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# Integrability and non-integrability in $\mathcal{N}=2$ SCFTs and their holographic backgrounds 

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Abstract: We show that the string worldsheet theory of Gaiotto-Maldacena holographic duals to $\mathcal{N}=2$ superconformal field theories generically fails to be classically integrable. We demonstrate numerically that the dynamics of a winding string configuration possesses a non-vanishing Lyapunov exponent. Furthermore an analytic study of the Normal Variational Equation fails to yield a Liouvillian solution. An exception to the generic nonintegrability of such backgrounds is provided by the non-Abelian T-dual of $A d S_{5} \times S^{5}$; here by virtue of the canonical transformation nature of the T-duality classical integrability is known to be present.

Keywords: AdS-CFT Correspondence, Gauge-gravity correspondence

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## 1 Introduction

One of the most attractive features of the correspondence between $\mathcal{N}=4$ supersymmetric Yang-Mills theory and type IIB superstrings in $A d S_{5} \times S^{5}$ [1], is the presence of integrability in the planar limit. ${ }^{1}$ This triumph begs the question to what extent can integrable structures be found in less symmetric gauge theories. In this note we will concern ourselves with the study of integrable structures - or lack thereof - in the worldsheet theories that describe, via holography, quantum field theories.

For the $A d S_{5} \times S^{5}$ superstring classical integrability is ensured since the Lagrangian equations of motion can be expressed as the flatness of Lax connection [3]. A similar

[^0]Lax formulation of the dynamics is available for strings propagating in the gravity duals to marginal Leigh-Strassler real $\beta$-deformation that preserve $\mathcal{N}=1$ supersymmetry [4]. However, when the $\beta$-deformation is complex, integrability is not present [5-8].

In the same line, the dual to the $\mathcal{N}=1$ superconformal Klebanov-Witten theory is given by strings in $A d S_{5} \times T^{1,1}$ [9]. Despite the preserved supersymmetry, the large nonabelian isometry group and geodesic integrability of $T^{1,1}$, it has been shown that certain classical string configurations are chaotic and hence integrability is not present in the corresponding gauge theory [10]. Conversely the comparatively recent development of so-called $\lambda$-deformed [11, 12] and $\eta$-deformed [13, 14] theories provide backgrounds as developed in [15-19] with fewer/no isometries and no supersymmetries but yet are integrable; such theories are thought to be described by quantum group deformations of the $A d S_{5} \times S^{5}$ superstring. ${ }^{2}$

How then do we determine if a worldsheet theory is integrable? In the absence of general criteria for a string worldsheet theory to admit a Lax formulation, and hence classical integrability, this is rather hard. Disproving integrability would seem to be a more tractable problem, at least in principle. Even this task presents challenges, since it requires a rigorous analysis of the full non-linear PDE's that arise from a string sigma model. A more practical approach is to study particular wrapped string configurations upon which the worldsheet theory admits a consistent one-dimensional truncation and whose equations of motion (consisting of non-linear ODE's) can be analysed either numerically or analytically. Should this truncation display non-integrability then one can conclude that the parent string worldsheet theory is also non-integrable. This strategy was applied in the papers [8, 20-34]. The method is to propose a string soliton with $l$-degrees of freedom, write its classical equations of motion, find simple solutions for $(l-1)$ of these equations and replace the solutions in a fluctuated version the last equation. One arrives to a linear second order differential equation known as the normal variation equation (NVE) which takes the form $y^{\prime \prime}+\mathcal{B} y^{\prime}+\mathcal{A} y=0$. The existence of closed form or Liouvillian solutions depends on the characteristics of a combination of the functions $\mathcal{A}$ and $\mathcal{B}$ and its derivatives. The work of Kovacic [35] gives criteria for a Liouvillian solution to exist and even provides a algorithmic construction of such solutions (we review this technology in appendix A). When the NVE is not Liouvillian integrable then we can conclude the string worldsheet theory is also non-integrable.

Our goal in this paper is to use both analytic and numerical techniques to illustrate that a wide class of gravity duals to $\mathcal{N}=2$ SCFTs are non-integrable. These SCFTs were introduced by Gaiotto in [36]. We will show the non-integrability by examining string winding configurations in Type IIA Gaiotto-Maldacena spacetimes [37]. Such spacetimes are classified by a single function $V(\sigma, \eta)$ that solves a Laplace equation with a given charge density. With appropriate boundary conditions, which ensure regularity of the spacetime, the quiver gauge theory can be directly extracted from the charge density. Two choices of $V(\sigma, \eta)$ will be studied in greater detail; first is the case of the SfetsosThompson background [38] where $V=V_{S T}$ corresponds to the spacetime obtained as

[^1]the non-Abelian T-dual of $\operatorname{Ad} S_{5} \times S^{5}$. The second corresponds to the Maldacena-Nunez solution $V=V_{M N}$ obtained in [39]. Though neither of these strictly satisfy the boundary conditions, the two solutions are somewhat fundamental. Indeed, any such IIA GaiottoMaldacena spacetime can be build of an appropriate superposition of $V_{M N}$ potentials see for example [40, 41]. Also, a particular scaling limit of a generic Gaiotto-Maldacena solution approximates to the Sfetsos-Thompson one, see [42]. We will see that rotating and wrapped strings in the background generated by $V_{M N}$, fail to be integrable. However no-evidence of non-integrability is found for rotating and wrapped strings in the space time generated by $V_{S T}$. In fact it has long been known that the non-Abelian T-duality that relates this background to $A d S_{5} \times S^{5}$ is, at least in the bosonic sector, a canonical transformation [43]. Hence one would anticipate from the outset that strings in the SfetsosThompson spacetime are classically integrable. This integrability was proven by Borsato and Wulff $[44,45]$ who explained how such dualities act on the Lax connection of the $\mathbb{Z}_{4}$ graded super-coset formulation of the $\operatorname{AdS} S_{5} \times S^{5}$ superstring.

The structure of the rest of the paper is as follows: in section 2 we will place the general strategy of the paper - described above - in the context of the simple example of strings in a background obtained by non-Abelian T -duality on $\mathbb{R} \times S^{3}$. Both numerical and analytic approaches used to diagnose non-integrability turn out to be negative, a result which we explain by explicitly illustrating the Lax-formulation of the dynamics. ${ }^{3}$ This study is complemented with the material in appendix A, where we review the analytical techniques required. In section 3 we move to the study of wrapped string configurations in gravity duals to $\mathcal{N}=2$ SCFTs as outlined in the strategy described above. We find that in generic Gaiotto-Maldacena backgrounds, dual to $\mathcal{N}=2$ SCFTs, the string soliton's equations of motion have non-Liouvillian solutions. As a consequence, the background and the dual SCFT are non-integrable. We conclude in section 5 . Other appendices complement the presentation, making our technical results more solid.

## 2 A warm up: the non-Abelian T-dual of $R \times S^{3}$

In this paper we will consider bosonic strings propagating in a curved target spacetime endowed with a metric $G_{\mu \nu}$ and an NS two-form $B_{\mu \nu}$ described by the non-linear sigma-model ${ }^{4}$

$$
\begin{align*}
S & =\frac{1}{4 \pi \alpha^{\prime}} \int d \tau d \tilde{\sigma}\left[G_{\mu \nu} \eta^{\alpha \beta}+B_{\mu \nu} \epsilon^{\alpha \beta}\right] \partial_{\alpha} X^{\mu} \partial_{\beta} X^{\nu} \\
& =\frac{1}{\pi \alpha^{\prime}} \int d \tau d \tilde{\sigma}\left[G_{\mu \nu}+B_{\mu \nu}\right] \partial_{+} X^{\mu} \partial_{-} X^{\nu}, \tag{2.1}
\end{align*}
$$

supplemented by the Virasoro constraints that require the vanishing of the stress tensor $T_{\alpha \beta}$ :

$$
\begin{equation*}
G_{\mu \nu} \dot{X}^{\mu} X^{\prime \nu} \approx 0, \quad G_{\mu \nu}\left(\dot{X}^{\mu} \dot{X}^{\nu}+X^{\mu} X^{\prime \nu}\right) \approx 0 . \tag{2.2}
\end{equation*}
$$

[^2]One of the simplest such theories is given by a target space $\mathbb{R} \times S^{3}$ with no NS flux. Introducing an $\mathrm{SU}(2)$ valued worldsheet field $g(\tilde{\sigma}, \tau)$ to parametrise the $S^{3}$, of radius $\kappa$, and $\mathfrak{s u}(2)$ current $l_{\alpha}=-i g^{-1} \partial_{\alpha} g$ the non-linear sigma model can be cast as

$$
\begin{equation*}
S=\frac{1}{4 \pi} \int d \tau d \tilde{\sigma} \eta^{\alpha \beta}\left(\frac{\kappa^{2}}{2} \operatorname{tr}\left(l_{\alpha} l_{\beta}\right)-\partial_{\alpha} X_{0} \partial_{\beta} X_{0}\right) . \tag{2.3}
\end{equation*}
$$

Here we are working classically and have fixed the world sheet metric to be flat Minkowski space. The field $X_{0}$ is decoupled and free however by exploiting (fixing) residual timereparametrisations it can be placed in static gauge $X_{0}=E \tau$. This leaves only undetermined the dynamics of $g$, for which we express the equations of motion and Bianchi identity of the current $l_{\alpha}$ in terms of the flatness of an $\mathfrak{s u}(2)^{\mathbb{C}}$ valued Lax connection

$$
\begin{equation*}
\mathfrak{L}_{ \pm}[z]=\frac{i}{1 \mp z} l_{ \pm}, \quad\left[\partial_{+}+\mathfrak{L}_{+}, \partial_{-}+\mathfrak{L}_{-}\right]=0 \tag{2.4}
\end{equation*}
$$

leading to the integrability of the theory. The Hamiltonian analysis here is somewhat involved since the imposition of static gauge and the Virasoro constraints necessitate the introduction of Dirac brackets (a particularly comprehensive review of this system can be found in [47]).

The sigma-model defined by eq. (2.3) has an $\mathrm{SU}(2)_{L} \times \mathrm{SU}(2)_{R}$ global symmetry arising from the isometries of the $S^{3}$. We can T-dualise the $\mathrm{SU}(2)_{L}$ symmetry using a non-Abelian extension of the Buscher procedure [48]. This is achieved by introducing $\mathfrak{s u}(2)$ valued gauge fields to gauge the isometry and Lagrange multipliers $v^{a}$ that enforce the flatness of the corresponding field strength. After integrating out the non-propagating gauge fields one obtains a sigma-model of the form

$$
\begin{equation*}
S=\frac{1}{\pi} \int d \tau d \tilde{\sigma} \partial_{+} v^{a}\left(M^{-1}\right)_{a b} \partial_{-} v^{b}-\partial_{+} X_{0} \partial_{-} X_{0}, \quad M_{a b}=(4 \kappa)^{2} \delta_{a b}+4 f_{a b}^{c} v_{c} . \tag{2.5}
\end{equation*}
$$

After making a coordinate transformation

$$
\begin{equation*}
v^{1}=2 \sqrt{2} \kappa^{2} r \cos \chi, \quad v^{2}=2 \sqrt{2} \kappa^{2} r \sin \chi \cos \xi, \quad v^{3}=2 \sqrt{2} \kappa^{2} r \sin \chi \sin \xi, \tag{2.6}
\end{equation*}
$$

we can express the target space geometry of the dual theory as

$$
\begin{equation*}
d s^{2}=-\left(d X_{0}\right)^{2}+\frac{\kappa^{2}}{2}\left[d r^{2}+\frac{r^{2}}{1+r^{2}}\left(d \chi^{2}+\sin ^{2} \chi d \xi^{2}\right)\right], \quad B_{2}=-\frac{\kappa^{2}}{2} \frac{r^{3}}{1+r^{2}} \sin \chi d \chi \wedge d \xi \tag{2.7}
\end{equation*}
$$

There is also a non-constant dilaton coming from the Gaussian elimination of gauge fields but this will play no role in the classical analysis we continue with here.

From this construction it is natural to expect that the T-dual theory eq. (2.5) is integrable and this is indeed the case as we shall explain shortly. Let us first examine how the diagnostic tools developed in references [20] -[34] and further explained in appendix A can be applied and will fail to detect non-integrability.

