have been performed successfully in a variety of applications, see Hill $\&$ Carr (2010) and [Singh \(2010\).](#page-14-0) However, for a considerable range of problems, the standard energy approach does not generate sharp results (i.e. where the linear and nonlinear thresholds coincide), see [Carr \(2003\),](#page-13-0) [Sunil & Mahajan \(2008\)](#page-14-0) and [Saravanan & Brindha \(2010\).](#page-14-0) This may, of course, be due to the significant possibility that the linear theory has not predicted

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the onset of instability accurately; however, there are several new nonlinear energy techniques in development which show that sharp thresholds can still be achieved.

In the field of porous media convection, a differential constraint approach (cf. [van Duijn](#page-14-0) *et al.* 2002) has been introduced, where the fluid-governing Darcy equation is retained as a constraint. This has been shown to reduce the number of required coupling parameters (cf. [Hill 2009\)](#page-14-0) and to provide sharp nonlinear thresholds in contrast to the standard approach, see [Capone](#page-13-0) *et al.* (2010).

In a separate recent development, an operative technique (for systems with constant coefficients) was introduced by [Mulone \(2004\)](#page-14-0) that involves a change of variable in the energy method, based on the eigenvalues of the linearized system. In this paper, which was concerned with porous media convection, the coincidence of the critical linear and nonlinear stability parameters was proved. Further applications of this technique have seen sharp thresholds being produced for biological systems with diffusion [\(Mulone & Straughan 2009\)](#page-14-0).

To assess the feasibility of an operative approach that could include systems with spatially dependent coefficients (which extends the use of the technique considerably), [Hill \(2008\)](#page-14-0) further developed the method in the context of penetrative double-diffusive convection. This is an interesting test case, as the standard energy method approach yields a nonlinear threshold that is independent of the salt field. By adopting this new approach, a nonlinear threshold that was dependent on the salt field was shown, although coincidence between the linear and nonlinear thresholds was not achieved.

It is important to note that the operative technique yields regions of conditional stability, for which the initial data must be bounded by a threshold. Conditional stability does not necessarily preclude subcritical instabilities if this restriction on the initial data is not considered.

The purpose of this paper is twofold. Firstly, to improve on the design of the method from [Hill \(2008\),](#page-14-0) now that its feasibility has been established, by basing the transformation solely on the linear eigenvalues, making the method much easier to solve numerically and, therefore, more widely useable. Secondly, to apply it to a physical system where all previous attempts to generate sharp nonlinear thresholds have not been successful in regions where the instability is monotonic. The technique, at present, has not been developed to be applicable in regions of oscillatory instability. The energy method itself (which underpins all of the nonlinear techniques) does not tend to generate sharp thresholds in such regions (cf. [Straughan 2004\)](#page-14-0). The operative technique will be used in conjunction with the differential constraint approach, which is a new development in the current literature.

Due to the competing temperature and salt gradients, double-diffusive convection poses a challenging nonlinear stability problem, and is therefore a good class of problems to test the method on. The introduction of a concentrationbased internal heat source yields a system that forms an accurate model of cumulus convection as it occurs in the atmosphere [\(Krishnamurti 1997\)](#page-14-0), and can also be interpreted as a simplified model of the salt gradient layer in a solar pond [\(Kudish & Wolf 1978\)](#page-14-0). Although there has been considerable interest in this problem, see [Krishnamurti \(1997\),](#page-14-0) [Straughan \(2002\),](#page-14-0) Hill [\(2003,](#page-14-0) [2005\)](#page-14-0) and [Chang \(2004\),](#page-13-0) the coincidence of the linear and nonlinear thresholds for the double-diffusive case has not been achieved, even with the application of the

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differential constraint approach [\(Hill 2009\)](#page-14-0). In this paper, we apply our operative technique to this problem (based on the results of [Hill 2005\)](#page-14-0), the purpose of which is to generate sharp nonlinear thresholds.

All numerical results were derived using the Chebyshev tau-QZ method [\(Dongarra](#page-13-0) *et al.* 1996), which is a spectral method coupled with the QZ algorithm, and were checked by varying the number of polynomials to verify convergence.

2. Formation of the problem

Let us consider the same configuration as [Hill \(2005\),](#page-14-0) where a fluid-saturated porous layer is contained between the horizontal parallel planes $z = 0$ and L. O*xyz* is the standard Cartesian frame of reference with unit vectors **i**, **j** and **k**, respectively. The model is of double-diffusive convection induced by the selective absorption of radiation, where the convection mechanism is essentially a penetrative one effectively modelled via an internal heat source.

The Darcy equation is assumed to govern the fluid motion in the layer, such that

$$
\frac{\mu}{K}\mathbf{v} = -\nabla p - \mathbf{k}g\rho,\tag{2.1}
$$

where $\mathbf{v} = (u, v, w)$ and p are velocity and pressure, q is the acceleration due to gravity, μ is the dynamic viscosity of the fluid and K is the permeability.

Denoting *T* to be the temperature and *C* to be the concentration of the dissolved species, the density $\rho(T, C)$ is given by

$$
\rho(T, C) = \rho_0 (1 - \alpha_t (T - T_0) + \alpha_c (C - C_0)),
$$

where ρ_0 , T_0 and C_0 are reference values of density, temperature and concentration, respectively, and α_t and α_c are the coefficients of thermal and solutal expansion.

