## College of Arts and Sciences

Drexel E-Repository and Archive (iDEA)
http://idea.library.drexel.edu/

Drexel University Libraries
www.library.drexel.edu

The following item is made available as a courtesy to scholars by the author(s) and Drexel University Library and may contain materials and content, including computer code and tags, artwork, text, graphics, images, and illustrations (Material) which may be protected by copyright law. Unless otherwise noted, the Material is made available for non profit and educational purposes, such as research, teaching and private study. For these limited purposes, you may reproduce (print, download or make copies) the Material without prior permission. All copies must include any copyright notice originally included with the Material. You must seek permission from the authors or copyright owners for all uses that are not allowed by fair use and other provisions of the U.S. Copyright Law. The responsibility for making an independent legal assessment and securing any necessary permission rests with persons desiring to reproduce or use the Material.

Please direct questions to archives@drexel.edu

# Flexion and Skewness in Map Projections of the Earth 

David M. Goldberg (Drexel University Dept. of Physics)<br>J. Richard Gott III (Princeton University Dept. of Astrophysical Sciences)<br>goldberg@drexel.edu, jrg@astro.princeton.edu

April 12, 2007


#### Abstract

Tissot indicatrices have provided visual measures of local area and isotropy distortions. Here we show how large scale distortions of flexion (bending) and skewness (lopsidedness) can be measured. Area and isotropy distortions depend on the map projection metric, flexion and skewness, which manifest themselves on continental scales, depend on the first derivatives of the metric. We introduce new indicatrices that show not only area and isotropy distortions but flexion and skewness as well. We present a table showing error measures for area, isotropy, flexion, skewness, distances, and boundary cuts allowing us to compare different world map projections. We find that the Winkel-Tripel projection (already adopted for world maps by the National Geographic), has low distortion on most measures and excellent quality overall.


Keywords: Maps, Earth, Projection, Curvature

## 1 Introduction

Tissot (1881) indicatrices have been very useful for providing a visual presentation of local distortions in map projections in a simple and compelling fashion. A small circle of tiny radius (say 0.1 degree of arc in radius) is constructed at a given location, and then enlarged and projected on the map at that location. This always produces an ellipse. Usually one favors conformal map projections that minimize the changes in scale factor, or equal area projections that minimize anisotropy, or recently, map projections that are neither conformal nor equal area, but which have a judicious combination of minimizing both scale and isotropy errors (like the Winkel-tripel used by the National Geographic Society for world maps).

The Tissot ellipse at a given location is specified by three parameters, the major axis, the minor axis, and the orientation angle $\theta$ of the major axis of the ellipse. Geometrically, from differential geometry (and General Relativity) we know that the measurement of local distances is measured by the metric tensor $g_{a b}$, where a and b can each take the values x or y , yielding three independent components. Locally for two nearby points separated by infinitesimal map coordinate differences dx and dy , the true distance between these two points on the globe is given by:

$$
\begin{equation*}
d s^{2}=g_{x x} d x^{2}+2 g_{x y} d x d y+g_{y y} d y^{2} \tag{1}
\end{equation*}
$$

The Tissot ellipse (major axis, minor axis, and orientation angle, $\theta$ ) can be calculated from the three components of the metric tensor. Thus, the Tissot ellipse essentially carries the information on the metric tensor for the map. It tells us how local infinitesimal distances on the map correlate with local infinitesimal distances on the globe.

As we will see, however, the Tissot ellipse does not carry all of the information related to distortions, which has been noted previously. Others (Stewart 1943; Peters 1975, 1978; Albinus 1981; Canters 1989, 2002; Laskowski 1997ab) have noted that there are finite distortions apart from those of the Tissot ellipse.

Previous authors have previously depicted finite size distortions by several methods, including showing faces on a globe (Reeves 1910; Gedymin 1946), 30 degree x 30 degree equiangular quadralaterals (Chamberlin 1947), a net of twenty spherical triangles, in an icosahedral arrangement (Fisher \& Miller 1944), 150 degree great circle arcs, and circles of 15 degrees radius Tobler (1964).

In the Oxford Hammond Atlas of the World (1993), new conformal map projections ("Hammond Optimal Conformal Projections") were designed for the continents. Following the Chebyshev criterion (Snyder 1993), the rms scale factor errors were minimized by producing a constant scale factor along the boundary of the continent. (For a circular region this conformal map would be a stereographic projection.) By tailoring the boundary to the shape of each continent, the errors could be reduced relative to those in a simple stereographic projection. In touting the advantages of their projection the Hammond Atlas did the following experiment. Following Reeves (Reeves 1910), they constructed a face on the globe with a triangular nose, a straight (geodesic) mouth, and eyes that were pairs of concentric circles on the globe. They then showed this face with various map projections. In the gnomonic, the mouth was straight, but the eye circles were not circular and were not concentric (what we would call skewness). The Mercator projection had the mouth smiling (what we would call flexion) and although the eyes were circular they were not concentric. The Hammond Optimal Conformal projections did a bit better on these qualities because the gradients of the scale factor changes were small so the flexion and skewness were small, although of course not zero.

In Sec. 2 we introduce the concept of "flexion", by which a map projection can cause artificial bending of large structures. In Sec. 3, we show another distortion: skewness, which represents lopsidedness and an asymmetric stretching of large structures. We show a simple way to visualize these distortions in Sec. [4, in which we introduce the "Goldberg-Gott Indicatrices." In Sec. 5 we derive a differential geometry approach to measuring flexion and skewness. While readers interested in computing flexion and skewness on projections not included in this paper should refer to Sec. 5 it is highly technical, and those interested in seeing results may skip directly to Sec. 6, in which we discuss Monte Carlo estimates of the distortions for a number of projections, a ranking of map projections, and our conclusions.

## 2 First Distortion: Flexion

The local effects shown by the Tissot Ellipse are not the only distortions present in maps. There are also "flexion" (or bending) and "skewness" (or lopsidedness; discussed in the next section), which describe curvature distortions visible on world maps (the terminology stems from a similar effect in gravitational lensing; Goldberg \& Bacon 2005).

One can think about flexion in the following way. Imagine a truck going along a geodesic of the globe at unit angular speed (say one radian per day). Now imagine the image of that truck on the map, moving along. If the map were perfect, if it had zero flexion and zero skewness, then that truck would move in a straight line on the map with constant speed. Its velocity vector on the map:

$$
\begin{equation*}
\mathbf{v}=\frac{d \mathbf{x}}{d \tau} \tag{2}
\end{equation*}
$$

would be a constant, where $\tau$ is the angle of arclength in radians traveled by the truck along the geodesic on the globe. Its acceleration:

$$
\begin{equation*}
\mathbf{a}=\frac{d \mathbf{v}}{d \tau} \tag{3}
\end{equation*}
$$

would be zero. Of course, this cannot be true for a general geodesic. In the general case, the image of the truck suffers an acceleration as it moves along. The acceleration vector, $\mathbf{a}$, in the two-dimensional map has two independent components: $a_{\perp}$ (which is perpendicular to the truck's velocity vector at that point), and $a_{\|}$(which is parallel to its velocity vector at that point).

