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Independent patterns of water mass anomalies over Australia from satellite data and models

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Abstract

The Gravity Recovery and Climate Experiment (GRACE) products allow the quantification of total water storage (TWS) changes at global to regional scales. However, the quantity measured by GRACE represents mass signals integrated over vertical columns, requiring their separation into their original sources. Such a separation is vital for Australia, for which GRACE estimates are affected by leakage from the surrounding oceans. The independent component analysis (ICA) method that uses higher-order statistics, is implemented here to separate GRACE-derived water storage signals over the Australian continent from its surrounding oceans, covering October 2002 to May 2011. The performance of ICA applied to GRACE is then compared to the ICA of WaterGAP Global Hydrology Model (WGHM) and the ICA of the Australian Water Resources Assessment (AWRA) system. To study the influence of rainfall variability on the derived independent patterns, use is made of Tropical Rainfall Measuring Mission (TRMM) data set, from January 2000

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to May 2011. Implementing ICA on GRACE-TWS showed a remarkable improvement in separating the continental hydrological signals from the surrounding oceanic anomalies which was not achievable using a conventional principle component analysis. Reconstructing the continental TWS changes using only those independent components of GRACE that were located over the continent, showed a high correlation with WGHM-TWS and AWRA-TWS. Mass concentrations over the oceans and particularly S2 semi-diurnal aliased pattern were separated as independent modes. Correlation analysis between the independent components of GRACE and climate teleconnections showed that the mass anomalies over the northern ocean, Gulf of Carpentaria and north-eastern parts of Australia were significantly correlated with the El Niño-Southern Oscillation, while those over south and south-eastern parts of Australia were mainly linked to the Indian Ocean Dipole.

Key words: GRACE-TWS, ICA, Signal Separation, Australia

1. Introduction

Water availability in Australia is highly variable from year to year, with various parts of the continent (e.g., the southern and eastern regions) having suffered from severe drought conditions during the last decade (e.g., Ummenhofer et al., 2011). These drought episodes have possibly been worsened by higher temperatures (see, e.g., Nicholls, 2004; Ummenhofer et al., 2009; Leblanc et al., 2009). The long-term and inter-annual climate variabilities of Australia are affected by ocean-atmospheric phenomena such as the El Niño-Southern Oscillation (ENSO) (Risbey et al., 2009) and the Indian Ocean Dipole (IOD) teleconnections (Cai et al., 2011). The IOD, for example, nor-

11 mally affects most of East Africa and parts of the Indonesian and Australian
12 regions (Saji et al., 1999), while the ENSO mainly affects the north and east
13 parts of Australia (Nicholls, 1991) and parts of Africa (Becker et al., 2010).

14 Such variabilities, therefore, affect water availability over large regions of
15 Australia. Implementing a regional hydrological model, for instance van Dijk
16 et al. (2011) showed that during 2006 to 2009, ENSO had a high influence
17 on Australian terrestrial water variations.

18 Sustainable conservation and management of the Australian water re-
19 source, particularly in areas with arid or semi-arid climates, which include
20 many parts of Australia, requires implementing efficient monitoring tool(s)
21 (see e.g., Ellett et al., 2006 and Awange et al., 2009). One such a technique is
22 the Gravity Recovery And Climate Experiment (GRACE) twin satellite mis-
23 sion that routinely observes time-variable gravity signals within the Earth's
24 system (e.g., Tapley et al., 2004a,b). GRACE has found numerous applica-
25 tions in hydrology and ocean sciences, as documented, e.g., in Ramillien et
26 al. (2004), Schmidt et al. (2008), Awange et al. (2008a,b), Werth et al.
27 (2009a) and Chambers (2006).

28 The possibility of using GRACE monthly gravity field solutions (Tapley
29 et al., 2004) to monitor Australia-wide total water storage (TWS) signals was
30 assessed by Rodell and Famiglietti (1999) prior to the launch of GRACE. This
31 view was continued by Ellett et al. (2006) who showed that GRACE monthly
32 solutions are suitable for monitoring annual and inter-annual groundwater
33 variations in Australia. Leblanc et al. (2009) combined GRACE solutions
34 with hydrological observations and modeling results, and estimated a loss
35 of groundwater from the Murray-Darling Basin (MDB) of 104 km^3 for the

36 period between 2001 and 2007. Chen et al. (2005) found good agreement
37 between the GRACE-based estimates of terrestrial water storage variability
38 for the Victoria Basin and Northern Australia with the values of Global Land
39 Data Assimilation System (GLDAS) hydrological model (Rodell et al., 2004).

40 Though such possibility of GRACE application to Australian hydrology
41 had been realized, Awange et al. (2009) pointed out that (i) much of Aus-
42 tralia has a relatively small TWS signal, which is very difficult to detect using
43 the current GRACE system and processing strategies, and (ii) the effect of
44 considerable spatial and spectral leakage from the surrounding oceans masks
45 the GRACE-derived TWS over the land. Note that the impact of leakage is
46 not the same for different parts of the region. For instance, Tregoning et al.
47 (2008) found a notable land hydrological signal over the Gulf of Carpentaria,
48 while Brown and Tregoning (2010) reported that the hydrological variabil-
49 ity over the MDB does not need to account for contamination emanating
50 from other geophysical sources. This result was confirmed by van Dijk et
51 al. (2011) who showed that the dry inland of Australia exhibits little impact
52 from ocean leakage.

53 Awange et al. (2009) suggested, therefore, that reprocessing of GRACE
54 data tailored to Australia would be desirable for extracting hydrological sig-
55 nals. In this regard, Awange et al. (2011) evaluated a regional solution com-
56 puted with the mass concentration (*mascon*) method (Lemoine et al., 2007),
57 using the Principal Component Analysis (PCA; von Storch and Zwiers, 1999)
58 and multilinear regression analysis (MLRA) methods. Their study indicated
59 that the mascon products slightly improved the identification of TWS over
60 the Australian continent. The PCA method that was used to extract the

61 large scale spatial and temporal TWS patterns, however, resulted in a clus-
62 tered behaviour of its derived orthogonal components (e.g., showing seasonal
63 components and trend overlaid in the temporal components). This problem
64 is known as the ‘mixing problem’ of PCA, which has been addressed in other
65 studies, e.g., in Jolliffe (1989), Hyvärinen (1999a) and Forootan and Kusche
66 (2011).

67 Furthermore, the condition of water storage in Australia during the last
68 decade, is linked to various climatic factors, such as large-scale atmospheric
69 pressure changes (Nicholls, 2009), higher air temperatures (Nicholls, 2004),
70 Indian Ocean conditions (Cai et al., 2009) and Pacific Ocean variability
71 (Cai and Cowan, 2008). Using the Complex Empirical Orthogonal Func-
72 tion (CEOF) method, for example, García-García et al. (2011) showed that
73 the spreading of the annual water variations from the north to the southeast
74 of Australia are linked to the ENSO and IOD teleconnections.

75 The successful performance of higher-order decomposition methods such
76 as *Independent Component Analysis (ICA)* (Hyvärinen, 1999 a,b; Hyvärinen
77 and Oja, 2000) for detecting slow dynamic and inter-annual phenomena (e.g.,
78 ENSO) from climatic data sets (e.g., surface temperature, sea-level pressure,
79 and precipitation data) was shown by Ilin et al. (2005). Frappart et al.
80 (2010a, b) and Forootan and Kusche (2011) recently applied ICA to decom-
81 pose GRACE-derived TWS products. In Forootan and Kusche (2011), for
82 example, it is demonstrated that while PCA clusters the GRACE-derived
83 long-term and periodical TWS in several modes, the ICA algorithm, without
84 making prior assumptions about the signal, improves the results.

85 The present contribution applies the ICA approach of Forootan and Kusche

86 (2011) to decompose the Australian GRACE-derived TWS data from the
87 GeoForschungsZentrum (GFZ), Center for Space Research (CSR) and Bonn
88 University for the period of October 2002 to May 2011. For comparison, use
89 is made of the WaterGAP Global Hydrology Model (WGHM) (Döll et al.,
90 2003), Australian Water Resources Assessment (AWRA) system (van Dijk
91 and Renzullo, 2011 and van Dijk et al., 2011) and the Tropical Rainfall Mea-
92 suring Mission (TRMM) precipitation data set (Kummerow et al., 1998).
93 The main objective is to extract statistically independent water storage pat-
94 terns from satellite observations and models which, we believe, is necessary
95 for better understanding the large scale variability of the Australian hydrolog-
96 y. Once ICA separates the signals over the continent from the surrounding
97 oceans, those independent components that are located over the land will
98 be used to reconstruct the continental water storage changes. This is an
99 important step for comparison and calibration of hydrological models with
100 GRACE-TWS measurements (see e.g., Werth et al., 2009a), in which the
101 leakage of ocean signals on terrestrial signals has not been accounted for.
102 Moreover, the links between the derived independent patterns and ENSO
103 and IOD were also investigated to identify those areas that are controlled by
104 teleconnections and those that are not.

105 Therefore, ICA was applied to (i) decompose GRACE-derived TWS sig-
106 nals over Australia, and to (ii) decompose WGHM, AWRA and TRMM data
107 sets. The obtained results in (i) were then compared to those of PCA and the
108 individual independent components of (ii). To understand the impact of the
109 climate teleconnections, the correlation coefficients between GRACE-derived
110 independent components of (i) and the ENSO and IOD indices were derived.

111 This contribution is organized as follows; in section 2, we briefly explain
112 the ICA method starting with a discussion of PCA. The data used in the
113 study are presented in section 3, followed in section 4 by the ICA results from
114 the GRACE, WGHM, AWRA and TRMM data sets for Australia. Section
115 5 discusses the link between water variability in Australia and the climate
116 teleconnection (ENSO/IOD phenomena). The study is concluded in section
117 6.

118 **2. PCA and ICA Methods**

119 PCA is a standard data analysis method based on eigenvalue decompo-
120 sition of the auto-covariance or auto-correlation matrix of a data set. Given
121 an $(n \times p)$ data matrix \mathbf{X} , where n is the number of observations over time,
122 and p the number of locations from which the observations are carried out,
123 PCA decomposes the data matrix as

$$\mathbf{X} \simeq \mathbf{P}_j \mathbf{E}_j^T, \quad (1)$$

124 where \mathbf{E}_j contains the unit length eigenvectors (i.e., empirical orthogonal
125 functions EOFs) in its columns arranged with respect to the magnitude of
126 eigenvalues, \mathbf{P}_j are their corresponding temporal components (i.e., principle
127 components PCs), and $j < n$ is the number of retained dominant modes. A
128 discussion on the selection of the retained modes is addressed in detail, e.g.,
129 in Preisendorfer (1988).

130 ICA represents a ‘blind’ source separation method based on the assump-
131 tion of the independence of sources and the non-Gaussian distribution of the
132 observations. ‘Blindness’ refers to the recovering of source signals from ob-

133 served mixtures without knowing how they have mixed (Hyvärine, 1999a).
 134 Therefore, ICA can be viewed as an extension of PCA (Comon, 1994; Han-
 135 nachi et al., 2009). While PCA decomposes a spatio-temporal data set (re-
 136 duced by the long-term average) into sets of orthogonal modes, ICA decom-
 137 poses it into statistically independent sources. From a statistical point of
 138 view, orthogonality and independence are equivalent for Gaussian signals.
 139 However, for non-Gaussian signals, independence implies orthogonality, but
 140 the reverse is not true (Hyvärine, 1999a,b).

141 The non-Gaussianity of time series can be computed using the *kurtosis*
 142 of time series (Comon, 1994)

$$k = E(\mathbf{x}^4)/E(\mathbf{x}^2)^2 - 3, \quad (2)$$

143 where \mathbf{x} stands for each time series (columns of \mathbf{X}), E is the expectation
 144 function that is usually approximated by the time average. For our time
 145 series with a length of more than 90 months, a kurtosis greater than 0.5
 146 indicates a non-Gaussian distribution (Westra et al., 2007).

147 To derive statistically independent components (ICs) from the uncorre-
 148 lated modes detected by PCA, we follow the ICA algorithm (Fig. 1) proposed
 149 in Forootan and Kusche (2011). Similar to Comon (1994), the ICA algorithm
 150 is implemented as a rotation method, which involves the PCA approach as
 151 the starting point of an iterative process. ICs are computed while a measure
 152 of joint statistical independence of the PCA modes is maximized (Cardoso
 153 and Souloumiac, 1993).

154 To this end, fourth-order cumulants (a tensor of quantity indicated by
 155 $\mathbf{Q}_{\mathbf{x}}$ in Fig. 1) provide us with a generalized form of the fourth-order mo-

156 ments (Cardoso and Souloumiac, 1993). For time series with Gaussian dis-
 157 tributions, all fourth-order cumulants vanish (cf. Papoulis, 1991). Cardoso
 158 (1998) showed that when the time series are independent the fourth-order
 159 cumulants tensor \mathbf{Q}_x becomes diagonal.

