

# Mapping magnetic sources at the millimeter to micrometer scale in dunite and serpentinite by high-resolution magnetic microscopy

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## ABSTRACT

1 Rock samples can have wide range of magnetic properties depending on composition, amount of ferromagnetic  
2 minerals, grain sizes and microstructures. Here, we used scanning magnetic microscopy, a highly sensitive and  
3 high-resolution magnetometric technique to map remanent magnetic fields over a planar surface of a rock sample.  
4 The technique allows for the investigation of discrete magnetic mineral grains, or magnetic textures and structures  
5 with submillimeter scale resolution. Here, we present a case-study of magnetic scans of pristine and serpentized  
6 dunite thin sections from the Reinfjord Ultramafic Complex, in northern Norway. The magnetic mineralogy is  
7 characterized by electron microprobe, scanning electron- and optical-microscopy, and with rock magnetic  
8 methods. In serpentized samples the magnetic carrier is end-member magnetite occurring as large discrete  
9 grains and small grains in micron scale veins. By contrast, the pristine dunite sample contains large Cr-spinel  
10 grains with very fine equant exsolutions ranging in composition from ferrichromite to end-member magnetite.  
11 Forward and inverse modeling of the magnetic anomalies is used to determine the remanent magnetization  
12 directions and intensities of discrete magnetic sources observed in the scanning magnetic microscopy. The fine-  
13 scale magnetization of the rock sample is used to investigate the magnetic carriers and the effect of  
14 serpentization on the magnetic properties of the dunite. Modeling shows that the dipolar magnetic anomalies  
15 that are mapped by scanning magnetic microscopy are caused by grains with heterogeneous magnetic sources.  
16 The intensity of the magnetization and the amount of magnetic minerals are higher in the serpentized sample  
17 than the pristine dunite sample, consistent with the measured bulk magnetic properties. Furthermore, the  
18 serpentized samples show a larger variability in the direction of the magnetization and a stronger heterogeneity  
19 with respect to the pristine sample. The ability to rigorously associate components of the bulk magnetic properties  
20 to individual mineral phases creates new possibilities for rock magnetic, paleomagnetic, and exploration  
21 applications.

## 1. INTRODUCTION AND GEOLOGICAL SETTING

22 Geological samples have a wide range of magnetic properties depending on quantity of ferromagnetic minerals,  
23 and their compositions, grain sizes and microstructures. These properties influence magnetic anomalies from the  
24 micro- to the planetary scale. The natural remanent magnetization (NRM) of a sample is additionally dependent  
25 on the time and conditions of magnetic acquisition, so it reflects and can record the geological history of the  
26 sample. Secondary processes such as serpentinization or metamorphism can significantly alter both mineralogical  
27 characteristics and NRM, dramatically affecting rock magnetic properties and in turn changing the nature of the  
28 magnetic anomalies. Therefore, a comprehensive characterization of the magnetic petrology of the rock and its  
29 thermal history is needed for accurate interpretation of magnetic anomalies.

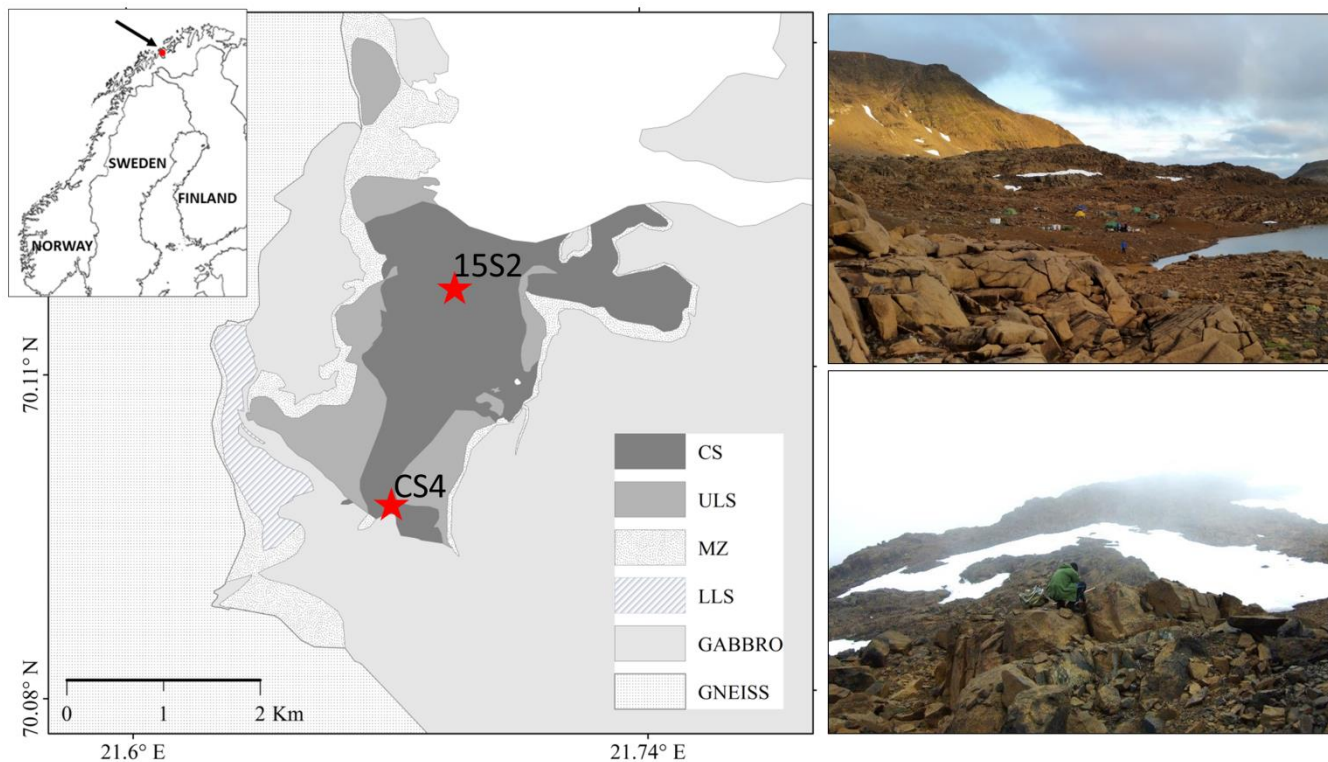
30 Rock magnetic methods are widely applied to measure rocks magnetic properties and characterize the carriers.  
31 While indispensable, traditional methods are bulk measurements that do not directly relate magnetic properties to  
32 individual mineral phases or microstructures. To attribute specific magnetic signals to the underlying mineralogy,  
33 techniques must be employed that can resolve magnetic properties at a fine scale. The ability to discriminate  
34 differing behavior of constituent phases is necessary for a complete understanding of the origin of bulk behavior  
35 measured in both the laboratory and in magnetic surveys, and provides vital evidence about primary and  
36 secondary geological processes and their role in determining magnetic response. One such technique that offers  
37 spatially-resolved measurements of magnetic signals is scanning magnetic microscopy (e.g. Fu et al., 2014;  
38 Fukuzawa et al., 2017; Hankard et al., 2009; Lima et al., 2014; Noguchi et al., 2017; Oda et al., 2011; Tominaga  
39 et al., 2017; Weiss et al., 2000; Weiss et al., 2007). This emerging technique generates an accurate map of the  
40 magnetic field distribution over a planar surface of a rock sample with sub-millimeter resolution. Previous  
41 applications of scanning magnetic microscopy have been primarily as an extension of, and complement to  
42 traditional paleointensity and paleomagnetic techniques. Weiss et al. (2008) used the technique to investigate fine  
43 scale heterogeneity of magnetization in Martian meteorites and estimate the ancient Martian field strength.  
44 Paleointensity estimates are commonly based on bulk measurements, and do not account for the non-  
45 unidirectional orientation of the fine scale magnetization of the sample, which controls the bulk properties.

46 Tominaga et al. (2017) investigated changes in the magnetic field intensity with the mineralogy during a  
47 carbonation sequence and used scanning magnetic microscopy to trace the reaction front. Oda et al. (2011) and  
48 Noguchi et al. (2017) used SQUID magnetic microscopy to generate a fine scale magnetostratigraphy to estimate  
49 ages and growth rate of ferromanganese crust. While scanning magnetic microscopy has been applied in Earth  
50 sciences for several decades (Thomas et al., 1992), interpretation of the data remains an unresolved problem. Most  
51 previous work modeled the data acquired by scanning magnetic microscopy in terms of dipole moment intensity  
52 and directions, which is sufficient for paleomagnetic and paleointensity studies and appropriate for sources with  
53 simple geometries. However, laterally extensive sources, or those with non-uniform magnetization, common in  
54 nature, may not be accurately described by single dipole field. Furthermore, comparison with bulk measurements  
55 requires that the volume of the magnetic material be taken into account, hence calculating magnetization, rather  
56 than moment. Weiss et al. (2002) describe an inversion method to calculate the magnetization distribution from  
57 magnetic scan data using a grid of evenly spaced dipoles, with a fixed volume for grid voxels. The size of the data  
58 sets, however, means that such methods are computationally intensive and time-consuming. Here, we propose an  
59 alternative approach to a full scan inversion, or to the dipole moment determination. This consists of forward  
60 modeling of the magnetization of a three-dimensional source geometry using the compositional and geometrical  
61 constraints given by optical and electron microscopy. This approach, by including the geometry of the source,  
62 limits the degrees of freedom which characterize the inherently non-unique magnetic data inversions. We propose  
63 forward modelling to estimate the magnetization of discrete sources. When applied to the entire sample, the  
64 magnetization estimates of discrete grains can then be compared to the bulk properties of the sample.

65 Here, scanning magnetic microscopy and forward magnetic modeling are used in combination with chemical and  
66 magnetic properties analyses to characterize the magnetic carriers in three samples; a pristine dunite sample  
67 (CS4), and two highly serpentinized samples (15S2D and 15S2B).

68 The samples are from the Reinfjord Ultramafic complex (RUC)(Fig. 1), of the Seiland igneous province (SIP) in  
69 northern Norway. This complex was emplaced during the Ediacaran at a depth of 25–35 km (Larsen et al., 2018)  
70 and later uplifted. The SIP is now exposed in the Middle Allochthon of the Norwegian Caledonian belt. Although  
71 the SIP has a complex geodynamic history, most of the magmatic textures of its rocks are well preserved. The

72 samples were selected to examine the origin of the primary magnetization in the dunite, and the effect of later  
73 serpentinization on the bulk properties. The pristine samples are believed to preserve primary NRM carriers of the  
74 lower crust, a topic of debate that is strongly linked to the thermal history of the rocks (McEnroe et al., 2018). In  
75 addition, there are local serpentinized areas within the ultramafic outcrops. Serpentinization is a relatively low  
76 temperature ( $\leq 400$  °C) fluid-mediated hydration process and in ultramafic rocks commonly leads to the  
77 production of magnetite. The creation of secondary magnetite could result in a composite NRM of the rock, a  
78 combination of the primary and secondary magnetizations, or may completely overprint the original NRM. The  
79 characterization of the NRM of discrete magnetite grains could be useful to distinguish different stages of  
80 serpentinization. By directly relating the micrometer scale anomalies to the mineralogy, we can improve our  
81 understanding of processes that control the magnetism of a rock and link these to the geological history. A greater  
82 understanding of the processes and features at the mineral scale will enhance our interpretation of magnetic  
83 anomalies on outcrop, regional, and planetary scales.



84  
85 **Fig. 1. Geological map of the Reinjfjord Ultramafic Complex (modified after Grannes, 2016) with samples**  
86 **localities (red stars); CS4 sample is taken from the southern side of the complex, while the 15S2 locality is from**

87 *the northern side. Right: outcrops photographs of the dunite rocks from the Central Series formation. The*  
88 *ultramafic complex, surrounded by gabbroic rocks and gneisses, consists of three ultramafic series: the Central*  
89 *series (CS), the upper layered series (ULS) and the lower layered series (LLS).*

## 2. METHODS

90 We investigated three samples of the Reinfjord Ultramafic Complex with optical and electron microscopy, rock  
91 magnetic methods, and magnetic modeling. Microscopy provides precise measurements of size, shape, and  
92 chemical composition of oxide and sulphide particles in a thin section, which had been surveyed in the scanning  
93 magnetic microscope before exposure to electron microscope fields. Bulk magnetic properties were measured on  
94 chips or cores of companion samples. Magnetic modeling of the magnetic microscopy scans of the thin sections  
95 was applied to isolated anomalies associated with discrete grains to estimate the magnetization intensity and  
96 direction of the magnetic grains. Modal mineralogy was investigated using optical and scanning electron  
97 microscopy (SEM) imaging by backscattered electrons at the NTNU NanoLab using a FEI Helios G4 UX  
98 scanning electron microscope (SEM). Chemical analyses were made using a JEOL 8200 SuperProbe (Electron  
99 Probe Microanalyzer-EPMA) at the University of Milan using wavelength-dispersive spectroscopy (WDS)  
100 techniques. All samples were analyzed at the microprobe with a spot current of 5 nA and 15 keV accelerating  
101 voltage. Points were spot analyzed with a beam diameter of 1  $\mu\text{m}$  and measuring time of 10 s on background and  
102 30 s on peak. Thin-sections magnetic scans were made with a scanning SQUID microscope at the Geological  
103 Survey of Japan, National Institute of Advanced Industrial Science and Technology (AIST), and on a newly built  
104 scanning magnetic tunnel junction instrument (here after referred to as the MTJ microscope) at the NTNU Rock-  
105 and Paleomagnetism laboratory. Both instruments measure the vertical component of the field and all imaging  
106 was carried out at room temperature ( $\sim 20^\circ\text{C}$ ) in field-free conditions. Therefore the signals represent remanent  
107 behavior. The nominal sampling step for all scans is 100  $\mu\text{m}$  in x and y. The SQUID microscope system uses a 200  
108  $\times$  200  $\mu\text{m}$  square washer type pickup coil, which has a field resolution of 1.1 pT/ $\sqrt{\text{Hz}}$  at 1 Hz (Kawai et al., 2016)  
109 and a sample stage with positioning accuracy of  $\sim 10 \mu\text{m}$  (Oda et al., 2016). Measurements were conducted with a

