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Multifaceted Quadruplet of Low-Lying Spin-Zero States in <sup>66</sup>Ni: Emergence of Shape Isomerism in Light Nuclei

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A search for shape isomers in the  $^{66}$ Ni nucleus was performed, following old suggestions of various mean–field models and recent ones, based on state-of-the-art Monte Carlo shell model (MCSM), all considering  $^{66}$ Ni as the lightest nuclear system with shape isomerism. By employing the two-neutron transfer reaction induced by an  $^{18}$ O beam on a  $^{64}$ Ni target, at the sub-Coulomb barrier energy of 39 MeV, all three lowest-excited  $0^+$  states in  $^{66}$ Ni were populated and their  $\gamma$  decay was observed by  $\gamma$ -coincidence technique. The  $0^+$  states lifetimes were assessed with the plunger method, yielding for the  $0^+_2$ ,  $0^+_3$ , and  $0^+_4$  decay to the  $2^+_1$  state the B(E2) values of 4.3, 0.1, and 0.2 Weisskopf units (W.u.), respectively. MCSM calculations correctly predict the existence of all three excited  $0^+$  states, pointing to the oblate, spherical, and prolate nature of the consecutive excitations. In addition, they account for the hindrance of the E2 decay from the prolate  $0^+_4$  to the spherical  $2^+_1$  state, although overestimating its value. This result makes  $^{66}$ Ni a unique nuclear system, apart from  $^{236,238}U$ , in which a retarded  $\gamma$  transition from a  $0^+$  deformed state to a spherical configuration is observed, resembling a shape-isomerlike behavior.

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The concept of potential energy surface (PES) is central in many areas of physics. Usually, the PES displays the potential energy of the system as a function of its geometry. As an example, the PES of a molecule expressed in such coordinates as bond length, valence angles, etc., can be used for finding the minimum energy shape or calculating chemical reaction rates [1]. The idea of potential energy surface in deformation space has also been widely applied to the nucleus at a given spin. For an even-even nucleus at spin 0, the lowest PES minimum corresponds to the ground state (g.s.), while there may exist additional (secondary) minima in which excited  $0^+$  states can reside: they can be interpreted as ground states of different shapes [2–6]. When a secondary minimum is separated from the main minimum by a high barrier, in the extreme case a long-lived isomer, called *shape isomer*, can be formed [7]. Shape isomerism at spin zero, so far, has clearly been observed only in actinide nuclei - these isomers decay mainly by fission, and in two cases only,  $^{236}\text{U}$  and  $^{238}\text{U}$ , by very retarded  $\gamma$ -ray branches [8–11].

The existence of shape isomers in lighter systems has been a matter of debate for a long time. Already in the 1980s, a study based on microscopic Hartree-Fock plus BCS calculations, in which a large number of nuclei was surveyed, identified ten isotopes in which a deformed  $0^+$  state is separated from a spherical structure by a significantly high barrier:  $^{66}$ Ni and  $^{68}$ Ni,  $^{190,192}$ Pt,  $^{206,208,210}$ Os, and  $^{194,196,214}$ Hg [12]. Other investigations [13,14], which used a nonaxial Hartree-Fock-Bogoliubov approach, selected a rather restricted number of candidates, among which the lightest were  $^{66,68}$ Ni,  $^{74,76}$ Kr,  $^{78,80,98}$ Sr,  $^{80,82,100}$ Zr, and  $^{86}$ Mo. Quite recently, Möller *et al.* [15] presented a global study of potential energy surfaces in 7206 nuclei from A = 31 to A = 290 by employing a well-benchmarked macroscopic-microscopic finite-range

liquid-drop model [16]. Here, secondary PES minima at spin 0 were found in a few tens of nuclei, among which <sup>66</sup>Ni was also present. To summarize, <sup>66</sup>Ni is the lightest system for which all three models, discussed above, suggest the existence of a pronounced secondary PES minimum which may give rise to shape isomerism.

