

# Fitting gravimetric geoid models to vertical deflections

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**Abstract.** Regional gravimetric geoid and quasigeoid models are now commonly fitted to GPS-levelling data, which simultaneously absorbs levelling, GPS and quasi/geoid errors due to their inseparability. We propose that independent vertical deflections are used instead, which are not affected by this inseparability problem. The formulation is set out for geoid slopes and changes in slopes. Application to 1080 astrogeodetic deflections over Australia for the AUSGeoid98 model shows that it is feasible, but the poor quality of the historical astrogeodetic deflections led to some unrealistic values.

**Keywords:** Gravimetric geoid errors, vertical deflections, vertical datum errors

## 1. Introduction

Fitting regional gravimetric geoid or quasigeoid models to GPS-levelling data has become a widespread practice. A principal objection to this is the inseparability of errors among the levelling and local vertical datum (LVD), GPS and gravimetric quasi/geoid model (cf. Featherstone 2004). While numerous different parameterisations

29 have been devised for this fitting (e.g., Milbert 1995; Jiang and Duquenne 1996;  
30 Forsberg 1998, Kotsakis and Sideris 1999, Fotopoulos 2005; Featherstone and Sproule  
31 2006; Soltanpour et al. 2006, etc.), it only ever models the reference surface of the LVD  
32 for GPS-based levelling, rather than the classical quasi/geoid (cf. Featherstone 1998,  
33 2006b).

34 On the other hand, astrogeodetically observed deflections (or deviations) of the  
35 vertical (i.e., from precisely timed observations to the stars) provide a source of  
36 terrestrial gravity field information that is independent of errors in the LVD (e.g.,  
37 Featherstone 2006a). Also, Jekeli (1999), Kütreiber (1999), Hirt and Flury (2007), Hirt  
38 et al. (2007), Hirt and Seeber (2008), Kütreiber and Abd-Elmotaal (2007), Marti  
39 (2007) and Müller et al. (2007b) demonstrate the utility of vertical deflections for  
40 gravity field determination and validation. Moreover, modern digital zenith cameras  
41 can now observe astrogeodetic vertical deflections to 0.1 arc-second in about 20 mins  
42 (e.g., Hirt and Bürki 2002, Hirt and Seeber 2007, Müller et al. 2007a). As such, vertical  
43 deflections will probably become more important for gravity field model validation (cf.  
44 Jekeli 1999; Featherstone and Morgan 2007, Pavlis et al. 2008).

45 In this short note, we propose that astrogeodetic vertical deflections are used to  
46 ‘correct/control’ errors in regional gravimetric quasi/geoid models, as a preferable  
47 alternative to the widespread use of using only GPS-levelling data because of the  
48 inseparability problem. This is akin to the classical orientation of a reference ellipsoid  
49 to a regional geodetic datum (e.g., Mather 1970, Mather and Fryer 1970). We present  
50 functional models for the two-, three- and four-parameter vertical deflection fitting  
51 (essentially geoid slopes and changes in slopes), which are then applied to 1080

52 historical astrogeodetic vertical deflections and vertical deflections derived from  
 53 AUSGeoid98 (Featherstone et al. 2001) over Australia.

54

## 55 **2. Background & Definitions**

56 Vertical deflections can either be absolute or relative, depending respectively on  
 57 whether a geocentric or local reference ellipsoid (and datum) is used in their definition  
 58 (Jekeli 1999; Featherstone and Rieger 2000). Here, we will only deal with absolute  
 59 vertical deflections since modern gravimetric quasi/geoid models refer to a geocentric  
 60 reference ellipsoid, and geodetic coordinates (used to compute the astrogeodetic vertical  
 61 deflections; see below) are directly or indirectly (i.e., by datum transformation) on a  
 62 geocentric datum and geocentric reference ellipsoid.

