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1	Reduction Spheroids Preserve a Uranium Isotope Record of the Ancient Deep
2	Continental Biosphere
3	
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ZZ	Life on Earth extends to several kilometres below the land surface and seafloor.
23	This deep biosphere is second only to plants in its total biomass, is metabolically
24	active and diverse, and is likely to have played critical roles over geological time
25	in the evolution of microbial diversity, diagenetic processes, and biogeochemical
26	cycles. However, these roles are obscured by a paucity of fossil and geochemical
27	evidence. Here, we apply the recently developed uranium-isotope proxy for
28	biological uranium reduction to reduction spheroids in continental rocks (red
29	beds). Although these common palaeo-redox features have previously been
30	suggested to reflect deep bacterial activity, unequivocal evidence for biogenicity
31	has been lacking. Our analyses reveal that the uranium present in reduction
-	g , o
32	spheroids is isotopically heavy, which is most parsimoniously explained as a
33	signal of ancient hacterial uranium reduction revealing a compelling record of
55	Signar of ancient bacteriar aramam reduction, revealing a compening record of
31	Farth's deen higsnhere
54	Darm suce prospinere.

36 The subsurface represents a vast habitat containing up to a fifth of Earth's current total biomass including diverse bacteria, archaea and fungi^{1,2}. In recent decades, 37 accelerating exploration of the deep biosphere has revealed new microbial groups, 38 new ecological niches, and new modes of microbe-mineral interaction^{3,4,5}. However, 39 40 a full understanding of the limits of the deep biosphere, its contribution to total 41 planetary biomass, and its biogeochemical significance requires the detection and 42 analysis of robust deep biosignatures in the rock record. Such biosignatures could also 43 refine search images for past or present life on Mars, whose surface has long been 44 uninhabitable but may conceal a warm, wet interior⁶.

45

46 It is difficult to confirm that possible traces of the ancient deep biosphere are truly 47 biogenic, post-burial in origin, and pre-modern. Few reported cellular or molecular 48 fossils pass all three tests unequivocally, and most candidates represent subseafloor environments^{7,8,9,10}. Microbially mediated diagenetic phenomena on palaeo-redox 49 50 fronts offer a potentially powerful alternative record that could extend to continental 51 settings^{11,12}. Of longstanding interest in this connection are reduction spheroids, 52 which are very common mm-dm-scale bleached spots found most commonly (but not 53 exclusively) in Proterozoic-Phanerozoic red beds, i.e., sedimentary rocks deposited in oxidizing terrestrial environments and rich in early diagenetic hematite 13,14 . The 54 bleached colour reflects localised Fe(III) reduction and loss¹⁵. Many examples contain 55 56 small, dark, central "cores" where redox-sensitive elements including uranium, 57 vanadium, and nickel are highly concentrated, and organic matter may also be present^{16,17}. 58

59

60 The mechanisms that produce reduction spheroids in the subsurface have hitherto 61 been unclear. They are sometimes attributed to the oxidation of organic-rich cores, 62 but the cores are usually darkened not by organic matter but by opaque metalliferous 63 minerals. Although organic carbon is present in some reduction spheroid cores and 64 may have stimulated their formation, organic carbon present prior to reduction is typically too scarce to have reduced the surrounding halos^{16,17}. It has thus been 65 66 proposed instead that spheroids form around localized chemolithotrophic microbial 67 populations, which would catalyse the oxidation of mobile reductants supplied through groundwater^{13,18,19,20}. These reductants could include H_2 derived 68 69 radiolytically from porewater, which could establish a positive-feedback mechanism for spheroid growth following initial uranium precipitation²⁰. Confirmation of a biotic 70 71 mode of origin would distinguish these common geological features as perhaps the 72 most widely accessible, recognisable and distinctive traces of the ancient deep 73 biosphere. It would also add weight to previous suggestions that reduction spheroids 74 could be a target for astrobiological sampling on Mars, where iron reduction is regarded as a plausible metabolic strategy for past or present life^{21,22,23}. 75 76 77 Direct evidence for the biogenicity of reduction spheroids has hitherto been lacking or

equivocal. Authigenic pyrite present in some spheroids has a sulphur isotope

79 composition consistent with but not diagnostic of a bacterial $\operatorname{origin}^{21}$; however, most

80 spheroids lack pyrite altogether¹⁹. Molecular biomarkers can be extracted from the

81 organic matter commonly associated with reduction spheroids (e.g., Supplementary

82 Figure 1), but are likely to pre-date the origin of the spheroids themselves, so cannot

83 shed light on their biogenicity. This organic matter is also damaged and isotopically

84 modified by exposure to ionising radiation commonly emitted by uranium in

reduction spheroids^{17,19,24}. Hence, determination of reduction spheroid biogenicity is
non-trivial and necessitates the analysis of authigenic phases demonstrably associated
with spheroid formation.

