Naturalness and a light Z'

Bin Zhu,¹ Florian Staub,^{2,3} and Ran Ding⁴

¹Department of Physics, Yantai University, Yantai 264005, P. R. China ²Institute for Theoretical Physics (ITP), Karlsruhe Institute of Technology, Engesserstra?e 7, D-76128 Karlsruhe, Germany ³Institute for Nuclear Physics (IKP), Karlsruhe Institute of Technology, Hermann-von-Helmholtz-Platz 1, D-76344 Eggenstein-Leopoldshafen, Germany ⁴Center for High-Energy Physics, Peking University, Beijing, 100871, P. R. China

> Models with a light, additional gauge boson are attractive extensions of the standard model. Often these models are only considered as effective low energy theory without any assumption about an UV completion. This leaves not only the hierarchy problem of the SM unsolved, but introduces a copy of it because of the new fundamental scalars responsible for breaking the new gauge group. A possible solution is to embed these models into a supersymmetric framework. However, this gives rise to an additional source of fine-tuning compared to the MSSM and poses the question how natural such a setup is. One might expect that the additional fine-tuning is huge, namely, $O(M_{\text{SUSY}}^2/m_{Z'}^2)$. In this paper we point out that this is not necessarily the case. We show that it is possible to find a focus point behaviour also in the new sector in co-existence to the MSSM focus point. We call this 'Double Focus Point Supersymmetry'. Moreover, we stress the need for a proper inclusion of radiative corrections in the fine-tuning calculation: a tree-level estimate would lead to predictions for the tuning which can be wrong by many orders of magnitude. As showcase, we use the $U(1)_{B-L}$ extended MSSM and discuss possible consequence of the observed ⁸Be anomaly. However, similar features are expected for other models with an extended gauge group which involve potentially large Yukawa-like interactions of the new scalars.

INTRODUCTION

Supersymmetry (SUSY) is still one of the best candidates for physics beyond standard model (BSM), which provides an elegant solution to hierarchy problem, achieving gauge coupling unification at a high scale, and providing a dark matter candidate [1]. However, the current null results from LHC direct SUSY searches together with the measured 125 GeV Higgs mass have exacerbated the little hierarchy problem and put pressure on the Minimal Supersymmetric Standard Model (MSSM) as natural extension of the SM. The little hierarchy is usually encoded in the following equation:

$$m_Z^2 \sim -2m_{H_{\gamma}}^2 - 2\mu^2 \;, \tag{1}$$

in order to prevent large fine-tuning, one needs $m_Z^2\sim m_{H_u}^2\sim \mu^2$ at the SUSY scale what becomes more and more disfavoured. To quantify the resulting amount of tuning, one can use one of the common measures like the one of Barbieri-Giudice [2]

$$\Delta^Z = \max \left| \frac{\partial \ln m_Z^2}{\partial \ln \alpha} \right| \tag{2}$$

where α represents the fundamental parameters of the model. With this measure one finds that the tuning stemming from μ is at tree-level given by $\Delta_{\mu} \simeq 2\mu^2/m_Z^2$. More recently, it was found that a more accurate prediction is actually $\Delta \simeq \mu^2/m_Z^2$ once loop corrections are taken into account [3] ¹. The other source of the

The situation becomes more complicated when a new gauge group is introduced. We consider here the case of an additional Abelian group which gets broken by a pair of scalars η_1 , η_2 with charges ± 1 under this group. In this case, one should also consider the tadpole equation associated with these new states and calculate the corresponding amount of tuning needed to fulfil them. Under the assumption that the new superfields also receive their SUSY mass from a dimensionful term in the superpotential (called μ_{η} in the following), one finds at tree-level a very similar connection as eq. (1) between the mass of the new gauge boson $m_{Z'}^2$ and the SUSY parameters

$$m_{Z'}^2 \sim -m_{\eta}^2 - \mu_{\eta}^2 ,$$
 (3)

If one assumes that the U(1) gauge coupling g' is of the same order as the weak coupling g, the current LHC

necessarily needs a light Higgsino mass. However, it was also pointed out that this strong conclusion can be avoided by introducing a non-holomorphic soft-term μ' [3, 4]

fine-tuning is $m_{H_u}^2$ which shows a large sensitive on the radiative corrections from (s)tops and the gluino. Consequently, the fine-tuning in the MSSM has a significant correlation with the value of the SM-like Higgs mass which also depends strongly on these particles. The SUSY masses necessary to explain the measured mass of 125 GeV [5, 6] unavoidable introduce a sizeable amount of fine-tuning which is typically very large in most regions of the parameter space. This is especially the case in unified scenarios like the constrained MSSM (CMSSM) with only a small set of free parameters at the scale of grand unification (GUT). The most important exception is the focus point region where the fine-tuning has only a mild dependence on the stop masses [7–18].

¹ This connection between the supersymmetric μ -term and the fine-tuning usually leads to the assumption that natural SUSY

bound for $m_{Z'}$ is in the multi-TeV range [19]. For such heavy Z', one does not need to worry too much about the fine-tuning induced by eq. (3) [20, 21]. However, a very light Z' is still allowed for tiny g'. In general, there are some motivations to consider such a light vector boson, e.g.

- Self-interacting dark matter (SIDM) [22–24] provides a possibility to reconcile the tension between the small scale structure observations and the conventional cold DM (CDM) predictions [25–27], see for instance Ref. [28] for a recent summary. One possible explanation for the self-interaction of DM is the exchange of light gauge bosons [29].
- An anomaly has been reported from isoscalar ${}^8\mathrm{Be}^*(1^+) \to {}^8\mathrm{Be}(0^+) + e^+e^-$ transitions: a bump in the opening angle distributions of e^+e^- pairs with 6.8σ significance was observed [30]. This conflicts with the standard expectation, which predicts that the distribution of opening angles of e^+e^- pairs should follow a smoothly downward curve. In addition, in the related isovector ${}^8\mathrm{Be}^{*'}(1^+) \to {}^8\mathrm{Be}(0^+) + e^+e^-$ transition no excess is visible. If one takes this excess serious, it could be originate from a new light vector boson X with the decay channel ${}^8\mathrm{Be}^*(1^+) \to {}^8\mathrm{Be}(0^+) + X \to {}^8\mathrm{Be}(0^+) + e^+e^-$ [30, 31]. The necessary mass to fit the data is

$$m_X = 16.7 \pm 0.35 \text{(stat)} \pm 0.5 \text{(sys)} \text{ MeV}$$
. (4)

Refs. [31, 32] have examined the case of a purely vector interaction with quarks. It has been shown that in order to be compatible with existing experimental constraints, the new vector boson should be protophobic [31], i.e., the coupling to proton is highly suppressed compared with the coupling to neutron. This proposal has stimulated many related works [33–41].

Such a light Z' is expected to dominate the overall fine-tuning in the model if m_{η} and μ_{η} are $O(M_{\rm SUSY})^2$.

We want to study in the following for the first time explicitly the impact of a light Z' on the fine-tuning in supersymmetric models. As example, we consider an $U(1)_{B-L}$ extension of the MSSM which was in the past mainly studied in the context of heavy Z' masses [42–49]. In particular, we discuss the necessary

Superfield	Spin $\frac{1}{2}$	Spin 1	Gauge group	Coupling
\hat{B}	$\lambda_{ ilde{B}}$	B	$U(1)_Y$	g_1
\hat{W}	$\lambda_{ ilde{W}}$	W	$SU(2)_L$	g_2
\hat{g}	$\lambda_{ ilde{g}}$	g	$SU(3)_c$	g_3
\hat{B}'	$\lambda_{ ilde{B}'}$	B'	$U(1)_{B-L}$	g_B

TABLE I. Vector superfields of the BLSSM and corresponding gauge couplings.