63

### 64 *2.1 Astrogeodetic deflections*

65 Astrogeodetic observations to the stars lead to natural/astronomic coordinates (latitude  
 66  $\Phi$ , longitude  $\Lambda$ ) of a point on or just above the Earth's surface, which when compared  
 67 with geocentric geodetic coordinates (latitude  $\phi$ , longitude  $\lambda$ ) of the same point yield  
 68 absolute Helmert (i.e., at the Earth's surface; cf. Jekeli 1999) north-south ( $\xi$ ) and east-  
 69 west ( $\eta$ ) deflections according to (e.g., Bomford 1980):

$$70 \quad \xi_H = \Phi - \phi \quad (1)$$

$$71 \quad \eta_H = (\Lambda - \lambda) \cos \phi \quad (2)$$

72 where subscript  $H$  is used to distinguish these as Helmert deflections. Sign conventions  
 73 mean that the deflection in the meridian  $\xi$  is positive north and negative south, and the  
 74 deflection in the prime vertical  $\eta$  is positive east and negative west.

75

76 *2.2 Gravimetric deflections*

77 Absolute Pizzetti deflections (i.e., deflections at the geoid; cf. Jekeli 1999) can be  
 78 computed directly by Vening-Meinesz's integral (e.g., Heiskanen and Moritz 1967), or  
 79 can be computed indirectly from horizontal gradients of a gravimetric geoid model by  
 80 (e.g., Torge 1991)

$$81 \quad \xi_P = \frac{-\Delta N}{\rho \Delta \phi} \quad (3)$$

$$82 \quad \eta_P = \frac{-\Delta N}{\nu \Delta \lambda \cos \phi} \quad (4)$$

83 where subscript  $P$  is used to distinguish these as Pizzetti deflections. The same sign  
 84 conventions as for astrogeodetic deflections also apply here. In Eqs. (3) and (4),  $\Delta N$  is  
 85 the change in the geoid height between grid nodes of latitude spacing ( $\Delta \phi$ ) and  
 86 longitude spacing ( $\Delta \lambda$ ),  $\rho$  is the radius of curvature of the [geocentric] reference  
 87 ellipsoid in the meridian,

$$88 \quad \rho = \frac{a(1-e^2)}{\left(\sqrt{1-e^2 \sin^2 \phi}\right)^3} \quad (5)$$

89 and  $\nu$  is the radius of curvature in the prime vertical

$$90 \quad \nu = \frac{a}{\sqrt{1-e^2 \sin^2 \phi}} \quad (6)$$

91 where  $e$  is the first numerical eccentricity and  $a$  is the semi-major axis length of the  
 92 reference ellipsoid; GRS80 (Moritz 1980) is used here.

93

94 *2.3 Curvature and torsion of the plumbline*

95 The curvature and torsion of the plumbline (cf. Grafarend 1997) cause a [small] angular  
 96 difference between Helmert and Pizzetti deflections, which is a function of 3D position.

97 However, the curvature and torsion are rather difficult to estimate accurately because  
 98 they require detailed knowledge of the shape of and mass-density distribution in the  
 99 topography (e.g., Heiskanen and Moritz 1967; Bomford 1980). Here, they are assumed  
 100 to be small (less than one arc-second) and thus neglected in the sequel, but in order to  
 101 achieve the best results in terms of theoretical consistency, they should be computed  
 102 and applied to the [astrogeodetic] Helmert deflections to give Pizzetti deflections  
 103 consistent with the geoid model.

104

### 105 **3. Functional Model**

106 A common mathematical model used to fit regional gravimetric quasi/geoids to GPS-  
 107 levelling has been a bias (simultaneously accounting for the deficient zero-degree term  
 108 in the quasi/geoid, LVD offsets and other constant biases (cf. Prutkin and Klees 2007))  
 109 and two orthogonal tilts (simultaneously accounting for the deficient first-degree terms  
 110 in the quasi/geoid, long-wavelength quasi/geoid errors, long-wavelength distortions in  
 111 the LVD and other tilts between the data). These all reflect the inseparability problem.

112 The origin of this popular four-parameter functional model can be traced back to  
 113 Heiskanen and Moritz (1967, Sects 2-18 and 2-19), where the scale and origin  
 114 deficiencies in a gravimetric geoid model  $\delta N$ , due to the inadmissible zero- and first-  
 115 degree terms, may be determined using external geometrical control via

$$116 \quad \delta N = N_0 + \Delta X \cos \phi \cos \lambda + \Delta Y \cos \phi \sin \lambda + \Delta Z \sin \phi \quad (7)$$

117 where  $N_0$  is the zero-degree term in the geoid representing the scale deficiency, and  
 118  $\Delta X, \Delta Y, \Delta Z$  are the three orthogonal origin shifts of the geocentre from the centre of the  
 119 reference ellipsoid (Heiskanen and Moritz 1967). This model is analogous with a four-  
 120 parameter geodetic datum transformation (cf. Kotsakis 2008).