88

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89 Here, we focus on a new low-temperature palaeo-redox proxy, the isotopic composition of uranium (238 U/ 235 U: δ^{238} U, in standard delta notation, relative to the 90 91 CRM-112a standard; Equation 1). Uranium enrichment appears to be a universal 92 feature of reduction spheroids, occurring both in the cores and in the halos as a result 93 of the highly localised reduction of soluble U(VI) to insoluble U(IV) (ref. 19). Thus, 94 uranium phases (both mineralized and non-mineralized) in reduction spheroids and analogous low-temperature redox-front uranium deposits have been shown^{12,19} to 95 contain predominantly U(IV). Uranium reduction can occur via many pathways^{18,25}. 96 97 both abiotic (coupled to the oxidation of various aqueous, mineral, and organic 98 species) and biotic (i.e., enzymatic catalysis by chemolithotrophic microorganisms 99 capable of facultatively utilising U(VI) as an electron acceptor, including iron- and 100 sulphate-reducers). The uranium isotope system is controlled by low-temperature 101 redox reactions that significantly fractionate the uranium isotope composition preserved in environmental samples away from a crustal (high-T) average δ^{238} U of – 102 0.29 ± 0.03 %, often concentrating the heavier isotope in the reduced product^{26,27}. 103 104 105 Several experimental studies have shown that bacterially reduced and precipitated

107 phase 28,29,30 . Bhattacharyya et al. (ref. 12) recently determined a bacterial origin for

uranium is isotopically heavier (i.e. records higher δ^{238} U) than the dissolved precursor

108 isotopically heavy authigenic uranium phases in roll front ore deposits, which

109 resemble reduction spheroids inasmuch as they are mineralised paleo-redox fronts

110 formed at low temperatures in subsurface aquifers. Field studies confirm that modern 111 groundwaters inoculated with metal reducing bacteria become isotopically lighter in uranium as the heavier isotope is preferentially precipitated^{31,32}. Experimental studies 112 113 so far have shown that, by contrast, abiotically reduced uranium either remains 114 unfractionated or is isotopically lighter, regardless of the reductant responsible^{29,30,33,34}. Consequently, the U-isotopic composition (δ^{238} U) of reduced 115 116 uranium phases in nature is emerging as a new and potentially powerful proxy for their mode of origin^{12,30,31} (Supplementary Note 1). 117 118 Here, we report δ^{238} U analyses of the dark cores, bleached halos, and surrounding 119 120 matrix of reduction spheroids collected from continental red beds in outcrop, 121 primarily at Dingwall in northern Scotland and Budleigh Salterton in southwest 122 England, sites where spheroids are both especially uraniferous, and can be linked to 123 unusually well constrained formation depths from their geological context. Spheroids 124 from other localities of diverse ages were analysed for comparison, as were

125 uraniferous hydrothermal veins expected to yield near-crustal δ^{238} U values reflecting

126 their high-temperature origin 30,34 . We find that reduction spheroids are enriched

127 towards their cores in uranium characterised by high δ^{238} U values. This result is most

128 parsimoniously explained as a signal of ancient bacterial U(VI) reduction, implying

that the spheroids themselves are most likely bacterial in origin.

130

131 **Results and Discussion**

132 Uranium isotope values

133 The cores of reduction spheroids have uniformly higher uranium concentrations and

134 heavier uranium isotope compositions (δ^{238} U) compared to the host rock in all

