

# Nanomagnetic properties of the meteorite cloudy zone

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**Meteorites contain a record of their thermal and magnetic history, written in the intergrowths of iron-rich and nickel-rich phases that formed during slow cooling. Of intense interest from a magnetic perspective is the "cloudy zone", a nanoscale intergrowth containing tetrataenite - a naturally occurring hard ferromagnetic mineral, which has potential applications as a sustainable alternative to rare-earth permanent magnets. Here we use a combination of high-resolution electron diffraction, electron tomography, atom probe tomography and micromagnetic simulations to reveal the three-dimensional architecture of the cloudy zone with subnanometre spatial resolution, and model the mechanism of remanence acquisition during slow cooling on the meteorite parent body. Isolated islands of tetrataenite are embedded in a matrix of an ordered superstructure. The islands are arranged in clusters of three crystallographic variants, which control how magnetic information is encoded into the nanostructure. The cloudy zone acquires paleomagnetic remanence via a sequence of magnetic domain state transformations (vortex to two-domain to single-domain), driven by Fe-Ni ordering at 320°C. Rather than remanence being recorded at different times at different positions throughout the cloudy zone, each sub-region of the cloudy zone records a coherent snapshot of the magnetic field that was present at 320°C. Only the coarse and intermediate regions of the cloudy zone are found to be suitable for paleomagnetic applications. The fine regions, on the other hand, have properties similar to those of rare-earth permanent magnets, providing potential routes to synthetic tetrataenite-based magnetic materials.**

Meteorite | Paleomagnetism | Cloudy Zone | Tomography | Rare Earth Magnets

Meteorites are fragments of asteroids – the remnants of planetesimals that formed during the first few million years of the solar system. Magnetic minerals in meteorites, such as Fe-Ni alloy, preserve evidence that magnetic fields were generated by the liquid iron cores of differentiated planetesimals shortly after their formation, much like the magnetic field of the Earth is generated by its liquid iron core today<sup>(1)</sup>. Slow cooling ( $\sim 10^{-1}$  to  $\sim 10^3$  °C per million years) of meteoritic Fe-Ni alloy gives rise to distinctive metallurgical features spanning a range of length scales, from the familiar centimetre-scale Widmanstätten intergrowth of kamacite (bcc Fe-Ni) and taenite (fcc Fe-Ni) to a nanometre-scale intergrowth of tetrataenite islands (ordered Fe<sub>0.5</sub>Ni<sub>0.5</sub>) in an Fe-rich fcc or bcc matrix, known as the "cloudy zone" (Fig 1). The cloudy zone forms at  $T < 450$  °C by a process of spinodal decomposition<sup>(2)</sup>. Islands of tetrataenite can vary in diameter from over 500 nm to less than 10 nm, depending on the cooling rate and the local Ni content. Small island sizes promote uniformly magnetized (i.e. single-domain) states, while the L10 ordered tetragonal symmetry of tetrataenite generates high magnetic

coercivity (up to 2 T). These two properties combine to make the cloudy zone a potent carrier of paleomagnetic information in meteorites, as well as a potential sustainable replacement for rare-earth permanent magnet materials<sup>(3, 4)</sup>. As islands of tetrataenite form in the presence of the meteorite parent body's internally generated magnetic field, it has been proposed that the cloudy zone preserves a record of the field's intensity and polarity<sup>(5, 6)</sup>. The ability to extract this paleomagnetic information only recently became possible with the advent of high-resolution X-ray magnetic imaging methods, which are capable of quantifying the magnetic state the cloudy zone on sub-micrometer length scales<sup>(7)</sup>. Despite notable successes of this new "nanopaleomagnetic" approach<sup>(8–11)</sup>, precisely how paleomagnetic information is recorded by the cloudy zone remains unknown, and further questions about the timing of primary remanence acquisition during cooling, and the susceptibility of the cloudy zone to acquire secondary remanence post cooling, remain open. Providing answers to these questions is an essential step in the quest to develop a quantitative theory linking the magnetic state of the cloudy zone to the intensity of the parent-body magnetic field.

Here, we begin to address these issues by performing a combined tomographic and micromagnetic study of the cloudy zone in the Tazewell IAB sLH iron meteorite, which cooled

## Significance Statement

The cloudy zone is naturally occurring nanocomposite found in Fe-Ni metal bearing meteorites. It is not only a potent carrier of paleomagnetic information from the early solar system, but shows promise as a sustainable alternative to rare-earth based permanent magnets. Here we explain how the remarkable magnetic properties of the cloudy zone are linked to its three-dimensional chemical, crystallographic and magnetic architecture, using a state-of-the-art combination of nanometre to sub-nanometre resolution tomography and micromagnetic simulations. We discover the mechanism by which paleomagnetic information becomes encoded into the cloudy zone, and, inspired by our findings, point towards potential pathways to optimise synthetic analogues of the cloudy zone for industrial applications.

R.J.H. conceived the 3D-EDS and APT experiments, contributed to writing the manuscript and performed the micromagnetic simulations. J.F.E. wrote the manuscript. P.A.M., J.F.E. and A.S.E. conceived the SPED experiments. J.F.E. and A.S.E. conducted the SPED experiments. S.M.C. developed the EDS thickness mapping technique. B.H.M. developed the fuzzy clustering algorithm for SPED data. Z.S., R.B. and J.F.E. conducted the 3D-EDS experiments. R.B. and P.A.J.B. conducted the APT experiments. Z.S. reconstructed the 3D-EDS tomographic data. J.F.E., R.B. S.M.C., P.A.J. and A.S.E. analyzed the results. P.A.M. developed the anti-phase boundary model. All authors reviewed the manuscript.

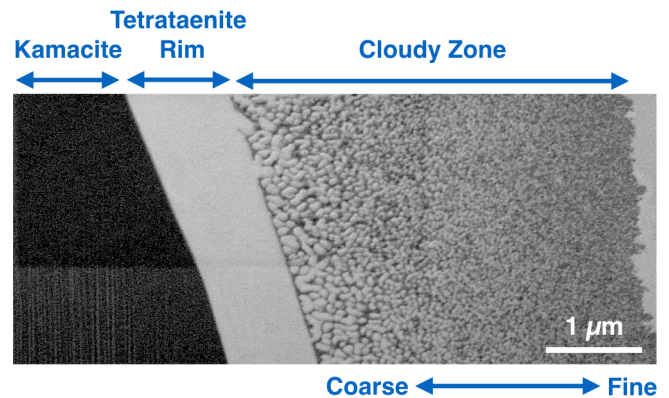
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125 at a rate that is conducive to the preservation of nanopaleomagnetic remanance. Cooling rate affects both the size  
 126 of the tetrataenite islands and the degree of Fe-Ni ordering within them (5). Alternating (002) planes of Ni and Fe, in  
 127 nearly equal ratios, are required to form the ordered tetragonal structure of tetrataenite, with the magnetic easy axis  
 128 aligning along the crystallographic c-axis. If the cooling rate is too fast, insufficient time is available for the Fe-Ni ordering  
 129 to take place, and the soft ferromagnetic phase taenite is retained that is unsuitable for nanopaleomagnetic analysis. The  
 130 quickly cooled Bishop Canon IVA meteorite ( $\sim 2500$  °C/Myr) demonstrates such soft magnetic behavior, characterised by the  
 131 presence of large, meandering magnetic domains throughout the cloudy zone, consistent with the absence of tetrataenite  
 132 (10). If the cooling rate is too slow ( $< 0.5$  °C/Myr), the tetrataenite islands grow to such size that they develop multiple  
 133 magnetic domains. In addition, the coarse length scale of the cloudy zone in slowly cooled meteorites leads to the  
 134 conversion of the metastable, paramagnetic, fcc matrix phase to a more stable, soft ferromagnetic, bcc martensite phase,  
 135 which further degrades and complicates the paleomagnetic properties. Both phenomena are observed in the slowly cooled  
 136 mesosiderites (12), making them unsuitable for nanopaleomagnetic study. A "Goldilock's zone" for cooling rates exists,  
 137 whereby cooling is slow enough to enable tetrataenite formation, yet fast enough to prevent the tetrataenite islands  
 138 growing to sizes that are beyond their single-domain threshold, and to maintain the metastable, paramagnetic, fcc structure  
 139 of the matrix phase. These considerations constrain usable meteorite samples to those which have experienced parent  
 140 body cooling rates between  $\sim 0.5$  °C/Myr and  $\sim 150$  °C/Myr (12). With a cooling rate of  $\sim 20.8$  °C/Myr at 500-600 °C  
 141 (13), the Tazewell IAB-sLH meteorite displays a well preserved cloudy zone that has not been affected by shock and is well  
 142 within the desired cooling-rate window for nanopaleomagnetic studies. We present a multi-scale, multi-dimensional study  
 143 of the cloudy zone in the Tazewell meteorite, yielding structural, crystallographic and chemical information across the  
 144 entire cloudy zone, in three dimensions and at nanometre to sub-nanometre spatial resolution. This new information  
 145 provides the input for micromagnetic simulations of both individual tetrataenite islands and small clusters of islands, which  
 146 have far-reaching implications for nanopaleomagnetic studies of planetesimals as well as for our understanding of the  
 147 thermodynamic and structural behaviour of possible replacements for rare-earth permanent magnet materials.