Equation (2.1), together with the incompressibility condition and the equations of energy and solute balance, yield the following system of governing equations:

$$
\frac{\mu}{K}\mathbf{v} = -\nabla p - \mathbf{k}g\rho_0(1 - \alpha_t(T - T_0) + \alpha_c(C - C_0)),\tag{2.2}
$$

$$
\nabla \cdot \mathbf{v} = 0,\tag{2.3}
$$

$$
\frac{1}{M}\frac{\partial T}{\partial t} + \mathbf{v} \cdot \nabla T = \kappa_t \Delta T + \beta C \tag{2.4}
$$

and

$$
\varepsilon \frac{\partial C}{\partial t} + \mathbf{v} \cdot \nabla C = \kappa_c \Delta C. \tag{2.5}
$$

The internal heat source is modelled linearly with respect to concentration, which is represented by the introduction of the βC term in equation (2.4), where β is some constant of proportionality. In these equations, ε is the porosity, κ_t and κ_c are the thermal and solutal diffusivities, and $M = (\rho_0 h_p)_f/(\rho_0 h)_{\text{m}}$ where $(\rho_0 h)_{\text{m}} =$ $(1 - \varepsilon)(\rho_0 h_s + \varepsilon(\rho_0 h_p)_f$, h_p is the specific heat of the fluid and h is the specific heat of the solid, with the subscripts 'f', 's' and 'm' referring to the fluid, solid and porous components of the medium, respectively.

The boundary conditions for the problem are $\mathbf{v} = \mathbf{0}$, $T = T_L$ and $C = C_L$ at $z = 0$ and $\mathbf{v} = \mathbf{0}$, $T = T_U < T_L$, $C = 0$ at $z = L$.

Assuming a steady-state solution $(\overline{v}, \overline{p}, \overline{T}, \overline{C})$ that corresponds to no fluid flow (i.e. $\overline{\mathbf{v}} = \mathbf{0}$), and introducing a perturbation to the steady state of the form

$$
\mathbf{v} = \overline{\mathbf{v}} + \mathbf{u}, \quad p = \overline{p} + \hat{p}, \quad T = \overline{T} + \theta \text{ and } C = \overline{C} + \phi,
$$

we may derive the governing non-dimensionalized perturbation equations,

$$
\mathbf{u} = -\nabla \hat{p} + \mathbf{k} R_{\rm T}^2 \theta - \mathbf{k} R_{\rm S}^2 L e \phi, \qquad (2.6)
$$

$$
\nabla \cdot \mathbf{u} = 0,\tag{2.7}
$$

$$
\frac{\partial \theta}{\partial t} + \mathbf{u} \cdot \nabla \theta = wF(z) + \Delta \theta + \tau \, Le \, \phi \tag{2.8}
$$

and
$$
\hat{\phi} Le \frac{\partial \phi}{\partial t} + Le \mathbf{u} \cdot \nabla \phi = w + \Delta \phi.
$$
 (2.9)

In these equations, $Le = \frac{\kappa_t}{\kappa_c}$ is the Lewis number, $\phi = \varepsilon M$ is the normalized porosity, $\tau = \beta c_L L^2 / (k_t (T_L - T_U))$ is a measure of the internal heat source generated by the radiation absorbing concentrate and $R_{\rm T}^2$ = $\alpha_t g \rho_0 (T_L - T_U) KL/(\mu \kappa_t)$ and $R_S^2 = \alpha_c g \rho_0 C_L KL/(\mu \kappa_t)$ are the thermal and solute Rayleigh numbers, respectively. The term $F(z) = 1 - \tau/3 + \tau(z - z^2/2)$ is the non-dimensionalized temperature gradient. Although a slightly different nondimensionalization is used in this paper, a detailed derivation of equivalent governing equations and explicit parameter definitions may be found in [Hill \(2005\).](#page-14-0)

The boundary conditions that follow from the perturbed quantities are $w = 0$, $\theta = 0$ and $\phi = 0$ at $z = 0, 1$, where $(\mathbf{u}, \theta, \phi)$ have a periodic plan-form tiling the (x, y) plane, and Ω is the period cell for the perturbations.

3. Linear analysis

The linearized equations are derived from equations (2.6) – (2.9) by discarding the nonlinear terms. Introducing normal modes of the form

$$
w = w(z) e^{\sigma t + ia_1 x + ia_2 y}
$$
, $\theta = \theta(z) e^{\sigma t + ia_1 x + ia_2 y}$ and $\phi = \phi(z) e^{\sigma t + ia_1 x + ia_2 y}$,

where $\sigma \in \mathbb{C}$ is the growth rate, and the wavenumber $a^2 = a_1^2 + a_2^2$ is a measure of the width of the convection cells that form at the onset of instability, and taking the double *curl* of equation (2.6) to remove the pressure term, system (2.6) – (2.9) yields

$$
\left(\frac{d^2}{dz^2} - a^2\right)w + a^2 R_{\rm T}^2 \theta - a^2 R_{\rm S}^2 \phi = 0,
$$
\n(3.1)

$$
\left(\frac{\mathrm{d}^2}{\mathrm{d}z^2} - a^2\right)\theta + wF(z) + \tau \, Le \,\phi = \sigma\theta\tag{3.2}
$$

and
$$
\left(\frac{\mathrm{d}^2}{\mathrm{d}z^2} - a^2\right)\phi + w = \hat{\phi} \, Le \, \sigma\phi, \tag{3.3}
$$

with boundary conditions $w = \theta = \phi = 0$ at $z = 0, 1$. If $Re(\sigma) > 0$, then the perturbation will grow exponentially in time, clearly leading to linear instability. The sixth-order system (3.1) – (3.3) was solved using the Chebyshev-tau method (cf. [Dongarra](#page-13-0) *et al.* 1996), which is a spectral technique coupled with the QZ algorithm. Numerical results for the linear theory are presented in §5.