Superfield	N_G	$U(1)_Y \otimes SU(2)_L \otimes SU(3)_c \otimes U(1)_{B-L}$
\hat{Q}	3	$egin{array}{c} rac{1}{6}\otimes 2\otimes 3\otimes rac{1}{6} \ -rac{2}{3}\otimes 1\otimes \overline{3}\otimes -rac{1}{6} \ rac{1}{3}\otimes 1\otimes \overline{3}\otimes -rac{1}{6} \end{array}$
\hat{U}	3	$-rac{2}{3}\otimes 1\otimes \mathbf{\overline{3}}\otimes -rac{1}{6}$
\hat{D}	3	$rac{1}{3}\otimes 1\otimes \overline{3}\otimes -rac{1}{6}$
\hat{L}	3	$-rac{1}{2}\otimes {f 2}\otimes {f 1}\otimes -rac{1}{2}$
\hat{E}	3	$\begin{array}{ c c c c c }\hline 1\otimes 1\otimes 1\otimes \frac{1}{2}\\ 0\otimes 1\otimes 1\otimes \frac{1}{2}\\ \hline \end{array}$
$\hat{\nu}_R$	3	$0 \otimes 1 \otimes 1 \otimes \frac{1}{2}$
\hat{H}_u	1	$\frac{1}{2}\otimes 2\otimes 1\otimes 0$
\hat{H}_d	1	$-rac{1}{2}\otimes 2\otimes 1\otimes 0$
$\hat{\eta}_1$	1	$0 \otimes 1 \otimes 1 \otimes -1$
$\hat{\eta}_2$	1	$0 \otimes 1 \otimes 1 \otimes 1$

TABLE II. Chiral superfields of the BLSSM and their charges under $U(1)_Y \otimes SU(2)_L \otimes SU(3)_c \otimes U(1)_{B-L}$ gauge group.

conditions to reach a focus-point like behaviour in the new sector of the model. We show that the new focus point can co-exist with the well-know MSSM focus point resulting in a 'Double Focus Point' (DFP) scenario. We also discuss the radiative corrections to the fine-tuning. These corrections cause significant deviations in the actual fine-tuning prediction from the tree-level estimate $\Delta_{\mu_{\eta}}^{Z'} \sim \mu_{\eta}^2/m_{Z'}^2$.

The rest of paper is organized as follows: in section II we present the details of the considered model. In section III we show the analytical derivation of the DFP and calculate the dominant radiative corrections to the fine-tuning. In section IV we perform a numerical study of the fine-tuning to validate our analytical results. Afterwards, we analyse the impact of the ⁸Be anomaly. We conclude in section V.

II. THE $U(1)_{B-L}$ EXTENDED MSSM

In the simplest $U(1)_{B-L}$ extension of the MSSM, the chiral superfields are extended by a pair bileptons $(\hat{\eta}_1, \hat{\eta}_2)$ and three generations of right-handed neutrino superfield $\hat{\nu}_{R_i}$. The complete particle contents and charge assignments are listed in table I and II. This model is known as the BLSSM and its superpotential is given by

$$W = Y_u^{ij} \hat{U}_i \hat{Q}_j \hat{H}_u - Y_d^{ij} \hat{D}_i \hat{Q}_j \hat{H}_d - Y_e^{ij} \hat{E}_i \hat{L}_j \hat{H}_d + \mu \hat{H}_u \hat{H}_d$$
$$Y_{\eta}^{ij} \hat{\nu}_{Ri} \hat{\eta}_1 \hat{\nu}_{Rj} + Y_{\nu}^{ij} \hat{L}_i \hat{H}_u \hat{\nu}_{Rj} - \mu_{\eta} \hat{\eta}_1 \hat{\eta}_2 . \tag{5}$$

 $^{^2}$ We don't make in the following any assumption how the different fine-tunings shall be combined, but we discuss them separately. Often, the tunings are dominates by the different μ -terms. In these cases the overall fine-tuning is given by $\Delta = \Delta_{\mu}^Z \times \Delta_{\mu_{\eta}}^{Z'}$. The situation is more complicated if the same parameter (like an universal scalar mass m_0) contributes significantly to the tuning in both sectors at the same time.

Here i, j denote family indices and all colour and isospin indices are suppressed. The soft-breaking terms are

$$\mathcal{L}_{BLSSM} = \mathcal{L}_{MSSM} - M_{BB'} \lambda_{\tilde{B}} \lambda_{\tilde{B}'} - \frac{1}{2} M_{B'} \lambda_{\tilde{B}'} \lambda_{\tilde{B}'}$$
$$- m_1^2 |\eta_1|^2 - m_2^2 |\eta_2|^2 - m_{\nu,ij}^2 (\tilde{\nu}_{Ri}^c)^* \tilde{\nu}_{Rj}^c$$
$$- B_{\mu_\eta} \eta_1 \eta_2 + T_{\nu'}^{ij} H_u \tilde{\nu}_{Ri}^c \tilde{L}_j + T_{\eta}^{ij} \eta_1 \tilde{\nu}_{Ri}^c \tilde{\nu}_{Rj}^c$$
(6)

After Higgs states and bileptons receive vacuum expectation values (VEVs), the electroweak and $U(1)_{B-L}$ symmetry are broken to $U(1)_{em}$. After symmetry breaking, the complex scalars are parametrised by

$$H_d^0 = \frac{1}{\sqrt{2}} \left(i\sigma_d + v_d + \phi_d \right) , H_u^0 = \frac{1}{\sqrt{2}} \left(i\sigma_u + v_u + \phi_u \right) ,$$

$$\eta_1 = \frac{1}{\sqrt{2}} \left(i\sigma_1 + v_1 + \phi_1 \right) , \eta_2 = \frac{1}{\sqrt{2}} \left(i\sigma_2 + v_2 + \phi_2 \right)$$
(7)

Following the MSSM definition $\tan \beta = v_u/v_d$, we denote the ratio of the two bilepton VEVs as $\tan \beta' = v_1/v_2$.

The particle content of the BLSSM gives rise to gauge-kinetic mixing even if it is absent at a given scale. This introduces two additional gauge couplings $g_{YB} \equiv g_{YB-L}$ and $g_{BY} \equiv g_{B-LY}$, i.e. the general form of the covariant derivatives is

$$D_{\mu}\phi = \left(\partial_{\mu} - i\sum_{i,j} Q_{\phi}^{i} g_{ij} V_{j}^{\mu}\right) \phi \qquad i, j = Y, B - L$$
(8)

Here, Q_{ϕ}^{i} is the U(1) charge of the particle ϕ under the gauge group $U(1)_{i}$ (i = Y, B - L).

However, as long as the two Abelian gauge groups are unbroken, we are allowed to make a change of basis. This freedom is used go to a basis where electroweak precision data is respected in a simple way: by choosing a triangle form of the gauge coupling matrix, the bilepton contributions to the Z mass vanish:

$$\begin{pmatrix} g_{YY} & g_{YB} \\ g_{BY} & g_{BB} \end{pmatrix} \to \begin{pmatrix} g_1 & \tilde{g} \\ 0 & g_B \end{pmatrix} , \qquad (9)$$

and the gauge couplings are related by [50]:

$$g_{1} = \frac{g_{YY}g_{BB} - g_{YB}g_{BY}}{\sqrt{g_{BB}^{2} + g_{BY}^{2}}} ,$$

$$\tilde{g} = \frac{g_{YB}g_{BB} + g_{BY}g_{YY}}{\sqrt{g_{BB}^{2} + g_{BY}^{2}}} ,$$

$$g_{B} = \sqrt{g_{BB}^{2} + g_{BY}^{2}} . \tag{10}$$

In addition, After electroweak and $U(1)_{B-L}$ breaking, the gauge-kinetic mixing further induces a mixing between the neutral SUSY particles from the MSSM and from the new sector, i.e. there are seven neutralinos in this model. In the gauge sector, the three neutral gauge bosons B, W^3 and B' are rotated to the mass eigenstates γ , Z, Z' via:

$$\begin{pmatrix} B \\ W \\ B' \end{pmatrix} = R(\theta, \ \theta') \begin{pmatrix} \gamma \\ Z \\ Z' \end{pmatrix} , \tag{11}$$

where the rotation matrix $R(\theta, \theta')$ depends on two angles θ and θ' with following expression

$$\begin{pmatrix}
\cos \theta_W & -\cos \theta'_W \sin \theta_W & \sin \theta_W \sin \theta'_W \\
\sin \theta_W & \cos \theta_W \cos \theta'_W & -\cos \theta_W \sin \theta'_W \\
0 & \sin \theta'_W & \cos \theta'_W
\end{pmatrix} . (12)$$

The entire mixing between the $U(1)_{B-L}$ and the SM gauge depends on mixing angle θ' , which can be approximately expressed as [51]

$$\tan 2\theta'_W \simeq \frac{2\tilde{g}\sqrt{g_1^2 + g_2^2}}{\tilde{g}^2 + 16\left(v_\eta/v\right)^2 g_B^2 - g_2^2 - g_1^2} \ . \tag{13}$$

with
$$v = \sqrt{v_u^2 + v_d^2}$$
 and $v_{\eta} = \sqrt{v_1^2 + v_2^2}$.