135	samples (Supplementary Table 1; Figure 1). All hydrothermal vein samples and
136	most of the red-bed matrix from the reduction-spheroid localities yielded δ^{238} U values
137	near the average crustal value of $-0.29 \pm 0.03\%$ (ref. 27). In most reduction
138	spheroids, both uranium concentration and δ^{238} U increased from the matrix through
139	the halo into the core. At Budleigh Salterton, the large size of the spheroids made it
140	possible to discriminate between isotopically heavier black inner cores (mean
141	+0.78‰; $n = 4$) and isotopically less heavy dark grey core margins (+0.07‰; $n = 2$),
142	as well as greenish outer halos (-0.24‰; n=4). Similarly, the spheroid cores from
143	Dingwall yielded much heavier values of δ^{238} U (mean +0.45‰; n=6) than the halos
144	(+0.04%; n = 4), the matrix $(-0.21%; n = 2)$, and a nearby bitumen vein $(-0.16%; n = 2)$
145	1). Spheroid cores from the other localities, where palaeodepth was less well
146	constrained, also recorded values heavier than the crustal range, and all were heavier
147	than their respective matrices by at least 0.10‰ (Supplementary Table 1).

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149 *Reduction spheroid biogenicity*

150 The reduction spheroids analysed here are enriched in uranium and show increasingly heavy isotopic compositions (δ^{238} U) towards their reduced cores. In some cases, the 151 152 matrix to core interval of reduction spheroids expresses U isotope variation 153 approximating almost the entire natural range of low-temperature systems on Earth^{26,35}. Field-based, experimental and geological studies to-date strongly suggest 154 that these high δ^{238} U values are best explained by bacterial uranium reduction and 155 precipitation within the spheroids^{12,29,30,31,32,33,34}. In particular, given the 156 157 environmental similarity between reduction spheroids and roll-front ore deposits, our 158 interpretation receives support from the recent measurement of isotopically heavy biogenic uranium phases associated with the latter¹². 159

161 Since Fe(III) and U(VI) reduction are carried out by the same groups of

microorganisms using the same reductants, and occur coextensively and concurrently in modern aquifers³⁶, our results strongly imply that the reduction and dissolution of ferric iron responsible for the presence of the bleached spheroids themselves was also bacterially mediated. We infer that reduction spheroids, which are both spatially and temporally widespread, represent an important record of the geological history of the deep biosphere, which was potentially Earth's largest reservoir of biomass prior to the proliferation of land plants³⁷.

169

170 A record of the ancient deep biosphere

171 The deep biosphere conventionally extends from ~metres depth to several kilometres^{4,38,39}. There is clear evidence that many—perhaps most—reduction 172 173 spheroids form at the deeper end of this range. In brief: first, halos are commonly 174 spherical, whereas shallow non-nodular features would be flattened by compaction; 175 second, radiometric ages of authigenic minerals concentrated within some spheroids are $>10^7$ years younger than the host rock^{16,40}; third, some spheroids occur in 176 177 hematite-stained igneous basement, hundreds of metres below the uppermost basement¹⁴: fourth, at many localities, the distribution of spheroids was clearly 178 179 influenced by pre-existing faults, fractures, cataclastic zones and cleavages younger than the host rock^{14,41,42}. Our findings evince a clear signal of bacterial uranium 180 181 reduction in spheroids demonstrably formed at multi-km depth, including one locality 182 (Dingwall) where they appear to be related to the early-stage biodegradation of hydrocarbons, and a weaker but consistent signal at all other localities. We conclude 183 184 that reduction spheroids represent an important and widespread archive of the deep

- 185 continental biosphere, present through much of Earth's geological record. This
- 186 finding lends weight to the suggestion that reduction spheroids be targeted for
- 187 analysis and sample return were they to be discovered on $Mars^{22}$.
- 188
- 189
- 190 Methods
- 191 Sample localities

192The Dingwall spheroids are hosted by red mudstones of the middle Devonian

193 Millbuie Sandstone Group, which forms part of a thick continental succession (the

- 194 Old Red Sandstone). As previously described by ref. 24, the cores are black, spherical
- nodules a few mm across, composed of solid hydrocarbons with uranium present as

196 microscopic inclusions of uraninite and other minerals (impure xenotime and possibly

- 197 brannerite). These cores occupy green-grey non-nodular halos that can extend for
- several centimetres, and occur through a stratigraphic thickness of 10 m (Figure 2).
- 199 The post-compaction origin of the Dingwall spheroids is confirmed by (1) the

spherical shape of the cores; (2) the lack of compaction drapes over them; and, (3) the

201 presence of solid hydrocarbon residues (bitumen) within the cores that clearly derive

202 from source rocks in the underlying kilometre of stratigraphy 24 , which must have been

203 deeply buried in order to reach thermal maturity and generate hydrocarbons (i.e.,

about 3 km assuming a normal geothermal gradient). Migration occurred while the

succession was still deeply buried, as demonstrated by the presence of bitumen-

206 bearing quartz veins through a stratigraphic thickness of ~10 m in the conglomerate

- 207 directly underlying the spheroid-hosting mudstone²⁴. These veins, which yield fluid-
- 208 inclusion temperatures close to 100 °C (ref. 43)— equivalent to ~3–4 km depth
- assuming a normal geothermal gradient—occur within 5 m of the spheroids

themselves. We therefore infer that the Dingwall spheroids formed at depths of
several kilometres and may be genetically related to hydrocarbon migration and
(bio)degradation.

