

1 **Quantitative Relationships between Basalt Geochemistry, Shear**
2 **Wave Velocity and Asthenospheric Temperature Beneath**
3 **Western North America**

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9 **Key Points:**

- 10 • Correlation between basalt geochemistry and shear wave velocities
11 • Mantle potential temperatures of 1340 ± 20 °C
12 • Uplift and magmatism generated by modest thermal anomalies beneath thinned plate

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Abstract

Western North America has an average elevation that is ~ 2 km higher than cratonic North America. This difference coincides with a westward decrease in average lithospheric thickness from ~ 240 km to < 100 km. Tomographic models show that slow shear wave velocity anomalies lie beneath this region, coinciding with the pattern of basaltic magmatism. To investigate relationships between magmatism, shear wave velocity and temperature, we analyzed a suite of > 260 basaltic samples. Forward and inverse modeling of carefully selected major, trace and rare earth elements were used to determine melt fraction as a function of depth. Basaltic melt appears to have been generated by adiabatic decompression of dry peridotite with asthenospheric potential temperatures of 1340 ± 20 °C. Potential temperatures as high as 1365 °C were obtained for the Snake River Plain. For the youngest (i.e. < 5 Ma) basalts with a sub-plate geochemical signature, there is a positive correlation between shear wave velocities and trace element ratios such as La/Yb. The significance of this correlation is explored by converting shear wave velocity into temperature using a global empirical parameterization. Calculated temperatures agree with those determined by inverse modeling of rare earth elements. We propose that regional epeirogenic uplift of western North America is principally maintained by widespread asthenospheric temperature anomalies lying beneath a lithospheric plate, which is considerably thinner than it was in Late Cretaceous times. Our proposal accounts for the distribution and composition of basaltic magmatism and is consistent with regional heatflow anomalies.

Plain Language Summary

Marine fossils from western North America show that this region, which includes the states of Arizona, Colorado and Utah, was below sea level 80 million years. Since that time, large-scale uplift of about 2 kilometers has occurred. This uplift coincides with massive outpourings of lava, the youngest of which occurred 1000 years ago at Sunset Crater outside Flagstaff and was witnessed by Native Americans. Seismic (i.e. acoustic) images of the deep structure beneath western North America show that the tectonic plate is only about 50 to 100 kilometers thick— much thinner than beneath the Great Plains further east. Here, we use chemical analyses of different lavas to calculate the temperature of mantle rocks that lie beneath the tectonic plate and which generated the lavas. Remarkably, this temperature is consistent with the temperature that is estimated from the speed of sound through these mantle rocks. We con-

44 clude that the whole of Western North America is supported by a combination of moderately
45 warm mantle rocks and a thinner than expected tectonic plate.

46 **1 Introduction**

47 It is recognized that convective circulation of the Earth's mantle generates and maintains
48 some fraction of surface topography, referred to as dynamic topography [Anderson *et al.*, 1973;
49 Parsons and Daly, 1983; Hager and Richards, 1989; Lithgow-Bertelloni and Silver, 1998; Moucha
50 *et al.*, 2008; Flament *et al.*, 2013]. A significant corollary is that the spatial and temporal evo-
51 lution of dynamic topography can help to constrain the behavior of mantle convection on ge-
52 ologic timescales. Western North America constitutes a dramatic example where large-scale
53 regional uplift appears to be supported by sub-plate processes [Ashwal and Burke, 1989]. Rem-
54 nants of extensive marine deposits, such as the Mancos shale of the Mesaverde Group that crops
55 out across Wyoming, Utah and Colorado, demonstrate that a Mid-Cretaceous seaway origi-
56 nally connected the Beaufort Sea and the Gulf of Mexico [Roberts and Kirschbaum, 1995].
57 This observation implies that present-day regional elevation principally grew during Cenozoic
58 times. Thermochronologic observations from the Grand Canyon area combined with clumped
59 isotopic studies of Paleogene lacustrine deposits suggest that this elevation grew in several dis-
60 crete stages [e.g. ~ 70 and ~ 30 Ma; Moucha *et al.*, 2009; Liu and Gurnis, 2010]. This infer-
61 ence is supported by inverse modeling of drainage networks and by the history of sedimen-
62 tary flux into basins, such as the Gulf of Mexico [Galloway *et al.*, 2011; Roberts *et al.*, 2012].

63 A region encompassing the Basin and Range Province, the Snake River Plain and the
64 Colorado Plateau sits ~ 2 km higher than cratonic North America (Figure 1a). Crustal thick-
65 ness of the Colorado Plateau is ~ 45 km, which is similar to that of the Great Plains [Shen and
66 Ritzwoller, 2016]. Given that their respective crustal velocities and inferred densities are sim-
67 ilar, this difference in elevation cannot easily be explained by crustal isostasy [e.g. Sheehan
68 *et al.*, 1995; Spencer, 1996; Shen and Ritzwoller, 2016]. Instead, regional elevation of west-
69 ern North America is probably supported either by a thinner lithosphere, by convective up-
70 welling of anomalously hot asthenospheric mantle, or by some combination of both mecha-
71 nisms [e.g. Bradshaw *et al.*, 1993; Hyndman and Currie, 2011; Becker *et al.*, 2013; Afonso *et al.*,
72 2016]. Surface wave tomographic models show that beneath most of western North America
73 the continental lithosphere is less than 120 km thick, whereas beneath the interior craton the
74 lithosphere is about 240 km thick [Priestley and McKenzie, 2013]. Receiver function analy-
75 ses suggest that a thin lithosphere underlies western North America since Sp conversions have

76 been reported at depths of 60–80 km [Lekić and Fischer, 2014; Hopper et al., 2018]. The rea-
77 son for such a large difference in lithospheric thickness across the continent remains obscure.
78 It has been proposed that mechanical thinning of the lithosphere is somehow linked to shal-
79 low subduction of the Farallon slab [Humphreys, 1995; Spencer, 1996]. More speculatively,
80 delamination of lithospheric mantle following slab hydration has been invoked [Humphreys
81 et al., 2003]. A long wavelength (i.e. 500–4000 km) positive free-air gravity anomaly of +40
82 mGal is centered on the Yellowstone area [Figure 1b; Bruinsma et al., 2014]. This cruciform
83 anomaly reaches across most of western North America, coinciding with the distribution of
84 Cenozoic magmatism and with the planform of a regional heatflow anomaly [Pollack et al.,
85 1993].

86 Tomographic models show that there are large negative shear wave velocity anomalies
87 beneath most of western North America at depths of 50–500 km [Figure 1c; e.g. Crow et al.,
88 2010; Schmandt and Humphreys, 2010; Obrebski et al., 2011; Shen et al., 2013; Burdick et al.,
89 2014; Schaeffer and Lebedev, 2014; Shen and Ritzwoller, 2016]. Different models disagree on
90 the detailed horizontal and vertical structure of these anomalies and on the amplitudes of phase
91 velocity measurements, but the general pattern is both consistent and striking. The slowest ve-
92 locity anomalies occur beneath the Yellowstone area, consistent with a protruding finger of hot
93 material that extends along the Snake River Plain towards cratonic lithosphere located further
94 east. A similarly slow finger lies beneath the Rio Grande Rift and a horseshoe-shaped anomaly
95 fringes the Colorado Plateau. In conjunction with positive free-air gravity anomalies, these re-
96 gional velocity anomalies are consistent with the presence of a shallow convective upwelling
97 beneath the North American plate.

98 Magmatism of western North America is spatially distributed over a region that broadly
99 coincides with elevated regional topography. Volcanism reaches far into the continental plate
100 and reveals age progressions that can largely be accounted for by horizontal plate translation
101 over a relatively stationary source of melting within the asthenospheric mantle. Basaltic vol-
102 canism commenced at ~80 Ma and its subsequent temporal evolution has several distinct phases,
103 the most significant of which are a dramatic increase in the volume of magmatism at ~40 Ma
104 and a marked switch from a lithospheric to an asthenospheric signature at ~5 Ma [Fitton et al.,
105 1991; Kempton et al., 1991]. The spatial distribution of Neogene basaltic volcanism closely
106 coincides with the pattern of shear wave velocity anomalies (Figure 1c).

107 Two classes of models have been proposed to account for these large-scale observations.
108 One school of thought invokes an upwelling mantle plume located beneath present-day Yel-
109 lowstone with secondary plumes triggering volcanism further south [e.g. *Leat et al.*, 1991; *Par-*
110 *sons et al.*, 1994; *Saltus and Thompson*, 1995; *Camp and Hanan*, 2008; *Hanan et al.*, 2008; *Moucha*
111 *et al.*, 2009; *Pierce and Morgan*, 2009; *Huang et al.*, 2015]. An alternative view is that man-
112 tle material flows off the edge of cratonic lithosphere and around complex remnants of the sink-
113 ing Farallon slab, triggering shallower convective upwelling [e.g. *Roy et al.*, 2009; *van Wijk*
114 *et al.*, 2010; *James et al.*, 2011; *Levander et al.*, 2011; *Levander and Miller*, 2012; *Refayee et al.*,
115 2013; *Ballmer et al.*, 2015]. The principal difference between these plume and flow models
116 centers on the temperature of the underlying asthenospheric mantle. In a plume model, man-
117 tle material is expected to be hotter than in flow models where convection could be edge-driven
118 or generated by shallow return flow of mantle material. In this contribution, our principal aim
119 is to shed some light on these different hypotheses by calculating depths and temperatures of
120 mantle melting beneath western North America from the geochemistry of mafic igneous rocks.
121 A significant part of our strategy is to combine a quantitative geochemical approach with the
122 results of shear wave tomographic studies. By integrating geochemical and geophysical ob-
123 servations, we hope to illuminate aspects of upper mantle processes that may have influenced
124 the spatial and temporal evolution of western North America.

125 **2 Basaltic Magmatism**

126 In order to isolate the source of intracontinental volcanism, it is important to identify
127 and remove crustal and/or mantle lithospheric contamination so that we can focus attention
128 on the most primitive (i.e. asthenospheric) melts that contain information about initial melt-
129 ing conditions. Many contributions highlight the bimodal nature of potential source compo-
130 sitions beneath western North American basalts. For example, Hf and Nd isotopes from some
131 basalts of the Western Transition Zone that fringes the Colorado Plateau, and from the Col-
132 orado Plateau itself, suggest some overlap with the oceanic mantle array. The isotopic com-
133 position of other samples from the same volcanic fields is also consistent with the presence
134 of components of Paleoproterozoic peridotitic lithosphere [*Reid et al.*, 2012]. Given our aims,
135 we are principally concerned with basaltic rocks that are as compositionally similar to ocean
136 island basalts (OIBs) as possible. It is therefore appropriate to filter out contributions from meta-
137 somatized lithospheric mantle, or from subduction-influenced magmatism, where hydrous melt-
138 ing and contamination with sedimentary material can be significant.

2.1 Sample Selection and Screening

We have assembled a substantial and comprehensive database of Cenozoic mafic igneous rocks. This database comprises >1000 analyses from the western North American volcanic and intrusive rock catalogue (NAVDAT; <http://www.navdat.org>), 215 samples collected by *Fitton et al.* [1991], 29 samples from the Western Transition Zone generously provided by T. Plank [written communication, 2015; *Plank and Forsyth*, 2016], as well as 65 samples collected across Arizona and Colorado during December 2014 and April 2015, respectively. The geographic distribution of all analyses and samples is summarized in Figure 2. Samples collected in the field and those selected from the catalogue of *Fitton et al.* [1991] were analyzed for trace and rare earth elements (REEs) using inductively coupled plasma-mass spectrometry (ICP-MS). The 65 samples collected from Arizona and Colorado were analyzed for major and trace elements using X-ray fluorescence (XRF). Based on their respective precisions, XRF measurements of V, Y, Zr, Nb, Cr, Cu, Sc, Ni, Sr and Zn together with ICP-MS measurements of Ba, REEs, Rb, Th and Pb were used for further study. Detailed analytical procedures and data tables are provided in Supplementary Information.

The combined database is sub-divided into ten geographic provinces shown in Figure 2: Snake River Plain (SRP), Great Plains (GP), Eastern Transition Zone (ETZ), Southern Transition Zone (STZ), Sentinel Plain (SE), Northern Basin and Range (NBR), Western Transition Zone (WTZ), Basin and Range (BR), Rio Grande Rift (RGR), and Colorado Plateau (CP). Samples from the Cascades and from the western Great Basin were excluded due to their proximity to the present-day subduction zone. Compositions of remaining samples range from basaltic andesite to picobasalt and basanite. The majority of these samples fall within the basaltic field. Samples from the Colorado Plateau are the most enriched in terms of alkaline and incompatible elements, while samples from the Snake River Plain have the most depleted signatures. Major and trace element contents were used to identify the most primitive, least fractionated rocks from each province. A cut-off of MgO ≥ 7 wt% was deemed appropriate in order to minimize the effects of pyroxene and plagioclase fractionation (Figure 3). For provinces with large numbers of high MgO samples, it was feasible to adopt a more severe (i.e. more primitive) cut-off value. For example, samples with MgO ≥ 8 wt% were selected from the Snake River Plain and from the Southern Transition Zone.