III. FOCUS POINT BEHAVIOUR AND LOOP CORRECTED FINE-TUNING IN THE BLSSM

After setting up the model, we begin to investigate the focus point property. Compared to the MSSM, the tadpole equations become more complicated and we just present the most relevant ones:

$$m_{H_d}^2 = \frac{1}{8}D_H + \frac{1}{\tan\beta}B_\mu - \mu^2 \tag{14}$$

$$m_{H_u}^2 = -\frac{1}{8}D_H + \tan\beta B_\mu - \mu^2 \tag{15}$$

$$m_{\eta_1}^2 = \frac{1}{4} D_{\eta} + \frac{1}{\tan \beta'} B_{\mu_{\eta}} - \mu_{\eta}^2 ,$$
 (16)

$$m_{\eta_2}^2 = -\frac{1}{4} D_{\eta} + \tan \beta' B_{\mu_{\eta}} - \mu_{\eta}^2 .$$
 (17)

with $D_H = (g_1^2 + \tilde{g}^2 + g_2^2)v^2\cos 2\beta + 2\tilde{g}g_Bv_\eta^2\cos 2\beta'$ and $D_\eta = (v^2\cos 2\beta g_B\tilde{g} + 2v_\eta^2\cos 2\beta'g_B^2)$. It is easy to see that for small kinetic mixing coupling \tilde{g} , the MSSM sector and $U(1)_{B-L}$ sector are decoupled. Therefore, the little hierarchy problems for two sectors can be treated separately. Nevertheless, there will be a non-trivial link between the fine-tuning of both sectors because of the relations of soft-breaking terms at the GUT scale. Under this assumption eq. (14) reduces to the standard MSSM tadpole equation and the little hierarchy problem can be handled as usual. In the B-L sector we have to distinguish the cases of small $\tan \beta'$ ($\simeq 1$) and large $\tan \beta'$ ($\gg 1$) which we discuss separately in the following.

A. Small $\tan \beta'$

In analogy to the MSSM, the tadpole equation of the B-L sector, eq. (17), can be simplified for $\tan \beta' \simeq 1$ to

$$m_{Z'}^2 = -m_{\eta_2}^2 - \mu_{\eta}^2 ,$$
 (18)

Here, we used $m_{Z'}^2 = g_B^2 v_\eta^2$. The important question is now, if it possible to obtain naturally values for m_η

and μ_{η} which are significantly smaller than the ordinary SUSY parameters during the renormalisation group equation (RGE) evolution. Or, to phrase it differently: is it possible to find a focus point behaviour? To answer this question, one needs to check the running of $m_{\eta_2}^2$. The one-loop beta function is given by

$$\frac{dm_{\eta_2}^2}{dt} = \frac{g_B^2}{16\pi^2} (-12M_{BB'}^2 + M_{B'}^2) \ . \tag{19}$$

where $M_{BB'}$ is a soft-breaking term mixing the two gaugino fields. Due to the absence of any Yukawa interaction in the running, it is impossible to obtain a focus point behaviour in this scenario. m_{η_2} will always increase during the evaluation from the GUT to the SUSY scale. Thus, in this simplest realisation, it is not possible to obtain a DFP for small $\tan \beta'$. However, this would become possible in the singlet extension of the model (N-BLSSM), with an additional superpotential term

$$\delta W = \lambda \, \hat{S} \, \hat{\eta}_1 \, \hat{\eta}_2 \tag{20}$$

 λ will give new contributions to the running of $m_{\eta_2}^2$. We discuss this briefly in appendix A.

B. Large $\tan \beta'$

We turn now to the case of large $\tan \beta'$ which is usually not considered in the case of heavy Z' masses. The reason is that the new D-term contributions to the sfermion masses result in tachyonic states once $\tan \beta'$ is too large. However, this is not the case for a light Z'. The corresponding tadpole equation in this case is given by

$$m_{Z'}^2 = -m_{\eta_1}^2 - \mu_{\eta}^2 \ . \tag{21}$$

The one-loop running of $m_{\eta_1}^2$ is given by

$$\frac{dm_{\eta_1}^2}{dt} = \frac{y_{\eta}^2}{8\pi^2} (4m_{\eta_1}^2 + 4m_{\nu_R}^2 + 4A_{\eta}^2) \ . \tag{22}$$

Here the soft trilinear term $A_{\eta}=T_{\eta}^{ij}/Y_{\eta}^{ij}$. We assumed here that y_{η} as well as $m_{\nu_R}^2$ are diagonal and degenerated. We defined $y_{\eta}\equiv Y_{\eta}^{ii}$. We can now investigate the necessary conditions to obtain a focus point in the running. Before we proceed, the following comments are at place

- We assume the beta functions of $m_{H_u}^2$ and $m_{\eta_1}^2$ to be independent from each other. This is justified as long as gauge kinetic mixing is small.
- Since the beta function of $m_{H_u}^2$ depends on m_q^2 , m_u^2 and A_t^2 , we must solve the coupled system of equations. In the B-L sector we need to consider m_{η_1} , A_{η} and m_{ν_R} .
- In contrast to the MSSM sector, where the top Yukawa coupling is fixed by experiment, we can treat y_{η} as free parameter. If we demand perturbativity up to $M_{\rm GUT}$), the maximal allowed value for y_{η} is about 0.42.

The relevant one-loop beta functions can be written into matrix form. The MSSM part is given by

$$\frac{d}{dt} \begin{bmatrix} m_{H_u}^2 \\ m_u^2 \\ m_g^2 \\ A_t^2 \end{bmatrix} = \frac{y_t^2}{8\pi^2} \begin{bmatrix} 3 & 3 & 3 & 3 \\ 2 & 2 & 2 & 2 \\ 1 & 1 & 1 & 1 \\ 0 & 0 & 0 & 12 \end{bmatrix} \begin{bmatrix} m_{H_u}^2 \\ m_u^2 \\ m_g^2 \\ A_t^2 \end{bmatrix} ,$$
(23)

and the BLSSM one by

$$\frac{d}{dt} \begin{bmatrix} m_{\eta_1}^2 \\ m_{\nu_R}^2 \\ A_x^2 \end{bmatrix} = \frac{y_{\eta}^2}{8\pi^2} \begin{bmatrix} 6 & 12 & 6 \\ 4 & 8 & 4 \\ 0 & 0 & 28 \end{bmatrix} \begin{bmatrix} m_{\eta_1}^2 \\ m_{\nu_R}^2 \\ A_\eta^2 \end{bmatrix}$$
(24)

The coupled beta functions can be solved in terms of eigenvectors and eigenvalues. We obtain

$$\begin{bmatrix} m_{H_u}^2 \\ m_u^2 \\ m_q^2 \\ A_t^2 \end{bmatrix} = \kappa_{12} \begin{bmatrix} 3 \\ 2 \\ 1 \\ 6 \end{bmatrix} e^{12I[t]} + \kappa_6 \begin{bmatrix} 3 \\ 2 \\ 1 \\ 0 \end{bmatrix} e^{6I[t]} + \kappa_0 \begin{bmatrix} -1 \\ 0 \\ 1 \\ 0 \end{bmatrix} + \kappa_0 \begin{bmatrix} -1 \\ 1 \\ 0 \\ 0 \end{bmatrix}, \quad (25)$$

as well as

$$\begin{bmatrix} m_{\eta_1}^2 \\ m_{\nu_R}^2 \\ A_{\eta}^2 \end{bmatrix} = \epsilon_{28} \begin{bmatrix} 3 \\ 2 \\ 7 \end{bmatrix} e^{28K[t]} + \epsilon_{14} \begin{bmatrix} 3 \\ 2 \\ 0 \end{bmatrix} e^{14K[t]} + \epsilon_0 \begin{bmatrix} -2 \\ 1 \\ 0 \end{bmatrix}$$
(26)

with so far arbitrary coefficients κ_i and ϵ_i . The functions I and K are defined in Appendix B. Approximate values are $\mathrm{e}^{14K[t]} \simeq 1/10$ and $\mathrm{e}^{6I[t]} \simeq 1/3$. This results in

$$\begin{bmatrix} m_{H_u}^2[Q_{\text{GUT}}] \\ m_q^2[Q_{\text{GUT}}] \\ m_u^2[Q_{\text{GUT}}] \\ A_t^2[Q_{\text{GUT}}] \end{bmatrix} = \begin{bmatrix} m_0^2 \\ m_0^2 + \kappa_0' - \frac{2\kappa_{12}}{3} \\ m_0^2 - \kappa_0' - \frac{4\kappa_{12}}{3} \end{bmatrix}$$

