

Textural insights into the evolving lava dome cycles at Santiaguito lava dome, Guatemala.

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12 **Abstract**

13 The structures and textures preserved in lava domes reflect underlying magmatic and eruptive
14 processes, and may provide evidence of how eruptions initiate and evolve. This study explores the
15 remarkable cycles in lava extrusion style produced between 1922 and 2012 at the Santiaguito lava
16 dome complex, Guatemala. By combining an examination of eruptive lava morphologies and textures
17 with a review of historical records, we aim to constrain the processes responsible for the range of
18 erupted lava type and morphologies. The Santiaguito lava dome complex is divided into four domes
19 (El Caliente, La Mitad, El Monje, El Brujo), containing a range of proximal structures (e.g. spines)
20 from which a series of structurally contrasting lava flows originate. Vesicular lava flows (with a 'a
21 like, yet non-brecciated flow top) have the highest porosity with interconnected spheroidal pores and
22 may transition into blocky lava flows. Blocky lava flows are high volume and texturally variable with
23 dense zones of small tubular aligned pore networks and more porous zones of spheroidal shaped
24 pores. Spines are dense and low volume and contain small skeletal shaped pores, and subvertical
25 zones of sigmoidal pores. We attribute the observed differences in pore shapes to reflect shallow
26 inflation, deflation, flattening or shearing of the pore fraction. Effusion rate and duration of the
27 eruption define the amount of time available for heating or cooling, degassing and outgassing prior to
28 and during extrusion, driving changes in pore textures and lava type. Our new textural data when
29 reviewed with all the other published data allow a cyclic model to be developed. The cyclic eruption

30 models are influenced by viscosity changes resulting from (1) initial magmatic composition and
31 temperature, and (2) effusion rate which in turn affects degassing, outgassing and cooling time in the
32 conduit. Each lava type presents a unique set of hazards and understanding the morphologies and
33 dome progression is useful in hazard forecasting.

34 **1 Introduction**

35 The morphology and textures of lava preserve a story of the conditions and processes taking place
36 during its ascent. Recognising and understanding conditions for various extrusive and eruptive
37 products can provide insights to future timing of events and associated hazards (Calder et al., 2015).
38 At lava domes worldwide, lava morphology and surface textures have been shown to form in response to
39 several emplacement variables, including variations in effusion rate (Watts et al., 2002), emplacement
40 stress regime (Fink et al., 1990; Hale and Wadge, 2008), changes in volatile content (Anderson and Fink,
41 1990), crystallinity or composition (Watts et al., 2002), fragmentation processes (Wadge et al., 2009), and
42 sintering (Kendrick et al., 2016). Insights into the interacting variables are preserved in the crystal and
43 pore textures of the lava and further supplemented by compositional analysis and observations during
44 eruption. At dome volcanoes, the volatile content, and consequently bubbles, are considered
45 particularly important as the system is often on the boundary between explosive and effusive
46 behavior (e.g., Edmonds and Herd, 2007; Mueller et al., 2011).

47 Bubbles are considered stationary rather than mobile in highly-crystalline high viscosity magmas, as
48 the viscous forces dominate over the buoyant forces acting on the bubbles, promoting a coupled rise
49 of melt and bubbles (Sparks, 2003). Where magma is rising slowly, permeable magma foam
50 (Eichelberger et al., 1986; Jaupart and Allègre, 1991; Rust and Cashman, 2011), or cracked magma
51 (Tuffen et al., 2003; Lavallée et al., 2013) allows outgassing along the outer shear margin of the
52 magmatic column (Gaunt et al., 2014; Hornby et al., 2015) and/ or out into the surrounding country
53 rock, developing a partially or periodically open system (Castro et al., 2012; Cashman and Sparks,
54 2013; Kendrick et al., 2013). Evidence of continual outgassing and a periodically open system can
55 both be observed at Santiaguito (Stoiber and Rose, 1969; Johnson, 2004; Bluth and Rose, 2004;
56 Holland et al., 2011; Scharff et al., 2012; Ball et al., 2013).

57 Here we consider degassing as the shallow process of volatiles exsolving from the melt— typically
58 forming bubbles (pores). This process is controlled by the volatile overpressure and solubility, and
59 equilibrium conditions permitting bubble nucleation and growth (Cashman and Sparks, 2013). In

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60 contrast, we view outgassing as the process of the exsolved volatiles escaping from the lava, which
61 requires a permeable pathway. A prerequisite for the effusion of high-viscosity lava domes is the
62 partial development of open system degassing and outgassing (Calder et al., 2015). In an open
63 system, gas is able to escape from the magma, reducing gas pressure and decreasing the chance of
64 explosive eruptions (Cashman et al., 2000; Gonnermann and Manga, 2007; Holland et al., 2011;
65 Mueller et al., 2011). Both degassing and outgassing are affected by temperature, pressure, stress
66 conditions and the residence time of magma in the shallow conduit. In particular, residence time in
67 the conduit is important for high-viscosity magmas for which the physical processes associated with
68 bubble growth, coalescence and collapse, is sluggish.

69 Santiaguito comprises four domes (El Caliente, La Mitad, El Monje, El Brujo) which are aligned
70 along an overall E-W trend (Figure 1). Extrusion began in 1922, and has continued uninterrupted,
71 with nine documented cycles of extrusion (Harris et al., 2003; Rose, 1973b; Ebmeier et al., 2012).
72 Santiaguito has experienced a range of activity during this time including lava dome, lava flow and
73 spine extrusions. Extrusive activity is punctuated by explosive events which has resulted in
74 Vulcanian gas- and ash-plumes (Figure 1A) and occasionally pyroclastic density currents. El Caliente
75 vent is the only site to have been continuously active since 1922. The three other domes are mostly
76 inactive, except for mild gas emission, and are thus accessible for textural analysis. Evolution of the
77 domes has been extensively described and documented (Sapper, 1926; Williams, 1932; Stoiber and
78 Rose, 1969, 1970; Rose, 1972b, 1973a,b, 1987b; Rose et al., 1970, 1976; Smithsonian Institution,
79 1980–present; Anderson et al., 1995; Andres and Rose, 1995; Harris et al., 2002, 2003, 2004; Bluth
80 and Rose, 2004; Sahetapy-Engel et al., 2004, 2008; Sahetapy-Engel and Harris, 2008; Forbes, 2010;
81 Sanderson et al., 2010; Brill, 2011; Holland et al., 2011; Ebmeier et al., 2012) providing a first order
82 basis for the detailed textural and morphological analysis of lava dome emplacement attempted
83 herein. Extrusion cycles have been classified based on extrusion rate (Rose, 1973b; Harris et al.,
84 2003), but as yet classification of lava types and how they fit into the eruption sequence has not been
85 investigated. Thus, here we aim to classify lava types according to their texture and morphology,
86 place them in within the eruptive sequence, and explain the origins of these textural variations and
87 lava cycles.

88 **2 Geological and volcanological context**

89 Santa Maria is located on the southern edge of the Xela Caldera and is part of the Central American
90 Volcanic Arc which extends from Mexico to El Salvador (Figure 2; Duffield et al., 1993; Bennati et

91 al., 2011). Eruptive activity at Santa Maria began around 103 ka, and the edifice was constructed in
92 four stages before a 25 thousand-year period of quiescence prior to 1902 (Rose, 1987a; Rose et al.,
93 1977; Conway et al., 2013). Numerous lineations cut through the Santa Maria and Santiaguito
94 edifices with a rough east to west orientation (Figure 2; Escobar-Wolf et al., 2010).

95 In 1902 Santa Maria underwent a devastating Plinian eruption, generating at least 8.5 km³ dense rock
96 equivalent (DRE) of dacite as ash- and pumice lapilli-fall deposits, and leaving an explosion crater or
97 collapse scar on the southern base of the edifice (Figures 1 and 2; Rose, 1972a; Williams and Self,
98 1983, Singer et al., 2013; Andrews, 2014). After a period of quiescence of 20 years, volcanic activity
99 resumed in 1922, with the extrusion of Caliente dome in the crater excavated in 1922 (Figure 1B).
100 Volcanic activity has occurred without interruption, although at varying effusion rates. Thus far, four
101 domes have been constructed— El Caliente (1922–1939; 1972–present day), La Mitad (1939–1949),
102 El Monje (1949–1958) and El Brujo (1958–1986), extruding a combined total of ~1 km³ of magma
103 (Harris et al., 2003). Following dome growth at El Caliente, La Mitad, El Monje and El Brujo grew
104 sequentially, moving progressively to the west until activity renewed at El Caliente coincident with
105 activity at El Brujo in 1972 (Figure 1; Harris et al., 2003). The growth of the domes has been
106 described by Rose et al. (1970) and Rose (1972b, 1987b) and units mapped with respect to time by
107 Escobar-Wolf et al. (2010). El Caliente has been termed the main vent as it remained active with
108 intense fumarolic activity throughout growth of the lateral vents, and has been the only centre of
109 activity since 1977 (Rose, 1972b, 1987b).

110 2.1 Cyclic Lava Extrusion

111 This study builds on and supports the view of Harris et al. (2003) that lava extrusion is cyclic, with 3
112 to 6 year-long episodes of high extrusion rate (0.5–2.1 m³s⁻¹) followed by 3 to 11 year periods of low
113 extrusion rate (≤ 0.2 m³s⁻¹). El Caliente is the most active dome at Santiaguito, experiencing multiple
114 cycles of extrusion, at times coeval with activity at the other domes. Visual observations have
115 documented multiple active vents at Caliente (e.g. April 1967: Rose, 1973b). To date, nine different
116 cycles have occurred (Harris et al., 2003), with the latest emplacement of four lava flows >2 km in
117 length in 2011 to 2015 signaling the start of a new cycle. [Note that during the writing of this study
118 (late 2015 to early 2017), Santiaguito lava dome has experienced a period of heightened activity as
119 explosions have lessened in frequency, increased in magnitude and excavated a large crater inside El
120 Caliente, before resuming a lava dome growth phase.]

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121 Activity prior to 1929 was endogenous, characterised by subsurface build-up of magma to inflate and
122 uplift the carapace of the El Caliente dome. Between 1929 and 1958 a combination of endogenous
123 and exogenous behavior occurred. Post 1958 growth has been solely exogenous, where lava extrudes
124 at the surface (Rose, 1987b; Harris et al., 2003). In December 2012 (i.e., at the time of our field
125 campaign), activity consisted of regular explosions ($\leq 2 \text{ h}^{-1}$) producing gas- and ash-plumes reaching
126 heights of $< 1 \text{ km}$ (Lavallée et al., 2015; De Angelis et al., 2016), and simultaneous lava flow
127 extrusion accompanied by frequent incandescent rock falls. The recent observed behavior is
128 consistent with activity at Santiaguito throughout its growth.