160 Generally, there are two alternative ways of applying ICA to spatio-
 161 temporal data sets (e.g., GRACE-TWS monthly maps) in which temporally
 162 independent time series or spatially independent components are estimated
 163 (Forootan and Kusche, 2011). Here, we are interested in the spatially inde-
 164 pendent components since the goal of our analysis is to separate GRACE-
 165 derived TWS signals over the Australian continent from the surrounding
 166 oceans.

167 To derive the spatially independent components, the cumulants matrix
 168 \mathbf{Q}_x is built using the EOFs of Eq. (1) as discussed in detail in Forootan
 169 and Kusche (2011). \mathbf{Q}_x is then decomposed using PCA to provide the initial
 170 values for the rotation matrix \mathbf{R}_j used in

$$f(\hat{\mathbf{R}}_j) = \sum_{i=1}^{n^2} (\mathbf{R}_j^T \mathbf{Q}_x(\mathbf{M}) \mathbf{R}_j), \quad (3)$$

171 where n^2 is the number of the off-diagonal elements, and \mathbf{M} an (arbitrary)
 172 $n \times n$ matrix. $\hat{\mathbf{R}}_j$ is an orthogonal rotation matrix that minimizes the off-
 173 diagonal elements of \mathbf{Q}_x in Eq. (3) and makes it as diagonal as possible
 174 (Cardoso, 1999). Finally, $\hat{\mathbf{R}}_j$ obtained from Eq. (3) is used in Eq. (1), i.e.,

$$\mathbf{X} \simeq \mathbf{P}_j \hat{\mathbf{R}}_j \hat{\mathbf{R}}_j^T \mathbf{E}_j^T = \hat{\mathbf{P}}_j \hat{\mathbf{E}}_j^T, \quad (4)$$

175 to rotate the retained modes and make them as spatially independent as
 176 possible (Forootan and Kusche, 2011). Finally, $\hat{\mathbf{E}}_j = \mathbf{E}_j \hat{\mathbf{R}}_j$ contains the de-

177 rived spatial independent components in its columns, while the columns of
178 $\hat{\mathbf{P}}_j = \mathbf{P}_j \hat{\mathbf{R}}_j$ contain their corresponding temporal evolutions. The schematic
179 illustration of the ICA algorithm is shown in Fig. 1.

FIGURE 1

180 3. Data

181 3.1. ENSO and IOD indices

182 Australian inter-annual climate variability is strongly affected by the dom-
183 inant tropical phenomena, namely the El Niño-Southern Oscillation (ENSO)
184 and the Indian Ocean Dipole (IOD) (Risbey et al., 2009) teleconnections.
185 El Niño is known as a global-scale phenomenon occurring in the ocean and
186 atmosphere, which is counted as the most prominent source of inter-annual
187 variability in the global weather and climate. The ENSO phenomenon con-
188 sists of an extension of warm weather from the central and eastern tropical
189 Pacific Ocean to Indonesia leading to a major shift in weather patterns across
190 the Pacific (Trenberth and Hoar, 1996). ENSO has long-period (three to eight
191 years) and inter-annual teleconnection rainfall impact mainly over north and
192 eastern Australia (Cai et al., 2011).

193 ENSO is usually measured by a simple index which represents a large-
194 scale oscillation of the air mass between the southeastern tropical Pacific
195 and the Australian-Indonesian region (Ropelewski and Jones, 1987). Sus-
196 tained positive values of the index can be interpreted as indicative of La
197 Niña (drought) conditions, while sustained negative values indicate El Niño
198 (wet) conditions. In this study, we used the monthly Southern Oscillation

199 Index (SOI) from 2002 to 2011, provided by the Australian Bureau of Mete-
200 orology¹, where the pressure differences have been measured between Tahiti
201 and Darwin. It should be mentioned here that there are several ENSO indi-
202 cators which are computed based on sea surface temperature or pressure data
203 sets. Among of these indices, Nino3.4 is an ENSO indicator with generally
204 stronger correlation to Australian rainfall compared to SOI (e.g., Timbal et
205 al., 2010). In the present contribution, however, SOI was selected in order to
206 enable comparison of our results with those of García-García et al. (2011).
207 The index is scaled using its standard deviation resulting in a unitless time
208 series.

209 IOD is a coupled ocean and atmosphere phenomenon, a basin-scale mode
210 of sea surface temperature (SST) and wind anomalies, in the equatorial In-
211 dian Ocean that affects the climate of Australia and other countries surround-
212 ing the Indian Ocean (Saji et al., 1999). An IOD event is an inter-annual
213 variability that usually starts around May or June in Australia, peaks be-
214 tween August and October and then rapidly decays (Australian Bureau of
215 Meteorology). A positive IOD year is associated with the cooler than normal
216 SST in the tropical eastern Indian Ocean near Indonesia, and warmer than
217 normal water in the tropical western Indian Ocean near Africa. A positive
218 IOD results in a decrease of rainfall over parts of Australia. In fact, a negative
219 IOD usually evolves following a positive IOD, with a reverse configuration.

220 The gradient of the IOD event is termed the Dipole Mode Index (DMI)
221 which is usually considered as a measure of the Indian Ocean influence

¹<http://www.bom.gov.au/climate/enso/>

222 on Atmospheric pressure variabilities (Ajayamohan and Rao, 2008). Our
223 study made use of the available DMI data provided by the Japan Agency
224 for Marine-Earth Science and Technology (JAMSTEC)¹ covering the period
225 from 2002 to September 2010. Similar to the SOI, the DMI pattern is also
226 scaled with its standard deviation to be unitless and multiplied by -1 to be
227 consistent with the previous studies, e.g., García-García et al. (2011).

228 *3.2. GRACE solutions*

229 The GRACE satellites were launched on 17th March 2002 as a joint U.S.
230 National Aeronautics and Space Administration (NASA)/German Aerospace
231 Center (DLR) space mission dedicated to monitoring temporal and spatial
232 variations of the Earth's gravity field on a global scale. GRACE gravity field
233 solutions are generated at regular intervals (e.g., daily to monthly) by several
234 institutions, three of which are used in this study as discussed below.

235 The GFZ solutions: For this study, we used all available release 4 (RL4)
236 monthly spherical harmonic gravity products from August 2002 to May 2011
237 computed by the GFZ Potsdam (Flechtner, 2007). These models are derived
238 as fully normalized spherical harmonic coefficients of the geopotential com-
239 puted to degree and order 120, and have been augmented by the degree-1
240 term from Rietbroek et al. (2009) in order to account for the variation of the
241 Earth's center of mass with respect to a crust-fixed reference system. Note
242 that the degree-1 values, for the months after Rietbroek et al. (2009)'s study,
243 were extrapolated using its annual and semi-annual dominant frequencies.

244 The CSR and ITG2010 solutions: Two other data sets from the CSR

¹<http://www.jamstec.go.jp>

245 at the University of Texas, USA (Bettadpur, 2007) and Bonn University,
246 Germany (ITG2010) (Mayer-Gürr et al., 2010) were incorporated in order to
247 validate our findings from GFZ data set. These two solutions are based on the
248 same GRACE L1 (GPS and K-band ranging data) as the GFZ solutions. The
249 CSR solutions are provided as gravitational spherical harmonic coefficients
250 up to degree and order 60, available for the same period as the GFZ products.
251 The ITG2010 solutions have been derived using a different functional model
252 based on short arcs of the satellite orbit, which allows the set-up of full
253 empirical variance-covariance matrices for the observations. Furthermore,
254 daily solutions based on a Kalman filter approach have been calculated as
255 a part of the ITG2010 time series and have been introduced as additional
256 background model for the calculation of the monthly solutions. This different
257 treatment of the background models might cause differences in the solutions
258 provided by the different processing centers. A study by Bonin and Chambers
259 (2011), however, showed no significant discrepancy between the background
260 models in the region around Australia. Therefore we conclude that possible
261 differences between the solutions cannot be contributed to the background
262 modeling. The monthly ITG models are available for the period between
263 September 2002 and August 2009, and are complete to degree and order 120
264 (Mayer-Gürr et al., 2010)¹.

¹GFZ and CSR RL04 gravity fields are available from the Information Systems and Data Center (<http://isdc.gfz-potsdam.de>). ITG2010 gravity fields are available from the website of Astronomische, Physikalische und Mathematische Geodäsie (APMG) group at Bonn University (<http://www.igg.uni-bonn.de/apmg/index.php?iditg-grace2010>)

265 *3.2.1. Validation of GRACE products over Australia*

266 There have been several studies to validate GRACE signal in the Aus-
267 tralian region by independent data sets, from which it can be concluded
268 that GRACE is able to determine geophysically relevant signal over Aus-
269 tralia. Tregoning et al. (2009), for example, have found a good correlation of
270 GRACE surface mass variations with vertical crustal deformations as mea-
271 sured by GPS for a permanent GPS site located in Darwin. Kurtenbach
272 (2011) provided correlations of up to 0.6 between the daily ITG2010 solu-
273 tions and daily height change observations at different Australian GPS sites.
274 Tregoning et al. (2008) have found that the annual amplitude of non-tidal
275 mass changes at the Gulf of Carpentaria estimated from GRACE shows good
276 agreement with tide gauge observations at Groote Eylandt. For the MDB,
277 strong correlations of GRACE inter-annual decrease in total water storage
278 with modelled decreases of groundwater levels were reported by Leblanc et
279 al. (2009). Unfortunately, it has not been possible to validate the GRACE
280 solutions against in-situ measurements of ocean bottom pressure, as no such
281 recorders are available in the oceans around Australia (see Macrander et al.,
282 2010).

283 *3.3. WGHM hydrological model*

284 WGHM (Döll et al., 2003) is the global hydrological part of the Wa-
285 terGAP (Water-Global Assessment and Prognosis) global model of water
286 availability and use. The WGHM model represents the major hydrologi-
287 cal components, such as soil moisture, rainfall, snow accumulation, melting,
288 evaporation, runoff, and the lateral transport of water within river networks.
289 Detailed information about the modeling concept and its corresponding as-

290 sumptions can be found, e.g., in Güntner et al. (2007). In this study, we
291 used global TWS products provided on a 0.5° by 0.5° grid (with the ex-
292 ceptions of Greenland and Antarctica) covering the period between January
293 2000 and December 2009. The WGHM model has been previously used, e.g.,
294 by Awange et al. (2011) to study the Australian water storage variations.
295 Since WGHM represents, after vertical aggregation and forming temporal
296 anomalies, the same TWS parameter as GRACE-TWS, the two data sets
297 can be directly compared.

298 *3.4. AWRA regional hydrological model*

299 The Australian Water Resources Assessment (AWRA) is a system which
300 combines an operational hydrological model with meteorological and remote
301 sensing data to estimate water storages in the soil, surface water and ground-
302 water (van Dijk and Renzullo, 2011). AWRA water balance, using stream
303 flow data from several catchments and incorporating evapotranspiration mea-
304 surements along with several remotely sensed parameters, has shown a re-
305 liable performance to model Australian hydrological variations (van Dijk et
306 al., 2011). To study the terrestrial water variations using a local hydrological
307 model, therefore, this study applied the same TWS grids covering from Jan-
308 uary 2000 to December 2010 that have previously been used in van Dijk et
309 al. (2011). Similar to WGHM, TWS changes from AWRA are also directly
310 comparable to GRACE-TWS.

311 *3.5. Tropical Rainfall Measuring Mission*

312 TRMM is a joint NASA/Japan Aerospace Exploration Agency mission,
313 which was designed to monitor and study tropical rainfall in the latitude

314 range $\pm 50^\circ$ over inaccessible areas such as the oceans and un-sampled ter-
315 rains (Kummerow et al., 1998, 2000; Huffman et al., 2007). To study the
316 monthly total precipitation over Australia, we used the global average prod-
317 ucts (3B43), which is derived from TRMM instruments, as well as data from
318 a number of other satellites and ground-based rain-gauge data (Kummerow
319 et al., 2000; Huffman et al., 2007). The 3B43 data¹ is originally provided as
320 mm/hour rainfall, and covers a period from January 1998 to May 2011. We
321 converted the data between January 2002 and May 2011 to mm precipitation
322 for each month. The suitability of using TRMM for studying rainfall pat-
323 terns over Australia has been assessed e.g., in Fleming et al. (2011) showed
324 good correlation between TRMM and rain gauge data over Australia.

