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Article

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Direct Observation of High Densities of Antisite Defects in $\text{Ag}_2\text{ZnSnSe}_4$

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Abstract

The limited efficiency of $\text{Cu}_2\text{ZnSn}(\text{SSe})_4$ (CZTSSe) solar cells has been widely attributed to band tailing due to high densities of Cu_{Zn} and Zn_{Cu} antisite defects. It has been proposed that partial replacement of Cu by Ag should reduce the antisite defect density, leading to reduced band tailing and increased cell efficiencies. This paper examines antisite defects in $\text{Ag}_2\text{ZnSnSe}_4$ (AZTSe) crystals grown at high temperatures from a stoichiometric mixture of elements by scanning transmission electron microscopy (STEM), X-ray diffraction (XRD) and photoluminescence (PL). The elemental distribution was examined directly by atomic resolution STEM energy-dispersive X-ray (EDX) mapping and simultaneous annular dark field (ADF) imaging. EDX mapping suggested complete intermixing of Ag and Zn on the Wyckoff 2c and 2d sites in the AZTSe kesterite unit cell, and around 14% substitution of Zn for Ag on the 2a site, while ADF images showed evidence for local nanometer-sized regions within the disordered matrix with partial ordering of Zn and Ag on the 2c and 2d sites. These observations show that the AZTSe had a high density of Ag_{Zn} and Zn_{Ag} antisite defects, in contrast with room temperature photoluminescence showing relatively narrow emission lines close to the band edge and thus minimal band tailing. The interpretation of these results, and their wider significance for understanding the role of antisite defects in CZTSSe and Ag-substituted CZTSSe cells, are discussed.

1. Introduction:

Thin film photovoltaic (PV) solar cells based on CdTe and CuInGaSe_2 (CIGS) account for nearly 5% of the solar cell market. However, there is increasing concern that these absorber materials

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3 contain elements which are either in short supply (particularly In, Te) or toxic (Cd). There is
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5 interest in developing replacement materials with similar structure and electronic properties,
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7 but composed of less toxic and more earth-abundant elements. One such material is
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9 $\text{Cu}_2\text{ZnSn}(\text{SSe})_4$ (CZTSSe), which has a band-gap in the range 1.0-1.5 eV. However, whereas CdTe
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11 and CIGS cells with similar band-gaps have efficiencies up to about 22%, the best CZTSSe cells
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13 are still below 13% efficiency.^[1] This has been widely ascribed to antisite defects, particularly
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15 Cu_{Zn} and Zn_{Cu} which are known to be present at high densities in the equilibrium kesterite
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17 structure from neutron and synchrotron X-ray diffraction, Raman and scanning transmission
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19 electron microscopy (STEM) studies.^[2-6] These defects, either in the form of Cu-Zn antisite pairs,
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21 or as single antisite defects which carry a net charge, have been considered to account for the
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23 “band tailing” apparent in optical data from CZTSSe.^[7,8] This, in turn, is theorized to account for
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25 the low open circuit voltage (V_{OC}), typically around 60% of the theoretical maximum, which is a
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27 principal reason for the limited cell efficiency.^[9] Cu-Zn antisite defects should form easily in
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29 CZTSSe as the estimated energy for exchange of an isolated pair of Cu-Zn ions is around 0.2 eV,
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31 partly due to the Cu^+ having a covalent radius around 5% larger than Zn^{2+} , leading to a relatively
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33 low strain energy contribution.^[10] Replacement of Cu^+ by Ag^+ , which has a covalent radius
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35 around 17% greater than Zn^{2+} , has been estimated to increase this ion exchange energy by a
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37 factor of about 4, such that antisite defect densities should be reduced by more than an order
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39 of magnitude.^[11] Experimentally, photoluminescence (PL) studies of Ag-substituted CZTSSe
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41 show a progressive reduction in band tailing as the Ag content increases, and for low levels of
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43 Ag substitution, modest increases in cell efficiencies have been noted.^[12] The implication, as yet
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3 unverified, is that Ag-substitution has reduced densities of antisite defects compared with
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5 CZTSSe, and that this is a key factor in improving cell efficiencies.
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9 In this paper we use STEM to directly examine Ag-Zn ordering in $\text{Ag}_2\text{SnZnSe}_4$ (AZTSe) for
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11 the first time. In previous STEM work, we have used atomic resolution high angle ADF (HAADF)
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13 imaging in the [010] orientation to examine cation ordering in CZTS.^[13] For a completely
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15 ordered kesterite structure viewed along [010] (Figure 1), the individual end-on atom columns
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17 in CZTS are composed of a single element, Sn, Cu, Zn or S. Since HAADF images show contrast
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19 which increases with the atomic number Z, [010] atomic resolution images distinguish the Sn
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21 ($Z=50$) columns as the sites of highest intensity from the Cu ($Z=29$), Zn ($Z=30$) or S ($Z=16$)
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23 columns. The images revealed high densities of antisite domain boundaries through relative
24
25 shifts of the Sn columns across edge-on boundaries (e.g. see figures 3,4 in ^[13]). Although the Cu
26
27 and Zn columns are essentially indistinguishable in HAADF owing to their close atomic numbers,
28
29 Mendis et al.^[5] used atomic resolution electron energy loss spectroscopy (EELS) in STEM to map
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31 the Cu and Zn distribution in [010]-oriented CZTS. They found some evidence of Cu-Zn
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33 intermixing on the Wyckoff 2c and 2d sites (defined in Figure 1), as well as Cu and Zn clustering
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35 on a 1-5nm scale. Aguiar et al.^[6] used atomic resolution EELS mapping in a [112] orientation to
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37 show further evidence for clustering of Cu and Zn in CZTS, albeit on a slightly broader (20nm)
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39 scale.
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48 Here, we combine HAADF imaging with atomic resolution elemental mapping by energy
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50 dispersive X-ray microanalysis (EDX) in the [010] orientation. The results, which enable us to
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52 map the distribution of all four elements, show that there are regions of near-complete
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54 intermixing of Ag and Zn on the Wyckoff 2c and 2d sites, along with less but significant
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intermixing of Ag and Zn on the Wyckoff 2a sites. Despite this evidence for high densities of Ag-Zn antisite defects, photoluminescence measurements show a relatively narrow emission close to the expected band edge. The interpretation of these results, and the wider implications for understanding the properties and the role of antisite defects in CZTSSe and Ag-substituted cells, are discussed.