121 Equation (7) has often been recast in the simpler equivalent form of a biased, tilted  
 122 and warped plane (cf. Forsberg 1998), giving

$$123 \quad \delta N = A + B\phi + C\lambda + D\phi\lambda \quad (8)$$

124 where  $A$  is the bias term (equivalent to  $N_0$  in Eq. (7)),  $B$  and  $C$  describe the tilted plane  
 125 in  $\phi$  and  $\lambda$ , and  $D$  allows for the tilted plane to be warped into a hyperbolic paraboloid  
 126 (e.g., Farin 2001, p.246).

127

128 The difference between astrogeodetic and geoid-derived deflections is parameterised  
 129 similarly here to give for the north-south (N-S) component

$$130 \quad \delta\xi = a_{00} + a_{10}\phi + a_{01}\lambda + a_{11}\phi\lambda \quad (9)$$

131 and for the east-west (E-W) component

$$132 \quad \delta\eta = b_{00} + b_{10}\phi + b_{01}\lambda + b_{11}\phi\lambda \quad (10)$$

133 where  $\delta\xi = \delta\xi_{\text{astro}} - \delta\xi_{\text{grav}}$  and  $\delta\eta = \delta\eta_{\text{astro}} - \delta\eta_{\text{grav}}$  are the N-S and E-W deflection  
 134 differences, respectively. Simplifications of these models down to two and three  
 135 parameters will be tested later.

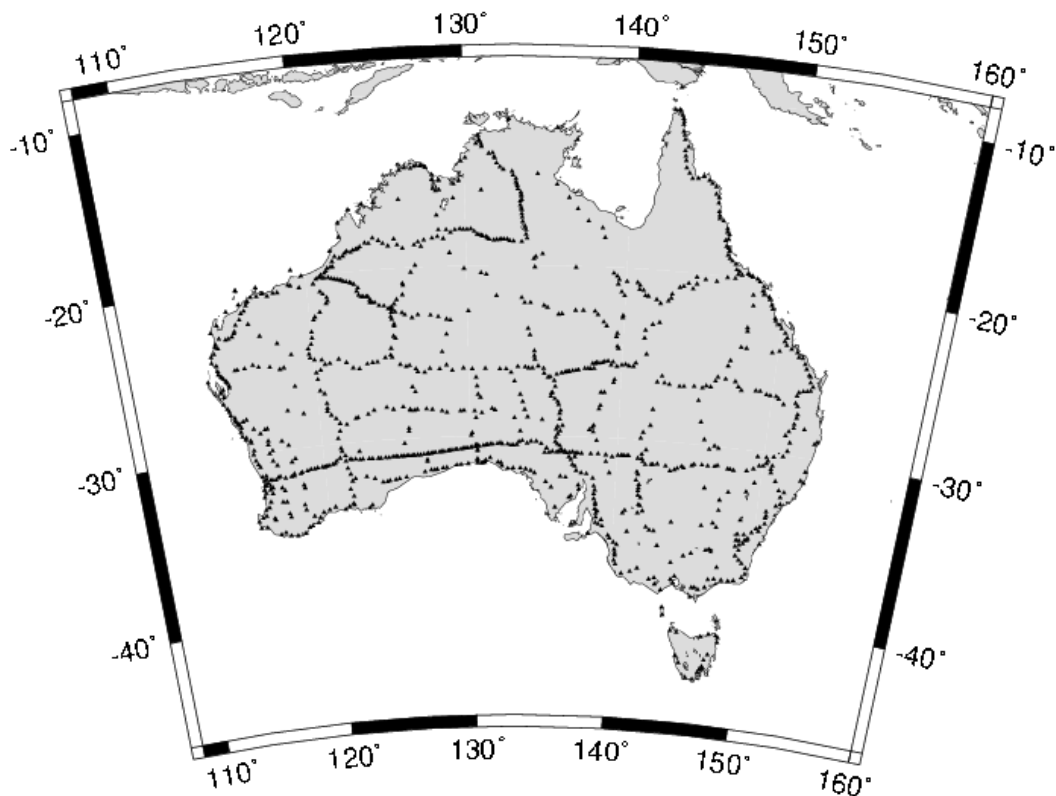
136 Since vertical deflections are second derivatives of the Earth's disturbing potential,  
 137 the interpretation of the parameters in Eqs. (9) and (10) is slightly different to that for  
 138 Eqs. (7) or (8). Firstly, the zero-degree term in the geoid (or LVD offset or other  
 139 constant biases) is indeterminate from vertical deflections; since they are angular  
 140 measures, they are insensitive to a scale change. The bias terms  $a_{00}$  and  $b_{00}$  in Eqs. (9)  
 141 and (10) represent the average difference in N-S and E-W tilts between the gravimetric  
 142 geoid and the [orthogonal] astrogeodetic deflections. The higher order terms in Eqs. (9)  
 143 and (10) represent latitudinal and longitudinal changes in the differences, thus