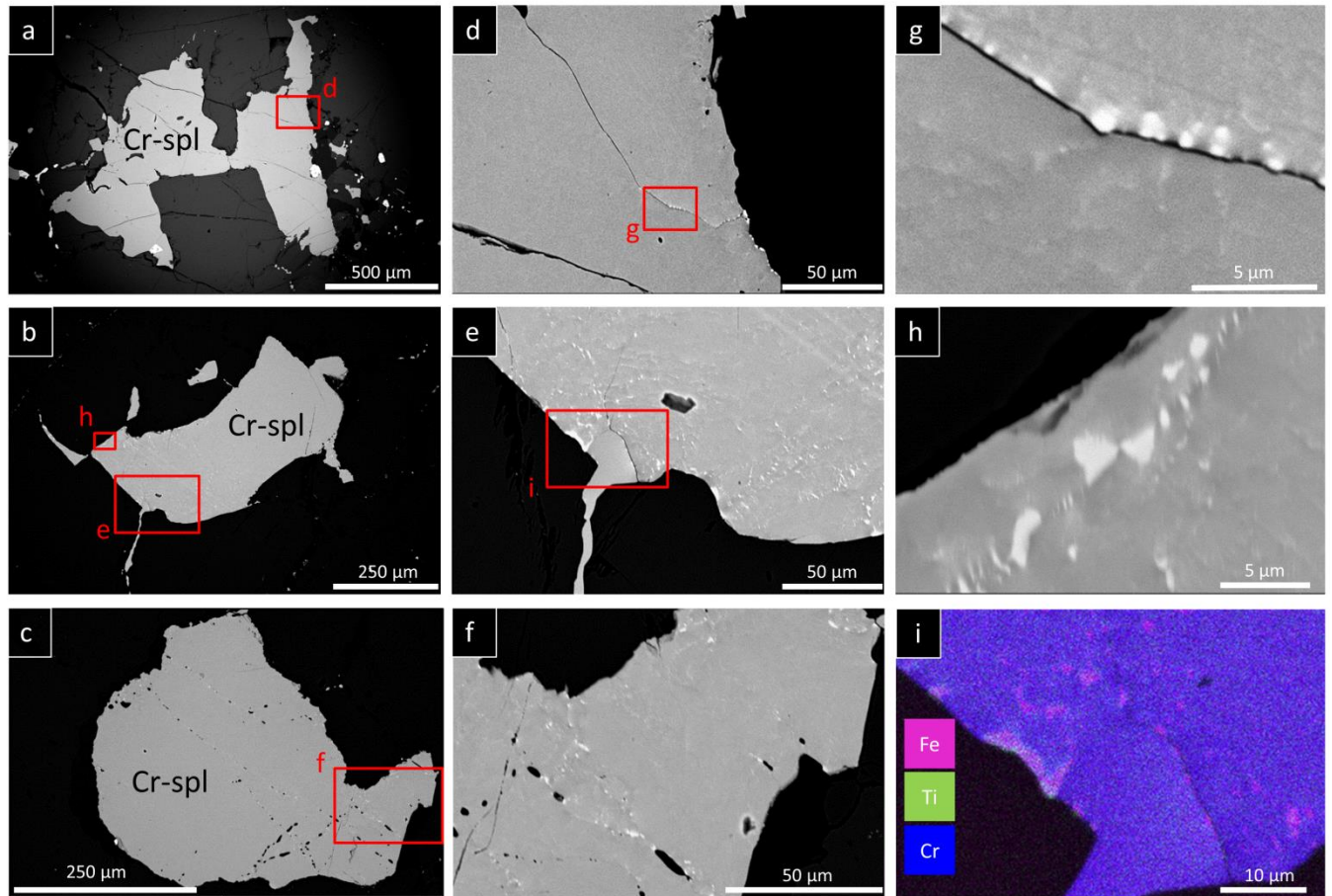
110 sensor-to-sample distance of approximately 253 $\mu$ m. The MTJ microscope has field noise of  $\sim 70$  nT/ $\sqrt{\text{Hz}}$  at 1 Hz,  
111 and positioning accuracy of  $\sim 100$  nm; the poorer noise performance of this instrument is partially offset by aver-  
112 aging 5 identical measurements. A sensor-to-sample distance of approximately 200 $\mu$ m was used for the surveys  
113 with this instrument. The spatial resolution of discrete magnetic sources in both devices is dependent on the  
114 sensor active area, scanning step size, positioning accuracy, measurement speed, sensor-to-sample distance and on  
115 the thickness of the magnetization distribution. Modeling of the magnetic data acquired by magnetic scanning  
116 microscopy was made using Tensor Research ModelVision software. Bulk rock magnetic property analyses were  
117 performed at NTNU using a variety of techniques. NRM was measured on sample cores of 2.5 cm diameter and  
118 2.2 cm height, or sample chips using an AGICOJR6 spinner magnetometer with sensitivity of 2 $\mu$ A/m.  
119 Susceptibility values were measured using a Sapphire susceptibility bridge on sample cores and an  
120 AGICOMFK1-A Kappabridge on sample chips with sensitivity of 6 $\cdot 10^{-8}$ SI. Temperature dependence of AC  
121 susceptibility was measured in argon, and in air using an AGICOMFK1-A-Kappabridge on powdered samples.  
122 For high temperature measurements, samples were heated from room temperature (RT) to 700  $^{\circ}\text{C}$  before cooling  
123 again to ambient temperature at an interval of 11  $^{\circ}\text{C}/\text{min}$ ; for low temperature run, samples were cooled from  
124 room temperature to  $-194$   $^{\circ}\text{C}$  and then heated back to room temperature. High- and low-field susceptibility,  
125 saturation remanence ( $M_r$ ) and saturation magnetization ( $M_s$ ) curves were measured as a function of temperature  
126 using a Princeton PMC Model 3902/F MicroMag Vibrating Sample Magnetometer (VSM) with a flowing helium  
127 furnace instead. The instrument measures the magnetic moment with an average sensitivity of 0.5 nA $\cdot$ m $^2$ .  
128 Measurements were made on chips of the samples using the quarter-hysteresis loop method of Fabian et al. (2013)  
129 here with a maximum field of 1 Tesla. Curie temperatures and blocking behavior were estimated from  
130 thermomagnetic curves. Room temperature hysteresis measurements were acquired before and after each thermal  
131 experiment.

### 3. DATA AND RESULTS

#### 3.1. PETROGRAPHY AND MINERAL CHEMISTRY

132 The three RUC samples discussed here are a pristine dunite sample CS4, and two serpentinized dunite samples  
133 15S2D and 15S2B. Samples contain minor opaque phases including oxides and sulfides (pentlandite, pyrrhotite  
134 and chalcopyrite). Accessory amounts of amphibole, pyroxenes, calcite, dolomite and biotite are also present.  
135 Based on image analysis of optical images, mineral phases abundances of sample CS4 result is 92.3% olivine in  
136 large subhedral crystals (1–3 mm in size), 7% pyroxenes (diopside and enstatite), occurring as interstitial grains  
137 between olivine grains, and the remaining 0.7% opaque minerals, including oxides and sulfides. The dominant  
138 opaque mineral is Cr-spinel. Minor amounts of ilmenite and pentlandite are present. Opaque grain sizes are  $\leq 0.1$ –  
139 mm. Backscattered electron (BSE) images of the Cr-spinel grains show these are not homogeneous and host fine-  
140 grained Fe-rich intergrowths (Fig. 2) with sizes varying from  $\leq 200$  nm to  $5\mu\text{m}$ . The Fe-rich exsolution 'blebs' are  
141 all designated as ferrichromite; however there is a variation in Fe-rich compositions and some are near or end-  
142 member magnetite. In the heavily serpentinized samples (15S2D and 15S2B) olivine has been mostly replaced by  
143 lizardite, brucite and magnetite; however, some relict olivine and pyroxene grains are recognizable in cross-  
144 polarized light. Based on image analyses of optical images, sulfides and oxides constitute up to the 7% of the  
145 serpentinized samples. The sulfides are mostly pentlandite and pyrrhotite with minor chalcopyrite. The main  
146 oxides are magnetite, ilmenite and Cr-spinels. Backscattered electron (BSE) images of opaque minerals show that  
147 spinel grains from serpentinized samples are homogeneous, unlike the Cr-spinel grains in CS4. Magnetite is  
148 present throughout the thin section, in small few-micrometer-thick veins (Fig. 3g, h), or in large grains (up to  
149  $700\mu\text{m}$ ) together with pentlandite (Fig. 3), Al/Cr spinel and ilmenite (Fig. 3a, c, d). Of the opaque grains,  
150 ferrichromite, magnetite and monoclinic pyrrhotite retain a remanent magnetization. Because the magnetic  
151 properties of these phases are strongly controlled by their composition, precise chemical analyses were measured.  
152 Measurements were taken from a homogeneous spot area at the microprobe scale. Analyses were calculated as  
153 weight percent of oxides. Representative analyses of spinels are shown in Table 1. Cations ratios are given per  
154 formula unit (p.f.u.).





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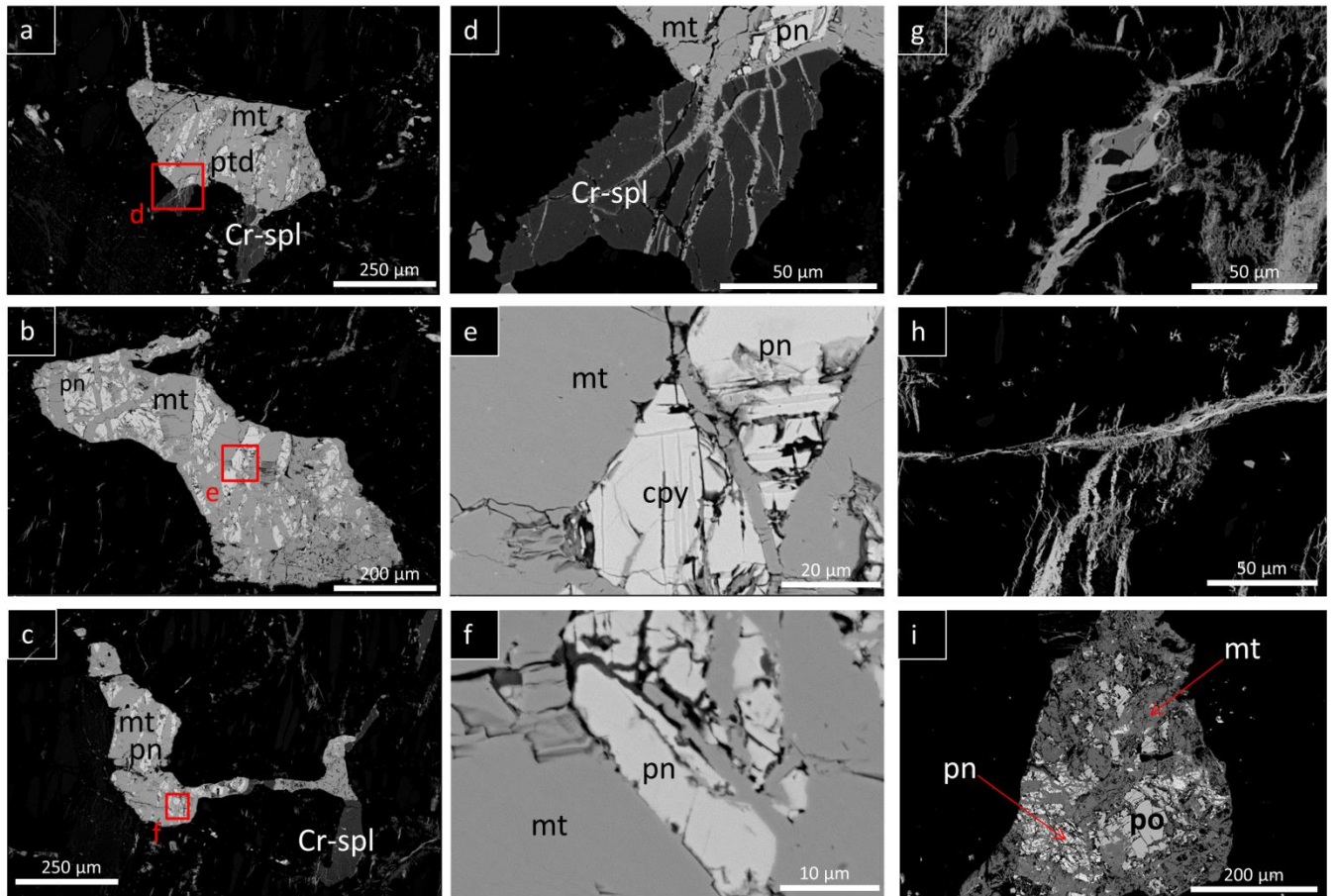
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*Fig. 2. SEM electron backscatter images from sample CS4 (a-h) and element (Fe, Ti, Cr) map from EMP (i). a) Cr-spinel and (d, g) close up images on the same grain. Cr-spinel is in dark gray, and ferrichromite exsolution blebs are light gray. b) Cr-spinel grain and (e, h) close up images on the same grain. Larger ferrichromite intergrowths occur on the margins at the contacts, or along fractures of the hosting grain. c) Cr-spinel grain and (f) close up images on the same grain. i) Elements map for selected area (red box in Fig. 2e) showing ferrichromite (pink) and areas enriched in Ti, enclosed within a matrix of a Cr- and Al rich spinel (blue).*



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*Fig. 3. SEM electron backscatter images from sample 15S2D. a) Magnetite (mt), pentlandite (pn) and Cr-spl assemblage and close up images on the grain (d). b) Magnetite, pentlandite assemblage and close up on the grain (e). c) Magnetite, pentlandite and Cr-spl assemblage and close up on the grain (f). g) Magnetite and Cr-spl in serpentine vein. h) Magnetite in vein. i) Magnetite (dark gray), pentlandite (light gray) and pyrrhotite (po) (medium gray) assemblage.*