Rigorous screening is used to exclude samples that are obviously contaminated by interaction with lithospheric melts. Trace element composition was used to identify samples de-

171 rived from the asthenosphere, following an approach similar to that described by *Fitton et al.*
 172 [1991]. Thus, samples were deemed to be of asthenospheric origin if their La, Ba and Nb com-
 173 positions fall within the fields expected for global suites of mid-ocean ridge basalts (MORB)
 174 and/or of OIBs [Figure 3d; *Stracke et al.*, 2005; *Willbold and Stracke*, 2006]. Partition coef-
 175 ficients show that these particular elements are highly incompatible in olivine and remain within
 176 the liquid phase [i.e. $D_{La} = D_{Nb} = 5 \times 10^{-4}$, $D_{Ba} = 5 \times 10^{-6}$; *Salter and Stracke*, 2004].
 177 Their relative abundances are largely insensitive to fractionation processes and, instead, are
 178 broadly reflective of source composition and of melting conditions. Arc magmas are typically
 179 enriched in large ion lithophiles, such as Ba, and depleted in Nb relative to MORB or OIB.
 180 In contrast, OIB, MORB and subduction zone melts have similar concentrations of La for a
 181 given melt fraction [e.g. *Pearce*, 1982]. Consequently, screening on the basis of La/Ba and La/Nb
 182 ratios helps to identify the chemical influence of subducting slabs (Figure 3c). It is possible,
 183 however, that volatile-rich fluids that are not produced by a slab could pass this form of screen-
 184 ing. Where available, Sr and Nd isotope ratios were used to check the efficacy of the screen-
 185 ing process for identifying only MORB- or OIB-type compositions.

186 This screening strategy reduces the combined database to 177 acceptable samples: 12
 187 out of 272 for Snake River Plain; 8 out of 76 for Great Plains; 13 out of 32 for the Eastern
 188 Transition Zone; 18 out of 42 for the Southern Transition Zone; 5 out of 9 for Sentinel Plain;
 189 2 out of 11 for Northern Basin and Range; 40 out of 102 for Western Transition Zone; 7 out
 190 of 28 for Basin and Range; 65 out of 150 for Rio Grande Rift; and 7 out of 14 for Colorado
 191 Plateau. 77 of the total number of the chosen samples and analyses are taken from *White et al.*
 192 [2004], *Thompson et al.* [2005], *Leeman et al.* [2009] and *Plank and Forsyth* [2016]. A total
 193 of 100 samples were extracted and analyzed from the catalogue of *Fitton et al.* [1991] and from
 194 the inventory collected during the two field campaigns. Two of these samples duplicate those
 195 of *Leeman et al.* [2009] and so the average composition was used. The majority of these sam-
 196 ples are younger than 5 Ma [*Fitton et al.*, 1991]. Trace element values for the screened database
 197 are summarized in Figure 4.

198 2.2 Melting Model and its Application

199 Relative abundances of incompatible elements can be used to determine primary melt-
 200 ing conditions provided that a series of assumptions are made about the nature of the source
 201 region and about the process of melt extraction. Here, trace elements from the screened database,
 202 notably the REEs, are used to determine the degree of mantle melting as a function of depth.

203 As a result of their differences in compatibility (i.e. partitioning behavior between solid and
204 liquid phases), REEs are sensitive to the cumulative amount of melting and to the relative pro-
205 portions of melting that occur within the garnet and spinel stability fields. The cumulative vol-
206 ume of generated melt is strongly influenced by the temperature at the time of melting. Higher
207 temperatures give rise to a larger solidus overstep and so produce larger melt fractions [*McKenzie*
208 *and Bickle*, 1988]. Ratios of light REEs to medium or heavy REEs, such as La/Sm or La/Yb,
209 are negatively correlated with melt fraction due to different degrees of incompatibility of these
210 elements. The partition coefficients of La, Sm and Yb in the mantle at 2 GPa are 6.6×10^{-3} ,
211 6×10^{-2} and 1.15×10^{-1} , respectively [*Salter and Stracke*, 2004]. The smaller the melt
212 fraction, the larger the differences in behavior of the relatively more compatible Sm and Yb
213 with respect to the more incompatible La. Hence, large values of La/Sm and La/Yb are as-
214 sociated with small melt fractions. Depth of melting is determined relative to the stability fields
215 of spinel and garnet. For example, large ratios of Sm/Yb indicate melting of garnet peridotite
216 since a greater proportion of Yb is retained within garnet of the solid phase and does not par-
217 tition into the melt phase.

218 The exact depth of the spinel-garnet transition is a subject of ongoing debate. Until re-
219 cently, it was thought that this transition was highly sensitive to temperature such that the greater
220 the temperature, the deeper and narrower the transition zone should be. A conservative esti-
221 mate of this temperature sensitivity is 40 ± 10 °C/kbar [*Klemme and O'Neill*, 2000; *Walter*
222 *et al.*, 2002]. However, *Green et al.* [2012] and *Jennings and Holland* [2015] argue that the pres-
223 sure of the garnet-spinel transition was overestimated in previous experimental studies, largely
224 due to the simplicity of the phase systems used (i.e. Mg-Al-Si rather than Ca-Mg-Al-Si). *Jen-*
225 *nings and Holland* [2015] demonstrate that their model compares well to existing studies, pro-
226 vided that simplifying corrections are applied, notably allowing for Ca activity within garnet.
227 In their thermodynamic calculations, the depth and thickness of the transition zone at, or above,
228 the solidus is not especially sensitive to temperature. Instead, a variation of up to 5 kbar to-
229 ward lower pressures for a temperature range of 880–1300 °C is found. By contrast, increas-
230 ing concentrations of Cr and Fe³⁺ within peridotite tend both to increase the thickness of the
231 transition and to cause a shift to greater pressures due to the greater stability of spinel [*Klemme*
232 *and O'Neill*, 2000; *Jennings and Holland*, 2015]. For the KLB-1 peridotite, *Jennings and Hol-*
233 *land* [2015] calculate the pressure at the top and bottom of the spinel-garnet transition where
234 it intersects the solidus and obtain values of 21.4 and 21.7 kbar, respectively.

235 An inverse modeling strategy enables REE compositions to be fitted by varying melt frac-
236 tion as a function of depth for a specified source composition. Here, we apply the INVMEL-
237 v12.0 algorithm, the first version of which was originally described by *McKenzie and O’Nions*
238 [1991]. This approach is especially sensitive to the relative amount of melting that occurs within
239 the garnet and spinel stability fields. Distributions of REE compositions are matched by as-
240 suming isentropic decompression melting of a dry aluminous peridotite mantle source. For a
241 given inverse model, the depth to the top of the melting region, the depth interval, and an ini-
242 tial distribution of melt fraction as a function of depth can be specified. An optimal fit is ob-
243 tained by iteratively computing the point-and-depth average composition using a continuous
244 melting curve. The root mean squared (rms) misfit between observed and calculated REE dis-
245 tributions is minimized using a conjugate direction search routine called Powell’s algorithm
246 [Press *et al.*, 1992]. When the optimal melt fraction as a function of depth is determined, the
247 composition of other trace and major elements can be predicted by forward modeling. In gen-
248 eral, melting interval and total melt fraction are the most reliable outputs of this inverse mod-
249 eling approach. The calculated melt fraction distribution is compared with a set of predicted
250 isentropic curves to estimate the potential temperature of melting, where potential tempera-
251 ture is calculated at the Earth’s surface using an adiabatic gradient of $0.48 \text{ }^\circ\text{C km}^{-1}$. These
252 curves are determined for different potential temperatures using a decompression melting model
253 with a dry solidus parameterization described by *Katz et al.* [2003]. For our purposes, an en-
254 tropy of melting, $J = 400 \text{ J K}^{-1} \text{ kg}^{-1}$, is used in order to be self-consistent [*Kojitani and*
255 *Akaogi*, 1997]. If calculated melt fraction distributions deviate from an isentropic path, a range
256 of potential temperatures is gauged from the deepest and shallowest melt fractions.

257 The INVMEL algorithm exploits partition coefficients calculated using the lattice strain
258 model of *Blundy and Wood* [2003]. We assume that the pressures at the top and the bottom
259 of the spinel-garnet transition are 21 and 24 kbar, which correspond to depths of 63 and 72
260 km, respectively. This transition zone is thicker than that proposed by *Jennings and Holland*
261 [2015] in order to stabilize the inverse algorithm— a difference that does not materially af-
262 fect our results. The combination of a different solidus parameterization together with differ-
263 ent depth and thickness of the spinel-garnet transition zone compared with *McKenzie and O’Nions*
264 [1991] means that temperature estimates at any given depth are generally 30–50 $^\circ\text{C}$ lower. Thus
265 whilst cumulative melt fractions are generally comparable, our results yield minimum estimates
266 of both temperature and lithospheric thickness.

267 We assume that asthenospheric mantle can be regarded, to a first approximation, as ho-
268 mogeneous beneath western North America. Modeling is generally carried out using a mix-
269 ture of primitive and depleted MORB mantle. Source composition is gauged using ε_{Nd} val-
270 ues of samples from each volcanic field published in the NAVDAT catalogue. For example,
271 if $\varepsilon_{Nd} = 10$ a depleted mantle source is used and if $\varepsilon_{Nd} = 0$ a primitive mantle source is
272 used [White and McKenzie, 1995]. An important exception is the Colorado Plateau which is
273 characterized by high concentrations of the most incompatible elements that cannot easily be
274 fitted by inverse modeling. Additional enrichment of the source region by adding a small frac-
275 tion of melt generated within the garnet stability field was required to optimize the fit between
276 observed and calculated concentrations. Published isotopic measurements and mantle sources
277 used for inverse modeling are summarized in Table 1. Compositions of depleted and primi-
278 tive mantle, as well as the small fraction of melt generated within the garnet stability field are
279 provided in the Supplementary Information.

280 Judicious sample selection is an important prerequisite since only near-fractional melt-
281 ing of a uniform dry peridotite source is accounted for during inverse modeling. Once sam-
282 ples have been selected, amounts of olivine fractionation are determined using the differences
283 between observed MgO and FeO concentrations and those calculated for a primitive melt of
284 the specified source composition [McKenzie and O’Nions, 1991]. In this way, the final melt
285 fraction distribution is appropriately corrected. This approach also holds, within limits, for clinopy-
286 roxene fractionation. However, it cannot be used to correct for the crystallisation of non-Mg/Fe
287 bearing phases such as plagioclase, which is the reason why sample selection is so important.
288 No corrections are applied for contamination by crust and/or lithospheric mantle (i.e. melt-
289 ing is assumed to be generated from a homogeneous asthenospheric source). The effect of volatiles,
290 or of a non-peridotitic source composition, on melting beneath western North America is sep-
291 arately addressed. This general strategy is used to determine the depth and degree of melting
292 beneath 26 volcanic fields from ten geographic provinces. Average major, trace and rare earth
293 element compositions for these provinces are provided in Table 1.

294 **2.3 Results**

295 Inverse models for each province are shown in Figures 5 and 6 and summarized in Ta-
296 ble 2. The observed REE concentrations are fitted such that the rms misfit between observed
297 and calculated ratios with respect to the source is < 0.9 . Forward-modeled fits to other trace
298 element concentrations are largely within the degree of uncertainty for geochemical compo-

299 sitions with minor exceptions (Figures 5 and 6b,e,h,m,p). For all provinces, more compatible
300 elements are better matched than highly incompatible ones. We stress that only fractionation
301 of olivine has been formally corrected for and so hydrous phases (e.g. amphibole, phlogopite)
302 that are observed in basalts from the Hopi Buttes volcanic field of the Colorado Plateau could
303 account for depletion of Na, Rb, P and K.

304 Cumulative melt fraction and depth of melting systematically vary across western North
305 America. Volcanic fields from the Snake River Plain represent the largest degrees of melting
306 ($\sim 10\%$) at the shallowest melting depths (~ 50 km), corresponding to the highest potential
307 temperatures for this region (~ 1365 °C). In contrast, basalts from the Colorado Plateau have
308 the smallest melt fractions ($\sim 1\%$) that formed at the greatest depths (> 62 km), correspond-
309 ing to the lowest potential temperatures (~ 1320 °C). These differences between Snake River
310 Plain and Colorado Plateau basalts are significant and reflect different concentrations of light
311 REEs relative to heavy REEs. Analyses from the Great Plains, from the Eastern and South-
312 ern Transition Zones, and from Sentinel Plain yield melt fractions of $\sim 6\text{--}7\%$ at depths of 53–
313 84 km, corresponding to potential temperatures of $\sim 1350\text{--}1360$ °C. Analyses from the Basin
314 and Range, from the Western Transition Zone and from the Rio Grande Rift yield 2–4 % melt-
315 ing at depths between 54 and 74 km, equivalent to potential temperatures of 1320–1330 °C.
316 Degrees of olivine fractionation generally vary between 16 and 31%. Note that in all cases,
317 most melt production occurs either within the spinel-only stability field or within the spinel-
318 garnet transition zone.

319 Errors associated with these results can be gauged by considering a combination of ran-
320 dom and systematic uncertainties [White *et al.*, 1992; Brodie *et al.*, 1994]. First, the typical stan-
321 dard deviation of geographically averaged sample concentrations is less than 10%, which gives
322 rise to an uncertainty in cumulative melt fraction of less than 2%. Secondly, the top of the melt-
323 ing column can be adjusted in each case by $\pm 2\text{--}5$ km, which contributes an uncertainty in cu-
324 mulative melt fraction of less than 2%. Thirdly, the depth and thickness of the spinel-garnet
325 transition can be varied by ± 5 km and ± 10 km, respectively. These variations yield a com-
326 bined uncertainty in cumulative melt fraction of less than 3%. It is important to emphasise that
327 more significant excursions in the values of the top of the melting column and in the depth
328 and thickness of the transition lead to degraded fits to the observed REE concentrations. Un-
329 certainties associated with source composition constrained by ε_{Nd} act to change the depth to
330 the top of the melting column by < 10 km, which yields an uncertainty in cumulative melt

331 fraction of less than 5%. Together, these estimates of the range of uncertainties for cumula-
332 tive melt fraction generate potential temperature variations of $\pm 10\text{--}30$ °C.

