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# Scaling limits of non-isentropic Euler-Maxwell equations for plasmas

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#### Abstract

In this paper, we will discuss asymptotic limit of non-isentropic compressible Euler-Maxwell system arising from plasma physics. Formally, we give some different limit systems according to the corresponding different scalings. Furthermore, some recent results about the convergence of non-isentropic compressible Euler-Maxwell system to the compressible Euler-Poisson equations will be given via the non-relativistic regime.

Keywords: non-isentropic Euler-Maxwell system, asymptotic limit, convergence

#### 1 Introduction and the formal limits

Let *n*, *u*, and  $\theta$  denote the scaled macroscopic density, mean velocity vector, and temperature of the electrons and *E* and *B* the scaled electric field and magnetic field, respectively. They are functions of a three-dimensional position vector  $x \in \mathbb{T}$  and of the time t > 0, where  $\mathbb{T} = (\mathbb{R}/2\pi\mathbb{Z})^3$  is the torus. The fields *E* and *B* are coupled to the particles through the Maxwell equations and act on the particles via the Lorentz force  $E + \gamma u \times B$ . These variables satisfy the scaled non-isentropic Euler-Maxwell system for plasma physics in a uniform background of non-moving ions with fixed density b(x) (see [1-3]):

$$\partial_t n + \nabla \cdot (nu) = 0, \tag{1.1}$$

$$\partial_t u + (u \cdot \nabla) u + \nabla \theta + \theta \nabla \ln n = -(E + \gamma u \times B), \tag{1.2}$$

$$\partial_t \theta + u \cdot \nabla \theta + \frac{2}{3} \theta \nabla \cdot u = 0, \tag{1.3}$$

$$\gamma \varepsilon \partial_t E - \nabla \times B = \gamma n u, \quad \gamma \partial_t B + \nabla \times E = 0, \tag{1.4}$$

$$\varepsilon \nabla \cdot E = b(x) - n, \quad \nabla \cdot B = 0, \tag{1.5}$$

$$(n, u, \theta, E, B)|_{t=0} = (n_0^{\gamma}, u_0^{\gamma}, \theta_0^{\gamma}, E_0^{\gamma}, B_0^{\gamma}).$$
(1.6)

In the system, Equations 1.1-1.3 are the mass, momentum, and energy balance laws, respectively, while (1.4)-(1.5) are the Maxwell equations. It is well known that two equations in (1.5) are redundant with two equations in (1.4) as soon as they are satisfied by the initial conditions. The non-dimensionalized parameters  $\gamma$  and  $\varepsilon$  can be

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chosen independently on each other, according to the desired scaling. Physically,  $\gamma$  and  $\varepsilon$  are proportional to  $\frac{1}{c}$  and the Debye length, where *c* is the speed of light. Thus, the limit  $\gamma \rightarrow 0$  is called the non-relativistic limit while the limit  $\varepsilon \rightarrow 0$  is called the quasi-neutral limit. Starting from one fluid and non-isentropic Euler-Maxwell system, we can derive some different limit systems according to the corresponding different scalings.

Case 1: Non-relativistic limit, Quasi-neutral limit

In this case, we first perform non-relativistic limit and then quasi-neutral limit. *Step 1:* Let  $\varepsilon$  be fixed and  $\gamma \rightarrow 0$ . Formally, we get the following system:

$$\partial_t n + \nabla \cdot (nu) = 0, \tag{1.7}$$

 $\partial_t u + (u \cdot \nabla) u + \nabla \theta + \theta \nabla \ln n = -E, \tag{1.8}$ 

$$\partial_t \theta + u \cdot \nabla \theta + \frac{2}{3} \theta \nabla \cdot u = 0, \tag{1.9}$$

$$\nabla \times B = 0, \quad \nabla \cdot B = 0, \tag{1.10}$$

$$\nabla \times E = 0, \quad \varepsilon \nabla \cdot E = b(x) - n.$$
 (1.11)

This limit is the Euler-Poisson system of compressible electron fluid.