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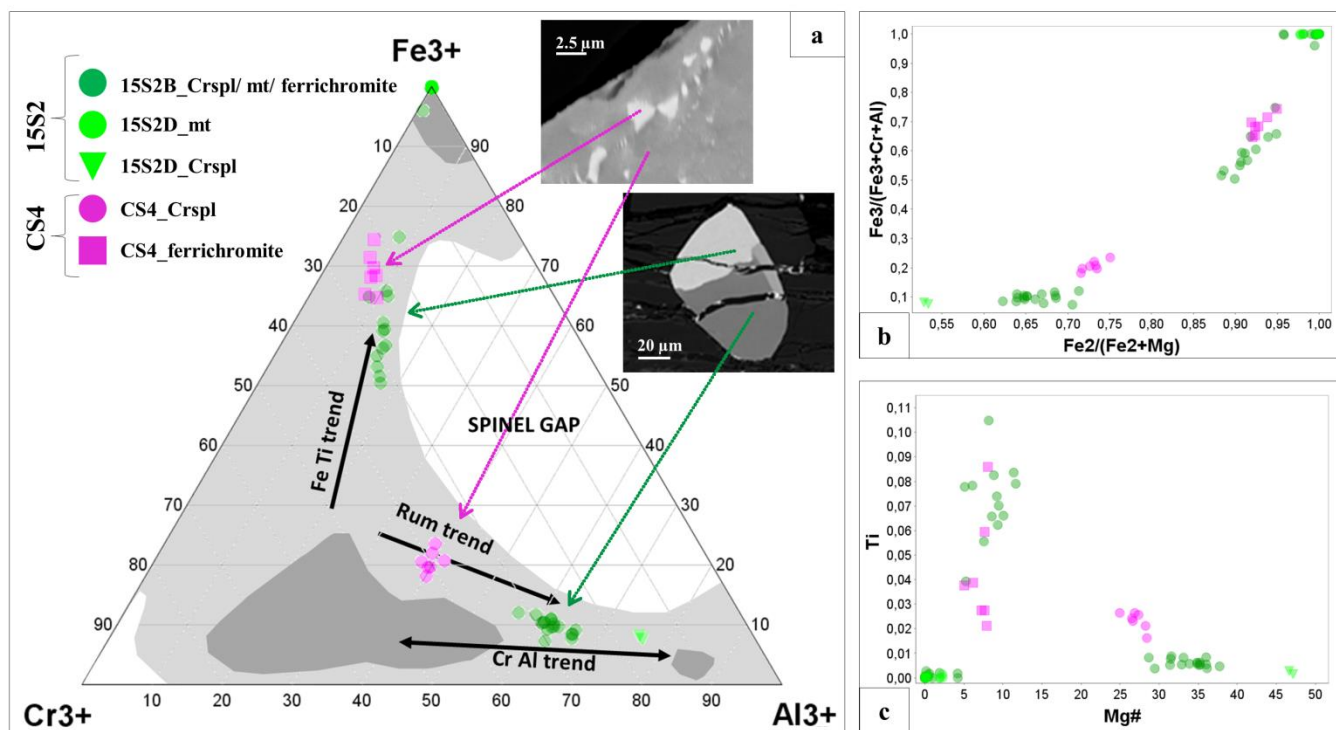
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The larger oxides grains in the pristine dunite sample (CS4) contain two phases (Fig. 2). The host grain consists of Cr-spinel with modal composition  $(\text{Fe}^{2+}_{0.76} \text{Mg}_{0.25})_{1.01} (\text{Fe}^{3+}_{0.46} \text{Cr}_{0.74} \text{Al}_{0.75} \text{Ti}_{0.03})_{1.97} \text{O}_4$ , while the exsolved phase is ferrichromite with modal composition  $(\text{Fe}^{2+}_{0.94} \text{Mg}_{0.07} \text{Ni}_{0.01})_{1.02} (\text{Fe}^{3+}_{1.33} \text{Cr}_{0.46} \text{Al}_{0.15} \text{Ti}_{0.03})_{1.97} \text{O}_4$ . Magnetite is the predominant Fe-Ti oxide in the serpentinized samples, and occurs in a wide range of grain sizes, varying from  $< 1 \mu\text{m}$  to  $> 400 \mu\text{m}$ . Cr-spinel is also present, however, it is more enriched in Al and Mg, and less enriched in Fe and Cr than in CS4, with modal composition  $(\text{Fe}^{2+}_{0.53} \text{Mg}_{0.46})_{0.99} (\text{Fe}^{3+}_{0.15} \text{Cr}_{0.32} \text{Al}_{1.52})_{1.99} \text{O}_4$ . The Cr-rich spinel

175 grains in the serpentinized samples do not contain exsolution microstructures observed in the dunite sample, CS4  
176 (see Fig. 2).

177 Spinel compositions from the dunite and serpentinite samples are shown on three plots in Fig. 4. In both dunite  
178 and serpentinized samples there is a bimodal distribution with a slight shift of the serpentinized samples' spinels  
179 (green symbols, Fig. 4a) towards higher Al content with respect to those from CS4 (pink) (Fig. 4a). Ti is present  
180 in only small amounts (0.0–0.33 p.f.u.); however spinels with higher Fe<sup>3+</sup> content or Mg# in the range 5–12 (Fig.  
181 4c), have more variable Ti content. A ternary plot of trivalent cations for the samples analyzed is shown  
182 in Fig. 4 and data are compared with the mafic and ultramafic spinel compositional fields from Barnes and Roeder  
183 (2001). The host and exsolved phases in the pristine dunite sample and the co-existing spinels in the serpentinite  
184 samples are separated due to the miscibility gap (“Spinel Gap” in Fig. 4a) in the solid solution. Data from our  
185 study plot along the three trends indicated as Cr\Al, Fe\Ti and Rum trends. According to Barnes and Roeder  
186 (2001) the Cr\Al trend is the result of equilibria between Al-bearing pyroxenes and Mg-rich spinels. Among the  
187 Fe-poor spinel grains, we observe a general enrichment in Al in spinels from the serpentinized sample with  
188 respect to those from the pristine dunite. This difference could be related to primary local heterogeneity in the  
189 melts, or to reaction of the primitive spinels with the silicates during later serpentinization. The formation of  
190 lizardite at the expense of diopside would cause an increased Al/Mg ratio in the fluid. The increased mobility of  
191 Fe, generally observed after serpentinization, can further favor iron oxide production and particularly the  
192 crystallization of magnetite (top corner in Fig. 4a and top right corner in Fig. 4b), the endpoint of the Fe\Ti trend.  
193 Barnes and Roeder (2001) attribute the Fe-Ti trend to either the evolution of spinel compositions during fractional  
194 crystallization of silicates with consequent increase of Fe/Mg ratio and Ti content of the melt (Fig. 4c), or to the  
195 exchange of Fe<sup>2+</sup> and Mg between spinel and coexisting silicates. In Fig. 4a spinels from both the dunite and  
196 serpentinized samples plot along this trend and show variable Ti contents (Fig. 4c). The ferrichromite lamellae  
197 (pink squares), previously shown in Fig. 2 are hosted by the Fe-poor spinels. The CS4 spinels are well grouped  
198 and have similar composition with Mg# between 24 and 28. The spinels from the serpentinized sample have a  
199 larger Mg# range and slightly lower Ti content. These differences result in an overlap with the Rum trend. This  
200 trend describes an increase in Al at the expense of Cr with decreasing Fe<sup>3+</sup>, and has been attributed to the reaction

201 between cumulus spinel grains and intercumulus liquid (Barnes and Roeder, 2001). We do not exclude that during  
 202 serpentinization primary spinels have incorporated Al at the expense of  $Fe^{3+}$ , which has been accommodated  
 203 instead the newly formed iron oxides.



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205 *Fig.4. Spinel compositional data for CS4 (pink) and 15S2 (green). a) Ternary plot of trivalent cations (atomic*  
 206  *$Cr^{3+}$ ,  $Fe^{3+}$ ,  $Al^{3+}$ ) with spinels compositional fields (different gray colors on the background are for different*  
 207 *data concentrations) and trends (black arrows) from Barnes and Roeder (2001). Illustrative SEM backscatter*  
 208 *images for spinels compositions from CS4 and 15S2 samples. (b)  $Fe^{2+}/(Fe^{2+} + Mg)$  versus  $Fe^{3+}/(Fe^{3+} + Cr + Al)$ .*  
 209 *d) Ti versus Mg# ( $Mg/(Mg + Fe^{2+})$ ).*

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211 **Table 1. Representative Wavelength-dispersive (WDS) chemical compositions of spinels in pristine dunite**  
 212 **(CS4) and serpentized (15S2D) samples. Values in parentheses correspond to 1 $\sigma$  standard deviation; n is the**  
 213 **number of analyses. Detection limits ranges on analyzed elements are indicated for each group of analyses.**

Sample	CS4				15S2D			
	n = 5		n = 5		n = 6		n = 2	
TiO <sub>2</sub>	0.97	(0.17)	1.81	(0.83)	0.01	(0.01)	0.10	(0.04)
Al <sub>2</sub> O <sub>3</sub>	20.22	(0.77)	2.79	(0.58)	0.00	(0.01)	45.63	(0.24)
Cr <sub>2</sub> O <sub>3</sub>	30.13	(1.70)	15.63	(1.43)	0.00	(0.01)	14.51	(0.11)
Fe <sub>2</sub> O <sub>3</sub>	16.79	(1.42)	48.11	(2.99)	70.25	(0.31)	7.37	(0.44)
FeO	27.39	(0.37)	31.37	(0.26)	31.53	(0.18)	22.38	(0.11)
MnO	0.28	(0.04)	0.21	(0.03)	0.03	(0.03)	0.17	(0.02)
MgO	5.64	(0.29)	1.31	(0.25)	0.02	(0.02)	11.09	(0.11)
NiO	0.08	(0.05)	0.21	(0.05)	0.08	(0.06)	0.23	(0.01)
<b>Total</b>	<b>101.50</b>		<b>101.44</b>		<b>101.94</b>		<b>101.46</b>	
<b>Normalized to 3 Cations and 4 Oxygens</b>								
Ti	0.023	(0.005)	0.050	(0.023)	0.000	(0.000)	0.002	(0.001)
Al	0.778	(0.026)	0.121	(0.025)	0.000	(0.000)	1.516	(0.010)
Cr	0.763	(0.040)	0.452	(0.040)	0.000	(0.000)	0.323	(0.002)
Fe <sup>3+</sup>	0.412	(0.043)	1.325	(0.080)	1.996	(0.002)	0.156	(0.009)
Fe <sup>2+</sup>	0.742	(0.016)	0.960	(0.013)	0.996	(0.003)	0.527	(0.003)
Mn	0.008	(0.001)	0.007	(0.001)	0.001	(0.001)	0.004	(0.001)
Mg	0.271	(0.014)	0.072	(0.013)	0.001	(0.001)	0.466	(0.004)
Ni	0.003	(0.001)	0.006	(0.001)	0.003	(0.002)	0.005	(0.000)
<b>Total</b>	<b>3.000</b>		<b>2.993</b>		<b>2.997</b>		<b>2.999</b>	
<b>Detection limits ranges (ppm)</b>								
Ti	310-353		331-367		343-407		291-320	
Al	143-162		141-153		141-155		170-173	
Cr	381-410		380-412		408-423		334-350	
Fe	373-399		413-457		450-468		351-351	
Mn	411-469		392-420		362-403		338-368	
Mg	136-178		132-158		118-153		169-167	
Ni	407-433		424-473		439-468		367-402	

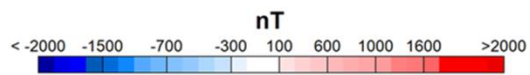
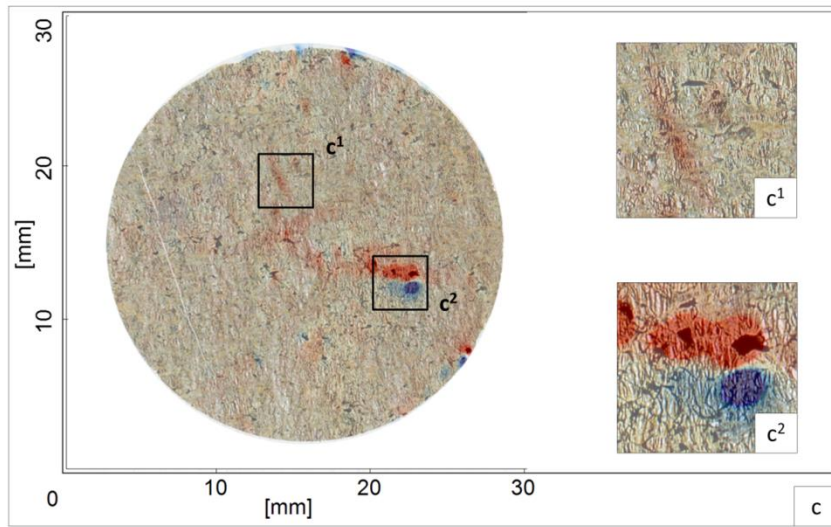
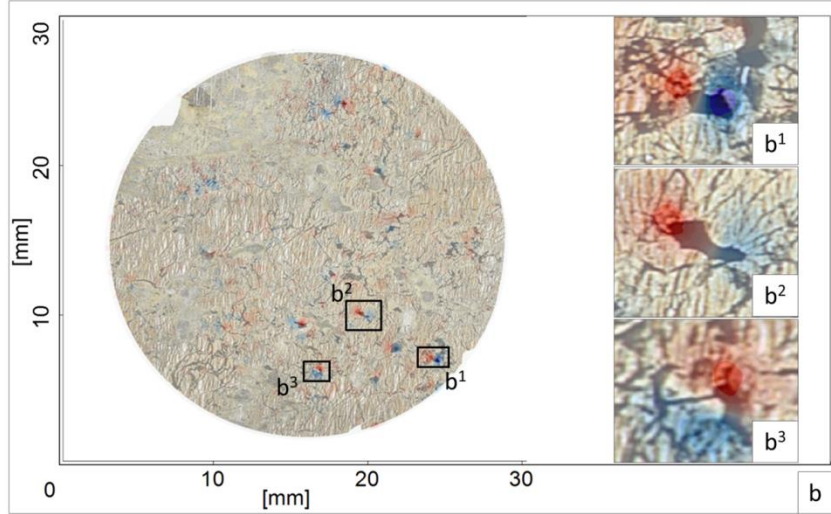
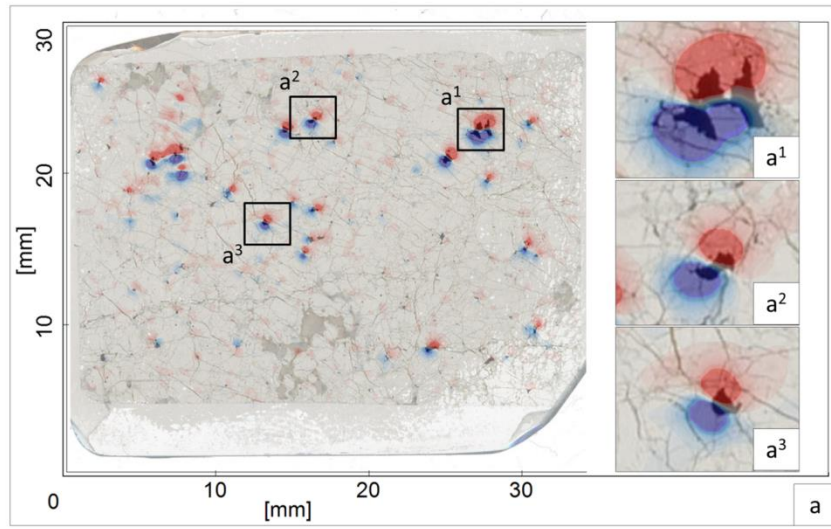
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### 3.2. SCANNING MAGNETIC MICROSCOPY

215 The magnetic scans of pristine sample CS4 and serpentized 15S2D were acquired with a scanning SQUID  
 216 microscope at the Geological Survey of Japan (GSJ) (Oda et al., 2016). The third scan (15S2B, serpentinite) was

217 made at the NTNU laboratory of rock magnetism and paleomagnetism using a scanning magnetic microscope  
218 equipped with a magnetic tunnel junction sensor (Church and McEnroe, 2018).The thickness of the thin sections  
219 is 30 $\mu$ m.

