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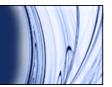
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### Infrared diode laser spectroscopy of the CCO radical: The 2v 1 -v 1 and v 1 +v 3 -v 3 difference bands

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#### Infrared diode laser spectroscopy of the CCO radical: The $2\nu_1 - \nu_1$ and $\nu_1 + \nu_3 - \nu_3$ difference bands

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The  $2\nu_1 - \nu_1$  and  $\nu_1 + \nu_3 - \nu_3$  difference bands of the CCO radical in the gas phase have been studied using a tunable infrared diode laser spectrometer. The CCO radical was produced using a flowing mixture of carbon suboxide and helium subjected to a hollow-cathode discharge. The spectral region between 1920–1960 cm<sup>-1</sup> was probed. Ninety and seventy-five rovibrational transitions were measured in the  $2\nu_1 - \nu_1$  and  $\nu_1 + \nu_3 - \nu_3$  bands, respectively. The analyses of these bands yielded spectroscopic constants for the (001), (200), and (101) vibrational states. The band origins for  $2\nu_1 - \nu_1$  and  $\nu_1 + \nu_3 - \nu_3$  were determined to be 1941.85761(54) cm<sup>-1</sup> and 1936.79402(56) cm<sup>-1</sup>, respectively. © 1998 American Institute of Physics. [S0021-9606(98)01416-0]

#### I. INTRODUCTION

The CCO radical is postulated to act as an intermediate in a variety of chemical reactions<sup>1</sup> and has been detected in interstellar space toward the dark molecular cloud TMC.<sup>2</sup> The first spectroscopic detection of this molecule was made in a matrix isolation study by Jacox et al.<sup>3</sup> who showed that CCO is an asymmetric linear molecule and determined the vibrational frequencies in the ground electronic state to be 381, 1074, and 1978 cm<sup>-1</sup>. Devillers and Ramsay<sup>4</sup> reported the first gas-phase spectrum of CCO, observing the  $\tilde{A}^{3}\Pi_{i}$  $-\tilde{X}^{3}\Sigma^{-}$  electronic transition. From the analysis of the rovibronic spectra, the term energy  $(T_0)$ , vibrational frequencies, rotational constants, and the Renner-Teller coupling constant for the  $\tilde{A}^{3}\Pi_{i}$  electronic state were determined. These earlier spectroscopic studies have been followed more recently by near-infrared diode laser spectroscopy to reinvestigate several bands of  $\widetilde{A} {}^{3}\Pi_{i} - \widetilde{X} {}^{3}\Sigma^{-}$  electronic transition<sup>5,6</sup> and photoelectron spectroscopy to determine the energetics of the singlet electronic states.<sup>7</sup>

Several groups have investigated the spectra of CCO in the ground electronic state. Yamada *et al.*<sup>8</sup> have measured the  $\nu_1$  band of CCO using diode laser kinetic spectroscopy. Here, the CCO radical was generated by the 193 or 248 nm excimer laser photolysis of carbon suboxide. The molecular parameters obtained from the observed infrared spectrum were then used to search and successfully record the pure rotational spectrum of the radical both in laboratory and space.<sup>9,2</sup> The attempt by Yamada *et al.* to detect CCO by an electrical discharge of C<sub>3</sub>O<sub>2</sub> was not successful.

Most recently, two of the low *N*-value transitions of CCO in the ground vibrational state were measured using a Fabry–Perot type Fourier-transform microwave spectrometer and a pulsed discharge nozzle.<sup>10</sup> An extended study of the  $\nu_1$ 

fundamental band and the first observation of the  $\nu_1 + \nu_2$ - $\nu_2$  sequence band were reported by Moazzen-Ahmadi *et al.*<sup>11</sup> All the spectroscopic investigations have confirmed the structure of the ground electronic state of CCO.

In the present work, we used a hollow cathode discharge cell with a flowing mixture of carbon suboxide and helium to generate the CCO radical. The absorption of the  $2\nu_1 - \nu_1$ and  $\nu_1 + \nu_3 - \nu_3$  difference bands were recorded using a tunable infrared diode laser spectrometer. Ninety and seventyfive rovibrational transitions were measured in the  $2\nu_1 - \nu_1$  $v_1 + v_3 - v_3$  bands, respectively. Unambiguous and N-numbering of the bands were made by combination differences and/or the observation of the weak low N-value lines near the band origin. The nonlinear least-squares analyses of the two bands yielded spectroscopic constants for the (001), (200), and (101) vibrational states. The results reported here and those of our earlier work<sup>11</sup> on the (000), (100), (010), and (110) vibrational states provide a more comprehensive set of spectroscopic constants for CCO in the ground electronic state. In the present study we highlight the usefulness of a hollow cathode discharge as a source for study of the hot band spectra of short-lived molecules.

