

1     **The effect of dynamic topography and gravity on lithospheric**  
2             **effective elastic thickness estimation: a case study**

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12     **Abbreviated title:** Improving  $T_e$  estimations by dynamic modelling

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## 14 **SUMMARY**

15       Lithospheric effective elastic thickness ( $T_e$ ), a proxy for plate strength, is helpful  
16 for the understanding of subduction characteristics. Affected by curvature, faulting and  
17 magma activity, lithospheric strength near trenches should be weakened but some  
18 regional inversion studies have shown much higher  $T_e$  values along some trenches  
19 than in their surroundings. In order to improve  $T_e$  estimation accuracy, here we discuss  
20 the long-wavelength effect of dynamic topography and gravity on  $T_e$  estimation by  
21 taking the Izu-Bonin-Mariana (IBM) Trench as a case study area. We estimate the  
22 long-wavelength influence of the density and negative buoyancy of the subducting  
23 slab on observed gravity anomalies and seafloor topography. The residual topography  
24 and gravity are used to map  $T_e$  using the fan-wavelet coherence method. Maps of  $T_e$ ,  
25 both with and without the effects of dynamic topography and slab gravity anomaly,  
26 contain a band of high- $T_e$  values along the IBM Trench, though these values and their  
27 errors are lower when slab effects are accounted for. Nevertheless, tests show that the  
28  $T_e$  map is relatively insensitive to the choice of slab-density modelling method, even  
29 though the dynamic topography and slab-induced gravity anomaly vary considerably  
30 when the slab density is modelled by different methods. The continued presence of a  
31 high- $T_e$  band along the trench after application of dynamic corrections shows that,  
32 before using 2D inversion methods to estimate  $T_e$  variations in subduction zones, there  
33 are other factors that should be considered besides the slab dynamic effects on the  
34 overriding plate.

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36 **Key words:** effective elastic thickness; lithospheric strength; Izu-Bonin-Mariana  
37 subduction system; slab density; dynamic topography; slab gravity anomaly

# 38 1 INTRODUCTION

39 The physical properties of the lithosphere govern intra-plate convergence and  
40 inner-plate deformation and therefore exert a strong control on trench shape and  
41 back-arc spreading ([Forsyth and Uyeda, 1975](#); [Niu, 2014](#)). One of the most important  
42 of these physical properties is the flexural rigidity, which is a measure of the  
43 lithosphere's resistance to bending under applied loads, whether from above, below or  
44 within ([Burov and Diament, 1995](#); [Watts et al., 1980](#)). The effective elastic thickness  
45 ( $T_e$ ), a quantity which can be estimated from the coherence between topography and  
46 Bouguer gravity anomaly, is related to lithospheric flexural rigidity ([McKenzie and](#)  
47 [Fairhead, 1997](#)) and its rheological structure ([Oruç and Sönmez, 2017](#)). Available  
48 gravity and topography/bathymetry data can be used to compute lateral variations of  
49  $T_e$ , and therefore study the geodynamic characteristics of a region, including  
50 subduction zones. For example, [Pérez-Gussinyé et al. \(2008\)](#) estimated the  
51 geodynamic relationship between subduction geometry and lithospheric strength along  
52 the Andean margin by estimating effective elastic thickness based on the Bouguer  
53 coherence technique, while [Ratheesh Kumar et al. \(2013\)](#) evaluated the convergence  
54 characteristics of subduction along the Indonesian active continental margin, also by  
55 estimating  $T_e$  from the Bouguer coherence.

56 Oceanic plate strength is strongly related to its temperature structure ([Kalnins and](#)  
57 [Watts, 2009](#)) because thermal softening induced by magmatic activity reduces  
58 lithospheric strength ([Ebinger et al., 1989](#)). The existence of present-day back-arc  
59 spreading, and active magmatic arcs and forearcs across the Izu-Bonin-Mariana (IBM,

60 Figure 1) arc region ([Stern et al., 2004](#)) demonstrates that there exists considerable  
61 magmatic activity. Hence, the lithosphere of the IBM regions surrounded by arcs  
62 should be weak. The downward bending of the incoming plate by subduction  
63 introduces significant fracturing and so reduces the strength of the incoming plate near  
64 trench ([Bry and White, 2007](#); [McNutt and Menard, 1982](#)). Therefore, the lithosphere at  
65 the trench and neighboring regions should have a reduced  $T_e$  value. Indeed, studies  
66 using forward and inverse modelling of one-dimensional (1D) gravity/bathymetry  
67 profiles along the IBM subduction zone found reduced  $T_e$  values ([Billen and Gurnis,](#)  
68 [2005](#); [Bry and White, 2007](#); [Contreras-Reyes and Osses, 2010](#); [Hunter and Watts,](#)  
69 [2016](#); [Zhang et al., 2014](#)). However, 2D  $T_e$ -estimation methods have shown a band  
70 along the IBM Trench with obviously higher  $T_e$  values than the surroundings ([Chen et](#)  
71 [al., 2013](#); [Kalnins and Watts, 2009](#)).

72 Along the Andean margin, the effects of slab gravity and dynamic topography on  
73 estimated  $T_e$  values are negligible ([Pérez-Gussinyé et al., 2008](#)). The subducting Nazca  
74 plate (younger than 50 m.y.) has a relatively small slab density contrast with the  
75 mantle and a small sinking force, resulting in a small slab gravity anomaly and  
76 dynamic topography. However, the IBM trench sees the subduction of the oldest  
77 oceanic plate (older than 130 m.y.) on the Earth and is under strong back-arc  
78 spreading ([Stern et al., 2004](#)). Therefore, the effects of dynamic topography and slab  
79 gravity along the IBM Trench are, in contrast, non-negligible for  $T_e$  studies. Here we  
80 estimate these effects on  $T_e$  estimations based on different slab density modelling  
81 methods.

82

## 83 2 METHODOLOGY

84 The effective elastic thickness ( $T_e$ ) is related to the lithosphere's flexural rigidity  
85 ( $D$ ) by:

$$86 \quad D = \frac{ET_e^3}{12(1-\nu^2)}, \quad (1)$$

87 where  $E$  is Young's modulus,  $\nu$  is Poisson's ratio, and both  $E$  and  $\nu$  can be treated as  
88 constants ([Walcott, 1970](#)). Under long-wavelength loading, the lithosphere is  
89 isostatically compensated by a mechanism such as Airy isostasy, and the coherence  
90 between topography and Bouguer anomaly tends to a value of 1 ([Chen et al., 2015](#);  
91 [Pérez-Gussinyé et al., 2004](#)). Short-wavelength loads that can be supported by  
92 lithospheric strength result in almost no lithospheric flexure, in which case the  
93 coherence between loading and Bouguer anomaly tends to zero ([Forsyth, 1985](#);  
94 [Pérez-Gussinyé et al., 2004](#)). [Simons and Olhede \(2013\)](#) give an expression for the  
95 coherence transition wavelength, being the wavelength separating compensated from  
96 supported loads, and at which the coherence has a value of 0.5. A related expression is  
97 provided by the flexural wavelength ( $\lambda$ ), being the wavelength of the depression due  
98 to a point surface load, and related to  $T_e$  ([Macario et al., 1995](#)):

$$99 \quad \lambda = 2\pi \left( \frac{ET_e^3}{12(1-\nu^2)} / \Delta\rho g \right)^{1/4}, \quad (2)$$

100 where  $\Delta\rho$  is the density contrast across the compensation surface (usually taken to  
101 be the Moho), and  $g$  is the acceleration due to gravity. Although the coherence  
102 transition and flexural wavelengths are of the same order of magnitude, they do differ

103 for very high or very low values of the initial subsurface-to-surface loading ratio  
104 ([Kirby and Swain, 2008](#)).

105

## 106 **2.1 $T_e$ estimation using fan wavelet**

107 There are two main widely-used methods for mapping spatial variations in  $T_e$   
108 based on the coherence between the Bouguer gravity anomaly and  
109 topography/bathymetry: the multitaper ([McKenzie and Fairhead, 1997](#); [Simons et al.,  
110 2000](#)) and the fan wavelet coherence method ([Kirby, 2005](#); [Kirby and Swain, 2004](#)).

111 The multitaper method uses a moving window of fixed dimensions to obtain spatially  
112 varying coherence values across an area and from this constructs a  $T_e$  map. The need  
113 to select a window size limits the multitaper method in two ways: firstly, flexure  
114 cannot be resolved when the window size is smaller than the flexural wavelength;  
115 secondly, a large window size cannot image detailed spatial variations in  $T_e$  ([Chen et  
116 al., 2013](#); [Kalnins and Watts, 2009](#); [Pérez-Gussinyé et al., 2008](#); [Pérez-Gussinyé et al.,  
117 2004](#)). The fan wavelet method avoids these shortcomings, so we choose this method  
118 to estimate  $T_e$  distribution in our research area. Nevertheless the fan wavelet method  
119 faces the problem that the resolution of the  $T_e$  map depends on the central  
120 wavenumber (CW) of the Morlet wavelet used in the wavelet analysis ([Chen et al.,  
121 2015](#); [Kirby, 2014](#)).