325 *3.6. Data processing*

326 In order to prepare the data sets for analysis, the following processing
327 steps were applied.

- 328 • The GRACE spherical coefficients at higher degrees are too strongly
329 affected by correlated noise and are, therefore, smoothed by apply-
330 ing the DDK2 decorrelation filter (Kusche et al., 2009). This filter is
331 anisotropic which makes it difficult to be compared with Gaussian fil-
332 ters; an in-depth discussion about shape of the filter is pointed out in
333 Kusche (2007). Applying DDK2 filter reduces the north-south striping
334 in GRACE monthly solutions (Werth et al., 2009b), which were then

¹TRMM is available from the Goddard Earth Sciences Data and Information Services Center (<http://disc.sci.gsfc.nasa.gov/precipitation>)

335 used to generate the global TWS values according to the approach of
336 Wahr et al. (1998).

337 • Similar to the GRACE products above, the DDK2 filter was applied to
338 the WGHM, AWRA and TRMM data sets to derive exactly the same
339 spectral content expansion as the GRACE filtered products. Other-
340 wise, the differing smoothness of the data sets would make comparison
341 between the TWS quantities difficult.

342 • After filtering, all data sets were converted to $0.5^\circ \times 0.5^\circ$ grids.

343 • From each data set, a rectangular region was selected encompassing
344 Australia, with the latitude between -44.5° to $-10.5^\circ N$ and longitude
345 between 112.5° to $154.5^\circ E$.

346 Each derived monthly map contained 5865 non-zero elements except for the
347 WGHM and AWRA products that had 4083 non-zero elements. Less non-
348 zero elements of WGHM and AWRA is due to the fact that hydrological
349 models do not cover the oceans. The data from GFZ covered 98 months of
350 GRACE solutions, the CSR solutions covered 99 months, while the ITG2010
351 included 84 months. The solutions of January 2003 and July to October
352 2004 were excluded from GRACE data during the data processing because
353 of their poor quality (Flechtner et al., 2010). WGHM covers 120 months,
354 AWRA covers 132 months while TRMM covers 124 months.

355 An overview of the signal root-mean-square (RMS) of the four main data
356 sets (GFZ-TWS, WGHM-TWS, AWRA-TWS and TRMM rainfall) is shown
357 in Fig. 2 in order to compare the signal strength over different regions of
358 Australia. The RMS indicates that the main water storage signal is detected

359 over northern Australia, where GRACE and AWRA showed stronger vari-
360 ability than WHGM and TRMM data sets. According to the RMS in Fig. 2,
361 WGHM showed a weak TWS variability over southern and western Australia
362 except a smaller region of the southwest coast. The TWS variability over
363 southern Australia detected by AWRA was also weaker than GRACE.

FIGURE 2

364 Before implementing the PCA and ICA methods on the data sets, in order
365 to account for the meridian convergence, an area weighting, with respect to
366 the square root of the cosine of the latitudes, was performed. Then the grid
367 maps were arranged in a matrix with grid points in its rows (same as the
368 matrix \mathbf{X} in Eq. (1)). \mathbf{X} was then centered by reducing the mean from each
369 of the columns and checked for non-Gaussianity using (Eq. (2)).

370 The derived kurtosis (Eq. (2)) shows that 48.8% of the GFZ, 49% of the
371 CSR, 52% of the ITG2010, 76% of the WGHM, 78% of the AWRA, and 78%
372 of the TRMM grid-point time series exhibit a non-Gaussian distribution.
373 We removed the dominant annual cycle from the data sets and then checked
374 the kurtosis of their residuals. The results indicated that still 38%, 42%,
375 43%, 51%, 51% and 53% of the inter-annual variabilities of GFZ, CSR, ITG,
376 WGHM, AWRA and TRMM time series respectively exhibit non-Gaussian
377 distribution.

378 4. Numerical results

379 4.1. Synthetic example

380 To show how ICA can help to extract true sources from a superposition
381 (summation) case, we designed a simple simulation before applying the PCA
382 and ICA methods to analyse the real hydrological data over Australia. Let
383 us assume that there are four anomaly concentrations over the north, west,
384 east and southeast of Australia, and that the regions have no overlap with
385 each other. Then we assume that the northern part exhibits only an annual
386 signal while the other parts exhibit a superposition of a low amplitude annual
387 signal and a linear trend (Fig. 3). Synthetic data is then generated by scaling
388 the known spatial anomalies with their corresponding temporal components
389 (see Fig. 3).

390 Now, both PCA and ICA methods are employed to separate the signals
391 in Fig. 3. The results are presented in Fig. 4, which clearly indicates that
392 the output of the PCA method contains a mixture of both linear trend and
393 annual signals over all regions, while the ICA method succeeds in separating
394 them into the original introduced signals. The mixture behaviour detected
395 in PCA happens because the method maximizes the variance explained by
396 each individual orthogonal component. Therefore, PCA clusters the linear
397 and annual signals in both modes (see Fig. 4, PCA results). The ICA
398 algorithm, however, uses the statistical information contained in the fourth
399 order cumulants to rotate the PCA components (Eq. (4)) so that they are
400 as statistically independent as possible. As a result, the introduced signals
401 (Fig. 3) that have different probability density characteristics are recovered
402 in two different modes (see Fig. 4, ICA results).

FIGURE 3

FIGURE 4

403 *4.2. PCA results from the GFZ data*

404 The PCA method in Eq. (1) was applied on the centered data matrices
405 derived from the GFZ TWS data. The derived eigenvalues along with their
406 corresponding cumulative variance percentages are shown in Fig. 5, GFZ.
407 The graph shows that the first two eigenvalues, corresponding to 57% of
408 the total variance, are dominant and well separated from the rest of the
409 eigenvalues. The other eigenvalues are considerably smaller than the first
410 two, as indicated by the flat shape of the graph (Fig. 5, GFZ).

FIGURE 5

411 Fig. 6 presents the first 6 PCA components of GFZ-TWS to be compared
412 with the ICA results of Figs. 7 and 8. According to Fig. 5, GFZ, selecting the
413 first 6 components reconstructs about 78.5% of the total variation (see Fig.
414 5, GFZ). The temporal patterns (PCs) are scaled using their corresponding
415 standard deviations such that they are unitless. The spatial patterns have
416 been multiplied by the standard deviation of their corresponding temporal
417 components and are given in millimeters. Therefore, one can consider each
418 of the spatial patterns as an anomaly map from which each mode of TWS
419 variability is reconstructed by multiplying the spatial patterns with their
420 corresponding temporal components using Eq. (1). This strategy is used for
421 all the decomposition results shown in this paper.

422 As shown in Fig. 6, the spatial pattern of each PCA component consists of
423 several concentrations of the hydrological signals over land and ocean. This is
424 more evident in the spatial pattern of EOF2 to EOF6. The temporal pattern
425 of PCs (see PCs in Fig. 6) also indicates that the long-term trend along with
426 annual and semi-annual cycles exist in nearly all of the components. Overall,
427 our PCA results contain more mixing behaviour than the one derived in,
428 e.g., Awange et al. (2011) from the GRACE data over Australia. This is
429 probably due to the fact that in Awange et al. (2011), the oceanic signals were
430 masked before the implementation of the PCA method, thus constraining
431 any contamination to the computed PCA modes from the ocean leakage.
432 However, this spatial masking of the mass signals itself could be a source of
433 error that needs to be considered (e.g., Fenoglio-Marc et al., 2007). In this
434 study, we did not mask the GRACE signals over the oceans but instead we
435 implemented ICA, by rotating PCA results (including the continental and
436 oceanic mass signals) towards independence, to spatially separate the mass
437 signals over the continent from the surrounding oceans.

FIGURE 6

438 *4.3. ICA results from GRACE, WGHM, AWRA and TRMM data sets*

439 Implementing the ICA algorithm on the data sets starts by applying PCA
440 (Eq. (1)) as the first step on each data sets before rotating them towards inde-
441 pendence (Fig. 1). The eigenvalues of the GFZ, WGHM, AWRA and TRMM
442 data sets along with their corresponding cumulative variance percentages are
443 shown in Fig. 5. Note that implementing ICA on WGHM and AWRA in-
444 dividually verifies the GRACE-ICA results from another perspective. Since

445 TWS from hydrological models represents the same GRACE-TWS over the
446 lands, theoretically, the result of the decomposition should be comparable
447 to that derived from GRACE. Following this section, we will show that the
448 ICA results from GRACE in several basins are in agreement with the ICA of
449 WGHM and AWRA. Note that the ICA results of TRMM provide additional
450 information on rainfall variability during the study period.

451 Fig. 5 shows that reconstructing more than 90% of the total variability
452 of the GRACE-TWS required selecting the first 11 PCA components. To
453 reconstruct more than 90% of WGHM, AWRA and TRMM, however, only 5
454 , 5 and 6 components were needed, respectively (see Figs. 5). The rest 10%
455 of the variance in data sets was assumed to be noise.

456 The eigenvalue spectrum of the TRMM data set is quite different from
457 those of GRACE, WGHM and AWRA (see Figs. 5, TRMM). For TRMM,
458 only the first component (with 71.3% of total variance) was well separated
459 and the rest of spectrum appears almost flat. For GRACE, WGHM and
460 AWRA the first 2 components with 57%, 79% and 69% of total variance
461 respectively were dominant.

462 In order to reduce the sensitivity of the ICA rotation (Eq. (3)) with
463 respect to unequal variance representation of the components, the first 2
464 dominant modes of GRACE, WGHM and AWRA were rotated first, then re-
465 spectively their remaining 9, 3 and 3 components. For TRMM that contained
466 only one dominant component, the ICA method was implemented in a single
467 step using the first 5 components. Constructing the fourth order cumulant
468 matrix \mathbf{Q}_x on the PCA results of each of the four data sets showed that
469 the matrix was not diagonal meaning that the PCA components were not

470 statistically independent (see Fig. 1). The ICA components were, therefore,
471 computed using Eq. (4) for all the data sets.

472 For brevity from GRACE data sets, we only present the spatial inde-
473 pendent patterns of the GFZ-TWS data in Fig. 7. To compare the signal
474 amplitudes, we projected the other GRACE-TWS data sets (from CSR and
475 ITG2010), WGHM-TWS, AWRA-TWS and TRMM rainfall data on to the
476 presented spatial base-functions derived from the ICA of GFZ-TWS (Fig.
477 7). The derived temporal amplitudes of those products are shown along
478 with the ICs of the GFZ-TWS data in Fig. 8. For TRMM, the results rep-
479 resent only the precipitation amounts over Australia and cannot therefore
480 be directly compared to the TWS values derived from GRACE, WGHM or
481 AWRA. However, projecting the TRMM results along with the TWS values
482 (from GRACE, WGHM and AWRA) provides information about the rainfall
483 contribution in each independent component. For those components that
484 were concentrated over the ocean, the WGHM and AWRA time series were
485 not projected, since hydrological models are limited only to the land and
486 do not cover the oceans. Additional results of ICA application to WGHM,
487 AWRA and TRMM are presented in Figs. 10, 11 and 12, respectively. IC5
488 of WGHM is not shown here since its corresponding variance magnitude was
489 considerably less than those of the first four components.

FIGURE 7

FIGURE 8

490 For comparing the ICA results of GRACE over the continent to the hy-
491 drological models, first, we selected the independent components 1, 4, 5, 6, 8

492 and 10 of Fig. 7 and Fig. 8 (the independent components of GRACE that
493 were located over the Australian continent). Then, GRACE-TWS changes
494 were reconstructed using Eq. (3). The RMS of the results is shown in Fig
495 9,(A). Comparing Fig. 9, A to Fig. 2, GFZ, it is clear that implementing ICA
496 has considerably isolated the signals of surrounding oceans (e.g., no anoma-
497 lies over the Gulf of Carpentaria and the eastern oceans can be seen). Fig
498 9, B shows the differences between the linear trend computed from the ICA-
499 reconstructed GRACE-TWS and AWRA-TWS, covering the period of 2003
500 to 2011. Comparing to AWRA, GRACE estimated stronger drying trend
501 in northwest Australia, as well as, stronger mass gain in east and northeast
502 Australia. These results confirmed van Dijk et al. (2011)'s findings, but they
503 associated the differences to the unexplained trend in AWRA. The magni-
504 tude of the differences derived in Fig. 9, B, however, were less than those
505 of van Dijk et al. (2011), which might be related to the different filtering
506 approaches. Finally, the GRACE-TWS variabilities separated by ICA over
507 the continent were temporally correlated with the time series of WGHM and
508 AWRA from 2003 to 2010 and 2003 to 2011, respectively. The correlation
509 results showed a high agreement between the two hydrological models and
510 the ICA-separated results of GRACE over the continent (see Fig. 9, C and
511 D).

