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Title: Shallow forearc mantle dynamics and geochemistry: new insights from the IODP expedition 366

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the clasts. This episode of alteration resulted in a flattening of REE spectra and an increase of Zn concentrations in serpentinites. Otherwise, no systematic changes of ultramafic clasts chemistry or mineralogy were observed with increasing depth to the slab. The samples document previously undescribed prograde metamorphic events in the shallow portions of the Mariana subduction zone, consistent with a continuous burial of the serpentinitized forearc mantle during subduction. Similar processes, induced by the interaction with fluids released from the downgoing slab, likely occur in subduction zones worldwide. At greater depth, breakdown of brucite and antigorite will result in the massive transfer of fluids and fluid mobile elements, such as As, Sb and B, to the source of arc magmas.

Research Data Related to this Submission

There are no linked research data sets for this submission. The following reason is given:

Data are already provided in the article and the appendixes.

Abstract

The Mariana forearc is a unique setting on Earth where serpentinite mud volcanoes exhume clasts originating from depths of 15 km and more from the forearc mantle. These peridotite clasts are variably serpentinitized by interaction with slab derived fluid, and provide a record of forearc mantle dynamics and changes in geochemistry with depth. During International Oceanic Discovery Program (IODP) Expedition 366, we recovered serpentinitized ultramafic clasts contained within serpentinite muds of three different mud volcanoes located at increasing distance from the Mariana trench and at increasing depth to the slab/mantle interface: Yinazao (distance to the trench: 55 km / depth to the slab/mantle interface: 13km), Fantangisña (62 km / 14 km) and Asùt Tesoru (72 km / 18 km). Four different types of ultramafic clasts were recovered: blue serpentinites, lizardite-serpentinites, antigorite/lizardite- and antigorite-serpentinites. Lizardite-serpentinites are primarily composed of orange serpentine, forming mesh and bastite textures. Raman and microprobe analyses revealed that these textures contain a mixture of Fe-rich brucite ($XMg \sim 0.84$) and lizardite/chrysotile. Antigorite/lizardite- and antigorite-serpentinites record the progressive recrystallization of mesh and bastite textures to antigorite, magnetite and pure Fe-poor brucite ($XMg \sim 0.92$). Oxygen isotope compositions of clasts and pore fluids showed that the transition from lizardite to antigorite is due to the increase in temperature from 200°C to about 400°C within the forearc area above the slab/mantle interface. Lizardite-, antigorite/lizardite- and antigorite-serpentinites displayed U-shaped chondrite normalized Rare Earth Element (REE) patterns and are characterized by high fluid mobile element concentrations (Cs, Li, Sr, As, Sb, B, Li) relative to abyssal peridotites and/or primitive mantle. The recrystallization of lizardite to antigorite is accompanied by a decrease in Cs, Li and Sr, and an increase in As and Sb concentrations in the bulk clasts, whereas B concentrations are relatively constant. Some clasts are overprinted by blue serpentine, often in association with sulfides. Most of these blue serpentinites were recovered at Yinazao and the uppermost units of Fantangisña and Asùt Tesoru suggesting alteration in the shallower portions of the forearc, possibly during exhumation of the clasts. This episode of alteration resulted in a flattening of REE spectra and an increase of Zn concentrations in serpentinites. Otherwise, no systematic changes of ultramafic clasts chemistry or

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Highlights:

- Clasts from Mariana mud volcanoes record three different stages of serpentinization
- Transition lizardite to antigorite enhanced by an increase of temperature from 200 to 400°C
- Evidences for a continuous burial of the serpentinized forearc during subduction
- Phase transitions accompanied by a modification of trace element chemistry

1 **Shallow forearc mantle dynamics and geochemistry: new insights**
2 **from the IODP expedition 366**

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53 **1. Introduction**

54 Serpentinization of the forearc mantle in subduction zones is intimately related to the devolatilization
55 of the downgoing slab. During the onset of subduction (i.e., less than ~80 km depth) volatiles, such as
56 H₂O, C, S, etc., are released from the slab, rise through and interact with the mantle wedge. This
57 process influences the physical and mechanical properties of the slab/mantle wedge interface (Gerya et
58 al., 2002; Rüpke et al., 2004; Van Keken et al., 2011), the dynamics of mantle flows (Hilaret and
59 Reynard, 2009; Reynard, 2013; Wada et al., 2008) and controls deep volatile and redox-sensitive
60 element cycles (Debret et al., 2018a; Deschamps et al., 2011; Hattori and Guillot, 2007; Ribeiro and
61 Lee, 2017; Savov et al., 2007, 2005). In addition, serpentinites, either in the slab and/or the mantle
62 wedge, have the capability to retain large amounts of water (up to 13 wt%) down to 100-200 km and
63 up to their transformation into chlorite bearing harzburgites (e.g., Ulmer and Trommsdorff, 1995;
64 Wunder and Schreyer, 1997). However, despite its importance, relatively little is known about the
65 extent of serpentinization, redox state and chemistry of serpentized forearc mantle wedges
66 worldwide.

67 A common way to study serpentized forearc mantle wedges is by measuring seismic velocities
68 (Bostock et al., 2002; Kamimura et al., 2002). Although the geometry of forearcs is highly variable
69 and strongly controlled by the age of the slab (e.g., Van Keken et al., 2011; Reynard, 2013; Wada et
70 al., 2008), it is widely assumed that the forearc mantle wedge is highly serpentized, typically more

71 than 50 % (Bostock et al., 2002; Nagaya et al., 2016), at depth ranging from 30 to 80 km. The physical
72 properties of the mantle wedge serpentinites are often approximated by the behaviour of antigorite
73 (i.e., the high temperature and pressure form of serpentine). Both modelling and experimental studies
74 suggest that antigorite should be the dominant phase crystallizing in the forearc mantle (e.g.,
75 Christensen, 2004), but the importance of other serpentine varieties (lizardite and/or chrysotile) and
76 minerals such as brucite, talc or chlorite, is poorly constrained (Reynard, 2013).

77 The buoyancy of serpentinitized peridotites in dense anhydrous peridotites has lead several studies to
78 propose that the serpentinitized forearc mantle (or the so-called “serpentinitization channel” along the
79 slab-wedge interface) may contribute to the exhumation of high pressure terranes in active subduction
80 zones (e.g., Chemenda et al., 1995; Guillot et al., 2000; Schwartz et al., 2001). This conclusion has
81 been supported by the common occurrence of serpentinites with eclogitic rocks in mountain ranges.
82 However, geochemical studies have shown that the forearc mantle wedge constitutes an essential
83 reservoir for fluid mobile elements and water in subduction zones that must be dragged down by
84 corner flow to contribute to the elemental and isotope budgets of subduction zone magmas (e.g. Savov
85 et al., 2005, 2007; Hattori and Guillot, 2007; Deschamps et al., 2011; Ribeiro and Lee, 2017; Debret et
86 al., 2018a). For example, recent mass balance calculations show that the serpentinitized fore-arc mantle
87 could provide enough water (~7–78% of the total water injected at the trenches) to account for the
88 water outfluxes beneath the volcanic arc (Ribeiro and Lee, 2017). Additionally, numerical models by
89 Nagaya et al. (2016) suggest that convection could develop in serpentinitized forearc mantle wedges.
90 This result is compatible with previous numerical modelling by Honda et al. (2010) indicating that
91 convective flow can be induced in the forearc wedge mantle when the viscosity of the wedge mantle is
92 sufficiently low ($< \sim 4 \times 10^{19} \text{ Pa s}^{-1}$); compatible with estimates for the effective viscosity of antigorite
93 (Hilaret et al., 2007). The discrepancies between these buoyancy and viscosity-controlled models
94 emphasize the difficulty in assessing the dynamics of mantle flow in the serpentinitized mantle wedge
95 in subduction zones.

96 The Mariana forearc is a unique setting to sample and study the serpentinitized mantle wedge as
97 protrusions of hydrated mantle form serpentinite mud volcanoes on the outer forearc of the Izu–

98 Bonin–Mariana intra-oceanic subduction system (Fryer & Mottl, 1992; Fryer et al., 2012; Taylor &
99 Smoot, 1984). The mud volcanoes are composed of serpentinite muds with embedded ultramafic clasts
100 from the forearc mantle as well as mafic clasts from the subducting crust, originating from depths
101 greater than 15 km (Fryer et al., 2000; Maekawa et al., 1993). Here we take advantage of the recent
102 International Ocean Discovery Program (IODP) Expedition 366 to study serpentinitized ultramafic
103 clasts contained in the serpentinite muds of three mud volcanoes: Yinazao (formerly known as Blue
104 Moon), Fantangisña (Celestial) and Asùt Tesoru (Big Blue; Fryer et al., 2018). We show that clasts
105 from the Mariana forearc mantle are variably serpentinitized and preserve various stages of lizardite
106 recrystallization into antigorite, brucite and magnetite. Oxygen isotope chemistry and thermometry of
107 the clasts and pore fluids show that the transition from lizardite to antigorite is likely to occur between
108 200 to 400°C, in good agreement with thermodynamic calculations (Evans, 2004) and field
109 observations in alpine meta-ophiolites (Schwartz et al., 2013). No obvious correlation between the
110 distance of the mud volcanoes to the trench and the lizardite to antigorite transition was observed. The
111 absence of correlation suggests that complex convective flows of material occur within the mantle
112 wedge area.

113 **2. Geological setting**

114 The non-accretionary Mariana subduction system, involving the subduction of the Mesozoic Pacific
115 plate below the Philippine Sea plate, is located in the Western Pacific Ocean (Fig. 1). At the surface,
116 the Mariana forearc is characterised by multiple horst and graben structures that developed under
117 extensional stress caused by a rapid slab roll-back (Fryer, 1996; Harry and Ferguson, 1991). As a
118 result, numerous serpentinite mud volcanoes are situated at varying distances from the trench (Fig. 1).
119 They are formed by the eruption of mud flows consisting of unconsolidated serpentinite mud and
120 containing variably serpentinitized ultramafic clasts, as well as minor amounts of recycled Pacific plate
121 and of Philippine Sea plate materials (Fryer et al., 2018; Fryer, 2012; Maekawa et al., 1993; Pabst et
122 al., 2011). The serpentinite muds are derived from the forearc mantle where slab derived fluids interact
123 with ultramafic lithologies and are buoyantly transported to the seafloor (Fryer, 2012 and reference
124 therein). IODP Expedition 366 drilled at Yinazao (formerly known as Blue Moon), Fantangisña

125 (Celestial) and Asùt Tesoru (Big Blue) serpentinite mud volcanoes, which are located at distances of
126 55 km, 65 km and 70 km to the trench, respectively. Two additional serpentinites mud volcanoes,
127 namely South Chamorro (78 km) and Conical (86 km from the trench), were previously drilled during
128 Ocean Drilling Leg 195 and 125, respectively, and data from these sites will be incorporated here.

129 Yinazao is the closest to the trench. It lies at 15°43'N latitude and 147°11'E longitude (Fig. 1), at
130 about 13 km above the subducting slab (Fryer et al., 2018; Hulme et al., 2010; Oakley, 2008).
131 Previous studies estimated the temperature of the slab/ mantle interface below Yinazao at about 80°C
132 (Oakley, 2008; Hulme et al., 2010). Drilling took place at the flank (Site U1491) and the summit (Site
133 U1492) of this mud volcano (Fryer et al., 2018). The recovered cores consisted mainly of an
134 alternation of an uppermost altered units of red-brown pelagic mud, ranging from a few cm to up to 4
135 m thick, and units of green and blue-grey serpentinite pebbly mud (Fig. 2). The red-brown pelagic
136 units are interpreted as paleo-seafloor horizons altered in contact with seawater between two mud
137 eruptions (Fryer et al., 2018). Each unit contains between 5 and 10% ultramafic clasts, most of which
138 are fully serpentinized. Clasts recovered from the upper unit are affected by brown weathering and can
139 display a high degree of carbonation (up to 80%, Figs 2a-b, Appendix A); whereas, clasts from the
140 lower units are characterised by a dark blue colour (Figs 2c-d) and are frequently crosscut by mm-
141 wide chrysotile veins with crack-seal like textures, indicative of a late-stage alteration event.

142 Fantangisña mud volcano is located to the north of Yinazao, at approximately 16°32'N and
143 147°13'E (Fig. 1). It is situated at about 14 km above the slab (Hulme et al., 2010). The temperature of
144 the slab/ mantle interface below Fantangisña mud-volcano was estimated at about 150°C (Fryer et al.,
145 2018; Hulme et al., 2010). During IODP Expedition 366, both the summit (Site U1497) and the flank
146 (Site U1498) of this mud volcano were drilled (Fryer et al. 2018). The recovered cores consisted of
147 alternating silt- or sand-rich layers containing ultramafic clasts with a brown weathering colour, and of
148 green and/or blue-grey serpentinite pebbly mud embedding a large amount (about 20%) of dark blue
149 ultramafic clasts (Figs 2c-d) that are predominantly harzburgites and dunites displaying a wide degree
150 of serpentinization degree, from about 50 to 100%.

151 Asùt Tesoru mud volcano lies to the north of Fantangisña, at approximately 18°06 N and 147.06 E
152 (Fig. 1). It is located at about 18 km above the slab (Hulme et al., 2010). Temperatures at the
153 slab/mantle interface are estimated at about 250°C (Fryer et al., 2018; Hulme et al., 2010). Three sites
154 were drilled on the flanks (Sites U1493, U1494 and U1495) and one at the summit (Site U1496; Fryer
155 et al., 2018). The uppermost recovered units consist of pelagic mud and fine grained sandstone or
156 siltstone containing weathered ultramafic clasts. Lower units are mainly composed of green to blue-
157 grey serpentinite mud with 2 to 15% lithic clasts mainly consisting of variably serpentinized
158 harzburgites and dunites with a dark blue colour (Figs 2c-d; serpentinization degree from 30 to 100%).

159 **3. Petrology of the ultramafic clasts**

160 Forty three ultramafic clasts were examined for their petrography onshore (see Appendix B for sample
161 IODP full names, locations and classification). The clasts are variably serpentinized harzburgites and
162 dunites of several centimetres to tens of centimetres long (Fig. 2). Many clasts preserve different
163 serpentinization stages reflecting various episodes of fluid infiltration within the mantle wedge.

164 *3.1 Identification of serpentine varieties*

165 The mineralogy of the ultramafic clasts was characterized by combining electron microprobe analyses
166 and Raman spectroscopy. These methods have been used to differentiate between serpentine varieties
167 and co-existing brucite (e.g., Debret et al., 2013; Groppo et al., 2006; Schwartz et al., 2013;
168 Schwarzenbach et al., 2016). *In situ* major element analyses were performed with a CAMECA SX 100
169 electron microprobe at the Laboratoire Magmas et Volcans (Clermont-Ferrand, France). Operating
170 conditions of 15 kV accelerating voltage, a sample current of 15 nA and a counting time of 10
171 s/element were used, except for Ni (20 s). Microprobe analyses are given in Appendix C. Raman
172 spectroscopy was performed at the University of Cambridge (UK). Raman spectra were collected in
173 the 150–1300 cm^{-1} and 3500–3800 cm^{-1} spectral ranges using a confocal Labram HR300 Raman
174 spectrometer (Horiba Jobin Yvon) of 300 mm focal length equipped with a holographic grating of
175 1800 gr mm^{-1} coupled to a Peltier cooled front illuminated CCD detector, 1024×256 pixels in size.
176 This configuration allowed for a spectral resolution of about 1.4 cm^{-1} per pixel. The excitation line at

177 532 nm was produced by a diode-pumped solid state laser (Laser Quantum) focused on the sample
178 using an Olympus 50 X objective (0.42 N.A.). Spectra were corrected from a linear baseline using the
179 fitting software Peakfit. The laser power was set at energies ranging from 5 mW to 500 μ W in order to
180 avoid degradation of serpentine or oxide minerals (Debret et al., 2013; Debret et al., 2014; Faria &
181 Vena, 1997), and the sample surface was checked after each analysis. In order to properly define the
182 different serpentine varieties, spectra of lizardite, chrysotile, antigorite and brucite were selected as
183 references (Appendix D).