220 Aligned overlays of optical and magnetic scans of the thin sections are shown in Fig. 5. All scans measure the  
221 vertical component of the magnetic field and are shown on the same color scale. The measured field intensity  
222 ranges from -7500 nT to +8100 nT for CS4, from -2200 nT to +2300 nT for 15S2D and from -5000 to +9900 nT  
223 for 15S2B.In each scan several isolated magnetic anomalies related to discrete opaque mineralogy are observed.  
224 In the dunite sample (CS4) the anomalies are commonly dipolar, in the plane of the sample, and oriented NE-SW  
225 (Fig. 5a). Slight variations in the directions are likely related to the shape and orientation of the grain. Most of the  
226 anomalies correlate with grains with an average surface of 300 $\mu$ m\*300 $\mu$ m. With thin section thicknesses of 30 $\mu$ m,  
227 this results in a high aspect ratio of the grains ( $\approx$ 10) which could influence the direction of magnetization. In  
228 15S2D the anomalies are dipolar and in the plane of the sample, but are more randomly oriented compared to  
229 those in CS4. The amplitude of the anomalies is comparable, or lower than those observed in CS4. The 15S2B  
230 scan is dominated by three high-intensity anomalies that correlate with large (> 200 $\mu$ m diameter) grains of  
231 magnetite with pentlandite and chalcopyrite, two at the edge of the sample and one in the SE quadrant (Fig. 5c).  
232 Weaker and elongated anomalies are also present and correlate with the diffuse, fine magnetite in the  
233 serpentized veins. The anomalies in the 15S2B thin section fall into two similar groups. The rounded, high  
234 intensity, in-plane signals are correlated with larger opaque grains and are similar to the stronger anomalies  
235 observed in 15S2D. The weaker, elongated anomalies correlate with the fine magnetite in serpentine veins. The  
236 two types of anomalies have two different directions, approximately 90° from each other (Fig. 5c) with the  
237 magnetite in the serpentine veins producing a signal approximately normal to the plane of the sample.



239 *Fig. 5. Overlays of the magnetic anomaly maps with the optical scans for (a) CS4, (b)15S2D and (c)15S2B.*  
240 *Shown in insets to the right are enlarged views of selected anomalies. Figs. 2 and 3 show high-magnification*  
241 *SEM backscatter images of the same opaque grains causing the anomalies in  $a^1$ ,  $a^2$ ,  $a^3$  (CS4) and  $b^1$ ,  $b^2$ ,  $b^3$*   
242 *(15S2D) respectively. All magnetic scans measure the vertical field and are displayed with the same color scale*  
243 *in nT*

### 3.3. ROCKS MAGNETIC AND PHYSICAL PROPERTIES

244 The concentration, composition, grain size, shape, and intergrowths or microstructures of magnetic minerals  
245 strongly control rock magnetic properties and particularly their ability to retain a stable magnetic memory over a  
246 geological time span (McEnroe et al., 2009a, 2009b; Robinson et al., 2016). Magnetically-ordered phases that  
247 possess spontaneous magnetization and are able to carry remanent magnetization are iron nickel alloys,  
248 uncommon in crustal rocks, or iron oxides and iron sulfides (monoclinic pyrrhotite). For the iron oxides in the  
249 chromite-magnetite solid solution the balance between Fe and Cr (or other cations) controls the magnetism of and  
250 affects intrinsic properties such as the Curie temperature (Robbins et al., 1971). Increasing substitution of cations  
251 such as Ti, Cr, Al and V in magnetite at the expense of Fe lowers the Curie temperature, although in whole-rock  
252 samples minor substitution does not necessarily produce weaker magnetic low-field susceptibility, or  
253 thermoremanent magnetization intensity, which in turn are more affected by grain size than composition (Clark,  
254 1997). Here we investigated magnetic properties of the samples and estimate the Curie temperature,  
255 concentration, and grain size of the magnetic carriers using established magnetic methods, described below.

#### 3.3.1 NATURAL REMANENT MAGNETIZATION, MAGNETIC SUSCEPTIBILITY AND DENSITY

256 The samples show distinctly different bulk magnetic and density properties, due to the effect of serpentinization  
257 on diverse petrophysical parameters (Table 2). The NRM varies between 0.6 and 1.0 A/m in the pristine sample  
258 and between 2.7 and 6.9 A/m in the serpentinite. The magnetic susceptibilities of the serpentinite samples are 1-2  
259 orders of magnitude higher than the dunite, consistent with the production of magnetite during serpentinization.  
260 Densities are significantly lower in the 15S2D and 15S2B cores (average density of  $2.8 \text{ g/cm}^3$ ) with respect to the



261 CS4 core (3.4 g/cm<sup>3</sup>). Most of the primary olivine and pyroxenes in the serpentinized samples have been replaced  
 262 by lizardite or brucite, which have a lower density.

### 3.3.2 HYSTERESIS PARAMETERS

263 Room temperature hysteresis behavior is primarily controlled by the magnetic mineralogy and the domain state.  
 264 The characteristic quantities calculated from the hysteresis loop are saturation magnetization Ms, the remanent  
 265 magnetization Mr, and the coercivity field Hc; the coercivity of remanence Hcr is measured using a separate  
 266 remanence analysis. Table 2 summarizes hysteresis parameters for the samples, companion specimens were used  
 267 for the high-temperature measurements and their properties are summarized in Table A1.

268 *Table 2. Samples densities and magnetic properties (left side of the table), and hysteresis measurements on*  
 269 *chips of the samples (right side of the table). Volume percent magnetite is calculated by dividing the magnetic*  
 270 *susceptibility (k) by 0.0347 (Clark, 1997) and is calculated from saturation moment by dividing the Ms of the*  
 271 *sample by the product of Ms of pure magnetite (480,000 A/m), sample mass and sample density. The*  
 272 *Königsberger ratio (Q) is the ratio between NRM and induced magnetization which is calculated multiplying*  
 273 *the susceptibility and the local magnetic field (43.0012 A/m). The precision used in the table is significant.*

SAMPLE		Density [g/cm <sup>3</sup> ]	Susceptibility [SI]	NRM [A/m]	Q	Volume % magnetite from susceptibility	Chip Mass [g]	Mr [Am <sup>2</sup> /kg]	NRM/Mr [%]	Ms [Am <sup>2</sup> /kg]	Hc	Hcr	Mr/Ms	Hcr/Hc	Volume % magnetite from Ms
CS4	C1 Chip	3.3	0.006	0.7	2.6	0.2	2.73	0.03	0.75	0.12	16.3	37.4	0.22	2.30	0.1
	C2 Chip	3.4	0.006	0.8	3.0	0.2	0.98	0.03	0.89	0.14	13.9	35.9	0.18	2.58	0.1
	C3 Chip	3.3	0.007	1.0	3.1	0.2	0.40	0.03	0.90	0.16	13.5	35.0	0.20	2.60	0.1
	CS4 Core	3.4	0.005	0.6	3.1	0.1									
15S2	D1 Chip	2.8	0.089	5.4	1.4	2.6	0.53	2.01	0.10	4.89	30.2	41.7	0.41	1.38	2.8
	D2 Chip	2.6	0.101	5.0	1.2	2.9	0.82	2.14	0.09	5.55	29.7	41.5	0.39	1.40	3.0
	D3 Chip	2.7	0.085	4.7	1.3	2.5	0.86	1.58	0.11	4.40	31.2	45.7	0.36	1.47	2.5
	D Core	2.8	0.048	4.8	2.3	1.4									
	B1 Chip	2.5	0.083	5.9	1.7	2.4	1.31	5.07	0.05	12.68	29.4	41.9	0.40	1.42	2.7
	B2 Chip	2.6	0.067	2.7	0.9	1.9	0.37	1.56	0.06	5.44	27.3	38.0	0.29	1.39	3.0
	B Core	2.8	0.049	6.9	3.2	1.4									

274

275

276 Ms is directly proportional to the magnetic content of the samples and is used to determine the volume magnetite  
277 content in the samples. This is estimated to be 2.5–3.0% in the serpentinized samples and 0.1% in the dunite. For  
278 the latter, 0.1% is a minimum estimate of the actual volume of ferromagnetic minerals within the sample, because  
279 the Ms of magnetite used in the calculation is higher than the Ms of cation-substituted compositions, in the  
280 magnetite-chromite solid solution. Therefore there could be slightly higher modal amount of ferromagnetic  
281 minerals. The electron microprobe analyses did not reveal endmember magnetite but rather ferrichromite with  
282 composition  $x=1.4$  in the system  $\text{Fe}^{2+}\text{Cr}_{2-x}\text{Fe}^{3+}_x\text{O}_4$ . Using this composition and corresponding Ms of 230,000 A/m  
283 (calculated from Robbins et al., 1971), the volume percent of magnetic oxides is approximately 0.2 %, double that  
284 obtained considering Ms of endmember magnetite. The volume percentage of magnetite can alternatively be  
285 calculated as a function of low-field susceptibility, volume % Mgt = volume susceptibility (k)/0.00347, an  
286 empirical estimation (Clark, 1997; Puranen, 1989). Estimates of magnetite content using this method included in  
287 Table 2 are similar to those calculated from Ms. The susceptibility calculation yields lower estimates for the whole  
288 serpentinite cores, which are approximately 30× the size of the individual chip samples. The magnetite estimates  
289 from both methods (0.1–0.2% for dunite CS4, 2.4–3.0% for serpentinites 15S2D and 15S2B) have similar trends  
290 to the estimate made by image analysis of 0.7% opaque minerals in the dunite and 7.0% in the serpentinite.  
291 Because the image analysis measures all opaque grains, including those non-magnetic, it is expected to yield a  
292 higher estimate than the magnetic measurements, however on both cases the ratio between the dunite and  
293 serpentinite is approximately 10×.

### 3.3.3 THERMOMAGNETIC BEHAVIOR

294 The Curie point ( $T_c$ ), is the temperature below which a magnetic ordering generates a net (spontaneous)  
295 magnetization and is a diagnostic tool for identification of magnetic minerals. Above this temperature the material  
296 is purely paramagnetic (Dunlop and Ozdemir, 1997). End-member magnetite is a commonly occurring natural  
297 magnetic oxide and has a distinct Curie point at 580 °C. However within solid solution series, the Curie point  
298 temperature varies over a wide temperature range and can be used to constrain the mineral composition, or

299 oxidation state (Fabian et al., 2013; Kądziałko-Hofmökł et al., 2008; Petersen and Bleil, 1982; Readman and  
300 O'Reilly, 1972). To measure the Curie temperature, samples are heated in a magnetic field and their susceptibility,  
301 or  $M_s$  is measured as a function of temperature. Below are the results of the thermomagnetic experiments.