333 A significant outcome of our study is that the bulk of melting beneath western North Amer-
334 ica occurred close to the garnet-spinel transition. Since this transition is fixed at a depth range
335 of 63–72 km, significant melting is required to occur shallower than ~ 70 km. Mafic compo-
336 sitions are consistent with mantle potential temperatures of 1320–1365 °C. The highest tem-
337 peratures are obtained for the youngest Snake River Plain samples whilst those from the Col-
338 orado Plateau do not record potential temperatures that are significantly different to that of am-
339 bient asthenospheric mantle. Previous inverse modeling yielded potential temperatures of ~ 1400 °C
340 at depths of 60–100 km beneath the Snake River Plain and beneath the Rio Grande Rift [*White*
341 *and McKenzie*, 1995; *Thompson et al.*, 2005]. It has been suggested that the top of the melt-
342 ing region corresponds to the base of the lithospheric plate [*McKenzie and O’Nions*, 1991; *White*
343 *et al.*, 1992]. The average plate thickness inferred by inverse modeling is 55 ± 10 km. Beneath
344 the Snake River Plain, melts are generated at depths as shallow as 48 km and beneath the Col-
345 orado Plateau, melts are generated at depths of > 62 km. Although our results suggest that
346 basaltic melting is generated within the asthenospheric mantle layer immediately beneath the
347 lithospheric plate, elevated $^3\text{He}/^4\text{He}$ ratios from hot-spring gases on the Snake River Plain and
348 from parts of the Basin and Range Province indicate that deeper, more primitive, mantle also
349 plays a role [*Craig et al.*, 1978; *Kennedy et al.*, 1985; *Welhan et al.*, 1988; *Jordan*, 2002; *Kennedy*
350 *and van Soest*, 2007; *Graham et al.*, 2009].

351 **3 Earthquake Tomographic Models**

352 Slow wave-speed anomalies have been identified at depths of greater than ~ 50 km be-
353 neath western North America [e.g. *Crow et al.*, 2010; *Schmandt and Humphreys*, 2010; *Obreb-*
354 *ski et al.*, 2011; *Shen et al.*, 2013; *Burdick et al.*, 2014; *Schaeffer and Lebedev*, 2014; *Shen and*
355 *Ritzwoller*, 2016]. Here, we target a subset of four tomographic models which reveal the de-
356 tailed structure of the shallow mantle where melts are inferred to have been generated (Fig-
357 ure 7). The chosen models are PM2012, SL2013NA, DNA13 and WUSA16 that were de-
358 veloped by *Priestley and McKenzie* [2013], by *Schaeffer and Lebedev* [2014], by *Porritt et al.*
359 [2014], and by *Shen and Ritzwoller* [2016], respectively. *Porritt et al.* [2014] and *Shen and Ritz-*
360 *woller* [2016] exploit the USArray database for western North America. *Schaeffer and Lebe-*
361 *dev* [2014] address the North American continent and also included USArray data. *Priestley*
362 *and McKenzie* [2013] constructed a lower resolution global model. Despite differences in the

363 wavelength and amplitude of velocity anomalies, these models mostly agree with respect to
 364 the gross pattern of anomalies beneath western North America. Here, we use these models to
 365 investigate the relationship between shear wave velocity, V_s , basalt geochemistry and temper-
 366 ature.

367 The region of western North America addressed by this study is similar to that discussed
 368 by *Afonso et al.* [2016] who carried out a joint inversion of the gravity field, shear wave ve-
 369 locity, together with major element compositions of basaltic rocks and other geophysical ob-
 370 servables by employing a Monte Carlo scheme. This ambitious approach tends to conceal the
 371 major variations in sensitivity possessed by different types of observations. For example, *Priest-*
 372 *ley and McKenzie* [2006] and *Schutt and Leshner* [2006] found that depletion of fertile upper
 373 mantle by removal of a basaltic melt changes shear wave velocity by less than 1%. In con-
 374 trast, a reduction of $\sim 20\%$ occurs as temperature approaches the solidus temperature. Since
 375 the functional form of $V_s(T)$ is both uncertain and controversial, the dependence of V_s on the
 376 extent of depletion can be safely ignored. Two different approaches have been used to deter-
 377 mine $T(V_s, P)$, both of which are empirical and suffer from the lack of any detailed physi-
 378 cal understanding of the grain boundary processes involved. *Faul and Jackson* [2007] param-
 379 eterized detailed laboratory experiments. The problem with this approach is that $T(V_s)$ is strongly
 380 dependent on grain size and the mantle grain size is likely to be two orders of magnitude greater
 381 than that used in laboratory experiments. The other approach, which is exploited here, is to
 382 use geophysical estimates of $V_s(T, P)$ to determine the relevant parameters by exploiting the
 383 functional form for this relationship proposed by *McCarthy et al.* [2011]. This approach is sim-
 384 ilar to that of *Priestley and McKenzie* [2013] with two modifications.

385 The first modification concerns the solidus temperature and melt fraction as a function
 386 of temperature and pressure. Here, we use the parameterization of these quantities described
 387 by *Katz et al.* [2003] to calculate the initial temperature at a spreading ridge, and the average
 388 interior potential temperature of mantle required to generate 7 km of oceanic crust. The re-
 389 sultant changes from the estimates of *McKenzie and Bickle* [1988] are small. For example, the
 390 revised average potential temperature is 1326 °C (instead of 1315 °C). The second modifica-
 391 tion involves using two activation energies to describe the Maxwell viscosity, η , so that

$$\frac{1}{\eta} = \frac{1}{\eta_1} + \frac{1}{\eta_2}, \quad (1)$$

392 where

$$\eta_i = A_i \exp \left[\left(\frac{E_i + (P - P_r)V_i}{R} \right) \left(\frac{1}{T} - \frac{1}{T_r} \right) \right]. \quad (2)$$

393 In this equation, $A_1 = 3.846 \times 10^{21}$ Pa s, $A_2 = 4.201 \times 10^{27}$ Pa s, $E_1 = 402 \times 10^3$ kJ/mol,
 394 $E_2 = 2805 \times 10^3$ kJ mol⁻¹, $V_1 = 0$ m³, $V_2 = 3.112 \times 10^{-5}$ m³, $P_r = 1.5$ GPa, and
 395 $T_r = 1473$ K. R is the gas constant and T the temperature in Kelvin. The purpose of us-
 396 ing two viscosities, where the second one has a large activation energy, is to model the rapid
 397 decrease in V_s with increasing temperature which occurs near the solidus. The unrelaxed shear
 398 modulus, μ , is given by

$$\mu = \mu_0 + \left(\frac{\partial\mu}{\partial P}\right)_T P + \left(\frac{\partial\mu}{\partial T}\right)_P T, \quad (3)$$

399 where $\mu_0 = 69.27$ GPa, $\left(\frac{\partial\mu}{\partial P}\right)_T = 2.679$, $\left(\frac{\partial\mu}{\partial T}\right)_P = -9.231 \times 10^{-3}$ GPa K⁻¹,
 400 and P is pressure in GPa. These expressions for V_s are affected by the presence of melt. For
 401 example, experimental studies show that the presence of melt fractions as small as 0.25% cause
 402 viscosity to decrease by about two orders of magnitude [Faul and Jackson, 2007; McCarthy
 403 and Takei, 2011]. The amount of melt that is retained within the mantle is unlikely to exceed
 404 $\sim 0.1\%$ [Priestley and M^cKenzie, 2006]. This assumption is corroborated by U-series disequi-
 405 librium studies at mid-oceanic ridges [M^cKenzie, 2000]. Although the physics of melt extrac-
 406 tion and retention is poorly understood, we have allowed for an appropriate reduction in vis-
 407 cosity when temperature is close to that of the dry solidus by including η_2 in the parameter-
 408 ization.

409 In this way, $T(z)$ is calculated from $V_s(z)$, which constrains the potential temperature
 410 beneath each volcanic field. We can also estimate lithospheric thickness from the $V_s(z)$ re-
 411 lationship using the method described by Priestley and M^cKenzie [2006] and Priestley and M^cKenzie
 412 [2013]. A geothermal profile is fitted to $T(z)$ and the lithospheric thickness is obtained by ex-
 413 trapolating the conductive portion of the geothermal profile to the depth where it intersects the
 414 adiabatic profile. Finally, we point out that the empirical parameterization used is calibrated
 415 against an updated 2016 version of the PM2012 model (<http://ds.iris.edu/ds/products/emc-cam2016>).
 416 This parameterization yields satisfactory results when applied to all of the models under con-
 417 sideration.

418 3.1 Velocity-Temperature Calibration

419 We extract vertical profiles of $V_s(z)$ at ten locations for each one of the four tomographic
 420 models. These locations are chosen as representative of the average velocity structure for each
 421 volcanic province (Figures 7 and 8). While there are significant differences between these ve-
 422 locity profiles, there are also important commonalities. In all four models, the slowest shear

423 wave velocities are observed beneath the Snake River Plain. The fastest velocities are observed
424 beneath the Colorado Plateau and beneath the Great Plains. In general, V_s between 60 and 100
425 km is slower than that of the WUS reference model [Pollitz, 2008]. Notable exceptions are
426 velocity profiles for the Colorado Plateau taken from the WUSA16 model and for both the
427 Colorado Plateau and the Great Plains taken from the SL2013NA model. Both of these pro-
428 files are positioned close to a sharp lateral change in shear wave velocity at the edge of cra-
429 tonic lithosphere, where velocities are faster than the relevant reference model. Hence, these
430 anomalously fast V_s profiles are probably not representative of the melting region beneath Col-
431 orado Plateau and Great Plains.

432 Temperature profiles correspond to potential temperatures of ~ 1320 – 1380 °C (Figure
433 7b,d,f,h). In accordance with the V_s profiles, the Snake River Plain is the hottest region, Col-
434 orado Plateau has ambient or only marginally elevated temperatures, and the other provinces
435 fall in between. The smallest variability is observed for profiles from the DNA13 model, which
436 can be attributed to damping effects (Figure 7c). The SL2013NA and WUSA16 models pre-
437 dict similar temperature ranges, although the specific order of provinces can differ due to small
438 variations in the proximity to lateral velocity gradients in each case. Lithospheric thicknesses
439 calculated from shear wave velocity profiles are less than ~ 100 km for all four models. Re-
440 cent (< 10 Ma) volcanic activity is focused within regions where the lithosphere is 50–55 km
441 thick (Figure 9). There is a reasonable correspondence between estimated temperatures at 75
442 km depth and the distribution of youthful volcanic activity. The DNA13 model is much smoother
443 than the other models and so its range of calculated temperatures is narrower and estimated
444 lithospheric thicknesses are probably too small.

445 **3.2 Comparing Temperature Estimates**

446 It is illuminating to compare seismically and geochemically determined temperatures.
447 There is a reasonable qualitative correlation between the location and amplitude of slow V_s
448 anomalies and the spatial distribution of basaltic volcanism (Figure 7). Comparison of V_s anoma-
449 lies and geochemical compositions for the screened volcanic database suggests that the ratio
450 of light to heavy REEs (e.g. La/Yb) correlates with shear wave velocities between depths of
451 ~ 60 and ~ 100 km with an optimal correlation at a depth of ~ 75 km where melting proba-
452 bly starts (Figure 10a,c,e,g). This ratio broadly reflects a combination of melt fraction and the
453 depth of melting. Since the depth of melting is similar for samples in the analytical database
454 used here (i.e. the bulk of melting occurs within the spinel field or within the spinel-garnet

455 transition zone), La/Yb can be regarded as a proxy for melt fraction. The highest values of
456 La/Yb are recorded for the Colorado Plateau where the fastest values of V_s are observed. The
457 smallest values of La/Yb are recorded for the Snake River Plain where the slowest values of
458 V_s are observed. Considerable variation of V_s is observed within volcanic fields. One possi-
459 ble cause of this scatter is that much volcanic activity is concentrated at boundaries between
460 anomalously fast and slow velocities. Inevitably, the melting process samples asthenospheric
461 mantle at a spatial resolution that is smaller than that resolved by teleseismic observations. Lat-
462 eral melt migration can cause additional uncertainties. Note that at near-solidus temperatures,
463 V_s rapidly decreases [Priestley and McKenzie, 2013]. Thus modest temperature excursions within
464 a given province can have a significant effect on shear wave velocities.

465 Despite these complications, it is useful to directly compare potential temperatures de-
466 termined from basalt geochemistry with those determined from coincident shear wave veloc-
467 ity profiles (Figure 10b,d,f,h). Temperatures calculated from basaltic geochemistry typically
468 have uncertainties of ± 15 °C, which reflect analytical errors and geographical averaging to-
469 gether with systematic errors associated with the depth and thickness of the spinel-garnet tran-
470 sition zone. Temperatures calculated from shear wave velocities typically have uncertainties
471 of up to ± 15 °C, which reflect geographical averaging. Uncertainties that are a consequence
472 of the velocity-temperature calibration have not been included [Priestley and McKenzie, 2013].
473 A reasonable correlation exists between both sets of potential temperatures with the highest
474 pair of values occurring beneath the Snake River Plain and the lowest pair of values occur-
475 ring beneath the Colorado Plateau. Differences in seismically determined temperatures from
476 different models can be attributed to variations in spatial resolution, in damping, and in the
477 spatial positioning of sharp lateral velocity gradients. Overall, the DNA13 model yields tem-
478 peratures that lie within a narrower band compared with the other three models. We suggest
479 that the WUSA16 model yields the optimal correlation.