We consider a wrapped string configuration specified by the ansatz

$$
\begin{equation*}
X_{0}=E \tau, \quad r=r(\tau), \quad \chi=\chi(\tau), \quad \xi=k \tilde{\sigma} . \tag{2.8}
\end{equation*}
$$

The equations of motion for this reduced system are

$$
\begin{align*}
& \ddot{r}=\frac{r\left(-k^{2} \sin ^{2} \chi+k r\left(r^{2}+3\right) \dot{\chi} \sin \chi+\dot{\chi}^{2}\right)}{\left(r^{2}+1\right)^{2}}  \tag{2.9}\\
& \ddot{\chi}=-\frac{k r \sin \chi\left(k\left(r^{2}+1\right) \cos \chi+\left(r^{2}+3\right) \dot{r}\right)+2 \dot{r} \dot{\chi}}{r^{3}+r}
\end{align*}
$$

and follow from the Hamiltonian

$$
\begin{equation*}
H=\frac{2 \pi}{\kappa^{2}}\left(p_{r}^{2}+\left(1+\frac{1}{r^{2}}\right) p_{\chi}^{2}\right)+k r \sin \chi\left(\frac{k r \kappa^{2}}{8 \pi}-p_{\chi}\right) . \tag{2.10}
\end{equation*}
$$

The non-trivial Virasoro constraint is

$$
\begin{equation*}
4 \pi H-E^{2} \approx 0 \tag{2.11}
\end{equation*}
$$

There are no secondary constraints since the derivative of the Virasoro constraint vanishes on the equations of motion.

Let us now trial the diagnostic tests of non-integrability on this reduced system.
The first test is numerical; we calculate the Lyapunov exponent. Given two initial points $X_{0}$ and $X_{0}+\Delta X_{0}$, arbitrarily close in the phase space, the Lyapunov exponent is defined as

$$
\begin{equation*}
\hat{\lambda}=\lim _{t \rightarrow \infty \Delta X_{0} \rightarrow 0} \lim _{X X_{0}} \hat{\lambda}(t), \quad \hat{\lambda}(t)=\frac{1}{t} \log \frac{\left|\Delta X\left(X_{0}, t\right)\right|}{\left|\Delta X\left(X_{0}, 0\right)\right|} . \tag{2.12}
\end{equation*}
$$

This provides a quantitative measure on the rate of increase (or decrease) in the separation between two infinitesimally close trajectories in the phase space. In our analysis, for obvious reasons, we will be concerned with the largest positive Lyapunov exponent and consider $\hat{\lambda}(t)$ measured at sufficiently late times. The largest Lyapunov exponent corresponding to dynamical systems exhibiting integrable trajectories in the phase space is identically zero. On the other hand, for systems exhibiting chaos $\hat{\lambda}$ is non zero and saturates to a positive value that guarantees an exponential growth in the separation between two nearby trajectories at sufficiently late times. We calculate numerically $\hat{\lambda}(t)$ by evolving a set of initial conditions that identically set the Hamiltonian constraint equal to zero, $H=0$. The result is displayed as the $\epsilon=0$ plot in figure 1. This indicates a vanishing Lyapunov exponent thus no evidence of non-integrability.

To use the analytic tests of non-integrability [22], we consider the normal variational equation (NVE) to a seed solution to the equations eq. (2.9):

$$
\begin{equation*}
r(\tau)=a \tau+b, \quad \chi(\tau)=0 . \tag{2.13}
\end{equation*}
$$

We then fluctuate $\delta \chi=0+\epsilon f(\tau)$ and take the equations to leading order in $\epsilon$ :

$$
\begin{equation*}
0=k\left(\frac{2 a}{(a \tau+b)^{2}+1}+a+k\right) f(\tau)+\frac{2 a}{(a \tau+b)^{3}+a \tau+b} \dot{f}(\tau)+\ddot{f}(\tau) . \tag{2.14}
\end{equation*}
$$

Defining $a \tau+b=\tilde{\tau}, a \hat{k}=k$ and $f(\tau)=\frac{\sqrt{1+\tilde{\tau}^{2}}}{\tilde{\tau}} \psi(\tilde{\tau})$ the NVE can be placed in normal form

$$
\begin{equation*}
\psi^{\prime \prime}(\tilde{\tau})+V(\tilde{\tau}) \psi=0, \quad V(\tilde{\tau})=\frac{3}{\left(\tilde{\tau}^{2}+1\right)^{2}}+\frac{2 \hat{k}}{\tilde{\tau}^{2}+1}+\hat{k}(1+\hat{k}) . \tag{2.15}
\end{equation*}
$$

This last equation, can be readily integrated. In fact, one can check that the function $V(\tilde{\tau})$ satisfies the necessary conditions to be Liouville integrable discussed in appendix A. See in particular the analysis in appendix A.3. Indeed, we can find two independent explicit solutions of the NVE,

$$
\begin{equation*}
f_{1}(\tau)=\frac{\hat{k} \cos (\tilde{\tau} \sqrt{x})}{\tilde{\tau}}-\sqrt{x} \sin (\tilde{\tau} \sqrt{x}), \quad f_{2}(\tau)=\frac{\hat{k} \sin (\tilde{\tau} \sqrt{x})}{\tilde{\tau}}+\sqrt{x} \cos (\tilde{\tau} \sqrt{x}) \tag{2.16}
\end{equation*}
$$

where $x=\hat{k}(1+\hat{k})$. Whilst this solution may not be completely trivial it is composed of appropriate elementary building blocks and is Liouvillian. This analytic test also provides no evidence of non-integrability. Actually, as we discuss now, the integrability of the system can be proven.

Indeed, the reason both tests for non-integrability returned negative results is that the Hamiltonian in eq. (2.10) is, in fact, integrable in the Liouvillian sense. In addition to the Hamiltonian in eq. (2.10), one can readily check that a second conserved charge is given by

$$
\begin{equation*}
Q=p_{r}^{2}+\frac{p_{\chi}^{2}}{r^{2}}+\frac{k \kappa^{2}}{2 \pi}\left(p_{r} \cos \chi-\frac{p_{\chi}}{r} \sin \chi\right) \tag{2.17}
\end{equation*}
$$

To make the integrability of eq. (2.10) we could here proceed to give a Lax pair of matrices $\{\mathbb{L}, \mathbb{M}\}$ obeying $d_{\tau} \mathbb{L}=[\mathbb{L}, \mathbb{M}]$ from which the conserved charges of $H$ and $Q$ are obtained via $\operatorname{Tr}\left(\mathbb{L}^{n}\right)$. The explicit form of the Lax matrices we found is not very enlightening and so we don't present it here. Instead, it is more powerful to see that parent two-dimensional theory defined by the T-dual action in eq. (2.5) is itself integrable. The equations of motion for the fields $v^{a}$ can be packaged into a Lax form

$$
\begin{align*}
\widehat{\mathfrak{L}}_{ \pm}[z] & =\frac{1}{1 \mp z} \widehat{l}_{ \pm}, & {\left[\partial_{+}+\widehat{\mathfrak{L}}_{+}, \partial_{-}+\widehat{\mathfrak{L}}_{-}\right]=0 }  \tag{2.18}\\
\widehat{l}_{+} & =-4\left(M^{-T} \partial_{+} v\right)^{a} T_{a}, & \widehat{l}_{-}=4\left(M^{-1} \partial_{-} v\right)^{a} T_{a}
\end{align*}
$$

This Lax formulation [44] follows from the mapping $l_{ \pm} \rightarrow \widehat{l}_{ \pm}$between world-sheet derivatives obtained as a consequence of the Buscher procedure and which defines a canonical transformation map between the starting theory and its dual [43]. Of course one should be concerned about the transformation of the Virasoro constraints between the two theories, however it is easy to see using the definition of the matrix $M_{a b}$ in eq. (2.5) that,

$$
\begin{equation*}
\frac{1}{2} \partial_{ \pm} v^{T}\left(M^{-1}+M^{-T}\right) \partial_{ \pm} v=\kappa^{2} \widehat{l}_{ \pm} \cdot \widehat{l}_{ \pm} \tag{2.19}
\end{equation*}
$$

and hence the Virasoro constraints are equivalent to

$$
\begin{equation*}
\kappa^{2} \widehat{l}_{ \pm} \cdot \widehat{l}_{ \pm} \approx\left(\partial_{ \pm} X_{0}\right)^{2} \tag{2.20}
\end{equation*}
$$

These follow from the Virasoro constraints of the starting theory by the canonical transformation $l_{ \pm} \rightarrow \widehat{l}_{ \pm}$.

Below, we use the same diagnostic tool kit to study non-integrability in a set of holographic duals to $\mathcal{N}=2$ SCFTs. These Type IIA backgrounds share some of the structure
of the background described around eq. (2.7). The non-Abelian T-dual of $\operatorname{AdS} S_{5} \times S^{5}$ however stands out as a unique example of these geometries since it provides an integrable theory $[44,45]$ (see appendix C for an explicit construction). ${ }^{5}$ In contrast we will see that generic backgrounds dual to $\mathcal{N}=2$ SCFTs will exhibit non-integrability.

## 3 Integrability of $\mathcal{N}=2$ conformal field theories

In this section, we study the (non)-integrability of a generic family of $\mathcal{N}=2$ SCFTs. Our results are presented in the language of the holographic string dual to these SCFTs.

The strategy we adopt follows the ideas in [10, 22]: we propose a string configuration, and study its classical equations of motion following from the action of eq. (2.1). We shall demonstrate analytically the non-integrability of these equations (non-existence of Liouvillian solutions as described in appendix A) and show the presence of chaos in the dynamical evolution. The existence of one such string configurations rules out the integrability of the whole CFT.

The super conformal field theories we focus our attention on, were introduced by Gaiotto in [36]. We start by summarising the holographic string dual to these conformal theories. The general form of Type IIA backgrounds was first presented by Gaiotto and Maldacena in [37]. These solutions are completely determined in terms of a potential function $V(\sigma, \eta)$. Denoting

$$
\dot{V}=\sigma \partial_{\sigma} V, \quad \ddot{V}=\sigma^{2} \partial_{\sigma}^{2} V+\sigma \partial_{\sigma} V ; \quad V^{\prime}=\partial_{\eta} V, \quad V^{\prime \prime}=\partial_{\eta}^{2} V,
$$

one can write the Type IIA background as

$$
\begin{align*}
d s_{I I A, s t}^{2}= & \alpha^{\prime}\left(\frac{2 \dot{V}-\ddot{V}}{V^{\prime \prime}}\right)^{1 / 2} \\
& \times\left[4 A d S_{5}+\mu^{2} \frac{2 V^{\prime \prime} \dot{V}}{\Delta} d \Omega_{2}^{2}(\chi, \xi)+\mu^{2} \frac{2 V^{\prime \prime}}{\dot{V}}\left(d \sigma^{2}+d \eta^{2}\right)+\mu^{2} \frac{4 V^{\prime \prime} \sigma^{2}}{2 \dot{V}-\ddot{V}} d \beta^{2}\right], \\
A_{1}= & 2 \mu^{4} \sqrt{\alpha^{\prime}} \frac{2 \dot{V} \dot{V}^{\prime}}{2 \dot{V}-\ddot{V}} d \beta, \quad e^{4 \phi}=4 \frac{(2 \dot{V}-\ddot{V})^{3}}{\mu^{4} V^{\prime \prime} \dot{V}^{2} \Delta^{2}}, \quad \Delta=(2 \dot{V}-\ddot{V}) V^{\prime \prime}+\left(\dot{V}^{\prime}\right)^{2}, \\
B_{2}= & 2 \mu^{2} \alpha^{\prime}\left(\frac{\dot{V} \dot{V}^{\prime}}{\Delta}-\eta\right) d \Omega_{2}, \quad C_{3}=-4 \mu^{4} \alpha^{\prime 3 / 2} \frac{\dot{V}^{2} V^{\prime \prime}}{\Delta} d \beta \wedge d \Omega_{2} . \tag{3.1}
\end{align*}
$$

The radius of the space is $\mu^{2} \alpha^{\prime}=L^{2}$. We use the two-sphere metric $d \Omega_{2}^{2}(\chi, \xi)=d \chi^{2}+$ $\sin ^{2} \chi d \xi^{2}$, with corresponding volume form $d \Omega_{2}=\sin \chi d \chi \wedge d \xi$. The usual definition $F_{4}=d C_{3}+A_{1} \wedge H_{3}$ is also used.

To write backgrounds in this family, one should find the function $V(\sigma, \eta)$ that solves a Laplace problem with a given charge density $\tilde{\lambda}(\eta)$ and boundary conditions,

$$
\begin{align*}
& \partial_{\sigma}\left[\sigma \partial_{\sigma} V\right]+\sigma \partial_{\eta}^{2} V=0, \\
& \tilde{\lambda}(\eta)=\left.\sigma \partial_{\sigma} V(\sigma, \eta)\right|_{\sigma=0} . \quad \tilde{\lambda}(\eta=0)=0, \quad \tilde{\lambda}\left(\eta=N_{c}\right)=0 . \tag{3.2}
\end{align*}
$$

[^3]The boundary condition at $\eta=N_{c}$ ensures that the corresponding SCFT quiver will have finite length, in this work we will also want to relax this boundary condition such that e.g. the Maldacena-Nunez solution is incorporated.