4. An operative method approach to derive sharp nonlinear thresholds

The operative method approach that is used in this work is a development of the argument presented in [Mulone \(2004\)](#page-14-0) (and later in [Mulone & Straughan](#page-14-0) [2006\)](#page-14-0) to address systems where the coefficients have spatial dependence. The presence of spatially dependent coefficients is highly common in doublediffusive porous convection problems, which makes this technique particularly applicable. A version of this technique was first applied to a simple penetrative double-diffusive convection problem in [Hill \(2008\)](#page-14-0) to assess its feasibility.

(*a*) *Transforming the nonlinear system*

Our aim is to construct a transformation associated, in a canonical way, to the linear system via the first eigenvalue of the Laplacian operator that we can apply to our full nonlinear system (2.6) – (2.9) .

However, in systems with spatially dependent coefficients (in our case, the $F(z)$ term in equation (3.2)), the eigenvalues of the Laplacian operator cannot be directly evaluated. To overcome this, we construct the transform based on the eigenvalues of a modified version of linearized equations. For equations (3.1) – (3.3) , this is achieved by replacing $F(z)$ with the average of its value over the range $z \in [0, 1]$ (namely 1).

Under this modified system, all the even derivatives of w, θ and ϕ now vanish at the boundaries, so we may take solutions of the form

$$
w = \sum_{p=1}^{\infty} w_p(t) \sin p\pi z, \quad \theta = \sum_{p=1}^{\infty} \theta_p(t) \sin p\pi z \quad \text{and} \quad \phi = \sum_{p=1}^{\infty} \phi_p(t) \sin p\pi z.
$$

It can easily be shown that the thermal Rayleigh number of this modified linearized system is minimized by the mode $p = 1$. Since the sin functions are linearly independent, we may study only the most unstable mode $p = 1$. Letting $s = a_{\rm L}^2 + \pi^2$ and $\alpha = a_{\rm L}^2/(a_{\rm L}^2 + \pi^2)$, equation (3.1) may be written

$$
w = \alpha (R_{\rm TL}^2 \theta - R_{\rm SL}^2 \, Le \, \phi),\tag{4.1}
$$

where a_{L}^2 , R_{TL}^2 and R_{SL}^2 are the wavenumber, and thermal and solute Rayleigh numbers associated with the modified linear system, respectively. Using equation (4.1) , the remaining linearized equations (3.2) and (3.3) may be written in the matrix form

$$
\mathbf{X}_{,t}=M\mathbf{X},
$$

where $\mathbf{X} = (\theta, \phi)^{\mathrm{T}}$ and

$$
M = \begin{pmatrix} -s + \alpha R_{\text{TL}}^2 & -\alpha \, Le \, R_{\text{SL}}^2 + \tau \, Le \\ \frac{\alpha R_{\text{TL}}^2}{\hat{\phi} \, Le} & -\frac{1}{\hat{\phi}} \left(\frac{s}{Le} + R_{\text{SL}}^2 \alpha \right) \end{pmatrix}.
$$

The eigenvalues of matrix *M* are given by

$$
eig(M) = \frac{1}{2} \left[\alpha \left(R_{\text{TL}}^2 - \frac{R_{\text{SL}}^2}{\hat{\phi}} \right) - s \left(\frac{1}{\hat{\phi} L e} + 1 \right) \pm \zeta_2 \right],\tag{4.2}
$$

where

$$
\zeta_1 = \alpha \left(R_{\rm TL}^2 + \frac{R_{\rm SL}^2}{\hat{\phi}} \right) + s \left(\frac{1}{\hat{\phi} L e} - 1 \right) \quad \text{and} \quad \zeta_2 = \sqrt{\zeta_1^2 + \frac{4\alpha R_{\rm TL}^2}{\hat{\phi}} (\tau - \alpha R_{\rm SL}^2)}.
$$

Under the condition

$$
\zeta_1^2 + \frac{4\alpha R_{\text{TL}}^2}{\hat{\phi}} (\tau - \alpha R_{\text{SL}}^2) > 0, \qquad (4.3)
$$

the matrix Q of eigenvectors associated with eigenvalues equation (4.2) can be constructed, such that

$$
Q = \begin{pmatrix} \hat{\phi} \, Le \\ \frac{2\alpha R_{\text{TL}}^2}{1} (\zeta_1 + \zeta_2) & \frac{\hat{\phi} \, Le}{2\alpha R_{\text{TL}}^2} (\zeta_1 - \zeta_2) \\ 1 & 1 \end{pmatrix}
$$

and

$$
Q^{-1} = \begin{pmatrix} \frac{\alpha R_{\text{TL}}^2}{\hat{\phi} L e \, \zeta_2} & \frac{-\zeta_1 + \zeta_2}{2 \zeta_2} \\ -\frac{\alpha R_{\text{TL}}^2}{\hat{\phi} L e \, \zeta_2} & \frac{\zeta_1 + \zeta_2}{2 \zeta_2} \end{pmatrix}.
$$

Introducing the new variables

$$
\begin{pmatrix} \gamma \\ \psi \end{pmatrix} = Q^{-1} \begin{pmatrix} \theta \\ \phi \end{pmatrix},
$$

we can now transform the full nonlinear system (2.6) – (2.9) . Note that as all the values in the transformation are taken from the linear system (which are derived first), matrices *Q* and *Q*−¹ contain only real numbers.

