Published in "Review of Scientific Instruments 90(2): 023104, 2019" which should be cited to refer to this work.

Time- and angle-resolved photoemission spectroscopy of solids in the extreme ultraviolet at 500 kHz repetition rate

M. Puppin,^{1,2,a)} Y. Deng,^{1,3} C. W. Nicholson,^{1,4} J. Feldl,¹ N. B. M. Schröter,^{1,5} H. Vita,¹ P. S. Kirchmann,⁶ C. Monney,^{1,4} L. Rettig,¹ M. Wolf,¹ and R. Ernstorfer^{1,b)} L.

AFFILIATIONS

- Department of Physical Chemistry, Fritz Haber Institute of the Max Planck Society, Faradayweg 4-6, 14195 Berlin, Germany
- ²Laboratory of Ultrafast Spectroscopy, ISIC, and Lausanne Centre for Ultrafast Science (LACUS), École Polytechnique Fédérale de Lausanne, CH-1015 Lausanne, Switzerland
- ³Paul Scherrer Institute, SwissFEL, 5232 Villigen PSI, Switzerland
- ⁴Département de Physique and Fribourg Center for Nanomaterials, Chemin du Musée 3, Université de Fribourg, CH-1700 Fribourg, Switzerland
- ⁵Paul Scherrer Institute, 5232 Villigen PSI, Switzerland
- ⁶Stanford Institute for Materials and Energy Sciences, SLAC National Accelerator Laboratory, 2575 Sand Hill Road, Menlo Park, California 94025, USA
- ^{a)}Electronic mail: michele.puppin@epfl.ch
- b) Electronic mail: ernstorfer@fhi-berlin.mpg.de

ABSTRACT

Time- and angle-resolved photoemission spectroscopy (trARPES) employing a 500 kHz extreme-ultraviolet light source operating at 21.7 eV probe photon energy is reported. Based on a high-power ytterbium laser, optical parametric chirped pulse amplification, and ultraviolet-driven high-harmonic generation, the light source produces an isolated high-harmonic with 110 meV bandwidth and a flux of more than 10¹¹ photons/s on the sample. Combined with a state-of-the-art ARPES chamber, this table-top experiment allows high-repetition rate pump-probe experiments of electron dynamics in occupied and normally unoccupied (excited) states in the entire Brillouin zone and with a temporal system response function below 40 fs.

Published under license by AIP Publishing. https://doi.org/10.1063/1.5081938

I. INTRODUCTION

The temporal evolution of the microscopic properties of a solid brought out of equilibrium by an ultrashort laser pulse provides fundamental insights into the couplings between its electronic, spin, and lattice degrees of freedom. Timeresolved spectroscopy allows disentangling experimentally the interplay and coupled evolution of these subsystems, whose characteristic time scales are set by their microscopic interactions. Ultrashort light pulses can resolve the fast electronic evolution occurring on femtosecond time scales. Timeand angle-resolved photoemission (trARPES) directly accesses electronic states of a material with momentum resolution, as the system is driven out of equilibrium by a femtosecond

optical pump pulse. ARPES measures the angular distribution and the kinetic energy of photoemitted electrons; it is frequently assumed that the ARPES intensity $I(\mathbf{k},\omega)$ can be written as the product between the occupation probability of the electronic state $f(\mathbf{k},\omega)$, the single-particle spectral function $A(\mathbf{k},\omega)$, and a matrix element between the initial and final state $|M_{ij}^{\mathbf{k}}|^2$; here \mathbf{k} and ω denote the electron's wavevector and angular frequency, respectively. Many-body effects are encoded in the spectral function $A(\mathbf{k},\omega)$ and manifest themselves in renormalization of the bare electronic bands and in the observed lineshape. In a trarress experiment, the distribution $I(\mathbf{k},\omega)$ is collected for a series of delays (τ) between pump and probe pulses; after perturbation, the population distribution $f(\mathbf{k},\omega,\tau)$

evolves toward a quasi-thermal distribution and energetically relaxes on femto- to picosecond time scales.² During relaxation, the concomitant many-body interactions affect the transient spectral function $A(\mathbf{k},\omega,\tau)$ and even the photoemission matrix elements might change if the final state's orbital symmetry is altered.³ trARPES accesses at once the population dynamics, the evolution of the spectral function, and the evolution of matrix elements. trARPES has found increasingly successful applications in the past few decades;4-6 among many examples, trARPES was used to study photo-induced phase transitions⁷⁻¹¹ and to observe electronic states above the Fermi level, unoccupied under equilibrium conditions. 12-16 Energy conservation in the photoemission processes imposes that a femtosecond light source for trARPES must possess a photon energy $\hbar\omega_{ph}$ exceeding the work function Φ , which in most materials lies in the range between 4 and 6 eV. Ultraviolet femtosecond light sources are thus required for these experiments. The conservation of the electrons' in-plane momentum $(\hbar k_{\parallel})$ in the photoemission process allows reciprocal space resolution. The advantage of a probe with high photon energy is the increased range of observable reciprocal space; lowphoton-energy sources are limited to regions close to the Brillouin zone center.

The scope of this work is to describe an experimental setup for trARPES based on a newly developed light source which operates at an energy of 21.7 eV and provides a monochromatized photon flux exceeding 10¹¹ ph/s on the sample. The light source is embedded in a beamline equipped with a state-of-the-art ARPES end-station, where the trARPES experiment can be performed with a system response function (pump-probe cross-correlation) better than 40 fs, a source linewidth of 110 meV, and at a repetition rate of 500 kHz. This is achieved by performing a change of the employed laser technology, from a conventional titanium:sapphire laser to an optical parametric chirped pulse amplifier (OPCPA), entirely based on sub-picosecond ytterbium lasers.¹⁷ The experimental apparatus presented in this work bridges the existing technology gap between widespread high-flux, high repetition rate sources with low photon energy^{15,18-22} and conventional high photon energy sources, based on high-order harmonic generation and operating at lower repetition rates,23-31 thereby enabling a vast class of new experiments to be performed in the whole Brillouin zone of most materials. The structure of the paper is as follows: Sec. II will be dedicated to the description of the light source, whereas the trARPES beamline will be described and characterized in Sec. III.

II. THE LIGHT SOURCE

For each trARPES experiment, a multi-dimensional data set is recorded; the photoelectron intensity distribution is measured as a function of energy, parallel momentum, and pump-probe delay time. In order to collect sufficient statistics, data have to be accumulated over numerous laser pulses since the probe pulse intensity cannot be increased arbitrarily. In fact, multiple photoelectrons per pulse lead to space-charge effects;³²⁻³⁴ the photoemitted electrons are initially confined in a small volume leading to strong Coulomb

repulsion and a spreading of the photoelectron cloud, worsening the energy and momentum resolution. The best way to mitigate this effect is to reduce the photon flux and accordingly increase the experimental repetition rate to compensate for the reduced count rate.