480 **4 Discussion**

481 We infer that a combination of anomalously hot asthenosphere and lithospheric thinning
482 has caused regional uplift of western North America during Cenozoic times. It is less obvi-
483 ous what role the foundering and fragmenting Farallon slab plays. Despite a wealth of geo-
484 logic, geophysical and geochemical observations, there has been much debate about possible
485 mechanisms of melt generation. For example, it is suggested that location and style of basaltic
486 magmatism are mainly controlled by the thickness and basal topography of the lithosphere so

487 that sharp gradients at the base of the lithosphere trigger edge-driven convection or shear-driven
488 upwelling [e.g. *van Wijk et al.*, 2010; *Ballmer et al.*, 2015]. It has also been suggested that mag-
489 matism is triggered by tearing of the subducting Farallon slab and/or by melting of metaso-
490 matized lithospheric drips [e.g. *Humphreys et al.*, 2003; *van Wijk et al.*, 2010; *James et al.*, 2011;
491 *Zhou et al.*, 2018]. *Roy et al.* [2009] argue that conductive heating and thinning of the litho-
492 sphere following slab removal could produce uplift and magmatism. The principal difficulty
493 with some of these proposals is their inability to generate both kilometer-scale regional up-
494 lift and basaltic volcanism. Here, we elaborate on four general observations that help to sup-
495 port our results. First, we summarize additional evidence for lithospheric thickness changes
496 and for elevated sub-plate temperatures beneath western North America. We then test alter-
497 native schemes of generating basaltic melts. Finally, we consider the relationship between our
498 results, regional heatflow anomalies, and the spatial and temporal pattern of regional uplift.

499 **4.1 Lithospheric Thickness**

500 There have been significant advances in our understanding of the crustal, lithospheric
501 and sub-lithospheric structure beneath western North America [e.g. *Lin et al.*, 2011; *Hansen*
502 *et al.*, 2013; *Hopper et al.*, 2014; *Shen and Ritzwoller*, 2016]. A striking observation is that crustal
503 thicknesses beneath the Great Plains are similar to, or exceed, those beneath the Colorado Plateau,
504 even though their respective elevations are <500 m and >2000 m [Figure 1; *Shen and Ritz-*
505 *woller*, 2016]. This substantial elevation difference can only be maintained by crustal isostasy
506 if crust beneath the Great Plains is 0.15 Mg/m³ denser than crust beneath the elevated plateaux.
507 This density difference would require crustal velocities beneath the Great Plains to be faster
508 by ~1 km/s, which is not observed [*Hansen et al.*, 2013; *Schmandt et al.*, 2015]. Thus sim-
509 ple isostatic constraints indicate that the topographic elevation of western North America is
510 supported by density variations within the lithospheric and/or the sub-lithospheric mantle [*Levandowski*
511 *et al.*, 2018].

512 Tomographic models demonstrate that the continental lithosphere beneath western North
513 America is approximately one half of the thickness of the cratonic lithosphere beneath the Great
514 Plains [e.g. *Priestley and McKenzie*, 2013; *Schaeffer and Lebedev*, 2014]. These models also
515 indicate that slow shear wave velocity anomalies exist in the upper mantle beneath western
516 North America, although the spatial distribution of these slow anomalies is complicated by the
517 presence of fast anomalies at depths of 300–600 km beneath the Colorado Plateau that are prob-
518 ably associated with the Farallon plate [*Obrebski et al.*, 2011]. Existence of continental litho-

519 sphere that is 50–100 km thick is corroborated to some extent by receiver function studies of
520 the lithosphere-asthenosphere boundary [Kumar *et al.*, 2012; Hopper *et al.*, 2014; Lekić and
521 Fischer, 2014; Hopper *et al.*, 2018]. For example, *S_p* receiver functions place this boundary
522 at 55–65 km beneath the Snake River Plain, and at 60–80 km beneath the Basin and Range,
523 the Rio Grande Rift and the Transition Zone. Beneath the Colorado Plateau, the putative base
524 of the lithosphere appears to occur at 90–140 km depth [Levander *et al.*, 2011; Kumar *et al.*,
525 2012; Levander and Miller, 2012; Hopper *et al.*, 2014; Lekić and Fischer, 2014; Hopper *et al.*,
526 2018]. Using probabilistic inverse modeling of multiple observations, Afonso *et al.* [2016] pre-
527 dict lithospheric thicknesses at the edge of the Colorado Plateau that are in close agreement
528 with those obtained from receiver functions. In the Rio Grande Rift, compositions of mantle
529 xenoliths from ~45 km depth are characteristic of both Proterozoic sub-continental lithosphere
530 as well as of younger depleted upper mantle [Byerly and Lassiter, 2012]. Gao *et al.* [2004] as-
531 sociate the existence of anomalously slow velocities with lithospheric thinning. In contrast,
532 mantle xenoliths from the Zuni-Bandera volcanic field in the Southern Transition Zone are ex-
533 humed from depths of 55–60 km. These xenoliths have sub-continental lithospheric mantle
534 compositions [Byerly and Lassiter, 2012]. Leeman and Rogers [1970] and Lachenbruch and
535 Sass [1977] use anomalous heatflow measurements to constrain melting depths in the Basin
536 and Range and in the Rio Grande Rift to depths of 40–60 km.

537 The origin of thin lithosphere beneath western North America is poorly understood [see,
538 e.g., Kay and Mahlburg-Kay, 1991; Levander and Miller, 2012; Havlin *et al.*, 2013]. There are
539 two possible end-members. First, continental lithosphere beneath western North America may
540 have been thinner than cratonic lithosphere for ~0.5 Ga. Secondly, thick lithosphere may have
541 been thinned, which is more likely for stratigraphic reasons. Thick piles of Paleozoic sedimen-
542 tary rocks are recorded across North America and near identical strata can be traced from the
543 Grand Canyon area toward the center of the continent [e.g. Illinois and Michigan basins; Cross
544 and Pilger, 1978]. Marine sedimentary rocks of the Late Cretaceous Seaway demonstrate that
545 tracts of western North America were below sea level until ~70 Ma, after which regional up-
546 lift occurred [Roberts and Kirschbaum, 1995]. Rapid removal of the lower portion of the litho-
547 sphere might occur by thermal erosion, as a result of the growth of Rayleigh-Taylor instabil-
548 ities [Houseman *et al.*, 1981; Conrad and Molnar, 1997; Lee *et al.*, 2001].

4.2 Other Temperature Estimates

549
550 *Lee et al.* [2009] developed a thermobarometric scheme based upon silica activity and
551 upon Mg exchange between olivine and melt, which can be used to constrain the temperature
552 and pressure of melting within the source region. They propose that melting within the San
553 Francisco volcanic field of the Western Transition Zone takes place at a mantle potential tem-
554 perature of 1480 °C beneath a lithospheric plate that is 120–150 km thick. In the western Basin
555 and Range, they argue that basaltic melts are produced at temperatures of 1350–1450 °C and
556 depths of 60–90 km. *Reid et al.* [2012] apply the same thermobarometer to basaltic samples
557 from the Transition Zone fringing the Colorado Plateau. Based upon the results of *Li et al.* [2008],
558 they assumed that these melts have a water content of 0.05 wt%. They report mantle poten-
559 tial temperatures of > 1465 °C at depths that are mostly shallower than 75 km. *Plank and Forsyth*
560 [2016] adapted the expressions of *Lee et al.* [2009], specifically to exploit a more accurate pa-
561 rameterization of the role of volatiles during melting, and obtained largely similar results. By
562 taking into account water and CO₂ concentrations of basaltic melts, they calculated potential
563 temperatures of ~1300–1500 °C, with an average value of ~1370 °C, at depths of 55–75 km
564 across the Basin and Range and Western Transition Zone.

565 We have applied the method of *Lee et al.* [2009] to our screened database and find that
566 samples from the Western Transition Zone yield temperatures of 1470–1500 °C at ~70 km
567 depth. These values are consistent with the results of *Lee et al.* [2009] but require that the bulk
568 of melting occurs within the garnet stability field, in contrast to our conclusions. Samples from
569 the Basin and Range and from the Rio Grande Rift yield similar, or slightly lower, temper-
570 atures and pressures. For the Snake River Plain, the thermobarometric scheme yields a litho-
571 spheric thickness of ~55 km and mantle potential temperatures of > 1500 °C. The approach
572 of *Lee et al.* [2009] assumes that all melt equilibrates at a single pressure, in contrast to the
573 polybaric fractional melting approach. Furthermore, it is well known that this thermobarom-
574 eter is very sensitive to the Fe³⁺/Fe^T ratio. An average Fe³⁺/Fe^T ratio of 0.2 for western North
575 America is reported for samples < 5 Ma old from the NAVDAT database, which is consis-
576 tent with ratios reported by *Plank and Forsyth* [2016] for the Basin and Range and for the West-
577 ern Transition Zone. In contrast, [*Lee et al.*, 2009] use a ratio of 0.1. Recalculated tempera-
578 tures for our screened database using Fe³⁺/Fe^T = 0.2 are 50 °C lower and better match tem-
579 peratures predicted by our inverse modeling.

580 A range of alternative methodologies have been proposed. Here, we have also tested the
 581 PRIMELT-3 algorithm, which uses a mass balance approach to constrain primary magma com-
 582 positions [Herzberg and Asimow, 2015]. This approach yields potential temperatures of 1340–
 583 1480 °C for our screened database with the greatest spread of temperatures obtained for the
 584 Snake River Plain samples. By applying the scheme of *Hole and Millett* [2016], we estimated
 585 final pressures and temperatures of melting, obtaining results that are similar to those of *Lee*
 586 *et al.* [2009]. *Rudzitis et al.* [2016] applied a different thermobarometric scheme based upon
 587 clinopyroxene fractionation to Western Transition Zone samples. They obtain crystallisation
 588 temperatures that are 100–200 °C lower than primary melt temperatures calculated by *Lee et al.*
 589 [2009]. We conclude that many strategies yield broadly similar ranges of mantle potential tem-
 590 peratures and depths of melting. Nevertheless, it is notable that inverse modeling of REE con-
 591 centrations yields melting temperatures that are lower by up to 100 °C compared with ther-
 592 mobarometric temperature estimates. This systematic disparity is partly accounted for by vari-
 593 ations in the ratio of $\text{Fe}^{3+}/\text{Fe}^T$.

594 4.3 Alternative Mechanisms of Melt Generation

595 Basaltic melting beneath continental lithosphere can be produced by elevating mantle
 596 temperature, thinning the lithosphere, and/or introducing volatiles to the source region [*Green*
 597 *and Ringwood*, 1967; *McKenzie and Bickle*, 1988]. Distribution of volcanic activity across west-
 598 ern North America is evidently correlated with the planform of shear wave velocity anoma-
 599 lies where lithospheric thickness is <100 km. However, melt fractions are significantly higher
 600 than those typically generated by melting of dry, peridotitic mantle at ambient potential tem-
 601 peratures. It is well known that hydration reduces melting temperatures by ~50 °C [*Katz et al.*,
 602 2003]. To assess the role that water could play in generating slow seismic anomalies, as well
 603 as accounting for the distribution and composition of observed volcanism, we used the alphaMELTS
 604 algorithm to generate forward models of melting at 0–4 GPa for mantle potential temperatures
 605 of 1250, 1300 and 1350 °C with source water contents of 0 to 10⁴ ppm [*Ghiorso et al.*, 2002].
 606 Assuming near-fractional isentropic melting and a residual porosity of 0.5%, we calculated trace
 607 element compositions, temperatures and melt fraction profiles, together with the changing wa-
 608 ter content of both source and melt. Shear wave velocity profiles are calculated using an ap-
 609 propriate correction for source water content [*Karato*, 2003].

610 We find that 10⁴ ppm of water in the source region is required to give rise to a gradi-
 611 ent change in REE concentrations that is similar to that produced by a temperature increase

612 of ~ 50 °C. However, melt fraction distributions and $V_s(z)$ profiles are significantly different
613 when water content is varied instead of temperature. In the hydrous example, the depth of on-
614 set of melting is deeper but the cumulative melt fraction remains similar to that generated by
615 dry melting at the same temperature. Since water content decreases rapidly with continued melt-
616 ing, seismic velocities first increase with decreasing depth due to loss of water before slowly
617 declining with decreasing pressure, once water is exhausted from the source. Dry melting at
618 higher temperatures also leads to deepening of the onset of melting but produces much larger
619 cumulative melt fractions and a smoother velocity profile that decreases with pressure. With
620 regard to inverse modeling of REE compositions, a significantly hydrated source region could
621 be simulated with a much more enriched source composition and/or with a low melt fraction
622 tail within the garnet field but no change in the potential temperature estimate (i.e. final melt
623 fraction and depth of melting). We conclude that the presence of minor amounts of water within
624 the mantle do not significantly change our results.

625 We cannot entirely preclude water as a contributing factor to mantle melting beneath west-
626 ern North America. There are, however, several arguments suggesting that water content does
627 not play a significant role with regard to the modeling of analyses presented here. There is undis-
628 puted evidence for water in the source region beneath the Basin and Range and beneath the
629 Colorado Plateau, based upon melt inclusion observations [e.g. *Plank and Forsyth, 2016; Gazel*
630 *et al., 2012*], upon geochemical signatures [e.g. *Reid et al., 2012; Rudzitis et al., 2016*], and
631 upon the presence of water in nominally anhydrous minerals of xenoliths [*Li et al., 2008*]. How-
632 ever, those who favor the importance of water content for generation of basaltic volcanism in
633 these regions also agree that temperature anomalies are required [e.g. *Dixon et al., 2004*]. There
634 is a consensus that some combination of long-lived hydration of the upper mantle caused by
635 the presence of the subducting Farallon plate and temperature anomalies are needed in order
636 to account for geochemical observations. If water content were the primary cause of melting,
637 a homogeneous distribution of water within the upper mantle over a considerable area would
638 be required. This signature would necessarily have to be either preserved or constantly replen-
639 ished during the 80 Ma period over which volcanism has occurred. Furthermore, water con-
640 tent can only affect the initial stages of melting—it starts deeper and compositions are more
641 enriched than for dry melting at the same temperature but the cumulative melt fraction is al-
642 most identical.