After making this substitution (and taking the third component of the double curl of equation (2.6) to remove the pressure term), we have the system

$$
\Delta w = \Delta^* (R_{\rm T}^2 Q_{11} - Le \, R_{\rm S}^2) \gamma + \Delta^* (R_{\rm T}^2 Q_{12} - Le \, R_{\rm S}^2) \psi, \tag{4.4}
$$

$$
\frac{\partial \gamma}{\partial t} + (1 + A_1)\mathbf{u} \cdot \nabla \gamma + A_1 \mathbf{u} \cdot \nabla \psi = F_1 w + (1 + B_1)\Delta \gamma + B_1 \Delta \psi
$$

+ $C_1 \gamma + C_1 \psi$ (4.5)

and $\frac{\partial \psi}{\partial t}$

$$
\frac{\partial}{\partial t} + (1 + A_2)\mathbf{u} \cdot \nabla \psi + A_2 \mathbf{u} \cdot \nabla \gamma = F_2 w + (1 + B_2)\Delta \psi + B_2 \Delta \gamma
$$

+ $C_2 \gamma + C_2 \psi$, (4.6)

where $\Delta^* = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2}$ and

$$
A_j = Q_{j2}^{-1} \left(\frac{1}{\hat{\phi}} - 1 \right), \quad B_j = Q_{j2}^{-1} \left(\frac{1}{\hat{\phi}} I e^{-1} \right),
$$

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$$
C_j = \tau Q_{j1}^{-1} Le \text{ and } F_j = F(z) Q_{j1}^{-1} + \frac{Q_{j2}^{-1}}{\hat{\phi} Le},
$$

with boundary conditions $w = \gamma = \psi = 0$ at $z = 0, 1$.

(*b*) *Deriving the sharp nonlinear stability bounds*

To obtain sharp nonlinear stability bounds in the stability measure $L^2(\Omega)$, (following an analogous argument to [Lombardo](#page-14-0) *et al.* (2001)), we multiply equation (4.5) by γ and $\Delta \gamma$, respectively, and equation (4.6) by ψ and $\Delta \psi$, respectively, and integrate over Ω to obtain

$$
\frac{1}{2} \frac{d}{dt} ||\gamma||^2 + A_1 \langle \mathbf{u} \cdot \nabla \psi, \gamma \rangle = \langle F_1 w, \gamma \rangle - (1 + B_1) ||\nabla \gamma||^2 - B_1 \langle \nabla \psi, \nabla \gamma \rangle
$$

+ $C_1 ||\gamma||^2 + C_1 \langle \psi, \gamma \rangle$,

$$
-\frac{1}{2} \frac{d}{dt} ||\nabla \gamma||^2 + (1 + A_1) \langle \mathbf{u} \cdot \nabla \gamma, \Delta \gamma \rangle + A_1 \langle \mathbf{u} \cdot \nabla \psi, \Delta \gamma \rangle
$$

= $\langle F_1 w, \Delta \gamma \rangle + (1 + B_1) ||\Delta \gamma||^2 + B_1 \langle \Delta \psi, \Delta \gamma \rangle - C_1 ||\nabla \gamma||^2 - C_1 \langle \nabla \psi, \nabla \gamma \rangle$,

$$
\frac{1}{2} \frac{d}{dt} ||\psi||^2 + A_2 \langle \mathbf{u} \cdot \nabla \gamma, \psi \rangle = \langle F_2 w, \psi \rangle - (1 + B_2) ||\nabla \psi||^2 - B_2 \langle \nabla \gamma, \nabla \psi \rangle
$$

+ $C_2 \langle \gamma, \psi \rangle + C_2 ||\psi||^2$

and

$$
-\frac{1}{2}\frac{\mathrm{d}}{\mathrm{d}t}\|\nabla\psi\|^2 + (1+A_2)\langle\mathbf{u}\cdot\nabla\psi,\Delta\psi\rangle + A_2\langle\mathbf{u}\cdot\nabla\gamma,\Delta\psi\rangle = \langle F_2w,\Delta\psi\rangle
$$

$$
+\left(1+B_2\right)\|\Delta\psi\|^2 + B_2\langle\Delta\gamma,\Delta\psi\rangle - C_2\langle\nabla\gamma,\nabla\psi\rangle - C_2\|\nabla\psi\|^2,
$$

where $\|\cdot\|$ and $\langle \cdot, \cdot \rangle$ denote the norm and inner product on $L^2(\Omega)$. Adopting the differential constraint approach of [van Duijn](#page-14-0) *et al.* (2002) and [Hill \(2009\),](#page-14-0) equation (4.4) is retained as a constraint. This is the first instance, in the current literature, in which the differential constraint approach has been used in conjunction with an operative method.