Recently, state-of-the-art shell model calculations became capable of calculating shape coexistence in nuclei with masses A = 60-100 [17–22]. In particular, it became possible to assess transition probabilities between states of different shapes, and to search for retarded decays which would be the signature of a shape-isomerlike structure. Monte Carlo shell model (MCSM) studies [23-26] have been performed for the neutron-rich <sup>68-78</sup>Ni isotopes and coexistence of low-lying spherical, oblate, and strongly deformed prolate shapes has been found in <sup>68</sup>Ni and <sup>70</sup>Ni [17,18]. A significant hindrance for the E2 transition deexciting the prolate deformed 0+ state was predicted in <sup>68</sup>Ni only. On the experimental side, the lifetimes of the supposedly well-deformed prolate 0<sup>+</sup> states, located at 2511 and 1567 keV in <sup>68</sup>Ni [27,28] and <sup>70</sup>Ni [29,30], respectively, were assessed in a very recent  $\beta$ -decay study of <sup>68</sup>Co and <sup>70</sup>Co [31]. No hindrance was observed in both cases, pointing to a substantial mixing (or a low potential barrier) between deformed and spherical configurations in <sup>68</sup>Ni and <sup>70</sup>Ni isotopes.

At this point, we turned our attention to <sup>66</sup>Ni, for which mean-field calculations already predicted a secondary PES deep minimum [12-16]. We performed new theoretical and experimental studies of <sup>66</sup>Ni low-lying states, with a particular focus on 0<sup>+</sup> states. The theoretical investigation was carried out within the Monte Carlo shell model [23–26], using the same Hamiltonian and model space as the ones previously employed for <sup>68–78</sup>Ni [17]. Figure 1 shows the potential energy surface of <sup>66</sup>Ni, obtained similarly to Refs. [17,18,32]. In comparison with <sup>68</sup>Ni, the barrier height is similar, but the prolate minimum is lower in  $^{66}$ Ni, due to stronger p-ninteraction coming from more neutron holes in the N=40closed shell. As a consequence, a more strict hindrance to the decay to the main spherical minimum, from the prolate one, is expected in <sup>66</sup>Ni. The circles on the plot of Fig. 1 indicate the intrinsic quadrupole moments of a given basis vector of the MCSM eigenstate, and their size implies the overlap probability of this basis vector with the eigenwave function [17,18].

In Fig. 1, the circles are concentrated in spherical  $(0_1^+)$  and  $(0_3^+)$ , oblate  $(0_2^+)$ , and prolate  $(0_4^+)$  domains, with  $\beta_2 \approx 0.0$ , -0.2, and (0.3), respectively. While the local minimum in the prolate domain is profound, due to a shell evolution of type II [18], the one in the oblate domain is very shallow, being almost a shoulder.

With respect to the  $0_1^+$  state, the  $0_{2,4}^+$  states are found to be composed of sizable excitations of protons, from  $f_{7/2}$  to  $f_{5/2}$  and  $p_{3/2,1/2}$ , as well as neutrons, from  $f_{5/2}$  and  $p_{3/2,1/2}$  to  $g_{9/2}$  and  $d_{5/2}$ . In the  $0_2^+$  state,  $\sim 1$  proton and  $\sim 1.5$ 

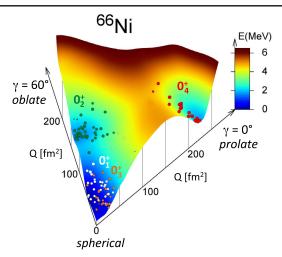


FIG. 1. Potential energy surface (PES) for the lowest  $0^+$  states of  $^{66}$ Ni, as a function of prolate and oblate quadrupole moments. Circles on the PES represent shapes in the MCSM basis vectors: white (orange) circles for  $0_1^+$  ( $0_3^+$ ) spherical states, green (red) for  $0_2^+$  oblate ( $0_4^+$  prolate) states, respectively.

neutrons are excited, whereas these numbers are doubled in the  $0_4^+$  state. This large difference in particle-hole excitations leads to very different deformations between  $0_2^+$  and  $0_4^+$  states. On the contrary, the  $0_{1,3}^+$  and  $2_1^+$  states have similar occupation numbers, being spherical.