184 3.2 Serpentinization at Yinazao mud volcano

185 Close to the seafloor, the clasts mainly consisted of carbonated breccia made of dusty calcite and/or
186 aragonite embedding serpentinite fragments of various sizes, from 100 μ m to several centimetres (Figs
187 2a-b, 3a). The degree of carbonation varies from 20 to 80% in the different clasts. Serpentine minerals
188 display pseudomorphic mesh and bastite textures that replace mantle olivines and orthopyroxenes,
189 respectively. Large veins of chrysotile with crack seal like textures crosscut mesh and bastite textures.
190 The rims of many clasts are affected by an episode of late alteration, consisting of brownish clay
191 minerals, that overprints previous textures (Fig. 2b; Appendix A). This late alteration stage is in
192 accordance with results from other studies on Conical and South Chamorro mud volcanoes drilled
193 during previous ODP expeditions (e.g., Kahl et al., 2015).

194 In the deeper units, clasts consist mainly of blue serpentine, spinel and sulfides with rare
195 hydrogrossular and relicts of mantle olivine and orthopyroxene (subsequently called ‘blue
196 serpentinites’; Figs 3b-d). Spinel are homogeneous and display euhedral shapes. They can be
197 surrounded by $\leq 50 \mu$ m wide coronas of magnetite. Spinel cores have XCr (Cr / [Cr + Al]) of 53-54
198 and XMg (Mg / [Fe + Mg]) of 54-57 (Fig. 4). Serpentine forms mesh and bastite textures, replacing
199 olivine and orthopyroxene, and display a pale blue colour in plane polarized light (Figs 3b,c). The
200 Raman spectra of blue serpentine are characterized by peaks at 225, 381, 525, 692, 1095 and 3697 cm^{-1}
201 ¹ characteristic of chrysotile. Brucite is intergrown with chrysotile at the microscale as indicated by
202 additional Raman peaks at 280, 440, 3643 and 3650 cm^{-1} (Fig. 3b). Blue serpentine compositions are
203 characterized by XMg of 0.85 to 0.94, and SiO₂ and FeO contents ranging from 17.9 to 40.3 wt% and

204 from 4.2 to 11.3 wt%, respectively (Fig. 5). The most Si-rich analyses display a broad match to
205 serpentine mineral stoichiometry, whereas regression analysis of the whole data set suggests variable
206 proportions of Mg-Si-serpentine and Si-free, Fe-rich brucite endmembers (Fig. 5). The low Si analyses
207 can be attributed to the presence of Fe-brucite at nanoscale, as this mineral has a low XMg (~ 0.84 as
208 shown by regression analyses in Fig. 5a) relative to serpentine and does not incorporate silica (see also
209 Kahl et al., 2015; Schwarzenbach et al., 2016). Bastite textures are occasionally associated with
210 roundish hydrogrossular crystals of about 50 μm width. Several studies have reported the presence of
211 hydrogrossular in different serpentinization environments and this has, in most cases, been attributed
212 to an excess in Al during the final stages of serpentine growth after pyroxene (e.g., Beard et al., 2009).

213 The centre of mesh and bastite textures is often associated with large opaque aggregates of ~ 200
214 μm width (Figs 3b,c). These aggregates are composed of lamellar minerals associated with framboidal
215 Fe-sulfides (pyrite), ranging in size from 0.5 to 2 μm , formed inside the intergranular porosity (Fig.
216 3d). Raman analyses of the lamellar minerals show three main peaks at 280, 440 and 3639 cm^{-1}
217 corresponding to brucite, with four small additional peaks at 369, 521, 689 and 3689 cm^{-1} suggesting
218 the presence of small amounts of serpentine (Fig. 3b). Occasionally, 50 to 200 μm wide veins
219 consisting mainly of euhedral pyrite crosscut the serpentine textures (Fig. 3c).

220 *3.3 Serpentinization at Fantangisña and Asùt Tesoru mud volcanoes*

221 Clasts recovered from Fantangisña and Asùt-Tesoru mud volcanoes display similar textures to each
222 other. Samples from the uppermost units are also affected by sub-surface alteration such as clay
223 mineral crystallization and carbonation. However, in contrast to Yinazao, the formation of blue
224 serpentine and sulfides was mainly observed close to the seafloor at Fantangisña and Asùt Tesoru. The
225 Blue serpentine can either completely replace former textures or be limited to clast rims (Fig. 2b;
226 Appendix A). In lower units, three different types of ultramafic clasts have been identified (Figs 2c-d;
227 6-8); their distribution appears to be random within the mud volcanoes (see Appendix B for more
228 details):

229 (1) The first group (lizardite-rich serpentinites, referred to as Liz-serpentinites from here on)
230 corresponds to variably serpentinized peridotites that can preserve mantle minerals (olivine,
231 orthopyroxene, spinel and rare clinopyroxene). These Liz-serpentinite clasts were the most commonly
232 reported during previous ODP expeditions, i.e. in South Chamorro and Conical mud volcanoes (e.g.,
233 D'Antonio and Kristensen, 2004; Savov et al., 2005, 2007; Kahl et al., 2015). Olivine XMg and NiO
234 contents range from 0.91 to 0.93 and from 0.37 to 0.44 wt%, respectively, whereas MnO contents are
235 below 0.1 wt%. Orthopyroxene is characterized by Cr₂O₃ and Al₂O₃ contents ranging from 0.1 to 1.88
236 wt% and 0.1 to 0.5 wt%, respectively. Spinel relicts have euhedral shapes with dark cores and narrow
237 (< 100 µm) lighter rims that correspond to a compositional zoning from Mg and Al-rich cores (XMg =
238 0.30 - 0.59; XCr = 0.51 - 0.95) to Cr-rich rims (XMg ~ 0; XCr ~ 1; Fig. 4). Small magnetite crystals
239 are present at the boundary between serpentine and spinel. The serpentinization degrees of the clasts
240 vary from about 30 to 90%. In slightly serpentinized clasts, serpentine crystallizes as brownish veins
241 of 20 to 300 µm width, with regular shape, crossing olivine and orthopyroxene (Fig. 6a). Raman
242 spectra indicate mixtures of brucite and lizardite and/or chrysotile inside the veins. Serpentine fully
243 replaces olivine and orthopyroxene by forming mesh and bastite textures where the local
244 serpentinization degree is > 60 %. No magnetite was observed in the mesh centres. Serpentine
245 compositions are characterized by low SiO₂ (31.8 - 40.6 wt%), XMg (0.88 - 0.95) and high FeO (3.5 -
246 8.1 wt%; Fig. 5) suggesting variable proportions of Fe-brucite and serpentine at microscale. Bastite is
247 associated with amphibole minerals, which form pale green needles about 50 µm in length. Rare
248 magnetite grains have been observed in these clasts; they crystallize mainly in the centre of brucite
249 veins of 20 to 200 µm width crossing mantle minerals, lizardite veins, mesh or bastite textures (Fig.
250 6b). Towards the rims, the brucite + magnetite veins are surrounded by a corona of small antigorite
251 lamellae, about 30 µm in length.

252 (2) The second group (antigorite-/lizardite-rich serpentinites, referred as Atg/Liz-serpentinites
253 from here on) corresponds to highly serpentinized peridotites (serpentinization degree > 70%)
254 displaying mesh and bastite textures associated with antigorite (Fig. 7). Few primary mantle spinel
255 relicts have been observed. XMg and XCr of spinel range from 0.19 to 0.39 and from 0.63 to 0.81,

256 respectively. Spinel frequently displays thin, about 50 μm wide, coronas of chromite ($X_{\text{Cr}} > 0.9$) and
257 magnetite toward the rim (Fig. 4). Serpentine mesh centres are often several hundred micrometres in
258 size and characterised by a homogenous grey colour with weak birefringence and occasionally
259 undulatory extinction. Raman spectra of the mesh and bastite textures denote mixtures between
260 lizardite (or chrysotile) and antigorite, with pronounced peaks at 1044 and 3700 cm^{-1} and at 1100 and
261 3685 cm^{-1} , corresponding to antigorite and lizardite, respectively (Fig. 7a). The mesh rims are
262 composed of antigorite needles associated with thin veinlets of pure brucite and magnetite in their
263 centres (Fig. 7a). Chemical compositions of serpentine minerals are homogeneous throughout the
264 samples. They are characterized by higher SiO_2 (39.8 – 45.4 wt%) and X_{Mg} (0.92-0.97) and lower
265 FeO (1.9-5.2 wt%) relative to serpentine/brucite textures in the Liz-serpentinites (Fig. 5). This
266 suggests the absence of brucite at microscale. The amount of antigorite is highly variable from one
267 sample to another, varying from about 40 to 80 modal %. The serpentinites are crosscut by 200 to 500
268 μm wide veins of brucite and magnetite (Fig. 7b). Mesh and bastite textures in contact with these veins
269 are recrystallized to antigorite lamellae and needles about 30 μm in length.

270 (3) The third group corresponds to antigorite-bearing serpentinites (Atg-serpentinites from here on)
271 mainly made of antigorite lamellae, about 50 μm long, with interstitial brucite and magnetite (Fig. 8).
272 The existence of antigorite bearing clasts have also been reported at Conical and South Chamorro mud
273 volcanoes (e.g., Alt and Shanks, 2006; Murata et al., 2009). Antigorite have higher SiO_2 contents (42.2
274 - 44.9 wt%) and X_{Mg} (0.96-0.99) and lower FeO contents (1.2-3.2 wt%) relative to serpentine in the
275 Liz- or Atg/Liz-serpentinites (Fig. 5). Brucite has crystallized as millimetre-sized patches containing
276 euhedral magnetite grains about 200 μm in width (Fig. 8). Brucite has MgO and FeO contents ranging
277 from 63.3 to 64.9 wt% and 9.7 to 9.9 wt%, respectively, and an average X_{Mg} of 0.92, which is
278 significantly higher than that of the Fe-Brucite in Liz-serpentinites ($X_{\text{Mg}} \sim 0.84$, Fig. 5). The Atg-
279 serpentinites are characterized by the presence of euhedral hydro-andradite crystals that display a
280 diamond shape and contain numerous inclusions of magnetite. The crystallization of andradite in
281 serpentinites has been observed in several settings (e.g. Frost, 1985) and can be attributed to a low
282 silica activity during alteration (Frost & Beard, 2007). Magnetite is abundant throughout the samples.

283 **4. Methods**

284 *4.1 Oxygen isotope geochemistry*

285 *4.1.1 Mineral separates*

286 The oxygen isotope compositions of serpentine and magnetite mineral separates were measured at the
287 University of Texas at Austin using a ThermoElectron MAT 253 mass spectrometer. Serpentinite
288 samples were crushed and handpicked under a binocular microscope in order to visually assess
289 mineral separate purity. In many samples, serpentine and magnetite were extensively intergrown and
290 pure magnetite could not be separated, therefore most samples could not be confidently analysed for
291 the $\delta^{18}\text{O}$ value of magnetite. Oxygen isotope ratios were determined on ~ 2.0 mg of mineral separates
292 using the laser fluorination method of Sharp (1990). Standard UWG-2 ($\delta^{18}\text{O}_{\text{garnet}} = +5.8\text{‰}$; Valley et
293 al., 1995) and in-house standard Lausanne-1 ($\delta^{18}\text{O}_{\text{quartz}} = +18.1\text{‰}$) were analysed to verify precision
294 and accuracy. All $\delta^{18}\text{O}$ values are reported relative to SMOW, where the $\delta^{18}\text{O}$ value of NBS-28 is
295 $+9.7\text{‰}$. The error on $\delta^{18}\text{O}$ values is $\pm 0.1\text{‰}$, based on the long-term average of standard analyses.

296 *4.1.2 Pore fluids*

297 Whole-round (WR) core samples were taken immediately after core sectioning on the deck for the
298 subsequent extraction of interstitial (pore) water (IW). The length of the WR core taken for IW
299 analyses varied from 10 cm in the upper units, to 40 cm in the deeper units where the volume of
300 extracted IW was limited. Although this sacrificed a large amount of core material, this was the only
301 way to get sufficient volume of pore fluids for the deeper samples, and only core sections which
302 appeared to be highly homogenous were used. Typically one WR per section was collected between 0
303 and 10 mbsf, and 2 WR were selected every 10 m of depth from 10 mbsf to total depth of the core.
304 Whole-round samples were processed in a nitrogen filled glove-bag after cooling in a refrigerator for
305 about one hour. First the cored material was extruded from the core liner, then portions of the material
306 that were potentially contaminated by seawater and sediment smearing were removed by scraping the
307 core's outer surface with a spatula. For advanced piston (APC) cores about 0.5 cm of material from
308 the outer diameter and the top and bottom faces was removed. In contrast, material recovered by

309 extended core barrel (XCB) coring required additional removal of material, and as much as two-thirds
310 of the sediment was removed from each WR sample. The remaining inner core of uncontaminated
311 material (~150–300 cm³) was placed into a titanium squeezer (modified after Manheim and Sayles,
312 1974) and compressed using a laboratory hydraulic press to extract pore water, using a total pressure
313 up to 30 MPa. Fluids extracted from the compressed sediment sample were filtered through a pre-
314 washed Whatman No. 1 filter situated above a titanium mesh screen. Approximately 10-80 mL of IW,
315 depending on the length of the WR being processed, was collected in acid-cleaned plastic syringes
316 attached to the squeezing assembly and filtered again through a Gelman polysulfone disposable filter
317 (0.45 µm). After extraction, the squeezer parts were thoroughly cleaned with shipboard water, rinsed
318 with de-ionized water, and dried. Pore fluids were syringe filtered into small, pre-cleaned (acid
319 washed), Nalgene plastic vials, capped and then immediately flash frozen in liquid nitrogen to prevent
320 evaporation.

321 Samples were measured via a Thermo Gas Bench II connected to a Thermo Delta Advantage mass
322 spectrometer in continuous flow mode at Union College (Schenectady, New York – sample analyses
323 are displayed in Appendix E). Three inhouse laboratory standards were used for isotopic corrections,
324 and to assign the data to the appropriate isotopic scale using linear regression. These standards were
325 calibrated directly to VSMOW (0.0‰) and SLAP (-55.50‰). The inhouse standards have δ¹⁸O values
326 that range from -0.6‰ to -16.52‰. The combined uncertainty (analytical uncertainty) for δ¹⁸O of IW
327 samples is ± 0.02‰ (SMOW), based on 8 internal tap water standards over two analytical sessions.

328 *4.2 Bulk rock major and trace elements analyses*

329 A suite of 25 representative serpentinized ultramafic clasts were analysed for major and selected trace
330 elements by Inductively Coupled Plasma Optical Emission Spectrometry (Thermo ICP-OES Icap
331 6500) at the SARM (Service d'Analyses des Roches et des Minéraux Nancy, France – Appendix F).
332 This sample set includes 5 Blue-serpentinites recovered at Yinazao (Site U1492, samples: M2, M3,
333 M5, M6, M7), 1 Blue-serpentinite (Site U1496, sample: M24), 4 Liz-serpentinites (Site U1493,
334 samples: TSB55, M9, M10; Site U1496, sample: M19), 2 Atg/Liz-serpentinites (Site U1495, samples:
335 M12, M13), 4 Atg-serpentinites (Site U1495, samples: M14, M15, M16, M17) and 1 Brucitite (Site

336 U1496, sample: M20) recovered at Asùt Tesoru, 2 Blue-serpentinities (Site U1497, sample: M30; Site
337 U1498, sample: M38), 2 Atg/Liz-serpentinities (Site U1497, sample: M32; Site U1498, sample: M45)
338 and 3 Atg-serpentinities (Site U1497, samples: M50, M51, TSB102) recovered at Fantangisña. Sample
339 digestions for major (SiO_2 , Al_2O_3 , Fe_2O_3 , MnO , MgO , CaO , Na_2O , K_2O , TiO_2) and trace elements
340 (Co, Cr, Ga, Ge, Ni, Sc, V, Zn, Cu) were performed on LiBO_2 fluxed fusions following the procedures
341 described by Carignan et al. (2001). Boron concentrations were measured by spectrophotometric
342 determination at the SARM. Samples were dissolved by fusion with anhydrous sodium carbonate. The
343 reproducibility of the standard was better than 2% for major elements, 5% for Co, Cr, Ga, Ge, Ni, Sc,
344 V, Zn, B and Li, and 10% for Cu based on repeated analyses of UB-N (serpentine standard from the
345 Centre de Recherches Pétrographiques et Géochimiques (CRPG) of Nancy, France). The analyses
346 were accurate within 1-5% for SiO_2 , Al_2O_3 , Fe_2O_3 , MnO , MgO , CaO , Na_2O , K_2O and within 1-10%
347 for TiO_2 based on repeated analyses of U.S. Geological Survey, CRPG and Geological Survey of
348 Japan standards BIR-1, UB-N and JP-1, and within 1-10% for Co, Cr, Ga, Ge, Ni, Sc, Cu, Zn, B, Li
349 and better than 15% for V based on UB-N analyses (see Appendix F for comparison with standard
350 preferred values).