129 **2.2 Lava Composition**

130 The chemical composition of Santiaguito dome lavas has been thoroughly analysed by Scott et al.
131 (2012, 2013), as well as by Rose (1972b), Avard and Whittington (2012), and Singer et al. (2013).
132 The first Santiaguito dome lavas extruded in 1922 were very similar to the 1902 eruption products:
133 porphyritic dacites dominated by phenocrysts of plagioclase feldspar ($\sim 20\text{--}30\%$), and less than five
134 percent in total of pyroxene, titanomagnetite, and sometimes amphibole (Rose, 1972b; Singer et al.,
135 2011; Scott et al., 2013). Although the phenocryst assemblage and abundance has not notably
136 changed over time, the bulk SiO_2 content of erupted magma has become progressively less evolved
137 with a decrease from 66 to 62 wt.% SiO_2 (Scott et al., 2013). Samples from Harris et al. (2003)
138 indicated that this decrease happened gradually post 1970, but more recent research by Scott et al.
139 (2013) suggest that the shift began as early as the 1930's and has been driven by magma mixing and
140 fractional crystallisation. Scott et al. (2012) suggest that the magma chamber developed in the lower
141 crust, 12 to 24 km deep, with no evidence of a shallow storage zone. A stratified magma chamber
142 model has been proposed with a vertical gradient in composition from a deep basaltic magma to a
143 shallow dacitic magma. This model may explain the progressive depletion in silica content in
144 eruptive products through time and the eruption of andesite at this stage (Scott et al., 2012).

145 Previous studies have highlighted that the various lava dome structures may be associated with
146 distinct geochemical signatures (e.g., Rose, 1972b; Avard and Whittington, 2012). Most lava
147 structures occur over the entire compositional range with a couple of exceptions. Spine structures
148 only seem to have developed from dacitic lava with 64 to 66 wt.% SiO_2 . Blocky lava flows that have
149 reached distances exceeding 2.5 km from the vent seem to be characteristic of andesitic lavas with 62
150 to 64 wt.% SiO_2 (Rose, 1972b; Harris et al., 2003; Scott et al., 2012, 2013). Although there is a
151 general trend associated with lava types and composition, various lava types and structures are

152 observed within short periods of constant magmatic composition, and throughout the eruptive history
153 at Santiaguito.

154 **3 Material and methods**

155 **3.1 Characterisation**

156 Mapping and sample collection at Santiaguito was carried out during a four-week field campaign in
157 November to December 2012. The outlines of units were based on high-resolution aerial
158 orthophotographs taken in 2006 by the Instituto Geographico Nacional (IGN) of Guatemala. Features
159 such as the outline of the main lava flows were adapted (in consultation with the authors) from Escobar-
160 Wolf et al. (2010) and Ball et al. (2013). The focus of this study was on mapping summit features such as
161 individual spines and lobes, their strikes and dips, and unique textural or structural features associated
162 with them. Samples were taken from a range of these features on the dome summits and the recent lava
163 flows, ranging in age from 1940 to 2012.

164 **3.2 Porosity**

165 The density and porosity of samples were measured by pycnometry of 25 mm diameter cylindrical
166 samples using a Helium Ultrapycnometer 1000 at Massey University, New Zealand. We used dry rock
167 equivalent densities of $2616 \pm 4 \text{ kg m}^{-3}$ for summit dome samples, and $2630 \pm 4 \text{ kg m}^{-3}$ for the 2012 lava
168 flows (the closest spatially and temporally corresponding units from Avard and Whittington (2012)).

169 **3.3 2D and 3D textural analysis**

170 Textural analysis was undertaken to constrain the range of physical attributes of the lavas erupted. This
171 was achieved quantitatively in 2D and 3D using a combination of techniques. Thin sections were
172 prepared from rocks impregnated with epoxy containing a fluorescent dye, which eased the observation
173 and delineation of pore space in UV light source. Complementarily, the crystal fraction was imaged under
174 plane polarised light (PPL). A combination of manual tracing and ImageJ processing were used to
175 threshold out and calculate the phenocryst crystallinity. Microlite crystal content required further analysis
176 and 3–4 mm samples were prepared for backscatter imagery at the Scanning Electron Microscope (SEM)
177 at the University of Canterbury. The microlites were traced out in Inkscape and quantified using ImageJ.

178 3D microtomography reconstructions aided with the interpretation of pore shapes from the
179 fluorescent thin sections and visually displayed their connectivity. Cross-sectional images of 10-mm
180 diameter samples were obtained using the Imaging and Medical beamline (IMBL) at the Australian
181 Synchrotron in Clayton, Victoria and reconstructed and analysed as 3D-image stacks in ImageJ. The

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182 scan energy of the beam was changed from 30–45 KeV as the samples varied in density. Pixel sizes
183 of 2D images were either 6.10 μm or 13.73 μm , equivalent to voxel sizes in reconstructed 3D stacks
184 of 227 μm^3 and 2,588 μm^3 .

185 3.4 Timing

186 The timing and interpretation of events was constructed from an exhaustive review of the literature and
187 was aided by bulletins and photographs from: Smithsonian Institution Scientific Event Alert Network
188 (SEAN) bulletins, Santiaguito Volcano Observatory (OVSAN) photographs, Instituto Nacional de
189 Sismología, Vulcanología, Meteorología e Hidrología (INSIVUMEH) volcanic alert bulletins and NASA
190 Landsat images. Photos and information from visiting and local scientists and observers also contributed
191 (Bill Rose, Jessica Ball, Kyle Brill, Rudiger Escobar-Wolf, Julio Cornejo).

192 4 Results

193 4.1 Lava type characterisation

194 Our historical summary, fieldwork, and textural analysis divide the lava types at Santiaguito into
195 three main types: vesicular lava flows, blocky lava flows and spines. Crystallinity (percentage of
196 phenocrysts and microlites) are consistent between the lava types (Table 1). Typically lava domes
197 with high crystallinities have pore shapes controlled by crystals, producing irregularly shaped pores
198 with ragged pore walls reflecting the inward protrusion of crystals (microlites and phenocrysts)
199 (Hammer et al., 2000; Kendrick et al., 2013). This is also the case for the Santiaguito lavas. Despite
200 this, a significant range of porosity and pore textures are exhibited and we use these in combination
201 with morphology to characterise the lava types.

202 Here, we present the morphology (Figure 3) and distribution (Figure 4), bubble and crystal textures
203 (Figures 5–7) and relationships between the lava types (Figure 8). A summary of results is provided
204 in Table 1.

205 4.1.1 Vesicular Lava Flows

206 These refer to lava flows with coherent, vesicular (a’-like, yet unbrecciated) flow top and a steep
207 rounded front (Figure 3). [Note that this is not easily classified by existing lava flow morphology
208 nomenclature, i.e it is not a’ or blocky, thus we are limited to the use of “vesicular lava” in this
209 study] Vesicular lava flows are restricted to the summit zone of the domes (Figure 4). An example of
210 vesicular lava was witnessed extruding from El Caliente after moderate dome and summit lava flow
211 collapse on November 28th 2012 (this field expedition; Figure 3A). In most cases vesicular lava flows

212 are overprinted by subsequent eruptions and not preserved (i.e. the November 2012 flow was
213 overprinted by unit Rcs). The units Rbe (erupted in 1978) and the beginning of Rma (erupted in the
214 1950's) are examples of vesicular lava flows (Figure 4). There are numerous historical accounts and
215 photographs of vesicular lava at Santiaguito, most notably following the explosion and dome collapse
216 in 1929, and when activity returned to El Caliente in 1972 (Anderson et al., 1995).

217 In our samples of vesicular lava from El Brujo (flow top and interior) porosity is high, ranging from
218 35 to 70% (Figure 5). The pores are spheroidal and interconnected, with dominantly convex edges
219 and with concave remnant pore wall protrusions (Figure 6). In thin section, pores reach a maximum
220 diameter of 10 mm in size, however tomography (Figure 6) and connected porosity data (Figure 5)
221 show that these pores form a connected network reaching total lengths of several centimeters and
222 extending beyond the dimensions of our 40-mm samples. Pores from the top of the flow are not
223 strongly aligned, though thicker horizontal bands of coalesced pores are present.

224 **4.1.2 Blocky Lava Flows**

225 Blocky lava flows generally have a coherent interior, and a top surface composed of smooth sided,
226 angular–subrounded blocks 0.1 m to 2 m in size (Figure 3). Blocky lava flows extrude from low
227 points in the active dome, however their thickness causes their upper surface to be ~10 m above the
228 vent rim. The flows reach 70 m thick, 500 m wide and vary in length from 400 m to 4 km (Figure 4).
229 These flows follow local depressions, typically river channels. Blocky lava flows vary significantly
230 in volume and have been divided up by length into summit blocky flows and long blocky flows for
231 some of our results (Figures 4 and 8).

232 Summit blocky lava flows are low-volume flows, which extend a short distance down the sides of the
233 dome, but never reach the base. The Units Rbf and Rbh (erupted in 1975) are examples of summit
234 blocky lava flows (Figure 4). These flows repeatedly collapse at the flow toe due to their low volume
235 and the high slope angle ($\geq 35^\circ$) from the eruptive summit to the flank of the dome, producing block
236 and ash flows. This collapse dynamic may have prevented the flow front from reaching the base of
237 the dome.

238 Long blocky lava flows are the highest volume end member, and largest volume extrusive products at
239 Santiaguito. The units Rbb (erupted in 1959–1963) and Rcm (erupted in 2001–2004) are examples of
240 long blocky lava flows (Figure 4). A single long blocky lava flow may continue growing for a few
241 years and as a result these flows have the most pronounced ogives (pressure ridges) and levees

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242 (Figures 3 and 4). Regular small block and ash flows and avalanches ($>1 \text{ min}^{-1}$, reaching up to 400
243 m) initiate from collapse of both the flow front and sides. Flow fronts of long flows may collapse to
244 form block and ash flows reaching kilometers from the flow front such as in 1976 (Rose et al., 1976).

245 Vesicularity is heterogeneous, at both outcrop scale and micro scale. The top 5–10 m of long blocky
246 lava flows are more vesicular than the bulk of the interior as seen on La Mitad (this study) and on El
247 Brujo (Rose 1972). This vesicularity difference is reflected in the samples from the blocky top (3–
248 79%; Figure 5). Pores range in shape and size between and within samples, including larger round
249 pores as seen in vesicular lava flows, and slim tube shaped interconnected pores with varying degrees
250 of flattening and shearing that are unique to blocky lava flows (Figure 6).