For the purposes of the classical analysis of our bosonic sigma model string solution, we need only the metric and the $B_{2}$-field in the configuration of eq. (3.1). We introduce the notation

$$
\begin{align*}
d s^{2} & =4 f_{1}(\sigma, \eta) A d S_{5}+f_{2}(\sigma, \eta)\left(d \sigma^{2}+d \eta^{2}\right)+f_{3}(\sigma, \eta) d \Omega_{2}(\chi, \xi)+f_{4}(\sigma, \eta) d \beta^{2} \\
B_{2} & =f_{5}(\sigma, \eta) \sin \chi d \chi \wedge d \xi \tag{3.3}
\end{align*}
$$

In what follows, we set $\alpha^{\prime}=L=\mu=1$. Comparing with eq. (3.1), the functions $f_{i}(\sigma, \eta)$ are,

$$
\begin{equation*}
f_{1}=\left(\frac{2 \dot{V}-\ddot{V}}{V^{\prime \prime}}\right)^{1 / 2}, f_{2}=f_{1} \frac{2 V^{\prime \prime}}{\dot{V}}, f_{3}=f_{1} \frac{2 V^{\prime \prime} \dot{V}}{\Delta}, f_{4}=f_{1} \frac{4 V^{\prime \prime} \sigma^{2}}{2 \dot{V}-\ddot{V}}, f_{5}=2\left(\frac{\dot{V} \dot{V}^{\prime}}{\Delta}-\eta\right) . \tag{3.4}
\end{equation*}
$$

Let us now propose a string configurations and study its dynamical evolution.

### 3.1 Study of strings in Gaiotto-Maldacena backgrounds

We consider a string that sits in the center of $A d S_{5}$ rotates and wraps on the following coordinates ( $\tau, \tilde{\sigma}$ are the world-sheet coordinates),

$$
\begin{equation*}
t=t(\tau), \quad \sigma=\sigma(\tau), \quad \eta=\eta(\tau), \quad \chi=\chi(\tau) ; \quad \xi=k \tilde{\sigma}, \quad \beta=\lambda \tilde{\sigma} \tag{3.5}
\end{equation*}
$$

With $(k, \lambda)$ being integer numbers that indicate how many times the string wraps each of the corresponding directions.

To study the equations of motion, we write an effective Lagrangian using eq. (2.1) and the associated Virasoro constraint. For the configuration in eq. (3.5), we have, ${ }^{6}$

$$
\begin{align*}
L & =4 f_{1} \dot{t}^{2}-f_{2}\left(\dot{\sigma}^{2}+\dot{\eta}^{2}\right)+f_{3}\left(k^{2} \sin ^{2} \chi-\dot{\chi}^{2}\right)+f_{4} \lambda^{2}+2 k f_{5} \dot{\chi} \sin \chi  \tag{3.6}\\
T_{\tau \tau} & =T_{\tilde{\sigma} \tilde{\sigma}}=-4 f_{1} \dot{t}^{2}+f_{2}\left(\dot{\sigma}^{2}+\dot{\eta}^{2}\right)+f_{3}\left(k^{2} \sin ^{2} \chi+\dot{\chi}^{2}\right)+f_{4} \lambda^{2}=0 . \quad T_{\tilde{\sigma} \tau}=0
\end{align*}
$$

There is also an effective Hamiltonian that reads,

$$
\begin{equation*}
-H=-\frac{p_{t}^{2}}{16 f_{1}}+\frac{p_{\sigma}^{2}+p_{\eta}^{2}}{4 f_{2}}+\frac{1}{f_{3}}\left(\frac{p_{\chi}}{2}+k f_{5} \sin \chi\right)^{2}+k^{2} f_{3} \sin ^{2} \chi+\lambda^{2} f_{4} \tag{3.7}
\end{equation*}
$$

[^4]The equations of motion derived from the effective Lagrangian are,

$$
\begin{aligned}
f_{1} \dot{t}= & E . \\
f_{3} \ddot{\chi}= & -f_{3} k^{2} \cos \chi \sin \chi+k \sin \chi\left(\dot{\eta} \partial_{\eta} f_{5}+\dot{\sigma} \partial_{\sigma} f_{5}\right)-\dot{\chi}\left(\dot{\eta} \partial_{\eta} f_{3}+\dot{\sigma} \partial_{\sigma} f_{3}\right) . \\
f_{2} \ddot{\sigma}= & -\dot{\eta} \dot{\sigma} \partial_{\eta} f_{2}-2 \frac{E^{2}}{f_{1}} \partial_{\sigma} \log f_{1}+\frac{1}{2}\left(\dot{\eta}^{2}-\dot{\sigma}^{2}\right) \partial_{\sigma} f_{2}+\frac{1}{2}\left(-k^{2} \sin ^{2} \chi+\dot{\chi}^{2}\right) \partial_{\sigma} f_{3} \\
& -\frac{\lambda^{2}}{2} \partial_{\sigma} f_{4}-k \sin \chi \dot{\chi} \partial_{\sigma} f_{5} . \\
f_{2} \ddot{\eta}= & -\dot{\eta} \dot{\sigma} \partial_{\sigma} f_{2}-2 \frac{E^{2}}{f_{1}} \partial_{\eta} \log f_{1}+\frac{1}{2}\left(-\dot{\eta}^{2}+\dot{\sigma}^{2}\right) \partial_{\eta} f_{2}+\frac{1}{2}\left(-k^{2} \sin ^{2} \chi+\dot{\chi}^{2}\right) \partial_{\eta} f_{3} \\
& -\frac{\lambda^{2}}{2} \partial_{\eta} f_{4}-k \sin \chi \dot{\chi} \partial_{\eta} f_{5} .
\end{aligned}
$$

Here $E$ is a constant of motion. Notice that the $t$-equation of motion - the first of eqs. (3.8) was used in the other equations.

The reader can check that the derivative of the Virasoro constraint vanishes when evaluated on the second order equations (3.8). Hence the constraint is a constant 'on shell'. We choose the integration constant $E$ such that $T_{\alpha \beta}=0$.

The equations (3.8) define the $\tau$-evolution of our string configuration. Below, we discuss the possibility of finding simple solutions. We also study the non-integrability of these simple solutions and the chaotic dynamics of the configuration proposed in eq. (3.5).

### 3.2 Finding simple solutions

Since the equations (3.8) depend on the functions $f_{1}, \ldots, f_{5}$ and these in turn depend on the function $V(\sigma, \eta)$, our analysis should start by specifying the potential function $V(\sigma, \eta)$.

Inspired by the paper [49], we choose a potential function such that expanded close to $\sigma=0$ reads,

$$
\begin{equation*}
V(\sigma, \eta)=F(\eta)+a \eta \log \sigma+\sum_{k=1}^{\infty} h_{k}(\eta) \sigma^{2 k} . \tag{3.9}
\end{equation*}
$$

This potential satisfies the Laplace equation (3.2) if,

$$
\begin{equation*}
4 h_{1}(\eta)=-F^{\prime \prime}(\eta), \quad h_{k}(\eta)=-\frac{1}{4 k^{2}} h_{k-1}^{\prime \prime}(\eta), \quad k=2,3, \ldots \tag{3.10}
\end{equation*}
$$

Which determines all the functions $h_{k}(\eta)$ in terms of $F(\eta)$. This potential function gives a charge density $\tilde{\lambda}(\eta)=a \eta$, as defined by eq. (3.2). Hence, it does not satisfy the second boundary condition in eq. (3.2).

Let us now find a simple solution to eqs. (3.8). We observe that the potential in eq. (3.9) implies that close to $\sigma(\tau)=0$, the functions $\left.\partial_{\sigma} f_{i}\right|_{\sigma=0}=0$. This suggest a simple configuration,

$$
\sigma(\tau)=\dot{\sigma}(\tau)=\ddot{\sigma}(\tau)=0
$$

that solves automatically the equation for the $\sigma$-variable, the third in eqs. (3.8).

Expanding close to $\sigma(\tau)=0$, the functions $f_{i}(\sigma=0, \eta)$ are only functions of $\eta(\tau)$. The remaining two equations for the variables $\chi(\tau)$ and $\eta(\tau)$ read,

$$
\begin{align*}
& f_{3} \ddot{\chi}=-f_{3} k^{2} \cos \chi \sin \chi+k \sin \chi \dot{\eta} \partial_{\eta} f_{5}-\dot{\chi} \dot{\eta} \partial_{\eta} f_{3}  \tag{3.11}\\
& f_{2} \ddot{\eta}=-2 \frac{E^{2}}{f_{1}} \partial_{\eta} \log f_{1}-\frac{1}{2} \dot{\eta}^{2} \partial_{\eta} f_{2}+\frac{1}{2}\left(-k^{2} \sin ^{2} \chi+\dot{\chi}^{2}\right) \partial_{\eta} f_{3}-\frac{\lambda^{2}}{2} \partial_{\eta} f_{4}-k \sin \chi \dot{\chi} \partial_{\eta} f_{5} . \tag{3.12}
\end{align*}
$$

With these two equations, we follow the procedure described in [22]. First, consider the situation in which $\chi(\tau)=\dot{\chi}(\tau)=\ddot{\chi}(\tau)=0$. This solves eq. (3.11) and using eq. (3.12) gives,

$$
\begin{equation*}
f_{2} \ddot{\eta}=-2 \frac{E^{2}}{f_{1}} \partial_{\eta} \log f_{1}-\frac{1}{2} \dot{\eta}^{2} \partial_{\eta} f_{2}-\frac{\lambda^{2}}{2} \partial_{\eta} f_{4} . \tag{3.13}
\end{equation*}
$$

With the function $V(\sigma, \eta)$ proposed in eq. (3.9), we find expressions for the $f_{i}(0, \eta)$ and $\partial_{\eta} f_{i}(0, \eta)$. Using these expressions the $\eta$-equation (3.13) reads (we are always expanding close to $\sigma(\tau)=0$ ),

$$
\begin{equation*}
\ddot{\eta}=\left(\dot{\eta}^{2}-E^{2}\right)\left(\frac{F^{\prime \prime}-\eta F^{\prime \prime \prime}}{4 \eta F^{\prime \prime}}\right) . \tag{3.14}
\end{equation*}
$$

This equation can be easily solved for all functions $F(\eta)$. The solution is $\eta_{s}(t)=E \tau+\beta$ with $(E, \beta)$ constants, compare this with eq. (2.13). We use this simple solution in what follows.

Fluctuating eq. (3.12) by $\chi(t)=0+z(t)$, to first order in the fluctuation we get,

$$
\begin{align*}
& \ddot{z}(\tau)+\mathcal{B} \dot{z}(\tau)+\mathcal{A} z(\tau)=0,  \tag{3.15}\\
& \mathcal{A}=\left.\left(k^{2}-k \dot{\eta} \frac{\partial_{\eta} f_{5}}{f_{3}}\right)\right|_{\eta=\eta_{s}}, \quad \mathcal{B}=\left.\left(\dot{\eta} \partial_{\eta} \log f_{3}\right)\right|_{\eta=\eta_{s}} .
\end{align*}
$$

The integrability depends on the behaviour of this last equation respect to Kovacic's criterium, see appendix A for an explanation and application to some examples. In terms of the function $F(\eta)$ that defines the function $V$ in eq. (3.9), the explicit expression of the coefficients $\mathcal{A}, \mathcal{B}$ is

$$
\begin{aligned}
\mathcal{A} & =k^{2}+k \dot{\eta} \frac{\sqrt{2}}{\sqrt{a \eta F^{\prime \prime}}} \times\left.\frac{2 a F^{\prime \prime}+2 \eta\left(F^{\prime \prime}\right)^{2}+a \eta F^{\prime \prime \prime}}{\left(a+2 \eta F^{\prime \prime}\right)}\right|_{\eta_{s}} \\
\mathcal{B} & =\dot{\eta} \frac{1}{2 \eta F^{\prime \prime}\left(a+2 \eta F^{\prime \prime}\right)} \times\left.\left(3 a F^{\prime \prime}+2 \eta F^{\prime \prime 2}+a \eta F^{\prime \prime \prime}-2 \eta^{2} F^{\prime \prime} F^{\prime \prime \prime}\right)\right|_{\eta_{s}}
\end{aligned}
$$

To gain more understanding of the integrability (or not) of our string soliton, we should study the application of Kovacic's criterium to eq. (3.15) for different backgrounds dual to $\mathcal{N}=2$ SCFTs. In what follows, we will specify different functions $V(\sigma, \eta)$ that define various well known backgrounds. We shall observe that whilst the generic case turns out to be non-integrable, there is a particular background - the Sfetsos-Thompson solution, for which the string soliton of eq. (3.5) is integrable. After that, we complement this analytic study with the numerical study of the eqs. (3.8) or equivalently, those derived from the Hamiltonian in eq. (3.7).