#### **II. EXPERIMENTAL DETAILS**

The measurements were carried out using equipment and procedures very similar to those used in the infrared studies of  $C_3^{12}$  and  $C_4^{13,14}$  The spectra were recorded using a tunable diode laser spectrometer and a multipass hollow cathode absorption cell. The cathode (70 cm long, 5 cm diam) was constructed from two concentric cylinders and was cooled to -10 °C by flowing methanol from a refrigerated bath through the space between the cylinders.

Three diodes were used in order to get reasonably complete coverage of the 1920–1960 cm<sup>-1</sup> region of interest. Wave numbers of the CCO lines were measured relative to known lines of allene<sup>15</sup> using fringes from a temperature-

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stabilized solid Ge étalon with a free spectral range of 0.0488 cm<sup>-1</sup>. The absolute uncertainty of the measured lines is  $\pm 0.002$  cm<sup>-1</sup>.

The CCO radicals were generated in a 1 kHz modulated discharge through a flowing mixture of carbon suboxide (0.2 Torr) and helium (1 Torr). The discharge modulation technique was employed for detection.<sup>16</sup> Carbon suboxide was prepared by dehydration of malonic acid with  $P_2O_5^{17}$  or by hot-tube pyrolysis of molten diacetyltartaric anhydride<sup>18</sup> and purified by fractional distillation through traps cooled by dry-ice (-78 °C) and ethanol (-115 to -110 °C) slush baths, the latter retaining pure C<sub>3</sub>O<sub>2</sub>.

In contrast to the relatively high discharge current (~400 mA) required for the optimum production of  $C_3$ ,<sup>12</sup>  $C_4$ ,<sup>13,14</sup> and  $C_5$ ,<sup>19</sup> the maximum signal for the CCO radical was achieved at a much lower current (40 mA). See Ref. 11 for more experimental details.

#### **III. SPECTRA, ANALYSIS, AND RESULTS**

The ground electronic state of CCO is  ${}^{3}\Sigma^{-}$ . Each rotational level, characterized by the rotational angular momentum quantum number N, is split into three fine structure levels  $F_1$ ,  $F_2$ , and  $F_3$  characterized by total angular momentum J=N+1, J=N, and J=N-1 (except for N=0, where there is only a single component with J=1), respectively. For all but the lowest values of N, almost all of the intensity of the vibration–rotation P(N), Q(N), or R(N) transition is contained in the three components with  $\Delta J = \Delta N$ .

The general features of the difference bands  $2\nu_1 - \nu_1$ and  $\nu_1 + \nu_3 - \nu_3$  are similar to those for the  $\nu_1$  band.<sup>11</sup> For transitions with  $N'' \ge 8$ , each spectral line appears as a doublet with a 2:1 intensity ratio and an essentially constant spacing between the two components. This spacing which measures the difference between the values of the spin-spin parameter ( $\lambda$ ) in the upper and lower states is about 205 MHz for the  $2\nu_1 - \nu_1$  band and 175 MHz for the  $\nu_1 + \nu_3$  $-\nu_3$  band. The corresponding spacing for the  $\nu_1$  band is slightly larger (248 MHz).<sup>11</sup>

The doubling of the spectral lines is due to the fact that the  $F_1$  and  $F_3$  components are close in energy and that the contribution to energy from the spin-rotation effect is nearly the same for both components. Moreover, the  $F_2$  component is widely separated from the other two components and is at higher energy.<sup>20</sup> Since all three components are expected to have similar intensities, a doublet with the higher frequency component half as intense as the lower frequency component is observed. Figure 1 illustrates a small portion of the spectra recorded for the  $2\nu_1 - \nu_1$  band. Here, the trace shows the observed spectrum near the R(18) transition which exhibits a normal ("zeroth derivative") line shape due to the discharge modulation technique employed.<sup>16</sup> The stronger doublet to the higher frequency side of the R(18) transition is the P(19) transition of the  $\nu_1$  band.<sup>11</sup>