The reduced probabilities for E2 transitions deexciting the  $0_2^+$  oblate,  $0_3^+$  spherical, and  $0_4^+$  prolate states to the spherical  $2_1^+$  state are 4.1, 0.01, and 0.006 Weisskopf units (W.u.), respectively. While the retardation of the  $0_4^+$  decay arises from the prolate to spherical shape change, the hindrance of the spherical  $0_3^+$  decay is caused by cancellation effects in the E2 transition matrix elements due to different structures within spherical states.

Experimentally, a first inspection of <sup>66</sup>Ni data, established prior to our study, already shows remarkable features in line with the MCSM calculations: three excited 0<sup>+</sup> states (out of a total of 6 excitations below 3 MeV) have been located at 2444, 2664, and 2965 keV by using a (t,p) reaction [33]. Out of those, the first two were confirmed by the discrete  $\gamma$ -ray studies of [27,34,35] and their energies were determined with higher precision at 2443 and 2671 keV. On the contrary, the  $0_4^+$  state was not observed. We propose that these three  $0^+$ excited states correspond to the three 0<sup>+</sup> excitations predicted by the MCSM calculations at 1971 (oblate), 2596 (spherical), and 3296 keV (prolate), respectively [see Figs. 2(d) and (e)]. This experiment-theory correspondence is supported by two independent observations: (i) both experiment [27,35,36] and MCSM calculations show that, out of the four 0<sup>+</sup> states in  $^{66}$ Ni, the  $\beta$  decay of the 1<sup>+</sup> ground state of  $^{66}$ Co feeds only the spherical ground state and the  $0^+_3$  state, what points to similarly spherical (although configuration-wise different) structures of these two states; (ii) in the (t,p) reaction study [33], the population of the  $0_2^+$  and  $0_3^+$  states is strongly

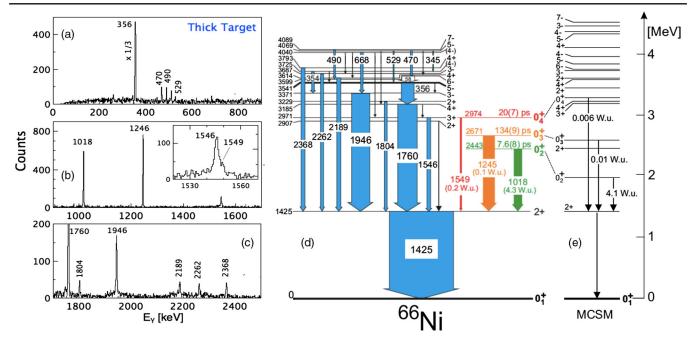


FIG. 2. (a)–(c)  $\gamma$ -ray spectrum of  $^{66}$ Ni gated by  $2_1^+ \rightarrow 0_1^+ 1425$ -keV transition, observed in the thick-target measurement. (d) Decay scheme of  $^{66}$ Ni observed in the experiment (black-thin arrows for known transitions [34] not seen here). (e) Results from MCSM calculations.

favored over the population of the  $0_4^+$  excitation (by a factor of 7 and 12, respectively).

To establish experimentally more detailed characteristics of the 0<sup>+</sup> excitations in <sup>66</sup>Ni, we undertook a new experimental study of their lifetimes, at the Tandem Laboratory of the Horia Hulubei National Institute for Physics and Nuclear Engineering (IFIN-HH) in Bucharest. The states of interest were populated in a 2*n*-transfer reaction induced by a beam of <sup>18</sup>O on a <sup>64</sup>Ni target, at incident energy (39 MeV) below the Coulomb barrier. The power of this approach resides in the fact that the fusion-evaporation channel, which is highly favored above the Coulomb barrier, is severely hindered in this case - its cross section becomes comparable to the one of the 2*n*-transfer process (few mb).