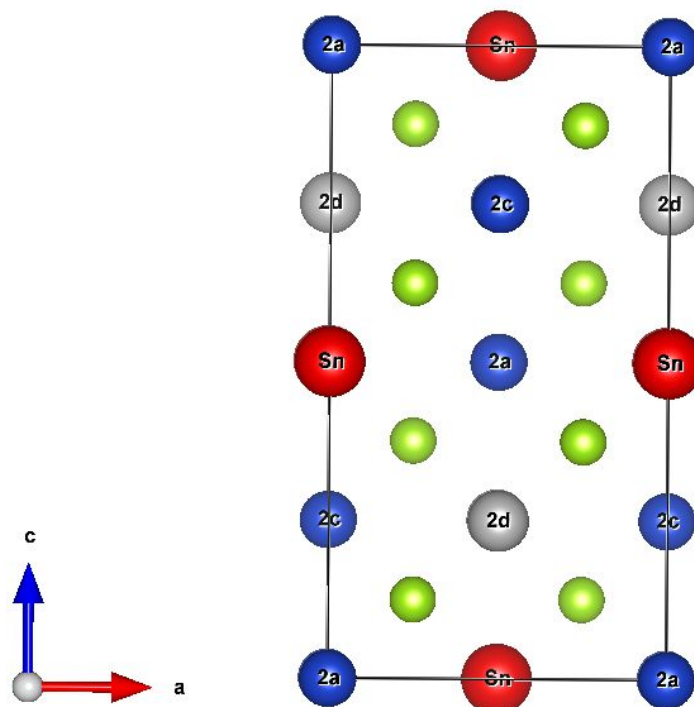


Figure 1: The unit cell for kesterite CZTSSe or AZTSSe viewed down the b-axis [010]. Red: Sn columns; green (unlabelled): S or Se columns. The remaining columns are Wyckoff sites 2a (Cu/Ag, blue), 2c (Cu/Ag, blue) and 2d (Zn, grey)

2. Experimental methods:

$\text{Ag}_2\text{SnZnSe}_4$ was grown at high temperature in an evacuated sealed ampoule from a stoichiometric mixture of the pure elements.^[14] The elemental mixture was sealed in a quartz

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3 ampoule at 10^{-6} torr and annealed for 20 days at 480°C. The material was then extracted and
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5 ground to a powder to homogenise it, re-sealed under vacuum in a second ampoule and
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7 annealed for a further 26 days at 510°C. The resulting material was then divided into two
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9 samples, one of which was quenched into ice-water, and the other slow-cooled to room
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11 temperature, at 3°C/hr, to maximise ordering.
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16 Both samples were powders with occasional crystals up to 1 mm or more across, along
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18 with micron-sized crystals. The general morphology and composition were studied in an Amray
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20 1810T Digital Scanning Electron Microscope (SEM) with Energy Dispersive Spectroscopy. Bulk
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22 structural analysis was performed via X-ray diffraction (XRD) using a Philips/Norelco scanning
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24 wide angle X-ray diffractometer with Bragg-Brentano geometry. XRD specimens were gently
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26 pulverized in an agate crucible with 125 mesh in accordance with ASTM standards. Lattice
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28 parameters were determined using JADE XRD analysis software (Materials Data, Inc. (MDI) JADE
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30 2010) for unit cell Rietveld refinements.
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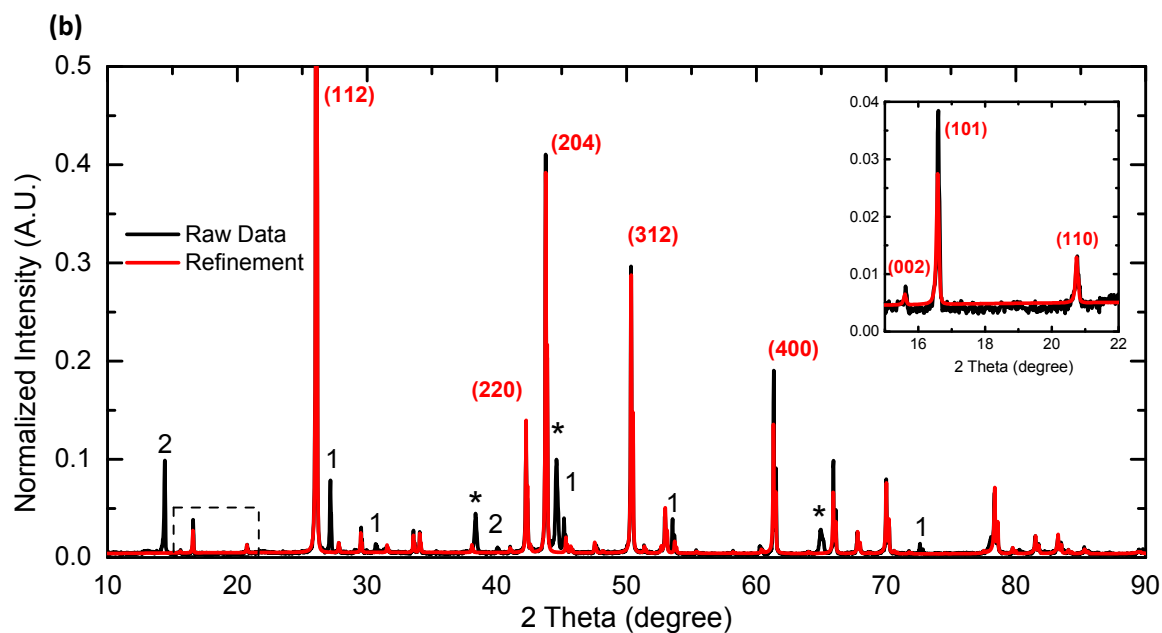
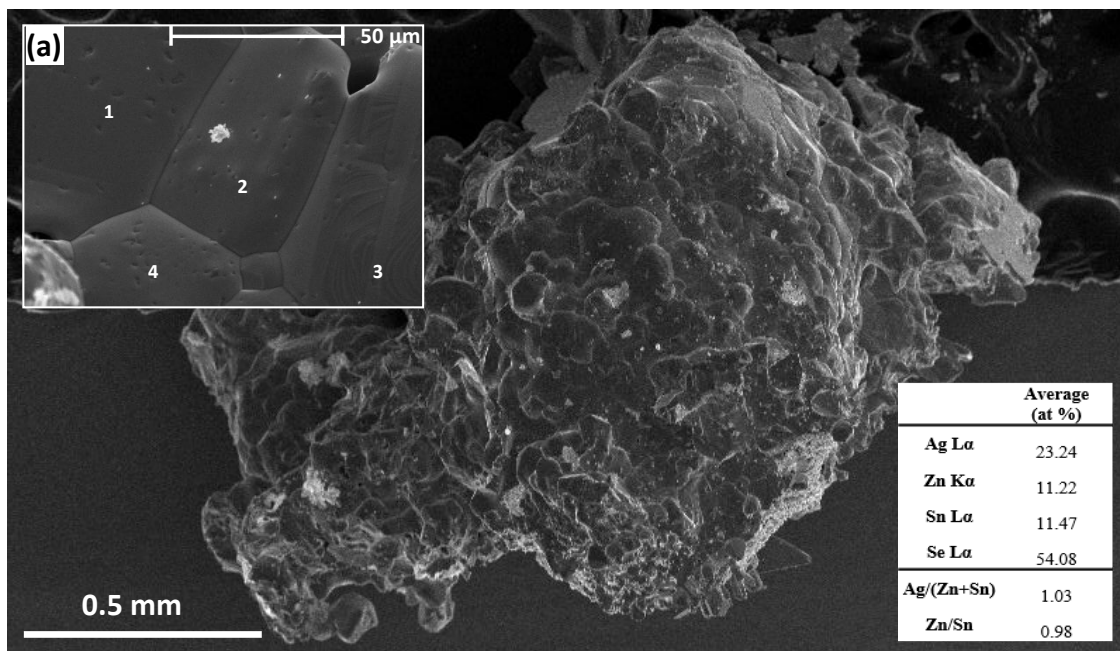
36 For STEM, samples were ground into a fine powder and dispersed in pure ethanol onto
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38 holey carbon films. The films were then heated to 90°C under vacuum for 12hr to remove
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40 hydrocarbons and examined in a JEOL ARM200CF (S)TEM operating with a cold field emission
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42 source and aberration correction on the probe, giving a probe size less than 0.1nm. HAADF
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44 imaging and simultaneous EDX microanalysis using a single Centurio 100mm² detector were
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46 carried out in scanning (STEM) mode. HAADF images shown in this paper have inner and outer
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48 cut-off angles of 73 mrad and 236 mrad respectively.
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3 Photoluminescence was carried out on the as-grown powders at room temperature
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5 using a 532 nm pulsed laser operating at 15 kHz with a photomultiplier tube and InGaAs
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7 detector. A neutral density filter wheel was used to change incident laser intensity.
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10 11 **3. Results and discussion:**