The perpendicular acceleration, $a_{\perp}$, causes the truck to turn without changing its speed on the map. This causes flexion, or bending, of geodesics. We define the flexion along a given geodesic at a given point to be:

$$
\begin{equation*}
f=\frac{a_{\perp}}{v}, \tag{4}
\end{equation*}
$$

or, more usefully:

$$
\begin{equation*}
f=\frac{d v_{\perp}}{d \tau} \frac{1}{v} . \tag{5}
\end{equation*}
$$

In this form, we may define:

$$
\begin{equation*}
d \alpha=\frac{d v_{\perp}}{v} . \tag{6}
\end{equation*}
$$

where $\alpha$ is the angle of rotation suffered by the velocity vector. Remember, if $a_{\perp}$ is the only acceleration present, then the velocity vector of the truck on the map does not change in magnitude, but just rotates in angle. Thus $f=d \alpha / d \tau$, and represents the angular rate at which the velocity vector rotates divided by the angular rate at which the truck moves on the globe.

Skewness and flexion only express themselves on large scales. They are not noticeable on infinitesimal scales where the metric contains all the information one needs, but become noticeable on finite scales, with their importance growing with the size of the area being examined. Flexion and skewness are thus important on continental scales and larger in a world map.

It is possible to design a map projection that has zero flexion: the gnomonic projection shows all great circles as straight lines. However it does exhibit anisotropy, scale changes, and skewness, and at best can show only one hemisphere of the globe.

### 2.1 Example 1: The Stereographic Projection

As an example, consider a truck traveling on the equator as seen by in the polar stereographic projection (see Fig [1). In the stereographic projection, the north pole is in the center of the map and the equator is a circle around it. As the truck circles the equator (the equator is a geodesic-so the truck drives straight ahead on the globe), it travels around a circle on the map. By azimuthal symmetry, the truck circles the equatorial circle on the map at a uniform rate. The velocity vector of the truck on the map rotates a complete $360^{\circ}$ ( $2 \pi$ radians), as the truck circles the equator, traversing $360^{\circ}$ of arc on the globe. So the flexion is $\mathrm{f}=1$, for a point on the equator, for a geodesic pointing in the direction of the equator. The flexion is defined at a point, and for a specific geodesic traveling through that point.

For an arbitrary point on the equator in the stereographic projection and a geodesic pointing in the direction of a meridian of longitude (also a geodesic) the flexion is zero, because these geodesics are shown as radial straight lines in the polar stereographic projection, and the velocity vector of the truck does not turn as it travels north.

The stereographic projection has the property that every great circle (geodesic) on the globe is shown as a circle on the map, except for a set of measure zero that pass through the north pole
(i.e. the meridians of longitude). Thus by the argument given above, the average flexion integrated around a random great circle must be $\langle f\rangle=1$, because the truck's velocity vector on a random great circle must rotate by $360^{\circ}$ as it circles the $360^{\circ}$ of arc completing that great circle on the globe. The magnitude of the velocity vector of the truck on the map is larger the further from the north pole it is and so its rotation per angle of arc of truck travel on the globe is larger there as well, and so the flexion along that random geodesic is larger the further away from the pole one is, with the integrated average value along the whole great circle being $\langle f\rangle=1$.

### 2.2 Example 2: The Mercator Projection

The Mercator projection (see Fig (2) is conformal and so only the scale factor changes as a function of position on the map (i.e. $g_{x x}=g_{y y}$, and $g_{x y}=0$ and the Tissot ellipses are all circles with radii proportional to $1 / g_{x x}$ ). But there is bending. The northern boundary between the continental United States and Canada at the $49^{\text {th }}$ parallel of latitude is shown as a straight line in the Mercator Map, but really it is a small circle that is concave to the north. If one drove a truck down that border from west to east, one would have to turn the steering wheel slightly to the left so that one was continually changing direction. The great circle route (the straightest route) connecting the Washington State and Minnesota (both at the 49th parallel) is a straight line which goes entirely through Canada. This straight line on the globe when extended, passes south of the northern part of Maine, so the continental United States is bend downward like a frown in the Mercator Map. (See Figure 3 and 4). Likewise, Maine on a Mercator map Maine sags below the line connecting Washington State and Minnesota, while on the globe this is not true.

In the Mercator projection, the flexion along the equator is zero, also along all meridians of longitude, but these are a set of measure zero. A random geodesic is a great circle that is inclined at some angle between $0^{\circ}$ and $90^{\circ}$ with respect to the equator. On the Mercator map this is a wavy line that bends downward in the northern hemisphere, and by symmetry, upward in the southern. Since the curvatures are equal and opposite in the two hemispheres, the average flexion $\langle f\rangle=0$, but this is misleading because the flexion at each point off the equator is not zero. So if we are rating map projections by the amount of flexion they contain we should use the absolute value of the flexion instead: $|f|$. In a region where the flexion does not change sign (such as the northern hemisphere in the Mercator projection or the entire stereographic map) the total bending of a geodesic segment will be the integral of the flexion $|f|$ over that segment. In fact, in Section 6.1 we will evaluate the overall flexion on a map by simply picking random points on the sphere and random directions for geodesics going through them, and then calculating the absolute value for the flexion for all random points on the globe and random directions through them.

### 2.3 A Global Flexion Measure

We can calculate the flexion for any point in the Mercator (or any other) projection through any geodesic using spherical trigonometry. As a reminder to the reader, the Mercator projection uses the mapping:

$$
\begin{align*}
x & =\lambda  \tag{7}\\
y & =\ln (\tan [\pi / 4+\phi / 2]) \tag{8}
\end{align*}
$$

where here and throughout, $\lambda$ is the longitude expressed in radians, and $\phi$ is the latitude expressed in radians.