However, in pump-probe experiments, the pump excitation also has to be taken into consideration when designing the experiment. First, it is important that the sample reequilibrates within the laser's duty cycle; second, in many experiments, a minimum energy density has to be reached to initiate a certain non-equilibrium process, for example, to photo-induce a phase transition. 7,9,35 In many cases, the photo-degradation of the sample occurs faster at higher repetition rates, making a vast variety of studies unfeasible. The ideal repetition rate depends therefore on the sample under investigation; however, the need for high photon energies complicates trARPES when going beyond a few tens of kHz. This stems from the fact that femtosecond lasers, typically in the visible to near-infrared (VIS-NIR) spectral ranges, have to undergo a high-order frequency up-conversion process, requiring laser amplifiers with high peak powers. Titanium:sapphire laser amplifiers are currently the workhorse of femtosecond science, providing terawatt pulses at the kHz level.³⁶ Intensities on the order of 10¹⁴ W/cm² are easily reached, routinely enabling the generation of extreme ultraviolet radiation (XUV, 20-100 eV) via high harmonic generation

Unfortunately, it is difficult to scale XUV generation with Ti:sapphire lasers above a repetition rate of a few tens of kHz; some of the highest repetition rates demonstrated so far for trARPES experiments are 50-100 kHz, 14,38 whereas static ARPES experiments have been demonstrated at the MHz level.³⁹ Several approaches have been devised to increase the repetition rate of table-top light sources, often employing different type of femtosecond lasers: for example, cavityenhanced XUV generation provides repetition rates of several tens of MHz and space-charge-free photoemission; 40-42 in this regime also the pump excitation has to be modest to allow for sample's relaxation. In the present work, we aim at increasing the repetition rate of single-pass harmonic generation by employing an OPCPA based on ytterbium lasers, with a sufficient average power to efficiently drive the process at 500 kHZ.

A. Optical parametric chirped pulse amplification

In the past decades, several ytterbium lasers operating at 100s of watts of average power and MHz-level repetition rates were demonstrated:⁴³ such lasers are already capable of directly producing XUV radiation via high-harmonic generation⁴⁴ or can be brought to the few-cycle pulse regime by nonlinear compression.⁴⁵ Short pulse duration and high mode quality make amplified ytterbium lasers ideally suited for OPCPA.⁴⁶ In OPCPAs, by controlling the spectral phase of the amplified optical pulses, the central frequency and the bandwidth can be tuned at will. In the VIS-NIR range, several ytterbium-based OPCPAs were demonstrated either with a broad bandwidth spectrum, supporting few cycle pulses⁴⁷⁻⁵¹ or with a frequency-tunable spectrum over a wide

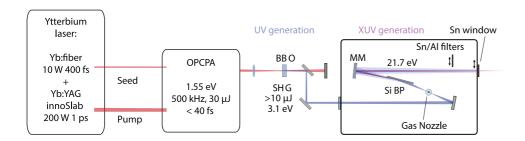


FIG. 1. Schematic layout of the 21.7 eV light source. OPCPA: optical parametric chirped pulse amplifier, SHG: second harmonic generation, BP: Brewster plate, and MM: multilayer mirror.

range.^{17,52-55} This combination of high average power and time-bandwidth flexibility makes OPCPAs promising candidates for the next generation of femtosecond lasers⁵⁶ and high-repetition rate sources.⁵⁷

The approach followed in this work comprises an OPCPA based on an Yb:YAG laser which produces femtosecond NIR light at 1.55 eV (Fig. 1) at 500 kHz. The details of the laser system has been described elsewhere 17 and will only be summarized here. The OPCPA is seeded by a broad-bandwidth white light continuum, which is generated in a YAG crystal 58 by a fiber laser with a pulse duration of 400 fs (full-width at half maximum, FWHM) at a carrier wavelength of 1030 nm. The pump for the parametric amplification is the second harmonic (515 nm) of a slab amplifier 43 capable of 200 W average power and with a picosecond pulse duration. This pump laser is seeded by the same fiber laser producing the broadband seed for the OPCPA; this ensures an inherent optical synchronization between the seed and the pump pulses, important for parametric amplification of sub-ps pulses. 59

A single amplification stage in a beta-barium borate (β -BBO) crystal is sufficient to saturate the pump conversion efficiency, simplifying the optical setup. The central frequency of the amplifier, tunable within 650 nm–950 nm, was set to 800 nm for practical reasons to access the widespread optical components available for Ti:sapphire lasers. After a prism compressor, 30 μ J pulses at 500 kHz are available for further conversion to the XUV range. For typical trARPES experiments, the spectrum of the laser is set to a bandwidth of 80 meV (FWHM), around the photon energy of 1.55 eV and compressed to a pulse duration below 35 fs (FWHM) with a mode quality M² < 1.5.

B. UV-driven high-harmonic generation at 500 kHz

In the process of high-harmonic generation in gases, the XUV radiation originates from electrons recombining with their parent ion after being accelerated in the optical field. The hall-mark of this effect in the frequency domain is the appearance of odd harmonics of the fundamental laser frequency ω_0 , co-propagating with the driving radiation. The irradiance for several harmonic orders is nearly constant in a wide energy region, the so-called plateau region, which extends up to a cut-off energy determined by the atomic ionization potential and the electron's ponderomotive potential. The temporal structure of the harmonics, consisting of bursts of radiation at every half-cycle of the driving electric

field, has proven to be extremely valuable for femtosecond⁴⁻⁶ and attosecond spectroscopies.⁶²

Although it is possible to directly generate high harmonics with the NIR pulses of the OPCPA, there are advantages in frequency up-converting the pulses in a nonlinear crystal and generating high harmonics with UV pulses. 28,30,63 The conversion losses introduced are largely compensated by an increase in the single atom response⁶⁴ and an improvement of macroscopic phase-matching conditions. 63 At a given XUV energy, the doubled driver's photon energy reduces the order q of the nonlinear processes. This leads to a lower Gouy phase for the focused laser beam and a lower phase mismatch resulting from the free-electron plasma; an overall efficiency increase in nearly two orders of magnitude was reported using Ti:sapphire lasers.63 On top of this increased efficiency, UV -driven HHG simplifies selection of a single harmonic out of the XUV frequency comb, as will be explained in more detail in Sec. II D.