$$\rightarrow \qquad \qquad \rightarrow$$

$$\begin{bmatrix} m_{H_u}^2[Q_{\text{SUSY}}] \\ m_q^2[Q_{\text{SUSY}}] \\ m_u^2[Q_{\text{SUSY}}] \\ A_t^2[Q_{\text{SUSY}}] \end{bmatrix} = \begin{bmatrix} 0 \\ \frac{m_0^2}{3} + \kappa_0' - \frac{2\kappa_{12}}{5} \\ \frac{2m_0^2}{3} + \kappa_0' - \frac{4\kappa_{12}}{5} \end{bmatrix} , \qquad (27)$$

and

$$\begin{bmatrix} m_{\eta_{1}}^{2}[Q_{\text{GUT}}] \\ m_{\nu_{R}}^{2}[Q_{\text{GUT}}] \\ A_{\eta}^{2}[Q_{\text{GUT}}] \end{bmatrix} = \begin{bmatrix} m_{0}^{2} \\ \frac{43}{54}m_{0}^{2} - \frac{7}{20}\epsilon_{28} \end{bmatrix}$$

$$\rightarrow \begin{bmatrix} m_{\eta_{2}}^{2}[Q_{\text{SUSY}}] \\ m_{\nu_{R}}^{2}[Q_{\text{SUSY}}] \\ A_{\eta}^{2}[Q_{\text{SUSY}}] \end{bmatrix} = \begin{bmatrix} 0 \\ \frac{7}{54}m_{0}^{2} - \frac{7}{20}\epsilon_{28} \\ \frac{7}{100}\epsilon_{28} \end{bmatrix}$$
 (28)

Eq. (A5) and (28) show that $m_{H_u}^2$ and $m_{\eta_1}^2$ evolve to zero at low scale no matter what value we take for m_0 which means that the weak scale and U_{B-L} breaking scale are insensitive to variation of fundamental parameters. As a consequence, we obtain DFP SUSY even for several TeV sfermions which are induced by large m_0 . A few more comments:

- 1. Besides m_0 , there are three parameters: κ_{12} , κ'_0 , and ϵ_{28} . Here κ_{12} and ϵ_{28} represent large A-term generated by gravity mediation.
- 2. The parameters κ'_0 and κ_{14} give deviation from the soft masses are predicted by minimal supergravity. The source of the deviation could be for instance hybrid anomaly or gauge mediation.
- 3. The parameters κ_{12} , κ_0 and ϵ_{28} are all dimensionful. They can be related to m_0 via dimensionless parameters x,y and z respectively. This turns out to be helpful for the numerical calculation.

C. Radiative corrections to the fine-tuning

Up to now, we have shown that it is possible to find naturally parameter regions in which $|\mu|$ and $|\mu_n|$ at the same time are significantly smaller than the ordinary SUSY scale. Nevertheless, μ_{η} is still expected to be of the same size as μ , i.e. O(100 GeV). In that case, the tree-level estimate for the fine-tuning would be $\Delta_{\mu_{\eta}} \sim \mu_{\eta}^2/m_Z^{\prime,2}$ which predicts values for the fine-tuning above 10^6 for $\mu_{\eta} > 100$ GeV and $m_Z' < 100$ MeV. In that case, the overall fine-tuning would be completely dominated by the new sector and all the considerations about DFP wouldn't have been necessary at all. However, it has rather recently been pointed out in Ref. [3] that loop corrections to the fine-tuning are very important in the MSSM. The same kind of corrections is even more important here. The starting point of the discussion is the one-loop corrected tadpole equation which is in general given by

$$0 = m_{\eta_1}^2 v_{\eta} + \mu_{\eta}^2 v_{\eta} + \frac{1}{2} g_B^2 v_{\eta}^3 + \delta t_{\eta}$$
 (29)

with

$$\delta t_{\eta} = \frac{\partial V^{(1)}}{\partial v_{\eta}} \tag{30}$$

Here, $V^{(1)}$ is the one-loop effective potential which can be calculated as usual as [52]

$$V^{(1)} = \frac{1}{16\pi^2} \sum_{i}^{\text{all fields}} r_i s_i C_i m_i^4 \left(\log \frac{m_i^2}{Q^2} - c_i \right)$$
 (31)

with $r_i=1$ for real bosons, otherwise 2; C_i is a colour factor; $\{s_i,c_i\}=\{-\frac{1}{2},\frac{3}{2}\}$ for fermions, $\{\frac{1}{4},\frac{3}{2}\}$ for scalars and $\{\frac{3}{4},\frac{5}{6}\}$ for vector bosons. In the case of large y_η as needed for DFP, the dominant contributions are due to right (s)neutrinos.

In general, δt_{η} can be parametrised as

$$\delta t_{\eta} = v_{\eta} \delta_1 + v_{\eta}^2 \delta_2 + v_{\eta}^3 \delta_3 \tag{32}$$

Because of symmetry reasons, δ_2 always vanishes. Therefore, the fine-tuning with respect to μ_{η} becomes

$$\Delta_{\mu_{\eta}} = \frac{\mu_{\eta}^2}{(g_B^2 + 2\delta_3)v_\eta^2} \tag{33}$$

Thus, for a reliable calculation of the fine-tuning, the knowledge of δ_3 is crucial. One fiends for three generations of degenerated right sneutrinos with masses $O(M_{\rm SUSY})$ that δ_3 at the SUSY scale is given by

$$\delta_3 = -\frac{3y_\eta^4}{2\pi^2} \log \frac{2v_\eta^2 y_\eta^2}{M_{\text{SUSY}}^2} \tag{34}$$

From this one obtains an improvement in the fine-tuning at the loop level of

$$\frac{\Delta_{\mu_{\eta}}^{\text{Tree}}}{\Delta_{\mu_{n}}^{\text{Loop}}} \sim \frac{3y_{\eta}^{4}}{g_{B}^{2}\pi^{2}} \log \frac{g_{B}^{2} M_{\text{SUSY}}^{2} + 2m_{Z}^{\prime,2} y_{\eta}^{2}}{2m_{Z}^{\prime,2} y_{\eta}^{2}}$$
(35)

The gain in the fine-tuning as function of y_{η} and g_{B} for $m_{Z'} = 100 \text{ MeV}$ is shown in Figure 1. One might be surprised by these huge changes in the fine-tuning due to the radiative corrections. However, the same radiative corrections push also the mass of the light B-Lscalar, which is at tree-level $O(m_{Z'})$, into the multi GeV range. Thus, the fine-tuning regarding the new vector boson mass $(\partial \ln m_{Z'}^2/\partial \ln \mu_{\eta})$ becomes comparable to the fine-tuning regarding the pole mass of the scalars $(\partial \ln \log m_{h_{\eta}}^2/\partial \ln \mu_{\eta})$. Moreover, since higher order corrections don't modify the general form of the radiatively corrected tadpole equation but are only corrections to the coefficients c_i , similar huge changes in the fine-tuning won't occur by going to a higher loop level. Therefore, the numerical most important effects are already caught by the one-loop corrections 3 .

³ It has been discussed in the context of other models that the fine-tuning measure with respect to m_Z can clearly deviate from a measure with respect to m_h [53]. It would be interesting to see if a proper inclusion of loop corrections in the fine-tuning calculation reduces this discrepancy.

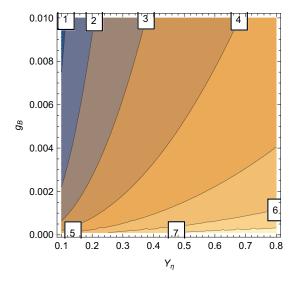


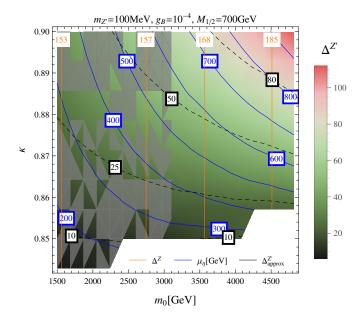
FIG. 1. The improvement in the fine-tuning as function of g_B and y_η when including radiative corrections to the tadpole equations. The lines correspond to contours of constant $\log_{10}\left(\Delta_{\mu_\eta}^{\rm Tree}/\Delta_{\mu_\eta}^{\rm Loop}\right)$. We used here $m_{Z'}=100$ MeV and $M_{\rm SUSY}=2.5$ TeV.