## 172 Results

173 **3D structure of the cloudy zone.** A two-dimensional cross section through the interface region of the Widmansätten  
 174 intergrowth in the Tazewell meteorite is shown in Fig. 1. A region of kamacite (far left) is followed by a  $\sim 1\mu\text{m}$  thick  
 175 rim of tetrataenite, which is then followed by the island/matrix nanostructure of the cloudy zone. The size of the tetrataenite  
 176 islands (bright regions) decreases systematically from 150 nm adjacent to the rim to less than 10 nm at a distance of  
 177 several microns away from the rim. This change from 'coarse' to 'fine' island sizes is caused by a decrease in the local Ni  
 178 concentration with increasing distance from the rim, which in turn lowers the temperature at which spinodal decomposition  
 179 initiates during cooling (thereby reducing the diffusion

length and time available for coarsening of the islands). The matrix phase (dark regions) appears to percolate throughout  
 the entire cloudy zone. We reveal the true three-dimensional structure of the cloudy zone through the complementary  
 approaches of 3D Energy Dispersive Spectroscopy (3D-EDS) and Atom Probe Tomography (APT) (see Methods).

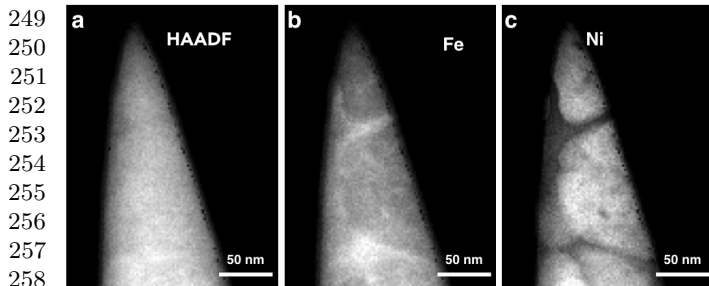


**Fig. 1.** FIB-secondary electron micrograph of the interface region between kamacite (bcc iron) and taenite (fcc Fe-Ni), which forms part of the Widmanstätten intergrowth in the Tazewell meteorite. Kamacite (left) is followed by a rim of pure tetrataenite (centre), which then gives way to the cloudy zone (right), a nanoscale intergrowth of tetrataenite islands surrounded by an Fe-rich matrix. The cloudy zone transitions from coarse tetrataenite particles ( $> 150$  nm) to fine tetrataenite particles ( $< 50$  nm) with increasing distance from the tetrataenite rim. The matrix phase (dark in the image) percolates throughout the entire cloudy zone.

Due to the similarity in mean atomic number between the two different Fe-Ni alloy phases, conventional Scanning Transmission Electron Microscopy (STEM) using the High-Angle Annular Dark-Field (HAADF) signal does not resolve the internal microstructure in the cloudy zone (Fig ??a). The Fe-rich matrix (Fig ??b) and the Ni-rich tetrataenite islands (Fig ??c) are clearly resolved through the acquisition of an EDS chemical map at each angle of a tilt series. The spectral signals are de-noised using principal component analysis (PCA) and the total intensity associated with the  $\text{Fe}_{K\alpha}$  peak (Fig ??b) and the  $\text{Ni}_{K\alpha}$  peaks are integrated (Fig ??c). The resulting tilt series of chemical maps are reconstructed to form a quantitative 3D volume of Fe and Ni concentrations using a compressive sensing algorithm (14, 15).

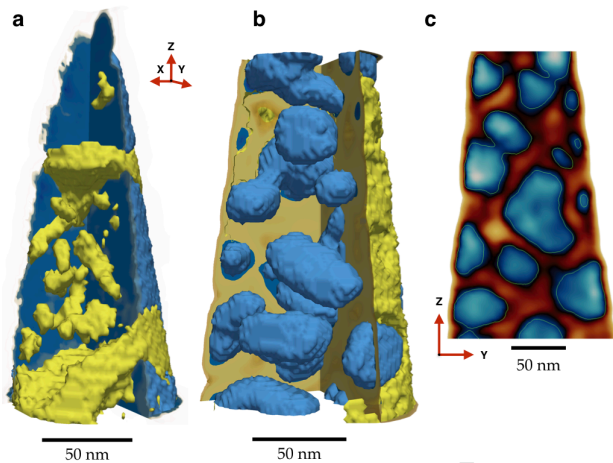
In the coarse cloudy zone (Fig 2a and SI Appendix, Movie S1) we observe two large tetrataenite islands (blue). The sample volume is shown as reconstructed by STEM-EDS tilt-series tomography and visualized as surfaces of constant intensity, termed "isosurfaces." The intensities defining the isosurfaces were defined by the strong contrast between the two phases in the Ni maps to reveal the morphology of the matrix and tetrataenite phases. In the coarse cloudy zone (Fig. 3a), the tetrataenite islands contain continuous threads of matrix phase (yellow) running through them. The matrix threads form a mostly continuous network of 10 to 30 nm diameter structures, with occasional small isolated patches of matrix contained within the tetrataenite islands.

Figure 2b and SI Appendix, Movie S2 show a region of the intermediate-fine cloudy zone. With smaller island sizes, the full three-dimensional shape of some of the tetrataenite islands



**Fig. 2.** (a) STEM HAADF micrograph of the coarse cloudy zone. (b) Iron concentration map collected with 3D-EDS. (c) Nickel concentration map collected with 3D-EDS.

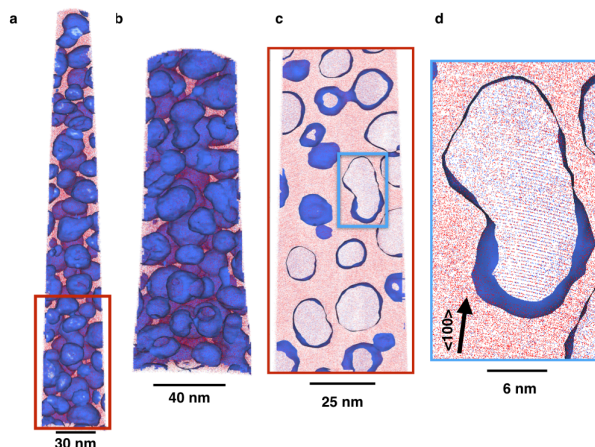
in this tomographic volume are recovered. We use best-fit ellipsoids as a representative metric for island size within the tomographic volume (16, 17). In the intermediate region of the cloudy zone we find the major axis diameter to range from 18 nm up to 100 nm. Comparison of the major, intermediate and minor ellipsoid axes yields a prolate tri-axial symmetry with the ratio of major/intermediate and intermediate/minor axes ranging from  $\sim 1$  to  $\sim 2$ , with an average of  $\sim 1.2$  (we caution, however, that the small number of complete particles observed means that these values may not be representative).



**Fig. 3.** (a,b,c) Three-dimensional visualizations of tomographic reconstructions from STEM-EDS tilt-series using the characteristic Ni K X-ray emission lines. In (a, b) surface of constant intensity, termed 'isosurfaces' are shown. A rectangular subvolume ('corner cut') has been removed to enhance the visibility of small features, with the tetraenaite particles depicted in blue and the Fe-rich matrix phase in yellow for (a) the coarse and (b) the medium cloudy zones. (c) A single plane or 'orthoslice' from the center to the reconstruction volume of the medium cloudy zone.

The three dimensional nature of a fine region of the cloudy zone was explored through the use of APT (Fig 3, SI Appendix and SI Appendix, Movies S3 and S4). Figures 3a and b show APT data from the fine cloudy zone (tetraenaite particles with long diameter of 30 nm or less). These are two data sets collected from the same needle (additional APT studies of other parts of the cloudy zone are provided in the SI Appendix). Isosurfaces defined by a Ni content of 32.5 % identify 260 tetraenaite islands within the APT needle; of these only 8 of these volumes are fully contained in the tip shown in Fig. 3a. For these particles we estimate the major axis of the best fitting ellipse at 20 nm, based on the extent of the isosurface. In general the islands measured in the volume seen

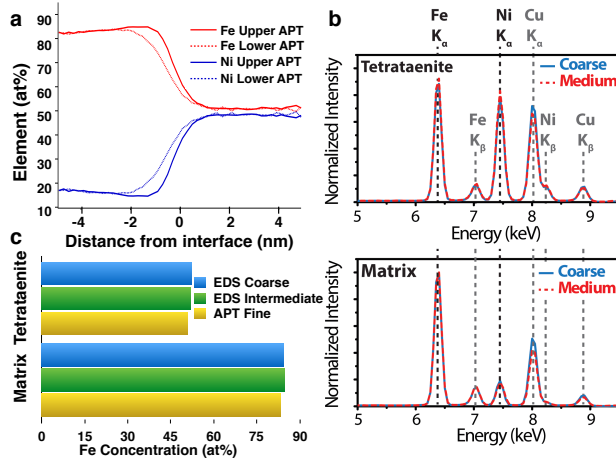
here possess an oblate spheroid aspect ratio. It should be noted that particle geometry is poorly constrained by APT, as this is highly sensitive to assumptions made during the reconstruction process. However, these reconstruction artefacts do not affect the robustly determined chemistry or absolute volume fraction of the two phases. For this reason, 3D-EDS provides the most reliable geometrical information about the cloudy zone. In Fig. 3d we are able to observe the atomic planes in the tetraenaite particle, demonstrating that the needle was aligned close to  $\langle 100 \rangle$ .