C. High-pressure gas target for HHG in a tight-focusing geometry

A high photon flux for trARPES can only be achieved if the harmonic radiation produced across a macroscopic volume of the gas target adds coherently.37 This phase-matching condition is more difficult to achieve in a tight-focusing regime;65 this is a direct consequence of the Gouy phase of the laser beam, which becomes an increasingly important term as the focal spot gets smaller. A detailed analysis of the scaling properties of the phase matching relations for HHG65-67 reveals that the theoretical conversion efficiency can be made invariant of the spot size w_0 , provided that the pulse energy ϵ , the gas pressure p, and the gas target length L are suitably rescaled. More precisely,66 the efficiency reached in a loose focusing geometry at certain intensity can be also expected for a lower pulse energy $\epsilon' = \epsilon s$ and a smaller spot size $\sqrt{s}w_0$, where s < 1 is the scaling factor. For this to happen, a higher pressure $p' = ps^{-1}$ has to be achieved in a gas target confined in a shorter length L' = Ls; the main challenge of a tight-focused HHG light source is therefore to realize a high-pressure, confined gas target within a high-vacuum beamline.68

In our setup this is realized by focusing the beam at the output of a small diameter converging nozzle, ⁶⁹ connected to a gas line with a pressure of several bars; the gas nozzle configuration is schematized in Fig. 2(a). The gas jet freely expands in a high-vacuum chamber; to minimize absorption losses, the

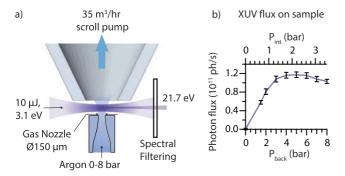


FIG. 2. (a) Side view of the differential pumping system for the high-pressure gas nozzle. (b) Measured flux at 21.7 eV on the sample as a function of the backing pressure P_{back} , the estimated pressure at the focus of the laser beam P_{int} is shown in the mirror axis.

lowest residual pressure is desirable along the beam path of the strongly ionizing XUV radiation. To reduce the overall gas load, a skimmer with an aperture of 3 mm diameter is placed in front of the nozzle. A three-axis manipulator holds the skimmer and is used for precise alignment. A scroll pump with a pumping speed of $35~\text{m}^3/\text{h}$ is connected to the rear-side of the skimmer through a flexible hose. For a typical nozzle diameter of $150~\mu\text{m}$, the differential pumping setup ensures a long-term operation of the 700 l/min turbomolecular pump of the vacuum chamber, with up to 9 bars backing pressure of argon; the chamber's pressure during operation is in the 10^{-2} mbar range. The beam path following the gas target consists of several optical components to re-collimate the beam, suppress the 3.1 eV UV driver, and isolate a single harmonic around 22 eV.

D. Single harmonic selection

The separation between neighboring harmonics is $2\omega_0$; in UV-driven HHG, the spacing between different orders also increases, facilitating the monochromatization of the XUV radiation. For a trARPES experiment, a single harmonic must be isolated with good spectral contrast; the contamination of neighboring harmonics produces unwanted replicas of the photoelectron spectrum. In particular, if the qth harmonic is selected, a residual (q+2) order produces photoelectrons from deeper valence states which overlap in energy with states in the vicinity of the Fermi level.

Monochromators based on a single diffraction grating, widely adopted in XUV synchrotron beamlines, have the drawback of introducing a pulse front tilt which hinders the temporal resolution in pump-probe experiments. Special designs must be adopted to minimize this effect without sacrificing the transmission considerably.^{27,70} The alternative approach is to use a combination of reflective and transmissive optics to isolate a single harmonic from the fundamental beam.

In this kind of monochromator, a harmonic is selected by multilayer-coated mirrors, designed for the specific wavelength; this preserves the pulse duration, sacrificing however the tunability. Multilayer mirrors have a typical reflectivity of few tens percent at the wavelength for which they are designed; unfortunately, away from the reflectivity peak, a non-negligible residual reflectivity decreases the contrast between neighboring harmonics. In the case of NIR-driven HHG, the contrast is improved by a second, lossy reflection.²⁴ A transmissive filter is required to fully suppress the co-propagating fundamental radiation, which is several orders of magnitude brighter than the harmonics. A free-standing metal foil with a thickness of some hundreds of nanometers effectively suppresses the fundamental radiation and still transmits a reasonable proportion of the XUV radiation. Aluminum is commonly used as it acts as a high-pass filter above 20 eV; the estimated⁷¹ transmission of 200 nm Al is shown in Fig. 3(a).

The spacing between UV-driven harmonics (6.2 eV in our case) allows for a simplified setup based on a single reflection on a multilayer mirror, followed by a transmission through a free-standing tin foil.⁶³ Tin has a transmission window centered approximately at 22 eV close to the 7th harmonic of the 3.1 eV driver; the theoretical transmission through 200 nm of Sn is shown in Fig. 3(a) together with the position of the closest harmonics (vertical dashed lines). Experimentally it was determined that a Sn foil (Lebow) of 200 nm nominal thickness had a transmission of 9% at a photon energy of 21.7 eV. The reduced transmission compared to the theoretical value is likely due to oxide layers forming at the surfaces.

The optical setup following the gas target is shown schematically in Fig. 1. High-repetition rate systems have an inherently higher thermal load which could easily damage the

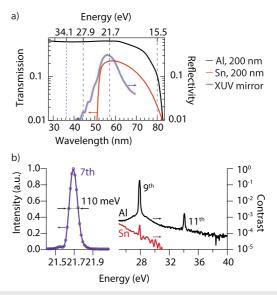


FIG. 3. (a) Left vertical axis: Theoretical transmission 71 of 200-nm-thick Sn (red line) and Al (black line) foils. Right vertical axis: theoretical reflectivity of the multi-layer mirror. The vertical dashed lines indicate the position of the harmonics, from order 5 to 11. (b) Left vertical axis: Normalized spectrum of the 7th harmonic at 21.7 eV, highlighting the FWHM of 110 meV. Right vertical axis: intensity of the 9th and 11th harmonics, relative to the main 21.7 eV line. The contrast of Al (black line) and Sn (red line) is compared.

thin metal filter. To solve this problem, a Brewster plate consisting of a silicon single crystal was used as the first optical element in the beamline, placed approximately 100 mm after the focus. Most of the fundamental power is absorbed in the silicon wafer, preventing damage in the following components. The thermal load did not cause a sensible mode degradation of the XUV beam; the characterization of the XUV beam profile is discussed in Sec. III. The reflectivity of the Si plate in the XUV was estimated numerically using the free software IMD, 72 including the effects of a passivating SiO₂ layer of about 1 nm, as expected for a commercial high-purity silicon single crystal. 73 At the Brewster angle for 400 nm (80° angle of incidence), the measured reflected power for the driving radiation is below 100 mW, while the calculated reflectivity for the 7th harmonic is 80%.