351 A subset of 20 serpentinites was analysed for rare earth elements (REE), Y, Sr, Li, Rb, Cs, Cd, Pb,
352 As, Sb, Ba, U, Th, Nb, Ta, Hf, Zr, V, Ga, Cu and Zn using a High Resolution Sector Field ICP-MS
353 (Element XR) at the Vrije Universiteit Brussel (VUB, Belgium - Appendix F). This sample subset
354 includes 4 Blue-serpentinities recovered at Yinazao (site U1492, samples: M2, M3, M6, M7), 1 Blue-
355 serpentine (site U1496, sample: M24), 3 Liz-serpentinities (site U1493, samples: TSB55, M9, M10),
356 2 Atg/Liz-serpentinities (site U1495, samples: M12, M13), 4 Atg-serpentinities (site U1495, samples:
357 M14, M15, M16, M17) and 1 Brucitite (site U1496, sample: M20) recovered at Asùt Tesoru, 2 Blue-
358 serpentinities (site U1497, sample: M30; U1498, sample: M38) and 2 Atg/Liz-serpentinities (site
359 U1497, sample: M32; site U1498, sample: M45) recovered at Fantangisña. Samples were digested
360 with a 1:1 mixture of HCl and HF for 4 days in par bombs. The samples were analysed in low
361 resolution mode after dilution in 2% HNO_3 of 500 for most of trace elements and in medium
362 resolution mode after dilution in 2% HNO_3 of 2000 for Cu, Zn, As, Ba and Pb. The external precision

363 and accuracy were determined by analysing known ultramafic rock standards (UB-N from the CRPG
364 Nancy, JP-1 from Geological Survey of Japan, PCC-1 and DTS-2b from US Geology Survey). One of
365 the challenges of measuring refractory peridotites is their very low abundance of many trace elements
366 (e.g., REE, HFSE, U, Th, Pb, As). As a consequence, there is a lack of precise and accurate data for
367 reference materials of refractory peridotites (PCC-1, JP-1 and DTS-2b). On the basis of UB-N
368 analysis, reproducibility is better than 10% for most of the trace elements and between 10 and 15% for
369 Hf and Th (see Appendix F). The values obtained for rock standards UB-N, PCC-1, JP-1 and DTS-2b
370 during this study are reported in Appendix F and are in good agreement with previously published data
371 within a 2 standard deviation error.

372 **5. Results and discussion of bulk rock oxygen isotope geochemistry**

373 Results of whole rock analyses are given in Appendix F and are in good agreement with previous
374 studies of serpentinite clasts from nearby seamounts (South Chamorro and Conical; e.g., Kodolanyi et
375 al., 2011; Parkinson & Pearce, 1998; Savov et al., 2005, 2007; Fig. 9) and with the shipboard analyses
376 (Fryer et al., 2018). Relatively low Al_2O_3/SiO_2 (<0.03) and high MgO/SiO_2 (>0.98) of the ultramafic
377 clasts are consistent with a refractory protolith, i.e., dunite or pyroxene-poor harzburgite (e.g., Godard
378 et al., 2008). The high loss on ignition values (>15 wt%) and low SiO_2 contents (<36 wt%) of some
379 samples are in agreement with the crystallization of high amounts of sulfides (e.g. blue serpentinites)
380 and/or brucite during peridotite serpentinization.

381 Chondrite-normalized REE plots are presented in Fig. 8a. Overall, the studied serpentinites are
382 depleted in REE concentrations compared to chondrite values, and LREE (Light REE, 0.002 to 0.143
383 chondritic values) are more depleted than HREE (Heavy REE, 0.03 to 0.3 times chondritic values).
384 Among the recovered clasts, two different patterns are observed. Group 1 corresponds to most of the
385 blue serpentinite samples. They are characterized by relatively flat REE patterns with slightly higher
386 concentrations of HREE compared to LREE ($La_N/Yb_N = 0.3-0.6$; $Gd_N/Yb_N = 0.2-0.5$; N: Chondrite
387 normalized). This group reflects the pattern of the serpentinite mud erupted at the mud volcanoes (e.g.,
388 Savov et al., 2004; Fig. 10a) and may therefore reflect late stage re-equilibration between the mud and
389 the ultramafic clasts.

390 Group 2 includes blue-, Liz-, Atg/Liz- and Atg- serpentinites. These samples display U-shape REE
391 patterns, with depletion in the MREE (Middle Rare Earth Elements) relative to the LREE ($La_N/Yb_N =$
392 $0.03-2$) and HREE ($Gd_N/Yb_N = 0.02-0.4$; Fig. 10a), and also possess positive Eu anomalies ($Eu/Eu^* =$
393 $Eu_N/[(Sm_N+Gd_N)/2]$). The LREE concentrations progressively increase from the Liz-serpentinites
394 ($La_N/Yb_N = 0.04-0.13$) to the Atg/Liz- ($La_N/Yb_N = 0.12-2$) and Atg-serpentinites ($La_N/Yb_N = 0.1-0.4$).
395 The Eu anomaly is more pronounced in Liz- and Atg/Liz-serpentinites ($Eu/Eu^* = 0.9-6.8$) relative to
396 Atg-serpentinites ($Eu/Eu^* = 0.7-1.8$). These patterns are similar to those reported in the Conical mud
397 volcano and have been interpreted as inherited from the peridotite protolith (Parkinson and Pearce,
398 1998). In Fig. 10b, Group 2 serpentinites are characterized by a strong depletion in most incompatible
399 elements with respect to primitive mantle. The studied clasts show enrichments in Cs ($Cs_N/La_n > 21$; n:
400 Primitive Mantle normalization) and positive anomalies in U ($U_n/Th_n > 2$), Pb ($Pb_n/La_n > 4$), Sr
401 ($Sr_n/Pr_n > 2$), As, Sb ($As_n/Pr_n > 1500$) and Li ($Li^* > 60$, $Li^* = Li / [Dy/2 + Y/2]$).

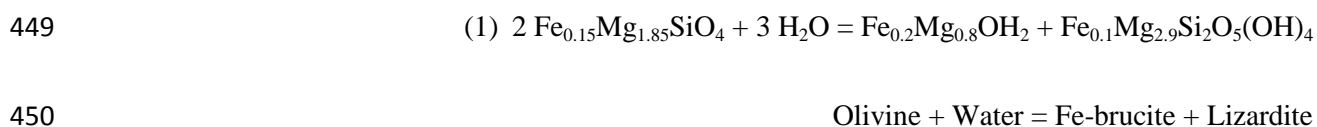
402 Variations of fluid mobile elements (FME) and metal concentrations are observed for the different
403 serpentinite types. The concentrations of As and Sb are lower in Liz-serpentinites (As = 0.7-1.3 ppm;
404 Sb < 0.001 ppm;) compared to Atg-serpentinites (As = 1.5-9.2 ppm; Sb = 0.03-0.21 ppm; B = 10-97
405 ppm), whereas Li (from 4-6 ppm to 0.4-0.8 ppm), Sr (from 1-8.5 ppm to 0.1-1 ppm) and Cs (from
406 0.14-0.22 ppm to <0.01-0.03 ppm) concentrations decrease from Liz- to Atg- serpentinites (Fig. 11). B
407 and Zn concentrations are relatively constant in Liz- (B = 20-30 ppm; Zn = 35-65 ppm), Atg/Liz- (B =
408 48-59 ppm, Zn = 48-55 ppm) and Atg-serpentinites (B = 10-49 ppm; Zn = 36-52 ppm). It should be
409 noted that boron (B) is highly enriched on all the studied samples relative to primitive mantle ($B_{PM} =$
410 0.19 ± 0.02 ppm, PM: Primitive Mantle; Marschall et al., 2017). Of all samples, blue serpentinites are
411 characterized by the highest Zn (51-92 ppm), and highly variable Sr (2-28 ppm) and B (7-250 ppm)
412 concentrations (Fig. 11c; Appendix F). It is unknown whether or not these elements are solely carried
413 by serpentine minerals, which can incorporate these elements in its structure (e.g., Pabst et al., 2011;
414 Debret et al., 2017), or by other accessory (micro- to nano-) phases (e.g., sulfides, spinels, hydro-
415 garnets...).

416 Eight serpentine separates have $\delta^{18}\text{O}_{\text{Srp}}$ values of 5.8 to 8.3‰ (Table 1), overlapping with the range of
417 analyses from previous Mariana mud volcano studies (6.5 to 10.8‰, reported by Alt & Shanks, 2006;
418 5.8 to 8.5‰ by Sakai et al., 1990; 6.1 to 10.5‰ by Kahl et al., 2015), and one magnetite separate has
419 $\delta^{18}\text{O}_{\text{Mgt}}$ value of 1.8‰ (Table 1), in good agreement with Alt and Shanks (2006) who reported $\delta^{18}\text{O}_{\text{Mgt}}$
420 values of 0 to 2‰ in antigorite-rich samples. No obvious changes of $\delta^{18}\text{O}_{\text{Srp}}$ are observed with
421 increasing distance from the trench in this study ($\delta^{18}\text{O}_{\text{Srp}}$ [Yinazao] = 6.4 ‰; $\delta^{18}\text{O}_{\text{Srp}}$ [Fantangisña] =
422 8.3 ‰; $\delta^{18}\text{O}_{\text{Srp}}$ [Asùt Tesoru] = 5.8-8.3 ‰) or in previous studies ($\delta^{18}\text{O}_{\text{Srp}}$ [S. Chamorro] = 6.4-10.5 ‰;
423 $\delta^{18}\text{O}_{\text{Srp}}$ [Conical] = 6.1-10.8 ‰; Alt and Shanks, 2006; Kahl et al., 2015; Sakai et al., 1990). However,
424 systematic variations of $\delta^{18}\text{O}_{\text{Srp}}$ relative to sample mineralogy and in pore fluids are observed. Blue
425 serpentine displays $\delta^{18}\text{O}$ values ranging from 6.4 to 7.4‰, whereas higher $\delta^{18}\text{O}$ values are observed in
426 Liz-serpentinites ($\delta^{18}\text{O}_{\text{serp}}$ 6.8 to 7.6‰) compared to Atg/Liz-serpentinites ($\delta^{18}\text{O}_{\text{serp}}$ 5.8 to 6.1‰).
427 Antigorite in the Atg-serpentinites records $\delta^{18}\text{O}_{\text{Srp}}$ values varying from 7.1 to 8.3‰, whereas
428 associated magnetite has a $\delta^{18}\text{O}_{\text{Mgt}}$ value of 1.8 ‰. Pore fluid $\delta^{18}\text{O}$ values range from -1.39 to -0.14‰
429 in Yinazao, from -0.03 to 0.25‰ in Fantangisña and from 1.73 to 1.97‰ in Asùt Tesoru. Although
430 highly variable with depth, these values suggest an increase in the $\delta^{18}\text{O}$ values of the pore fluid from
431 shallow Yinazao (mean $\delta^{18}\text{O}_{\text{fluid}}$ = -0.90‰, n = 10) to the deeper Fantangisña (mean $\delta^{18}\text{O}_{\text{fluid}}$ = 0.11‰,
432 n = 6) or Asùt Tesoru (mean $\delta^{18}\text{O}_{\text{fluid}}$ = 1.83‰, n = 4; Fig. 12). These values are significantly lower
433 than those reported for South Chamorro (mean $\delta^{18}\text{O}_{\text{fluid}}$ = 2.5‰) and Conical (mean $\delta^{18}\text{O}_{\text{fluid}}$ = 4‰;
434 Mottl et al., 2003), which are further from the trench.

435 **6. Reconstructing forearc serpentization conditions**

436 Textural relationships between the different serpentine generations allow the reconstruction of a semi-
437 quantitative temperature evolution of the serpentization conditions within the Marianas forearc
438 mantle wedge and thus a discussion of subduction dynamics. Forearc mantle wedge peridotites are
439 former sub-arc peridotites that underwent extensive partial melting before being dragged into the
440 forearc by mantle convection (e.g. Parkinson & Pearce, 1998). These peridotites are hydrated by slab-
441 derived fluids and progressively transformed into serpentinites. The formation of early brown lizardite
442 bearing veins crosscutting olivine and orthopyroxene in Liz-serpentinites (Fig. 6a) constitutes the first
443 stage of the forearc mantle wedge hydration and serpentization. These textures have been observed

444 at the Fantangisña and Asùt Tesoru mud volcanoes as well as at the South Chamorro and Conical mud
445 volcanoes during previous IODP expeditions (e.g. Kahl et al., 2015). The presence of significant
446 amounts of brucite in Liz-serpentinites (Figs 5, 6a) can be attributed to the refractory composition of
447 forearc mantle peridotites. In Mg-rich systems and at low-temperatures (e.g., Klein et al., 2014), the
448 serpentinization of olivine results in magnesium excess allowing brucite precipitation:



451 Here the XMg of brucite is assumed to be 0.8 based on linear regression of microprobe analyses (Fig.
452 5). Although lizardite and brucite can coexist over a large range of temperatures during the
453 serpentinization process, the absence of magnetite during brown serpentine crystallization suggests
454 rather low serpentinization temperature (< 200°C; Bonnemains et al., 2016; Klein et al., 2014).
455 Assuming equilibria between lizardite and associated pore fluids, serpentine crystallisation
456 temperatures (Table 1) were calculated based on the serpentine–water oxygen isotope fractionation
457 from Saccocia et al. (2009). In agreement with petrographic observation, the temperature estimates of
458 lizardite crystallization vary between 203 and 211°C (Table 1). It should, however, be noted that we
459 cannot exclude an intergrowth of the serpentine separates with microscale brucite; such impurities
460 would then lead to an overestimation of the calculated temperatures, as discussed by Alt and Shanks
461 (2006). The crystallization temperatures estimates must therefore be considered as maxima.