122 [Kirby and Swain \(2004\)](#) constructed the fan wavelet by superposition of a series  
123 of rotated Morlet wavelets, giving isotropic and complex wavelet coefficients. The  
124 Fourier transform of the two dimensional (2D) Morlet wavelet is

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$$\hat{\psi}_{s\theta}(\mathbf{k}) = s e^{-[(su - k_0 |\cos\theta|)^2 + (sv - k_0 |\sin\theta|)^2]^{1/2}}, \quad (3)$$

126

where  $|\mathbf{k}_0|$  is the central wavenumber (CW),  $u$  and  $v$  are the components of the 2D

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wavenumber vector  $\mathbf{k}$ ,  $s$  is the scale of the wavelet, and  $\theta$  is its azimuth. Each Morlet

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wavelet scale is related to an equivalent Fourier wavenumber,  $k_F$ , by  $k_F = |\mathbf{k}_0|/s$

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([Kirby, 2005](#)).

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The wavelet transform is a space-domain convolution between the signal to be

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analyzed and the wavelet. This convolution is performed at all the selected scales and

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azimuths, and as such, reveals the harmonic content of the signal at each space-domain

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grid node and azimuth. The wavelet transform is most efficiently performed in the

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wavenumber domain where convolution becomes multiplication, yielding wavelet

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coefficients  $B_{s\mathbf{x}\theta}$  via the equation:

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$$B_{s\mathbf{x}\theta} \equiv B(s, \mathbf{x}, \theta) = \mathbf{F}^{-1} [\hat{b}(\mathbf{k}) \hat{\psi}_{s\theta}(\mathbf{k})], \quad (4)$$

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where  $\mathbf{F}^{-1}$  is the inverse 2D Fourier transform operator, and  $\hat{b}(\mathbf{k})$  is the Fourier

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transform of the space domain signal  $b(\mathbf{x})$ .

139

The 2D observed wavelet coherence ( $\gamma_{obs}^2(s, \mathbf{x})$ ) between observed Bouguer gravity

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anomaly and topography at scale ( $s$ ) and location ( $\mathbf{x}$ ) is calculated by summing the

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wavelet cross- and co-spectra over azimuth ([Kirby and Swain, 2008](#); [Kirby and Swain,](#)

142

[2011](#)):

143

$$\gamma_{obs}^2(s, \mathbf{x}) = \frac{\left| \langle B_{s\mathbf{x}\theta} H_{s\mathbf{x}\theta}^* \rangle_{\theta} \right|^2}{\langle B_{s\mathbf{x}\theta} B_{s\mathbf{x}\theta}^* \rangle_{\theta} \langle H_{s\mathbf{x}\theta} H_{s\mathbf{x}\theta}^* \rangle_{\theta}}, \quad (5)$$

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where  $B_{s\mathbf{x}\theta}$  and  $H_{s\mathbf{x}\theta}$  are the complex wavelet coefficients of Bouguer anomaly and

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topography, respectively, the \* indicates the complex conjugate, and  $\langle \rangle_{\theta}$  represents



146 averaging over Morlet wavelet azimuth.

147 The observed coherence is then inverted against the predictions of a thin, elastic  
 148 plate model of plate flexure using the wavelet adaptation of the method of [Forsyth](#)  
 149 [\(1985\)](#) by [Swain and Kirby \(2006\)](#). The inversion for  $T_e$  is performed by comparing  
 150 the observed coherence with the predicted coherence at each space-domain grid node,  
 151 where the predicted coherence can be computed by ([Chen et al., 2015](#); [Kirby and](#)  
 152 [Swain, 2011](#))

$$153 \quad \gamma_{pre}^2(s, \mathbf{x}) = \frac{(\mu_T \kappa_T + \mu_B \kappa_B f^2 r^2)^2}{(\mu_T^2 + \mu_B^2 f^2 r^2)(\kappa_T^2 + \kappa_B^2 f^2 r^2)}. \quad (6)$$

154 Here,  $f$  is the subsurface to surface loading ratio which can be determined by:

$$155 \quad f^2(s, \mathbf{x}) = \frac{\langle |W_i|^2 \rangle_\theta}{r^2 \langle |H_i|^2 \rangle_\theta}, \quad (7)$$

156 where  $W_i$  and  $H_i$  are the wavelet coefficients of initial surface and subsurface loads  
 157 which are assumed to be statistically independent (have random phase) ([Forsyth, 1985](#))  
 158 and are products of the inversion ([Forsyth, 1985](#); [Swain and Kirby, 2006](#)). Other  
 159 variables used in equation (6) are calculated from:

$$160 \quad \begin{cases} \kappa_B = \Delta\rho_2 / \varphi \\ \kappa_T = 1 - \Delta\rho_1 / \varphi \\ \mu_B = 2\pi G \Delta\rho_2 (1 - \Delta\rho_2 / \varphi) e^{-kZ_m} \\ \mu_T = 2\pi G \Delta\rho_2 (-\Delta\rho_1 / \varphi) e^{-kZ_m} \\ \varphi = Dk^4 / g + (\rho_m - \rho_w) \end{cases}, \quad (8)$$

161 where  $\Delta\rho_1 = \rho_b - \rho_*$ ,  $\Delta\rho_2 = \rho_m - \rho_b$ ;  $r = \Delta\rho_1 / \Delta\rho_2$ ;  $\rho_*$  is seawater or air density  
 162 depending on the setting (continent or ocean);  $\rho_m$ ,  $\rho_b$  and  $\rho_w$  are densities of mantle,  
 163 basement and seawater respectively;  $Z_m$  is the depth to the compensating interface (the  
 164 Moho);  $g$  is the gravitational acceleration,  $G$  is the Newtonian gravitational

165 constant; and  $k \equiv |\mathbf{k}|$  is the 1D wavenumber which can be replaced by the equivalent  
166 Fourier wavenumber in the fan wavelet case ([Kirby, 2005](#)).

167 The  $T_e$  value at each grid node is obtained by minimizing the misfit between  
168 predicted and observed coherence ([Chen et al., 2015](#); [Kirby and Swain, 2011](#)). Since  
169 the final  $T_e$  resolution would be affected by central wavenumber values, we compute  
170  $T_e$  maps based on four different  $|\mathbf{k}_0|$  values (2.668, 3.081, 3.773 and 5.336) as done  
171 by [Chen et al. \(2015\)](#). See [Kirby and Swain \(2011\)](#) for an explanation of the  $|\mathbf{k}_0|$   
172 values. The values of the constants used for estimating  $T_e$  are shown in Table 1.

173 Error estimates on the coherence were obtained using the jackknife method of  
174 error estimation ([Thomson and Chave, 1991](#)) applied to wavelets ([Kirby and Swain,](#)  
175 [2013](#)). The coherence errors were then used as weights on the difference between  
176 observed and predicted coherence in minimization of the square of the  $L_2$ -norm  
177 (chi-squared statistic) ([Kirby, 2014](#)). Errors on  $T_e$  were obtained from the width of the  
178 misfit curve in the parameter space (where  $T_e$  is the only parameter) corresponding to  
179 a confidence level of 95% ([Press et al., 1992](#)).

180 Kirby and Swain (2008, 2011) have performed extensive tests on the accuracy of  
181 the method under different scenarios using synthetic, fractal loading models. Results  
182 have been encouraging. Of particular note here is the test on a plate with a  $T_e$   
183 discontinuity, representing a fault or subduction zone ([Kirby and Swain, 2008](#)).

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## 185 **2.2 Dynamic topography and slab gravity anomaly**

186 Loading of lithosphere is a key parameter for  $T_e$  evaluations and loading is not

187 limited to topographic variations. Figure 2 illustrates factors which include surface  
188 loading such as mountains and seawater, and internal loading such as high-density  
189 magmatic intrusions into the basement and dynamic forces from beneath the  
190 lithosphere ([Bry and White, 2007](#)), referred to as under-loading in this paper. Here we  
191 convert the bathymetry to a rock-equivalent topography:

$$192 \quad h_{eq}(\mathbf{x}) = \frac{\rho_c - \rho_w}{\rho_c} d(\mathbf{x}), \quad \forall d(\mathbf{x}) < 0 \quad (9)$$

193 where  $d(\mathbf{x})$  is the bathymetry or topography at location  $\mathbf{x}$  and  $h_{eq}(\mathbf{x})$  is the  
194 equivalent topography at the same position. A similar conversion also should be used  
195 to calculate the equivalent topography of sediment cover when it is thick ([Shi et al.,](#)  
196 [2017](#)):

$$197 \quad h_{eq}(\mathbf{x}) = \frac{\rho_c - \rho_s}{\rho_c} d(\mathbf{x}), \quad \forall d(\mathbf{x}) < 0. \quad (10)$$

198 Where  $\rho_s$  is sediment density and  $d(\mathbf{x})$  is sediment thickness. Since it is hard to  
199 accurately constrain the detailed basement density distribution, we will assume that  
200 basement density in the research area is uniform.