The  $\gamma$  transitions of  $^{66}$ Ni were measured using the ROSPHERE array [37], consisting of 14 Ge detectors and 11 LaBr<sub>3</sub>(Ce) scintillators. A first part of the experiment was carried out with a 5 mg/cm<sup>2</sup> thick target of  $^{64}$ Ni. A total of about  $10^8$   $\gamma$ - $\gamma$  coincidences were collected in a 6 day-long run with a beam intensity of more than 30 pnA. Figures 2(a)–2(c) display a spectrum obtained by setting a gate on the  $2^+ \rightarrow 0^+$ , 1425 keV, ground state transition in  $^{66}$ Ni. All visible  $\gamma$  rays correspond to transitions previously located in  $^{66}$ Ni [34] and deexciting all known states below  $\approx$ 4.1 MeV excitation energy [Fig. 2(d)].

In the spectrum of Figs. 2(a)–2(c), a number of transitions shows tails arising from the emission occurring in flight during the stopping process of the <sup>66</sup>Ni reaction product inside the target material. This happens when a state is fed directly and its lifetime is of the order of 1 to 2 ps.

The transitions to the first 2<sup>+</sup> state (at 1425.1 keV) from the  $0^+_2$  and  $0^+_3$  excited states are clearly visible in Fig. 2(b), with energies of 1018 and 1246 keV, respectively. A weaker 1546-keV line, depopulating the 3<sup>+</sup> state at 2971.0 keV, is also seen. It was verified that both 1018- and 1246-keV lines do not show any broadening, i.e., lifetimes longer than 2 ps are expected for the  $0^+_2$  and  $0^+_3$  states. In contrast, the 1546-keV peak (shown in the inset), is more complex: besides the broadening, which was used to extract a lifetime of 1.4(2) ps for the 3<sup>+</sup> state at 2971.0 keV (with the Doppler shift attenuation method [38]), a weak satellite line is observed on the right-hand side shoulder. We hypothesized that this satellite line, at approximate energy of 1549 keV, may be a transition from the  $0_4^+$  state of  $^{66}$ Ni to the first excited 2<sup>+</sup> state, which would then place this 0<sup>+</sup> excitation at 2974 keV. Our hypothesis was based on the fact that a  $0_4^+$  state located at 2965(10) keV in the low energy resolution (t, p) reaction study of Ref. [33], after recalibrating the (t, p) energy spectrum (using the precise transition energies of Ref. [34]), turned out to correspond to a value of 2973 keV. This agrees, within 1 keV, with the level deexcited by the 1549-keV  $\gamma$  ray. This interpretation is also supported by the intensity ratios between the  $\gamma$  decay from the excited  $0^+$  states, i.e., I(1018)/I(1246) = 0.7 and I(1549)/I(1246) = 0.05, which are of the same order as the ones established in the (t,p) reaction (0.6 and 0.15), respectively.

As the next step, we performed a lifetime measurement of the three  $0^+$  excited states decaying via the observed 1018-, 1246-, and 1549-keV transitions, by applying

the plunger technique. We used the same reaction,  $^{18}\mathrm{O}(39~\mathrm{MeV}) + ^{64}\mathrm{Ni}$ , but a thinner 1 mg/cm² target, in order to assure that all  $^{66}\mathrm{Ni}$  recoils exit the target material. The average velocity of the recoils was  $v/c \approx 2.2\%$ , corresponding to a time-of-flight of  $\approx 155~\mathrm{ps}$  over a 1 mm distance. For each  $\gamma$ -ray transition, the fraction of  $\gamma$  decay occurring in a Ta stopper (5 mg/cm² thick), placed at 12 distances from the target (i.e.,  $10, 20, 25, 40, 50, 60, 100, 200, 500, 1000, 2000, and 3000 <math>\mu$ m) was measured.