FIGURE 9

FIGURE 10

FIGURE 11

FIGURE 12

512 Below we present a detailed a detailed comparison between GRACE and
513 the hydrological models (e.g., in term of trend and seasonal components).
514 From Fig. 7, it was seen that the first three components of GRACE show
515 TWS concentrations over the northern part of Australia, where their cor-
516 responding ICs show an annual cycle (Fig. 8). Particularly, IC1 isolates
517 the annual signal over north Australia, IC2 represents the oceanic mass over
518 the north of the Gulf of Carpentaria (c.f. Tregoning et al., 2008), and IC3
519 mainly shows annual mass change over the Timor Sea. The computed long-
520 term linear rates for IC1 and IC2 between October 2002 and January 2011
521 showed that these regions gained mass at rates of 6 ± 2 *mm/year* and 5 ± 2
522 *mm/year*, respectively. Vinogradova et al. (2011) pointed out that the
523 variability detected over the Gulf of Carpentaria may be well related to the
524 self-attraction and loading physics, that is missing e.g., in the Ocean Model
525 for Circulation and Tide (OMCT, Thomas, 2002) which is often removed
526 as a background model during the processing of GRACE products. The
527 computed linear rate for IC3 (Timor Sea) was not statistically significant,
528 showing an almost steady mass balance.

529 A comparison between the first three ICs of GFZ with the projected
530 temporal evolutions from WGHM, AWRA and TRMM, also showed a good
531 agreement (see Fig. 8). This was, to some extent, evident when the computed
532 RMS in Fig. (2) had a similar concentration with a comparable magnitude
533 over the northern regions. However, IC1 indicates that the signal magnitude
534 of WGHM is still less than that of GRACE. Other studies have similarly
535 shown that WGHM generally tends to have smaller seasonal amplitude than
536 GRACE (see e.g., Schmidt et al., 2008). Comparing to WGHM, projection of

537 AWRA on IC1 shows that it fits better to IC1 of GRACE. This result confirms
538 van Dijk et al, (2011)'s findings that decomposed GRACE and AWRA signals
539 into their linear trend and seasonal components. Temporal evolutions of
540 TRMM, from projection, shows a 1-month lag between precipitation and
541 mass changes in these areas agreeing with previous studies, e.g., Rieser et al.
542 (2010).

543 IC1 from WGHM and AWRA, respectively shown in Fig. 10 and Fig. 11,
544 are comparable to IC1 of GRACE in Fig. 7, showing a similar annual TWS
545 signal over the northern regions. IC2 of WGHM represents the southern
546 annual cycle with a 6-month phase difference to the north due to the pre-
547 dominant rainfall season (cf. Rieser et al., 2010). The southeastern part of
548 IC2 from WGHM is comparable to IC6 of GRACE in Fig. 7. IC4 of AWRA
549 shows the same annual signal as WGHM over southern regions. Its signal
550 amplitude, however, is larger than that of WGHM and closer to the GRACE
551 results. The IC1, IC2 and IC3 of TRMM are also related to the rainfall
552 pattern over the northern areas with three separate spatial concentrations
553 (shown in Fig. 12). The derived ICs of TRMM are mostly annual and are
554 comparable with the projected evolutions of TRMM (in Fig. 8) in terms
555 of phase and amplitude (see Fig. 12, middle). These results, individually,
556 confirm our findings in Figs. 7 and 8.

557 IC4 from GRACE isolates a long-term mass loss along with an annual
558 cycle over the northwestern Australia. A linear regression analysis of the
559 GRACE data from October 2002 to May 2011 shows a rate of loss of -19.2 ± 2
560 *mm/year* in the region. The linear rate derived from the projected time evo-
561 lution of WGHM is around one third of the GRACE linear trend (see IC4

562 in Fig. 8). The reason can be due to the fact that hydrological models
563 are not designed to represent the trends, rather, their strengths are to illus-
564 trate cyclic behaviours. The projected evolution of AWRA on IC4, however,
565 agrees better than WGHM with the GRACE estimate. The linear rate of
566 mass decrease of AWRA over the northwest region was $-13.2 \pm 2 \text{ mm/year}$.
567 The smaller linear rate derived from AWRA has been already reported in
568 van Dijk et al. (2011). IC3 of WGHM (see Fig. 10) shows a mass loss at
569 a rate of $-7 \pm 1 \text{ mm/year}$ over the northwestern regions, which is less than
570 the rate computed from GRACE (see IC4 in Fig. 8). IC3 of WGHM also
571 shows a water loss over the southeastern Australia (i.e., the MDB) at a rate
572 of $-8 \pm 1 \text{ mm/year}$, which was larger than the GRACE estimate (-5.1 ± 1
573 mm/year). IC2 of AWRA confirms IC4 of GRACE by isolating the mass
574 loss over northwest of the continent. IC4 of TRMM (see Fig. 12) localizes
575 the rainfall in the northwestern and western part of the northern Australia
576 showing less precipitation from 2004 to 2006, and a decline in rainfall after
577 the year 2007 in these regions (i.e. the amplitude of annual rainfall in 2004-
578 2006 is considerably less than the other years). TRMM results from 2002 to
579 2011, however, do not show any significant trend in the integrated amount of
580 precipitation. Considering a long period observation of rain gauges, van Dijk
581 et al. (2011) and McGrath et al. (2012) linked the mass loss in northwest
582 Australia to the dry period (2003 – 2010) after the unusually wet conditions
583 during ($\sim 1997 - 2001$).

584 IC5 derived from GRACE in Figs. 7 and 8 represents inter-annual TWS
585 changes along with an increasing rate over the eastern and northeastern parts
586 of Australia due to the 2010-2011 wet conditions. A linear regression analysis

587 of the data from October 2002 to May 2011 shows that the rate of water gain
588 was $12 \pm 2 \text{ mm/year}$ (see IC5 in Fig. 8). The computed average TWS change
589 rate from October 2002 to January 2010, the period before the floods, was
590 $7.4 \pm 3.6 \text{ mm/year}$. Considering the projected temporal evolutions (IC5 in
591 Fig. 8), amplitude of WGHM is less than GRACE in this region. This is due
592 to the fact that the dominant annual cycle of WGHM is underestimated as
593 well as the fact that the WGHM data used here does not cover the years 2010
594 and 2011 when the region exhibits a large mass change. Projected evolution
595 of AWRA on IC5 shows the similar pattern as that of GRACE (see van Dijk
596 et al., 2011)). IC3 and IC5 of AWRA show respectively mass variability over
597 northeastern and eastern Australia, thus confirming GRACE results. The
598 temporal evolution from projection of TRMM indicates the influence of the
599 2010 rainfall in the region, which was also evident from implementing ICA on
600 TRMM (see IC5 in Fig. 12). The temporal pattern of TRMM is very different
601 from those of GRACE and hydrological models, indicating the importance
602 of using other hydrological parameters in this region. The drought patterns
603 of the eastern regions during 2006 and 2007 can also be identified by IC5 of
604 TRMM.

605 IC6 of GRACE in Fig. 7 localizes TWS concentrated over the southeast
606 of Australia. The corresponding IC6 (in Fig. 8) shows that the annual cycle
607 is dominant. The temporal pattern of IC6 also contains a considerable inter-
608 annual variabilities in its evolution. The long-term trend of IC6 can be split
609 into three sections; one from 2002 to the last months of 2005 that shows a
610 mass gain of $8.2 \pm 4 \text{ mm/year}$, then a decline in mass storage with a rate
611 of $-45.2 \pm 8 \text{ mm/year}$ is detected until the starting months of 2007, and

612 finally a mass gain at a rate of 12 ± 4 *mm/year* to May 2011. From the
613 projected patterns, WGHM and AWRA show the same patterns as GRACE,
614 however, their signal amplitudes are smoother than those of GRACE. The
615 temporal projected pattern of TRMM does not follow the pattern of IC6
616 of GRACE except for the annual peaks that still show a 1-month lag with
617 GRACE results. IC6 of TRMM, individually, represents the rainfall anomaly
618 over the southeastern regions (see IC6 in Fig. 12).

619 During the GRACE processing procedure high frequency mass variations,
620 e.g. caused by oceanic tides, are reduced as they cannot be resolved by
621 monthly gravity field solutions. However, current ocean tide models are
622 not accurate enough to fully reduce the tidal signal in the GRACE data
623 (Knudsen, 2003). Due to the sampling characteristics defined by the GRACE
624 orbit configuration, these unmodelled high frequency tidal signals occur in
625 the monthly gravity field solutions at alias periods. One well-recognized
626 example is the S2 semi-diurnal tide, which is mapped onto a 161-day period,
627 and thus does not cancel out in the monthly solutions (Chen et al., 2008, Ray
628 and Luthke, 2006). IC7 of GRACE separates the S2 aliasing effect over the
629 ocean located in the northwest of Australia (Figs. 7 and 8). This pattern was
630 previously reported by Melachroinos et al. (2009) who fitted a predetermined
631 cyclic signal (with a period of 161 days) to the GRACE time series. One of
632 the contributions of this study, therefore, is exploring such a pattern as an
633 independent component without using any predefined deterministic model.
634 Exploring such pattern as an individual mode was not possible using PCA.
635 Fitting a sinusoidal function to IC7 showed a period of 161.4 days, agreeing
636 with the theoretical derivation of Ray and Ponte (2003), and matching the

637 observations of Melachroinos et al (2009).

638 The water loss in west Australia is summarized in the IC8 of GRACE (see
639 Fig. 7). This pattern is also extended towards the ocean. This shows that
640 the performance of ICA, for separating the relatively lower amplitude signals,
641 is decreased. A linear regression of IC8 shows a rate of -8.2 ± 3 *mm/year*
642 from March 2002 to December 2009. The computed linear rate of GRACE-
643 TWS change in the area during the years 2010 to 2011 was 6 ± 4 *mm/year*,
644 which was due to the increased precipitation in this period. IC8 also shows an
645 opposite TWS gain over central Australia corresponding to 7 ± 2 *mm/year*
646 from March 2002 to December 2009. The projected temporal evolution from
647 WGHM shows similarities in phase with GRACE-IC8, but its amplitude is
648 around one fifth that of GRACE. The same statement is also true for the
649 projected evolution of AWRA with an amplitude of 70% of GRACE-IC8.
650 The temporal evolution of TRMM shows similarity only for the years 2004
651 to 2006. Our results confirm Rieser et al. (2010)'s study that showed the
652 poor link between the TRMM pattern and the GRACE-TWS changes in the
653 central, western, and eastern parts of Australia.

654 IC9 and IC10 of GRACE in Fig. 7 show mainly inter-annual mass fluc-
655 tuations over the ocean in the northeastern and southwestern parts of Aus-
656 tralia (with a frequency of 99.7 and 121 days, respectively). Note that the
657 reported frequencies are derived by fitting a sinusoidal cycle. For deriving
658 more reliable results together with their associated uncertainty, one might
659 use advanced methods (e.g., in Schmidt et al., 2008). Finally IC11 localizes
660 the TWS anomaly over southeast Australia. The linear rates of TWS change
661 in these regions were not statistically significant. The temporal evolution of

662 WGHM derived from its projection on to the spatial pattern of IC8 shows
663 the same temporal pattern in southeast but its amplitude was 60% of that of
664 GRACE. Unlike the GRACE results, no clear TWS anomalies were detected
665 over the southern and western Australia from the ICA of WGHM (see Fig.
666 10) and AWRA. This can be related to the fact that WGHM and AWRA,
667 according to the derived RMS in Fig. 2, show less TWS variability over these
668 regions. Further research will need to address these differences.

669 **5. Effects of ENSO and IOD on Australia**

670 In the preceding section, the relation between the patterns of GRACE-
671 TWS in different regions of Australia and the values of WGHM, AWRA
672 and TRMM were studied (see Fig. 8). Since the derived GRACE-ICs are
673 spatially independent, one can investigate the water variability of each com-
674 ponents individually (without considering the others). This gives the unique
675 opportunity to study the links between climate teleconnections (i.e. ENSO
676 and IOD) and the derived GRACE-ICs.

677 To this end, first, the ICs of GRACE-TWS from GFZ, SOI, and DMI
678 were smoothed with a 12-month moving average filter and interpolated to
679 a regular monthly time steps covering October 2002 to May 2011 (see Fig.
680 13). Then, the long-period temporal correlations at 95% level of confidence
681 between ICs and the indices, for the periods of October 2002 to 2011 as well
682 as 2006 to 2011, were computed. Selecting the period 2006 to 2011, besides
683 the study period, for computing the correlations was due to the high influence
684 of teleconnections on the Australian TWS changes (see, e.g., van Dijk et al.,
685 2011). The significant correlation values are reported in Fig. 13. Note that

686 selecting SOI and DMI indices for studying teleconnections as well as the 12-
687 month filter for smoothing were done to make the results comparable with
688 previous studies e.g., García-García et al. (2011).