201 There also exists a downward pulling force on the overlying lithosphere caused  
202 by the sinking of the subducted slab and transmitted by the asthenospheric wedge  
203 ([Flament et al., 2013](#); [Gvirtzman et al., 2016](#); [Husson, 2006](#)). The topography changes  
204 induced by this slab dynamic force are called ‘dynamic topography’ in this paper; the  
205 variation of this topography is predominantly long wavelength. The density contrast  
206 between the subducting slab and the mantle also causes a long-wavelength gravity  
207 perturbation, with this gravity anomaly referred to as ‘dynamic gravity’ or ‘slab

208 gravity anomaly' in this paper.

209 To evaluate the long-wavelength downward pulling force on the overlying  
210 lithosphere by slab sinking, the sinking slab is separated into discrete elements. Based  
211 on the Stokes stream function, the pulling force ( $F$ ) on a lithospheric surface point can  
212 be calculated by ([Husson, 2006](#); [Morgan, 1965](#)):

$$213 \quad F = \frac{3\Delta\rho_{ij}v_{ij}gz_{ij}^3}{\pi r_{ij}^5}, \quad (11)$$

214 where  $i, j$  is the slab element index,  $\Delta\rho_{ij}$  is the density difference between this slab  
215 unit and the background mantle,  $v_{ij}$  is the volume of the unit,  $g$  is the gravitational  
216 acceleration,  $z_{ij}$  is the depth of the unit and  $r_{ij}$  is the distance between the slab unit  
217 and the surface point. The overlying lithosphere will be bent downward, so when the  
218 surface deflection is  $h$  then the buoyancy force change of the overlying lithosphere  
219 induced by this deflection is:

$$220 \quad \sigma = (\rho_m - \rho_*)gh, \quad (12)$$

221 where  $\rho_m$  is mantle density and  $\rho_*$  is seawater or air density. If it is a stress-free  
222 surface then the buoyancy force change ( $\sigma$ ) should be equal to the slab pulling force  
223 ( $F$ ). Hence the total deflection ( $H$ ) of the lithosphere surface point by the  
224 three-dimensional slab can be calculated by equating eqs (11) and (12) and summing  
225 over all elements:

$$226 \quad H = \sum_j \sum_i \frac{3\Delta\rho_{ij}v_{ij}z_{ij}^3}{\pi r_{ij}^5(\rho_m - \rho_*)}. \quad (13)$$

227 Except for the density perturbation of the slab unit ( $\Delta\rho_{ij}$ ), all the other parameters can  
228 be assumed (Table 1) or calculated according to element coordinates.

229 Here we compare three methods for slab-density modelling. First, [Deschamps et](#)  
230 [al. \(2001\)](#) noted that both the lithosphere density ( $\rho$ ) and seismic shear-wave velocity  
231 ( $V_s$ ) are dependent on similar parameters, such as temperature, composition, pressure,  
232 etc.. This led [Isaak et al. \(1989\)](#) to derive a crude quantitative relationship between  $\rho$   
233 and  $V_s$  when only temperature affects  $\rho$  and  $V_s$ :

$$234 \quad \zeta = \frac{\ln \rho - \ln \rho_0}{\ln V_s - \ln V_{s_0}} \quad (14)$$

235 where the ratio ( $\zeta$ ) is 0.1 when depth is shallower than 400 km, and 0.2 when depth is  
236 between 400 km and 700 km, according to the study by [Karato \(1993\)](#). The parameters  
237  $\rho_0$  and  $V_{s_0}$  are reference density and shear-wave velocity, respectively; we call this  
238 method “Isaak1989”. A second method for slab density modelling is based on  
239 laboratory data by [Karato and Karki \(2001\)](#) which assigns a constant velocity-density  
240 conversion factor of  $0.15 \text{ cm}^{-3} \text{ km}^{-1} \text{ s}$ ; we call this method “KK2001”. The third  
241 method we consider assigns a slab density anomaly of  $25 \text{ kg m}^{-3}$  relative to its  
242 surroundings, as [Pérez-Gussinyé et al. \(2008\)](#) assumed; we call this “PG2008”.

243 After assuming slab thickness to be uniformly 100 km as [Husson \(2006\)](#) did, and  
244 extracting the shear-wave velocity of the subducted slab, the slab density perturbation  
245 can be derived from  $V_s$  data and equation (14) by taking the Earth model of  
246 [Dziewonski and Anderson \(1981\)](#) as reference. The dynamic topography due to the  
247 slab density perturbation can then be calculated using equation (13). Additionally, the  
248 dynamic gravity due to the slab density perturbation can be calculated in the frequency  
249 domain ([Oldenburg, 1974](#); [Parker, 1973](#)). The topography and gravity after removal of

250 the slab effects are called here the ‘residual’ topography and gravity, respectively.

251

## 252 **3 CASE STUDY AT THE IBM SUBDUCTION ZONE**

### 253 **3.1 Geologic setting**

254 Subduction of the Pacific plate under the West Philippine Sea plate began by  
255 ~52-45 Ma ([Arculus et al., 2015](#); [Stern et al., 2004](#)) and this subduction resulted in the  
256 formation of the IBM Trench. The IBM subduction hinge roll back commenced with  
257 the opening of two back-arc basins, the Shikoku and Parece Vela Basins, in the Late  
258 Oligocene ([Hall, 2002](#); [Hilde et al., 1977](#); [Honza and Fujioka, 2004](#)). Part of the initial  
259 volcanic arc remains on the western margin of these basins as a linear bathymetric  
260 high named the Kyushu Palau Ridge (Figure 1). The seafloor spreading direction of  
261 the Shikoku Basin changed from ENE–WSW to NE–SW around 19 Ma; the Parece  
262 Vela Basin spreading direction changed around 20 Ma, from E–W to NE–SW, possibly  
263 related to rotation of the Philippine Sea plate ([Hall, 1996](#); [Sdrolias et al., 2004](#)). These  
264 mid-ocean ridge spreading direction changes induced the formation of S-shaped  
265 fracture zones in these two basins ([Sdrolias et al., 2004](#)). Back-arc spreading in the  
266 Shikoku Basin and Parece Vela Basins ceased at ~15 Ma.

267 A hiatus in back-arc spreading then occurred in this area until the Mariana  
268 Trough’s initiation at ~7 Ma, which is still spreading at the back of the Mariana  
269 Trench ([Jolivet et al., 1989](#)). Continuous subduction after cessation of the Shikoku and  
270 Parece Vela Basin’s back-arc spreading resulted in widening of the Izu-Bonin-Mariana  
271 Arc. The previous Mariana Arc was split into two parts by the newly formed Marina

272 Trough which is bounded by the West Mariana Arc and East Mariana Arc (Figure 1).

273 The region of the subducting Pacific plate in our research area which is older than  
274 ~130 Ma ([Müller et al., 2008](#)) has cooled and subsided to great depths. Within this  
275 region, the subduction systems form deep trenches and contain the lowest point (the  
276 Challenger Deep) on the Earth's surface, at more than 10 km depth. The incoming  
277 western Pacific plate is much older and denser than the over-riding Shikoku and  
278 Parece Vela Basins (30 -15 Ma) ([Sdrolias et al., 2004](#)) and provides a large enough  
279 negative buoyancy to sustain subduction ([Hall et al., 2003](#); [Niu, 2014](#)). The subduction  
280 process is, however, affected by the sea-bottom topography and density distribution of  
281 the incoming Pacific plate ([Zhang et al., 2016](#)). For example, the docking of Bonin  
282 (Ogasawara) Plateau (Figure 1b) which is about 2 km to 3 km higher than the  
283 surrounding seafloor has already modified the shape of the IBM Trench ([Kong et al.,](#)  
284 [2018](#); [Mason et al., 2010](#)).