251 As with vesicular lava flows, in thin section pores reach a maximum diameter of 10 mm; however
252 tomography and connected porosity show that these pores form connected networks reaching total
253 lengths of several centimeters and extending beyond the dimensions of our 40-mm samples (Figures
254 5 and 6). Pores in the samples with the highest porosity have dominantly convex pore surfaces with
255 concave remnants of bubble-wall protrusions, which indicate pore coalescence and/ or collapse by
256 shearing. These convex pores are the largest pores in blocky lava flows and are usually connected to
257 other large pores. Many of these interconnected pores show preferential long-axis alignment, and are
258 more aligned than the vesicular lava flow pores.

259 In the densest samples the pores have smaller maximum diameters. Some pores have both concave
260 and convex surfaces and evidence of remnant pore-wall protrusions, but have little alignment, which
261 are similar to some pores seen in spines. However, in most cases this pore morphology is aligned,
262 with dominantly convex pore walls and little evidence of pore-wall protrusions. These pores form
263 networks of interconnected tubes, and are unique to blocky lava flows. All pore types are dominantly
264 located adjacent to phenocrysts and lithic inclusions (Figure 6). Blocky lava flows can be preceded or
265 followed by spine extrusion.

266 4.1.3 Spines

267 Spines are dense, coherent masses of lava that extrude sub-vertically above the vent along linear
268 fault-controlled planes (Figure 3). They are low volume, and limited to the summit zone of the
269 domes. The largest spines reach 200 m long, 50 m wide and 70 m tall and the smallest spines are only
270 3 m x 5 m (Figure 4). The spines exposed on the domes today are commonly characterised by zones
271 of fractured dense material ($>3 \text{ m}$ wide), separated by regularly spaced subvertical scoriaceous or

272 brecciated zones (<1 m, here termed “shear zones”) (Figure 3). Observations of spine extrusion at
273 Santiaguito report that sections of spines regularly collapse forming block and ash flows that reach
274 the base of the dome, or are repeatedly disrupted by explosive events (Williams, 1932; Rose, 1973b).

275 Most spines appear to have extruded near-vertically, whereas a few spines have curvilinear surfaces
276 that appear drooped and folded exhibiting some features similar to whaleback structures at
277 Monsterrat described by Watts et al. (2002). The difference between these low angle spines and the
278 lava flows is that they are much denser and have shear zones on the outer edges. Similar features
279 have been described at Unzen volcano (Smith et al., 2001; Hornby et al., 2015), Soufriere Hills
280 (Watts et al., 2002) and Mount St. Helens (Gaunt et al., 2014; Kendrick et al., 2014).

281 The largest of the exposed spines form the prominent ridges on La Mitad and El Monje domes (units
282 Re (erupted in 1931) and Rm (erupted in 1950); Figure 4). These spines tend to have multiple widely
283 spaced shear zones marked by more vesicular rinds at the edge of each dense core (Figure 3).
284 Commonly a vertical gap is observed in the centre of, and parallel to, a shear zone. The gaps widen
285 towards the top of the spine, implying that the denser sections of the spines may have moved apart
286 laterally at the weaker shear zones during extrusion. On a larger scale the major shear zones
287 interweave. Note that shear zones are distinct from cooling joints, which are not marked by a change
288 of texture or vesicularity.

289 The inside of the spines are dense, but the outside surface may either be smooth, or have a thin
290 brecciated or vesicular rind— “shear zones” (Figure 3). The dense zones within the spines have low
291 porosity (4–22%; Figure 5), and the pores are similar in shape and size between samples. In thin
292 section the pores reach a maximum of 5 mm in diameter; however tomography and connected
293 porosity data show that the pores form large interconnected networks within the 40-mm sample size.

294 The most common pore type is very thin, skeletal-shaped interconnected pores (Figure 6). These
295 have both convex and concave pore surfaces, evidence of remnant bubble-wall protrusion, and the
296 largest examples always border phenocrysts. Tomography shows that these pores form skeletal-like
297 networks of interconnected pores that are aligned extrusion-parallel in our samples. In contrast to the
298 thin tubular pores in the lava flows, the pore networks predominantly align in thin, tightly spaced
299 vertical planes, with lesser lateral connectivity than vertical. Occasional convex surfaces, bulbous
300 pores, and pore-wall remnants characteristic of more round interconnected pores are observed within
301 the porous network.

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302 Shear zones within the spines generally have a higher porosity (11–79%) than the dense spine
303 interior (4–22%; Figure 5) and are texturally variable. The shear zones can broadly be differentiated
304 into two end members: vesicular and brecciated. Tomography and thin-section show that the
305 vesicular samples have an en-echelon pattern of interconnected sigmoidal pores (Figure 7). The
306 sigmoidal pores have convex and concave pore surfaces. Pores may be large and widely spaced or
307 thin and closely spaced and the largest examples always border phenocrysts and have evidence of
308 remnant bubble-wall protrusion.

309 In thin section the shear zone pores reach a maximum of 10 mm in diameter; however tomography
310 and connected porosity data show that the pores form interconnected networks within the 40-mm
311 sample size (Figures 5 and 7). Small sigmoidal interconnected pores also form along flow bands
312 within the dense spine interior. In thin section the dense interior shows local en-echelon bands with
313 remnants of previous rounded or thin interconnected pores. The length of the chain often extends
314 beyond the length of the 40 mm thin sections and thus cannot account for the full scale of the feature
315 observed in the field. Yet, the geometrical relationship observed in the micro- as well as the macro-
316 textures, such as S-C fabrics (showing the intersection of S-planes oblique to shear surface and C-
317 planes parallel to shear surface), sigmoidal pores and crystal alignments suggest simple shear in these
318 regions of strain localisation.

319 The brecciated shear zones host the dense rocks from the spine interior. We note that the porosity
320 increases slightly and the crystal size decreases along the edge of the spine. The clasts in the shear
321 zones are generally highly brecciated at the macro-scale (clasts 1–30 mm in size) and partially
322 fragmented internally at the microscopic scale. This may be analogous to the cataclasite zones
323 described at Mount St. Helens (Pallister et al., 2012) and have similar properties to those described in
324 the 1994–95 spine erupted at Unzen (e.g., Smith et al., 2001).

325 4.2 Timing of lava types at Santiaguito

326 The characterisation of lava types allowed a re-examination of the timing and progression of lava
327 types at Santiaguito. This synthesis of data is presented in Figure 7, with an accompanying map of all
328 the currently exposed extrusive products in Figure 4. The timeline builds on and supports Harris et al.
329 (2003) view that lava extrusion is cyclic.

330 A typical eruption cycle begins with an increase in extrusion rate and the early extrusion of spines,
331 often from a new vent. Spine growth is generally preceded by or coincident with endogenous growth

332 as the extrusion rate was low but steadily increasing. As the extrusion rate increases, lava flows of
333 increasing length are emplaced. Towards the end of a cycle the extrusion rate declines again, and the
334 trend is reversed culminating in spine growth. The next phase of activity typically resumes at a new
335 location.

336 Over time the lava type end members in the cycles changed. Prior to 1980, alternation between spine
337 formation and blocky lava flows of moderate lengths dominated the eruptive sequence. Since 1980,
338 blocky lava flows of variable lengths were erupted from the summit of Caliente dome. Blocky lava
339 flows in excess of 2.5 km in length were not extruded prior to 1965 and have become increasingly
340 dominant since. This transition has coincided with the less frequent spines, becoming significantly
341 smaller in size and volume, and no spines have been extruded since 1990.

342 Infrequent larger eruptions and dome or lava flow collapse events have caused larger ash plumes (<6
343 km) and block and ash flows (Rose, 1973a; Rose et al., 1976; Fink and Kieffer, 1993). The most
344 notable of which occurred in 1929 when a collapse and explosion at El Caliente sent a pyroclastic
345 density current 11 km downstream, killing ~ 5000 inhabitants in the town of El Palmar (Rose et al.,
346 1976; Simkin et al., 1994). Vesicular lava often follows collapse events.

347 The timeline highlights how lava types correspond with and respond to the extrusion rate and
348 eruption history, including the effects of slope, cooling, degassing and crystallization during
349 emplacement.

350 **5 Discussion**

351 **5.1 Controls on lava dome morphology**

352 This review, synthesis and new data concerning timing of lava types has revealed remarkable cycles
353 in lava extrusion style. Lava domes may extrude varying lava styles, from endogenous lobes to
354 exogenous lobes or spines (e.g., Manley, 1996; Fink and Griffiths, 1998; MacKay et al., 1998;
355 Nakada et al., 1999; Harris et al., 2002, 2004; Navarro-Ochoa et al., 2002). The development of
356 varying lava styles depends on lava viscosity and extrusion rate (Anderson and Fink, 1989; Swanson
357 and Holcomb, 1990; Zobin et al., 2002; Watts et al., 2002; Pallister et al., 2012), which dictates the
358 development of the shear zones that control extrusion dynamics (e.g., Lavallée et al., 2007, 2013;
359 Tuffen et al., 2013; Kendrick et al., 2014). We attribute the changes in morphology to (1) an evolving
360 source composition and temperature decrease over time (Scott et al., 2013), and (2) conduit processes

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361 such as degassing, outgassing and strain rate driven viscosity changes and the evidence for this is
362 preserved in the pore shapes and sizes.

363 Several variables have been suggested to control the viscosity of dome lavas; principally melt
364 composition and temperature (Giordano et al., 2008; Mueller et al., 2009) as well as crystallinity
365 (Caricchi et al., 2007), porosity (Caricchi et al., 2011) and strain rate (Cordonnier et al., 2009;
366 Lavallée et al., 2007, 2013; Kendrick et al., 2013).

367 Consistent with the bulk rock composition, interstitial glass analyses of Santiaguito lavas display a
368 large range in composition (rhyolite–trachyte–dacite), broadly decreasing in SiO₂ with time, although
369 the SiO₂ content of the glass in individual samples (within a given eruptive unit) may range by up to
370 10 wt.% (Scott et al., 2012). Along with the composition, the temperature of the reservoir feeding
371 Santiaguito has likely increased (Table 2). The early Santiaguito lavas were very similar to the 1902
372 eruption products (Rose, 1972b; Singer et al., 2011; Scott et al., 2013) and may represent a pocket of
373 left-over dacite in the magma reservoir (e.g., Scott et al., 2013; Singer et al., 2013); thus we assume
374 that they erupted at a temperature similar to the Santa Maria eruption of around 850 °C (Scaillet et
375 al., 1998; Singer et al., 2013; Andrews, 2014;). Over time as progressively more andesitic magma
376 was erupted we expect that the temperature of the magma increased, to reach a maximum of 950 °C,
377 as estimated by Sahetapy-Engel et al. (2004), and supported by Scott et al. (2013) based on
378 amphibole geothermometry. Similarly, there have been <8% changes in crystal fractions and Scott et
379 al. (2012) hypothesised that crystallisation may be limited at shallow depths due to a “final quench”
380 where microlites stop nucleating and growing.