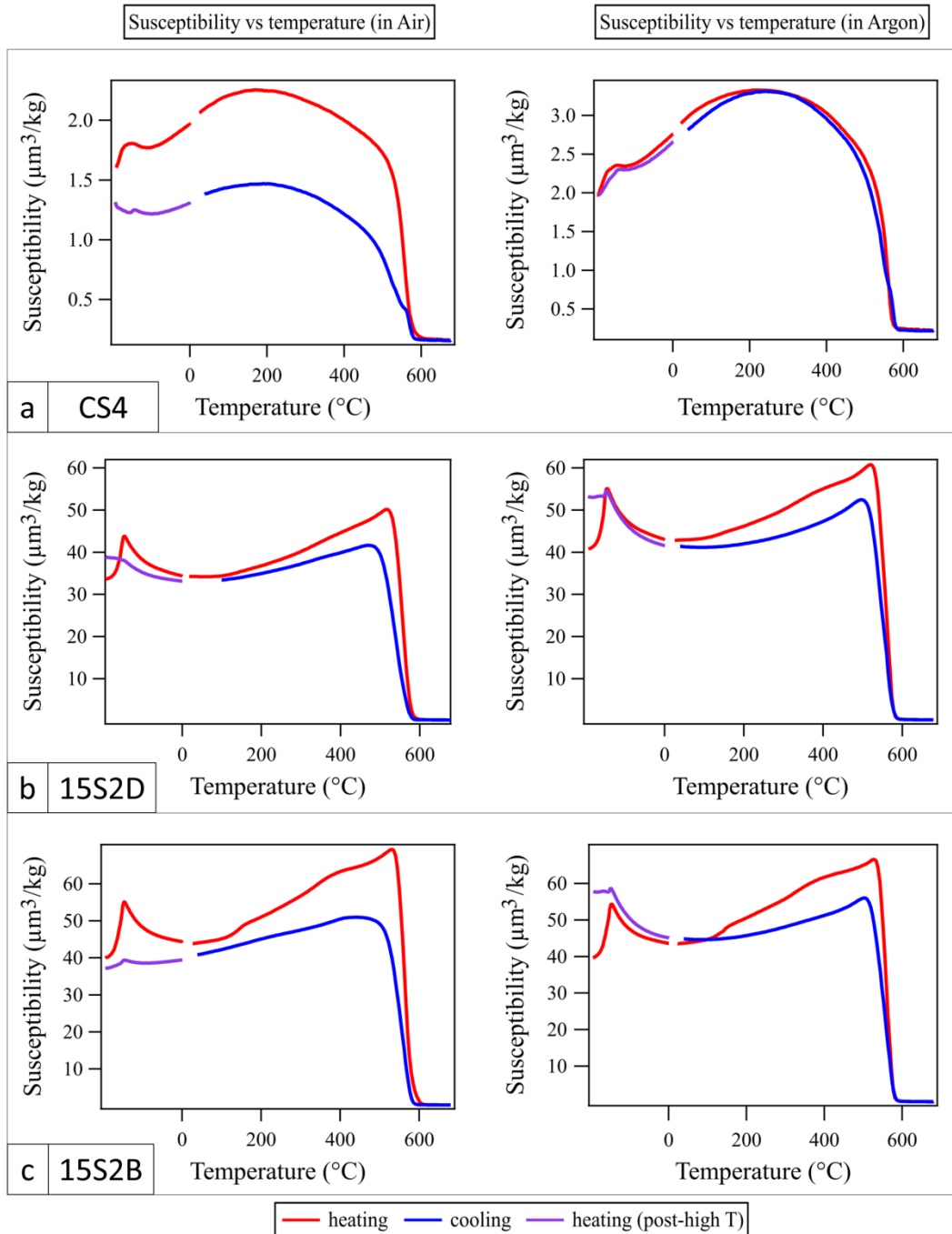
302

### 3.3.3.1 *LOW FIELD SUSCEPTIBILITY*

303 Sample chips of CS4 and 15S2 were powdered for high-temperature susceptibility measurements. Companion  
304 specimens were measured in air and argon, to check for reduction or oxidation during the heating processes on  
305 specimens (Fig. 6). Each plot shows a curve for heating and cooling. Low temperature measurements were made  
306 before and after heating the sample to high temperature as a further check for alteration that may have occurred.  
307 With the exception of the CS4 sample measurement in air (Fig. 6a, left), all other plots show that the susceptibility  
308 at room temperature is similar before and after heating. However, in both argon and air measurements, the  
309 comparison between the cooling and the heating curves suggests an alteration of the sample with a loss in  
310 susceptibility after the heating process. Serpentinized samples show well-defined Hopkinson peak both on heating  
311 and cooling at  $\sim 570$  °C, near the Curie temperature of endmember magnetite. This peak, typical of magnetite and  
312 other magnetic materials, manifests as an increase in magnetic susceptibility between the blocking and the Curie  
313 temperatures and is often indicative of fine-grained particles. The heating branches show a “hump” centered at  
314  $370$  °C in both specimens (Fig. 6b,c) suggesting ferrichromite (Table3). The lower temperature hump between  $150$   
315 and  $170$  °C may be due to the  $\lambda$ -transition from antiferro- to a ferrimagnetic behavior in pyrrhotite (Minyuk et al.,  
316 2013). From optical and electron microscopy pyrrhotite is more abundant in the 15S2B sample with respect to  
317 15S2D, consistent with the more pronounced excursions around  $150$ °C in the latter sample. Horen et al. (2014)  
318 describe a similar hump as consequence of oxidation and destabilization of ferrichromite with a mechanism of  
319 dynamic segregation (Domenichini et al., 2002). This mechanism describes the formation of a new low-  
320 temperature magnetic phase and destruction/rehomogenization of the material within the same heating cycle. This  
321 rehomogenization would explain the irreversibility of the thermomagnetic curve that shows only magnetite in the

322 cooling curve (Fig. 6b, c). Dehydration and/or reduction of hydroxides may also cause variations in the  
323 susceptibility (Funaki et al., 2000) and the irreversibility of the chemical reaction is consistent with the absence of  
324 these variations after heating. The dunite sample (CS4) shows smooth heating and cooling curves, concave down,  
325 with no Hopkinson peak. The maximum susceptibility is reached around 240 °C for runs carried out in argon, and  
326 190 °C for those in air. While the steepest descent in both measurements occurs at 560 °C, indicating abundant,  
327 impure but near-endmember magnetite, the steady decline in susceptibility from ~240 °C is interpreted as a wide  
328 range of Curie temperatures, and hence, compositional variations. The heating and cooling curves are reversible  
329 when measured in argon atmosphere, indicating little change in the magnetic mineralogy during heating.  
330 However, there is clear evidence for mineralogical change during the run in air, as indicated by the irreversibility  
331 of the thermomagnetic curves (Fig. 6a). On the cooling run, this sample exhibits a small inflection near 520 °C,  
332 below the highest Curie temperature. This second Curie temperature suggests a new phase was created during the  
333 experiment, possibly due to cation diffusion between the Fe-poor and Fe-rich regions of the Cr-spinel or the onset  
334 of spinodal decomposition. All samples exhibit the Verwey transition, an abrupt excursion in magnetic  
335 susceptibility near -153 °C (120 K, Walz, 2002, and references therein), which is diagnostic of endmember  
336 magnetite. The transition in the pristine dunite is less sharp and begins at a lower temperature than the literature  
337 value, suggesting that magnetite has some small degree of cation substitution or non-stoichiometry. By contrast,  
338 the Verwey transition in the serpentinite samples (measured before heating) is sharp, indicating a near-endmember  
339 composition. Curie temperature estimates (T<sub>c</sub>), based on representative composition of the magnetic mineralogy  
340 estimated from EMP data and on temperature dependent susceptibility curves, are listed in Tables 3 and  
341 4, respectively. For the T<sub>c</sub> estimates based on composition data we considered magnetite-spinel (Fe<sub>3</sub>O<sub>4</sub>)<sub>x</sub>  
342 (MgAl<sub>2</sub>O<sub>4</sub>)<sub>1-x</sub> with 0 ≤ x ≤ 1 and magnetite-chromite Fe<sub>3-x</sub>Cr<sub>x</sub>O<sub>4</sub> 0 ≤ x ≤ 2 solid solution series (Table 3). For the  
343 magnetite-spinel solid solution series we refer to Harrison and Putnis (1996) where the T<sub>c</sub> varies as a function of  
344 the mole fraction of magnetite (x) approximated by: T<sub>c</sub> (°C) = -853 + 2410x - 970x<sup>2</sup>. For the magnetite-chromite  
345 solid solution series we refer to the combined data sets of Francombe (1957) and Robbins et al. (1971) fit with a  
346 logistic function. For comparison T<sub>c</sub> estimates based on an interpolation between the two solid solution series are  
347 listed in Table 3. In both dunite and serpentinitized samples the Fe-poor spinel compositions are paramagnetic at

348 room temperature. Tc estimates for the ferrichromite vary between 200 °C and 400 °C depending on the cation  
349 ratios. To estimate Tc from the initial susceptibility, we used a derivative method that uses the maximum negative  
350 slope in the thermomagnetic curve (Table 4).For the serpentinized samples these Tc estimates on the heating  
351 curves are in the temperature range of 557–565 °C, whereas the cooling curves show Tc estimates in the range  
352 520 °C–572 °C. This variability may indicate that some non-stoichiometric, non-endmember magnetite is present  
353 and that paramagnetic elements such as Al, Ti, Cr and/or Mg are substituted in the crystal lattice (Dunlop and  
354 Özdemir, 1997), to a greater extent in the dunite than in the serpentinite.



355

356

*Fig. 6. Temperature dependent susceptibility curves in air (to the left) and argon (to the right) for CS4*

357

*(a), 15S2D (b), and 15S2B (c) samples chips. Susceptibility is normalized by the mass of the specimen.*

358

359 *Table 3. Tc estimates from representative oxides compositions, determined considering magnetite-spinel*  
360 *(M-S) (Harrison and Putnis, 1996), magnetite-chromite (M-C) (Francombe, 1957; Robbins et al., 1971)*  
361 *solid solution series and an interpolation between the two solid solution series. Chemical formulas do not*  
362 *account for elements with < 0.01 p.f.u. Pm in the table indicates that the corresponding composition is*  
363 *paramagnetic at all temperatures.*

Sample	Spinel Composition	Tc (°C) from composition		
		(Francombe 1957; Robbins et al. (1971))	(Harrison & Putnis 1996)	(considering a combination of the M-C and M-S solid solutions)
CS4	$(\text{Fe}^{2+}_{0.76} \text{Mg}_{0.25})_{1.01} (\text{Fe}^{3+}_{0.46} \text{Cr}_{0.74} \text{Al}_{0.75} \text{Ti}_{0.03})_{1.97} \text{O}_4$ SPINEL/HERCYNITE	-120	-33	-90
	$(\text{Fe}^{2+}_{0.94} \text{Mg}_{0.07} \text{Ni}_{0.01})_{1.02} (\text{Fe}^{3+}_{1.33} \text{Cr}_{0.46} \text{Al}_{0.15} \text{Ti}_{0.03})_{1.97} \text{O}_4$ FERRICHROMITE	400	417	345
15S2D	$(\text{Fe}^{2+}_{0.53} \text{Mg}_{0.46})_{0.99} (\text{Fe}^{3+}_{0.15} \text{Cr}_{0.32} \text{Al}_{1.52})_{1.99} \text{O}_4$ SPINEL/HERCYNITE	-180	<i>pm</i>	-180
	$(\text{Fe}^{2+}_{0.98} \text{Fe}^{3+}_2)_{2.98} \text{O}_4$ MAGNETITE (LARGE GRAINS)	585	586	580
	$(\text{Fe}^{2+}_1 \text{Mg}_{0.02})_{1.02} (\text{Fe}^{3+}_{1.92})_{1.92} \text{O}_4$ MAGNETITE (VEIN)	580	583	580
15S2B	$(\text{Fe}^{2+}_{0.71} \text{Mg}_{0.29})_{1.00} (\text{Fe}^{3+}_{0.24} \text{Cr}_{0.62} \text{Al}_{1.11})_{1.97} \text{O}_4$ SPINEL/HERCYNITE	-150	-185	-125
	$(\text{Fe}^{2+}_{0.93} \text{Mg}_{0.12})_{1.05} (\text{Fe}^{3+}_{0.94} \text{Cr}_{0.57} \text{Al}_{0.30} \text{Ti}_{0.08} \text{V}_{0.02} \text{Mn}_{0.01})_{1.92} \text{O}_4$ FERRICHROMITE	205	276	196
	$(\text{Fe}^{2+}_{0.97} \text{Mg}_{0.05})_{1.02} (\text{Fe}^{3+}_{1.42} \text{Cr}_{0.33} \text{Al}_{0.15} \text{Ti}_{0.04} \text{V}_{0.01})_{1.92} \text{O}_4$ FERRICHROMITE	440	455	414
	$(\text{Fe}^{2+}_{0.99} \text{Fe}^{3+}_{1.98})_{2.97} \text{O}_4$ MAGNETITE	585	585	580
	$(\text{Fe}^{2+}_1 \text{Fe}^{3+}_2)_3 \text{O}_4$ MAGNETITE (LARGE GRAINS)	585	586	580

364

365

366

367 *Table 4. Tc estimates from thermal experiments. Tc estimates based on high temperature measurements*  
 368 *are determined from susceptibility (k) and Ms versus temperature curves using the derivative method*  
 369 *described in the text. Blocking temperatures (Tb) are derived from Mr. Some samples have multiple Tc*  
 370 *estimates in the cooling curve, which reflect the inception of spinodal decomposition during the*  
 371 *experiment.*

Sample	High-T experiment	Tc (°C) from k		Tc (°C) _Ms	Tb (°C) _Mr		
		Heating	Cooling				
CS4	KB air	556	520 568				
	KB Ar	560	545 572				
	VSM	579				576	560
		574					
15S2D	KB air	557	541				
	KB Ar	559	542 561				
	VSM					567	150 561
						565	156 555
15S2B	KB air	565	562				
	KB Ar	148	544				
		561					
VSM			576	153 567			

372

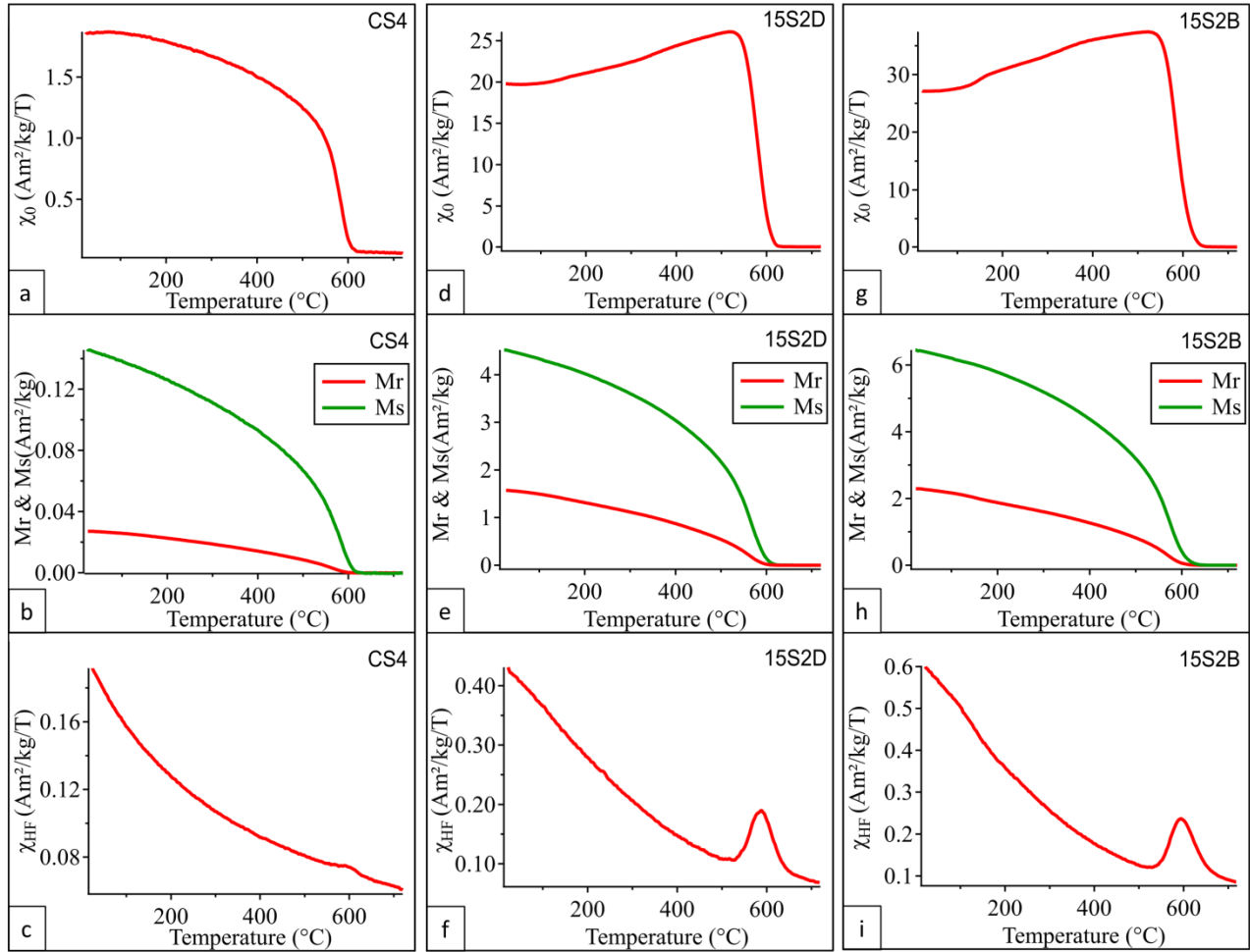
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### 3.3.3.2 HIGH TEMPERATURE VSM MEASUREMENTS

374 An alternative calculation of the Curie temperature and a method to observe the evolution of other informative  
 375 metrics as a function of temperature is provided by high-temperature VSM measurements. The advantage of the