12 13 14 **3.1 SEM and XRD**

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17 Figure 2 shows SEM and XRD results from the slow-cooled sample. Figure 2a illustrates the as-
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19 grown material and a magnified view of a group of individual AZTSe grains, with near
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21 stoichiometric composition as confirmed by EDX spot measurements. In Figure 2b, the
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23 experimental XRD pattern is compared with a simulation in red for the fully ordered $\text{Ag}_2\text{SnZnSe}_4$
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25 kesterite structure (Figure 1). The main peaks in the XRD are consistent with the kesterite
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27 phase, with lattice constants of $a = 6.03 \pm 0.01 \text{ \AA}$ and a c/a value of 1.88 ± 0.01 similar to
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29 previous reports.^[15] Evidence of secondary phases, including SnSe_2 and ZnSe can also be
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31 detected in the powder due to the large XRD sampling volume. The major $\text{Ag}_2\text{SnZnSe}_4$ peaks,
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33 such as $hkl = 112, 220, 312$ etc, represent the underlying near-cubic face-centred lattice
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35 obtained if the ordering of the cations, Sn, Ag and Zn, is ignored (i.e. corresponding to a unit cell
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37 half that in Figure 1, with dimensions $a, a, c/2$). Minor $\text{Ag}_2\text{SnZnSe}_4$ peaks such as 002, 101 and
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39 110, with intensities typically a few percent of the major peak intensities, are also present due
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41 to cation ordering. It is worth noting that the 101 reflection is sensitive to Ag-Zn ordering on the
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43 Wyckoff 2c and 2d sites, and is kinematically forbidden in a model where Ag and Zn are
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45 completely intermixed on these sites. The 101 reflection was present for the quenched as well
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47 as the slow-cooled sample but with an intensity about 30% lower.
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46 Figure 2: (a) SEM micrograph of synthesized AZTSe powder. Insert: Magnified AZTSe grains with
47 numbered regions of interest. Table: Average EDX composition of regions 1-4, (b) XRD pattern
48 for the slow-cooled sample (black) and fit for ordered AZTSe model (red).
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50 3.2 STEM

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3 While XRD is relatively insensitive to cation ordering in this system and provides only a
4 view of aggregate structure across thousands of crystals/ grains, STEM allows us to select single
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6 crystal particles of $\text{Ag}_2\text{SnZnSe}_4$ and tilt them to “structure sensitive” orientations such as [010]
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8 where cation ordering can be examined in more detail.
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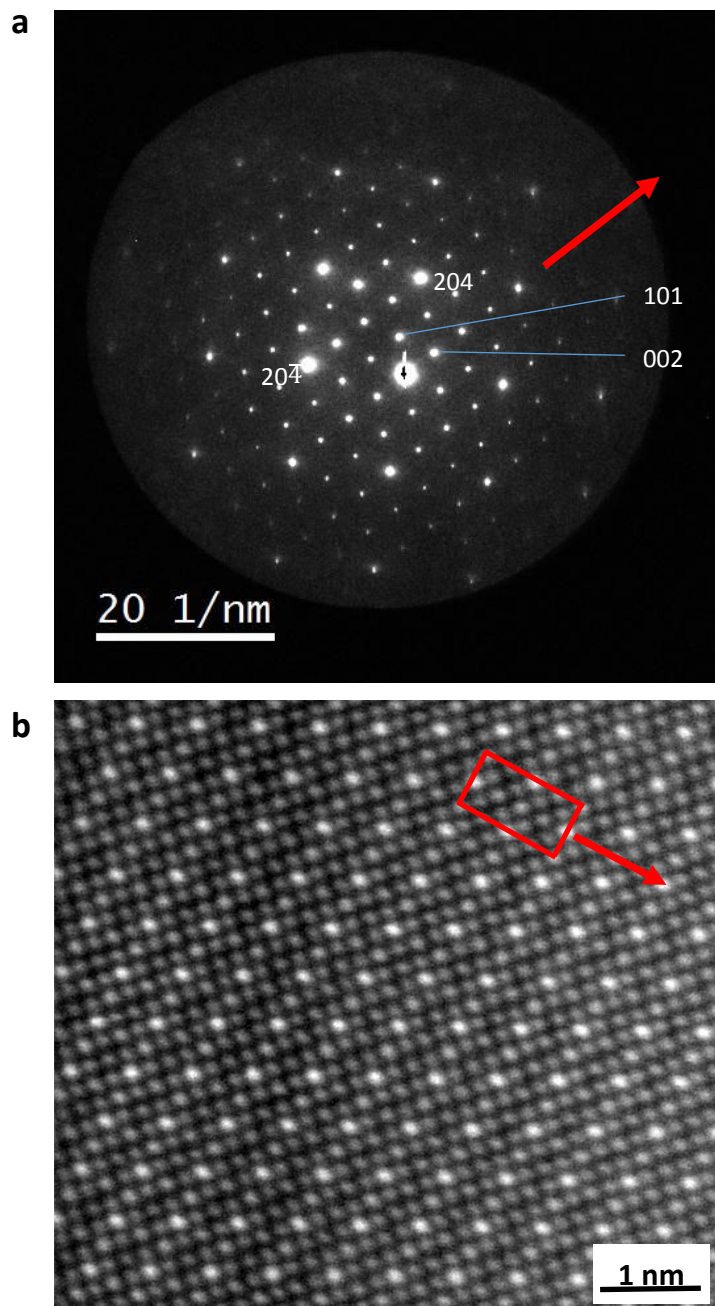