285

### 286 **3.2 Data sets**

287 [Forsyth \(1985\)](#) coherence method requires the use of Bouguer, rather than free air  
288 gravity anomalies, since the coherence between the former and the topography is  
289 much less dependent on the loading ratio and model parameters such as layer depths  
290 and densities than is the coherence between the latter and the topography (e.g., Kirby,  
291 2014). The ETOPO1 model is used for the land topography and ocean bathymetry in  
292 this study ([Amante and Eakins, 2009](#)), with a 1 arc minute grid spacing. In marine  
293 areas, the ETOPO1 is the combination of bathymetry measured from ships, and

294 bathymetry inverted from satellite altimetry data. The Bouguer anomaly grid by  
295 Bureau Gravimétrique International ([Balmino et al., 2012](#)) is derived from the  
296 EGM2008 Geopotential model and the ETOPO1 Global Relief model, and this  
297 Bouguer grid is with 2 arc minutes resolution. The EGM2008 model includes surface  
298 gravity measurements, satellite altimetry and gravimetry measurements. Since our  
299 research area has thick water but thin sediment coverage ([Amante and Eakins, 2009](#);  
300 [Divins, 2004](#)), equivalent topography is computed only from the bathymetry using the  
301 seawater depth from ETOPO1. Global shear-wave velocity structure from surface to  
302 700 km depth by [Schaeffer and Lebedev \(2013\)](#) is used for computing the dynamic  
303 topography and slab gravity anomaly. This velocity volume has a horizontal resolution  
304 of  $0.5^\circ$  and a vertical resolution of 25 km.

305

### 306 **3.3 Elastic thickness results**

307 We first computed the effective elastic thickness from the fan-wavelet method  
308 using the equivalent topography and Bouguer anomaly without the slab corrections.  
309 Four  $T_e$  maps were computed with four different central wavenumbers (CW, or  $|\mathbf{k}_0|$ ).  
310 The  $T_e$  range for CW = 2.668 is 0.5 - 152.0 km, for CW = 3.081 it is 0.5 - 148.3 km,  
311 for CW = 3.773 it is 0.8 - 139.9 km, and for CW = 5.336 the range is 1.0 - 96.4 km.  
312 These initial results, before consideration of slab effects, are shown in Figure 3. Note  
313 the obvious high  $T_e$  band along, and to both sides of, the IBM Trench in each result.

314 The  $T_e$  map generated with the lowest central wavenumber (2.668) shows the  
315 most detail in the spatial variations of  $T_e$  (Figure 3a), while the 5.336 CW map (Figure



316 3d) is much longer-wavelength in nature due to such wavelets tending to smear  
317 information out over larger spatial scales ([Kirby and Swain, 2011](#)). It is this property  
318 of higher-CW wavelets, despite their superior wavenumber-domain resolution, that  
319 make them less suited for studies where  $T_e$  can change rapidly and by large amounts  
320 over short distances, as seems to be the case in this study area. To this end, we chose  
321 the smallest allowable value of  $|\mathbf{k}_0|$ , 2.668, in the following sections (which must  
322 be  $>2.5$ ; [Kirby and Swain, 2011](#)).

323 Figure 4 shows the dynamic topography and gravity when the subducted slab's  
324 density structures are modelled with the PG2008, Isaak1989 and KK2001 methods,  
325 respectively. The dynamic topography and gravity results have a similar distribution  
326 but with reverse polarity because the positive density anomaly of the subducting slab  
327 induces a negative dynamic topography but positive gravity anomaly. Relatively large  
328 absolute dynamic topography and slab gravity anomaly values occur behind the trench,  
329 controlled by the density and position of the subducting slab (eq. (13)).

330 From the comparison between these three different groups of long-wavelength  
331 dynamic modelling results, we see that dynamic topography maps have a similar  
332 distribution trend but with different amplitudes. The smallest dynamic topography  
333 amplitude, with a value of 0.79 km, is based on the Isaak1989 method; the largest  
334 dynamic topography amplitude, with a value of 2.32 km, is based on the PG2008  
335 method. The latter is almost three times the amplitude of the former. The dynamic  
336 gravity maps also have similar comparison characteristics. Therefore, different slab  
337 density modelling methods produce very different dynamic calculation results.

338 Slab effects on gravity and topography were removed from the Bouguer gravity  
339 anomaly and topography, giving residual quantities. The effective elastic thickness  
340 was then recomputed using the residual topography and gravity anomaly when  $CW =$   
341 2.668 (Figure 5). A comparison between  $T_e$  maps based on different slab density  
342 modelling methods (Figure 5) shows that these three  $T_e$  maps have very similar trends  
343 and similar amplitudes even though the dynamic topography and gravity vary  
344 considerably when the slab density modelling method changes. These three  $T_e$  maps  
345 also have very similar error distributions (Figure 6). Therefore, while dynamic  
346 topography and gravity values are relatively sensitive to slab density modelling,  $T_e$   
347 estimations derived from them are not. The  $T_e$  values after application of the  
348 corrections are very much reduced, especially in the region surrounding the trench;  
349 this will be discussed in Section 3.5.

350 Figure 6 shows four  $T_e$  error maps before and after application of dynamic  
351 corrections to the gravity and topography data. The errors are uniformly low over the  
352 majority of the study area, while the largest errors are distributed along the high- $T_e$   
353 band surrounding the trench and on its boundary. The  $T_e$  estimations without dynamic  
354 corrections have the largest errors of the four; the mean value of this  $T_e$  error map  
355 (Figure 6a) is 2.7 km with a maximum error of 49.5 km. The mean  $T_e$  error after  
356 dynamic corrections when slab density is modelled based on the PG2008, Isaak 1989  
357 and KK2001 methods, is reduced to 0.94 km, 0.93 km and 0.96 km, respectively, and  
358 the maximum error of each  $T_e$  map is 23.8 km, 21.0 km and 23.5 km, respectively.  
359 Therefore, both the mean and maximum values of  $T_e$  error have been considerably

360 reduced by application of the dynamic corrections.

361

### 362 **3.4 Comparison with previous studies**

363 Two other studies have also mapped spatial variations of effective elastic  
364 thickness in this region ([Chen et al., 2013](#); [Kalnins and Watts, 2009](#)). The  $T_e$  map by  
365 [Chen et al. \(2013\)](#) also used the Bouguer coherence method, though they used the  
366 multitaper method of spectral estimation, and used ETOPO2 for the  
367 topography/bathymetry data and free-air gravity anomalies from the EIGEN-6C model  
368 (<http://icgem.gfz-potsdam.de/ICGEM/modelstab.html>) to derive the Bouguer anomaly  
369 is; we call this study Chen2013. The  $T_e$  map by [Kalnins and Watts \(2009\)](#) used the  
370 free-air admittance method with a satellite altimetry-derived free-air gravity anomaly  
371 grid (V16.1) ([Smith and Sandwell, 1994](#)) and the GEBCO bathymetry grid ([IOC,](#)  
372 [2003](#)); we refer to this map as KW2009. The western section of our research area  
373 coincides with the study area of Chen2013, while the eastern section coincides with  
374 that of KW2009.

375 There is no unusual  $T_e$  high along the Nankai Trough either in the Chen2013 map  
376 or in our  $T_e$  map (Figure 7b). Our average  $T_e$  value of Parece Vela Basin is 5.6 km and  
377 for the Shikoku Basin is 3.1 km (Figure 7b). Chen2013 also shows that the Parece  
378 Vela Basin has a stronger lithosphere than the Shikoku Basin. The  $T_e$  value of  
379 KW2009 for the incoming Pacific plate in the overlapping area between our model  
380 and KW2009 has a range of 0 - 15 km; our model has a similar range.

381 Both KW2009 and Chen2013 both have a very similar high- $T_e$  band along the

382 IBM Trench, and neither considered the effect of dynamic topography and gravity by  
383 the subducting slab. The spatial distribution of our high- $T_e$  band (Figure 7b) matches  
384 their geometries even though our methodology and input data are different from  
385 KW2009 and Chen2013.  $T_e$  maps from all studies have a sharp  $T_e$  transition on the  
386 boundaries of the high- $T_e$  band. The  $T_e$  value along this band in Chen2013 is larger  
387 than 60 km which is similar to our first-step result (Figure 7a). The  $T_e$  values along  
388 this band in KW2009 lie between 25 - 50 km, a relatively small range compared to  
389 Chen2013, but still much larger than the surrounding  $T_e$  values. Our peak value in the  
390 region overlapping with KW2009 is 37.2 km. These observations demonstrate that,  
391 irrespective of which spectral analysis method is used (multitaper (Chen2013 and  
392 KW2009) or wavelet (our model)), high- $T_e$  values along the IBM Trench are the  
393 outcome. But as noted above, removal of dynamic topography and gravity from the  
394 bathymetry/topography and gravity reduces the amplitude of this high- $T_e$  band.