376 VSM is that one measures four parameters ( $M_s$ ,  $M_r$  and low- and high-field susceptibility). In strong-field  
377 measurements, the variation of the saturation magnetization with temperature is rigorous and more accurate  
378 indicator for the Curie temperatures than initial susceptibility and for the identification of ferro-, ferri- and  
379 antiferromagnetic states (Fabian et al., 2013; McEnroe et al., 2016). High-temperature curves of  $M_s$ ,  $M_r$ , high-  
380 field ( $\chi_{HF}$ ), and low-field ( $\chi_0$ ) susceptibility are shown in Fig. 7 for samples CS4 (a-c), 15S2D(d-f) and 15S2B  
381 (g-i) chips. Temperature-dependent initial susceptibility (Fig. 7a, d and g) curves are comparable to those  
382 measured on the Kappabridge (Fig. 6), albeit measured with different protocols. Sample CS4 shows a gradual  
383 decrease in initial susceptibility followed by a rapid decrease in the temperature range 560–580 °C. By contrast  
384 the serpentinized samples (15S2D and 15S2B) show an initial increase in susceptibility followed by a rapid  
385 decrease close to the  $T_c$  of magnetite. The thermal curves of  $M_r$  and  $M_s$  are similar in both the 15S2D and 15S2B  
386 samples;  $M_s$  and  $M_r$  slowly decrease with temperature up to 530°C and at a faster rate above this temperature.  
387 The point of maximum descent in  $M_s$  is a widely-applied technique for estimating the Curie temperature (Tauxe,  
388 1998), though yields a slight underestimation (Fabian et al., 2013), and can be compared to similar calculations  
389 from low-field susceptibility measurements in the Kappabridge.  $T_c$  estimates based on  $M_s$  and unblocking  
390 temperatures from  $M_r$  are summarized in Table 4. In all samples, the  $T_c$  estimated from  $M_s$  is higher than that  
391 estimated from Kappabridge measurements, to a small degree in sample 15S2D and to a larger extent in CS4 and  
392 15S2B. The  $M_r$  curves provide unblocking temperatures, the temperature above which particles have a  
393 spontaneous magnetization but due to thermal activation at high temperatures do not carry a stable remanence.  
394 The serpentinized samples demonstrate weak unblocking behaviour (with consequent loss of remanent  
395 magnetization) near 150 °C. At this temperature there is also a weak enhancement in susceptibility measured in  
396 both Kappabridge (Fig. 6) and VSM (Fig. 7) in sample 15S2B. The high-field (HF) susceptibility curve for 15S2  
397 samples shows a clear Landau peak, slightly above Curie temperature for magnetite indicating ferrimagnetic  
398 ordering. These temperatures of 588 °C for sample 15S2D, and 595 °C for sample 15S2B, are higher than  
399 determinations from  $M_s$ ,  $M_r$  or initial susceptibility, in agreement with Landau theory (Fabian et al., 2013). The  
400 Landau peak is visible, though small in the CS4 high-field susceptibility curve.



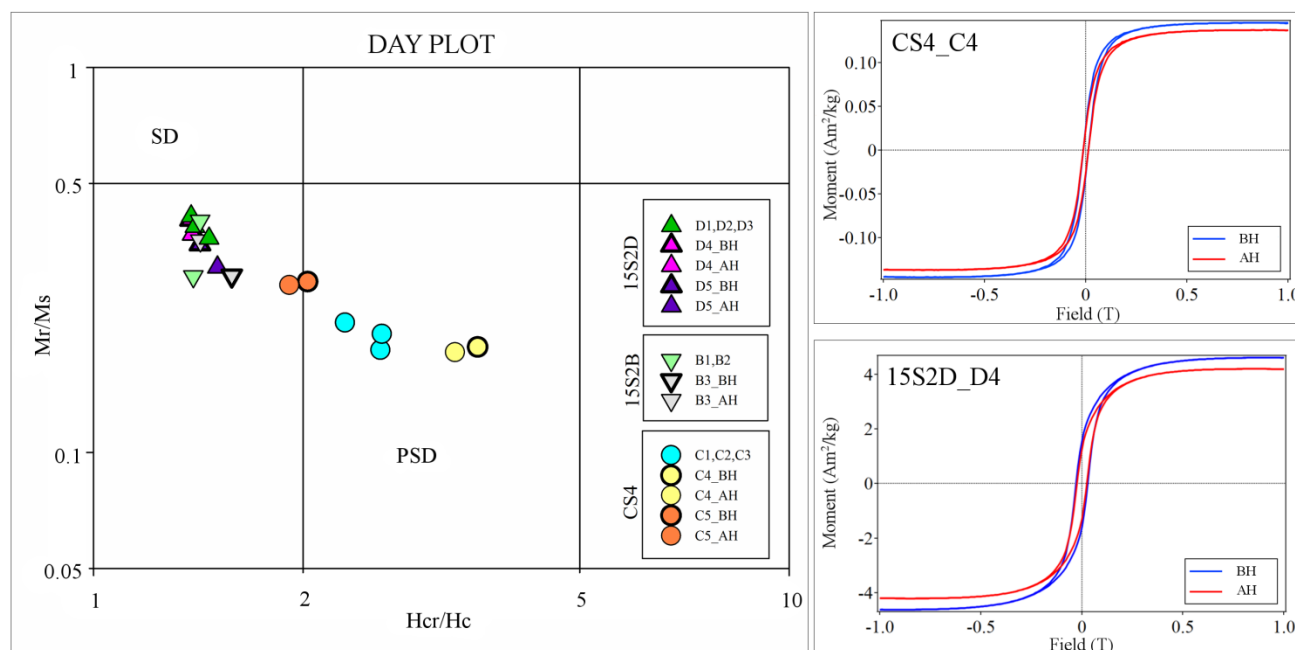
401

402 **Fig. 7. High temperature experiments on 15S2 (15S2D and 15S2B) and CS4 samples chips. (a, d, g) Low-field**  
 403 **susceptibility ( $\chi_0$ ) versus temperature. (b, e, h)  $M_r$  and  $M_s$  versus temperature. (c, f, i) High-field susceptibility**  
 404 **( $\chi_{HF}$ ) versus temperature.**

405

406 Room temperature hysteresis measurements were acquired before and after each thermal experiment, to  
 407 characterize the sample, and to check for alteration that may have occurred during the heating. Representative  
 408 hysteresis loops, measured before and after the high-temperature VSM experiment are shown in Fig. 8. Hysteresis  
 409 loops have been corrected for the paramagnetic contribution to isolate the ferromagnetic response. Both loops  
 410 show a decrease in  $M_s$  after the high-temperature measurement, indicating some limited mineralogical change.  
 411 Comparing the loops before heating, serpentinite sample 15S2D has a wider hysteresis loop than that of the

412 pristine dunite CS4. The wider loop indicates a higher coercivity ( $H_c$ ) in the serpentinized samples likely related to  
 413 the presence of fine-grained magnetite, such as that observed in the veins. Magnetic behavior, particularly  
 414 remanence, is strongly influenced by magnetic domain state, which for magnetite can be estimated on the Day  
 415 plot (Day et al., 1977). The  $M_r/M_s$  and  $H_{cr}/H_c$  ratios before and after heating are shown in Fig. 8, with additional  
 416 measurements of other chips not used for thermomagnetic measurements, and limits of single-domain and  
 417 pseudo-single domain behavior calculated by Dunlop (2002). The serpentinized samples cluster at high  $M_r/M_s$   
 418 and low  $H_{cr}/H_c$  values, approaching ideal single-domain behavior. By contrast, the pristine CS4 samples lie  
 419 outside this cluster with parameters that generally fall within the pseudo-single domain region, which implies that  
 420 at least some magnetic particles that contribute to the bulk properties are larger than single-domain. After the high  
 421 temperature experiments the values of  $M_s$ ,  $M_r$ ,  $H_c$  and  $H_{cr}$  decrease slightly, indicating a change in chemical  
 422 composition, grain size, and/or shape. This change is visible in the hysteresis loops and leads to slight shifts in the  
 423 before-(BH) and after-heating (AH) positions of samples on the Day plot.



424  
 425 **Fig. 8. Left: Day plot of  $M_r/M_s$  versus  $H_{cr}/H_c$  for specimens from the serpentinized samples (15S2D and**  
 426 **15S2B) and the pristine dunite sample (CS4) (after Day et al., 1977; Dunlop, 2002). CS4 sample plots in the**  
 427 **pseudo-single domain (PSD) region while 15S2 samples plot closer to the single-domain (SD) region. These**

428 *parameters are listed in Table 2 and Table A1. BH is before and AH after heating. Right: BH (blue) and AH*  
429 *(red) hysteresis loops for CS4\_C4 (top) and 15S2D\_D4 (bottom).*

### 3.4. MAGNETIC DATA MODELING

430 Scanning magnetic microscopy provides high resolution mapping of magnetic fields above the thin sections. The  
431 technique measures the vertical component of the magnetic field in field-free condition originating from the  
432 sample NRM. The measurements, made at room temperature can be correlated with the bulk magnetic properties  
433 of the sample. The magnetic field varies across the study thin sections, and is used to locate the magnetic carriers.  
434 These were investigated in both magnetic properties and mineral chemistry. The pristine dunite sample has  
435 distinctly different NRM intensity and susceptibility values than the serpentized samples (Table 2). Here, we  
436 used magnetic modeling of these anomalies to characterize NRM directions and intensity of selected discrete  
437 grains.

#### 3.4.1 MAGNETIC DATA AND PROCESSING

438 Isolated magnetic anomalies are observed in all magnetic scans (Fig. 5) associated with the opaque mineralogy.  
439 Here we model three grains, one for each thin section, which correlate with strong dipolar anomalies (Fig. 9a, b  
440 and c). For each grain the magnetic anomaly and analytic signal maps are compared and shown overlaying the  
441 optical scan in Fig. 9. The analytic signal of the magnetic field was calculated to locate and distinguish among  
442 multiple sources contributing to each dipolar anomaly. This field transformation, obtained through a combination  
443 of the horizontal and vertical gradients of the magnetic field, does not depend on the direction of magnetization of  
444 the anomaly source, but depends on its location and shape. Assuming a rectangular, or Cartesian, coordinate  
445 system with the top plane of the thin section in the xy-plane and with the z-axis (measurement direction)  
446 perpendicular to it and directed downward into the thin section, we calculated the analytic signal of the vertical  
447 magnetic field ( $B_z$ ). The amplitude of the analytic signal at (x,y) is calculated as the square root of the sum of the  
448 squares of the derivatives in the x, y, and z directions of the magnetic field  $B_z$  according to (Roest et al., 1992):

449

$$(1) \quad |AS(x,y)| = \sqrt{\left(\frac{\partial Bz}{\partial x}\right)^2 + \left(\frac{\partial Bz}{\partial y}\right)^2 + \left(\frac{\partial Bz}{\partial z}\right)^2}$$

450

Considering the high aspect ratio of our modeled grains, the analytic signal amplitude should have a maximum

451

centered above the magnetic source (Nabighian 1972). A comparison between the analytic signal and grain shapes

452

shows that:

453

- For the grain in Fig. 9d, the analytic signal reflects the shape of the left side of the grain and indicates that

454

region as the main source of the magnetic anomaly.

455

- For the grain in Fig. 9e, the analytic signal is dominantly centered over the grain with smaller highs

456

towards the left side of the grain that could reflect compositional variations.

457

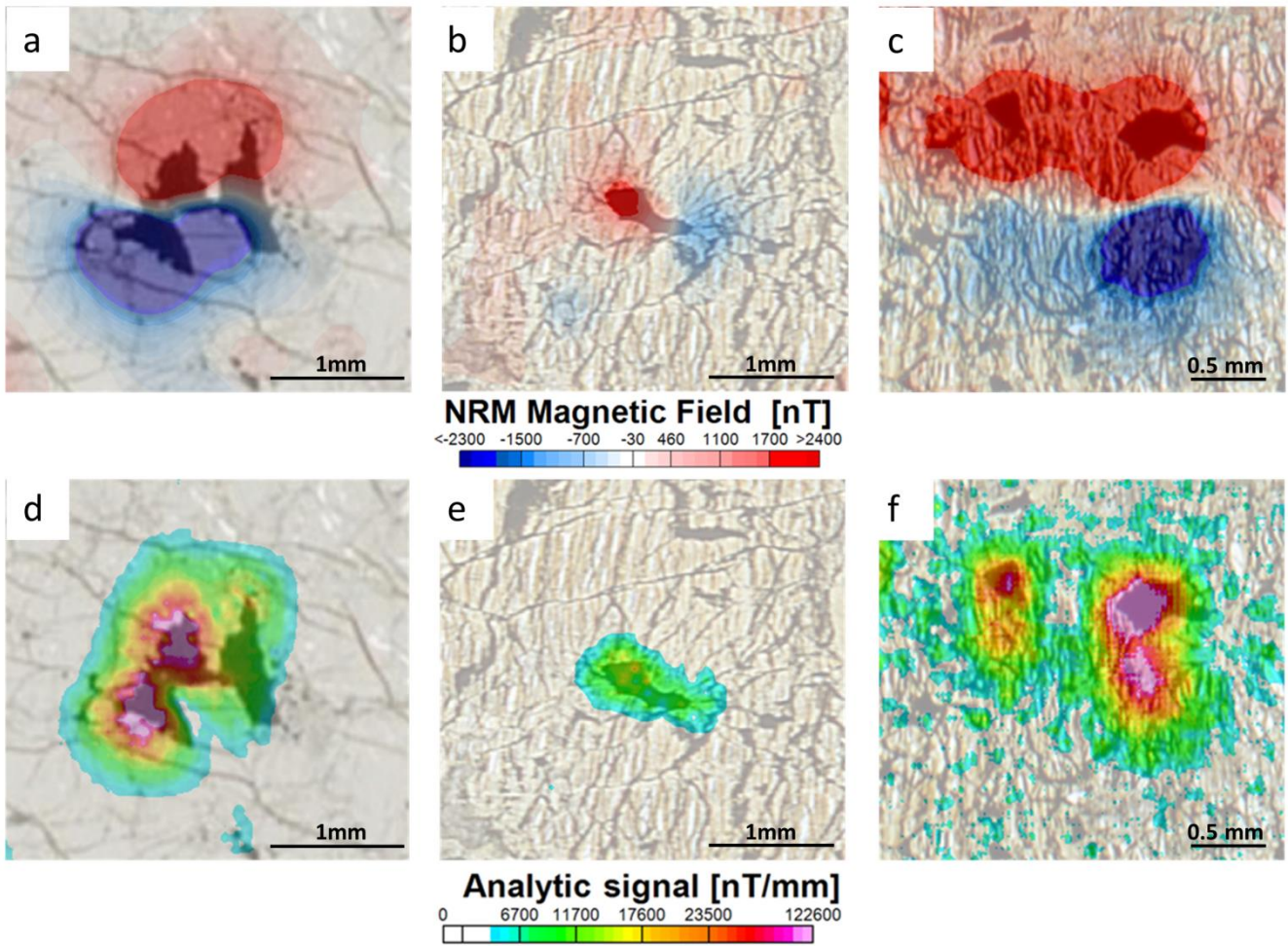
- For the grain in Fig. 9f, there are three different highs in the analytic signal, which can be distinguished,

458

two correlating with large opaque grains, and one below one of the large grains. This may indicate an

459

additional source of fine magnetite within the serpentized veins.