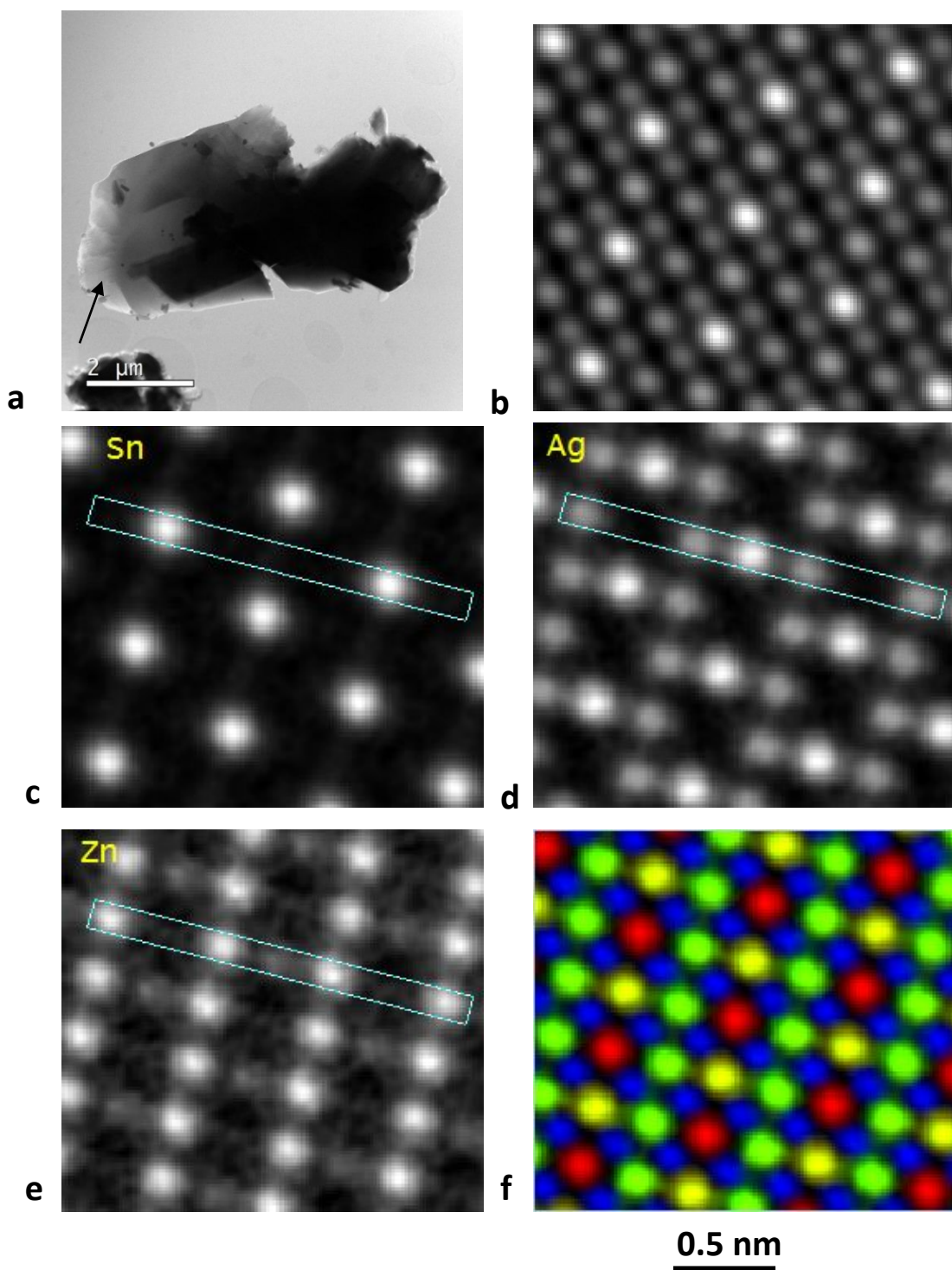
Figure 3: (a) selected area diffraction pattern and (b) HAADF image of AZTSe (unit cell outlined) viewed down the b-axis. The c-axis is arrowed in (a) and (b). As for CZTS in Figure 1(b), the brightest spots in the HAADF image, which form a near-hexagonal array

The STEM samples contained well separated micrometer-sized particles. Selection and tilting of particles to the [010] orientation was carried out in direct imaging (TEM) mode, with a near-parallel incident beam. Figure 3a shows a typical electron diffraction pattern taken from