395

### 396 **3.5 Influence of dynamic correction on $T_e$ evaluation**

397 Both faulting and magmatic activity reduce lithospheric strength ([Billen and](#)  
398 [Gurnis, 2005](#); [Bry and White, 2007](#); [Chen et al., 2013](#); [McNutt and Menard, 1982](#);  
399 [Zhang et al., 2014](#)). Therefore, one would expect  $T_e$  at the trench and its surrounding  
400 areas to have relatively small values. This is not seen in the  $T_e$  maps computed before  
401 corrections of the gravity and topography data (Figure 3).

402 The mean value of the high- $T_e$  band bounded by black solid lines in Figure 7a is  
403 76.7 km; the removal of slab effects from the data reduces the mean value within this

404 high- $T_e$  band by 56.9 km to 19.8 km (Figure 7b). Another difference between these  
405 two  $T_e$  maps (Figure 7a and Figure 7b) is that the boundaries between the high- $T_e$  band  
406 and the adjacent regions are sharper in Figure 7a than in Figure 7b. Similar  
407 characteristics appear on the  $T_e$  error map (Figure 6). The high- $T_e$  band of Figure 7a is  
408 much wider than the high- $T_e$  band of Figure 7b. Reduction of the amplitude and width  
409 of the high- $T_e$  band has brought the  $T_e$  value closer to the expectations of a curved,  
410 heated and faulted plate. This means that the dynamic correction could improve  $T_e$   
411 estimation results along oceanic trenches.

### 412

### 413 **3.6 Other possible influencing factors on $T_e$ evaluation**

414 The presence of a high- $T_e$  band after dynamic corrections were made (Figure 5)  
415 illustrates that there are still some other factors affecting  $T_e$  estimates. Besides slab  
416 sinking, the surrounding mantle flow and density variations could also modify  
417 topography and gravity as shown in Figure 2, and thus could affect  $T_e$  estimates. But  
418 because the density and velocity of the upper mantle are affected not only by  
419 temperature but also by composition variations, it is difficult to image density  
420 distributions from seismic tomography. Therefore, density contrasts in the upper-most  
421 mantle have been ignored for dynamic studies ([Conrad and Husson, 2009](#); [Heine et al.,  
422 2008](#); [Steinberger, 2007](#)). The bottom depth of this section varies between different  
423 studies, for example, 220 km by [Steinberger \(2007\)](#) and [Heine et al. \(2008\)](#), and 300  
424 km by [Conrad and Husson \(2009\)](#). Figure 8 shows that the mantle wedge has a clear  
425 low seismic velocity, but the mantle under the slab has a relatively uniform structure

426 and small changes in velocity. These uncertainties provide the reason why we only  
427 consider slab dynamic effects.

428 Oceanic trenches and back-arc basins over a subducting slab are dynamically  
429 compensated by viscous mantle flow due to the force on the overriding plate induced  
430 by slab sinking and mantle wedge flow ([Zhong and Gurnis, 1994](#)). Stresses induced by  
431 slab sinking are transmitted along the slab to the surface, resulting in trench deepening,  
432 so the trench topography is compensated by the subducted slab ([Davies, 1981](#)). In this  
433 case, the compensation surface is not the Moho as supposed by the 2D  $T_e$  inversion  
434 method. Thus,  $T_e$  estimates along the trench based on 2D inversion would be biased.  
435 Therefore, before estimating  $T_e$  variations in subduction zones, other factors should be  
436 considered besides the slab dynamic effects on the overriding plate.

437

## 438 **4 CONCLUSIONS**

439 In this paper, we have attempted to study the effect of long-wavelength dynamic  
440 topography and gravity upon lithospheric effective elastic thickness estimation by  
441 taking the Izu-Bonin-Mariana Trench as a case study region. Dynamic topography and  
442 slab gravity anomaly induced by the subducted slab was estimated and then removed  
443 from the topography and gravity anomaly. The resulting residual topography and  
444 gravity anomaly were used to generate a  $T_e$  map using the fan wavelet coherence  
445 method.  $T_e$  maps generated before the removal of dynamic topography and slab  
446 gravity anomaly, and also the results of other studies ([Chen et al., 2013](#); [Kalnins and](#)  
447 [Watts, 2009](#)), reveal a high- $T_e$  band along and surrounding the IBM Trench, in

448 disagreement with expectations of reduced  $T_e$  due to heating, faulting and plate  
449 curvature. However when the long-wavelength dynamic effects are applied, the  
450 amplitude and width of this high- $T_e$  band is strongly reduced, in keeping with  
451 expectations. Our tests show that effective elastic thickness estimation results are  
452 relatively insensitive to the particular method of slab-density estimation, despite the  
453 model topography and gravity being more sensitive to model choice. The continued  
454 presence of a high- $T_e$  band along the trench after application of dynamic corrections  
455 shows that, before using 2D inversion methods to estimate  $T_e$  variations in subduction  
456 zones, there are other factors that should be considered besides the slab dynamic  
457 effects on the overriding plate.

458

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466

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639

Table 1 Physical constants for  $T_e$  estimation after [Ratheesh Kumar et al. \(2013\)](#)

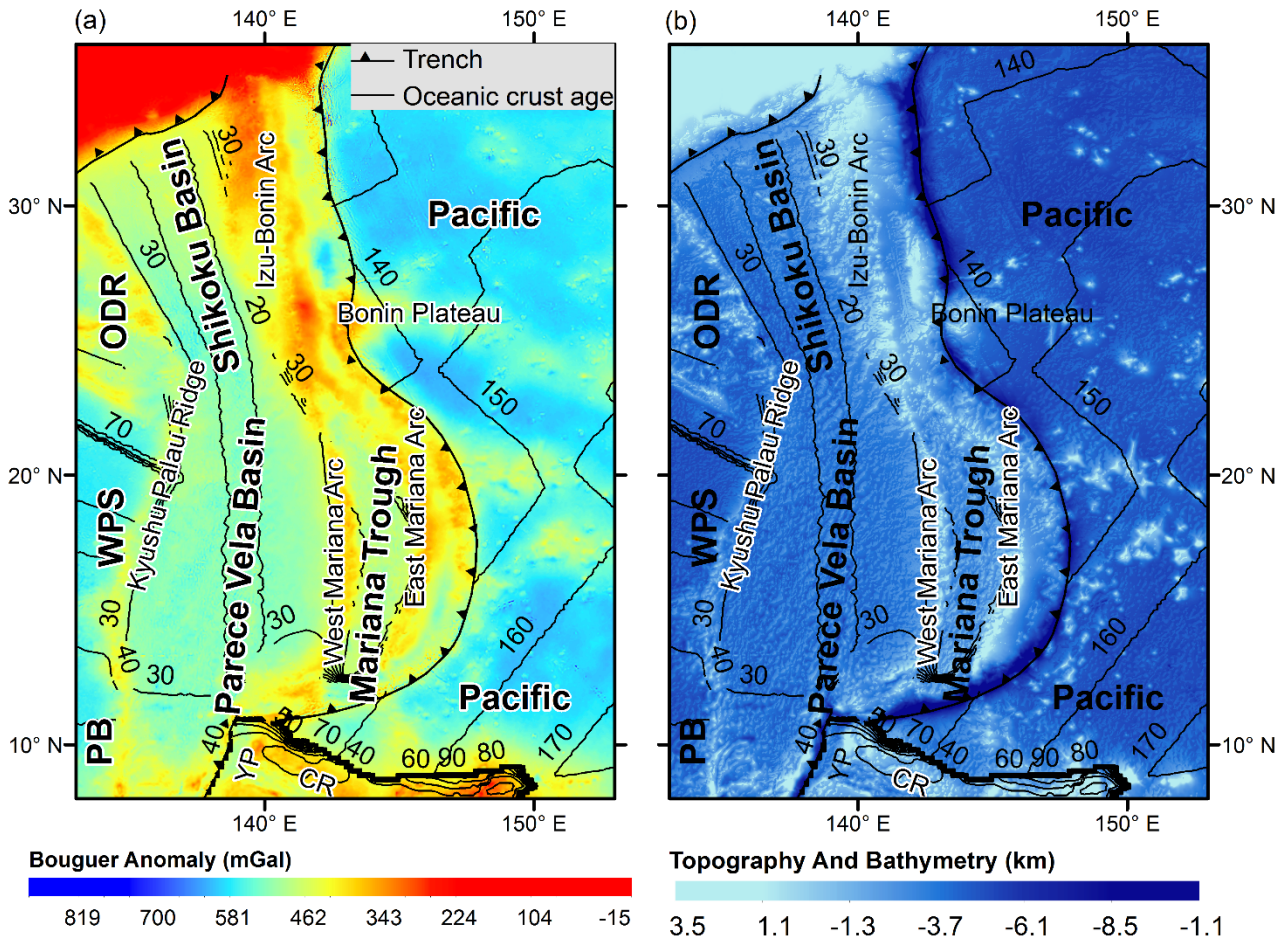
Constant	symbol	value	Unit
Young modulus	$E$	$10^{11}$	Pa
Newtonian gravitational constant	$G$	$6.67 \times 10^{-11}$	$m^3 kg^{-1} s^{-2}$
Poisson ratio	$\nu$	0.25	
gravitational acceleration	$g$	9.81	$m / s^2$
seawater density	$\rho_w$	1030	$kg / m^3$
basement density	$\rho_b$	2800	$kg / m^3$
mantle density	$\rho_m$	3300	$kg / m^3$