144 permitting medium-wavelength errors in the gravimetric geoid model to be controlled  
 145 by the approach proposed.

146

#### 147 **4. Data**

148 1080 astrogeodetic deflections (Fig. 1) were compiled from data held by Geoscience  
 149 Australia and Landgate (the Western Australian geodetic agency). Most of these  
 150 historical data were observed over 40 years ago so as to provide azimuth control on the  
 151 long-line traverses for the Australian Geodetic Datum 1966 (Bomford 1967); also see  
 152 Featherstone (2006) and Featherstone and Morgan (2007). No digital zenith camera  
 153 observations are yet available in Australia.



154

155 **Fig 1.** Coverage of the 1080 astrogeodetic vertical deflections (triangles) over Australia

156

[Lambert projection]

157

158 The accuracy of the Australian astrogeodetic deflections is very difficult to ascertain  
 159 because original records appear to be unavailable. Given the era of the observations, the  
 160 main limiting factors are precise timing and the accuracy of the star catalogues then  
 161 available, which will be substantiated later in Fig 2 by a larger spread in the E-W  
 162 deflections. Using crude hand-waving arguments, as well as comparisons with  
 163 AUSGeoid98, the accuracy of these astrogeodetic deflections is cautiously estimated to  
 164 be one arc-second (Featherstone and R ueger 1999; Featherstone 2006; Featherstone and  
 165 Morgan 2007); also see Kearsley (1976). The geodetic coordinates are on the  
 166 Geocentric Datum of Australia 1994, thus yielding absolute Helmert deflections (Eqs 1  
 167 and 2).

168

	All 1080 stations		After removal of 39 outliers	
	N-S ( $\delta\xi$ )	E-W ( $\delta\eta$ )	N-S ( $\delta\xi$ )	E-W ( $\delta\eta$ )
Max	17.83	9.11	2.92	3.00
Min	-7.76	-12.65	-3.36	-3.62
Mean	-0.25	-0.17	-0.25	-0.16
STD	$\pm 1.28$	$\pm 1.36$	$\pm 0.80$	$\pm 1.05$

169

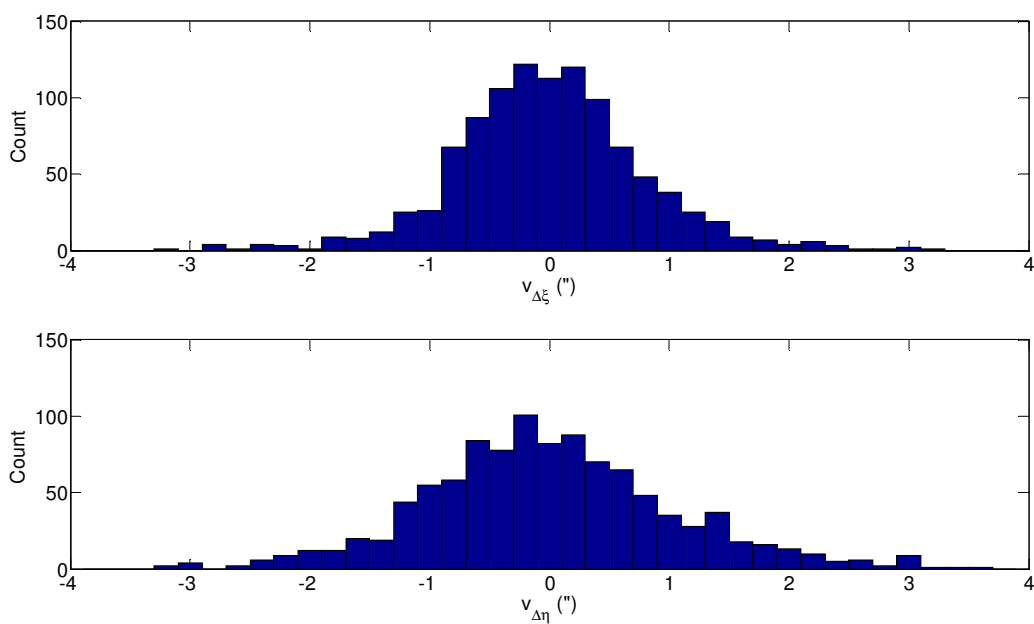
170 **Table 1.** Statistics (in arc-seconds) of the difference between AUSGeoid98-derived and  
 171 astrogeodetic deflections. Outlier detection used Baarda's (1968) data-snooping technique.