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3 one such particle in the [010] orientation. This illustrates that, in contrast to XRD, the intensities
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5 of the cation ordering reflections 002, 101 etc are relatively increased, although not directly
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7 interpretable owing to dynamical scattering, compared with the major reflections 204 and 20-4
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9 (Figure 2b). The microscope was then changed to scanning (STEM) mode, and HAADF and EDX
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11 mapping were carried out. Figure 3b shows a HAADF image of the slow-cooled AZTSe taken
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13 with [010] parallel to the beam. The atomic numbers for Sn (Z=50) and Ag (Z=47) are relatively
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15 close, and therefore the intensities on Sn and Ag (Wyckoff 2a and 2c) sites should be similar,
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17 with much lower intensity for Zn (Z=30, Wyckoff 2d) sites. However, the HAADF image shows a
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19 near-hexagonal array of bright spots, which we identify later as the Sn columns, and lower, and
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21 relatively similar intensities for other cation columns (Wyckoff 2a, 2c and 2d sites)
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25 corresponding to mixtures of Ag and Zn.
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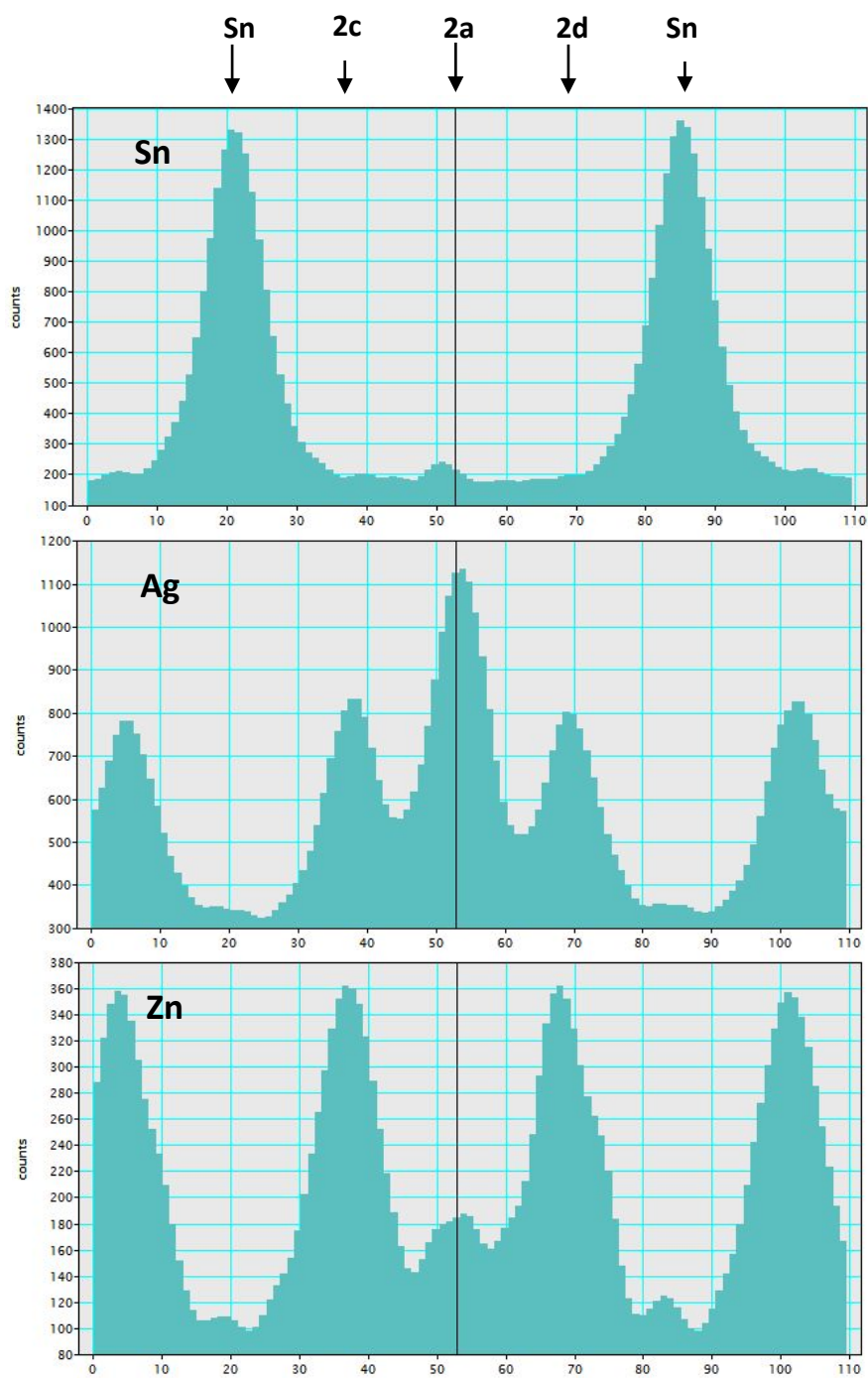
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31 In order to unambiguously identify the elemental distribution, we have correlated
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33 HAADF images with simultaneous EDX mapping, using methods of data superposition described
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35 by Jones et al.^[16,17] The final results are shown in Figure 4 and more detail is shown in the
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37 Supplementary data. In summary, HAADF images and EDX spectra were recorded by scanning a
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39 0.1nm probe, with 17pA current, over an 11nm x 11nm area of single crystal. Owing to low X-
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41 ray count rates, repeated scans were carried out for the same or a closely adjacent area to
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43 minimise radiation damage effects, and the HAADF (along with the EDX) maps aligned by non-
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45 rigid registration, correcting for the effects of specimen drift and scan distortions.^[17] In this
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47 procedure, the bright spots in the HAADF images of AZTSe (Sn columns) act as a reliable
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49 reference lattice to bring successive frames into coincidence. To obtain the results shown in
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Figure 4, a tile was then defined in the extended scan, and



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3 Figure 4: Simultaneous HAADF and EDX mapping, (a) area selected (arrowed), (b) HAADF, (c-e)
4 EDX maps for Sn L, Ag L and Zn L X-rays following data superposition, (f) colour montage: red -
5 Sn; yellow - Ag in 2a site; green - Zn and Ag in 2c/2d sites; blue - Se K
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7
8 template matching was used to identify 111 other unique tiles. Summing over these further
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10 improves the signal to noise ratio, giving an effective electron dose corresponding to the
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12 HAADF and EDX maps in Figure 3 of $8.9 \times 10^{11} \text{ e}^- \text{ nm}^{-2}$.
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50 Figure 5: Intensity line scans parallel to the c-axis for the areas outlined in Figure 4. The
51 line scans pass through the 2a, 2c, 2d and Sn sites as indicated

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53 The results in Figure 4 show the average distributions of Sn, Ag and Zn from an area
54 some tens of nanometers across and estimated to be 20-30nm thick from a t/λ fit to EELS
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3 spectra.^[18] Linescans in the [001] direction across the EDX maps in Figure 4 are shown in Figure
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5 5, and compare the distributions of Sn, Ag and Zn across all three cation sites. The higher
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7 concentration of Ag on the 2a site compared with the 2c and 2d sites confirms that the basic
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9 structure is kesterite rather than stannite, which is a slightly higher energy structure with Zn on
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11 the 2a site.^[19] We note that the distinction between sites 2c and 2d is arbitrary, as, within the
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13 experimental error, neither the HAADF nor the EDX maps show a clear difference between
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15 these sites. Despite the superimposition of many images, Fourier analysis gave the real-space
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17 resolution in the HAADF images as 0.094nm, and the resolution in the EDX maps in the range
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19 0.16-0.20 nm. In the thin crystal limit, EDX spatial resolution is expected to be not much larger
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21 than the probe size as X-ray generation depends on a localised interaction of the incident
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23 electron beam with tightly bound inner electron shells. As the focussed probe is scanned across
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25 a given atom column, electron dechannelling (beam spreading) due to thermal diffuse
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27 scattering should contribute to some broadening of the X-ray peaks and also cause some X-rays
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29 to be generated on neighbouring atom columns. Calculations^[20] suggest that is likely to be a
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31 small contribution at the film thicknesses here (20-30nm), but this will depend on the Debye-
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33 Waller factors for AZTSe, which are not known at present, and might include a significant static
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35 component as Ag and Zn intermixing should cause ions to be displaced from lattice sites owing
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37 to the disparity in the Ag and Zn ionic radii.^[21] Taking the Sn map and line profiles in Figures 5,
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39 there is a good correspondence between the bright spots in the HAADF image peaks and the
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41 main peaks in the Sn L map, confirming our earlier assumption that the bright spots in HAADF
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43 maps represent Sn columns. There is also a small Sn peak on the 2a column (less than 4% of the
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45 main peaks), and similar small Ag and Zn peaks on the Sn sites. These subsidiary peaks will
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3 include the dechannelling contribution, confirming that the dechannelling contribution is small.
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6 However, whether these subsidiary peaks are wholly due to dechannelling or include a
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8 proportion of elemental intermixing remains uncertain.
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11 The EDX maps confirm that Ag and Zn are both present in the 2c and 2d sites in about
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13 equal amounts within random error limits (i.e. $N \mp \sqrt{N}$ where N is the number of X-ray counts)
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15 and that there is also a significant Zn peak in the 2a site. Table 1 shows the atomic fractions of
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17 Ag and Zn in all three sites, calculated separately from the Ag and Zn profiles. In calculating
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19 these fractions, it has been assumed that in total, over all three sites, the Ag and Zn were
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21 present in stoichiometric amounts (i.e. two Ag and one Zn atom per unit cell) and that the
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23 subsidiary Ag and Zn peaks on the Sn sites measure the dechannelling contribution from
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25 neighbouring 2a, 2c and 2d sites only, i.e. no Ag or Zn in the Sn columns. The dechannelling
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27 contribution to each of the 2a, 2c and 2d peaks will depend on the location of neighbouring
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29 columns with the same element and the extent of dechannelling. As the latter is not known, we
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31 have assumed, for simplicity, that the dechannelling peak on the Sn site represents a fixed
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33 contribution to be subtracted from all peaks.
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41 **Table 1:** Zn and Ag atomic fractions estimated from the respective Zn and Ag maps
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	2a site	2c site	2d site
Zn fraction from Zn map	0.14 ± 0.03	0.43 ± 0.02	0.43 ± 0.02
Ag fraction from Ag map	0.89 ± 0.02	0.57 ± 0.02	0.55 ± 0.02