460

461 *Fig. 9. Magnetic anomaly maps (top) and analytic signal maps (bottom) for three selected grains. Displayed*  
 462 *grains are from CS4 (left), 15S2D (middle) and 15S2B (right) thin sections. Grains locations are indicated*  
 463 *in Fig. 5 with black boxes labelled as  $a^1$ ,  $b^2$  and  $c^2$ . These show the respective grain location within each thin*  
 464 *section respectively.*

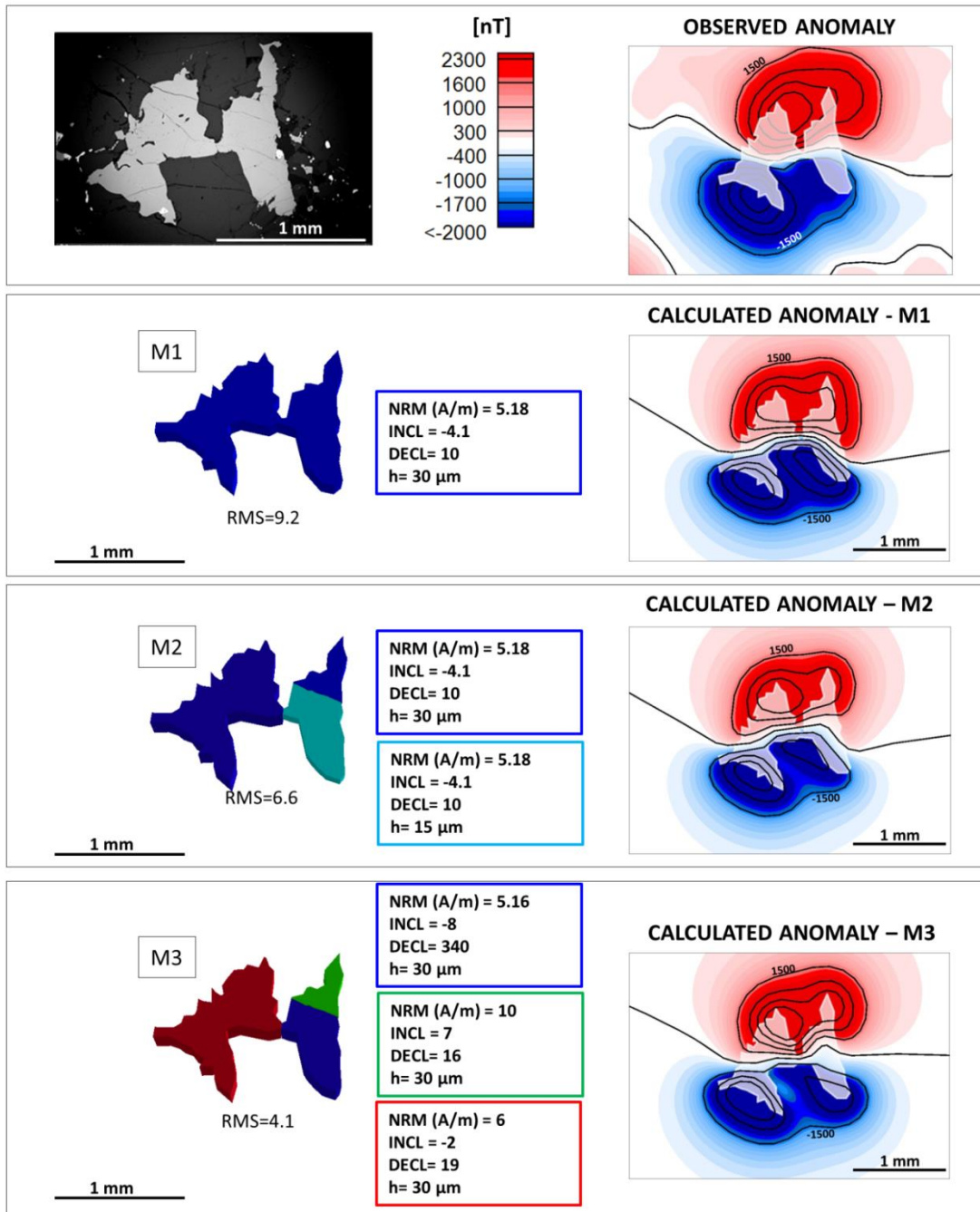
### 3.4.2 REMANENT MAGNETIZATION MODELING AND RESULTS

465 We modeled the NRM intensity and direction of the three grains shown in Fig. 9, using a forward modeling  
 466 approach with Model Vision. Modeling of each grain was made using homogeneously magnetized 3D frustum  
 467 bodies with top and bottom constrained to be horizontal and defined by polygons, or tabular bodies with a  
 468 maximum thickness of 30 $\mu$ m (thin section thickness). Each body was then inverted for magnetization intensity

469 and direction to obtain the best fit between observed and calculated anomaly field. Several tests were made on  
470 selected grains. The best fitting models, for each test, are shown in Figs. 10, 11 and 12, with their respective  
471 modeling parameters. Modeling tests for a Cr-spinel grain from CS4 (Figs. 2a&5a) are shown in Fig. 10. Three  
472 different tests have been run (M1, M2 and M3). The first test considers a single large homogeneously magnetized  
473 grain, whose shape is constrained by the SEM images and by the thin section thickness of 30 $\mu$ m. Inversions for  
474 magnetization intensities and direction gave a percentage root mean square error (RMS) of 9.2 for the best fitting  
475 model (M1). The RMS is expressed as a percentage of the dynamic range of the active data. It is calculated, for all  
476 positions at which the field values are used in the inversion, as the root mean square difference between the model  
477 input and output field values at a specific point, divided by the total range of the input magnetic field data in the  
478 modeled area. The second test was made using two sub-grains with same magnetization intensities and direction  
479 as in the M1 model, and inverting only for vertical extents of the two sub-grains and limiting this extent to a  
480 maximum of 30 $\mu$ m. This test gave a lower RMS error of 6.6; the best fitting model M2 requires a lower thick-ness  
481 for the sub-grain to the right of the larger grain. The third test was made assuming three sub-grains of fixed  
482 thickness (30 $\mu$ m) and freely variable magnetization intensity and directions. The best fitting model gave a RMS  
483 error to the observed anomaly of 4.1 and suggests a variable intensity of the remanent magnetization within the  
484 modeled grain, but broadly similar magnetization directions. For the magnetite grain with an intergrowth of  
485 pentlandite from 15S2D (Fig. 5b) three modeling tests were run and the best fitting models are shown in Fig. 11.  
486 Model M1 assume a homogeneously magnetized grain and gave a RMS error of 13. In model M2 the grain is  
487 subdivided in three smaller grains with same vertical extent but freely variable remanent magnetization intensities  
488 and directions; the model gave a RMS error to the observed anomaly of 8.9. In the last model M3 the modeled  
489 grain is subdivided in multiple tabular bodies, each one homogeneously magnetized, which are inverted to obtain  
490 the best fit between the observed and the calculated anomaly. The model M3 gave a RMS error of 3 and similar to  
491 the grain in CS4 suggests multiple sources of magnetization within the larger grains with higher intensities on the  
492 right side of the grain. This can be explained by variations in the amount of magnetite versus pentlandite. The  
493 stereoplot in Fig. 11 shows the magnetization direction of each sub-grain, which indicates extremely variable  
494 magnetization within the composite particle. Two grains from the serpentinized thin section 15S2B were modeled

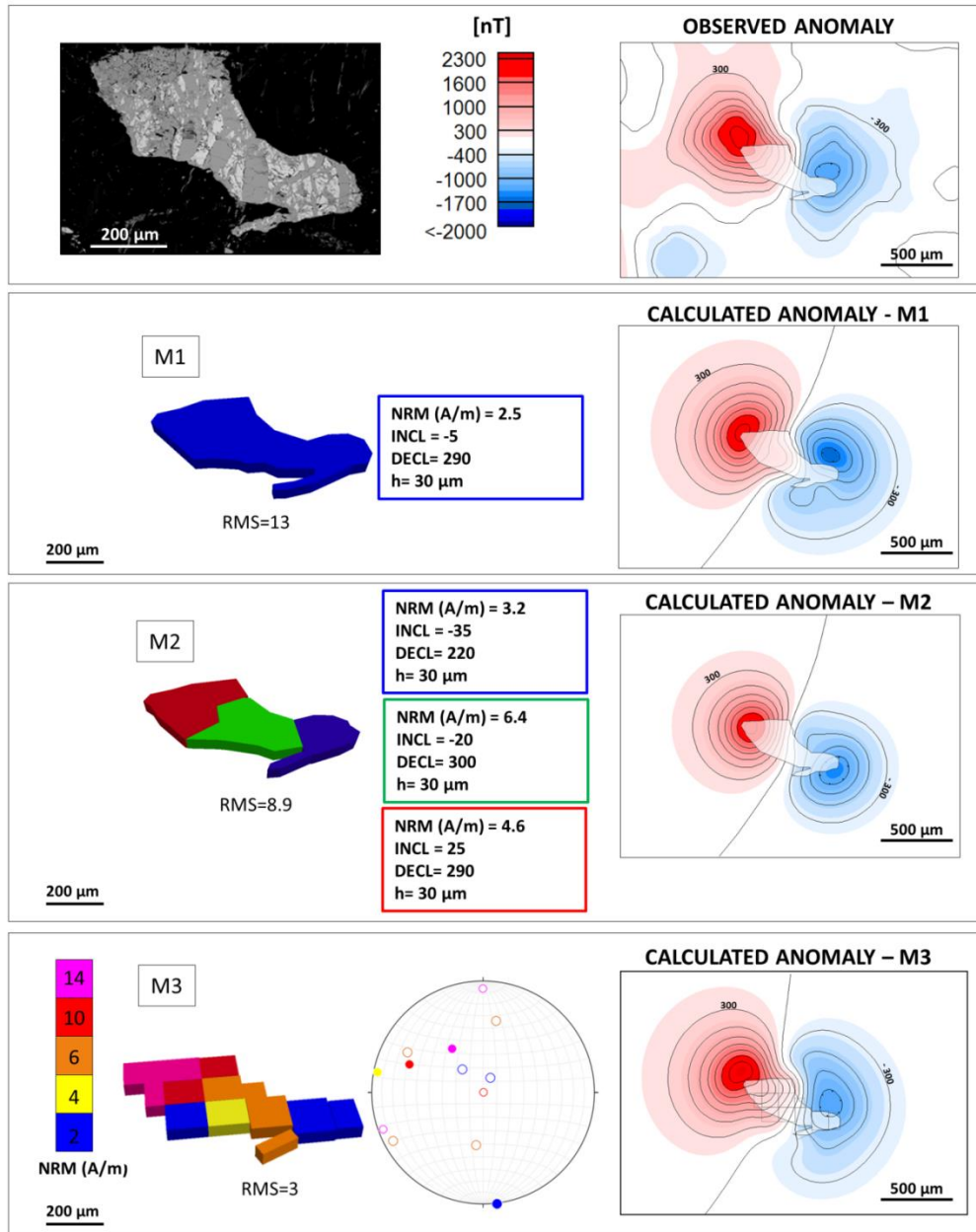
495 (Fig. 12). The opaque grains are magnetite intergrown with pentlandite, pyrrhotite, and minor chalcopyrite. Two  
496 tests were performed: the first test assumes the two grains are homogeneously magnetized, and that the two larger  
497 grains are the main source of the observed anomalies. The second test (analogously to the M3 in the previous  
498 modeled grains) inverts a set of tabular bodies for remanent magnetization direction and intensity. The best fitting  
499 model in the first test gave a RMS error of 14 with a notable mismatch on the right side of the modeled area. This  
500 test suggests that an important source of magnetization is located below the large grain to the right of the modeled  
501 area, which is reversely magnetized (black box in Fig. 12). In reflected light, at high magnification, smaller  
502 magnetite grains (up to 10 $\mu$ m) are visible on the surface of the thin section, and additional particles are likely  
503 below the surface of the sample. This area has been modeled in the second test by means of tabular bodies, which  
504 gave a localized NRM intensity of 12 A/m, and steep negative inclination and an improved RMS error of 5.  
505 Modeling of the magnetic anomalies over the three isolated grains in the thin section indicates heterogeneous  
506 sources of remanent magnetization with intensities varying between 2 and 12 A/m and variable directions. In the  
507 CS4 sample this variability may be associated with a variable amount of small ferrichromite exsolved within the  
508 hosting Cr- spinel. In the two serpentized thin sections generally weaker NRMs correlate with a larger amount  
509 of pentlandite versus magnetite. In these samples the magnetization was acquired when magnetite was produced  
510 during serpentization at lower temperature than the blocking temperature. The variable direction of the NRMs  
511 may also reflect the multidomain behavior expected for such large grains, which results in a less efficient  
512 acquisition of the magnetization.