213

214 The reduction spheroids at Budleigh Salterton are hosted by red mudstone in the latest 215 Permian Littleham Mudstone Formation, a ~200-m thick unit within the New Red 216 Sandstone Supergroup, a laterally and stratigraphically extensive Permo-Triassic continental succession⁴⁴. The specimens analysed here are spherical pale green 217 218 nodules with diffuse black centres; uranium is present as fine-grained coffinite and is not associated with organic matter^{45,46}. Cross-cutting relationships described by ref. 219 220 38 show that these spheroids formed penecontemporaneously with sheet-like copper 221 nodules, which themselves replaced an earlier generation of crack-seal calcite veins 222 generated by overpressure during compaction dewatering. These relationships suggest 223 the Budleigh Salterton reduction spheroids formed relatively early, at depths of up to 224 around 1-2 km (ref. 46).

225

226 In addition to Budleigh Salterton and Dingwall, reduction spheroids were collected 227 from the field at four other localities: (1) the Mesoproterozoic (~1.4 Ga) red beds of the Sibley Group of Ontario, sampled at a road cut near Nipigon⁴⁷; (2) Devonian red 228 sandstone at Millport, Great Cumbrae, Scotland; (3) Carboniferous white sandstone at 229 Heysham, Lancaster, England²⁴; (4) Triassic red siltstone at Hartlepool, Co. Durham, 230 231 England. The sample from Ord Burn was collected from a hydrothermal vein in Caledonian granite, Sutherlandshire, Scotland⁴⁸. The sample from Great Orme's Head 232 is a copper ore deposit in Carboniferous limestone, North Wales⁴⁹. The sample from 233

Laxey is a hydrothermal vein-hosted lead-zinc ore deposit in Lower Palaeozoic slates
above Caledonian granite, Isle of Man⁵⁰.

236

237 Sample preparation

238 Samples were cut, cleaned and crushed in an agate mill, or micro-drilled with a

tungsten carbide drill bit (previously tested to not contaminate U isotope analysis, and

240 cleaned between samples) to target the specific components of reduction spheroids,

host rocks and other samples (Figure 3). Approximately 0.1-0.4 g of each powdered

and homogenised sample were ashed in a 100 °C oven for 24 hours. Samples were

243 digested in a 3:1 mixture of concentrated HNO₃ and HF on a hotplate for 24 hours.

244 Samples were dried and re-digested in concentrated HCl and HNO₃.

245

246 Isotope analyses

247 Trace metal concentrations were measured at the Yale Metal Geochemistry Center on

a Themo-Finnigan Element XR ICP-MS on splits from each digest. The ²³⁶U-²³³U

249 double spike was added based on uranium concentrations $(^{238}U/^{236}U \sim 30)$, prior U

250 purification via ion exchange methods. The spiked samples were dried and taken up

in 3N HNO₃. The U was then purified using the UTEVA column chemistry method

252 (after ref. 51; see methods of ref. 39, 52). Purified U was dissolved in 0.75 N HNO₃

with 50 ppb concentration. Uranium isotopes were measured at the Yale Metal

254 Geochemistry Center on a Themo-Finnigan Neptune Plus Multi-Collector ICP-MS at

low mass resolution using a Jet sampler cone and a standard skimmer cone. Sample

256 were introduced through an Elemental Scientific μ Flow PFA nebulizer at ~50 μ L/min

via an Elemental Scientific Apex IR. A 50 ppb sample solution yielded 32–40 volts of

258 238 U signal on a $10^{11} \Omega$ amplifier.



261 monitored to have a negligible effect on measurement of 233 U. Measurements

262 consisted of five blocks, each block 10 cycles, each cycle 4.19 s. Blank U level was

less than 50 pg. External reproducibility was assessed using full protocol duplicates of

the geostandard NOD-A-1, which yielded an average δ 238U of -0.52 +/-0.08‰ based

on nine repeats (2σ error = 0.12). Duplicate samples agreed within error.