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3 In interpreting the data in Table 1, it is necessary to consider whether the Ag and Zn are
4 randomly distributed in the 2a, 2c and 2d columns or whether there is local ordering. To
5 investigate this, we have examined individual HAADF images. In general, the HAADF intensity is
6 known to be relatively simply related to Z, varying as approximately $Z^{1.7}$ over a wide range of
7 imaging conditions.^[22,23] However, static atomic displacements from idealized kesterite
8 coordinates can affect quantitative analysis of column compositions, and may be expected in
9 our case.^[21] Nevertheless, HAADF images taken from different areas of both the slow-cooled
10 and quenched samples showed clear evidence for localised regions with partial ordering, as
11 indicated by an asymmetry in the intensity between 2c and 2d sites. This is evident from Figure
12 6, which illustrates a HAADF image from the quenched sample, along with a linescan taken
13 along the [100] direction through the 2c and 2d sites. This shows two localised regions A and B
14 with evidence of ordering on the alternating 2c/2d sites, i.e. every second peak is more intense,
15 indicating relatively more Ag. Between A and B there is a relative reversal in the ordering.
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35 In order to reconcile the apparently different results in Figures 5 and 6, one possibility is
36 that the AZTSe has neighbouring regions of ordered kesterite with a relative rotation of 180°
37 about [010] which bring the Sn columns back into coincidence. Structurally this corresponds to
38 a twin relationship between domains. For a single end-on column containing both domains, this
39 would superimpose the 2c and 2d sites, giving apparently equal Zn and Ag occupancy if the
40 domains were present in equal amounts. However, this model would not account for the
41 presence of Zn in the 2a columns since these columns would be unaffected by the
42 transformation. In addition, if twins were present, we should also expect them to be limited
43 laterally along the [100] direction, which is crystallographically equivalent to [010]. This would
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3 lead to local reversals in the 2c/2d asymmetry of the type seen in Figure 6. In this case, the
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5 transition between neighbouring domains takes place over 1-2nm. Whether this reflects the
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7 width of the boundary region, or the boundary is atomically sharp but inclined is, however,
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9 unclear. Indeed, we found no clear evidence in the HAADF images from either the quenched
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11 and slow-cooled samples for abrupt twin boundaries. This suggests that the transition region
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13 between ordered domains has a finite width, but further studies are needed to confirm this
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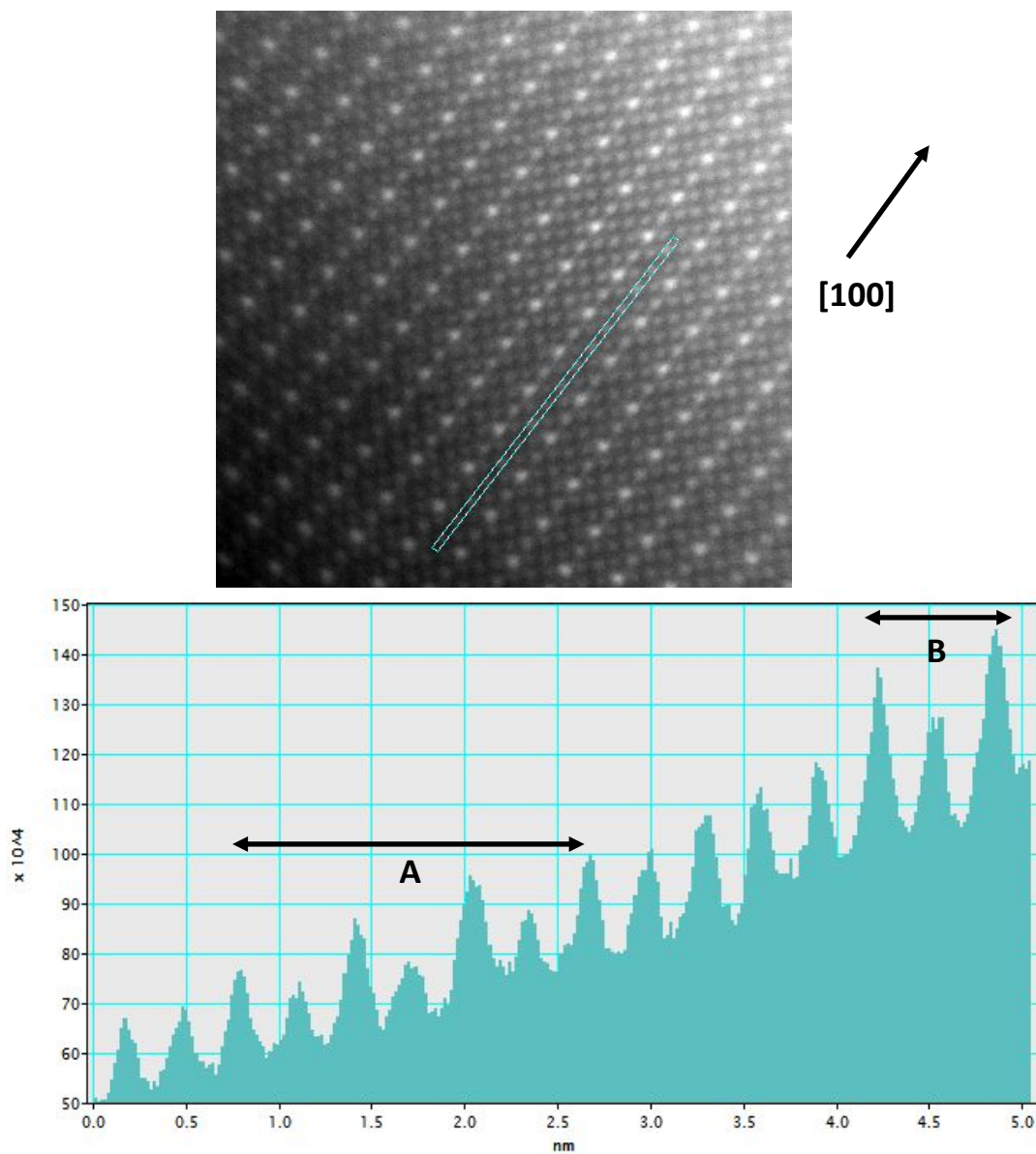