After the silicon wafer, the XUV radiation is re-collimated by a spherical mirror with a focal length of 200 mm, coated with a multilayer with reflectivity centered at the 7th harmonic. The theoretical reflectivity of the XUV mirror is on the order of 30% and is plotted in Fig. 3(a). The mirror's coating consists of a silicon layer, covering a multilayer composed of chromium and scandium, realized on an XUV-grade substrate (flatness $\lambda/20$, roughness <0.2 nm RMS). The metallic Sn filter is mounted as a window of a gate valve, and a motorized filter wheel can be used to insert additional filters (Al or Sn, 200 nm thickness).

A gold mirror can be inserted in the beam to reflect the beam into a grating spectrometer⁷⁵ to measure the XUV spectrum. A typical spectrum is shown in Fig. 3(b). The lower limit for the instrumental resolution under the measurement conditions is 60 meV; the measured line width of the 7th harmonic is 110 meV. The contrast between the 7th harmonic at 21.7 eV and the 9th harmonic at 27.9 eV was measured with a single 200 nm-thick Al filter, with a flat transmission above 22 eV and compared with the one of a single 200 nm-thick Sn foil. The contrast between the 7th harmonic and the 9th improves by more than two orders of magnitude in the latter case; an additional 200 nm-thick Sn foil reduces the signal at 28 eV to the noise level of the spectrometer's detector, corresponding to a flux on the sample below 10⁷ ph/s.

To characterize the 7th harmonic radiant power, an XUV diode⁷⁶ mounted on a linear transfer arm can be inserted in the beam. It is important to note that no additional optical elements are present in the beamline after the XUV diode (see Sec. III); the radiant flux measured in this position corresponds therefore to the one at the sample's position during trARPES experiments. The measured photon flux at 21.7 eV, calculated using the diode's factory responsivity of 0.15 A/W, is plotted as a function of the backing pressure P_{back} of the nozzle in Fig. 2(b). Optimal phase matching for a 150 μ m-diameter nozzle was observed at a backing pressure of approximately 5 bars for argon. The gas pressure in the interaction region P_{back} is shown in the mirror axis of Fig. 2; these results assume an ideal supersonic expansion of the gas,⁷⁷ and the minimal distance of the beam axis from the nozzle front face is assumed to be twice the beam radius w(z), with a beam waist of $2w_0 = 25 \,\mu\text{m}$. The overall transmission of the transmissive monochromator with a single Sn foil is estimated to be 2.2%; this allows us to

TABLE I. Relative maximum 21.7 eV flux for different nozzle throat diameters together with the measured flux on the sample. The source flux is calculated from the estimated monochromator transmission of 2.2%.

| Nozzle throat (μ m) | 40 | 80 | 150 | 500 |
|---|------|------|------|------|
| Relative flux | 1.00 | 0.74 | 0.60 | 0.13 |
| Source flux (×10 ¹² ph/s) | 9.2 | 6.8 | 5.6 | 1.2 |
| Flux on the sample ($\times 10^{11}$ ph/s) | 2.0 | 1.5 | 1.2 | 0.3 |

calculate the photon flux from the source before monochromatization, which is on the order of 5.6×10^{12} photons/s for the 7th harmonic.

E. Comparison between different nozzles and noble gases

The optimal gas medium length was studied by testing nozzles of different throat diameters (40, 80, 150, and 500 μ m) under the same focusing conditions in argon. The position of each nozzle was optimized to maximize the radiant flux with a pressure up to 9 bars, which was the maximal pressure for the gas fittings in the gas line. The best results for each configuration are reported in Table I. A saturation of the XUV flux with pressure was observed only for the 150 μ m -diameter nozzle, whereas in the case of the 500 μ m-diameter nozzle, the turbomolecular pump overloaded before HHG saturation. For the 40 μ m-diameter nozzle, the signal was still monotonically increasing at the maximum backing pressure. The highest flux was measured with the smallest nozzle, indicating that for a medium length of $\approx 100 \mu m$ the harmonic re-absorption already limits the signal build-up. In practice and in view of performing long measurements, a 150 μ mdiameter nozzle was preferred as HHG was found to be less sensitive to beam-pointing drifts. With this nozzle size and a typical backing pressure of 2 bars, the argon gas flow is below 300 sccm/min.

Other noble gases were also tested for the 150 μ m nozzle; in each case, pressure and nozzle position were optimized to maximize the XUV flux. The relative intensities are listed in Table II. The best results were obtained with argon, even though krypton is expected to yield a stronger single-atom response. ^{14,63} It is possible that the higher ionization in the case of Kr prevented phase matching at the pulse peak, resulting in a shorter coherence length. A longer focal length could not be tested in the setup due to geometrical constrains; nonetheless, it is expected that a higher XUV radiant power would be possible, using the present driver laser and krypton as gas medium. Due to the scarcity and higher cost of this

TABLE II. Comparison of the relative radiant flux for the 21.7 eV harmonic with different gases for 150 μ m nozzle diameter and measured on the sample's position. The source's flux is calculated from the estimated monochromator transmission of 2.2%.

| Noble gas | Ne | Ar | Kr |
|--|------|-----|------|
| Relative flux Source flux (×10 ¹² ph/s) Flux on the sample (×10 ¹¹ ph/s) | 0.02 | 1.0 | 0.78 |
| | 0.1 | 5.6 | 4.3 |
| | 0.02 | 1.2 | 0.9 |

gas, a gas recycling system^{14,78} is needed for a high pressure nozzle, complicating its adoption in the current setup. Overall, the source's flux using Ar exceeds 10¹¹ ph/s at 21.7 eV and at 500 kHz, which is well suited for trARPES experiments.³⁵

III. THE trARPES BEAMLINE

An XUV beamline connects the light source to the experimental chamber; the setup is depicted in Fig. 4. The XUV beam, generated in the high-pressure gas target in the HHG chamber, is reflected by the silicon Brewster plate and relayed by the XUV mirror toward a second chamber. In this chamber, a manual 3-axis manipulator holds a motorized iris used in the experiments to attenuate the XUV flux without changing the phase-matching conditions. This chamber also hosts a motorized filter wheel with 6 slots, containing Sn and Al filters, which can be inserted in the beam to improve the contrast between neighboring harmonics. After the filter wheel, a gate

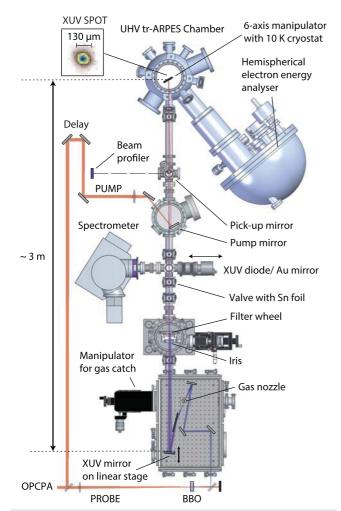


FIG. 4. Scaled layout of the beamline. Refer to Sec. III in the main text for a detailed description. Inset: Measured XUV profile.

valve hosts the Sn foil, used both as a transmissive filter for selecting the 7th harmonics and as a pressure barrier before the ultra-high-vacuum (UHV) experimental chamber.