**Remark 1.1**. Equations 1.10 implies B = 0 when the mean value of B(x, t) vanishes, *i.e.* m(B) = 0. Here,

$$m(v) = \frac{1}{(2\pi)^3} \int_{\mathcal{T}^3} v(x, \cdot) \, dx$$

denotes the mean value of a given scalar or vector function v(x, t) in  $\mathcal{T}^3$  with respect to x. Furthermore, equation  $\nabla \times E = 0$  in (1.11) with  $\mathbf{m}(E) = 0$  implies the existence of a potential function  $\varphi^0$  such that

 $E = -\nabla\phi.$ 

Step 2: Set  $\varepsilon = 0$  in Euler-Poisson system (1.7)-(1.11), we can obtain n - b(x) = 0, which is so-called quasi-neutrality in plasma physics. Then,  $(u, \theta, \varphi)$  satisfy the following equations:

$$\nabla \cdot (bu) = -\partial_t b, \tag{1.12}$$

 $\partial_t u + (u \cdot \nabla) u + \nabla \theta + \theta \nabla \ln b = \nabla \phi, \qquad (1.13)$ 

$$\partial_t \theta + u \cdot \nabla \theta + \frac{2}{3} \theta \nabla \cdot u = 0. \tag{1.14}$$

**Remark 1.2.** If the ion density b(x) is a constant, say b(x) = 1 for simplicity; then from (1.12)-(1.14), we see that  $(u, \theta, \varphi)$  satisfy the non-isentropic incompressible Euler equations of ideal fluid:

$$\nabla \cdot u = 0, \tag{1.15}$$

 $\partial_t u + (u \cdot \nabla) u + \nabla \theta = \nabla \phi, \tag{1.16}$ 

$$\partial_t \theta + u \cdot \nabla \theta + \frac{2}{3} \theta \nabla \cdot u = 0.$$
(1.17)

Hence, one derives the non-isentropic incompressible Euler equations.

Case 2: Quasineutral limit, Non-relativistic limit

In this case, we take b(x) = 1 for simplicity. Contrarily to Case 1, we first perform quasineutral limit and then non-relativistic limit.

Step 1: Let  $\gamma$  be fixed and  $\varepsilon$  turns to zero, one can gets n - 1 = 0 (quasineutrality), and then we get from the Euler-Maxwell system (1.1)-(1.5) that

$$\partial_t u + (u \cdot \nabla)u + \nabla\theta + E = -\gamma u \times B, \tag{1.18}$$

$$\partial_t \theta + u \cdot \nabla \theta + \frac{2}{3} \theta \nabla \cdot u = 0, \tag{1.19}$$

$$\nabla \times B = \gamma u, \quad \nabla \cdot B = 0, \tag{1.20}$$

$$\gamma \ \partial_t B + \nabla \times E = 0. \tag{1.21}$$

This is so-called the non-isentropic e-MHD equations. *Step 2:* We set  $\gamma = 0$  and get that

$$\nabla \times E = 0, \quad \nabla \times B = 0, \quad \nabla \cdot B = 0$$

and the non-isentropic incompressible Euler equations 1.16-1.18 of ideal fluid from the e-MHD system (1.18)-(1.21).

Case 3: Combined quasineutral and non-relativistic limits

Similarly to Case 2, we still take b(x) = 1 for simplicity. Choose  $\varepsilon = \gamma$  and let  $\varepsilon = \gamma \rightarrow 0$ , first it is easy to get from the Maxwell system (1.4)-(1.5) that n = 1 (quasi-neutrality) and

 $\nabla \times E = 0, \quad \nabla \times B = 0, \quad \nabla \cdot B = 0.$ 

Then one gets the non-isentropic incompressible Euler equations 1.15-1.17 of ideal fluid from the Euler-Maxwell system (1.1)-(1.5).

The above formal limits are obvious, but it is very difficult to rigorously prove them, even in isentropic case, see [4-6]. Since usually it is required to deal with some complex related problems such as the oscillatory behavior of the electric fields, the initial layer problem, the sheath boundary layer problem, and the classical shock problem. The proofs of these convergence are based on the asymptotic expansion of multiple-scale and the careful energy methods, iteration scheme, the entropy methods, etc. In the following section, we will provide a rigorous convergence result when  $\varepsilon$  is fixed (especially we take  $\varepsilon = 1$ ) and  $\gamma \rightarrow 0$ . We state our result in the following section. For detail, see [7]. For the other results, see [4-6] and references therein.