643 Finally, if a pyroxenite source is assumed, melt productivity increases without requir-
644 ing anomalously elevated mantle temperatures [e.g. *Hirschmann et al., 2003*]. Pyroxenite is

645 significantly more fusible than peridotite, which means that melting is initiated at greater depths
 646 leading to the generation of larger melt fractions throughout the entire melting column. *Reid*
 647 *et al.* [2012] and *Rudzitis et al.* [2016] suggest that the mantle source region for representa-
 648 tive primitive basalts from the San Francisco and Mormon Mountain Volcanic Fields of the
 649 Western Transition Zone is predominantly peridotitic, based upon Hf-Nd isotopic ratios, as well
 650 as Zn/Fe and Fe/Mn ratios for olivine. Our screened database was carefully selected on the
 651 basis of its similarity to OIB compositions, which are thought to be primarily derived from
 652 peridotitic sources [*Shorttle et al.*, 2014]. If the approach of *Shorttle and Maclennan* [2011]
 653 is applied, which uses major elemental compositions to constrain potential end-member sources,
 654 our screened samples are inferred to have been generated by melting of dry lherzolitic rocks.

655 4.4 Regional Uplift & Heatflow

656 Isostatic calculations help to gauge whether or not our estimates of mantle temperature
 657 and lithospheric thickness are sufficient to generate the elevated topography of western North
 658 America (Figure 11). Following the approach of *McNab et al.* [2018] and many published con-
 659 tributions, we balance idealized columns of continental lithosphere against the density struc-
 660 ture of a mid-ocean ridge. Elevation, e , of continental lithosphere is given by

$$e = t_w \left(\frac{\rho_w - \rho_a}{\rho_{ca}} \right) + t_{oc} \left(\frac{\rho_{oc} - \rho_a}{\rho_{ca}} \right) + t_{cc} \left(\frac{\rho_m - \rho_{cc}}{\rho_{ca}} \right) + 200 \left(\frac{\rho_a - \rho_{ca}}{\rho_{ca}} \right) - t_m \left(\frac{\rho_m - \rho_{ca}}{\rho_{ca}} \right), \quad (4)$$

661 where $t_w = 2.8$ km and $\rho_w = 1$ Mg m⁻³ are the thickness and density of water at the mid-
 662 ocean ridge, $t_{oc} = 7.1$ km, $\rho_{oc} = 2.86$ Mg m⁻³ and $t_{cc} = 35$ – 50 km, $\rho_{cc} = 2.8$ Mg m⁻³
 663 are variable thicknesses and densities of oceanic and continental crust, respectively. The vari-
 664 able lithospheric thickness, $t_m = 60$ – 200 km. ρ_m , ρ_a and ρ_{ca} are densities of lithospheric
 665 mantle and of asthenospheric mantle beneath the mid-ocean ridge and beneath continental litho-
 666 sphere, respectively. Their values depend upon temperature and are calculated using $\rho_T =$
 667 $\rho_0(1 - \alpha T)$ where $\alpha = 3.15 \times 10^{-5}$ °C⁻¹ and $\rho_0 = 3.33$ Mg m⁻³ is the density of mantle
 668 material at $T = 0$ °C. Note that ρ_a and ρ_{ca} differ in order to account for the putative ther-
 669 mal anomaly beneath western North America. The value of 200 km refers to the compensa-
 670 tion depth which is taken to be the typical thickness of cratonic lithosphere beneath central
 671 North America.

672 First, we calculate the value of e for each of the ten volcanic provinces shown in Fig-
 673 ure 7, using values of mantle potential temperature and lithospheric thickness determined by
 674 geochemical inverse modeling. At each location, average densities of lithospheric and astheno-

675 spheric mantle are determined by assuming a simple linear gradient and an adiabatic gradi-
676 ent, respectively. The effects of thermal expansion and compressibility were taken into account.
677 Chemical depletion of continental lithospheric mantle due to extraction of 1.5% melt was ac-
678 counted for by reducing its density by 15 kg m^{-3} [Crosby *et al.*, 2010]. The crustal thickness
679 profile is extracted from the model of Shen and Ritzwoller [2016] and the average crustal den-
680 sity is taken to be 2.8 Mg m^{-3} which is in close agreement with the results of Schmandt *et al.*
681 [2015]. At each location, lithospheric thickness is varied by $\pm 10 \text{ km}$ and mantle potential tem-
682 perature by $\pm 20 \text{ }^\circ\text{C}$, and in this way a mean elevation and its standard deviation are computed.
683 Secondly, we have constructed profiles at regular intervals along a curved transect that inter-
684 sects the principal volcanic provinces under consideration and terminates at the craton. On these
685 profiles, crustal thickness is also taken from Shen and Ritzwoller [2016] but lithospheric thick-
686 nesses are set in accordance with the results of Hopper *et al.* [2018] for western North Amer-
687 ica and of Priestley and McKenzie [2013] for the cratonic regions. Mantle densities are esti-
688 mated by converting V_s profiles from the SL2013NA model into temperature and density. Litho-
689 spheric thickness beneath the craton is varied by $\pm 25 \text{ km}$.

690 Elevations calculated in these two ways agree to within $\sim 300 \text{ m}$ and match the observed
691 topography to better than 500 m with three exceptions (Figure 11c). Large misfits occur at the
692 transition between western and cratonic North America toward the eastern end of this tran-
693 sect, on the Colorado Plateau, and adjacent to the Great Plains volcanic province. These mis-
694 fits arise from uncertainties in lithospheric thickness, thermal or density structure, as well as
695 the assumption of a simplified crustal structure [Rodgers *et al.*, 2002; Hopper *et al.*, 2018]. Across
696 western North America, the proportion of elevation that is generated and maintained by as-
697 thenospheric thermal anomalies appears to be $< 300 \text{ m}$, in agreement with previous isostatic
698 studies [Levandowski *et al.*, 2014, 2018]. We infer that the bulk of the topographic difference
699 between western and cratonic North America is caused by a $\sim 100 \text{ km}$ difference in lithospheric
700 thickness. A contribution from mantle flow is not specifically required to match these obser-
701 vations in agreement with Roy *et al.* [2009], Hyndman and Currie [2011] and Afonso *et al.* [2016].

702 The distribution of heatflow anomalies broadly matches the pattern of Cenozoic basaltic
703 magmatism and of shear wave velocity anomalies [Figure 11; Christiansen and Yeats, 1992;
704 Pollack *et al.*, 1993]. Lee and Uyeda [1965] and Roy *et al.* [1968] showed that heatflow mea-
705 surements are twice as high as the continental average. For example, heatflow through the Snake
706 River Plain is about 100 mW m^{-2} with geothermal gradients as high as $\sim 70 \text{ }^\circ\text{C km}^{-1}$. These
707 values increase eastward toward Yellowstone [Blackwell, 1989]. Average surface heatflow across

708 the Colorado Plateau is 65 mW m^{-2} with values of $\sim 83 \text{ mW m}^{-2}$ near the Jemez lineament
 709 (e.g. Zuni-Bandera field, STZ) and $\sim 95 \text{ mW m}^{-2}$ within the central Rio Grande Rift [Reiter
 710 and Mansure, 1983; Eggleston and Reiter, 1984; Reiter et al., 1986]. We calculate conductive
 711 heatflow at the surface for different columns of continental lithosphere using

$$\rho C_P \frac{\partial T}{\partial t} = \frac{-\partial Q}{\partial z} + A, \quad (5)$$

712 where ρ is density, $C_P = 1.2 \times 10^3 \text{ J mol}^{-1}$ is specific heat capacity, and $A = 0.75 \text{ } \mu\text{W}$
 713 m^{-3} is crustal heat production [Michaut et al., 2009; McKenzie et al., 2005]. Heatflow, Q , is
 714 related to the continental temperature gradient by

$$Q = -k \frac{\partial T}{\partial z}, \quad (6)$$

715 where k is thermal conductivity. Average mantle density is calculated directly from the assumed
 716 temperature structure. Thermal conductivity is assumed to vary as a function of temperature.
 717 Within the mantle, $k(T)$ is parameterized using conductivity measurements of olivine [Xu et al.,
 718 2004]. In the crust, the experimentally constrained expression of Whittington et al. [2009] is
 719 employed. In this way, surface heatflow is calculated both for lithospheric columns within each
 720 of the ten volcanic provinces shown in Figure 7 and for the transect shown in Figure 11a. At
 721 appropriate spot locations, we used temperature and lithospheric thickness estimates from geo-
 722 chemical inverse modeling. Along the transect, we use temperature estimates derived from the
 723 SL2013NA tomographic model of Schaeffer and Lebedev [2014]. Lithospheric thicknesses are
 724 from Priestley and McKenzie [2013] and Hopper et al. [2018].

725 We calculate an average surface heatflow of $65\text{--}80 \text{ mW m}^{-2}$ for western North Amer-
 726 ica and of $50\text{--}60 \text{ mW m}^{-2}$ for cratonic lithosphere (Figure 11d). Estimates determined from
 727 the results of geochemical inverse modeling are consistently higher than those determined from
 728 tomographic and receiver function models. This difference of $\sim 30 \text{ mW m}^{-2}$ appears to be re-
 729 solvable and probably reflects the presence of a thermal anomaly beneath a thin plate. We ac-
 730 knowledge that these heat flow calculations are simplistic and do not account for any lateral
 731 heterogeneities of internal heat production. In the Basin and Range and Snake River Plain provinces,
 732 the existence of significant sediment-filled basins and shallow aquifers act to reduce surface
 733 heatflow measurements [Blackwell, 1989]. A detailed treatment of near-surface conductivity
 734 structure would probably yield better fits to heatflow observations but it is beyond the scope
 735 of this investigation. We conclude that a combination of elevated asthenospheric temperature
 736 and thin lithosphere significantly elevates surface heatflow.

5 Conclusions

We analyze and model a comprehensive database of Cenozoic basaltic volcanic rocks from western North America. Our principal aim is to show that rare earth and other incompatible trace element measurements can be used to determine melt fraction as a function of depth, which enables asthenospheric temperature and plate thickness to be estimated. Basaltic magmatism is generated by adiabatic decompression at, or close to, the spinel-garnet transition zone. The average lithospheric thickness constrained by rare earth element inverse modeling is 55 ± 10 km with melt generation beneath the Colorado Plateau being as deep as 70 ± 10 km and melt generation beneath the Snake River Plain being as shallow as 50 ± 10 km. Most of this melting occurs at depths shallower than ~ 70 km. The average mantle potential temperature is 1340 ± 20 °C which is slightly hotter than the ambient asthenospheric value of ~ 1330 °C. Potential temperatures as high as 1365 °C occur beneath the Snake River Plain but the Colorado Plateau is underlain by mantle of ambient temperature.

These geochemical results are compared with shear wave velocity anomalies from a suite of regional and global tomographic models. We find that there is a positive correlation between shear wave velocities and trace element ratios, such as La/Yb, which act as proxies for the degree of melting. This correlation is confirmed by using an empirical calibration method to convert shear wave velocities into sub-plate temperatures. Seismically determined potential temperatures broadly agree with potential temperatures constrained by geochemical inverse modeling. We believe that this result is not significantly affected by variations in source rock composition, or by the presence of water in the source region.

Simple isostatic calculations highlight the overall consistency between regional epeirogeny, anomalously slow shear wave velocities, thinner lithosphere, and elevated heatflow across western North America. This consistency implies that sub-vertical mantle flow may not be a necessary prerequisite for generating and maintaining the observed regional topography. Instead, a combination of thin lithosphere and moderately elevated mantle potential temperature could be sufficient to explain ~ 2 km of regional elevation. The existence of temperature anomalies suggests that edge-driven convection along the cratonic lithospheric keel may not be the primary cause of regional uplift and basaltic volcanism. Instead, our results bolster the notion that in this instance large-scale dynamic topography is generated and maintained by temperature anomalies within asthenospheric mantle directly beneath a thin plate.