Figure 6: HAADF image from the quenched sample, and intensity linescan in the $[100]$ direction along the 2c-2d row (Figure 1) for the region indicated. A and B show regions of local asymmetry in intensity between the 2c and 2d sites

The HAADF results therefore suggest a model where there is local (nanoscale) Zn-Ag ordering on the 2c and 2d sites and a twin relationship between ordered regions. This is therefore consistent with the observation of reflections such as 101 in XRD which are

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2
3 kinematically forbidden for perfect Ag-Zn disorder on the 2c and 2d sites but have finite
4
5 structure factors otherwise.
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9 It is possible that regions of local ordering could be linked to local changes in stoichiometry.
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11 This is expected in CZTS, where theoretical studies suggest that reduced Cu content (Cu-poor
12
13 material) should reduce the density of Cu_{Zn} antisite defects and Cu-Zn disorder.^[10] However,
14
15 large area EDX mapping here showed no clear evidence for spatial variations in the Ag/Zn ratio.
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18 19 **3.2 Photoluminescence** 20

21
22 Figure 7 shows PL spectra obtained for the slow-cooled and quenched samples. Figure 7a
23
24 shows PL from the slow-cooled sample as a function of laser power, indicating a main peak at
25
26 1.345eV and a subsidiary peak at 1.27eV. The 1.345eV peak is relatively narrow and similar to
27
28 the main emission lines recorded for AZTSe thin films (in the range 1.33-1.36eV), which have
29
30 been interpreted as the band-edge emission.^[11] The PL intensity from the quenched sample
31
32 was an order of magnitude lower compared to the slow-cooled sample (Figure 7b), suggesting
33
34 higher rates of non-radiative recombination. In addition, although the 1.345eV peak is clearly
35
36 still present, the 1.27eV peak is either absent or has much reduced intensity relative to the
37
38 dominant emission at 1.345 eV in the quenched sample.
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45 Although attribution of the lower 1.27 eV peak is inconclusive, the most likely
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47 candidates appear to be either a SnSe secondary phase, which has a direct gap energy of
48
49 1.27eV,^[24] or emission related to Ag-Zn disorder. Evidence of SnSe however, was not found in
50
51 the XRD or TEM. The expected energies of other secondary phases are also not near the
52
53 observed 1.27 eV peak ($\text{Ag}_2\text{Se} = 0.15$ eV, $\text{ZnSe} = 2.7$ eV, $\text{Ag}_2\text{SnSe}_3 = 0.81$ eV, $\text{SnSe}_2 = 1.2 - 3.0$ eV).
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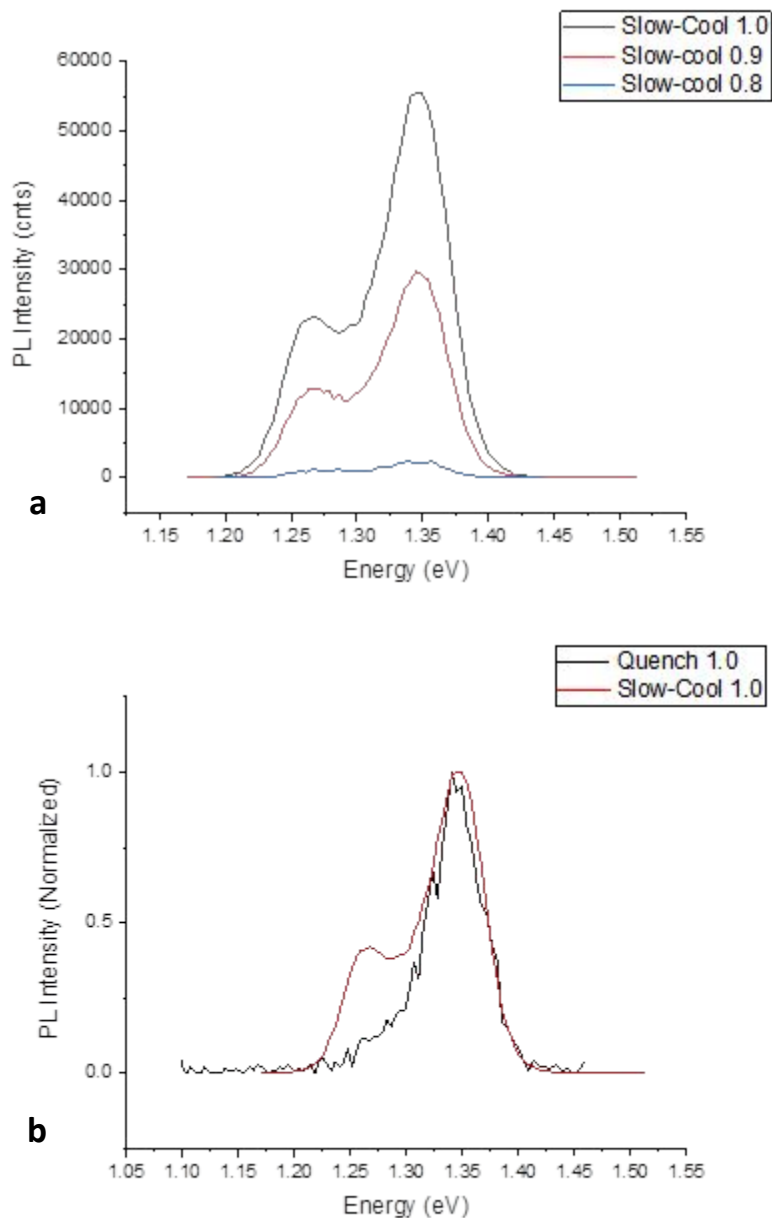


Figure 7: Room temperature PL (a) as a function of the laser power for the slow-cooled sample, (b) comparison of the quenched and slow-cooled samples at the same laser power

Considering intrinsic defects, a number of points in the literature suggest Ag-Zn related defect complex may be responsible. First, for all stable stoichiometries, the $[\text{Ag}_{\text{Zn}}+\text{Zn}_{\text{Ag}}]$ defect complex is predicted to have the lowest formation energy.^[25] Second, the shallow defect energy level measured by PL (0.07 eV from band edge) is the approximate energy predicted for