Table I lists the ninety lines measured for the  $2\nu_1 - \nu_1$ band. As can be seen, large spin splittings (0.01-0.03 cm<sup>-1</sup>) were observed for P(2)-P(6) and R(1)-R(7). The absolute *N*-numbering of the band was easily determined by com-

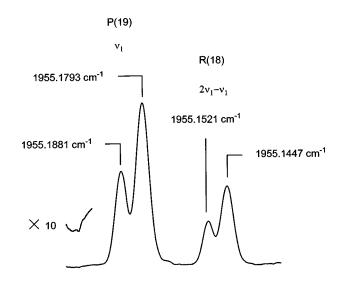


FIG. 1. The observed spectrum of CCO in the region of the R(18) line of the  $2\nu_1 - \nu_1$  difference band. For the identification of the individual components see Table I. The doublet to the higher frequency side of R(18) is the P(19) of the  $\nu_1$  band.

paring the lower level combination differences R(N-1) – P(N+1) with those of the upper state combination differences R(N) - P(N) derived from the  $\nu_1$  band. The data in Table I were analyzed using a full matrix treatment. The effective Hamiltonian and the matrix elements are given in Ref. 21. The lower state parameters were held fixed at the values given in Ref. 11. For the unresolved lines, the calculated frequencies were obtained from the average of the theoretical frequencies of the individual members. The resulting molecular parameters are listed in Table II; a comparison of the calculated and the measured frequencies are given in Table I.

The wave numbers for 75 lines observed in the  $\nu_1 + \nu_3 - \nu_3$  band are listed in Table III. Since the  $\nu_3$  band of CCO in the gas phase has not been observed, the *N*-numbering of the band could not be assigned using the combination differences and had to be otherwise determined.

Initially, the high *N*-value transitions in the  $v_1 + v_3 - v_3$  band were recorded. These lines appear as unresolved doublets with an  $F_3$  component visible as a shoulder to the higher frequency side. An example of spectra recorded for this band is shown in Fig. 2. The unresolved doublet near 1944.09 cm<sup>-1</sup> is the R(9) of the  $v_1 + v_3 - v_3$  band. Also shown in this figure are the P(31) of the  $v_1$  band and R(2) of the  $2v_1 - v_1$  band which occur in the same frequency region. As can be seen, the triplet for R(2) of the  $2v_1 - v_1$  band is fully resolved and exhibits relative intensities within the triplet that are in accordance with the approximate theoretical values.<sup>22</sup> The weak unidentified doublet, which is almost certainly due to the CCO radical, has not been identified.

To establish the *N*-numbering of the band, a preliminary analysis of the high *N*-value transitions was made. The assignment was then chosen, and is given in Table II, so that it resulted in  $\alpha_3(=B_{000}-B_{001})$  which was positive and comparable in magnitude to  $\alpha_1(=B_{000}-B_{100})$ . This assignment was later confirmed by the observation of the low *N*-value

TABLE I. Observed transitions in the  $2\nu_1 - \nu_1$  band of CCO (in cm<sup>-1</sup>).<sup>a</sup>