IV. NUMERICAL RESULTS

A. General results

We are going to compare the analytical results of the last section with a fully numerical calculation to check the validity of our results. For this purpose, we have implemented the considered model into the Mathematica package SARAH [54–58]⁴. We used this implementation to generate a spectrum generator for the model based on SPheno [59, 60]. SPheno solves numerically the full twoloop RGEs, calculates the mass spectrum at the full oneloop level and includes all important two-loop correction to the neutral scalar masses [61–63]. In order to keep the uncertainty of the Higgs mass to a low level also in the presence of very heavy SUSY scales, SPheno provides an effective calculation within the SM [64] where all SUSY effects are absorbed into λ via a pole mass matching of the Higgs masses at the SUSY scale [65]. Also a routine to obtain the electroweak fine-tuning is available out-ofthe-box. This routine has been extended to calculate also the fine-tuning with respect to Z' as

$$\Delta^{Z'} = \max \left| \frac{\partial \ln m_{Z'}^2}{\partial \ln \alpha} \right| \tag{36}$$



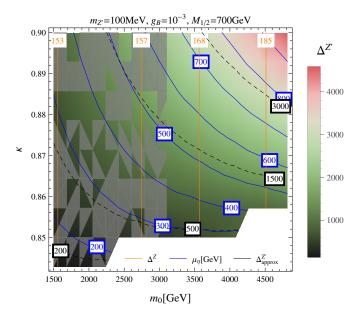


FIG. 2. The fine-tuning in the (m_0, κ) plane for $m_{Z'} = 100$ MeV and $g_B = 10^{-4}$ (top) or $g_B = 10^{-3}$ (bottom). We set $M_{1/2} = 700$ GeV, $\tan \beta = 20$, $\tan \beta' = 10$, $Y_{\eta} = 0.4$, $s_y = -0.1$, $s_z = s_x = 0$, $\tilde{g}(Q_{\rm GUT}) = 0$. In the gray shaded areas the SM-like Higgs doesn't fall into the desired range of [122,128] GeV. The orange line is the electroweak fine-tuning with respect m_Z (which is dominated by $M_{1/2}$), the blue lines show μ_{η} (in GeV) and the dashed black line give the calculated fine-tuning using the approximations eqs. (33) and (34).

with $\alpha = \{m_0, M_{1/2}, \mu'\}$. To obtain the loop corrected fine-tuning, we calculate the one-loop corrected tadpole equations with a diagrammatic approach and solve them numerically with respect to all four VEVs using a broydn routine. The VEVs obtained in that way are then used to calculate the dependence of the vector boson masses

⁴ We could use the already existing implementation of the BLSSM which we had slightly to modify: since the eigenstates are mass ordered, the definition of the neutral gauge boson mixing was changes from $\gamma - Z - Z'$ to $\gamma - Z' - Z$ and the boundary conditions needed to be adjusted.

on a finite variation of the parameters at the GUT scale. We checked that the fine-tuning obtained in that way is independent of the chosen size of the finite variation if a sufficiently small value of 10^{-6} is chosen. All parameter scans have been carried out using the package SSP [66]. One important question is how accurate the coefficient of $\frac{43}{54} \simeq 0.8$ for the soft-term of the right sneutrino soft-term is at the GUT scale. So far, we discussed only the one-loop running, but haven't considered the impact of two-loop RGEs. In order to check this, we use a more general parametrisation

$$m_{\nu_R}^2[Q_{\text{SUSY}}] = \kappa \, m_0^2 \, \mathbf{1}$$
 (37)

and treat κ is free parameter. The other GUT conditions are

$$m_{H_u}^2 = m_{H_d}^2 = m_{\eta_1}^2 = m_{\eta_2}^2 \equiv m_0^2$$

$$m_l^2 = m_d^2 = m_e^2 \equiv m_0^2 \mathbf{1}$$

$$m_{q,11}^2 = m_{q,22}^2 = m_{u,11}^2 = m_{u,22}^2 \equiv m_0^2$$

$$m_{q,33} \equiv (1-x)m_0^2 \qquad m_{u,33} \equiv (1+x-3y)m_0^2$$

$$M_1 = M_2 = M_3 = M_B \equiv M_{1/2}$$

$$T_i \equiv 3ym_0 Y_i (i=e,d,u) \qquad T_\eta \equiv \sqrt{7}zY_\eta \qquad (38)$$

The calculated fine-tuning in the (m_0, κ) plane for a Z'mass of 100 MeV and different values of g_B is shown in Figure 2. We find that the numerical calculation confirms our analytical results to a large extent: the focus point behaviour can be observed for κ values below 0.86 even for large m_0 . Also the numerically calculated fine-tuning agrees with our analytical approximation. It is also found that the fine-tuning has a strong dependence on the value of g_B for fixed $m_{Z'}$ as expected from eq. (35). These results confirm that a light Z' mass in supersymmetric models doesn't lead unavoidably to a huge fine-tuning as a one might expect. However, two conditions must be fulfilled: (i) the presence of a focus point, (ii) a Yukawa-like coupling to the scalars which give mass to the Z' which is much bigger than the corresponding gauge coupling. A more detailed parameter scan of the model is beyond the scope of this paper and we will only discuss one more aspect: what is the size of the expected fine-tuning if the ⁸Be should be explained within this model.

B. The ⁸Be anomaly

It has been shown in Ref. [32] that the 8 Be could be explained by a B-L gauge boson with a mass of about 17 MeV. However, strong constraints on the new couplings exist, especially on the one induced by gauge kinetic mixing. In general, the couplings to the SM fermions for the B-L vector boson are given by

$$g_u = \frac{1}{3}g_B + \frac{2}{3}\tilde{g} , \qquad g_{\nu} = -g_B$$

 $g_d = \frac{1}{3}g_B - \frac{1}{3}\tilde{g} , \qquad g_e = -g_B - \tilde{g} .$ (39)

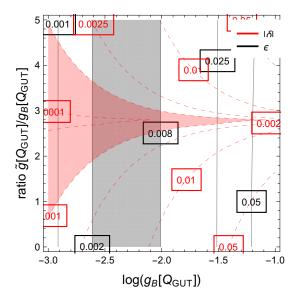


FIG. 3. Dependence of the ratio between g_0 and g_d on g_d at GUT scale. Regions shaded in grey are the allowed region that explains Be anomaly. The red dashed contour corresponds the abstract value of δ and black corresponds to ϵ .

which can be rewritten to

$$g_u = -\frac{1}{3}\epsilon + \frac{2}{3}\delta , \qquad g_\nu = -\epsilon$$

$$g_d = +\frac{2}{3}\epsilon - \frac{1}{3}\delta , \qquad g_e = -\delta . \qquad (40)$$

by defining $g_B = \epsilon$ and $\delta = -g_B + \tilde{g}$. The authors in Ref. [32] showed that, if the signal is real, δ and ϵ must fulfil

$$0.002 < |\epsilon| < 0.008$$
, $|\delta| < 0.001$ (41)

These constraints for ϵ and δ can be translated into constraints on the gauge couplings at the GUT scale. The preferred parameter ranges for $g_B(GUT)$ and $\tilde{g}(GUT)$ are shown in Fig. 3. Here, we used the two-loop SUSY RGEs with gauge kinetic mixing as calculated by SARAH based on the generic results of Ref. [67]. One can see that g_B must be smaller than 0.01 at the GUT scale. In contrast, \tilde{g} must be bigger by a factor 2 to 4. This is already an unexpected hierarchy in the gauge couplings which might be hard to realize in a full model like $E_6 \times E_6$ which gets broken to $G_{SM} \times U(1)_{B-L}$ [68–70]. Nevertheless, we take these parameter ranges as given. Since g_B and \tilde{g} are still sufficiently small to have only a weak link between the MSSM and BLSSM section, the discussion in the last section about the DFP behaviour remains fully valid. However, as expected from the analytical discussion and from the general results in the last subsection one must expect that the fine-tuning is rather large: for the values of q_B necessary to explain the anomaly one expects only an improvement of the fine-tuning at the loop level by