381 To a certain extent, magma chamber stratification and resultant shift in composition and temperature
382 has influenced the lava structures and extrusion cycles over time. In particular spines only formed
383 when dacitic lava was extruded and flows extending over 2.5 km only formed in andesitic lavas.
384 However, various lava types and structures are observed within short periods of similar and arguably
385 relatively constant pre-eruptive composition, temperature and crystallinity, and indicate that the pre-
386 eruptive conditions (i.e the physico-chemical state in the reservoir before ascent) are not the only
387 control on lava morphology.

388 Our evidence suggests that in addition to composition, conduit processes such as degassing, pore
389 configurations, strain rate and thermal shifts during ascent also drive the viscosity changes
390 responsible for the variations in dome eruption style at Santiaguito. The relatively open magmatic

391 system (Bluth and Rose, 2004; Holland et al., 2011) and the slow magma ascent likely allowed for
392 efficient outgassing through the interconnected porous network, which is observed in all lava types
393 (Figure 6), and obvious from continuous gas emission observations at Santiaguito. The amount of
394 dissolved water in the melt during extrusion is unknown at Santiaguito owing to a number of
395 challenging issues arising in these dome lavas. As such, here, we turn to pore textures as evidence of
396 degassing and outgassing and a schematic model of pore development is presented in Figure 9.

397 We attribute the observed differences in pore shapes and volume to be initially reflecting shallow
398 inflation, deflation, and coalescence of the exsolved bubble volume. Inflated interconnected pores
399 such as in vesicular lava flows and high porosity zones of blocky lava flows show complex pore
400 shapes and remnant pore wall protrusions from coalescence. In deflated interconnected pores such as
401 in spines and dense zones of blocky lava flows, the remnant pore-wall protrusions are smoothed as
402 decreasing gas pressure due to outgassing allows surface-tension-driven deformation (e.g., Kennedy
403 et al., 2016). Inflated pores likely indicate a lower viscosity and perhaps a residual water content in
404 the vesicular and blocky lavas.

405 Pores may have additionally undergone a degree of flattening and stretching, indicating preferential
406 compaction or shear (Rust et al., 2003; Wright et al., 2006; Kendrick et al., 2013; Ashwell et al.,
407 2015). Stretching of deflated pores results in a network of tube-shaped pores with a preferential
408 alignment in lava flows, or skeletal-shaped pore networks that are aligned along “extrusion-parallel”
409 planes developed in spines (Figure 6). Stretching of inflated pores produces sheet-like pores that
410 form flow bands in the lavas. Shear can additionally facilitate outgassing and deflation by increasing
411 the connectivity between pores (e.g., Okumura et al., 2009; Caricchi et al., 2011). This was shown
412 experimentally on Santiaguito samples by Avard and Whittington (2012) at strain rates representative
413 of long blocky lava flows, and is observed in our dense zones of blocky lava flow samples (Figure 6).

414 A distinct form of flow bands with en-echelon sigmoidal pores form along spine margins (Figure 7).
415 These are characteristic of dilation where brittle failure of magma results in tension gashes that rotate
416 into sigmoidal shapes and are influenced by phenocryst and remnant pore location (Kendrick et al.,
417 2012; Lavallée et al., 2013). The ragged pore walls in places indicate the magma was torn slowly
418 (unlike straight failure planes produced at high shear rates; see Lavallée et al., 2013). Such tearing
419 structures have also been observed at Unzen, Yakedake and Daisen by Smith et al. (2001). The high
420 porosity and connectivity of these shear zones likely contribute to outgassing. The considerable

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421 evidence of pore collapse in the dense core of spines, bordered by elongate, shear-driven porous
422 bands in these lavas suggest efficient outgassing during ascent and extrusion.

423 5.2 Advances to the understanding of cyclic lava extrusion at Santiaguito

424 Here we incorporate our observations and textural analysis with previously published models to build
425 on the growing understanding of cyclic lava extrusion at Santiaguito. The textures and eruption
426 history support a model of a gradually changing reservoir composition and temperature, which is
427 modified by conduit outgassing and localised shear, to dramatically change the viscosity and eruption
428 style of the lava. The timing of the lava types illustrated in Figure 8 highlights how lava types
429 correspond with and respond to the extrusion rate and extrusion history documented by others.

430 At Santiaguito, extrusion rate is cyclical with 3 to 6 year-long phases of high extrusion rate ($0.5\text{--}2.1$
431 $\text{m}^3 \text{s}^{-1}$), followed by a longer (3–11 year) phase of low extrusion rate ($0.2 \text{m}^3 \text{s}^{-1}$) (Harris et al., 2003).
432 Extrusion rate determines the time in the shallow conduit available for (1) outgassing, reflected in the
433 pore volume and structures, and (2) cooling, influencing viscosity. Lavas at Santiaguito have a
434 viscosity measured to range between $\sim 10^9$ and 10^{12} Pa.s (Harris et al., 2002; Avard and Whittington,
435 2012), with spine structures likely on the upper end of this, as at other domes worldwide spines are
436 generally lower temperature and higher viscosity than flows (Nakada and Motomura, 1999;
437 Schneider and Vallance, 2008; Cordonnier et al., 2009). The lava types reveal a progression from
438 porous short flows with inflated pores and coherent flow tops through long lava flows with varied
439 pore morphology and blocky flow tops to spines with deflated pores and well developed porous shear
440 bands. A model for lava type progression is presented in Figure 10.

441 Activity at Caliente, Mitad and Monje all commenced with endogenous growth (Figure 10).
442 Although we do not have good time constraints of endogenous growth it probably accompanied all
443 lava types prior to 1958, and we refer to this phase as ‘leaky’ endogenous. Growth at El Caliente was
444 mostly endogenous until the second extrusion cycle began in 1929 (Harris et al., 2003).

445 Perhaps the most notable event in extrusion cycle II was the November 1929 dome collapse, as the
446 extrusion rate was increasing at El Caliente. The collapse was followed by the extrusion of vesicular
447 lava.

448 The most porous lava type, vesicular lava, is extruded after large eruptions or dome collapses (Figure
449 10). The clearing of overlying, more-degassed lava allows fresh magma to ascend relatively quickly,
450 without significant degassing; as a result the lavas vesiculate late at (or near) the surface forming

451 spheroidal vesicles. The vesicular lava flow extrusions coincide with less regular gas explosions, e.g.,
452 after the November 2012 collapse. Upon continued extrusion and flow, the flow top of the initially
453 summit-limited, vesicular lava (Figure 11a) apparently evolves in time and space into blocky lava
454 (Figure 11b). Simultaneously in the flow core of the blocky lavas, vesicles progressively collapse and
455 outgas during shearing and flow advancement. In Figure 11c, we show that other proximal lavas
456 exhibit a moderate degree of flow-top brecciation as they spill from a collapse scar with a surface
457 morphology intermediate between the coherent flow tops of the vesicular lavas and the blocky lavas.

458 Blocky lava flows extrude at high and low extrusion rates and have a connected network of inflated
459 and deflated pores. The longest lava flows were emplaced during constant, high-extrusion rates that
460 can last up to 2 years (Figure 8; Harris et al., 2003). Eruptions and fumarolic activity occur
461 simultaneously with long blocky lava flow extrusion (Rose 1973a; Bluth and Rose 2004; Johnson
462 2004; Brill 2011; Holland et al. 2011) indicating outgassing in conduit. Outgassing continues as the
463 lava flows downslope. This is evidenced by further vesicle reconfiguration, localised shear flattening,
464 stretching and deflation (e.g., Manley and Fink, 1987; Avarad and Whittington, 2012, Ashwell and
465 Kendrick et al., 2015; Kennedy et al., 2016) in the flow core. Gas percolates through the permeable
466 network locally creating layers of highly vesicular lava immediately below the dense blocky top.
467 These spatially heterogeneous processes are reflected in the spatial heterogeneity of porosities
468 (vesicles and fractures) in the flows. At any time during flow, lava may be quenched at the flow front
469 or edges, preserving a complex history of pore inflation, deflation and shear from different parts of
470 the flow.

471 During periods of low extrusion rates, lava flows of decreasing length develop, and prior to 1990 the
472 dacitic composition favoured conditions suitable for spine growth (Figure 8). Spines precede and
473 follow blocky lava flows and have dense interiors containing small, connected, deflated pore
474 networks. Intermittent planar vesicular shear zones with en-echelon sigmoidal pores separate the
475 dense interiors. The low extrusion rate allows time for prolonged degassing, outgassing (evidenced
476 by pore deflation) and cooling in the conduit likely causing viscosity increases that may prompt spine
477 extrusion (Figure 10). Extrusion is accompanied by mild steam eruptions (Williams, 1932; Rose,
478 1973b), likely initiated along the planar vesicular shear zones, further aiding outgassing.

479 The recent andesitic eruptive period (1990–2016) switches between the extrusion of vesicular and
480 blocky lavas with length corresponding to extrusion rate, dome collapses and eruptions. Hence the

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481 progressive change in the geochemistry of the erupted lavas in the last century from dacite to andesite
482 (e.g., Scott et al., 2013) may have influenced the temporal occurrence of switches in eruptive activity.

483 **6 Conclusions and Implications**

484 Our new mapping and textural data allowed us to categorise lava types at Santiaguito. Following this
485 an extensive review of all lava types and timing was conducted to create a timeline of extrusive
486 activity at Santiaguito.

487 (1) There are three main lava types at Santiaguito: vesicular lava flows, blocky lava flows and
488 spines. These types reveal a progression from porous short flows with inflated pores through
489 to long lava flows with varied pore morphology to spines with deflated pores and well
490 developed porous shear bands. There is no significant differences in crystal content between
491 these lava types.

492

493 (2) The timing of the lava types highlights how they may respond to the extrusion rate and
494 extrusion history (including the effect of slope, cooling and degassing). Blocky lava flows of
495 increasing length develop at high extrusion rates and spines extrude at the beginning and end
496 of an extrusion cycle at low extrusion rates. A collapse or eruption may cause a shift to
497 extrusion of vesicular lava flows.

498

499 (3) Our new textural data when reviewed with all the other published data allows a model to be
500 developed: The model demonstrates the importance of (1) initial magmatic composition and
501 temperature, and (2) effusion-rate-driven degassing, outgassing and cooling in the conduit,
502 and hence viscosity, in controlling the eruption style and associated lava types.

503 Each lava type presents a unique set of hazards:

504 (1) Vesicular lava has the highest temperature and may correlate with higher volatile contents,
505 and therefore the greatest potential for decompression-driven explosive fragmentation.

506 However, they are short lived and low volume following vent-clearing eruptions or collapses
507 until conditions for blocky lava flows again prevail.

508

509 (2) Blocky lava flows produce regular small collapses from the flow fronts and flow sides. The
510 higher volume the flow (i.e. the higher the extrusion rate), the greater the hazard potential.