### 3.3 Some interesting examples

Below, we apply Kovacic's procedure (discussed in detail in appendix A) to eq. (3.15). We discuss various examples in turn by specifying particular potential functions $V(\sigma, \eta)$.

### 3.3.1 The Sfetsos-Thompson background

Let us start with the potential describing the Sfetsos-Thompson background, obtained by application of non-Abelian T-duality on $A d S_{5} \times S^{5}$ [38]. The field theoretical dual to this background was discussed in [50].

For the Sfetsos-Thompson (ST) solution, the potential function reads

$$
\begin{equation*}
V_{S T}=\eta \log \sigma-\eta \frac{\sigma^{2}}{2}+\frac{\eta^{3}}{3} \tag{3.16}
\end{equation*}
$$

In the notation of eqs. (3.9)-(3.10), we have $a=1, F(\eta)=\frac{1}{3} \eta^{3}, h_{1}=-\frac{\eta}{2}, h_{k>2}=0$. The coefficients $\mathcal{A}, \mathcal{B}$ are,

$$
\mathcal{A}_{S T}=k^{2}+2 E k \frac{\left(4 \eta^{2}+3\right)}{\left(4 \eta^{2}+1\right)}, \quad \mathcal{B}_{S T}=\frac{2 E}{\eta\left(4 \eta^{2}+1\right)},
$$

and eq. (3.15) becomes,

$$
\begin{equation*}
\ddot{z}+\frac{2 E}{\eta\left(4 \eta^{2}+1\right)} \dot{z}+\left(k^{2}+2 E k \frac{\left(4 \eta^{2}+3\right)}{\left(4 \eta^{2}+1\right)}\right) z=0 \tag{3.17}
\end{equation*}
$$

This equation admits Liouvillian solutions, the system is the same, up to a rescaling of the $\tau$-coordinate, to that analysed around eq. (2.14). Indeed, as we did with eq. (2.14), we can transform eq. (3.17) into a Schroedinger-like form defining $z(\tau)=e^{-\frac{1}{2} \int d \tau \mathcal{B}(\tau)} \psi(\tau)$,

$$
\begin{equation*}
\ddot{\psi}+V \psi=0, \quad V=\frac{1}{4}\left(4 \mathcal{A}-\mathcal{B}^{2}-2 \mathcal{B}^{\prime}\right)=\frac{12}{\left(4 \tau^{2}+1\right)^{2}}+k^{2}+2 k+\frac{4 k}{4 \tau^{2}+1} . \tag{3.18}
\end{equation*}
$$

In the last line, we have set the integration constant $E=1$. This last equation can be easily solved, as we did with eq. (2.15). Transforming back to the function $z(\tau)$, we find that eq. (3.17) admits Liouvillian solutions, written as a combination of trigonometric and rational functions. See also appendix A for the application of Kovacic's criteria to this case.

The presence of a string soliton that is Liouville-integrable is certainly not enough to claim the integrability of the theory. Nevertheless, this result together with the analysis in section 2, reinforce the point that the Sfetsos-Thomspson background is dual to an integrable CFT (see [44, 45]). In other words, the non-Abelian T-duality does not spoil the integrable character of a background-QFT pair.

Let us now study the effect of deforming the Sfetsos-Thompson solution.

### 3.3.2 Deforming the Sfetsos-Thompson solution

As anticipated, we consider 'deformations' away from the Sfetsos-Thompson solution. To parametrise these deformations, we propose a potential written in terms of a parameter $\epsilon$,

$$
\begin{equation*}
V_{\mathrm{def}}=V_{S T}+\epsilon\left(\frac{\eta^{4}}{12}+\frac{\sigma^{4}}{32}-\frac{\sigma^{2} \eta^{2}}{4}\right) \tag{3.19}
\end{equation*}
$$

This potential function $V_{\text {def }}$ satisfies the Laplace equation. But, like the Sfetsos-Thompson potential, it does not satisfy the second boundary condition in eq. (3.2). Notice that this


Figure 1. For various values of the parameter $\epsilon$ defining the deformation of the Sfetsos-Thompson solution given by eq. (3.19) we show the evolution of the Lyapunov coefficient (whose $t \rightarrow \infty$ limit is the Lyapunov exponent). The initial conditions for our analysis are, $\chi(0)=0.5, \eta(0)=0$, $\dot{\chi}(0)=0.1, \dot{\eta}(0)=0.1$ with the parameter $E$ fixed such that the Hamiltonian vanishes.
deformation is a solution for all values of the parameter $\epsilon$. In the notation of eq. (3.10), we have

$$
\begin{equation*}
F(\eta)=\frac{\eta^{3}}{3}+\frac{\epsilon}{12} \eta^{4}, \quad h_{1}=-\frac{\eta}{2}-\frac{\epsilon}{4} \eta^{2}, \quad h_{2}=\frac{\epsilon}{32}, \quad h_{k>2}=0 . \tag{3.20}
\end{equation*}
$$

It would be interesting to study the geometrical properties of the background generated using eq. (3.1). One can calculate the functions

$$
\begin{align*}
& \mathcal{A}_{\mathrm{def}}=k^{2}+\frac{2 \sqrt{2} E k}{\sqrt{2+\eta \epsilon}} \frac{\left(3+4 \eta^{2}+2 \epsilon \eta+4 \epsilon \eta^{3}+\epsilon^{2} \eta^{4}\right)}{\left(1+4 \eta^{2}+2 \epsilon \eta^{3}\right)} \\
& \mathcal{B}_{\mathrm{def}}=\frac{\left(8+5 \epsilon \eta-4 \epsilon \eta^{3}-2 \epsilon^{2} \eta^{4}\right)}{2 \eta(2+\epsilon \eta)\left(1+4 \eta^{2}+2 \epsilon \eta^{3}\right)} E \tag{3.21}
\end{align*}
$$

and plug them in eq. (3.15). Things go quickly awry. Indeed, even if we simplify these expressions by expanding $\mathcal{A}$ and $\mathcal{B}$ for small values of the deformation parameter $\epsilon$, the differential equation is very hard to solve exactly and we could not find Liouvillian solutions. A more refined analysis that we postpone to appendix A indicates that (at least for small values of the deformation parameter) the solutions to the NVE are non-Liouvillian.

We can complement this with a numerical exploration of the Lyapunov exponent $\hat{\lambda}$ for various values of the deformation parameter $\epsilon$. The results of figure 1 indicate that for the Sfetsos-Thompson background, the corresponding Lyapunov exhibits a sharp fall towards zero, as expected from the analysis in section 2 . On the other hand, when $\epsilon$ is non zero the Lyapunov exponent saturates to some positive value.

Together these point towards the non-integrability of the string soliton moving on this deformation of the Sfetsos-Thompson background. In turn this translates into the non-integrability of the associated $\mathcal{N}=2$ SCFT.

Below, we repeat this analysis for a solution that is qualitatively different to those in eq. (3.9).

### 3.3.3 Study of the Maldacena-Núñez solution

Here, we repeat the study in the previous sections, applying it to the solution of [39], for the case in which there are $N$ D4 branes in Type IIA. In this case, it is useful to work with the $\sigma$-derivative of the potential function $\dot{V}(\sigma, \eta)$,

$$
\begin{equation*}
2 \dot{V}_{M N}(\sigma, \eta)=\sqrt{\sigma^{2}+(N+\eta)^{2}}-\sqrt{\sigma^{2}+(N-\eta)^{2}}, \tag{3.22}
\end{equation*}
$$

and calculate all other derivatives appearing in the Gaiotto-Maldacena background, $\ddot{V}, \dot{V}^{\prime}, V^{\prime \prime}$ close to $\sigma(\tau)=0$ using an expansion of eq. (3.22)

$$
\begin{aligned}
\dot{V}_{M N}(\sigma \sim 0, \eta) & \sim \frac{1}{2} \lambda(\eta)+\frac{Q(\eta)}{4} \sigma^{2}+\frac{Z(\eta)}{16} \sigma^{4}, \\
\lambda(\eta) & =|N+\eta|-|N-\eta|, \quad Q(\eta)=\frac{|N+\eta|}{(N+\eta)^{2}}-\frac{|N-\eta|}{(N-\eta)^{2}}, \\
Z(\eta) & =\frac{|N-\eta|}{(N-\eta)^{4}}-\frac{|N+\eta|}{(N+\eta)^{4}} .
\end{aligned}
$$

The reader can check that $\left.\partial_{\sigma} f_{i}\right|_{\sigma(\tau)=0}=0$ and the $\sigma$-equation in (3.8) is solved for $\sigma(\tau)=$ $\dot{\sigma}(\tau)=\ddot{\sigma}(\tau)=0$. It can also be checked that $\eta=E \tau$ solves the $\eta$-equation close to $\sigma(\tau)=0$. The $\chi(\tau)$-equation after a fluctuation reads as in eq. (3.15) with,

$$
\begin{align*}
& \mathcal{A}_{M N}=k^{2}+E k \frac{1}{\left(2 Q \lambda-\lambda^{\prime 2}\right) \sqrt{2 \lambda Q}}\left[4 Q^{2} \lambda+2 Q \lambda \lambda^{\prime \prime}-4 \lambda^{\prime 2} Q+\lambda^{\prime 2} \lambda^{\prime \prime}-2 Q^{\prime} \lambda \lambda^{\prime}\right] \\
& \mathcal{B}_{M N}=\frac{E}{2 Q \lambda\left(2 Q \lambda-\lambda^{\prime 2}\right)}\left[2 Q^{2} \lambda \lambda^{\prime}-\lambda \lambda^{\prime 2} Q^{\prime}+4 \lambda \lambda^{\prime} \lambda^{\prime \prime} Q-3 \lambda^{\prime 3} Q-2 Q Q^{\prime} \lambda^{2}\right] \tag{3.23}
\end{align*}
$$

Here again, the equation is complicated enough and we don't find Liouvillian solutions. This strongly suggests the non-integrability of the string soliton moving on this background. A more refined analysis, involving necessary conditions for the equations to be integrable, is given in appendix A. There we show that, for evolution in the domain $0 \leq \eta \leq N$ we can make a redefinition that transforms $\mathcal{A}_{M N}$ and $\mathcal{B}_{M N}$ into rational functions and we can see that the criteria of [35] are not satisfied. This shows the Liouvillian non-integrability of the NVE which in turn indicates that the CFT is not-integrable.

The solution of [39] was used by the authors of [40] to construct the general solutions to the Laplace problem in eq. (3.2). Heuristically one can think of the general GaiottoMaldacena background as being provided by a superposition of multiple MN profiles. Thus since a single MN profile leads to non-integrability this is a strong indication of the nonintegrability the string soliton in a generic Gaiotto-Maldacena background. This translates to the non-integrability of the general Gaiotto $\mathcal{N}=2$ SCFT.

Let us close the section by summarising the results. We found that for a generic background in the family of Gaiotto-Maldacena solutions, it is easy to find a string soliton whose equations of motion lead to non-Liouvillian solutions. This implies the non-integrability of the dual $\mathcal{N}=2$ SCFTs. A very interesting exception is the Sfetsos-Thompson solution [38], obtained via non-Abelian T-duality on $A d S_{5} \times S^{5}$. Parts of our analysis used that the potential function $V(\sigma, \eta)$ can be written as in eq. (3.9). This is an important limitation to our approach that we amend with the discussion in appendix B.



Figure 2. Charge density $\lambda(\eta)$ of a "one-kink" spacetime (left) and an "Uluru" spacetime (right) with the values used for the numerical analysis.

## 4 Numerical analysis

Here, complementing the material in previous sections, we provide a light numerical analysis supporting the analytic study of non-integrability in some the Gaiotto-Maldacena backgrounds.

### 4.1 The backgrounds

We shall consider two representative backgrounds, that capture all the features of the Gaiotto-Maldacena solutions and CFTs. We write the associated potential functions $\dot{V}$ as a sum of $\dot{V}_{M N}$ in eq. (3.22).

The first background is dual to a CFT with gauge group $\mathrm{SU}(N) \times \mathrm{SU}(2 N) \times \ldots \times$ $\mathrm{SU}(P N)$ closed by a $\mathrm{SU}((P+1) N)$ flavour group illustrated by the quiver:


The function $\dot{V}(\sigma, \eta)=\sigma \partial_{\sigma} V(\sigma, \eta)$ is given by,

$$
\begin{align*}
\dot{V}= & \frac{N}{2} \sum_{n=-\infty}^{\infty}(P+1)\left[\sqrt{\sigma^{2}+(\eta+P-2 n(1+P))^{2}}-\sqrt{\sigma^{2}+(\eta-P-2 n(1+P))^{2}}\right]+ \\
& +P\left[\sqrt{\sigma^{2}+(\eta-1-P-2 n(1+P))^{2}}-\sqrt{\sigma^{2}+(\eta+1+P-2 n(1+P))^{2}}\right] . \tag{4.1}
\end{align*}
$$

In the following we shall refer to this Gaiotto-Maldacena geometry as a "one-kink" spacetime due to the profile of its charge density shown on the left of figure 2 . Notice that the kink where the charge density changes slope is associated to the flavour group of the quiver. This QFT is of particular interest since it has been proposed and studied in [50] as a completion of the solution of [38]. When performing the numerical study displayed below, we implemented a cut-off on the summation over the index $n$ taking it to range $-10 \leq n \leq 10$ and we took the value $P=7$.