On the map, the angle rotation along the geodesic with an azimuth, $\theta$, is calculated by constructing thin spherical triangle with side-angle-side given by $(\pi / 2-\phi, \theta, d \tau)$.

$$
\begin{equation*}
d \alpha=\pi-\beta-\theta \tag{9}
\end{equation*}
$$

because the geodesic intersects the north-south meridian (a vertical line in the Mercator) at an angle of $\theta$ initially, and at an angle $\pi-\beta$ at the other end, where $\beta$ is the angle in the spherical triangle at the other end of the $d \tau$ side. Solving for $d \alpha$ using spherical trigonometry in the limit as $d \tau$ goes to zero, we find that

$$
\begin{equation*}
f=\sin \theta \tan \phi \tag{10}
\end{equation*}
$$

Thus, for $\theta=\pi / 2$, an east-west geodesic, we find that

$$
\begin{equation*}
f_{E W}=\tan \phi \tag{11}
\end{equation*}
$$

so that in the northern hemisphere, traveling east one's geodesic is bending clockwise with $d \alpha / d \tau=$ $\tan \phi$. Therefore, the east-west geodesic bends downward. For, $\theta=0$, a north-south geodesic, the flexion is zero, as we expect, since the meridians of longitude are straight in the Mercator map. If we average over all azimuths at a given point, we find:

$$
\begin{equation*}
\langle | f(\phi)\left\rangle=|\tan \phi| \frac{2}{\pi}\right. \tag{12}
\end{equation*}
$$

Now we can integrate this over all points on the sphere to produce the average flexion over the whole sphere F. Taking advantage of the symmetry between the northern and southern hemisphere we can integrate only over the northern hemisphere (where $d A=2 \pi \cos \phi d \phi$ ), yielding:

$$
\begin{align*}
F & =\langle | f| \rangle  \tag{13}\\
& =\frac{\int 2 \pi \cos \phi \tan \phi \frac{2}{\pi} d \phi}{2 \pi} \\
& =\frac{2}{\pi}
\end{align*}
$$

The flexion is less than that of the stereographic because of the $180^{\circ}$ boundary cut along the longitude line at the international date line. A geodesic is a great circle on the globe, and if this is shown as a closed curve on the map that is always concave inward (the best possible case) it will always have a total rotation of the velocity vector of $360^{\circ}$ and so will have an average integrated flexion of 1 . If there is a boundary cut, the great circle does not have to close on the map (it has two loose ends at the boundary cut) and so need not completely rotate by $360^{\circ}$.

## 3 Second Distortion: Skewness

Acceleration in the direction parallel to the velocity vector of the truck $a_{\|}$, causes the truck to increase its speed along the geodesic curve without causing any rotation. This causes skewness, because as the truck accelerates, it covers more distance on the map on one side of a point than on the other, so the point in question will not be at the center of the line segment of arc centered on that point on the sphere.

Consider a segment of a meridian of longitude on the globe centered at $45^{\circ}$ north latitude. Going from South to North along that geodesic in the Mercator map the truck is accelerating with $a_{\|}>0$, because the scale factor is getting larger and larger the further north one goes, so as the
truck continues to cover equal arc length on the globe it covers larger and larger distances on the map. Thus, the center of the segment (at $45^{\circ}$ latitude) is not centered on the segment on the map.

We define the skewness:

$$
\begin{equation*}
s \equiv \frac{a_{\|}}{v} \tag{14}
\end{equation*}
$$

Taking the explicit case of the Mercator projection, we find:

$$
\begin{align*}
v_{y} & =\frac{d y}{d \phi}  \tag{15}\\
& =\frac{1}{\cos \phi}
\end{align*}
$$

and thus:

$$
\begin{equation*}
a_{\|}=\frac{\tan \phi}{\cos \phi}, \tag{16}
\end{equation*}
$$

so the skewness (for a vector pointed N-S) is simply:

$$
\begin{equation*}
s=\tan \phi . \tag{17}
\end{equation*}
$$

The skewness at $45^{\circ}$ is 1 , showing a lopsidedness toward the north. Given this relation, the skewness is positive (northward lopsidedness) in the northern hemisphere and negative (southward lopsidedness) in the southern hemisphere.

Consider a geodesic through a point in the northern hemisphere tipped at an azimuth angle of $\theta$ with respect to north. The only thing increasing the speed of the truck is the gradient of the scale factor as one moves northward, so the amplitude of the parallel acceleration is equal to the maximum acceleration (obtained going straight north) times $\cos \theta$. To get the average of the absolute value of the skewness for all geodesics through that point at all random angles $\theta$, one simply integrates over $\theta$ :

$$
\begin{align*}
<|s|> & =|\tan \phi| \frac{\int_{0}^{\pi / 2} \cos \theta d \theta}{\pi / 2}  \tag{18}\\
& =|\tan \phi| \frac{2}{\pi}
\end{align*}
$$

As with the flexion, we can integrate this over all points on the sphere to produce the average skewness over the whole sphere S. Similarly, we find:

$$
\begin{equation*}
S=\frac{2}{\pi} \tag{19}
\end{equation*}
$$

Notice that this is exactly the same value as the average flexion, $F$, for the Mercator. We will find that for conformal projections, the average absolute value of the skewness and flexion at a given point and over the whole globe are always equal. (This is only true for conformal projections, for general projections the skewness and flexion can be different, as illustrated by the gnomonic projection which has zero flexion but non-zero skewness.)

In the Mercator projection, at the equator, the skewness is zero, as we would expect from symmetry considerations. Indeed, because any geodesic crossing equator has a symmetric shape in the northern and southern hemisphere, the skewness $s=0$ for any geodesic line evaluated at a point on the equator. Likewise, the flexion is zero for any geodesic line evaluated at a point on the equator. So the Mercator map has perfect local shapes along the equator, uniform scale along the
equator (Tissot ellipses all equal size circles) and zero flexion and skewness along geodesics in any direction from points on the equator.

While much of the analysis in this work specifically addresses distortions in maps of the earth, these effects must also be taken into account in other maps as well. One of us (JRG) has recently produced a conformal "map of the universe" (Gott et al. 2005) based on the logarithm map of the complex plane. The horizontal coordinate is the celestial longitude in radians, yielding a $360^{\circ}$ panorama from left to right. The vertical coordinate is:

$$
\begin{equation*}
y=\ln \left(d / r_{\oplus}\right) \tag{20}
\end{equation*}
$$

where $d$ is the distance, and $r_{\oplus}$ is the radius of the earth. The distance scale goes inversely as the distance, allowing us to plot everything from satellites in low earth orbit to stars and galaxies, to the cosmic microwave background on one map. The map is conformal, having perfect local shapes. However, it does have flexion and skewness. Circles of constant radius from the earth are bent into straight lines for example, and a rocket going out from the earth at constant speed would be slowing down on the map.

In Figure 3, note that the continental United States also appears lopsided on the Mercator map. The geographical center of the continental United States (which is in Kansas) appears in the lower half of the continental United States on the Mercator map because the scale factor on the map gets larger and larger the further north one goes on the Map. Thus, the continental United States is lopsided toward the north in the Mercator map. Flexion or bending is manifest on the map as a bending of geodesics on the map, and skewness or lopsidedness is manifest on the map as the midpoint of a geodesic line segment on the map not being at the midpoint of that geodesic arc as shown on the map.

## 4 Goldberg-Gott Indicatrices

We began this discussion with the virtues of the Tissot indicatrices. Likewise, we have produced a simple indicatrix to show the flexion and skewness in a map as well as the isotropy and area properties indicated by the Tissot indicatrices. We will refer to them as "Goldberg-Gott" indicatrices. These are constructed as follows. At a specific point on the map draw a circle on the globe of radius $12^{\circ}$, and then plot it on the map. Inside this circle, plot the north-south, and east-west geodesics through the central point on the map. This leaves a $\oplus$ symbol on the map. If the map were perfect, this would be a perfect circle and the cross arms would be perfectly straight, with their intersection at the center of the circle.