| N'' | P(N'')    | ObsCalc.        | R(N'')    | ObsCalc         |
|-----|-----------|-----------------|-----------|-----------------|
| 1   |           |                 | 1943.3514 | -0.0008         |
|     |           |                 | 1943.3707 | -0.0004         |
| 2   | 1940.3301 | 0.0003          | 1944.1020 | -0.0000         |
|     | 1940.3301 | 0.0003          | 1944.1152 | -0.0007         |
|     |           |                 | 1944.1365 | -0.0004         |
| 3   | 1939.5511 | 0.0007          | 1944.8432 | 0.0006          |
|     | 1939.5511 | 0.0007          | 1944.8558 | -0.0010         |
|     | 1939.5183 | 0.0014          | 1944.8558 | -0.0011         |
| 4   | 1938.7662 | 0.0003          | 1945.5754 | -0.0003         |
|     | 1938.7662 | 0.0003          | 1945.5866 | 0.0005          |
|     | 1938.7482 | -0.0009         | 1945.5866 | 0.0005          |
| 5   | 1937.9739 | -0.0017         | 1946.3012 | -0.0007         |
|     | 1937.9739 | -0.0017         | 1946.3133 | 0.0018          |
|     | 1937.9671 | 0.0029          | 1946.3087 | -0.0005         |
| 6   | 1937.1783 | -0.0007         |           |                 |
| 0   | 1937.1783 | -0.0007         |           |                 |
|     | 1937.1693 | -0.0008         |           |                 |
| 7   | 1936.3733 | -0.0004         | 1947.7352 | -0.0012         |
|     | 1936.3775 | -0.0012         | 1947.7447 | 0.0012          |
|     | 1936.3700 | 0.0012          | 1947.7352 | -0.0012         |
| 8   | 1935.5635 | 0.0011          | 1948.4403 | -0.0020         |
| 0   | 1935.5691 | -0.0006         | 1948.4476 | -0.0013         |
| 9   | 1934.7480 | 0.0009          | 1949.1413 | -0.0005         |
| 9   | 1934.7543 | 0.0001          | 1949.1486 | 0.0005          |
| 10  | 1933.9263 | 0.0009          | 1949.8352 | 0.0001          |
| 10  | 1933.9203 | 0.0009          | 1949.8352 | 0.0004          |
| 11  | 1933.0967 | -0.0006         | 1949.8430 | 0.0002          |
| 11  |           |                 |           |                 |
| 12  | 1933.1033 | -0.0010         | 1950.5288 | 0.0007          |
| 12  | 1932.2632 | 0.0005 - 0.0000 | 1951.2015 | 0.0000 - 0.0008 |
| 10  | 1932.2697 |                 | 1951.2074 |                 |
| 13  | 1931.4227 | 0.0009          | 1951.8745 | -0.0006         |
| 14  | 1931.4284 | -0.0004         | 1951.8809 | -0.0010         |
| 14  | 1930.5745 | -0.0001         | 1952.5433 | 0.0011          |
| 1.5 | 1930.5810 | -0.0005         | 1952.5503 | 0.0013          |
| 15  | 1929.7219 | 0.0010          | 1953.2026 | -0.0002         |
|     | 1929.7285 | 0.0007          | 1953.2094 | -0.0002         |
| 16  | 1928.8618 | 0.0009          | 1953.8563 | -0.0006         |
|     | 1928.8681 | 0.0003          | 1953.8645 | 0.0007          |
| 17  | 1927.9930 | -0.0015         | 1954.5050 | 0.0004          |
|     | 1928.0014 | -0.0001         | 1954.5116 | 0.0002          |
| 18  | 1927.1219 | 0.0001          | 1955.1447 | -0.0010         |
|     | 1927.1289 | 0.0002          | 1955.1521 | -0.0004         |
| 20  | 1925.3579 | 0.0005          |           |                 |
|     | 1925.3650 | 0.0007          |           |                 |
| 21  | 1924.4654 | -0.0003         |           |                 |
|     | 1924.4722 | -0.0004         |           |                 |
| 22  | 1923.5684 | 0.0007          | 1957.6455 | 0.0005          |
|     | 1923.5748 | 0.0002          | 1957.6529 | 0.0010          |
| 23  | 1922.6635 | 0.0002          | 1958.2537 | 0.0001          |
|     | 1922.6702 | -0.0000         | 1958.2606 | 0.0001          |
| 24  | 1921.7506 | -0.0021         | 1958.8566 | 0.0010          |
|     | 1921.7577 | -0.0019         | 1958.8625 | 0.0000          |
| 25  | 1920.8355 | -0.0002         | 1959.4512 | 0.0001          |
|     | 1920.8409 | -0.0017         | 1959.4592 | 0.0012          |

<sup>a</sup>For N'' < 8, the frequencies for  $F_1$ ,  $F_2$ , and  $F_3$  components are the first, second, and third entries, respectively. Otherwise, only two frequencies are given; the first entry corresponds to the unresolved  $F_1$  and  $F_3$  components and the second entry is for the  $F_2$  component.

transitions whose wave numbers were predicted using the molecular parameters from the preliminary analysis. Figure 3 shows an example of the spectra recorded near the  $\nu_1 + \nu_3$   $-\nu_3$  band center. The weak triplet for R(2) of the  $\nu_1 + \nu_3$   $-\nu_3$  band which partially overlaps with R(25) of the  $\nu_1$ 

 $+\nu_2-\nu_2$  band is shown. The lines labelled by \* have not been identified. The results of the least-squares fit of the data in Table III, based on the chosen assignment, are listed in Table II. Comparison of the calculated and the measured frequencies are given in Table III.