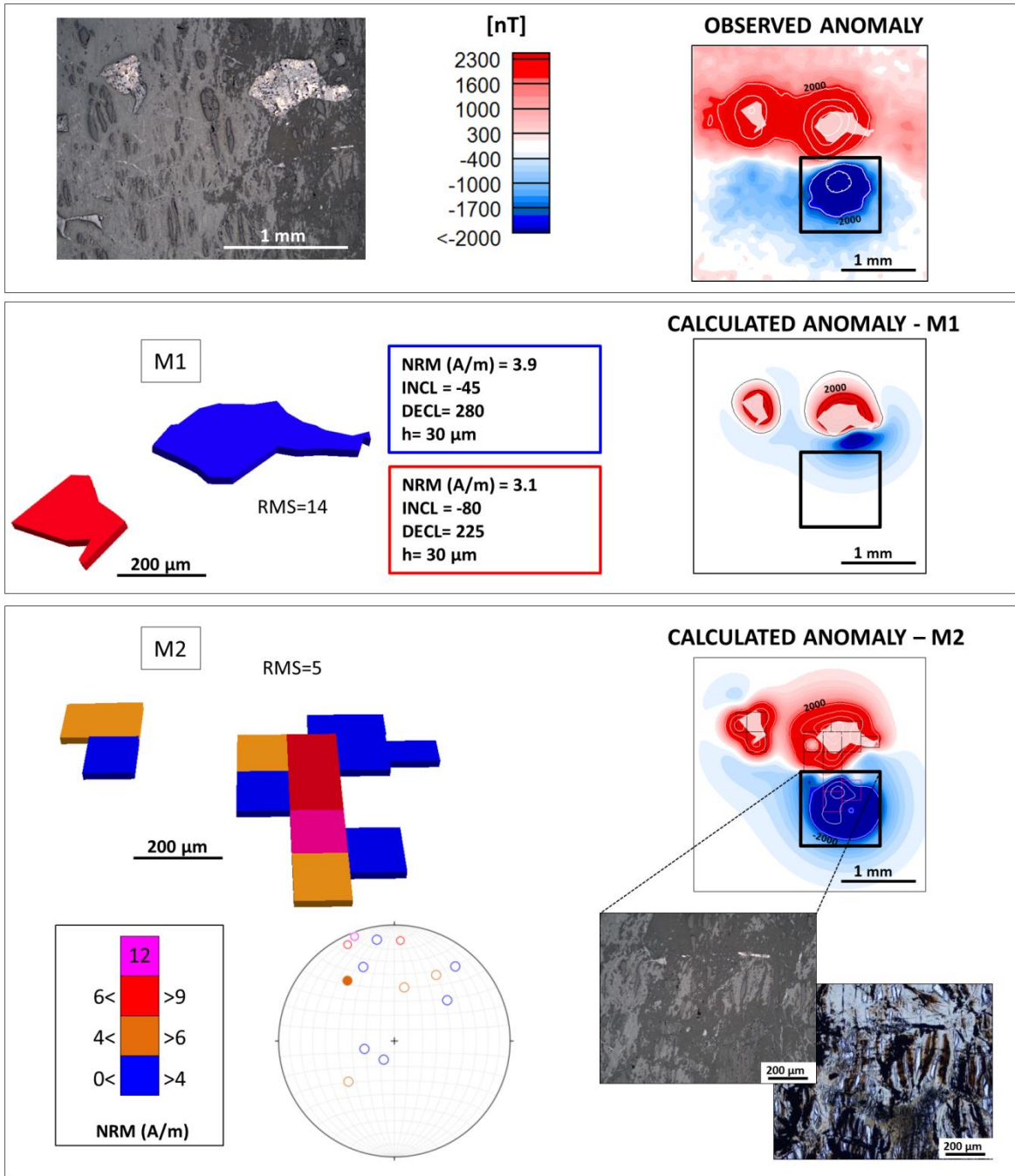
513

514 *Fig. 10. NRM modeling tests for CS4 grain. Top panel: SEM image of the modeled Cr-spinel grain and*  
 515 *correlative magnetic anomaly. Below for each test are: 3D image of the grain color coded by number of bodies*  
 516 *used in the modeling (left) and respective modeling parameters (center) and resulting calculated anomaly*  
 517 *(right) with contours interval of 1500 nT. RMS =root means square error between calculated and observed*  
 518 *anomaly grids.*



519

520 *Fig. 11. NRM modeling test results for 15S2D grain. Top an SEM image of the modeled grain. For each test*  
 521 *there is a panel with 3D image of the grain color coded by number of bodies used in the modeling (left) and*  
 522 *respective modeling parameters (center) and resulting calculated anomaly (right) with contours interval of 300*  
 523 *nT. RMS = root means square error between calculated and observed anomaly grids. The stereoplot in the M3*  
 524 *model is for NRM directions of the modeled tabular bodies; closed circles are for positive inclinations and open*  
 525 *circles for negative inclinations, colors are for NRMs intensities.*



526

527 *Fig. 12. NRM modeling tests for 15S2B grain. Above are reflected light image of the modeled grains and*  
 528 *measured anomaly. Below (for each test) are: 3D image of the grain color coded by number of bodies used in*  
 529 *the modeling (left) and respective modeling parameters (center) and resulting calculated anomaly (right) with*  
 530 *contours interval of 2000 nT. RMS = root means square error between calculated and observed anomaly grids.*

531 *The stereoplot in the M2 model is for NRM directions of the modeled tabular bodies; closed circles are for*  
532 *positive inclinations and open circles for negative inclinations, circles' colors indicate NRM intensities*

#### 4. DISCUSSION

533 We investigated three samples of the Reinfjord Ultramafic Complex using rock magnetic methods, optical and  
534 electron microscopy and magnetic modeling. The bulk susceptibility and NRM of the serpentinized and pristine  
535 dunite samples vary by more than one order of magnitude. These properties were investigated with respect to the  
536 magnetic mineralogy, composition, fabric and texture. Magnetic scans of thin sections were used to locate the  
537 magnetic sources that result in distinct magnetic anomalies. We have shown that the magnetic mineralogy, the  
538 source of the magnetic anomalies, is significantly different in the pristine and serpentinized dunite samples. In the  
539 pristine dunite sample, where the primary magnetic mineralogical assemblage is preserved, the predominant  
540 source is Cr-spinel with fine exsolution microstructures (1–3 $\mu$ m) of iron-rich ferrichromite to end-member  
541 magnetite. The Cr-spinel host is paramagnetic at room temperature due to its composition, therefore the  
542 exsolution intergrowths are the source of the magnetic anomalies. Thermomagnetic experiments confirm that  
543 there is a compositional variation within the grains which is reflected in the wide T<sub>c</sub> temperature range between  
544 200 °C and 579 °C (Table 3). The Verwey transition in the temperature versus susceptibility curves indicates the  
545 presence of near end-member magnetite. By contrast the predominant magnetic carrier in the serpentinized  
546 samples is end-member magnetite. Hysteresis parameters on the CS4 sample indicates that the bulk signal is a  
547 mixture of pseudo-single domain and single-domain size grains with relatively high coercivity, which can be  
548 explained by the presence of fine exsolution blebs. These are not homogeneously distributed within the Cr-spinel  
549 grains and likely the cause of the heterogeneous sources of magnetization shown in the analytic signal map and in  
550 the modeling results. Our modeling results confirm the microscopic observation of stronger magnetic intensity in  
551 areas of greater occurrences of ferrichromite exsolution within the Cr-rich spinel and weaker intensity where these  
552 exsolved phases were not observed. In the serpentinized dunite thin sections, the largest anomalies correlate with  
553 larger grains of magnetite commonly found together with pentlandite or other sulfides. Thermal experiments

554 indicate magnetite is the main magnetic carrier in the sample. Characteristic features which indicate magnetite, are  
555 the well-defined Hopkinson peaks and Verwey transitions, observed in the temperature-susceptibility curves, and  
556 the  $T_c$  estimates close to the  $T_c$  of endmember magnetite. The magnetite in veins, observed at the SEM, is fine  
557 grained, in agreement with the hysteresis parameters ( $M_r/M_s$  and  $H_{cr}/H_c$ ) approaching values that are diagnostic  
558 of very fine single-domain particles. Most of the fine magnetite grains have a relatively small effect on the  
559 observed magnetic scans' anomalies. This is possibly related to the resolution of the magnetic scans; lowering the  
560 sensor height and increasing the sampling density could better resolve the sources of the magnetization but this  
561 has instrumental challenges. It is also possible that, although the field intensity is weaker on the magnetic scan,  
562 the sources of these anomalies may contribute to the bulk properties if they preserve a consistent magnetization  
563 direction throughout the sample. Here, we modeled only the anomalies caused by larger grains that are associated  
564 with the highest field intensities. Most of the fine-grained magnetite in the serpentized samples formed during  
565 serpentization and, as suggested by the ratio between NRM and  $M_r$  (Table 2), yields a lower efficiency of  
566 magnetization with respect to the magnetic carriers in the pristine dunite sample. The Cr-spinel in the dunite likely  
567 developed the magnetic Fe-rich exsolution microstructures when cooling through the solvus at temperatures near,  
568 or above their respective  $T_c$ , therefore the NRM in this sample can be considered a thermoremanent  
569 magnetization (TRM). For non-interacting single-domain particles, Stacey and Banerjee (1974) calculate that  
570 remanence acquired during growth through chemical alteration (e.g. during serpentization), a chemical remanent  
571 magnetization (CRM), must be smaller than the TRM of the same particles. While a similar calculation cannot be  
572 made for grains with other domain states, which are common in natural samples, Smirnov and Tarduno (2005)  
573 argue that the CRM of particles with other domain states is likely to be even weaker than those that are single-  
574 domain. This suggests that magnetization acquired during serpentization (hereb400 °C) must have a lower  
575 efficiency than a TRM, consistent with our observations on the efficiency of pristine dunite and serpentinite  
576 samples. Modeling of the large discrete magnetic grains in the serpentized samples confirms the microscopic  
577 observation of stronger magnetic intensity in areas of discrete magnetite, and weaker intensity where magnetite is  
578 found together with pentlandite, or other sulfides (chalcopyrite and pyrrhotite). The orientation of the magnetic  
579 anomalies varies across these thin sections and in future work this variation could be tied to the magnetic history,

580 or the stability of NRM and acquisition of recent magnetic components. Such lines of enquiry could provide  
581 information on the timing of the serpentinization reactions. Scanning magnetic microscopy could be used to  
582 distinguish between primary and later magnetite formed during serpentinization, by measuring the magnetization  
583 response of individual, or assemblages of grains.

584 To summarize, modeling of the magnetic anomalies over isolated grains indicated heterogeneous sources of NRM  
585 within the grains in serpentinite with intensities varying from 2 to 14 A/m and variable directions. These estimates  
586 are slightly higher for the serpentinized sample grains than in the pristine dunite. This result together with the  
587 composition, the percentage of magnetic minerals and the fine-grainsize of the magnetic material in the  
588 serpentinized sample explain its high bulk magnetic susceptibility and NRM.

## 5. CONCLUSIONS

589 Scanning magnetic microscopy was used here to map the magnetic mineralogy of serpentinized and pristine  
590 dunite samples. Magnetic modeling in combination with chemical and magnetic properties analyses allowed  
591 characterization of the main magnetic carriers. The main results are summarized below:

- 592 • Magnetic carriers were identified based on Curie temperatures estimates and microscopic observations. In the  
593 serpentinized samples the magnetic carriers are end-member magnetite found both in veins, and as discrete  
594 large grains. Minor pyrrhotite was also observed. In the pristine dunite sample, the magnetic carriers are  
595 exsolution blebs with ferrichromite to end-member magnetite compositions in the Cr-spinel grains.
- 596 • The bulk NRM and magnetic susceptibility values are one order magnitude lower in the dunite sample than the  
597 serpentinized samples. This is explained by differences in magnetic mineralogy, content, grain size and texture  
598 between the serpentinized and the dunite samples. In the dunite sample the percentage of magnetic minerals  
599 calculated from  $M_s$  values is approximately 0.2%. Here, ferrichromite and minor magnetite exsolution  
600 microstructures in the Cr-spinel contribute to the magnetization. In the serpentinized sample the percentage of  
601 magnetic oxides is significantly larger at 2.8%, as calculated from  $M_s$ . Hysteresis properties indicate that the  
602 magnetite grains range from single domain to pseudo-single domain in size.

- 603 • Detailed modeling of a magnetic anomaly over an isolated grain indicates that there are heterogeneous sources  
604 of magnetic direction and intensity within the grain, in both pristine and serpentinized dunite samples.
- 605 • In the pristine dunite sample the heterogeneity is limited to the magnetization intensity which we interpret to  
606 be caused by variable concentration and composition of the ferrichromite exsolution within the Cr-spinel.  
607 However, the direction of magnetization is similar through-out the grain, implying similar timing of  
608 acquisition of magnetization for the exsolution microstructures.
- 609 • In the anomaly modeled over the serpentinized sample this heterogeneity applies both to the intensity and  
610 direction of the magnetization.
- 611 • Evaluating a larger view of the magnetic scan of the pristine dunite sample shows dipolar anomalies of similar  
612 orientation across the entire thin section. This suggests a consistent magnetization direction and may imply that  
613 the bulk magnetization direction of the sample is consistent with the fine-scale magnetization mapped here,  
614 and that this was acquired during initial cooling through the magnetite–Cr-spinel solvus, after emplacement. In  
615 the serpentinized samples the scans show variability in the orientation of the large dipolar anomalies among  
616 the different magnetite grains in the thin section. This heterogeneity will lower the total NRM. However, the  
617 fine-grained magnetite located in the veins will also add to the NRM. Future work would include  
618 demagnetization of the thin sections followed by magnetic mapping. This could be used to investigate the  
619 stability of the magnetization, and to explain better the non-unidirectional magnetization in the serpentinized  
620 samples.

621 We conclude that the bulk magnetic properties of these samples can be explained by the observed magnetic  
622 mineralogy. Forward modeling was effective in defining areas of higher magnetization that was constrained by  
623 the 3D geometry of the magnetic grains. Serpentinization clearly affected the heterogeneity of the magnetization  
624 direction across the thin section. The observation of numerous anomalies with variable directions and intensities  
625 over the larger grains in the magnetic scans of the serpentinites may suggest that the magnetization was acquired  
626 over a long time interval. In order to investigate further the link between the bulk NRM and the magnetization of  
627 the discrete grains, modeling of the entire thin section should be made. Curvature analysis (Phillips et al., 2007) of

628 the magnetic data acquired by scanning magnetic microscopy could be used to automatically determine the  
629 boundaries of magnetic sources across the entire thin section. This is planned for future work.



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640

641 Appendix A. Characteristic hysteresis parameters on sample chips before (BH) and after (AH) high temperature  
 642 VSM experiments. The precision used in the table is significant.

Sample	Chip	Chip Mass [g]	Mr [Am <sup>2</sup> /kg]	Ms [Am <sup>2</sup> /kg]	Hc	Hcr	Mr/Ms	Hcr/Hc	
CS4	C4	0.11	0.03	0.15	12.0	42.8	0.19	3.57	BH
		0.11	0.03	0.14	12.4	41.1	0.18	3.31	AH
	C5	0.09	0.03	0.10	17.8	36.2	0.28	2.03	BH
		0.09	0.03	0.10	17.5	33.4	0.27	1.91	AH
15S2	D4	0.06	1.58	4.56	29.4	41.8	0.35	1.42	BH
		0.06	1.28	4.21	24.8	37.4	0.30	1.51	AH
	D5	0.06	2.48	6.15	29.6	40.8	0.40	1.38	BH
		0.06	2.06	5.60	24.1	33.4	0.37	1.39	AH
	B3	0.07	2.30	6.49	32.0	45.5	0.36	1.42	BH
		0.07	1.63	5.65	25.8	40.8	0.29	1.58	AH

643

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