We have produced such a Goldberg-Gott indicatrix located at the geographic center of the continental United States for using a Mercator projection in Figure 3,

We note that our technique is similar to that of Tobler (1964), who projected circles of 15 degrees in radius on to maps. However, it differs in that Tobler's circles do not have their centers marked nor perpendicular geodesics drawn from their centers like our indicatrices, and so do not convey the information on flexion and skewness. But they do show the finite shape distortions of the circles themselves which our indicatrices also cover. Tobler includes tables showing for particular locations on the sinusoidal projection, the maximum and minimum scale radius and maximum difference of radial directions for circles of various radii on the sphere (which address the Tissot issues of size and isotropy on finite circles but not flexion and skewness). Tobler also calculated errors in miles and angular errors for 300 random triangles within spherical circles of various radii in different map projections, and considered area errors for such random triangles within land areas in various world map projections.

Using our indicatrices, one can see in Fig. 3 that the north-south geodesic is straight, but that the east-west geodesic is bent downward. This shows dramatically the flexion in this region of the Mercator map. One can even read off the average value of the flexion by hand. Take a protractor and measure the tangent to the east-west geodesic at the two ends of the cross bar. Measure the difference in the angle orientation of the two. That gives the integrated flexion along $24^{\circ}$ of the globe. Divide that angle difference by $24^{\circ}$ and you will have the average value of the flexion along that curve.

The skewness is also visible in that the center of the cross is below the center of the circle, showing the lopsidedness to the north. In fact, one can observe the skewness in any direction from the center by seeing how far off center the center of the cross is with respect to the center of the circle in different directions. For comparison, we have in Figure 4 shown the continental United States in a oblique Mercator projection where the east-west geodesic through the geographic center of the continental United States is now the equator of the Mercator projection.

The flexion and skewness along the equator of a Mercator map are indeed zero, so the arms of the cross are now straight, and the circle is now nearly a perfect circle centered on the center of the cross. This gives a "straight on" view of the continental United States, that more accurately portrays its appearance on the globe.

One can place the Goldberg-Gott indicatrices every $60^{\circ}$ in longitude and every $30^{\circ}$ in latitude on the globe to show how the flexion and skewness vary over the map. In Figs. 55 [27, we provide G-G indicatrix maps for a number of well-known projections.

In fact, the Goldberg-Gott almost indicatrices can just replace the Tissot indicatrices because the shape and size of the oval in the Goldberg-Gott indicatrix is approximately the size and shape of the Tissot ellipse. The Tissot ellipse shows the [magnified] shape and size of an infinitesimal circle on the globe, the oval in the Goldberg-Gott indicatrix $\oplus$ shows the shape and size of a finite circle (radius $12^{\circ}$ ) on the map itself at correct scale. Thus, if the map is equal area, the GoldbergGott indicatrices, will all have equal area on the map. If the map is conformal the Goldberg-Gott indicatrices will all be nearly perfectly circular. If there is a $2: 1$ anisotropy in the Tissot ellipses in a given region the Goldberg-Gott indicatrices ovals will have that same 2:1 axis ratio.

## 5 A differential geometry approach

Thus far, we have defined the general properties of skewness and flexion, given a few analytic results for particular map projections, and given a graphic approach for describing and interpreting flexion and skewness on maps. In this section, we approach the matter somewhat differently, and produce general analytic results for all projections as well as a prescription for measuring the flexion and skewness analytically.

### 5.1 Coordinate Transforms

Let's consider a spherical globe with coordinates:

$$
\begin{equation*}
x^{\bar{a}}=\binom{\phi}{\lambda} \tag{21}
\end{equation*}
$$

Note that here and throughout, we will use $x^{\bar{a}}$ to refer to coordinates in the globe frame, and $x^{a}$ to refer to coordinates in the map frame.

On the globe, the metric is:

$$
g_{\bar{a} \bar{b}}=\left(\begin{array}{cc}
1 & 0  \tag{22}\\
0 & \cos ^{2} \phi
\end{array}\right)
$$

such that, as always, the distance between two points can be expressed as:

$$
\begin{equation*}
d l^{2}=d x^{\bar{a}} d x^{\bar{b}} g_{\bar{a} \bar{b}} \tag{23}
\end{equation*}
$$

Now, consider an arbitrary 2-d coordinate transformation:

$$
\begin{equation*}
\binom{x}{y}=\binom{x^{1}(\phi, \lambda)}{x^{2}(\phi, \lambda)} \tag{24}
\end{equation*}
$$

Of course, from this definition, we may easily compute a local transformation matrix:

$$
\begin{equation*}
\Lambda_{\bar{a}}^{a}=\frac{\partial x^{a}}{\partial x^{\bar{a}}} \tag{25}
\end{equation*}
$$

The inverse matrix is $\Lambda_{a}^{\bar{a}}=\partial x^{\bar{a}} / \partial x^{a}$. The metric in the map frame is:

$$
\begin{equation*}
g_{a b}=\Lambda_{a}^{\bar{a}} \Lambda_{b}^{\bar{b}} g_{\bar{a} \bar{b}} \tag{26}
\end{equation*}
$$

From this, we may then compute the Christoffel Symbols of general relativity:

$$
\begin{equation*}
\Gamma_{b c}^{a}=\frac{1}{2} g^{a e}\left(g_{e b, c}+g_{e c, b}-g_{b c, e}\right) \tag{27}
\end{equation*}
$$

where standard convention tells us to sum over identical indices in the upper and lower positions, and where a comma indicates a partial derivative with respect to a coordinate.

In practice, actually computing the Christoffel symbols for an arbitrary projection is not simple. To do this analytically requires that we have an analytic form for the map inversion. However, we make available a numerical code to compute the Christoffel symbols for all map projections discussed in this work on our projections website (see below).

### 5.2 Analytic forms of Flexion and Skewness

The whole point of computing the Christoffel symbols is that we want to address a very simple question: How are large structures distorted when projected onto a map? Clearly to an observer on the globe, a straight line is easy to generate. Point in a particular direction, and start driving (assuming your car can drive on the ocean) with the steering wheel set straight ahead. Drive for a fixed distance in units of angles or radian. Record all points along the way.

Geometers, of course, know this route as a geodesic, and if we consider $\tau$ to represent a physical distance on the surface of the earth, then the geodesic equation may be expressed as:

$$
\begin{equation*}
\frac{d u^{a}}{d \tau}=-\Gamma_{b c}^{a} u^{b} u^{c} \tag{28}
\end{equation*}
$$

where

$$
\begin{equation*}
u^{a}=\frac{d x^{a}}{d \tau} \tag{29}
\end{equation*}
$$

Equation (28)) describes the bending of straight lines on a particular map projection, and thus, if all of the Christoffel symbols could vanish, we would clearly have a perfect Cartesian map. Not possible, of course.