Letting

$$
E(t) = \frac{1}{2}(\|\gamma\|^2 + \lambda \|\psi\|^2) + \frac{b}{2}(\|\nabla \gamma\|^2 + \|\nabla \psi\|^2),
$$

where $\lambda > 0$ is a coupling parameter, and b is a positive constant, we can write

$$
\frac{\mathrm{d}E}{\mathrm{d}t} = \mathcal{I}_0 - \mathcal{D}_0 + \mathcal{N}_0 + b\mathcal{I}_1 - b\mathcal{D}_1 + b\mathcal{N}_1,\tag{4.7}
$$

where

$$
\mathcal{I}_0 = \langle F_1 w, \gamma \rangle + \lambda \langle F_2 w, \psi \rangle + C_1 ||\gamma||^2 + \lambda C_2 ||\psi||^2
$$

+
$$
(C_1 + \lambda C_2) \langle \psi, \gamma \rangle - (B_1 + \lambda B_2) \langle \nabla \psi, \nabla \gamma \rangle,
$$

$$
\mathcal{D}_0 = (1 + B_1) ||\nabla \gamma||^2 + \lambda (1 + B_2) ||\nabla \psi||^2,
$$

$$
\mathcal{N}_0 = -A_1 \langle \mathbf{u} \cdot \nabla \psi, \gamma \rangle - \lambda A_2 \langle \mathbf{u} \cdot \nabla \gamma, \psi \rangle,
$$

\n
$$
\mathcal{I}_1 = -\langle F_1 w, \Delta \gamma \rangle - \langle F_2 w, \Delta \psi \rangle + C_1 ||\nabla \gamma||^2 + C_2 ||\nabla \psi||^2
$$

\n
$$
+ \left(1 - \frac{1}{\hat{\phi} Le}\right) \langle \Delta \psi, \Delta \gamma \rangle,
$$

\n
$$
\mathcal{D}_1 = (1 + B_1) ||\Delta \gamma||^2 + (1 + B_2) ||\Delta \psi||^2
$$

\nand
\n
$$
\mathcal{N}_1 = -(1 + A_1) \langle \mathbf{u} \cdot \nabla \gamma, \Delta \gamma \rangle - A_1 \langle \mathbf{u} \cdot \nabla \psi, \Delta \gamma \rangle
$$

\n
$$
- (1 + A_2) \langle \mathbf{u} \cdot \nabla \psi, \Delta \psi \rangle - A_2 \langle \mathbf{u} \cdot \nabla \gamma, \Delta \psi \rangle.
$$

To ensure that \mathcal{D}_0 , $\mathcal{D}_1 > 0$ (i.e. $1 + B_1 > 0$ and $1 + B_2 > 0$), we have the condition

$$
\frac{4\hat{\phi} \, Le}{(\hat{\phi} \, Le + 1)^2} \zeta_1^2 + \frac{4\alpha R_{\rm TL}^2}{\hat{\phi}} (\tau - \alpha R_{\rm SL}^2) > 0. \tag{4.8}
$$

Note that if equation (4.8) holds, then condition (4.3) follows.

By defining

$$
\mathcal{D}_2 = \left(\frac{R_E - 1}{2R_E}\right)\mathcal{D}_0 + \frac{b}{2}\mathcal{D}_1 \quad \text{with} \quad \frac{1}{R_E} = \max_{\mathcal{H}} \frac{\mathcal{I}_0}{\mathcal{D}_0},
$$

where H is the space of admissible functions, it follows that $b\mathcal{I}_1 \leq D_2$ and \mathcal{N}_0 + $b\mathcal{N}_1 \leq p_1 \mathcal{D}_2 E^{1/2}$ for a positive constant p_1 (see appendix A), where $R_E > 1$.

From equation (4.7), we see that

$$
\frac{\mathrm{d}E}{\mathrm{d}t} \le -\mathcal{D}_2(1-p_1 E^{1/2}).
$$

If $R_E > 1$, then by the Poincaré inequality (e.g. $\pi^2 ||\gamma||^2 \le ||\nabla \gamma||^2$), $\mathcal{D}_2 \ge p_2 E$ for some positive constant p_2 . Hence, it follows that

$$
\frac{dE}{dt} \le -p_2 E(1 - p_1 E^{1/2}).\tag{4.9}
$$

We will show that equation (4.9) ensures conditional nonlinear stability as long as

$$
R_E > 1\tag{4.10a}
$$

and

$$
E^{1/2}(0) < \frac{1}{p_1}.\tag{4.10b}
$$

For equation (4.10_b) to hold, there are two possibilities as follows:

(i)
$$
E^{1/2}(t) < 1/p_1, t \ge 0,
$$

or

(ii) there exists $\zeta > 0$ such that $E^{1/2}(\zeta) = 1/p_1$, with $E^{1/2}(t) < 1/p_1$, for $t \in [0,\zeta)$.