	Input									
m_0	$M_{1/2}$	x	y	z	κ					
4600	680	0.22	-0.45	0.1	0.83					
$\tan \beta$	$\tan \beta'$	y_{η}	$m_{Z'}$	g_B	$ ilde{g}$					
18.5	10	0.41	0.017	0.01	0.003					
	Running parameters									
μ	μ_{η}	g_B	$ ilde{g}$							
613	395	0.0036	0.0026							
	Masses									
h_1	h_2	h_3	h_4	A_1^0	A_2^0					
1.4	122.4	4236	4640	4264	4640					
$ ilde{t}_1$	$ ilde{t}_2$	$ ilde{q}_{1,2}$	$ ilde{G}$	$ ilde{l}$	$ ilde{ u}_R$					
2063	3024	\sim 4700	1750	$\sim \! 4500$	~ 2000					
$ ilde{\chi}_1^0$	$\tilde{\chi}^0_{2,3}$	$ ilde{\chi}_4^0$	$ ilde{\chi}^0_5$	$ ilde{\chi}^0_6$	$ ilde{\chi}_{7}^{0}$					
284	397	531	611	657	680					
	Fine-Tuning									
$\Delta_{m_0}^Z$	$\Delta^Z_{M_{1/2}}$	Δ^Z_μ	$\Delta_{m_0}^{Z'}$	$\Delta^{Z'}_{\mu_{\eta}}$						
83	172	45	3347	3695						

TABLE III. An example for a parameter point explaining the ⁸Be anomaly within the BLSSM. All dimensionful parameters are given in units of GeV.

a few orders of magnitude, i.e. the tiny Z' mass still causes an enormous fine-tuning even when the radiative corrections are included.

We performed a random scan for this model fixing $m_{Z'}$ as well as the new gauge couplings. The features of an representative parameter point which is in agreement with the Higgs mass measurement and which could explain the ⁸Be anomaly is summarised in Table III. The fine-tuning in the MSSM sector is O(170) and dominated by $M_{1/2}$ because of the gluino mass limit. Thus, to further improve this tuning, it would be necessary to give up gaugino unification to find a 'gaugino focus point' [12, 71]. In the B-L sector the tuning with respect to m_0 and μ_{η} are of similar size and above 3000. This is not terribly good, but better than one might have expected. One possibility to further improve the tuning would be to assume a cut-off scale well below 10^{16} GeV: in that case the tuning with respect to m_0 becomes smaller because of the shorter running, and the tuning with respect to μ_{η} could be further reduce by larger y_{η} because of the relaxed perturbativity condition.

V. CONCLUSION

In this paper, we have considered the naturalness in supersymmetric models with a light Z' gauge boson. We have shown that the additional fine-tuning due to the new sector can be much smaller than the expected value $O(M_{\rm SUSY}^2/m_{Z'}^2)$. Two mechanisms are used to reduce the fine-tuning. First, it is possible to find relations of the SUSY breaking parameters at the GUT scale for which

the a focus point in the MSSM and in the new sector coexist. We call this Double Focus Point Supersymmetry. Second, we have discussed the importance of radiative corrections to the fine-tuning calculation which can alter the prediction by many orders of magnitude. For both effects one needs a large Yukawa-like (Y) coupling to the scalars which are responsible for the gauge symmetry breaking of the new group. We have discussed this explicitly at the example of the $U(1)_{B-L}$ extended MSSM (BLSSM). In particular, the radiative corrections alter the prediction of the fine-tuning by a factor Y^4/g^2 , where g is the new gauge coupling. We have confirmed this analytical estimate by a fully numerical calculation. We found that for $m_{Z'} = 100$ MeV an additional finetuning of O(100) is easily possible for $g \simeq 10^{-4}$. Finally, we have considered the ⁸Be anomaly in this model. Because of the necessary coupling strength to explain this excess in the BLSSM, a fine-tuning with respect to Z'above 10^3 seems to be unavoidable as long as perturbativity up to $M_{\rm GUT} \sim 10^{16}$ GeV is demanded.

We expect that similar features are present in other SUSY models with extended gauge sectors which involve potentially large Yukawa-like couplings to the new scalars like left-right models [72, 73] or separately gauged baryon and lepton number [74].

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Appendix A: NBLSSM with small mixing

We now turn to calculate the DFP for $m_{H_u}^2$ and $m_{\eta_2}^2$ in N-BLSSM. After taking the gauge coupling and Yukawa coupling Y_{ν} to zero, the total one-loop beta functions for $m_{H_u}^2$ and $m_{\eta_2}^2$ are given as

$$\begin{split} \beta_{m_{H_u}^2} &= \frac{1}{16\pi^2} \left(6m_{H_u}^2 y_t^2 + 6m_q^2 y_t^2 + 6m_u^2 y_t^2 + 6T_t^2 \right) \\ \beta_{m_{\eta_2}^2} &= \frac{1}{16\pi^2} \left(2m_{\eta_2}^2 \lambda^2 + 2m_{\eta_1}^2 \lambda^2 + 2m_s^2 \lambda^2 + 4T_\lambda^2 \right) \end{split} \tag{A1}$$

where $T_t = y_t A_t$ and $T_{\lambda} = \lambda A_{\lambda}$ are the soft trilinear terms. The calculation of the MSSM section is completely analog to the BLSSM. The matrix form of the RGEs in the B-L sector are

This can be solved as

$$\begin{bmatrix} m_{\eta_2}^2 \\ m_{\eta_1}^2 \\ m_s^2 \\ A_\lambda^2 \end{bmatrix} = \lambda_6 \begin{bmatrix} 1 \\ 1 \\ 1 \\ 3 \end{bmatrix} e^{6J[t]} + \lambda_3 \begin{bmatrix} 1 \\ 1 \\ 1 \\ 0 \end{bmatrix} e^{3J[t]} + \lambda_0 \begin{bmatrix} -1 \\ 0 \\ 1 \\ 0 \end{bmatrix} + \lambda_0' \begin{bmatrix} -1 \\ 1 \\ 0 \\ 0 \end{bmatrix} . \tag{A3}$$

The interesting point is that the exponent $e^{6I[t]}$ remains approximately 1/3 even though considering the extended gauge group. Furthermore the exponent $e^{3J[t]}$ is approximately 4/5. For the derivation of $e^{6I[t]}$ and $e^{3J[t]}$, see appendix B. The subtle point is that the $e^{6I[t]}$ and $e^{3J[t]}$ are only calculable when they become Bounlli type. The price we should pay is the smallness of gauge-kinetic mixing couplings i.e. $g_{BY} \sim g_{YB} \sim \mathcal{O}(10^{-2})$. For now

$$\begin{bmatrix} m_{H_u}^2[Q_{\text{GUT}}] \\ m_q^2[Q_{\text{GUT}}] \\ m_u^2[Q_{\text{GUT}}] \\ A_t^2[Q_{\text{GUT}}] \end{bmatrix} = \begin{bmatrix} m_0^2 \\ m_0^2 + \kappa_0' - \frac{2\kappa_{12}}{3} \\ m_0^2 - \kappa_0' - \frac{4\kappa_{12}}{3} \end{bmatrix}$$

$$\rightarrow$$

$$\begin{bmatrix} m_{H_u}^2[Q_{\text{SUSY}}] \\ m_q^2[Q_{\text{SUSY}}] \\ m_u^2[Q_{\text{SUSY}}] \\ A_t^2[Q_{\text{SUSY}}] \end{bmatrix} = \begin{bmatrix} 0 \\ \frac{m_0^2 + \kappa_0' - \frac{2\kappa_{12}}{3} \\ \frac{2m_0^2 + \kappa_0' - \frac{4\kappa_{12}}{5} \\ \frac{2m_0^2 + \kappa_0'}{3} + \kappa_0' - \frac{4\kappa_{12}}{5} \end{bmatrix} , \quad (A4)$$

as well as

$$\begin{bmatrix} m_{\eta_{2}}^{2}[Q_{\text{GUT}}] \\ m_{\eta_{1}}^{2}[Q_{\text{GUT}}] \\ m_{s}^{2}[Q_{\text{GUT}}] \\ A_{\lambda}^{2}[Q_{\text{GUT}}] \end{bmatrix} = \begin{bmatrix} m_{0}^{2} \\ 5m_{0}^{2} + \lambda'_{0} - \frac{4\lambda_{6}}{5} \\ 9m_{0}^{2} - \lambda'_{0} - \frac{8\lambda_{6}}{5} \end{bmatrix}$$

$$\rightarrow$$

$$\begin{bmatrix} m_{\eta_{2}}^{2}[Q_{\text{SUSY}}] \\ m_{\eta_{1}}^{2}[Q_{\text{SUSY}}] \\ m_{s}^{2}[Q_{\text{SUSY}}] \\ A_{\lambda}^{2}[Q_{\text{SUSY}}] \end{bmatrix} = \begin{bmatrix} 0 \\ 4m_{0}^{2} + \lambda'_{0} - \frac{4\lambda_{6}}{5} \\ 8m_{0}^{2} - \lambda'_{0} - \frac{8\lambda_{6}}{5} \end{bmatrix} . \quad (A5)$$