266

267 Uranium isotope variations of samples and standards are reported as

268 δ^{238} U_{CRM 112a}, which is defined as:

269
$$\delta^{238} U = \left(\left[\frac{\left(\frac{2^{38} U}{2^{35} U}\right)_{sample}}{\left(\frac{2^{238} U}{2^{35} U}\right)_{CRM-112a}} \right] - 1 \right) \times 1000 \% (Equation 1)$$

270

271 Biomarker analyses

272 Quantitative biomarker data (Supplementary Note 2) were obtained by gas 273 chromatography-mass spectrometry (GC-MS) from the spheroids from Dingwall as 274 follows. Core samples were prepared by rinsing twice with distilled water and again 275 with dichloromethane (DCM), and ultrasonicated with DCM and methanol. All 276 glassware was thoroughly cleaned with a 93:7 mixture of DCM/MeOH. Crushed 277 samples were weighed, recorded, transferred into pre-extracted thimbles, dried with a 278 rotary evaporator, and separated into aliphatic, aromatic and polar fractions via silica 279 column chromatography using hexane, hexane/DCM in the ratio 3:1 and DCM/MeOH 280 respectively. Prior to GC-MS analysis, an internal standard (5β-Cholane, Agilent 281 Technologies) was added to the saturated fraction before injection into the GC-MS 282 machine, and subsequent biomarker identification. This was done using an Agilent

283	6890N gas chrom	atograph fitted	with a J&W DB	-5 phase 50 m MS	SD and a quadruple
	0				

mass spectrometer operating in SIM mode (dwell time 0.1 s per ion and ionisation

energy 70eV). Samples were injected manually using a split/splitless injector

- 286 operating in splitless mode (purge 40 ml min-1 for 2 min). The temperature
- programme for the GC oven was 80–295 °C, holding at 80 °C for 2 min, rising to 10
- ²⁸⁸ °C min⁻¹ for 8 min and then 3 °C min⁻¹, and finally holding the maximum temperature
- for 10 min. Data were obtained by comparing with the response of the internal
- standard.
- 291

292 Data availability statement

All data generated during and/or analysed during this study are available from thecorresponding author upon reasonable request.

295

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307

308 Author contributions

309 SM conceived the study and wrote the manuscript. SM, JP and AvSH prepared the

310 samples. AvSH conducted the isotopic analyses and wrote the methods section. JP

- 311 contributed the samples and supporting locality information. SB conducted the GC-
- 312 MS analysis. All authors contributed to the design of the study and edited the
- 313 manuscript.

314

315 Competing interests statement

316 The Authors declare no competing interests.

317

318

319 **References**

- 321 1. Colwell, F.S. & D'Hondt, S. Nature and extent of the deep biosphere. *Reviews in Mineralogy and Geochemistry* 75, 547-574 (2013).
- 323 2. McMahon, S. & Parnell, J. Weighing the deep continental biosphere. *FEMS*324 *Microbiology Ecology* 87,113-120 (2014).
- 325 3. Anantharaman, K., et al. Thousands of microbial genomes shed light oninterconnected biogeochemical processes in an aquifer system. *Nature*
- 326 interconnected biogeochemical processes in an aquifer system. *Natur* 327 *Communications* 7, 13219 (2016).
- 328 4. Moser, D.P., et al. Desulfotomaculum and Methanobacterium spp. dominate a 4-to 5kilometer-deep fault. *Applied and Environmental Microbiology* 71, 8773-8783
 (2005).
- 331 5. Edwards, K.J., Becker, K. & Colwell, F. The deep, dark energy biosphere:
 intraterrestrial life on earth. *Annual Review of Earth and Planetary Sciences* 40, 551-568 (2012).
- 334 6. Michalski, J.R., et al. Groundwater activity on Mars and implications for a deepbiosphere. *Nature Geoscience* 6, 133 (2013).
- 336 7. Hofmann, B.A. & Farmer, J.D. Filamentous fabrics in low-temperature mineralassemblages: are they fossil biomarkers? Implications for the search for a
- subsurface fossil record on the early Earth and Mars. *Planetary and Space Science*48, 1077-1086 (2000).
- 340 8. Contreras, S., et al. Cyclic 100-ka (glacial-interglacial) migration of subseafloor redox
 341 zonation on the Peruvian shelf. *Proceedings of the National Academy of Sciences*342 110, 18098-18103 (2013).
- 343 9. Klein, F., Humphris, S.E., Guo, W., Schubotz, F., Schwarzenbach, E.M. & Orsi, W.D.,
- 344 Fluid mixing and the deep biosphere of a fossil Lost City-type hydrothermal
- 345 system at the Iberia Margin. Proceedings of the National Academy of
- *Sciences* **112**, 12036-12041 (2015).
- 347 10. Bengtson, S., et al. Fungus-like mycelial fossils in 2.4-billion-year-old vesicular
 basalt. *Nature Ecology & Evolution* 1, 0141 (2017).
- 349 11. Meister, P. For the deep biosphere, the present is not always the key to the past: whatwe can learn from the geological record. *Terra Nova* 27,400-408 (2015).