**Fig. 4.** (a,b) Atom Probe Tomography reconstruction of two data sets of same needle from the fine cloudy zone showing the 32.5 at% Ni isosurface. (c) 1 nm orthoslice from the region highlighted in (a). (d) Zoomed in detail of the tetraenaite particle highlighted in (c). Atomic planes along a  $\langle 100 \rangle$  can be resolved in the interior of the 32.5 at% Ni isosurface.

**Chemical composition of the cloudy zone.** The composition of tetraenaite islands and the matrix phase in coarse, medium and fine regions of the cloudy zone was measured using APT (Fig. 3a and b and SI Appendix). The transition from tetraenaite to matrix corresponds to a sharp compositional gradient that is just 2 nm wide, as seen in the proxigrams in Fig. 4a. The average composition of tetraenaite particles is found to be  $49.61 \pm 0.16$  at% Fe,  $50.00 \pm 0.16$  at% Ni and  $0.39 \pm 0.02$  at% Co. The matrix phase is found to be  $81.95 \pm 0.15$  at% Fe,  $17.69 \pm 0.16$  at% Ni and  $0.36 \pm 0.02$  at% Co. The error presented is equal to two standard deviations based on counting statistics. The chemical composition information contained in the 3D-EDS data sets was also examined (see Methods), and agrees very well with APT measurements (albeit with significantly greater uncertainties). Our quantification of the 3D-EDS measurements produces an average spectrum for each phase as seen in Fig. 4b. Cliff-Lorimer quantification of the average spectra gives the composition of coarse cloudy zone phases as tetraenaite at  $52 \pm 2$  at% Fe,  $48 \pm 2$  at% Ni, and the matrix at  $84 \pm 2$  at% Fe,  $16 \pm 2$  at% Ni (errors reflect uncertainties in x-ray counting statistics). Likewise for the medium cloudy zone we observe the tetraenaite composition to be  $52 \pm 2$  at% Fe and  $48 \pm 2$  at% Ni, with the matrix phase measured at  $85 \pm 2$  at% Fe and  $15 \pm 2$  at% Ni. The elemental compositions of islands and matrix do not change significantly between the coarse, medium and fine regions (Fig 4c). Instead, the variations in both APT and 3D-EDS data sets correspond to the uncertain-

ties in the quantification methods. Ultimately, APT provides us with the most accurate chemical picture of the cloudy zone, whereas 3D-EDS provides us with the most accurate geometric information.

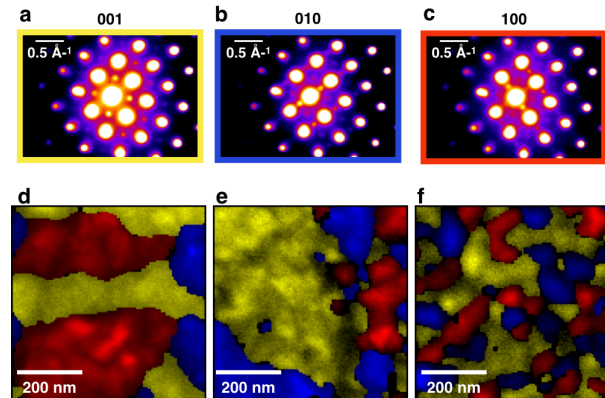


**Fig. 5.** (a) Proxigrams for the two APT datasets in Fig. 3a and b reveal a sharp interface between the tetrataenite and matrix phases. (b) The thickness-map-derived spectra for the two 3D-EDS studies presented. (c) Comparison of Fe composition quantification results by APT and EDS for the tetrataenite and matrix phases. APT analyses revealed the average composition of tetrataenite in the fine cloudy zone to be  $48.4 \pm 0.4$  at% Fe to  $51.1 \pm 0.4$  at% Ni and the matrix phase in the fine region to be  $81.8 \pm 0.2$  at% Fe and  $17.8 \pm 0.2$  at% Ni. EDS quantification yielded a composition of  $52.1 \pm 2$  at% Fe and  $47.9 \pm 2$  at% Ni for the tetrataenite islands and  $84.4 \pm 2$  at% Fe and  $15.6 \pm 2$  at% Ni for the matrix.

**Crystallographic analysis of the cloudy zone.** Scanning precession electron diffraction (SPED) allows a two-dimensional array of electron diffraction patterns to be recorded with a step size of 5 nm, while minimizing dynamical scattering effects and increasing the recorded area of the Ewald sphere. SPED experiments produce a series of diffraction patterns, used to map out the distribution of tetrataenite c-axis orientations, as well as to identify an ordered superstructure in the matrix phase. Recording a diffraction pattern at every pixel results in a statistically dense data set, to which we applied machine learning strategies to deconvolve signals arising from overlapping features and noise. Our approach relies on first denoising the data using PCA. The processed maps are then analyzed using a cluster analysis scheme, to isolate and identify regions that are most self-similar (see Methods).

A  $\langle 110 \rangle$  lamella was tilted parallel to  $\langle 100 \rangle$  for the collection of SPED data. The effective lamella thickness at this tilt angle inevitably leads to the overlap of multiple islands in the recorded data, requiring statistical unmixing post acquisition. Cluster analysis identifies three  $\langle 100 \rangle_{\text{tet}}$  diffraction patterns, each containing a distinct set of superlattice peaks corresponding to one of the three choices of c-axis orientation for tetrataenite (Fig. 5a-c). The orientation maps for the coarse, medium and fine regions of the cloudy zone (Fig 5d-f) show groups of neighboring islands with a uniform c-axis orientation. Variations in the proportions of different easy axes in the cloudy zone are thought to reflect the strength and direction of magnetic field present during cooling (see Supplemental Information Table S2). Due to the limited field of view, however, it is not possible to extract meaningful paleomagnetic information from this data. Due to the projection

thickness at this high tilt angle, a unique matrix diffraction pattern could not be determined from this lamella.

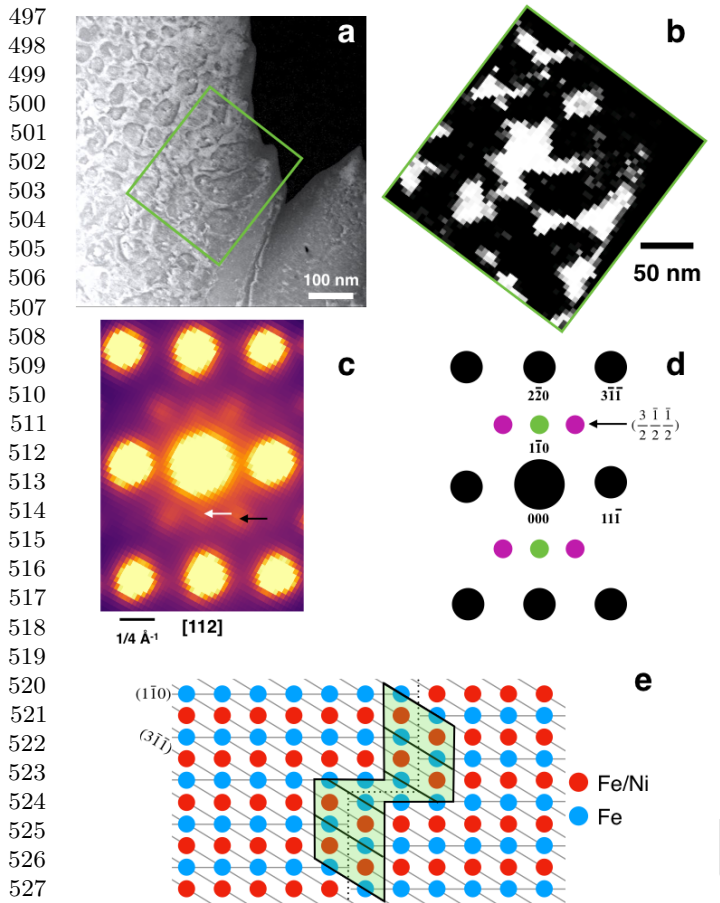


**Fig. 6.** SPED cluster analysis results along the  $[001]$  direction of the cloudy zone. (a-c) The three dominant cluster centers for the three different  $\{100\}$  tetrataenite orientations. (d-f) Cluster loading maps, showing how the three different orientations group in the (d) coarse, (e) intermediate and (f) fine regions of the cloudy zone.

To isolate the crystal structure of the matrix phase, a second lamella was fabricated with a surface normal approximately parallel to the  $[112]$  zone axis of the parent taenite crystal structure. In the field of view studied, we observed mainly  $[211]$  and  $[121]$  orientations of tetrataenite islands, both of which lack superlattice peaks associated with their ordered structure. In contrast, a superstructure in the matrix becomes apparent, leading to a stronger overall diffraction condition, as seen in the STEM annular dark field image (Fig 6a). Application of the cluster analysis on this SPED data reveals the presence of a diffraction pattern (Fig 6c) unique to the matrix phase (Fig 6b). The representative pattern contains strong reflections associated with the cubic lattice and additional superlattice reflections, which arise due to the chemical ordering of the matrix phase.