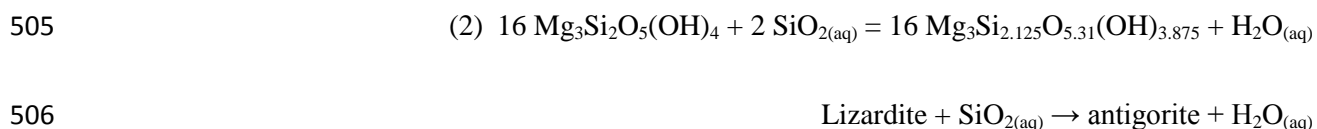
462 The crystallization of antigorite at the expense of lizardite-bearing textures has been observed at
463 Fantangisña and Asùt Tesoru in Atg/Liz-serpentinites and has also been reported at South Chamorro
464 and Conical mud volcanoes (e.g. Alt and Shanks, 2006). The formation of antigorite corresponds to a
465 second stage of serpentinization that is accompanied with the precipitation of large amounts of
466 magnetite and the crystallization of Mg-rich brucite (XMg ~ 0.9; Figs 7 and 8). The transition of
467 lizardite to antigorite in subduction settings is commonly interpreted to result from increasing P-T
468 conditions during prograde metamorphism (Debret et al., 2013; Evans, 2004; Scambelluri et al., 2004;
469 Schwartz et al., 2013; Wunder et al., 2001). The $\delta^{18}\text{O}$ values of serpentine minerals crystallizing in

470 Atg/Liz-serpentinites indicate crystallisation temperatures of 230-240°C if we assume equilibria
471 between serpentine and associated pore fluids. Those estimates are however at the lower range of
472 previous thermodynamic estimates or natural observations that predict the coexistence of lizardite and
473 antigorite during subduction between about 250-350°C (Evans, 2004; Schwartz et al., 2013). Indeed,
474 based on combined δD and $\delta^{18}O$ values of serpentine from Conical ultramafic clasts, Alt and Shanks
475 (2006) propose that the serpentinizing fluids released by the subducting slab at depth are likely to
476 increase with temperature and to be progressively enriched in ^{18}O ($\delta^{18}O_{\text{fluid}} \sim 5.5\%$ at 250°C and \sim
477 9.0% at 400°C; Fig. 12). Similarly, Sakai et al. (1990) also suggest a slab-derived fluid with a $\delta^{18}O$
478 value of approximately 3.0% based on the oxygen and hydrogen isotope composition of Izu and
479 Mariana forearc serpentinite clasts. In agreement with those studies, we observed a progressive
480 increase of $\delta^{18}O$ values of pore fluids with increasing distance from the trench (Fig. 12). However,
481 pore fluid measurements from Yinazao, Fantangisña or Asùt Tesoru (-1.39 to $+1.97 \%$) are always
482 lower than $\delta^{18}O$ estimates of slab-derived fluids suggesting that the $\delta^{18}O$ values of pore fluids are
483 driven toward to lower values during serpentinization processes occurring within the forearc (Alt and
484 Shanks, 2006). In agreement with this scenario, the $\delta^{18}O_{\text{Srp}}$ analyses are more or less constant and
485 enriched in ^{18}O (5.8 - 10.8% ; e.g., Alt and Shanks, 2006; Kahl et al., 2015; Sakai et al., 1990; this
486 study) relative to pore fluids analyses, regardless of the slab depth. Higher $\delta^{18}O$ values for pore fluids
487 than those recorded here have been reported by Mottl et al. (2003) for the more distant South
488 Chamorro ($\delta^{18}O_{\text{fluid}} = 2.5 \pm 0.5 \%$) and Conical ($\delta^{18}O_{\text{fluid}} = 4 \pm 0.5 \%$) mud volcanoes (Fig. 12). If
489 these higher pore fluid $\delta^{18}O$ values are used, temperature estimates are somewhat higher at 243-283°C
490 for serpentine crystallization in Atg/Liz-serpentinites and therefore in better agreement with previous
491 estimates of the transition lizardite to antigorite in subduction zones (250-350°C; Schwartz et al.,
492 2013).

493 The observation of euhedral magnetite embedded in brucite in Atg-serpentinites (Figs 6b, 7b and 8b)
494 suggests equilibrium with brucite, which in turn formed in equilibrium with antigorite. We therefore
495 use the $\delta^{18}O$ values of serpentine–magnetite pairs to estimates the temperature of antigorite
496 crystallization (thermometer of Wenner & Taylor, 1971 revised by Früh-Green et al., 1996) as 322 to

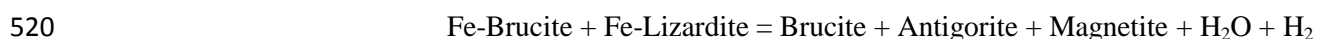
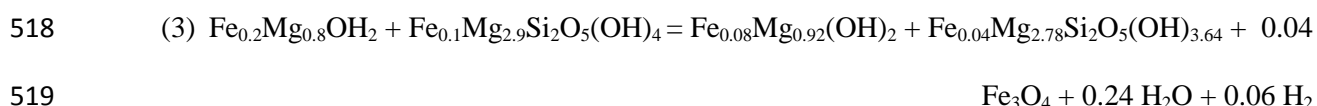
497 409°C. These temperatures overlap or are slightly higher compared to previous estimates of 300–
 498 375°C from Alt & Shanks (2006), and suggest that the transition of lizardite to antigorite in the forearc
 499 mantle likely occurs in a temperature range of 200 to 320°C. These estimates are also in accordance
 500 with thermodynamic calculations by Evans (2004) predicting the assemblage antigorite + brucite to be
 501 more stable than lizardite and chrysotile at temperatures > 300°C.

502 The transition from lizardite to antigorite can follow different reactions (e.g., Evans, 2004; Vils et al.,
 503 2011). For example, in a water saturated open system, the transition can be written in a MASH system
 504 as (Evans, 2004):



507 whereby the required influx of SiO₂ could be generated by sediment dehydration (e.g. Deschamps et
 508 al., 2011; Schwartz et al., 2013). However, the absence of a SiO₂-rich phase (e.g. talc, diopside) and
 509 the high amounts of brucite (Figs 7 and 8) rather suggest a system with low SiO₂ activity. The high
 510 amounts of brucite formed during mantle wedge serpentinization are also in line with former studies of
 511 the Mariana forearc (e.g., D'Antonio & Kristensen, 2004; Murata et al., 2009) and mantle wedge
 512 relicts from the Sanbagawa Belt (southwest Japan, Kawahara et al., 2016). These observations suggest
 513 a limited transfer of SiO₂ during slab dehydration at shallow depth.

514 In our samples, the antigorite-forming alteration stage is accompanied by a decrease of FeO in
 515 serpentine and brucite minerals suggesting a redistribution of Fe between Fe-bearing minerals. In
 516 order to account for the production of brucite and magnetite during the recrystallization of lizardite
 517 into antigorite without the addition of SiO₂ by fluids, we propose the following equation:



521 with the XMg of antigorite and brucite derived from microprobe analyses. It should be noted that this
522 equation does not take into account the potential incorporation of Fe³⁺ into serpentine minerals
523 (Andreani et al., 2013; Debret et al., 2014), the potential mobility of Fe in slab derived fluids (Debret
524 et al., 2016), nor the role of other redox sensitive elements (e.g., C or S). Hence the production of H₂
525 in the equation is speculative.

526 The studied samples were partly affected by late serpentinization stages characterised by the
527 crystallization of blue serpentine and sulfides (Fig. 3). If present, this late serpentinization stage
528 largely replaces former textures, i.e. serpentine or mantle minerals (see Appendix A). This alteration
529 texture is highly developed in Yinazao, whereas it is limited to the uppermost units of Fantangisña and
530 Asùt Tesoru mud volcanoes suggesting that the crystallization of blue serpentine mainly occurs as a
531 late stage of serpentinization. Large amounts of sulfides in some of these samples may indicate
532 ongoing reduction of sulfates to sulfides through the activity of microbial communities (e.g., Mottl et
533 al., 2003) during this stage. These sulfides display framboidal textures that are commonly interpreted
534 as microbe-derived textures formed during bacterially mediated sulfate reduction (e.g., Thiel et al.,
535 1999; Wilkin and Barnes, 1997). However, based on δ¹⁸O analyses, the temperature estimates of blue
536 serpentine crystallization vary between 183 and 194°C (Table 1), higher than those considered
537 feasible for life (~ 122°C; Kashefi and Lovley, 2003). It should be noted that these temperature
538 estimates are maxima due to brucite/serpentine intergrowths in serpentine textures in the blue-
539 serpentinites, as discussed above. It is thus possible that the blue serpentinites formed at temperatures
540 lower than 122°C. In addition, recent investigations of Mariana ultramafic clasts show the possible
541 existence of microbial ecosystems within or below the Mariana mud volcanoes (Plümper et al., 2017).
542 Microbial activity may therefore have taken place at a similar time to the crystallization of blue-
543 serpentine in these sub-surface environments.

544 The temperatures of the slab / forearc mantle interface have previously been estimated to be ~80°C
545 below Yinazao and to be ~250°C below Asùt Tesoru (Hulme et al., 2010), in both cases using pore
546 water chemistry. These estimates are significantly lower than ours derived from oxygen isotope
547 thermometry (up to about 400°C) and incompatible with antigorite crystallization in a subduction

548 setting (Evans, 2004; Schwartz et al., 2013). Although recent studies have reported the crystallization
549 of antigorite in an oceanic setting at relatively low temperatures (e.g., Rouméjon et al., 2014), this has
550 been attributed to Si-metasomatism, possibly following pyroxene serpentinization, and is not
551 compatible with SiO₂ undersaturated systems. The occurrence of antigorite in our samples rather
552 suggests a progressive increase in temperature leading to the formation of antigorite, magnetite and
553 Fe-poor brucite at the expense of lizardite. This scenario is compatible with a progressive burial of the
554 forearc mantle wedge during subduction.

555 The deep burial of forearc rocks can be due to either corner flow enhanced by the low viscosity of
556 serpentinite (e.g. Nagaya et al., 2016) or to frictional stresses mechanically disaggregating the slab
557 surface and eroding the mantle wedge above the décollement zone, incorporating serpentinized mantle
558 into the aggregated subducting inventory (e.g., King et al., 2006). Both processes could potentially
559 explain a prograde metamorphic path and the coexistence of ultramafic clasts displaying various
560 mineralogical assemblages equilibrated at various temperatures in each of the mud volcanoes. The
561 ultimate mechanism responsible for carrying the clasts to depth remains unclear and requires a detailed
562 textural investigation, which is beyond the scope of this study. However, retrograde processes leading
563 to the overprinting of high-temperature serpentine phases by lower-temperature chrysotile (\pm lizardite)
564 such as those observed here are interpreted to occur during the rise of the clasts towards shallower
565 levels in the forearc, e.g., in the mud volcano conduits. The recovered clasts display a high degree of
566 serpentinization (most of the clasts are serpentinized to almost 100%) compared to geophysical data
567 based on seismic velocities which suggest serpentinization \geq 30 % in the forearc (Reynard, 2013).
568 These conflicting observations can be reconciled if only low density material (i.e., highly
569 serpentinized parts of the forearc) can be exhumed, probably by buoyancy (e.g. Guillot et al., 2000),
570 during subduction.

571 Similar observations of prograde and retrograde metamorphism in the Mariana forearc peridotites
572 were described by Murata et al. (2009), who recognized lower-temperature chrysotile veins in
573 antigorite-rich clasts that both pre- and post-date high-temperature antigorite growth. The authors
574 concluded that this reflected a complex process of tectonic cycling of shallow mantle wedge

575 peridotites to depth and then back again to the surface. Several other studies have also noted the
576 possibility that serpentinite formed in the shallower parts of the subduction zone may be carried
577 deeper into the subduction zone (e.g., Kawahara et al., 2016; Kerrick & Connolly, 2001; Savov et al.,
578 2005, 2007; Snyder et al., 2005; Tamblyn et al., 2018). This suggests that the serpentinitized forearc
579 mantle wedge can significantly contribute to arc magmas isotope and elemental budget during
580 subduction (e.g., Ribeiro and Lee, 2017; Debret et al., 2018a).

581 **7. Evolution of forearc mantle wedge composition during subduction**

582 Mariana forearc mud volcanoes are formed by the interaction of slab-derived fluids and forearc
583 peridotites. The nature of the fluids released during slab dehydration is expect to change considerably
584 with increasing depth and associated increase of pressure and temperature at depth (Bebout, 2013 and
585 reference therein). Previous analyses of pore water chemistry of Mariana ultramafic clasts have shown
586 an increase in K, sulfate, carbonate alkalinity, Na/Cl, B, Mn, Fe, Co, Rb, Cs, Gd/Tb, Eu, and LREE
587 and a decrease in Ca, Sr, and Y with depth to the slab/mantle interface (Fryer et al., 2018; Hulme et
588 al., 2010; Mottl et al., 2003). In agreement with these observations, we observed an increase of $\delta^{18}\text{O}$
589 values of pore fluids from Yinazao to Asùt Tesoru and the deeper-sources South Chamorro and
590 Conical mud volcanoes (Fig. 12). However, the bulk-rock major and trace element compositions of
591 serpentinitized clasts recovered from Yinazao, Fantangisña, Asùt Tesoru, South Chamorro and Conical
592 mud volcanoes largely overlap and therefore do not reflect these strong variations (Fig. 9 and 11). This
593 is in good agreement with petrographic observations showing the existence of up and down
594 movements beneath serpentinite mud volcanoes (e.g. Kawahara et al., 2016 or this study). We do
595 however observe modifications of clasts chemistry according to mineralogy, i.e., Blue-, Liz-, Atg/Liz-
596 and Atg- serpentinites.

597 The first step of serpentinitization of the forearc mantle corresponds to the formation of brown
598 serpentine (Liz-serpentinite), a mixture of chrysotile/lizardite with Fe-rich brucite, at the expense of
599 mantle minerals (Fig. 13). As previously documented (e.g., Kahl et al., 2015; Peters et al., 2017;
600 Savov et al., 2005, 2007), lizardite-dominated samples are characterized by high concentrations of
601 fluid mobile elements (FME), such as B, Li, Cs, As, Sb, relative to primitive mantle and/or abyssal

602 serpentinite/peridotite (Fig. 11). These enrichments have been attributed to the influx of slab-derived
603 fluids in the forearc mantle at intermediate temperatures (200-500°C). The onset of subduction is
604 accompanied by large amounts of compaction, deformation and metamorphic reactions (e.g. clay
605 mineral or carbonate breakdown) in the slab resulting in the release of FME-rich fluids (e.g., Barnes et
606 al., 2014; Bebout, 2013; Cannà et al., 2015; Debret et al., 2013, 2018a; Hattori & Guillot, 2007). The
607 transfer of such fluids to the overlying fore-arc mantle wedge allows its serpentinization and the
608 storage of FME in forearc serpentinites.

609 The progressive burial of ultramafic material is accompanied by the recrystallization of Liz-
610 serpentinites to antigorite, i.e., Atg/Liz- and Atg-serpentinites (Fig. 13). During this transformation,
611 Cs, Li and Sr concentrations in serpentinites progressively decrease. Indeed, although a relatively
612 small amount of water is released during the transition of lizardite to antigorite (e.g. Evans, 2004), this
613 reaction can be accompanied with dissolution and leaching of C- and/or S- bearing phases (e.g. Debret
614 et al., 2014) and FME released in fluids (e.g., Debret et al., 2013; Kodolányi & Pettke, 2011; Vils et
615 al., 2011) suggesting that these elements are highly mobile in fluids during forearc burial. Boron
616 concentrations remain high in Atg/Liz- and Atg-serpentinites relative to slab serpentinites (e.g., Vils et
617 al., 2011; Debret et al., 2013) and/or abyssal peridotites/serpentinites (Andreani et al., 2009). This
618 suggests that antigorite also crystallizes in equilibrium with B-rich slab derived fluids. *In situ* analyses
619 of partly serpentinized forearc peridotites reveal that, in low temperature lizardite bearing serpentinites
620 (< 200°C), serpentine textures integrate high amounts of B (up to 200 ppm, Pabst et al, 2011; Kahl et
621 al., 2014). These values are close to those reported in antigorite bearing textures in mantle wedge
622 settings (e.g., Deschamps et al., 2010). Hence, the absence of correlation between B concentrations
623 and indices of prograde metamorphism suggests that, during the serpentinization of the forearc,
624 saturation (or exchange equilibrium – e.g., Pabst et al., 2011) is rapidly reached in the serpentine, i.e.,
625 before the transition from lizardite to antigorite. In agreement with these observations, sediment pore
626 water chemistry reveals progressive enrichments in Cs, Li and B concentrations with increasing
627 distance to the trench (Fryer et al., 2018; Hulme et al., 2010; Mottl et al., 2003) confirming that these
628 elements become progressively abundant in fluids during the progressive burial of the forearc (e.g., De

629 Hoog and Savov, 2018). In contrast to other fluid mobile elements, the concentrations of As, Sb and
630 LREE progressively increase in whole rocks during the transition of lizardite to antigorite (Figs 9 and
631 10). Previous studies have shown that these elements can be incorporated in antigorite (e.g., Hattori et
632 al., 2005). Fluids released during the early (shallow) stage of slab devolatilization are likely dominated
633 by diagenesis and opal dehydration, whereas later (deeper) processes included decarbonation and clay
634 mineral decomposition resulting in a modification of slab derived fluid composition (Bebout, 2013 and
635 reference therein). As, Sb and the LREE are enriched in fluids derived from sediment decarbonation
636 (Bebout, 2013; Debret et al., 2018a). Hence, we interpret the high As, Sb and LREE concentrations of
637 the antigorite dominated serpentinites in terms of the onset of carbonate dissolution within the slab
638 (e.g., carbonated altered oceanic crust). This scenario is in good agreement with high dissolved
639 inorganic carbon concentrations of pore fluids from the furthest mud volcanoes (Asùt Tesoru, South
640 Chamorro and Conical; Fryer et al., 2018).