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| Province | SRP avg (n=12) | GP avg (n=8) | ETZ avg (n=13) | STZ avg (n=18) | SE avg (n=5) |
|---|-------------------|-----------------|-------------------|-------------------|-----------------|
| SiO ₂ (wt%) | 47.34 ± 0.78 | 48.84 ± 0.62 | 49.13 ± 1.94 | 46.71 ± 1.56 | 49.06 ± 1.44 |
| Al ₂ O ₃ | 15.43 ± 0.40 | 15.88 ± 0.71 | 15.02 ± 0.31 | 15.39 ± 0.58 | 14.36 ± 0.33 |
| Fe ₂ O ₃ ^T | 12.62 ± 0.85 | 11.72 ± 0.47 | 12.30 ± 0.44 | 11.95 ± 0.86 | 11.71 ± 0.37 |
| MgO | 9.37 ± 0.81 | 8.62 ± 1.09 | 9.08 ± 0.86 | 8.79 ± 0.91 | 8.10 ± 0.35 |
| CaO | 10.52 ± 0.43 | 9.24 ± 0.26 | 9.10 ± 0.35 | 9.45 ± 1.03 | 9.52 ± 0.82 |
| Na ₂ O | 2.41 ± 0.22 | 3.36 ± 0.11 | 2.90 ± 0.27 | 3.16 ± 0.55 | 3.05 ± 0.19 |
| K ₂ O | 0.53 ± 0.15 | 0.97 ± 0.25 | 0.83 ± 0.39 | 1.11 ± 0.50 | 0.79 ± 0.15 |
| TiO ₂ | 1.74 ± 0.39 | 1.54 ± 0.19 | 1.63 ± 0.40 | 2.05 ± 0.34 | 1.77 ± 0.24 |
| MnO | 0.19 ± 0.01 | 0.17 ± 0.01 | 0.16 ± 0.02 | 0.18 ± 0.01 | 0.16 ± 0.01 |
| P ₂ O ₅ | 0.40 ± 0.08 | 0.41 ± 0.03 | 0.28 ± 0.14 | 0.45 ± 0.16 | 0.31 ± 0.06 |
| Li (ppm) | 6.99 ± 0.80 | 8.42 ± 1.24 | 6.99 ± 1.47 | 10.74 ± 8.56 | 10.54 ± 2.31 |
| Be | 0.68 ± 0.14 | 0.99 ± 0.23 | 0.93 ± 0.38 | 1.43 ± 0.51 | 0.95 ± 0.14 |
| P | 1425 ± 366 | 1791 ± 125 | 1244 ± 656 | 2031 ± 751 | 1475 ± 204 |
| K | 3574 ± 790 | 8893 ± 2447 | 6999 ± 3292 | 9386 ± 4040 | 7476 ± 1422 |
| Sc | 32.23 ± 3.26 | 27.85 ± 2.63 | 23.86 ± 1.99 | 25.62 ± 4.06 | 22.44 ± 1.44 |
| Ti | 10139 ± 2539 | 9470 ± 1291 | 9771 ± 2537 | 12299 ± 2351 | 11550 ± 1340 |
| V | 247.1 ± 14.8 | 191.6 ± 17.1 | 198.5 ± 22.5 | 221.2 ± 31.9 | 199.7 ± 7.4 |
| Cr | 391.3 ± 132.9 | 245.4 ± 72.3 | 275.3 ± 28.2 | 274.5 ± 57.0 | 296.9 ± 34.0 |
| Mn | 1447 ± 55 | 1379 ± 181 | 1396 ± 117 | 1487 ± 115 | 1355 ± 201 |
| Co | 51.70 ± 2.91 | 58.95 ± 6.56 | 77.77 ± 13.70 | 70.33 ± 11.59 | 242.2 ± 406.0 |
| Ni | 139.8 ± 46.7 | 156.0 ± 47.0 | 199.5 ± 31.6 | 173.7 ± 45.7 | 221.0 ± 39.8 |
| Cu | 52.72 ± 16.29 | 58.50 ± 16.03 | 85.15 ± 21.32 | 59.72 ± 11.63 | 96.28 ± 15.76 |
| Zn | 92.33 ± 12.10 | 87.84 ± 5.01 | 92.79 ± 4.82 | 81.04 ± 9.27 | 99.32 ± 5.11 |
| Ga | 17.15 ± 1.30 | 18.42 ± 0.60 | 18.29 ± 0.81 | 18.46 ± 0.67 | 19.80 ± 0.74 |
| Rb | 9.42 ± 3.36 | 14.47 ± 7.15 | 14.96 ± 4.66 | 17.11 ± 7.82 | 17.79 ± 2.36 |
| Sr | 239.5 ± 44.0 | 600.9 ± 68.5 | 376.0 ± 169.5 | 590.2 ± 188.8 | 368.8 ± 55.9 |
| Y | 27.91 ± 3.33 | 24.08 ± 1.65 | 21.98 ± 1.93 | 25.58 ± 2.18 | 23.54 ± 0.77 |
| Zr | 159.1 ± 32.0 | 155.2 ± 22.6 | 134.6 ± 49.3 | 182.6 ± 50.7 | 129.4 ± 15.7 |
| Nb | 16.73 ± 3.68 | 23.95 ± 8.29 | 19.70 ± 11.70 | 43.76 ± 17.72 | 25.28 ± 6.21 |
| Sn | 0.80 ± 0.17 | 1.06 ± 0.20 | 1.00 ± 0.26 | 1.30 ± 0.30 | 1.08 ± 0.23 |
| Cs | 0.10 ± 0.04 | 0.13 ± 0.15 | 0.37 ± 0.32 | 0.45 ± 0.59 | 0.12 ± 0.06 |
| Ba | 263.7 ± 50.5 | 410.5 ± 84.0 | 193.8 ± 79.5 | 345.7 ± 96.5 | 303.5 ± 67.2 |
| La | 15.53 ± 3.05 | 22.46 ± 2.52 | 15.31 ± 7.70 | 27.52 ± 8.61 | 17.44 ± 2.99 |
| Ce | 33.60 ± 6.85 | 45.74 ± 4.09 | 32.04 ± 15.20 | 54.16 ± 15.69 | 35.69 ± 5.72 |
| Pr | 4.44 ± 0.93 | 5.61 ± 0.39 | 4.04 ± 1.78 | 6.39 ± 1.69 | 4.51 ± 0.67 |
| Nd | 19.61 ± 4.12 | 23.04 ± 1.63 | 17.44 ± 6.92 | 26.03 ± 6.22 | 19.71 ± 2.71 |
| Sm | 4.69 ± 0.93 | 4.81 ± 0.29 | 4.16 ± 1.22 | 5.46 ± 0.98 | 4.83 ± 0.58 |
| Eu | 1.66 ± 0.30 | 1.58 ± 0.10 | 1.40 ± 0.35 | 1.81 ± 0.33 | 1.64 ± 0.17 |
| Gd | 5.04 ± 0.87 | 4.72 ± 0.21 | 4.36 ± 0.87 | 5.42 ± 0.78 | 5.21 ± 0.40 |
| Tb | 0.83 ± 0.13 | 0.74 ± 0.04 | 0.70 ± 0.10 | 0.84 ± 0.11 | 0.83 ± 0.05 |
| Dy | 5.08 ± 0.72 | 4.31 ± 0.23 | 4.09 ± 0.45 | 4.83 ± 0.50 | 4.68 ± 0.31 |
| Ho | 1.05 ± 0.13 | 0.85 ± 0.06 | 0.79 ± 0.07 | 0.92 ± 0.09 | 0.87 ± 0.03 |
| Er | 2.97 ± 0.32 | 2.38 ± 0.22 | 2.18 ± 0.17 | 2.53 ± 0.27 | 2.35 ± 0.13 |
| Tm | 0.43 ± 0.05 | 0.34 ± 0.04 | 0.31 ± 0.03 | 0.36 ± 0.04 | 0.32 ± 0.02 |
| Yb | 2.62 ± 0.24 | 2.06 ± 0.22 | 1.90 ± 0.17 | 2.21 ± 0.27 | 1.90 ± 0.17 |
| Lu | 0.40 ± 0.04 | 0.30 ± 0.04 | 0.28 ± 0.02 | 0.32 ± 0.04 | 0.27 ± 0.03 |
| Hf | 3.49 ± 0.69 | 3.08 ± 0.37 | 3.10 ± 0.93 | 4.05 ± 0.92 | 3.18 ± 0.34 |
| Ta | 0.94 ± 0.39 | 1.27 ± 0.44 | 1.55 ± 0.74 | 3.02 ± 1.41 | 3.10 ± 3.60 |
| Tl | 0.02 ± 0.01 | 0.03 ± 0.01 | 0.04 ± 0.01 | 0.03 ± 0.01 | 0.02 ± 0.01 |
| Pb | 2.76 ± 0.65 | 4.16 ± 2.29 | 2.80 ± 0.89 | 2.63 ± 1.10 | 1.78 ± 0.41 |
| Th | 0.98 ± 0.28 | 2.11 ± 0.50 | 1.88 ± 0.64 | 3.44 ± 1.46 | 1.90 ± 0.36 |
| U | 0.31 ± 0.12 | 0.44 ± 0.28 | 0.68 ± 0.25 | 0.95 ± 0.43 | 0.58 ± 0.14 |
| ϵ_{Nd} | -5.09 ± 0.06 | | 5.11 | | |

| Province | NBR avg (n=2) | WTZ avg (n=40) | BR avg (n=7) | RGR avg (n=65) | CP avg (n=7) |
|---|------------------|-------------------|-----------------|-------------------|-----------------|
| SiO ₂ (wt%) | 47.01 ± 0.24 | 46.36 ± 1.89 | 46.13 ± 1.46 | 45.24 ± 1.13 | 41.37 ± 1.21 |
| Al ₂ O ₃ | 16.76 ± 0.39 | 14.18 ± 1.06 | 15.21 ± 0.59 | 14.89 ± 0.63 | 11.07 ± 0.52 |
| Fe ₂ O ₃ ^T | 11.88 ± 0.18 | 12.15 ± 0.91 | 12.85 ± 1.17 | 11.68 ± 0.57 | 13.39 ± 0.61 |
| MgO | 7.40 ± 0.29 | 9.98 ± 1.47 | 8.28 ± 0.90 | 9.88 ± 1.23 | 9.28 ± 1.44 |
| CaO | 9.51 ± 0.04 | 9.90 ± 1.00 | 9.29 ± 0.41 | 10.29 ± 0.63 | 11.26 ± 0.46 |
| Na ₂ O | 3.25 ± 0.00 | 3.22 ± 0.48 | 3.60 ± 0.38 | 3.48 ± 0.48 | 3.60 ± 0.76 |
| K ₂ O | 1.12 ± 0.15 | 1.24 ± 0.44 | 1.50 ± 0.50 | 1.60 ± 0.39 | 0.86 ± 0.28 |
| TiO ₂ | 2.15 ± 0.04 | 2.06 ± 0.53 | 2.53 ± 0.35 | 2.31 ± 0.17 | 3.72 ± 0.22 |
| MnO | 0.19 ± 0.01 | 0.18 ± 0.01 | 0.19 ± 0.01 | 0.18 ± 0.01 | 0.19 ± 0.02 |
| P ₂ O ₅ | 0.54 ± 0.04 | 0.60 ± 0.19 | 0.55 ± 0.13 | 0.54 ± 0.08 | 1.45 ± 0.44 |
| Li (ppm) | 11.63 ± 2.98 | 9.98 ± 3.66 | 8.31 ± 2.00 | 8.33 ± 1.70 | 32.66 ± 14.02 |
| Be | 1.28 ± 0.06 | 1.63 ± 0.56 | 1.47 ± 0.52 | 1.61 ± 0.15 | 2.77 ± 0.85 |
| P | 2303 ± 118 | 2655 ± 942 | 2517 ± 567 | 2378 ± 253 | 6393 ± 2118 |
| K | 8737 ± 1209 | 9326 ± 3495 | 12970 ± 4787 | 13815 ± 2265 | 7126 ± 2262 |
| Sc | 29.05 ± 1.34 | 23.55 ± 4.32 | 21.14 ± 2.43 | 29.09 ± 3.47 | 18.10 ± 2.88 |
| Ti | 12603 ± 229 | 8048 ± 4786 | 15469 ± 2246 | 14099 ± 1691 | 21706 ± 1529 |
| V | 251.2 ± 34.1 | 222.4 ± 25.0 | 218.9 ± 12.6 | 222.7 ± 23.2 | 244.6 ± 28.4 |
| Cr | 164.2 ± 63.2 | 360.0 ± 148.3 | 213.8 ± 85.2 | 283.0 ± 78.6 | 233.8 ± 81.9 |
| Mn | 1496 ± 57 | 1395 ± 138 | 1496 ± 88 | 1531 ± 150 | 1576 ± 122 |
| Co | 39.64 ± 0.25 | 57.14 ± 17.00 | 58.37 ± 8.70 | 63.41 ± 6.84 | 52.36 ± 5.51 |
| Ni | 115.3 ± 32.2 | 206.9 ± 51.5 | 138.8 ± 36.9 | 183.4 ± 49.7 | 190.8 ± 64.8 |
| Cu | 51.25 ± 2.05 | 61.40 ± 11.09 | 40.33 ± 4.24 | 53.09 ± 2.93 | 53.99 ± 9.90 |
| Zn | 77.15 ± 0.07 | 100.3 ± 14.7 | 89.73 ± 6.43 | 76.68 ± 5.84 | 138.9 ± 23.6 |
| Ga | 17.47 ± 0.12 | 17.77 ± 3.31 | 19.64 ± 1.22 | 19.47 ± 1.06 | 22.36 ± 2.03 |
| Rb | 21.06 ± 0.59 | 20.03 ± 7.92 | 36.47 ± 15.40 | 33.72 ± 9.57 | 13.92 ± 7.22 |
| Sr | 365.7 ± 30.1 | 752.0 ± 236.1 | 644.8 ± 96.3 | 652.1 ± 94.7 | 1654 ± 482 |
| Y | 37.00 ± 0.42 | 23.59 ± 2.08 | 28.86 ± 1.77 | 27.71 ± 2.33 | 37.87 ± 8.92 |
| Zr | 206.2 ± 22.3 | 195.8 ± 57.6 | 227.1 ± 76.1 | 192.3 ± 28.4 | 532.0 ± 153.1 |
| Nb | 41.10 ± 5.94 | 50.13 ± 21.21 | 55.16 ± 19.70 | 57.11 ± 9.20 | 119.6 ± 31.9 |
| Sn | 1.34 ± 0.05 | 1.32 ± 0.34 | 1.43 ± 0.44 | 1.33 ± 0.13 | 2.69 ± 0.46 |
| Cs | 0.25 ± 0.14 | 1.41 ± 5.88 | 0.30 ± 0.17 | 0.30 ± 0.11 | 1.34 ± 1.38 |
| Ba | 292.2 ± 62.2 | 648.7 ± 290.5 | 387.4 ± 54.9 | 490.0 ± 87.2 | 1070 ± 137 |
| La | 21.14 ± 1.39 | 42.01 ± 17.03 | 32.00 ± 10.39 | 36.40 ± 6.03 | 104.5 ± 34.1 |
| Ce | 46.84 ± 2.40 | 80.88 ± 29.95 | 64.82 ± 19.44 | 69.28 ± 10.28 | 215.0 ± 68.7 |
| Pr | 6.20 ± 0.20 | 9.47 ± 3.23 | 7.90 ± 2.03 | 8.72 ± 1.24 | 25.57 ± 8.00 |
| Nd | 27.07 ± 0.30 | 36.55 ± 11.60 | 32.80 ± 7.33 | 34.85 ± 4.27 | 102.1 ± 30.4 |
| Sm | 6.10 ± 0.16 | 6.86 ± 1.72 | 6.81 ± 1.16 | 6.90 ± 0.63 | 17.63 ± 4.51 |
| Eu | 2.03 ± 0.06 | 2.11 ± 0.46 | 2.26 ± 0.35 | 2.20 ± 0.18 | 5.08 ± 1.29 |
| Gd | 6.38 ± 0.16 | 6.05 ± 1.12 | 6.54 ± 0.76 | 6.46 ± 0.44 | 13.51 ± 2.84 |
| Tb | 1.04 ± 0.03 | 0.88 ± 0.14 | 1.00 ± 0.10 | 0.97 ± 0.06 | 1.73 ± 0.40 |
| Dy | 6.35 ± 0.23 | 4.67 ± 0.56 | 5.52 ± 0.39 | 5.32 ± 0.31 | 8.34 ± 1.77 |
| Ho | 1.33 ± 0.04 | 0.86 ± 0.08 | 1.03 ± 0.07 | 1.02 ± 0.06 | 1.37 ± 0.30 |
| Er | 3.80 ± 0.12 | 2.19 ± 0.24 | 2.82 ± 0.16 | 2.62 ± 0.17 | 3.23 ± 0.66 |
| Tm | 0.56 ± 0.01 | 0.30 ± 0.03 | 0.39 ± 0.03 | 0.40 ± 0.03 | 0.40 ± 0.08 |
| Yb | 3.50 ± 0.11 | 1.78 ± 0.27 | 2.33 ± 0.11 | 2.36 ± 0.17 | 2.23 ± 0.46 |
| Lu | 0.54 ± 0.01 | 0.26 ± 0.04 | 0.34 ± 0.02 | 0.37 ± 0.03 | 0.29 ± 0.06 |
| Hf | 4.58 ± 0.27 | 4.49 ± 1.22 | 4.90 ± 1.31 | 4.68 ± 0.60 | 10.85 ± 2.66 |
| Ta | 2.19 ± 0.30 | 2.89 ± 1.18 | 3.21 ± 1.14 | 3.86 ± 0.72 | 6.44 ± 1.54 |
| Tl | 0.02 ± 0.01 | 0.04 ± 0.02 | 0.03 ± 0.01 | 0.03 ± 0.01 | 0.07 ± 0.06 |
| Pb | 1.64 ± 0.33 | 5.88 ± 2.16 | 1.91 ± 0.68 | 1.88 ± 0.76 | 7.17 ± 2.32 |
| Th | 1.90 ± 0.03 | 6.12 ± 2.92 | 3.62 ± 1.81 | 4.34 ± 0.88 | 11.60 ± 3.45 |
| U | 0.63 ± 0.06 | 1.69 ± 0.76 | 1.02 ± 0.48 | 1.18 ± 0.26 | 3.75 ± 1.96 |
| ϵ_{Nd} | | 2.38 | | 6.04 0.49 | |