The second background we consider is defined by a periodic "Uluru" profile [40] with the function $\dot{V}$ given as:

$$
\begin{equation*}
\dot{V}(\sigma, \eta)=\frac{N}{2} \sum_{n=-\infty}^{\infty} \sum_{l=1}^{3} \sqrt{\sigma^{2}+\left(\nu_{l}+n \Lambda-\eta\right)^{2}}-\sqrt{\sigma^{2}+\left(\nu_{l}-n \Lambda+\eta\right)^{2}}, \tag{4.2}
\end{equation*}
$$

with $V^{\prime \prime}=-\frac{1}{\sigma^{2}} \ddot{V}$. Here the extra parameters $\Lambda=2 K+4, \nu_{1}=1, \nu_{2}=K+1, \nu_{3}=-K-2$ define a quiver consisting of $K \operatorname{SU}(N)$ gauge group nodes terminated on each end by an $\mathrm{SU}(N)$ flavour node:


In some sense, this is a relative of $\mathcal{N}=2 \mathrm{SQCD}$ with $N_{f}=2 N_{c}$. The charge density, shown on the right of figure 2 , in this case reads

$$
\lambda(\eta)=\left.\dot{V}\right|_{\sigma=0}=\left\{\begin{array}{cc}
N \eta & 0<\eta<1  \tag{4.3}\\
N & 1<\eta<K+1 \\
N-N(\eta-K-1) & K+1<\eta<K+2
\end{array} .\right.
$$

For numerical evaluations we set $N=3$ and $K=10$ and again restrict the summation over the index $n$ ranges $-10 \leq n \leq 10$.

### 4.2 The observables

For both the one-kink and the Uluru spacetimes we will present three numerical plots.
First we will consider the actual profile of trajectories in phase space. Studying eq. (3.14), we found the simple solution $\eta=E \tau$. The angular coordinate was taken to be fixed at $\chi=0$. This trajectory is certainly possible, but when reaching the end of the space (the points $\eta=8$ in the first background and the point $\eta=12$ in the second one), the trajectory should bounce back. Here instead we will study configurations for which $\chi$ is not constant in time. We can plot numerically the trajectories in the $(\eta, \cos \chi)$-plane. We observe the trajectories are not periodic and become very disordered when raising the energy or when more generic initial conditions for $0<\chi \leq \pi$ are imposed. The case of one-kink spacetime is shown in figure 3 and that of the Uluru in figure 6 .

We can then numerically perform a Fourier analysis to obtain the power spectrum [59] of these trajectories. For periodic trajectories we should see well defined (peaked) frequencies. For non-periodic and chaotic trajectories the Fourier analysis reveals a continuum of frequencies. This is borne out as shown in figure 4 for the one-kink spacetime and figure 7 for the Uluru.

Let us move to study the Poincaré sections. Recall that an $N$-dimensional integrable system possesses $N$ independent integrals of motion that are in involution. That is the Poisson bracket of any two of these conserved quantities vanish. The corresponding phase space trajectories are confined to the surface of an $N$-dimensional KAM torus (for a review see [59]). To learn whether a system is integrable or not, one should take cross-sections of its phase-space trajectories. Such a cross-section is known as a Poincaré section [59]. The KAM theorem tells us how these KAM curves will change when we perturb an integrable Hamiltonian with a deformation $H^{\prime}$. The resonant tori, for which these trajectories close on themselves, will be destroyed by this deformation, the motion becomes seemingly random and loosing all of the structure of the KAM curves in our Poincaré section. We fix an energy, find suitable boundary conditions and run a numerical analysis of the Poincaré sections, that is displayed in figure 5 for the one-kink spacetime and figure 8 for the Uluru.

## Numerical plots for the one-kink spacetime



Figure 3. Plots of example trajectories in the $\eta(\tau), \cos (\chi(\tau))$ plane in the one-kink space time. Left we have $E=0.05$ and on the right $E=5.0$.


Figure 4. Power spectra associated to the trajectories in the $\eta(\tau), \cos (\chi(\tau))$ plane displayed in figure 3 for the one-kink space time. Left we have $E=0.05$ and on the right $E=5.0$.


Figure 5. $\eta-P_{\eta}$ plane projections of the Poincare section at $\chi=0$ for the one-kink spacetime. Clockwise from top left we vary the parameter $E=\{0.1,0.5,1,2.5\}$. The plots fill an area bounded by maximal value of $P_{\eta}$ compatible with the constraint that $H=0$ indicated with a grey dashed line. The 100 different seed initial conditions that are numerically evolved to generate these sections are indicated by colour. For small values of $E$ we have a perturbation around an integrable Hamiltonian (for $E=0$ the Hamiltonian is trivial and vanishing) and one sees vestiges of KAM tori which as $E$ is increased dissolve away.

## Numerical plots for the Uluru spacetime



Figure 6. Plots of trajectories in the $\eta(\tau), \cos (\chi(\tau))$ plane in the Uluru space time. Left we have $E=0.1$ with see trajectory confined to the region of space to $11<\eta<12$ and on the right with $E=5.0$ the trajectory wildly explores all of space.


Figure 7. Power spectra associated to the trajectories shown in figure 6. Left we have $E=0.05$ with a see a comparatively clean spectrum and on the right $E=5.0$.


Figure 8. $\eta-P_{\eta}$ plane projections of the Poincare section at $\chi=0$ for the Uluru spacetime. Clockwise from top left $E=\{0.1,0.5,1,2.5\}$. The plots fill an area bounded by maximal value of $P_{\eta}$ compatible with the constraint that $H=0$ indicated with a grey dashed line. The 100 different seed initial conditions that are numerically evolved to generate these sections are indicated by colour. For small values of $E$ we have a perturbation around an integrable Hamiltonian and one sees vestiges of KAM tori which as $E$ is increased dissolve away.

## 5 Conclusions

Let us start with a summary of the paper. We studied the non-integrability properties of generic $\mathcal{N}=2$ SCFTs. The procedure we employed is based on the use of a particular classical string soliton, that rotates and wraps around various compact dimensions associated with the R-symmetry of the CFT. The corresponding field theory operator should have large dimension and R-charges. The study of the Hamiltonian system describing the constrained dynamics of the string, reduce to two nonlinear and coupled ordinary differential equations. We have discussed the presence or not of Liouvillian solutions for these equations using well-established mathematical techniques. We found that non-integrability is generic among the very large family of $\mathcal{N}=2$ SCFTs. Nevertheless, there is one notable exception that is the field theory defined holographically by the Sfetsos-Thompson background [38]. This exceptional case was expected to be singled-out by our approach in light of the results of [45]. Indeed, our string solitons shows no sign of non-integrability. For any deformation away from that background or any other background dual to a generic quiver CFT, we have proven that they will show analytic signs of non-integrability in the Liouville sense.

We have complemented our study with a numerical analysis where the characteristic chaotic indicators were calculated. In fact, we found that for deformations away from the Sfetsos-Thompson solution, the Lyapunov exponent is nonzero. For more generic quiver CFTs we computed the Poincare sections and power spectra, that also display signs of chaotic (hence non-integrable) dynamics.

In the paper [65] the authors showed that $\mathcal{N}=2 \mathrm{SQCD}$, that is the theory with one gauge group $\operatorname{SU}\left(N_{c}\right)$ and $N_{f}=2 N_{c}$, is not integrable. They considered operators of the form

$$
\begin{equation*}
\mathcal{O}=\operatorname{Tr}\left(\phi^{k_{1}} M^{k_{2}} \phi^{k_{3}} M^{k_{4}} \ldots\right), \tag{5.1}
\end{equation*}
$$

where $\phi$ is some adjoint operator made out of fields in the vector multiplet and $M$ is a dimer operator in the adjoint, constructed out of fundamental field with the flavour index contracted. Calculating two body S-matrix scattering, they showed that these dimer fields play an important role in the absence of integrability.

The conjecture that there may be a subsector of the theory that preserves integrability was proposed in [65] and made more concrete in [66]. In fact, when only fields belonging to the vector multiplet are considered, the integrable structure is inherited from that of $\mathcal{N}=4$ SYM, up to a rescaling of the gauge coupling.

These perturbative results complement the findings of our study, that is nonperturbative in nature. We suggest that for quiver theories with fundamental fields, the presence of the flavour group is at the root of the non-integrability we found. Indeed, when our strings explore the regions close to the end of the quiver our numerical study shows the KAM tori becoming diffuse. More analytically, a deviation from the Sfetsos-Thompson solution (by the addition of a flavour group), also displayed non-integrability. Also, the conjecture that the sub-class of operators not containing dimer fields inherit their integrable properties from $\mathcal{N}=4$ SYM is nicely mirrored in our approach. In fact, strings uncharged under $\mathrm{SU}(2)_{R}$ - with $k=0$ in eq. (3.5), but charged under $\mathrm{U}(1)_{r}$ (with nonzero $\lambda$ ), are
dual to an long operator made out of vector multiplet fields only. The analysis shows integrability of those solitons.

In summary, this work belongs to a line of studies of $\mathcal{N}=2$ SCFTs and holographically shows that generically they are not-integrable. The paper opens some interesting topics for exploration. We list a few of these below and hope to return to them in future work.

- It would be interesting to understand in more detail the holographic background we associated with the deformation away from the Sfetsos-Thompson solution see around eq. (3.9). The calculation of some observables, charges and structure of singularities in terms of the deformation parameter should give clues about the dual CFT.
- It seems natural to extend our study to CFTs with similar characteristics. Namely to holographic backgrounds with AdS-factors and some $\operatorname{SU}(2)$ isometry, preserving eight Poincare supercharges. Some examples come to mind. The Gaiotto-Witten CFTs with holographic dual summarised in [61] and the recently discussed five dimensional CFTs based on intersections of D5-NS5-D7 branes summarised in [62]. Notice that both systems admit a background obtained via non-Abelian T-duality (hence analog to Sfetsos-Thompson), given in the works [63] and [64] respectively.
- The zero-winding case - see eq. (3.15) for the case $k=0$ - becomes integrable in the Liouville sense. This translates to the integrability of this particular geodesic motion. It seems plausible, following the ideas of [25], to study generically the integrability of geodesics in Gaiotto-Maldacena backgrounds.
- Finally and more interestingly, it would be nice to clearly establish what characteristic of the Gaiotto CFT is introducing the non-integrability and/or the chaotic behaviour. One intuitive answer indicates that is the presence of fundamental fields (flavours) that is producing Poincare sections with signs of chaos. One should like a more precise identification of the corresponding operators, like that in eq. (5.1) that drive this feature.

We hope to report on these issues in the future.

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## A Kovacic's algorithm and rudiments of differential Galois theory

In this article we need to establish if certain linear ordinary differential equations are integrable in the Liouvillian sense. That is to say we wish to know if a differential equation is "symmetric" enough to admit solutions given in closed form in terms of a finite composition of elementary functions. The symmetries here are linear transformations of the space of solutions that respect both algebraic and differential relations between solutions. These symmetries are described by Picard-Vessiot or differential-Galois theory. Since this is a concept less familiar amongst physicists we give a brief orientation below. ${ }^{7}$

We will describe two different types of arguments. One is based on usual lore and manipulations with differential equations, the other on Lie-group theoretical arguments (Picard-Vessiot theory). Both developments are familiar to mathematicians and were used by Kovacic in his work [35]

## A. 1 The differential equation approach

Here, we briefly describe Kovacic's algorithm [35]. Let us start considering a linear differential equation,

$$
\begin{equation*}
y^{\prime \prime}(x)+\mathcal{B}(x) y^{\prime}(x)+\mathcal{A}(x) y(x)=0 \tag{A.1}
\end{equation*}
$$

where $\mathcal{A}(x), \mathcal{B}(x)$ are complex rational functions. We are concerned with the existence of solutions that can be expressed in terms of algebraic functions, exponentials, trigonometric and integrals of the above. These are called 'Liouvillian' solutions.

The algorithm of [35] provides such solutions or shows that there are none (in such case we refer to the equation (A.1) as non-integrable). We will not describe the algorithm in itself, as it is already implemented in different softwares. We limit ourselves to explain the 'logic' behind the derivation and some necessary but not-sufficient conditions that a combination the functions $\mathcal{A}, \mathcal{B}, \mathcal{B}^{\prime}$ must satisfy, for the eq. (A.1) to be Liouville integrable.