Appendix B: Appendix

We show in this appendix the derivation of $e^{6I[t]}$ and $e^{3J[t]}$ which are the integral between GUT and SUSY scale for coupling y_t and λ .

$$e^{6I[t]} = \exp\left(6 \int_{\log Q}^{\log Q_0} \frac{y_t[\rho]^2}{8\pi^2} d\log \rho\right) , \qquad (B1)$$

$$e^{3J[t]} = \exp\left(3 \int_{\log Q}^{\log Q_0} \frac{\lambda[\rho]^2}{8\pi^2} d\log \rho\right) . \qquad (B2)$$

$$e^{3J[t]} = \exp\left(3\int_{\log Q}^{\log Q_0} \frac{\lambda[\rho]^2}{8\pi^2} d\log \rho\right) .$$
 (B2)

Since y_t and λ are the dominant couplings in the model, their beta functions belong to the Bernoulli type as long as the kinetic mixing couplings become negotiable compared with other gauge couplings. This is subtle in realization of protophobic vector boson model but is suitable for SIDM with light hidden photon mediator.

$$\frac{d\alpha_a}{dt} = 2b_a \alpha_a^2 ,$$

$$\frac{d\alpha_t}{dt} = 2(s\alpha_t - \sum_a r_a \alpha_a) ,$$

$$\frac{d\alpha_\lambda}{dt} = 2(s\alpha_\lambda - \sum_a r_a \alpha_a) ,$$
(B3)

with the definition $\alpha_t=y_t^2/16\pi^2, \alpha_a=g_a^2/16\pi^2$ and $\alpha_\lambda=\lambda^2/16\pi^2$. Here s=6 for y_t and s=3 for λ . a ranges from 1 to 4. There is a formal solution for eq. (B3),

$$\alpha_t[Q] = \frac{\alpha_t[Q_0]\mathcal{E}_t[Q]}{1 - 2s\alpha_t[Q_0]\mathcal{F}_t[Q]} ,$$

$$\alpha_{\lambda}[Q] = \frac{\alpha_{\lambda}[Q_0]\mathcal{E}_{\lambda}[Q]}{1 - 2s\alpha_{\lambda}[Q_0]\mathcal{F}_{\lambda}[Q]} ,$$
(B4)

where

$$\mathcal{E}[Q] = \prod_{a} \left(1 - 2b_a \alpha_a [Q_0] \log \left(\frac{Q}{Q_0} \right) \right)^{\frac{r_a}{b_a}} ,$$

$$\mathcal{F}[Q] = \int_{\log Q_0}^{\log Q} \mathcal{E}[\rho] d \log \rho . \tag{B5}$$

 r_a and b_a can be extracted from the beta function of corresponding Yukawa coupling. Different ${\mathcal E}$ and ${\mathcal F}$ correspond to different r_a and b_a . Substituted eq. B5 into the definition of the exponent gives

$$e^{sI[t]} = \exp\left[2s \int_{\log[Q_0]}^{\log[Q]} \alpha_t[\rho] d\log[\rho]\right]$$

$$= \frac{1}{2s\alpha_t[Q_0]\mathcal{F}_t[Q]}$$

$$= 1 + \frac{2s\alpha_t[Q]\mathcal{F}_t[Q]}{\mathcal{E}_t[Q]}, \qquad (B6)$$

$$e^{sJ[t]} = \exp\left[2s \int_{\log[Q_0]}^{\log[Q]} \alpha_\lambda[\rho] d\log[\rho]\right]$$

$$= \frac{1}{2s\alpha_\lambda[Q_0]\mathcal{F}_\lambda[Q]}$$

$$= 1 + \frac{2s\alpha_\lambda[Q]\mathcal{F}_\lambda[Q]}{\mathcal{E}_\lambda[Q]}. \qquad (B7)$$

As a consequence, the ratio \mathcal{F}/\mathcal{E} determining the location of focus point. Though \mathcal{F} and \mathcal{E} are varied with the extended gauge group, the ratio remains invariant which is proven numerically. In our case when the strict gauge coupling unification is imposed, we have $\mathcal{E}_t[174] = 13.5$ and $\mathcal{F}_t[174] = -135.721$ where the top quark pole mass is chosen as low energy scale. Then the exponent of $\exp(6I[t])$ is approximately 1/3 which is the same as literature. For $\exp(3J[t])$, we not only need the ratio $\mathcal{F}_{\lambda}/\mathcal{E}_{\lambda}$

but the input value of λ at low scale. In order to escape the dangerous landau pole for λ , it is natural to set $\lambda=0.5$ at low scale. Thus the exponent $\exp(3J[t])$ is approximately 4/5. The same procedure can be applied to BLSSM with large $\tan\beta'$, the exponent $\exp(10K[t])$ is 1/10.

- S. P. Martin, (1997), [Adv. Ser. Direct. High Energy Phys.18,1(1998)], arXiv:hep-ph/9709356 [hep-ph].
- [2] R. Barbieri and G. F. Giudice, Nucl. Phys. B306, 63 (1988).
- [3] G. G. Ross, K. Schmidt-Hoberg, and F. Staub, (2017), arXiv:1701.03480 [hep-ph].
- [4] G. G. Ross, K. Schmidt-Hoberg, and F. Staub, Phys. Lett. B759, 110 (2016), arXiv:1603.09347 [hep-ph].
- [5] G. Aad et al. (ATLAS Collaboration), Phys.Lett. B716, 1 (2012), arXiv:1207.7214 [hep-ex].
- [6] S. Chatrchyan et al. (CMS), Phys. Lett. B716, 30 (2012), arXiv:1207.7235 [hep-ex].
- [7] J. L. Feng, K. T. Matchev, and T. Moroi, Phys. Rev. D61, 075005 (2000), arXiv:hep-ph/9909334 [hep-ph].
- [8] J. L. Feng, Ann. Rev. Nucl. Part. Sci. 63, 351 (2013), arXiv:1302.6587 [hep-ph].
- [9] J. L. Feng, K. T. Matchev, and D. Sanford, Phys. Rev. D85, 075007 (2012), arXiv:1112.3021 [hep-ph].
- [10] J. L. Feng and D. Sanford, Phys. Rev. **D86**, 055015 (2012), arXiv:1205.2372 [hep-ph].
- [11] J. L. Feng and K. T. Matchev, Phys. Rev. D63, 095003 (2001), arXiv:hep-ph/0011356 [hep-ph].
- [12] D. Horton and G. G. Ross, Nucl. Phys. B830, 221 (2010), arXiv:0908.0857 [hep-ph].
- [13] F. Brummer and W. Buchmuller, JHEP 05, 006 (2012), arXiv:1201.4338 [hep-ph].
- [14] T. T. Yanagida and N. Yokozaki, Phys. Lett. B722, 355 (2013), arXiv:1301.1137 [hep-ph].
- [15] K. Agashe, Phys. Rev. D61, 115006 (2000), arXiv:hep-ph/9910497 [hep-ph].
- [16] P. Draper, J. L. Feng, P. Kant, S. Profumo, and D. Sanford, Phys. Rev. D88, 015025 (2013), arXiv:1304.1159 [hep-ph].
- [17] R. Ding, T. Li, F. Staub, and B. Zhu, JHEP 03, 130 (2014), arXiv:1312.5407 [hep-ph].
- [18] D. Kim, P. Athron, C. Balázs, B. Farmer, and E. Hutchison, Phys. Rev. **D90**, 055008 (2014), arXiv:1312.4150 [hep-ph].
- [19] ATLAS (ATLAS Collaboration), ATLAS-CONF-2016-045 (2016), http://cds.cern.ch/record/2206127.
- [20] P. Athron, M. Binjonaid, and S. F. King, Phys. Rev. D87, 115023 (2013), arXiv:1302.5291 [hep-ph].
- [21] P. Athron, D. Harries, and A. G. Williams, Phys. Rev. D91, 115024 (2015), arXiv:1503.08929 [hep-ph].
- [22] H. Davoudiasl, D. E. Morrissey, K. Sigurdson, and S. Tulin, Phys. Rev. Lett. 105, 211304 (2010), arXiv:1008.2399 [hep-ph].
- [23] S. Tulin, H.-B. Yu, and K. M. Zurek, Phys. Rev. D87, 115007 (2013), arXiv:1302.3898 [hep-ph].
- [24] K. K. Boddy, J. L. Feng, M. Kaplinghat, and T. M. P. Tait, Phys. Rev. **D89**, 115017 (2014), arXiv:1402.3629 [hep-ph].