Suppose (ii) holds. From equation (4.9), it follows that $dE/dt < 0$ for $t \in [0,\zeta)$, hence, $E^{1/2}(t) \leq E^{1/2}(0) < 1/p_1$ for any $t \in [0, \zeta)$. As $E(t)$ is a continuous function of *t* on [0, ζ], then it is impossible that $E^{1/2}(\zeta) = 1/p_1$. This contradiction implies (i) and consequently $dE/dt \leq 0$ for $t \geq 0$ such that $E^{1/2}(t) \leq E^{1/2}(0)$. Thus,

$$
\frac{\mathrm{d}E}{\mathrm{d}t} \le -p_2(1 - p_1 E^{1/2}(0))E.
$$

Integrating, we have

$$
E(t) \le E(0) e^{-p_2(1-p_1 E^{1/2}(0))t} \to 0 \text{ as } t \to \infty,
$$

which shows the decay of γ and ψ . Using equation (2.6),

$$
\|\mathbf{u}\| \le |R_{\rm T}^2 Q_{11} - R_{\rm S}^2 L e| \|\gamma\| + |R_{\rm T}^2 Q_{12} - R_{\rm S}^2 L e| \|\psi\|.
$$
 (4.11)

The decay of **u** clearly follows.

It is important to note that we have derived conditional stability (under the restriction (4.10*b*)), such that the initial data $E^{1/2}(0)$ must be bounded by $1/p_1$. From the definition of p_1 (see appendix A) as $R_E \rightarrow 1$, the bound on the initial data tends to 0. However, for $\hat{\phi} = 1$, we may take $b = 0$, which yields unconditional nonlinear stability, such that there is no bound on the initial data.

Following the differential constraint approach of [van Duijn](#page-14-0) *et al.* (2002) and [Hill \(2009\)](#page-14-0) (where we retained equation (4.4) as a constraint), we introduce an Euler–Lagrange multiplier μ such that

$$
\mu(\mathbf{x})(\Delta w + \Delta^*(Le R_S^2 - R_T^2 Q_{11})\gamma + \Delta^*(Le R_S^2 - R_T^2 Q_{21})\psi) = 0.
$$

The Euler–Lagrange equations for the maximization problem $1/R_E$ = $\max_{\mathcal{H}}(\mathcal{I}_0/\mathcal{D}_0)$ are now given by

$$
R_E(\Delta w + \Delta^*(Le R_S^2 - R_T^2 Q_{11})\gamma + \Delta^*(Le R_S^2 - R_T^2 Q_{21})\psi) = 0,
$$

\n
$$
R_E(\Delta \mu + F_1\gamma + F_2\lambda \psi) = 0,
$$

\n
$$
R_E(F_1 w + (C_1 + \lambda C_2)\psi + 2C_{11}\gamma + (B_1 + \lambda B_2)\Delta \psi
$$

\n
$$
+ (Le R_S^2 - R_T^2 Q_{11})\Delta^*\mu) + 2(1 + B_1)\Delta\gamma = 0
$$

\nand
\n
$$
R_E(\lambda F_2 w + (C_1 + \lambda C_2)\gamma + 2C_2\lambda \psi + (B_1 + \lambda B_2)\Delta\gamma
$$

\n
$$
+ (Le R_S^2 - R_T^2 Q_{12})\Delta^*\mu) + 2(1 + B_2)\Delta\psi = 0,
$$

\n(4.12)

with boundary conditions $w = \mu = \gamma = \psi = 0$ at $z = 0, 1$.

After adopting normal modes, the eighth-order eigenvalue problem (4.12) for *RE* was solved using the Chebyshev-tau method (cf. [Dongarra](#page-13-0) *et al.* 1996). Under the analysis of this section, the ranges of the remaining parameters (restricted by equation (4.8)) with eigenvalues $R_E > 1$ correspond to regions of stability. An overview of the numerical approach and the results for the operative nonlinear energy approach are presented in §5.

5. Results and conclusions

For fixed values of *Le*, τ , $\hat{\phi}$ and $R_{\rm S}^2$, the region (in terms of $R_{\rm T}^2$) that yielded conditional stability under the operative method was numerically derived by adopting the following approach:

- (1) solve the linear instability problem to yield the critical values of R_{TL}^2 and $a_{\rm L}^2$ (noting that $R_{\rm SL}^2 = R_{\rm S}^2$);
- (2) if condition (4.8) is satisfied, then the nonlinear operative approach may be applied;
- (3) use the derived R_{TL}^2 and a_{L}^2 values to define the constants A_j , B_j , C_j and the functions F_j ($j=1,2$) in equation (4.12);
- (4) fix the value of $R_{\rm T}^2$;
- (5) minimize R_E over the wavenumber a^2 and maximize R_E over the coupling parameter λ by repeatedly solving eigenvalue problem (4.12) for R_E ;
- (6) if $R_E > 1$, then we have stability; and
- (7) repeating steps 5 and 6 for a variety of $R_{\rm T}^2$ values yields the region of stability.

Eigenvalue problem (4.12) was solved using the Chebyshev tau-QZ method [\(Dongarra](#page-13-0) *et al.* 1996), which is a spectral method coupled with the QZ algorithm. The results were checked by varying the number of polynomials to verify convergence. To compare our operative technique with the existing energy methods, standard nonlinear stability results are also derived (see [Hill 2005](#page-14-0) for more details).