1224 **Table 1.** Average compositions used for inverse modeling. Major elements are reported as oxide wt%, trace
1225 elements as ppm. SRP = Snake River Plain; GP = Great Plains; ETZ = Eastern Transition Zone; STZ = South-
1226 ern Transition Zone; SE = Sentinel Plain; NBR = Northern Basin and Range; WTZ = Western Transition
1227 Zone; BR = Basin and Range; RGR = Rio Grande Rift; CP = Colorado Plateau. Samples from SRP include
1228 L73-64 and L73-112 from *Leeman et al.* [2009] and I-2725 from *White et al.* [2004]. Samples from WTZ
1229 include AZ-09 UK-1, 2, 11, 13b, 18, 19b, 22, 23b, 26, 27, 30, 31b, 32–35, SC 07 03, and SC 07 05 generously
1230 provided by T. Plank [written communication, 2015; *Plank and Forsyth*, 2016]. Samples from RGR include
1231 671, 672, 674–676, 678, 695, 699, 860, 864–866, 869, 870, 875, 879, 882, 883, 888, 894, 895, 898, 6100,
1232 6102–6104, 6108, 6110, 6130, 6140, 6143, 6151, 6152, 6155, 6157, 6158, 6185, 6187, 8101, 8103, 8107,
1233 8109–8112, 8128, 8129, 8134, 8136, 8138–8140, 8144, 8157, 8159–8161, and 8164 from *Thompson et al.*
1234 [2005].

1235 **Table 2.** Summary of REE inverse modeling for 26 volcanic fields from ten geographic provinces. F =
 1236 cumulative melt fraction; T_P = potential temperature; SRP = Snake River Plain; GP = Great Plains; ETZ =
 1237 Eastern Transition Zone; STZ = Southern Transition Zone; SE = Sentinel Plain; NBR = Northern Basin and
 1238 Range; WTZ = Western Transition Zone; BR = Basin and Range; RGR = Rio Grande Rift; CP = Colorado
 1239 Plateau; A = Albuquerque; E = eastern SRP; G = Geronimo; GC = Grand Canyon; HB = Hopi Buttes; J =
 1240 Jornada del Muerto; L = Lucero; LC = Lunar Crater; M = Mormon Mountain; MD = Mojave Desert; MR
 1241 = Magic Reservoir; MT = Mount Taylor; N-E = north-eastern SRP; NN = Northern Nevada; O = Ocate;
 1242 P = Potrillo; RC = Raton-Clayton; SC = San Carlos; S-E = south-eastern SRP; SF = San Francisco; SP =
 1243 Springerville; Y = Yellowstone; W-C = west-central SRP; WP = Washington-Panguitch; ZB = Zuni-Bandera.
 1244 T_P WUSA16 refers to potential temperature calculated from shear wave velocity at 75 km depth.

| Province | Field | Age Ma | Latitude °N | Longitude °W | Depth km | F | T_P °C | T_P WUSA16 °C |
|----------|-------|-----------|----------------|-----------------|-------------|------|-------------|--------------------|
| SRP | E | 6.51 | 43.50 | -113.00 | 48–80 | 0.10 | 1354–1376 | 1372 |
| | MR | 9 | 43.00 | -114.20 | | | | 1371 |
| | N-E | 5 | 43.57 | -112.05 | | | | 1366 |
| | S-E | 7 | 42.94 | -111.33 | | | | 1333 |
| | W-C | 13 | 43.10 | -115.73 | | | | 1350 |
| | Y | 1 | 44.00 | -110.50 | | | | 1356 |
| GP | O | 3 | 36.03 | -104.93 | 53–83 | 0.07 | 1355–1361 | 1341 |
| | RC | 2.9 | 36.78 | -103.84 | | | | 1329 |
| ETZ | A | 0.2 | 34.83 | -106.90 | 57–77 | 0.06 | 1345–1357 | 1350 |
| | L | 1.3 | 34.95 | -107.21 | | | | 1347 |
| | MT | 2 | 35.00 | -108.50 | | | | 1328 |
| | ZB | 0.5 | 34.99 | -108.26 | | | | 1333 |
| STZ | SC | 1 | 33.34 | -110.39 | 54–79 | 0.06 | 1350–1352 | 1337 |
| | SP | 2 | 34.10 | -109.59 | | | | 1335 |
| | G | 0.26 | 31.00 | -109.30 | | | | 1321 |
| SE | SE | 3 | 33.05 | -113.02 | 53–78 | 0.06 | 1347–1354 | 1311 |
| NBR | NN | 4.5 | 40.51 | -117.12 | 54–69 | 0.04 | 1328–1331 | 1342 |
| WTZ | GC | 1 | 36.34 | -113.10 | 62–72 | 0.02 | 1327–1328 | 1335 |
| | M | 14 | 34.65 | -111.64 | | | | 1330 |
| | SF | 1 | 35.43 | -112.05 | | | | 1335 |
| | WP | 1 | 37.23 | -113.35 | | | | 1336 |
| BR | LC | 4.65 | 38.47 | -115.95 | 60–70 | 0.02 | 1322–1326 | 1337 |
| | MD | 2.7 | 34.91 | -115.90 | | | | 1330 |
| RGR | J | 1 | 33.41 | -107.05 | 59–72 | 0.02 | 1321–1322 | 1344 |
| | P | 0.2 | 31.97 | -107.16 | | | | 1343 |
| CP | HB | 2.1 | 35.38 | -110.25 | 62–72 | 0.01 | 1318–1322 | 1318 |

1245 **Figure 1.** (a) Topographic map of western North America where colored circles show spatial and temporal
 1246 distribution of mafic volcanism (sample locations with MgO >4 wt% are colored by age and taken from
 1247 NAVDAT database. Arrow = velocity of North American plate with respect to Pacific plate [26.8 ± 7.8 mm
 1248 yr^{-1} ; *Gripp and Gordon, 2002*]; Y = Yellowstone. (b) Long wavelength (500–4000 km) free-air gravity
 1249 anomalies [*Bruinsma et al., 2014*]. Mafic volcanism as before. (c) Map showing shear wave velocities at 100
 1250 km depth from SL2013NA tomographic model [*Schaeffer and Lebedev, 2014*]. Mafic volcanism as before.

1251 **Figure 2.** Locations of samples used in study. Small black circles = Cenozoic mafic samples from NAV-
 1252 DAT database; red circles = screened samples from Snake River Plain (SRP); pink hexagons = samples from
 1253 Great Plains (GP); orange stars = samples from Eastern Transition Zone (ETZ); light green right-pointing
 1254 triangles = samples from Southern Transition Zone (STZ); light blue diamonds = samples from Sentinel Plain
 1255 (SE); turquoise squares = samples from Northern Basin and Range (NBR); green inverted triangles = samples
 1256 from Western Transition Zone (WTZ); blue squares = samples from Basin and Range; dark green triangles =
 1257 samples from Rio Grande Rift (RGR); dark blue stars = samples from Colorado Plateau; gray circles = sam-
 1258 ples excluded due to suspected lithospheric contamination; black lines = physiographic regions [*Thompson*
 1259 *and Zoback, 1979*]; Y = Yellowstone. Database includes 77 samples from *White et al. [2004]*, *Thompson et al.*
 1260 *[2005]*, *Leeman et al. [2009]*, and *Plank and Forsyth [2016]* together with 100 new samples.

1261 **Figure 3.** Selection criteria for most primitive (i.e. asthenospheric) samples. (a) SiO₂ concentrations
 1262 plotted as function of MgO. Small black circles = Cenozoic samples from NAVDAT database; red circles
 1263 = samples from Snake River Plain (SRP); pink hexagons = samples from Great Plains (GP); orange stars =
 1264 samples from Eastern Transition Zone (ETZ); light green right-pointing triangles = samples from Southern
 1265 Transition Zone (STZ); light blue diamonds = samples from Sentinel Plain (SE); turquoise squares = samples
 1266 from Northern Basin and Range (NBR); green inverted triangles = samples from Western Transition Zone
 1267 (WTZ); blue squares = samples from Basin and Range; dark green triangles = samples from Rio Grande Rift
 1268 (RGR); dark blue stars = samples from Colorado Plateau (CP); gray circles = samples excluded after applying
 1269 selection criteria; dashed line = sample cut-off at MgO <7 wt% to exclude highly fractionated samples (note
 1270 that higher cut-off of 8 wt% was used for SRP and STZ samples). (b) CaO concentrations plotted as function
 1271 of MgO. Symbols and dashed line as before. (c) Fe₂O₃^T (i.e. total Fe expressed as Fe₂O₃) concentrations
 1272 plotted as function of MgO. Symbols and dashed line as before. (d) La/Ba ratios plotted as function of La/Nb
 1273 ratios. Symbols as before. Dashed ellipse = delineation of range of ratios for Ocean Island Basalts [samples
 1274 outside of ellipse deemed to be affected by lithospheric and/or subduction fluid contamination; *Fitton et al.,*
 1275 *1991*; *Stracke et al., 2003*].

1276 **Figure 4.** (a) Averaged trace element distribution of basaltic samples from Snake River Plain (SRP) chosen
 1277 in accordance with selection criteria (Figure 3). Compositions normalized to primitive mantle [*McDonough*
 1278 *and Sun, 1995*]. Red line with gray band = mean values for province $\pm 1\sigma$; pair of dashed lines = range of
 1279 compositions from all provinces; (b-l) Averaged trace element distributions for Great Plains (GP), Eastern
 1280 Transition Zone (ETZ), Southern Transition Zone (STZ), Sentinel Plain (SE), Northern Basin and Range
 1281 (NBR), Western Transition Zone (WTZ), Basin and Range (BR), Rio Grande Rift (RGR), and Colorado
 1282 Plateau (CP).