- [25] R. A. Flores and J. R. Primack, Astrophys. J. 427, L1 (1994), arXiv:astro-ph/9402004 [astro-ph].
- [26] M. Boylan-Kolchin, J. S. Bullock, and M. Kaplinghat, Mon. Not. Roy. Astron. Soc. 415, L40 (2011), arXiv:1103.0007 [astro-ph.CO].
- [27] M. Boylan-Kolchin, J. S. Bullock, and M. Kapling-hat, Mon. Not. Roy. Astron. Soc. 422, 1203 (2012), arXiv:1111.2048 [astro-ph.CO].
- [28] S. Tulin and H.-B. Yu, (2017), arXiv:1705.02358 [hep-ph].
- [29] C. Mahoney, A. K. Leibovich, and A. R. Zentner, (2017), arXiv:1706.08871 [hep-ph].
- [30] A. J. Krasznahorkay et al., Phys. Rev. Lett. 116, 042501 (2016), arXiv:1504.01527 [nucl-ex].
- [31] J. L. Feng, B. Fornal, I. Galon, S. Gardner, J. Smolinsky, T. M. P. Tait, and P. Tanedo, Phys. Rev. Lett. 117, 071803 (2016), arXiv:1604.07411 [hep-ph].
- [32] J. L. Feng, B. Fornal, I. Galon, S. Gardner, J. Smolinsky, T. M. P. Tait, and P. Tanedo, Phys. Rev. **D95**, 035017 (2017), arXiv:1608.03591 [hep-ph].
- [33] P.-H. Gu and X.-G. He, Nucl. Phys. B919, 209 (2017), arXiv:1606.05171 [hep-ph].
- [34] U. Ellwanger and S. Moretti, JHEP 11, 039 (2016), arXiv:1609.01669 [hep-ph].
- [35] Y. Liang, L.-B. Chen, and C.-F. Qiao, Chin. Phys. C41, 063105 (2017), arXiv:1607.08309 [hep-ph].
- [36] C.-H. Chen and T. Nomura, Phys. Lett. B763, 304 (2016), arXiv:1608.02311 [hep-ph].
- [37] L.-B. Jia and X.-Q. Li, Eur. Phys. J. C76, 706 (2016), arXiv:1608.05443 [hep-ph].
- [38] T. Kitahara and Y. Yamamoto, Phys. Rev. **D95**, 015008 (2017), arXiv:1609.01605 [hep-ph].
- [39] Y. Kahn, G. Krnjaic, S. Mishra-Sharma, and T. M. P. Tait, JHEP 05, 002 (2017), arXiv:1609.09072 [hep-ph].
- [40] J. Kozaczuk, D. E. Morrissey, and S. R. Stroberg, Phys. Rev. **D95**, 115024 (2017), arXiv:1612.01525 [hep-ph].
- [41] O. Seto and T. Shimomura, Phys. Rev. **D95**, 095032 (2017), arXiv:1610.08112 [hep-ph].
- [42] S. Khalil and A. Masiero, Phys. Lett. B665, 374 (2008), arXiv:0710.3525 [hep-ph].
- [43] J. E. Camargo-Molina, B. O'Leary, W. Porod, and F. Staub, Phys. Rev. D88, 015033 (2013), arXiv:1212.4146 [hep-ph].
- [44] L. Basso and F. Staub, Phys. Rev. D87, 015011 (2013), arXiv:1210.7946 [hep-ph].
- [45] L. Basso, B. O'Leary, W. Porod, and F. Staub, JHEP 09, 054 (2012), arXiv:1207.0507 [hep-ph].
- [46] F. Staub, Adv. High Energy Phys. 2015, 840780 (2015), arXiv:1503.04200 [hep-ph].
- [47] G. Bélanger, J. Da Silva, U. Laa, and A. Pukhov, JHEP 09, 151 (2015), arXiv:1505.06243 [hep-ph].
- [48] C. S. Un and O. Ozdal, Phys. Rev. D93, 055024 (2016), arXiv:1601.02494 [hep-ph].

- [49] L. Delle Rose, S. Khalil, S. J. D. King, C. Marzo, S. Moretti, and C. S. Un, (2017), arXiv:1702.01808 [hep-ph].
- [50] P. H. Chankowski, S. Pokorski, and J. Wagner, Eur. Phys. J. C47, 187 (2006), arXiv:hep-ph/0601097 [hep-ph].
- [51] L. Basso, S. Moretti, and G. M. Pruna, Phys. Rev. D82, 055018 (2010), arXiv:1004.3039 [hep-ph].
- [52] S. R. Coleman and E. J. Weinberg, Phys. Rev. D7, 1888 (1973).
- [53] A. Kaminska, G. G. Ross, K. Schmidt-Hoberg, and F. Staub, JHEP 1406, 153 (2014), arXiv:1401.1816 [hep-ph].
- [54] F. Staub, (2008), arXiv:0806.0538 [hep-ph].
- [55] F. Staub, Comput.Phys.Commun. 181, 1077 (2010), arXiv:0909.2863 [hep-ph].
- [56] F. Staub, Comput.Phys.Commun. 182, 808 (2011), arXiv:1002.0840 [hep-ph].
- [57] F. Staub, Computer Physics Communications 184, pp. 1792 (2013), arXiv:1207.0906 [hep-ph].
- [58] F. Staub, Comput.Phys.Commun. 185, 1773 (2014), arXiv:1309.7223 [hep-ph].
- [59] W. Porod, Comput. Phys. Commun. 153, 275 (2003), arXiv:hep-ph/0301101 [hep-ph].
- [60] W. Porod and F. Staub, Comput.Phys.Commun. 183, 2458 (2012), arXiv:1104.1573 [hep-ph].
- [61] M. D. Goodsell, K. Nickel, and F. Staub, Eur. Phys. J. C75, 32 (2015), arXiv:1411.0675 [hep-ph].
- [62] M. Goodsell, K. Nickel, and F. Staub, Eur. Phys. J. C75, 290 (2015), arXiv:1503.03098 [hep-ph].

- [63] J. Braathen, M. D. Goodsell, and F. Staub, (2017), arXiv:1706.05372 [hep-ph].
- [64] F. Staub and W. Porod, Eur. Phys. J. C77, 338 (2017), arXiv:1703.03267 [hep-ph].
- [65] P. Athron, J.-h. Park, T. Steudtner, D. Stöckinger, and A. Voigt, JHEP 01, 079 (2017), arXiv:1609.00371 [hep-ph].
- [66] F. Staub, T. Ohl, W. Porod, and C. Speckner, Comput.Phys.Commun. 183, 2165 (2012), arXiv:1109.5147 [hep-ph].
- [67] R. M. Fonseca, M. Malinsky, W. Porod, and F. Staub, Nucl. Phys. B854, 28 (2012), arXiv:1107.2670 [hep-ph].
- [68] W. Buchmuller, K. Hamaguchi, O. Lebedev, and M. Ratz, Nucl. Phys. B785, 149 (2007), arXiv:hep-th/0606187 [hep-th].
- [69] M. Ambroso and B. A. Ovrut, Int. J. Mod. Phys. A25, 2631 (2010), arXiv:0910.1129 [hep-th].
- [70] M. Ambroso and B. A. Ovrut, Int. J. Mod. Phys. A26, 1569 (2011), arXiv:1005.5392 [hep-th].
- [71] A. Kaminska, G. G. Ross, and K. Schmidt-Hoberg, JHEP 11, 209 (2013), arXiv:1308.4168 [hep-ph].
- [72] M. Hirsch, M. Malinsky, W. Porod, L. Reichert, and F. Staub, JHEP 1202, 084 (2012), arXiv:1110.3037 [hep-ph].
- [73] M. Hirsch, W. Porod, L. Reichert, and F. Staub, Phys.Rev. D86, 093018 (2012), arXiv:1206.3516 [hepph].
- [74] P. Fileviez Perez and M. B. Wise, Phys. Rev. D82, 011901 (2010), [Erratum: Phys. Rev.D82,079901(2010)], arXiv:1002.1754 [hep-ph].