We start by rewriting the differential equation as,

$$
\begin{align*}
y(x) & =e^{\int w(x)-\frac{\mathcal{B}(x)}{2} d x} \\
w^{\prime}(x)+w(x)^{2} & =V(x)=\frac{2 \mathcal{B}^{\prime}+\mathcal{B}^{2}-4 \mathcal{A}}{4} \tag{A.2}
\end{align*}
$$

It was shown that if the function $w(x)$ is algebraic of degrees $1,2,4,6$, or 12 , then the eq. (A.1) is Liouville integrable [35]. This results comes from the application of Galois theory to differential equations (which is Piccard-Vessiot theory). As we shall discuss below, this formalism studies the most general group of invariances of the differential equation (A.1), the transformations that act on the solutions of the equation, that is a subgroup of $\operatorname{SL}(2, C)$. Kovacic [35] showed that there are four possible cases of subgroups of $\operatorname{SL}(2, C)$ that can arise

- Case 1: the subgroup is generated by the matrix of the form

$$
G=\left[\begin{array}{ll}
a & 0 \\
b & \frac{1}{a}
\end{array}\right],
$$

with $a, b$ complex numbers. In this case $w(x)$ is a rational function of degree 1 .

[^5]- Case 2: the subgroup of $\mathrm{SL}(2, C)$ is generated by matrices of the form

$$
G=\left[\begin{array}{cc}
c & 0 \\
0 & \frac{1}{c}
\end{array}\right], \quad G=\left[\begin{array}{cc}
0 & c \\
-\frac{1}{c} & 0
\end{array}\right],
$$

in this case the functions $w(x)$ is rational of degree 2

- Case 3: the situation in which $G$ is a finite group, not included in the two above cases. In such case, the degree of $w(x)$ is either 4,6 or 12 .
- Case 4: the group is $\mathrm{SL}(2, C)$ and the solution for $w(x)$, if they exist are nonLiouvillian.

Kovacic provided not only an algorithm to find the solutions in the first three cases above, but also a set of necessary but not sufficient conditions that the function $V(x)$ in eq. (A.2) must satisfy to be in any of the first three cases detailed above [35]. For each of the cases as ordered above, the conditions are:

- Case 1: every pole of $V(x)$ has order 1 or has even order. The order of the function $V(x)$ at infinity ${ }^{8}$ is either even or greater than 2.
- Case 2: $V(x)$ has either one pole of order 2 , or poles of odd-order greater than 2 .
- Case 3: the order of the poles of $V$ does not exceed 2 , and the order of $V$ at infinity is at least 2 .

If none of the above is satisfied, the analytic solution (if it exists), is non-Liouvillian. In section A.3, we shall explore some examples in the main part of this paper in light of these statements. Before that, let us discuss the group theoretical Picard-Vessiot viewpoint of the ideas in this section.

## A. 2 The group theoretical approach

Consider a homogenous $n^{\text {th }}$ order linear differential equation $L(y)=y^{(n)}+a_{n-1} y^{(n-1)}+\cdots+$ $a_{0} y=0$ with coefficients in some differential field $k$ with constants $\mathbb{C}$. The $n$-independent solutions of $L(y)$ form a vector space $V$ over $\mathbb{C}$. A differential field $K$ is said to be a PicardVessiot (PV) extension of $k$ for $L$ if $K$ is generated over $k$ by the solutions of $L$. A PV extension - whose existence and uniqueness was shown by Kolchin - is thus the smallest such extension of $k$ that contains the $n$-independent solutions of $L(y)$. The differential Galois group $G=\partial \operatorname{Gal}(K / k)$ is the group of automorphisms of the PV extension $K$ that commute with the derivative and which leave elements of $k$ fixed. The condition that $L(y)$ be Liouvillian integrable is now more formally stated as demanding that $G^{0}$, the identity connected component of the differential Galois group is solvable. If $G^{0}$ is Abelian then $L(y)=0$ is integrable.

[^6]To make this rather more concrete let us do a trivial example. $y^{\prime \prime}(x)+\frac{1}{x} y^{\prime}(x)=0$ is an integrable equation with a solution space $V=1 \oplus \log (x)$. Since 1 is just a constant of the base field any $\sigma \in G$ should obey $\sigma(1)=1$. For the action of $\sigma$ on $\log (x)$ we have $\partial \sigma(\log (x))=\sigma(\partial \log (x))=\sigma\left(\frac{1}{x}\right)=\frac{1}{x}$ in which we used that $\sigma$ commutes with derivation and that $\frac{1}{x}$ is by definition left invariant by $\sigma$. We can then extract $\sigma(\log (x))$ by integration to find

$$
\sigma:(1, \log (x)) \mapsto(1, \log (x))\left(\begin{array}{ll}
1 & c  \tag{A.3}\\
0 & 1
\end{array}\right), \quad c \in \mathbb{C},
$$

and so that the Galois group is simply the Abelian additive group on $\mathbb{C}$ formed by the composition of such $\sigma$. This is an example of Case 1 in the preceding discussion. In general, unlike in this example, the construction of differential Galois groups is not facile, however can be achieved algorithmically (for second order equations one has an algorithm due to Kovacic and in generality from Hrushovski).

The concept of differential Galois group is at first sight similar to that of monodromy. Transporting the Wronskian matrix of fundamental solutions $\mathbf{Z}$ around some closed path $\gamma$ avoiding singularities of $L(y)$ generates some new fundamental solution matrix $\mathbf{Z}_{\gamma}=\mathbf{Z} M_{\gamma}$. Providing that the singularities of $L$ are regular (i.e. $a_{i}$ has a pole of at most order $n-i$ at singular points ) then monodromy is dense in the Galois group. However in general singularities need not be regular. In this case a theorem by Ramis establishes the Galois group as being generated by formal monodromy that comes from $x \rightarrow x e^{2 \pi i}$, the so-called Ramis/exponential torus (a subgroup of $\left(\mathbb{C}^{*}\right)^{n}$ whose precise definition we shall not need) and the Stokes multipliers.

This is well exemplified by the Bessel equation,

$$
\begin{equation*}
x^{2} y^{\prime \prime}(x)+x y^{\prime}(x)+\left(x^{2}-n^{2}\right) y=0 \tag{A.4}
\end{equation*}
$$

in which $x=0$ is a regular point and $x=\infty$ is irregular. To simplify matters let us just consider $\nu$ not an integer for which $J_{+\nu}(x)$ and $J_{-\nu}(x)$ are independent solutions. Around $x=0$ these are described by convergent series $J_{\nu}(x)=\frac{1}{\Gamma(1+\nu)}\left(\frac{x}{2}\right)^{\nu}\left(1-\frac{1}{1+\nu}\left(\frac{x}{2}\right)^{2}+O\left(x^{4}\right)\right)$ with monodromy $J_{\nu}\left(e^{2 \pi i} x\right)=e^{2 \pi i \nu} J_{\nu}(x)$. Hence the monodromy around the origin is simply $M_{0}=\operatorname{diag}\left(e^{2 \pi i \nu}, e^{-2 \pi i \nu}\right)$. On the other hand expansions around $x=\infty$ are asymptotic. Translating into first order equations by introducing the vector $\mathbf{Y}(x)=\left(y(x), y^{\prime}(x)\right)$ and letting $x=z^{-1}$ we have

$$
d_{z} \mathbf{Y}=\mathbf{A} \cdot \mathbf{Y}, \quad \mathbf{A}=\left(\begin{array}{cc}
0 & 1  \tag{A.5}\\
-\frac{1}{z}-\frac{1}{z^{4}}+\frac{\nu^{2}}{z^{2}}
\end{array}\right) .
$$

The formal fundamental solution matrix $\mathbf{Z}$ which obeys the same equation can be factored as

$$
\begin{equation*}
\mathbf{Z}=\hat{\phi}(z) z^{\mathbf{L}} e^{Q(z)} \tag{A.6}
\end{equation*}
$$

in which $\hat{\phi}(z)$ is a matrix consisting of formal (i.e. asymptotic) power-series $\sum_{0}^{\infty} a_{n} z^{n}$ and in the case at hand $\mathbf{L}=\frac{1}{2} \mathbf{1}$ and $Q=\operatorname{diag}\left(q_{1}, q_{2}\right)=\operatorname{diag}\left(\frac{i}{z}, \frac{-i}{z}\right)$. The formal monodromy at infinity is sensitive only to the square root in $z^{L}$ and so $\hat{M}_{\infty}=-\mathbf{1}$. Within a given sector
the formal $\hat{\phi}(z)$ can be Borel resumed however as the phase of $z$ is varied the result of this procedure changes exactly as one crosses singular directions (anti-stokes rays). These occur at in a direction $\arg (z)=\theta$ for which $\operatorname{Re}\left(q_{1}-q_{2}\right)=0$. It is in these directions that dominant and sub-dominant asymptotic behaviours switch roles. The summation before and after crossing the singular direction specified by $\arg (z)=\theta$ must then be related via $\phi_{+}(z)=\phi_{-}(z) \mathcal{S}_{\theta}$ where $\mathcal{S}_{\theta}$ called a Stokes multiplier. For the Bessel function the singular directions are $\theta=\frac{\pi}{2}, \frac{3 \pi}{2}$ and corresponding multipliers are

$$
\mathcal{S}_{\frac{\pi}{2}}=\left(\begin{array}{ll}
1 & \lambda  \tag{A.7}\\
0 & 1
\end{array}\right), \quad \mathcal{S}_{\frac{3 \pi}{2}}=\left(\begin{array}{cc}
1 & 0 \\
-\mu & 1
\end{array}\right)
$$

We don't need to evaluate the mulitpliers directly, instead we follow the nice argument in [53] to relate the actual monodromy at the origin to the formal monodromy at $\infty$ via $M_{0}=\mathcal{S}_{\frac{\pi}{2}} \cdot \hat{M}_{\infty} \cdot \mathcal{S}_{\frac{3 \pi}{2}}$ and by taking a trace of this relation we constrain the Stokes multipliers to obey

$$
\begin{equation*}
\lambda \mu=4 \cos ^{2}(\pi n) \tag{A.8}
\end{equation*}
$$

If $n \notin \mathbb{Z}+\frac{1}{2}$ the Stokes multipliers are not-vanishing and $\mathcal{S}_{\frac{\pi}{2}}, \mathcal{S}_{\frac{3 \pi}{2}}$ do not commute; hence $G^{0}$ is non-Abelian and the Bessel equation is not Liouvillian integrable. On the other-hand when $n \in \mathbb{Z}+\frac{1}{2}$ the formal series in $\hat{\phi}(z)$ actually terminate and the Stokes multipliers vanish and the Bessel equation becomes integrable. Indeed we have that e.g.

$$
\begin{equation*}
J_{\frac{1}{2}}(x)=\left(\frac{2}{\pi x}\right)^{\frac{1}{2}} \sin x \tag{A.9}
\end{equation*}
$$

## A. 3 Some examples

Along the lines of the discussion above, let us study the function $V(x)$ in eq. (A.2) for some of the examples in this paper. We start with the Sfetsos-Thompson solution. The associated NVE differential equation was discussed in eq. (3.17). The coefficientes $\mathcal{A}, \mathcal{B}$ read in this case (we take $k=E=1$ to avoid a cluttered notation),

$$
\begin{equation*}
\mathcal{A}=1+2 \frac{\left(4 \tau^{2}+3\right)}{4 \tau^{2}+1}, \quad \mathcal{B}=\frac{2}{4 \tau^{3}+\tau} \tag{A.10}
\end{equation*}
$$

The coefficients are rational, so we construct the potential function

$$
\begin{equation*}
4 V(\tau)=4 \mathcal{A}^{2}-\mathcal{B}^{2}-2 \mathcal{B}^{\prime}=3+\frac{12}{\left(4 \tau^{2}+1\right)^{2}}+\frac{4}{4 \tau^{2}+1}+\frac{\gamma}{\tau^{2}} \tag{A.11}
\end{equation*}
$$

The last term has a coefficient $\gamma=E(E-1)$, that vanish for our particular choice of constants. Note that his function is the one appearing in eq. (3.18) and after a rescaling and choice of constants, in eq. (2.15). Analysing the $V$-function, we see that it satisfies all the three possible necessary conditions listed above. Hence, if there is a Liouvillian solution, Kovacic's algorithm should find it. Indeed, this happens when feeding the differential equation to any software, the (Liouvillian) solution in eq. (2.16) is found.