The next section of the beamline contains tools for HHG characterization; after the Sn foil, a linear translation stage hosts the XUV diode, which measures the flux incident on the sample. The same linear translation arm holds a gold mirror, used to reflect the harmonics into the XUV spectrometer. A third vacuum chamber along the beamline hosts a fixed mirror mount, used to in-couple the pump beam, at an angle $\approx\!3^\circ$ relative to the XUV beam.

The pump and the probe beams finally arrive to the UHV trARPES chamber where they overlap on the sample. The trARPES chamber reaches a base pressure in the upper 10^{-11} mbar range; thanks to the Sn window no significant increase is observed during measurements when the high-pressure gas nozzle is operating. The XUV focal spot was characterized using a micro-channel-plate electron multiplier, followed by a phosphor screen imaged by a camera. The resulting beam profile is shown in the inset of Fig. 4 and measures $130~\mu m$ full-width at half maximum.

The UHV chamber is equipped with a hemispherical electron energy analyzer.⁷⁹ The sample is installed on a 6-axis manipulator⁸⁰ which allows liquid helium cooling down to 10 K. The XUV beamline is connected to the analysis chamber in the same plane of the analyzer's entrance slit; the analyzer's input axis is fixed at an angle of 40° relative to the beamline. The manipulator, holding the sample during the experiments, can be moved to an upper chamber, where samples are stored and prepared before the measurements. The samples are inserted in vacuum through a load-lock chamber and are moved to the sample storage chamber using a magnetic transfer arm.

Just before the trARPES chamber, a cross joint holds a linear translation stage with a metallic pick-up mirror; this can be inserted in the beam path to simultaneously reflect the pump and the residual 400 nm HHG driver, which is still observable if the Sn foil is removed. This allows for characterization of the pump beam profile and coarse temporal overlap of pump and probe beams on a photodiode. The XUV beam is aligned in the photoemission chamber by controlling a motorized mirror mount holding the multilayer mirror. The mount is translated along the beam axis with a linear translation stage to precisely focus the source on the sample. Once the XUV beam is aligned relative to the electron energy analyzer focus, the pump beam can be overlapped on the probe beam by imaging with a CCD camera a Ce:YAG scintillator, installed on the manipulator.

IV. CHARACTERIZATION OF THE trARPES SETUP

First trARPES experiments were demonstrated on WSe₂, a member of the well-studied transition-metal dichalcogenide semiconductors family; $^{12-14,81}$ the results are summarized in Fig. 5. The photon energy of the light source is sufficient to reach electrons with a parallel momentum of 1.7 Å⁻¹ (with a photoemission angle of 60°). While data were recorded over extended regions in momentum space, here we focus on the trARPES signal in the vicinity of the high symmetry point \bar{K} ,

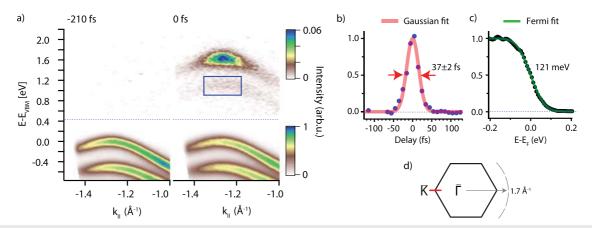


FIG. 5. (a) False color plots of trARPES data 210 fs before and during temporal overlap in WSe₂ at the \overline{K} point. The photoelectron energy E is referenced to the top of the valence band at the \overline{K} point, E_{VBM} . Above 0.45 eV, marked by a black line, a different color scale was used in the image to allow visualizing the weaker excited states signal, two color bars indicate the signal level, and the same arbitrary units are used for both the images. The blue box indicates the integration area used to estimate the pump-probe cross-correlation. (b) Integrated intensity of the photoemission signal in the blue box of panel (a), showing the result of a fit using a Gaussian function, indicating a temporal cross-correlation between the pump and probe signal of 37 \pm 2 fs FWHM. (c) Energy distribution curve at the Fermi edge of TbTe₃, at a temperature of 100 K. A Fermi function at a temperature of 100 K was convoluted with a Gaussian and fitted to the data, yielding a broadening of 121 meV. (d) Surface-projected Brillouin zone of WSe₂. The red line indicates the position in reciprocal space of the ARPES plots of panel (a). The gray line indicates the accessible momentum space.

a corner of the hexagonal surface-projected Brillouin zone [Fig. 5(d)]. For the experiment, a split portion of the fundamental output of the OPCPA at 1.55 eV was used as a pump, with a peak fluence on the sample of 1.1 mJ/cm². The trARPES spectrum is shown in Fig. 5(a), before (left) and during (right) pump-probe temporal overlap. The energy zero is set for convenience to the local maximum of the valence band at K, an inherently two-dimensional state. 12 During temporal overlap between the XUV pulse and the 1.55 eV s-polarized pump, non-resonant two-photon photoemission from the spin-orbit split valence band produces a short-lived signal in the material's bandgap, separated by 1.55 eV from the initial valence band states.¹³ We use this non-resonant signal to estimate the temporal cross-correlation between the pump and probe; the intensity in a small region at 1.1 eV and -1.27 Å^{-1} [blue box in Fig. 5(a)] is shown in Fig. 5(b). The maximum of this signal is used as a zero for the temporal axis, and the trace is normalized. This signal is fitted with a Gaussian function; the curve's FWHM is $3\overline{7} \pm 2$ fs. This value can be compared with the pump pulse duration of 32 fs, which was characterized by frequency-resolved optical grating; the system response function is apparently dominated by the pump pulse duration. Assuming both pump and probe pulses to have a Gaussian envelope, with a time-bandwidth product of 1824 meV fs, one obtains a probe pulse width of 19 fs FWHM, not far from the Fourier-limited pulse duration of 17 fs extracted from the spectral width of 110 meV. This suggests that an even better temporal resolution, without losses in energy resolution, would be possible by further shortening of the pump pulses. The dataset, acquired with an integration time of 31 s per delay point, clearly demonstrates the ability of the setup to follow population dynamics of excited states, unoccupied at equilibrium, in the entire Brillouin zone. The detectable in-gap two-photon photoemission signal has an intensity of $\approx 10^{-4}$ relative to the simultaneously measured VB; this high dynamic range can be used to follow in a single data acquisition the evolution of states below and above the Fermi level. The excited states in this proof-of-principle experiment show rich momentum dependent relaxation dynamics which will be discussed in detail in a future publication.