#### 2 Rigorous convergence

Let  $(n^{\gamma}, u^{\gamma}, \theta^{\gamma}, E^{\gamma}, B^{\gamma})$  be the classical solutions to problem (1.1)-(1.6) and assume that the initial conditions have the following asymptotic expansion with respect to  $\gamma$ :

$$\big(n_{0}^{\gamma}, u_{0}^{\gamma}, \theta_{0}^{\gamma}, E_{0}^{\gamma}, B_{0}^{\gamma}\big) = \sum_{j=0}^{m} \gamma^{j} \big(n_{j}, u_{j}, \theta_{j}, E_{j}, B_{j}\big) + O(\gamma^{m+1}).$$

Plugging the following ansatz:

$$(n^{\gamma}, u^{\gamma}, \theta^{\gamma}, E^{\gamma}, B^{\gamma}) = \sum_{j \ge 0} \gamma^{j} (n^{j}, u^{j}, \theta^{j}, E^{j}, B^{j})$$

$$(2.1)$$

into system (1.1)-(1.6), we obtain:

(1) The leading profiles  $(n^0, u^0, \theta^0, E^0, B^0)$  satisfy the following equations:

$$\partial_t n^0 + \operatorname{div}(n^0 u^0) = 0,$$
 (2.2)

$$\partial_t u^0 + (u^0 \cdot \nabla) u^0 + \nabla \theta^0 + \theta^0 \nabla \ln n^0 = -E^0, \qquad (2.3)$$

$$\partial_t \theta^0 + (u^0 \cdot \nabla) \theta^0 + \frac{2}{3} \theta^0 \nabla \cdot u^0 = 0, \qquad (2.4)$$

$$\nabla \times E^0 = 0, \quad \text{div}E^0 = b(x) - n^0,$$
 (2.5)

$$\nabla \times B^0 = 0, \quad \operatorname{div} B^0 = 0, \tag{2.6}$$

$$(n^{0}, u^{0}, \theta^{0})|_{t=0} = (n_{0}, u_{0}, \theta_{0}).$$
(2.7)

From (2.6), we may take  $B^0 = 0$ , and equation  $\nabla \times E^0 = 0$  in (2.5) implies the existence of a potential function  $\varphi^0$  such that  $E^0 = -\nabla \varphi^0$ . Then Equations 2.2-2.5 become a non-isentropic compressible Euler-Poisson system and determine a unique smooth solution  $(n^0, u^0, \varphi^0)$  in the class  $\mathbf{m}(\varphi^0) = 0$  well defined on  $\mathbb{T} \times [0, T_*]$  with  $T_* > 0$ . Here, we need the following compatibility conditions on  $(E_0, B_0)$ :

$$E_0 = -\nabla \phi_0, \quad B_0 = 0, \tag{2.8}$$

where  $\varphi_0$  satisfies

$$-\Delta\phi_0 = b(x) - n_0 \quad \text{in } \mathbb{T} \quad \text{and} \quad \mathbf{m}(\phi^0) = 0.$$
(2.9)

(2) For any  $j \ge 1$ , the profiles  $(n^j, u^j, \theta^j, E^j, B^j)$  can be obtained by induction. Now, we assume that  $(n^k, u^k, \theta^k, E^k, B^k)_{0 \le k \le j-1}$  are smooth and already determined in previous steps. Then  $(n^j, u^j, \theta^j, E^j, B^j)$  satisfy the following linearized equations:

$$\partial_t n^j + \sum_{k=1}^j \operatorname{div}(n^k u^{j-k}) = 0,$$
 (2.10)

$$\partial_t u^j + \sum_{k=0}^j \left( u^k \cdot \nabla \right) u^{j-k} + \theta^0 \nabla \left( \frac{n^j}{n^0} \right) + \theta^j \nabla (\ln n^0)$$

$$+ \nabla \theta^j + E^j + \sum_{k=0}^{j-1} u^k \times B^{j-1-k} + f^{j-1} = 0,$$

$$(2.11)$$

$$\partial_t \theta^j + \sum_{k=0}^j u^k \cdot \nabla \theta^{j-k} + \frac{2}{3} \sum_{k=0}^j \theta^k \operatorname{div} u^{j-k} = 0, \qquad (2.12)$$