689 The correlation results indicate a strong stable influence of ENSO for the
690 period 2006 to 2011 on IC1, IC2, and IC3 (with correlations of 0.57, 0.76,
691 and 0.64, respectively). Computing the long-period correlation coefficients
692 (October 2002 to May 2011) between the first three ICs and SOI index also
693 showed significant correlations of 0.51, 0.71, and 0.61, respectively. These
694 correlations confirm the effect of the tropical ocean-atmosphere variability
695 associated with ENSO rainfall in the northern regions, where IC2 and IC3
696 (indicating the mass variabilities over the ocean north of Australia) have
697 stronger correlation than the northern land signal (IC1). García-García et
698 al. (2011) and Ummenhofer et al. (2009) reported similar results for the
699 northern region using Complex EOF.

700 Computed correlations between -DMI and the first three ICs were high
701 in some years, e.g., 2006 and 2009. Their long-period correlation values for
702 October 2002 to May 2011, however, were 0.35, 0.26, and 0.32, respectively.
703 The correlations decreased for the period of October 2002 to May 2011. This
704 might show that the effect of IOD on TWS variations over the northern
705 regions is relatively less compared to that of ENSO.

706 Correlations of IC4 and IC8 with SOI, show the contribution of ENSO
707 to the short-period mass gain between 2010 and 2011. Their long-period
708 correlation for October 2002 to May 2011, however, were not statistically
709 significant. This statement was also true for the long-period correlation of
710 IC4 and IC8 with -DMI.

711 The effect of ENSO is also evident in IC5 (i.e., TWS variations in the
712 eastern and northeastern regions) and IC9 (which concentrates over the ocean
713 in east Australia) with a significant correlation of 0.79 and 0.81 with SOI for
714 October 2002 to May 2011. For the same period, the correlation between
715 IC5 and -DMI was -0.36, while the correlation value between IC9 and -DMI
716 was -0.31. These values indicate the less long-period effect of IOD on TWS
717 variations over the eastern and northeastern regions compared to that of
718 ENSO.

719 A significant influence of IOD from middle of 2005 to 2009 over the south-
720 eastern and southern parts of Australia is shown in IC6 and IC11. The com-
721 puted correlation between IC6 and -DMI for the period of 2006 to 2011 was
722 0.82. This correlation, for the same period, between SOI and IC6 was 0.51.
723 The correlation value between IC11 and -DMI for October 2002 to May 2011
724 was 0.54. This value for IC11 and SOI was 0.44.

725 Correlation between IC10 and IOD, for the period 2006 to 2011 was 0.53,
726 while in this period no significant correlation with ENSO was found. Note
727 that -DMI is selected for computing the correlations since negative IOD indi-
728 cates increase in water budget and positive IOD indicates decrease in water
729 budget. Since the spatial anomalies are positive (as they are in Fig. 7, the
730 6th and 11th patterns), -DMI should follow the pattern of the derived ICs
731 (see Fig. 13). These results confirm the study of Ummenhofer et al. (2009)
732 which shows the effect of the Indian Ocean Dipole in the southern regions
733 (see McGrath et al., 2012).

FIGURE 13

734 6. Conclusion

735 In this contribution, large scale statistically independent hydrological pat-
736 terns over Australia have been extracted from remote sensing data and mod-
737 els using the ICA method. Our results indicate that:

738 ICA has the capability to isolate the effects of spatial and spectral leakages
739 from the surrounding oceans that mask potential terrestrial hydrological sig-
740 nals detectable by GRACE in a regional case such as Australia with weaker
741 hydrological signal, what was unachievable using mascon or PCA and its
742 ordinary extensions (see e.g., Awange et al., 2011 and PCA results in Fig.
743 6).

744 In this study, instead of spatial masking of GRACE-TWS over the oceans
745 before implementing the analysis, we extracted a rectangular region includ-
746 ing Australia from GRACE-TWS products, covering October 2002 to May
747 2011. Then an ICA algorithm (Fig. 1) was implemented on the GRACE
748 time series to spatially separate the continental TWS changes from the sur-
749 rounding oceans. The results appear to be well separated, while the derived
750 spatially independent patterns are localized over the continent or the oceans
751 (see Figs. 8 and 7). Reconstructing the GRACE-TWS over the Australian
752 continent using the ICs that were located over the land showed a significant
753 high correlation with the TWS time series derived from WGHM and AWRA
754 (see Fig. 9).

755 Some hidden physical processes such as S2 tidal-aliasing along with other
756 model deficiencies were also detected and localized from the data without
757 fitting any pre-defined deterministic model, which was what was done in
758 previous studies.

759 Verifying the ICs of GRACE-TWS by implementing ICA on the WGHM
760 and AWRA data sets individually showed good agreement in the results,
761 mainly for the dominant variabilities, e.g., mass change over the top north
762 of the continent or the mass loss over the northwest. We could not, however,
763 find the same patterns of e.g., IC2 of WGHM or IC5 of AWRA summarized
764 in one particular IC of GRACE (IC2 of WGHM or IC5 of AWRA were seen
765 in IC6 and IC8 of GRACE). This was due to the difference between the
766 components of GRACE and the hydrological models (i.e., GRACE compo-
767 nents cover both the continental and oceanic signals while those of WGHM
768 and AWRA only cover the continent). For instance, for reconstructing 90%
769 of the GRACE signal, we had to select 11 components while for WGHM
770 and AWRA, selecting only 5 components was enough. Following Eq. (4),
771 implementing the ICA rotation on GRACE was done using the first 2 then
772 the remaining 9 components, while for WGHM and AWRA, this was done
773 by selecting the first 2 then the remaining 3 components (see Section 4.3).
774 Selecting more componets for the GRACE case, therefore, forced the op-
775 timization procedure (Eq. (3)) to decompose some less dominant patterns
776 (e.g., over the southeast and southwest) into two modes (e.g., IC6 and IC8).
777 However, the reconstruction of GRACE signal over the continent (Fig. 9,
778 A) showed that the separation from the surrounding oceans was succesful.
779 Comparing RMS of the reconstruction and those of AWRA also showed a
780 good agreement. The derived high temporal correlations with the signals of
781 WGHM and AWRA also confirmed that, although some components are not
782 directly compareable, the reconstruction has performed well (see Fig. 9, A
783 and B).

784 Studying the correlation between the ICA-localized TWS results with the
785 ENSO and IOD phenomena revealed the influences of the climatic telecon-
786 nections on each individual statistically independent hydrological region. As
787 a result, a strong link between SOI and the TWS variations in the northern
788 regions and the relation of the IOD to the eastern and southeastern Australia
789 was established. ICA thus presents an alternative method of analysing the
790 relationship between hydrological changes and climate variabilities.

791 The presented ICA algorithm might be suitable for analysing other hy-
792 drological areas suffering from the leakage of surrounding oceanic signals,
793 and weaker hydrological signals.

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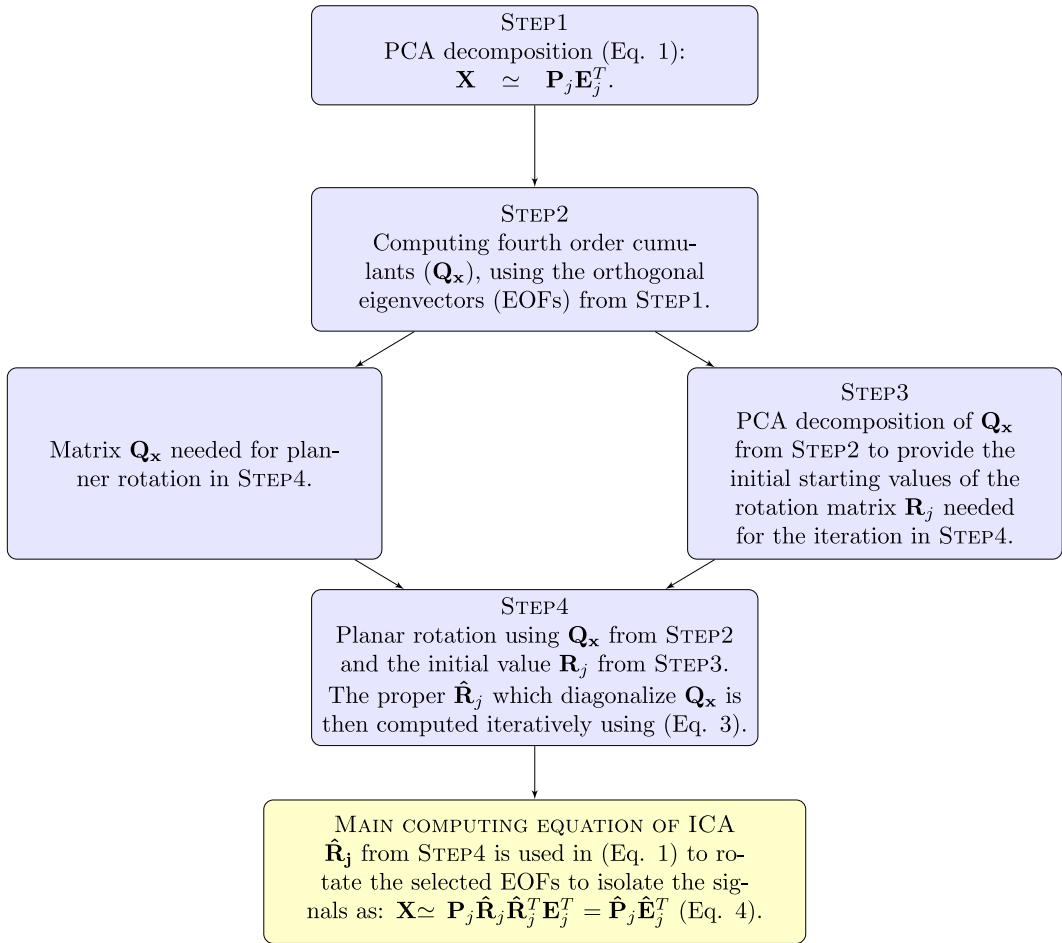


Figure 1: A schematic illustration of the ICA algorithm. For more details, we refer to Forootan and Kusche (2011).

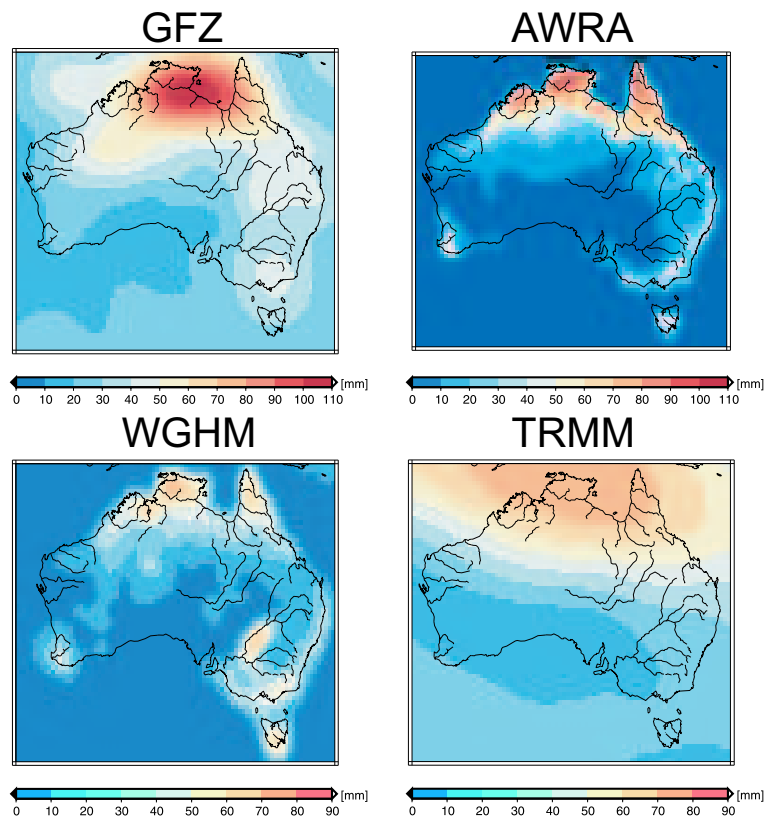


Figure 2: Comparing the signal variability (RMS) of the four main data sources used in this study after smoothing using the Kusche et al. (2009)'s DDK2 filter; GFZ-data (top-left), WGHM-data (bottom-left), AWRA-data (top-right) and TRMM-data (bottom-right).