- 351 12. Bhattacharyya, A., Campbell, K.M., Kelly, S.D., Roebbert, Y., Weyer, S., Bernier-
- Latmani, R. & Borch, T., Biogenic non-crystalline U (IV) revealed as major component in uranium ore deposits. *Nature Communications* **8**, 15538 (2017).
- 354 13. Hofmann, B.A. Reduction spheroids. In *Encyclopedia of Geobiology* (pp. 761-762).
 Springer Netherlands (2011).
- 356 14. Parnell, J., Brolly, C., Spinks, S. & Bowden, S. Metalliferous biosignatures for deep
 subsurface microbial activity. *Origins of Life and Evolution of Biospheres* 46,
 107-118 (2016).
- 359 15. Turner, P. Continental Red Beds (Elsevier, Amsterdam, 1980).
- 360 16. Hofmann, B.A. Reduction spheroids from northern Switzerland: mineralogy,
 geochemistry and genetic models. *Chemical Geology* 81, 55-81 (1990).
- 362 17. Hofmann, B.A. Organic matter associated with mineralized reduction spots in red
- beds. In: J. Parnell, H. Kucha, P. Landais (eds) Bitumens in ore deposits
 (Springer), 362-378 (1993).
- 365 18. Lovley, D.R., Phillips, E.J., Gorby, Y.A. & Landa, E.R. Microbial reduction of uranium. *Nature* **350**, 413-416 (1991).
- 367 19. Hofmann, B.A. Mineralogy and geochemistry of reduction spheroids in red
 beds. *Mineralogy and Petrology* 44, 107-124 (1991).
- 369 20. Hofmann, B.A. Isolated reduction phenomena in red-beds: a result of porewater
 radiolysis. In Proceedings of the 7th International Symposium on Water–Rock
 Interaction (503-506) (1992).
- 372 21. Spinks, S. C., Parnell, J. & Bowden, S. A. Reduction spots in the Mesoproterozoic
 age: implications for life in the early terrestrial record. *International Journal of Astrobiology* 9, 209-216 (2010).
- 375 22. Mustard, J.F., et al. Appendix to the Report of the Mars 2020 Science Definition
 376 Team, 51 pp., posted July, 2013, by the Mars Exploration Program Analysis
 377 Group (MEPAG) at
- 378 <u>http://mepag.jpl.nasa.gov/reports/MEP/Mars_2020_SDT_Report_Appendix.pdf</u>
 379 (2013).
- 380 23. Hurowitz, J.A., Grotzinger, J.P., Fischer, W.W., McLennan, S.M., Milliken, R.E.,
 381 Stein, N., Vasavada, A.R., Blake, D.F., Dehouck, E., Eigenbrode, J.L. & Fairén,
 382 A.G. Redox stratification of an ancient lake in Gale crater, Mars. *Science* 356,
 383 p.eaah6849 (2017).
- 384 24. Parnell, J. & Eakin, P. The replacement of sandstones by uraniferous hydrocarbons:
 significance for petroleum migration. *Mineralogical Magazine* 51, 505-515
 (1987).
- 387 25. Murphy, M. J., Stirling, C. H., Kaltenbach, A., Turner, S. P. & Schaefer, B. F.
 388 Fractionation of 238 U/235 U by reduction during low temperature uranium
 389 mineralisation processes. *Earth and Planetary Science Letters* 388, 306-317
 390 (2014).
- 391 26. Andersen, M.B., Stirling, C.H. & Weyer, S. Uranium isotope fractionation. *Reviews* 392 *in Mineralogy and Geochemistry* 82, 799-850 (2017).