Bryson et al.(6) proposed an ordered matrix phase, based on the presence of superlattice peaks in  $[001]$  diffraction patterns collected from island/matrix regions containing only one or two out of the possible three tetrataenite orientations. When combined with EDS analysis, their observations were most consistent with a  $\text{Fe}_3\text{Ni}$  cubic primitive structure. The (cluster center) diffraction pattern in Fig 6c contains  $(1\bar{1}0)$  superlattice peaks that are consistent with such an ordering of the matrix phase (green spots in Fig 6d). However, the presence of the  $(\frac{3}{2}, \frac{1}{2}, \frac{1}{2})$  (purple spots in Fig 6d) require an additional doubling of the spacing of the  $(3\bar{1}\bar{1})$  planes that is not consistent with the  $\text{Fe}_3\text{Ni}$  structure. Furthermore, the arrows in Fig 6c indicate that the  $(\frac{3}{2}, \frac{1}{2}, \frac{1}{2})$  and  $(110)$  superlattice peaks are not exactly aligned. Close examination of these peaks (in both the raw data and cluster based results) reveals that their relative positions change as the beam position is rastered across the sample. While these observations demonstrate unequivocally that the matrix phase is an ordered structure, the precise nature of the ordered structure is clearly more complex than the pure  $\text{Fe}_3\text{Ni}$  structure proposed by Bryson et al.(20014a).

**Micromagnetic state of the cloudy zone.** Three neighboring tetrataenite islands were extracted from the tomography stack



**Fig. 7.** (a) STEM-HAADF of cloudy zone along the [112] direction. Tetraetaenite particles are dark against the bright Fe-rich matrix. (b) Cluster analysis loading map associated with the cluster center diffraction pattern shown in (c). (c) Cluster center for the matrix phase. Misalignment of superlattice reflections highlighted with white and black arrows. (d) Schematic diffraction pattern along [112] based on the proposed ordered structure. Black and green spots are in line with kinematic simulations of  $\text{Fe}_3\text{Ni}$ . Purple spots are added based on experimental results. (e) Model for the matrix phase when viewed down the [112] with an anti-phase boundary highlighted in green.

of the intermediate cloudy zone and converted to tetrahedral volume meshes (Fig. 7a). Two of the islands are approximately prolate ellipsoids, with major axis diameters of  $\sim 80\text{-}90$  nm and minor axis diameters of  $\sim 40\text{-}50$  nm. The third island is similar in size but has an L-shaped geometry (Fig. 2c), possibly resulting from a merger of two smaller islands. Finite-element micromagnetic simulations were performed initially using room-temperature micromagnetic parameters appropriate to cubic disordered  $\text{Fe}_{0.5}\text{Ni}_{0.5}$  ( $M_s = 1273$  kA/m,  $K_1 = 1$  kJ/m<sup>3</sup>,  $A_{ex} = 1.13 \times 10^{-11}$  J/m), which is a soft ferromagnet below its Curie temperature of  $\sim 450$  °C (18, 19). The matrix phase is not included in the micromagnetic models since Mössbauer spectroscopy demonstrates that the matrix phase is paramagnetic in the bulk cloudy zone (20, 21). All three islands adopted either single or double vortex states at remanence (Fig. 7b), regardless of whether simulations were performed for each island separately or whether all three islands were simulated together (including, thereby, the effects of magnetostatic interactions between the islands). Magnetostatic interactions were sufficiently large to change significantly the values of vortex

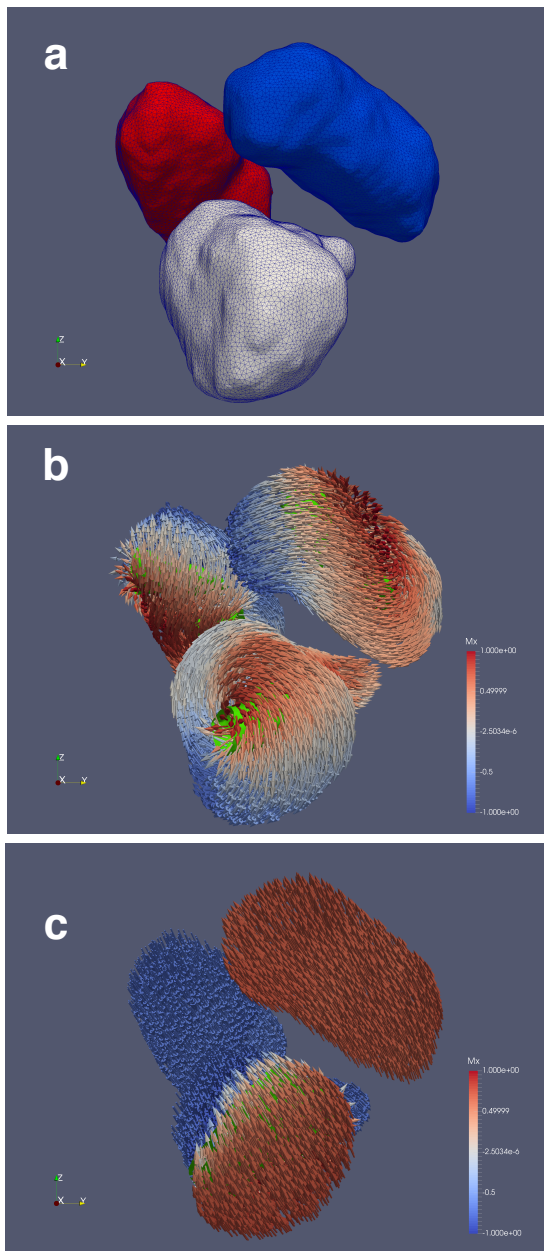
nucleation, switching and annihilation fields during hysteresis cycles, and also to influence the sense of vortex rotation and core magnetization adopted at remanence. The strength of interactions was not sufficient, however, to force the islands into a single-domain state. "High-temperature" simulations were performed by rescaling the micromagnetic parameters.  $M_s$  was reduced by a factor of 2,  $A_{ex}$  was reduced by a factor of 4 and  $K_1$  was reduced to 0, corresponding to a normalized temperature of  $T/T_c = 0.9$ . The remanence states observed at room temperature were all retained at high temperature, indicating that above 320 °C (the critical temperature for Fe-Ni ordering in tetraetaenite) the cloudy zone can best be considered as an ensemble of weakly to moderately interacting single or double vortex states.

The effect of Fe-Ni ordering below 320 °C was explored in a series of simulations in which the uniaxial anisotropy was increased from  $K_u = 0$  to  $K_u = 1370$  kJ/m<sup>3</sup> (the room-temperature value for pure tetraetaenite (22)) in increments of 10 kJ/m<sup>3</sup>. This procedure approximates the effect of a continuous and homogeneous increase in Fe-Ni ordering throughout the system. The orientation of the uniaxial easy axis was chosen to lie along either the long, intermediate or short axis of the islands (see SI Appendix, Movies S6-S9). When simulating a cluster of three islands, the c-axis of all three islands was oriented in the same direction, thereby mimicking the small clusters of islands with equal c-axis observed using SPED (Fig. 5). For isolated islands, placing the uniaxial easy axis along the long axis of the island lead to a transition to a stable single-domain state. Placing the easy axis along either the intermediate or short axes lead to the formation of a two-domain state, consisting of two equal and oppositely magnetized domains separated by a central domain wall. When magnetostatic interactions between the islands were included, however, there was a greater tendency for single-domain states to be adopted, as the interaction field caused walls to displace and eventually annihilate at the boundary of the islands (Fig. 7c). The vortex to two-domain to single-domain transition represents a complete rearrangement of the magnetic state of the system; there is no obvious relation between either the intensity or direction of remanence before and after the transition (see SI Appendix, Movies S10-S11).

## Discussion

**The chemical and structural and state of cloudy zone.** The combination of transmission electron microscopy and atom probe tomography has yielded the most comprehensive picture of the cloudy zone to date. Chemically, a rather simple picture emerges: well-defined and constant compositions for the islands at 49.6 at % Fe : 50.0 at % Ni, and matrix 82.0 at % Fe : 17.7 at % Ni are observed throughout the entire cloudy zone, with the two phases separated by a sharp ( $< 2$  nm wide) chemical gradient. This observation implies that chemical equilibrium between the islands and matrix was maintained at least down to the temperature of formation of the fine cloudy zone. The chemistry of the islands is consistent with tetraetaenite, as expected. However, the chemical composition of the matrix is revealed to be much richer in Fe than has been suggested in some previous studies (2, 6, 23–26), and inconsistent with the proposal that the matrix is pure ordered  $\text{Fe}_3\text{Ni}$  (6). The improvement in chemical analysis is explained by the more sophisticated methods of 3D-EDS quantification used here,

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**Fig. 8.** (a) Finite-element meshes of three neighboring islands extracted from the 3D-EDS tomography reconstruction of the intermediate cloudy zone. Average mesh element size is 1.6 nm. Best fitting ellipsoids yield island sizes of: (red) 82 x 48 x 38 nm, (white) 92 x 61 x 40 nm, and (blue) 92 x 47 x 41 nm. The white island has a distinct "L" shape due to coalescence of two smaller islands. (b) Remanent micromagnetic state for interacting islands of taenite, representing the likely magnetic state of the cloudy zone above 320 °C. Islands adopt predominantly single (occasionally double) vortex states. Green isosurfaces outline the approximate position of the vortex core. (c) Remanent micromagnetic state for interacting islands of tetraetaenite, representing the likely magnetic state of the cloudy zone below 320 °C. The tetraetaenite easy axis for all three islands is parallel to the long axis of the red island in (a). Two islands show ideal single domain states. The third island adopts a two-domain state with sharp domain wall at the intersection of the two coalesced islands.

combined with the enhanced capability to isolate signals from each phase through the use of high-resolution APT. This brings our composition data into agreement with the work of Miller and Russell (27), who examined meteorites with a similar cooling history. Our results contrast with the APT

results of Rout et al (2017), who examined a more rapidly cooled IVA meteorite. This suggests, that the composition of the matrix phase is a function of the cooling rate, with slower cooling rates leading to a more Fe-rich matrix, consistent with expectations based on the low-temperature phase diagram (2).