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511 Flow fronts of long flows may collapse to form block and ash flows such as in September
512 1976 (Rose et al., 1976). In addition, large eruptions historically occur during an increase in
513 extrusion rate and blocky lava flow development (i.e. November 1929, April 1973, November
514 2012). Loose debris associated with large volume flows is transported down river channels by
515 destructive lahars (Harris et al., 2006).

516

517 (3) Spines are cooler and degassed prior to extrusion and collapse regularly along fractures
518 during growth, producing a broad talus blanket and historically only small volume block and
519 ash flows. Spines may be destroyed by subsequent vent-clearing eruptions.

520

521 Eruptions, small pyroclastic flows and collapses are associated with all lava types. However, the
522 trends presented above provide an added understanding on the development of lava types and
523 sequencing useful in continued hazard forecasting.

524

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777 **Tables**

778 Tables should be inserted at the end of the manuscript. Tables must be provided in an editable format
779 e.g., Word, Excel. Tables provided as jpeg/tiff files will **not be accepted**. Please note that very large
780 tables (covering several pages) cannot be included in the final PDF for reasons of space. **These**
781 **tables will be published as [Supplementary Material](#) on the online article page at the time of**
782 **acceptance. The author will be notified during the typesetting of the final article if this is the**
783 **case.**

784

785 **Table 1: Summary of lava properties by type**

	Vesicular Summit Lava Flow	Blocky Lava Flow	Spine
Description	Short very vesicular flow with a rounded coherent top	Short to long lava flow with mixed porosity and a blocky top	Dense sub-vertically extruded units
Rough Volume	Low <300 000 m ³	Low - High 200 000–1 600 000 m ³	Low <400 000 m ³
Dip	Low 5°–30° With steep dipping flow fronts (approx. >50°)	Low 5°–30° With shallow 32° ^v to steep 70° dipping flow fronts	High 38°–90°
Porosity (range)	35–70%	3 [†] –79%	4–23%
Pore Textures	Round, interconnected pores	Zones of larger round and smaller tubular interconnected pores with evidence of flattening and stretching	Small skeletal interconnected pores (primary) and dilational pores with ragged pore walls (minor)
Crystallinity	31–34%	30–35%	27–34%
Microlites (% of groundmass)	37–40%	40–43%	41–45%
Composition	62–65 wt.% SiO ₂	62–65 wt.% SiO ₂ Note blocky lava flows that exceed distances of 2.5 km from the vent: 62–64 wt.% SiO ₂	64–65 wt.% SiO ₂
Extrusion Rate	High	Low – High (0.2–2.1 m ³ s ⁻¹)	Low (≤0.2 m ³ s ⁻¹).
	[†] Avard and Whittington (2012)	^v Harris et al. (2006)	

786

787 **Table 2: Temperatures at Santiaguito.**

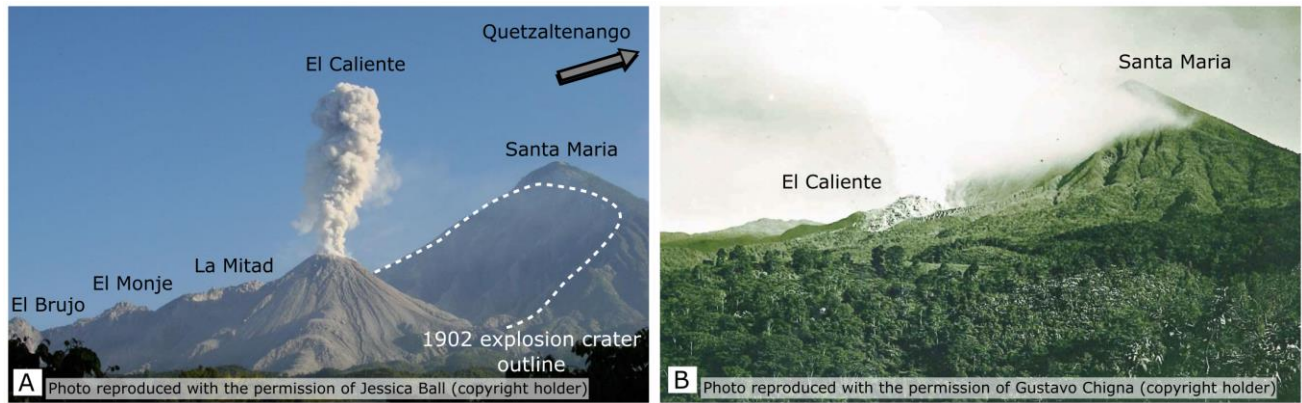
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Reference	Temperature	Extrusive Unit and Method	Eruptive/ Pre-eruptive Temperature
Andrews (2014)	840–850 °C	Phase equilibria experiments of natural pumice and phenocryst compositions of the 1902 Santa Maria dacite.	Pre-eruptive (Santa Maria)
Singer et al. (2013)	870 °C	Fe-Ti oxide + two pyroxene geothermometry and melt inclusion compositions.	Pre-eruptive (Santa Maria)
Scott et al. (2013)	900–1000 °C	Phenocrysts from Santa Maria pumice using the thermobarometer of Ridolfi, et al. (2010).	Pre-eruptive (Santa Maria)
Scott et al. (2012)	940–980 °C (± 20 °C)	Matches amphibole rim widths and best represents geochemistry conditions at Santiaguito over its history.	Pre-eruptive (Santiaguito)
Sahetapy-Engel et al. (2004)	850–950 °C	Maximum temperatures calculated at the surface of the El Caliente vent in 2002 using an infrared thermometer, spectroradiometer and digital video camera.	Eruptive at vent (Andesite, high eruption rate and extrusion of 4km blocky lava flow)
Harris et al. (2002, 2003)	496–531 °C	Highest direct measurements taken of the 1999–2002 (Rcm) flow front core 2.5 km from vent. Note: used eruption temperatures 833 °C from Scaillet et al. (1998) for thermal modelling.	Eruptive at the flow front (Andesite, high eruption rate and extrusion of 4km blocky lava flow)
Scaillet et al. (1998)	833 °C (800–850 °C)	Fe-Ti oxide geothermometry on the Santa Maria pumice. Used method from Ghiorso and Sack (1991).	Pre-eruptive (Santa Maria)

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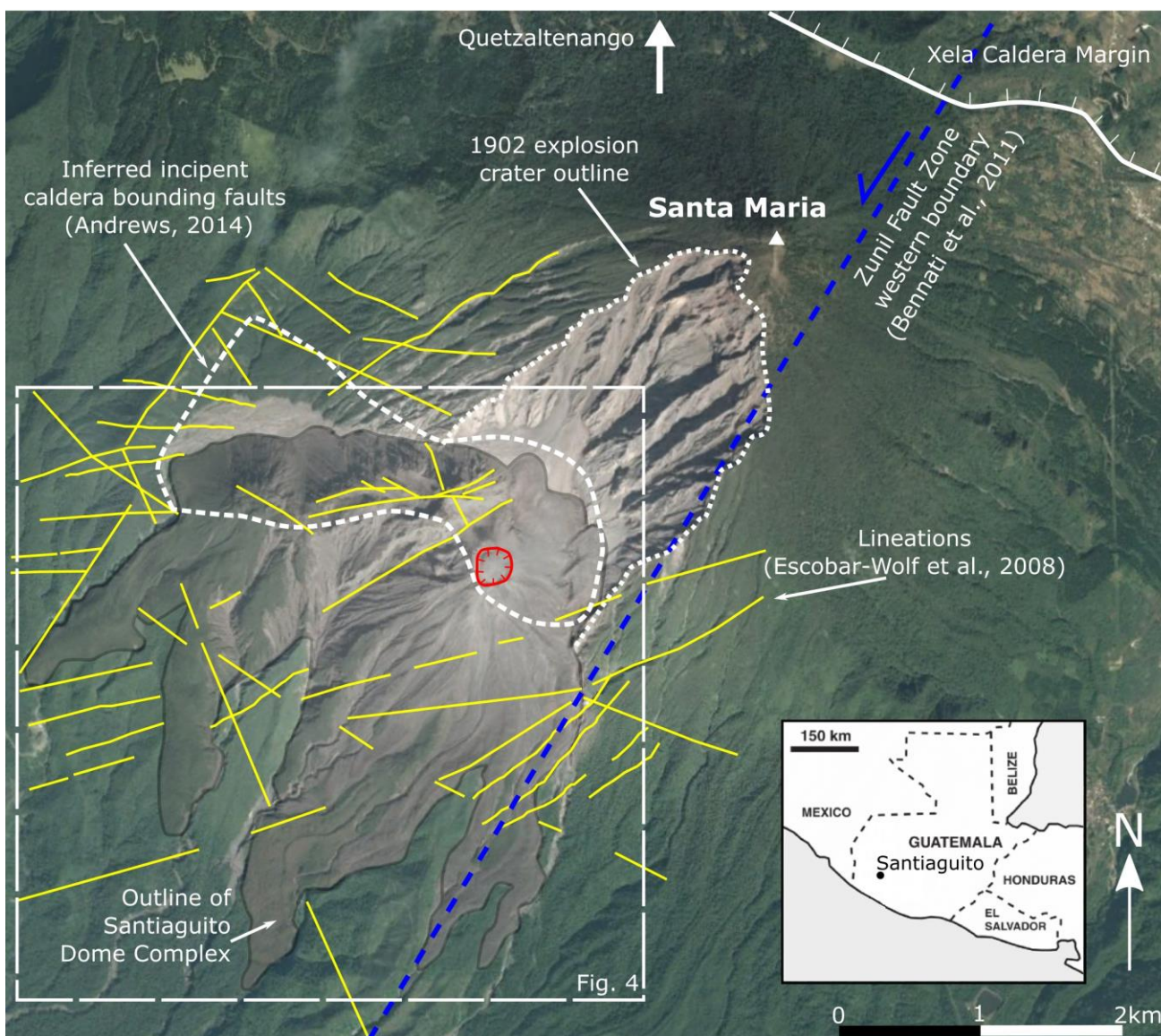
791 6.1 Figures