Figures 3(a) and 3(b) show partial  $\gamma$ -ray spectra gated on the 1425-keV  $2^+ \rightarrow 0^+$  transition of  $^{66}$ Ni, which were measured at short and long target-to-stopper distances, respectively. The spectrum in panel (a) is the sum of all spectra taken at stopper distances 10–100  $\mu$ m, which can be translated in the time-of-flight range 1.5–15 ps. Two well pronounced, sharp peaks correspond to the decay in the stopper of the  $0^+_2$  and  $0^+_3$  states, via the 1018- and 1246-keV  $\gamma$  rays, respectively. No in-flight components are visible due to the large spreading of the products velocity vectors. In turn, Fig 3(b) displays a summed spectrum for the distances 500–3000  $\mu$ m - it corresponds to the time-of-flight range 75–450 ps. Here, only one pronounced line at 1246 keV is seen.

It was clear then that the  $0_2^+$  state would have a lifetime of the order of 10 ps while the  $0_3^+$  excitation would decay with

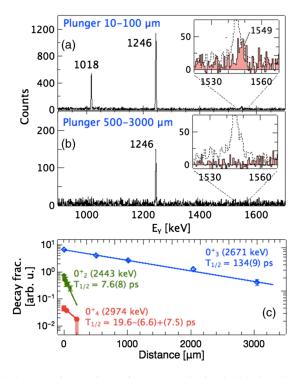


FIG. 3. (a)–(b) Portions of  $\gamma$  spectra obtained with the plunger setup, in coincidence with the 1425-keV transition [(a) sum over the short, (b) sum over the long distances]. The zoom around the 1546-keV transition is shown in the insets. The dashed histograms display the corresponding thick target spectrum, for comparison. (c) Decay curves of the three excited  $0^+$  states of  $^{66}$ Ni, as a function of the target-to-stopper distance.

a lifetime longer than 100 ps. The case of the  $0_4^+$  state, deexciting via a 1549-keV  $\gamma$  ray, required special attention. The insets in Figs. 3(a) and 3(b) show the portion of the corresponding spectra around the energy of 1549 keV. While in the thick-target spectrum [inset of Fig 2(b)] the peak around 1546 keV is composed of the 1546-keV  $3^+ \rightarrow 2^+$  line and the satellite transition at 1549 keV, as discussed earlier, in the plunger spectrum, corresponding to measurements at short distances [inset of Fig. 3(a)], only the 1549-keV  $\gamma$  ray is seen. This observation is in line with both the previously established 1.4 ps lifetime of the  $3^+$  state which deexcites via the 1546-keV  $\gamma$  ray and the existence of the 1549-keV  $0_4^+ \rightarrow 2_1^+$  transition, which exhibits a lifetime longer than a few picoseconds. The nonobservation of the 1549-keV line in the spectrum associated with larger plunger distances [inset in Fig. 3(b)] narrows the upper limit of the  $0_4^+$  state lifetime value to a few tens of picoseconds.

In order to extract precise values of the lifetime for the three  $0^+$  excited states, we developed an analysis procedure based on the Monte Carlo simulations of the reaction, energy loss in the target, plunger-target geometry, and  $\gamma$  emission in the plunger measurement. The initial velocity vector distribution of  $^{66}$ Ni products, after the 2n-transfer reaction, was calculated using the semiclassical GRAZING model [39] - the average value of the velocity, with respect to c, perpendicular to the plunger was 2.19% with  $\sigma=0.31\%$ .

The fraction of  $\gamma$ -ray decays in the stopper could then be simulated and a minimization procedure, with respect to the lifetime, was applied to the corresponding experimental data, at the considered distances. The lifetime uncertainty was determined by a Monte Carlo error estimate, taking into account the experimental errors.

The experimental values for the decay fraction of the 1246-keV line, deexciting the  $0_3^+$  state, were obtained from the intensities of the stopped components (observed in coincidence with the  $2^+ \rightarrow 0^+1425$ -keV transition, at each plunger distance), normalized to the 147-keV line from the 78.7 ns isomeric decay of <sup>79</sup>Kr, produced in the fusion channel. The results are shown in Fig. 3(c), together with the decay curve corresponding to the half-life  $T_{1/2}=134\pm 9$  ps, obtained from the minimization procedure described above.