641 During their exhumation forearc serpentinites are partly to fully recrystallized into blue serpentine.
642 This episode is accompanied with a flattening of REE patterns (Fig. 10a – Group 1), probably linked
643 to a high mobility of LREE in fluids (e.g., Fryer et al., 2018), and an increase in Zn concentrations
644 (Fig. 11c). This observation suggests that the composition serpentinites recorded at depth can be partly
645 overprinted by low-temperature reactions (< 250°C) during clast exhumation. However, the exact
646 source of the added Zn is unclear as Zn can be mobile in fluids either at depth during serpentinite
647 devolatilization (Pons et al., 2016), decarbonation processes (Debret et al., 2018a; Inglis et al., 2017)
648 as well as near the seafloor through hydrothermal fluid circulation (Debret et al., 2018b) and microbial
649 activity via sulfate reduction (Kelley et al., 2009).

650 **8. Conclusions**

651 Our study reveals that the ultramafic clasts recovered in serpentinite mud volcanoes record three main
652 stages of serpentinitization (Fig. 13): the crystallization of brown serpentine bearing textures composed
653 of a mixture of Fe-rich brucite and chrysotile and/or lizardite (stage 1 in Fig. 13); the formation of
654 antigorite in equilibrium with magnetite and Fe-poor brucite at the expense of brown serpentine
655 (stages 2 to 3 in Fig. 13); the late formation of blue serpentine associated with frambroidal sulfides

656 during the exhumation of the ultramafic clasts (stage 4 in Fig. 13). The transition of lizardite to
657 antigorite is enhanced by an increase of temperature from 200°C up to about 400°C within the forearc
658 area. These estimates are in good agreement with thermodynamic calculations carried out by Evans
659 (2004). The crystallization of antigorite at the expense of lizardite has been observed in different mud-
660 volcanoes. Although most of the blue serpentine was observed at Yinazao, there is no evidence for a
661 systematic serpentine phase change according to depth of the slab/mantle interface. These observations
662 suggest the existence of complex transport mechanisms below the mud volcanoes, with the
663 serpentinites being progressively dragged down to greater depth before their exhumation, potentially
664 controlled by buoyancy, toward to the surface. In agreement with this scenario, no obvious changes of
665 serpentinite clasts chemistry is observed according to depth to the slab/mantle interface (Fig. 9-11). In
666 contrast, the crystallization of lizardite and then antigorite in serpentinites is accompanied with a
667 decrease of Cs, Li, Sr and an increase in As, Sb and LREE concentrations in whole rock, whereas B
668 concentrations are relatively constant. This suggests that the serpentinitized mantle wedge acts as a filter
669 for trace elements and controls the fluxes of these elements between the surface and the deep mantle.
670 The fluid release during shallow metamorphic reactions (i.e., transition lizardite to antigorite) are
671 likely to feed hydrothermal circulation near the surface. In contrast, the dragged down residue (Atg-
672 serpentinites) will undergo a progressive increase in temperature, coupled to the ongoing burial, until
673 its dehydration at greater depth. This process will release large amounts of fluid, thereby contributing
674 to arc magmas genesis in subduction zones.

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990

991 **Figure Captions**

992 Fig. 1: Bathymetry map of the Mariana subduction system showing the locations of Yinazao,
993 Fantangisña and Asùt Tesoru drilled during IODP Exp. 366 as well as Conical and South Chamorro
994 that were drilled during previous ODP legs.

995 Fig. 2: Representative core images of the main lithostratigraphic units observed at Yinazao,
996 Fantangisña and Asùt Tesoru flanks or summits. (a) Core image of the upper level unit recovered at
997 the site U1492A (Yinazao summit, section 1H2-99/139cm). The uppermost units of the mud-volcano
998 is made of red-brown pelagic mud containing carbonated serpentinite breccia and weathered
999 serpentinites clasts (brown weathering). (b) Core image of the contact between upper and lower level
1000 units at the Site U1496A (Asùt Tesoru summit, section 1F1-1/41cm). The uppermost units of the mud-
1001 volcano is made of red-brown pelagic mud containing partly weathered serpentinite clasts. Note that
1002 the clast displays a carbonated and brown weathering crust toward to the core. The clast itself is also
1003 rimmed by a halo of Blue-serpentinite while the core correspond to a Liz-serpentinite. The square
1004 indicates the localization of the macroscopic picture and thin section observations presented in
1005 Appendix A. The lower unit is made of blue-grey serpentinite mud containing serpentinite clasts of
1006 various size. (c) Core image of the lower level unit recovered at the Site 1492A (Yinazao summit,
1007 section 4F2-50/90cm). The lower unit is made of blue-grey serpentinite mud containing serpentinite
1008 clasts of various size. (d) Core image of a large clast of 40 cm long recovered at the Site 1498B
1009 (Fantangisña flank, section 3R3, 38-78cm).

1010 Fig. 3: Photomicrographs (a: crossed polarized light; b and c: plane polarized light), back scattered
1011 electron image and Raman spectra of the carbonated (a) and blue (b-d) serpentinites recovered at the
1012 Yinazao mud volcano (photo taken by the Shipboard Scientists, 2018). (a) Carbonated breccia
1013 (shipboard sample U1491B-2H5-85/88) showing dismembered crack-seal like veins (mainly made of
1014 chrysotile, Ctl) embedded into a calcite (Cal) matrix. (b) Mesh textures with a pale blue colour
1015 (sample M7) and corresponding Raman spectra. The meshes centres are replaced by opaque
1016 aggregates composed of sulfide (pyrite) and brucite \pm serpentine. Raman spectra of mesh rim (black
1017 line) and core (grey line) are mixtures between serpentine and brucite. In the high frequency region,

1018 the core of the mesh is mainly dominated by brucite. White square: localisation of the Fig. 2s. (c)
1019 Opaque vein made of pyrite crossing mesh and bastite textures (shipboard sample U1492C-8F3-
1020 108/112). The centres of the mesh and bastite textures are associated with sulfides. (d) SEM
1021 observation of a mesh core (sample M7). The core is composed of brucite lamellae with interstitial
1022 framboidal pyrite.

1023 Fig. 4: Plot of XCr vs. XMg of spinels in forearc ultramafic clasts from Yinazao, Fantangisña and
1024 Asùt Tesoru. Compositions broadly overlap with those of forearc peridotites and are more Fe- and Cr-
1025 rich with respect to abyssal peridotites (abyssal and forearc peridotite fields are from Dubois-Côté et
1026 al. (2005).

1027 Fig. 5: Major element contents and normalized cations of the different serpentine phases. (a-b)
1028 Variations of XMg ($= \text{Mg} / [\text{Fe} + \text{Mg}]$) and FeO with SiO₂. (c-d) Variations of Mg and Fe cations per
1029 formula unit (p.f.u.) with Si + Al cations p.f.u. The decrease of XMg and SiO₂ in serpentine
1030 crystallizing in Blue serpentinites and Liz-serpentinites reflects the presence of Si-free, Fe-rich brucite
1031 at microscale. The brucite trend intercepts SiO₂ = 0 wt% value at #Mg = 0.84. Crystallization of
1032 antigorite (Atg/Liz-serpentinites and Atg-serpentinites) is associated with the disappearance of the
1033 brucite component in serpentine analyses, an increase in SiO₂ and XMg and a decrease in FeO in
1034 serpentine.

1035 Fig. 6: Photomicrographs and corresponding Raman spectra of Liz-serpentinite recovered from Asùt
1036 Tesoru (Photomicrographs taken by the Shipboard Sci. Party, 2018). (a) Serpentine forms mesh
1037 textures with a brownish colour and preserved olivine relicts in their centres (plane polarized light,
1038 sample M19). Note the presence of euhedral and unaltered spinel on the microphotograph bottom.
1039 Mesh Raman spectra are mixed analyses of lizardite and brucite. (b) Antigorite vein with brucite and
1040 magnetite in its centre (crossed polarized light, shipboard sample U1497A-13G-CC-W 61/63). The
1041 vein crosscuts a lizardite/brucite-bearing vein.

1042 Fig. 7: Photomicrographs (crossed polarized light) and corresponding Raman spectra of the Atg/Liz-
1043 serpentinites recovered from Fantangisña and Asùt Tesoru. (a) Antigorite crystallizes as several

1044 hundred microns long lamellae penetrating mesh textures (sample M32). Centres of the antigorite
1045 veins are composed of magnetite and brucite. Note that the centres of the mesh textures show mixed
1046 Raman spectra between lizardite and antigorite. (b) Wide vein of brucite and magnetite crosscutting
1047 relicts of mesh textures (sample M13). The mesh texture is fully recrystallized into pure antigorite
1048 where it is in contact with the vein.

1049 Fig. 8: Photomicrographs of Atg-serpentinites recovered from Asùt Tesoru (both in crossed polarized
1050 light). (a) Antigorite lamellae with interstitial brucite and magnetite (sample M16). (b) Brucite patch
1051 associated with euhedral grains of magnetite (sample M15).

1052 Fig. 9: Bulk rock major element composition of Mariana ultramafic clasts illustrated in (a) Al_2O_3 vs
1053 MgO/SiO_2 and (b) MgO (wt.%) vs FeO (wt.%). South Chamorro (grey crosses) and Conical (black
1054 crosses; data from Geldmacher et al., 2008; Kodolanyi et al., 2011; Parkinson and Pearce, 1998;
1055 Pearce et al., 2000; Savov et al., 2007) are shown for comparison. On Fig. 9a, the dark line represents
1056 the silicate Earth differentiation trend and the primitive mantle ratio (PM; Godard et al., 2008).
1057 Changes in whole-rock ratios of both MgO/SiO_2 and $\text{Al}_2\text{O}_3/\text{SiO}_2$ accompany the transition (left to
1058 right) of depleted (e.g., dunite) to enriched (e.g., lherzolite) peridotites. On Fig. 9b, the dark line
1059 represents the stoichiometric variations of olivine Fe–Mg composition. Abyssal peridotite
1060 endmembers of dunite and lherzolite (Godard et al., 2008) are shown for comparison. Note that several
1061 samples display abnormal high MgO/SiO_2 and MgO contents, such reflect the ultra-refractory
1062 compositions of the ultramafic protoliths and/or the high amount of brucite in the samples.

1063 Fig. 10: Whole-rock trace elemental compositions of the different ultramafic clasts (Blue-, Liz-,
1064 Atg/Liz- and Atg-serpentinites). (a) and (b) patterns are normalized to chondrite and primitive mantle
1065 (PM), respectively, using normalization values from Sun and McDonough (1989). Serpentine mud
1066 analysis is from Kodolanyi et al. (2011). Group-1 correspond to blue serpentinites with flat patterns
1067 similar to that of serpentine muds and Group-2 correspond to blue serpentinites, Liz-, Atg/Liz- and
1068 Atg- serpentinites with U-shaped patterns.

1069 Fig. 11: Plots of Cs/Yb vs Li* (a), As vs Sb (b) and Sr vs Zn (c) of studied ultramafic clasts.
1070 Concentrations overlap well with those of ultramafic clasts from South Chamorro (grey crosses) and
1071 Conical (black crosses; data from Geldmacher et al., 2008; Kodolanyi et al., 2011; Parkinson and
1072 Pearce, 1998; Pearce et al., 2000; Savov et al., 2007, 2005). Abyssal peridotites (white circles; data
1073 from Andreani et al., 2014; Augustin et al., 2008; Boschi et al., 2013; Delacour et al., 2008; Jöns et al.,
1074 2010; Kodolanyi et al., 2011; Niu, 2004; Paulick et al., 2006; Rouméjon et al., 2014) have consistently
1075 lower Cs/Yb and Li* contents with respect to Liz- and Atg/Liz-serpentinites. Sr (~0.4-1000 ppm) and
1076 Zn (~30-200 ppm) concentrations of abyssal peridotites and serpentinites overlap with those of the
1077 forearc but were not presented here for sake of clarity.

1078 Fig. 12: Plot of $\delta^{18}\text{O}$ (‰) variations in ultramafic clast pore fluids vs distance to the trench (km). The
1079 $\delta^{18}\text{O}_{\text{fluid}}$ values progressively increase passing from Yinazao to Fantangisña, Asùt Tesoru, South
1080 Chamorro and Conical. M: data from Mottl et al. (2003); A: calculated $\delta^{18}\text{O}_{\text{fluid}}$ in equilibrium with
1081 antigorite below Conical by Alt and Shanks (2006).

1082 Fig. 13: Conceptual model illustrating serpentinisation processes in relation to fluid circulation and
1083 mantle flow within the Mariana forearc. Numbers in diagram correspond to those in the P-T diagram,
1084 where pressures have been estimated according to slab/mantle interface depth estimates below the mud
1085 volcanoes (Hulme et al., 2010). As clasts from the forearc mantle are dragged down to depth (stages 1
1086 to 3), they undergo an increase in temperature from about 200 to 400°C and the associated
1087 transformation of lizardite into antigorite. During uplift, the clasts are variably retromorphosed into
1088 blue serpentinites (stage 4).