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TABLE II. Molecular parameters for the  $\nu_1$ ,  $2\nu_1 - \nu_1$ , and  $\nu_1 + \nu_3 - \nu_3$  bands of CCO.<sup>a</sup>

| Parameter         | $\nu_1$         | $2\nu_1 - \nu_1$    | $\nu_1 + \nu_3 - \nu_3$ |
|-------------------|-----------------|---------------------|-------------------------|
| $v_0 / cm^{-1}$   | 1970.86332(41)  | 1941.85761(54)      | 1936.79402(56)          |
| B'/MHz            | 11453.508(50)   | 11357.05(14)        | 11380.71(34)            |
| D'/kHz            | 5.814(37)       | 5.71(21)            | 5.831(70)               |
| $\lambda'/MHz$    | 11848(24)       | 12054(18)           | 12025(20)               |
| $\lambda'_D/MHz$  | $-0.00537^{b}$  | $-0.00537^{b}$      | $-0.00537^{b}$          |
| $\gamma'/MHz$     | -17.5(4.2)      | -17.5 <sup>c</sup>  | -17.5 <sup>c</sup>      |
| B"/MHz            | 11545.59684(61) | 11453.508°          | 11479.62(36)            |
| D"/kHz            | 5.819(67)       | 5.814 <sup>c</sup>  | 5.814 <sup>c</sup>      |
| $\lambda''/MHz$   | 11600(18)       | 11848 <sup>c</sup>  | 11848 <sup>c</sup>      |
| $\lambda_D''/MHz$ | -0.00537(46)    | $-0.00537^{\circ}$  | $-0.00537^{\circ}$      |
| $\gamma''/MHz$    | -17.8166(40)    | - 17.5 <sup>c</sup> | -17.5 <sup>c</sup>      |

<sup>a</sup>The uncertainties in parentheses correspond to  $1\sigma$  from the least-squares fit. <sup>b</sup>This was held fixed at its value for (000) state. See Ref. 11.

<sup>c</sup>This was held fixed at its value for (100) vibrational state. See Ref. 11.

#### **IV. RESULTS AND DISCUSSION**

The molecular parameters resulting from the analyses of the two bands reported in this paper are compared with those of the  $\nu_1$  fundamental<sup>11</sup> in Table II. The band origin for  $2\nu_1 - \nu_1$  was determined to be 1941.85761(54) cm<sup>-1</sup>, 29.01 cm<sup>-1</sup> lower than that for the  $\nu_1$  fundamental. For comparison, the band origin for  $\nu_1 + \nu_3 - \nu_3$ was 1936.79402(56)  $\text{cm}^{-1}$ , some 34  $\text{cm}^{-1}$  below the band origin for the  $v_1$  band. This indicates that the excitation of  $v_3$  by one quantum results in a larger red-shift in  $\nu_1$  frequency than the excitation of  $v_1$  by one quantum. The change in B values with increasing  $v_1$  exhibits a smooth trend, namely  $\alpha_1(=B_{000}-B_{100})=92.09$ MHz and  $B_{100} - B_{200}$ = 96.45 MHz. The difference in the *B* values between the upper and lower states of the  $\nu_1 + \nu_3 - \nu_3$  is also in line with the above values;  $B_{001} - B_{101} = 98.91$  MHz.

From the lower state *B* values of the  $\nu_1$  and  $\nu_1 + \nu_3 - \nu_3$  bands, we find that  $\alpha_3(=B_{000}-B_{001})=65.98$  MHz. This indicates that the excitation of  $\nu_3$  vibration causes a smaller decrease in *B* than the excitation of the  $\nu_1$  vibration.

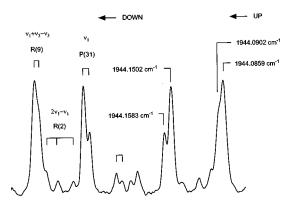


FIG. 2. The observed spectrum of CCO in the region of R(9) of the  $\nu_1 + \nu_3 - \nu_3$  difference band. Two scans (up and down) are shown. Because the difference between the spin-spin parameter in the lower and upper states for this transition are smaller than those for the  $\nu_1$  band, the  $F_2$  component appears as a shoulder to the high frequency side. The P(31) of the  $\nu_1$  band and R(2) of the  $2\nu_1 - \nu_1$  band which occur in the same frequency region are also shown. The weak doublet, which is almost certainly due to CCO, has not been identified.