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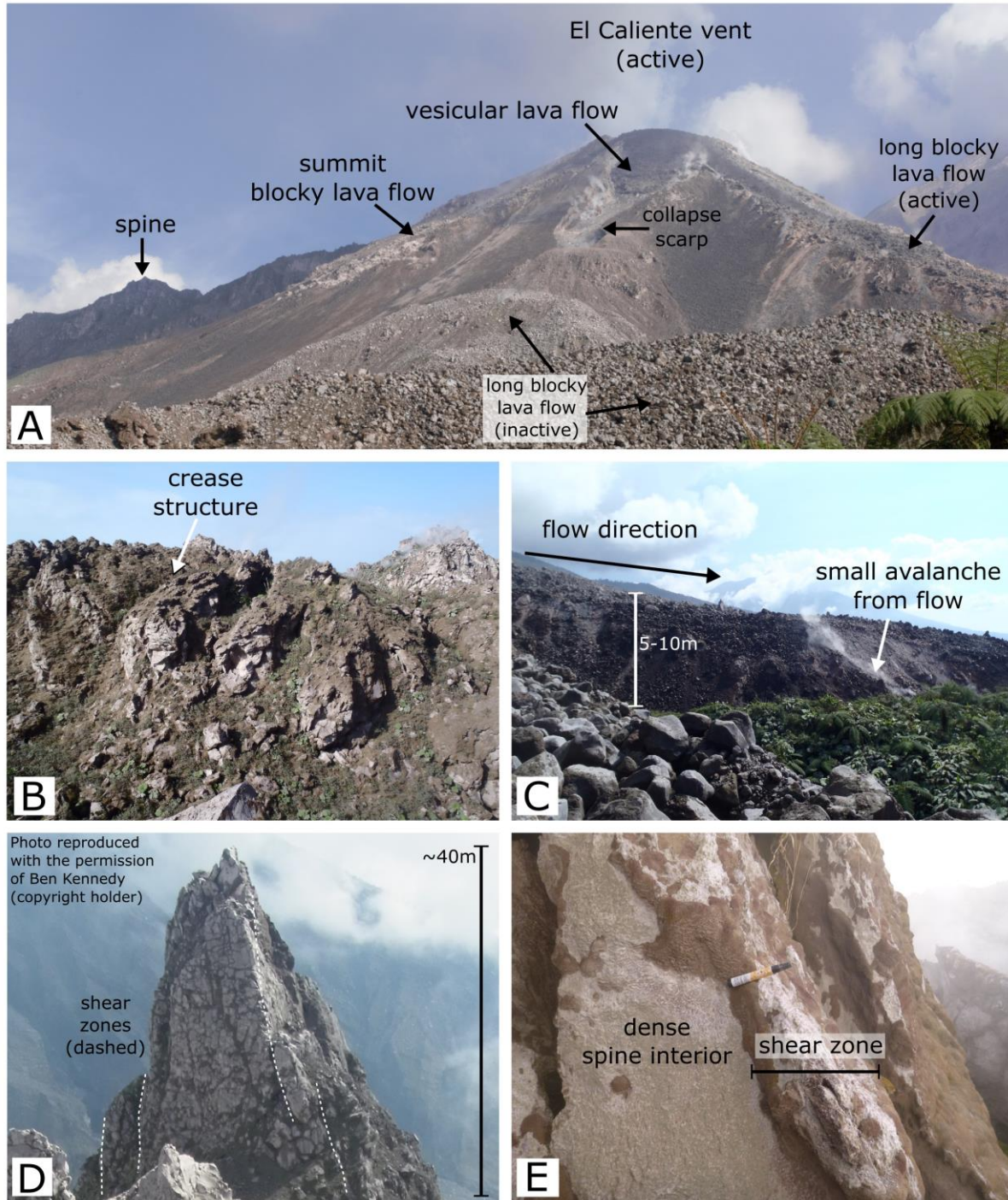
Figure 1. (A) Photo of Santiaguito and Santa Maria from the south, showing a typical eruption from El Caliente. The four different domes are labelled. Quetzaltenango is Guatemala’s second largest city and is located 10 km north. Photo from Jessica Ball. (B) Photo from a similar location in 1922, the first year of extrusion at El Caliente Dome inside the Santa Maria explosion crater. Photo source unknown, 1922.

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800

801 **Figure 2.** Major structural features at Santa Maria and Santiaguito. Santa Maria lies to the south-east
802 of the Xela Caldera Margin (84–12 Ma; Duffield et al., 1993), and sits on the western margin of the
803 Zunil Fault Zone, both mapped by Bennati et al. (2011). The ~1000 m high erosional explosion crater
804 and inferred incipient caldera bounding margin (Andrews, 2014) formed during the 1902 eruption are
805 marked with a white dotted and dashed line, respectively. Lineations, defined by the alignment of
806 topographic and or geomorphic features, possibly related to faults or joints, are displayed in yellow
807 and were mapped by Escobar-Wolf et al. (2011). Santiaguito is shaded grey with the active El
808 Caliente crater marked in red. Aerial photograph sourced from Google Map data ©2018.



809

810 **Figure 3.** Photos of lava types. (A) El Caliente from the south, November 2012, where a vesicular
 811 lava is infilling a recent collapse scarp, and long blocky lava flow and summit blocky lava flow are
 812 also active. Spines from El Monje can be seen in the distance. (B) A vesicular lava flow on El Brujo,
 813 this flow is semi-circle shaped and exhibits a crevasse with features similar to crease structures
 814 described at Mount St. Helens by Anderson and Fink (1992). (C) A long blocky lava flow
 815 approximately 1km from the vent. (D) A prominent spine on La Mitad. (E) Close up of a spine
 816 showing the dense spine interior and shear zone rim.

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Extrusive Lava-types

- Vesicular Lava Flow
- Blocky Lava Flow >2.5km
- Blocky Lava Flow <2.5km
- Blocky Lava Flow - Summit
- Spine

Other Units

- Talus
- Summit Rubble: large broken up blocks from collapsing spines and lava flows
- Tephra (~1972-2015)
- Playas: low areas on the dome infilled with alluvium and ash
- Alluvial and laharc deposits

Boundaries

- Unit boundary: including associated summit rubble from broken spines and lava flows
- In-situ outcrop

Structural Features

- Active Crater Rim
- Crater Rim
- Major Arcuate Feature: semi-circular ridge that surrounds part of a vent
- Collapse Scarp

Unit names are labelled following Rose et al. (1972) and Escobar-Wolf et al. (2010). "R" stands for Recent, and is followed by a prefix which denotes the dome. "c" = El Caliente, "e" = La Mitad, "m" = El Monje, and "b" = El Brujo. The third letter represents the age of the unit. Note some older units have been eroded or buried by subsequent units. The number at the end denotes the subunits, if any.

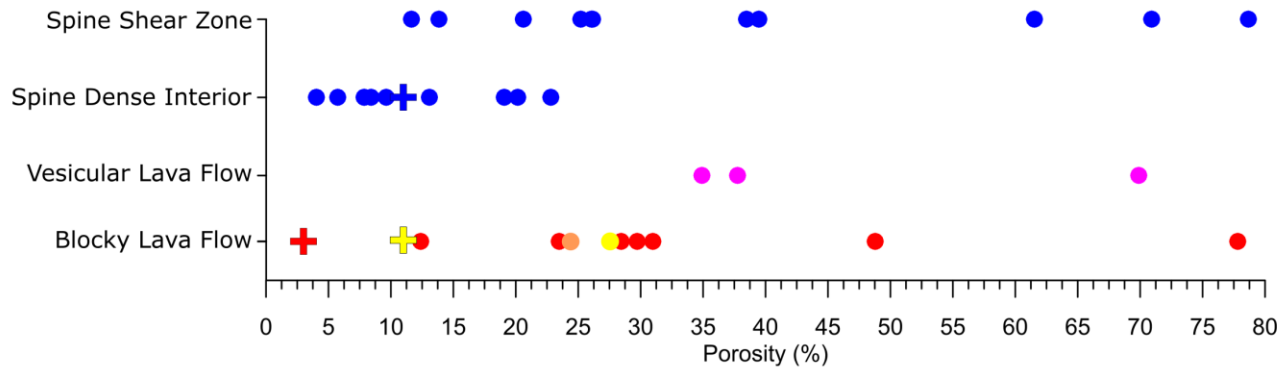
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819 **Figure 4.** Lava type map of the Santiaguito Lava Dome Complex. This map is developed as a
820 companion map to Escobar-Wolf et al. (2010) and Ball et al. (2013). Pre-2010 dome and lava flow

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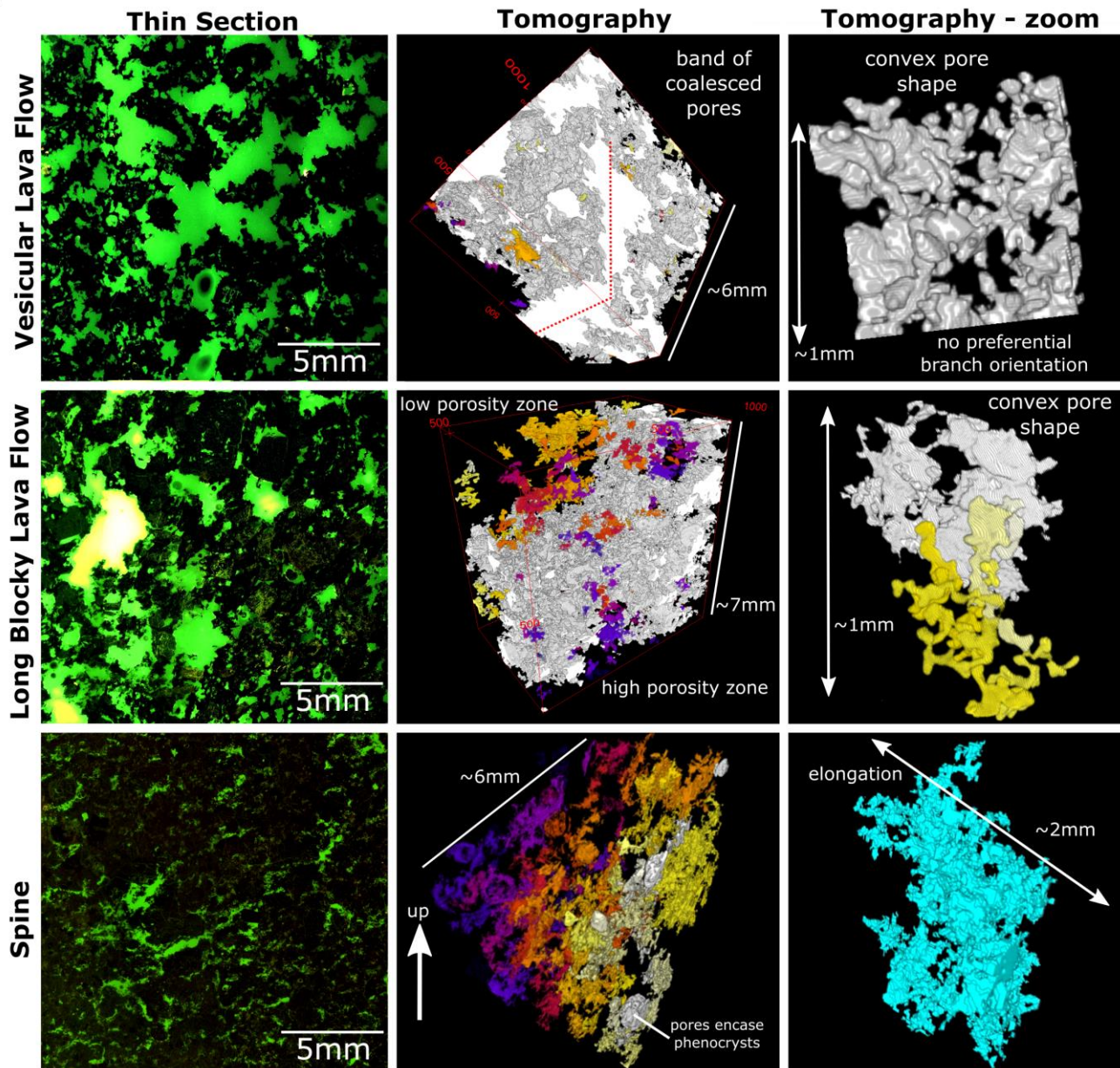
821 features were mapped by Escobar-Wolf et al. (2010) and Ball et al. (2013) from aerial 0.5 m pixel
822 resolution aerial orthophotos 1860-II-14 and 1860-II-19 (acquired between November 2005 and
823 April 2006 by the by the Instituto Geográfico Nacional) as well as by direct field mapping. Recent
824 post-2011 collapses and lava flows were mapped from NASA ASTER images, field observations,
825 INSIVUMEH photos, and Google Earth.