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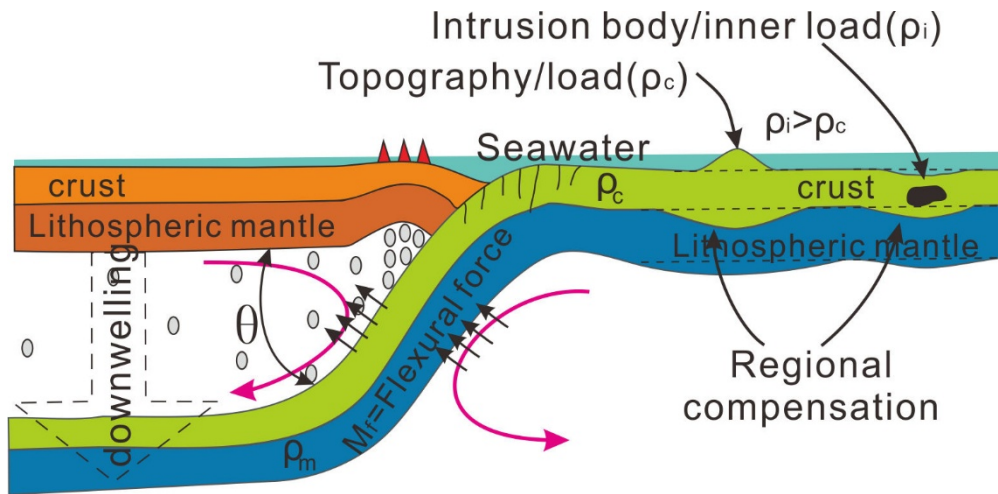
642

643 **FIGURES**



644  
 645 Figure 1. (a) The Bouguer gravity anomaly from the Bureau Gravimétrique  
 646 International ([Balmino et al., 2012](#)), and (b) the topography and bathymetry from  
 647 ETOPO1 ([Amante and Eakins, 2009](#)), of the IBM subduction system and surroundings.  
 648 The oceanic crustal age contour (labelled in m.y.) is based on the age grid provided by  
 649 [Müller et al. \(2008\)](#). Abbreviations: WPS, West Philippine Sea; ODR, Oki-Daito  
 650 Ridge; PB, Palau Basin; YT, Yap Trench; CR, Caroline Ridge.

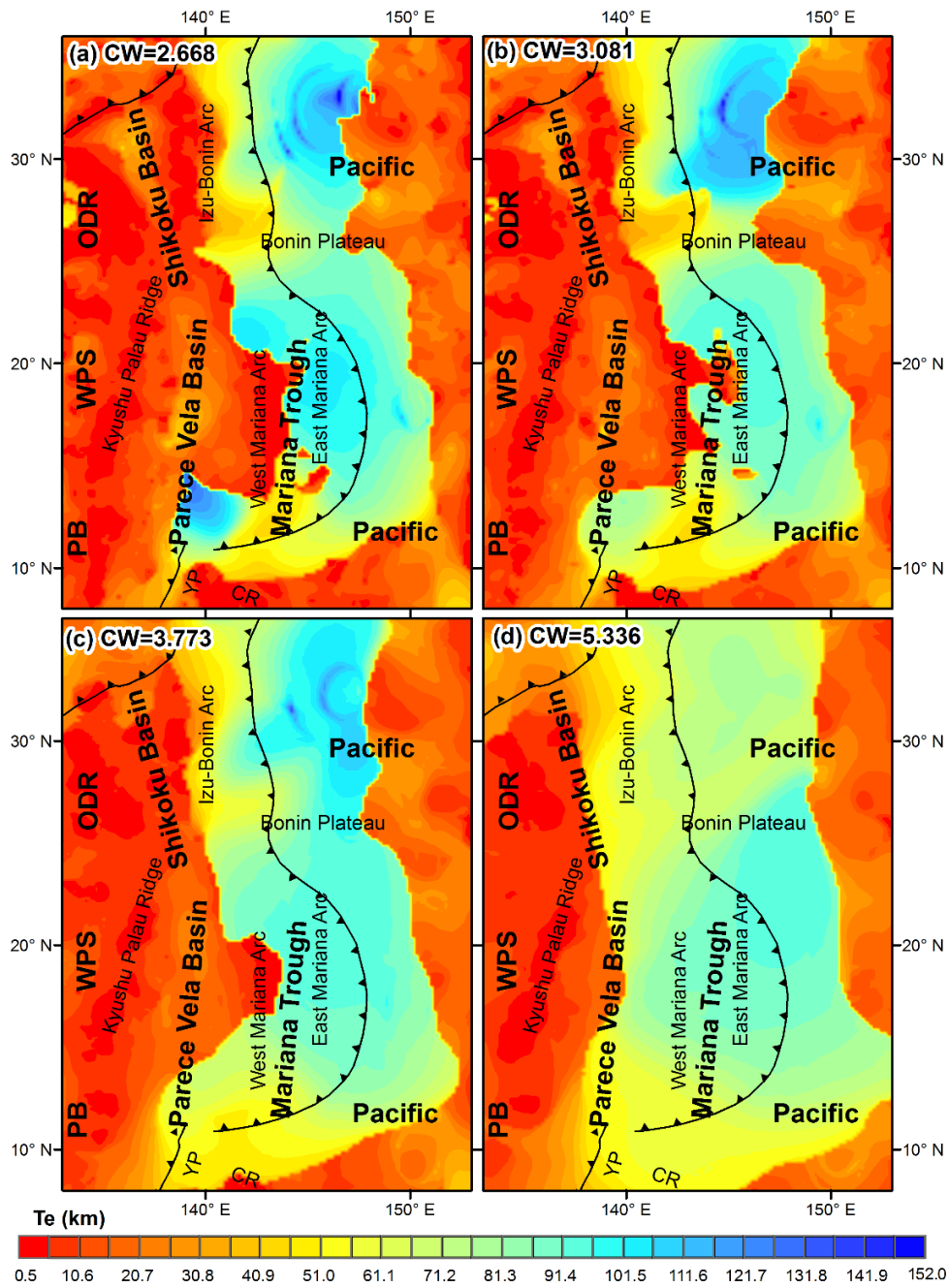
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652

653 Figure 2. Cartoon showing compensation mode and subduction dynamics. The dashed  
 654 lines represent the initial topography, Moho and lithospheric bottom. Subsequent  
 655 magmatic intrusion and mountain formation results in lithospheric bending and  
 656 compensation surface adjustment. Viscous asthenosphere transmits the dynamic forces  
 657 due to slab sinking to the overlying plate and induces topography changes. Water  
 658 released from dehydration of the subducted slab results in lithospheric partial melting  
 659 and reduction of lithospheric strength ([Chen et al., 2013](#)). Pink arrows represent  
 660 mantle flow at different locations ([Pérez-Gussinyé et al., 2008](#)).  $\theta$  is the subduction  
 661 angle.