The experimental energy resolution was checked on a sample with bands crossing the Fermi level, TbTe $_3$; a momentum-integrated energy distribution curve (EDC) is shown in Fig. 5(c); the data set was collected at a temperature of 100 K, with a photocurrent of 130 pA, corresponding to 1.6×10^3 electrons emitted per pulse. The EDC was fitted by a Fermi-Dirac distribution function with the temperature of 100 K as a fixed parameter, convolved by a Gaussian broadening as a fit parameter; the total resolution is 121 meV. This value includes the dominant 110 meV source's linewidth, broadened by the analyzer resolution of 35 meV and a residual space-charge contribution, on the order of 35-40 meV.

V. CONCLUSIONS

In conclusion, we have demonstrated a novel setup for high-repetition-rate trARPES at XUV energies capable of band mapping in the whole Brillouin zone of most materials. The light source developed for the experiment is based on a Ytterbium-based OPCPA operating at 500 kHz, frequency upconverted to 21.7 eV by a UV-driven HHG source. Time-resolved two-photon photoemission studies are feasible in a broad reciprocal space region with a system response function below 40 fs. The novel setup can map the excited-state band structure and follow its evolution on a femtosecond time scale. 35

ACKNOWLEDGMENTS

We thank S. Kubala, M. Krenz, D. Bauer, R. Franke, P. Heyne, W. Erlebach., J. Malter, and C. Cacho for technical support. This work was funded by the Max Planck Society, the European Research Council (ERC) under the European Union's Horizon 2020 research and innovation program (Grant No. ERC-2015-CoG-682843), the European Union's FP7 program through the CRONOS project, Grant No. 280879, and the German Research Foundation (DFG) within the Emmy Noether program (Grant No. RE 3977/1). C.M. acknowledges support by the Swiss National Science Foundation under Grant No. PZ00P2-154867.

REFERENCES

- ¹A. Damascelli, Phys. Scr. **T109**, 61 (2004).
- ²M. Lisowski, P. Loukakos, U. Bovensiepen, J. Stähler, C. Gahl, and M. Wolf, Appl. Phys. A: Mater. Sci. Process. **78**, 165 (2004).
- Boschini, D. Bugini, M. Zonno, M. Michiardi, R. P. Day, E. Razzoli,
 B. Zwartsenberg, E. H. d. S. Neto, S. dal Conte, S. K. Kushwaha, R. J. Cava,
 S. Zhdanovich, A. K. Mills, G. Levy, E. Carpene, C. Dallera, C. Giannetti,
 D. J. Jones, G. Cerullo, and A. Damascelli, e-print arXiv:1810.06571v1 (2018).
 R. Haight, Surf. Sci. Rep. 21, 275 (1995).
- ⁵U. Bovensiepen and P. S. Kirchmann, Laser Photonics Rev. **6**, 589 (2012).
- $^6\mathrm{C}.$ L. Smallwood, R. A. Kaindl, and A. Lanzara, Europhys. Lett. 115, 27001 (2016).
- ⁷F. Schmitt, P. S. Kirchmann, U. Bovensiepen, R. G. Moore, L. Rettig, M. Krenz, J.-H. Chu, N. Ru, L. Perfetti, D. H. Lu, M. Wolf, I. R. Fisher, and Z.-X. Shen, Science **321**, 1649 (2008).
- ⁸T. Rohwer, S. Hellmann, M. Wiesenmayer, C. Sohrt, A. Stange, B. Slomski, A. Carr, Y. Liu, L. M. Avila, M. Kalläne, S. Mathias, L. Kipp, K. Rossnagel, and M. Bauer, Nature **471**, 490 (2011).
- ⁹C. Monney, M. Puppin, C. W. Nicholson, M. Hoesch, R. T. Chapman, E. Springate, H. Berger, A. Magrez, C. Cacho, R. Ernstorfer, and M. Wolf, Phys. Rev. B **94**, 165165 (2016).
- ¹⁰ A. Zong, A. Kogar, Y.-Q. Bie, T. Rohwer, C. Lee, E. Baldini, E. Ergeçen, M. B. Yilmaz, B. Freelon, E. J. Sie, H. Zhou, J. Straquadine, P. Walmsley, P. E. Dolgirev, A. V. Rozhkov, I. R. Fisher, P. Jarillo-Herrero, B. V. Fine, and N. Gedik, Nat. Phys. 15, 27 (2018).
- ¹¹C. W. Nicholson, C. Monney, R. Carley, B. Frietsch, J. Bowlan, M. Weinelt, and M. Wolf, Phys. Rev. Lett. 117, 136801 (2016).
- ¹²R. Bertoni, C. W. Nicholson, L. Waldecker, H. Hübener, C. Monney, U. De Giovannini, M. Puppin, M. Hoesch, E. Springate, R. T. Chapman, C. Cacho, M. Wolf, A. Rubio, and R. Ernstorfer, Phys. Rev. Lett. 117, 277201 (2016).
- ¹³P. Hein, A. Stange, K. Hanff, L. X. Yang, G. Rohde, K. Rossnagel, and M. Bauer, Phys. Rev. B **94**, 205406 (2016).
- ¹⁴R. Wallauer, J. Reimann, N. Armbrust, J. Güdde, and U. Höfer, Appl. Phys. Lett. **109**, 162102 (2016).
- ¹⁵J. A. Sobota, S. Yang, J. G. Analytis, Y. L. Chen, I. R. Fisher, P. S. Kirchmann, and Z.-X. Shen, Phys. Rev. Lett. **108**, 117403 (2012).
- ¹⁶F. Mahmood, C.-K. Chan, Z. Alpichshev, D. Gardner, Y. Lee, P. A. Lee, and N. Gedik, Nat. Phys. **12**, 306 (2016).
- ¹⁷M. Puppin, Y. Deng, O. Prochnow, J. Ahrens, T. Binhammer, U. Morgner, M. Krenz, M. Wolf, and R. Ernstorfer, Opt. Express 23, 1491 (2015).
- ¹⁸C. L. Smallwood, C. Jozwiak, W. Zhang, and A. Lanzara, Rev. Sci. Instrum. 83, 123904 (2012).
- 19Y. Ishida, T. Togashi, K. Yamamoto, M. Tanaka, T. Kiss, T. Otsu, Y. Kobayashi, and S. Shin, Rev. Sci. Instrum. 85, 123904 (2014).
- ²⁰J. D. Koralek, J. F. Douglas, N. C. Plumb, J. D. Griffith, S. T. Cundiff, H. C. Kapteyn, M. M. Murnane, and D. S. Dessau, Rev. Sci. Instrum. 78, 053905 (2007).
- ²¹F. Boschini, H. Hedayat, C. Dallera, P. Farinello, C. Manzoni, A. Magrez, H. Berger, G. Cerullo, and E. Carpene, Rev. Sci. Instrum. 85, 123903 (2014).