1089

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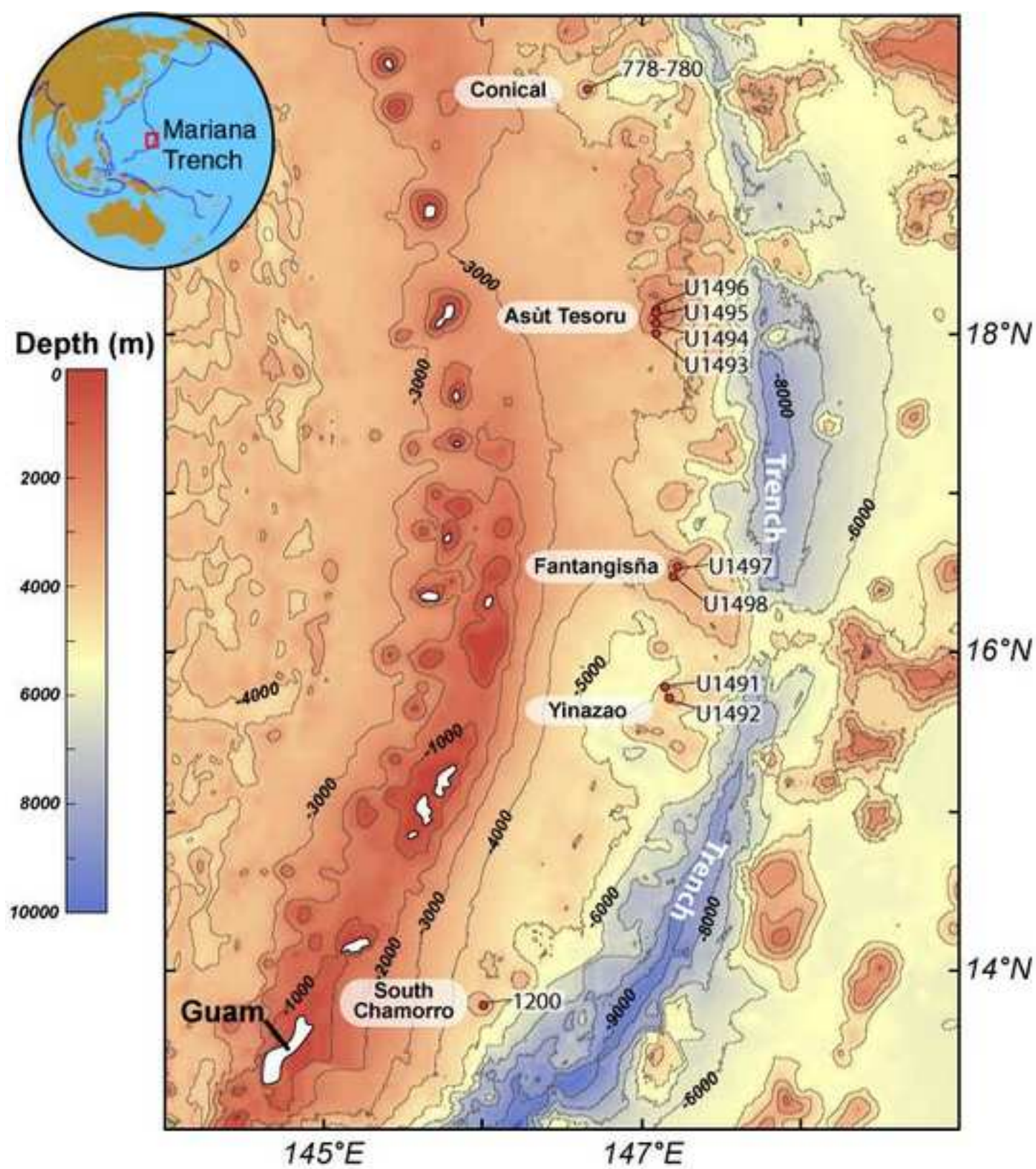


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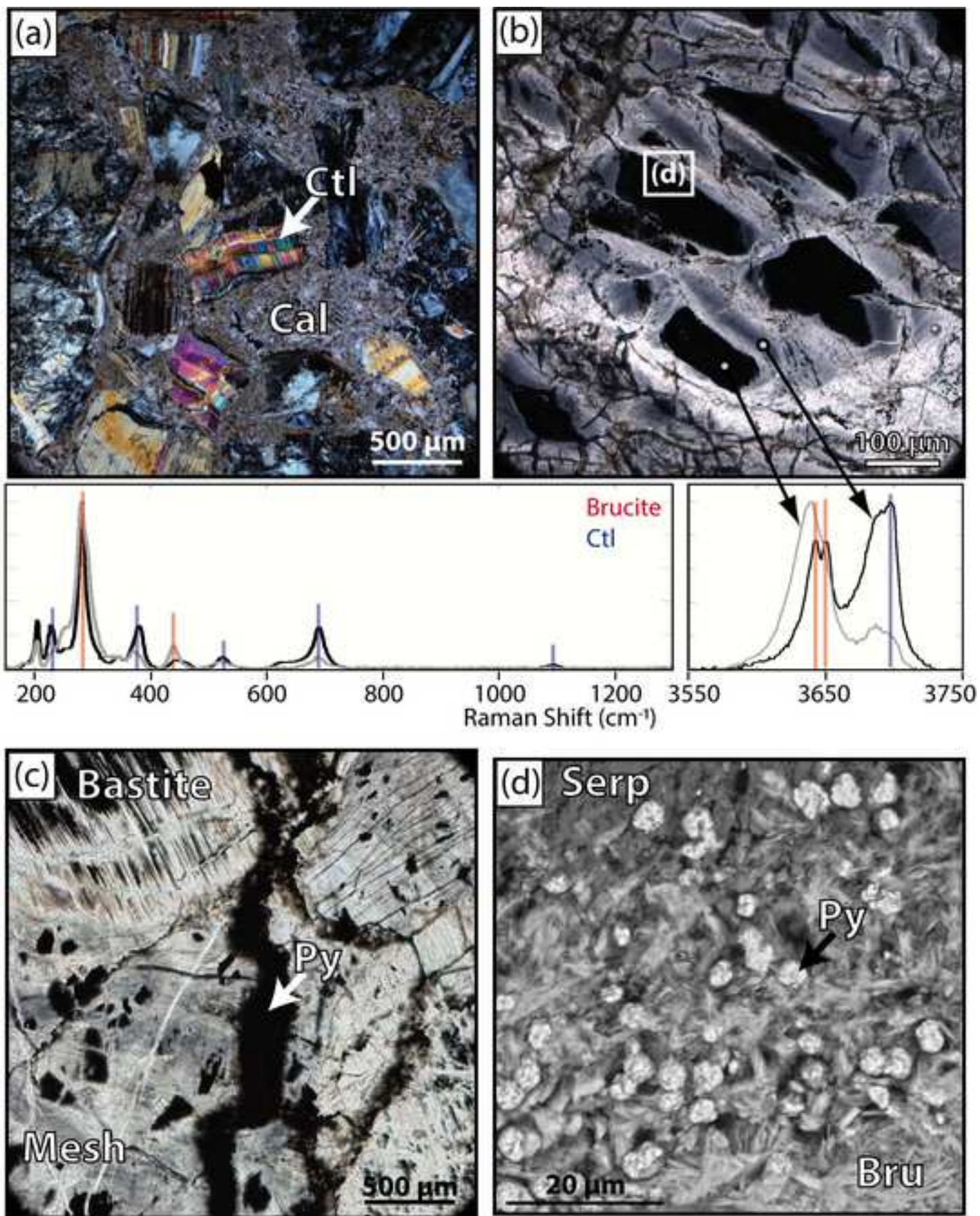


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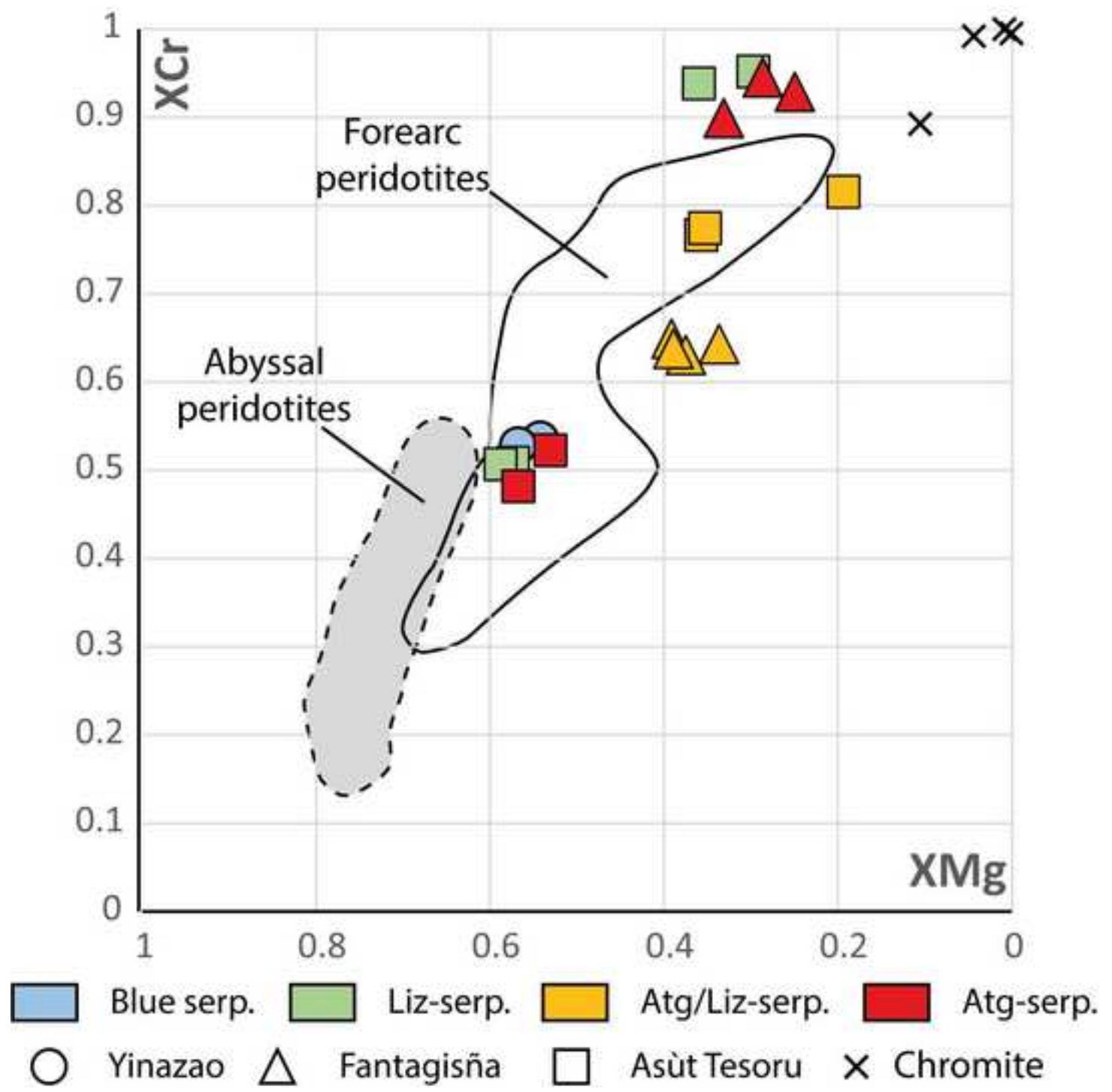


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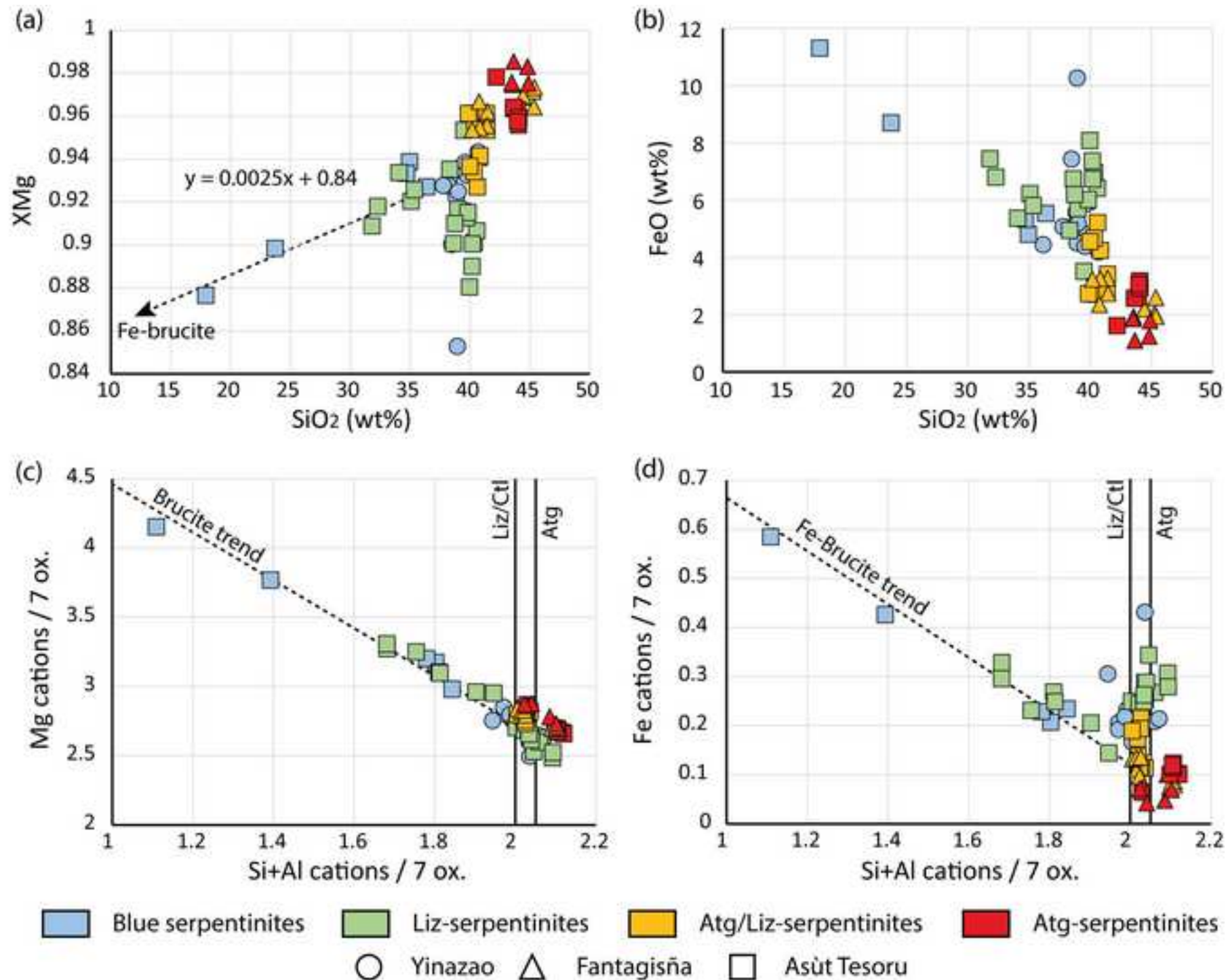


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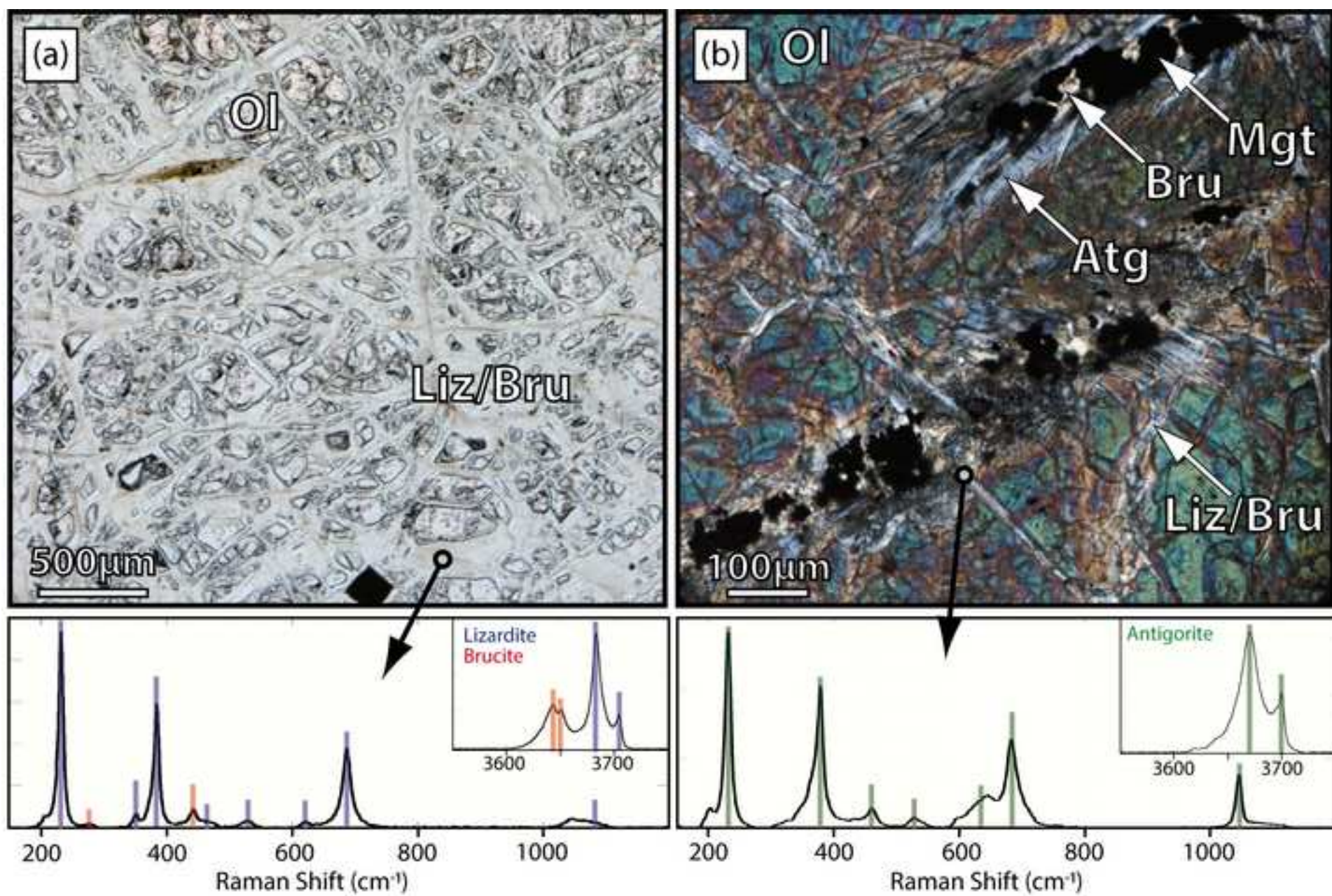