662

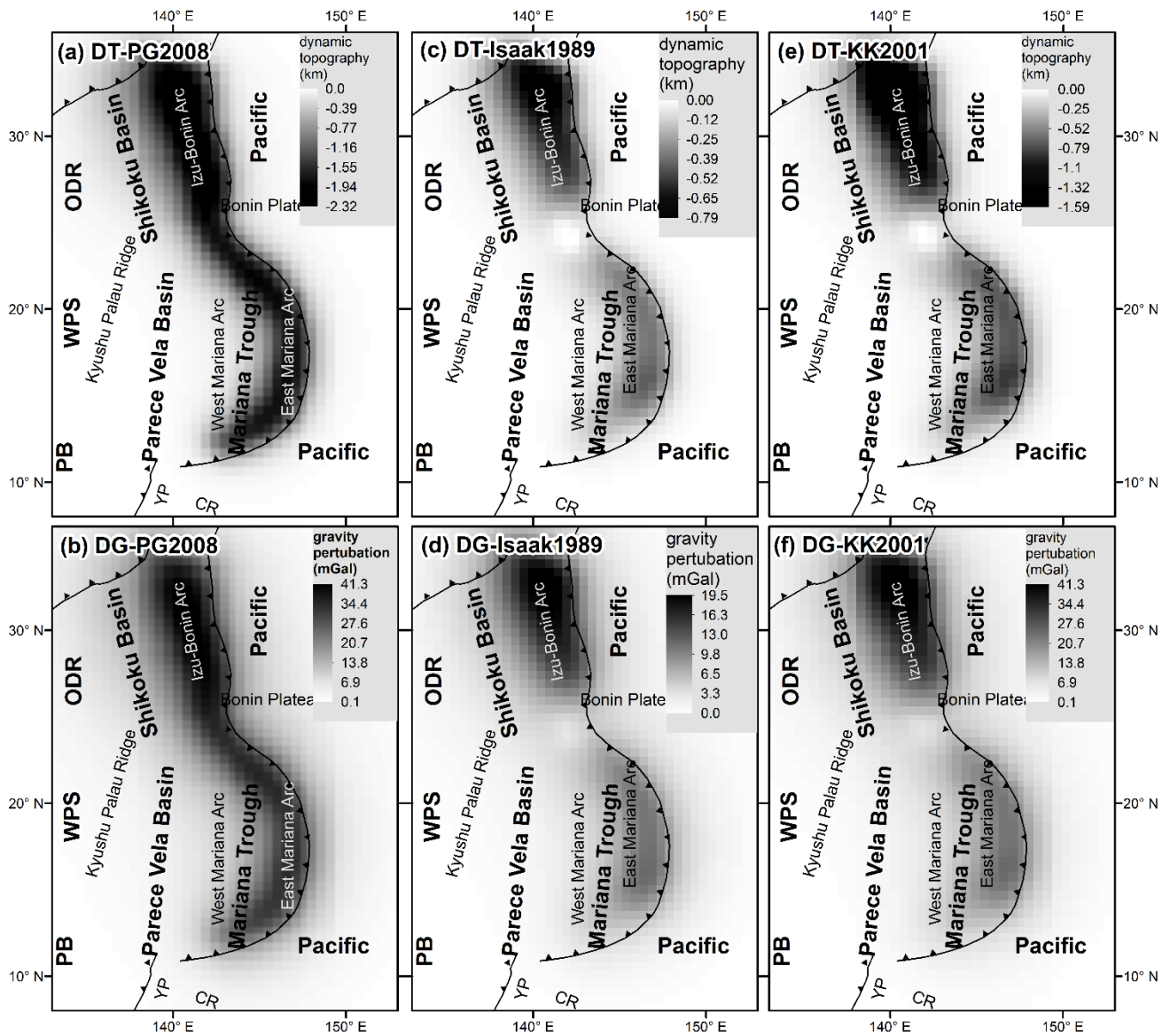


663

664 Figure 3.  $T_e$  estimation results (without slab corrections) based on four different  
 665 central wavenumber (CW) values. Abbreviations can be found in the caption to Figure

666 1.

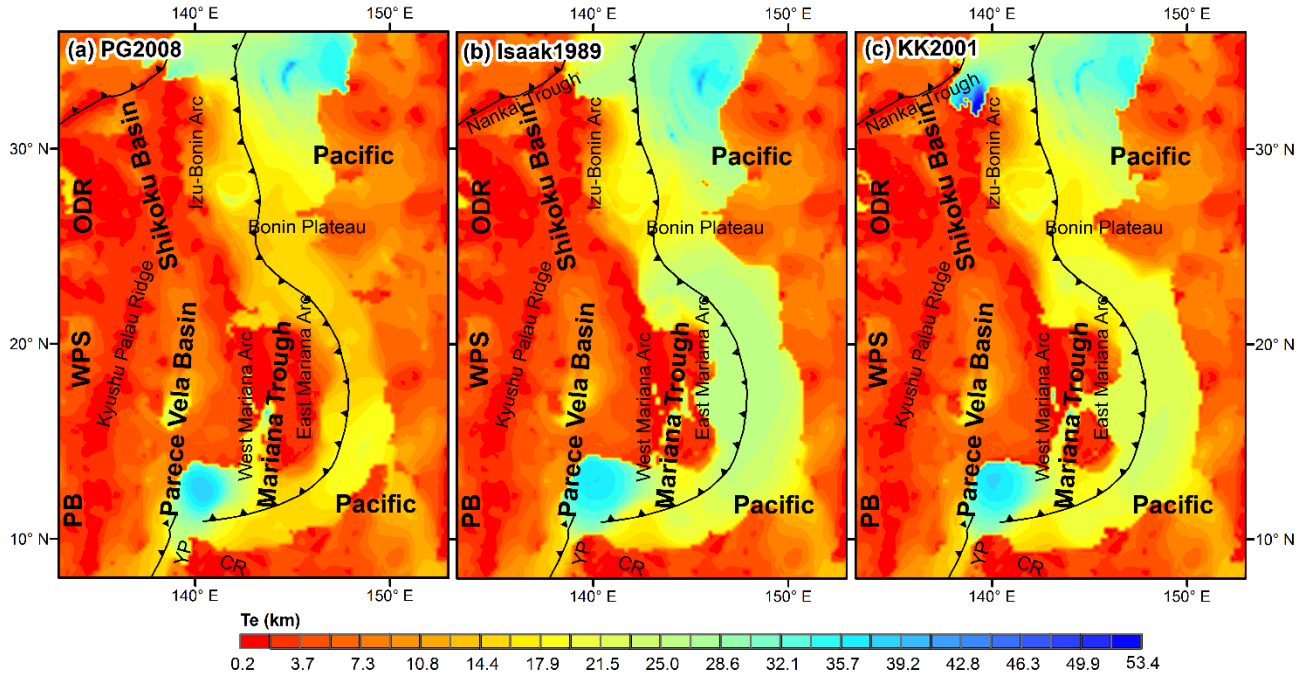
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668

669 Figure 4. Dynamic topography (DT) on the top panels and dynamic gravity (DG) on  
 670 the bottom panels induced by the sinking of the subducted Pacific plate when the slab  
 671 density is based on the PG2008, Isaak1989 and KK2001 methods, respectively. See  
 672 Figure 1 for abbreviations.

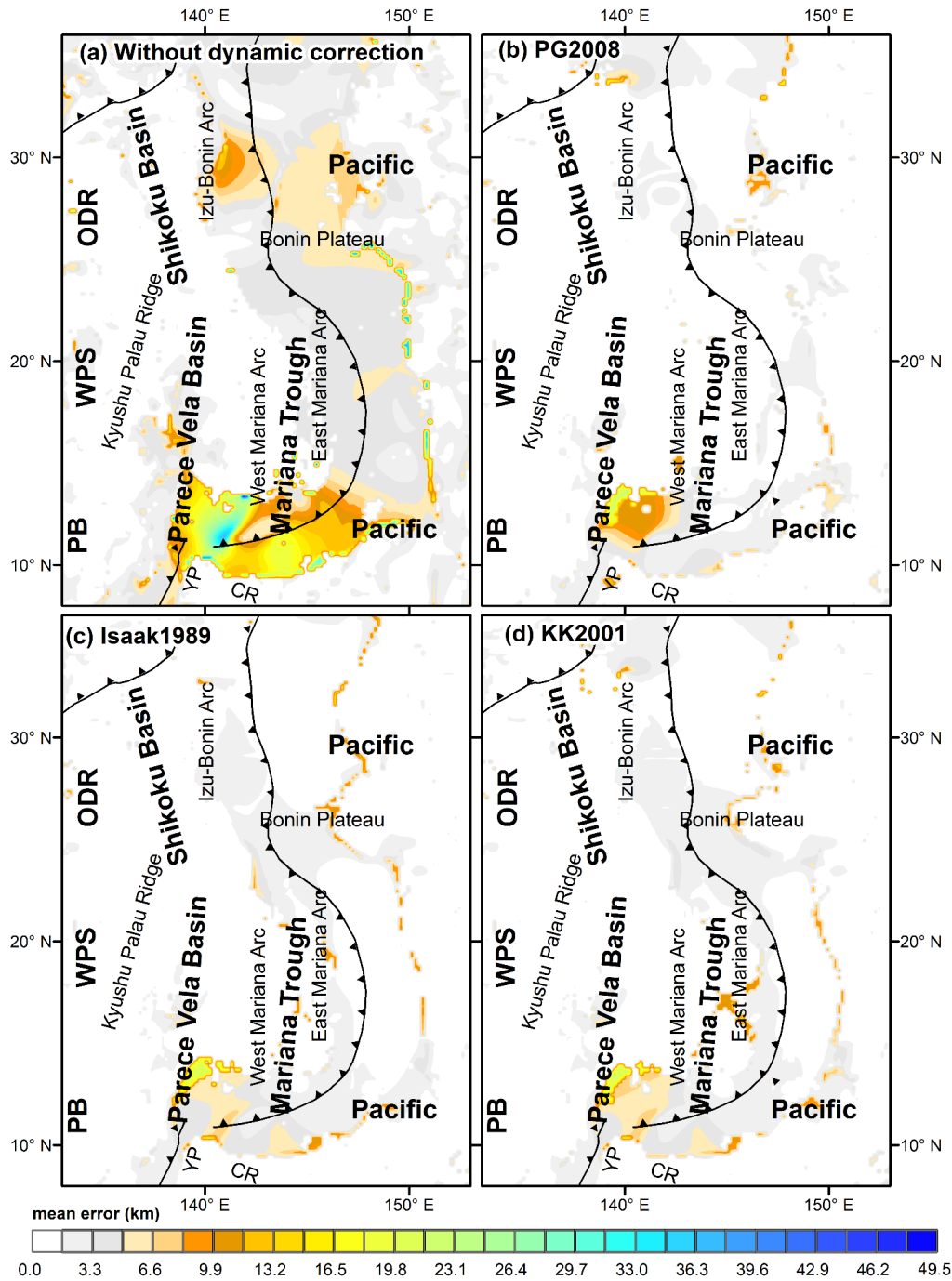
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675 Figure 5.  $T_e$  maps using residual gravity and topography, when the subducted slab  
 676 density is modelled using the PG2008, Isaak1989 and KK2001 methods, respectively  
 677 (central wavenumber is 2.668 in all). Abbreviations are the same as Figure 1.