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2
3 the Ag_{Zn} defect complex. Third, sample stoichiometry is expected to be near stoichiometric due
4
5 to element weights during fabrication, which should limit formation of significant quantities of
6
7 stoichiometry-altering substitutions (i.e. $\text{V}_{\text{Ag}+\text{Zn}_{\text{Ag}}}$). It is worth noting that the theoretical
8
9 predictions^[25] have been made for the sulphide system ($\text{Ag}_2\text{ZnSnS}_4$), and do not exist for the
10
11 selenide, but significant deviations from the sulphide case are not expected, and are unlikely to
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13 change the points noted above.
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18 It is interesting to note that the lower energy peak is less resolved in the quenched
19
20 sample where more disorder is expected. This could be due to increased non-radiative
21
22 recombination in the Ag-Zn anti-site domains for the quenched sample, or changes in the
23
24 extent of bandgap fluctuations, which can be larger for slow-cooled samples due to the time
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26 allowed for domains to reach minimum energy state. Clearly further work is needed to
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28 investigate this point.
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33 34 **4. Conclusions** 35

36 Although the rationale for Ag substitution of Cu in CZTSSe was to reduce antisite defect
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38 formation, the results here show that, in both the quenched and slow-cooled AZTSe, there is
39
40 direct evidence for substantial Ag-Zn intermixing between the kesterite 2c (Ag) and 2d (Zn) sites
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42 and some Zn on the 2a (Ag) sites. Direct elemental mapping of the slow-cooled sample by EDX
43
44 suggests that, over an area a few tens of nanometers across, there is complete intermixing of
45
46 Ag and Zn on the 2c and 2d sites. However, individual HAADF images from both the quenched
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48 and slow-cooled samples show at least partial ordering of Ag and Zn between the 2c and 2d
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50 columns at the nanometer level, in agreement with XRD results. We have shown that these
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3 results can be explained by a model where partially ordered twin-related domains exist in a
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5 more disordered matrix.
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9 The results therefore imply that our AZTSe has high densities, albeit non-uniformly
10 distributed, of Ag_{Zn} and Zn_{Ag} antisite defects, both at the growth temperature of 550°C, since
11 the quenched sample should largely freeze in the high temperature structure, and despite slow-
12 cooling which should maximise ordering. XRD suggests some evidence that slow-cooling
13 increases the degree of Ag-Zn ordering on the 2c and 2d sites, although TEM suggests a similar
14 domain structure is present in both the quenched and slow-cooled samples. The PL shows
15 emission close to the expected bandgap, i.e. minimal band tailing, despite the Ag-Zn disorder,
16 and that, for the slow-cooled sample, there is an emission peak at around 0.07eV below the
17 expected band edge, which could be attributed to $[\text{Ag}_{\text{Zn}} + \text{Zn}_{\text{Ag}}]$ antisite complex, although
18 further evidence to assign a definitive attribution should be considered.
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34 It is worth considering the implications of these results for understanding solar cells based
35 on CZTSSe and Ag-substituted CZTSSe absorbers. Our results show that, even after a slow-
36 cooling treatment, Ag substitution of Cu does not suppress extensive antisite defect formation,
37 either in AZTSe, and, by implication, also not in less Ag-substituted ACZTSe alloys. Thus the
38 progressive improvement in band tailing as the Ag content is increased is not due to a
39 progressive reduction in the overall antisite defect density. As shown by Gerschon et al¹², Ag-
40 substituted CZTSe films with Ag contents above about 50% substitution are not promising for
41 use as solar cell absorbers owing to an extremely low carrier density, which should lead to poor
42 junction formation. However, films with low Ag contents (5-10%) have much higher carrier
43 densities and a significant improvement in ACZTSe solar cell efficiencies compared with CZTSe
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3 cells has been demonstrated, consistent with the lower band tailing observed in the ACZTSe
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5 alloys.^{12,26} The lower formation energies for Cu_{Zn} and Zn_{Cu} antisite defects suggest that there
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7 are also high densities of antisites in CZTSSe, as confirmed experimentally by a variety of
8
9 techniques (see earlier). Neutron diffraction measurements also shows that slow-cooling of
10
11 CZTSSe samples reduces but doesn't eliminate the total number of Cu-Zn antisite defects,
12
13 although whether this reflects isotropic changes or the formation of ordered and disordered
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15 domains as in our case remains uncertain.^[27] Given therefore that both CZTSSe and AZTSe
16
17 contain substantial numbers of antisite defects, the generally greater band tailing observed in
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19 CZTSSe, up to 0.2-0.3eV below the band edge, needs explanation. In both CZTSSe and AZTSe,
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21 the Cu(Ag)-Zn antisite defect levels are similarly close to the band edge.^[25] This implies that
22
23 either there may be other active point defects present in CZTSSe that are not in AZTSe, or that
24
25 longer range extrinsic microstructural features play a role and need to be considered in
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27 materials preparation treatments. Although the primary emphasis of the work here has been
28
29 on Ag-Zn disorder, a preliminary analysis of microstructural defects showed that individual
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31 grains of AZTSe rarely contained either low- or high-angle grain boundaries, or dislocations.
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33 However, in previous work of CZTS, we have observed high densities of antisite domain
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35 boundaries,^[13] and recent work by Mendis et al^[28] have demonstrated similar results for
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37 CZTSe. Some of these antisite boundaries contain Sn antisite defects and can be charged. This
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39 should lead to local potential fluctuations and carrier trap states, both of which could
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41 contribute to band-tailing. The AZTSe samples here were found to contain antisite domain
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43 boundaries, but at densities less than 10⁸ cm⁻², a factor of at least 10² times lower than we
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45 reported for CZTS.^[13] Alternatively, theoretical studies have suggested that Cu and Zn clustering
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3 is energetically favourable in CZTSSe, leading to local bandgap variations.^[7,29] There is also
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5 strong experimental evidence for defect clustering in CZTSSe, for example from Raman and
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7 nuclear magnetic resonance (NMR) studies,^[30] and from EELS studies, which suggest
8
9 segregation into Cu-rich and Zn-rich regions over the 5-20nm range as noted earlier.^[5,6] On the
10
11 basis of their similar electronic structures, we would expect clustering to be favoured in both
12
13 CZTSSe and AZTSe. However, we have found no evidence from EDX mapping for long range
14
15 fluctuations in the Ag and Zn concentrations. It is possible that the size difference between the
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17 Ag^+ and Zn^{2+} ions inhibits long-range Ag-Zn segregation in AZTSe compared to Cu-Zn
18
19 segregation in CZTSSe. This could occur because Ag-rich and Zn-rich regions would be regions of
20
21 compressional and dilatational strain respectively, leading to an increase in strain energy.
22
23 However, more work is needed to clarify whether this is a significant energy contribution.
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32
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38
39 EP/K040375/1 and EP/L017792/1 for access to the JEOL ARM200CF (S)TEM (South of England
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41 Analytical Electron Microscope)
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46 ASSOCIATED CONTENT Supporting Information Available: This shows more detail on the
47
48 processing steps used in converting raw HAADF and EDX data into the final elemental maps in
49
50
51 Figure 4
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