$$\nabla \times E^{j} = -\partial_{t}B^{j-1}, \quad \text{div} \ E^{j} = -n^{j}, \tag{2.13}$$

$$\nabla \times B^{j} = \partial_{t} E^{j-1} - \sum_{k=0}^{j-1} n^{k} u^{j-1-k}, \quad \operatorname{div} B^{j} = 0,$$
(2.14)

$$(n^{j}, u^{j}, \theta^{j})|_{t=0} = (n_{j}, u_{j}, \theta_{j}),$$
(2.15)

where  $f^0 = 0$ ,  $\sum_{k=0}^{-1} = 0$  and  $f^{j-1}((n^k, \theta^k)_{k \le j-1})$  is defined by

$$\begin{pmatrix} \theta^0 + \sum_{j \ge 1} \gamma^j \theta^j \end{pmatrix} \nabla \ln \left( n^0 + \sum_{j \ge 1} \gamma^j n^j \right)$$
  
=  $\sum_{j \ge 0} \gamma^j (\theta^j \nabla \ln n^0 + \theta^0 \nabla (\ln' n^0 n^j)) + \sum_{j \ge 2} \gamma^j f^{j-1}$ 

Equations 2.14 are of curl-div type and they determine a unique smooth  $B^{j}$  in the class  $\mathbf{m}(B^{j}) = 0$  in  $\mathbb{T} \times [0, T_{*}]$ . Moreover, from div  $B^{j} = 0$ , we deduce the existence of a given vector function  $\omega^{j}$  such that  $B^{j} = -\nabla \times \omega^{j}$ . Then, the first equation in (2.13) becomes  $\nabla \times (E^{j} - \partial_{t}\omega^{j-1}) = 0$ . It follows that there is a potential function  $\varphi^{j}$  such that  $E^{j} = \partial_{t}\omega^{j-1} - \nabla \varphi^{j}$  with  $\omega^{0} = 0$ .

Then,  $(n^{j}, u^{j}, \theta^{j}, \varphi^{j})$  solve a compressible linearized Euler-Poisson system:

$$\partial_t n^j + \sum_{k=1}^j \operatorname{div}(n^k u^{j-k}) = 0,$$
 (2.16)

$$\partial_{t}u^{j} + \sum_{k=0}^{j} (u^{k} \cdot \nabla)u^{j-k} + \theta^{0}\nabla\left(\frac{n^{j}}{n^{0}}\right) + \theta^{j}\nabla(\ln n^{0})$$

$$+\nabla\theta^{j} - \nabla\phi^{j} + \sum_{k=0}^{j-1} u^{k} \times B^{j-1-k} + g^{j-1} = 0,$$
(2.17)

$$\partial_t \theta^j + \sum_{k=0}^j u^k \cdot \nabla \theta^{j-k} + \frac{2}{3} \sum_{k=0}^j \theta^k \operatorname{div} u^{j-k} = 0, \qquad (2.18)$$

$$-\Delta \phi^{j} = -n^{j} - \partial_{t} \operatorname{div} \omega^{j-1}, \qquad (2.19)$$

$$(n^{j}, u^{j}, \theta^{j})|_{t=0} = (n_{j}, u_{j}, \theta_{j})$$
(2.20)

where  $g^{j-1} = f^{j-1} + \partial_t \omega^{j-1}$ . Then system (2.16)-(2.20) determines a unique smooth solution  $(n^j, u^j, \theta^{j}, \varphi^{j})_{j\geq 1}$  in the class  $\mathbf{m}(\varphi^{j}) = 0$ , in the time interval [0,  $T_*$ ]. Since  $E^j = \partial_t \omega^{j-1} - \nabla \varphi^j$ , we need the following compatibility conditions on  $(E_j, B_j)$ :

 $E_j = \partial_t \omega_{j-1} - \nabla \phi_j, \quad B_j = B^j(0, x), \tag{2.21}$ 

where  $\varphi_i$  is determined by

$$-\Delta\phi_i = -n_i - \partial_t \operatorname{div}\omega^{j-1}(0, x), \qquad (2.22)$$

for  $x \in \mathbb{T}$  and  $\mathbf{m}(\varphi^{j}) = 0$ .