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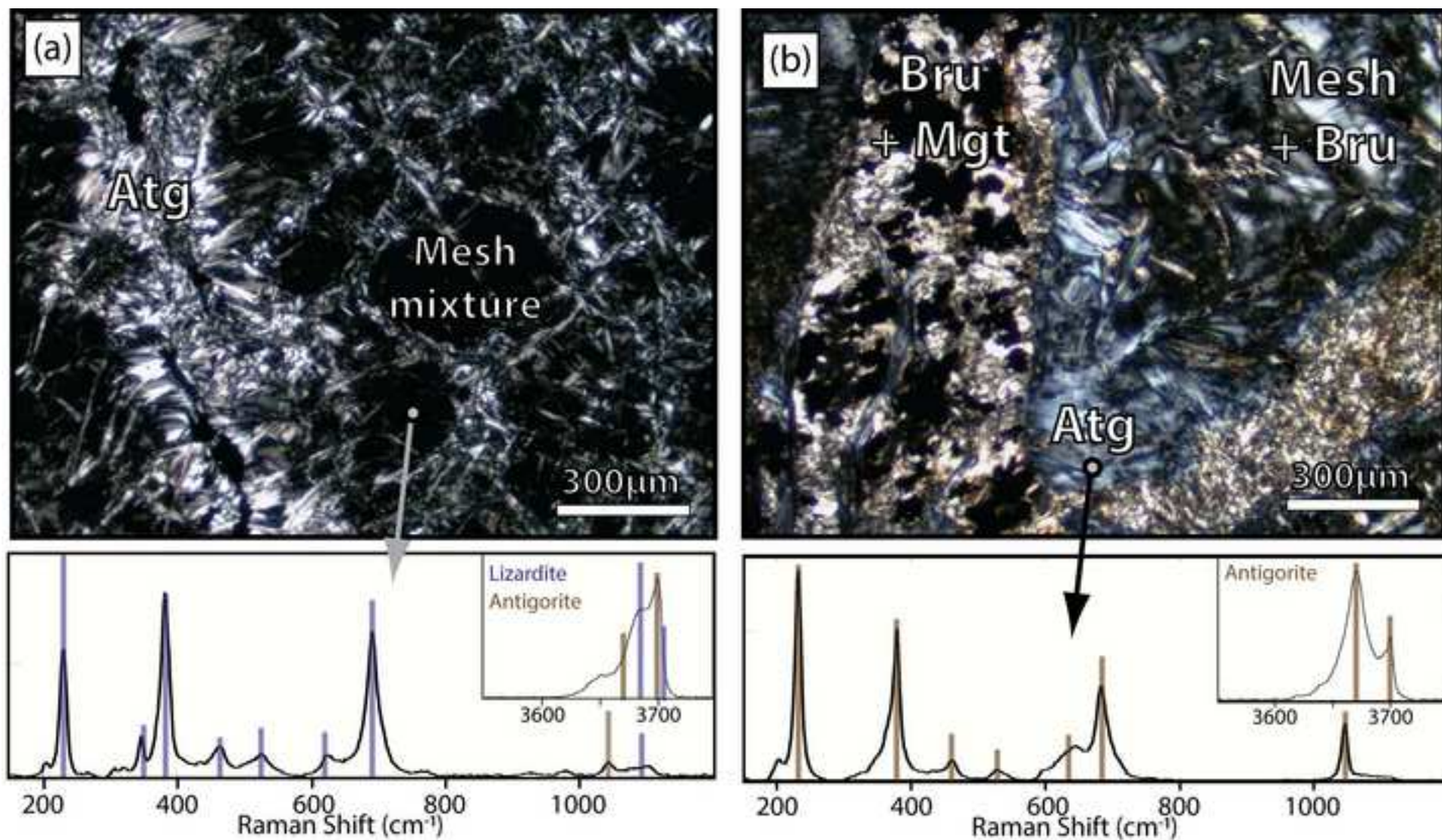


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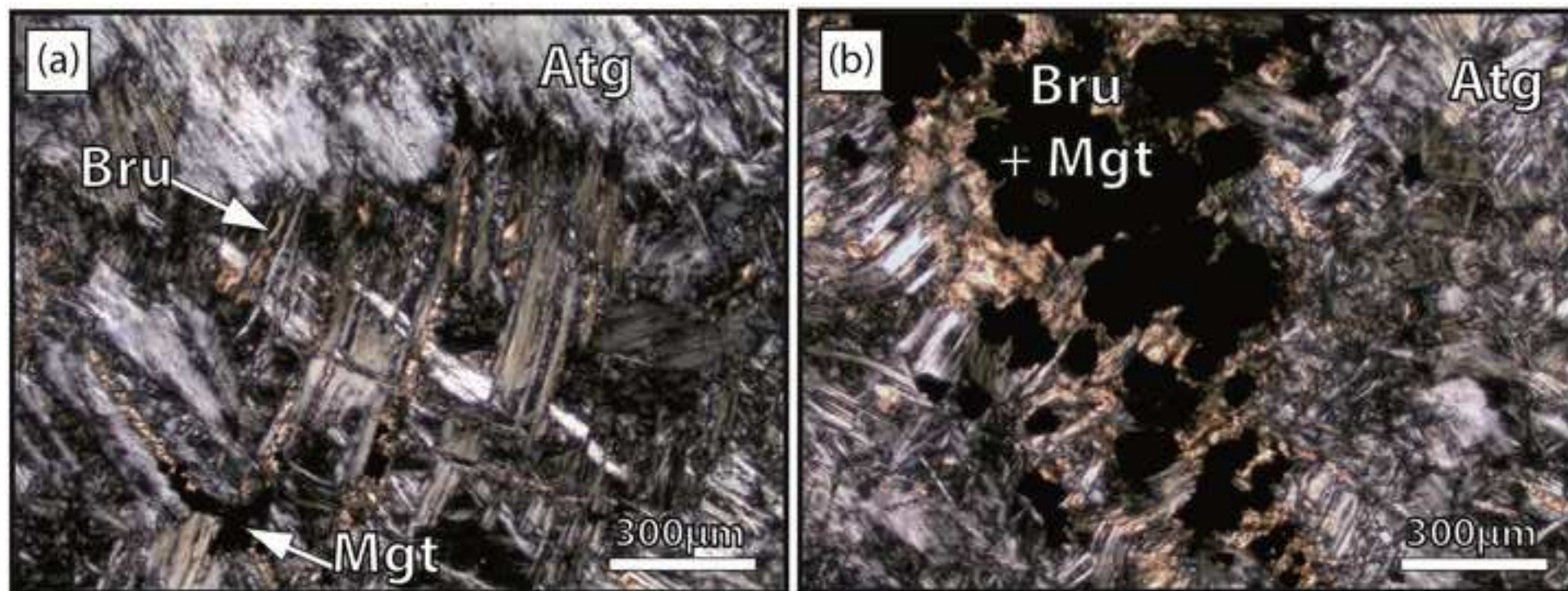


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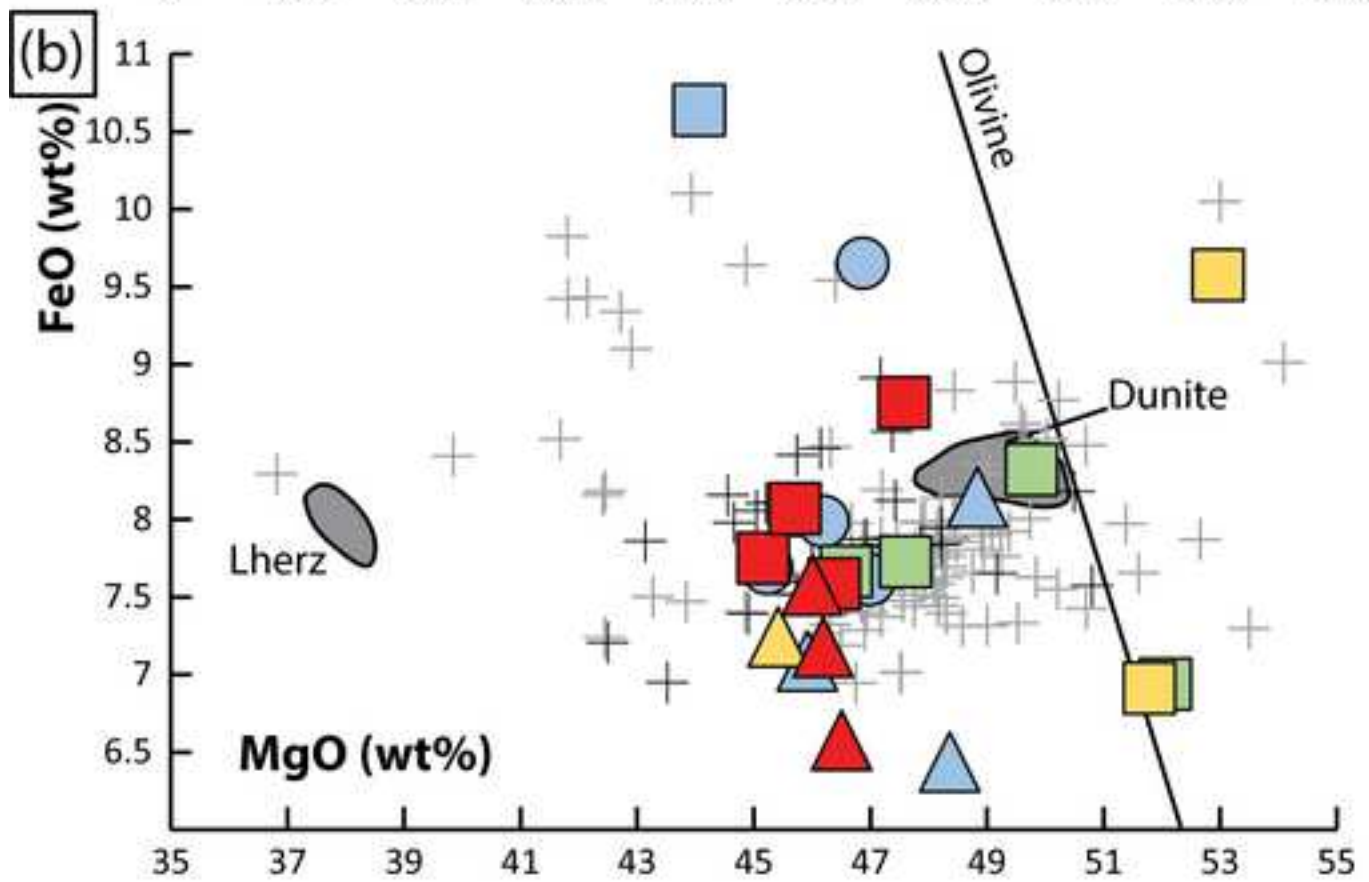
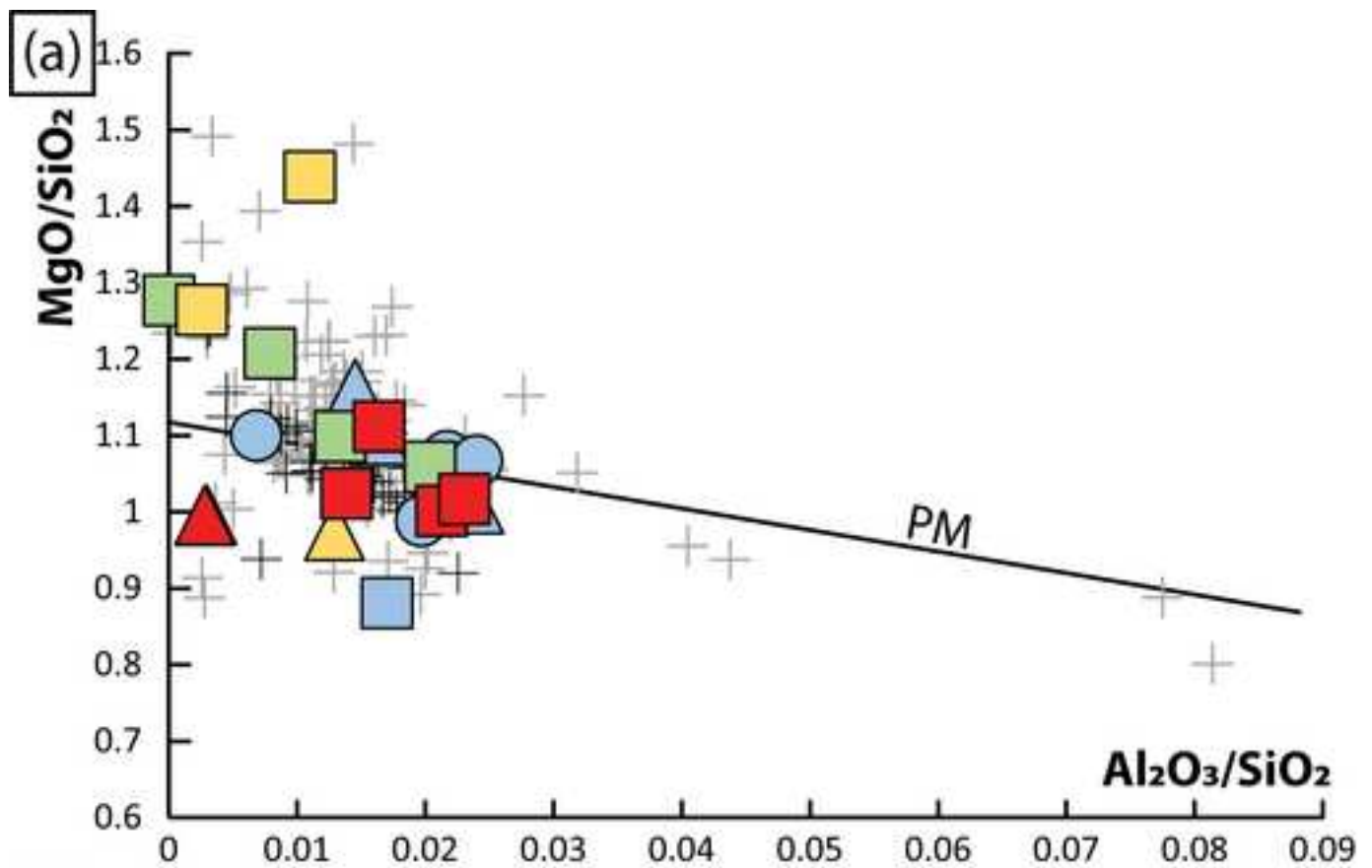