But what is the physical significance of the Christoffel symbols? Since the lower indices are symmetric by inspection, there are 6 unique symbols. What do they mean?

### 5.2.1 Analytic Flexion

We define a vector oriented in a particular direction:

$$
\begin{equation*}
\tilde{u}^{a}(\theta)=\binom{\cos \theta}{\sin \theta} \tag{30}
\end{equation*}
$$

In reality, this is not the unit vector, since:

$$
\begin{equation*}
|\tilde{u}|^{2}=g_{a b} \tilde{u}^{a} \tilde{u}^{b} \neq 1 \tag{31}
\end{equation*}
$$

Of course, we could define a true unit unit vector:

$$
\begin{equation*}
u(\theta)=l(\theta) \tilde{u}(\theta) \tag{32}
\end{equation*}
$$

where $l$ is the "length" in grid coordinates of the unit vector. This has a value of:

$$
\begin{equation*}
l(\theta)=\frac{1}{\sqrt{\cos ^{2} \theta g_{11}+\sin ^{2} \theta g_{22}+2 \sin \theta \cos \theta g_{12}}} \tag{33}
\end{equation*}
$$

Of course, it is clear that at any point on the map, the set of all $u(\theta)$ represents an ellipse - the Tissot ellipse.

We define the flexion in the following manner: Follow a particular geodesic a distance, $d \tau$ (measured in, for example, radians). On the map, the geodesic will change direction by an angle $d \theta$. The ratio:

$$
\begin{equation*}
f=\frac{d \theta}{d \tau} \tag{34}
\end{equation*}
$$

is the flexion. Note that for all polar projections, the equator will have a flexion of 1.
In terms of the geodesic equation, the flexion can be expressed as:

$$
\begin{align*}
f(\theta) & =l(\theta)\left(\frac{d \tilde{\mathbf{u}}}{d \tau} \times \tilde{\mathbf{u}}\right) \\
& =l(\theta)\left(\Gamma_{a b}^{2} \tilde{u}^{a} \tilde{u}^{b} \tilde{u}^{1}-\Gamma_{a b}^{1} \tilde{u}^{a} \tilde{u}^{b} \tilde{u}^{2}\right) \tag{35}
\end{align*}
$$

where $\tilde{\mathbf{u}}$ denotes a vector.
Some light can be shed on equation (35) if we expand the expression explicitly into trigonometric functions:

$$
\begin{align*}
f(\theta)= & l(\theta)\left[\Gamma_{11}^{2} \cos ^{3} \theta+\left(2 \Gamma_{12}^{2}-\Gamma_{11}^{1}\right) \cos ^{2} \theta \sin \theta\right. \\
& \left.-\left(2 \Gamma_{12}^{1}-\Gamma_{22}^{2}\right) \sin ^{2} \theta \cos \theta-\Gamma_{22}^{1} \sin ^{3} \theta\right] \tag{36}
\end{align*}
$$

This immediately reveals an important idea. If:

$$
\begin{align*}
\Gamma_{11}^{2} & =\Gamma_{22}^{1}=0 \\
\Gamma_{11}^{1} & =2 \Gamma_{12}^{2} \\
\Gamma_{22}^{2} & =2 \Gamma_{12}^{1} \tag{37}
\end{align*}
$$

then $f \theta$ ) in all directions will be zero. That is, all geodesics will be straight lines in this projection. As we will see, this is true for the Gnomonic projection.

Inspection of equation (36) shows it is clearly anti-symmetric about exchanges of $\theta$ and $-\theta$, since $\theta$ is not the normal vector to the geodesic, but rather runs along it.

### 5.2.2 Analytic Skewness

Essentially, a skewness means that if you walk (initially) north (for example) for 1000 miles, or walk south for 1000 miles, you will cover different amounts of map coordinate.

The skewness along a geodesic can be defined similarly to the flexion:

$$
\begin{align*}
s(\theta) & =l(\theta)\left(\frac{d \tilde{\mathbf{u}}}{d \tau} \cdot \tilde{\mathbf{u}}\right) \\
& =l(\theta)\left(\Gamma_{b c}^{1} \tilde{u}^{b} \tilde{u}^{c} \tilde{u}^{1}+\Gamma_{b c}^{2} \tilde{u}^{b} \tilde{u}^{c} \tilde{u}^{2}\right) \tag{38}
\end{align*}
$$

As with the flexion, we may expand these out explicitly:

$$
\begin{align*}
s(\theta)= & l(\theta)\left[\Gamma_{11}^{1} \cos ^{3} \theta+\left(2 \Gamma_{12}^{1}+\Gamma_{11}^{2}\right) \cos ^{2} \theta \sin \theta\right. \\
& \left.+\left(2 \Gamma_{12}^{2}+\Gamma 1_{22}\right) \sin \theta \cos ^{2} \theta+\Gamma_{22}^{2} \sin ^{3} \theta\right] \tag{39}
\end{align*}
$$

A map with no skewness will have:

$$
\begin{align*}
& \Gamma_{11}^{1}=\Gamma_{22}^{2}=0 \\
& \Gamma_{11}^{2}=-2 \Gamma_{12}^{1} \\
& \Gamma_{22}^{1}=-2 \Gamma_{12}^{2} \tag{40}
\end{align*}
$$

Unlike the flexion, we know of no projections with zero skewness everywhere.

### 5.3 Projections with straightforward analytic results

### 5.3.1 The Gnomonic Projection

The Gnomonic Projection is particularly interesting. It has a coordinate transformation:

$$
\begin{equation*}
\binom{x}{y}_{g}=\binom{\cot \phi \cos \lambda}{\cot \phi \sin \lambda} \tag{41}
\end{equation*}
$$

which is directly invertible to yield:

$$
\begin{equation*}
\binom{\phi}{\lambda}=\binom{\cos ^{-1} \sqrt{\frac{x^{2}+y^{2}}{1+x^{2}+y^{2}}}}{\tan ^{-1}\left(\frac{y}{x}\right)} \tag{42}
\end{equation*}
$$

The coordinate transformation is thus:

$$
\Lambda_{\bar{a}}^{a}=\left(\begin{array}{cc}
-\frac{x}{\sqrt{x^{2}+y^{2}}\left(1+x^{2}+y^{2}\right)} & -\frac{y}{\sqrt{x^{2}+y^{2}}\left(1+x^{2}+y^{2}\right)}  \tag{43}\\
-\frac{y}{x^{2}+y^{2}} & \frac{x}{x^{2}+y^{2}}
\end{array}\right)
$$

We thus have the map metric:

$$
g_{a b}=\left(\begin{array}{cc}
\frac{1+y^{2}}{\left(1+x^{2}+y^{2}\right)^{2}} & \frac{-x y}{\left(1+x^{2}+y^{2}\right)^{2}}  \tag{44}\\
\frac{-x y}{\left(1+x^{2}+y^{2}\right)^{2}} & \frac{1+x^{2}}{\left(1+x^{2}+y^{2}\right)^{2}}
\end{array}\right)
$$

We can compute the Christoffel symbols in the normal way. We find:

$$
\begin{align*}
\Gamma_{11}^{1} & =-\frac{2 x}{1+x^{2}+y^{2}} \\
\Gamma_{11}^{2} & =0 \\
\Gamma_{12}^{1} & =-\frac{y}{1+x^{2}+y^{2}} \\
\Gamma_{12}^{2} & =-\frac{x}{1+x^{2}+y^{2}} \\
\Gamma_{22}^{1} & =0 \\
\Gamma_{22}^{2} & =-\frac{2 y}{1+x^{2}+y^{2}} \tag{45}
\end{align*}
$$

This clearly satisfies the requirements of equations (37), but not (40), and thus, the Gnomonic projection produces straight, but skewed geodesics.