- ²²J. Faure, J. Mauchain, E. Papalazarou, W. Yan, J. Pinon, M. Marsi, and L. Perfetti, Rev. Sci. Instrum. 83, 043109 (2012).
- ²³R. Haight and D. Peale, Rev. Sci. Instrum. **65**, 1853 (1994).
- ²⁴S. Mathias, L. Miaja-Avila, M. Murnane, H. Kapteyn, M. Aeschlimann, and M. Bauer, Rev. Sci. Instrum. 78, 083105 (2007).
- ²⁵G. Dakovski, Y. Li, T. Durakiewicz, and G. Rodriguez, Rev. Sci. Instrum. 81, 073108 (2010).
- ²⁶I. C. E. Turcu, E. Springate, C. A. Froud, C. M. Cacho, J. L. Collier, W. A. Bryan, G. R. A. J. Nemeth, J. P. Marangos, J. W. G. Tisch, R. Torres, T. Siegel, L. Brugnera, J. G. Underwood, I. Procin, W. R. Newell, C. Altucci, R. Velotta, R. B. King, J. D. Alexander, C. R. Calvert, O. Kelly, J. B. Greenwood, I. D. Williams, A. Cavalleri, J. C. Petersen, N. Dean, S. S. Dhesi, L. Poletto, P. Villoresi, F. Frassetto, S. Bonora, and M. D. Roper, Proc. SPIE 7469, 746902 (2010).
- ²⁷B. Frietsch, R. Carley, K. Döbrich, C. Gahl, M. Teichmann, O. Schwarzkopf, P. Wernet, and M. Weinelt, Rev. Sci. Instrum. 84, 075106 (2013).
- ²⁸S. Eich, A. Stange, A. Carr, J. Urbancic, T. Popmintchev, M. Wiesenmayer, K. Jansen, A. Ruffing, S. Jakobs, T. Rohwer, S. Hellmann, C. Chen, P. Matyba, L. Kipp, K. Rossnagel, M. Bauer, M. M. Murnane, H. C. Kapteyn, S. Mathias, and M. Aeschlimann, J. Electron Spectrosc. Relat. Phenom. 195, 231 (2014).
- ²⁹S. Plogmaker, J. a. Terschlüsen, N. Krebs, M. Svanqvist, J. Forsberg, U. B. Cappel, J.-E. E. Rubensson, H. Siegbahn, and J. Söderström, Rev. Sci. Instrum. **86**, 123107 (2015).
- ³⁰G. Rohde, A. Hendel, A. Stange, K. Hanff, L.-P. Oloff, L. X. Yang, K. Rossnagel, and M. Bauer, Rev. Sci. Instrum. 87, 103102 (2016).
- ³¹J. Ojeda, C. A. Arrell, J. Grilj, F. Frassetto, L. Mewes, H. Zhang, F. van Mourik, L. Poletto, and M. Chergui, Struct. Dyn. 3, 023602 (2016).
- ³²S. Passlack, S. Mathias, O. Andreyev, D. Mittnacht, M. Aeschlimann, and M. Bauer, J. Appl. Phys. **100**, 024912 (2006).
- ³³S. Hellmann, K. Rossnagel, M. Marczynski-Bühlow, and L. Kipp, Phys. Rev. B 79, 035402 (2009).
- ³⁴J. Graf, S. Hellmann, C. Jozwiak, C. Smallwood, Z. Hussain, R. Kaindl, L. Kipp, K. Rossnagel, and A. Lanzara, J. Appl. Phys. **107**, 014912 (2010).
- ³⁵C. W. Nicholson, A. Lücke, W. G. Schmidt, M. Puppin, L. Rettig, R. Ernstorfer, and M. Wolf, Science **362**, 821 (2018).
- ³⁶S. Backus, C. G. Durfee, M. M. Murnane, and H. C. Kapteyn, Rev. Sci. Instrum. **69**, 1207 (1998).
- ³⁷T. Popmintchev, M.-C. Chen, P. Arpin, M. M. Murnane, and H. C. Kapteyn, Nat. Photonics **4**, 822 (2010).
- ³⁸J. Heye Buss, H. Wang, Y. Xu, J. Maklar, F. Joucken, L. Zeng, S. Stoll, C. Jozwiak, J. Pepper, Y.-D. Chuang, J. D. Denlinger, Z. Hussain, A. Lanzara, and R. A. Kaindl, e-print arXiv:1811.00715 (2018).
- ⁵⁹C.-T. Chiang, M. Huth, A. Trützschler, M. Kiel, F. O. Schumann, J. Kirschner, and W. Widdra, New J. Phys. **17**, 13035 (2015).
- ⁴⁰A. K. Mills, S. Zhdanovich, A. Sheyerman, G. Levy, A. Damascelli, and D. J. Jones, Proc. SPIE **9512**, 95121I (2015).
- ⁴¹C. Corder, P. Zhao, J. Bakalis, X. Li, M. D. Kershis, A. R. Muraca, M. G. White, and T. K. Allison, Struct. Dyn. **5**, 054301 (2018).
- ⁴²Y. Ishida, T. Otsu, A. Ozawa, K. Yaji, S. Tani, S. Shin, and Y. Kobayashi, Rev. Sci. Instrum. 87, 123902 (2016).
- ⁴³P. Russbueldt, T. Mans, G. Rotarius, J. Weitenberg, H. Hoffmann, and R. Poprawe, Opt. Express 17, 12230 (2009).
- 44J. Boullet, Y. Zaouter, J. Limpert, S. Petit, Y. Mairesse, B. Fabre, J. Higuet, E. Mével, E. Constant, and E. Cormier, Opt. Lett. 34, 1489 (2009).
- ⁴⁵S. Hädrich, A. Klenke, J. Rothhardt, M. Krebs, A. Hoffmann, O. Pronin, V. Pervak, J. Limpert, and A. Tünnermann, Nat. Photonics 8, 779 (2014).
- ⁴⁶A. Vaupel, N. Bodnar, B. Webb, L. Shah, and M. Richardson, Opt. Eng. **53**, 051507 (2013).
- ⁴⁷T. V. Andersen, O. Schmidt, C. Bruchmann, J. Limpert, C. Aguergaray, E. Cormier, and A. Tünnermann, Opt. Express **14**, 4765 (2006).
- ⁴⁸C. Homann, C. Schriever, P. Baum, and E. Riedle, Opt. Express **16**, 5746 (2008).