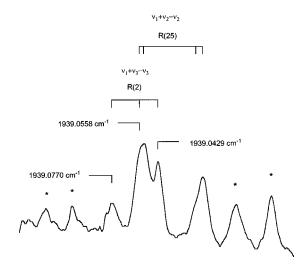


FIG. 3. The observed spectrum of CCO in the region of R(2) of the  $\nu_1 + \nu_3 - \nu_3$  difference band. For the identification of the individual components see Table III. The R(25) transition of the  $\nu_1 + \nu_2 - \nu_2$  band is also identified. The lines labeled by \* have not been identified.

The analogous quantities for other three-heavy-atom molecules follow the same trend, i.e.  $\alpha_1 = 85.49$  MHz and  $\alpha_3 = 76.16$  MHz for CCN<sup>23</sup> and  $\alpha_1 = 103.5$  MHz and  $\alpha_3 = 52.6$  MHz for N<sub>2</sub>O.<sup>24</sup> Thus the values of  $\alpha_1 = 98.91$  MHz for the  $v_3$  vibration and  $\alpha_3 = 65.98$  MHz obtained here provide further evidence in support of the *N* assignment for the  $v_1 + v_3 - v_3$  band.

Although several low *N*-value transitions were observed for each of the two bands, the analyses of the data did not permit the determinations of  $\lambda_D$  and  $\gamma$  for the (200), (001), and (101) vibrational states. This indicates that the vibrational dependence of these parameters are too small. These parameters were, therefore, held fixed at their values for (100) vibrational state.<sup>11</sup> Furthermore, only one of the spin– spin parameters for the  $\nu_1 + \nu_3 - \nu_3$  band could be determined. Since a value for  $\lambda_{001}$  is not available, this was held fixed at  $\lambda_{100}$ =11848 MHz.

The *normal* values of the centrifugal distortion constants observed here are indicative of the fact that none of the vibrational states analyzed here are strongly perturbed and that  $C_2O$  is a well behaved rigid molecule unlike  $C_3$  which shows substantial nonrigidity.

Comparison of the intensities between transitions in the  $\nu_1$  band with those in the  $2\nu_1 - \nu_1$  band shows that the strength of the  $2\nu_1 - \nu_1$  band is only about two times smaller than that for the  $\nu_1$  band. See Fig. 1. The same is true for the relative intensity of  $\nu_1 + \nu_3 - \nu_3$  and  $\nu_1$ . This implies that a significant fraction of the CCO radicals are created in vibrationally excited states in a hollow cathode discharge. This might also be the case for the  $C_n$  clusters, in which case many of the excited vibrational states of these species can be probed.

In summary, the  $2\nu_1 - \nu_1$  and  $\nu_1 + \nu_3 - \nu_3$  difference bands of the CCO radical have been observed in a hollowcathode discharge of a flowing C<sub>3</sub>O<sub>2</sub>+He mixture by means of a tunable diode laser spectrometer. Rovibrational transitions with *N*-values up to 34 were assigned and fitted. Analyses of the two bands yielded band centers of 1941.85761(54)

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TABLE III. Observed transitions in the  $\nu_1 + \nu_3 - \nu_3$  sequence band of CCO (in cm<sup>-1</sup>).<sup>a</sup>