### 5.3.2 Stereographic

The Stereographic projection is conformal, and thus, all of the Tissot ellipses are circles. Does this mean there is no skewness in the projection? No, as we've already seen. The Stereographic projection has the coordinate transformation:

$$
\begin{equation*}
\binom{x}{y}_{g}=\binom{\tan (\pi / 4+\phi / 2) \cos \lambda}{\tan (\pi / 4+\phi / 2) \sin \lambda} \tag{46}
\end{equation*}
$$

which, again, is directly invertible to yield:

$$
\begin{equation*}
\binom{\phi}{\lambda}=\binom{\frac{\pi}{2}-2 \tan ^{-1}\left(\sqrt{x^{2}+y^{2}}\right)}{\tan ^{-1}\left(\frac{y}{x}\right)} \tag{47}
\end{equation*}
$$

The coordinate transformation is:

$$
\Lambda_{\bar{a}}^{a}=\left(\begin{array}{cc}
-\frac{2 x}{\sqrt{x^{2}+y^{2}}\left(1+x^{2}+y^{2}\right)} & -\frac{2 y}{\sqrt{x^{2}+y^{2}}\left(1+x^{2}+y^{2}\right)}  \tag{48}\\
-\frac{x}{x^{2}+y^{2}} & \frac{x}{x^{2}+y^{2}}
\end{array}\right)
$$

This can be used to compute the metric on the map:

$$
g_{a b}=\left(\begin{array}{cc}
\frac{4}{\left(1+x^{2}+y^{2}\right)^{2}} & 0  \tag{49}\\
0 & \frac{4}{\left(1+x^{2}+y^{2}\right)^{2}}
\end{array}\right)
$$

This clearly indicates that all Tissot ellipses will be circular.
From these, of course, we can compute the Christoffel symbols:

$$
\begin{align*}
\Gamma_{22}^{1} & =-\Gamma_{11}^{1}=-\Gamma_{12}^{2}=\frac{2 x}{1+x^{2}+y^{2}}  \tag{50}\\
\Gamma_{11}^{2} & =-\Gamma_{22}^{2}=-\Gamma_{12}^{1}=\frac{2 y}{1+x^{2}+y^{2}} \tag{51}
\end{align*}
$$

It is clear from inspection that geodesics are generally neither straight nor unskewed.

Moreover, it is clear that $l(\theta)$ is independent of orientation since the map is conformal. In general it can be shown to be:

$$
\begin{equation*}
l=\frac{1}{2}\left(1+x^{2}+y^{2}\right) . \tag{52}
\end{equation*}
$$

In the stereographic projection, a circle of radius $12^{\circ}$ on the globe is a perfect circle on the map but the center of the circle on the globe is not at the center of the circle on the globe (see Figure 11), and thus, there is skewness.

## 6 Discussion

### 6.1 Numerical Analysis of Standard Map Projections

Not all projections produce such simple results. Thus, in general, we will want to compute the local flexion and skewness numerically. Our approach is as follows: For each projection we chose 30,000 points selected randomly on the surface of a globe. For each of these points, we chose a random direction to start a geodesic. We follow that geodesic using small steps ( $d \tau \simeq 10^{-5} \mathrm{rad}$ ) numerically, and use standard difference methods to compute the map velocity and acceleration along the geodesic. We are thus able to compute the metric and the Christoffel symbols (and thus the flexion and skewness) directly. We make our IDL (Interactive Data Language) code available to the interested reader at our projection webpage (see below). Likewise, we also do a distance test, in which pairs of points, ( $\mathrm{i}, \mathrm{j}$ ), are chosen at random and the distance is measured both on the globe and on the map.

This is a somewhat different perspective than simply inspecting the Goldberg-Gott indicatrices at a few locations, since we are now doing a uniform sample over the surface of the globe, rather than a uniform sampling over the map. When looking at the indicatrix map we can occasionally get a distorted view as to the quality of a particular projection. Some (like the Mercator) have relatively good fits over most of the globe, but the high latitudes can, in principle, be projected to infinite areas, and thus, the reader may erroneously think the Mercator infinitely bad. By sampling uniformly over the globe, we get a fair assessment of the overall quality of a particular projection.

We define a number of fit parameters: I, corresponding to errors in the local isotropy (zero for conformal projections), A, corresponding to errors in the Area (zero for equal area projections), F, corresponding to flexion (defined in the discussion of flexion, above), S, corresponding to skewness (also defined above), D, corresponding to distance errors, and B , corresponding to the average number of map boundary cuts crossed by the shortest geodesic connecting a random pair of points.

$$
\begin{align*}
I & =R M S\left(\ln \frac{a_{i}}{b_{i}}\right)  \tag{53}\\
A & =R M S\left(\ln a_{i} b_{i}-\left\langle\ln a_{i} b_{i}\right\rangle\right)  \tag{54}\\
F & =\langle | f_{i}| \rangle  \tag{55}\\
S & =\langle | s_{i}| \rangle  \tag{56}\\
D & =R M S\left(\ln \frac{d_{i j, \text { map }}}{d_{i j, g l o b e}}\right)  \tag{57}\\
B & =\frac{L_{B}}{4 \pi} \tag{58}
\end{align*}
$$

where $a_{i}$ and $b_{i}$ are the major and minor axes of the local Tissot ellipses of random point, $i,\left\langle X_{i}\right\rangle$ indicates the mean of property, X , and $L_{B}$ is the total length of the boundary cuts.