172

173 The Pizzetti vertical deflections were derived from AUSGeoid98 (Featherstone et al.  
 174 2001) using Eqs. (3) to (6) for GRS80. The accuracy of these deflections is also  
 175 difficult to ascertain, but they are also cautiously estimated to be around one arc-second  
 176 (Featherstone 2006; Featherstone and Morgan 2007). However, this becomes  
 177 immaterial if the astrogeodetic vertical deflections are to be used as control. The  
 178 AUSGeoid98-derived deflections were bi-cubically interpolated from a pre-computed  
 179 grid (Featherstone 2001), then subtracted from the astrogeodetic deflections. Bi-cubic



180 interpolation proved to be better than bi-linear interpolation, which is consistent with  
 181 expectation because vertical deflections contain more power in the high frequencies.  
 182 The statistics of these differences are in Table 1, before and after rejection of 39 outliers  
 183 that were identified with Baarda's (1968) data-snooping test at 99.9% confidence (cf.  
 184 Kuang 1996). Descriptive statistics are acceptable metrics because the differences are  
 185 reasonably normally distributed (Fig. 2).



186  
 187 **Fig 2.** Histograms (in arc-seconds) of the difference between AUSGeoid98-derived and  
 188 astrogeodetic deflections (top: N-S; bottom: E-W). The larger spread in the E-W deflection  
 189 differences probably reflects the poorer astrogeodetic measurements due to timing and star-  
 190 catalogue errors in these historical data.

191

## 192 5. Results

193 Equations (9) and (10) were applied to the differences between the AUSGeoid98-  
 194 derived and astrogeodetic deflections, but in stages to determine the relative statistical

195 significance of each of the parameters. This involved a two-, three- and four-parameter  
196 model variants of Eqs. (9) and (10) for each deflection component (Sect. 5.1).

197 Standard parametric least-squares was used to estimate the parameters in each case  
198 with the stochastic models  $C_{\delta\xi} = \sigma_{\delta\xi}^2 I$  and  $C_{\delta\eta} = \sigma_{\delta\eta}^2 I$ , where  $\sigma_{\delta\xi} = \sigma_{\delta\eta} = \pm 1''$  based on  
199 the earlier crude estimate of the accuracy of the astrogeodetic deflection data. All data  
200 were first reduced to their 2D centroid (i.e., mean  $\phi$  and mean  $\lambda$  of the stations in Fig. 1)  
201 to improve the conditioning of the normal equation matrices.

202

### 203 5.1 Adjustment cases

204 In the first case tested, Eqs. (9) and (10) reduce to

$$205 \quad \delta\xi = a_{00} + a_{10}\phi \quad (11)$$

$$206 \quad \delta\eta = b_{00} + b_{01}\lambda \quad (12)$$

207 while for the second case, they reduce to

$$208 \quad \delta\xi = a_{00} + a_{01}\lambda \quad (13)$$

$$209 \quad \delta\eta = b_{00} + b_{10}\phi \quad (14)$$

210 For the three-parameter model, Eqs. (9) and (10) reduce to

$$211 \quad \delta\xi = a_{00} + a_{10}\phi + a_{01}\lambda \quad (15)$$

$$212 \quad \delta\eta = b_{00} + b_{10}\phi + b_{01}\lambda \quad (16)$$

213 The least-squares parameter estimates, without the 39 outliers, from these cases (Eqs.  
214 11 to 16) as well as the four-parameter model (Eqs. 9 and 10) are given in Table 2.  
215 Only significant parameters are reported. Significance was evaluated by testing the  
216 ratio of the parameter estimate and its estimated standard deviation at 95% confidence  
217 for which the critical value was taken from the Gaussian distribution tables due to the