[Figure 1](#page-11-0) presents stable and unstable regions, which have been generated numerically, in a graphical form. As the linear instability thresholds (which give the unstable regions) have been presented before for a range of parameter values in [Hill \(2005\),](#page-14-0) [figure 1](#page-11-0) simply gives a representative example of how the operative method performs. The conditionally stable and globally stable regions correspond to the results generated by the operative and standard nonlinear stability approaches, respectively.

It is immediately obvious from [figure 1](#page-11-0) that the operative technique produces sharp thresholds. In figure $1a, b$, as R_S^2 is increased, the region between the linear and standard nonlinear thresholds becomes very substantial. This region is entirely removed by the operative approach, clearly demonstrating the highly substantial improvement this new technique brings over the existing methods.

However, in contrast to the operative method, the standard approach yields unconditional stability (such that there is no bound on the initial data). The conditional stability generated by the operative approach does not necessarily preclude sub-critical instabilities if the restriction on the initial data is not considered.

From [figure 1](#page-11-0)*a*, we can see that the sharpness of the threshold is maintained in nearly all the regions where the method is valid (under condition (4.8)). This behaviour was repeated in the parameter ranges explored. It is important to note, though, that condition (4.8) only appeared to fail near to regions where the instability becomes oscillatory. The energy method itself (which underpins all of the nonlinear techniques) does not tend to generate sharp thresholds in such regions (cf. [Straughan 2004\)](#page-14-0).

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Figure 1. Visual representation of unstable, conditionally stable (nonlinear operative method) and globally stable (standard nonlinear method) regions, with critical thermal Rayleigh number $R_{\rm T}^2$ plotted against R_S^2 . The solid line represents the onset of linear instability, such that the region above it is unstable. The dashed line represents the point at which the transformation is no longer valid for increasing R_S^2 . The graphs correspond to (*a*) $\tau = 1, \hat{\phi} = 0.1, Le = 20$ and (*b*) $\tau = 5, \hat{\phi} =$ $0.1, Le = 20.$

It has been demonstrated that for a system with spatially dependent coefficients, the operative technique presented in this paper can generate sharp thresholds, which is not achieved by the standard approach. This strongly advocates its wider use, with the potential to sharpen nonlinear stability thresholds in a substantial number of problems. The stability results generated by this approach are conditional, though, such that the initial data must be bounded by a defined threshold.

The authors would like to thank three anonymous referees whose comments have led to improvements in the paper.

Appendix A

Recall that for $R_E > 1$ and $b > 0$,

$$
\mathcal{D}_0 = (1 + B_1) \|\nabla \gamma\|^2 + \lambda (1 + B_2) \|\nabla \psi\|^2, \quad \mathcal{D}_2 = \left(\frac{R_E - 1}{2R_E}\right) \mathcal{D}_0 + \frac{b}{2} \mathcal{D}_1,
$$
\n
$$
\mathcal{D}_1 = (1 + B_1) \|\Delta \gamma\|^2 + (1 + B_2) \|\Delta \psi\|^2
$$
\nand\n
$$
\mathcal{I}_1 = -\langle F_1(z)w, \Delta \gamma \rangle - \langle F_2(z)w, \Delta \psi \rangle + C_1 \|\nabla \gamma\|^2 + C_2 \|\nabla \psi\|^2
$$
\n
$$
+ \left(1 - \frac{1}{\hat{\phi} \, Le}\right) \langle \Delta \psi, \Delta \gamma \rangle,
$$

where under condition (4.8),
$$
1 + B_1 > 0
$$
 and $1 + B_2 > 0$.

Letting $g_1 = \max_{z \in [0,1]} F_1^2(z)$ and $g_2 = \max_{z \in [0,1]} F_2^2(z)$, using Young's inequality (for constants $c_1, c_2 > 0$) and noting that $C_2 < 0$, it follows that

$$
b\mathcal{I}_1 \leq \frac{b}{2} \left(\frac{1}{c_1} + \frac{1}{c_2} \right) \|w\|^2 + bC_1 \|\nabla \gamma\|^2
$$

+
$$
\frac{b}{2} \left(\left(c_1 g_1 + 1 - \frac{1}{\hat{\phi} \, Le} \right) \|\Delta \gamma\|^2 + \left(c_2 g_2 + 1 - \frac{1}{\hat{\phi} \, Le} \right) \|\Delta \psi\|^2 \right).
$$

Choosing *c*¹ and *c*² such that

$$
c_1 g_1 + 1 - \frac{1}{\hat{\phi}} L e = 1 + B_1
$$
 and $c_2 g_2 + 1 - \frac{1}{\hat{\phi}} L e = 1 + B_2$,

which holds for $1 < \phi \, Le < 3$, we have

$$
b\mathcal{I}_1 \leq \frac{b}{2} \left(\frac{g_1 \hat{\phi} \, Le}{B_1 \hat{\phi} \, Le + 1} + \frac{g_2 \hat{\phi} \, Le}{B_2 \hat{\phi} \, Le + 1} \right) \|w\|^2 + bC_1 \|\nabla \gamma\|^2 + \frac{b}{2} \mathcal{D}_1.
$$