We now move to study the case of the deformed Sfetsos-Thompson solution in eqs. (3.19)-(3.21). For this analysis to be simpler, we shall consider the deformation parameter $\epsilon \sim 0$ and keep only up to linear order in a series expansion in $\epsilon$. In this case the coefficients in eq. (3.21) are

$$
\begin{equation*}
\mathcal{A}=\frac{\left(12 \tau^{2}+7\right)}{4 \tau^{2}+1}+\epsilon\left(\frac{5 \tau+8 \tau^{3}+16 \tau^{5}}{2\left(4 \tau^{2}+1\right)^{2}}\right), \quad \mathcal{B}=\frac{2}{4 \tau^{3}+\tau}+\epsilon\left(\frac{1-16 \tau^{2}-16 \tau^{4}}{4\left(4 \tau^{2}+1\right)^{2}}\right) \tag{A.12}
\end{equation*}
$$

The associated potential function is,

$$
\begin{equation*}
V=\frac{19+40 \tau^{2}+48 \tau^{4}}{\left(4 \tau^{2}+1\right)^{2}}+\frac{\epsilon}{4 \tau\left(4 \tau^{2}+1\right)^{3}}\left(128 \tau^{8}+96 \tau^{6}+40 \tau^{4}+50 \tau^{2}-1\right) \tag{A.13}
\end{equation*}
$$

We can check that the three criteria are failed. So the solution to the NVE equation is non-Liouvillian.

Finally, let us analyse the case of the Maldacena-Núñez solution in section 3.3.3. In this case $V_{M N}$ solution of the Laplace equation dictates the functions $\mathcal{A}, \mathcal{B}$ for $0 \leq \eta \leq N$ to be

$$
\begin{align*}
\mathcal{A} & =k^{2}+\sqrt{2} E k \frac{\left(3 N^{2}+\eta^{2}\right)}{\left(N^{2}+\eta^{2}\right) \sqrt{\left|\eta^{2}-N^{2}\right|}} \\
\mathcal{B} & =E \frac{\left(\eta^{4}+3 N^{2} \eta^{2}-2 N^{4}\right)}{\eta\left(\eta^{4}-N^{4}\right)} . \tag{A.14}
\end{align*}
$$

For $N \leq \eta$ we find,

$$
\begin{aligned}
& \mathcal{A}=k^{2}+\sqrt{2} E k \frac{1}{\sqrt{\left|\eta^{2}-N^{2}\right|}} \\
& \mathcal{B}=E \frac{\eta}{\left(\eta^{2}-N^{2}\right)} .
\end{aligned}
$$

Here we understand that $\eta=E \tau$. We simplify the analysis by choosing again all arbitrary constants $E=k=N=1$ and considering the system in the interval $[0,1]$ by choosing appropriate initial conditions. Thus

$$
\begin{equation*}
\mathcal{A}=1+\sqrt{2} \frac{\left(\tau^{2}+3\right)}{\left(\tau^{2}+1\right) \sqrt{1-\tau^{2}}}, \quad \mathcal{B}=-\frac{\tau^{4}+3 \tau^{2}-2}{-\tau^{5}+\tau} . \tag{A.15}
\end{equation*}
$$

In this case, the coefficients are not rational. We should then change variables as,

$$
\begin{equation*}
\tau=\sqrt{1-z^{2}}, \quad \frac{d z}{d \tau}=-\frac{\sqrt{1-z^{2}}}{z}, \quad \frac{d}{d z}\left(\frac{d z}{d \tau}\right)=\frac{1}{z^{2} \sqrt{1-z^{2}}}, \quad \frac{d}{d z} \frac{d z}{d \tau}=\frac{1}{z^{2} \sqrt{1-z^{2}}} \tag{A.16}
\end{equation*}
$$

The derivatives change according to,

$$
\begin{aligned}
& \dot{x}=\frac{d x}{d \tau}=\frac{d x}{d z} \frac{d z}{d \tau}=-x^{\prime}(z) \frac{\sqrt{1-z^{2}}}{z} \\
& \ddot{x}=x^{\prime \prime}\left(\frac{d z}{d \tau}\right)^{2}+x^{\prime} \frac{d}{d z}\left(\frac{d z}{d \tau}\right) \times \frac{d z}{d \tau}
\end{aligned}
$$

The NVE equation changes according to,

$$
\begin{align*}
& \ddot{x}(\tau)+\mathcal{B} \dot{x}(\tau)+\mathcal{A} x(\tau)=0 \rightarrow x^{\prime \prime}(z)+\mathcal{C} x^{\prime}(z)+\mathcal{D} x(z)=0, \\
& \mathcal{C}=\frac{\mathcal{B}+\frac{d}{d z}\left(\frac{d z}{d \tau}\right)}{\frac{d z}{d \tau}}=-z \frac{\mathcal{B}(z)+\frac{1}{z^{2} \sqrt{1-z^{2}}}}{\sqrt{1-z^{2}}}=\frac{z^{3}-4 z}{z^{4}-3 z^{2}+2}, \\
& \mathcal{D}=\frac{\mathcal{A}}{\left(\frac{d z}{d \tau}\right)^{2}}=\frac{z^{2}}{1-z^{2}} \mathcal{A}(z)=-\frac{\sqrt{2} z\left(z^{2}-4\right)+z^{4}-2 z^{2}}{z^{4}-3 z^{2}+2} . \tag{A.17}
\end{align*}
$$

Plugging this in the function $-4 V(z)=2 \mathcal{C}^{\prime}+\mathcal{C}^{2}-4 \mathcal{D}$, we find

$$
\begin{equation*}
V(z)=4+\frac{10-z^{2}}{\left(z^{4}-3 z^{2}+2\right)^{2}}+\frac{4 \sqrt{2} z^{3}+5 z^{2}-16 \sqrt{2} z-13}{z^{4}-3 z^{2}+2} \tag{A.18}
\end{equation*}
$$

In this case we see that the three necessary criteria are failed. The solution to the NVE is non-Liouvillian.

## B Non-integrability in generic Gaiotto-Maldacena backgrounds

We will consider a generic Gaiotto-Maldacena background. All the elements of the metric and other fields can be written in terms of the function $\dot{V}(\sigma, \eta)$ and its derivatives.

Generically, the potential function and its derivatives are [40, 41],

$$
\begin{array}{ll}
\dot{V}(\sigma, \eta)=\sum_{n=1}^{\infty} A_{n}\left(\omega_{n} \sigma\right) K_{1}\left(\omega_{n} \sigma\right) \sin \left(\omega_{n} \eta\right), & \dot{V}^{\prime}(\sigma, \eta)=\sum_{n=1}^{\infty} A_{n} \omega_{n}\left(\omega_{n} \sigma\right) K_{1}\left(\omega_{n} \sigma\right) \cos \left(\omega_{n} \eta\right), \\
\ddot{V}(\sigma, \eta)=\sum_{n=1}^{\infty} A_{n}\left(\omega_{n} \sigma\right)^{2} K_{0}\left(\omega_{n} \sigma\right) \sin \left(\omega_{n} \eta\right), & V^{\prime \prime}(\sigma, \eta)=-\sum_{n=1}^{\infty} A_{n} \omega_{n}^{2} K_{0}\left(\omega_{n} \sigma\right) \sin \left(\omega_{n} \eta\right) .
\end{array}
$$

Here, we used that $\omega_{n}=\frac{n \pi}{N_{5}}$ and that for $\sigma=0$ we have,

$$
\tilde{\lambda}(\eta)=\dot{V}(0, \eta)=\sum_{n=1}^{\infty} A_{n} \sin \left(\omega_{n} \eta\right) .
$$

In other words, the coefficients $A_{n}$ are the Fourier decomposition of $\tilde{\lambda}(\eta)$.
The authors of the paper [41] proposed an expansion close to $\sigma=0$ for a value of $\eta=\eta_{i}$ where a change in slopes is found. They set

$$
\sigma=r \cos \theta, \quad \eta=\eta_{i}+r \sin \theta,
$$

and found the expression for the metric and background fields close to that point - see section 4.2.2 in [41]. The fields relevant to our purposes are,

$$
\begin{align*}
d s^{2} & =\sqrt{g(r)}\left[4 A d S_{5}+d \chi^{2}+\sin ^{2} \chi d \xi^{2}\right]+\frac{1}{\sqrt{g(r)}}\left[d r^{2}+r^{2} d \theta^{2}+r^{2} \sin ^{2} \theta d \beta^{2}\right] \\
B_{2} & =-2 \eta_{i} \sin \chi d \chi \wedge d \xi \tag{B.1}
\end{align*}
$$

The function $g(r)=4 \frac{\lambda\left(\eta_{i}\right)}{\left(a_{i-1}-a_{i}\right)} r$.

As in previous sections, we propose a configuration of the form,

$$
t=t(\tau), \quad r=r(\tau), \quad \theta=\theta(\tau), \quad \chi=\chi(\tau), \quad \xi=k \sigma, \quad \beta=\nu \sigma .
$$

The effective Lagrangian is,

$$
\begin{equation*}
L_{\mathrm{eff}}=\sqrt{g}\left[4 \dot{t}^{2}+k^{2} \sin ^{2} \chi-\dot{\chi}^{2}\right]+\frac{1}{\sqrt{g}}\left[-\dot{r}^{2}-r^{2} \dot{\theta}^{2}+\nu^{2} r^{2} \sin ^{2} \theta\right]+4 k \eta_{i} \sin \chi \dot{\chi} . \tag{B.2}
\end{equation*}
$$

Notice that the last term, the one induced by the B-field is a total derivative, as expected, since the B-field turns out to be pure gauge in the expansion.

We calculate the equations of motion. We find the usual conservation equation, $\dot{t}=$ $\frac{E}{\sqrt{g(r)}}$. Replacing this in the other equations, we have,

$$
\begin{align*}
4 g(r) \ddot{r} & =g^{\prime}(r)\left(\dot{r}^{2}+4 E^{2}\right)+\left(4 r g(r)-r^{2} g^{\prime}(r)\right)\left(\dot{\theta}^{2}-\nu^{2} \sin ^{2} \theta\right)+g(r) g^{\prime}(r)\left(\dot{\chi}^{2}-k^{2} \sin ^{2} \chi\right) . \\
2 g(r) r \ddot{\theta} & =-\nu^{2} g(r) r \sin (2 \theta)-\left(r g^{\prime}(r)-4 g(r)\right) \dot{r} \dot{\theta}, \\
2 g(r) \ddot{\chi} & =-k^{2} g(r) \sin (2 \chi)-g^{\prime}(r) \dot{r} \dot{\chi} . \tag{B.3}
\end{align*}
$$

We observe that for $\theta=\dot{\theta}=\ddot{\theta}=0$ and $\chi=\dot{\chi}=\ddot{\chi}=0$, we solve automatically the $\theta$ and $\chi$-equations in (B.3). If we replace this in the $r$-equation we find,

$$
\begin{equation*}
4 g(r) \ddot{r}-g^{\prime}(r)\left(4 E^{2}+\dot{r}^{2}\right)=0 . \tag{B.4}
\end{equation*}
$$

This equation has a complicated solution in terms of exponential, cubic roots, etc. We will not write it here.

A small fluctuation around the $\chi=0+\epsilon x(\tau)$ and $\theta=0+\epsilon y(\tau)$, gives at leading order in $\epsilon$ the equations,

$$
\begin{array}{r}
\ddot{x}+\frac{g^{\prime}(r) \dot{r}}{2 g(r)} \dot{x}+k^{2} x=0, \\
\ddot{y}+\frac{\left(2 g(r) \dot{r}-g^{\prime}(r) r \dot{r}\right)}{g(r) r} \dot{y}+\nu^{2} x=0 . \tag{B.5}
\end{array}
$$

One can calculate the time-dependent coefficients of the 'friction terms' and attempt to solve the complicated equations. This leads to solutions involving a complicated combination of special functions.

In the very simple case in which the integration constant $E=0$, the solution of eq. (B.4) is $r=t^{4 / 3}$. Then eq. (B.5) leads to a simple solution in terms of Bessel functions. These are not Liouvillian solutions, hence the string soliton is non-integrable.

In summary, we find non-integrability in Gaiotto Maldacena backgrounds, with the string soliton at a generic kink-point.

## C Integrability in the non-Abelian T-dual of $A d S_{5} \times S^{5}$

The preservation of integrability under non-Abelian T-duality of various (super)-coset string sigma-models was shown in $[44,45]$. In this appendix we reinforce these results
by giving a very explicit treatment of the integrability of the bosonic theory with an $S^{5}$ target space dualised with respect to an $\operatorname{SU}(2)_{L}$ isometry acting inside the sphere. Our considerations apply equally well to the bosonic sector of the $\operatorname{Ad} S_{5} \times S^{5}$ string since the $A d S_{5}$ space is decoupled from the five sphere (except via Virasoro constraints).