Morphologically, we observe distinct characteristics in the coarse, intermediate and fine cloudy zones. In the coarse cloudy zone, islands contain multiple threads of fine-scale matrix material, forming a network that connects back to the main matrix on either side of an island. We interpret this network of threads as relics of matrix material that became trapped as smaller islands grew and coalesced during the process of coarsening while the parent body cooled (SI Appendix, Figure S13). Tetraetaenite particles are often observed to be composed of two particles intersecting at 90° (Fig. 1, and Fig. 2c). In the intermediate cloudy zone, tetraetaenite islands contain more isolated regions of matrix, which appear like secondary precipitates, but may also be simple relics of the coarsening process. Intermediate islands have a range of sizes and shapes, but are, on average, classified as prolate ellipsoids. A weak {100} texture is observable, providing the first hint that there is crystallographic control of the shape and elongation direction of the tetraetaenite islands. We speculate that there is a preference for the elongation direction of an island to align with one of the <100> directions of the cubic parent phase, and that this elongation direction is parallel to the c-axis of tetraetaenite. Islands in the fine zone show few, if any, secondary precipitates, indicating that the fine zone has undergone less coarsening.

Crystallographically, we have demonstrated that nearby tetraetaenite islands are able to adopt different c-axis orientations, thereby confirming one of the fundamental assumptions of nanopaleomagnetic studies of the cloudy zone. The distribution of c-axis orientations is not random, but consists of ~ 50-500 nm clusters of islands with uniform c-axis orientation. Clustering was first predicted by Bryson et al. (7) in order to explain the unexpectedly large size of magnetic domains observed in the coarse and intermediate cloudy zone of the Tazewell meteorite using XPEEM. The close match between the shape and size of the c-axis clusters observed with SPED and the shape and size of magnetic domains observed with XPEEM confirms that the magnetic state of the cloudy zone is controlled by the underlying crystallography and clustering of the islands. This is most evident in the fine zone, which typically displays a strong and uniform magnetization with localized deviations in magnetization direction causing a fine-scale mottling in both electron holography and XPEEM images. A high degree of c-axis alignment is required to explain these observations. The origin of this enhanced c-axis alignment is not well understood, but various possibilities are considered below.

The electron diffraction pattern for the matrix phase (Fig 6) contains two distinct sets of superlattice reflections. The first set of reflections (110 type; green spots in Fig. 6d) are consistent with one of the three possible tetraetaenite orientations, as well as Fe<sub>3</sub>Ni when viewed along the [112]. The application of data clustering to the scanning diffraction data, however, spatially localizes this diffraction pattern to the matrix region. Additionally, the cluster center presented in Fig. 6c contains a second set of reflections that are incompatible with either Fe<sub>3</sub>Ni or tetraetaenite (pink spots in Fig. 6d). These

745 reflections require a further doubling of the lattice periodicity.  
746 Although, such a doubling could, in principle, be explained by  
747 an ordered structure of the Fe<sub>7</sub>Ni type, this structure does not  
748 provide a good fit to the intensities of superlattice peaks, nor  
749 does it explain the observation that the row of superlattice  
750 spots are clearly misaligned with respect to each other. We  
751 propose an alternative model, whereby the approximate dou-  
752 bling of lattice periodicity occurs inside antiphase boundaries  
753 created by incommensurate ordering of Fe and Ni (Fig. 6e).  
754 Local doubling of the periodicity of (3 $\bar{1}\bar{1}$ ) planes occurs within  
755 the antiphase boundary regions. The presence of antiphase  
756 boundaries also provides a mechanism for short-range chemical  
757 segregation, enabling the incorporation of additional Fe into  
758 the structure needed to match the experimentally observed  
759 composition of the matrix (note that excess Fe is found in  
760 the central part of the kinked antiphase boundary in Fig.  
761 6e). Such incommensurate ordering, creating high-densities  
762 of antiphase boundaries and local chemical segregation is well  
763 known in off-stoichiometry A<sub>3</sub>B alloys (28, 29), which create  
764 similar patterns of superlattice reflections. Conventional Fe-Ni  
765 systems in this composition range usually undergo a martensitic  
766 transformation upon cooling. However, it appears that,  
767 in this case, the high degree of lattice coherency within the  
768 intergrowth, and the relatively small volume of the matrix  
769 phase, provides a sufficiently large barrier to this transfor-  
770 mation, allowing the newly identified ordered phase to exist  
771 metastably.

772 Compositionally, the matrix lies in between Fe<sub>3</sub>Ni and  
773 Fe<sub>7</sub>Ni. The magnetic ground state of Fe<sub>3</sub>Ni is predicted to be  
774 ferromagnetic at 0 K, with a moment of 2.066  $\mu_B$  per atom  
775 (compared to the 2.2  $\mu_B$  per atom for Fe in kamacite) (30).  
776 The magnetic groundstate of Fe<sub>7</sub>Ni is predicted to be ferri-  
777 magnetic with a moment of 0.942  $\mu_B$  (30). We can speculate  
778 that the magnetic groundstate of the matrix lies somewhere  
779 between these two cases. The matrix is paramagnetic at  
780 room temperature, due to its combination of high Fe con-  
781 tent and fcc structure, which is incompatible with strong  
782 ferromagnetism (31). This property matches the proposed  
783 properties of antitaenite, a low-moment phase of taenite that  
784 is the most widely accepted explanation for the matrix phase.  
785 It is tempting, therefore, to state that antitaenite and the  
786 Fe-enriched, incommensurate ordered structure identified here  
787 are one and the same thing, although this statement would  
788 have to be corroborated by further studies involving a wider  
789 range of Fe-Ni meteorites.

### 791 **The mechanism of remanence acquisition in the cloudy zone.**

792 Magnetically, we have demonstrated that the coarse to inter-  
793 mediate cloudy zone adopts vortex states for  $T > 320$  °C.  
794 Weak to moderate magnetostatic interactions between taeni-  
795 te islands have no effect on their domain state, so vortex  
796 states will dominate in all regions containing islands larger  
797 than the single-domain threshold (22.5 nm diameter for spher-  
798 ical particles; 65 nm long axis for a prolate ellipsoid with  
799 aspect ratio 2:1 - see methods). Below 320 °C, ordering of  
800 Fe and Ni atoms to form tetrataenite leads to a dramatic  
801 increase in uniaxial anisotropy and a transition from vortex  
802 to single-domain states, via an intermediate two-domain state.  
803 A two-domain state is preserved in isolated islands when the  
804 c-axis is chosen to be perpendicular to the elongation direction  
805 of the island, or at the intersection of two coalescing islands.  
806 However, the presence of an externally applied field, as well as

magnetostatic interactions between neighboring islands, plays  
an important role, causing the domain wall to displace to  
the island boundary, where it annihilates to produce a single  
domain. Once a single-domain state has been achieved, the  
field needed to re-nucleate the domain wall is very high (of  
the order of 1 T). This means that the magnetisation state  
of each island becomes effectively fixed once the domain wall  
annihilates, i.e. this is the blocking point for the acquisition  
of paleomagnetic remanence. This mechanism of blocking is  
very different from the thermal blocking process envisaged  
in previous models (32). We predict a fundamental differ-  
ence in the mechanism of remanence acquisition within the  
coarse/intermediate cloudy zones versus the fine cloudy zone,  
caused by the presence of vortex states in the former and  
single-domain states in the latter. The domain state transition  
from single-vortex to single-domain state, via an intermediate  
two-domain state, driven by the phase transformation from  
taenite to tetrataenite, provides an effective mechanism for an  
ancient field to influence the paleomagnetic state of the cloudy  
zone, by biasing the movement of domain walls during the  
transition. Once remanence is blocked, by annihilation of the  
domain wall, the high domain wall nucleation field prevents  
remagnetisation, explaining the apparent lack of secondary  
magnetisation in the cloudy zone. In the fine cloudy zone,  
the influence of magnetostatic interactions between uniformly  
magnetised islands is likely to be dominant. Further work is  
now needed to explore these complexities, such as extending  
the simulations to include much larger numbers of interacting  
islands and also investigating the quantitative relationship  
between the final remanence state and the intensity of the  
ancient magnetic field present on the meteorite parent body.