Using equation (4.11) yields

$$
b\mathcal{I}_1 \leq \frac{b}{2}c_5(c_3 + c_4)(c_3\|\gamma\|^2 + c_4\|\psi\|^2) + bC_1\|\nabla\gamma\|^2 + \frac{b}{2}\mathcal{D}_1,
$$

where

$$
c_3 = |R_{\rm T}^2 Q_{11} - R_{\rm S}^2 L e|, \quad c_4 = |R_{\rm T}^2 Q_{12} - R_{\rm S}^2 L e|
$$

and

$$
c_5 = \frac{g_1 \hat{\phi} L e}{B_1 \hat{\phi} L e + 1} + \frac{g_2 \hat{\phi} L e}{B_2 \hat{\phi} L e + 1}.
$$

Using the Poincaré inequality and letting

$$
b = \left(\frac{R_E - 1}{R_E}\right) \min\left(\frac{(1 + B_1)\pi^2}{c_3 c_5 (c_3 + c_4) + 2C_1 \pi^2}, \frac{\lambda (1 + B_2)\pi^2}{c_4 c_5 (c_3 + c_4)}\right),
$$

we have

$$
b\mathcal{I}_1 \leq \left(\frac{R_E - 1}{2R_E}\right)\mathcal{D}_0 + \frac{b}{2}\mathcal{D}_1 = \mathcal{D}_2,
$$

as required.

Concerning \mathcal{N}_0 , using the inequality $\sup_{\Omega} |f| \leq c_6 \|\Delta f\|$, where

$$
c_6 = \frac{\sqrt{3}}{(\pi^5 h^3 \sqrt{2}(\sqrt{2}-1))^{1/2}} + \frac{2^{5/2} 5 (1+\pi^2)^{1/2} h^{3/5}}{3\pi}, \text{ with } h = \min(a_1, a_2, 1)
$$

and $f = 0$ on $z = 0, 1$, (cf. [Straughan 2004\)](#page-14-0), the Cauchy–Schwarz inequality and equation (4.11) , we have

$$
\mathcal{N}_0 = -A_1 \langle \mathbf{u} \cdot \nabla \psi, \gamma \rangle - \lambda A_2 \langle \mathbf{u} \cdot \nabla \gamma, \psi \rangle \n\leq |A_1|c_6 \|\Delta \gamma\| \|\mathbf{u}\| \|\nabla \psi\| + \lambda |A_2|c_6 \|\Delta \psi\| \|\mathbf{u}\| \|\nabla \gamma\| \n\leq c_6 (c_3 \|\gamma\| + c_4 \|\psi\|) (|A_1| \|\Delta \gamma\| \|\nabla \psi\| + \lambda |A_2| \|\Delta \psi\| \|\nabla \gamma\|) \n\leq c_6 \left(c_3 \sqrt{2} E^{1/2} + c_4 \sqrt{\frac{2}{\lambda}} E^{1/2} \right) \left(2|A_1| \mathcal{D}_2 \left(\frac{R_E}{\lambda b (1 + B_1)(1 + B_2)(R_E - 1)} \right)^{1/2} \right) \n+ 2\lambda |A_2| \mathcal{D}_2 \left(\frac{R_E}{b (1 + B_1)(1 + B_2)(R_E - 1)} \right)^{1/2} \right) \n\leq r_1 E^{1/2} \mathcal{D}_2,
$$

where

$$
r_1 = 2^{3/2} c_6 \left(\frac{R_E}{b(1+B_1)(1+B_2)(R_E-1)} \right)^{1/2} \left(c_3 + \frac{c_4}{\sqrt{\lambda}} \right) \left(\frac{|A_1|}{\sqrt{\lambda}} + \lambda |A_2| \right).
$$

It is important to note that $r_1 \rightarrow \infty$ as $R_E \rightarrow 1$.

Again, using the inequality $\sup_{\Omega} |u| \leq c_6 \|\Delta u\|$, the Cauchy–Schwarz inequality and equation (4.11), we obtain

$$
b\mathcal{N}_1 = -b(1 + A_1) < \mathbf{u} \cdot \nabla \gamma, \ \Delta \gamma > -bA_1 < \mathbf{u} \cdot \nabla \psi, \ \Delta \gamma > \\
\quad - b(1 + A_2) < \mathbf{u} \cdot \nabla \psi, \ \Delta \psi > -bA_2 < \mathbf{u} \cdot \nabla \gamma, \ \Delta \psi > \\
\leq bc_6(c_3 \|\Delta \gamma\| + c_4 \|\Delta \psi\|)(1 + A_1 \|\nabla \gamma\| \|\Delta \gamma\| + |A_1| \|\nabla \psi\| \|\Delta \gamma\| \\
\quad + |1 + A_2| \|\nabla \psi\| \|\Delta \psi\| + |A_2| \|\nabla \gamma\| \|\Delta \psi\|)
$$
\n
$$
\leq r_2 E^{1/2} \mathcal{D}_2,
$$

where

$$
r_2 = \frac{2\sqrt{2}c_6}{\sqrt{b}} \left(\frac{c_3}{\sqrt{1+B_1}} + \frac{c_4}{\sqrt{1+B_2}} \right) \left(\frac{|1+A_1|+|A_1|}{\sqrt{1+B_1}} + \frac{|1+A_2|+|A_2|}{\sqrt{1+B_2}} \right).
$$

Hence,

$$
\mathcal{N}_0 + b\mathcal{N}_1 \le p_1 \mathcal{D}_2 E^{1/2},
$$

where $p_1 = \max(r_1, r_2)$, as required, with $p_1 \rightarrow \infty$ as $R_E \rightarrow 1$.

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