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827

828 **Figure 5.** Total porosity of extrusive products at Santiaguito. Circles are data from this study, and
829 crosses are values from Avard and Whittington (2012). Blocky lava flows are separated into: summit
830 blocky flows (yellow), blocky lava flows that reach the base of the cone but do not extend further
831 than 2.5 km length (orange), and blocky lava flows >2.5 km (red). Less than 1% of the porosity was
832 isolated in all the samples used in this study, so only total porosity is shown, equivalent to the
833 connected porosity.



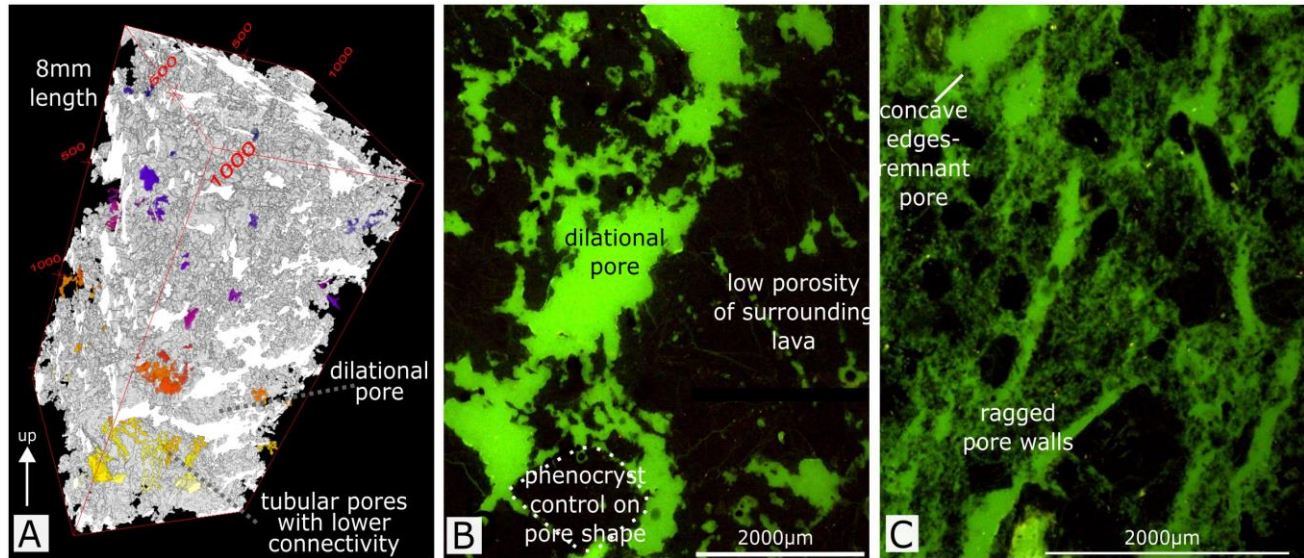
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835 **Figure 6.** Ultra-violet pictomicrographs and tomography imaging of pores in vesicular and blocky
 836 lava flows, and of the dense interior of spines. Different colours in tomography images represent
 837 connected pores, and the 100 largest pores are displayed in the images.

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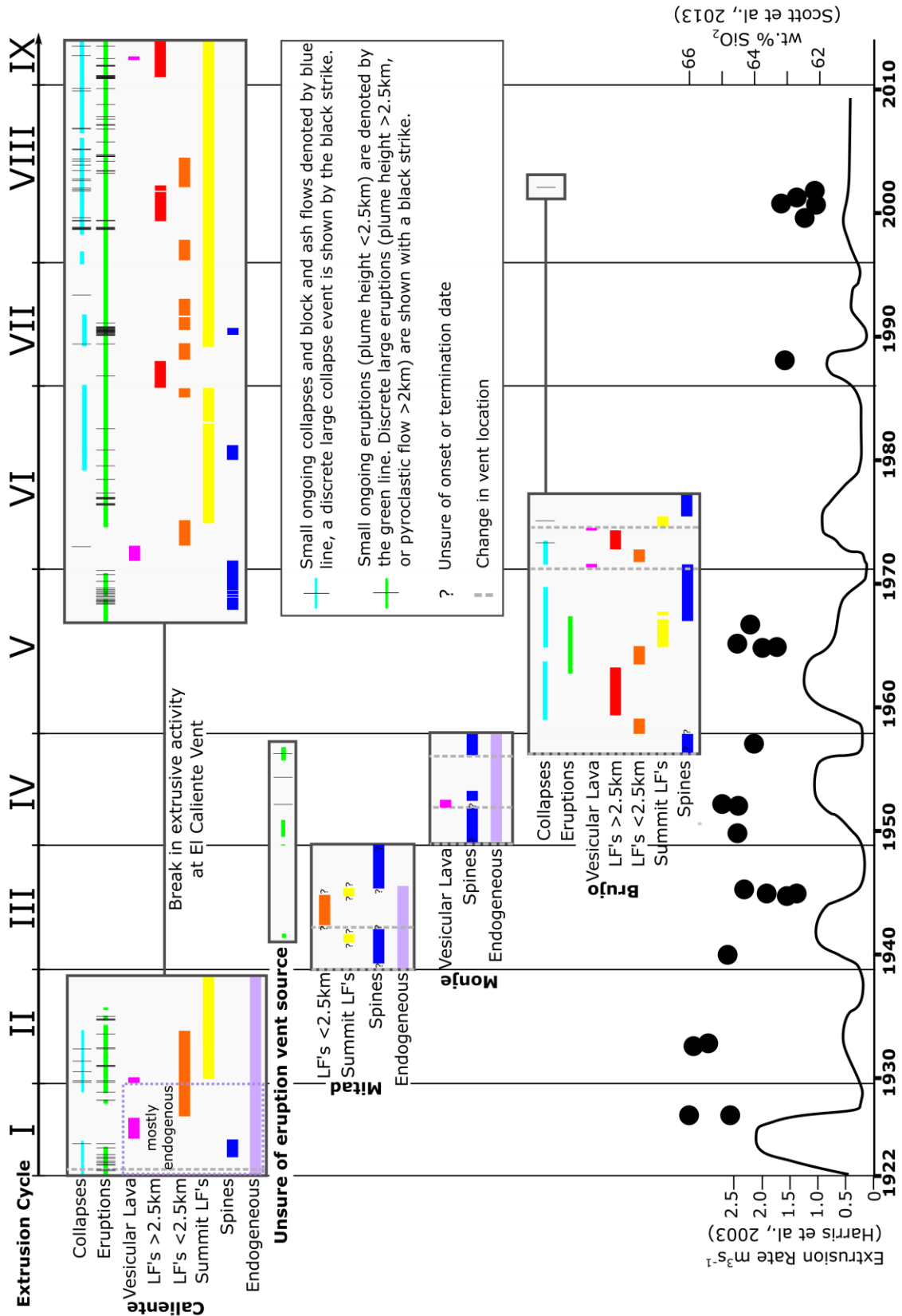
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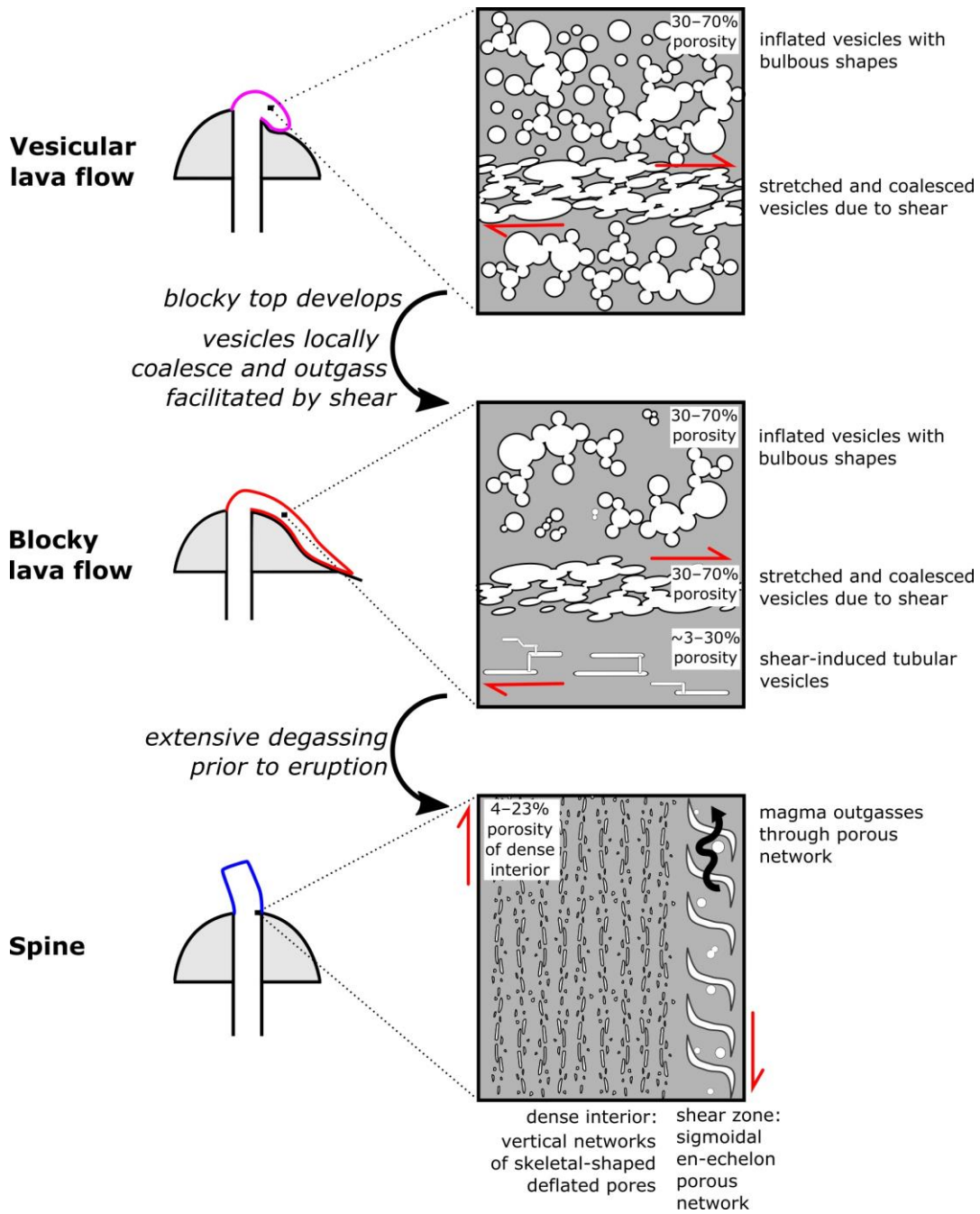
842 **Figure 7.** (A) Tomography of pores in a vesicular zone bordering a spine. The 100 largest pores are
 843 displayed in the image in different colours. Note the high connectivity of dilatational pores (white) and
 844 zone of lower porosity with lower connectivity (bottom corner). (B) UV photo of flow band in a
 845 spine showing interconnected dilatational pores and phenocryst control on pore shape. Most pore walls
 846 are concave and convex in this example. (C) UV photo of a shear zone with ragged pore walls
 847 marked by the protrusion of groundmass fragments and phenocrysts. Remnant concave pore walls of
 848 previous pores are also visible. Note subvertical alignment of pores.