- 393 27. Tissot, F.L. & Dauphas, N. Uranium isotopic compositions of the crust and ocean: 394 Age corrections, U budget and global extent of modern anoxia. Geochimica et
- 395 Cosmochimica Acta 167, 113-143 (2015).
- 396 28. Basu, A., Sanford, R.A., Johnson, T.M., Lundstrom, C.C. & Löffler, F.E. Uranium
- 397 isotopic fractionation factors during U (VI) reduction by bacterial
- 398 isolates. Geochimica et Cosmochimica Acta 136, 100-113 (2014).
- 399 29. Stirling, C.H., Andersen, M.B., Warthmann, R. & Halliday, A.N. Isotope
- 400 fractionation of 238U and 235U during biologically-mediated uranium reduction. 401 Geochimica et Cosmochimica Acta 163, 200-218 (2015).
- 402 30. Stylo, M., et al. Uranium isotopes fingerprint biotic reduction. *Proceedings of the* National Academy of Sciences 112, 5619-5624 (2015). 403
- 404 31. Bopp, C. J., Lundstrom, C. C., Johnson, T. M. & Glessner, J. J. Variations in 238U/235U in uranium ore deposits: Isotopic signatures of the U reduction 405 process?. Geology 37, 611-614 (2009). 406
- 407 32. Shiel, A.E., Johnson, T.M., Lundstrom, C.C., Laubach, P.G., Long, P.E. & Williams, K.H. Reactive transport of uranium in a groundwater bioreduction study: Insights 408 409 from high-temporal resolution 238 U/235 U data. Geochimica et Cosmochimica 410 Acta 187, 218-236 (2016).
- 411 33. Rademacher, L.K., Lundstrom, C.C., Johnson, T.M., Sanford, R.A., Zhao, J. & 412 Zhang, Z. Experimentally determined uranium isotope fractionation during 413 reduction of hexavalent U by bacteria and zero valent iron. Environmental Science 414 & Technology 40, 6943-6948 (2006).
- 415 34. Stirling, C.H., Andersen, M.B., Potter, E.K. & Halliday, A.N. Low-temperature 416 isotopic fractionation of uranium. Earth and Planetary Science Letters 264, 208-417 225 (2007).
- 418 35. Wang, X., Johnson, T.M. & Lundstrom, C.C. Low temperature equilibrium isotope 419 fractionation and isotope exchange kinetics between U (IV) and U (VI). 420
- Geochimica et Cosmochimica Acta 158, 262-275 (2015).
- 421 36. Campbell, K.M., Kukkadapu, R.K., Qafoku, N.P., Peacock, A.D., Lesher, E.,
- Williams, K.H., Bargar, J.R., Wilkins, M.J., Figueroa, L., Ranville, J. & Davis, 422
- 423 J.A. Geochemical, mineralogical and microbiological characteristics of sediment
- 424 from a naturally reduced zone in a uranium-contaminated aquifer. Applied
- Geochemistry 27, 1499-1511 (2012). 425
- 426 37. McMahon, S. & Parnell, J., The deep history of Earth's biomass. Journal of the Geological Society. Online First: https://doi.org/10.1144/jgs2018-061 (2018). 427
- 428 38. Teske, A. & Sørensen, K. B. Uncultured archaea in deep marine subsurface sediments: have we caught them all? The ISME journal 2, 3-18 (2008). 429
- 430 39. Biddle, J. F., Sylvan, J. B., Brazelton, W. J., Tully, B. J., Edwards, K. T., Mover, C. 431 L., Heidelberg, J. F. & Nelson, W. C. Prospects for the Study of Evolution in the
- 432 Deep Biosphere. Frontiers in Microbiology 2, 285 (2012).
- 433 40. Hofmann, B.A. & Frei, R. Age constraints of reduction spot formation from Permian 434 red beds sediments, northern Switzerland, inferred from U-Th-Pb
- 435 systematics. Schweizerische Mineralogische und Petrographische
- 436 Mitteilungen, 76, 235-244 (1996).