**Proposition 2.1.** Assume that the initial data  $(n_j, u_j, E_j, B_j)_{j<0}$  are sufficiently smooth with  $n_0 > 0$  in Tand satisfy the compatibility conditions (2.8)-(2.9) and (2.21)-(2.22). Then there exists a unique asymptotic expansion up to any order of the form (2.1), i.e. there exist the unique smooth profiles  $(n^j u^j, E^j, B^j)_{j<0}$ , solutions of the problems (2.2)-(2.7) and (2.10)-(2.15) in the time interval  $[0, T_*]$ . In particular, the formal non-relativistic limit

Set

 $\gamma \rightarrow 0$  of the non-isentropic compressible Euler-Maxwell system (1.1)-(1.6) is the non-isentropic compressible Euler-Poisson system.

$$\left(n_m^{\gamma}, u_m^{\gamma}, \theta_m^{\gamma}, E_m^{\gamma}, B_m^{\gamma}\right) = \sum_{j=0}^m \gamma^j (n^j, u^j, \theta^j, E^j.B^j),$$

where  $(n^{i}, u^{j}, \theta^{j}, E^{j}, B^{j})$  are those constructed in the previous Proposition 1.1.

For the convergence of the compressible Euler-Maxwell system (1.1)-(1.6), our main result is stated as follows.

**Theorem 2.1.** For any fixed integer  $s_0 > \frac{3}{2} + 1$  and  $m \ge 1$ , assume that the mean values of  $E^{\gamma}(x, t)$ ,  $B^{\gamma}(x, t)$  vanish and the ion density b(x) the initial data  $(n_j, u_j, \theta_j)_{j\ge 0}$ , satisfy the following conditions:

b(x), n<sub>j</sub>, u<sub>j</sub>, θ<sub>j</sub> ∈ H<sup>s</sup>(T), s ≥ N + 2, N ≥ j ≥ 0,
n<sub>0</sub>, θ<sub>0</sub> ≥ δ > 0 for some constant δ,
m(b(x) - n<sub>0</sub>) = m(n<sub>i</sub>) = 0, j ≥ 1

with N = m and  $s = m + s_0 + 3$  hold. Furthermore, if

$$(E_j, B_j)(x) = (E^j, B^j)(x, 0), \quad 0 \le j \le m,$$
 (2.23)

satisfy the compatibility condition

$$div E_0^{\gamma} = b(x) - n_0^{\gamma}, \quad div B_0^{\gamma} = 0, \quad x \in \mathbb{T},$$
(2.24)

and initial condition

$$\left\| \left( n_{0}^{\gamma}, u_{0}^{\gamma}, \theta_{0}^{\gamma}, E_{0}^{\gamma}, E_{0}^{\gamma} \right) - \sum_{j=0}^{m} \gamma^{j} \left( n_{j}, u_{j}, \theta_{j}, E_{j}, B_{j} \right) \right\|_{s_{0}} \le C \gamma^{m+1},$$
(2.25)

then, there exists  $T_* \in (0, T_*]$  such that problem (1.1)-(1.6) has a unique solution

$$(n^{\gamma}, n^{\gamma}, \theta^{\gamma}, E^{\gamma}, B^{\gamma}) \in C^{i}([0, T_{\star}], H^{s_{0}-i}(\mathbb{T})), \quad i = 0, 1.$$

Furthermore,

$$\left\| \left| n^{\gamma}, u^{\gamma}, \theta^{\gamma}, E^{\gamma}, B^{\gamma} \right| - \sum_{j=0}^{m} \gamma^{j} \left( n^{j}, u^{j}, \theta^{j}, E^{j}, B^{j} \right) \right\|_{s_{0}, T_{\star}} \leq C \gamma^{m+1},$$

where  $(n^{j}, u^{j}, \theta^{j}, E^{j}, B^{j})_{0 \le j \le m}$  are solutions to problems and C >0 is a constant independent of  $\gamma$ .

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All authors contributed equally to the manuscript and read and approved the final draft.

#### **Competing interests**

The authors declare that they have no competing interests.

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