Logarithmic errors have been used before by Kavrayskiy (1958). Bugayevskiy and Snyder (1995) adopted schemes that weighted isotropy errors, $(a / b-1)^{2}$, and area errors, $(a b-1)^{2}$, according to one's interest in isotropy and area. Airy (1861) had just given these terms equal weight.

| Projection | I | A | F | S | D | B |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |
| Non Interrupted Projections |  |  |  |  |  |  |
| Azimuthal Equidistant | 0.87 | 0.60 | 1.0 | 0.57 | 0.356 | 0 |
| Gott-Mugnolo Azimuthal | 1.2 | 0.20 | 1.0 | 0.59 | 0.341 | 0 |
| Lambert Azimuthal | 1.4 | 0 | 1.0 | 2.1 | 0.343 | 0 |
| Stereographic | 0 | 2.0 | 1.0 | 1.0 | 0.714 | 0 |
| 180 deg. Boundary Cut |  |  |  |  |  |  |
| Breisemeister | 0.79 | 0 | 0.81 | 0.42 | 0.372 | 0.25 |
| Eckert IV | 0.70 | 0 | 0.75 | 0.55 | 0.390 | 0.25 |
| Eckert VI | 0.73 | 0 | 0.82 | 0.61 | 0.385 | 0.25 |
| Equirectangular | 0.51 | 0.41 | 0.64 | 0.60 | 0.449 | 0.25 |
| Gall-Peters | 0.82 | 0 | 0.76 | 0.69 | 0.390 | 0.25 |
| Gall Stereographic | 0.28 | 0.54 | 0.67 | 0.52 | 0.420 | 0.25 |
| Gott Elliptical | 0.86 | 0 | 0.85 | 0.44 | 0.365 | 0.25 |
| Gott-Mugnolo Elliptical | 0.90 | 0 | 0.82 | 0.43 | 0.348 | 0.25 |
| Hammer | 0.81 | 0 | 0.82 | 0.46 | 0.388 | 0.25 |
| Hammer-Wagner | 0.687 | 0 | 0.789 | 0.518 | 0.377 | 0.25 |
| Kavrayskiy VII | 0.45 | 0.31 | 0.69 | 0.41 | 0.405 | 0.25 |
| Lagrange | 0 | 0.73 | 0.53 | 0.53 | 0.432 | 0.25 |
| Lambert Conic | 0 | 1.0 | 0.67 | 0.67 | 0.460 | 0.25 |
| Mercator | 0 | 0.84 | 0.64 | 0.64 | 0.440 | 0.25 |
| Miller | 0.25 | 0.61 | 0.62 | 0.60 | 0.439 | 0.25 |
| Mollweide | 0.76 | 0 | 0.81 | 0.54 | 0.390 | 0.25 |
| Polyconic | 0.79 | 0.49 | 0.92 | 0.44 | 0.364 | 0.25 |
| Sinusoidal | 0.94 | 0 | 0.84 | 0.68 | 0.407 | 0.25 |
| Winkel-Tripel | 0.49 | 0.22 | 0.74 | 0.34 | 0.374 | 0.25 |
| Winkel-Tripel (Times) | 0.48 | 0.24 | 0.71 | 0.373 | 0.39 | 0.25 |
| 1 360 deg. Boundary Cut |  |  |  |  |  |  |
| Lambert Azimuthal (2 hemisphere) | 0.36 | 0 | 0.52 | 0.11 | 0.432 | 0.5 |
| Stereographic (2 hemisphere) | 0 | 0.39 | 0.37 | 0.37 | 0.692 | 0.5 |
| Multiple Boundary Cut Projections |  |  |  |  |  |  |
| Gnomonic Cube | 0.22 | 0.37 | 0.12 | 0.87 | 0.43 | 0.686 |

Table 1: The errors in isotropy, area, flexion, skewness, distances, and boundary cuts for some standard projections.

In Table 6.1 we compare a these measures for a number of standard projections, which, for fairness of comparison we divide into a number of categories. The Gott-Elliptical, The GottMugnolo Elliptical, and the Gott-Mugnolo Azimuthal have been discussed in Gott, Mugnolo \& Colley (2007) and Gott, et al. (2007), where they have been applied to the earth, Mars, the moon, and the Cosmic Microwave all sky map.

First, we show projections which represent the complete globe without interrupts. These projections are azimuthal and the average flexion over these maps is 1 .

Second, we show the set of whole earth projections with one $180^{\circ}$ interrupt. These include
rectangular and elliptical projections. Note that among all of the complete projections with 0 or $1180^{\circ}$ interrupt, there are a number of "winners" with regards to performance for flexion and skewness. The Lagrange has the smallest flexion. The Winkel-Tripel has the smallest skewness. For all conformal projections, the skewness is equal to the flexion.

Of all of the whole earth projections, the most accurate for distance measure between points is the Gott-Mugnolo Azimuthal, followed very closely by the Lambert Azimuthal (Gott, Mugnolo \& Colley 2007).

In the third and fourth groups, we show 2-hemisphere and other multiple cut projections, respectively.

In the final group, we have a projection with multiple interrupts, the gnomonic cube, which is defined piecemeal. This is a particularly interesting projection since the gnomonic is locally flexion-free, but it is clear that geodesics will not trace out straight lines in the gnomonic cube cube map because they bend when they cross an edge between faces. The gnomonic cube is presented as a cross, so 5 edges are included in the map proper. Geodesics bend when they cross an edge in this laid out cross configuration.

The above comparisons do not depend on how important each of the criteria are (i.e. what weighting to give each measure). Laskowski (1997ab) suggested a means of ranking very disparate maps. Though his weighting scheme is not unique, as a simple illustration of how this can be done, we will minimize the sum of the squares of all 6 parameters, normalized to their values in the equirectangular projection:

$$
\begin{equation*}
\Sigma_{\varepsilon}=\left(\frac{I}{N_{i}}\right)^{2}+\left(\frac{A}{N_{a}}\right)^{2}+\left(\frac{F}{N_{f}}\right)^{2}+\left(\frac{S}{N_{s}}\right)^{2}+\left(\frac{D}{N_{d}}\right)^{2}+\left(\frac{B}{N_{b}}\right)^{2} \tag{59}
\end{equation*}
$$

Following Laskowksi (1997ab) we set the normalization constants equal to the values of these errors in the Equirectangular projection $(x=\lambda, y=\phi): N_{i}=0.51, N_{a}=0.41, N_{f}=0.64, N_{s}=0.60$, $N_{d}=0.449, N_{b}=0.25$.

The projections with the lowest values of $\Sigma_{\varepsilon}$ are:

1. Winkel-Tripel (4.5629)
2. Winkel-Tripel (Times Atlas) (4.5687)
3. Kavrayskiy VII (4.8390)
4. Gall Stereographic (5.7582)
5. Hammer-Wagner (5.7847)

## 6. Eckert IV (5.8519)

This approach is certainly not unique. One may take issue with the weighting of the individual parameters, the domain over which they are applied (the whole earth, as opposed to continents only, for example), or even how the parameters are computed, we present it as a simple example of how our results may be combined with previous studies of map projections. It is interesting that the "best" map, as selected by this criterion is the one already used by the National Geographic for it's whole world projection. It is also interesting to note that the Winkel-Tripel has especially low skewness.