We view the $S^{5}$ as a coset $\mathrm{SO}(6) / \mathrm{SO}(5)$ and introduce an explicit representation for the algebra $\mathfrak{g}=\mathfrak{s u}(4) \cong \mathfrak{s o}(6)$. We follow the conventions of [60] first define 4d Dirac matrices ${ }^{9}$ which obey the Clifford algebra $\left\{\gamma_{i}, \gamma_{j}\right\}=2 \delta_{i j}$ and supplement them with $\gamma_{5}=-\gamma_{1} \gamma_{2} \gamma_{3} \gamma_{4}$. Generators of $\mathfrak{s o}(5) \in \mathfrak{s o}(6)$ are given by $n_{i j}=-n_{j i}=\frac{1}{4}\left[\gamma_{i}, \gamma_{j}\right]$ and the remaining five coset generators are given by $n_{i 6}=\frac{i}{2} \gamma_{i}$. The $\mathfrak{s u}(4)$ algebra is $\mathbb{Z}_{2}$ graded by

$$
\begin{equation*}
\Omega(\mathfrak{g})=K \cdot \mathfrak{g} \cdot K^{-1}, \quad K=-\gamma_{2} \gamma_{4} . \tag{C.1}
\end{equation*}
$$

Denoting $\mathfrak{g}^{(k)}=\left\{X \in \mathfrak{g} \mid \Omega(X)=i^{k} X\right\}$ we have $\mathfrak{g}=\mathfrak{g}^{(0)}+\mathfrak{g}^{(2)}$ with $\mathfrak{g}^{(0)}$ being the $\mathfrak{s o}(5)$ subgroup and $\mathfrak{g}^{(2)}$ the coset generators. Notice that this is a symmetric coset since $\left[\mathfrak{g}^{(0)}, \mathfrak{g}^{(0)}\right] \in$ $\mathfrak{g}^{(0)},\left[\mathfrak{g}^{(2)}, \mathfrak{g}^{(2)}\right] \in \mathfrak{g}^{(0)}$ and $\left[\mathfrak{g}^{(0)}, \mathfrak{g}^{(2)}\right] \in \mathfrak{g}^{(2)}$. We let $P^{(k)}$ be the projector onto $\mathfrak{g}^{(k)}$.

To define the sigma model on the $S^{5}$ we introduce a coset representative $G$ from which we construct an algebra valued one-form

$$
\begin{equation*}
\mathfrak{a}=-G^{-1} d G=\mathfrak{a}^{(0)}+\mathfrak{a}^{(2)} . \tag{C.2}
\end{equation*}
$$

The Lagrangian is then given by the pull back

$$
\begin{equation*}
\mathcal{L}=\operatorname{Tr}\left(\mathfrak{a}_{+}^{(2)} \mathfrak{a}_{-}^{(2)}\right), \tag{C.3}
\end{equation*}
$$

where $\sigma^{ \pm}=\tau \pm \tilde{\sigma}$ are world sheet light cone coordinates. The equations of motion and Bianchi identities of this theory are encapsulated by a Lax connection with light cone components

$$
\begin{equation*}
\mathfrak{L}_{ \pm}=\mathfrak{a}_{ \pm}^{(0)}+\left(u_{1} \mp u_{2}\right) \mathfrak{a}_{ \pm}^{(2)}, \quad u_{1}^{2}-u_{2}^{2}=1 \tag{C.4}
\end{equation*}
$$

and which obeys the flatness condition

$$
\begin{equation*}
\left[\partial_{+}+\mathfrak{L}_{+}, \partial_{-}+\mathfrak{L}_{-}\right]=0 \tag{C.5}
\end{equation*}
$$

Various options are available for the parametrisation of the coset and to proceed explicitly we must pick one. We will make a choice that will most easily allow us to perform the T-dualisation of the sigma model with respect to an $\mathrm{SU}(2)$ generated by

$$
\begin{equation*}
T_{1}=-\frac{1}{2}\left(n_{12}+n_{34}\right), \quad T_{2}=-\frac{1}{2}\left(n_{13}-n_{24}\right), \quad T_{3}=-\frac{1}{2}\left(n_{14}+n_{23}\right), \quad\left[T_{a}, T_{b}\right]=\epsilon_{a b c} T_{c} . \tag{C.6}
\end{equation*}
$$

$$
\gamma_{1}=\left(\begin{array}{cccc}
0 & 0 & 0 & -1 \\
0 & 0 & 1 & 0 \\
0 & 1 & 0 & 0 \\
-1 & 0 & 0 & 0
\end{array}\right), \quad \gamma_{2}=\left(\begin{array}{cccc}
0 & 0 & 0 & i \\
0 & 0 & i & 0 \\
0 & -i & 0 & 0 \\
-i & 0 & 0 & 0
\end{array}\right), \gamma_{3}=\left(\begin{array}{cccc}
0 & 0 & 1 & 0 \\
0 & 0 & 0 & 1 \\
1 & 0 & 0 & 0 \\
0 & 1 & 0 & 0
\end{array}\right), \quad \gamma_{4}=\left(\begin{array}{cccc}
0 & 0 & -i & 0 \\
0 & 0 & 0 & i \\
i & 0 & 0 & 0 \\
0 & -i & 0 & 0
\end{array}\right)
$$

We will then mimic the standard Euler angles and define

$$
\begin{equation*}
G=H \hat{G}, \quad H=\exp \left[\phi T_{3}\right] \cdot \exp \left[\theta T_{2}\right] \cdot \exp \left[\psi T_{3}\right], \quad \hat{G}=\exp \left[\frac{i}{2} \tilde{\phi} \gamma_{5}\right] \cdot \frac{1}{\sqrt{1+r^{2}}}\left(1+i r \gamma_{3}\right) \tag{C.7}
\end{equation*}
$$

The target space metric of the Lagrangian eq. (C.3) reads

$$
\begin{equation*}
d s^{2}=d \alpha^{2}+\sin ^{2} \alpha d \tilde{\phi}^{2}+\frac{1}{4} \cos ^{2} \alpha\left(l_{1}^{2}+l_{2}^{2}+l_{3}^{2}\right) \tag{C.8}
\end{equation*}
$$

in which we use the $\mathrm{SU}(2)$ left invariant one-forms

$$
\begin{equation*}
l_{1}=\sin \theta \cos \psi d \phi-d \theta \sin \psi, \quad l_{2}=\cos \psi(\sin \theta \tan \psi d \phi+d \theta), \quad l_{3}=\cos \theta d \phi+d \psi \tag{C.9}
\end{equation*}
$$

and the angle $\sin \alpha=\frac{1-r^{2}}{1+r^{2}}$. The virtue of this parametrisation is that we have chosen a gauge for the Lax connection that makes the $\mathrm{SU}(2)$ isometry manifest. Explicitly we find that

$$
\begin{equation*}
\mathfrak{a}^{(0)}=\cos \alpha d \tilde{\phi} n_{35}+\frac{1}{2} l_{1}\left(n_{12}+\sin \alpha n_{34}\right)+\frac{1}{2} l_{2}\left(-n_{24}+\sin \alpha n_{13}\right)+\frac{1}{2} l_{3}\left(n_{14}+\sin \alpha n_{23}\right) \tag{C.10}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathfrak{a}^{(2)}=-d \alpha n_{36}+\sin \alpha d \tilde{\phi} n_{56}-\frac{1}{2} \cos \alpha\left[l_{1} n_{46}+l_{2} n_{16}+l_{3} n_{26}\right] \tag{C.11}
\end{equation*}
$$

Notice our choice is such that the Euler angles only appear in their left invariant oneform combinations, this is vital because the T-duality rules are local when acting on these one-forms but non-local when applied to the individual coordinates $(\theta, \phi, \psi)$.

We now proceed to the T-dualisation of the Lagrangian eq. (C.3) with respect to the $\mathrm{SU}(2)_{L}$ symmetry for which the $l_{i}$ are invariant one-forms. Some useful quantities are

$$
\begin{equation*}
\hat{J}=-\hat{G}^{-1} d \hat{G}, \quad T_{a}^{\hat{G}}=\hat{G}^{-1} T_{a} \hat{G}, \quad G_{a b}=-\operatorname{Tr}\left(T_{a}^{\hat{G}} P^{(2)} T_{b}^{\hat{G}}\right), \quad Q_{a}=\operatorname{Tr}\left(\hat{J} P^{(2)} T_{a}^{\hat{G}}\right) \tag{C.12}
\end{equation*}
$$

In the case at hand $G_{a b}=\frac{\cos ^{2} \alpha}{4} \delta_{a b}$ and $Q_{a}=0$. Then the T-dual theory is given by the Lagrangian

$$
\begin{equation*}
\mathcal{L}_{\text {dual }}=-\operatorname{Tr}\left(\hat{J}_{+} P^{(2)} \hat{J}_{-}\right)+\partial_{+} v_{a}\left(M^{-1}\right)^{a b} \partial_{-} v_{b} \tag{C.13}
\end{equation*}
$$

with

$$
\begin{equation*}
M_{a b}=G_{a b}+\epsilon_{a b c} v_{c} \tag{C.14}
\end{equation*}
$$

With the transformation

$$
\begin{equation*}
v_{1}=\frac{1}{4} r \cos \theta, \quad v_{2}=\frac{1}{4} r \sin \theta \cos \phi, \quad v_{3}=\frac{1}{4} r \sin \theta \sin \phi \tag{C.15}
\end{equation*}
$$

the T-dual metric is given by

$$
\begin{equation*}
d s^{2}=d \alpha^{2}+\sin ^{2} \alpha d \tilde{\phi}^{2}+\frac{1}{4}\left(\frac{d r^{2}}{\cos \alpha^{2}}+\frac{r^{2} \cos ^{2} \alpha}{r^{2}+\cos ^{4} \alpha}\left(d \theta^{2}+\sin ^{2} \theta d \phi^{2}\right)\right) \tag{C.16}
\end{equation*}
$$

This internal metric when augmented by the untouched $A d S_{5}$ space-time and supplemented by the dilaton and NS and RR fluxes provides the solution of [38] preserving $\mathcal{N}=2$ supersymmetry.

The Buscher procedure gives the T-duality rules for world-sheet derivatives which read

$$
\begin{equation*}
l_{+}^{a} \rightarrow-\left(M^{-1}\right)^{b a} \partial_{+} v_{b}, \quad l_{-}^{a} \rightarrow\left(M^{-1}\right)^{a b} \partial_{+} v_{b} . \tag{C.17}
\end{equation*}
$$

Upon making the above substitution into eqs. (C.10) and (C.11) the Lax connection of (C.4) becomes T-dualised to a Lax connection encoding the dynamics of the T-dual theory with Lagrangian eq. (C.13).

Since in the bosonic theory the $A d S_{5}$ is coupled to the $S^{5}$ only via the Virasoro constraints (which are preserved in the T-dualisation), this guarantees the classical integrability (in the bosonic sector) of the $\mathrm{SU}(2)_{L}$ non-Abelian T-dual of $A d S_{5} \times S^{5}$. The full $\mathfrak{p s u}(2,2 \mid 4)$ coset is treated in [45].

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[^0]:    ${ }^{1}$ For a not so recent review see ref. [2].

[^1]:    ${ }^{2}$ To be precise the $\eta$-deformed backgrounds satisfy conditions of global scale invariance that are weaker than full local conformal invariance and hence obey field equations of a modification to type II supergravity [16].

[^2]:    ${ }^{3}$ Despite this example, it should be emphasised that failure to detect non-integrability in a particular truncation does not in general imply integrability is present.
    ${ }^{4}$ We denote the spatial worldsheet coordinate as $\tilde{\sigma}$ to avoid conflict later. Light cone coordinates are $\sigma^{ \pm}=\tau \pm \tilde{\sigma}$ and we work in units where $\alpha^{\prime}=1$. Hermitian generators of $\mathfrak{s u}(2)$ are $T_{a}=\frac{\tau^{a}}{2}$ such that $l_{\alpha}=-i g^{-1} \partial_{\alpha} g \equiv l_{\alpha}^{a} T_{a}$ with $l_{\alpha}^{a}$ real. We choose $\epsilon^{01}=-1$.

[^3]:    ${ }^{5}$ For the non-Abelian T-dual of the principal chiral model a description of the Lax was provided in [11] and the Lax pair provided by [44, 45] follows, in a rather circuitous route, from one introduced in [46].

[^4]:    ${ }^{6}$ In the following, we denote with a dot the $\tau$-derivative, which should not be confused with the $\sigma \partial_{\sigma}$ derivative defined above.

[^5]:    ${ }^{7}$ Our presentation is at a rather telegraphic and superficial level and we recommend the reader to the books [51-53]. The articles [53-57] explain the application of these techniques to Hamiltonian systems.

[^6]:    ${ }^{8}$ The order of a rational function at infinity is the highest power of the denominator minus the highest power of the numerator.