| N'' | P(N'')     | ObsCalc. | R(N'')    | ObsCalc  |
|-----|------------|----------|-----------|----------|
| 2   |            |          | 1939.0429 | -0.0001  |
|     |            |          | 1939.0558 | -0.0001  |
|     |            |          | 1939.0770 | -0.0010  |
| 3   | 1934.4803  | -0.0006  | 1939.7847 | 0.0000   |
|     | 1934.4803  | -0.0006  | 1939.7970 | 0.0017   |
|     | 1934.4474  | -0.0005  | 1939.8003 | -0.0011  |
| 4   | 1933.6945  | 0.0003   | 1940.5187 | 0.0000   |
|     | 1933.6945  | 0.0003   | 1940.5291 | 0.0005   |
|     | 1933.6778  | 0.0000   | 1940.5291 | 0.0005   |
| 5   | 1932.9029  | 0.0014   |           |          |
|     | 1932.9029  | 0.0014   |           |          |
|     | 1932.8917  | 0.0012   |           |          |
| 6   | 1932.1021  | -0.0003  | 1941.9662 | 0.0005   |
|     | 1932.1021  | -0.0003  | 1941.9725 | -0.0001  |
|     | 1932.0944  | 0.0004   | 1941.9725 | -0.0001  |
| 8   |            |          | 1943.3872 | -0.0002  |
|     |            |          | 1943.3919 | -0.0011  |
| 9   |            |          | 1944.0859 | -0.0011  |
| -   |            |          | 1944.0902 | -0.0025  |
| 10  | 1928.8377  | 0.0004   | 171110702 | 0.0020   |
|     | 1928.8420  | -0.0014  |           |          |
| 11  |            |          |           |          |
|     | 1928.0114  | -0.0005  |           |          |
| 12  | 1927.1687  | 0.0010   | 1946.1460 | 0.0001   |
|     | 1927.1745  | 0.0008   | 1946.1516 | -0.0001  |
| 13  | 1926.3230  | -0.0001  | 1946.8188 | -0.0001  |
| 10  | 1926.3282  | -0.0009  | 1946.8241 | -0.0006  |
| 14  | 1925.4732  | 0.0013   | 1947.4848 | -0.0004  |
|     | 1925.4792  | 0.0013   | 1947.4913 | 0.0002   |
| 15  | 1924.6150  | 0.0008   | 1948.1442 | -0.0002  |
| 10  | 1924.6207  | 0.0005   | 1948.1506 | -0.0002  |
| 16  | 1923.7492  | -0.0007  | 1948.7983 | 0.0002   |
| 10  | 1923.7555  | -0.0004  | 1948.8048 | 0.0010   |
| 17  | 1725.15555 | 0.0001   | 1949.4449 | 0.0006   |
| 17  |            |          | 1949.4510 | 0.0009   |
| 18  | 1922.0032  | 0.0013   | 1950.0835 | -0.0004  |
| 10  | 1922.0095  | 0.0017   | 1950.0897 | -0.0001  |
| 19  | 1921.1182  | 0.0002   | 1950.7166 | -0.0003  |
|     | 1921.1232  | -0.0002  | 1950.7222 | -0.0006  |
| 20  | 1920.2266  | -0.0019  | 1951.3437 | 0.0006   |
|     | 1920.2317  | -0.0020  | 1951.3491 | 0.0001   |
| 21  |            |          | 1951.9637 | 0.0010   |
|     |            |          | 1951.9691 | 0.0005   |
| 22  |            |          | 1952.5750 | - 0.0006 |
|     |            |          | 1952.5812 | -0.0003  |
| 23  |            |          | 1953.1822 | 0.0005   |
| -   |            |          | 1953.1882 | 0.0006   |
| 24  |            |          | 1953.7816 | 0.0004   |
|     |            |          | 1953.7877 | 0.0004   |
| 25  |            |          | 1954.3738 | -0.0001  |
|     |            |          | 1954.3791 | -0.0007  |
| 28  |            |          | 1956.1107 | -0.0010  |
| _0  |            |          | 1956.1181 | 0.0010   |
| 33  |            |          | 1958.8738 | 0.0005   |
|     |            |          | 1958.8803 | 0.0000   |
| 34  |            |          | 1958.8805 | -0.0009  |
| 57  |            |          |           |          |
|     |            |          | 1959.4113 | 0.000    |

<sup>a</sup>For N'' < 8, the frequencies for  $F_1$ ,  $F_2$ , and  $F_3$  components are the first, second, and third entries, respectively. Otherwise, only two frequencies are given; the first entry corresponds to the unresolved  $F_1$  and  $F_3$  components and the second entry is for the  $F_2$  component.

cm<sup>-1</sup> and 1936.79402(56) cm<sup>-1</sup> for  $2\nu_1 - \nu_1$  and  $\nu_1 + \nu_3 - \nu_3$ , respectively. The results obtained here and those of our earlier work<sup>11</sup> on the (000), (100), (010), and (110) vibrational states provide a more comprehensive set of spectroscopic constants for CCO in the ground electronic state.

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