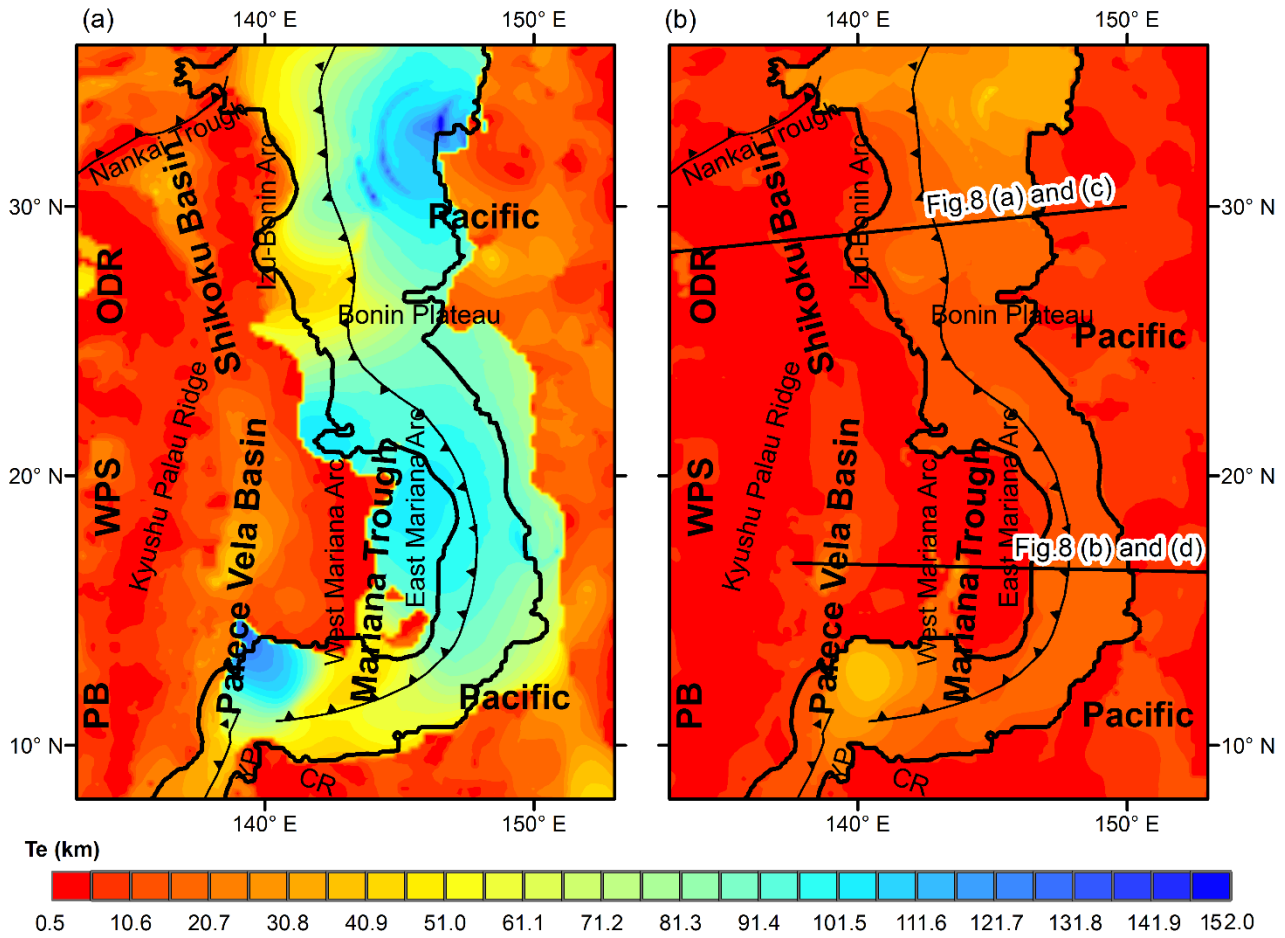
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680 Figure 6.  $T_e$  error maps (a) without dynamic corrections to the gravity and topography  
 681 data, and when dynamic corrections are based on the (b) PG2008, (c) Isaak 1989, and  
 682 (d) KK2001 methods, respectively. The central wavenumber for all maps is 2.668.

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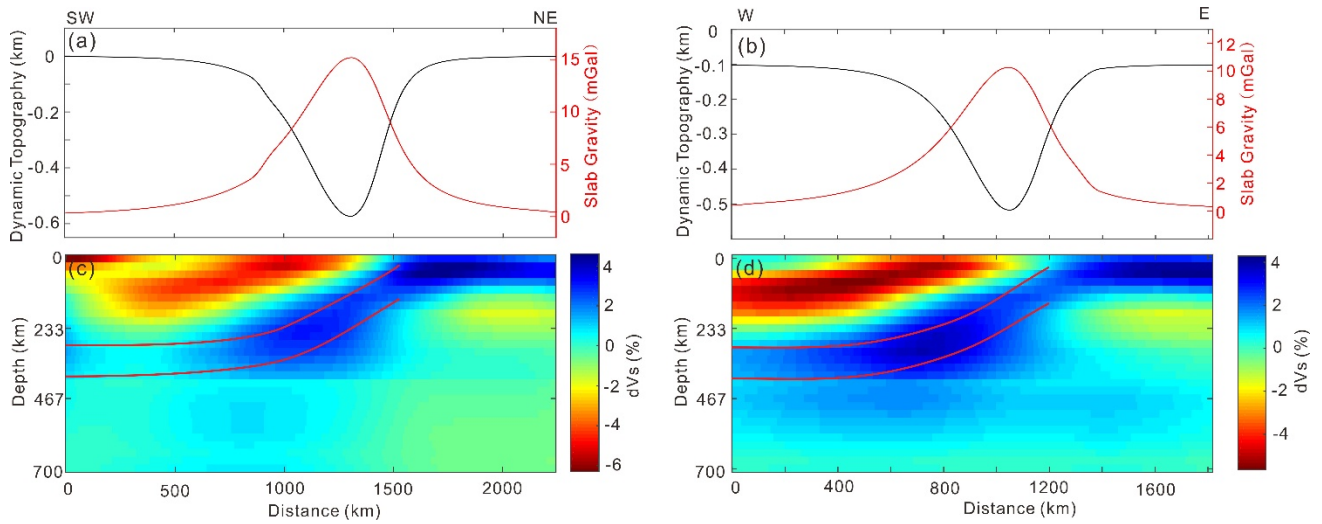
686 Figure 7.  $T_e$  maps (a) before and (b) after considering the influence of pull-force by the  
 687 subducted slab of the Pacific plate on topography and gravity ( $CW = 2.668$ ). The solid  
 688 black lines in (a) and (b) are the outer boundaries of the high- $T_e$  band from Figure 5b.

689 (a) is a reproduction of Figure 3a, while (b) is a reproduction of Figure 5b. All

690 abbreviations are the same as Figure 1.

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693 Figure 8. Two profiles show seismic velocity structure (on bottom panel), dynamic  
 694 topography and slab gravity (black and right curve respectively on top panel). Slab  
 695 density modelling is based on the Isaak1989 method for dynamic effect calculations.

696 The red curves on the bottom show the slab geometries with 100 km thickens. The  
 697 profile locations are shown in Figure 7.

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