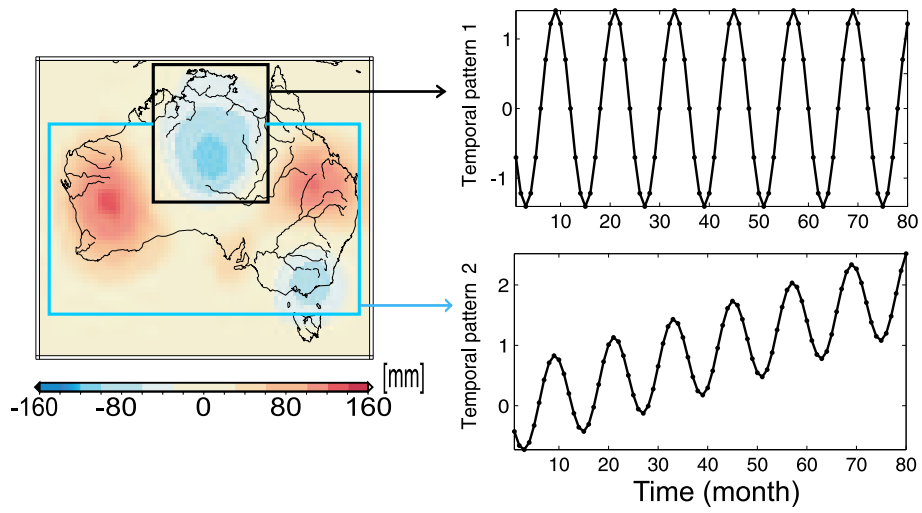


Figure 3: A synthetic example; the north of Australia exhibits only an annual signal while the west, east and southeast contain a superposition of a weaker annual signal and a linear trend. In order to reconstruct the synthetic data set, one should multiply the spatial pattern (left) with the temporal components (right).

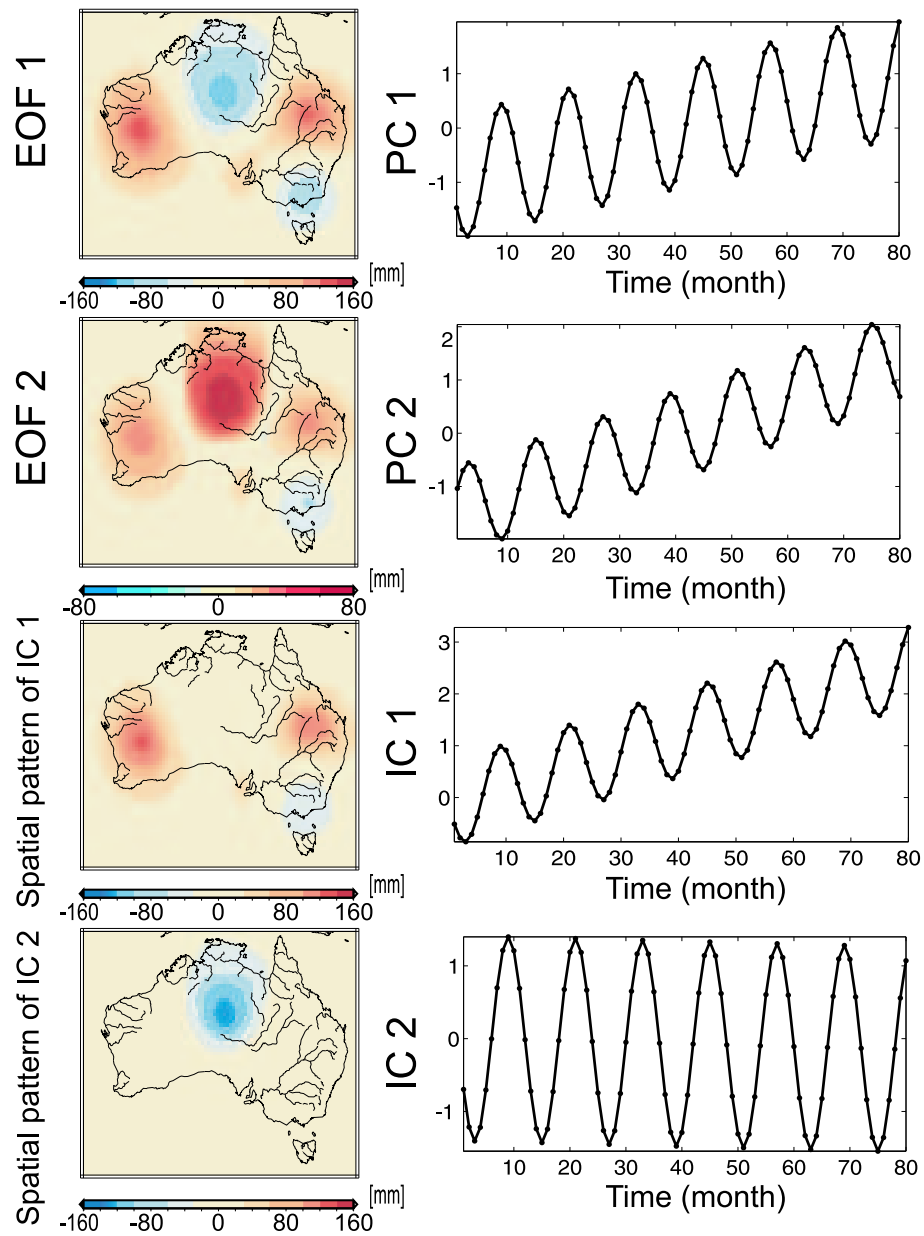


Figure 4: Separation of the simulated hydrological signals of Fig. (3) using PCA and ICA methods. The first two rows are related to the PCA results while the last 2 rows are related to the ICA results.

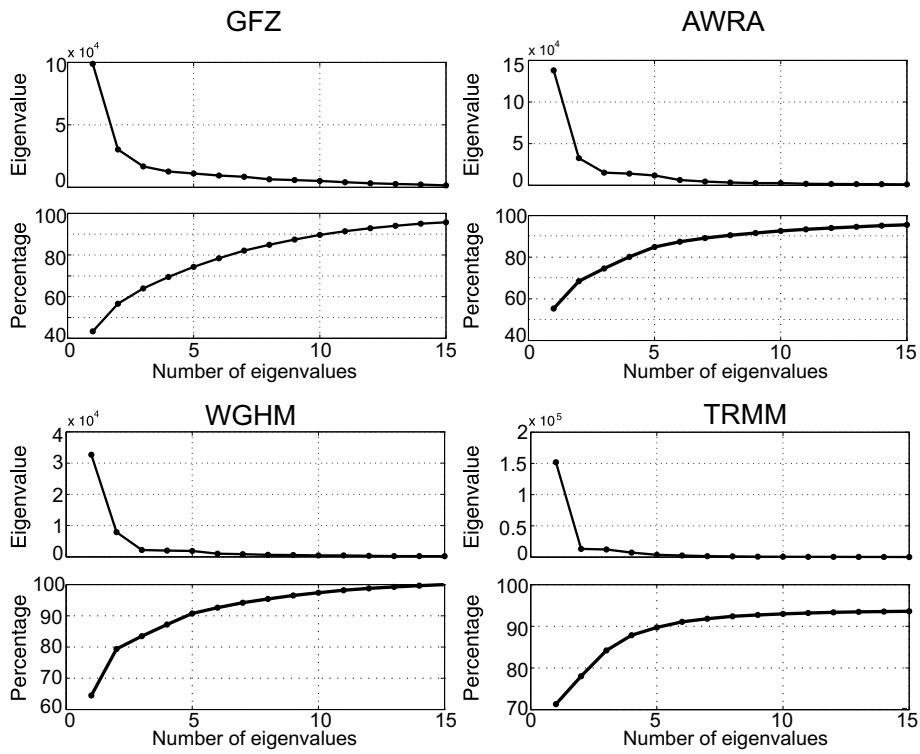


Figure 5: Eigenvalue results derived from implementing the PCA method on the GFZ, WGHM, AWRA and TRMM data sets are respectively shown in top-left, bottom-left, top-right and bottom right of the figure. Each sub-figure contains 2 graphs, including eigenvalue spectrum and cumulative variance contribution of the eigenvalues of each mentioned data sets.

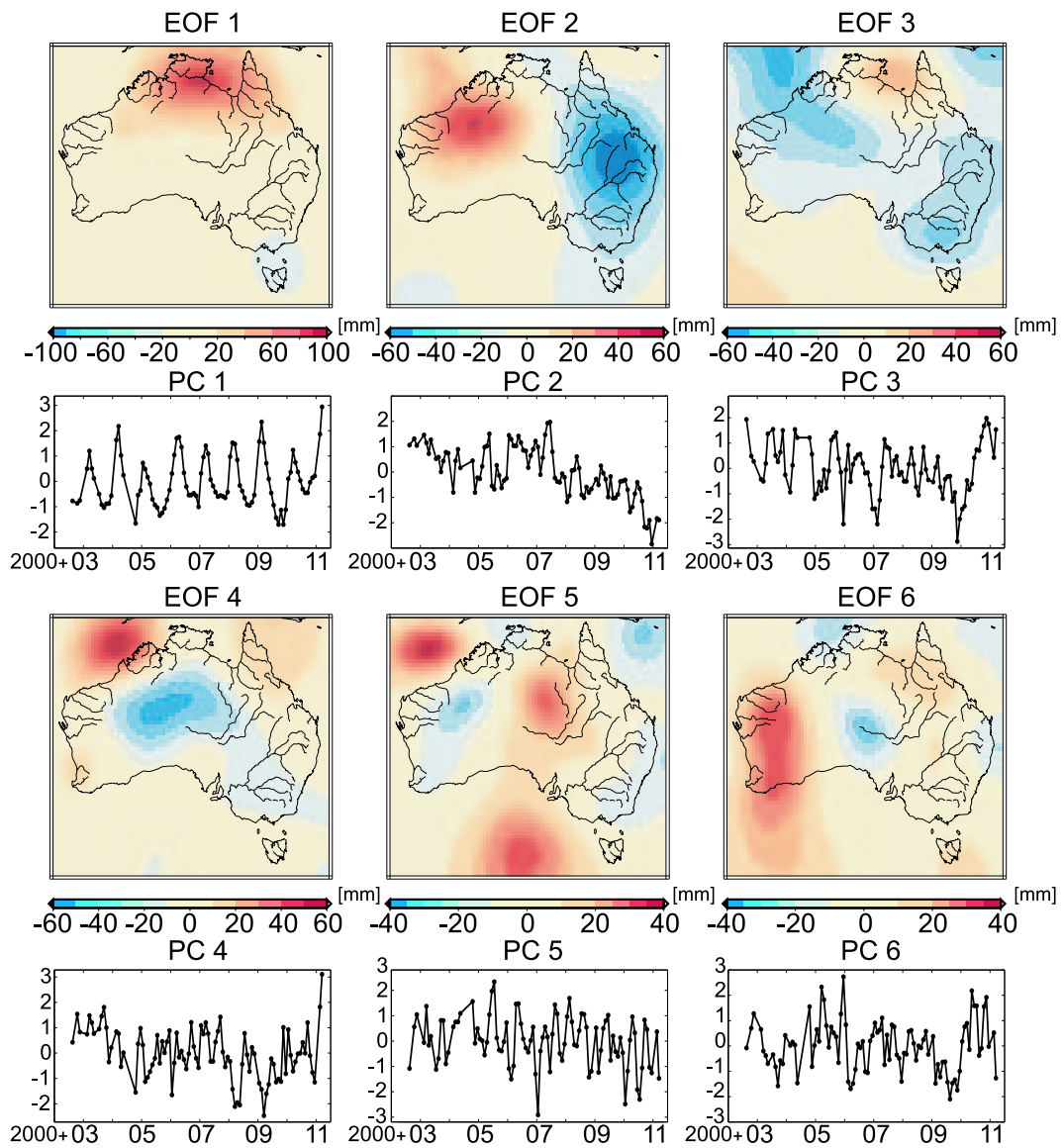


Figure 6: Results of PCA applied on total water storage maps of the GFZ center over Australia. The analysis covered the period from October 2002 to May 2011. The spatial anomalies are scaled using the standard deviation of their corresponding PCs to give millimeter unit. The temporal components are evolutions with standard deviation equal to 1.