218 high redundancy of the fitting and the distribution of the deflection differences (Fig 2).  
 219 Individual testing of terms is valid due to the low correlation among parameters: the  
 220 largest correlation coefficient magnitude was 0.29 from the four-term model. The  
 221 statistics of the post-fit residuals are in Table 3 (cf. Table 1).

222 Table 2 shows that in both two-parameter cases, only the bias term is significant  
 223 in the N-S deflection differences, while the bias and linear term are both significant in  
 224 the E-W deflection differences (discussed later in Sect 5.2). The significant terms in the  
 225 three-parameter model are the same as for the two-parameter models. The additional  
 226 longitudinal parameters ( $a_{01}$  and  $b_{01}$ ) are insignificant, which is also reflected in the  
 227 post-fit residuals, where the values are very similar (Table 3). The additional  
 228 parameterisation is not warranted here, mostly because of the data quality (discussed  
 229 later in Sect 5.2). In the four-parameter case, the significance of the parameters is  
 230 consistent with the two- and three-parameter models, with the exception of the latitude-  
 231 longitude cross term ( $b_{11}$ ) for the E-W vertical deflection difference.

232

deflection	parameter	2-term model	2-term model	3-term model	4-term model
		Eqs. (11, 12)	Eqs. (13, 14)	Eqs. (15, 16)	Eqs. (9, 10)
N-S ( $\delta\xi$ )	$a_{00}$ (")	$-0.245\pm 0.031$	$-0.245\pm 0.031$	$-0.245\pm 0.031$	$-0.249\pm 0.031$
	$a_{10}$ ("/rad)	--	n/a	--	--
	$a_{01}$ ("/rad)	n/a	--	--	--
	$a_{11}$ ("/rad <sup>2</sup> )	n/a	n/a	n/a	--
E-W ( $\delta\eta$ )	$b_{00}$ (")	$-0.161\pm 0.031$	$-0.161\pm 0.031$	$-0.161\pm 0.031$	$-0.173\pm 0.031$
	$b_{10}$ ("/rad)	n/a	$-1.214\pm 0.274$	$-1.158\pm 0.275$	$-0.879\pm 0.289$
	$b_{01}$ ("/rad)	$0.381\pm 0.159$	n/a	--	--
	$b_{11}$ ("/rad <sup>2</sup> )	n/a	n/a	n/a	$-5.181\pm 1.596$

233

234 **Table 2.** Summary of the significant parameter estimates for the two-, three- and four-parameter  
 235 deflection fitting models (n/a = not applicable; -- = insignificant)

236

237

	2-term model		2-term model		3-term model		4-term model	
	Eqs. (11, 12)		Eqs. (13, 14)		Eqs. (15, 16)		Eqs. (9, 10)	
	N-S ( $\delta\zeta$ )	E-W ( $\delta\eta$ )	N-S ( $\delta\zeta$ )	E-W ( $\delta\eta$ )	N-S ( $\delta\zeta$ )	E-W ( $\delta\eta$ )	N-S ( $\delta\zeta$ )	E-W ( $\delta\eta$ )
Max	3.11	3.52	3.11	3.29	3.10	3.20	3.10	3.14
Min	-3.16	-3.16	-3.19	-3.22	-3.18	-3.26	-3.14	-3.17
STD	$\pm 0.80$	$\pm 1.04$	$\pm 0.80$	$\pm 1.04$	$\pm 0.80$	$\pm 1.04$	$\pm 0.79$	$\pm 1.03$

238

239 **Table 3.** Residual statistics for the two-, three- and four-parameter deflection model fits (in arc-  
240 seconds) after rejection of 39 outliers

241

## 242 5.2 Deflection-derived geoid corrections and discussion

243 Only the statistically significant parameter estimates in Table 2 will be used to attempt  
244 to apply ‘corrections’ to the gravimetric model. For the N-S deflection differences, only  
245 the first term ( $a_{00}$ ) is significant for all parameterisations tested, which consistently  
246 shows an N-S-oriented misalignment of  $\sim -0.25$  arc-seconds between the astrogeodetic  
247 and geoid-derived deflections. For the E-W deflection differences, the first term ( $b_{00}$ ) is  
248 also significant for all parameterisations, showing an E-W-oriented misalignment of  $\sim$   
249 0.16 arc-seconds.