- ⁴⁹ M. Emons, A. Steinmann, T. Binhammer, G. Palmer, M. Schultze, and U. Morgner, Opt. Express 18, 1191 (2010).
- ⁵⁰M. Schulz, R. Riedel, A. Willner, T. Mans, C. Schnitzler, P. Russbueldt, J. Dolkemeyer, E. Seise, T. Gottschall, S. Hädrich, S. Duesterer, H. Schlarb, J. Feldhaus, J. Limpert, B. Faatz, A. Tünnermann, J. Rossbach, M. Drescher, and F. Tavella, Opt. Lett. 36, 2456 (2011).
- ⁵¹J. Nillon, O. Crégut, C. Bressler, and S. Haacke, Opt. Express 22, 14964 (2014).
- ⁵²C. Schriever, S. Lochbrunner, P. Krok, and E. Riedle, Opt. Lett. 33, 192 (2008).
- ⁵³ M. Pergament, M. Kellert, K. Kruse, J. Wang, G. Palmer, L. Wissmann, U. Wegner, and M. J. Lederer, Opt. Express 22, 22202 (2014).
- ⁵⁴R. Riedel, A. Stephanides, and M. Prandolini, Opt. Lett. **39**, 1422 (2014).
- ⁵⁵M. Bradler and E. Riedle, Opt. Lett. **39**, 2588 (2014).
- ⁵⁶H. Fattahi, H. G. Barros, M. Gorjan, T. Nubbemeyer, B. Alsaif, C. Y. Teisset, M. Schultze, S. Prinz, M. Haefner, M. Ueffing, A. Alismail, L. Vámos, A. Schwarz, O. Pronin, J. Brons, X. T. Geng, G. Arisholm, M. Ciappina, V. S. Yakovlev, D.-E. Kim, A. M. Azzeer, N. Karpowicz, D. Sutter, Z. Major, T. Metzger, and F. Krausz, Optica 1, 45 (2014).
- ⁵⁷M. Krebs, S. Hädrich, S. Demmler, J. Rothhardt, A. Zaïr, L. Chipperfield, J. Limpert, and A. Tünnermann, Nat. Photonics 7, 555 (2013).
- M. Bradler, P. Baum, and E. Riedle, Appl. Phys. B: Lasers Opt. 97, 561 (2009).
 R. Riedel, M. Schulz, M. J. Prandolini, A. Hage, H. Höppner, T. Gottschall, J. Limpert, M. Drescher, and F. Tavella, Opt. Express 21, 28987 (2013).
- $^{\bf 60}{\rm Z.}$ Chang, Fundamentals of Attosecond Optics (CRC Press, Boca Raton, 2011).
- ⁶¹ M. Lewenstein, P. Balcou, M. Y. Ivanov, A. L'Huillier, and P. B. Corkum, Phys. Rev. A 49, 2117 (1994).
- 62 F. Krausz and M. Ivanov, Rev. Mod. Phys. 81, 163 (2009).
- ⁶³ H. Wang, Y. Xu, S. Ulonska, J. S. Robinson, P. Ranitovic, and R. A. Kaindl, Nat. Commun. 6, 7459 (2015).
- ⁶⁴C.-J. Lai, G. Cirmi, K.-H. Hong, J. Moses, S.-W. Huang, E. Granados, P. Keathley, S. Bhardwaj, and F. X. Kärtner, Phys. Rev. Lett. **111**, 073901 (2013).

- ⁶⁵C. M. Heyl, J. Güdde, A. L'Huillier, and U. Höfer, J. Phys. B: At., Mol. Opt. Phys. 45, 074020 (2012).
- 66 J. Rothhardt, M. Krebs, S. Hädrich, S. Demmler, J. Limpert, and A. Tünnermann, New J. Phys. 16, 033022 (2014).
- ⁶⁷C. M. Heyl, H. Coudert-Alteirac, M. Miranda, M. Louisy, K. Kovacs, V. Tosa, E. Balogh, K. Varjú, A. L'Huillier, A. Couairon, and C. L. Arnold, Optica 3, 75 (2016).
- ⁶⁸A. Comby, S. Beaulieu, E. Constant, D. Descamps, S. Petit, and Y. Mairesse, Opt. Express **26**, 6001 (2018).
- ⁶⁹BioMedical Instruments GmbH, glass nozzles of type Wieland, see https://biomedical-instruments.com/.
- $^{70}\mathrm{F}.$ Frassetto and L. Poletto, Nucl. Instrum. Methods Phys. Res., Sect. A 635, S75 (2011).
- ⁷¹ B. L. Henke, E. M. Gullikson, and J. C. Davis, At. Data Nucl. Data Tables 54, 181 (1993).
- 72 D. L. Windt, Comput. Phys. 12, 360 (1998).
- $^{73}\mathrm{R.}$ Doering, Handbook of Semiconductor Manufacturing Technology (CRC Press, Boca Raton, 2008).
- 74Reflective X-ray Optics LLC, see http://www.rxollc.com/.
- $^{75} \rm McPherson~234/30,$ equipped with a CCD camera, ANDOR IDUS DO420A -OE-995 and a concave platinum-coated grating with 2400 lines/mm.
- $^{76}\mbox{AXUV}\mbox{100Al},$ Opto diode corporation, read by a picoammeter, Keithley 6485.
- $^{\bf 77}{\rm H.}$ Pauly, Atom, Molecule, and Cluster Beams I (Springer, Berlin, Heidelberg, 2000).
- ⁷⁸C.-T. Chiang, M. Huth, A. Trützschler, F. O. Schumann, J. Kirschner, and W. Widdra, J. Electron Spectrosc. Relat. Phenom. **200**, 15 (2015).
- ⁷⁹SPECS Surface Nano Analysis GmbH, product spectrometer PHOIBOSTM 150 (2013), see http://www.specs.de/cms/front_content.php?idart=122.
- 80 SPECS Surface Nano Analysis GmbH, product manipulator CARVINGTM (2013), see http://www.specs.de.
- ⁸¹ A. Rettenberger, P. Leiderer, M. Probst, and R. Haight, Phys. Rev. B **56**, 12092 (1997).