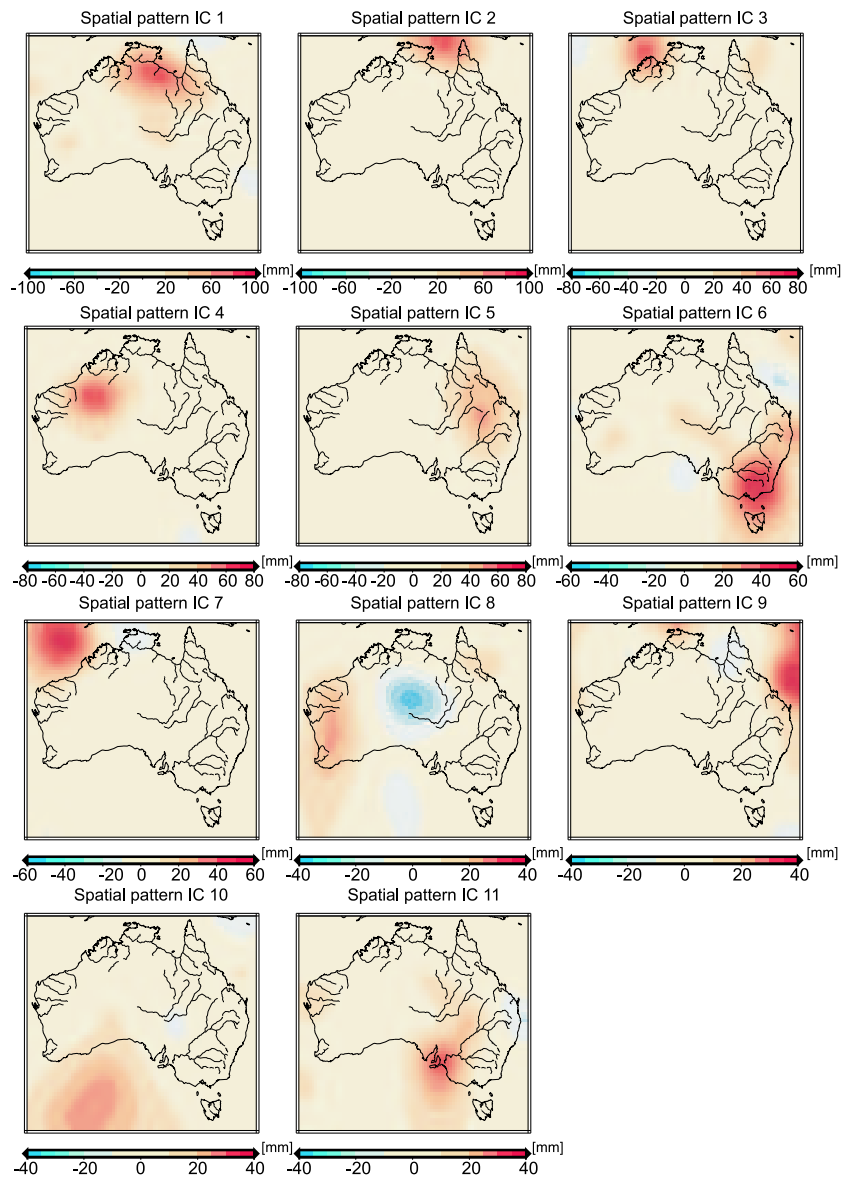


Figure 7: Results of the spatial ICA method applied to the GFZ total water storage maps over Australia. The spatial patterns are anomalies related to the GFZ data, which are scaled using the standard deviation of their corresponding temporal evolutions (shown in Fig. 8). The results are ordered according to the signal strength they represent such that they are comparable with those of PCA results in Fig. 6.

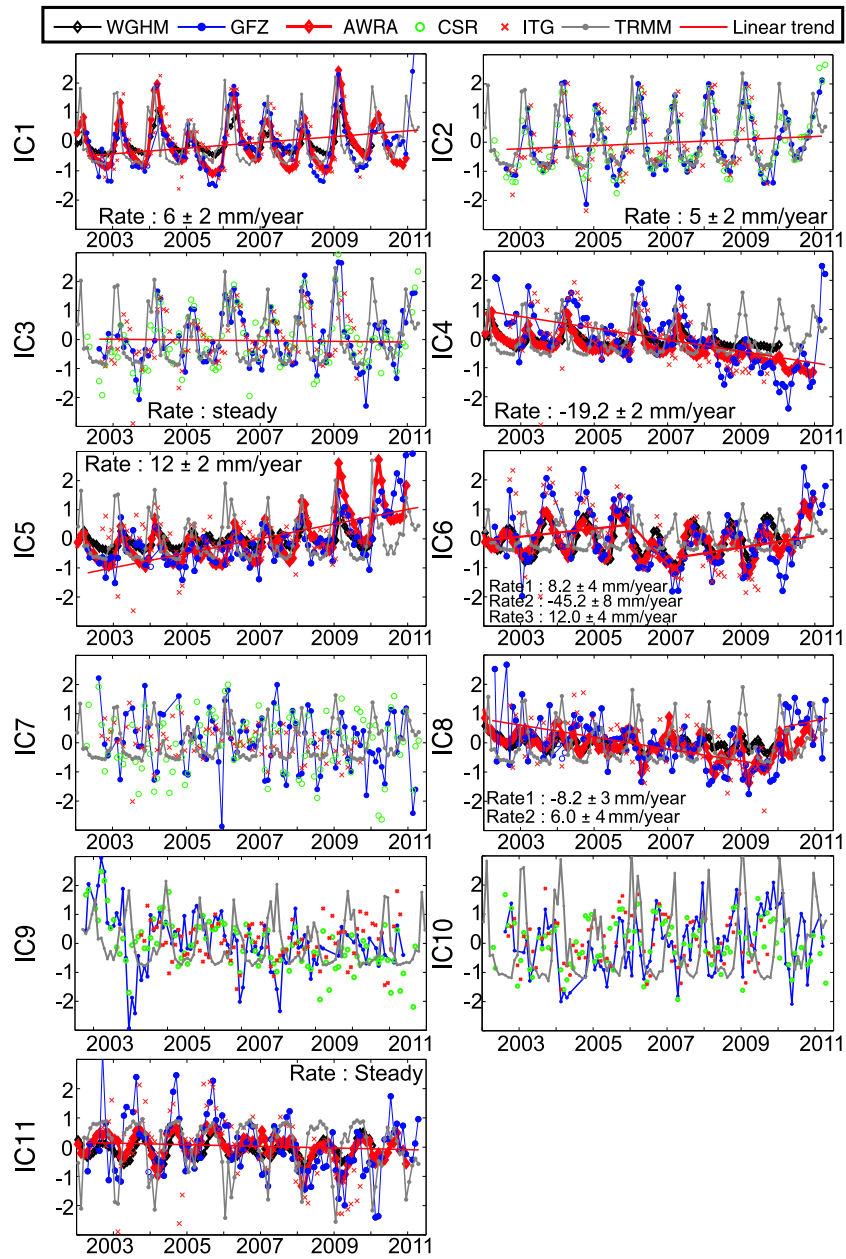


Figure 8: The evolutions of ICs corresponding to the spatial anomalies in Fig. 7 along with the projected temporal values of the CSR, ITG2010, WGHM, AWRA and TRMM data sets. For comparison purpose, the temporal evolutions are scaled using the standard deviation of the computed ICs of the GFZ data.

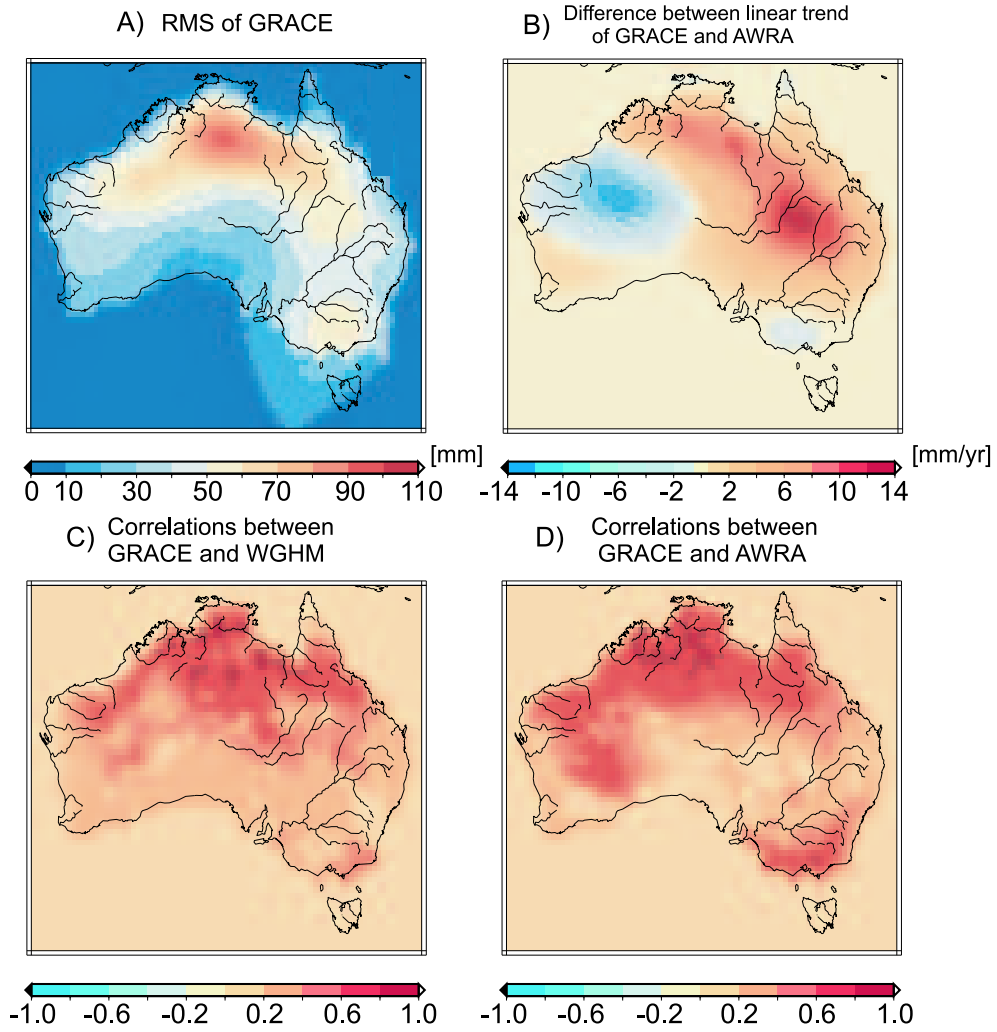


Figure 9: Overview of reconstructing GRACE-TWS variations over the Australian continent using ICA. (A) RMS of the reconstructed GRACE-TWS derived by multiplying the independent components of 1, 4, 5, 6, 8 and 10 of Fig. 7 to their corresponding temporal components in Fig. 8. (B) The difference between the linear trend computed from the ICA reconstructed time series of GRACE over the continent and the linear trend of AWRA, covering the period 2003 to 2011 (C) Temporal correlations between the ICA reconstructed time series of GRACE and the time series of WGHM for the period of 2003 to 2010. (D) Temporal correlations between the ICA reconstructed time series of GRACE and the time series of AWRA for the period of 2003 to 2011.

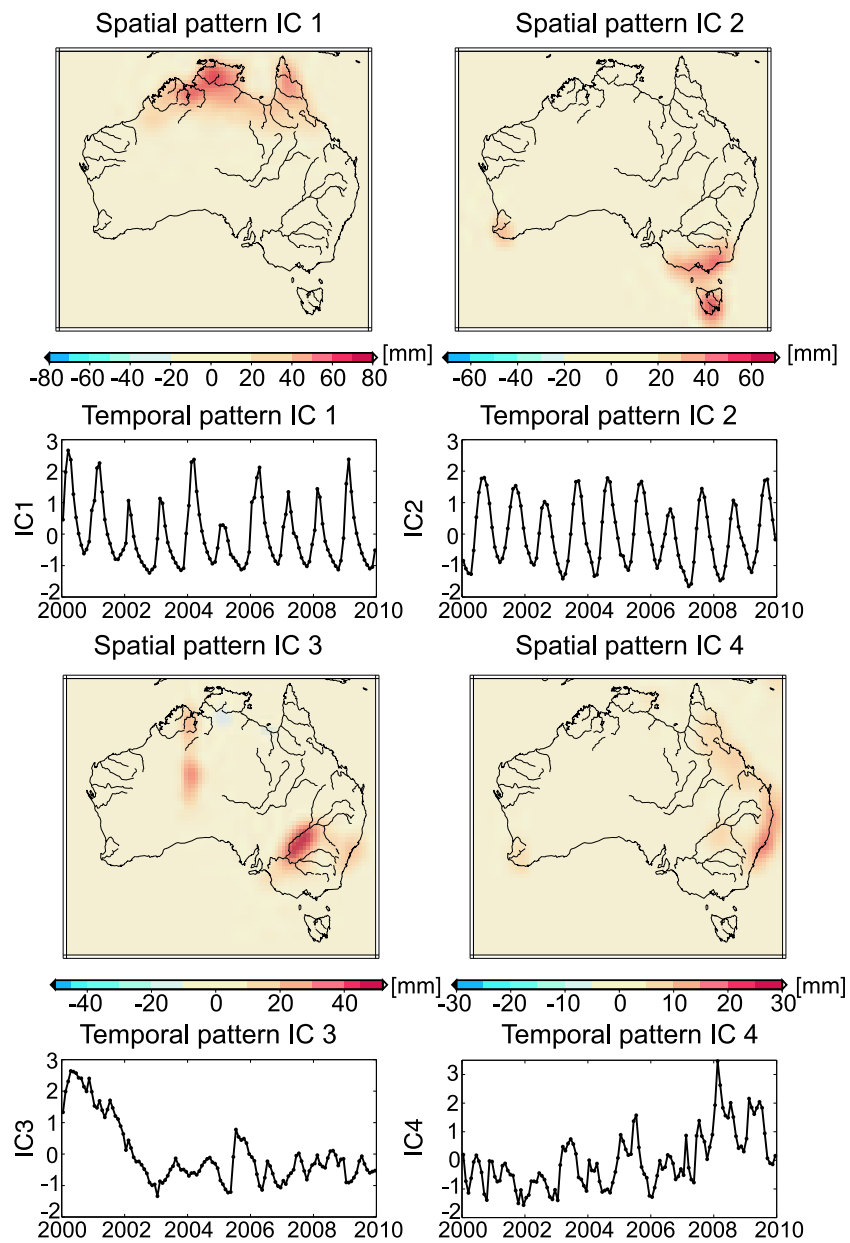


Figure 10: Overview of the performance of the ICA method on the WGHM hydrological data over Australia. The analysis covers the period between January 2000 and January 2010. The temporal components are scaled using their standard deviations such that they are unitless. The standard deviations are multiplied by spatial maps to give millimeter unit.