1283 **Figure 5.** Inverse modeling of screened samples. (a) Rare earth element (REE) concentrations for samples
 1284 from Snake River Plain (SRP) normalized to primitive mantle [*McDonough and Sun, 1995*]. Red circles with
 1285 vertical bars = average concentrations $\pm 1\sigma$; red line = best-fit concentrations calculated by inverse modeling.
 1286 (b) Trace element concentrations for SRP. Red circles with vertical bars = average concentrations $\pm 1\sigma$; red
 1287 line = concentrations predicted by forward modeling. (c) Melt fraction as function of depth. Red line = melt
 1288 fraction corrected for olivine fractionation obtained by fitting average REE concentrations shown in panel (a);
 1289 dashed line = same but uncorrected for olivine fractionation; solid black lines = isentropic curves calculated
 1290 using parameterization from *Katz et al. [2003]* and labeled according to potential temperature; vertical dashed
 1291 lines = phase transitions for spinel and garnet. Inset panel summarizes: (i) source composition where PM =
 1292 primitive mantle; (ii) average wt % of MgO plus its uncertainty; (iii) percentage of olivine fractionation; and
 1293 (iv) total melt thickness. (d-f) Great Plains (GP). Inset panel as before where source composition is now given
 1294 as percentages of Depleted MORB Mantle (DMM) and Primitive Mantle (PM) estimated from ε_{Nd} values
 1295 (Table 1). (g-k) Eastern Transition Zone (ETZ). (l-n) Southern Transition Zone (STZ). (o-q) Sentinel Plain
 1296 (SE).

1297 **Figure 6.** Inverse modeling of screened samples. **(a)** Rare earth element (REE) concentrations for sam-
 1298 ples from Northern Basin and Range (NBR) normalized to primitive mantle [*McDonough and Sun, 1995*].
 1299 Turquoise squares with vertical bars = average concentrations $\pm 1\sigma$; turquoise line = best-fit concentrations
 1300 calculated by inverse modeling. **(b)** Trace element concentrations for NBR. Turquoise squares with vertical
 1301 bars = average concentrations $\pm 1\sigma$; turquoise line = concentrations predicted by forward modeling. **(c)** Melt
 1302 fraction as function of depth. Turquoise line = melt fraction corrected for olivine fractionation obtained by
 1303 fitting average REE concentrations shown in panel **(a)**; dashed line = same but uncorrected for olivine frac-
 1304 tionation; solid black lines = isentropic curves calculated using parameterization from *Katz et al. [2003]* and
 1305 labeled according to potential temperature; vertical dashed lines = phase transitions for spinel and garnet.
 1306 Inset panel summarizes: (i) source composition given as percentages of Depleted MORB Mantle (DMM) and
 1307 Primitive Mantle (PM) estimated from ϵ_{Nd} values (Table 1); (ii) average wt % of MgO plus its uncertainty;
 1308 (iii) percentage of olivine fractionation; (iv) total melt thickness. **(d-f)** Western Transition Zone (WTZ). **(g-k)**
 1309 Basin and Range (BR). **(l-n)** Rio Grande Rift (RGR). **(o-q)** Colorado Plateau (CP). Inset panel as before but
 1310 source composition is now Depleted MORB Mantle (DMM) with 20% enrichment by small fraction melt
 1311 generated within garnet stability field.

1312 **Figure 7.** **(a)** Map of shear wave velocity at depth of 75 km from PM2012 tomographic model [*Priestley*
 1313 *and M^cKenzie, 2013*]. Small black circles = Cenozoic mafic samples from NAVDAT database; large black
 1314 circles = basaltic samples analyzed in this study; thin black lines = physiographic regions (Figure 2); Colored
 1315 squares = loci of velocity profiles shown in Figure 8. **(b)** Same for SL2013NA tomographic model [*Schaeffer*
 1316 *and Lebedev, 2014*]. **(c)** Same for DNA13 model [*Porritt et al., 2014*]. **(d)** Same for WUSA16 model [*Shen*
 1317 *and Ritzwoller, 2016*].

1318 **Figure 8.** Shear wave velocity and temperature profiles. (a) Shear wave velocity, V_s , as function of depth
 1319 for ten locations shown on Figure 7a with same color scheme [PM2012 tomographic model; *Priestley and*
 1320 *M^cKenzie*, 2013]. Red line = Snake River Plain (SRP); pink line = Great Plains (GP); orange line = Eastern
 1321 Transition Zone (ETZ); light green line = Southern Transition Zone (STZ); light blue line = Sentinel Plain
 1322 (SE); turquoise line = Northern Basin and Range (NBR); green line = Western Transition Zone (WTZ); blue
 1323 line = Basin and Range (BR); dark green line = Rio Grande Rift (RGR); dark blue line = Colorado Plateau
 1324 (CP); coarse dashed line = Western United States (WUS) reference velocity model [*Pollitz*, 2008]. (b) Tem-
 1325 perature as function of depth calculated from V_s profiles shown in panel (a). Colored scheme as before.
 1326 Dashed lines = isentropic curves labeled according to potential temperature, T_p ; horizontal line = maximum
 1327 crustal thickness of 50 km [*Shen and Ritzwoller*, 2016]; solid diagonal line = solidus for dry mantle peridotite
 1328 [*Katz et al.*, 2003]. (c-d) Same for SL2013NA tomographic model [*Schaeffer and Lebedev*, 2014]. (e-f) Same
 1329 for DNA13 [*Porrirt et al.*, 2014]. (g-h) Same for WUSA16 [*Shen and Ritzwoller*, 2016].

1330 **Figure 9.** (a) Map of lithospheric thickness calculated from $V_s(z)$ profiles for PM2012 tomographic
 1331 model [*Priestley and M^cKenzie*, 2013]. Black circles = volcanic activity younger than 10 Ma (NAVDAT).
 1332 (b) Map of potential temperature calculated from V_s values at 75 km depth for PM2012 model. (c-d) Same
 1333 for SL2013NA [*Schaeffer and Lebedev*, 2014]. (e-f) Same for DNA13 [*Porrirt et al.*, 2014]. (g-h) Same for
 1334 WUSA16 [*Shen and Ritzwoller*, 2016].

1335 **Figure 10.** Tomographic and geochemical temperatures. (a) Shear wave velocity, V_s , plotted as function
 1336 of La/Yb ratio for 177 individual samples (see Figure 2 for locations). Each value of V_s is averaged over
 1337 0.5° radius around each volcanic center at depth of 75 km, except for samples from CP, G (STZ) and SE
 1338 from PM2012 model that are averaged over 1.2° radius [*Priestley and M^cKenzie*, 2013]. Note that V_s values
 1339 >4.218 km/s were excised before averaging to mitigate effect of fast cratonic roots where no melting is ex-
 1340 pected to have occurred. Red circles = Snake River Plain (SRP); pink hexagons = Great Plains (GP); orange
 1341 stars = Eastern Transition Zone (ETZ); light green right-pointing triangles = Southern Transition Zone (STZ);
 1342 light blue diamonds = Sentinel Plain (SE); turquoise squares = Northern Basin and Range (NBR); green in-
 1343 verted triangles = Western Transition Zone (WTZ); blue squares = Basin and Range; dark green triangles =
 1344 Rio Grande Rift (RGR); dark blue stars = Colorado Plateau. (b) Potential temperatures calculated from V_s
 1345 anomalies at 75 km depth plotted as function of potential temperature calculated from geochemical inverse
 1346 modeling of rare earth element distributions. Colored symbols as in panel (a); horizontal and vertical error
 1347 bars = cumulative uncertainties for calculated potential temperatures; dotted line = 1:1 relationship for visual
 1348 guidance. (c-d) Same for SL2013NA [*Schaeffer and Lebedev*, 2014]. (e-f) Same for DNA13 [*Porrirt et al.*,
 1349 2014]. (g-h) Same for WUSA16 [*Shen and Ritzwoller*, 2016].

1350 **Figure 11.** (a) Horizontal slice at 100 km through SL2013NA tomographic model [Schaeffer and Lebedev,
 1351 2014]; small colored circles = heat flow measurements [Pollack *et al.*, 1993]; large colored circles = locations
 1352 of $V_s(z)$ profiles for each volcanic province shown in Figure 7 where red = Snake River Plain (SRP); pink =
 1353 Great Plains (GP); orange = Eastern Transition Zone (ETZ); light green = Southern Transition Zone (STZ);
 1354 light blue = Sentinel Plain (SE); turquoise = Northern Basin and Range (NBR); green = Western Transition
 1355 Zone (WTZ); blue = Basin and Range (BR); dark green = Rio Grande Rift (RGR); dark blue = Colorado
 1356 Plateau (CP); black line labeled $\mathbf{x-x'}$ = location of transect shown in panels (b-d). (b) Vertical slice through
 1357 SL2013NA model along transect shown in panel (a). Black line with gray band = topographic profile and
 1358 crustal thickness profile from Shen and Ritzwoller [2016]; dashed line = putative lithosphere-asthenosphere
 1359 boundary for western North America and craton from Hopper *et al.* [2018] and Priestley and *M^cKenzie*
 1360 [2013], respectively; colored circles = lithospheric thickness estimates from REE inverse modeling. (c) Gray
 1361 line and band = observed regional elevation $\pm 1\sigma$ within ± 10 km corridor along transect shown in panel (a);
 1362 dashed and pair of dotted lines = uplift $\pm 1\sigma$ calculated from shear wave velocity structure and lithosphere-
 1363 asthenosphere boundary from Hopper *et al.* [2018] and Priestley and *M^cKenzie* [2013]; red dashed line =
 1364 uplift calculated from temperature anomaly alone; colored circles = uplift $\pm 1\sigma$ calculated using results of
 1365 REE inverse modeling at locations shown in (a). (d) Gray line and small colored circles = averaged and spot
 1366 heat flow measurements within ± 100 km corridor along same transect; dashed and pair of dotted lines = heat
 1367 flow $\pm 1\sigma$ calculated from shear wave velocity structure and lithosphere-asthenosphere boundary from Hop-
 1368 per *et al.* [2018] and Priestley and *M^cKenzie* [2013]; large colored circles = heat flow $\pm 1\sigma$ calculated using
 1369 results of REE inverse modeling at locations shown in panel (a).

1370 **A: Analytic Procedures**

1371 65 samples from Arizona and Colorado were analyzed by X-ray fluorescence (XRF) on
1372 a Panalytical PW2404 wavelength-dispersive sequential X-ray spectrometer at the School of
1373 GeoSciences, University of Edinburgh. Method of analysis and estimates of precision are de-
1374 scribed in *Fitton et al.* [1998]. 280 samples were analyzed for trace elements on a PerkinElmer
1375 SCIEX Elan DRC II quadrupole ICP-MS Inductively Coupled Plasma-Mass Spectrometer (ICP-
1376 MS) at the Department of Earth Sciences, University of Cambridge. The method of analysis
1377 is similar to that used by *Eggins et al.* [1997], based upon the use of international rock stan-
1378 dards for matrix-matched calibration. The ICP-MS internal standards were 10 ppb Rh, In and
1379 Re and each sample was diluted 5000 times for analysis in 1% HNO₃. Under the conditions
1380 used, instrumental drift was less than 5% measured for the internal standard intensity during
1381 the entire analytical run (40 or more solutions per batch). Solutions were analyzed using a Mi-
1382 cromist nebulizer (FM05, Glass Expansion, Australia) and a quartz cyclonic baffled spray cham-
1383 ber with platinum sampler and skimmer cones. ICP-MS sensitivity for this configuration was
1384 5×10^4 cps/ppb In with CeO/Ce ratios = 2%. Appropriate corrections were made using ox-
1385 ide/metal correction factors calculated by analyzing pure single-element standard solutions. In-
1386 strument calibration was carried out using values from the GEOREM database (version 9, 2009;
1387 <http://georem.mpch-mainz.gwdg.de>) by analyzing matrix-matched United States
1388 Geologic Survey (USGS) rock standards BIR-1, AGV-1, BHVO-2, and BCR-2, which were
1389 dissolved using the same procedures as for samples. Concentrations were calculated on a spread-
1390 sheet where raw intensities were blank subtracted, internal standard normalized, and rare earth
1391 oxide corrected. The calibration method was a simple linear calibration curve fitted to calcu-
1392 lated slopes and the intercept was set at zero. All results (i.e. standards, unknowns) were ac-
1393 curately corrected for dilution by mass. Each sample was prepared by digesting 0.1 g of finely
1394 ground rock powder using 4 ml HF plus 1 ml HNO₃ in a sealed PFA vial. The acids used for
1395 sample preparation were ppb grade, which were further distilled in-house using Teflon or quartz
1396 stills. An Evapoclean (Analab, France) system consisting of a temperature-controlled Teflon-
1397 covered graphite block was used for digestions and evaporations within a closed, clean PFA
1398 environment thus avoiding the need for sample preparation to be carried out in a clean lab-
1399 oratory. Blanks and standards were prepared with each set of samples to monitor the quality
1400 of the sample preparation method. Total procedural blanks for all elements were very low, slightly
1401 higher than the ultra pure 1% HNO₃ rinse solution but negligible compared to sample inten-
1402 sities. External reproducibility, based on replicate analysis of standards and samples within batches,

1403 is 2–5% for all analytes. Accuracy for the analysis of rock standards such as BCR-2 during
1404 the run for most elements was within ~2% of the GEOREM-preferred values and better than
1405 5% (n=5) for the rest of the elements studied.





