Another nice property of this weighting scheme is that it gives a low score to pathological projections. For example, a series of $n$ gores (made using the polyconic projection) arranged in a sunflower pattern would approximate the Azimuthal equidistant projection in distance errors as $n$
became large but would have arbitrarily low values of I, A, F, and S. But if a boundary cut term B is included this term would blow up and save us from picking the bad subdivided map as better than the more visually pleasing projections described above. Stretching individual pixels at the edges of the gores to fill the gaps between them would eliminate the boundary cuts. However, it would also cause the skewness to blow up, preventing us from giving a good score to a bad projection.

Interested readers may visit www.physics.drexel.edu/~goldberg/projections/ to download a free IDL code to measure the flexion, area, and other measures discussed in this paper. We have not done all known projections, but have covered ones that have available mathematical formulas and we thought likely to do well.

### 6.2 Conclusions

We have developed two new measures of curvature distortions found in maps of the earth. The Skewness and the Flexion can be used to identify particularly warped sections of a map, or to identify features (such as the U.S. Canada border), which appear as a straight line on some map projections but do not, in fact, follow geodesics. We have developed a new graphical tool, called the Goldberg-Gott indicatrices, which can be used to show these distortions, along with those in area and shape, at many points in a map, and have produced indicatrix maps for a number of popular projections. Finally, we have used these measures to produce global distortion measures for different projections. We have found that the Winkel-Tripel produces low distortion on most measures, and, in particular, has the lowest skewness of all projections in our sample with a $180^{\circ}$ boundary cut.

## Acknowledgments

JRG is supported by National Science Foundation grant AST04-06713. DMG is supported by a NASA Astrophysics Theory Grant. We thank Wes Colley for useful discussions, as well as the helpful comments made by the anonymous referees.

## References

[1] Airy, G.B. (1861), London, Edinburgh, and Dublin Philosophical Magazine, 4th ser. 22(149), 409.
[2] Albinus, H.-J., (1981), Anmerkungen und Kritik zur Entfernungsverzerrung, Kartographische Nachrichten, 31, 179-83.
[3] Bugayevskiy, L. M. \& Snyder, J. P., (1995) Map Projections, A Reference Manual, Taylor \& Francis, London
[4] Canters, F. (1989) New projections for world maps, a quantitative-perceptive approach, Cartographica, 26,53-71
[5] Canters, F. (2002), Small-scale map projection design, Research Monographs in Geographical Information Systems, London and New York: Taylor and Francis.
[6] Chamberlin, W. (1947), The round Earth on flat paper: map projections used by cartographers, Washington: National Geographic Society
[7] Fisher, I. \& Miller, O.M. (1944) World maps and globes, New York: Essential Books
[8] Gedymin, A.B. (1946) Kartographiya, Moscow
[9] Goldberg, D.M. \& Bacon, D.J., (2005) Astrophys. J., 619, 741
[10] Gott, J. R., Colley, W., Park, C.-G., Park, C., Mugnolo, C., in preparation
[11] Gott, J.R., Juric, M. Schlegel, D., Hoyle, F., Vogeley, M., Tegmark, M., Bahcall, N. \& Brinkmann, J., (2005) Astrophys. J., 624, 463
[12] Gott, J. R., Mugnolo, C., Colley, W. (2007) accepted to Cartographica
[13] Kavrayskiy, V. V., (1958), Izbrannye trudy, vol. II: Mathematicheskaja kartografija, part 1: Obscaja teorija kartograficheskih projekcij, Izdaniye Nachal'nika Upraveleniya, Gidrograficheskoy Sluzhby Voyenno-morskoy Akademii.
[14] Laskowski, P., (1997a), Cartographica, 34, 3
[15] Laskowski, P., (1997b). Cartographica, 34, 19
[16] Oxford-Hammond World Atlas, (1993) Oxford, p 10-11.
[17] Peters, A.B. (1975) Wie man unsere Weltkarten der Erde ahnlicher machen kann, Kartographische Nachrichten 25, 173-83
[18] Peters, A.B. (1978) Uber Weltkartenverzerrunngen und Weltkartenmittelpunkte, Kartographische Nachrichten, 28, 106-13
[19] Reeves, E.A. (1910), Maps and Map-Making, London: Royal Geographical Society.
[20] Snyder, J.P. (1993) Flattening the Earth, University of Chicago Press, Chicago.
[21] Stewart, J.Q. (1943), The use and abuse of map projections, Geographical Review, 33, 589-604.
[22] Tissot, Nicolas A., (1881) Memoire sur la representation des surfaces et les projections des cartes geographiques. Gauthier Villars, Paris.
[23] Tobler, W.R. (1964), Geographical coordinate computations, Part II, Finite map projection distortions, Technical Report no. 3, ONR Task no. 389-137, University of Michigan, Department of Geography, Ann Arbor.


Figure 1: The Indicatrix map for a Stereographic projection.


Figure 2: The Indicatrix map for a Mercator projection.


Figure 3: A Mercator projection cutout of continental the United States. We have have put a Goldberg-Gott indicatrix (a circle of radius $12^{\circ}$ with N-S and E-W geodesics from the central point) at the geographic center of the continental U.S.


Figure 4: A oblique Mercator projection cutout of the United States. We have have put a GoldbergGott indicatrix at the geographic center of the continental U.S. Notice that this projection has no flexion or skewness at the center.


Figure 5: The Indicatrix map for an azimuthal equidistant projection.


Figure 6: The Indicatrix map for a Briesemeister projection.


Figure 7: The Indicatrix map for an Eckert IV projection.


Figure 8: The Indicatrix map for an Eckert VI projection.


Figure 9: The Indicatrix map for a Gall-Peters projection.


Figure 10: The Indicatrix map for an Equirectangular projection.


Figure 11: The Indicatrix map for a Gall Stereographic projection.


Figure 12: The Indicatrix map for a gnomonic cube projection.


Figure 13: The Indicatrix map for a Gott-Mugnolo projection.


Figure 14: The Indicatrix map for a Gott-Mugnolo Elliptical projection.


Figure 15: The Indicatrix map for a Gott Equal-Area Elliptical projection.


Figure 16: The Indicatrix map for a Hammer projection.


Figure 17: The Indicatrix map for a Hammer-Wagner projection.


Figure 18: The Indicatrix map for a Kavrayskiy VII projection.


Figure 19: The Indicatrix map for a Lagrange projection.


Figure 20: The Indicatrix map for a Lambert Azimuthal projection.


Figure 21: The Indicatrix map for a Lambert Conic projection.


Figure 22: The Indicatrix map for a Miller projection.


Figure 23: The Indicatrix map for a Mollweide projection.


Figure 24: The Indicatrix map for a Polyconic projection.


Figure 25: The Indicatrix map for a Sinusoidal projection.


Figure 26: The Indicatrix map for a Winkel-Tripel projection.


Figure 27: The Indicatrix map for a Winkel-Tripel (Times Atlas) projection.