849

850 **Figure 8.** Timing of extrusive activity and eruptive and collapse events at Santiaguito with reference
 851 to extrusion cycles (Harris et al., 2003), composition (Scott et al., 2013), and extrusion rate (Rose,
 852 1973b; Harris et al., 2003).

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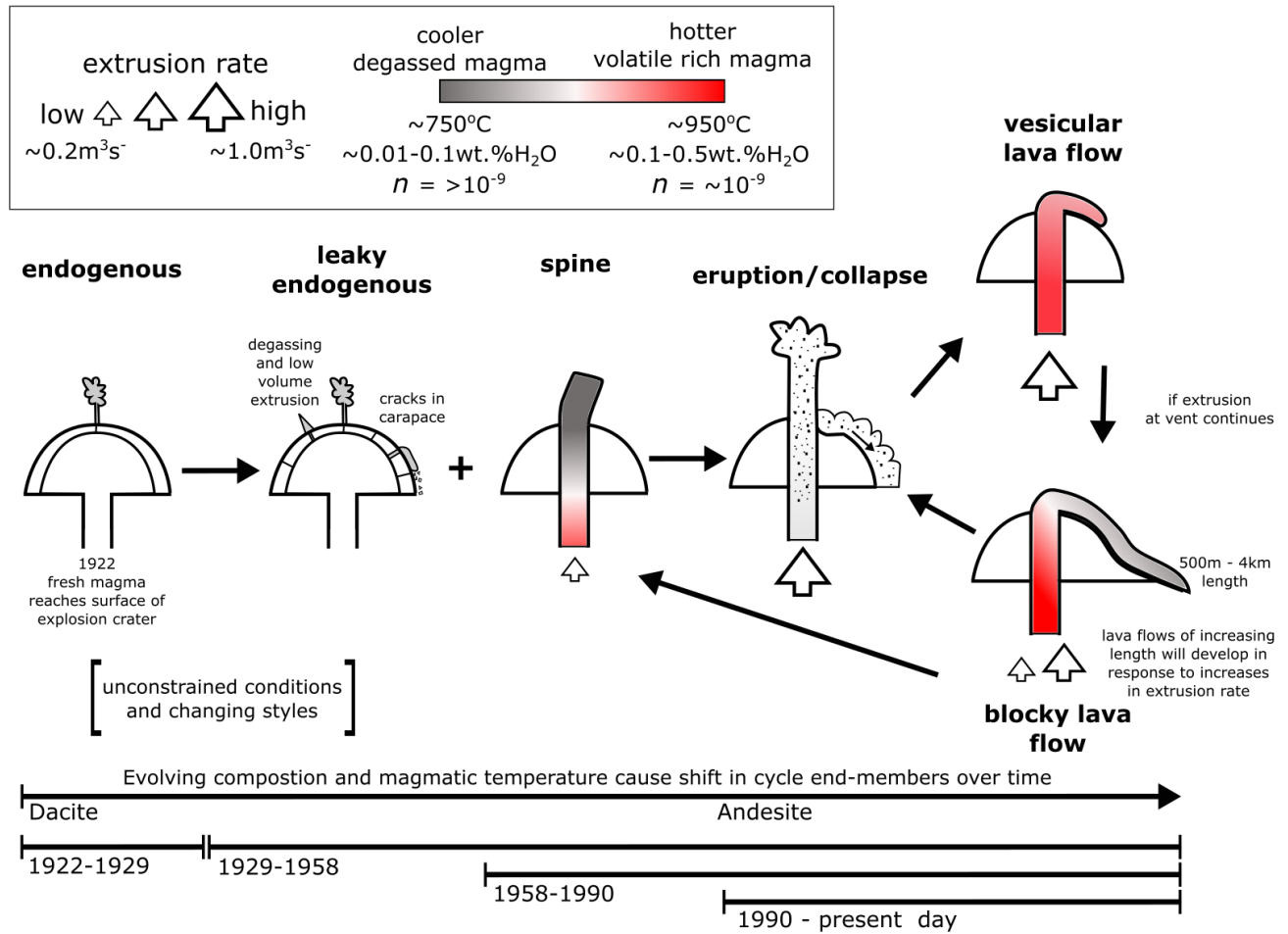


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854 **Figure 9.** Conceptual model of pore development at Santiaguito

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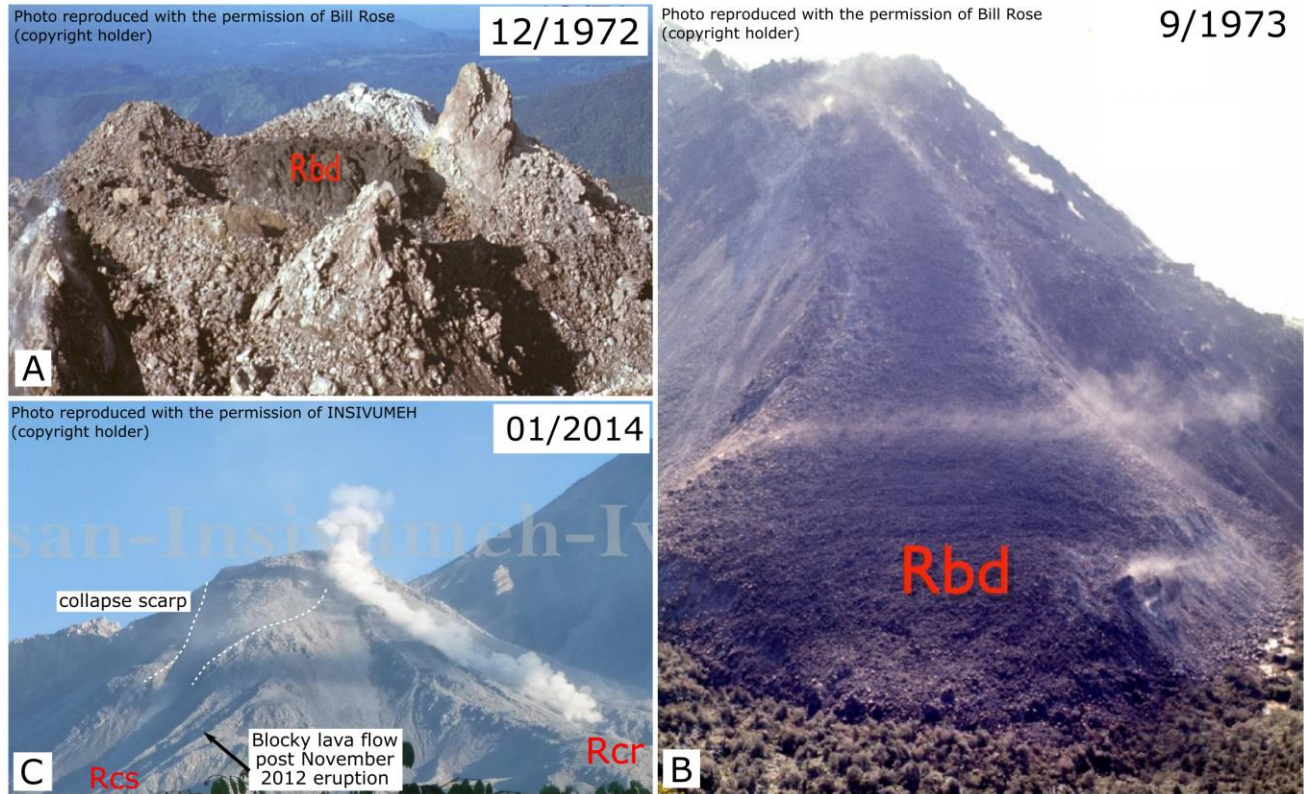


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858 **Figure 10.** Cyclical model of extrusive activity at Santiaguito.

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Textural insights into evolving lava dome cycles, Santiaguito Lava Dome, Guatemala



860

861 **Figure 11.** Evolution of vesicular lava into blocky lava. (A) Vesicular lava extruding on El Brujo,
862 surrounded by older spines. (B) Photo of the same flow front, 10 months later, with increased
863 brecciation of the flow top. (C) Proximal lava infills a collapse scarp; the lava flow top is made up of
864 large blocks in comparison to longer flows (as in B). See also the blocky lava flow (Rcs) that
865 advanced in the months after the November 2012 collapse, following the initial vesicular lava flow.
866 Note the small block-and-ash flow descending the dome.

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873 **7 Conflict of Interest**

874 The authors declare that the research was conducted in the absence of any commercial or financial
875 relationships that could be construed as a potential conflict of interest.

876 **8 Author Contributions**

877 ER undertook field mapping and sample collection, compiled the timeline, and carried out textural
878 analysis. BK and ER conceptualized models and wrote/edited the final manuscript. YL assisted with
879 the field campaign provided constructive reviews on/ edited the final manuscript. AH assisted with
880 field work. ME assisted with tomography processing. GH was our scientific liaison in Guatemala and
881 provided access to the domes.

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886 E613) in New-Zealand) and the Medical Imaging beamline at the Australian Synchrotron, Victoria.

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895 (MASSIVE) (www.massive.org.au)

896