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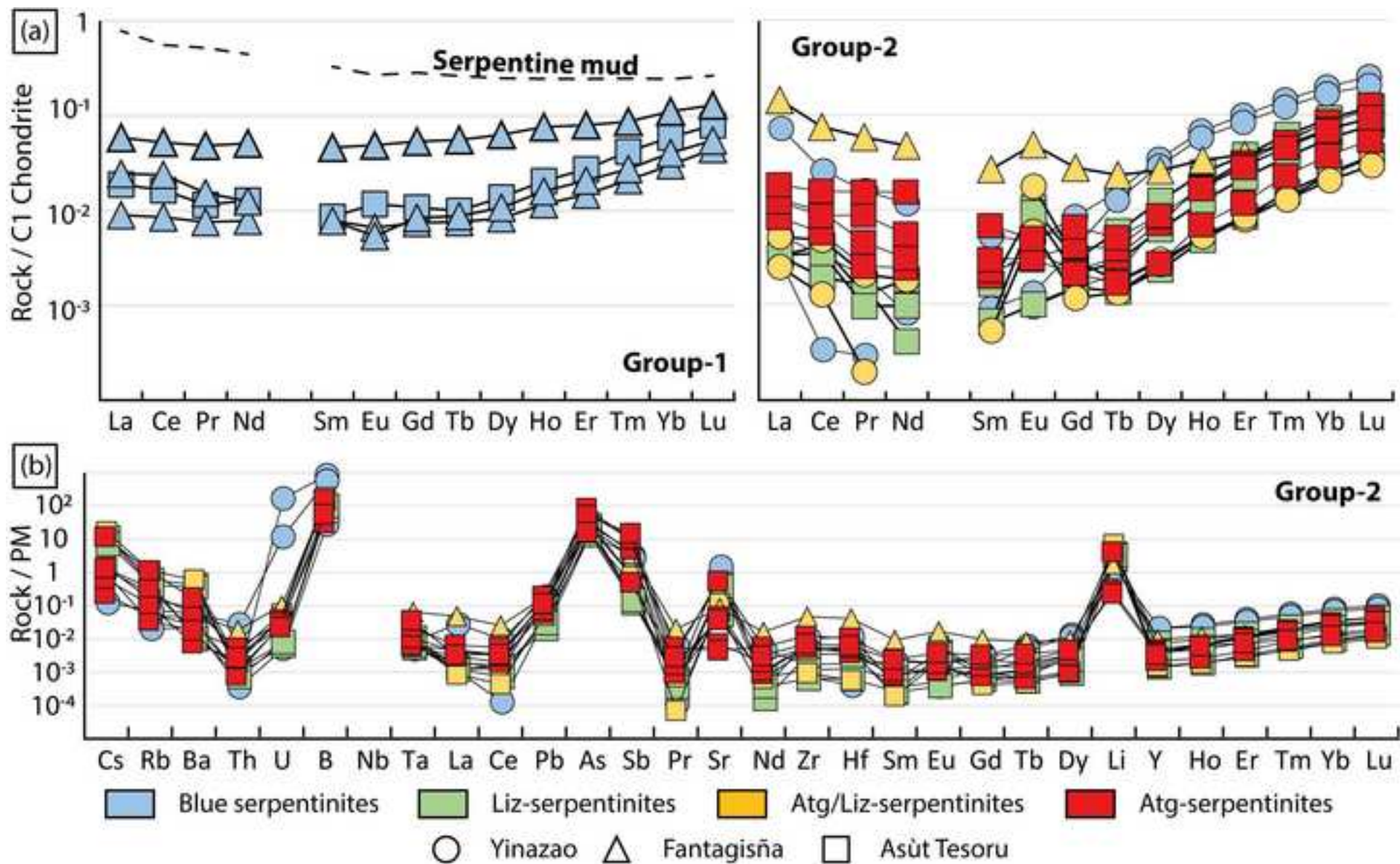


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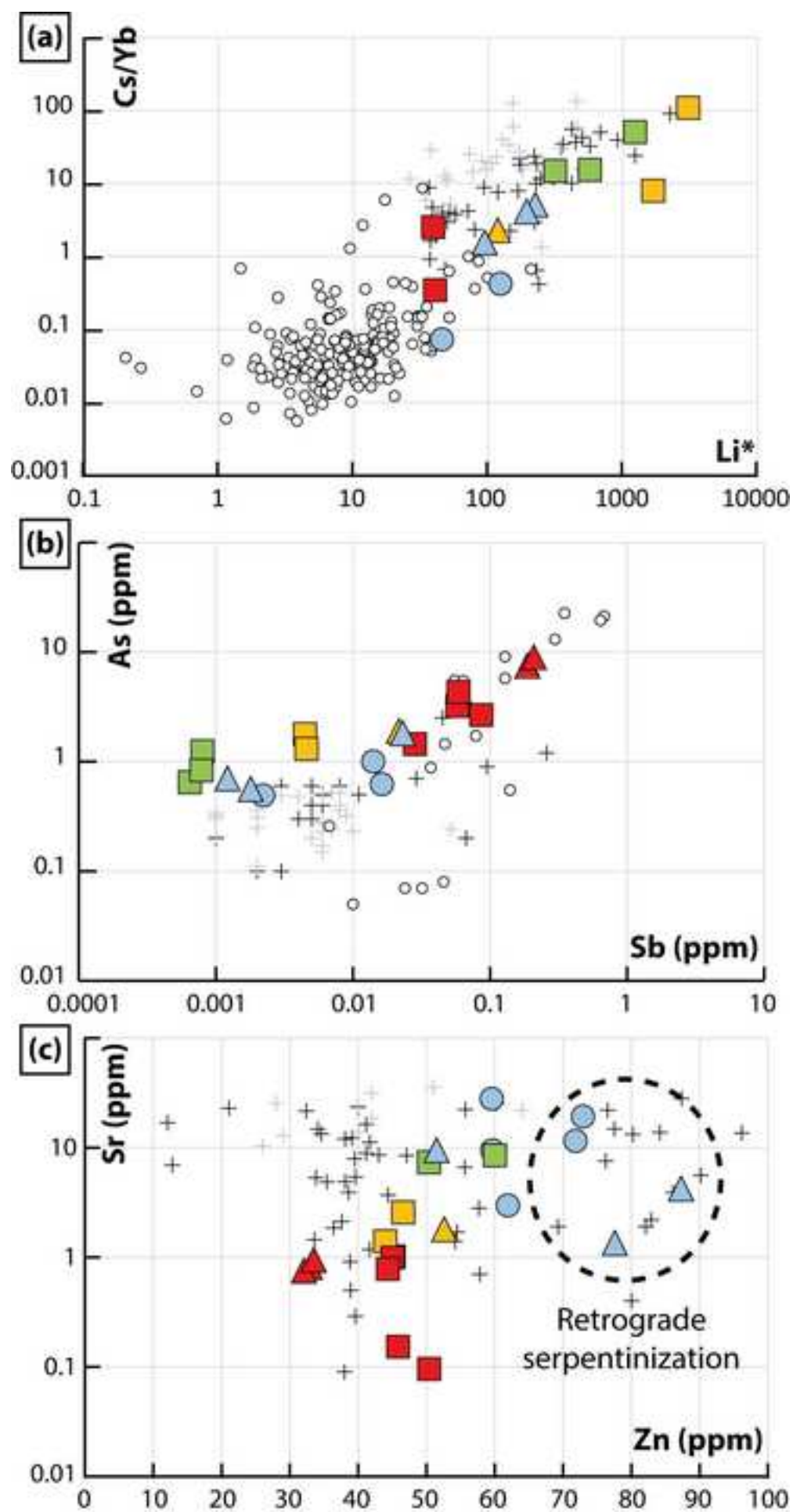


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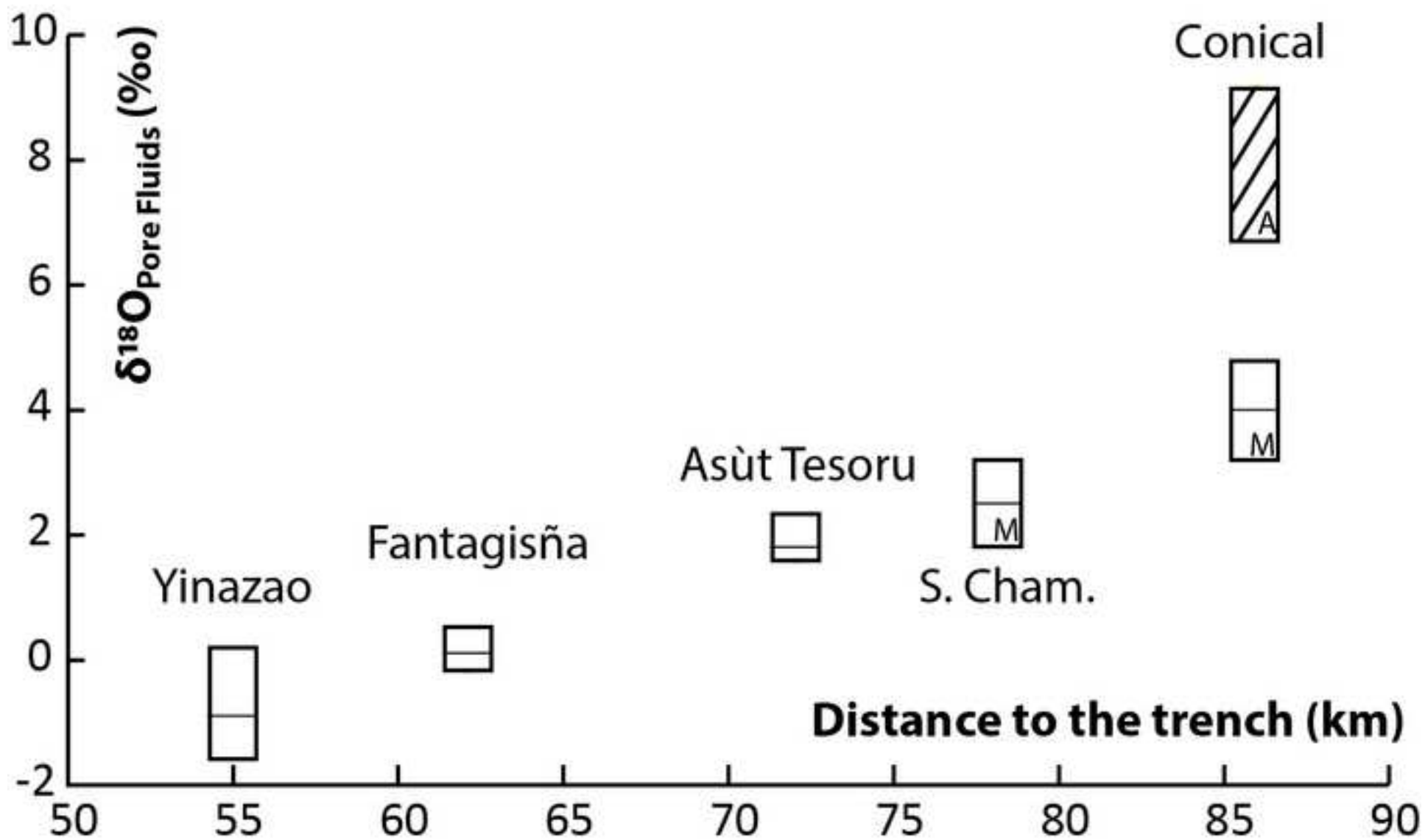


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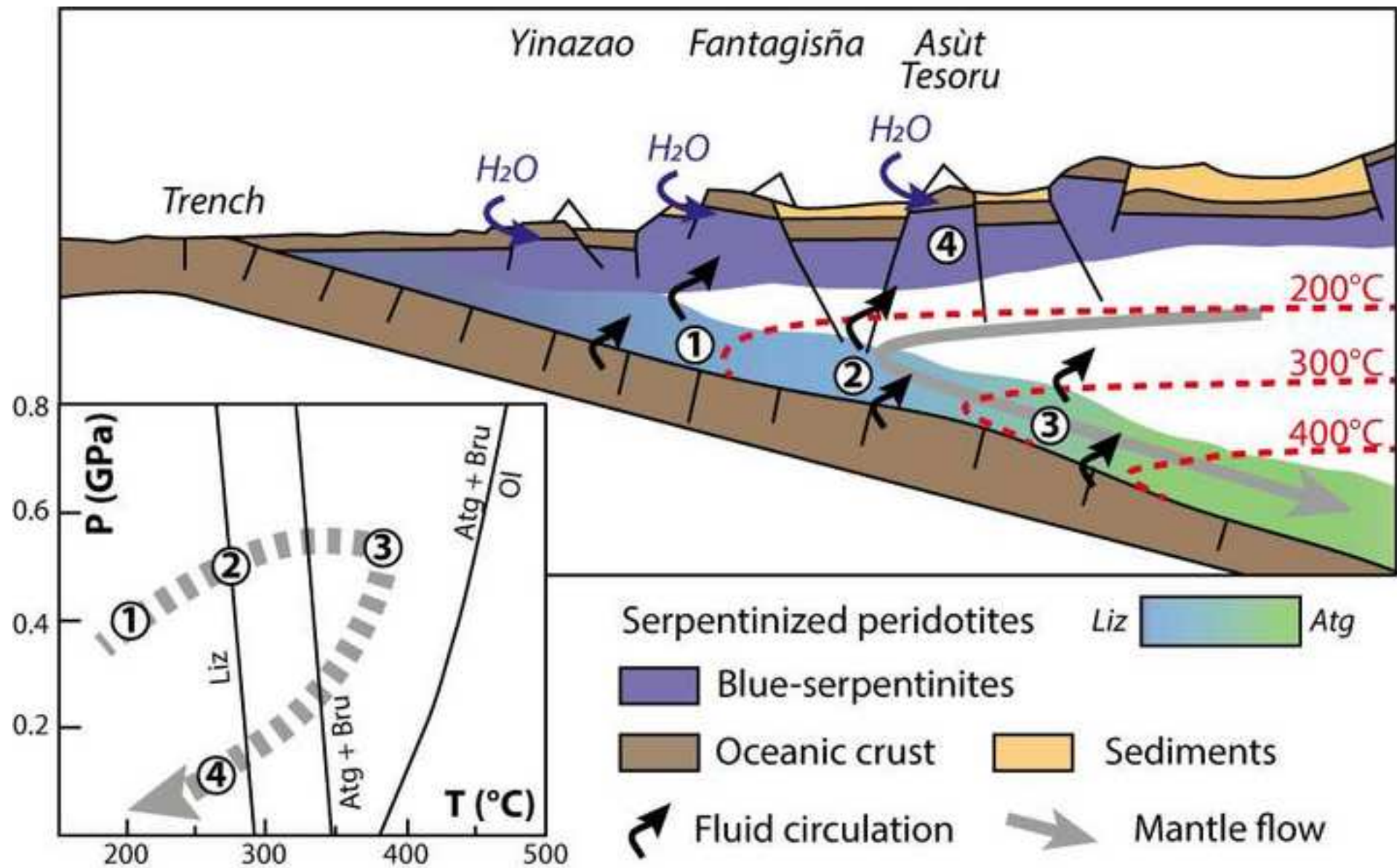


Table 1: Oxygen isotope data and resultant temperature estimates for selected samples. Pore fluid $\delta^{18}\text{O}$ are minimum, maximum and average values.

Reference	$\delta^{18}\text{O}_{\text{Srp}}$ (‰)	$\delta^{18}\text{O}_{\text{fluid}}$ (‰)			T (°C) ¹			$\delta^{18}\text{O}_{\text{fluid}}$ * (‰)	T(°C) ¹	$\delta^{18}\text{O}_{\text{Mgt}}$ (‰)	T(°C) ²
		mean	min	max	mean	min	max				
<i>Blue serpentinite</i>											
M3	6.4	-0.9	-1.4	-0.1	183	176	194	-	-	-	-
M30	7.4 (7.3)	0.1	0.0	0.3	183	181	184	-	-	-	-
<i>Liz-serpentinite</i>											
M9	7.6	1.8	1.7	2.0	205	203	206	-	-	-	-
M10	6.8	1.8	1.7	2.0	220	218	222	-	-	-	-
<i>Atg/Liz-serpentinite</i>											
M12	6.1	1.8	1.7	2.0	232	230	234	4.0	276	-	-
M13	5.8	1.8	1.7	2.0	238	236	240	4.0	283	-	-
<i>Atg-serpentinite</i>											
M14	8.0	-	-	-	-	-	-	-	-	1.8	340
M15	7.1	-	-	-	-	-	-	-	-	1.8	409
M51	8.3	-	-	-	-	-	-	-	-	1.8	322

All $\delta^{18}\text{O}$ values are given in SMOW. (*value*): duplicate; 1: thermometer serpentine/fluid of Saccocia et al (2009); 2: thermometer serpentine/magnetite of Wenner and Taylor (1971) revised by Früh-Green et al. (1996); *fluid value from Mottl et al. (2003).

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