250 The first of the two-parameter models for the E-W deflection differences shows a  
251 significant longitudinal term ( $b_{01}$ ), but which is not significant in the three- and four-  
252 parameter models (Table 2). This is explained when seeing that the latitudinal term  
253 ( $b_{10}$ ) is significant in the other two-parameter model, as well as in the three- and four-  
254 parameter models, and a significant latitude-longitude term ( $b_{11}$ ) occurs in the four-  
255 parameter model. Therefore, the longitudinal term in the two-parameter model is  
256 actually a part of the latitude-longitude dependency ( $b_{11}$ ) that becomes evident in the  
257 four-parameter model for the E-W deflection difference.

258 We now use these parameter estimates to apply ‘corrections’ to the gravimetric geoid  
259 model, akin to the use of GPS-levelling. The first terms ( $a_{00}$  and  $b_{00}$ ) are

260 straightforward to apply; they represent N-S and E-W tilts that should be applied to the  
261 gravimetric geoid model. Applying the estimated  $a_{00}$  and  $b_{00}$  terms over the data ranges  
262 of  $\Delta\phi=0.5948\text{rad}$  ( $34.0810^\circ$  or  $\sim 3783\text{km}$ ) and  $\Delta\lambda=0.7059\text{rad}$  ( $40.4449^\circ$  or  $\sim 4489\text{km}$ )  
263 gives a N-S tilt of  $-(4.49\pm 0.02)\text{m}$  and an E-W tilt of  $-(3.50\pm 0.02)\text{m}$ .

264 These values are much larger than could realistically be expected. For instance,  
265 comparisons of AUSGeoid98 with GPS-levelling data do not show such large tilts (e.g.,  
266 Featherstone et al. 2001; Featherstone and Sproule 2006; Soltanpour et al. 2006),  
267 especially not in the E-W direction, though there is evidence for a  $\sim -1\text{--}2\text{ m}$  N-S-  
268 oriented tilt (using the same sign convention) in the Australian Height Datum (e.g.,  
269 Featherstone 2004; 2006a). This exemplifies the problem of the inseparability when  
270 using GPS-levelling data. The only plausible reason for these unrealistically large N-S  
271 and E-W tilts comes from the poor quality of the historic astrogeodetic deflections over  
272 Australia.

273 Recall that their accuracy was estimated to be one arc-second, which is substantially  
274 larger than the parameter estimates summarised in Table 10. Applying this one arc-  
275 second uncertainty over the N-S and E-W data ranges, gives uncertainties in the tilts of  
276  $\pm 18.34\text{m}$  and  $\pm 21.55\text{m}$  respectively. Accordingly, the above-estimated tilts of  $-4.49\text{m}$   
277 and  $-3.50\text{m}$  are statistically insignificant when considering the quality of these  
278 historical deflection data. Therefore, very accurately known astrogeodetic deflections  
279 would be needed to utilise this method over a very large area like Australia. However,  
280 this accuracy requirement will be lessened over a smaller area, so may be attractive in  
281 geographically smaller countries.

282

283

284

285 **6. Summary and Conclusion**

286 We have presented an alternative and new method with which to control gravimetric  
287 geoid model errors using astrogeodetic deflections of the vertical. This is a preferable  
288 alternative to the current widespread use of GPS-levelling data, which suffers from the  
289 inseparability of height-related errors in that data combination strategy. Two-, three-  
290 and four-parameter functional models have been formulated here, but other  
291 parameterisations are possible, as has been the case for the GPS-levelling combination  
292 strategy. These are left for future work.

293 Numerical experiments with 1080 historical astrogeodetic deflections over Australia  
294 and AUSGeoid98 show that the approach presented is indeed feasible, but the poor  
295 quality of the astrogeodetic deflections, coupled with the size of the study area, causes  
296 unrealistically large values for the deflection-derived geoid corrections. However, using  
297 modern digital zenith cameras would provide much better results.

298

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303 publication number **xx**.

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