Both the intensity and direction of remanence change dra-  
matically as a result of the transition to tetrataenite, suggest-  
ing that the final remanence is recorded during this ordering  
process, potentially erasing any memory of remanence that  
may already have been recorded above 320 °C. This statement  
has profound implications for the concept of time-resolved  
paleomagnetic information recorded throughout the cloudy  
zone. If all remanence is recorded at 320 °C, rather than at the  
temperature of spinodal decomposition, then all regions record  
their remanence at the same instance in time, rather than at  
sequentially different times according to the local Ni content.  
Whilst this destroys the concept of time resolution as origi-  
nally postulated by Bryson et al. (2014), in which sequential  
subregions of the cloudy zone record remanence at different  
times, time-resolved records of dynamo activity can still be  
derived by studying several meteorites from a single parent  
body that cooled to 320 °C at different rates. For this reason,  
Bryson et al.'s observation of strong and weak magnetisation in  
the Imilac and Esquel meteorites, representing the early and  
late stages of core solidification on the pallasite parent body,  
respectively, still stand. Their approach was subsequently  
extended by Nichols et al. (2016) to observe the predicted  
quiescent period of dynamo activity on the pallasite parent  
body that preceded core solidification. Simulations containing  
much larger ensembles of interacting islands are now necessary  
to determine whether any fraction of preexisting remanence  
can survive the transition, thereby preserving a time-resolved  
paleomagnetic record. In addition, regions of the cloudy zone  
containing islands below the single-domain threshold size may  
have significantly enhanced capacity to retain a memory of a

869 preexisting remanence, since the choice of c-axis orientation  
870 during the transition to tetrataenite is likely to be strongly  
871 influenced by the direction of uniform magnetization in the  
872 parent taenite island. This case would apply, for example, to  
873 some rapidly cooled IVA meteorites, which preserve evidence  
874 of a time-varying magnetic field generated by their parent  
875 body (10).

876  
877 **The origin of optimal hard magnetic properties in the fine**  
878 **cloudy zone.** The high degree of magnetic alignment observed  
879 in the fine cloudy zone may have several possible origins.  
880 Previous suggestions that this effect was due to exchange  
881 coupling between islands via the matrix are contingent on  
882 the matrix phase being ferromagnetic (6, 7). High-resolution  
883 Mössbauer spectroscopy measurements have since shown that  
884 the matrix is paramagnetic in the bulk phase and only ap-  
885 pears ferromagnetic at surfaces and in thin films (21). With  
886 a paramagnetic matrix, magnetostatic interactions between  
887 islands provide an alternative mechanism leading to strong  
888 alignment. Smaller islands in the fine cloudy zone mean that  
889 single-domain states, rather than vortex states, are likely to  
890 dominate both above and below 320°C. Taenite islands with  
891 uniform magnetization will generate larger magnetostatic in-  
892 teraction fields, which may cause higher degrees of c-axis  
893 alignment (and, as suggested above, increase the chances of  
894 remanence being inherited through the taenite to tetrataenite  
895 transition). Secondly, the finest islands form in regions with  
896 lowest average Ni content. According to the most commonly  
897 used phase diagram for the Fe-Ni system (2), spinodal decom-  
898 position in these regions will initiate below 320 °C, in which  
899 case, islands will grow already within the tetrataenite stability  
900 field. The ability of magnetostatic interactions to influence  
901 the choice of c-axis orientation may be dramatically enhanced  
902 under these conditions, leading to the high degree of alignment  
903 through a process of self-organization.

904  
905 **A pathway to sustainable rare-earth free permanent magnets.**  
906 Synthetic rare-earth permanent magnets possess a wide range  
907 of applications. However, the scarcity and environmental im-  
908 pact of extracting rare earth elements, combined with the  
909 increasing demand for permanent magnets in the transport  
910 and renewable energy industries, means that sustainable al-  
911 ternatives need to be developed. Recently, a low-temperature  
912 nitrogen insertion and topotactic extraction (NITE) process  
913 has been developed to produce gram quantities of well ordered  
914 tetrataenite on short (i.e., laboratory) time scales (4), opening  
915 up the possibility of creating bulk synthetic tetrataenite for  
916 industrial use. Unfortunately, the magnetic coercivity of the  
917 material produced so far is below that needed to become a  
918 viable competitor to rare earth magnets (3). We observe much  
919 higher coercivity in localized regions of the meteorite cloudy  
920 zone, suggesting that there is scope to learn valuable lessons  
921 from nature as we strive to optimize the properties of synthetic  
922 tetrataenite (3, 33). The fine cloudy zone provides a suitable  
923 template for a sustainable permanent magnetic material. To  
924 achieve the maximum energy product it is necessary to max-  
925 imize both the saturation magnetization and the coercivity,  
926 both of which are observed using electron holography and  
927 XPEEM measurements in the fine cloudy zone (6, 7). The  
928 results of the present study suggest that these optimal con-  
929 ditions are achieved through a two-phase integrowth, where  
930 fully isolated islands of single-domain tetrataenite, with highly

aligned c axes, are coherently integrown with a metastable,  
partially ordered fcc paramagnetic matrix phase. Adapting  
the NITE process to produce a synthetic equivalent of the  
ordered matrix phase may be achievable by simple modifi-  
cation of the chemical composition of the pre-cursor Fe-Ni  
alloy. Mechanical mixing of synthetic tetrataenite and matrix  
nanoparticles, followed by low-temperature/high-pressure sin-  
tering in the presence of a strong magnetic field, would then  
provide a potential route to fabricating the same two-phase  
nanocomposite of tetrataenite and matrix phase in the labora-  
tory that nature took tens of millions of years to achieve on  
the meteorite parent body.

## Materials and Methods

**Sample preparation.** A section of the Tazewell meteorite was ac-  
quired from the Sedgwick Museum of Earth Sciences, University of  
Cambridge, Sample Number 16269. The same sample has previously  
been studied using electron holography, XPEEM and Mössbauer  
spectroscopy. Focused Ion Beam (FIB) milling with in-situ lift out  
was used to prepare samples for SPED, 3D-EDS and APT study,  
using an FEI Helios Dual-Beam FIB located in the Department of  
Materials Science and Metallurgy, University of Cambridge. Planar  
lamellae were prepared for SPED analysis using a basic in-situ TEM  
lamella preparation strategy. Final thinning and low kV cleaning  
were performed using a simplified version of the process outlined  
by Schaffer et al.(34). Tomography needle samples were fabricated  
using the block lift out approach described by Thompson et al.(35).  
For 3D-EDS and APT studies, blocks of the cloudy zone were lifted  
out and mounted in-situ onto tomography posts. For 3D-EDS, Cu  
tomography pins specifically designed to work with the Fischione In-  
struments 2050 full-tilt tomography sample holder were used. APT  
needles were mounted on a Cameca silicon lift-out coupon with  
36 posts. After welding with FIB-induced Pt deposition, needle  
geometries were formed using the procedure described by Larson et  
al.(36). Amorphous damage and Ga<sup>+</sup> implantation were removed  
using a 2 kV 100 pA polishing step. Several APT samples were  
fabricated with the tip axis parallel to either the [001] and [110]  
directions.

**3D-EDS.** Tomographic elemental mapping of the cloudy zone was  
performed using EDS on an FEI Tecnai Osiris 80-200, operating at  
an accelerating voltage of 200 kV. The high density of the Fe-Ni  
phase in the cloudy zone coupled with the variable thickness of  
a tomography needle geometry means that electron energy loss  
spectroscopy would suffer from multiple scattering events leading to  
non-quantitative results. Additionally, the four detector geometry  
of the Osiris instrument allows for rapid mapping of the sample,  
minimizing drift during the tilt series collection. Needle samples  
were loaded into the microscope using the Fischione Full tilt tomog-  
raphy holder. The EDS tilt series was collected every 5° from -70° to  
+75°. Elemental maps were collected in 246 nm x 327 nm region of  
interest on each of the tomography needles studied. With a pixel  
size of 2.46 nm, this resulted in the collection of 348,000 spectra.