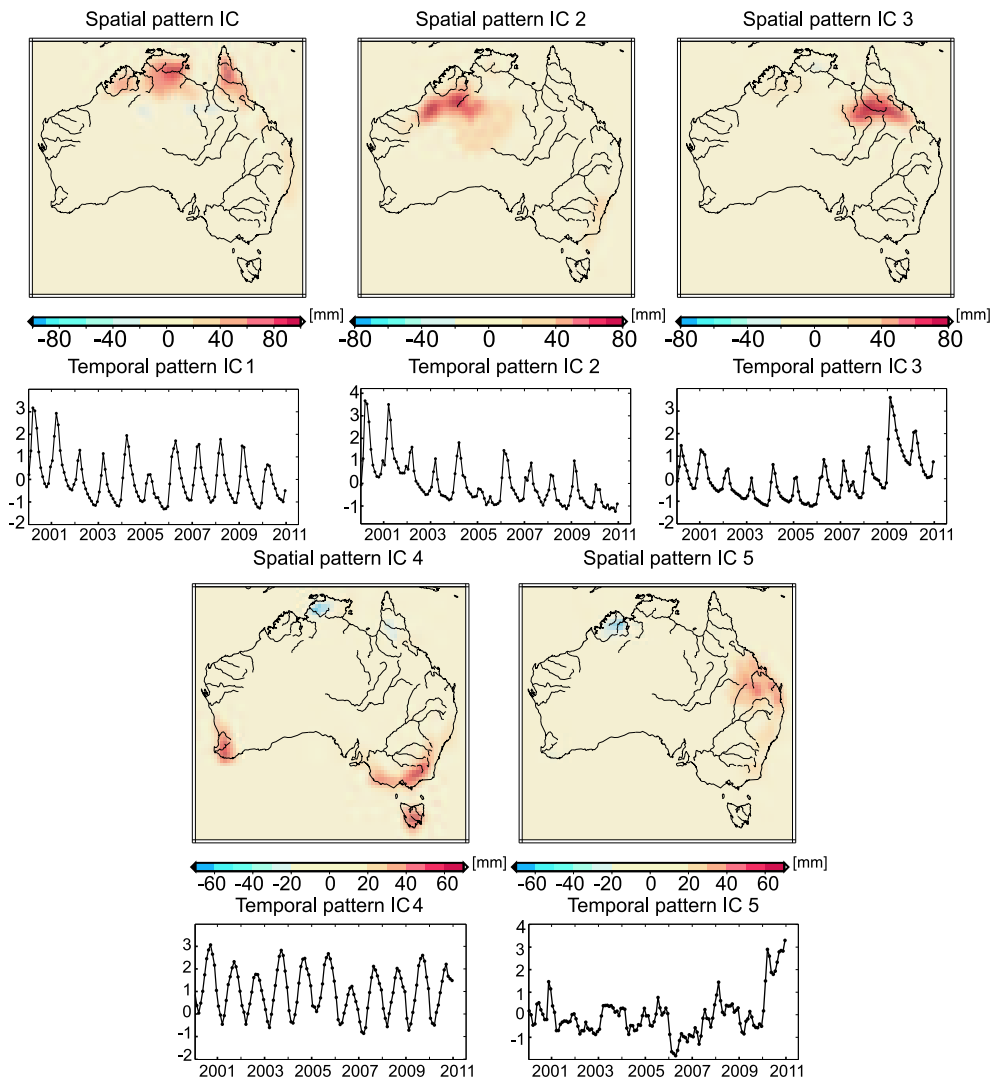


Figure 11: Overview of the performance of the ICA method on the AWRA hydrological data over Australia. The analysis covers the period between January 2000 and December 2010. The temporal components are scaled using their standard deviations such that they are unitless. The standard deviations are multiplied by spatial maps to give millimeter unit.

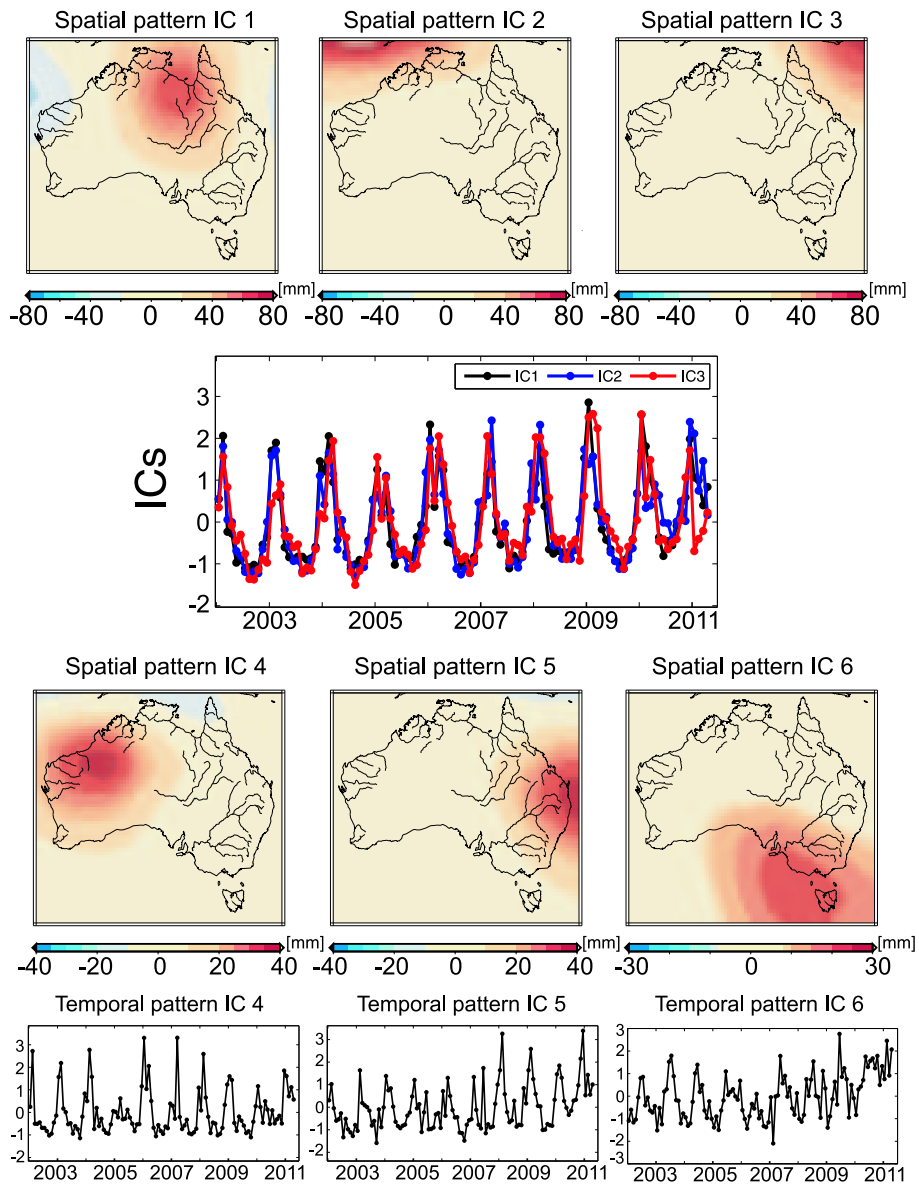


Figure 12: Overview of the performance of the ICA method on the TRMM data. The results are derived by rotating the first 5 EOFs, which contain more than 90% of the total variance of the data. The spatial patterns are anomalies that are scaled using the standard variation of their corresponding temporal evolutions. The temporal evolutions are unitless.

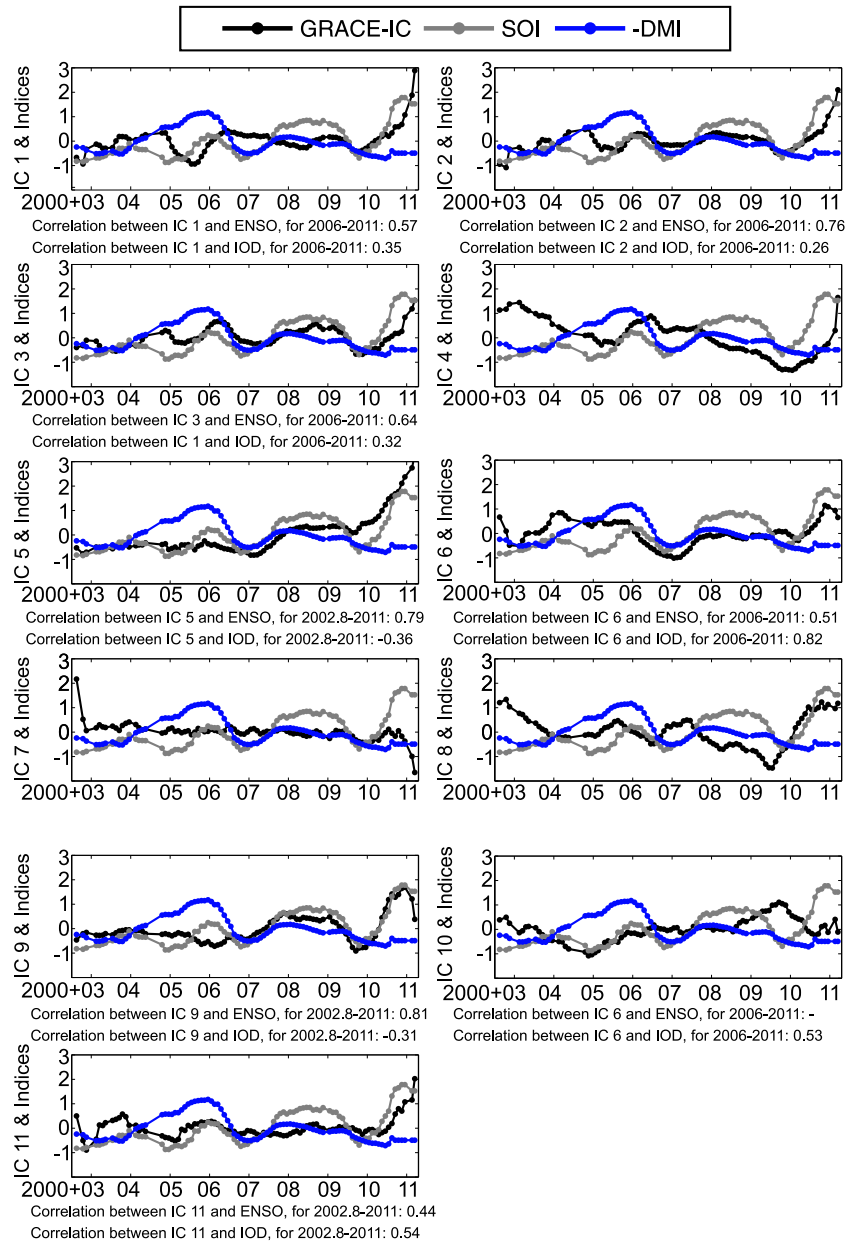


Figure 13: Overview of the temporal relation between ENSO and IOD with the Australian independent hydrological patterns derived from ICA of GFZ-TWS. In each graph, GFZ-ICs, SOI and -DMI indices are filtered using a 12-months moving average filter. The correlations are computed at 95% confidence level.