- 437 41. Bateson, J.H. & Johnson, C.C. Reduction and related phenomena in the New Red
- 438 Sandstone of south-west England. British Geological Survey, Technical
 439 report, WP/92/1 (1992).
- 440 42. Nakamura, N. & Borradaile, G. Do reduction spots predate finite strain? A magnetic
 diagnosis of Cambrian slates in North Wales. *Tectonophysics* 340, 133-139
 (2001).
- 44343. Parnell, J. Alteration of crystalline basement rocks by hydrocarbon-bearing fluids: 444 Moinian of Ross-shire, Scotland. *Lithos* **37**, 281–292 (1996).
- 445 44. Harrison, R.K. Concretionary concentrations of the rarer elements in Permo-Triassic
 red beds of south-west England. *Geological Survey of Great Britain Bulletin* 52,126 (1975). Accessed October 2017 at:
 http://pubs.bgs.ac.uk/publications.html?pubID=B04509
- 448 <u>http://pubs.bgs.ac.uk/publications.html?publD=B04509</u>
- 449 45. Kemp, A.J., Palmer, M.R. & Ragnarsdottir, K.V. The uranium-thorium and rare earth 450 element composition of reduction nodules from Budleigh Salterton, Devon.
- 451 Proceedings of the Ussher Society **8**, 214-218 (1994).
- 452 46. Milodowski, A.E., Styles, M.T. & Hards, V.L. A natural analogue for copper waste
 canisters: The copper-uranium mineralised concretions in the Permian mudrocks
 of south Devon, United Kingdom. Swedish Nuclear Fuel and Waste Management
 Co. No. SKB TR. 00, 11 (2000)
- 455 Co. No. SKB-TR--00-11 (2000).
- 456 47. Fralick, P. & Zaniewski, K. Sedimentology of a wet, pre-vegetation floodplain
 assemblage. *Sedimentology* 59, 1030-1049 (2012).
- 458 48. Tweedie, J.R. Origin of uranium and other metal enrichments in the Helmsdale
 Granite, eastern Sutherland, Scotland. *Transactions of the Institution of Mining and Metallurgy* 88, B145-B153 (1979).
- 461 49. Parnell, J. Mineralogy of uraniferous hydrocarbons in Carboniferous-hosted mineral
 deposits, Great Britain. *Uranium* 5, 197-218 (1988).
- 463 50. Davidson C.F. & Bowie S.H.U. On thucholite and related hydrocarbon-uraninite
 complexes with a note on the origin of the Witwatersrand gold ores. *Bulletin of the Geological Survey of Great Britain* 3, 1–19 (1951).
- 466 51. Weyer, S., Anbar, A.D., Gerdes, A., Gordon, G.W., Algeo, T.J. & Boyle, E.A.
 Natural fractionation of ²³⁸U/²³⁵U. *Geochimica et Cosmochimica Acta* 72, 345-359
 (2008).
- 469 52. Hood, A.v.S., Planavsky, N.J., Wallace, M.W., Wang, X., Bellefroid, E.J., Gueguen, 470 B. & Cole, D.B. Integrated geochemical-petrographic insights from component-471 selective δ 238U of Cryogenian marine carbonates. *Geology* 44, 935-938 (2016).
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478 Figure legends

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Figure 1: Results of uranium isotope analysis. Error bars represent two standard errors. δ^{238} U values are shown from reduction spheroid innermost cores (black), core margins (grey), bleached halos (cyan) and surrounding red-bed matrix (magenta). Boxes indicate physically contiguous samples. **a:** Results from reduction spheroids of constrained formation depth. **b:** Results from other reduction spheroids. **c:** Results from three hydrothermal veins.

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Figure 2: Schematic view of geological context of Dingwall samples. Key to

symbols: stipples = igneous basement; parallel lines = mudrock; dots = sandstone;
open circles = conglomerate; bulls' eyes = reduction spheroids. Bitumen is present at

the cores of the reduction spheroids and in the local fractures. Adapted from ref. 24.

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Figure 3: Example of a freshly exposed reduction spheroid. This spheroid shows a

distinctive dark grey core, a bleached halo and a red-brown matrix flecked with
smaller reduction spots. The coin is ~21 mm across. Devonian red sandstone,

- 495 Millport, Great Cumbrae, Ayrshire, Scotland.
- 496
- 497
- 498 Table
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Isotope	²³² Th	²³³ U	²³⁴ U	²³⁵ U	²³⁶ U	²³⁸ U
Cup	L3	L2	L1	С	H1	Н3
Amplifier	10 ¹¹	10^{11}	10^{12}	10^{11}	10^{11}	10^{11}

500

501 **Table 1.** Neptune Faraday Detector Setup

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