**EDS structural and chemical quantification .** The EDS tilt series col-  
lected produces a statistically dense spectral data set. Integrated Fe  
and Ni spectral peak intensities were extracted from the EDS tilt  
series using spectral and machine learning tools in the open source  
Python library, Hyperspy (37). The spectra were denoised using  
Principal Component Analysis (PCA) and background subtracted  
before performing the intensity integration on the two respective  
elemental peaks. Intensity integration produces a greyscale tilt  
series for each element, where the peak intensity is proportional  
to the quantity of iron or nickel present in each pixel. The tilt  
series are aligned using TomoJ (38) and then reconstructed using an  
in-house compressive sensing (CS) algorithm (14, 15). The two re-  
sulting volumes were individually segmented using a random walker  
algorithm (39) and the resulting volumes were then compared to  
produce the structural morphologies for tetrataenite and matrix  
phases seen in Fig ?? and 2. Morphological analysis was performed



993 on the segmented Ni maps, as these provided a stronger contrast  
994 between tetraetaenite particles and the Fe<sub>3</sub>Ni matrix.  
995 Compressive sensing reconstruction preserves the boundaries  
996 and physical morphologies of features found in the tilt series. How-  
997 ever, the resulting greyscale values in the reconstruction are not  
998 linearly preserved. This results in uncertainty for the use of the  
999 reconstructed greyscale information for EDS quantification. We  
1000 addressed this uncertainty by developing an alternative method  
1001 for the quantification of 3D chemical data. First, the CS recon-  
1002 struction volume was segmented to label voxels as either Ni-rich  
1003 or Fe-rich. These chemical phase-specific sub-volumes were then  
1004 re-projected, using a discrete Radon transform in Sci-Kit Image  
1005 (Python), giving a phase-specific thickness map at each tilt angle  
1006 (40, 41). These thickness maps were registered to the raw EDS spec-  
1007 tra using image processing routines in Matlab. The reconstruction  
1008 problem was then recast as a system of linear equations, assuming  
1009 the observed spectra were a linear combination of signals arising  
1010 from the island and matrix phases. At each energy channel, the  
1011 348,000 simultaneous equations provided for an over-determined  
1012 problem to recover the coefficients corresponding to the intensity at  
1013 each energy channel for each of the two phases. Across the entire  
1014 spectrum, this analysis determined the characteristic EDS spectra  
1015 of the island and matrix phases based on the physically segmented  
1016 tomography volume (Fig. 3a). Cliff-Lorimer methods were used  
1017 to quantify the Fe - Ni ratio from the resulting EDS spectra. We  
1018 estimate the relative uncertainty in the ratio to be 2 at% based on  
1019 counting statistics. Errors associated with the first principles de-  
1020 rived k-factors provided by the instrument manufacturer have been  
1021 reported at approximately 8 at% using Co-K and Pt-L lines(42).  
1022 However, calculated Fe and Ni k-factors for K lines are known to  
1023 have particularly low errors (1-2%) (43) and are systematic and  
1024 of similar small magnitudes for Fe and Ni (and will therefore be  
1025 substantially reduced in relative composition). As such, we report  
1026 2 at% errors for EDS quantification in line with the x-ray counting  
1027 statistics.

1028 **Atom probe tomography.** Atom Probe Tomography experiments  
1029 were carried out on a LEAP 3000X-HR instrument (University  
1030 of Oxford), running in laser mode with a 532nm beam operating  
1031 at 0.4nJ. A small subset of samples were also run in voltage mode,  
1032 to confirm the accuracy of composition data, although the data  
1033 yield was lower from these. The specimen stage temperature was  
1034 kept at 55K for all experiments. Data was reconstructed using  
1035 IVAS software (3.6.12). The composition of the islands and matrix  
1036 was obtained by placing numerous cylindrical regions of interest at  
1037 least 1 nm away from the nearest interface, in order to avoid sampling  
1038 neighboring phases (SI Appendix, Fig. S3). The atoms inside mul-  
1039 tiple cylinders across multiple datasets were summed to produce  
1040 the average compositions quoted. The compositional variations of  
1041 either phase between different APT datasets were not statistically  
1042 significant.

1043 **Scanning Precession Electron Diffraction.** SPED allows for detailed  
1044 crystallographic mapping of a sample by precessing the electron  
1045 beam about a small angle. This results in diffraction patterns where  
1046 the dynamical effects of scattering are minimized and we are able  
1047 to record a larger region of the Ewald sphere(44). Here, the electron  
1048 probe is also scanned across the region of interest to map out the  
1049 crystallography of the sample. We use a Phillips CM300 FEGTEM  
1050 equipped with a Nanomegas ASTAR precession system. Using 300  
1051 kV spot 7, we collect diffraction patterns every 5 nm in the region  
1052 of interest. For the sample oriented along the [100] direction, we  
1053 collected three sets of maps, one from the coarse, medium and fine  
1054 regions of the cloudy zone (Fig. 5). Each [100] map collected was  
600 nm by 600 nm. For the lamella oriented along the [112] zone axis  
of the parent taenite crystal structure, a single map of 220 nm by  
235 nm was collected (Fig 6). Once the data was collected the maps  
were then analysis using cluster analysis, to determine the dominant  
diffraction patterns(45). This allows for physical diffraction patterns  
to be analyzed and orientation maps to be produced.

1055 **Cluster analysis.** Cluster analysis is the unsupervised (or semi-  
1056 supervised) identification and classification of groups of points  
1057 which lie close together in space(46). In the context of SPED,

for a highly coherent crystal structure such as reported here, statisti-  
cal decomposition methods such as PCA and Non-negative Matrix  
Factorization (NMF) can be misleading, because so many of the  
reflections are common to all of the diffraction patterns in the scan.  
As a result the common reflections tend to be grouped into one  
significant component and variations in the structure (often limited  
to a small number of weak reflections) are associated with the higher  
components. For clustering, each diffraction pattern is assigned a  
position in high-dimensional space according to how much of each  
component is associated with it. Clusters are formed by grouping  
points in this space that are 'close together' and the average diffrac-  
tion pattern for each signal is calculated from that point in space.  
Although it is common to define distances using a Euclidean metric,  
custom metrics can be used as well. The key advantage here is that  
the component corresponding to the common reflections can and  
will be included in the cluster representations, leading to meaningful  
diffraction patterns that nevertheless highlight the key structural  
differences between the different phases in the microstructure.

For Figure 6, initial decomposition was performed using NMF  
retaining 6 components. Clustering was performed using the  
Gustafson-Kessel variation of the probabilistic fuzzy c-means al-  
gorithm, searching for 5 clusters within the data. For Figure 5,  
a custom distance metric was used, based on the identification of  
peaks and their cumulative distance from those in other patterns.  
The distances themselves form clusters; these were used to derive  
the localization maps. The representative patterns were derived as  
the weighted means of the patterns in those areas.

**Micromagnetic modeling.** The Finite Element Method/Boundary El-  
ement Method (FEM-BEM) micromagnetics package MERRILL  
(Micromagnetic Earth Related Rapid Interpreted Language Labo-  
ratory) was used to solve for the magnetic scalar potential inside  
each particle and thereby calculate the demagnetizing energy of the  
system(47, 48). This approach avoids the need to discretize the  
non-magnetic volume outside the particle. Tetrahedral meshes of  
average 1.6 nm spacing were used for modeling magnetic behavior  
of the particles, to ensure that exchange interactions were being  
accounted for appropriately. Simulations were performed using an  
Apple iMac with a 3.4 GHz Intel i7 processor and 24 GB of RAM.  
Each particle was initialized with a random magnetization state.  
Simulations then minimized the total micromagnetic energy at each  
applied field and/or anisotropy value, using a conjugate gradient  
method adapted to micromagnetic problems. The total micromag-  
netic energy consists of summing the exchange, cubic anisotropy (in  
the case of taenite), uniaxial anisotropy (in the case of tetraetaenite),  
magnetostatic and demagnetizing energies. Material parameters  
used were appropriate for taenite at room temperature: saturation  
magnetization  $M_s = 1273$  kA/m, exchange constant  $A_{ex} = 1.13$   
 $\times 10^{-11}$  J/m, and cubic anisotropy with  $K_1 = 1$  kJ/m<sup>3</sup> (18, 19).  
Similarly, for tetraetaenite the room temperature parameters were:  
saturation magnetization  $M_s = 1390$  kA/m, exchange constant  $A_{ex}$   
 $= 1.13 \times 10^{-11}$  J/m, and cubic anisotropy with  $K_u = 1370$  kJ/m<sup>3</sup>  
(22). For simulations performed on individual islands, the anisotropy  
vector was set to be parallel to either the long, intermediate or short  
axis found for the best fit ellipsoid for each island. For simulations  
performed on all three interacting islands, the anisotropy vector  
was set to be parallel to either the long, intermediate or short axis  
found for just one of the islands. To simulate the transition from  
taenite to tetraetaenite, the cubic anisotropy was set to zero and  
the uniaxial anisotropy was increased from  $K_u = 0$  to  $K_u = 1370$   
kJ/m<sup>3</sup> in increments of 10 kJ/m<sup>3</sup>. The converged solution at each  
step was used as the starting configuration for the next step. The  
room-temperature single-domain threshold for a spherical particle  
of taenite was calculated using the size-scaling method of (47). A  
spherical mesh of 25 nm diameter and 1 nm element size was used,  
with size scaling factors varied from 0.5 to 3 and back again in steps

of 0.1. The threshold for a prolate ellipsoid particle with aspect ratio 2:1 was calculated using an initial mesh of diameter 50 x 25 nm and 1 nm element size. Quoted values are the lower limit obtained during the decreasing-size portion of the size hysteresis loop.