In the case of the 1018-keV line, deexciting the  $0_2^+$  state, the experimental values of the decay fractions were obtained by using for normalization the long-lived 1246-keV transition mentioned above (correcting for its decay). The fitting procedure yielded the half-life value  $T_{1/2}=7.6\pm0.8\,$  ps, as shown in Fig. 3(c). For the 1549-keV  $\gamma$  ray, depopulating the  $0_4^+$  state, a dedicated procedure was applied, owing to its very low intensity. Three experimental points were used: (i) the decay intensity of the 1549-keV  $\gamma$  ray extracted from the thick target measurement, (ii) the decay fraction obtained by summing the plunger measurements between 10 and

100  $\mu$ m, and (iii) an upper limit for the decay fraction at 200 µm distance. For each point, again the 1246-keV transition was used for normalization. Here, the minimization procedure was based on the experimental weight of the individual measurements between 10 and 100  $\mu$ m (corresponding to an effective distance of 40  $\mu$ m). The value  $T_{1/2} = 19.6 - (6.6) + (7.5)$  ps was obtained, and the result is shown in Fig. 3(c). To estimate B(E2) reduced probabilities for the E2 transitions deexciting the  $0^+$  states, possible E0 branching into the  $0_{g,s}^+$  should be considered. We have calculated upper limits for the squared dimensionless monopole transition strengths  $\rho^2(E0)$  within the simple two-level model of Ref. [40], considering the shapes predicted by the MCSM discussed above, and assuming maximum mixing. We obtained  $\rho^2(E0) < 0.018$  and < 0.09for the oblate and prolate configurations, respectively, leading to E0 branches out of the  $0_2^+$  and  $0_4^+$  states  $\leq 0.1\%$  and  $\leq 3\%$ , respectively [41]. For the spherical  $0_3^+$ state the E0 branch is expected to be negligible. Such low values of E0 do not influence significantly the partial lifetimes for E2 decay from the excited  $0^+$  states - their effect is well within the experimental uncertainty. Therefore, we adopt the B(E2) values of  $4.3 \pm 0.5$ ,  $0.09 \pm 0.01$ , and  $0.21 \pm 0.07$  W.u., for the  $0_2^+$ ,  $0_3^+$ , and  $0_4^+$  states, respectively. It is remarkable to note that both E2 transitions from the  $0_3^+$ and  $0_4^+$  states are significantly retarded.

The MCSM calculations, discussed above, perfectly reproduce the  $0_2^+$  oblate state decay probability, while the calculated B(E2) values for transitions deexciting the spherical  $0_3^+$  and prolate  $0_4^+$  states are smaller than the experimental ones (i.e., 0.01 and 0.006 W.u., respectively). We remark that extremely small calculated B(E2) values ( $\ll$ 0.1 W.u.) may be related to the omission of tiny components of the wave function, which would be picked up in a larger calculation.

In summary, we have measured lifetimes for all three  $0^+$ excited states in <sup>66</sup>Ni by employing a two-neutron transfer reaction below the Coulomb barrier and the plunger technique. We found a substantial hindrance for the E2 decay from the second and third excited 0+ states. The MCSM calculations account well for the presence of all three excited 0<sup>+</sup> states and provide a good description of their E2 decay probabilities. Of special interest is the hindrance measured for the E2 transitions originating from the  $0_3^+$  spherical and  $0_4^+$  prolate state. While the retarded E2 decay from the  $0^+_3$  state is due to cancellation effects among E2 matrix elements, the retarded E2 decay from the  $0_4^+$ state arises from a sizable potential barrier between the prolate (secondary) and spherical ground state minima [18]. This result makes the <sup>66</sup>Ni nucleus a unique example of a nuclear system, apart from the actinides, in which a shapeisomerlike structure exists.

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