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- Scheinberg A, Fu RR, Elkins-Tanton LT, Weiss BP, Stanley S (2017) Magnetic Fields on Asteroids and Planetesimals in *Planetesimals*, eds. Elkins-Tanton LT, Weiss BP. (Cambridge University Press, Cambridge), pp. 180–203.
- Yang CW, Williams D, Goldstein J (1997) Low-temperature phase decomposition in metal from iron, stony-iron, and stony meteorites. *Geochimica et Cosmochimica Acta* 61(14):2943–2956.
- Lewis LH, et al. (2014) Inspired by nature: investigating tetraenaite for permanent magnet applications. *Journal of physics. Condensed matter : an Institute of Physics journal* 26(6):064213.
- Goto S, et al. (2017) Synthesis of single-phase L10-FeNi magnet powder by nitrogen insertion and topotactic extraction. *Scientific Reports* 7(1):1–7.
- Uehara M, Gattacceca J, Leroux H, Jacob D, van der Beek CJ (2011) Magnetic microstructures of metal grains in equilibrated ordinary chondrites and implications for paleomagnetism of meteorites. *Earth and Planetary Science Letters* 306(3-4):241–252.
- Bryson JF, Church NS, Kasama T, Harrison RJ (2014) Nanomagnetic intergrowths in Fe–Ni meteoritic metal: The potential for time-resolved records of planetesimal dynamo fields. *Earth and Planetary Science Letters* 388:237–248.
- Bryson JF, et al. (2014) Nanopaleomagnetism of meteoritic Fe–Ni studied using X-ray photoemission electron microscopy. *Earth and Planetary Science Letters* 396:125–133.
- Bryson JFJ, et al. (2015) Long-lived magnetism from solidification-driven convection on the pallasite parent body. *Nature* 517(7535):472–475.
- Nichols CI, et al. (2016) Pallasite paleomagnetism: Quiescence of a core dynamo. *Earth and Planetary Science Letters* 441:103–112.
- Bryson JF, Weiss BP, Harrison RJ, Herrero-Albillos J, Kronast F (2017) Paleomagnetic evidence for dynamo activity driven by inward crystallisation of a metallic asteroid. *Earth and Planetary Science Letters* 472:152–163.
- Harrison RJ, Bryson JFJ, Nichols CIO, Weiss BP (2017) Magnetic Mineralogy of Meteoritic Metal: Paleomagnetic Evidence for Dynamo Activity on Differentiated Planetesimals in *Planetesimals*, eds. Elkins-Tanton LT, Weiss BP. (Cambridge University Press, Cambridge), pp. 204–223.
- Elkins-Tanton LT, Weiss BPPB (2017) *Planetesimals : early differentiation and consequences for planets*. p. 381.
- Goldstein J, Scott ERD, Winfield T, Yang J (2013) Thermal histories of group IAB and related iron meteorites and comparison with other groups of irons and stony iron meteorites. *Lunar and Planetary Science Conference* 44:3–4.
- Saghi Z, et al. (2011) Three-dimensional morphology of iron oxide nanoparticles with reactive concave surfaces. A compressed sensing-electron tomography (CS-ET) approach. *Nano letters* 11(11):4666–73.
- Leary R, Saghi Z, Midgley PA, Holland DJ (2013) Compressed sensing electron tomography. *Ultramicroscopy* 131:70–91.
- Doube M, et al. (2010) BoneJ: Free and extensible bone image analysis in ImageJ. *Bone* 47(6):1076–9.
- Carriero A, et al. (2014) Altered lacunar and vascular porosity in osteogenesis imperfecta mouse bone as revealed by synchrotron tomography contributes to bone fragility. *Bone* 61:116–24.
- Hýřch MJ, et al. (2003) Vortex Flux Channeling in Magnetic Nanoparticle Chains. *Physical Review Letters* 91(25):257207.
- Gehrmann B (2005) Nickel–iron alloys with special soft magnetic properties for specific applications. *Journal of Magnetism and Magnetic Materials* 290:1419–1422.
- Rancourt D, et al. (1999) Experimental proof of the distinct electronic structure of a new meteoritic Fe–Ni alloy phase. *Journal of Magnetism and Magnetic Materials* 191(3):L255–L260.
- Blukis R, Harrison RJ (2017) A high spatial resolution synchrotron Mössbauer study of the Tazewell III CD and Esquel pallasite meteorites. *Meteoritics & Planetary Science*.
- Néel L, Pauleve J, Pauthenet R, Laugier J, Dautreppe D (1964) Magnetic Properties of an Iron–Nickel Single Crystal Ordered by Neutron Bombardment. *Journal of Applied Physics* 35(3):873–876.
- Leroux H, Doukhan JC, Perron C (2000) Microstructures of metal grains in ordinary chondrites: Implications for their thermal histories. *Meteoritics and Planetary Science* 35:569–580.
- Goldstein J, Scott E, Chabot N (2009) Iron meteorites: Crystallization, thermal history, parent bodies, and origin. *Chemie der Erde - Geochemistry* 69(4):293–325.
- Reuter KB, Williams DB, Goldstein JI (1988) Low temperature phase transformations in the metallic phases of iron and stony-iron meteorites. *Geochimica et Cosmochimica Acta* 52(3):617–626.
- Rout SS, et al. (2017) Atom-probe tomography and transmission electron microscopy of the kamacite-taenite interface in the fast-cooled Bristol IVA iron meteorite. *Meteoritics & Planetary Science* 52(12):2707–2729.
- Miller M, et al. (1989) An Atom Probe Field-Ion Microscopy Study Of Phase Separation In The Twin City And Santa Catharina Meteorites. *Le Journal de Physique Colloques* 50(11):413–418.
- Watanabe D, Ogawa S (1956) On the Superstructure of the Ordered Alloy Cu3Pd - I. Electron Diffraction Study. *Journal of the Physical Society of Japan* 11(3):226–239.
- Fujiwara K (1957) On the Period of Out-of-step of Ordered Alloys with Anti-phase Domain Structure. *Journal of the Physical Society of Japan* 12(1):7–13.
- Mishin Y, Mehl M, Papaconstantopoulos D (2005) Phase stability in the Fe–Ni system: Investigation by first-principles calculations and atomistic simulations. *Acta Materialia* 53(15):4029–4041.
- Rancourt D, Scorzelli RB (1995) Low-spin  $\gamma$ -Fe-Ni( $\gamma$ LS) proposed as a new mineral in Fe-Ni-bearing meteorites: epitaxial intergrowth of  $\gamma$ LS and tetraenaite as a possible equilibrium state at approximately 20–40 at% Ni. *Journal of Magnetism and Magnetic Materials* 150(1):30–36.
- Berndt T, Muxworthy AR, Fabian K (2015) Does size matter? Statistical limits of paleomagnetic field reconstruction from small rock specimens. *Journal of Geophysical Research: Solid Earth* 121(1):15–26.
- Makino A, et al. (2015) Artificially produced rare-earth free cosmic magnet. *Scientific reports* 5:16627.
- Schaffer M, Schaffer B, Ramasse Q (2012) Sample preparation for atomic-resolution STEM at low voltages by FIB. *Ultramicroscopy* 114:62–71.
- Thompson K, et al. (2007) In situ site-specific specimen preparation for atom probe tomography. *Ultramicroscopy* 107(2-3):131–9.
- Larson D, et al. (1999) Field-ion specimen preparation using focused ion-beam milling. *Ultramicroscopy* 79(1):287–293.
- (2017) hyperspy/hyperspy: HyperSpy 1.1.2.
- Messaoudil C, Boudier T, Sorzano C, Marco S (2007) TomoJ: tomography software for three-dimensional reconstruction in transmission electron microscopy. *BMC Bioinformatics* 8(1):288.
- Grady L (2006) Random walks for image segmentation. *IEEE Transactions on Pattern Analysis and Machine Intelligence* 28(11):1768–1783.
- Zhu Gz, Radtke G, Botton GA (2012) Bonding and structure of a reconstructed (001) surface of SrTiO3 from TEM. *Nature* 490(7420):384–387.
- Collins SM, Fernandez-Garcia S, Calvino JJ, Midgley PA (2017) Sub-nanometer surface chemistry and orbital hybridization in lanthanum-doped ceria nano-catalysts revealed by 3D electron microscopy. *Scientific Reports* 7(1):5406.
- MacArthur KE, et al. (2016) Quantitative Energy-Dispersive X-Ray Analysis of Catalyst Nanoparticles Using a Partial Cross Section Approach. *Microscopy and Microanalysis* 22(01):71–81.
- Metcalfe E, Broomfield JP (1984) Determination Of Cliff-Lorimer K Factors For A Hitachi H700h 200 Kv Scanning Transmission Electron Microscope. *Le Journal de Physique Colloques* 45(C2):2–407.
- Vincent R, Midgley P (1994) Double conical beam-rocking system for measurement of integrated electron diffraction intensities. *Ultramicroscopy* 53(3):271–282.
- Martineau BH, Johnstone DN, Einsle JF, Midgley PA, Eggeman AS (2017) Data clustering and scanning precession electron diffraction for microanalysis. *Microscopy and Microanalysis* 23(S1):116–117.
- Koch I (2013) *Analysis of multivariate and high-dimensional data*. p. 504.
- Ó Conbhuí P, et al. (2018) MERRILL: Micromagnetic Earth Related Robust Interpreted Language Laboratory. *Geochemistry, Geophysics, Geosystems* 19(4):1080–1106.
- Einsle JF, et al. (2016) Multi-scale three-dimensional characterization of iron particles in dusty olivine: Implications for paleomagnetism of chondritic meteorites. *American Mineralogist* 101(9):2070–2084.